

Identifying Dark Matter with MeV Gamma-Ray Telescopes

Chris Cappiello

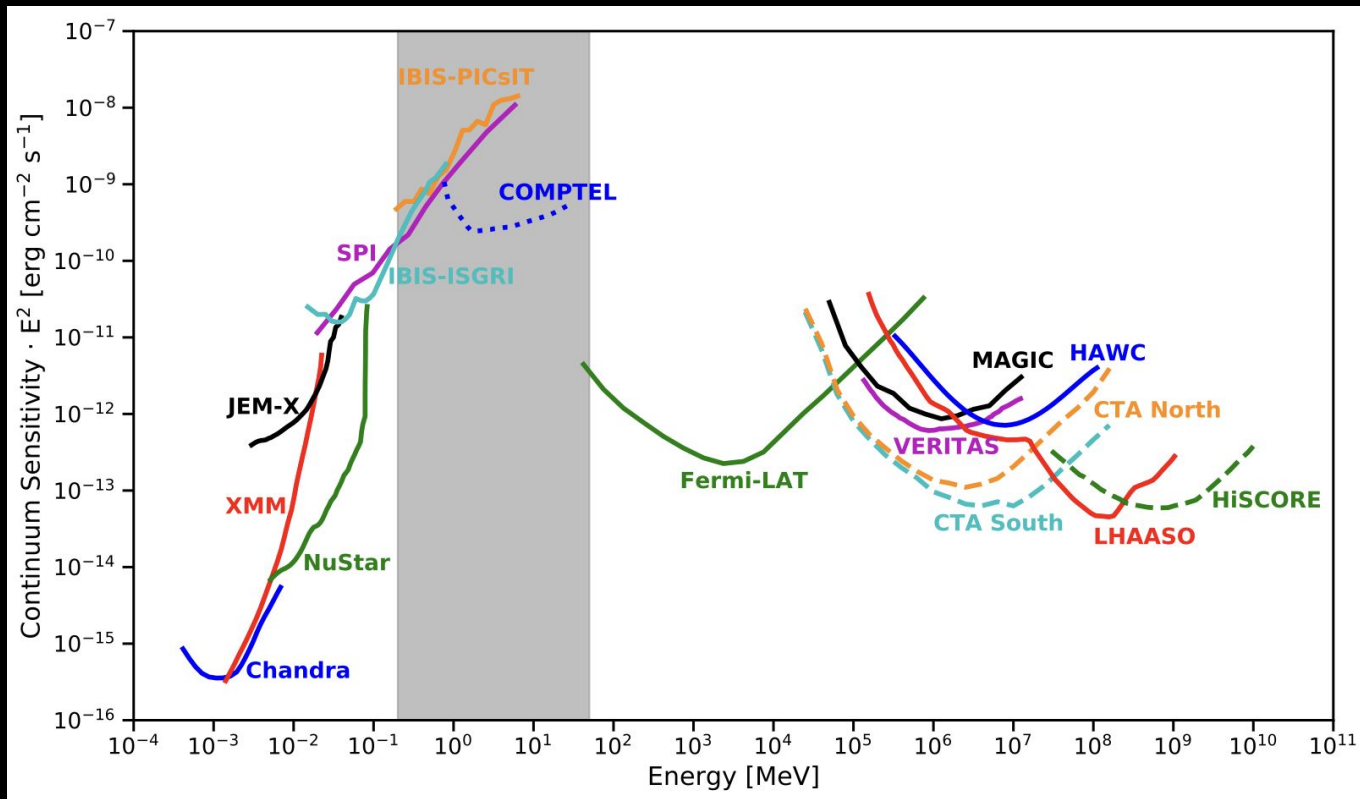
w/ James Dent, Bhaskar Dutta, and Jeremy Padvorac

Light Dark World

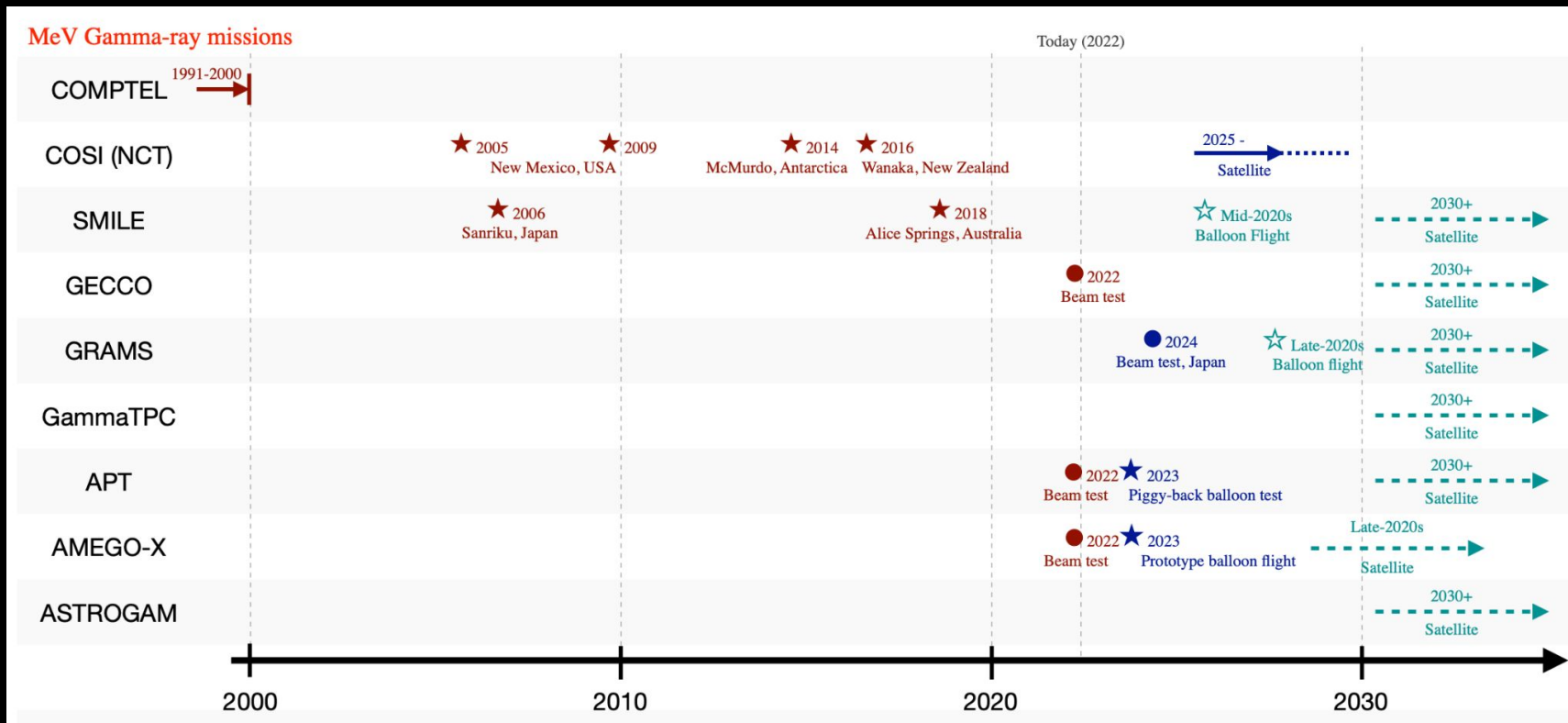
Carleton University, July 28, 2026



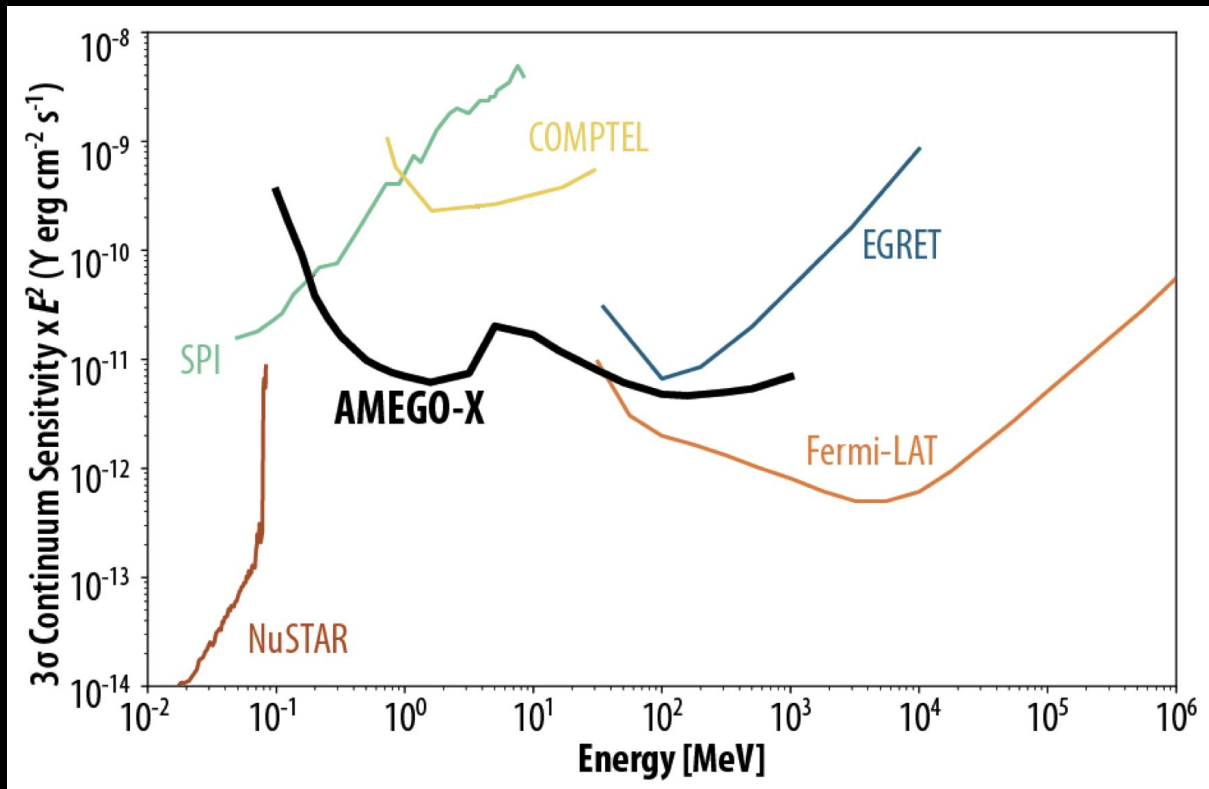
The MeV Gap



The MeV Gap

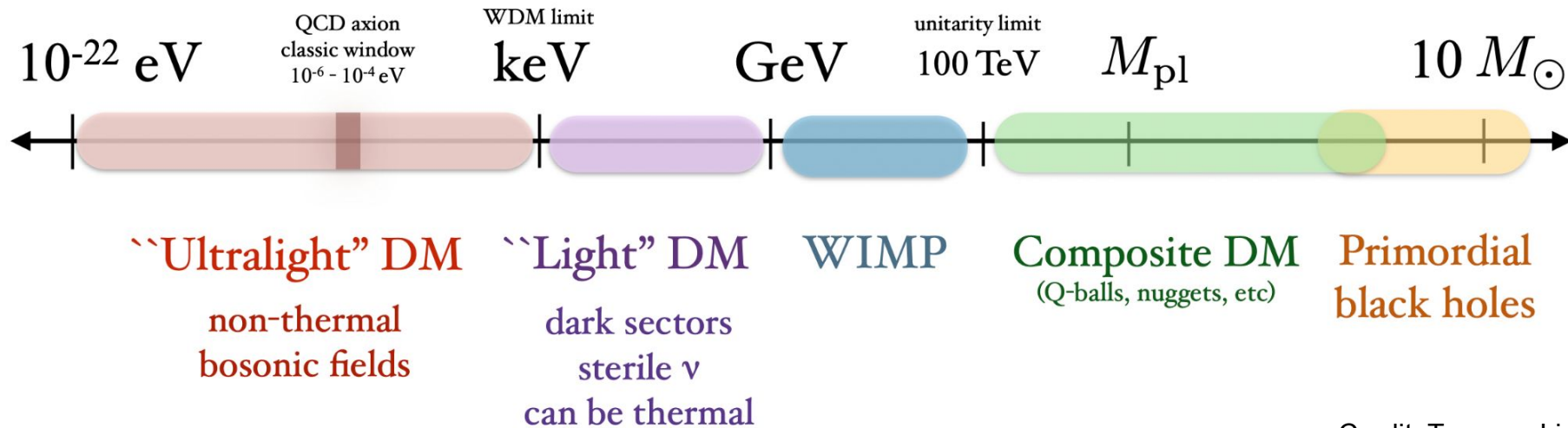


The MeV Gap



Mass scale of dark matter

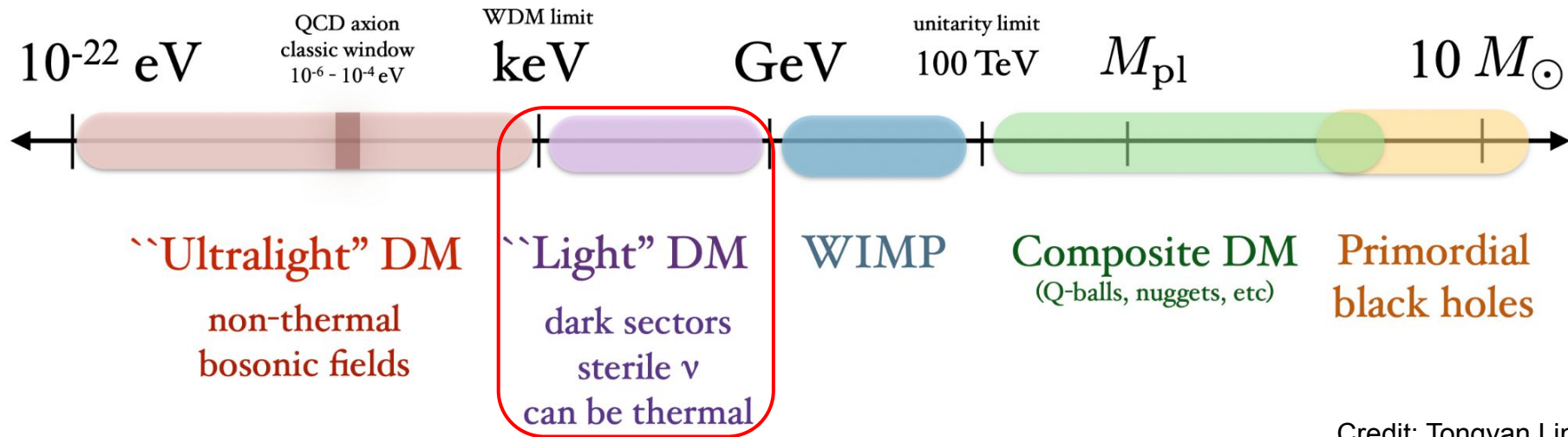
(not to scale)



Credit: Tongyan Lin

Mass scale of dark matter

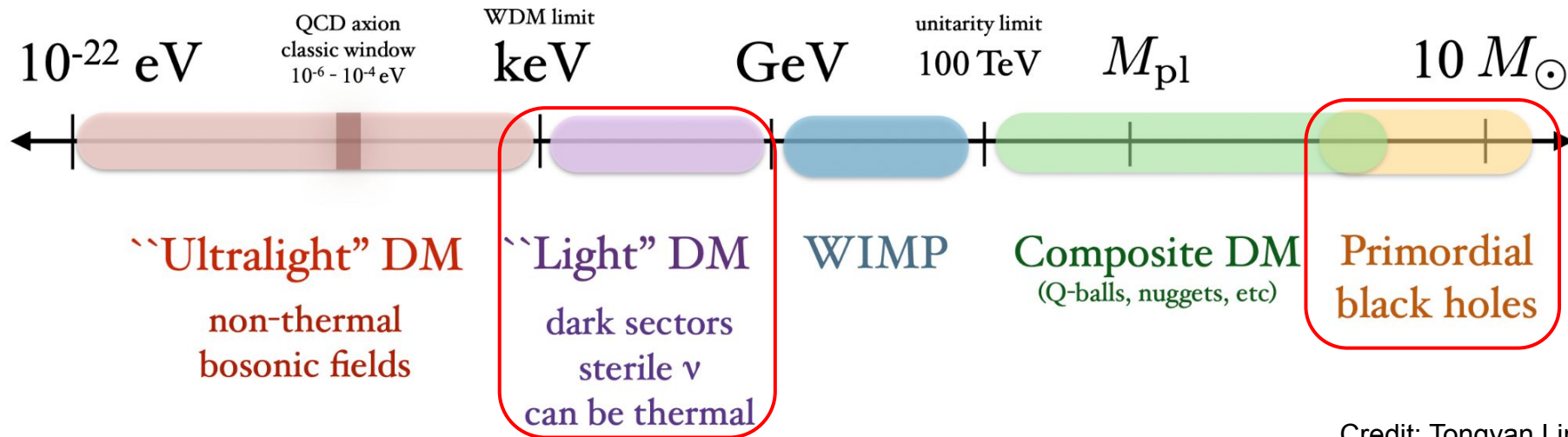
(not to scale)



Credit: Tongyan Lin

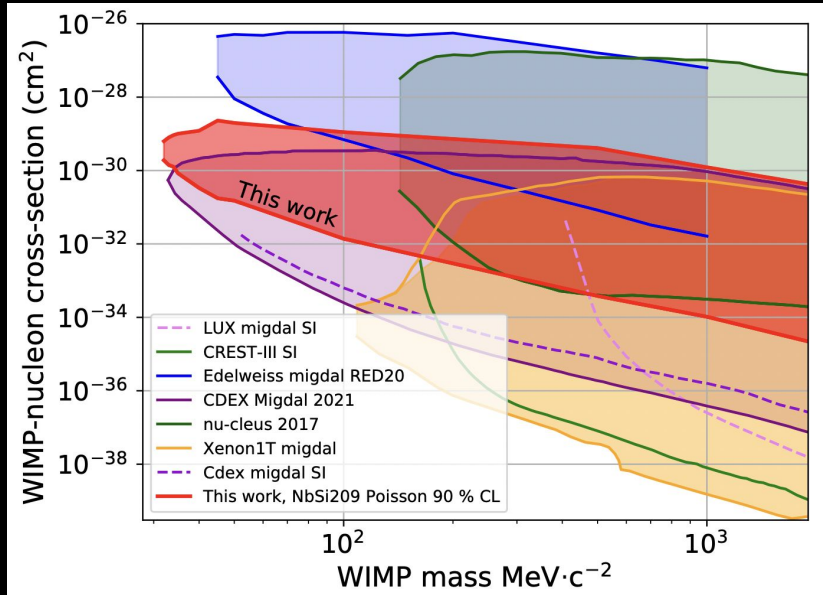
Mass scale of dark matter

(not to scale)

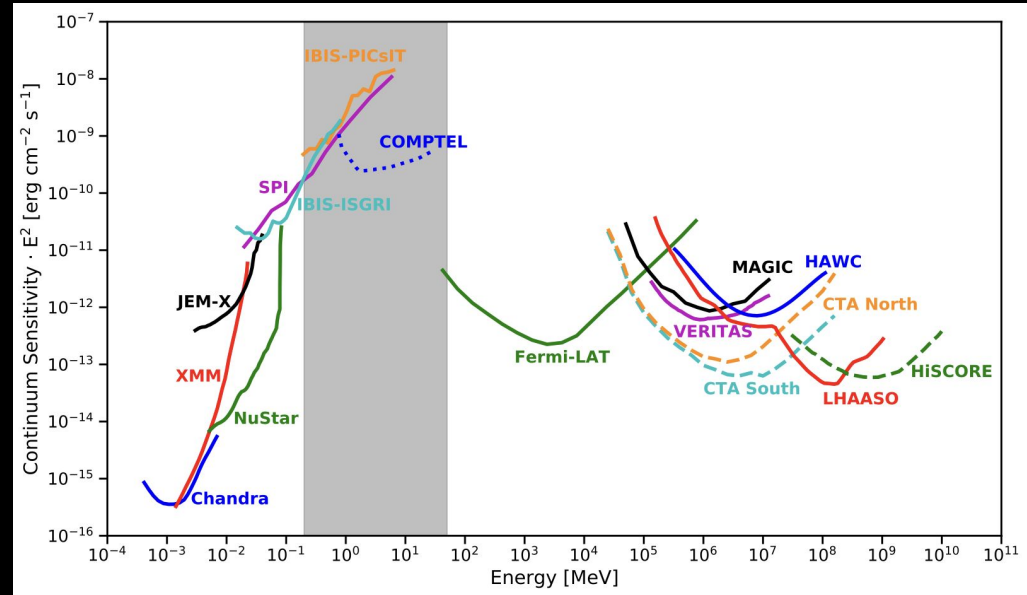


Credit: Tongyan Lin

Sub-GeV Dark Matter

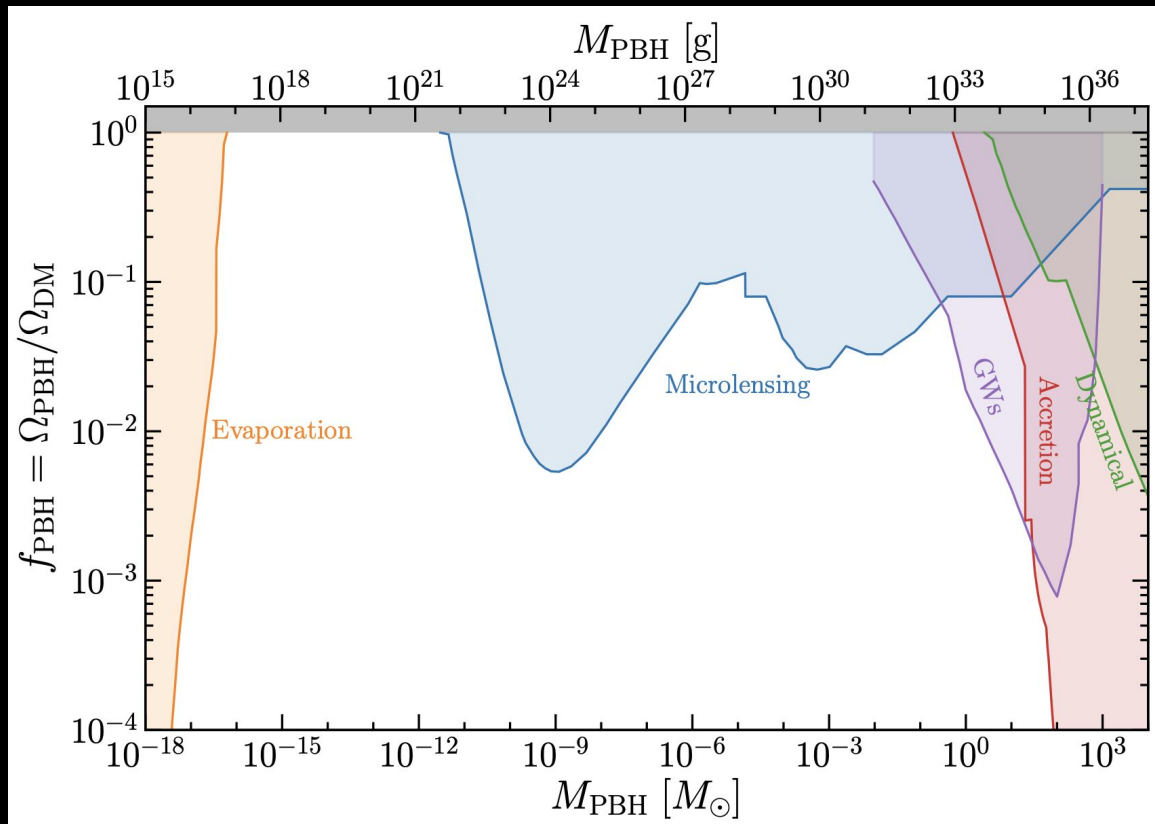


arXiv:2203.03993 (Edelweiss)



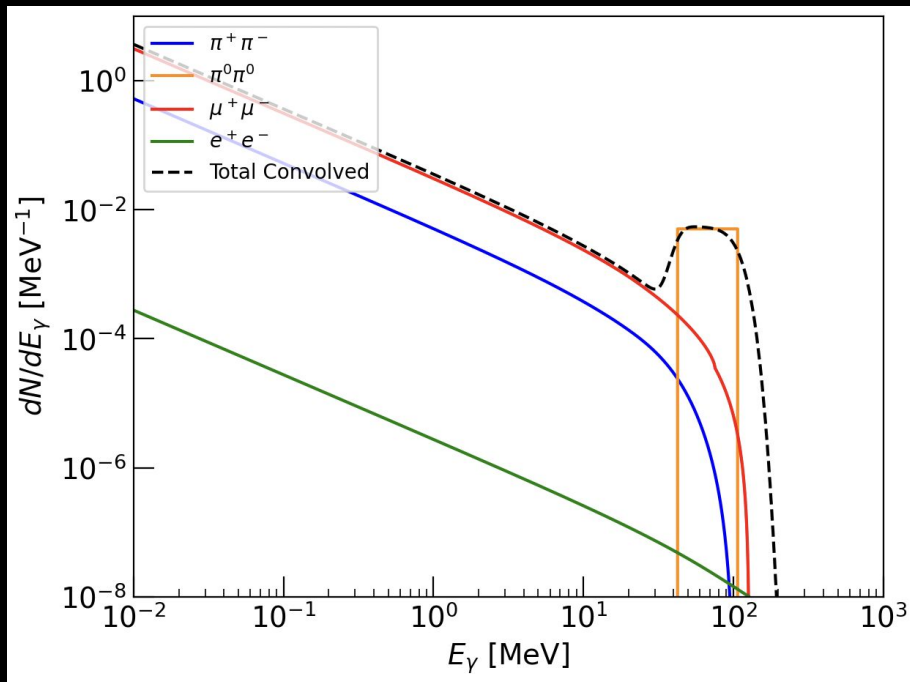
arXiv:2204.01325 (MeVCube)

Primordial Black Holes

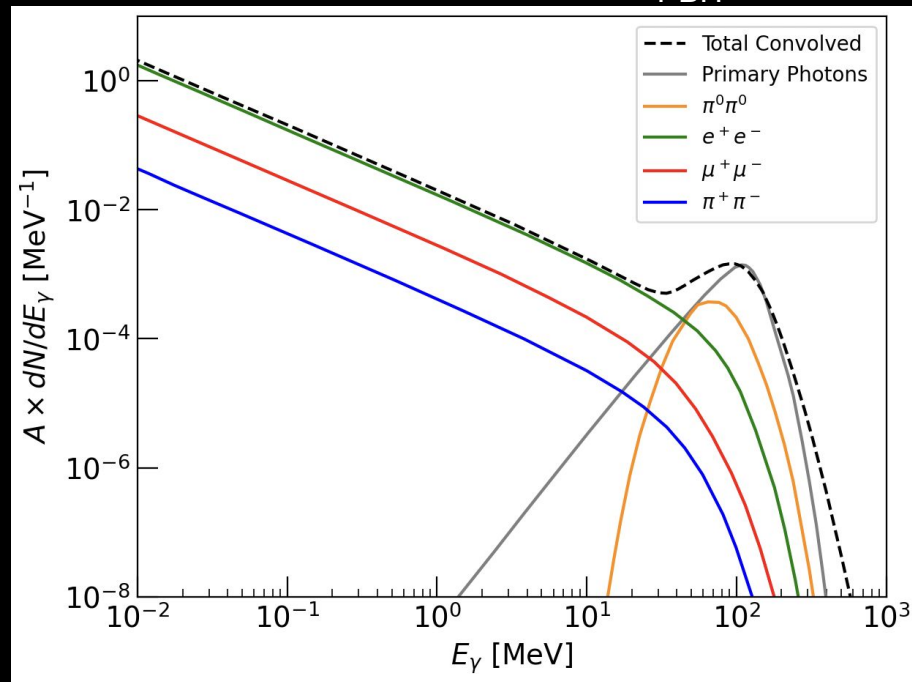


Gamma-Ray Spectra

Scalar DM, $M = 300 \text{ MeV}$, $T = 10^{28} \text{ s}$



PBH DM, $M = 5.5 \cdot 10^{14} \text{ g}$, $f_{\text{PBH}} < 10^{-9.5}$



Dark Matter Models

Particle Dark Matter

- Scalar w/ couplings equal to Standard Model Yukawas
- Vector (kinetically mixed dark photon)
- For both, assume lifetime = 10^{28} s

Primordial Black Holes

- Monochromatic distribution
- Spin 0, Charge 0
- Fit subject to constraints on f_{PBH}
- Extremal PBHs: work in progress

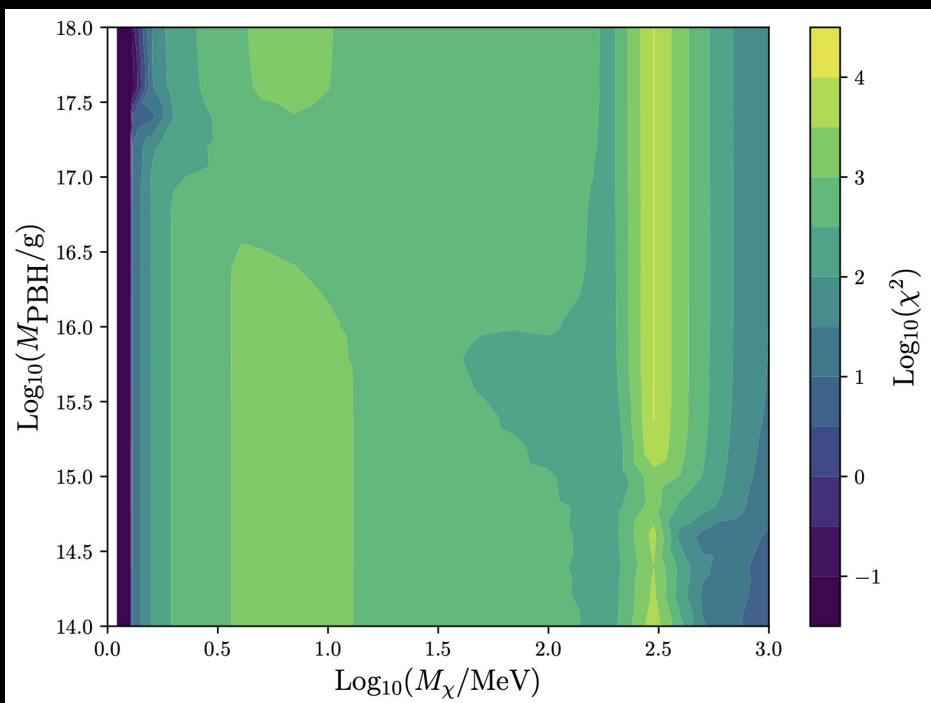
Statistical Analysis

Compare particle dark matter and PBH dark matter spectra

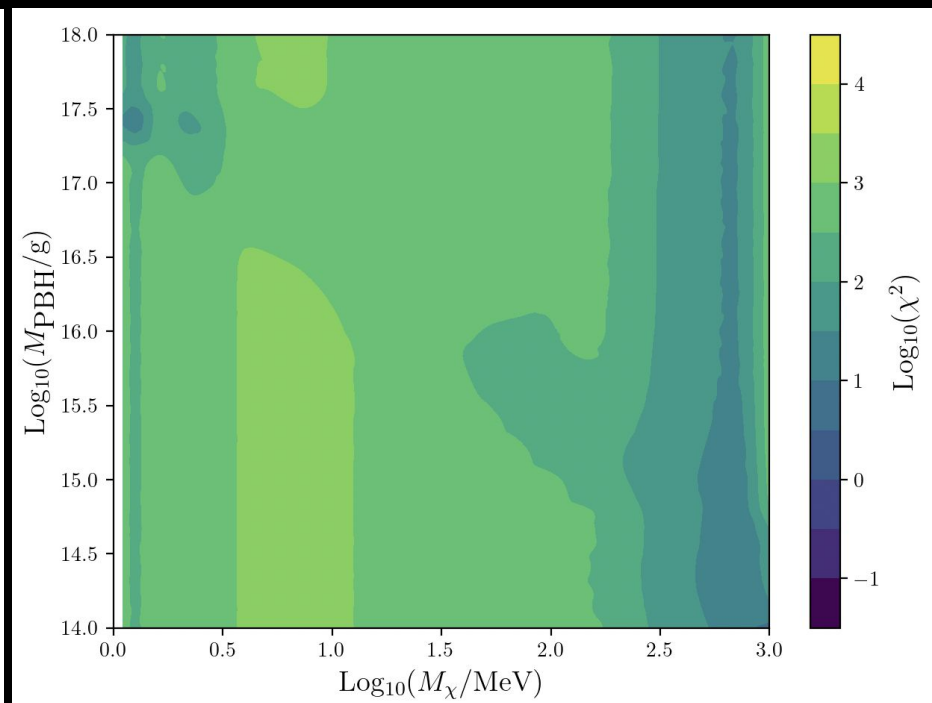
- Generate Asimov dataset for particle dark matter plus background
- Compute delta chi squared between particle dark matter plus background and PBH dark matter plus background
- Determine how well an instrument like AMEGO-X can distinguish these spectra, how much observation time is needed, etc.
- Start by assuming observation time of 10^8 s

PBH Model Fit to Particle DM

Scalar DM

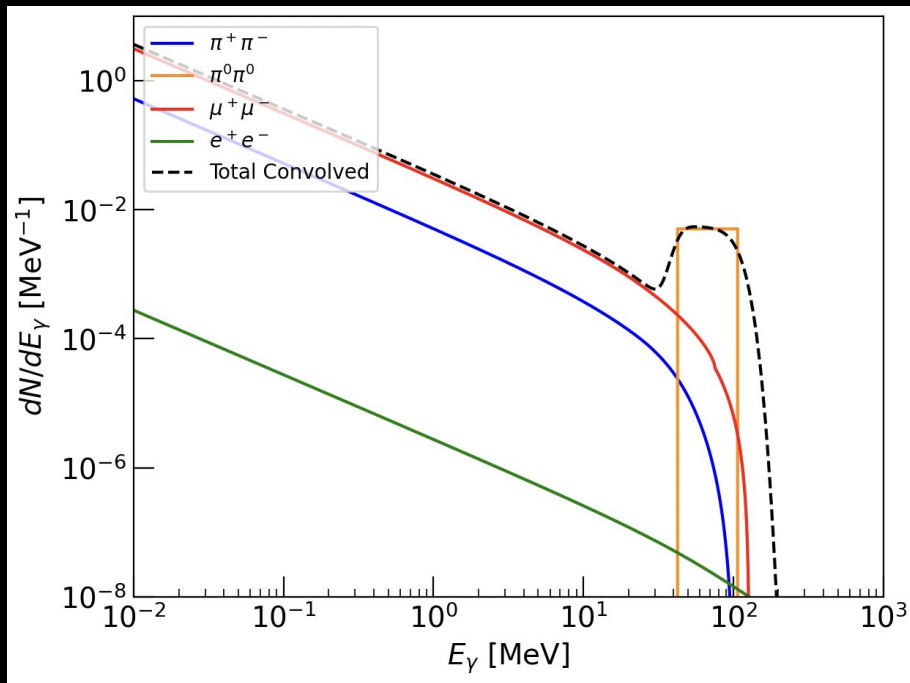


Vector DM

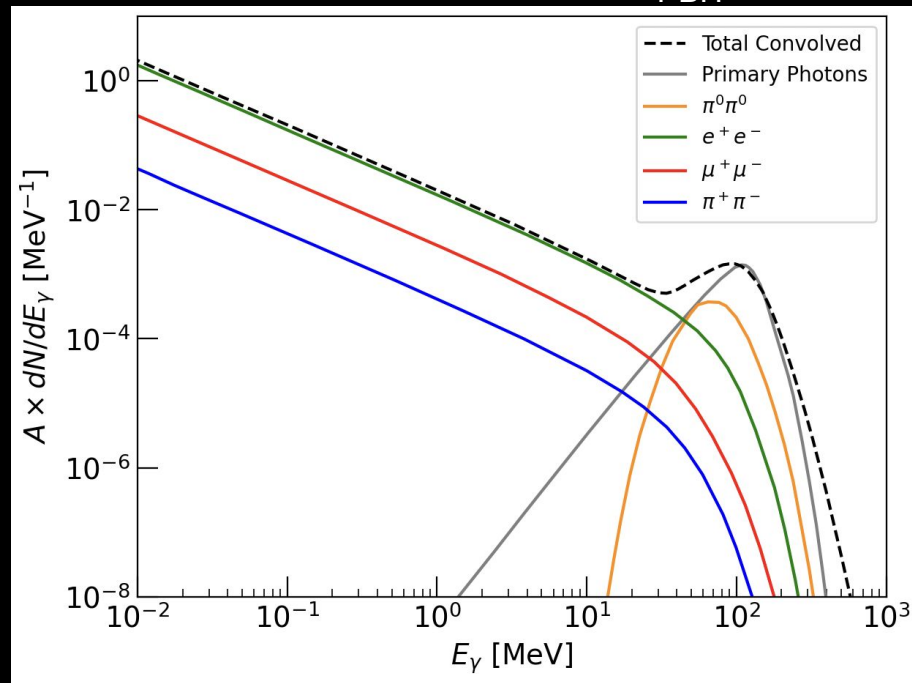


Gamma-Ray Spectra

Scalar DM, $M = 300 \text{ MeV}$, $T = 10^{28} \text{ s}$

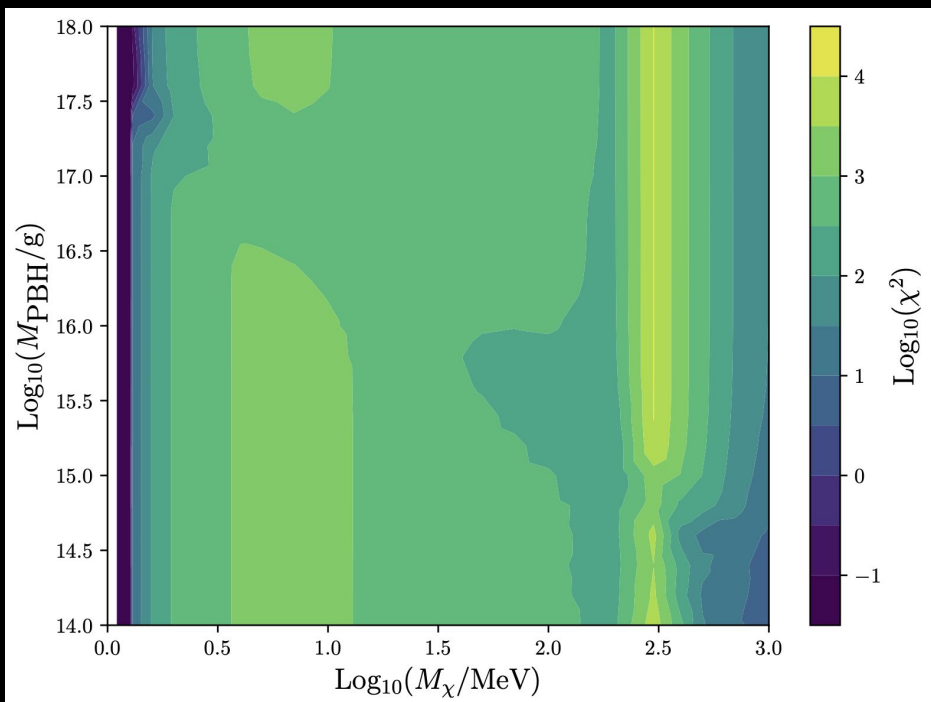


PBH DM, $M = 5.5 \cdot 10^{14} \text{ g}$, $f_{\text{PBH}} < 10^{-9.5}$

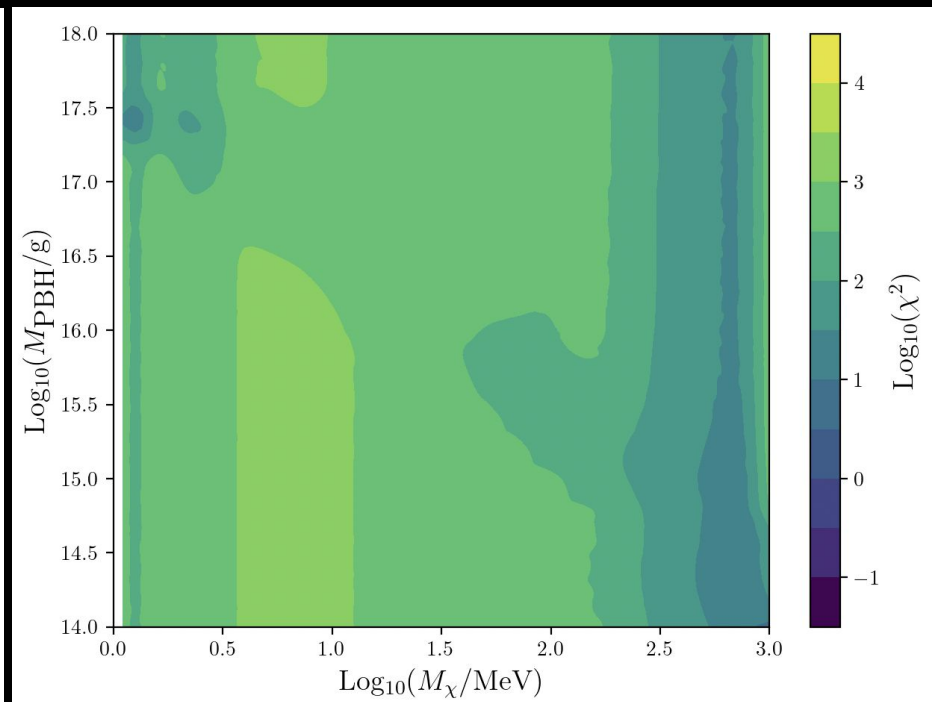


PBH Model Fit to Particle DM

Scalar DM

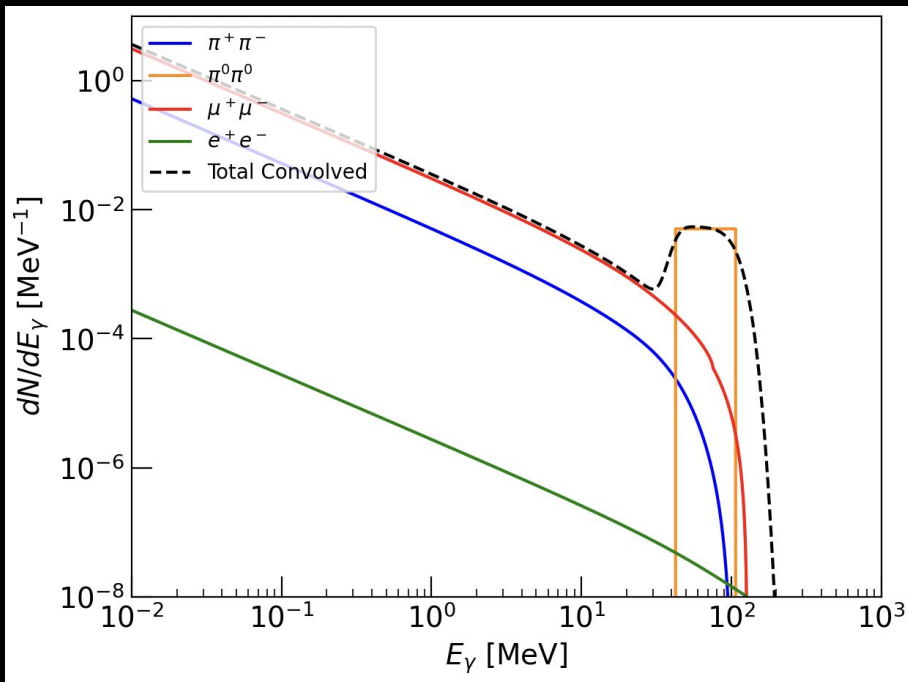


Vector DM

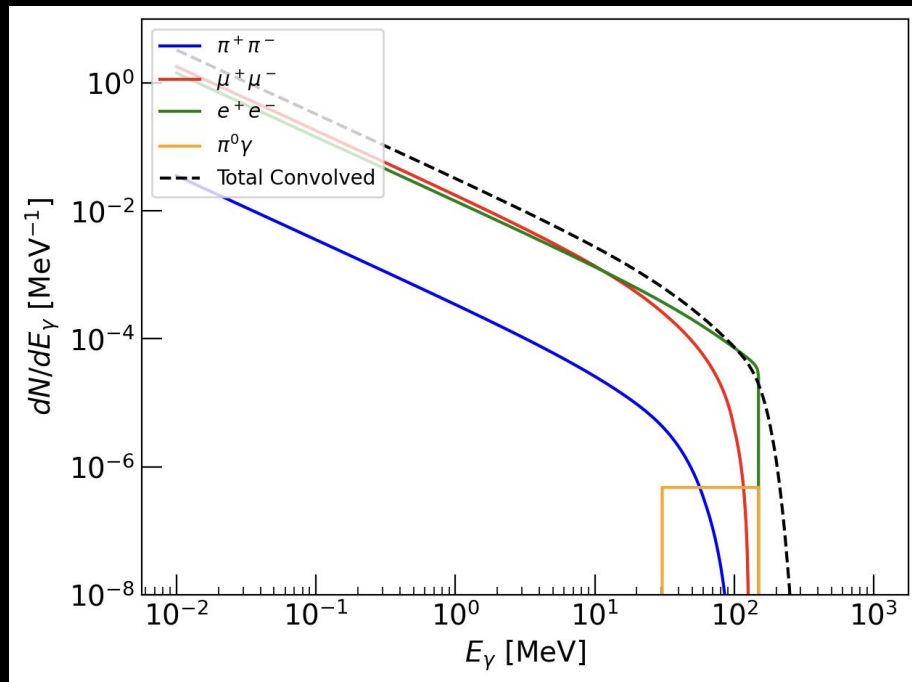


Gamma-Ray Spectra

Scalar DM, $M = 300$ MeV, $T = 10^{28}$ s

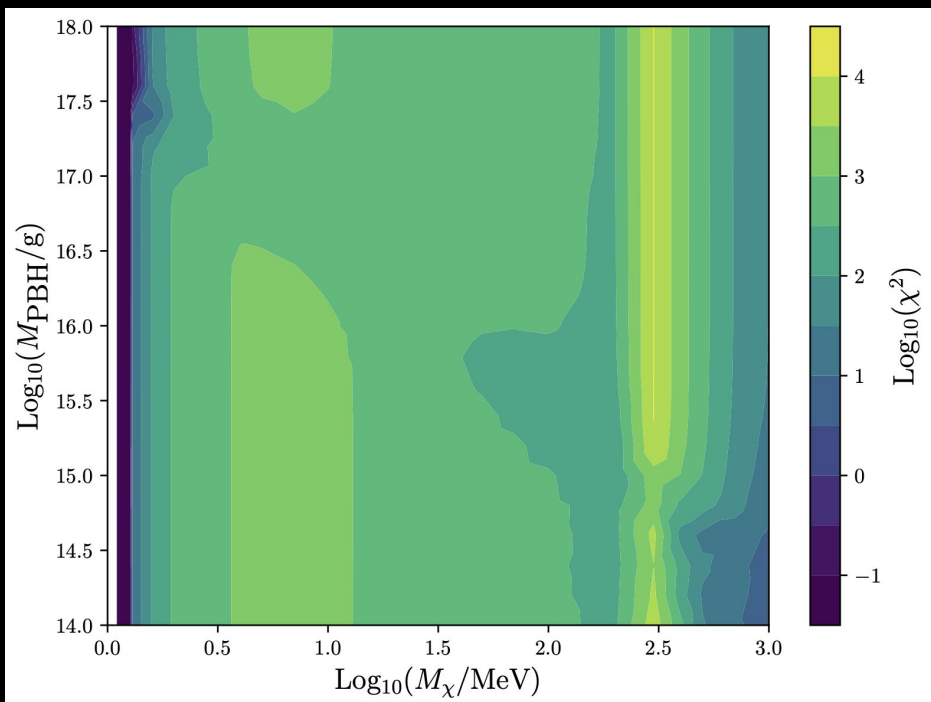


Vector DM, $M = 300$ MeV, $T = 10^{28}$ s

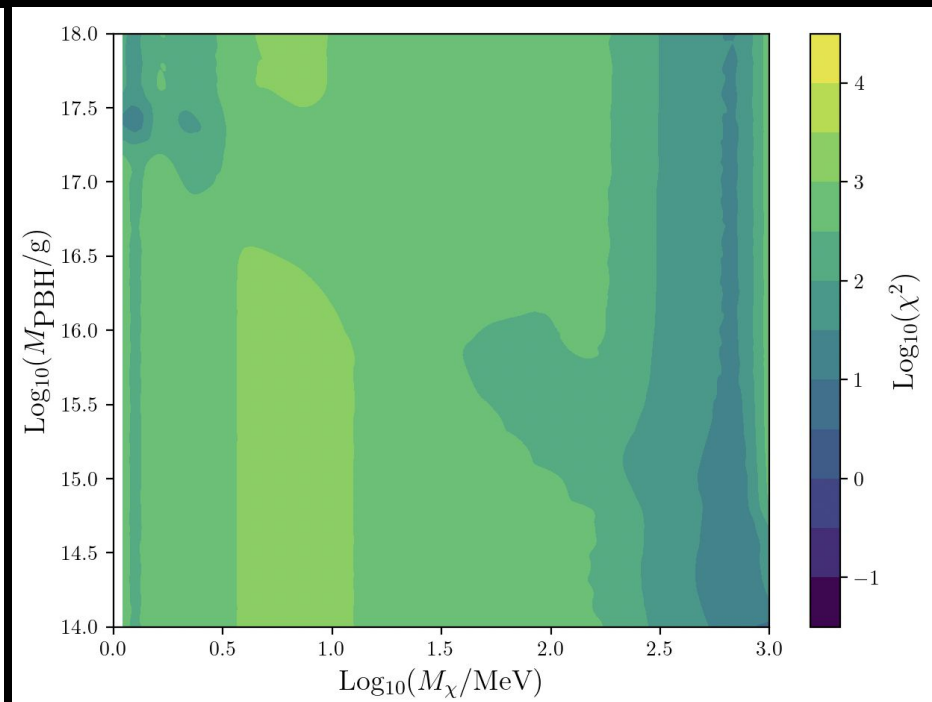


PBH Model Fit to Particle DM

Scalar DM



Vector DM



Required Observation Time

Scalar DM

Vector DM

