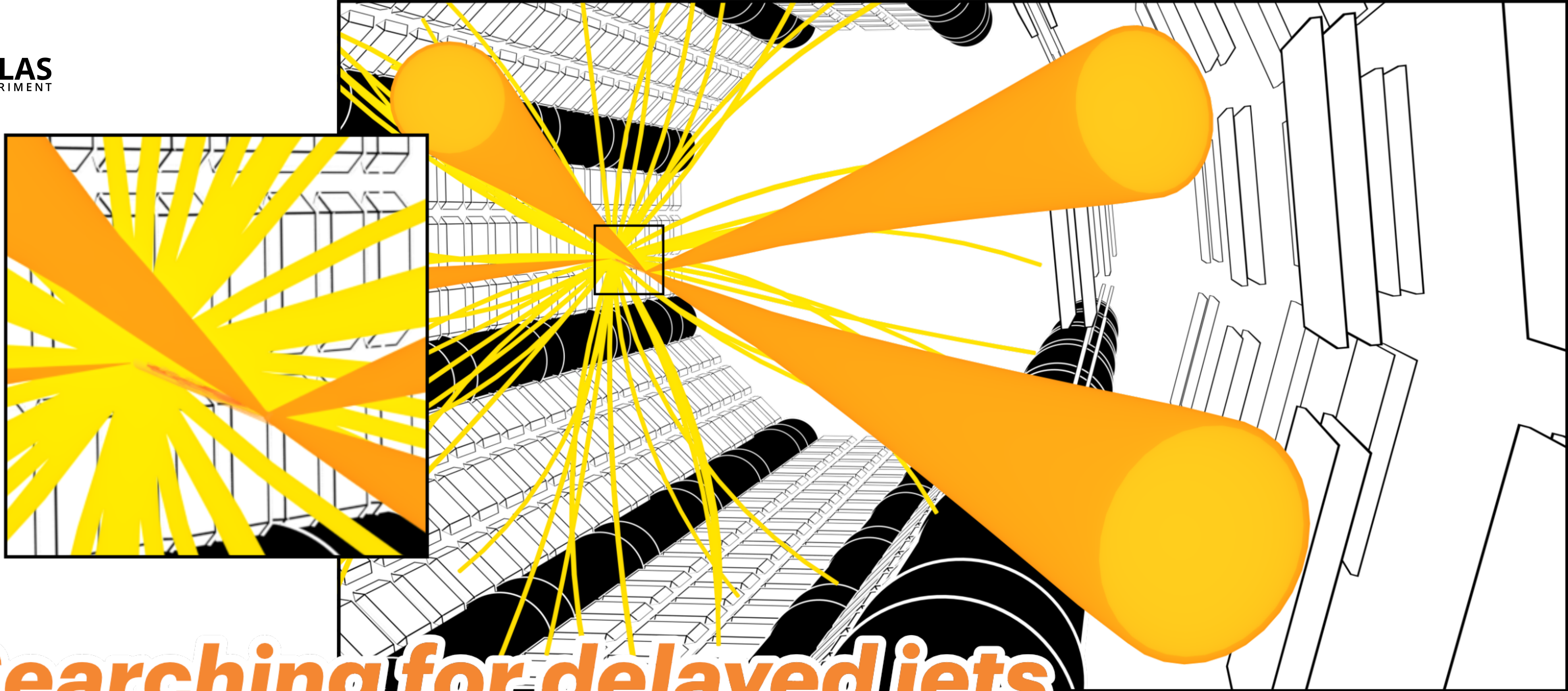




Searching for delayed jets from long-lived particles with the ATLAS detector

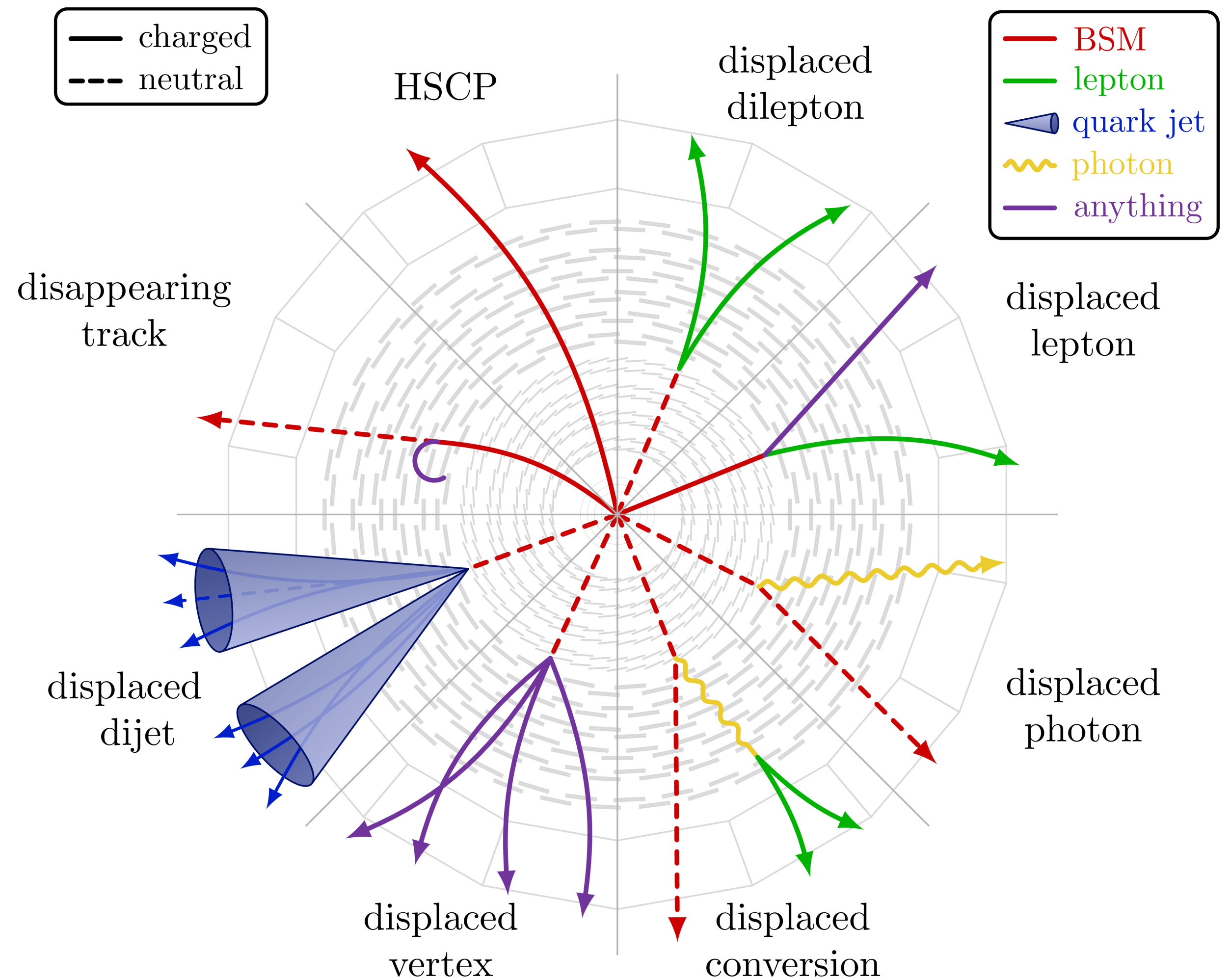
Renske Wierda

*Partikeldagarna
17/09/2026*

Long-lived particles

Long-lived particles (LLPs):
particles with finite lifetimes that decay noticeably away from the interaction point (IP)

Many searches target the displacement, e.g. an invisible LLP decaying to displaced vertices, leptons/photons or jets



Credit: Izaak Neutelings on tikz.net

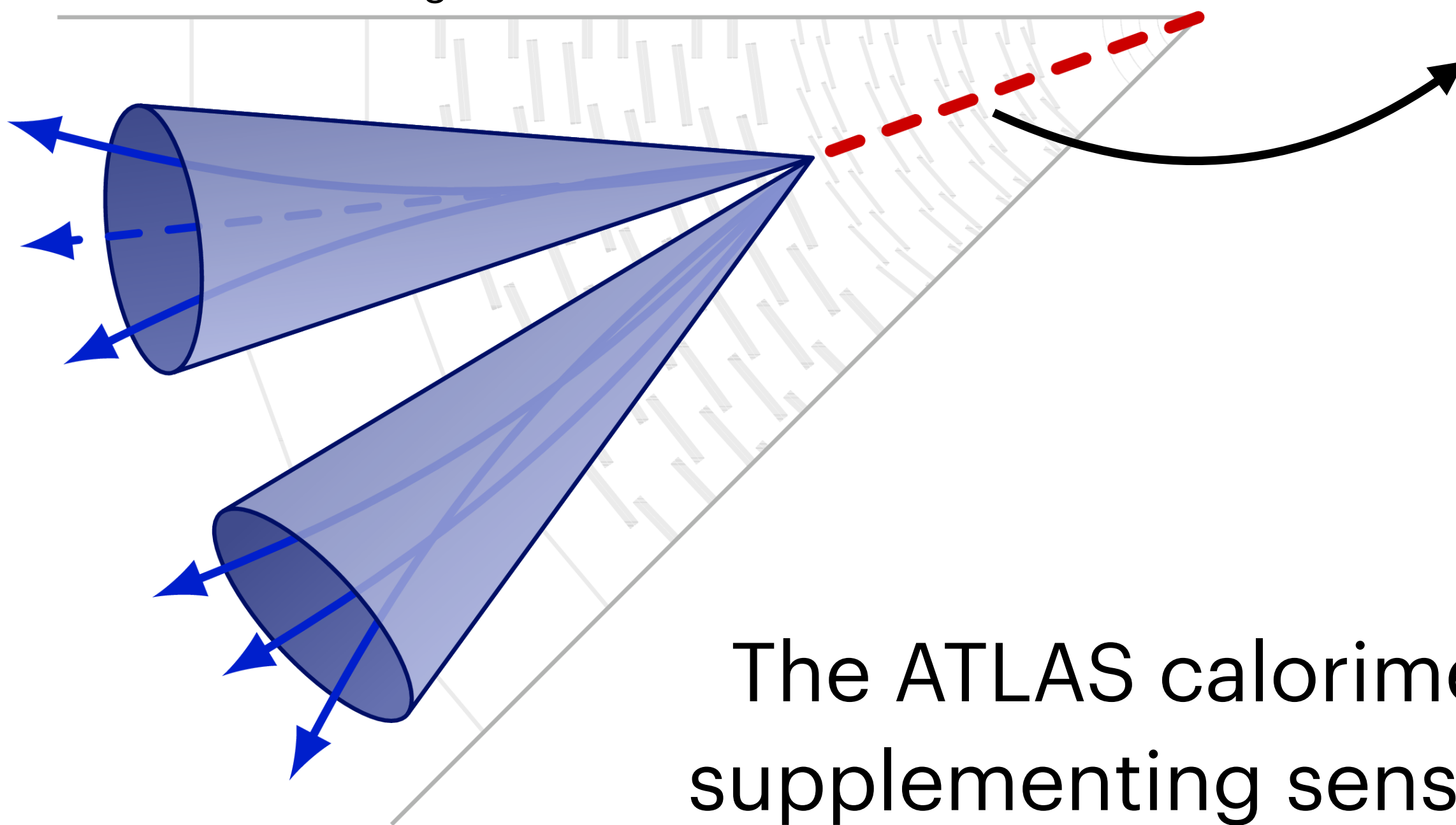
From displacement to delay

The time-of-arrival of a particle is calculated w.r.t. the speed of light, so a photon produced exactly at the IP arrives anywhere in the detector at $t = 0$

From displacement to delay

The time-of-arrival of a particle is calculated w.r.t. the speed of light, so a photon produced exactly at the IP arrives anywhere in the detector at $t = 0$

Credit: Izaak Neutelings on tikz.net



If LLP is low mass ($\beta \approx 1$), it's solely a displaced jet

However, if it's high mass ($\beta < 1$), it's ***both displaced and delayed***

The ATLAS calorimeters have good timing capabilities, supplementing sensitivity in a region where sensitivity to displacement is more limited

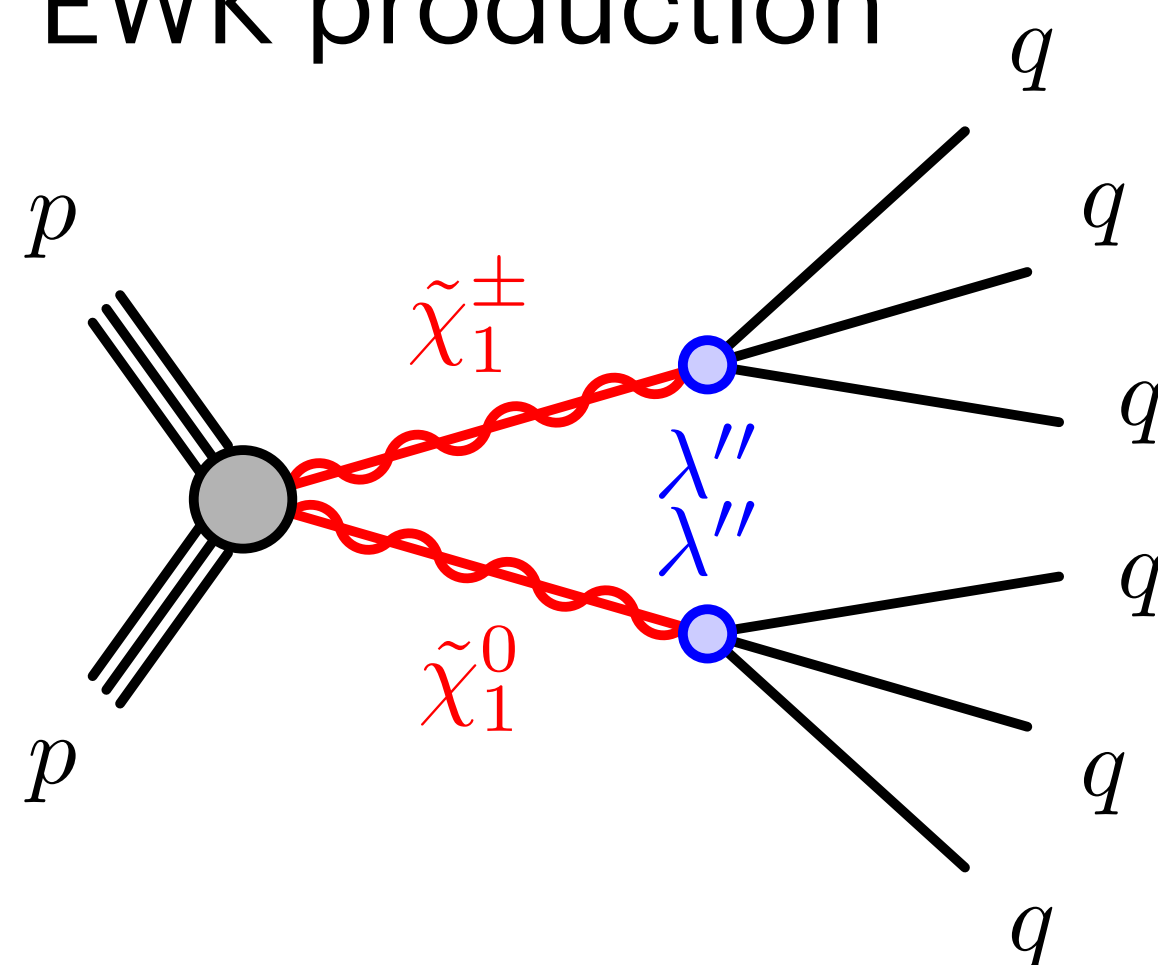
Signal models

Generally: we target models with heavy LLPs that decay into quarks before or inside the calorimeter systems

Specifically: we use three different signal models for benchmarking

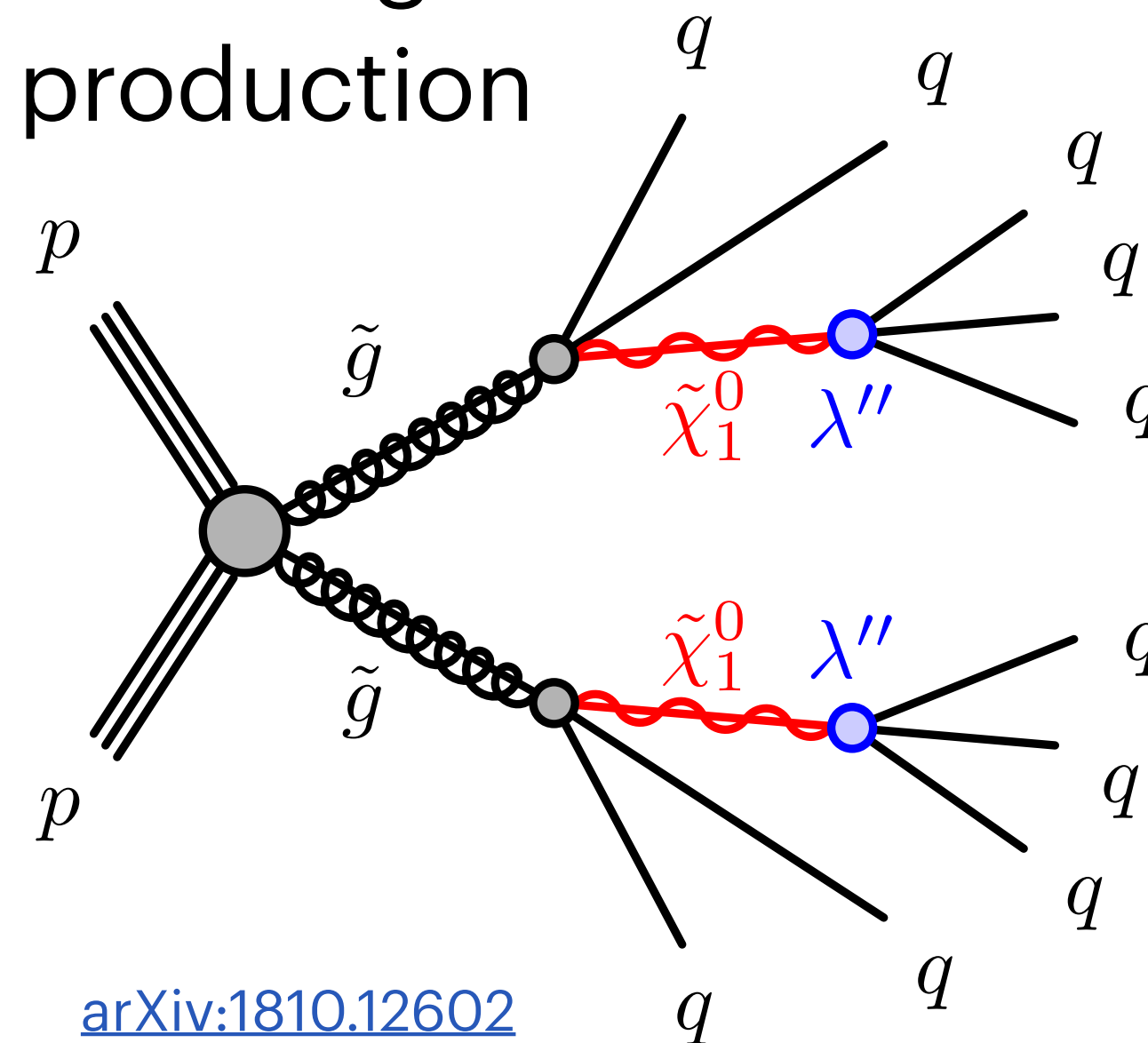
R-parity violating (RPV) neutralinos

EWK production



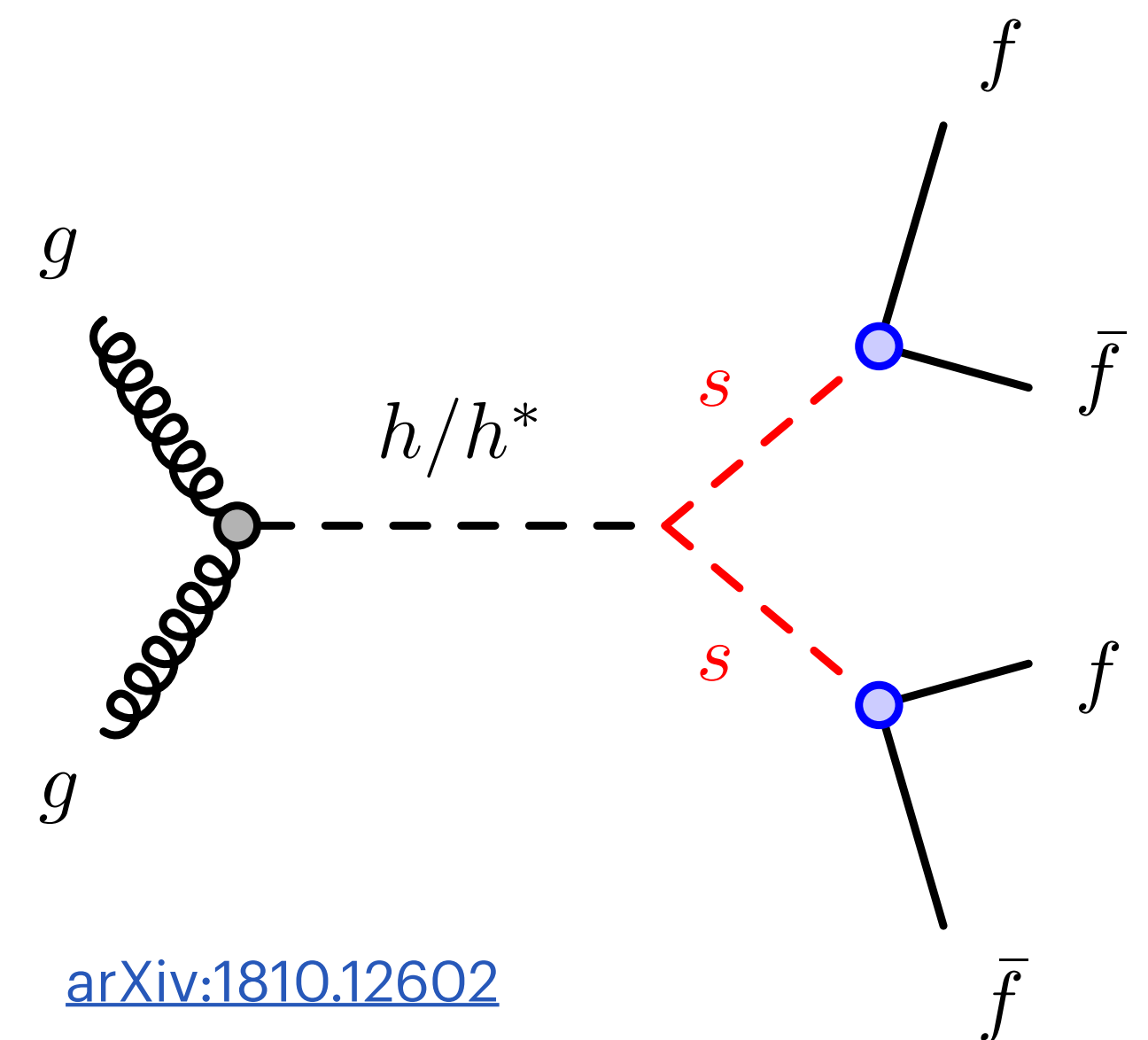
[arXiv:2301.13866](https://arxiv.org/abs/2301.13866)

Strong production



[arXiv:1810.12602](https://arxiv.org/abs/1810.12602)

Higgs to heavy scalar (HSS)

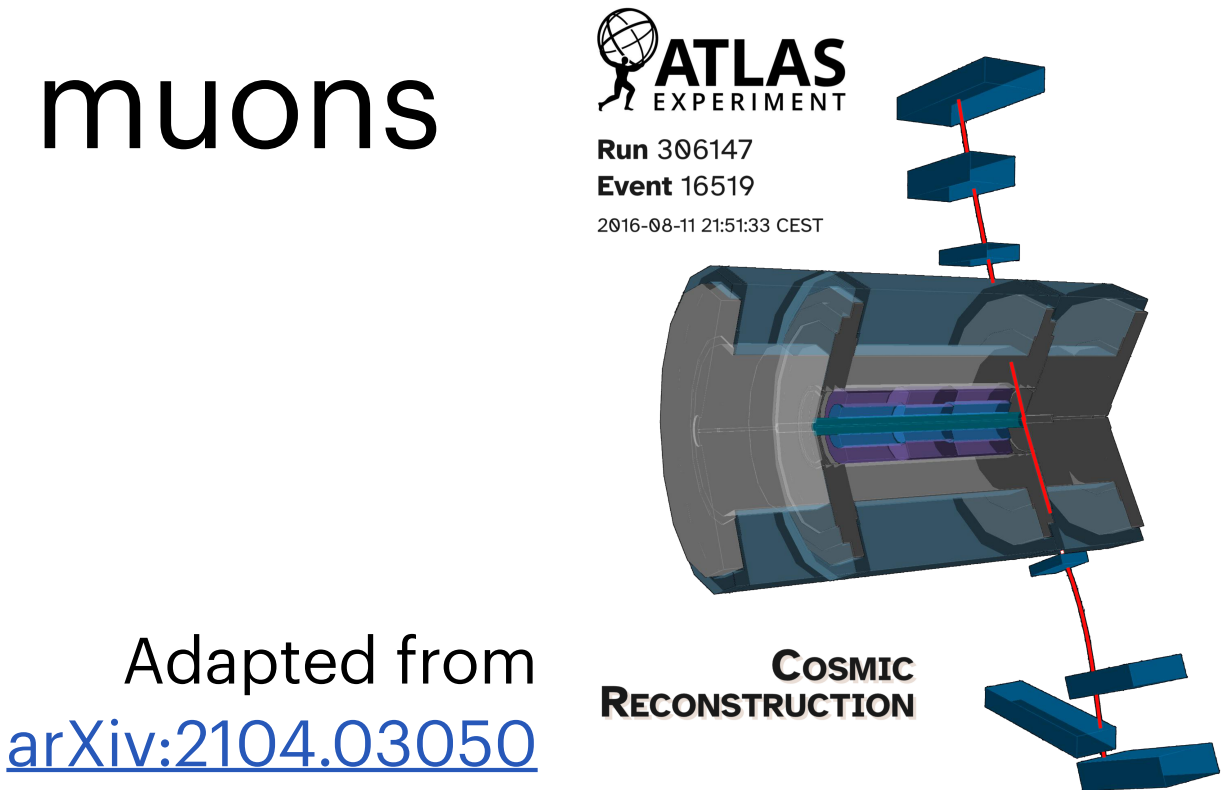


[arXiv:1810.12602](https://arxiv.org/abs/1810.12602)

Background overview

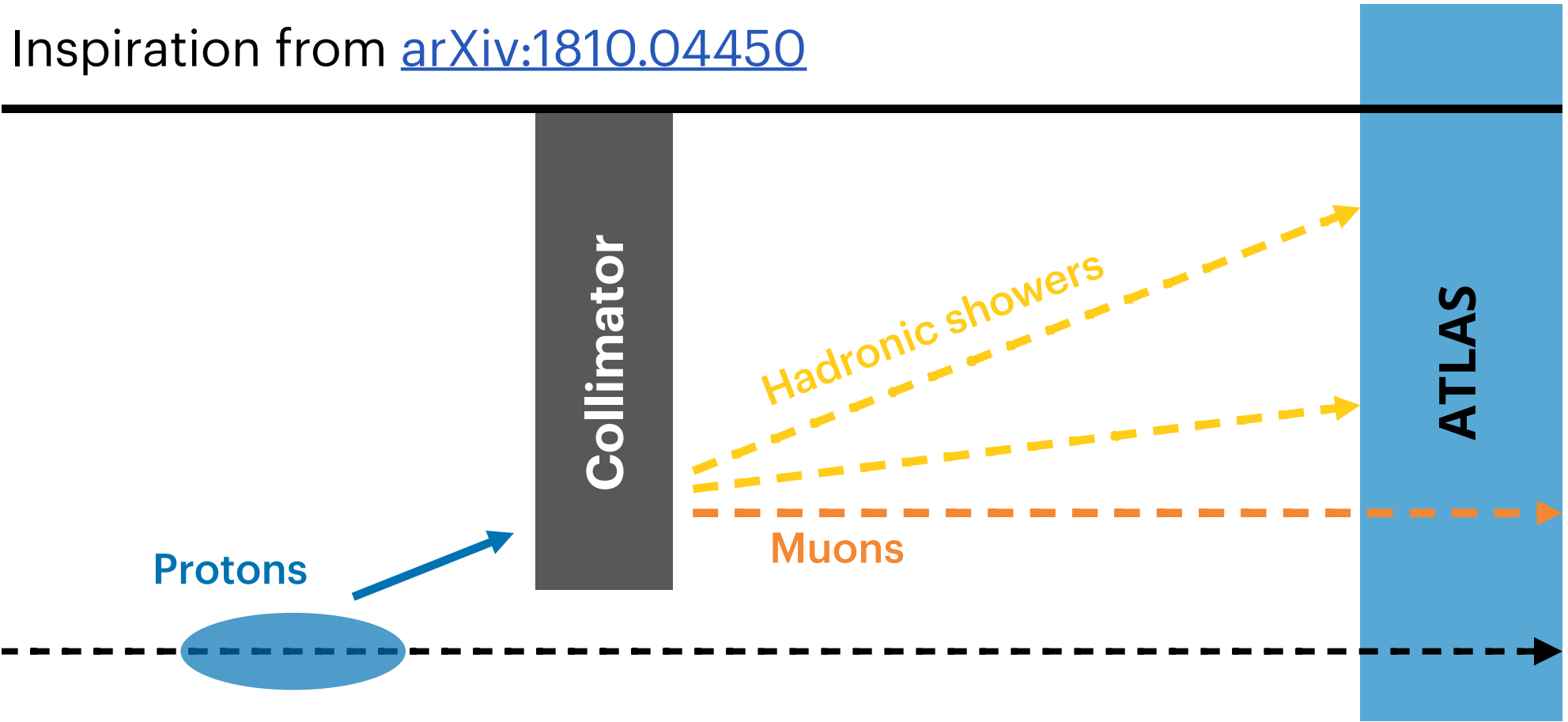
Non-collision backgrounds

Cosmic muons



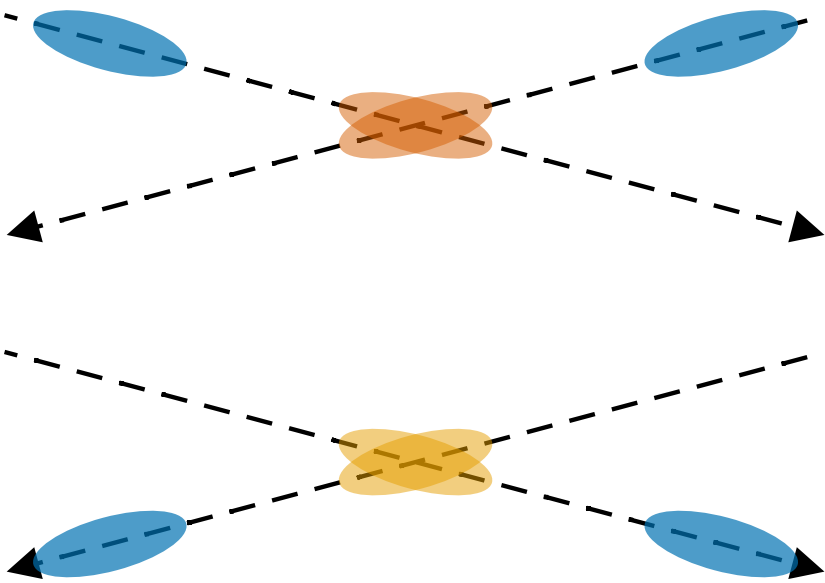
Beam-induced background (BIB)

Inspiration from [arXiv:1810.04450](https://arxiv.org/abs/1810.04450)

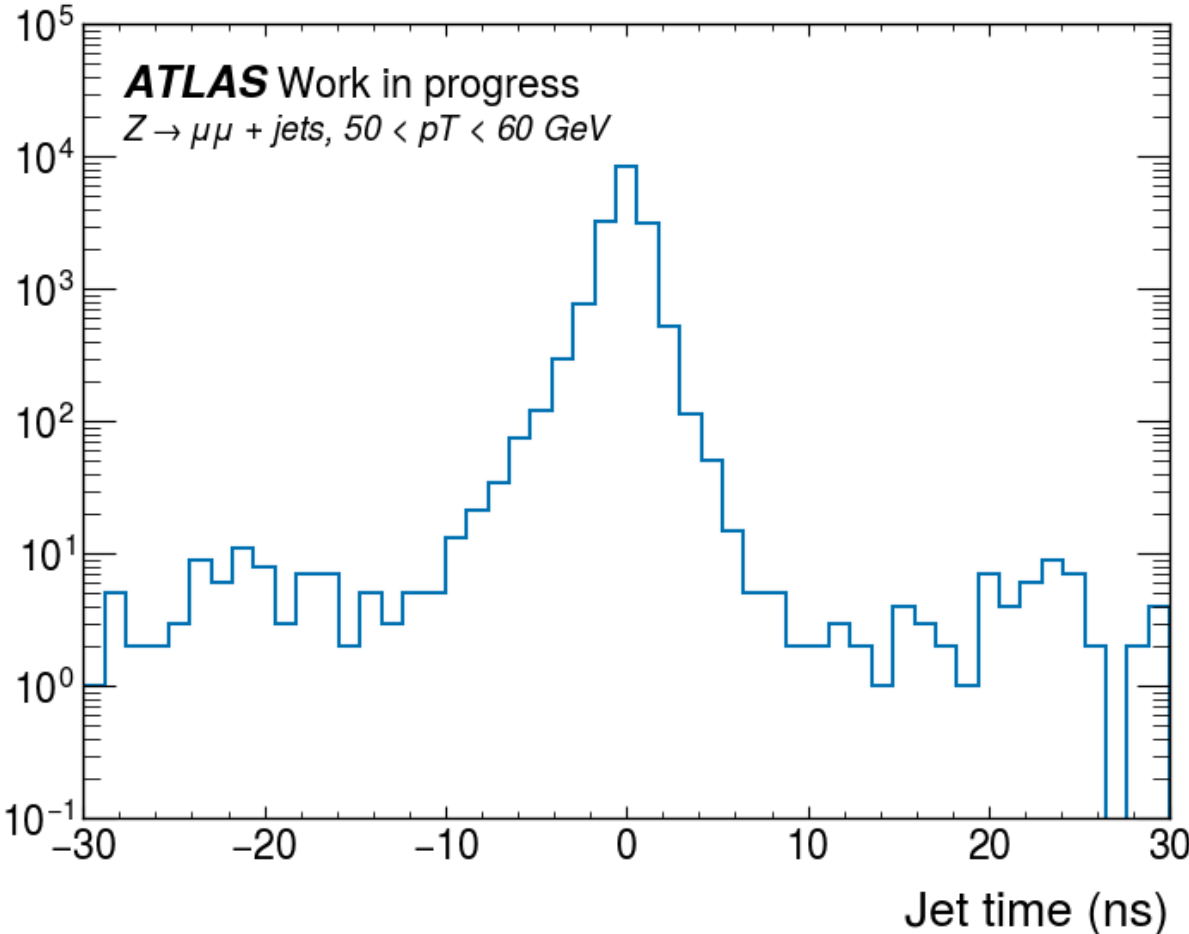


Collision backgrounds

Satellite collisions



Mis-timed SM jets

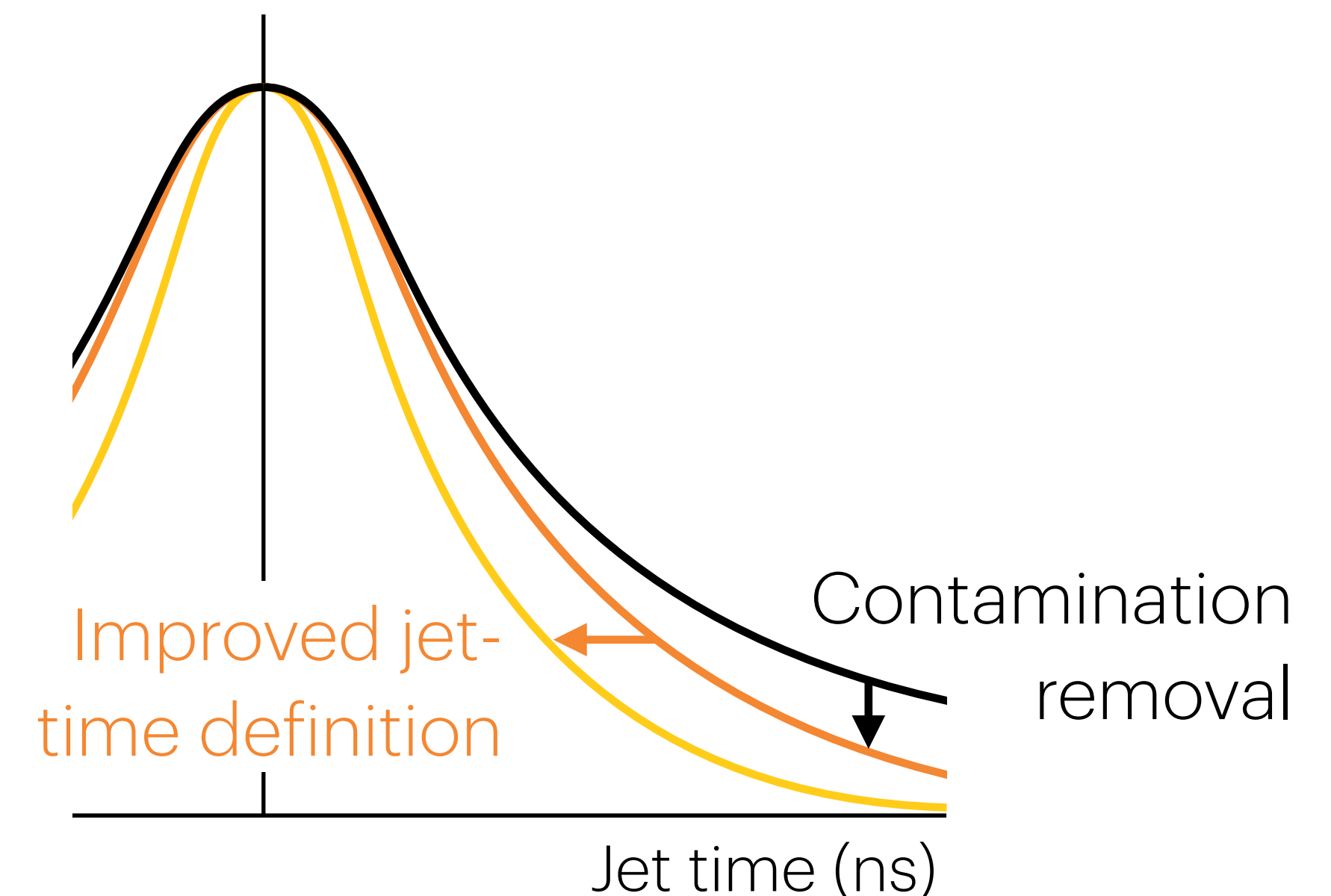


Analysis overview

Goal: low-background search for high- p_T and significantly-delayed jets

Backgrounds are not modelled in MC, so all require data-driven methods:

- **Cosmic muons:** custom tag to remove cosmons based on muon segments
- **BIB:** custom tag to remove BIB based on calorimeter cell positions
- **Mis-timed SM jets:**
 - which I also work on!
 - 1. Custom tag to remove contaminated jets (including anomaly detection R&D)
 - 2. Improved jet-timing calculation compared to ATLAS default (including GNN R&D)
- **Satellites:** up next!

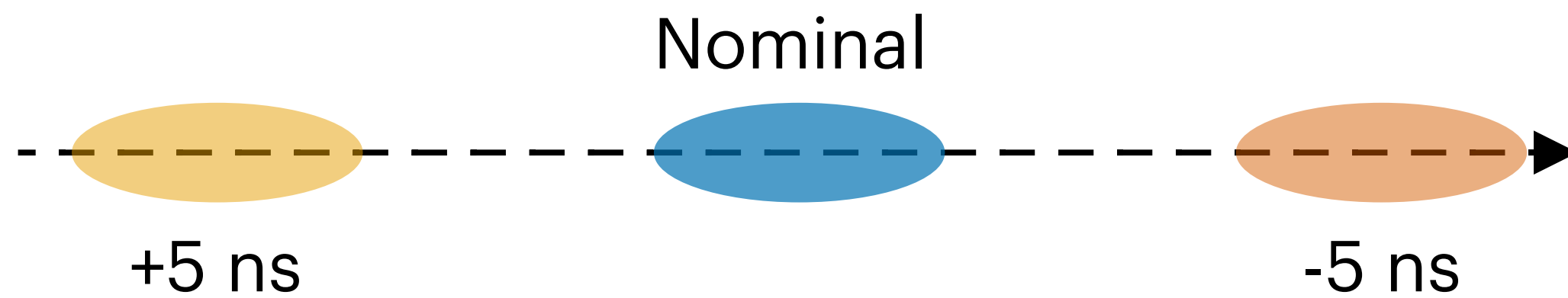


The image features a black background with a stylized white graphic of a satellite collision. On the left, a satellite is depicted with a circular body and several long, thin lines radiating from it, representing its orbital path or debris. To the right of the satellite is a large, white, teardrop-shaped area that represents the impact or the resulting debris field. In the center of the image, there is a rectangular orange box with a white border. Inside this box, the words "Satellite collisions" are written in a bold, white, sans-serif font. The text is slightly italicized and has a subtle drop shadow effect.

Satellite collisions

What are satellite collisions?

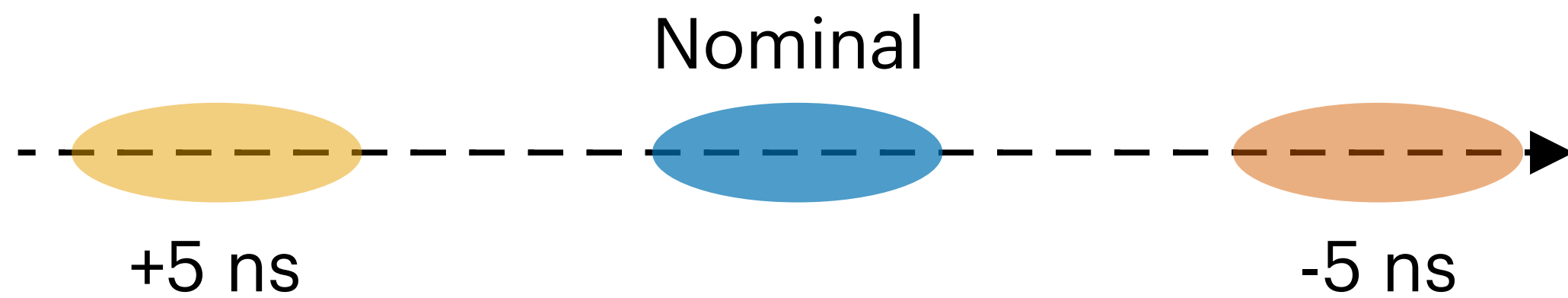
Bunch structure:



Satellite collisions are collisions
between protons in RF-buckets other
than nominal buckets

What are satellite collisions?

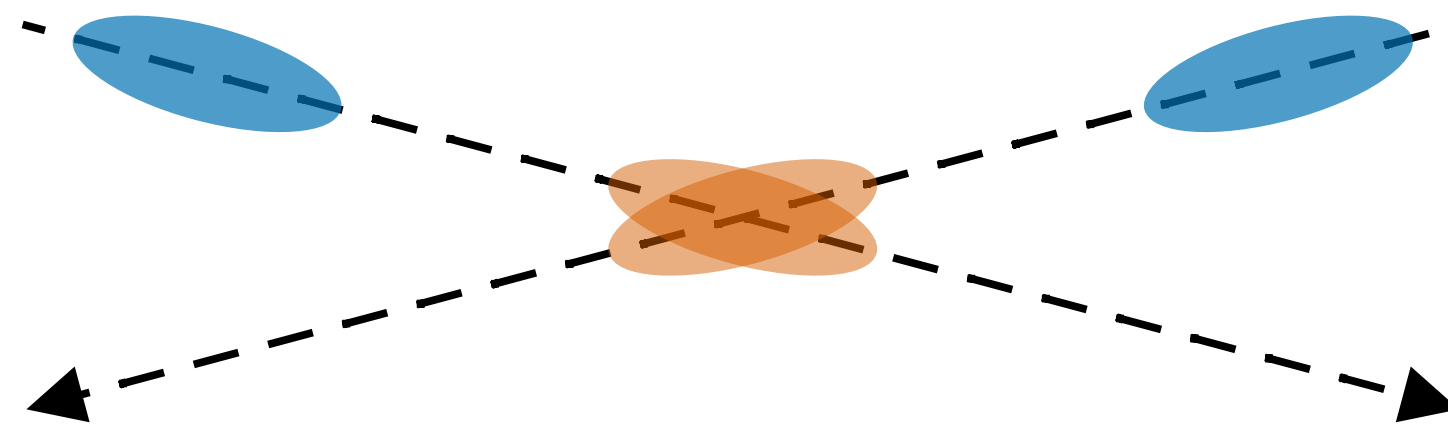
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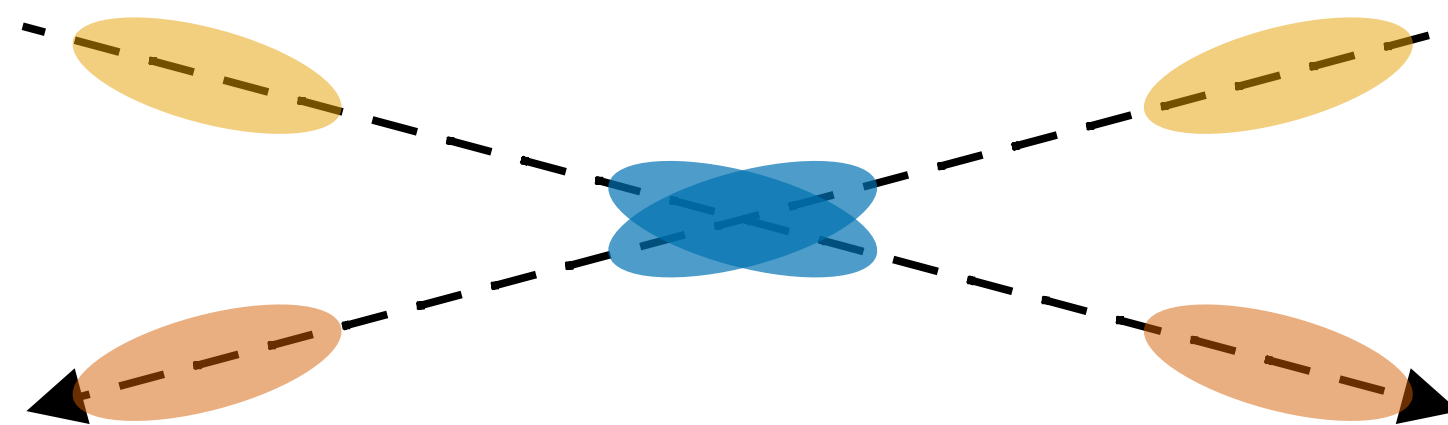
Satellite collisions are collisions between protons in RF-buckets other than nominal buckets

This produces genuinely delayed jets (but localised in time and with tracks in the inner detector)

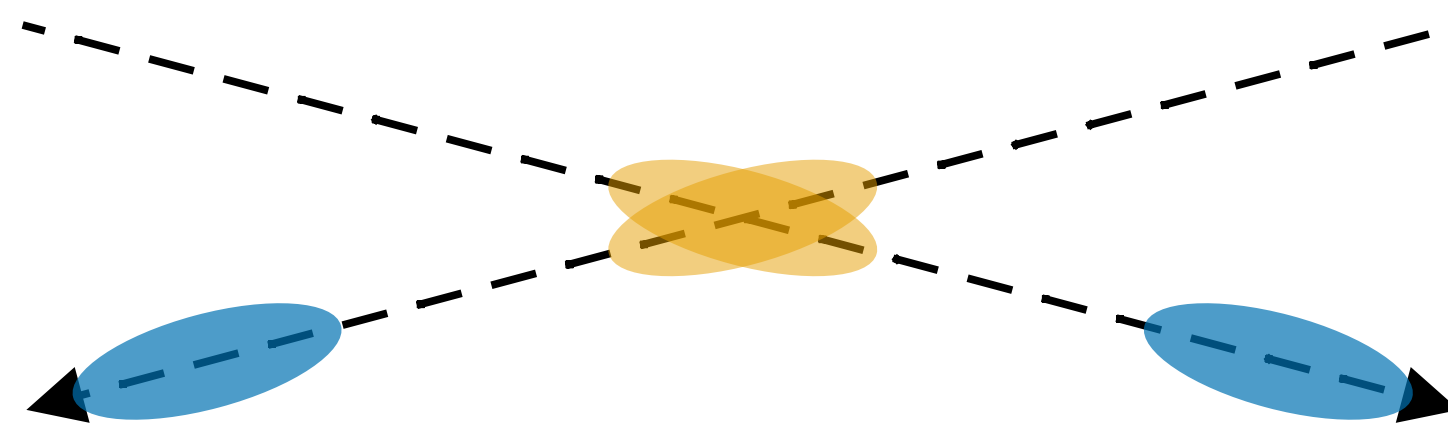
Collision types:



Satellite collision:
5 ns early



Nominal collision

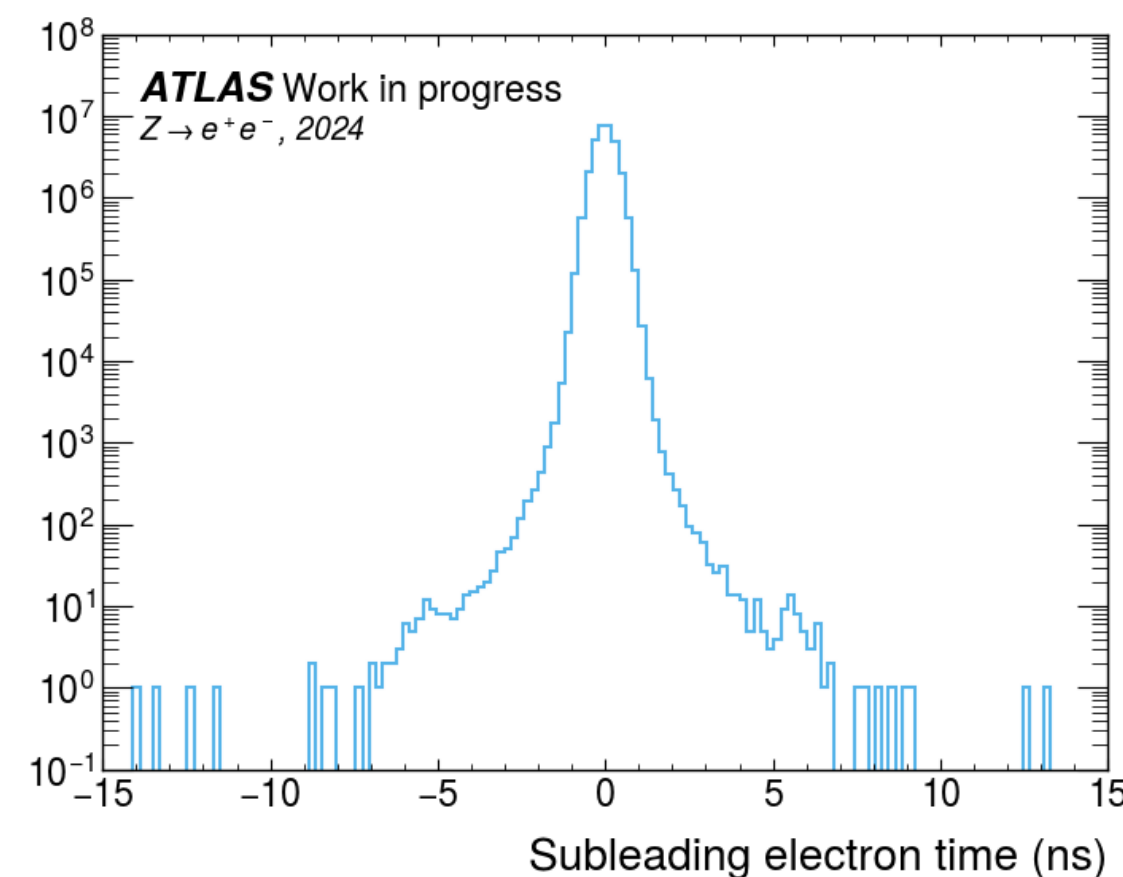
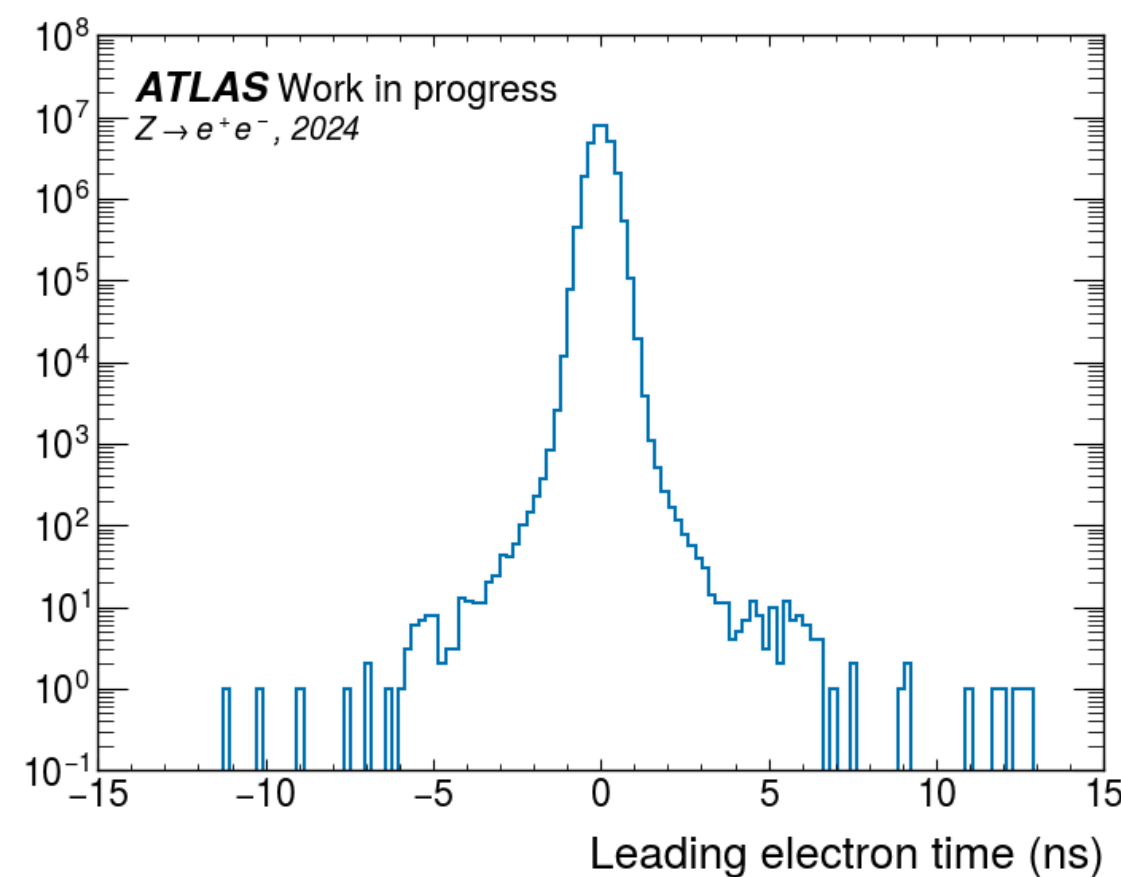


Satellite collision:
5 ns late

Satellites in $Z \rightarrow e^+e^-$

LAr has high precision (~ 200 ps) for high-energy electrons, so look at $Z \rightarrow e^+e^-$ (since it can be cleanly selected) and require cell energy > 10 GeV

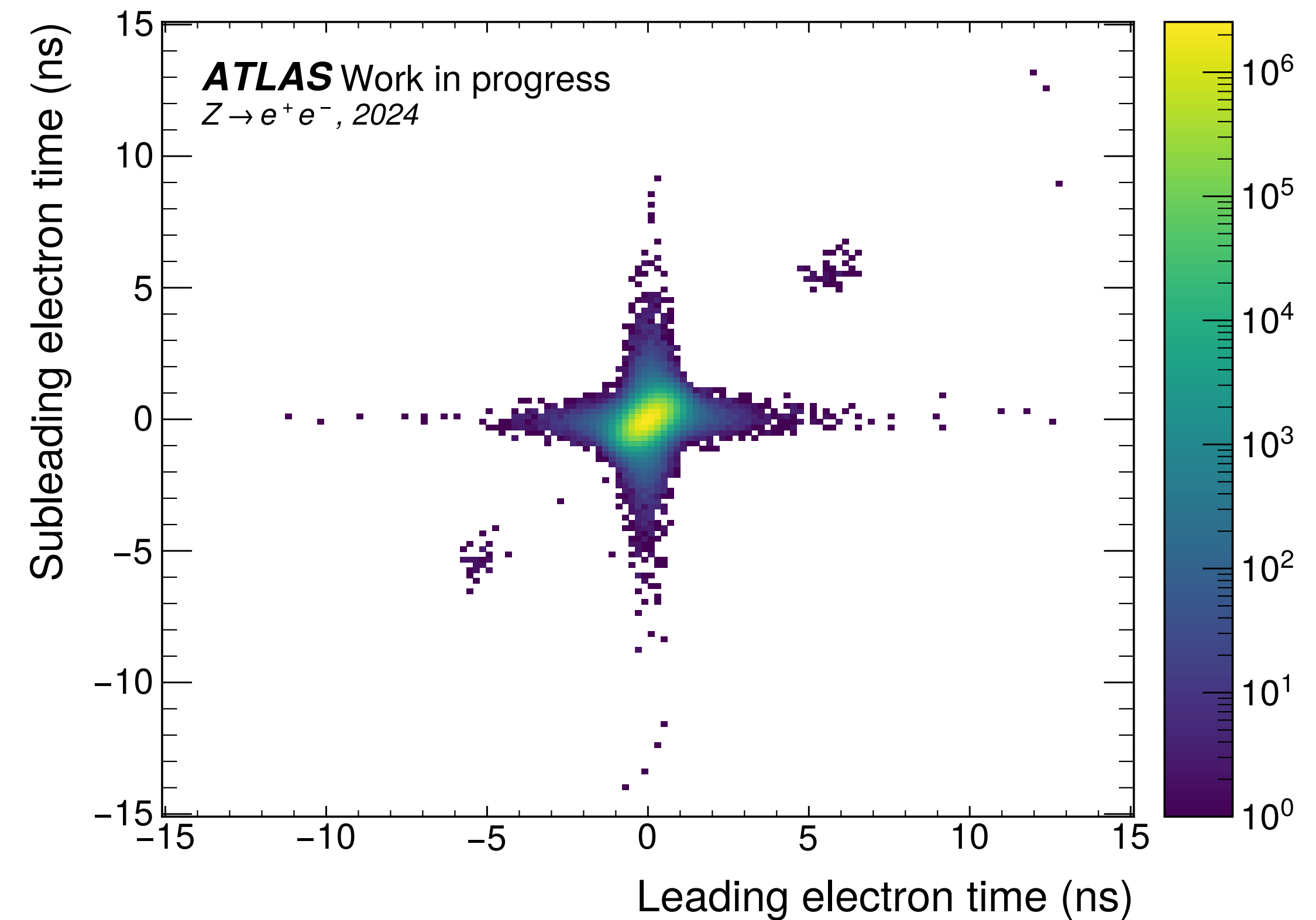
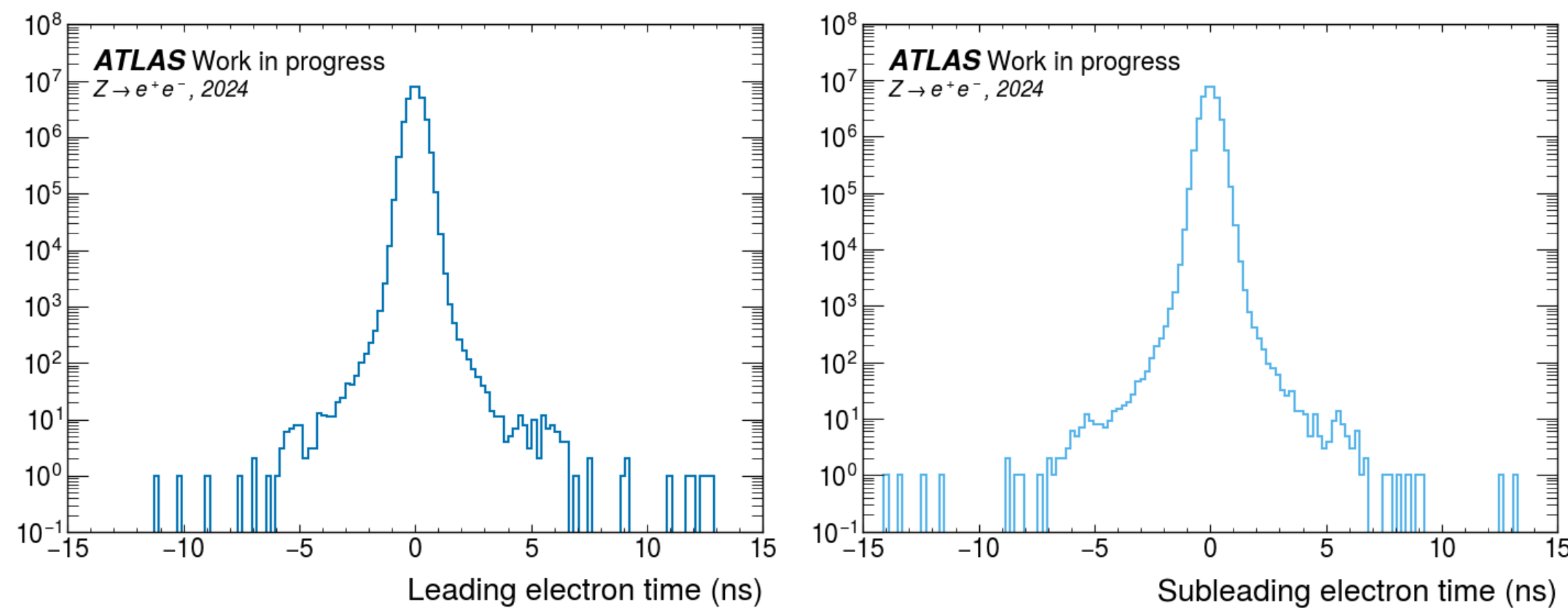
1D distributions are still too broad to clearly see electrons from satellite events



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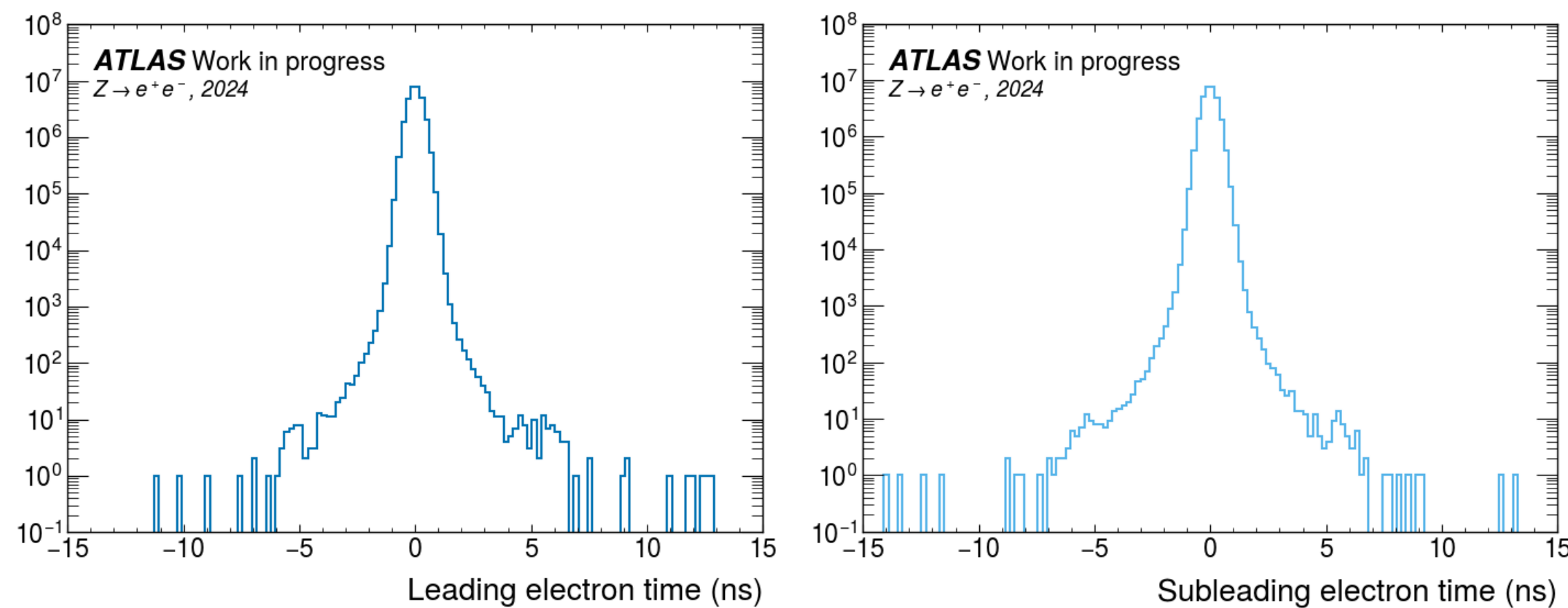
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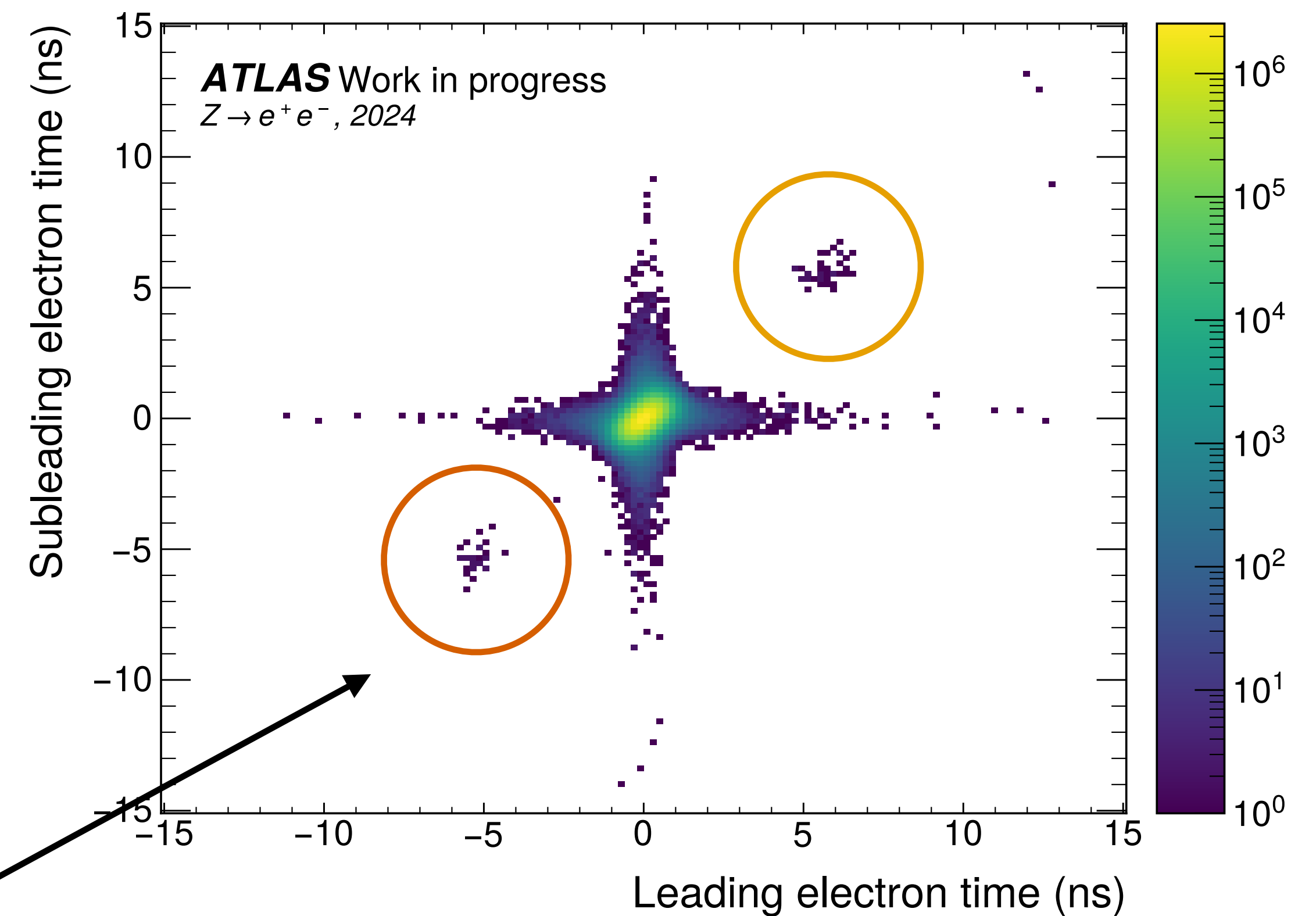
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However, leading vs. subleading electron time clearly shows clusters at ± 5 ns $\rightarrow$ **Z bosons produced in early and late satellite collisions!**



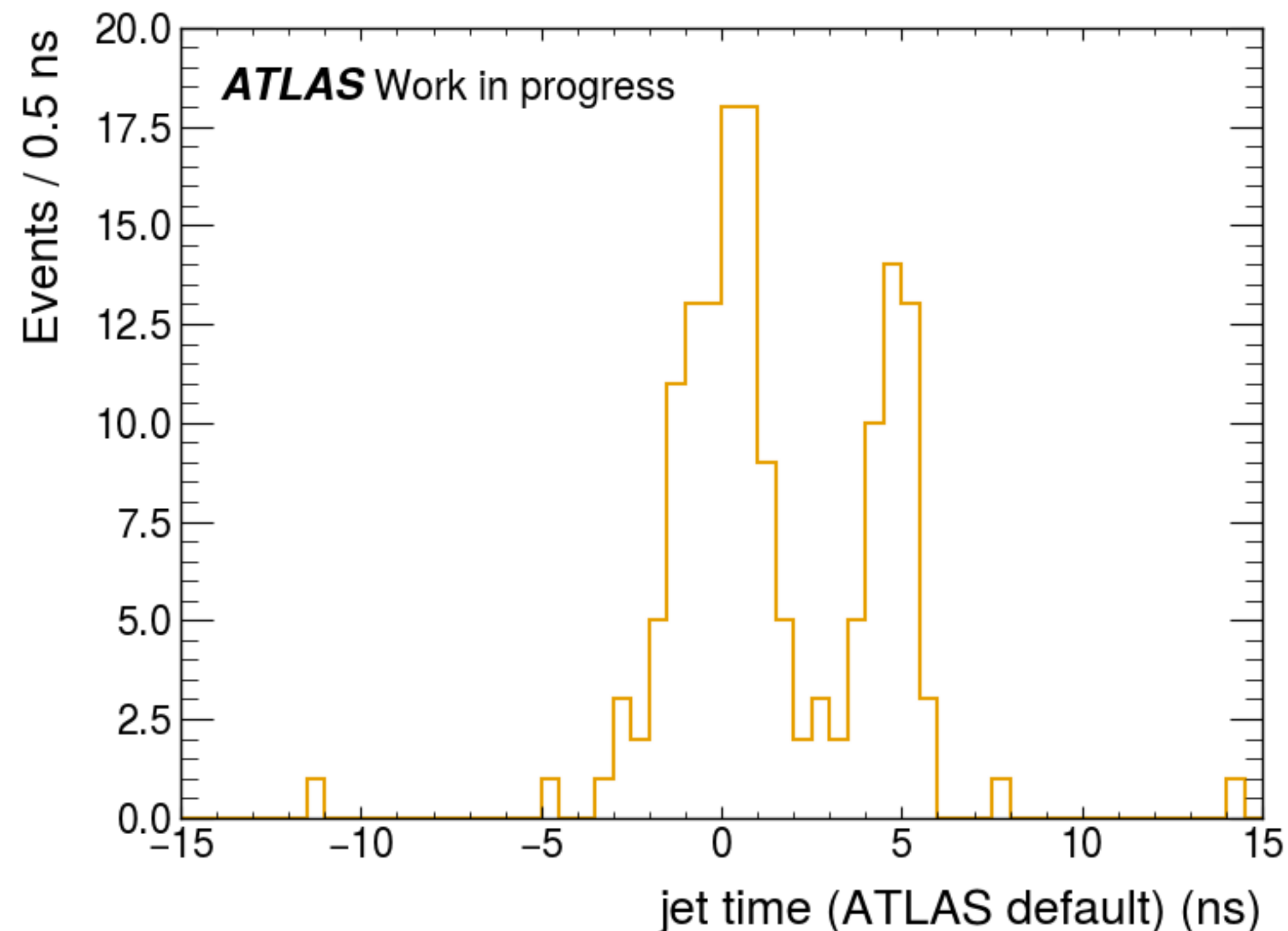
Relative counts: $r_{\text{early}} = (1.0 \pm 0.2) \times 10^{-6}$
(compared to nominal collisions)
 $r_{\text{late}} = (1.3 \pm 0.2) \times 10^{-6}$

Moving from electrons to jets

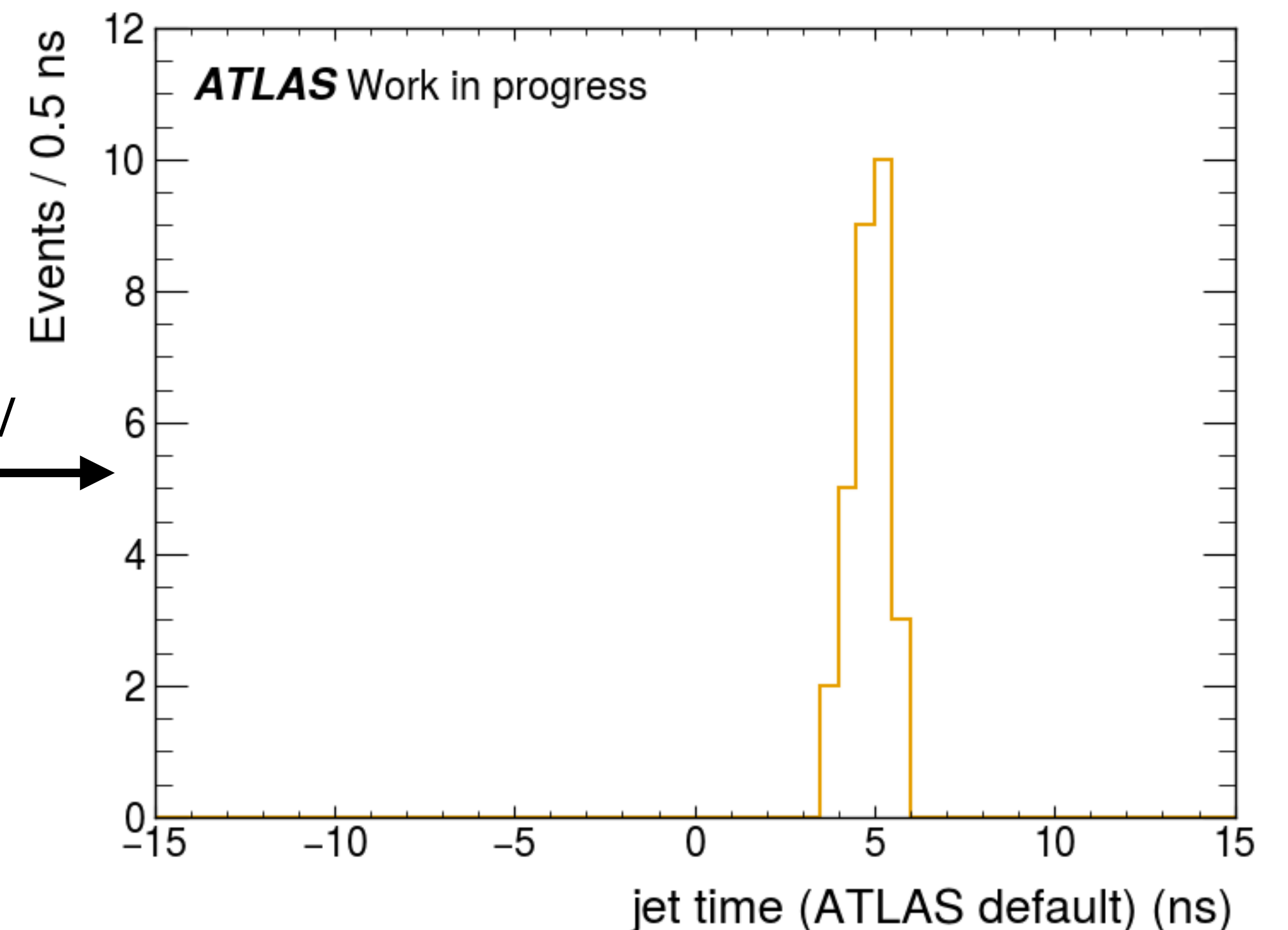
Jet-time distribution is different from electrons, but expected to follow SM jets (just shifted)

But shifted by how much? Exactly ± 5 ns, or are there some biases?

Let's look at jets in known satellite events from $Z \rightarrow e^+e^-$ (pictured here: [late](#))



$p_T > 50$ GeV

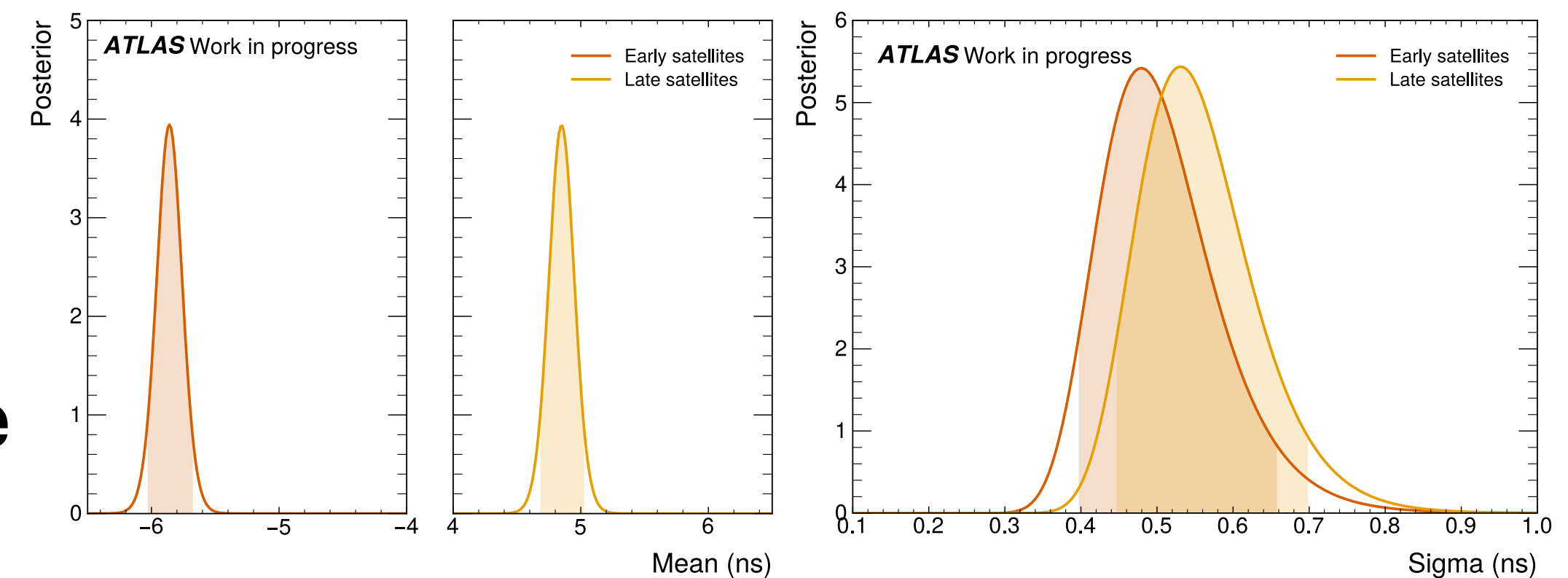
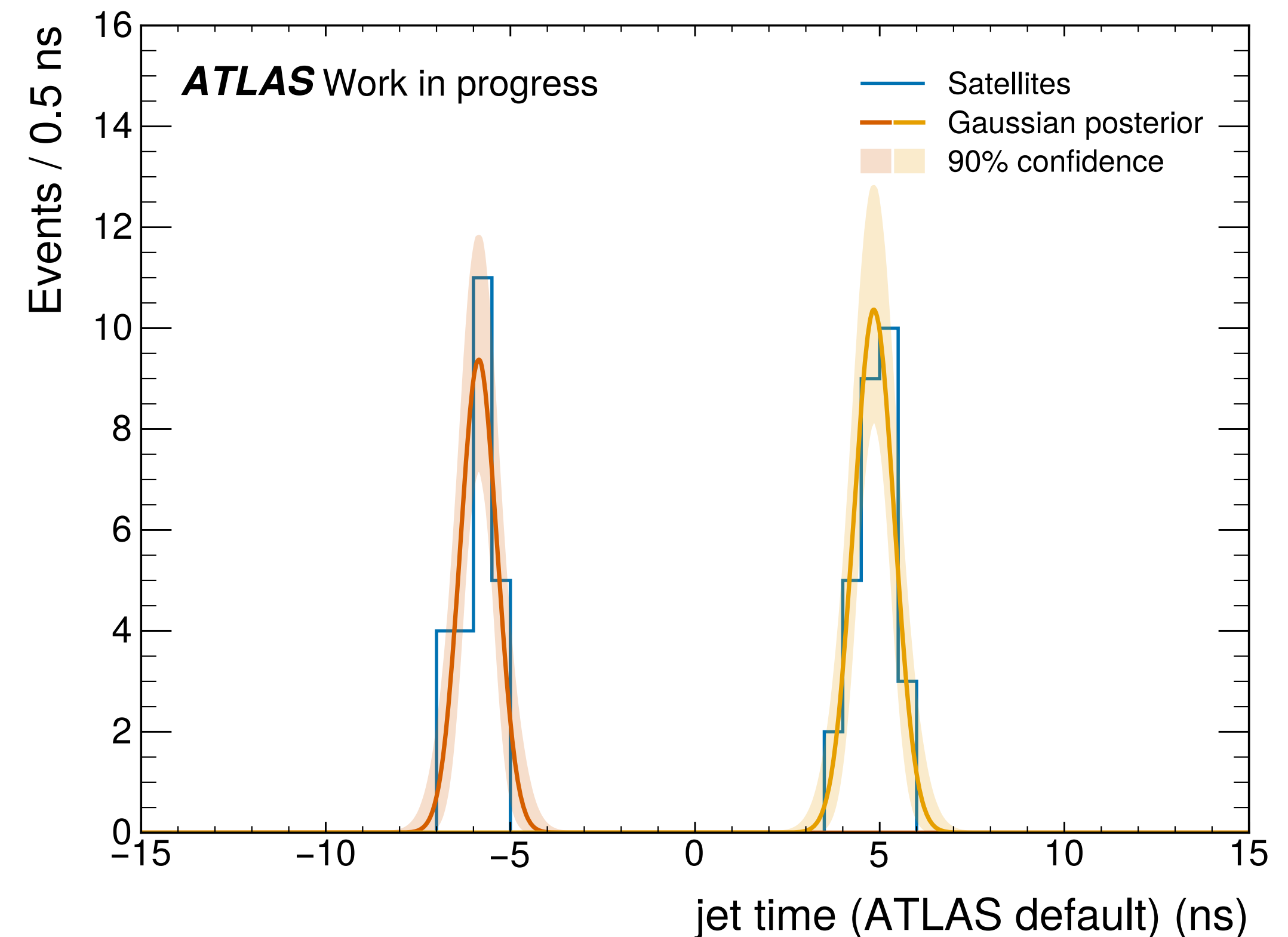


Bayesian analysis

- 24 jets in **early** and 29 in **late** satellites
—> Small sample size, ideal for Bayesian inference with a Gaussian model
- Jet times aren't perfectly centred on ± 5 ns
—> Jet-time calculation isn't independent of absolute time

Not enough data to fully model this effect, but we can use this analysis to inform a systematic uncertainty

Combining distribution posteriors with relative counts and a signal-region definition will give the final satellites estimate!



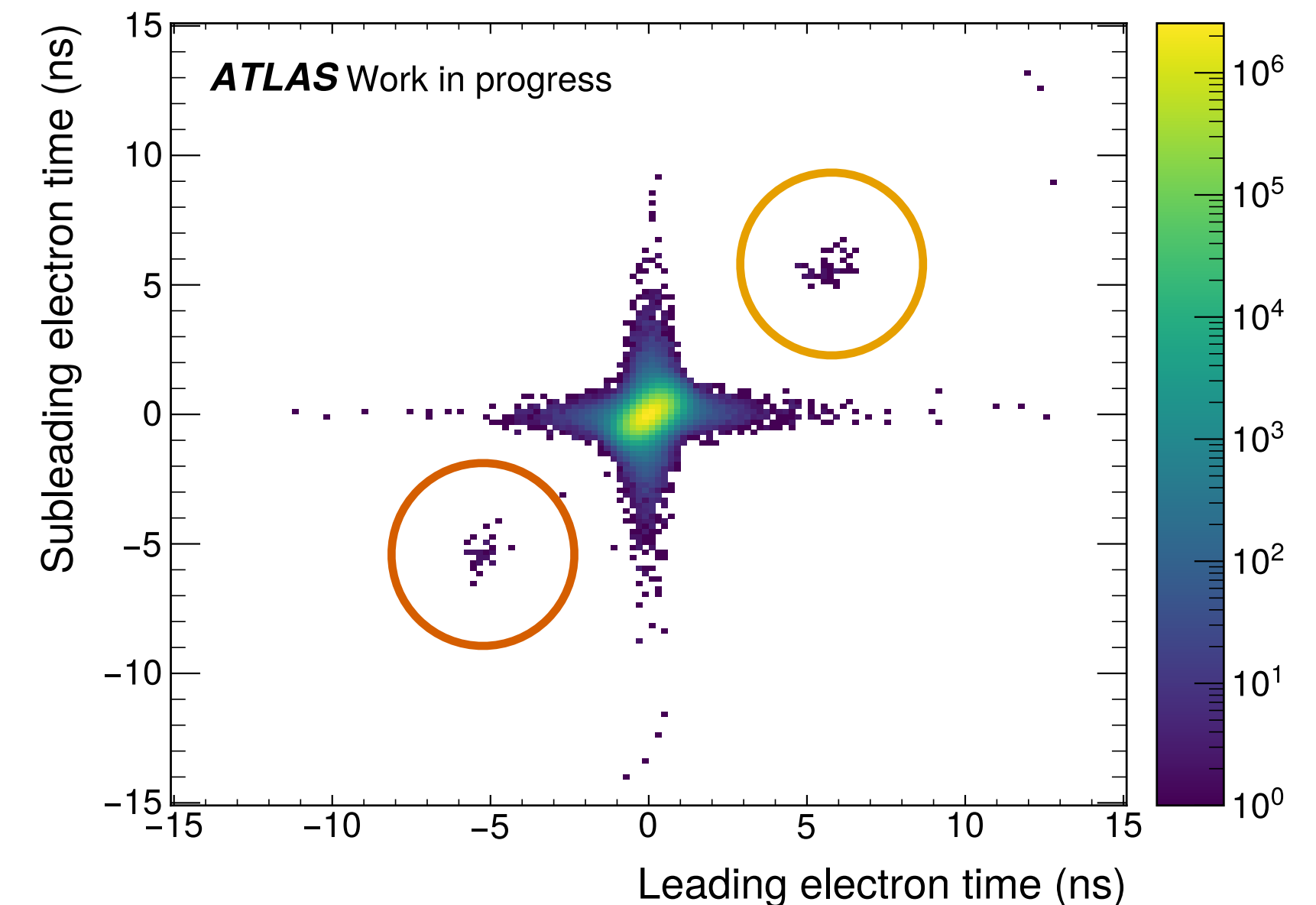
Recap & outlook

Heavy long-lived particles produce **delayed** decays, which complements their displacement and gives additional sensitivity in calorimeter systems

The search for **delayed jets** requires modelling of unconventional backgrounds, with almost all new estimation techniques

One such background: **satellite collisions** happening at ± 5 ns, estimated with:

- Relative count estimated at $\mathcal{O}(10^{-6})$ from $Z \rightarrow e^+e^-$
- Jet-time distribution and systematics estimated from SM jets from $Z \rightarrow e^+e^- + \text{jets}$



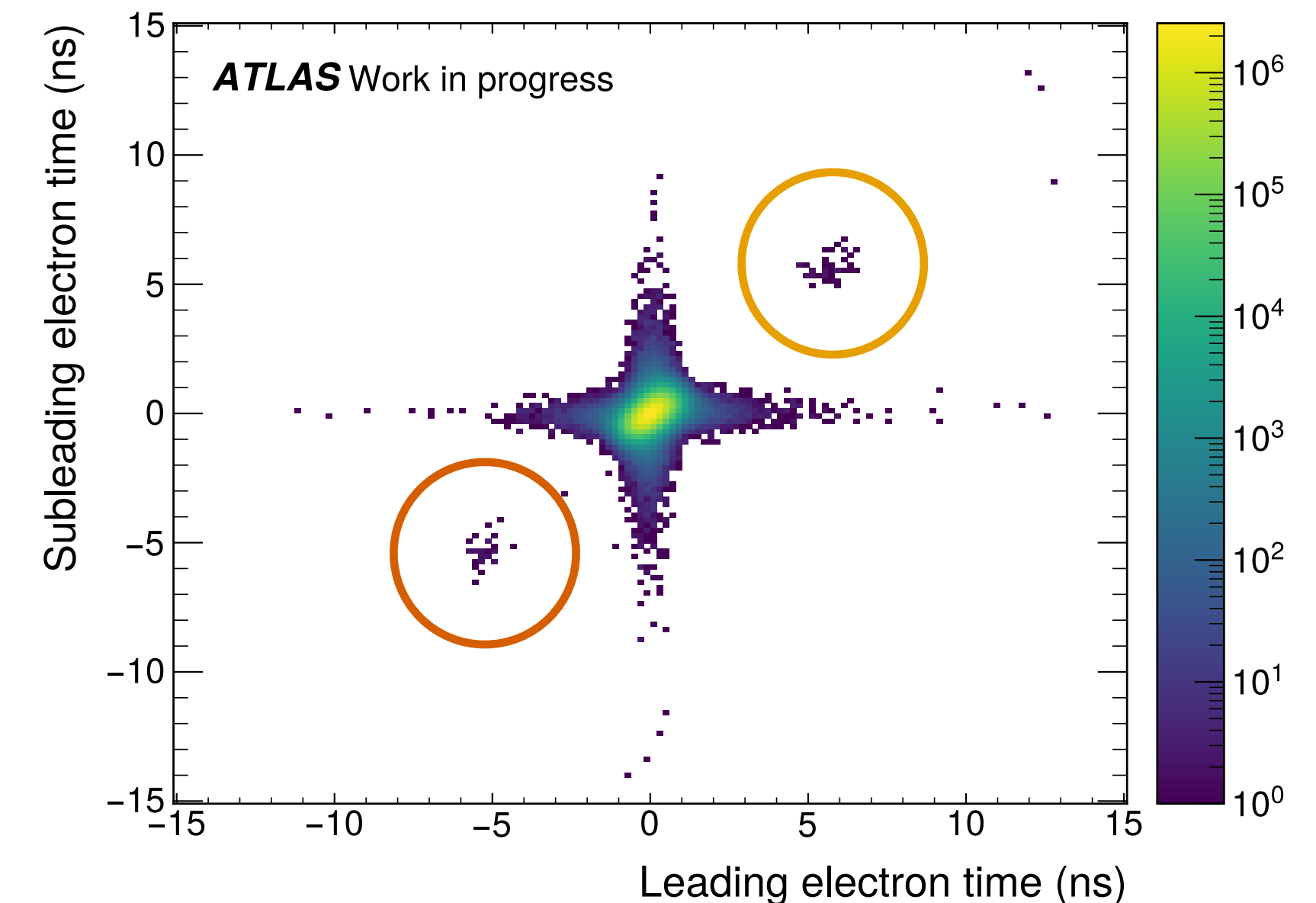
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*In Run 4, ATLAS will have a new timing detector: **HGTD**! Timing information will become more important to overall ATLAS reconstruction, and this search can provide important groundwork*

Ba

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up

Satellites in $W^\pm \rightarrow e^\pm \nu$

- $W^\pm \rightarrow e^\pm \nu$ has much more stats than $Z \rightarrow e^+ e^-$ so it's also a candidate for satellites estimation
- Two problems:
 - Satellite collisions are overlayed on the instrumental background, requiring a fitting procedure and making it harder to select satellite events to study the jets
 - LAr's timing response is calibrated on $W^\pm \rightarrow e^\pm \nu$, introducing biases in the satellite counts from fits (when compared to $Z \rightarrow e^+ e^-$)

