

# Quantum Geometry from Commutators

A Heisenberg-picture kinematical framework for metric operators, time observables,  
and non-commuting translations

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## The one-sentence question

Can spacetime geometry be encoded directly in the Heisenberg algebra?

### Standard instinct

Start with geometry  $g_{\mu\nu}(x)$ , then quantize fields on it, or try to quantize  $g_{\mu\nu}$  itself.

coordinates commute

metric is an operator

### This framework

Start from commutators. Let the metric be the operator measured by  $[\hat{x}, \hat{P}]$ .

translations need not commute

# What is new, and what is deliberately not claimed

## Established in the paper

- ▶ Closed operator algebra on a common dense domain  $\mathcal{D}$ .
- ▶ Jacobi identities imply an operator analogue of metric compatibility.
- ▶ Time can be treated either relationally through a POVM or self-adjointly after enlarging the sector structure.
- ▶ FRW and weak-field checks give transparent non-commuting translations.

## Not claimed

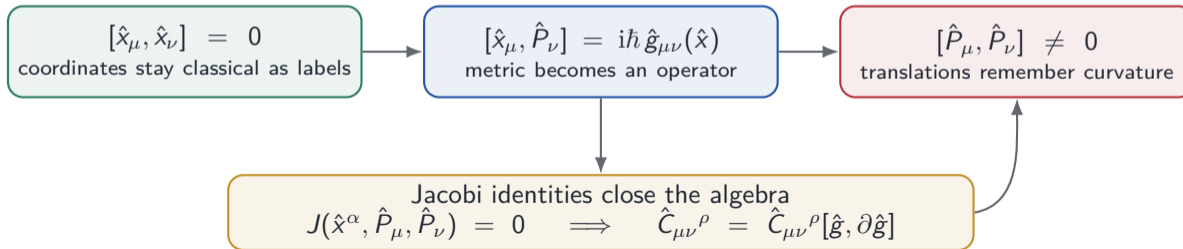
- ▶ No full interacting QFT yet.
- ▶ No derived quantum Einstein equations yet.
- ▶ No automatic black-hole singularity resolution theorem.
- ▶ The high- $z$  halo calculation is a toy propagation of a proxy parameter.

## Core algebra: geometry appears in the cross-commutator

$$[\hat{x}_\mu, \hat{x}_\nu] = 0, \quad (4a)$$

$$[\hat{x}_\mu, \hat{P}_\nu] = i\hbar \hat{g}_{\mu\nu}(\hat{x}), \quad (4b)$$

$$[\hat{P}_\mu, \hat{P}_\nu] = i\hbar \hat{C}_{\mu\nu}{}^\rho(\hat{x}) \hat{P}_\rho \quad \text{minimal scalar sector.} \quad (4c)$$



## The metric operator is Hermitian by construction

### Definition

$$\hat{g}_{\mu\nu} \equiv \frac{1}{2i\hbar} \left( [\hat{x}_\mu, \hat{P}_\nu] + [\hat{x}_\nu, \hat{P}_\mu] \right), \quad \eta_{\mu\nu} = \text{diag}(-, +, +, +).$$

$$\hat{g}_{\mu\nu} = \hat{g}_{\nu\mu}$$

symmetric

$$\langle \hat{g} \rangle \rightarrow g$$

classical limit

$$g \rightarrow \eta$$

flat limit

The commutator is not deformed by geometry; **geometry is what the commutator returns.**

# Jacobi identities do the heavy lifting

## The nontrivial Jacobi identity

$$[\hat{x}^\alpha, [\hat{P}_\mu, \hat{P}_\nu]] = [[\hat{x}^\alpha, \hat{P}_\mu], \hat{P}_\nu] + [\hat{P}_\mu, [\hat{x}^\alpha, \hat{P}_\nu]].$$

## Closure condition

$$\hat{C}_{\mu\nu}{}^\rho \hat{g}_{\alpha\rho} = \hat{g}_{\beta\nu} \partial_\beta \hat{g}_{\alpha\mu} - \hat{g}_{\beta\mu} \partial_\beta \hat{g}_{\alpha\nu}. \quad (7)$$

$$\hat{C}_{\mu\nu}{}^\rho = \hat{g}^{\rho\alpha} (\hat{g}_{\beta\nu} \partial_\beta \hat{g}_{\alpha\mu} - \hat{g}_{\beta\mu} \partial_\beta \hat{g}_{\alpha\nu}). \quad (9)$$

Interpretation: the  $\hat{C}$ 's act like connection/anhonomy coefficients for the algebraic translation frame.

## Including spin/tangent indices adds the Lorentz sector

For fields with tangent or spin indices

$$[\hat{P}_\mu, \hat{P}_\nu] = i\hbar \hat{C}_{\mu\nu}{}^\rho \hat{P}_\rho + \frac{i\hbar}{2} \hat{R}_{\mu\nu}{}^{ab} \hat{J}_{ab}. \quad (11)$$

Scalar sector

$\hat{J}_{ab}$  acts trivially

anholonomy / torsion-like part

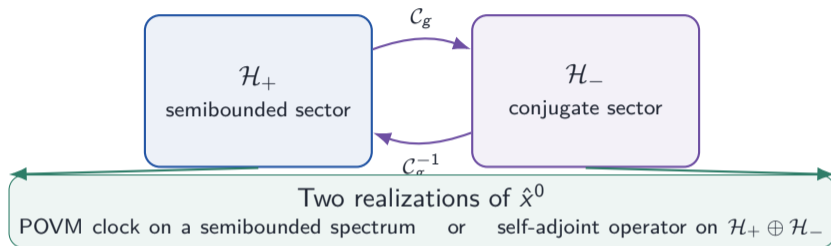
Tensor/spinor sector

$\hat{J}_{ab}$  acts on local indices

curvature enters explicitly

This reproduces the familiar geometry of covariant derivative commutators, but as an operator-algebra statement.

## Time as an observable: two consistent routes



### Route I: relational time

$\hat{x}^0$  is maximally symmetric and generates a POVM  $\Pi_t$ . Predictions are conditioned on clock outcomes.

### Route II: enlarged spectrum

With  $C_g$ , use  $\mathcal{H}_{\text{phys}} = \mathcal{H}_+ \oplus \mathcal{H}_-$ , so Pauli's semibounded-spectrum premise no longer applies.

## A concrete scalar representation

### Coordinate representation

On  $\mathcal{H} = L^2(\mathbb{R}^4, d^4x)$ ,

$$(\hat{x}^\mu \psi)(x) = x^\mu \psi(x), \quad (\hat{P}_\mu \psi)(x) = -i\hbar g_{\mu\nu}(x) \partial_\nu \psi(x). \quad (13)$$

- ▶ Realizes  $[\hat{x}_\mu, \hat{P}_\nu] = i\hbar g_{\mu\nu}$ .
- ▶ Gives  $[\hat{P}_\mu, \hat{P}_\nu] = i\hbar C_{\mu\nu}{}^\rho \hat{P}_\rho$ .
- ▶ Operator promotion means  $g \mapsto \hat{g}$ .

Useful check:  
the algebra has ordinary-looking representations before demanding full dynamics.

## Worked example: spatially flat FRW

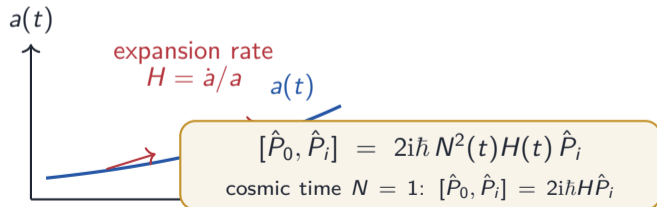
### Metric and momenta

$$ds^2 = -N^2(t)dt^2 + a^2(t)d\mathbf{x}^2, \quad (16)$$

$$\hat{P}_0 = i\hbar N^2(t)\partial_t, \quad \hat{P}_i = -i\hbar a^2(t)\partial_i. \quad (17)$$

### Translation algebra

$$[\hat{P}_i, \hat{P}_j] = 0, \quad [\hat{P}_0, \hat{P}_i] = 2i\hbar N^2(t)H(t)\hat{P}_i. \quad (19)$$



## What the FRW commutator is saying

$$H(t_*) = 0$$

translations commute at that instant

finite  $H$

algebra is well-defined on the representation domain

$$|H| \rightarrow \infty$$

background representation breaks down

### Interpretation

The non-commutativity is controlled by the expansion rate, not inserted by hand. In this scalar realization it is naturally read as frame/anhonomy data, while curvature for spin/tensor fields lives in the Lorentz sector.

Expansion becomes algebra.

## Second check: weak-field limit

### Perturb around Minkowski space

$$g_{\mu\nu}(x) = \eta_{\mu\nu} + h_{\mu\nu}(x), \quad |h_{\mu\nu}| \ll 1, \quad (20)$$

$$C_{\mu\nu}{}^\rho = \partial_\nu h^\rho{}_\mu - \partial_\mu h^\rho{}_\nu + \mathcal{O}(h^2). \quad (21)$$

### Static Newtonian specialization

$$ds^2 = -(1 + 2\Phi)dt^2 + (1 - 2\Psi)d\mathbf{x}^2 \quad \Longrightarrow \quad [\hat{P}_0, \hat{P}_i] = 2i\hbar(\partial_i\Phi)\hat{P}_0 + \mathcal{O}(h^2). \quad (25)$$

Inhomogeneity, not only cosmological expansion, appears directly as non-commuting translation generators.

# Gravitational conjugation symmetry $\mathcal{C}_g$

## Definition

Let  $\Theta$  be antiunitary. Then

$$\begin{aligned}\hat{x}^\mu &\mapsto \Theta \hat{x}^\mu \Theta^{-1}, \\ \hat{P}_\mu &\mapsto -\Theta \hat{P}_\mu \Theta^{-1}, \\ \hat{g}_{\mu\nu} &\mapsto \Theta \hat{g}_{\mu\nu} \Theta^{-1}, \\ \hat{J}_{ab} &\mapsto \Theta \hat{J}_{ab} \Theta^{-1}.\end{aligned}\tag{26}$$

$$\mathcal{C}_g: \text{antiunitary} \\ \Theta i \Theta^{-1} = -i, \quad \hat{P}_\mu \mapsto -\Theta \hat{P}_\mu \Theta^{-1}$$

$$[\hat{x}, \hat{P}] = i\hbar \hat{g} \quad \Longrightarrow \quad [\hat{x}', \hat{P}'] = i\hbar \hat{g}'$$

$$\begin{aligned}[\hat{P}, \hat{P}] &= i\hbar T \hat{P} + \frac{i\hbar}{2} R \hat{J} \quad \Longrightarrow \\ [\hat{P}', \hat{P}'] &= i\hbar T' \hat{P}' + \frac{i\hbar}{2} R' \hat{J}'\end{aligned}$$

Torsion odd, curvature even;  
algebraic form preserved.

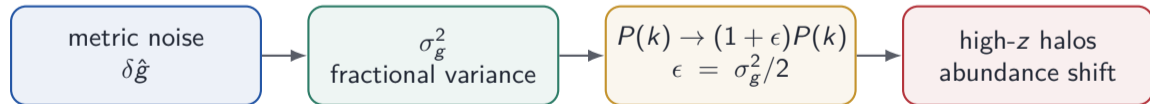
# Uncertainty relations and metric fluctuations

From commutator to uncertainty

$$\Delta x^\mu \Delta P_\nu \geq \frac{\hbar}{2} |\langle \hat{g}^\mu{}_\nu \rangle|.$$

If  $\hat{g} = \bar{g} + \delta\hat{g}$  fluctuates

$$\Delta x^\mu \Delta P_\nu \gtrsim \frac{\hbar}{2} \left( |\bar{g}^\mu{}_\nu| + \mathcal{O} \left( \langle (\delta\hat{g})^2 \rangle^{1/2} \right) \right). \quad (31)$$



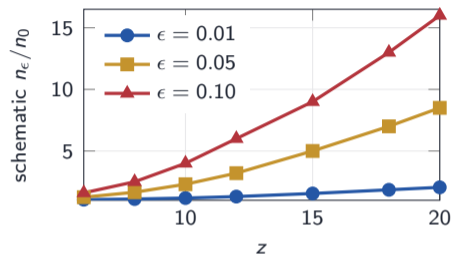
## Toy phenomenology: high- $z$ structure as an amplifier

### Proxy used in Appendix D

$$P(k) \rightarrow (1 + \epsilon)P(k), \quad \epsilon = \frac{1}{2}\sigma_g^2, \quad (\text{D1})$$

$$\sigma(M, z; \epsilon) = D(z)\sigma_0(M)\sqrt{1 + \epsilon}. \quad (\text{D3})$$

- ▶ Small amplitude shifts can be magnified in rare high- $z$  halos.
- ▶ A scale-dependent  $\epsilon(k, z)$  would be the real target.
- ▶ This is a forecast channel, not a detection claim.



Schematic sensitivity sketch, not a numerical reproduction of the paper's Fig. 1.

## Positioning relative to neighboring ideas

Approach	Primary deformation / input	Difference here
GUP / minimal length	Momentum-dependent deformation of canonical brackets	No momentum-dependent deformation is postulated; $[x, P]$ is geometric data.
Snyder / DFR	Non-commuting spacetime coordinates $[x, x] \neq 0$	Coordinates commute: $[\hat{x}_\mu, \hat{x}_\nu] = 0$ .
Teleparallel / coframe methods	Start with a coframe or connection	Algebra first; anholonomy emerges in useful representations.
Loop quantum gravity	Discrete quantum geometry from holonomies and fluxes	This deck uses a continuous coordinate manifold and an operator metric from $[x, P]$ .

The proposal is best understood as an algebraic kinematics for quantum geometry, not as a direct competitor to every existing program in one heroic slide.

# What a full theory still has to deliver

## Dynamics

Need equations for  $\langle \hat{g}_{\mu\nu} \rangle$  and  $\delta \hat{g}_{\mu\nu}$ : Einstein equations, deformations, or something more radical.

## Representations

Need classification of unitary representations, domains, and sector stability under interactions.

## Matter and fields

Need QFT on the algebra: how do  $(\phi, \pi)$  inherit modified commutators from  $[\hat{X}, \hat{P}]$ ?

## Phenomenology

Need constraints from clocks, interferometers, CMB spectra, and high-z structure.

Translation: the algebra gives a clean playground. Nature still gets a vote, because apparently that is how physics works.

## Take-home message

- 1 Geometry can be encoded kinematically by

$$[\hat{x}_\mu, \hat{P}_\nu] = i\hbar \hat{g}_{\mu\nu}(\hat{x}).$$

- 2 Jacobi closure fixes the translation algebra and yields an operator compatibility condition.
- 3 In FRW,

$$[\hat{P}_0, \hat{P}_i] = 2i\hbar N^2 H \hat{P}_i,$$

so expansion directly controls non-commuting translations.

- 4 Time can be treated operationally through POVMs or self-adjointly on a  $\mathcal{C}_g$ -enlarged Hilbert space.
- 5 Metric-operator fluctuations suggest observable proxy channels, especially high- $z$  structure, but only after dynamics is supplied.

Thank you.

## Backup: Jacobi derivation in one line

$$\begin{aligned}[\hat{x}^\alpha, [\hat{P}_\mu, \hat{P}_\nu]] &= [[\hat{x}^\alpha, \hat{P}_\mu], \hat{P}_\nu] + [\hat{P}_\mu, [\hat{x}^\alpha, \hat{P}_\nu]] \\ i\hbar \hat{C}_{\mu\nu}{}^\rho [\hat{x}^\alpha, \hat{P}_\rho] &= [i\hbar \hat{g}^\alpha{}_\mu, \hat{P}_\nu] + [\hat{P}_\mu, i\hbar \hat{g}^\alpha{}_\nu] \\ -\hbar^2 \hat{C}_{\mu\nu}{}^\rho \hat{g}^\alpha{}_\rho &= -\hbar^2 \left( \hat{g}^\beta{}_\nu \partial_\beta \hat{g}^\alpha{}_\mu - \hat{g}^\beta{}_\mu \partial_\beta \hat{g}^\alpha{}_\nu \right).\end{aligned}$$

Multiply by the inverse metric to recover Eq. (9).

## Backup: FRW commutator check

$$\hat{P}_0 \hat{P}_i \psi = i\hbar N^2 \partial_t (-i\hbar a^2 \partial_i \psi) = \hbar^2 N^2 (2a\dot{a} \partial_i \psi + a^2 \partial_t \partial_i \psi),$$

$$\hat{P}_i \hat{P}_0 \psi = -i\hbar a^2 \partial_i (i\hbar N^2 \partial_t \psi) = \hbar^2 a^2 N^2 \partial_i \partial_t \psi,$$

$$[\hat{P}_0, \hat{P}_i] \psi = 2\hbar^2 N^2 a\dot{a} \partial_i \psi = 2i\hbar N^2 H \hat{P}_i \psi.$$

### Symmetrized momentum operator

$$\hat{P}_\mu = -\frac{i\hbar}{2} (g_{\mu\nu} \partial_\nu + \partial_\nu g_{\mu\nu}).$$

- ▶ Improves Hermiticity with respect to the curved measure  $L^2(\sqrt{|g|} d^4x)$ .
- ▶ The structure-function statements in the algebra survive at the level needed for the kinematical construction.
- ▶ Domain questions are real, but they do not spoil the formal Jacobi closure identities.

## Backup: Sheth–Tormen proxy used in Appendix D

$$\nu(M, z; \epsilon) = \frac{\delta_c}{\sigma(M, z; \epsilon)}, \quad \sigma(M, z; \epsilon) = D(z)\sigma_0(M)\sqrt{1 + \epsilon},$$

$$f_{\text{ST}}(\nu) = A\sqrt{\frac{2a}{\pi}} \nu \exp\left(-\frac{a\nu^2}{2}\right) \left[1 + (a\nu^2)^{-p}\right],$$

$$\frac{dn}{d \ln M} = \frac{\bar{\rho}_{m0}}{M} f_{\text{ST}}(\nu) \frac{d \ln \sigma_0^{-1}}{d \ln M},$$

$$n(> M_{\min}, z; \epsilon) = \int_{\ln M_{\min}}^{\ln M_{\max}} \frac{dn}{d \ln M} d \ln M.$$

Here  $(A, a, p) = (0.3222, 0.707, 0.3)$  and  $\delta_c \simeq 1.686$  in the toy setup.






## Backup: likely questions and disciplined answers

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Question	Clean answer
Is this a full theory of quantum gravity?	Not yet. It is a kinematical operator algebra plus representations and checks.
Does it solve black-hole singularities?	Not as proven in the paper. The algebra suggests possible regularization routes, but no theorem is established.
Why keep $[x, x] = 0$ ?	To isolate geometry in $[x, P]$ and avoid assuming noncommutative spacetime coordinates at the start.
Where is curvature?	For scalars, the minimal sector gives anholonomy-like $C$ terms. For tensors/spinors, curvature enters through the Lorentz-generator term.
What would make it predictive?	A dynamical law for $\hat{g}$ , QFT on the algebra, and constrained $\epsilon(k, z)$ .

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# References

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