



SCALES 1st General Meeting - Superfluid Condensates
in Astrophysics and Laboratory Experiments
June 23, 2026

BEC phase transitions in a system of interacting relativistic bosonic particles and antiparticles at finite temperatures

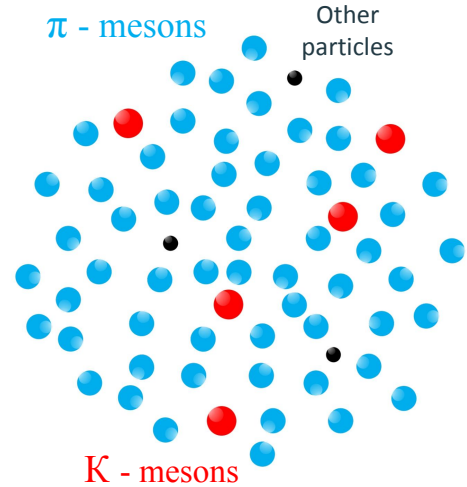
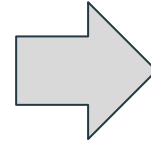
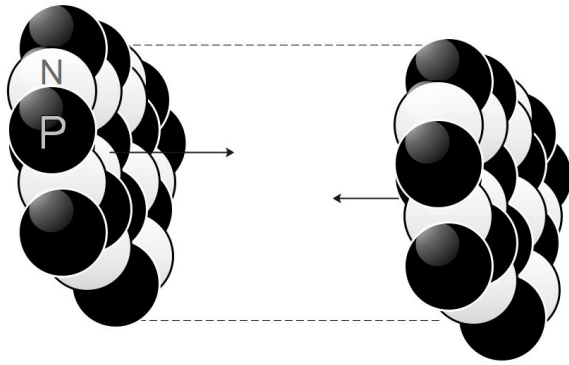
Speaker: Denys Zhuravel

In collaboration with **Dmytro Anchishkin**, **Volodymyr Gnatovskyy**,
Vladyslav Karpenko

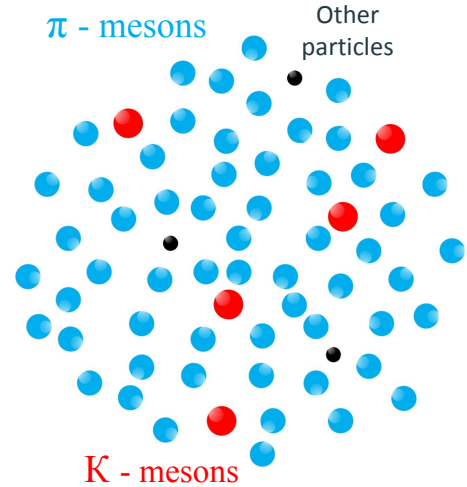
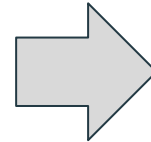
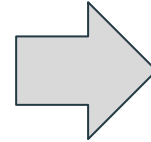
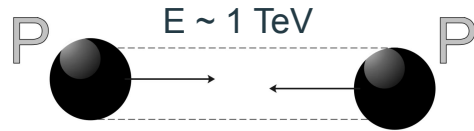
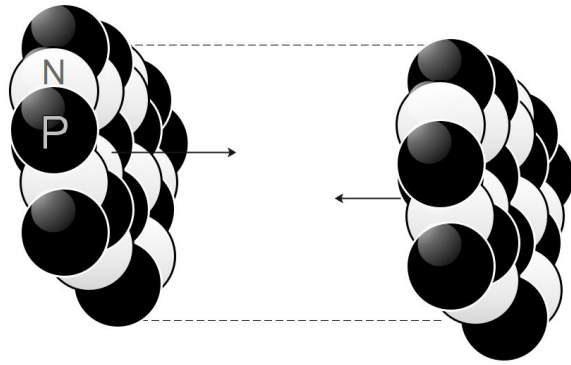
Bogolyubov Institute for Theoretical Physics



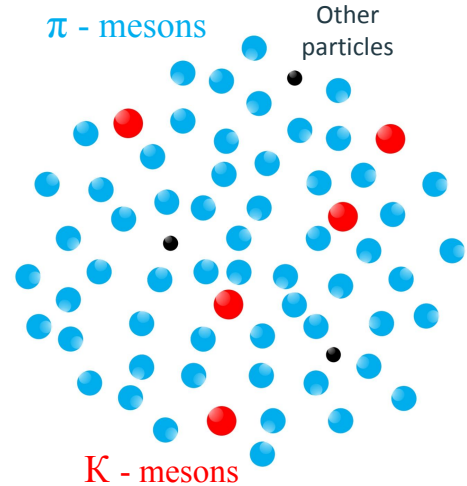
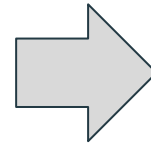
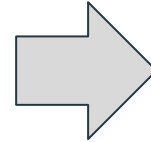
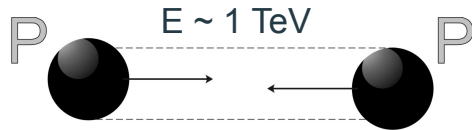
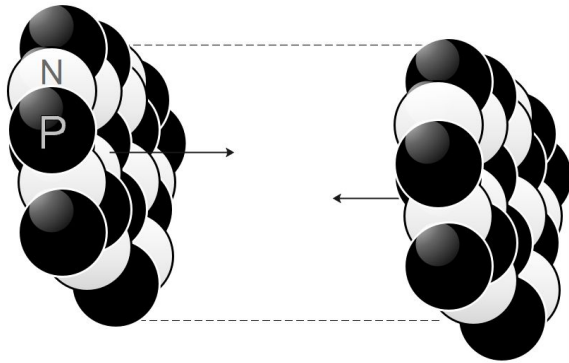
High energy collisions



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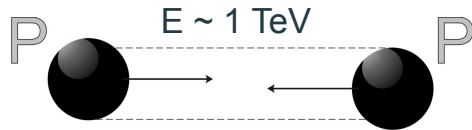
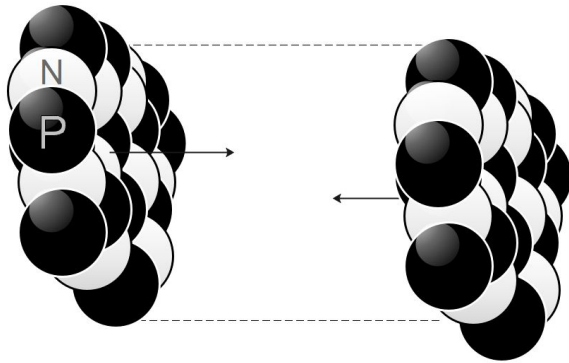
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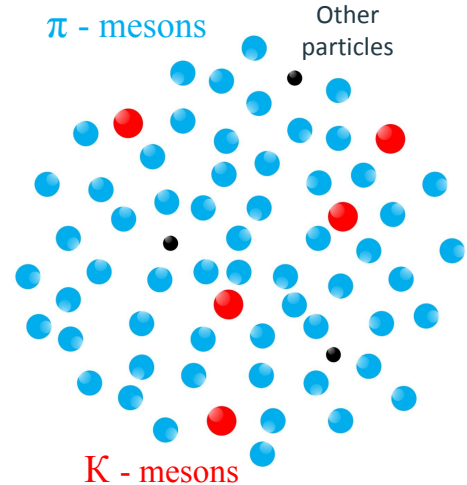
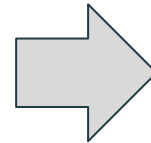
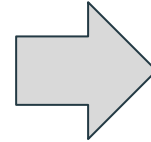
$N \sim 10^2 - 10^4$

Probability < 1%

High energy collisions



Task: Study of the conditions for the occurrence of Bose condensate in the pion-antipion system with interaction



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Probability < 1%

Historical remarks and Bose-Einstein distribution

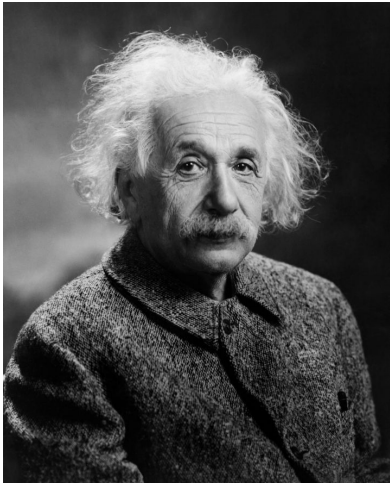
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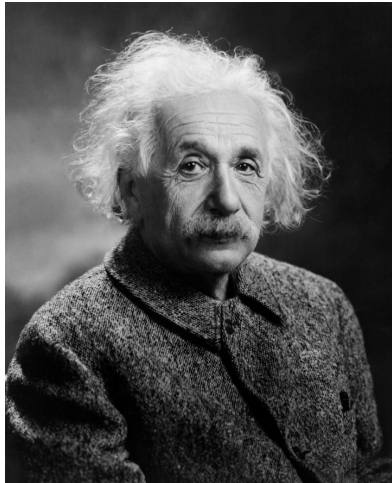
Satyendra Nath Bose

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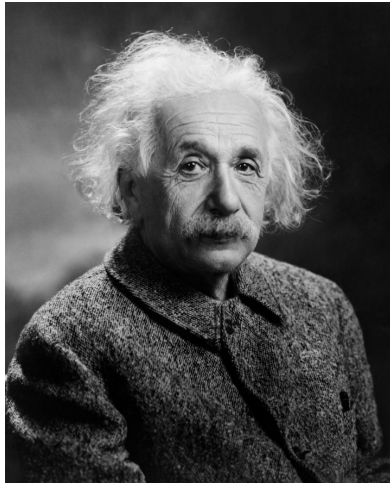


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Introduced a most remarkable derivation of Planck's radiation formula based on quantum theory and not a classical one.

$$E = \sum_s \frac{8\pi h\nu^{s3}}{c^3} V \frac{1}{e^{\frac{h\nu^s}{kT}} - 1} d\nu^s$$

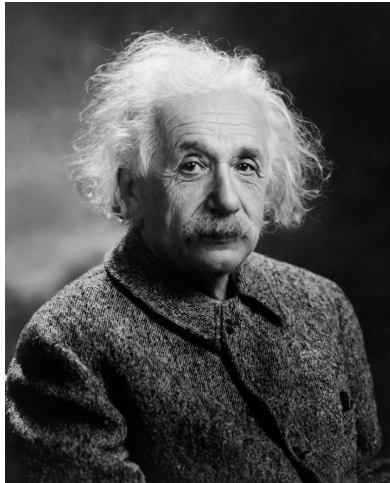


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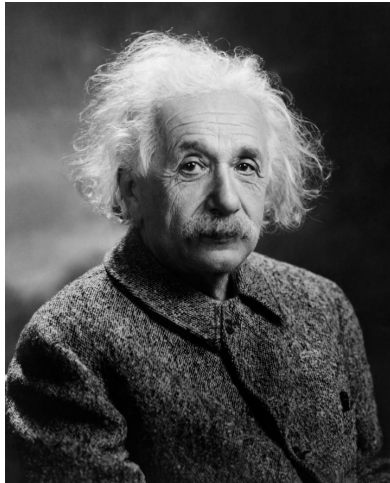


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1st paper: Extended Bose's statistical method from photons to massive Bose particles.
2nd paper: Introduced ideas of bosons condensation



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When does Bose-Einstein
condensation (BEC) occur?

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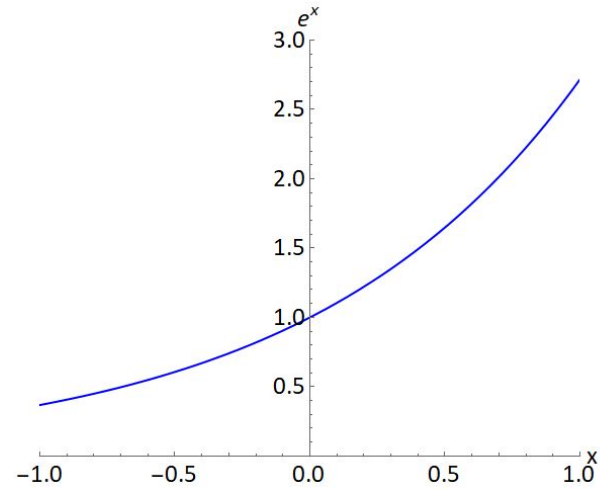
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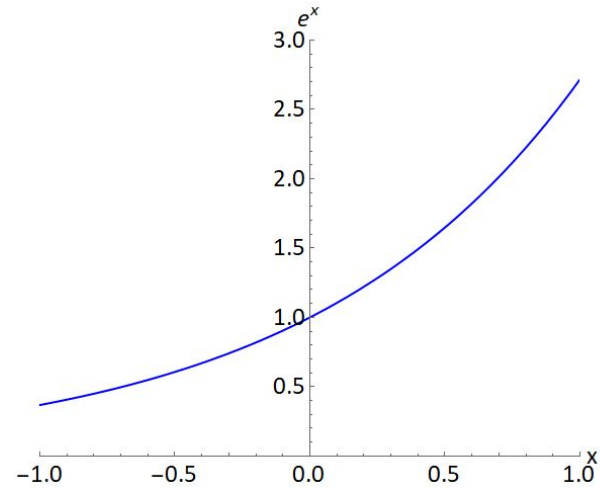


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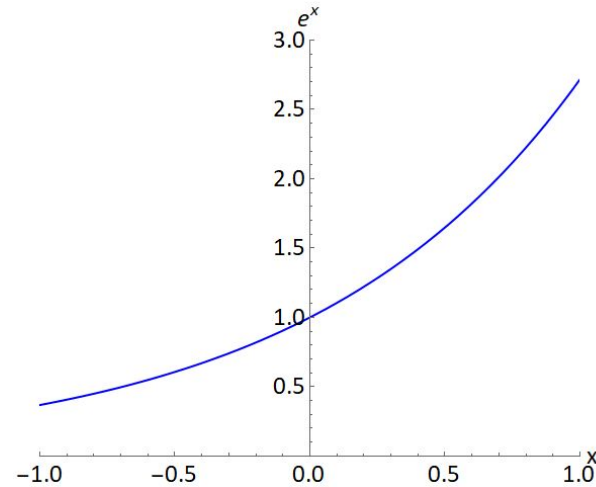
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When does Bose-Einstein condensation (BEC) occur?

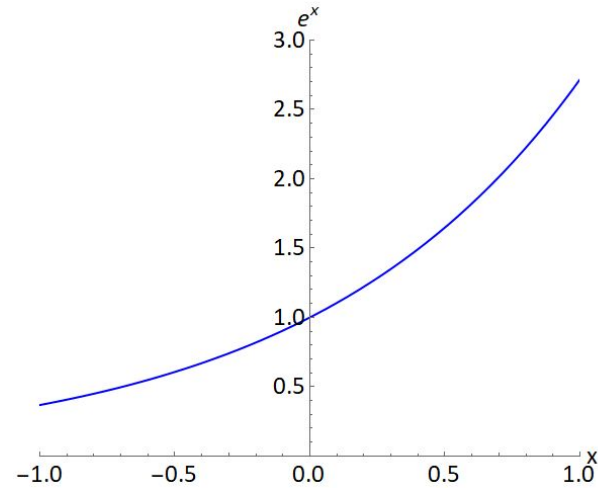
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Energy levels of system:

$$\varepsilon_0 < \varepsilon_1 < \varepsilon_2 < \dots$$

Ground
state



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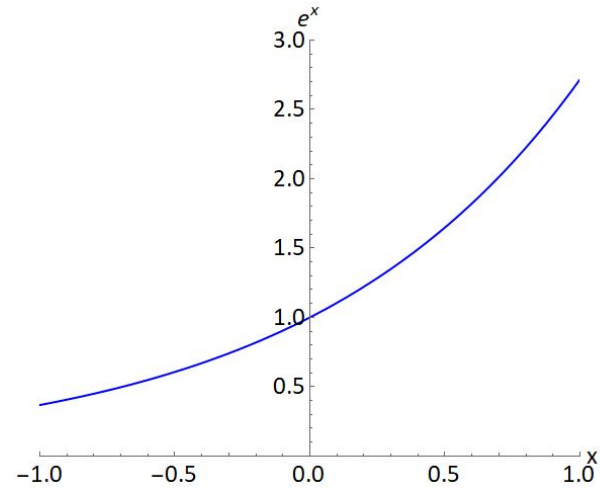
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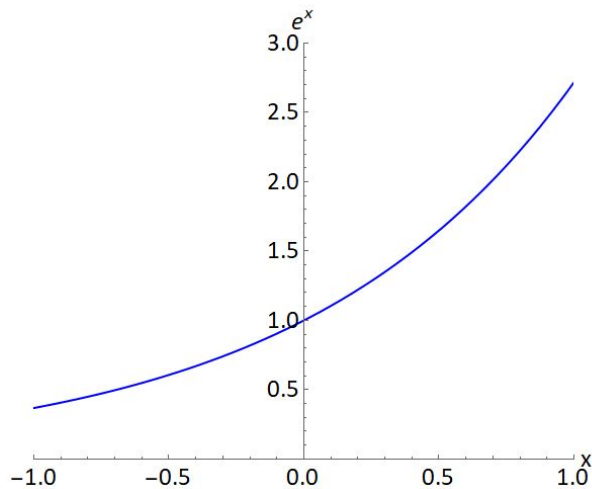
Ground
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$$\varepsilon_0 = \mu$$



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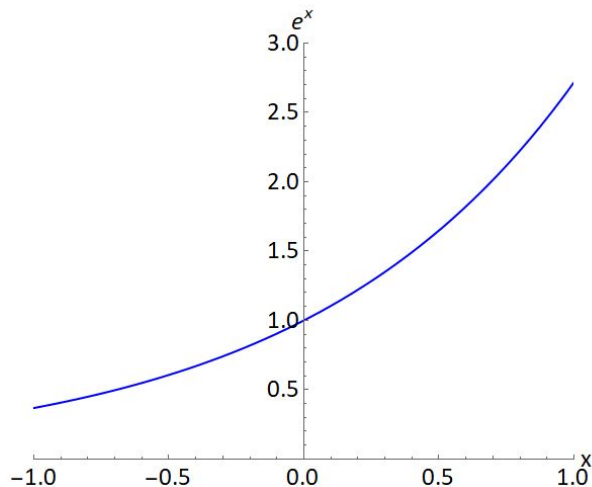


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Condition for the onset of Bose-Einstein condensation

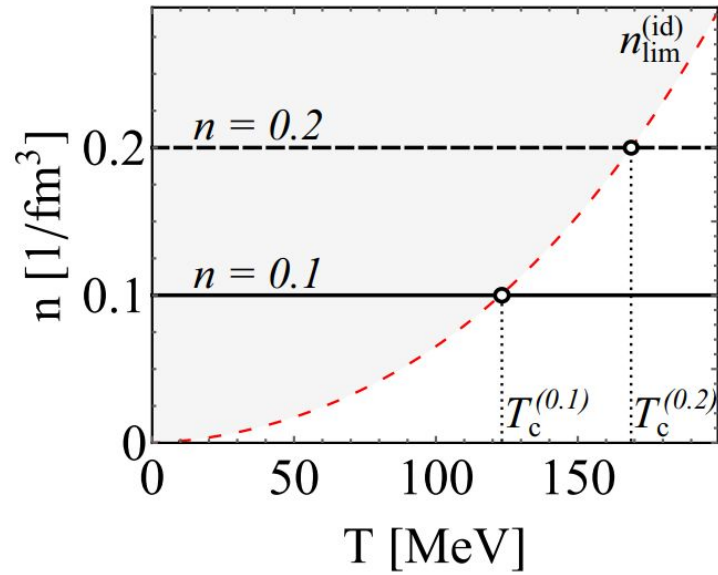


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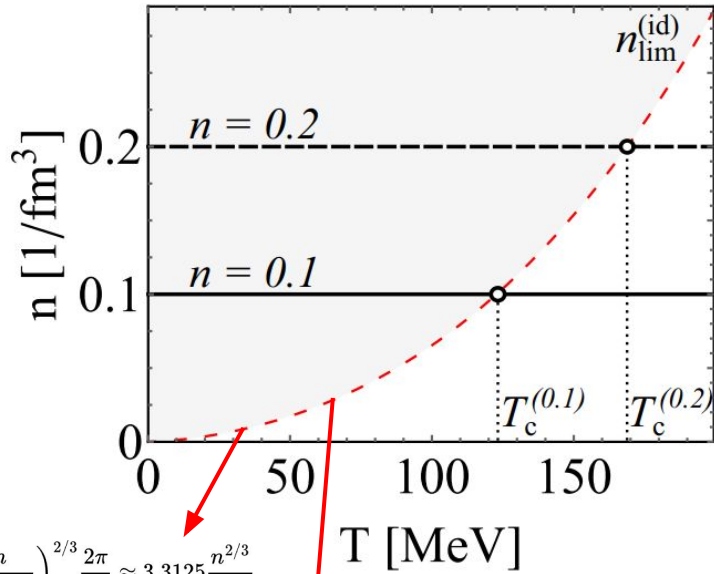
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1-component ideal gas



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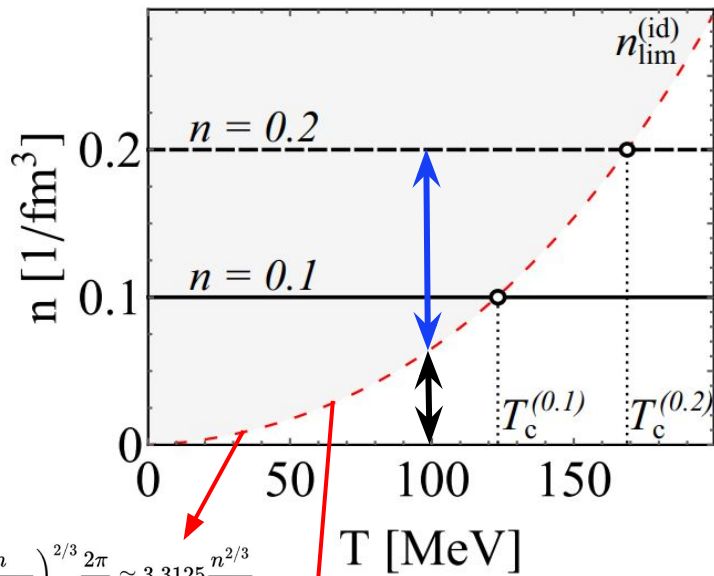
$$T_c = \left(\frac{n}{\zeta(3/2)} \right)^{2/3} \frac{2\pi}{m} \approx 3.3125 \frac{n^{2/3}}{m}$$

Classical system

$$n_{\text{lim}}^{(\text{id})}(T) = g \int_0^\infty \frac{d^3k}{(2\pi)^3} \frac{1}{\exp\left[\frac{\sqrt{m^2+k^2}-m}{T}\right] - 1}$$

Relativistic system

1-component ideal gas



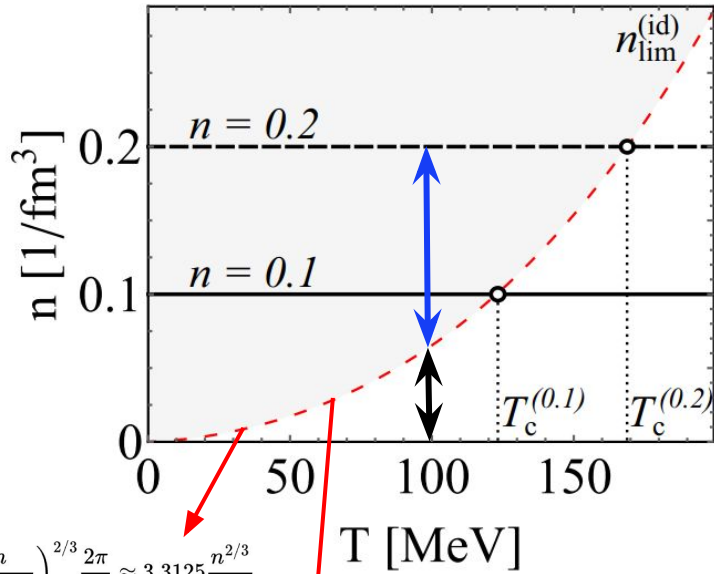
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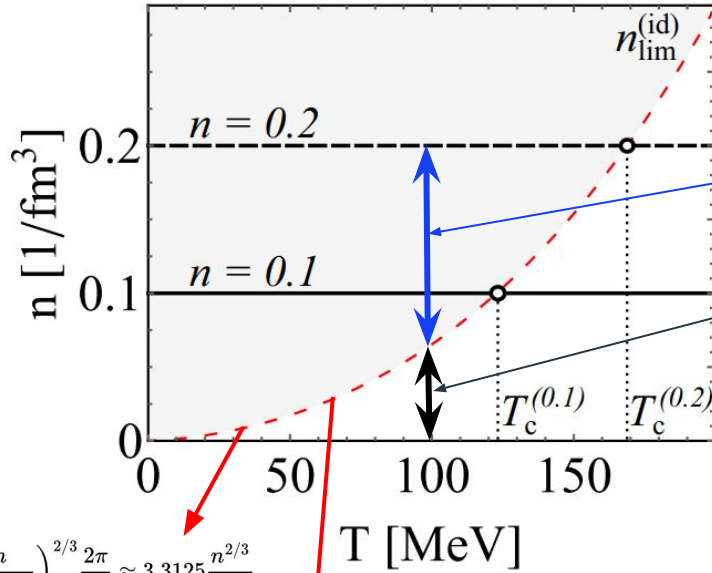
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Relativistic system

How to describe 2-component
system with interaction?

Scalar field model^[3,4]

Thermodynamic mean-field model^[1-3]

Starting
point

The Lagrangian density:

$$\mathcal{L}(x) = \partial_\mu \phi^+(x) \partial^\mu \phi(x) - m^2 \phi^+(x) \phi(x) + \mathcal{L}_{\text{int}}(x)$$

Free energy density:

$$\phi(n_1, n_2, T) = \phi_1^{(0)}(n_1, T) + \phi_2^{(0)}(n_2, T) + \phi_{\text{int}}(n, T),$$

Self-consistent set
of equations

Parameterization
of interaction

[1] D. Anchishkin and V. Vovchenko, J. Phys. G 42, 105102 (2015).

[2] D. Anchishkin, I. Mishustin, and H. Stoecker, J. Phys. G 46, 035002 (2019).

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[4] I. N. Mishustin, D. V. Anchishkin, L. M. Satarov, O. S. Stashko, and H. Stoecker, Phys. Rev. C 100, 022201(2019).

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Fixed

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$$f_{\text{BE}}(\omega, \mu) = \left\{ \exp \left[\frac{\omega - \mu}{T} \right] - 1 \right\}^{-1}, \quad \hat{\sigma}(x) = \phi^+(x) \phi(x)$$



Parameterization of interaction

$$m = 140 \text{ MeV}$$

$$\mathcal{L}_{\text{int}}(\sigma) = \frac{a}{2} \sigma^2 - \frac{b}{3} \sigma^3, \quad \omega - \mu_I = 0$$

$$\kappa = \frac{a}{2m\sqrt{b}}.$$

Thermodynamic mean-field model^[1-3]

Free energy density:

$$\phi(n_1, n_2, T) = \phi_1^{(0)}(n_1, T) + \phi_2^{(0)}(n_2, T) + \phi_{\text{int}}(n, T),$$



$$n = \int \frac{d^3k}{(2\pi)^3} [f_{\text{BE}}(E(k, n), \mu_I) + f_{\text{BE}}(E(k, n), -\mu_I)],$$

$$n_I = \int \frac{d^3k}{(2\pi)^3} [f_{\text{BE}}(E(k, n), \mu_I) - f_{\text{BE}}(E(k, n), -\mu_I)].$$

Fixed

$$E(k, n) = \sqrt{m^2 + k^2} + U(n)$$

$$f_{\text{BE}}(E, \mu) = \left\{ \exp \left[\frac{E - \mu}{T} \right] - 1 \right\}^{-1}$$



$$U(n) = -An + Bn^2, \quad E - \mu = 0$$

$$\kappa \equiv \frac{A}{2\sqrt{mB}}.$$

Scalar field model^[3,4]

Starting point

The Lagrangian density:

$$\mathcal{L}(x) = \partial_\mu \phi^+(x) \partial^\mu \phi(x) - m^2 \phi^+(x) \phi(x) + \mathcal{L}_{\text{int}}(x)$$



Self-consistent set of equations

$$\sigma = \int \frac{d^3k}{(2\pi)^3} \frac{1}{2\omega_k(\sigma)} \{ f_{\text{BE}}[\omega_k(\sigma), \mu_I] + f_{\text{BE}}[\omega_k(\sigma), -\mu_I] \},$$

$$n_I = \int \frac{d^3k}{(2\pi)^3} \{ f_{\text{BE}}[\omega_k(\sigma), \mu_I] - f_{\text{BE}}[\omega_k(\sigma), -\mu_I] \}.$$

Fixed

$$\omega_k(\sigma) = \sqrt{M(\sigma)^2 + k^2}, \quad M(\hat{\sigma}) = m^2 - \frac{\partial \mathcal{L}_{\text{int}}(\hat{\sigma})}{\partial \hat{\sigma}}$$

$$f_{\text{BE}}(\omega, \mu) = \left\{ \exp \left[\frac{\omega - \mu}{T} \right] - 1 \right\}^{-1}, \quad \hat{\sigma}(x) = \phi^+(x) \phi(x)$$



Parameterization of interaction

$$m = 140 \text{ MeV}$$

$$\mathcal{L}_{\text{int}}(\sigma) = \frac{a}{2} \sigma^2 - \frac{b}{3} \sigma^3, \quad \omega - \mu_I = 0$$

$$\kappa = \frac{a}{2m\sqrt{b}}$$

Interplay between a and b
Parameter b - fixed

Thermodynamic mean-field model^[1-3]

Free energy density:

$$\phi(n_1, n_2, T) = \phi_1^{(0)}(n_1, T) + \phi_2^{(0)}(n_2, T) + \phi_{\text{int}}(n, T),$$



$$n = \int \frac{d^3k}{(2\pi)^3} [f_{\text{BE}}(E(k, n), \mu_I) + f_{\text{BE}}(E(k, n), -\mu_I)],$$

$$n_I = \int \frac{d^3k}{(2\pi)^3} [f_{\text{BE}}(E(k, n), \mu_I) - f_{\text{BE}}(E(k, n), -\mu_I)].$$

Fixed

$$E(k, n) = \sqrt{m^2 + k^2} + U(n)$$

$$f_{\text{BE}}(E, \mu) = \left\{ \exp \left[\frac{E - \mu}{T} \right] - 1 \right\}^{-1}$$



$$U(n) = -An + Bn^2, \quad E - \mu = 0$$

$$\kappa \equiv \frac{A}{2\sqrt{mB}}$$

Interplay between A and B
Parameter B - fixed

Scalar field model^[3,4]

Thermodynamic mean-field model^[1-3]

Starting point

The Lagrangian density:

$$\mathcal{L}(x) = \partial_\mu \phi^+(x) \partial^\mu \phi(x) - m^2 \phi^+(x) \phi(x) + \mathcal{L}_{\text{int}}(x)$$



Self-consistent set of equations

$$\sigma = \int \frac{d^3k}{(2\pi)^3} \frac{1}{2\omega_k(\sigma)} \{ f_{\text{BE}}[\omega_k(\sigma), \mu_I] + f_{\text{BE}}[\omega_k(\sigma), -\mu_I] \},$$

$$n_I = \int \frac{d^3k}{(2\pi)^3} \{ f_{\text{BE}}[\omega_k(\sigma), \mu_I] - f_{\text{BE}}[\omega_k(\sigma), -\mu_I] \}.$$

Fixed

$$\omega_k(\sigma) = \sqrt{M(\sigma)^2 + k^2}, \quad M^2(\hat{\sigma}) = m^2 - \frac{\partial \mathcal{L}_{\text{int}}(\hat{\sigma})}{\partial \hat{\sigma}}$$

$$f_{\text{BE}}(\omega, \mu_I) = \left\{ \exp \left[\frac{\omega - \mu_I}{T} \right] - 1 \right\}^{-1}, \quad \hat{\sigma}(x) = \phi^+(x) \phi(x)$$



Parameterization of interaction

$$m = 140 \text{ MeV}$$

$$\mathcal{L}_{\text{int}}(\sigma) = \frac{a}{2} \sigma^2 - \frac{b}{3} \sigma^3, \quad \omega - \mu_I = 0$$

$$\kappa = \frac{a}{2m\sqrt{b}}$$

Interplay between a and b
Parameter b - fixed

Free energy density:

$$\phi(n_1, n_2, T) = \phi_1^{(0)}(n_1, T) + \phi_2^{(0)}(n_2, T) + \phi_{\text{int}}(n, T),$$

$\kappa < 1$ - weak attraction regime

$\kappa = 1$ - critical attraction regime

$\kappa > 1$ - strong(over-critical) attraction regime

$$U(n) = -An + Bn^2, \quad E - \mu = 0$$

$$\kappa \equiv \frac{A}{2\sqrt{mB}}$$

Interplay between A and B
Parameter B - fixed

Ideal 2-component relativistic gas

$$U(n) = 0$$

Mean field
model

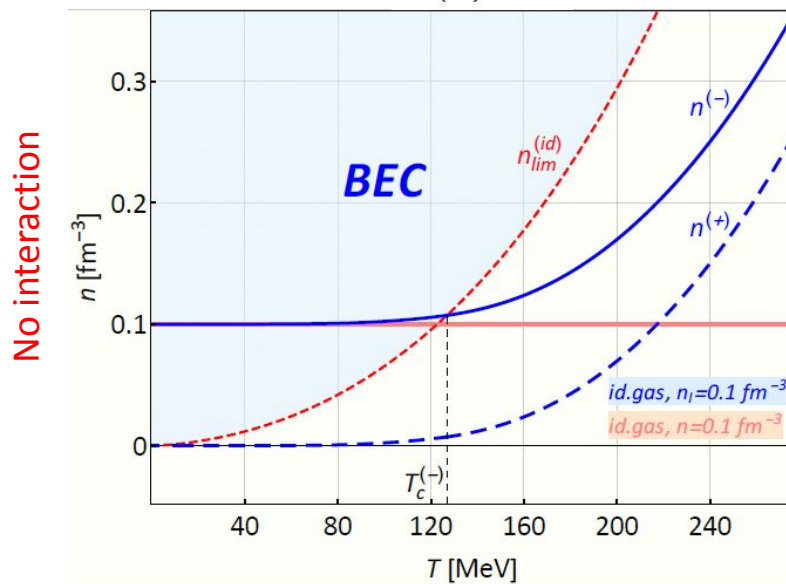


Figure 1. The particle-number densities $n^{(+)}$, $n^{(-)}$ versus temperature for the ideal $\pi^+ - \pi^-$ pion gas. The total isospin density is kept constant, $n_1 = 0.1 \text{ fm}^{-3}$.

Ideal 2-component relativistic gas

$$U(n) = 0$$

Mean field
model

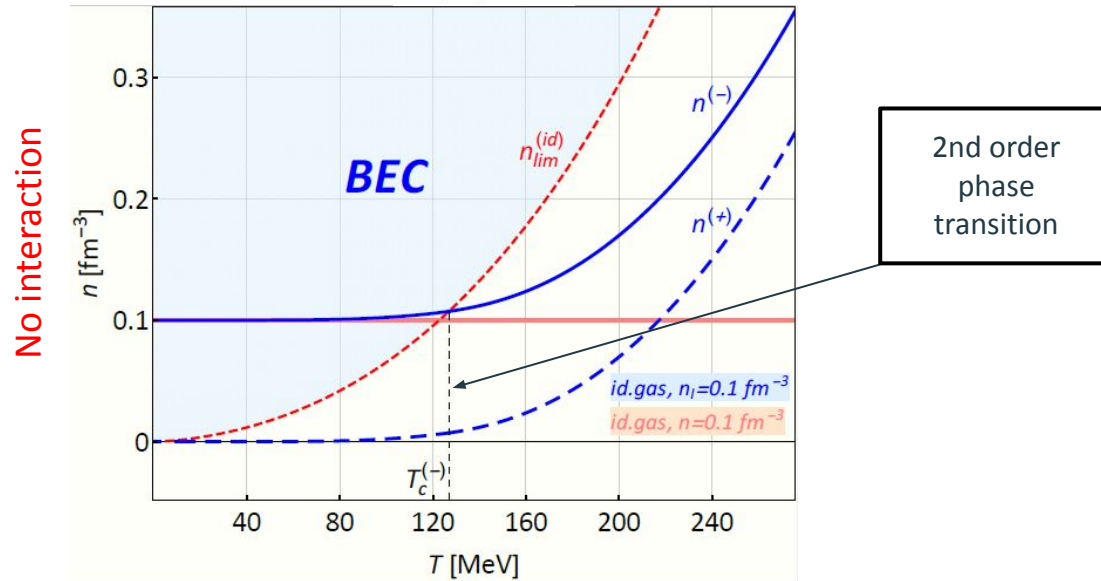


Figure 1. The particle-number densities $n^{(+)}$, $n^{(-)}$ versus temperature for the ideal $\pi^+\pi^-$ pion gas. The total isospin density is kept constant, $n_1 = 0.1 \text{ fm}^{-3}$.

Ideal 2-component relativistic gas

Scalar model

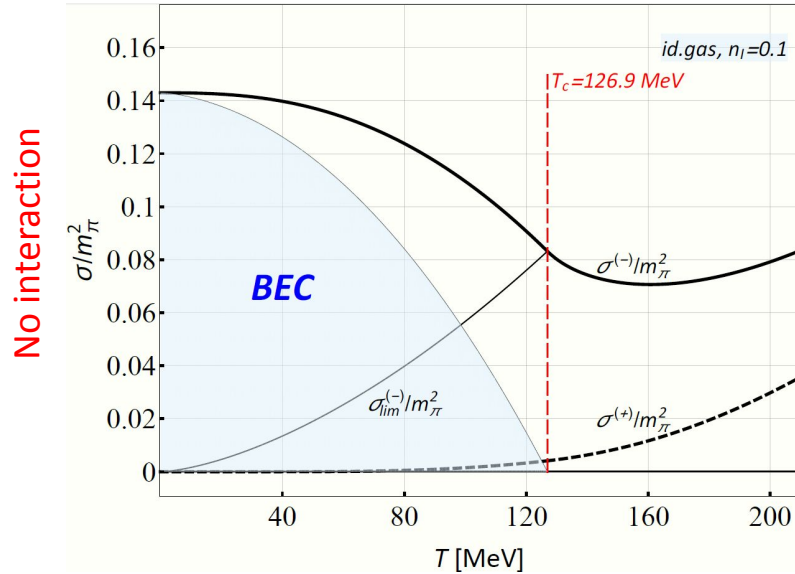


Figure 2. Dependence of scalar density $\sigma^{(+)}/m^2$ and $\sigma^{(-)}/m^2$ on temperature for ideal gas of $\pi^+\pi^-$ pions. Isospin density is considered to be fixed, $n_l = 0.1 \text{ fm}^{-3}$.

Ideal 2-component relativistic gas

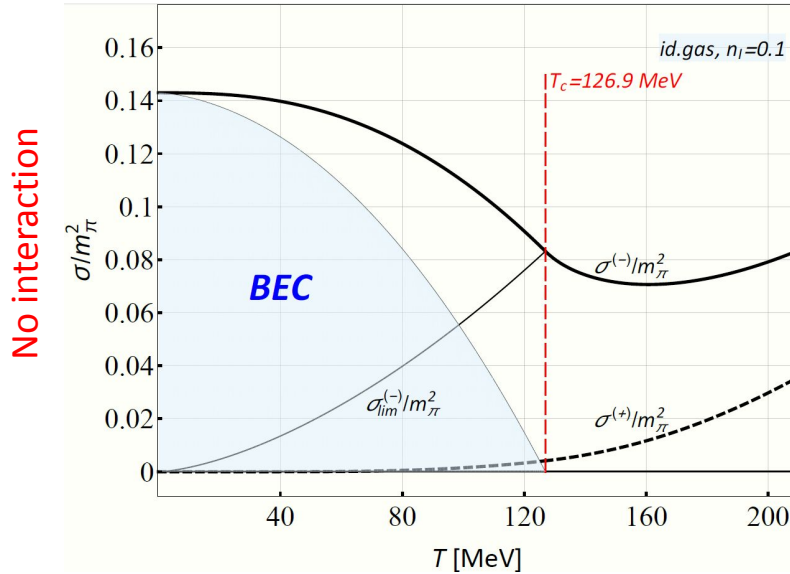


Figure 2. Dependence of scalar density $\sigma^{(+)} / m_\pi^2$ and $\sigma^{(-)} / m_\pi^2$ on temperature for ideal gas of $\pi^+ - \pi^-$ pions. Isospin density is considered to be fixed, $n_l = 0.1 \text{ fm}^{-3}$.

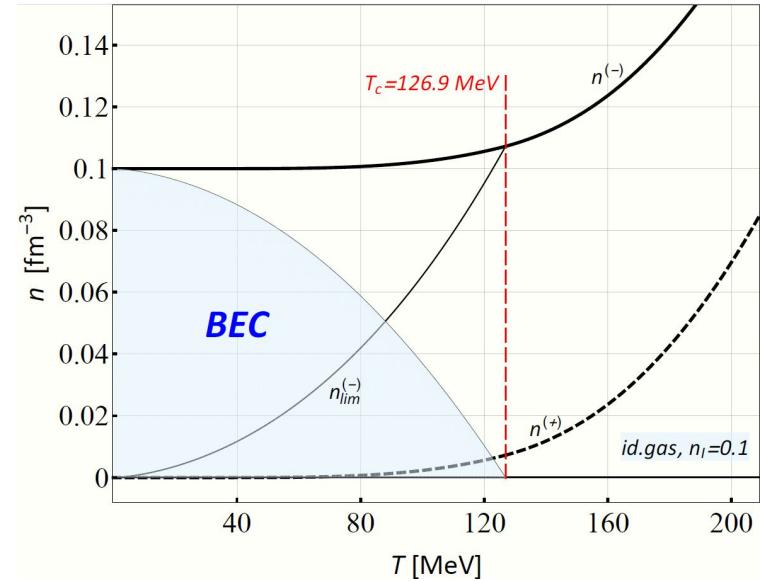


Figure 3. Dependence of particle density $n^{(+)}$, $n^{(-)}$ on temperature for ideal gas of $\pi^+ - \pi^-$ pions within the framework of the scalar field approach. Isospin density is considered to be fixed, $n_l = 0.1 \text{ fm}^{-3}$.

Ideal 2-component relativistic gas

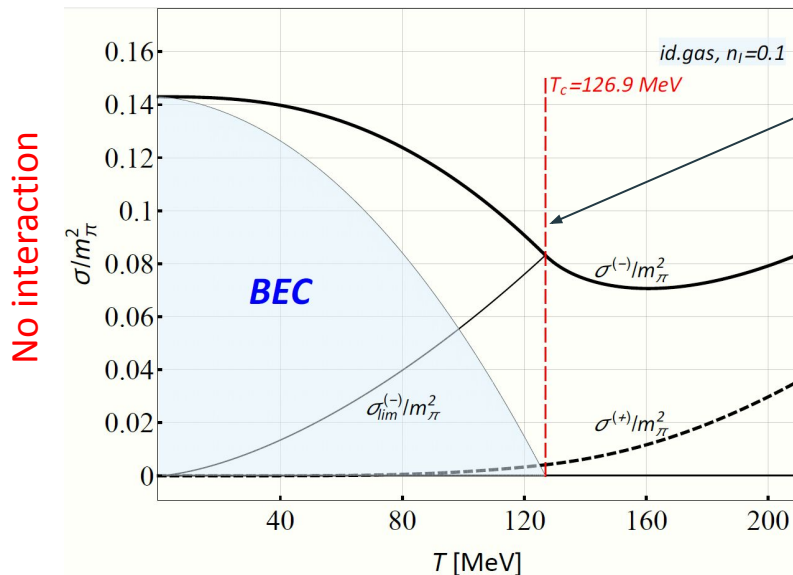


Figure 2. Dependence of scalar density $\sigma^{(+)}/m^2$ and $\sigma^{(-)}/m^2$ on temperature for ideal gas of $\pi^+\pi^-$ pions. Isospin density is considered to be fixed, $n_l = 0.1 \text{ fm}^{-3}$.

2nd order
phase
transition

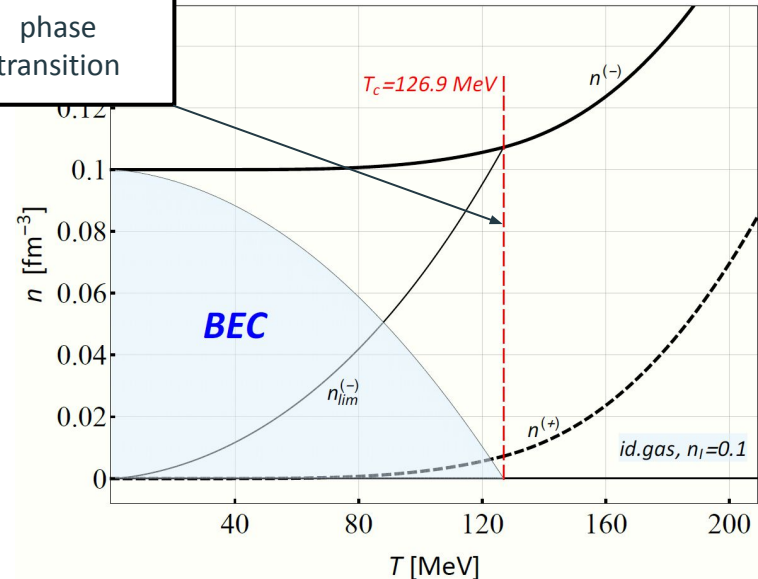


Figure 3. Dependence of particle density $n^{(+)}$, $n^{(-)}$ on temperature for ideal gas of $\pi^+\pi^-$ pions within the framework of the scalar field approach. Isospin density is considered to be fixed, $n_l = 0.1 \text{ fm}^{-3}$.

Thermodynamic approach

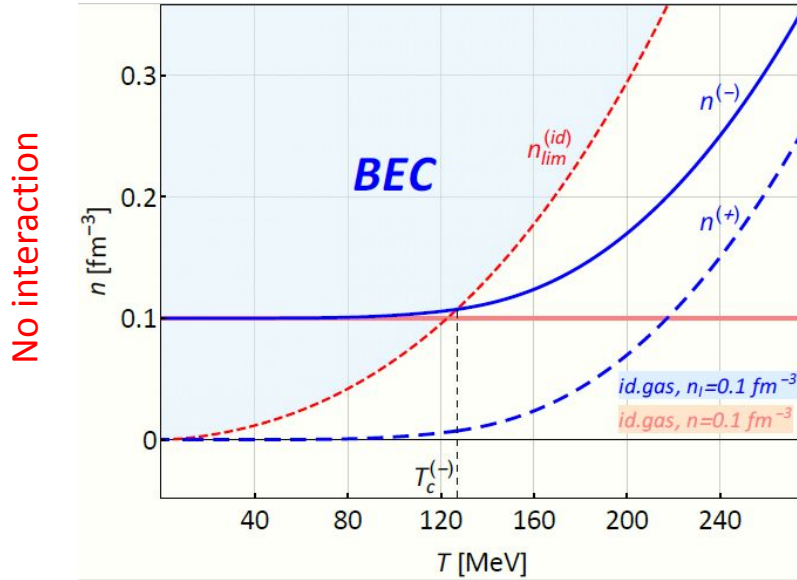


Figure 1. Dependence of particle density $n^{(+)}$, $n^{(-)}$ on temperature for ideal gas of $\pi^+\pi^-$ pions within the framework of the thermodynamic mean-field approach. Isospin density is considered to be fixed, $n_l = 0.1 \text{ fm}^{-3}$.

Field approach

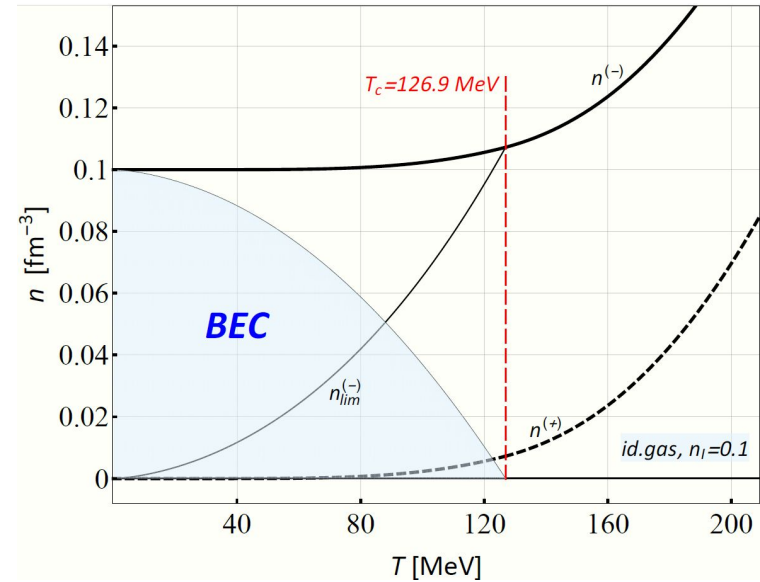


Figure 3. Dependence of particle density $n^{(+)}$, $n^{(-)}$ on temperature for ideal gas of $\pi^+\pi^-$ pions within the framework of the scalar field approach. Isospin density is considered to be fixed, $n_l = 0.1 \text{ fm}^{-3}$.

Thermodynamic approach

No interaction

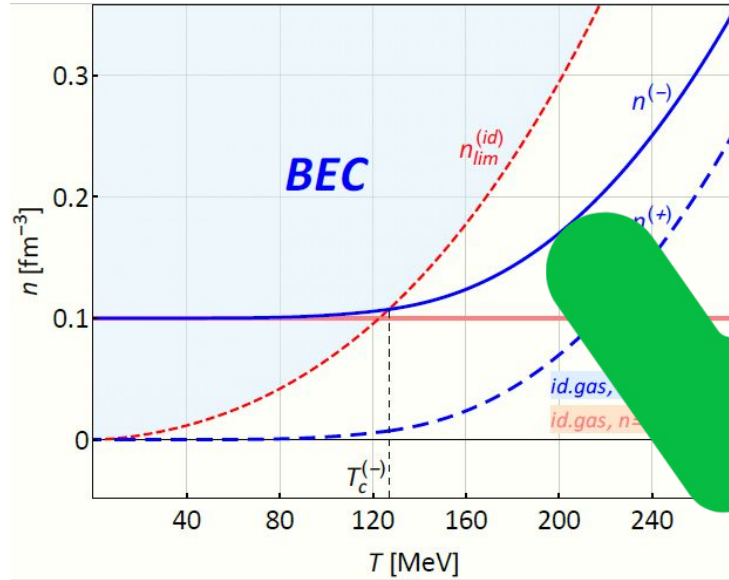


Figure 1. Dependence of particle density $n^{(+)}$, $n^{(-)}$ on temperature for ideal gas of $\pi^+\pi^-$ pions within the framework of the thermodynamic mean-field approach. Isospin density is considered to be fixed, $n_1 = 0.1 \text{ fm}^{-3}$.

Field approach

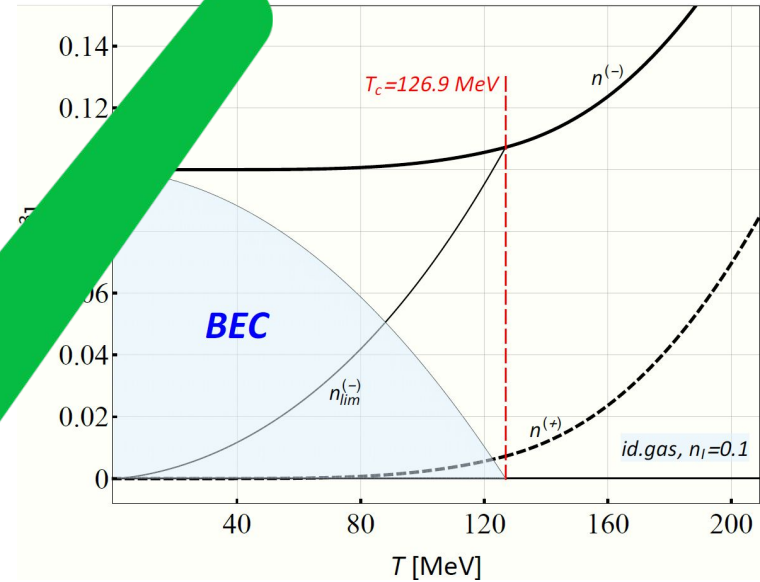
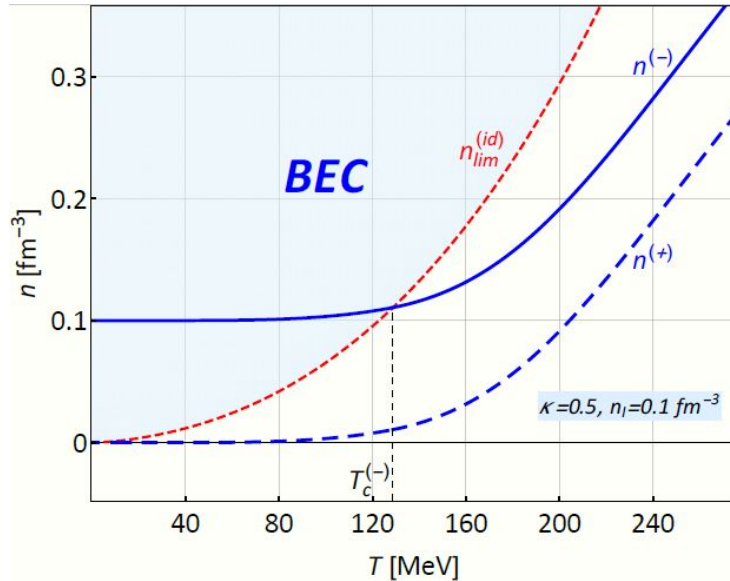


Figure 3. Dependence of particle density $n^{(+)}$, $n^{(-)}$ on temperature for ideal gas of $\pi^+\pi^-$ pions within the framework of the scalar field approach. Isospin density is considered to be fixed, $n_1 = 0.1 \text{ fm}^{-3}$.

Let's turn the attraction ON!

Type 1 phase transition (2nd order ph. tr.)

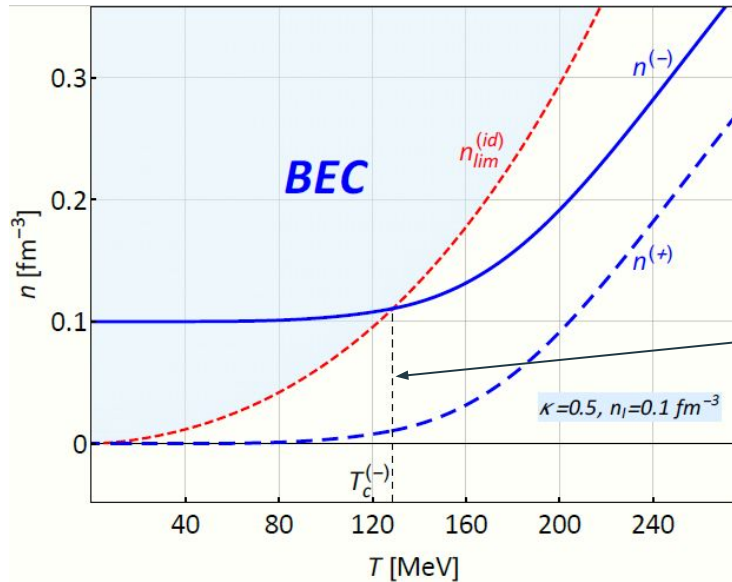


Weak attraction regime

$$\kappa = 0.5$$

Figure 4. The particle-number densities $n^{(+)}$, $n^{(-)}$ versus temperature for the interacting $\pi^+-\pi^-$ pion gas in the mean-field model. The total isospin density is kept constant, $n_1 = 0.1 \text{ fm}^{-3}$, and the attraction parameter is $\kappa = 0.5$.

Type 1 phase transition (2nd order ph. tr.)



Phase
transition of
the 2nd order

Weak attraction regime
 $\kappa = 0.5$

Figure 4. The particle-number densities $n^{(+)}$, $n^{(-)}$ versus temperature for the interacting $\pi^+\text{-}\pi^-$ pion gas in the mean-field model. The total isospin density is kept constant, $n_1 = 0.1 \text{ fm}^{-3}$, and the attraction parameter is $\kappa = 0.5$.

Type 1 phase transition (2nd order ph. tr.)

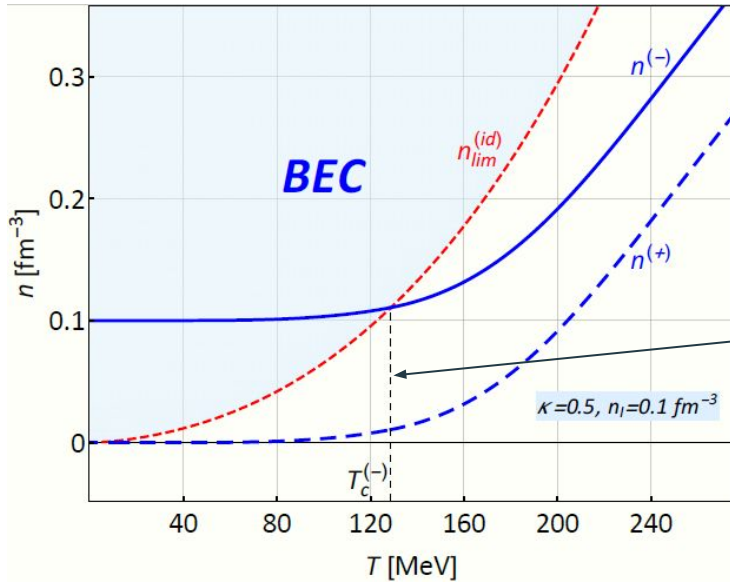
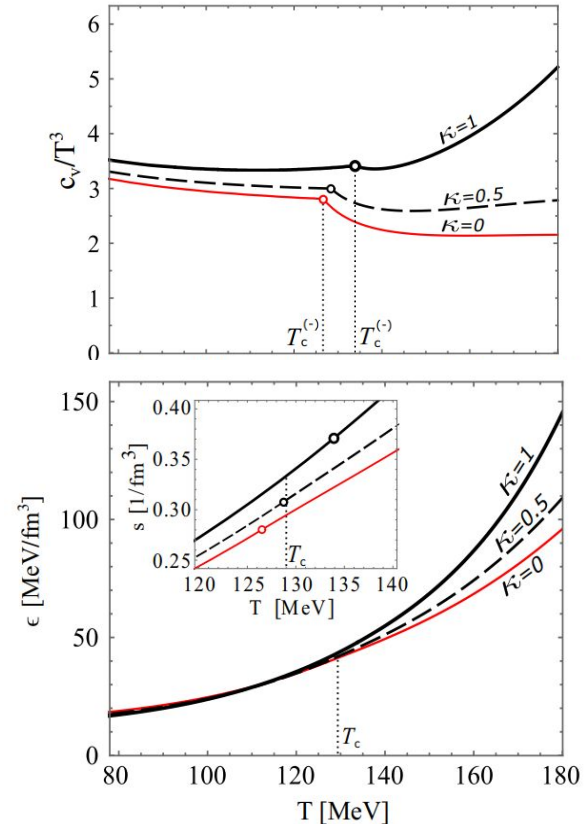
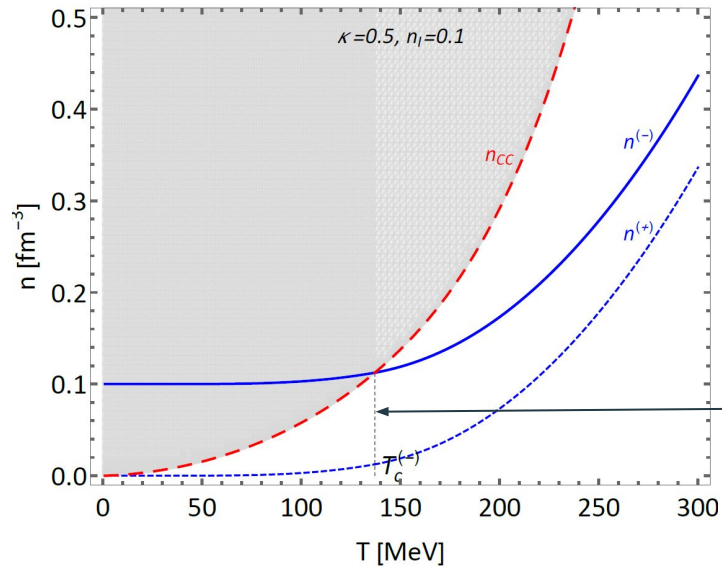


Figure 4. The particle-number densities $n^{(+)}$, $n^{(-)}$ versus temperature for the interacting $\pi^+\pi^-\pi^-$ pion gas in the mean-field model. The total isospin density is kept constant, $n_1 = 0.1 \text{ fm}^{-3}$, and the attraction parameter is $\kappa = 0.5$.



Type 1 phase transition (2nd order ph. tr.)



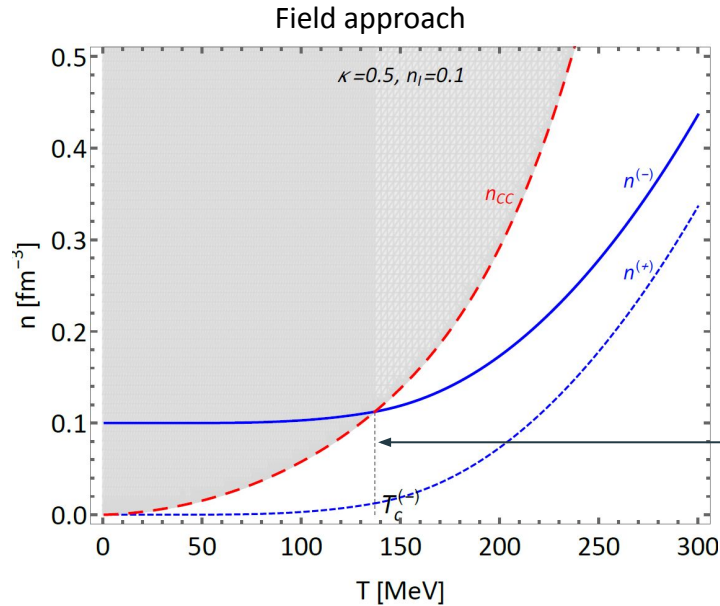
Weak attraction regime

$$\kappa = 0.5$$

2nd order
phase
transition

Figure 5. Dependence of particle density $n^{(+)}$, $n^{(-)}$ on temperature for $\pi^+\pi^-$ pion gas with interaction within the framework of the scalar field approach. Isospin density is considered to be fixed, $n_1 = 0.1 \text{ fm}^{-3}$, the attraction parameter $\kappa = 0.5$.

Type 1 phase transition (2nd order ph. tr.)



2nd order
phase
transition

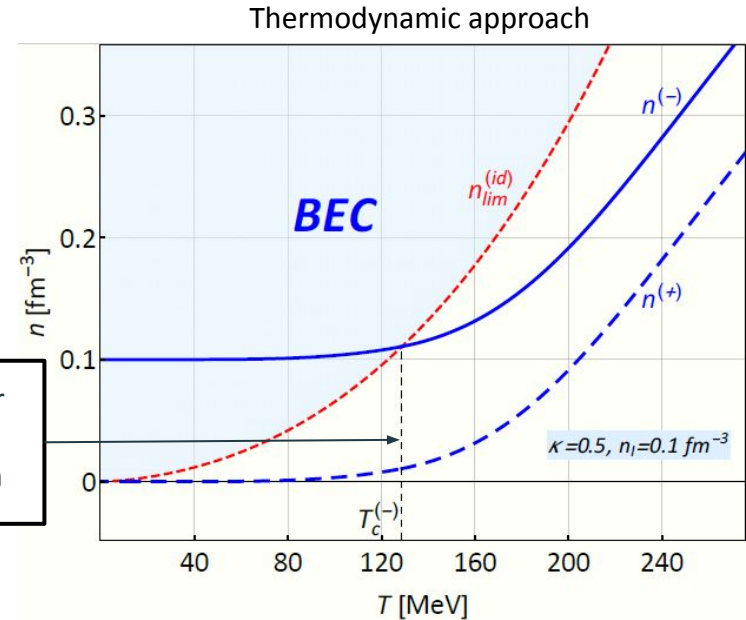


Figure 5. Dependence of particle density $n^{(+)}$, $n^{(-)}$ on temperature for $\pi^+\pi^-\pi^-\pi^+$ pion gas with interaction within the framework of the scalar field approach. Isospin density is considered to be fixed, $n_l = 0.1 \text{ fm}^{-3}$, the attraction parameter $\kappa = 0.5$.

Figure 4. The particle-number densities $n^{(+)}$, $n^{(-)}$ versus temperature for the interacting $\pi^+\pi^-\pi^-\pi^+$ pion gas in the mean-field model. The total isospin density is kept constant, $n_l = 0.1 \text{ fm}^{-3}$, and the attraction parameter is $\kappa = 0.5$.

Type 1 phase transition (2nd order ph. tr.)

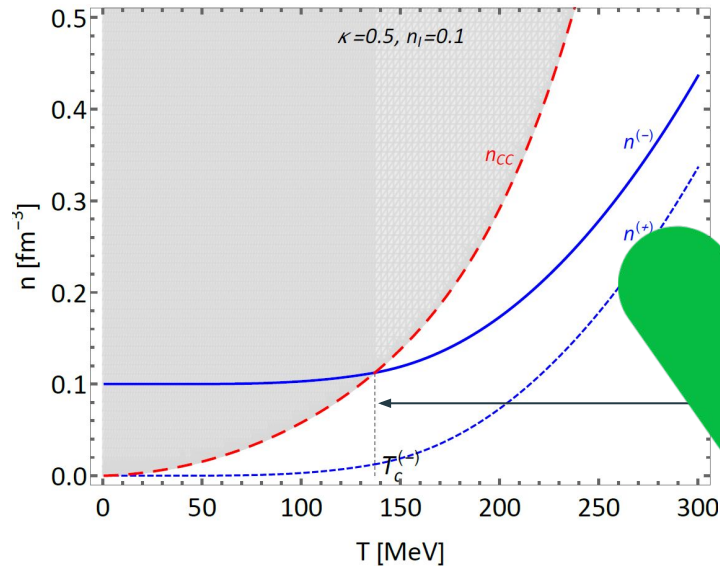


Figure 5. Dependence of particle density $n^{(+)}$, $n^{(-)}$ on temperature for $\pi^+\pi^-$ pion gas with interaction within the framework of the scalar field approach. Isospin density is considered to be fixed, $n_l = 0.1 \text{ fm}^{-3}$, the attraction parameter $\kappa = 0.5$.

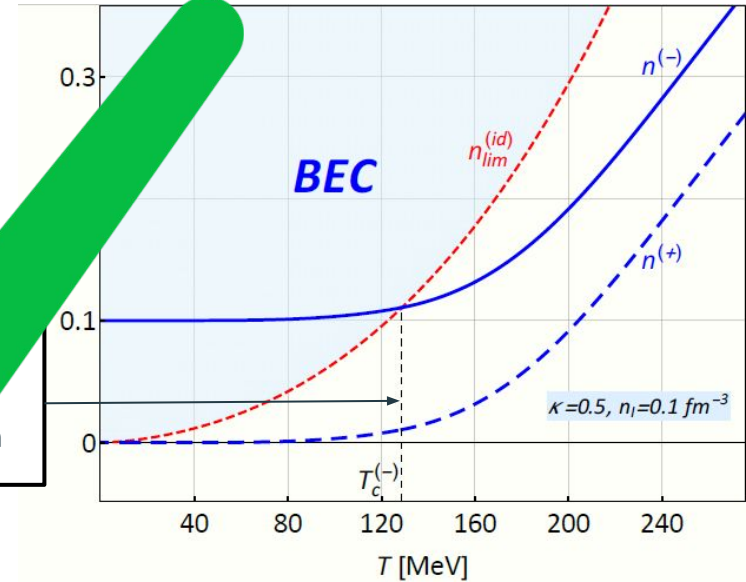
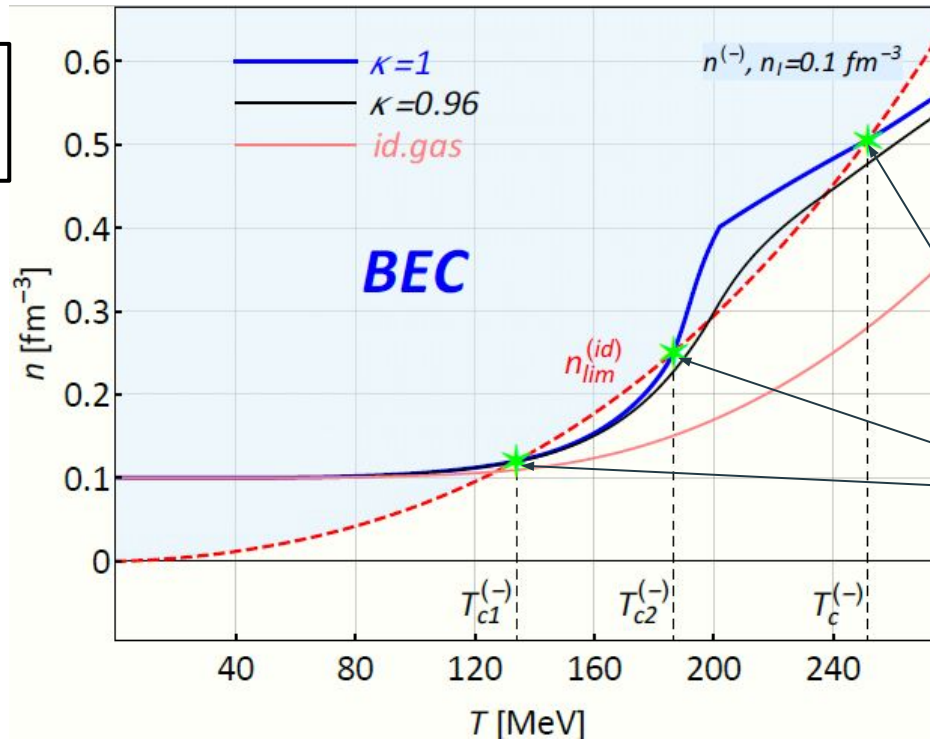


Figure 4. The particle-number densities $n^{(+)}$, $n^{(-)}$ versus temperature for the interacting $\pi^+\pi^-$ pion gas in the mean-field model. The total isospin density is kept constant, $n_l = 0.1 \text{ fm}^{-3}$, and the attraction parameter is $\kappa = 0.5$.

What will happen if we increase attraction parameter κ in such system?

Type 2 phase transition (2nd order ph. tr.)

Only particles π^- are depicted



2nd order phase transitions

Figure 6. The particle density $n^{(-)}$ versus temperature for the interacting $\pi^+\pi^-\pi^-$ pion gas, the attraction parameter is $\kappa = 1$.

Type 3 phase transition (2nd order ph. tr.)

Only antiparticles π^+
are depicted

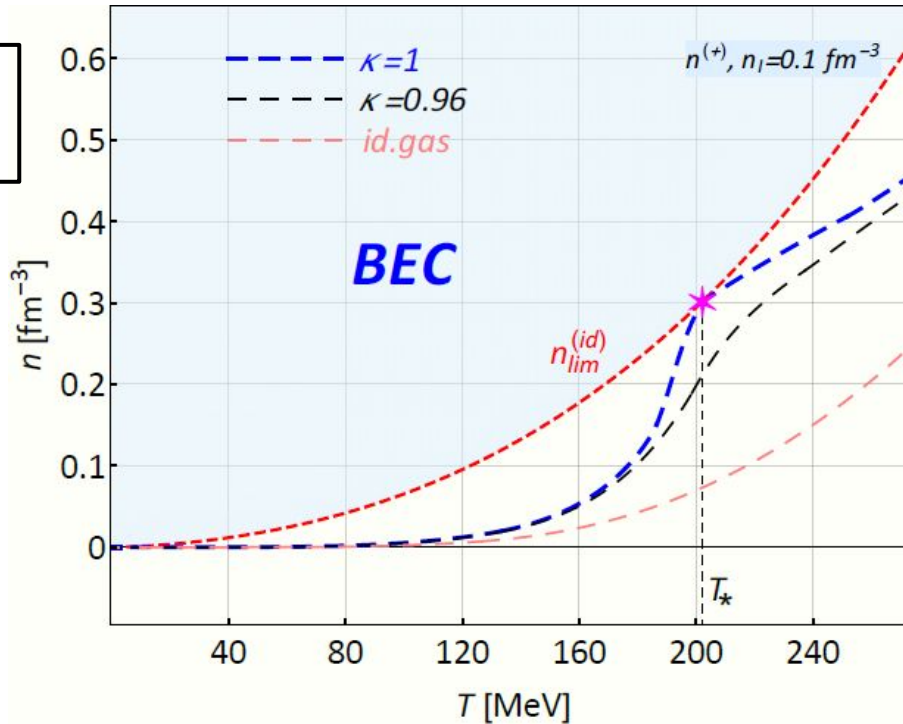
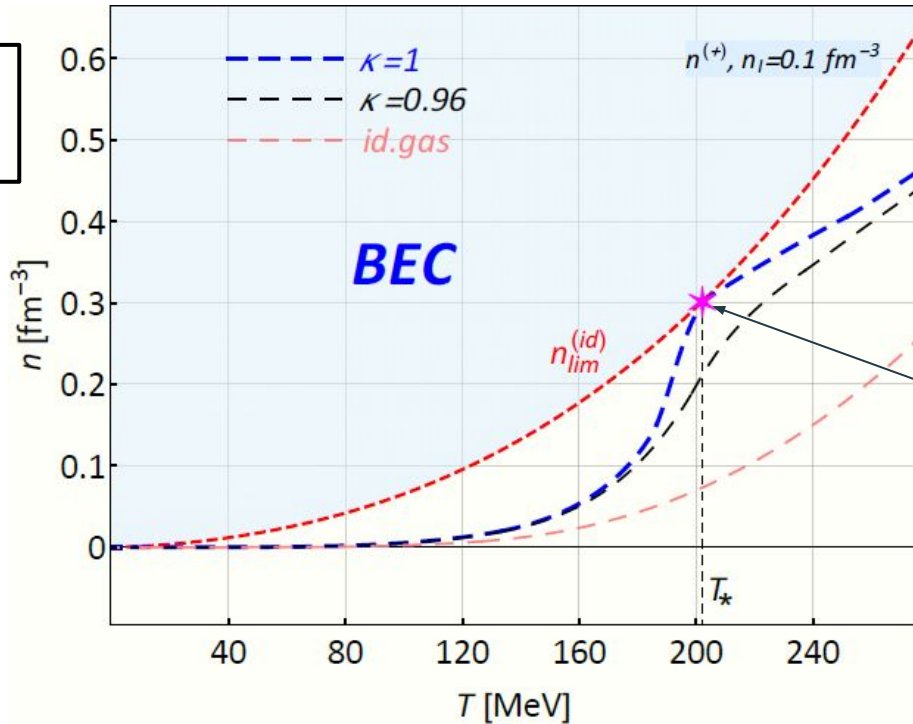


Figure 7. The particle density $n^{(+)}$ versus temperature for the interacting $\pi^+\pi^-$ pion gas, the attraction parameter is $\kappa = 1$.

Type 3 phase transition (2nd order ph. tr.)

Only antiparticles π^+
are depicted



2nd order
point-like phase
transition

Figure 7. The particle density $n^{(+)}$ versus temperature for the interacting $\pi^+\pi^-$ pion gas, the attraction parameter is $\kappa = 1$.

Type 3 phase transition (2nd order ph. tr.)

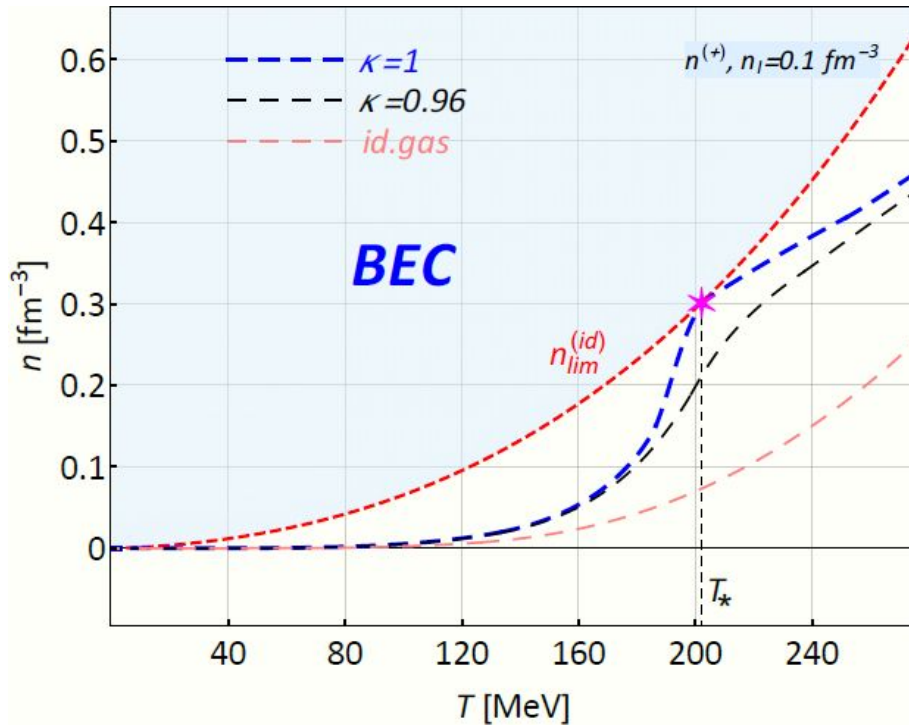
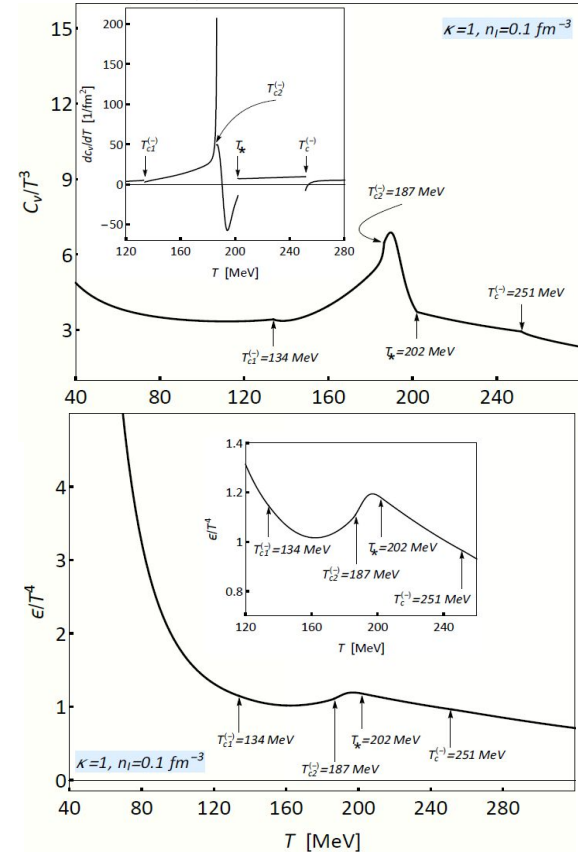


Figure 7. The particle density $n^{(+)}$ versus temperature for the interacting $\pi^+\pi^-$ pion gas, the attraction parameter is $\kappa = 1$.



Type 2,3 phase transition (2nd order ph. tr.)^[3]

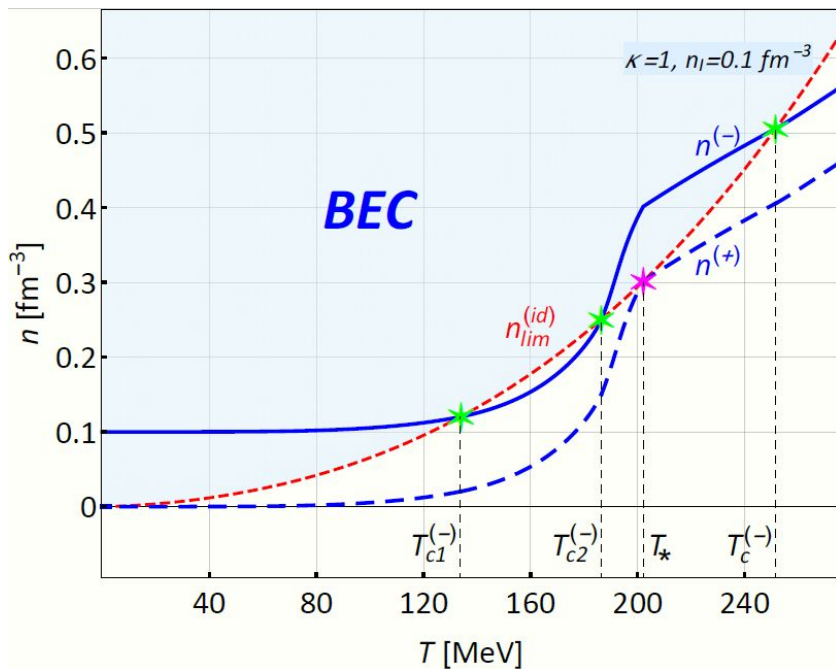


Figure 8. The particle-number densities $n^{(+)}$, $n^{(-)}$ versus temperature for the interacting $\pi^+\pi^-$ pion gas. The attraction parameter is $\kappa = 1$.

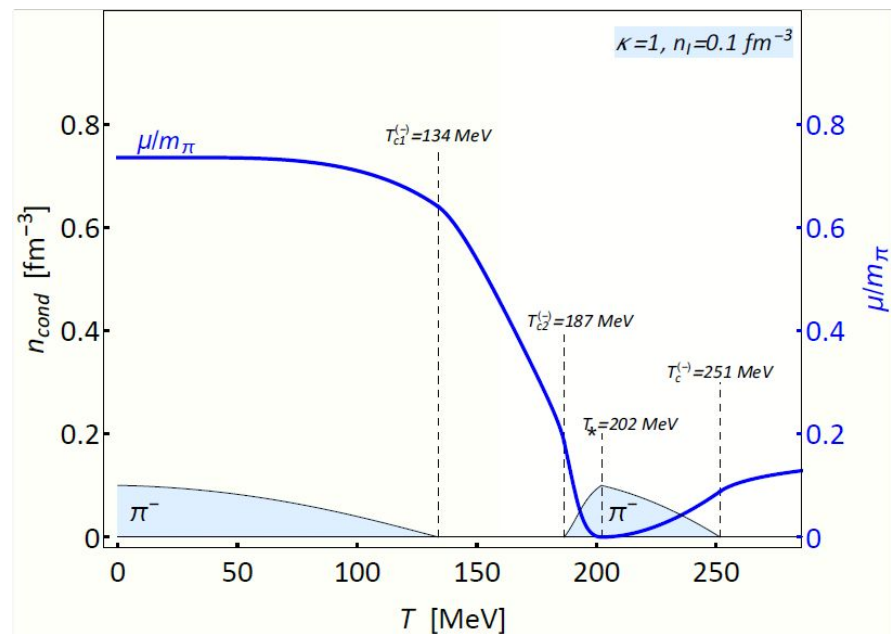


Figure 9. The condensate particle-number densities n_{cond} versus temperature. The attraction parameter is $\kappa = 1$.

Point-like phase transition in scalar model

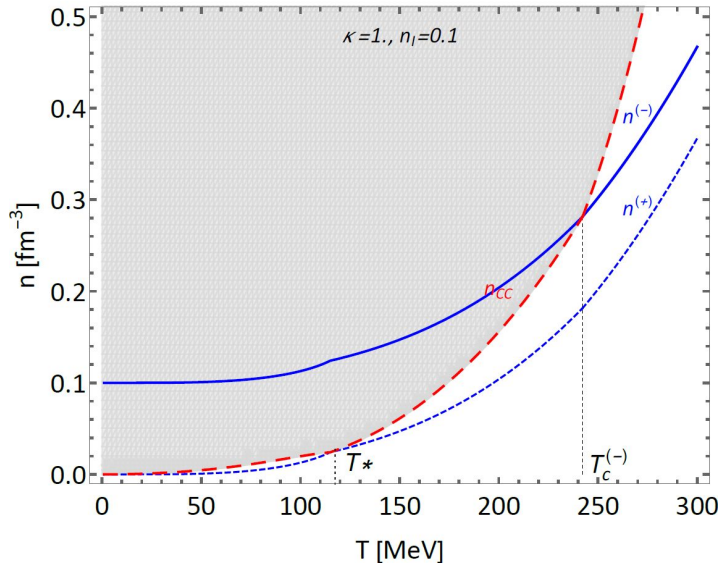
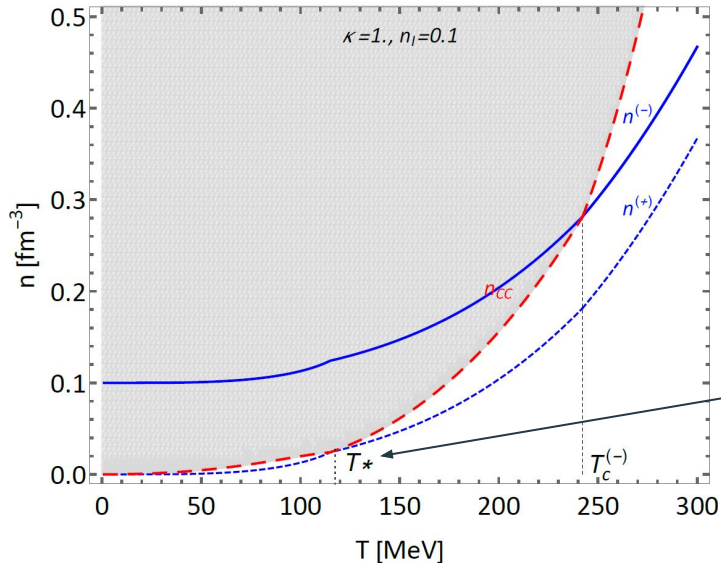


Figure 10. Dependence of particle density $n^{(+)}$, $n^{(-)}$ on temperature for $\pi^+\pi^-$ pion gas with interaction within the framework of the scalar field approach. Isospin density is considered to be fixed, $n_1 = 0.1 \text{ fm}^{-3}$, the attraction parameter $\kappa = 1.$

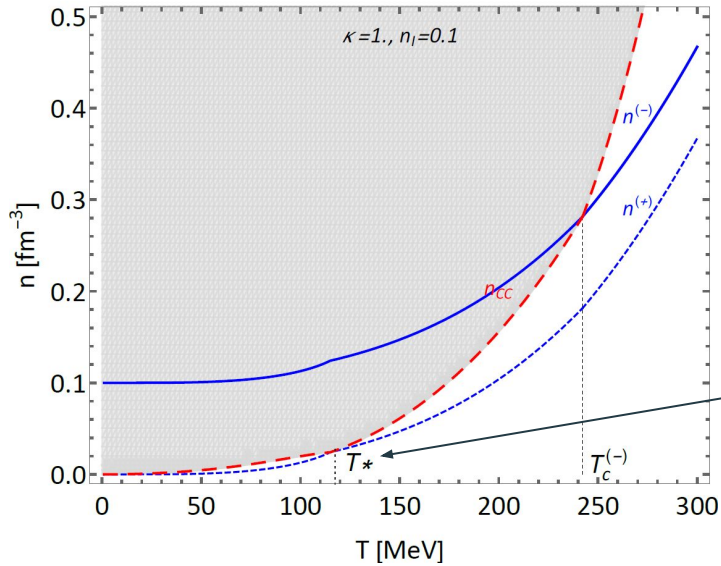
Point-like phase transition in scalar model



Point-like phase transition of the 2nd order?

Figure 10. Dependence of particle density $n^{(+)}$, $n^{(-)}$ on temperature for $\pi^+\pi^-$ pion gas with interaction within the framework of the scalar field approach. Isospin density is considered to be fixed, $n_l = 0.1 \text{ fm}^{-3}$, the attraction parameter $\kappa = 1$.

Point-like phase transition in scalar model

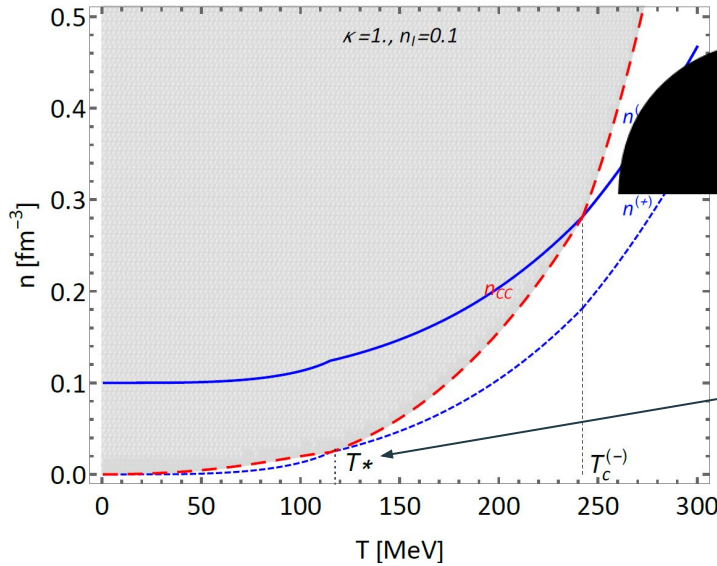


Technical difficulties with describing the condensate phase within the scalar field approach

Point-like phase transition of the 2nd order?

Figure 10. Dependence of particle density $n^{(+)}$, $n^{(-)}$ on temperature for $\pi^+\pi^-$ pion gas with interaction within the framework of the scalar field approach. Isospin density is considered to be fixed, $n_1 = 0.1 \text{ fm}^{-3}$, the attraction parameter $\kappa = 1$.

Point-like phase transition in scalar model



Technical difficulties with describing the condensate phase within the scalar field approach

Point-like phase transition of the 2nd order?

Figure 10. Dependence of particle density $n^{(+)}$, $n^{(-)}$ on temperature for $\pi^+\pi^-$ pion gas with interaction within the framework of the scalar field approach. Isospin density is considered to be fixed, $n_1 = 0.1$ fm⁻³, the attraction parameter $\kappa = 1$.

To be continued...

What if κ is in over-critical regime?

$$\kappa > 1$$

Type 4 phase transition (1st order ph. tr.)^[3,5]

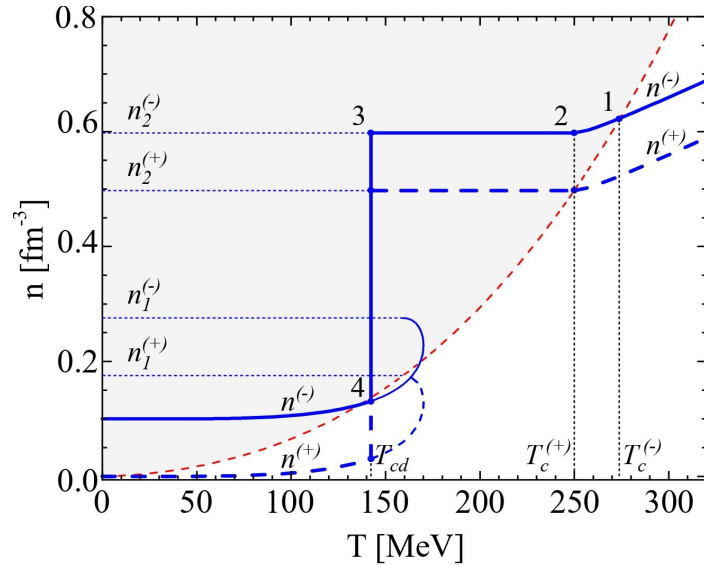


Figure 11. The particle-number densities $n^{(+)}$, $n^{(-)}$ versus temperature for the interacting $\pi^+\pi^-$ pion gas. The attraction parameter is $\kappa = 1.1$.

[3] D. Anchishkin, V. Gnatovskyy, D. Zhuravel, V. Karpenko, I. Mishustin, H. Stoecker, Universe 2023, 9, 411 (2023).

[5] D. Anchishkin, V. Gnatovskyy, D. Zhuravel, V. Karpenko, I. Mishustin, H. Stoecker, Ukr. J. Phys. 2024. Vol. 69, No. 1 (2024).

Type 4 phase transition (1st order ph. tr.)^[3,5]

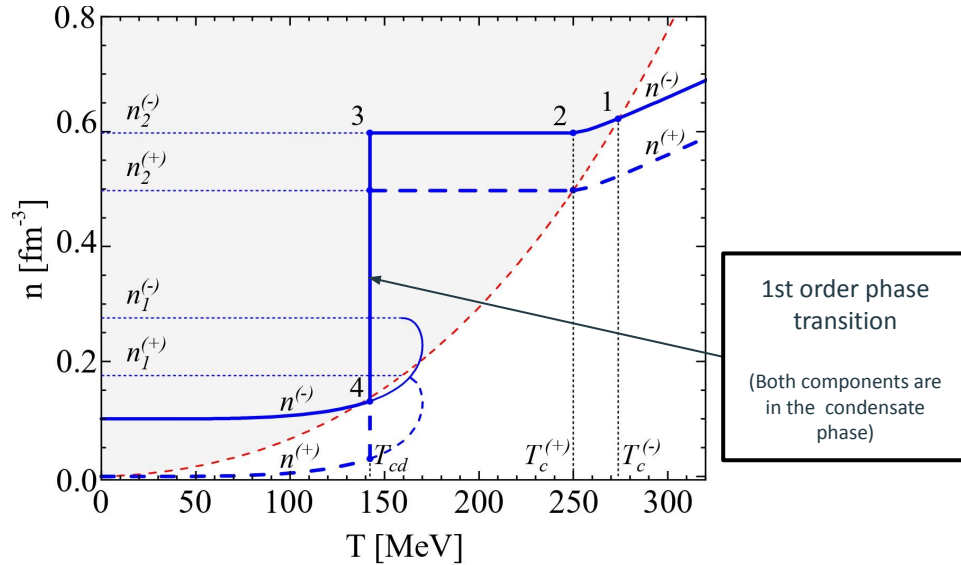


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[5] D. Anchishkin, V. Gnatovskyy, D. Zhuravel, V. Karpenko, I. Mishustin, H. Stoecker, Ukr. J. Phys. 2024. Vol. 69, No. 1 (2024).

Type 4 phase transition (1st order ph. tr.)^[3,5]

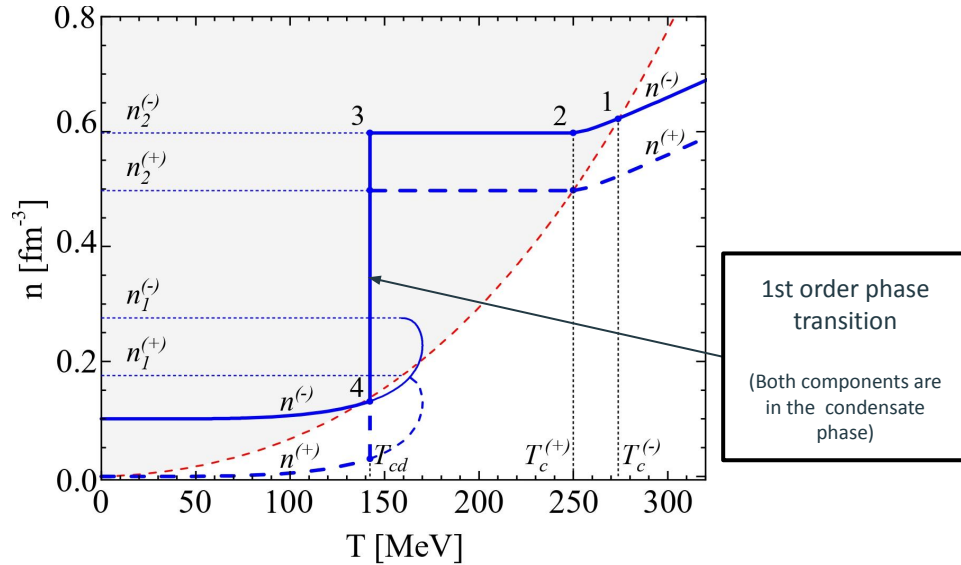
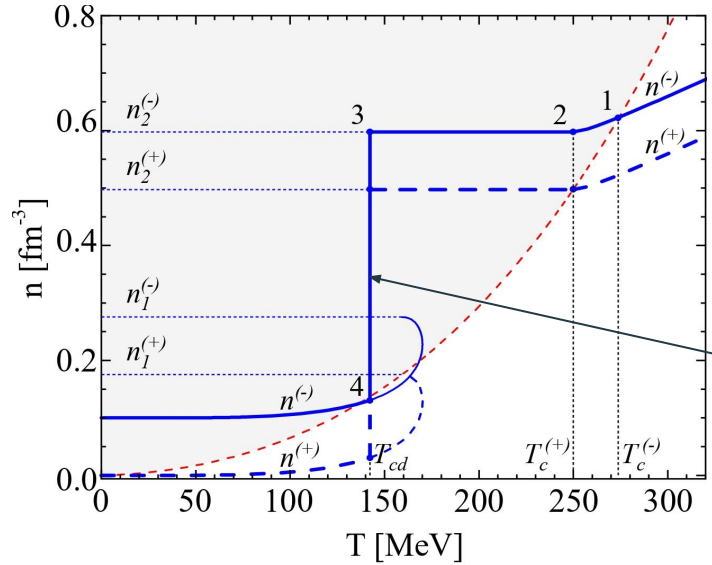


Figure 11. The particle-number densities $n^{(+)}$, $n^{(-)}$ versus temperature for the interacting $\pi^+\pi^-$ pion gas. The attraction parameter is $\kappa = 1.1$.

$$n_1 = \sqrt{\frac{m}{B}} \left(\kappa - \sqrt{\kappa^2 - 1} \right), \quad n_2 = \sqrt{\frac{m}{B}} \left(\kappa + \sqrt{\kappa^2 - 1} \right),$$

[3] D. Anchishkin, V. Gnatovskyy, D. Zhuravel, V. Karpenko, I. Mishustin, H. Stoecker, Universe 2023, 9, 411 (2023).
 [5] D. Anchishkin, V. Gnatovskyy, D. Zhuravel, V. Karpenko, I. Mishustin, H. Stoecker, Ukr. J. Phys. 2024. Vol. 69, No. 1 (2024).

Type 4 phase transition (1st order ph. tr.)^[3,5]



1st order phase transition
(Both components are in the condensate phase)

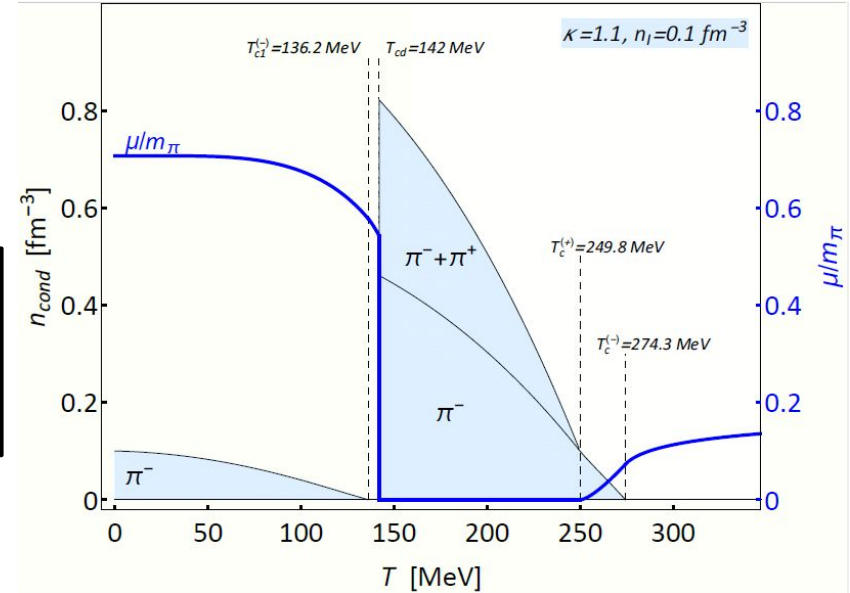


Figure 11. The particle-number densities $n^{(+)}$, $n^{(-)}$ versus temperature for the interacting $\pi^+\pi^-$ pion gas. The attraction parameter is $\kappa = 1.1$.

$$n_1 = \sqrt{\frac{m}{B}} \left(\kappa - \sqrt{\kappa^2 - 1} \right), \quad n_2 = \sqrt{\frac{m}{B}} \left(\kappa + \sqrt{\kappa^2 - 1} \right),$$

[3] D. Anchishkin, V. Gnatovskyy, D. Zhuravel, V. Karpenko, I. Mishustin, H. Stoecker, Universe 2023, 9, 411 (2023).
[5] D. Anchishkin, V. Gnatovskyy, D. Zhuravel, V. Karpenko, I. Mishustin, H. Stoecker, Ukr. J. Phys. 2024. Vol. 69, No. 1 (2024).

Comparison with lattice data

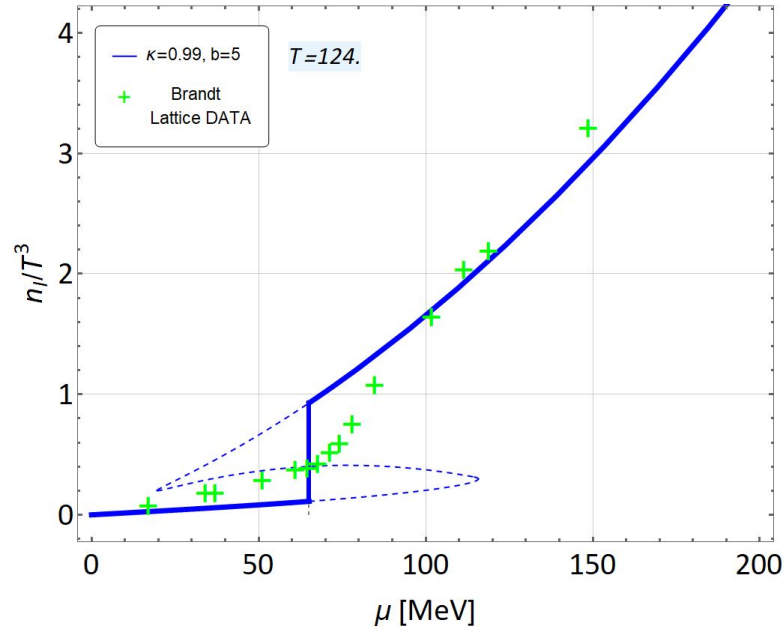


Figure 12. Dependence of normalized isospin density n_I/T^3 on chemical potential for $\pi^+\pi^-$ pions system within field approach showed with blue line. Lattice data for π -mesons configurations at $T=124$ MeV from [6].

Conclusions

- ❑ The mean-field thermodynamic model and the scalar field model were used to describe a 2-component system of interacting mesons (bosons) at high temperatures in the presence of condensate.
- ❑ 4 types of phase transitions in the system with repulsion and attraction were introduced: **type 1** - one component single phase transition (2nd order), **type 2** - one component multiple phase transition (2nd order), **type 3** - second component point-like phase transition (2nd order), **type 4** - both components phase transition (1st order).
- ❑ Parameters of interaction within scalar field approach were fitted to up-to-date lattice data. Attraction parameter k in this case is close to its critical value.

Thank you for
your attention!

Over-critical regime with $n_I = 0$

Термодинамічна модель
середнього поля

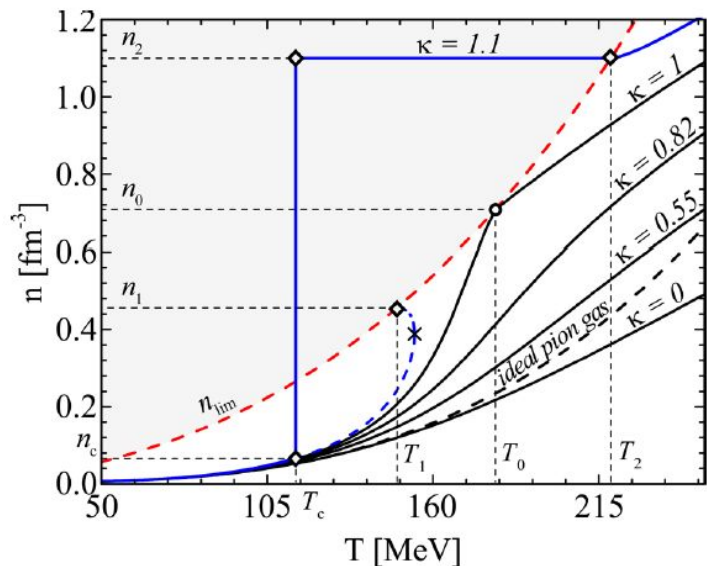


Figure 18. Залежність густин частинок n від температури для взаємодіючого газу $\pi^+\pi^-$ піонів в рамках термодинамічної моделі середнього поля. Густина ізоспіну вважається сталою, $n_I=0$.

Модель скалярного
поля

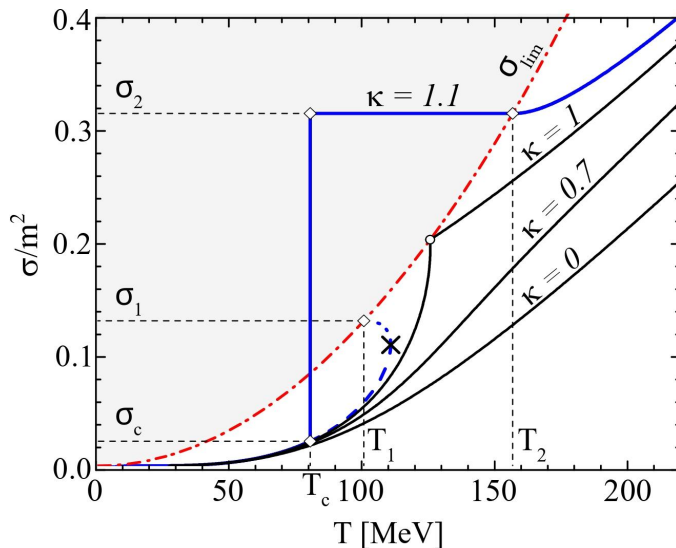
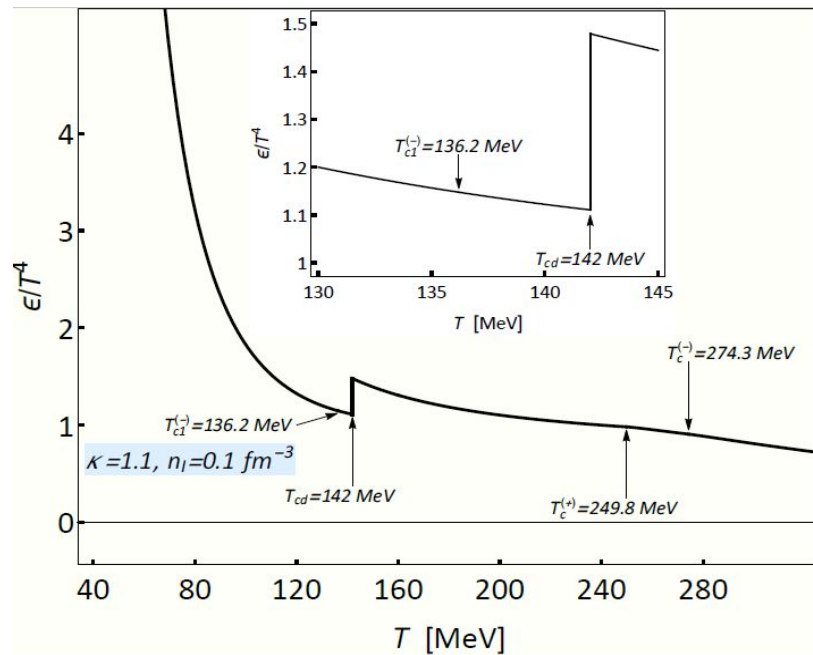
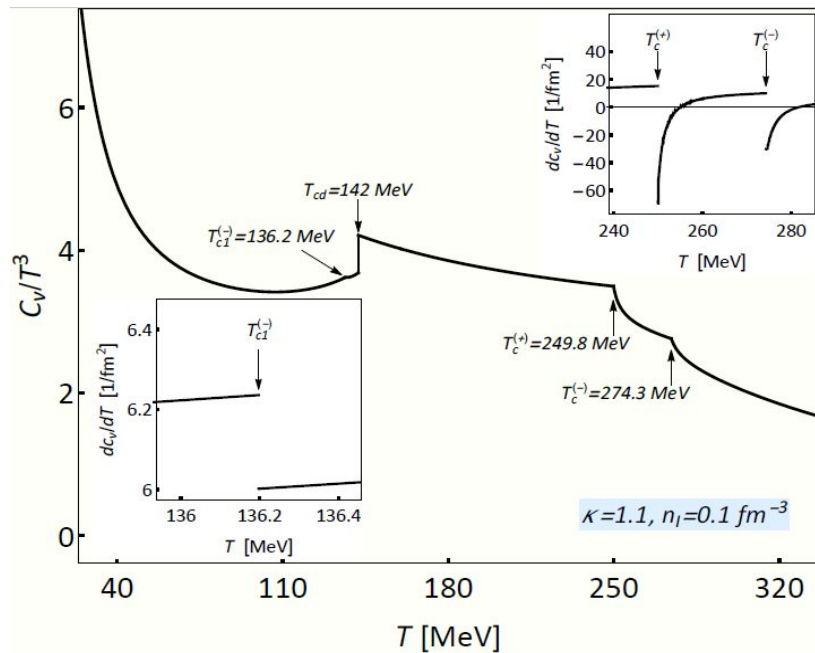


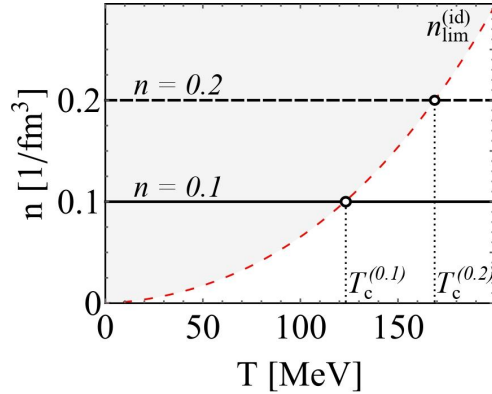
Figure 19. Залежність скалярної густини σ / m^2 від температури для газу $\pi^+\pi^-$ піонів із взаємодією. Густина ізоспіну вважається сталою, $n_I = 0$.

Type 4 phase transition (1st order ph. tr.)^[3,4]

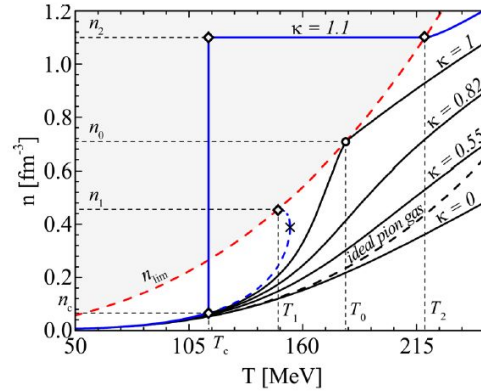
Proof



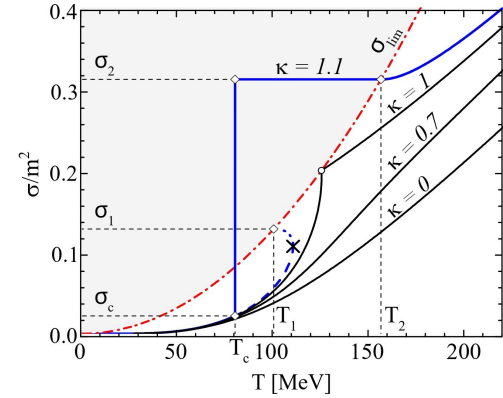
An ideal relativistic gas



Thermodynamic mean-field model



Scalar field model



Thermodynamic mean-field model

$$n = \int \frac{d^3k}{(2\pi)^3} [f_{\text{BE}}(E(k, n), \mu_I) + f_{\text{BE}}(E(k, n), -\mu_I)] ,$$

$$n_I = \int \frac{d^3k}{(2\pi)^3} [f_{\text{BE}}(E(k, n), \mu_I) - f_{\text{BE}}(E(k, n), -\mu_I)] .$$

$$E(k, n) = \sqrt{m^2 + \mathbf{k}^2} + U(n)$$

$$f_{\text{BE}}(E, \mu) = \left\{ \exp \left[\frac{E - \mu}{T} \right] - 1 \right\}^{-1}$$

Scalar field model

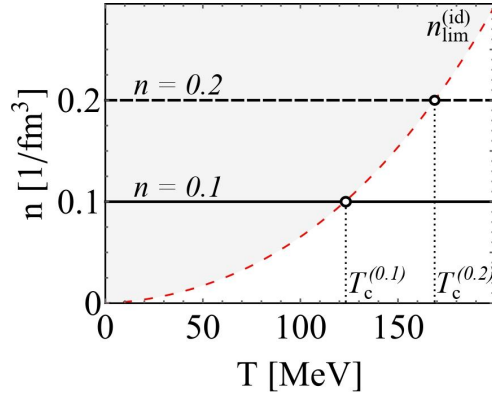
$$\sigma = \int \frac{d^3k}{(2\pi)^3} \frac{1}{2\omega_k(\sigma)} \{ f_{\text{BE}}[\omega_k(\sigma), \mu_I] + f_{\text{BE}}[\omega_k(\sigma), -\mu_I] \} ,$$

$$n_I = \int \frac{d^3k}{(2\pi)^3} \{ f_{\text{BE}}[\omega_k(\sigma), \mu_I] - f_{\text{BE}}[\omega_k(\sigma), -\mu_I] \} .$$

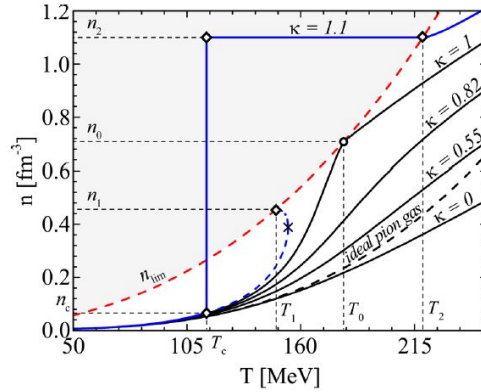
$$\omega_k(\sigma) = \sqrt{M(\sigma)^2 + \mathbf{k}^2}$$

$$f_{\text{BE}}(\omega, \mu_I) = \frac{1}{e^{(\omega - \mu_I)/T} - 1}$$

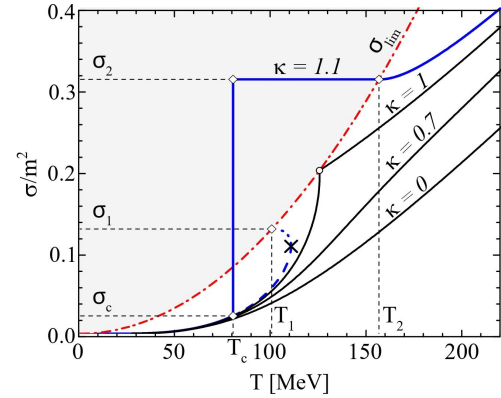
Ideal relativistic gas



Thermodynamic mean-field model



Scalar field model



Thermodynamic mean-field model

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$$E(k, n) = \sqrt{m^2 + \mathbf{k}^2} + U(n)$$

$$f_{\text{BE}}(E, \mu) = \left\{ \exp \left[\frac{E - \mu}{T} \right] - 1 \right\}^{-1}$$

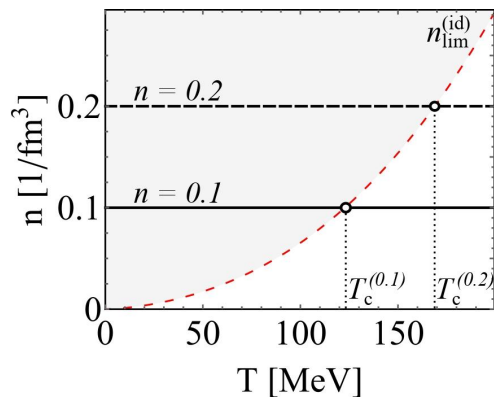
Scalar field model

$$\sigma = \int \frac{d^3k}{(2\pi)^3} \frac{1}{2\omega_k(\sigma)} \{ f_{\text{BE}}[\omega_k(\sigma), \mu_I] + f_{\text{BE}}[\omega_k(\sigma), -\mu_I] \} ,$$

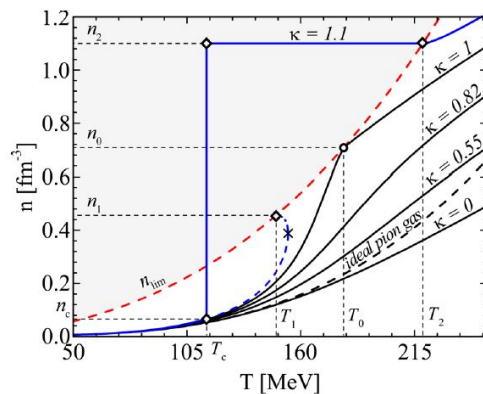
$$n_I = \int \frac{d^3k}{(2\pi)^3} \{ f_{\text{BE}}[\omega_k(\sigma), \mu_I] - f_{\text{BE}}[\omega_k(\sigma), -\mu_I] \} .$$

$$n^{(-)} = \int \frac{d^3k}{(2\pi)^3} f_{\text{BE}}[\omega_k(\sigma), \mu_I], \quad n^{(+)} = \int \frac{d^3k}{(2\pi)^3} f_{\text{BE}}[\omega_k(\sigma), -\mu_I]$$

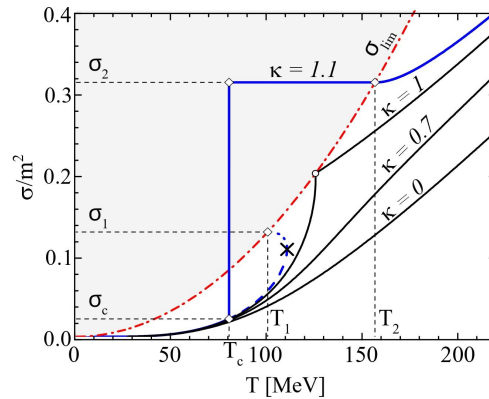
Ideal relativistic gas



Thermodynamic mean-field model



Scalar field model



Thermodynamic mean-field model

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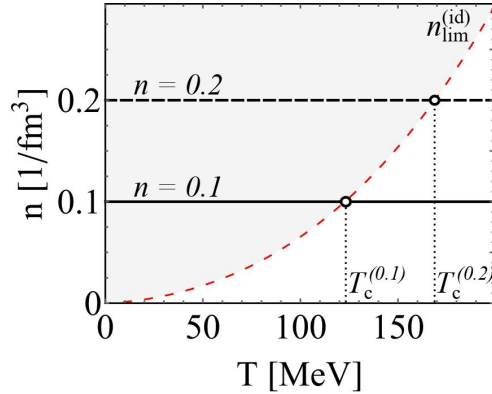
Scalar field model

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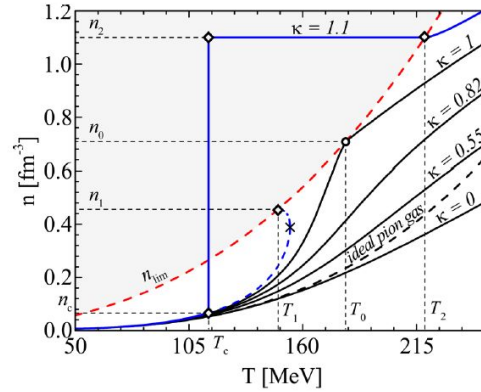
$$n_I = \int \frac{d^3k}{(2\pi)^3} \{ f_{\text{BE}}[\omega_k(\sigma), \mu_I] - f_{\text{BE}}[\omega_k(\sigma), -\mu_I] \} .$$

$$n^{(-)} = \int \frac{d^3k}{(2\pi)^3} f_{\text{BE}}[\omega_k(\sigma), \mu_I], \quad n^{(+)} = \int \frac{d^3k}{(2\pi)^3} f_{\text{BE}}[\omega_k(\sigma), -\mu_I]$$

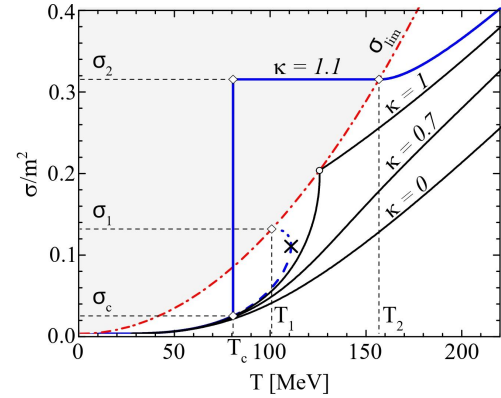
Ideal relativistic gas



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Thermodynamic mean-field model

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$$f_{\text{BE}}(E, \mu) = \left\{ \exp \left[\frac{E - \mu}{T} \right] - 1 \right\}^{-1}$$

Scalar field model

$$\sigma = \int \frac{d^3k}{(2\pi)^3} \frac{1}{2\omega_k(\sigma)} \{ f_{\text{BE}}[\omega_k(\sigma), \mu_I] + f_{\text{BE}}[\omega_k(\sigma), -\mu_I] \} ,$$

$$n_I = \int \frac{d^3k}{(2\pi)^3} \{ f_{\text{BE}}[\omega_k(\sigma), \mu_I] - f_{\text{BE}}[\omega_k(\sigma), -\mu_I] \} .$$

$$n = \int \frac{d^3k}{(2\pi)^3} \left\{ f_{\text{BE}}[\omega_k(\sigma), \mu_I] + f_{\text{BE}}[\omega_k(\sigma), -\mu_I] \right\}$$

Thermodynamic mean-field model^[1-3]

Starting point

Free energy density:

$$\phi(n_1, n_2, T) = \phi_1^{(0)}(n_1, T) + \phi_2^{(0)}(n_2, T) + \phi_{\text{int}}(n, T),$$



Self-consistent set of equations

$$n = \int \frac{d^3k}{(2\pi)^3} [f_{\text{BE}}(E(k, n), \mu_I) + f_{\text{BE}}(E(k, n), -\mu_I)],$$

$$n_I = \int \frac{d^3k}{(2\pi)^3} [f_{\text{BE}}(E(k, n), \mu_I) - f_{\text{BE}}(E(k, n), -\mu_I)].$$

Fixed

$$E(k, n) = \sqrt{m^2 + \mathbf{k}^2} + U(n)$$

$$f_{\text{BE}}(E, \mu) = \left\{ \exp\left[\frac{E - \mu}{T}\right] - 1 \right\}^{-1}$$

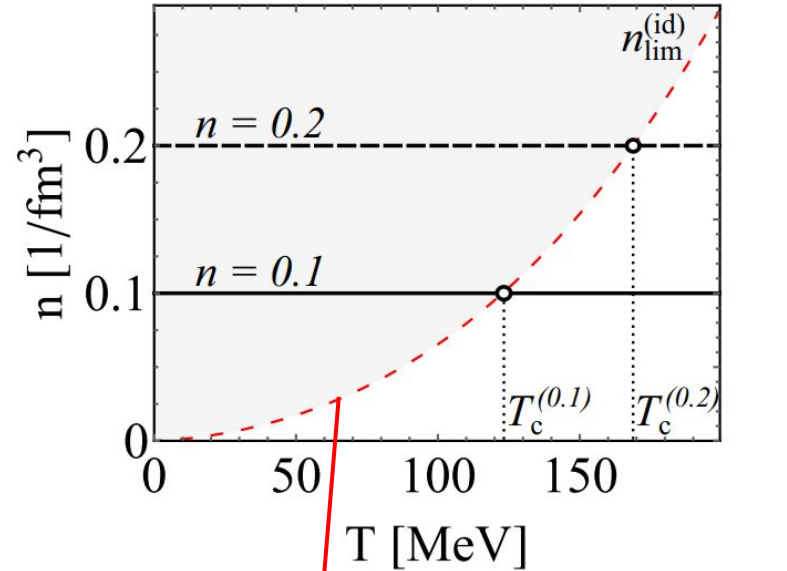


Parameterization of the mean-field

$$U(n) = -An + Bn^2, \quad n_1 = \sqrt{\frac{m}{B}}(\kappa - \sqrt{\kappa^2 - 1}), \quad n_2 = \sqrt{\frac{m}{B}}(\kappa + \sqrt{\kappa^2 - 1}),$$

$$\kappa \equiv \frac{A}{2\sqrt{mB}}$$

Interplay between A and B
Parameter B - fixed



$$n_{\text{lim}}^{(\text{id})}(T) = g \int_0^\infty \frac{d^3k}{(2\pi)^3} \frac{1}{\exp\left[\frac{\sqrt{m^2 + \mathbf{k}^2} - m}{T}\right] - 1}$$

[1] D. Anchishkin and V. Vovchenko, J. Phys. G 42, 105102 (2015).

[2] D. Anchishkin, I. Mishustin, and H. Stoecker, J. Phys. G 46, 035002 (2019).

[3] D. Anchishkin, V. Gnatovskyy, D. Zhuravel, V. Karpenko, I. Mishustin, H. Stoecker, Universe 2023, 9, 411 (2023).

Introduction

