

A new search for dark matter axions using quantum technologies

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DMUK meeting
22/9/2022

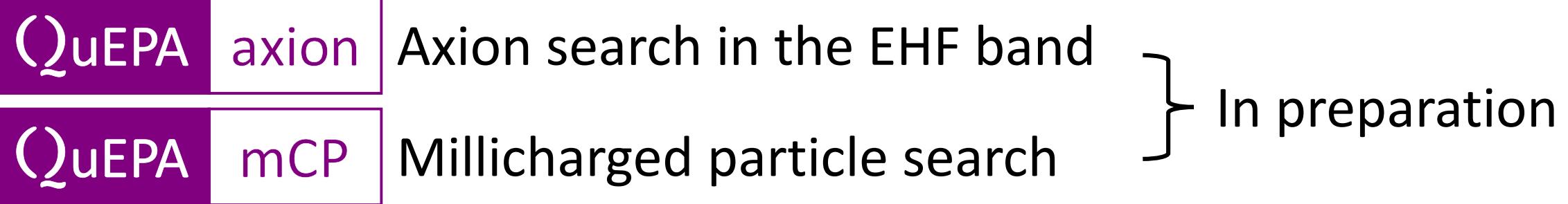


Imperial College
London



Who we are

QuEPA Quantum Enhanced Particle Astrophysics



Project started 1/9/2022

Local Team



Jack Devlin
PI



David Pitman
Mechanical
Engineer



Richard Thompson
Penning
trap expert



Norbert Klein
Microwave
expert

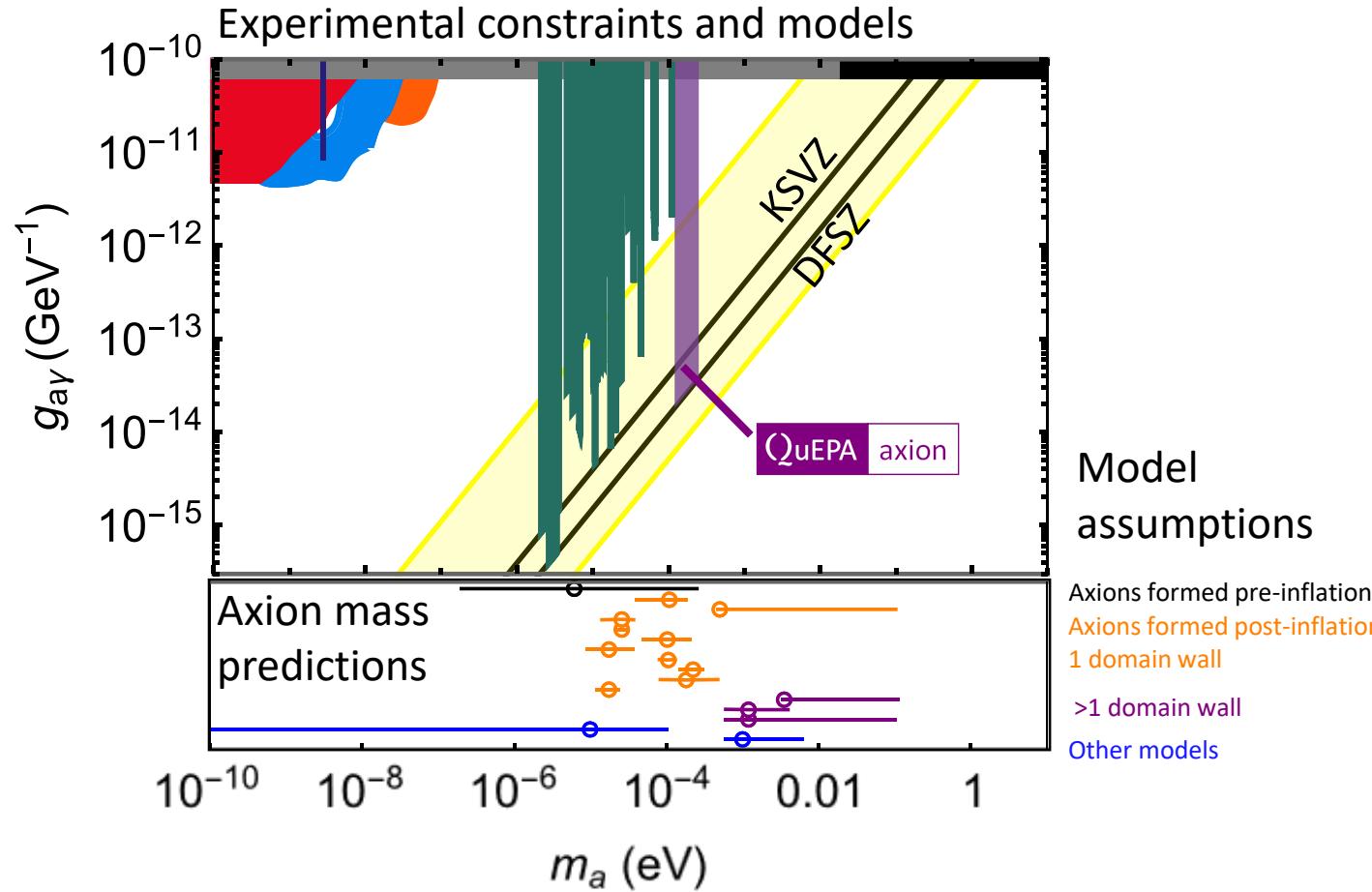
Support



Mike Tarbutt
Quantum
Science and
Fundamental
Physics

**Imperial College
London**

Overview



Use tools of AMO physics
to search for axions

Medium term
 $m_a = 120 \leftrightarrow 240 \mu\text{eV}$
@ DFSZ sensitivity

Figure adapted from:

G. Iraizozza and J. Redondo, Prog. Part. Nucl. Phys. **102**, 89 (2018).

Additions from <https://cajohare.github.io>

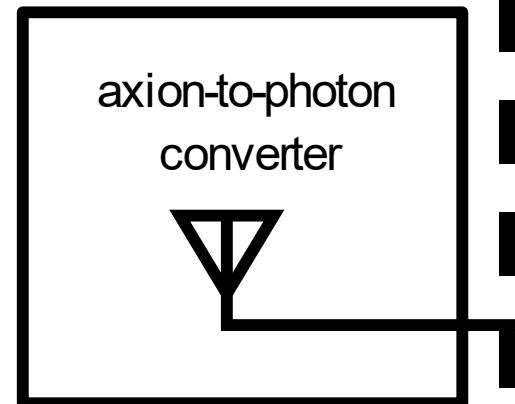
An axion haloscope

Converter $m_a \rightarrow \nu_a = m_a c^2/h$

$$P_{out} = \frac{\hbar \rho_{DM} g_{a\gamma\gamma}^2}{c^3 \mu_0 m_a} \times B^2 V_m Q$$

Desirable

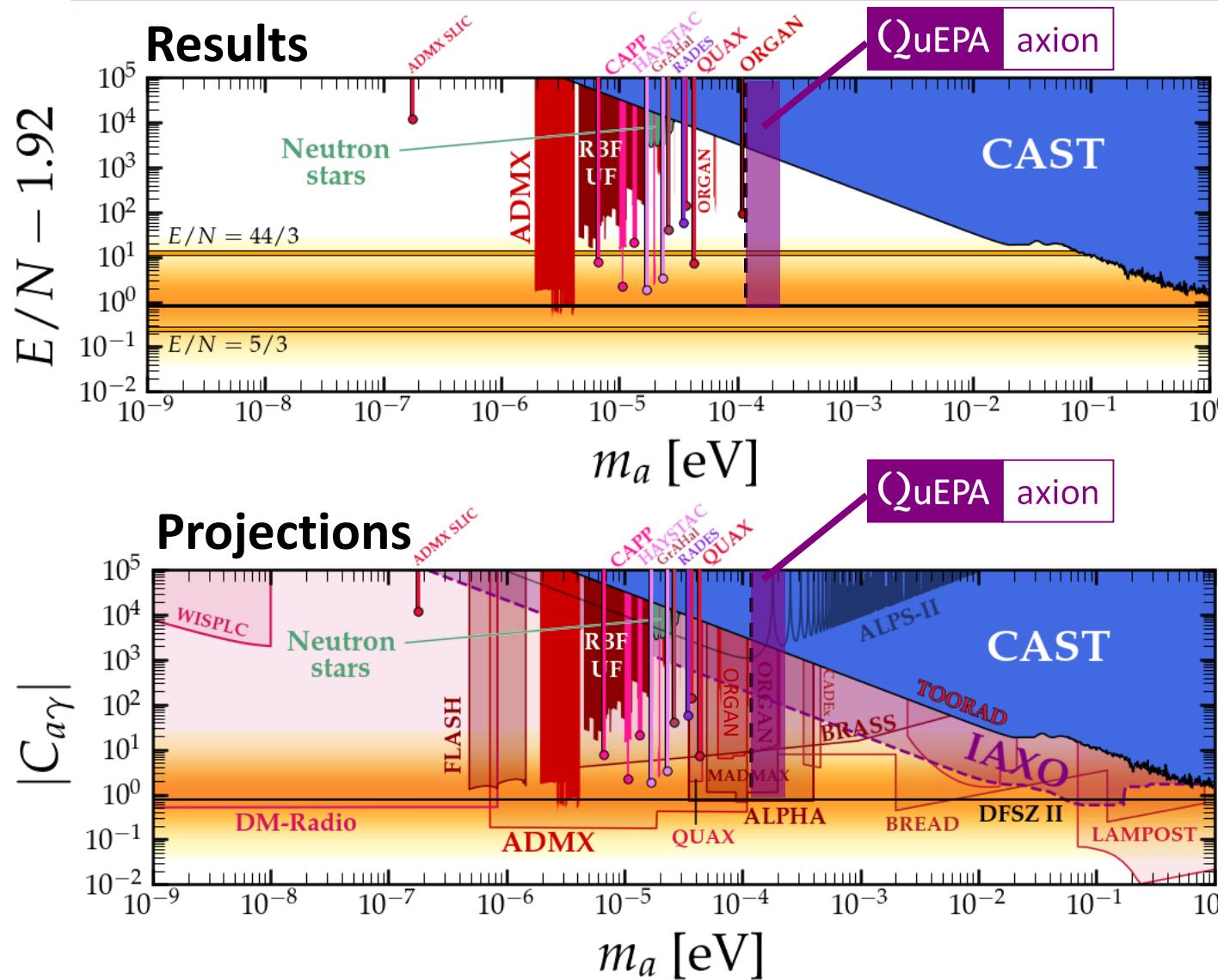
- High volume factor
- Q-factor up to $Q_a = 10^6$
- High B field compatible
- Easy to adjust frequency
- Broad tuning range
- Feasible



Why hard at higher masses/frequencies?

→ Typically $V_m \propto \frac{1}{\nu_a^3}$, $Q \propto \frac{1}{\nu_a^{2/3}}$ (for copper)

Other axion-photon converter efforts above 30 GHz



Dielectric cavities

ORGAN QDM lab (Univ. Western Australia)
Lead by Dr. Michael Tobar

TEM_{00q} Fabry-Perot

ORPHEUS Rybka et al.

Plasma haloscope

ALPHA Frank Wilczek et al.

“Magnetized mirrors”...

BRASS University of Hamburg

BREAD Cambridge/Fermilab, higher m_a

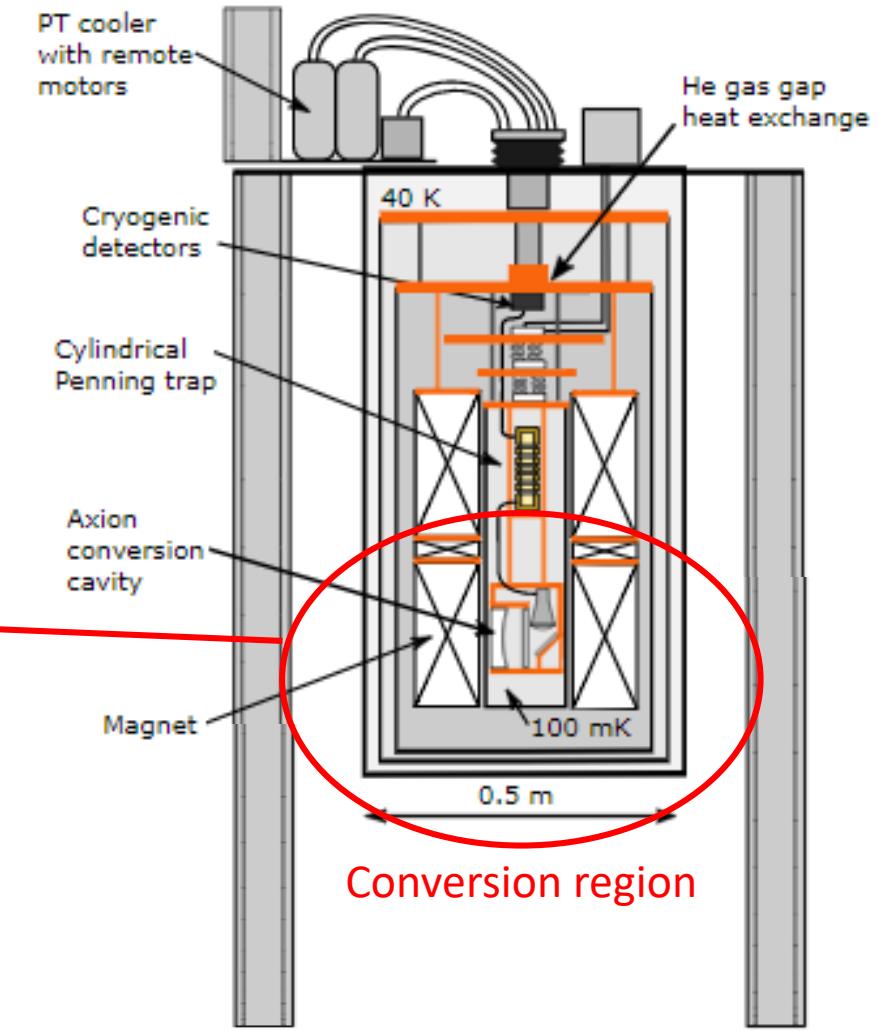
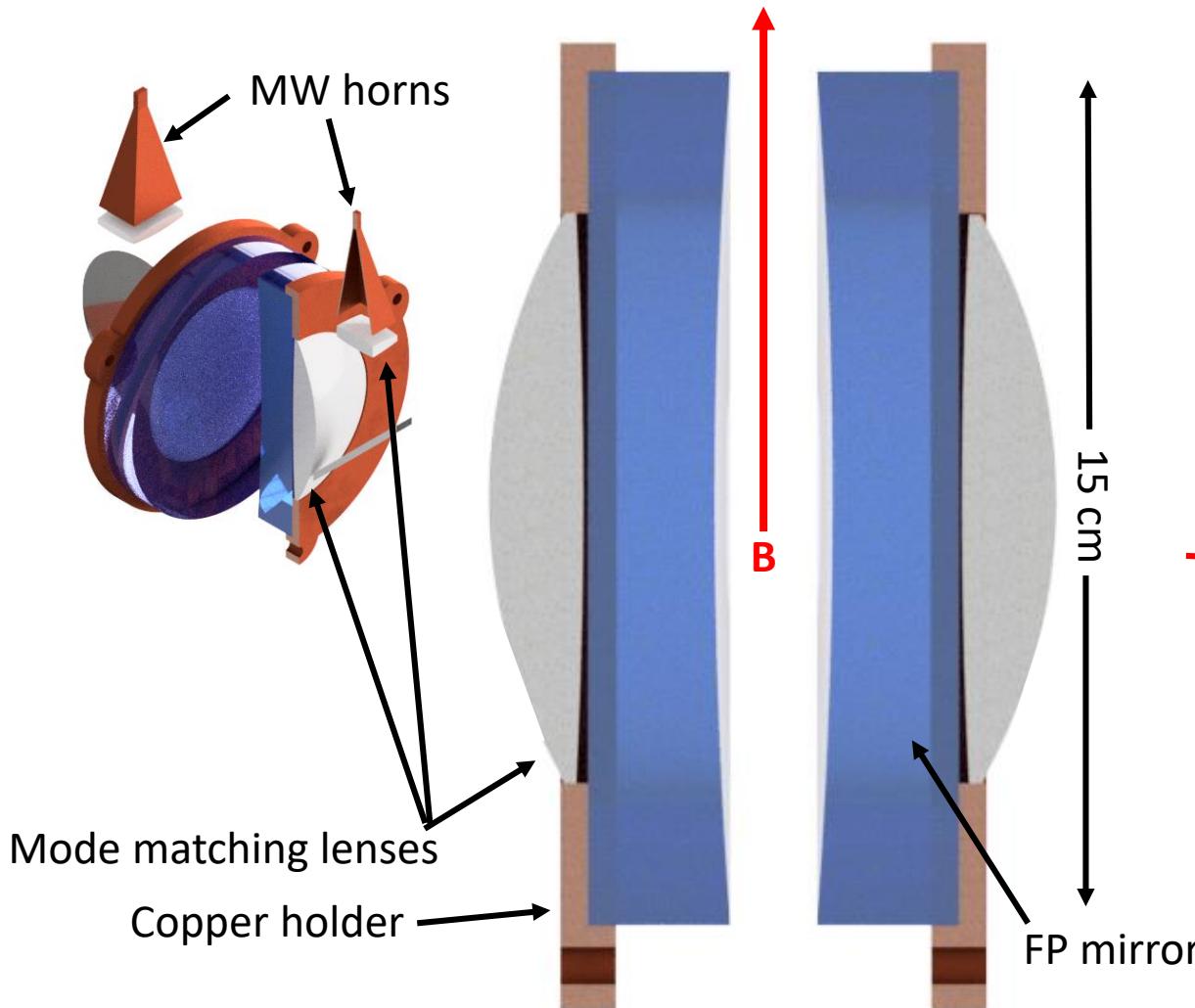
.. with dielectric boost

MADMAX Big effort at DESY, lower m_a
initially

Many more new ideas between 5-30 GHz

Our converter concept

Large mode area Fabry Perot cavity operating in TEM_{001} mode



See ORPHEUS for alternative F-P TEM_{00q} concept G. Rybka et al., Phys. Rev. D **91**, 011701 (2015)

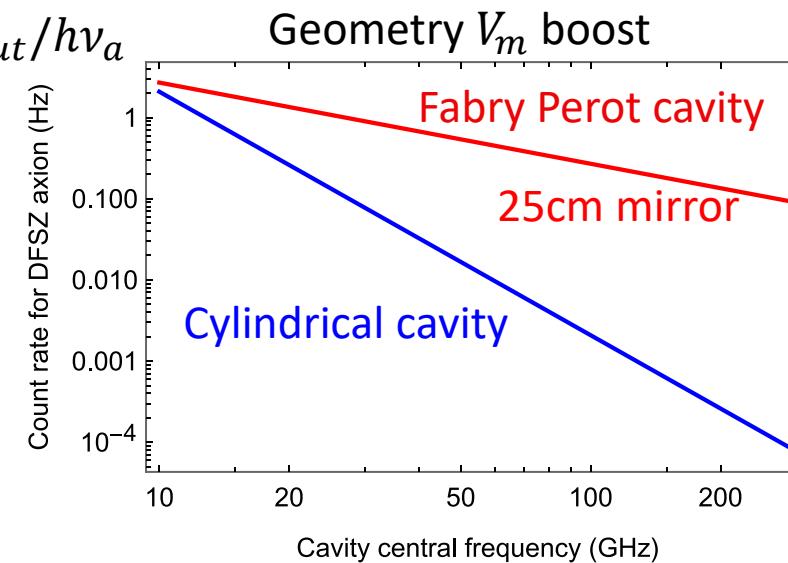
Performance

Desirable

- High volume factor ✓
- Q-factor up to $Q_a = 10^6$ ✓
- High B field compatible ✓
- Easy to adjust frequency ✓
- Broad tuning range ✓
- Feasible ✓

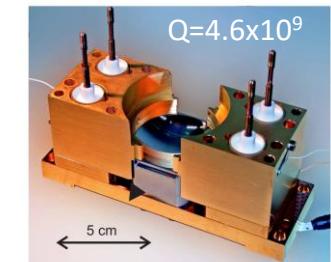
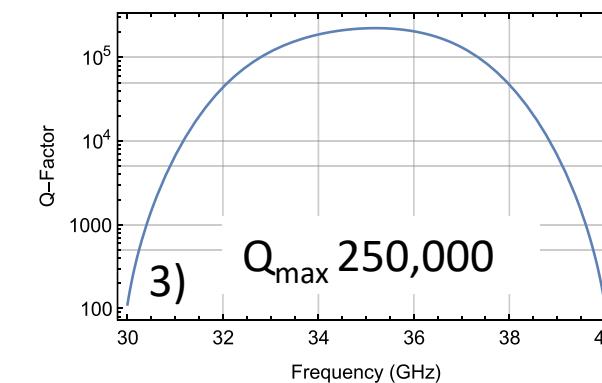
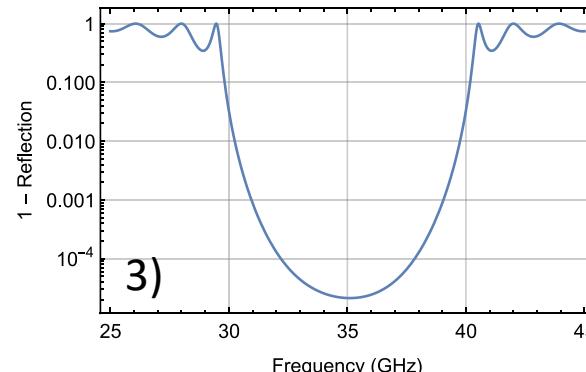
15-25 cm diameter, realistic B-fields

$$R = P_{out}/hv_a$$



Options for coatings:

- 1) High TC YBCO see J. Golm et al., arXiv:2110.01296
- 2) Sapphire/NbTi multilayer see Xiaoxiang Xi et al., Phys. Rev. Lett. **105**, 257006 (2010)
- 3) Sapphire/quartz Bragg mirror



At design frequency

$$R_{FP} = KC_{a\gamma} w^2 \frac{\pi c}{16 v_a} QB^2$$

$$R_{CC} = KC_{a\gamma} 0.69 \times 1.8 * \frac{c^3}{v_a^3} QB^2$$

Crossover at ~10 GHz

C. Boutan et al., Phys. Rev. Lett. **121** 261302 (2018)

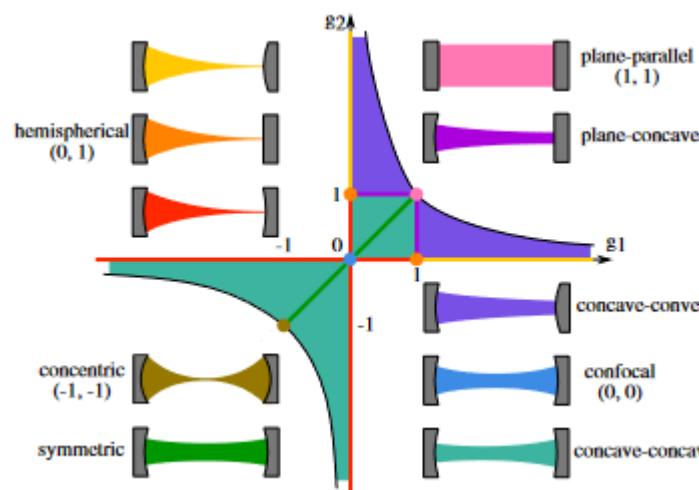
S. Kuhr et al., Appl. Phys. Lett. **90**, 164101 (2007)

B=10, Q=10⁶ for both

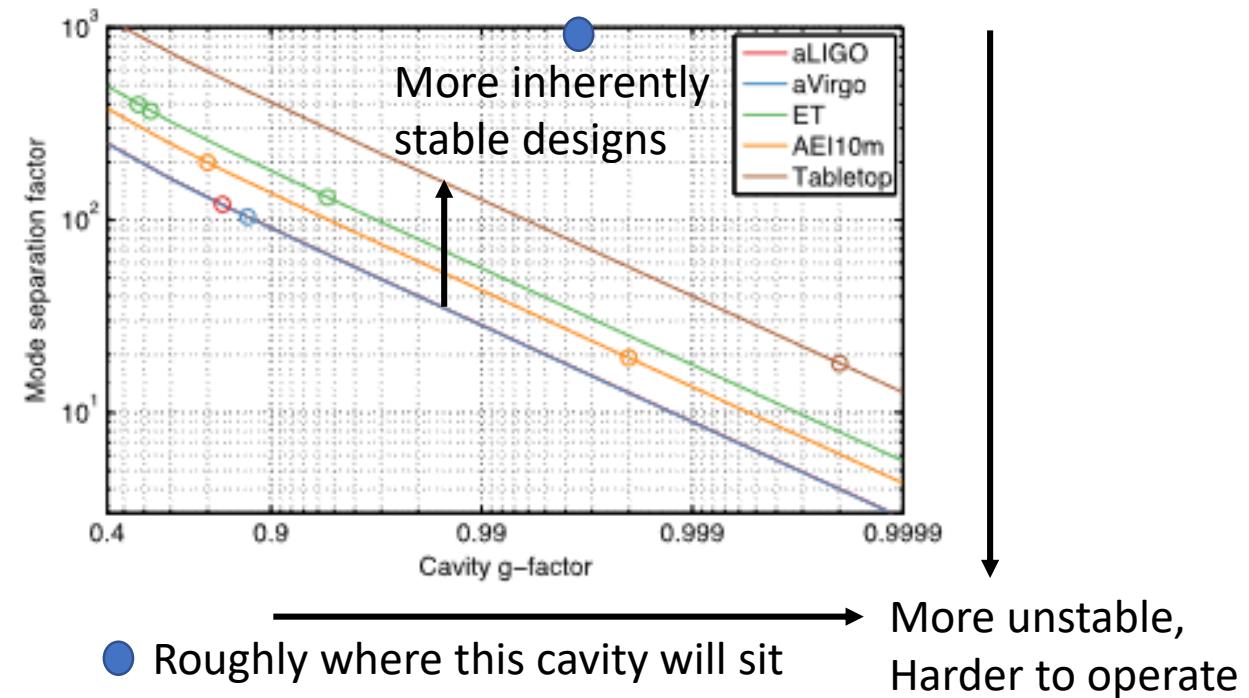
Research program

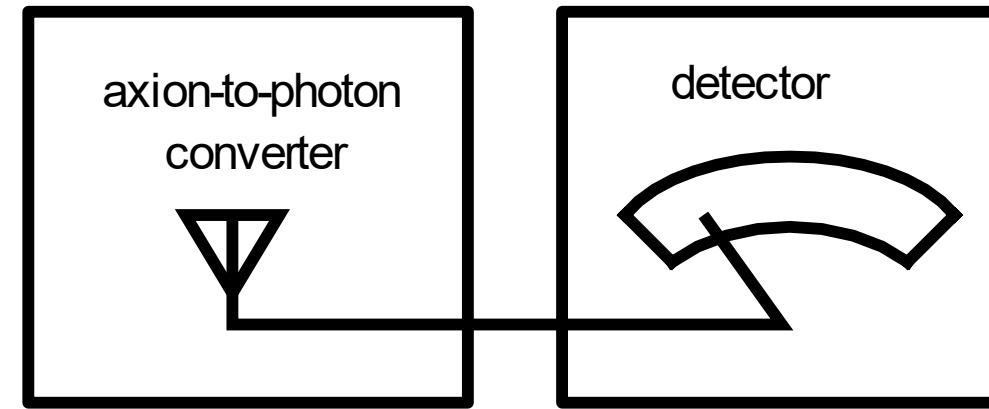
R&D program:

- 1) Which coating technology is best?
- 2) Test mirror construction techniques
- 3) How close to the instability limit can we reach?

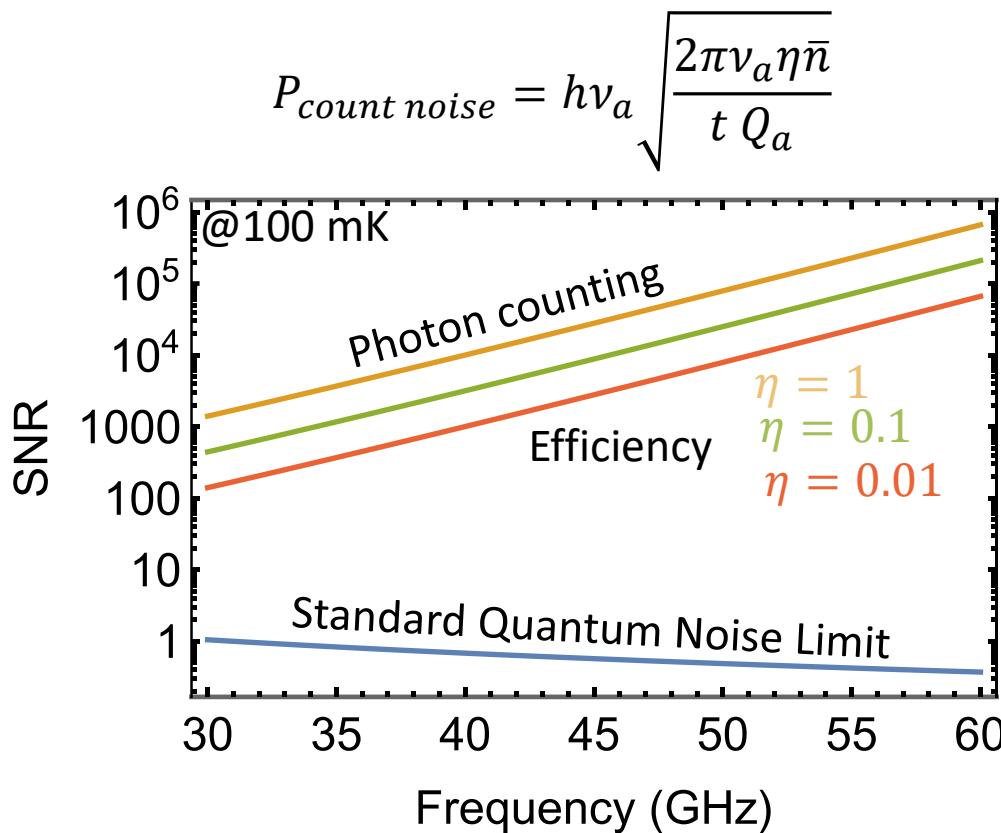


$$\delta = \frac{\Delta f_n}{\text{FWHM}/2}$$



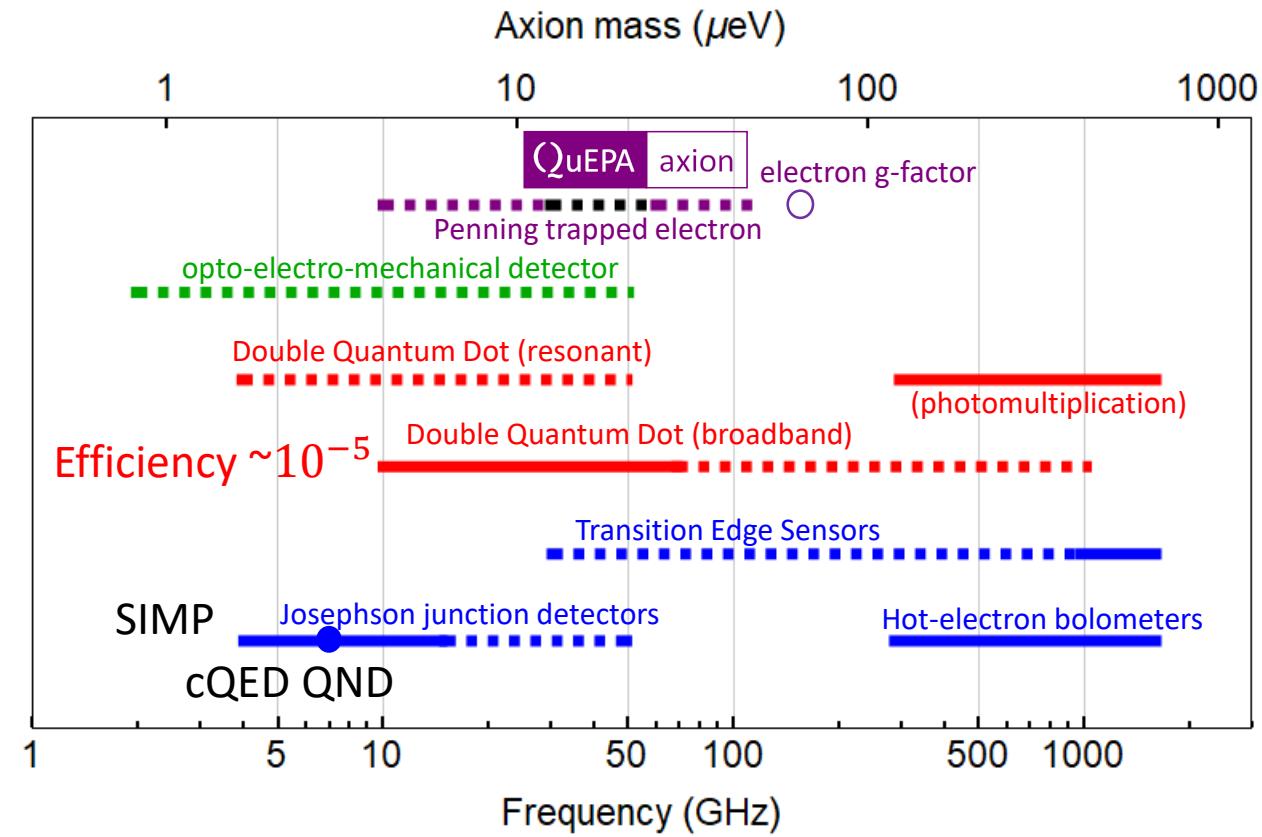


Making a better detector



$$P_{\text{SQL noise}} = h\nu_a (\bar{n} + 1) \sqrt{\frac{\nu_a}{t Q_a}}$$

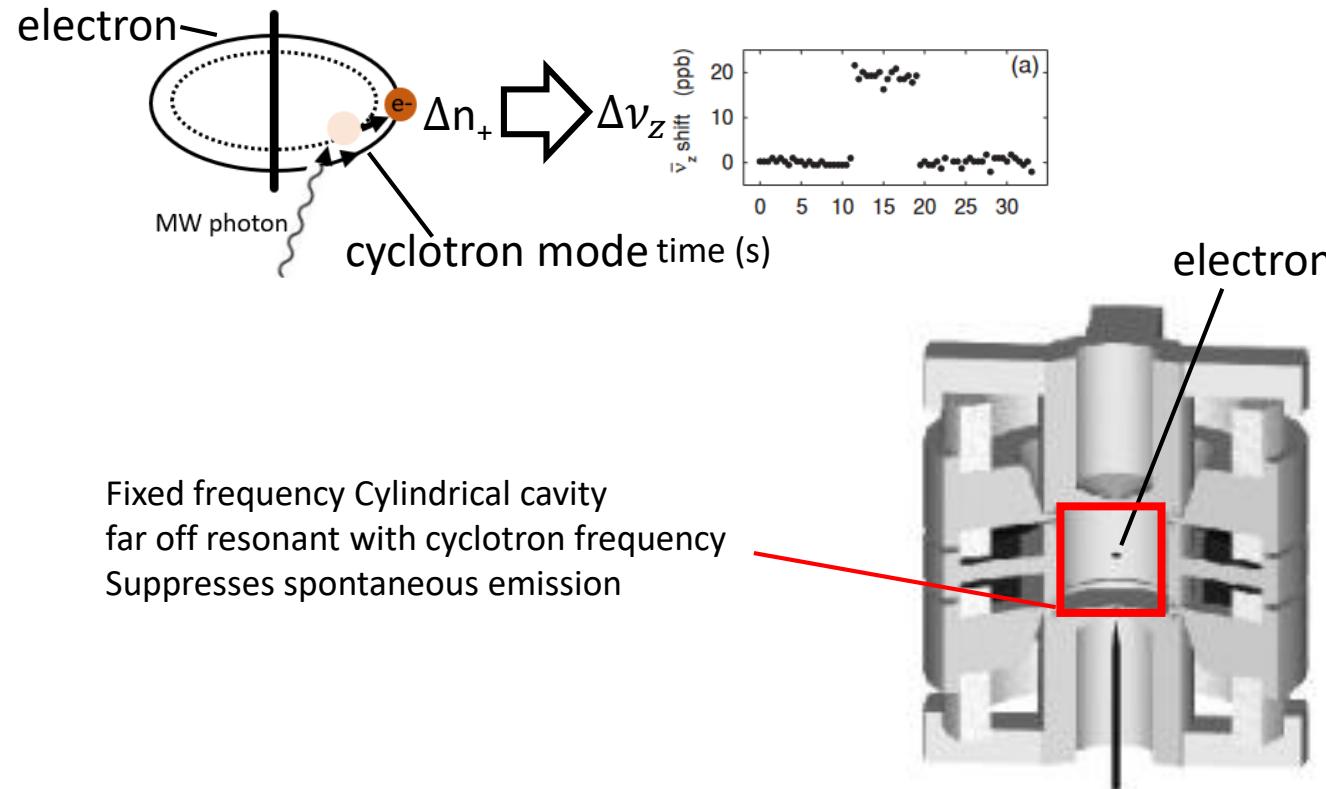
What technology to use?



AMO Quantum technology for Single Photon Counting

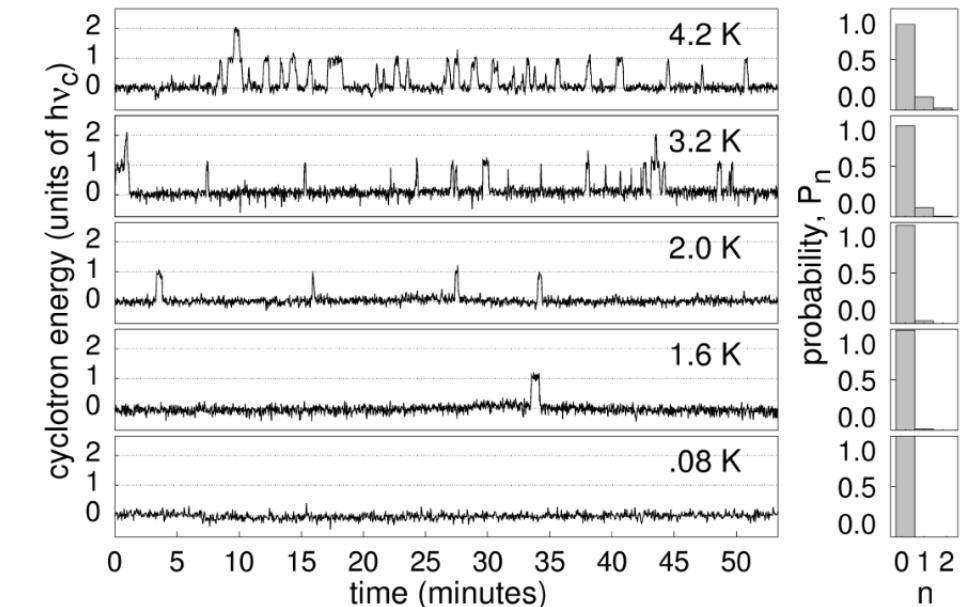
An old idea: CARRACK I & II Rydberg atoms

Counting photons with a single trapped electron



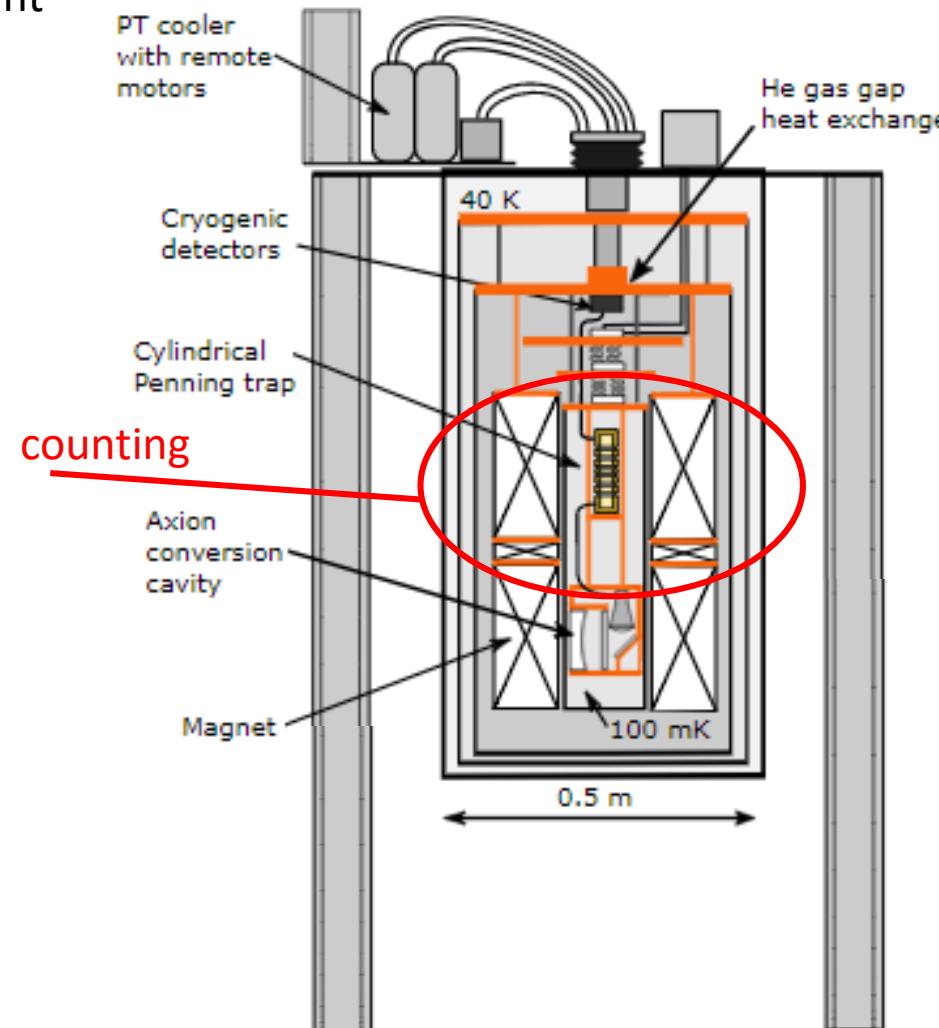
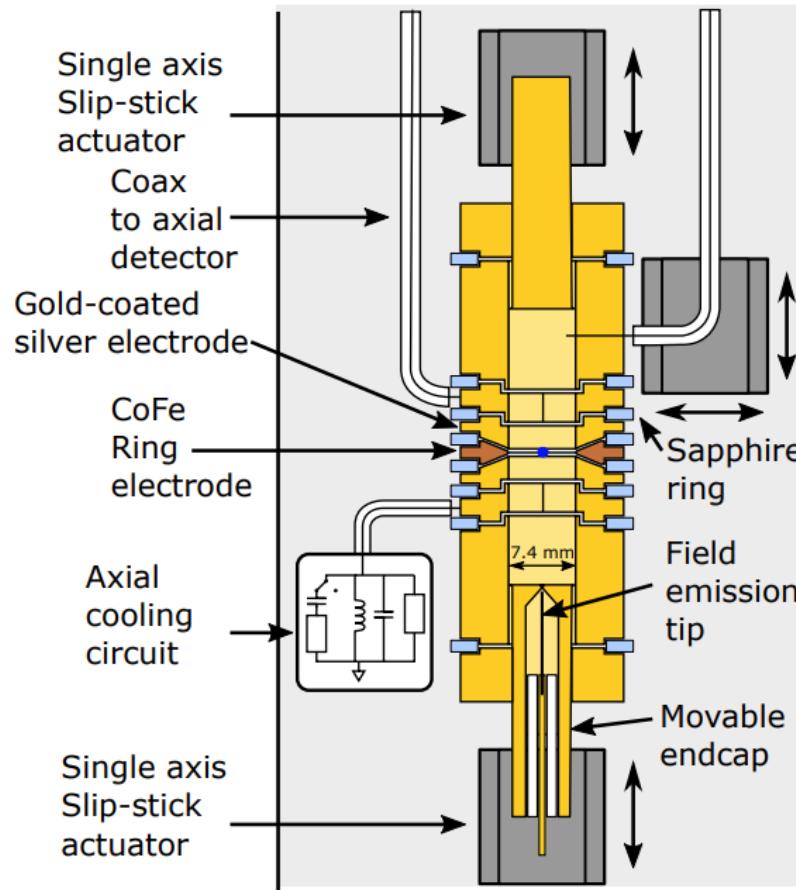
M. Tada et al., Nuclear Physics B (Proc. Suppl.) **72** 164 (1999)
 M. Tada et al., Physics Letters A **349** 6 488 (2006)

S. Peil and G. Gabrielse Phys. Rev. Lett. **83**, 1287 (1999)
 D. Hanneke, S. Fogwell Hoogerheide, and G. Gabrielse, PRA **83**, 052122 (2011)



Our approach

Moving endcaps allows TE_{11x} modes to be matched to cyclotron frequency for efficient detection



R&D goals

- Application of frequency measurement techniques compatible with on resonant cavity
- Medium-Q (25,000) cavity.

Some realistic numbers

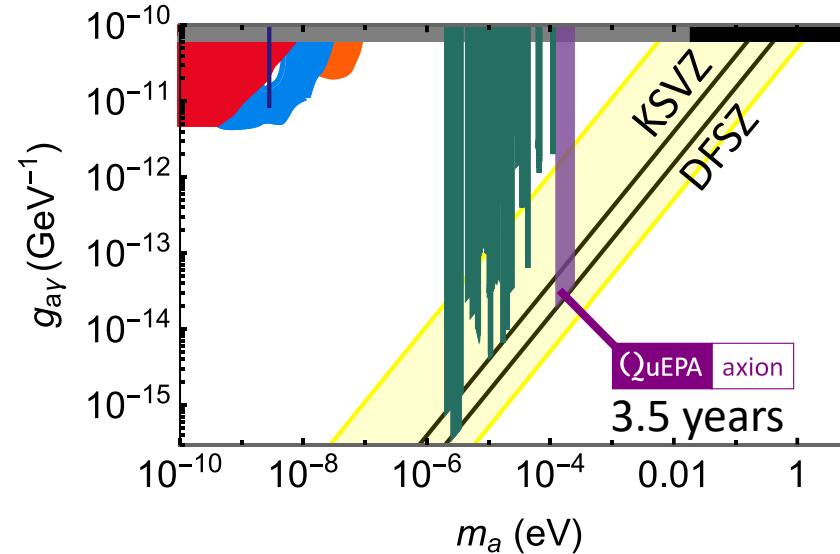
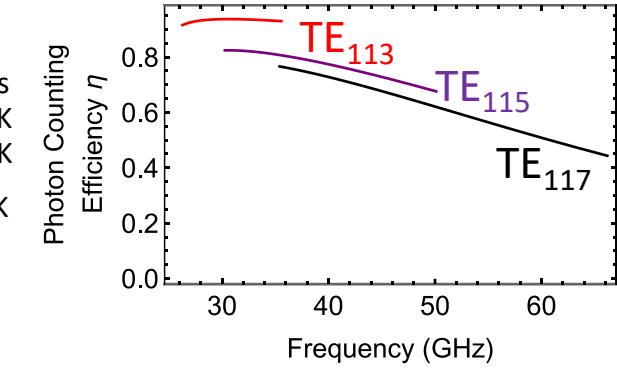
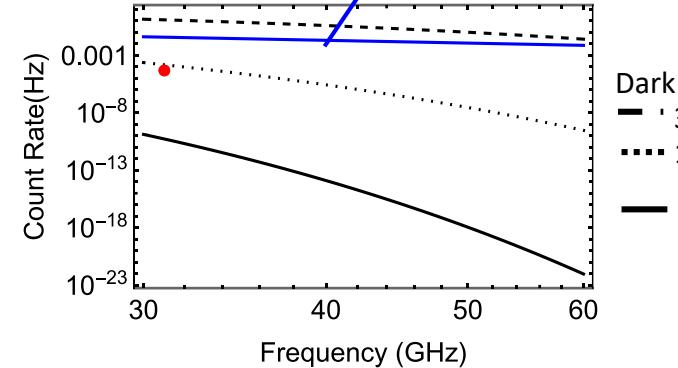
Axion production cavity assumptions

Parameter	Value	Why?
Q	250,000	Calculated Bragg mirror
B	7 T	Modest large bore Nb-Ti magnet field
Mirror diameter	15 cm	Standard semiconductor substrate size
Mirror radius of curvature	1 m	$g_{\max} = 0.9975$ $w_{\max} = 1.5 \text{ cm} = d/10$ $\Delta f_{\min} \frac{Q}{f} = 4 * 10^3$
In-coupling losses	50 %	Modest mode mismatch

Photon counting assumption

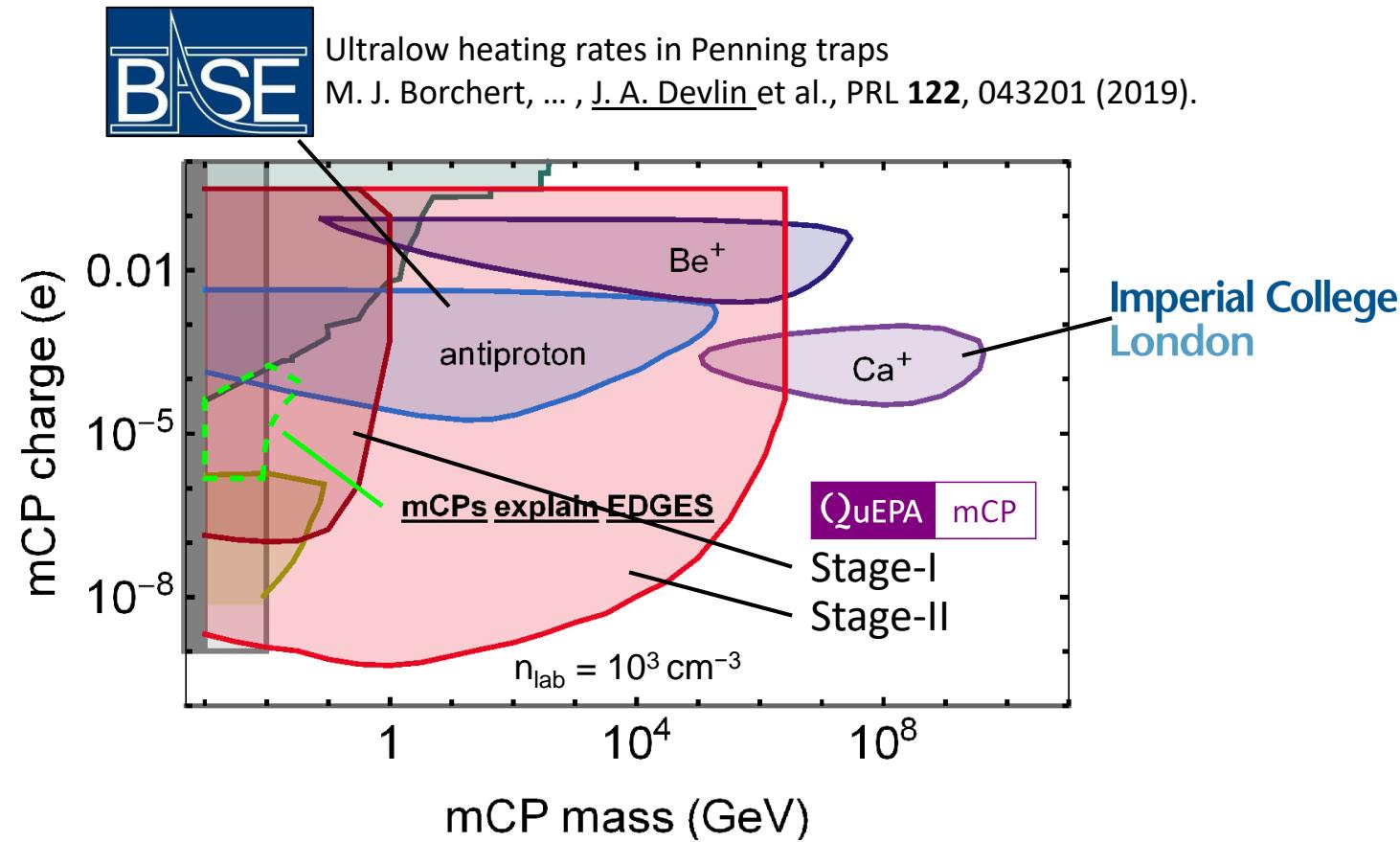
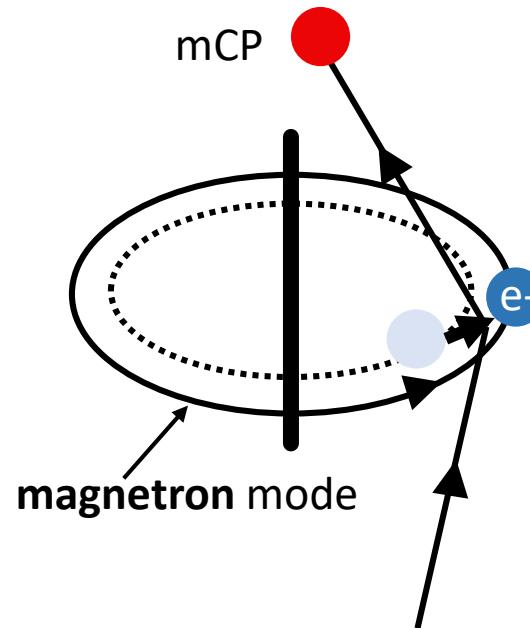
Parameter	Value	Why?
Q	25,000	Cavity wavemeter/electron g-factor
Missed fraction	5 %	< 1 ppm axial stability

DFSZ axion conversion rate



Other applications – millicharged particles

Look for collisions between millicharged particles (mCPs) and electrons that change the magnetron mode



Status and short-term goals

Project started 1/9/2022



£125 k for developing Fabry Perot cavity
+ 5-year University Research Fellowship for JAD

Imperial College London

40 m² lab space in Ion Trapping Group
Use of facilities in Dept. of Engineering
Equipment sharing with Centre for Cold Matter



Science and
Technology
Facilities Council

£500k fEC “Developing Quantum Technologies for Fundamental Physics” call for cryocooler (300 mK) + magnet and personnel



2-year R&D goals

Fabry-Perot cavity

- Cavity coating
- Construct cryogenic test cavity

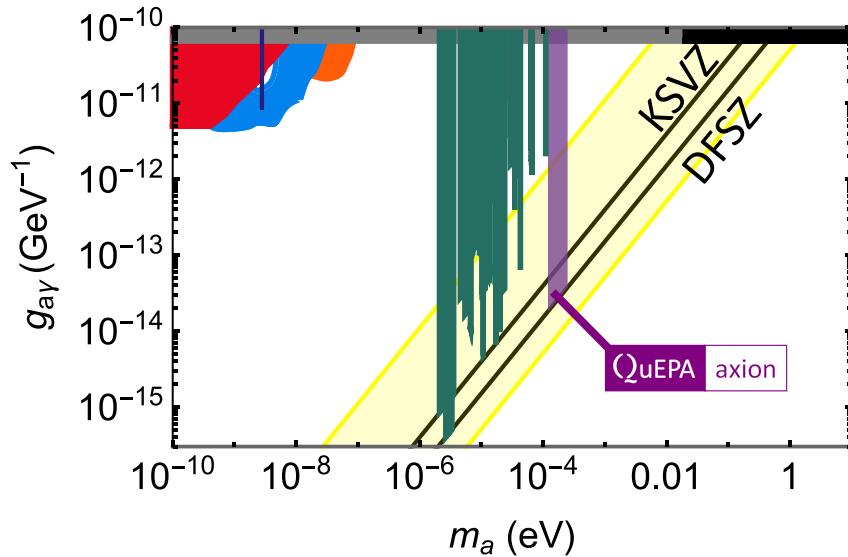
Penning trap

- Construct trap
- Trap and detect single electrons
- Measure magnetron heating rate
- Demonstrate principle of photon counting

2-year physics goal

→ mCP search

Thank you for listening



The local team



Jack Devlin

PI



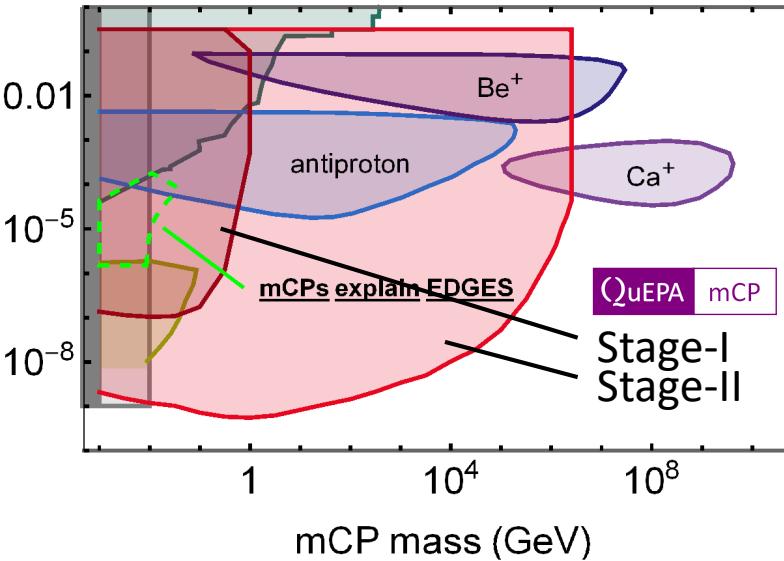
David Pitman

Mechanical
Engineer



Richard
Thompson

Penning
trap expert



Support



Norbert
Klein

Microwave
expert



Mike
Tarbutt

Quantum
Science and
Fundamental
Physics

We're recruiting: <https://www.imperial.ac.uk/ion-trapping/positions/> 1 PhD