



# BUBBLE GROWTH STUDIES FOR THE PICO EXPERIMENT

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**DETECTING DARK  
MATTER WITH BUBBLES?**

# How Did Dark Matter Come To Be?

# Galaxies Shouldn't **Shine** This Way

Light  $\neq$  Mass



Zwicky 1930's<sup>1</sup>

$$2\langle T \rangle = -U$$

Virial theorem

$$M = \frac{r\langle v^2 \rangle}{G}$$

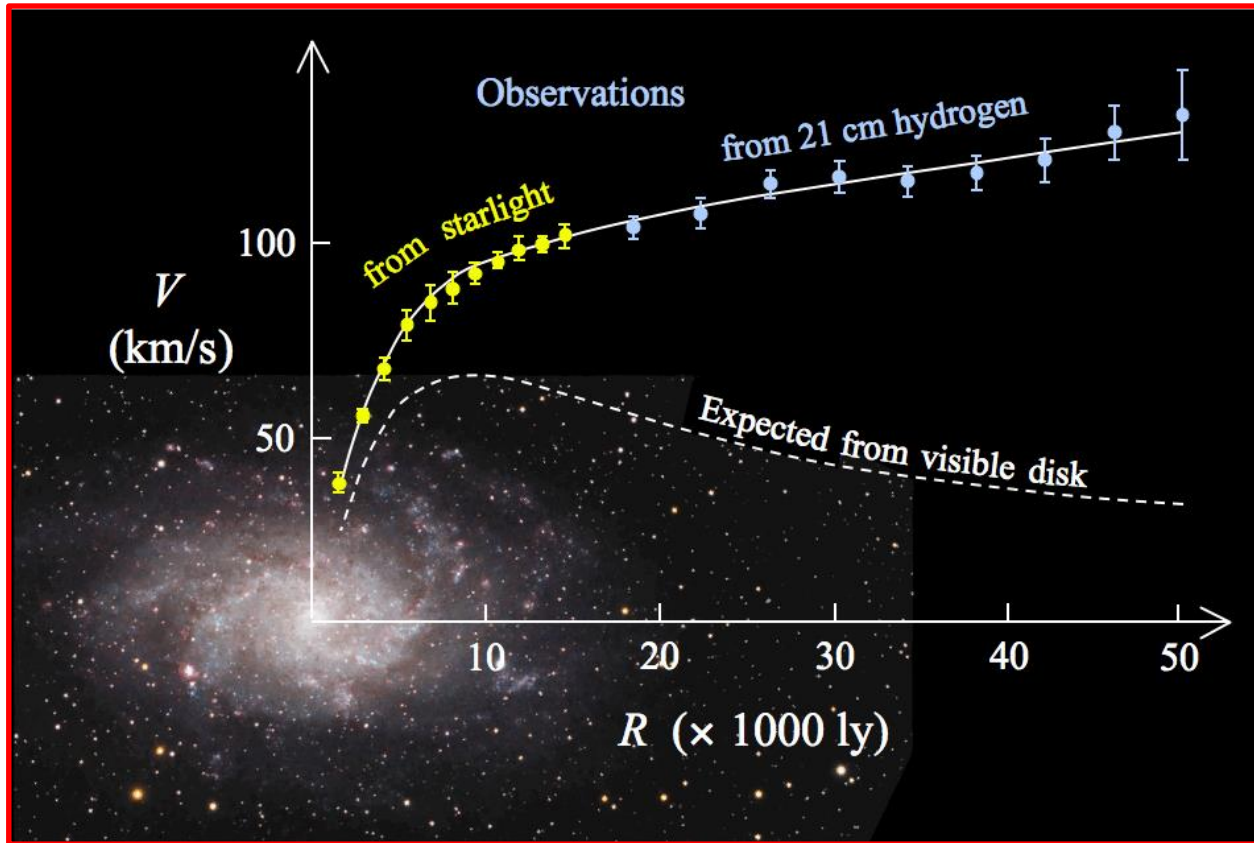
Mass of a cluster

$$L \approx M^{2.25}$$

Luminosity of a cluster

M was off by a factor of **400!**

# Stars Shouldn't **Move** This Way

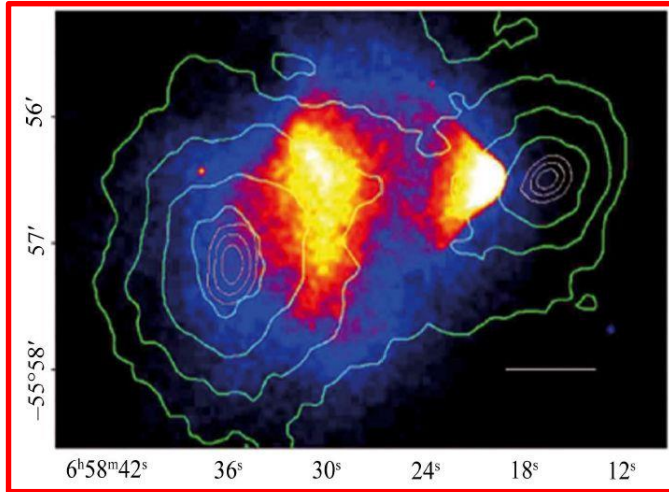


Vera Rubin 1970's<sup>2</sup>

Image, Corbelli, E. & Salucci, P. (2000). "The extended rotation curve and the dark matter halo of M33". [Monthly Notices of the Royal Astronomical Society](#). 311 (2): 441-447

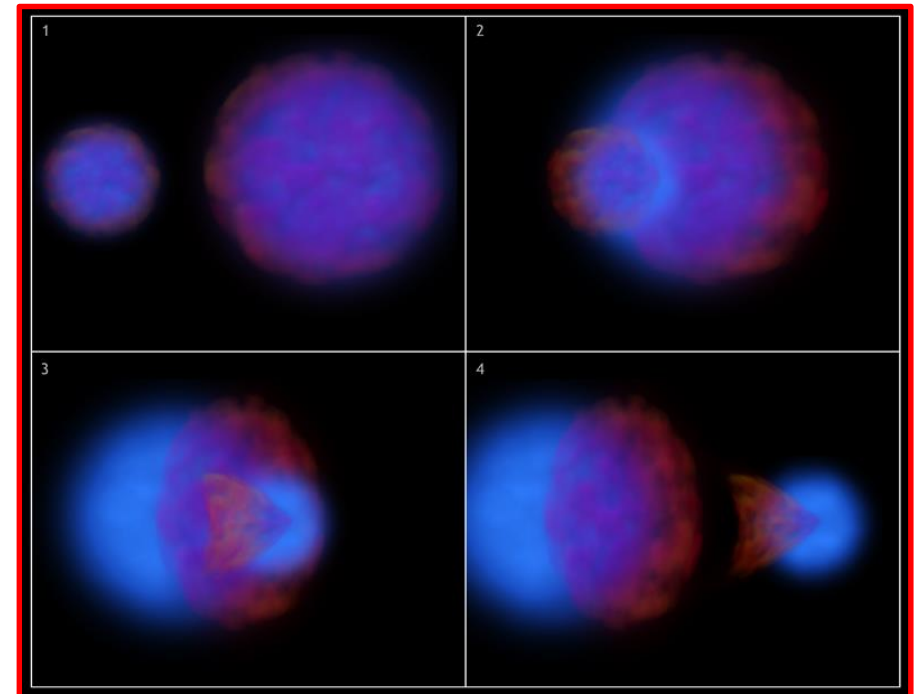
# Galaxies Shouldn't Collide This Way

Bullet Cluster<sup>3</sup>



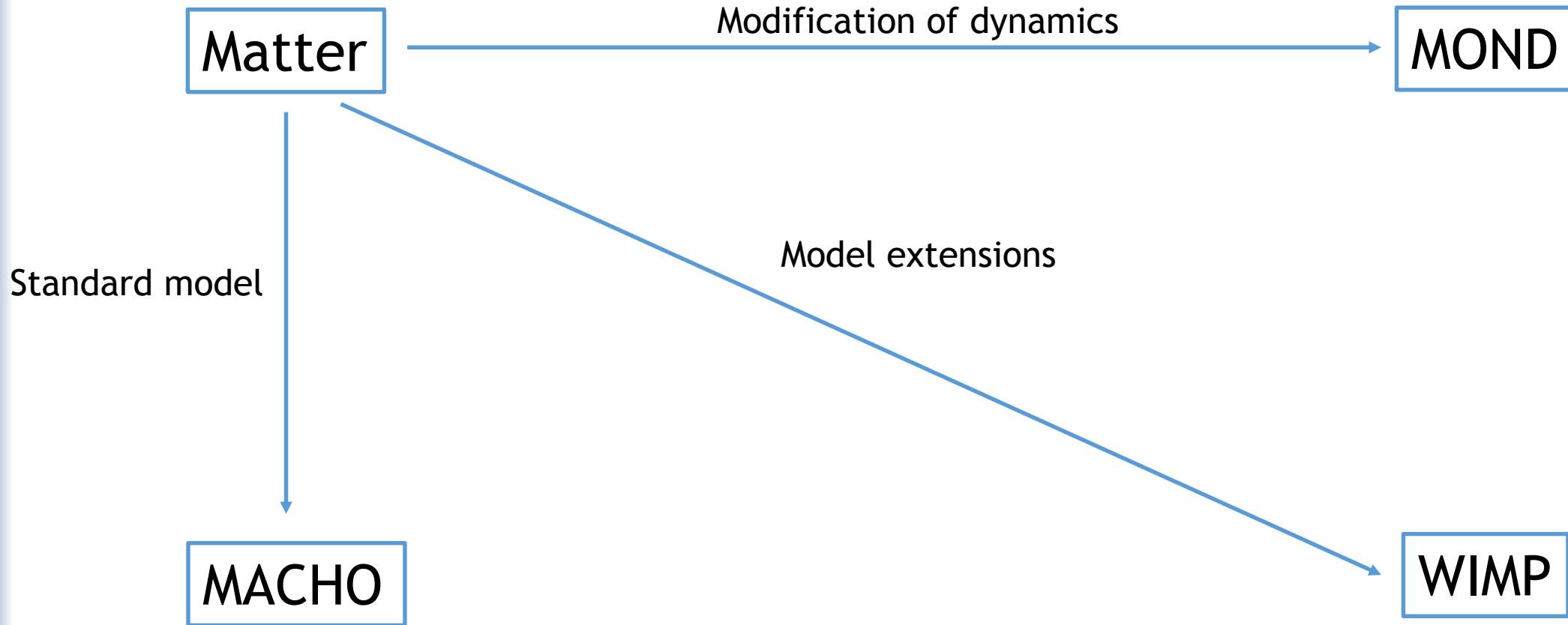
Equivalent to

Bullet Cluster Collision Simulation<sup>4</sup>



# Detection Philosophy

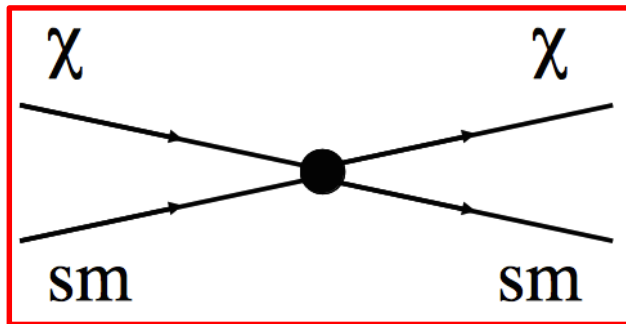
# How Do We Decide What's This Missing Thing



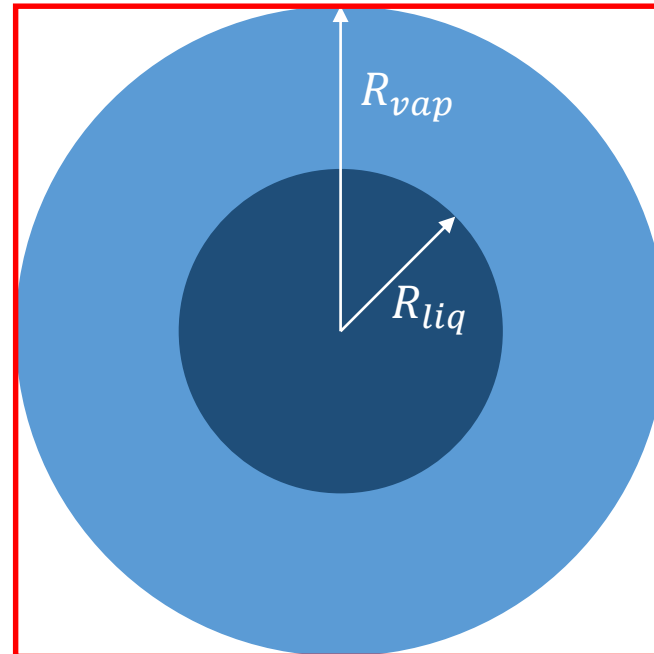


# The **PICO** Detection Principle

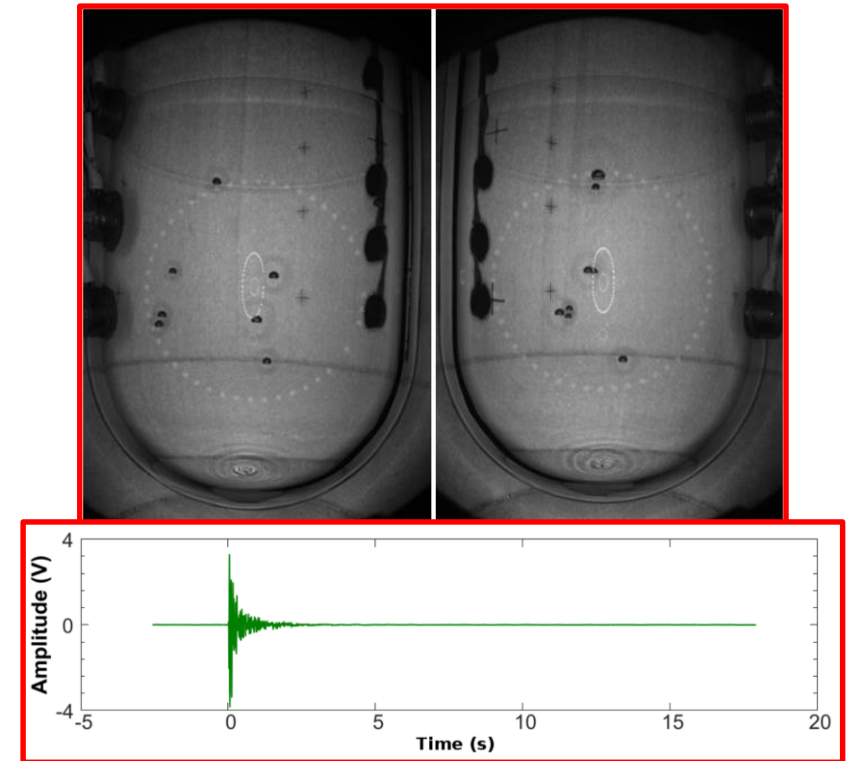
WIMP Interaction



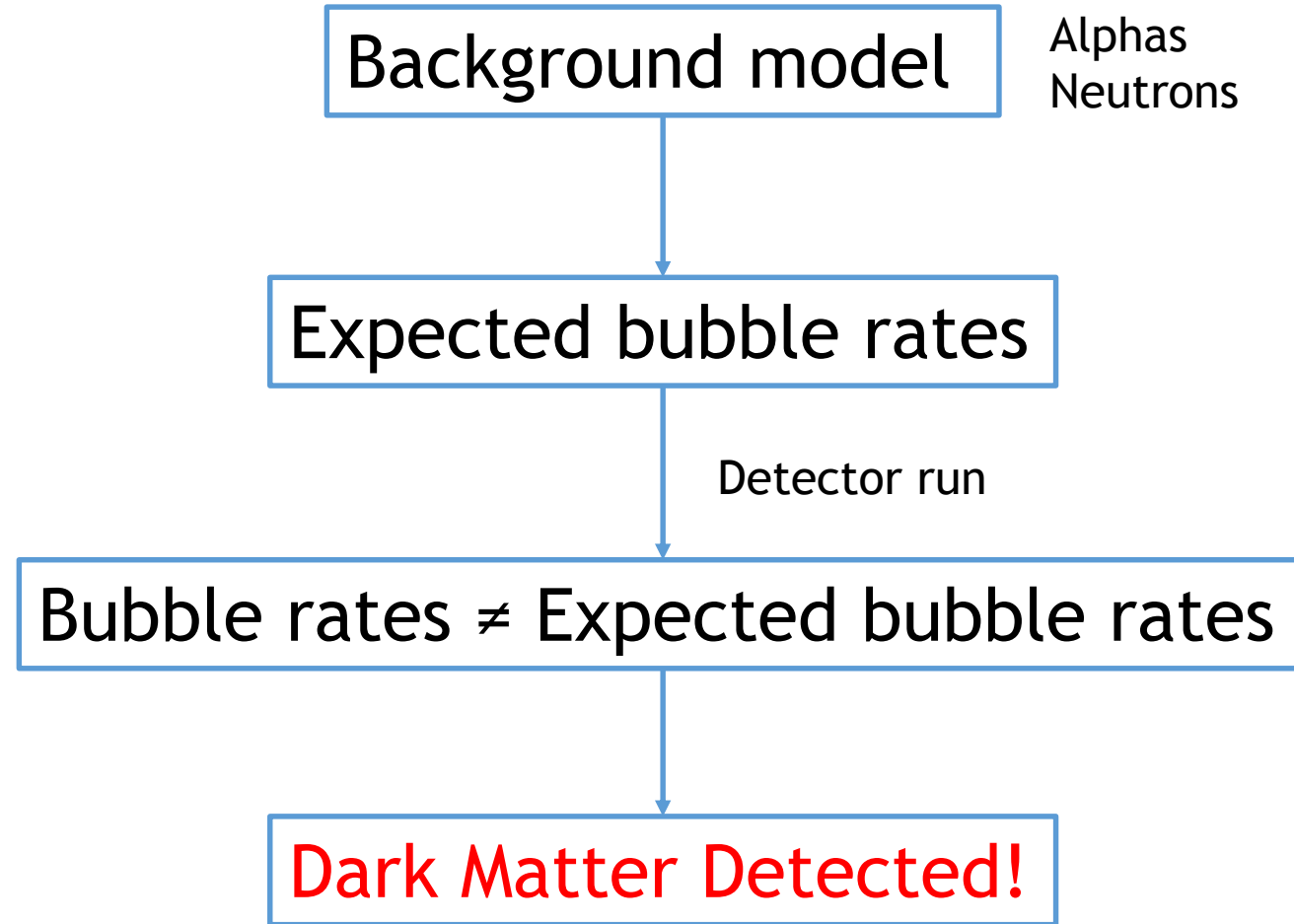
Phase Transition



Acoustic Signal

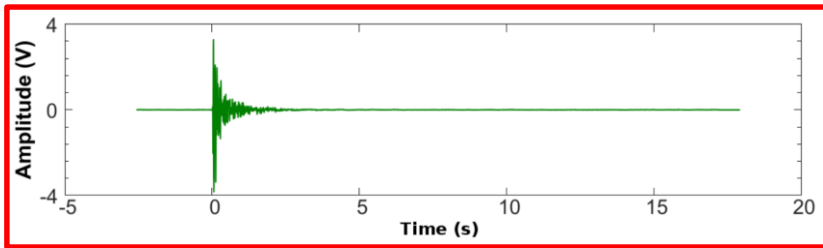


# The **PICO** Detection Principle



# Background Discrimination: AP

Acoustic signal

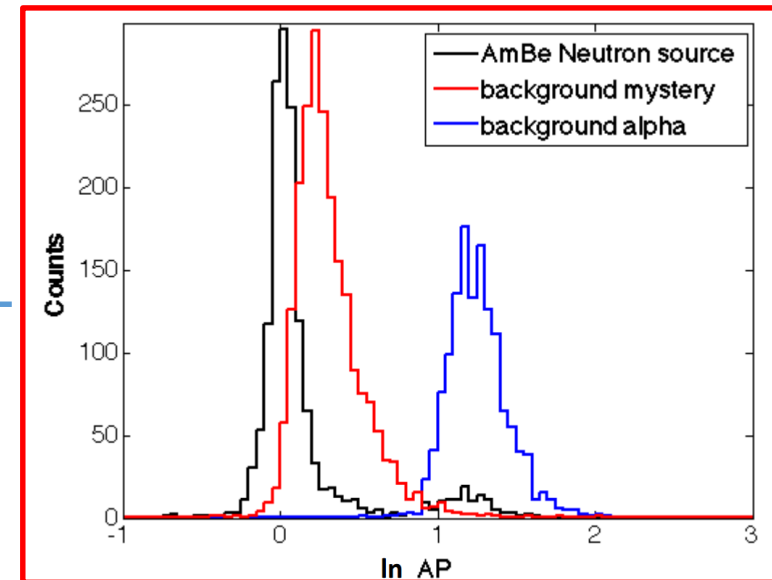


AP Calculation Tool

Output value

Histogram of values

Neutrons and alphas are different!



**What Am I Doing?**

# Bubble Growth

DISCRIMINATION TOOL

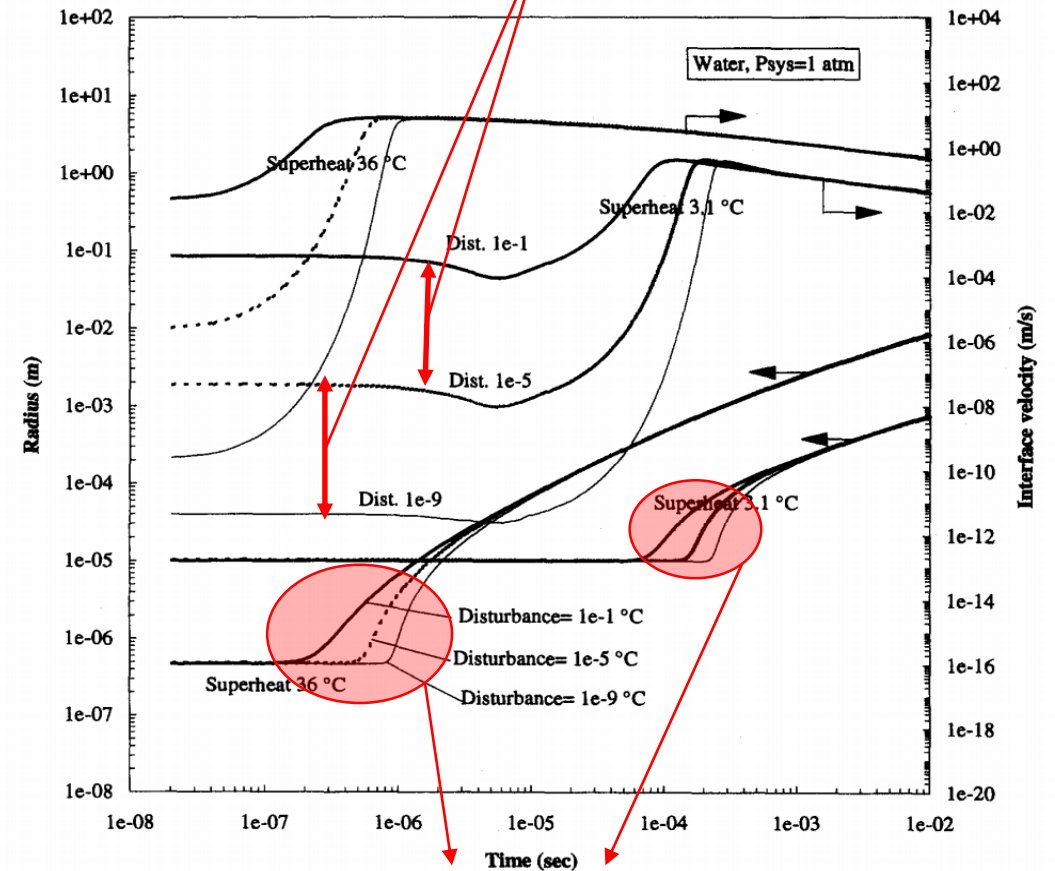
## Governing Equations

$$R \frac{d^2R}{dt^2} + \frac{3}{2} \left( \frac{dR}{dt} \right)^2 = \frac{P'(T) - P_\infty(T)}{\rho''} - \frac{2\sigma(T)}{\rho''R} - \frac{4\mu}{\rho''R} \frac{dR}{dt}$$

$$\frac{\partial T}{\partial t} + \frac{R^2}{r^2} \frac{dR}{dt} \nabla_s T = \alpha \nabla_s^2 T$$

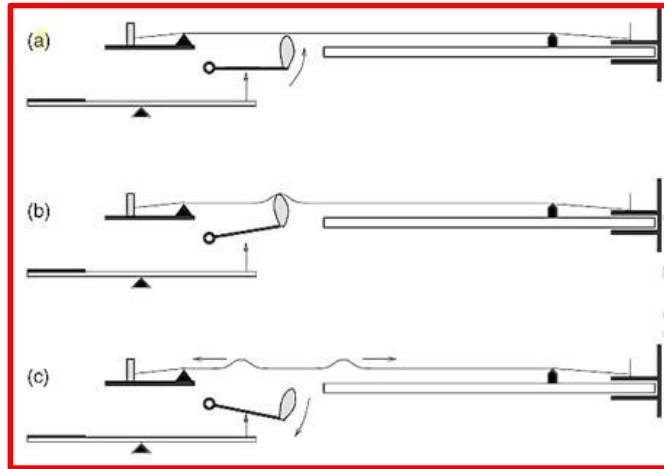
Spherical vapor bubble growth in superheated liquids

2431



IMPORTANT PART OF BUBBLE GROWTH

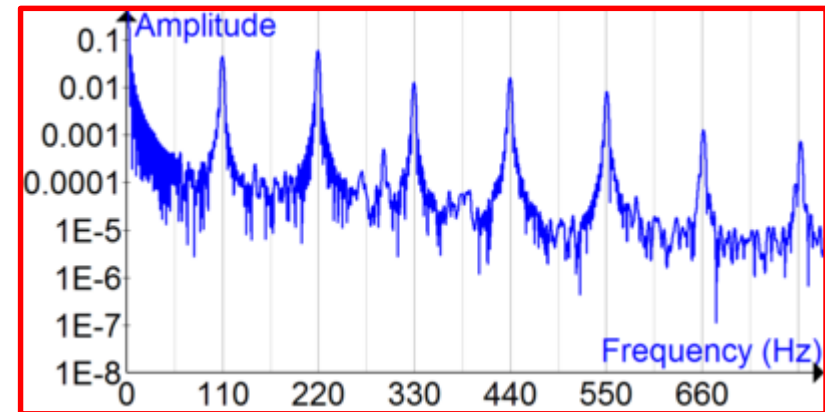
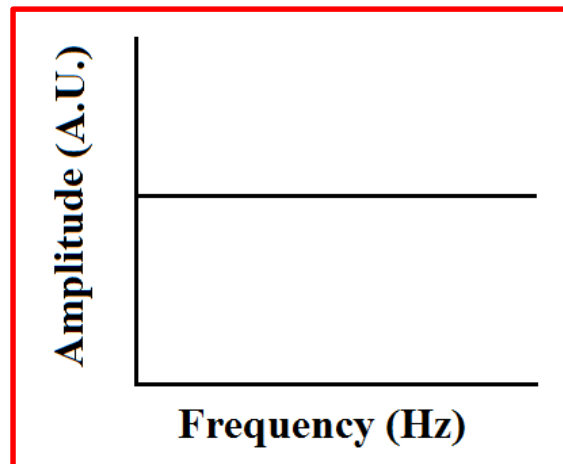
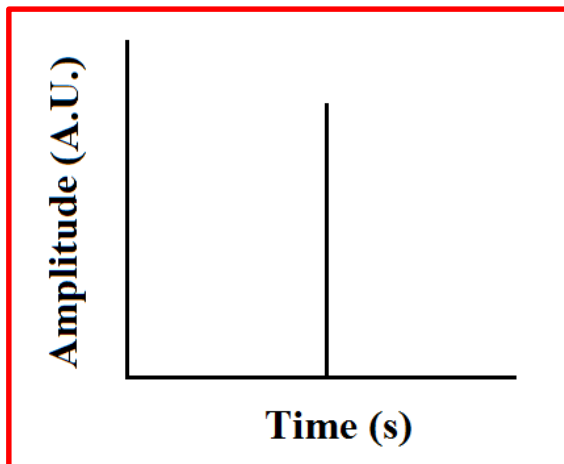
# How do You Make an **Acoustic Signal**?



Sound you can hear!

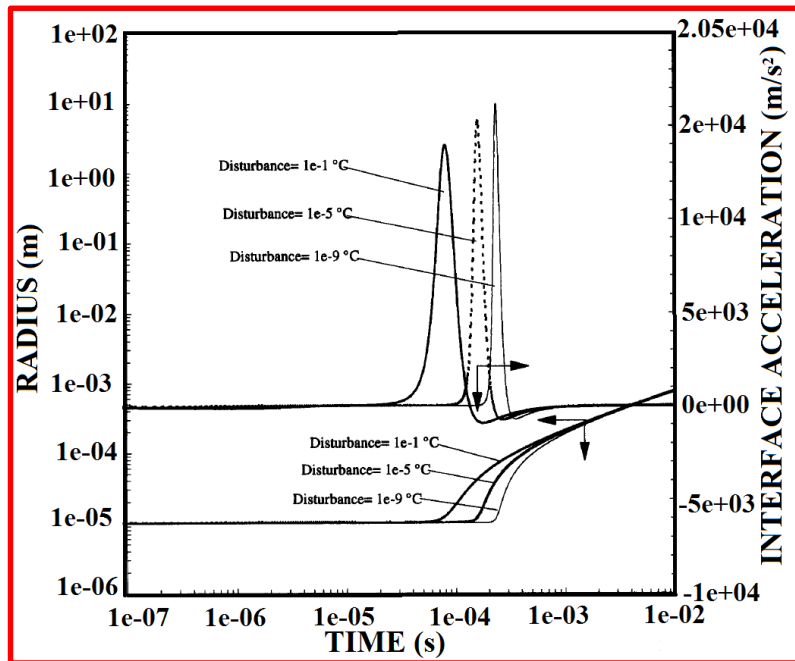
Fourier Transform

A string frequency response



# Acceleration Curves' Analysis

Acceleration  $\xrightarrow{\text{Is equal to}}$  Force  $\xleftarrow{\text{Is equivalent to}}$  Striking a piano string



Fourier Transform  
of acceleration curves

Key signatures  
in frequency

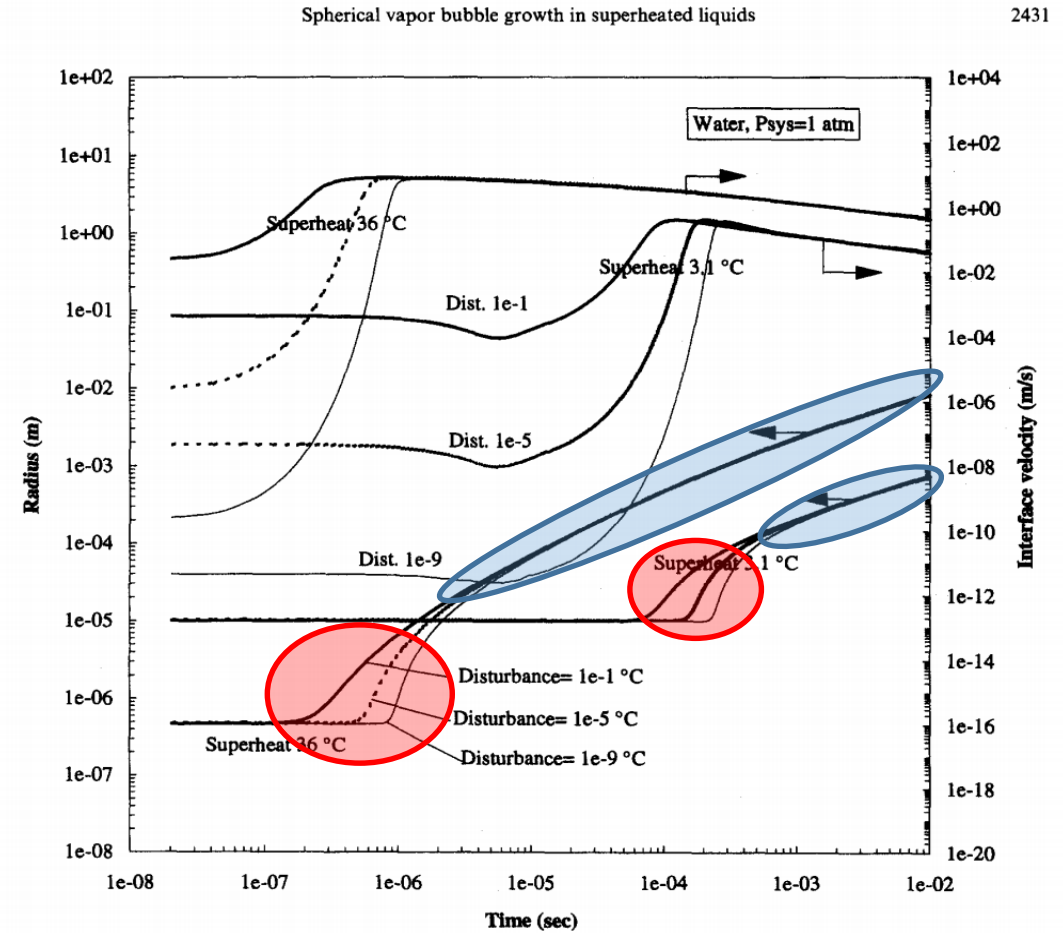
Improved AP

Image, H. S. Lee and H. Merte, "Spherical vapour bubble growth in uniformly superheated liquids", *Int. J. Heat Mass Transfer* Vol. 39 No. 12, 2427-2447 (1996)

# Optical Discrimination

Blue: curves are degenerate

Red: curves are non-degenerate!



Image, H. S. Lee and H. Merte, "Spherical vapour bubble growth in uniformly superheated liquids", *Int. J. Heat Mass Transfer* Vol. 39 No. 12, 2427-2447 (1996)

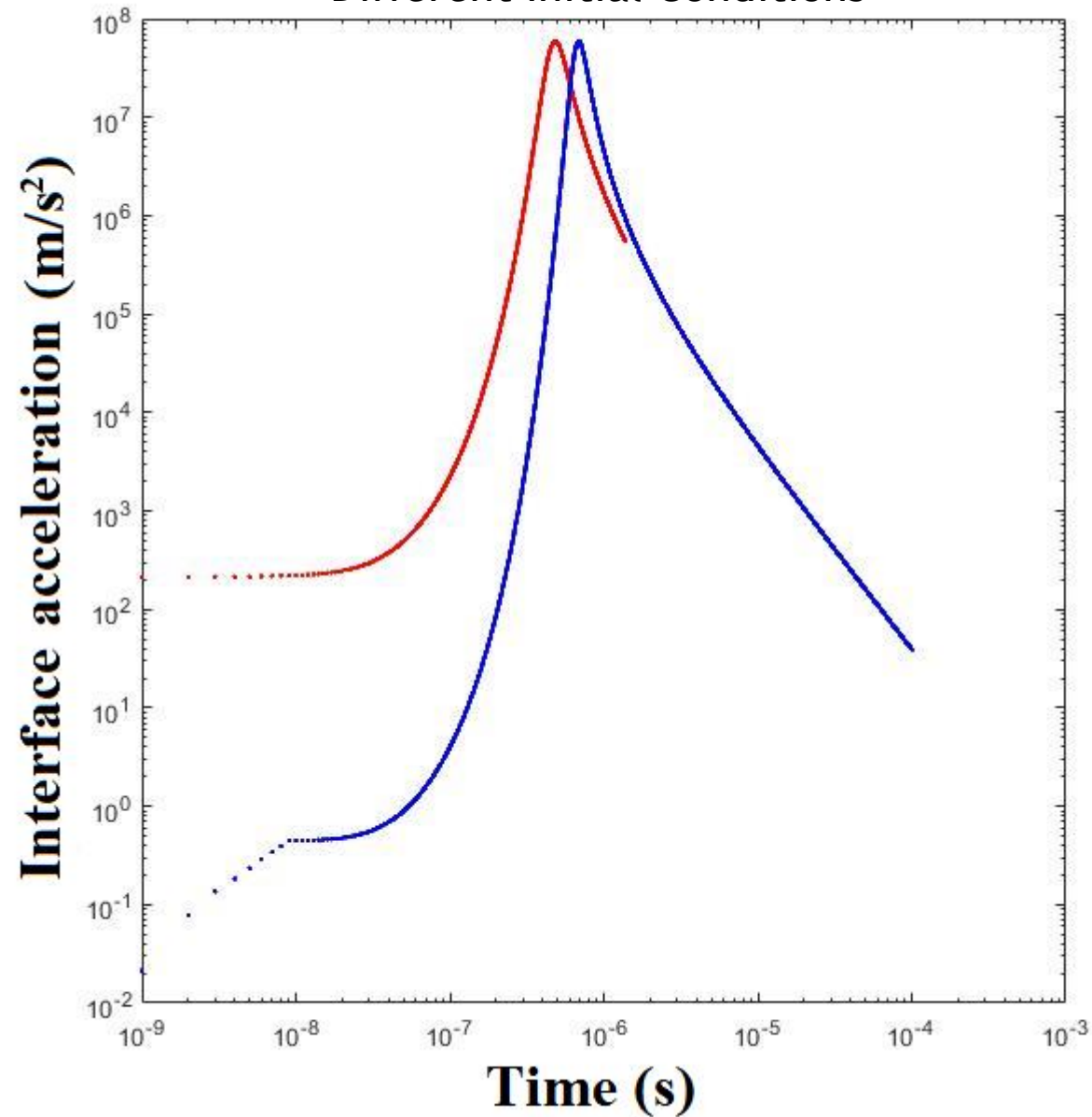


# THANKS!

## References/Acknowledgements:

1. [https://ned.ipac.caltech.edu/level5/March02/Abell/Abell3\\_3.html](https://ned.ipac.caltech.edu/level5/March02/Abell/Abell3_3.html)
2. Corbelli, E. & Salucci, P. (2000). "The extended rotation curve and the dark matter halo of M33". [Monthly Notices of the Royal Astronomical Society](#). **311** (2): 441-447
- 3.1 (Top) Deep Chandra image of the Bullet cluster. Shown in green are mass contours from weak lensing - reconstruction. From Clowe et al. 2006.
- 3.2 (Bot) X-ray: NASA/CXC/CfA/M.Markevitch et al.; Optical: NASA/STScI; Magellan/U.Arizona/D.Clowe et al.; Lensing Map: NASA/STScI; ESO WFI; Magellan/U.Arizona/D.Clowe et al.
4. NASA/CXC/M. Weiss
5. H. S. Lee and H. Merte, "Spherical vapour bubble growth in uniformly superheated liquids", *Int. J. Heat Mass Transfer* Vol. 39 No. 12, 2427-2447 (1996)

## Interface Acceleration Curves for Different Initial Conditions



In red: temperature disturbance of  $10^{-5} \text{ }^\circ\text{C}$   
In blue: temperature disturbance of  $10^{-9} \text{ }^\circ\text{C}$

36 °C Superheated water at 1 atmosphere