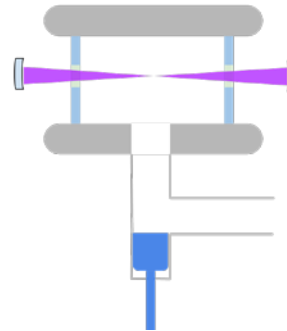
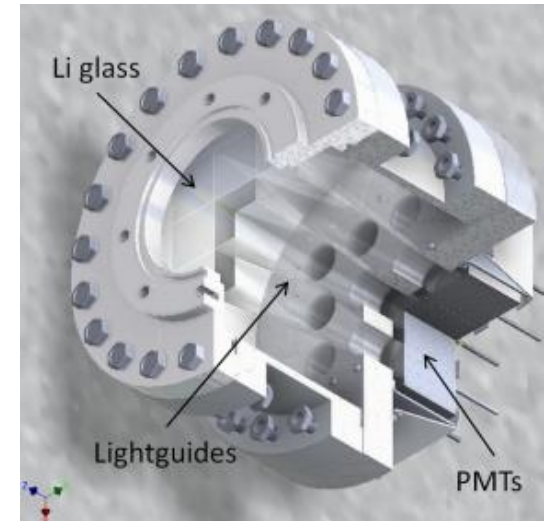
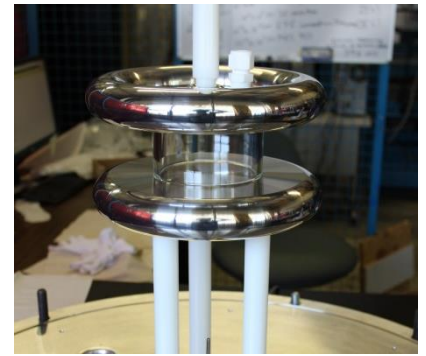


# Testing Fundamental Symmetries with the Next Generation Ultracold Neutron Source at TRIUMF

Russell Mammei

The University of Winnipeg



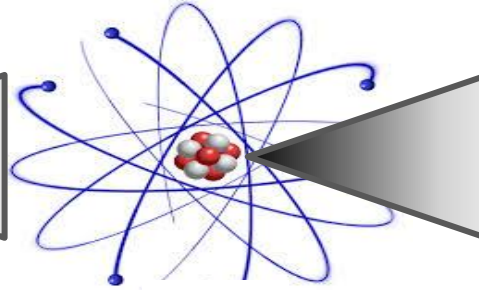
# THE STANDARD MODEL (SM)



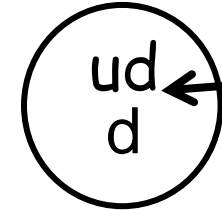
soda can



molecules

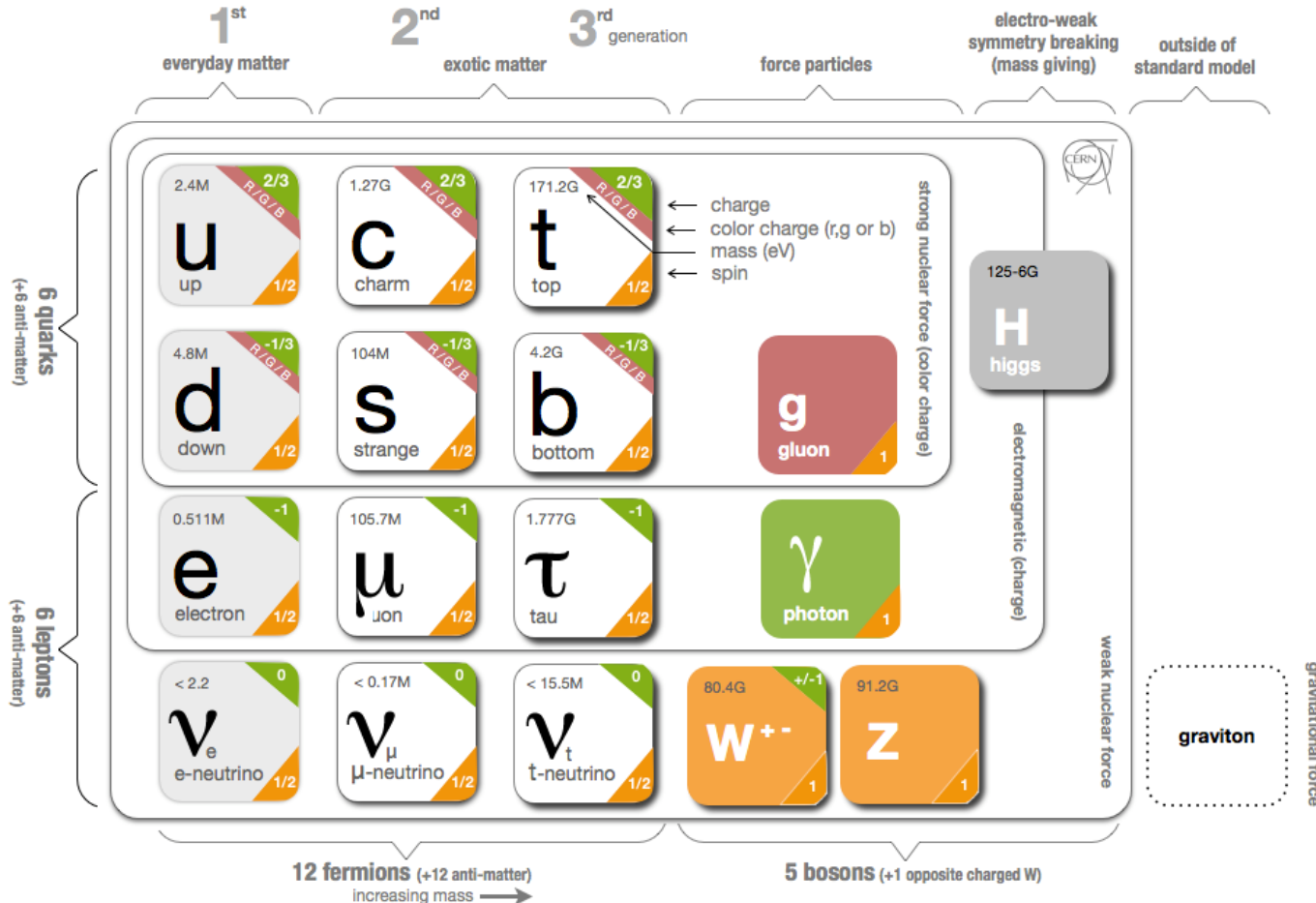


atom



neutron

quarks



Infographic from CERN's 2012 WEBFEST  
<http://cds.cern.ch/journal/CERNBulletin/2012/35/News%20Articles/1473657>

# DISCRETE SYMMETRIES

## Continuous Symmetries:

- Translation in space → momentum conservation
- Translation in time → energy conservation
- Rotation → angular momentum conservation

## Discrete Symmetries:

- Spatial Inversion (P) → P-invariance (parity)
- Charge Conjugation (C) → C-invariance
- Time reversal (T) → T-invariance

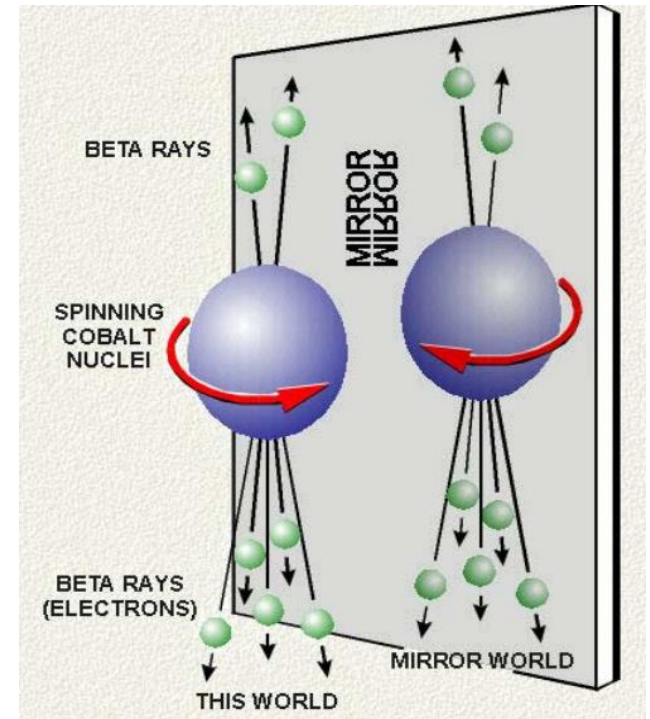
But Wait...

- Parity violation discovered in 1950s
- Charge parity (CP) violation discovered in 1960s
  - kaon sector (1964, 1990s)
  - B meson sector (2000's)

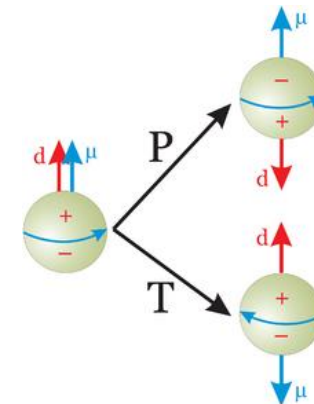
See. D. London's talk Thursday

A neutron electric dipole moment would be a violation of both parity and time reversal invariance

CPT still thought to be not violated ("good symmetry")

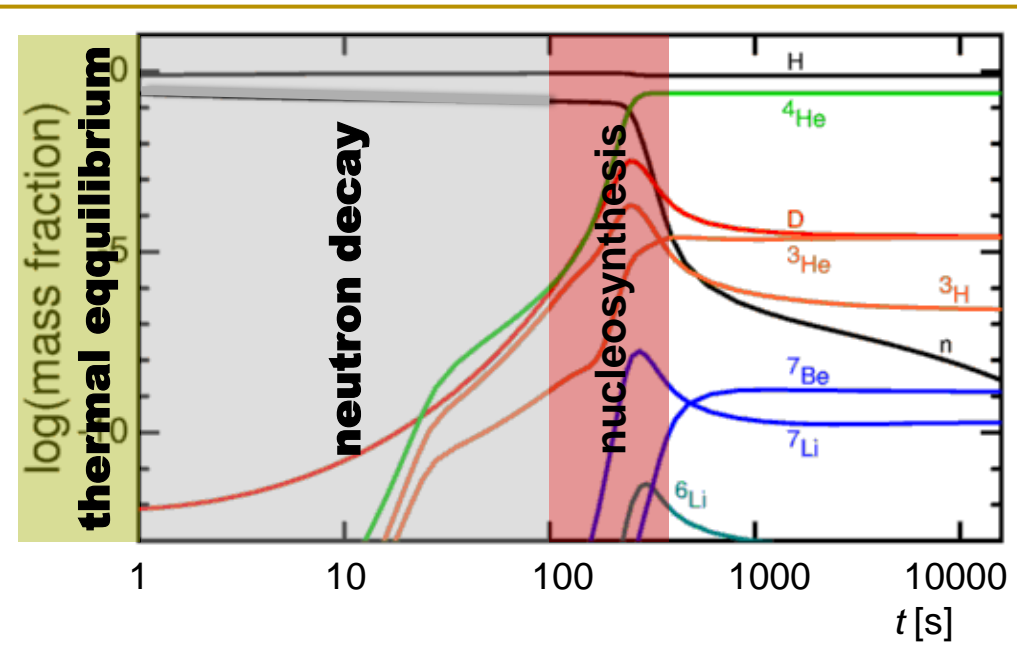
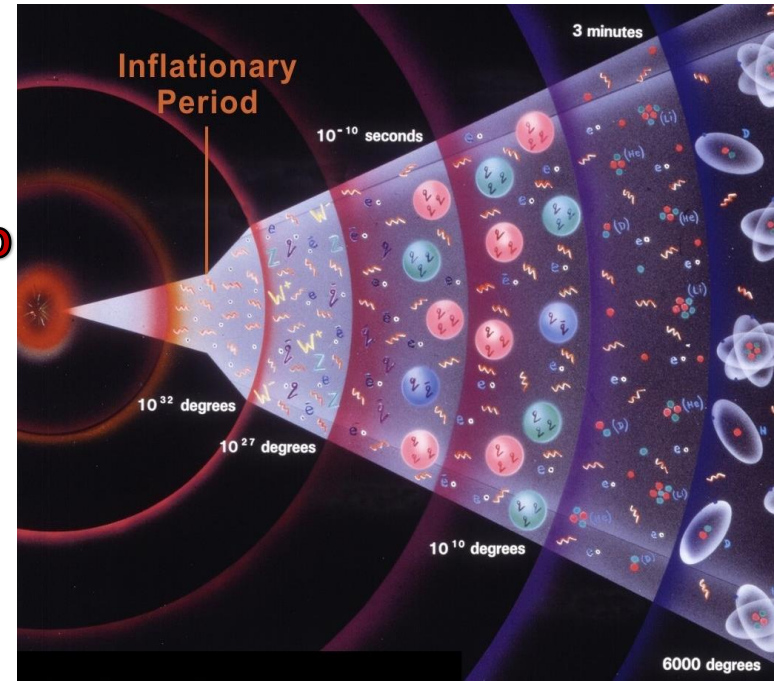


parity (P) violation in  $\beta$ -decay  
(Madam Wu, 1957)



# THE BIG QUESTIONS

- Baryon asymmetry of the universe**
  - Sakharov Criteria
  - EW Baryogenesis (Sphalerons) ✓
  - Departure from thermal equilibrium ✓
  - CP violation –need more **nEDM would help**
- Number of quark flavors/generations** (CKM unitarity) **Neutron Decay**
- Predictions of elemental abundances in the universe**
- Testing Short-range Gravitation Interactions**



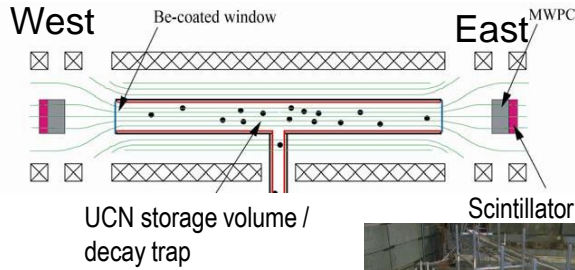
CKM Matrix

$$\begin{pmatrix} d' \\ s' \\ b' \end{pmatrix} = \begin{pmatrix} V_{ud} & V_{us} & V_{ub} \\ V_{cd} & V_{cs} & V_{cb} \\ V_{td} & V_{ts} & V_{tb} \end{pmatrix} \begin{pmatrix} d \\ s \\ b \end{pmatrix}$$

$$|V_{ud}|^2 + |V_{us}|^2 + |V_{ub}|^2 = 1 \quad \text{Unitarity}$$

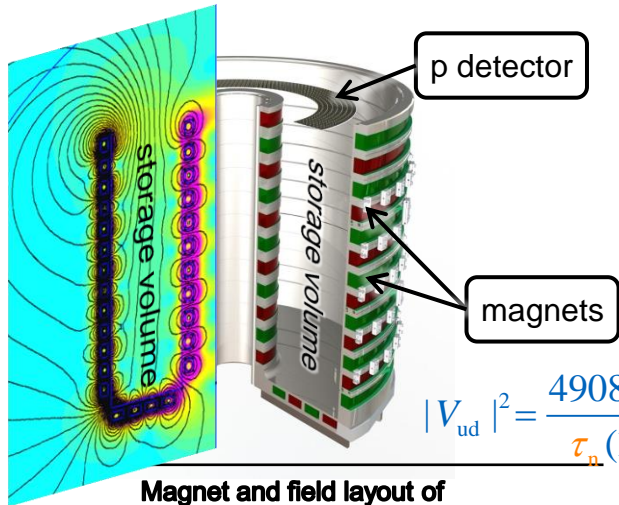
# UCN EXPERIMENTS

## Neutron decay correlations



$$A = -2 \frac{|\lambda|^2 + \Re(\lambda)}{1 + 3|\lambda|^2}$$

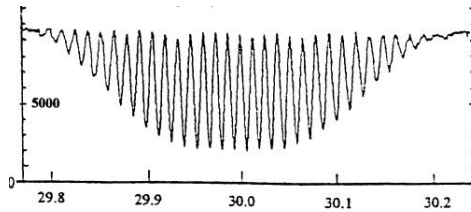
## Neutron lifetime



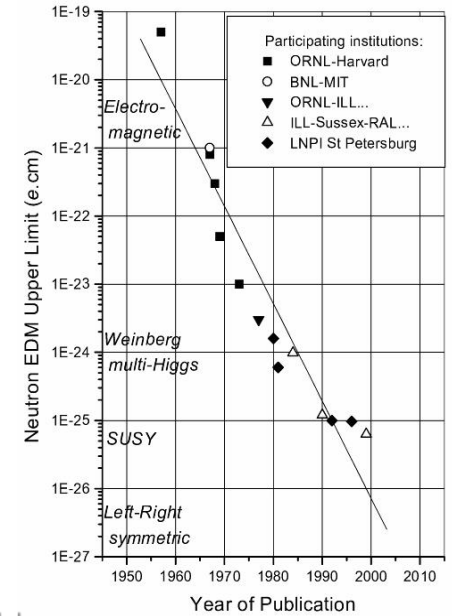
$$|V_{ud}|^2 = \frac{4908.7(1.9) \text{ s}}{\tau_n (1 + 3\lambda^2)}$$

Magnet and field layout of PENeLOPE

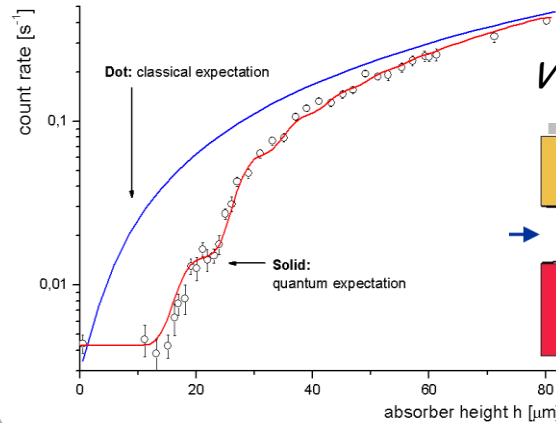
## Electric dipole moment



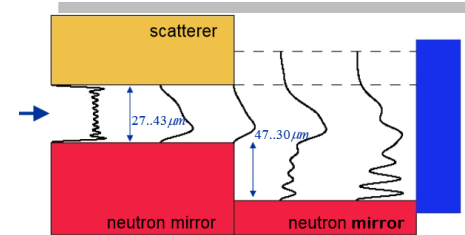
RAL Sussex EDM @ ILL



## Gravity



$$V(r) = G \frac{m_1 \cdot m_2}{r} (1 + \alpha \cdot e^{-r/\lambda})$$



# ULTRACOLD NEUTRONS (UCN)

Neutrons moving so slowly that they bounce off surfaces and can be bottled.

$$\begin{aligned}v &< 8 \text{ m/s} = 30 \text{ km/hr} \\KE &< 300 \text{ neV} \\ \lambda &> 50 \text{ nm}\end{aligned}$$

## Interactions

- **Strong interaction**

long wavelength samples over many atoms in materials  
⇒ average Fermi potential, total reflection

- **Electromagnetic Interaction**  $V_m = -\mu \cdot B = \pm 60 \text{ neV per Tesla}$

We can make beam beam of 100% polarized UCN !!!

- **Gravity**  $V_g = mgh \approx 100 \text{ neV per meter}$

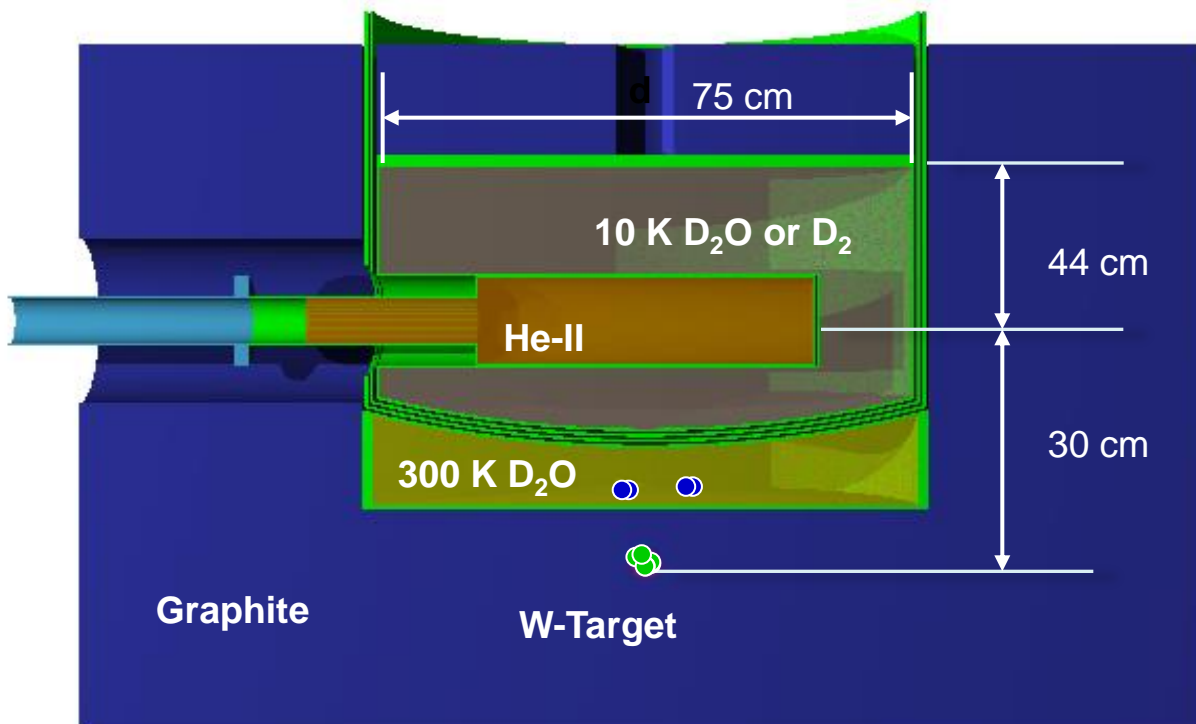
- **Weak Interaction**

$\beta$ -decay  $n \rightarrow p + e + \bar{\nu}_e, 728 \text{ keV lifetime } \sim 15 \text{ min}$

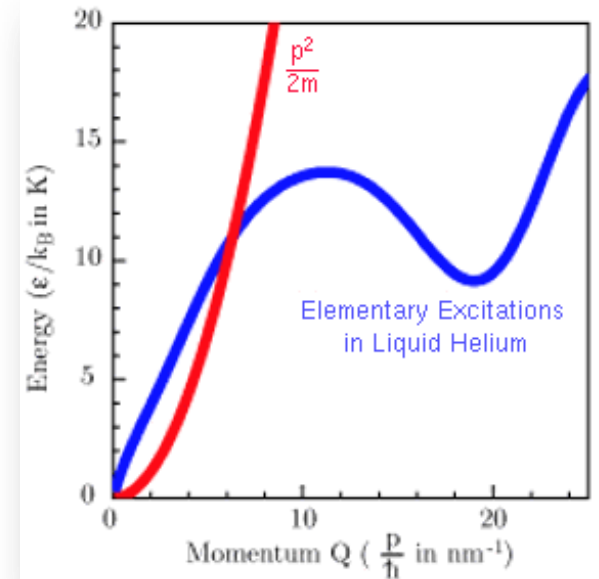
**Can store/transport UCN on times comparable to lifetime** 6

# HOW DO WE MAKE UCN AT TRIUMF?

- Spallation –Free the neutrons from W target
- Moderation –Cool the neutrons in  $D_2/D_2O$
- Conversion –Really cool the neutrons in He-II



Spallation target, thermal and cold moderator and He-II converter



MeV neutrons

(ultra-)cold neutrons

Technology developed at RCNP Osaka

480 MeV protons

# TRIUMF UCN HISTORY SO FAR

2006: UCN project was first introduced into the TRIUMF 5Y planning

2007: International Workshop UCN sources and Experiments at TRIUMF

2008: Positive review by TRIUMF's Experiments Evaluation Committee (EEC)

2009: **CFI-NIF Award for UCN Source**

2010: International Review endorses UCN program strongly

2011: MoU between Uwpg, KEK, RCNP and TRIUMF was signed...

- to build a He-II spallation source at KEK/RCNP and move it to TRIUMF
- to develop and conduct a neutron EDM experiment
- to build a dedicated beam line and target at TRIUMF

2011-2013: development of beam line in Meson hall

- Kicker, septum, bender, focusing elements, diagnostics, target
- Shielding upgrade
- clean-up of Meson hall

2012 Two new hires in Winnipeg that work on UCN

2013: TRIUMF hires are research scientist for UCN

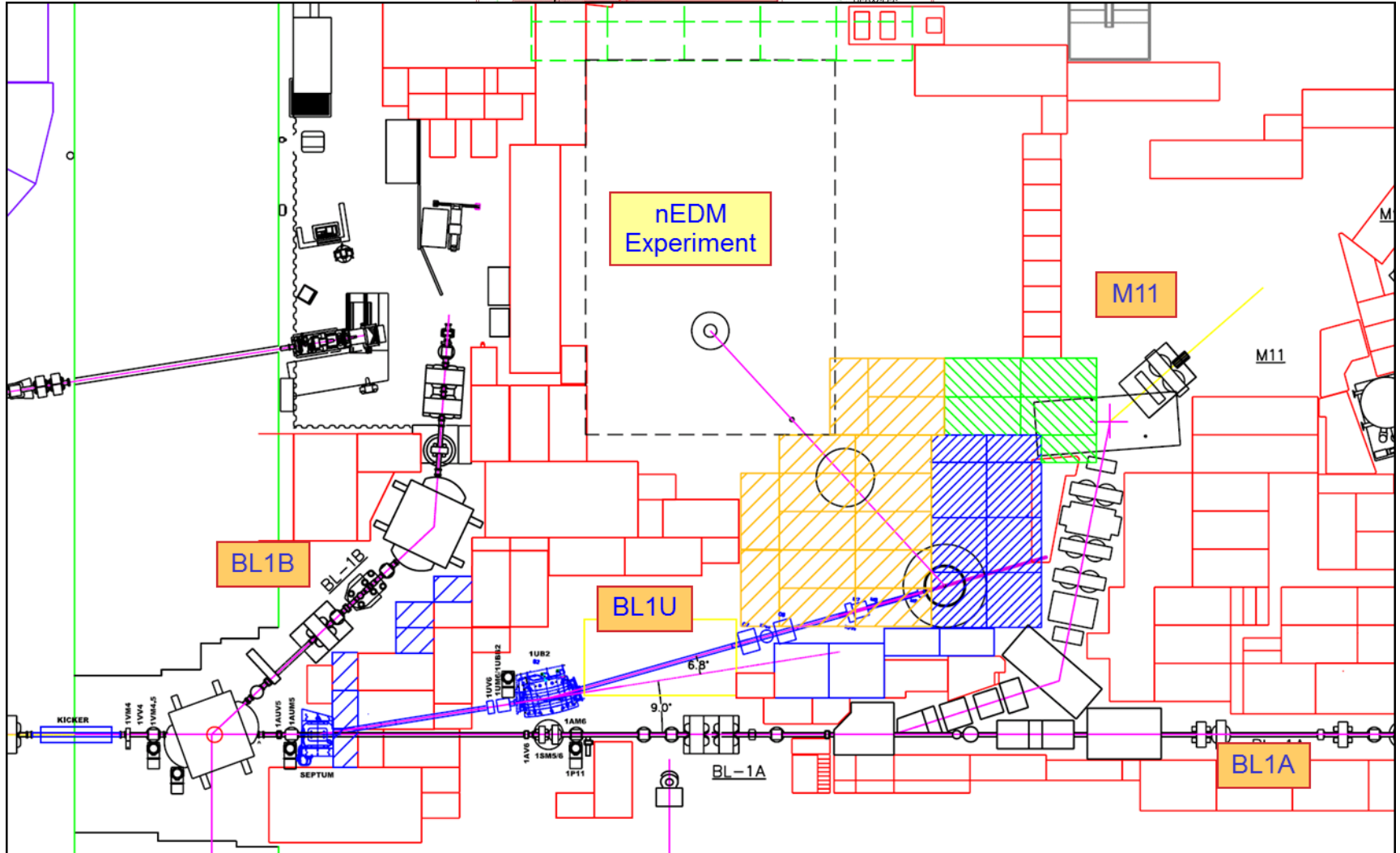
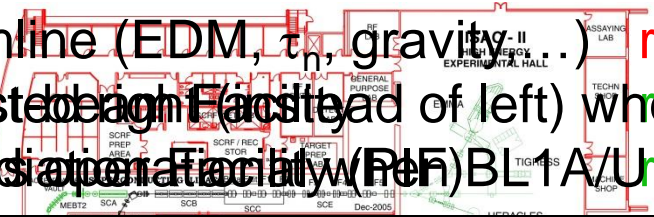
2014: First substantial installations occurred this spring

**Seeking CFI-IF for nEDM experiment, 2<sup>nd</sup> experiment port, & my coating facility**



# UCN FACILITY AT TRIUMF: MESON HALL

**BEAM LINES AND EXPERIMENTAL FACILITIES**  
 BL1B: UCN Beamline (EDM,  $\tau_n$ , gravity...) removed  
 M11: Detector deflected to right (instead of left) when space needed  
 BL1A: Photo irradiation facility (PFI) BL1A/U (off (Cyc. cooldown))



# INSTALLATION SCHEDULE

2015/16:

- target
- moderators
- He-II cryostat
- UCN guides
- UCN polarizer
- finish shielding

2015 & 2016

EDM experiment

TRIUMF

UC polarizer

Cold moderator cryostat

Spallation target

2015

Bending dipole

Lambertson septum

Kicker magnet

2014

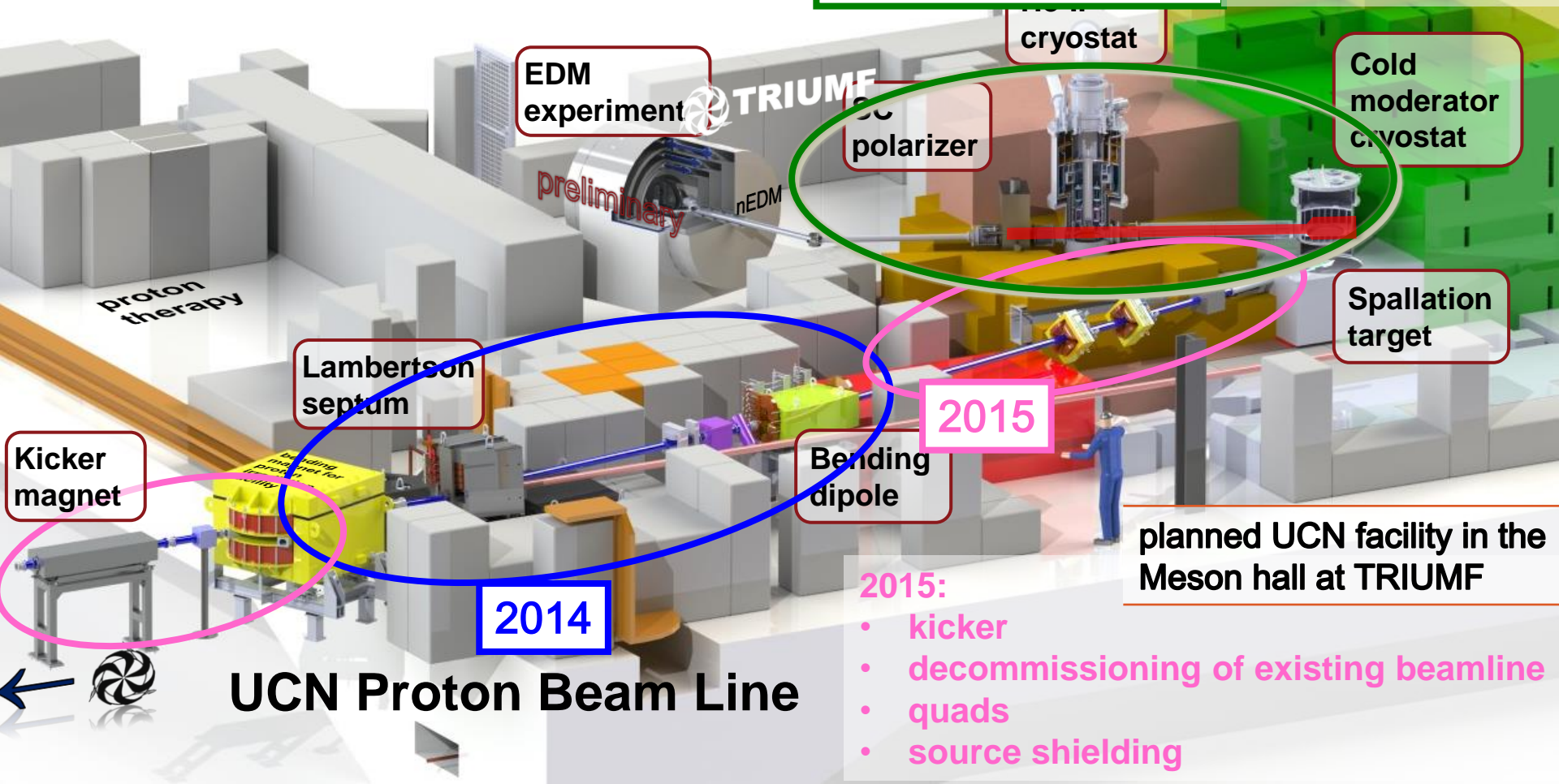
planned UCN facility in the Meson hall at TRIUMF

2015:

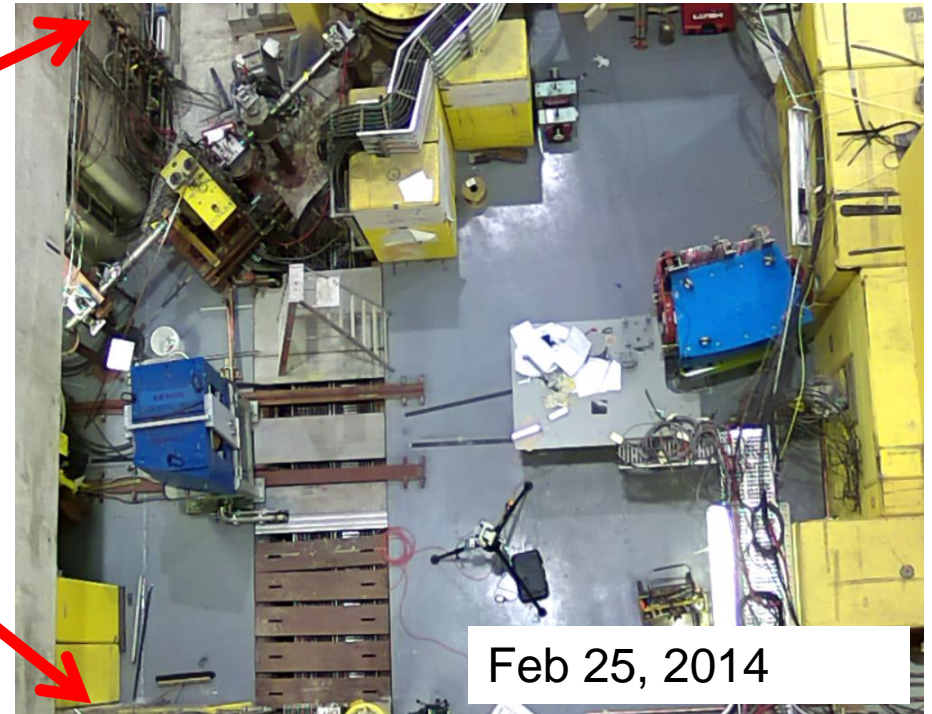
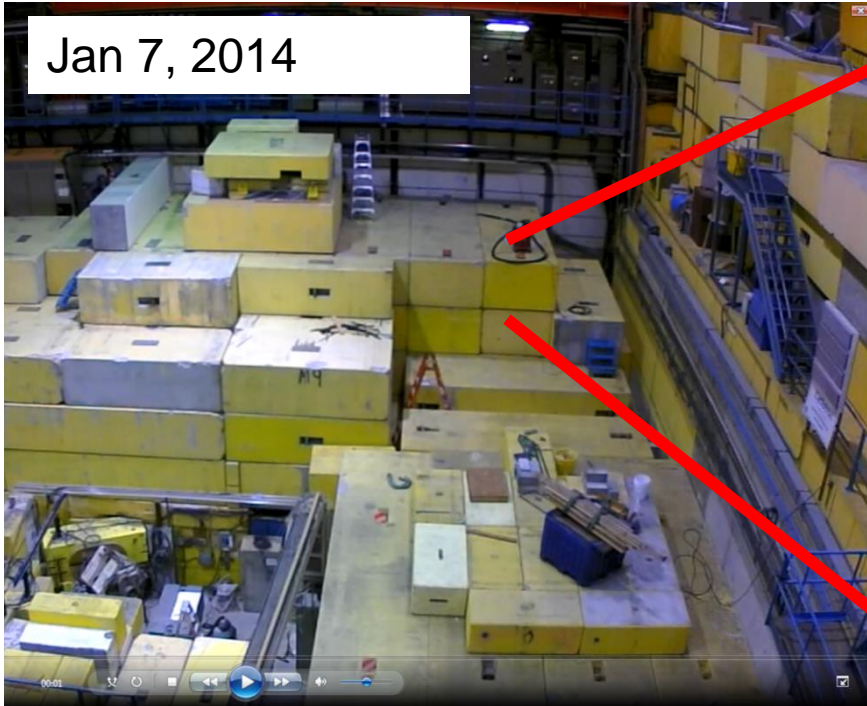
- kicker
- decommissioning of existing beamline
- quads
- source shielding

UCN Proton Beam Line

- 2014:
- septum
  - dipole
  - replacement of shielding towards cyclotron



# 2014 INSTALLATION



## Primary goals for 2014 installation met



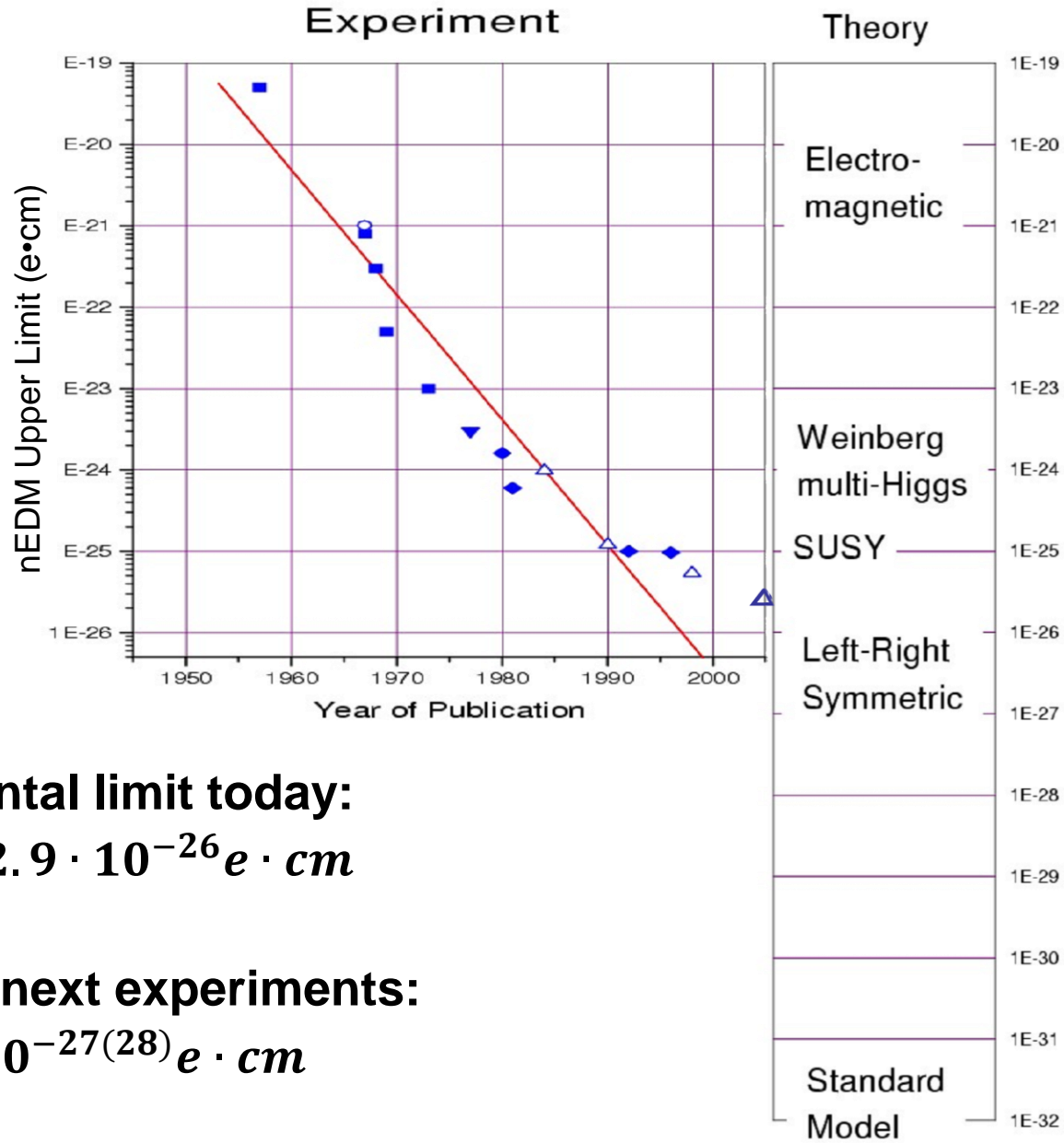
- Reconfigure Cyclotron shielding (Shield-Plug)
- Septum subsystem (1AM5: vacuum vessel only)
- Rough-in (trench) services
- BL1U → UCN-Dipole, girder, reconfiguration of BL1A

## “Best Efforts” goals 90-95 %



- Rough-in (non-trench) services
- Complete services
- BL1U girder components

# SEARCH FOR THE NEUTRON ELECTRIC DIPOLE MOMENT



**Experimental limit today:**

$$d_n < 2.9 \cdot 10^{-26} e \cdot \text{cm}$$

**Goals for next experiments:**

$$d_n < 10^{-27(28)} e \cdot \text{cm}$$

# nEDM EXPERIMENTAL SITES

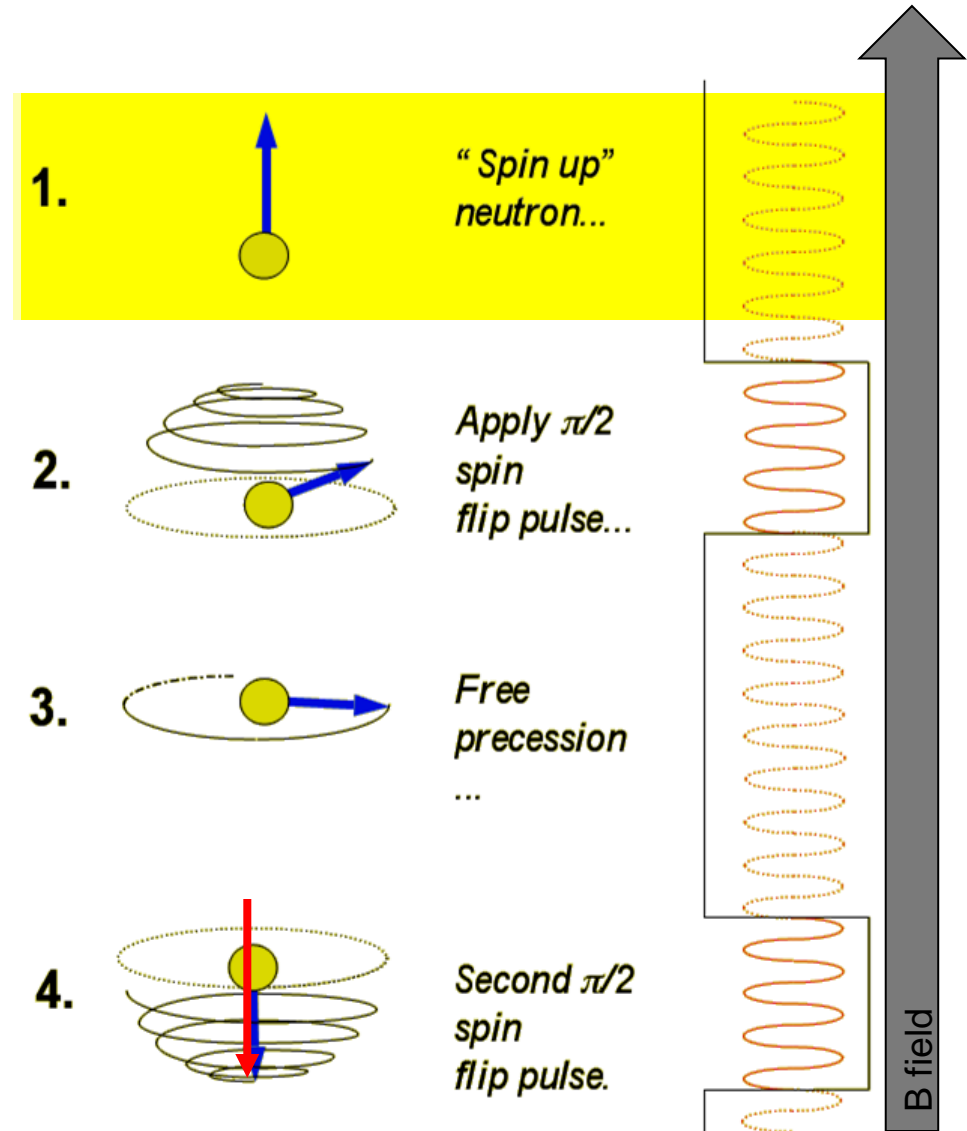


# RAMSEY'S METHOD

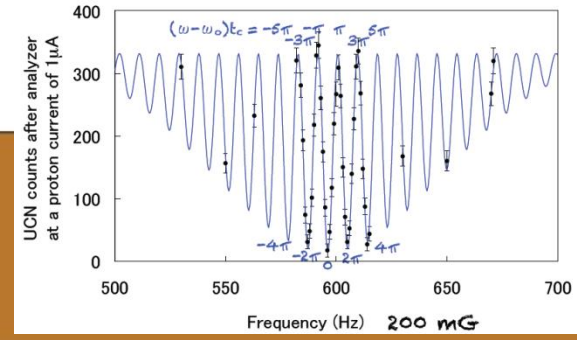
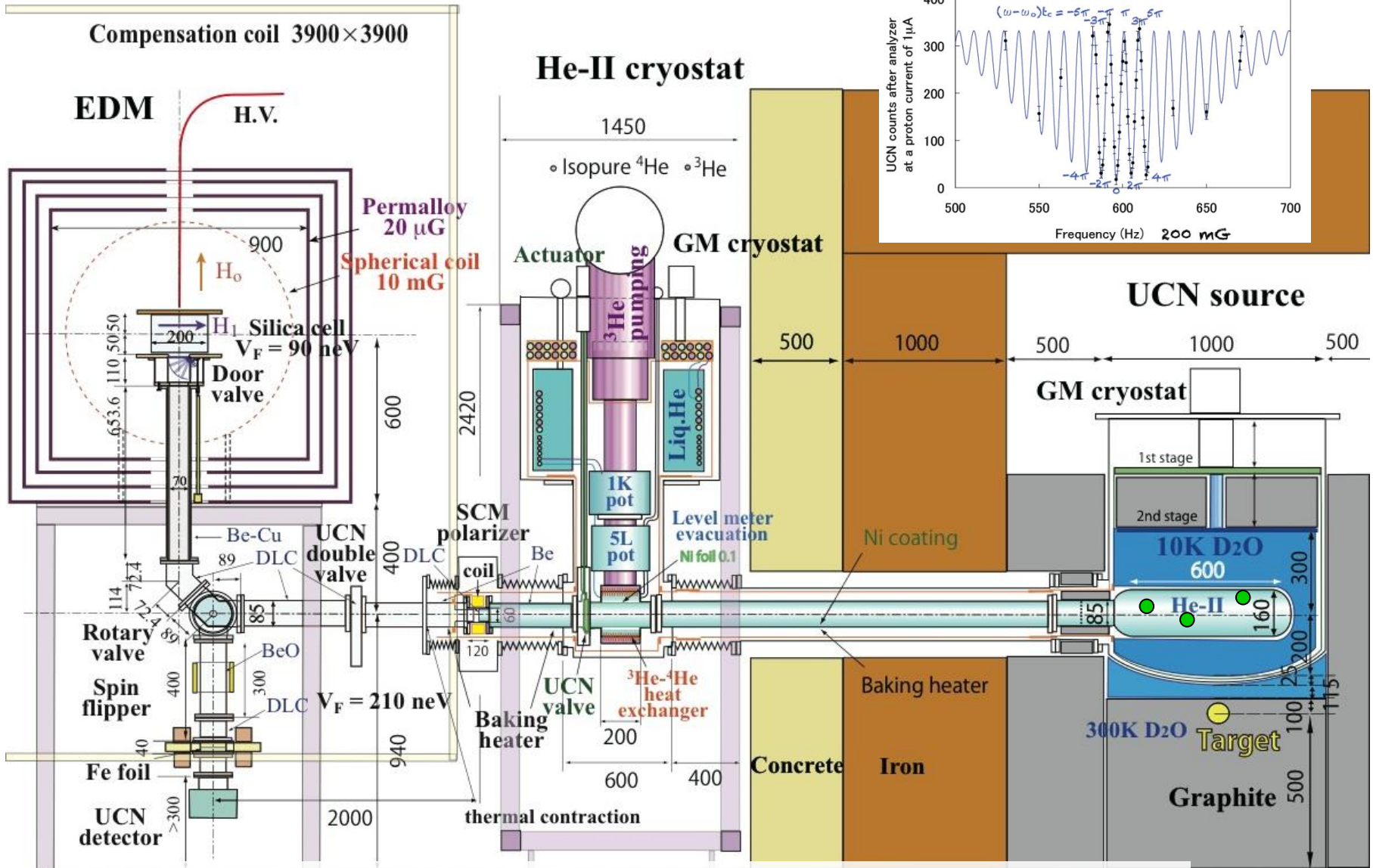
1. prepare a sample of polarized neutrons
2. make a  $\pi/2$  spin flip (“start clock”)
3. allow free spin precession in  
(anti-)parallel  $\mathbf{B}$  and  $\mathbf{E}$  static fields
4. make a  $\pi/2$  spin flip (“stop clock”)
5. analyze direction of neutron spin

look at energy (frequency) shift  
under Electric field inversion:

$$\Delta\varepsilon = h |\Delta\nu| = 4Ed_n$$



# THE EXPERIMENT



Look at energy (frequency) shift under E field inversion:

$$\Delta\varepsilon = h |\Delta\nu| = 4Ed_n$$

# ERROR BUDGET

Best nEDM limit so far is  $2.9 \cdot 10^{-26} \text{ e}\cdot\text{cm}$  (ILL/RAL/SUSSEX)

EDM statistical sensitivity:

$$\sigma_d = \frac{\hbar}{2\alpha ET\sqrt{N}}$$

$\alpha$ : visibility  
E: electric field  
T: observation time  
N: # of neutrons

ultra-cold neutrons are:

...totally reflected\*

⇒ long observation time  $T$

...enough from our new source

⇒ sufficient statistics  $\sqrt{N}$

...polarizable to 100%

⇒ good visibility  $\alpha$

Expect in one year

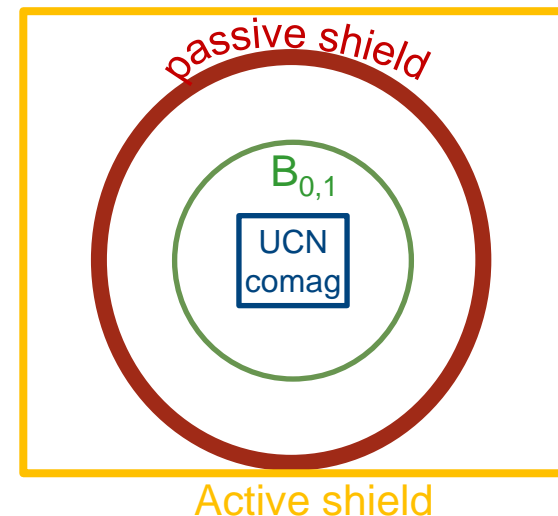
$$\sigma_d \sim 10^{-27} \text{ e}\cdot\text{cm}$$

Systematics Errors are Key

Biggest Error due in-homogeneities in the magnetic field (GPE effect)

Requirements for  $10^{-27} \text{ e}\cdot\text{cm}$

- $B_0 \sim 1 \mu\text{T}$
- Homogeneity  $< \text{nT/m}$   
⇒  $< 100 \text{ pT}$  across the cell
- Stability controlled to  $< \text{pT}$



\*by suitable materials under all angles of incidence



# CANADIAN EDM R&D

## Magnetic environment

- active shielding
- passive shielding
- creation of stable and homogeneous B fields
- magnetometry



Cylindrical shells of the 4 layer passive shield

Univ. Winnipeg

## UCN detection



- conventional  ${}^3\text{He}$  detectors too slow
- high rate capability
- Li glass scintillator + lightguide + PMTs



Univ. Winnipeg, Univ. Manitoba, TRIUMF

## Dual Co-magnetometer

	${}^{199}\text{Hg}$	${}^{129}\text{Xe}$	$n$
Spin	$\frac{1}{2}$	$\frac{1}{2}$	$\frac{1}{2}$
$\gamma$ (MHz/T)	7.65	-11.7	-29.16
UCN capt. $\sigma$ (bars)	150	21.0	
transition (nm)	253.7 nm	32.4 nm	
transition process	one-photon	two-photon	

Univ. Brit. Col., Simon Fraser Univ.

## Optical Magnetometry

- Conventional Sensors  
~ 1-10 pT sensitivity
- NMR  
Sensor  
  - All NMR
  - ~ fT Sensitivity



Univ. Winnipeg

# CONCLUSIONS

- ❑ Good physics can be done with neutrons
  - ❑ Search for new sector of CP violation (nEDM)
  - ❑ Search for beyond standard model interactions (via neutron decay)
- ❑ TRIUMF will have the best UCN source in the world
  - ❑ Proton beamline installation started this year
  - ❑ On track to do commissioning in 2016
  - ❑ first measurements in 2017
- ❑ nEDM at TRIUMF
  - ❑ Subsystem R&D well underway
  - ❑ Seeking CFI-IF
  - ❑ High discovery potential of a nEDM

# UCN CANADIAN COLLABORATION

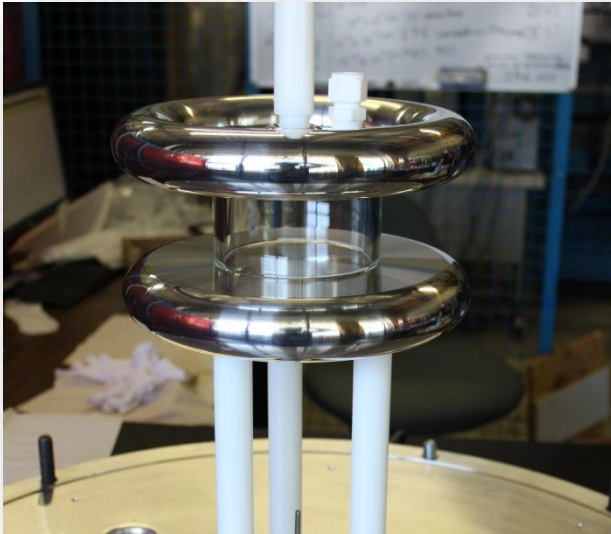
T. Adachi<sup>1</sup>, E. Altieri<sup>2</sup>, T. Andalib<sup>3,4</sup>, C. Bidinosti<sup>3</sup>, J. Birchall<sup>4</sup>, C. Davis<sup>5</sup>, F. Doresty<sup>4</sup>,  
W. Falk<sup>4</sup>, M. Gericke<sup>4</sup>, K. Hatanaka<sup>6</sup>, B. Jamieson<sup>3</sup>, S.C. Jeong<sup>1</sup>, D. Jones<sup>2</sup>, K. Katsika<sup>5</sup>,  
S. Kawasaki<sup>1</sup>, A. Konaka<sup>5</sup>, E. Korkmaz<sup>7</sup>, M. Lang<sup>3,4</sup>, L. Lee<sup>4,5</sup>, K. Madison<sup>2</sup>, J. Mammei<sup>4</sup>,  
R. Mammei<sup>3</sup>, J.W. Martin<sup>3</sup>, Y. Masuda<sup>1</sup>, R. Matsumiya<sup>6</sup>, K. Matsuta<sup>8</sup>, M. Mihara<sup>8</sup>,  
C.A. Miller<sup>5</sup>, E. Miller<sup>2</sup>, K. Mishima<sup>10</sup>, T. Momose<sup>2</sup>, W.D. Ramsay<sup>5</sup>, S.A. Page<sup>4</sup>,  
R. Picker<sup>5</sup>, E. Pierre<sup>8,5</sup>, L. Rebenitsch<sup>3,4</sup>, J. Sonier<sup>9</sup>, I. Tanihata<sup>6,10</sup>, W.T.H. van Oers<sup>4,5</sup>,  
Y. Watanabe<sup>1</sup>, and J. Weinands<sup>2</sup>

*<sup>1</sup>KEK, <sup>2</sup>UBC, <sup>3</sup>Winnipeg, <sup>4</sup>Manitoba, <sup>5</sup>TRIUMF,  
<sup>6</sup>RCNP Osaka, <sup>7</sup>UNBC, <sup>8</sup>Osaka, <sup>9</sup>SFU, <sup>10</sup>Beihan*

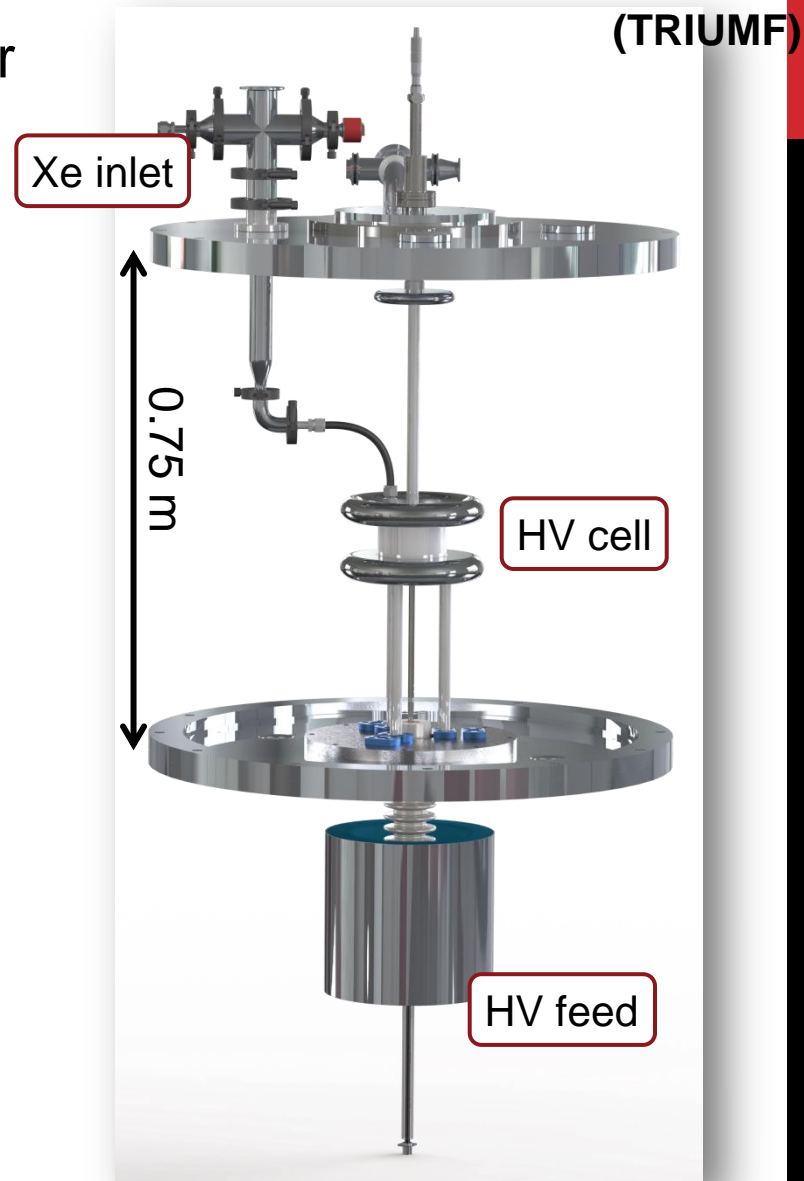
BACKUP SLIDES

# EDM CELL AND ELECTRIC FIELD

- dielectric strength of Xe at  $10^{-3}$  mbar unknown
- HV test setup at TRIUMF
- 50x100 mm cylindrical test cell
- field strength goal  $> 10$  kV/cm
- test of different cell materials
- **commissioned 8/2013**



**HV/EDM cell mock-up at TRIUMF**



**HV setup at TRIUMF**

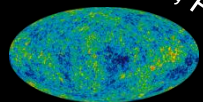
# Parameters of Big Bang Nucleosynthesis

$$Y_P = 0.228 + 0.023 \log \eta_{10} + 0.012 N_\nu + 0.018 (\tau_n - 10.28)$$

cosmic helium abundance  $Y_P$  from old stars

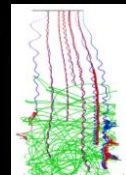


cosmic baryon density  $(n_B/n_\gamma) \cdot 10^{-10}$  WMAP, Planck



# of neutrino flavors 3

neutron lifetime PENELOPE

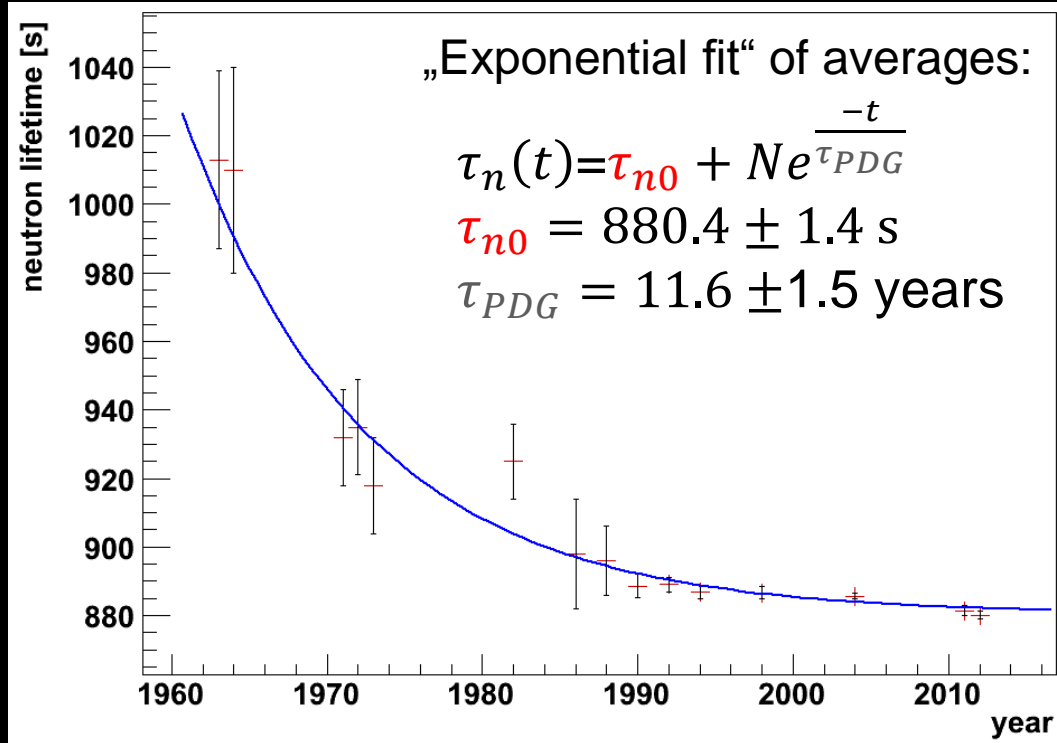


- the particle data group (PDG) reviews all major particle properties annually <http://pdg.lbl.gov/>
- PDG „world“ averages of the neutron lifetime for the last 50 years

## ⇒ PENELOPE

(Precision Experiment on the Neutron Lifetime Operating with Proton Extraction)

- Combination of **magnetic storage of ultra-cold neutrons** and in-situ **proton detection**
- Large volume
- Blind analysis
- Many knobs to turn to investigate systematic effects



# Transport and Storage of UCN

Fermi (Material) Potential

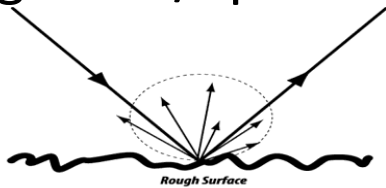
$$V = \frac{2\pi\hbar^2}{m} Nb \quad \text{UCN flux} \propto (V)^{3/2}$$

UCN Loss Probability

- Neutron absorption
- Inelastic upscattering

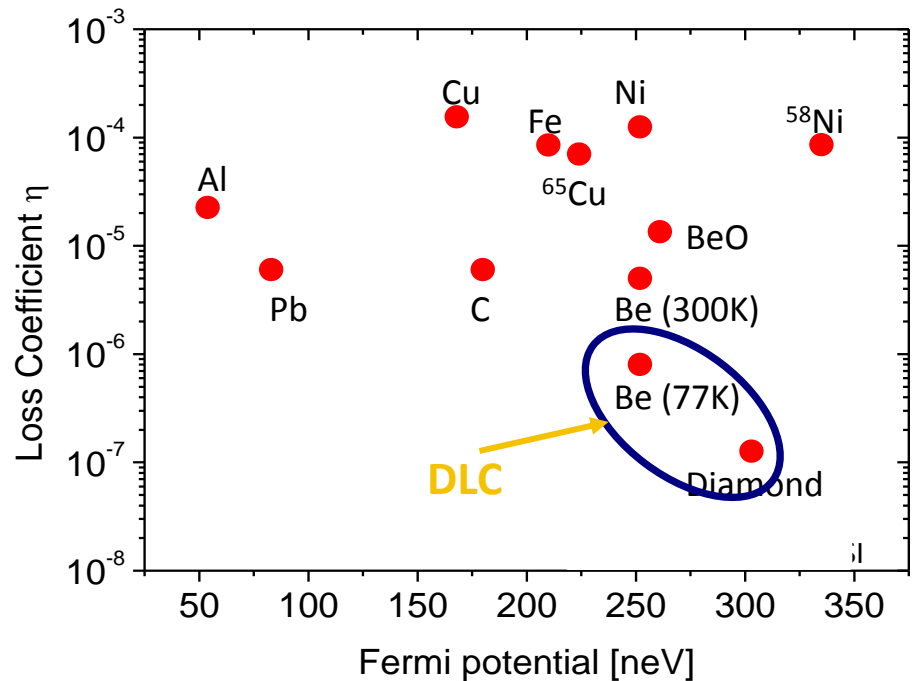
Depolarization Probability

Roughness/Specular Reflection

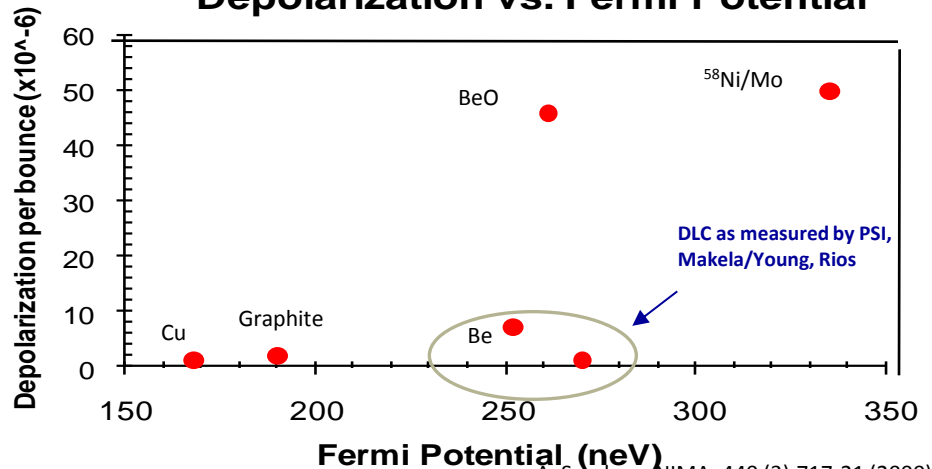


Usually means applying a coating on industrial, machined components

Need dedicated facility



Depolarization vs. Fermi Potential

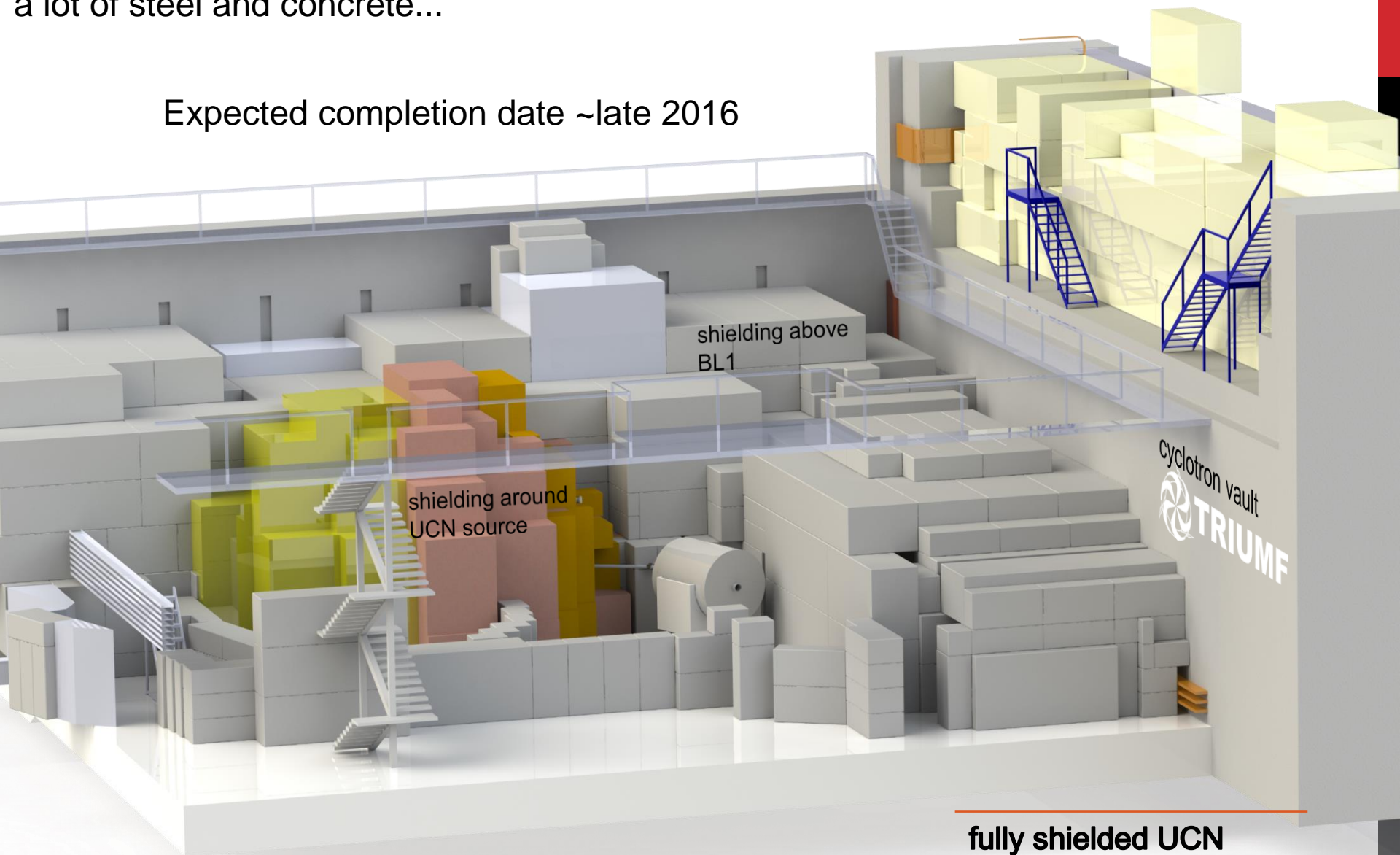


A. Serebrov NIMA, 440 (3):717-21 (2000)  
A. Serebrov PLA, 313 (5-6):373-79 (2003)

# UCN SOURCE FULLY SHIELDED

a lot of steel and concrete...

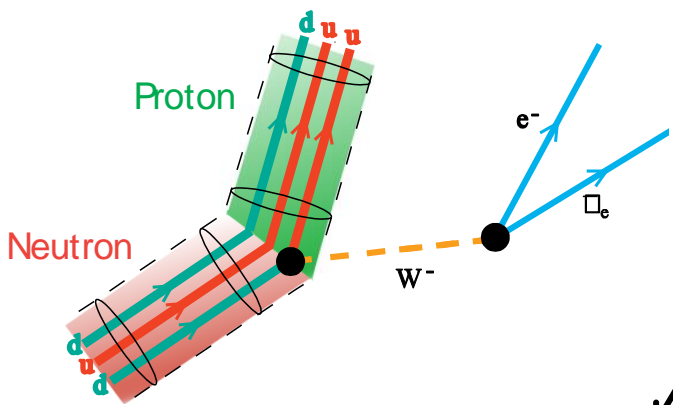
Expected completion date ~late 2016



fully shielded UCN  
source



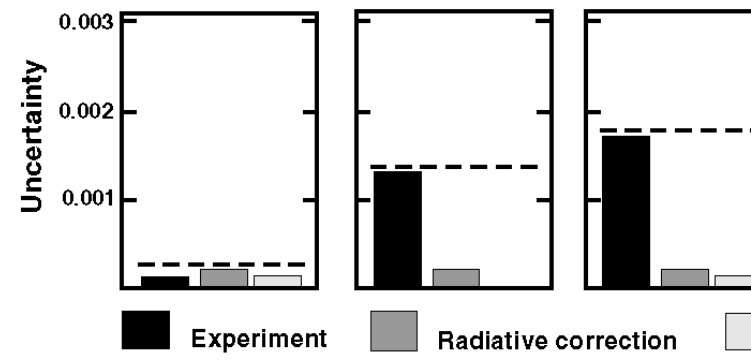
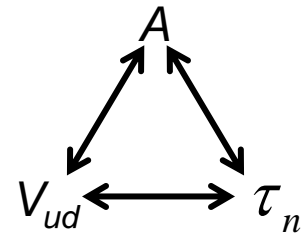
# Neutron Decay



$$\begin{pmatrix} d' \\ s' \\ b' \end{pmatrix} = \begin{pmatrix} V_{ud} & V_{us} & V_{ub} \\ V_{cd} & V_{cs} & V_{cb} \\ V_{td} & V_{ts} & V_{tb} \end{pmatrix} \begin{pmatrix} d \\ s \\ b \end{pmatrix} = V_{CKM} \begin{pmatrix} d \\ s \\ b \end{pmatrix}$$

nuclear  $0^+ \rightarrow 0^+$       neutron      nuclear mirrors

$|V_{ud}|$



## “Master Formula”

$$\frac{1}{\tau_n} = W = (G_V V_{ud})^2 g_V^2 \left( 1 + 3 \left( \frac{g_A}{g_V} \right)^2 \right) (1 + \Delta_R) f_n p_e E_e (E_0 - E_e)^2 \left[ 1 + m_e b \frac{J_b}{f_n} \right]$$

## Angular Correlations (directional distribution shown)

$$\frac{dW}{d\Omega_e d\Omega_{\nu} dE_e} \propto p_e E_e (E_0 - E_e)^2 \left[ 1 + a \frac{\vec{p}_e \cdot \vec{p}_\nu}{E_e E_\nu} + b \frac{m_e}{E_e} + \left\langle \frac{\vec{J}_n}{J_n} \right\rangle \cdot \left( A \frac{\vec{p}_e}{E_e} + B \frac{\vec{p}_\nu}{E_\nu} + D \frac{\vec{p}_e \times \vec{p}_\nu}{E_e E_\nu} \right) \right]$$

## Big Bang Nucleosynthesis –neutron lifetime matters

### Parameter s of Big Bang Nucleosynthesis

$$Y_P = 0.228 + 0.023 \log_{10} \left( \frac{n_\nu}{n_B} \right) + 0.012 N_\nu + 0.018 (\tau_n - 10.28)$$

cosmic helium abundance from observations  
 cosmic baryon density  $(n_B/n_\gamma) \cdot 10^{-10}$  WMAP, Planck  
 # of neutrino flavors  $N_\nu$   
 neutron lifetime  $\tau_n$  PENeLOPE

## Conclusion

- 1) **Current Exps & Th:**  $|V_{ud}|^2 + |V_{us}|^2 + |V_{ub}|^2 = 0.9999(4)_{Vud}(4)_{Vus}$   
**Great Unitarity Test & Success → No New Physics!**

**Nuclear Isospin Breaking? Needs Resolution**

**Radiative Corrections Stable (*Unchallenged!*)**

- 2) **Neutron Decay:**  $|V_{ud}| = [4908.7(1.9)s/\tau_n(1+3g_A^2)]^{1/2}$  **clean & precise**

**Neutron Lifetime Controversy ( $6\sigma$  discrepancies)**

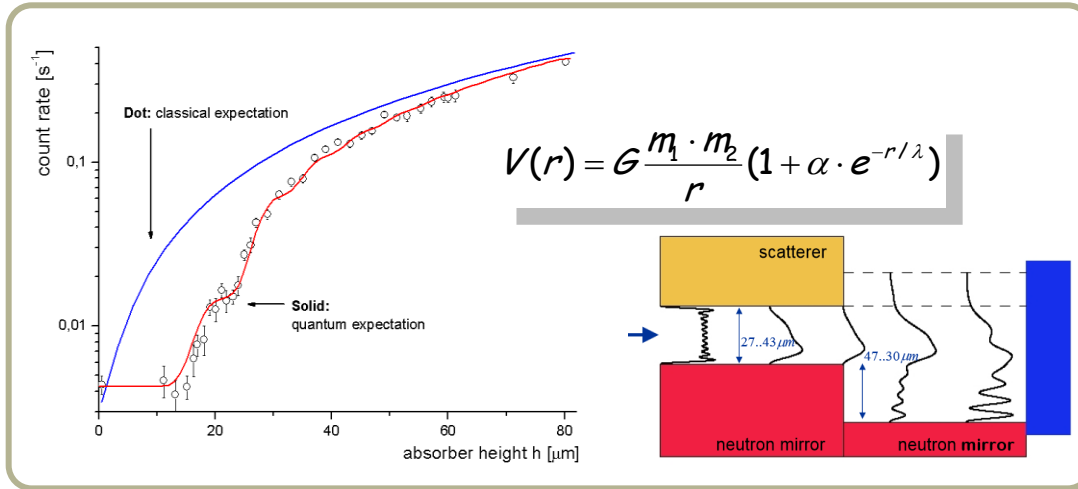
**2010  $\tau_n^{PDG} = 885.7(8)s$  vs  $\tau_n = 878.5(8)s$  Needs Resolution**

**$g_A$  larger? Perkeo Ave.  $1.2755(13)$  vs 2010  $g_A^{PDG} = 1.2695(29)$**

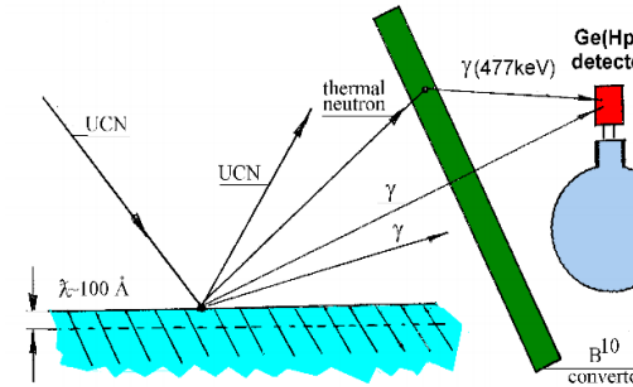
**Larger  $g_A$  & smaller  $\tau_n$  → Unitarity, solar neutrino flux, primordial nuclear abundances, proton spin, Goldberger-Treiman/Muon Capture, Bjorken Sum Rule, lattice calculation benchmark...**

***$V_{ud}$ ,  $\tau_n$  and  $g_A$  must be precisely determined!***

# Gravity



# Material science



Molecular rotor UCN inelastic scattering reflectometer get pictures

## Neutron Decay ( $n \rightarrow p e \bar{\nu}$ ) & $V_{ud}$

$$|V_{ud}|^2 = \frac{4908.7(1.9)\text{sec}}{\tau_n(1+3g_A^2)} \quad \text{Master Relation}$$

Measure  $\tau_n$  and  $g_A \equiv G_A/G_V$  (decay asymmetries)

2008 PDG  $\tau_n^{\text{ave}} = 885.7(8)\text{sec}$ ,  $g_A^{\text{ave}} = 1.2695(29)$

$$\Rightarrow |V_{ud}|^{\text{ave}} = 0.9746(4)_{\tau_n(18)g_A(2)}_{\text{RC}} \quad \text{reasonable but ...}$$

2012  $\tau_n^{\text{PDG}} \approx 880.1(1.1)\text{sec?}$  &  $g_A \approx 1.2755(13)$  Perkeo II

$$\Rightarrow |V_{ud}| = 0.9739(6)_{\tau_n(8)g_A(2)}_{\text{RC}}$$

**Agrees with superallowed!**  $0^+ \rightarrow 0^+$  Nuclear Beta  $V_{ud} = 0.97425(22)$

(Are  $\tau_n$  &  $g_A$  both shifting?)

History  $g_A = 1.18 \rightarrow 1.23 \rightarrow 1.25 \rightarrow 1.26 \rightarrow 1.27 \rightarrow 1.275?$

*Many New  $\tau_n$  &  $g_A$  Experiments Planned*

**Non-zero nEDM implies violation of time-reversal symmetry (T).**

**In QFT, this is equivalent to CP violation. (CPT=1)**

**The nEDM is sensitive to sources of CP violation:**

- within the Standard Model, via  $\theta_{\text{QCD}}$   $n\text{EDM} \approx 10^{-32}$  ecm
- beyond the Standard Model, e.g. as required by Electroweak Baryogenesis in order to generate the baryon asymmetry of the universe

Where is all the antimatter



**Sakharov Conditions:** (A.D. Sakharov, JETP Lett. 5, 24-27, 1967)

**(1) Baryon number violation (may imply proton decay)**

- **Baryon:** particle made out of 3 quarks (proton, neutron, lambda...)
- **proton is lightest baryon (uud), could only decay to leptons or mesons (2 quarks)**

**(2) Departure from thermal equilibrium**

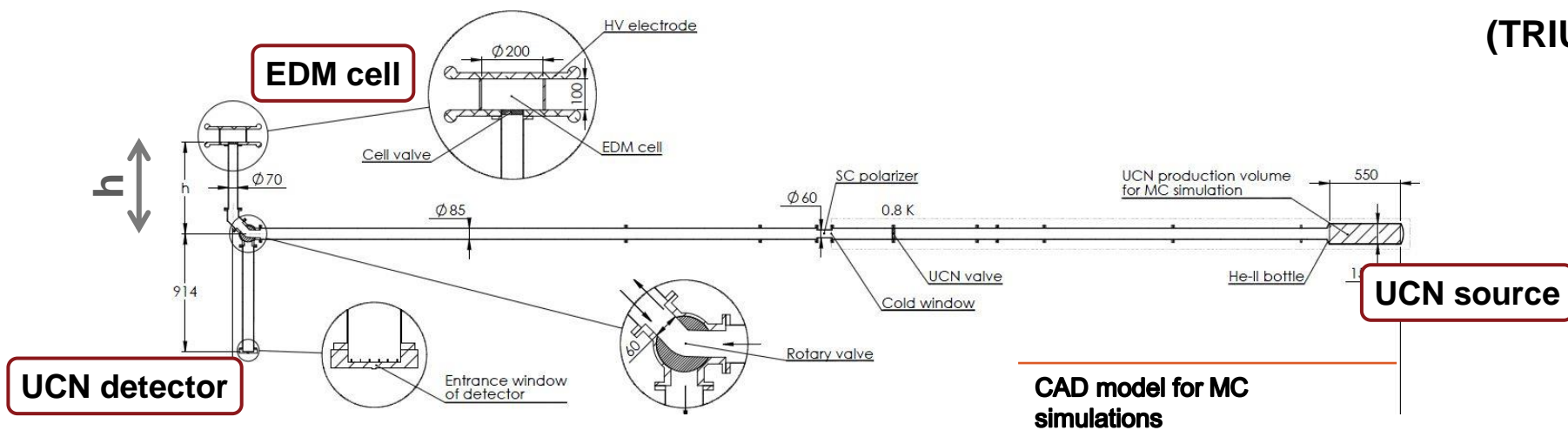
- **Phase transitions**
- **Expansion of the Universe (Inflation)**

**(3) Time reversal violation ( $\Rightarrow$  CP violation)**

- **not enough in Standard Model  $\Rightarrow$  electric dipole moment would help**

# MONTE CARLO SIMULATION: ONE EXAMPLE

(TRIUMF)



CAD model for MC simulations

Which height of EDM cell is best at what storage time?

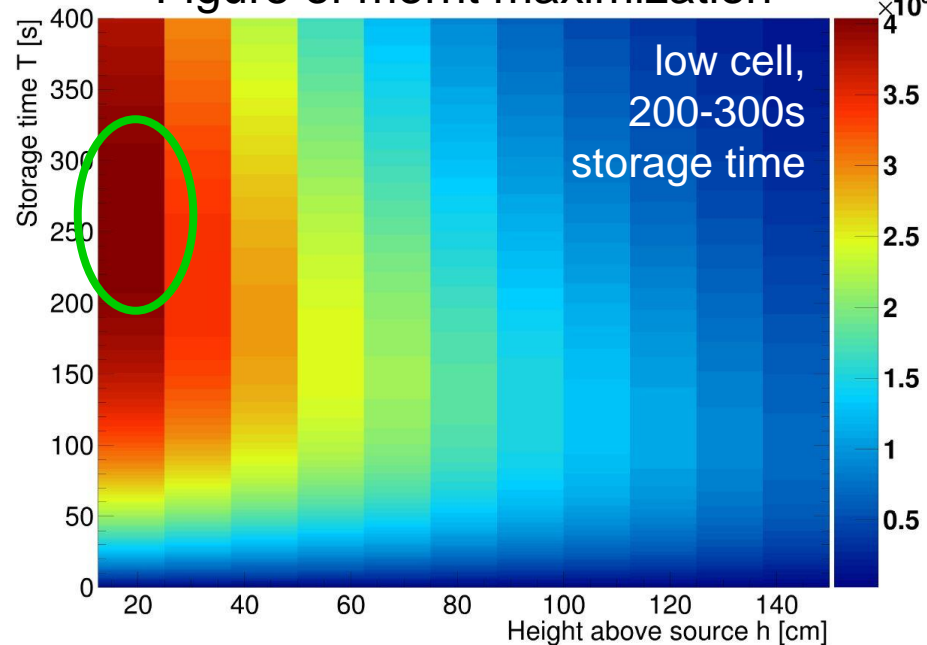
$$\sigma_{d_n} = \frac{\hbar}{2\alpha E T \sqrt{N}} \rightarrow \text{figure of merit} \rightarrow$$

$$\sqrt{N(T, h)} \cdot T = \sqrt{\sum_{\text{bin}=0}^n N_{0, \text{bin}}(h) \exp\left(-\frac{T}{\tau_{s, \text{bin}}}\right)} \cdot T$$

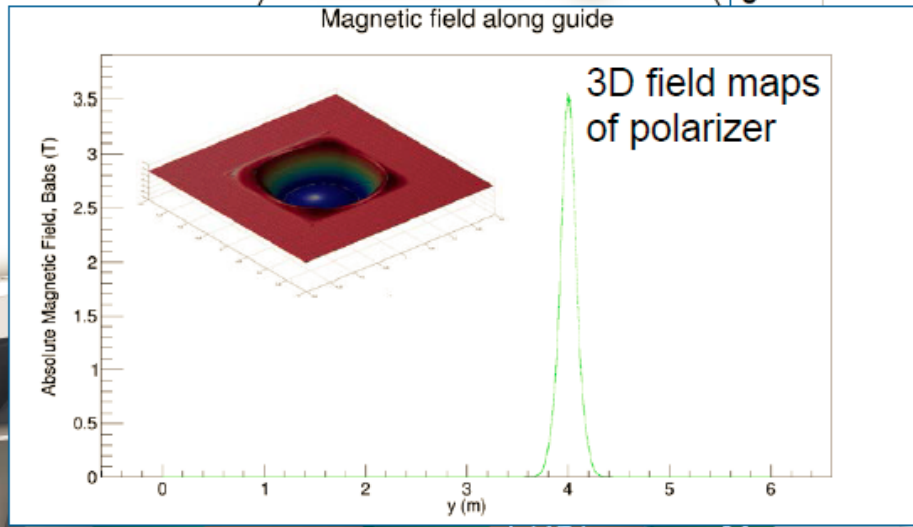
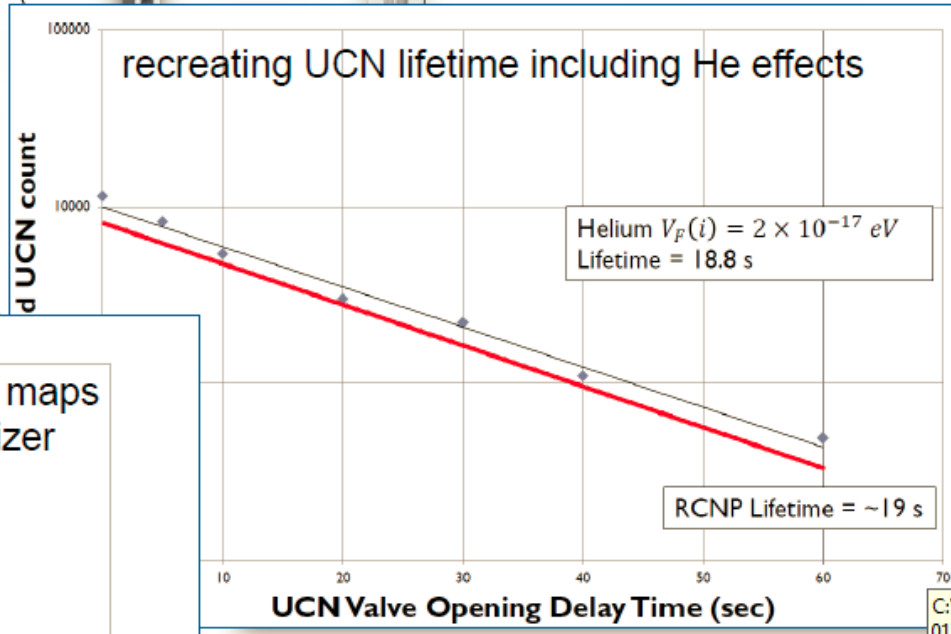
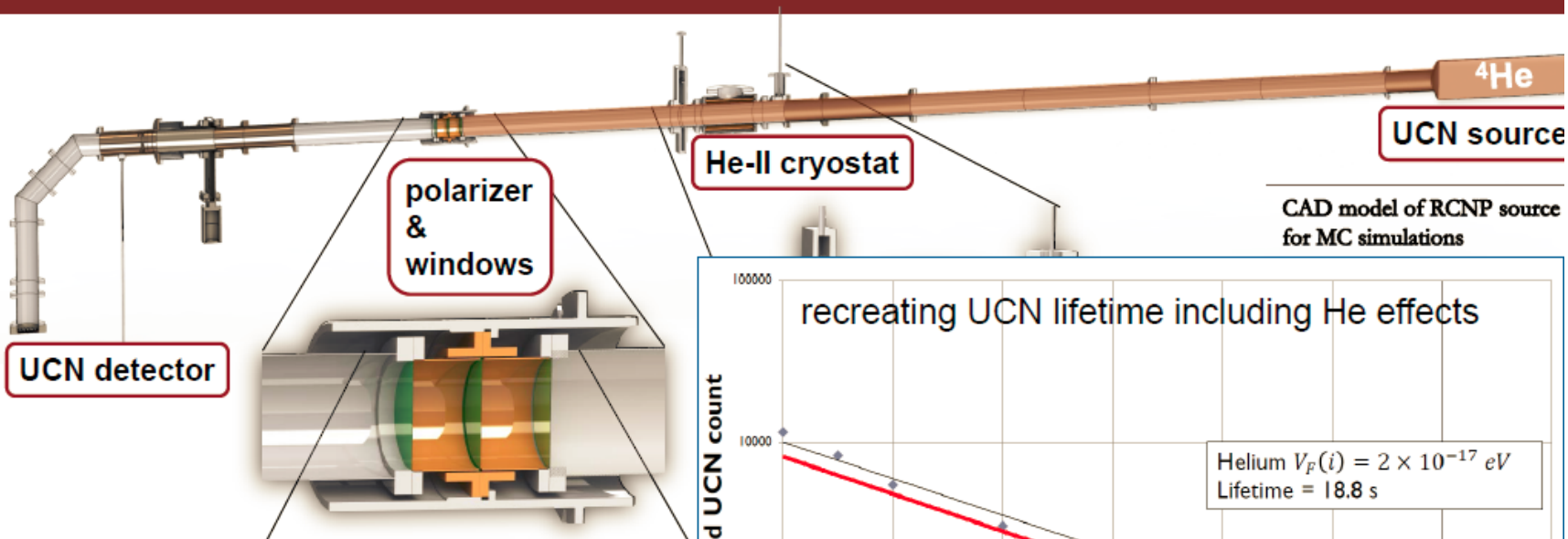
Plans to simulate:

- ⇒ depolarization
- ⇒ spin evolution
- ⇒ various GPEs

Figure of merit maximization

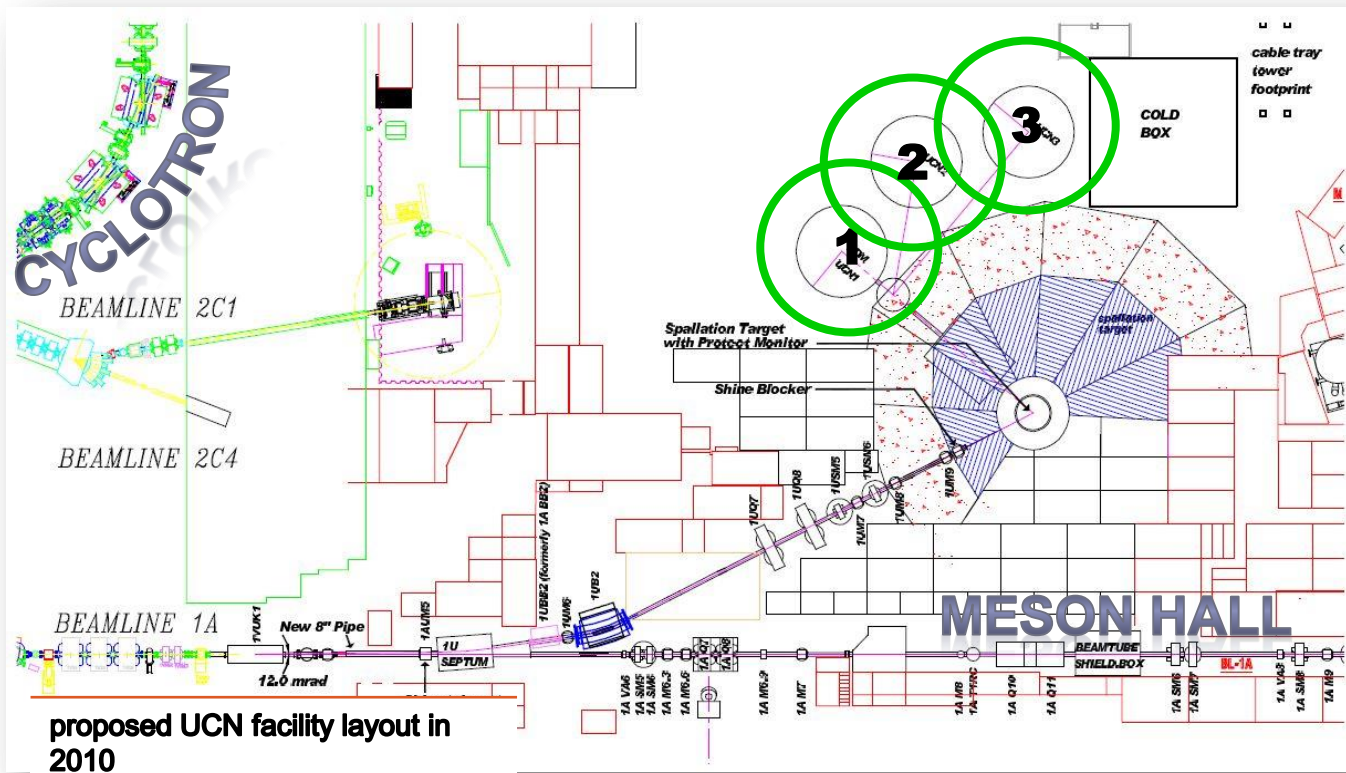


# Monte Carlo Simulation: Reproducing RCNP experiment

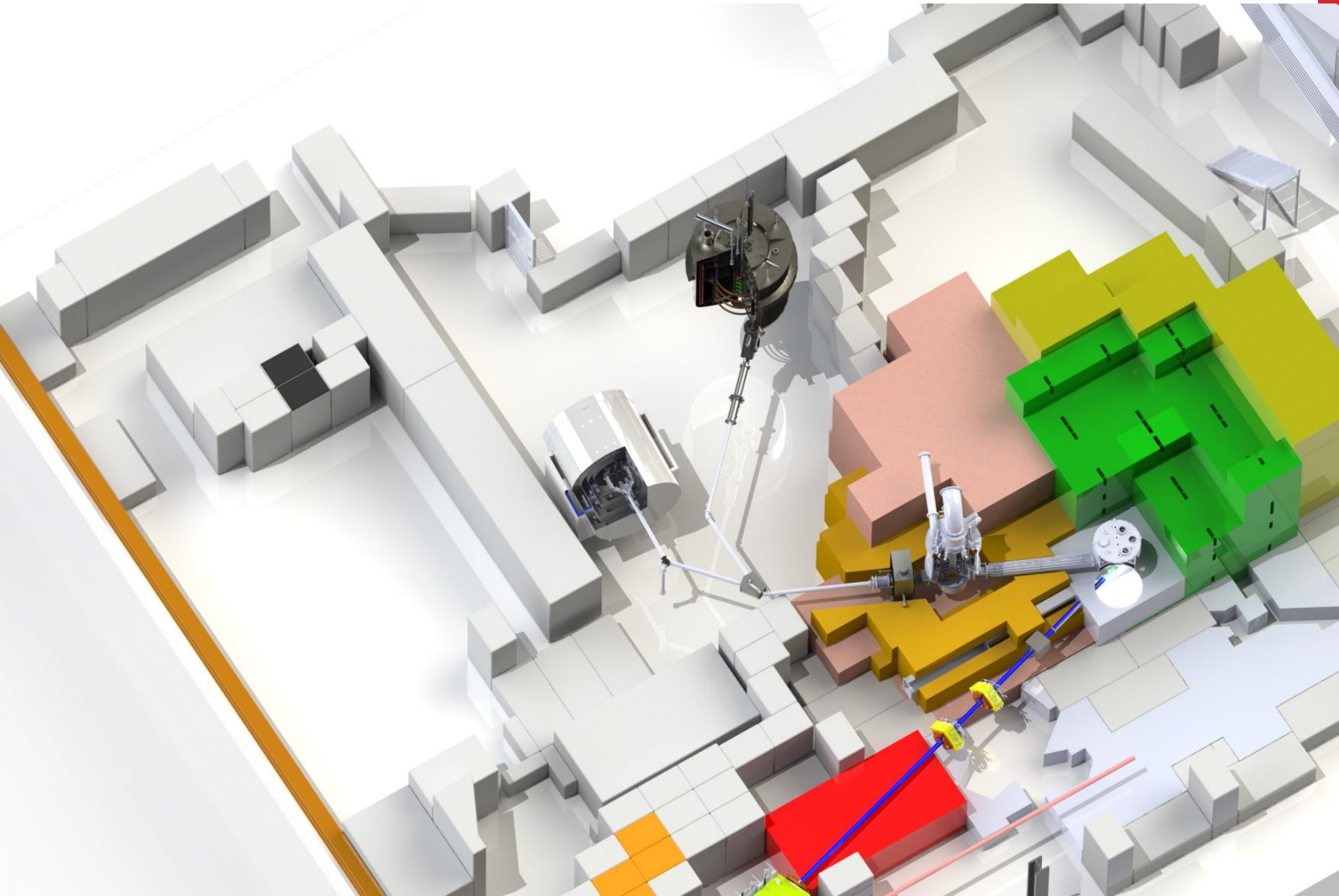


- detailed modelling of RCNP source run Nov 2013 to understand results
- geometry is meshed and used in PENTrack, our MC code

- **second UCN experiment port very valuable**
  - short term: for beam development, detector and guide tests
  - long term: for experiments besides EDM: lifetime, neutron decay, charge, gravity
- included in our upcoming CFI request
- big step towards a real user facility
- will attract UCN physicists from around the world

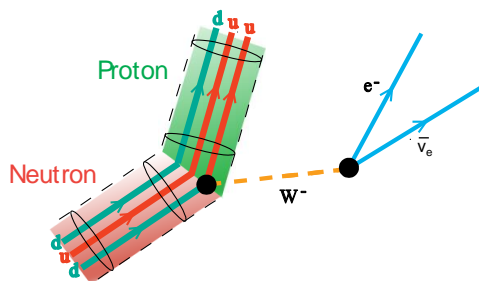






# PHYSICS WITH SLOW NEUTRONS

## $\beta$ -decay

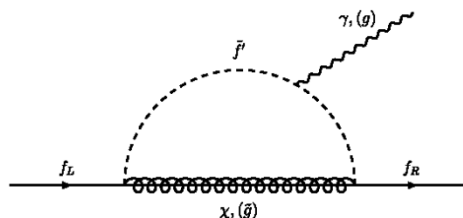


Charged Current parameters

BSM (S, V, A, T)  
 $\geq 7$  TeV

**best limits/values are possible**

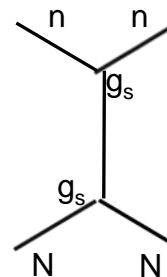
## Charge and Moments



nEDM: **best constraints**  
 CP, T violation

Charge: GUT-level effects

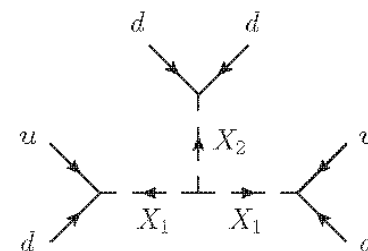
## Short range Forces



**best limits**  
 $\sim$ nm scale

Grav. states:  
 $\sim$  $\mu$ m scale

## $n \rightarrow \bar{n}$ Oscillation



Probes  $\Delta B=2$   
 $>100$  TeV  
 M.I. scale

**Best limits are possible**