



UNIVERSITY OF
BIRMINGHAM



INSTITUTE OF GRAVITATIONAL
WAVE ASTRONOMY



LIGO
Scientific
Collaboration
COLLABORATION

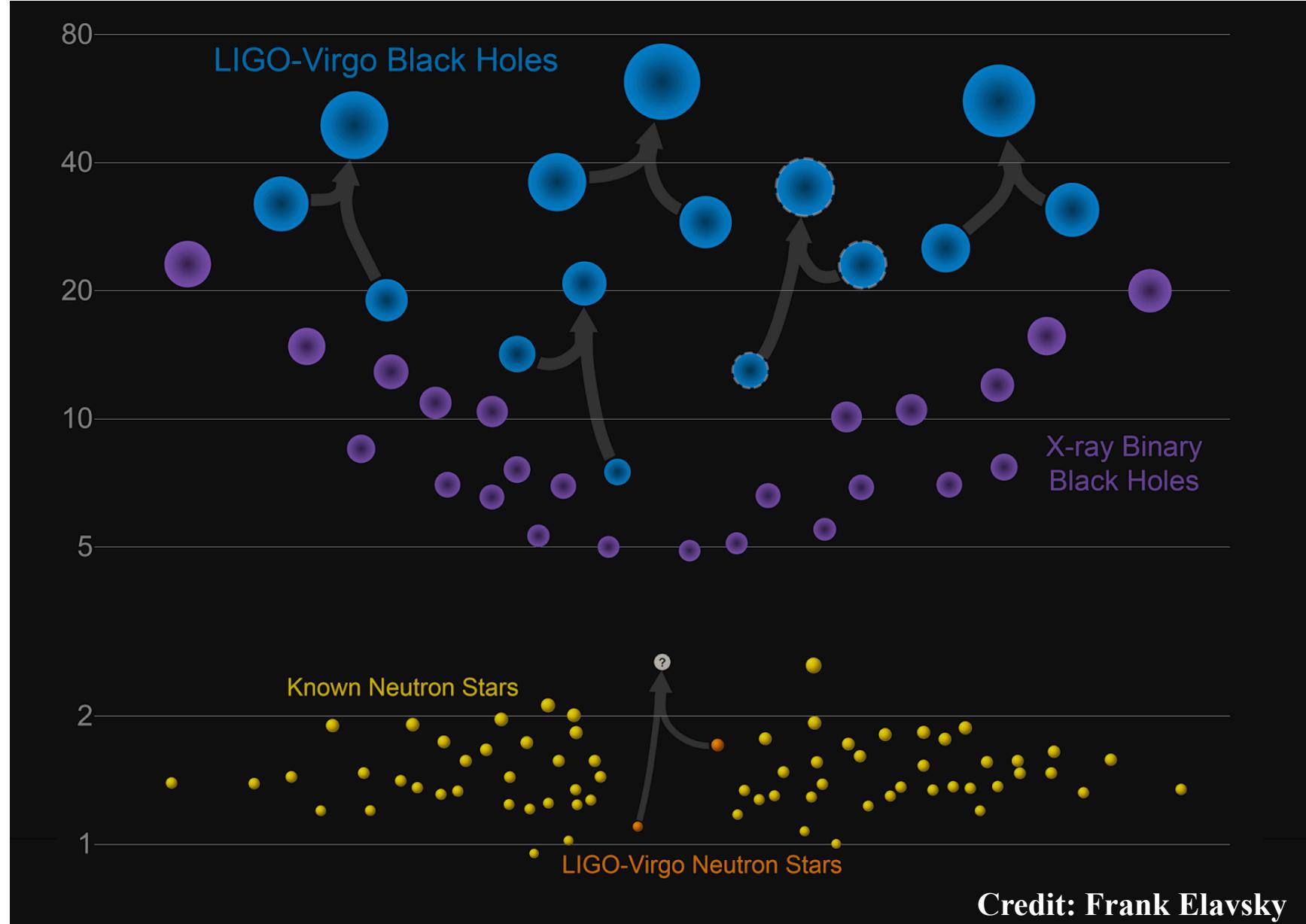
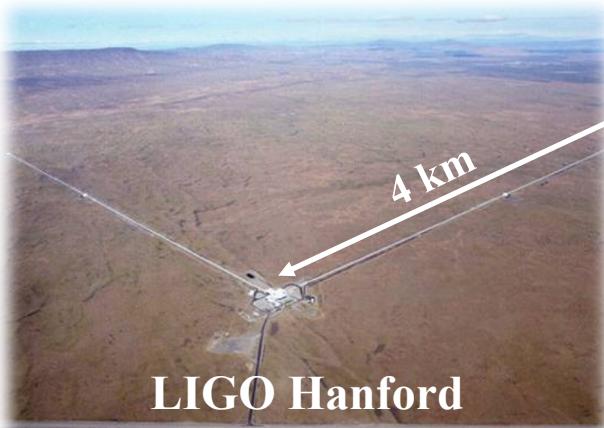
Quantum Techniques in Laser Interferometric Gravitational-wave Detectors

Haixing Miao

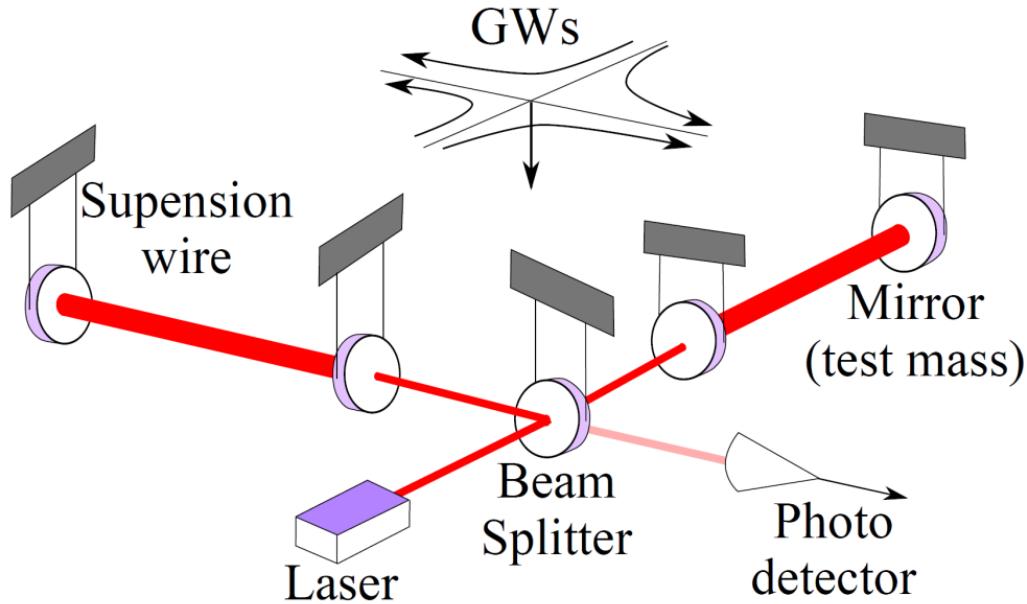
In collaboration with:

Rana Adhikari, Dominic Branford, Yanbei Chen, Animesh Datta, Stefan Danilishin, Matthew Evans, Andreas Freise, Farid Khalili, Yiqiu Ma, Denis Martynov, Nicholas Smith, Belinda Pang, Chunrong Zhao

Gravitational-wave discoveries



Why Quantum?



Strain sensitivity of km size detector:

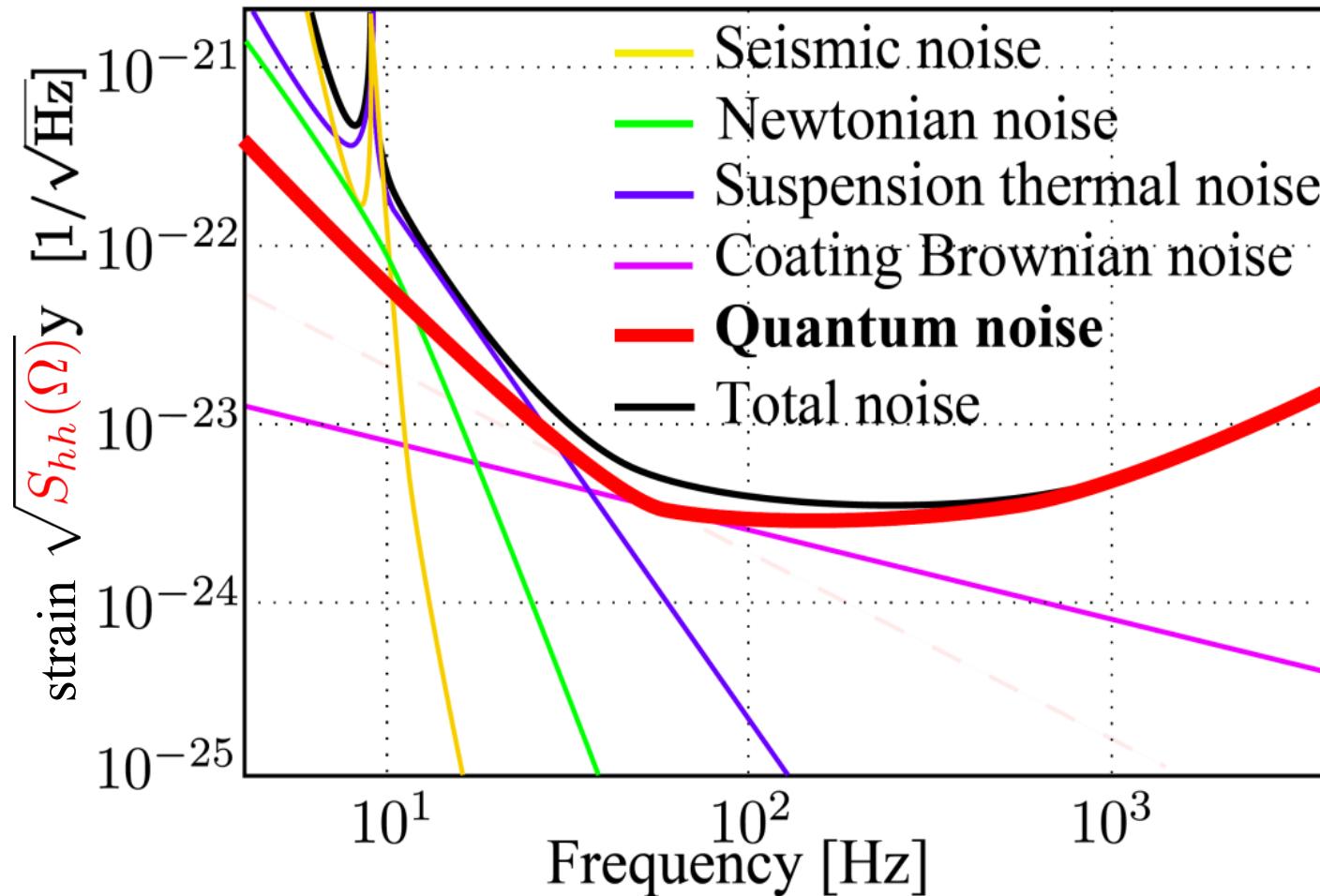
$$h = \frac{\Delta L}{L} \sim 10^{-22} \quad \Rightarrow \quad \Delta L \sim 10^{-19} \text{ m}$$

de Broglie wavelength of kg size test mass: $\lambda_d \sim \sqrt{\hbar/(2\pi m f)}|_{100\text{Hz}} \sim 10^{-19} \text{ m}$

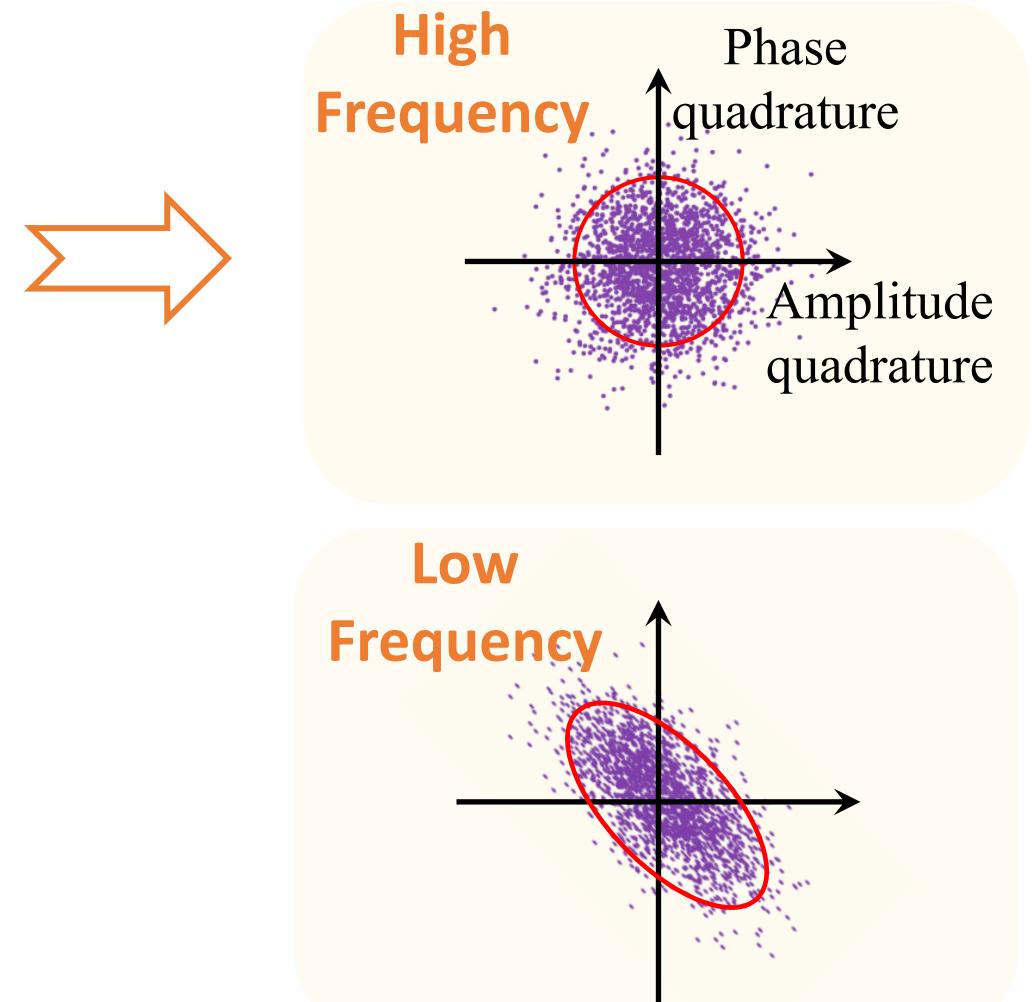
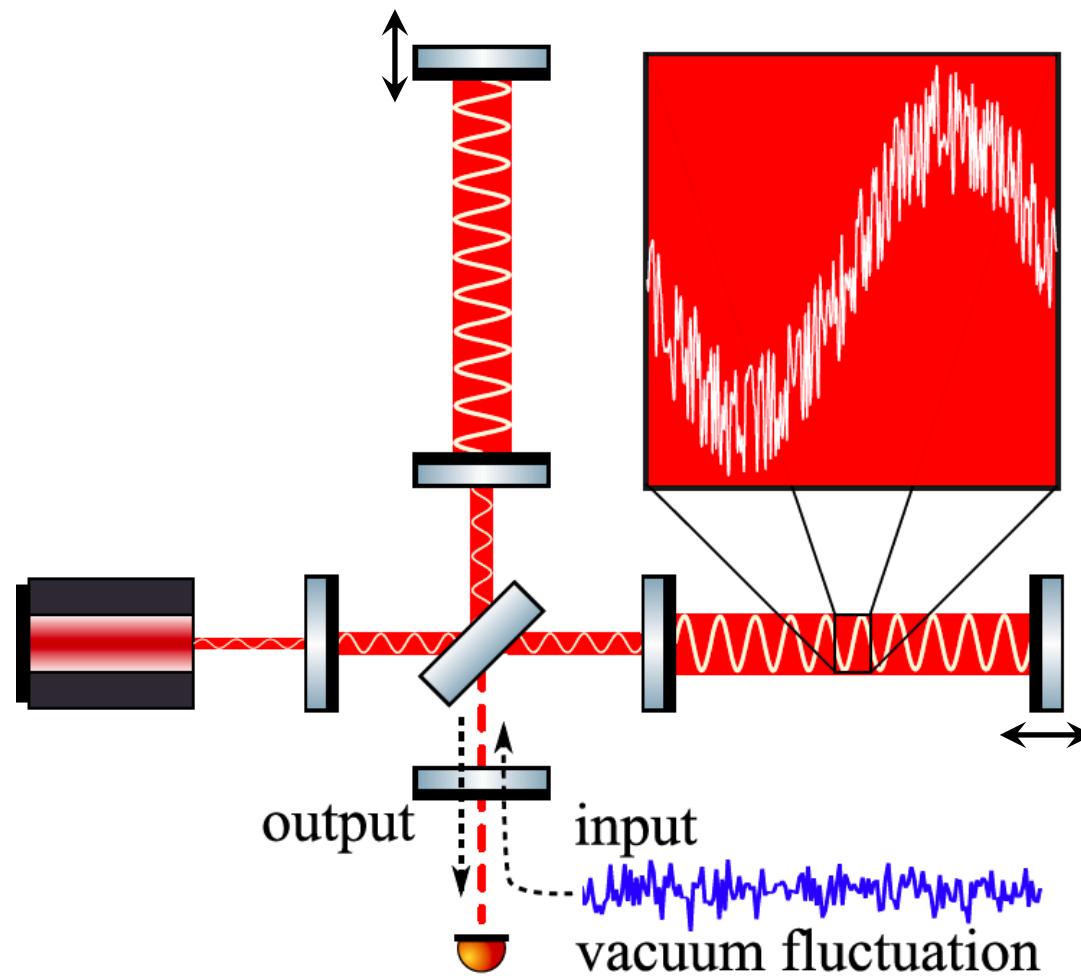
Quantum effects are indeed important!

Why Quantum?

Advanced LIGO design sensitivity curve [1]:



Origin of quantum noise



Ponderomotive (anti-)squeezing [2,3]
(manifestation of quantum back action) 5

Standard Quantum Limit (SQL)

Phase
fluctuation

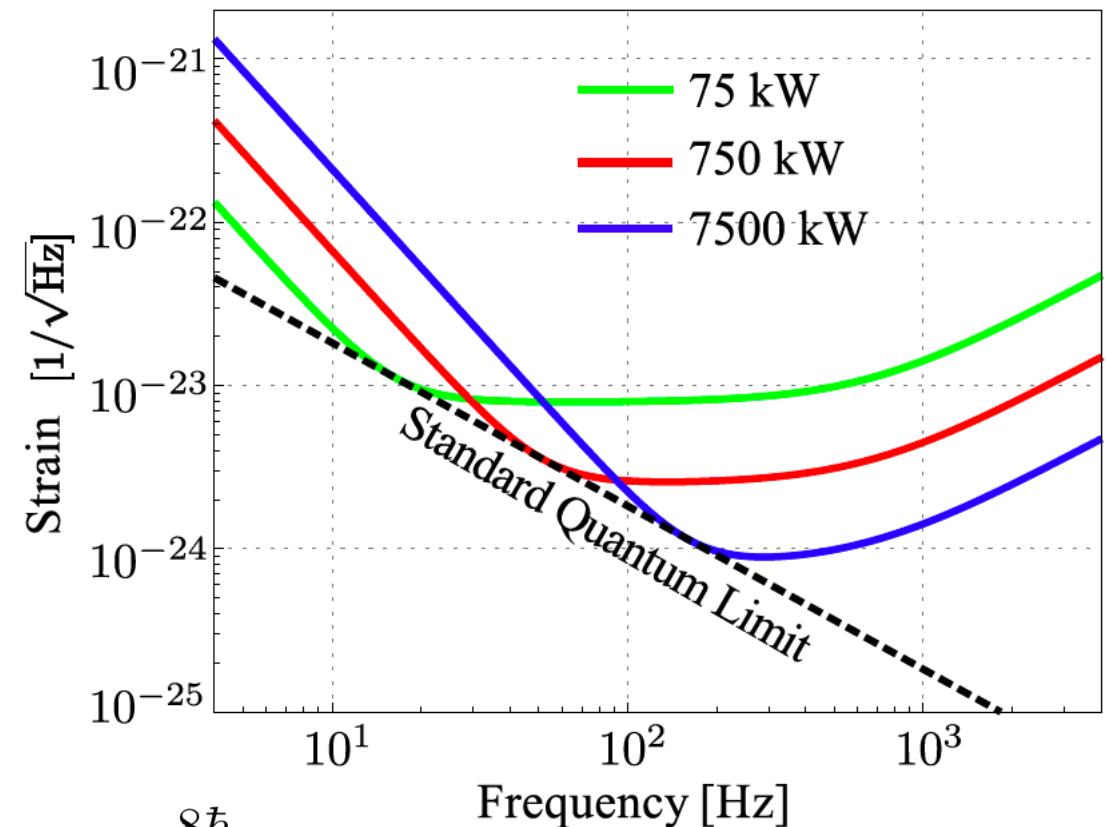
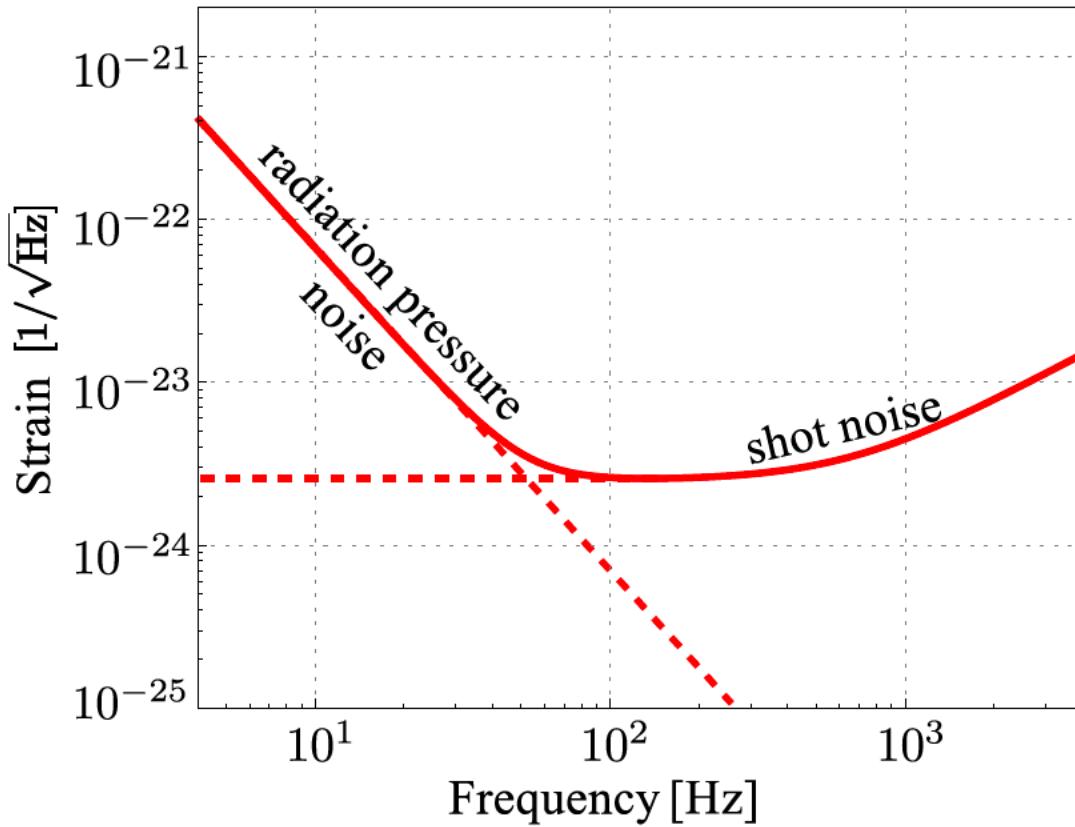


shot noise

Amplitude
fluctuation



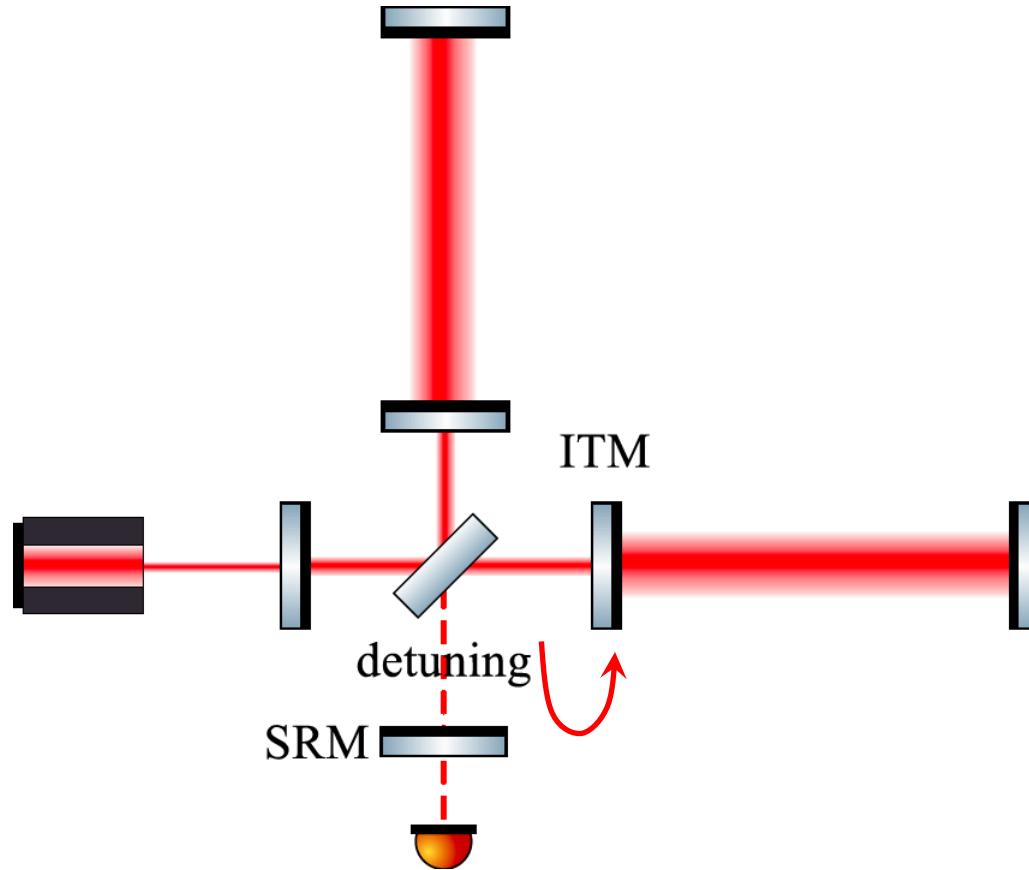
Radiation pressure
(back action) noise



$$S_{hh}^{\text{SQL}}(\Omega) = \frac{8\hbar}{M\Omega^2 L_{\text{arm}}^2} \quad [4]$$

$$\Omega = 2\pi f$$

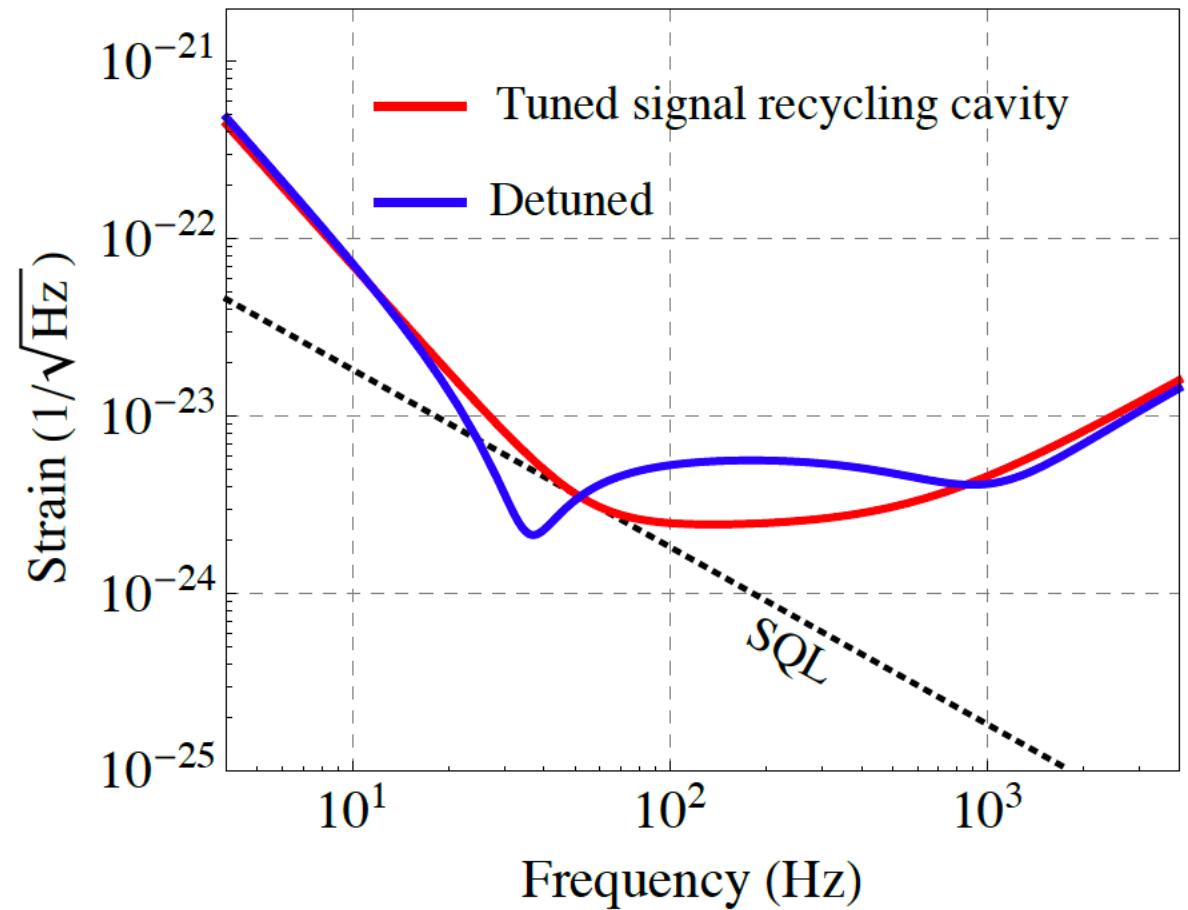
Surpassing the SQL by detuning the SRC



SRM: Signal recycling mirror

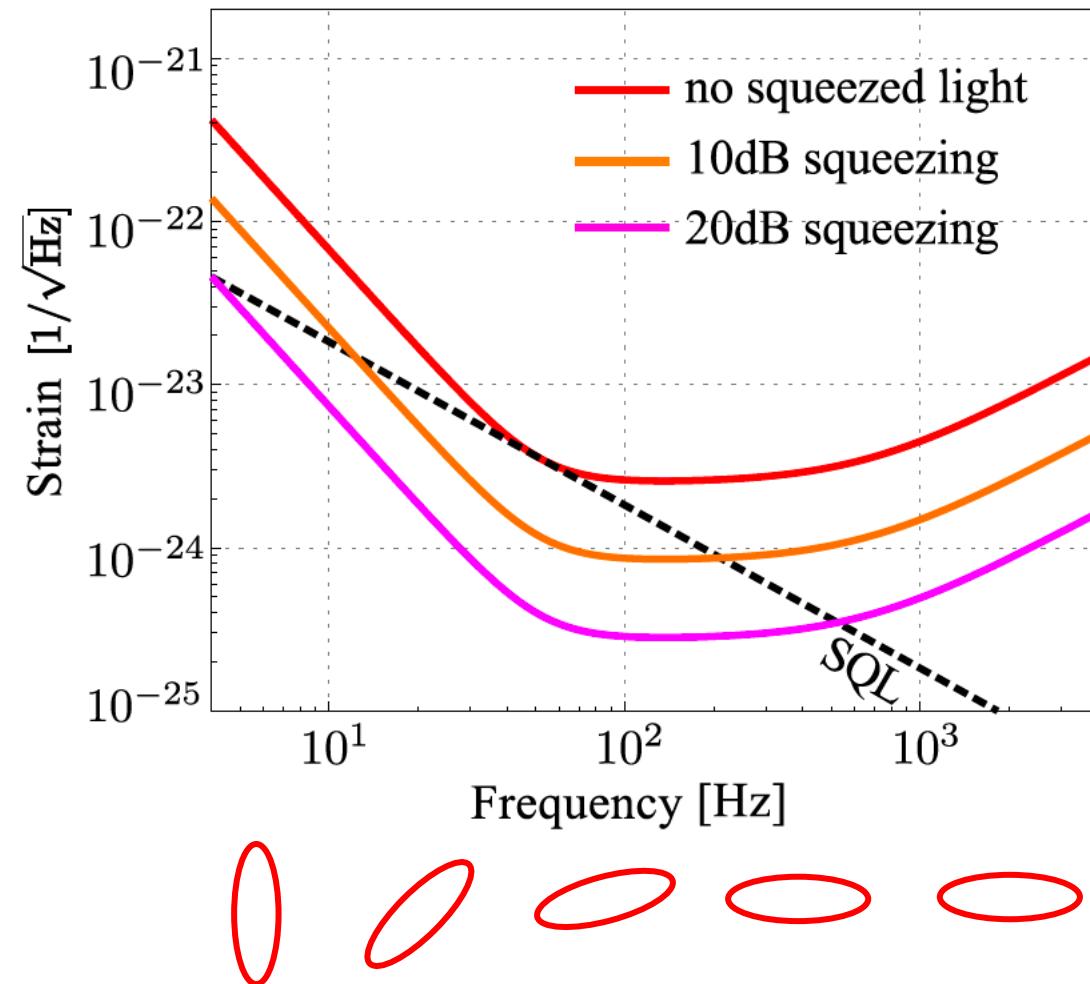
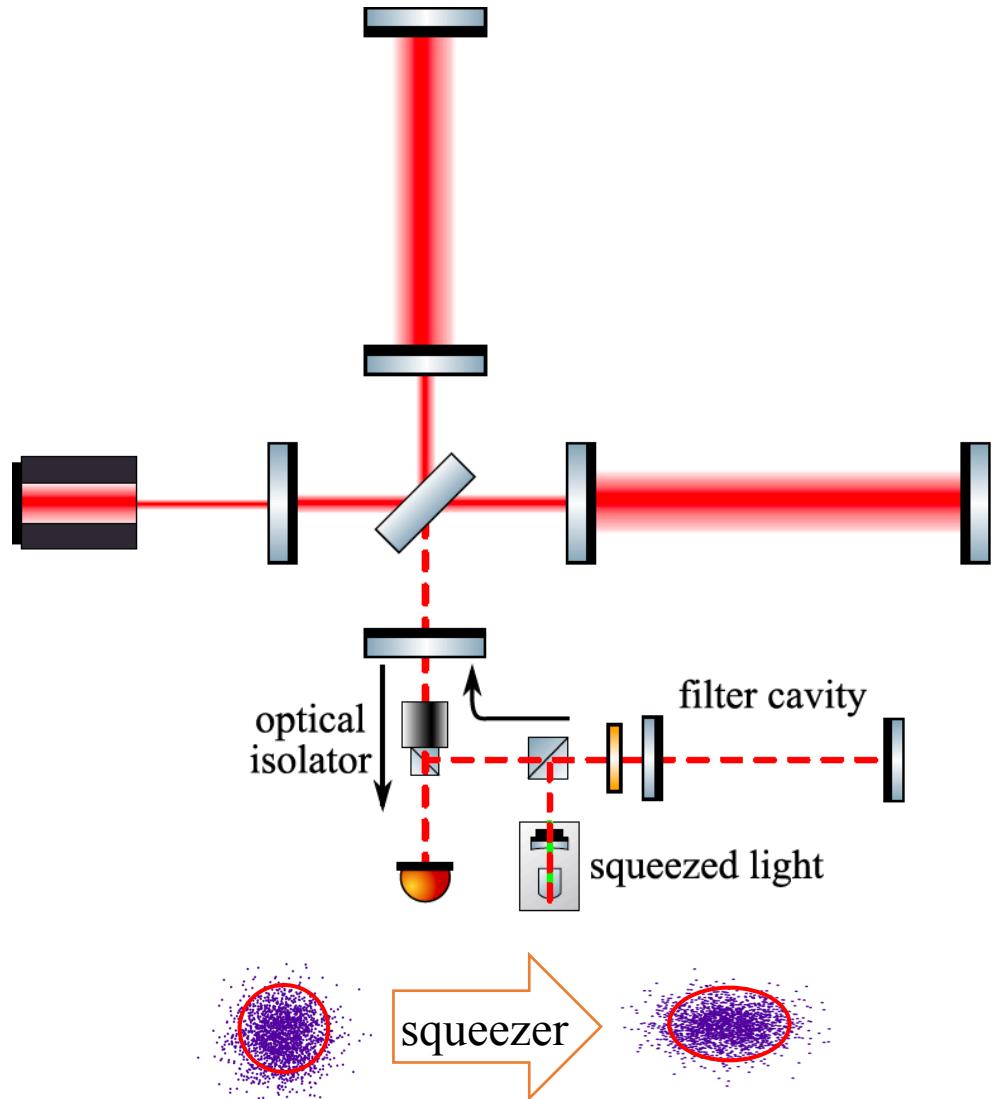
ITM: Input test mass mirror

SRC: Signal recycling cavity



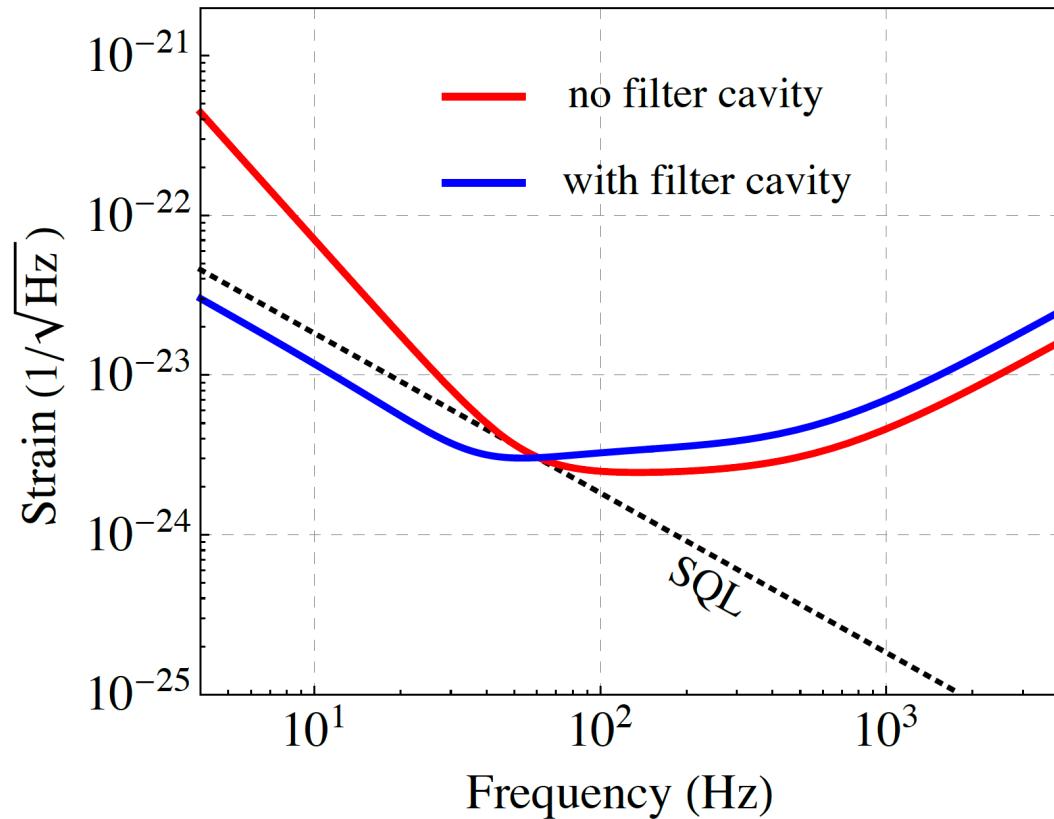
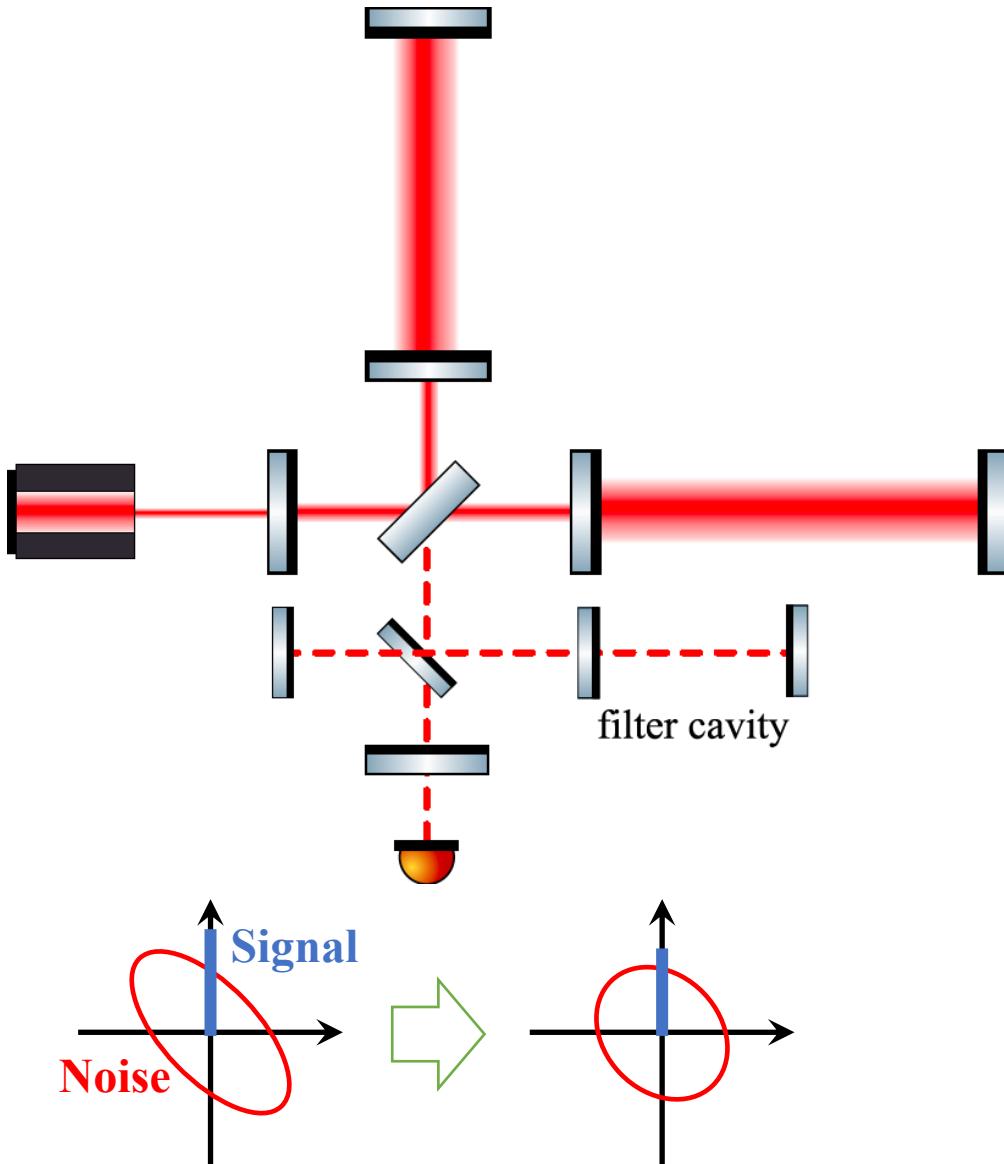
**Take advantage of the internal
ponderomotive squeezing
(back action as a resource) [5, 6]**

Input-filtering: Frequency-dependent squeezing



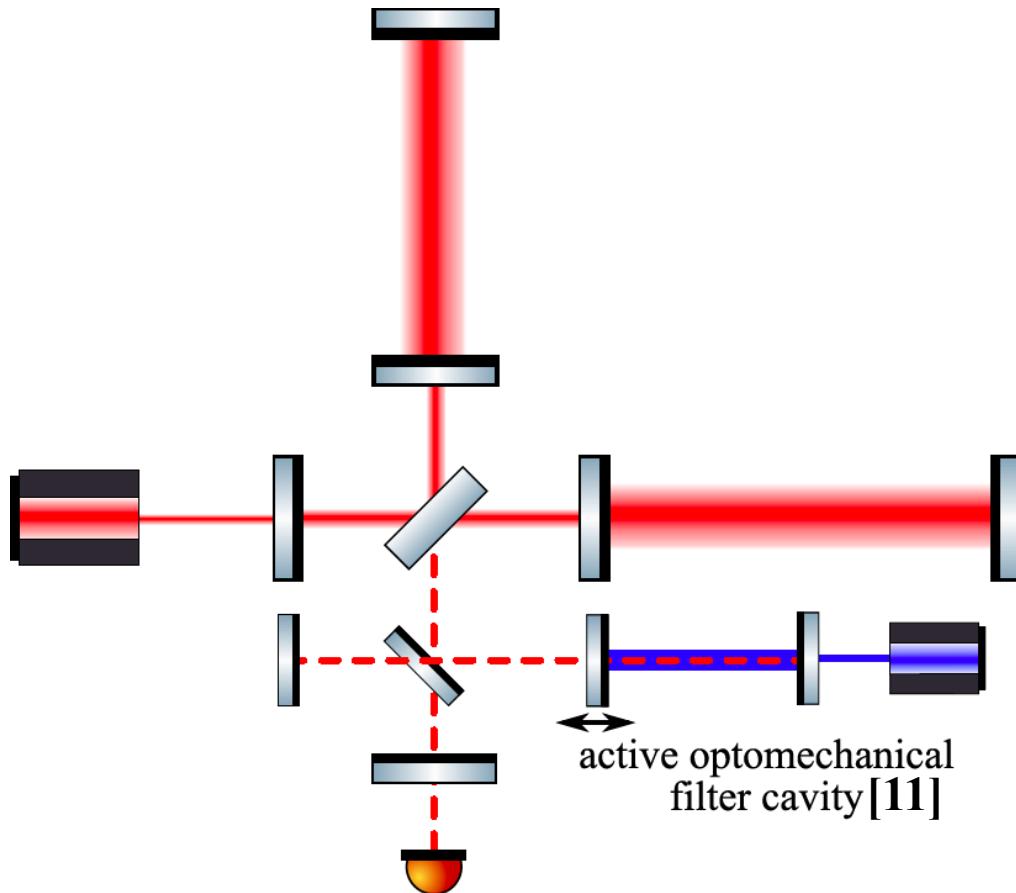
Filter cavity rotates input squeezing angle [7]
counteracts ponderomotive (anti-)squeezing

Intra-cavity filtering: Speed meter

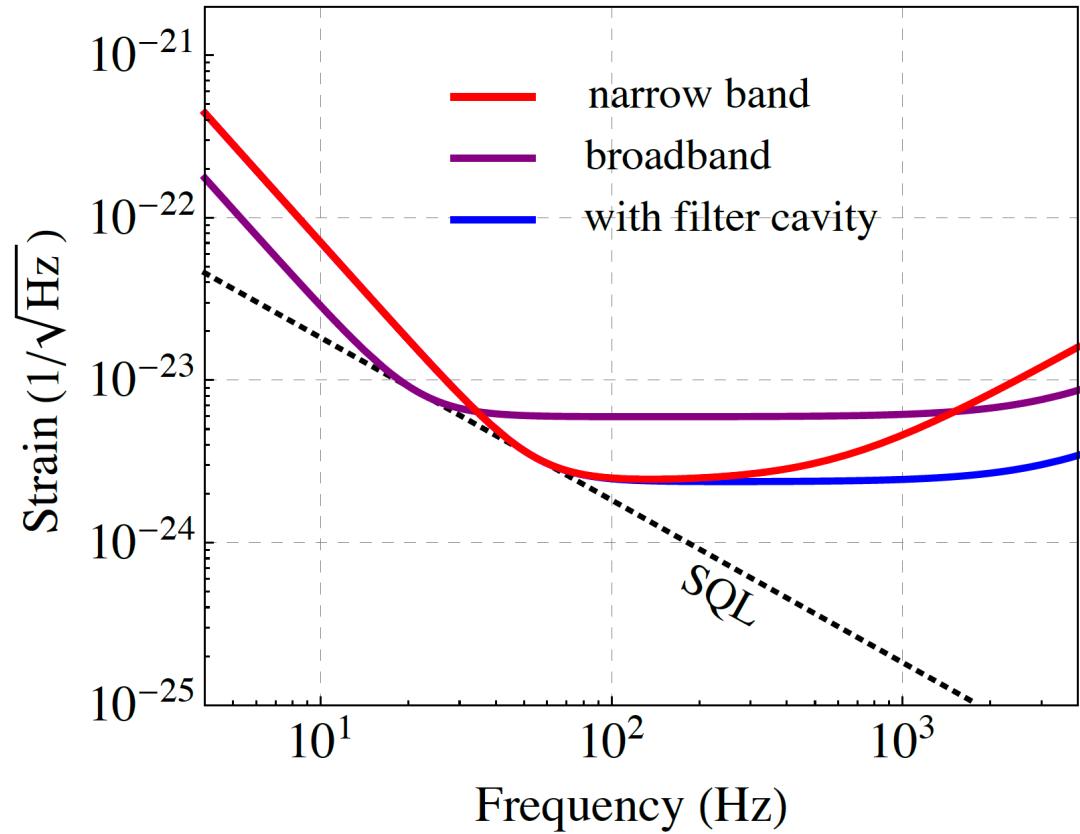


Sacrificing the signal response but decreasing the ponderomotive (anti-)squeezing more [8]
An equivalent realisation with Sagnac [9, 10]

Intra-cavity filtering: White-light cavity (“fast light”)

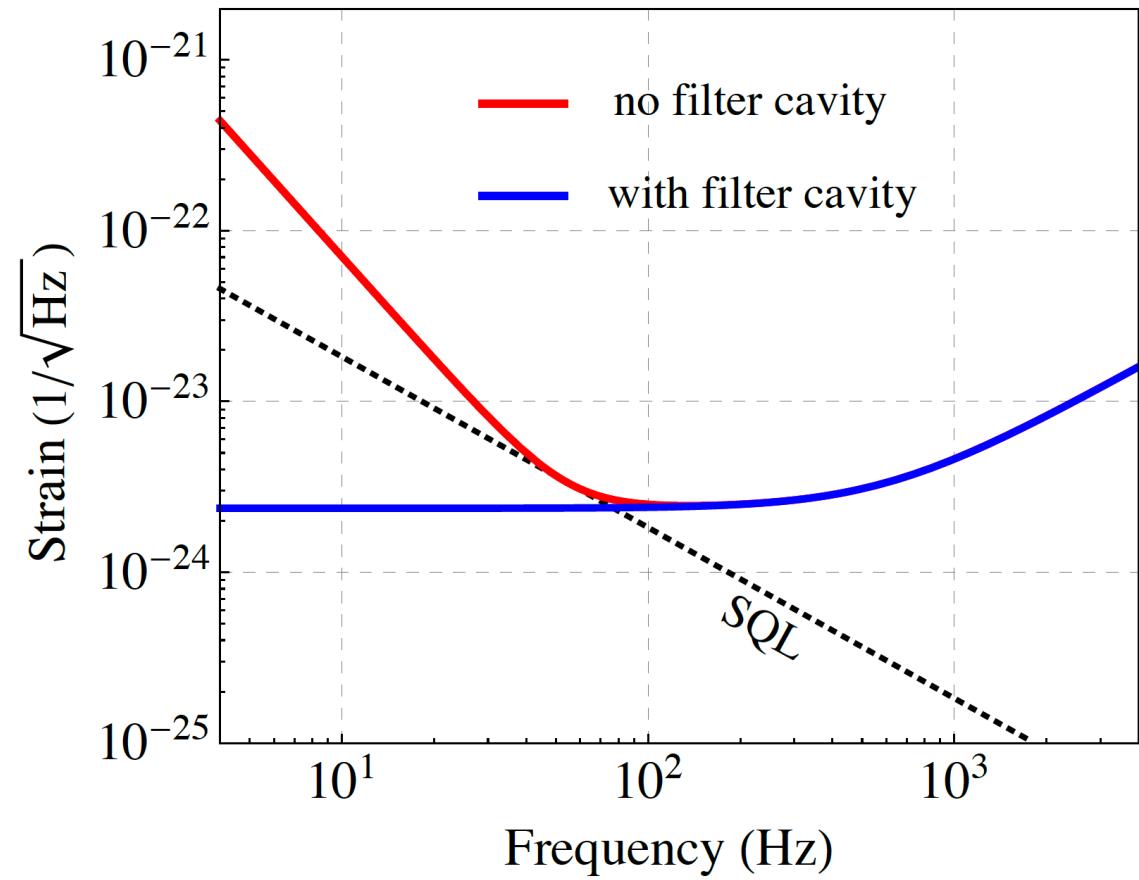
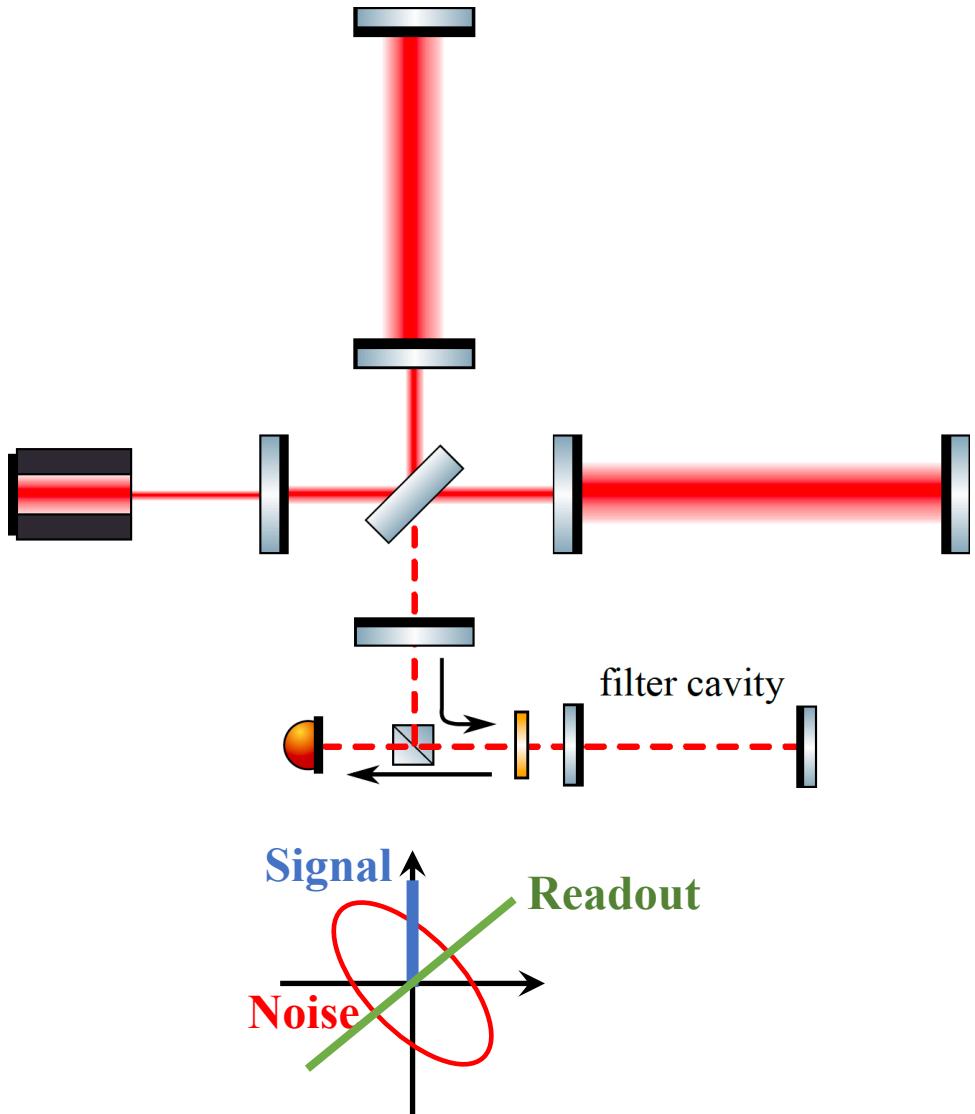


Can be realised also with
active atomic system [12, 13]



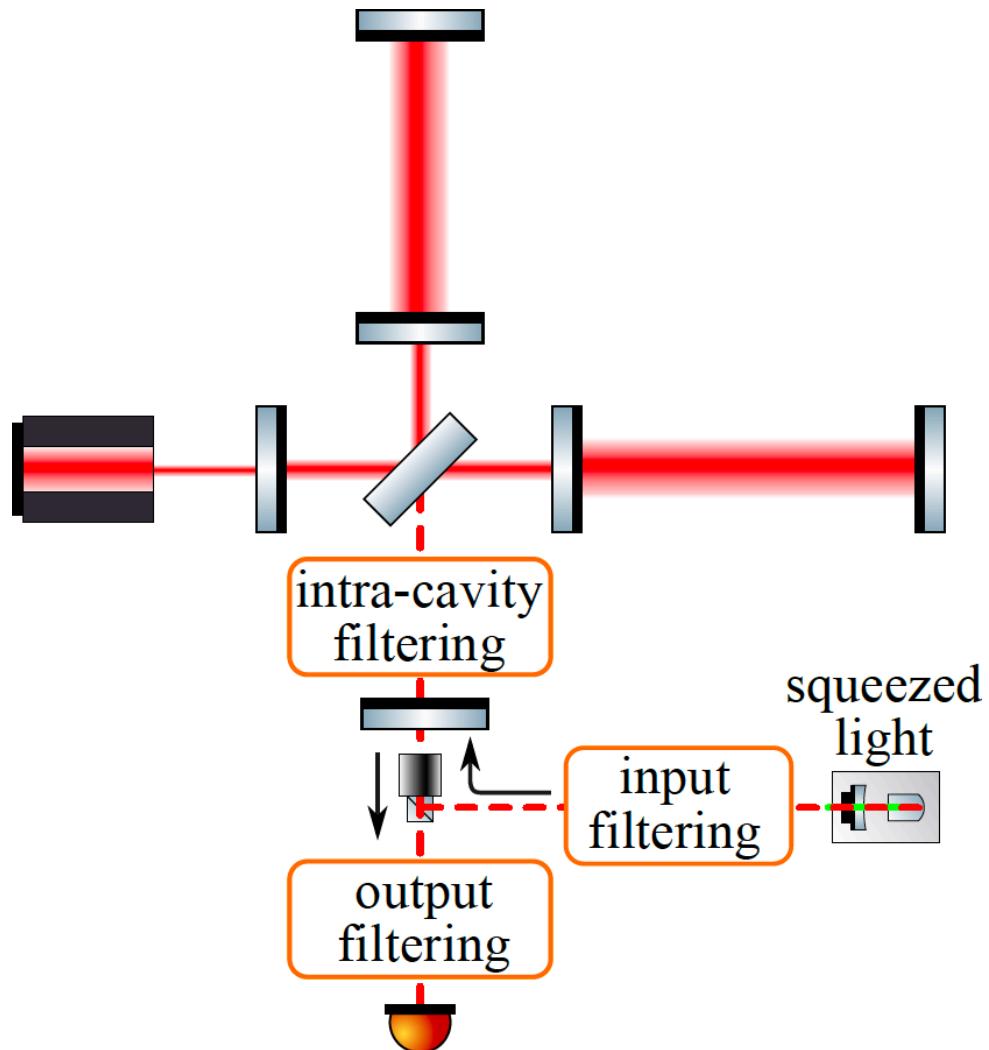
Enhancing the bandwidth without reducing
the peak sensitivity

Output filtering: Frequency-dependent readout



Read out the optimal quadrature that has the maximum SNR [7]

General case



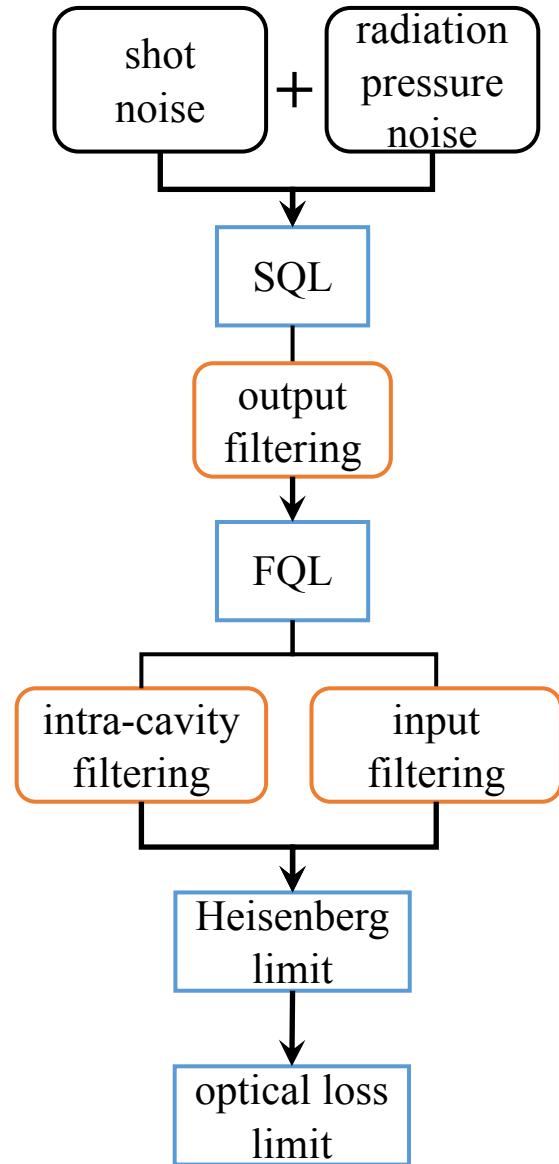
Each filtering module can be a cascade of passive/active filter cavities. **Infinite combinations!**

Question:

How do we combine techniques in a **systematic** way?

What is the **optimal** scheme?

A unified framework



Fundamental Quantum Limit (FQL) [6, 14 - 16]

$$S_{hh}^{\text{FQL}}(\Omega) = \frac{\hbar^2 c^2}{2L_{\text{arm}}^2 S_{PP}(\Omega)} = \frac{2\hbar^2}{S_{EE}(\Omega)}$$

Intuition: time-energy uncertainty relation

$$\Delta t \geq \hbar / \Delta E$$

Heisenberg Limit (all photons entangled)

$$S_{EE} \propto N_{\text{photon}}^2 \text{ rather than } N_{\text{photon}}$$

$$N_{\text{photon}}^{\text{LIGO}} \sim 10^{20}$$

Optical loss limit [17, 18]

$$S_{EE} \propto N_{\text{photon}} / \epsilon \quad \epsilon_{\text{LIGO}} \sim 10^{-4}$$

This framework applies to other linear quantum measurements.

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