

IOP HEPP Conference 2016, Sussex

Neutrinoless double beta decay at SNO+

Building on the success of SNO





LIP Coimbra
LIP Lisbon



Oxford University
Queen Mary,
University of London
University of
Liverpool
University of Sussex
University of
Lancaster



SNOLAB
TRIUMF
University of Alberta
Queens University
Laurentian University



Armstrong State University
Brookhaven National Lab
University of California, Berkley
University of California, Davis
Lawrence Berkeley National
Laboratory
University of Chicago
University of Pennsylvania
University of Washington



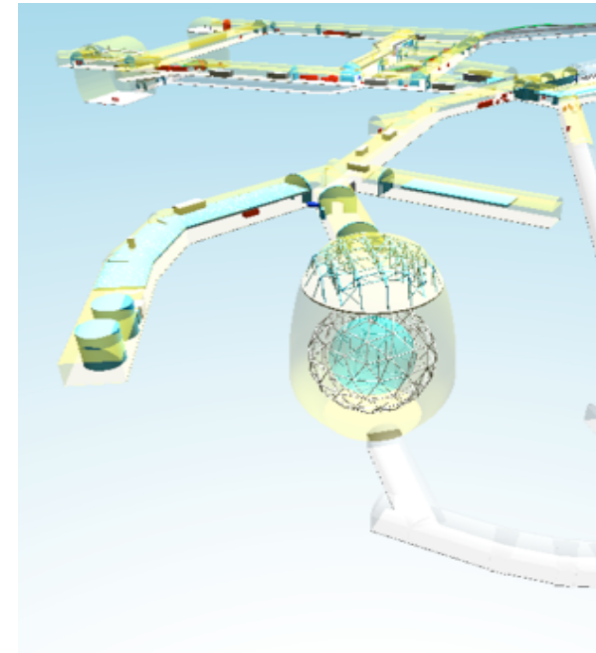
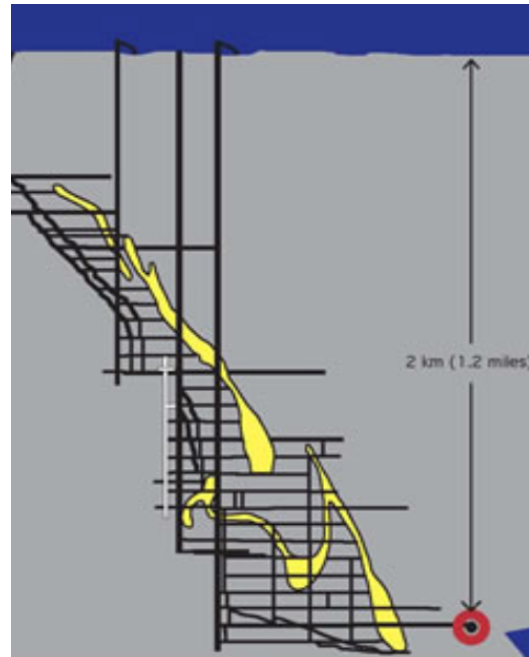
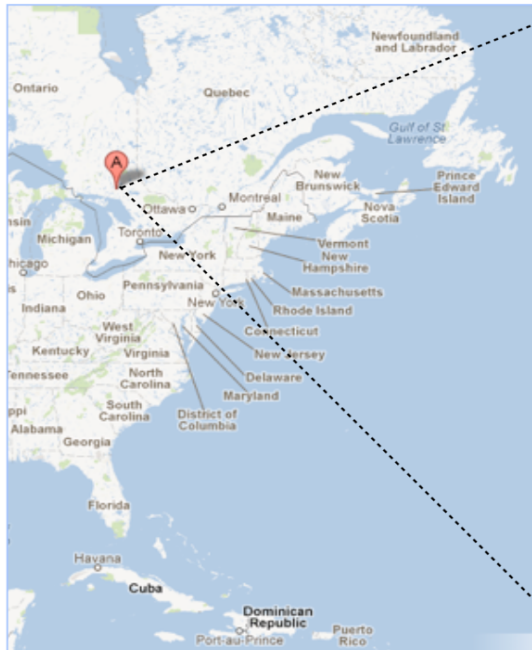
TU Dresden



UNAM



SNOLAB and SNO+



- Depth = 2070 m (6000 m w.e.)
- ~70 muons / day in SNO+
- 10,000 sq ft Class-2000 clean room

Located a SNOLAB inside the Creighton mine near Sudbury, Canada, SNO+ is the successor to Sudbury Neutrino Observatory (SNO)



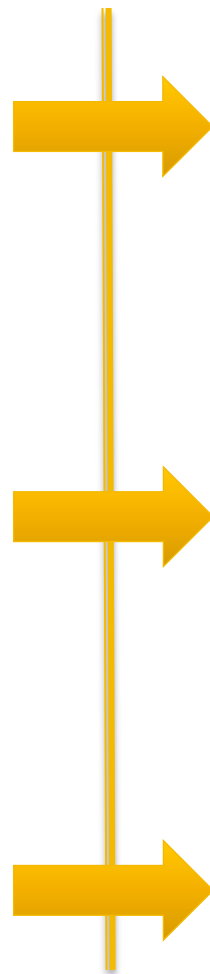
From SNO to SNO+

SNO

- Investigate the solar neutrino problem.
- Why do we observe $\sim 1/3$ less ν_e than predicted?

Answer:

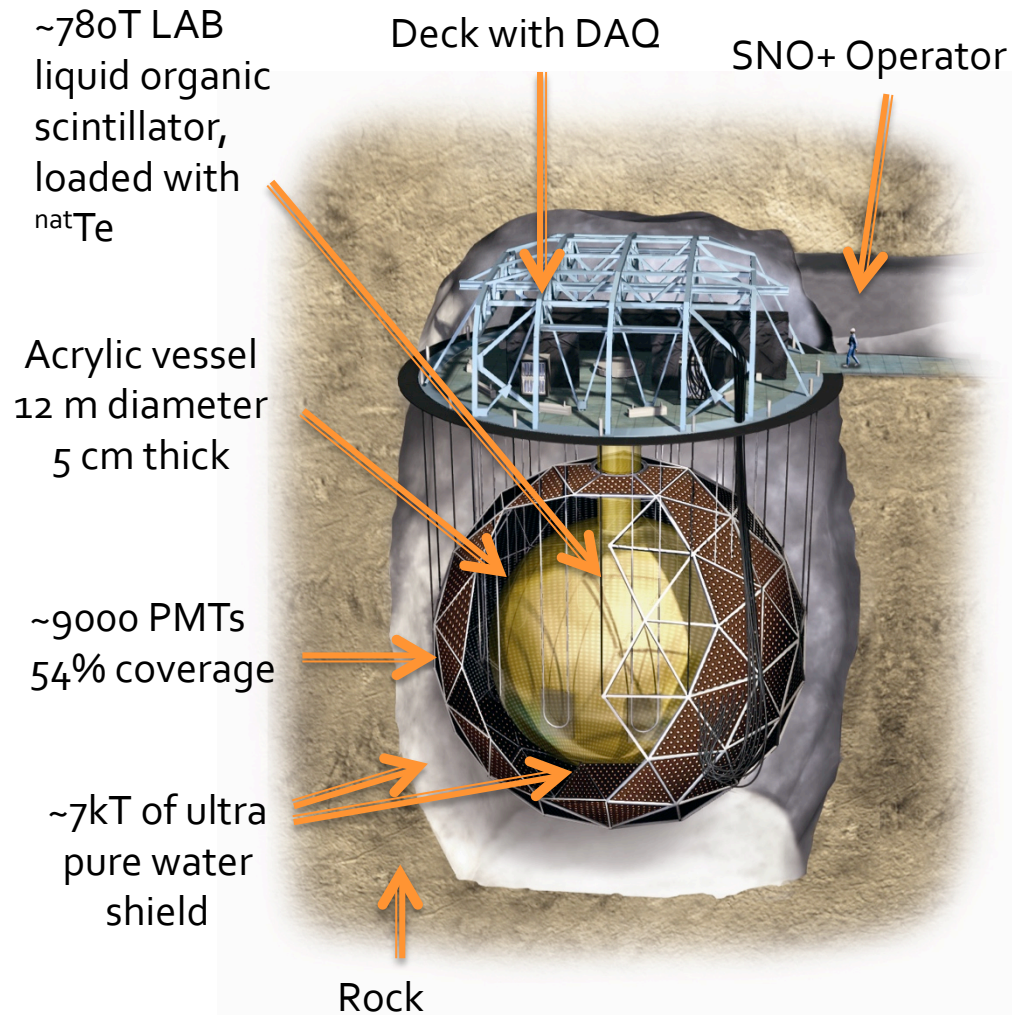
- Neutrinos change flavour!
(DOI: [10.1103/PhysRevC.88.025501](https://doi.org/10.1103/PhysRevC.88.025501))
- They must have mass....



SNO(+)

- Neutrinoless double beta decay.
Is the neutrino a Majorana particle?
(i.e. Are neutrinos and anti-neutrinos the same particle?)
- Further neutrino physics:
 - Solar (pep & CNO)
 - Geo
 - Reactor
 - Supernova
- Other physics
 - Nucleon decay
 - ALP search

SNO+ Detector



SNO+ upgrades the existing SNO detector to make a multi-purpose scintillator experiment.

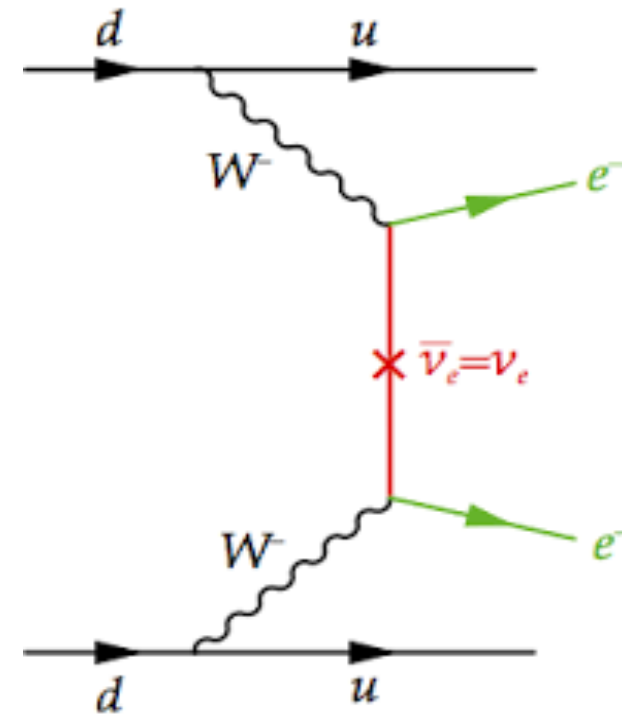
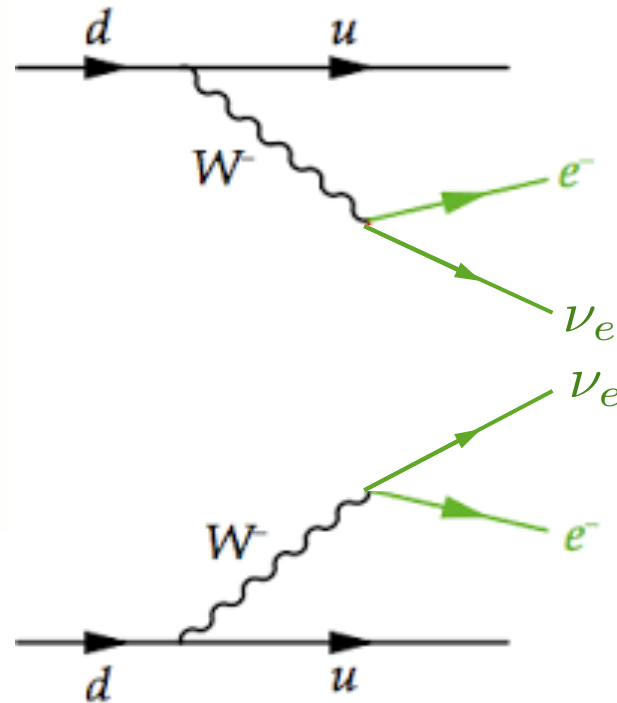
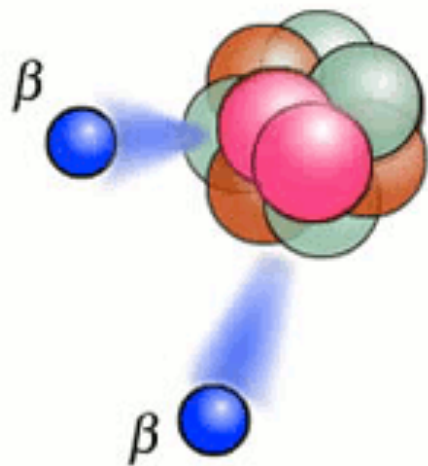
- Replace heavy water with Liquid scintillator (LAB)
- Upgrade trigger and electronics to handle order of magnitude increase in event rate
- Load scintillator mixture with ^{nat}Te ~34% of which is ^{130}Te , a known double beta decay isotope
 - Phase 1: 0.5% ^{nat}Te Loading
 - Phase 2: 3-5% ^{nat}Te loading (planned)



Neutrinoless double beta decay ($0\nu\beta\beta$)



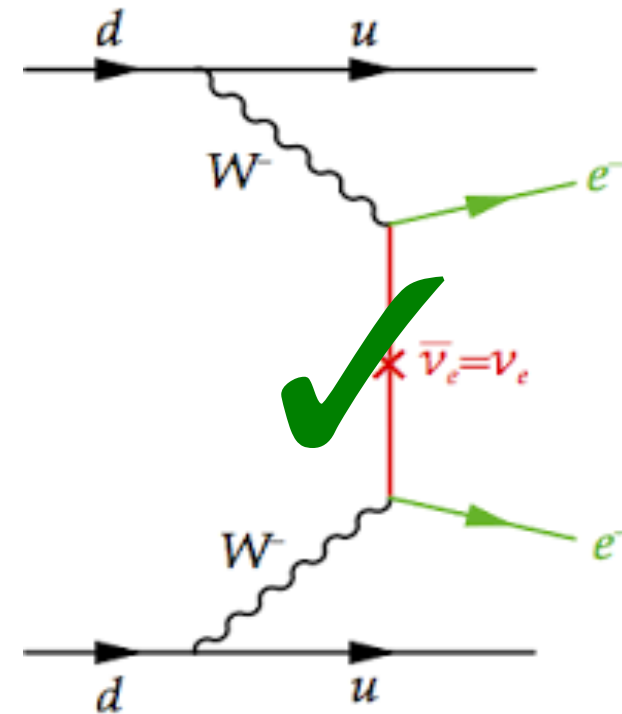
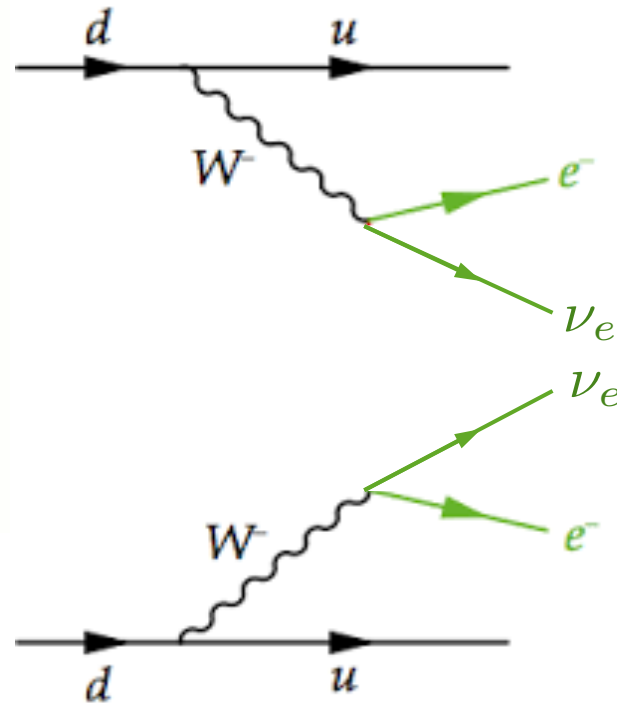
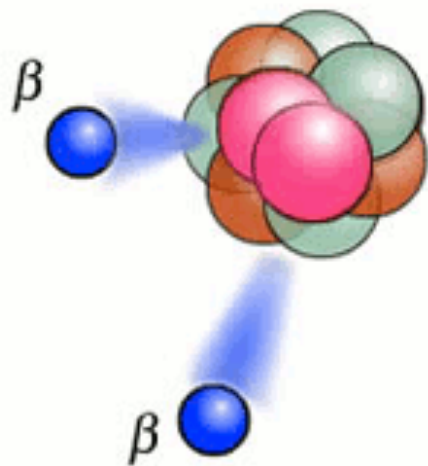
Neutrinoless double beta decay ($0\nu\beta\beta$)



Answers to the questions:

- Lepton number violation?
- Dirac or Majorana ?
- Normal or inverted hierarchy?

Neutrinoless double beta decay ($0\nu\beta\beta$)



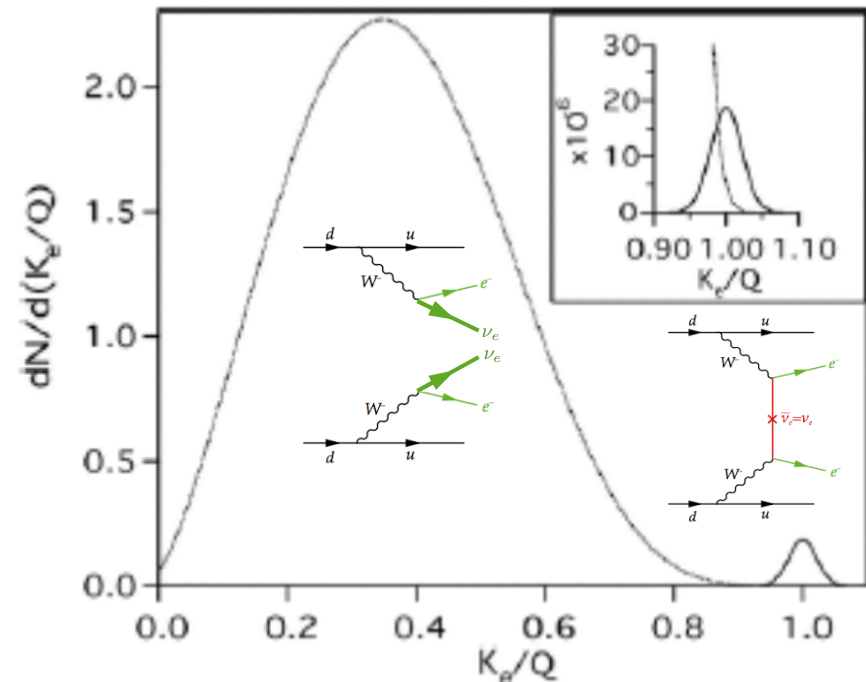
If we see it:

- Lepton number violation?
- Dirac or Majorana ?
- Normal or inverted hierarchy?

Neutrinoless double beta decay ($0\nu\beta\beta$)

- Allowed if $\nu = \bar{\nu}$, $m_\nu > 0$ for any mass eigenstate
- Explicitly violates lepton number conservation
- Rare: $\Gamma = G|M|^2|m_{\beta\beta}|^2, t_{1/2}^{0\nu} > 10^{21} y$
- Signature is a monoenergetic peak at $2\nu_{\text{BB}}$ Q-value
- ^{130}Te :
 - $Q_{\beta\beta} = 2527.518 \pm 0.013 \text{ keV}$

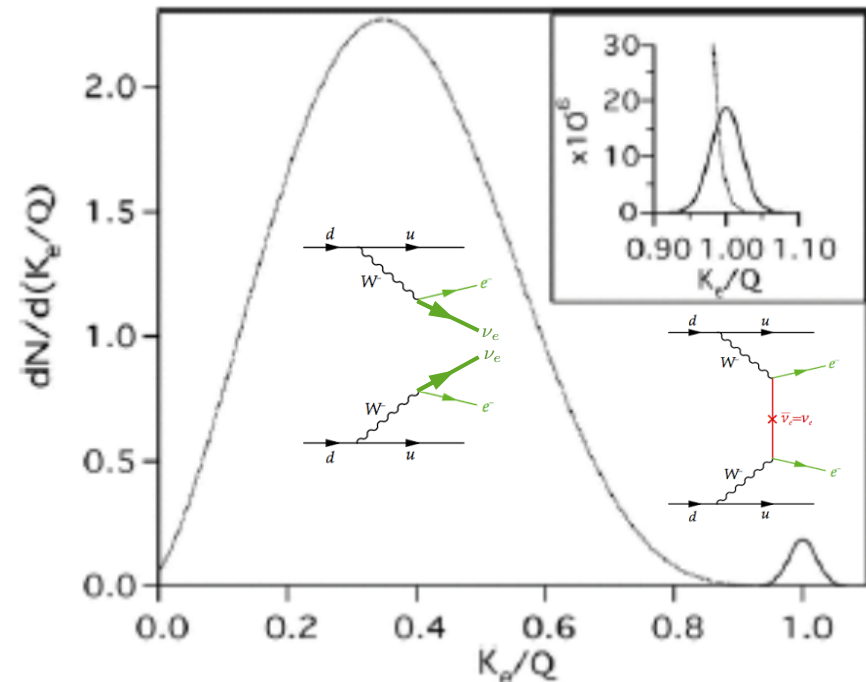
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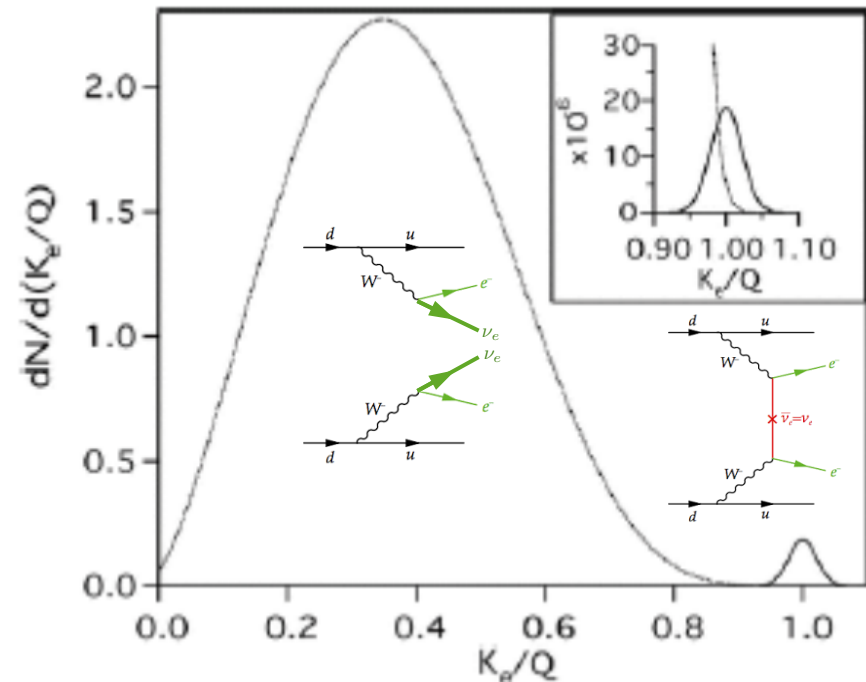
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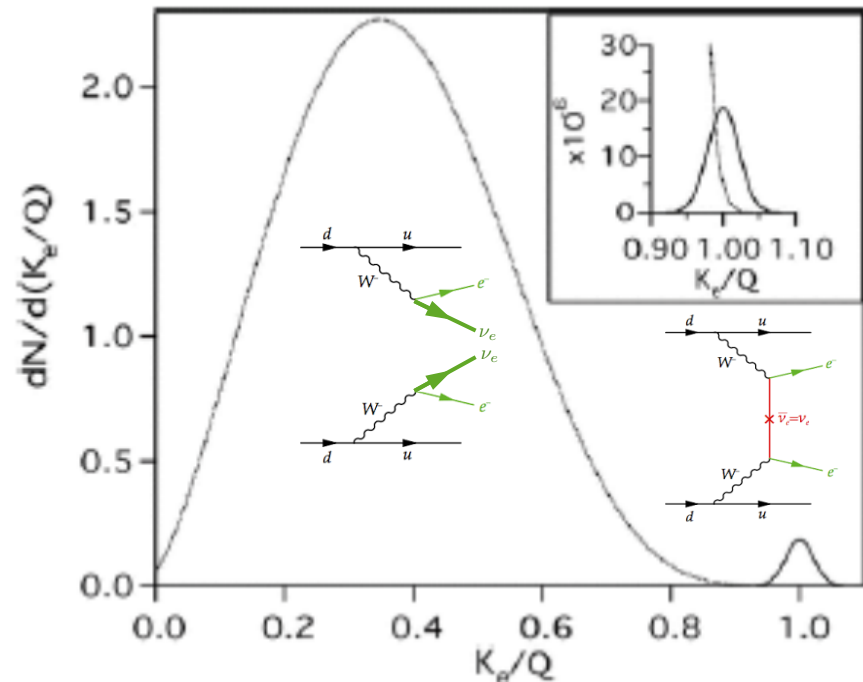
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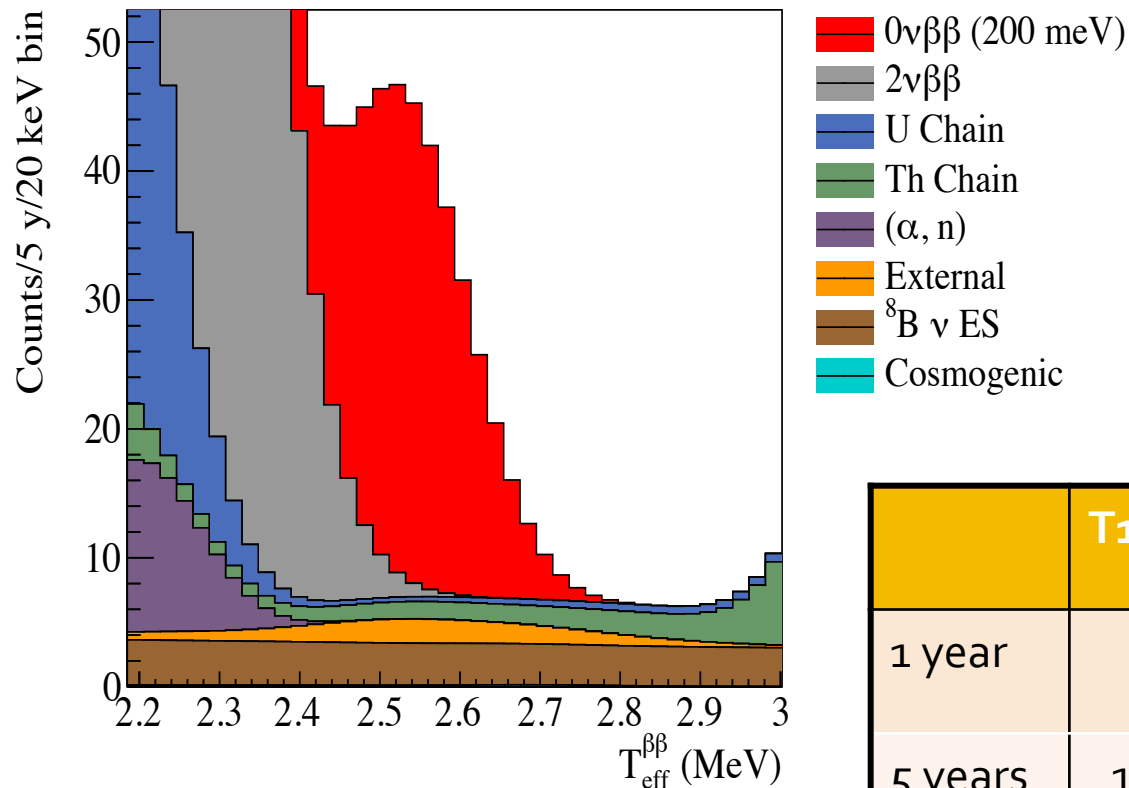
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SNO+ : ^{nat}Te loading 0.5 %

5 years



Assumptions:

- $M^{0\nu} = 4.03$ (IBM-2)
- $G^{0\nu} = 3.69 \times 10^{-14} \text{ y}^{-1}$
- $R < 3.5 \text{ m}$ (FV = 20%)
- $> 99.99\%$ (98%) rejection of ²¹⁴BiPo (²¹²BiPo)
- Light yield 390 NHits/MeV

	$T_{1/2}$ [yr]	$m_{0\nu\beta\beta}$ [meV]
1 year	8×10^{25}	75.2
5 years	1.96×10^{26}	38 - 92

SNO+ : ^{nat}Te loading 0.5 %

Two neutrino mode $2\nu\beta\beta$:

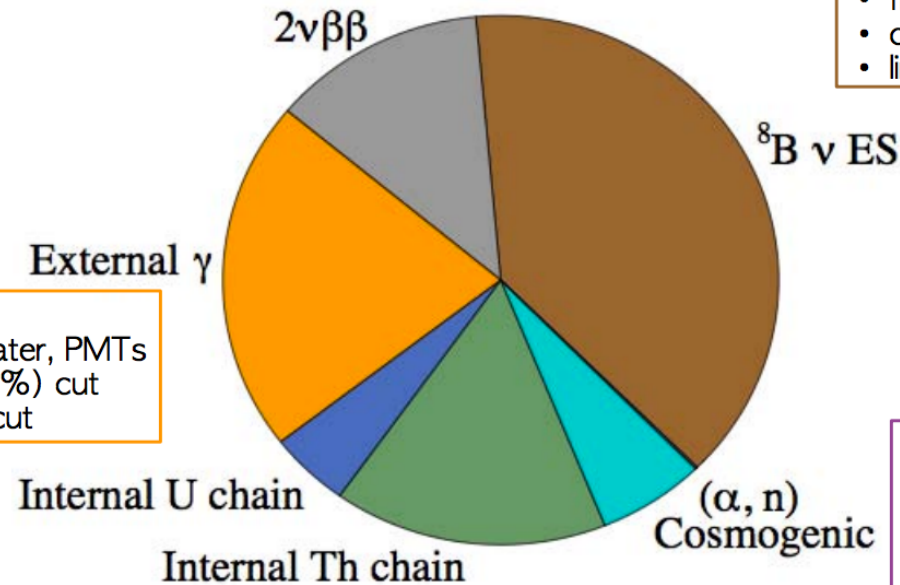
- asymmetric ROI around the $0\nu\beta\beta$ signal
- limited by energy resolution

^8B solar neutrinos:

- flat spectrum
- constrained by SNO/SK data
- limited by resolution

External gammas:

- from AV, ropes, water, PMTs
- fiducial volume (20%) cut
- 50% rejection LH cut



(a, n) :

- alpha-capture on $^{13}\text{C}/^{18}\text{O}$
- production of neutrons
- thermal neutron capture
- delayed coincidence tagging

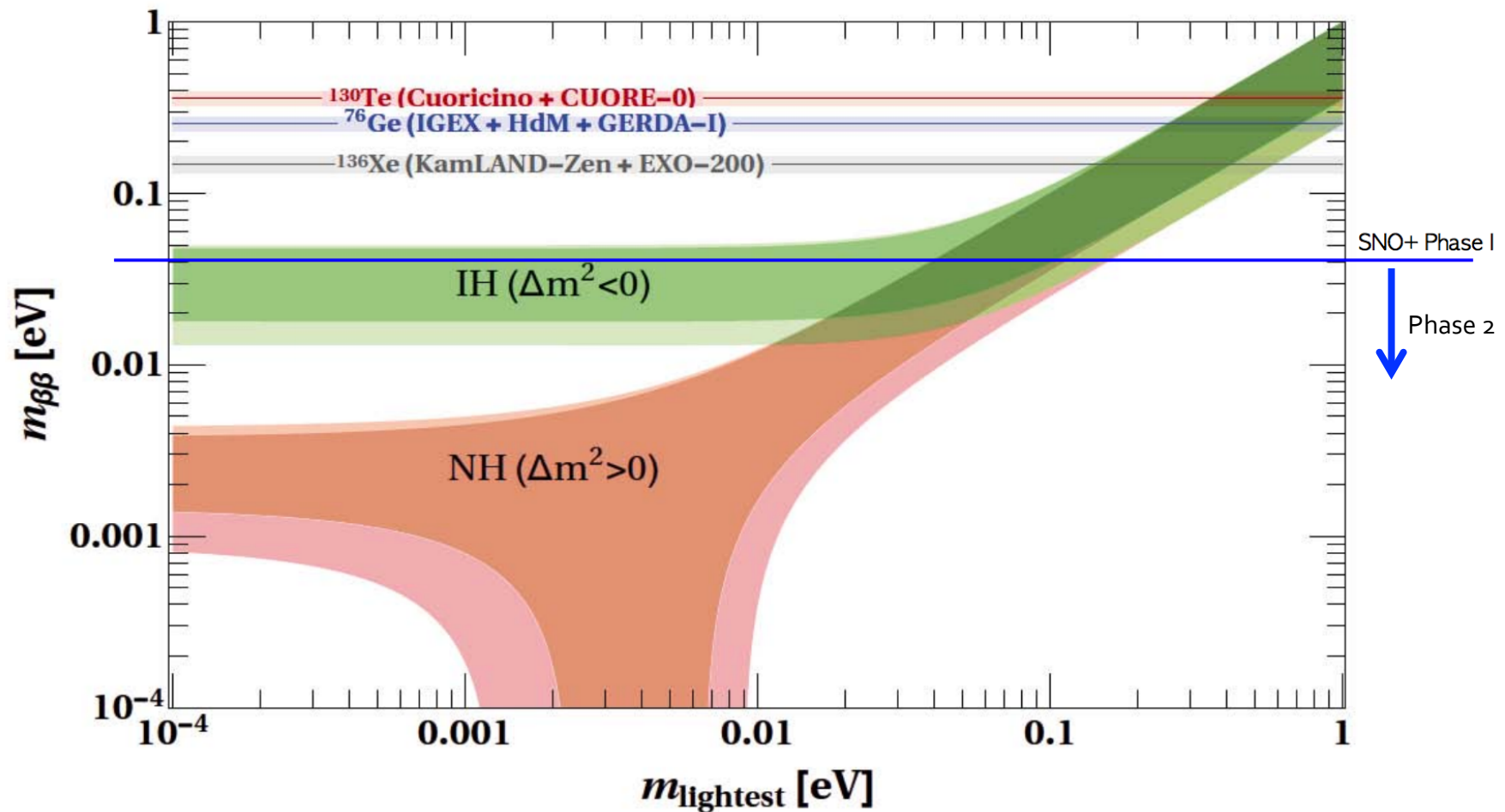
Internal U/Th chain:

- $^{214}\text{BiPo}$, $^{212}\text{BiPo}$
- $\beta - \alpha$ delayed coincidence tagging
- 100% rejection in ROI
- In-window trigger: x50 rejection

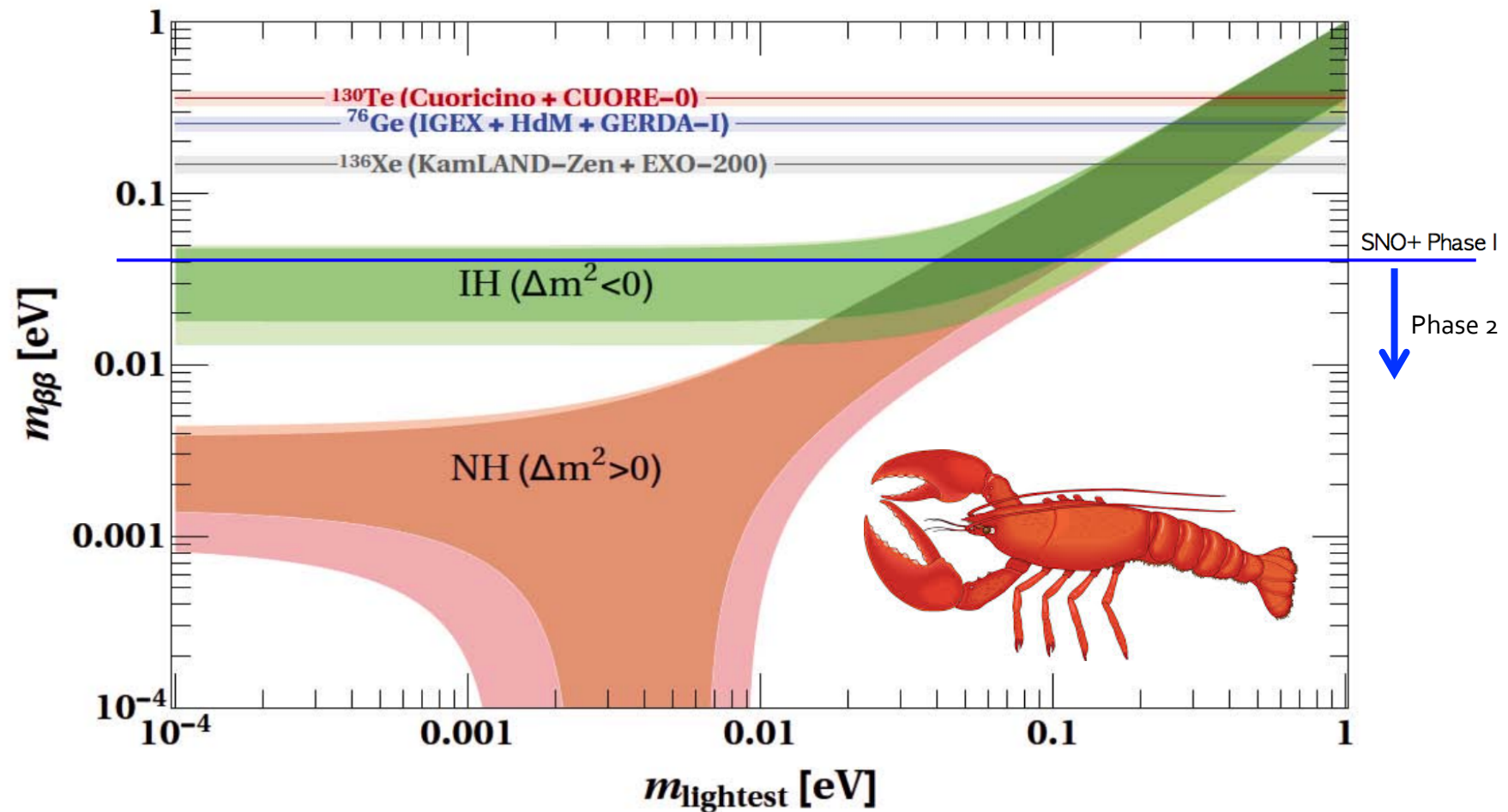
Cosmogenic:

- ^{60}Co , $^{110\text{m}}\text{Ag}$, ^{88}Y , ^{22}Na
- mitigation: purification + "cool-down" UG
- aimed: less than 1 ev/yr in ROI/FV

SNO+: ^{nat}Te loading 0.5 %



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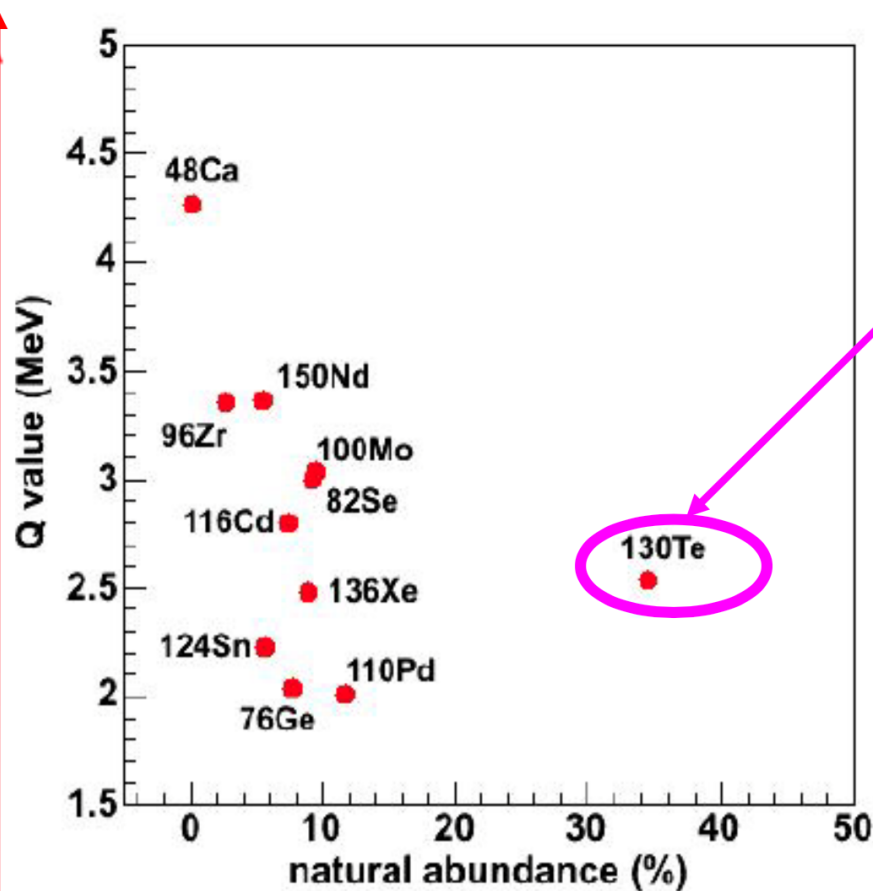
Conclusions & outlook

- SNO+ is a large liquid scintillator experiment, re-using the SNO detector.
- Physics priority: $\nu\beta\beta$ search with ^{130}Te
- A 'Phase 1' loading of 0.5 % $^{\text{nat}}\text{Te}$ will begin to probe into the inverted hierarchy, competitive with CUORE and KamLAND-Zen
- Inverted hierarchy largely explored in the planned phase 2
- Water fill data to start June 2016
- Te-loaded phase to start in 2017



Backup slides...

$\nu\beta\beta$ isotopes



Decay candidate	Q value (MeV)	natural abundance (%)
$^{48}\text{Ca} \rightarrow ^{48}\text{Ti}$	4.271	0.187
$^{76}\text{Ge} \rightarrow ^{76}\text{Se}$	2.040	7.8
$^{82}\text{Se} \rightarrow ^{82}\text{Kr}$	2.995	9.2
$^{96}\text{Zr} \rightarrow ^{96}\text{Mo}$	3.350	2.8
$^{100}\text{Mo} \rightarrow ^{100}\text{Ru}$	3.034	9.6
$^{110}\text{Pd} \rightarrow ^{110}\text{Cd}$	2.013	11.8
$^{116}\text{Cd} \rightarrow ^{116}\text{Sn}$	2.802	7.5
$^{124}\text{Sn} \rightarrow ^{124}\text{Te}$	2.228	5.64
$^{130}\text{Te} \rightarrow ^{130}\text{Xe}$	2.533	34.1
$^{136}\text{Xe} \rightarrow ^{136}\text{Ba}$	2.479	8.9
$^{150}\text{Nd} \rightarrow ^{150}\text{Sm}$	3.367	5.6

Heirarchy

