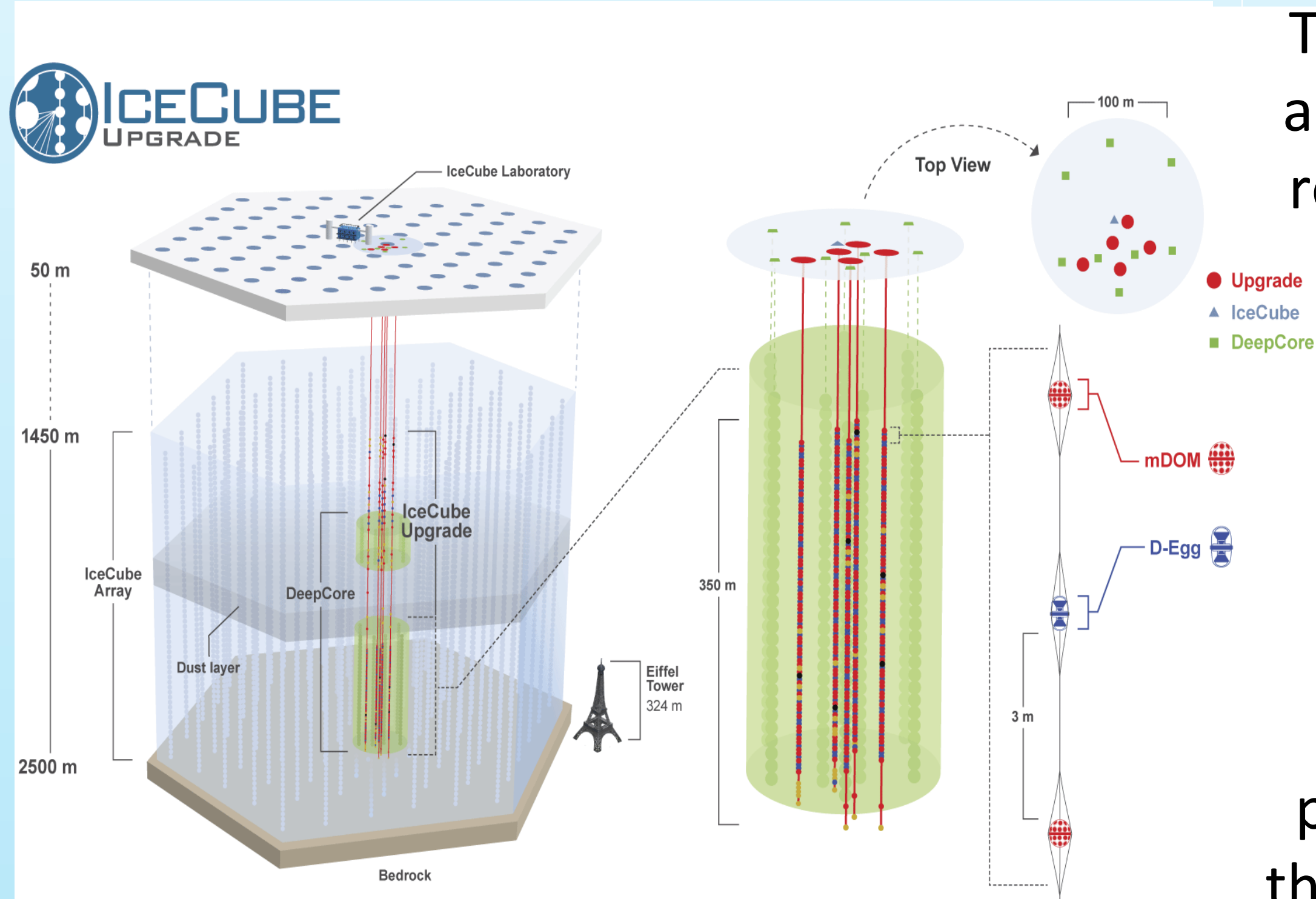


Justin Mitchell¹ for the Southern University Group — Ali R. Fazely¹, Sean Roubion¹

¹Southern University and A&M College | IceCube Collaboration

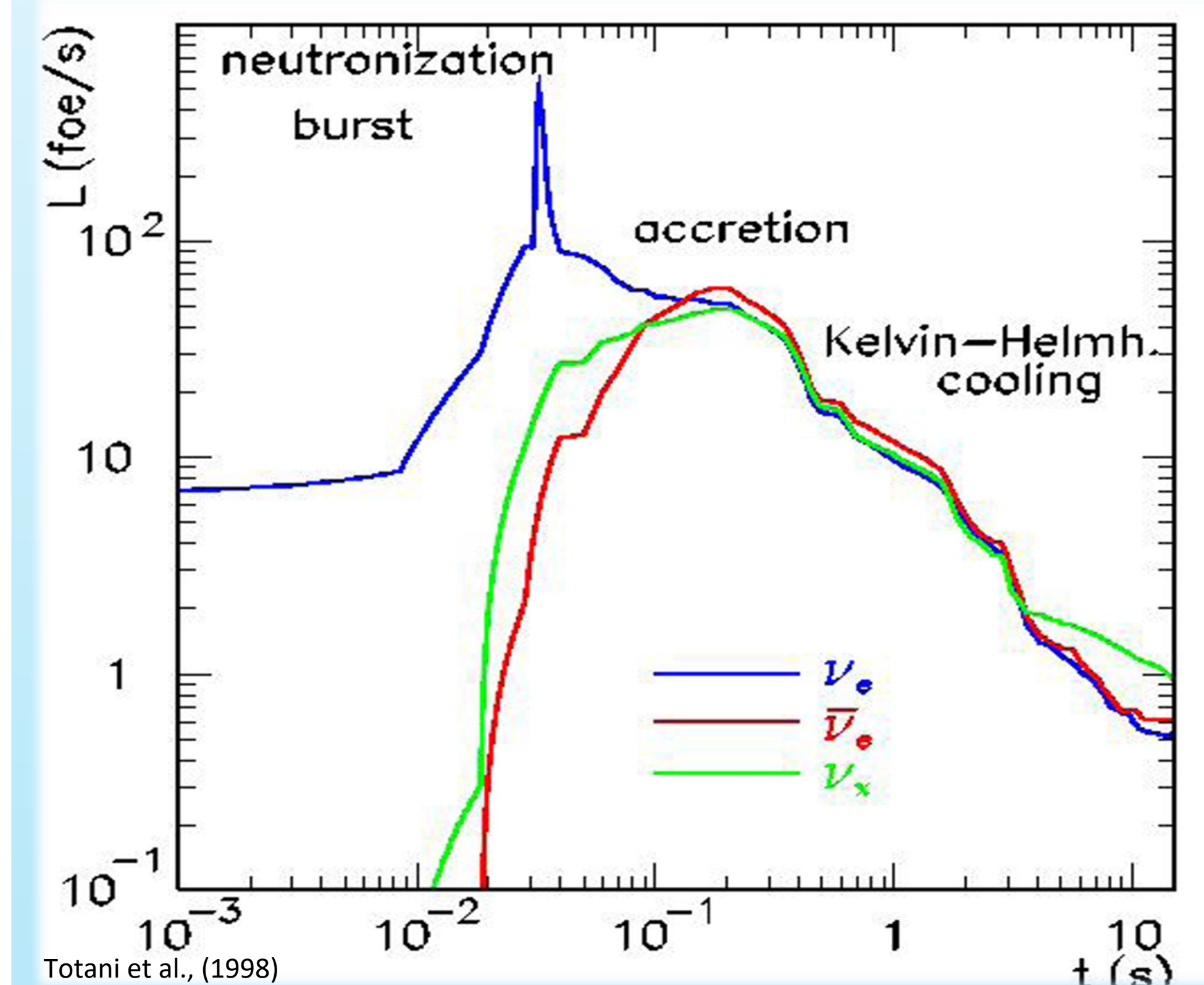
1 Background



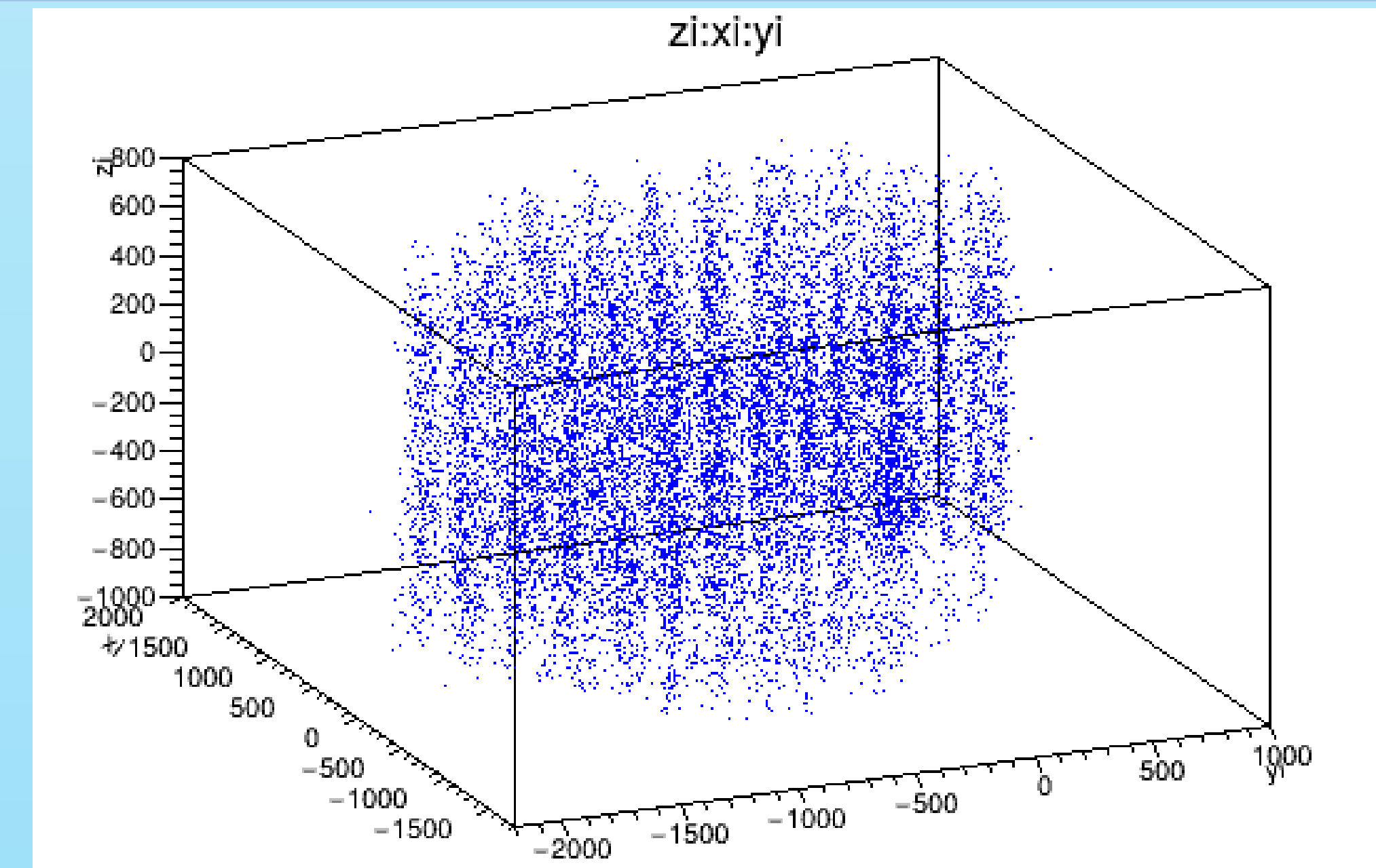
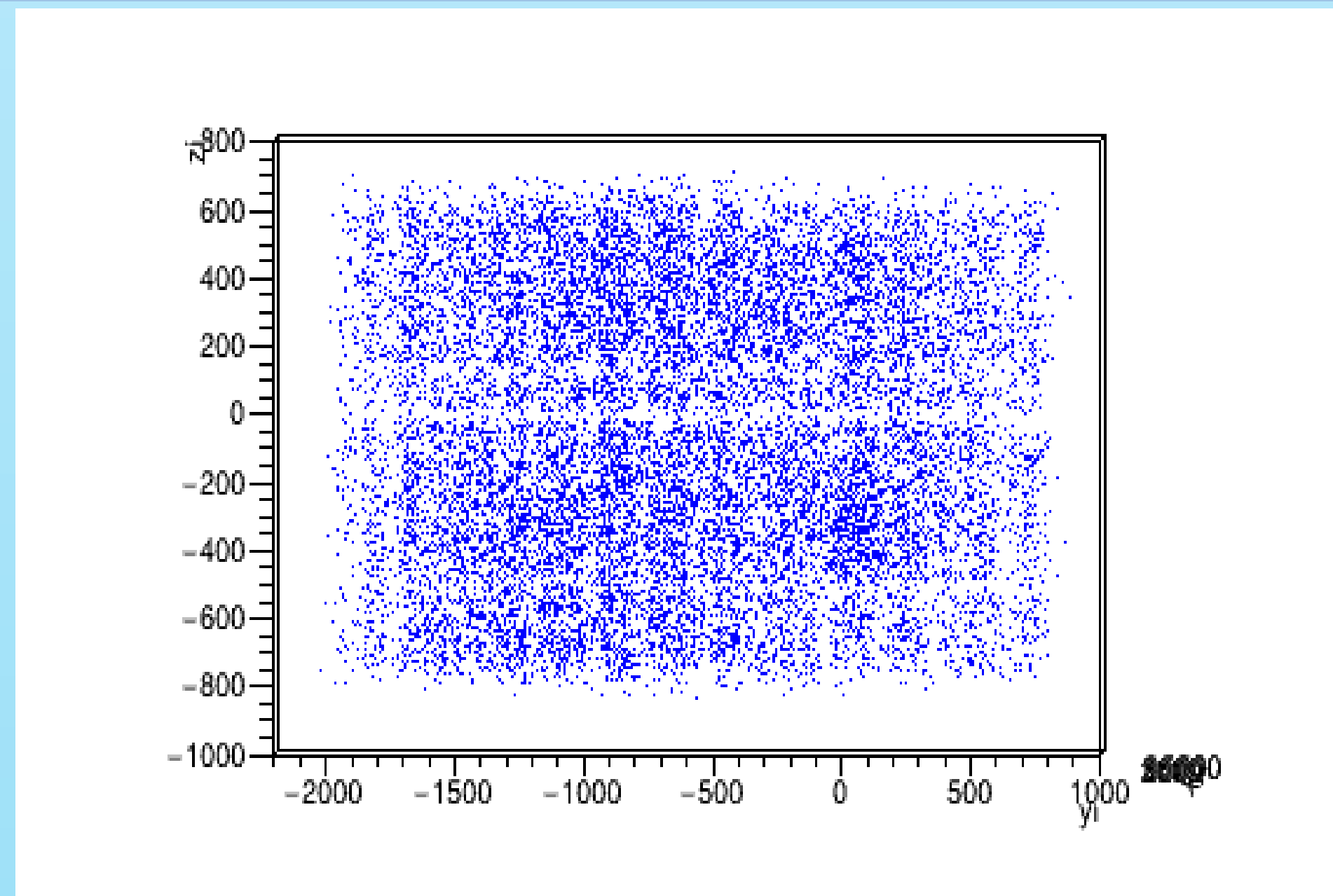
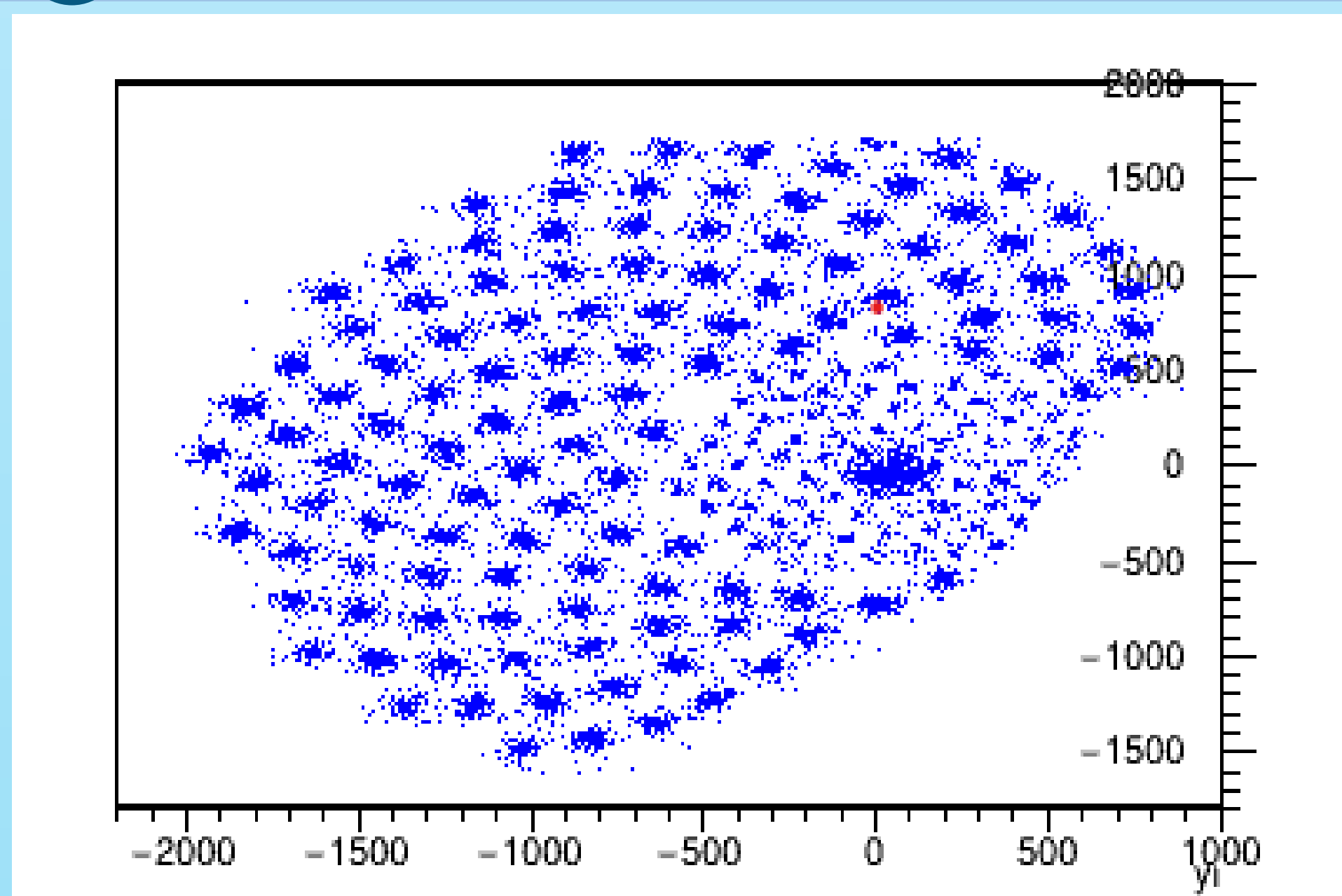
The **IceCube Upgrade** has recently deployed 5 additional, denser strings across the DeepCore region. Each decorated with **~700 new DOMs**, the **D-Egg** and **mDOM**.

D-Egg: 2 vertical opposing PMTs (8 in.)
mDOM: 24 PMTs in nearly uniform angular coverage (3 in.)

Both anticipated to outperform the original pDOM, we're currently comparing the two by their respective Cherenkov photon yield against expected positron-interaction behavior.



2 Simulated Hit Plots

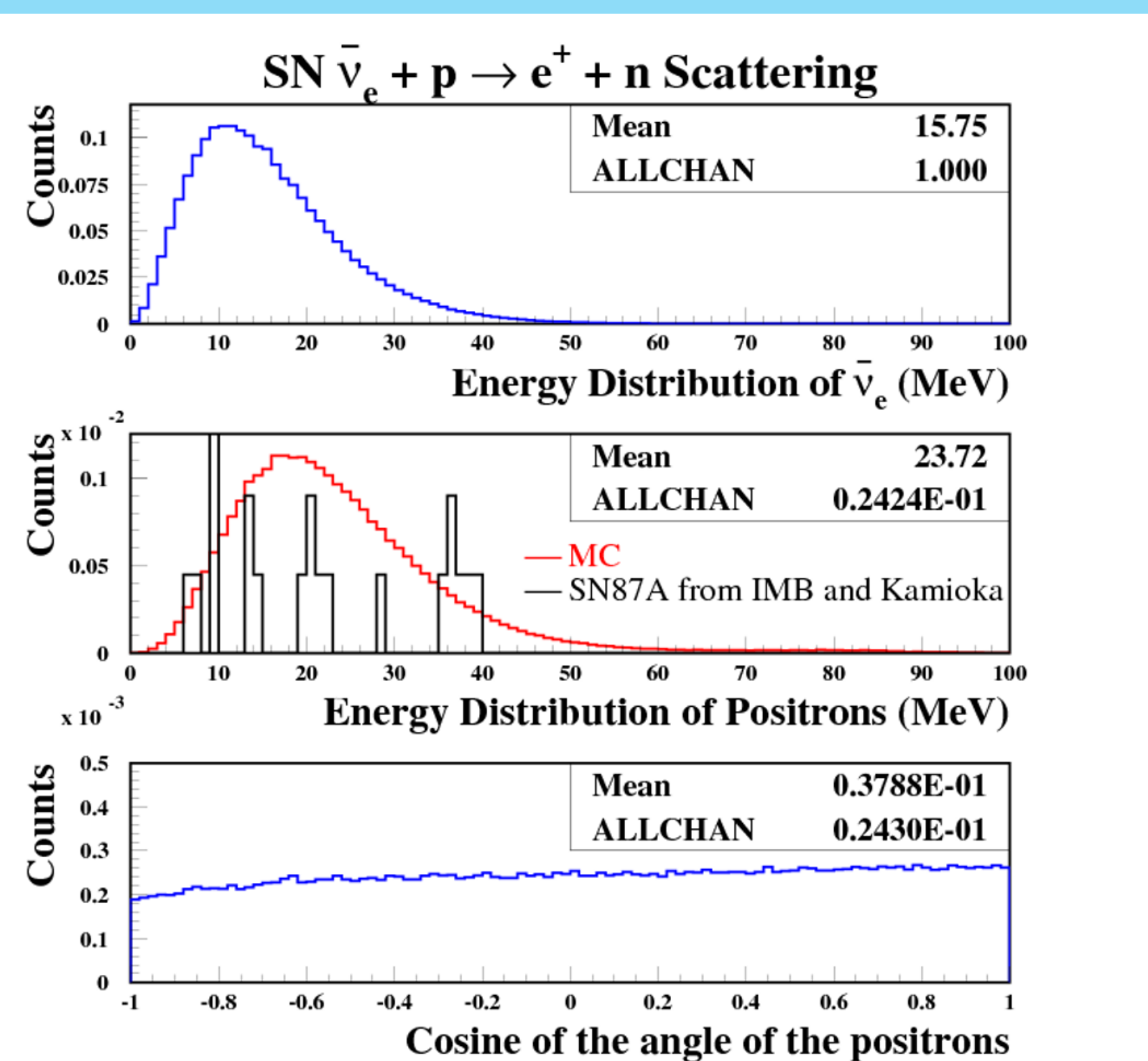


5 Results/ More Plots

3 Methodology

- $n(E_\nu) = (1/F(\alpha)T^3)(E_\nu^3/\exp(E_\nu/T - \alpha) + 1)$
We assume $T = 7$ MeV and $\alpha = 0$.
- $F(\alpha)$ normalizes the spectrum to unit flux
- This yields an average neutrino energy of ~ 24 MeV.
- 99% of the energy produced in Super Nova is carried off by the neutrinos.

4 Motivations



Simulated supernova interaction products ($\bar{\nu}_e$ energy, positron energy, positron angle)

Benchmarked against real SN1987A detections from IMB and Kamiokande

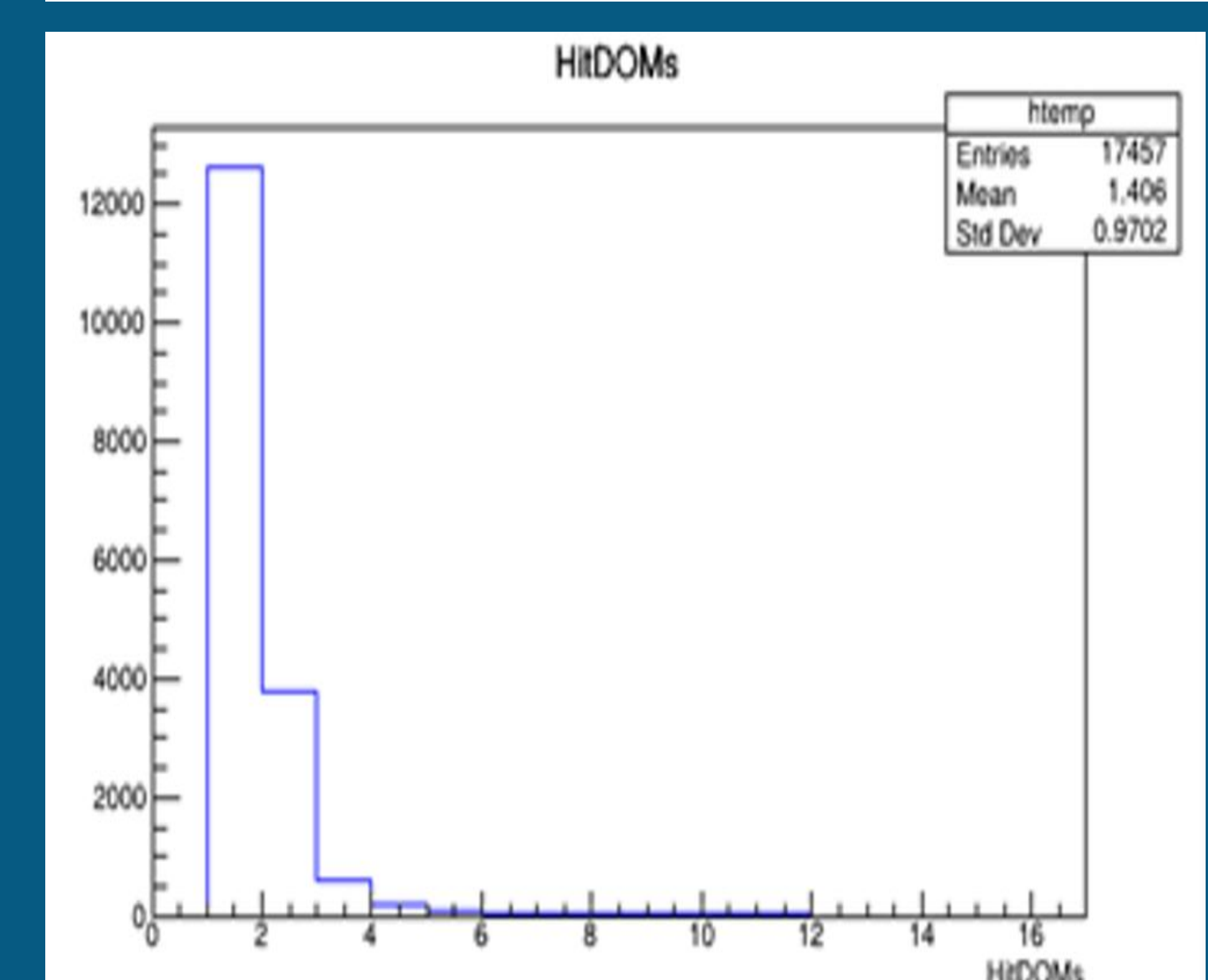
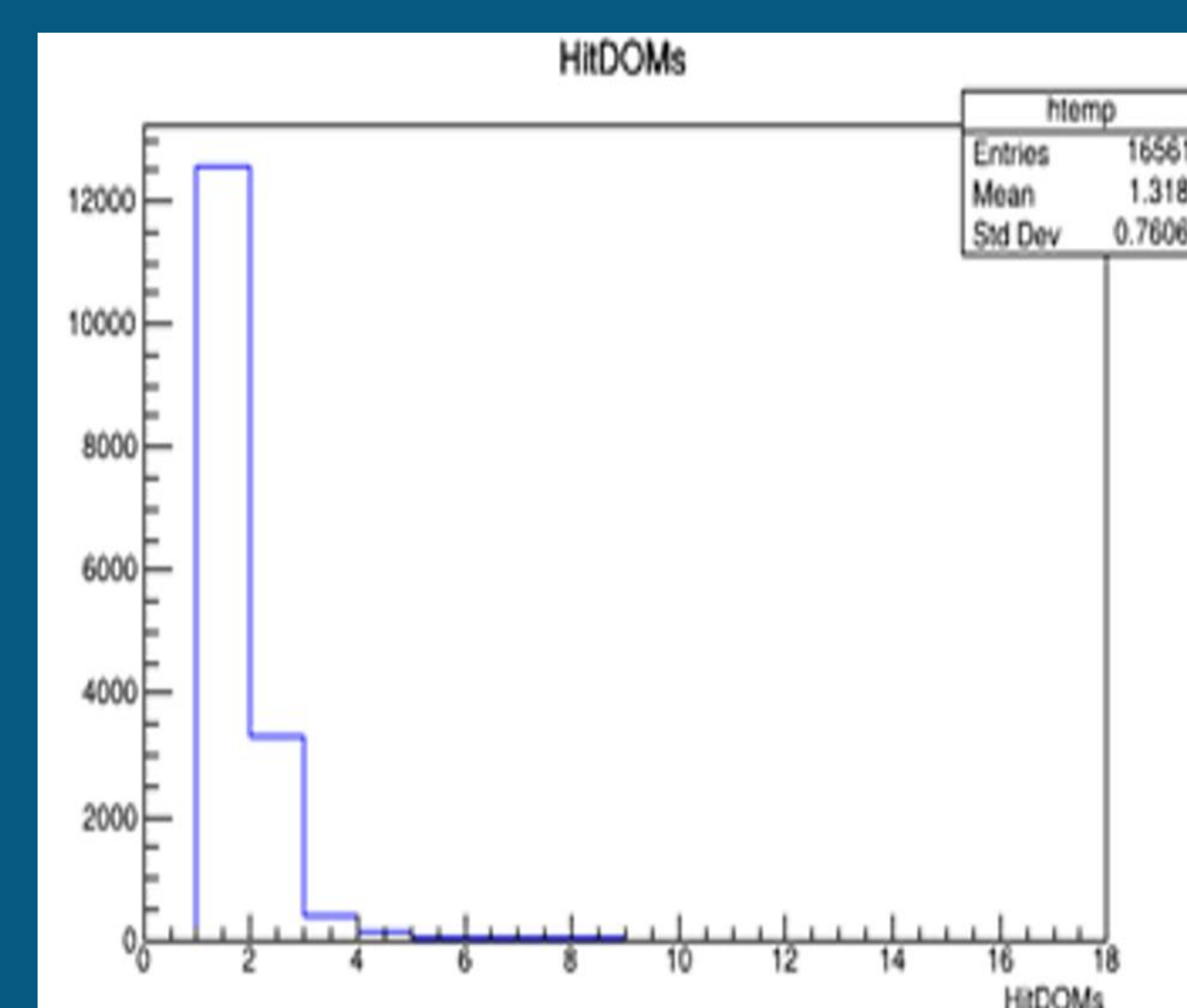
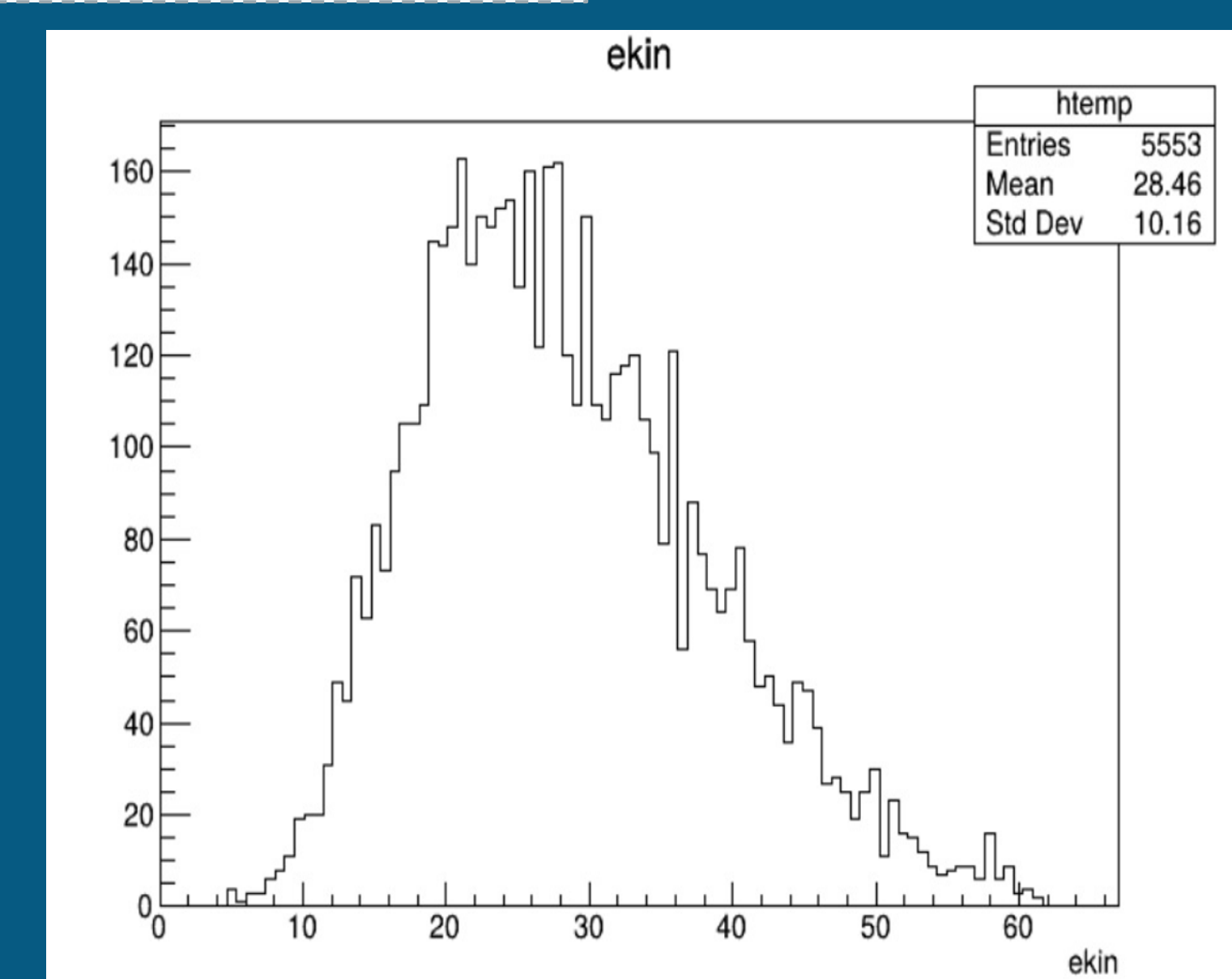
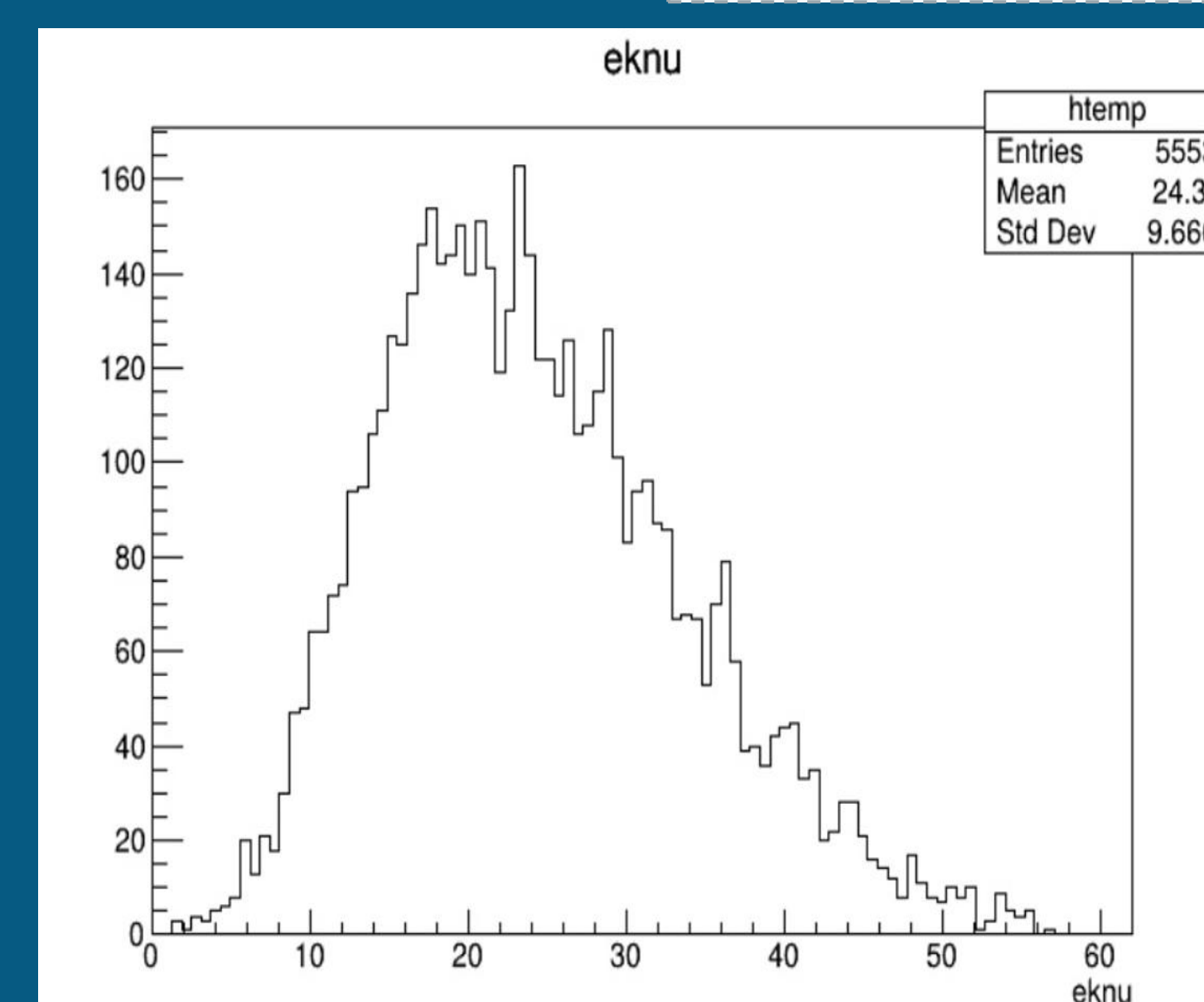
6 Summary/ Next Steps

- Detector geometry initializes correctly across the IceCube+DC, Upgrade, and Gen2 configurations.
- Simulated hit distributions to reproduce the expected detector profile, and the D-Egg configuration shows a higher two-module coincidence rate than mDOM (**28% vs. 24%**)
- Complete QE benchmarking against Collaboration and extend the comparison to angular acceptance and timing response in future.**

Acknowledgments

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SN Electron Antineutrino Spectrum (MeV)



2DOMs/1DOM = 24% , mDOM

2DOMs/1DOM = 28% , DEgg