

# Optimisation of the SHiP muon shield

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*“We know there is new physics,…”*

Dark matter, baryon asymmetry and neutrino masses are direct experimental evidence that we're missing something.

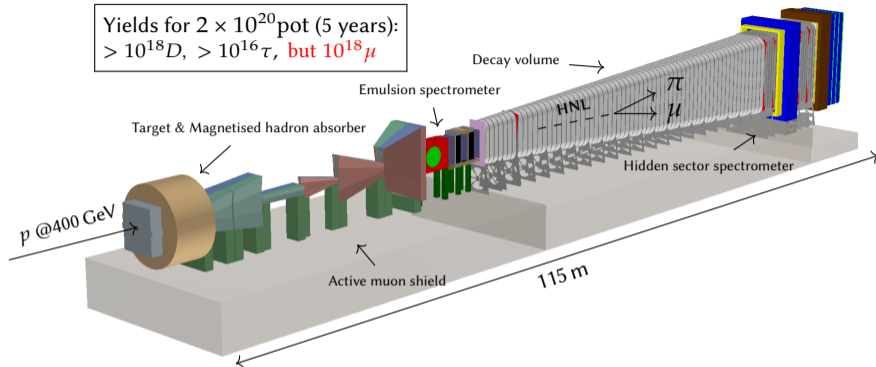
*“... We don't know where it is...”*

We do not know which energy scale to target: Very weakly coupled new physics could be hiding in plain sight — at energies already accessible!

*“... We need to be as broad as possible in our exploratory approach”*

— *Fabiola Gianotti*

# Overview of the Search for Hidden Particles



## Two signatures:

1. Via decay to visible particles in hidden sector spectrometer
  2. Via scattering in nuclear emulsion
- } Generic signatures predicted by many new physics models

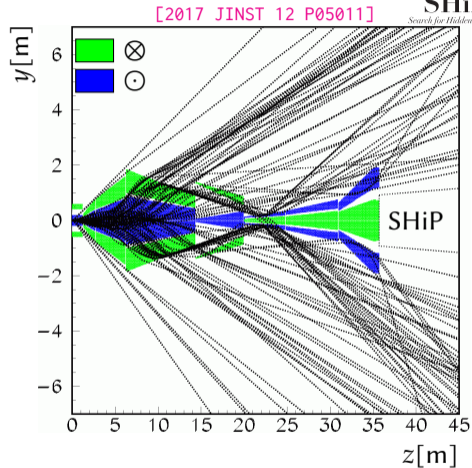
**Zero Background** crucial to study hidden sector decays

## Crucial challenge: Zero background

- › Passive hadron absorber
- › Active muon shield that has to reduce muon flux by at least 6 orders of magnitude
- › kinematic range of muons up to  $p \sim 350$  GeV
- › kinematic range of muons up to  $p_T \sim 8$  GeV

*The muon shield is the critical component to optimise to maximise the experimental acceptance*

- › A measurement of the muon spectrum for the SHiP target at the H4 test-beam at CERN's SPS is planned for this summer
  - › Obtain  $10^{11}$  protons on target, c.f.  $10^{10}$  currently available in simulation



# Goals & Challenges of the muon shield optimisation



**Goal:** Optimisation using *full simulation* with FairShip framework for every evaluation to *optimise performance vs. cost* and *provide robustness by optimising for a lower field strength*.

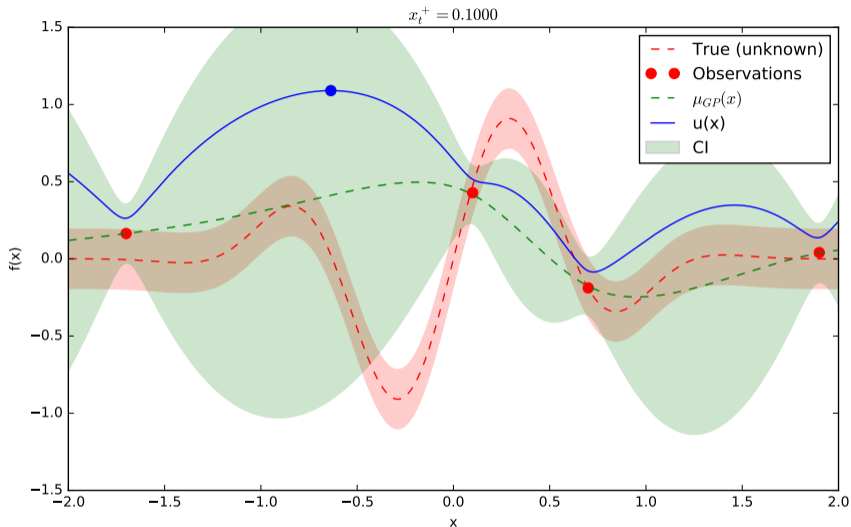
## Challenges

- › Doubly statistically limited
  - › Not enough simulation
  - › Not enough computing power to use entire simulation for optimisation
- › Underlying physics inherently stochastic
  - › Nearly identical configurations may have very different performance
  - › With a different random seed entirely different muons pass the shield

→ *Evaluation of points very expensive, gradient information not available and can not be approximated*

- › Even with a simple parametrisation we have ~50 free parameters (lengths), each varying from cm to m

# Introduction to Bayesian Optimisation using a 1D example\*



\*Based on [scikit-optimize documentation](#)

## Not quite as simple as this example:

- › Bayesian optimisation does not scale well for high-dimensional problems.
- › Computing model imposes additional constraints.
  - › 1600 cores available at YANDEX<sup>†</sup>
  - › Make up to 100 guesses at once (with 16 nodes parallelising every function evaluation)
- › Use scikit-optimize implementation of Bayesian optimisation DOI [10.5281/zenodo.1170575](https://doi.org/10.5281/zenodo.1170575).
- › Use Gaussian processes and random forests as surrogate models.
- › Reduce muon sample by factor  $\sim 40$  to speed up evaluation and even out coverage of phase space:
  - › Currently:
    1. study the importance of different regions of the phase-space
    2. reduce and re-weight manually
  - › Evaluating importance sampling and other options

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<sup>†</sup>Russian internet company which contributes to LHCb, COMET, CMS and SHiP with its machine learning expertise and computing power

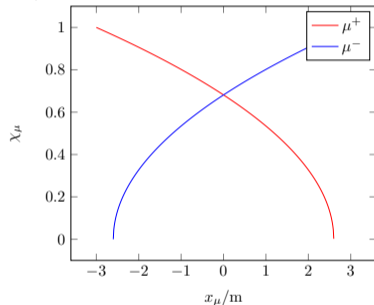
$$f(W, \chi_\mu) = \begin{cases} 10^8 & \text{if } W > 3 \text{ kt} \\ (1 + \exp(10 \times (W - W_0)/W_0)) \times (1 + \sum_\mu \chi_\mu(x_\mu)) & \text{otherwise,} \end{cases}$$

where:

- $W$  weight of the muon shield
- $W_0$  weight of the baseline
- $\chi_\mu$  weighted position of muon  $\mu$  passing a sensitive plane at position  $x_\mu$ .

*Note:*

- › Penalise muons entering the acceptance
- › Length optimised implicitly via the weight
- › Weight cut-off as regularisation



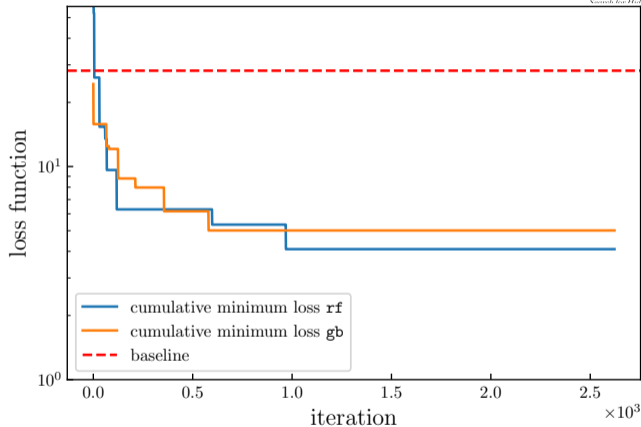
**Figure 1:**  $\chi_\mu(x_\mu)$

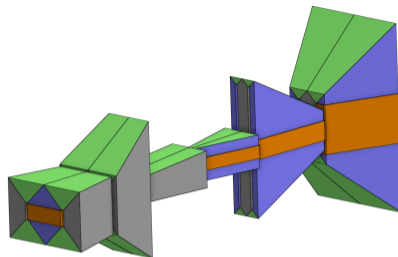
**Loss function continues to evolve with technological constraints and background studies.**



# Optimisation convergence

- › Cumulative loss: exploring points with high uncertainty part of algorithm, only cumulative loss is meaningful
- › Two optimisers shown here: still evaluating different regression algorithms to determine which performs best
- › Performance here is on the reduced muon sample: perform follow-up studies on the full dataset to confirm performance





- › Significant reduction in weight ( $\rightarrow$ cost)
- › Same performance with significantly reduced magnetic field

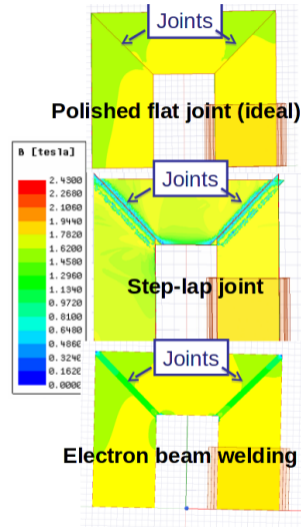
Configuration	length/m	weight/kt	reduced sample	full sample
baseline @1.8 T	34.60	1.72	27 $\pm$ 5	70 $\pm$ 15
new optimum @1.7 T	34.82	1.28	22 $\pm$ 3	42 $\pm$ 6

## Grain oriented steel

- › Allows to achieve fields of up to 1.8 T with warm magnets
- › Manufacturing of SHiP will push the limits of the technology:
  - › Scale of muon shield exceptional
  - › Several techniques need to be evaluated for the joints of the magnets

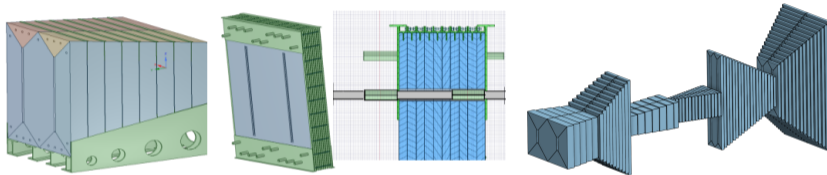
## *Optimise technology as well as geometry*

- › Several prototypes will be produced this year, and the most promising will be tested with beams at CERN → Part of the CERN/Imperial team testing the technology



## Conclusion and further work

- › Found new configuration for comprehensive design study.
- › Have an algorithm that works and can be used as base for further improvements.
- › Optimisation infrastructure is now also used for optimisation of other subsystems.



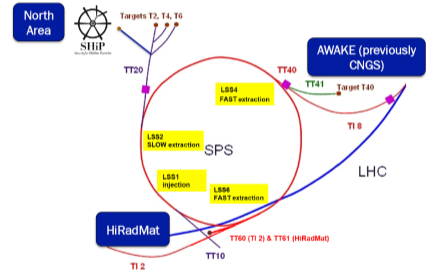
## Future work

- › Fully automate process, add additional constraints to loss function and improve the shield further!
- › Collaboration with engineers at MISIS to progress to a detailed engineering design and prototypes.

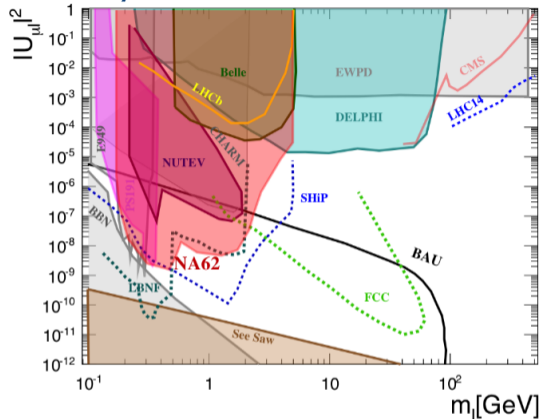
# Backup

## Maximise intensity and mass reach

- › Intense proton beam from the SPS @400 GeV at the new beam dump facility (BDF) in the North Area
- › Very dense target of  $12 \times \lambda_{\text{int}}$ 
  - › abundant production of heavy flavour
  - › reduced neutrino production from  $\pi$  and  $K$  decays
- › Number of protons per cycle similar to CNGS, but slow instead of fast extraction
- › Operation in parallel with LHC, other beam-lines at the SPS



## Sensitivity: HNL



**Figure 2:** HNL sensitivity at SHiP for  $\nu_{\text{MSM}}$  with  $U_e^2 : U_\mu^2 : U_\tau^2 = 1 : 16 : 3.8$  and a normal neutrino mass hierarchy.

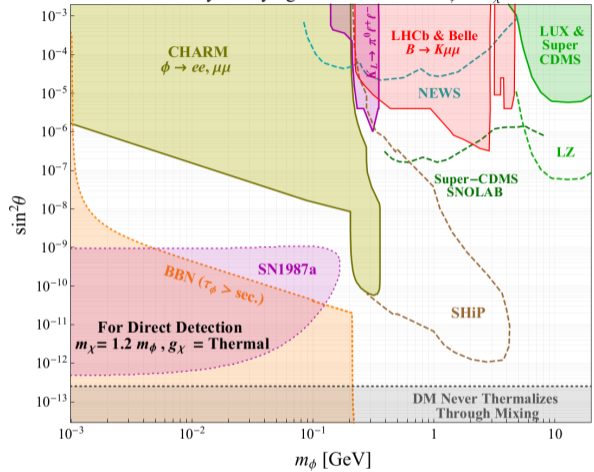
- › Best sensitivity up to charm kinematic limit
- › Significant contribution from  $B$ -decays

Theoretical limits from:

- › Baryon asymmetry of the universe (BAU)
- › Big bang nucleosynthesis (BBN)
- › Model-independent limit for any Seesaw model

# Sensitivity: Dark Scalars

Visibly Decaying Scalar Mediator  $m_\phi \approx m_\chi$



**Figure 3:** Dark scalar sensitivity at SHiP.

- › For short lifetimes  $B$ -factories and LHCb best
- › SHiP covers unique parameter space complementing other experiments
- › Large contribution from  $B$ -decays at SHiP
- › “Hole” at  $c\tau \sim O(m)$ , where lifetime is too short for SHiP and too long for  $B$ -experiments

*NB: Before re-optimisation*



# Sensitivity: Dark Photons

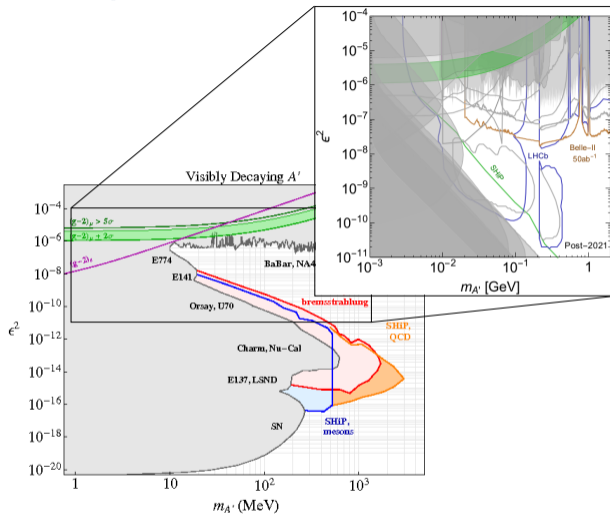
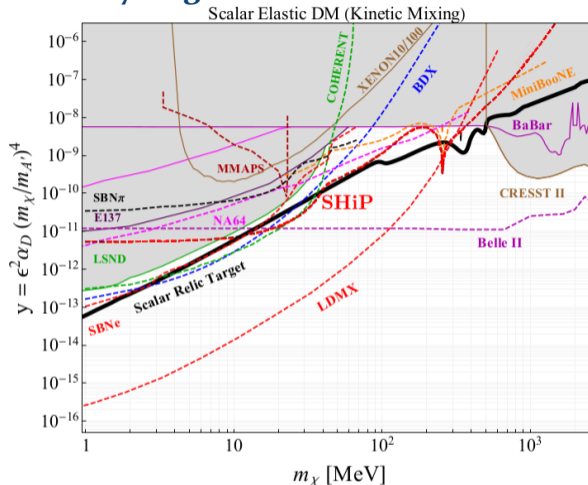


Figure 4: Dark photon sensitivity at SHiP.

- › Based on  $> 10^{20} \gamma$  at SHiP over 7 years
  - › Visible decays of dark photons
  - › Produced in qcd, bremsstrahlung and meson decays
  - › *No production via EM showers yet*  
→ *Work in progress*
  - › Complementary to regions studied by other experiments
  - › Top-right edge of sensitivity determined by short lifetime
- NB: Before re-optimisation*

# Sensitivity: Light Dark Matter



**Figure 5:** Light dark matter sensitivity at SHiP for  $\frac{m_{A'}}{m_{\chi}} = 3$ .

- › For dark matter lighter than WIMPs “direct detection” experiments quickly lose sensitivity.

## Two approaches:

- › missing mass/energy searches ( $\propto U^2$ )
- › scattering/recoil ( $\propto U^4$ )

SHiP: *Indirect* detection via electron and nuclear recoil in nuclear emulsion:

- › Main background for electron recoil from  $\nu_e$  scattering, but differences in the kinematics can be exploited.
- › *Preliminary; cascade production not yet implemented* → already best sensitivity for scattering

LDMX@SLAC:

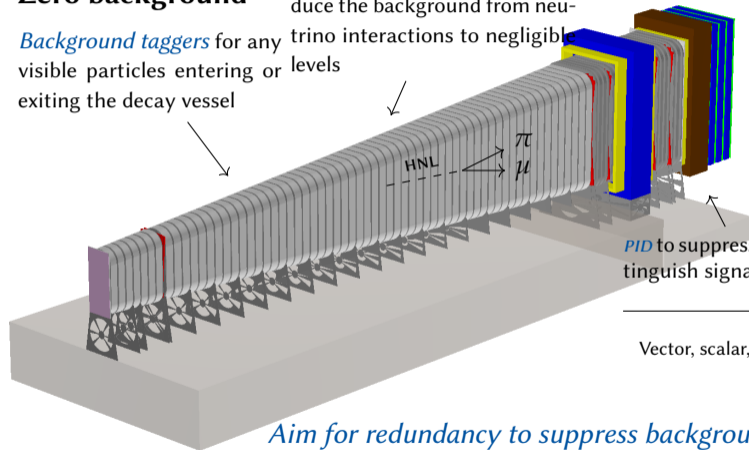
- › missing energy at electron beam

# Crucial challenges

## Zero background

*Background taggers* for any visible particles entering or exiting the decay vessel

*Evacuated decay vessel* to reduce the background from neutrino interactions to negligible levels



- > *Timing* to suppress combinatorial background from muons
- > *Tracking* for vertexing and impact parameter measurement

*PID* to suppress background and distinguish signal final states:

Particle	Final states
HNL, neutralino	$l^\pm \pi^\mp, l^\pm K^\mp, l^\pm \rho^\mp$
Vector, scalar, axion portals; goldstino	$l^\pm l^\mp$
HNL, neutralino, axino	$l^\pm l^\mp \nu_l$
Axion portal, sgoldstino	$\gamma\gamma$
Sgoldstino	$\pi^0 \pi^0$

*Aim for redundancy to suppress background*