
PriNCe

Propagation including Nuclear Cascade equations

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Hands-on sessions

ACME

09.04.2026



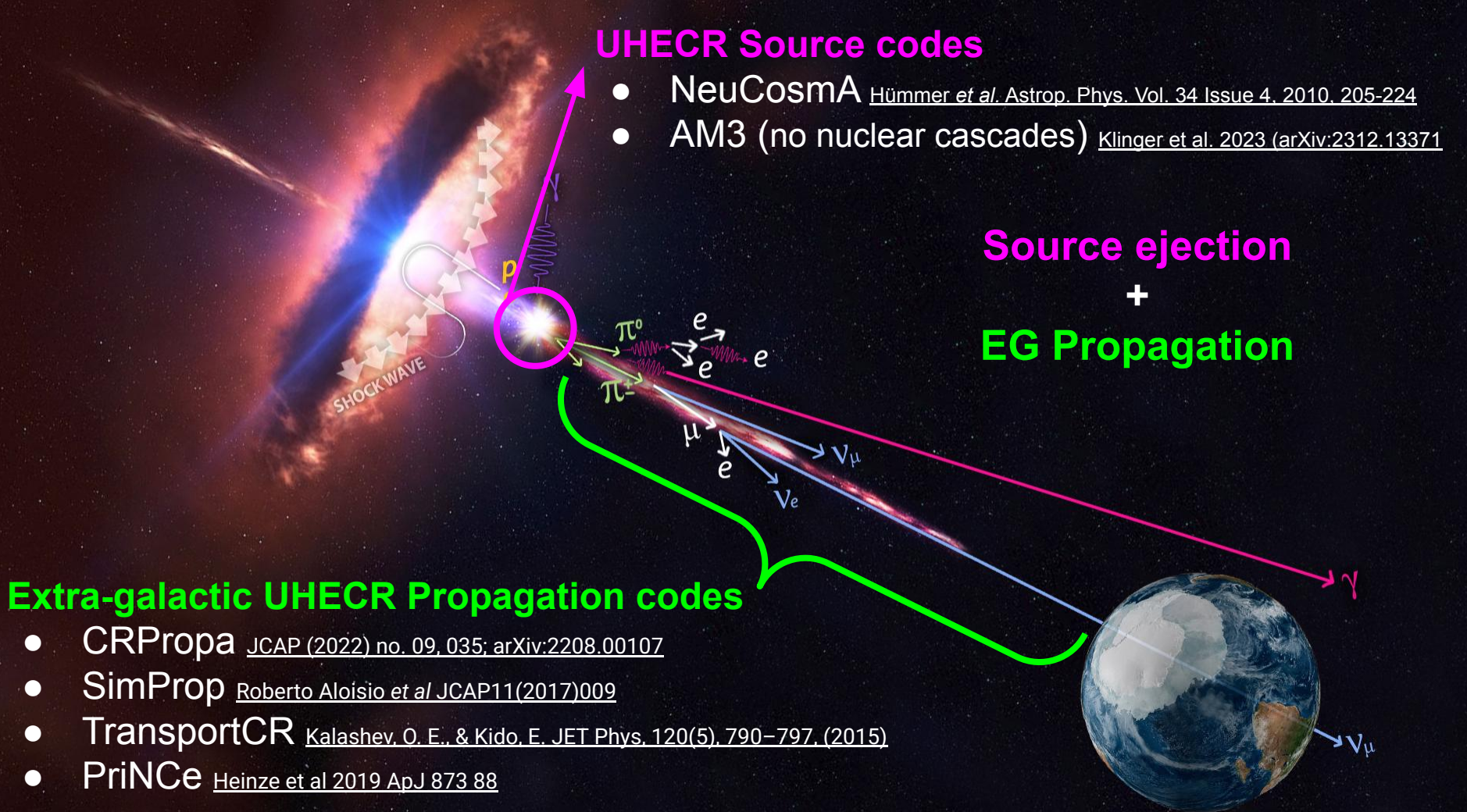
UHECR Source codes

- NeuCosmA [Hümmer et al. Astrop. Phys. Vol. 34 Issue 4, 2010, 205-224](#)
- AM3 (no nuclear cascades) [Klinger et al. 2023 \(arXiv:2312.13371\)](#)

Source ejection
+
EG Propagation

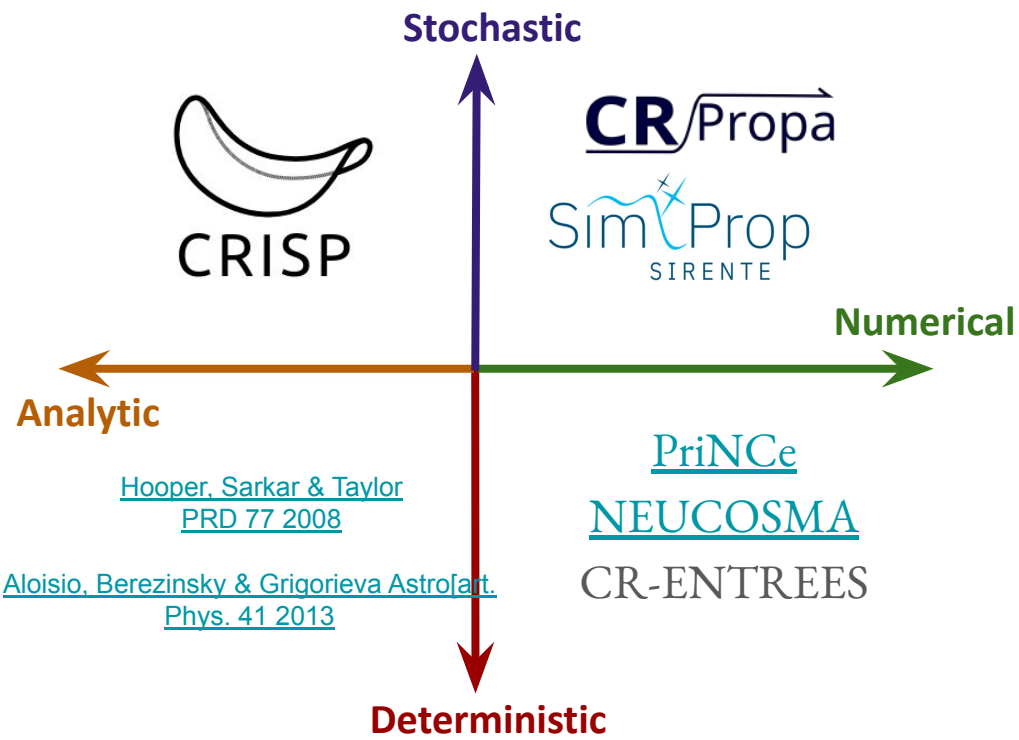
Extra-galactic UHECR Propagation codes

- CRPropa [JCAP \(2022\) no. 09, 035; arXiv:2208.00107](#)
- SimProp [Roberto Aloisio et al JCAP11\(2017\)009](#)
- TransportCR [Kalashev, O. E., & Kido, E. JET Phys. 120\(5\), 790–797, \(2015\)](#)
- PriNCe [Heinze et al 2019 ApJ 873 88](#)

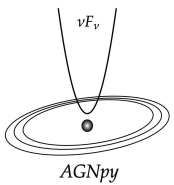




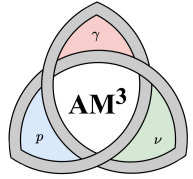
Overview

PriNCE: compared to other frameworks



CRISP could be coupled to...

SOPRANO **OneHaLe**


AGNpy **JetSeT**
Jets SED modeler and fitting Tool

γπ A Python package for **gamma-ray** astronomy



PriNCE: access and documentation

Main sources for PriNCE

- Documentation: prince.readthedocs.io/en/latest/
- Original Repository: github.com/joheinze/PriNCE
- Examples: github.com/joheinze/PriNCE-examples
- [PhD thesis, Jonas Heinze](#)

Resources for today's session on PriNCE

- Repository for today's session: github.com/mohller/PriNCE
- Examples: github.com/joheinze/PriNCE-examples

Nuclear Cascade Equations

$$\frac{\partial}{\partial t} N_i(E, t) = \frac{\partial}{\partial E} (b(E) N_i(E, t)) + Q_i^{\text{ext}}(E, t) - \Gamma_i N_i(E) + \sum_j Q_{j \rightarrow i}(N_j(E, t))$$

Energy losses
causing redistribution

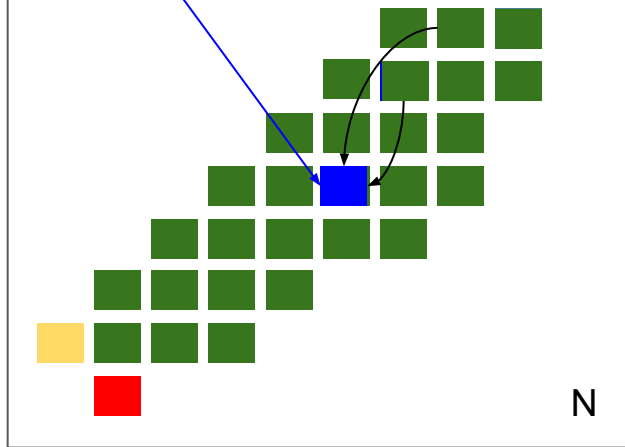
Injection
(accelerated particles)

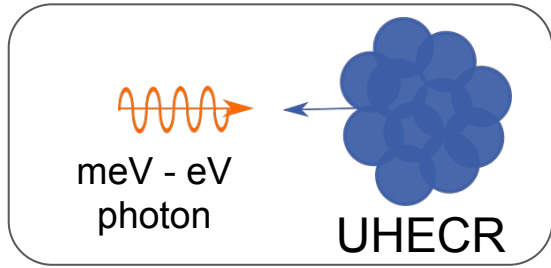
Z

System of first order PDEs.
Particle densities for different species.

- SODE describes time evolution
- CR injection constant in time
- Integration over source time-scale
- Production of all particle species (including nuclei, mesons and neutrinos)

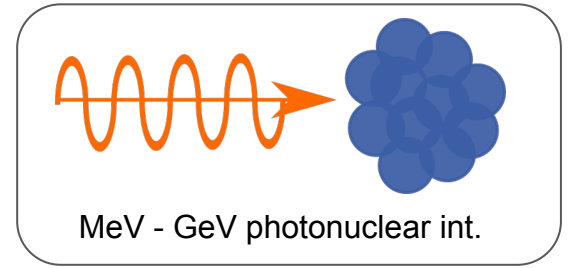
chart of nuclear species

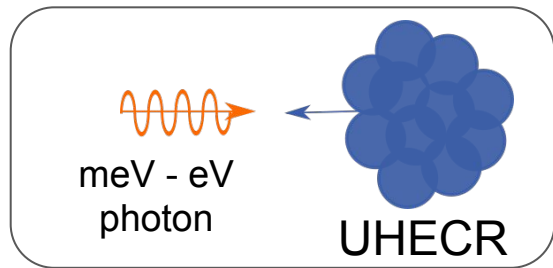




to center of mass
reference frame

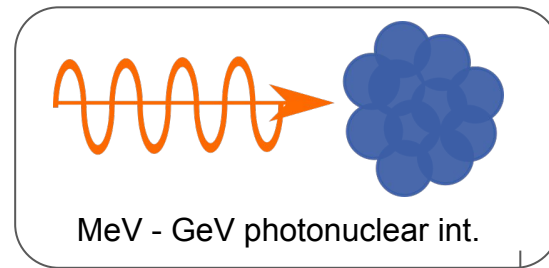
A dashed horizontal line with a grey arrowhead pointing to the right.



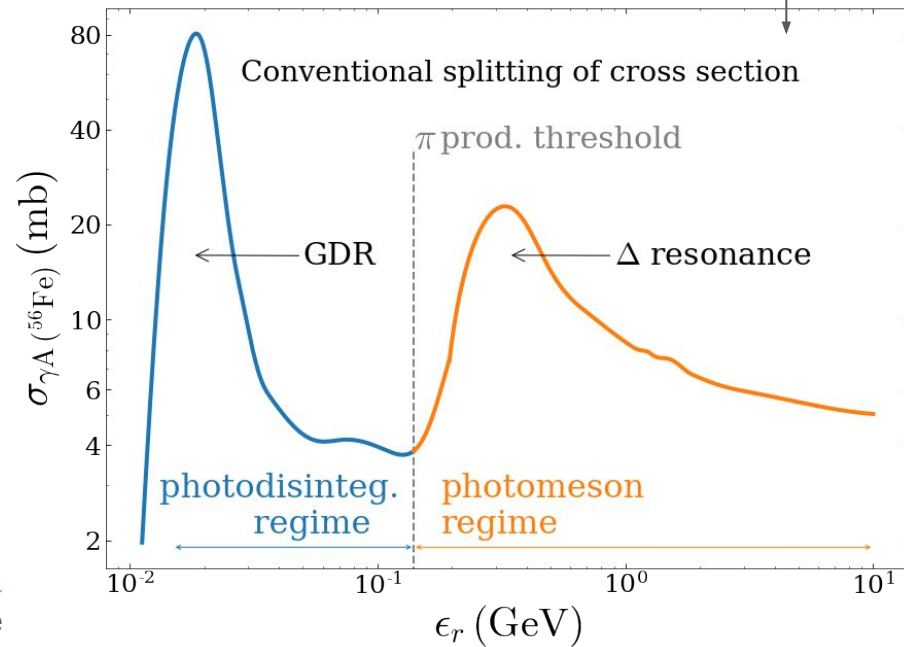


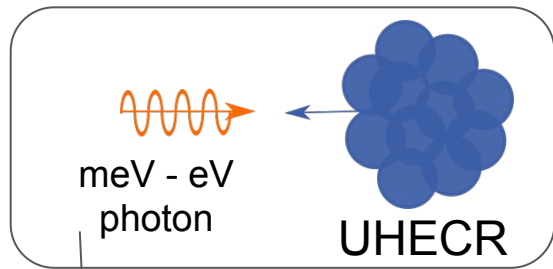
to center of mass reference frame

A dashed black arrow points from the left diagram to the right diagram, indicating a transition to the center of mass reference frame.

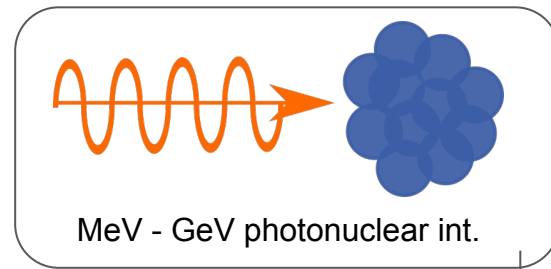


[L. Morejon et al JCAP11\(2019\)007](#)

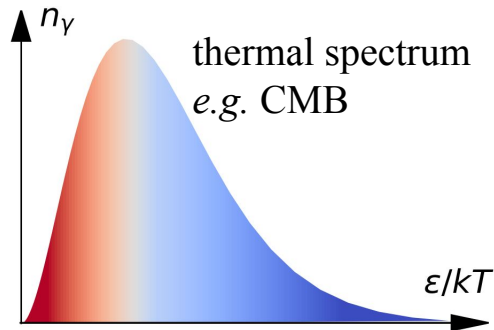




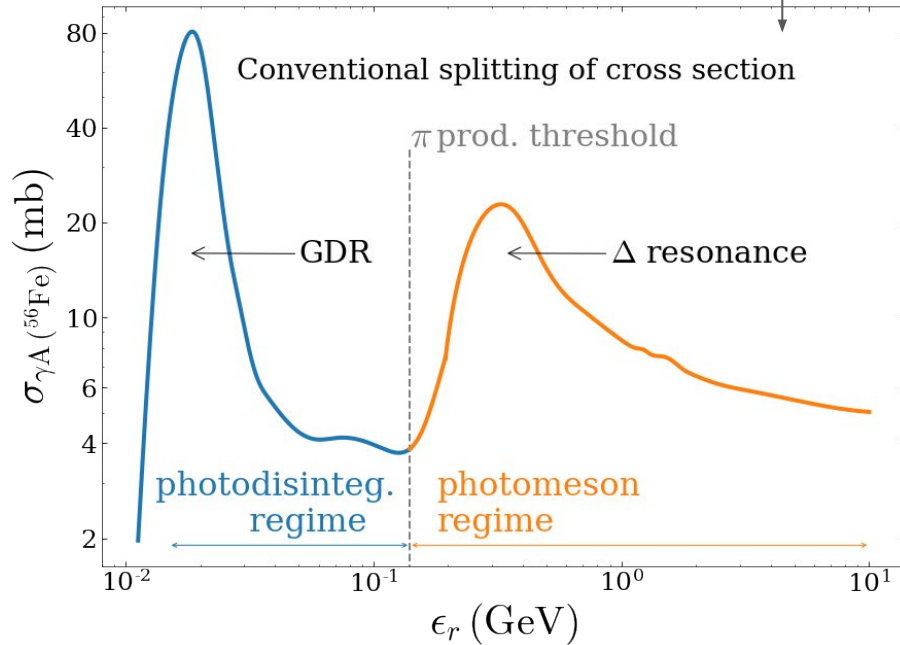
to center of mass reference frame

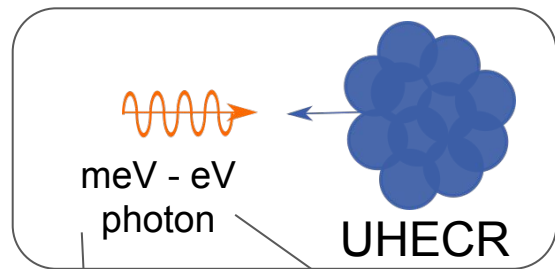


PROPAGATION



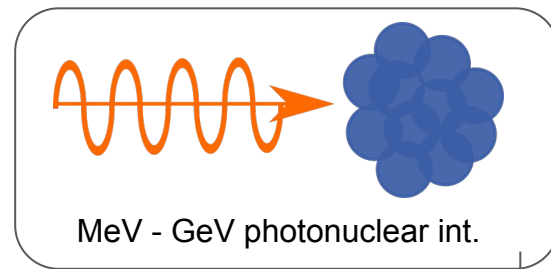
[L. Morejon et al JCAP11\(2019\)007](#)





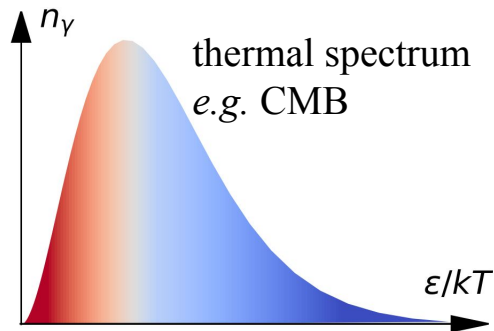
to center of mass reference frame

A dashed black arrow points from the left diagram to the right diagram.

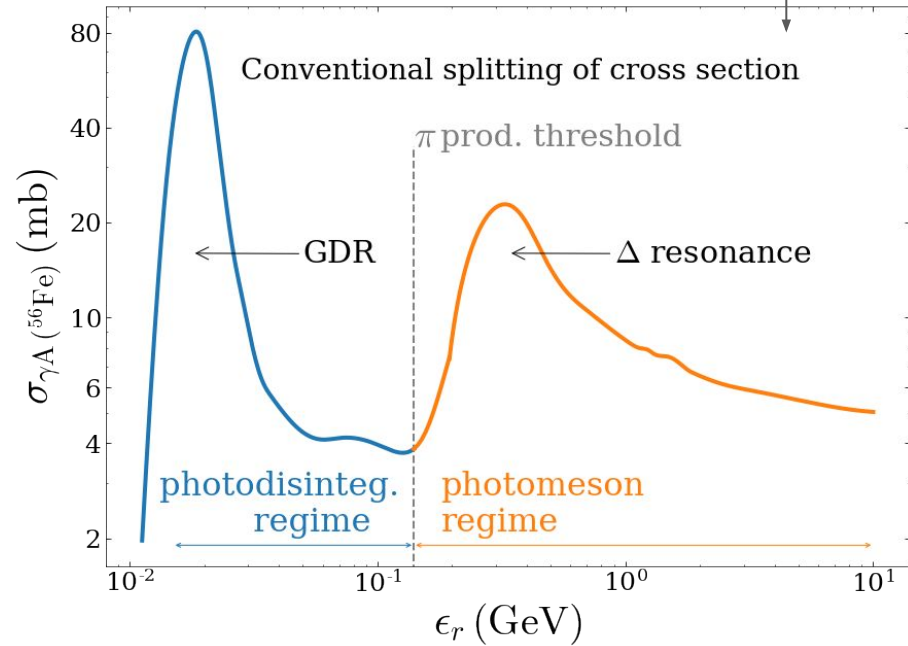
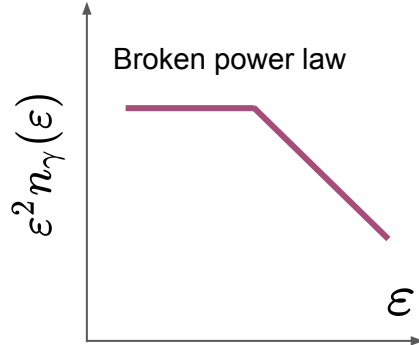


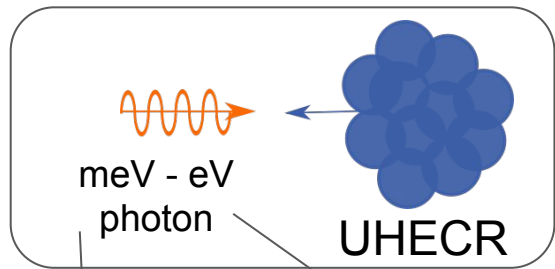
[L. Morejon et al JCAP11\(2019\)007](#)

PROPAGATION

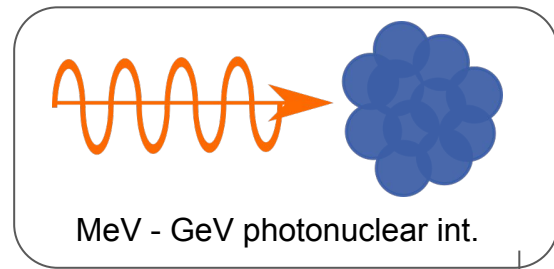


SOURCE



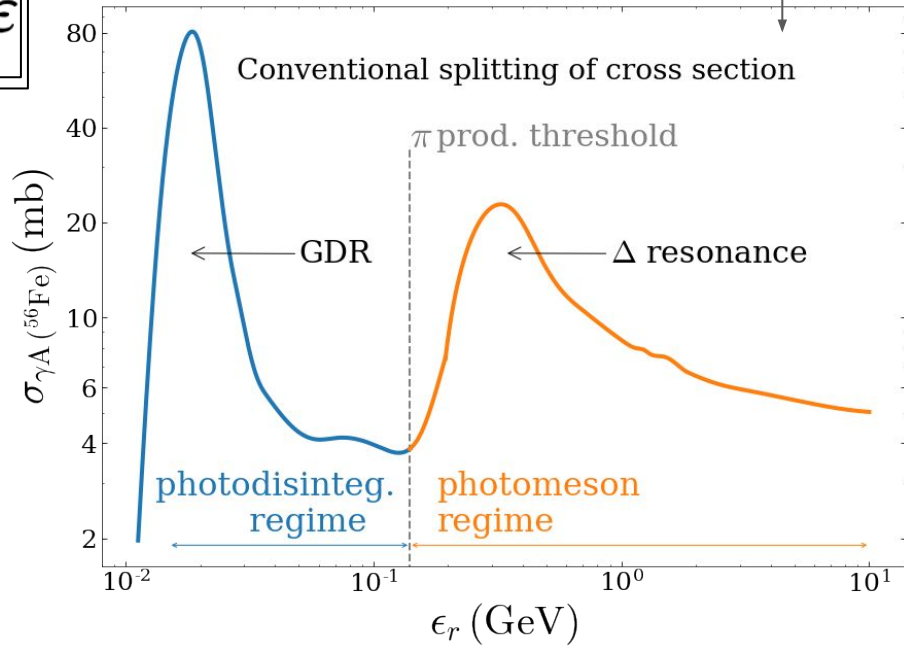


to center of mass reference frame



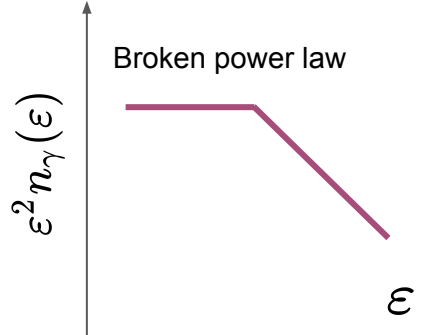
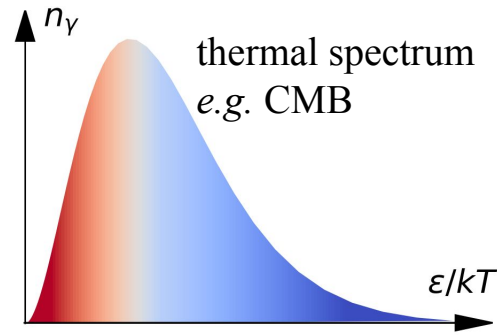
$$\lambda(\gamma) = \frac{1}{2\gamma^2} \int_0^\infty \frac{n(\epsilon)}{\epsilon^2} d\epsilon \int_0^{2\epsilon\gamma} \epsilon \sigma(\epsilon) d\epsilon$$

[L. Morejon et al JCAP11\(2019\)007](#)



PROPAGATION

SOURCE



Support & Feedback...

Feedback

Please use this link and provide feedback on this session...



<https://www.acme-astro.eu/survey-hands-on-sessions-xwj4jm9/>

Support

For questions, support and issues in the code you can open an issue on the [github repository](#).

Additionally you can open a ticket in the ACME support website:



<https://support.acme-astro.eu/>