Searches for Ultra Long-Lived Particles with

MATHUSLA

CAP CONGRESS

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MIRIAM DIAMOND

ASSISTANT PROFESSOR

Arthur B. McDonald Canadian Astroparticle Physics Research Institute

UNIVERSITY OF TORONTO

Outline

- Basic Concept
  - Backgrounds
  - Identifying LLPs
- LLP Sensitivity
- Cosmic Ray Telescope
- Detector Design

An Update to the Letter of Intent for MATHUSLA: Search for Long-Lived Particles at the HL-LHC (arXiv:2009.01693)

Basic Concept

MAssive **T**iming **H**odoscope for **U**ltra-**S**table Neutra**L** P**A**rticles

LLPs at the [HL-]LHC

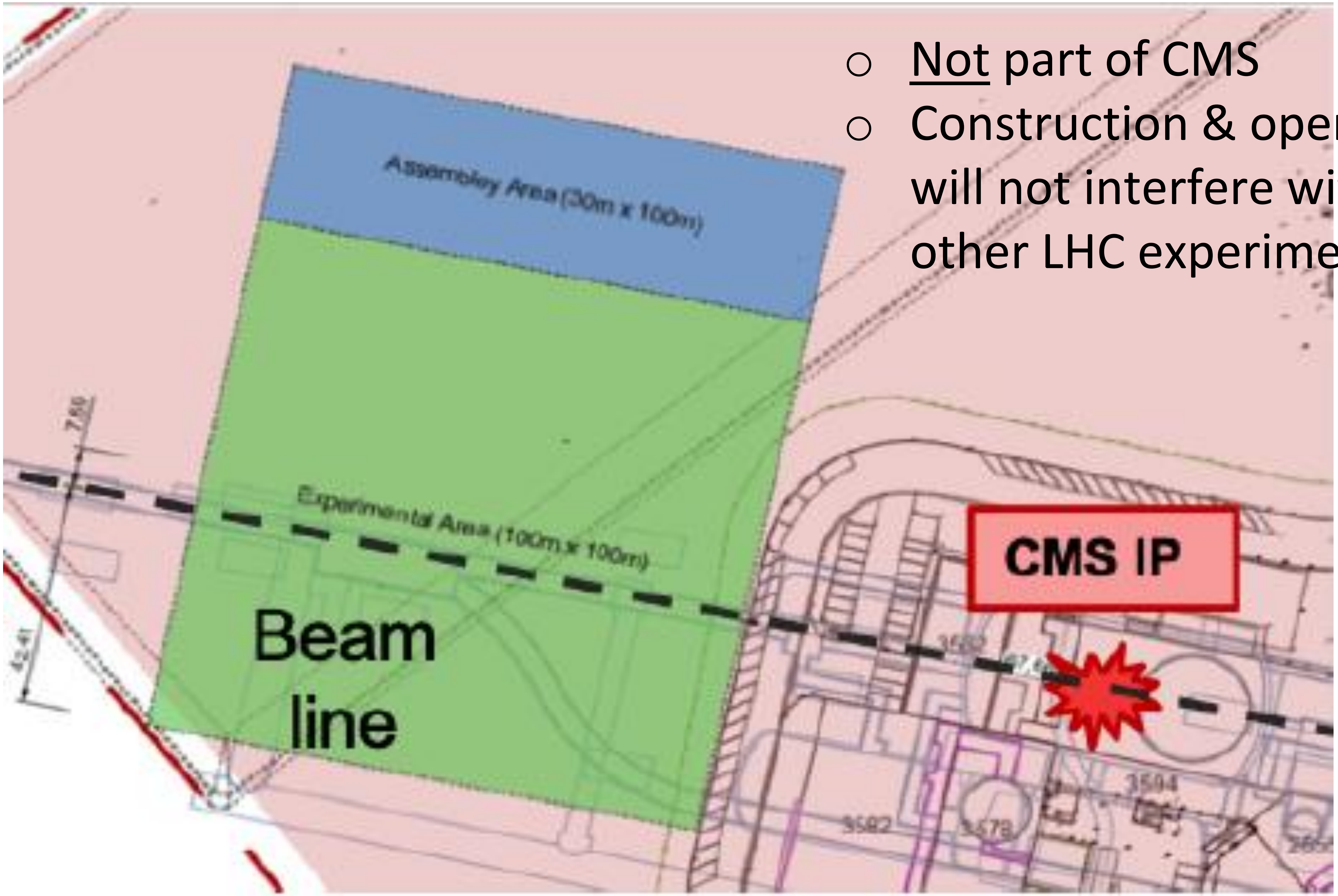
Seeking to go Beyond the Standard Model (BSM) motivates the possibility of so-far-undiscovered LLPs

- "**Top-down**": Various BSM theories (e.g. supersymmetry) constructed to explain the "fundamental mysteries" naturally include new LLPs
- "**Bottom-up**": LLPs occur in the SM (e.g. muons), and can occur via similar mechanisms when adding new particles to the model

The problem of long lifetimes: LHC could be making LLPs that are invisible to its main detectors!

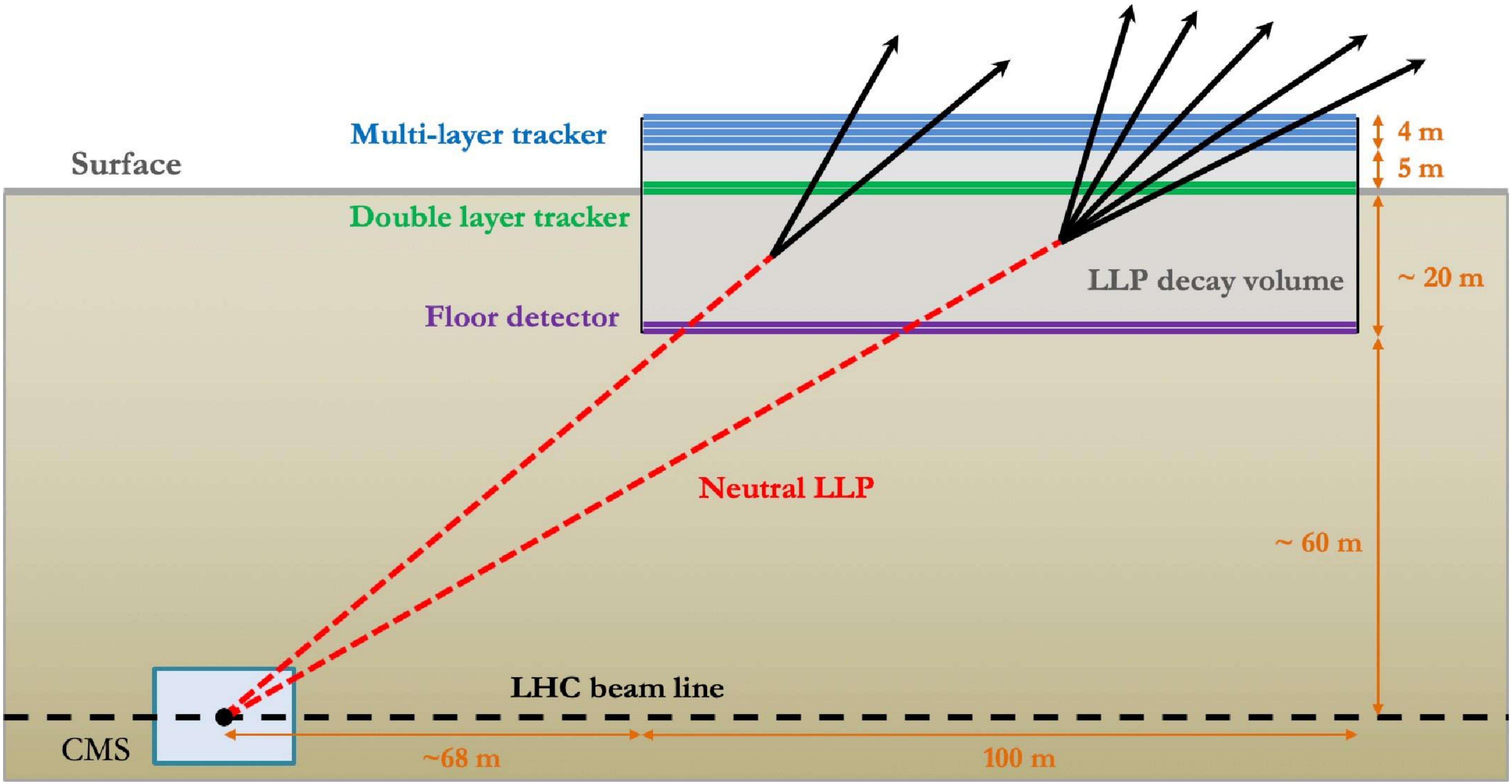
- **If the LLP has c · lifetime >> detector size, most escape the detector**
- **Even LLPs that decay in the detector, but a significant distance away from the Interaction Point, are difficult to spot**
- **If the LLPs decay in the detector with only a tiny rate, they get swamped by backgrounds**

An External LLP Detector for HL-LHC



- <u>Not</u> part of CMS
- Construction & operation will not interfere with any other LHC experiments

An External LLP Detector for HL-LHC

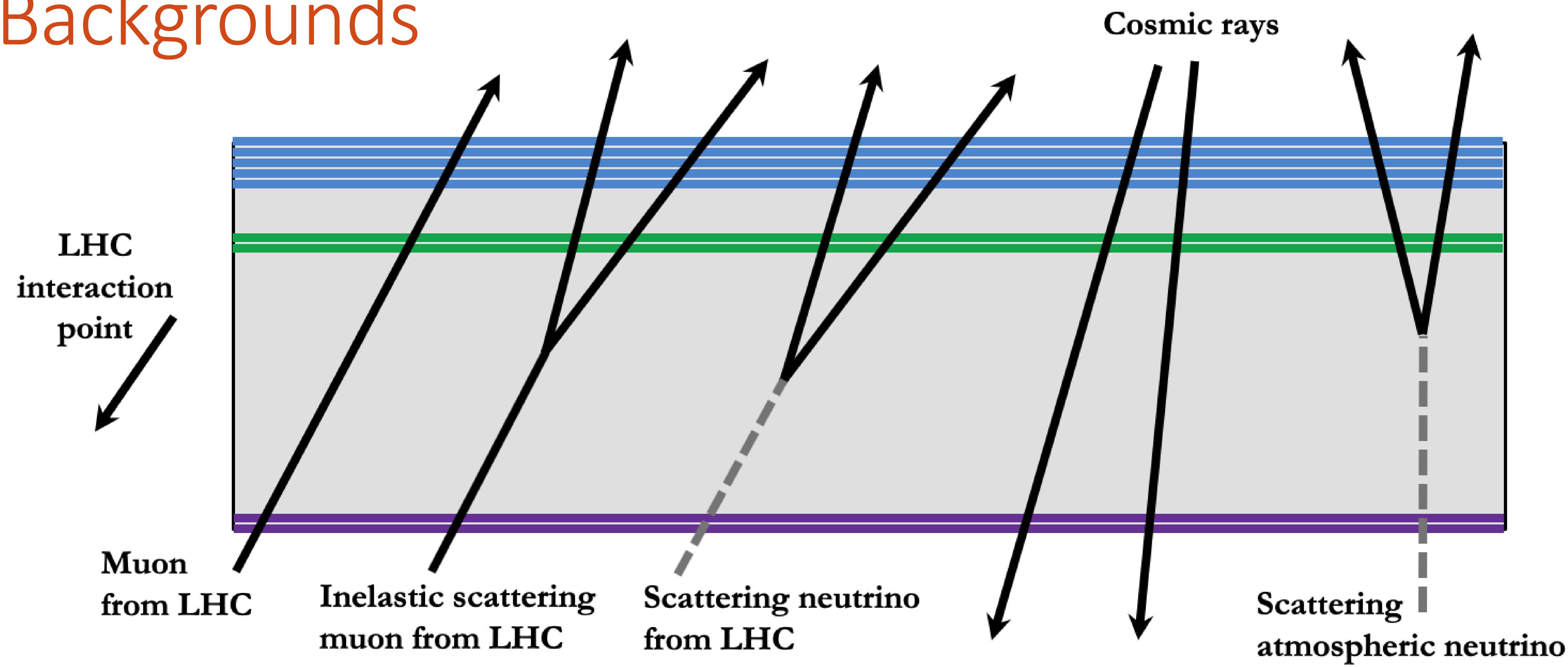


NOT TO SCALE

100m x 100m x 25m decay volume

Displacement from IP: 70m horizontally, 60m vertically

Backgrounds



LLP displaced vertex (DV) signal has to satisfy many stringent geometrical and timing requirements ("4D vertexing" with cm/ns precision)

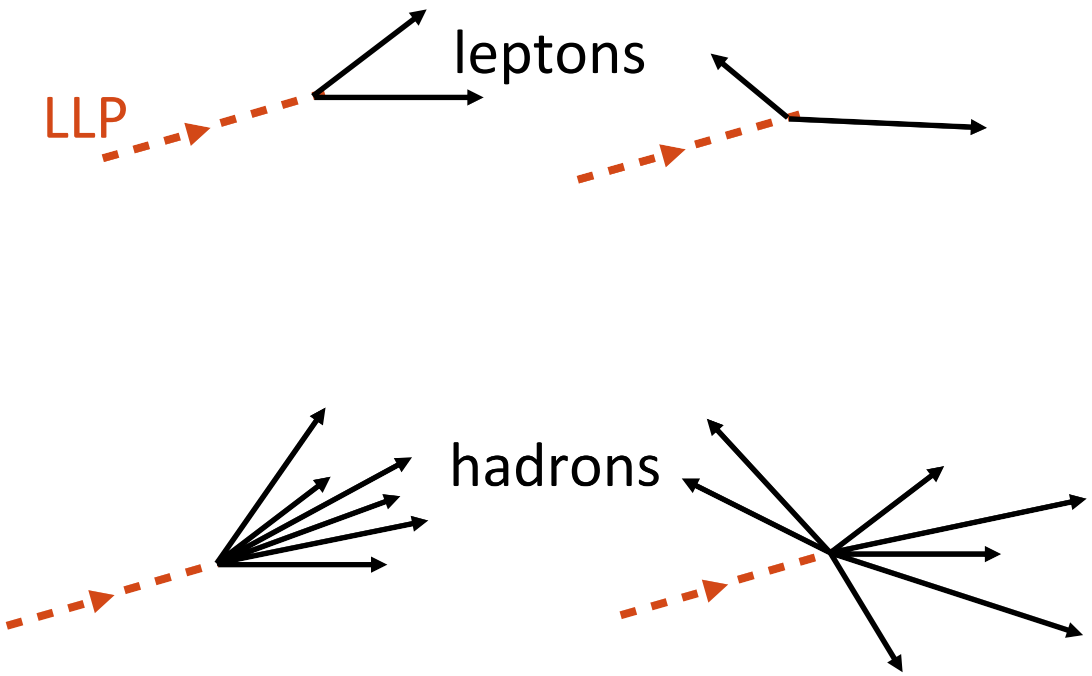
These requirements, plus a few extra geometry & timing cuts, provide "near-zero background" (< 1 event per year) for neutral LLP decays!

Backgrounds

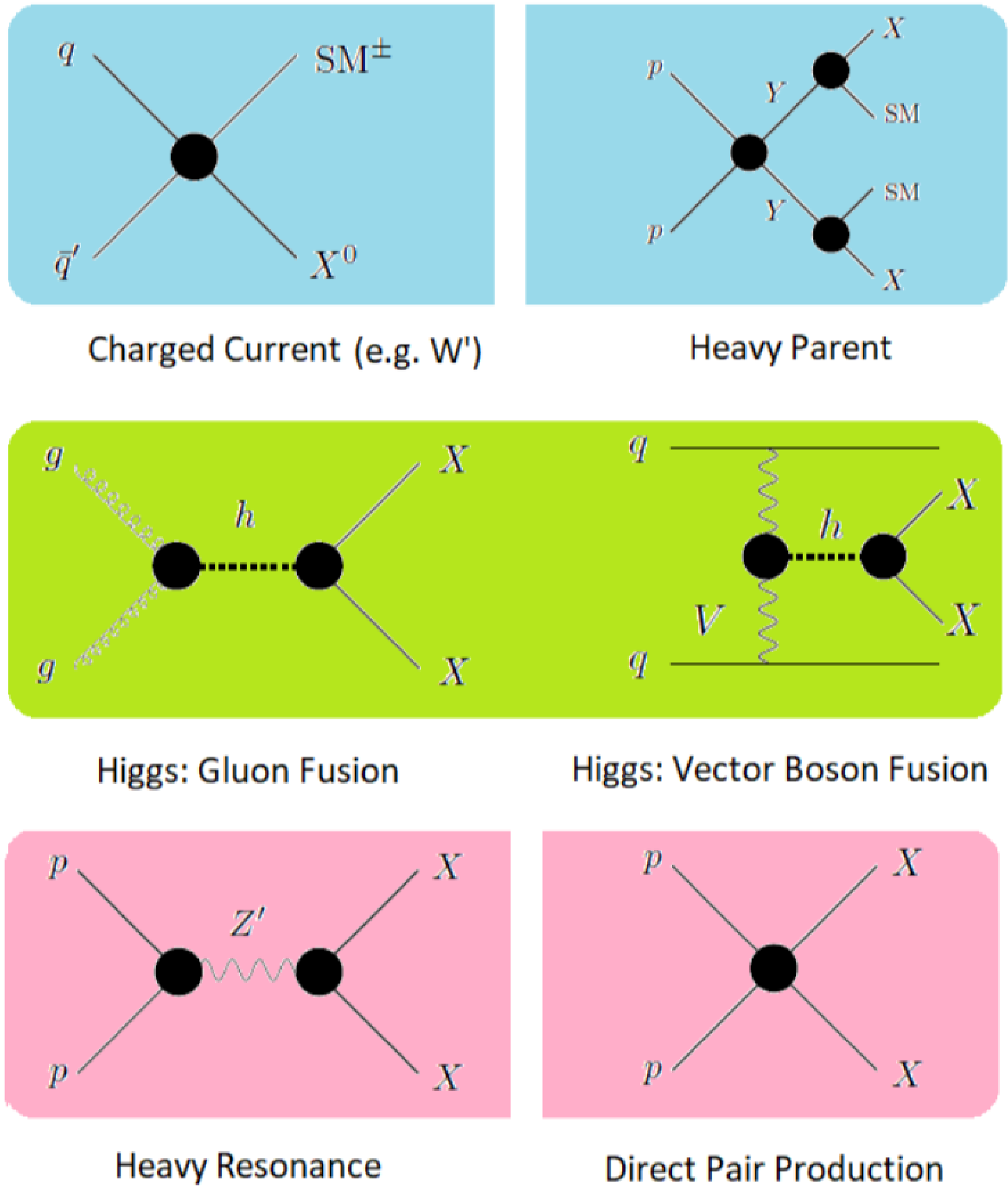
- Cosmic rays
  - Calibrations performed using Test Stand measurements (taken above ATLAS IP in 2018) arXiv: 2005.02018
  - Downward-going events ~3 x 10^{14} over entire HL-LHC run, distinguished from LLPs using timing cuts
  - Upward-going events ~2 x 10^{10} : inelastic backscatter from CRs hitting the floor, or decay of stopped muons in floor. Only tiny fraction (estimates underway) produce fake DV, via decay to 3 charged tracks
  - Rare production of K^0_L harder to estimate; work underway on veto strategies
- Rare decays of muons originating from HL-LHC collisions
  - Upward-going events ~2 x 10^8 , mostly from W and bbar production
  - Work underway for optimal rejection strategies
- Charged particles from neutrino scattering in decay volume
  - Neutrinos from HL-LHC collisions << 1 "fake" DV/year
  - Atmospheric neutrinos ~30 "fake" DV/year, reduced to < 1 with cuts

Identifying LLPs

MATHUSLA can't measure particle momentum or energy, but:
track geometry → measure of LLP boost event-by-event



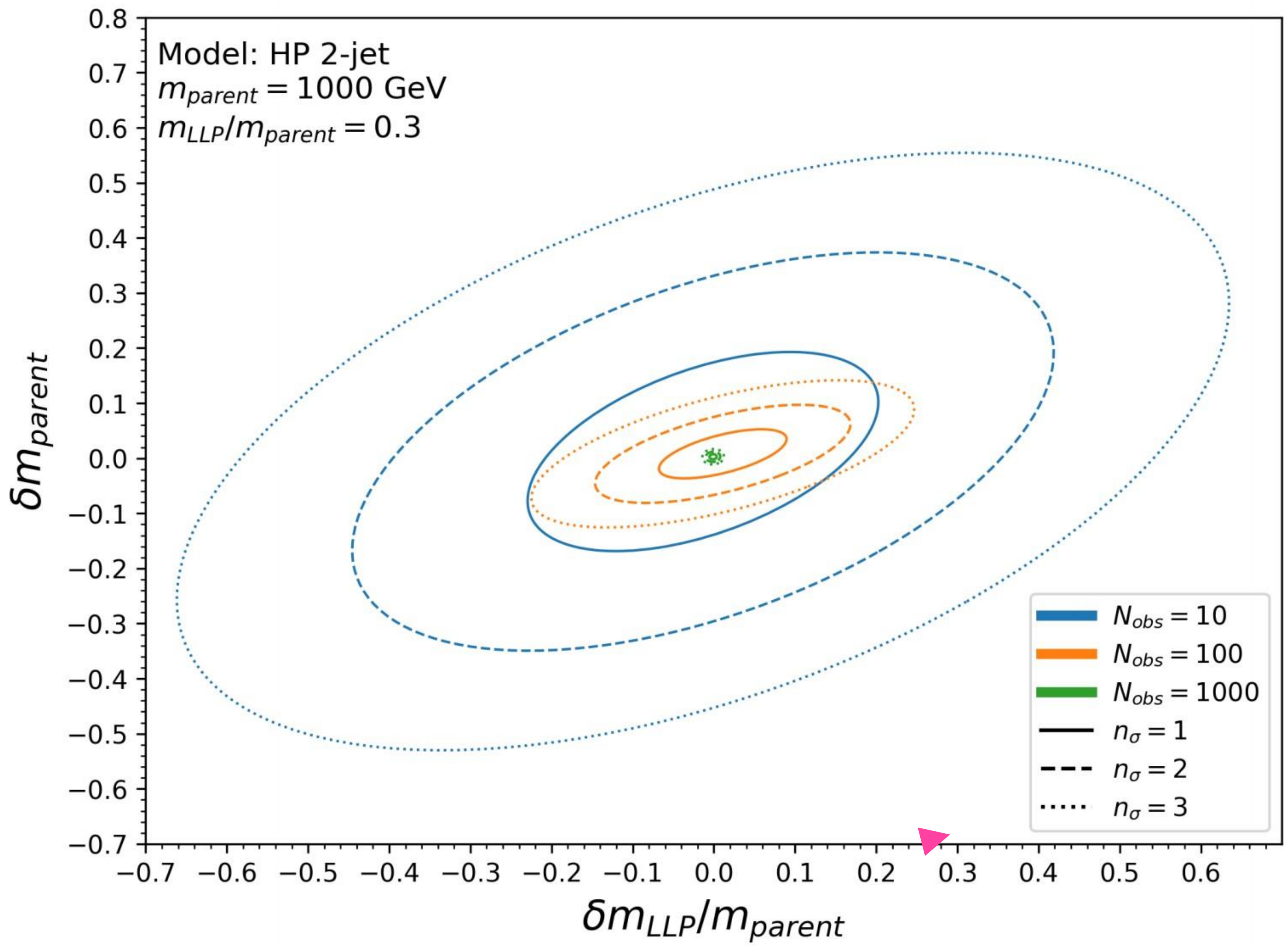
Incorporate MATHUSLA into CMS L1 Trigger
Correlate event info off-line → determine LLP production mode



Identifying LLPs

If production mode is known: Boost distribution → LLP mass
If LLP mass is known: Track multiplicity → LLP decay mode

MATHUSLA + CMS analysis will reveal model parameters (parent mass, LLP mass) with just ~ 100 observed LLP events!



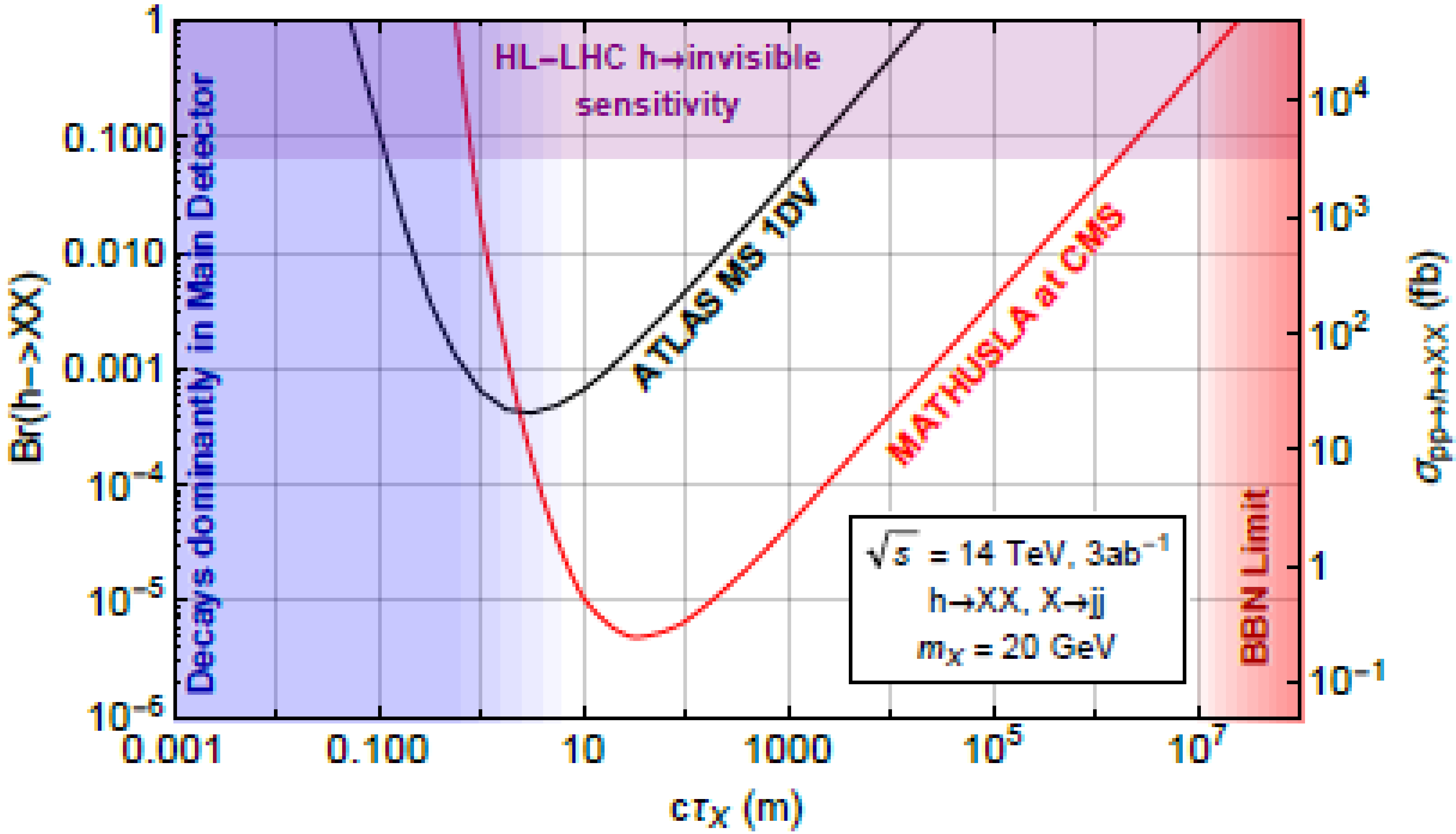
LLP Sensitivity

More benchmark models can be found in **Physics Beyond Colliders at CERN: Beyond the Standard Model Working Group Report** arXiv:1901.09966

LLP Sensitivity: Weak- to TeV- Scale

Up to 1000x better sensitivity than LHC main detectors

e.g. hadronically-decaying LLPs in exotic Higgs decay



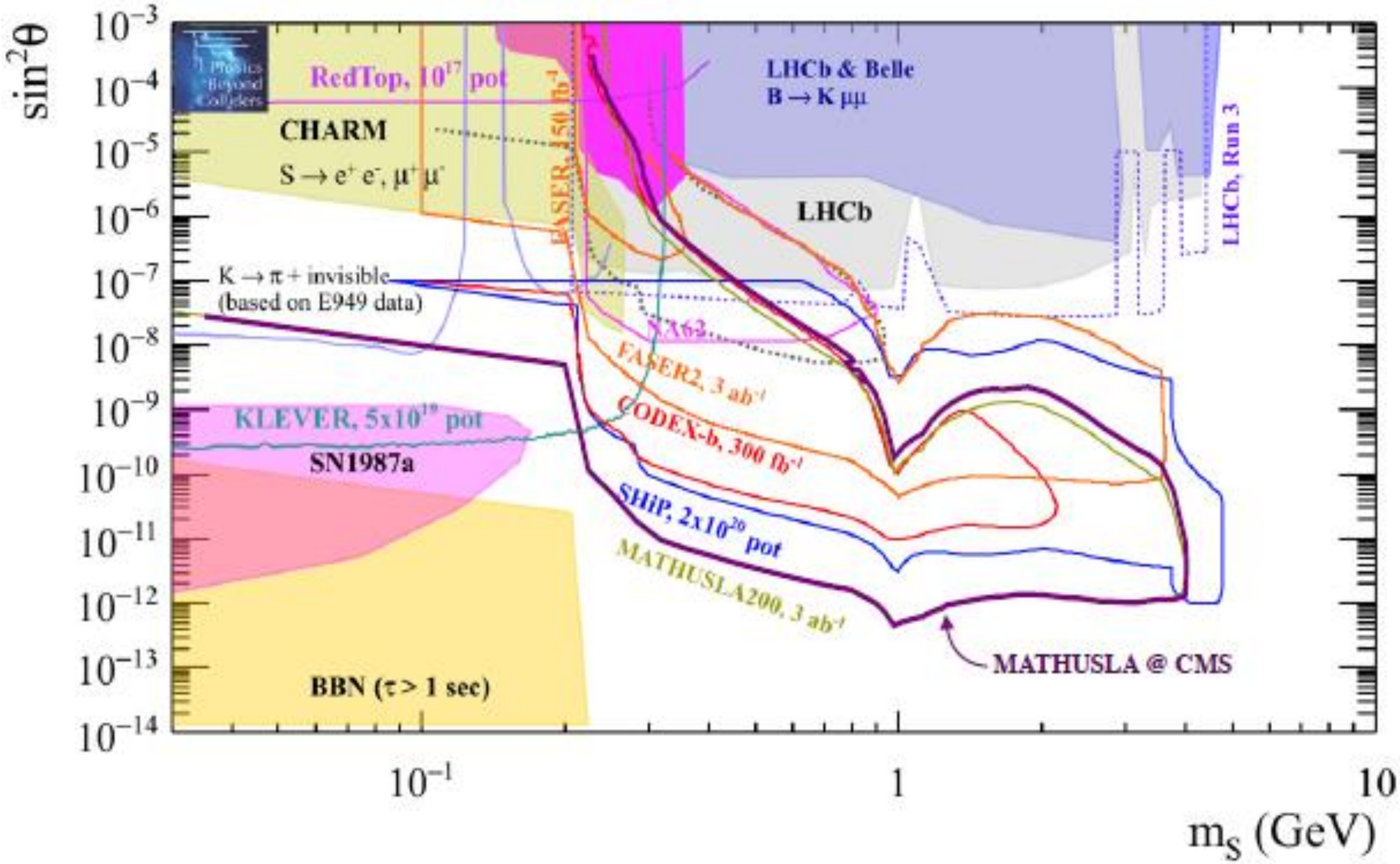
Any LLP production process with σ > fb can give signal in MATHUSLA

arXiv:2001.04750

LLP Sensitivity: GeV-Scale

For scenarios where the long-lifetime limit (>100m) is accessible, MATHUSLA is complementary to other planned experiments

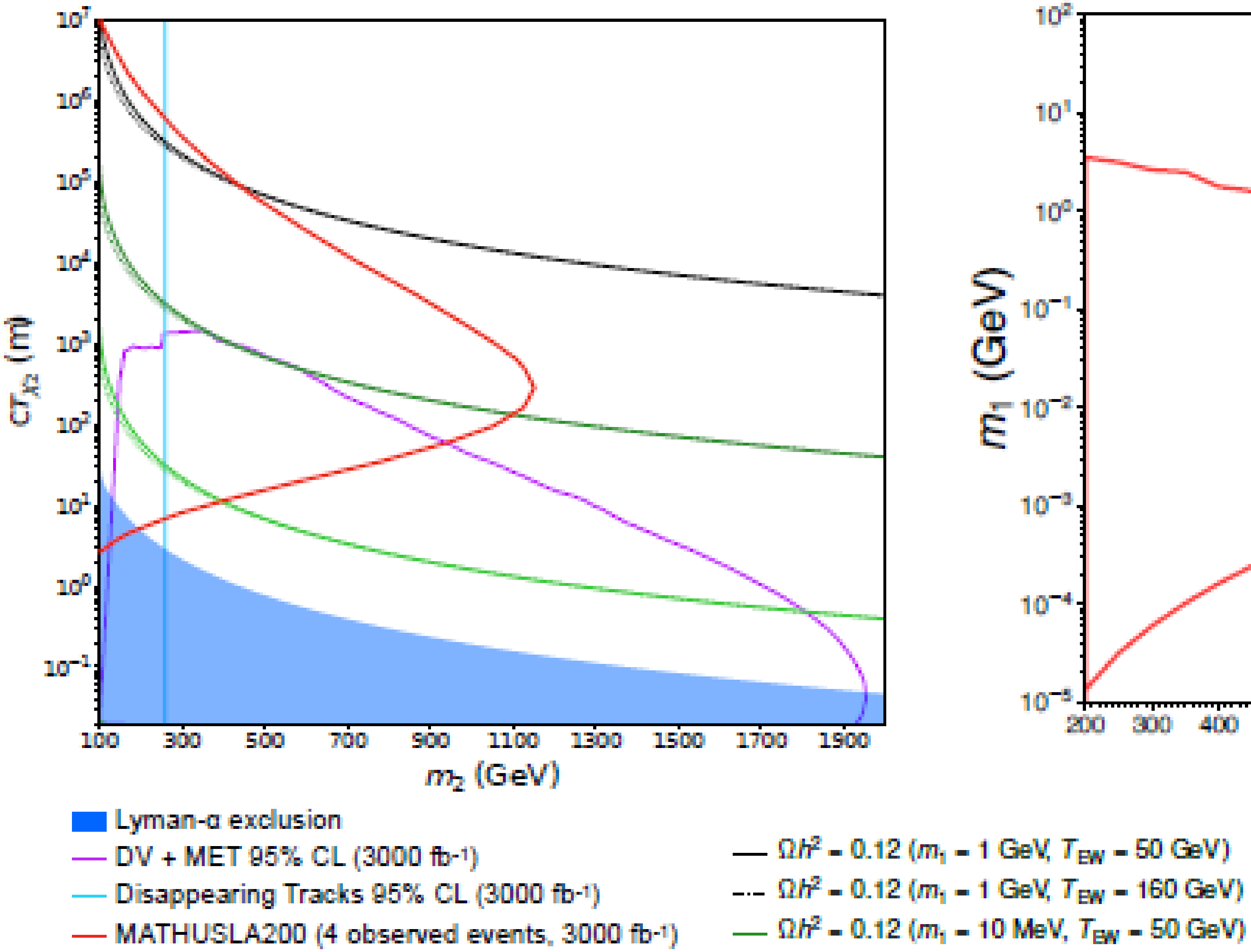
e.g. singlet dark scalar S, mixing angle θ with SM Higgs

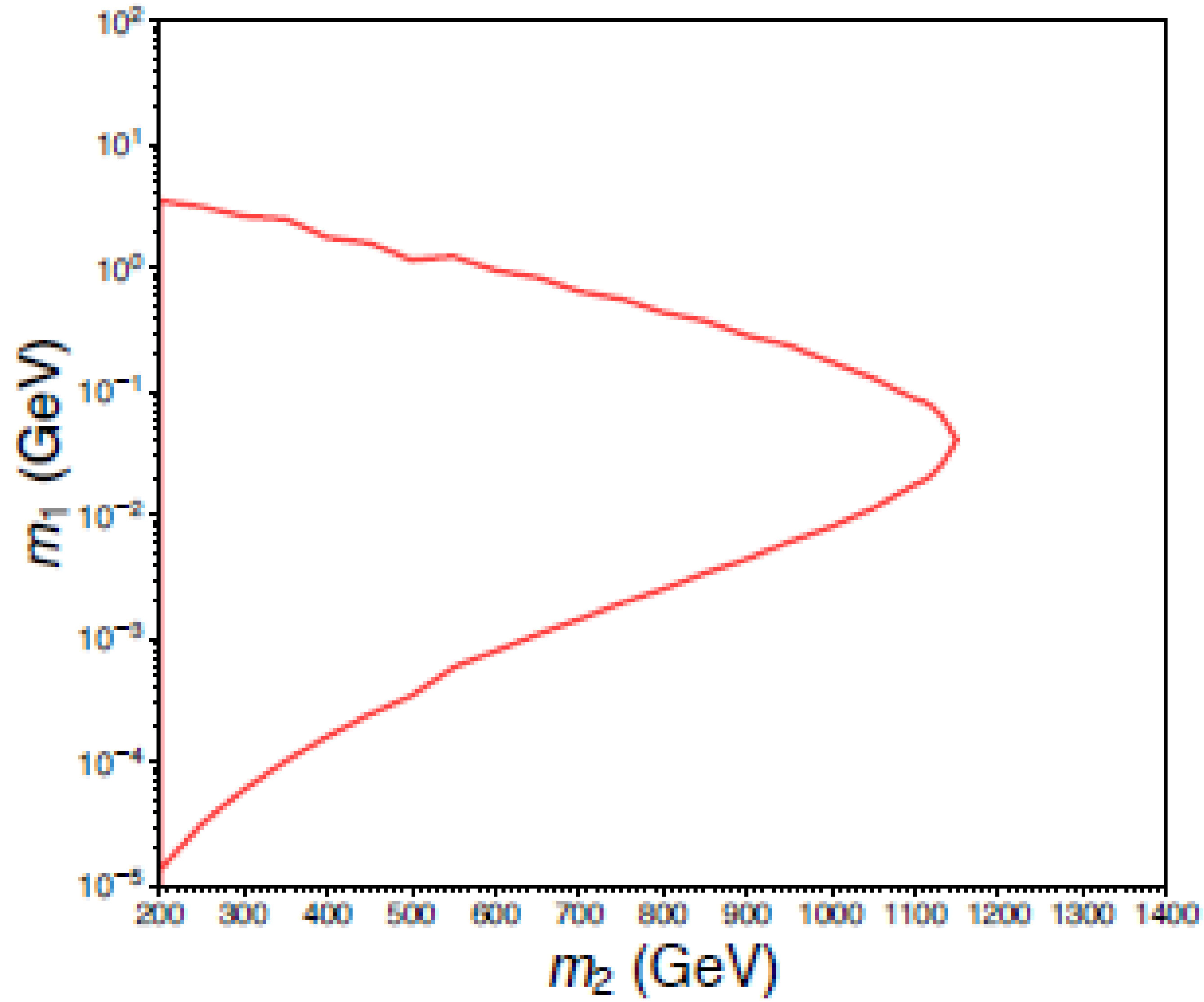


LLP Sensitivity: DM

Scenarios where LLP → DM + SM decay is the only way to see the DM

e.g. Freeze-In Dark Matter: BSM mass eigenstates χ_1 (DM) and χ_2 (LLP), where χ_2 was in thermal equilibrium with primordial plasma

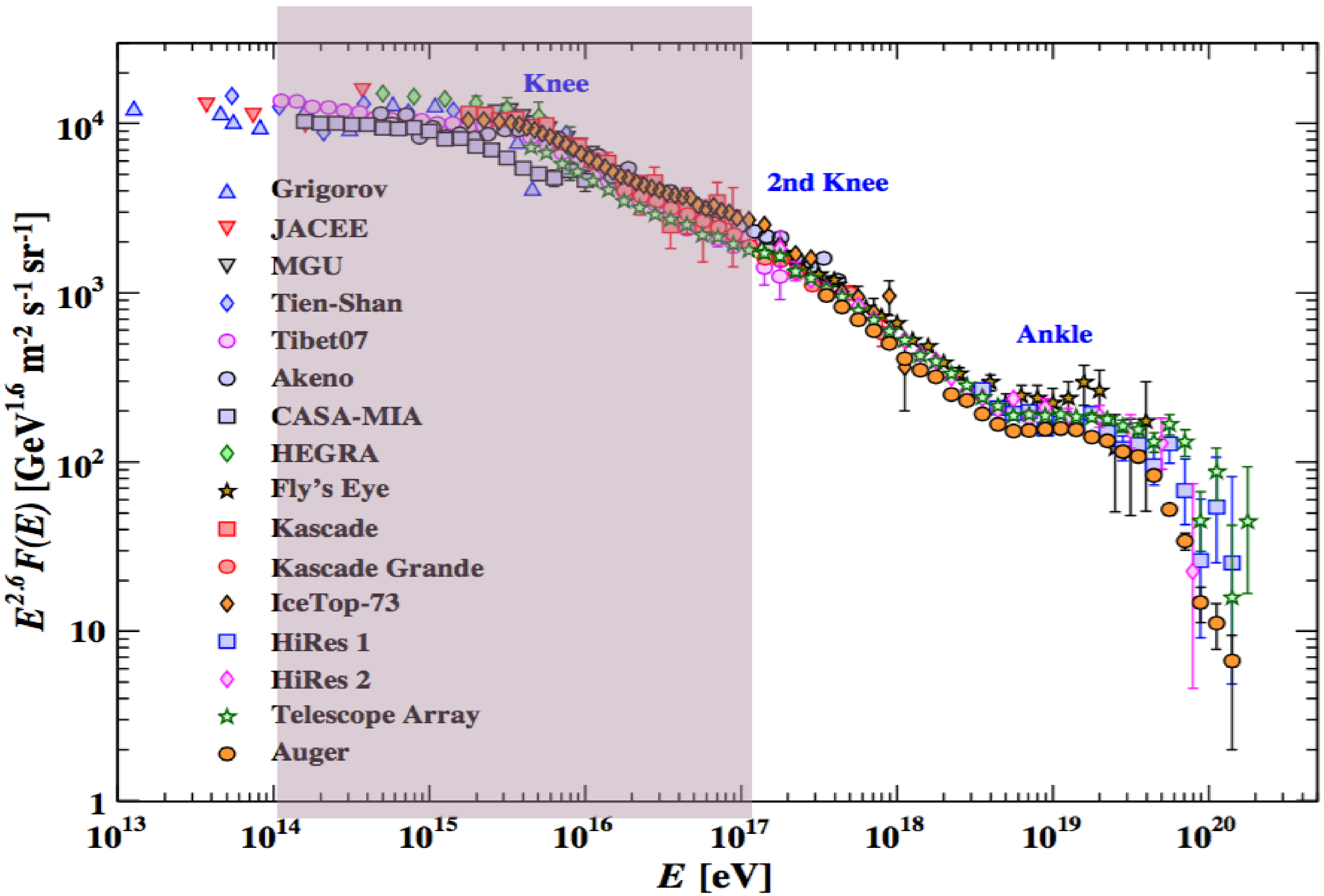




Cosmic Ray Telescope

MATHUSLA as a Cosmic Ray Telescope

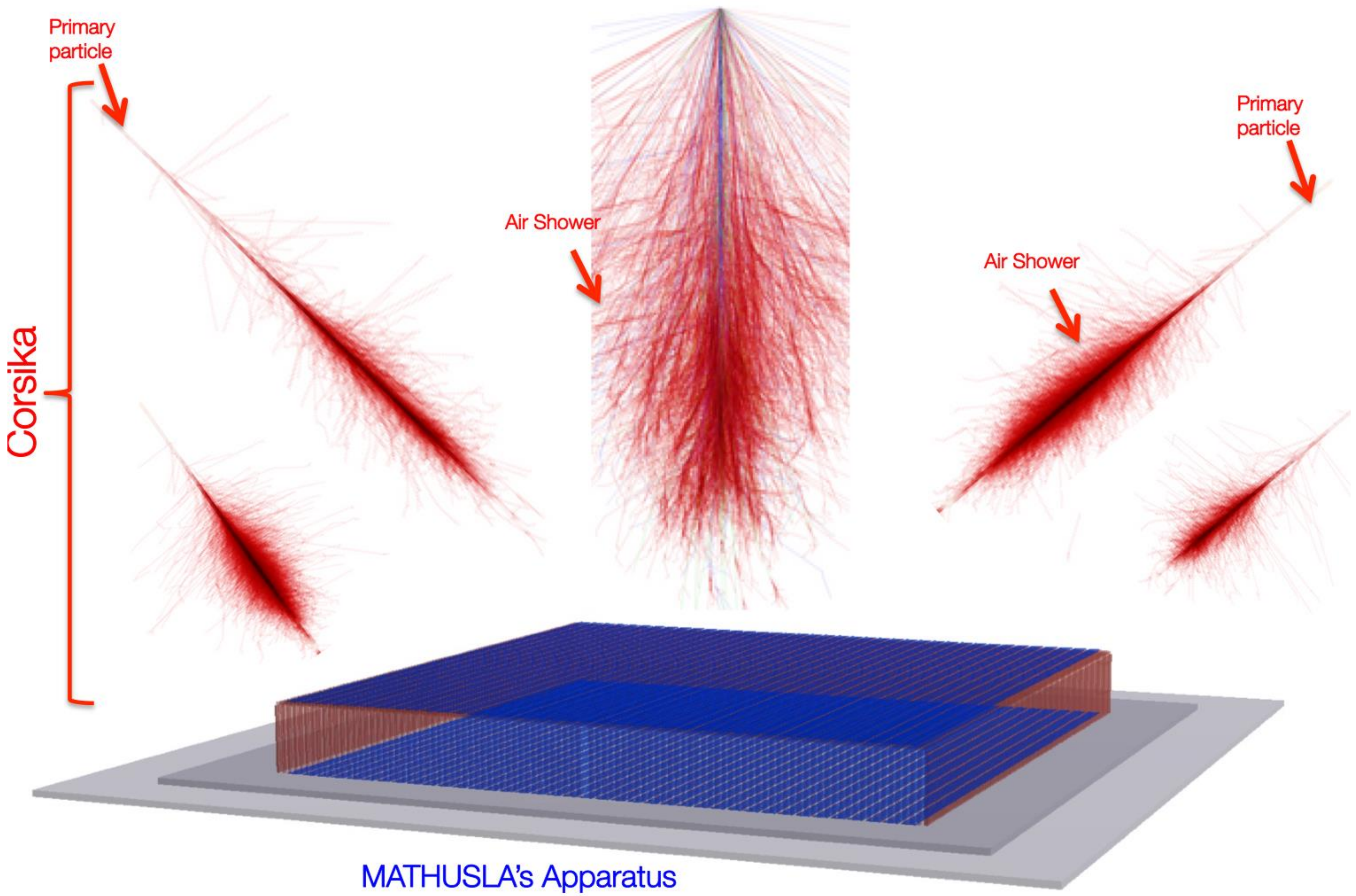
Unique abilities in CR experimental ecosystem (precise resolution, directionality, large-area coverage, interesting region CR energy spectrum)



Paper describing potential contributions to CR physics nearly completed, led by the Mexico MATHUSLA team

MATHUSLA as a Cosmic Ray Telescope

Reconstruction of shower core, direction, total # charged particles, slope of radial particle density distribution

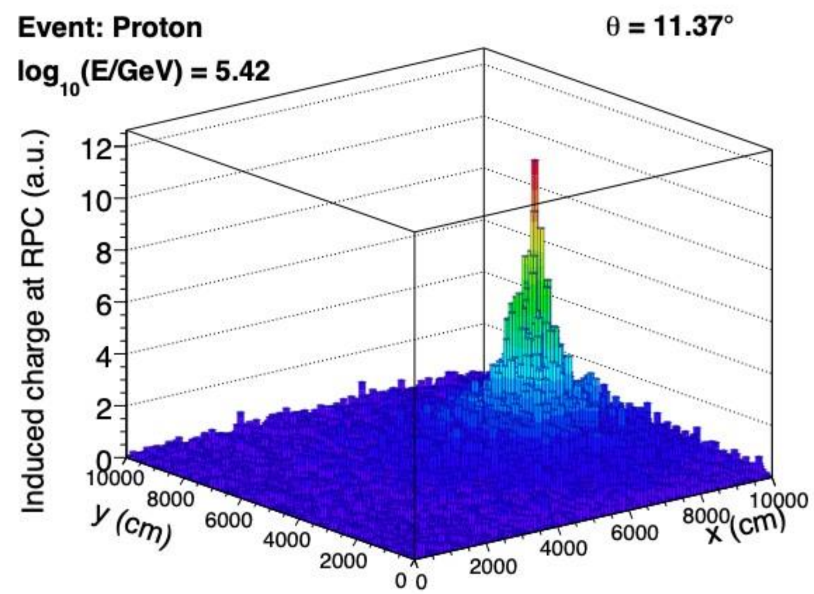


MC simulations using CORSIKA (https://www.iap.kit.edu/corsika/)

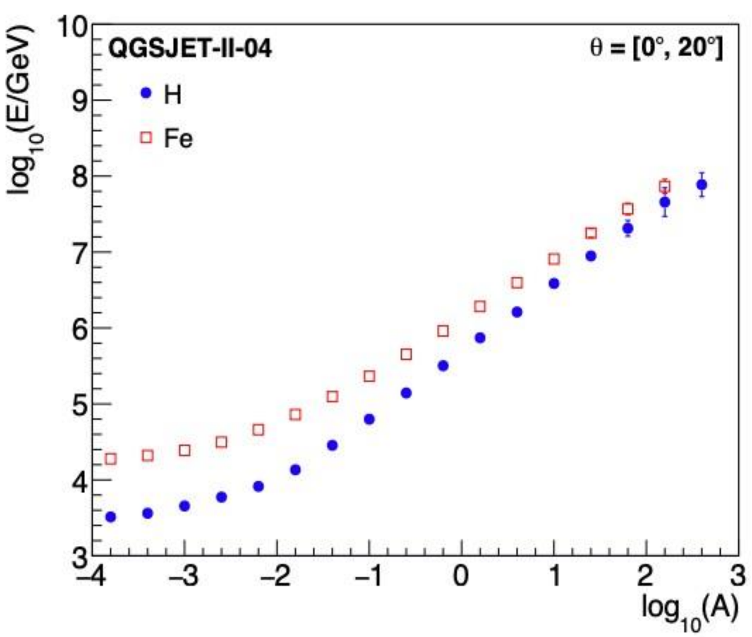
MATHUSLA as a Cosmic Ray Telescope

CR physics reach would be greatly enhanced by adding an analog RPC layer, due to scintillator saturation effects

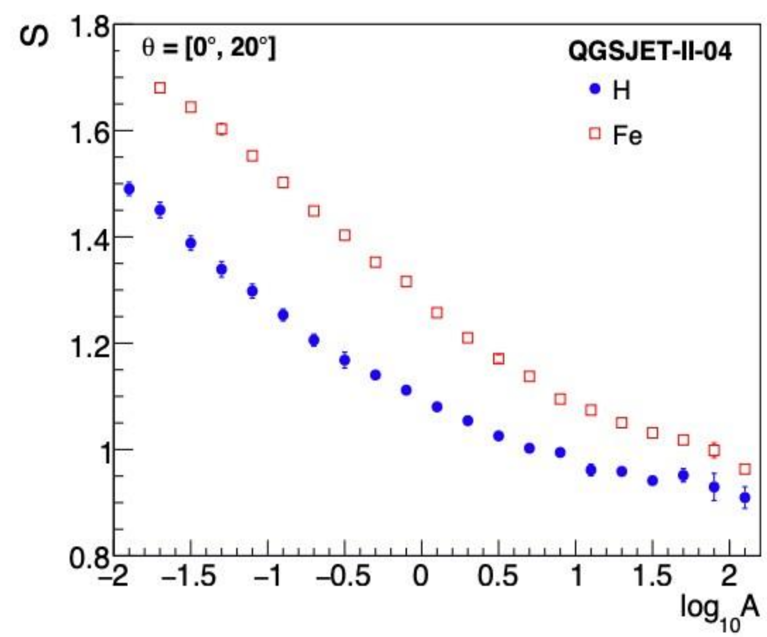
Charge density at the RPC



Amplitude of lateral distribution (LD)



Shower age (slope of LD) vs amplitude



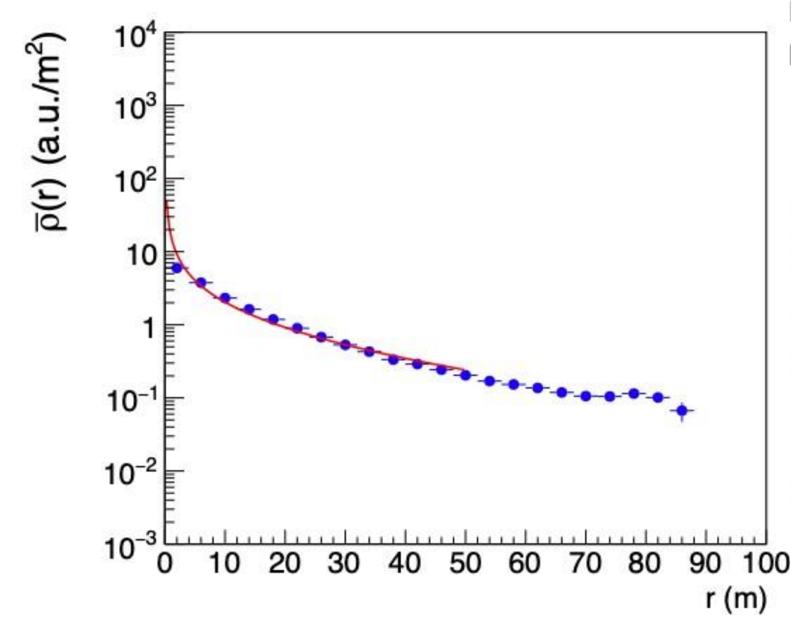
- In region of maximum efficiency linear dependence of logE with logA.

—> **It could provide energy scale**

- Shower age shows sensitivity to primary composition.

—> **Useful for composition studies**

Lateral charge density at RPC



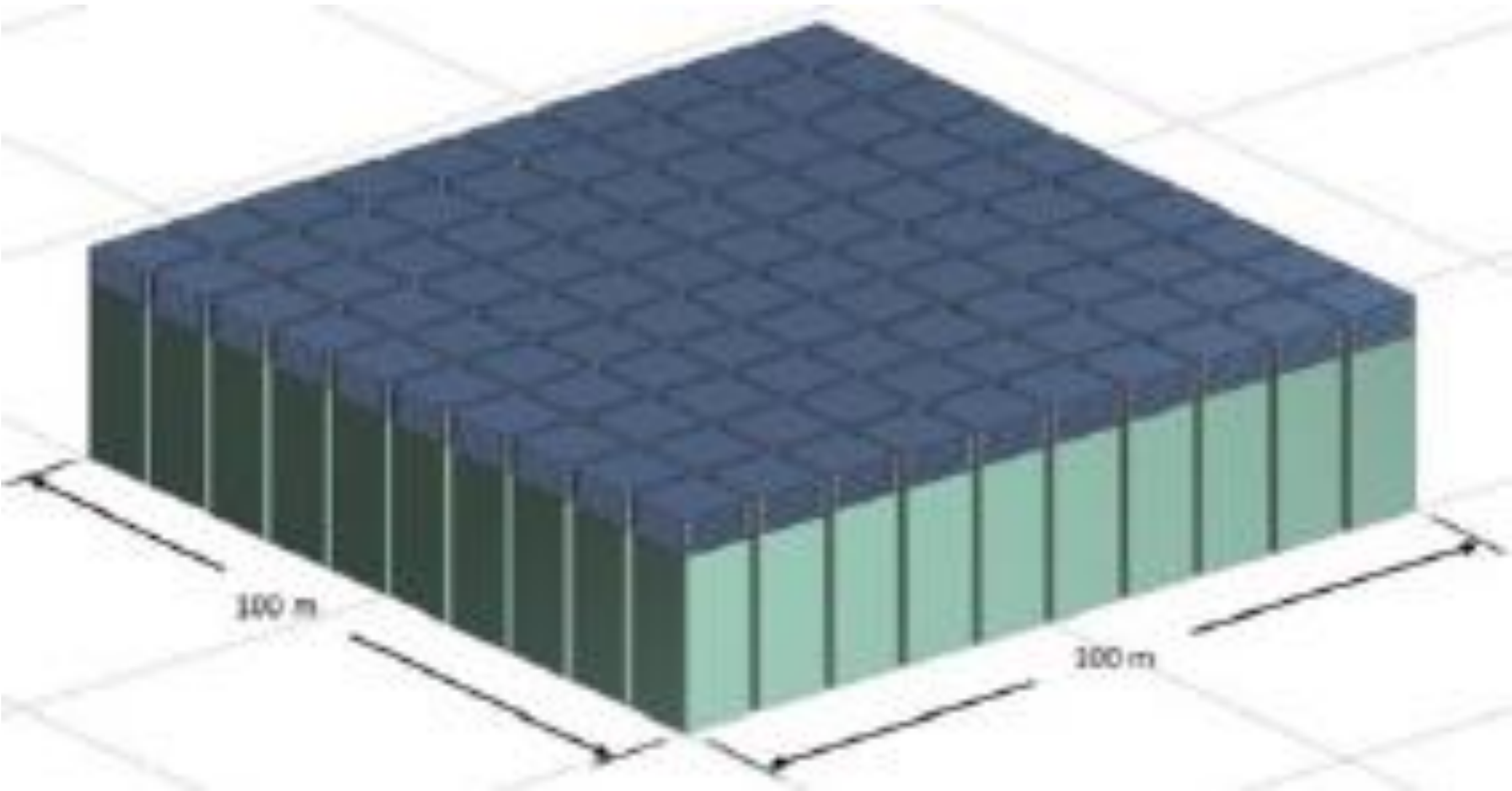
- RPC allows to **extend CR** energy and composition **studies above E = 10^{15} eV.**

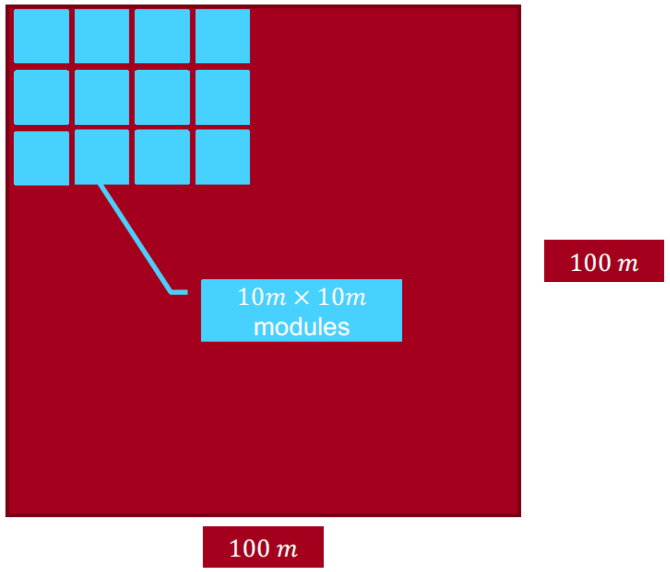
https://indico.cern.ch/event/980853/contributions/4361206/attachments/2251261/3819144/CRMathusla_LLP_25May2021_JC.pdf

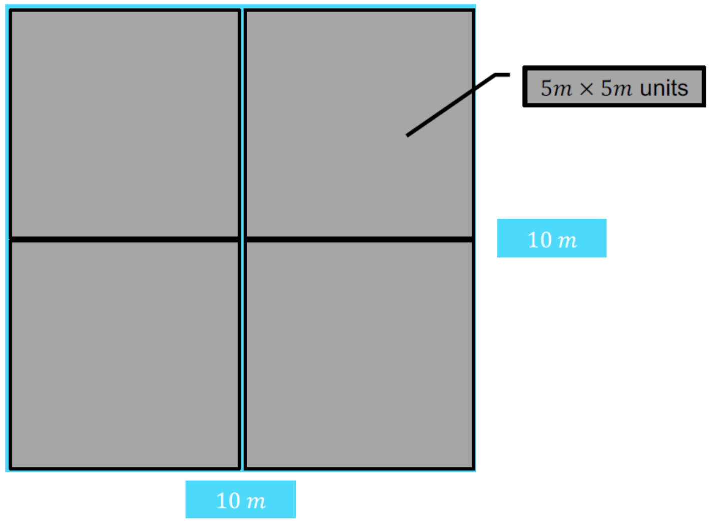
Detector Design

Detector Design

Modular design facilitates staged construction and commissioning



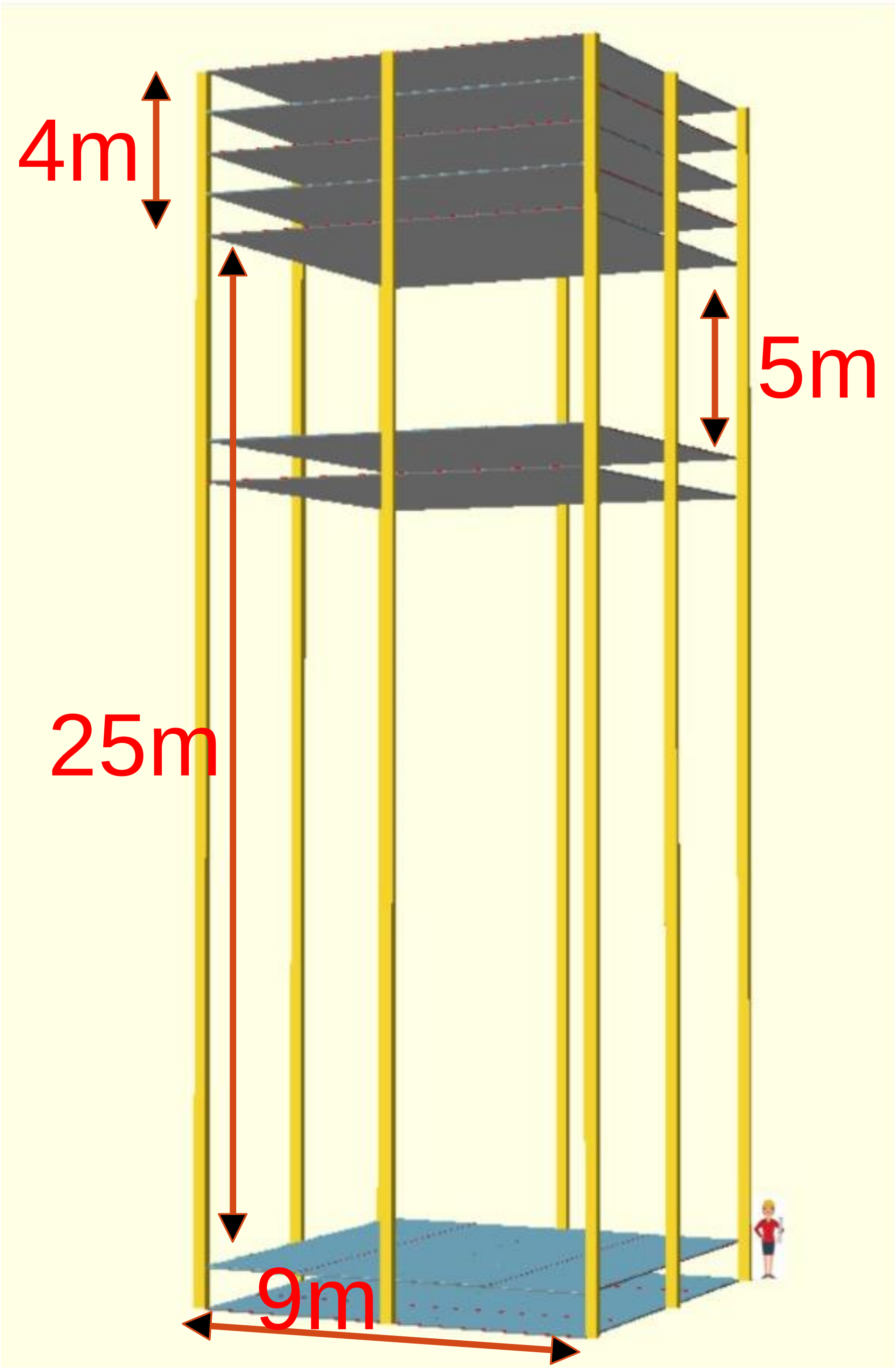




100 Modules in $\mathbf{100}\boldsymbol{m} \times \mathbf{100}\boldsymbol{m}$ Footprint

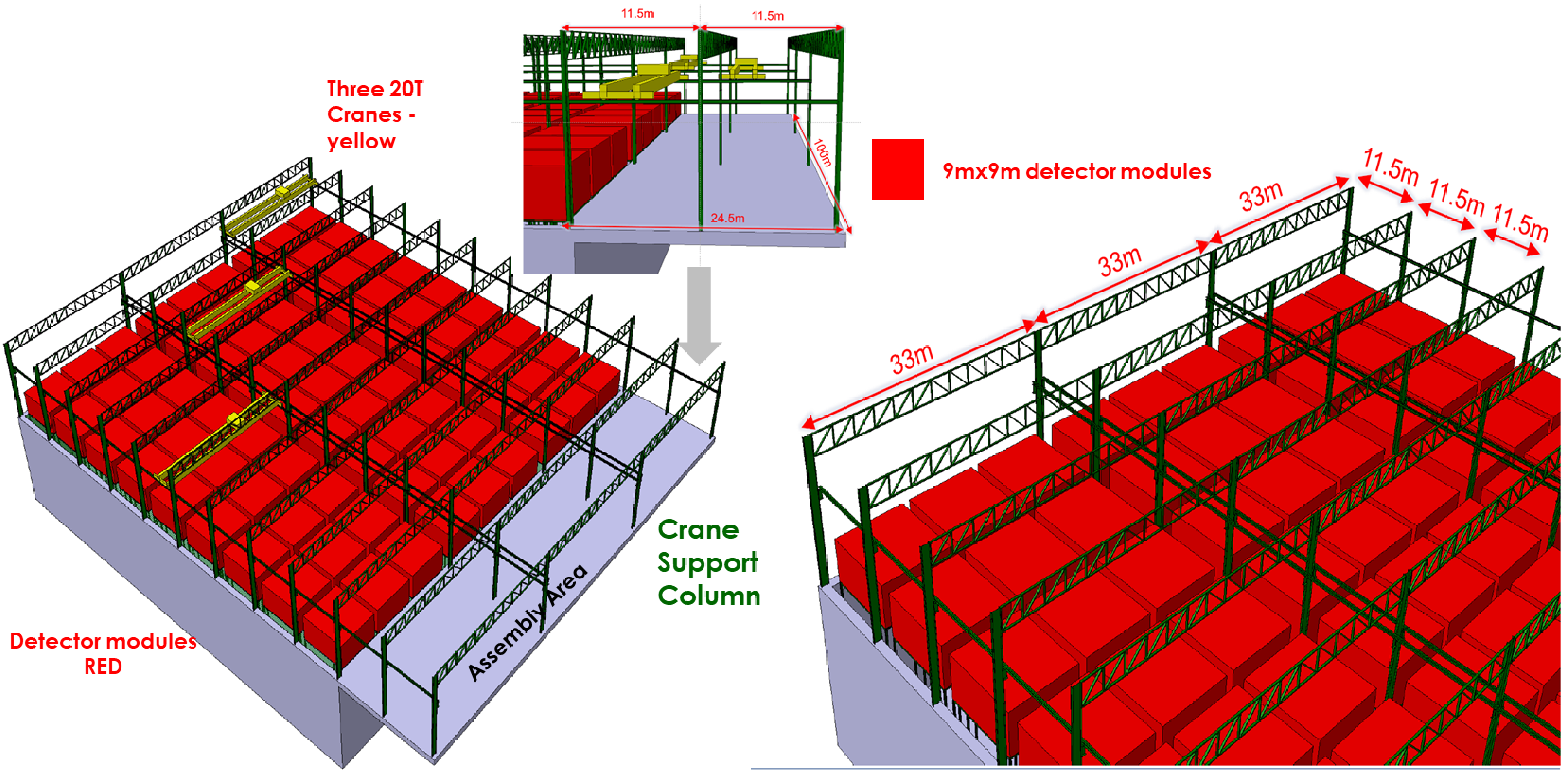
4 Detector Units per Module Plane

Detector Design

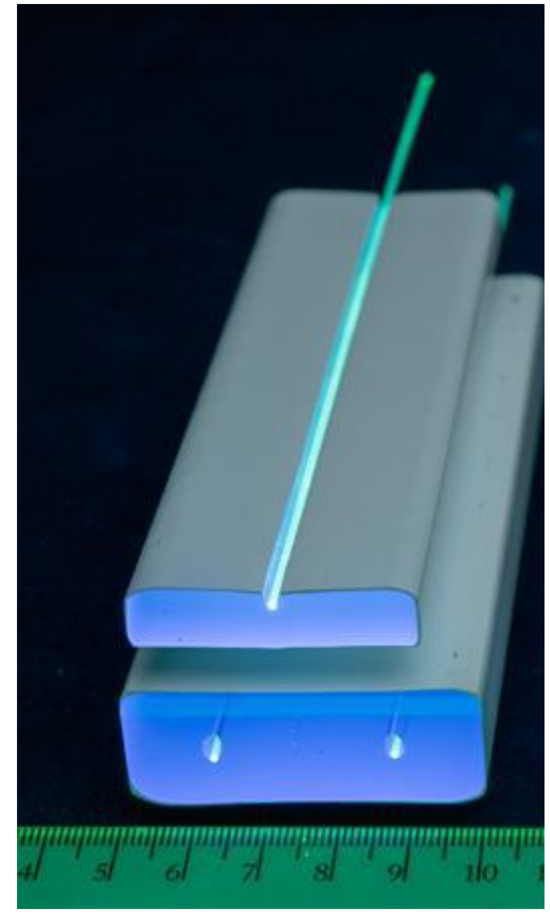


Each module has:
5 tracking layers on top
+ 2 floor layers
+ 2 mid-level layers

Detector Design



Trackers



Tracker layers: Composed of extruded scintillator bars with wavelength-shifting fibers coupled to Silicon Photo Multipliers

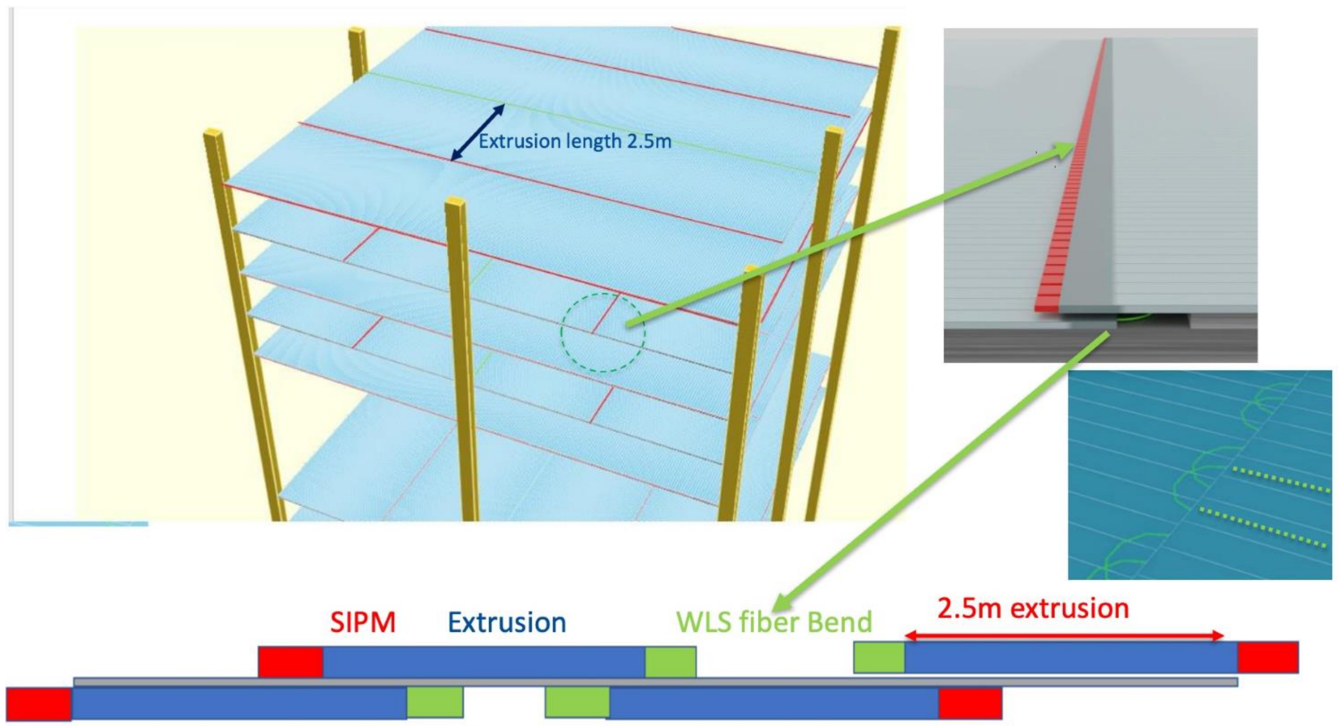
- Extrusion facilities in FNAL used for several experiments (e.g. Belle muon trigger upgrade, Mu2e)
- Possibility of adding Resistive Plate Chamber layers

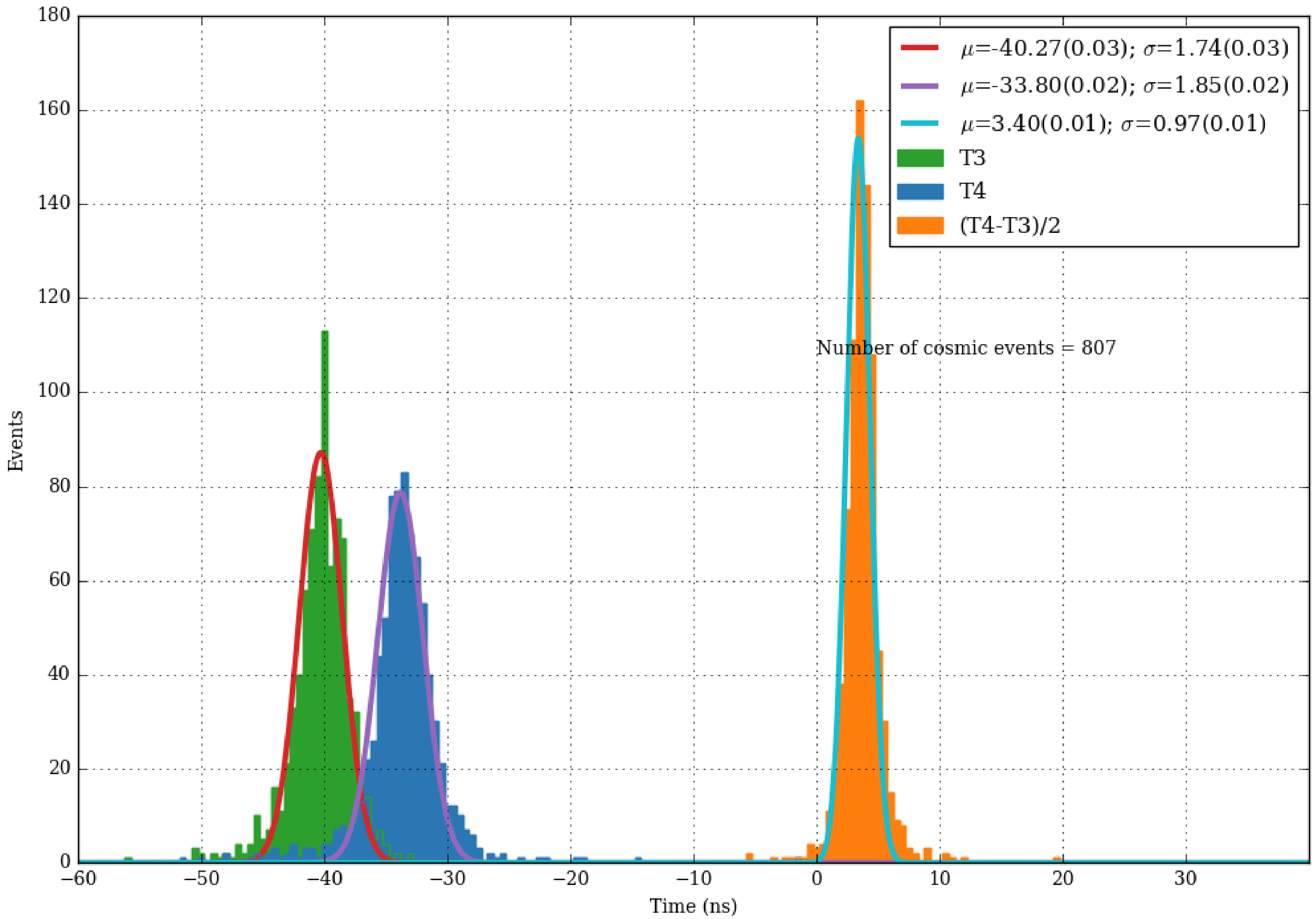
Each scintillator bar ~ 4.5m x 5cm x 2cm with readout at both ends

- Or 2.5m with looped fiber for readout at one end
- Transverse resolution $\sigma \approx 1$ cm
- Δt between two ends gives longitudinal resolution: need sub-ns precision

Trackers

~1ns timing resolution of cosmic ray hits recently achieved in ~5m bar test setup

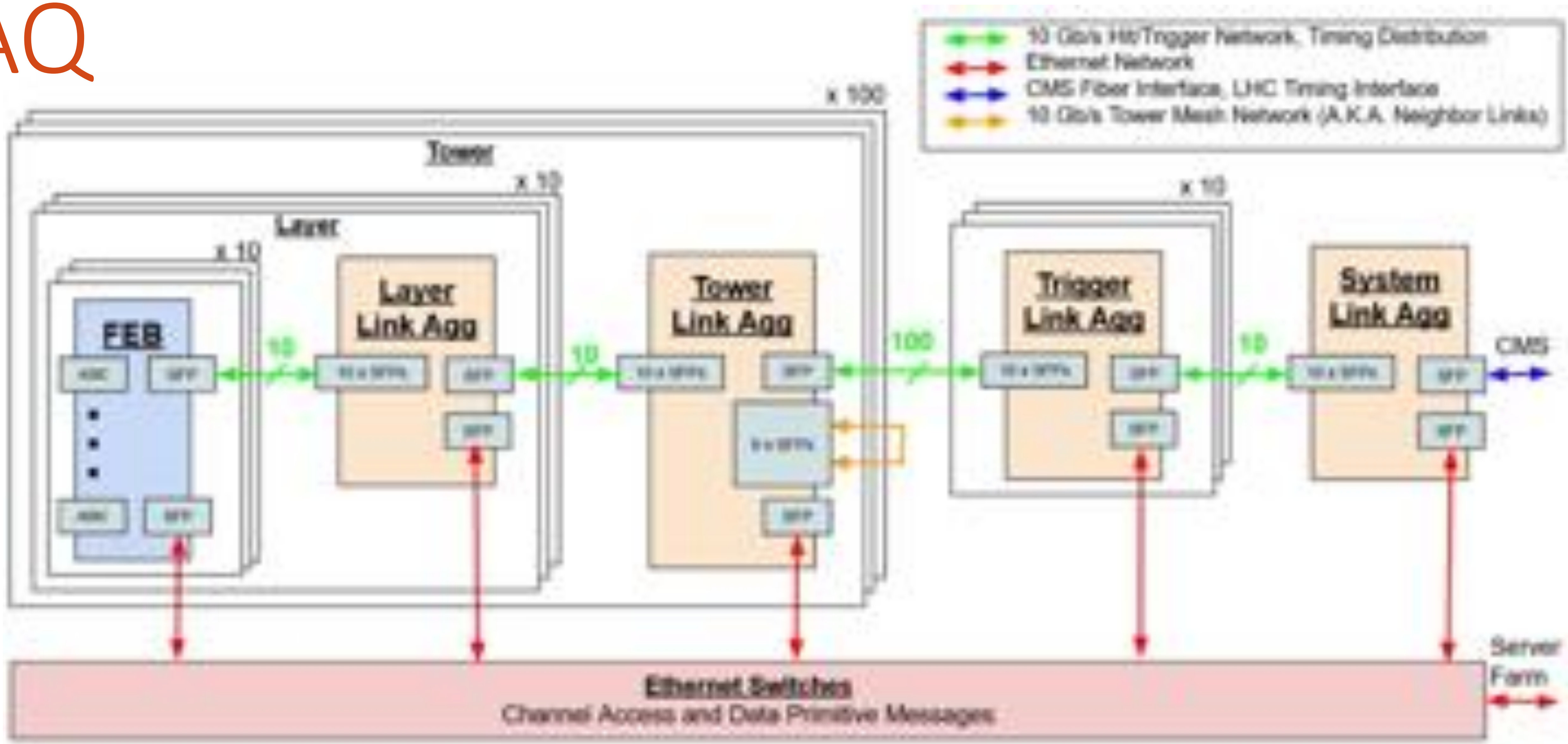




Ongoing R&D:

- Various WLSFs
  - Attenuation
  - Light collection
- Various SiPMs
  - Dark counts
- Scintillator bar geometry

DAQ



Preliminary design: tower-by-tower approach, with modular design of FEBs and link aggregation boards, is scalable and stage-able

Tower aggregation module triggers on upward-going tracks that form a vertex within a 3x3 tower module

LHC timing distributed across all modules to synchronize with CMS

Conclusions

- MATHUSLA is a planned external LLP detector for the HL-LHC that can probe deep into LLP parameter space in a variety of Beyond the Standard Model scenarios
  - Including many DM models
- Significant recent progress and ongoing efforts
  - Extruded scintillators, fibers, SiPMs, trigger, DAQ
  - Simulations studies of rare backgrounds
  - Tracking algorithms for MATHUSLA's unique environment
  - Cosmic ray physics case
- Aiming to produce TDR by early 2022, followed by prototype module and full detector for HL-LHC
- New collaborators always welcome!

References

John Paul Chou, David Curtin, and H.J. Lubatti. New detectors to explore the lifetime frontier. Physics Letters B, 767:29–36, Apr 2017.

Cristiano Alpigiani et al. A Letter of Intent for MATHUSLA: a dedicated displaced vertex detector above ATLAS or CMS, 2018, arXiv:1811.00927.

David Curtin and Michael E. Peskin. Analysis of long-lived particle decays with the MATHUSLA detector. Physical Review D, 97(1), Jan 2018.

David Curtin et al. Long-lived particles at the energy frontier: the MATHUSLA physics case. Reports on Progress in Physics, 82(11):116201, Oct 2019.

Imran Alkhatib. Geometric Optimization of the MATHUSLA Detector, 2019, arXiv:1909.05896.

Cristiano Alpigiani. Exploring the lifetime and cosmic frontier with the MATHUSLA detector, 2020, arXiv: 2006.00788.

M. Alidra et al. The MATHUSLA Test Stand, 2020, arXiv:2005.02018.

Jared Barron and David Curtin, On the Origin of Long-Lived Particles, 2020, arXiv:2007.05538.

Cristiano Alpigiani et al. An Update to the Letter of Intent for MATHUSLA: Search for Long-Lived Particles at the HL-LHC, 2020, arXiv:2009.01693.

BACKUP

Backgrounds: Recent Refined Estimates

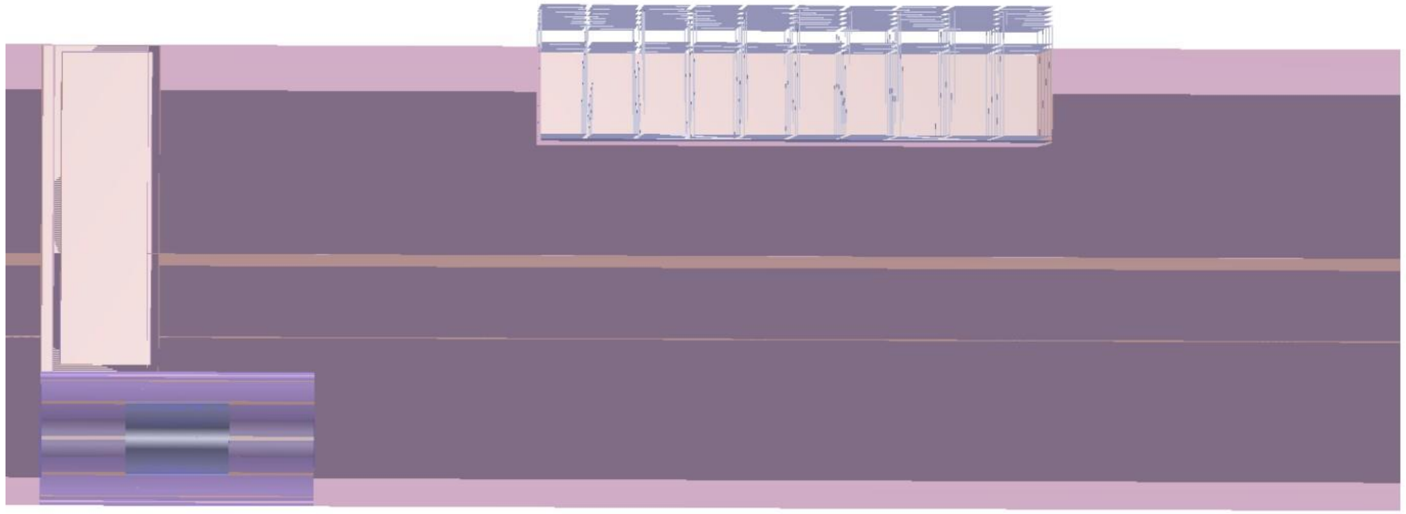
- Cosmic rays
  - Calibrations performed using Test Stand measurements (taken above ATLAS IP in 2018) arXiv: 2005.02018
  - Simulated using PARMA 4.0 + GEANT4
  - Downward-going events ~3 x 10^{14} over entire HL-LHC run, distinguished from LLPs using timing cuts
  - Upward-going events ~2 x 10^{10} , produced through inelastic backscatter from CRs that hit the floor, or through decay of stopped muons in floor. Tiny fraction can produce fake DV, via decay to 3 charged tracks
  - Rare production of K^0_L harder to estimate; veto strategies are available. Currently working on precise estimates and studying rejection

Backgrounds: Recent Refined Estimates

- Rare decays of muons originating from HL-LHC collisions
  - Expect ~2 x 10^8 upward-going muons over entire HL-LHC run, mostly from W and bbar production
  - Simulated using MadGraph & Pythia8
  - Full study underway to demonstrate optimal rejection while maintaining high LLP signal efficiency; test-bed for custom tracking algorithms in unique MATHUSLA environment

- Charged particles from neutrino scattering in decay volume
  - Simulated using GENIE
  - Neutrinos from HL-LHC collisions: using LHC minimum-bias samples, estimate << 1 “fake” DV/year
  - Atmospheric neutrinos: using flux measurements from Frejus experiment, estimate ~30 “fake” DV/year, reduced to < 1 with cuts

GEANT4 Simulations

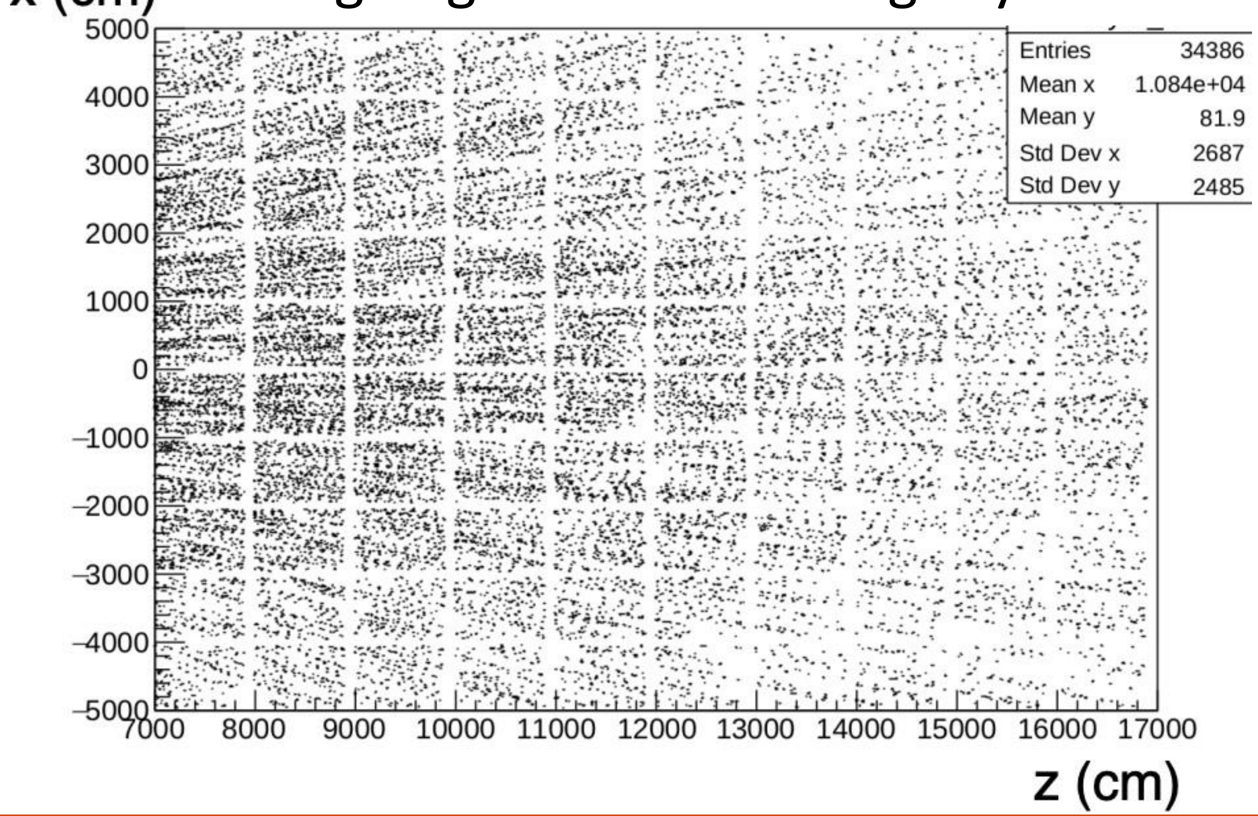


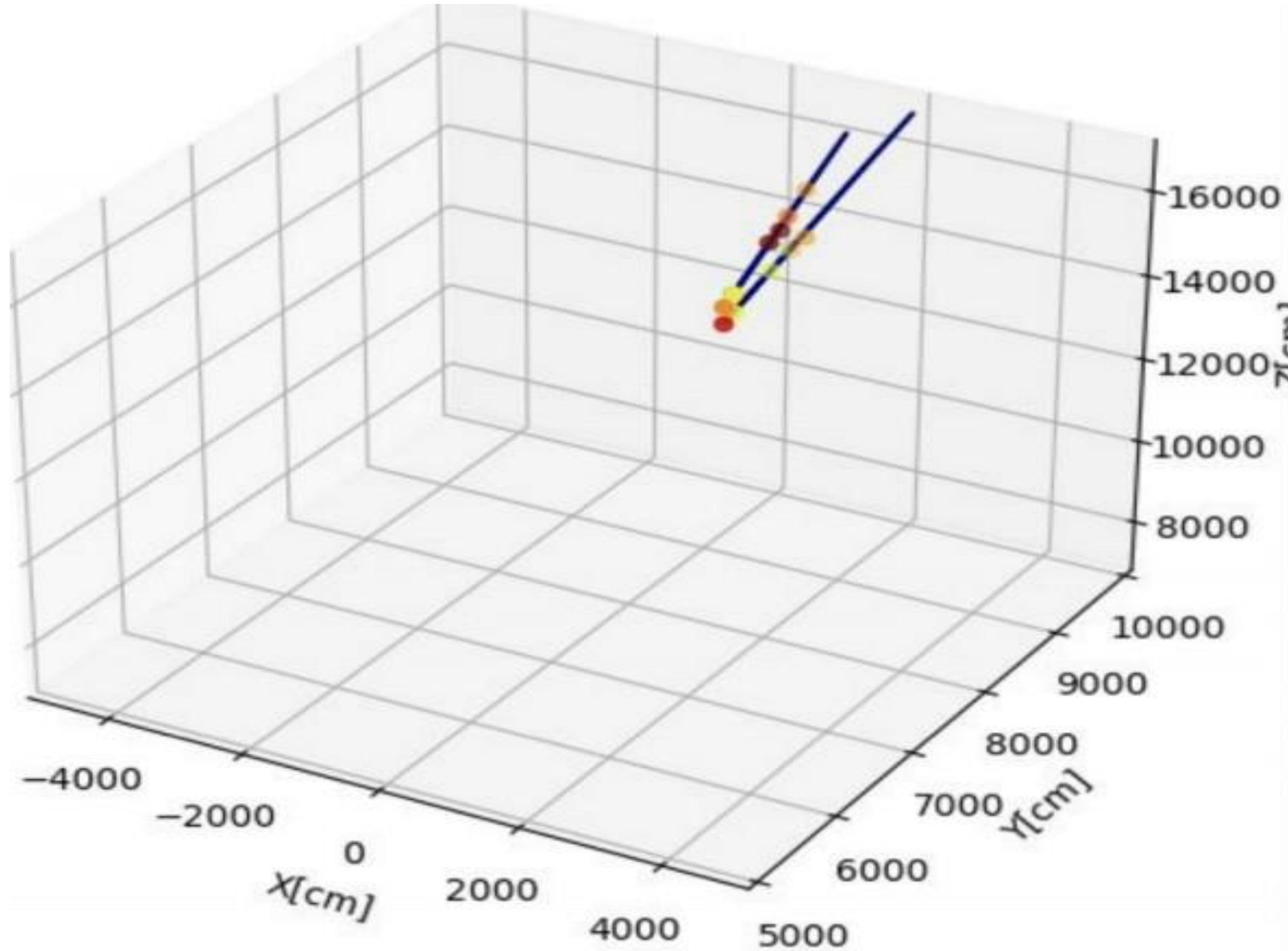
Cavern, access shaft, CMS, rock, detector all modeled

- Rock composition from a geological survey of LHC site
- CMS as hollow iron cylinder, ~10 interaction lengths

Preliminary track reconstruction finds displaced vertices; more sophisticated algorithms being implemented

Simulated Hits for W decays to upward-going muons: Tracking Layer 4

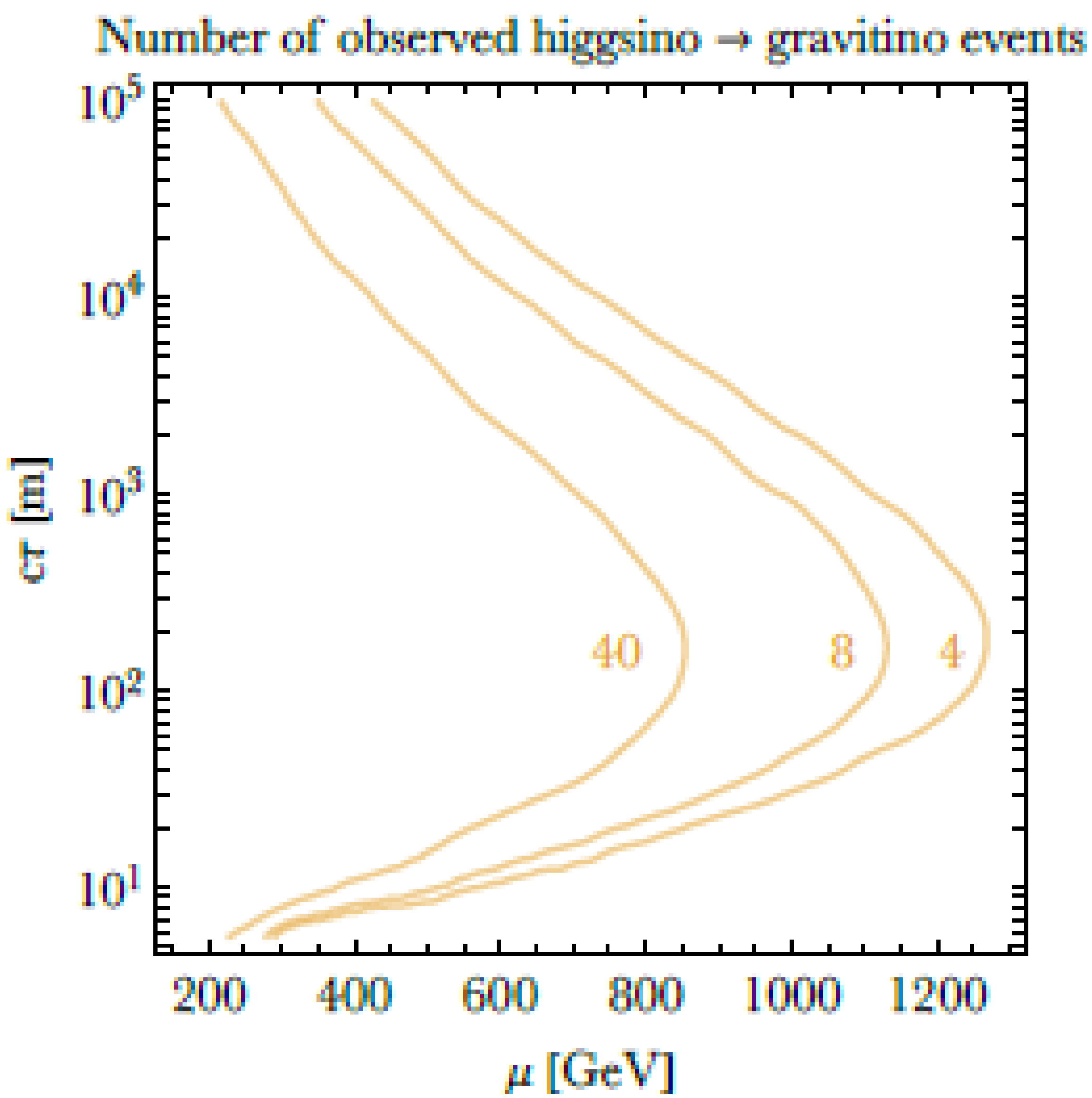


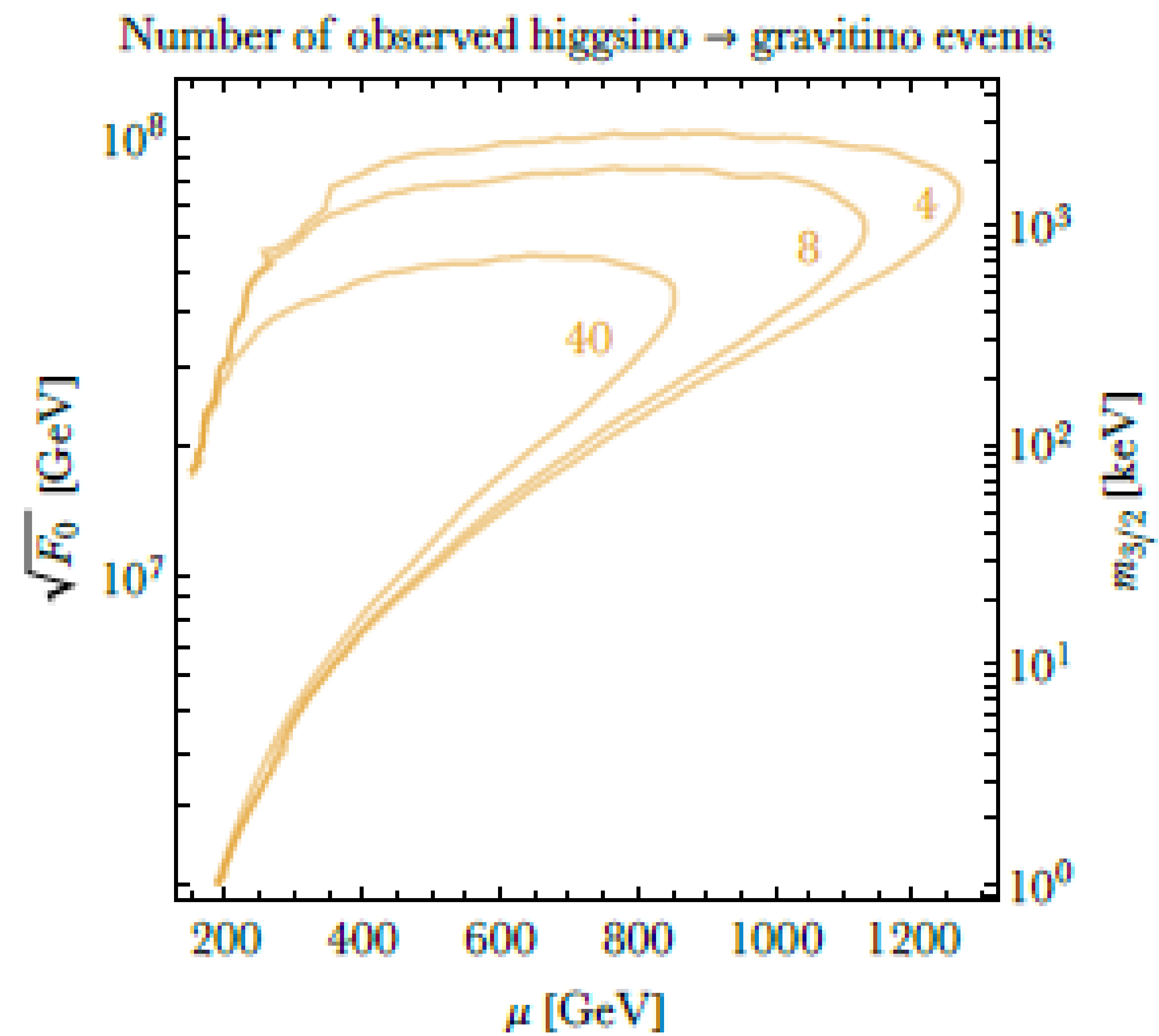


LLP Sensitivity: TeV-Scale

Any LLP production process with σ > fb can give signal.

e.g. meta-stable Higgsinos

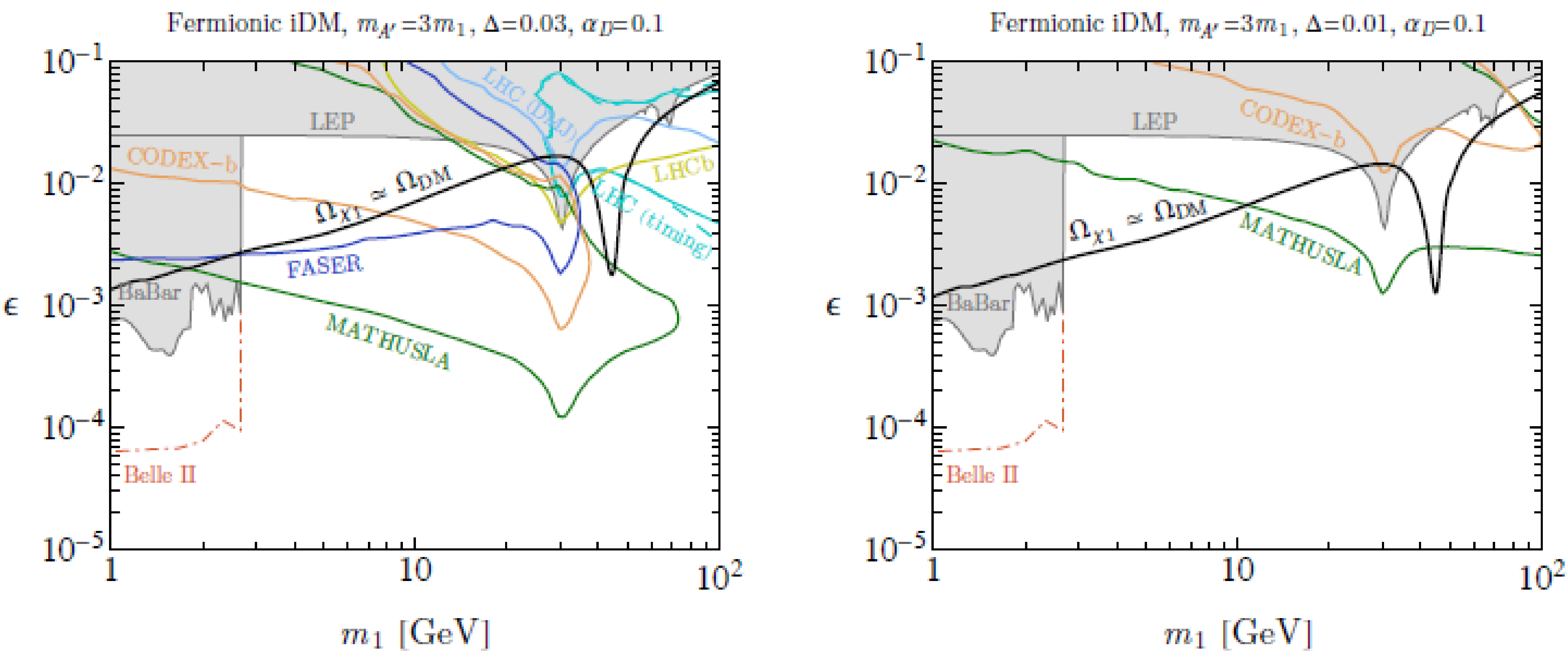




LLP Sensitivity: DM

Scenarios where LLP → DM + SM decay is the only way to see the DM

e.g. Inelastic Dark Matter: BSM mass eigenstates χ_1 (DM) and χ_2 (LLP) with mass splitting Δ , dark photon A' with mixing ϵ with SM photon



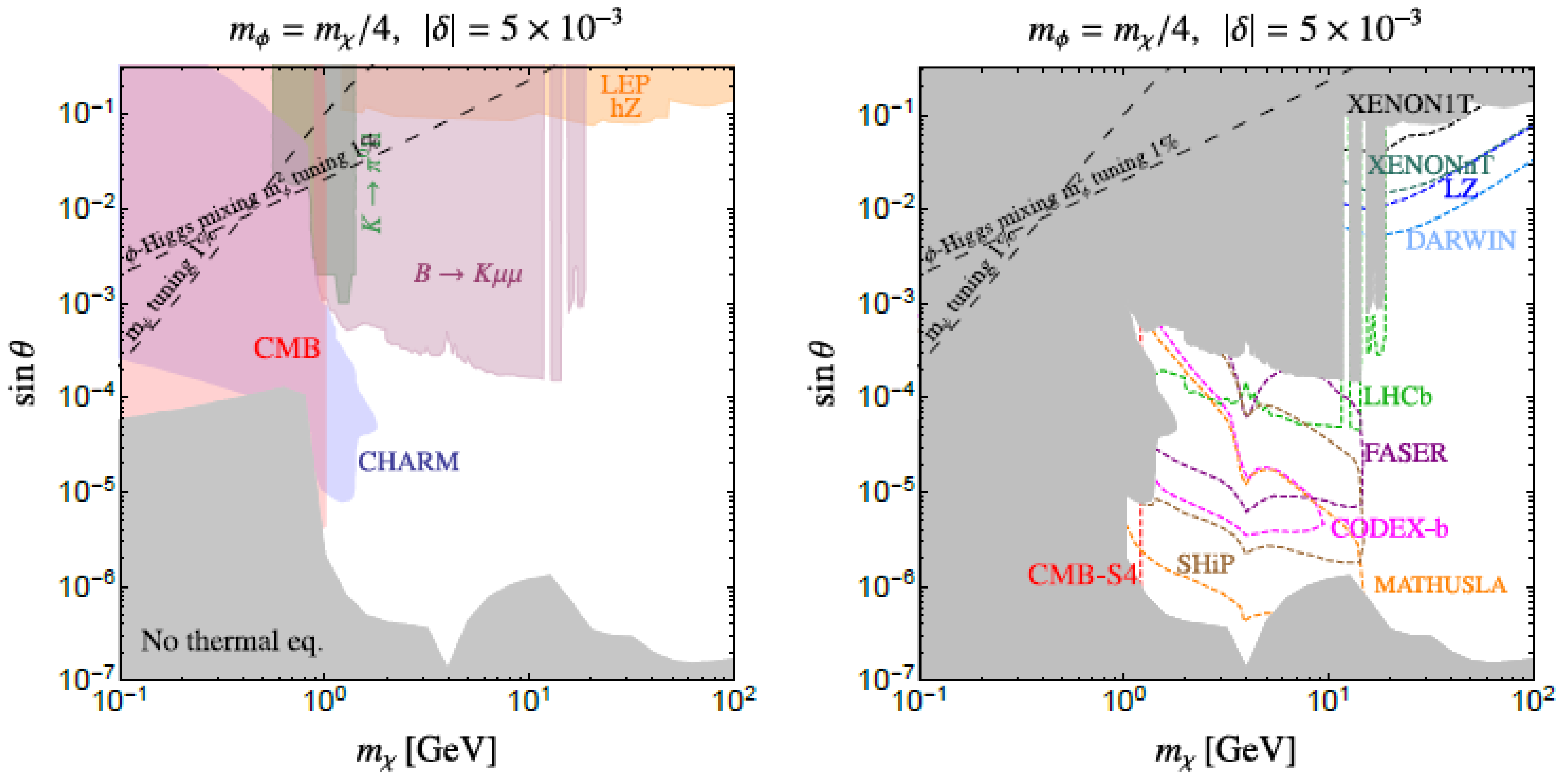
Black curve: thermal o-annihilations $\chi_2 \chi_1 \to A' \to f\bar{f}$ yield observed DM relic density

LLP Sensitivity: DM

Scenarios where DM model requires existence of LLP, but LLP signature does not involve the DM particle directly

e.g. Co-Annihilating DM: BSM χ and χ_2 with mass splitting δ,

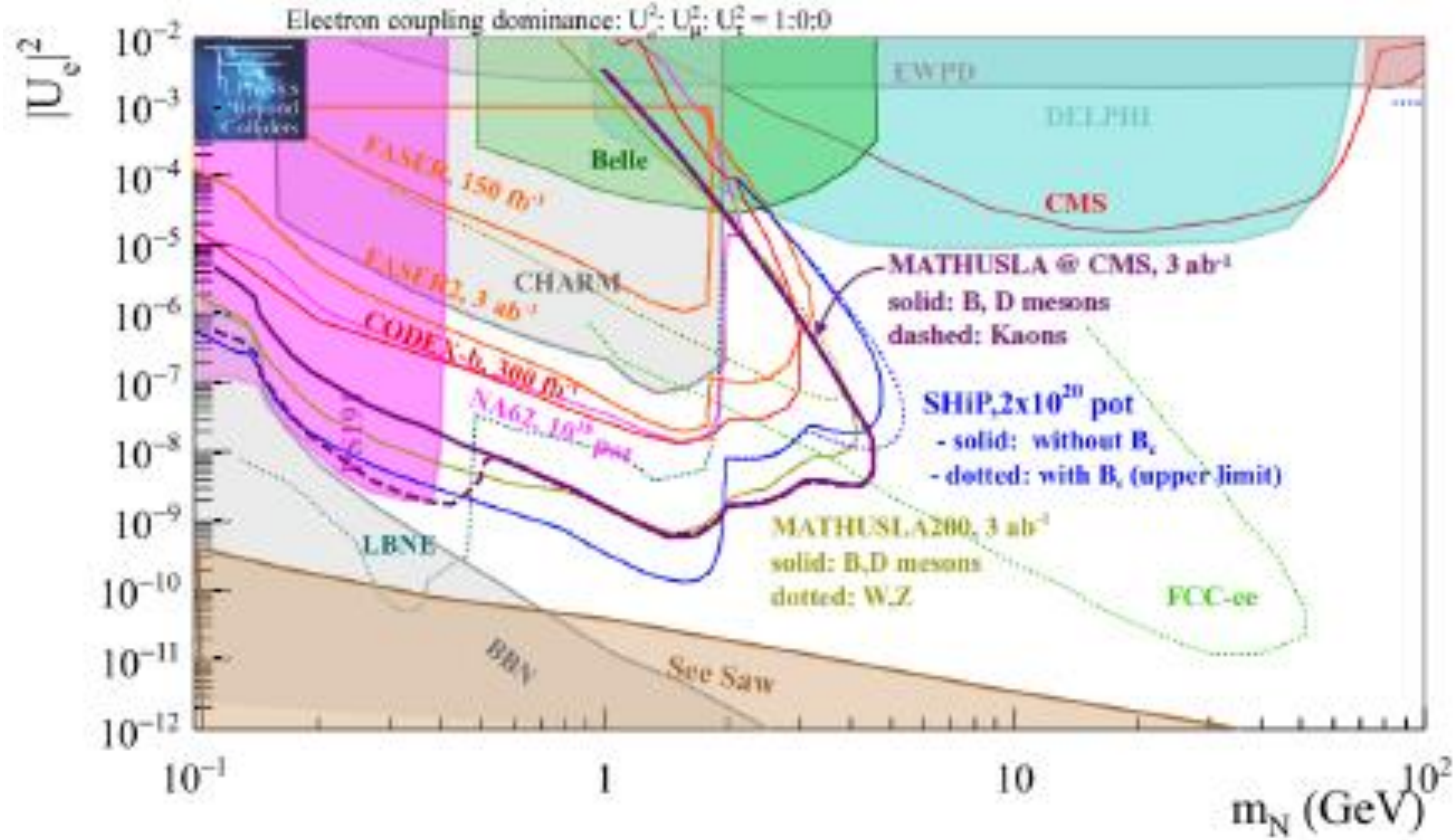
$\chi\,\chi_2 \rightarrow \phi\phi$ where scalar ϕ has mixing angle θ with SM Higgs



LLP Sensitivity: GeV-Scale

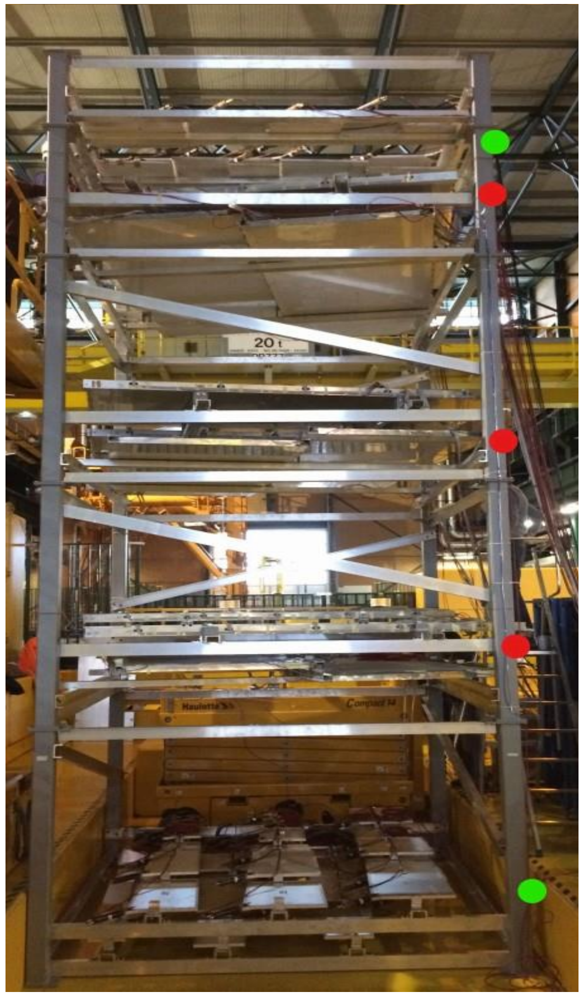
For heavy neutral leptons, reach is similar to SHiP

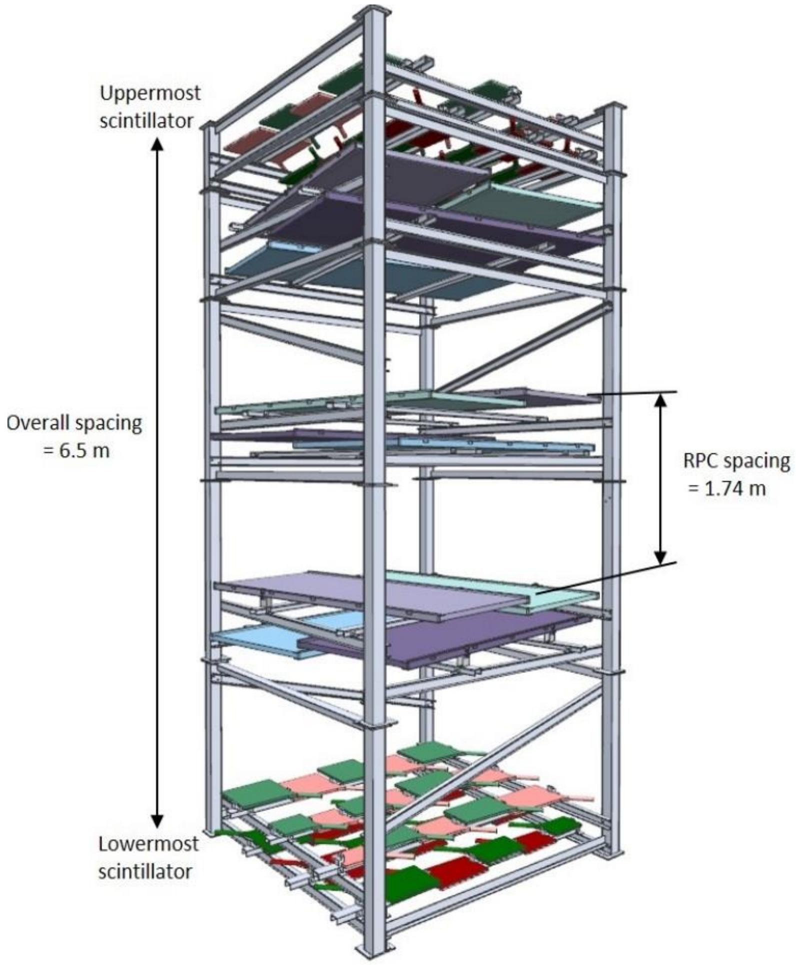
e.g. sterile neutrino N predominantly mixing with electron-neutrino

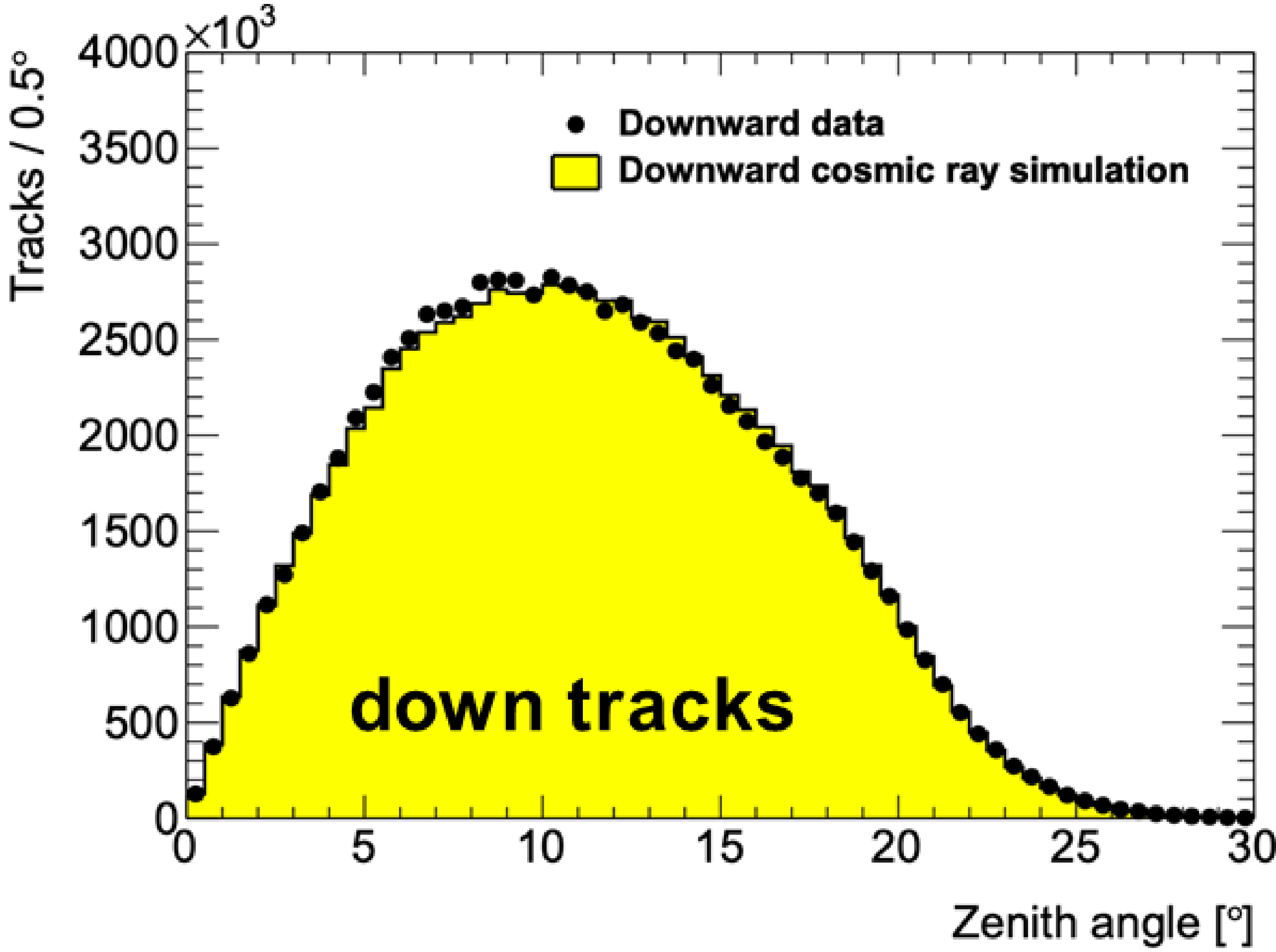


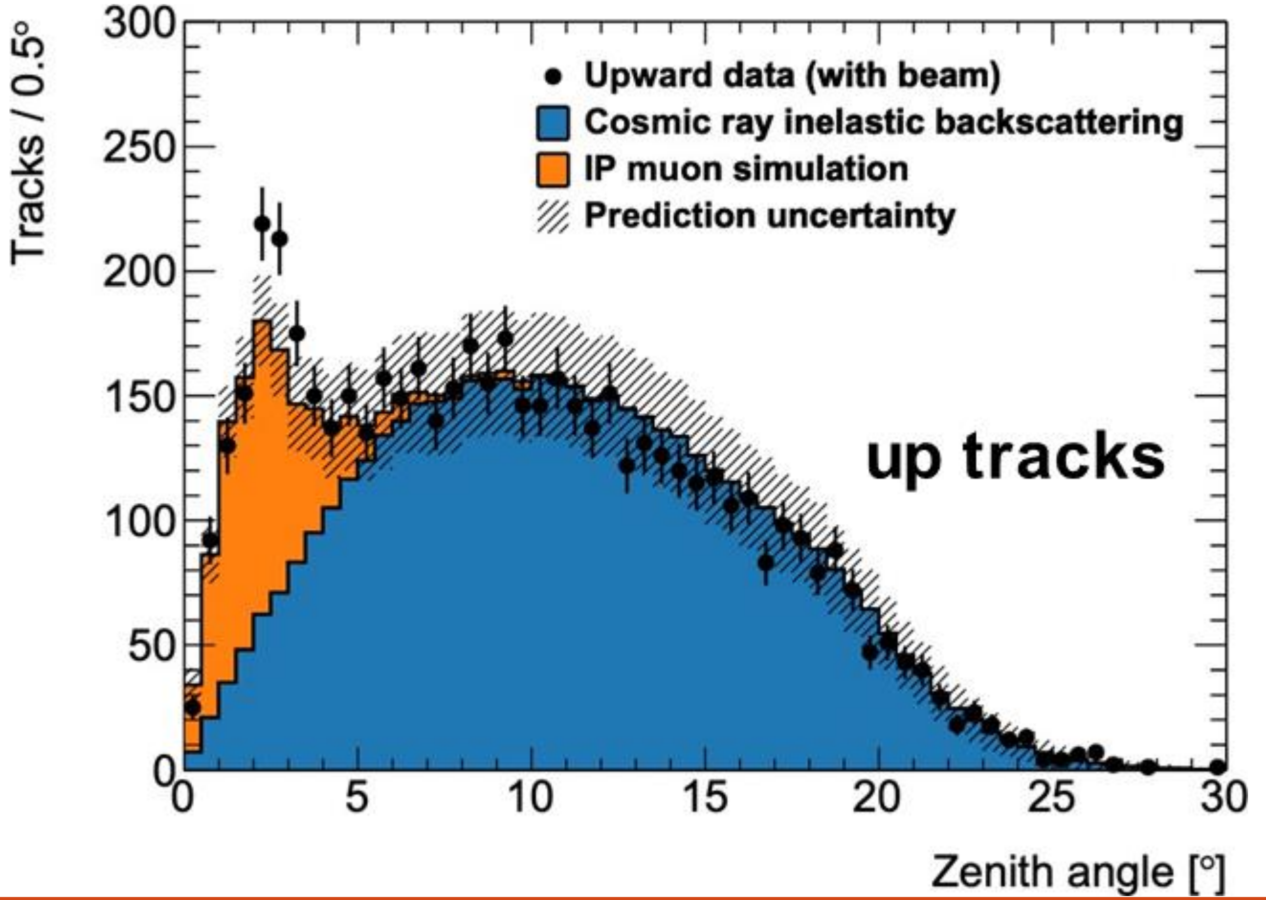
MATHUSLA Test Stand

Operated above ATLAS in 2018









Downward cosmic rays, upward LHC muons and upward CR backscatter well described by simulations