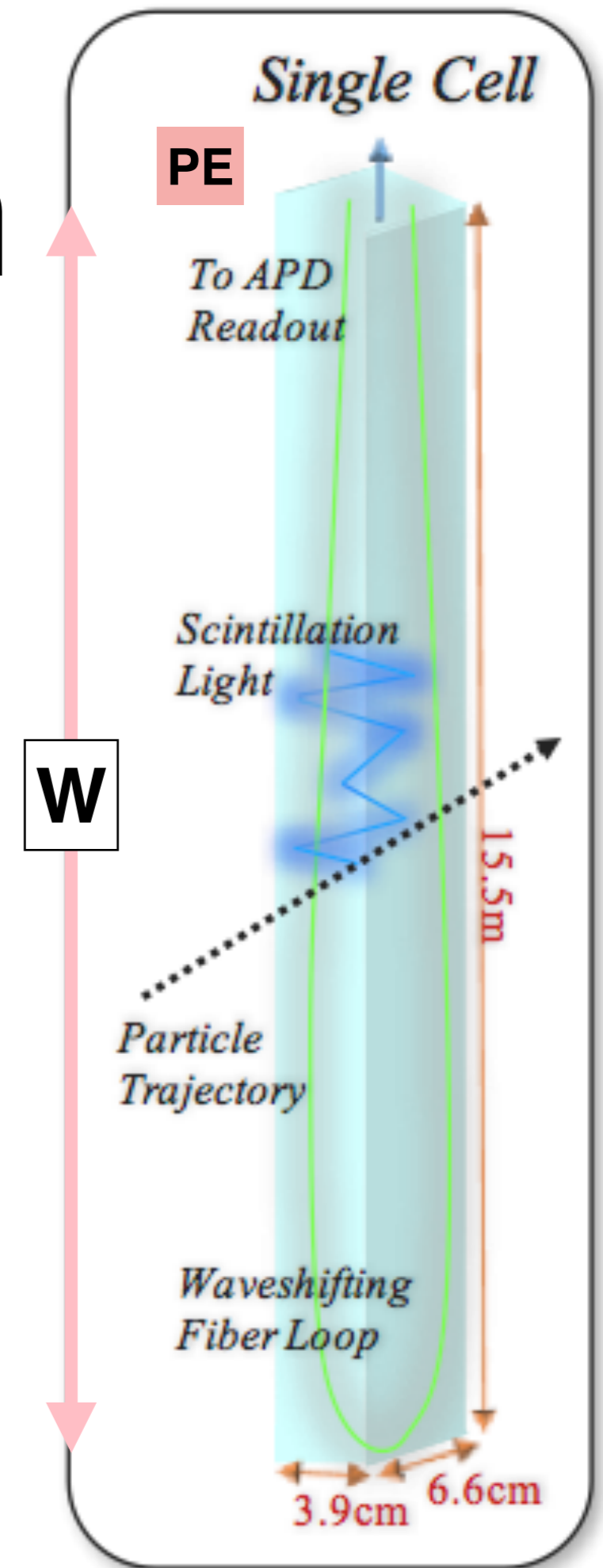


NOvA Calibration

*Calorimetric Energy Scale in the
NOvA Detectors*

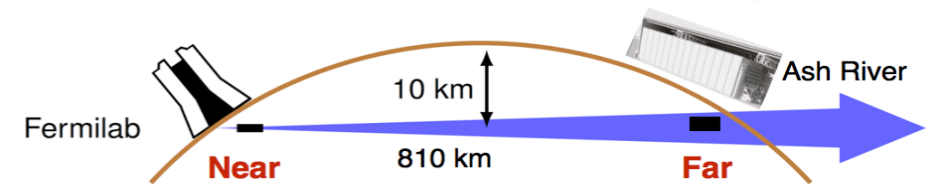
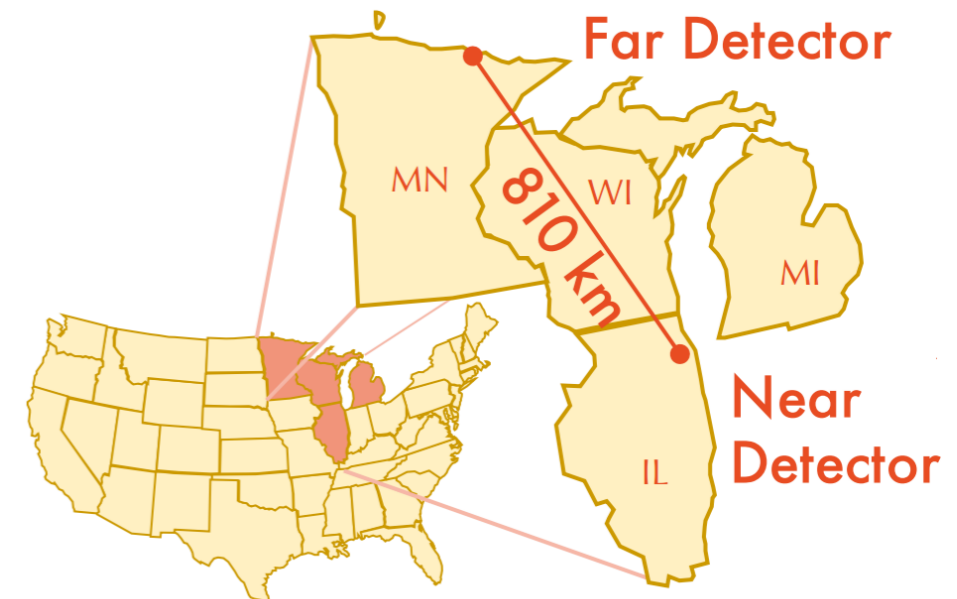
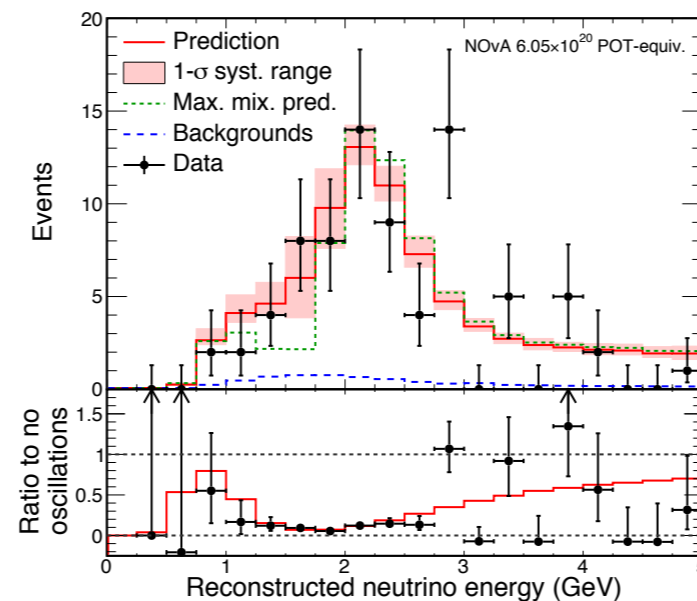
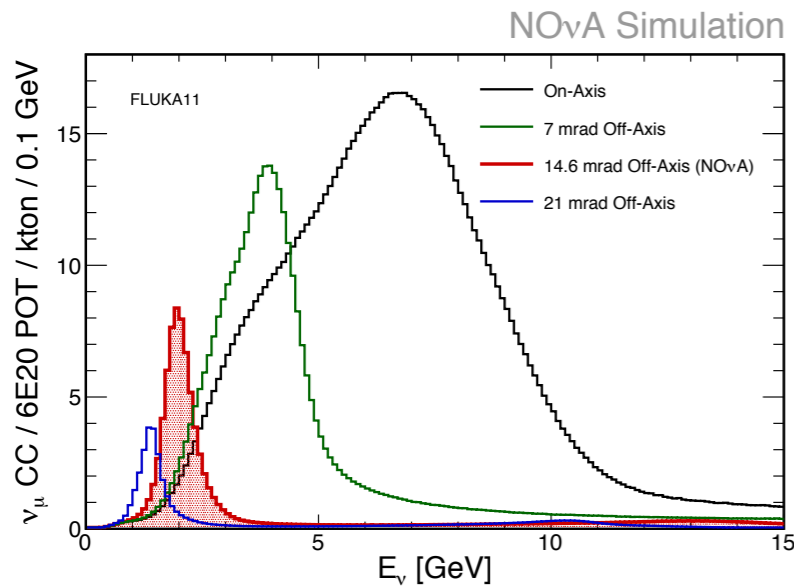
- ◆ The NOvA Detectors
- ◆ What impacts the energy of a hit
- ◆ Calibrate Relative Effects
- ◆ Calculate Energy Scale

Tyler Alion
IOP Bristol
26 March 2018





NuMI Off-Axis ν_{e^-} Appearance Experiment



Neutrino Beam: NuMI (Neutrinos at the Main Injector)
120 GeV Protons on a target produce kaons and pions decaying into muons and neutrinos

Beamline: 810 km between functionally similar Near and Far Detectors, both 14mrad off-axis from beam

Physics Program:

- Three-Flavour Oscillation
- Search for Sterile Neutrinos
- Inclusive ν_{μ} -Charged-Current Cross Section
- Other ND physics
- Supernova Neutrinos

ν_e Appearance

ν_{μ} Disappearance

— θ_{23} Octant?

Maximal?

— Mass Hierarchy?

— CP Violation?

The NOvA Detectors

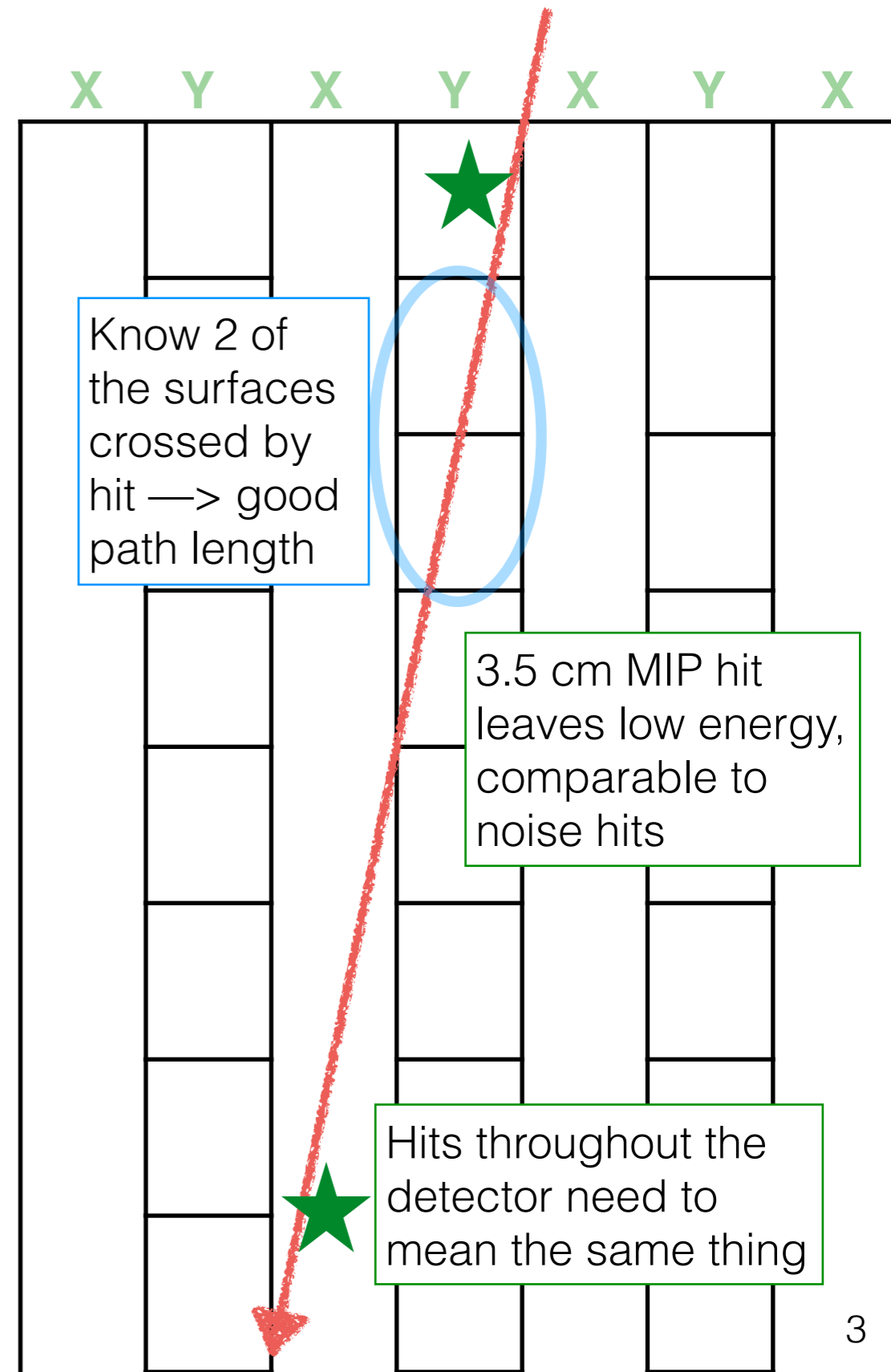
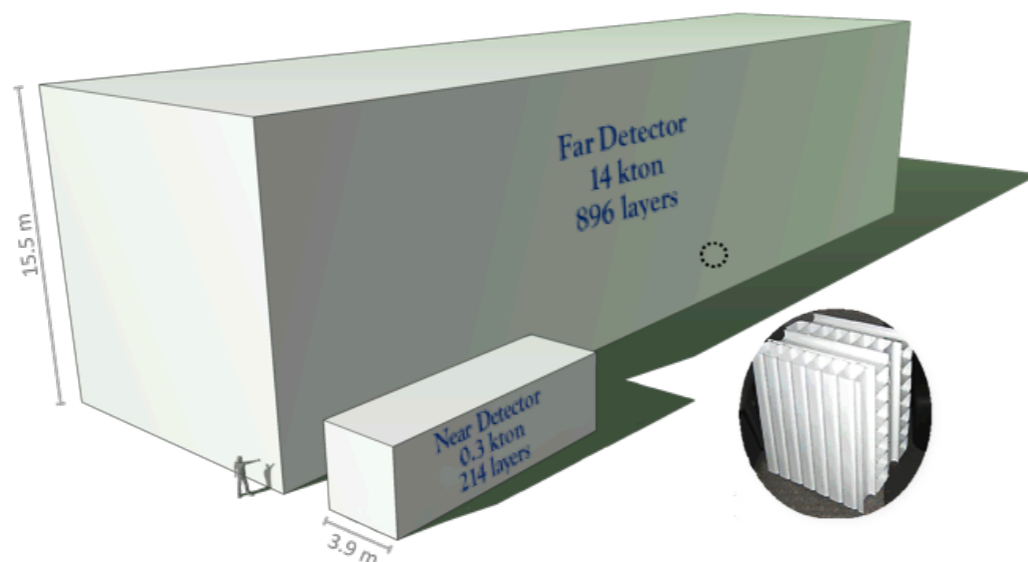
◆ Tracking scintillators

- Orthogonal views X & Y
- Light collection calorimetry

◆ Light collected exclusively through WLS fibers

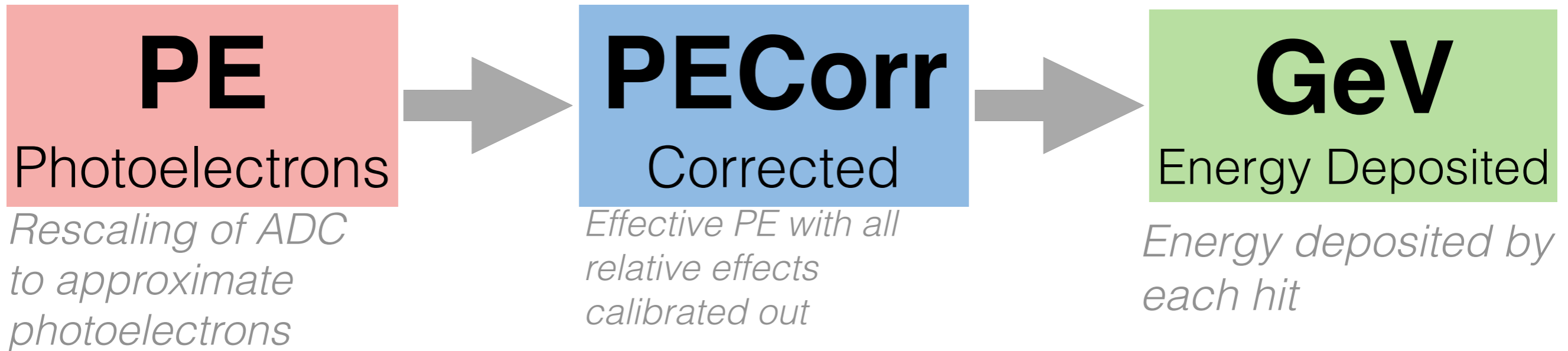
◆ FD about 4x the size of ND in each dimension

◆ “Slice” up continuous readout into events less than a few 100 ns long



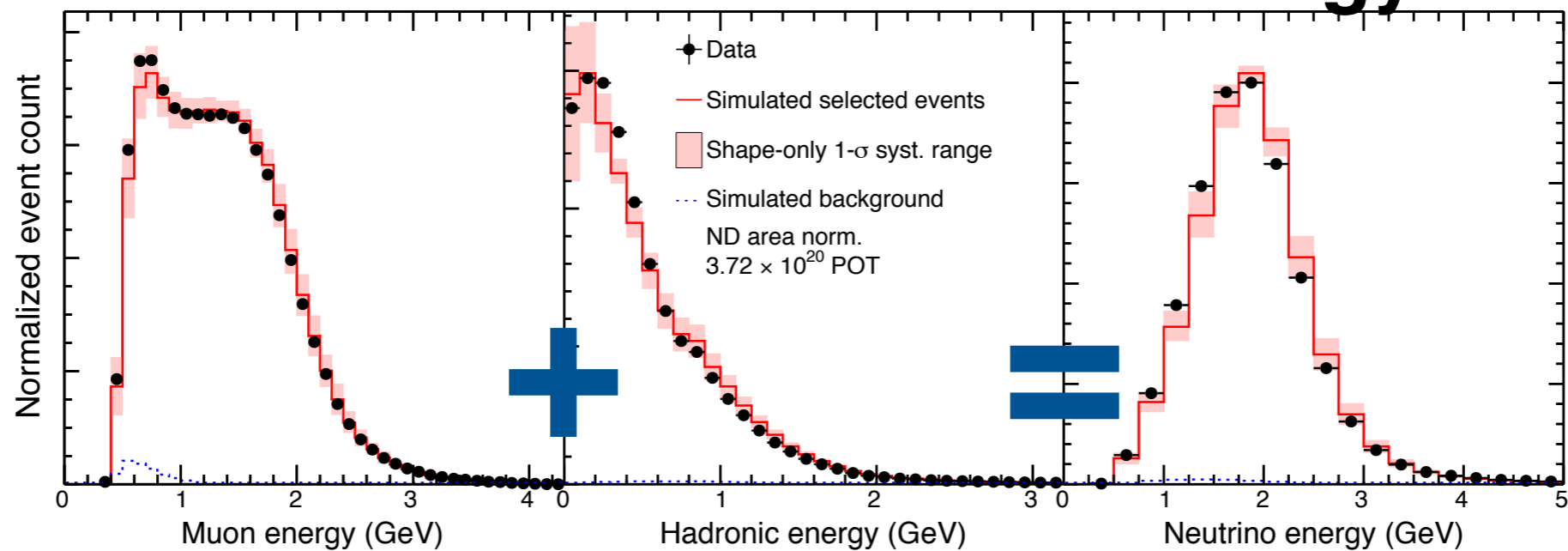
Energy Calibration

With our large sample of Cosmic Muon Data



Reconstructed Neutrino Energy

Muon Energy from Track Length



Energy Calibration directly impacts Reconstructed Hadronic Energy

Bias of Deposited Energy

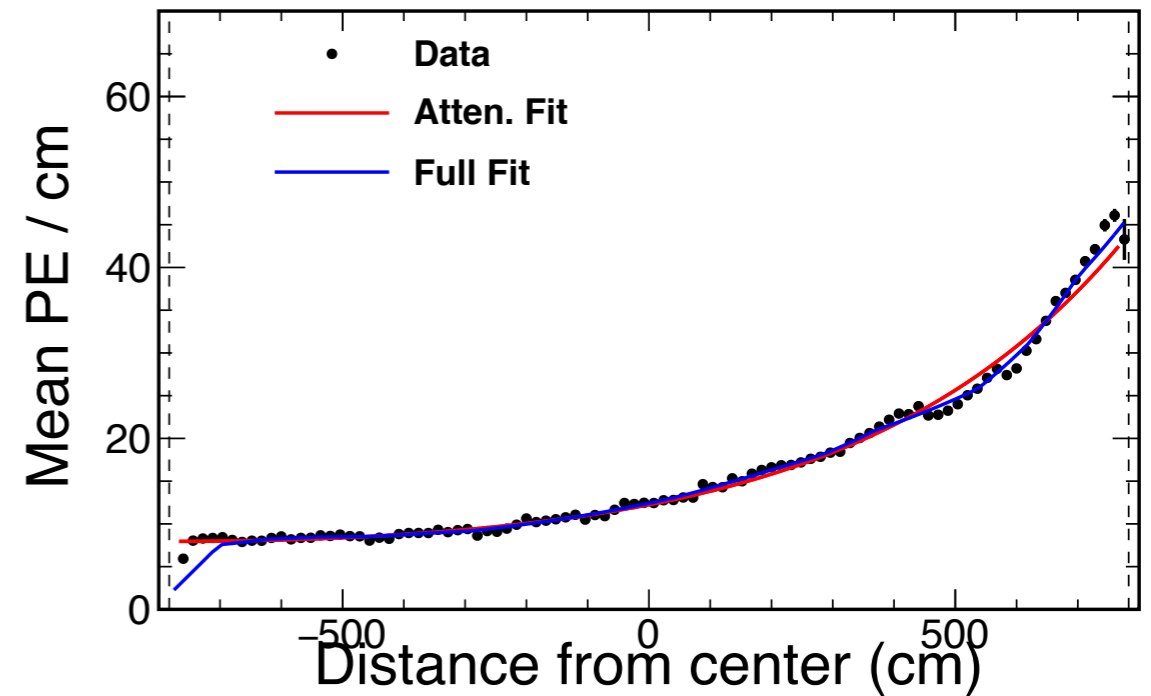
◆ Light Attenuation

* More light absorbed further away from the readout, FD > ND

- ◆ Threshold Bias
- ◆ Detector Shadowing
- ◆ Non-Uniform Reflectivity
- ◆ Twisted Fiber Loops
- ◆ Air bubbles (Y)
- ◆ Fiber sag (Y)

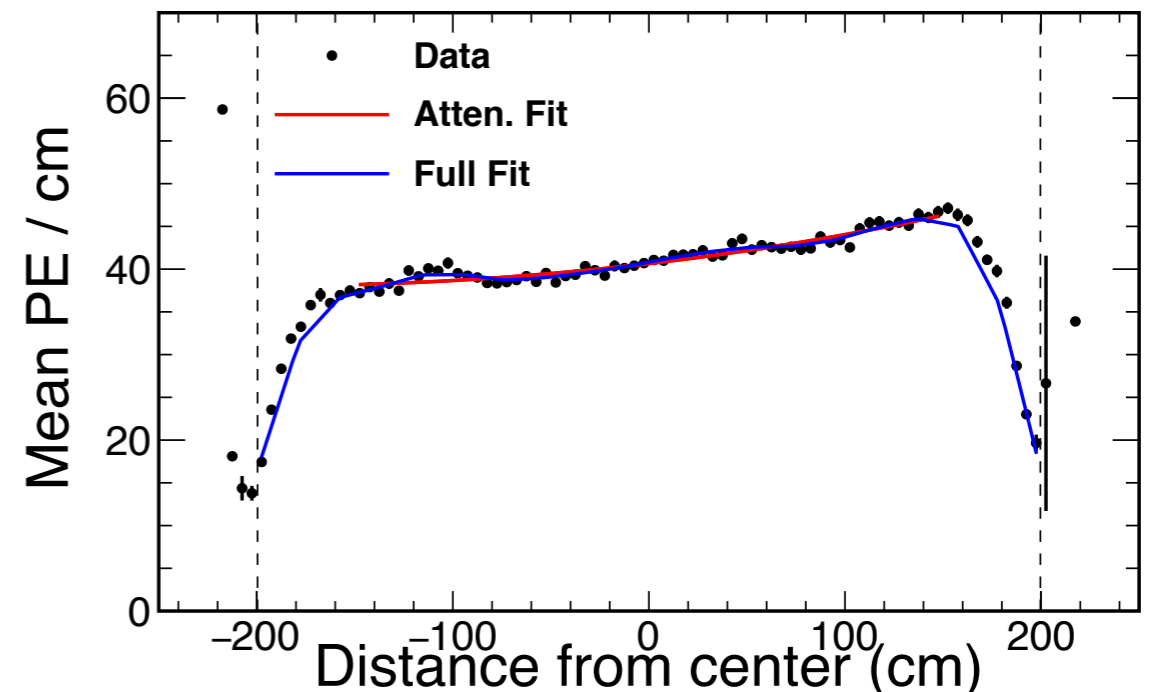
NOvA Preliminary

FD cosmic data - plane 2 (horizontal), cell 376



NOvA Preliminary

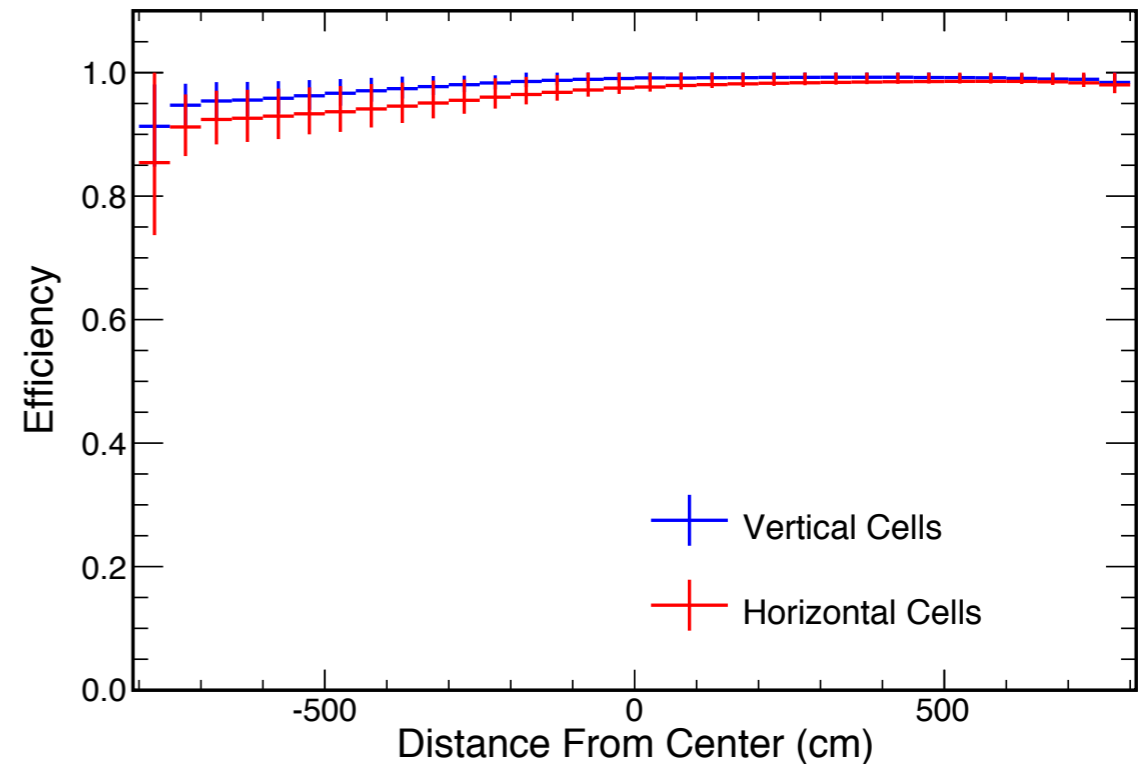
ND cosmic data - plane 48 (horizontal), cell 81



Bias of Deposited Energy

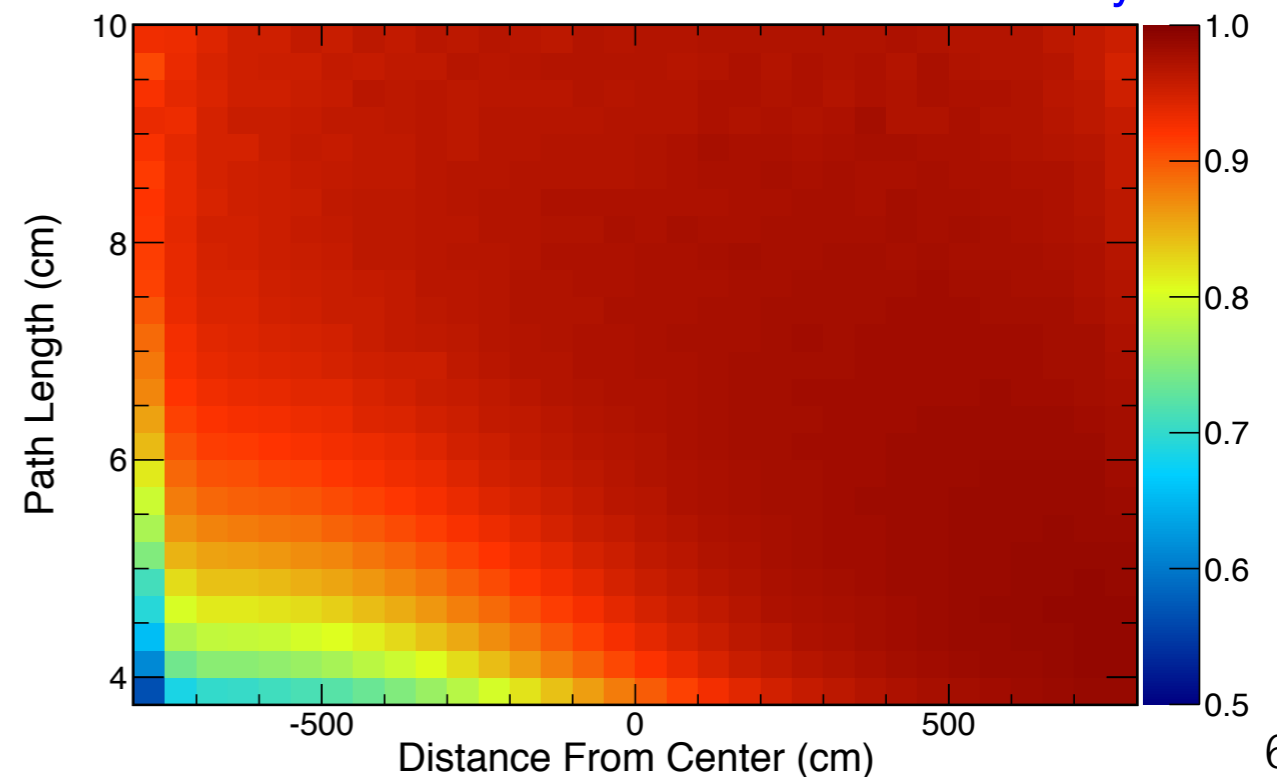
- ◆ Light Attenuation
- ◆ **Threshold Bias**
 - * Readout threshold filters noise
 - * Short hits far from readout can fluctuate below
 - * Hit not seen \rightarrow Energy bias

NOvA Preliminary



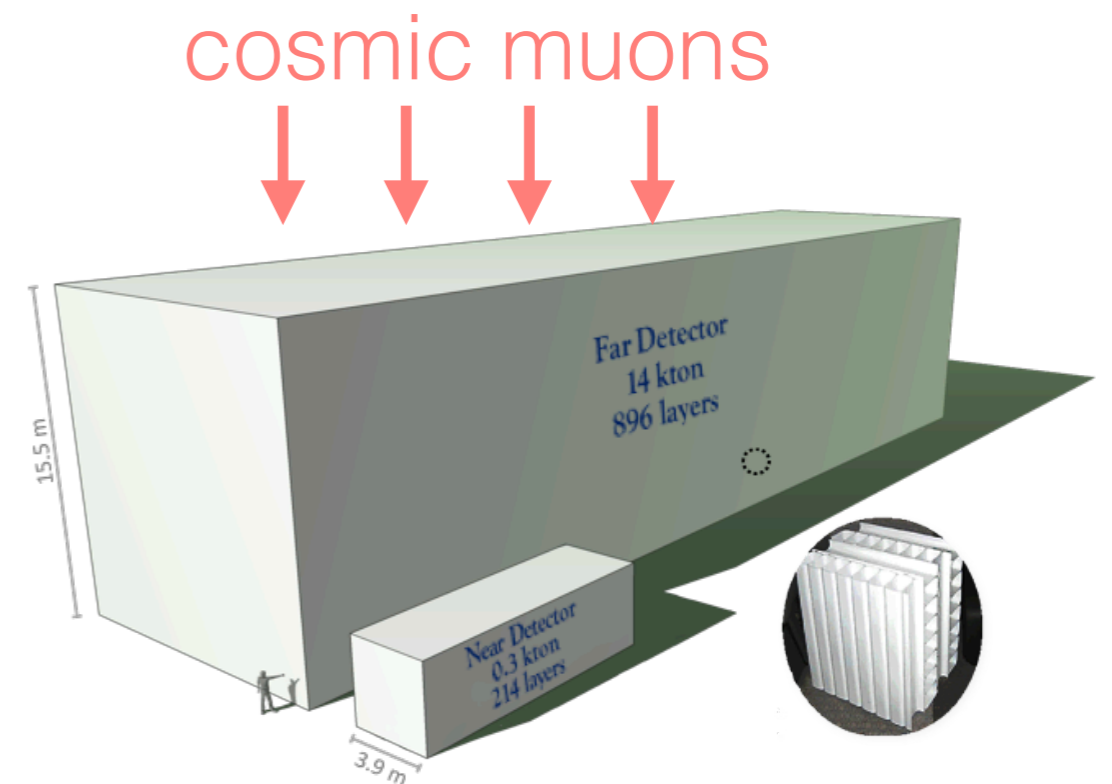
- ◆ Detector Shadowing
- ◆ Non-Uniform Reflectivity
- ◆ Twisted Fiber Loops
- ◆ Air bubbles (Y)
- ◆ Fiber sag (Y)

NOvA Preliminary



Bias of Deposited Energy

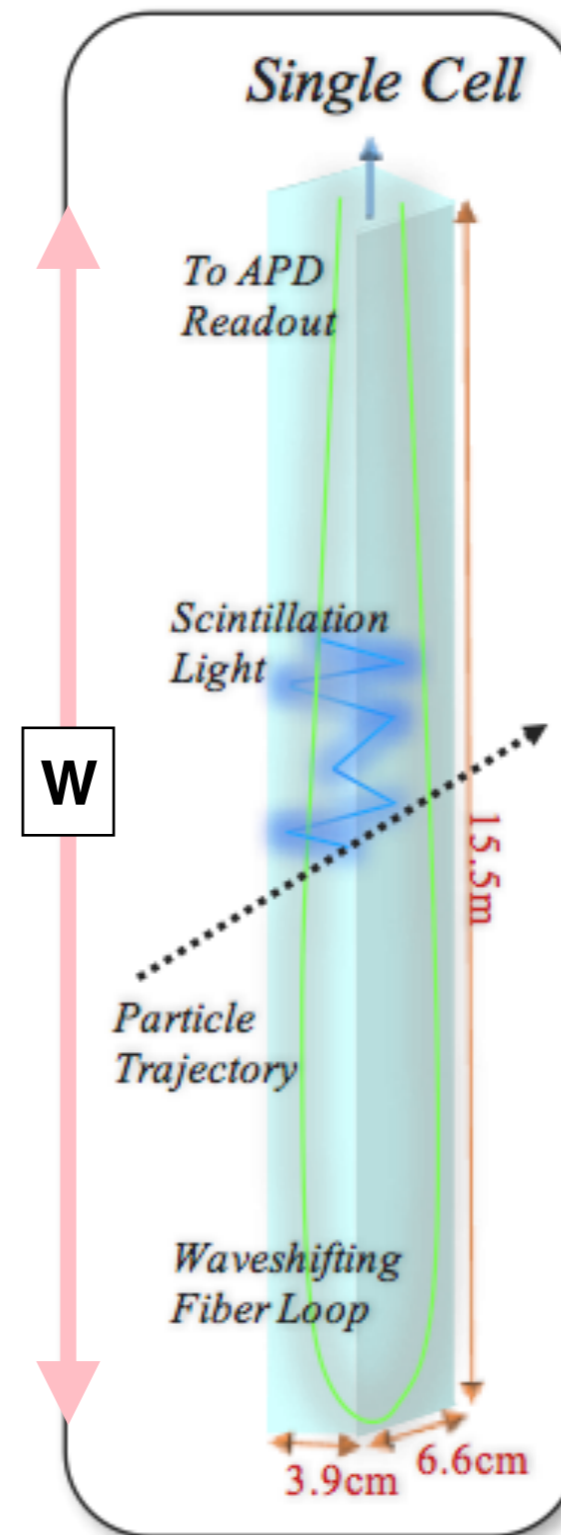
- ◆ Light Attenuation
- ◆ Threshold Bias
- ◆ **Detector Shadowing**
 - * Few-percent effect with big impact
 - * Entangled with threshold bias
- ◆ Non-Uniform Reflectivity
- ◆ Twisted Fiber Loops
- ◆ Air bubbles (Y)
- ◆ Fiber sag (Y)



Muons tend to be higher energy at the bottom, slightly higher dE/dx

Bias of Deposited Energy

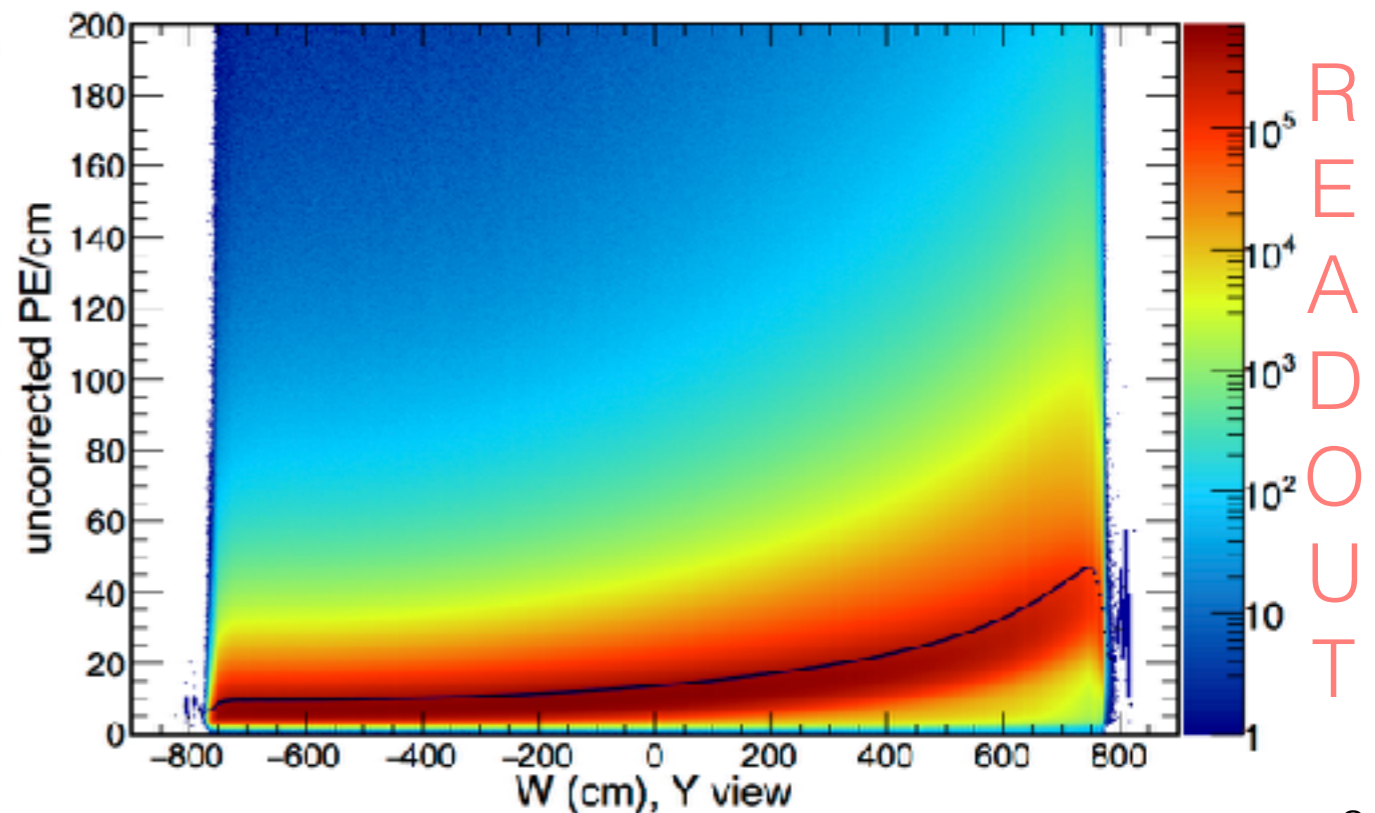
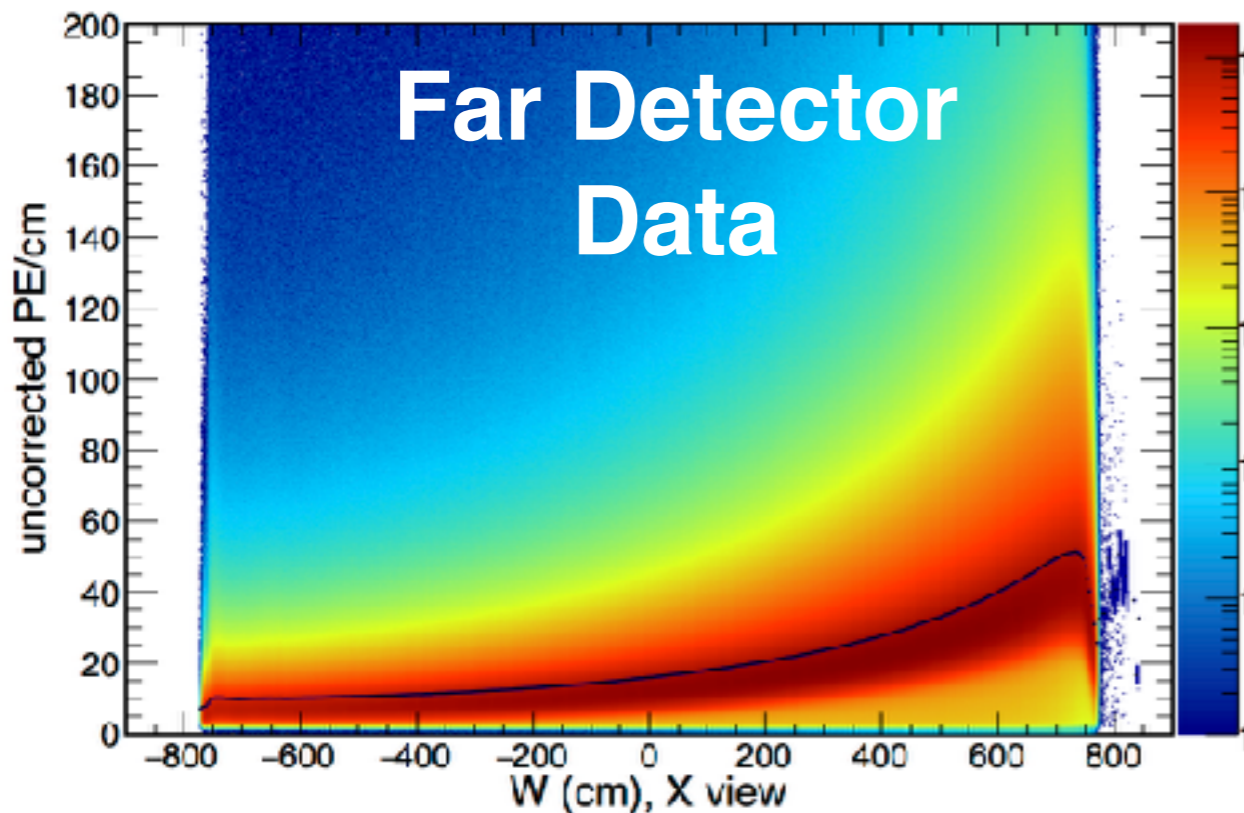
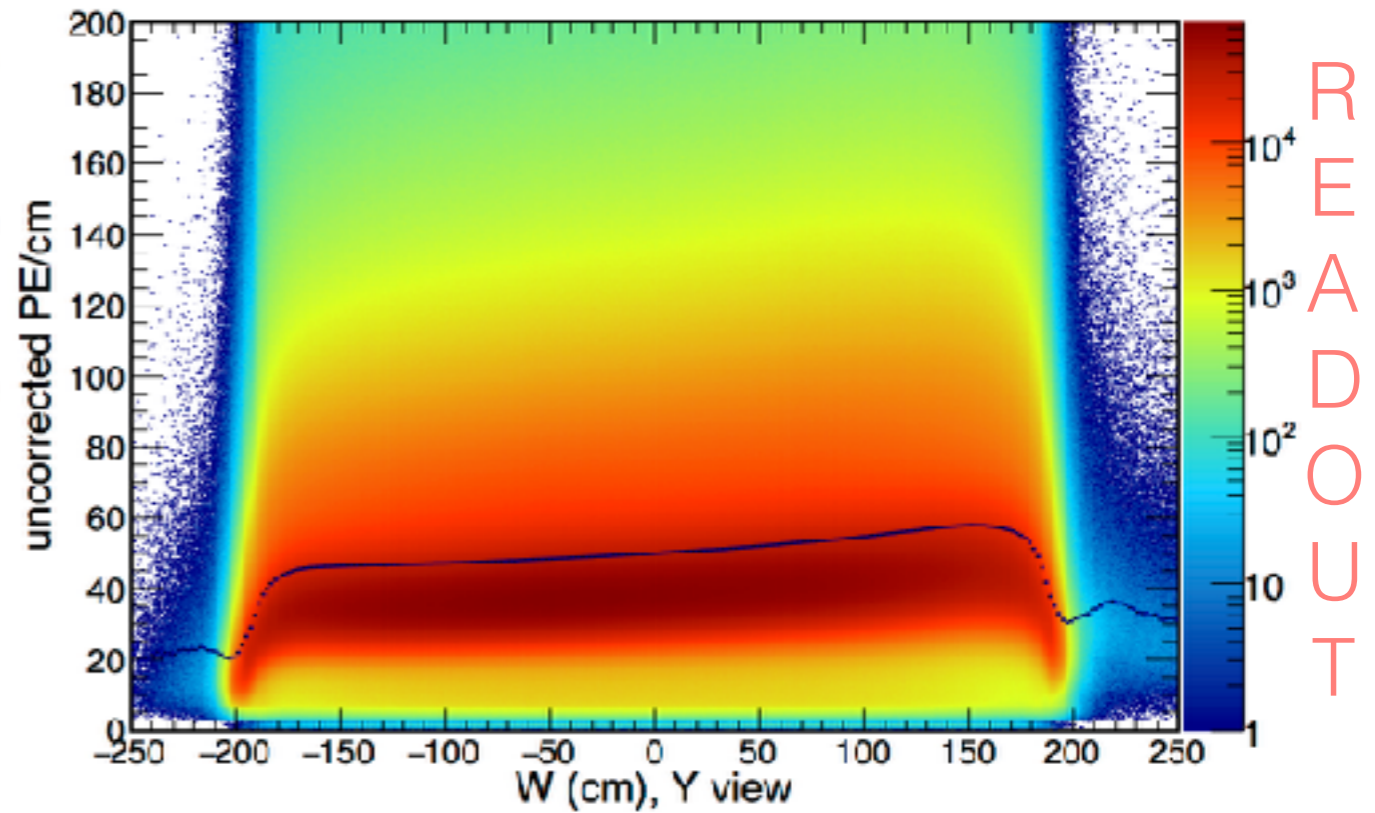
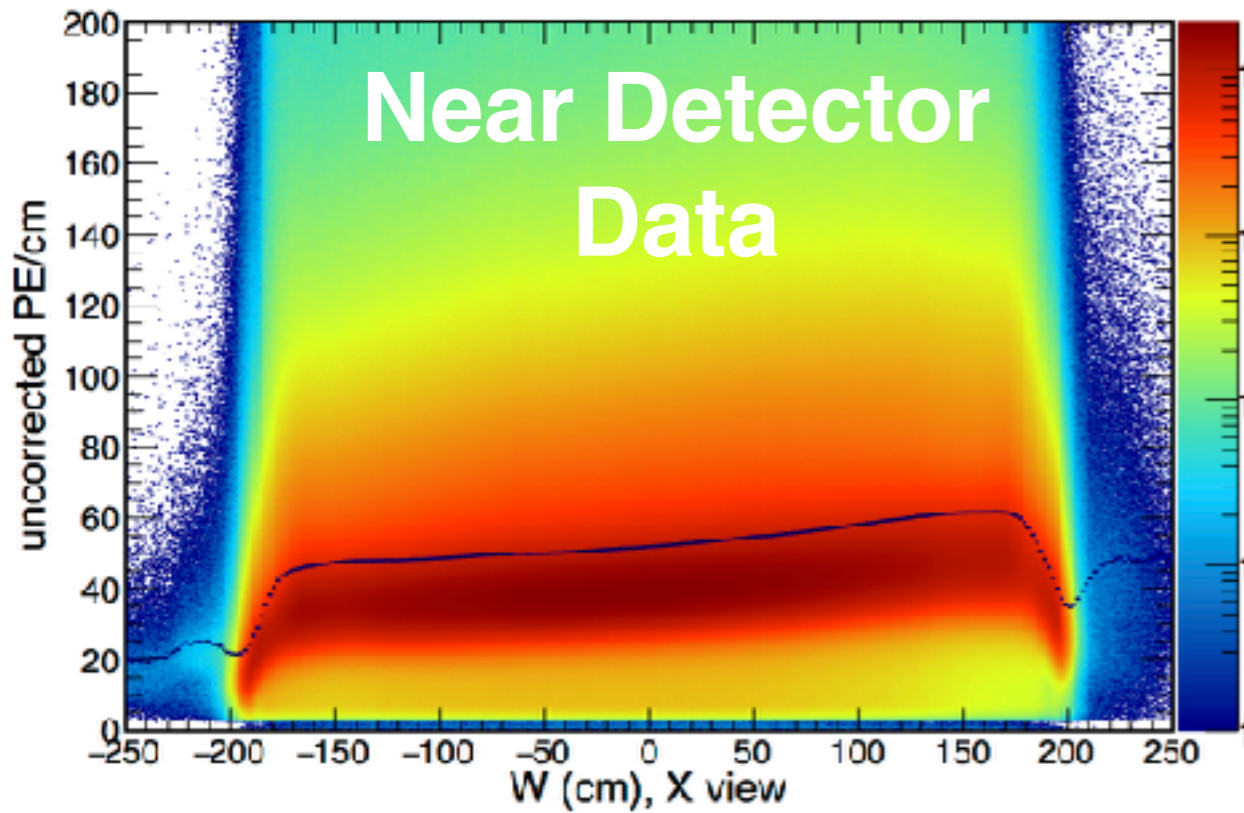
- ◆ Light Attenuation
- ◆ Threshold Bias
- ◆ Detector Shadowing
- ◆ **Non-Uniform Reflectivity**
 - * Brightness of fibers significantly impacts attenuation
- ◆ **Twisted Fiber Loops**
- ◆ **Air bubbles (Y)**
- ◆ **Fiber sag (Y)**
 - * Horizontal bars



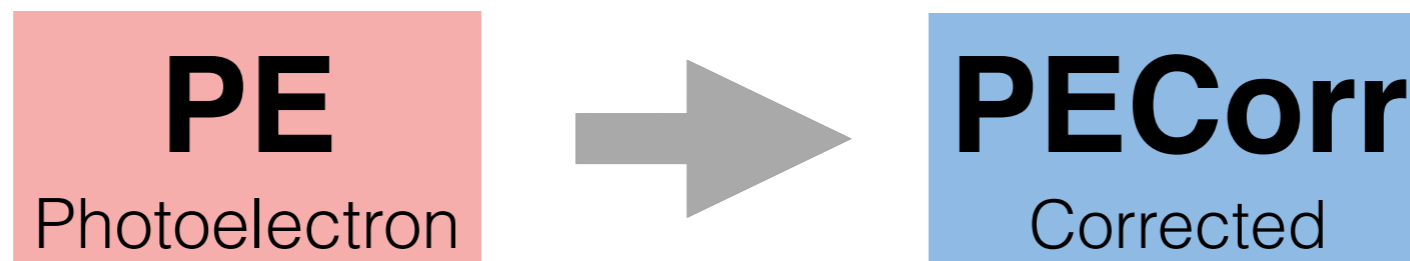
$$e^{-W/\lambda}$$

Attenuation length is different in every cell, due in part to different fiber brightness

Attenuation of Scintillation Light



Calibration Strategy



Attenuated Response $= C + A \left(\exp \left(\frac{W}{X} \right) + \exp \left(-\frac{L + W}{X} \right) \right)$

L: cell length
W: cell pos.
X: Att. Length

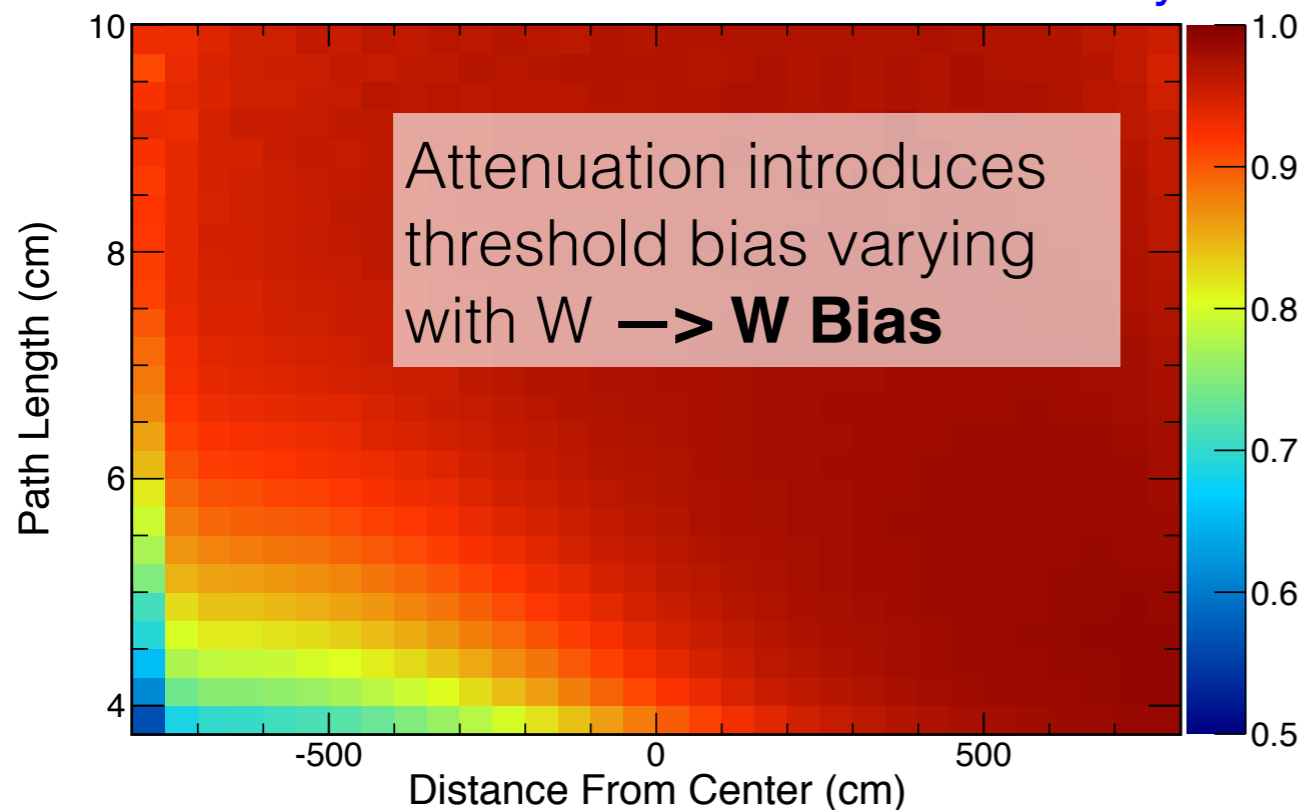
♦ **Fit Attenuation Profiles** (PE/cm vs Position)

- * Separately for every cell
- * Using Mean Response (Eventually use Median)
- * Free Parameters
 - * **A, C** adjust for cell-To-cell Scale/Efficiency differences
 - * Attenuation Length **X** different for each cell
- ♦ But first, need to remove bias from cosmic muon sample
 - * Threshold & Shadowing Correction *from MC*

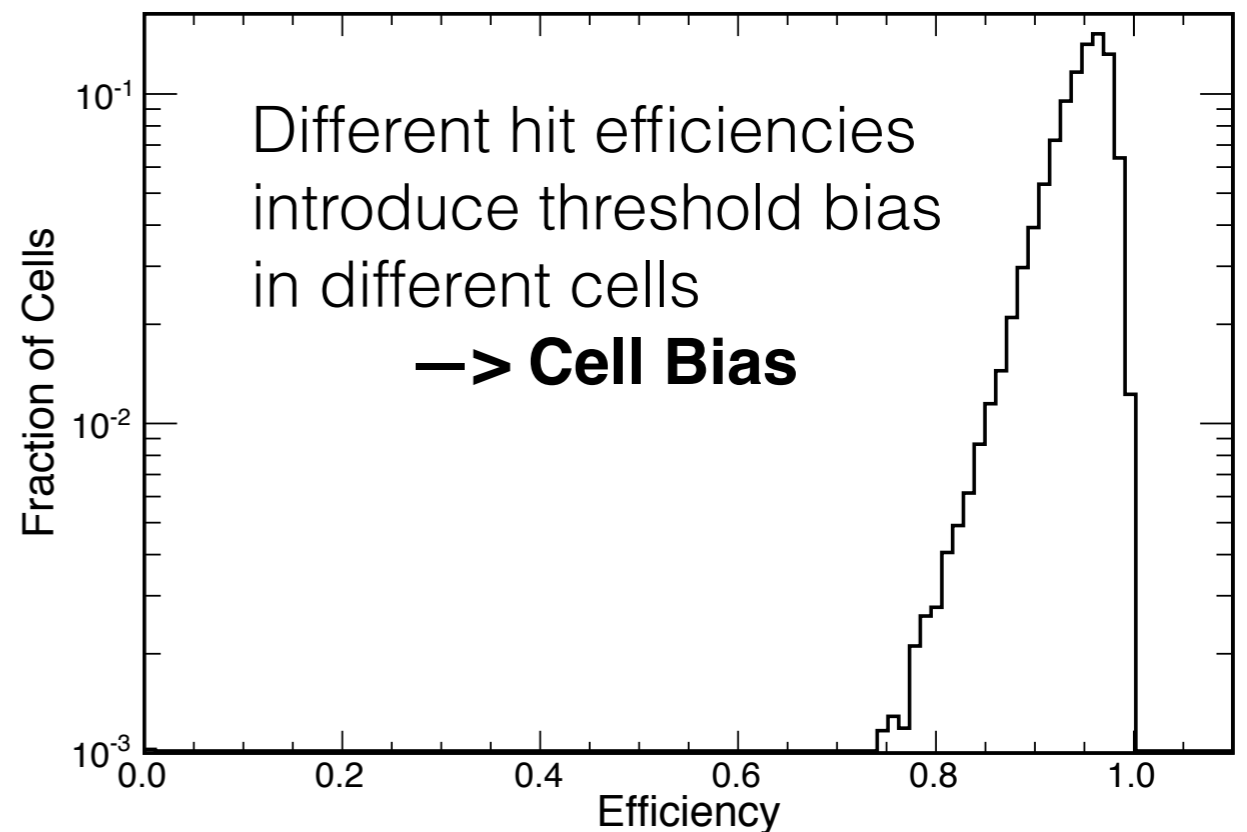
Hit Threshold

For a hit to be seen above threshold, the energy deposited may need to be an upwards fluctuation within the underlying Landau distribution

NOvA Preliminary



NOvA Preliminary



- Short MIP hits are impacted the most — Horizontal Y View
 - * Especially hits which have attenuated
- Different cells throughout the detector affected to a different degree

Combined Threshold & Shielding Correction

$$T = \frac{PE}{\lambda} * \frac{E_{\text{true}}}{E_{\text{MIP}}}$$

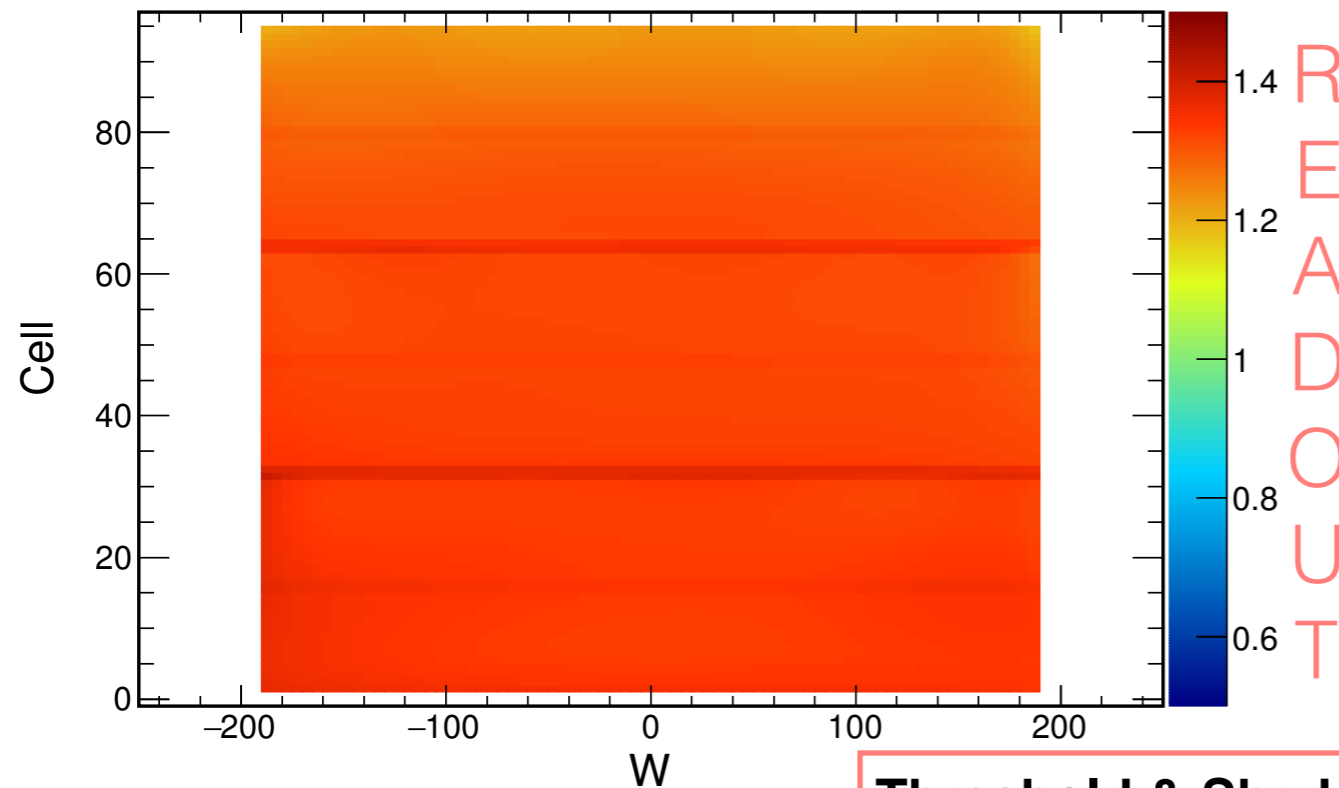
Threshold ——— Shielding

- PE: Simulated photoelectrons at the readout
- λ : number of simulated photons expected at readout in the absence of fluctuations (PE is Poisson distributed)
- E_{true} : True energy deposited in cell
- E_{MIP} : Path length * dE/dx of minimum ionising particle
- Use MC Sample to to construct T , depending on cell and position within,

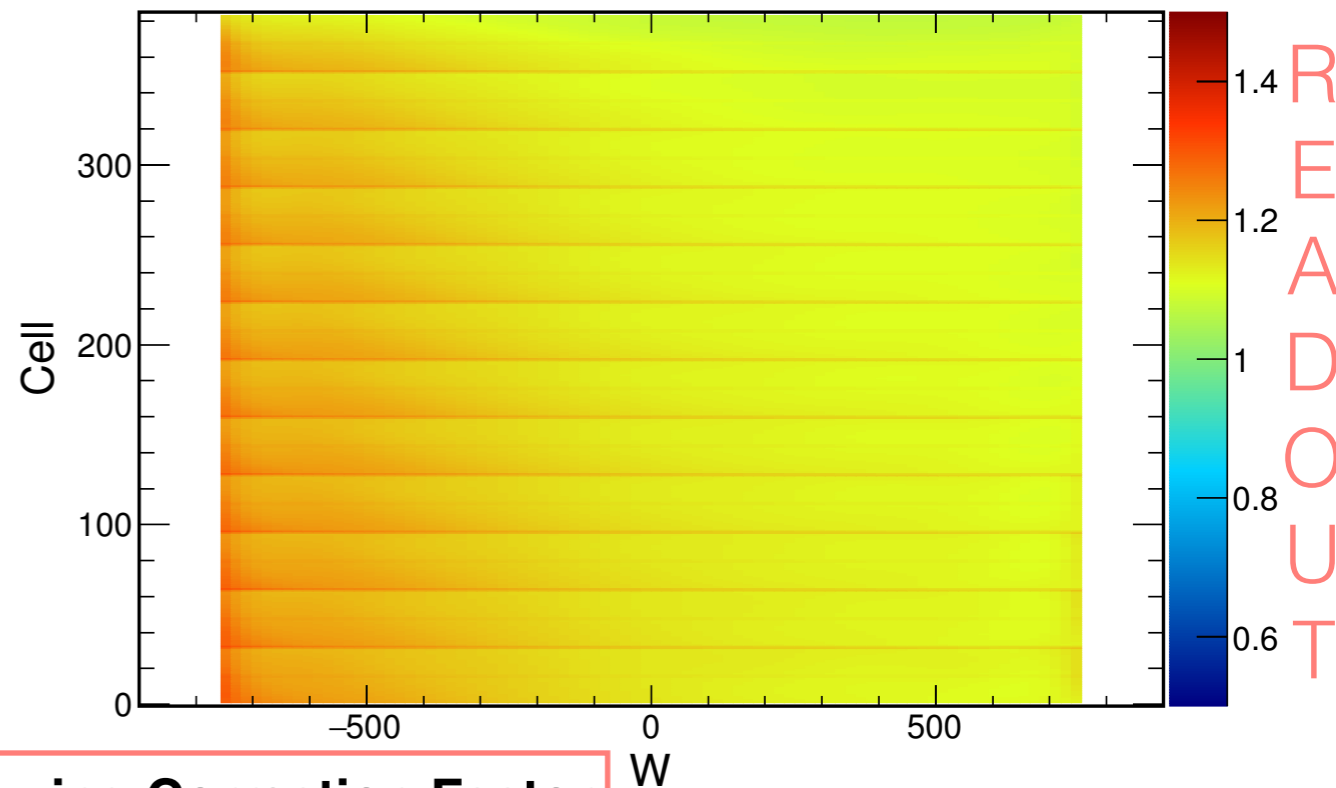
Applied to both MC and Data

Relative Correction Factors

Y View



Y View



Near Detector

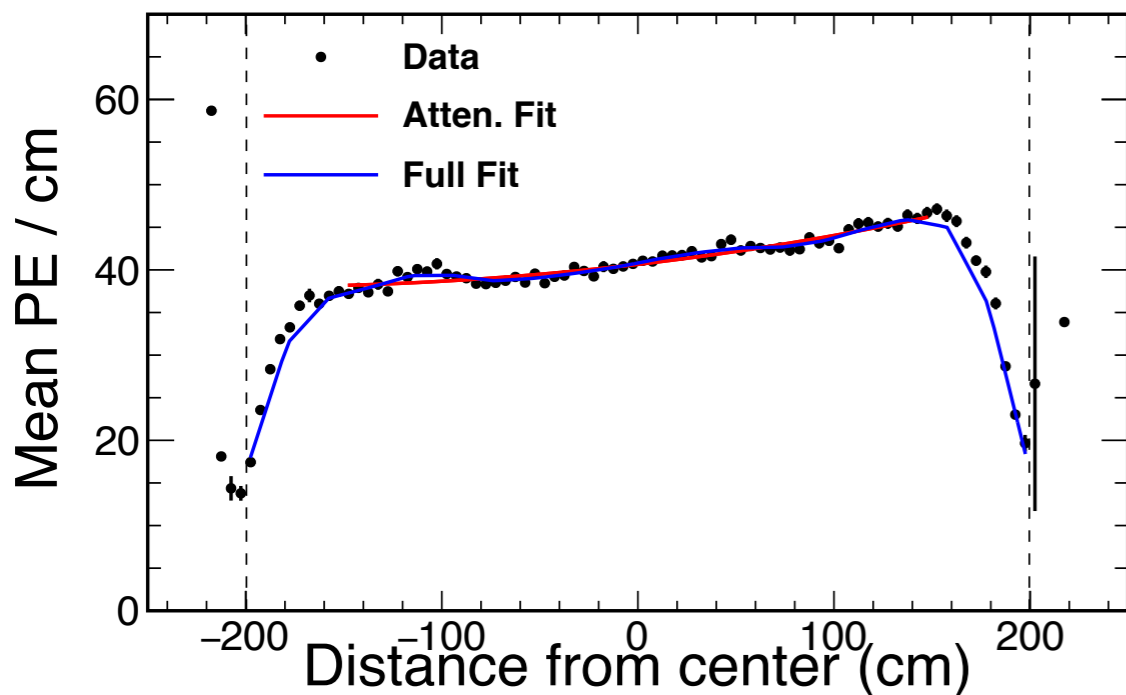
Far Detector

Threshold & Shadowing Correction Factor

Attenuation Fit, each cell

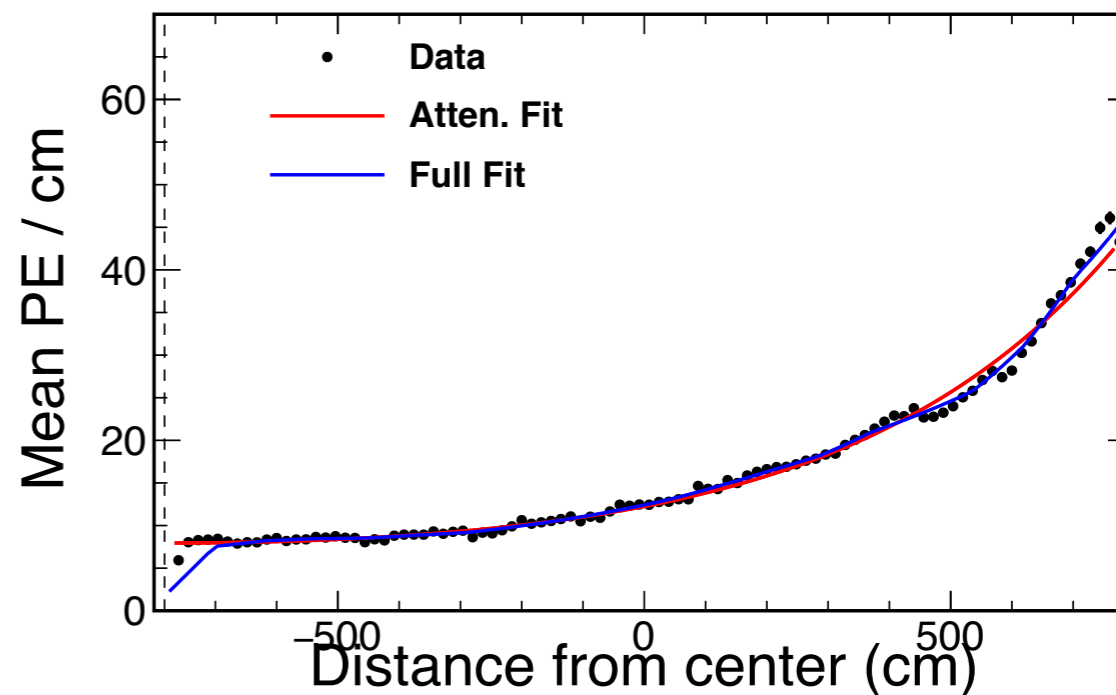
NOvA Preliminary

ND cosmic data - plane 48 (horizontal), cell 81



NOvA Preliminary

FD cosmic data - plane 2 (horizontal), cell 376

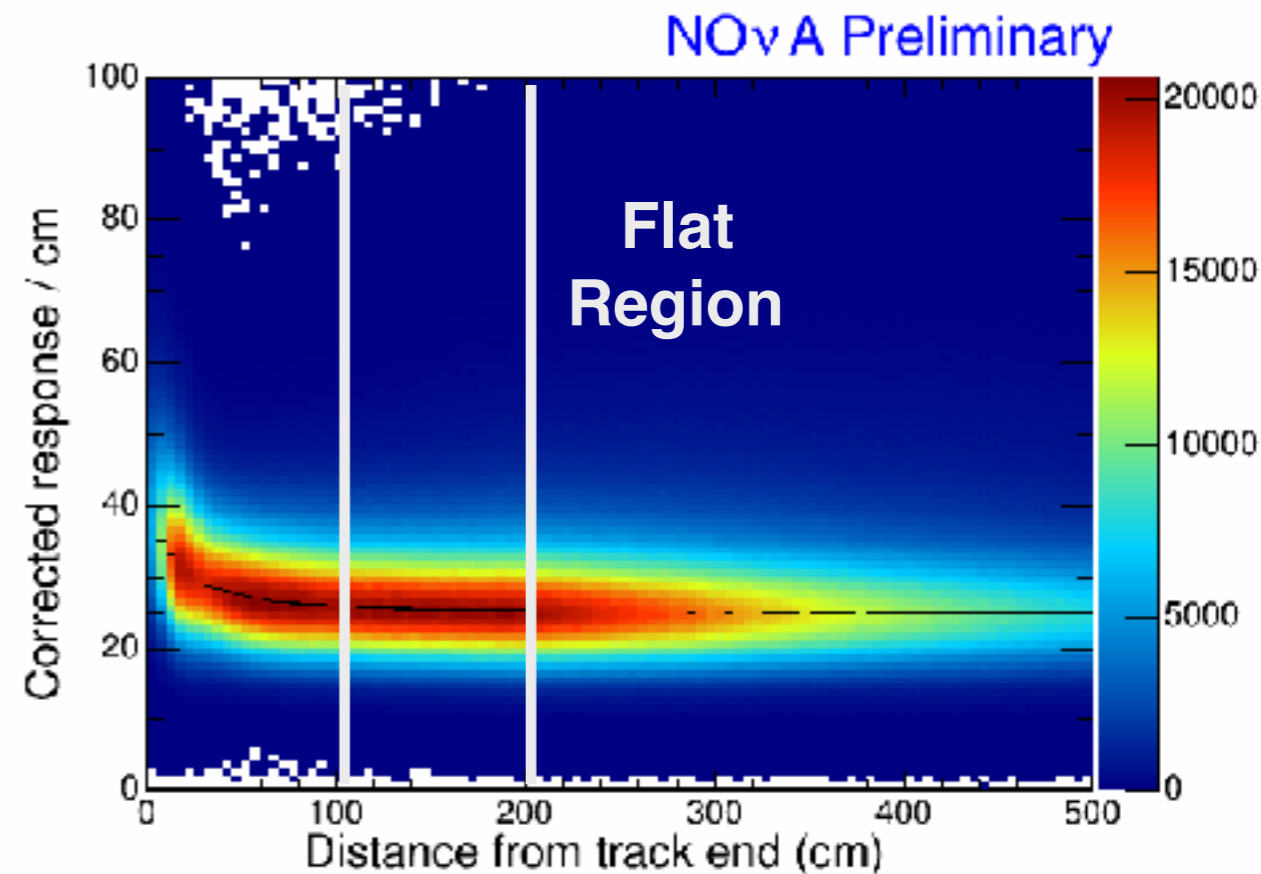


Absolute Calibration

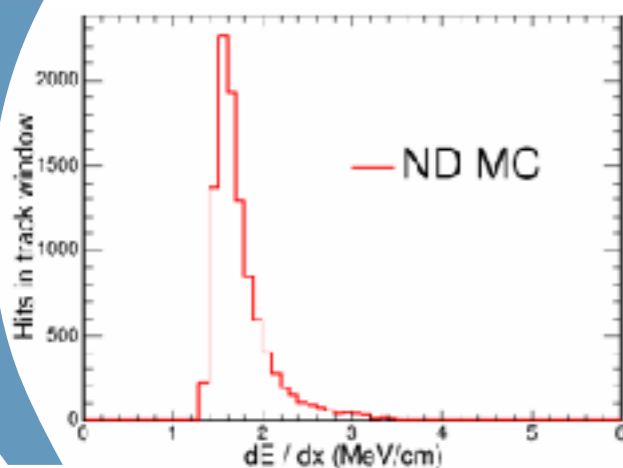
Energy Scale

$$PECCorr_{hit} * \left(\frac{MeV/cm}{PECCorr/cm} \right) = MeV_{hit}$$

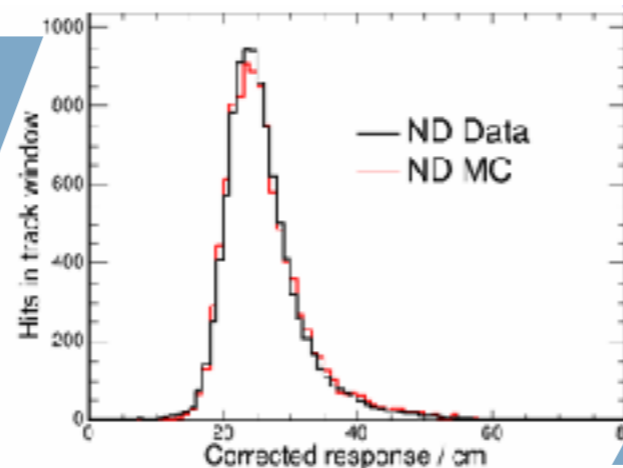
- Select **Stopping Muons**
- Tighten dE/dx peaks by selecting hits in the Bethe-Bloch flat region
- MC: True dE/dx and Response
- Data: Response (PECCorr/cm)



MeV/cm



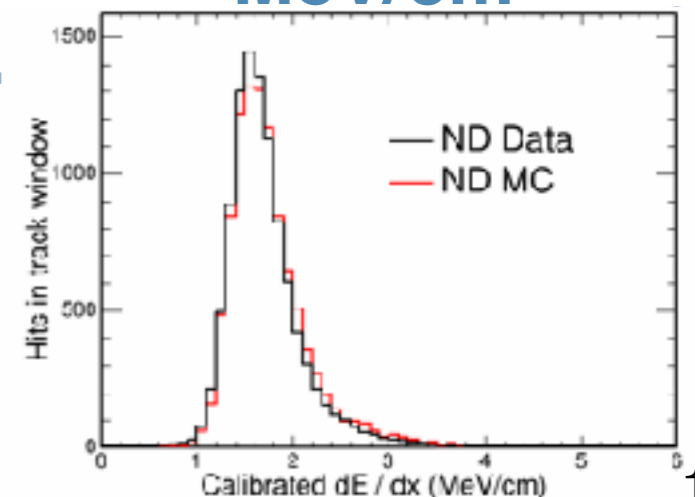
PECCorr/cm



* PECCorr

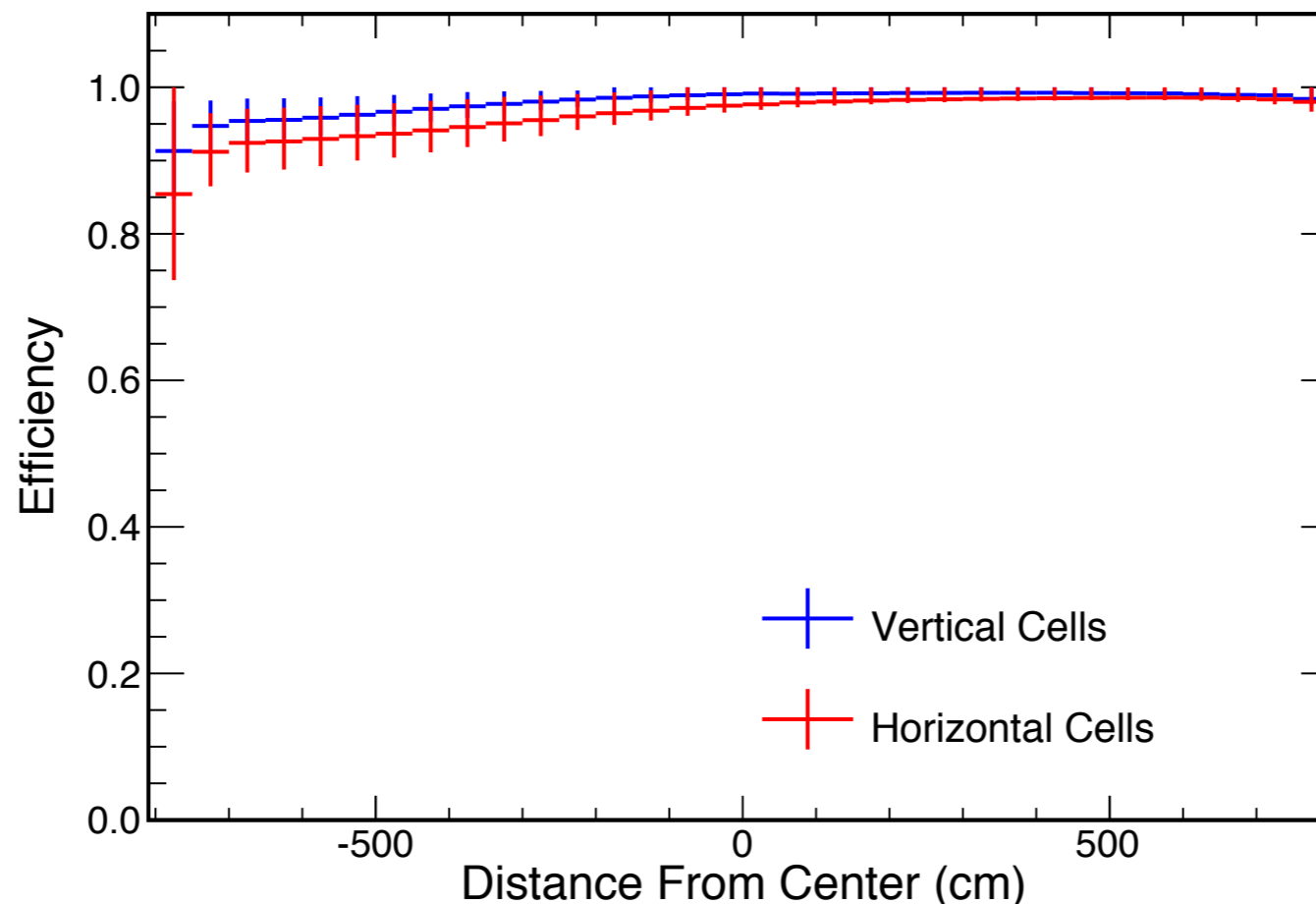
=

MeV/cm



Absolute Calibration

NO ν A Preliminary

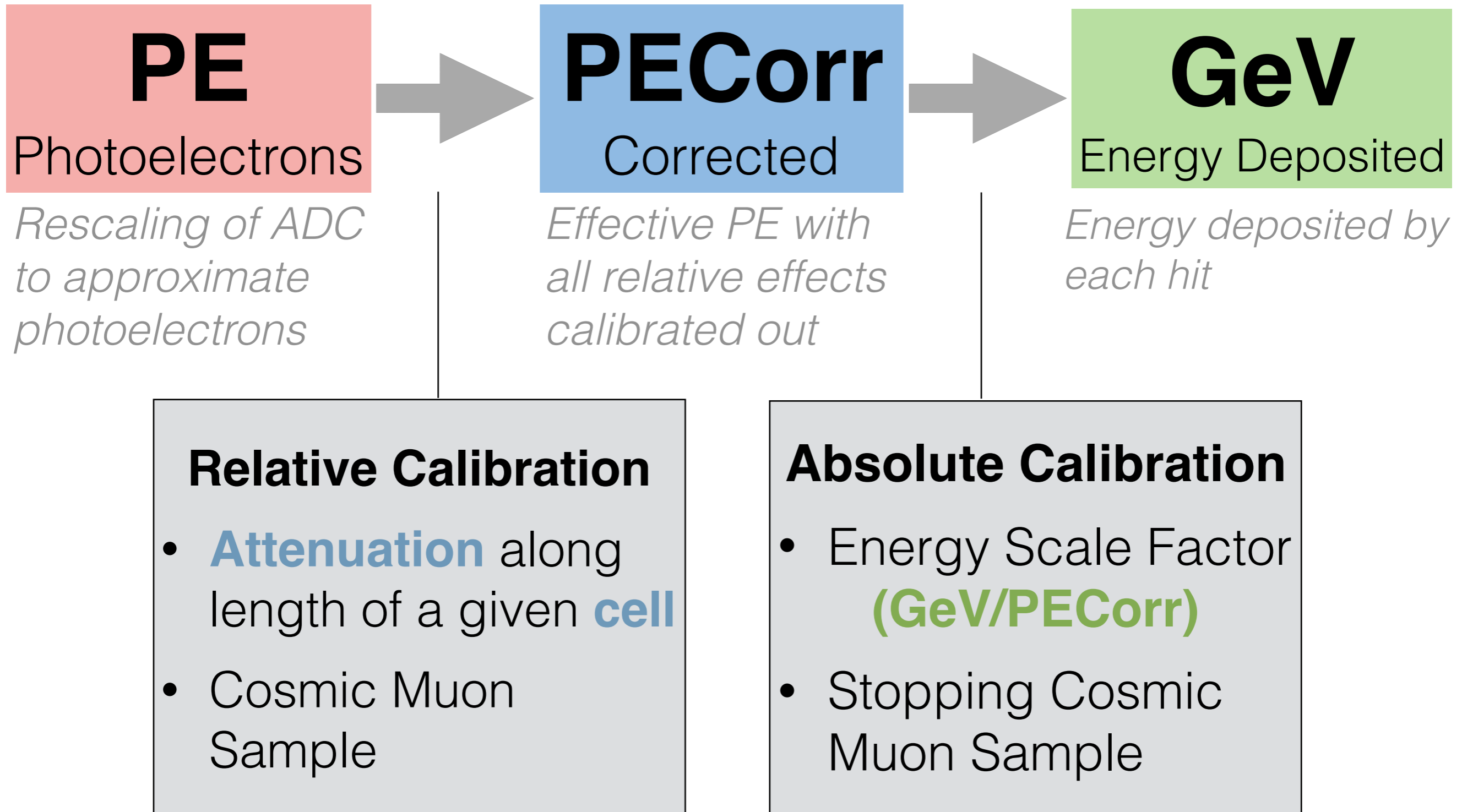


Far Detector
Hit Efficiency

- Threshold means selected hits were more likely to be up-fluctuations in *true* energy deposition
 - Select only flat region of *W* so Energy scale is not biased
- Horizontal Y View has many more hits, calibrate separately and average

Energy Calibration

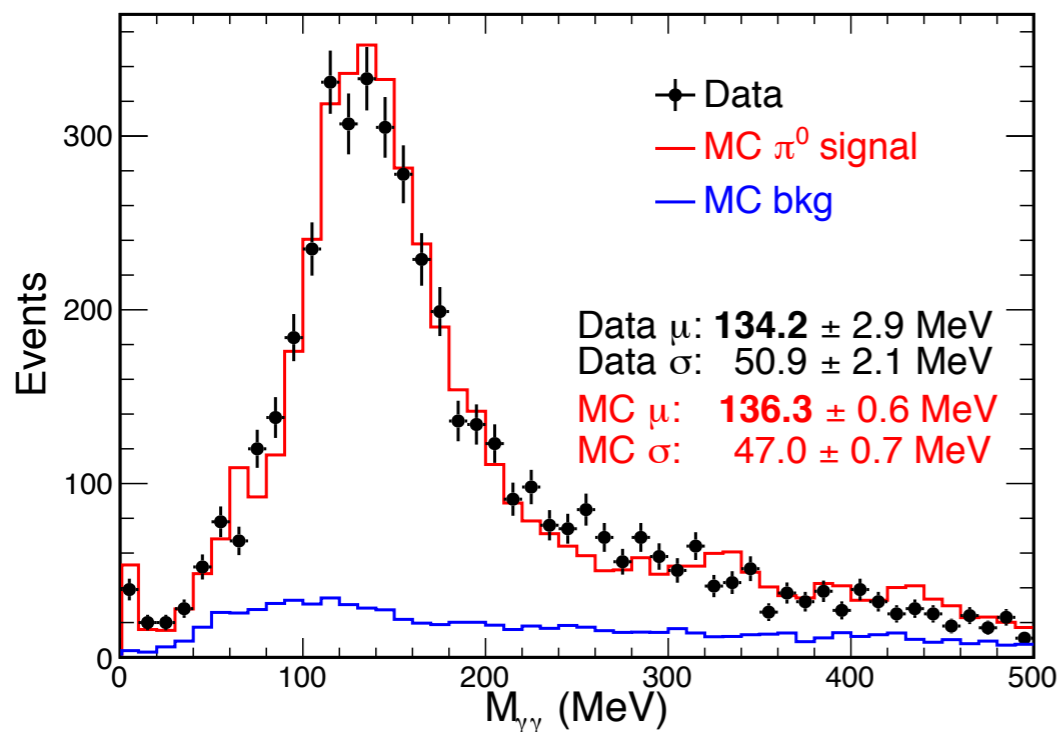
Cosmic Muon Data and MC



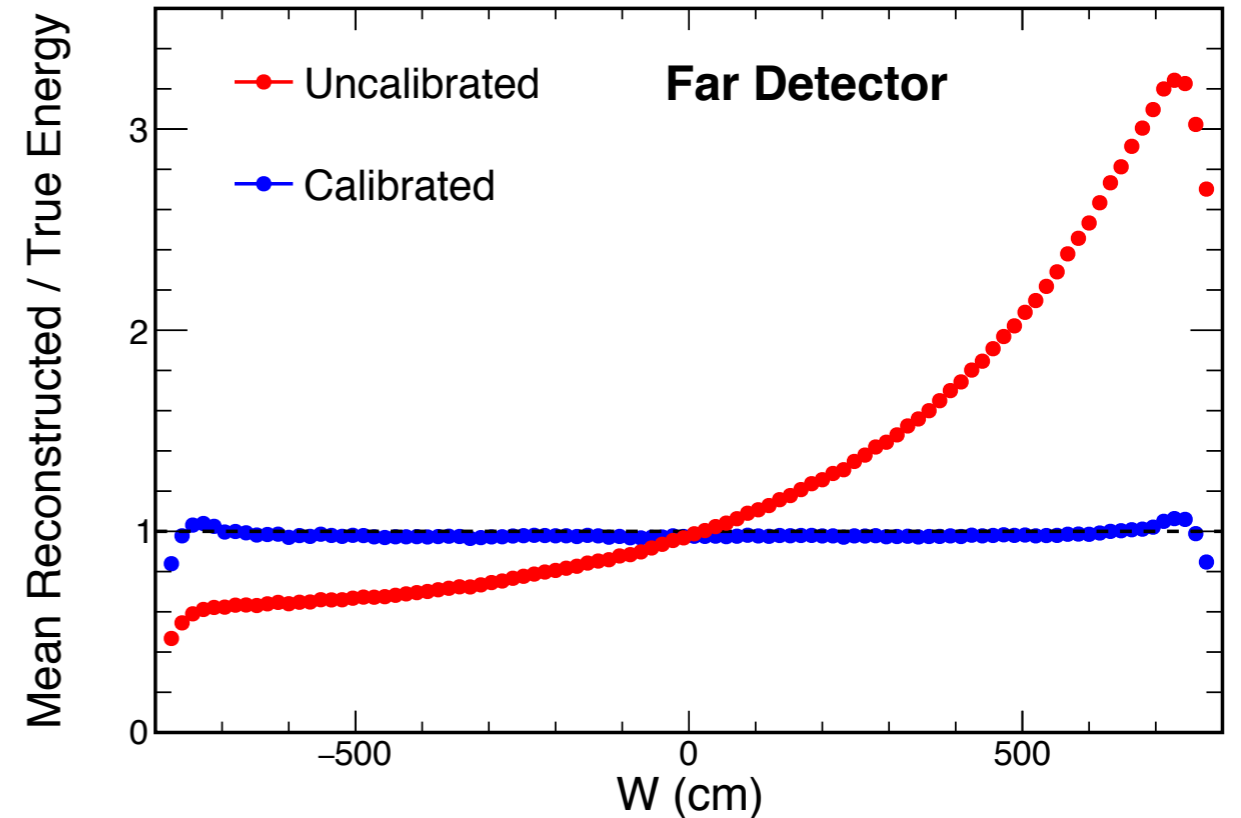
Verification with MC

- Profile Ratio of Reconstructed over True Energy
 - * **Energy Scale:** average vertical deviation from 1
 - * **Relative Calibration:** shape along W, cell, plane
- Pi-Zero Mass Peak
- Muon/Proton dE/dx
- Michel Electron Spectrum

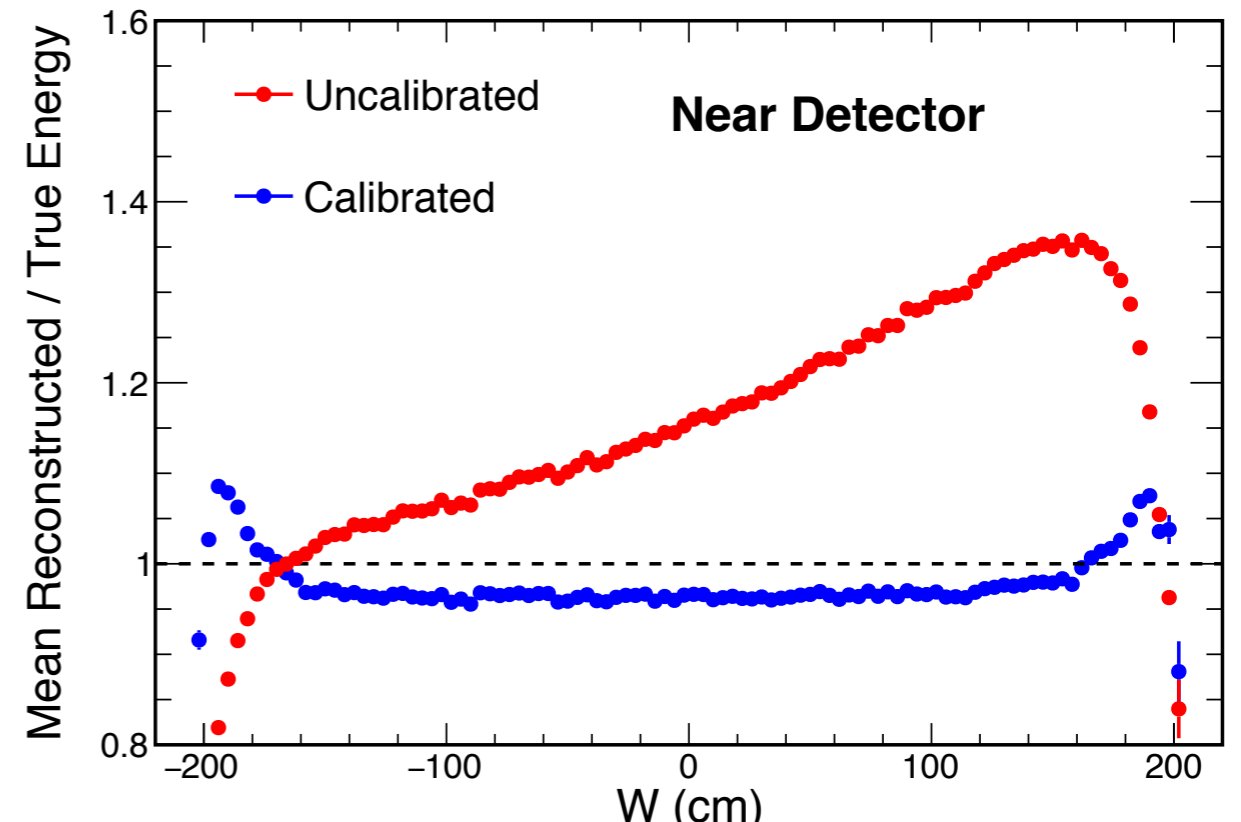
NOvA Preliminary



Y view, true deposits > 15MeV NOvA Simulation



Y view, true deposits > 15MeV Simulation



Takeaway Points

- **Full chain of correction factors** to take measured PE to a best estimate of energy deposited, GeV
- **Scintillation Light attenuates** to different degrees in each cell
 - * Fit a double-exponential form to PE/cm response for each cell: attenuation correction factor
- **Remove bias from the cosmic muon sample** before calibrating
 - * Threshold and Shielding Effects
- Do not let **threshold effects bias the absolute energy scale**
 - * $\langle \text{MeV} / \text{cm} \rangle / \langle \text{PECorr} / \text{cm} \rangle$
- Ongoing Work
 - * Threshold & Correction Factor with data instead of MC
 - * Fit to Response (PE/cm) Median
 - * Rigorous understanding of systematic uncertainty