In-Situ and Ex-Situ Observations of an Extremely Long-lived Tail in TPB Fluorescence Under Alpha Excitation in DEAP-3600

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#### **A Bit About DEAP**

- Located at SNOIab
- Over 3 tonnes LAr
- TPB Wavelength Shifter
- Viewed by 255 PMTs
- 50 cm acrylic light guides between PMTs and LAr
- Inside water tank muon veto





### **Dark matter interactions are rare**



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### Dark matter interactions are rare

## In order to detect dark matter, we must have as little background as possible



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- Particle scatters in LAr
- Singlet and triplet dimers form
- 128 nm photos emitted
- TPB shifts photons to visible
- Photons are detected





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 $\tau_1 \sim 6 \text{ ns}$  $\tau_2 \sim 1600 \text{ ns}$ 



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From DEAP-1 [Astroparticle Physics 85 (2016) 1-23]





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IPB



15

#### Too low energy



| PF

#### • $T_{Po-210}^{1/2}$ = 138 d • E( $\alpha$ ) = 5.304 MeV • E(<sup>206</sup>Pb) = 103 keV



Usually too high energy, but maybe...



#### • $T^{1/2}_{Po-210}$ = 138 d • E( $\alpha$ ) = 5.304 MeV • E(<sup>206</sup>Pb) = 103 keV





### TPB

#### Too high energy



IPF

#### • $T_{Po-210}^{1/2} = 138 d$ • $E(\alpha) = 5.304 MeV$ • $E(^{206}Pb) = 103 keV$



#### Possible background!



#### • $T^{1/2}_{Po-210}$ = 138 d • E( $\alpha$ ) = 5.304 MeV • E(<sup>206</sup>Pb) = 103 keV



## Both of these dangerous cases involve the α scintillating in the TPB



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## Both of these dangerous cases involve the α scintillating in the TPB

What does TPB scintillation look like under  $\alpha$  excitation?

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### PMT\_\_\_\_

Contains Spectralon Cup Filled with vacuum or LAr

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#### ~0.3 mg/cm<sup>2</sup> of Tetraphenyl Butadiene (TPB)



## $S(t) = (N_0 e^{-t/\tau_s} + f(t)) * Gaus(t,\sigma_r)$



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#### **Time resolution**

## $S(t) = (N_0 e^{-t/\tau_s} + f(t)) * Gaus(t,\sigma_r)$



# Singlet decays $S(t) = (N_0 e^{-t/\tau_s} + f(t)) * Gaus(t, \sigma_r)$



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# Triplet decays $S(t) = (N_0 e^{-t/\tau_s} + f(t)) * Gaus(t, \sigma_r)$

$$f(t) = \frac{k_s}{2t_b} \frac{k_{tt}}{\chi_{tt}} \frac{N_T(0) \exp\left(-\frac{2t/\tau_T}{T}\right)}{\left\{1 + \frac{t_a}{2t_b} \exp\left(\frac{t_a}{\tau_T}\right) \left[\operatorname{Ei}\left(-\frac{t+t_a}{\tau_T}\right) - \operatorname{Ei}\left(-\frac{t_a}{\tau_T}\right)\right]\right\}^2 (1 + t/t_a)}.$$

R. Volts and G. Laustriat, "Radioluminescence des Milieux Organiques I. Étude Cinétique." Le Journal de Physique (29). 1968

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## Different Waveforms for $\alpha$ and $\beta$ Scintillation



30

### Longer Tails and More Extreme Difference at Cryogenic Temperature



#### **Does this long tail appear in actual experiments?**



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#### 200 $\,\mu\text{s}$ DAQ Window Data



#### 200 $\,\mu\text{s}$ DAQ Window Data







5 n V1720 (sub

Counts

Additional 320 PE in following 190 us Consistent with TPB scintillation

Standard 10 us window – 643 PE contains most of standard LAr scintillation pulses

100 us

time

0 us

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 $200 \,\mathrm{us}$ 

XXX + 206 Pb

206 Pb

#### Position reconstruction is hard



206 Ph Position reconstruction is hard Reduces fiducial mass → Decreases sensitivity

## Can TPB scintillation pulse shape help?

#### Conclusion

- Surface α decays are an important background in low-background experiments
- We have seen an extremely long tail in TPB scintillation under  $\alpha$  excitation, significantly different from its wavelength shifting and  $\beta$  scintillation time constants
- Evidence that this tail appears in DEAP-3600
- Could be used as a powerful tool for discriminating these backgrounds

End



### Difference between Wavelength Shifting and Surface Backgrounds

