



Laurentian University
Université Laurentienne

Simulations of the Muon Veto for the PICO Experiment

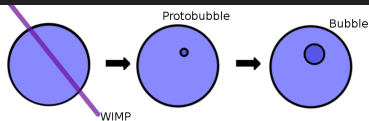
Olivia Scallon

2017 CAP Congress / Congrès de l'ACP 2017

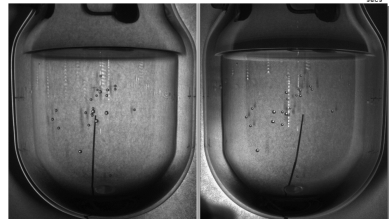
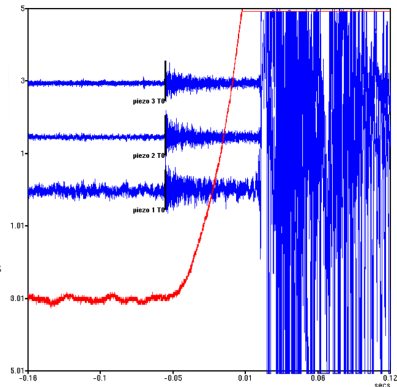
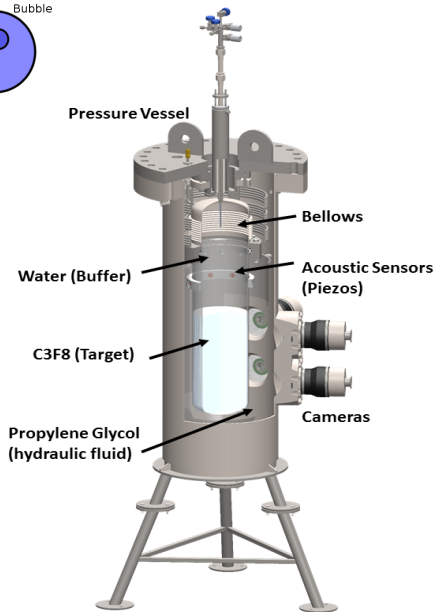
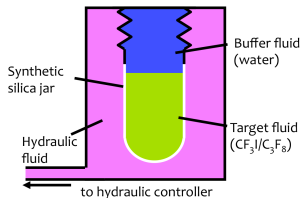
May 30, 2017

The PICO experiment

Detection Principles : Physics



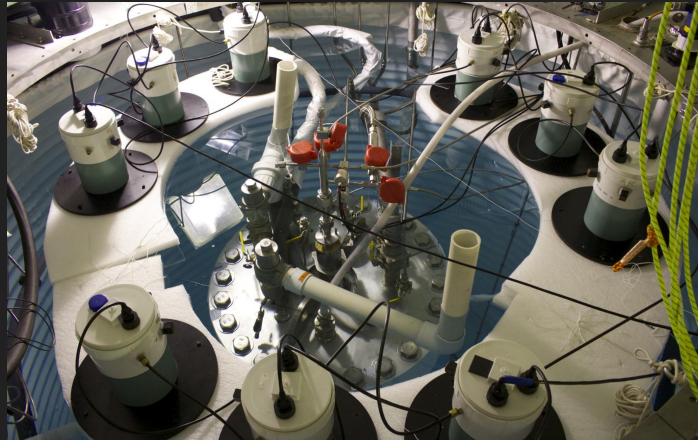
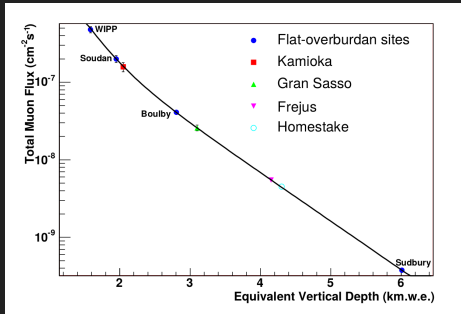
- Radiation induced boiling of superheated liquid : liquid-to-vapour phase transition.
- We put the detector in a metastable state and wait for the heat spike formed by a collision.



Physics and Backgrounds

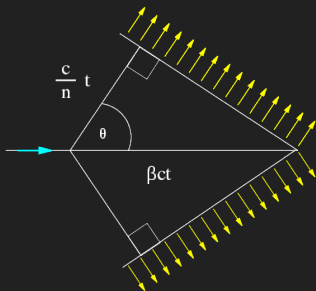
Cosmic Muons

- Muons induce high energy neutrons.
- Muon flux in SNOLAB : $3.77 \times 10^{-10} \text{ cm}^{-2} \text{ sec}^{-1}$
- Detector surrounded by a muon veto



Geant4 Muon Veto Simulations

Physics



Muons travelling faster than the speed of light in water produce Cherenkov radiation.

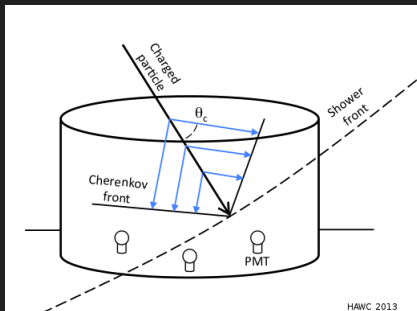
$$n = 1.33 \text{ for water}$$

$$\cos(\theta) = \frac{1}{n\beta}$$

$$\beta = v/c$$

number of optical photons produced :

$$N = 2\pi\alpha Z^2 \sin^2(\theta) \left(\frac{1}{\lambda_1} - \frac{1}{\lambda_2} \right) L$$



Geant4 Muon Veto Simulations

Geant4 Technicalities

- Full simulation of the already existing veto for PICO-60

- Geant4.10.03

- Goal : optimize a geometry for the future veto for PICO-500.

Particles :

- G4OpticalPhoton
- G4Meson
- G4Boson
- G4Baryon
- G4Ion

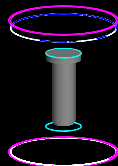
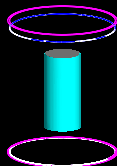
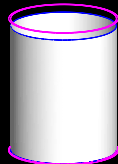
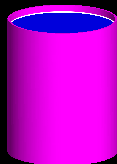
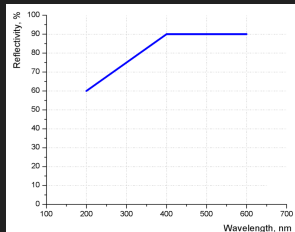
Processes :

- G4Cerenkov
- G4OpBoundaryProcess
- G4Scintillation
- G4OpAbsorption
- G4OpRayleigh

Geant4 Muon Veto Simulations

Detector Geometry

- Water (UPW)
- Water Tank
3.6m, d=2.8m, thickness=5cm
- PVC liner
(C₂H₃Cl) density 1.35 g/cm³
- Pressure vessel
Stainless Steel 302, density=8.03*g/cm³
- TYVEK cover

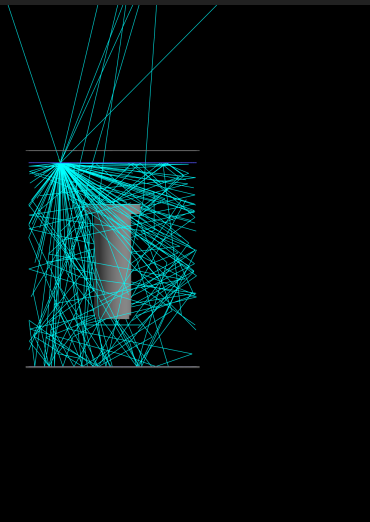


Geant4 Muon Veto Simulations

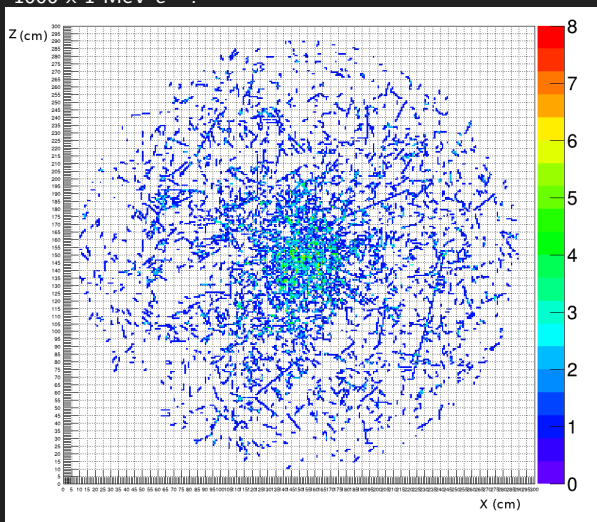
Scoring - Optical Photons Counting

3m x 3m scoring mesh with 1cm x 1cm bins on top of the water tank.

1 MeV e^- :



1000 \times 1 MeV e^- :



Geant4 SNOLAB Muon Source

D.-M. Mei, A. Hime. *Muon-Induced Background Study for Underground Laboratories.*

The energy spectrum :

$$\frac{dN}{dE_\mu} = A e^{-bh(\gamma_\mu - 1)} [E_\mu + \epsilon_\mu (1 - e^{-bh})]^{-\gamma_\mu}$$

With A the normalization constant with respect to the differential muon intensity at a given depth h , E_μ the muon energy after crossing h .

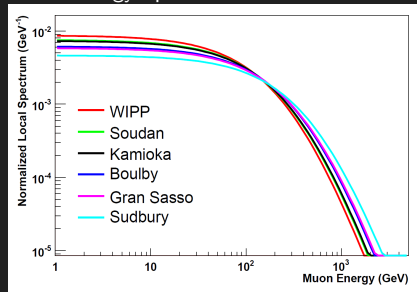
$b = 0.4/\text{km.w.e.}$, $\gamma_\mu = 3.77$ and $\epsilon_\mu = 693 \text{ GeV}$.

The muon angular distribution :

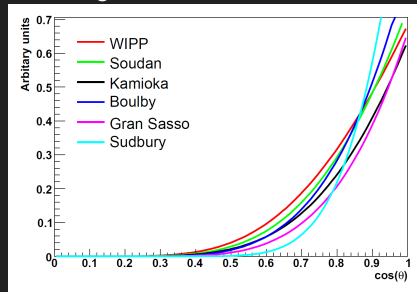
$$I_{th}(h, \theta) = \frac{I_1 e^{\left(\frac{-h_0}{\lambda_1 \cos(\theta)}\right)} + I_2 e^{\left(\frac{-h_0}{\lambda_2 \cos(\theta)}\right)}}{\cos(\theta)}$$

vertical depth h_0 , zenith angle θ , $I_1 = (8.60 \pm 0.53) \times 10^{-6} \text{ sec}^{-1} \text{ cm}^{-2} \text{ sr}^{-1}$, $I_2 = (0.44 \pm 0.0) \times 10^{-6} \text{ sec}^{-1} \text{ cm}^{-2} \text{ sr}^{-1}$, $\lambda_1 = 0.45 \pm 0.01 \text{ km.w.e.}$, $\lambda_2 = 0.87 \pm 0.02 \text{ km.w.e.}$

Muon Energy Spectrum :



Muon Angular Distribution :



Geant4 Muon Veto Simulations

Modelling The Muon Source : Angular distribution

- All muons must hit the detector to reduce computer time.
- Initial position of particles is set as a 8m diameter half dome around the detector.
- Initial momentum is set as "towards center of bottom of detector".

$$I_{th}(h, \theta) = \frac{I_1 e^{\left(\frac{-h_0}{\lambda_1 \cos(\theta)}\right)} + I_2 e^{\left(\frac{-h_0}{\lambda_2 \cos(\theta)}\right)}}{\cos(\theta)}$$

vertical depth h_0 , zenith angle θ ,

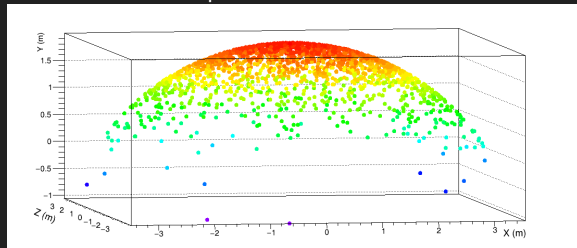
$$I_1 = (8.60 \pm 0.53) \times 10^{-6} \text{ sec}^{-1} \text{ cm}^{-2} \text{ sr}^{-1},$$

$$I_2 = (0.44 \pm 0.0) \times 10^{-6} \text{ sec}^{-1} \text{ cm}^{-2} \text{ sr}^{-1},$$

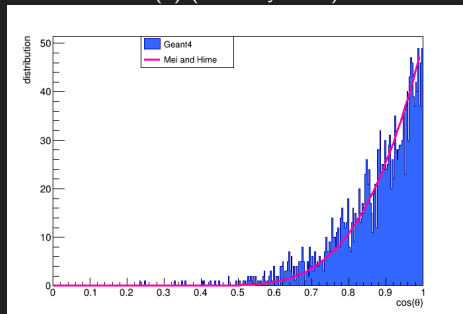
$$\lambda_1 = 0.45 \pm 0.01 \text{ km.w.e.}$$

$$\lambda_2 = 0.87 \pm 0.02 \text{ km.w.e.}$$

Initial position of 2000 muons :



$\cos(\theta)$ (arbitrary units) :



Geant4 Muon Veto Simulations

Modelling The Muon Source : Energy distribution

$$\frac{dN}{dE_\mu} = A e^{-bh(\gamma_\mu - 1)} [E_\mu + \epsilon_\mu (1 - e^{-bh})]^{-\gamma_\mu}$$

A the normalization constant with respect to the differential muon intensity at a given depth h ,

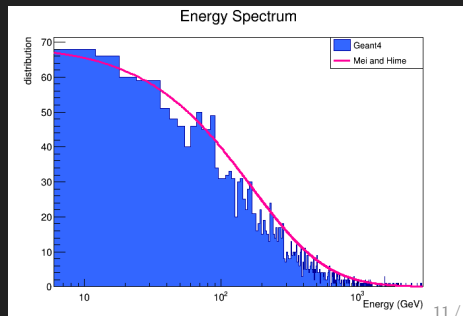
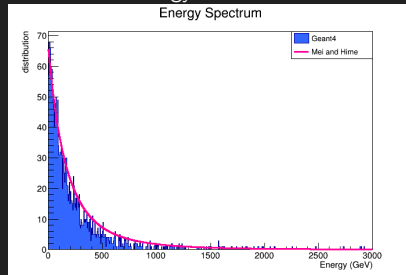
E_μ the muon energy after crossing h .

$b = 0.4/\text{km.w.e.}$,

$\gamma_\mu = 3.77$

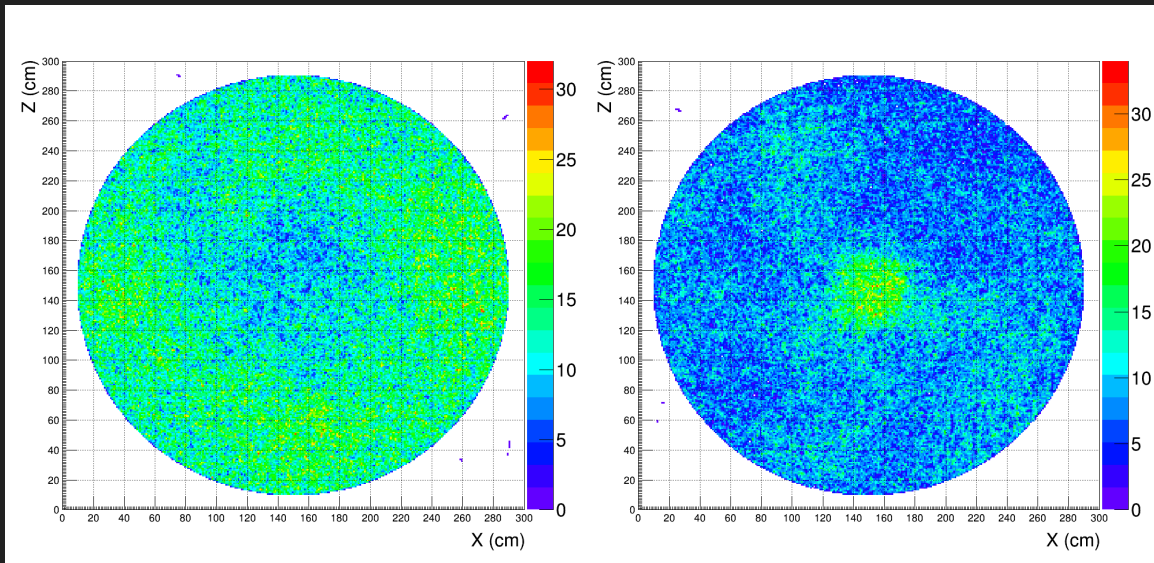
$\epsilon_\mu = 693 \text{ GeV}$

Initial Energy of 2000 muons :



Geant4 Muon Veto Simulations

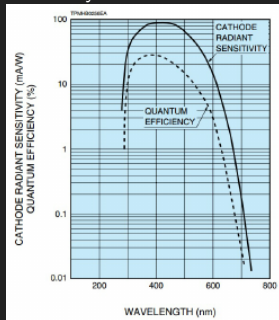
Results for the PICO-60 Veto (bottom and top scoring meshes)



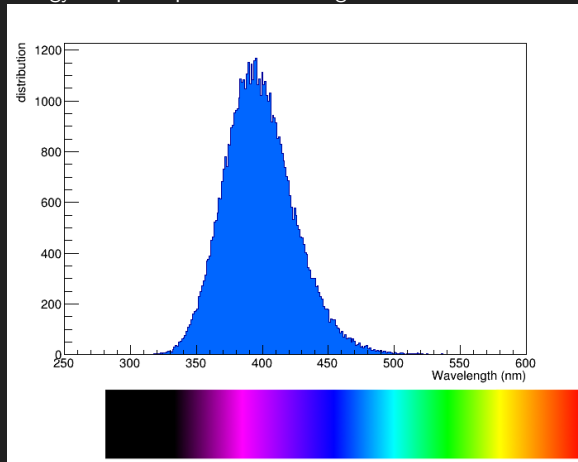
Results

Optical Photon Energy

Efficiency curve for PMT Hamamatsu :



Energy of optical photons in scoring mesh :



- Area of PMTs vs water tank $\sim 3.74\%$
- Total photons detected ~ 448 photons per muon.

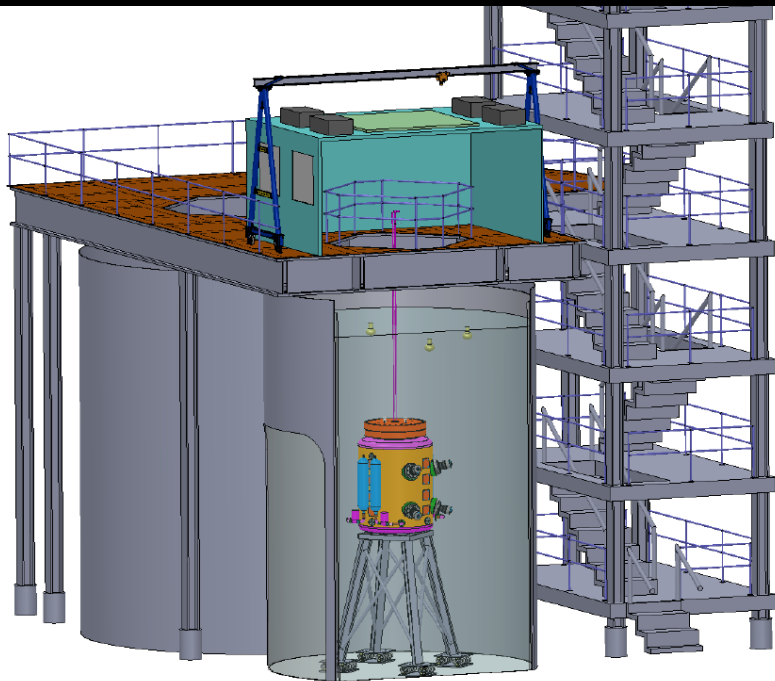
Future Veto for PICO 500

Next Steps

A geometry for a pico500 muon veto has been implemented in Geant4.

- Run simulations on SHARKNET (desktop computer takes ~ 10 h per muon)
- Extract Scoring Meshes to optimize placement of PMTs
- Try other geometries (spherical, cylindrical...)
- Test different reflectivities of tyvek and bottom of tank to maximise photon counting

Future Veto for PICO 500



Thank you for attending!

