

# First results from the NEWS-G direct dark matter search experiment

CAP Conference, Kingston, ON

28/05/2017

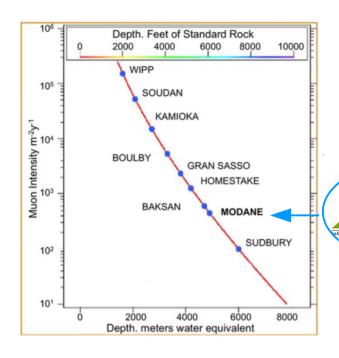
New Experiment With Spheres - Gas

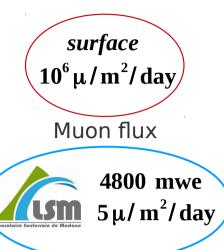




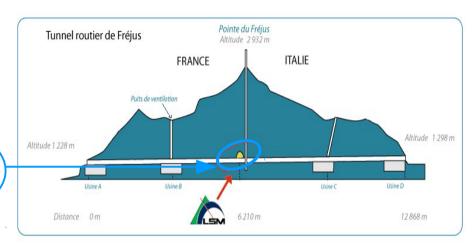


# NEWS-G @ LSM - Experimental Setup



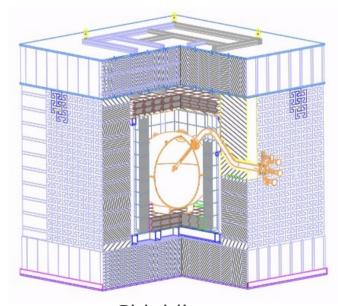


#### Laboratore Souterrain de Modane



#### **Data taking conditions**

9.7 kg.days of exposure with Neon+0.7 % CH<sub>4</sub> @ 3.1 bars ~280 g target mass, operated for 42.7 days in sealed mode



<u>Shieldings</u> **30 cm PE**, **10-15 cm Pb**, [**3-8**] **cm Cu** 

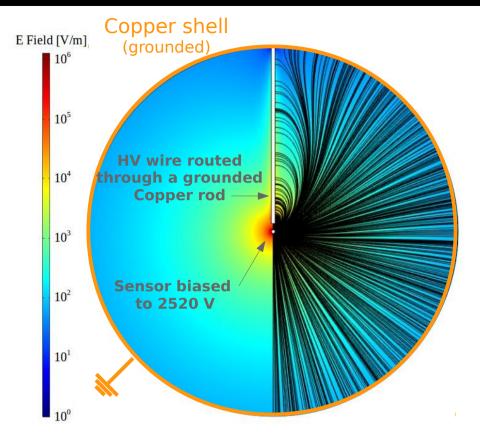


Vessel
60 cm Ø NOSV Copper



<u>Sensor</u> **6.3 mm** Ø

# Designed for low-mass WIMP search



#### Low threshold 10-40 eVee

#### High amplification gain arising from

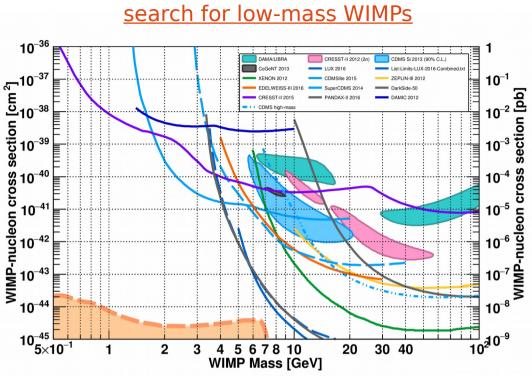
$$E(r) \propto \frac{1}{r^2}$$

#### **And Low Capacitance**

(doesn't depend on the size of the sphere)

$$C = \frac{4\pi\epsilon}{\left(\frac{1}{r_{sensor}} + \frac{1}{r_{vessel}}\right)} \approx 4\pi\epsilon r_{sensor} \approx 0.35 \,\mathrm{pF}$$

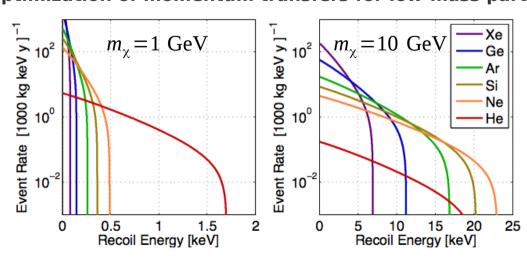
sensitivity to single electrons

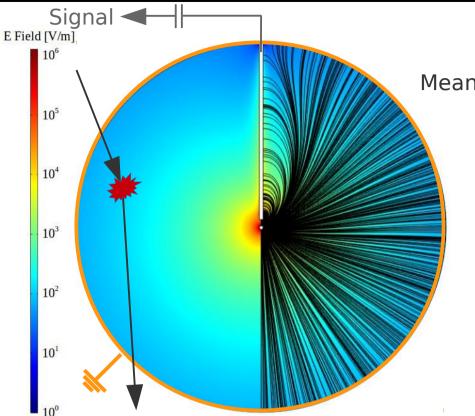


with Spherical Proportional Counters

#### **Light Targets**

### Optimization of momentum transfers for low-mass particles





# **Primary Ionisation**

Mean number of primary electrons created :  $\langle N \rangle = \frac{E_R}{\epsilon_i}$ 

With Neon: 
$$\epsilon_{\gamma} = 36 \text{ eV}$$
  $\epsilon_{n} = \frac{\epsilon_{\gamma}}{Q(E_{P})} \approx 5 \epsilon_{\gamma}$ 

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## Drift of the electrons toward the sensor

Typical drift time surface → sensor : ~500 μs

#### **Avalanche Process**

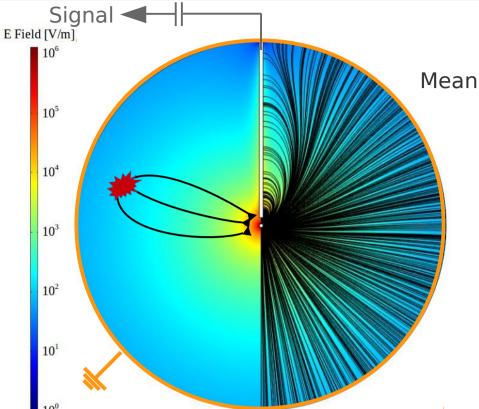
Each primary electron leads to thousands of secondary ionisations

#### **Signal Formation**

Current induced by Secondary ions

## Signal Read out

Induced current integrated by a resistive feedback charge sensitive pre-amplifier CAMBERRA (RC=50 µs) and digitized at 2.08 MHz



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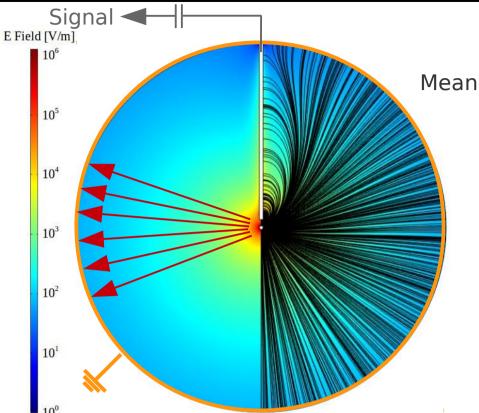
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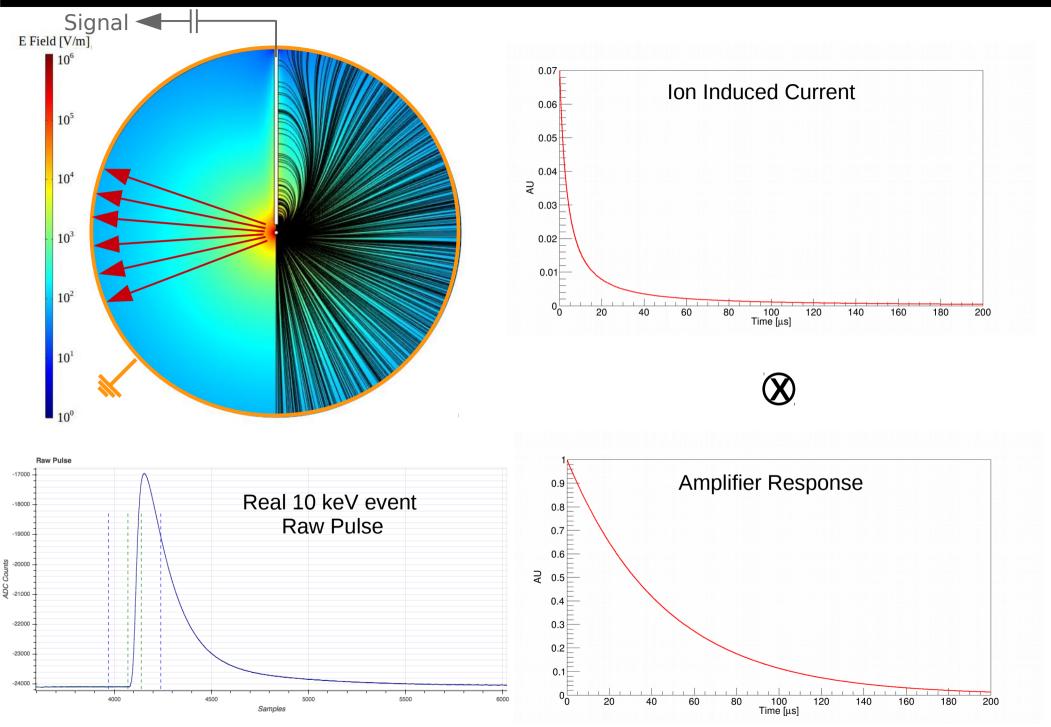
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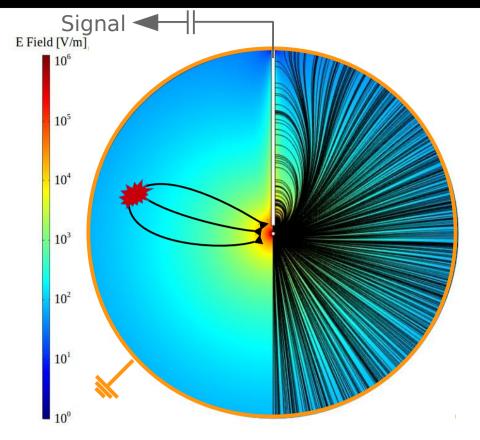
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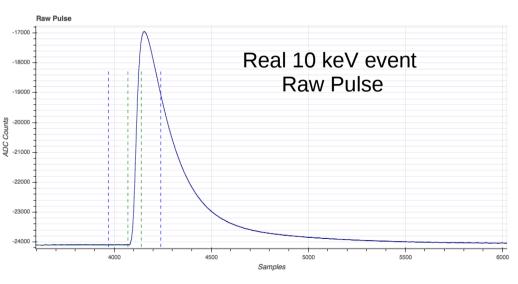
Typical drift time surface  $\rightarrow$  sensor (500  $\pm$  20) µs

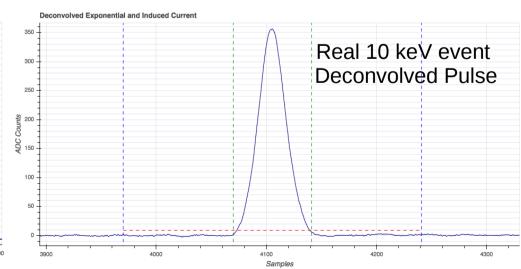
Gaussian dispersion in the arrival time of the primary electrons due to diffusion

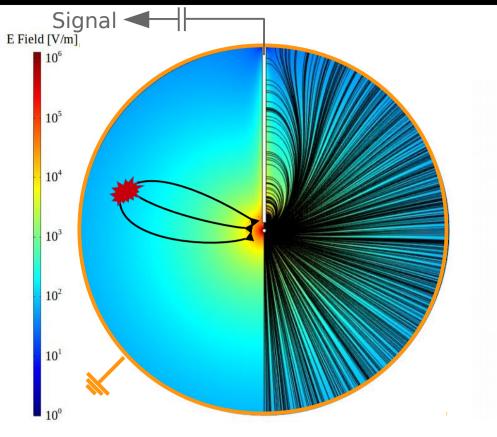
whose standard deviation increases with the radial location of the event

$$\sigma(\mathbf{r}) = \left(\frac{\mathbf{r}}{r_{sphere}}\right)^3 \times 20 \,\mu s$$

# **Fiducialisation**

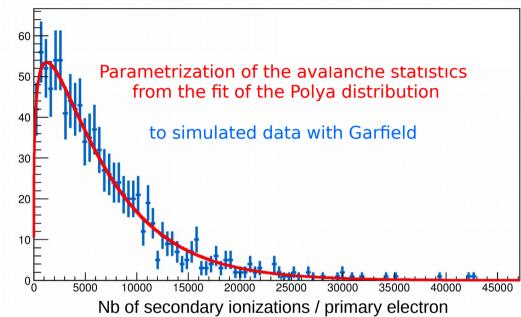


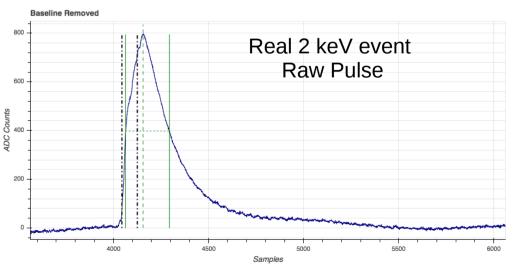


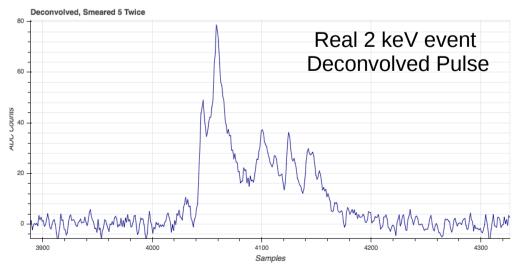


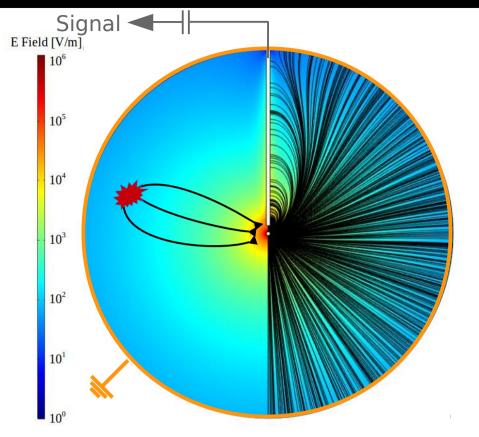
#### **Avalanche Process**

Each PE  $\rightarrow$  <N> ~ 7000 secondary ionisations



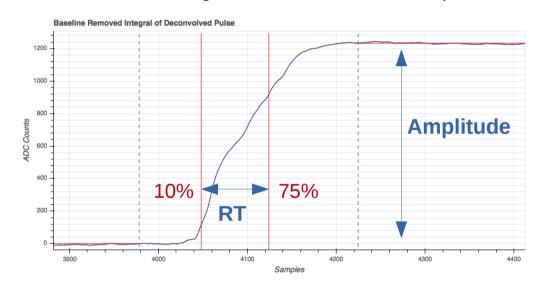


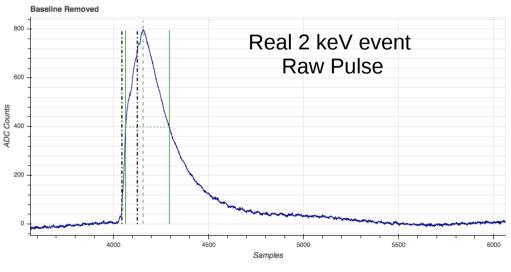


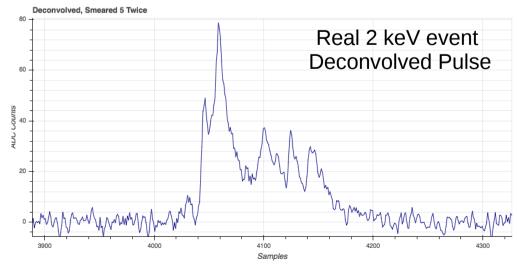


## **Energy and Diffusion Estimators**

#### Real 2 keV event Cumulative integration of the deconvolved pulse

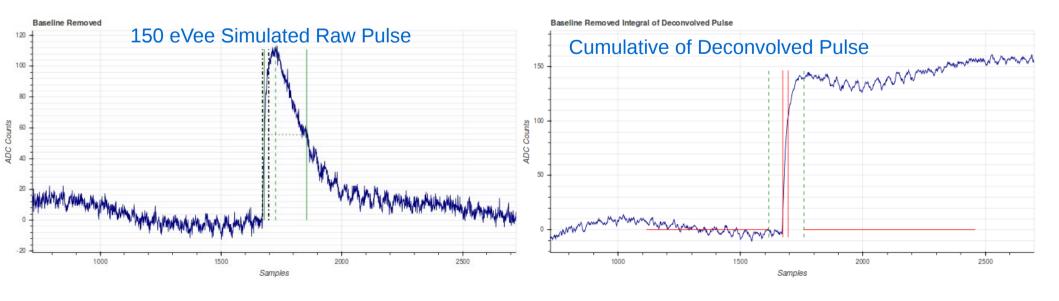




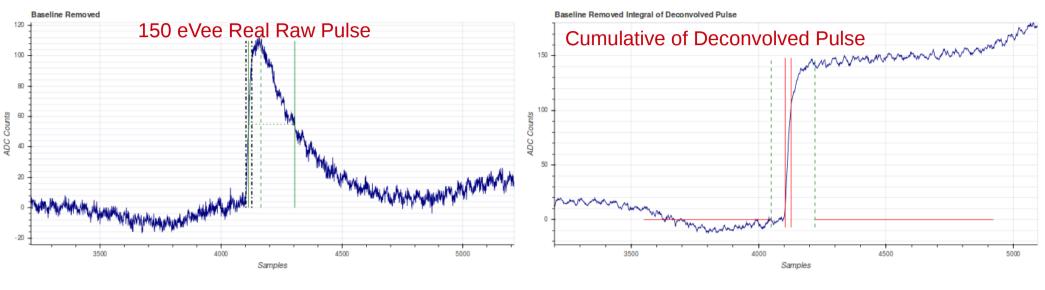


# **Simulations**

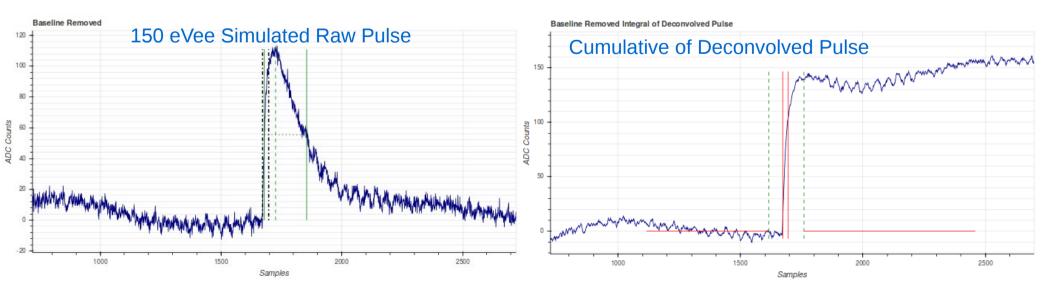
- Field map computed with COMSOL
- Quenching parametrization  $\,Q(E_{\scriptscriptstyle R})\,$  derived from SRIM (Stopping and Range of Ions in Matter)
- Primary electrons are individually drifted using drift parameters from Magboltz
- Number of secondary electrons drawn from the Polya distribution (parametrized with Garfield)
- Simulated pulses : ( Ion Induced current X Amplifier response )
- Noise templates taken from the pretraces of real pulses
- Same trigger algorithm and processing than for real pulses



#### **Data**

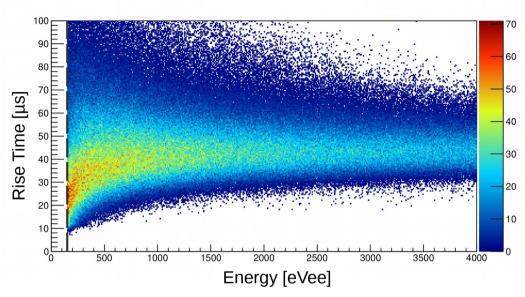


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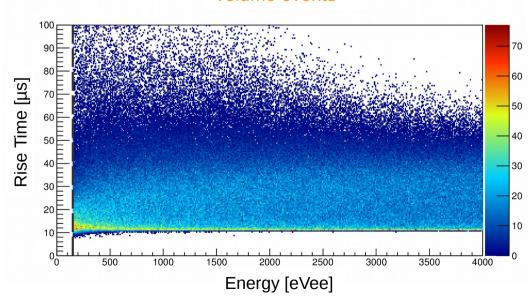


## **Simulations**

#### Surface events



#### Volume events



# Physics-run Data

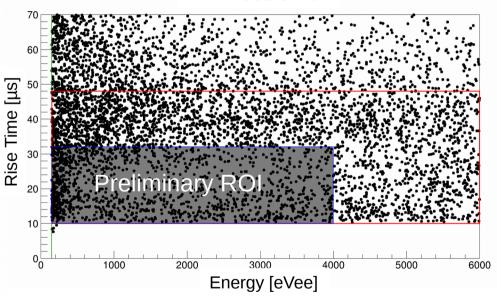
Quality cuts responsible for 20.1 % dead time

Total exposure = 34.1 live-days x 0.28 kg = 9.7 kg.days

Analysis threshold: 150 eVee (~720 eVnr)

100% trigger efficiency (trigger threshold @ ~35 eVee)

#### WIMP search run



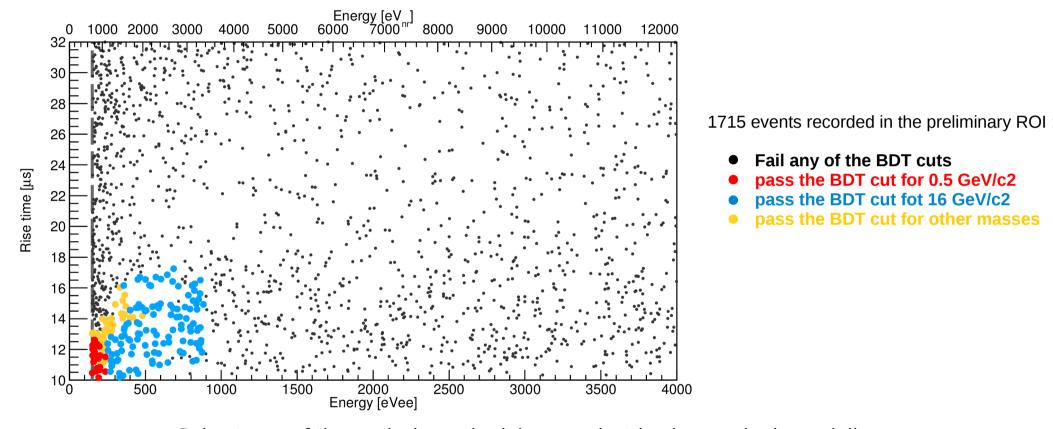
**Side Band region** used together with simulations to determine the number and distribution of background events expected in the **preliminary ROI** 

Further tuning of the ROI performed with a Boosted Decision Tree (BDT)

# Physics-run data analysis

We make use of a Boosted Decision Tree (BDT) algorithm that we train with our signal and background models to identify the fine-tuned ROI that maximizes our expected sensitivity for 8 different WIMP masses

We end with a WIMP-mass-dependent fine-tuned ROI in the rise time vs energy plane



Robustness of the analysis methodology against background mis-modeling

If the BDT were to be trained with inaccurate background models, the fine-tuned ROI would just end to be non-optimized for signal/background discrimination

Still, an accurate modelisation of the signal is critical for the exclusion limit to be unbiased.

# Comparison of Calibration data with Simulation

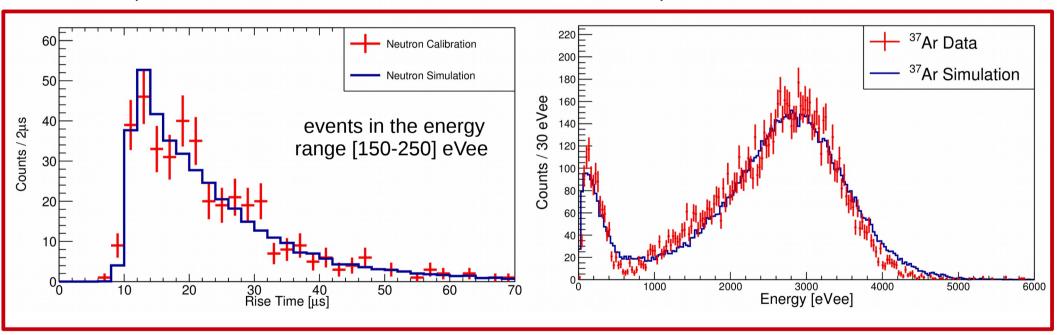
### Validation of the simulation through its agreement with calibration Data

37Ar gas added to the mixture: 2.82 keV and 270 eV X-rays from the electron capture in the K- and L-shells respectively.

Am-Be neutron source: nuclear recoils homogeneously distributed in the volume.

#### Comparison of Neutron data with simulation

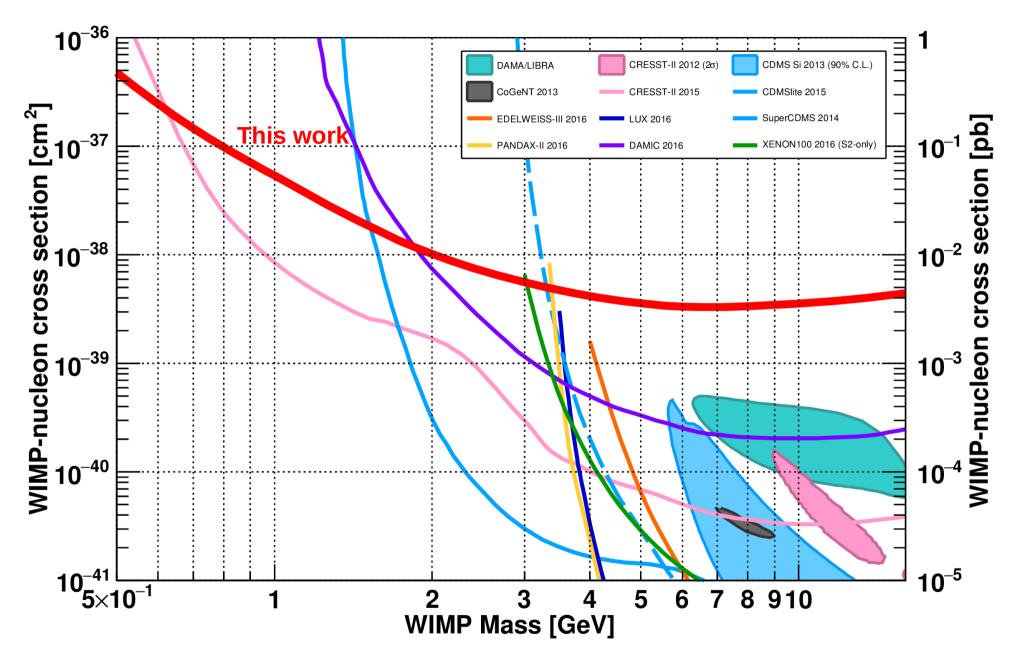
#### Comparison of 37Ar data with simulation



Response in rise time and drift parameters

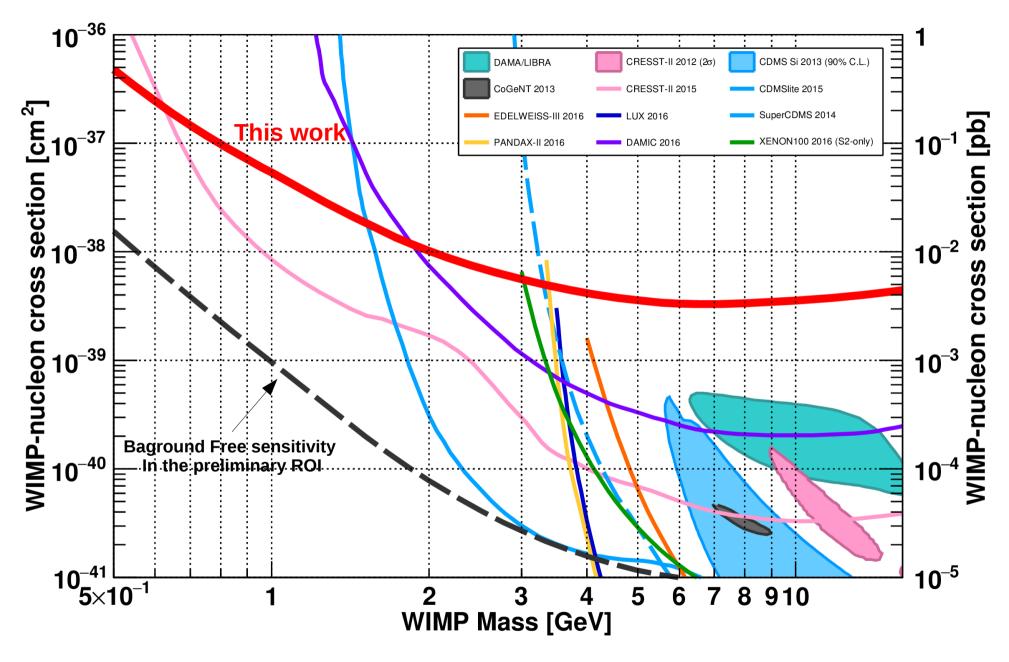
Response in energy

The overall agreement between the simulation and calibration data allows us to confidently derive our sensitivity from simulated WIMP events



90 % (CL) upper limit derived from Poisson statistics.

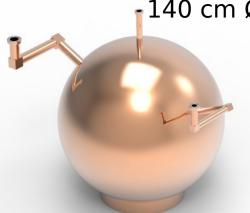
Standard halo model parameters :  $\rho_{DM} = 0.3 \text{ GeV/c}^2/\text{cm}^3$   $v_{esc} = 544 \text{ km/s}$   $v_0 = 220 \text{ km/s}$ 



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140 cm Ø detector @ 10 bars (Ne, He, CH<sub>4</sub>)

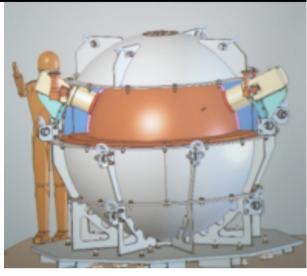


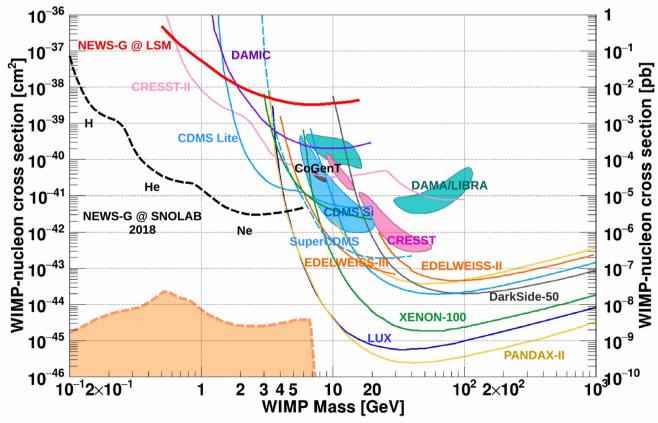
## **Copper vessel**

- Thickness ~12 mm
- Electropolishing cleaning

#### Lead shield

- 40 cm PE + Boron sheet
- 22 cm VLA (1 Bq/kg 210Pb)
- 3 cm archeological lead
- Air tight SS envelope to flush pure N





See G. Gerbier talk on Thursday for more details

# Conclusion and outlook

# **NEWS-G** @ LSM

The NEWS-G experiment sets competitive constraints on the spin-independent WIMP-nucleon scattering cross section in the sub-GeV range.

Paper to be submitted to the Astroparticle Physics Journal and uploaded on ArXiv by the end of next week

# **NEWS-G @ SNOLAB**

These results are very promising for the next phase of the NEWS-G experiment @ SNOLAB

See G. Gerbier 's talk on Thursday for details

#### Talks from the NEWS-G collaboration all in the same room (Botterell B139)

#### Today at 4.15 pm

A. Brossard: "Sensor optimisation and gas quality analysis for spherical gas detector operation"

D. Durnford: "Calibration schemes for Spherical Gas Detectors"

Thursday at 1.30 pm

G. Gerbier: "Status of NEWS-G experiment"

P. Di Stefano: "Quenching measurements for a spherical detector at the COMIMAC facility"



Queen's University Kingston - G Gerbier, P di Stefano, R Martin, T Noble, D. Durnford

A Brossard, A Kamaha, F Vazquez de Sola, Q Arnaud, K Dering, J Mc Donald, M Clark, M Chapellier



- Copper vessel and gas set-up specifications, calibration, project management
- Gas characterization, laser calibration, on smaller scale prototype
- Simulations/Data analysis

**IRFU** (Institut de Recherches sur les Lois fondamentales de l'Univers)/**CEA Saclay** -I Giomataris, M Gros, C Nones, I Katsioulas, T Papaevangelou, JP Bard, JP Mols, XF Navick,



- Sensor/rod (low activity, optimization with 2 electrodes)
- Electronics (low noise preamps, digitization, stream mode)
- DAQ/soft

LSM (Laboratoire Souterrain de Modane), IN2P3, U of Chambéry - F Piquemal, M Zampaolo, A DastgheibiFard

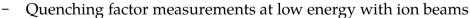


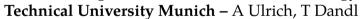
- Low activity archeological lead
- Coordination for lead/PE shielding and copper sphere

Thessaloniki University – I Savvidis, A Leisos, S Tzamarias, C Elefteriadis, L Anastasios

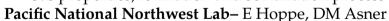
- Simulations, neutron calibration
- Studies on sensor

LPSC (Laboratoire de Physique Subatomique et Cosmologie) Grenoble - D Santos, JF Muraz, O Guillaudin

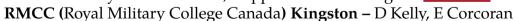




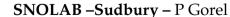
- Gas properties, ionization and scintillation process in gaz



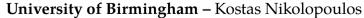
- Low activity measurements, Copper electroforming



- 37 Ar source production, sample analysis



- Calibration system/slow control



- Simulations, analysis, R&D





- Future R&D on light detection, sensor









