

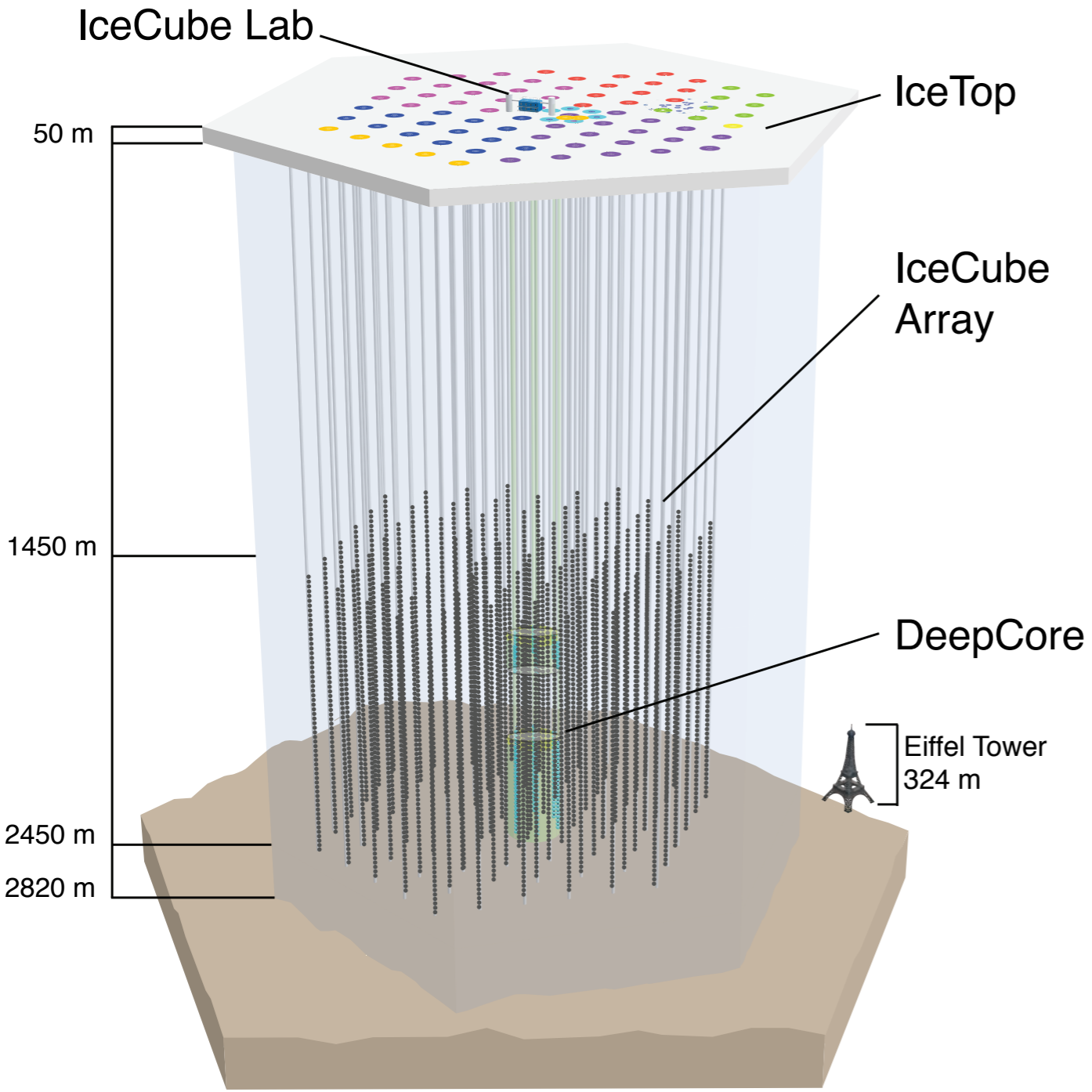
# Recent Results from IceCube

Chris Weaver for the IceCube Collaboration

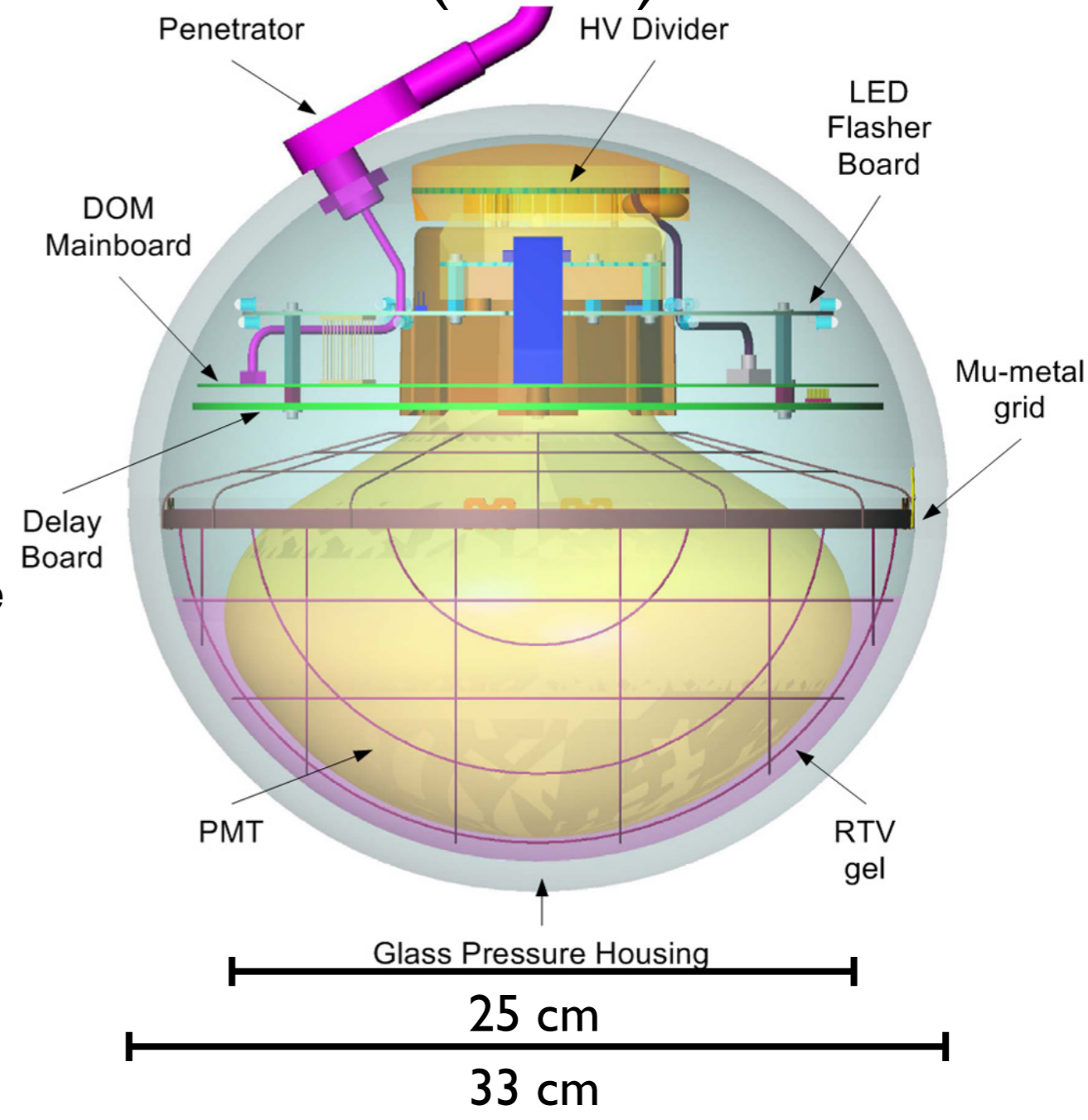


# Detector Hardware

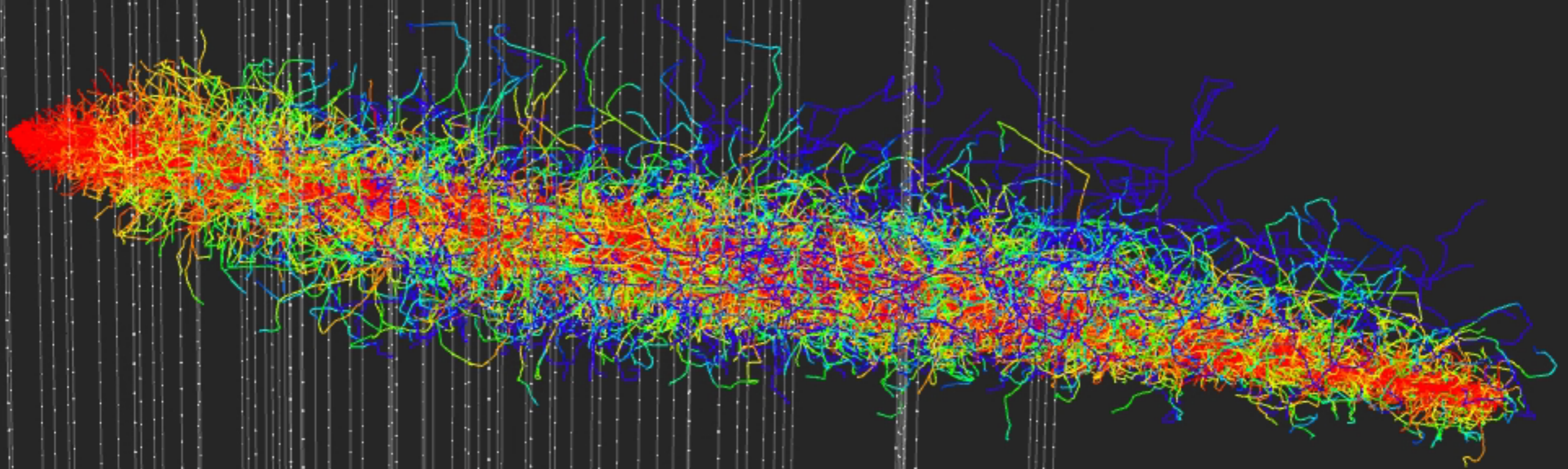
## Detector Layout



## Digital Optical Module (DOM)



# Particle Light Emission



Age

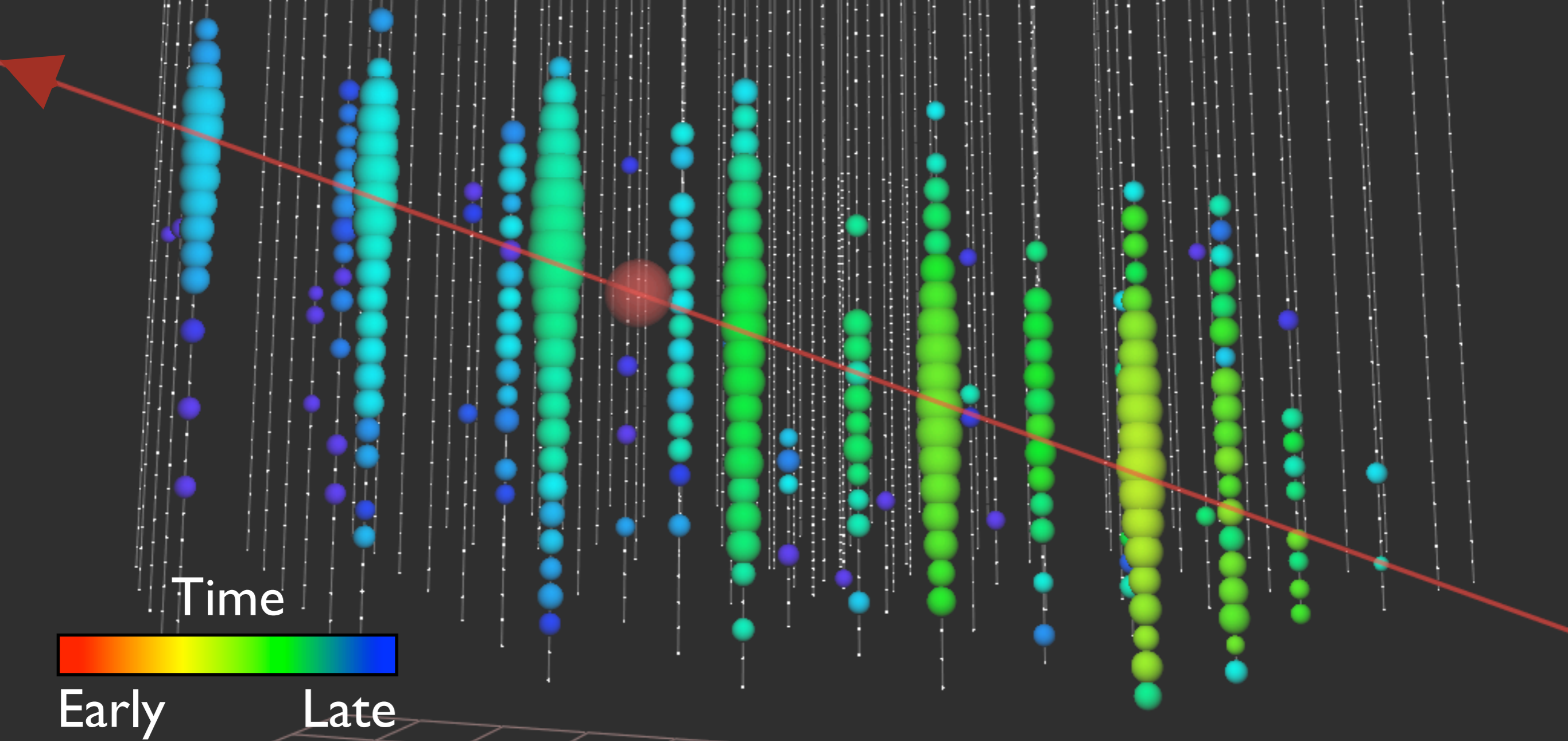


Young

Old



# Light Detection

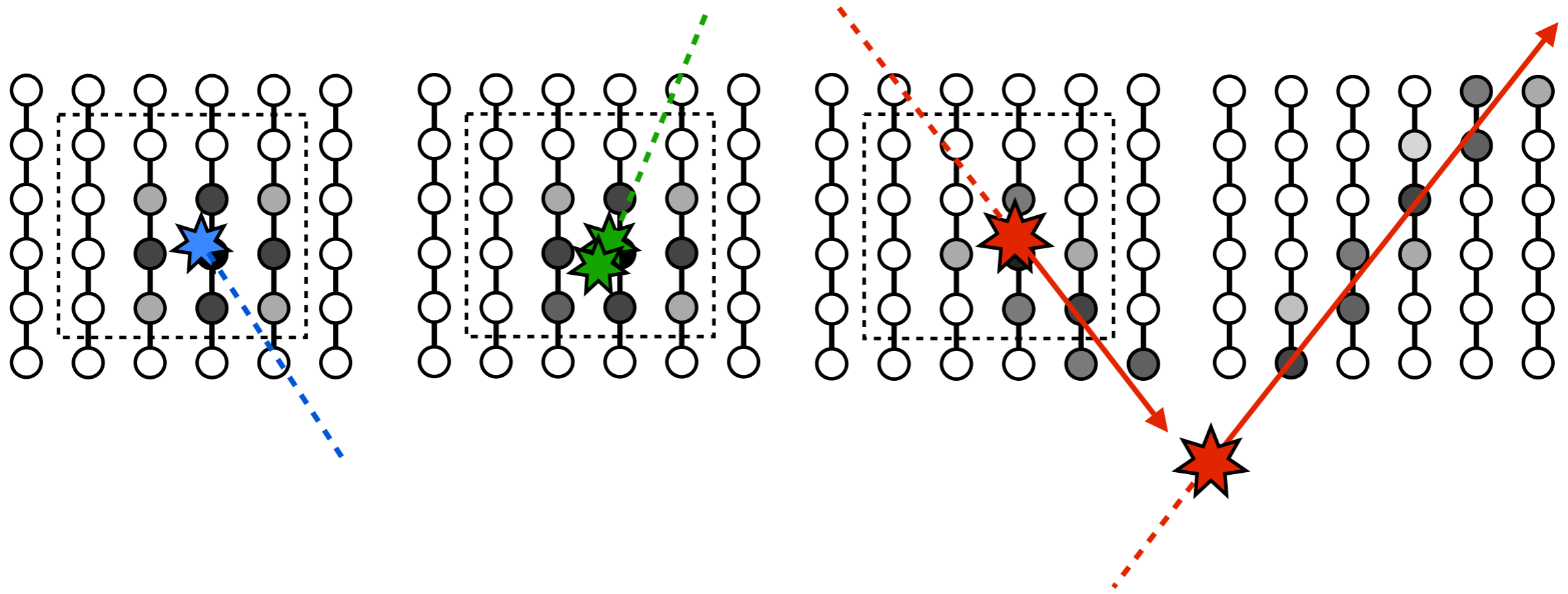


Time

Early Late

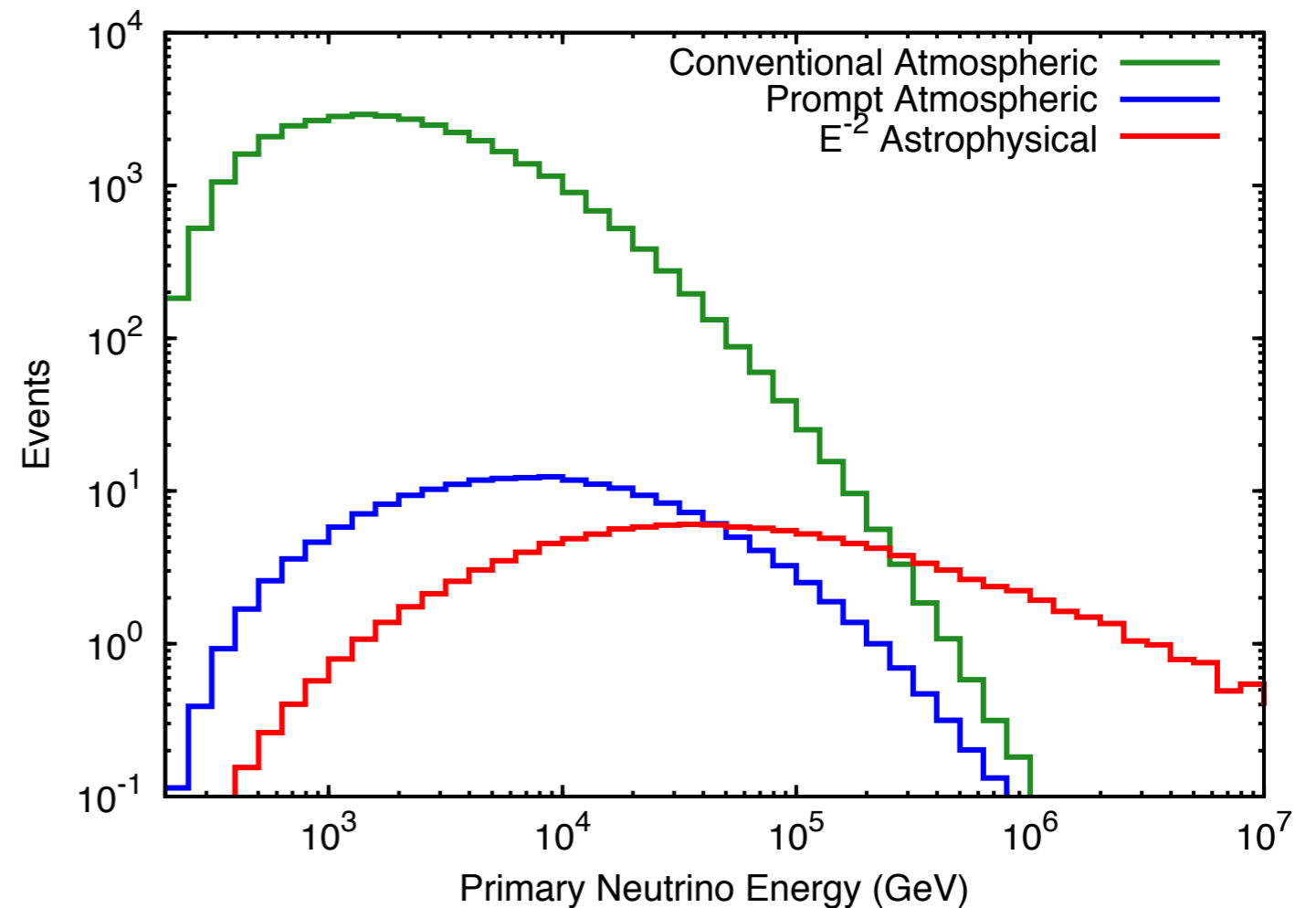
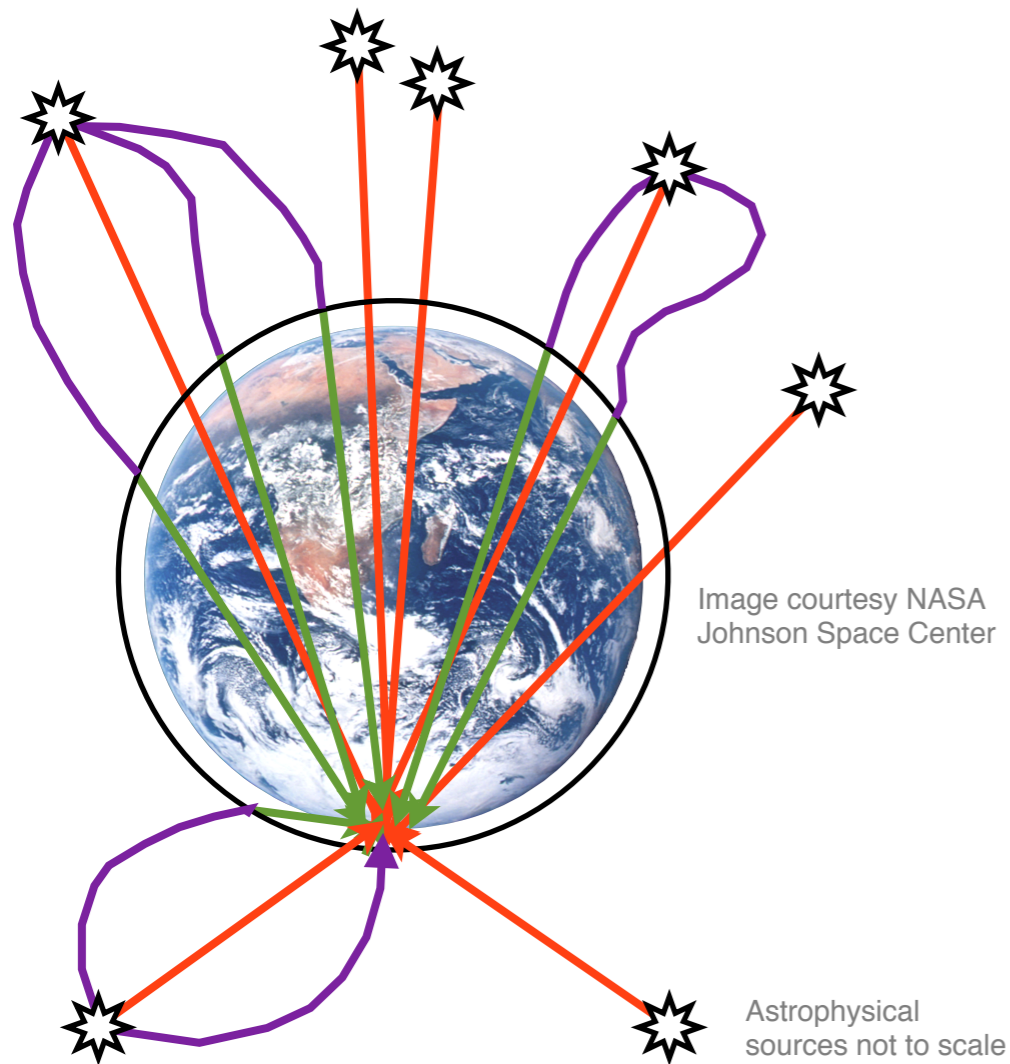


# IceCube Event Topologies



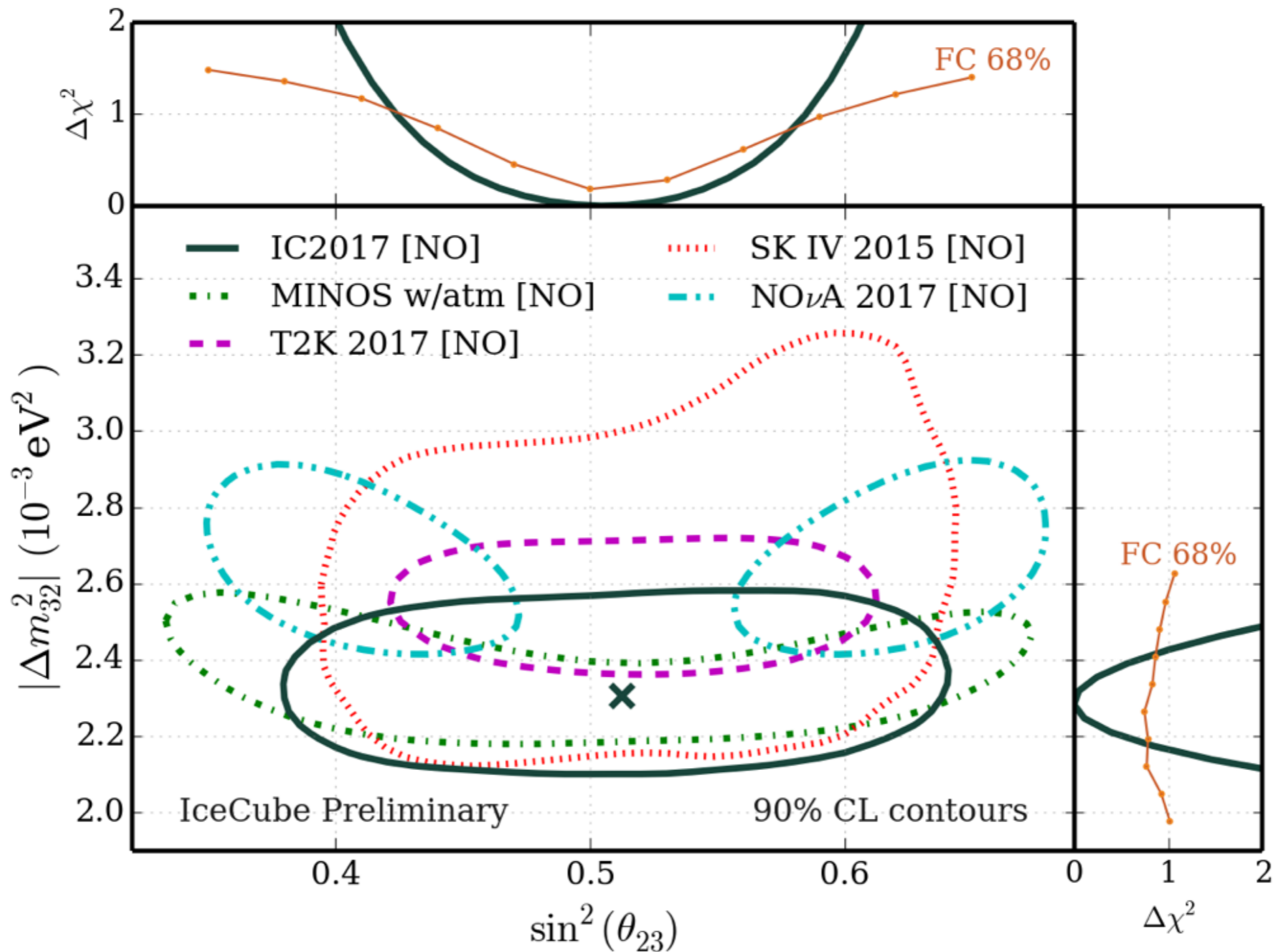
- Can select events starting in a fiducial volume or entering in a direction impossible for non-neutrinos
- Muons produce long tracks, almost everything else looks approximately point-like

# Signals in IceCube



- Cosmic ray air showers
- Atmospheric neutrinos—all directions, many ‘baselines’
- Astrophysical neutrinos

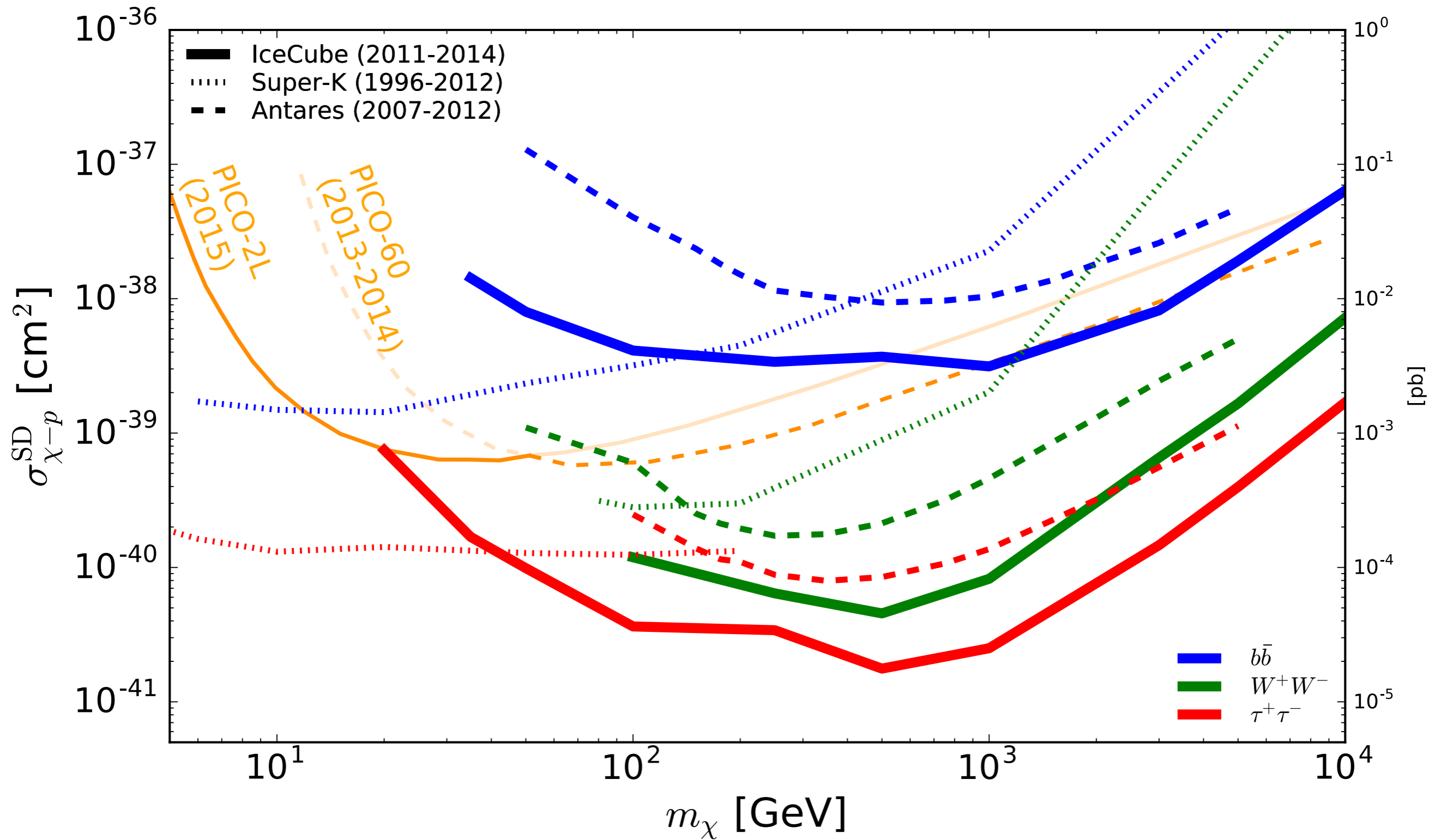
# Neutrino Oscillation



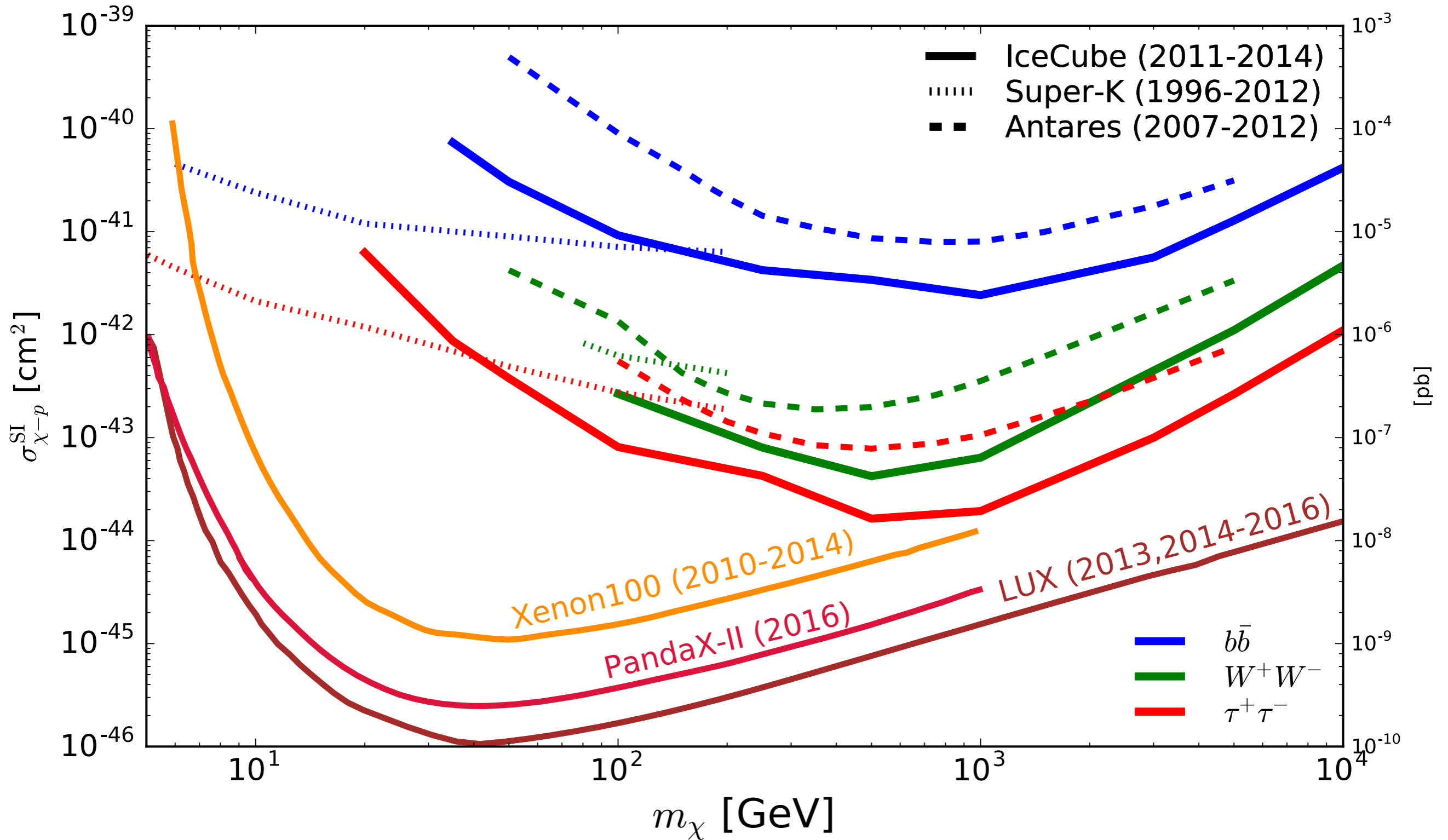
Atmospheric  $\nu_\mu$  disappearance—3 years of data  
Results very similar for Inverted ordering



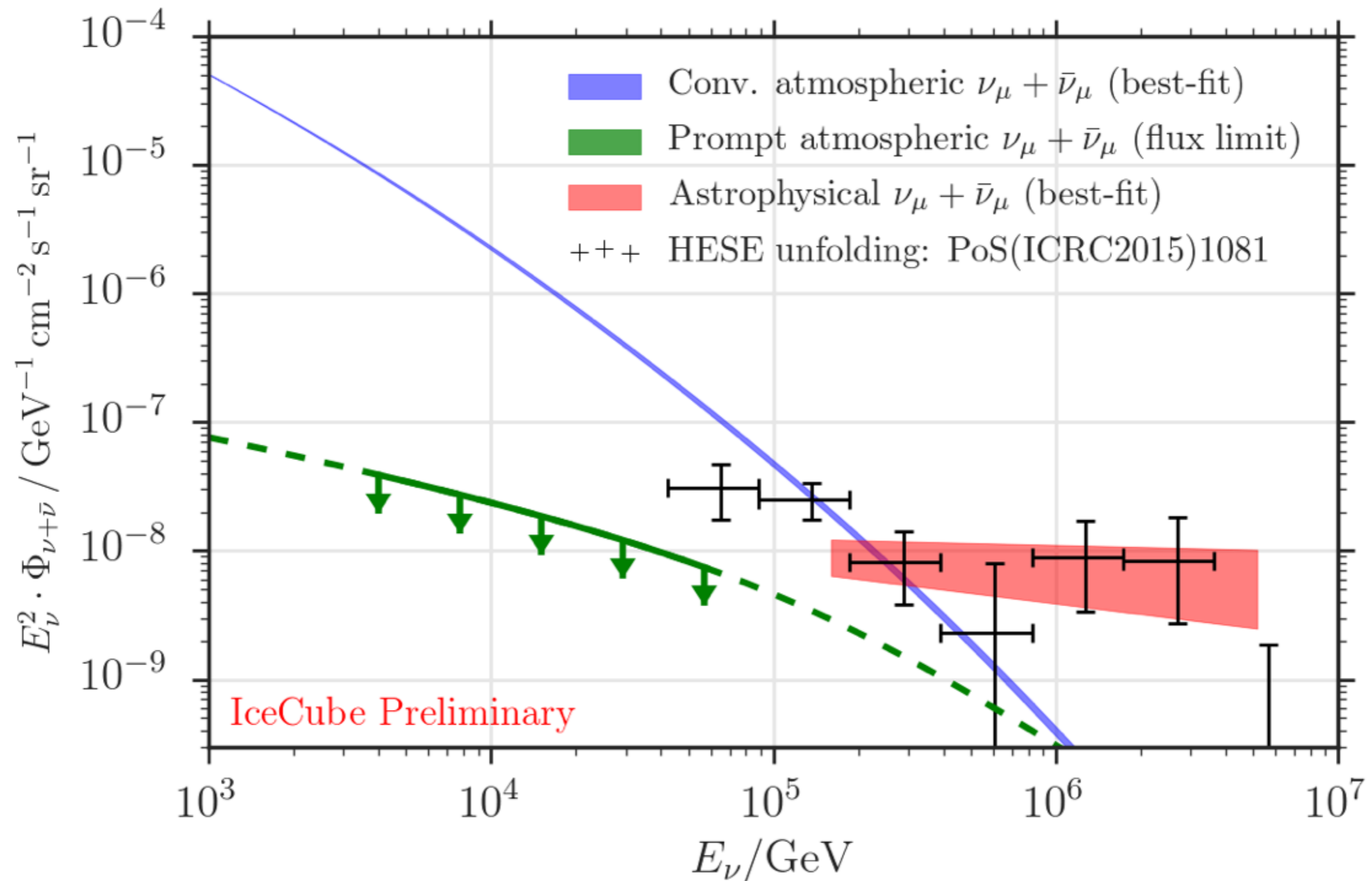
# Solar WIMP Annihilation



# Solar WIMP Annihilation



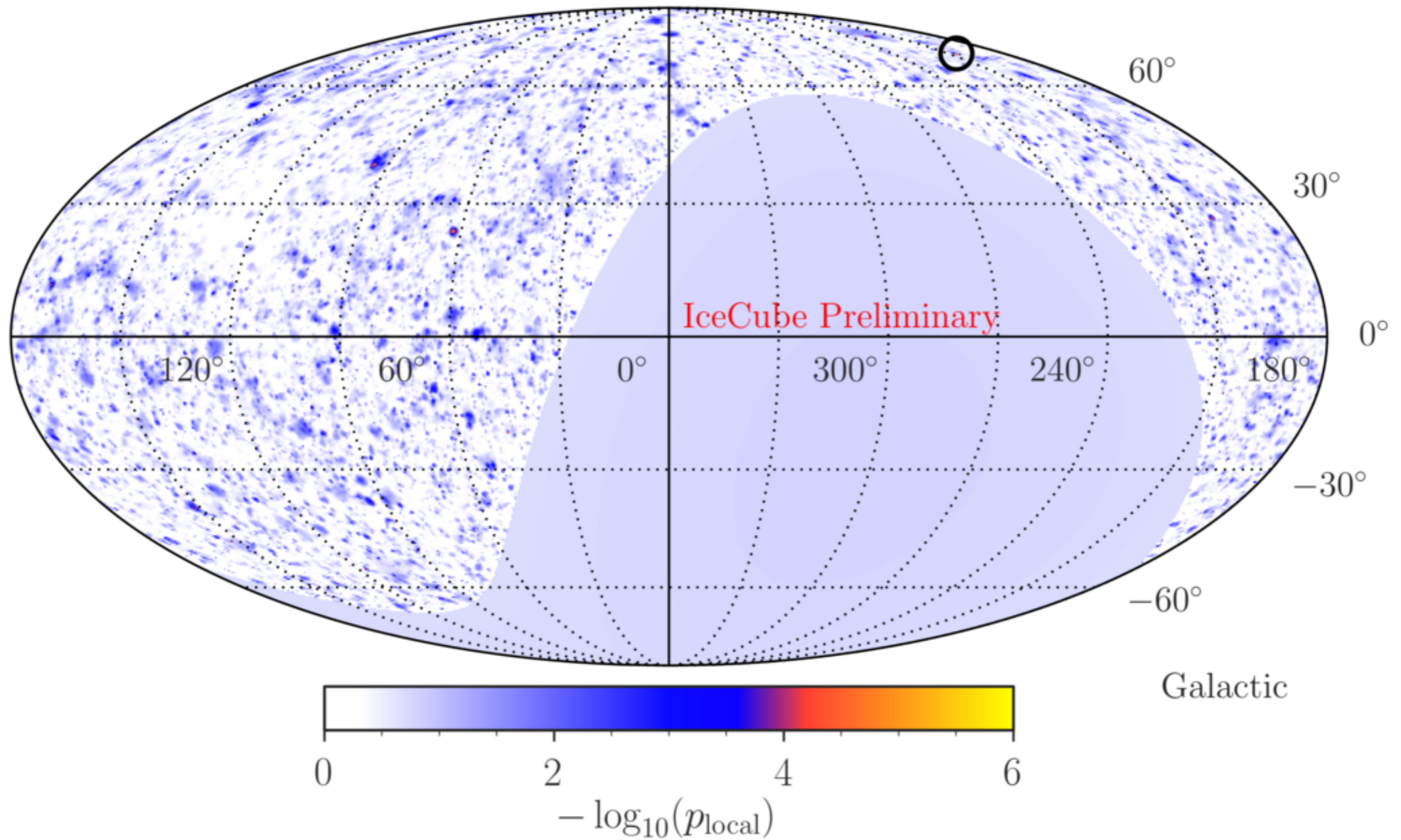
# Astrophysical Neutrino Spectrum



Now 7 years of through-going track data



# Search for Astrophysical Sources



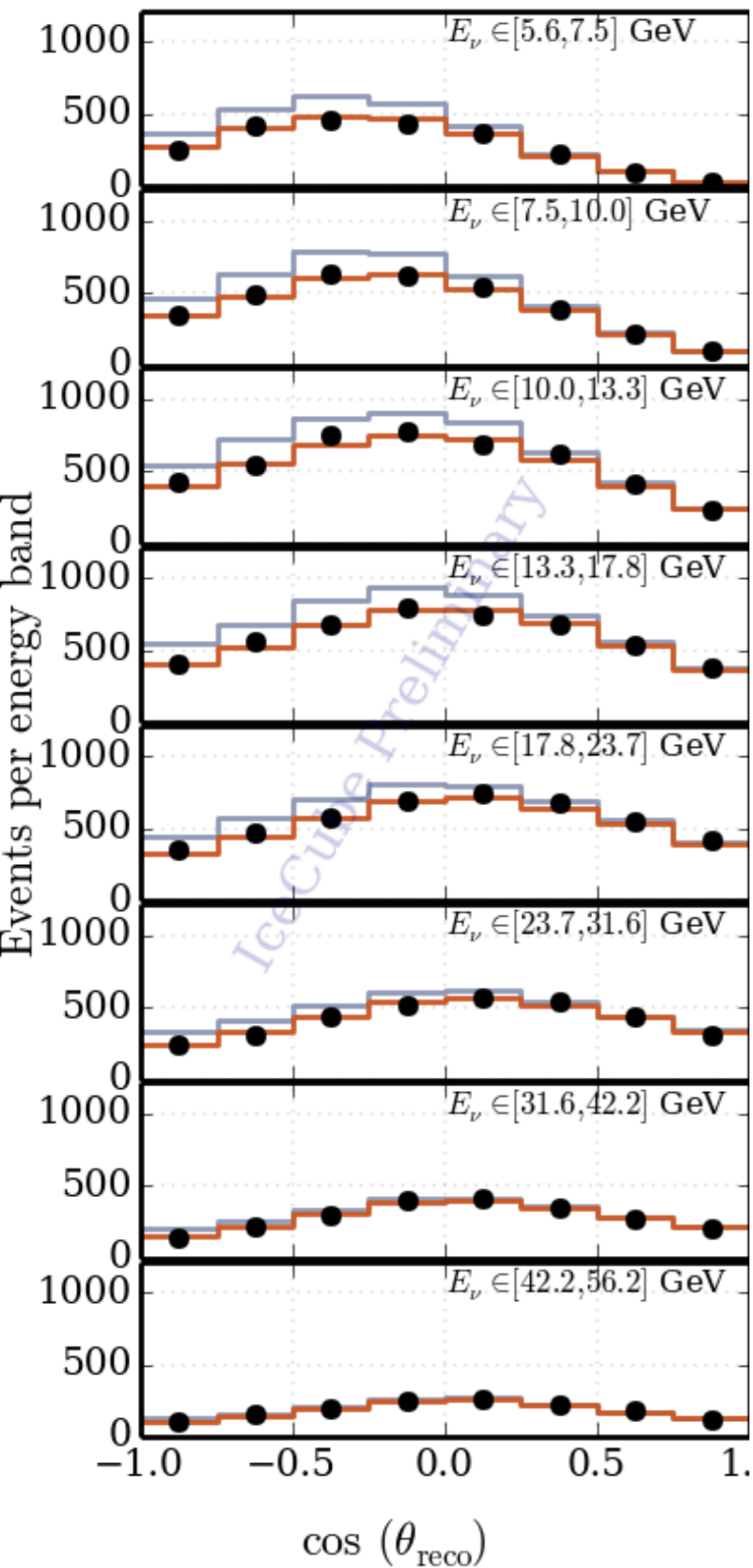
**Thank You!**  
**Questions?**





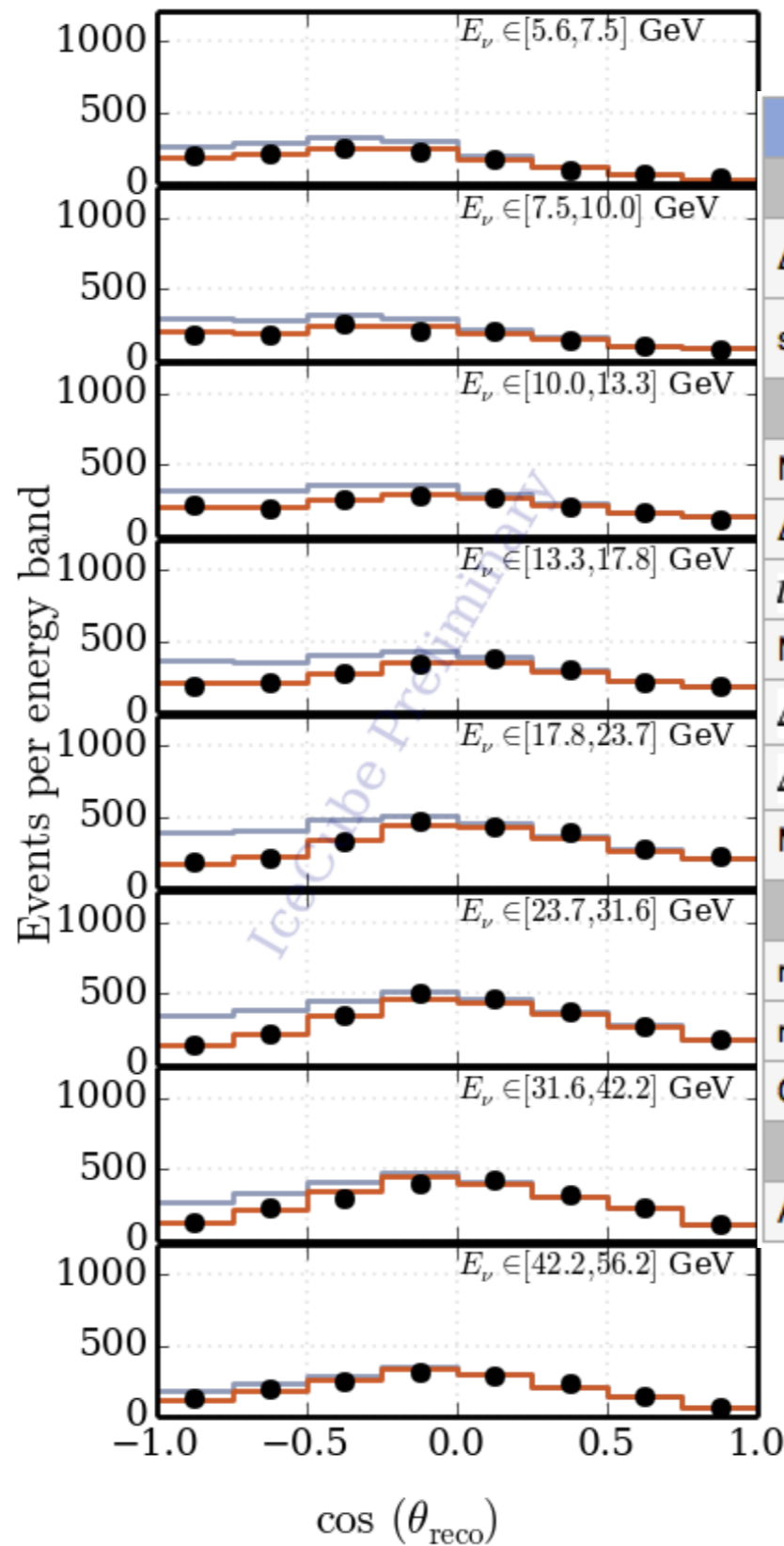
### Cascade-like

— no osc. — best fit  $\blacksquare$  data

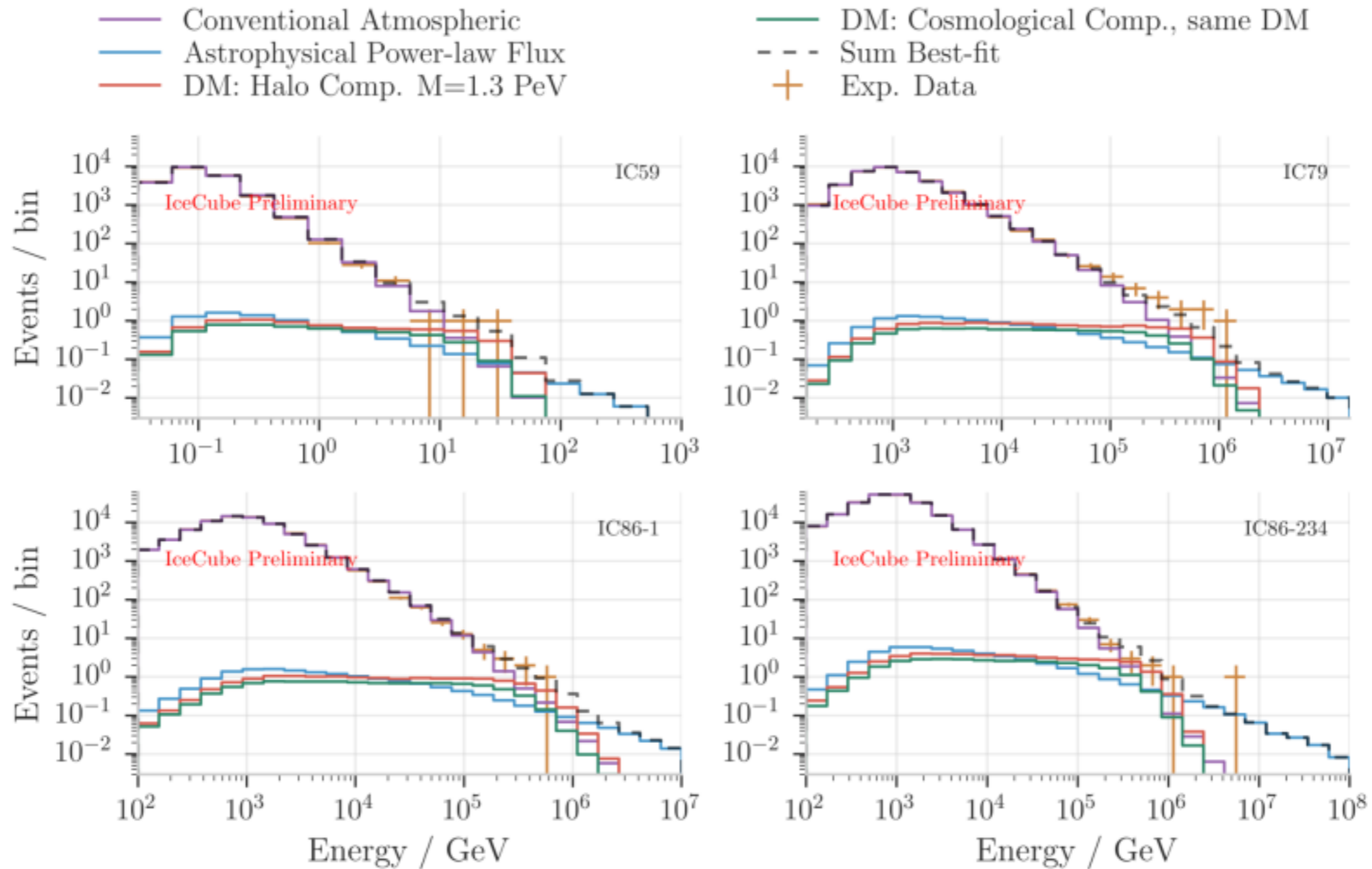


### Track-like

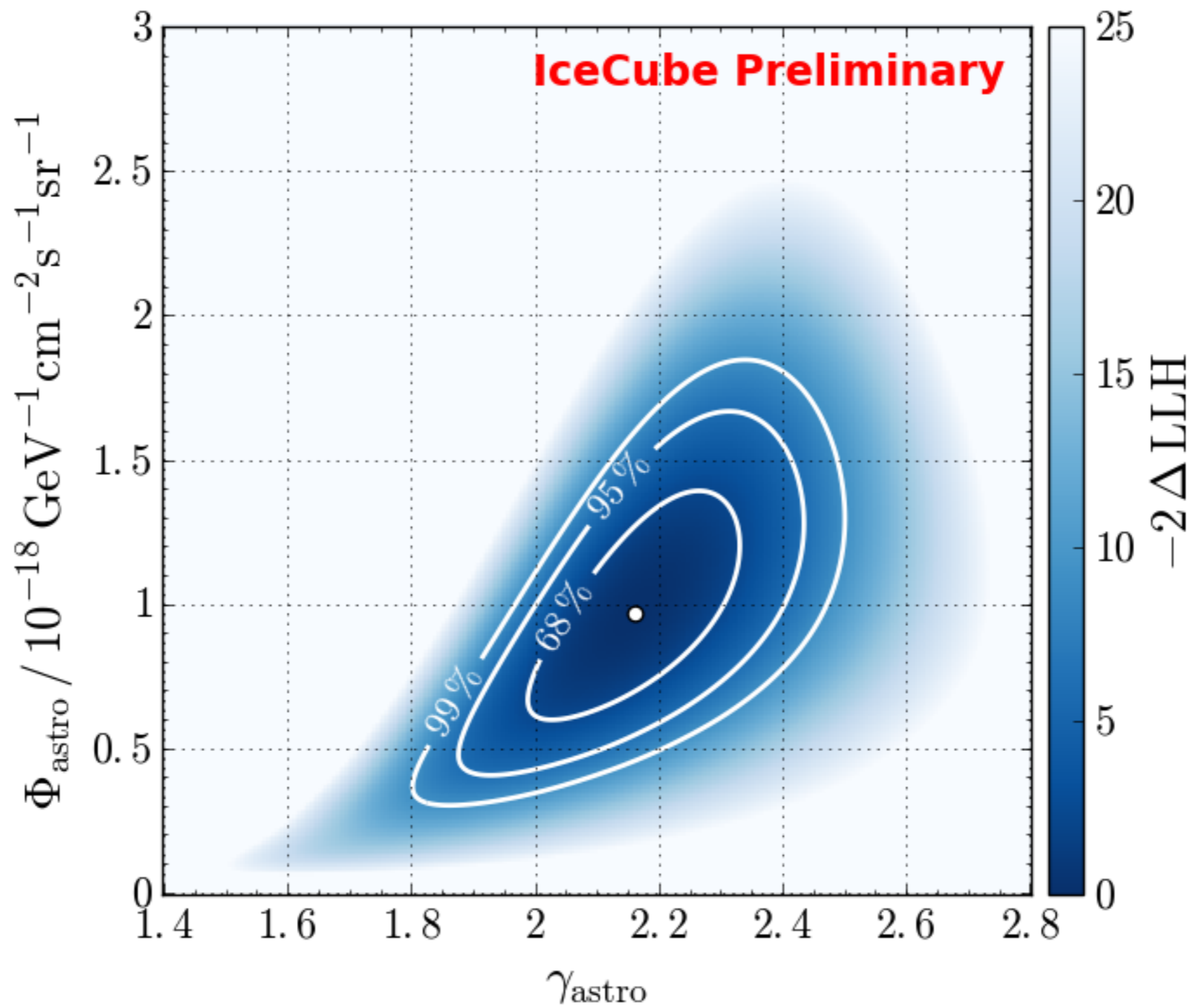
— no osc. — best fit  $\blacksquare$  data

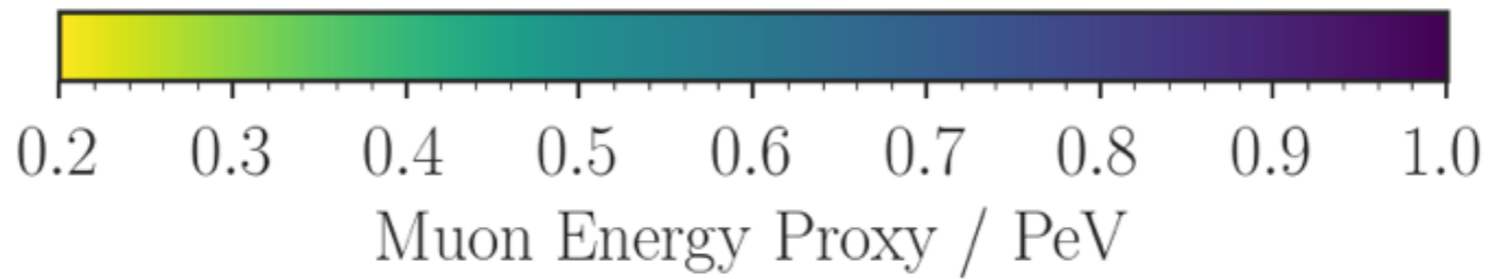
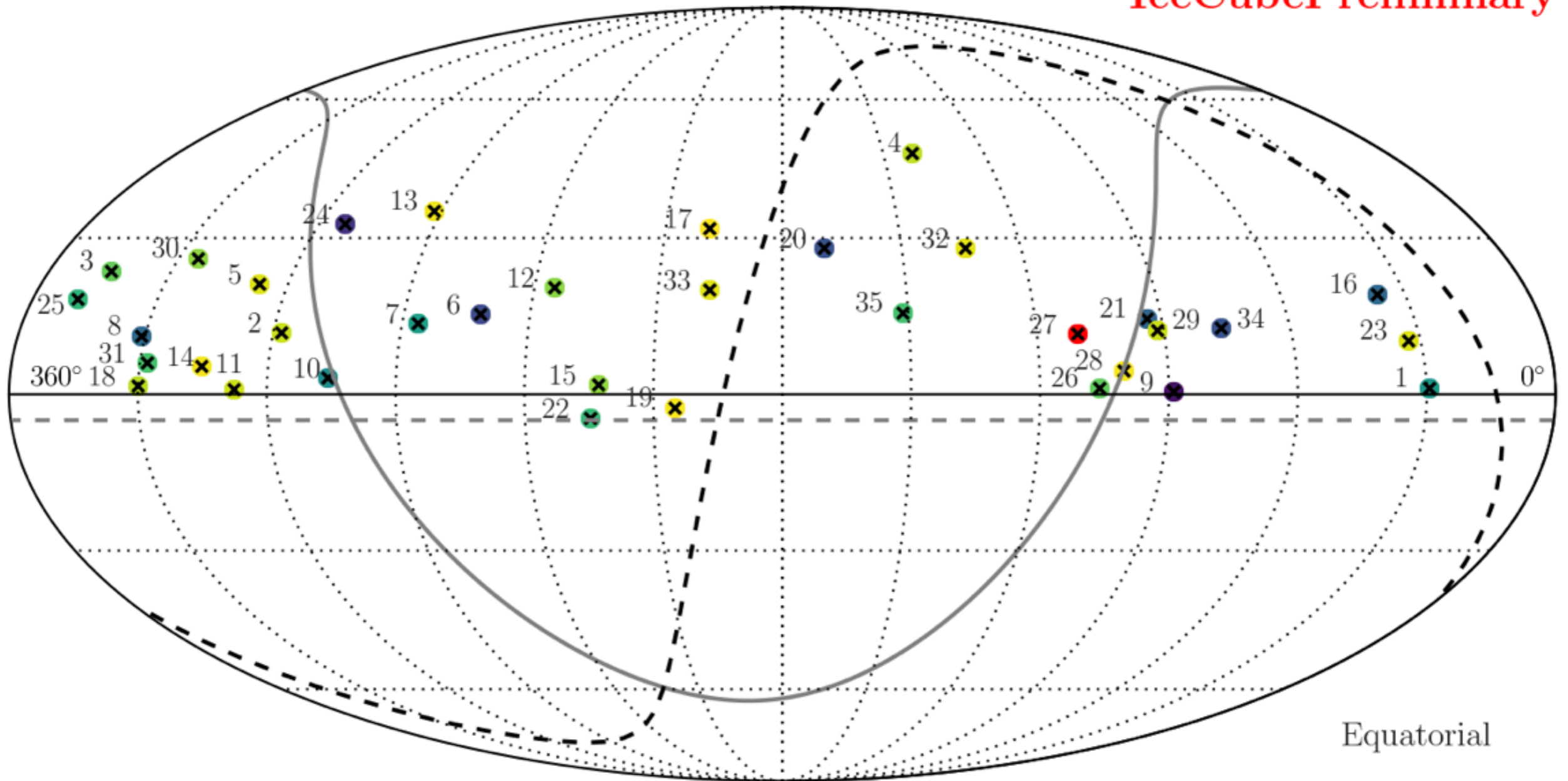


Parameter	Priors	NO
Standard mixing parameters		
$\Delta m^2_{32}$ ( $10^{-3} \text{ eV}^2/c^4$ )	no prior	$2.31^{+0.11}_{-0.13}$
$\sin^2 \theta_{23}$	no prior	$0.51^{+0.07}_{-0.09}$
Flux and cross section parameters		
Neutrino event rate [% of nominal]	no prior	85
$\Delta\gamma$ (spectral index change)	$0.00 \pm 0.10$	-0.02
$\nu_e + \bar{\nu}_e$ relative normalization (%)	$100 \pm 20$	125
NC relative normalization	$100 \pm 20$	106
$\Delta(\nu/\bar{\nu})(\sigma)$ , energy dependent	$0.00 \pm 1.00\sigma$	-0.56
$\Delta(\nu/\bar{\nu})(\sigma)$ , zenith dependent	$0.00 \pm 1.00\sigma$	-0.55
$M_A$ (resonance) (GeV)	$1.12 \pm 0.22$	0.92
Detector parameters		
relative DOM efficiency, lateral ( $\sigma$ )	$0.0 \pm 1.0$	0.2
relative DOM efficiency, head-on (a.u.)	no prior	-0.72
Overall DOM efficiency (%)	$100 \pm 10$	102
Atmospheric $\mu$ background		
Atmospheric $\mu$ contamination (% of sample)	no prior	5.5



Non-zero DM contribution does not appear to be significant;  
 only  $1.8\sigma$  when compared to background-only trials





# Galactic Halo Dark Matter

