

# AMPLITUDES, RESONANCES & THE ULTRAVIOLET COMPLETION OF GRAVITY

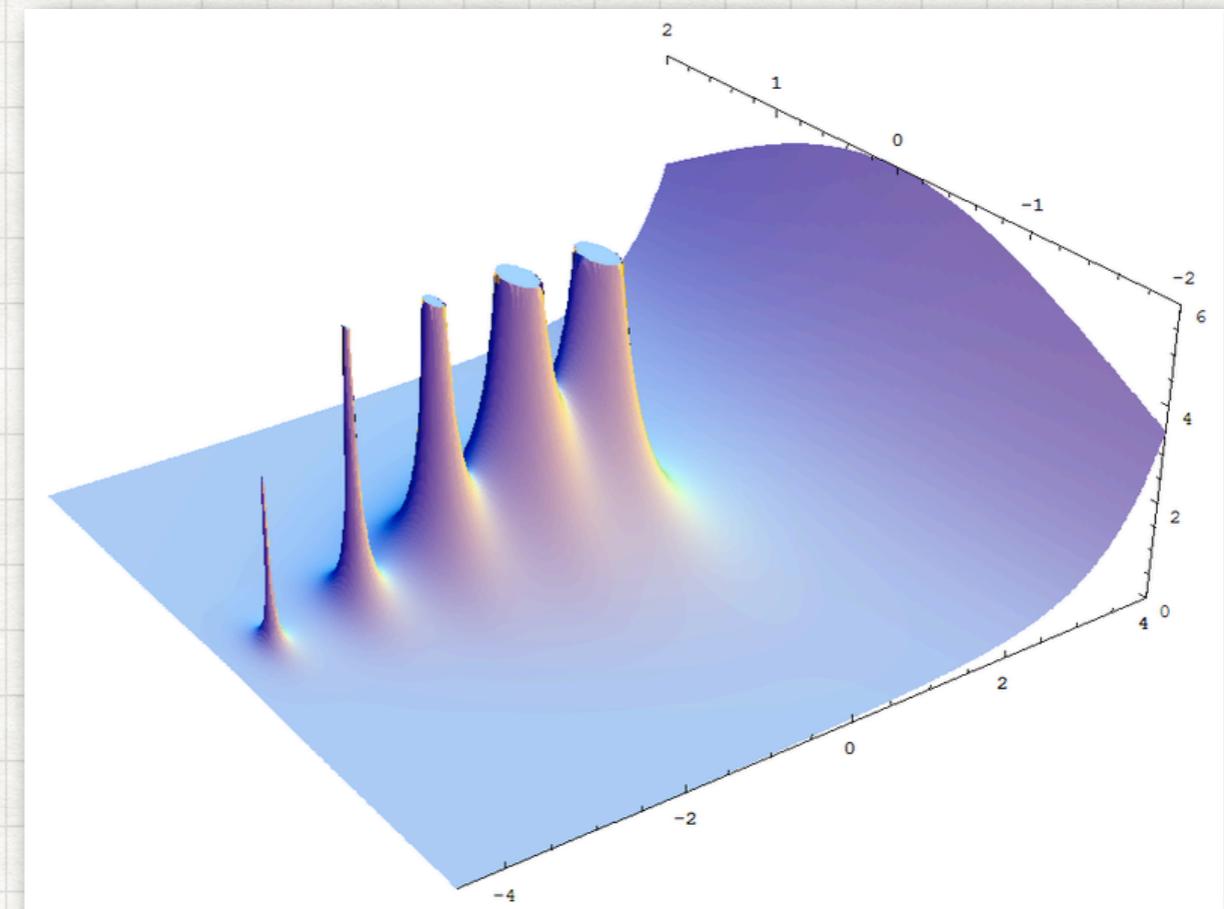
Rodrigo Alonso — IPPP



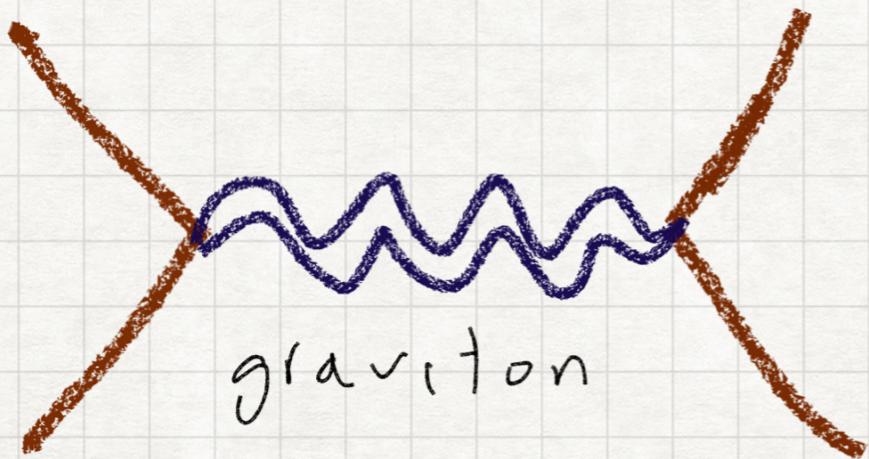
HEP seminar, DAMPT April 30th  
[R.A. & A. Urbano (906.((687)]

# OUTLINE

1. ON SHELL AMPLITUDES
2. UNITARITY
3. RESONANCES FOR GRAVITY
4. ANALYSIS.

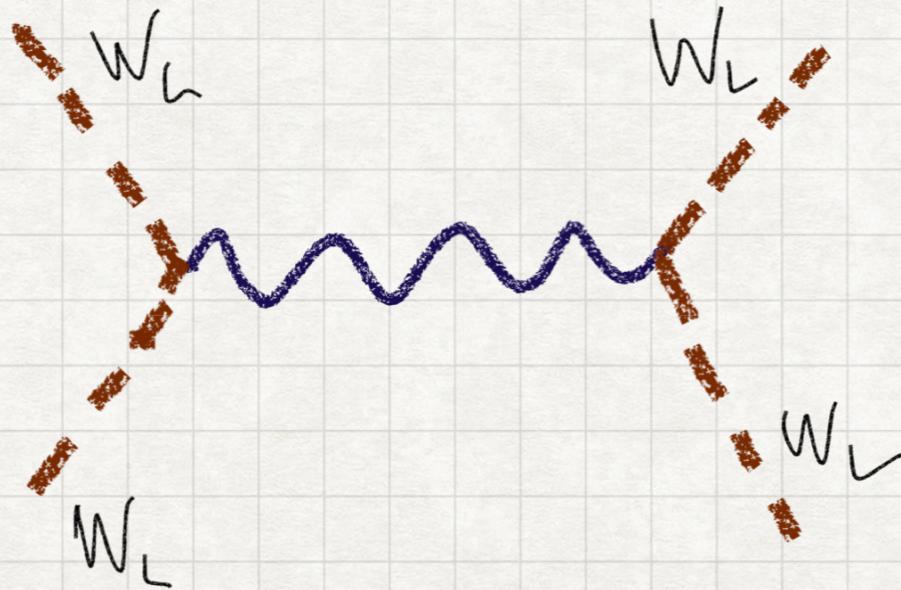


# This talk in a nutshell



$$\mathcal{A} \sim \frac{s}{M_{\text{pl}}^2}$$

Let's add  
Resonances  
to fix this



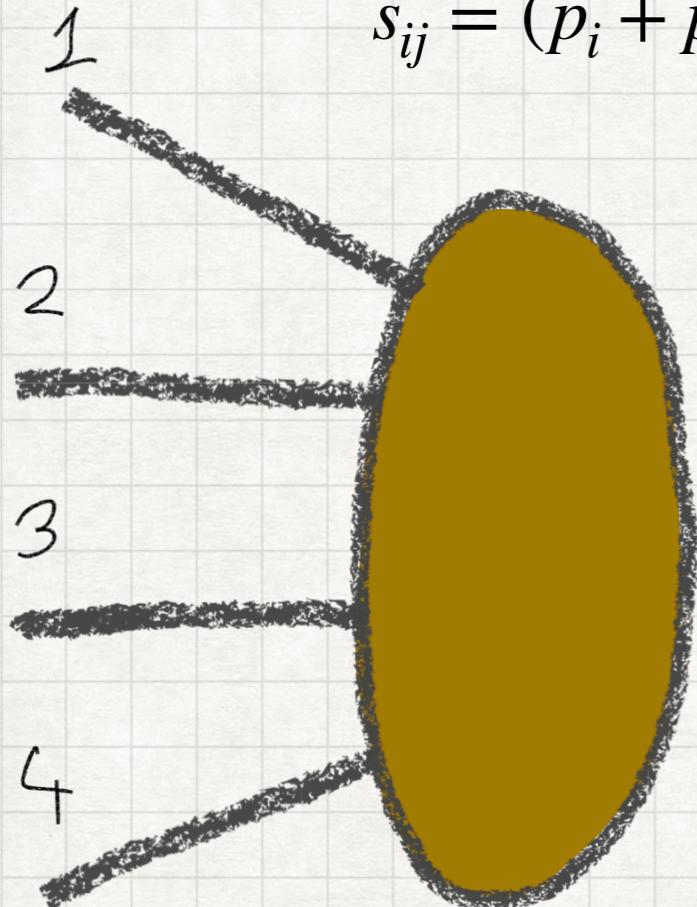
$$\mathcal{A} \sim \frac{s}{v_{\text{ew}}^2}$$

The Higgs

# I. On-shell amplitudes

# Setting the stage

$$s_{ij} = (p_i + p_j)^2$$



$$p_{2,4} \rightarrow -p_{2,4}$$

$$p_{3,4} \rightarrow -p_{3,4}$$

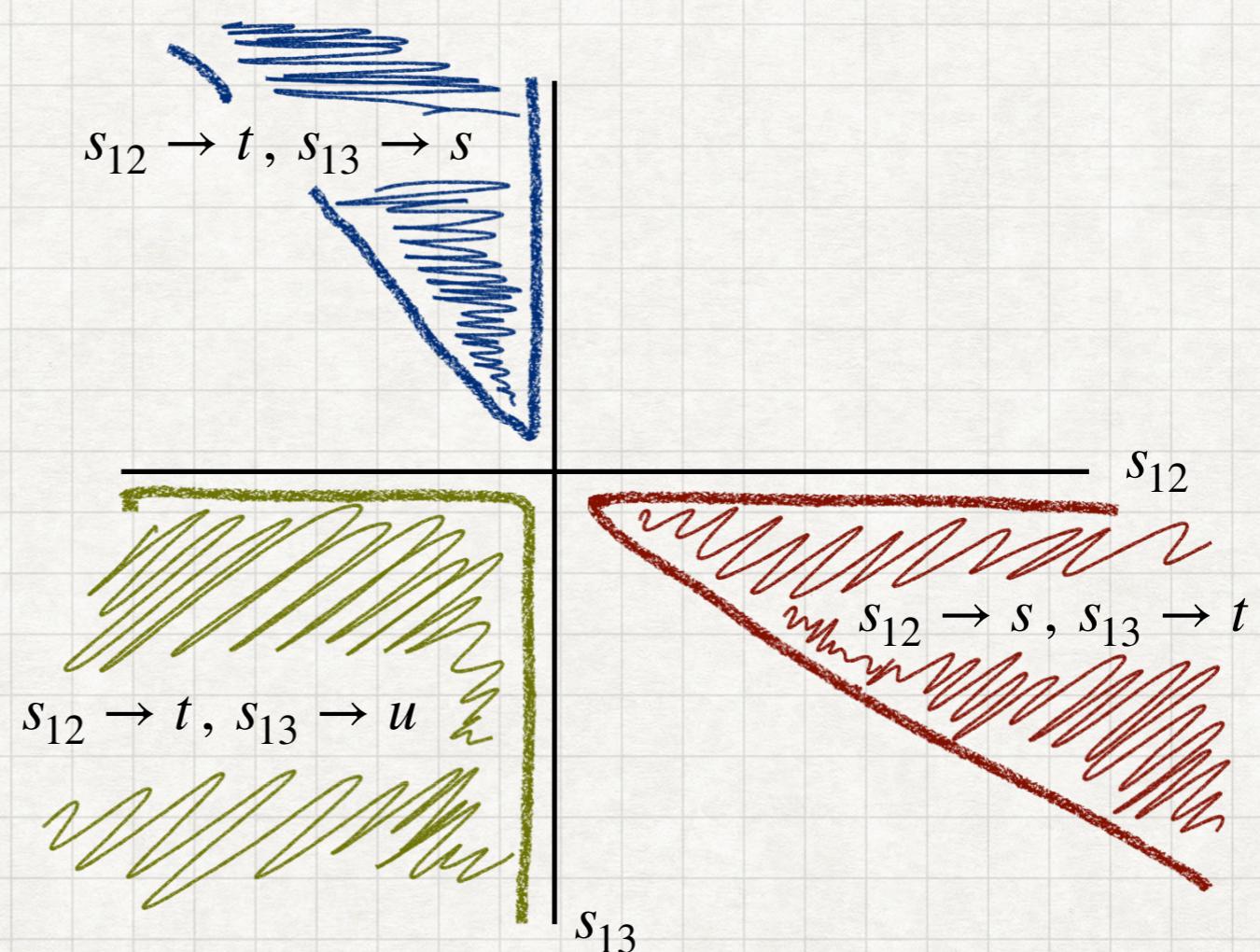
$$p_{2,3} \rightarrow -p_{2,3}$$

+

Helicity

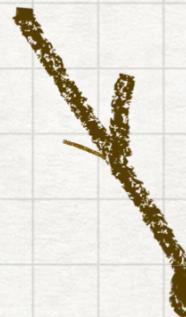
Flip

$$\mathcal{A}(s_{12}, s_{13})$$



$$s_{12} + s_{13} + s_{14} = \sum m_i^2$$

# Helicity scaling or Little group Reps.

 $\alpha |p\rangle$  $SO(1,3); \quad |p\rangle \rightarrow i\Lambda^{\mu\nu}\sigma_{\mu\nu} |p\rangle$  $U(1)_{LG}; \quad |p\rangle \rightarrow |p\rangle e^{-i\phi/2}$  $\dot{\alpha} |p]$  $SO(1,3); \quad |p] \rightarrow i\Lambda^{\mu\nu}\bar{\sigma}_{\mu\nu} |p]$ 

(Just Weyl spinors)

 $U(1)_{LG}; \quad |p] \rightarrow |p] e^{i\phi/2}$ 

$$\varepsilon_+^\mu = \frac{\langle \xi | \sigma^\mu | p ]}{\sqrt{2} \langle \xi | p } \quad$$

$$U(1)_{LG}; \quad \frac{\langle \xi | \sigma^\mu | p ] e^{i\phi/2}}{\sqrt{2} \langle \xi | p } e^{-i\phi/2}$$

In this talk all external  
lines  $\rightarrow$  massless

Learn more

[Dixon 9601359, Elvang & Huang 1308.1697]

# Helicity scaling or Little group Reps. Massive



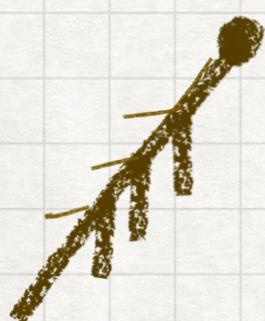
$$_{\alpha} \left| p_I \right\rangle \quad SO(1,3); \quad \left| p_I \right\rangle \rightarrow i\Lambda^{\mu\nu}\sigma_{\mu\nu} \left| p_I \right\rangle \quad SU(2)_{LG}; \quad \left| p \right\rangle \rightarrow \left| p_J \right\rangle i[T_a]_{IJ}\phi_a$$



$$^{\dot{\alpha}} \left| p_I \right\rangle \quad SO(1,3); \quad \left| p_I \right\rangle \rightarrow i\Lambda^{\mu\nu}\bar{\sigma}_{\mu\nu} \left| p_I \right\rangle \quad SU(2)_{LG}; \quad \left| p \right\rangle \rightarrow \left| p_J \right\rangle i[T_a]_{IJ}\phi_a$$

$$u(p) = \begin{pmatrix} \left| p_I \right\rangle \\ \left| p_I \right\rangle \end{pmatrix}$$

Just half of  
Dirac  
Spinors



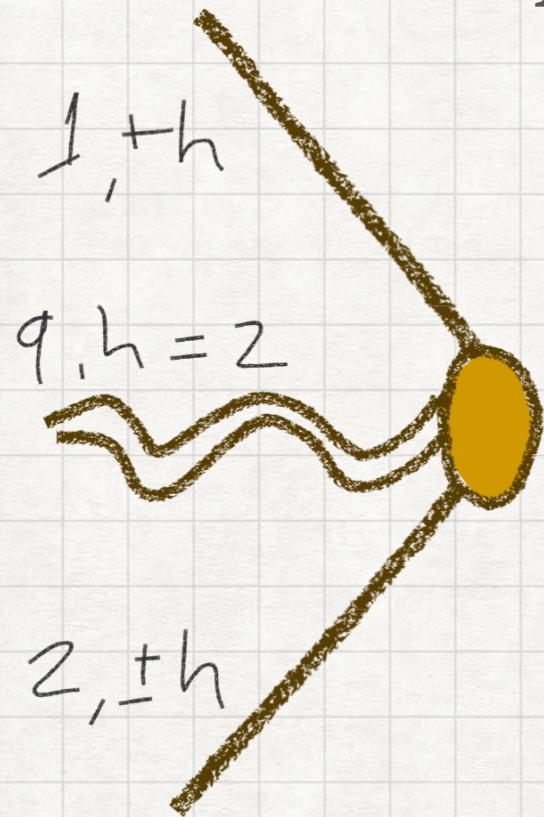
Spin J  
Massive  
State

$$\otimes_{\text{sym}}^{2J} |p] = |p_{I_1}] \times |p_{I_2}] \times \dots |p_{I_{2J}]} \quad (2J+1 \text{ elements})$$

Learn more

[Arkani-Hamed & Huang(x2), 1709.04891]

# 3 point amplitude for gravity



$$\mathcal{A}_3^{\text{GR}} = C [1]^{2h} \times [2]^{\pm 2h} \times [q]^4 = C [12]^a [1q]^b [2q]^c$$

$h, h$        $\downarrow$        $h, -h$

$$[12]^{2h-2} [1q]^2 [2q]^2, \quad [12]^{-2} [1q]^{2+2h} [2q]^{2-2h}.$$

Mass dimension of  $C$

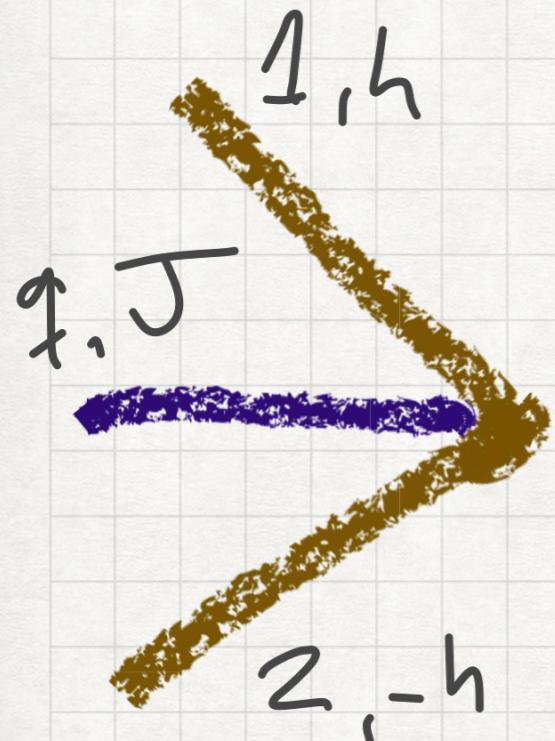
$$\dim(\mathcal{A}_n) = 4 - n \quad \Rightarrow \quad [C] = \binom{1 - 2h - 2}{1 - 2}$$

$$C = \frac{\sqrt{8\pi}}{M_{\text{pl}}}$$

Only opposite helicity case in GR

$$\boxed{\frac{\sqrt{8\pi}}{M_{\text{pl}}} \frac{[1q]^{2+2h} [2q]^{2-2h}}{[12]^2}}$$

# 3 point amplitude massive spin J

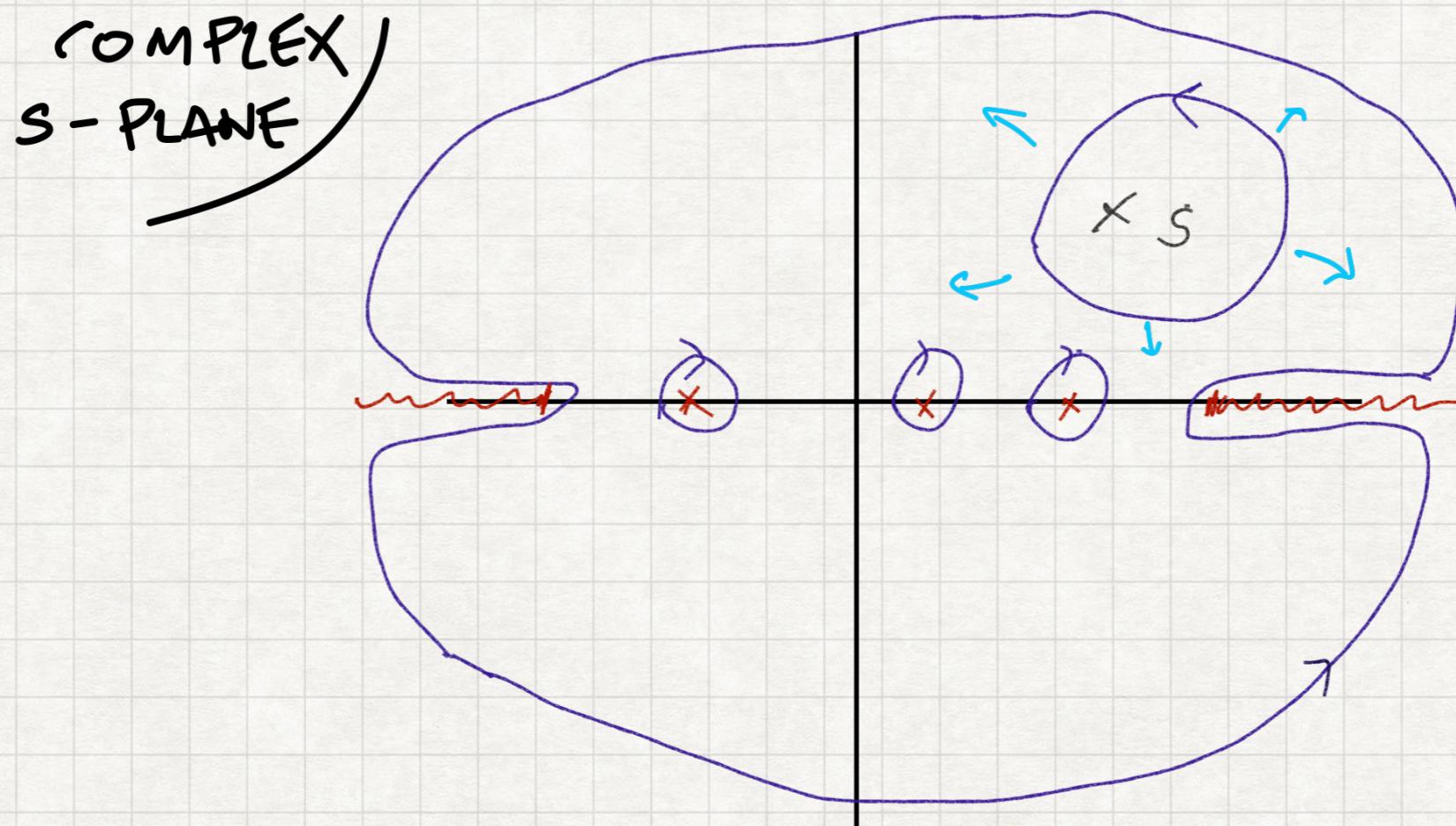


$$gM \frac{(\langle q_I \hat{P}_{12} q_I \rangle)^{J-2h} ([1q_I] \langle 2q_I \rangle)^{2h}}{M^{2J}}$$

$$\hat{P} = \sigma_\mu P^\mu \quad P_{ij} = p_i - p_j$$

# On shell methods @ tree level

$$2\pi i \mathcal{A}(s, t) = \oint \frac{\mathcal{A}(s', t)}{s' - s} ds'$$

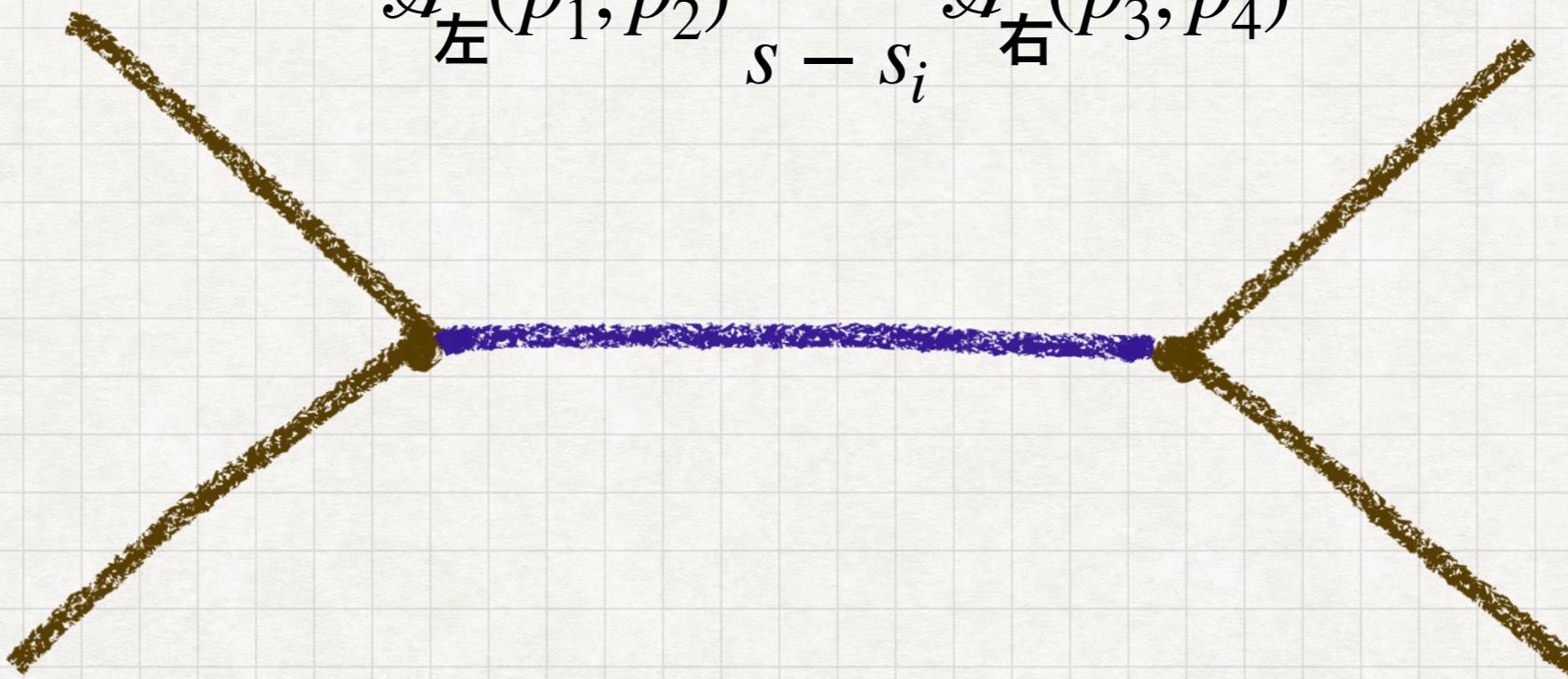


$$2\pi i \mathcal{A}(s, t) = - \oint_{\text{Poles}} \frac{\mathcal{A}(s', t)}{s' - s} ds' + \oint_{\infty} \frac{\mathcal{A}(s', t)}{s' - s} ds'$$

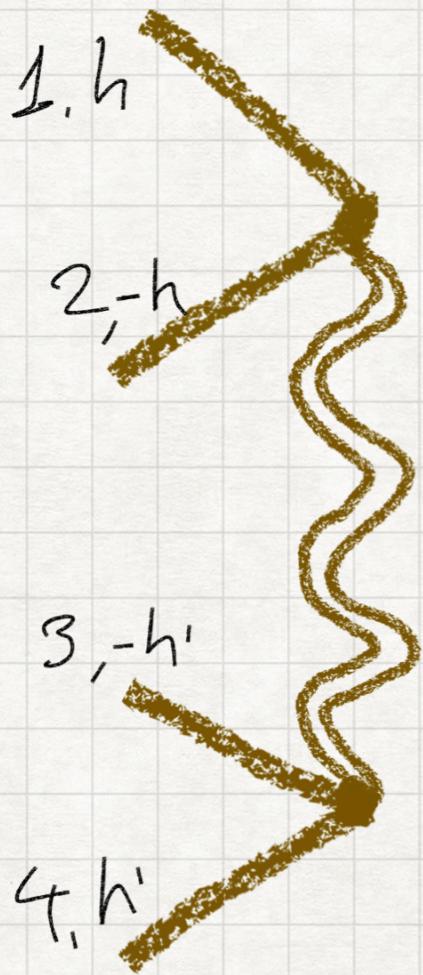
# On shell part of the amplitude

$$2\pi i \mathcal{A}(s, t) = - \oint_{\text{Poles}} \frac{\mathcal{A}(s', t)}{s' - s} ds' + B = - \frac{\text{Res}(\mathcal{A})_{s=s_i}}{s_i - s} + B$$

$$\mathcal{A}_{\text{左}}(p_1, p_2) \frac{1}{s - s_i} \mathcal{A}_{\text{右}}(p_3, p_4)$$



# 4 pt amplitude from 3 pt in Gravity



$$\frac{\sqrt{8\pi}}{M_{\text{pl}}} \frac{[1q]^{2+2h}[2q]^{2-2h}}{[12]^2}$$

$$\frac{1}{s_{12}}$$

$$\frac{\sqrt{8\pi}}{M_{\text{pl}}} \frac{\langle 3q \rangle^{2+2h'} \langle 4q \rangle^{2-2h'}}{\langle 34 \rangle^2}$$

Using on-shell relations like

$$|q\rangle[q| = \bar{\sigma}_\mu q^\mu = -\bar{\sigma}_\mu p_1^\mu - \bar{\sigma}_\mu p_2^\mu$$

$$\hat{P} = \sigma_\mu P^\mu$$

We obtain:

$$= \frac{8\pi}{M_{\text{pl}}^2} c_r \left( \frac{s_{14}}{s_{13}} \right)^r \frac{(s_{13}s_{14})^{1-h'}}{s_{12}} ([14]\langle 23 \rangle)^{2h} (\langle 3\hat{P}_{12}4 \rangle)^{2h'-2h}$$

$$h - 1 < r < 1 - h \quad h' - 1 < r < 1 - h'$$

# Ambiguity in fermions and scalars

-> Scalar

$$\frac{8\pi}{M_{\text{pl}}^2} \left( \frac{s_{13}s_{14}}{s_{12}} - as_{12} \right)$$



Contact terms

-> Fermion

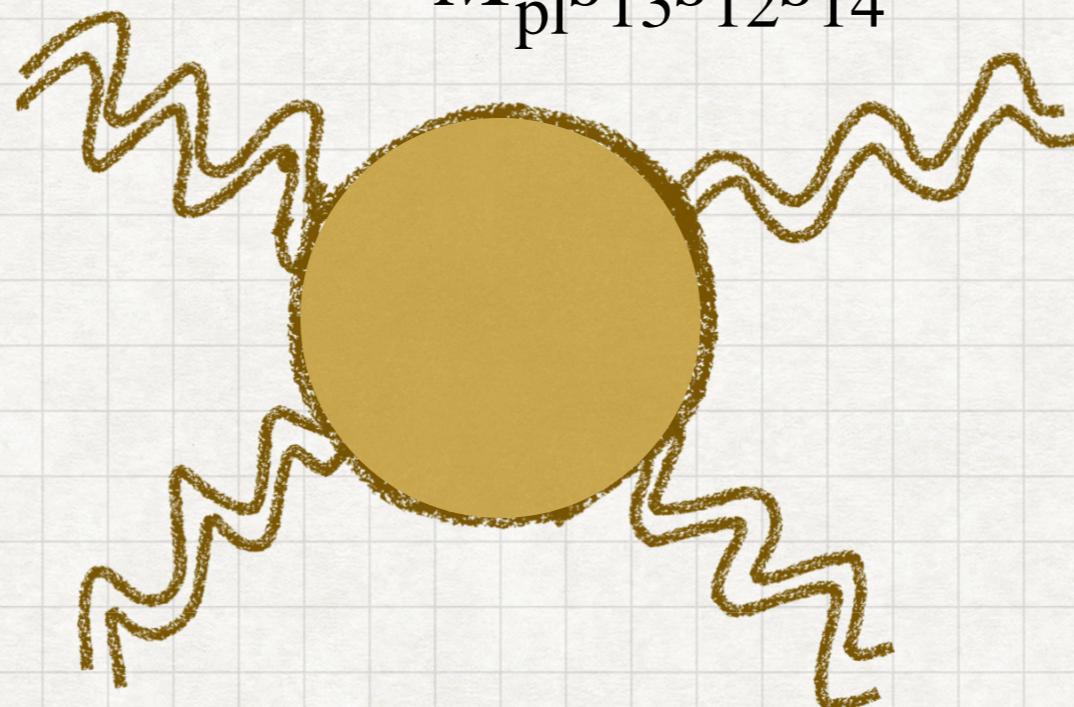
$$\frac{8\pi[14]\langle 23 \rangle}{M_{\text{pl}}^2} \left( \frac{s_{13}}{s_{12}} + \frac{b}{2} \right)$$

# We have computed more than we know!

$$\mathcal{A} = \frac{8\pi}{M_{\text{pl}}^2} \frac{(s_{13}s_{14})^{1-h'}}{s_{12}} ([14]\langle 23 \rangle)^{2h} (\langle 3\hat{P}_{12}4 \rangle)^{2h'-2h}$$

$$h = h' = 2$$

$$\frac{8\pi[14]^4\langle 23 \rangle^4}{M_{\text{pl}}^2 s_{13} s_{12} s_{14}}$$



The Feynman rule  
computation has  $\sim 1000$  terms!

[B. DeWitt, Phys. Rev. I62 I239, 1967]

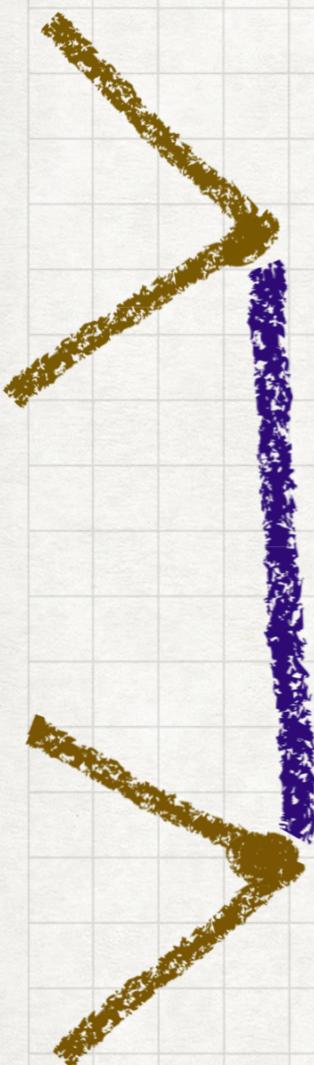
# These are all of them

| $\mathcal{A}_{1^h 2^{-h} 3^{-h'} 4^{h'}}$ | <b>Scalar</b>   | <b>Fermion</b>  | <b>Vector</b>   | <b>Graviton</b>   |
|---|---|---|---|---|
| <b>Scalar</b>                             | $\frac{8\pi}{M_{\text{Pl}}^2} \left( \frac{s_{13}s_{14}}{s_{12}} - as_{12} \right)$   | $\frac{8\pi \langle (3 \hat{P}_{12} 4) \rangle}{M_{\text{Pl}}^2} \left( \frac{s_{13}-s_{14}}{2s_{12}} \right)$              | $-\frac{8\pi \langle (3 \hat{P}_{12} 4) \rangle^2}{M_{\text{Pl}}^2 s_{12}}$                                       | $\frac{8\pi \langle (3 \hat{P}_{12} 4) \rangle^4}{M_{\text{Pl}}^2 s_{12} s_{13} s_{14}}$                                |
|   | $\frac{8\pi}{M_{\text{Pl}}^2} \left( \frac{s_{13}s_{14}}{s_{12}} + \frac{s_{12}s_{14}}{s_{13}} + \frac{s_{13}s_{12}}{s_{14}} \right)$ |   |   |   |
| <b>Fermion</b>                            |   | $-\frac{8\pi \langle (2 3)[1 4] \rangle}{M_{\text{Pl}}^2} \left( \frac{s_{13}}{s_{12}} + \frac{b}{2} \right)$               | $\frac{8\pi \langle (2 3)[1 4] \rangle \langle (3 \hat{P}_{12} 4) \rangle}{M_{\text{Pl}}^2 s_{12}}$               | $-\frac{8\pi \langle (2 3)[1 4] \rangle \langle (3 \hat{P}_{12} 4) \rangle^3}{M_{\text{Pl}}^2 s_{12} s_{13} s_{14}}$    |
|   |   | $-\frac{8\pi \langle (2 3)[1 4] \rangle}{M_{\text{Pl}}^2} \left( \frac{s_{13}}{s_{12}} + \frac{s_{12}}{s_{13}} + b \right)$ |   |   |
| <b>Vector</b>                             |   |   | $-\frac{8\pi \langle (2 3)^2[1 4] \rangle^2}{M_{\text{Pl}}^2 s_{12}}$   | $\frac{8\pi \langle (2 3)^2[1 4] \rangle^2 \langle (3 \hat{P}_{12} 4) \rangle^2}{M_{\text{Pl}}^2 s_{12} s_{13} s_{14}}$ |
|   |   |   | $-\frac{8\pi \langle (2 3)^2[1 4] \rangle^2}{M_{\text{Pl}}^2} \left( \frac{1}{s_{12}} + \frac{1}{s_{13}} \right)$ |   |
| <b>Graviton</b>                           |   |   |   | $\frac{8\pi \langle (2 3)^4[1 4] \rangle^4}{M_{\text{Pl}}^2 s_{12} s_{13} s_{14}}$                                      |

Maximal Helicity (non) Violation

$$\sum h_i = 0$$

# If we exchange a massive spin J? Legendre



$$gM \left( \frac{\langle q_I \hat{P}_{12} q_I \rangle}{M^2} \right)^J = \frac{1}{q^2 - M^2}$$

$|q^I\rangle [q_I] = \bar{\sigma}_\mu q^\mu = -\bar{\sigma}_\mu p_1^\mu - \bar{\sigma}_\mu p_2^\mu$

USING THIS —

$$\langle \mathbf{q}_{\text{I}_1} | \hat{P}_{12} | \mathbf{q}_{\text{I}_2} ] \times \langle \mathbf{q}_{\text{I}_3} | \hat{P}_{12} | \mathbf{q}_{\text{I}_4} ] \times \langle \mathbf{q}_{\text{I}_5} | \hat{P}_{12} | \mathbf{q}_{\text{I}_6} ] \times \dots$$

[                          ]                            [                          ]

$$\langle \mathbf{q}_{\text{I}_1} | \hat{P}_{34} | \mathbf{q}_{\text{I}_2} ] \times \langle \mathbf{q}_{\text{I}_3} | \hat{P}_{34} | \mathbf{q}_{\text{I}_5} ] \times \langle \mathbf{q}_{\text{I}_4} | \hat{P}_{34} | \mathbf{q}_{\text{I}_6} ] \times \dots$$

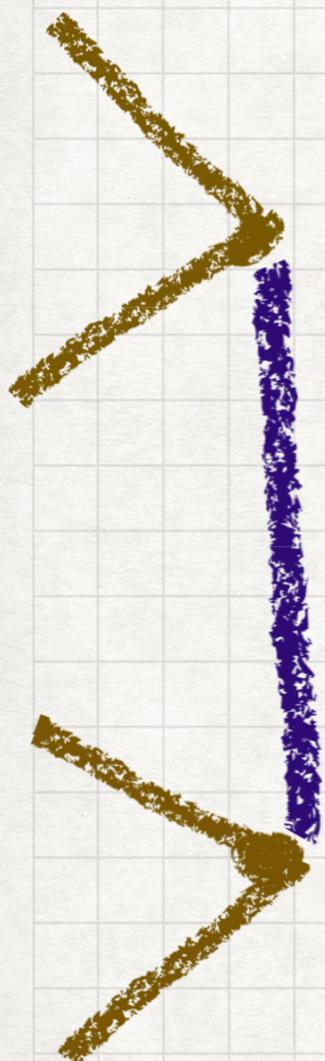
$$J = 2, \quad \frac{1}{4!} [ 8 \square\square + 16 \square\infty ]$$

$$J = 3, \quad \frac{3!2^3}{6!} \left[ (\square)^3 + 6 \square\infty + 8 \infty\infty \right]$$

$$J = 4, \quad \frac{4!2^4}{8!} \left[ (\square)^4 + 12 (\infty)^2 + 12 (\square)^2\infty + 32 \square\infty\infty + 48 \infty\infty\infty \right]$$

$$\text{tr}(\bar{\sigma}^\mu p_\mu \sigma^\mu k_\mu)^n = 2^n \sum_m \binom{n}{2m} (p \cdot k)^{n-2m} [(p \cdot k)^2 - p^2 k^2]^m.$$

If we exchange a massive spin J? Legendre



$$gM \left( \frac{\langle q_I \hat{P}_{12} q_I \rangle}{M^2} \right)^J = \frac{1}{q^2 - M^2}$$

$|q^I\rangle[q_I| = \bar{\sigma}_\mu q^\mu = -\bar{\sigma}_\mu p_1^\mu - \bar{\sigma}_\mu p_2^\mu$

USING THIS —

$$\frac{g^2(2J)!!}{(2J-1)!!} \frac{M^2}{s_{12} - M^2}$$

$$\times \frac{(2J-2m)!(-P_{12}^2 P_{34}^2)^m (P_{12} \cdot P_{34})^{J-2m}}{m!(J-m)!(J-2m)2^J M^{2J}}$$

$$= \frac{P_J(x)}{4^J}, \quad x = 1 + \frac{2s_{13}}{M^2}.$$

# If we exchange a massive spin J? Jacobi

1, h  
2, -h  
3, -h'  
4, h'

$$gM \frac{(\langle q_I \hat{P}_{12} q_I \rangle)^{J-2h} ([1q_I] \langle 2q_I \rangle)^{2h}}{M^{2J}}$$

$$\frac{1}{q^2 - M^2}$$

=

$$\frac{g^2(2J)!!(J-2h')!(J+2h')!}{(2J-1)!!J!^2} \left( \frac{[14]\langle 32 \rangle}{M^2} \right)^{2h} \left( \frac{\langle 3\hat{P}_{12}4 \rangle}{M^2} \right)^{2h'-2h} P_{J-2h'}^{(2h'-2h, 2h'+2h)}(x) \frac{M^2}{s_{12} - M^2}$$

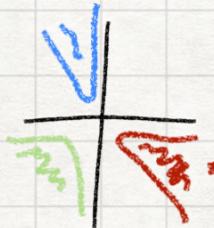
JACOBI POLYNOMIALS

$$P_n^{(a,b)}(x) \equiv \sum \binom{n+a}{n-k} \binom{n+b}{k} \left( \frac{x-1}{2} \right)^k \left( \frac{x+1}{2} \right)^{n-k}$$

$$x = 1 + \frac{2s_{13}}{M^2}.$$

## II. Unitarity

# Angular Analysis and unitarity? Wigner



$$s_{12} \rightarrow s, \quad s_{13} \rightarrow -s \sin^2(\theta/2), \quad \langle 13 \rangle \rightarrow \sqrt{s} \sin(\theta/2), \quad \dots$$

$$\mathcal{A} = \left( \frac{[14]\langle 32 \rangle}{M^2} \right)^{2h} \left( \frac{\langle 3\hat{P}_{12}4 \rangle}{M^2} \right)^{2h-2h} P_{J-2h'}^{(2h'-2h, 2h'+2h)}(x)$$

$$\left( \frac{s \cos^2(\theta/2)}{M^2} \right)^{2h} \left( \frac{s \sin(\theta/2) \cos(\theta/2)}{M^2} \right)^{2h-2h}$$

$$P_{J-2h'}^{(2h'-2h, 2h'+2h)}(1 - (1 - \cos \theta)s/M^2)$$

On-shell       $s = M^2$

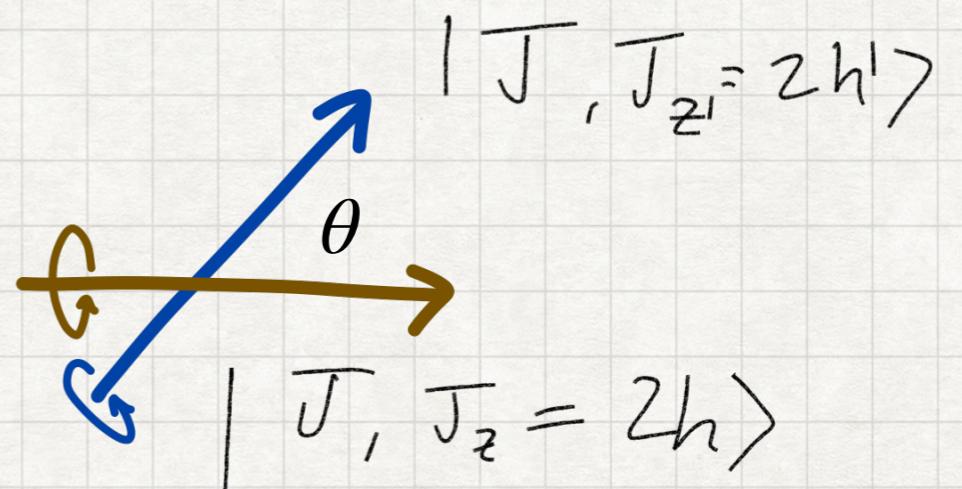
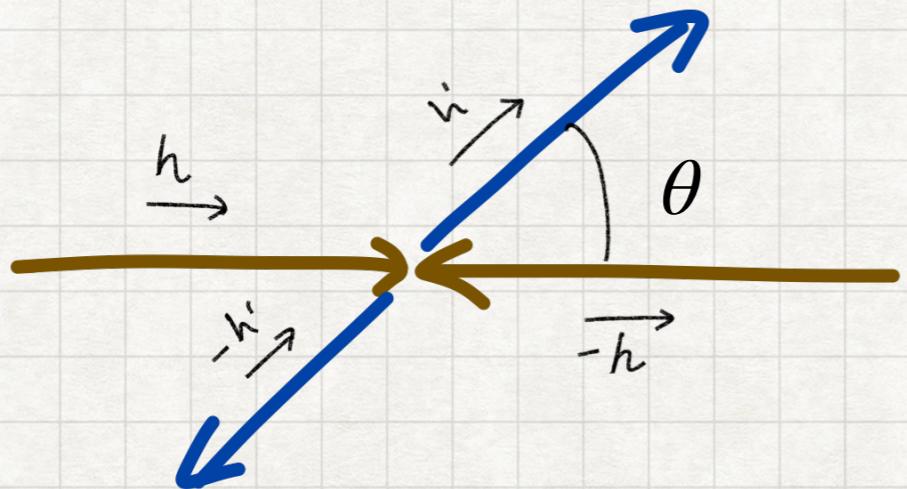
$$P_{J-2h'}^{(2h'-2h, 2h'+2h)}(\cos \theta)$$

# Wigner d-function and angular analysis

$$\sin^{2h'-2h}(\theta/2) \cos^{2h+2h'}(\theta/2) P_{J-2h'}^{(2h'-2h, 2h'+2h)}(\cos \theta)$$

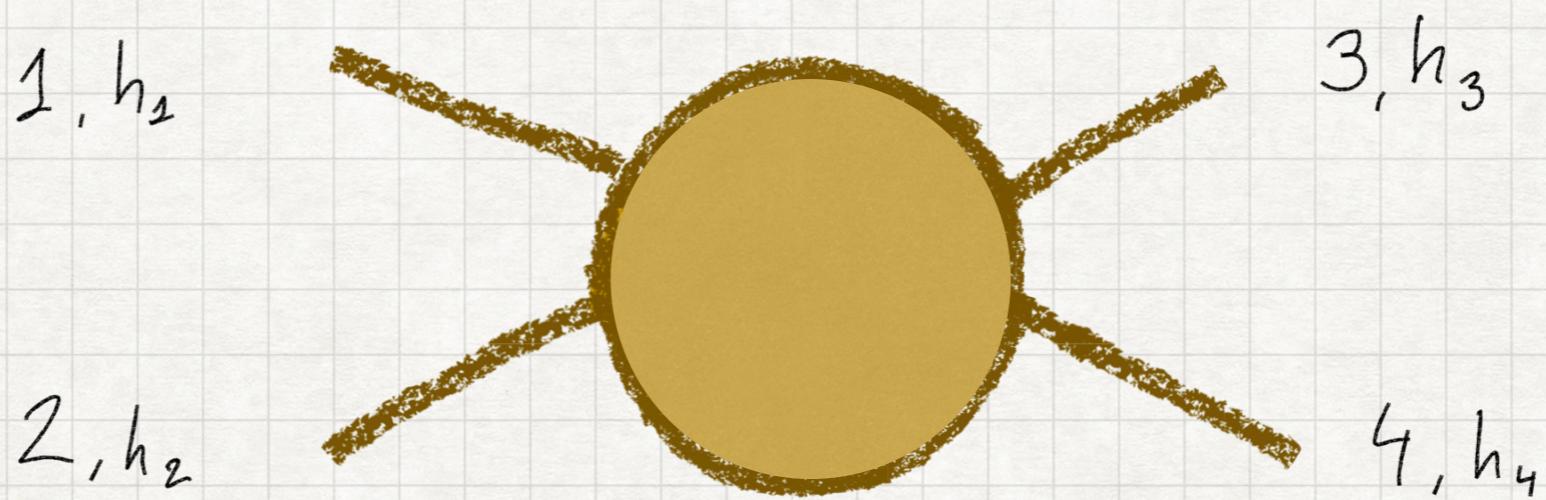
$$\propto d_{2h, 2h'}^J(\theta)$$

$$d_{m', m}^J = \langle J, m' | R(\theta) | J, m \rangle$$



# Wigner d-function and angular analysis

$$\mathcal{A} = \sum 16\pi(2J+1)a^J(s)d_{h_1-h_2,h_3-h_4}^J(\theta)$$



Partial wave expansion

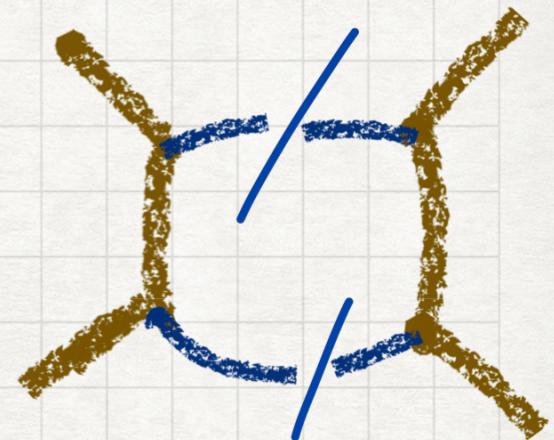
a's bounded to be less than one to conserve probability

$$a^J = \frac{1}{32\pi} \int d\cos(\theta) d_{h_1-h_2,h_3-h_4}^J(\theta) \mathcal{A}(s, \theta)$$

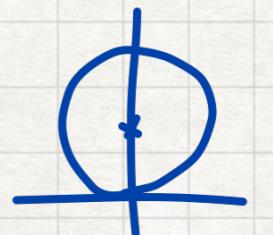
# ✳️ Unitarity ✳️

$$(1 + i\mathcal{A})(1 + i\mathcal{A})^\dagger = 1$$

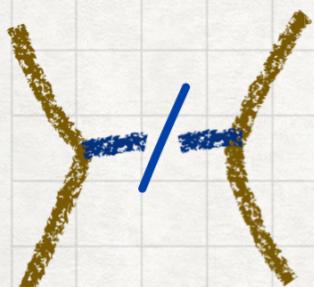
$$2 \operatorname{Im}(\mathcal{A}_{12 \rightarrow 34}) = \int \prod_i \frac{dp_i^3}{(2E_i)(2\pi)^3} \mathcal{A}_{34 \rightarrow n}^* \mathcal{A}_{12 \rightarrow n}$$



If  $12=34$  and we do angular dec.  $\operatorname{Im}(a^J) \leq 1$ ,  $\operatorname{Re}(a^J) \leq 1/2$ .



For a single particle intermediate state  $n$



$$2 \operatorname{Im}(\mathcal{A}_{12 \rightarrow 34}) = \pi \mathcal{A}_{34 \rightarrow J}^*(\theta) \mathcal{A}_{12 \rightarrow J}(\theta) \delta(s - M^2)$$

$$\frac{1}{s - M^2 + i\epsilon} = \mathcal{PV} \left[ \frac{1}{s - M^2} \right] - i\pi \delta(s - M^2)$$

# Unitarity

$$(1 + i\mathcal{A})(1 + i\mathcal{A})^\dagger = 1$$

$$2 \operatorname{Im}(\mathcal{A}_{12 \rightarrow 34}) = \pi \mathcal{A}_{34 \rightarrow J}^*(\theta) \mathcal{A}_{12 \rightarrow J}(\theta) \delta(s - M^2)$$

If  $12=34$

$$2 \operatorname{Im}(\mathcal{A}_{12 \rightarrow 12}) = \pi |\mathcal{A}_{12 \rightarrow J}(\theta)|^2 \delta(s - M^2)$$

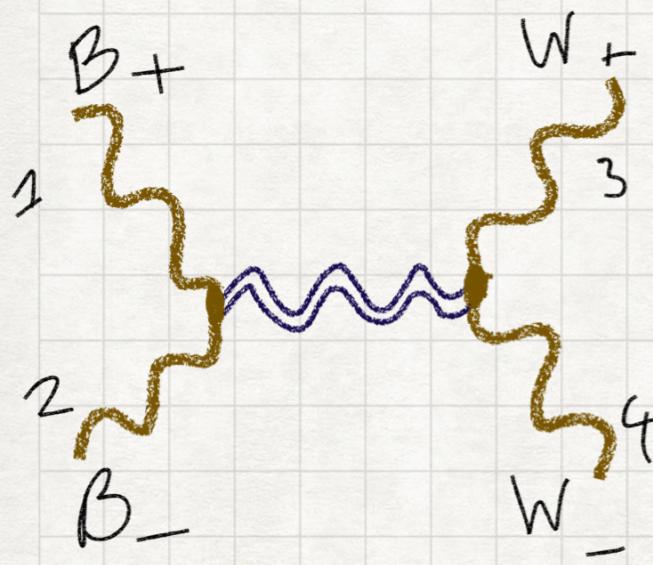
Positivity

If in the forward direction

$$2 \operatorname{Im}(\mathcal{A}_{12 \rightarrow 34})(\theta = 0) = 16\pi(2J + 1)M\Gamma_{J \rightarrow 12} \delta(s - M^2)$$

# III. Gravity in the ultraviolet

# Gravity in the ultraviolet

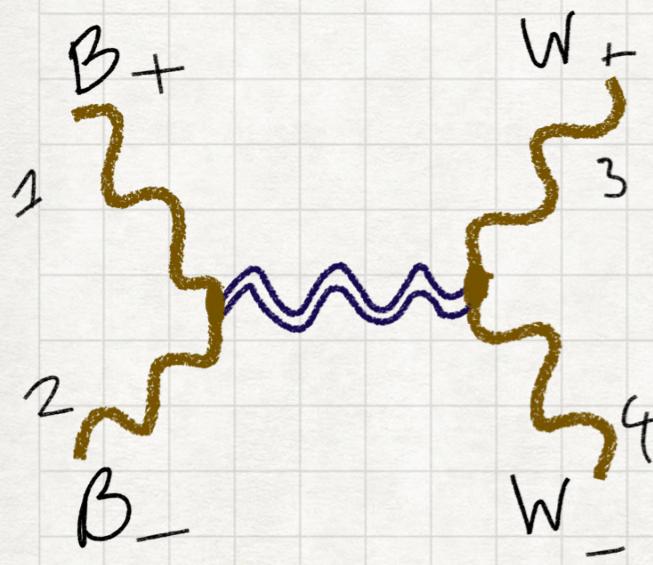


$$\mathcal{A}_{B,W} = -\frac{8\pi}{M_{\text{pl}}^2} \frac{\langle 32 \rangle^2 [41]^2}{s_{12}}$$

Second partial  
wave coeff.

$$a_{\text{GR}}^2 = \frac{s}{10M_{\text{pl}}^2}$$

# Gravity in the ultraviolet



$$\mathcal{A}_{B,W} = -\frac{8\pi}{M_{\text{pl}}^2} \frac{\langle 32 \rangle^2 [41]^2}{s_{12}}$$

Second partial  
wave coeff.

$$a_{\text{GR}}^2 = \frac{s}{10M_{\text{pl}}^2}$$

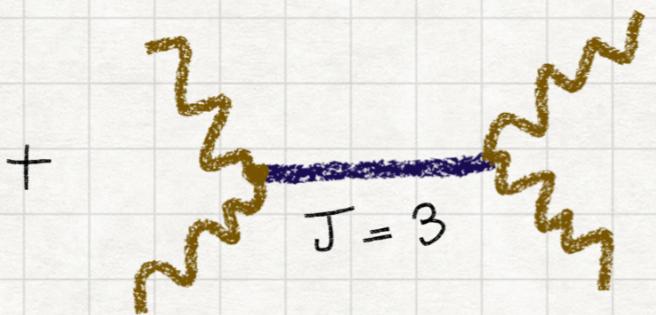


$$\mathcal{A}_{B,W}^{\text{GR}} + \mathcal{A}_{B,W}^{J=2} = \langle 23 \rangle^2 [14]^2 \left( \frac{8\pi}{M_{\text{pl}}^2 s} + \frac{g_B g_W P_0^{(0,4)}(x)}{M_2^2 (s - M_2^2)} \right)$$

$$g_B g_W = -8\pi \frac{M_2^2}{M_{\text{pl}}^2}$$

But opposite sign does not extend  
to a third species, e.g. gluon

# Gravity in the ultraviolet



$$\mathcal{A}_{B,W}^{\text{GR}} + \mathcal{A}_{B,W}^{J=3} = \langle 23 \rangle^2 [14]^2 \left( \frac{8\pi}{sM_{\text{pl}}^2} + \frac{g_B g_W P_1^{(0,4)}(x)}{M_3^2(s - M_3^2)} \right)$$

$$1 + 6t/M_3^2$$



$$a_{\text{GR}}^2 = \frac{s}{10M_{\text{pl}}^2} - \frac{g_B g_W s^2}{80M_3^4}$$

# Gravity in the ultraviolet

$$\mathcal{A}_{B,W}^{\text{GR}} + \sum_J \mathcal{A}_{B,W}^J = \langle 23 \rangle^2 [14]^2 \left( \frac{8\pi}{sM_{\text{pl}}^2} + \sum_J \frac{C_J P_J^{(0,4)}(1 + 2t/M_J^2)}{s - M_J^2} \right)$$

$$\equiv \frac{\langle 23 \rangle^2 [14]^2}{sM_{\text{pl}}^2} \frac{\prod_n (t - f_n(s))}{\prod_i (s - M_i^2)}$$

Solving for  $f_n$  is not straightforward so we will make  
a number of assumptions 

# Gravity in the ultraviolet

$$\frac{\langle 23 \rangle^2 [14]^2}{sM_{\text{pl}}^2} \frac{\prod (t - f_n(s))}{\prod (s - M_i^2)} \Big|_{s=M_i} \propto \prod_{n=1}^{\infty} (t - f_n(M_i^2))$$

 *f<sub>n</sub> analytic around M*

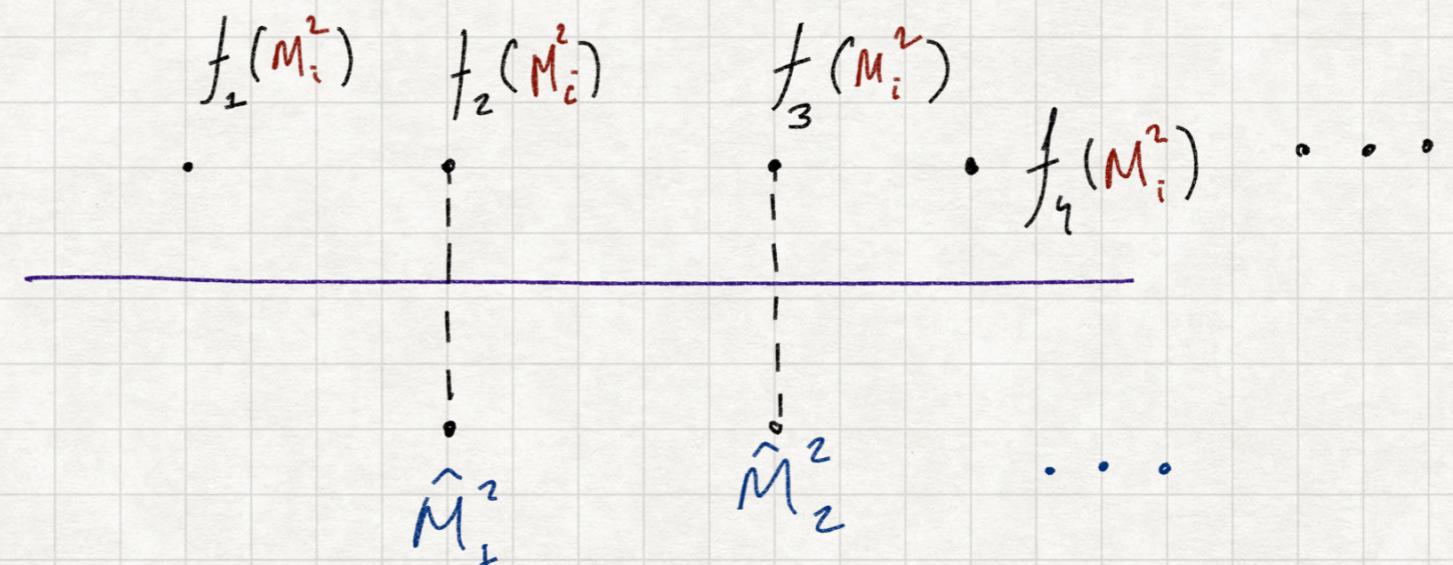
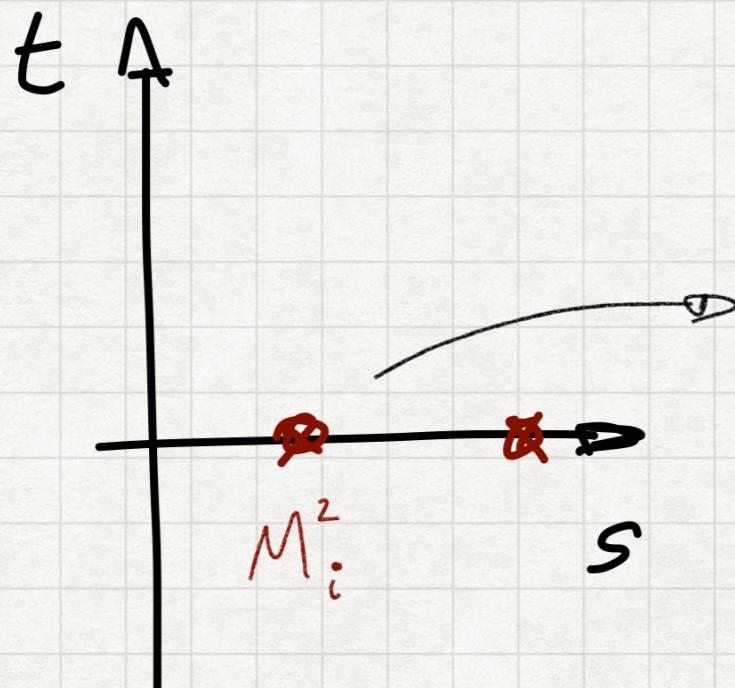
We don't want an infinite degree (infinite spin)

So we introduce inverse powers of t

$$\frac{\langle 23 \rangle^2 [14]^2}{sM_{\text{pl}}^2} \frac{\prod (t - f_n(s))}{\prod (s - M_i^2) \prod (t - \hat{M}_j^2)}$$

# Gravity in the ultraviolet

$$\frac{\langle 23 \rangle^2 [14]^2}{s M_{\text{pl}}^2} \frac{\prod (t - f_n(s))}{\prod (s - M_i^2) \prod (t - \hat{M}_j^2)} \Big|_{s=M_i^2} \propto \frac{\prod (t - f_n(M_i^2))}{\prod (t - \hat{M}_i^2)}$$

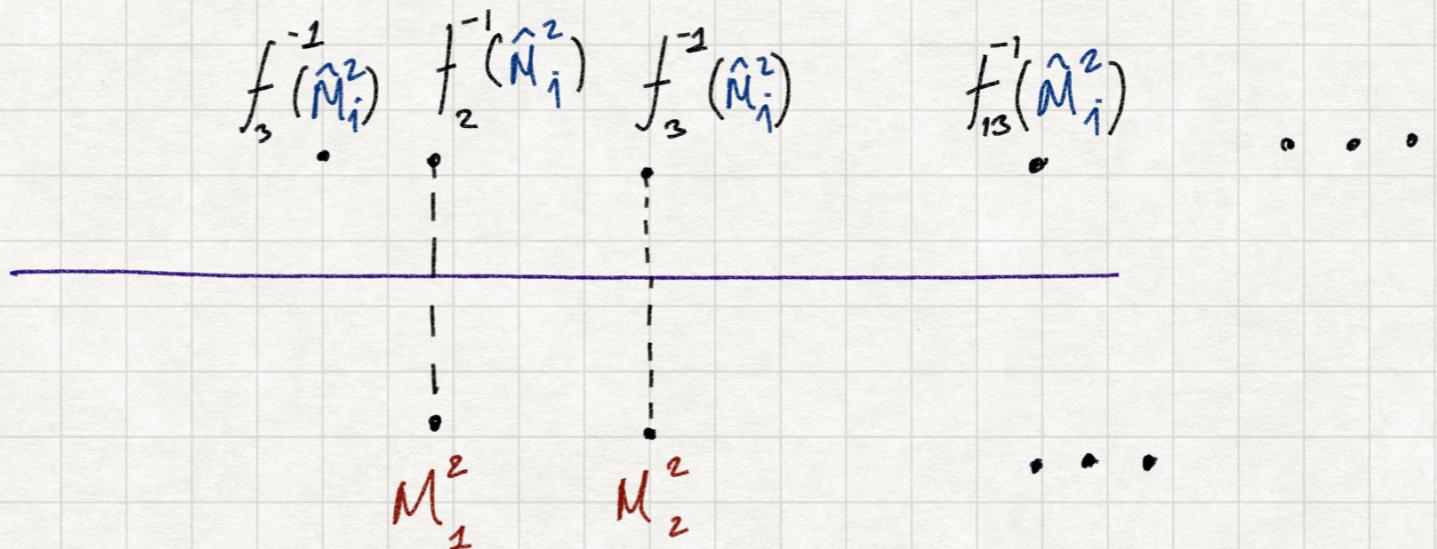
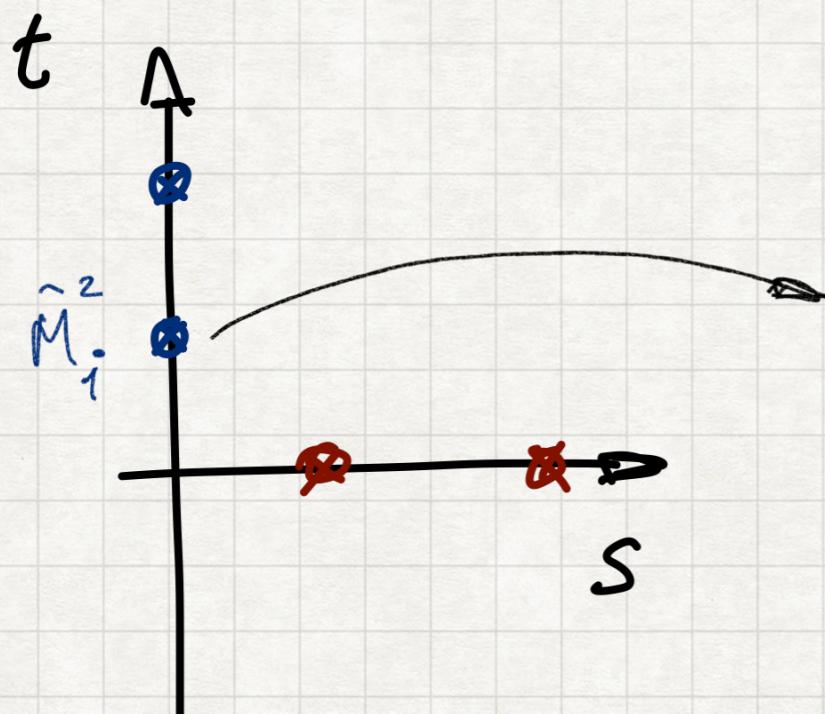


$$\{f_n(s_i)\} \supset \{\hat{M}_j^2\}, \quad \forall i.$$

The zeroes contain  
the poles

# Gravity in the ultraviolet

$$\frac{\langle 23 \rangle^2 [14]^2}{s M_{\text{pl}}^2} \frac{\prod (t - f_n(s))}{\prod (s - M_i^2) \prod (t - \hat{M}_j^2)} \Big|_{t=\hat{M}_j^2} \propto \frac{\prod (\hat{M}_j^2 - f_n(s))}{\prod (s - M_i^2)}$$

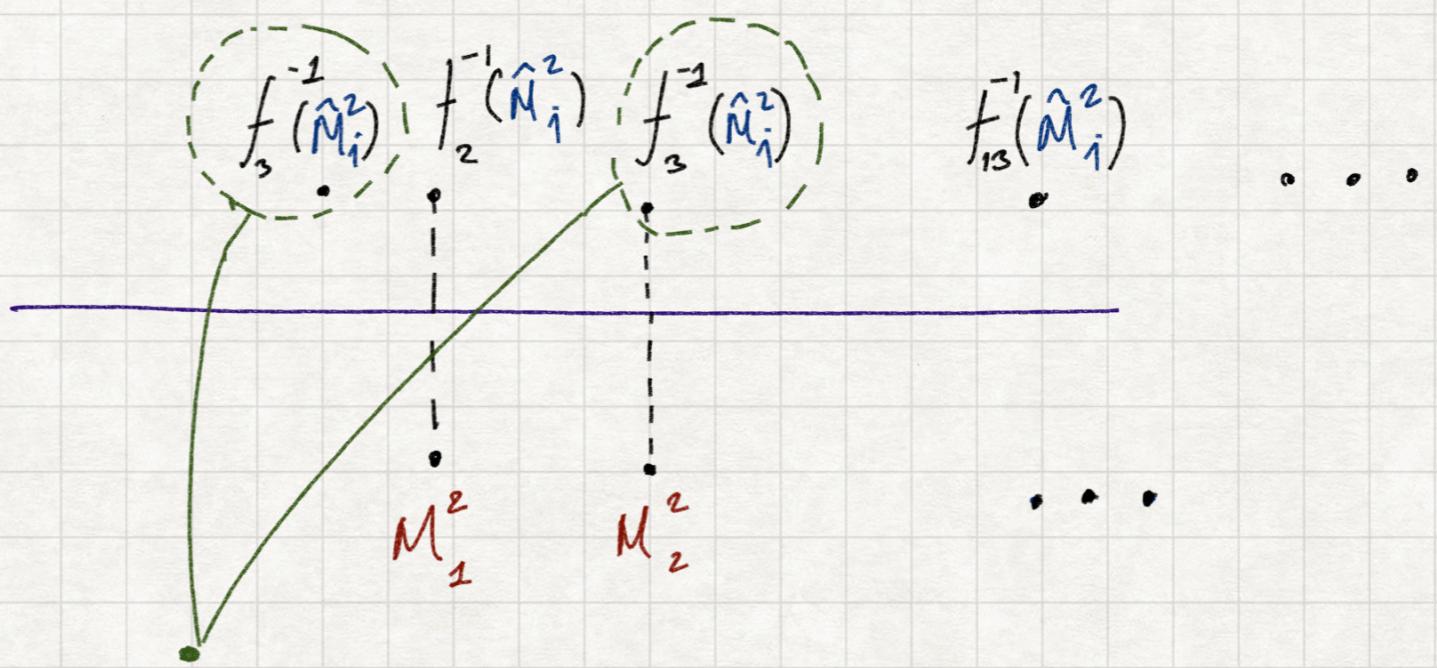
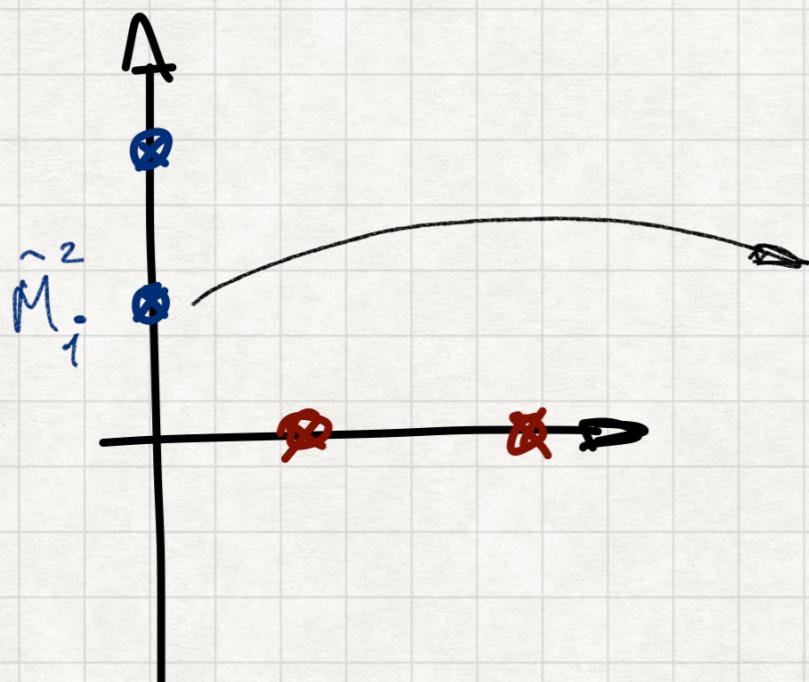


$$\{f_n^{-1}(t_j)\} \supset \{M_i^2\}, \quad \forall j.$$

The zeroes contain  
the poles (also in s)

# Gravity in the ultraviolet

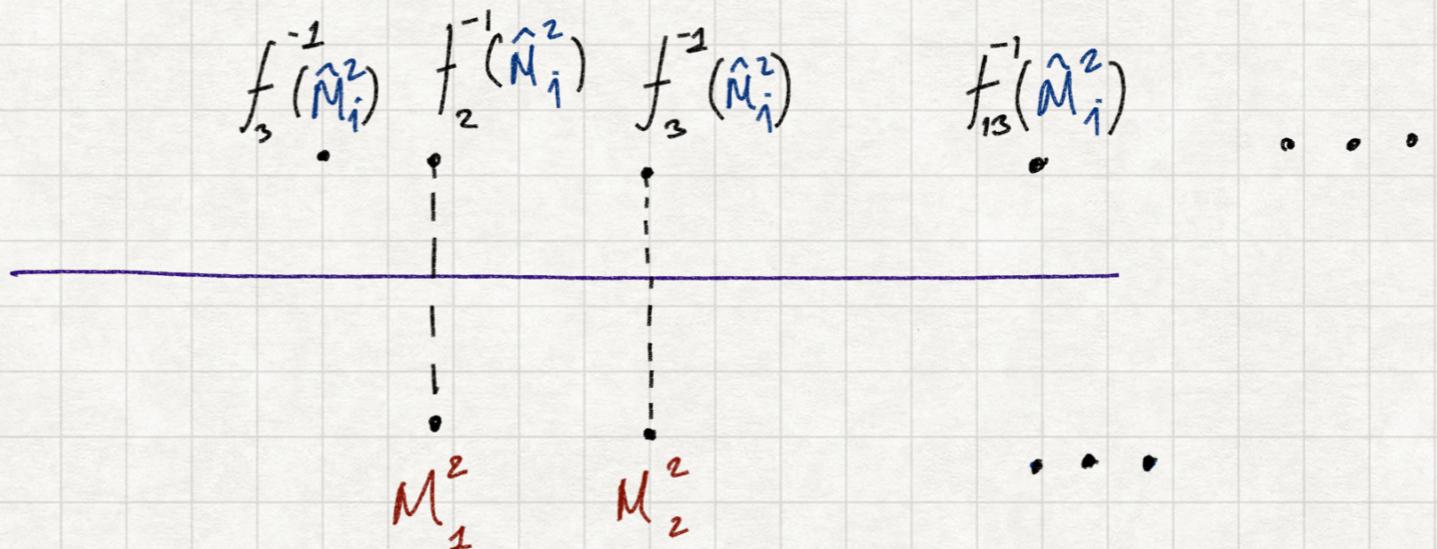
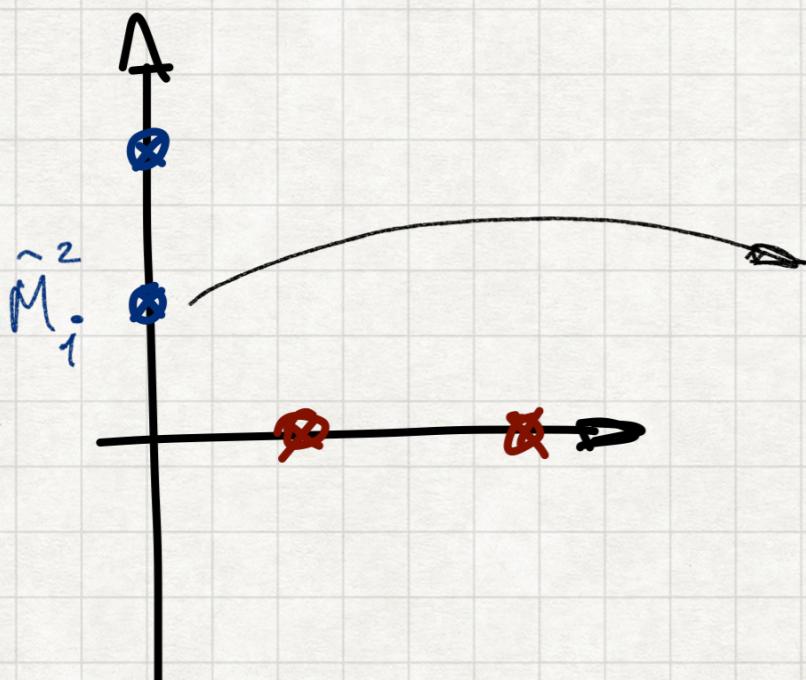
$$\frac{\langle 23 \rangle^2 [14]^2}{s M_{\text{pl}}^2} \frac{\prod (t - f_n(s))}{\prod (s - M_i^2) \prod (t - \hat{M}_j^2)} \Big|_{t=\hat{M}_j^2} \propto \frac{\prod (\hat{M}_j^2 - f_n(s))}{\prod (s - M_i^2)}$$



But how many zeroes for  
the inverse of  $f_n$ ?

# Gravity in the ultraviolet

$$\frac{\langle 23 \rangle^2 [14]^2}{s M_{\text{pl}}^2} \frac{\prod (t - f_n(s))}{\prod (s - M_i^2) \prod (t - \hat{M}_j^2)} \Big|_{t=\hat{M}_j^2} \propto \frac{\prod (\hat{M}_j^2 - f_n(s))}{\prod (s - M_i^2)}$$



Assume  $f_n^{-1}$  are single valued

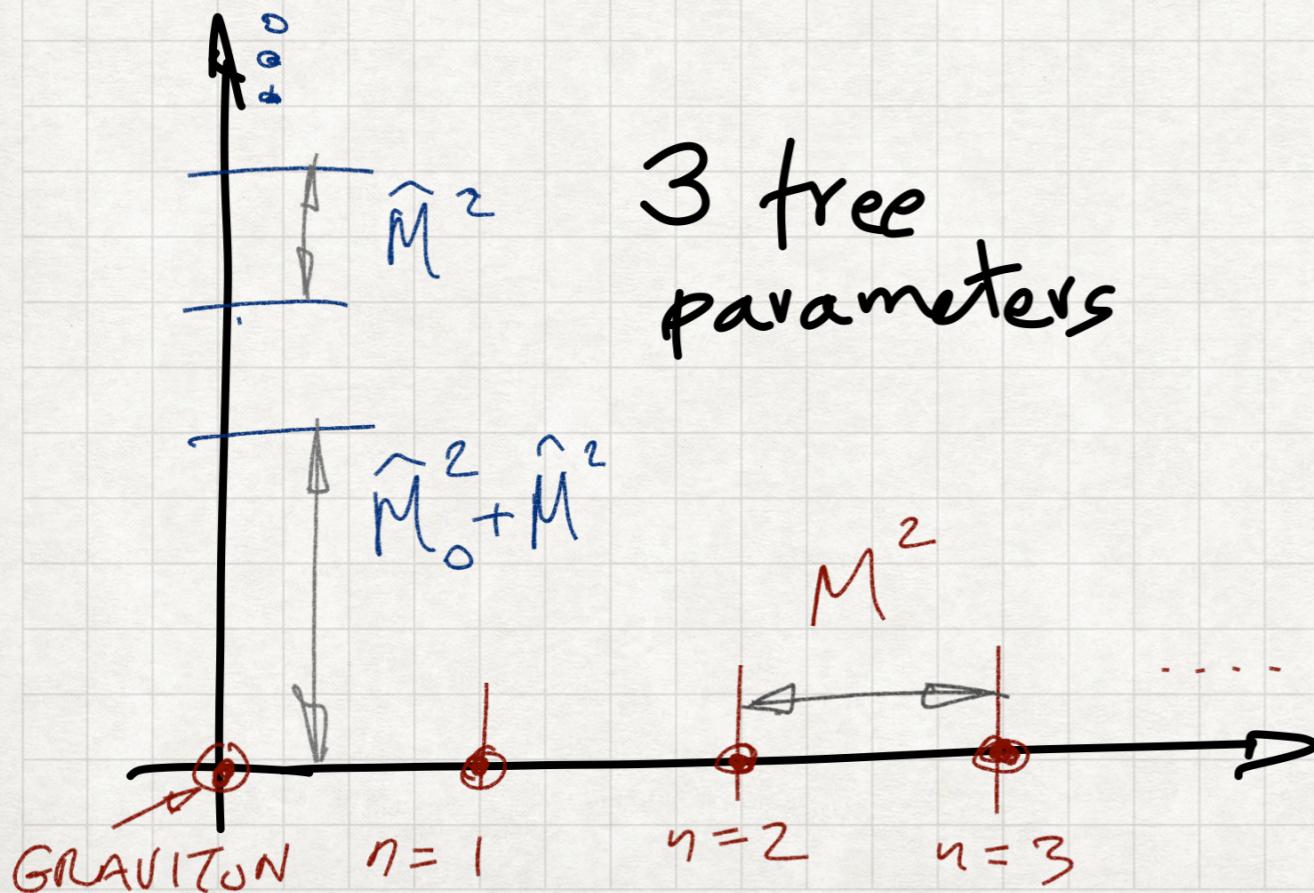
→ linear  $f_n = f'_n s + f_{n,0}$

# Gravity in the ultraviolet



Finally we assume the spin (degree of polynomial in  $t$ ) of resonances increases with  $n$

$$\frac{\langle 23 \rangle^2 [14]^2}{s M_{\text{pl}}^2} \prod_n \frac{M^2 t + \hat{M}^2 s - M^2(n \hat{M}^2 + \hat{M}_0^2)}{(s - n M^2)(t - \hat{M}_0^2 - n \hat{M}^2)}, \quad n \in \mathbb{N}.$$



Recalling the following definition of the Gamma function

$$\Gamma(z) = \frac{1}{z} \prod \frac{(1 + 1/n)^z}{(1 + z/n)}$$

# Gravity in the ultraviolet

$$\mathcal{A} = \frac{8\pi\langle 23\rangle^2[14]^2}{M_{\text{pl}}^2 s} \frac{\Gamma(1 - \tilde{s})\Gamma(1 - \hat{t})}{\Gamma(1 - \tilde{s} - \hat{t})}$$

$$\tilde{s} = s/\textcolor{red}{M}^2, \quad \hat{t} = (t - \hat{M}_0^2)/\hat{M}^2$$

This looks familiar but... did we solve the problem we set out to?

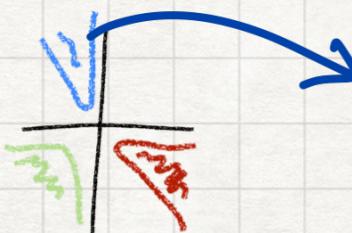
$$\mathcal{A} \rightarrow s e^{R\tilde{s}} \quad R = \log \left( (1 - \eta s_{\theta/2}^2)^{\eta s_{\theta/2}^2} (\eta s_{\theta/2}^2)^{\eta s_{\theta/2}^2} \right)$$

$\eta \equiv \textcolor{red}{M}^2/\hat{M}^2$  for  $\eta$  less (or =) than one it decays exponentially!

# IV. Analysis

# Positivity from unitarity

The  $t$  resonances we found are accessible in the crossed process which furthermore is subject to positivity



$$s_{12} \rightarrow t, \quad [14] \rightarrow \sqrt{s} \cos(\theta/2) \quad \text{etc}$$

$$\mathcal{A} = \frac{8\pi \langle 23 \rangle^2 [14]^2}{s_{12} M_{\text{pl}}^2} \frac{\Gamma(1 - \eta \tilde{s}_{13}) \Gamma(1 - \tilde{s}_{12})}{\Gamma(1 - \tilde{s}_{12} - \eta \tilde{s}_{13})}$$

Make (blue)  
substitutions  
& evaluate @ poles

$$\eta \tilde{s}_{13} = n$$

$$\alpha_{4,4}^J = \frac{1}{32\pi} \int dc_\theta d_{4,4}^J(\theta) \text{Res}(\mathcal{A}(s, \theta))_{s=M_n^2}$$

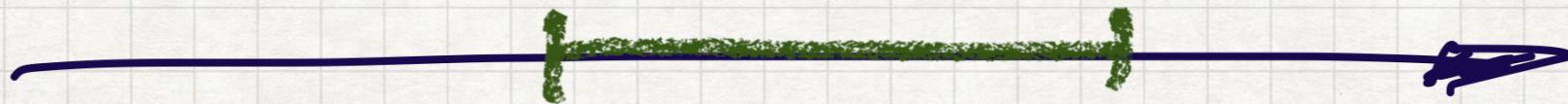
$$\alpha^J \geq 0 \Rightarrow (\eta^{-1} - 1) \leq \frac{3}{2n} \quad n = 1, 2, \dots \Rightarrow M = \hat{M}$$



# Fermions are special

$$\frac{8\pi\langle 23\rangle[14]}{M_{\text{pl}}^2 s_{12}} \left( \frac{s_{13}}{s_{12}} + \frac{s_{12}}{s_{13}} + b \right) \frac{\Gamma(1 - \tilde{s}_{12})\Gamma(1 - \tilde{s}_{13})}{\Gamma(1 - \tilde{s}_{12} - \tilde{s}_{13})}$$

$$\frac{2}{3} \leq b \leq \frac{22}{5}$$

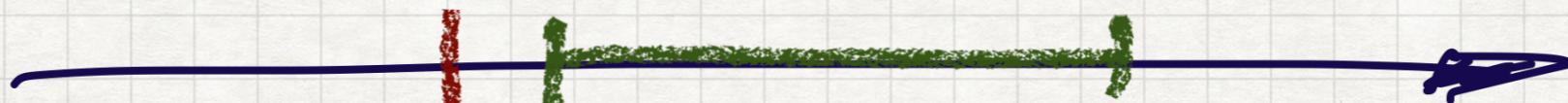


But  $b$  we can compute from Feynman rules of GR!

# Fermions are special

$$\frac{8\pi \langle 23 \rangle [14]}{M_{\text{pl}}^2 s_{12}} \left( \frac{s_{13}}{s_{12}} + \frac{s_{12}}{s_{13}} + b \right) \frac{\Gamma(1 - \tilde{s}_{12}) \Gamma(1 - \tilde{s}_{13})}{\Gamma(1 - \tilde{s}_{12} - \tilde{s}_{13})}$$

$$\frac{2}{3} \leq b \leq \frac{22}{5}$$



But  $b$  we can compute from Feynman rules of GR!

$$b = 1/2 !$$

# Modify the low energy content of gravity

We can fix this introducing a 3-form  $H$

$$-\frac{g}{M} \epsilon^{\mu\nu\rho\sigma} H_{\mu\nu\rho} \psi^\dagger \sigma_\mu \psi + (H^2)$$

Which nonetheless is not dynamical but it integrates out to

$$\frac{g^2}{M^2} (\psi^\dagger \sigma_\mu \psi)^2$$

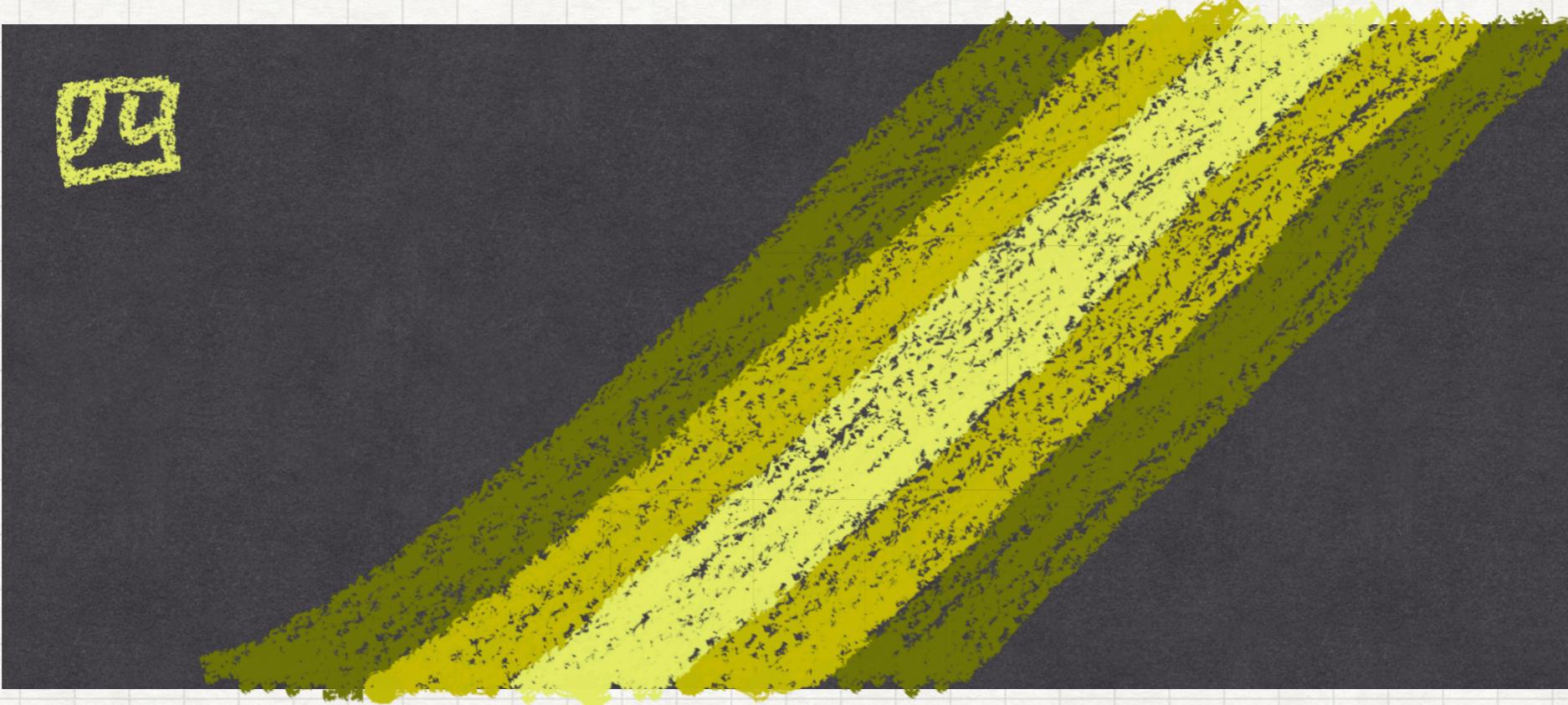
and the amplitude we obtain  
is reconciled with the low  
energy EFT for the range

$$\frac{1}{108} \leq \frac{g^2 M_{\text{pl}}^2}{\pi M^2} \leq \frac{13}{60}$$

# Summary

A sample of how amplitude methods  
Provide a new angle to approach gravity

... and quite remarkably give predictions ^-^



Analysis of results