

Axions in the Sky

Samuel J. Witte

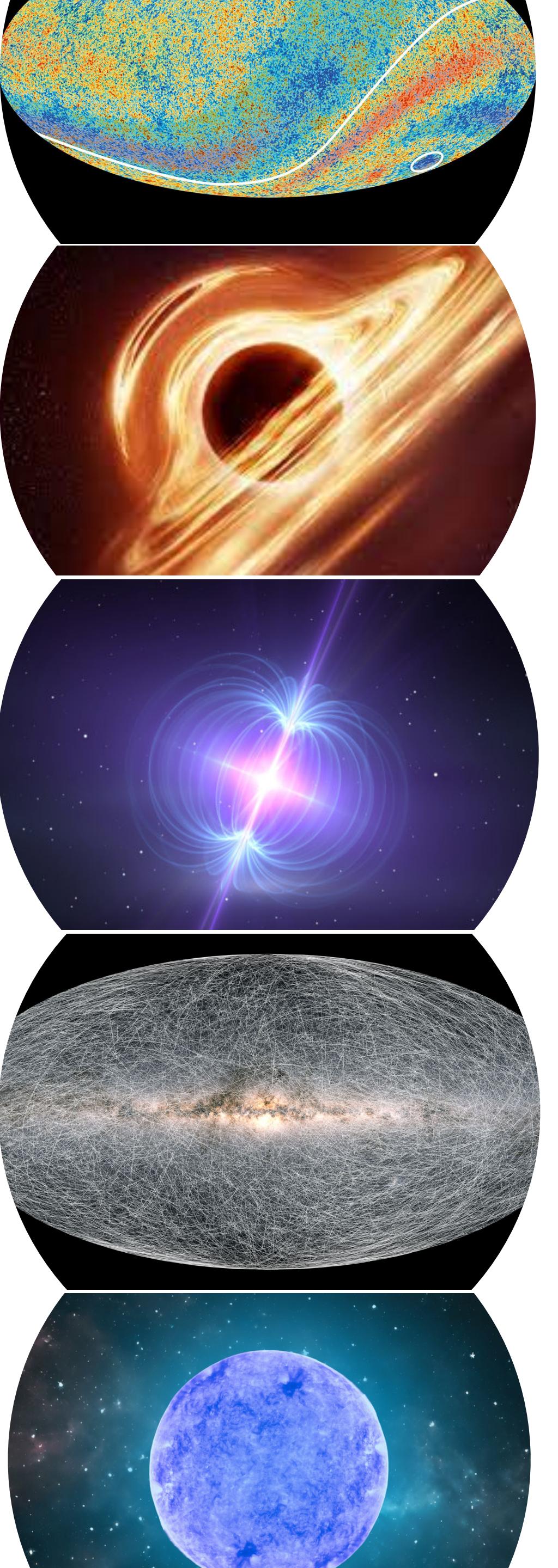
Oxford Theoretical Studies of Particles and Strings Retreat
March, 2024



THE ROYAL SOCIETY



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OXFORD



Astrophysics as a laboratory

Nature's Laboratories



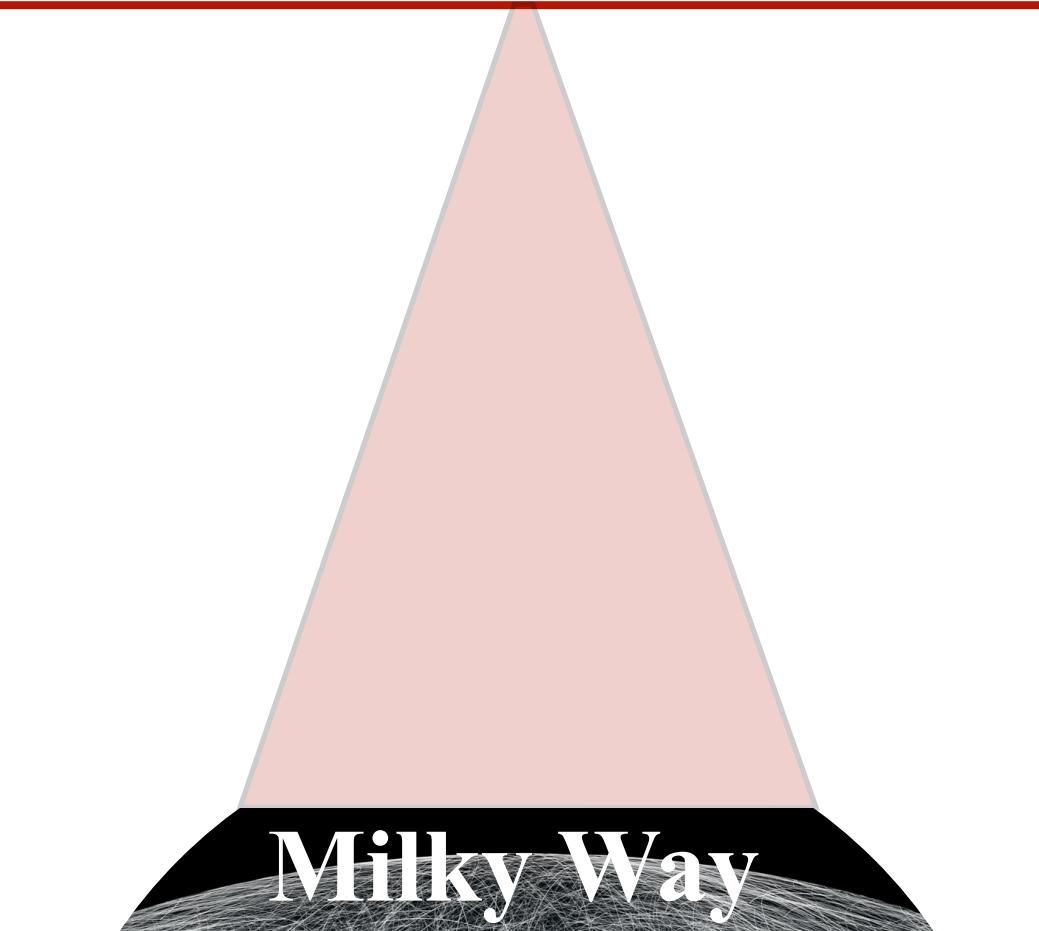
Black hole



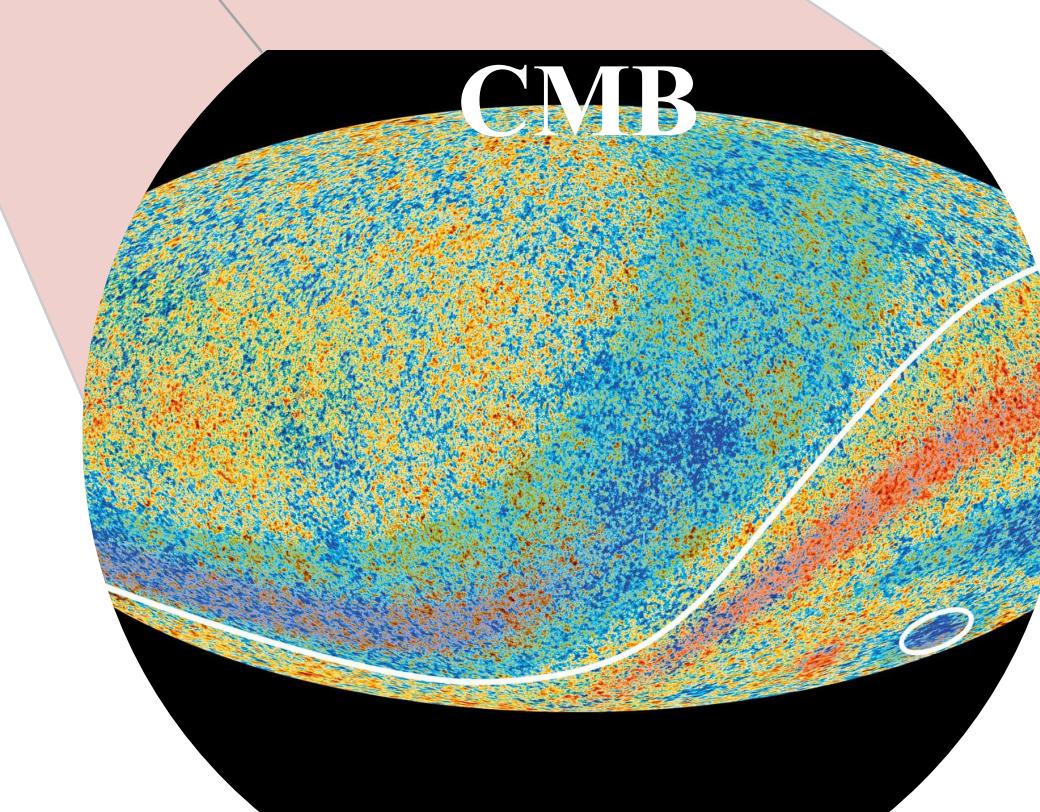
White Dwarf



Pulsar



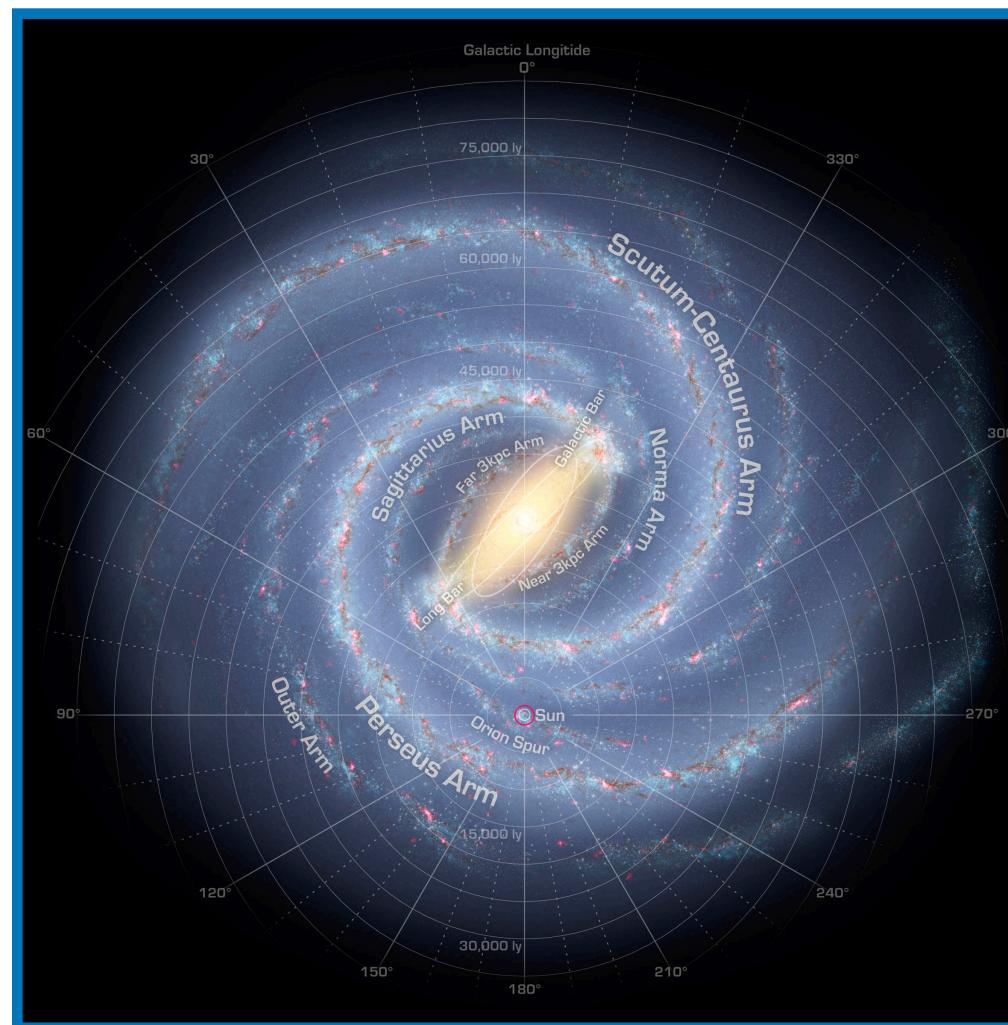
Milky Way



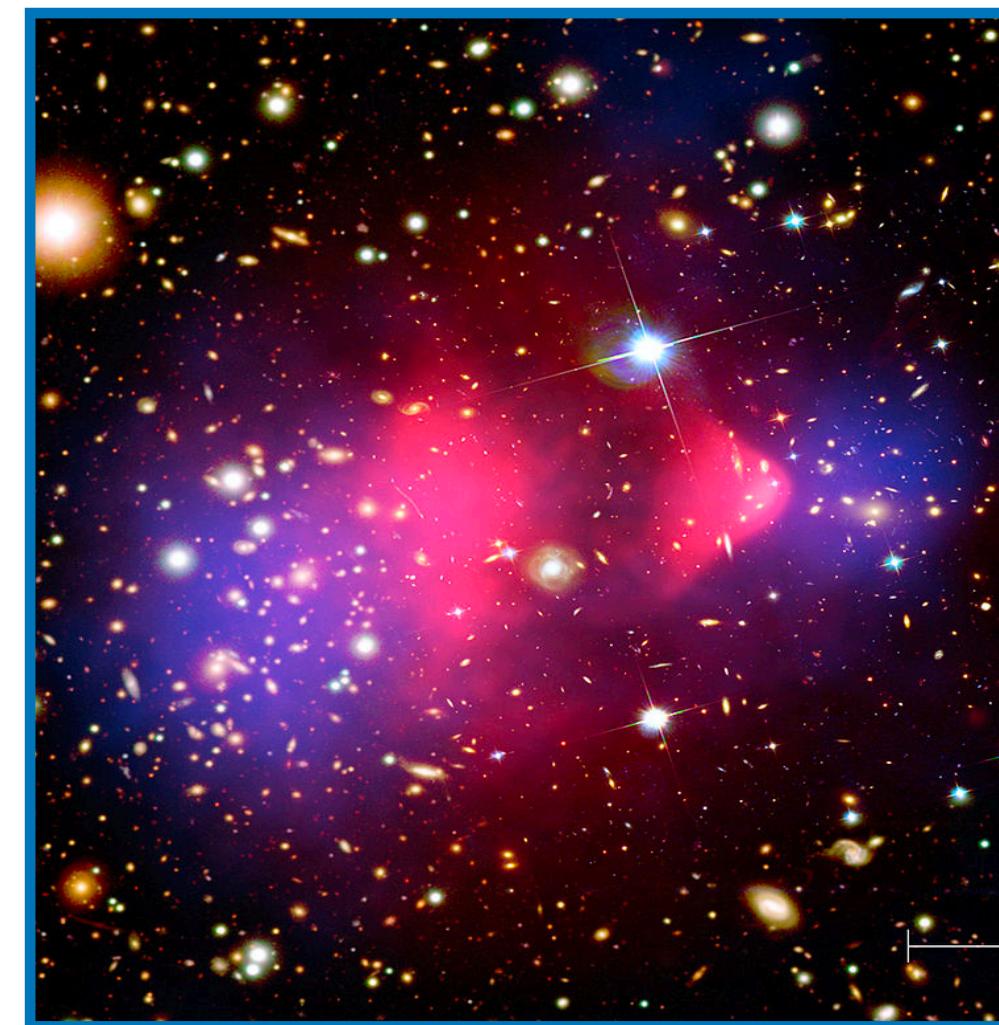
CMB

New particles in the sky: dark matter

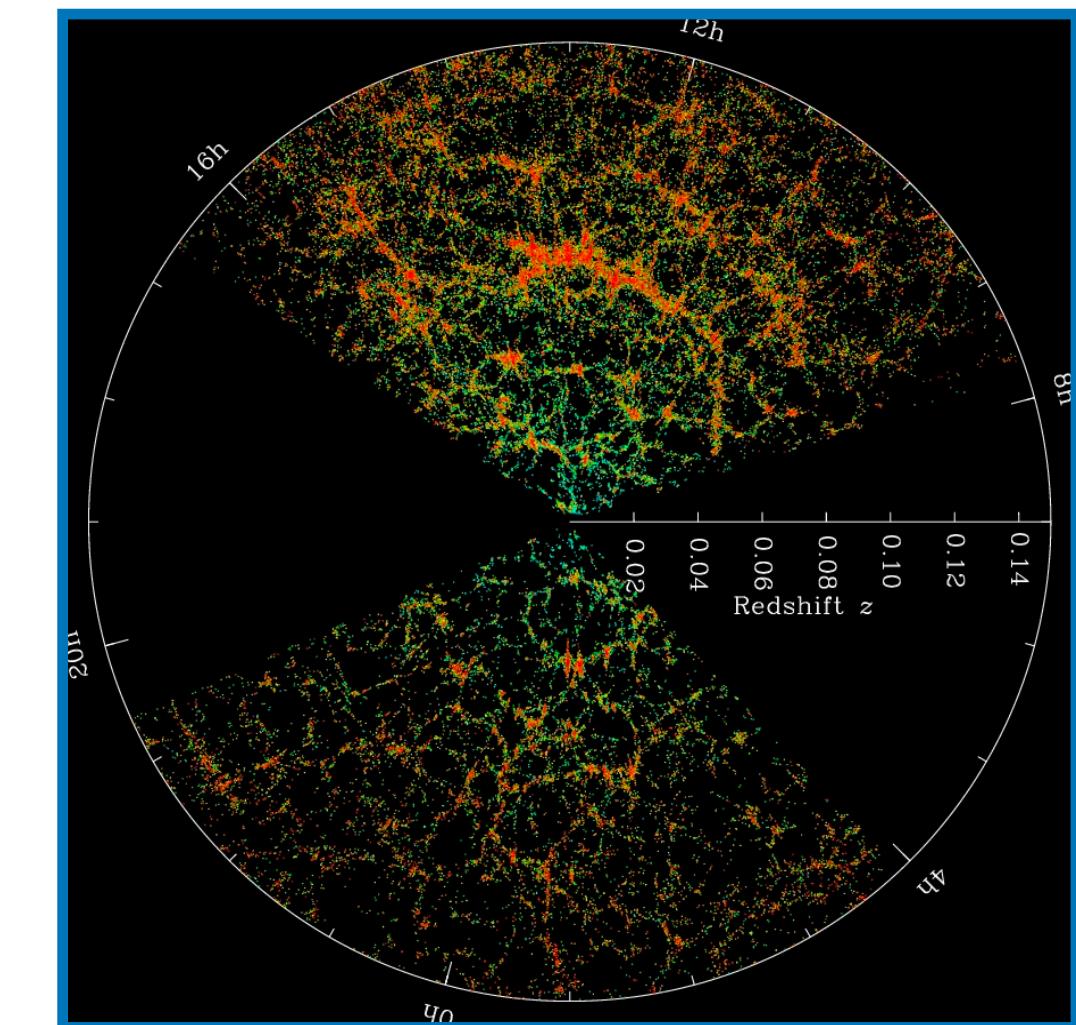
Galaxy rotation curves



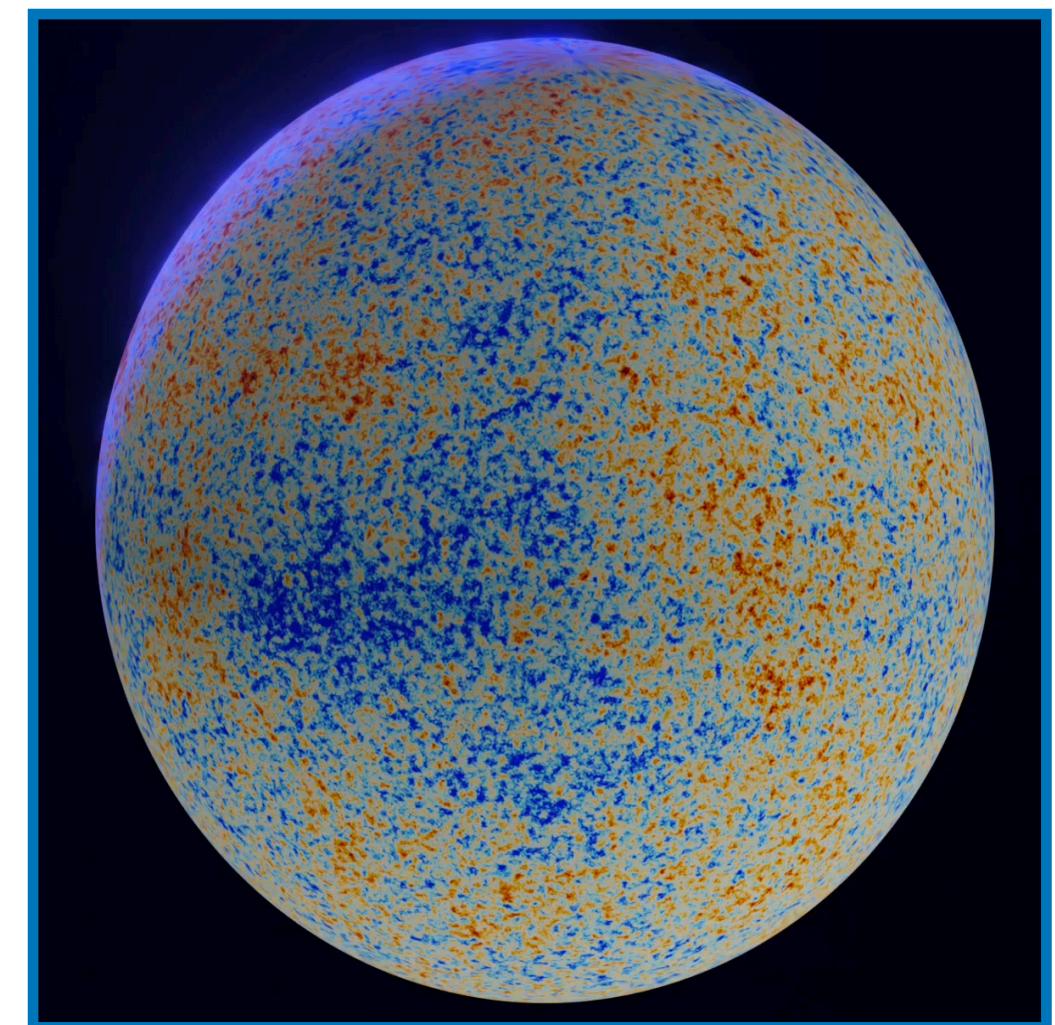
Merging galaxy clusters



Large scale structure



Cosmic microwave background



$\sim kpc$

$\xrightarrow{\hspace{1cm}}$
(Today)

$\sim Mpc$

$\xrightarrow{\hspace{1cm}}$

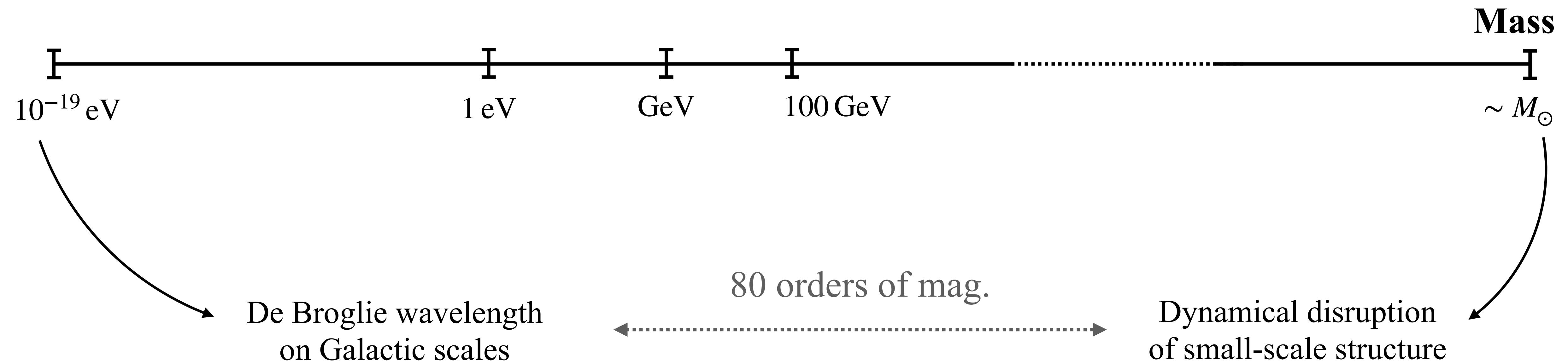
$\sim 100 \text{ Mpc}$

$\xrightarrow{\hspace{1cm}}$

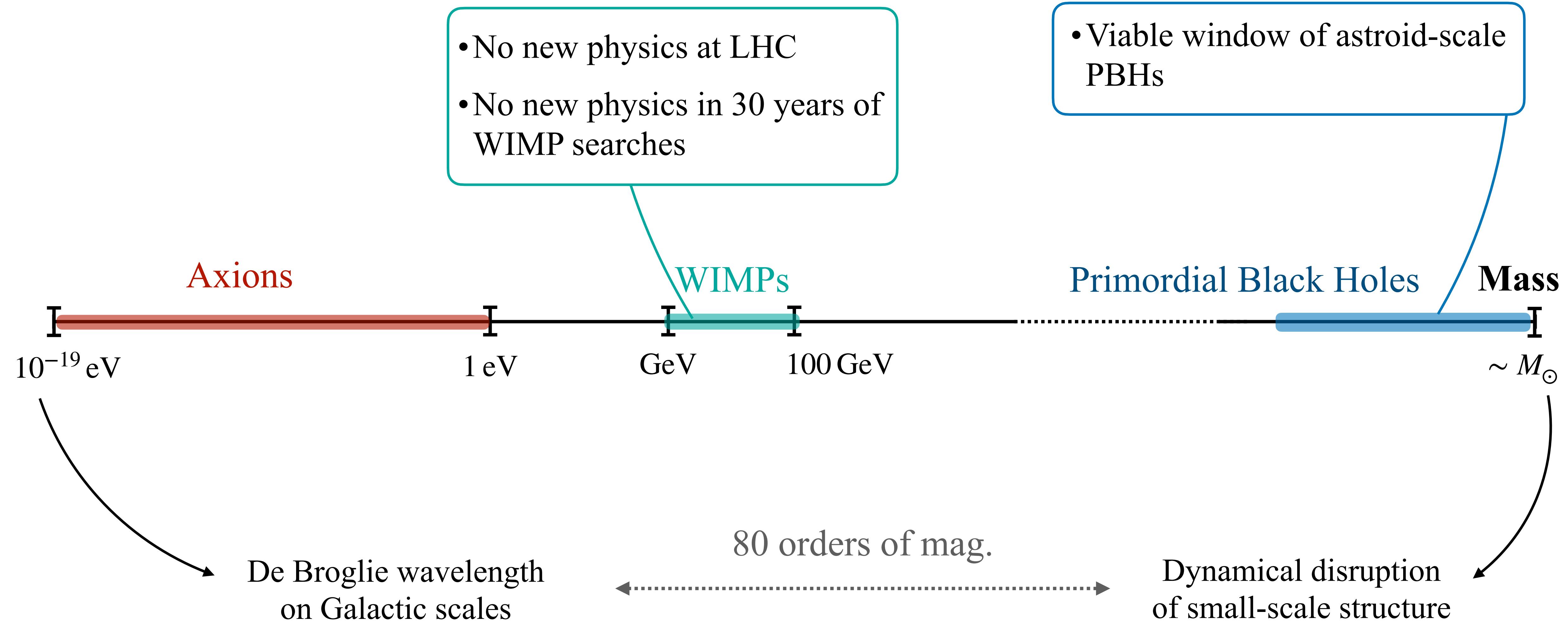
$\sim 10 \text{ Gpc}$
(370,000 yrs after big bang)

$\xrightarrow{\hspace{1cm}}$
Characteristic Scale

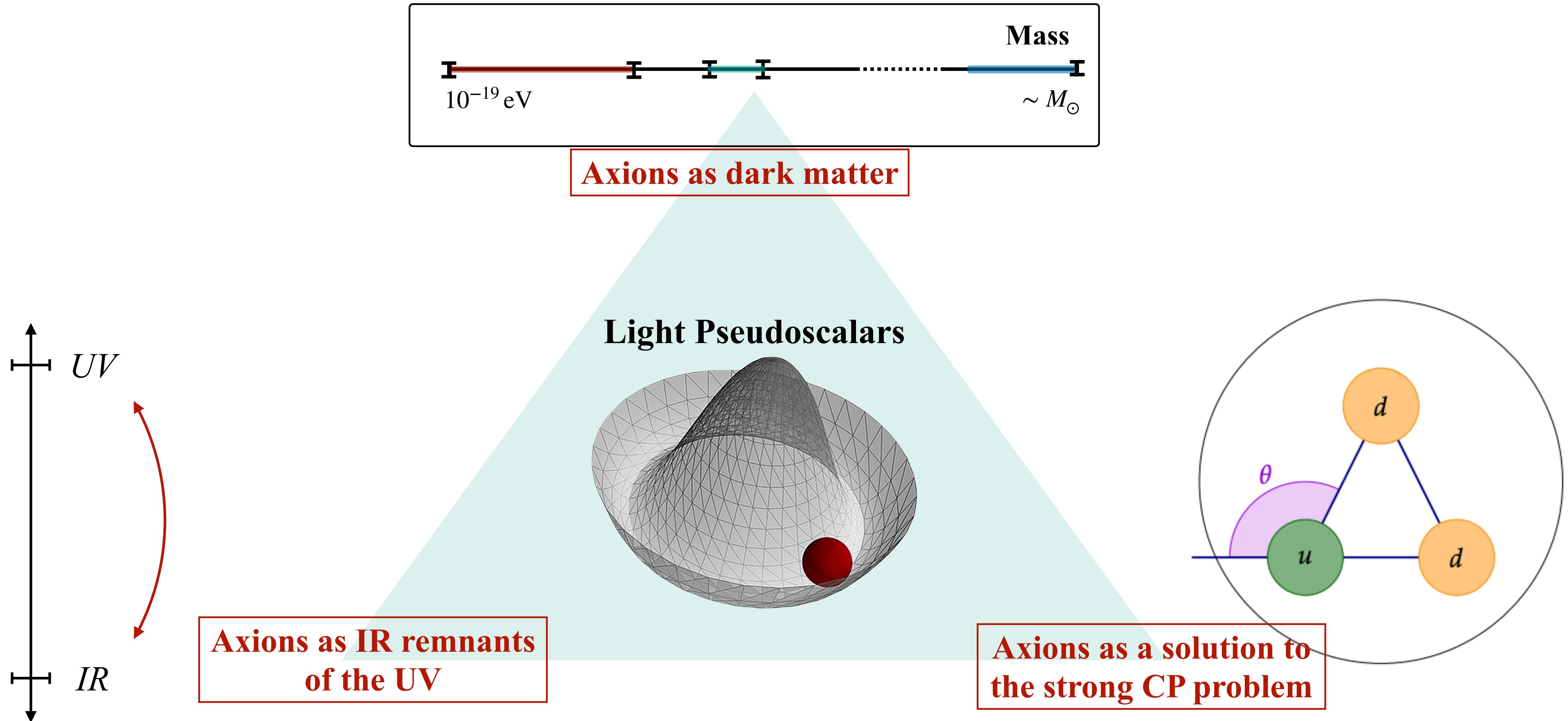
Dark matter candidates



Dark matter candidates



Axions



*How do we search for these particles and
what can we learn from their detection?*

Axions and the Standard Model

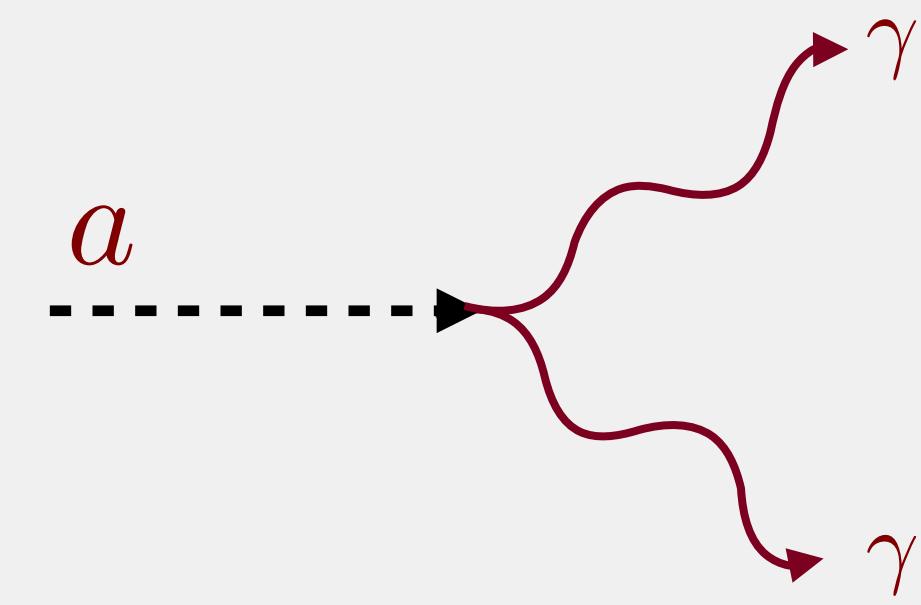
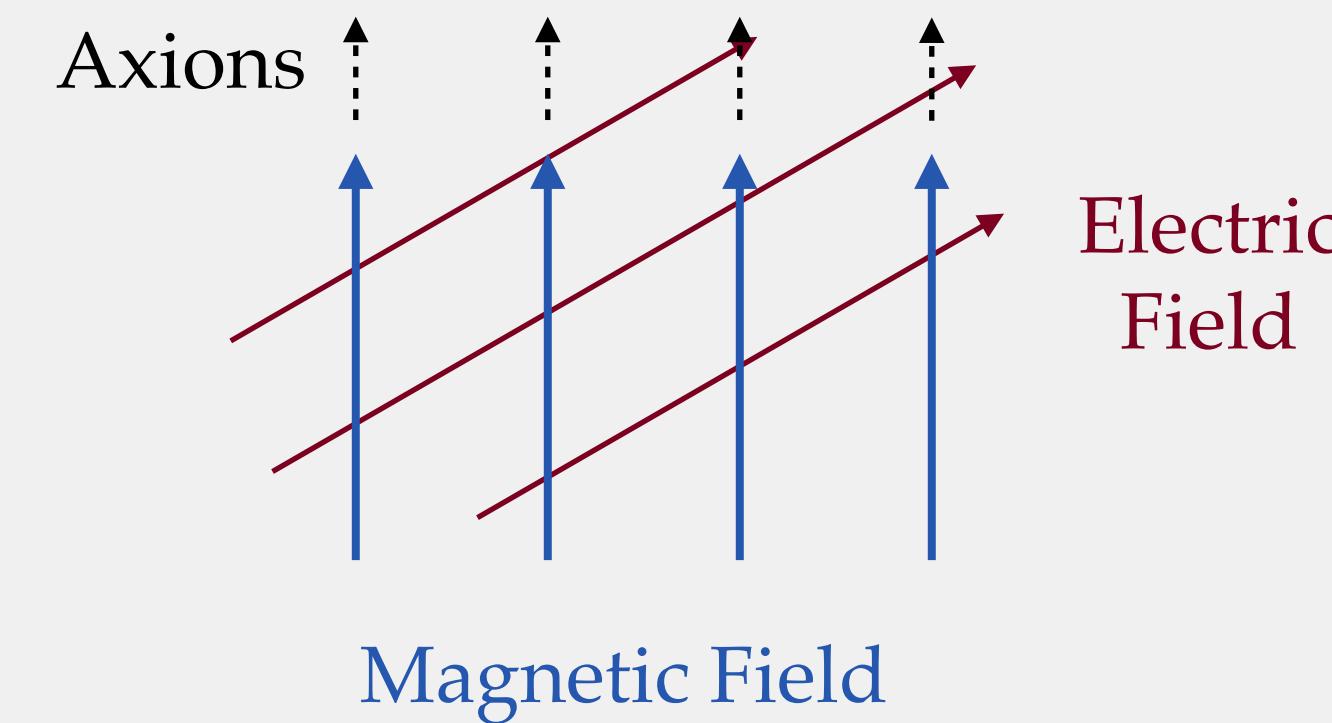
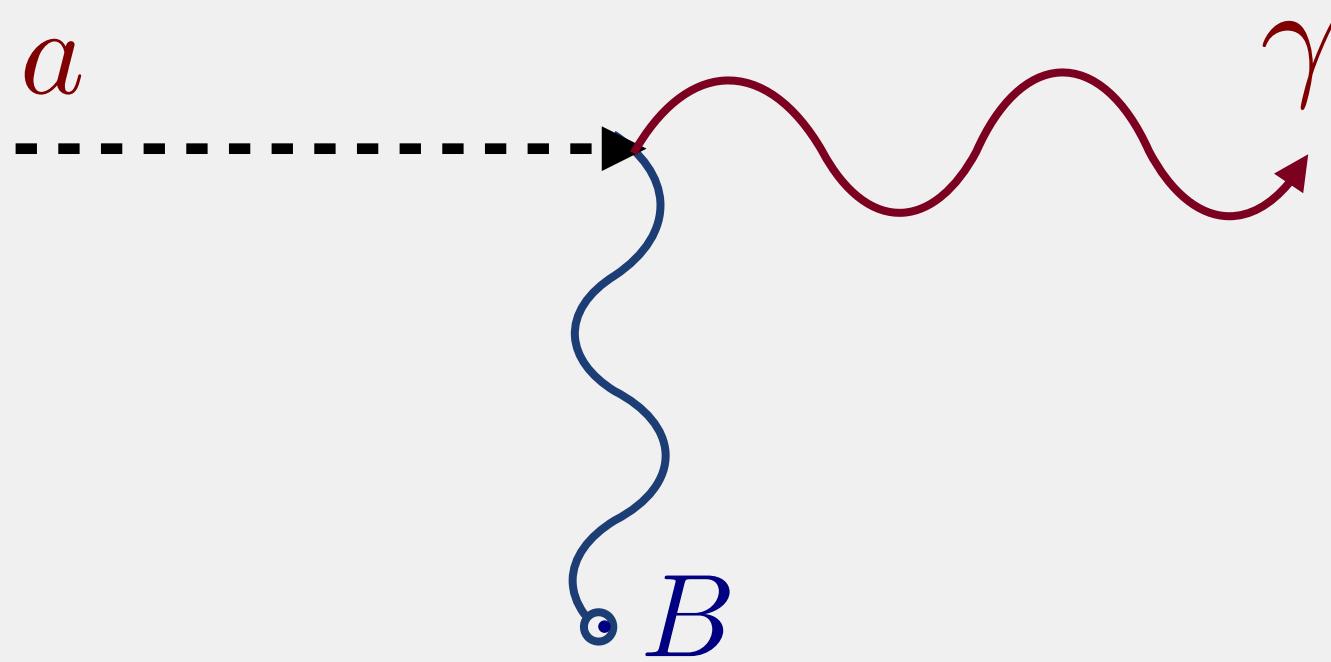
$$\mathcal{L} \supset$$

$$\mathcal{V}_{\text{QCD}}(a) +$$

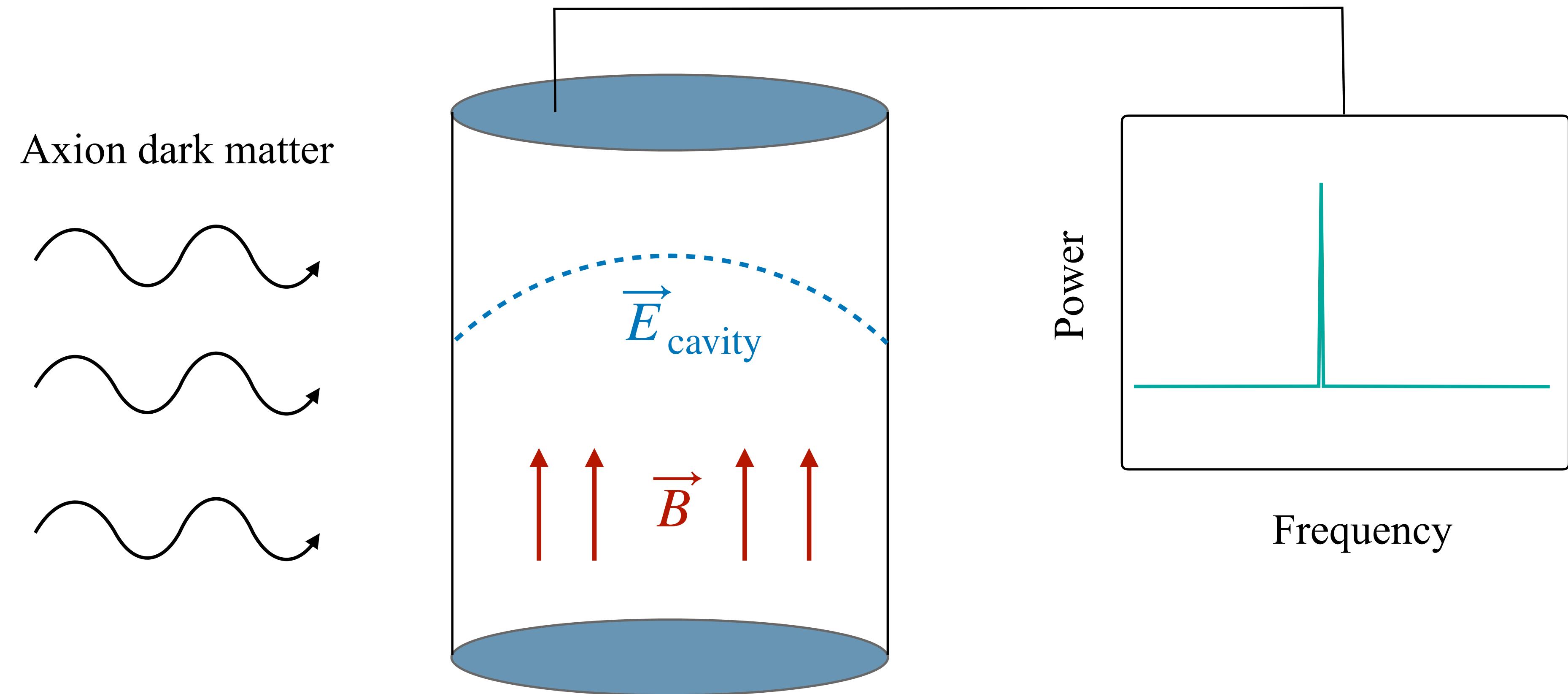
$$g_{a\gamma\gamma} a F_{\mu\nu} \tilde{F}^{\mu\nu}$$

$$+ \frac{1}{f_a} \partial_\mu a \bar{\psi} \gamma^\mu \gamma^5 \psi + \dots$$

$$\propto a (\vec{E} \cdot \vec{B})$$



Axion detection in the lab: the haloscope



Can understand as broken translational invariance allows $k_a^\mu \simeq k_\gamma^\mu + \delta k^i$

Fundamental Limitations: Magnetic field strength, must re-tune cavity for each mass

Predicting the axion mass

- Assume QCD axion is dark matter (& high inflationary scale)

Cosmic strings network in the early Universe

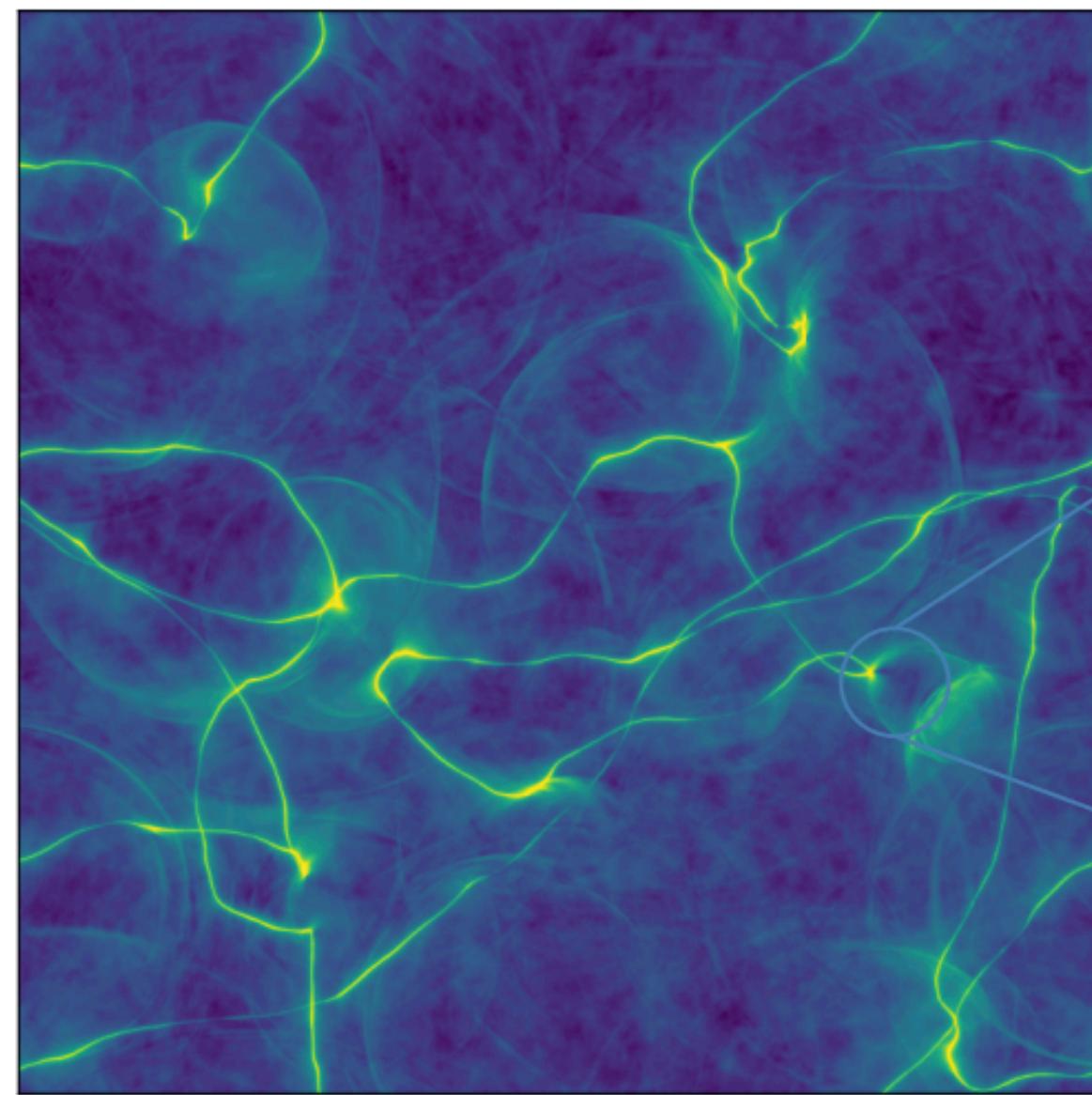


Image credit: Bernabou et al (2023)

Grilla di Cortana, **Hardy**, Pardo Vega, Villadoro (2016), Ghorgetto, **Hardy**, Villadoro (2018, 2021),
Bushmann et al (2022), Saikawa et al (2024),
Beyer & **Sarkar** (2023)

Axion density field

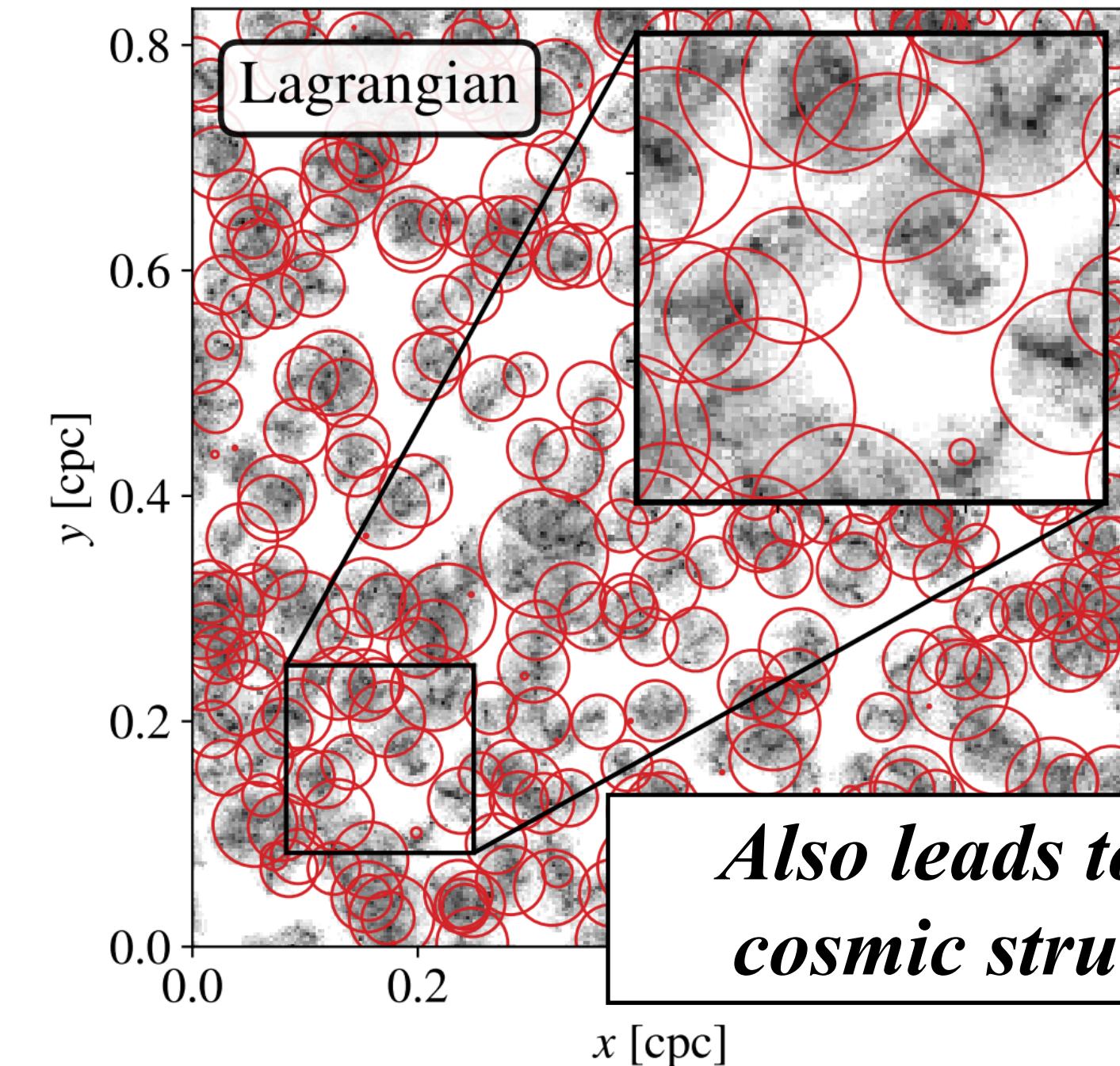


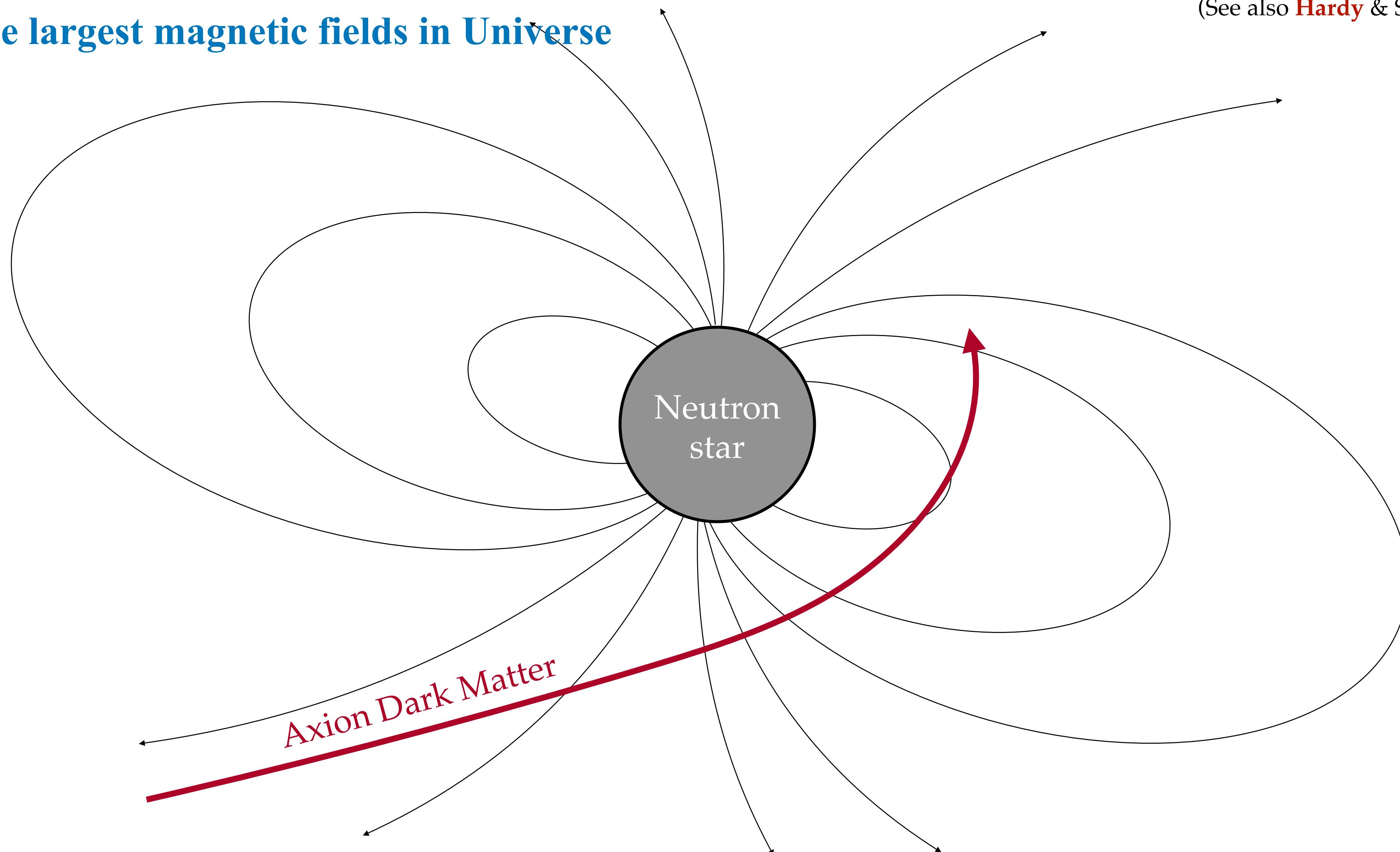
Image credit: Ellis et al (2022)

See also Ghorghetto, **Hardy**, **March-Russell**, Song and West (2023) for related idea with dark photon

Broadband resonators in Nature

In the largest magnetic fields in Universe

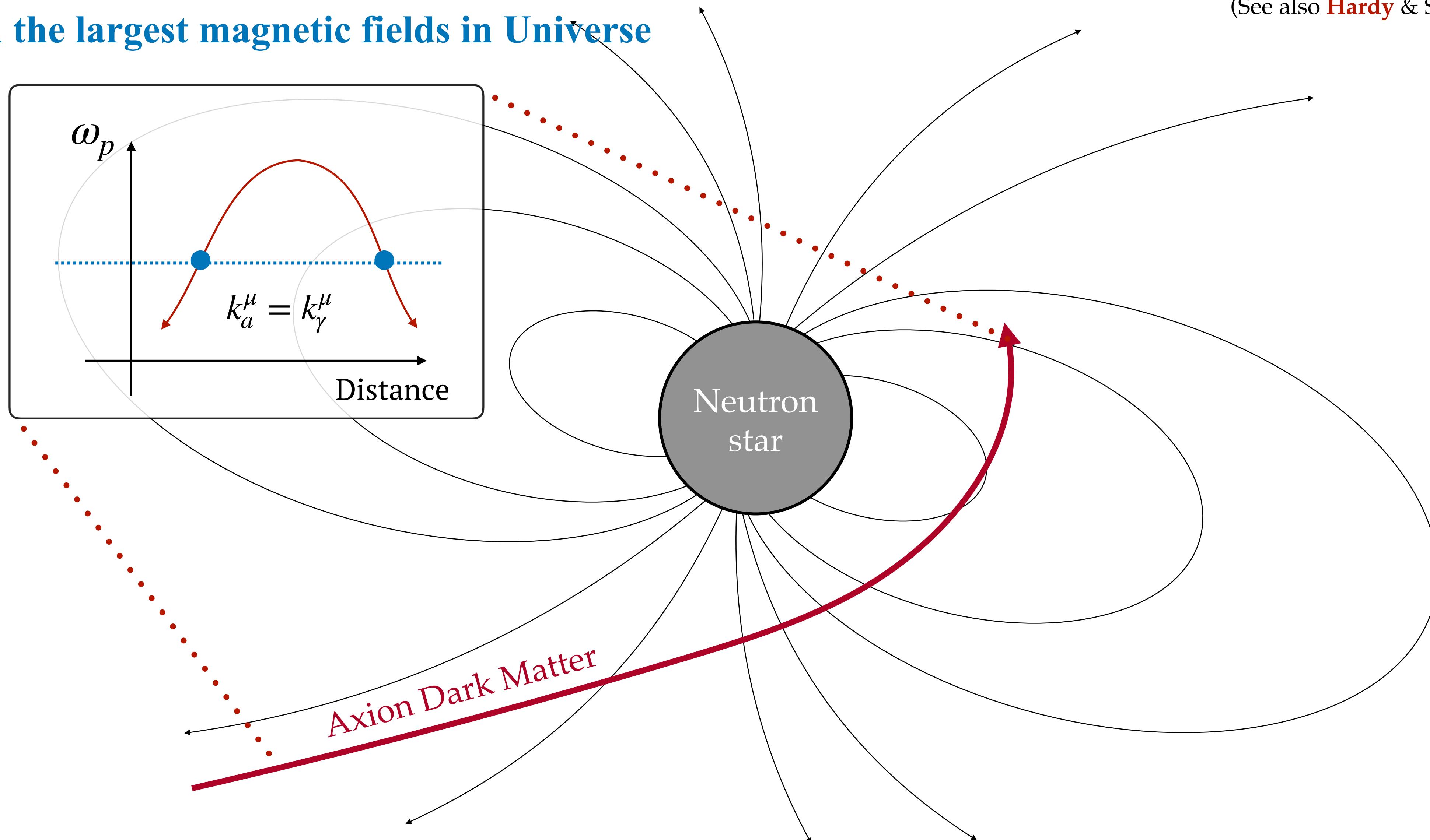
See e.g.: Pshirkov & Popov (2009), Hook et al. (2018),
Safdi et al. (2018), Battye et al. (2019, 2021, 2023),
SJW et al. (2021, 2022), Foster, **SJW** et al (2022), ...
(See also **Hardy** & Song (2023) for dark photon)



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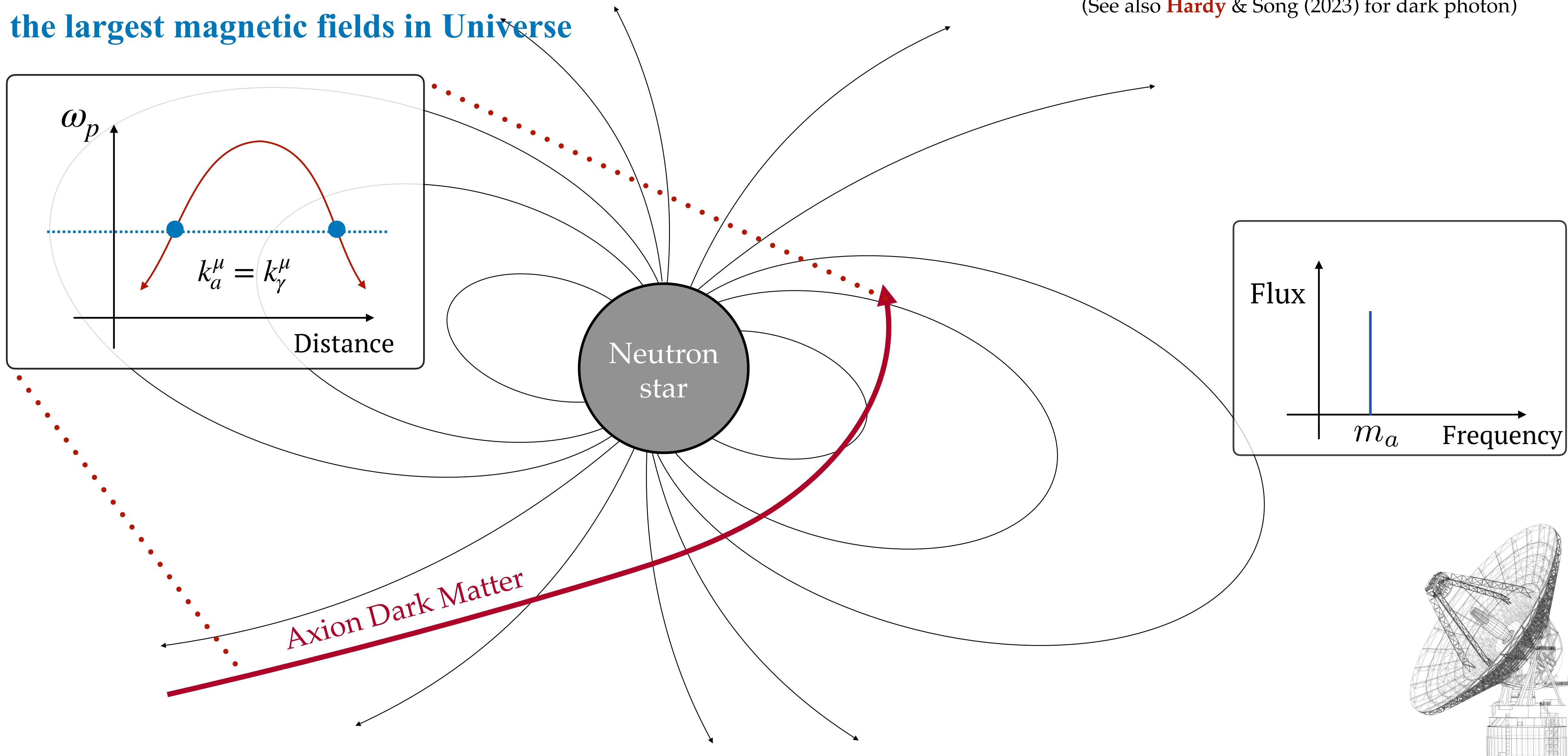
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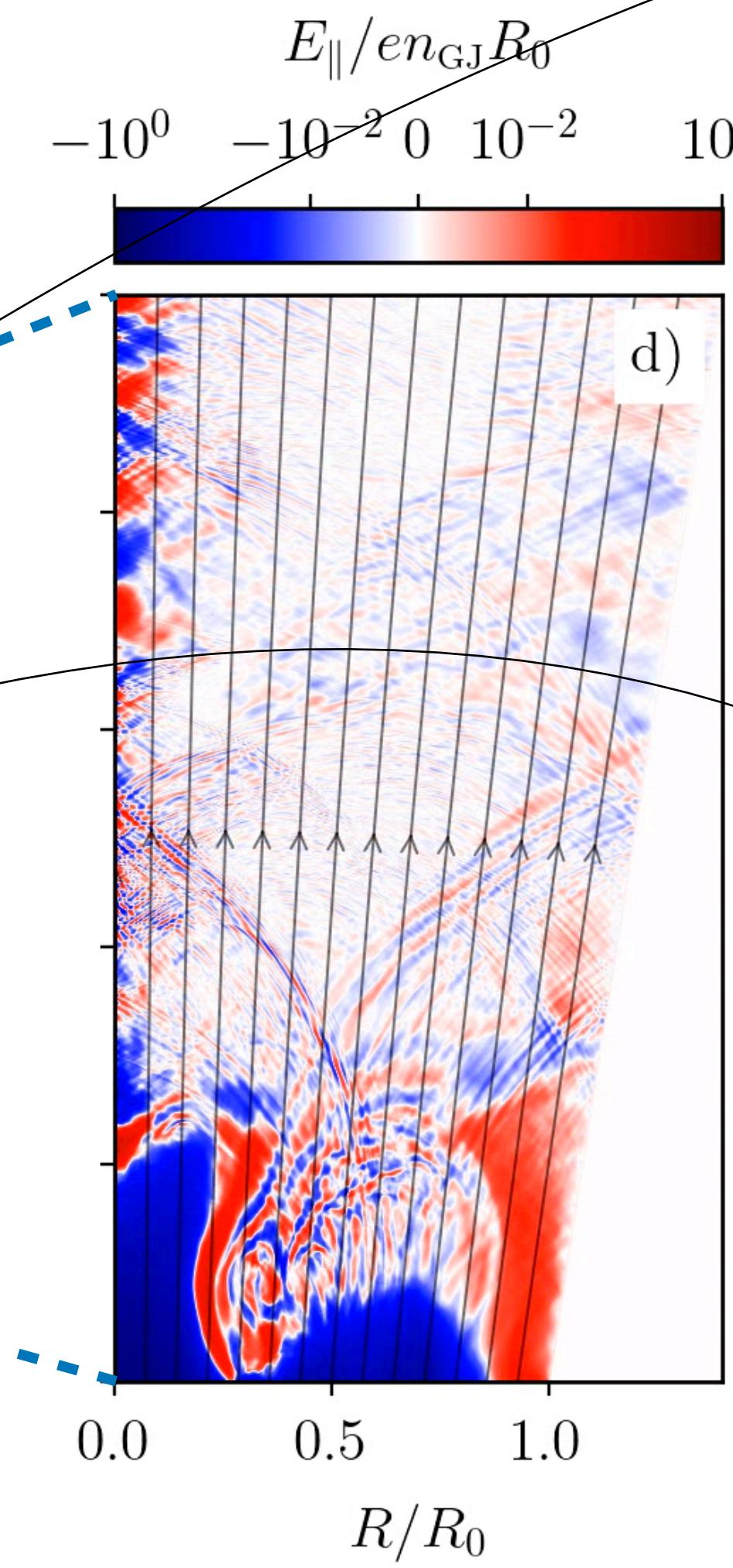
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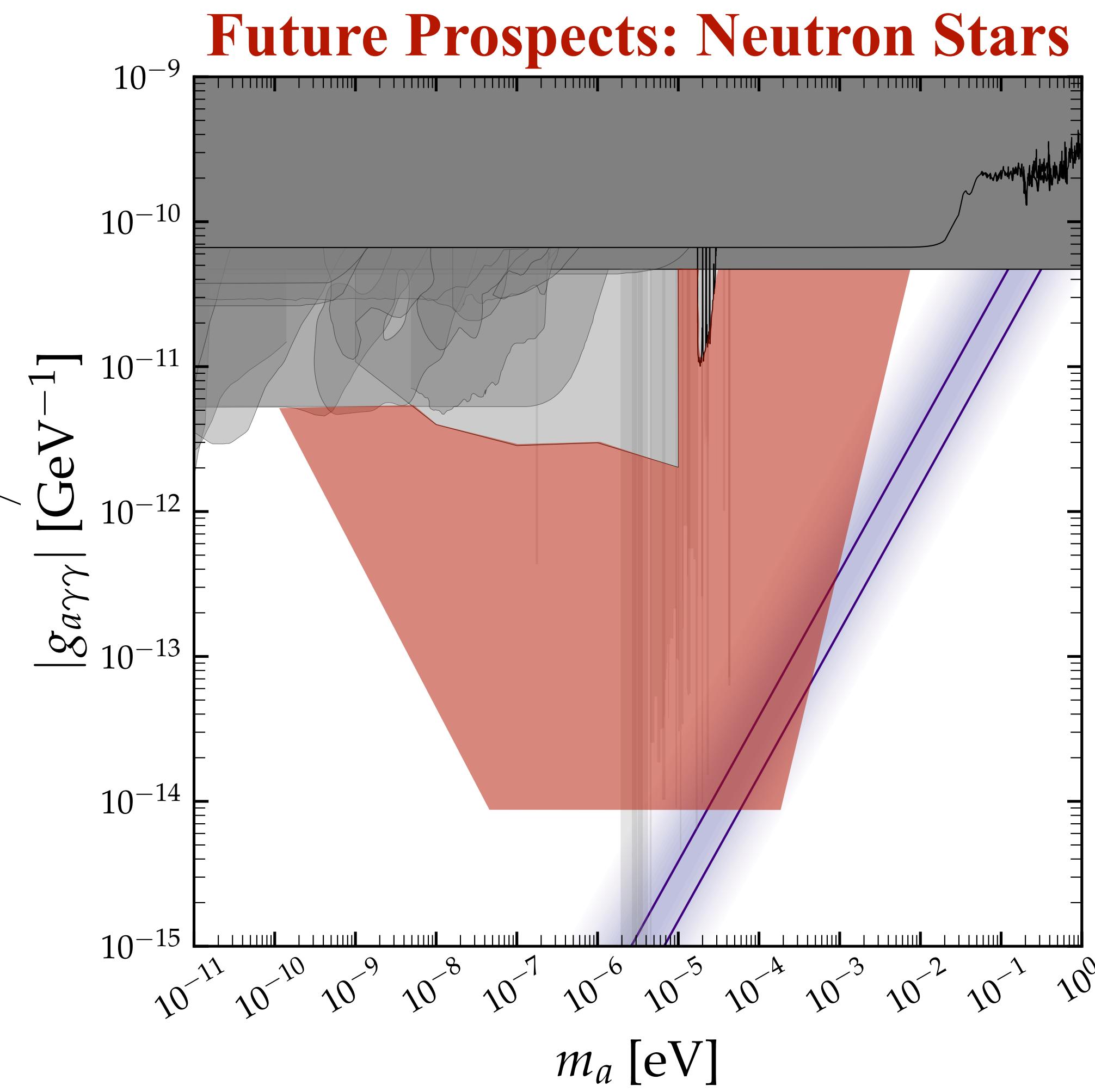
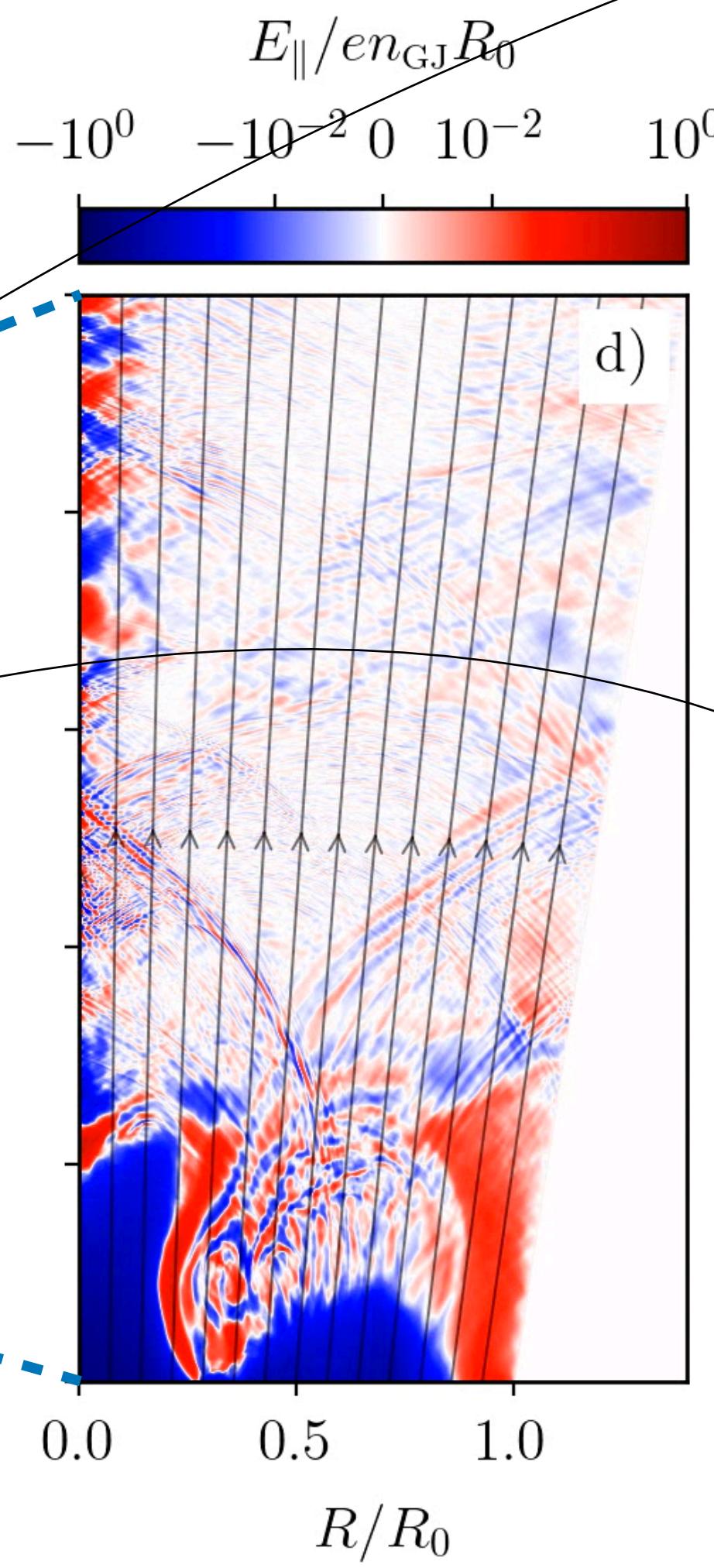
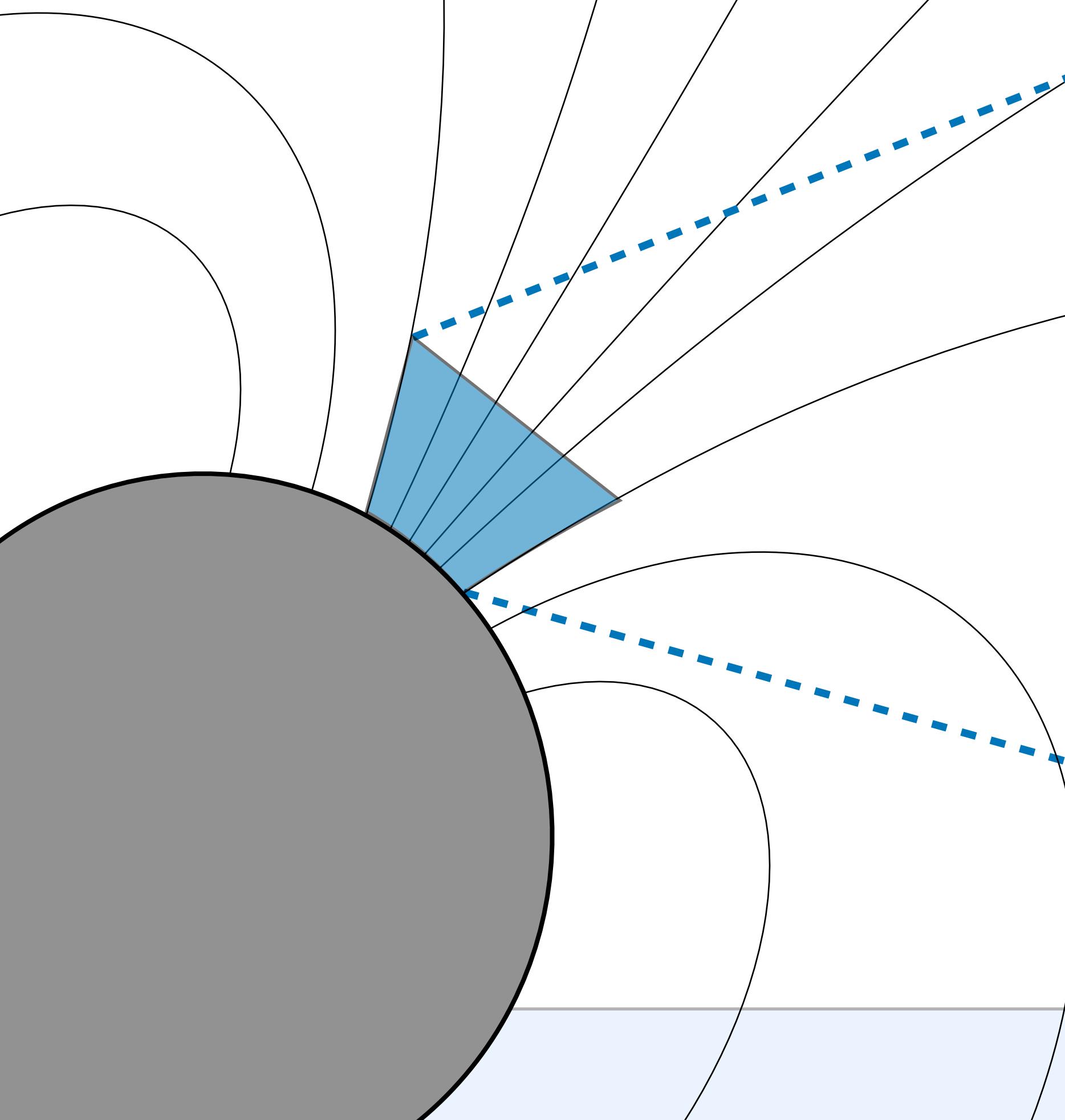
Axions beyond dark matter

Prabhu 2021, Noordhuis, Prabhu, **SJW**, Cruz, Chen, Weniger (2022), Noordhuis, Prabhu, Weniger, **SJW** (2023), Caputo, **SJW**, Philippov, Jacobson (2023)

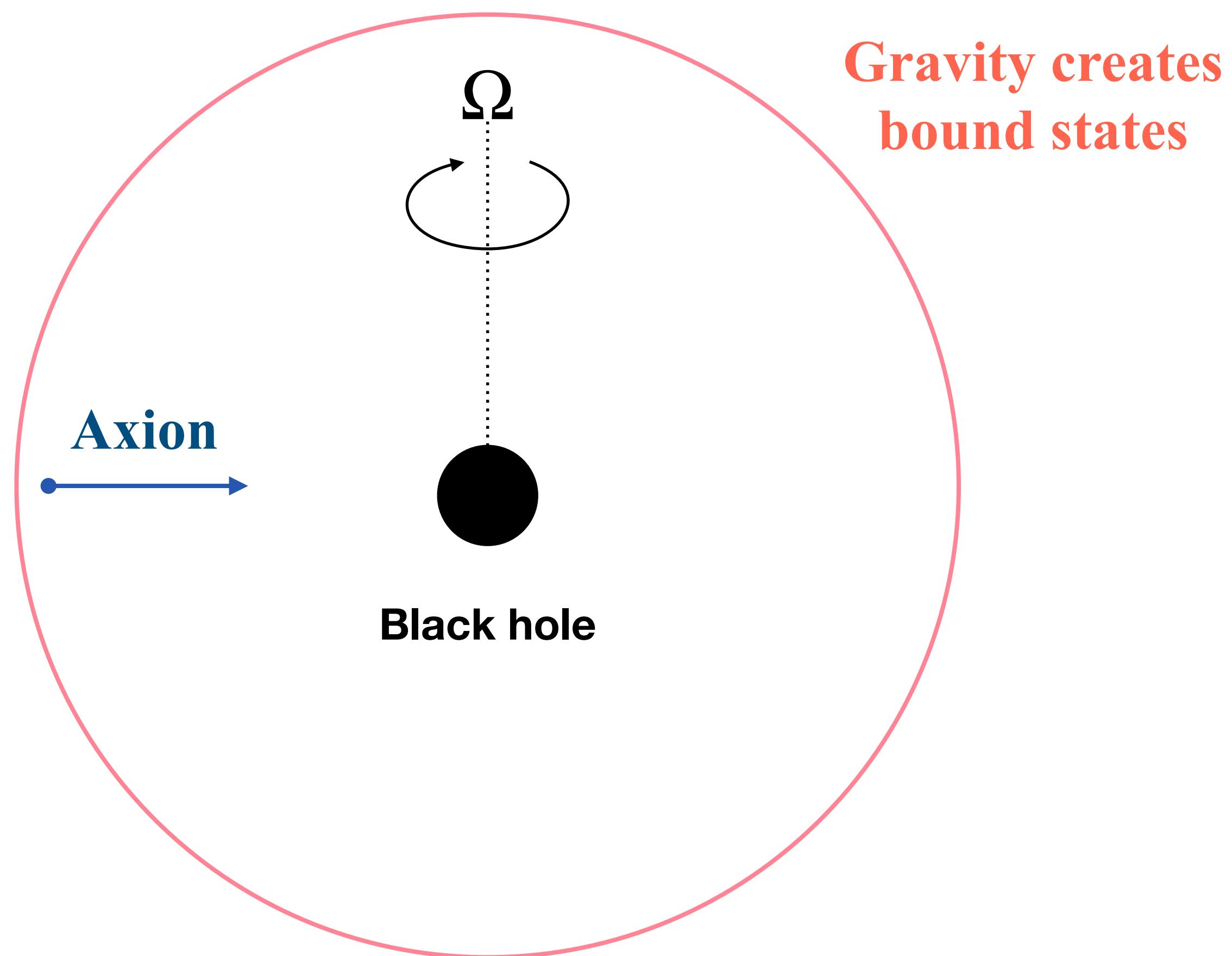


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Approaches for lighter axions: superradiance



Gravity creates
bound states

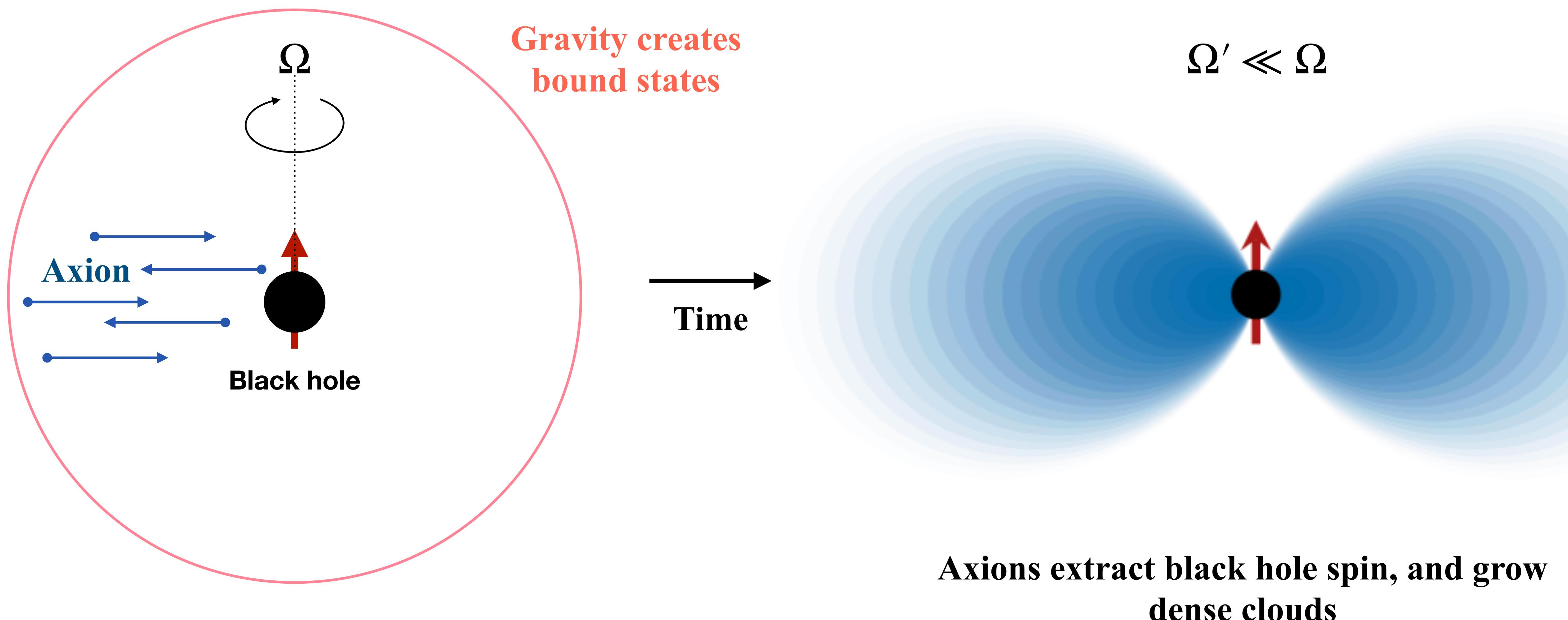
Zeldovich (1972) Press & Teukolsky (1972),

Related to **Penrose** process (1971)

Arvanitaki, Dimopoulos, Dubovsky, Kaloper, **J. March-Russell** (2010),

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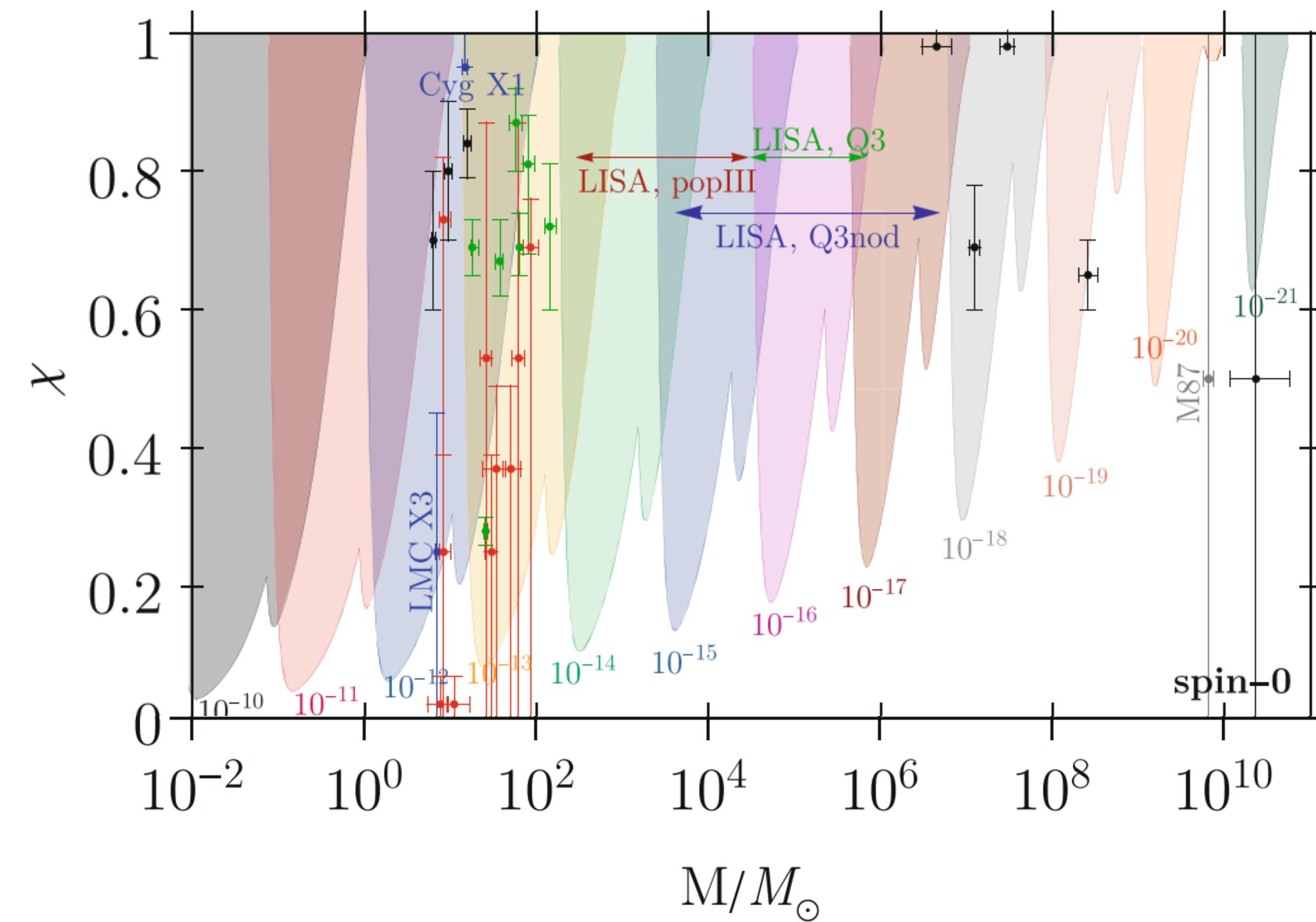
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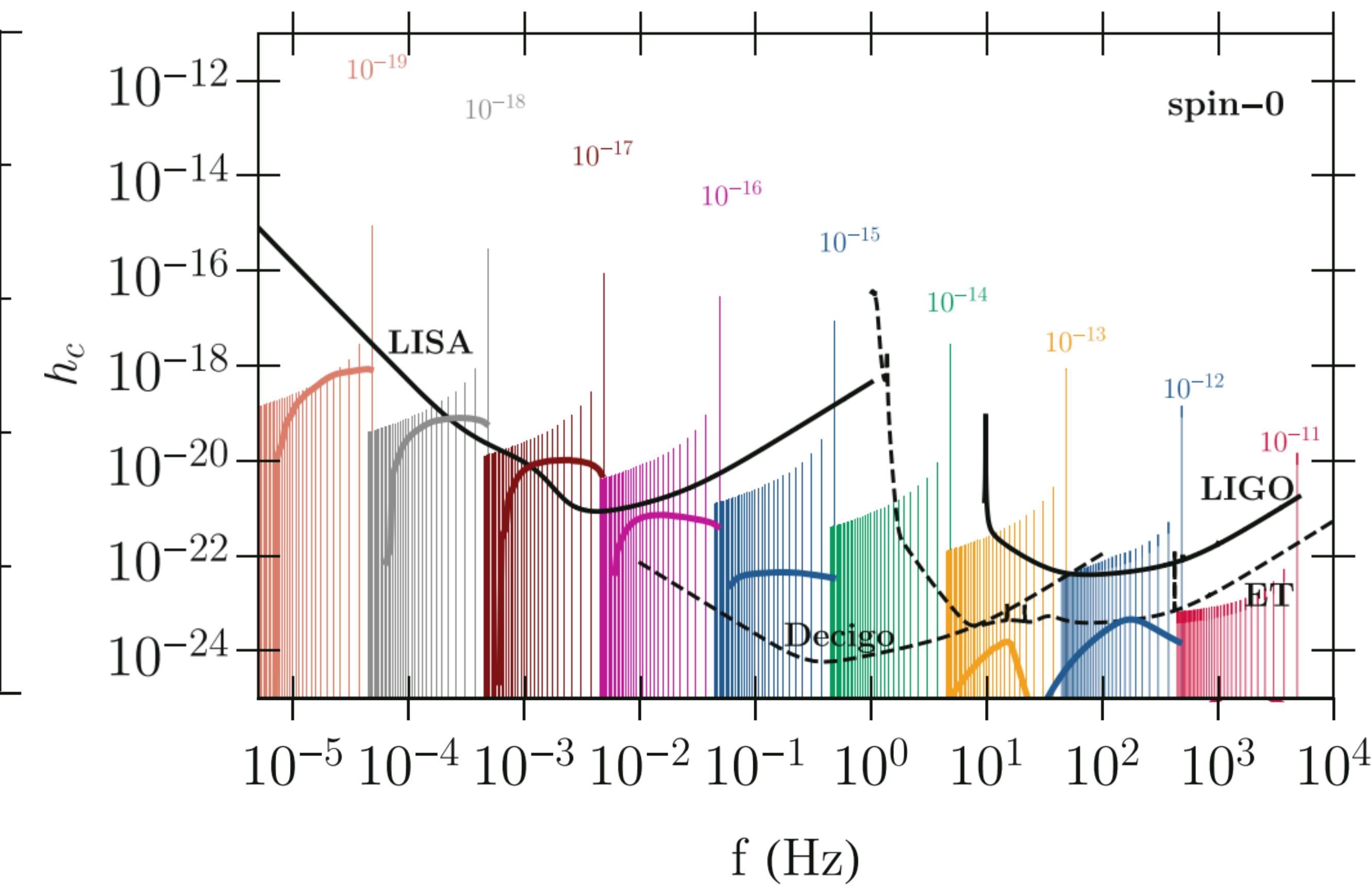
Axions extract black hole spin, and grow dense clouds

Black hole superradiance

Black hole spin distributions

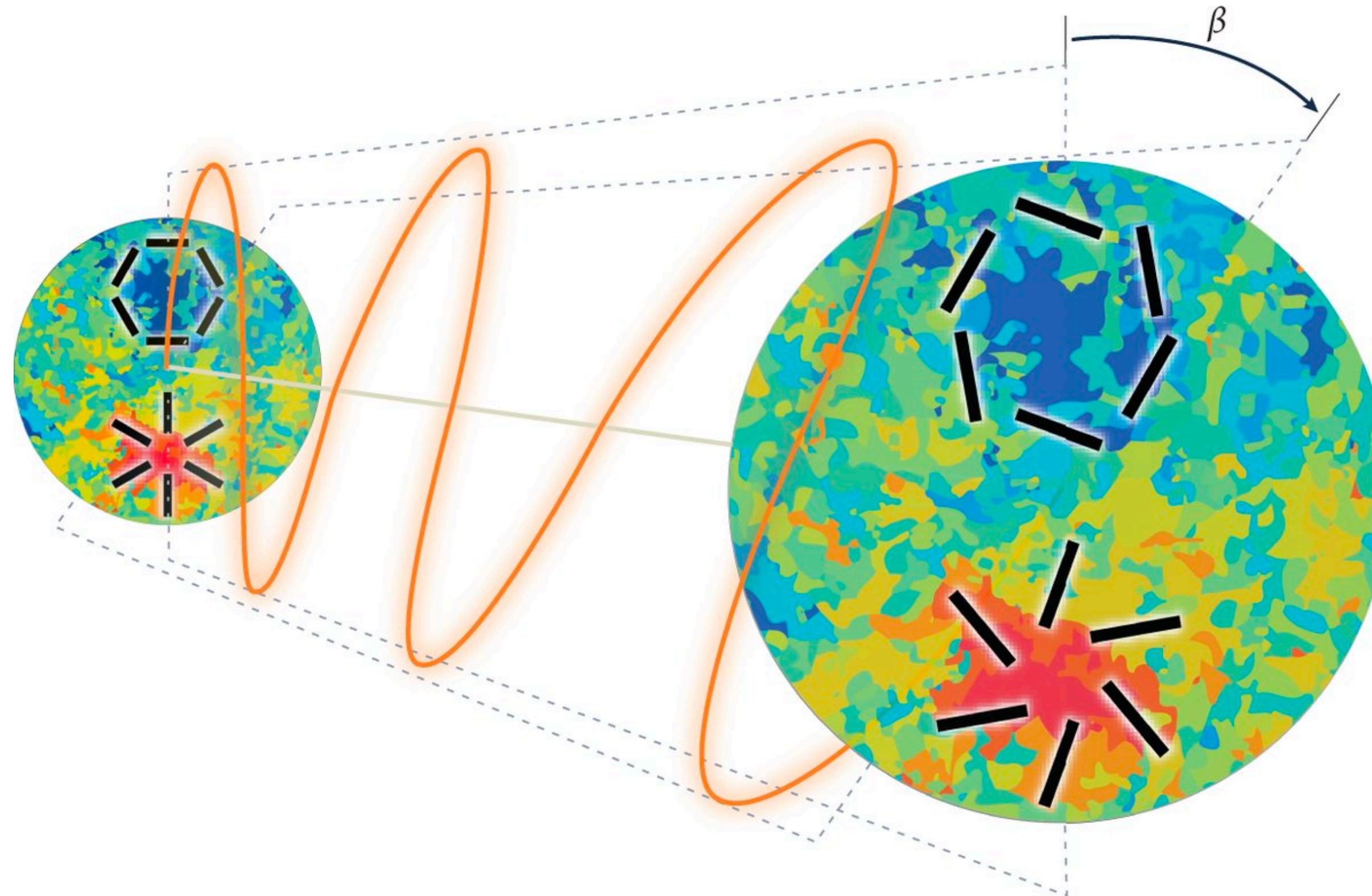


Gravitational waves from axion cloud



Approaches for *even* lighter axions

Cosmic Birefringence



Tentative evidence in Planck data @ $\sim 3.6\sigma$ See e.g. Komatsu (2022), Eskilt & Komatsu (2022)

Light axions can rotate polarisation of CMB

$$m_a \lesssim 10^{-27} \text{ eV}$$

- Quintessence

Carroll (1998)

- Axion string network

P. Agrawal, Hook, Huang (2020)

P. Agrawal, Hook, Huang, Marques-Tavares (2021)

Conclusions: so what can we learn?

- Strong CP problem?
- Dark Matter?
- What can a detection tell us about physics at high energies?

Axion detection can probe:

- GUTs / structure of gauge symmetries
- Existence of magnetic monopoles and heavy degrees of freedom

Multiple axions could hint toward stringy “axiverse”

See e.g.

- P. Agrawal, Nee, M. Reig (2022)
P. Agrawal & Platshorre (2023)
Sokolov & Ringwald (2022, 2023)
P. Agrawal, Hook, Huang (2020)
P. Agrawal, Hook, Huang, Marques-Tavares (2021)

Arvanitaki, Dimopoulos, Dubovsky, Kaloper, J. March-Russell (2010)

