

The Einstein Telescope: opportunities and challenges of a next-generation detector

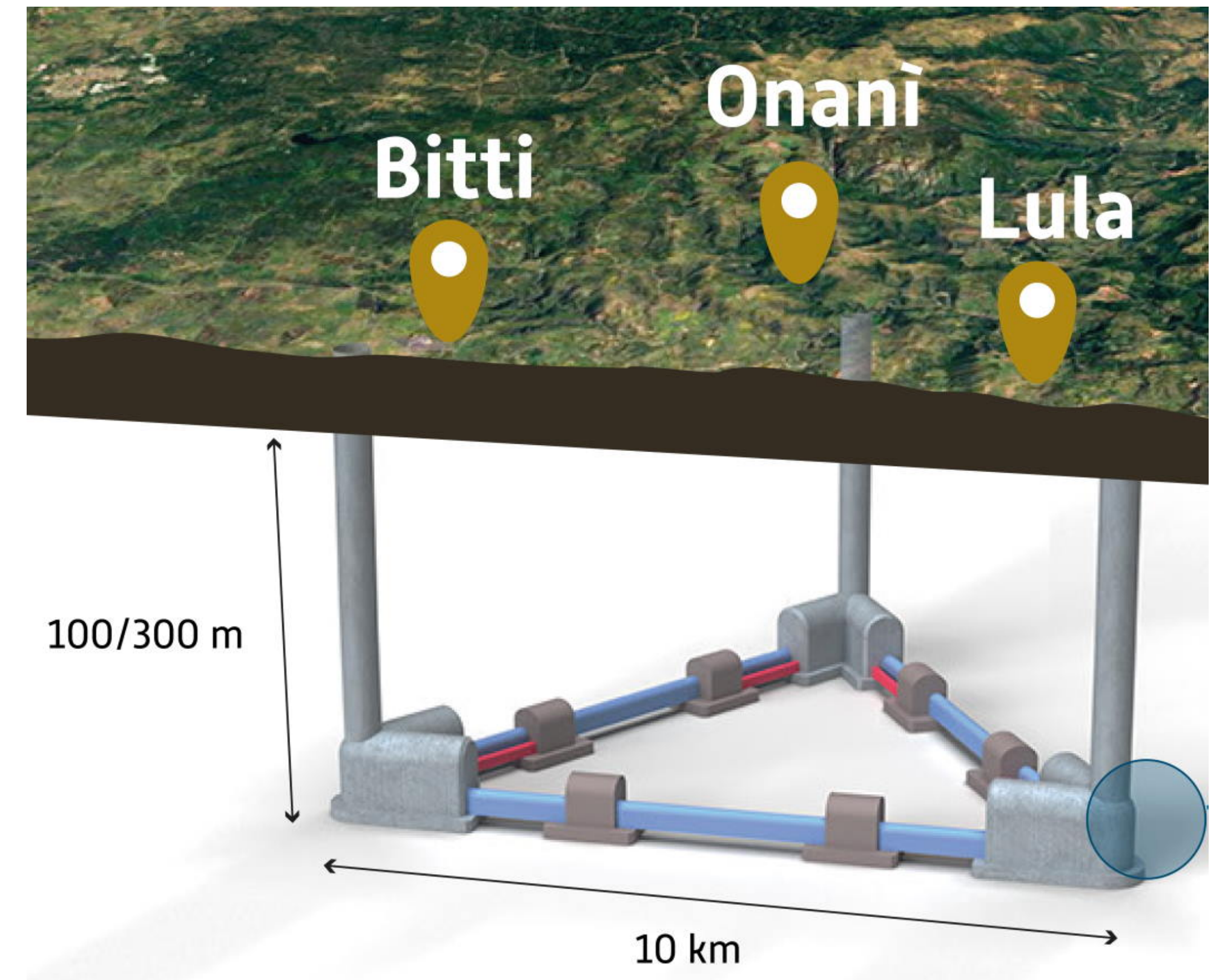
Filippo Santoliquido
GSSI, INFN

Talk on behalf of the ET collaboration

ÖPG-CMD Joint Meeting 2026
Graz, 21 September

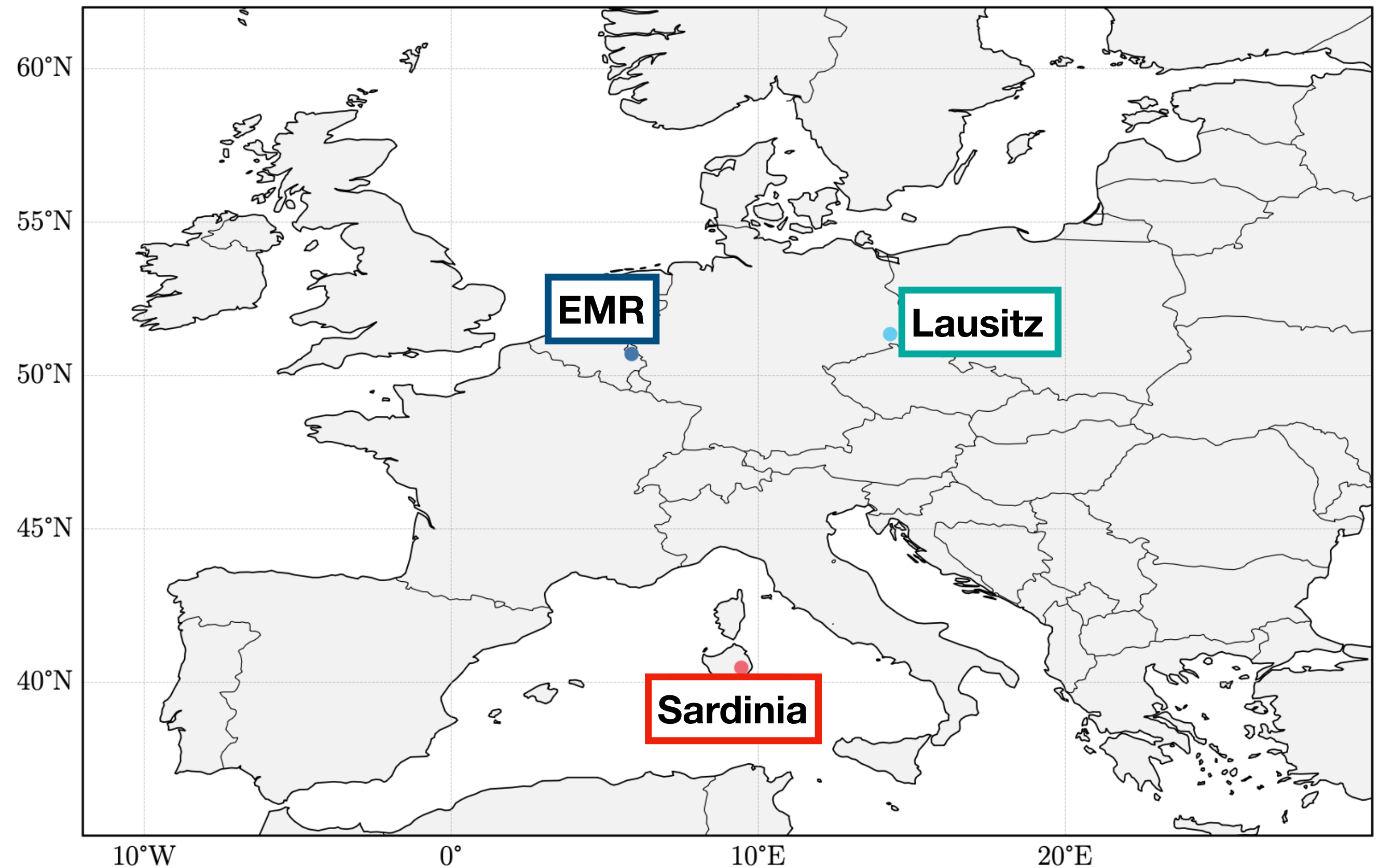
Where we stand

- Operations by 2040
- Estimated budget ~2 billion euros
- Geometry and site selection expected by 2027
- Three candidate sites and two geometries



Candidate sites

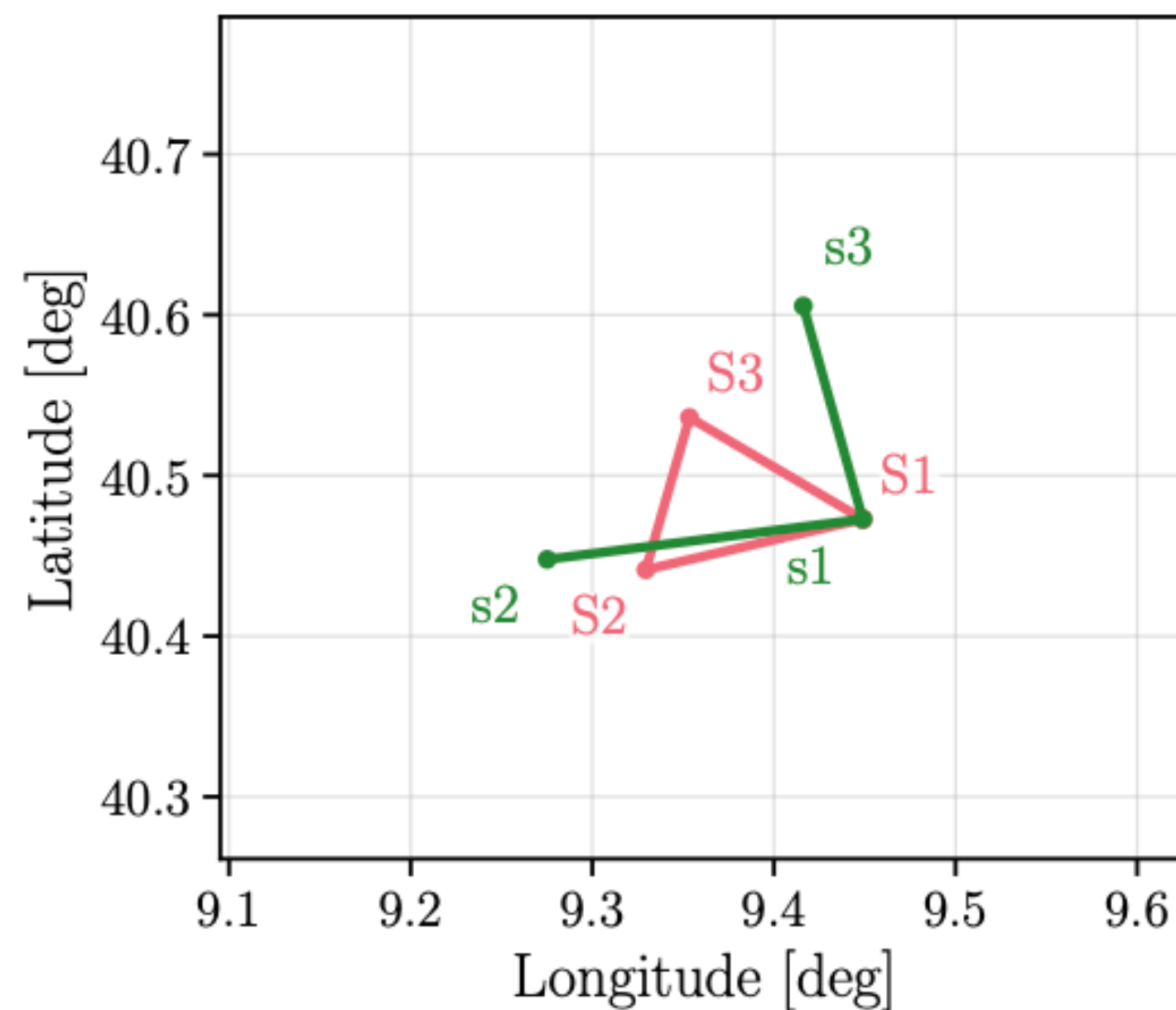
- Italy and the Netherlands-Belgium have pledged to cover nearly half of the cost if their proposed site is selected
- Sardinia and Saxony have signed a declaration of intent to strengthen scientific collaboration for the 2L design



Geometries

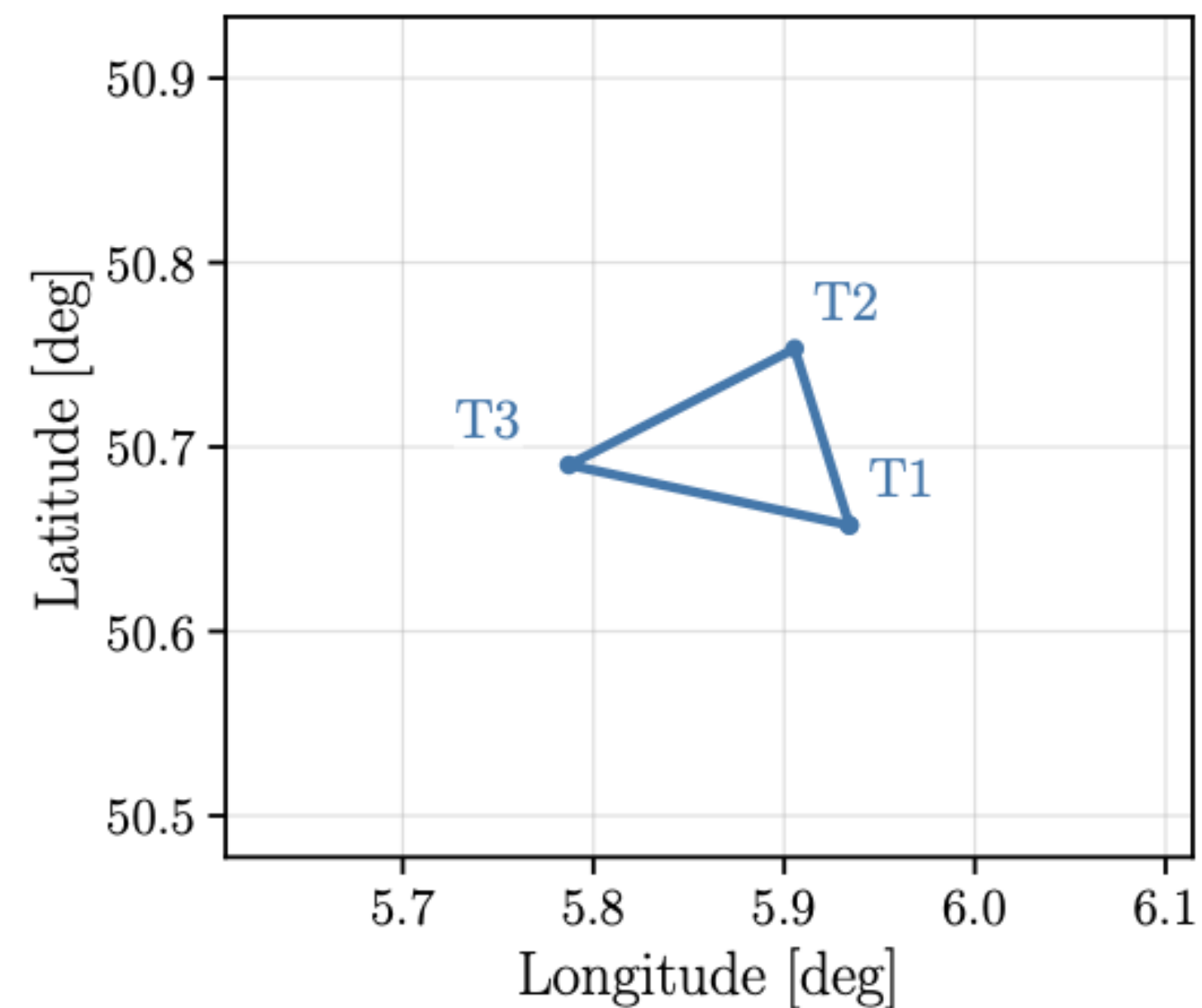
Triangular (10-km arms) vs. 2L misaligned (15-km arms)

Sardinia



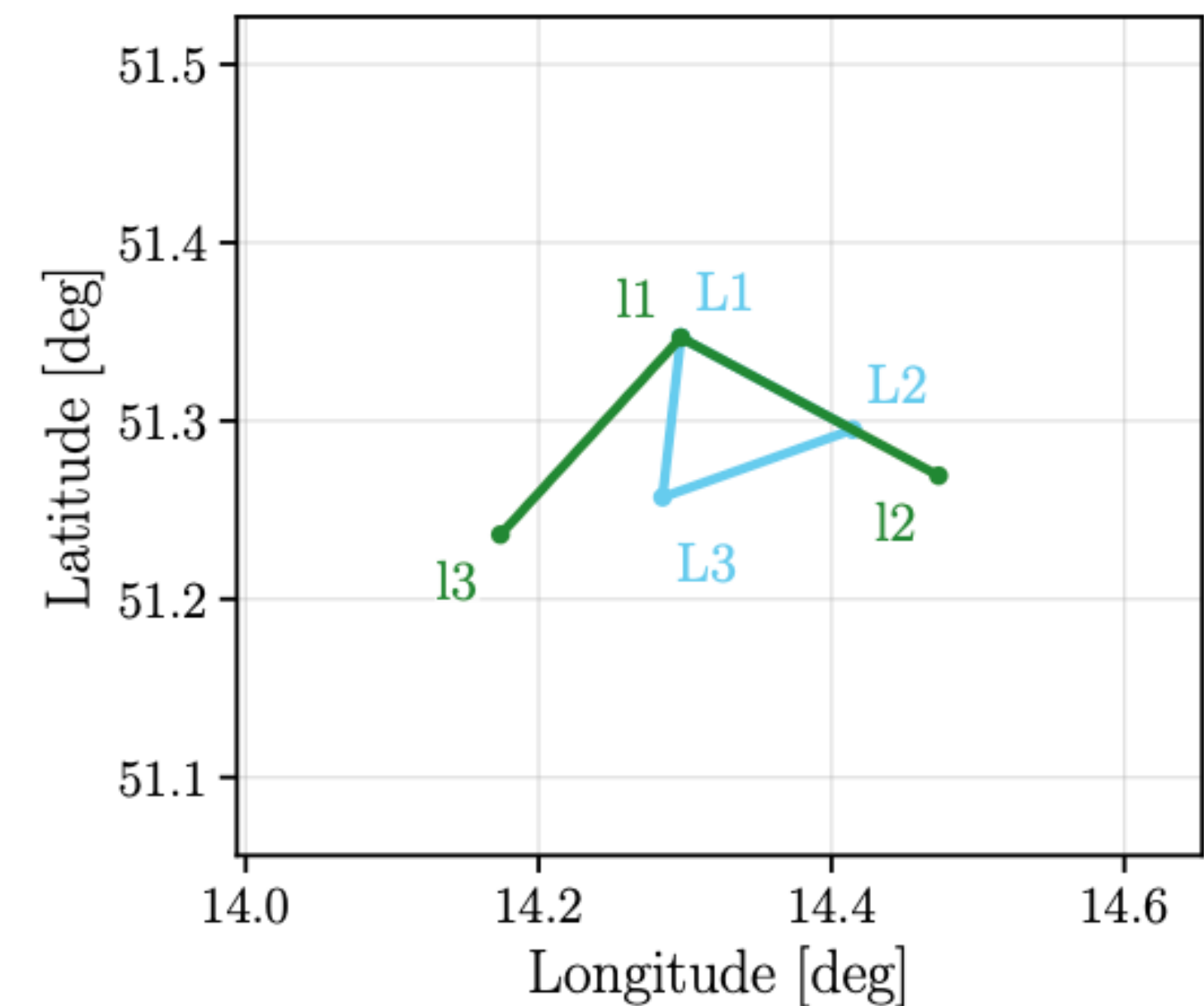
<https://www.einstein-telescope.it/>

EMR



<https://www.einsteintelelescope-emr.eu/en/>

Lausitz

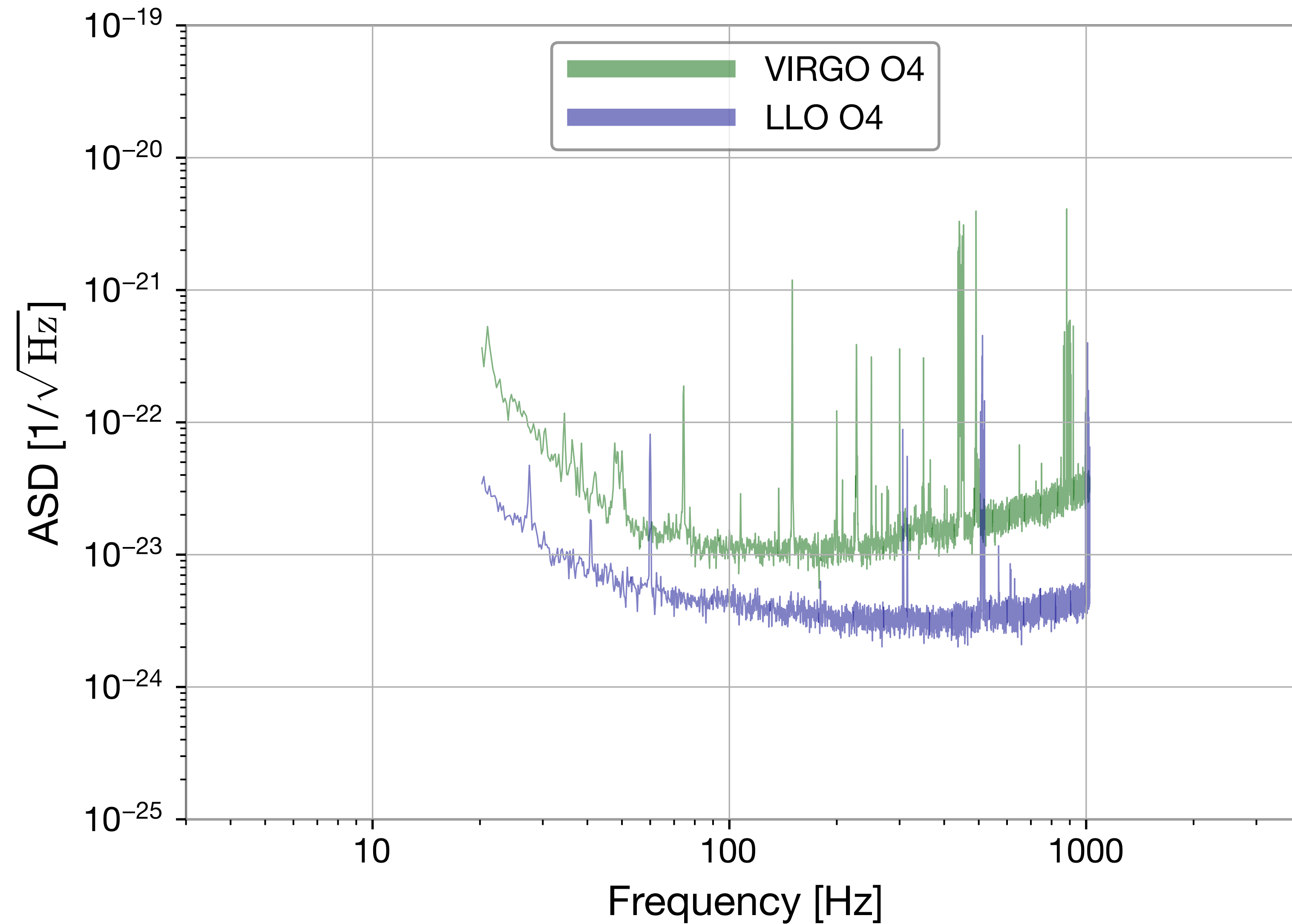


<https://einsteintelelescope-lausitz.de/>

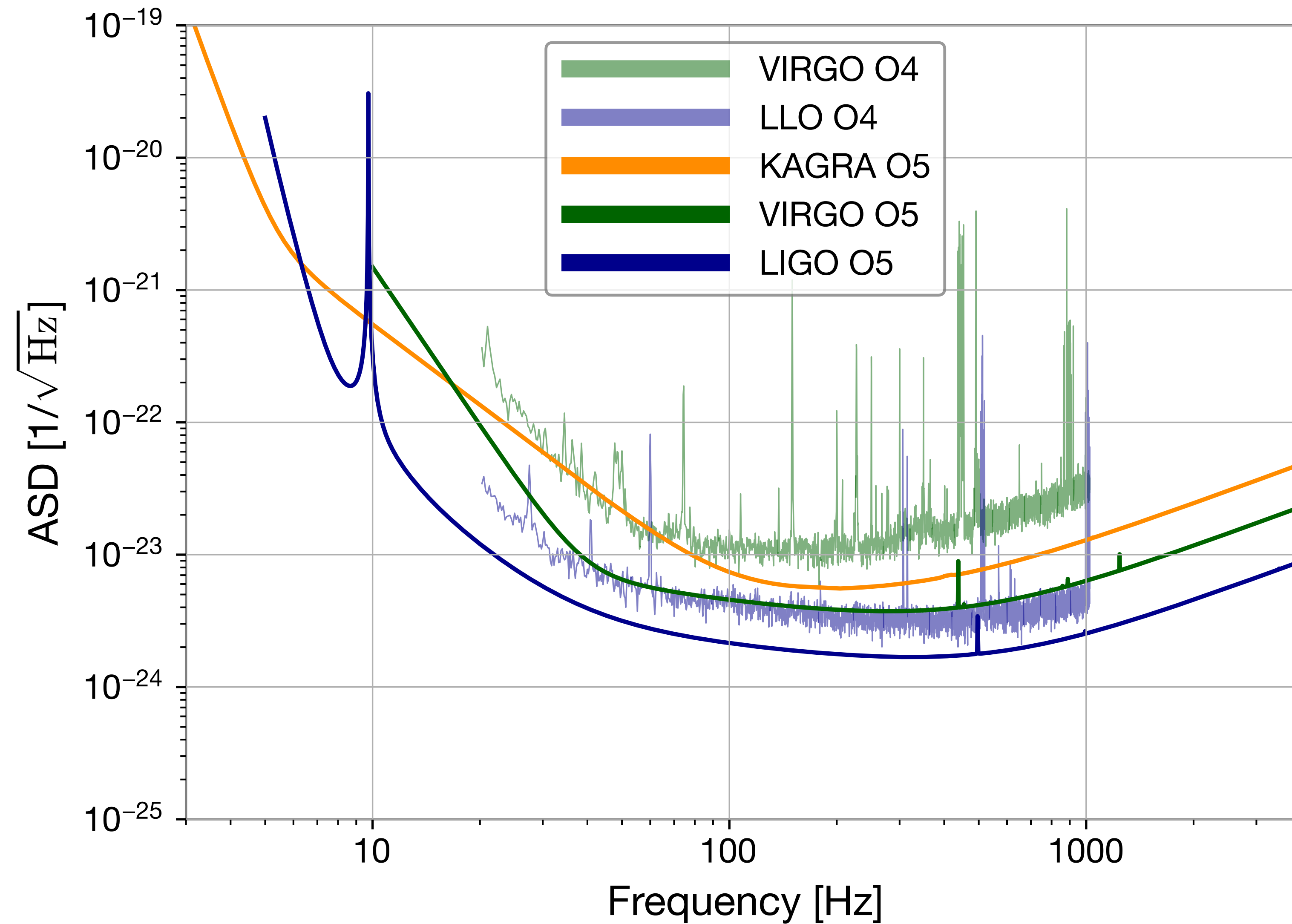
Opportunities

Science goals

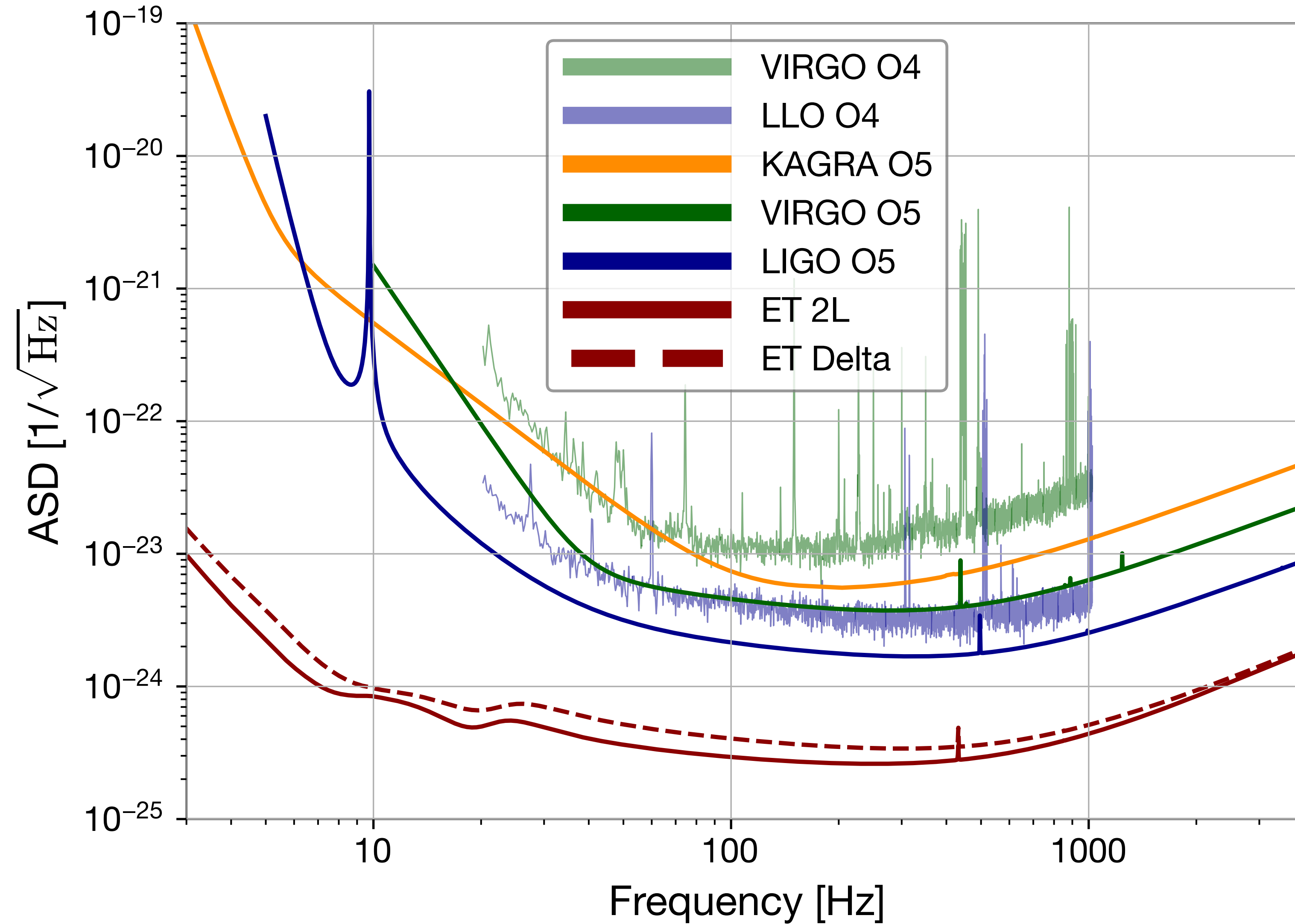
Sensitivity



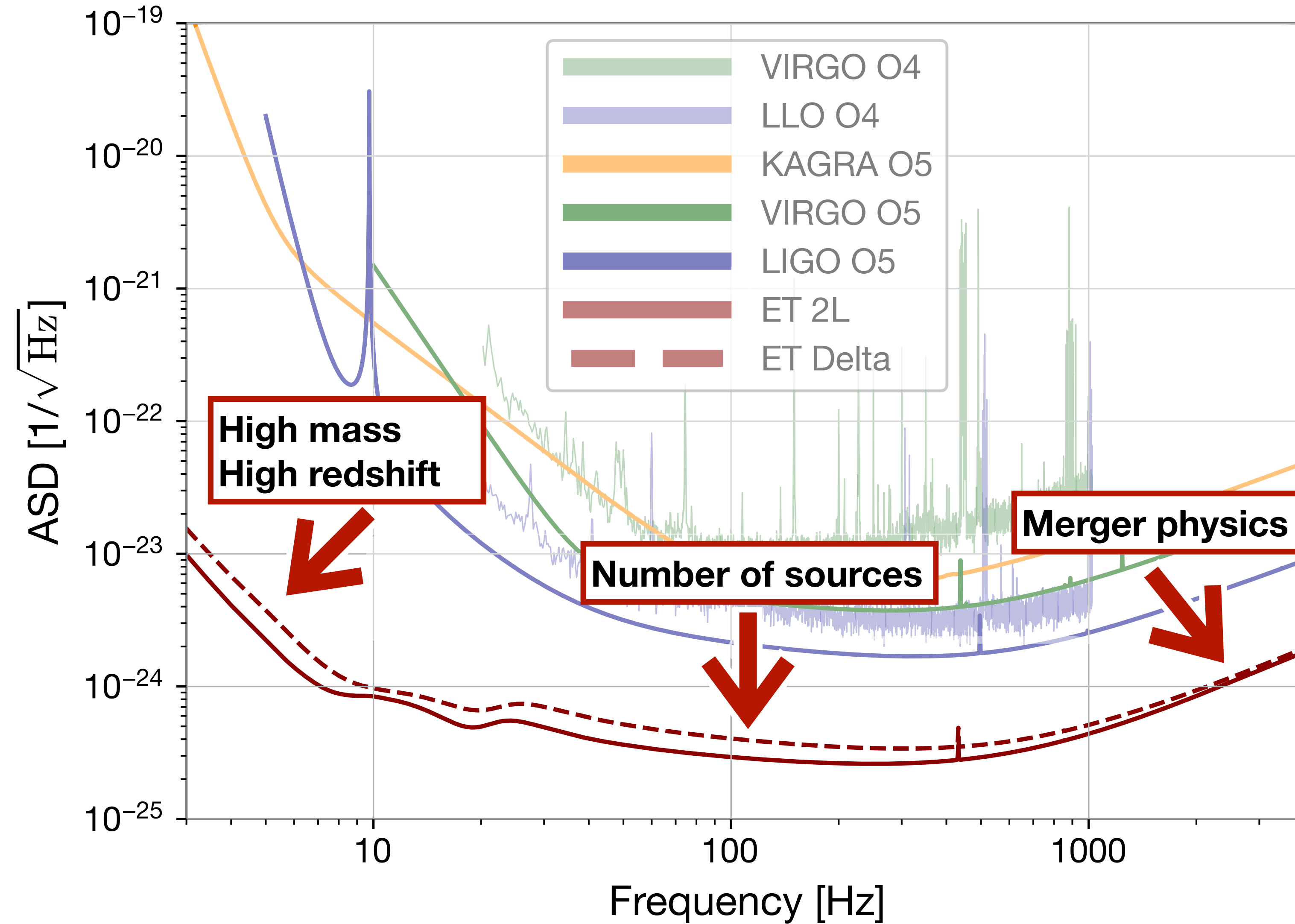
Sensitivity



Sensitivity



Sensitivity



ET science case in one book

- First collaboration paper
- Coordinated by the
Observational Science Board
- > 300 contributors, > 800 pages
- All the science objectives of the
Einstein Telescope in detail

Journal of **C**osmology and **A**stroparticle **P**hysics
An IOP and SISSA journal

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The Science of the Einstein Telescope

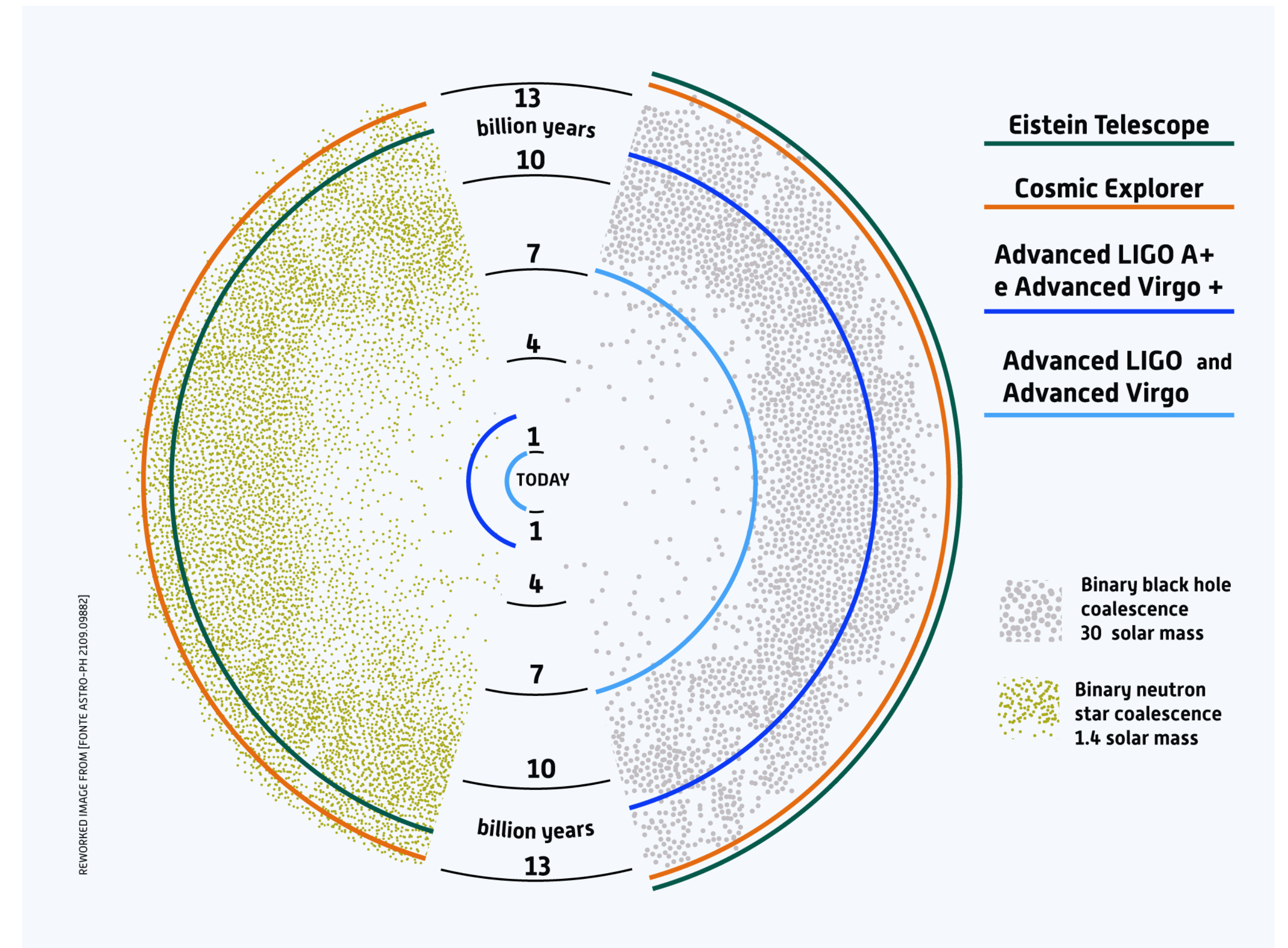
The Einstein Telescope collaboration

Full author list at the end of the paper

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ET science case

- Combination of:
 - Larger distances
➡ explore the early Universe
 - Increased number of detections
➡ advanced population studies
 - Detections with very high-SNR
➡ precise GW astronomy



ET science case

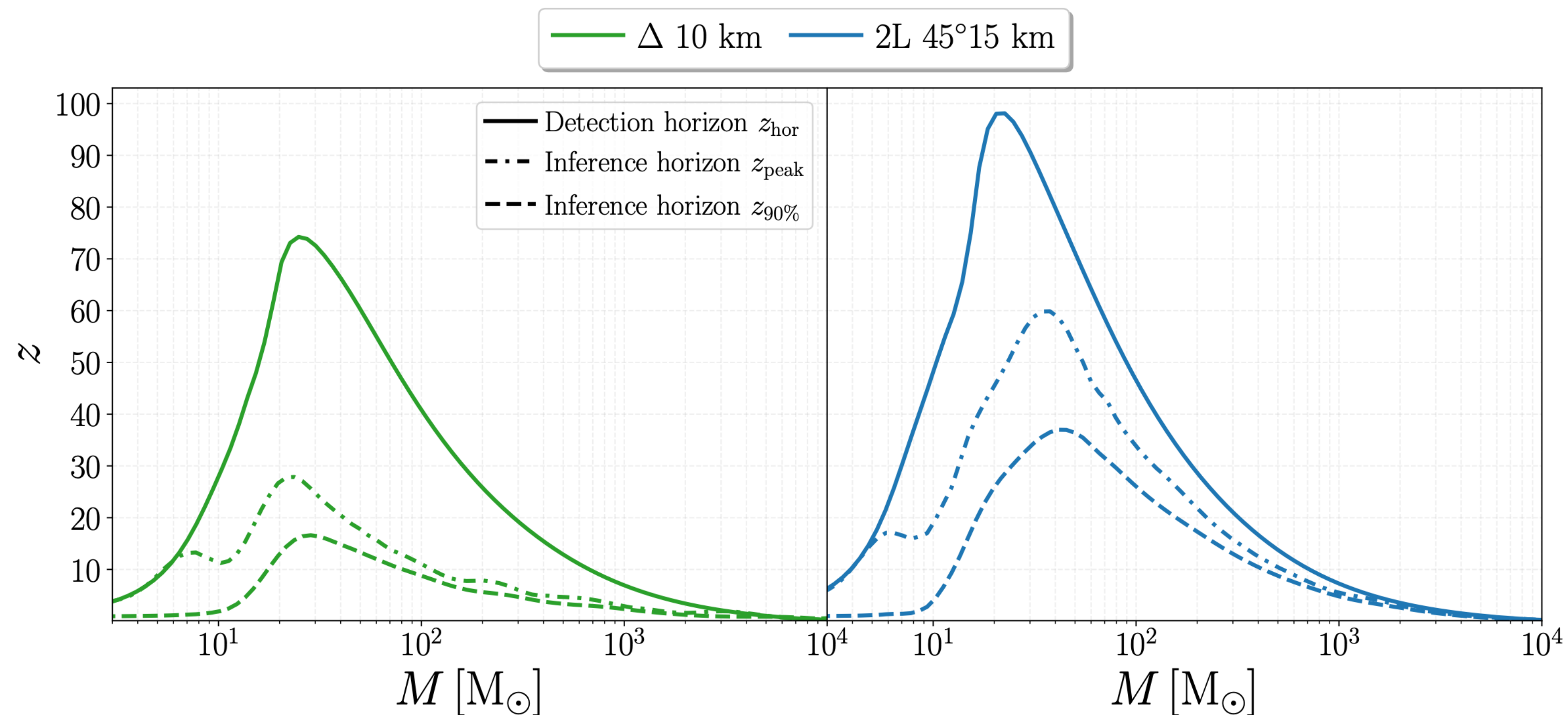
Astrophysics:

- **Black hole properties**
 - Origin (stellar vs. primordial)
 - Formation and demography
- **Neutron star properties**
 - Interior structure (QCD at ultra-high densities, exotic states of matter)
 - Demography
- **Multimessenger astronomy**
 - Joint GW/EM observations (GRBs, kilonovae)
 - Multiband GW astronomy (LISA)
 - Neutrinos
- **New GW sources:**
 - CC supernovae
 - Isolated neutron stars
 - Stochastic background of astrophysical origin

Fundamental physics and cosmology:

- **Nature of compact objects**
 - Near-horizon physics
 - Test of no-hair theorem
 - Exotic compact objects
- **Test of General Relativity**
 - post-Newtonian expansion
 - Strong-field regime
- **Dark matter**
 - Primordial BHs
 - Axion clouds, dark matter accreting on compact objects
- **Dark energy and modified gravity:**
 - Equation of state
 - Modified GW propagation
- **Stochastic background of cosmological origin**
 - Inflation, cosmic strings

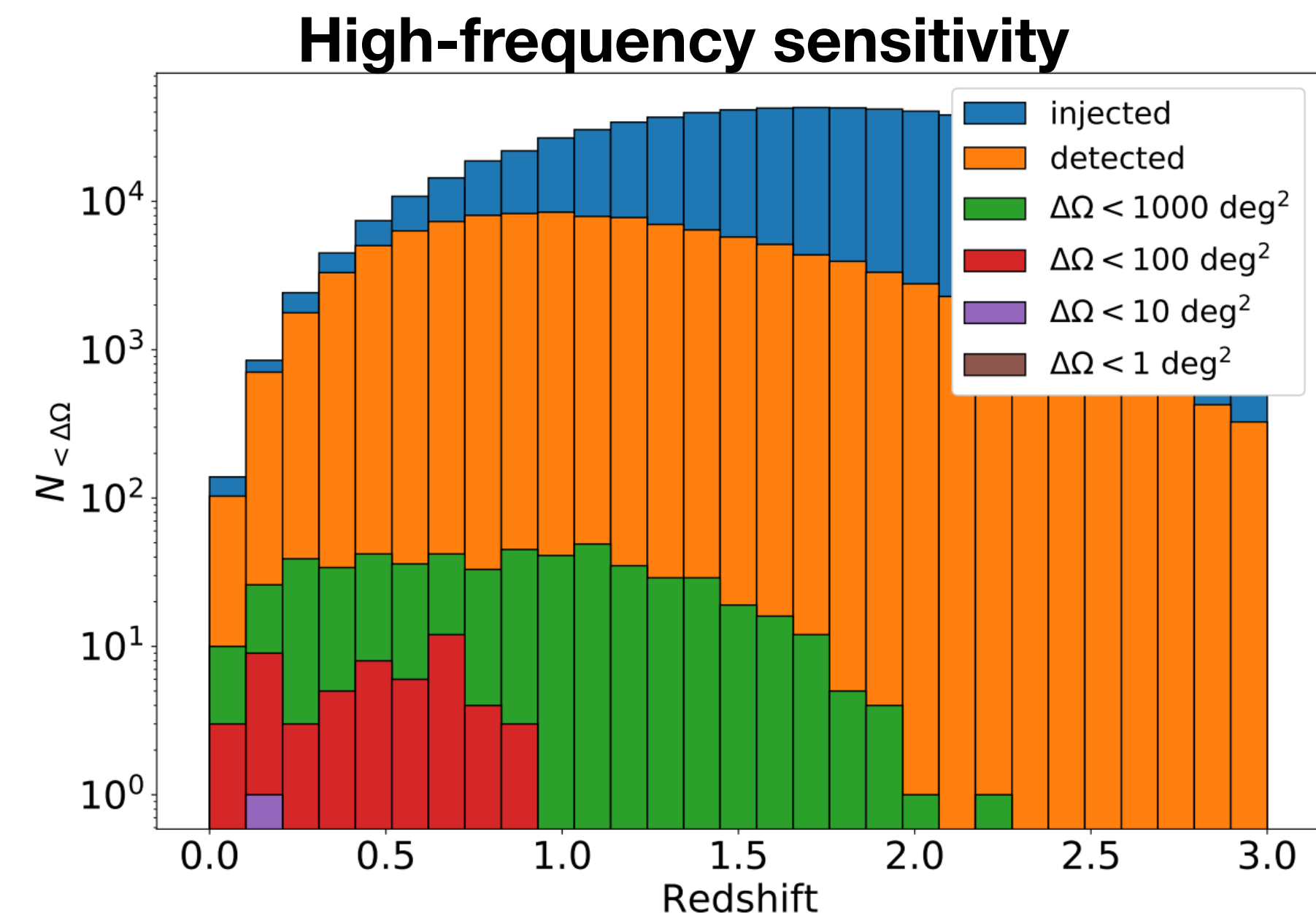
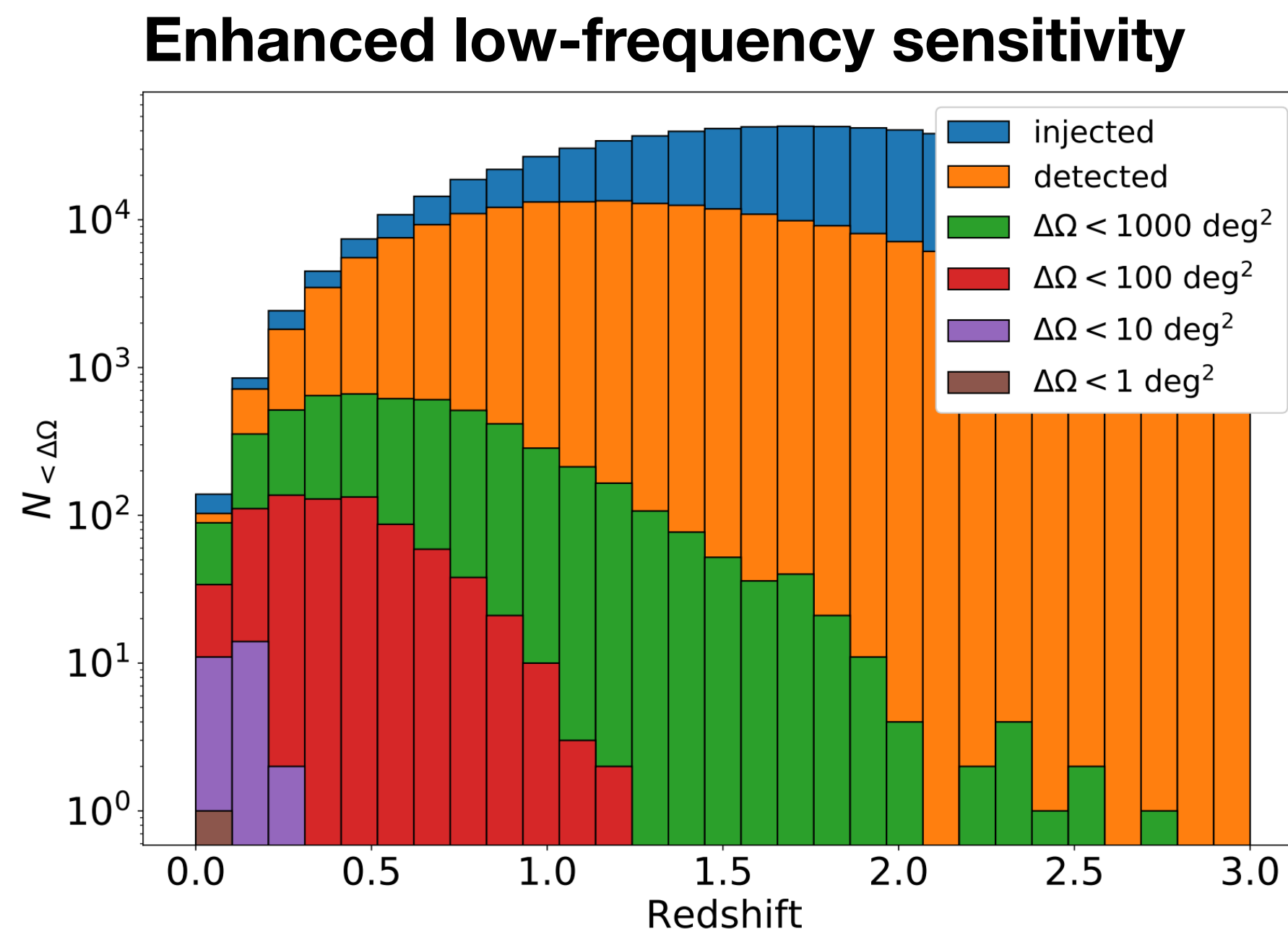
Horizon distance



- The first generation of black holes in the early Universe
- Intermediate-mass black holes (IMBHs) with masses between 100 - 10^4 solar masses → progenitors of supermassive black holes

Sky localization

- > 100 sources per year with sky localization error $< 100 \text{ deg}^2$
- Sky localization and pre-merger alerts  sensitivity at low frequency is paramount

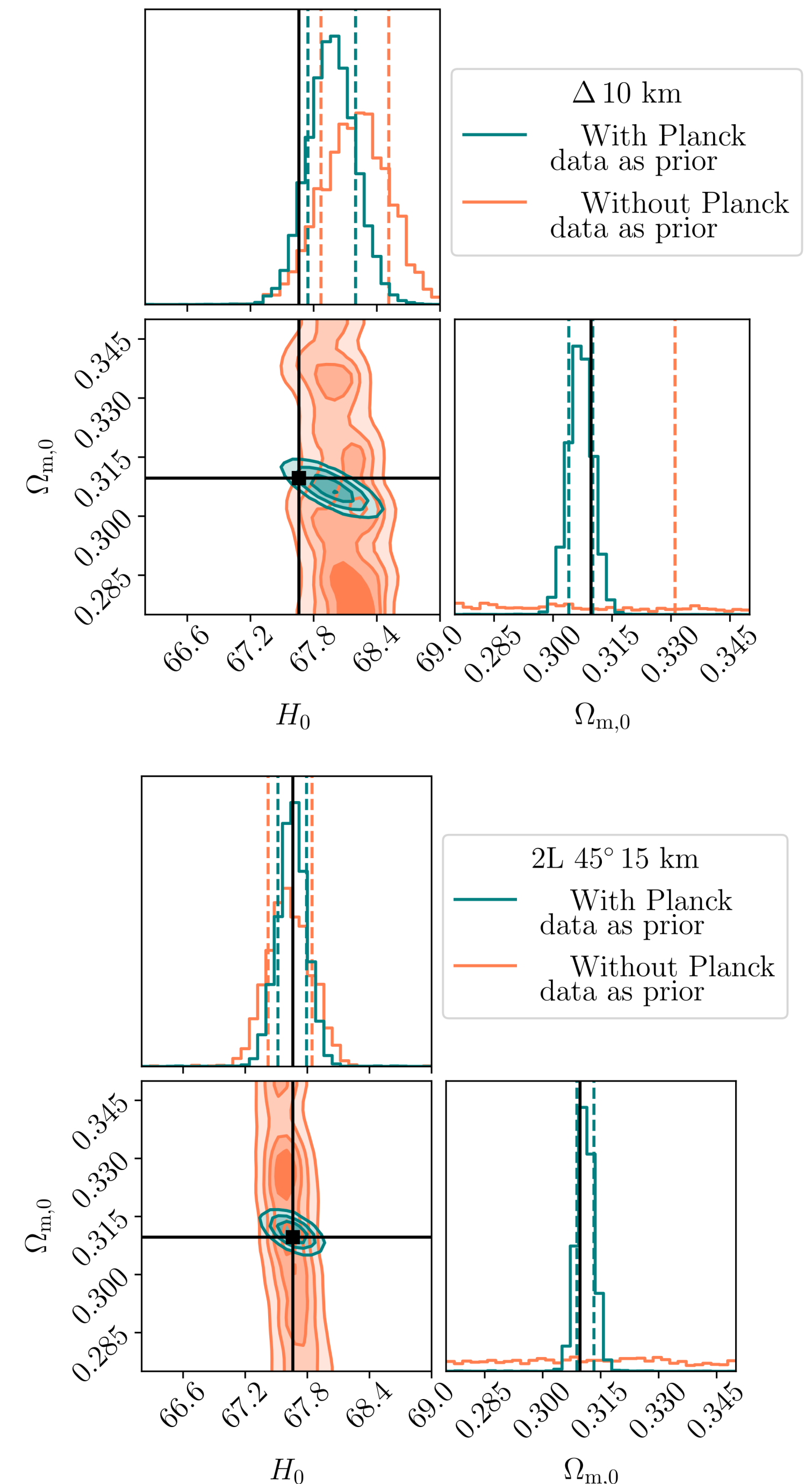


Cosmology

- Independent measurement of cosmological parameters

$$D_L(z) = (1 + z) \frac{c}{H_0} \int_0^z \frac{dz'}{\sqrt{\Omega_{m,0}(1 + z')^3 + (1 - \Omega_{m,0})}}$$

- In the example: H_0 , $\Omega_{m,0}$ from GW + Kilonova events in one year with Vera Rubin Observatory

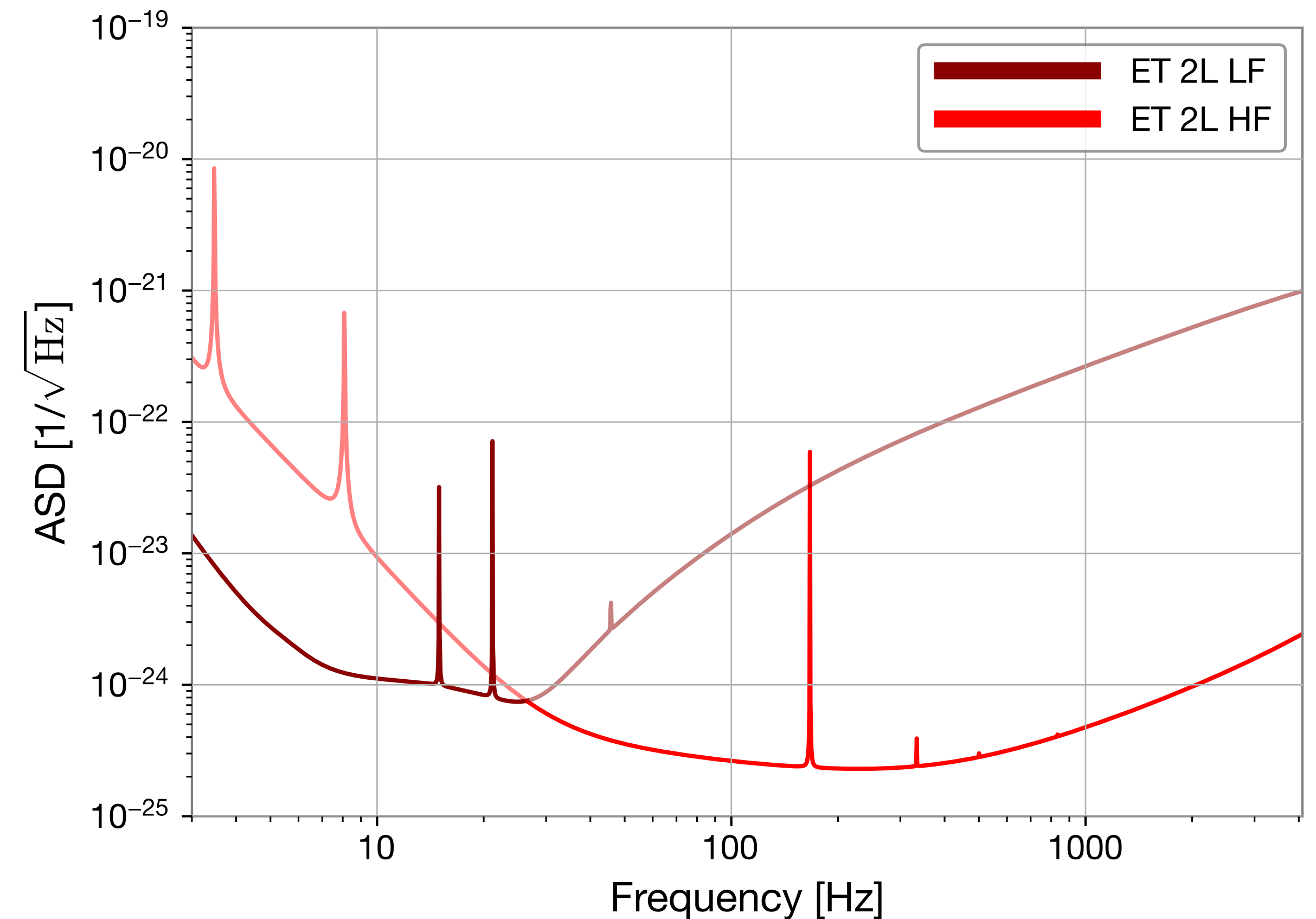


Challenges

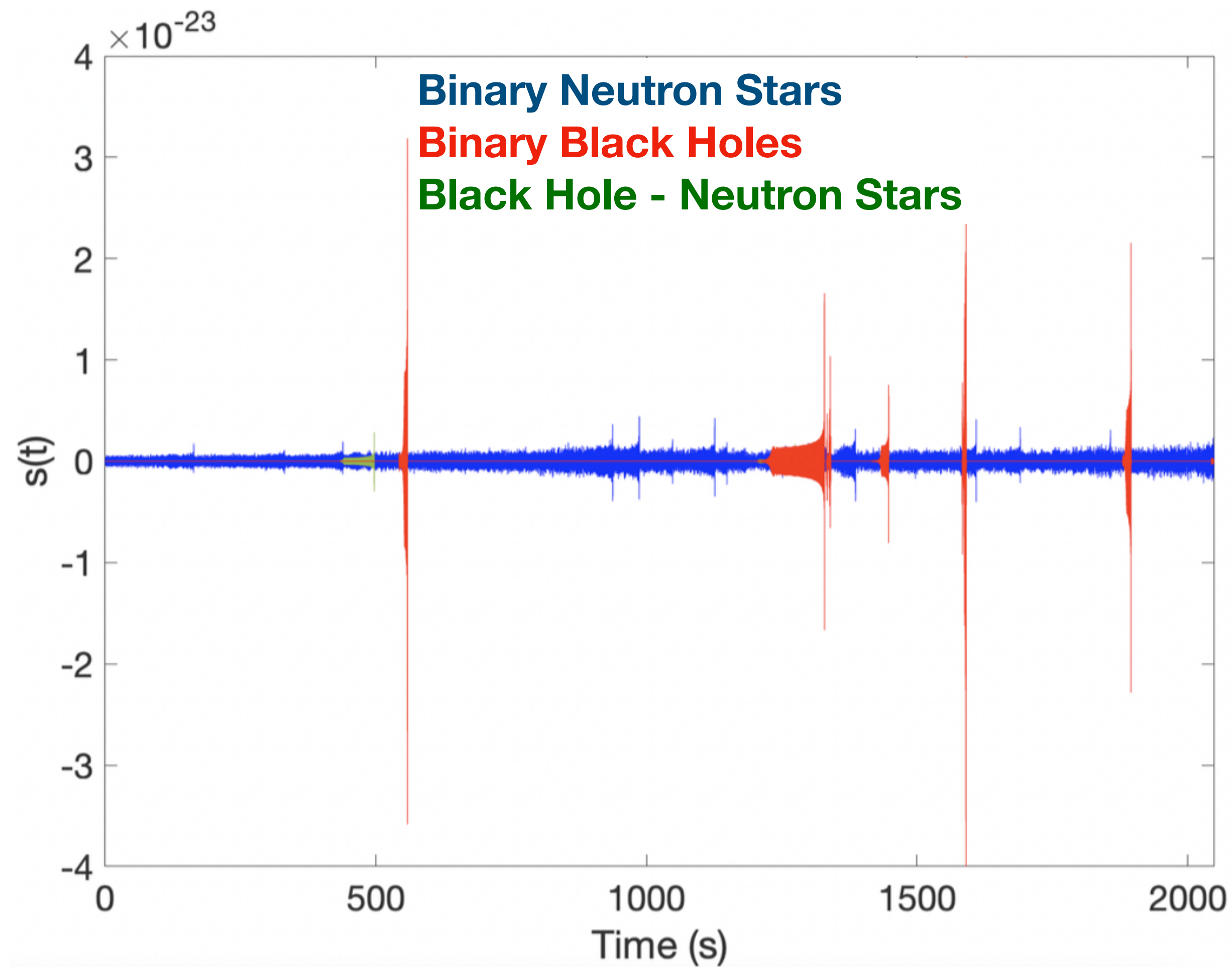
Technological development

Low frequency sensitivity

- The great potential of ET lies in its ability to access the 3-30 Hz regime
- Many sources of noise: seismic, thermal, Newtonian, electromagnetic
- Underground + cryogenics (10-20 K)
- **Xylophone configuration:** one instrument tuned to low frequency (LF) and one for high frequency (HF)



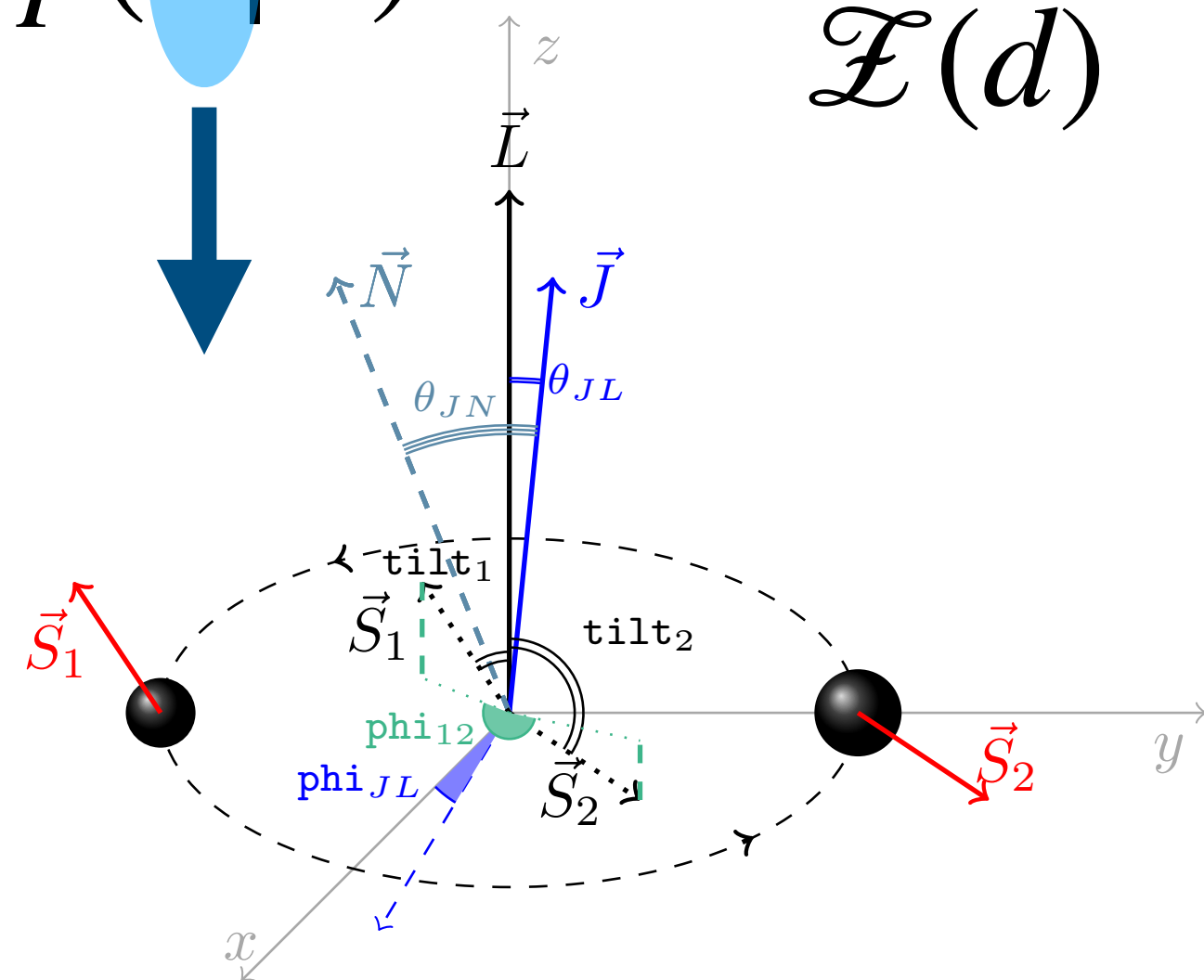
Data analysis challenge



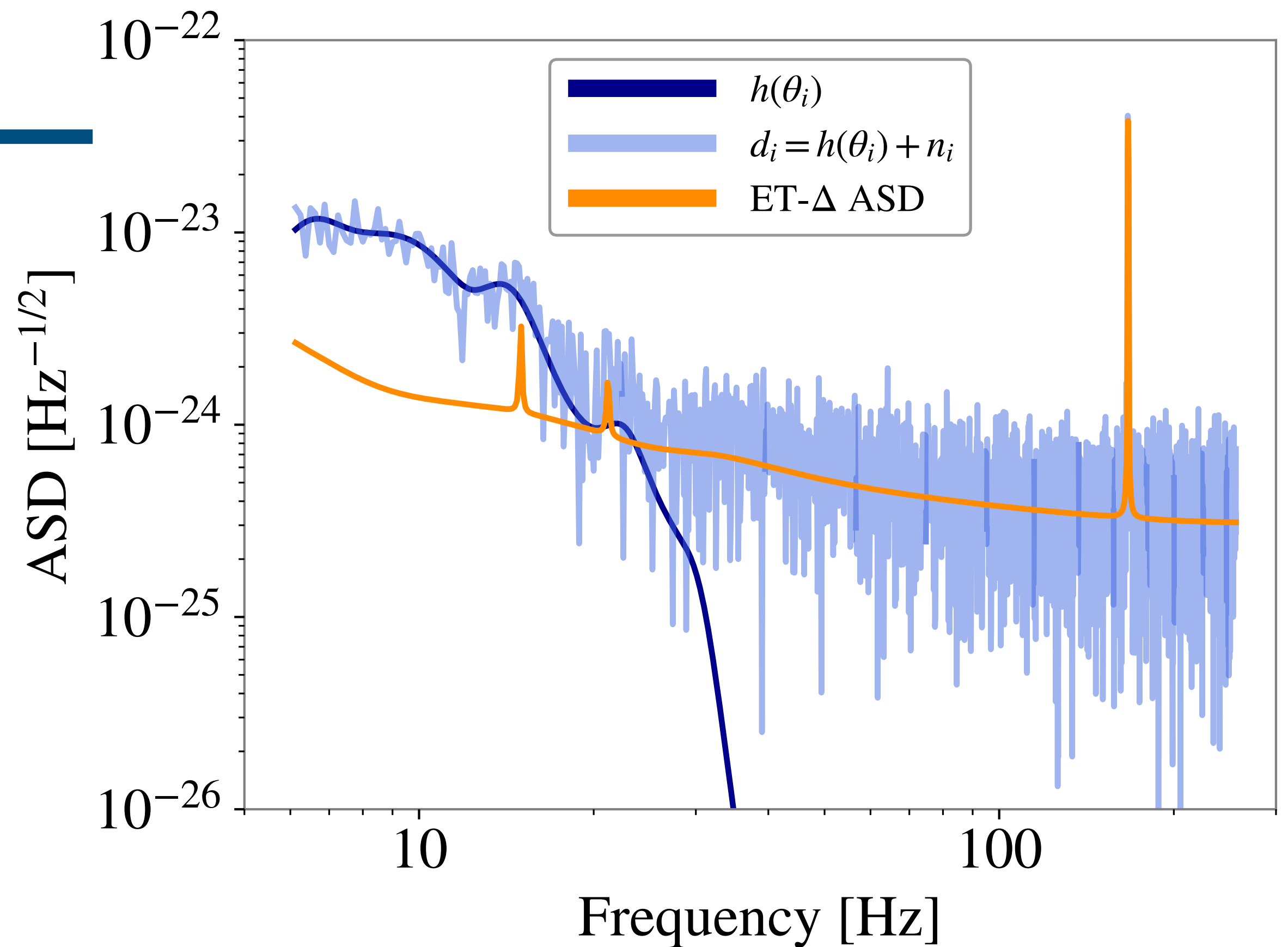
- 10^5 signals per year
- Long signals
- Non-stationary noise
- Overlapping signals

Bayesian inference

$$p(\theta | d) = \frac{\mathcal{L}(d | \theta) \pi(\theta)}{\mathcal{Z}(d)}$$

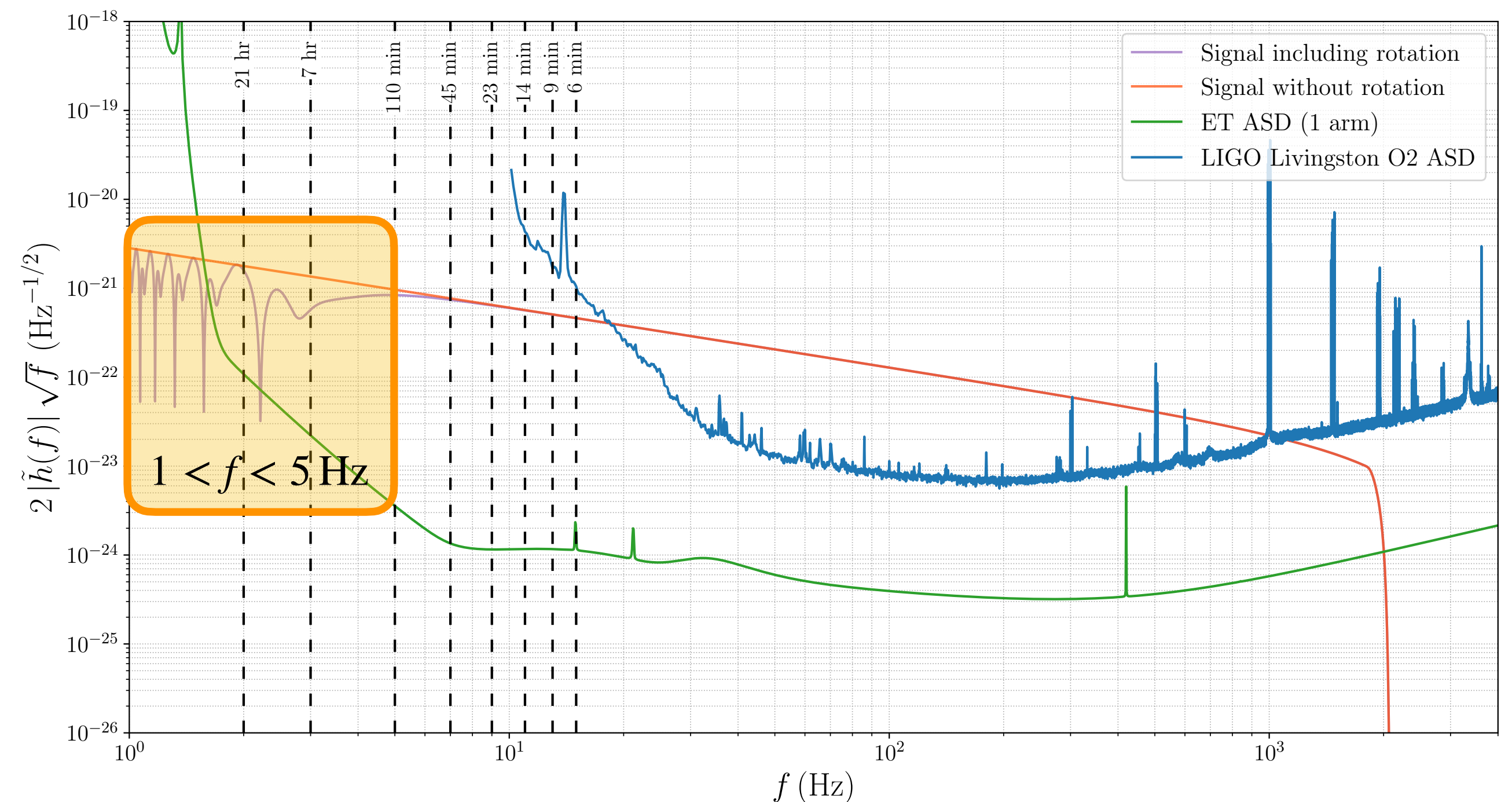


$$\theta = \{\mathcal{M}_d, q, d_L, \text{ra}, \text{dec}, \theta_{JN}, \psi, \phi, t_{\text{geocent}}, \chi_1, \chi_2\}$$

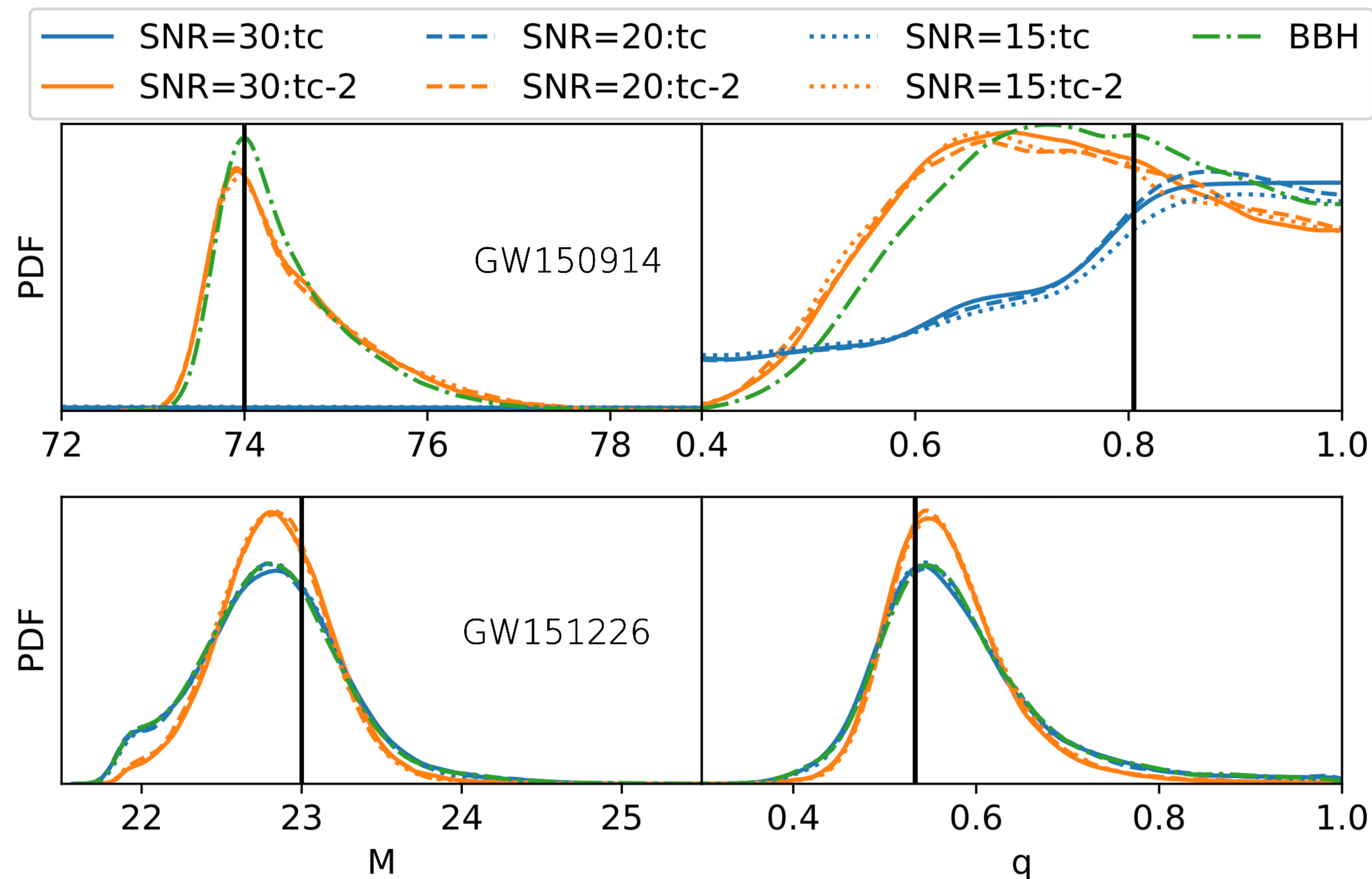


Long signals

- GW170817-like signal lasts ~ 7 hours starting from $f_{\min} = 3$ Hz. Challenges:
- Linear increase with signal duration in the evaluation cost of the likelihood
- Measured signal is modulated in amplitude and phase due to Earth's rotation
- Detector noise could change over time and signals could overlap



Overlapping signals

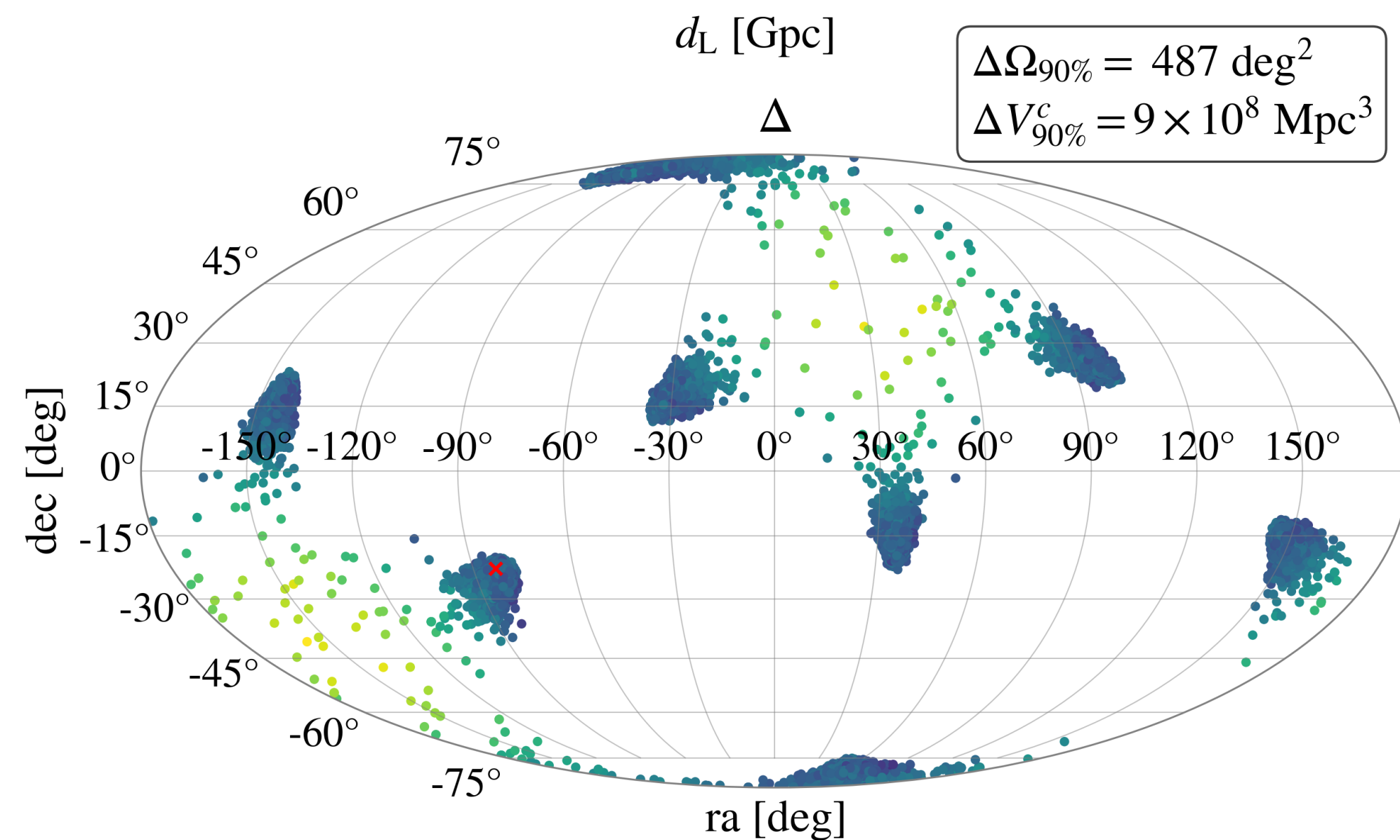


- Traditional **single-signal models** may introduce biases, particularly for closely timed or comparable signal amplitudes.
- New methods to separate and analyze overlapping signals simultaneously are crucial.

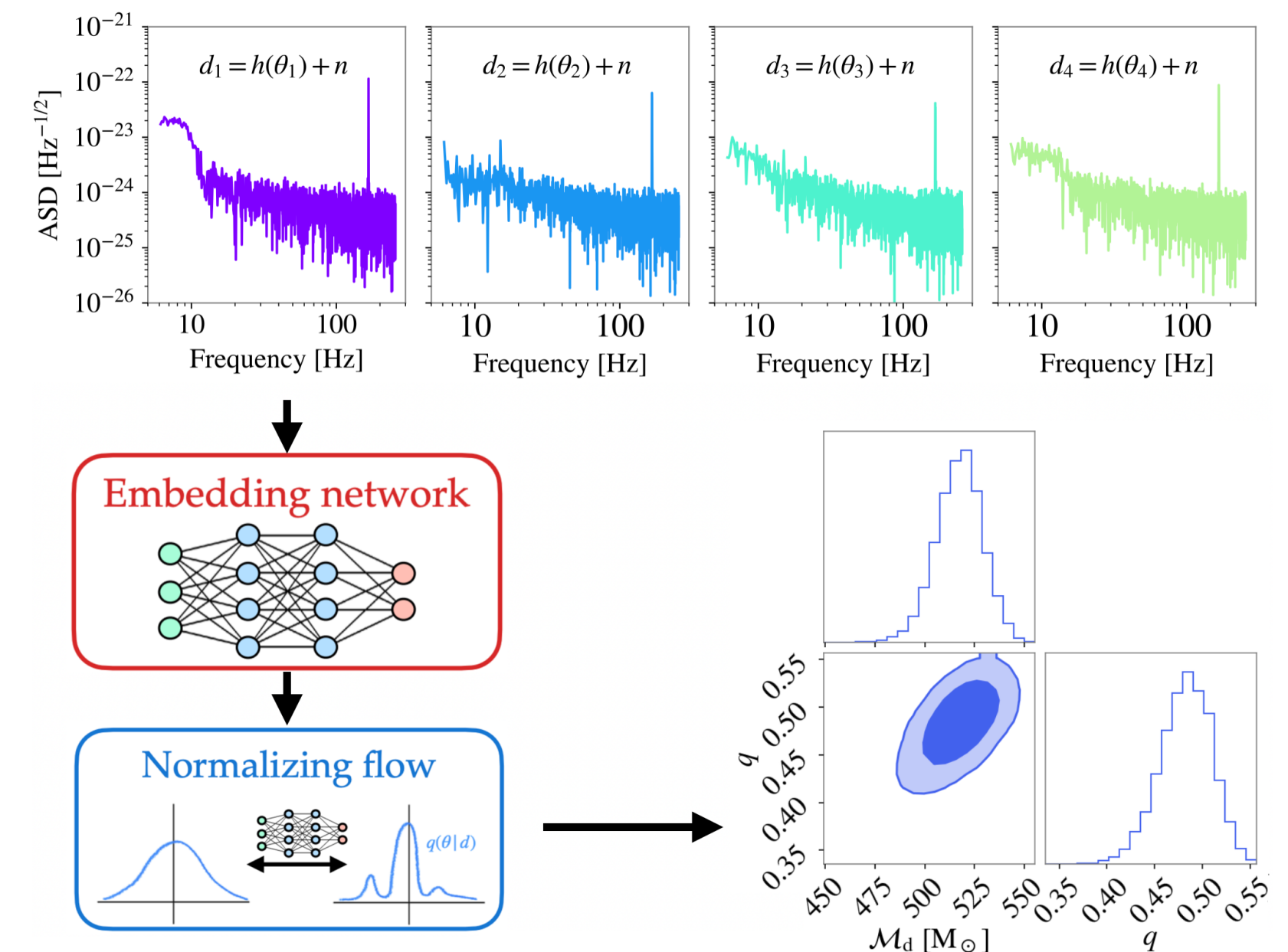
In the example: BBH signal overlapped with a BNS signal at **different SNRs** (line styles) and at **different coalescing times** (colors). Green-dotted lines are posterior of single BBH

Scaling inference

Large and disconnected
parameter space requires
long computational times
with standard methods



Necessary to develop new methods based on deep learning (simulation based inference)



Mock data challenges

The Data Analysis division stimulates, guides and validates the methods that will handle ET data

Why

- Benchmark algorithms, evaluate computational cost
- Predict ET performance in realistic scenarios

How

- [gwmock](#): Python package for generating simulated datasets
- Complexity of the data increases with each challenge

“The Einstein Telescope will open new frontiers in GW astronomy, but it demands a paradigm shift in data analysis methodologies”

The Science of the Einstein Telescope, [Abac et al. 2026](#)

Conclusions

- Unprecedented scientific reach ➡ Europe's science flagship of the next decades
- Precision GW astronomy, multimessenger astronomy, high-redshift sources ➡ potential for new discoveries
- Einstein Telescope is one of the greatest challenges ever, both in infrastructure and in data analysis