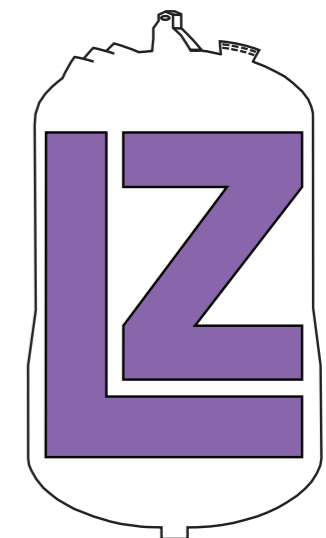


Direct Dark Matter Searches with the LUX and LZ Experiments

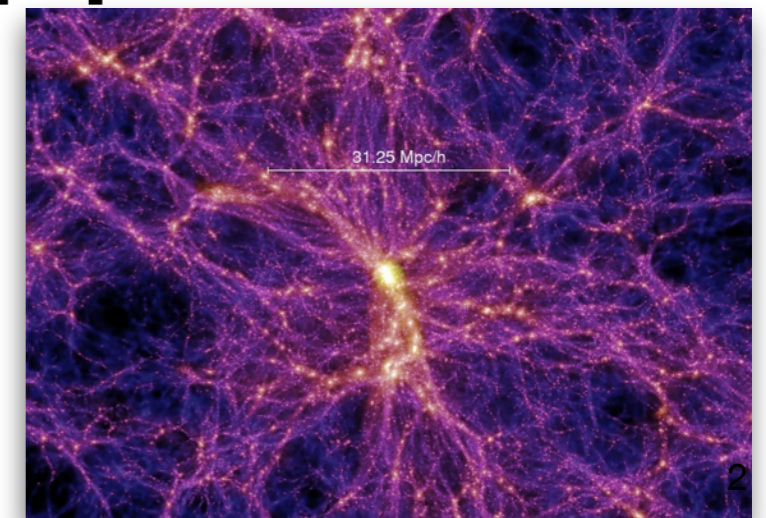
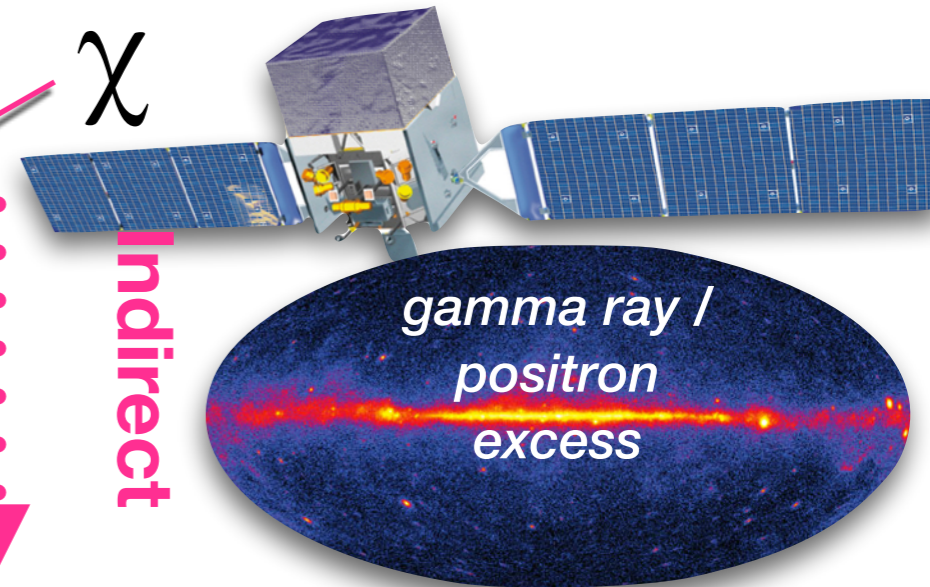
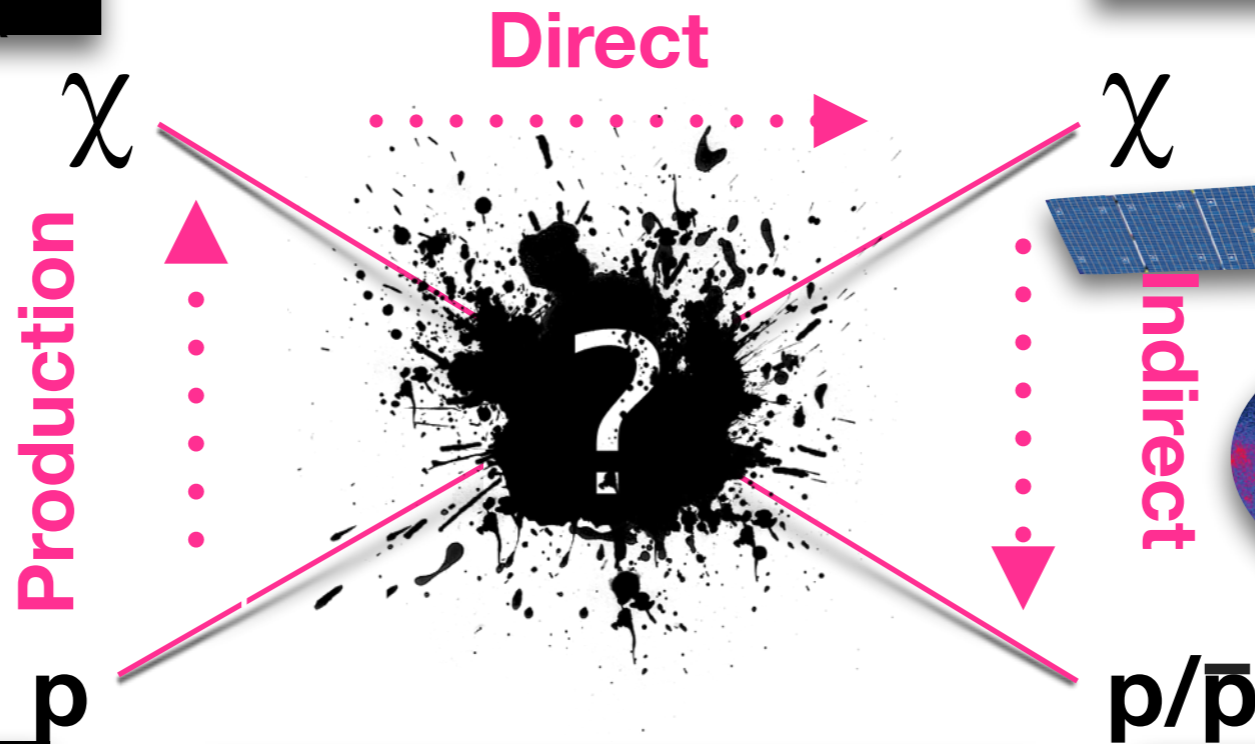
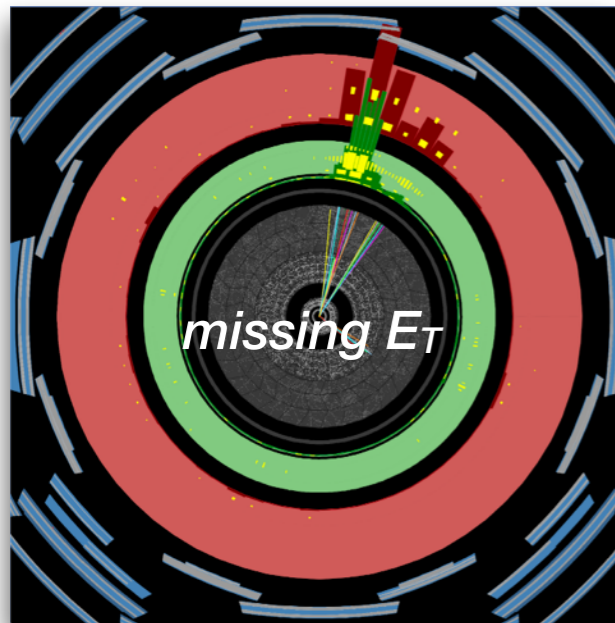
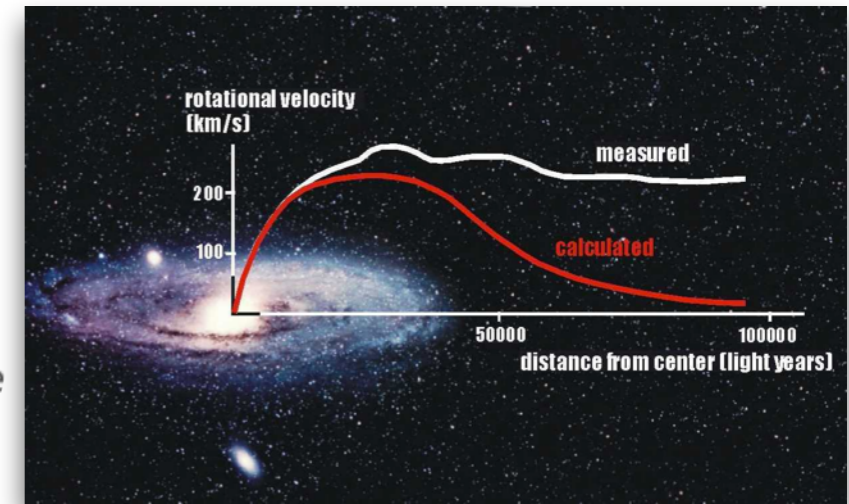
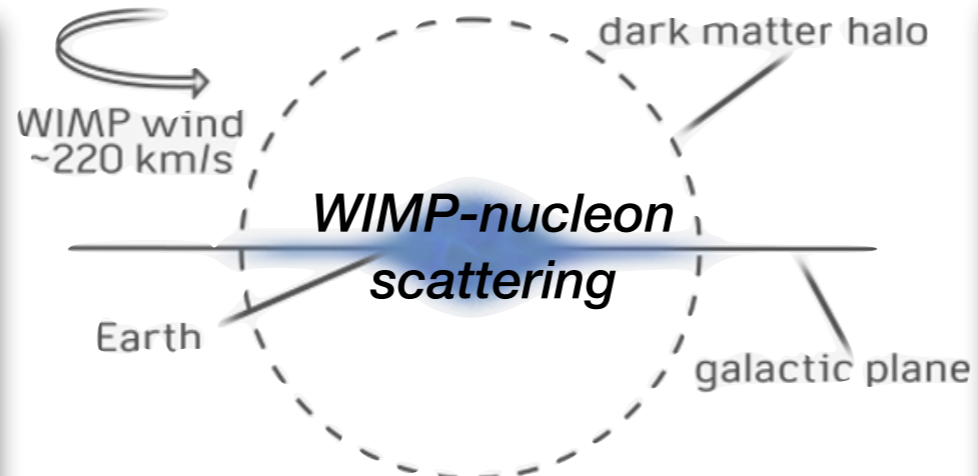
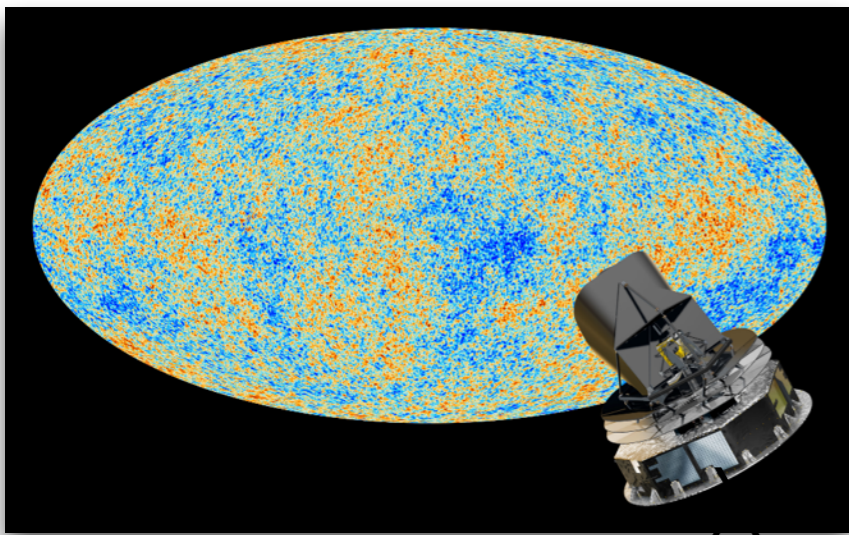
Sally Shaw

IoP HEP & APP Joint Meeting

22nd March 2016



Dark Matter Detection



- **Large Underground Xenon** detector
- Dual-phase Xenon TPC with 3D position reconstruction and 99.6% electron recoil (ER) nuclear recoil (NR) discrimination
- World's first sub-zeptobarn detector
- Status: WIMP search mode for Run 4, done by June 2016

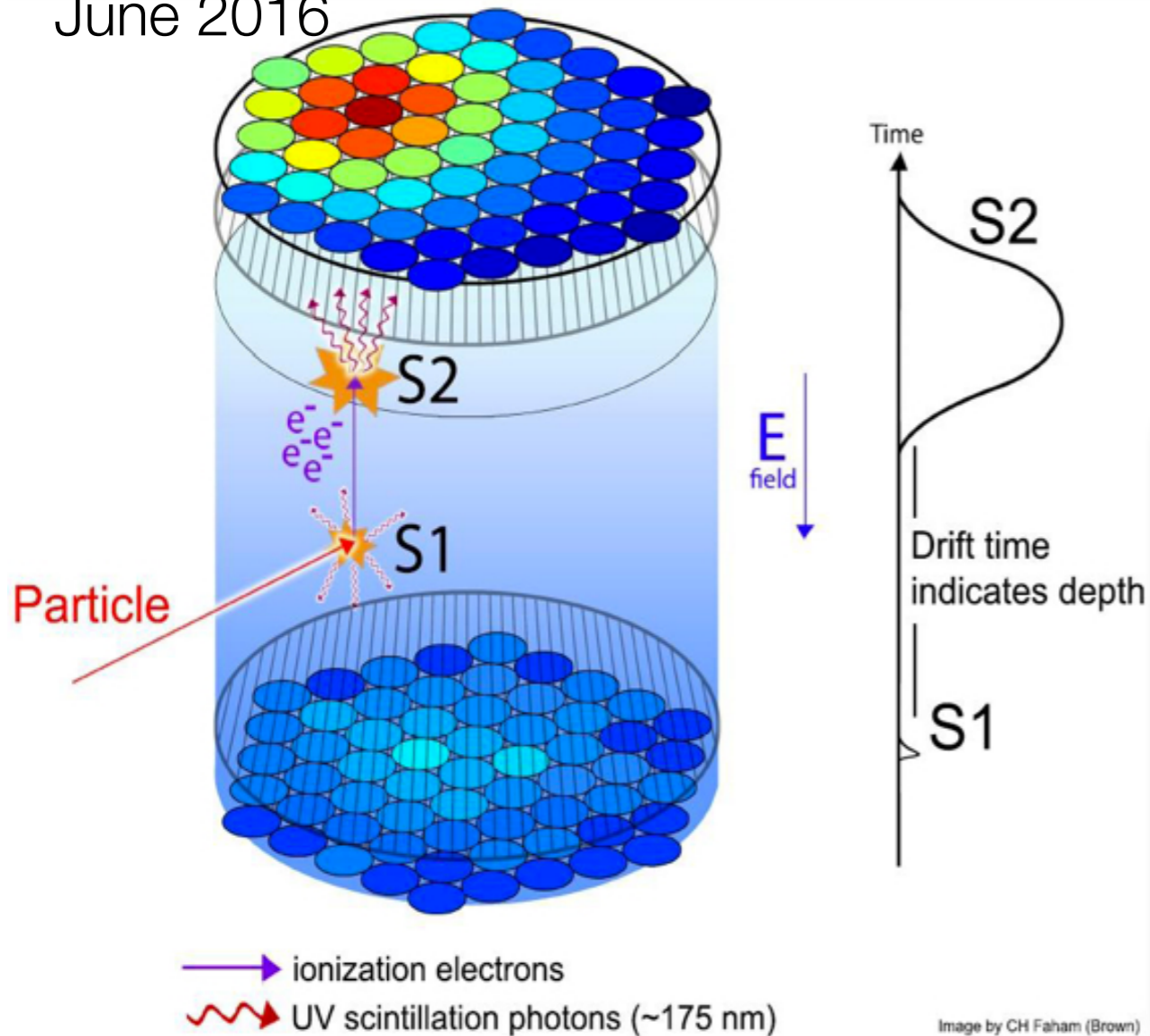
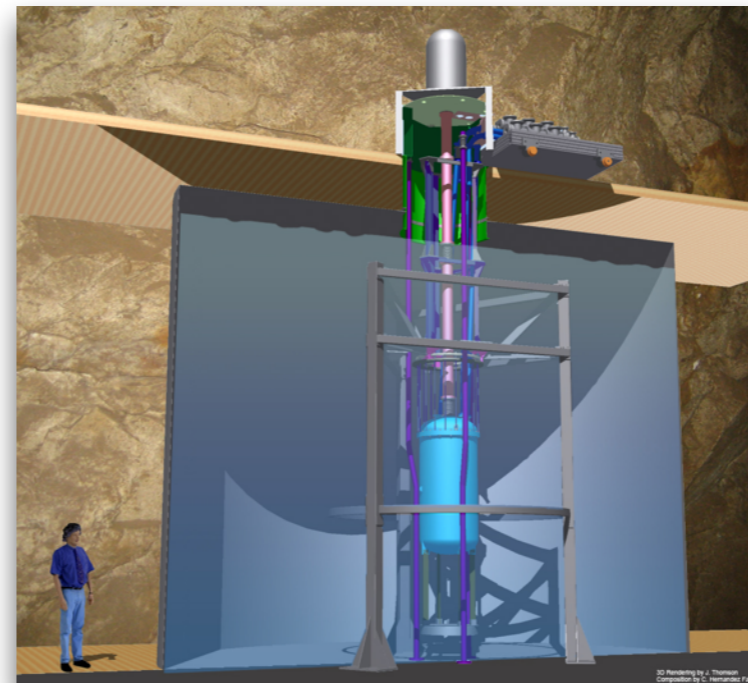
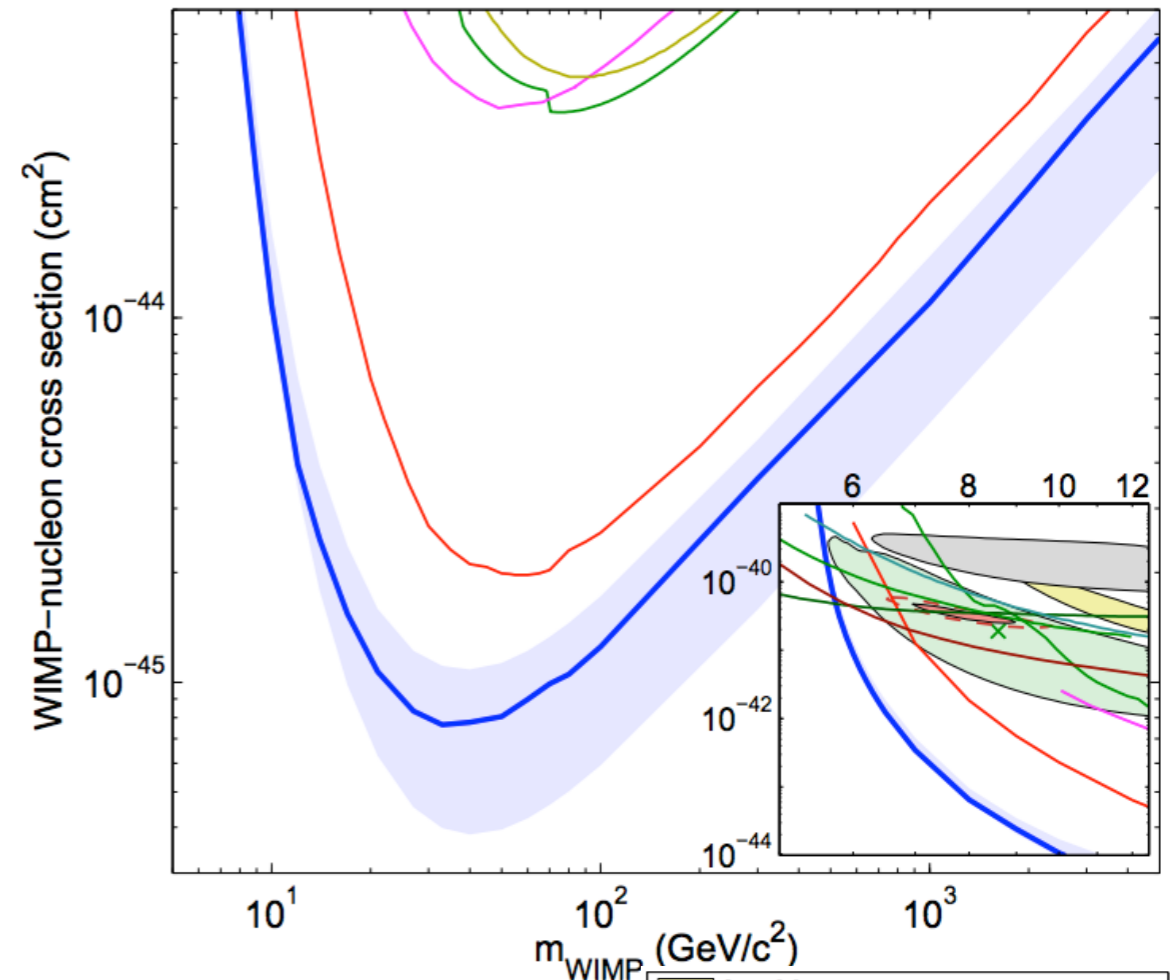
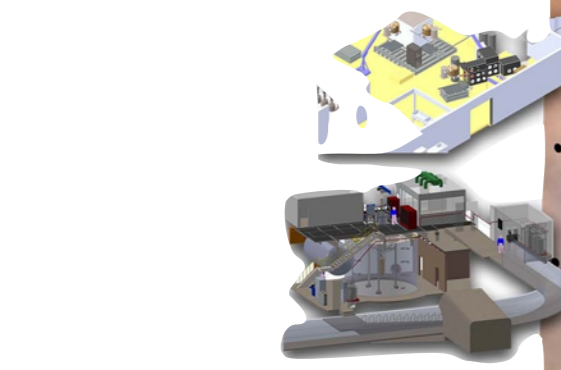


Image by CH Faham (Brown)



Sanford Lab, South Dakota



Davis Campus

- **MJD**
MAJORANA DEMONSTRATOR
Neutrinoless double-beta decay
- **LUX/LZ**
Large Underground Xenon experiment
First and second generation dark matter
- **CUBED**
Center for Ultra-Low Background
Experiments in the Dakotas
Low background counting

Yates Shaft

Ross Shaft

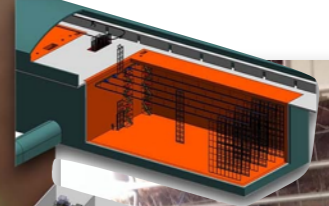
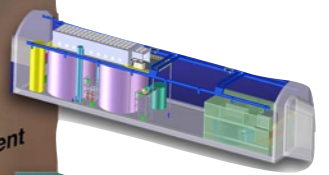
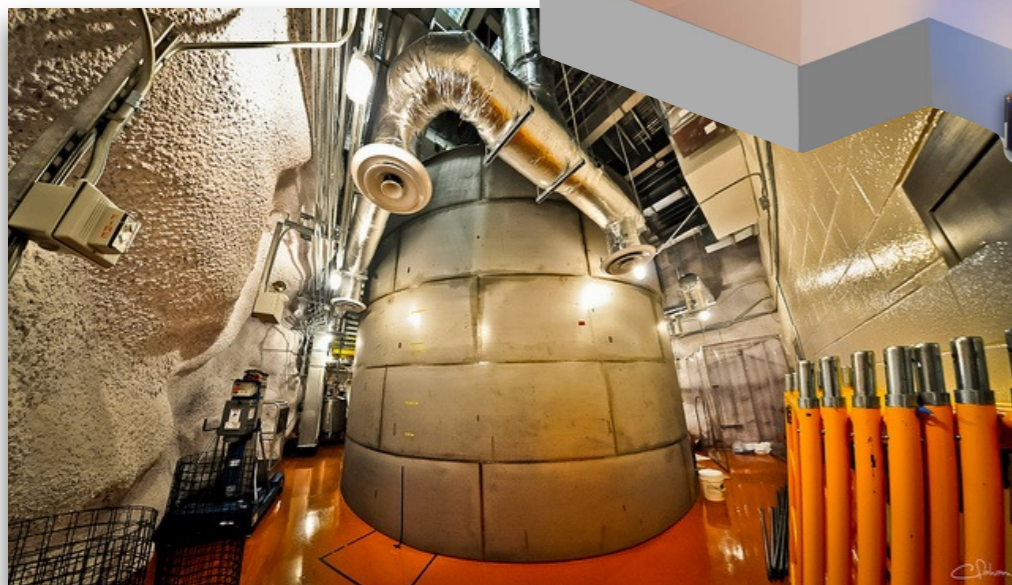
- **Experiment Hall**
Third generation dark matter experiment
1 T neutrinoless double-beta decay experiment

- **LBNE**
Long-Baseline Neutrino Experiment
4850 Level liquid argon
- **Low Background Counting**

Ross Campus

- **MJD**
MAJORANA DEMONSTRATOR
Electroforming laboratory

- **DIANA**
Dual Ion Accelerators for Nuclear Astrophysics
4850 Level DIANA Laboratory



- More accurate estimation of detected photons
- More livetime - further 10 days (1.4×10^4 kg · days)
- Improved background model allowing a larger fiducial volume (118kg \rightarrow 147kg)
- Better determination of g1 and g2 for energy reconstruction:

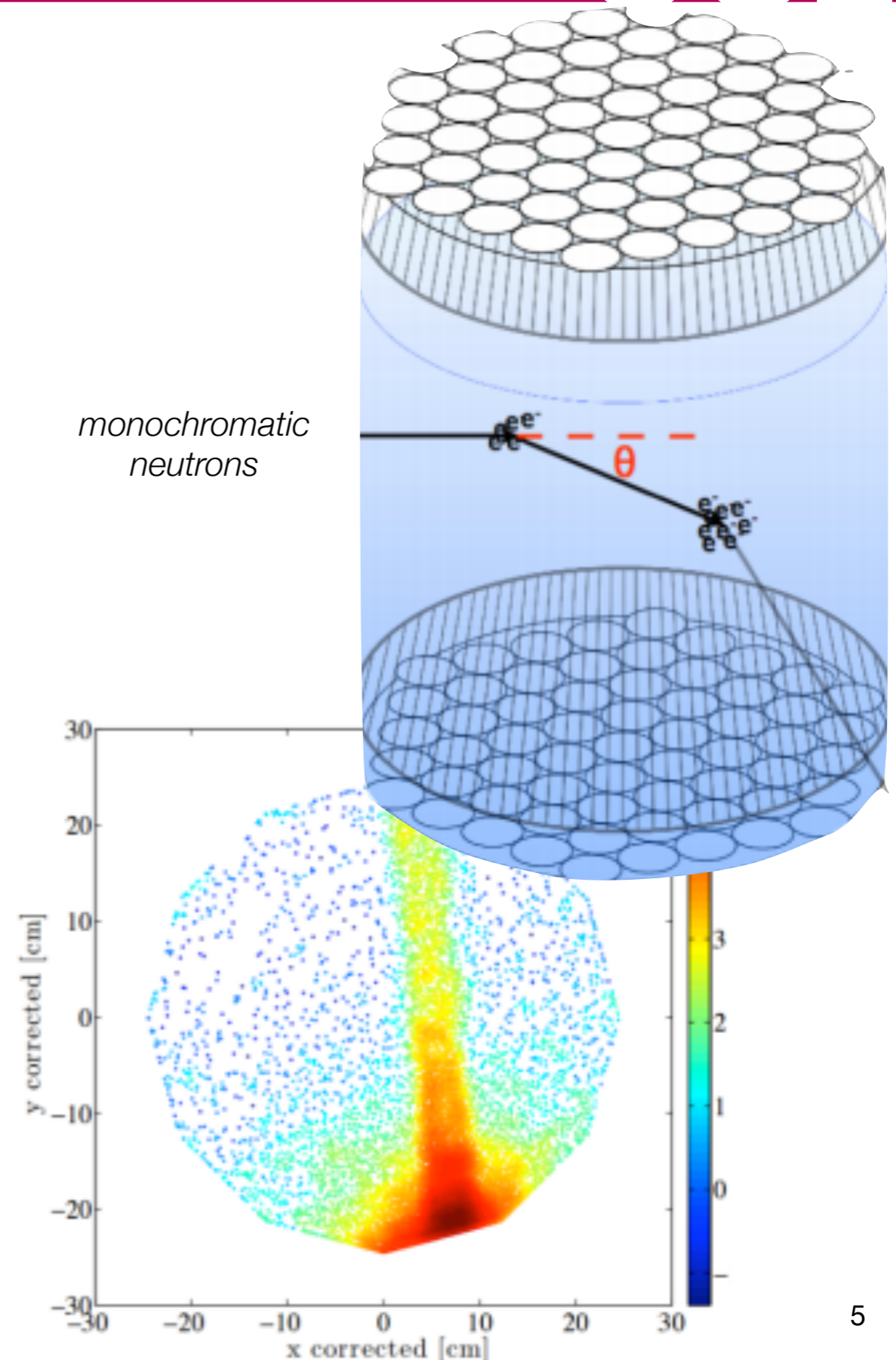
$$E = \frac{1}{\mathcal{L}(E)} W \left(n_{ph} + n_e \right)$$

Lindhard factor-energy loss to heat \rightarrow $\mathcal{L}(E)$ \leftarrow 13.7 keV_{ee}

$S1/g1$ \rightarrow n_{ph} \leftarrow $S2/g2$ \rightarrow n_e

- In situ NR calibration - light and charge yield measured, low energy threshold 3 keV \rightarrow 1.1 keV

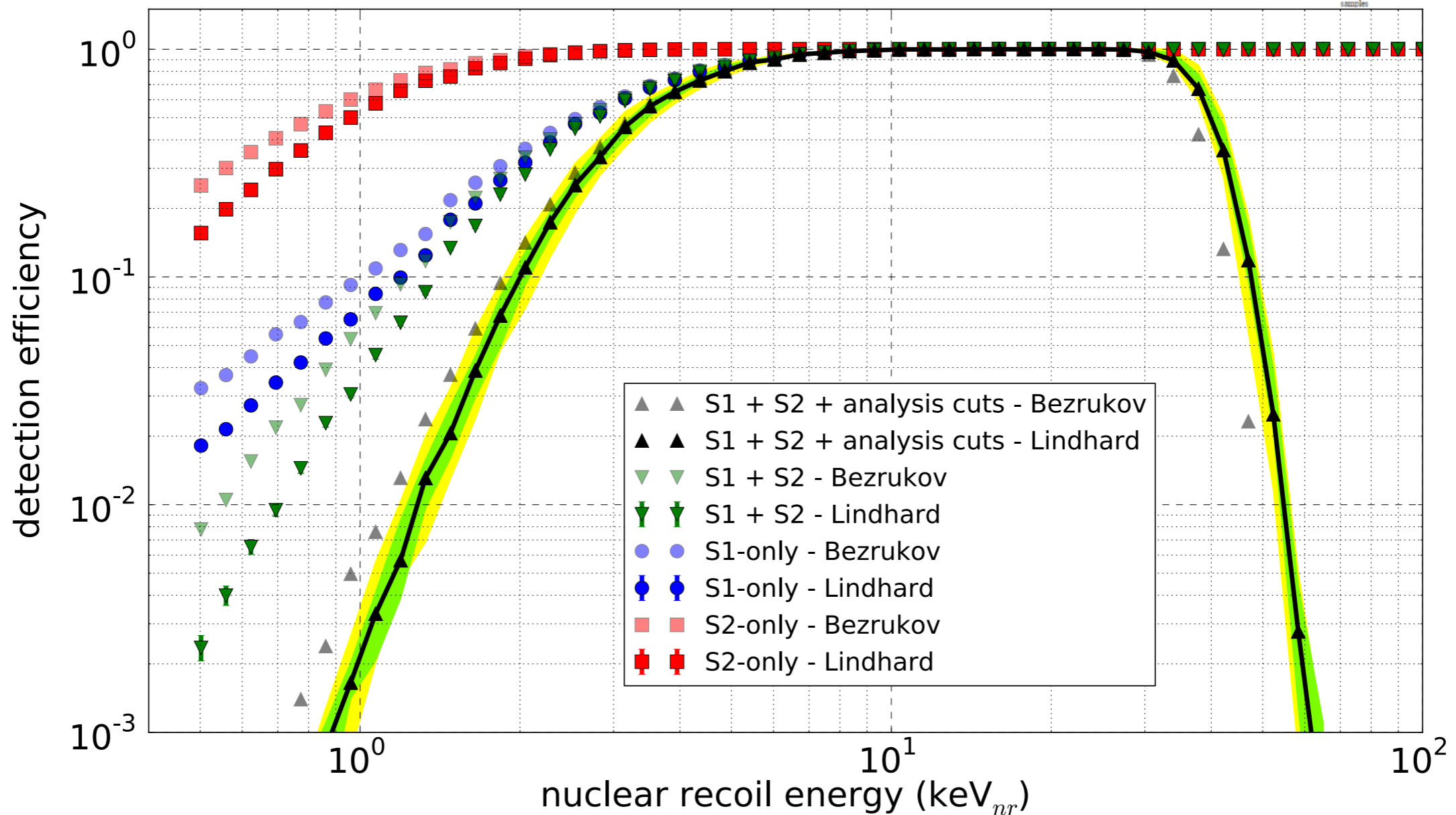
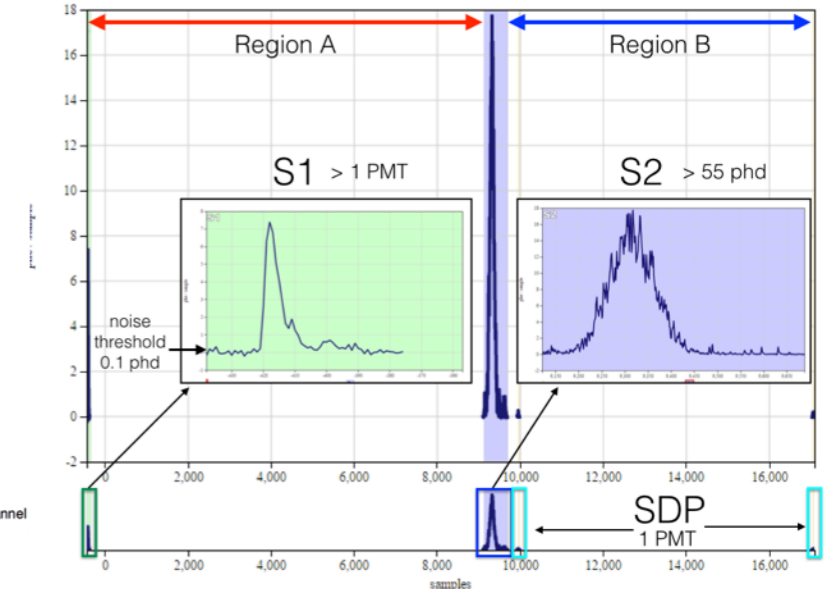
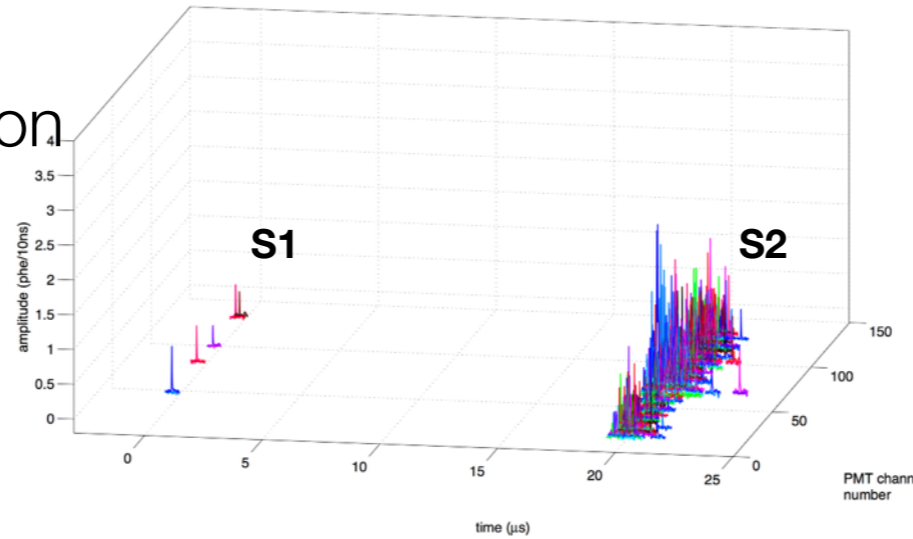
$$E_r = E_n \frac{4m_n m_{Xe}}{(m_n + m_{Xe})^2} \frac{1 - \cos\theta}{2}$$



Pulse Identification & Event Selection

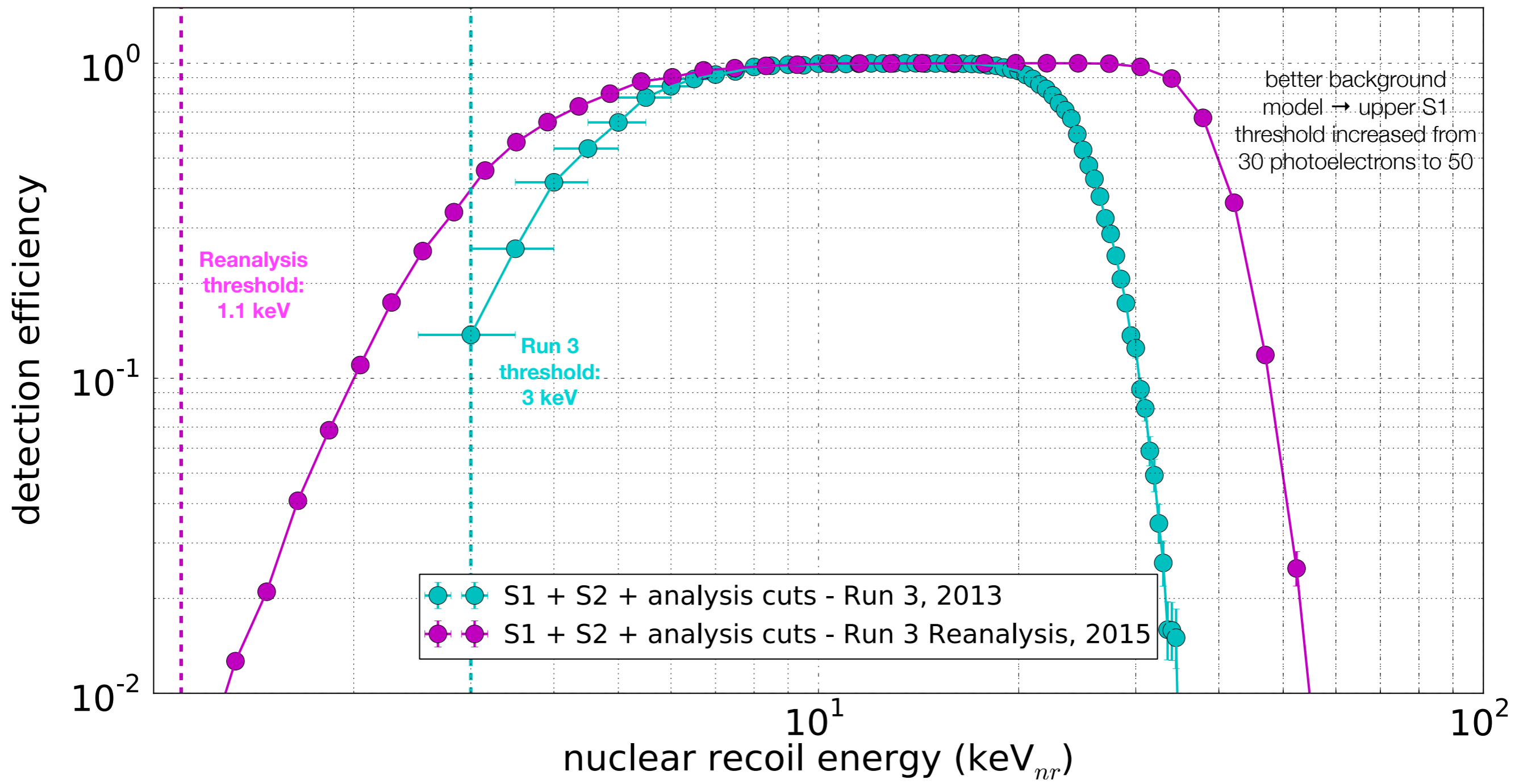
Contributions:

- Pulse finding and classification
- Event classification
- WIMP search selection cuts
- Nuclear recoil efficiencies



- **lowered S1 & S2 thresholds**
 - threshold: minimum size (in photoelectrons) for a given classification

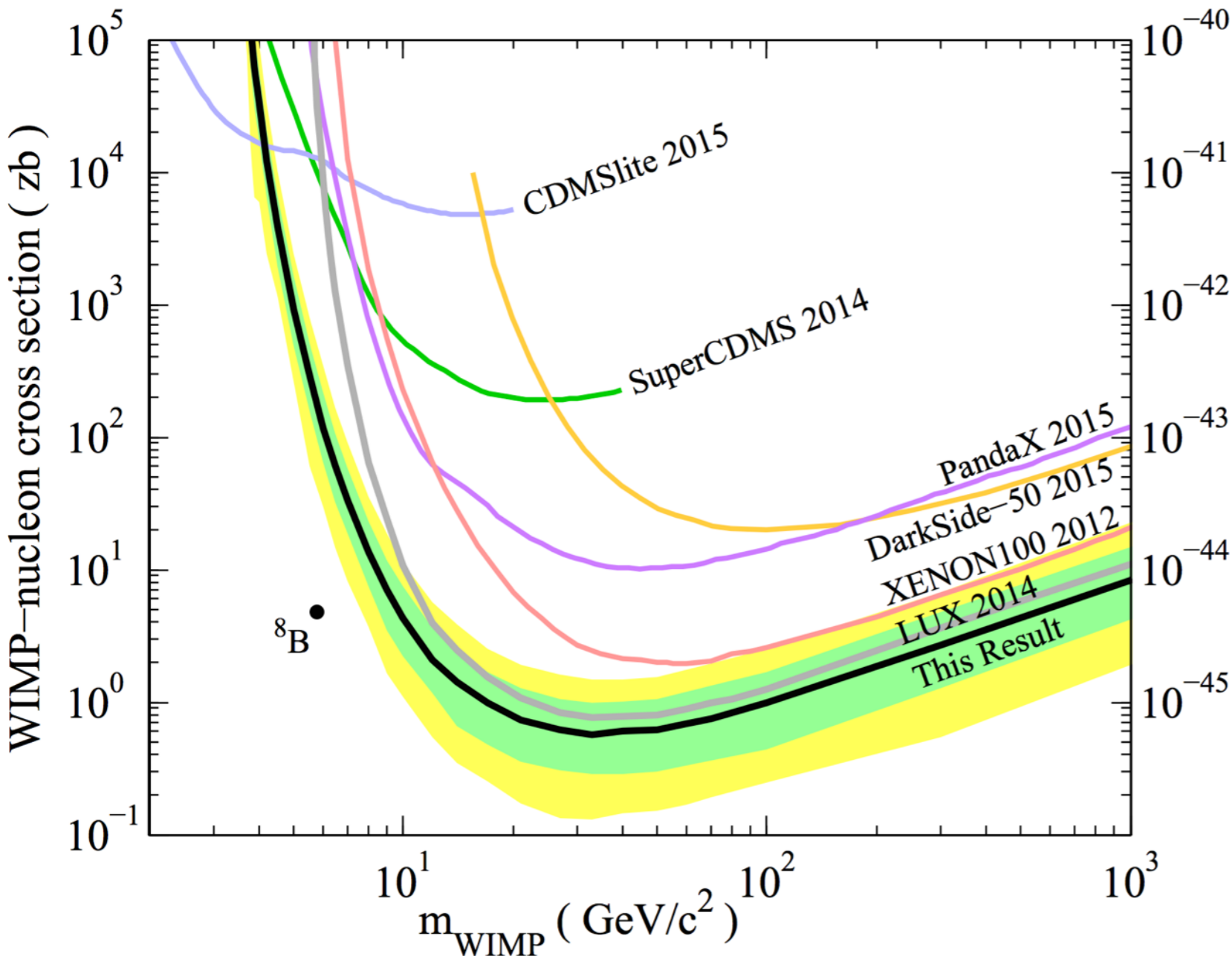
- **algorithm improvements for both pulse finder and classifier**
 - multiple scatter identification
 - bug fixes



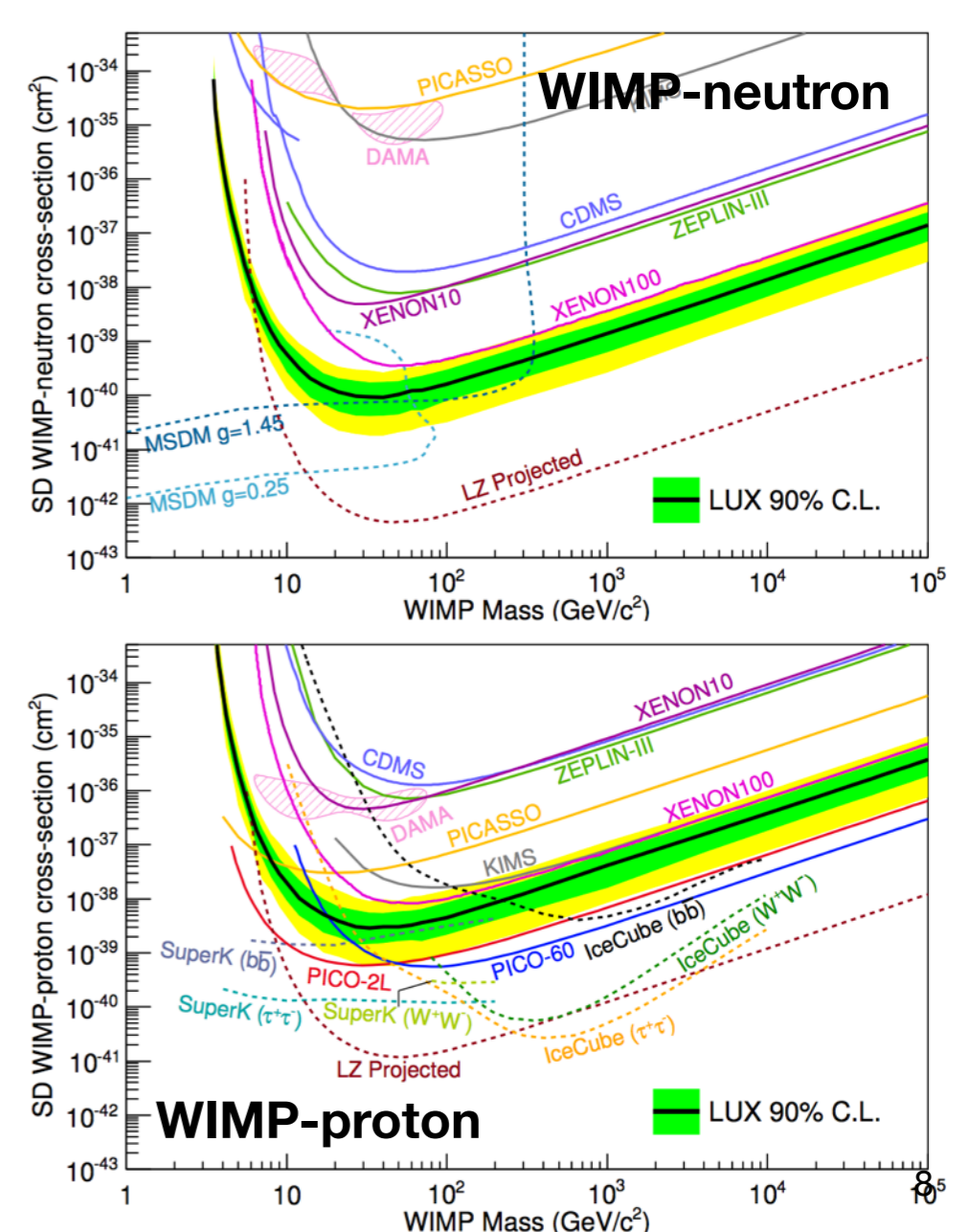
LUX New Results

- Main improvement at low WIMP mass (ie low recoil energy)
- 33 GeV WIMP: 90% CL upper limit minimum 0.6 zb

Better spin-independent result arXiv:1512.03506



First spin-dependent results arXiv:1602.03489

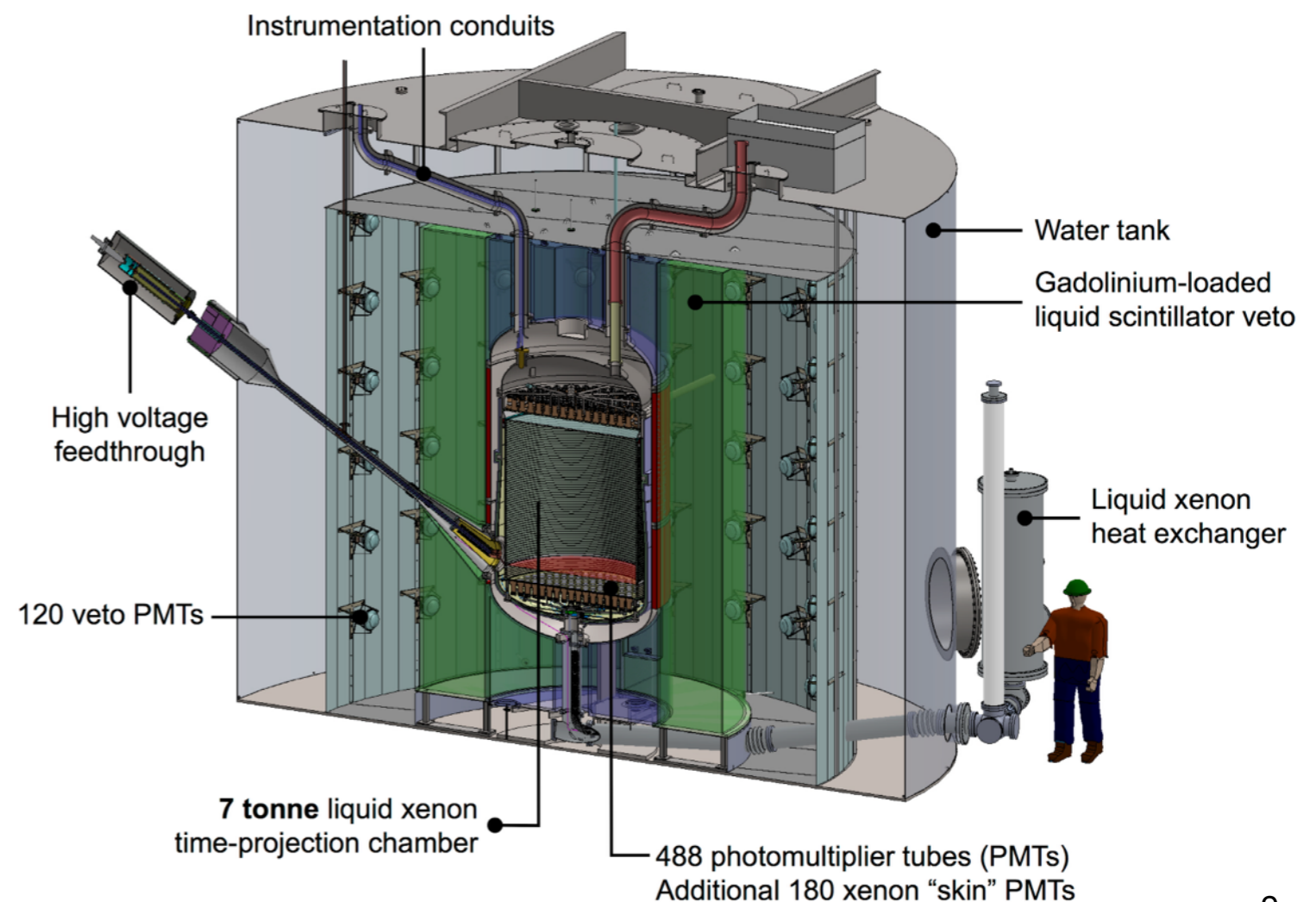


- Schedule: LUX removed late 2016, LZ commissioning 2019
- 10 tonnes of LXe - 7 tonne fiducial (c.f. LUX 118kg)
- 488 xenon PMTs (c.f. LUX 122)
- Outer detector system:
 - Gd-loaded LAB scintillator
 - 120 veto PMTs in water tank
 - Xenon skin veto
 - 180 skin PMTs
- Projected sensitivity: $2 \times 10^{-48} \text{ cm}^2$ for 50 GeV WIMP
- **High precision background model needed**

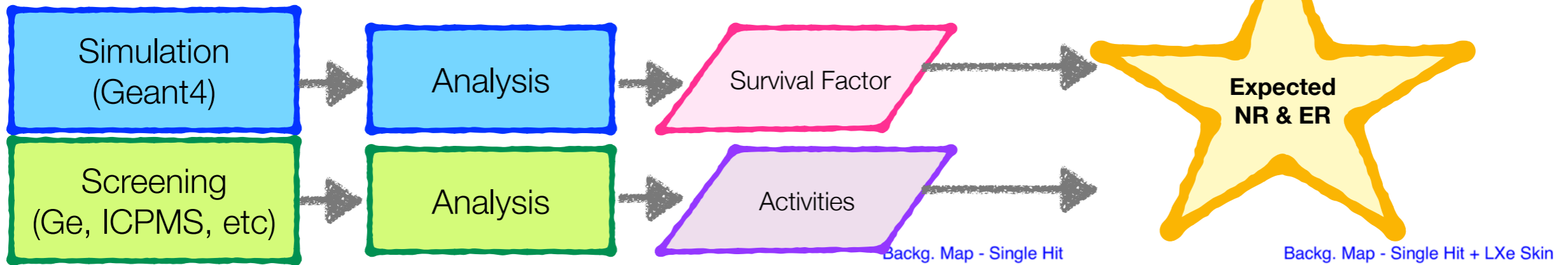
LZ CDR: arXiv:1509.02910



The LZ Dark Matter Experiment



LZ Background Model



Simulations

Detector components

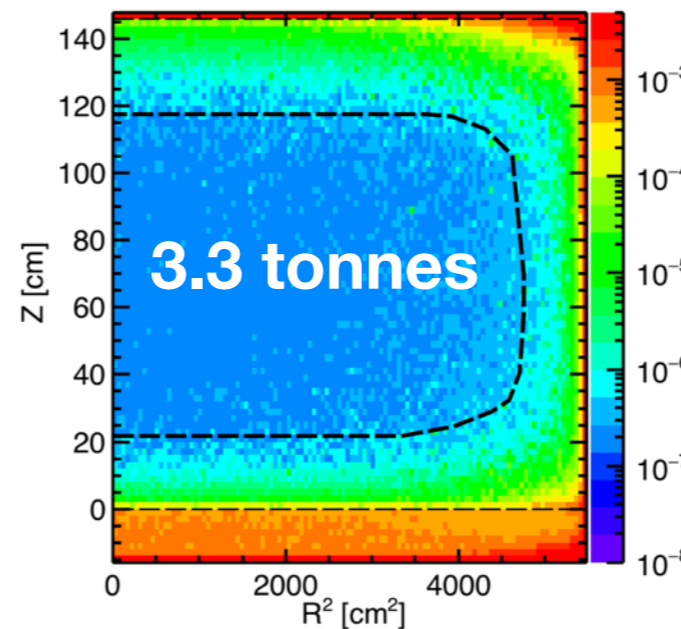
- Gammas:
 - U238 chain
 - Th232 chain
 - K40
 - Co60
 - Sc46
- Neutrons:
 - U early (α, n)
 - U late (α, n)
 - Th (α, n)

Other

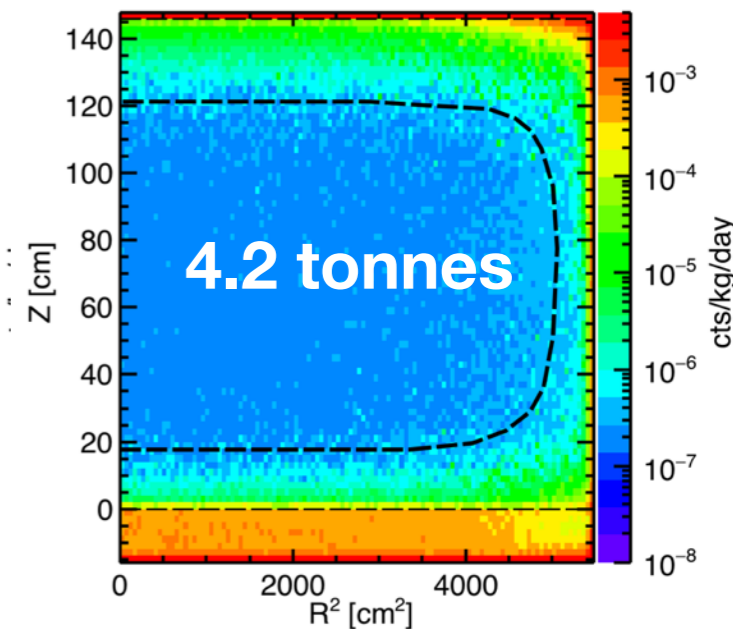
- Rock/cavern gammas
- Radon

Analytical

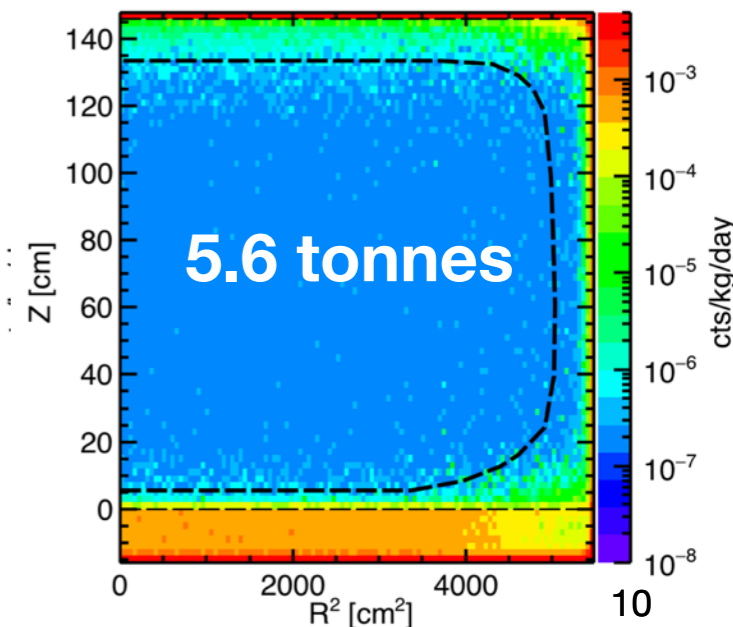
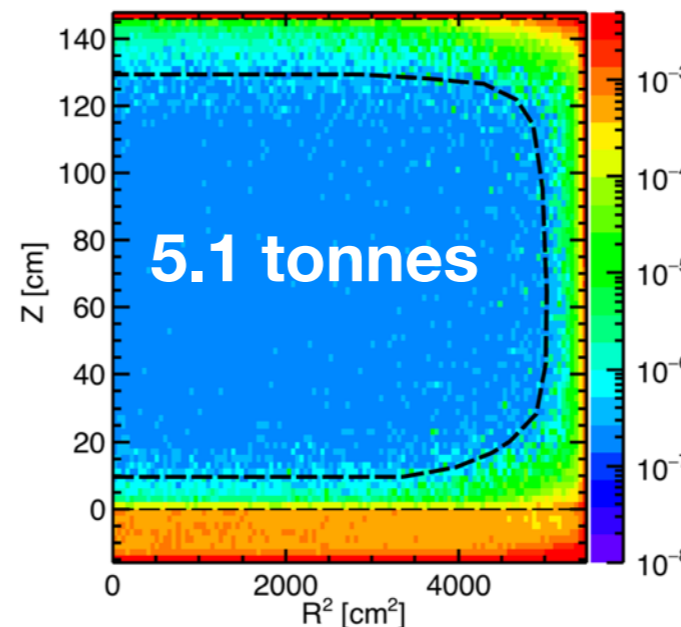
- Dust & plateout
- Intrinsic Xe (Kr, Ar)
- Neutrinos



Backg. Map - Single Hit + Gd-LS Outer Detector

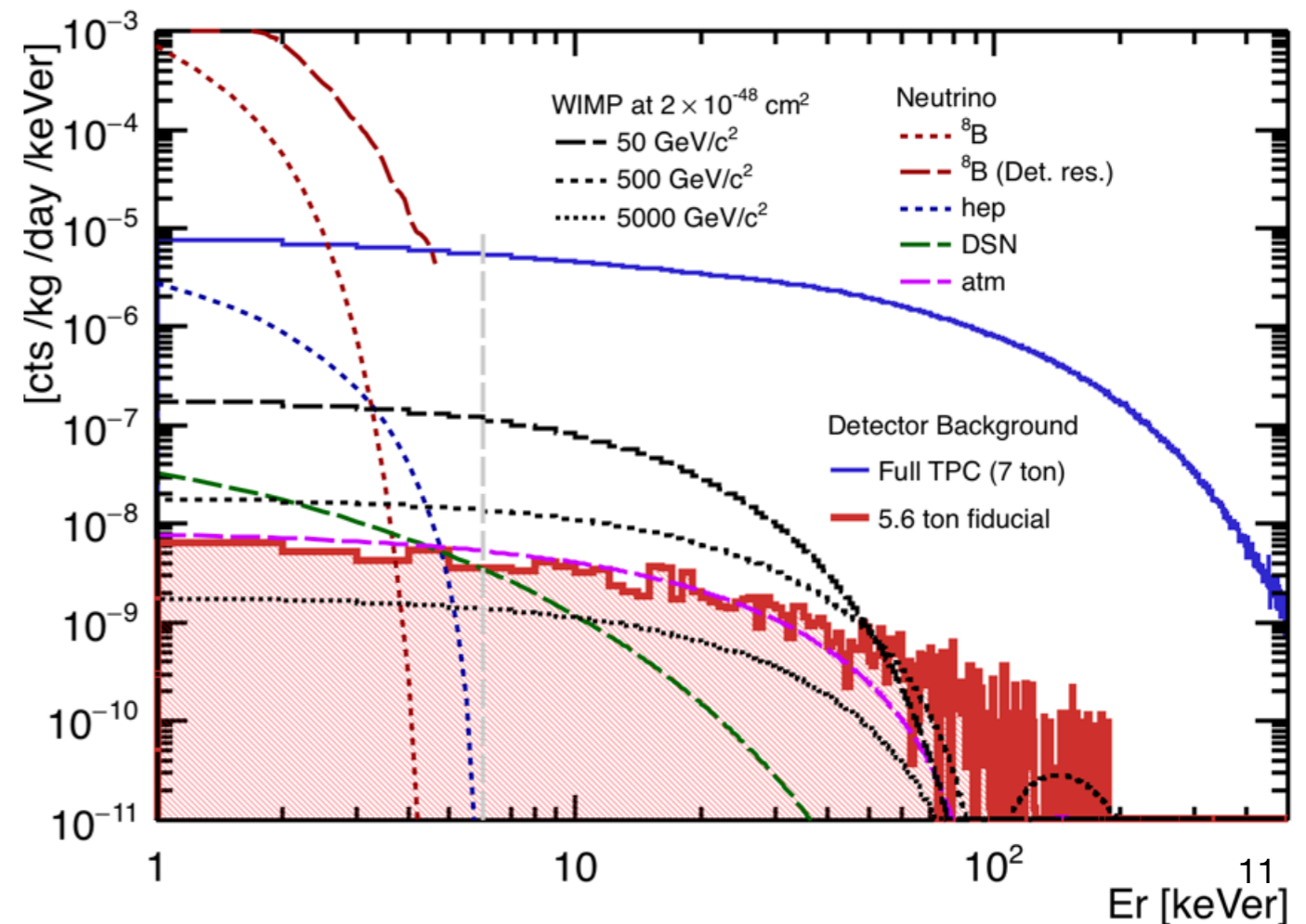
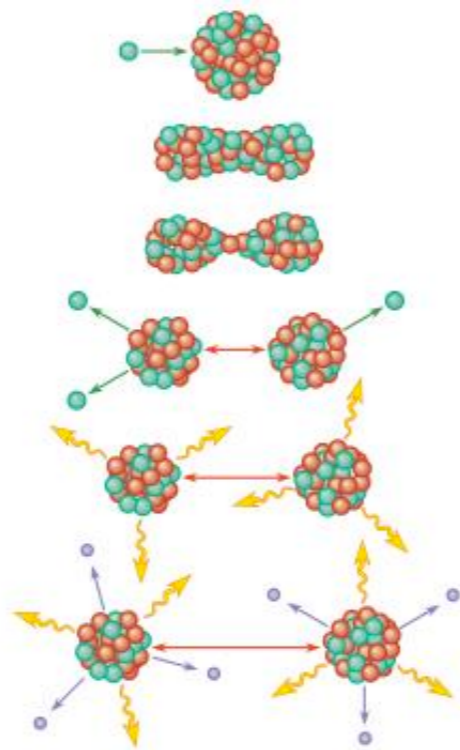
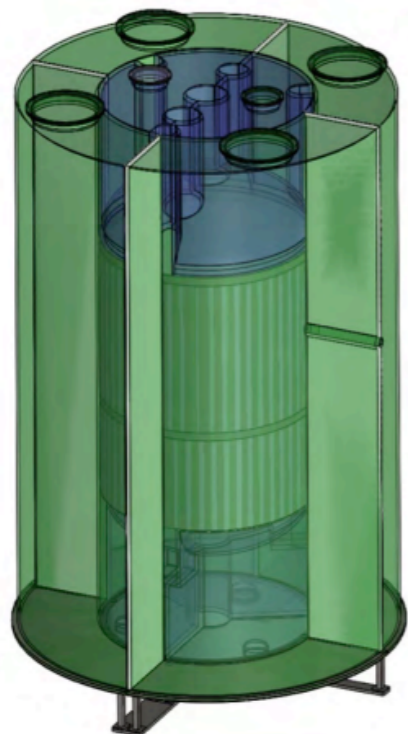
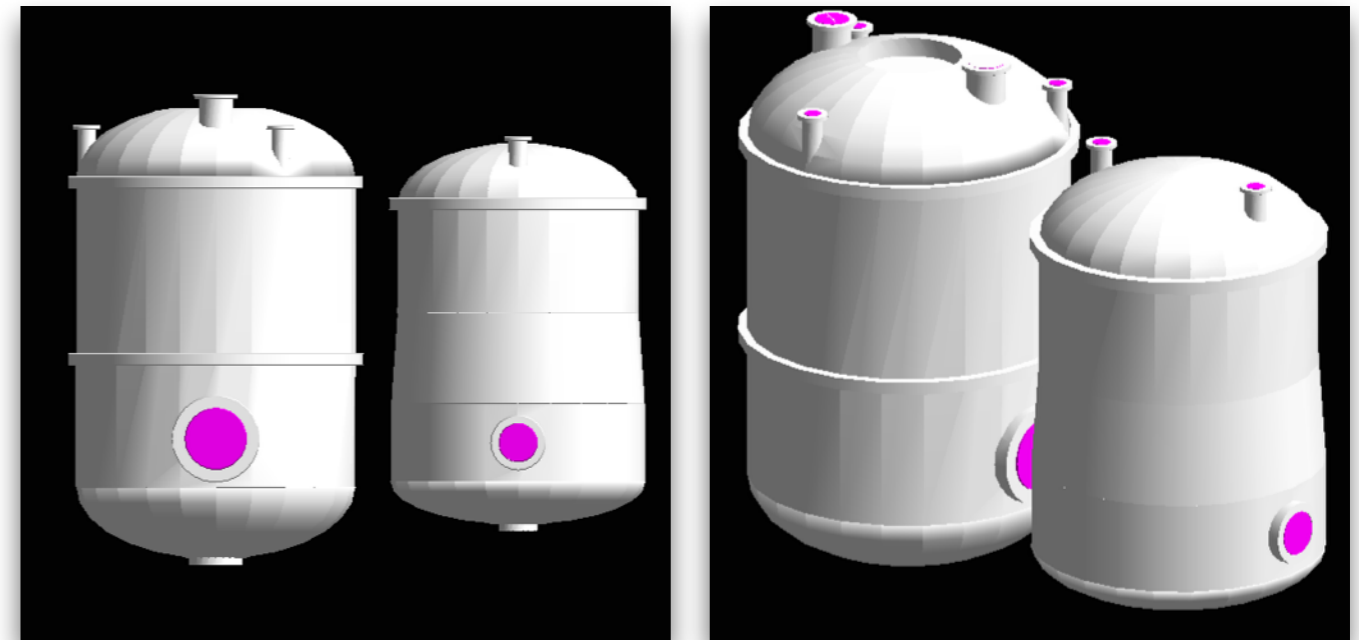


Backg. Map - Single Hit + LXe Skin + Gd-LS Outer Detector



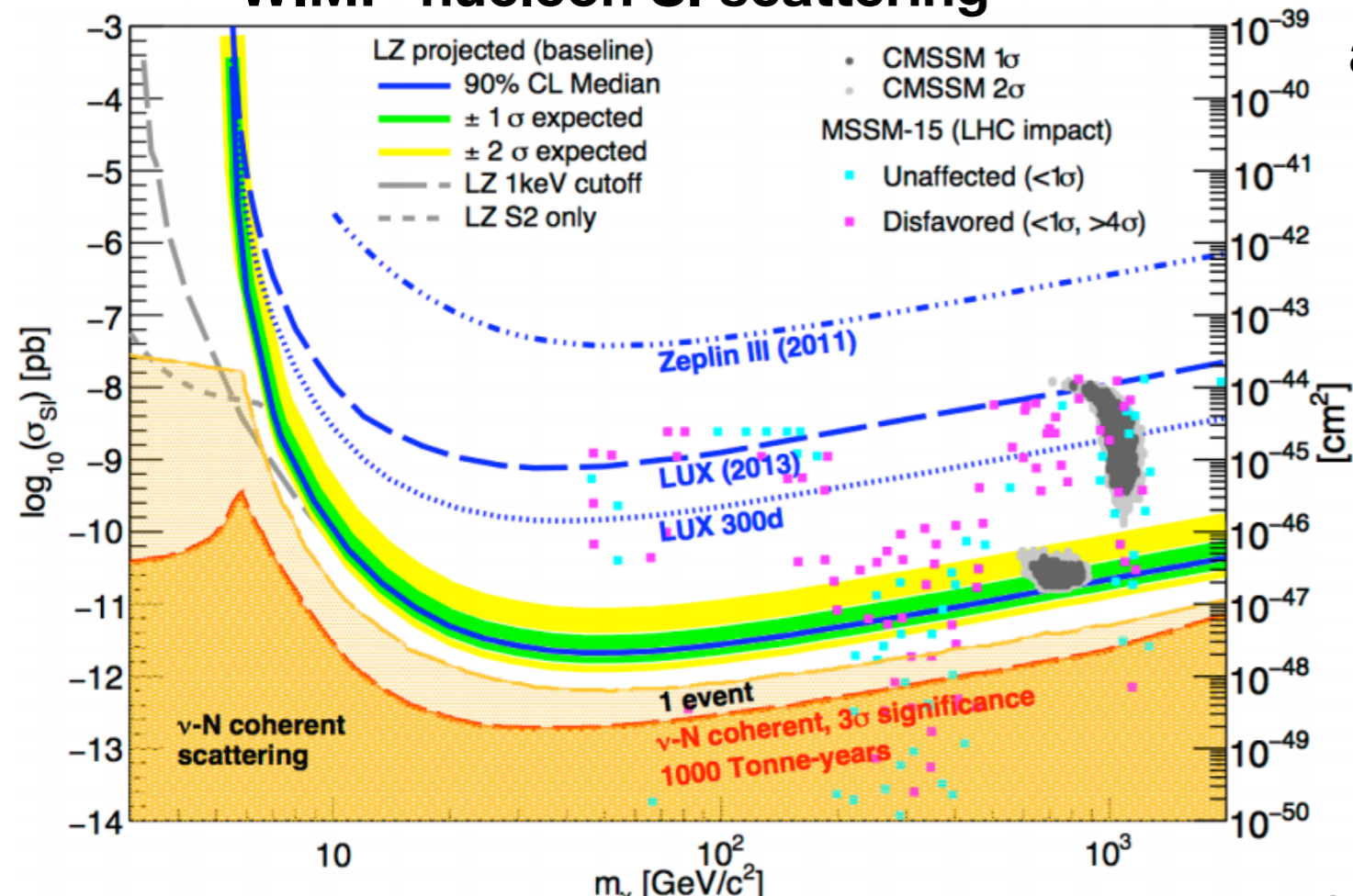
LZ Simulation Work

- Backgrounds analysis code and cut validation
- NR & ER counts for Technical Design Review Backgrounds Table
- Correct treatment of Uranium spontaneous fission events
- Geometry - updating LZSim (Geant4) to latest CAD models
- Expected background rate in scintillator veto
- Validation of neutron & gamma survival probabilities in different detector materials



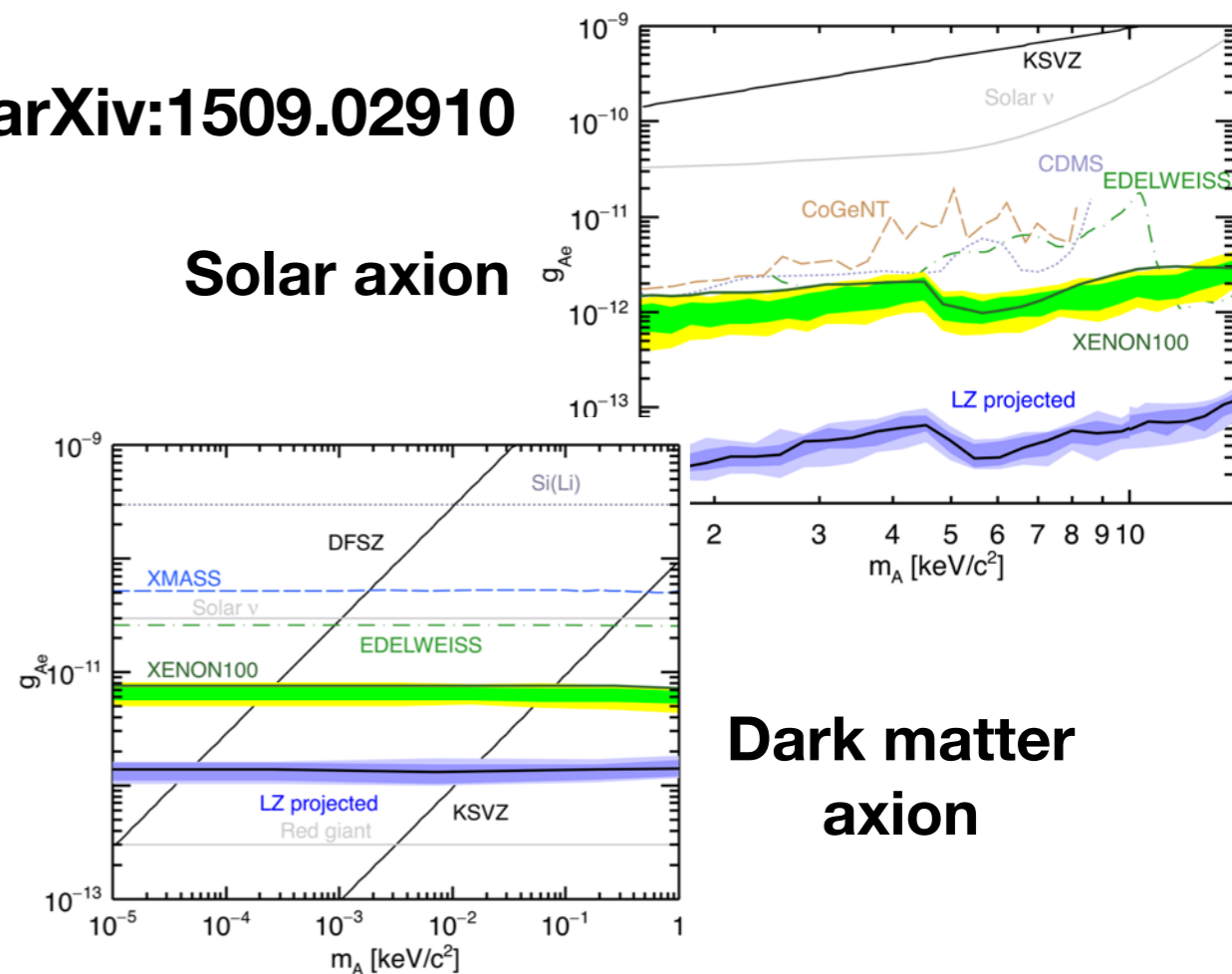
LZ Sensitivity

WIMP-nucleon SI scattering



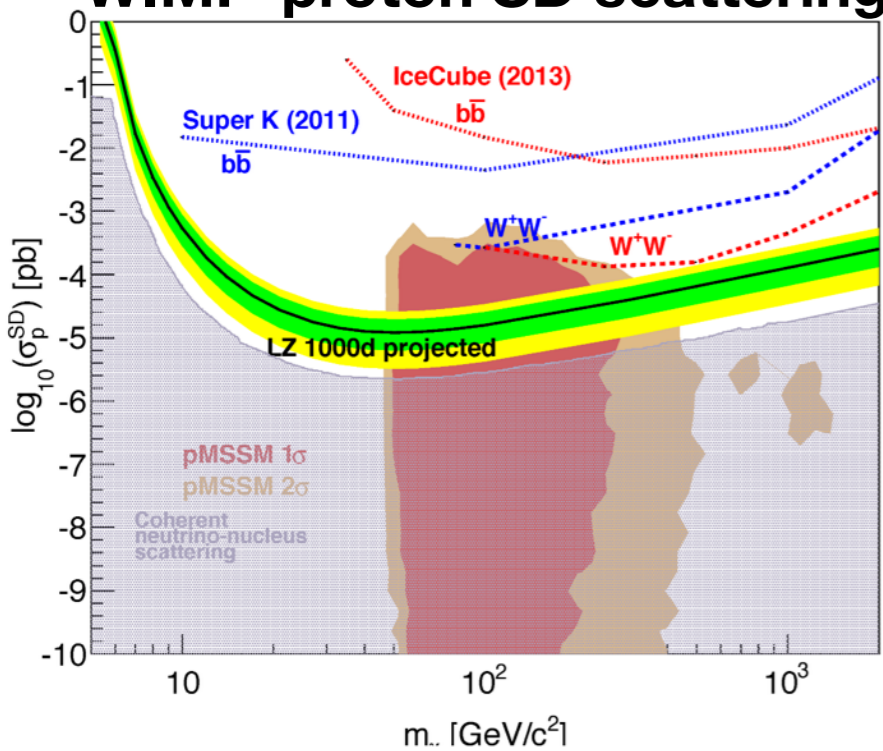
arXiv:1509.02910

Solar axion

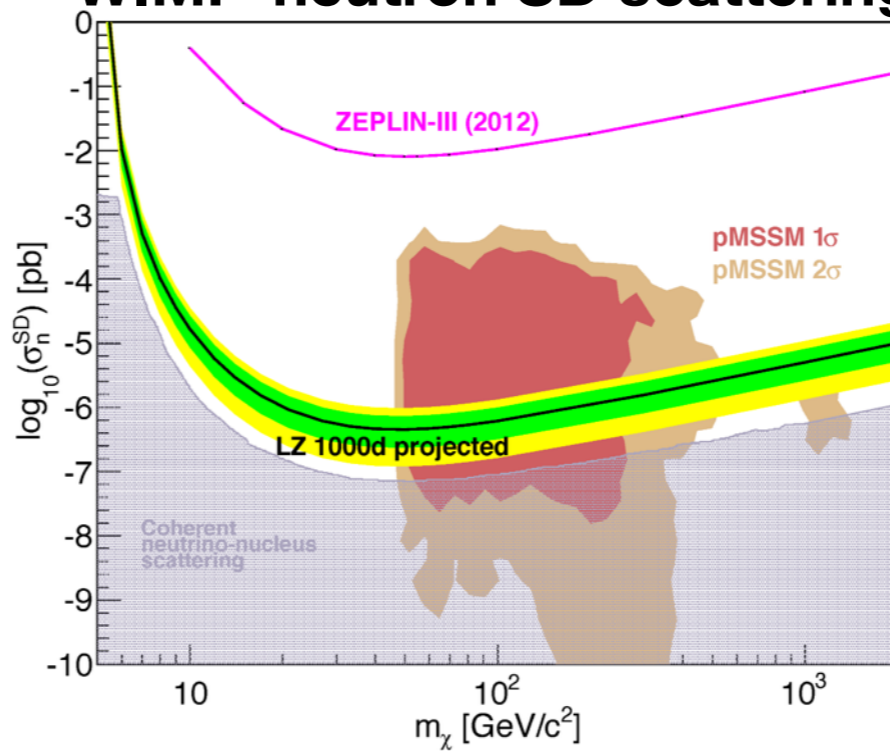


Dark matter axion

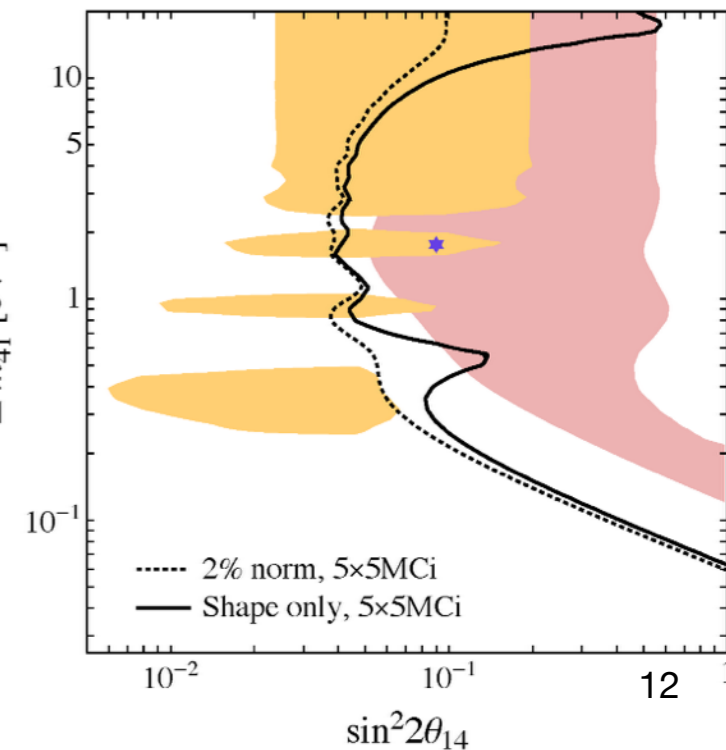
WIMP-proton SD scattering



WIMP-neutron SD scattering



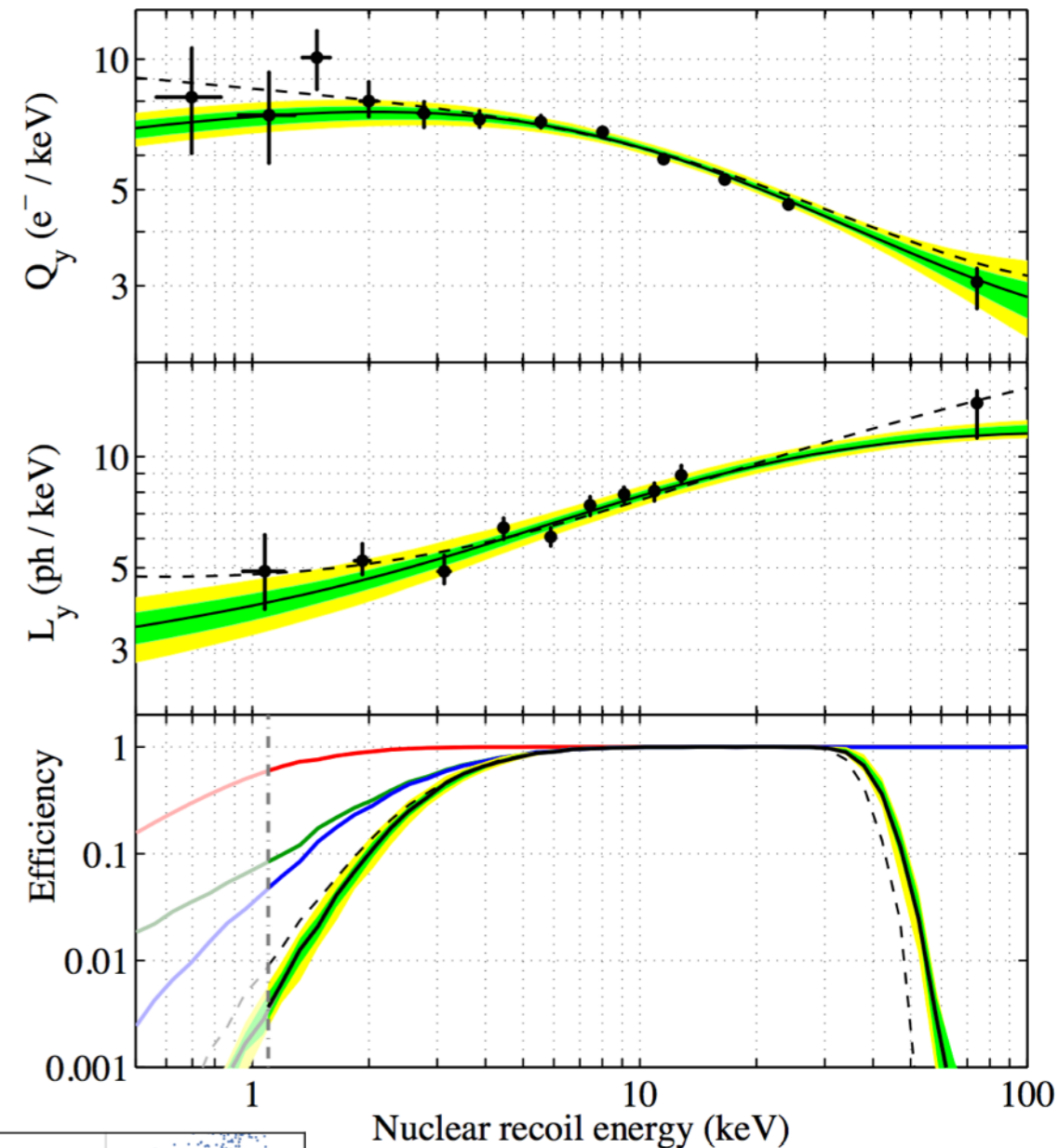
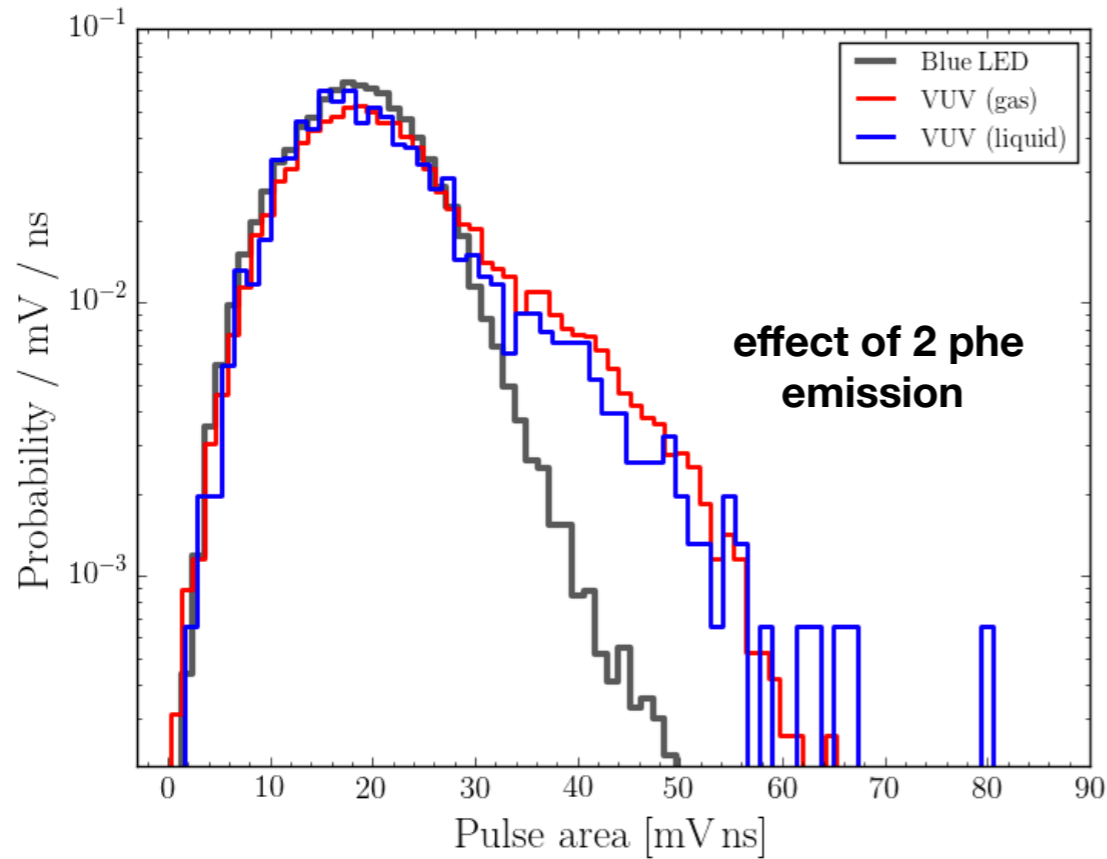
Neutrino physics



Thanks for listening!

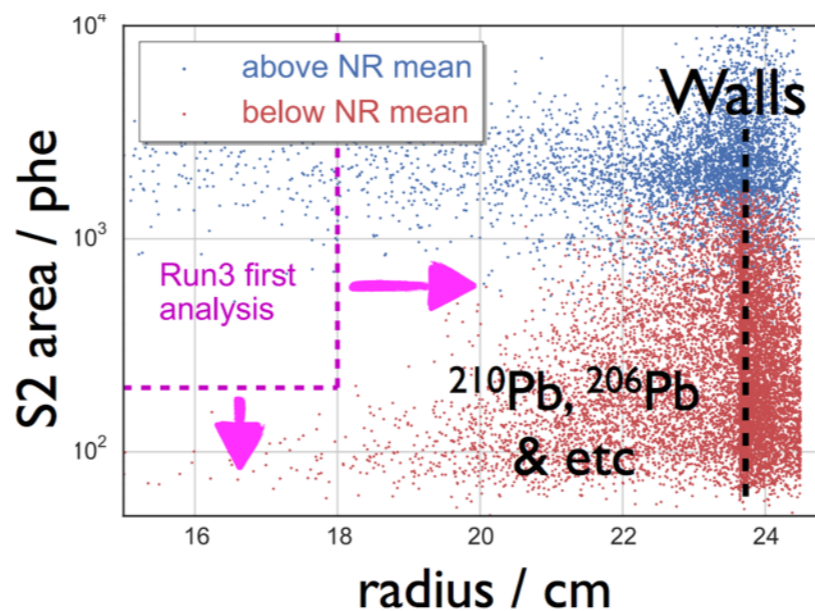
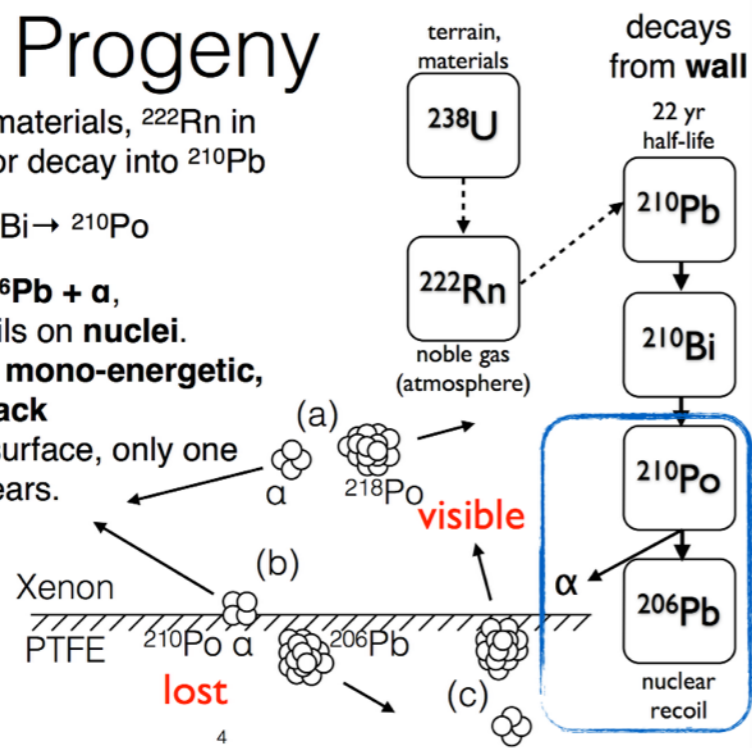


Back-up Slides



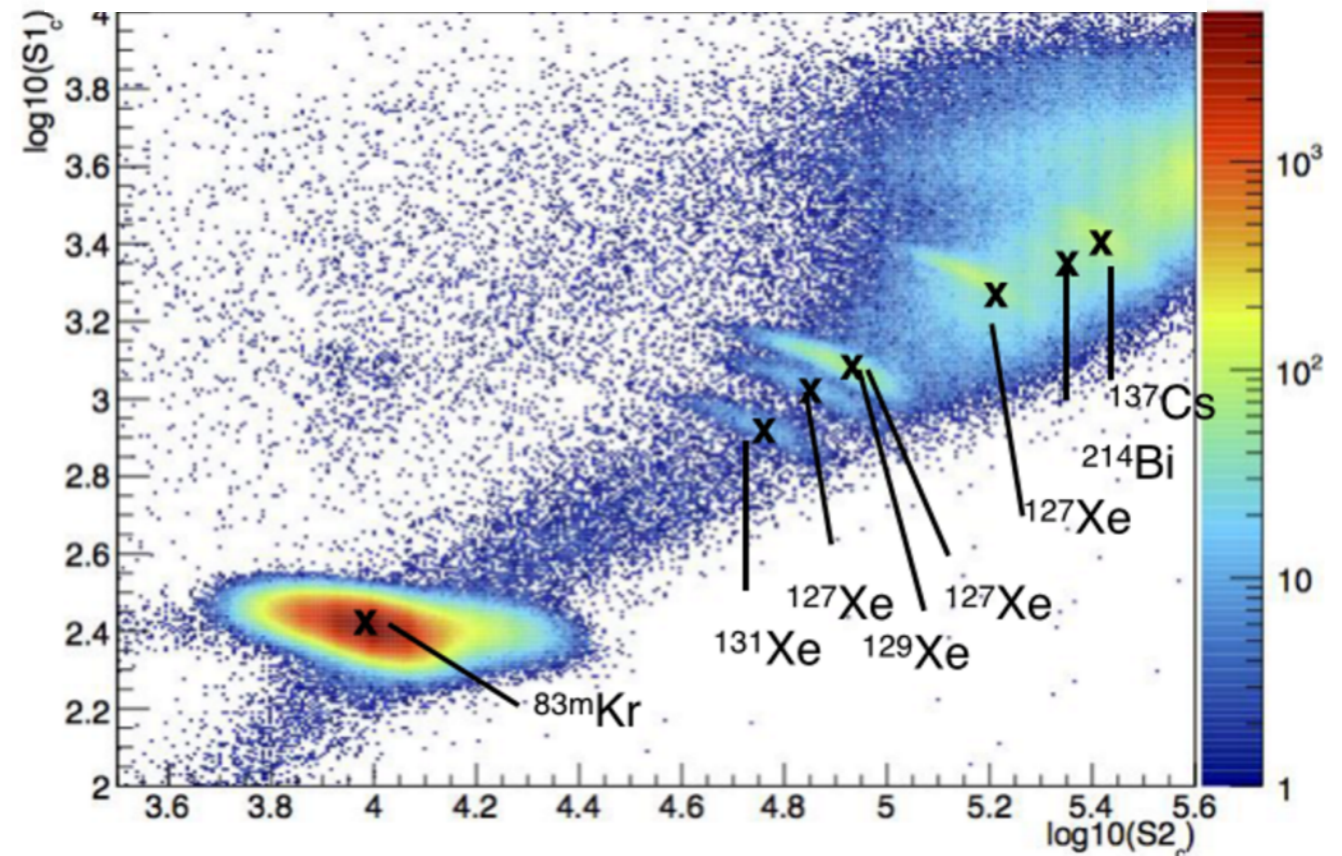
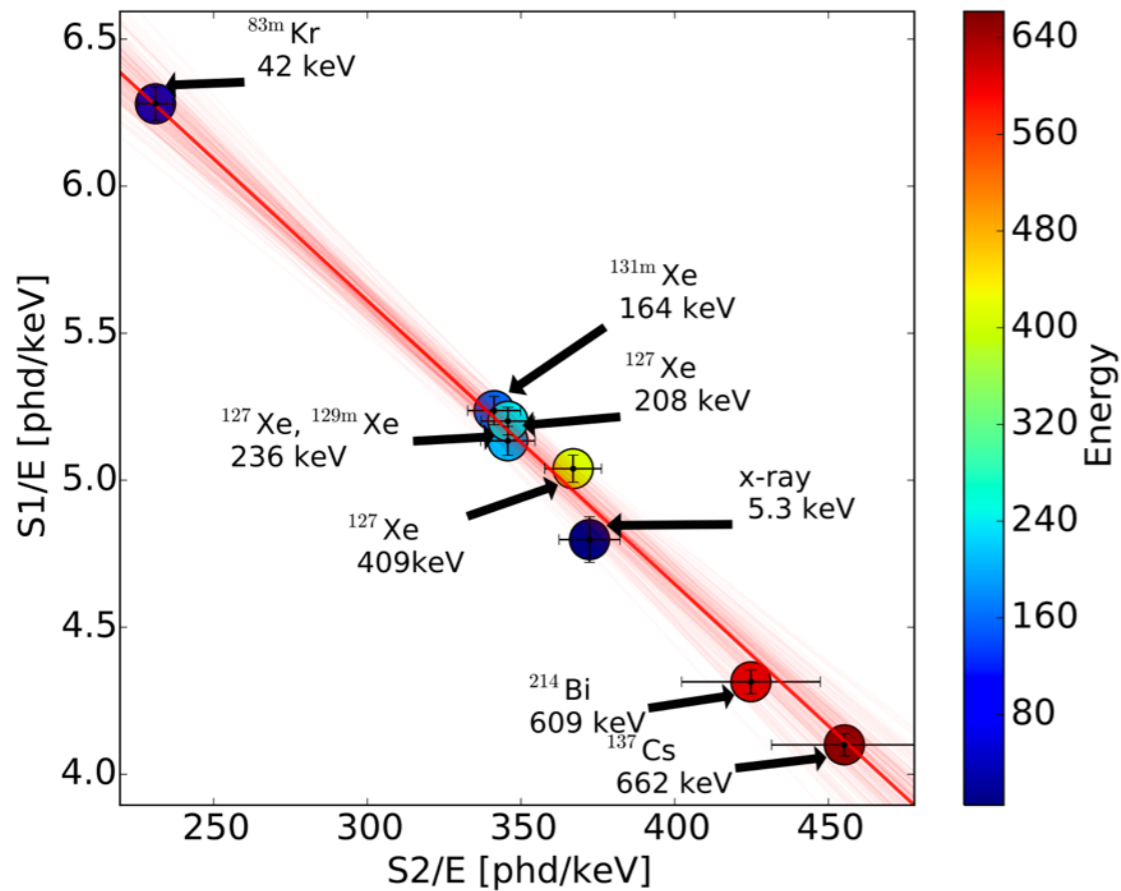
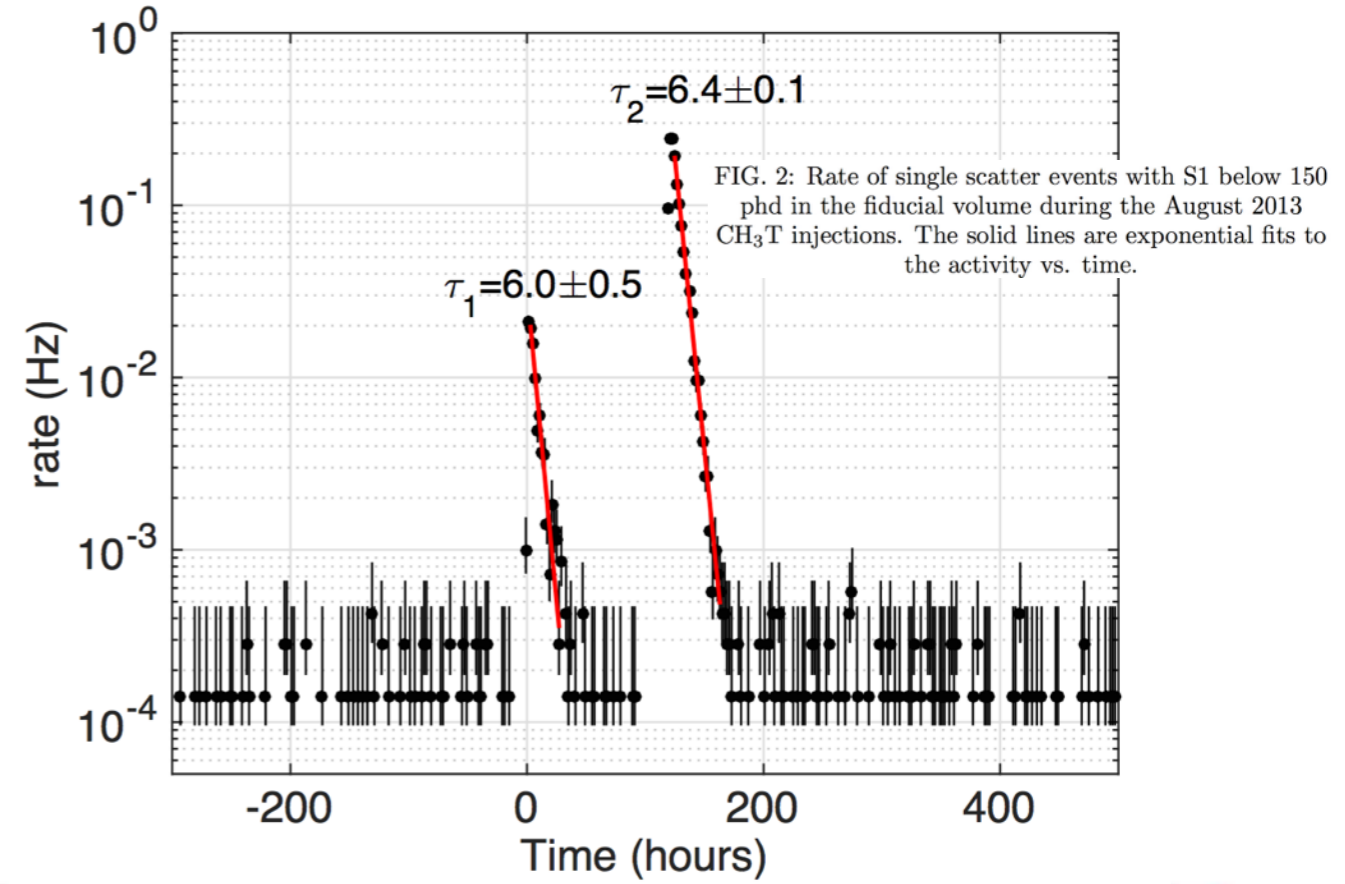
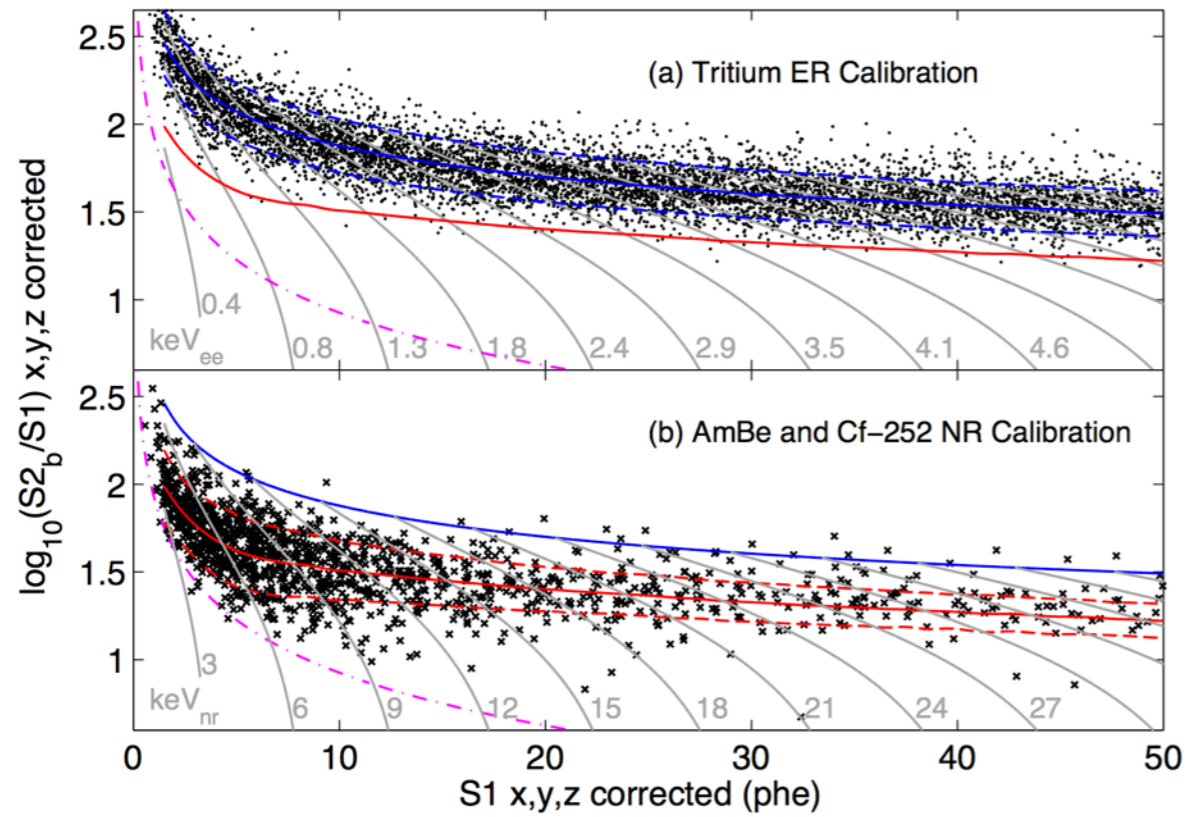
^{222}Rn Progeny

- ^{238}U from materials, ^{222}Rn in the detector decay into ^{210}Pb
- $^{210}\text{Pb} \rightarrow ^{210}\text{Bi} \rightarrow ^{210}\text{Po}$
- $^{210}\text{Po} \rightarrow ^{206}\text{Pb} + \alpha$, ^{206}Pb recoils on nuclei. Two-body: **mono-energetic, back-to-back**. From the surface, only one recoil appears.

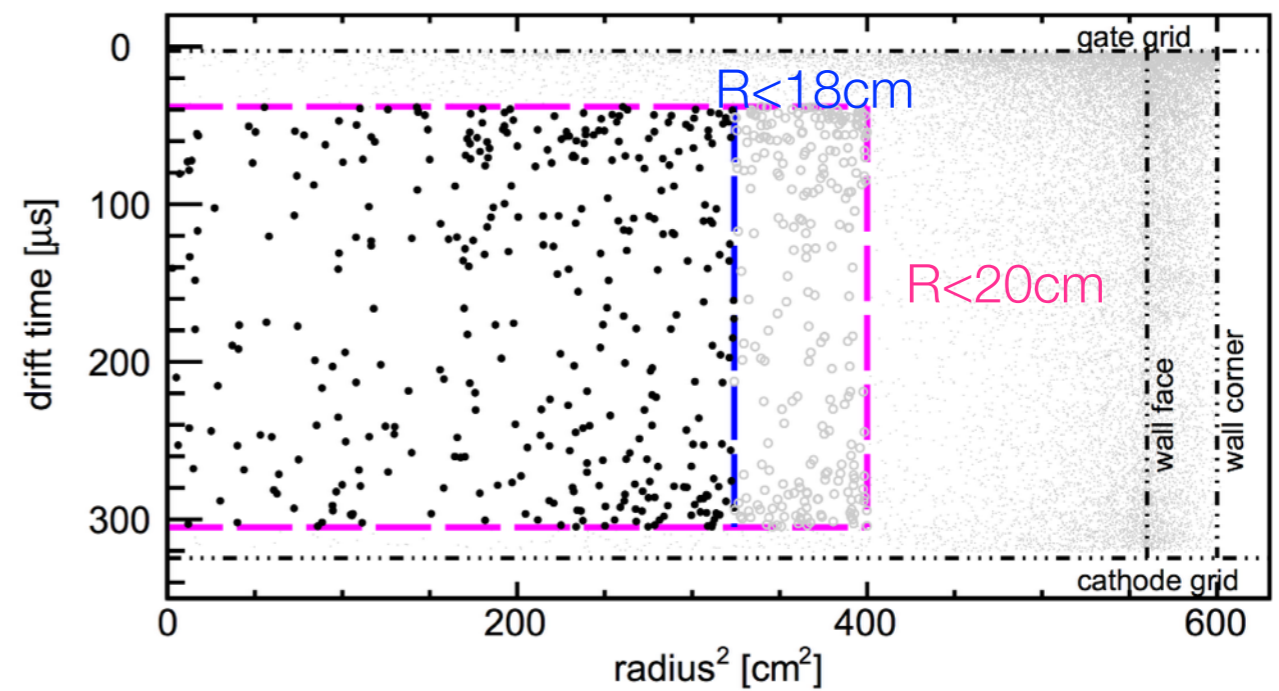
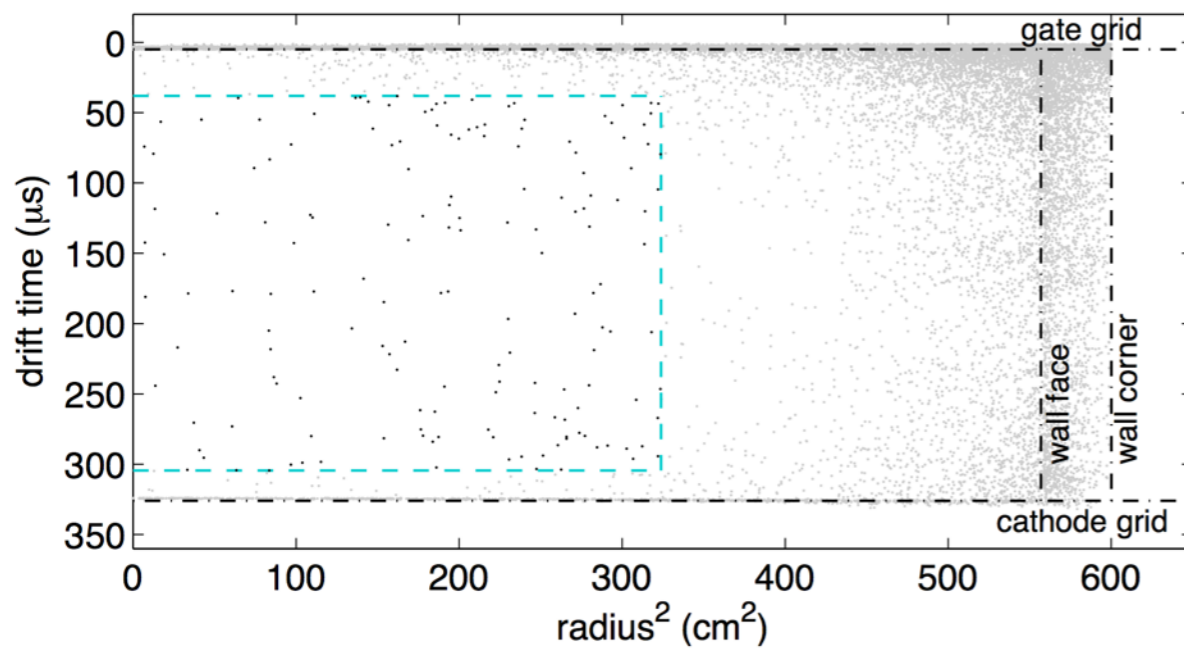
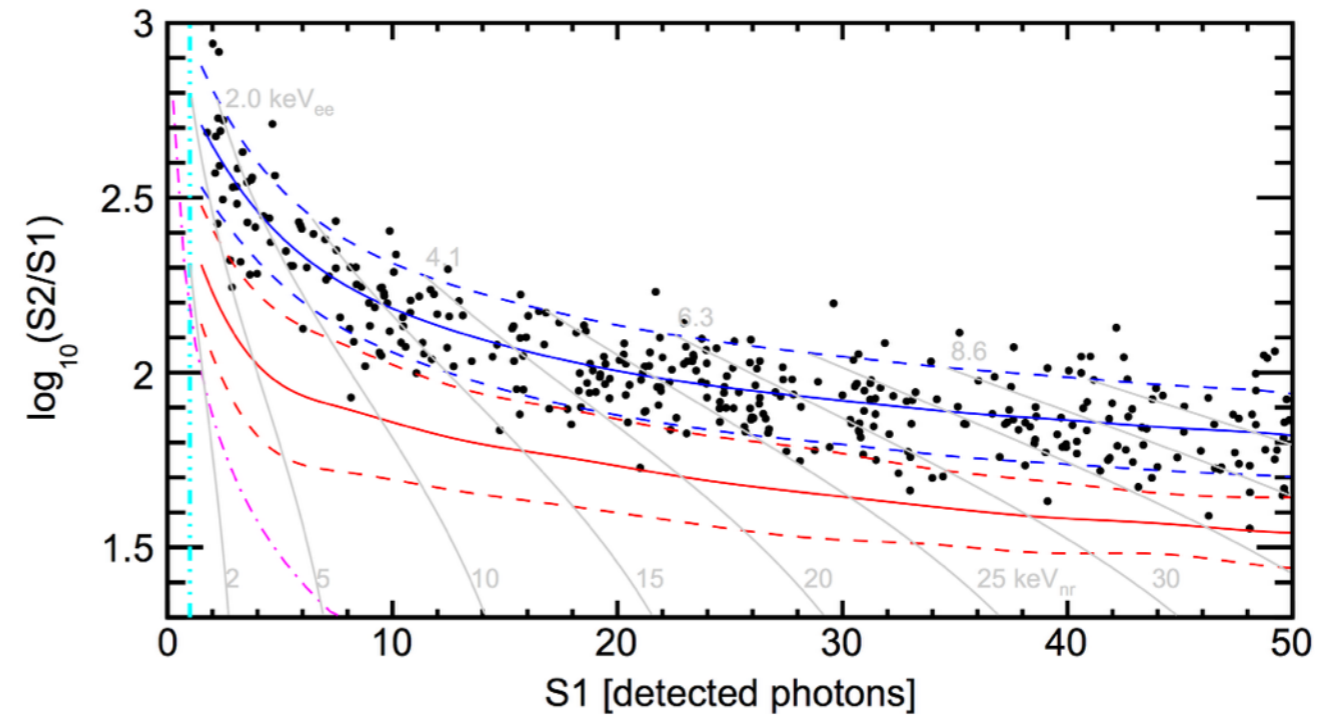
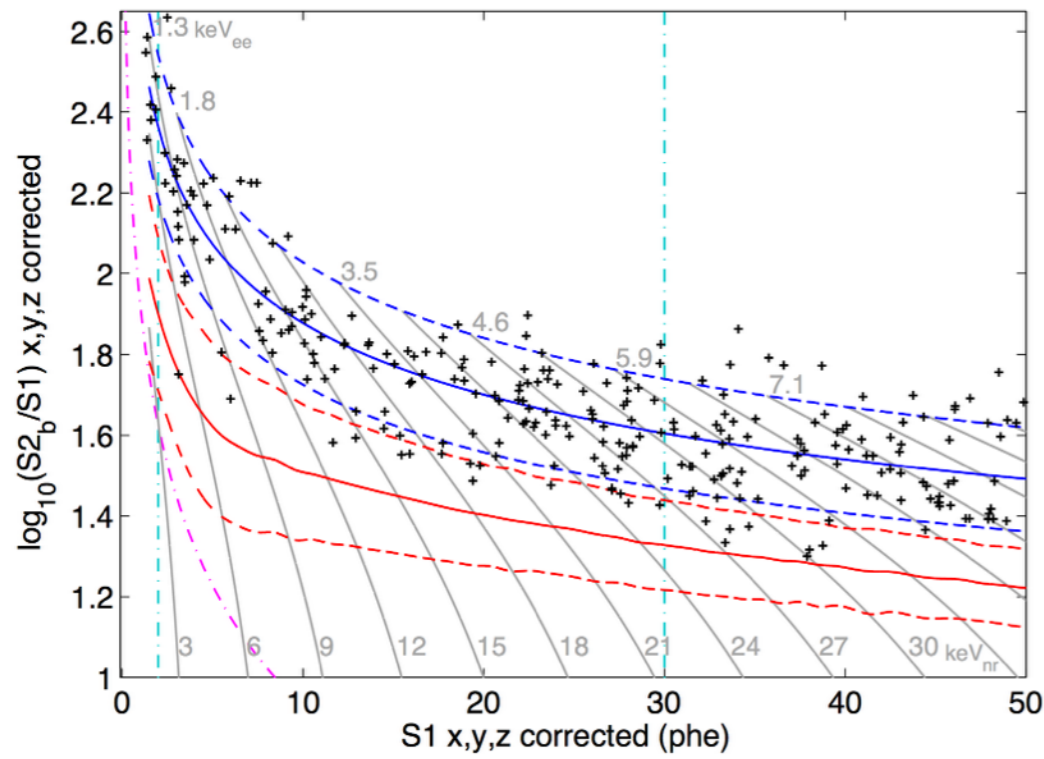


Light yield measured down to 0.7 keV, charge yield 1.1 keV

LUX Calibrations

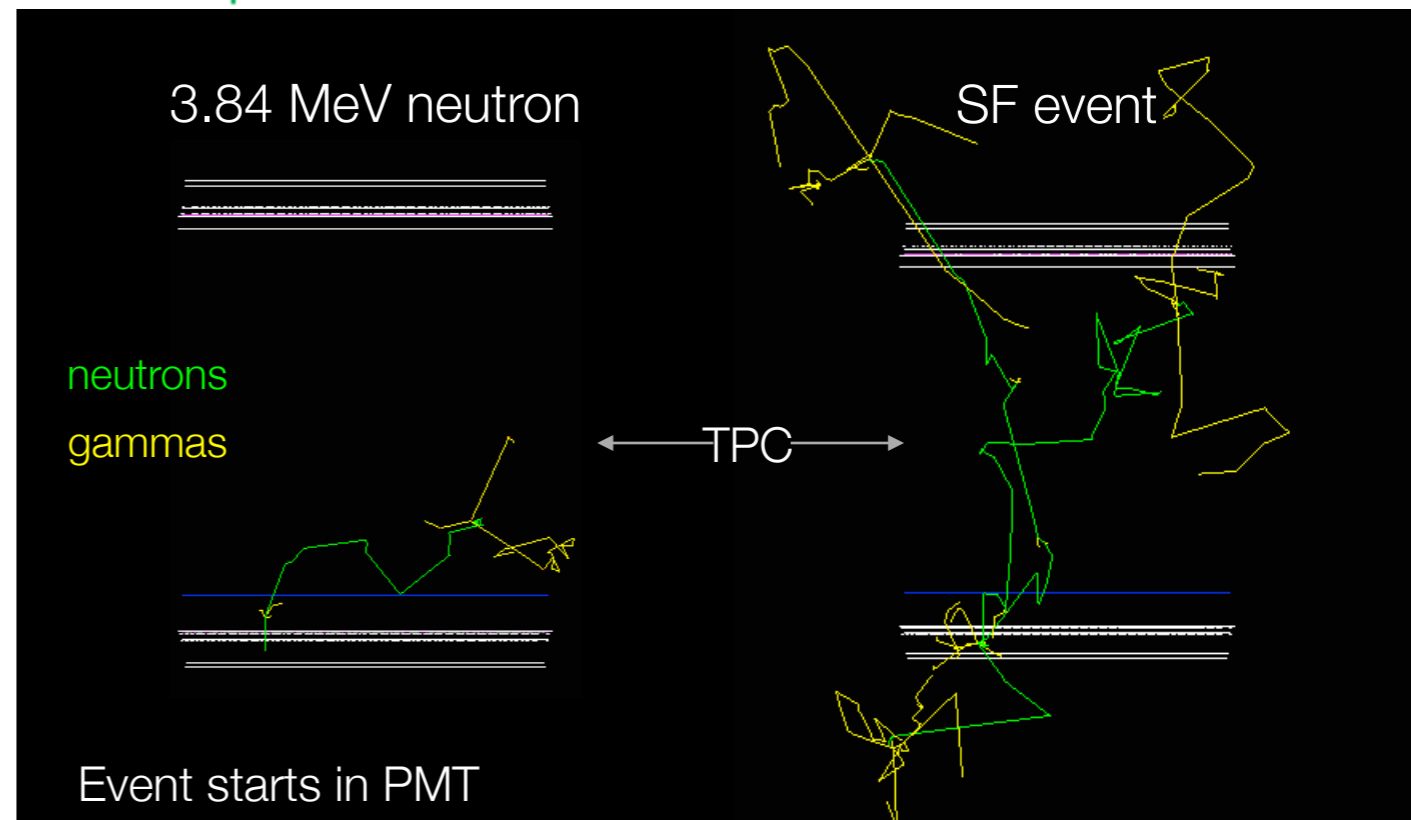
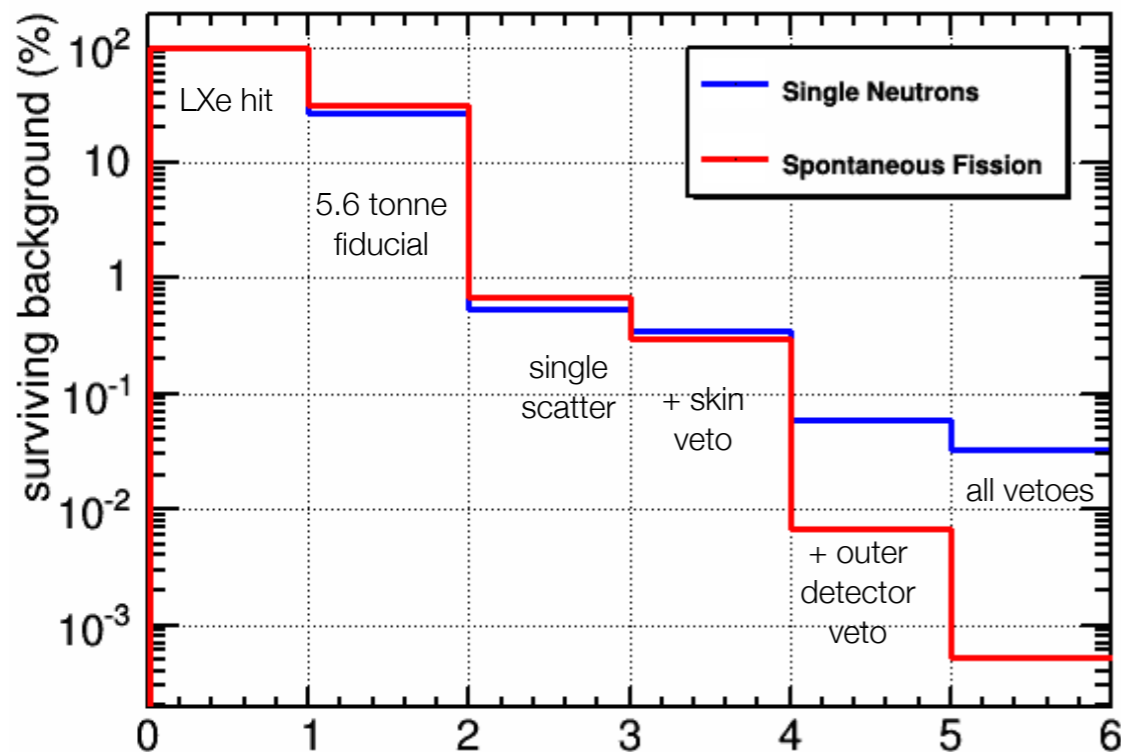
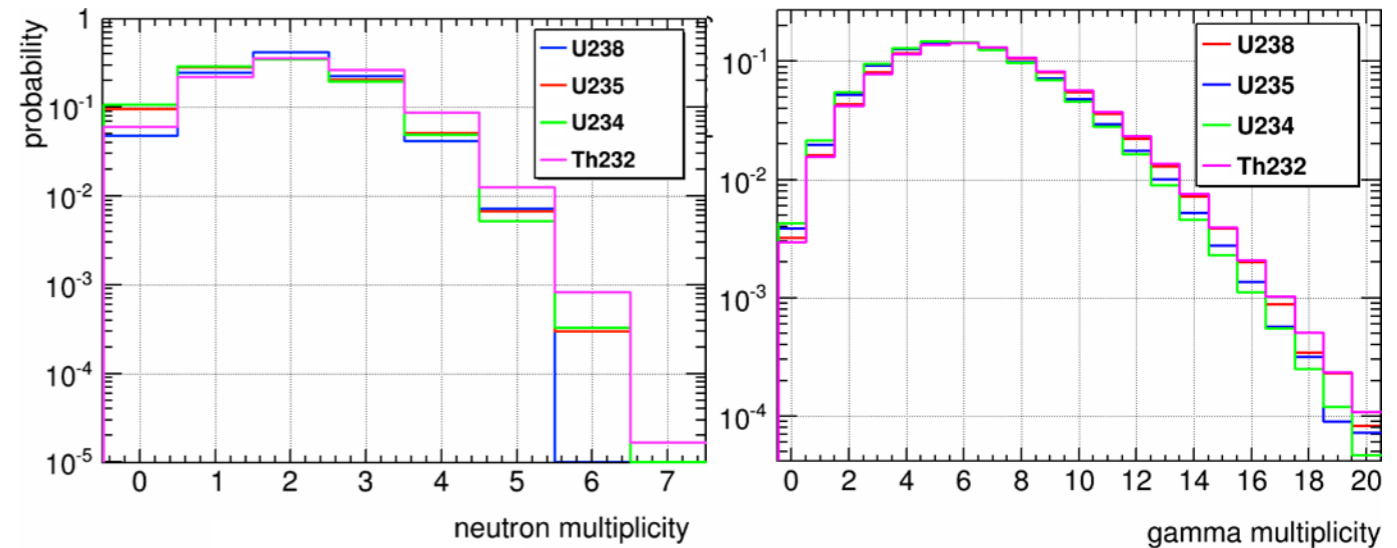


Run 3 vs Run 3 Reanalysis



Spontaneous Fission

- S.F. can be a dominant background source in some materials
- Aim: determine if multiplicity of neutrons and gammas allows better rejection
- Development of Geant4 SF generator for LZSim
- Scintillator veto has big impact on rejection ability - ~50-100x better for S.F.
- S.F. rejection good enough to remove it from background estimates



LZ Backgrounds Table

Intrinsic Contamination Backgrounds	Mass	U-early	U-late	Th-early	Th-late	⁶⁰ Co	⁴⁰ K	n/yr	ER	NR
	(kg)	(mBq/kg)						(cts)	(cts)	
Upper PMT structure	40.2	1.45	0.10	0.25	0.21	0.00	0.50	3.96	0.01	0.002
Lower PMT structure	64.1	0.85	0.06	0.15	0.12	0.00	0.33	5.49	0.01	0.003
R11410 3" PMTs	93.7	67.1	2.68	2.01	2.01	3.86	62.1	372.5	1.24	0.203
R11410 PMT bases	2.7	525.	74.6	29.1	29.1	3.60	109.	76.7	0.17	0.033
R8520 Skin 1" PMTs	4.2	60.5	5.19	4.75	4.75	24.2	333.	11.4	0.09	0.002
R8520 Skin PMT bases	0.9	513.	58.3	24.2	24.2	3.91	108.	23.3	0.06	0.003
PMT cabling	85.5	29.8	1.47	3.31	3.15	0.65	33.14	89.5	0.92	0.008
TPC PTFE	343.	0.02	0.02	0.01	0.01	0.00	0.10	24.1	0.17	0.007
Grid wires	0.33	1.20	0.27	0.33	0.49	1.60	0.40	0.02	0.01	0.000
Grid holders	69.6	1.60	0.09	0.28	0.23	0.00	0.54	6.92	0.02	0.003
Field-shaping rings	262.	5.89	1.81	1.13	1.08	0.00	1.83	32.2	1.22	0.004
TPC sensors	0.90	8.76	7.28	1.37	1.37	0.20	5.39	0.72	0.08	0.000
TPC thermometers	0.70	332.	329.	136.	136.	4.90	658.	85.2	3.67	0.010
Xe recirc. tubing	5.2	0.02	0.02	0.01	0.007	0.00	0.10	0.37	0.00	0.000
HV conduits – cables	138.	1.80	2.00	0.40	0.60	1.40	1.20	15.6	0.72	0.001
HX and PMT conduits	200.	1.05	0.21	0.27	0.38	1.18	0.60	11.9	0.41	0.000
Cryostat vessel	2.14E3	1.60	0.09	0.28	0.23	0.00	0.54	213.	0.86	0.019
Cryostat seals	4.5	102.	102.	34.0	34.0	7.27	22.6	40.3	0.79	0.001
Cryostat insulation	23.8	18.9	18.9	3.45	3.45	1.97	51.7	85.2	0.92	0.003
Cryostat Teflon liner	70.7	0.02	0.02	0.01	0.01	0.00	0.10	4.97	0.00	0.000
Outer detector tanks	4.00E3	0.15	0.37	0.02	0.06	0.04	4.32	101.	0.14	0.0002
Liquid scintillator	2.08E4	0.01	0.01	0.01	0.01	0.00	0.00	22.9	0.00	0.00
Outer detector PMTs	122.	1.50E3	1.50E3	1.07E3	1.07E3	0.00	3.90E3	2.09E4	0.08	0.022
OD PMT supports	620.	1.20	0.27	0.33	0.49	1.60	0.40	37.0	0.25	0.00
²²² Rn (0.67 mBq)									23.2	-
²²⁰ Rn (0.07 mBq)									4.68	-
^{nat} Kr (0.015 ppt g/g)									24.5	-
^{nat} Ar (0.45 ppb g/g)									2.47	-
Subtotal (Non-ν counts)									66.7	0.33
Physics Backgrounds										
¹³⁶ Xe 2νββF									53.8	0
Astrophysical ν counts (pp+ ⁷ Be)									271	0
Astrophysical ν counts (⁸ B)									0	0
Astrophysical ν counts (Hep)									0	0.002
Astrophysical ν counts (diffuse supernova)									0	0.113
Astrophysical ν counts (atmospheric)									0	0.385
Total									392	0.83
Total (with 99.5% ER discrimination, 50% NR efficiency)									1.96	0.41
Sum of ER and NR in LZ for 1,000 days, 5.6-tonne FV, with all analysis cuts									2.37	