

Experimental Test of the Leptonic Mixing (PMNS) matrix

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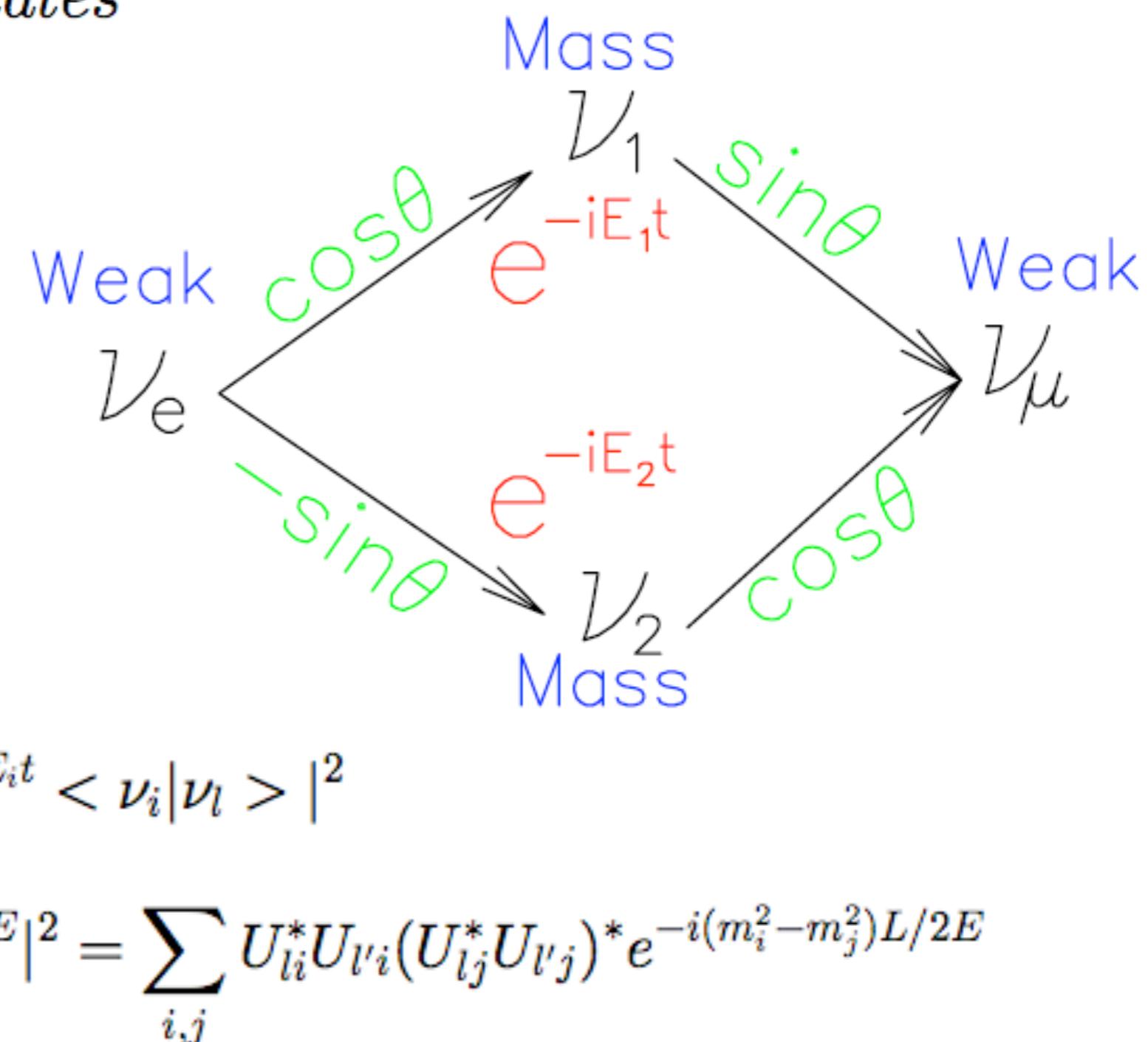
Neutrino Oscillation

$\langle \nu_l | \nu_i \rangle = U_{li}$: PMNS matrix (Unitary matrix)

$l = e, \mu, \tau$: weak eigenstates

$i = 1, 2, 3$: mass eigenstates

Neutrinos are produced and detected as weak eigenstates and travels as mass eigenstates:
Interference



$$\begin{aligned}
 P_{l \rightarrow l'} &= |\langle \nu_{l'}(t) | \nu_l(t) \rangle|^2 \\
 &= \left| \sum_i \langle \nu_{l'} | \nu_i \rangle e^{-iE_i t} \langle \nu_i | \nu_l \rangle \right|^2 \\
 &= \left| \sum_i U_{li}^* U_{l'i} e^{-im_i^2 L/2E} \right|^2 = \sum_{i,j} U_{li}^* U_{l'i} (U_{lj}^* U_{l'j})^* e^{-i(m_i^2 - m_j^2)L/2E}
 \end{aligned}$$

Unitarity of the PMNS matrix

- Unitarity condition: $U^\dagger U =$

$$\begin{pmatrix} 1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

- Normalization

$$|U_{e1}|^2 + |U_{e2}|^2 + |U_{e3}|^2 = 1$$

$$|U_{\mu 1}|^2 + |U_{\mu 2}|^2 + |U_{\mu 3}|^2 = 1$$

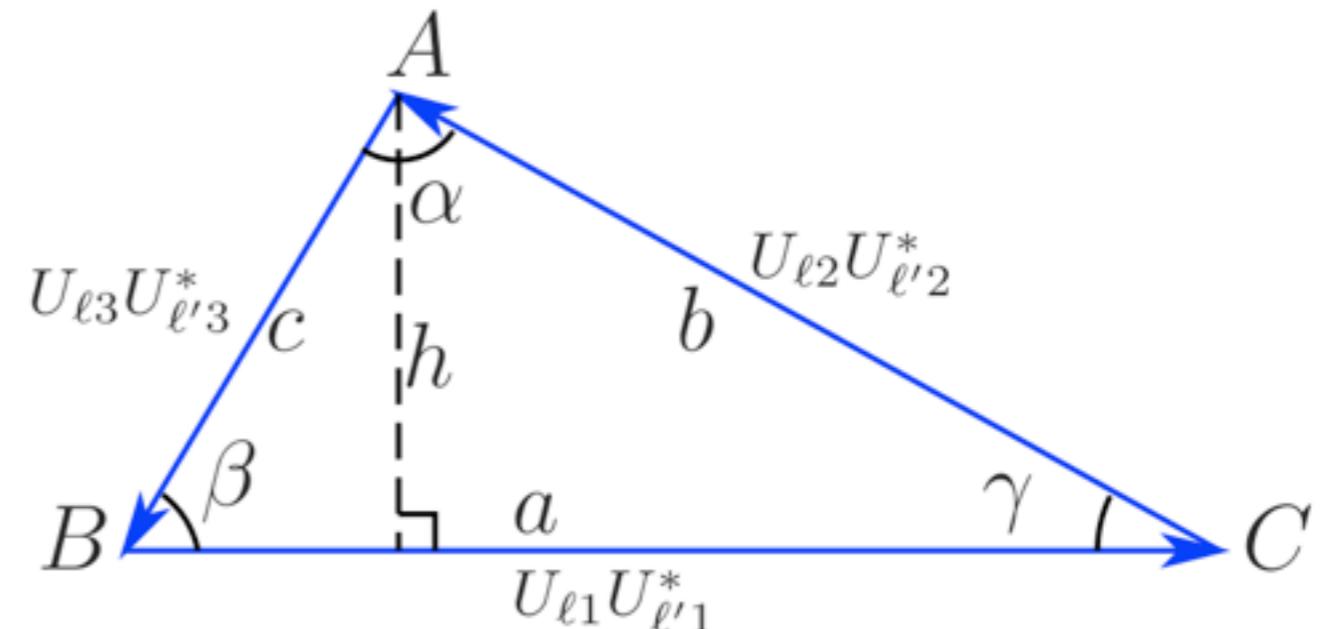
$$|U_{\tau 1}|^2 + |U_{\tau 2}|^2 + |U_{\tau 3}|^2 = 1$$

- Unitarity triangle

$$U_{e1}U_{\mu 1}^* + U_{e2}U_{\mu 2}^* + U_{e3}U_{\mu 3}^* = 0$$

$$U_{\mu 1}U_{\tau 1}^* + U_{\mu 2}U_{\tau 2}^* + U_{\mu 3}U_{\tau 3}^* = 0$$

$$U_{\tau 1}U_{e1}^* + U_{\tau 2}U_{e2}^* + U_{\tau 3}U_{e3}^* = 0$$



sum of three complex numbers (vectors) = 0

- New physics can break the Unitarity of PMNS
 - sterile neutrino or new interactions

e- μ unitarity of PMNS matrix

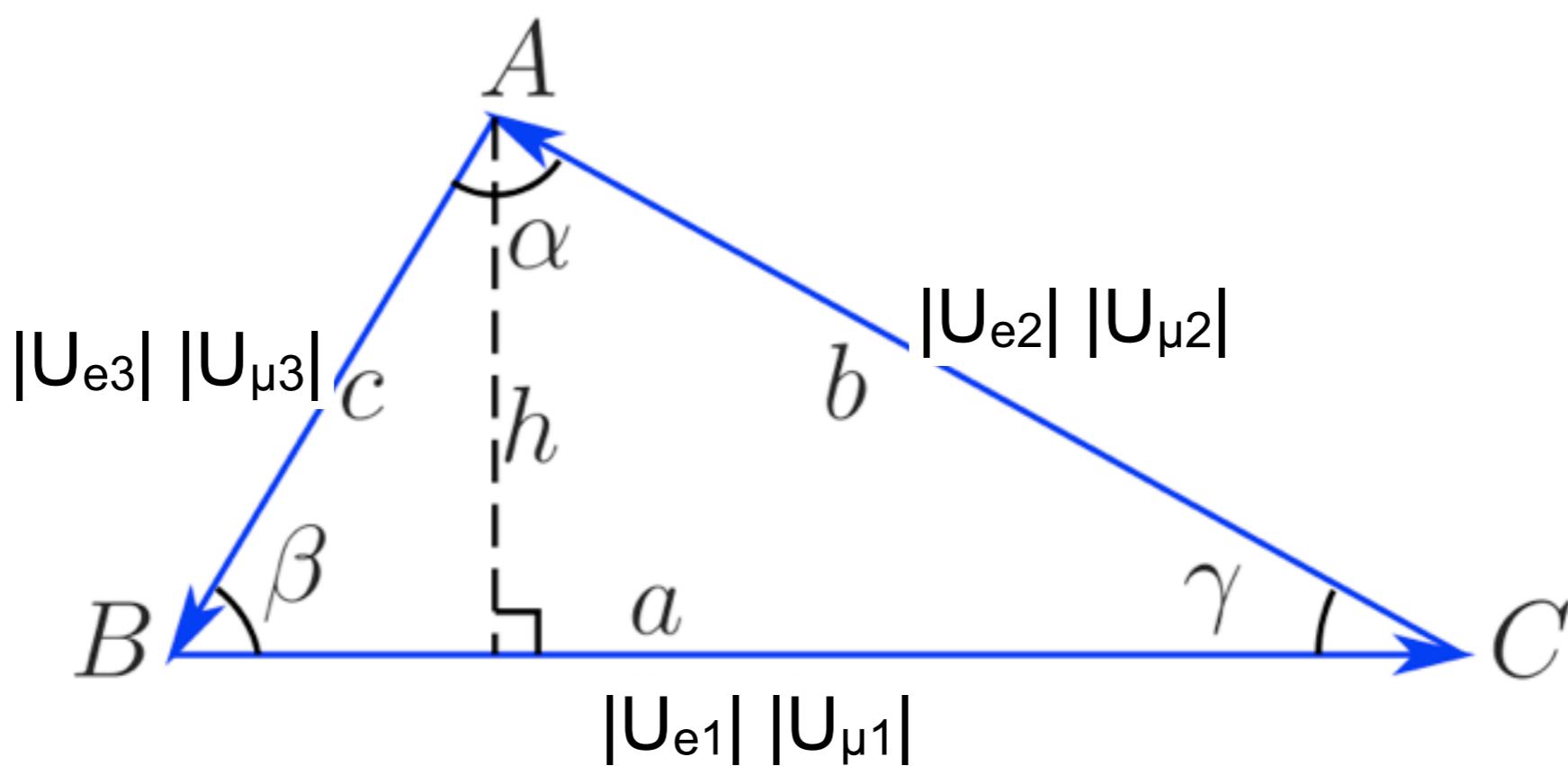
- e- μ part of unitarity can be studied the best:

$$|U_{e1}|^2 + |U_{e2}|^2 + |U_{e3}|^2 = 1$$

Normalization

$$|U_{\mu 1}|^2 + |U_{\mu 2}|^2 + |U_{\mu 3}|^2 = 1$$

$$U_{e1}U_{\mu 1}^* + U_{e2}U_{\mu 2}^* + U_{e3}U_{\mu 3}^* = 0 \quad \text{Unitarity triangle}$$



Disappearance

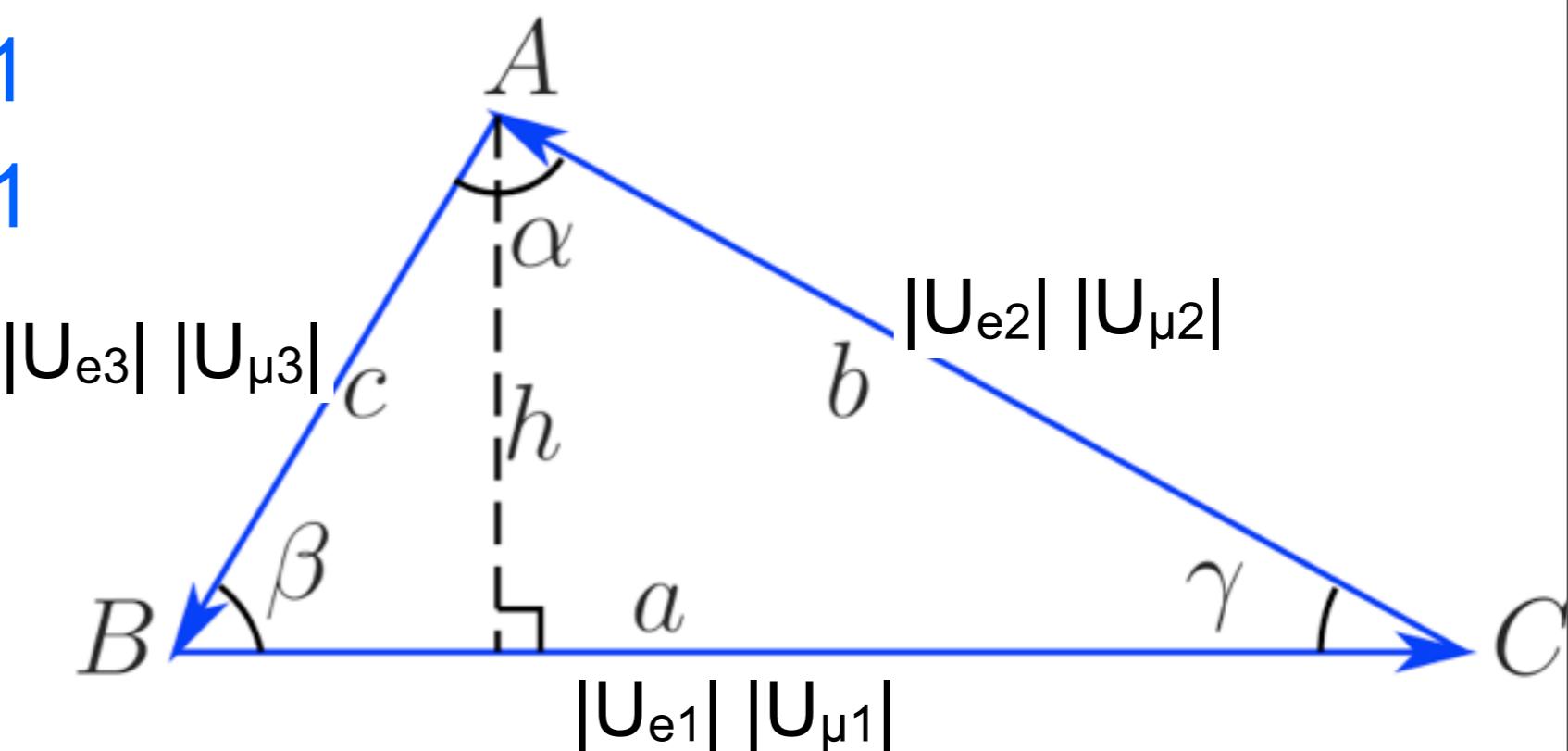
- Disappearance tests normalization part of unitarity and the sides of the unitarity triangle:

$$\begin{aligned}
 P_{\text{disapp}} &= 1 - P_{e \rightarrow e} \\
 &= 4|U_{e1}|^2 |U_{e2}|^2 \sin^2 \Delta_{12} \\
 &\quad + 4|U_{e2}|^2 |U_{e3}|^2 \sin^2 \Delta_{23} \\
 &\quad + 4|U_{e3}|^2 |U_{e1}|^2 \sin^2 \Delta_{31}
 \end{aligned}$$

$$\begin{aligned}
 \Delta_{12} &= \Delta m^2_{12} L / 4E \\
 \Delta_{23} &= \Delta m^2_{23} L / 4E \\
 \Delta_{31} &= \Delta m^2_{31} L / 4E
 \end{aligned}$$

$$|U_{e1}|^2 + |U_{e2}|^2 + |U_{e3}|^2 = 1$$

$$|U_{\mu 1}|^2 + |U_{\mu 2}|^2 + |U_{\mu 3}|^2 = 1$$



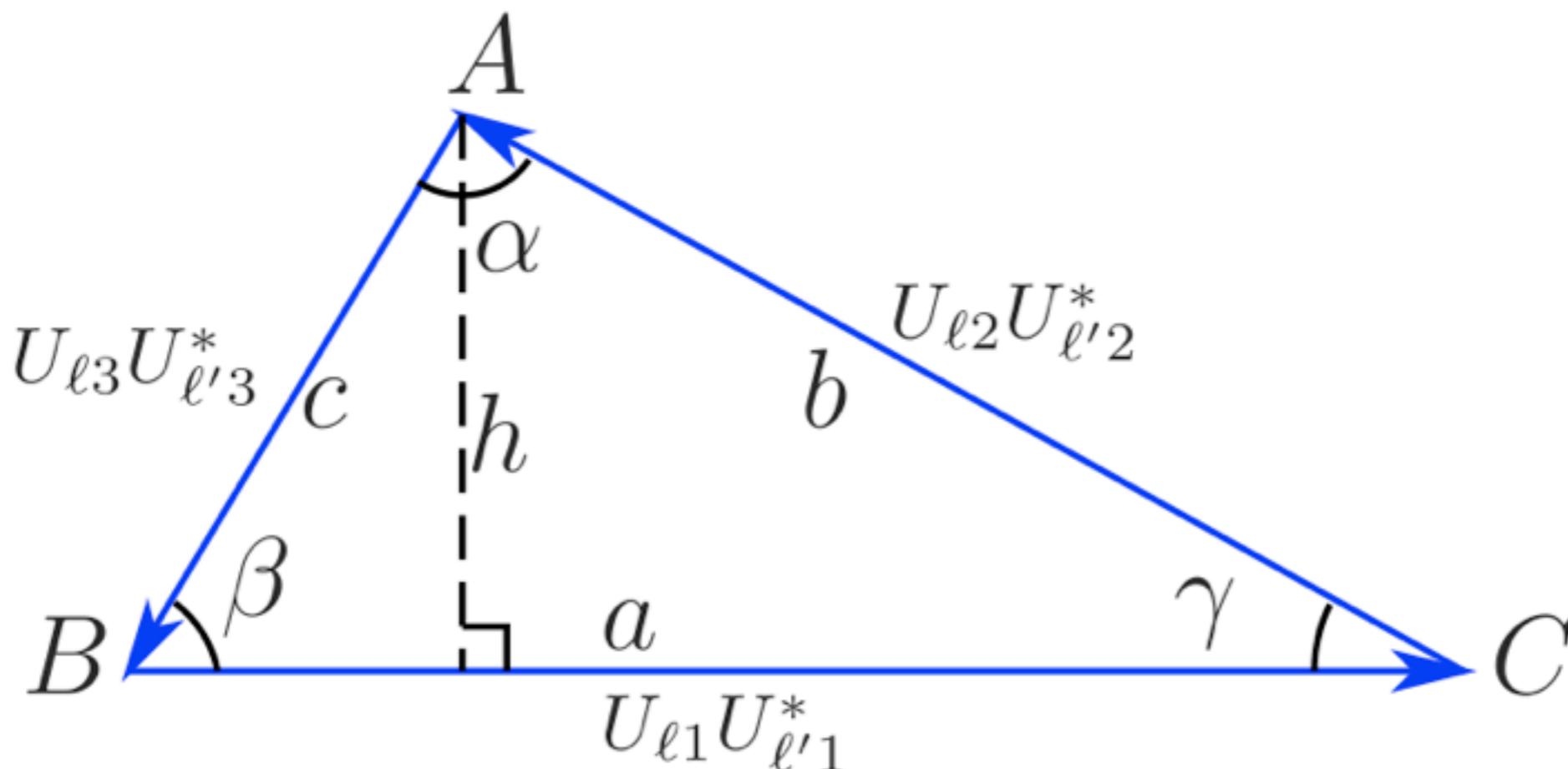
$$|U_{e1}|^2 + |U_{e2}|^2 + |U_{e3}|^2 = 1$$

- Solar neutrino: $|U_{e2}|^2$
 - MSW conversion in the sun: $\nu_e \rightarrow \nu_2$ and detect ν_e
- KamLand: $|U_{e1}|^2 |U_{e2}|^2$
 - Reactor ν_e disappearance at Δ_{12} scale of L/E
- Reactor θ_{13} : $|U_{e3}|^2 (|U_{e1}|^2 + |U_{e2}|^2)$
 - Reactor ν_e disappearance at Δ_{13} scale of L/E
- All the parameters are already measured:
 - Better precision for $|U_{e1}|^2$ and $|U_{e2}|^2$ are expected by JUNO, untangling Δ_{23} and Δ_{31} contributions

$$|U_{\mu 1}|^2 + |U_{\mu 2}|^2 + |U_{\mu 3}|^2 = 1$$

- LBL ν_μ disappearance: $|U_{\mu 3}|^2 (|U_{\mu 1}|^2 + |U_{\mu 2}|^2)$
 - T2K/MINOS/NOvA/SK/ICECUBE measure this.
- $\nu_\mu \rightarrow \nu_e$ atm. matter resonance (~ 6 GeV): $(r|U_{\mu 3}|^2 - 1)$
 - Amplitude of the atm. ν mass hierarchy study [SK/HK]
- Δ_{12} scale atm. ν_μ disapp. : $|U_{\mu 1}|^2 |U_{\mu 2}|^2$
 - sub-GeV up-going atm. ν_μ disapp. [HK]
- Good potential to untangle all three parameters:
HyperK will play a major role

Appearance (PRD89(2014)073002)



$$a = |U_{\ell 1}| |U_{\ell' 1}|$$

$$b = |U_{\ell 2}| |U_{\ell' 2}|$$

$$c = |U_{\ell 3}| |U_{\ell' 3}|$$

$$\Delta_{12} = \Delta m^2_{12} L / 4E$$

$$\Delta_{23} = \Delta m^2_{23} L / 4E$$

$$\Delta_{31} = \Delta m^2_{31} L / 4E$$

Appearance

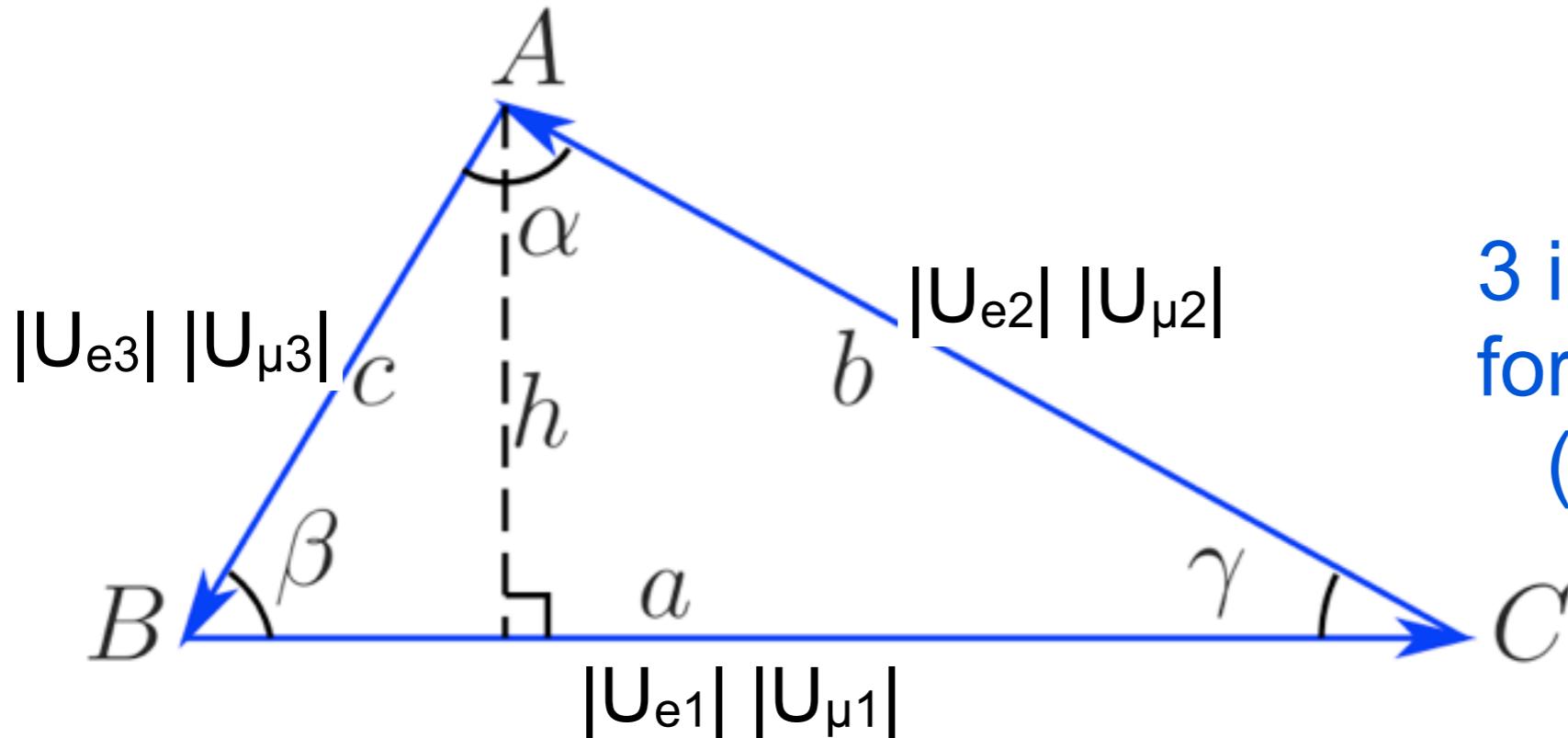
$$P_{\ell \rightarrow \ell'} = 4ab \sin(\Delta_{12} \pm \gamma) \sin \Delta_{12}$$

$$+ 4bc \sin(\Delta_{23} \pm \alpha) \sin \Delta_{23}$$

$$+ 4ac \sin(\Delta_{31} \pm \beta) \sin \Delta_{31}$$

Each of the CP angles (α, β, γ) correspond to each L/E oscillation length; Δ_{23} , Δ_{31} , Δ_{12}

$$U_{e1}U^* \mu_1 + U_{e2}U^* \mu_2 + U_{e3}U^* \mu_3 = 0$$



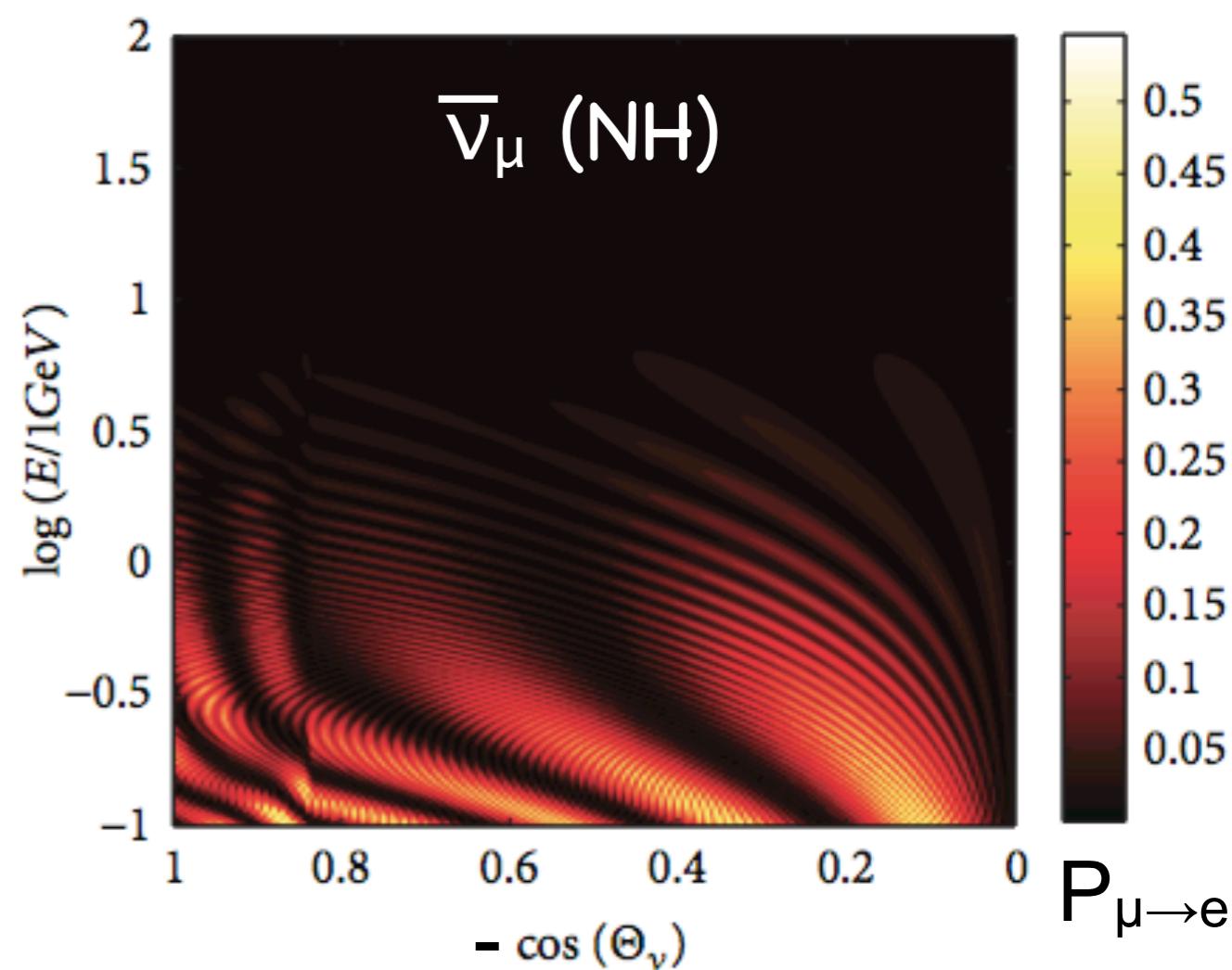
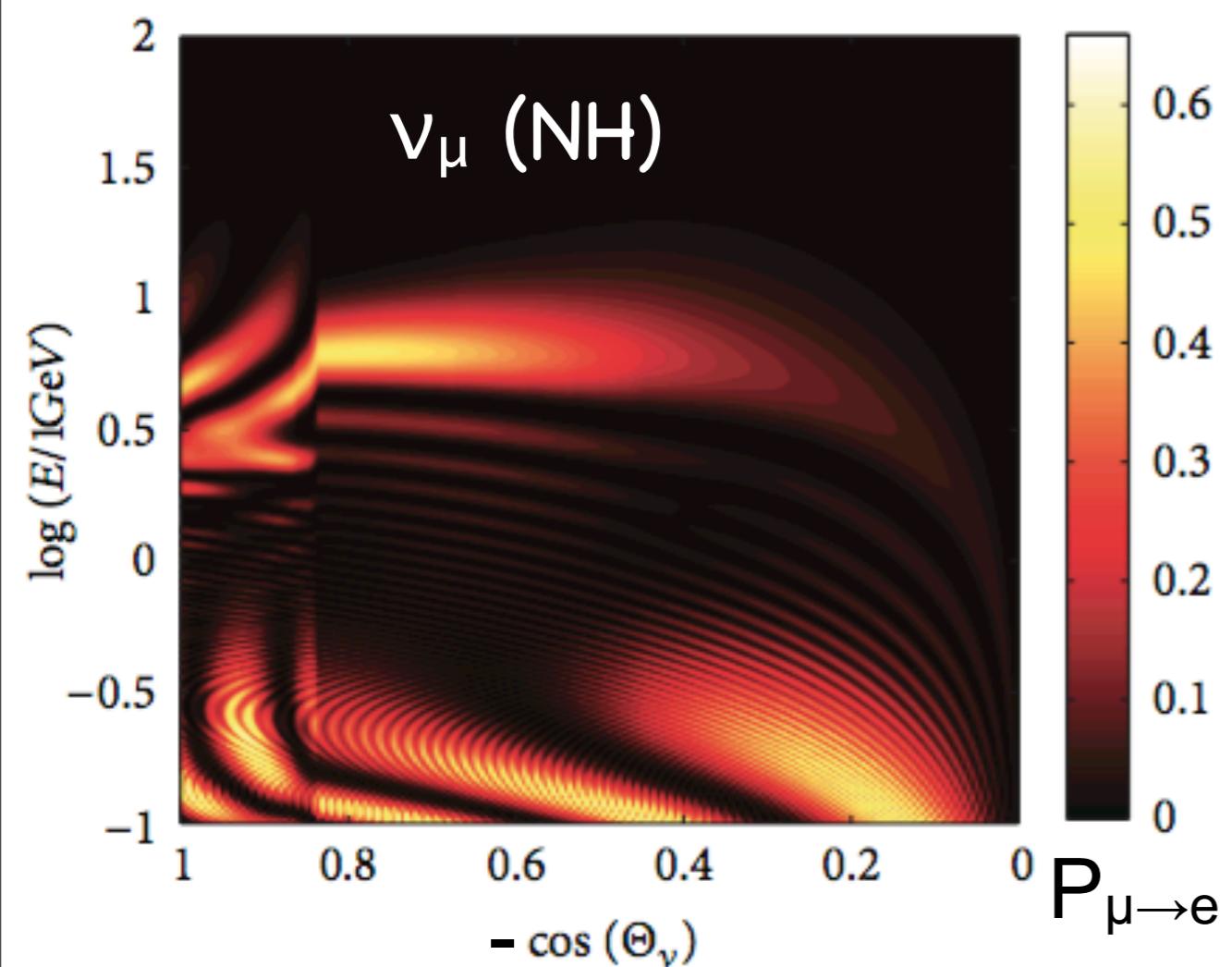
3 independent parameter
for 6 observables
(3 constraints)

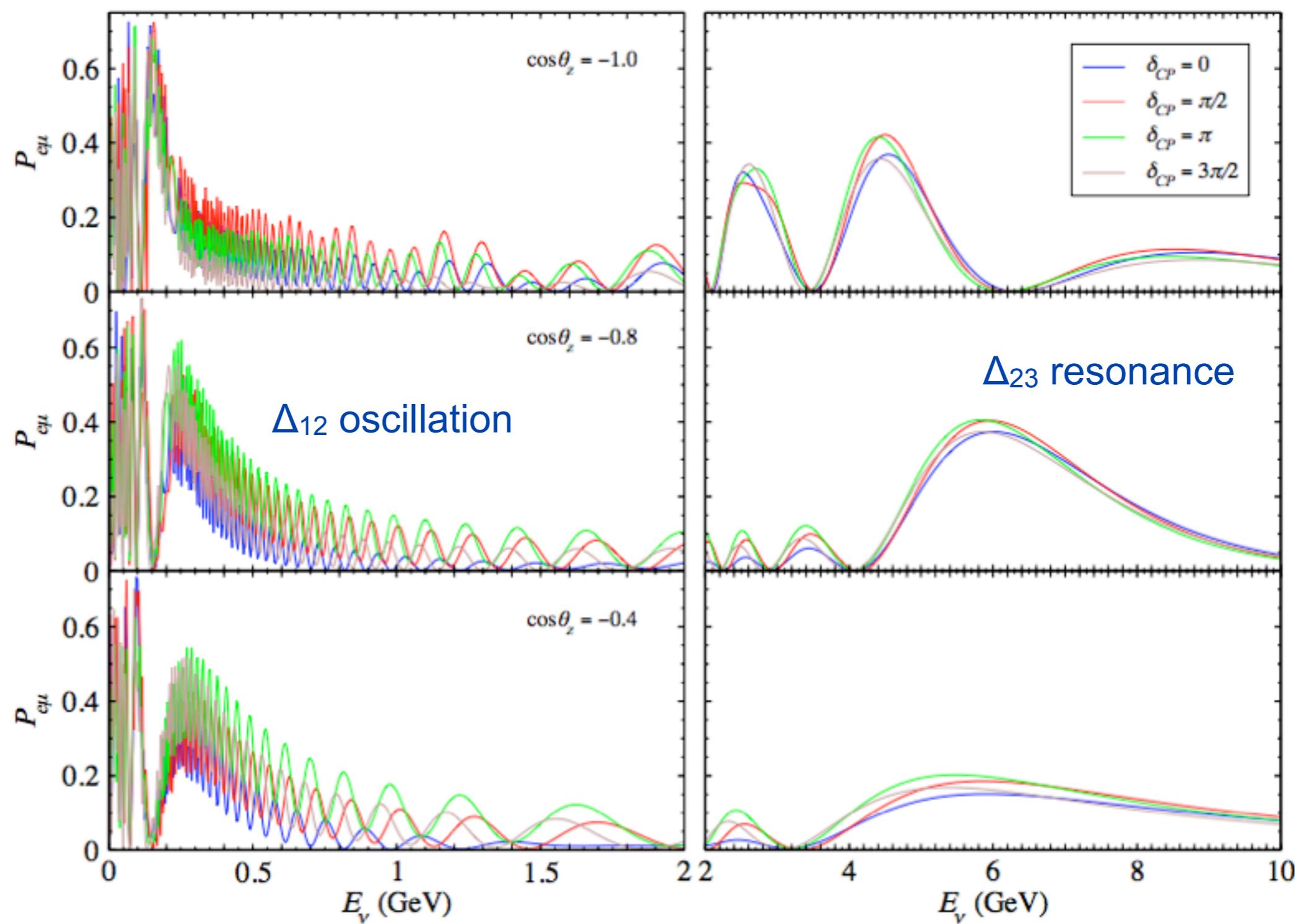
- a, b, c can be determined by disappearance
- α, β from LBL $\nu_\mu \rightarrow \nu_e$ CP violation, γ from atm. $\nu_\mu \rightarrow \nu_e$ CP
 - T2K/T2HK/DUNE will determine combination of α and β [T2K/HK]
 - γ can be determined by the 0.5–0.8GeV atm. $\nu_\mu \rightarrow \nu_e$ [SK/HK]
 - Atm. $\nu_\mu \rightarrow \nu_e$ (2nd to 10th max.) untangles $\beta(\Delta_{31})$ and $\alpha(\Delta_{23})$ [HK]

Atmospheric $\nu_\mu \rightarrow \nu_e$

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- $\nu_\mu \rightarrow \nu_e$ at several GeV: matter resonance
 - Mass hierarchy determination
- Δ_{12} oscillation in the sub-GeV region:
 - large θ_{12} effect



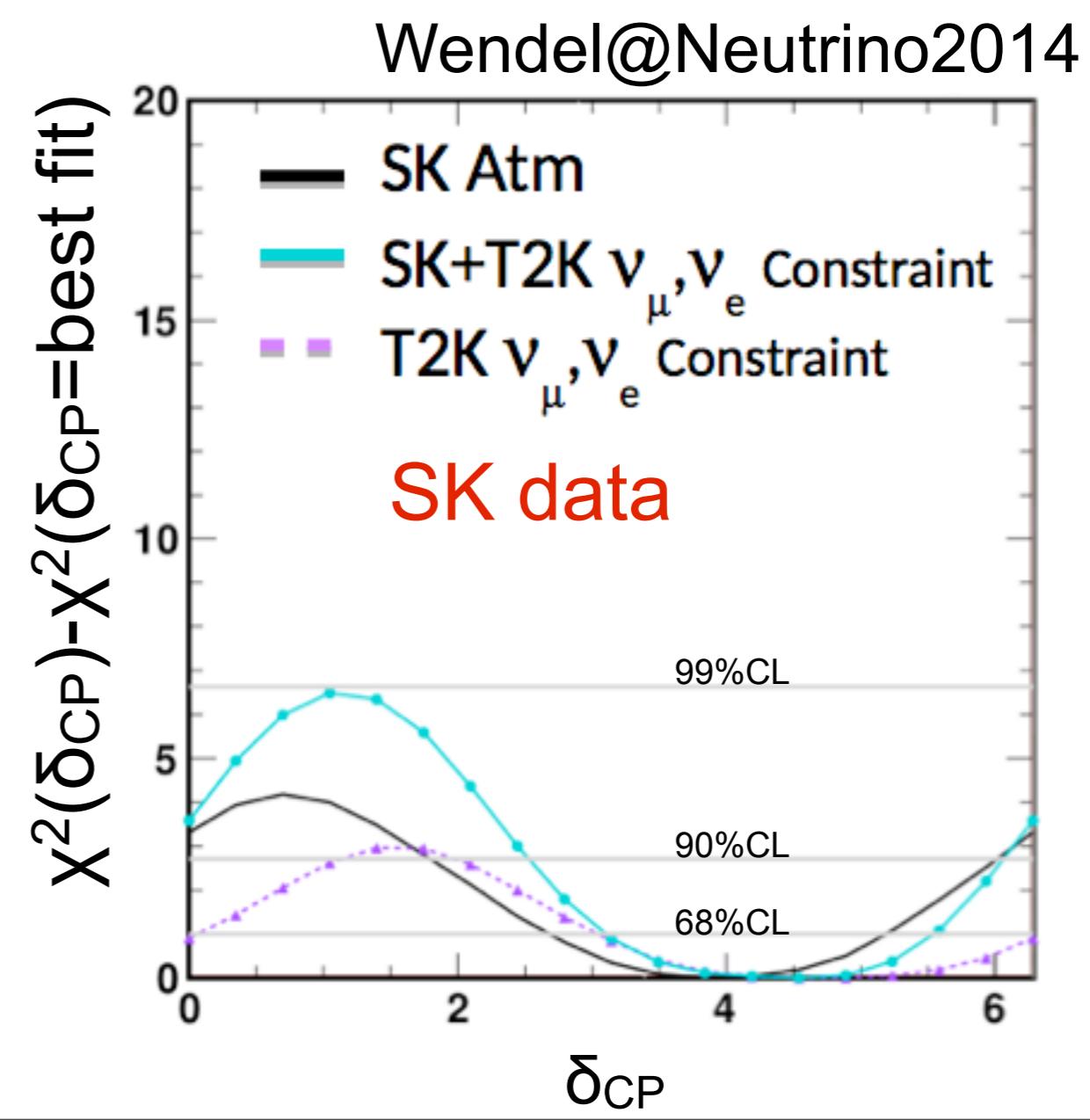
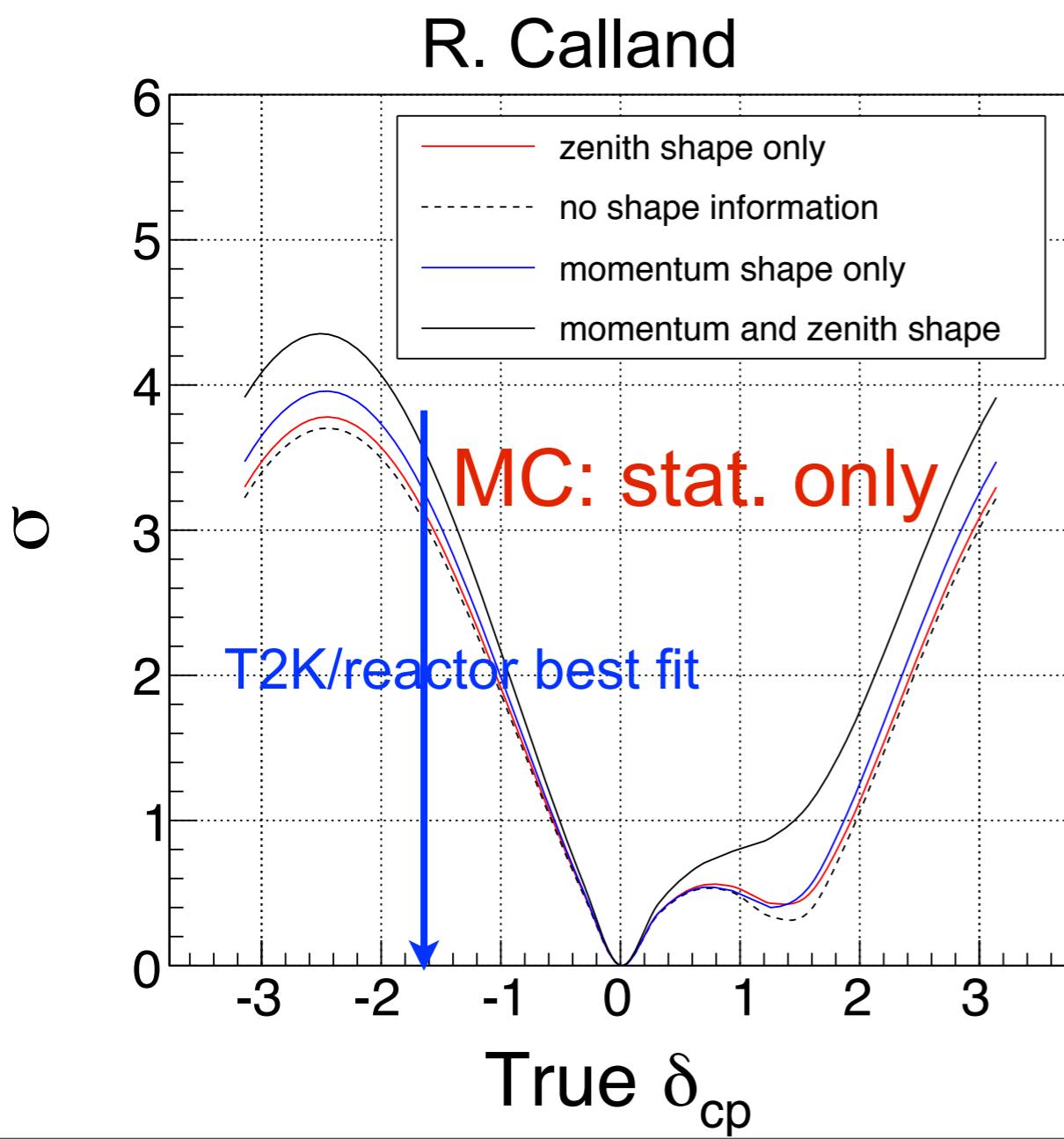


CP violation effect is large! $P_{\mu e} \sim$ up to 20%

Sub-GeV CP sensitivity of SK

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- Sensitivity to reject $\delta_{CP}=0$ with existing data:
 - Statistical: 3.6σ at the T2K-reactor best fit point
 - Current SK analysis with systematics: 1.6σ

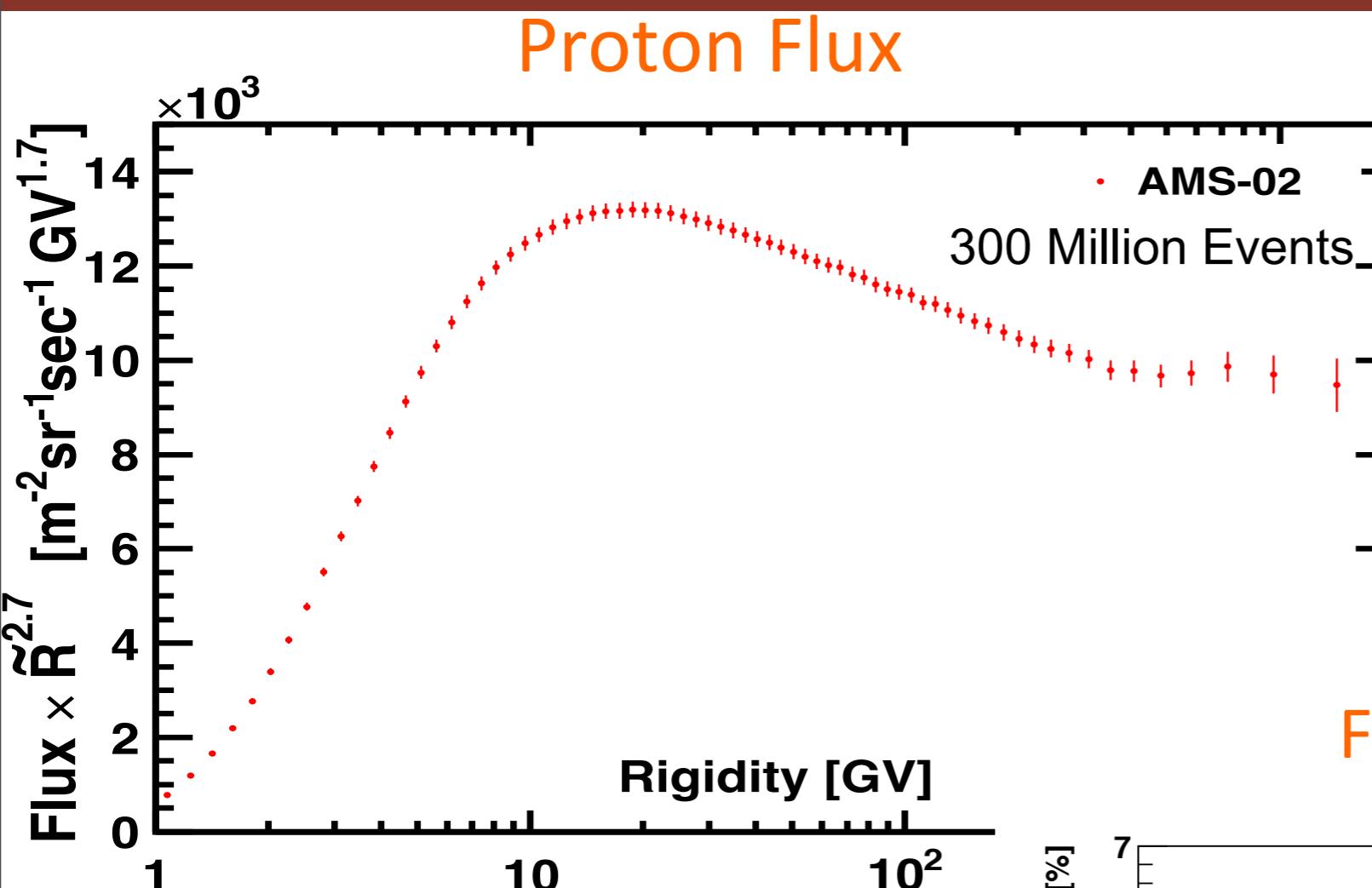


"Grain of Salt" ?

- Atm. ν offers wide range of L/E in oscillation
 - However, taken by the community with "grain of salt"
 - Good statistics, limited by "unknown" systematics
- Emerging opportunity for precision atm. ν:
Control systematic errors by
 - Neutrino flux
 - **AMS** data provides precise primary cosmic ray flux
 - Hadron production studies developed by **CERN-NA61**
 - Neutrino cross section
 - Model independent cross section proposal by **nuPRISM**

Primary Cosmic Ray Flux (AMS)

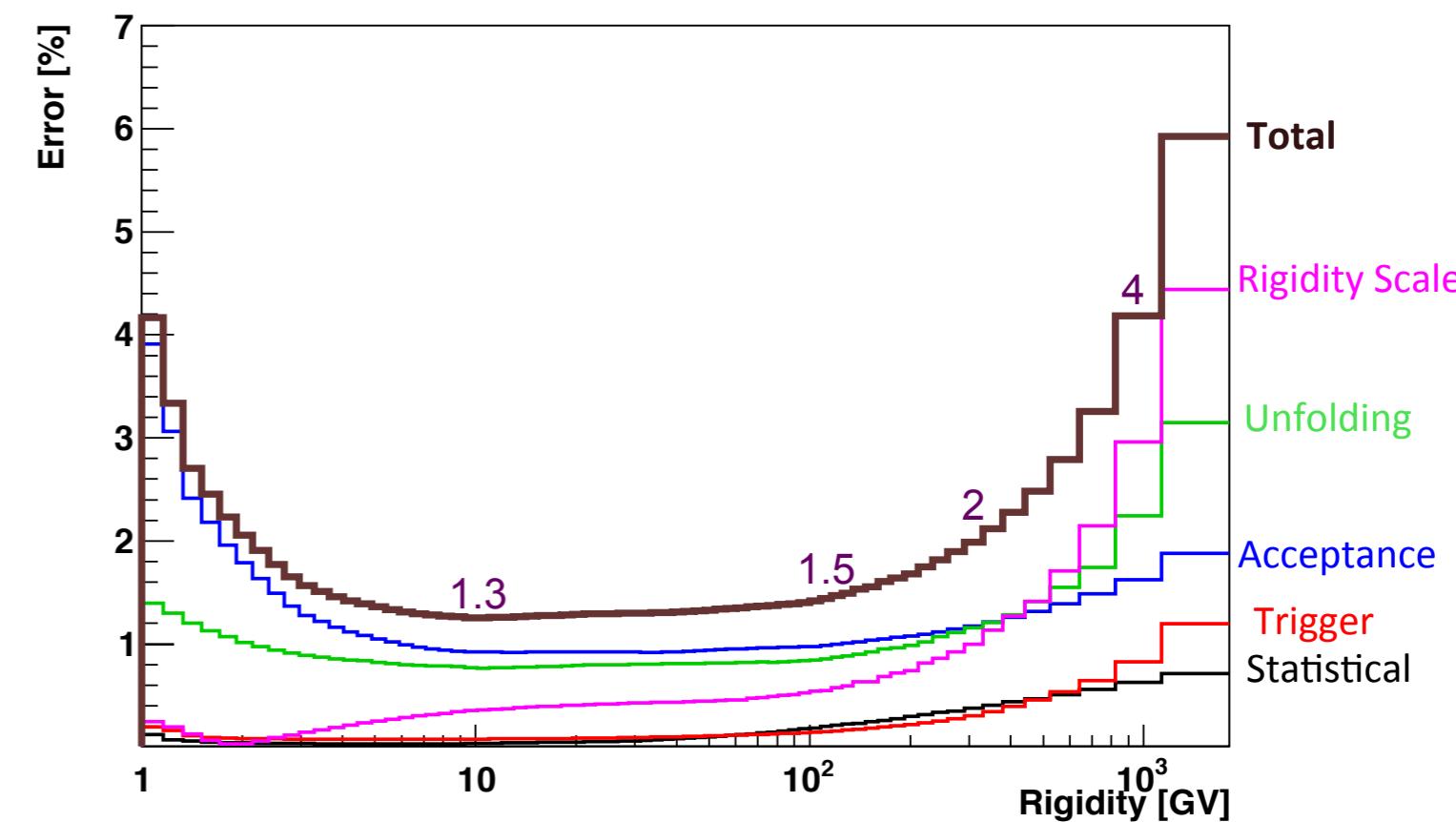
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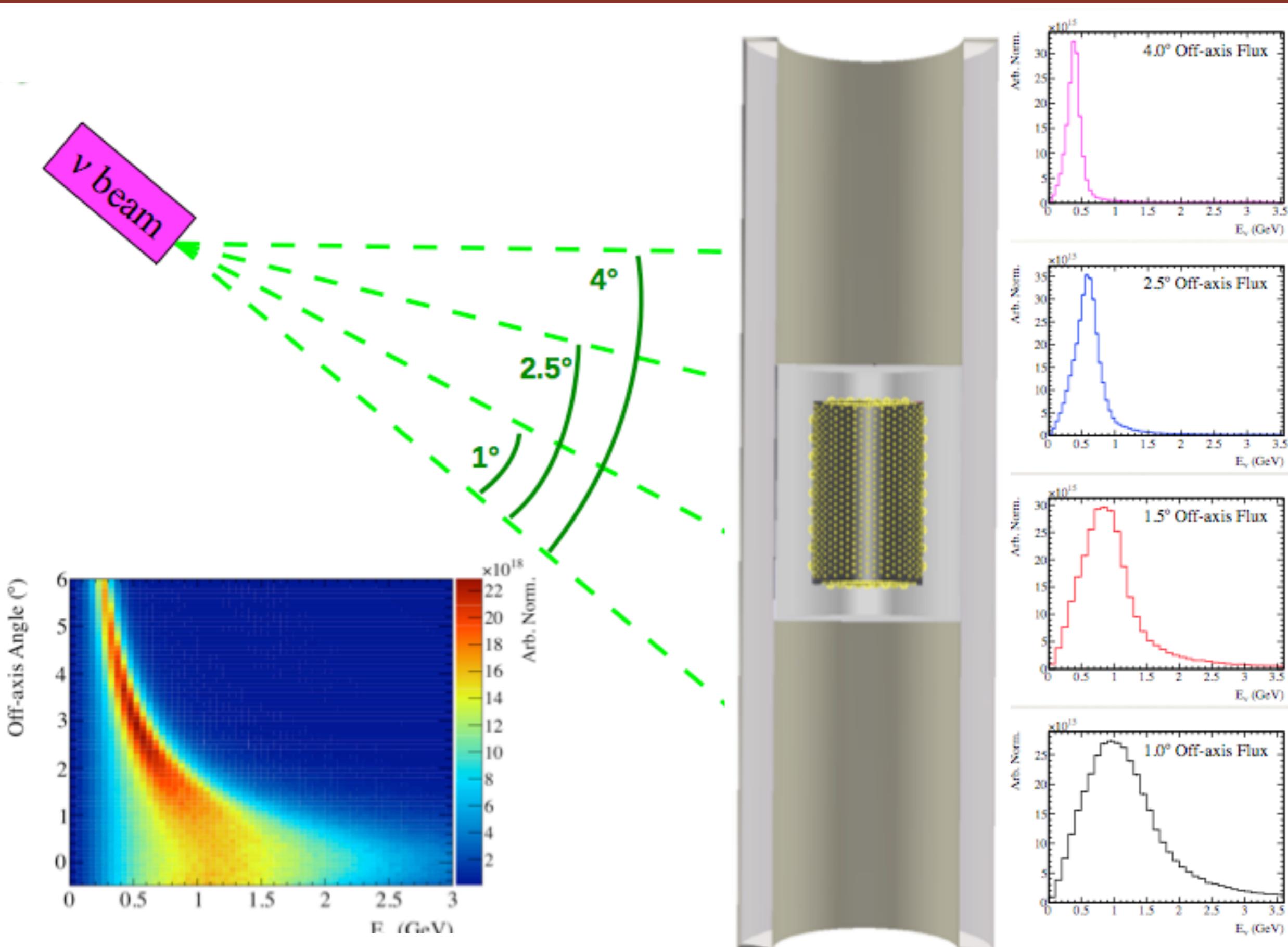
AMS: April 2015

Flux Errors Breakdown

- Primary cosmic ray flux is now measured at 1-2% level.
- Open the door for very precise atm.v studies



vPRISM (T2K/HyperK ND) concept



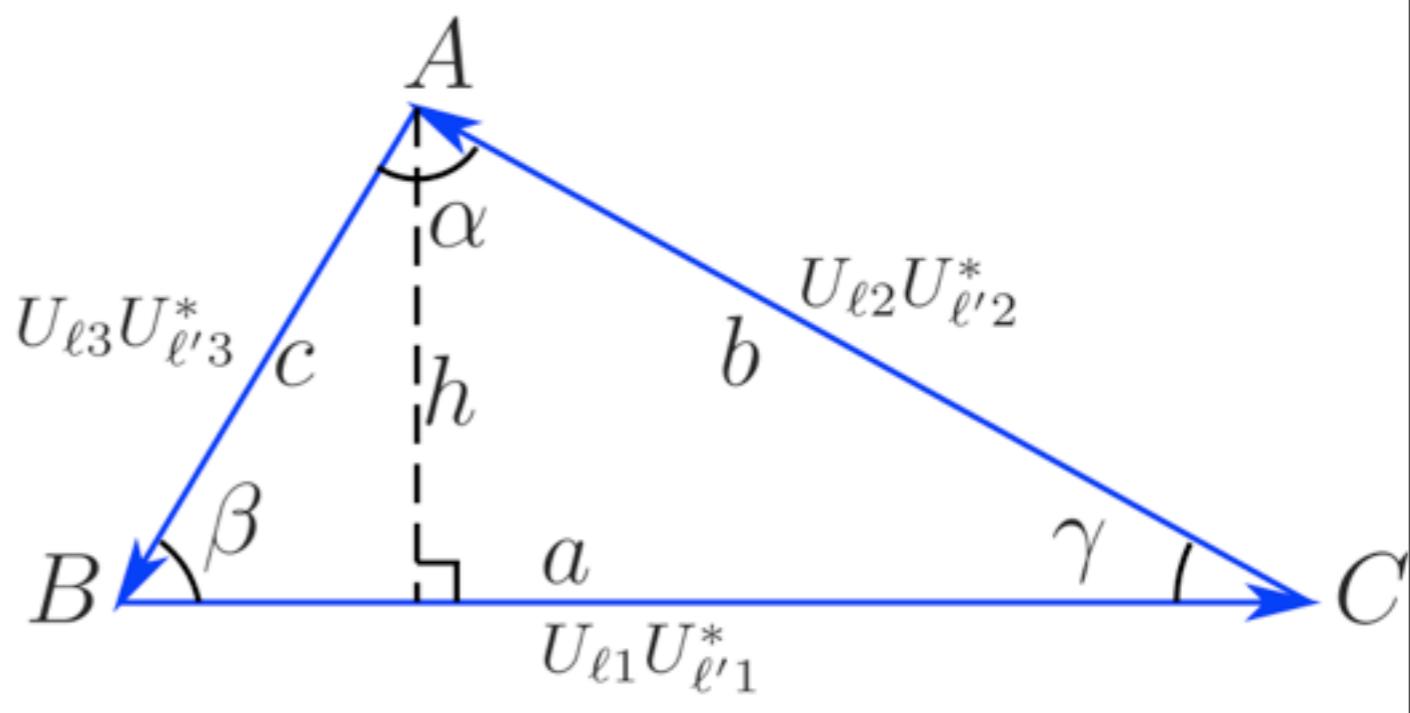
- Atm. v covers wide range of oscillation parameters
 - Realistic scenario emerges to control systematic uncertainties:
 - Precise AMS cosmic ray measurement
 - Precision hadron production study established by NA61: nuPRISM proposal for model independent v cross section study
 - Precision test of e- μ PMNS Unitarity
 - 9 parameters with 5 constraints to be tested
 - $|U_{e1}|^2, |U_{e2}|^2, |U_{e3}|^2, |U_{\mu 1}|^2, |U_{\mu 2}|^2, |U_{\mu 3}|^2, \alpha, \beta, \gamma$
 - Unitarity triangle (3 constraints), normalization (2)
 - All 9 parameters can be measured:
 - Typical expected accuracy would be 10-20%
 - HyperK will lead the new measurements: $|U_{\mu i}|^2, \alpha, \beta, \gamma$
JUNO will improve $|U_{ei}|^2$
 - HyperK will play a leading role in establishing PMNS paradigm, similar to what B-factory did for CKM

Backup slides

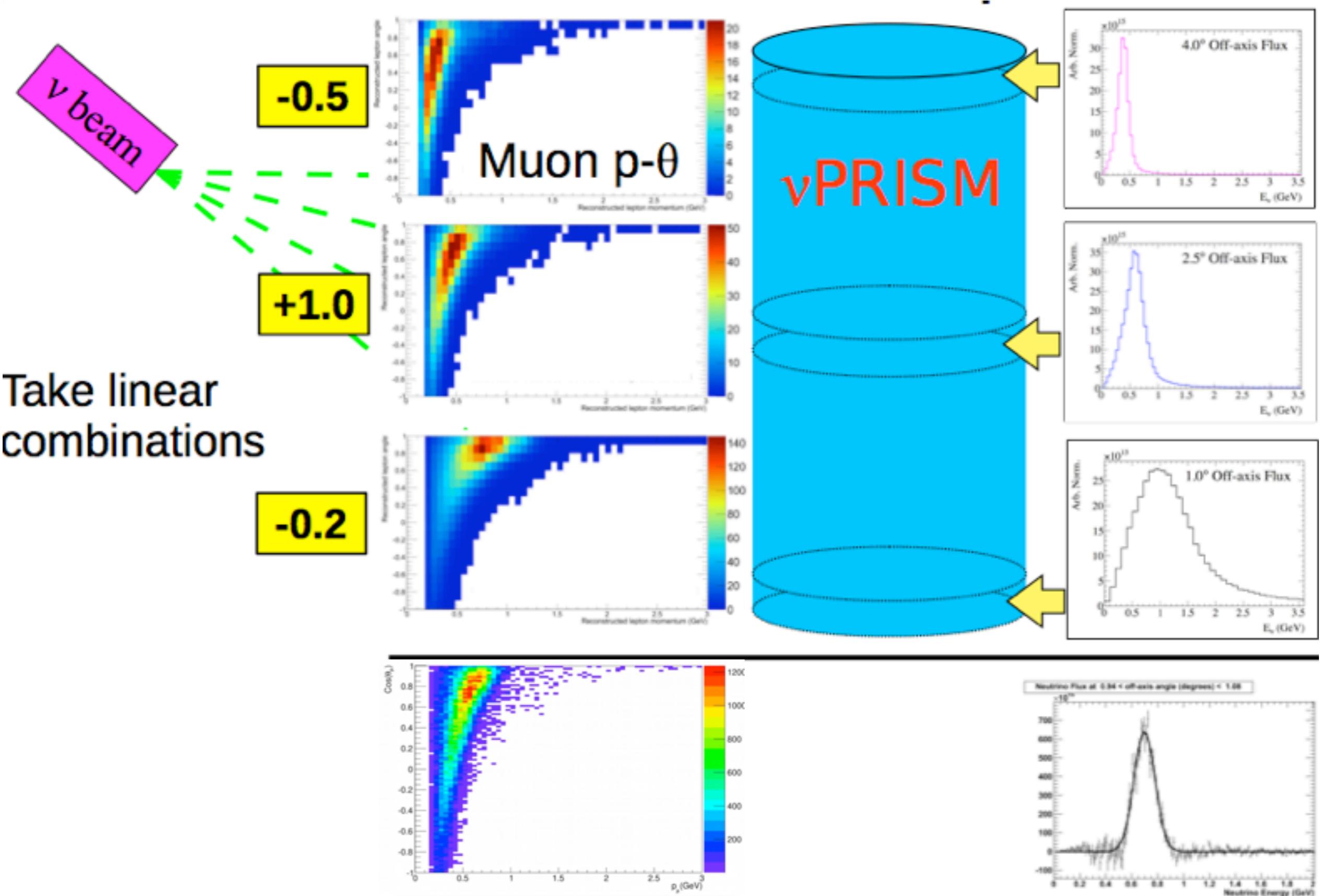
Unitarity triangle

$$\begin{aligned}
 P_{l \rightarrow l'} &= |<\nu_{l'}(t)|\nu_l(t)>|^2 \\
 &= |\sum_i <\nu_{l'}|\nu_i> e^{-iE_it} <\nu_i|\nu_l>|^2 \\
 &= |\sum_i U_{li}^* U_{l'i} e^{-im_i^2 L/2E}|^2 = \sum_{i,j} U_{li}^* U_{l'i} (U_{lj}^* U_{l'j})^* e^{-i(m_i^2 - m_j^2)L/2E} \\
 &= |U_{l1}^* U_{l'1}|^2 + |U_{l2}^* U_{l'2}|^2 + |U_{l3}^* U_{l'3}|^2 \\
 &\quad + 2Re[U_{l1}^* U_{l'1} [U_{l2}^* U_{l'2}]^* e^{-i(m_1^2 - m_2^2)L/2E}] \quad \Delta_{12} = \Delta m^2_{12} L / 4E \\
 &\quad + 2Re[U_{l2}^* U_{l'2} [U_{l3}^* U_{l'3}]^* e^{-i(m_2^2 - m_3^2)L/2E}] \quad \Delta_{23} = \Delta m^2_{23} L / 4E \\
 &\quad + 2Re[U_{l3}^* U_{l'3} [U_{l1}^* U_{l'1}]^* e^{-i(m_3^2 - m_1^2)L/2E}] \quad \Delta_{31} = \Delta m^2_{31} L / 4E \\
 &= a^2 + b^2 + c^2 \\
 &\quad - 2ab \cos(2\Delta_{12} + \gamma) \\
 &\quad - 2bc \cos(2\Delta_{12} + \alpha) \\
 &\quad - 2ca \cos(2\Delta_{12} + \beta) \\
 &= 4ab \sin(\Delta_{12} + \gamma) \sin \Delta_{12} \\
 &\quad + 4ab \sin(\Delta_{23} + \alpha) \sin \Delta_{23} \\
 &\quad + 4ab \sin(\Delta_{31} + \beta) \sin \Delta_{31}
 \end{aligned}$$

Calculation is simpler without Unitarity



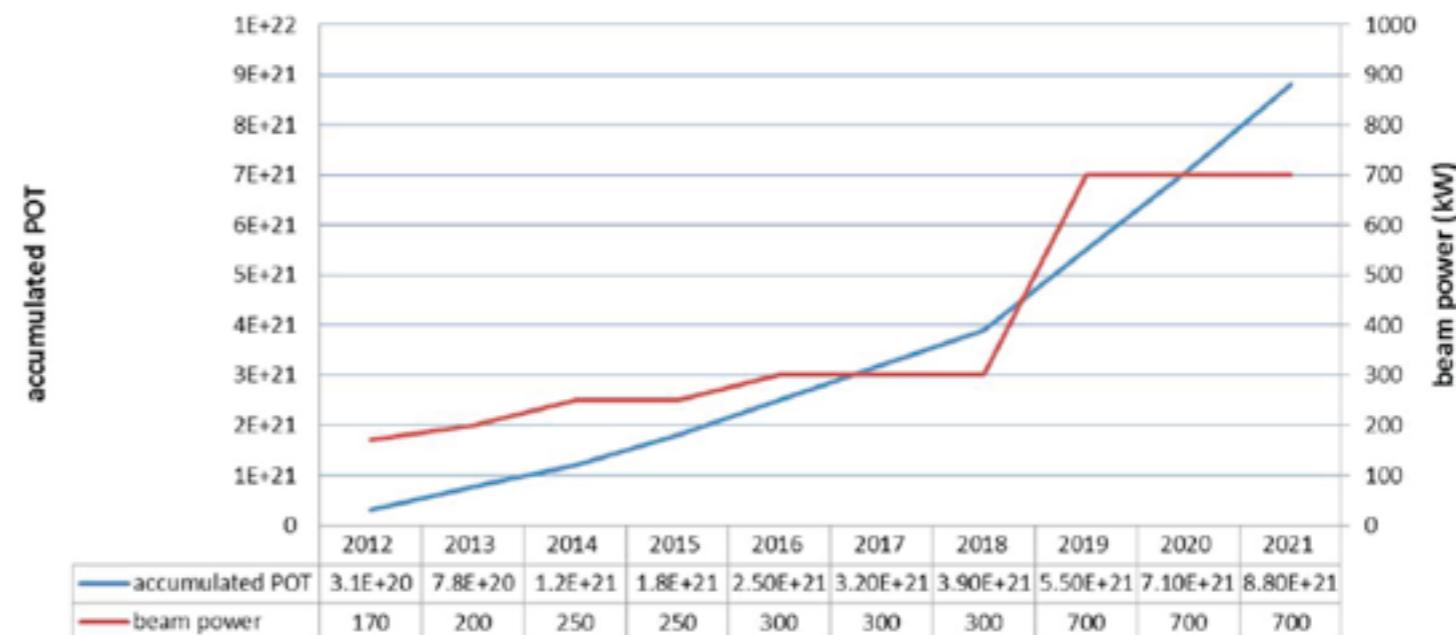
vPRISM (T2K/HyperK ND) concept



Increasing the statistics

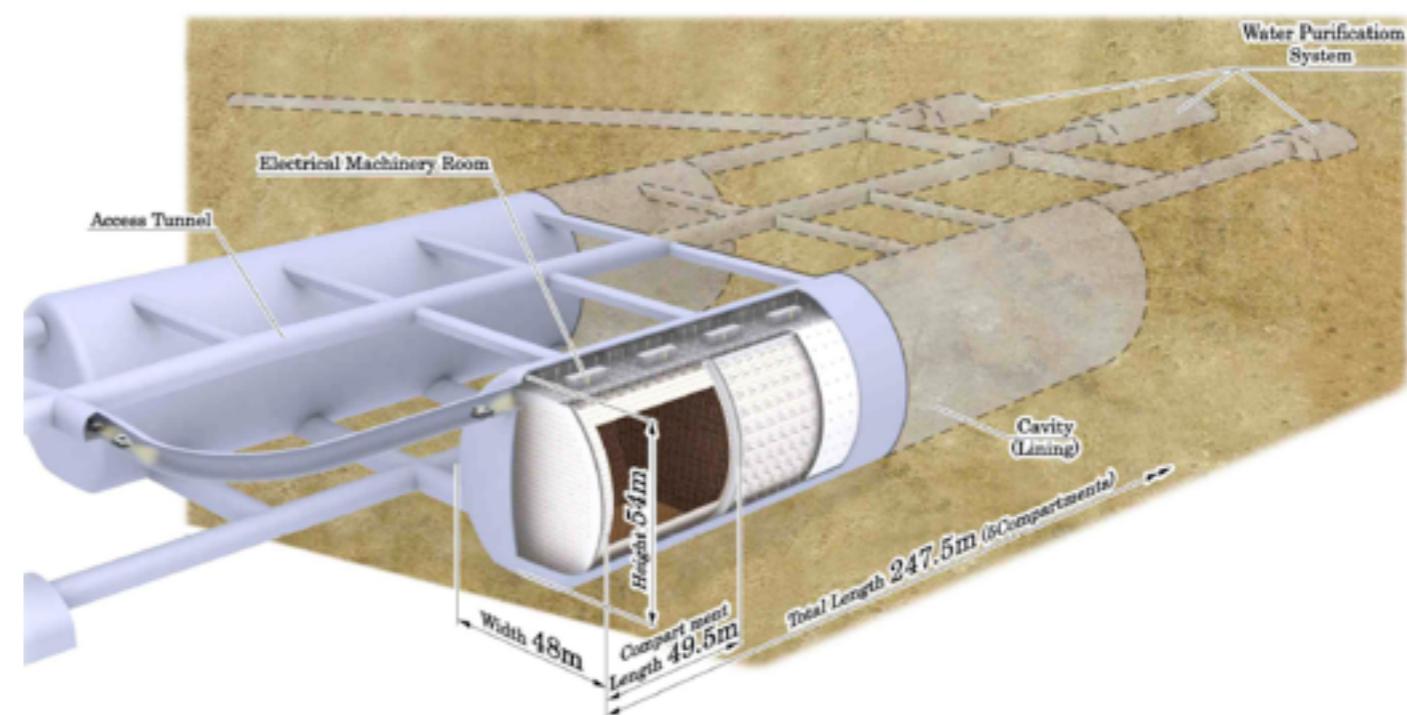
- Accelerator upgrades

- beam power:
 $250\text{ kW} \rightarrow 700\text{ kW}$
- by increased rep. rate:
 - every 3sec \rightarrow 1sec
 - keeping the similar instantaneous rate



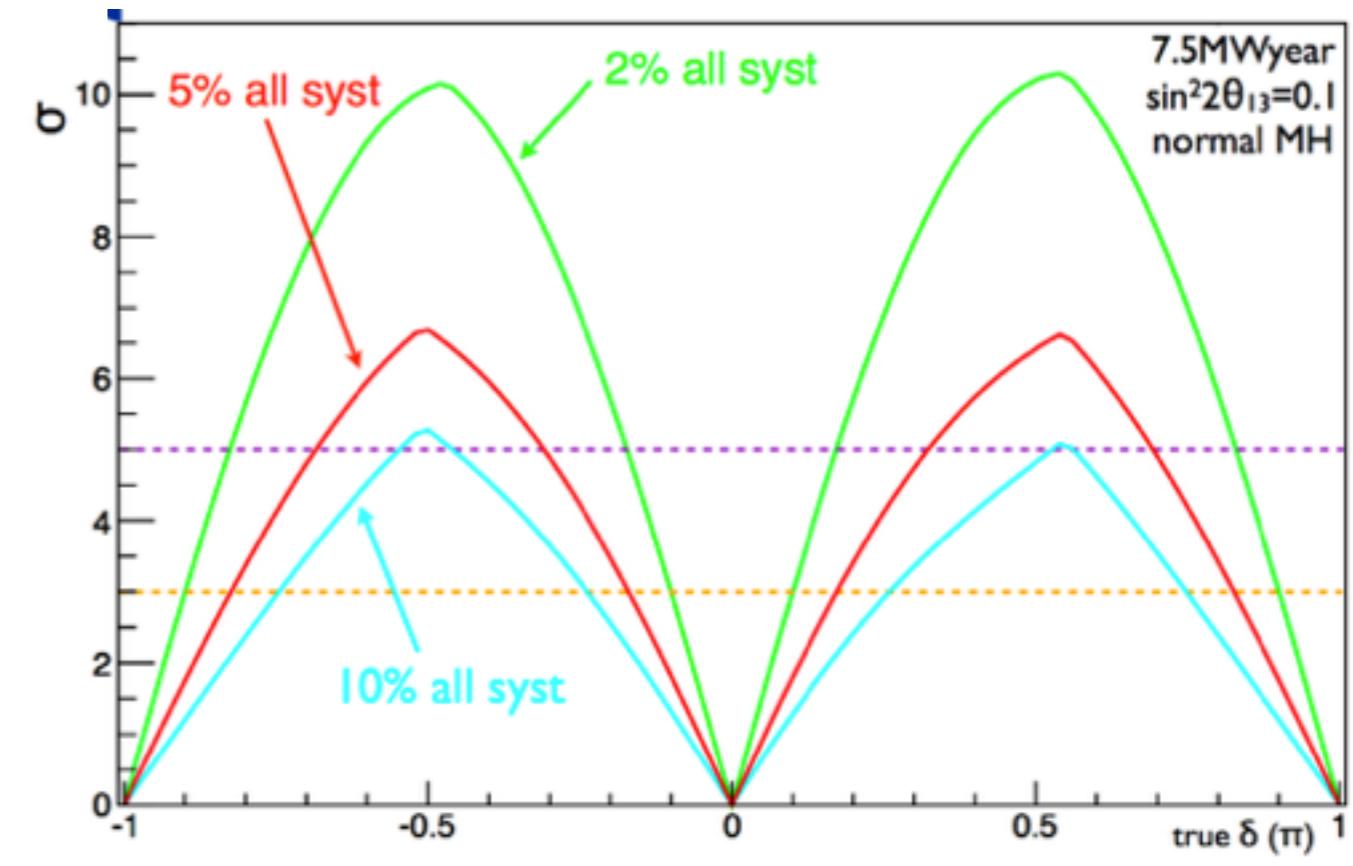
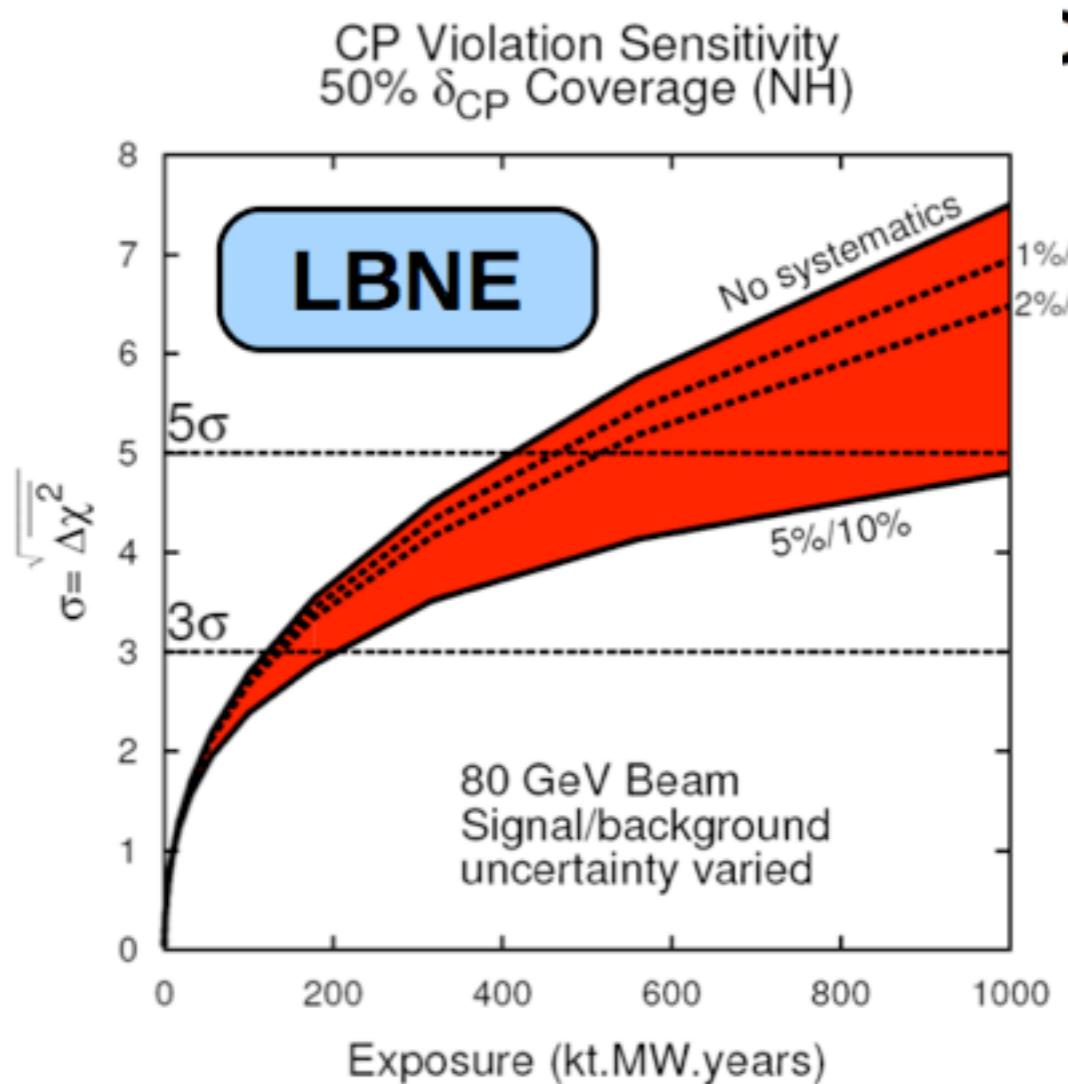
- Far detector upgrade:
Hyper-Kamiokande

- 1Mton water Cerenkov
- $\times 20$ larger fiducial volume than SK



Systematic uncertainties

- Systematic uncertainty of 1-2% is required for the wide range of CP discovery
 - Challenge: ν cross section systematics (multi-nucleon correlation, final state interaction)



T2K systematic uncertainties

- Systematic uncertainties in event rate at SK:

	ν_e app.	ν_μ disapp.
flux × cross section	3.1%	2.7%
cross section	5.3%	5.8%
SK detector	2.7%	4.0%
Total	6.8%	7.7%

- [Flux × cross section]: ND280 constrains
 - place near detector further away than 280m
- Cross section: Nuclear effects
 - water target with well controlled ν beam
- SK detector: SK detector efficiency
 - water Cherenkov calibration and ν beam test

CP cross section systematics

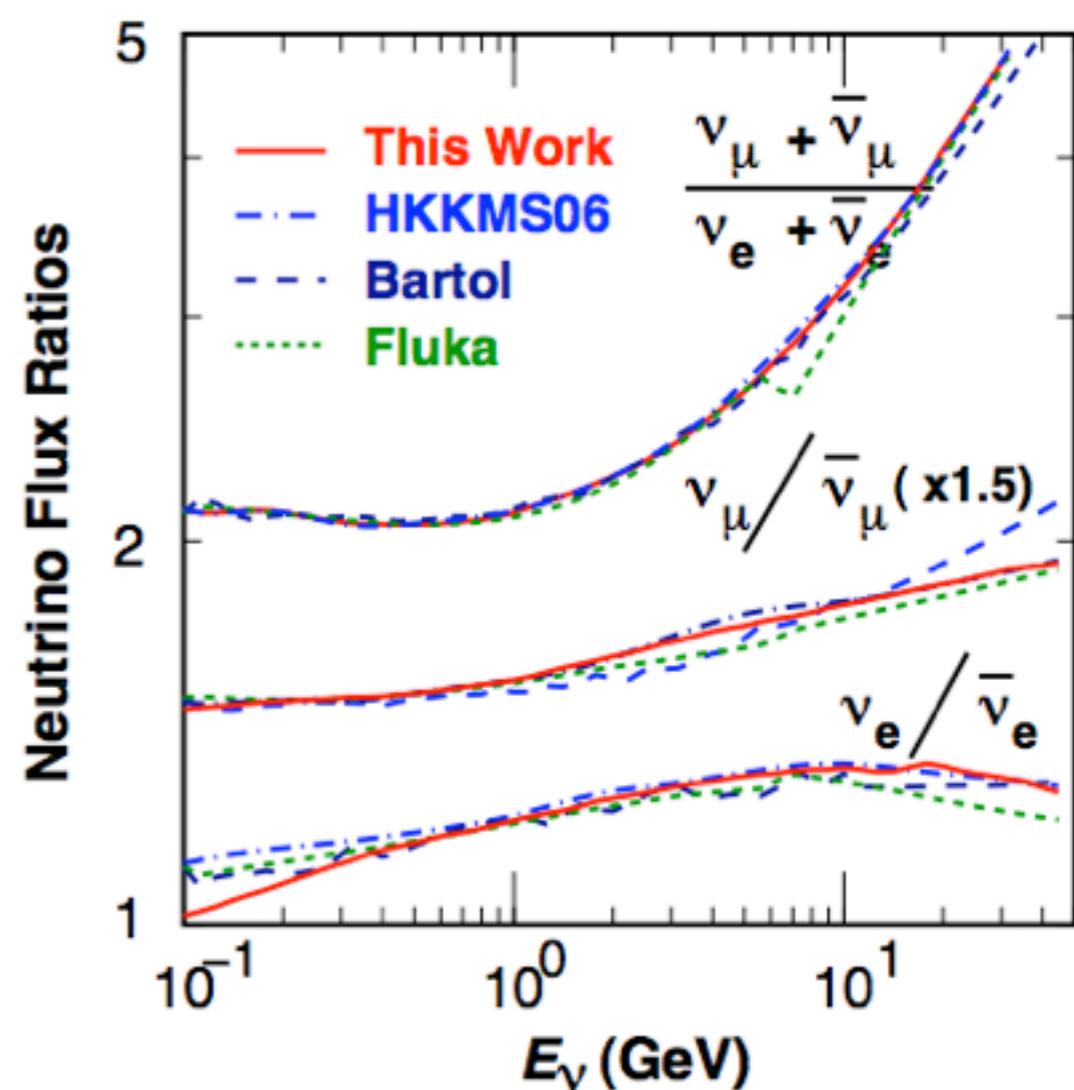
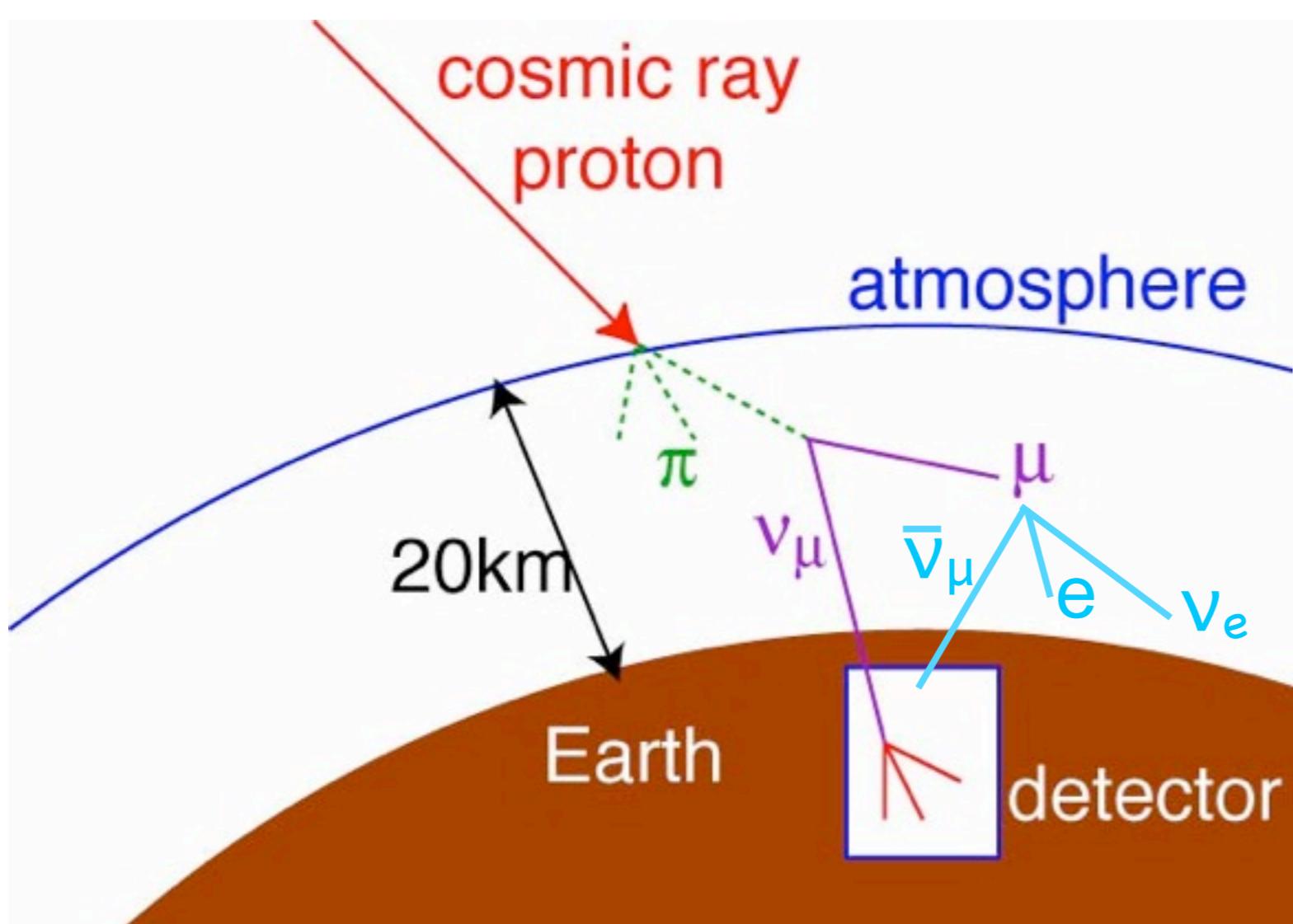
- $\sigma(\nu_e)/\sigma(\nu_\mu)$ ratio
 - For ν_e appearance, the near/far detectors observe ν_e/ν_μ , respectively
 - In the first order, it is the same: kinematical phase space, radiative correction
- $\sigma(\bar{\nu}_\mu)/\sigma(\nu_\mu)$ and $\sigma(\bar{\nu}_e)/\sigma(\nu_e)$ ratio
 - angular distribution is very different between neutrinos and anti-neutrinos
 - nuclear effect can be different
- nuPRISM to address these
 - nuSTORM (π storage ring) is another proposal

Atmospheric neutrinos

$$r = \Phi(\nu_\mu)/\Phi(\nu_e) \sim 2$$

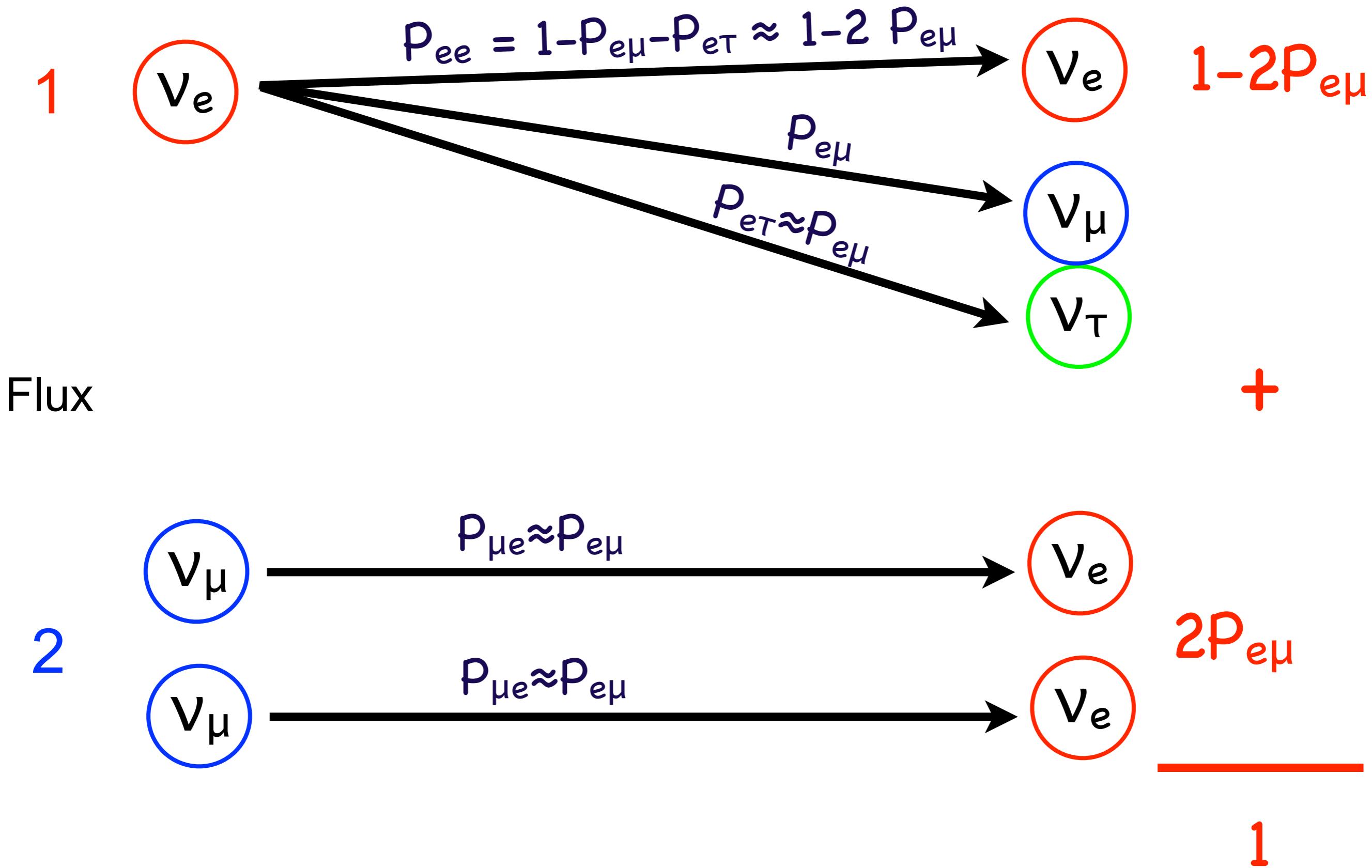
for $E_\nu < 3 \text{ GeV}$

Produced by cosmic ray proton interacting in earth atmosphere:

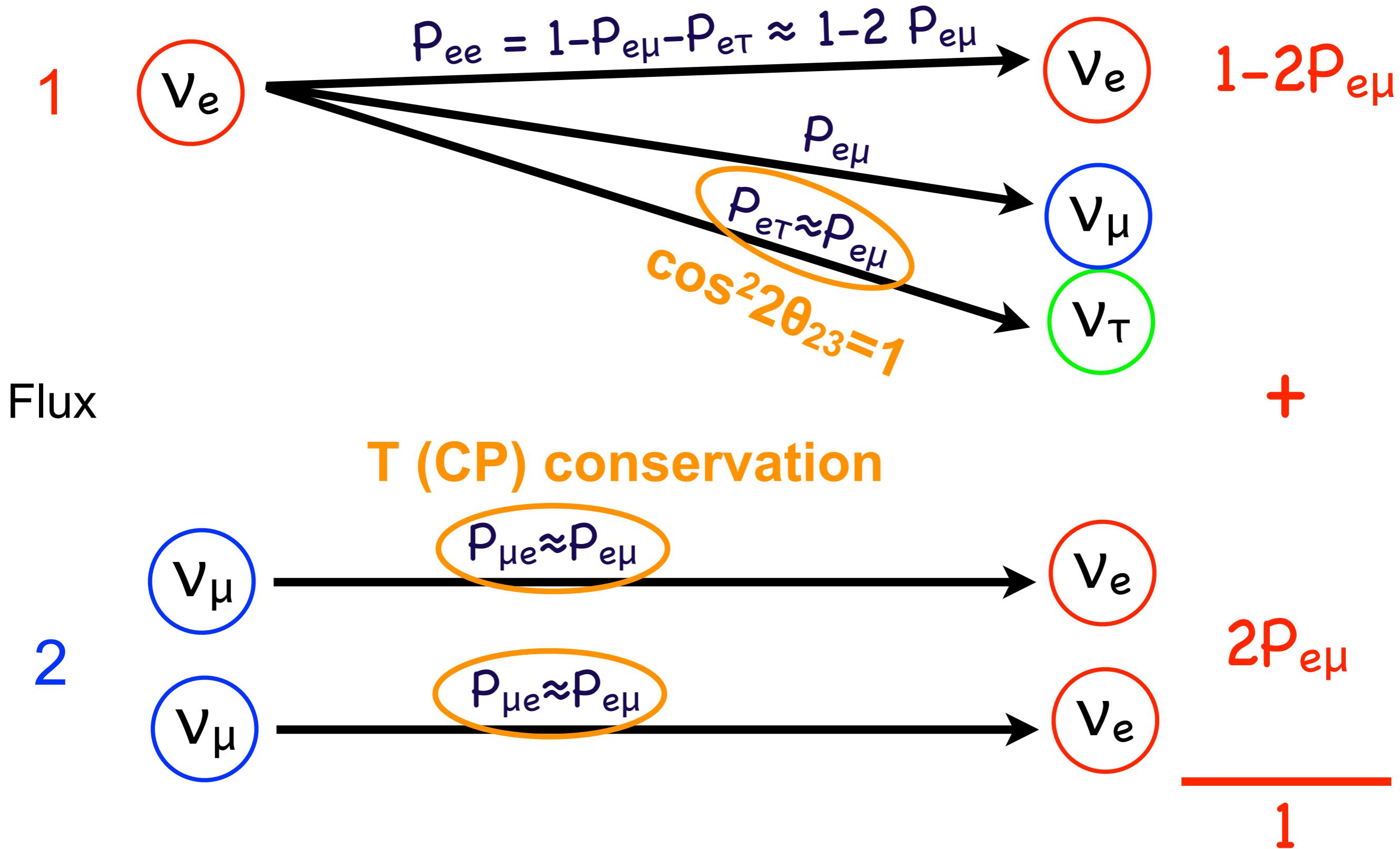


arXiv:1102.2688

Cancellation of atm. $\bar{\nu}_e$ appearance



atm. $\bar{\nu}_e$ app. sensitive to T(CP) viol.



atm. $\bar{\nu}_e$ app. sensitive to T(CP) viol.

