

Neural Simulation-Based Inference

Aishik Ghosh, Eddie McGrady, Tae Hyun Park

ML4FP School 2026

2 June 2026



**Georgia Institute
of Technology**

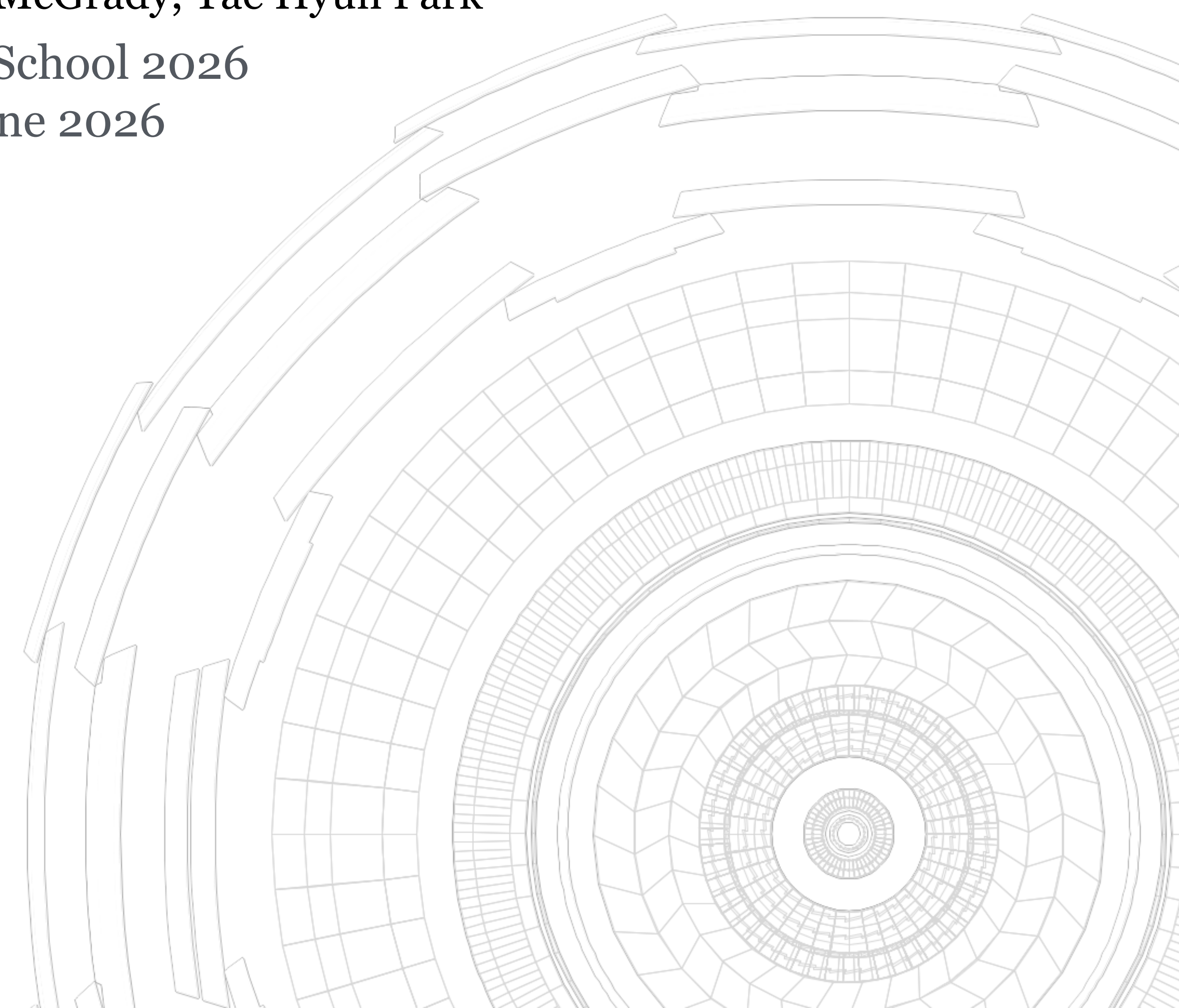


*Some new statistics concepts for some of you,
ask questions!*

 [Prof-Aishik-Ghosh](#)

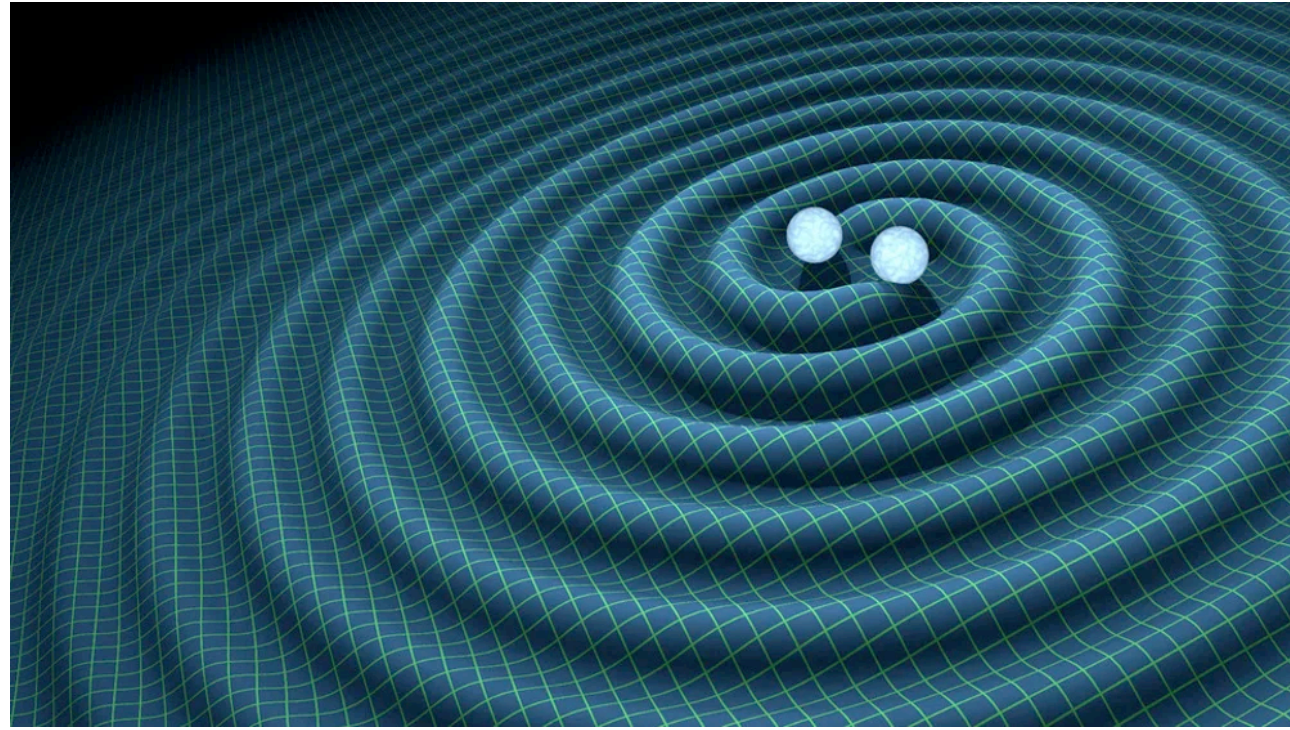
 [aishikghosh.bsky.social](#)

 [@Aishik_Ghosh_](#)



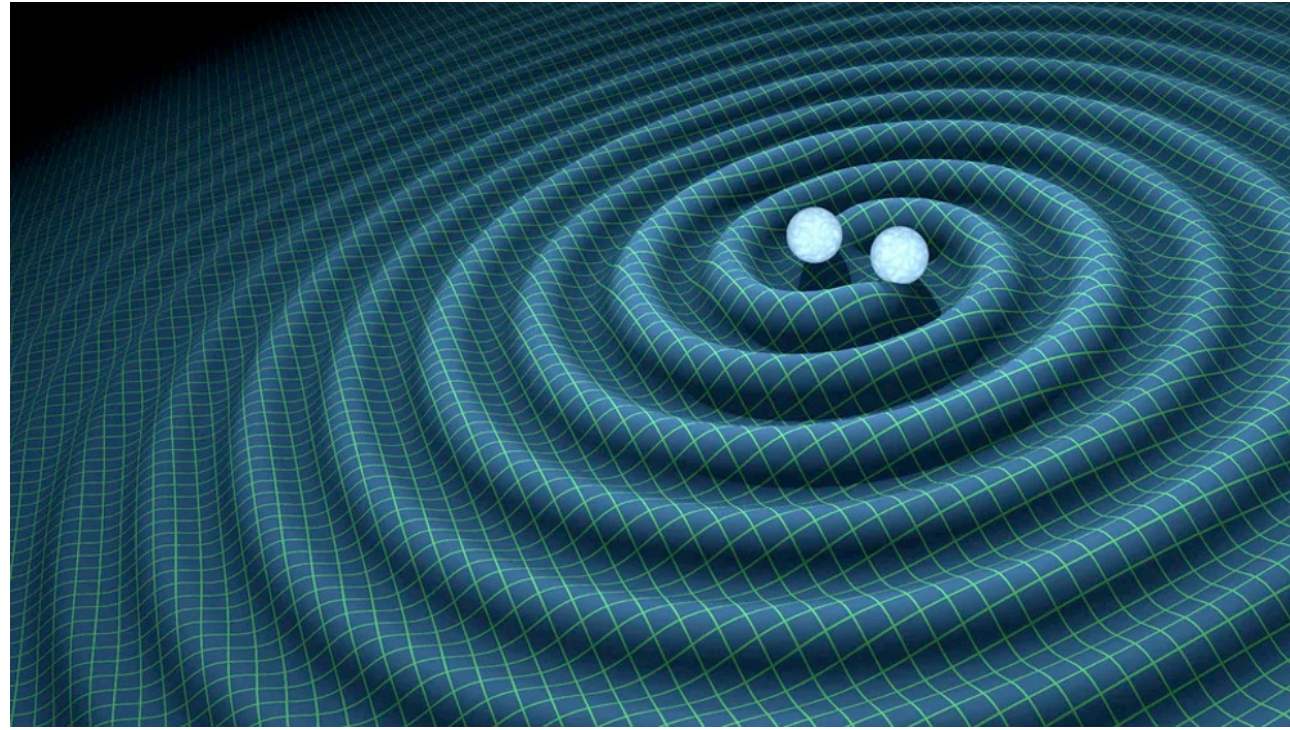
Gravitational Wave discovery (2016)

Astrophysics:

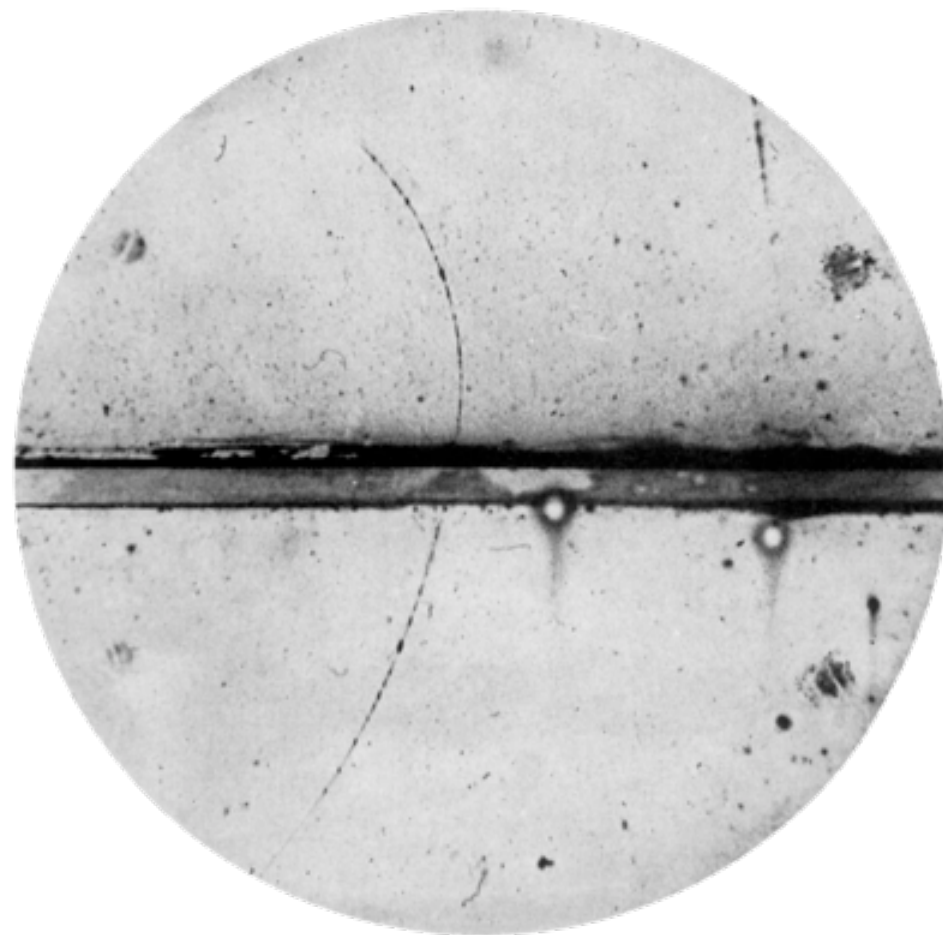


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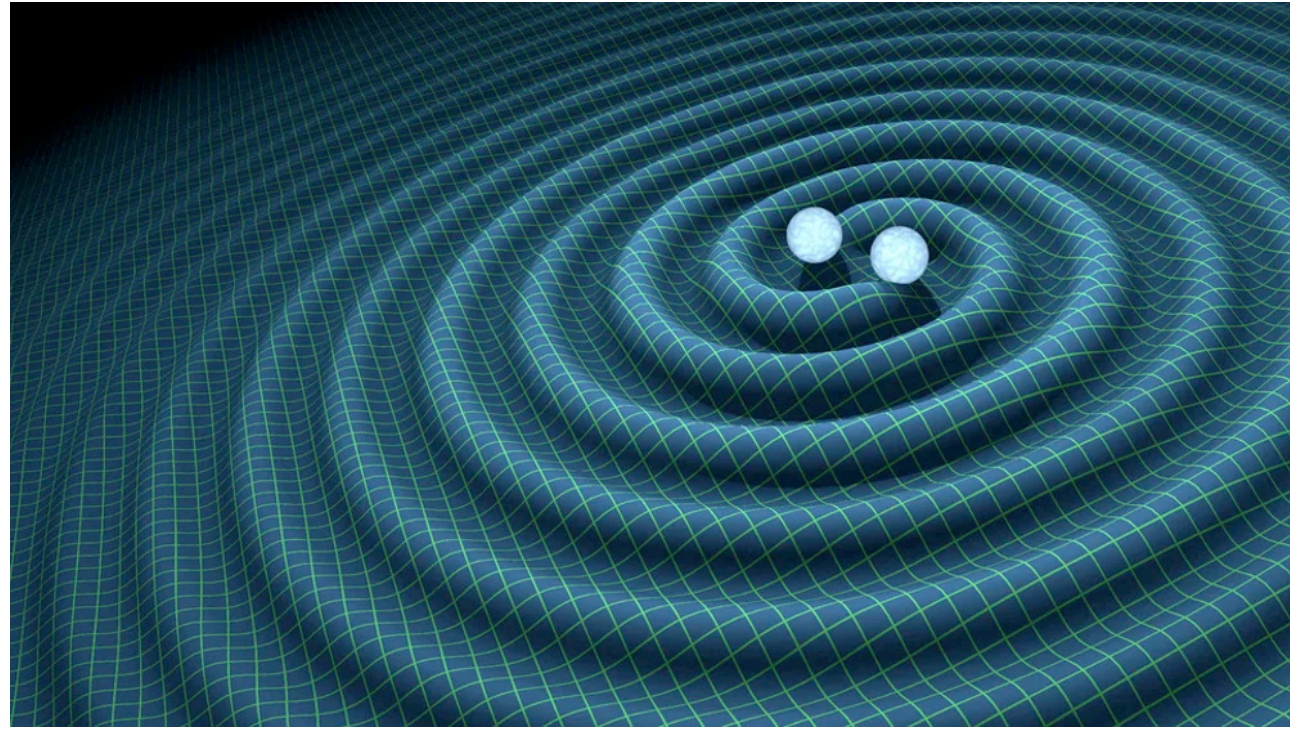
Particle physics:



Positron discovery (1930s)

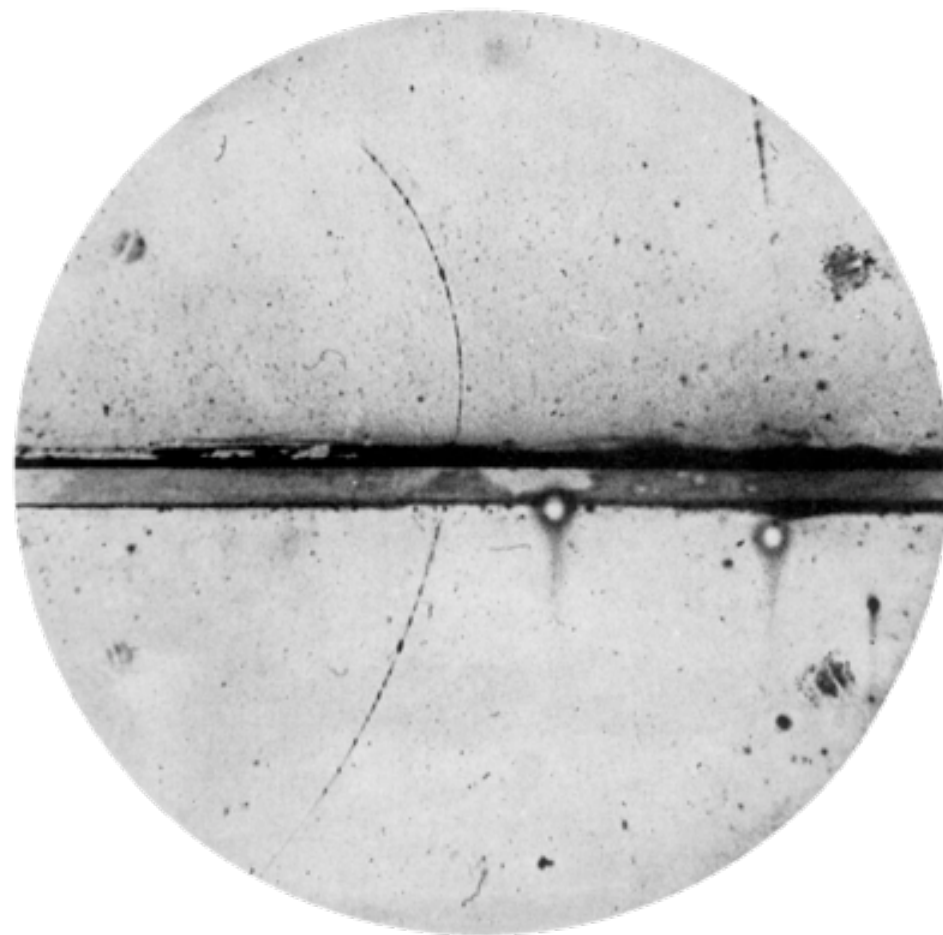
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Single event

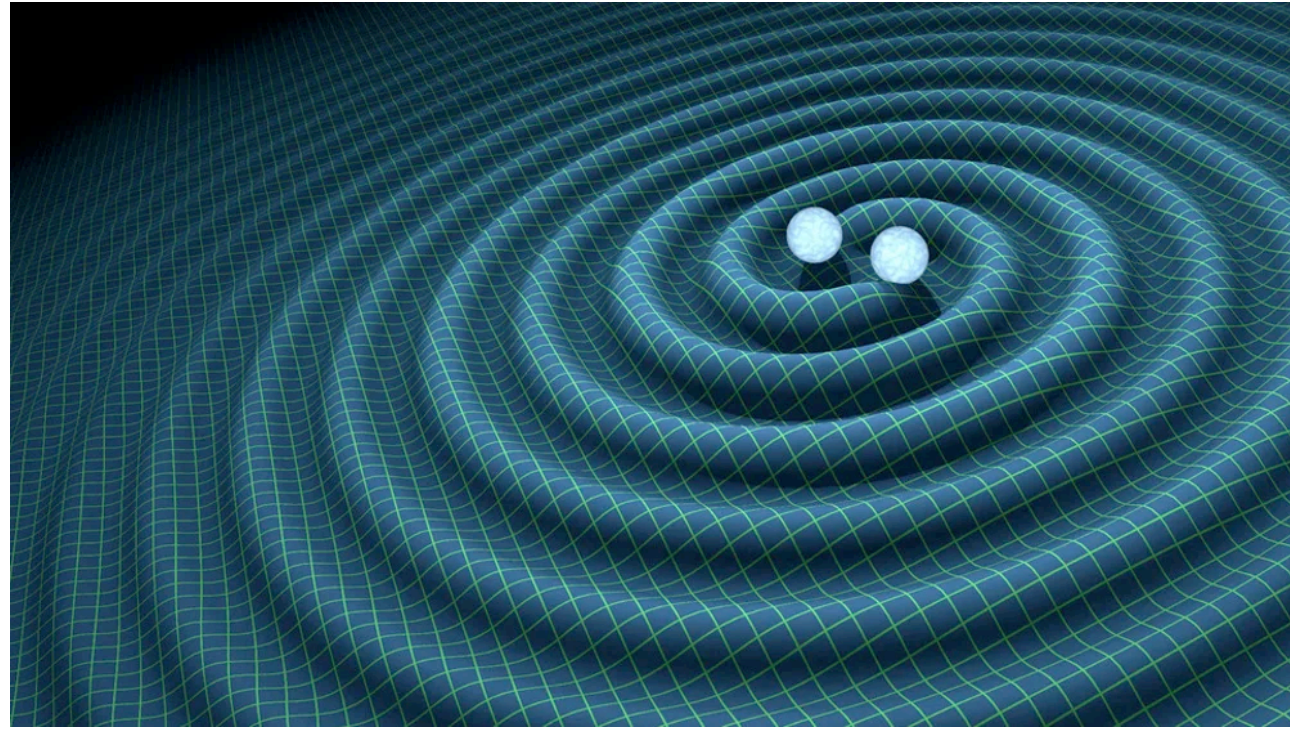
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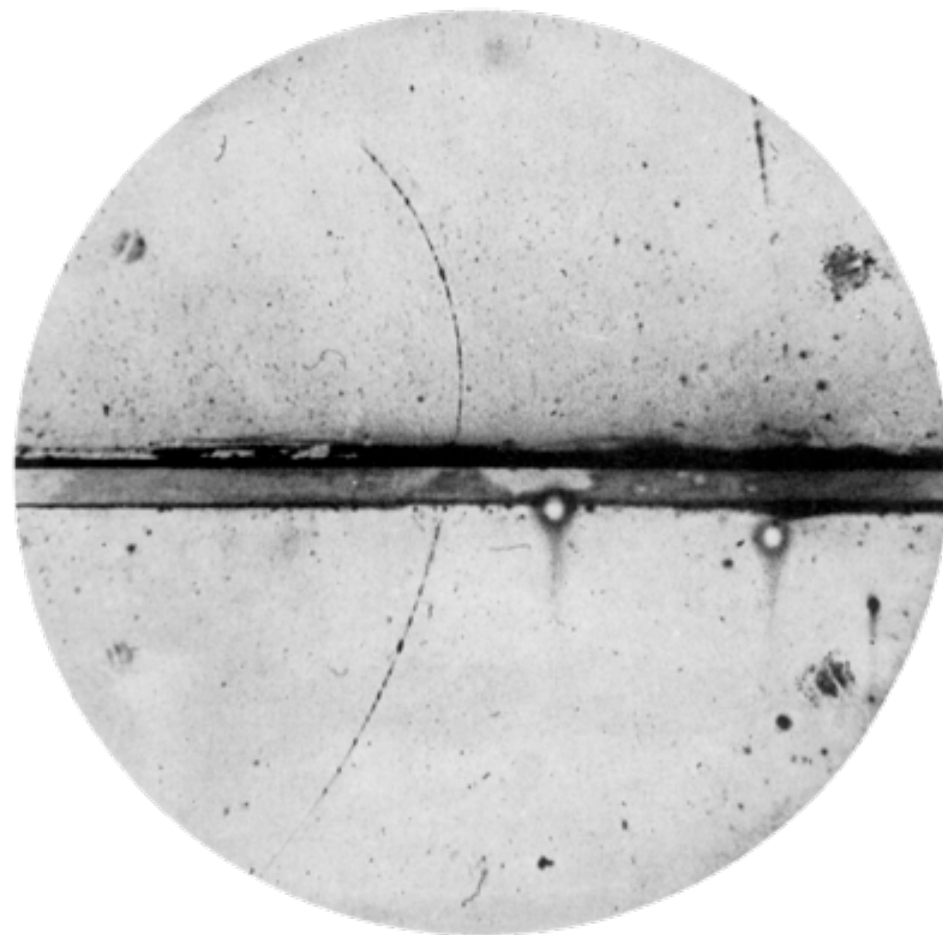
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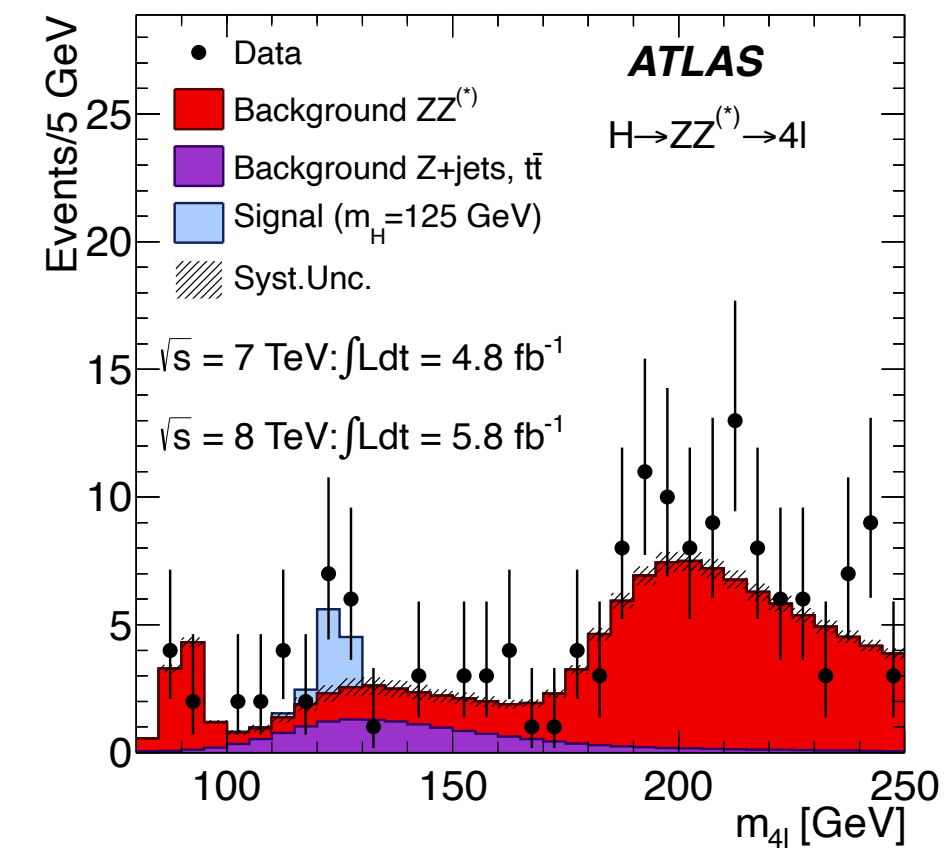
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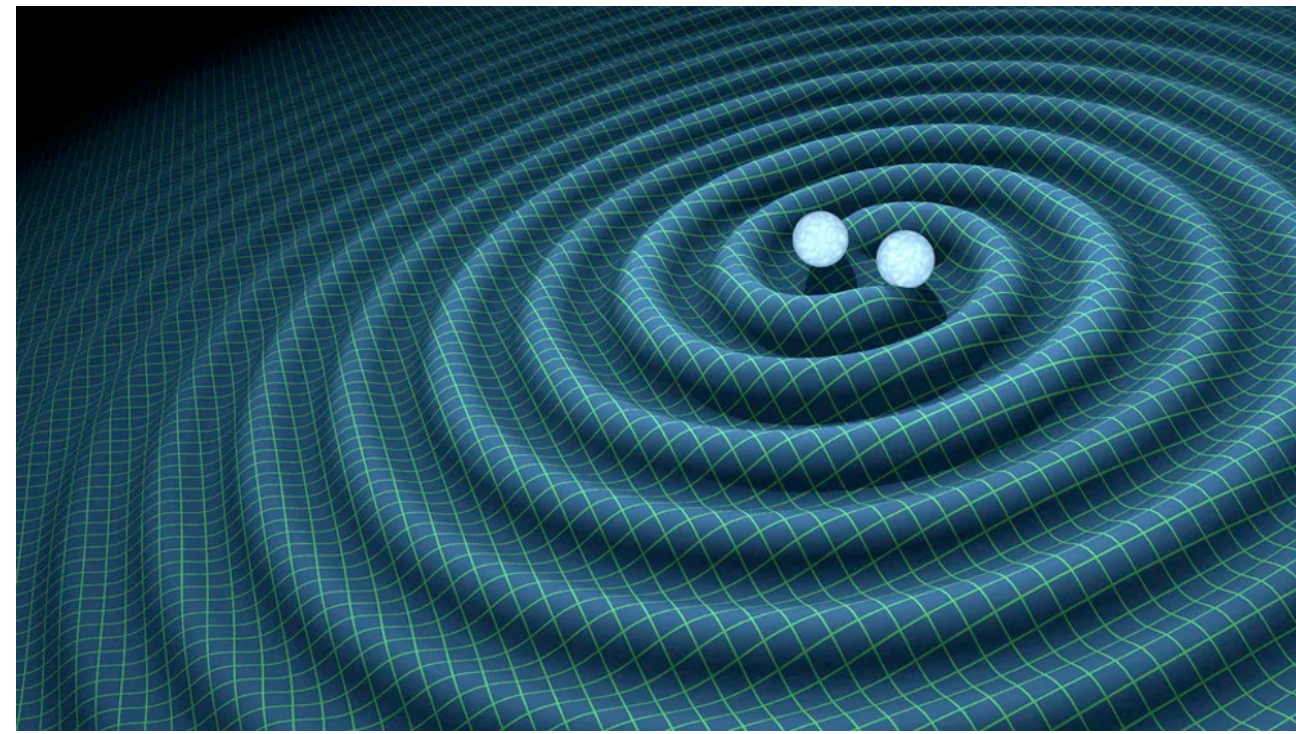


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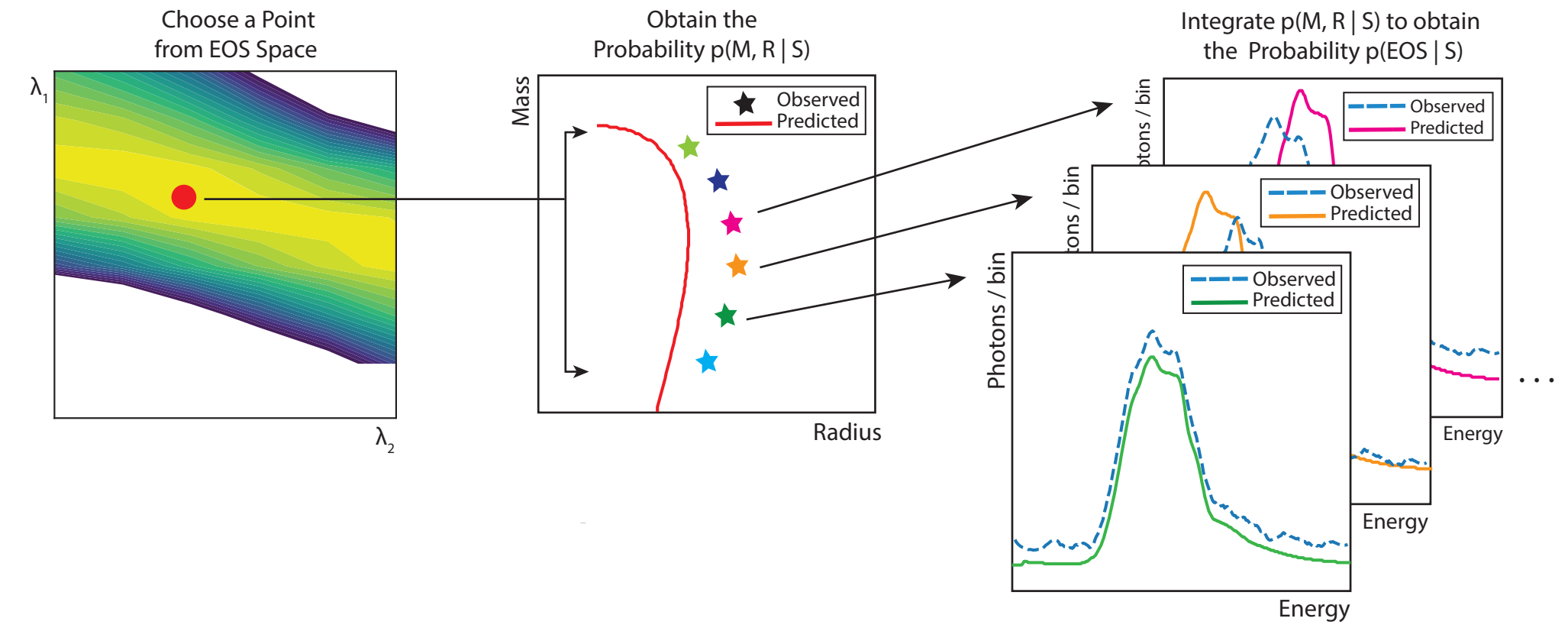
Higgs boson discovery (2012)

Gravitational Wave discovery (2016)



Astrophysics:

Probing the interior of neutron stars

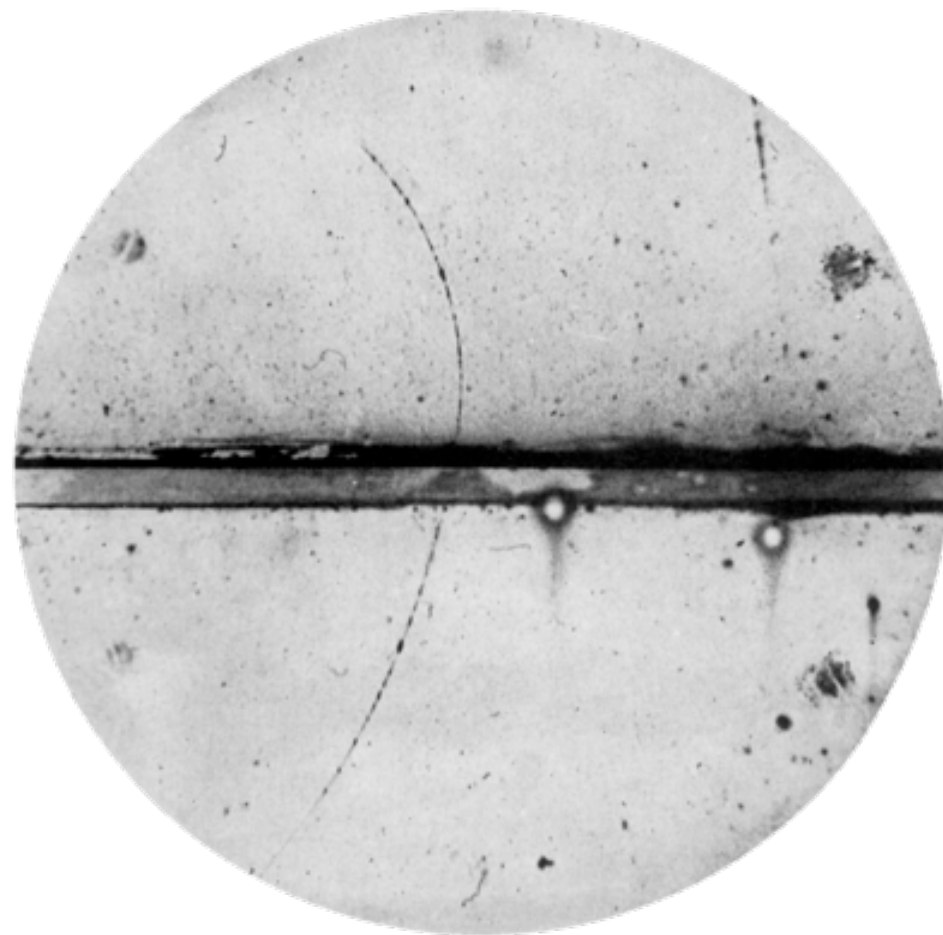


[JCAP12\(2023\)022](#): Farrell et al. (incl. Ghosh)

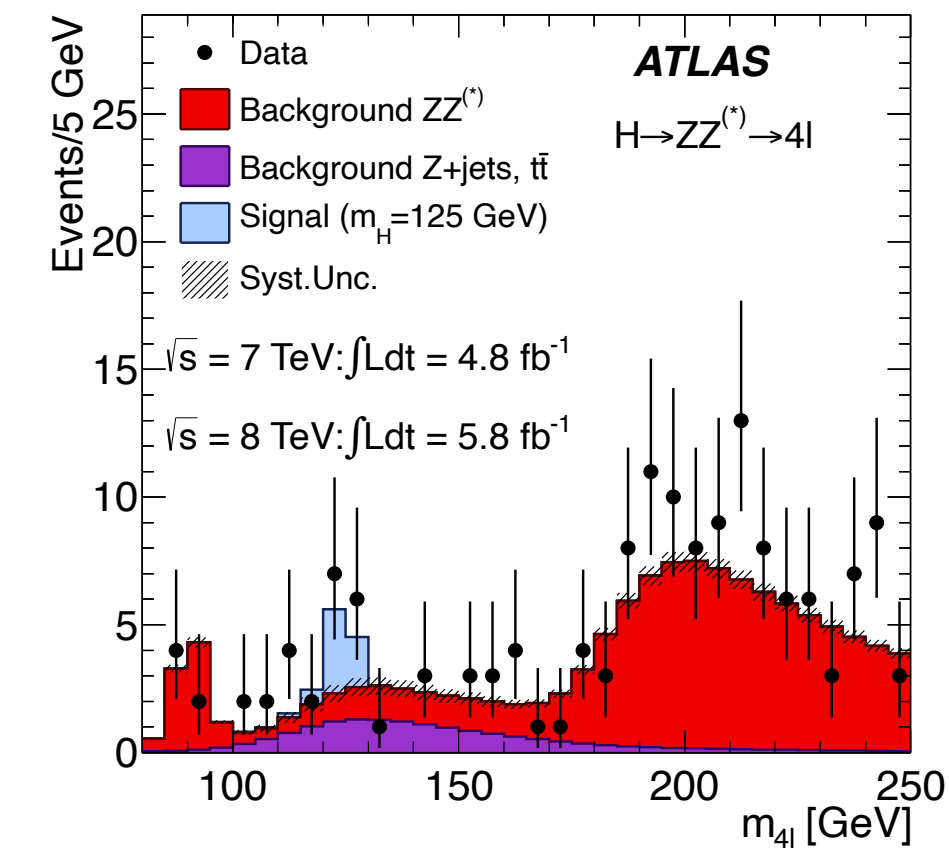
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Statistical inference of physics parameters

$$m_H = 125.25 \pm 0.17 \text{ GeV}$$

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“How sure am I ? How can I reduce my uncertainty ?”

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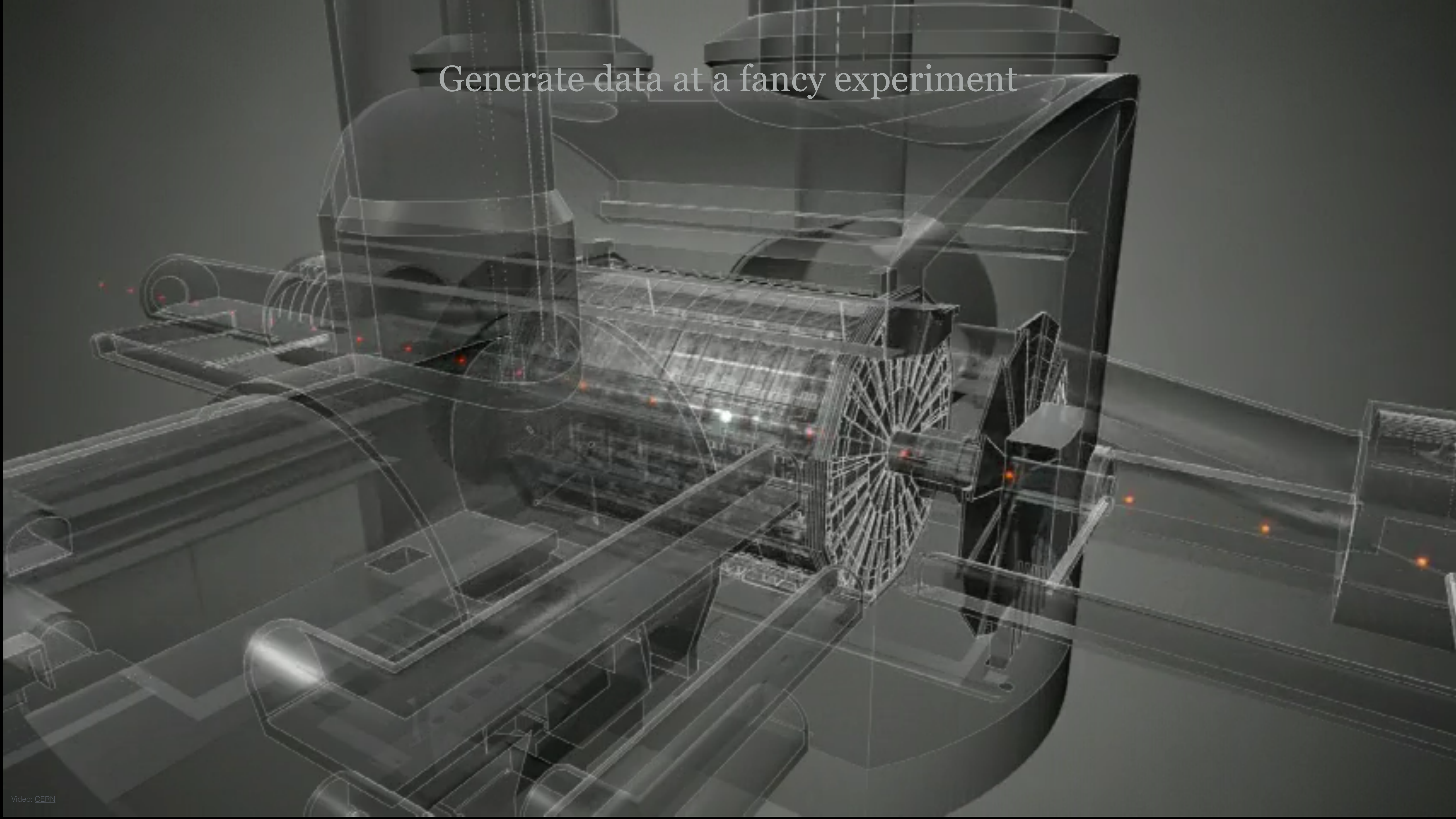
Statistical uncertainties, but also detector systematic, theory systematic, MC stat,



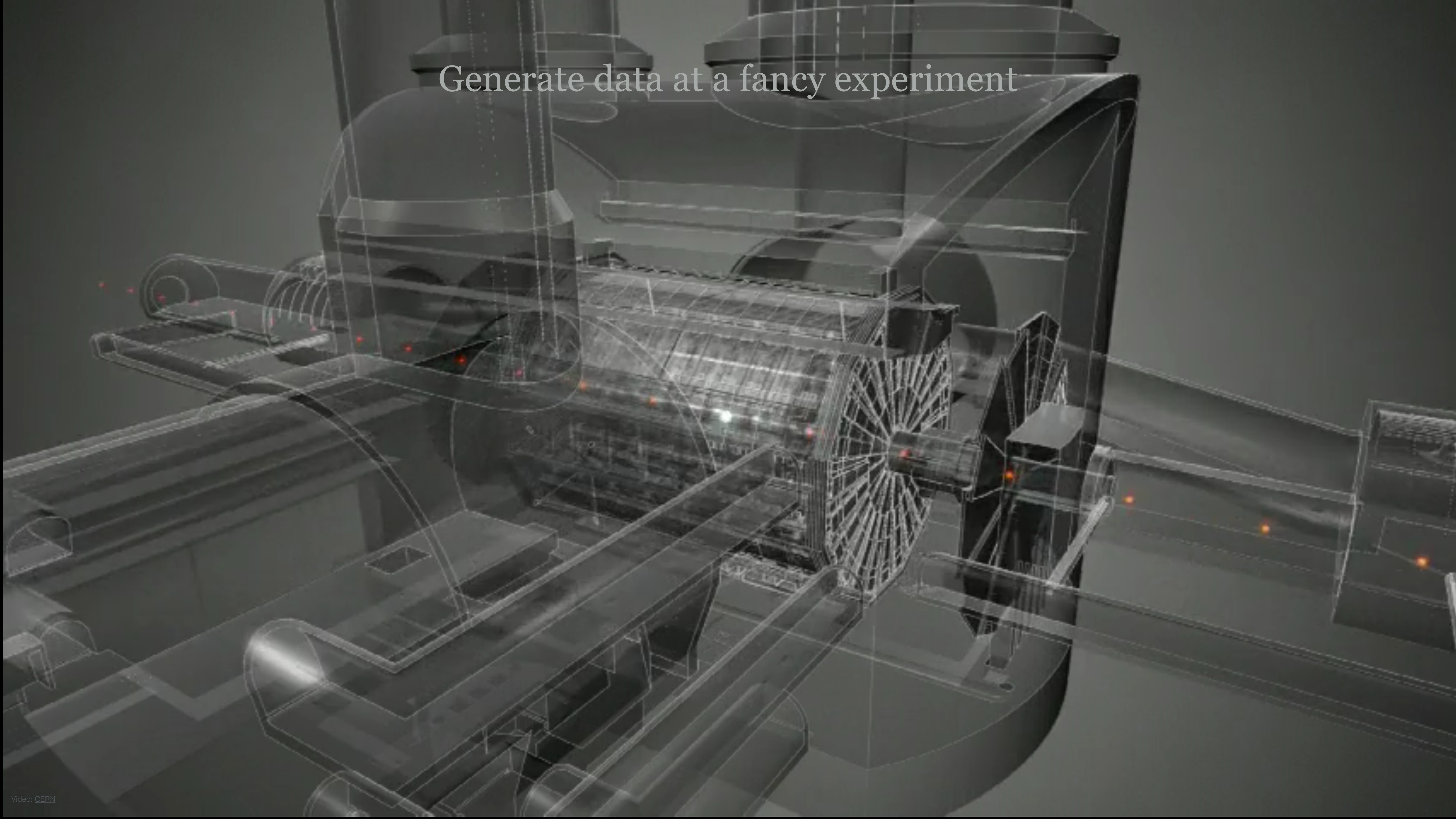
“How sure am I ? How can I reduce my uncertainty ?”

An experimentalist thinks about ‘confidence intervals’

Generate data at a fancy experiment

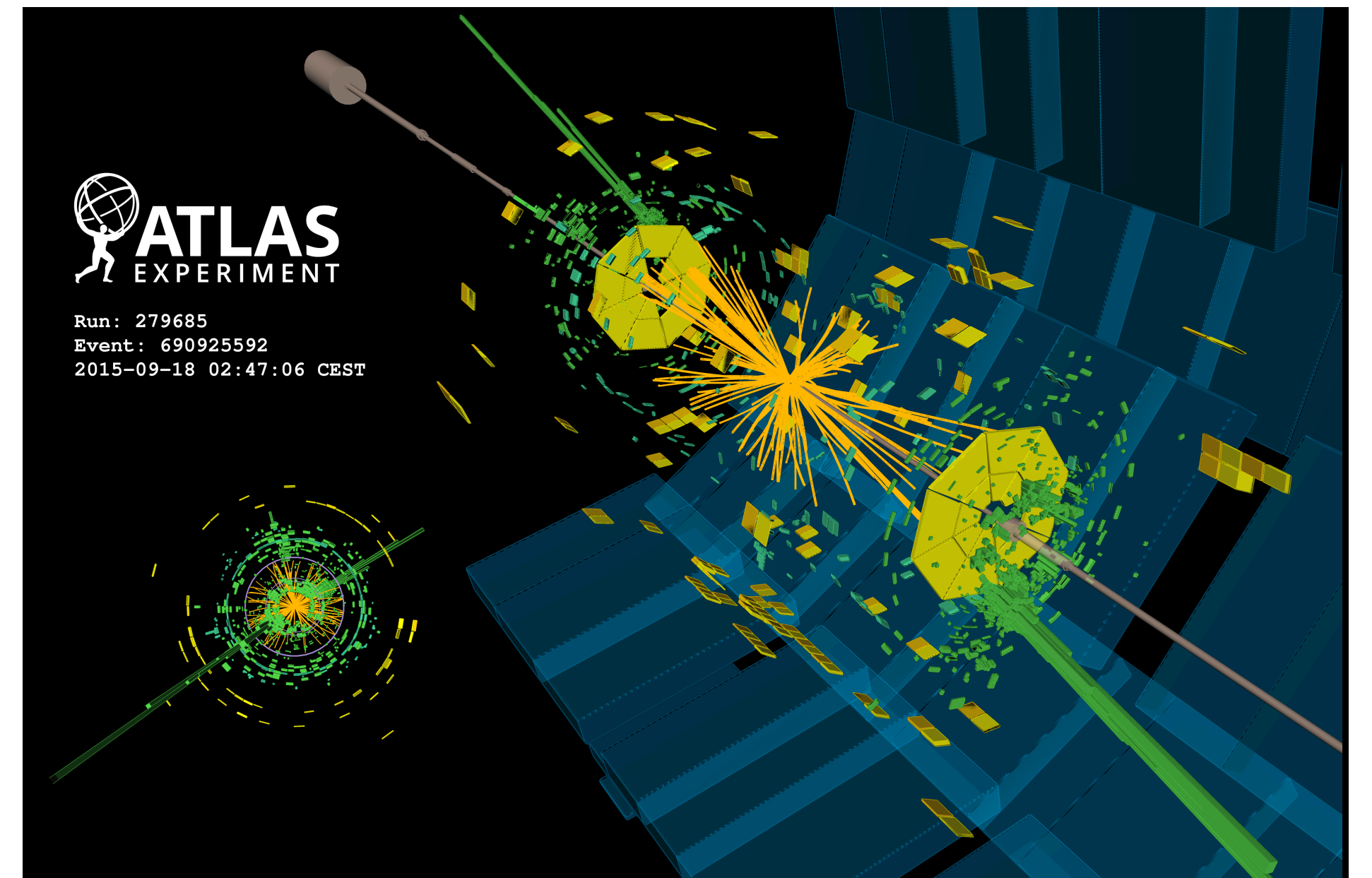


Generate data at a fancy experiment



Traditional Approach: Design one sensitive observable

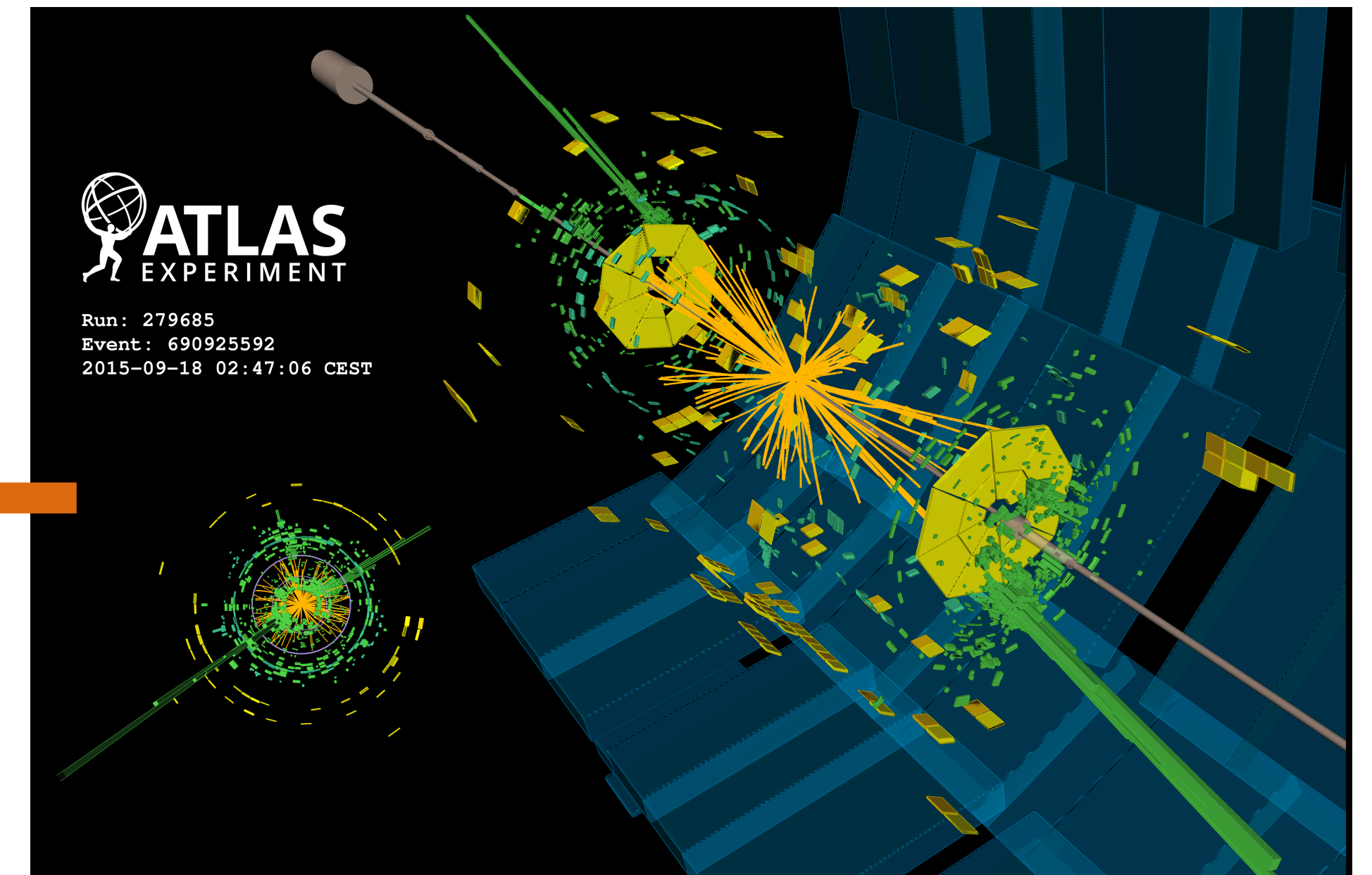
- Detector has $O(100 \text{ million})$ sensors
- Reconstruction pipeline, event selection
- Design one summary variable
 - Compression: $O(100 \text{ million}) \rightarrow 1$
- Build a 1-D histogram
- Estimate probability densities with histograms



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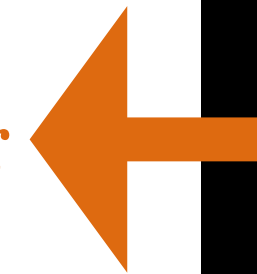
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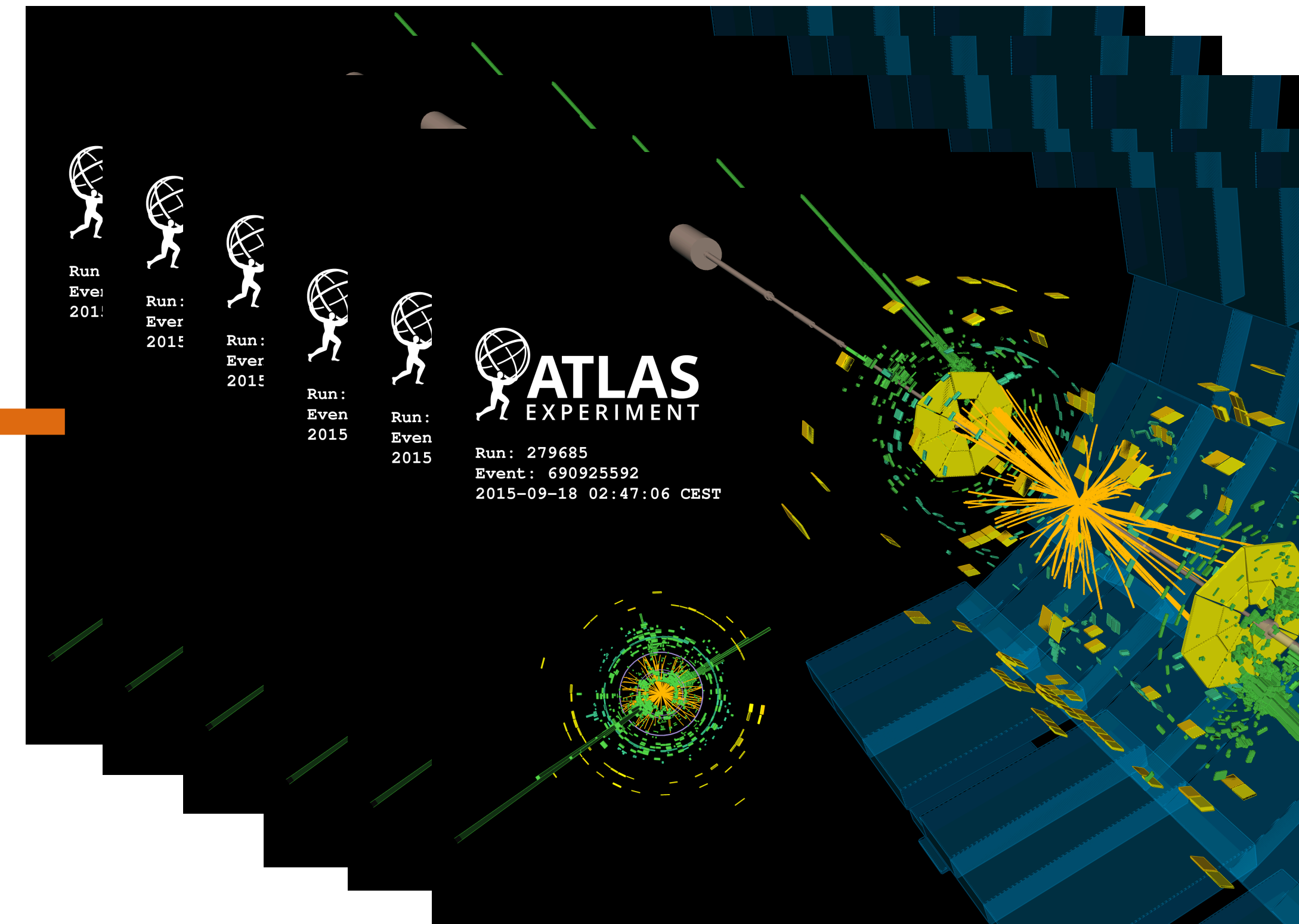
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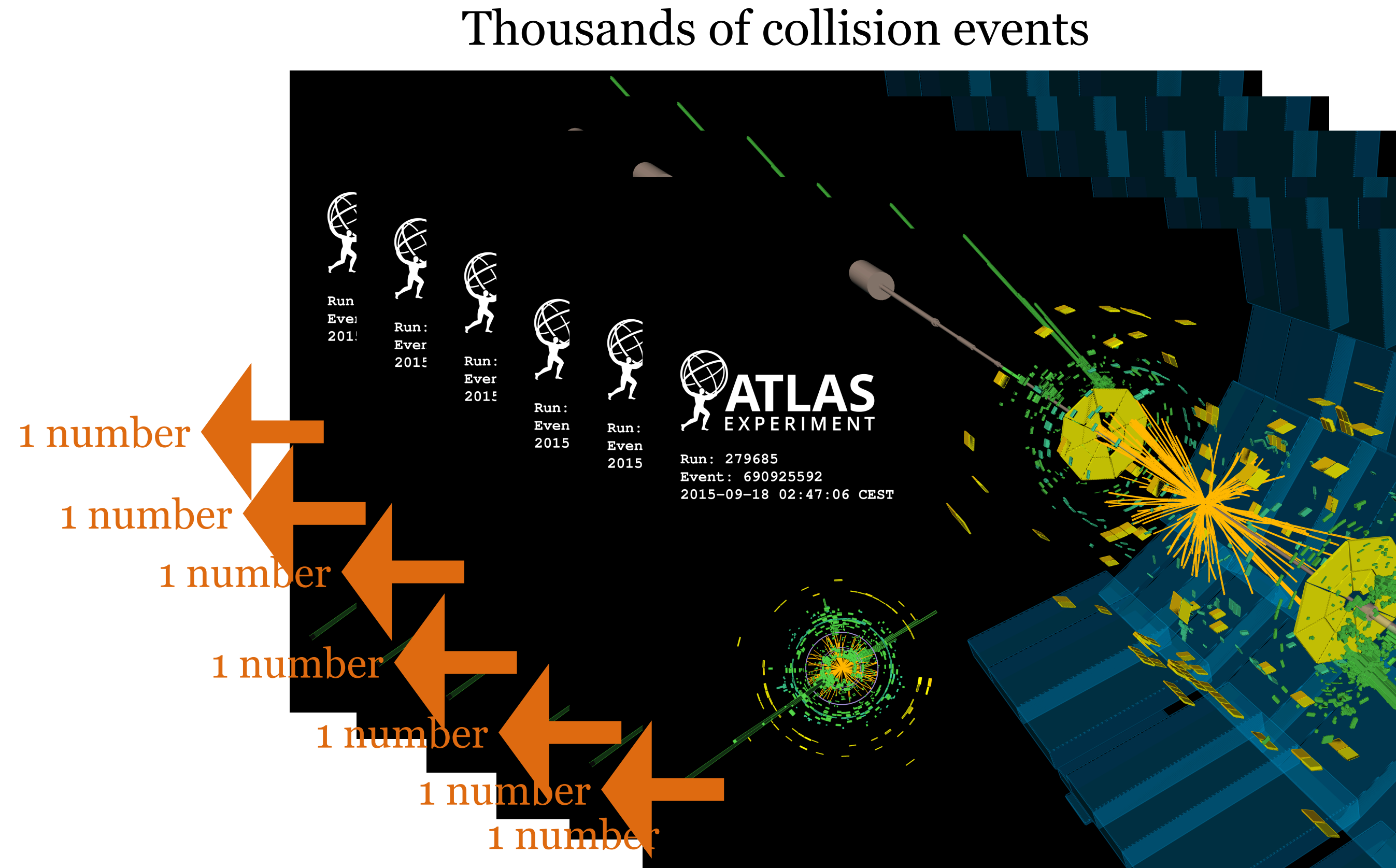


Thousands of collision events



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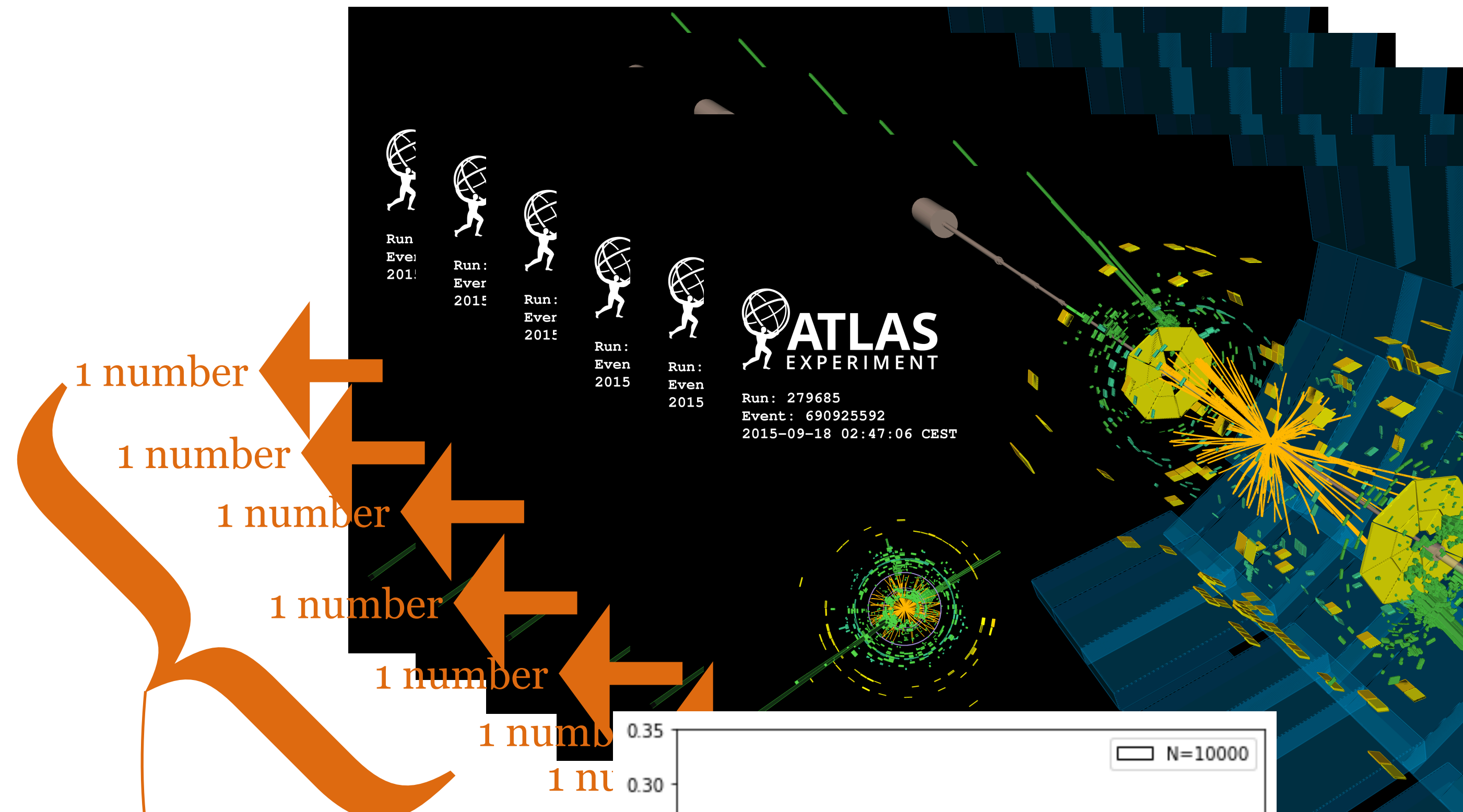
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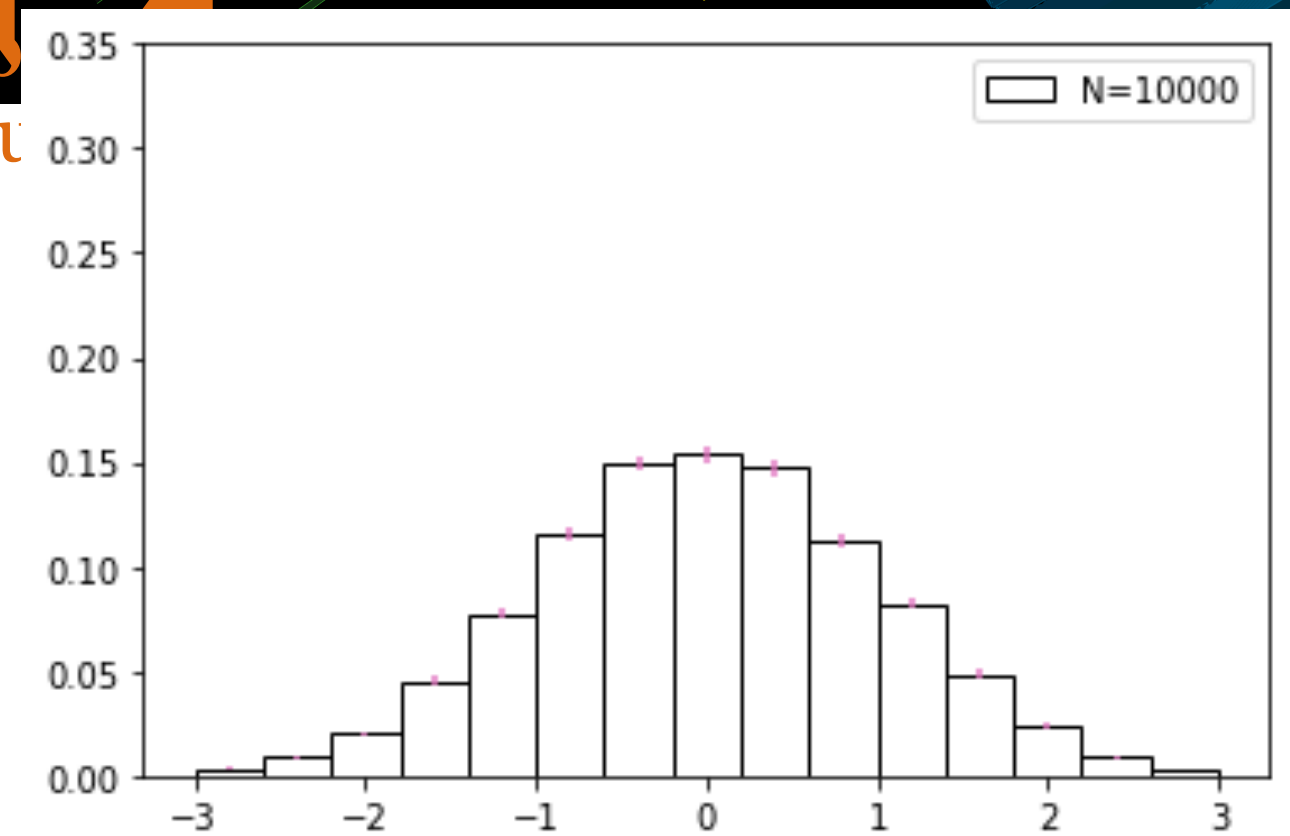
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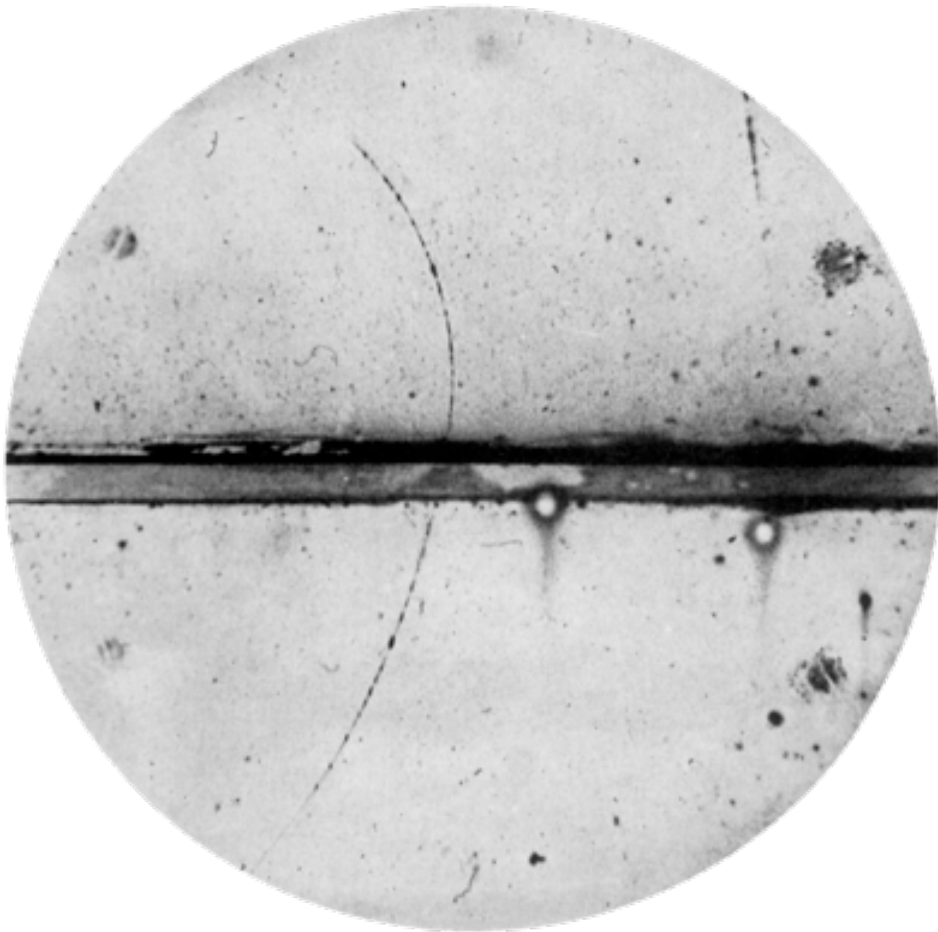


Why histogram help calculate probabilities?



Positron discovery (1930s)

Top quark discovery (1990s)



Channel:	SVX
observed	27 tags
expected background	6.7 ± 2.1
background probability	2×10^{-5}

Single event

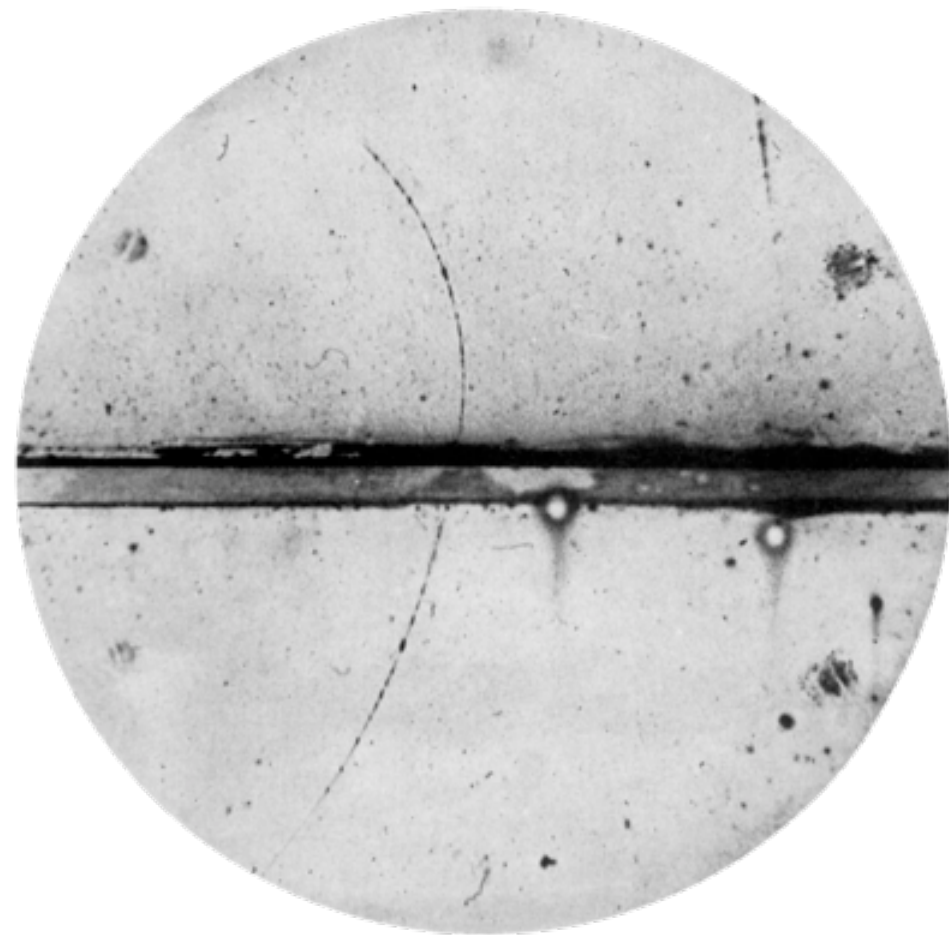
Multiple events:
Cut-and-count

Image: [Wikipedia / PhysRev.43.491](#)

CDF Collaboration: [arXiv:9503002](#)

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What is the probability of seeing 27 events if my null hypothesis predicts 7?

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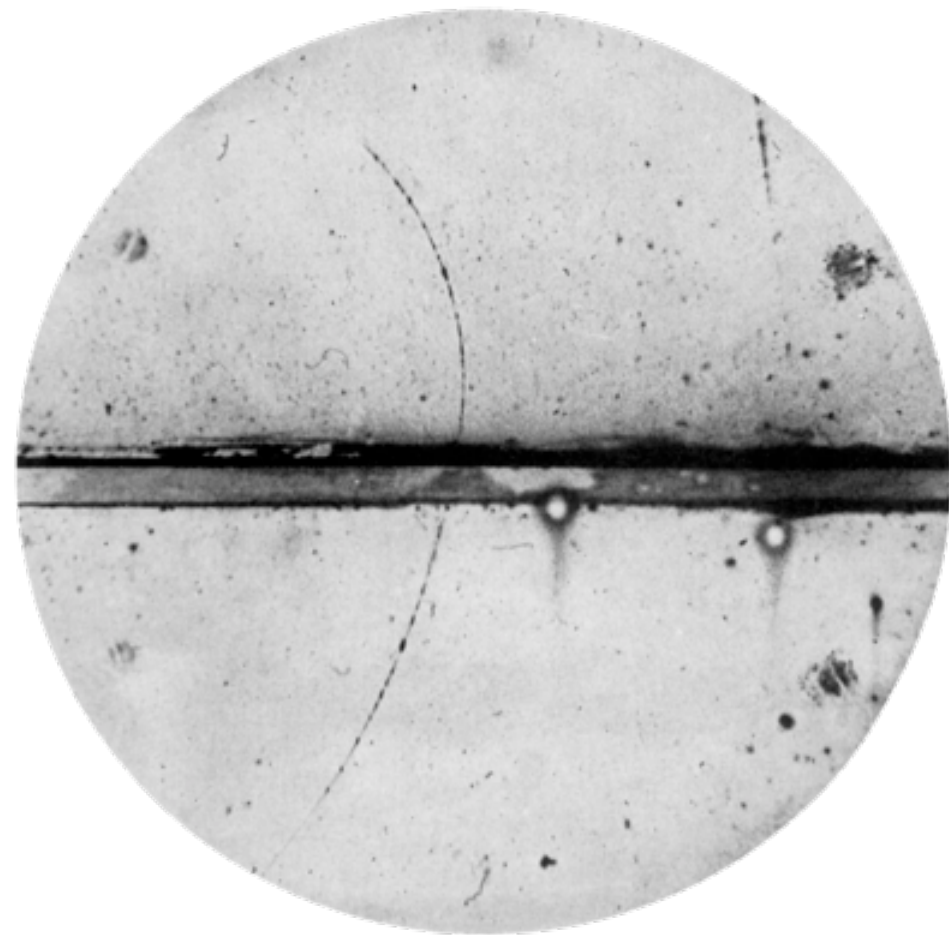
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$$p(n_{obs} | n_{exp}) = \frac{n_{exp}^{n_{obs}} \cdot e^{-n_{exp}}}{n_{obs}!}$$

Counting experiment → Poisson distribution

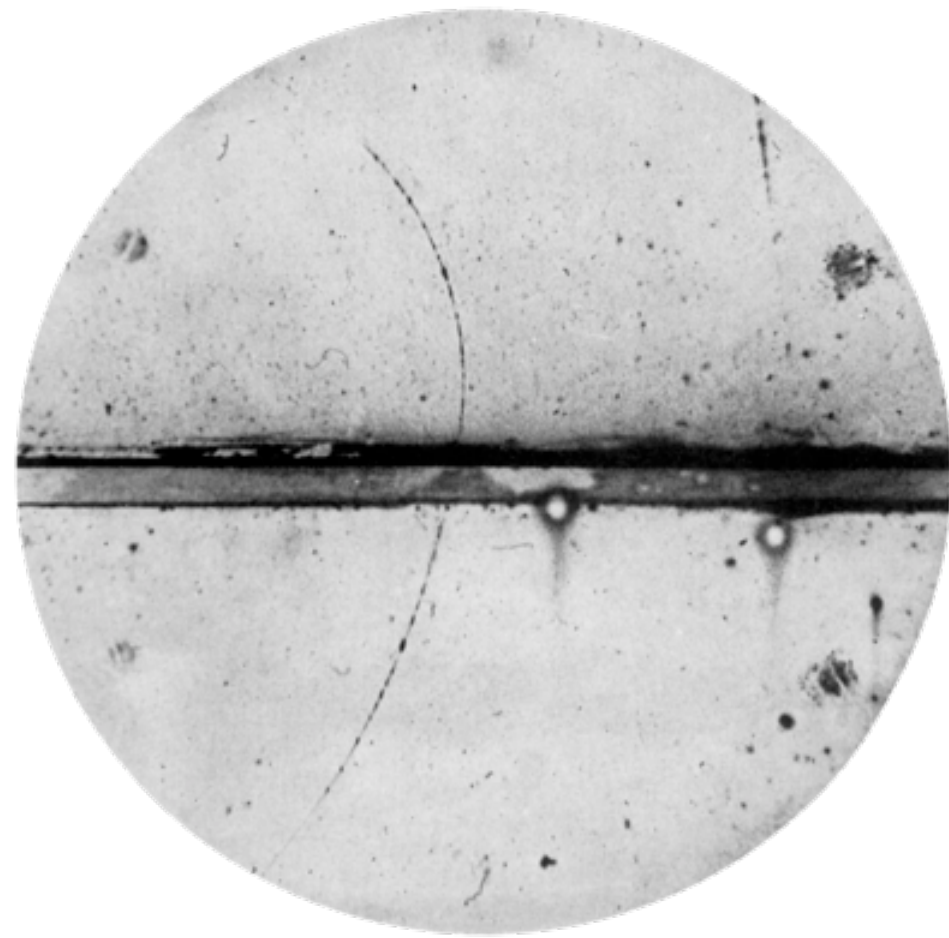
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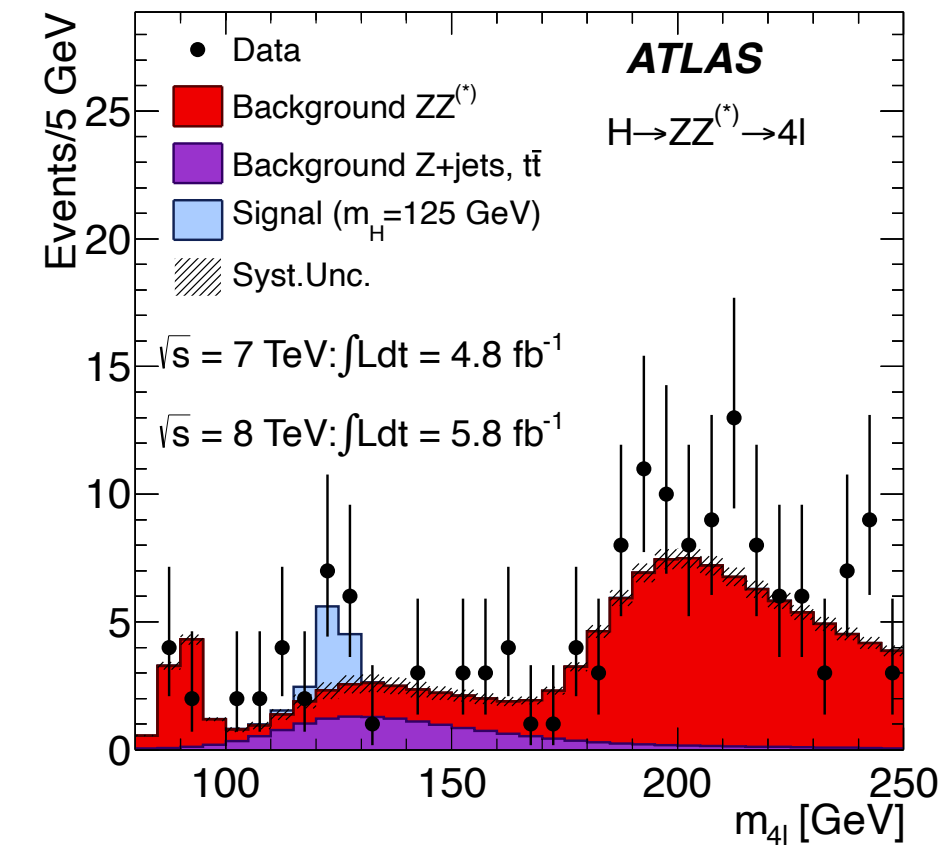
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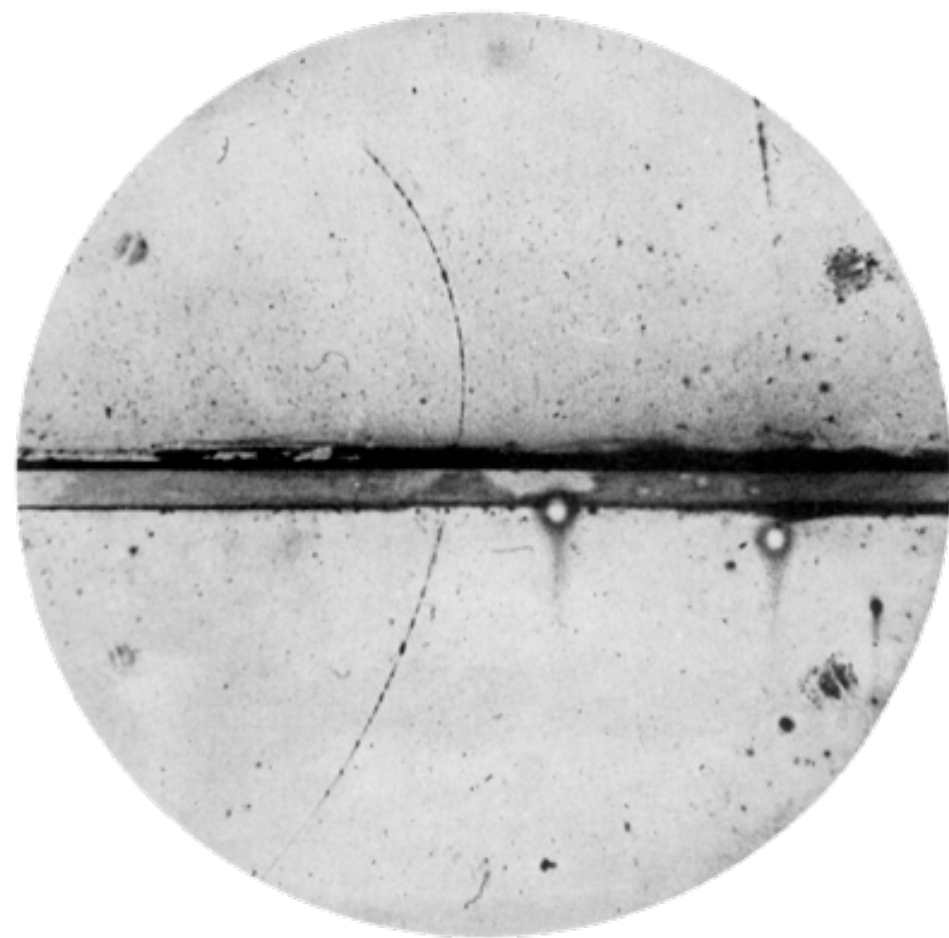
Histogram bins just a bunch of independent counting experiments

Single event

Multiple events:
Cut-and-count

Shape information:
Histogram

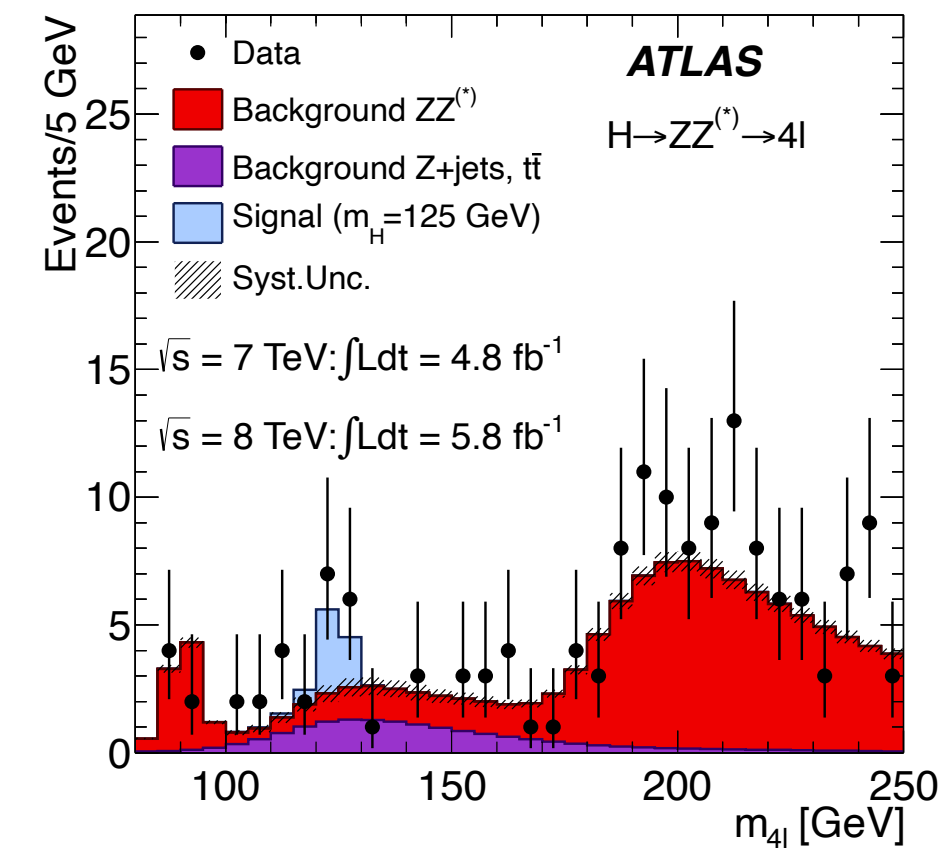
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$$p(\mathcal{D} | \mu) = \prod p_{bin}(n_{obs} | n_{exp})$$

Histogram bins just a bunch of independent counting experiments

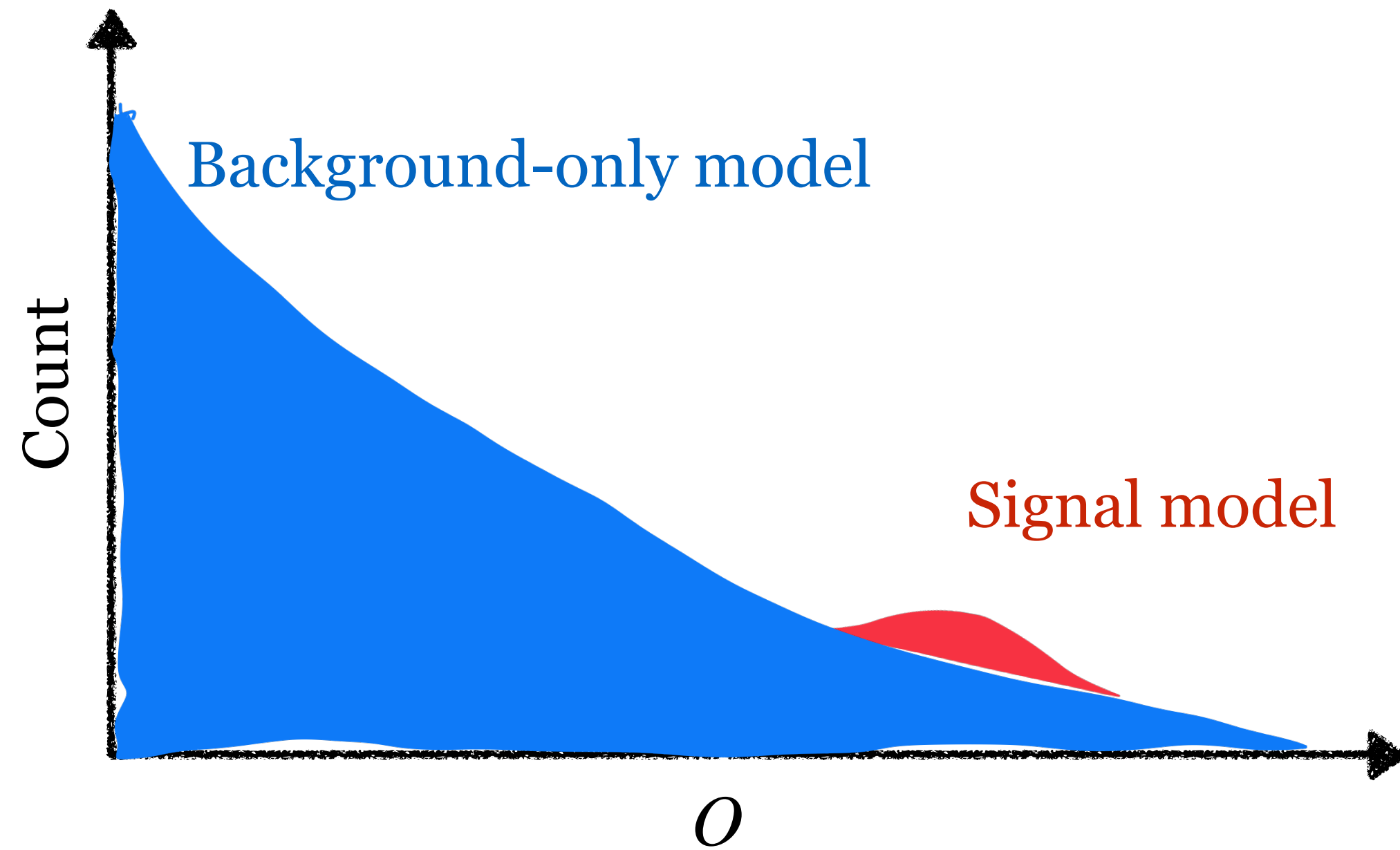
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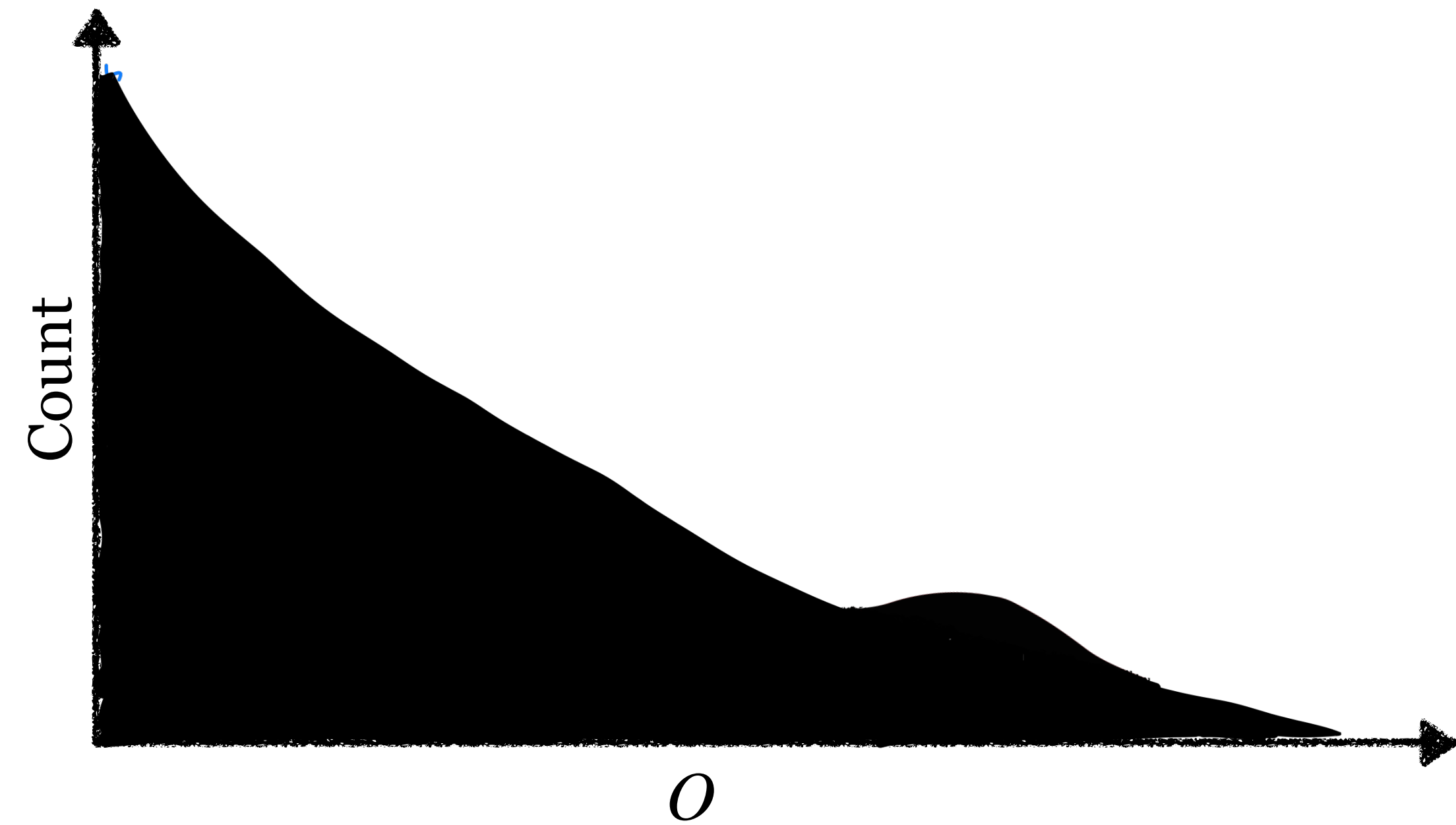
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Probability density estimation, with a histogram

Theory Predictions



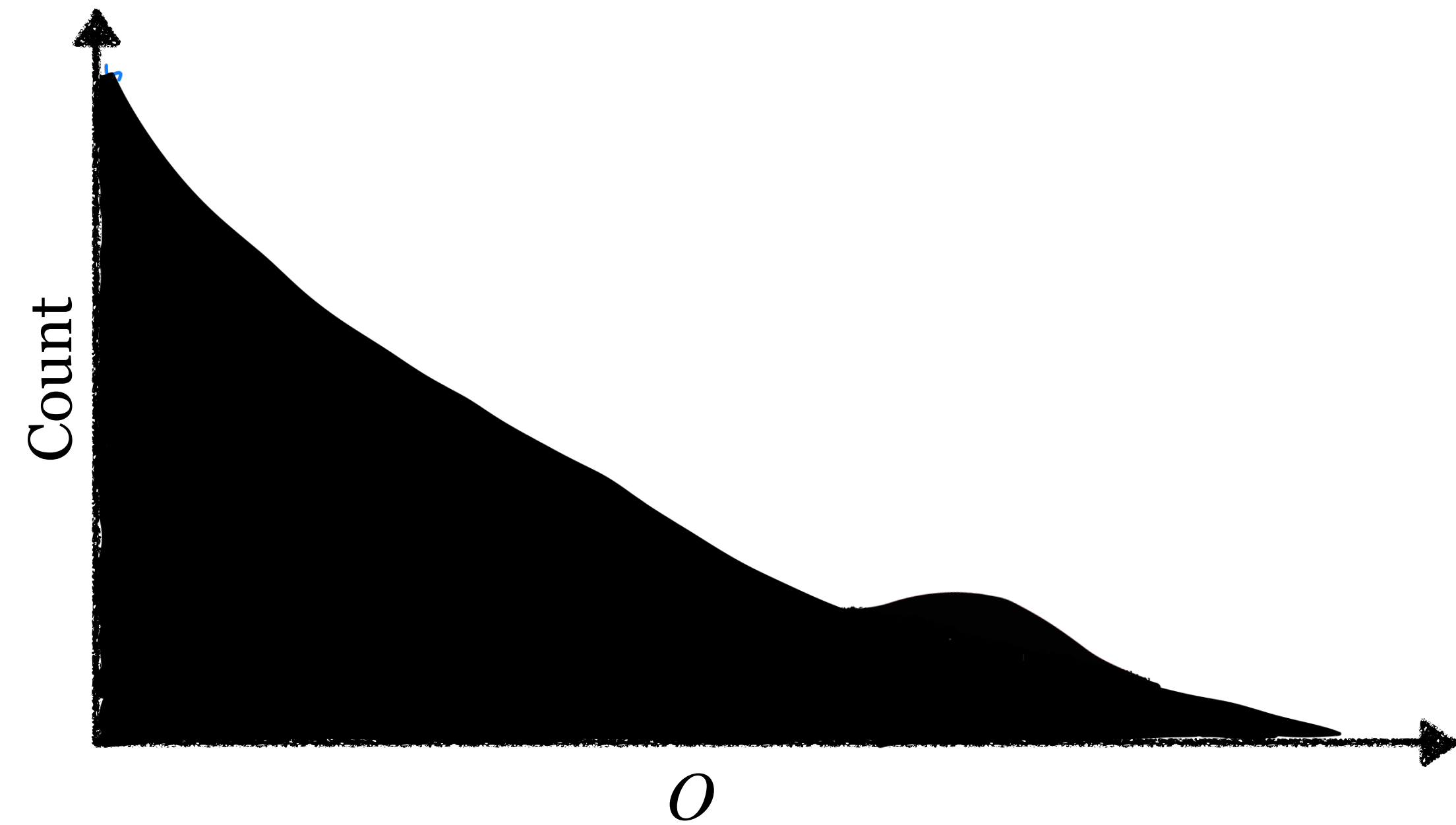
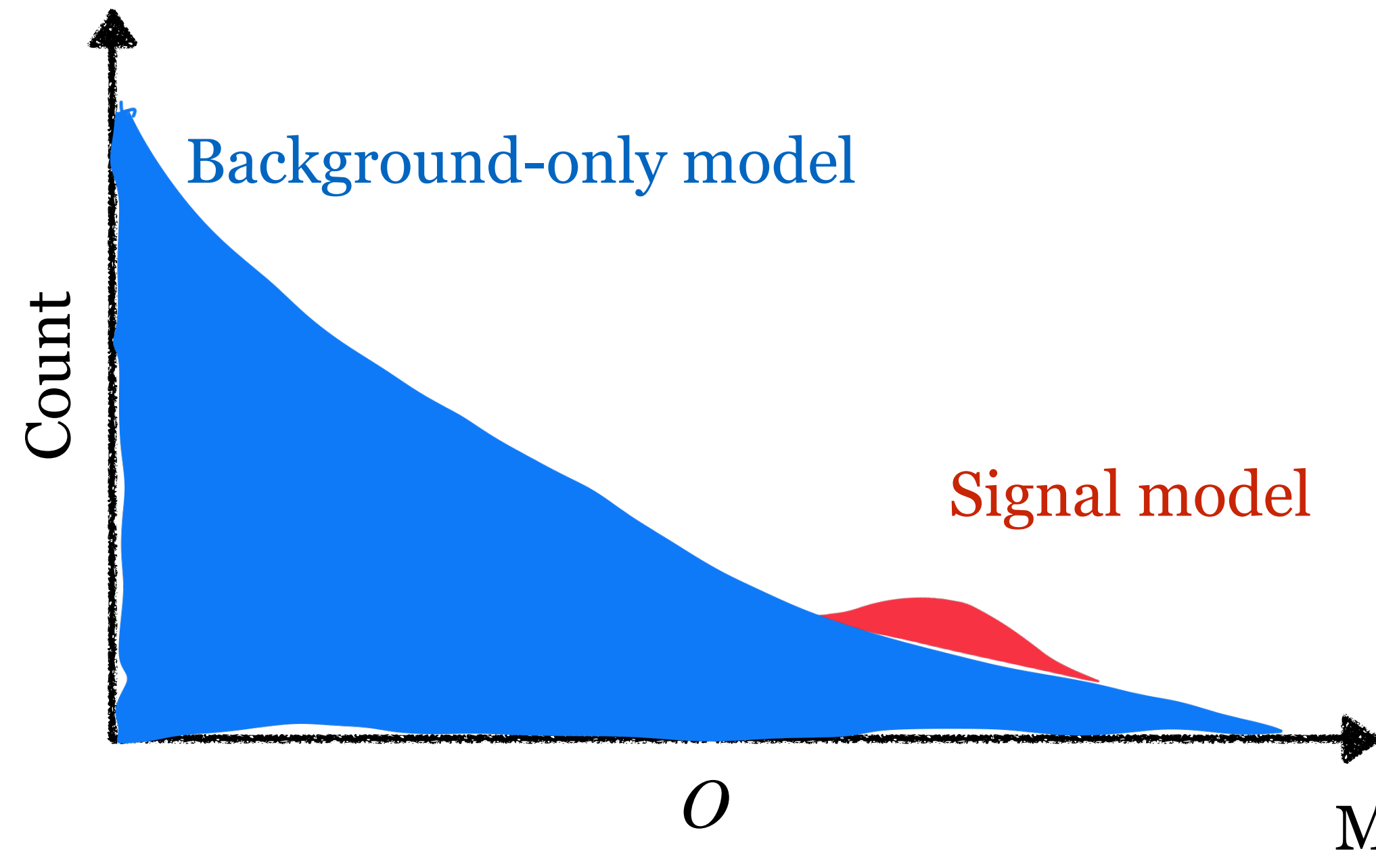
Data



Probability density estimation, with a histogram

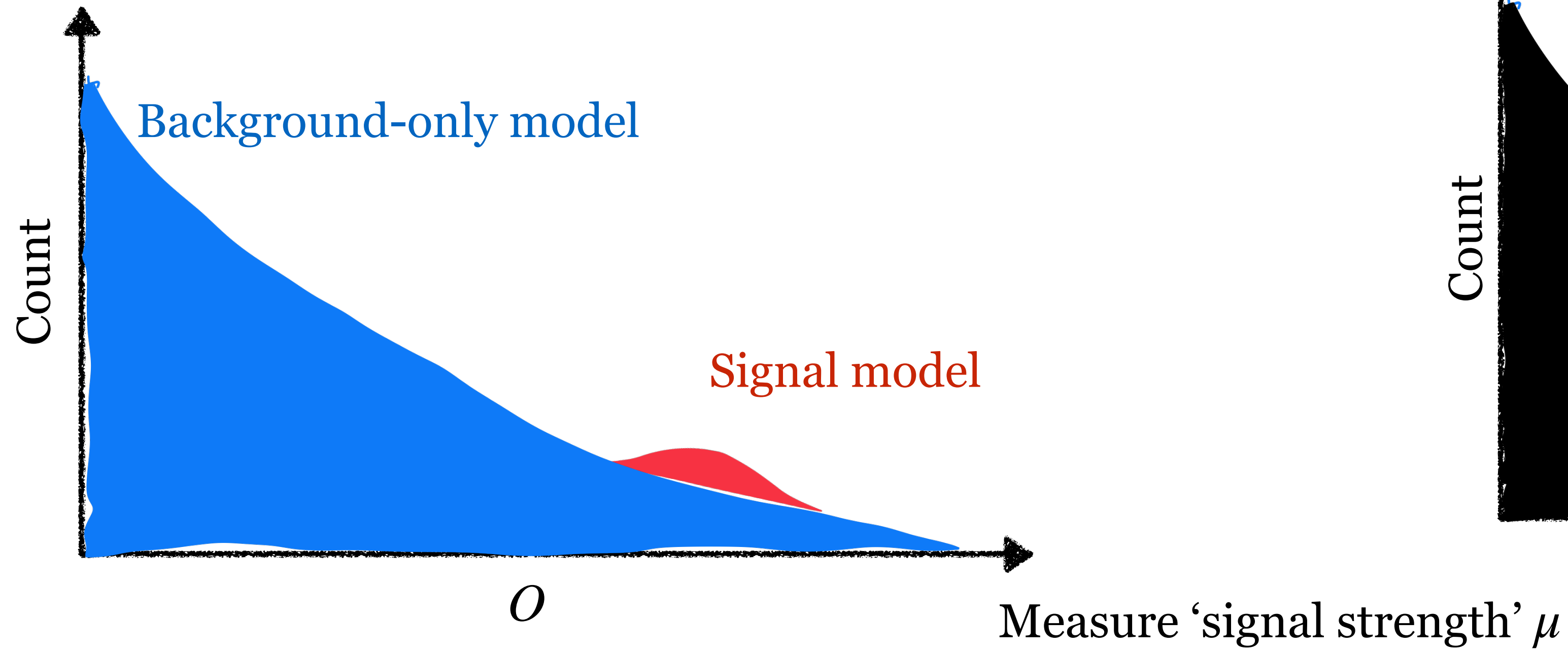
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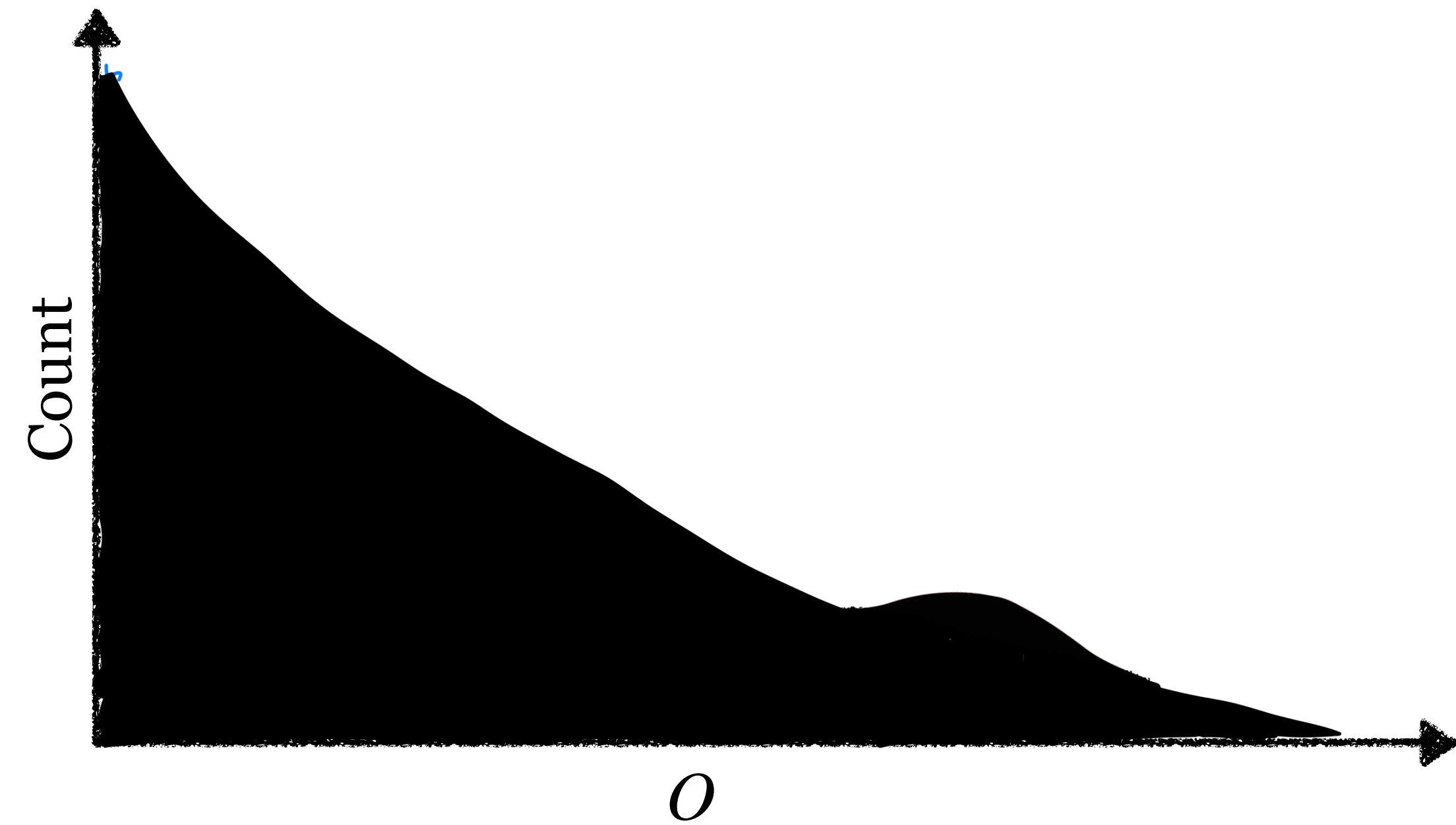


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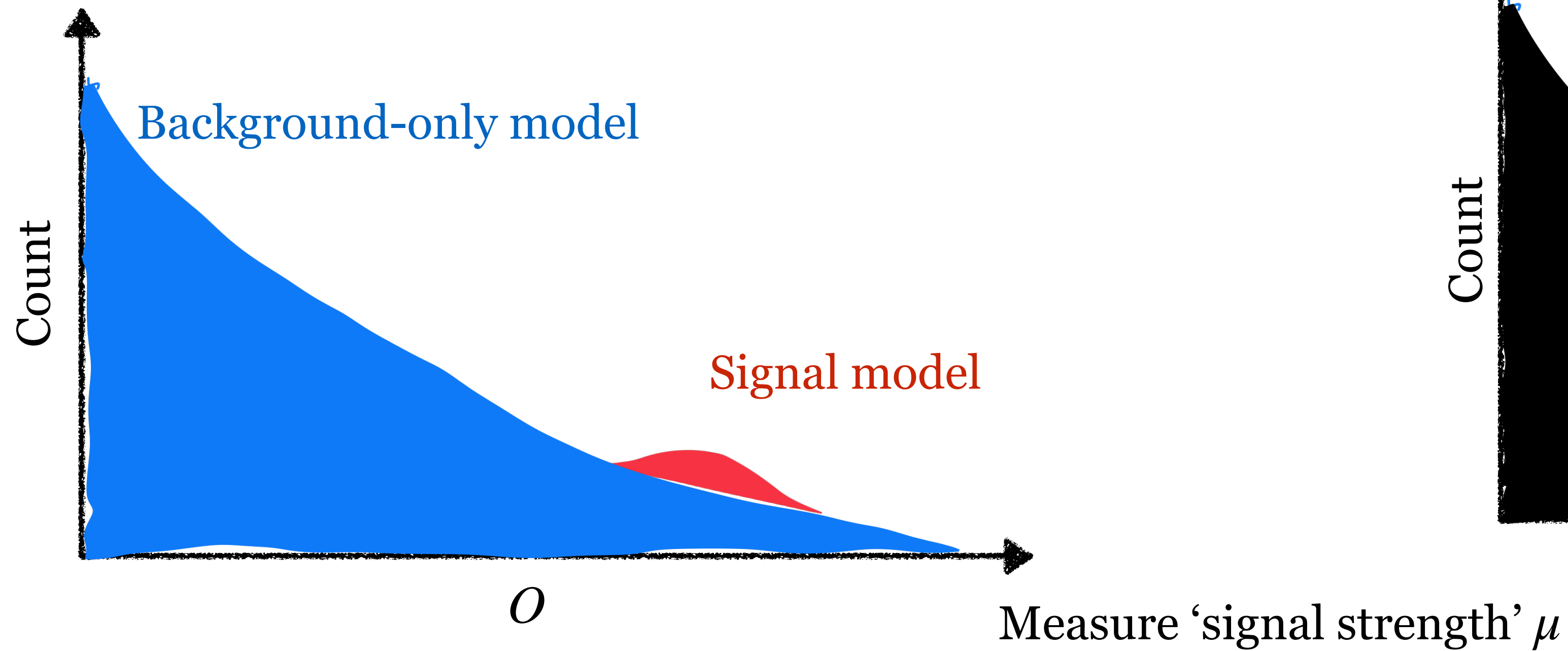


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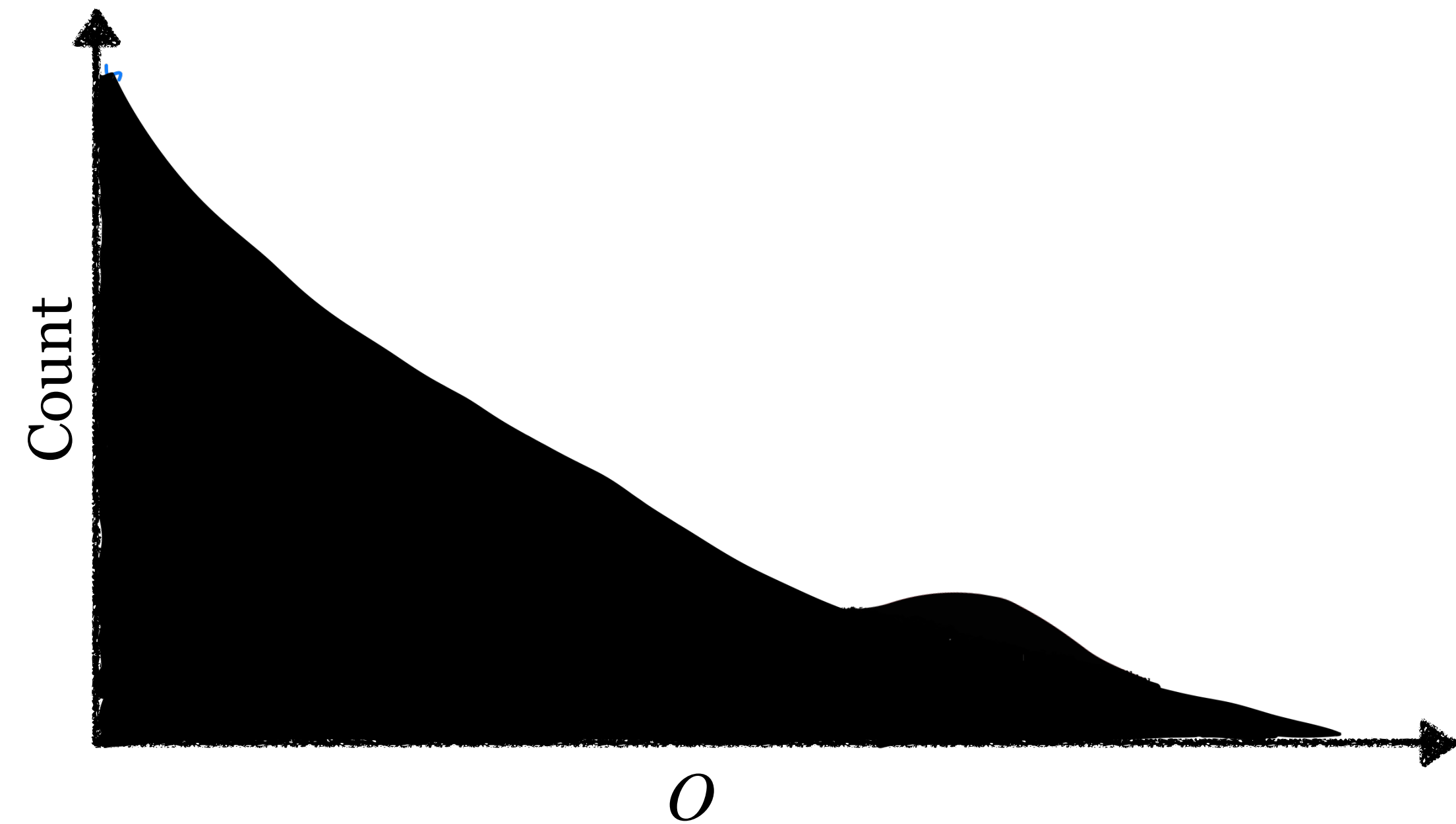


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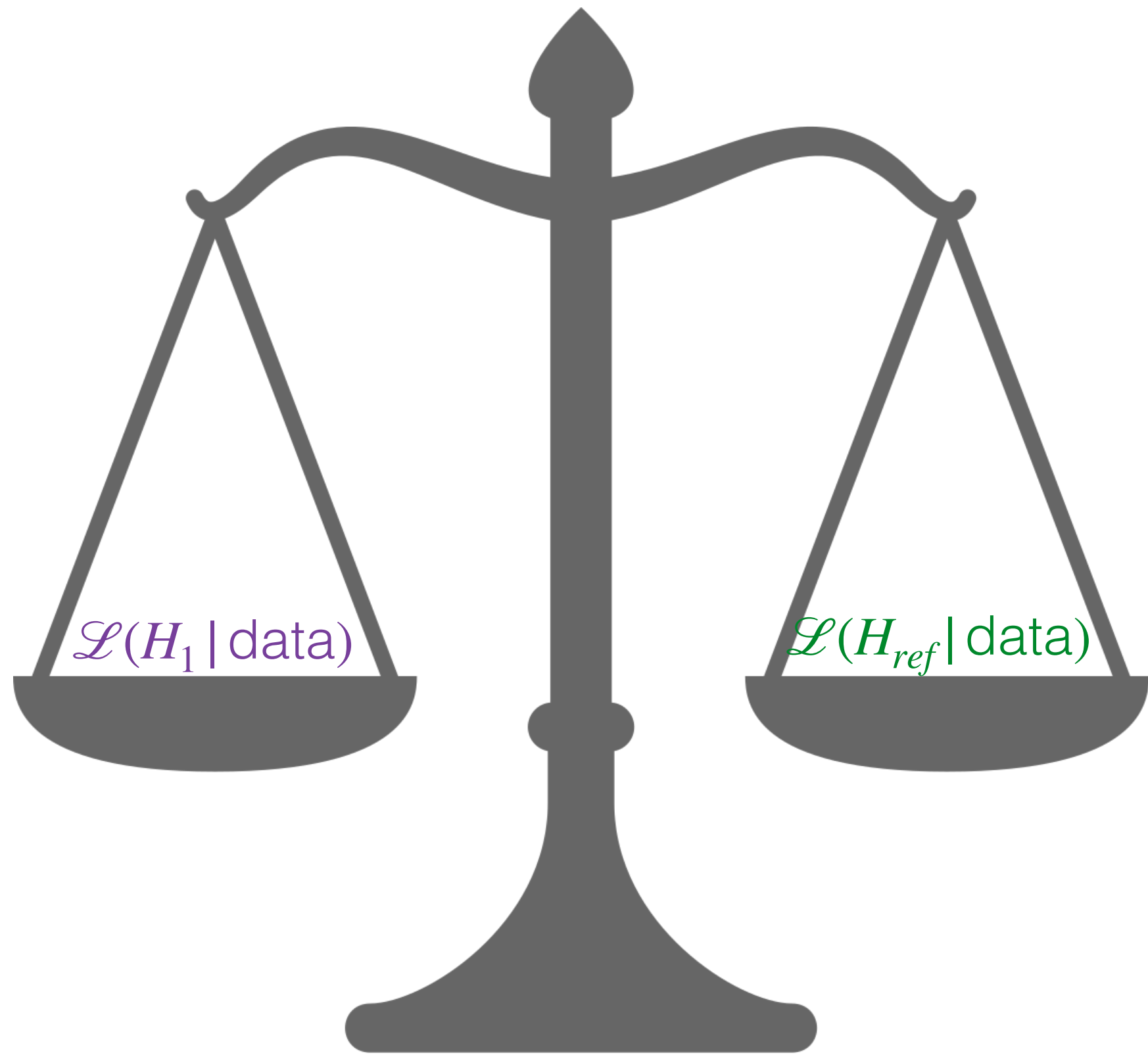


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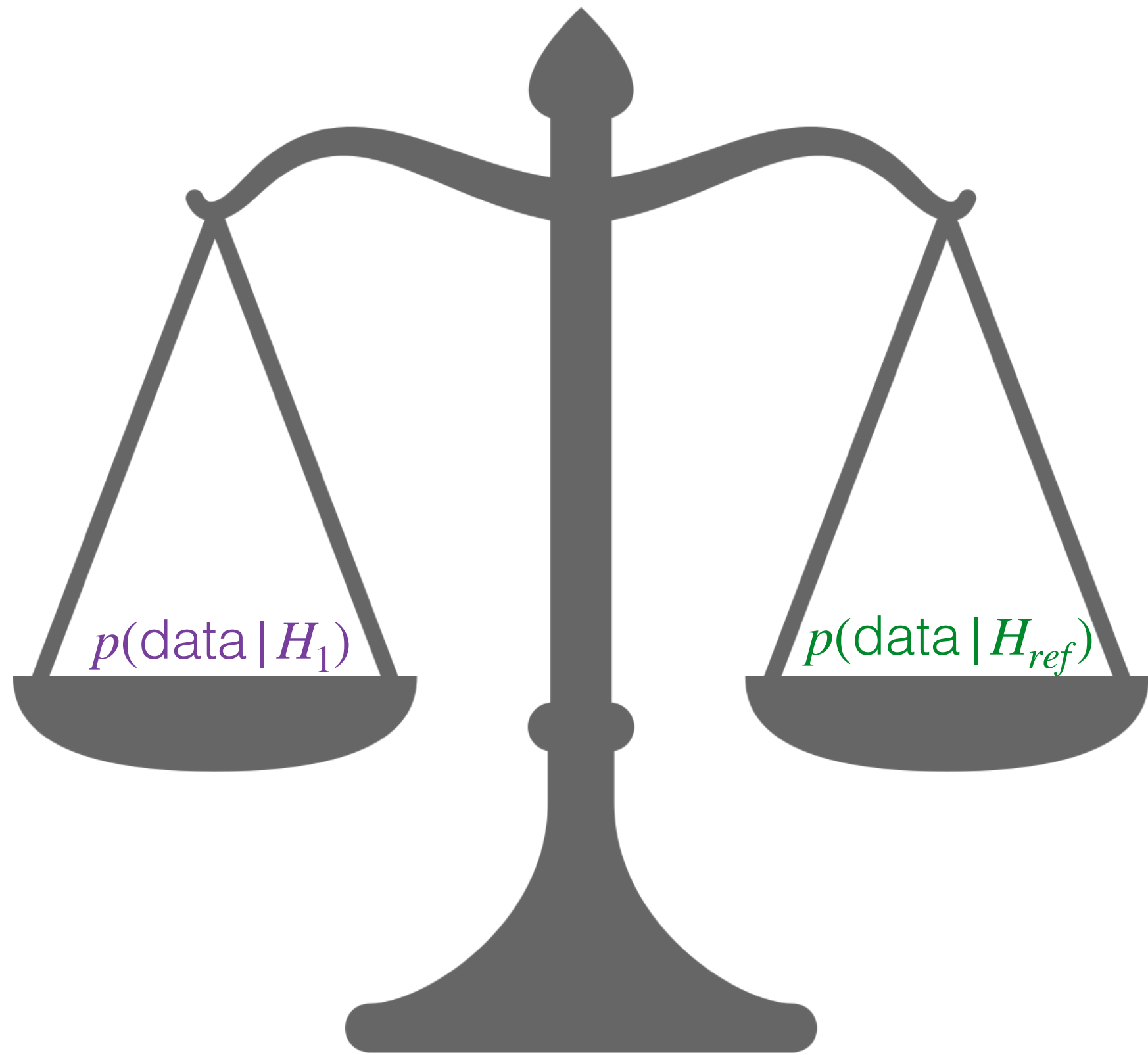
With histograms we can ask “Given the data, what is the likelihood of $\mu = 1$ hypothesis vs $\mu = 2$ hypothesis?”

(Frequentist) Hypothesis tests



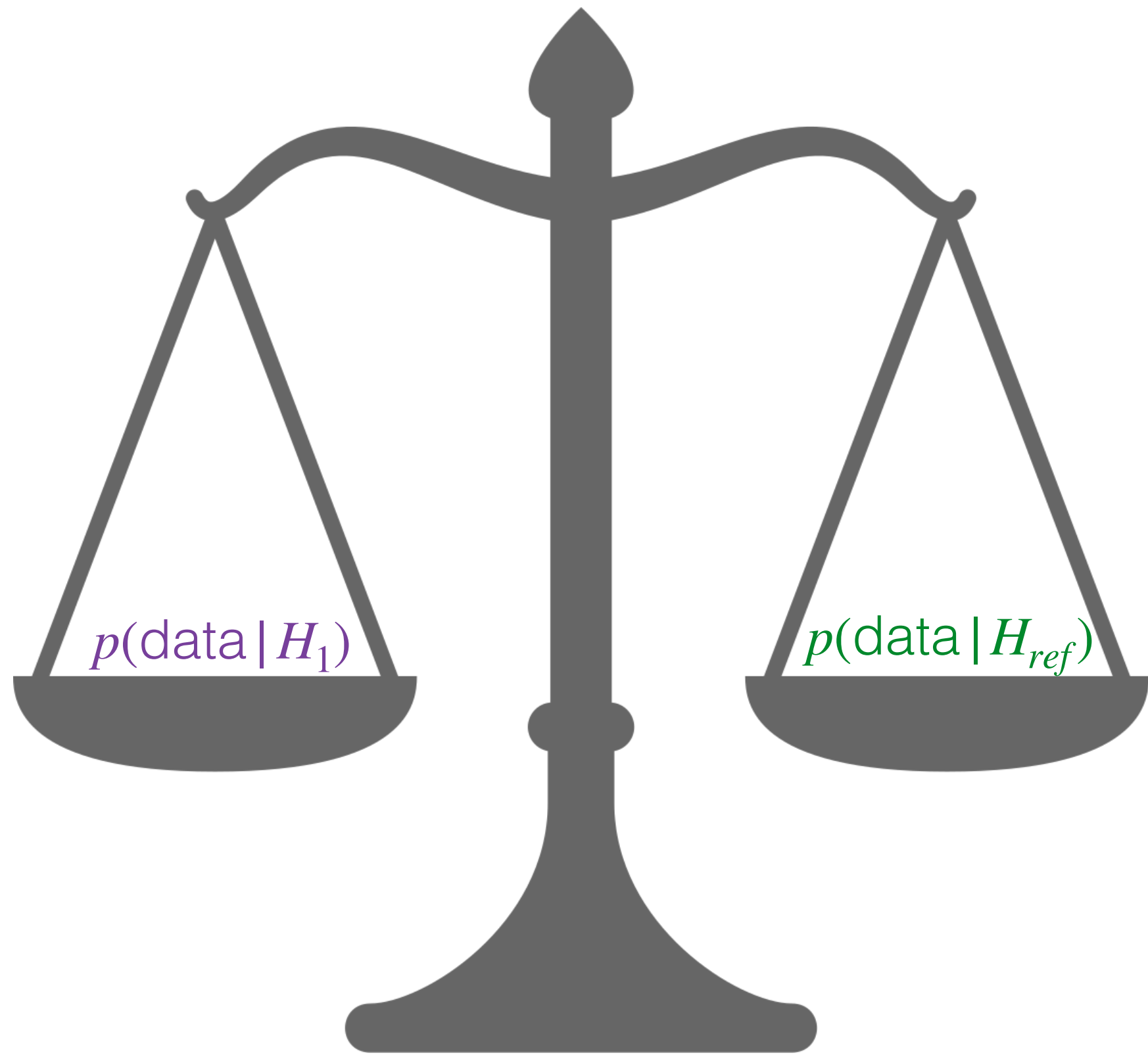
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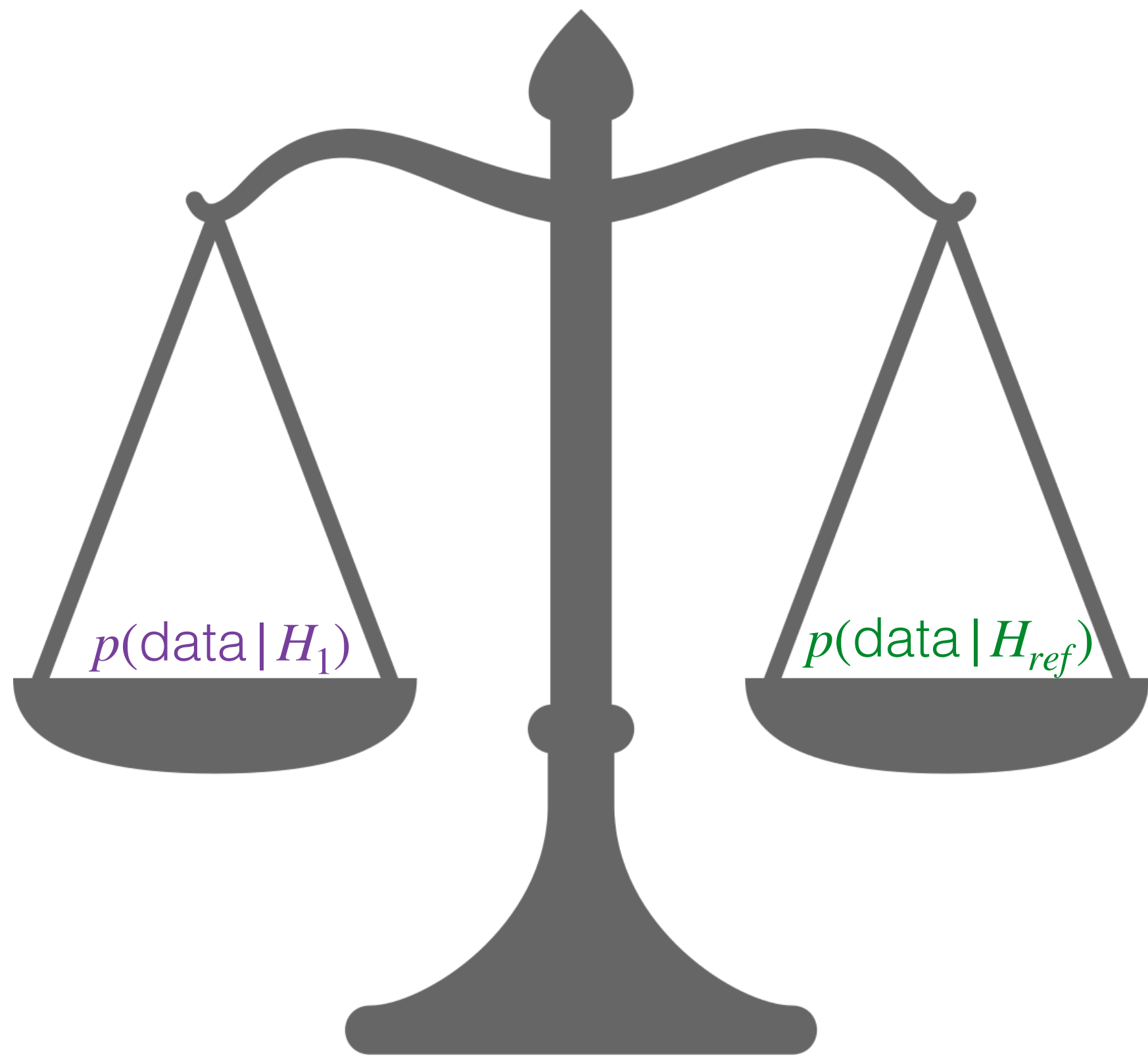
When comparing 2 hypotheses, LR guaranteed to be optimal test by Neyman-Person lemma

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$$H_1 = \{\mu_1\}$$

$$H_1^1 = \{\mu_1 \alpha_1\}$$



μ : Parameter of interest, you want to measure this

α : Nuisance parameters, you need to measure this to avoid biases. Systematic uncertainties like jet energy scale

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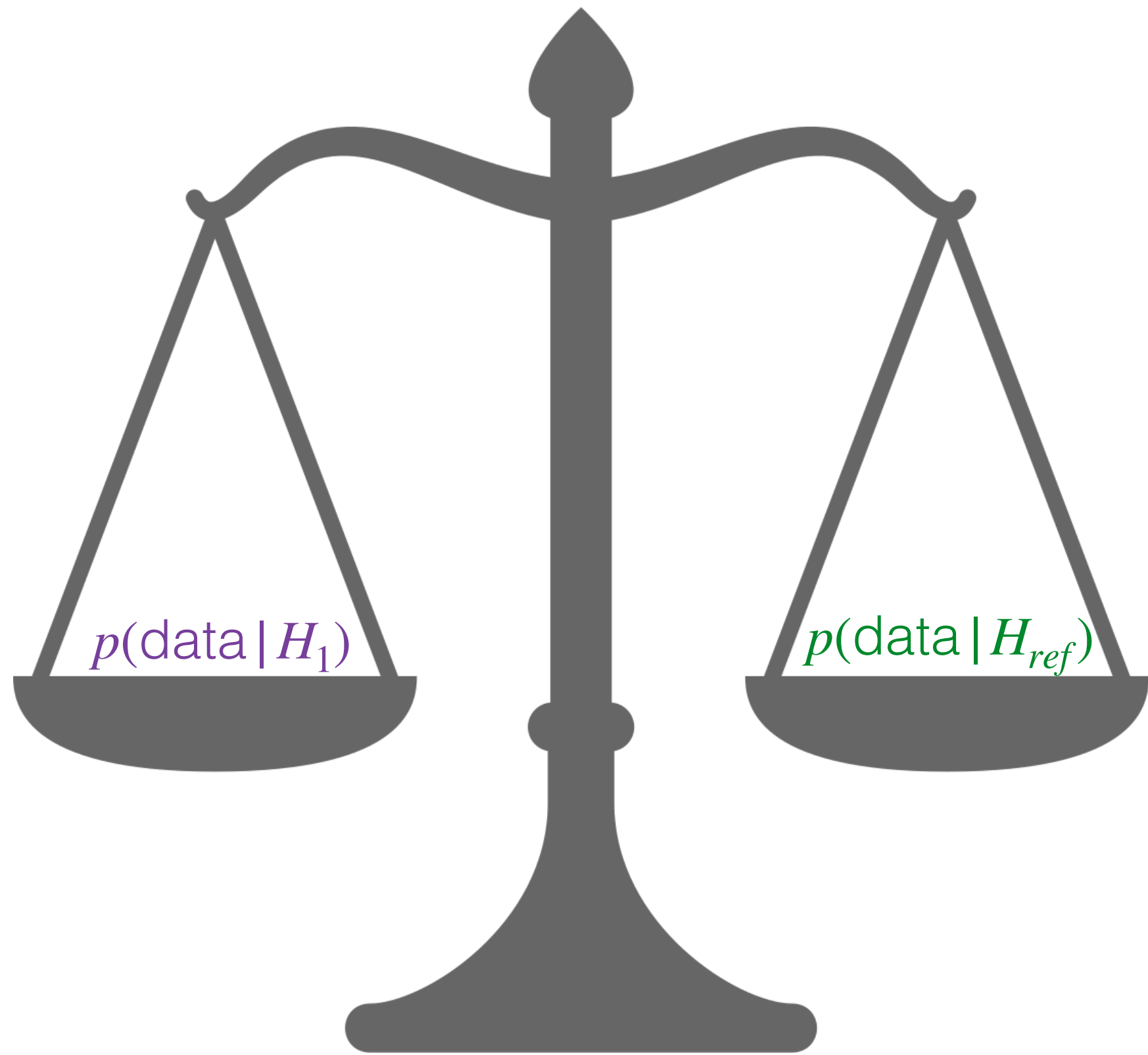
\mathcal{D} now is a 1D summary of the high-dimensional data

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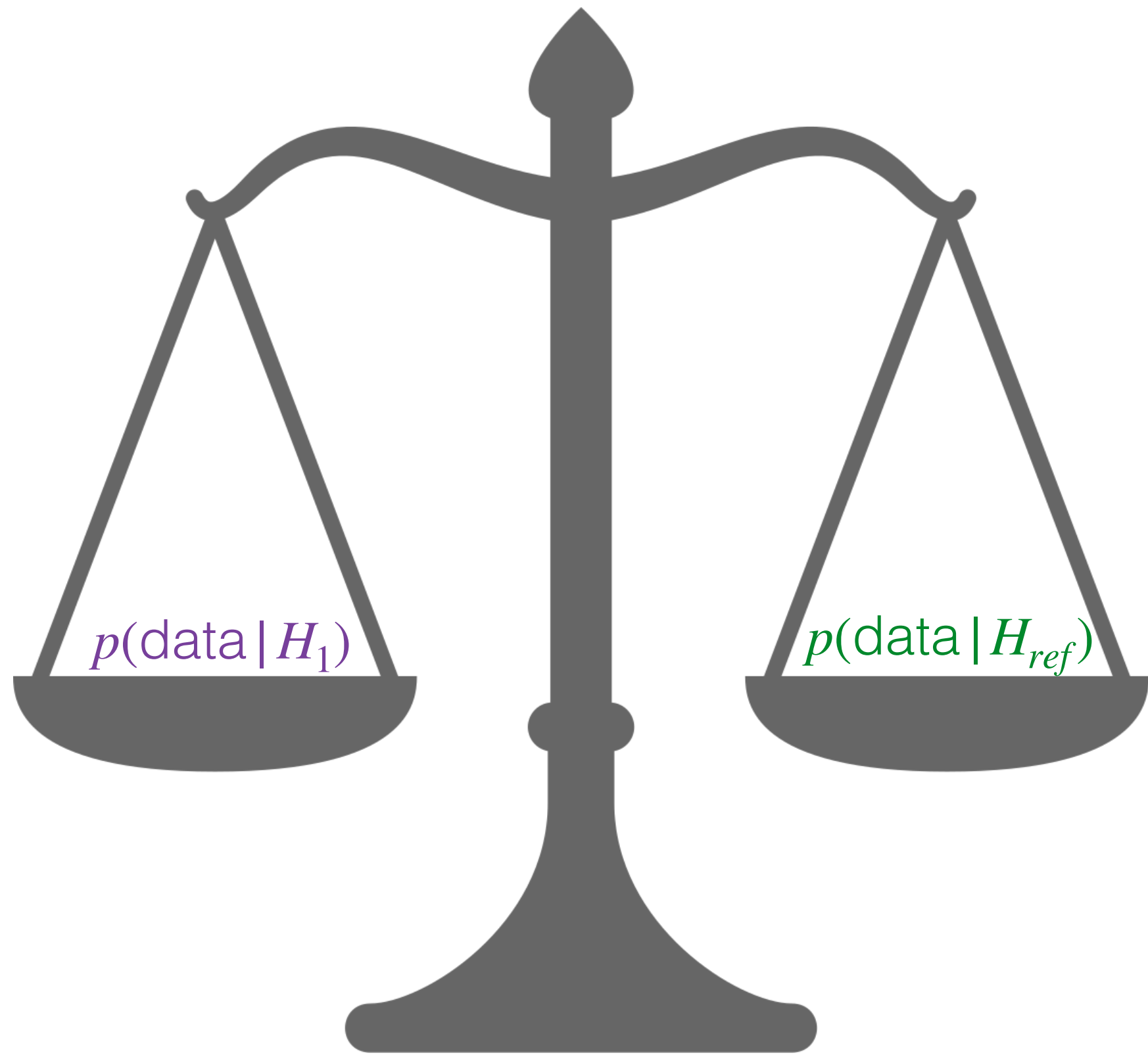
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$$\text{Profile likelihood ratio:} \quad \frac{p(\mathcal{D} | \mu, \hat{\alpha})}{p(\mathcal{D} | \hat{\mu}, \hat{\alpha})}$$



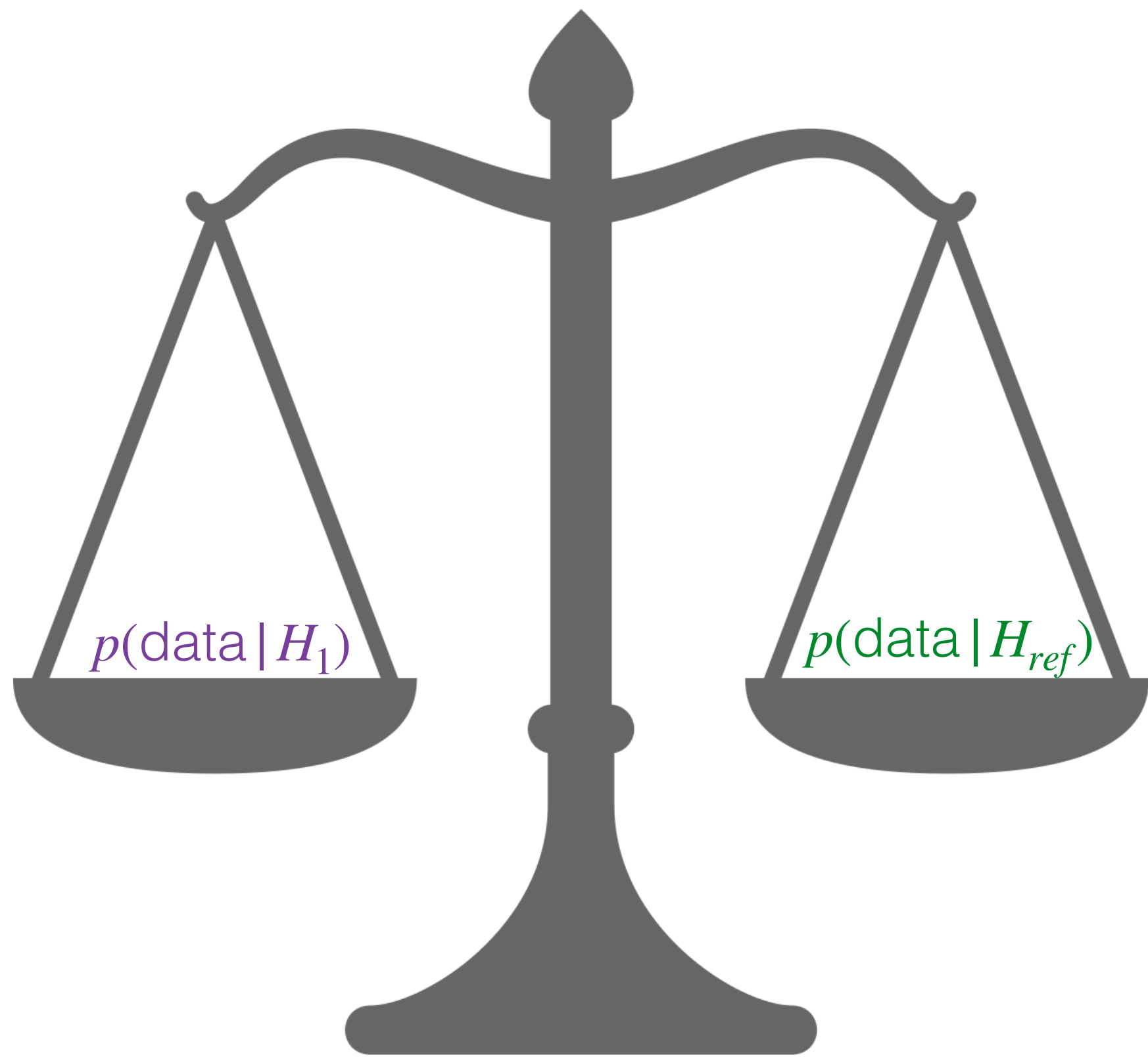
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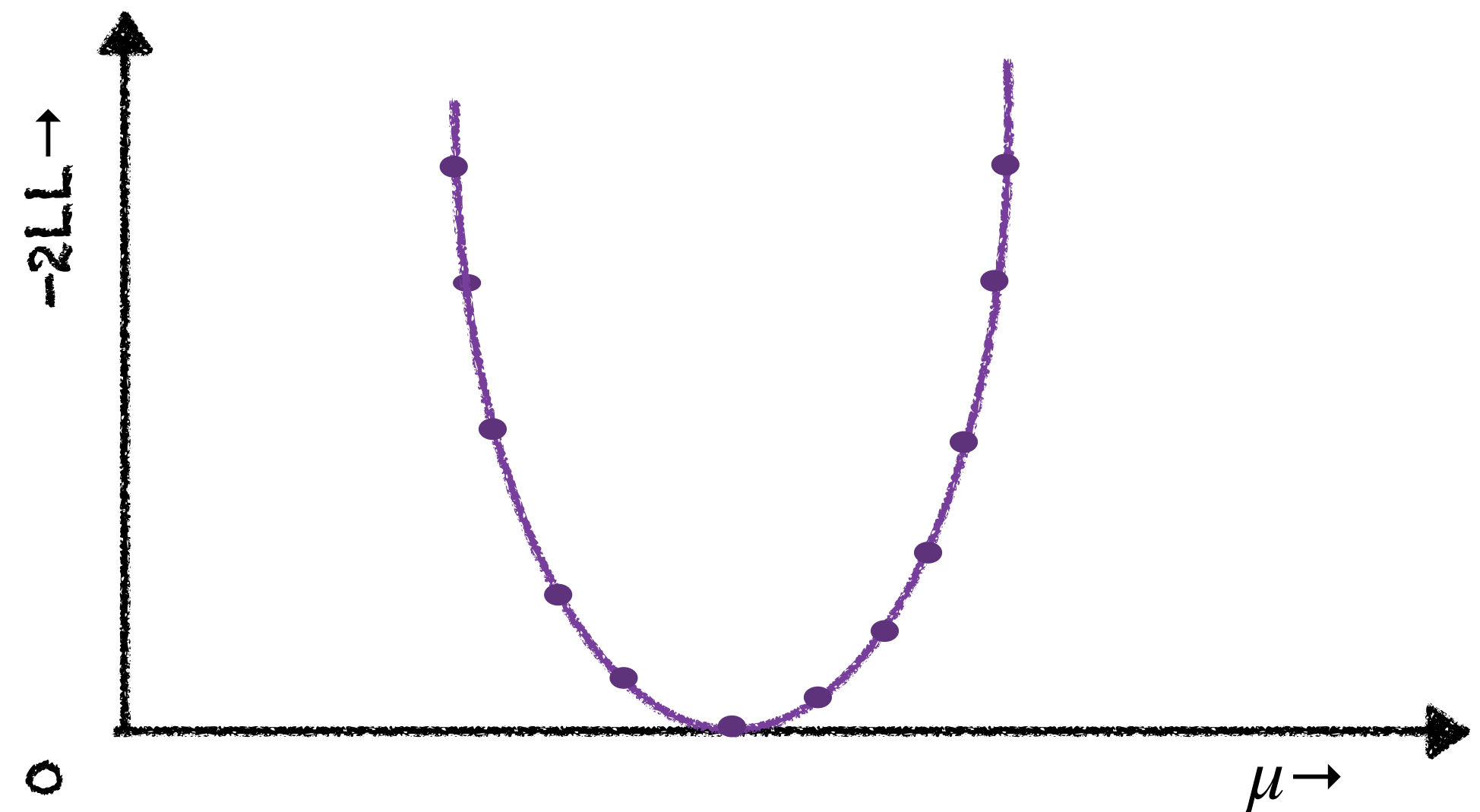
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$$-2 \log \frac{p(\mathcal{D} | \mu, \hat{\alpha})}{p(\mathcal{D} | \hat{\mu}, \hat{\alpha})}$$

Invert the test to get confidence intervals

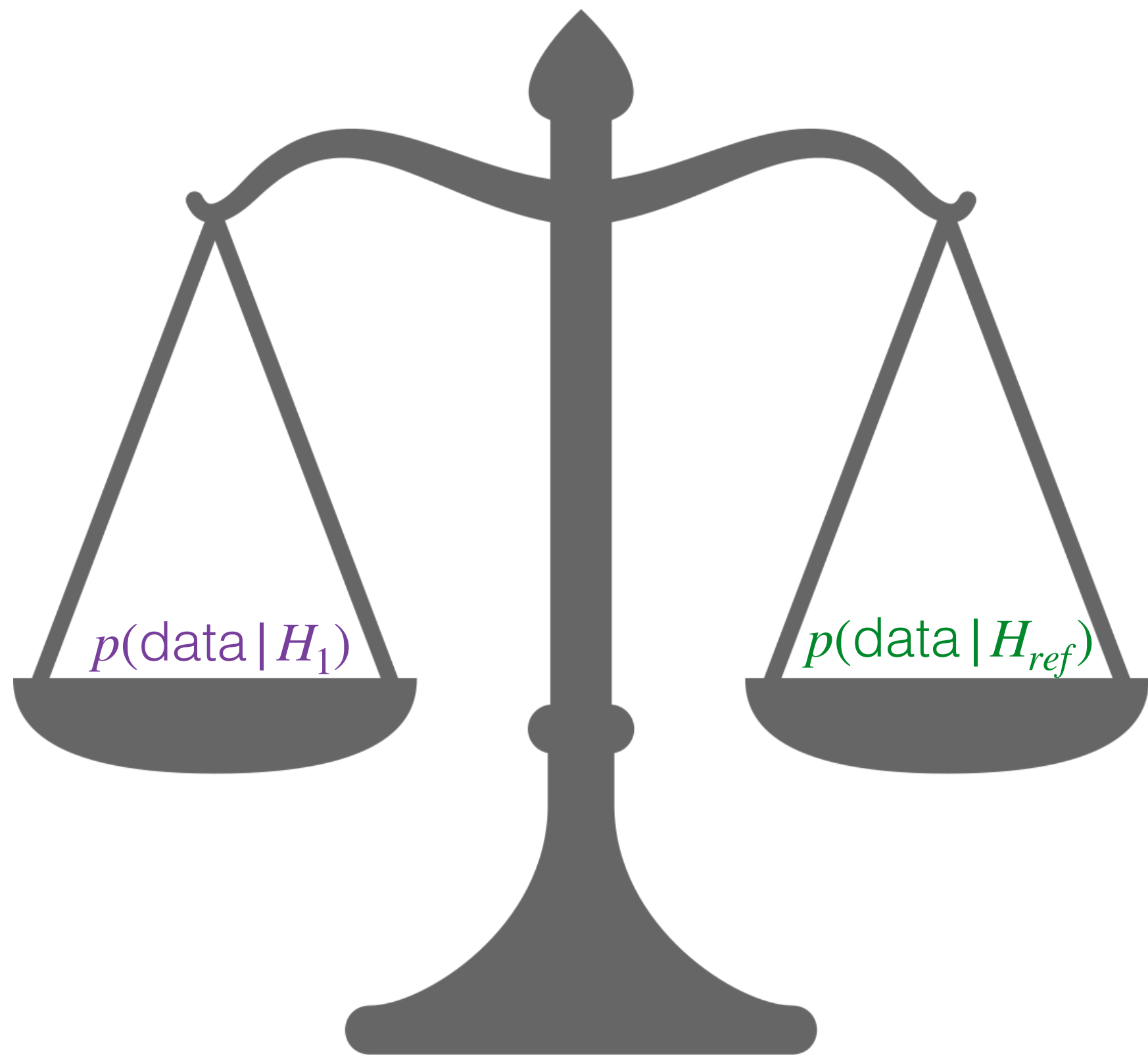


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Parameter estimation (infinite hypotheses)

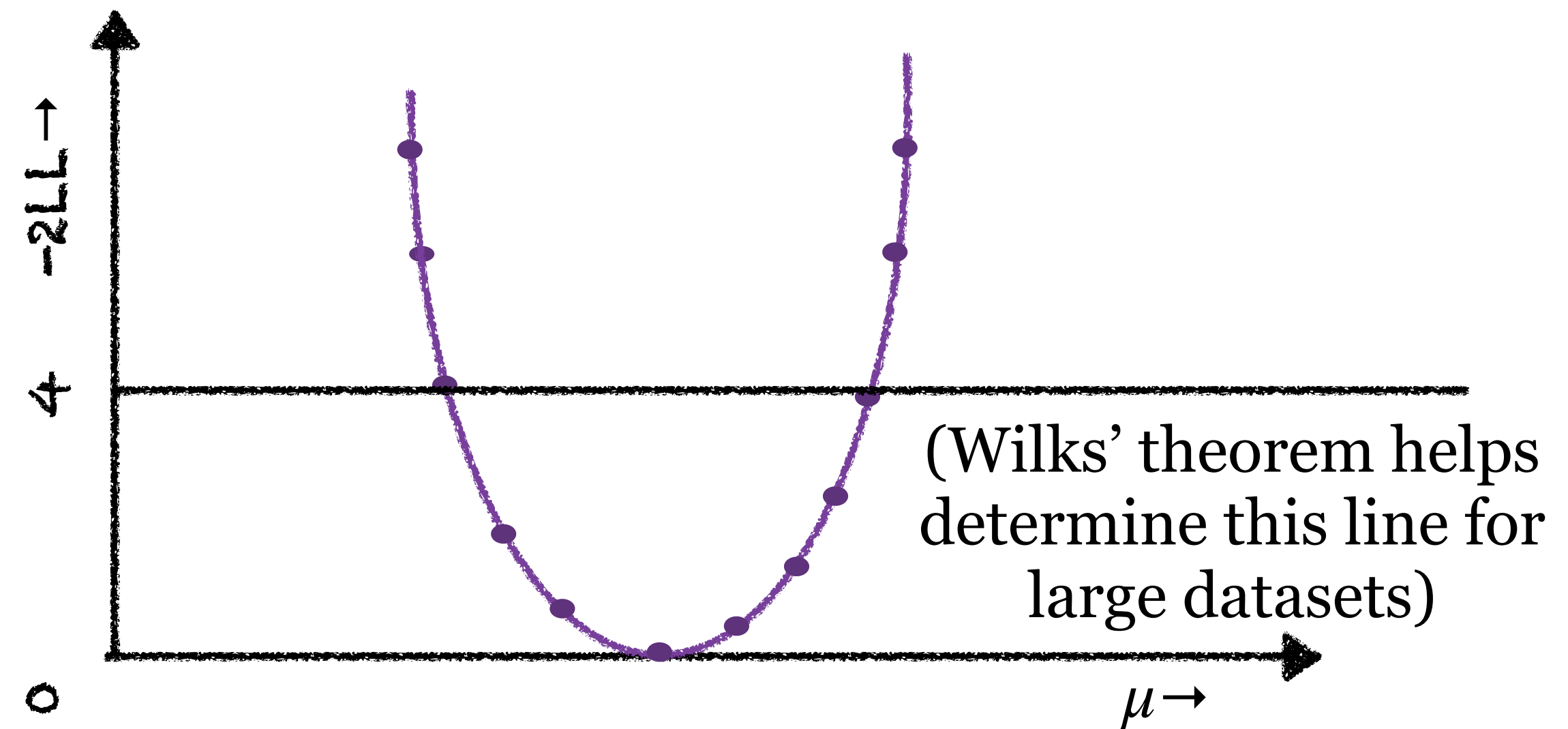
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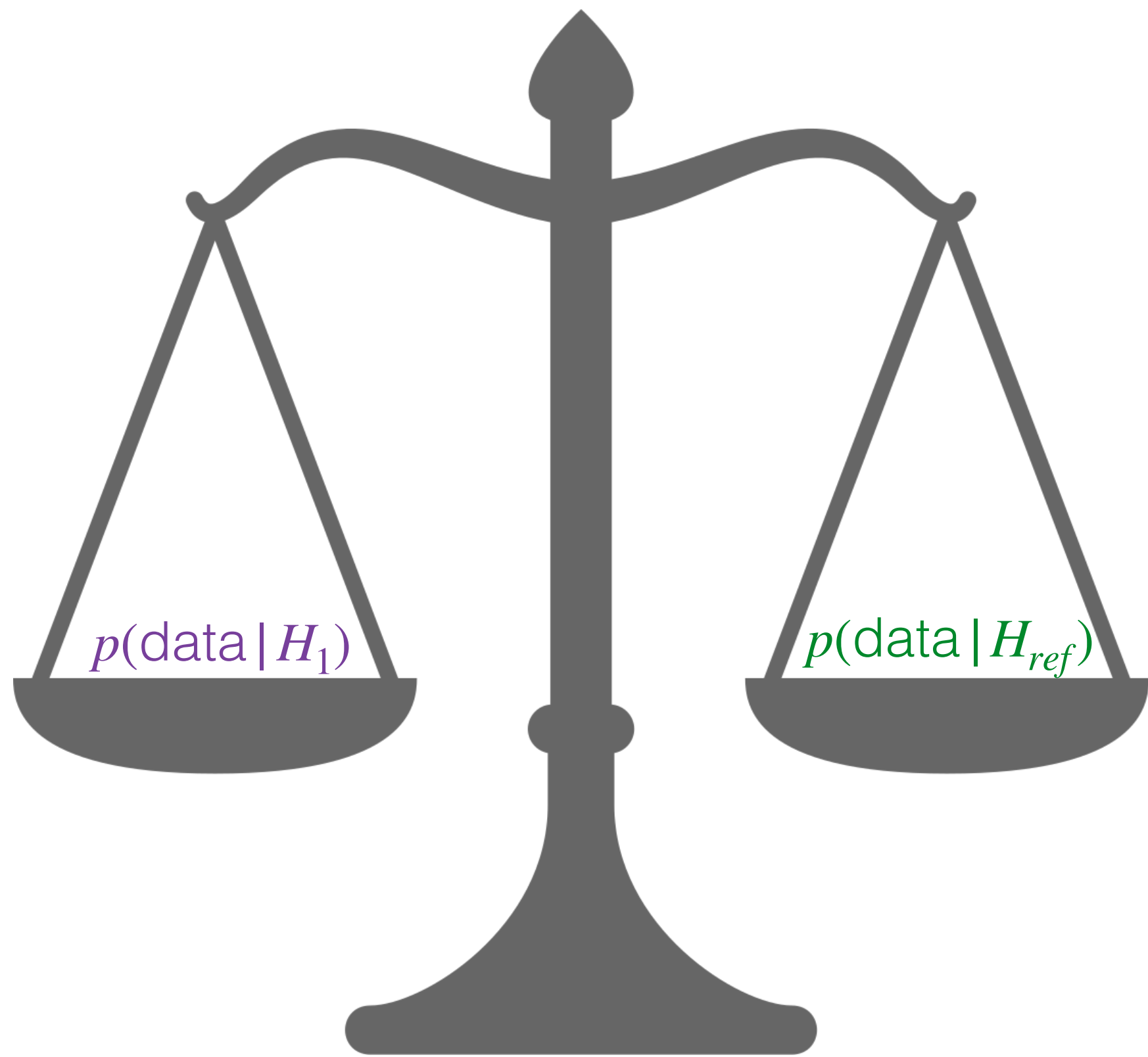


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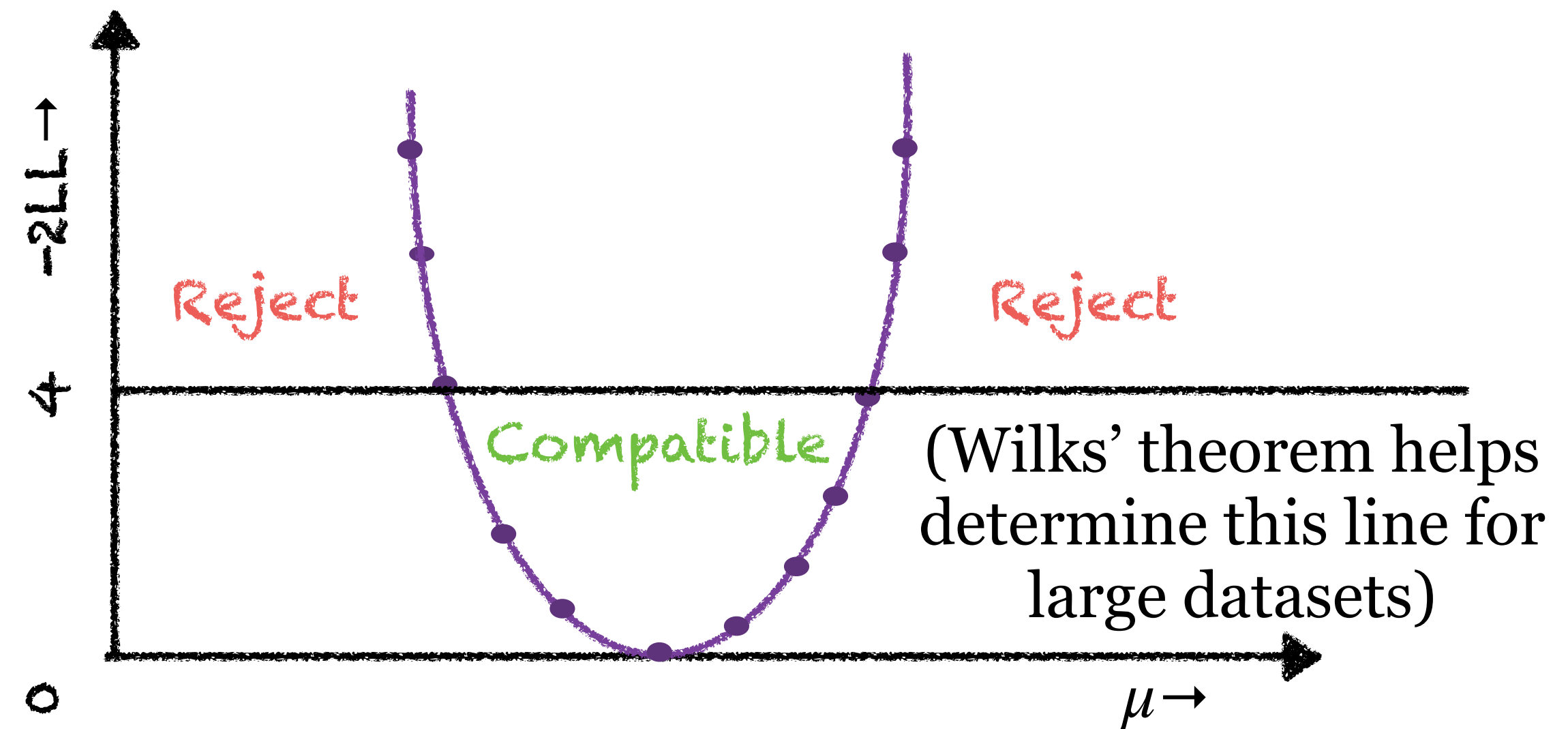
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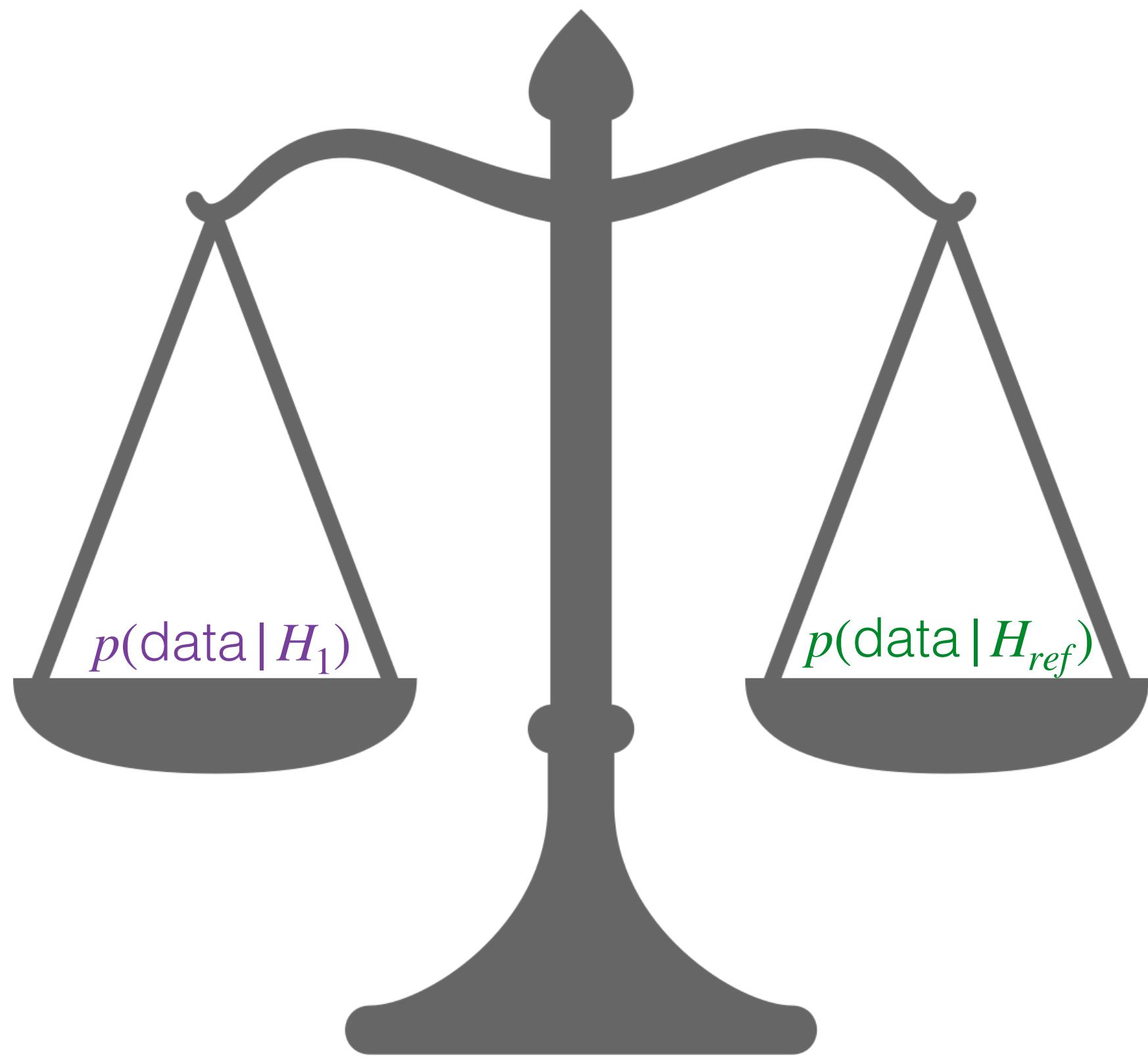


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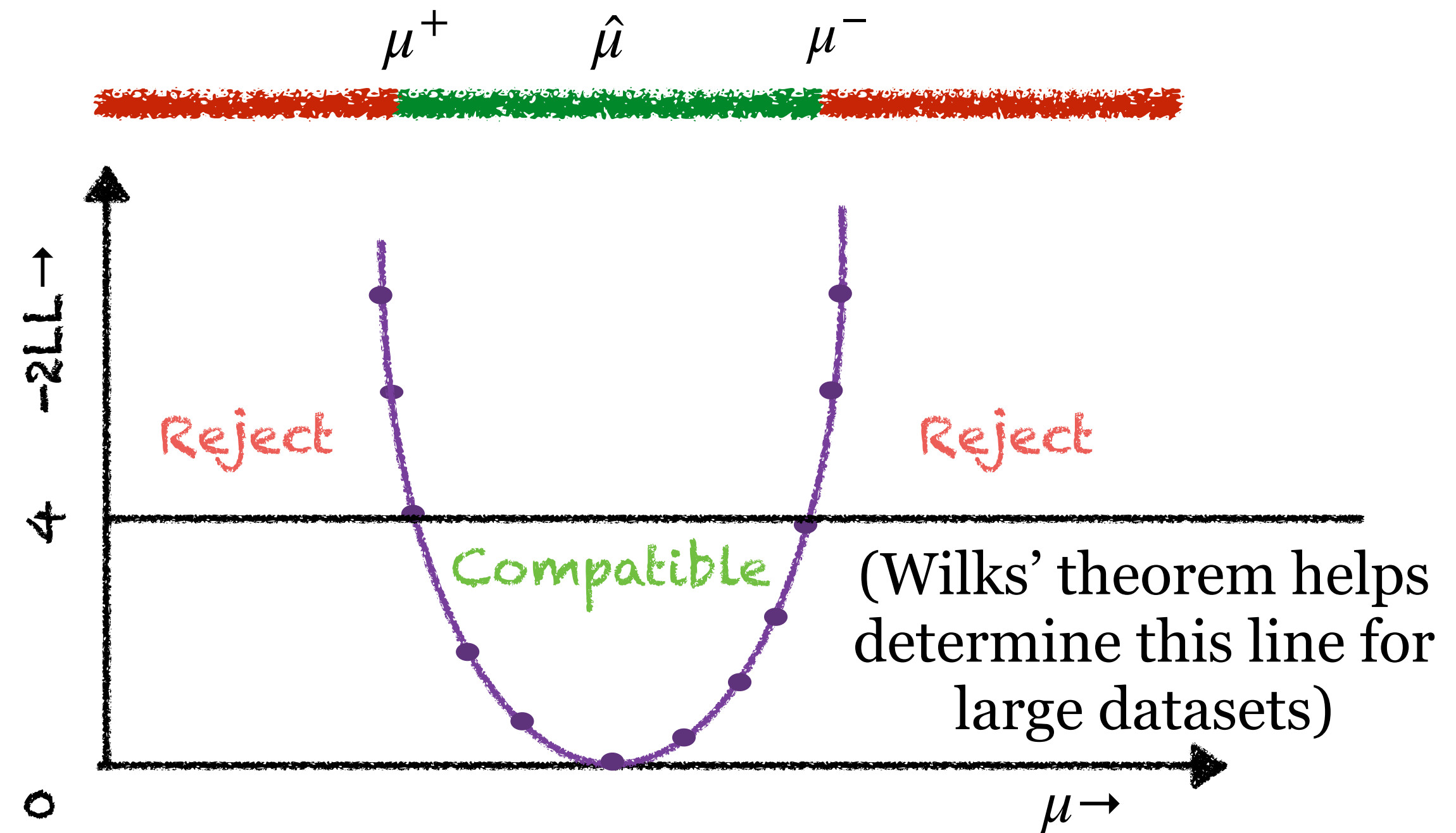
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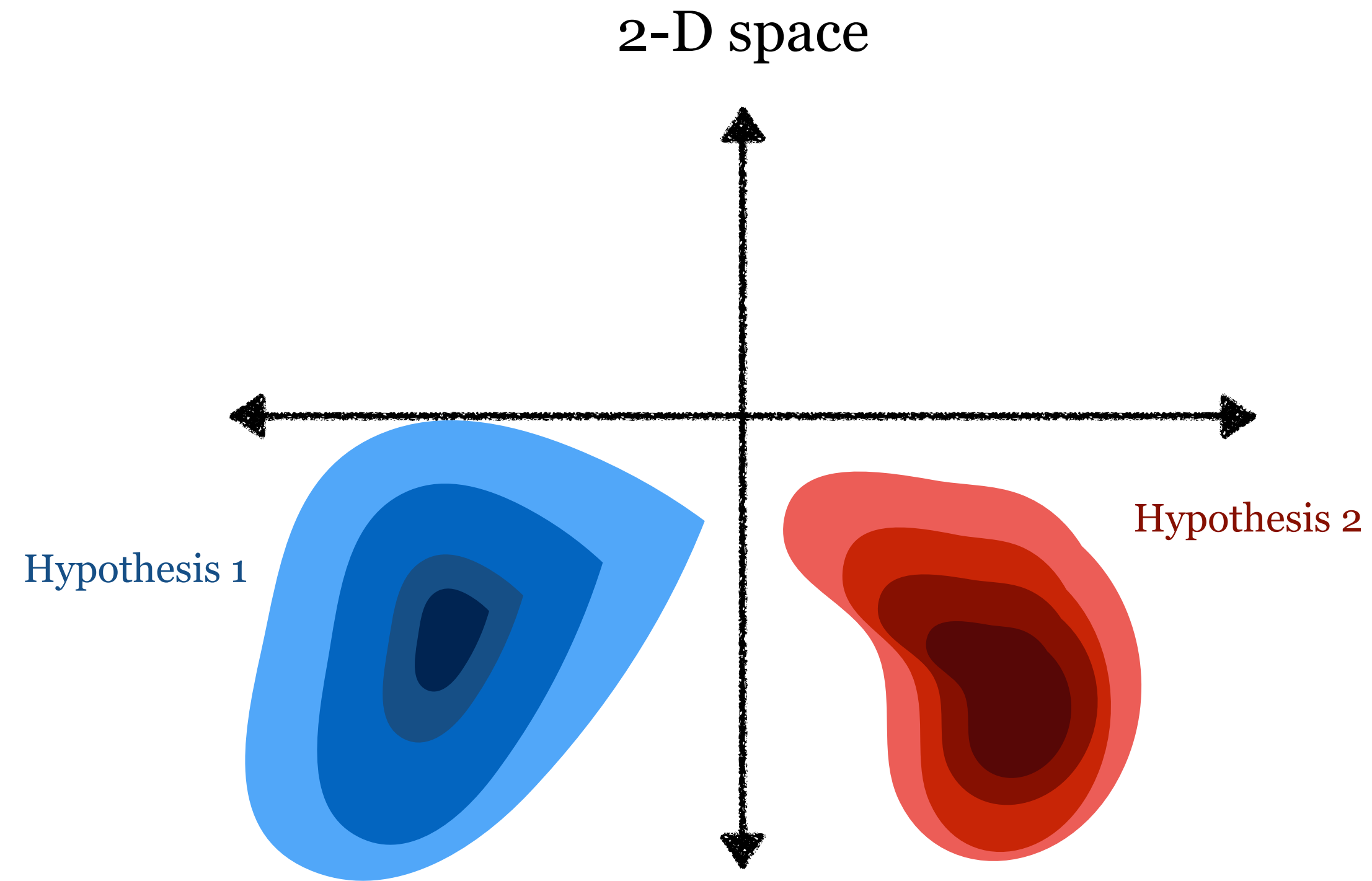
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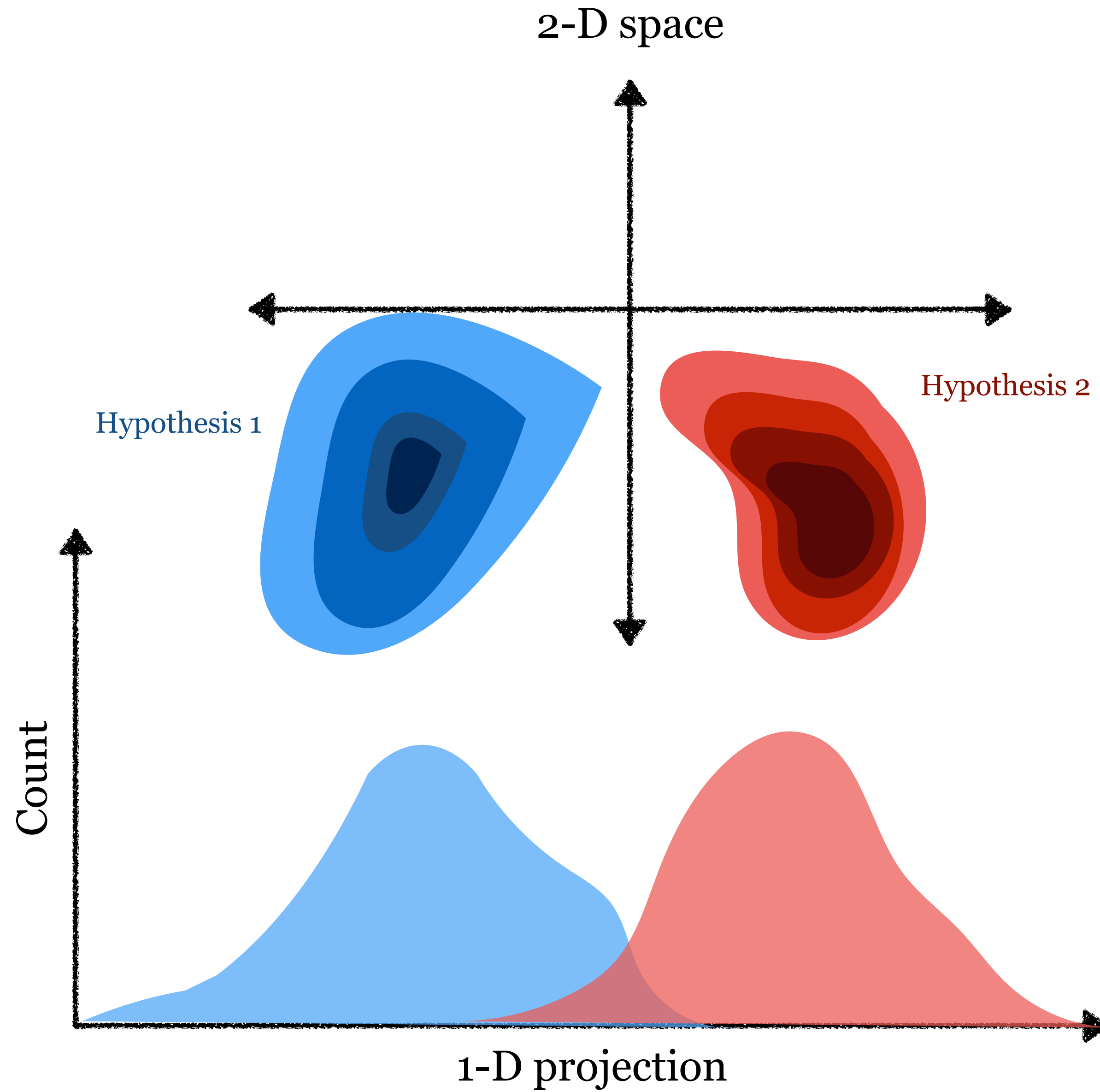
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Can we always find a sufficient 1-D summary?

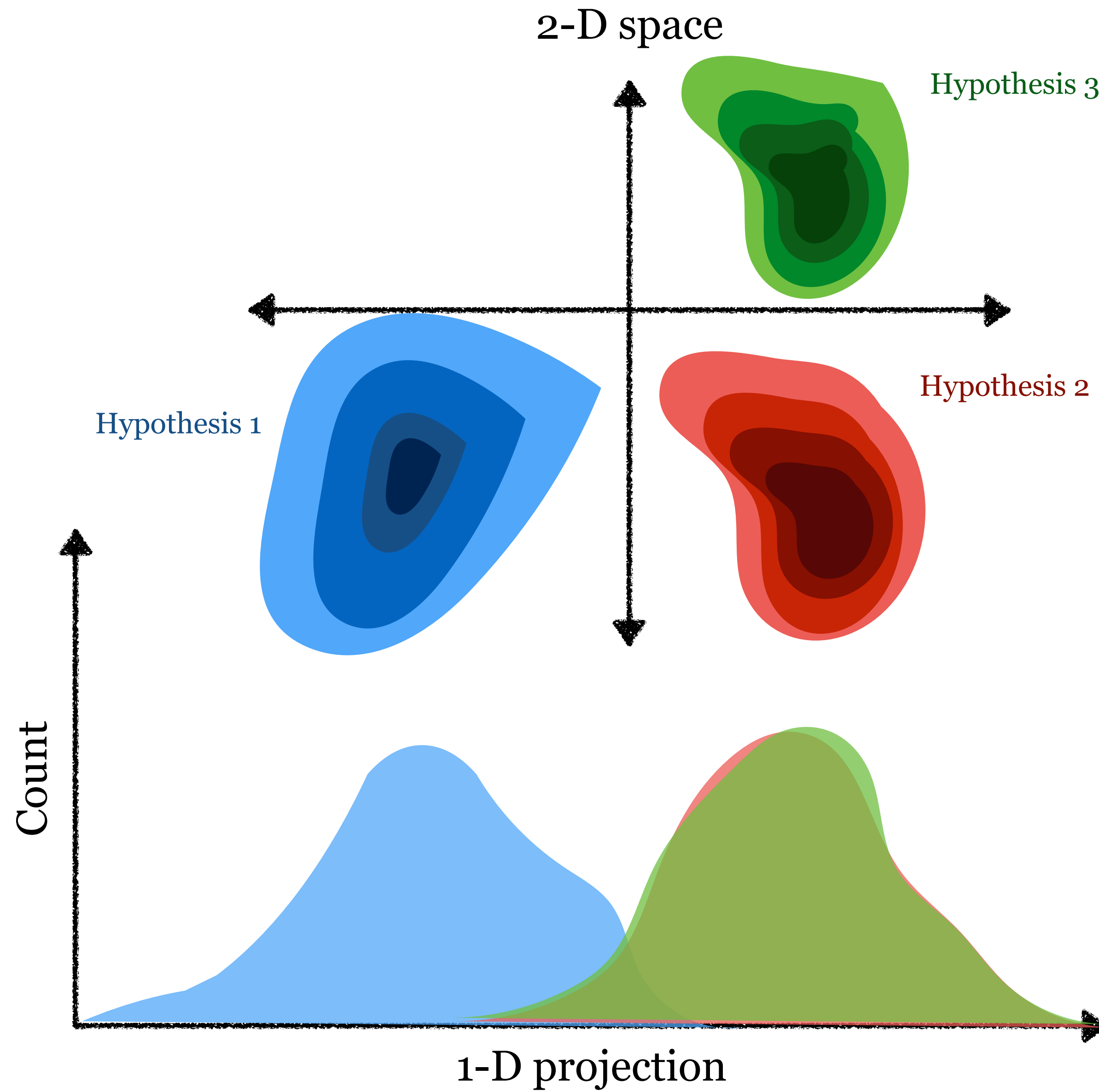
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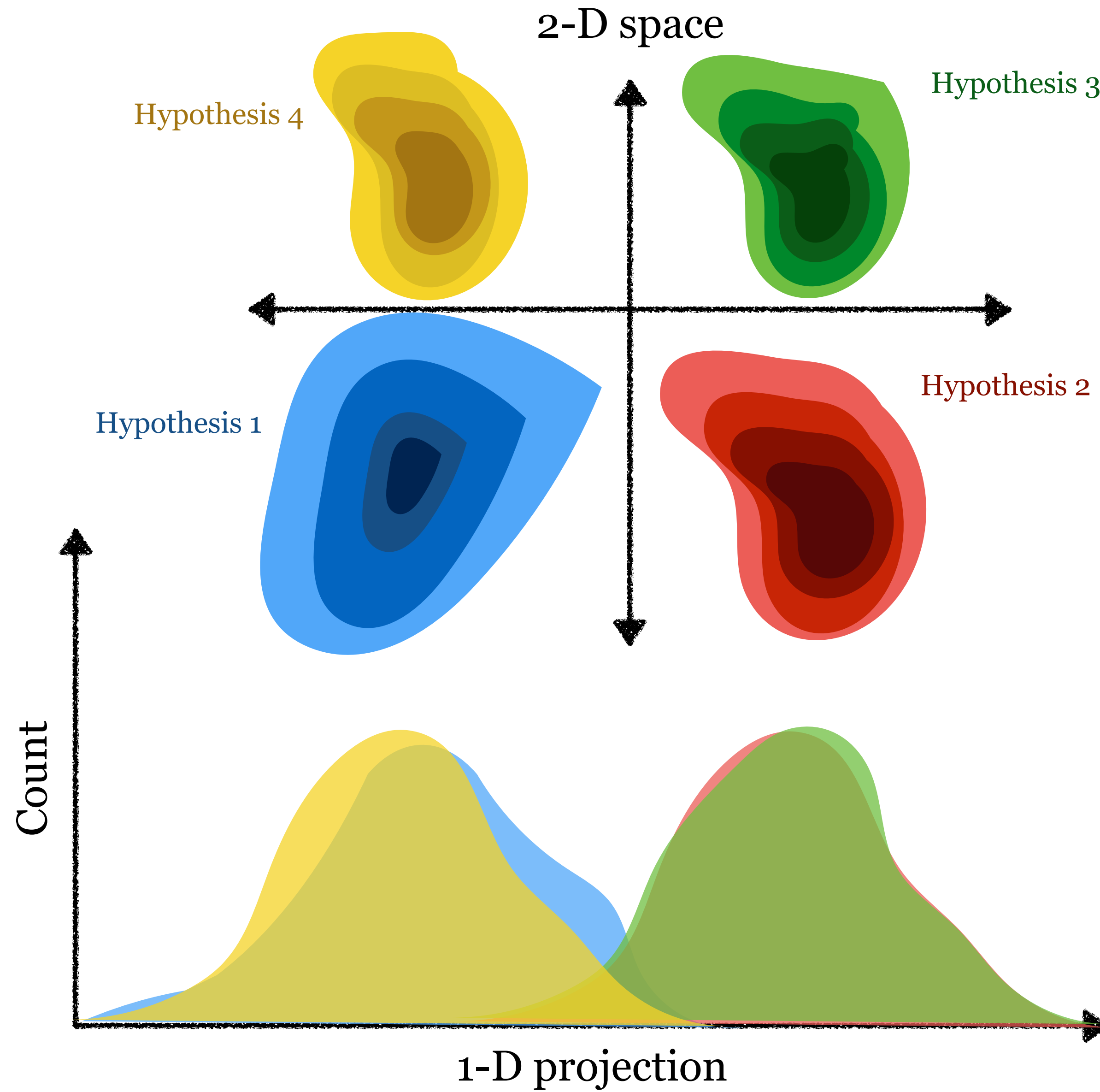
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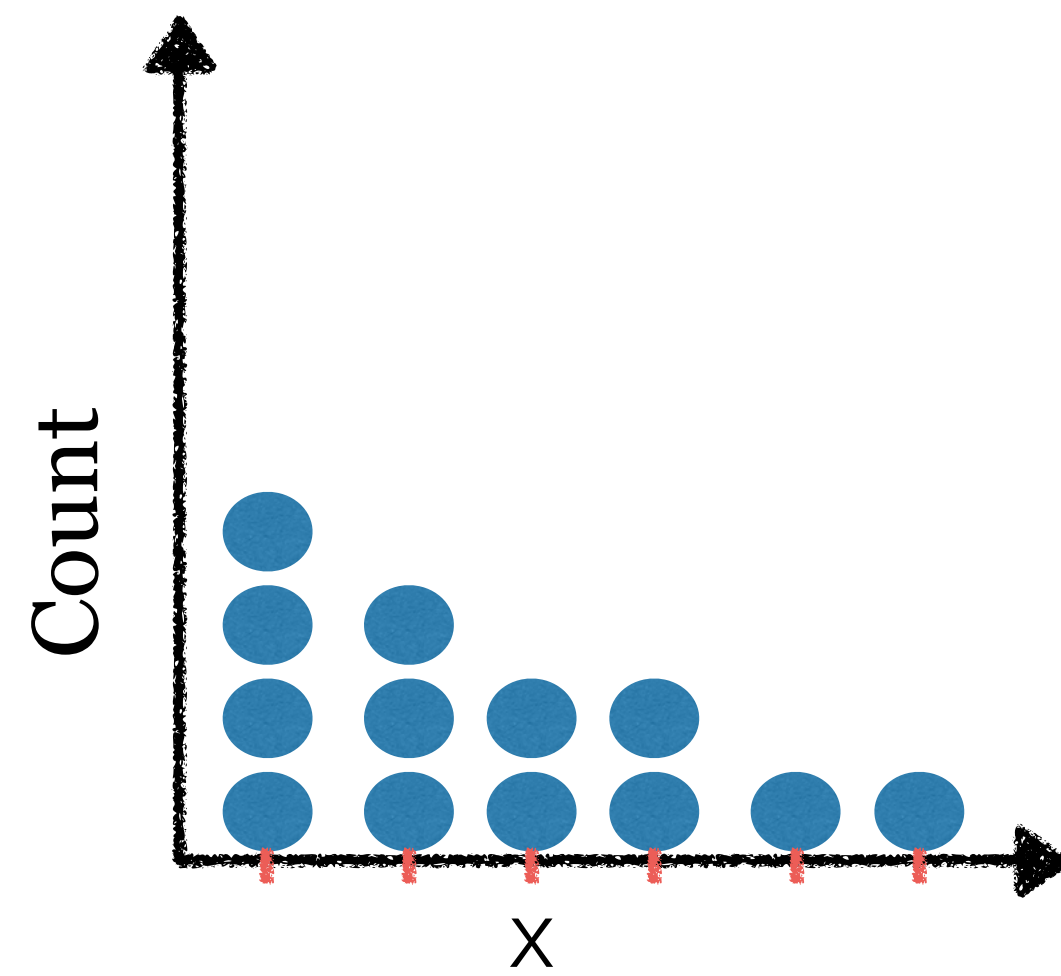
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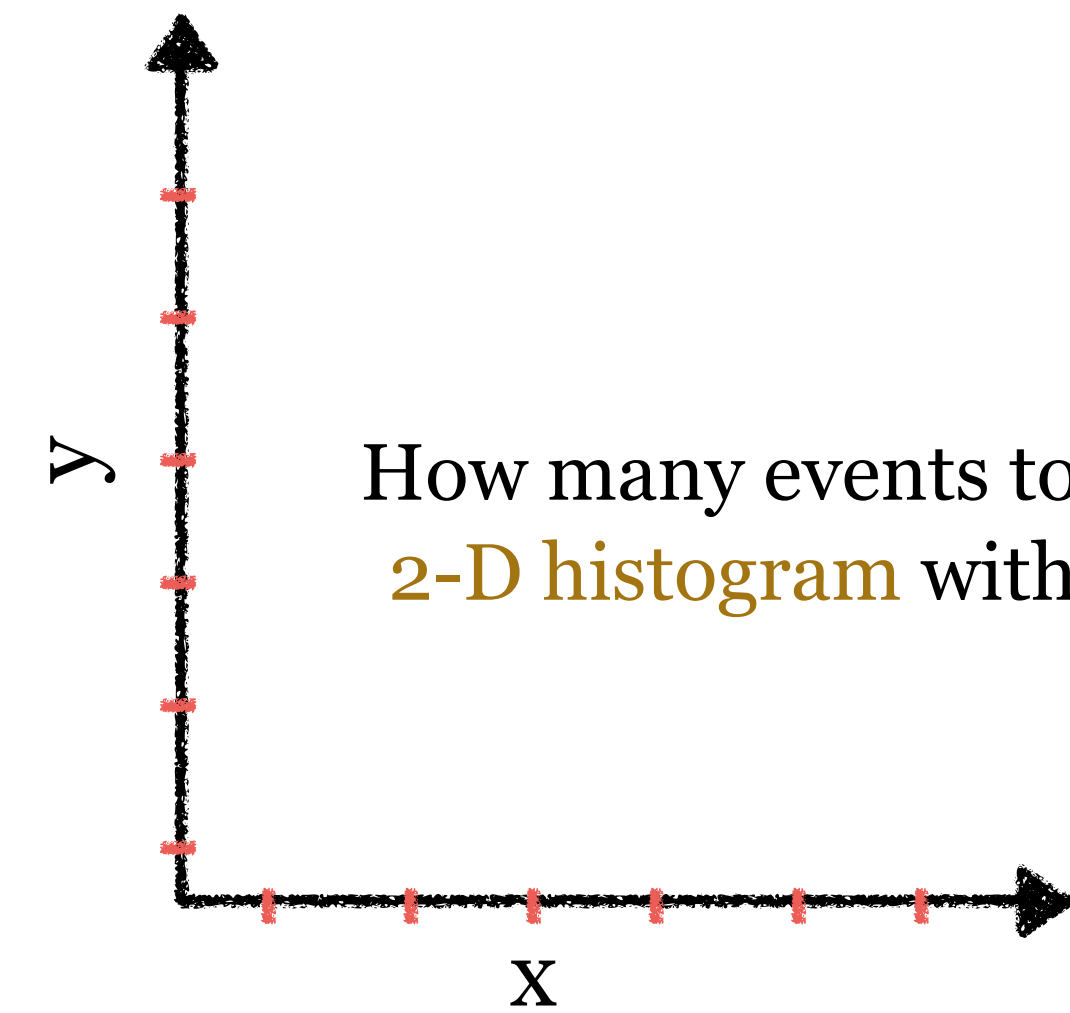
- Clearly separable in 2-D
- A 1-D sufficient summary statistic does not exist for the most important particle physics measurements
 - Higgs Width: [hal-02971995\(p172\)](https://arxiv.org/abs/1506.02169): Ghosh et al
 - Systematic uncertainties: [PRD 104.056026](https://arxiv.org/abs/1506.02169): Ghosh et al
 - Effective Field Theories: [PRD 98.052004](https://arxiv.org/abs/1506.02169): Brehmer et al, (Also see [talk](#) by Eddie McGrady at CHEP)

Need to perform inference directly in high dim

But probability density estimation in higher dimensions is hard...



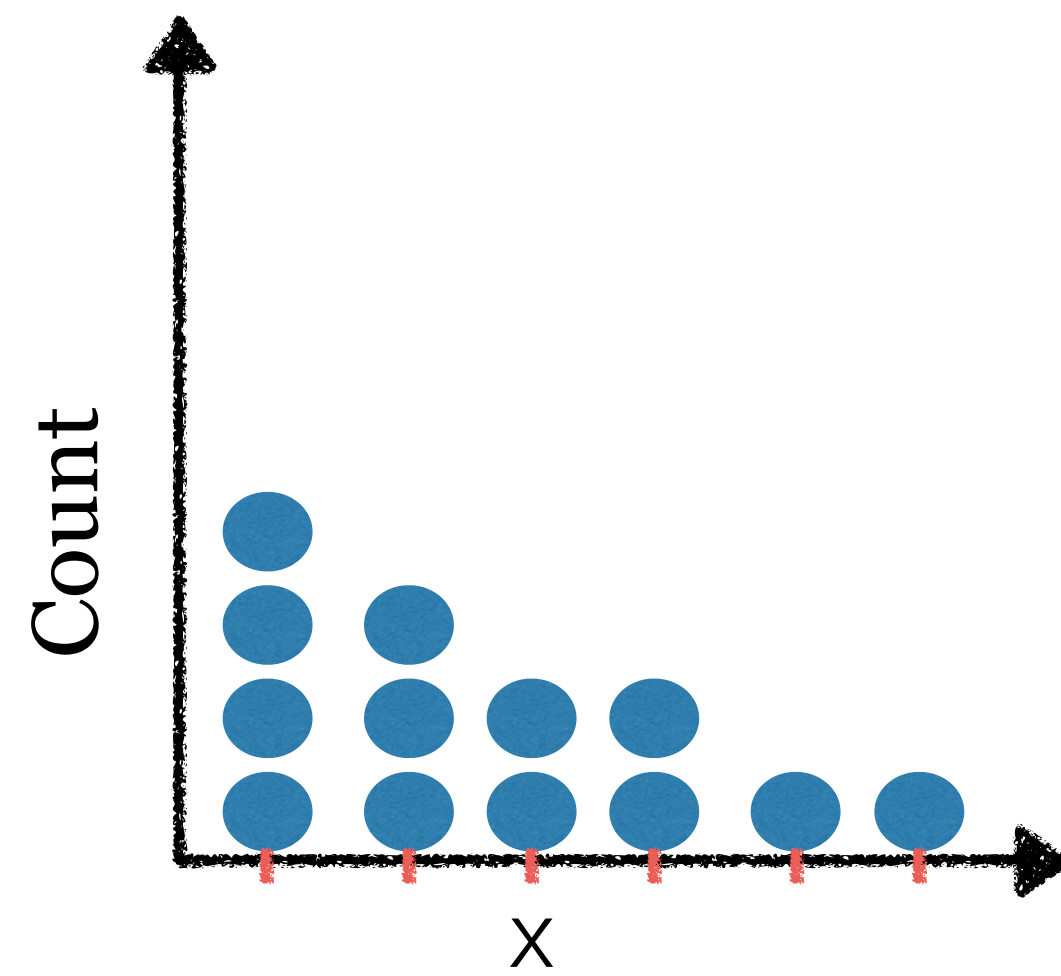
1-D histogram with 6 bins: few events enough to populate it



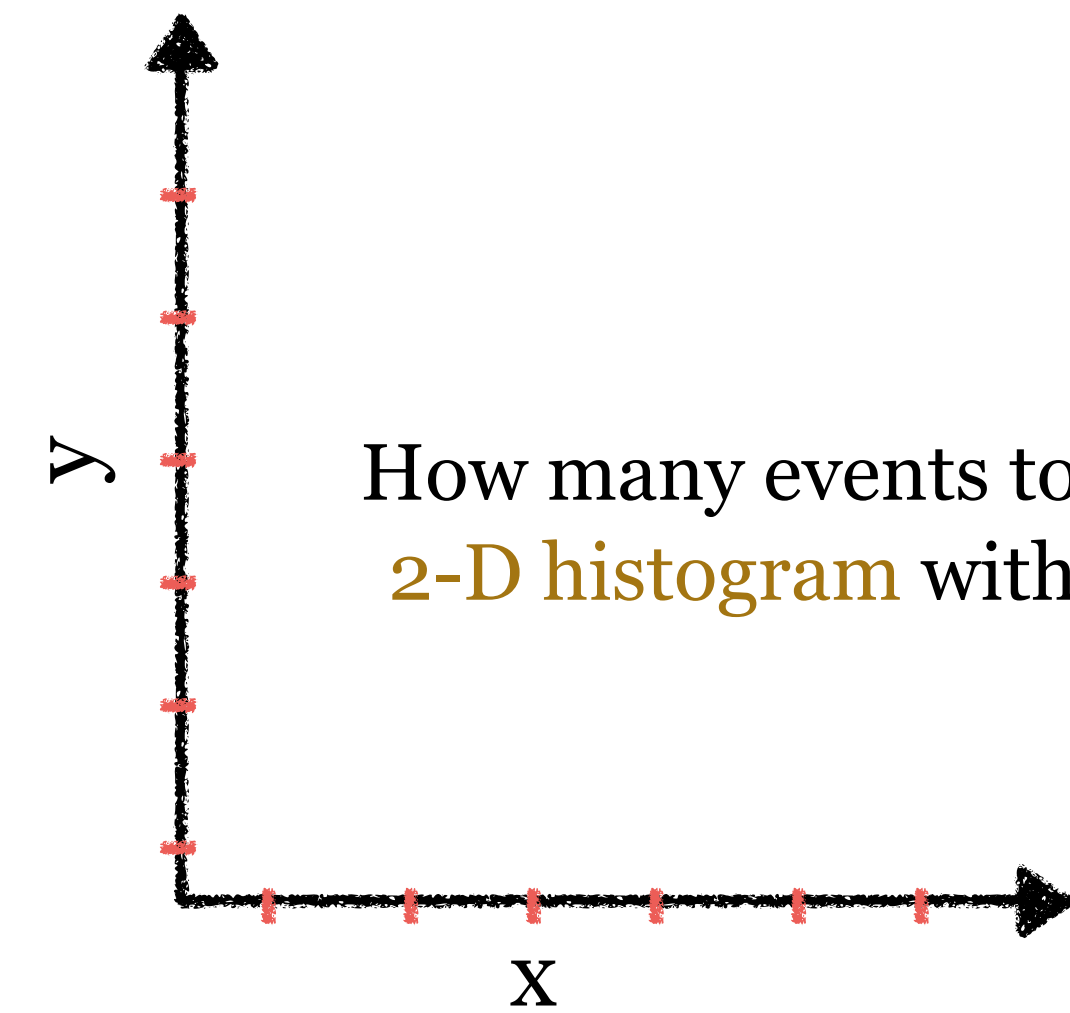
How many events to populate 2-D histogram with 6^2 bins ?

How many events for 50-D histogram with 6^{50} bins ?

But probability density estimation in higher dimensions is hard...



1-D histogram with 6 bins: few events enough to populate it



How many events to populate 2-D histogram with 6^2 bins ?

How many events for 50-D histogram with 6^{50} bins ?

Is there a better way to scale to higher dimensions ?

Neural networks can give us unbinned likelihood ratios

$$\mathcal{L}(\mu | \mathcal{D}) = p(\mathcal{D} | \mu)$$

Neyman–Pearson lemma: Likelihood ratio is the most powerful test statistic

We want to compare likelihoods:

$$\frac{p(\mathcal{D} | \mu)}{p(\mathcal{D} | ref)}$$

A neural network classifier trained on **simulated samples from μ_1** vs **simulated samples from ref** , estimates the decision function:

$$s(x_i) = \frac{p(x_i | \mu_1)}{p(x_i | \mu_1) + p(x_i | ref)}$$

Which contains all the information required for the likelihood ratio:

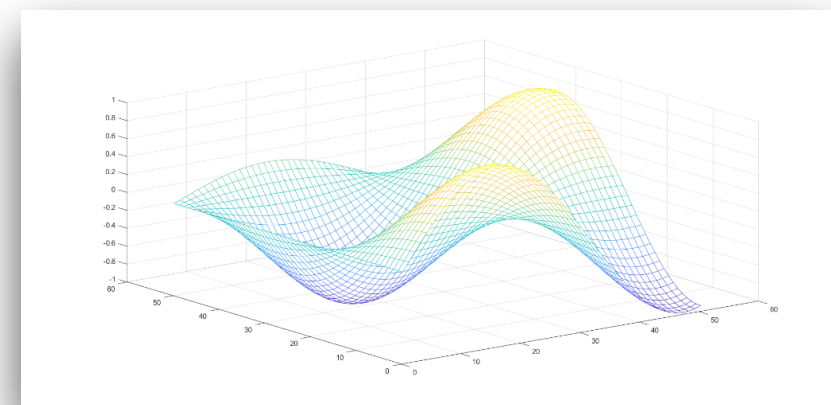
$$\frac{p(x_i | \mu_1)}{p(x_i | ref)} = \frac{s(x_i)}{1 - s(x_i)}$$

- * Optimal statistic to test each value of μ
- * We get the LR *per event* (unbinned)

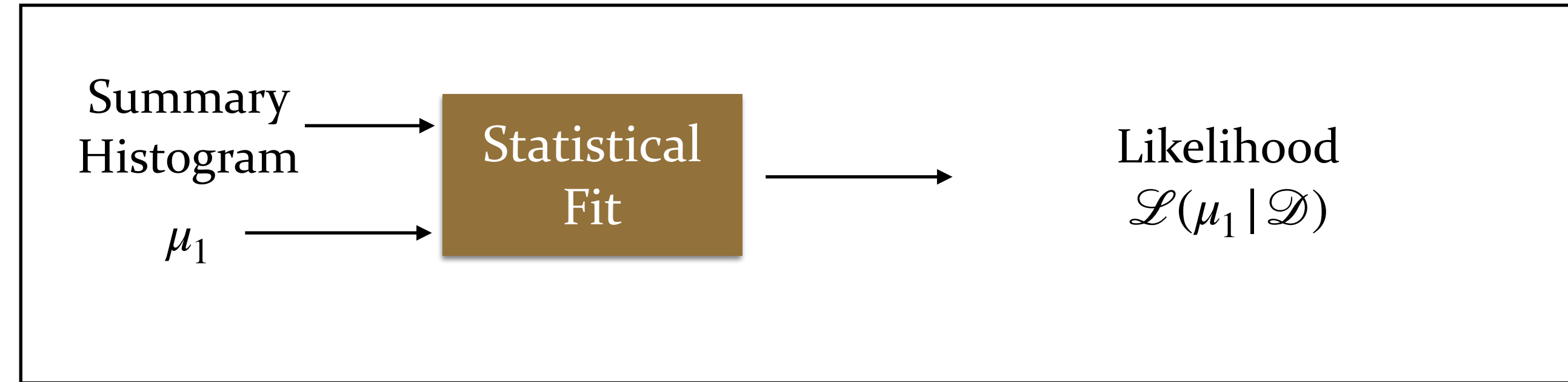
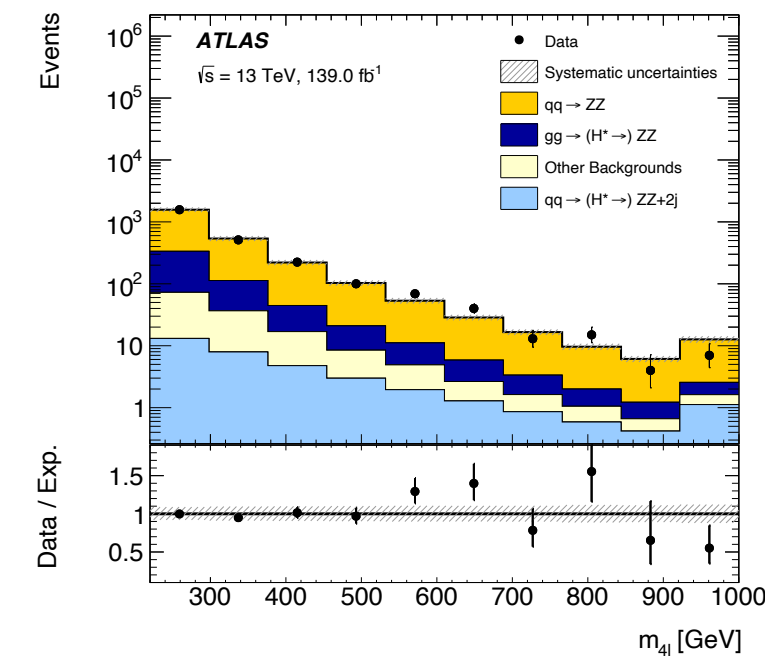
A new paradigm: Neural simulation-based inference (NSBI)

Traditional framework:

Summarisation
to histogram

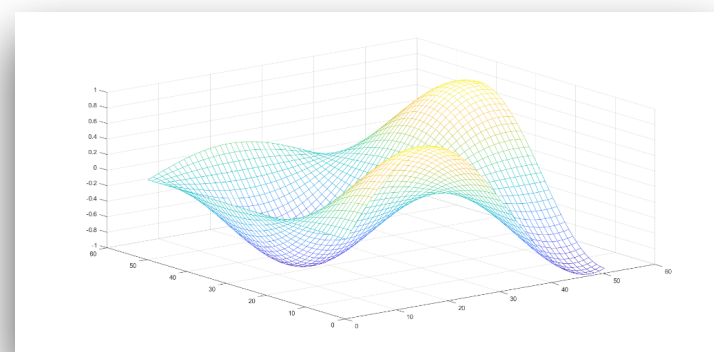


High-dim data



μ is now arbitrary parameter of interest(s)

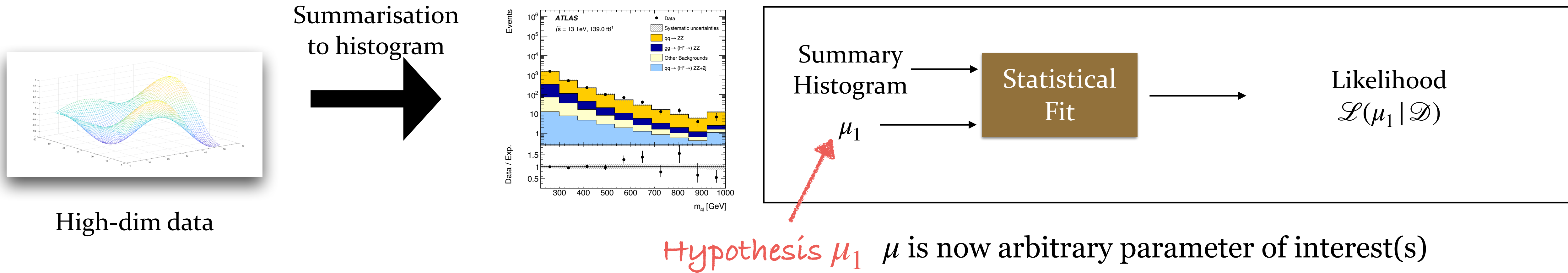
Neural simulation-based inference framework:



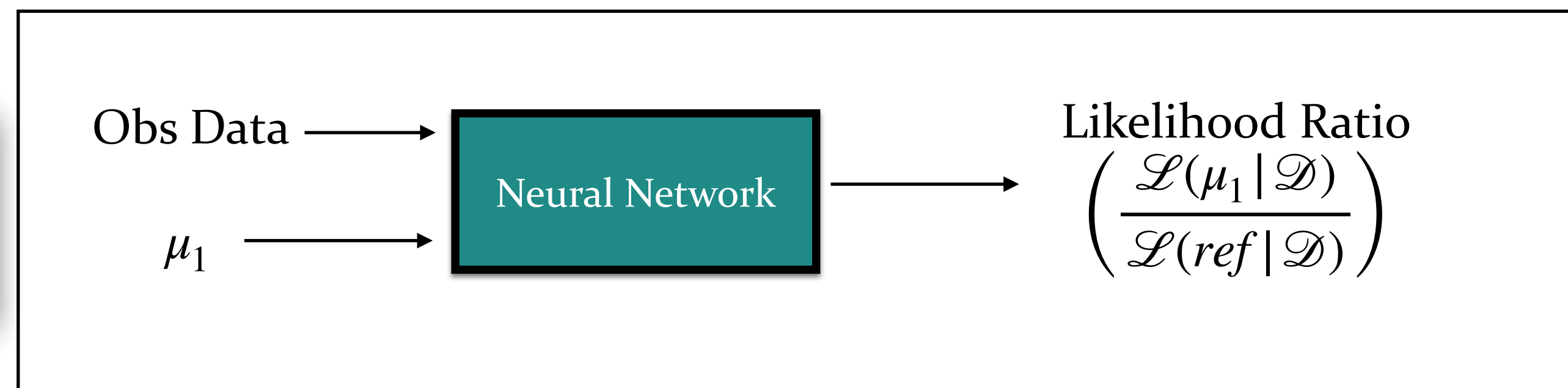
High-dim data

A new paradigm: Neural simulation-based inference (NSBI)

Traditional framework:

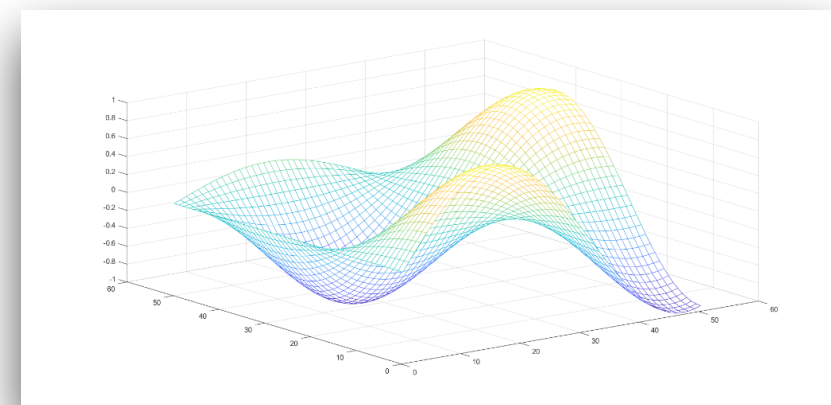


Neural simulation-based inference framework:

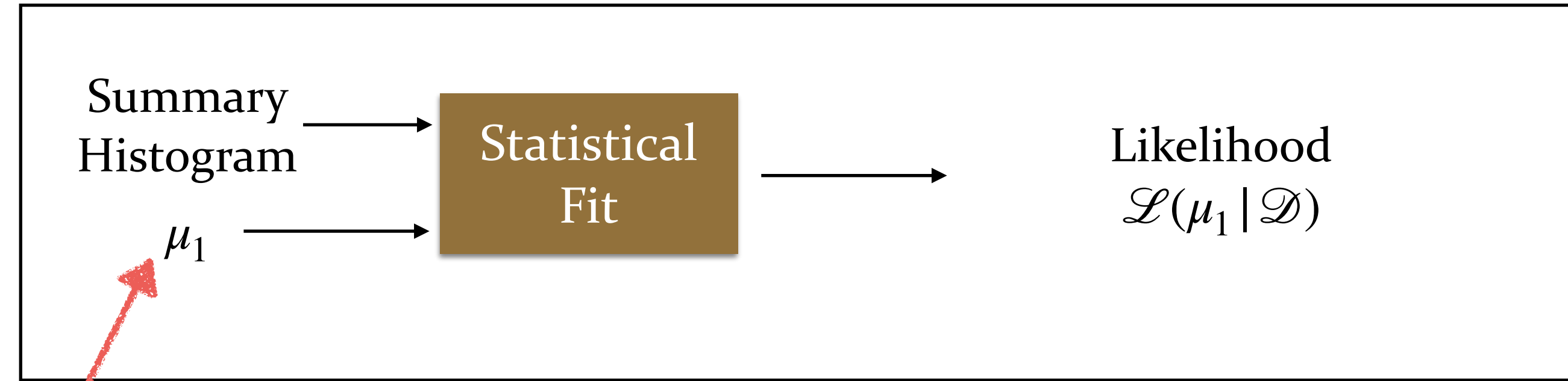
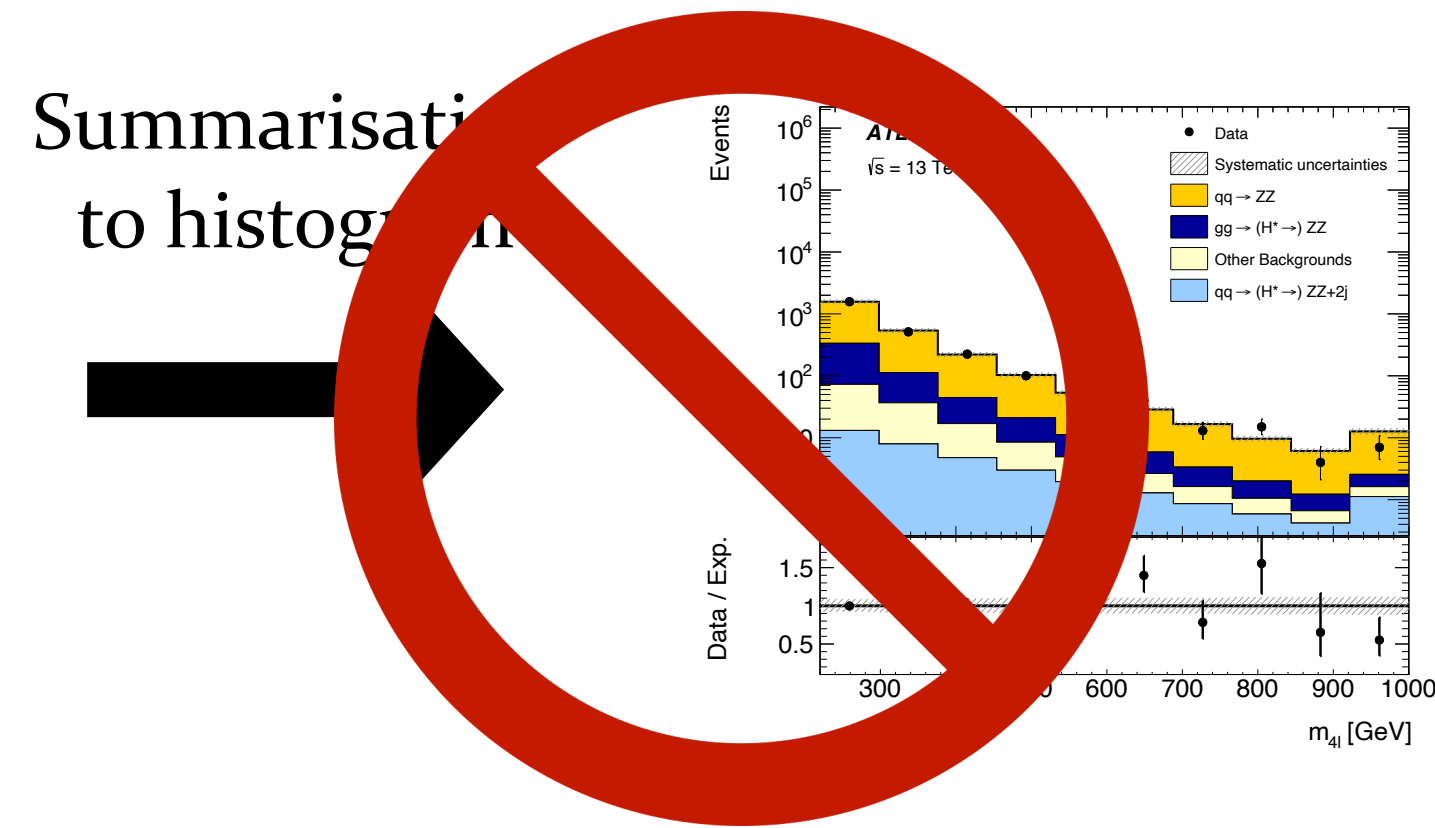


A new paradigm: Neural simulation-based inference (NSBI)

Traditional framework:

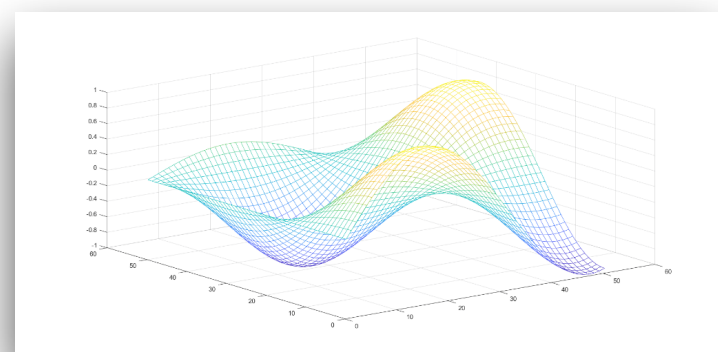


High-dim data

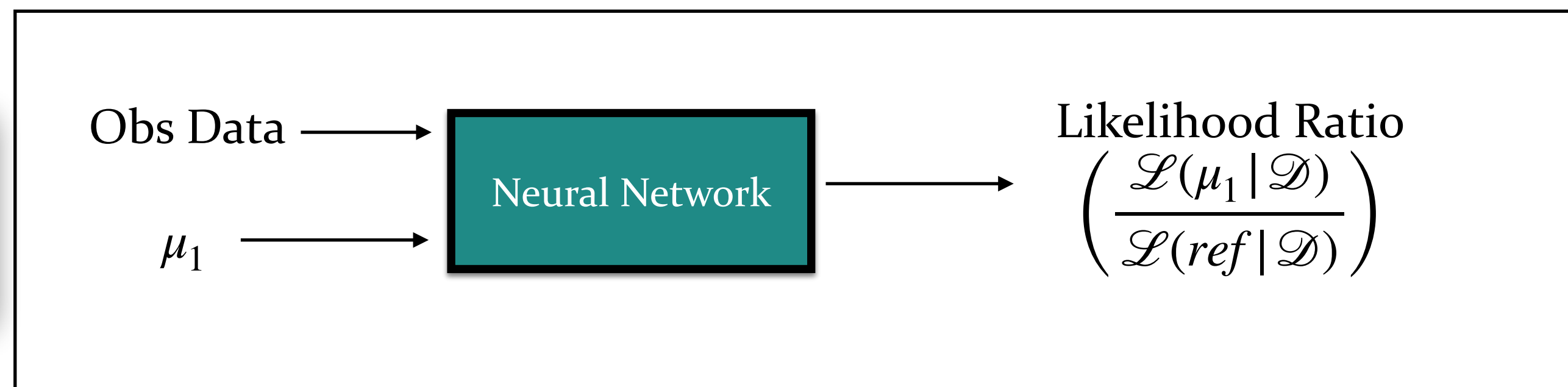


Hypothesis μ_1 μ is now arbitrary parameter of interest(s)

Neural simulation-based inference framework:

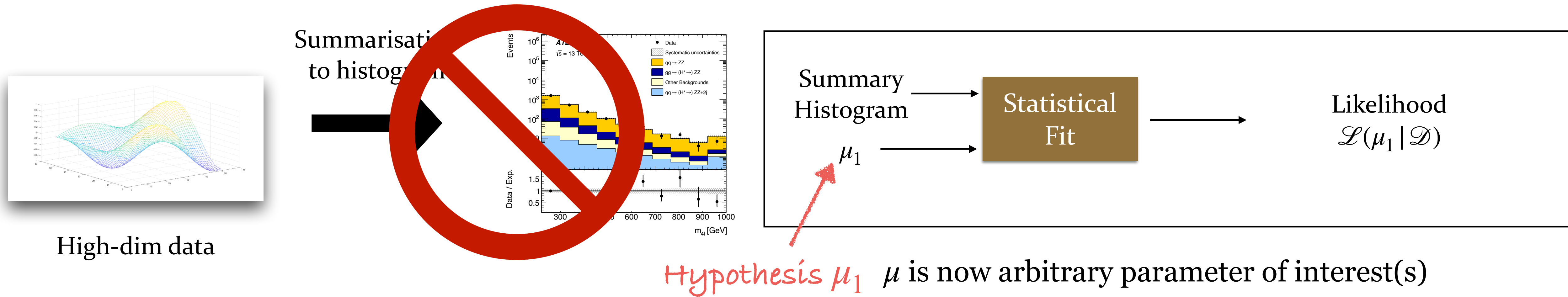


High-dim data

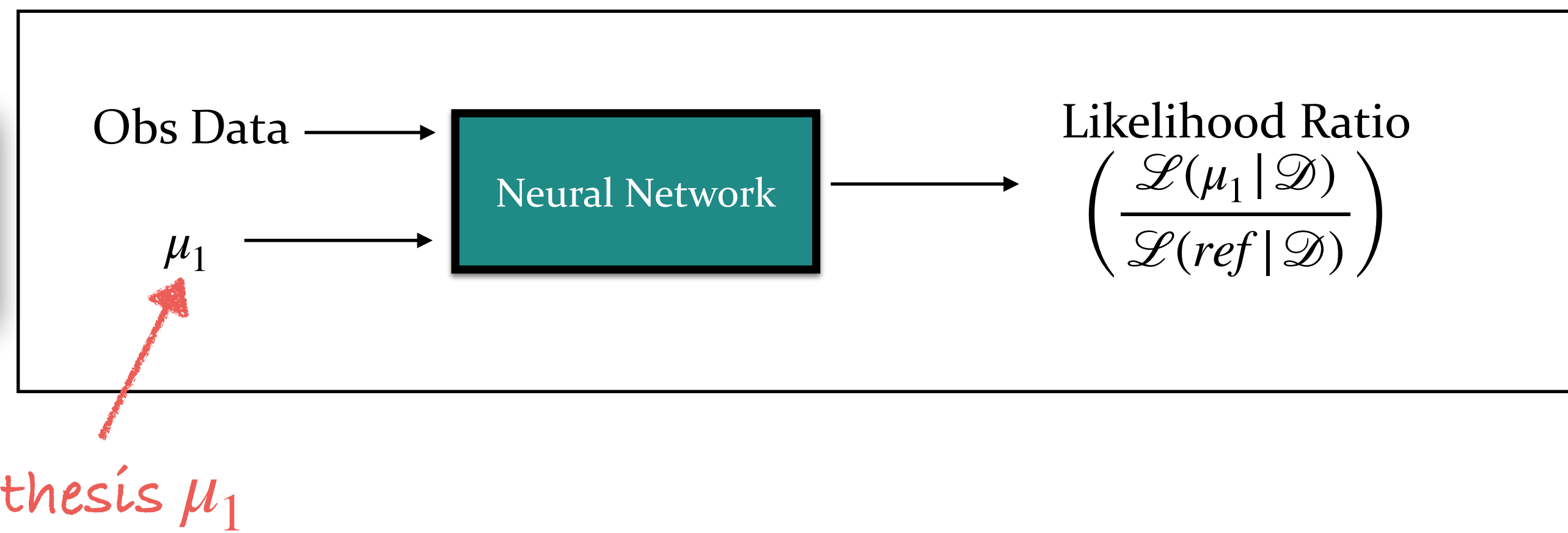


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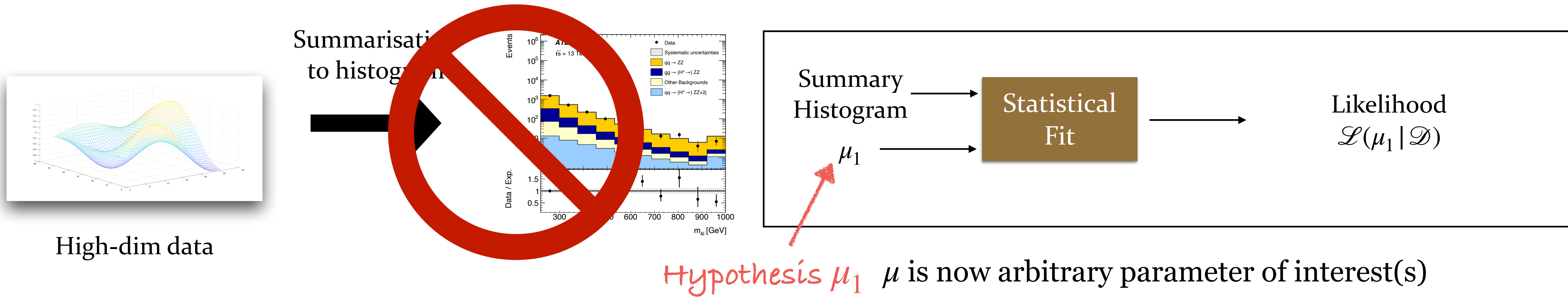


Neural simulation-based inference framework:

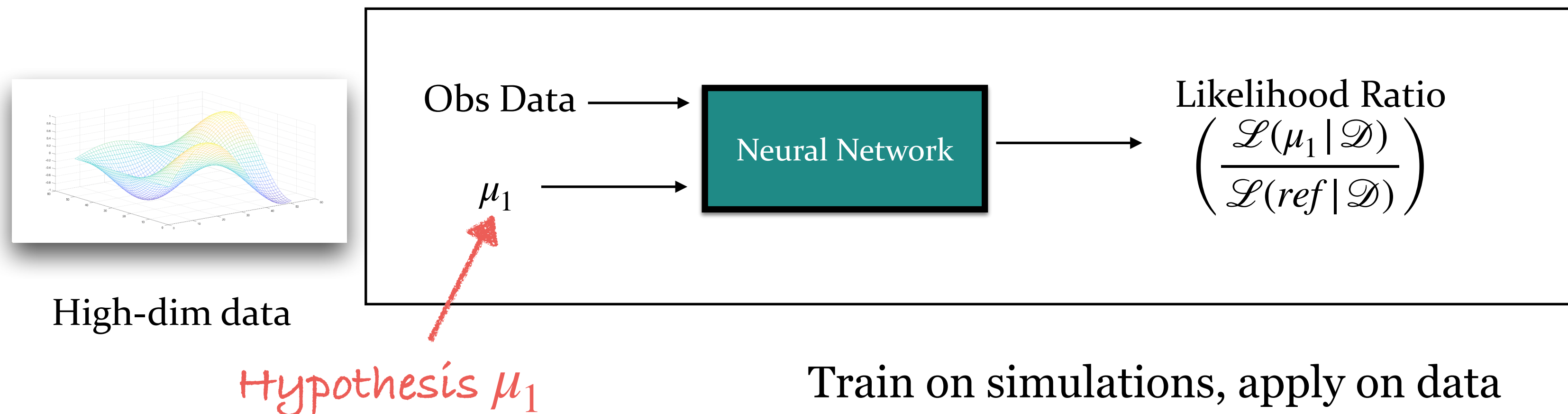


A new paradigm: Neural simulation-based inference (NSBI)

Traditional framework:



Neural simulation-based inference framework:





Is this dream attainable in practice?

- How to ensure robustness? By design and validation
- How to incorporate systematic uncertainties?
- How to perform Neyman construction in high dimensions?

Is this dream attainable in practice?

Yes!

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Volker Crede and John Yelton

Applied on Run2 data, superseding previous ATLAS paper on same data !

A VERY FUN PROBLEM

How a grad student got LHC data to play nice with quantum interference

New approach is already having an impact on the experiment's plans for future work.

MATT VON HIPPEL - 23 JUN 2025 11:00 70



Is this dream attainable in practice?

Yes!

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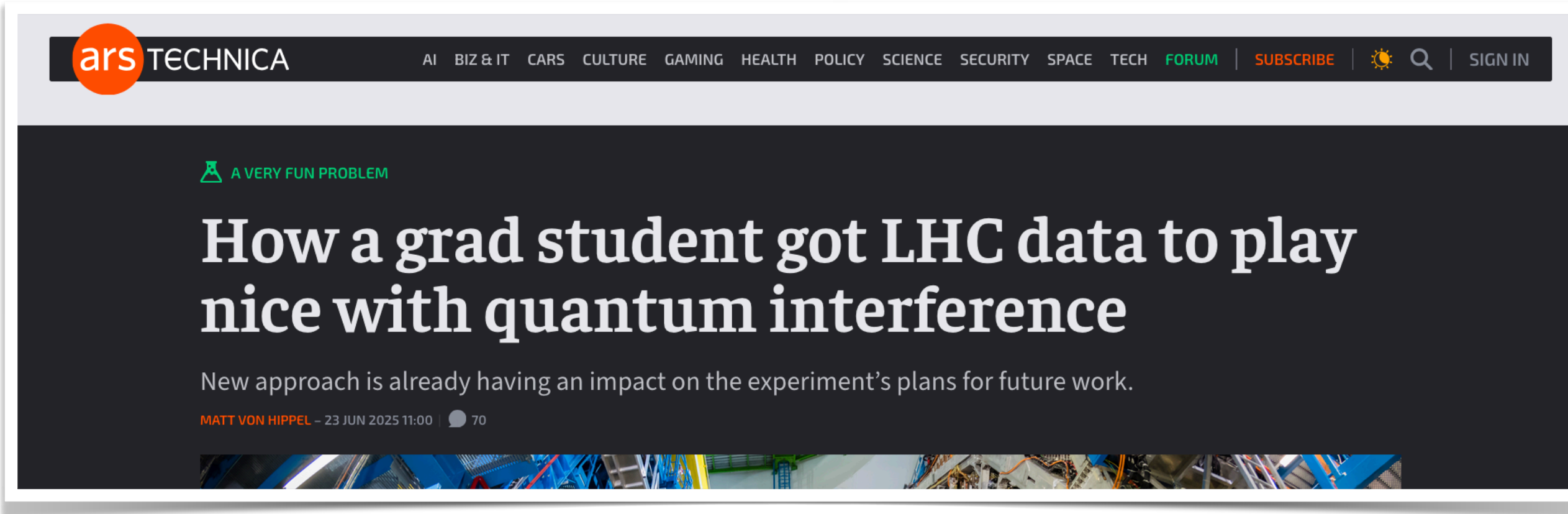
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Applied on Run2 data, superseding previous ATLAS paper on same data !



Today's tutorial will follow this approach, although many variations are possible

Is this dream attainable in practice?

Yes!

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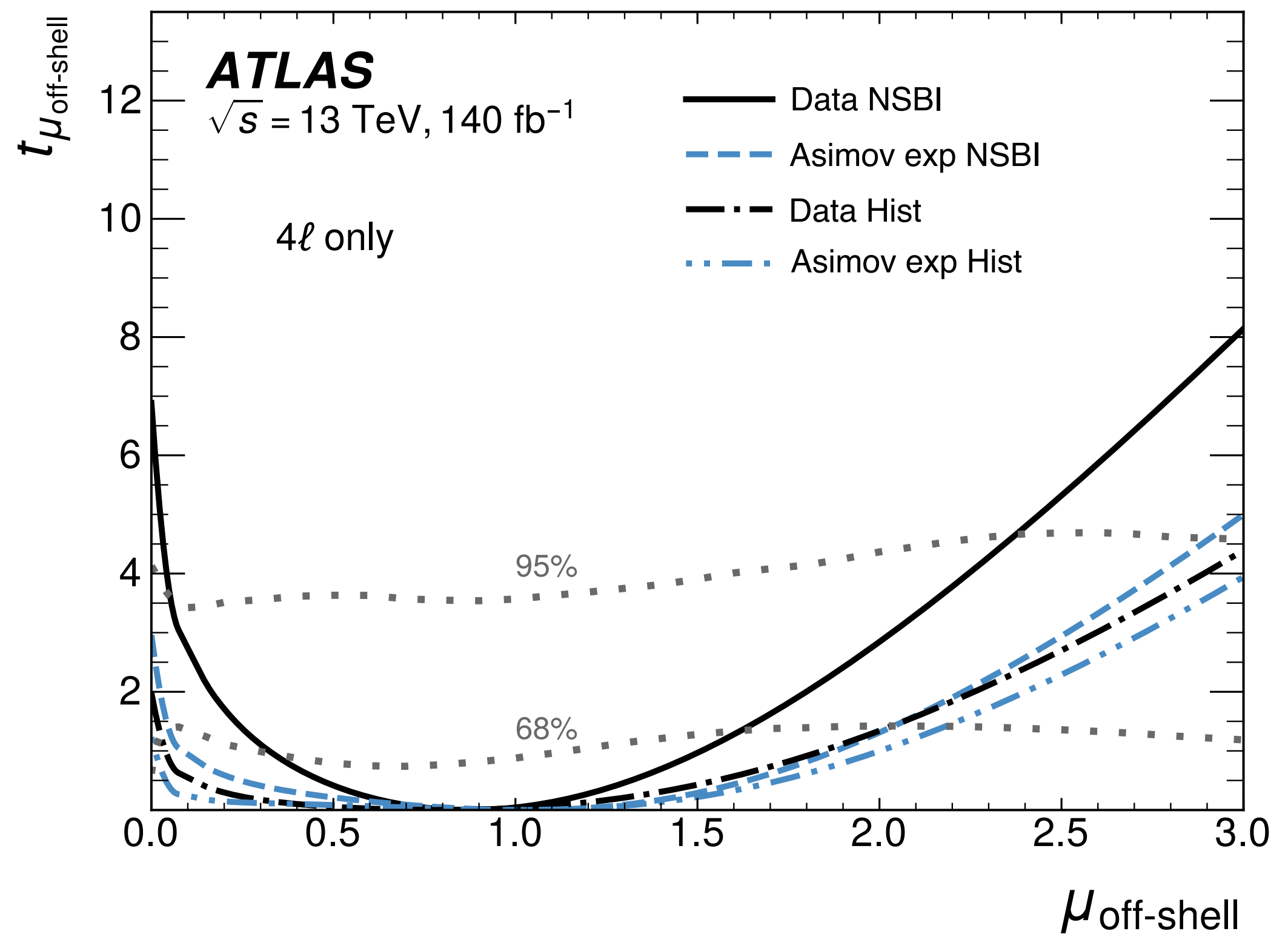
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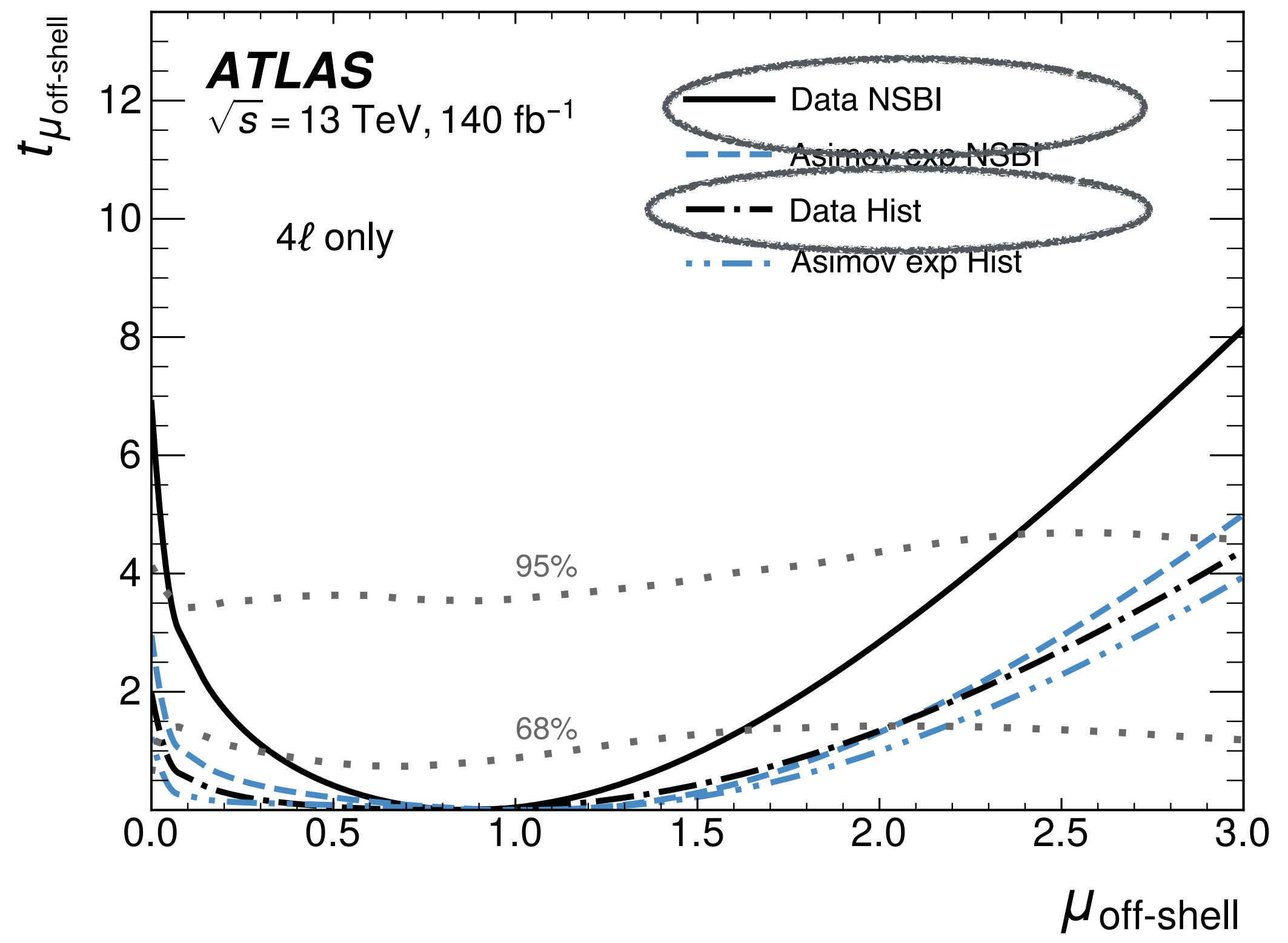
Application in off-shell Higgs Run 2 analysis: big improvement!

NSBI vs histogram analysis



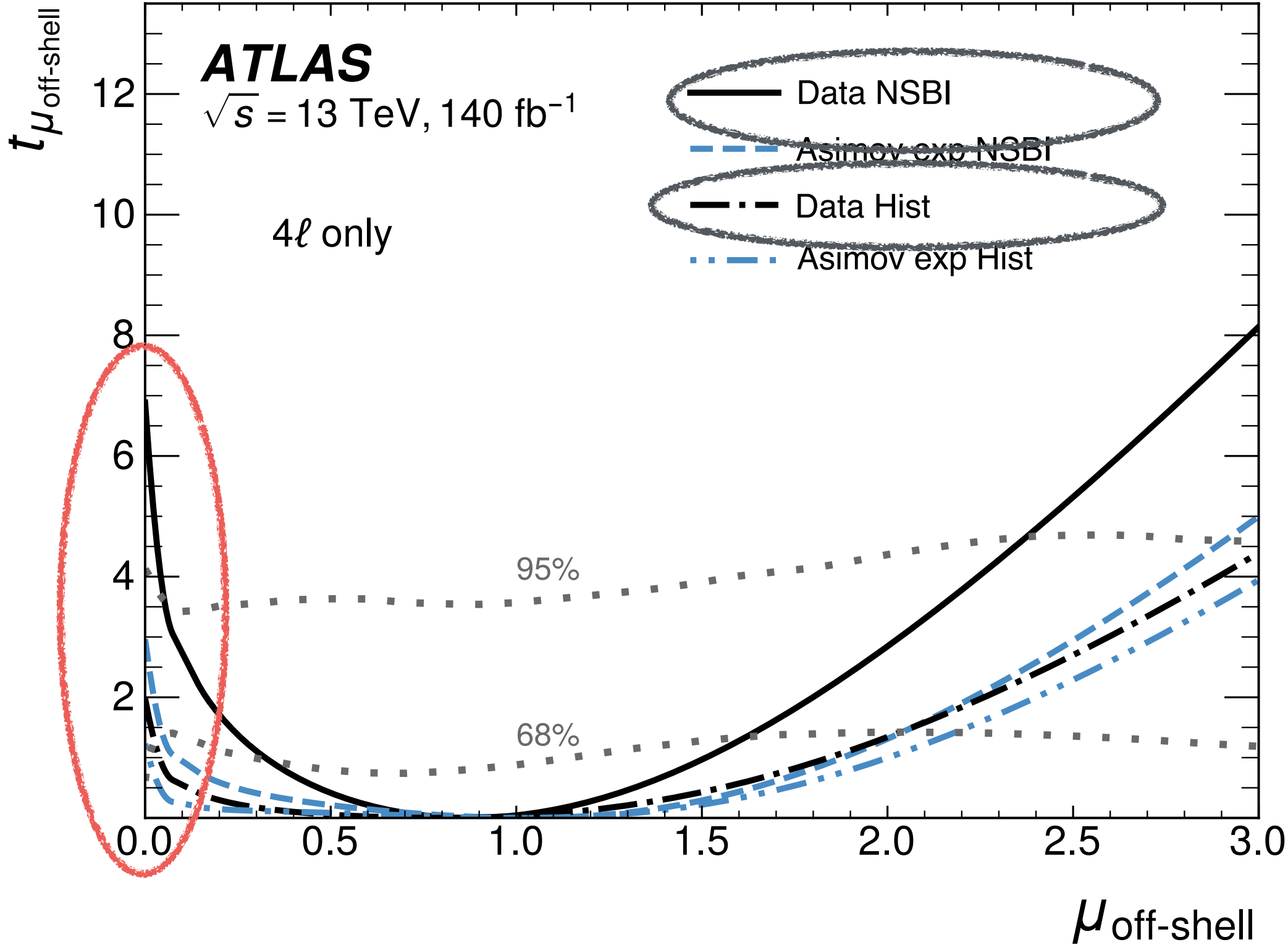
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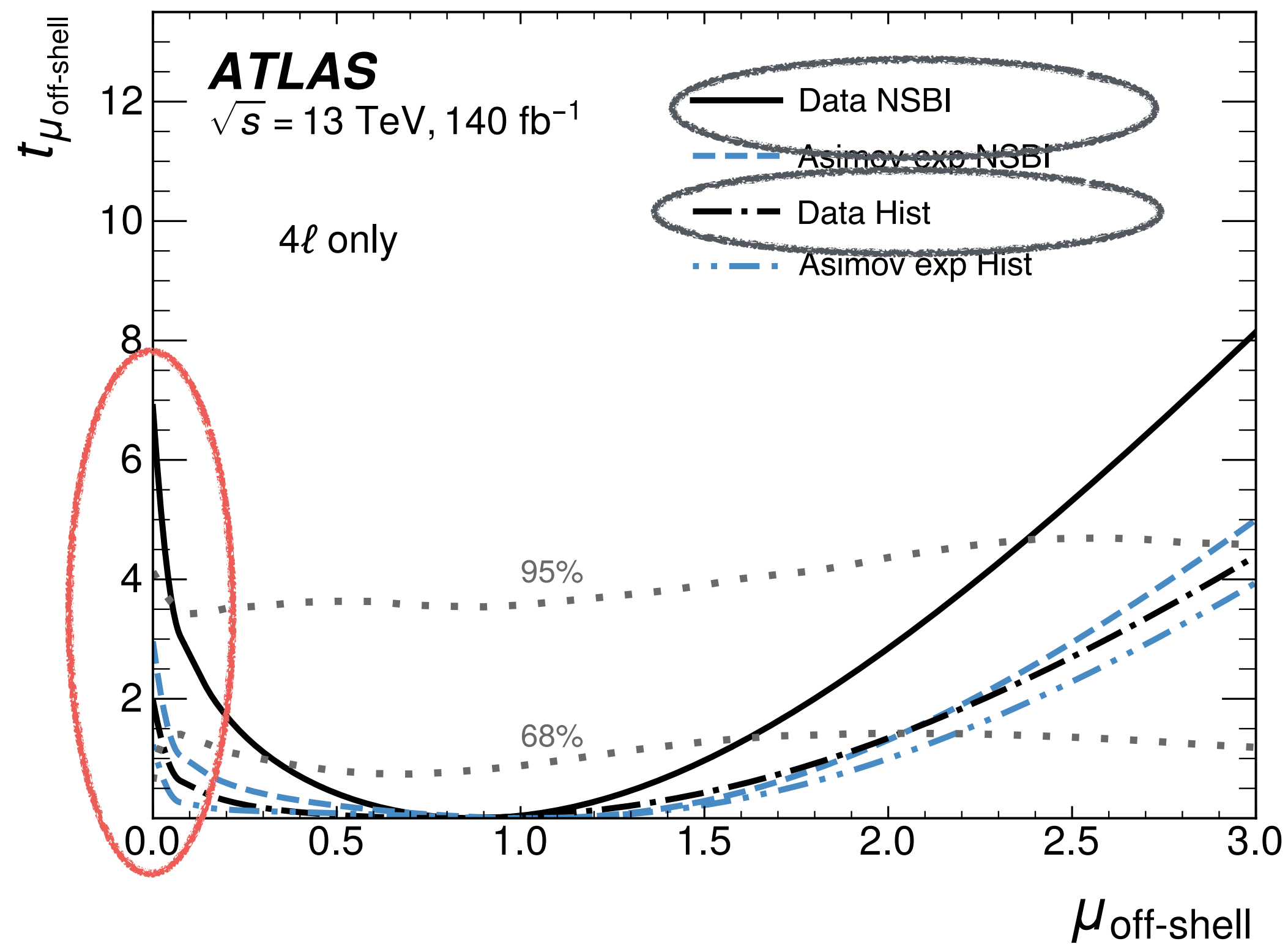
NSBI vs histogram analysis



Unprecedented improvement in ability to reject null hypothesis!

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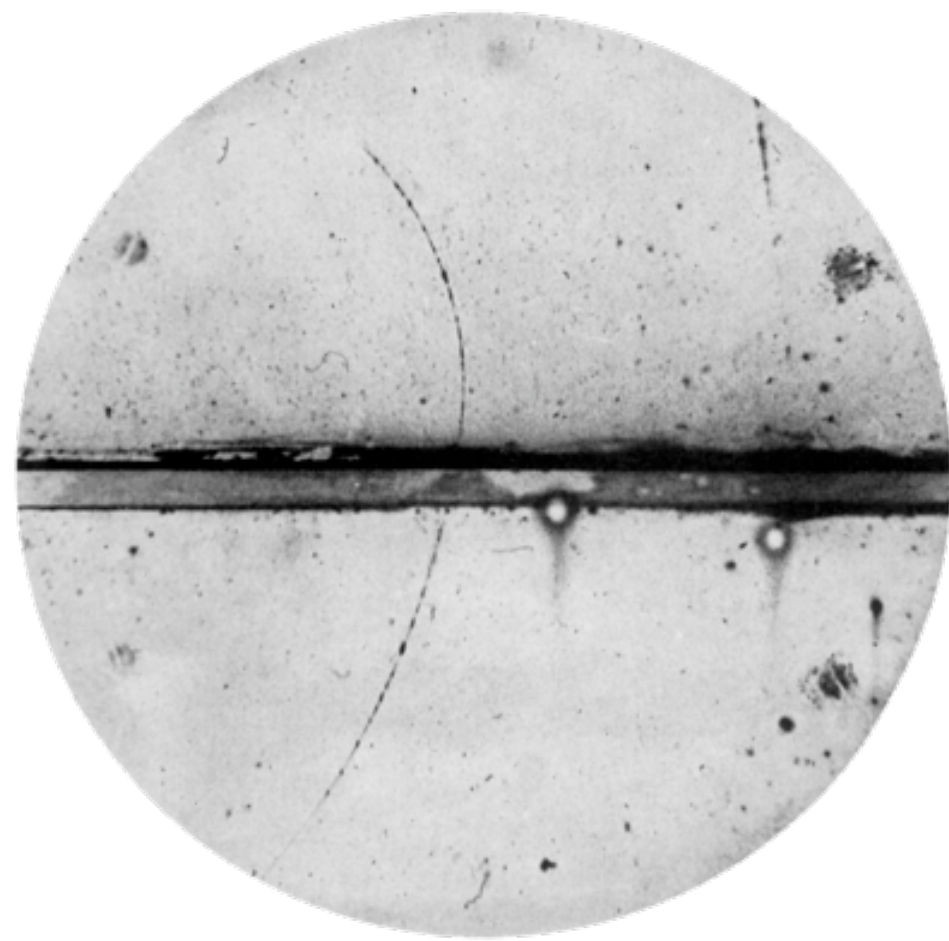
NSBI vs histogram analysis



Unprecedented improvement in ability to reject null hypothesis!

Observed data happens to provide stronger constraint than Asimov for both hist and NSBI

Positron discovery (1930s)



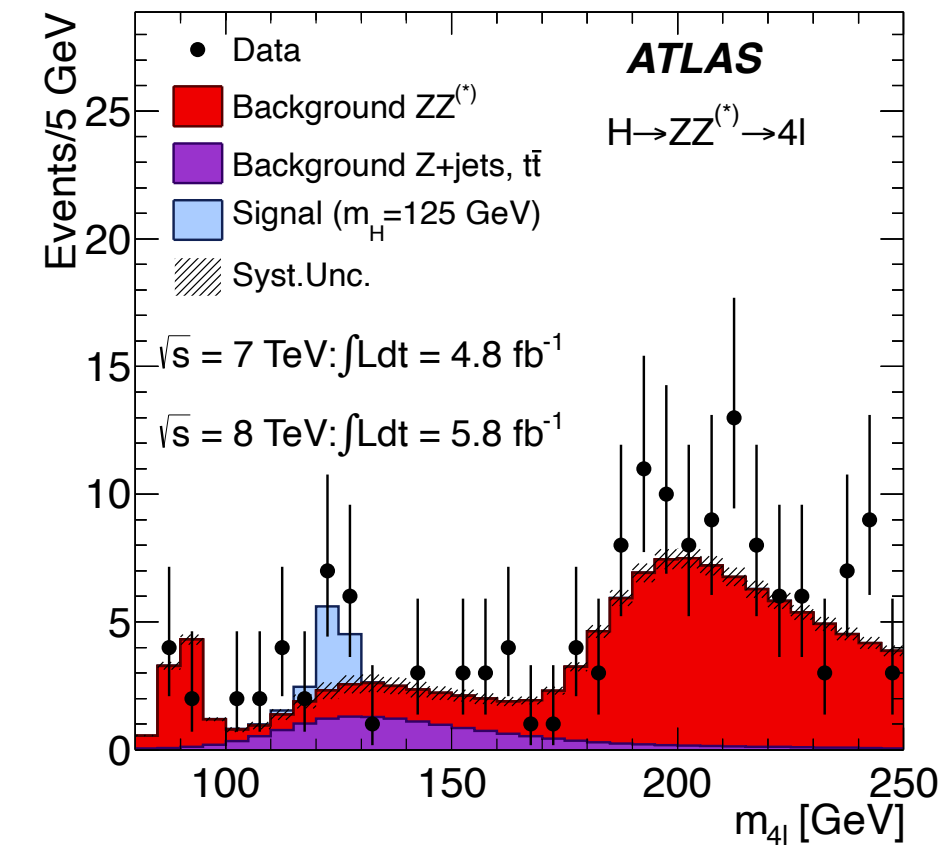
Single event

Top quark discovery (1990s)

Channel:	SVX
observed	27 tags
expected background	6.7 ± 2.1
background probability	2×10^{-5}

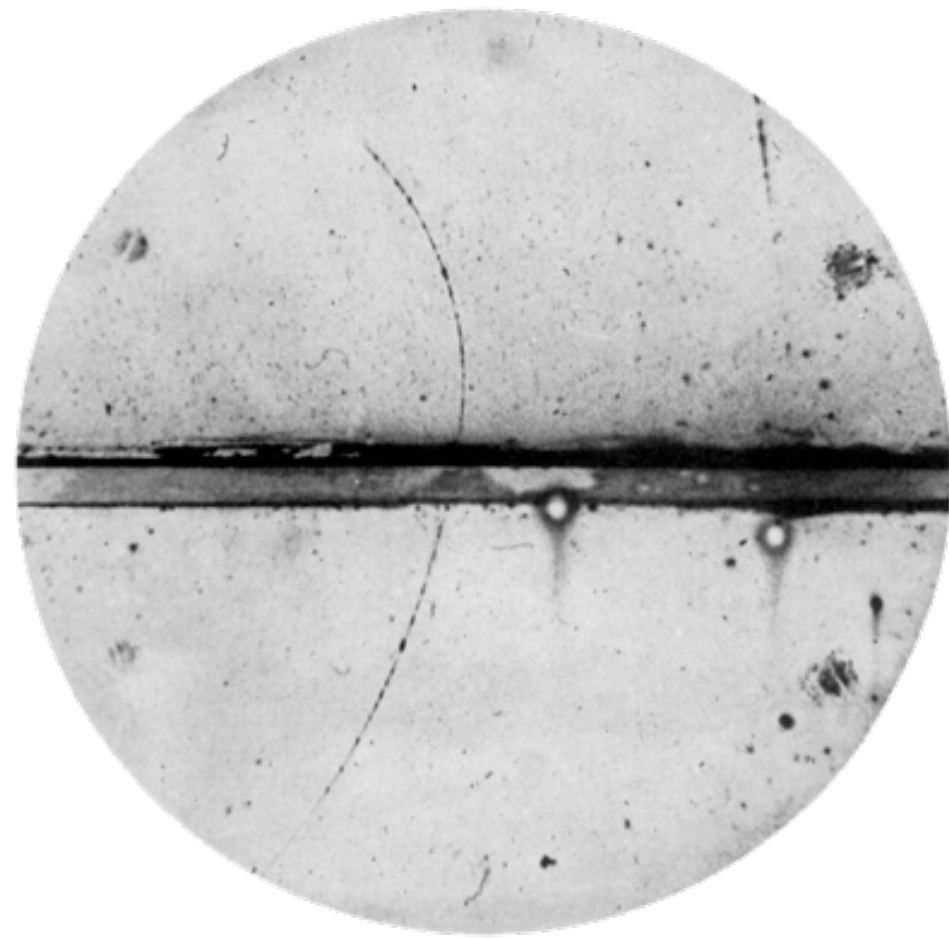
Multiple events:
Cut-and-count

Higgs boson discovery (2010s)



Shape information:
Histogram

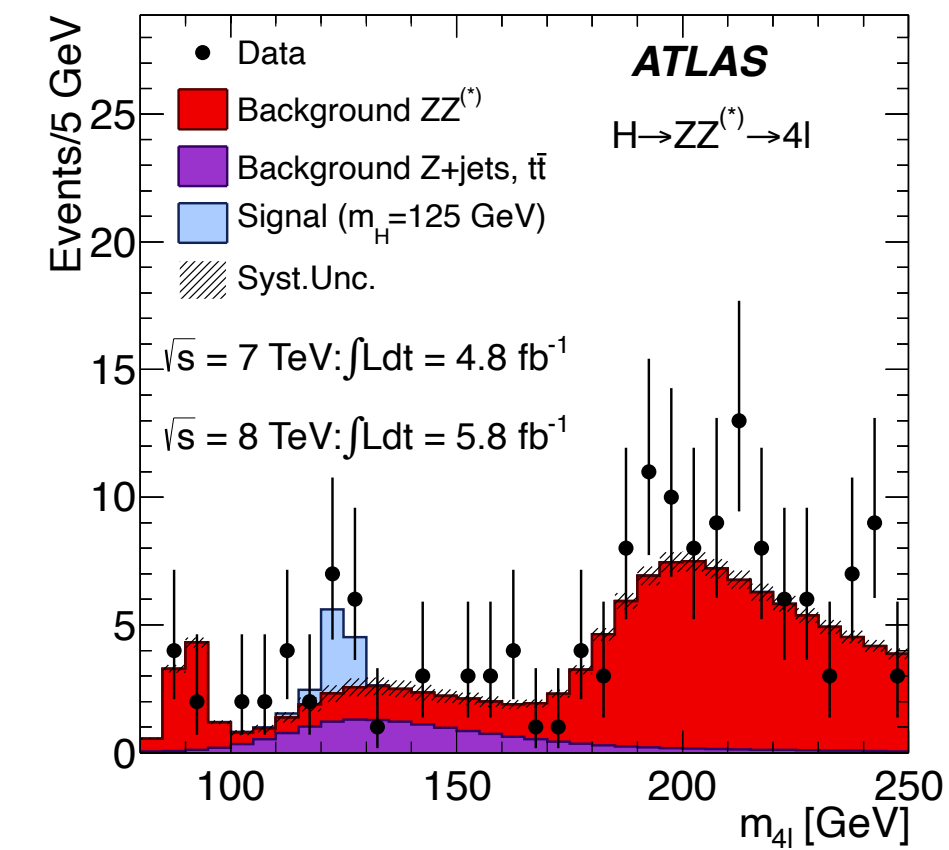
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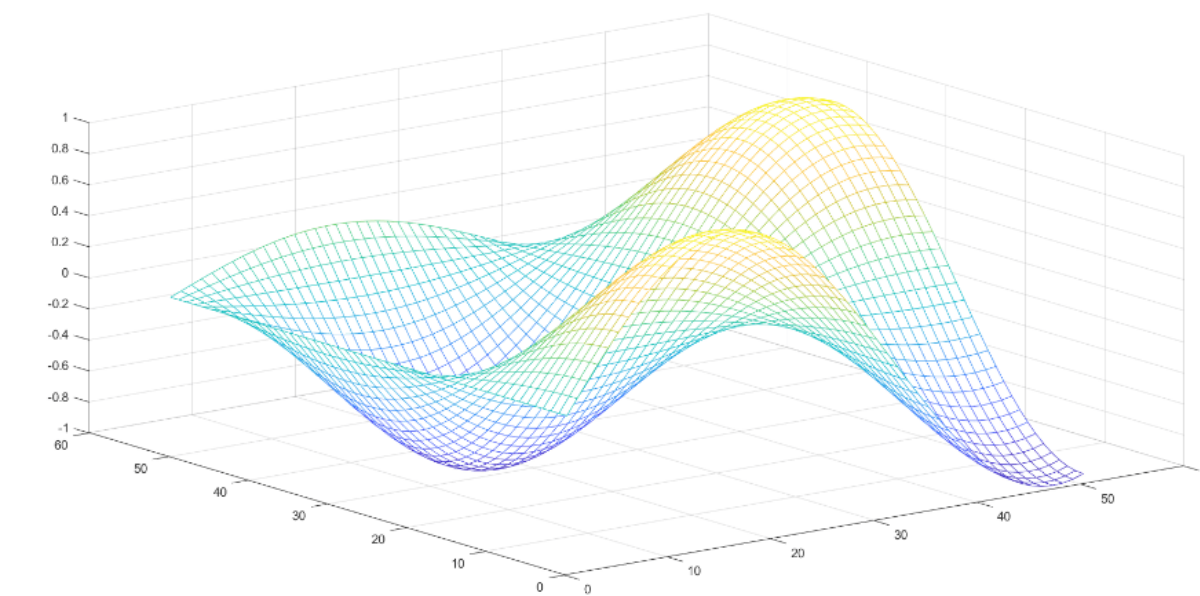
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Channel:	SVX
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Higgs boson discovery (2010s)



Future discovery (2020s ?)



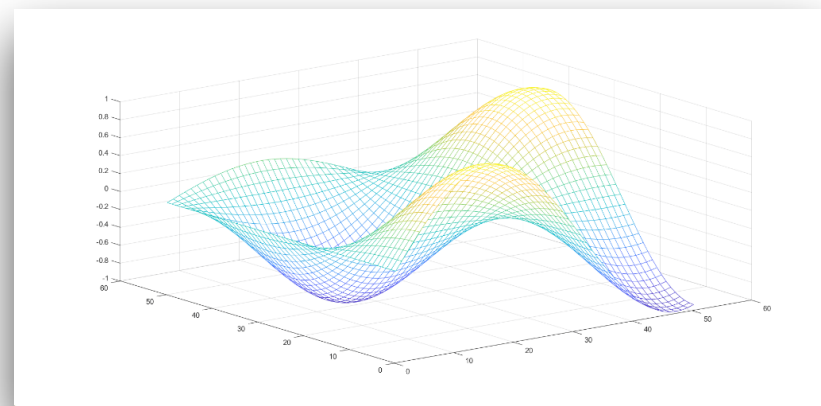
Single event

Multiple events:
Cut-and-count

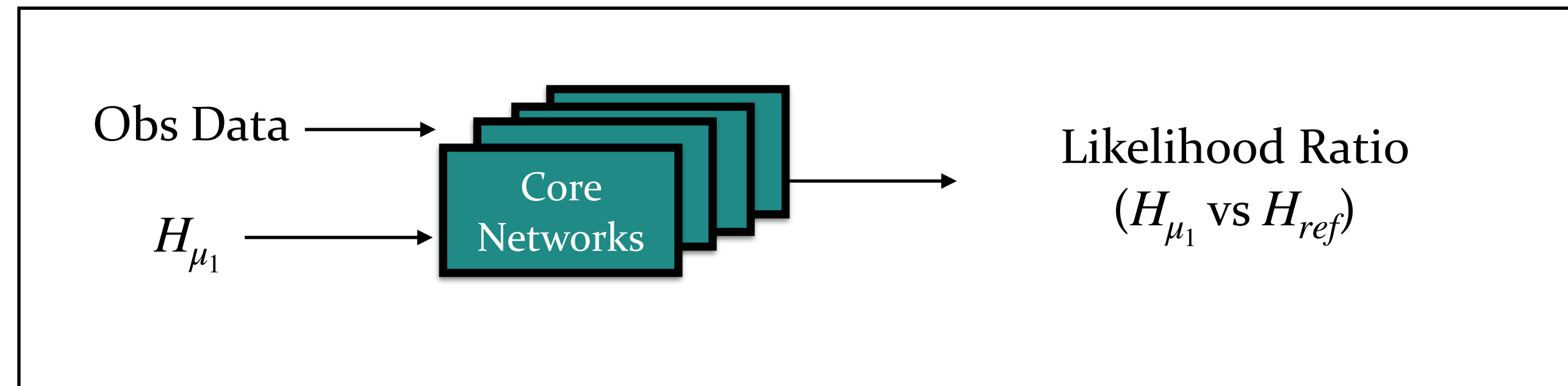
Shape information:
Histogram

High-dim shape information,
continuous (i.e. unbinned):
Neural inference

A robust NSBI implementation

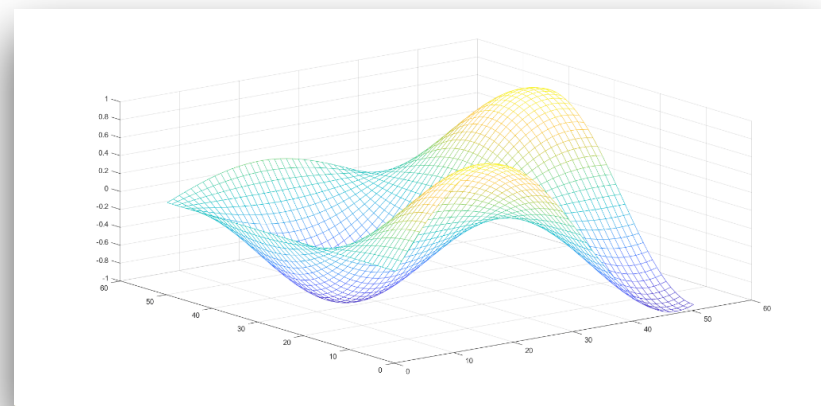


$O(16)$ observables

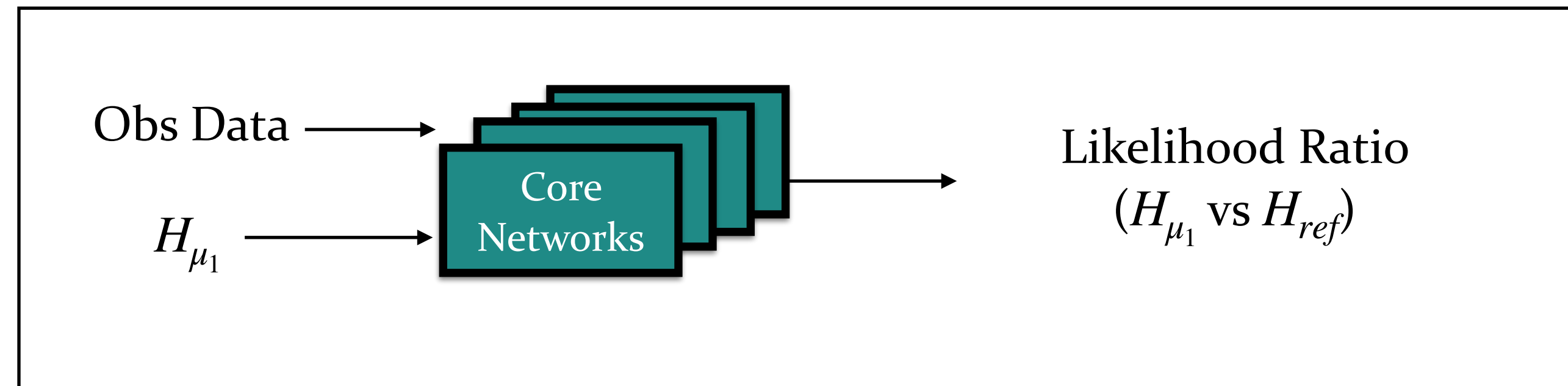


$$\frac{p(x_i | \mu, \alpha)}{p_{ref}(x_i)} = \frac{1}{\nu(\mu, \alpha)} \sum_j^C f_j(\mu) \cdot \nu_j \cdot \frac{p_j(x_i)}{p_{ref}(x_i)} \cdot \prod_k^{N_{syst}} G_j(\alpha_k) \cdot \frac{p_j(x_i, \alpha_k)}{p_j(x_i)}$$

A robust NSBI implementation

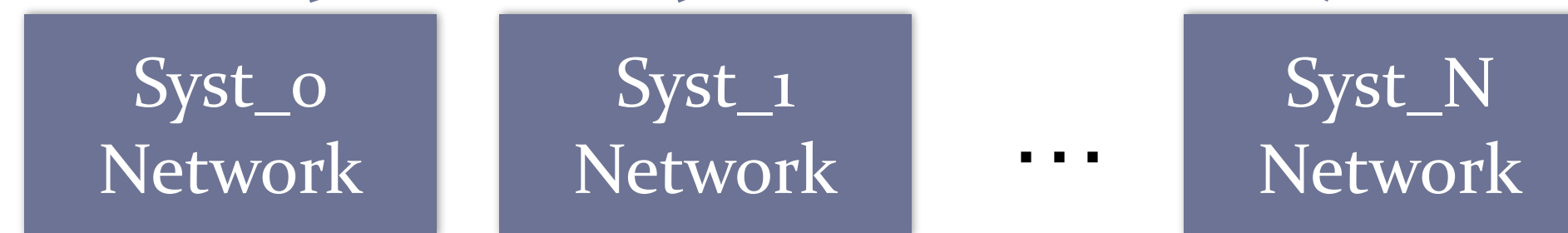
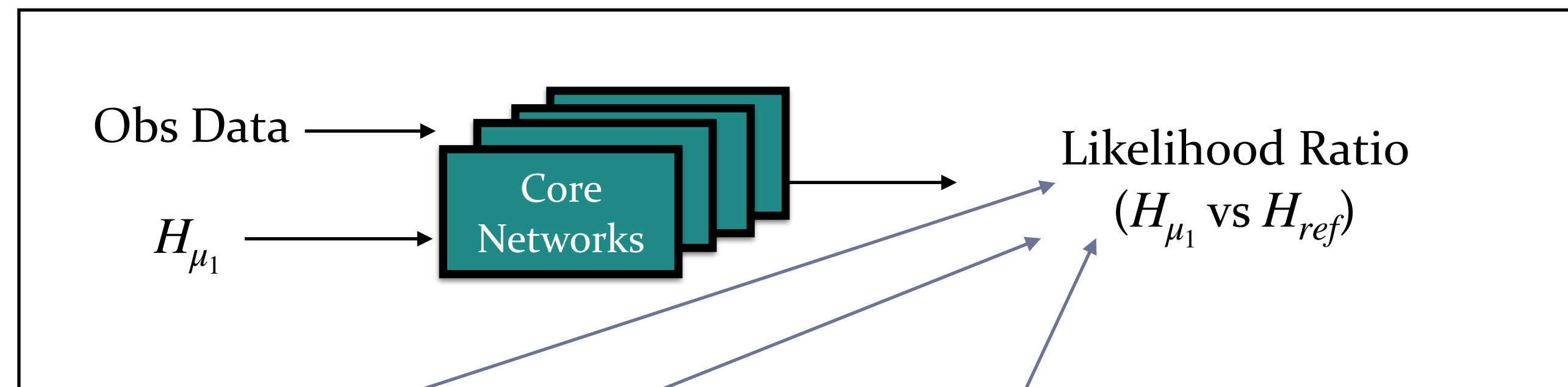


$O(16)$ observables



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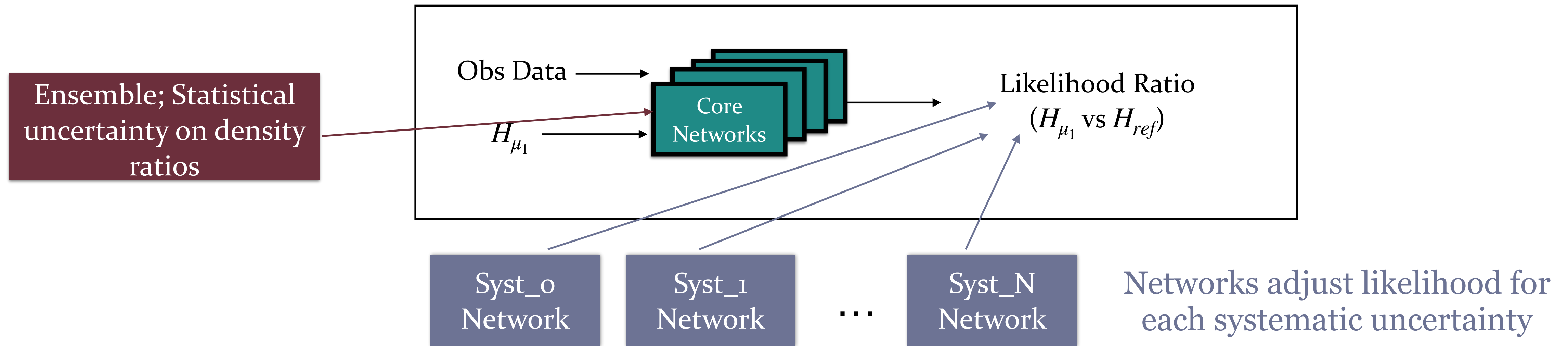
A robust NSBI implementation



Networks adjust likelihood for each systematic uncertainty

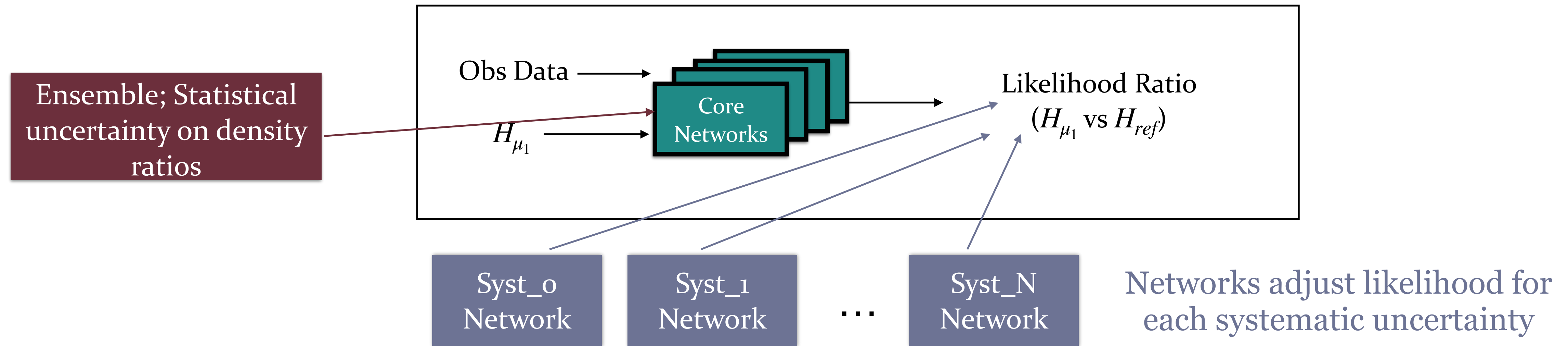
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A robust NSBI implementation



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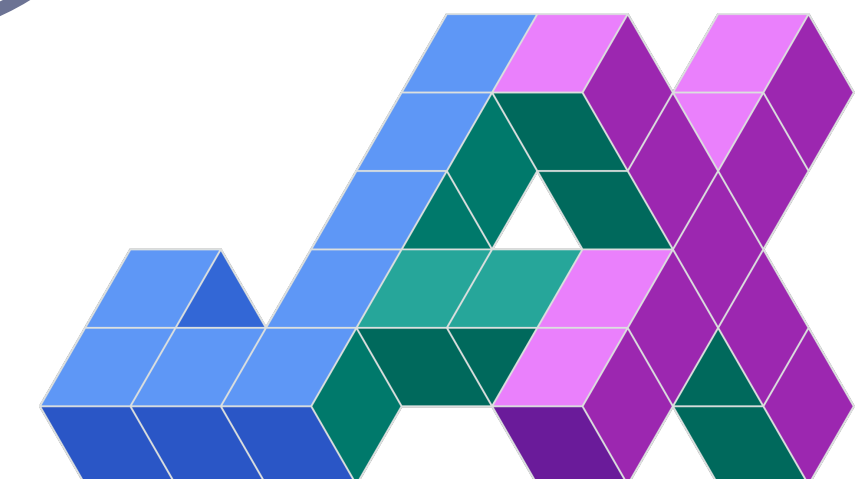
A robust NSBI implementation



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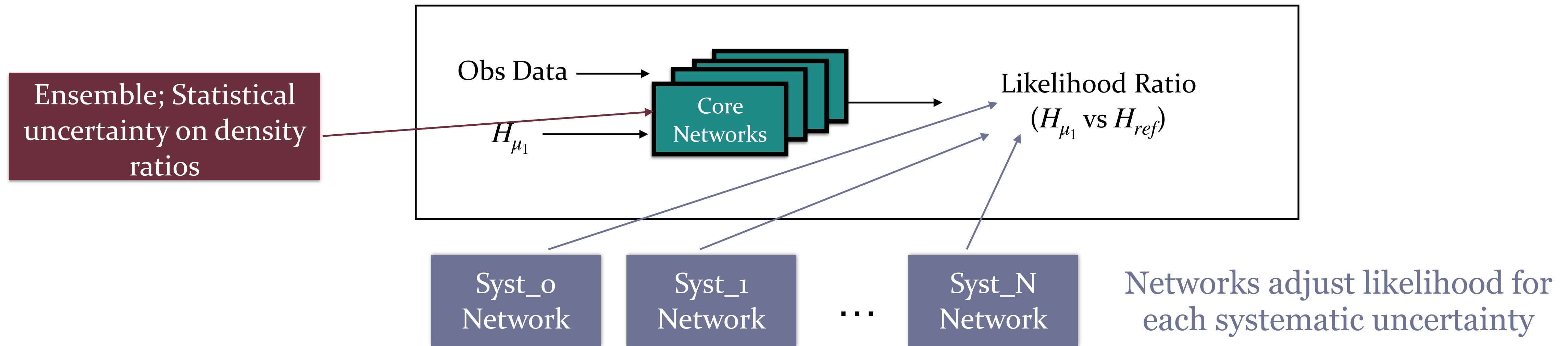
Training details

- ◆ Train $O(10^4)$ networks on TensorFlow
- ◆ Computing resources provided by Google, SMU, other HPC clusters
- ◆ Fits with JAX



A robust NSBI implementation

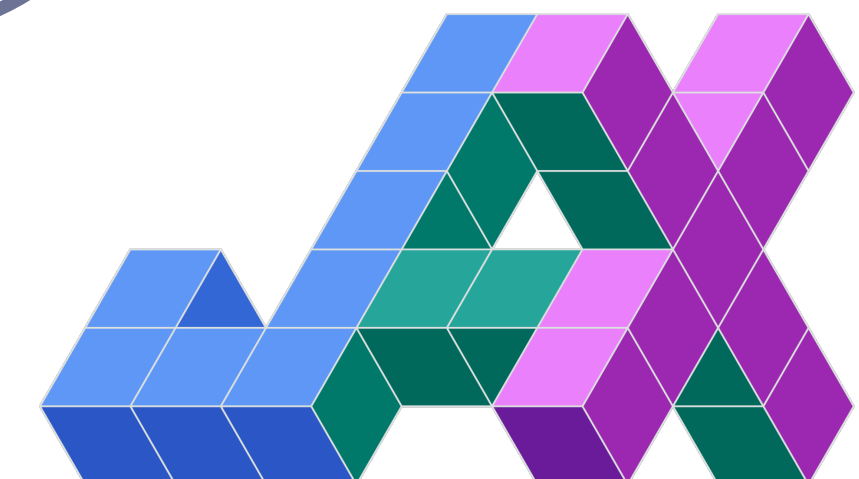
NN learns probability ratios in high dim



$$\frac{p(x_i | \mu, \alpha)}{p_{ref}(x_i)} = \frac{1}{\nu(\mu, \alpha)} \sum_j^C f_j(\mu) \cdot \nu_j \cdot \frac{p_j(x_i)}{p_{ref}(x_i)} \cdot \prod_k^{N_{syst}} G_j(\alpha_k) \cdot \frac{p_j(x_i, \alpha_k)}{p_j(x_i)}$$

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Is this dream attainable in practice?

- **How to ensure robustness? By design and validation**
- How to incorporate systematic uncertainties?
- How to perform Neyman construction in high dimensions?

Semi-parametric NSBI

x_i vector representing one individual event

General Formula

$$p(x_i|\mu) = \frac{1}{\nu(\mu)} \sum_j^{\mathcal{C}} f_j(\mu) \cdot \nu_j p_j(x_i)$$

j runs over different physics process
(Eg. $gg \rightarrow H^* \rightarrow 4l$, $gg \rightarrow ZZ \rightarrow 4l$)

Example use case

Semi-parametric NSBI

x_i vector representing one individual event General Formula

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Example use case

$$p_{\text{ggF}}(x|\mu) = \frac{1}{\nu_{\text{ggF}}(\mu)} \left[(\mu - \sqrt{\mu}) \nu_S p_S(x) + \sqrt{\mu} \nu_{\text{SBI}_1} p_{\text{SBI}_1}(x) + (1 - \sqrt{\mu}) \nu_B p_B(x) \right]$$

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Comes from theory model chosen to interpret data

Example use case

$$p_{\text{ggF}}(x|\mu) = \frac{1}{\nu_{\text{ggF}}(\mu)} \left[\underline{(\mu - \sqrt{\mu})} \nu_S p_S(x) + \underline{\sqrt{\mu}} \nu_{\text{SBI}_1} p_{\text{SBI}_1}(x) + \underline{(1 - \sqrt{\mu})} \nu_B p_B(x) \right]$$

$f_i(\mu)$ will depend on morphing bases points (which values of μ were used to simulate samples)

Semi-parametric NSBI

x_i vector representing one individual event

General Formula

$$p(x_i|\mu) = \frac{1}{v(\mu)} \sum_j^C f_j(\mu) \cdot v_j p_j(x_i)$$

Event rates estimated from simulations

Comes from theory model chosen to interpret data

j runs over different physics process
(Eg. $gg \rightarrow H^* \rightarrow 4l$, $gg \rightarrow ZZ \rightarrow 4l$)

Example use case

$$p_{\text{ggF}}(x|\mu) = \frac{1}{v_{\text{ggF}}(\mu)} \left[(\mu - \sqrt{\mu}) v_S p_S(x) + \sqrt{\mu} v_{\text{SBI}_1} p_{\text{SBI}_1}(x) + (1 - \sqrt{\mu}) v_B p_B(x) \right]$$

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Reference hypothesis j runs over different physics process
(Eg. $gg \rightarrow H^* \rightarrow 4l$, $gg \rightarrow ZZ \rightarrow 4l$)

Event rates estimated from simulations

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$$p_{\text{ggF}}(x|\mu) = \frac{1}{\nu_{\text{ggF}}(\mu)} \left[(\mu - \sqrt{\mu}) \nu_S p_S(x) + \sqrt{\mu} \nu_{\text{SBI}_1} p_{\text{SBI}_1}(x) + (1 - \sqrt{\mu}) \nu_B p_B(x) \right]$$

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Event rates estimated from simulations

Comes from theory model chosen to interpret data

Example use case

$$p_{\text{ggF}}(x|\mu) = \frac{1}{\nu_{\text{ggF}}(\mu)} \left[(\mu - \sqrt{\mu}) \nu_S p_S(x) + \sqrt{\mu} \nu_{\text{SBI}_1} p_{\text{SBI}_1}(x) + (1 - \sqrt{\mu}) \nu_B p_B(x) \right]$$

$$\frac{p(x|\mu)}{p_S(x)} = \frac{1}{\nu(\mu)} \left[(\mu - \sqrt{\mu}) \nu_S + \sqrt{\mu} \nu_{\text{SBI}_1} \frac{p_{\text{SBI}_1}(x)}{p_S(x)} + (1 - \sqrt{\mu}) \nu_B \frac{p_B(x)}{p_S(x)} \right]$$

$f_i(\mu)$ will depend on morphing bases points (which values of μ were used to simulate samples)

Semi-parametric NSBI

x_i vector representing one individual event

General Formula

$$p(x_i|\mu) = \frac{1}{\nu(\mu)} \sum_j^C f_j(\mu) \cdot \nu_j p_j(x_i) \quad ? \quad \xrightarrow{\text{Estimated using an ensemble of networks}} \quad \frac{p(x_i|\mu)}{p_{\text{ref}}(x_i)} = \frac{1}{\nu(\mu)} \sum_j^C f_j(\mu) \cdot \nu_j \frac{p_j(x_i)}{p_{\text{ref}}(x_i)}$$

Event rates estimated from simulations

Reference hypothesis

j runs over different physics process
(Eg. $gg \rightarrow H^* \rightarrow 4l, gg \rightarrow ZZ \rightarrow 4l$)

Comes from theory model chosen to interpret data

Example use case

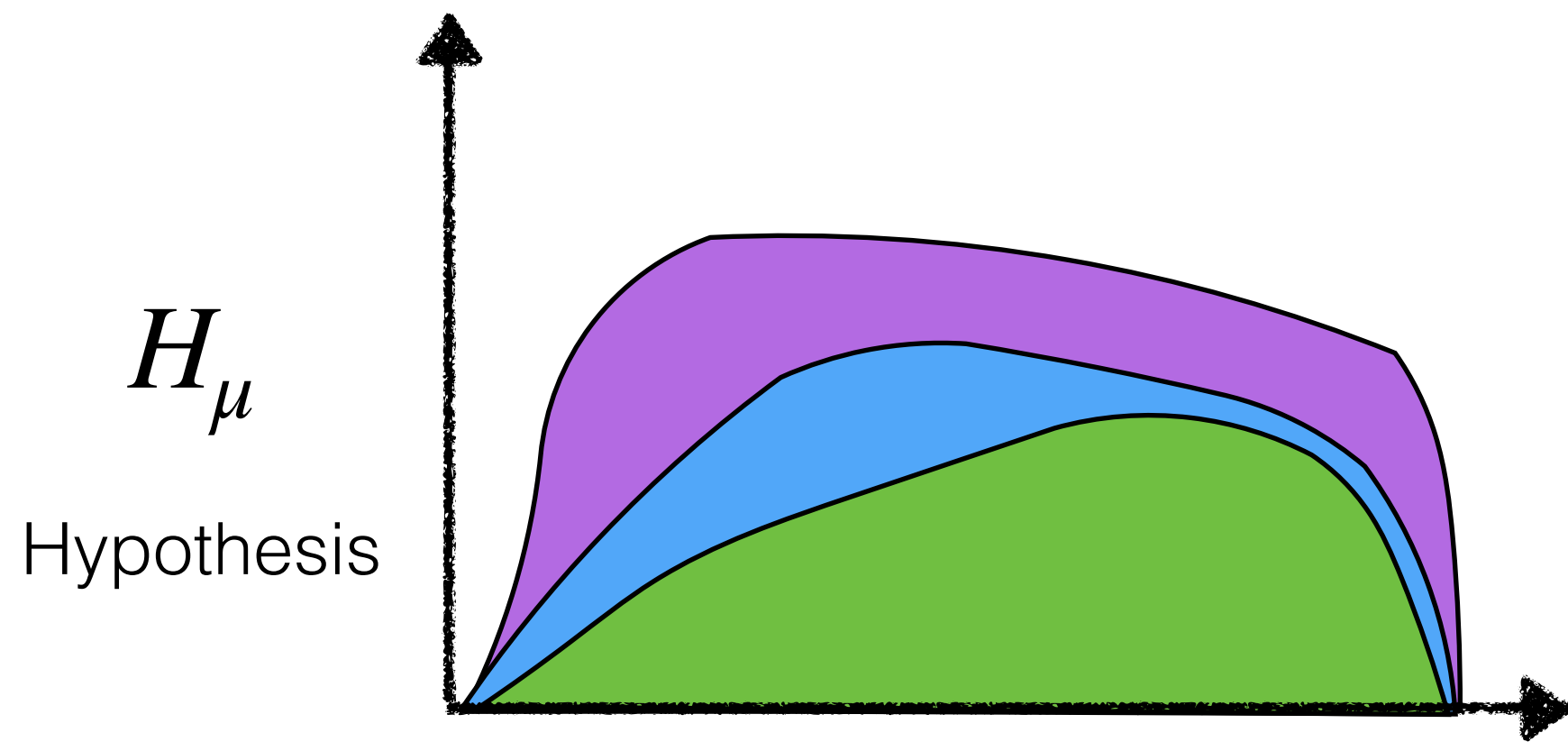
$$p_{\text{ggF}}(x|\mu) = \frac{1}{\nu_{\text{ggF}}(\mu)} \left[(\mu - \sqrt{\mu}) \nu_S p_S(x) + \sqrt{\mu} \nu_{\text{SBI}_1} p_{\text{SBI}_1}(x) + (1 - \sqrt{\mu}) \nu_B p_B(x) \right]$$

$$\xrightarrow{\quad} \quad \frac{p(x|\mu)}{p_S(x)} = \frac{1}{\nu(\mu)} \left[(\mu - \sqrt{\mu}) \nu_S + \sqrt{\mu} \nu_{\text{SBI}_1} \frac{p_{\text{SBI}_1}(x)}{p_S(x)} + (1 - \sqrt{\mu}) \nu_B \frac{p_B(x)}{p_S(x)} \right]$$

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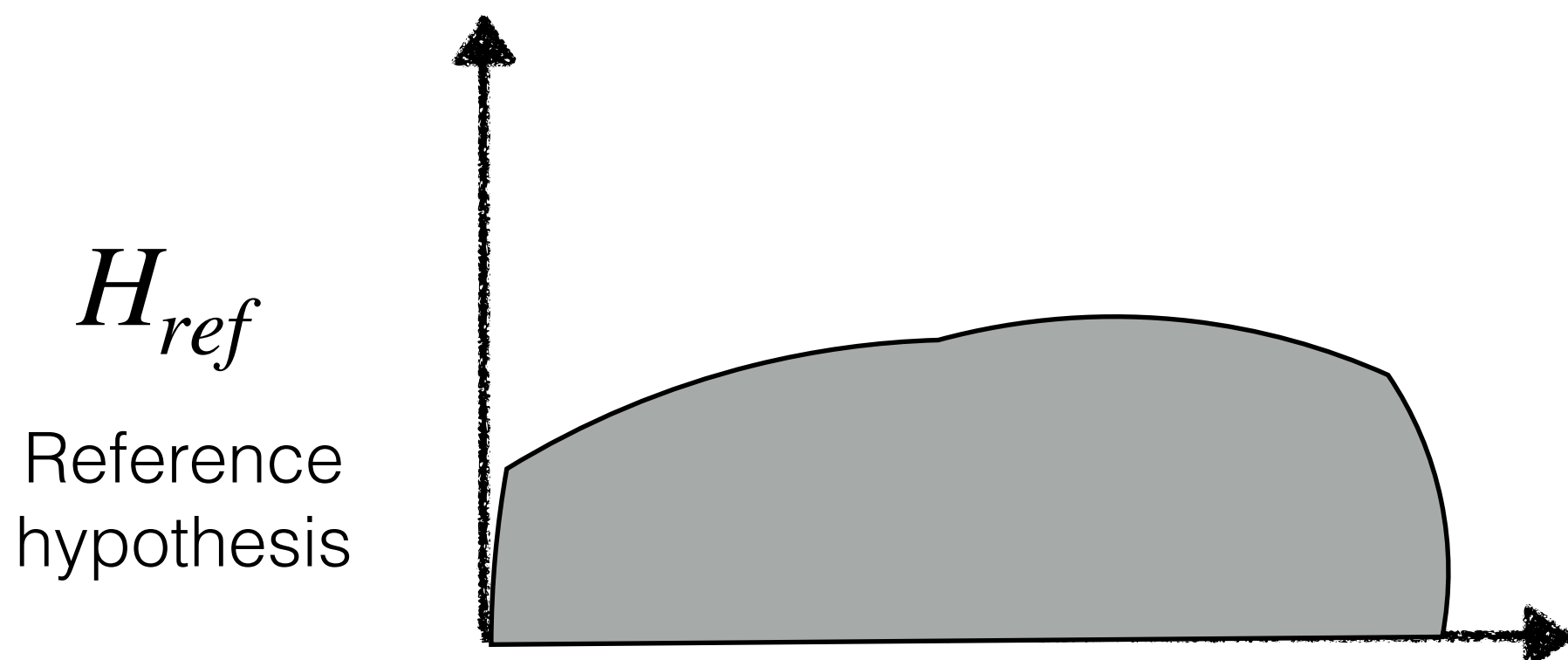
Robust, parameterised classifier without parameterising

H_{ref} : Reference hypothesis



$$\frac{p(x_i|\mu)}{p_{ref}(x_i)} = \frac{1}{v(\mu)} \sum_j^C f_j(\mu) \cdot v_j \frac{p_j(x_i)}{p_{ref}(x_i)}$$

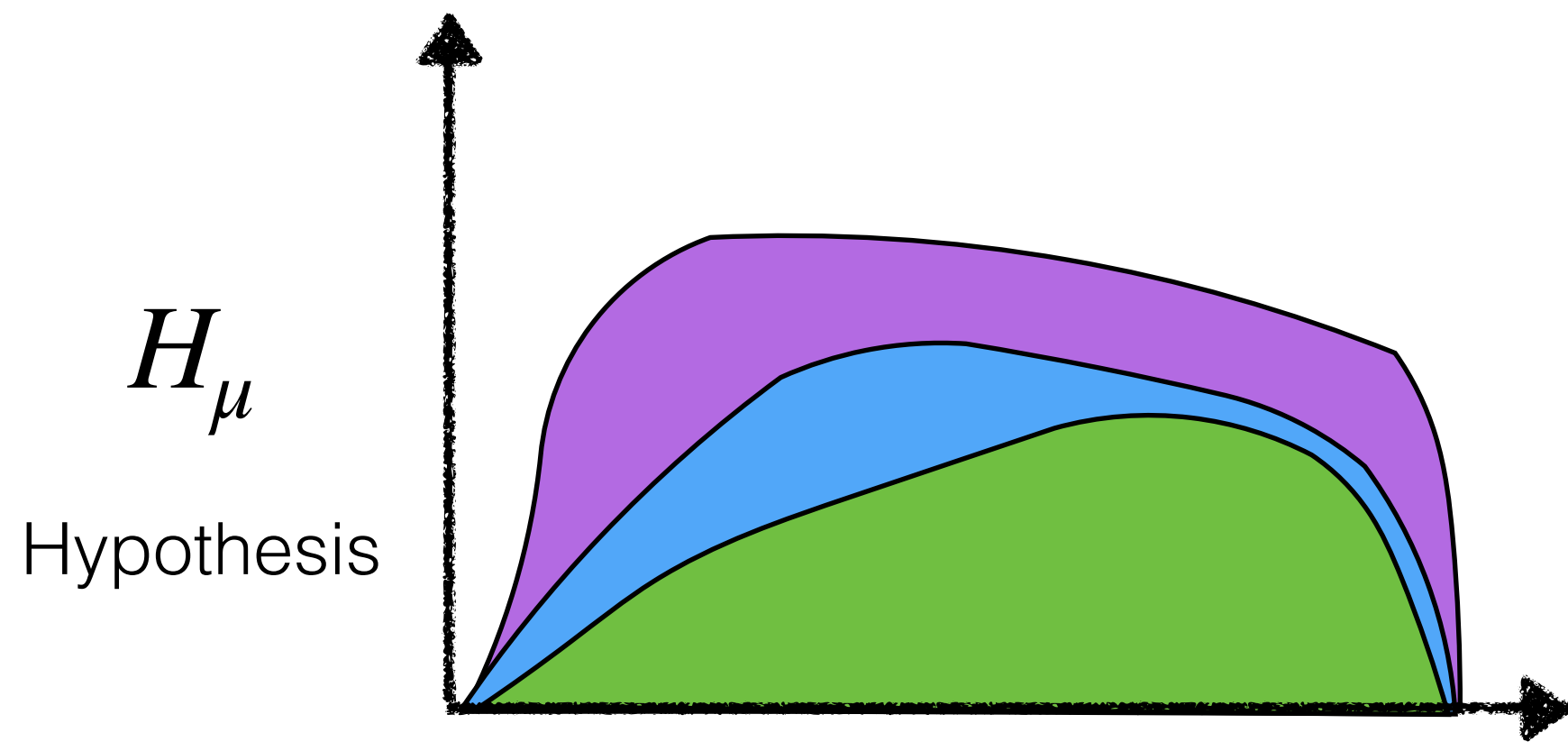
VS



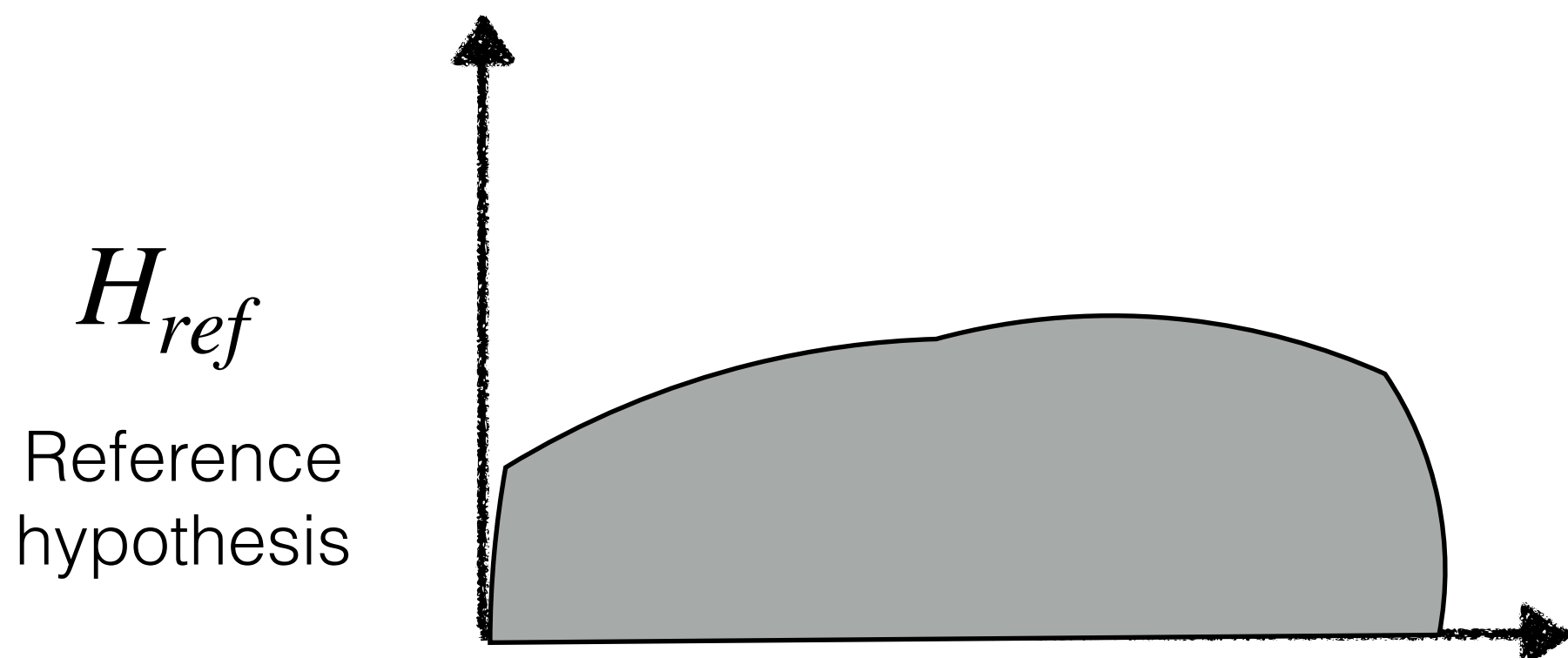
A separate classifier per physics process j
(Eg. $gg \rightarrow H^* \rightarrow 4l$, $gg \rightarrow ZZ \rightarrow 4l$)

Robust, parameterised classifier without parameterising

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VS

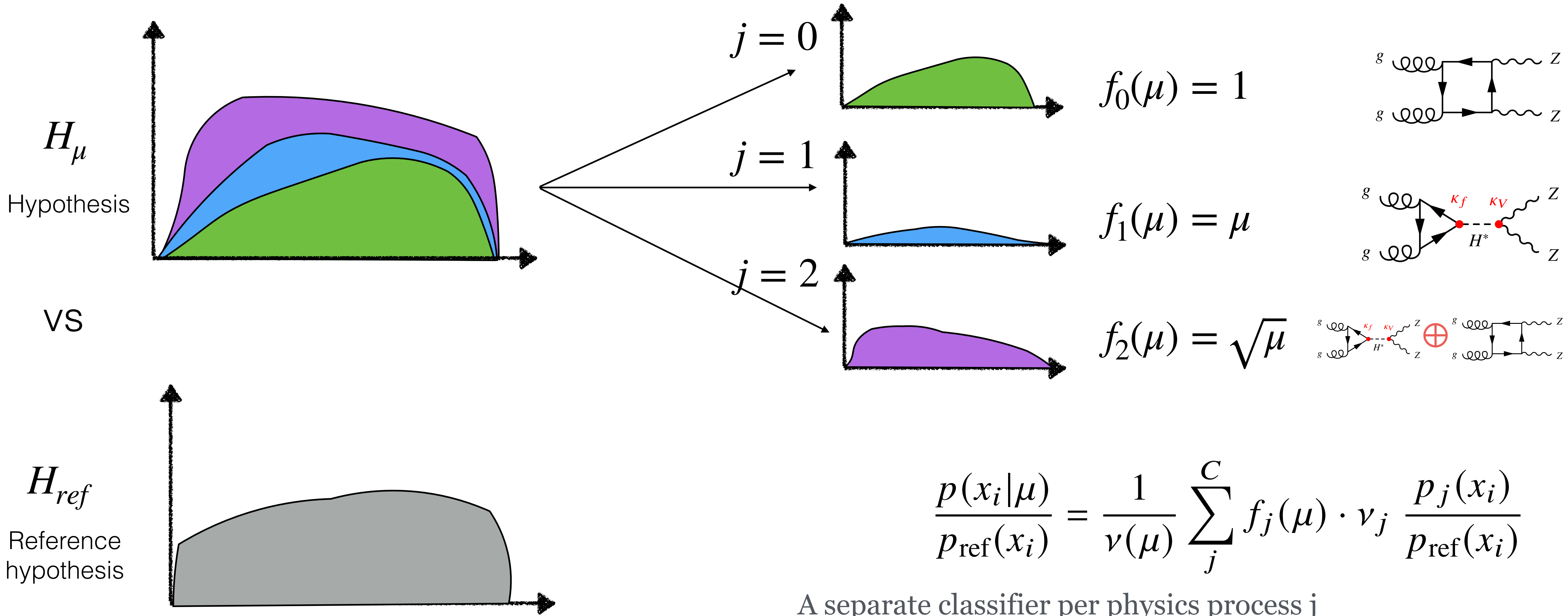


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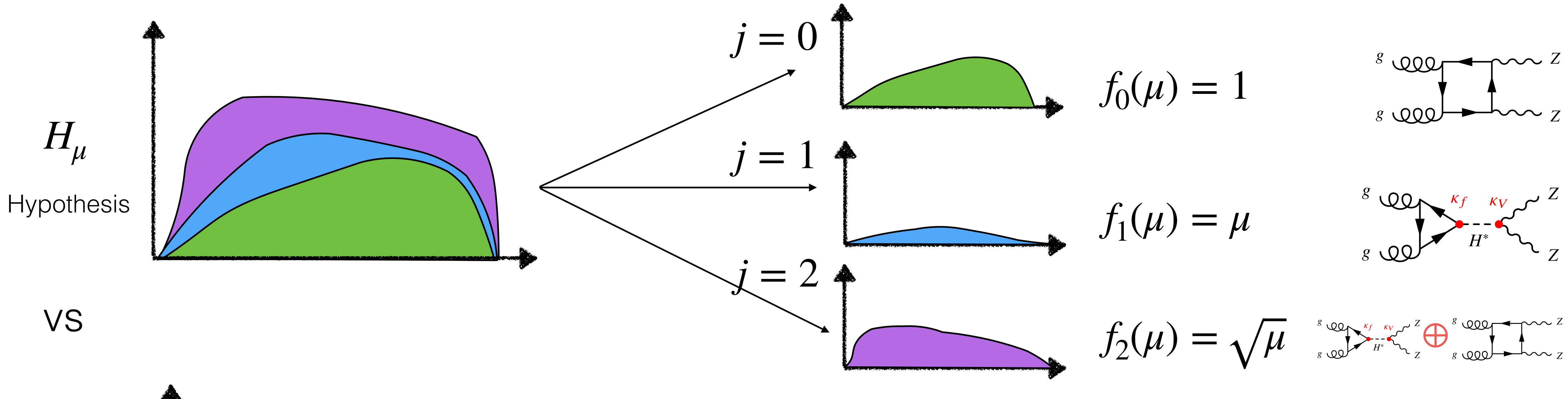
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A separate classifier per physics process j
(Eg. $gg \rightarrow H^* \rightarrow 4l$, $gg \rightarrow ZZ \rightarrow 4l$)

Robust, parameterised classifier without parameterising

H_{ref} : Reference hypothesis



VS



Analytically parameterised in μ , allows to get LR for any hypothesis μ without training parameterised networks!

$$\frac{p(x_i|\mu)}{p_{ref}(x_i)} = \frac{1}{v(\mu)} \sum_j^C f_j(\mu) \cdot v_j \frac{p_j(x_i)}{p_{ref}(x_i)}$$

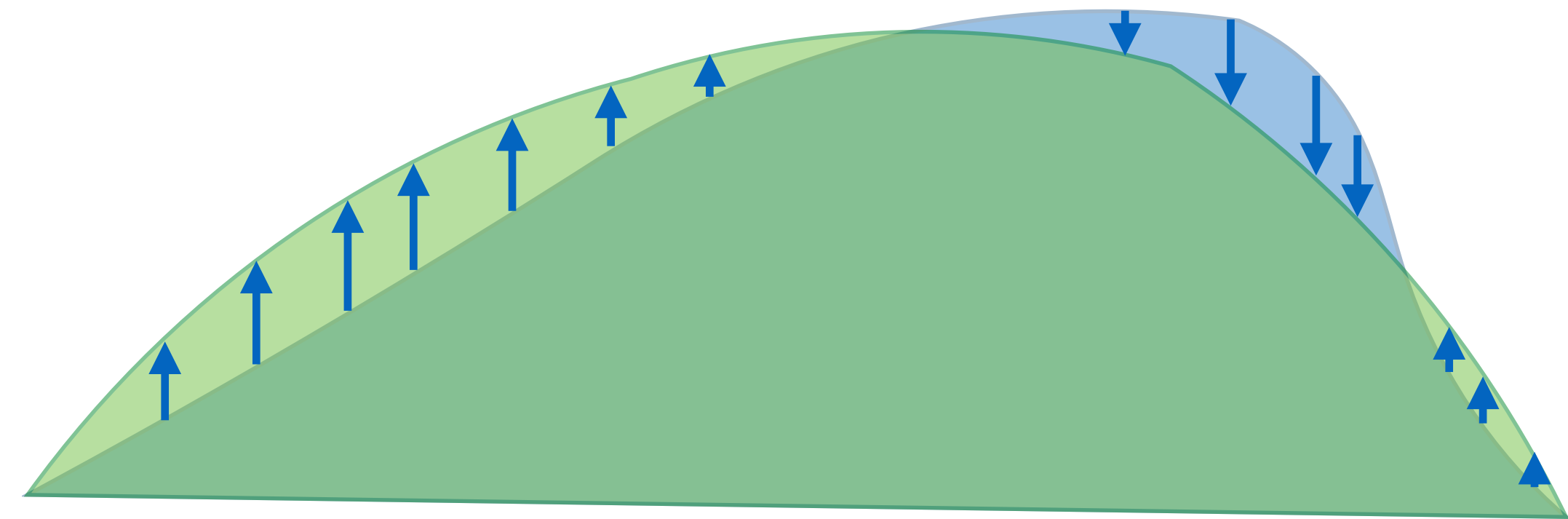
A separate classifier per physics process j
(Eg. $gg \rightarrow H^* \rightarrow 4l, gg \rightarrow ZZ \rightarrow 4l$)

Is this dream attainable in practice?

- **How to ensure robustness? By design and validation**
- How to incorporate systematic uncertainties?
- How to perform Neyman construction in high dimensions?

Validate quality of LR estimation with re-weighting task

Reweighting: Calculate weights w_i for events x_i in **blue sample** to match **green sample**

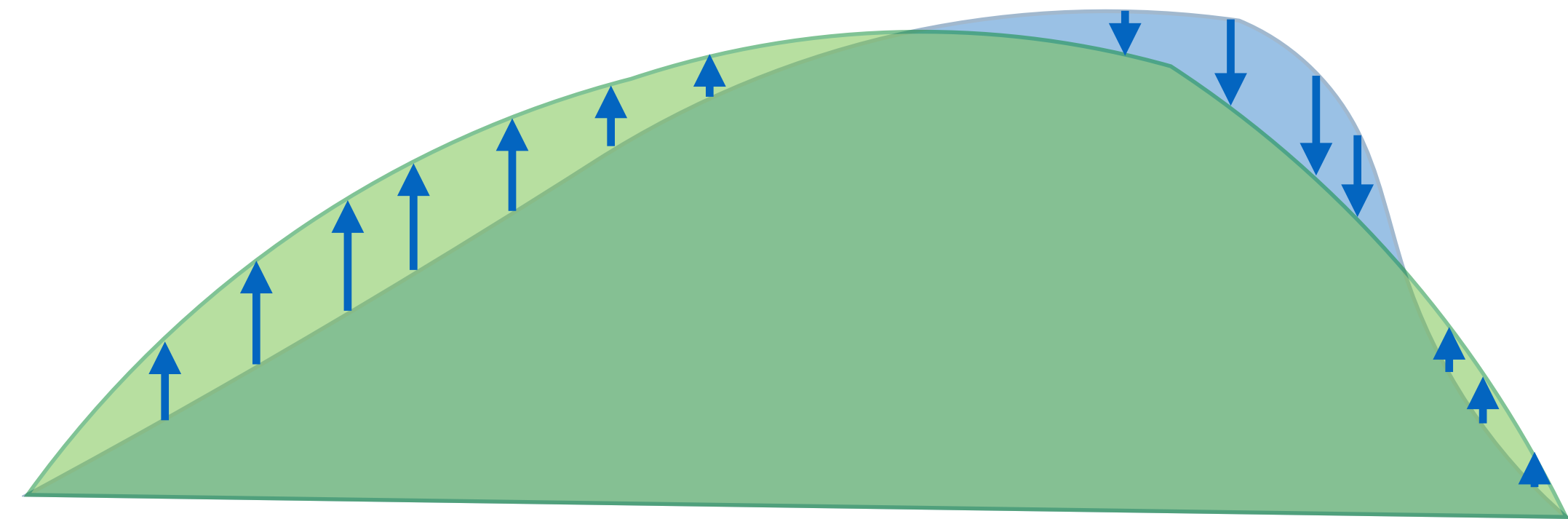


Validate quality of LR estimation with re-weighting task

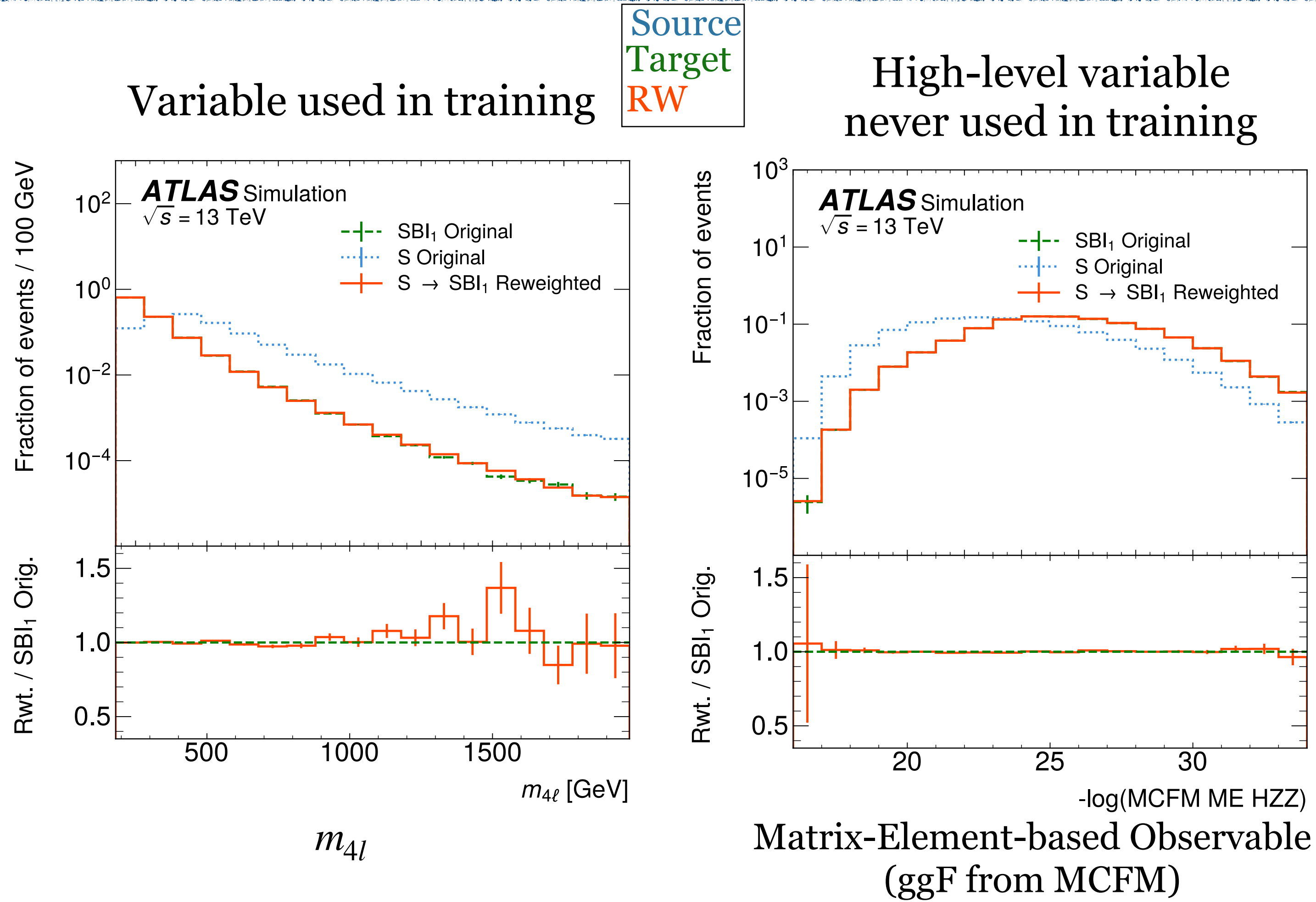
Reweighting: Calculate weights w_i for events x_i in **blue sample** to match **green sample**

$$w_i = r(x_i, \mu_0, \mu_1) = \frac{p(x_i | \mu_0)}{p(x_i | \mu_1)}$$

Probability ratios already estimated using an ensemble of networks



Re-weight closures

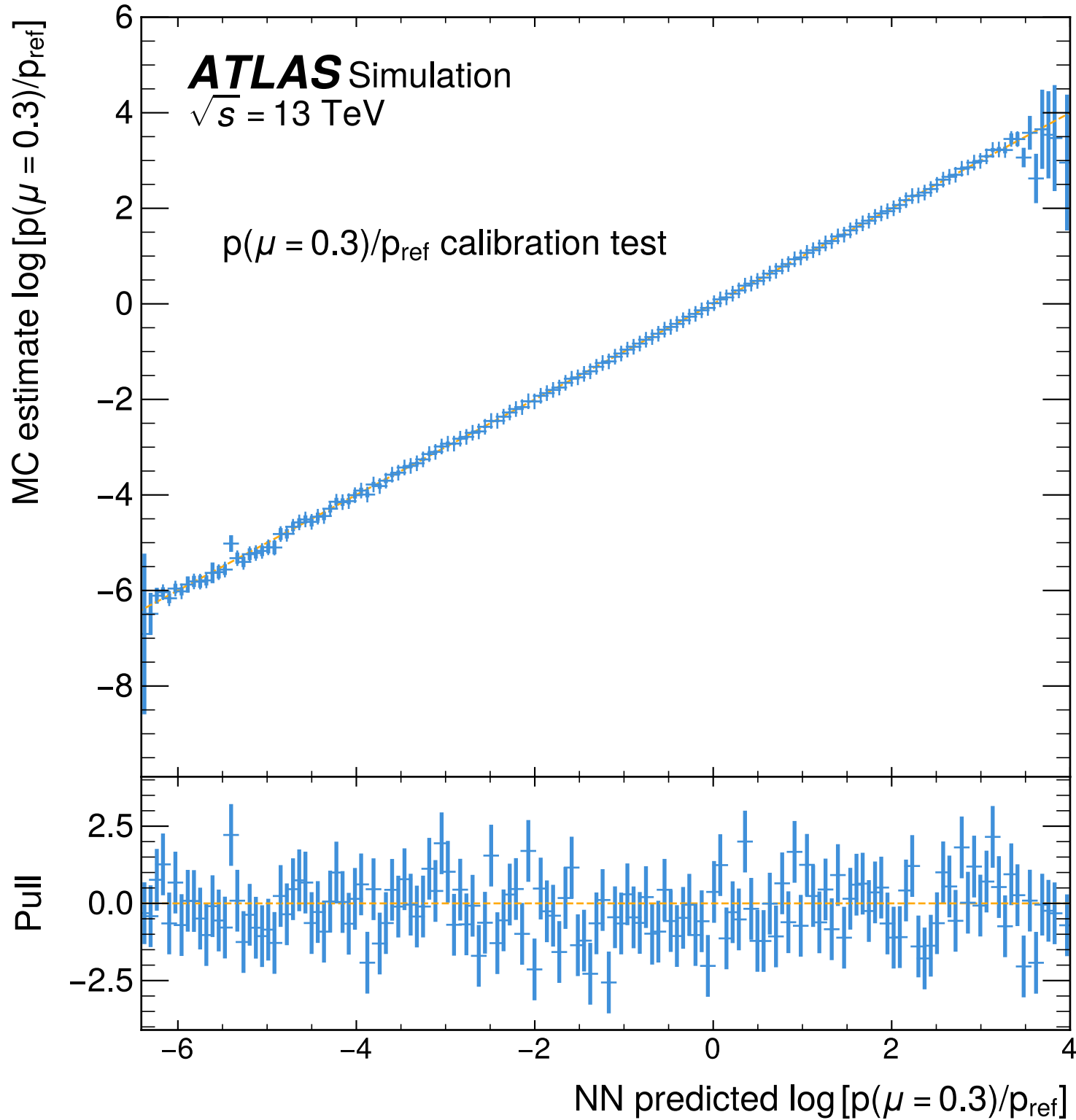


Calibration curves of probability density ratios

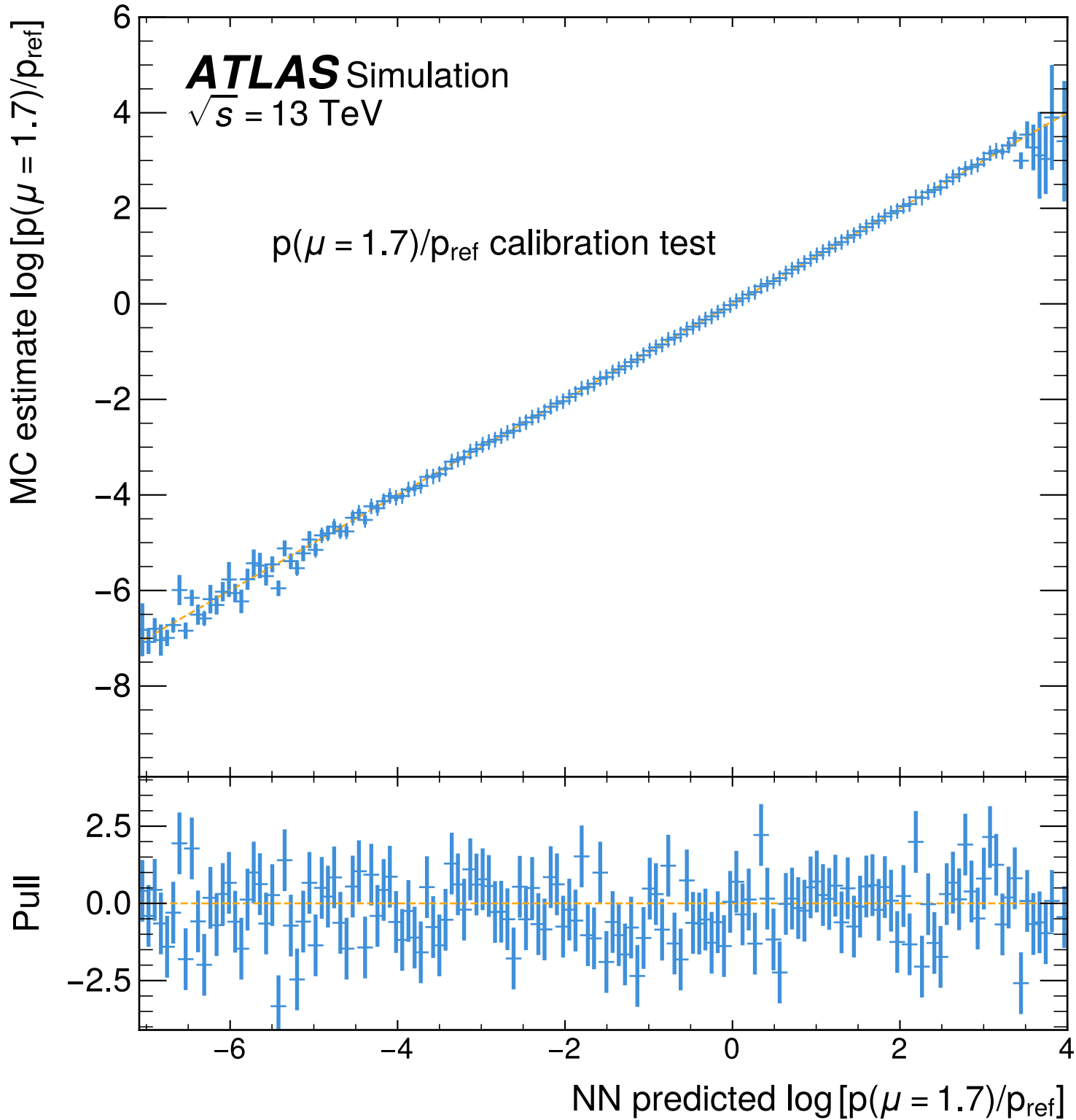
$$\frac{P_{\mu=0.3}(x_i)}{P_{ref}(x_i)}$$

$$\frac{P_{\mu=1.7}(x_i)}{P_{ref}(x_i)}$$

Binned estimate



Ensemble prediction

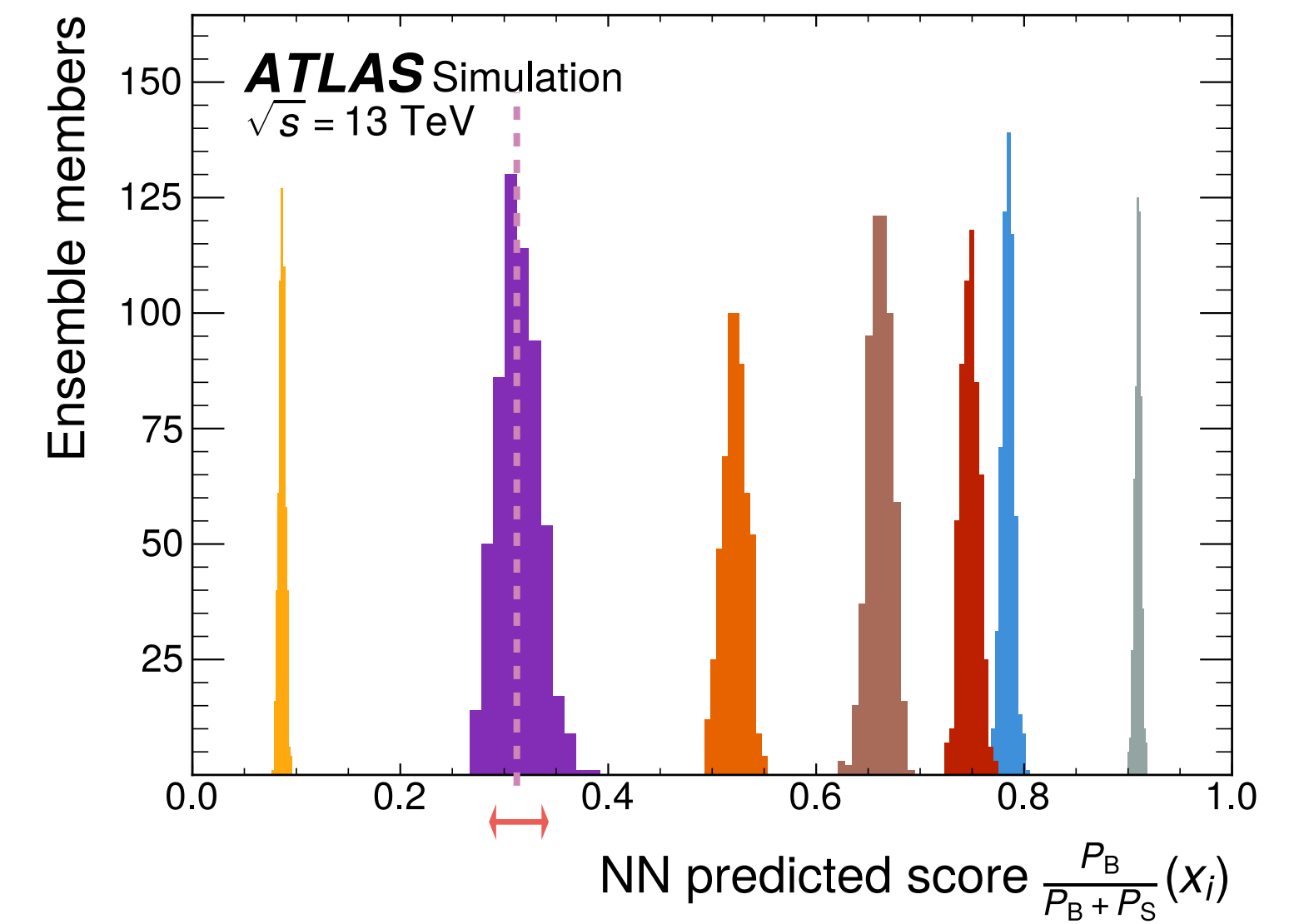


Ensemble prediction

Perfect calibration would give $y = x$

Uncertainties on the network estimates themselves

Distribution of NN predictions for example events



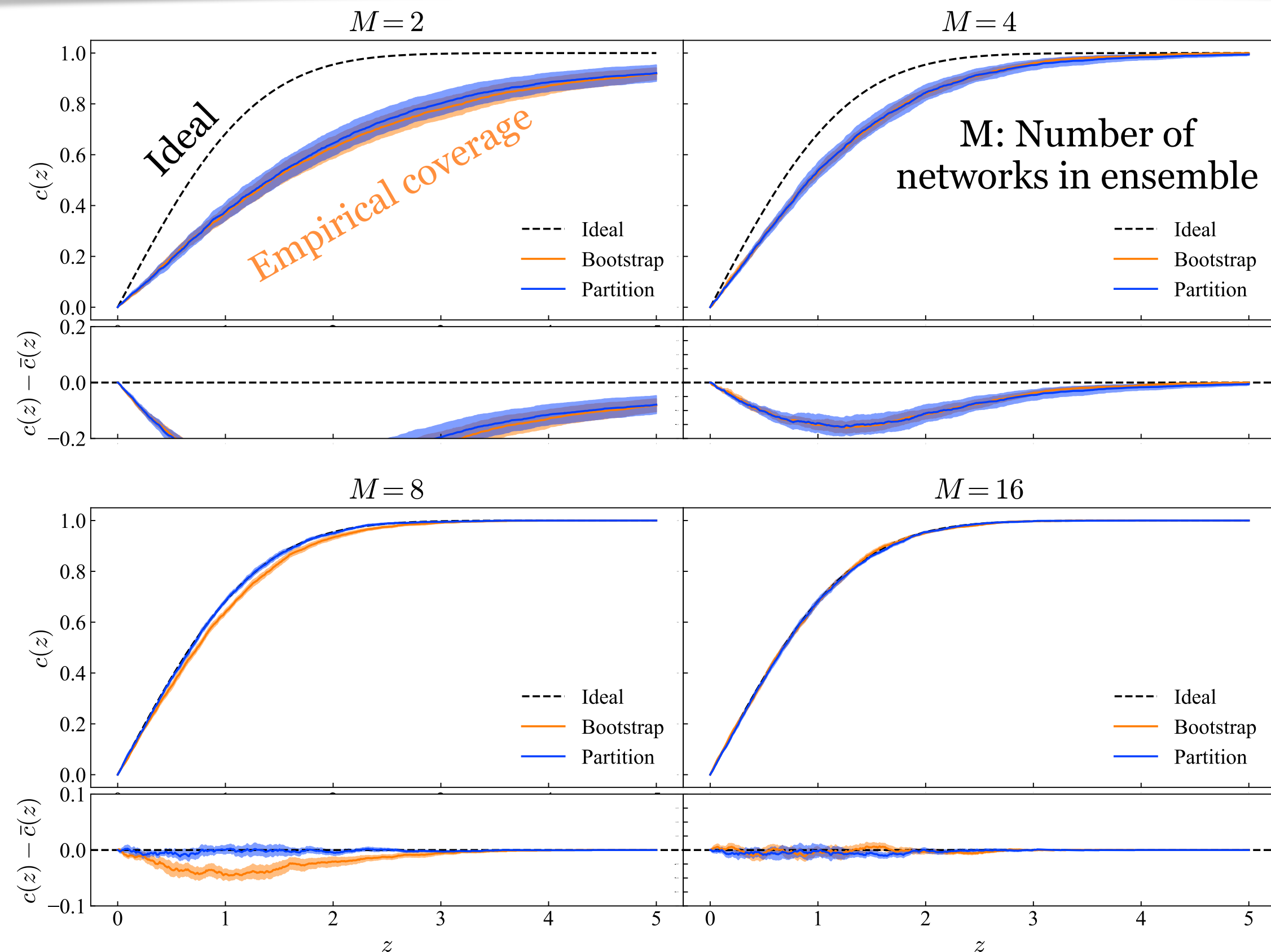
Uncertainties on the network estimates themselves

Frequentist Uncertainties on Neural Density Ratios with $w_i f_i$ Ensembles

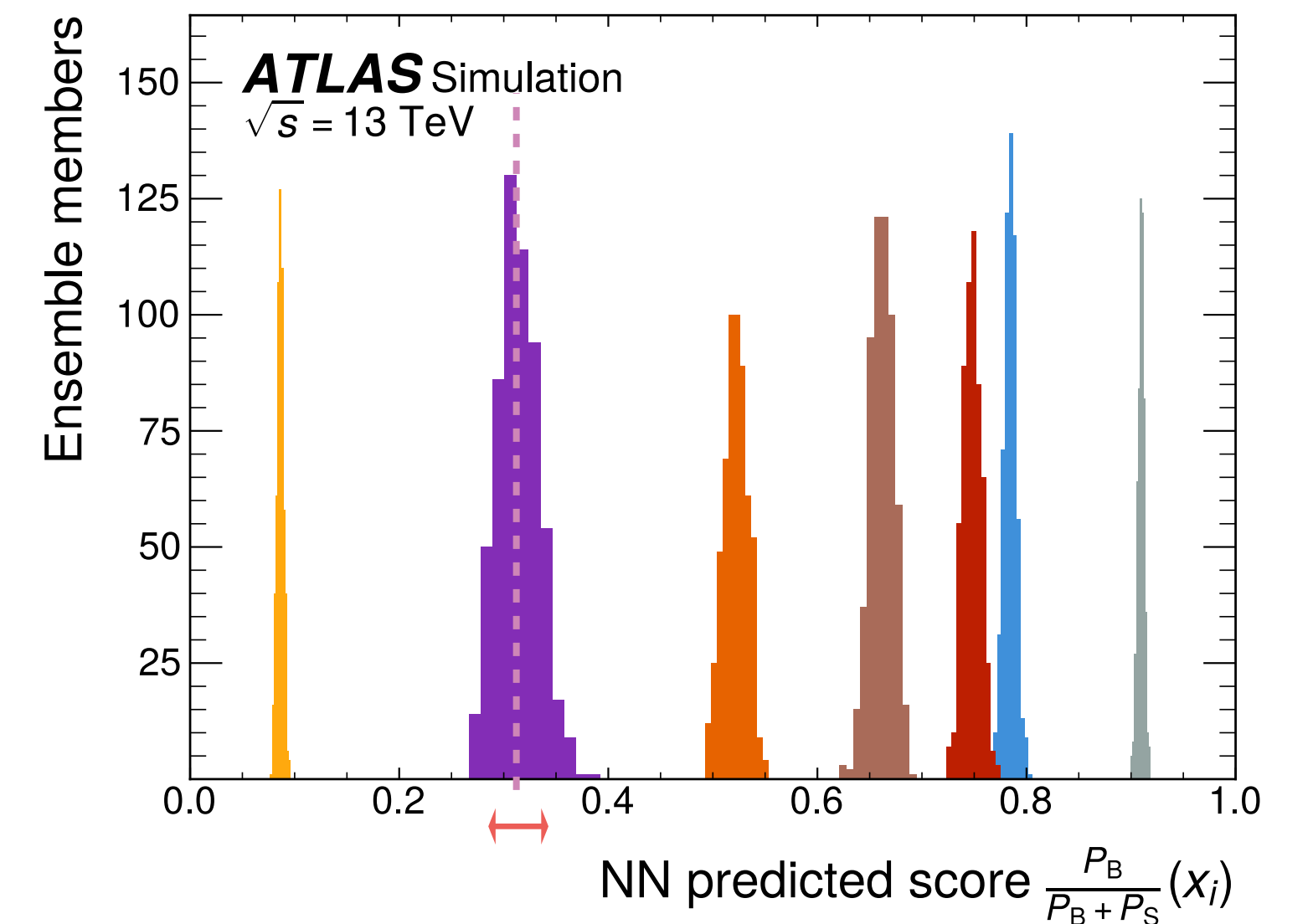
Sean Benevedes^{1,2,*} and Jesse Thaler^{1,2,†}

¹Center for Theoretical Physics, Massachusetts Institute of Technology,
Cambridge, Massachusetts, United States

²The NSF AI Institute for Artificial Intelligence and Fundamental Interactions



Distribution of NN predictions for example events

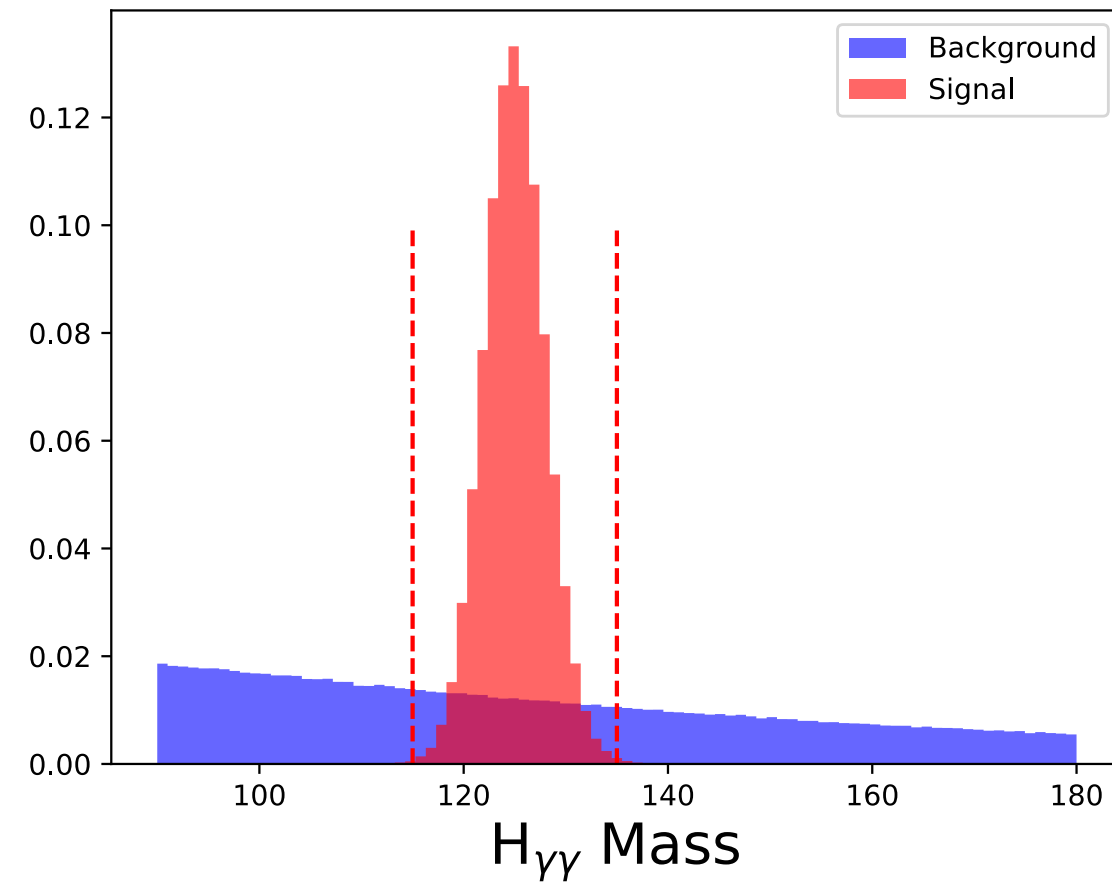


- Genuine 95% coverage, with much smaller ensembles
- Elegant, mathematically motivated method to estimate uncertainties

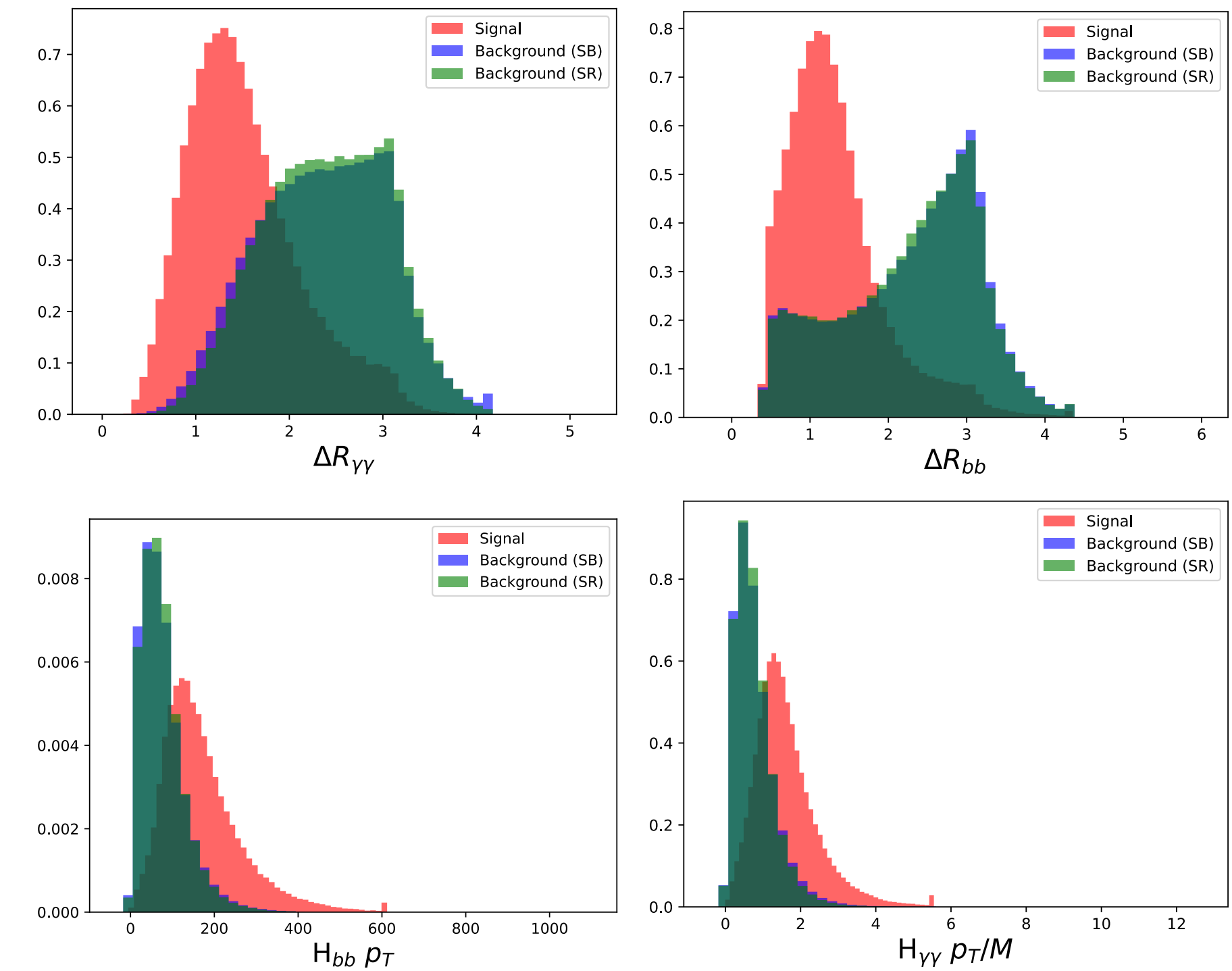
[arXiv:2506.00113](https://arxiv.org/abs/2506.00113): Benevedes & Thaler



Unreliable simulator ?



- Training data can come from control regions of real data
- **Data-driven background estimation** techniques work with NSBI
- Amram & Szwec uses generative models instead of classifiers for NSBI



Is there more to simulation-based inference?

So much more! Diagnostics, Neyman construction, focusing the power of a test...
Can discuss if there is time.

But, let's focus on the basics in today's tutorial

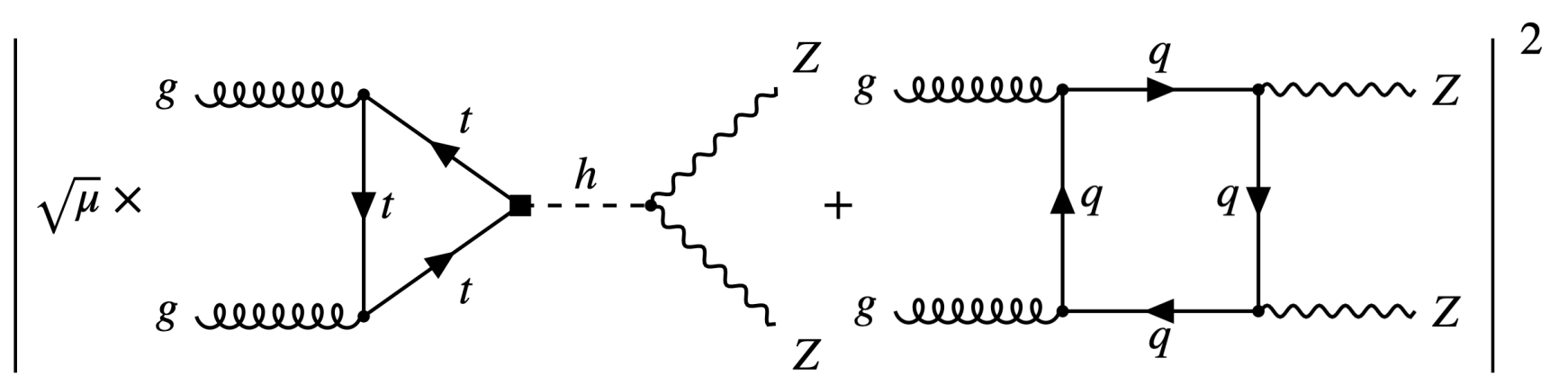
This tutorial (by Eddie McGrady)

Based on dataset and code from **Tae Hyoun Park**, originally for [arXiv:2507.02032](https://arxiv.org/abs/2507.02032)

- Simplified dataset:
 - Physics processes:
 - S: $gg \rightarrow H^* \rightarrow 4l$
 - B: $gg \rightarrow ZZ \rightarrow 4l$
 - Combined SBI
 - No dominant qqZZ background, no EW processes
 - No systematics

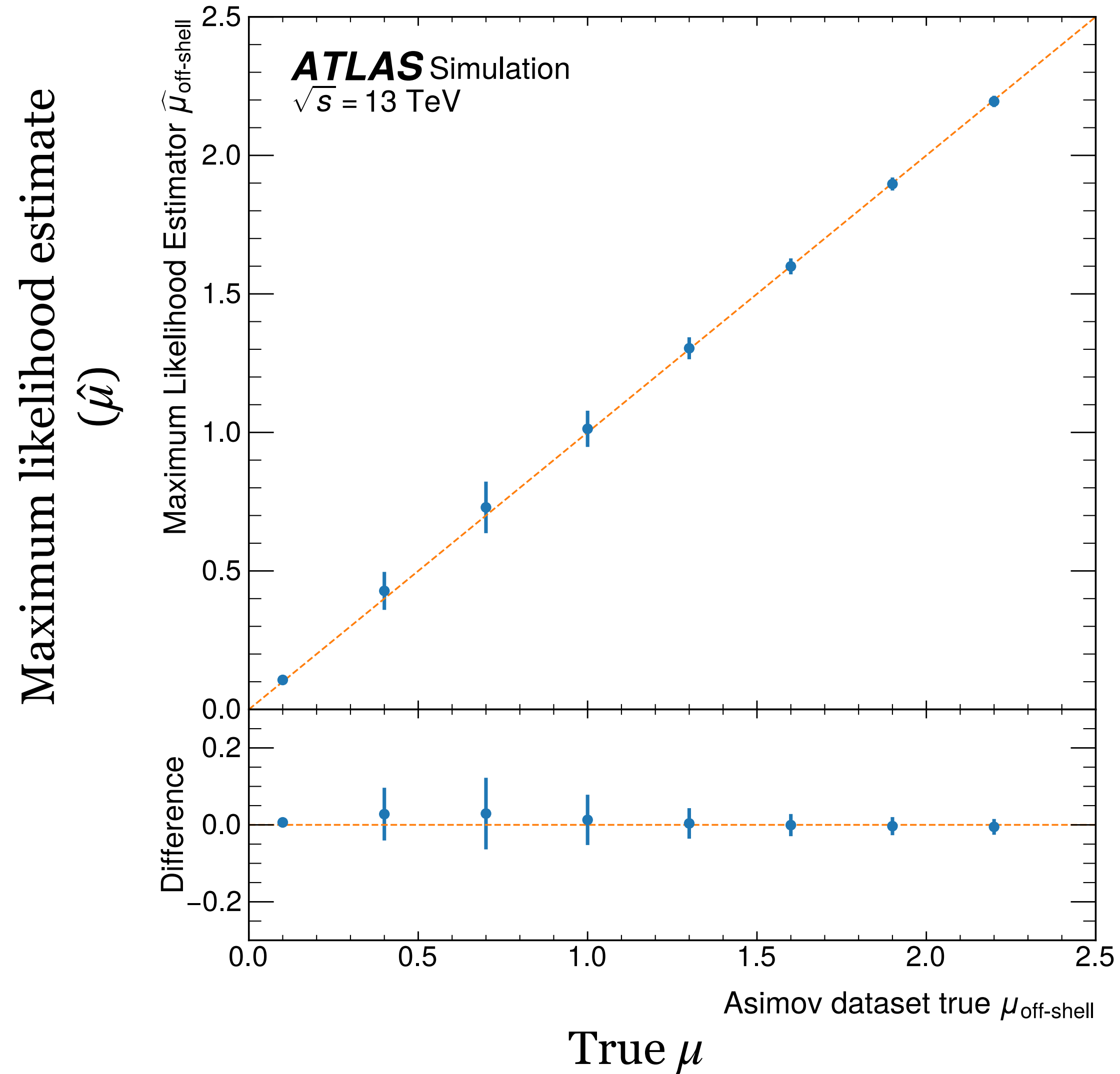
• The background sample will be used as 'reference', so $\frac{P_B}{P_{ref}} = 1$

• Only 2 density ratios to be estimated: $\frac{P_S}{P_{ref}}$ & $\frac{P_{SBI}}{P_{ref}}$

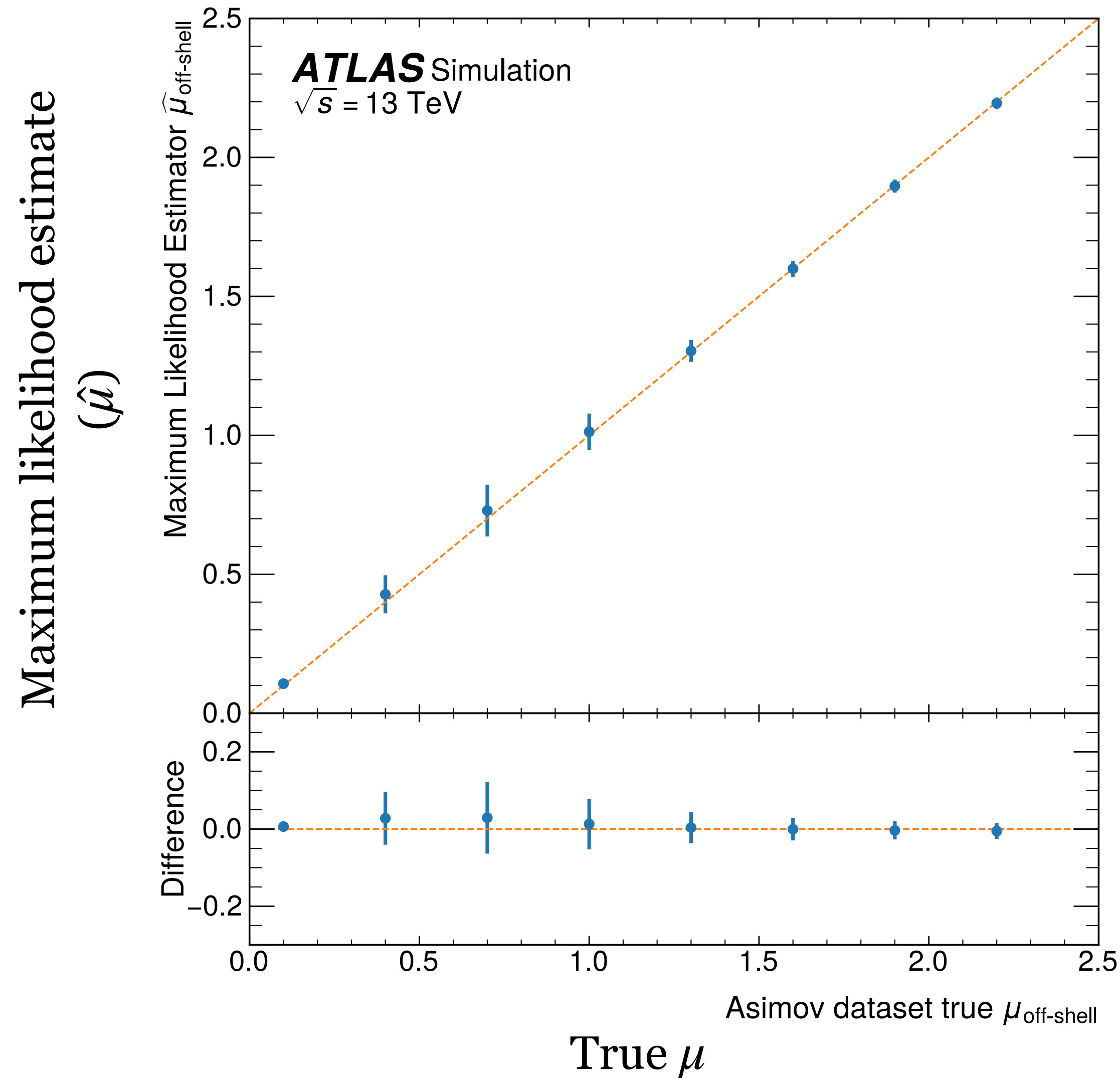


Let's play with the notebooks !

More diagnostics

Testing full analysis on samples from different values of μ 

Testing full analysis on samples from different values of μ



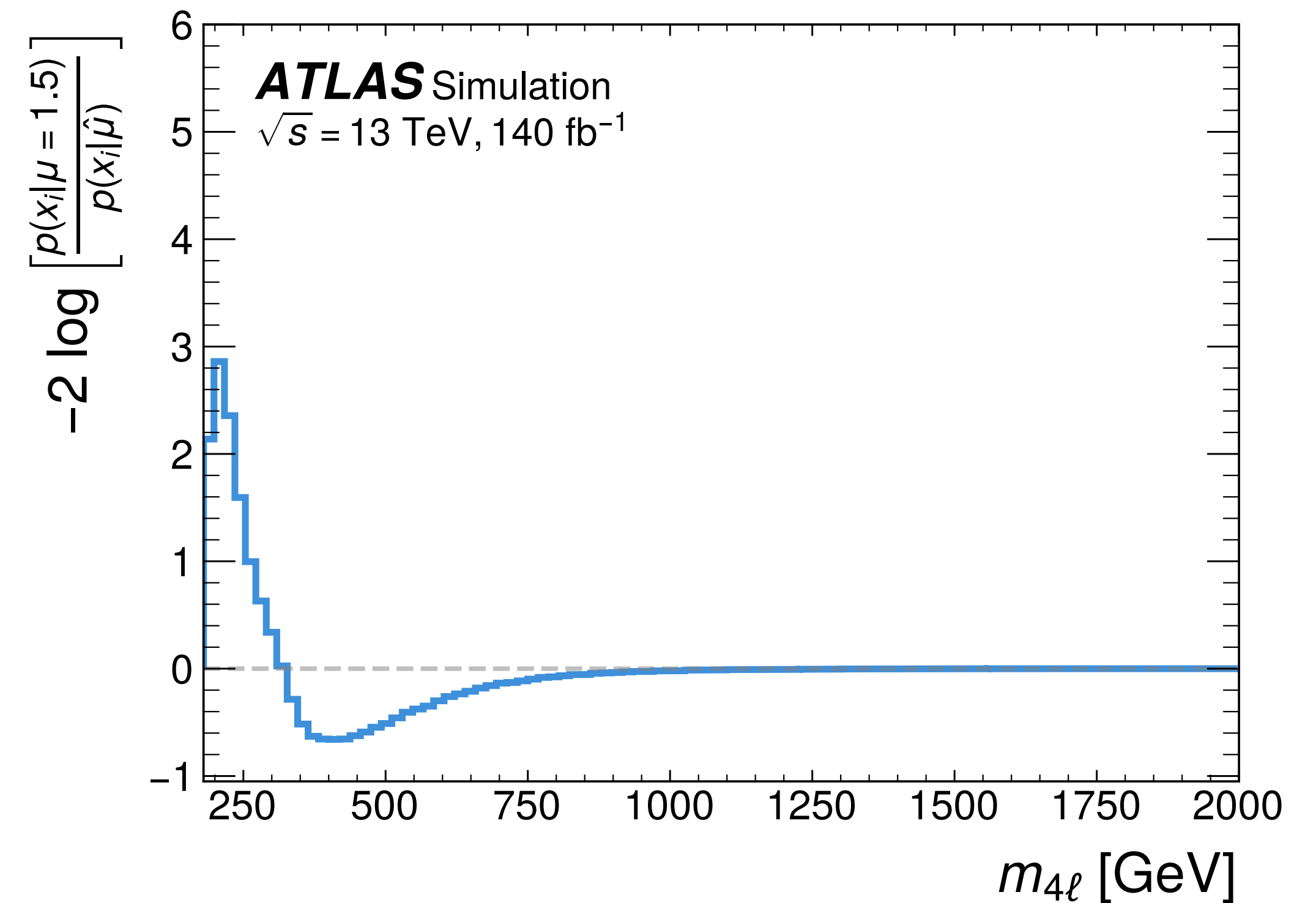
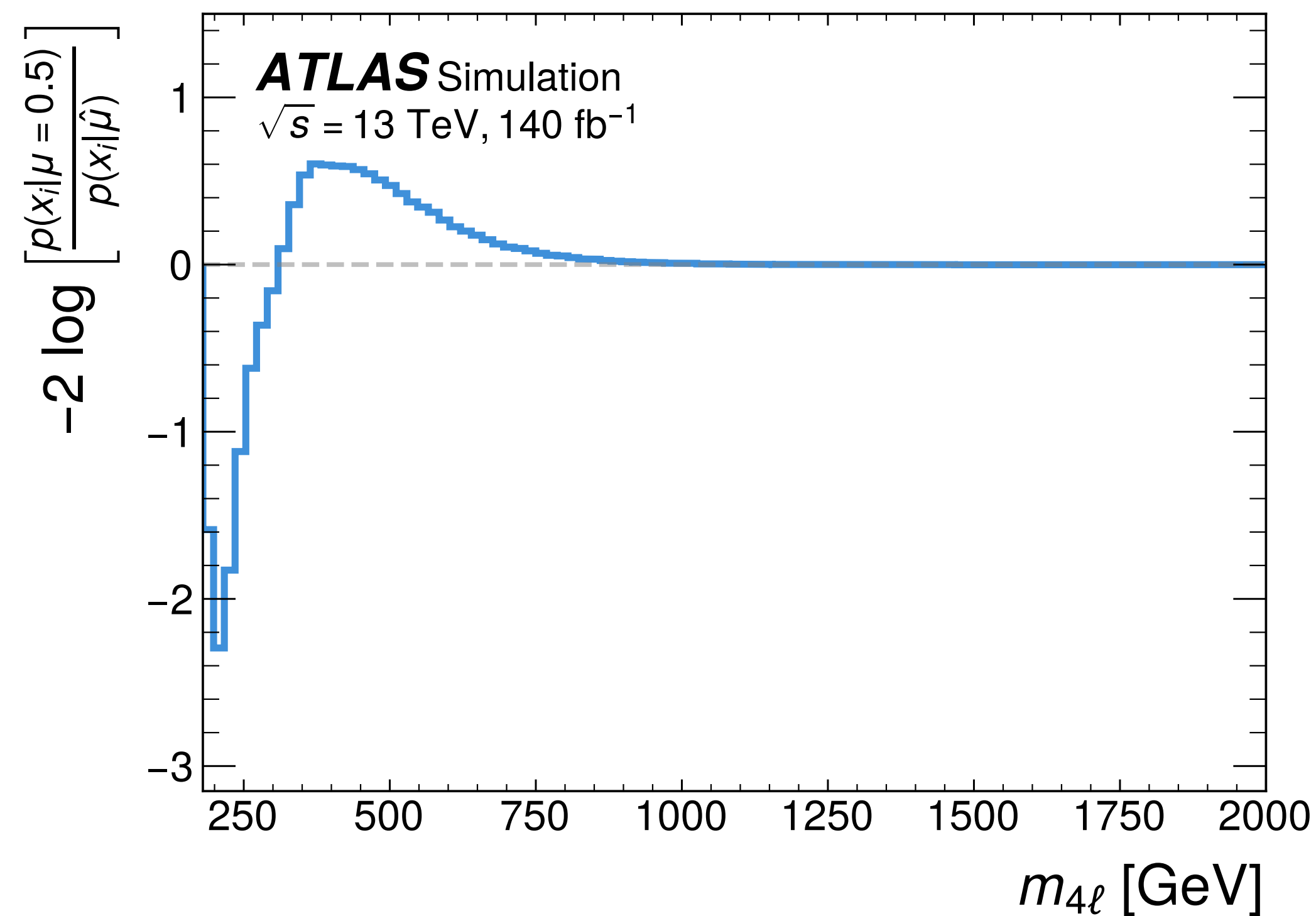
No bias: Method recovers correct value of μ on average
(Correct MLE when tested on the median ‘Asimov dataset’)

And many more diagnostics, data vs MC validation
(see [backup](#))

Interpretability: Which phase space favours one hypothesis over another?

$$-2 \cdot \log \frac{P(x_i | \mu = 0.5)}{P(x_i | \mu = 1)}$$

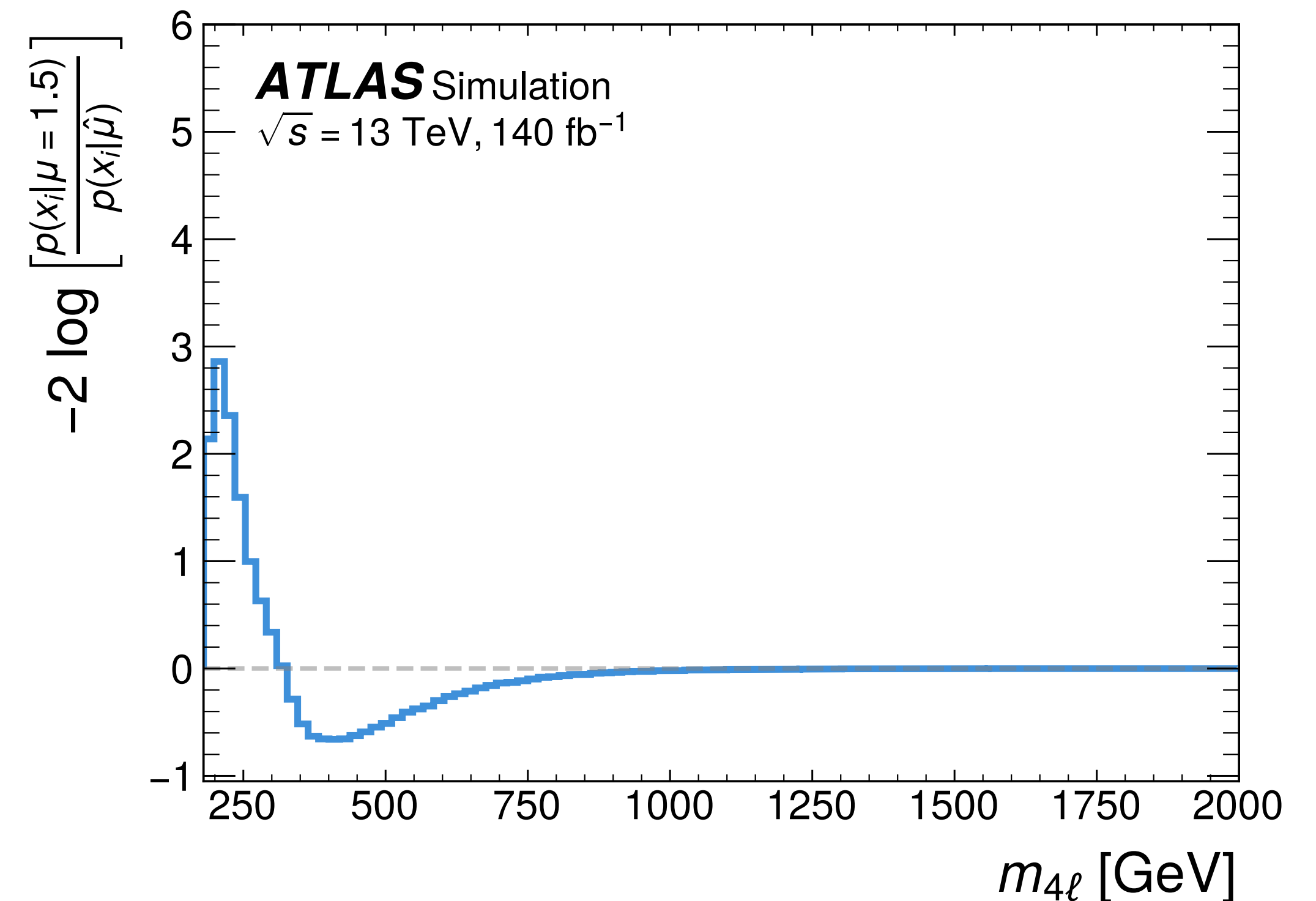
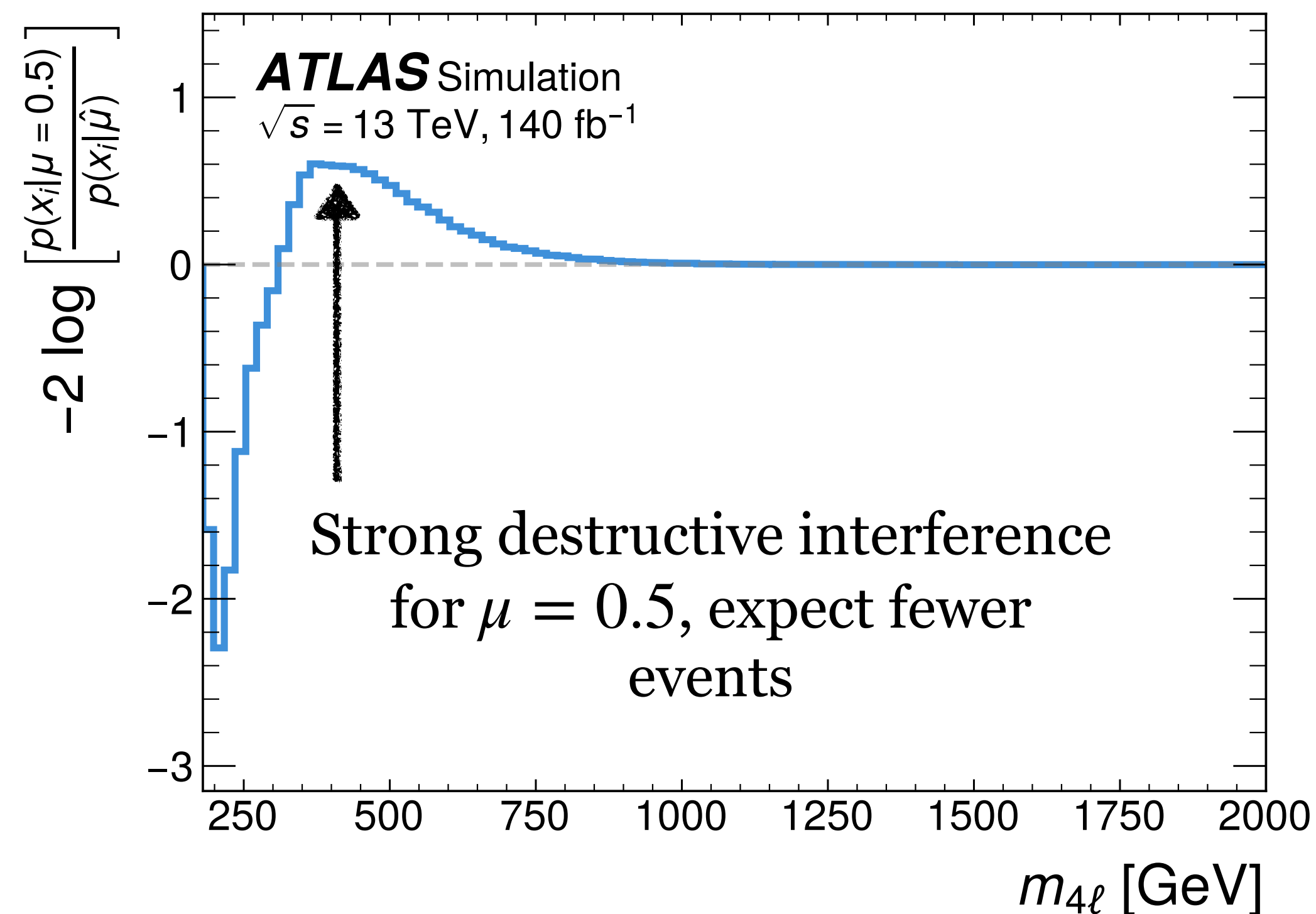
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Interpretability: Which phase space favours one hypothesis over another?

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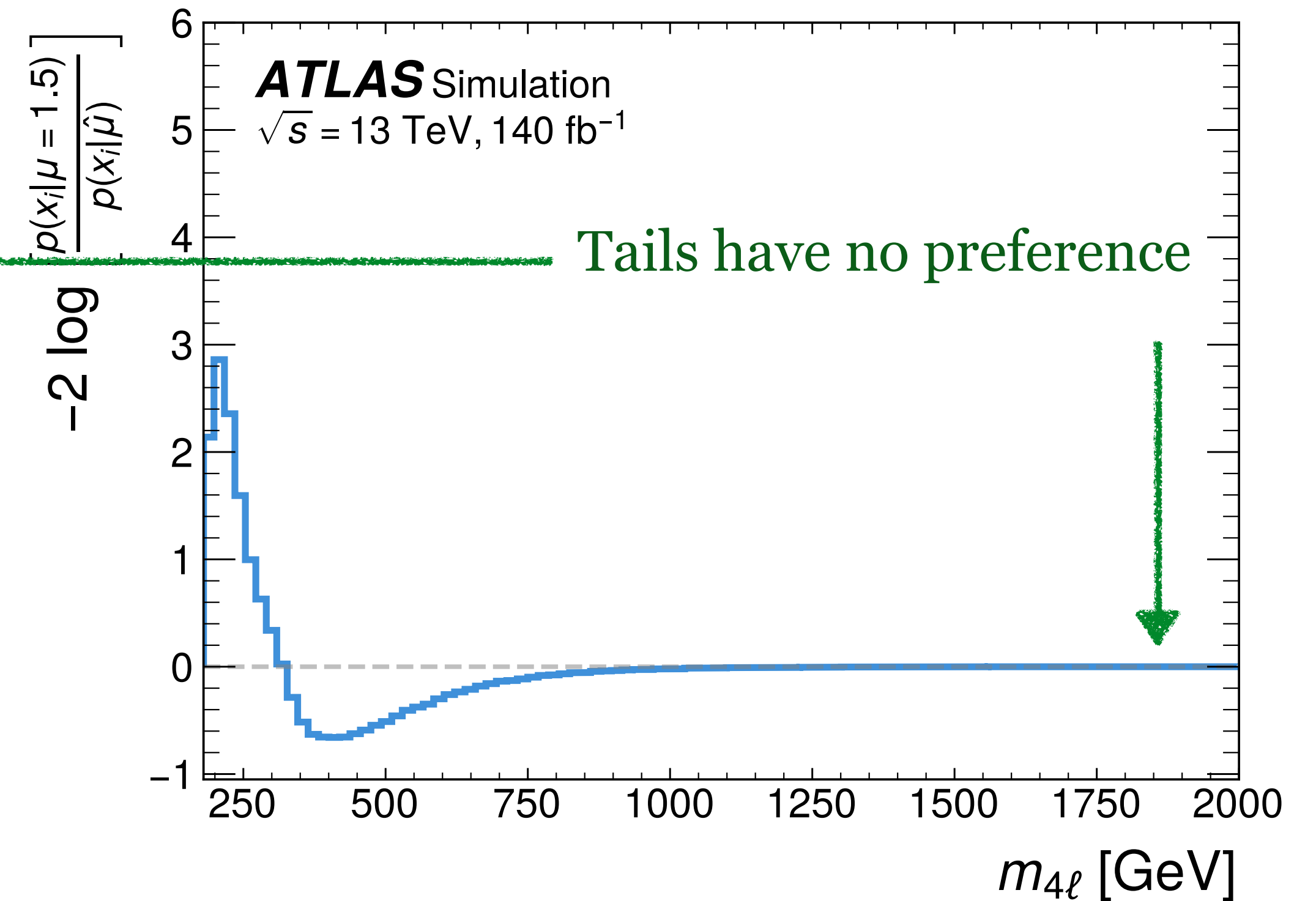
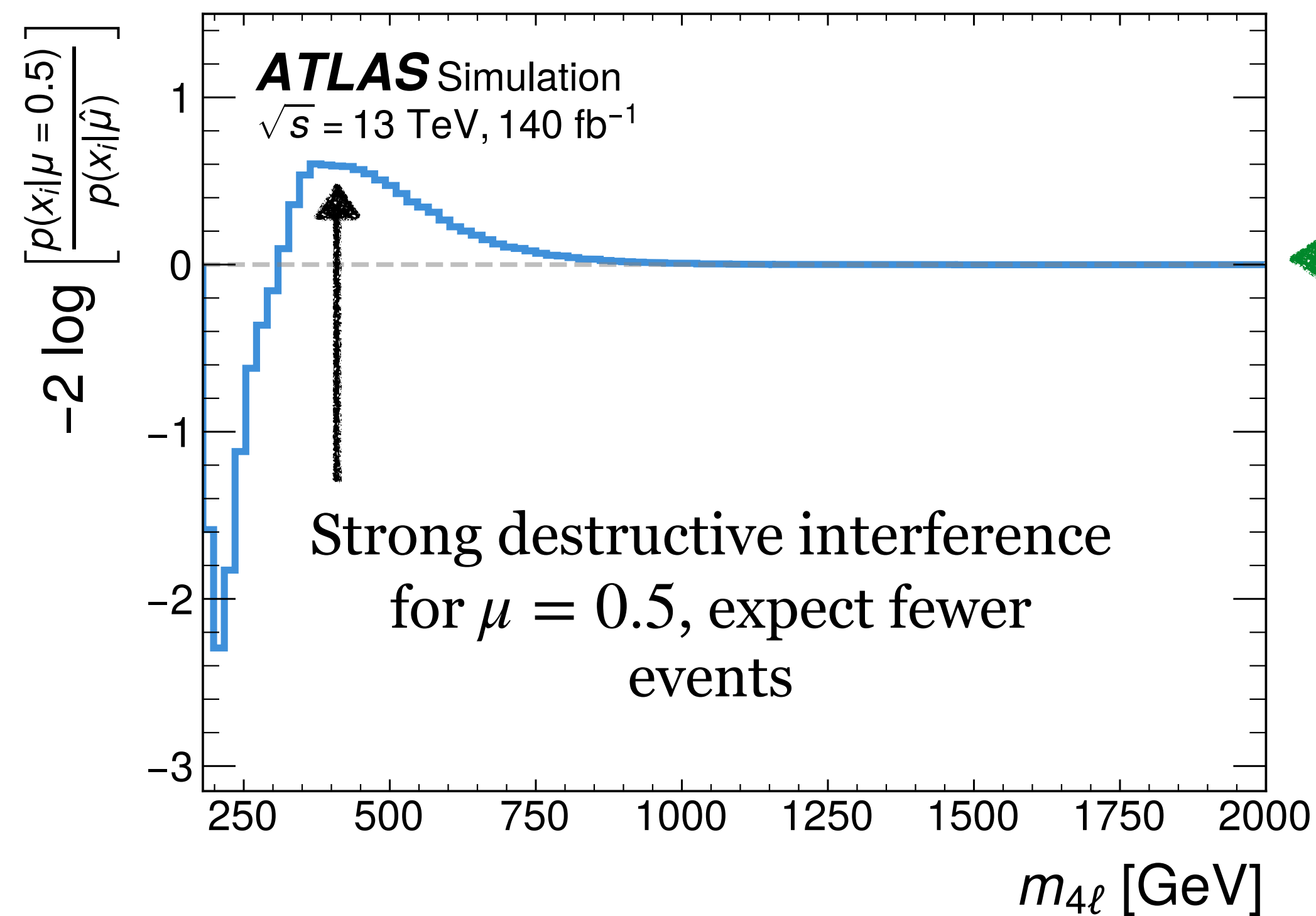
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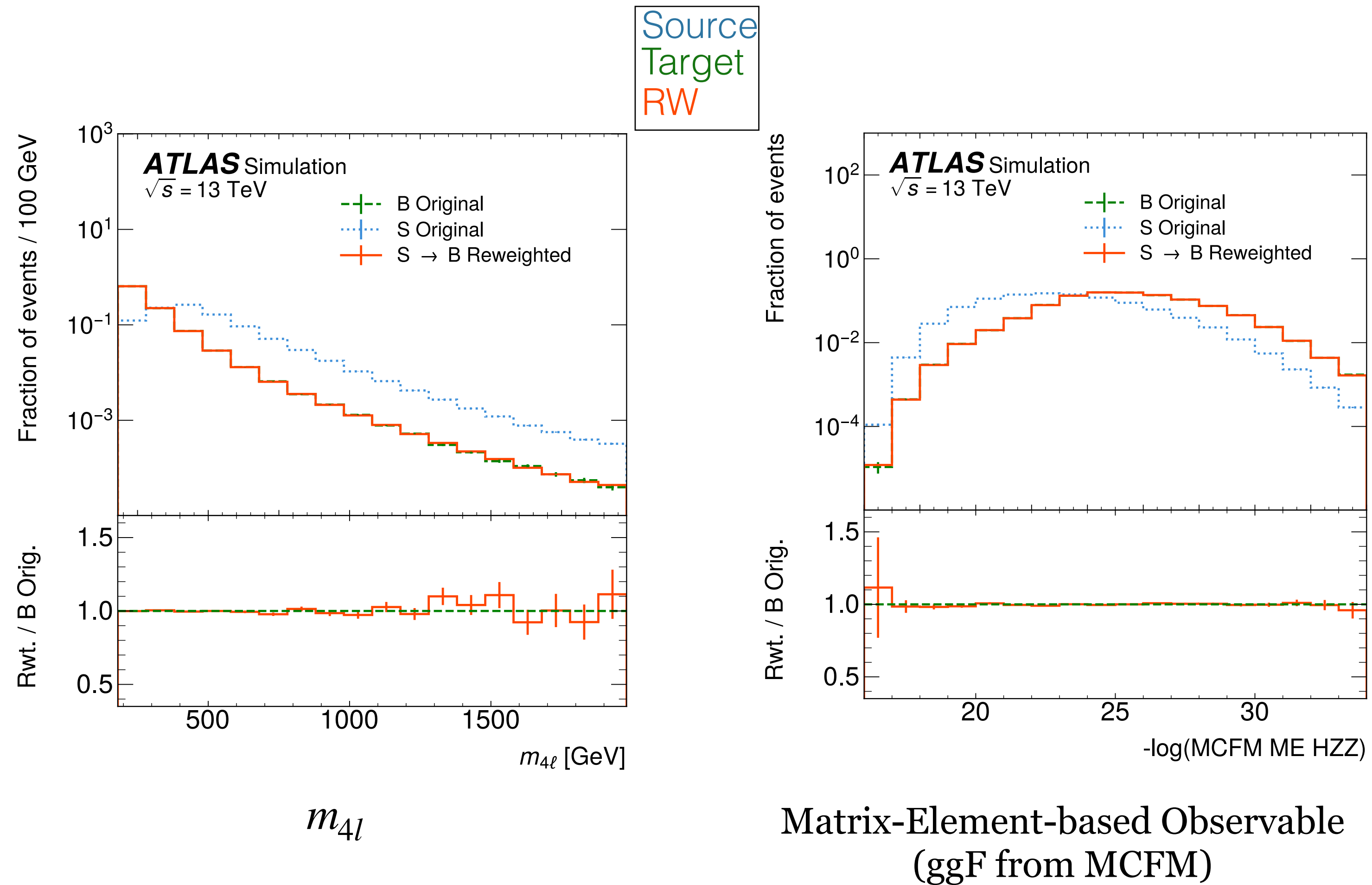
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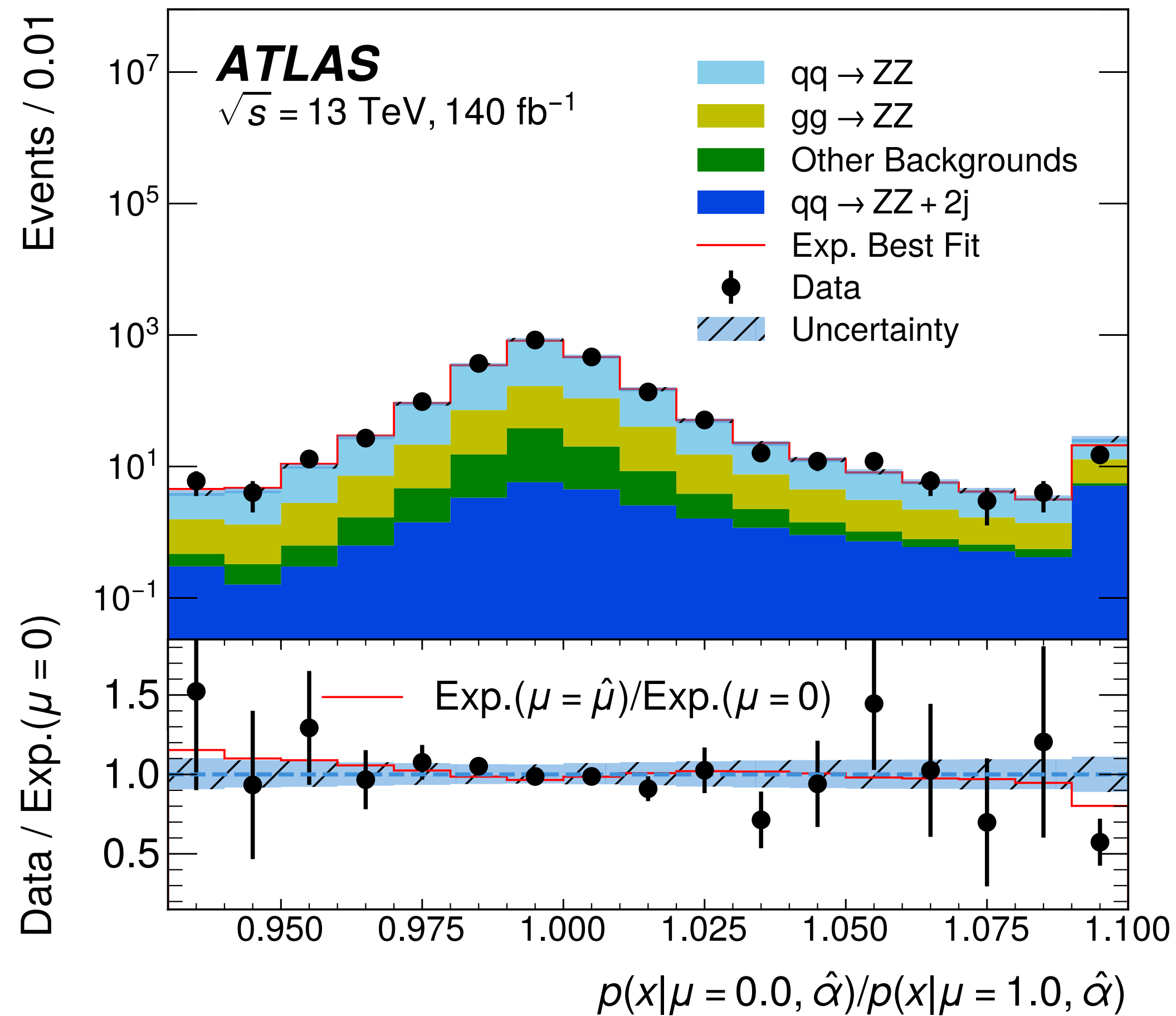
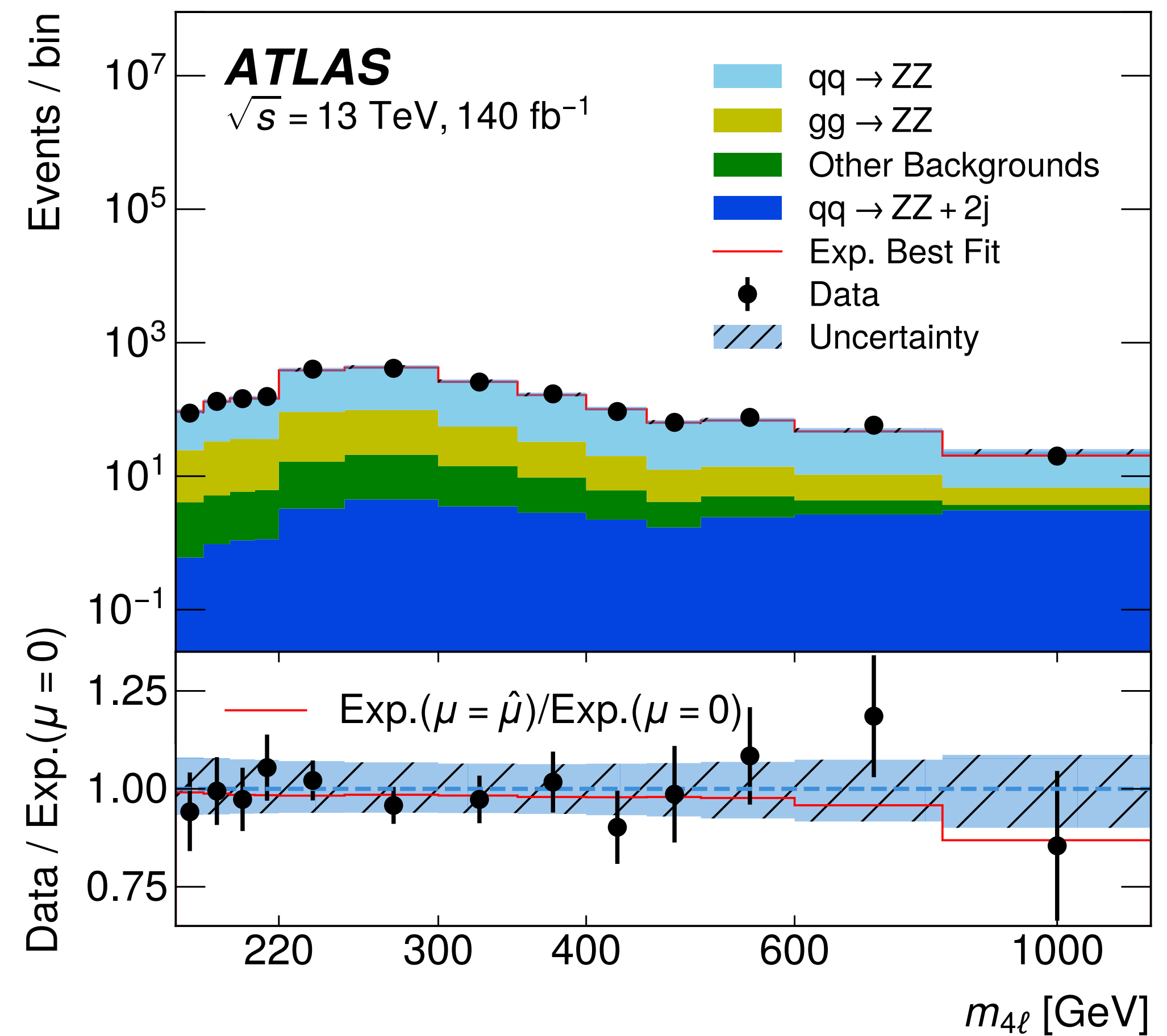


Re-weight closures for B



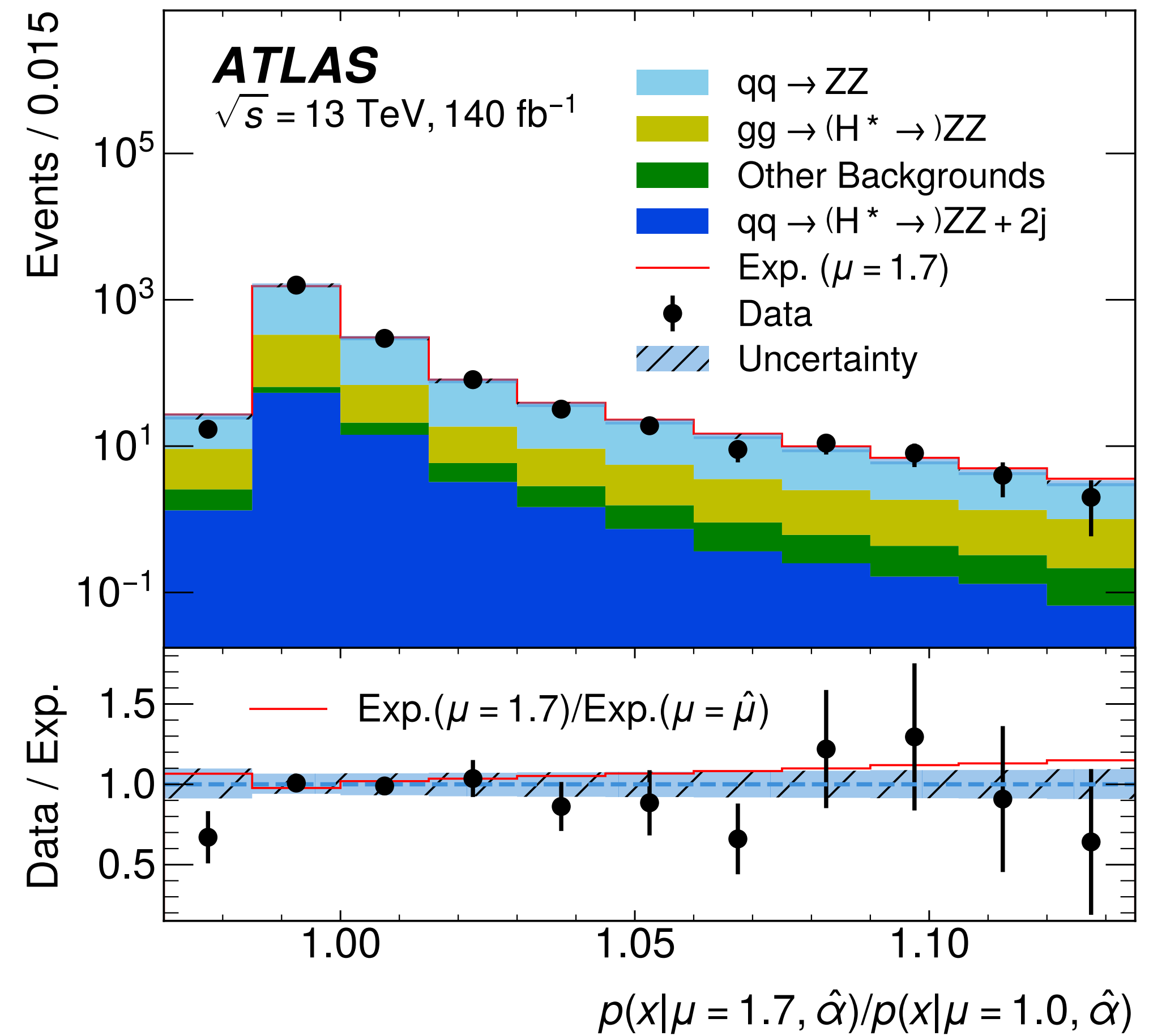
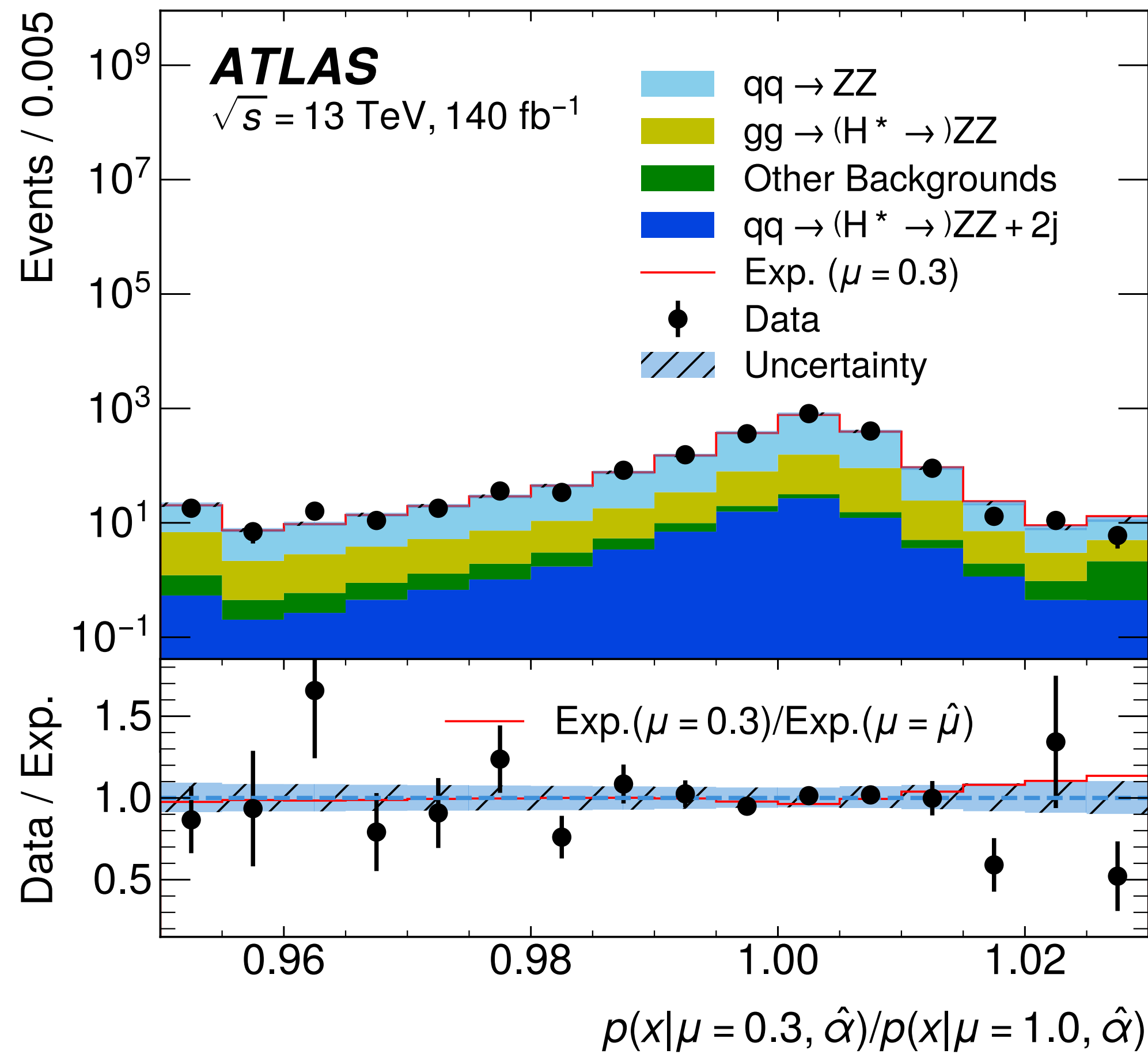
Data-MC validation

NN observable

 m_{4l} 

Data-MC validation

Different NN observables

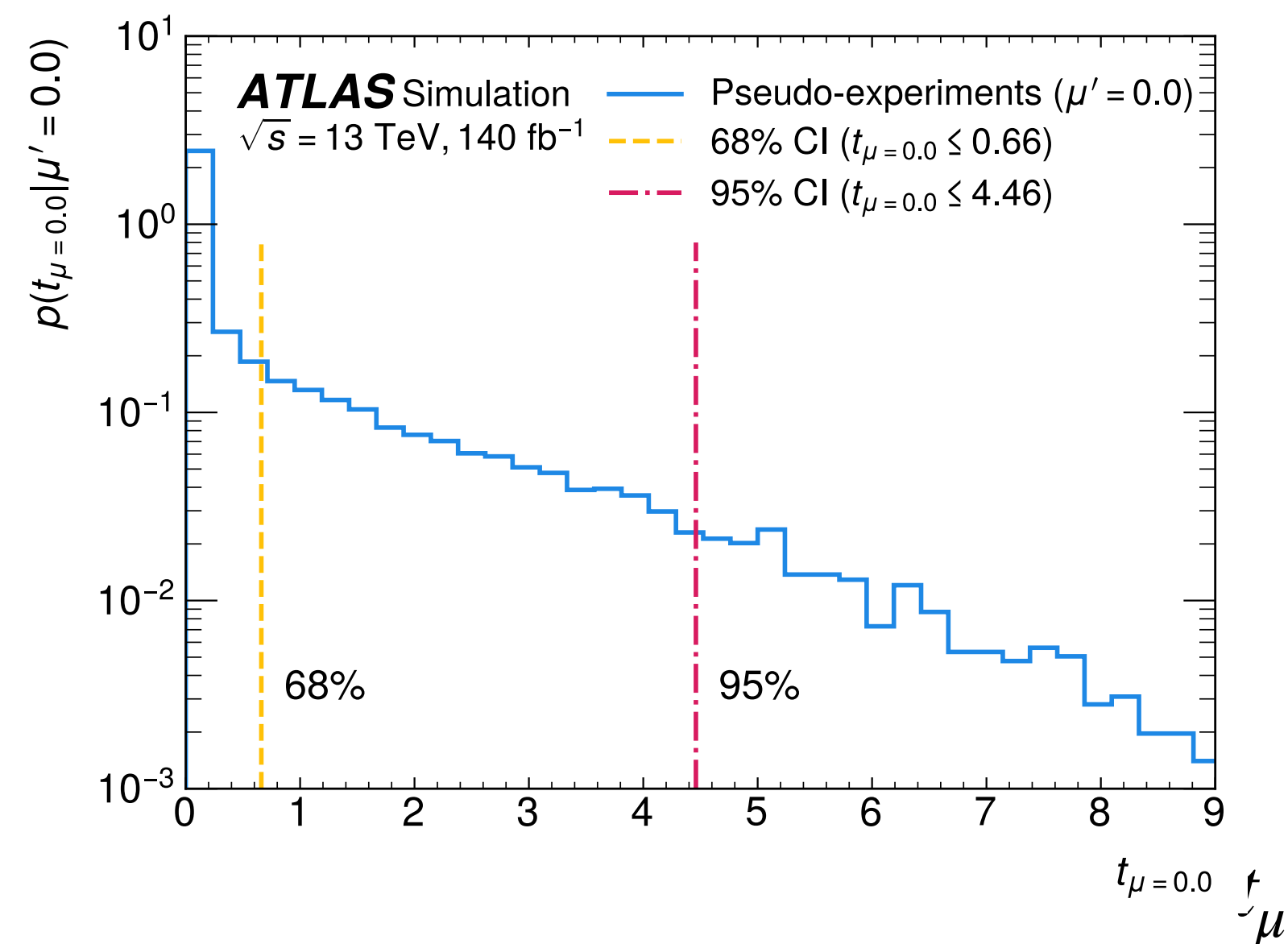


Computational Challenge: Inverting the test

- Determine 68 % & 95 % CI empirically from this distribution
- Do it for each value of μ

Distribution of test statistic t_μ over thousands of simulated pseudo-experiments

True $\mu = 0$



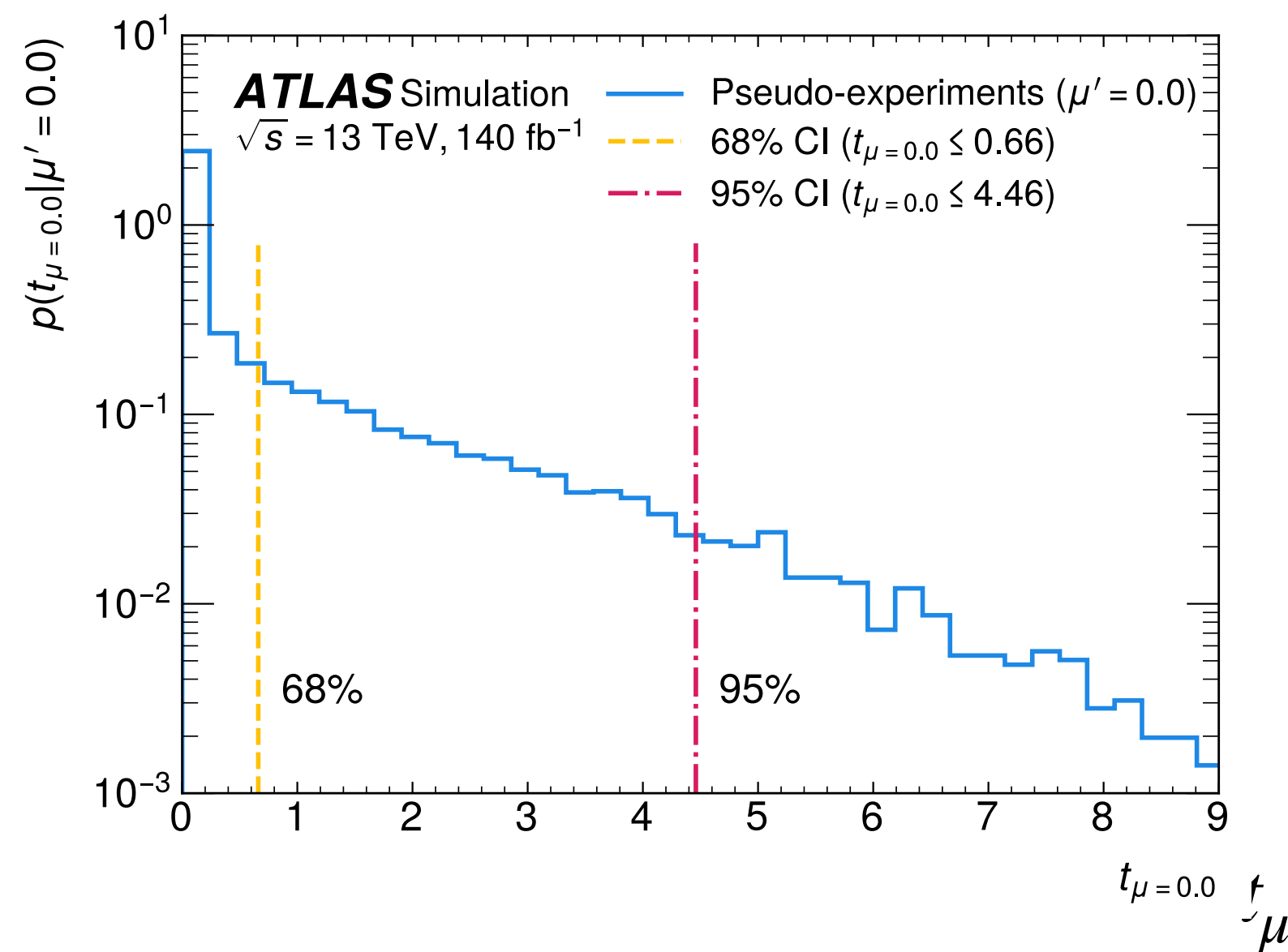
t_μ

Computational Challenge: Inverting the test

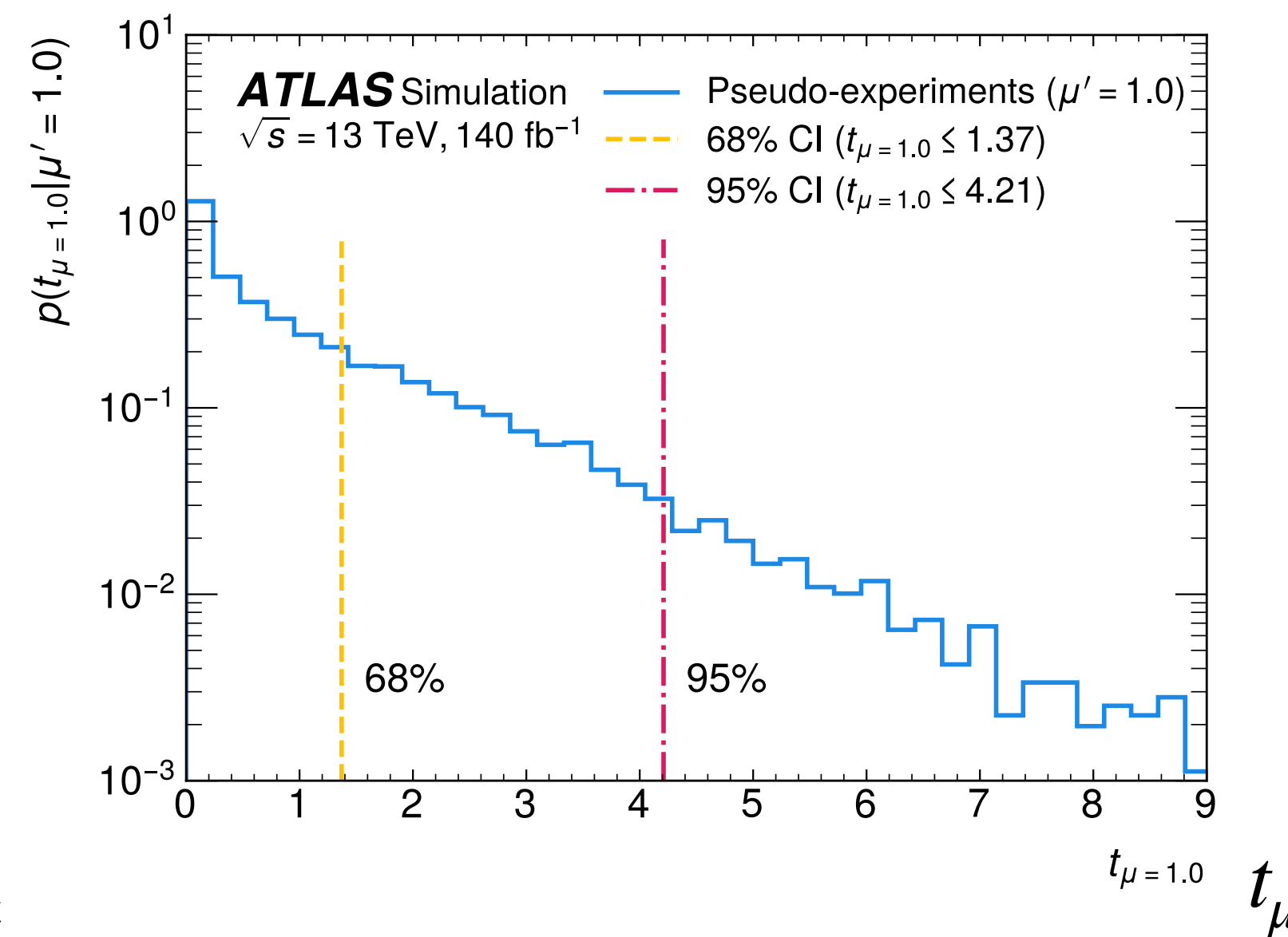
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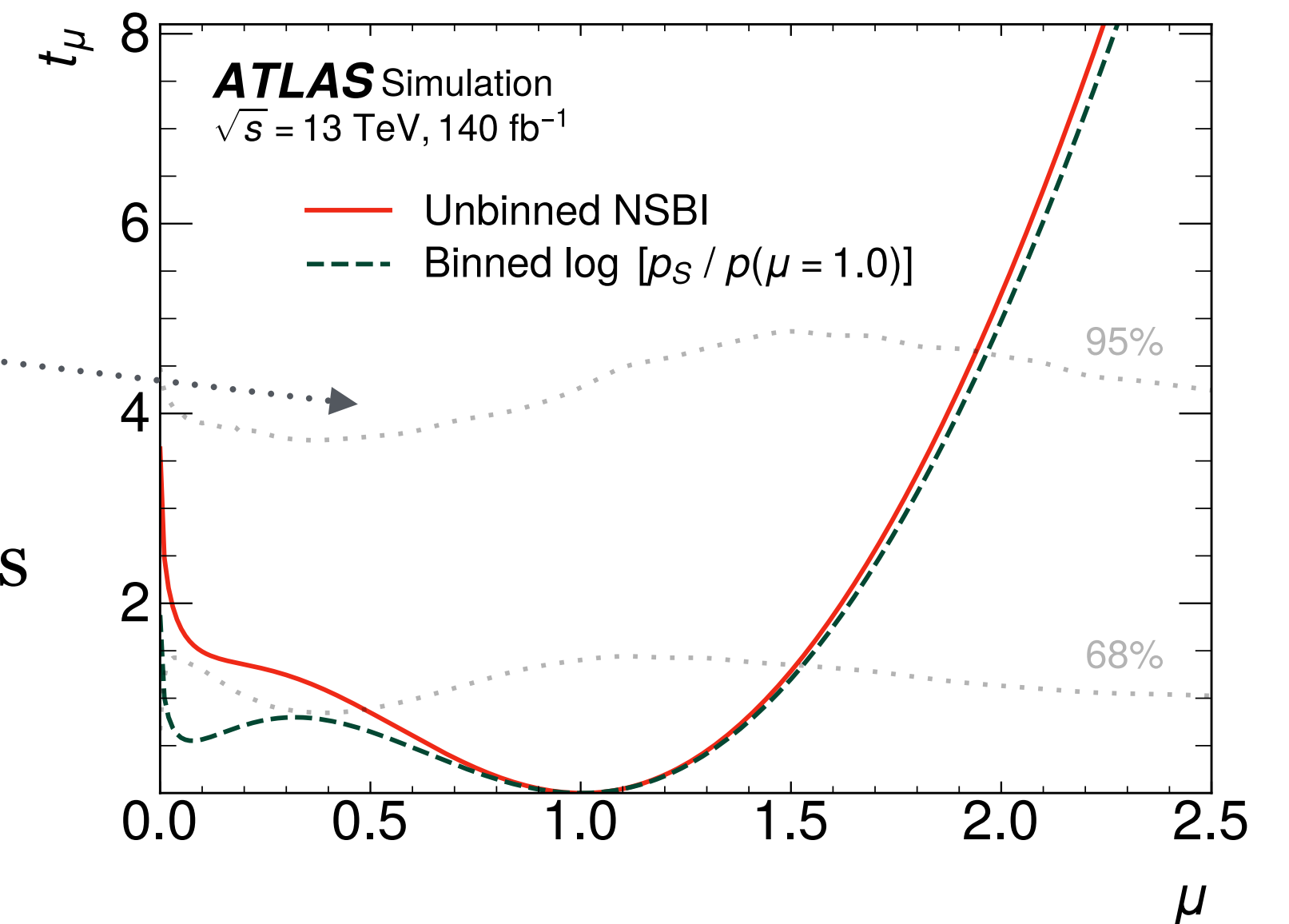


True $\mu = 1$



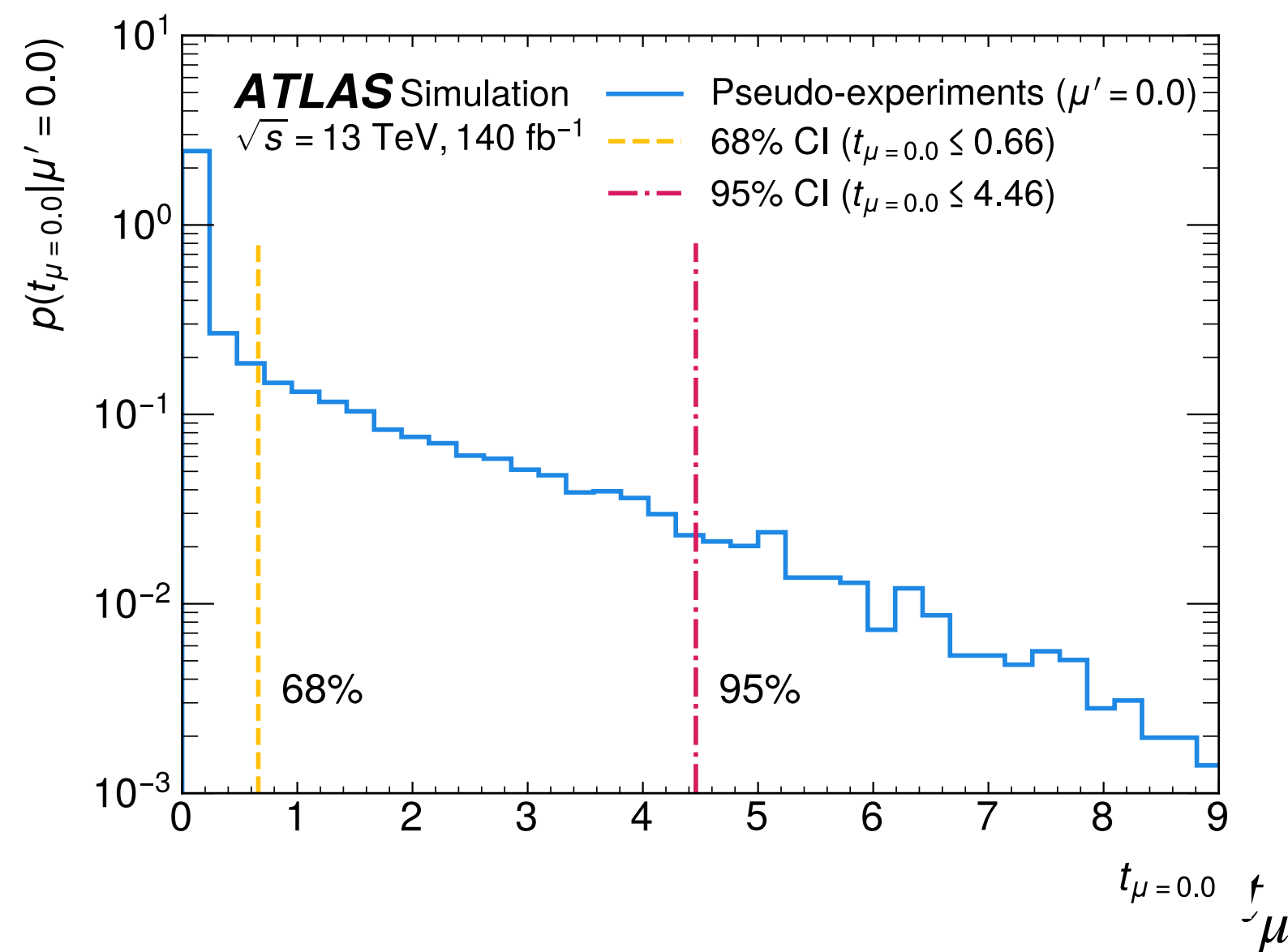
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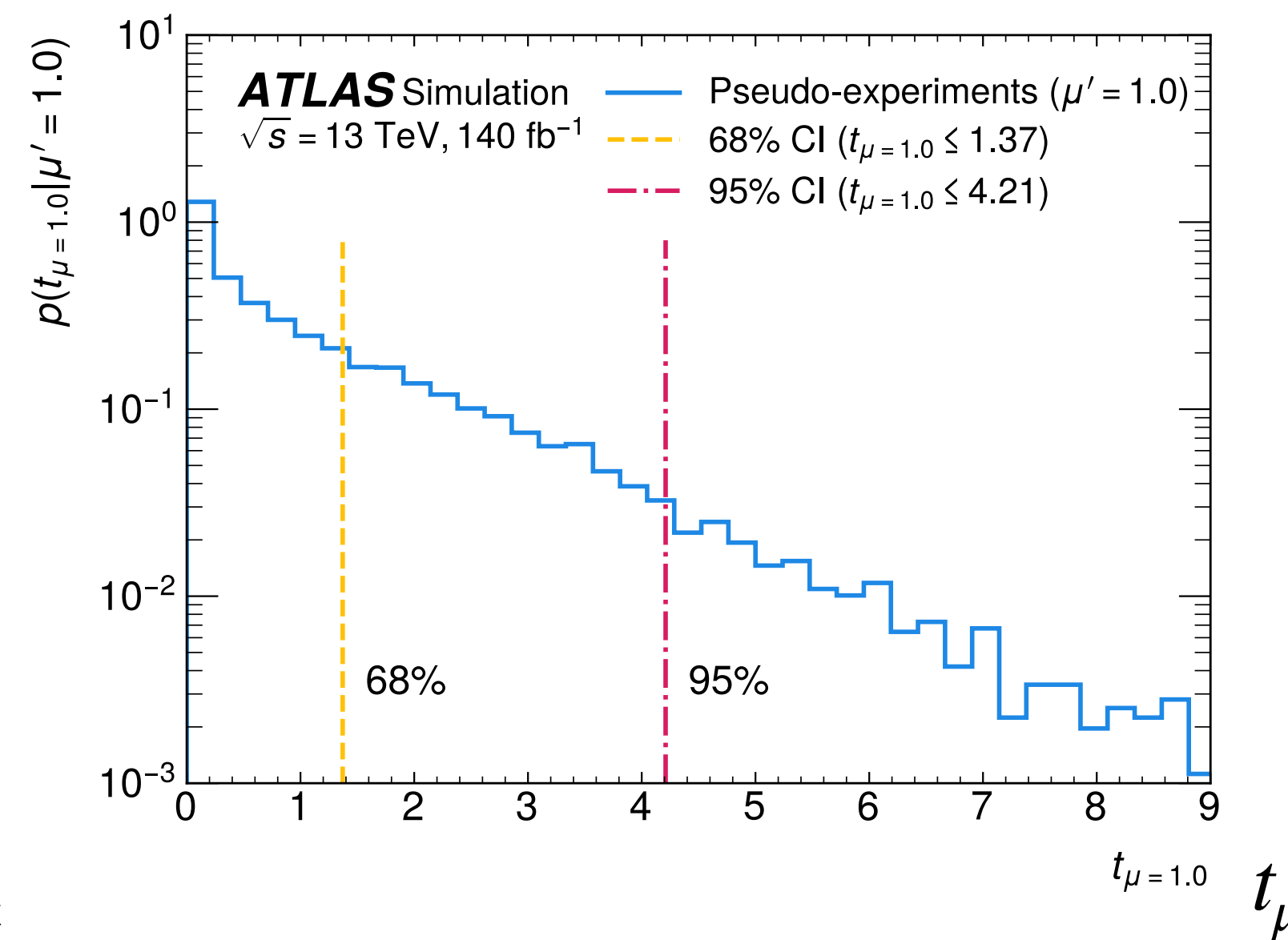


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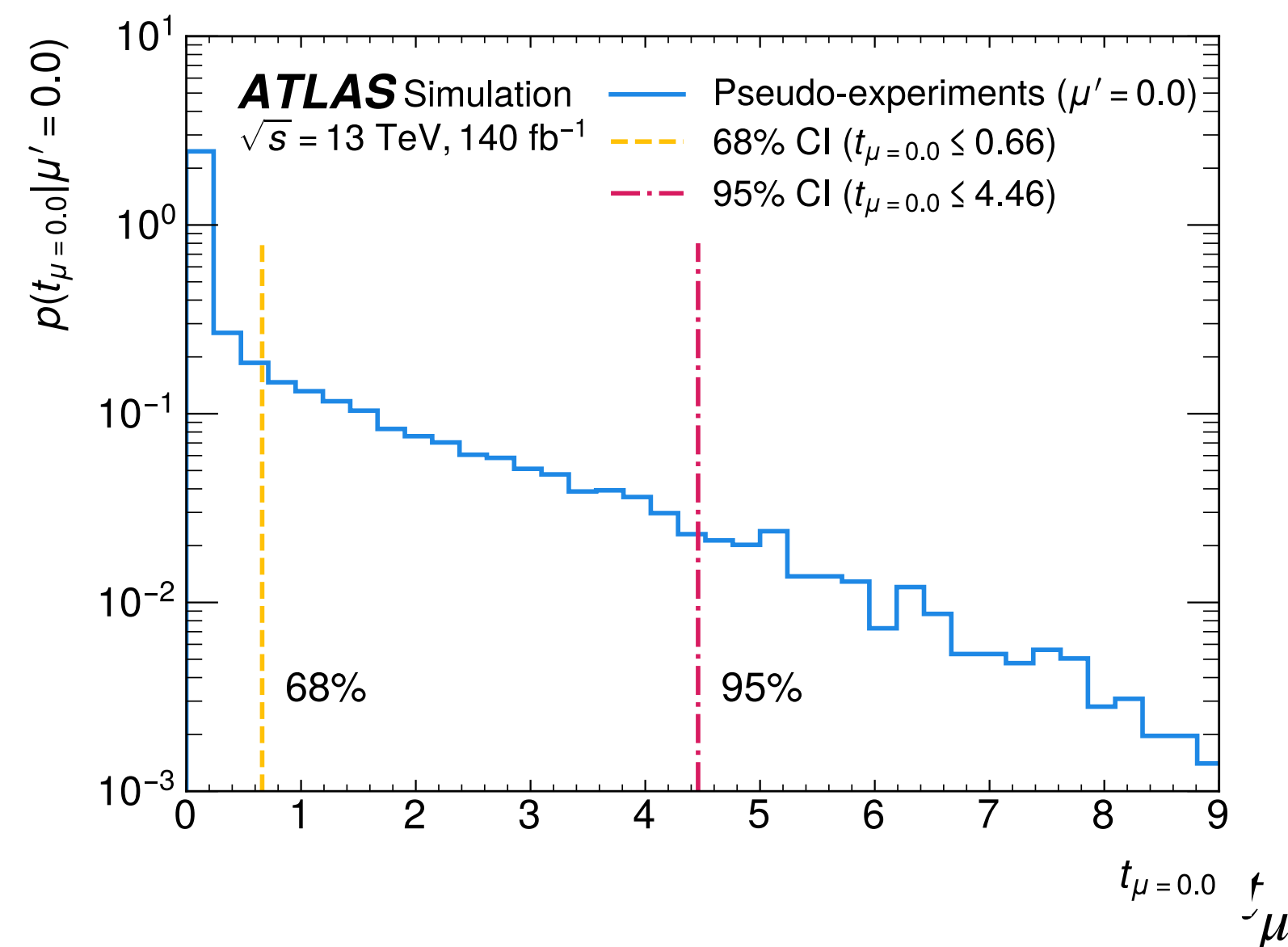


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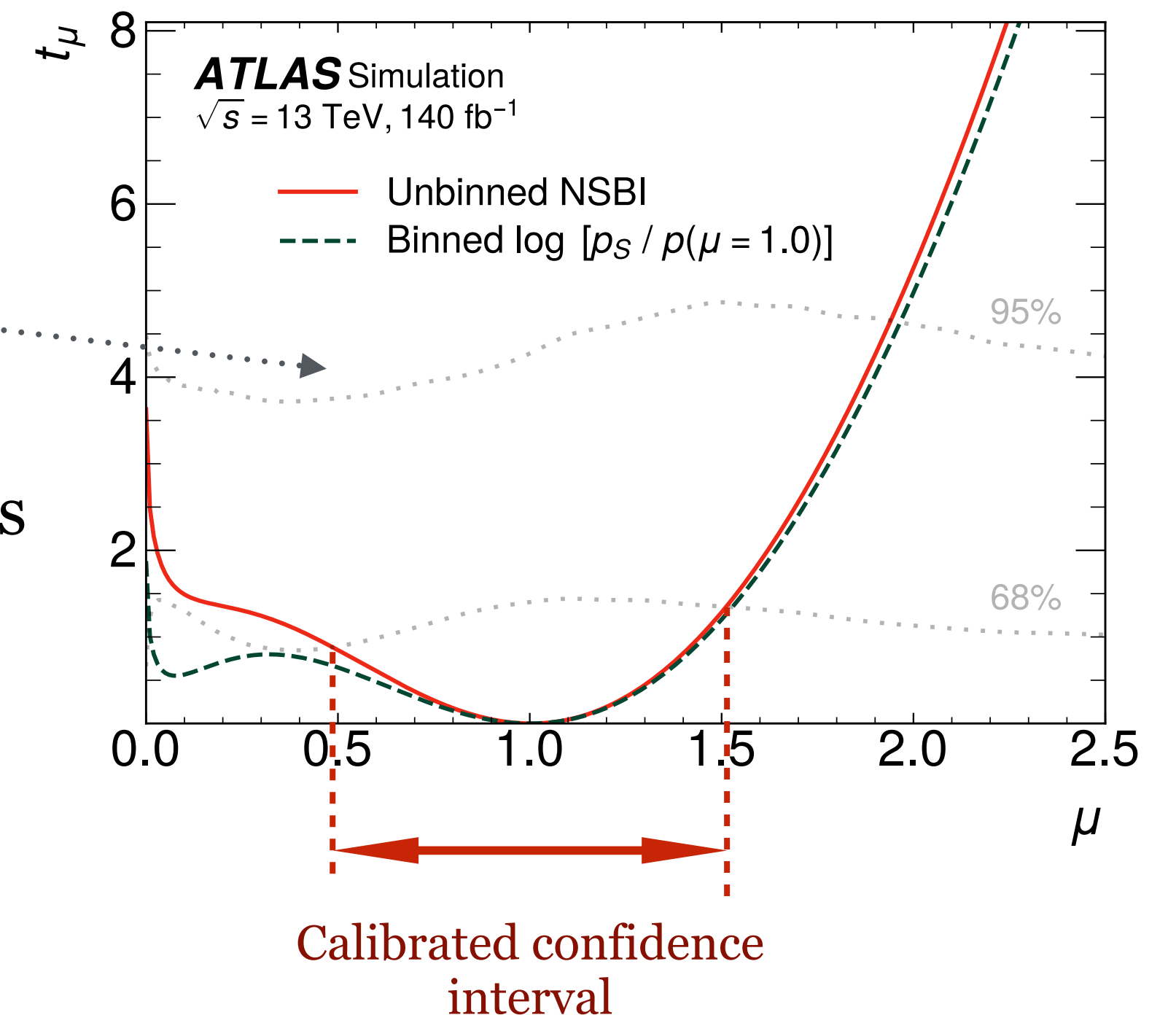
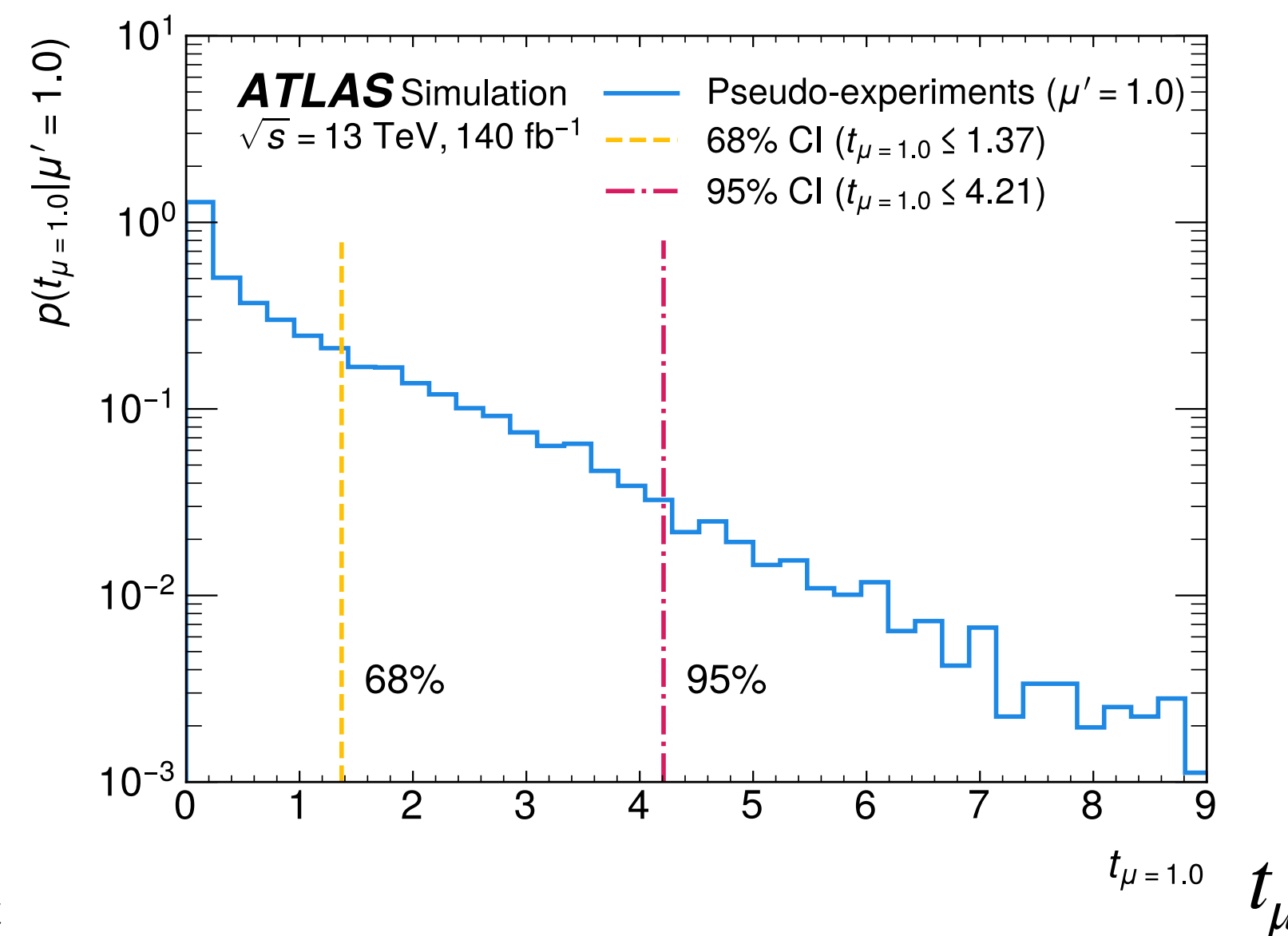
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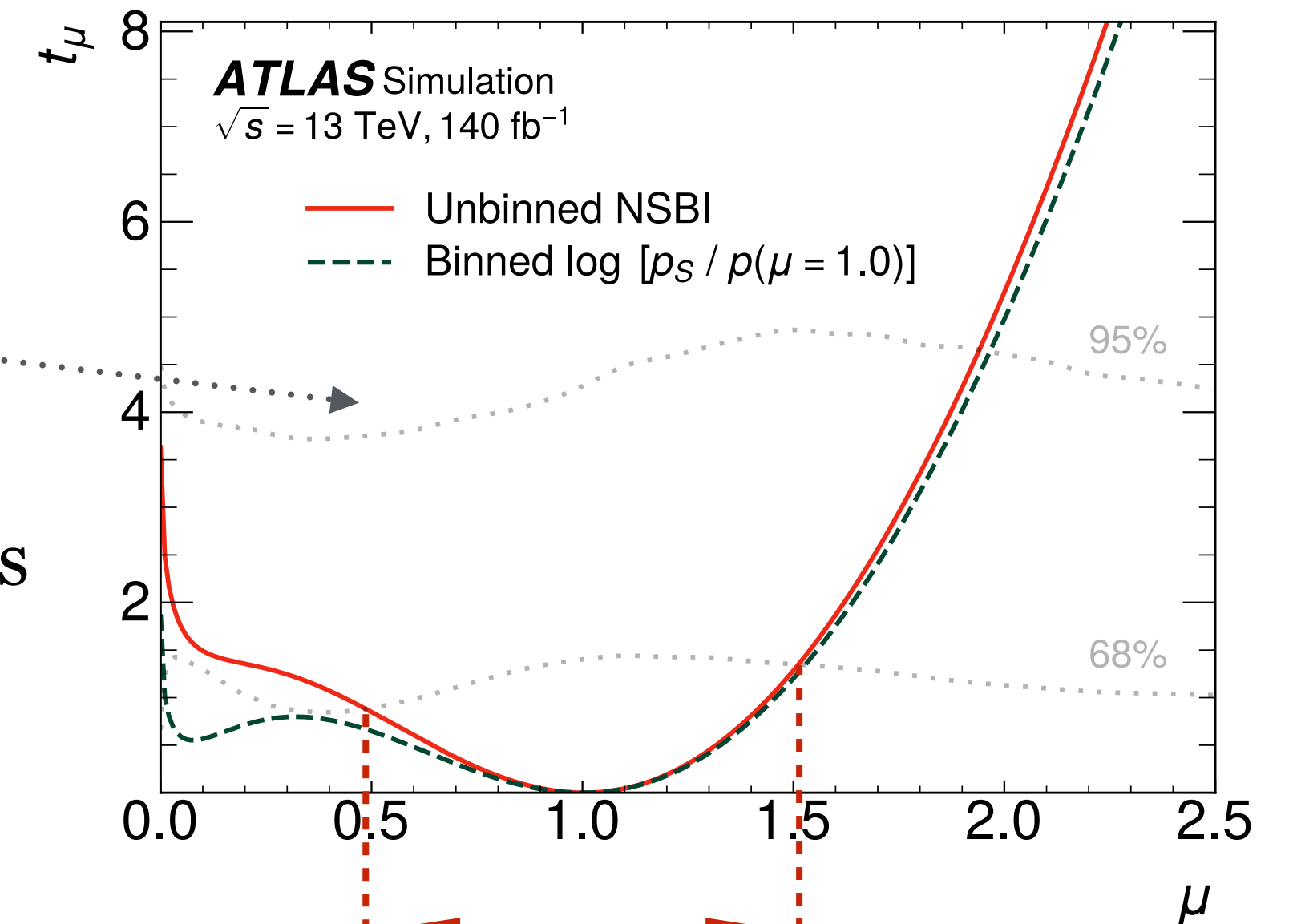
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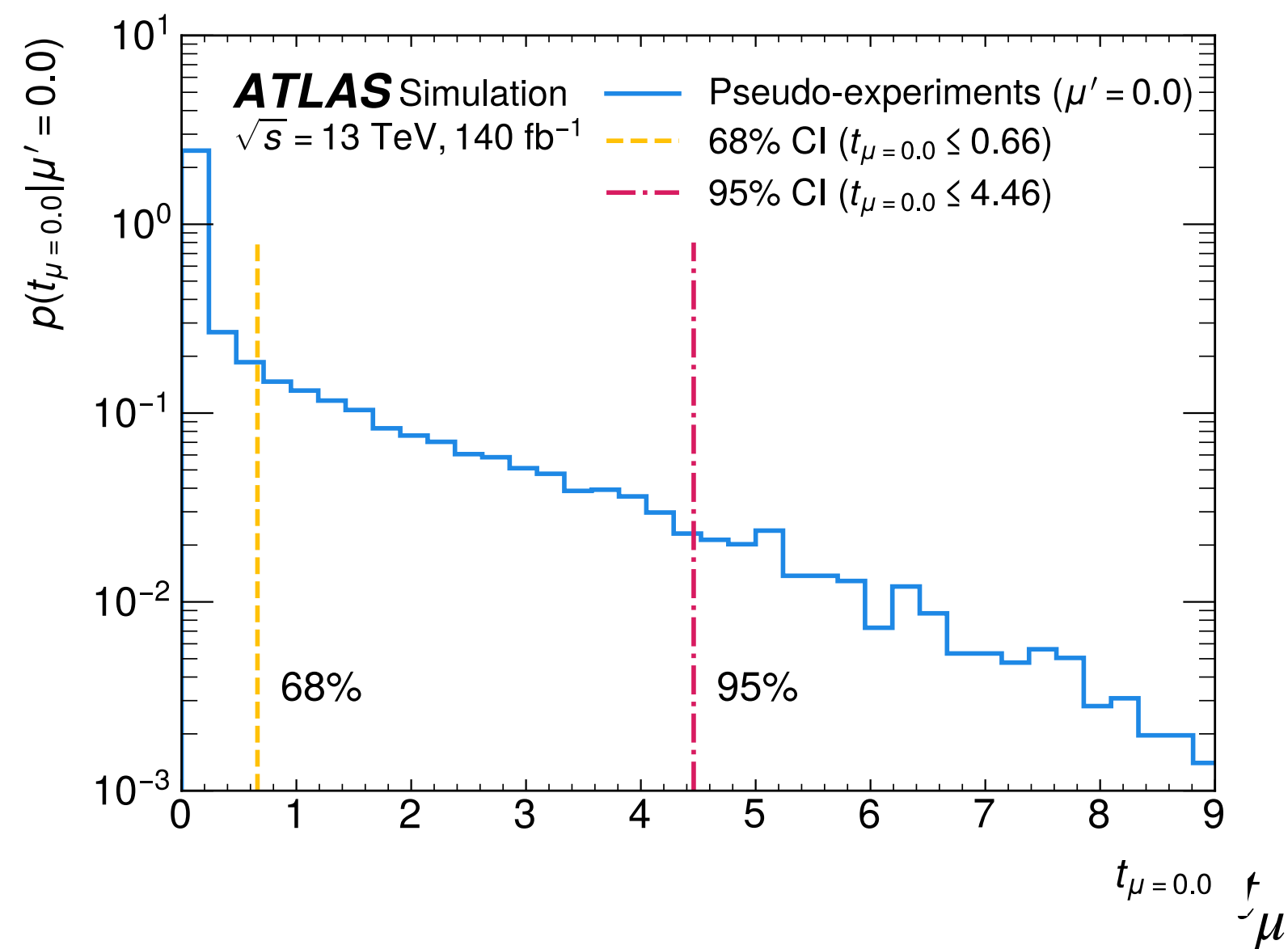
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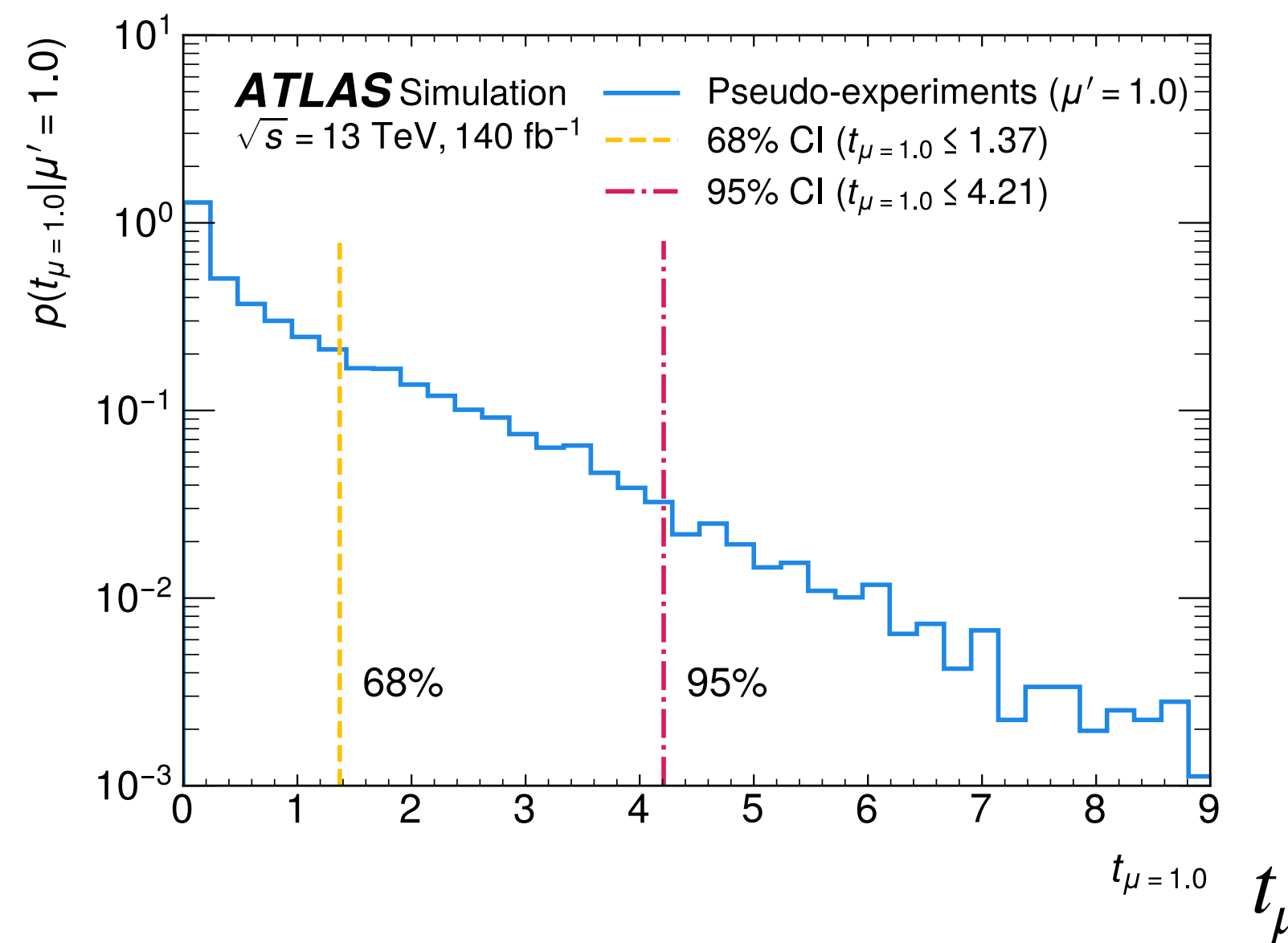
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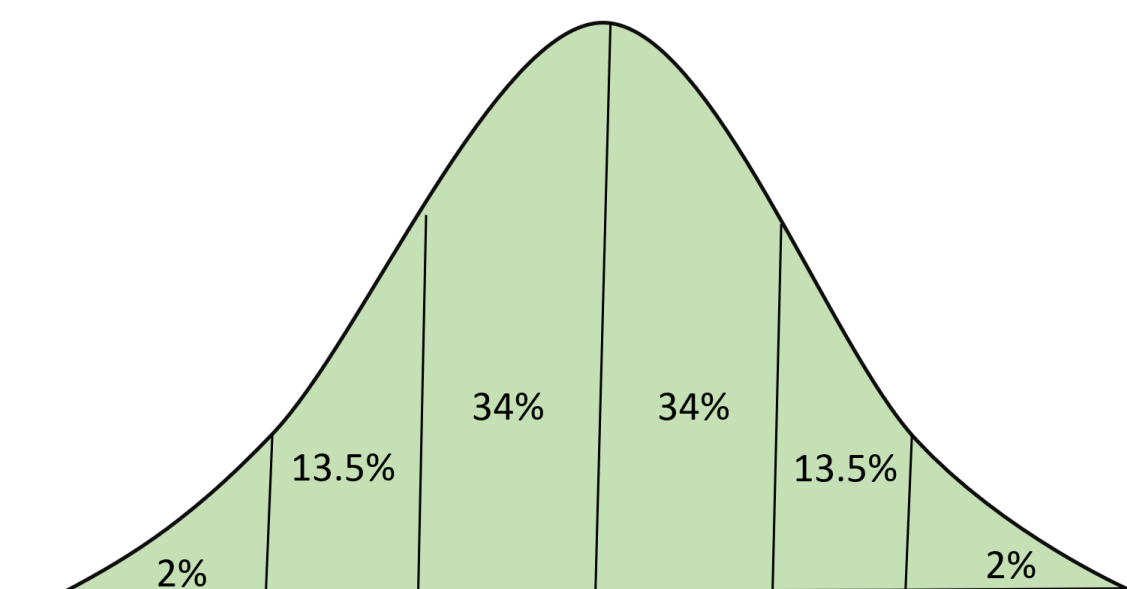
True $\mu = 0$



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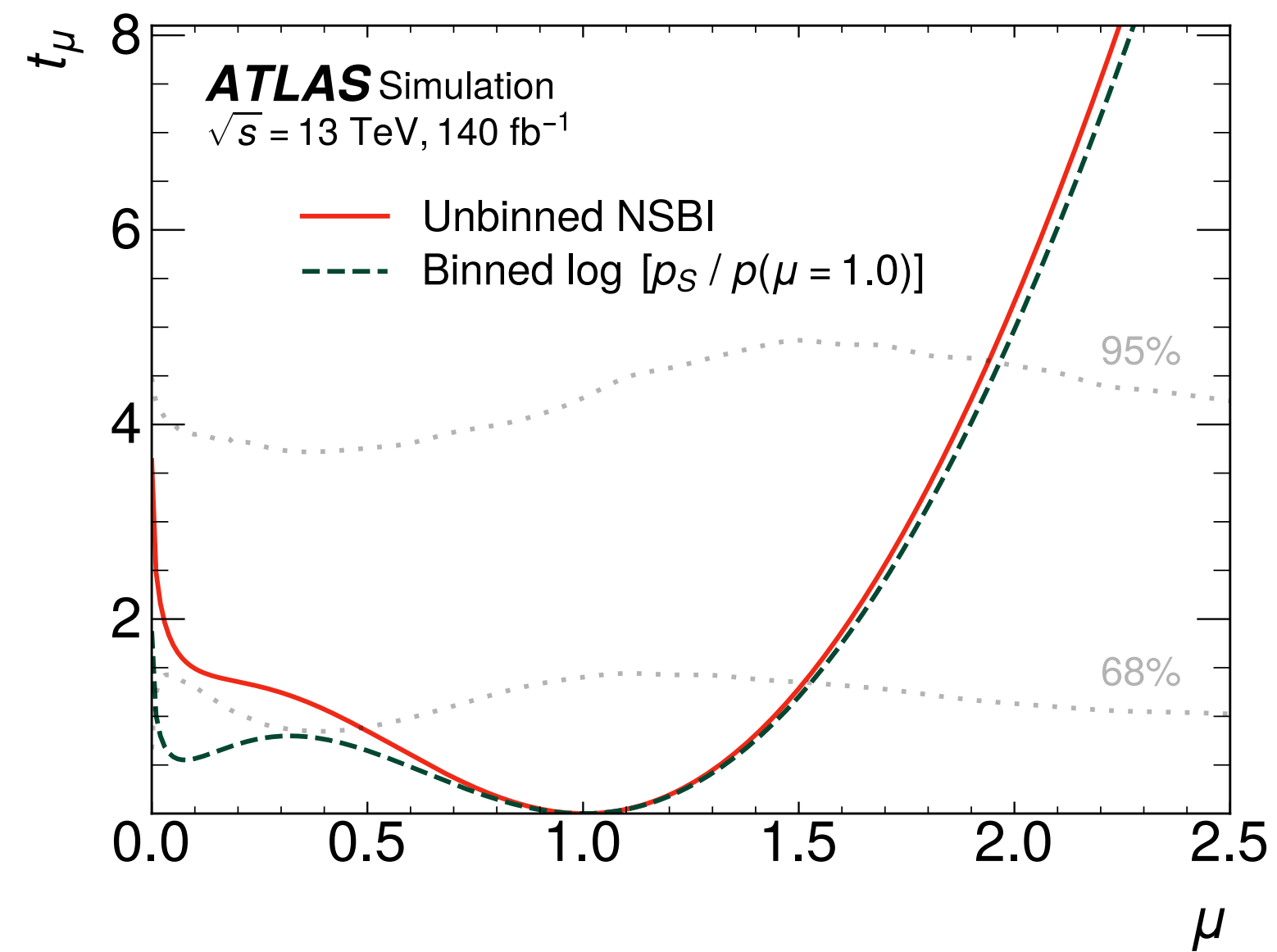
Task is to determine quantiles



Fast confidence belts with QR

Neyman construction guarantees accuracy of confidence intervals, even if you don't trust the NN estimates

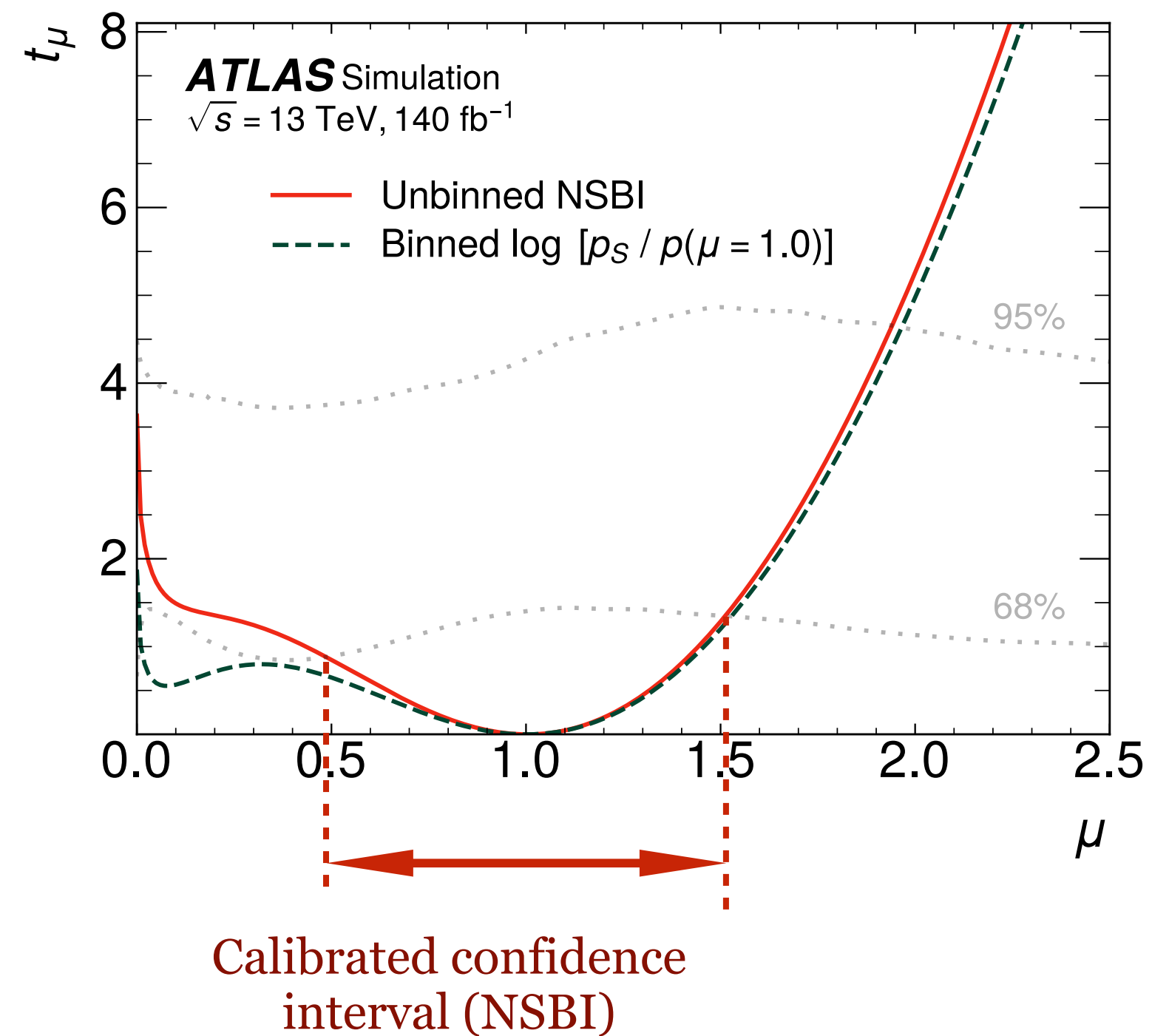
But very slow and expensive



Fast confidence belts with QR

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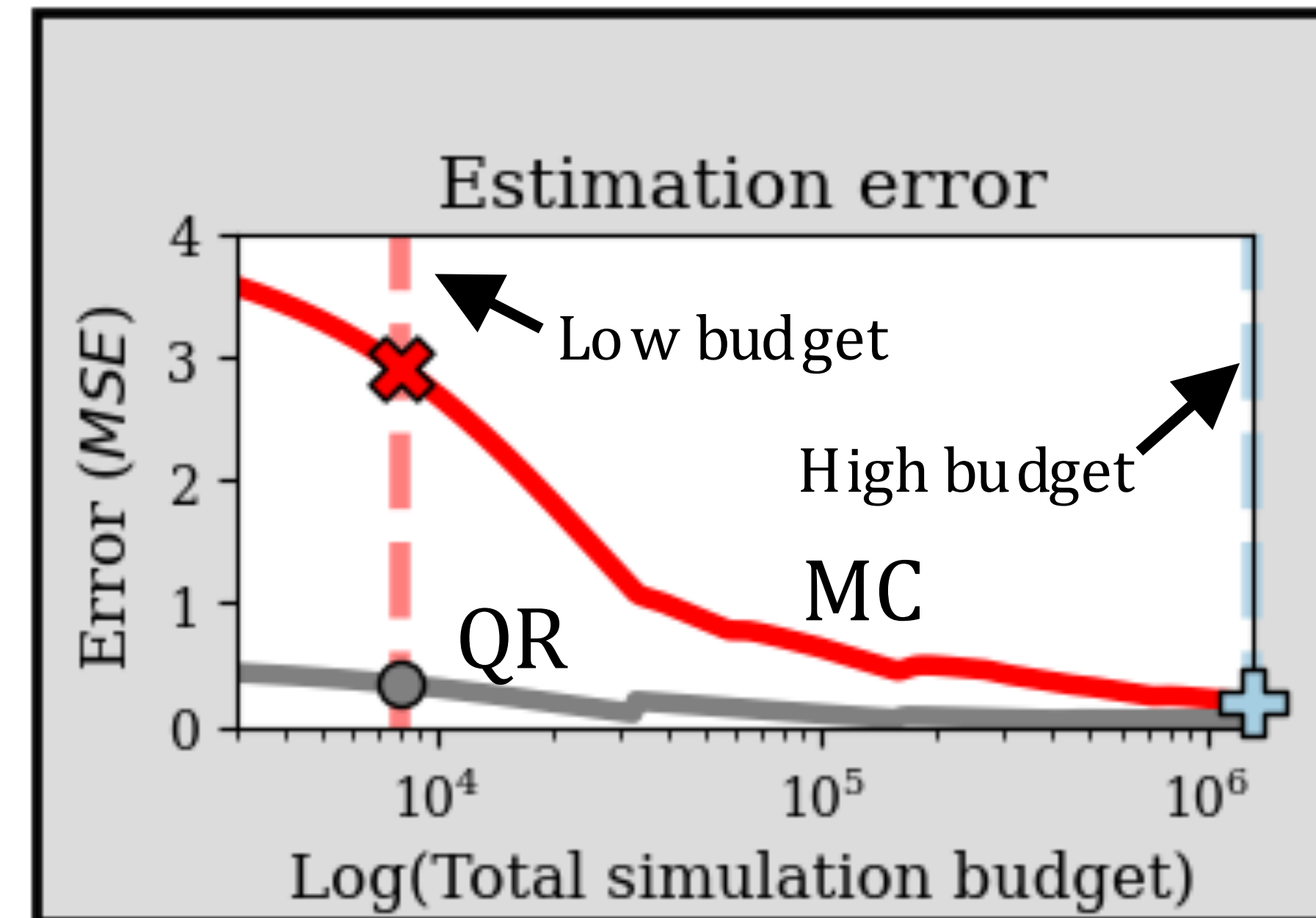
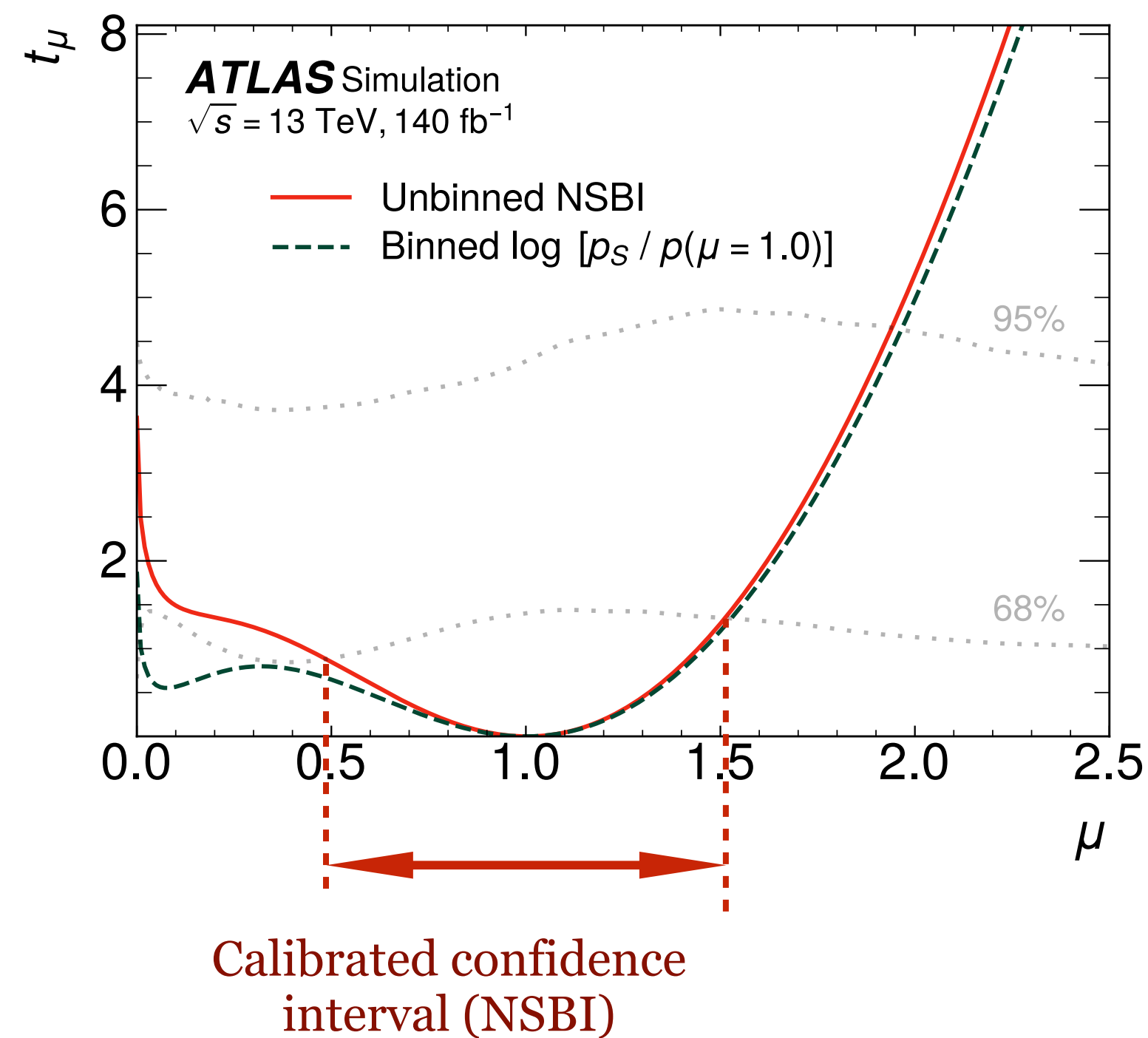


Fast confidence belts with QR

Neyman construction guarantees accuracy of confidence intervals, even if you don't trust the NN estimates

But very slow and expensive

Can do it faster and more accurately with quantile regression (QR)

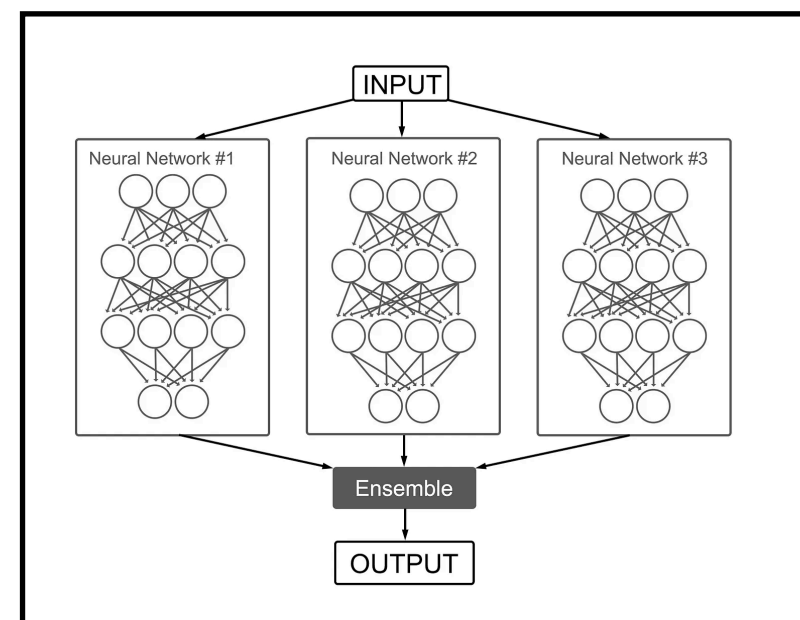


Workflows for the computational challenge

- HEP performs ~100 measurements per year that could benefit
- Careful design and validation of 10,000 networks per measurement is infeasible

Need fully automated workflows with Ray:

Train ensembles on
HPC

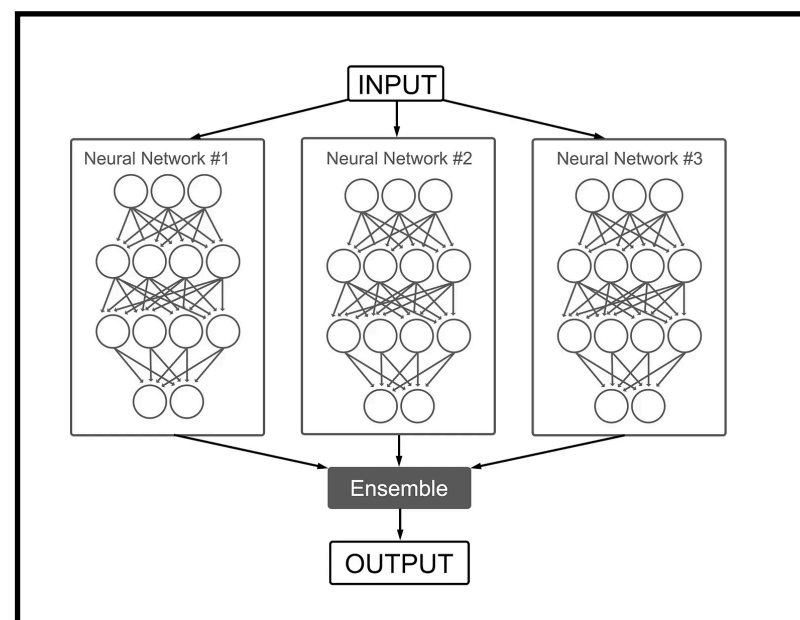


Workflows for the computational challenge

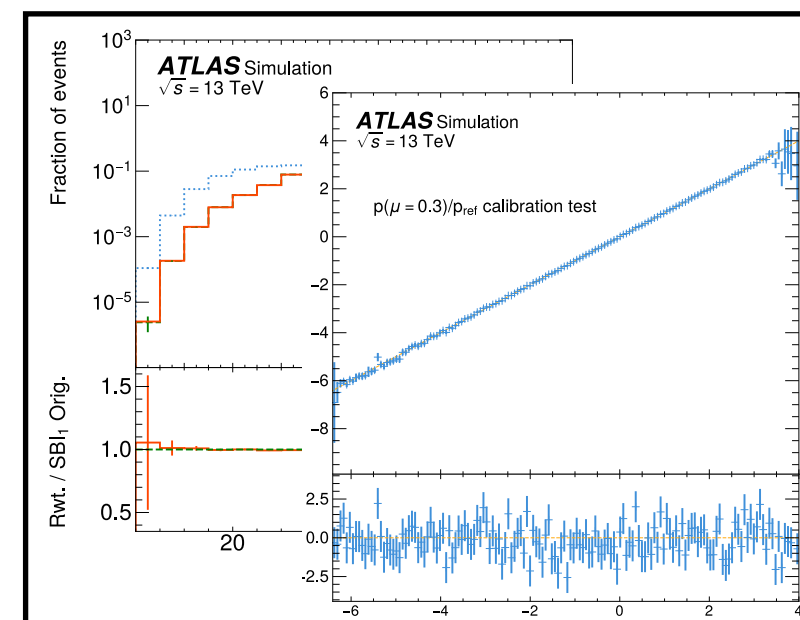
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Validation of density
estimates

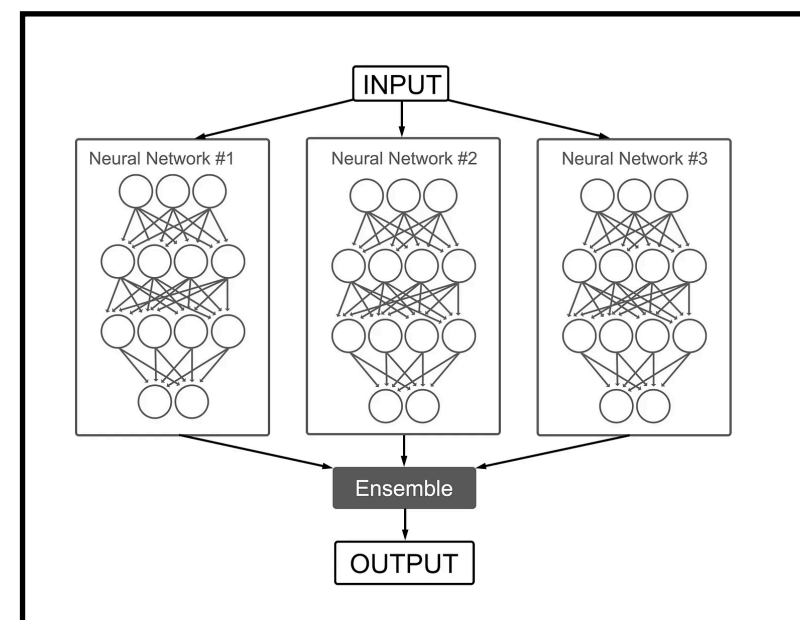


Workflows for the computational challenge

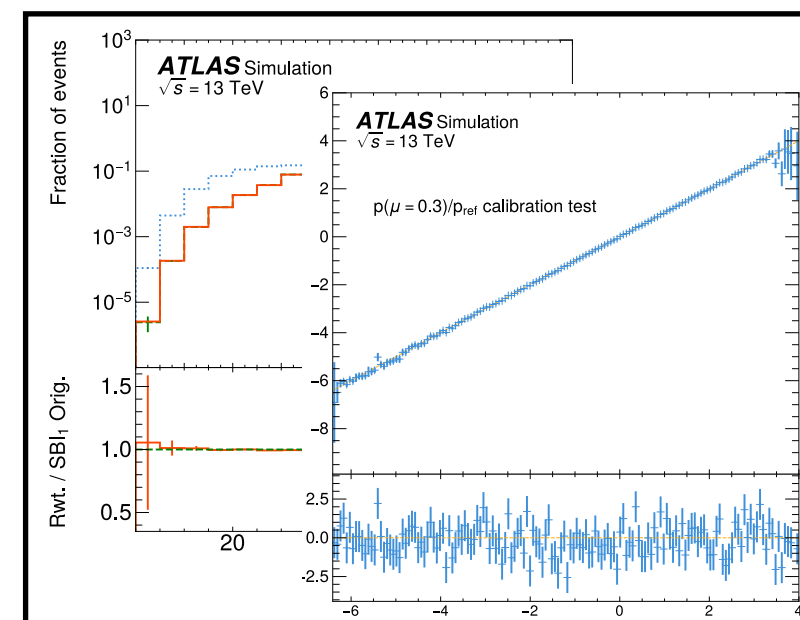
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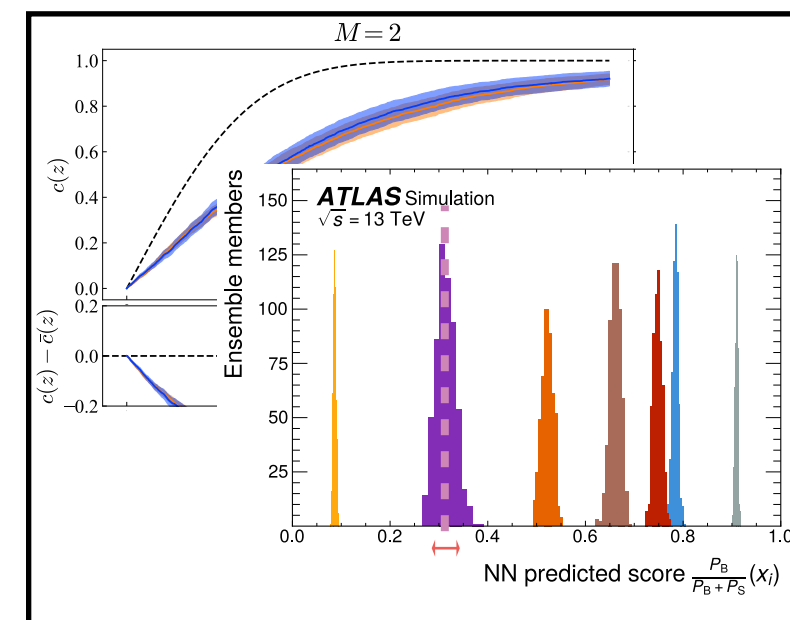
Train ensembles on HPC



Validation of density estimates



Uncertainty quantification

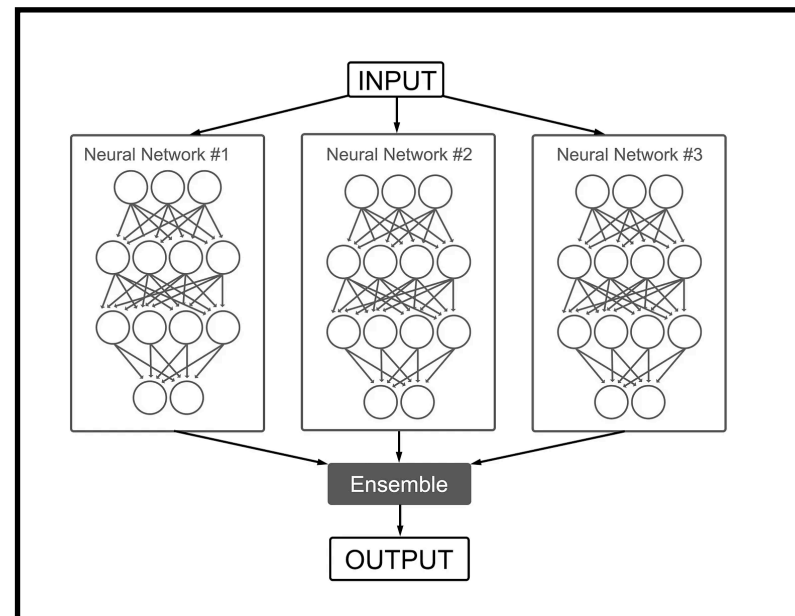


Workflows for the computational challenge

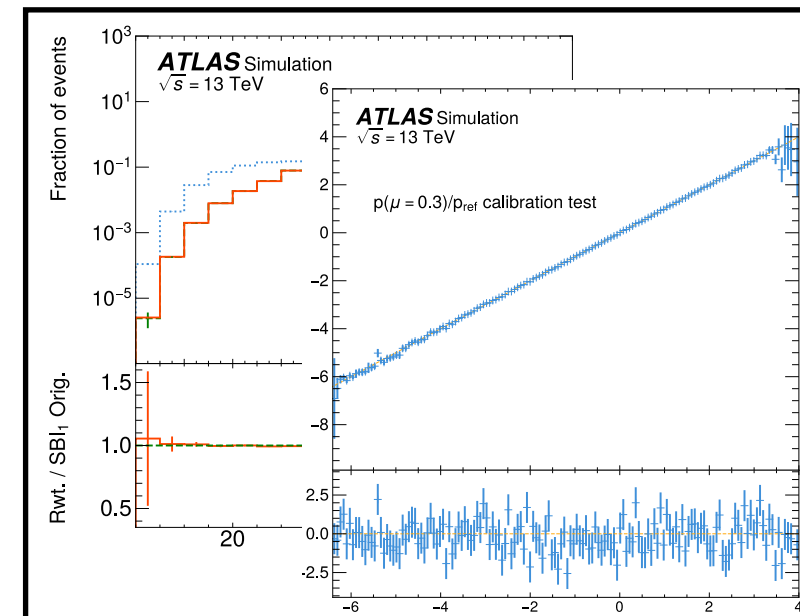
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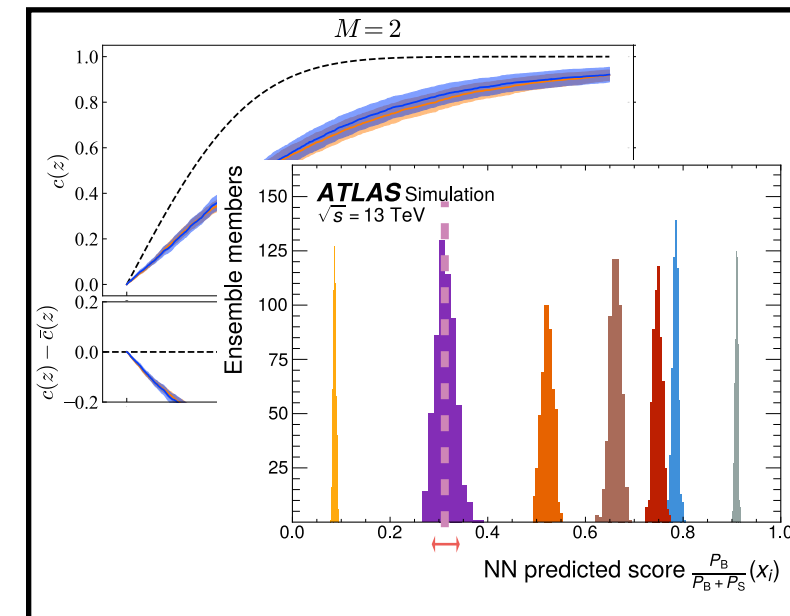
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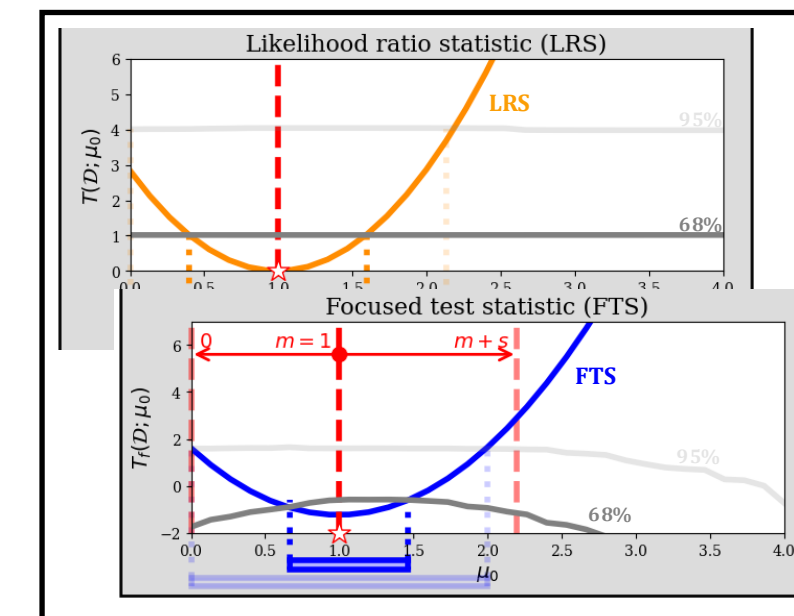
Validation of density estimates



Uncertainty quantification



Neyman inversion with pseudo-experiments

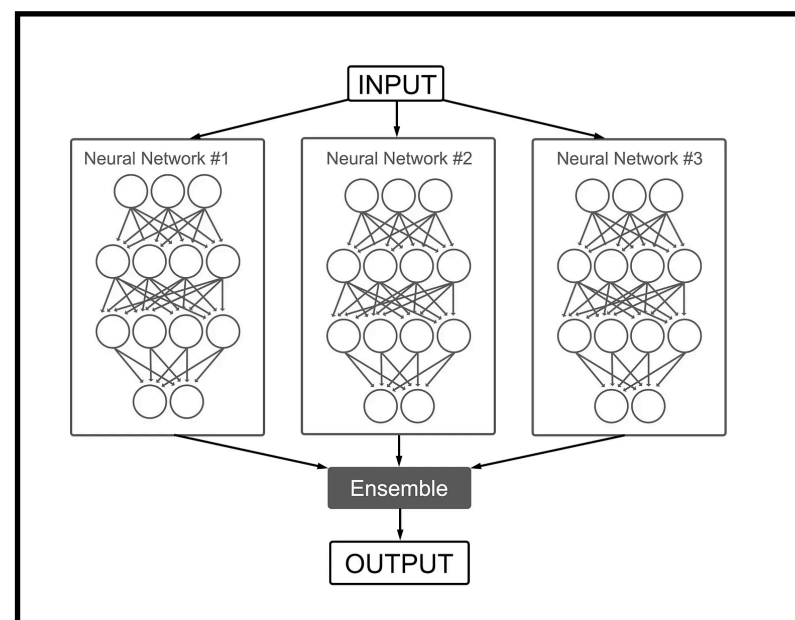


Workflows for the computational challenge

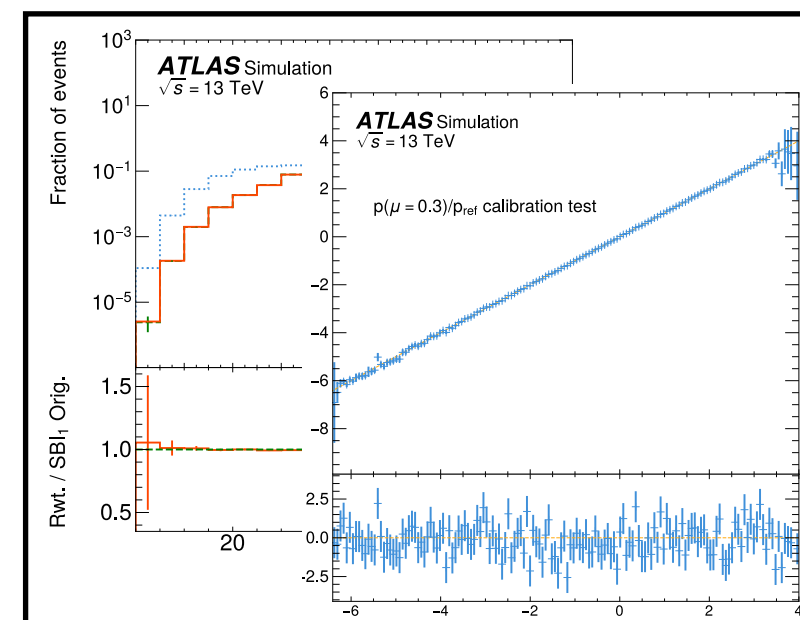
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- Careful design and validation of 10,000 networks per measurement is infeasible

Need fully automated workflows with Ray:

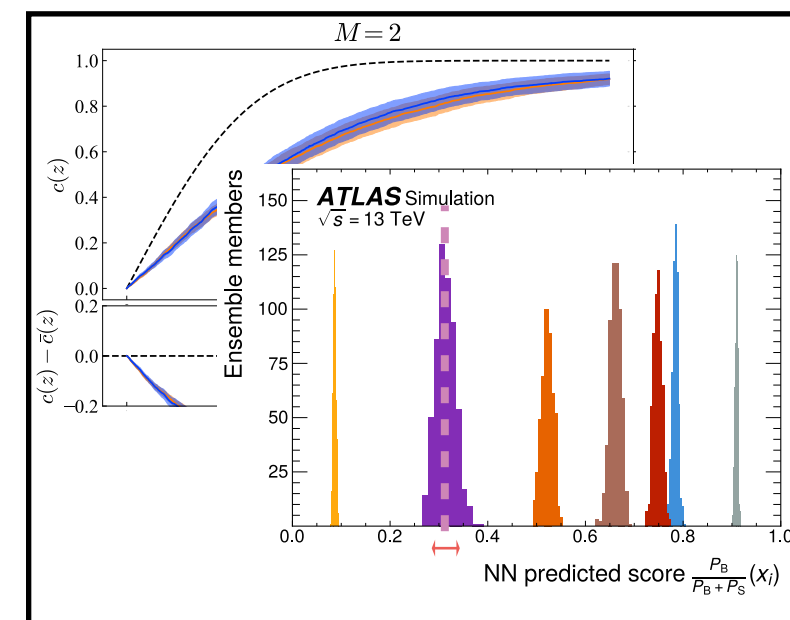
Train ensembles on HPC



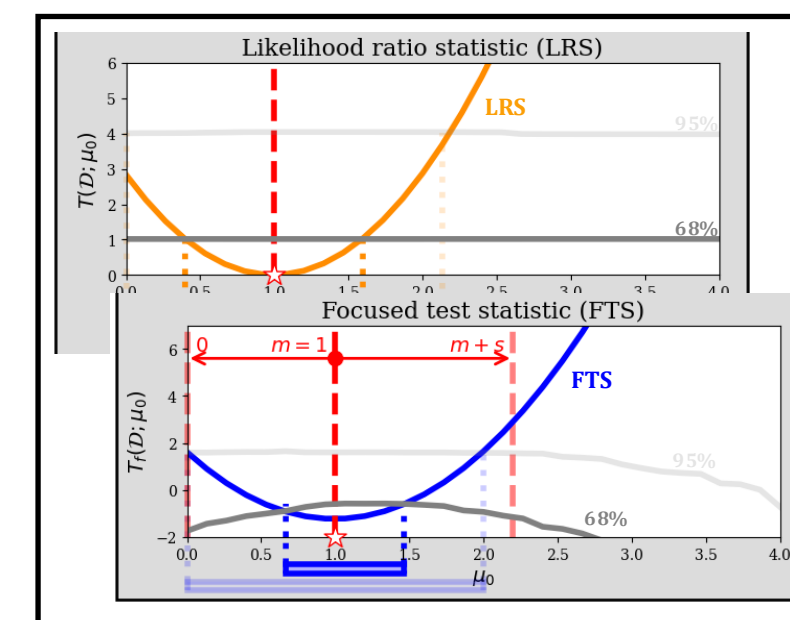
Validation of density estimates



Uncertainty quantification



Neyman inversion with pseudo-experiments



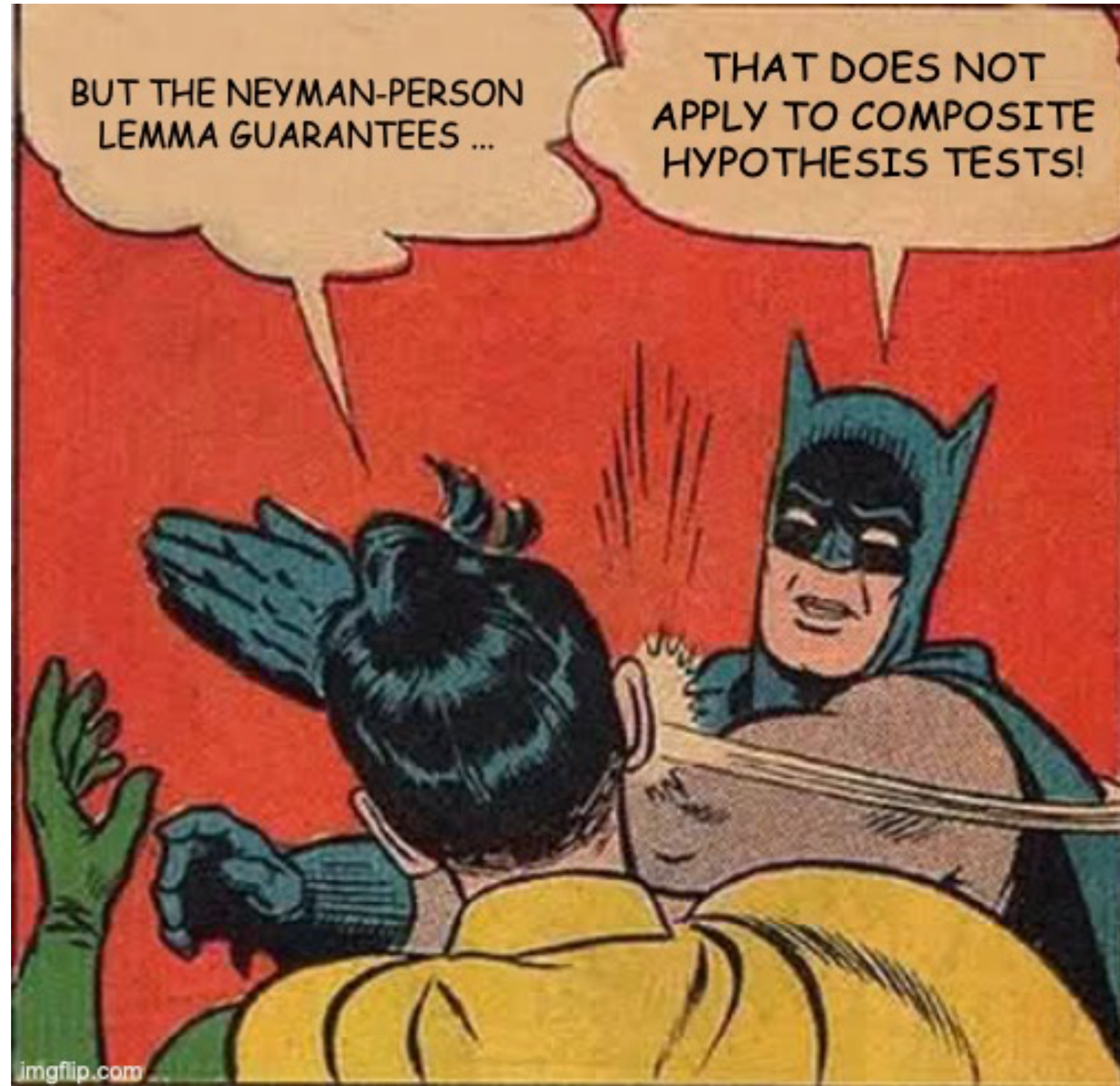
Optimize

Until now, we have replaced individual pieces with ML in age-old likelihood ratio test

Do we dare question the test itself?

Until now, we have replaced individual pieces with ML in age-old likelihood ratio test

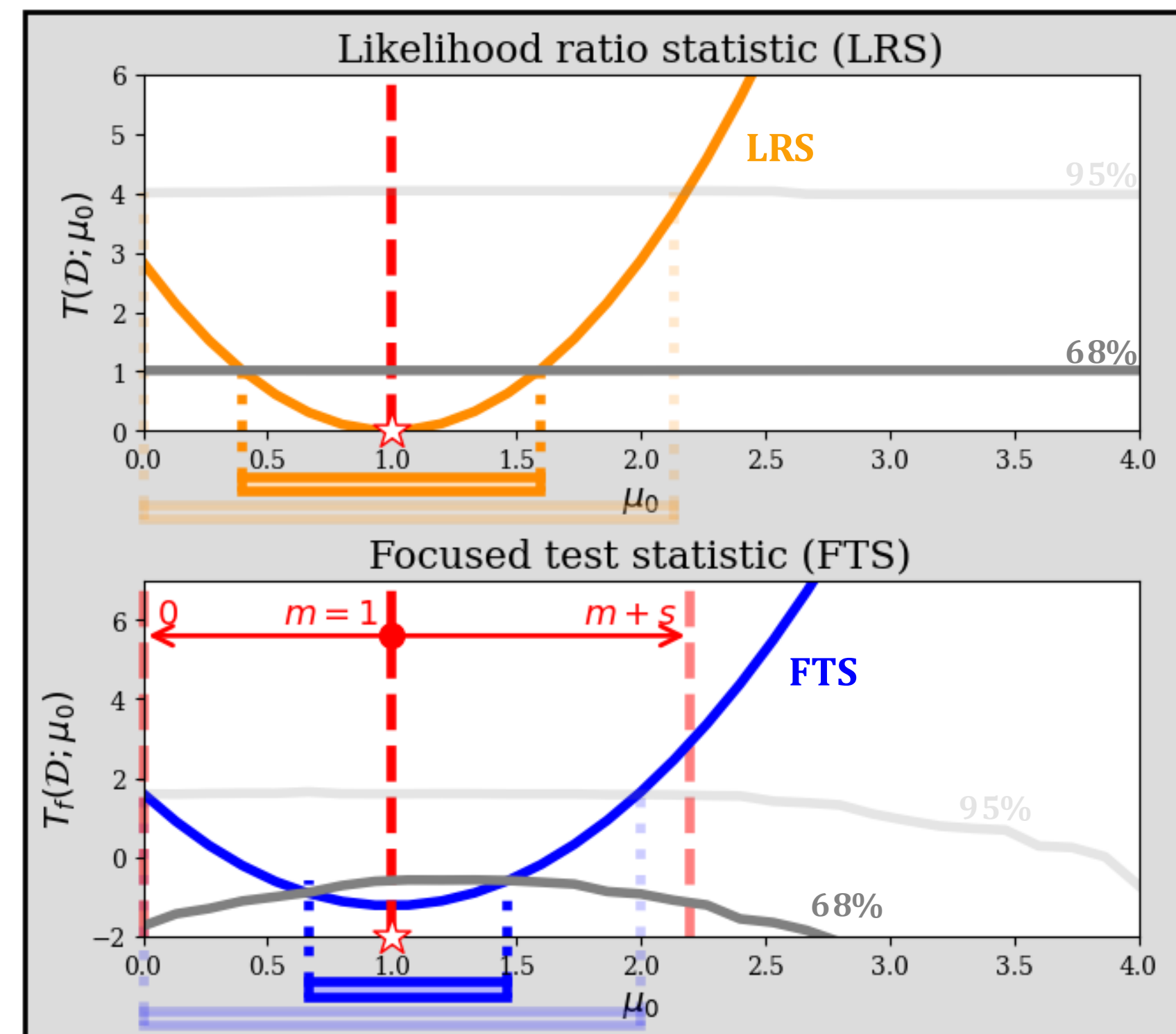
Do we dare question the test itself?



You *can* beat the likelihood ratio test statistic

$$LRS(\mathcal{D}; \mu_0) = -2 \log \left(\frac{p(\mathcal{D} | \mu_0)}{\sup_{\mu \in \Theta} p(\mathcal{D} | \mu)} \right) \longrightarrow FTS(\mathcal{D}; \mu_0) = -2 \log \left(\frac{p(\mathcal{D} | \mu_0)}{\int_{\Theta} p(\mathcal{D} | \mu) f(\mu) d\mu} \right)$$

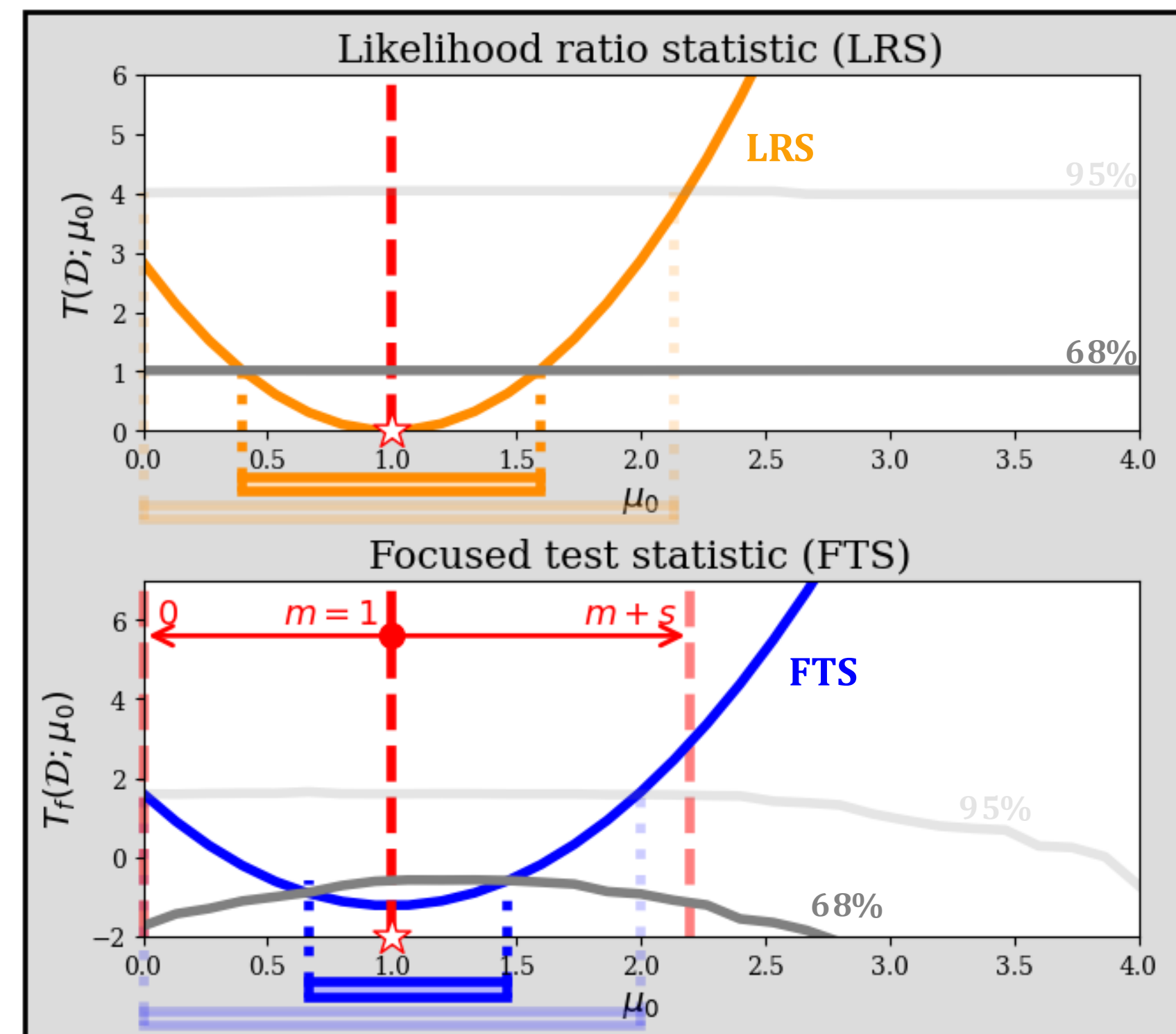
- Denominator in ‘Focused Test Statistic’ (FTS) knows about all alternate hypotheses



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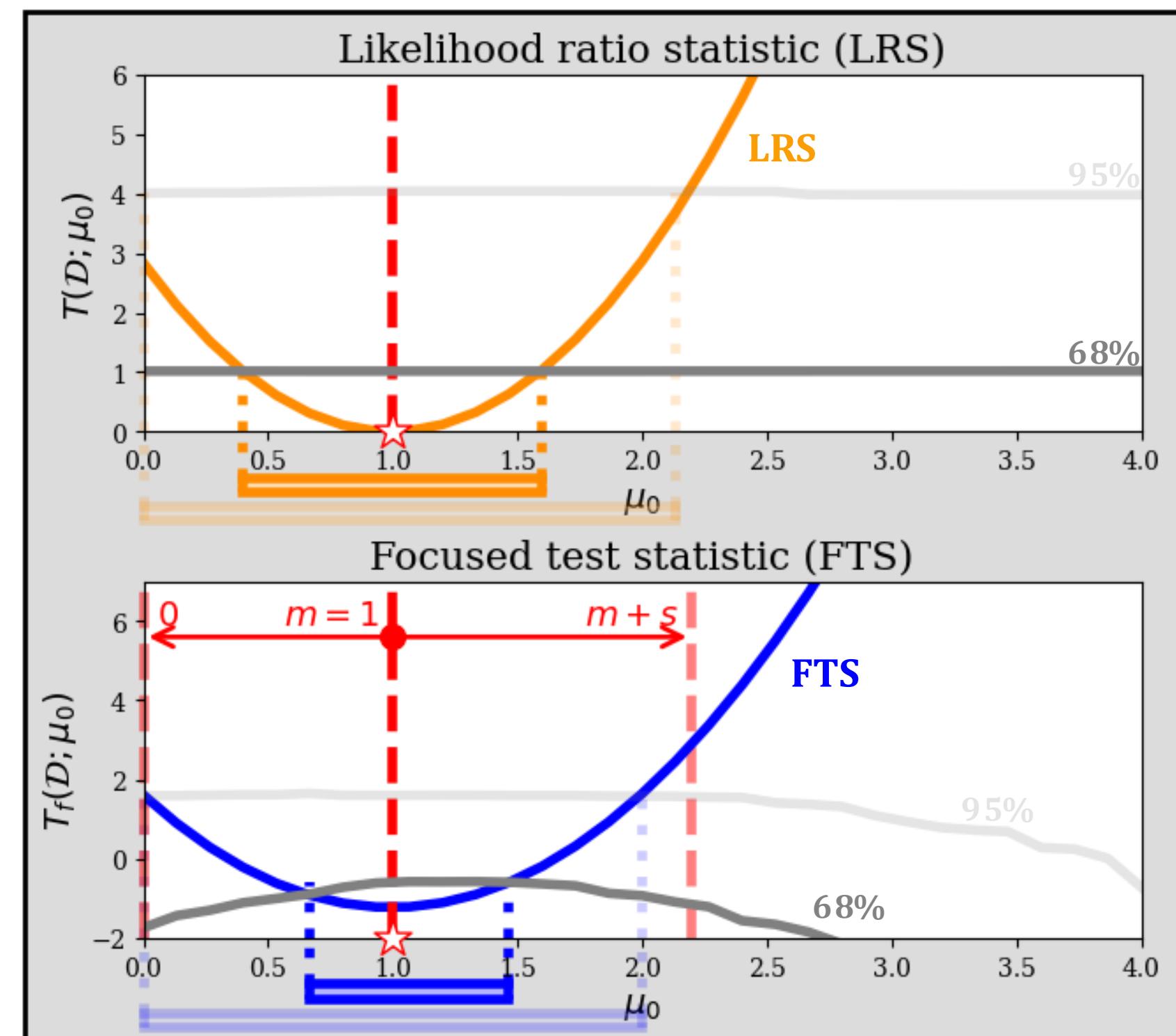


Shorter median length for confidence intervals with FTS even in 'asymptotic regime' where Wilks' applies

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- Denominator in ‘Focused Test Statistic’ (FTS) knows about all alternate hypotheses
- $f(\mu)$ focuses statistical power in meaningful regions of parameter space
 - Particularly useful in small sample / small signal regime
- Fast critical value estimation with ML

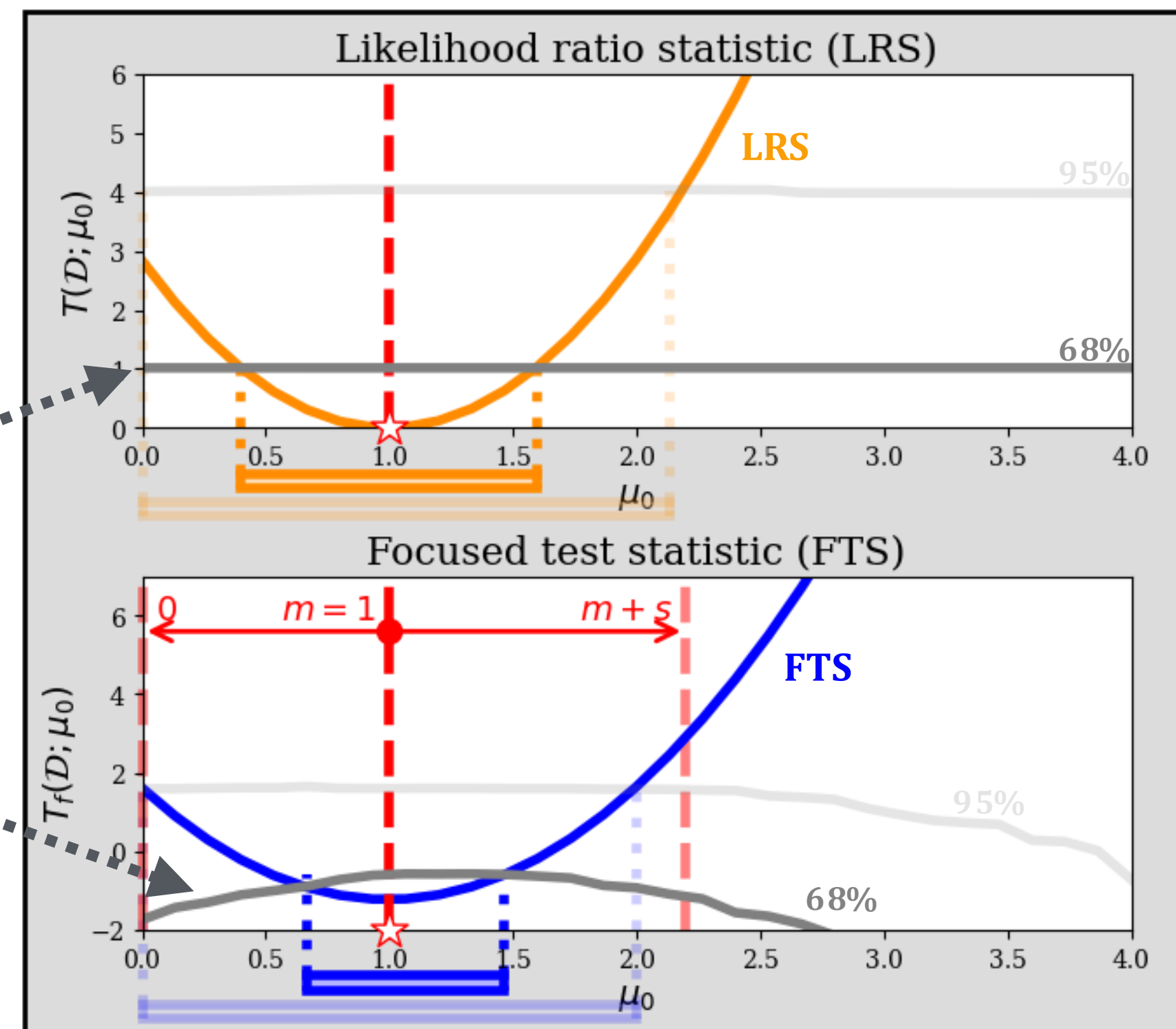


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Shorter median length for confidence intervals with FTS even in ‘asymptotic regime’ where Wilks’ applies

funding gets harder to secure, principal investigators are in their office writing grants while the trainees get to do the cool stuff.

Bryan W. Jones is a retinal neuroscientist at the University of Pittsburgh in Pennsylvania.

AISHIK GHOSH **STUDENTS OVERTURN** **LONG-HELD ASSUMPTION**

I have worked on experimental particle physics since 2015, searching for Higgs bosons at CERN, Europe's particle-physics lab near Geneva, Switzerland, and now also working on the Deep Underground Neutrino Experiment (DUNE) in the United States. For this research, there's one statistical test we've used for decades to confirm the existence of a new particle – the generalized likelihood ratio test (GLRT). This compares two models – a simple null hypothesis, which includes no new particle or matter being discovered, and a more complicated alternative model, which includes a new particle with many possible values of strength.

In December 2024, a couple of PhD students working with my collaborator, Ann Lee, a data scientist at Carnegie Mellon University in Pittsburgh, Pennsylvania, were confident they could disprove the assumption that the GLRT was optimal. In the corner of my mind, I hoped they would prove us wrong. I gave them one of the most famous Higgs boson data sets to play around with. By early 2025, they showed that, although our previous physics results weren't wrong, our use of the GLRT wasn't ideal because it assumed large sample sizes

are always generated, which is often not the case. Instead, the test left valuable information on the table. That day was special. I was still sceptical and I went through a battery of checks because I had to go back to my community and defend the PhD students' work, but it was all correct. The paper is currently in review, receiving a great deal of scrutiny.

Together, we produced a statistical test that will drastically improve our ability to make discoveries in particle physics, for example in searches for a new particle such as dark matter, where we expect to see only a few signal events at best. As a scientist, I want deeply held beliefs to be questioned. It was a real shock to the particle-physics community. Young people find it exciting. Senior members are still highly sceptical, as they should be, but they are coming around. As the DUNE experiment comes online, with this new statistical model in place, we hope to make precise measurements about neutrinos much sooner than anticipated.

Aishik Ghosh is a fundamental physicist at the Georgia Institute of Technology in Atlanta.

RAFIK TAREK NEME GARRIDO **SHOCKING** **CORAL FIND**

A couple of years ago, after a day of pouring rain, the water on the Caribbean coast of Colombia was crystal clear and my master's student, Jorge Mareno, managed to take pictures of corals that no one knew existed here. We could find no scientific reports of corals in the area. Typically, the water is pretty turbid because the

Magdalena River, which flows from the south of the country to the Caribbean Sea, brings chemicals and pollutants. It's an ongoing ecological and social challenge, but these corals must be adapting to these conditions. We did a sampling campaign across three days with a boat, using environmental DNA to find areas where corals, sponges and fish successfully survive the conditions. Most of the records are completely new for the region. It's super gratifying.

Rafik Tarek Neme Garrido is an evolutionary biologist at the University of the North in Barranquilla, Colombia.

TIM CURRAN **BURN** **PREDICTIONS**

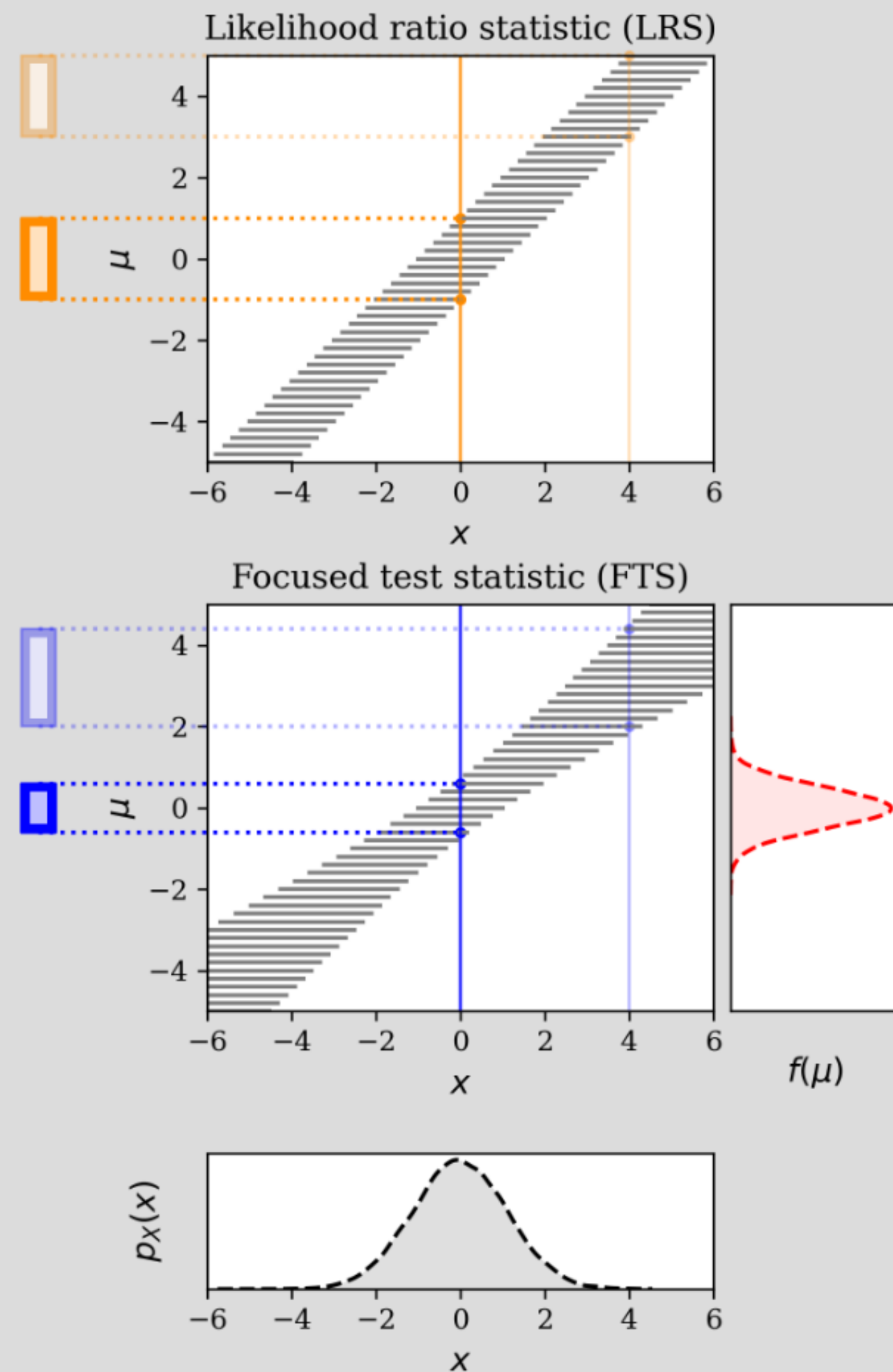
In my group, we test the flammability of plant species using a barbecue. The results can help with fire-mitigation policies and with understanding the evolution of flammability. As part of an outreach activity, we host school-children at the university who haven't had much exposure to academia before. We ask the kids to predict how a particular plant species will behave – for example, what characteristics will make it burn less or more – and then we see who is right. The kids get really into it. They ask amazing questions, the same kind that peer reviewers have asked us, including questioning our methodological assumptions, such as "why do you only blowtorch them for ten seconds?"

Most of the really good days doing science have been associated with young students having a light-bulb moment. In the rather



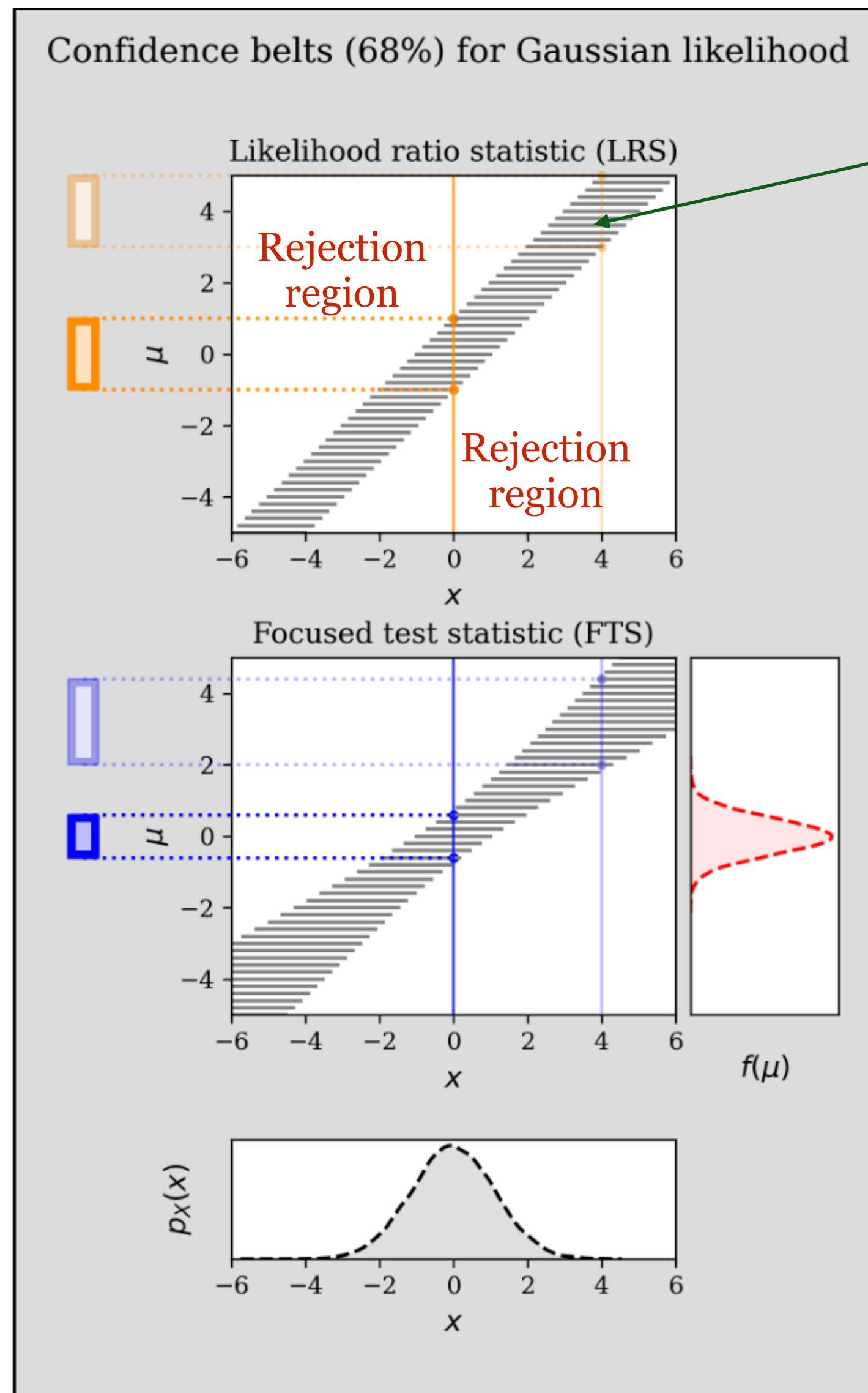
Test inversion for simplest measurement

Confidence belts (68%) for Gaussian likelihood



Data comprises single measurement x

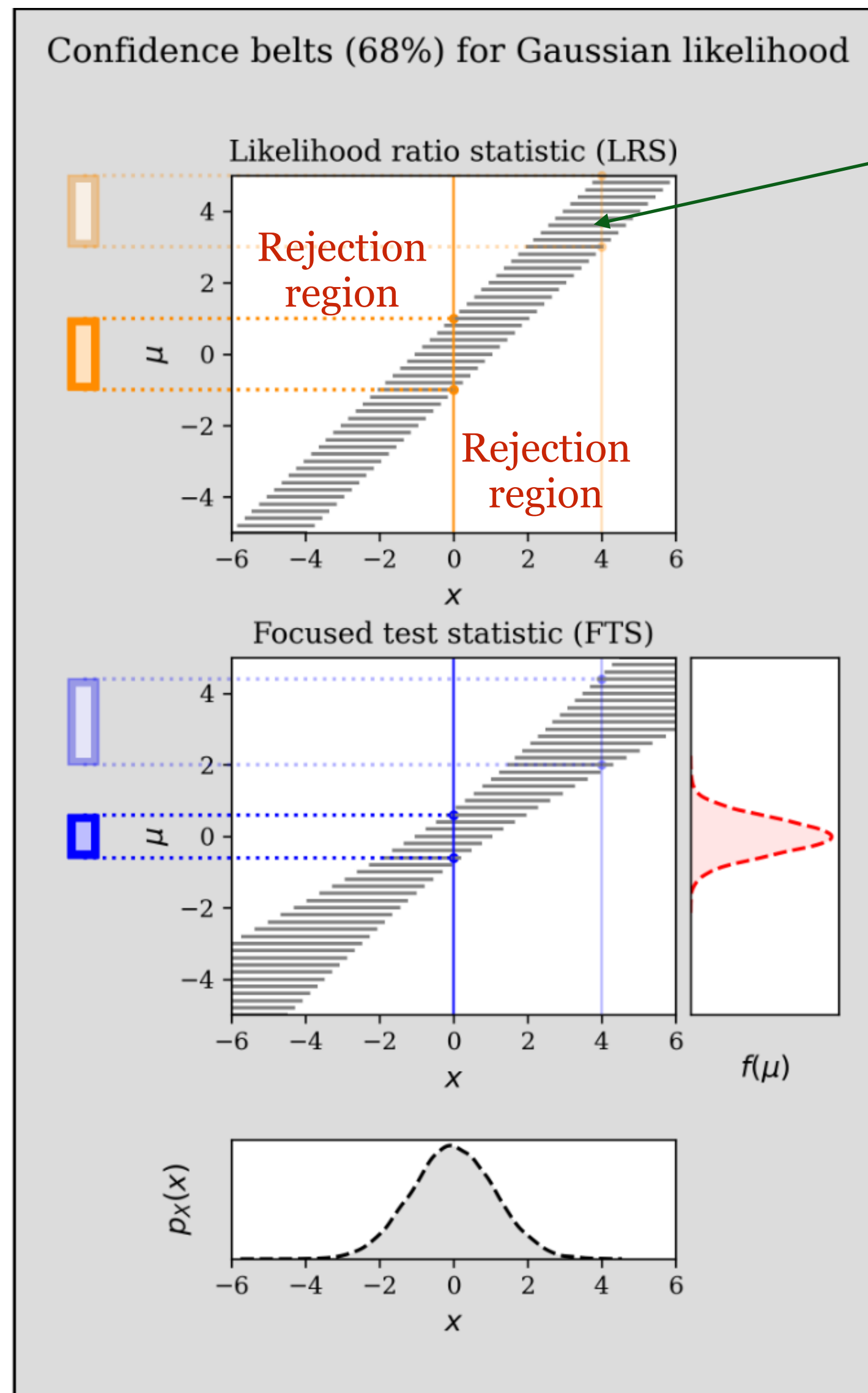
Test inversion for simplest measurement



Acceptance region

Data comprises single measurement x

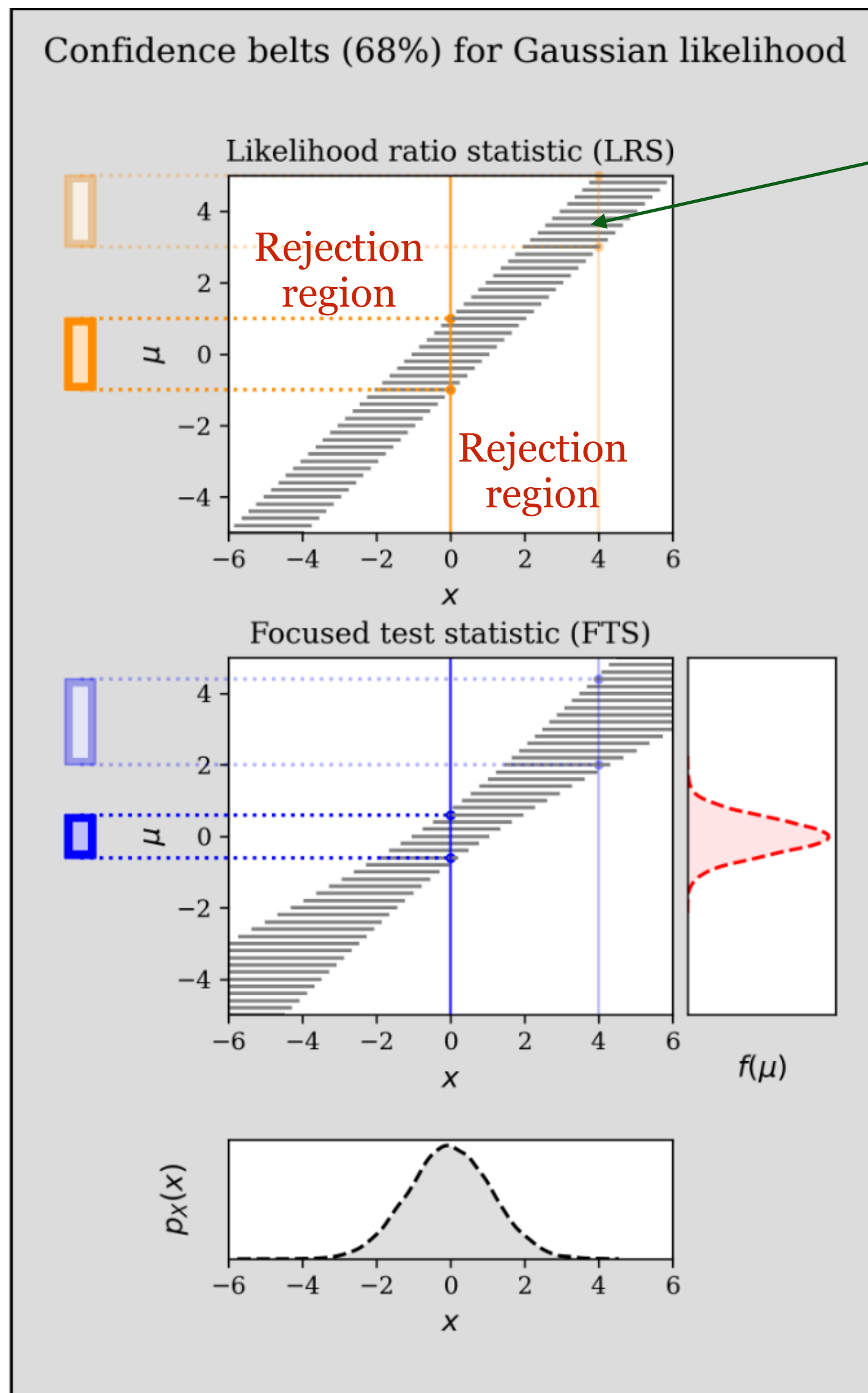
Test inversion for simplest measurement



Data comprises single measurement x

- LRS gives constant CI sizes
- FTS gives narrower near focus region by sacrificing power away from focus region

Test inversion for simplest measurement



Data comprises single measurement x

- LRS gives constant CI sizes
- FTS gives narrower near focus region by sacrificing power away from focus region
- You have full freedom to **choose any focus function** that gives you the **best expected sensitivity** in physics-motivated regions using simulated samples

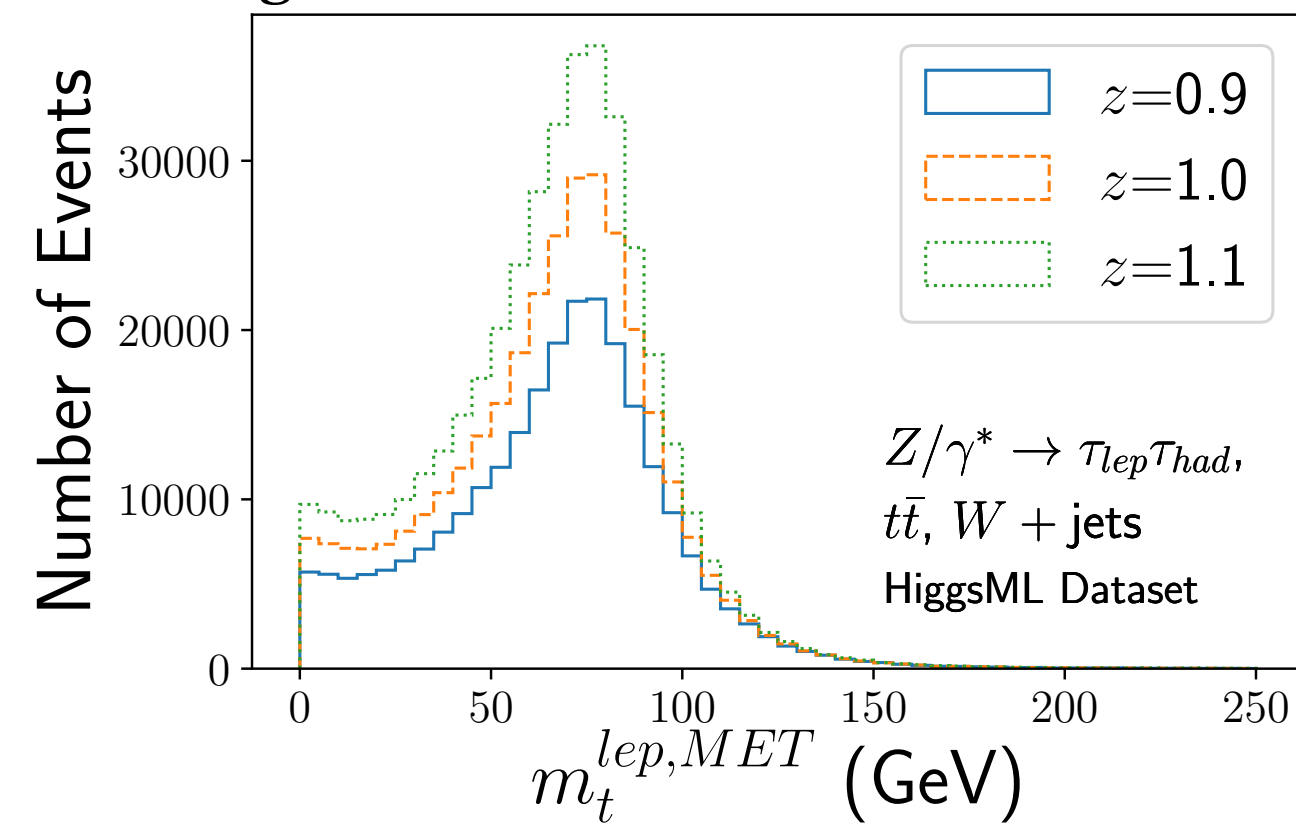
Open problems to extend to full ATLAS analysis:

- ✓ Robustness: Design and validation
- ▶ **Systematic Uncertainties: Incorporate them in likelihood (ratio) model**
- Neyman Construction: Throwing toys in a per-event analysis

Systematic uncertainties

Experimental uncertainties:

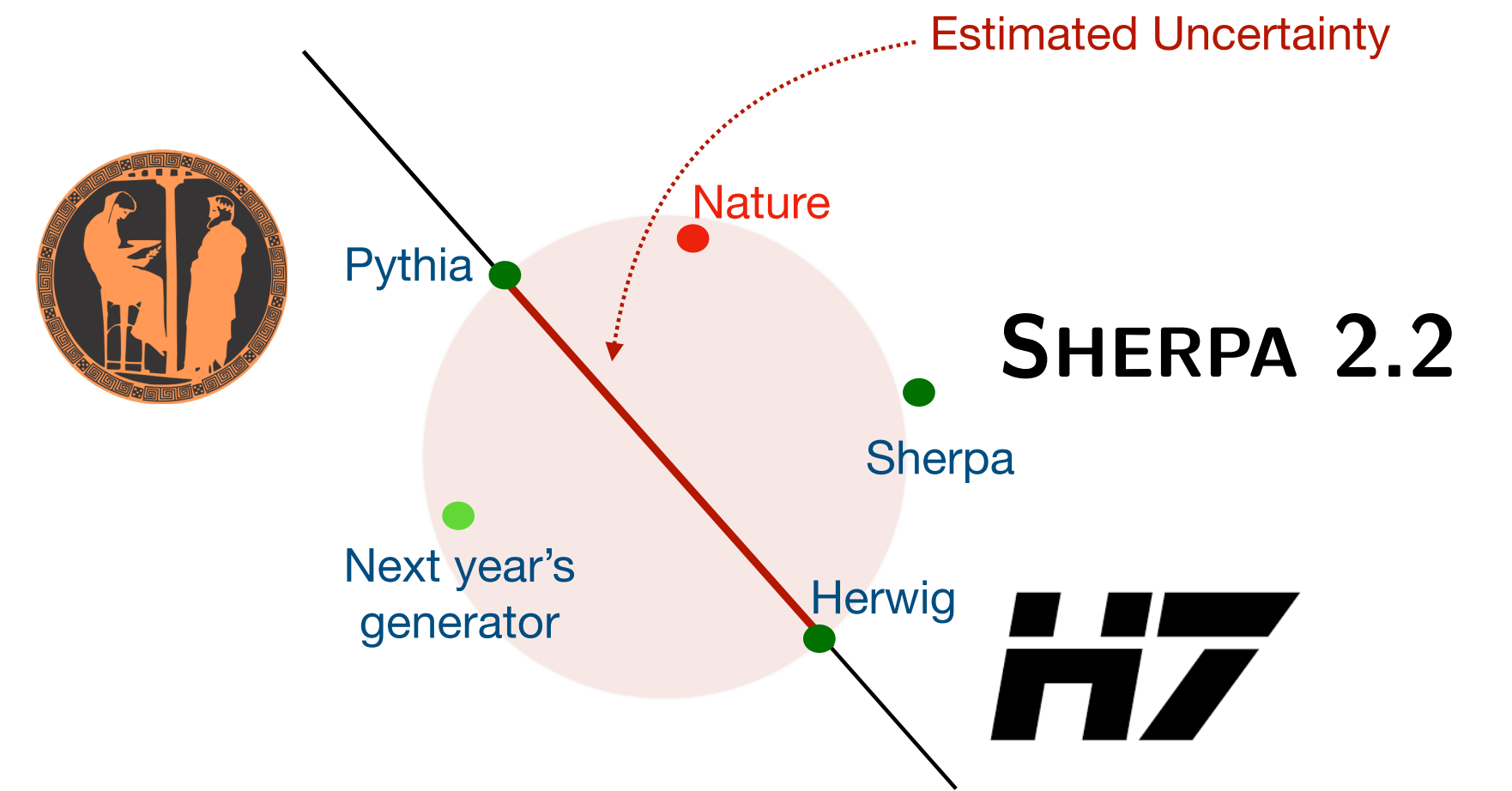
Eg. Inaccuracies in the calibration of our detector



Ghosh et al: [PhysRevD.104.056026](#)

Theory uncertainties:

Eg. Inability to compute QFT to infinite order

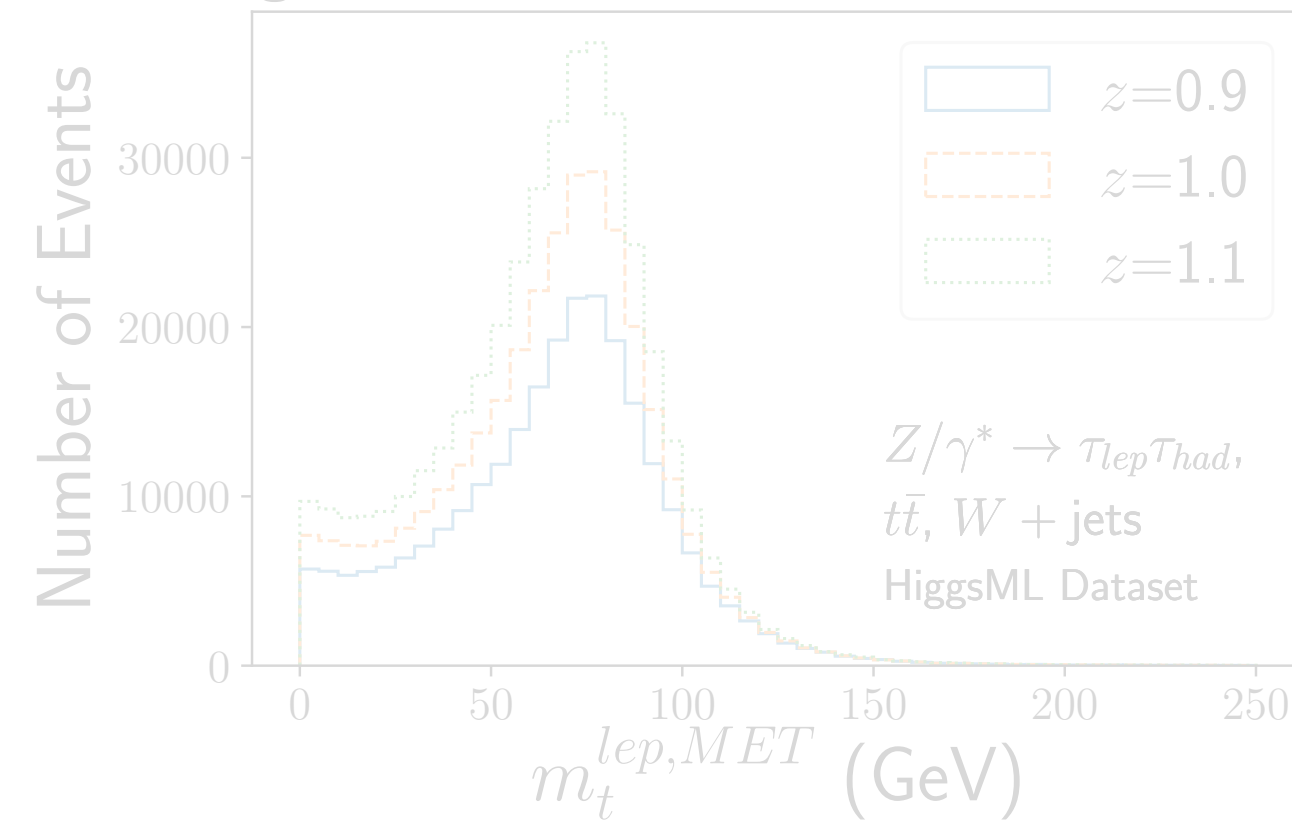


Ghosh and Nachman: [epjc.s10052.022.10012.w](#)

Systematic uncertainties

Experimental uncertainties:

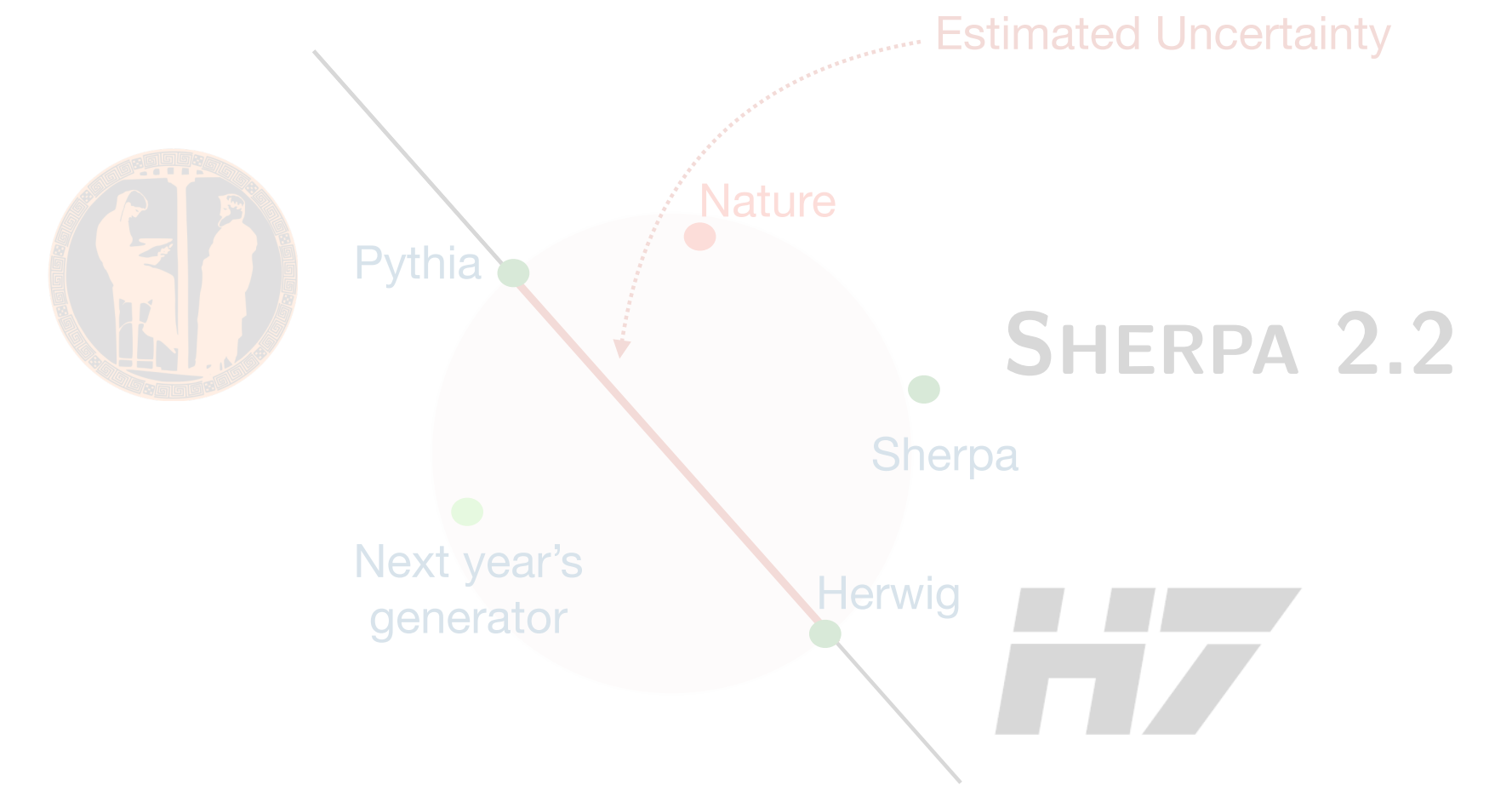
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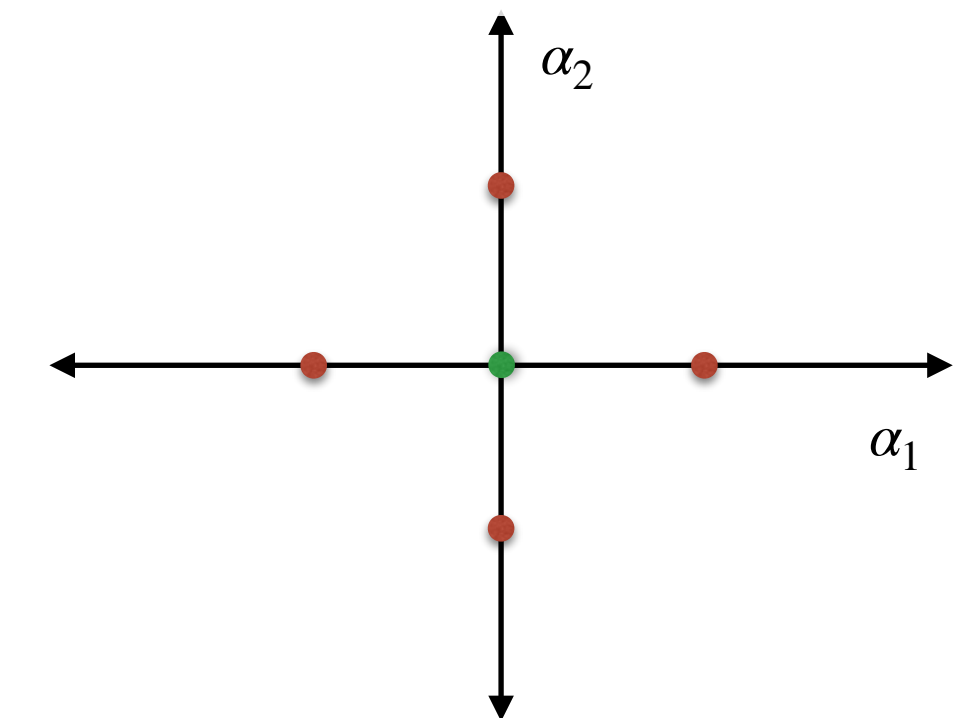
Theory uncertainties:

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Ghosh and Nachman: [epjc.s10052.022.10012.w](#)

- We only have simulations at 3 variations of each nuisance parameter α_k



Known interpolation strategies

See formula used in [backup](#)

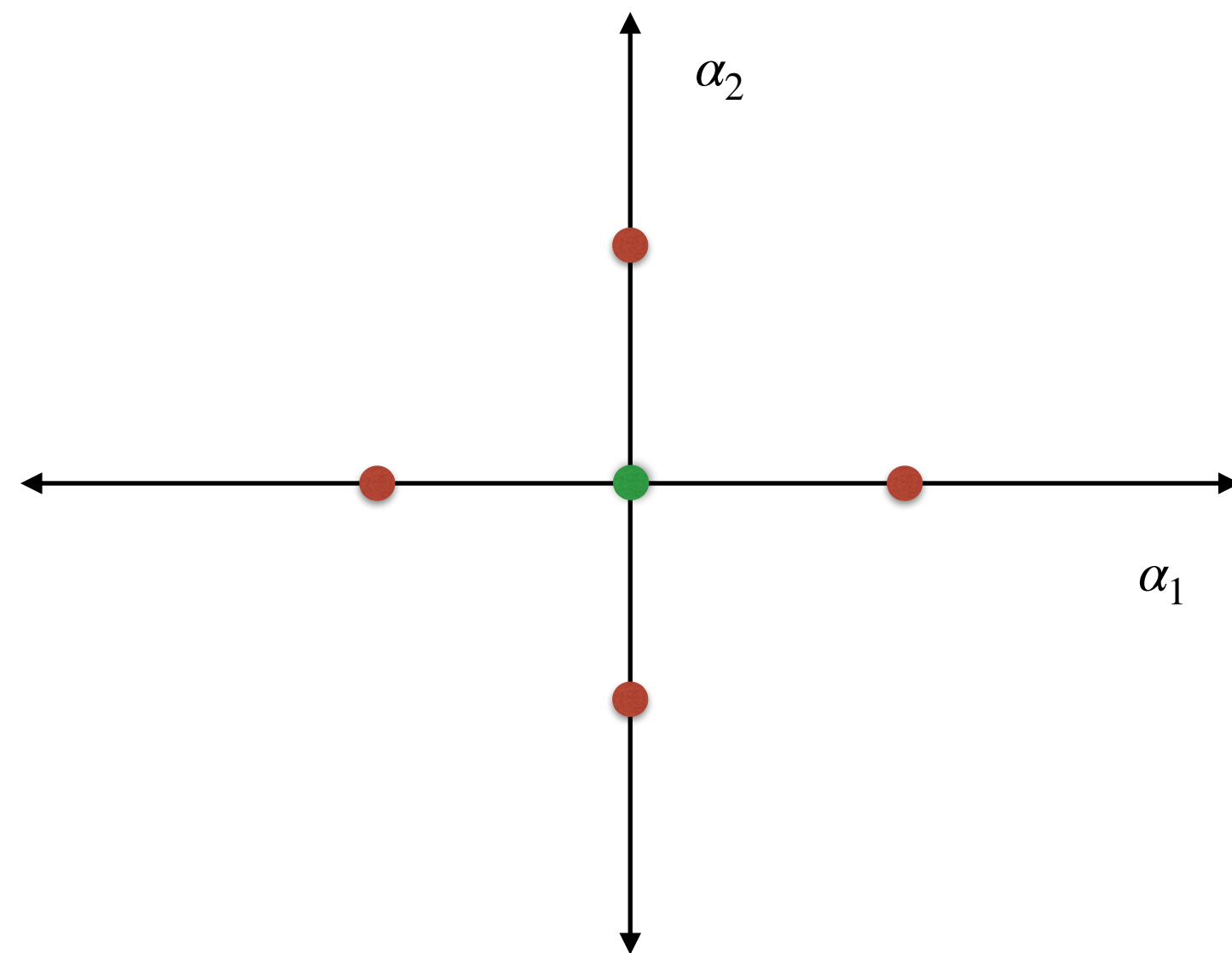
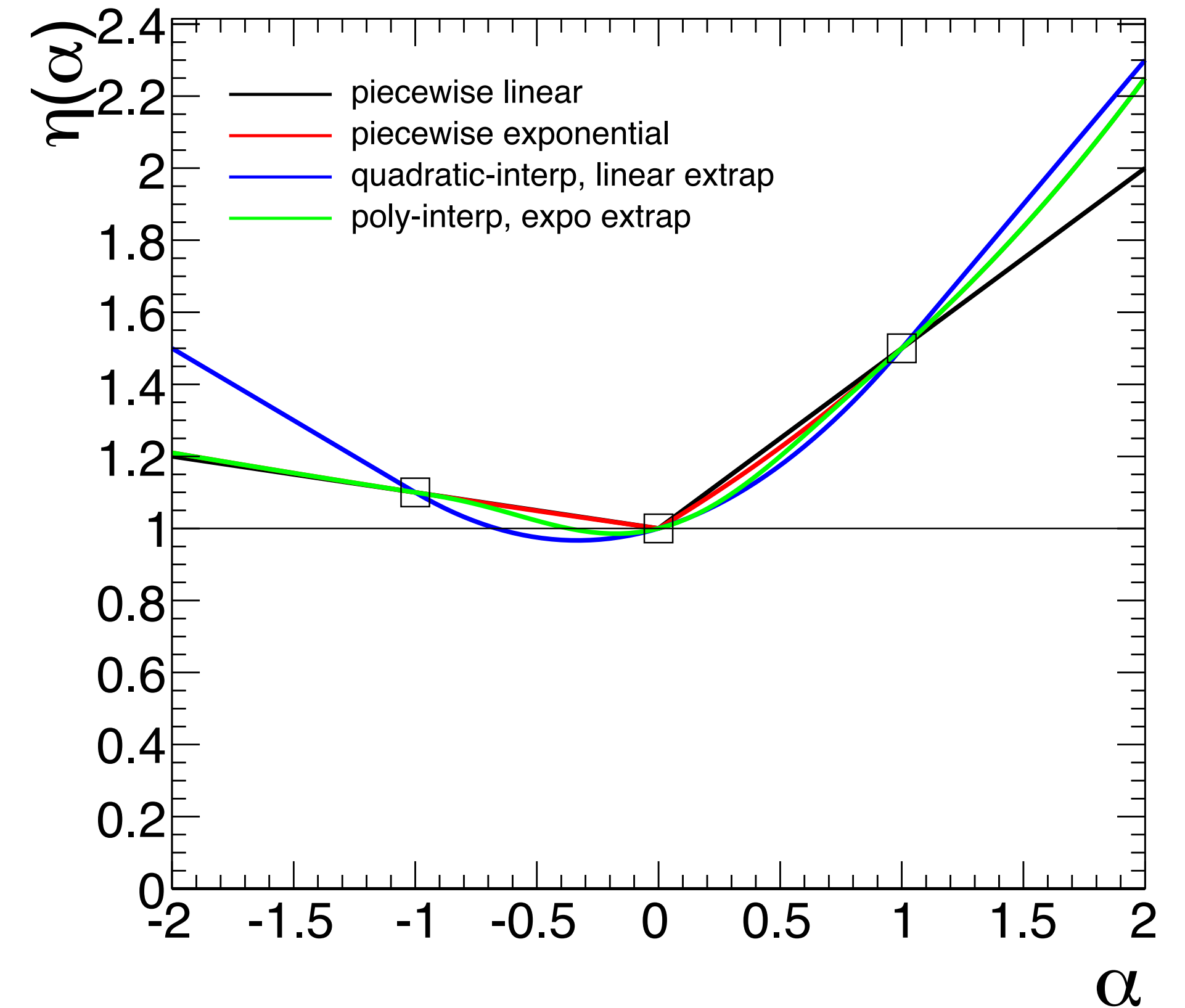


Image: arXiv:1503.07622

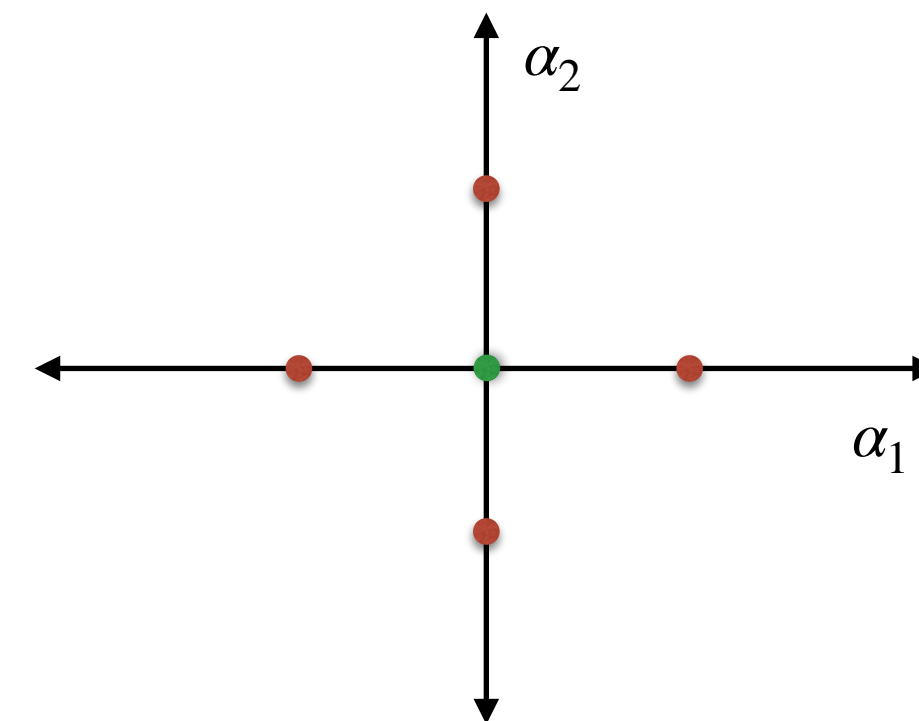


⇒ Combine these traditional interpolation with neural network estimation of per-event likelihood ratios

Probability density ratio including nuisance parameters (α)

x_i is one individual event

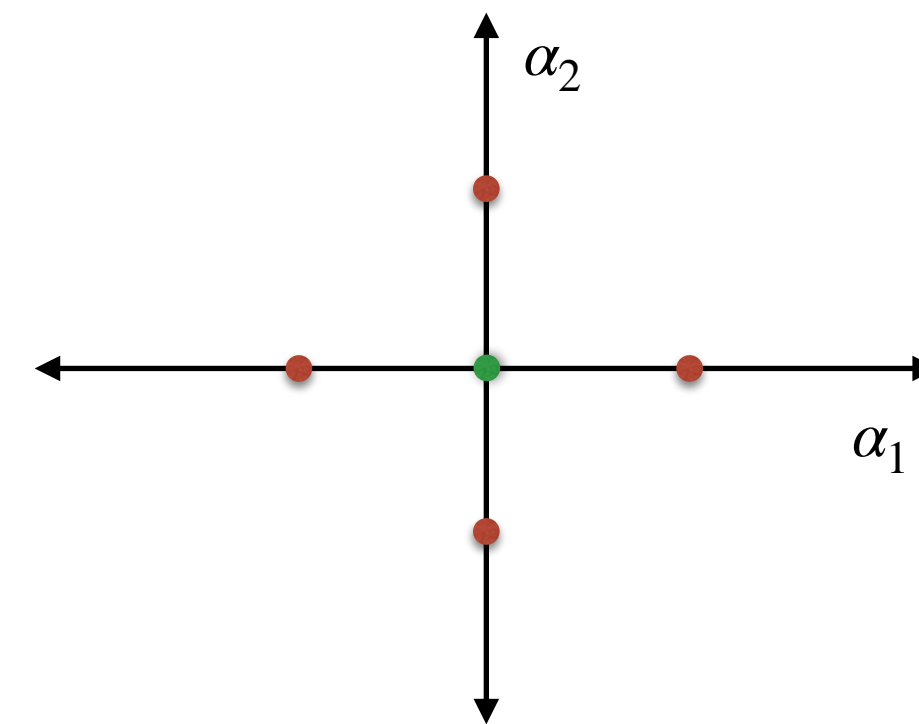
$$\frac{p(x_i | \mu, \alpha)}{p_{ref}(x_i)} =$$



Probability density ratio including nuisance parameters (α)

x_i is one individual event

$$\frac{p(x_i | \mu, \alpha)}{p_{ref}(x_i)} = \frac{1}{\nu(\mu, \alpha)} \sum_j^C f_j(\mu) \cdot \nu_j \cdot \frac{p_j(x_i)}{p_{ref}(x_i)} \cdot \prod_k^{N_{syst}} G_j(\alpha_k) \cdot g_j(x_i, \alpha_k)$$



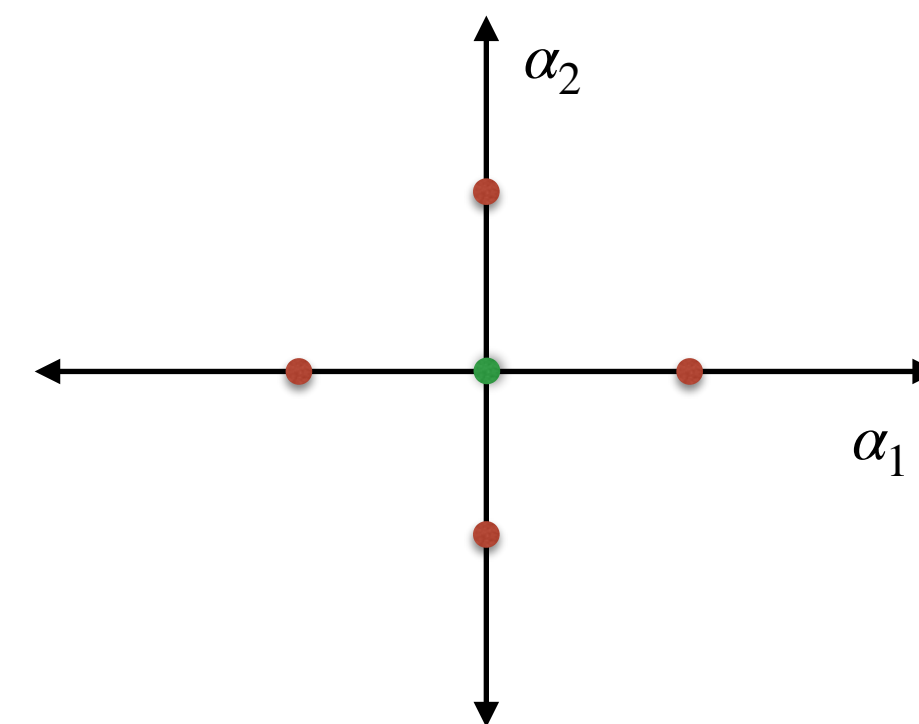
$$g_j(x_i, \alpha_k) = \frac{p_j(x_i, \alpha_k)}{p_j(x_i)}$$

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We have this already



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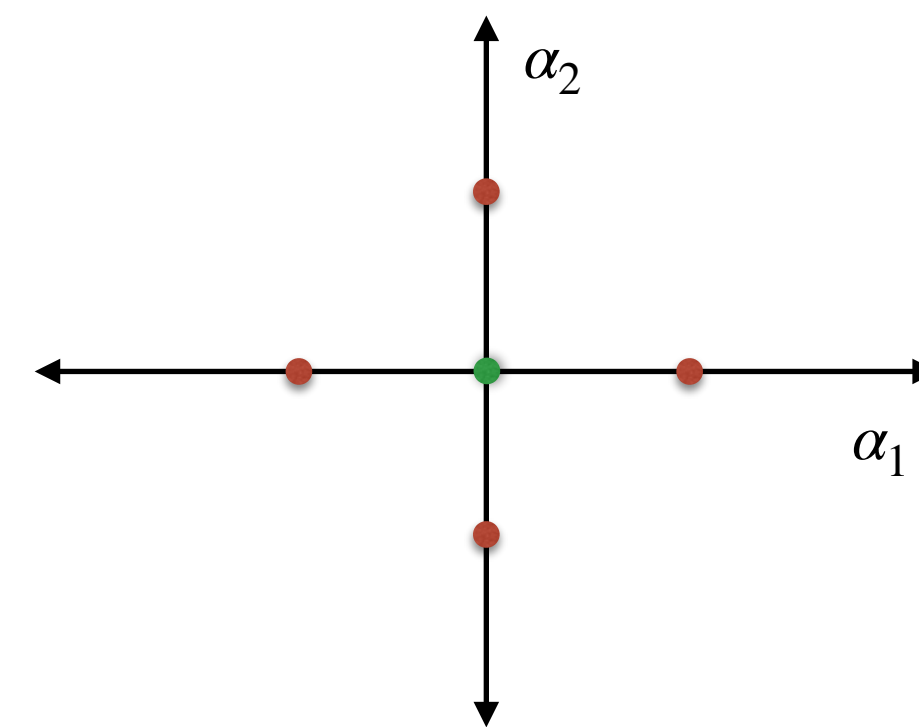
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We have this already

Estimate from simulations and existing interpolation methods



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Probability density ratio including nuisance parameters (α)

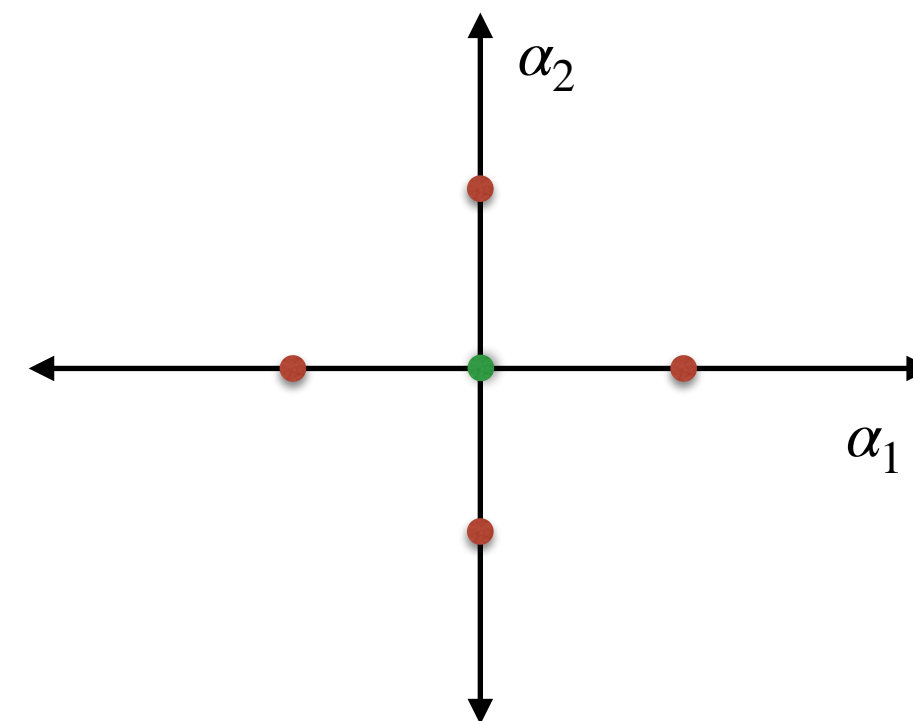
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We have this already

Per-event terms estimated using another ensemble of networks and interpolation methods

Estimate from simulations and existing interpolation methods



$$g_j(x_i, \alpha_k) = \frac{p_j(x_i, \alpha_k)}{p_j(x_i)}$$

Final test statistic

x_i is one individual event

$$\frac{L_{\text{full}}(\mu, \alpha | \mathcal{D})}{L_{\text{ref}}(\mathcal{D})} = \text{Pois}(N_{\text{data}} | \nu(\mu, \alpha)) \prod_i^{N_{\text{data}}} \frac{p(x_i | \mu, \alpha)}{p_{\text{ref}}(x_i)} \prod_k \text{Gaus}(a_k | \alpha_k, \delta_k)$$

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From previous slide

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Prod over events

From previous slide

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Rate term (points to $\text{Pois}(N_{\text{data}} | \nu(\mu, \alpha))$)
Prod over events (points to $\prod_i^{N_{\text{data}}}$)
From previous slide (points to $\frac{p(x_i | \mu, \alpha)}{p_{\text{ref}}(x_i)}$)

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Rate term (points to $\text{Pois}(N_{\text{data}} | \nu(\mu, \alpha))$)
Prod over events (points to $\prod_i^{N_{\text{data}}}$)
From previous slide (points to $\frac{p(x_i | \mu, \alpha)}{p_{\text{ref}}(x_i)}$)
Constrain term (points to $\prod_k \text{Gaus}(a_k | \alpha_k, \delta_k)$)

Final test statistic

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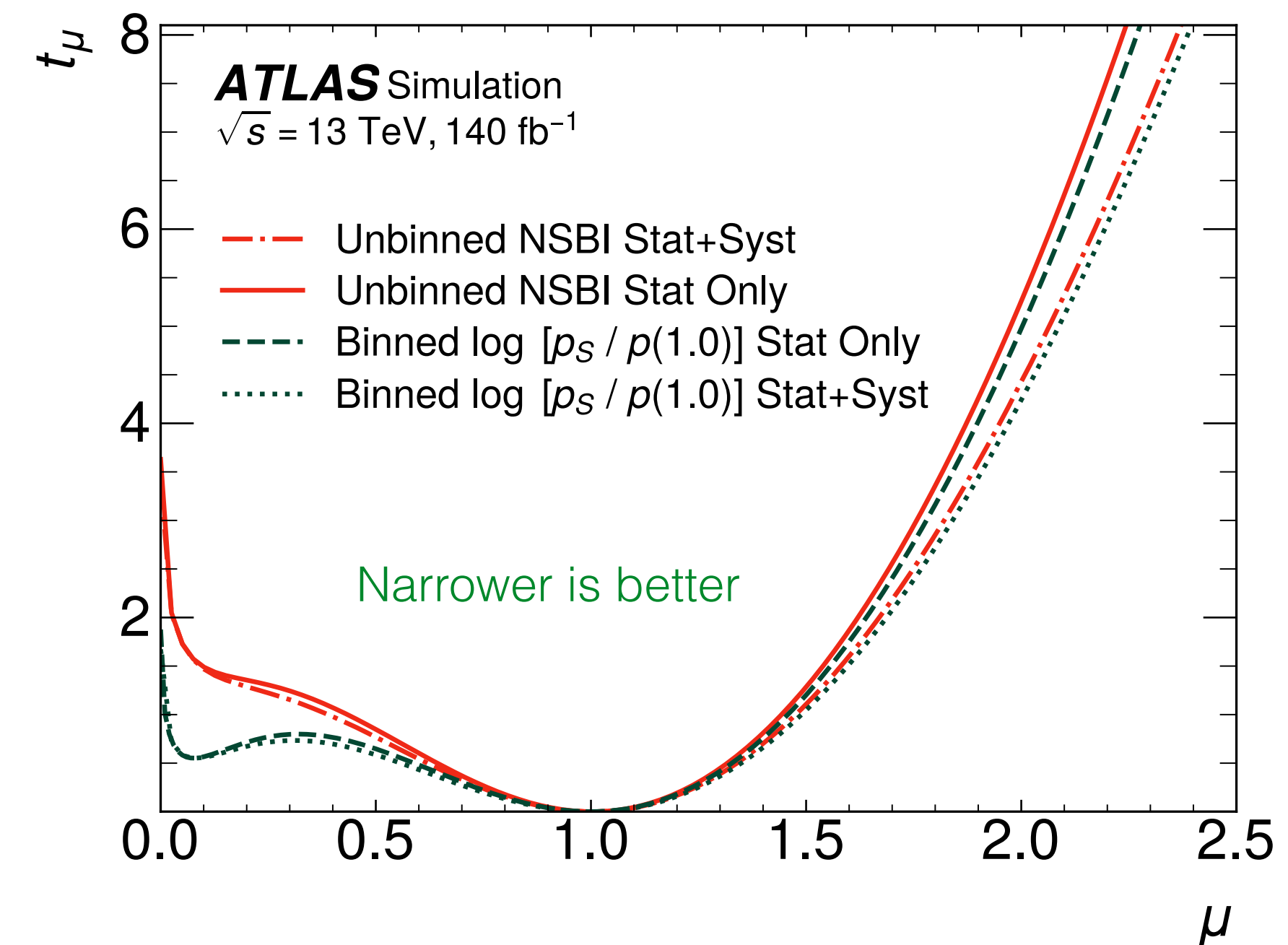
$$\frac{L_{\text{full}}(\mu, \alpha | \mathcal{D})}{L_{\text{ref}}(\mathcal{D})} = \text{Pois}(N_{\text{data}} | \nu(\mu, \alpha)) \prod_i^{N_{\text{data}}} \frac{p(x_i | \mu, \alpha)}{p_{\text{ref}}(x_i)} \prod_k \text{Gaus}(a_k | \alpha_k, \delta_k)$$

Rate term (points to Poisson term)
Prod over events (points to \prod_i)
From previous slide (points to $\frac{p(x_i | \mu, \alpha)}{p_{\text{ref}}(x_i)}$)
Constrain term (points to $\prod_k \text{Gaus}$)

Profiling:

$$t_\mu = -2 \ln \left(\frac{L_{\text{full}}(\mu, \hat{\hat{\alpha}}) / L_{\text{ref}}}{L_{\text{full}}(\hat{\mu}, \hat{\alpha}) / L_{\text{ref}}} \right)$$

This is why we define p_{ref} to be independent of μ



Final test statistic

x_i is one individual event

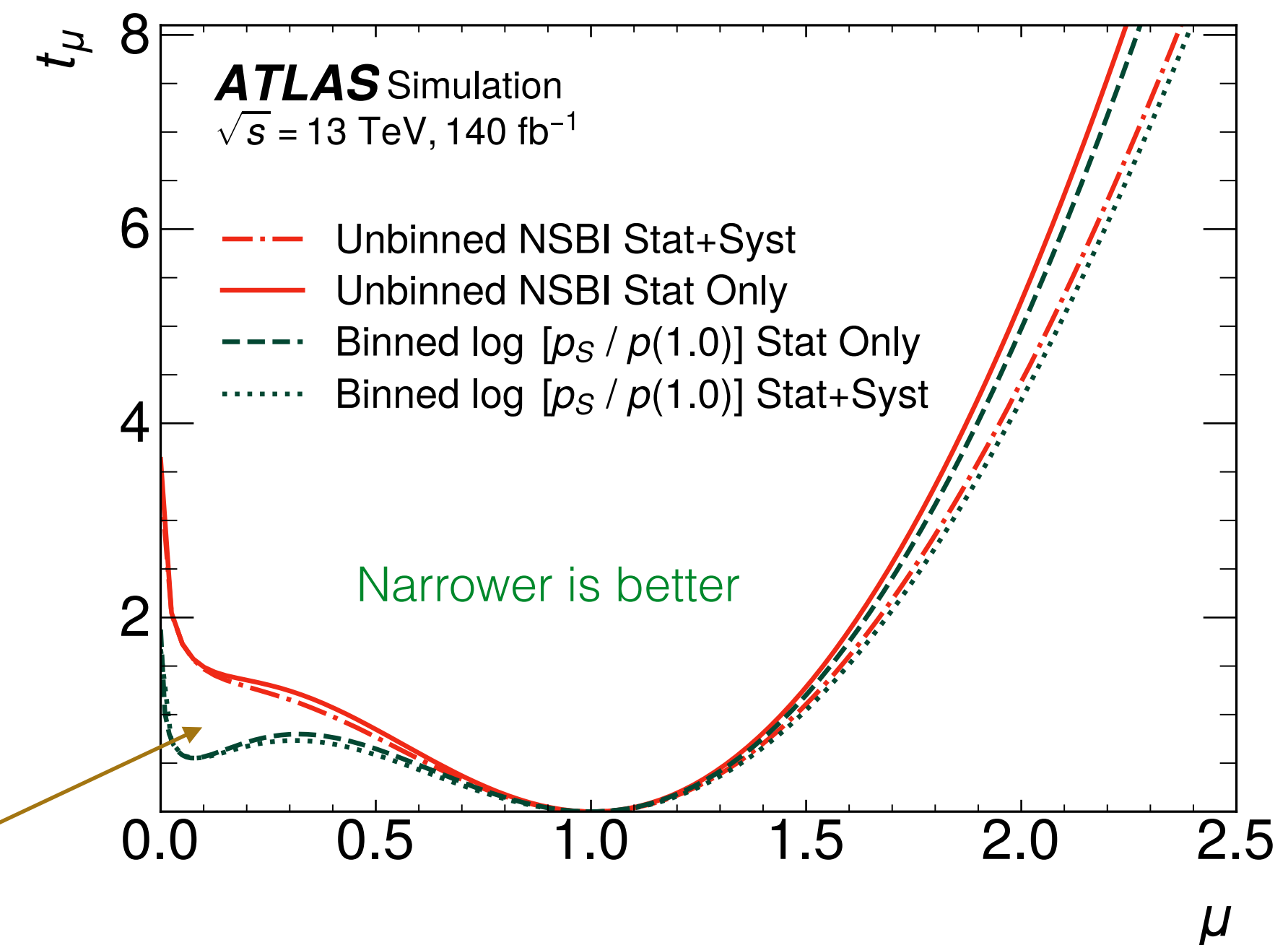
$$\frac{L_{\text{full}}(\mu, \alpha | \mathcal{D})}{L_{\text{ref}}(\mathcal{D})} = \text{Pois}(N_{\text{data}} | \nu(\mu, \alpha)) \prod_i^{N_{\text{data}}} \frac{p(x_i | \mu, \alpha)}{p_{\text{ref}}(x_i)} \prod_k \text{Gaus}(a_k | \alpha_k, \delta_k)$$

Rate term (points to Poisson term)
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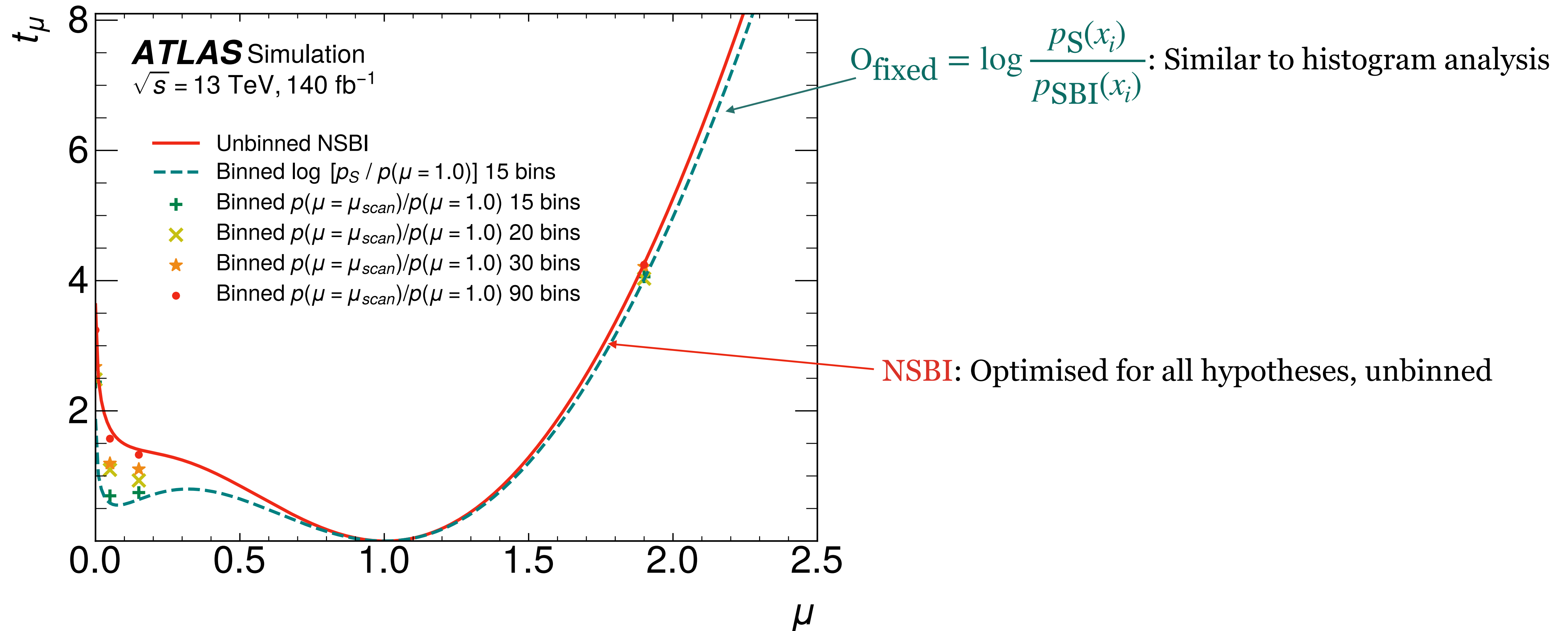
This is why we define p_{ref} to be independent of μ



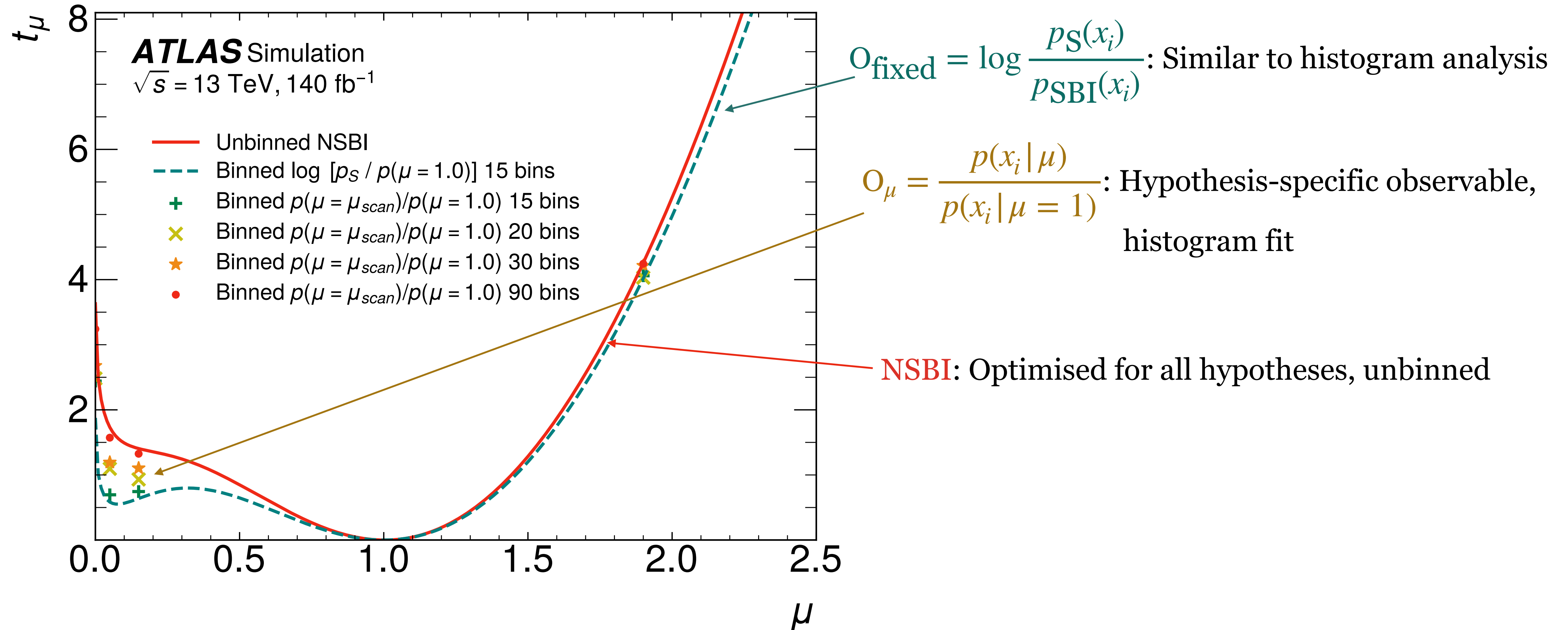
Non-parabolic shape due to non-linear effects from quantum interference

Why does NSBI work better than traditional analyses?

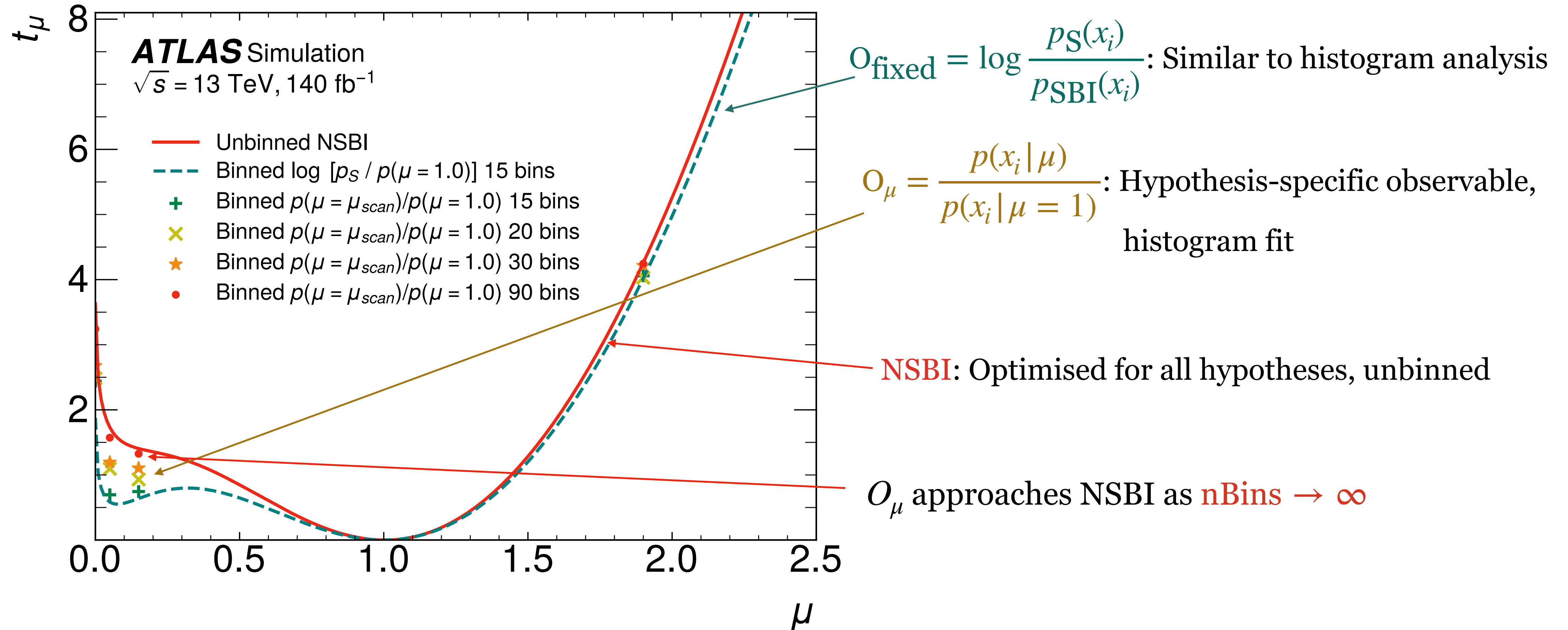
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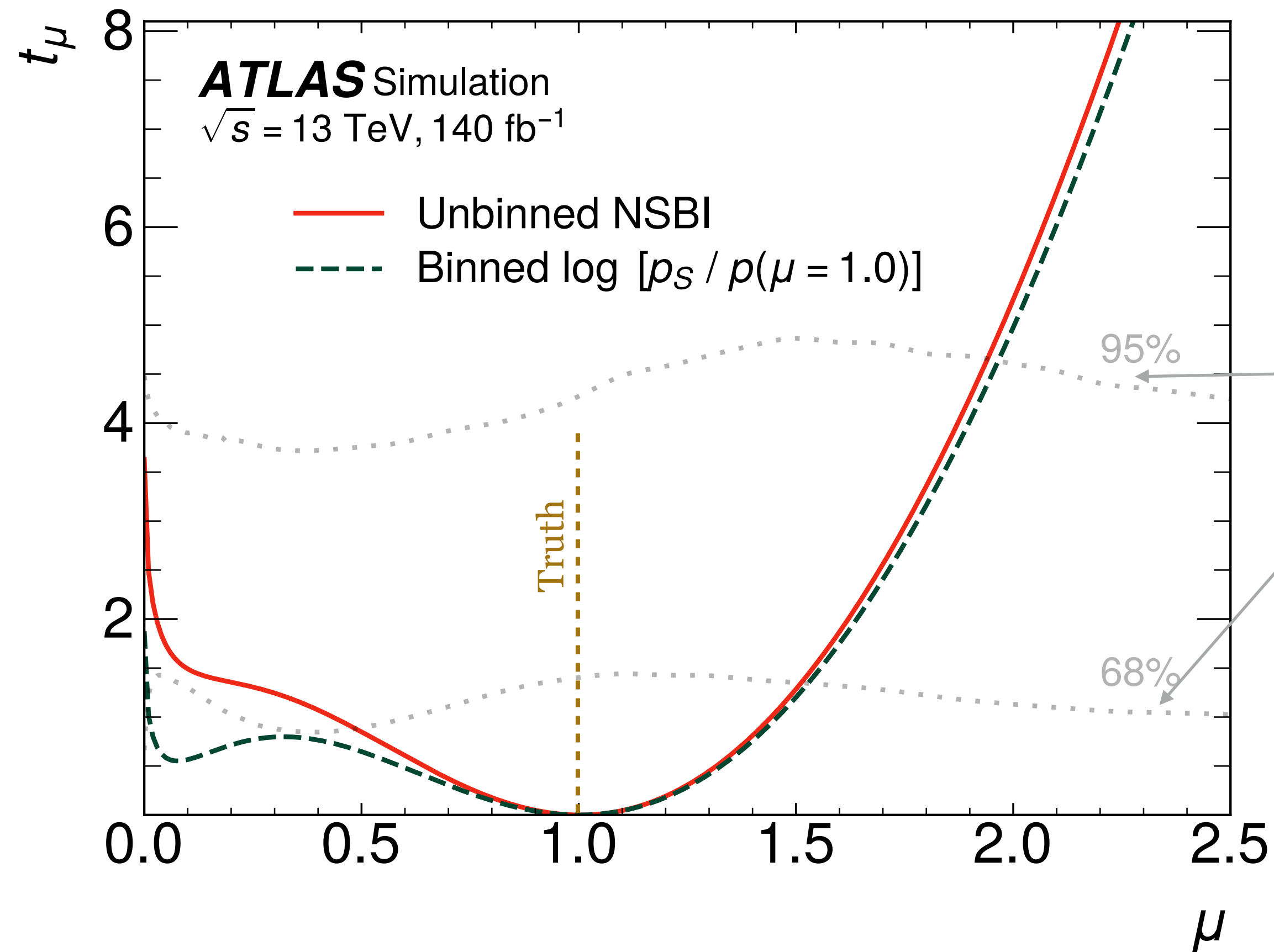
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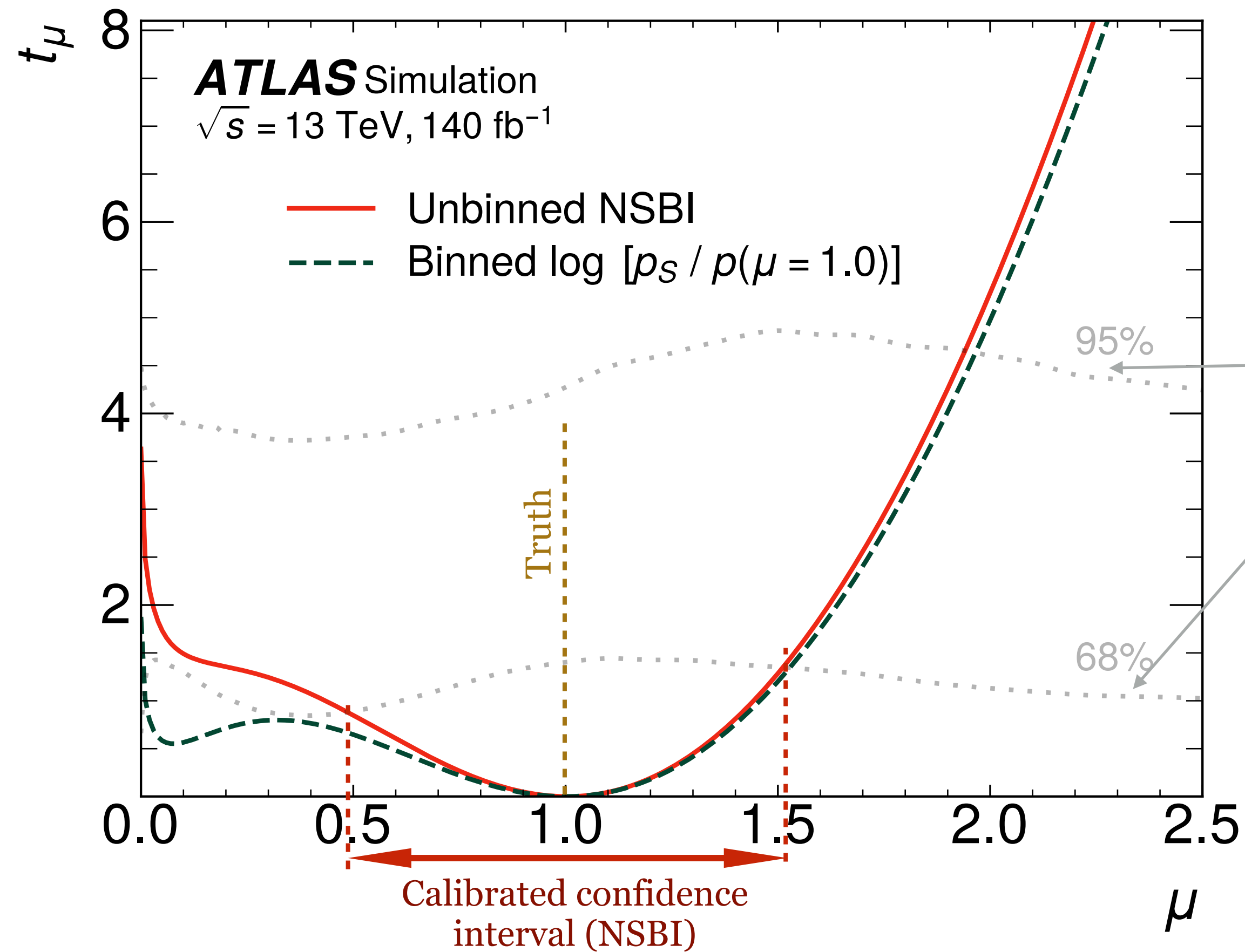
Critical values



Built with Neyman construction: Expensive process of running entire analysis on thousands of pseudo-experiments

Similar to structure seen in histogram analysis

Critical values



Built with Neyman construction: Expensive process of running entire analysis on thousands of pseudo-experiments

Similar to structure seen in histogram analysis

Vertical interpolation

$$G_j(\alpha_k) = \begin{cases} \left(\frac{v_j(\alpha_k^+)}{v_j(\alpha_k^0)} \right)^{\alpha_k} & \alpha_k > 1 \\ 1 + \sum_{n=1}^6 c_n \alpha_k^n & -1 \leq \alpha_k \leq 1 \\ \left(\frac{v_j(\alpha_k^-)}{v_j(\alpha_k^0)} \right)^{-\alpha_k} & \alpha_k < -1 \end{cases} \quad g_j(x_i, \alpha_k) = \begin{cases} (g_j(x_i, \alpha_k^+))^{\alpha_k} & \alpha_k > 1 \\ 1 + \sum_{n=1}^6 c_n \alpha_k^n & -1 \leq \alpha_k \leq 1 \\ (g_j(x_i, \alpha_k^-))^{-\alpha_k} & \alpha_k < -1 \end{cases}$$

With some continuity requirements