

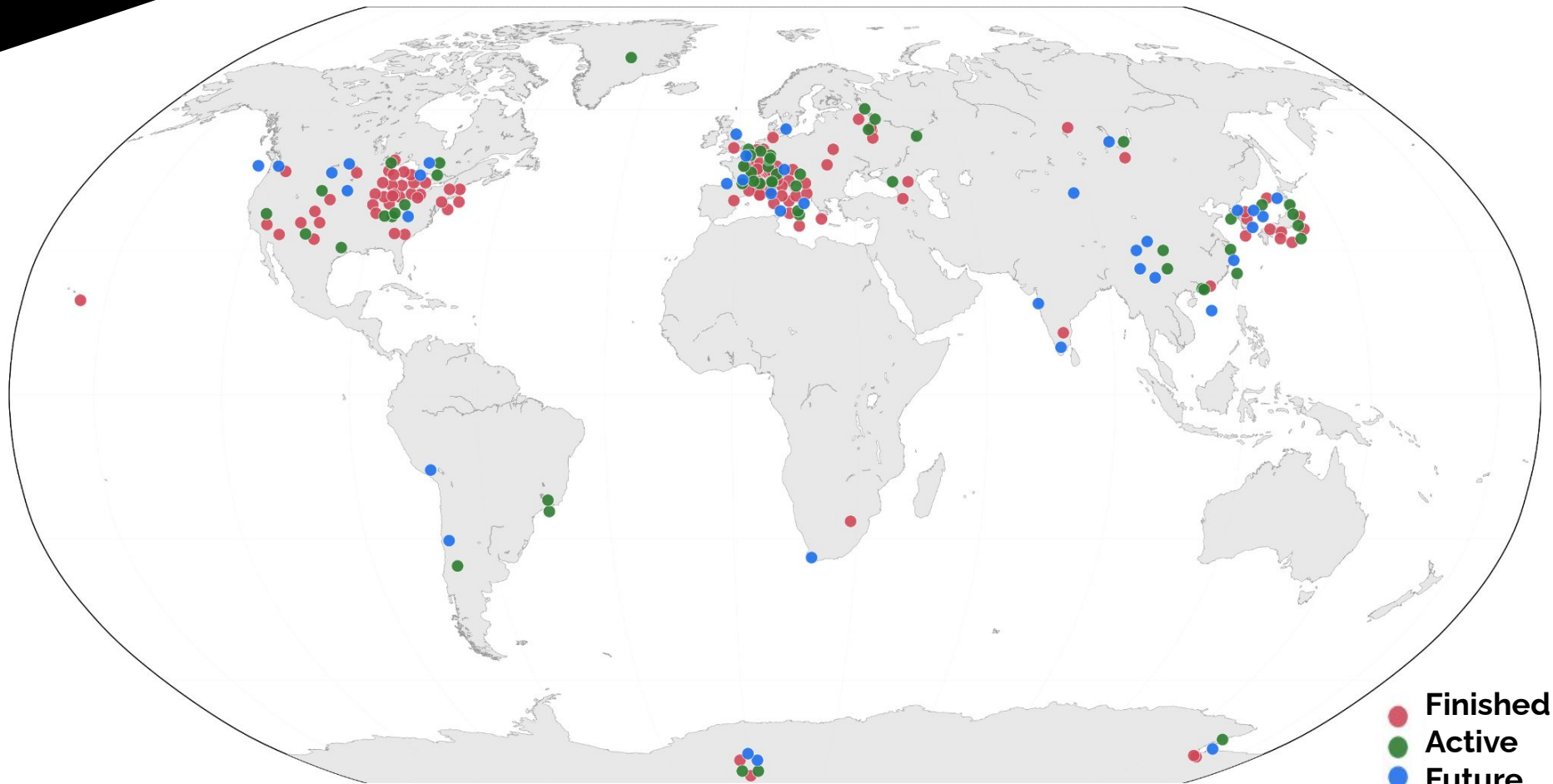
# Recent Highlights in Neutrino Physics

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**Mark Ross-Lonergan**

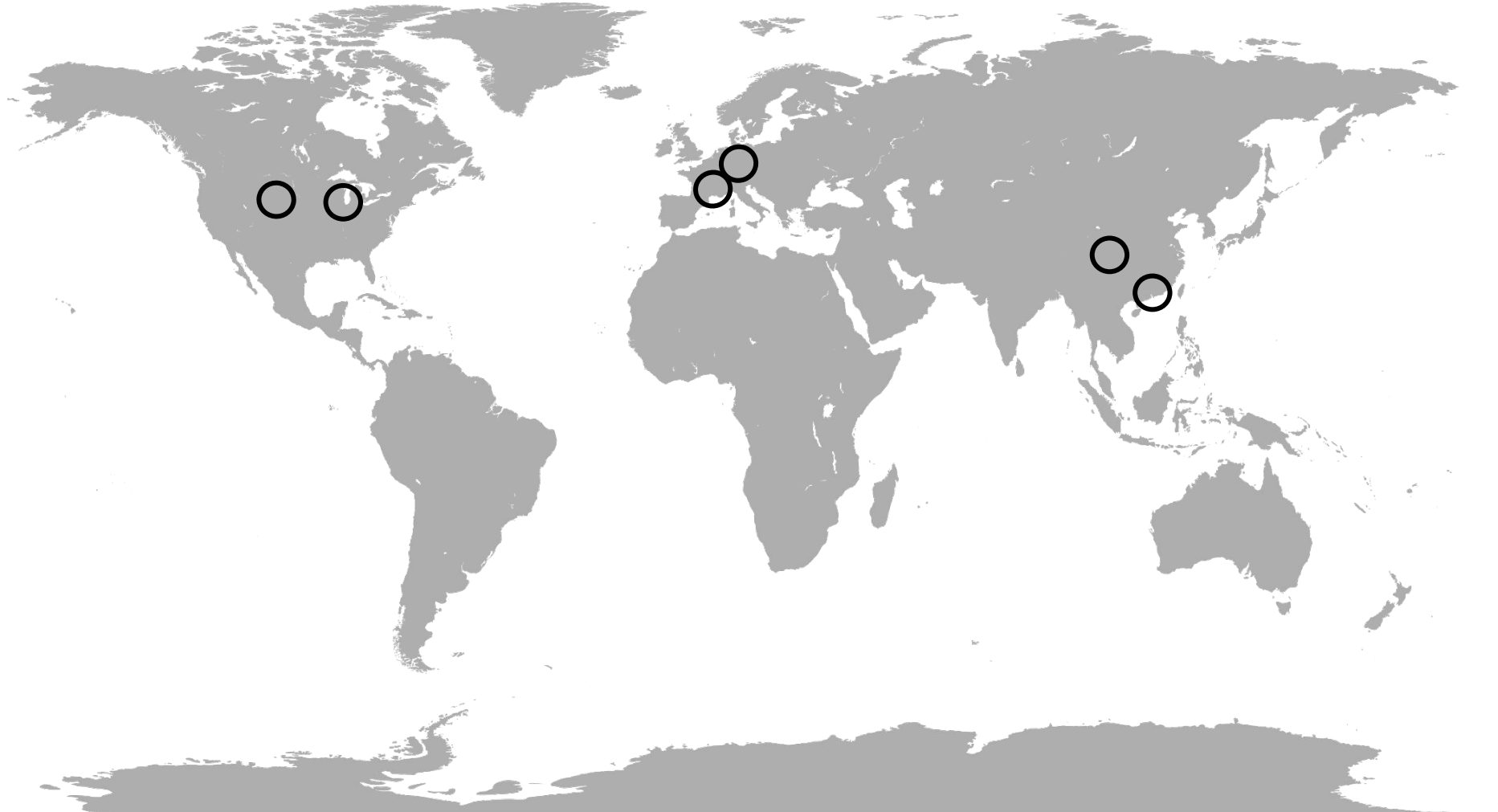
*PHENO 2026 - University of Pittsburgh*  
May 13<sup>th</sup> 2026

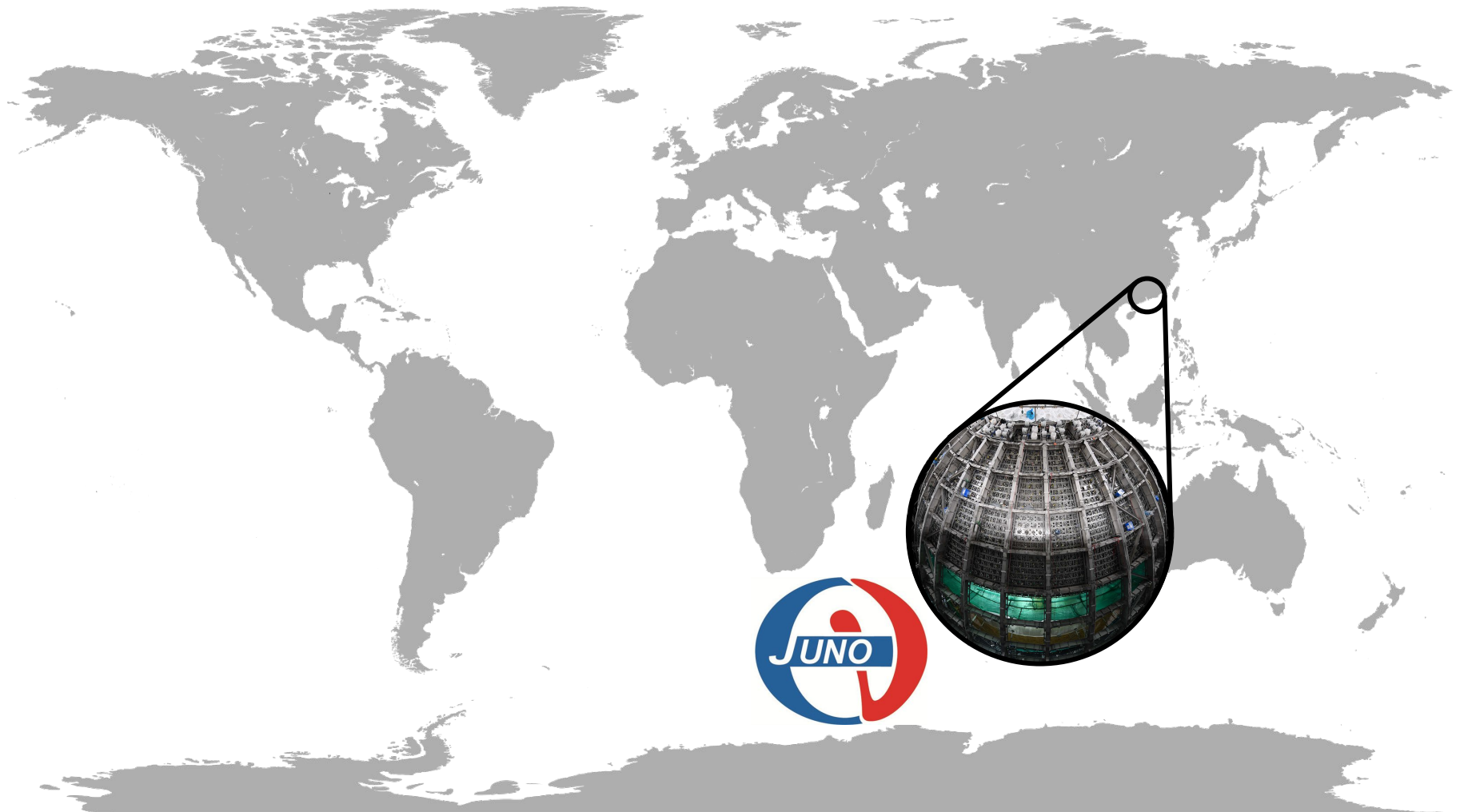
# Neutrino Experiments of the World > 200



- Finished
- Active
- Future

Focus on last ~6 months

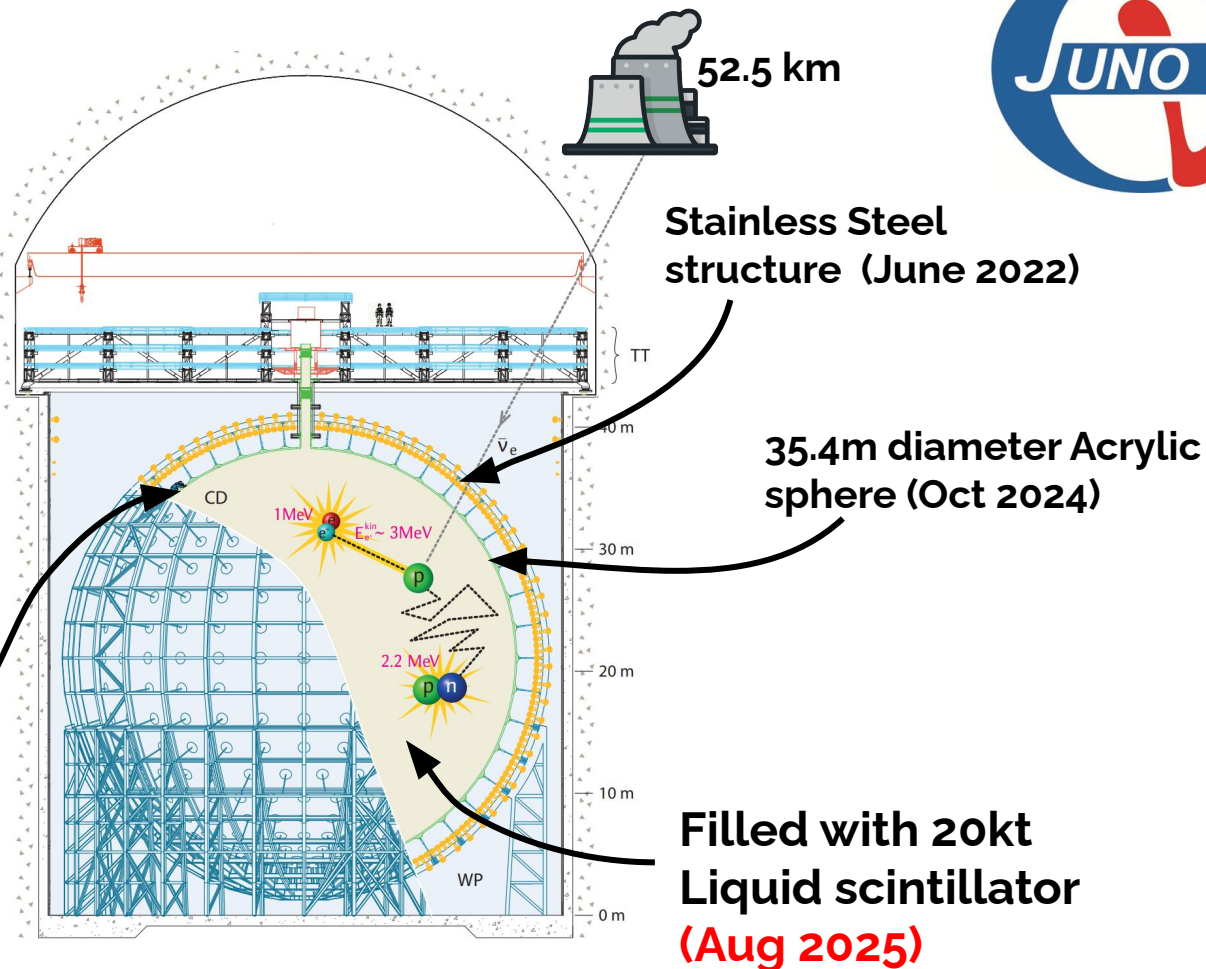




# JUNO: Jiangmen Underground Neutrino Observatory



20,012 20" PMTs  
25,600 3" PMTs  
(Dec 2024)  
75% photocathode coverage



Stainless Steel structure (June 2022)

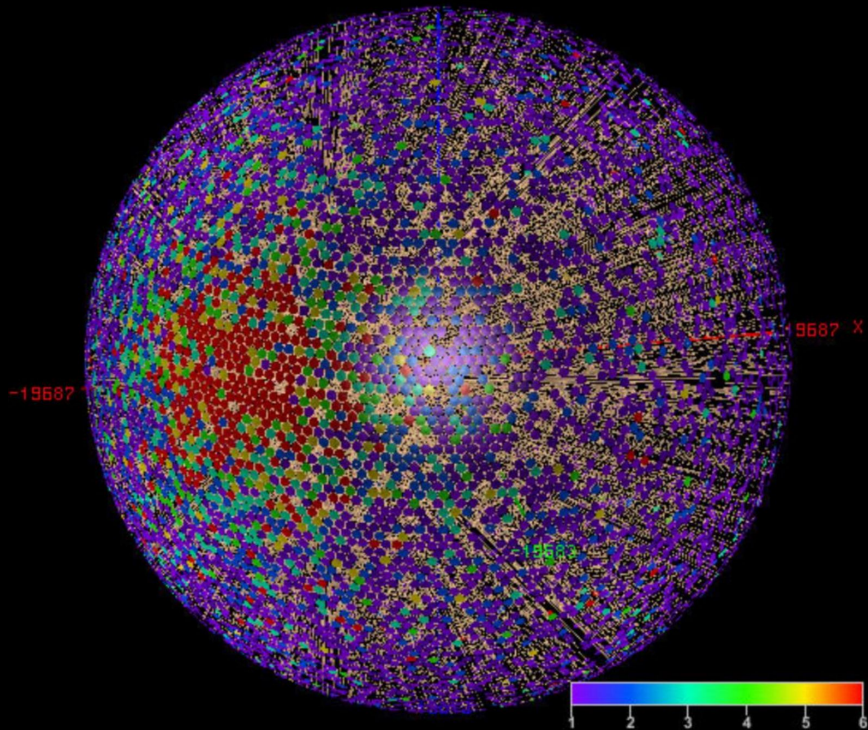
35.4m diameter Acrylic sphere (Oct 2024)

Filled with 20kt Liquid scintillator  
(Aug 2025)

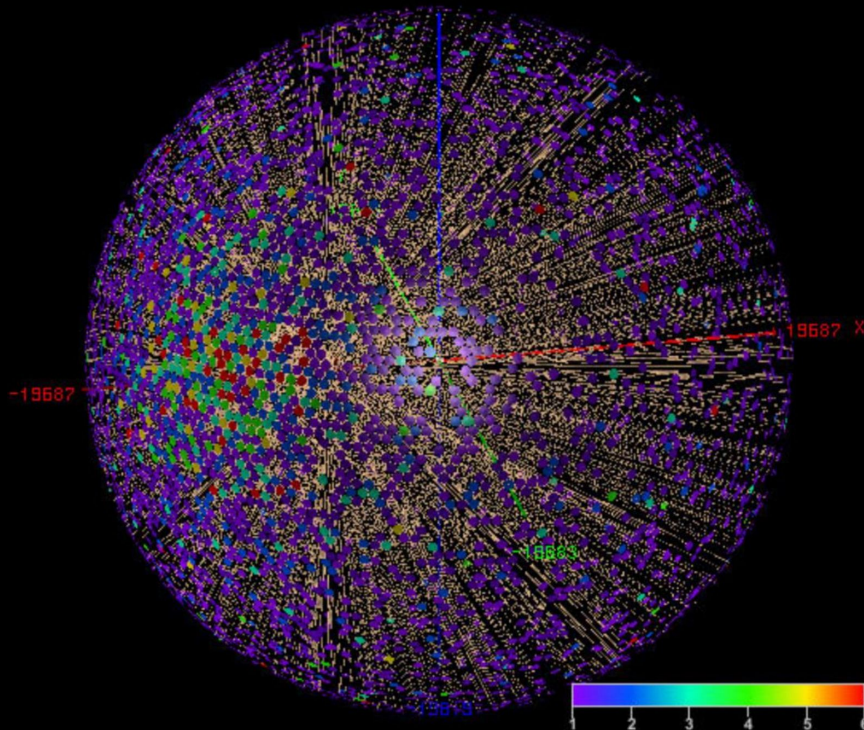
# JUNO : Reactor Inverse Beta Decay Event

Mon, 25 Aug 2025 22:50:45  
RecEnergy = 6.3 MeV

Mon, 25 Aug 2025 22:50:45  
RecEnergy = 2.4 MeV



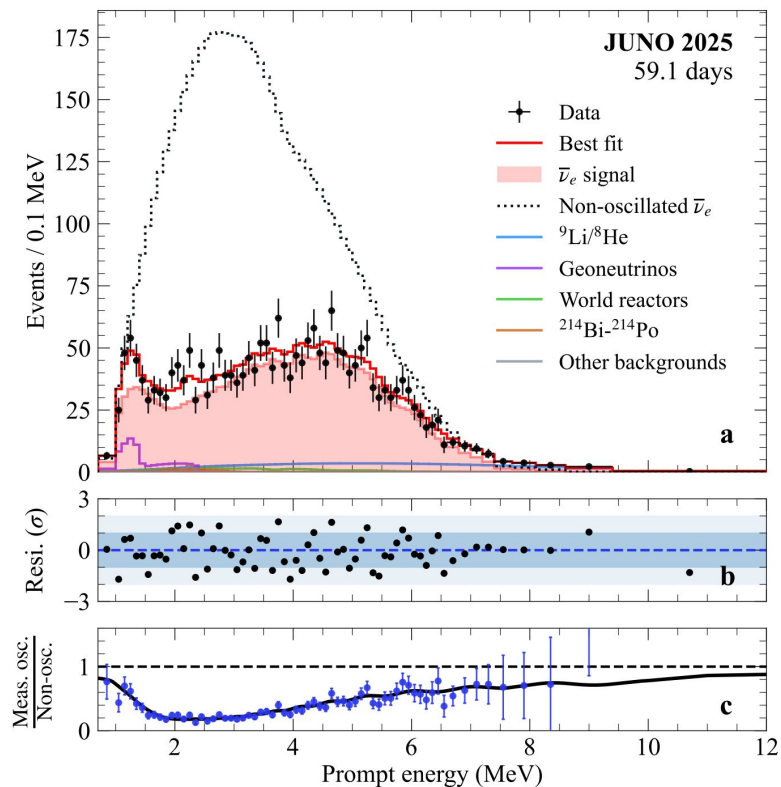
Prompt  $e^+$  signal



Delay neutron signal



- First results released with just **59.1 days of data collected**



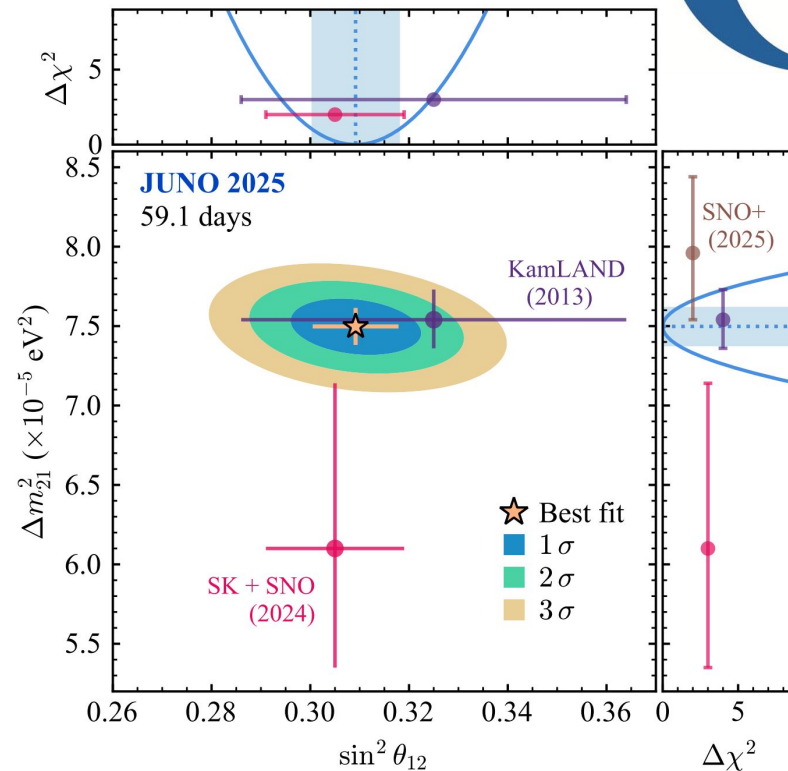
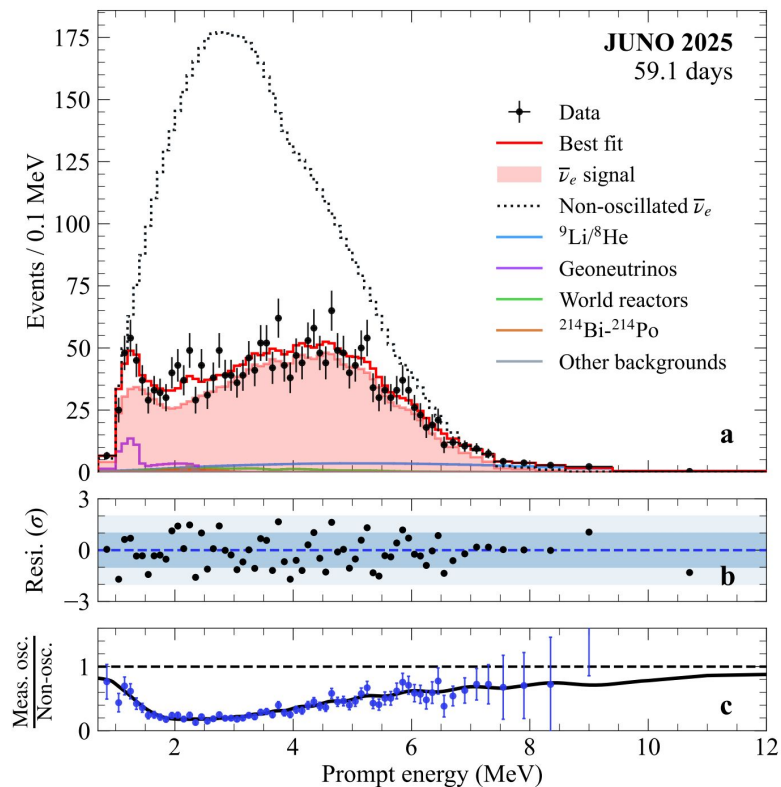
- **2,379  $\bar{\nu}_e$  IBD candidates**
- **69.9% Signal Efficiency**

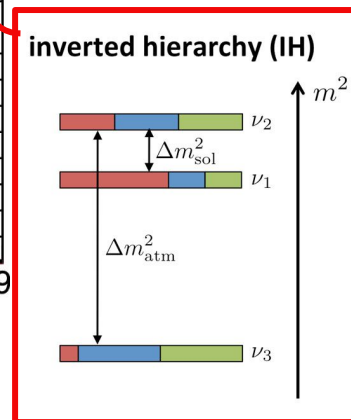
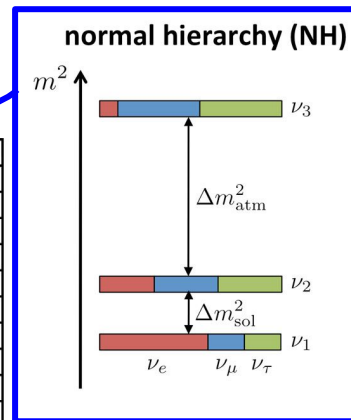
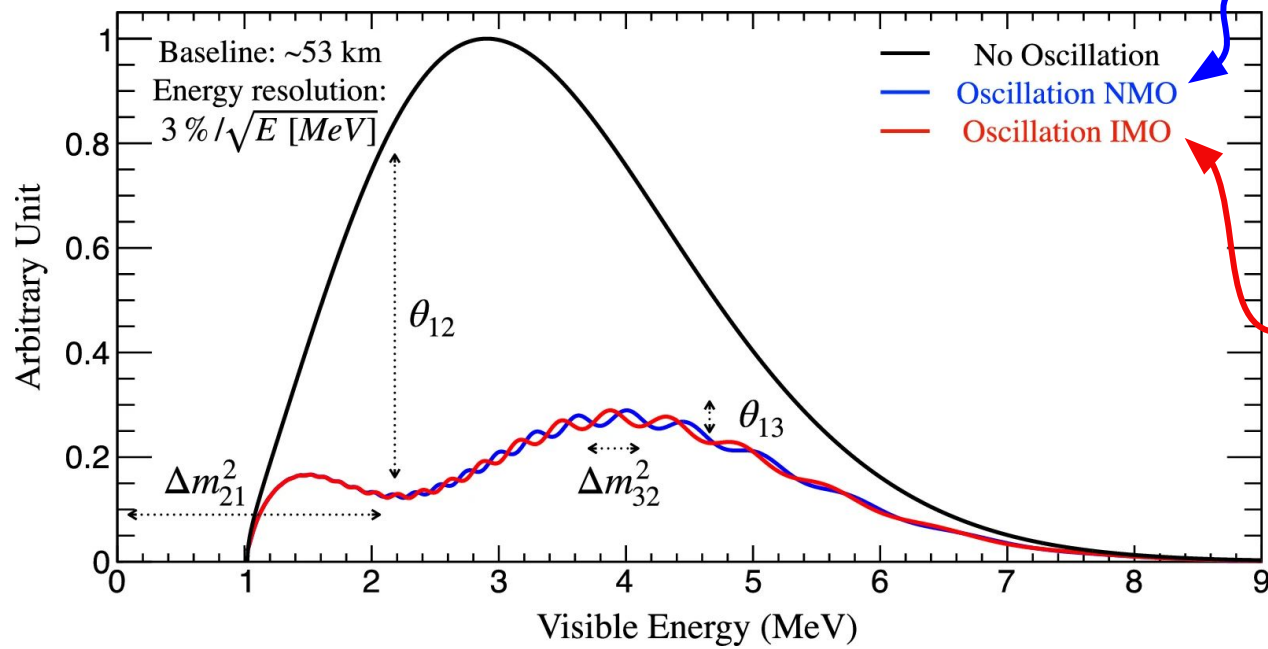
Largest **backgrounds** being

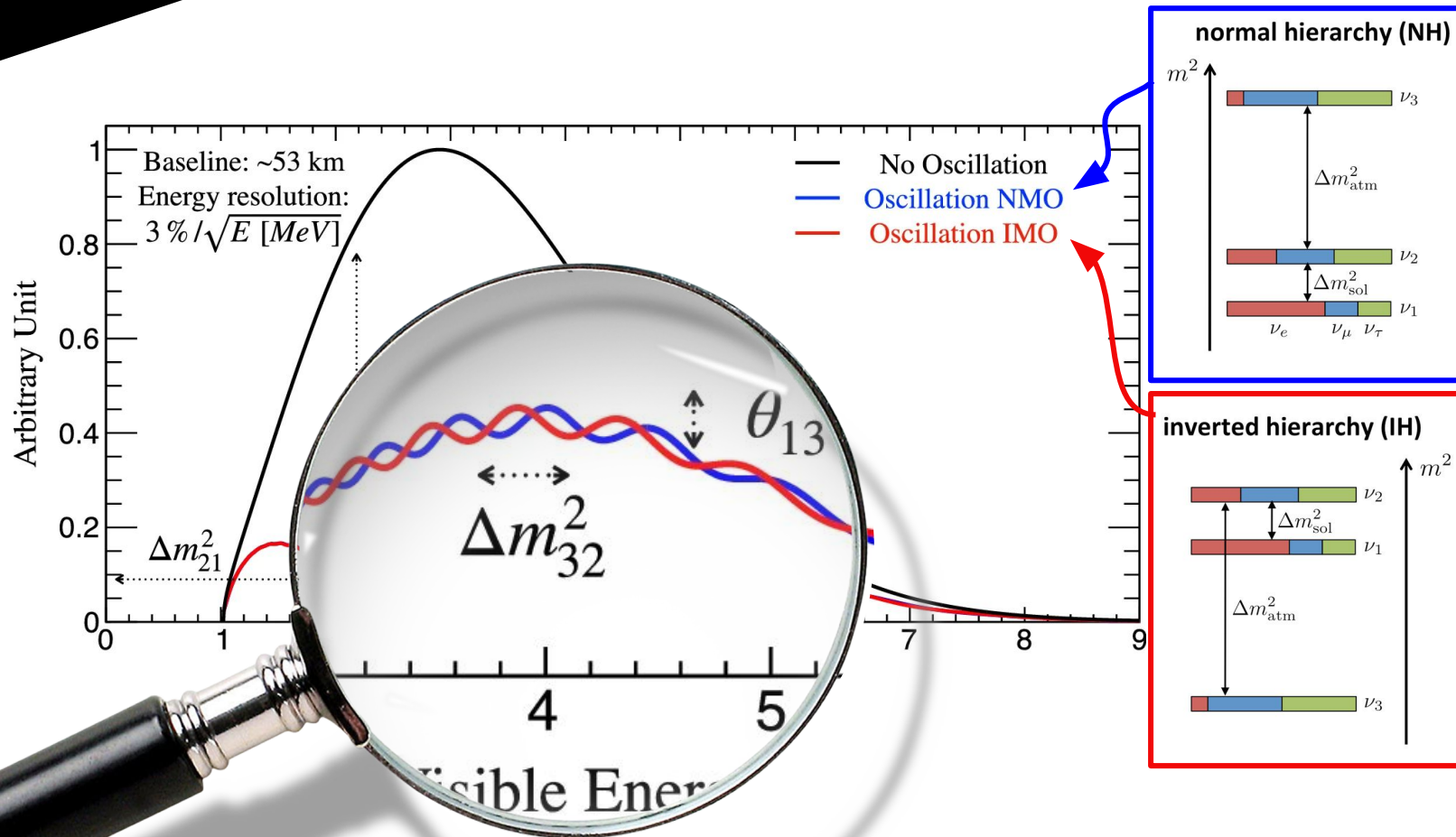
- **${}^9\text{Li}/{}^8\text{He}$**  (~4.3 events per day)
- **Geoneutrinos** (~1.2 events per day)
- **World Reactors** (<1 event per day)



- First results released with just **59.1 days** of data collected

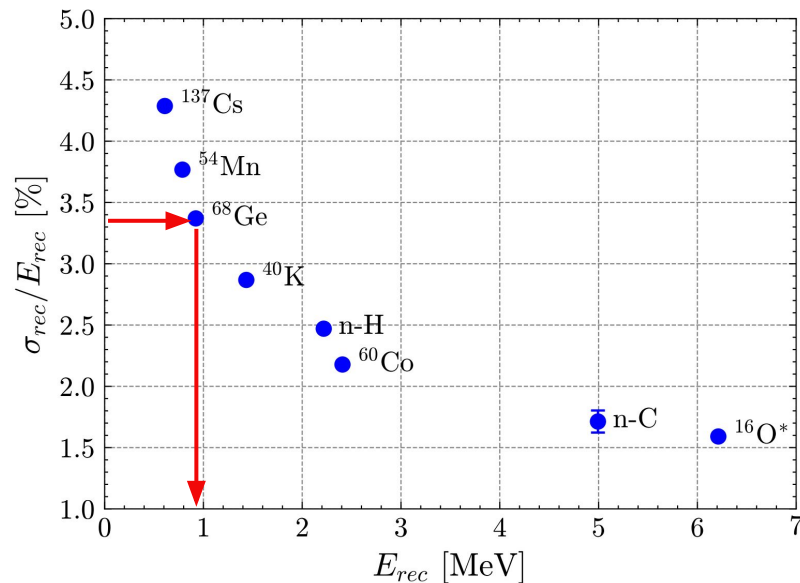








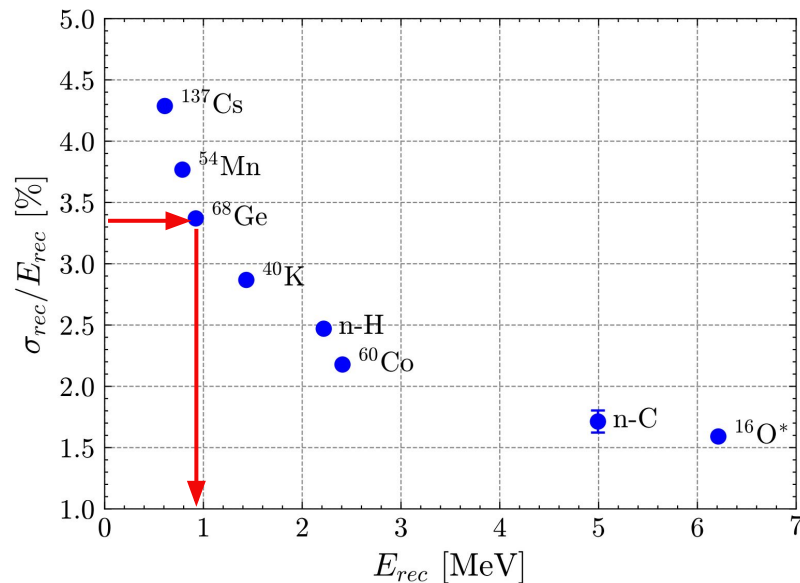
Energy Resolution **design goal** ( $\sim 3\%$  @ 1 MeV)



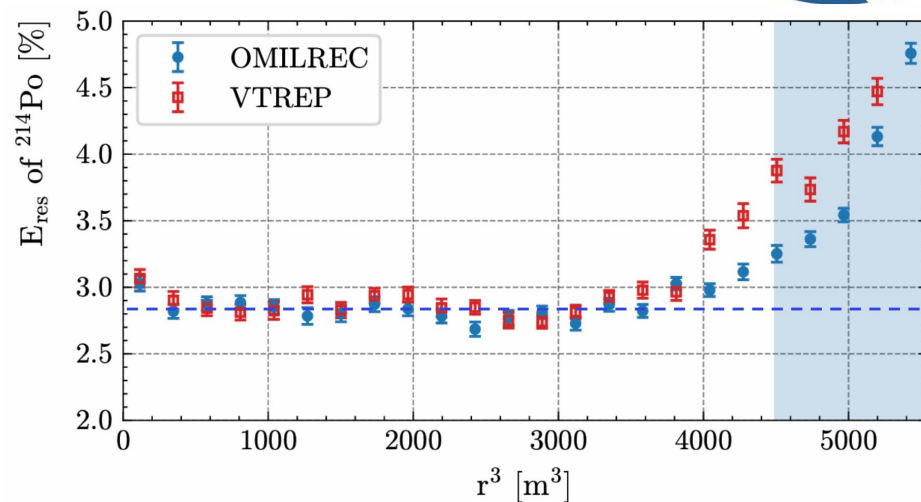
Calibration  $\gamma$  source efficiency at  
**center of detector**



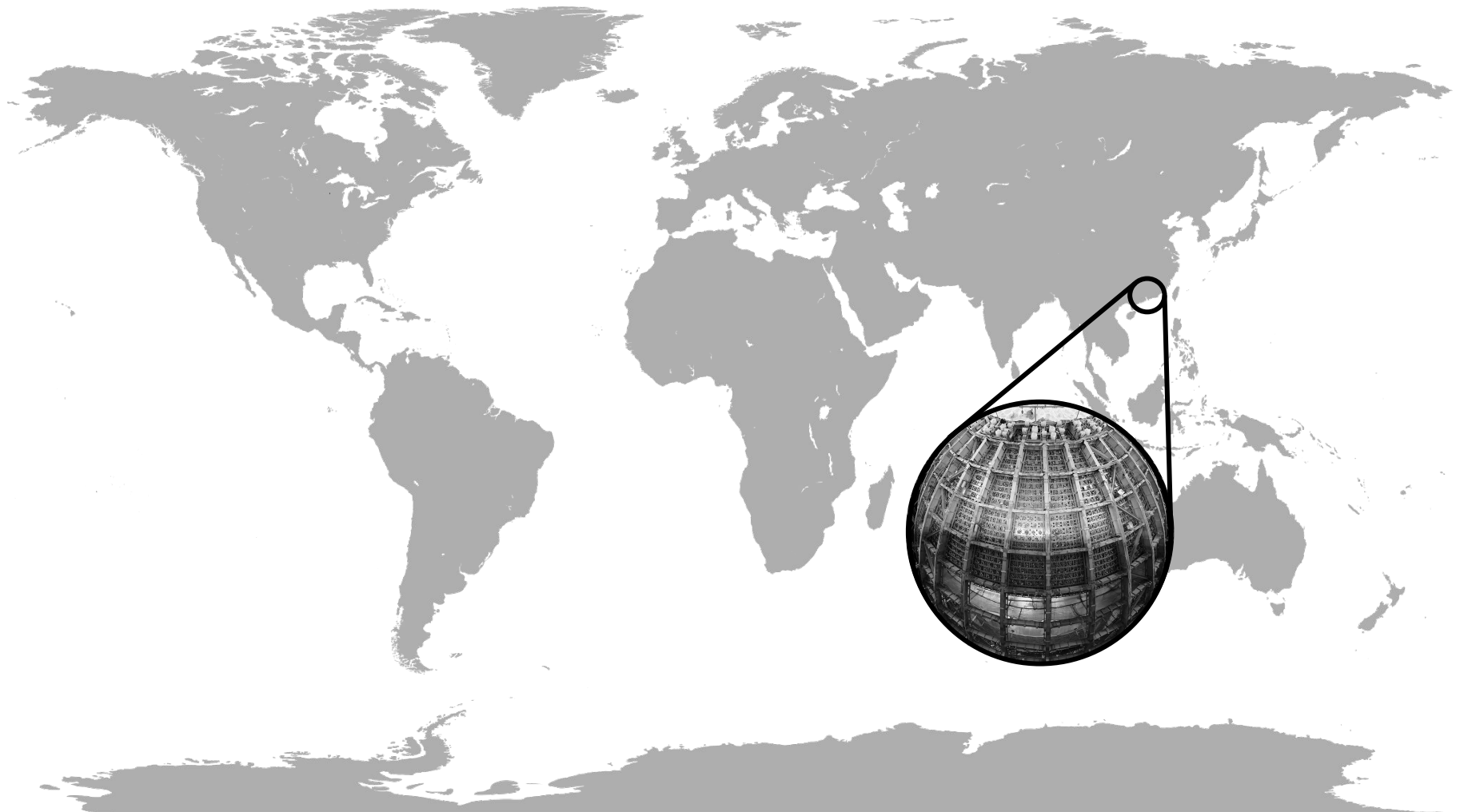
Energy Resolution **design goal** ( $\sim 3\%$  @ 1 MeV)

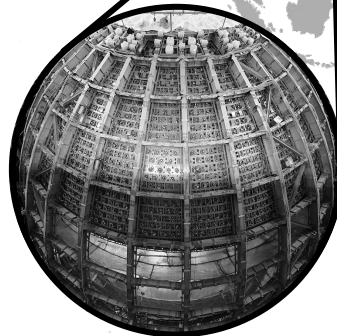
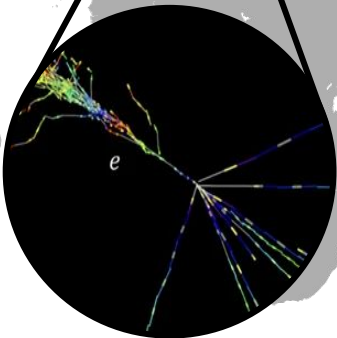


Calibration  $\gamma$  source efficiency at **center of detector**



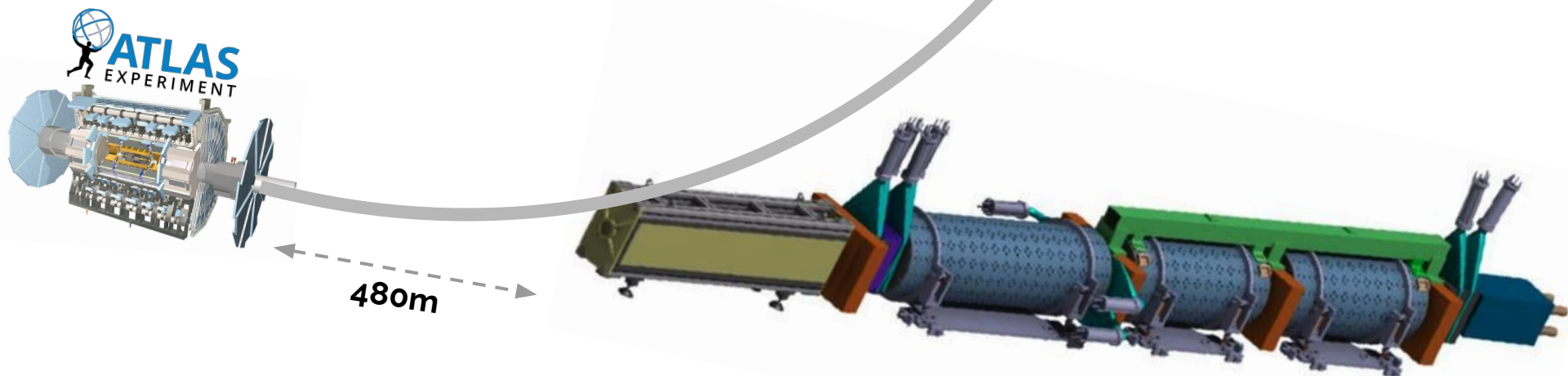
**2.9 % Energy Resolution** for a 0.94 MeV  $^{214}\text{Po}$   $\alpha$  decays from natural  $^{238}\text{U}$  decay chain



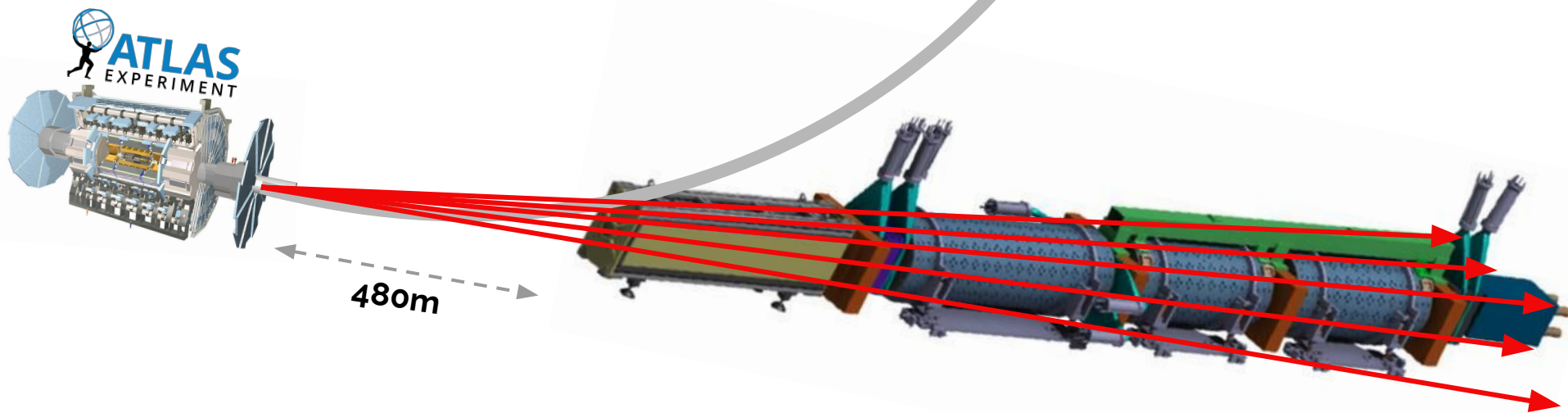


**FASER**

- **Collider Neutrinos** a rapidly growing area
  - **100 GeV** → **~few TeV** previously unexplored

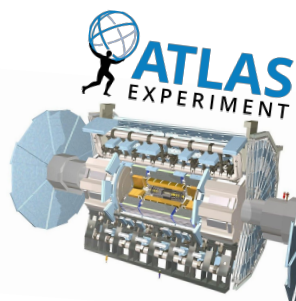


- **Collider Neutrinos** a rapidly growing area
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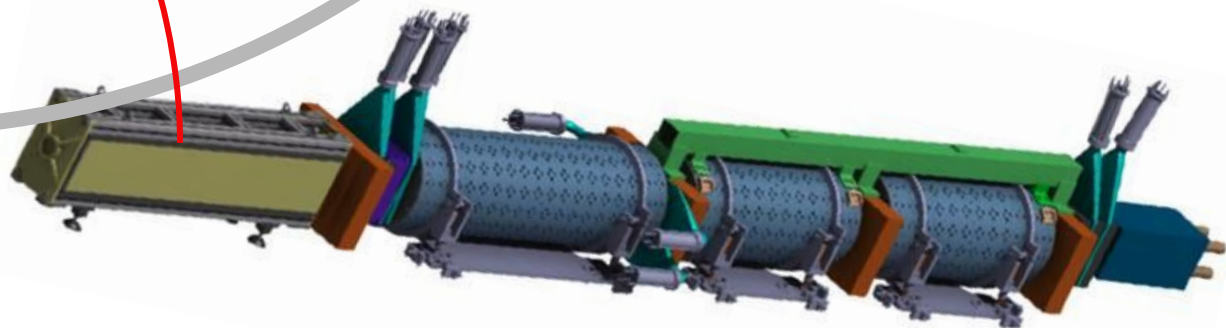
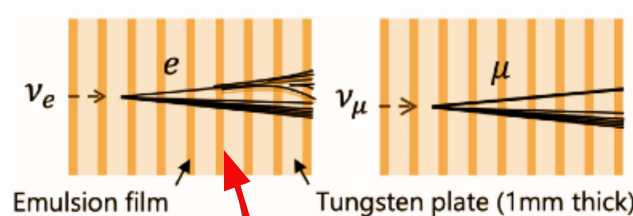


## Emulsion detector:

- Sensitive to **all neutrino flavours**
- **Excellent spatial resolution** ( $0.3\mu\text{m}$ )
- **Slow** ( film must be scanned, digitised, and processed)

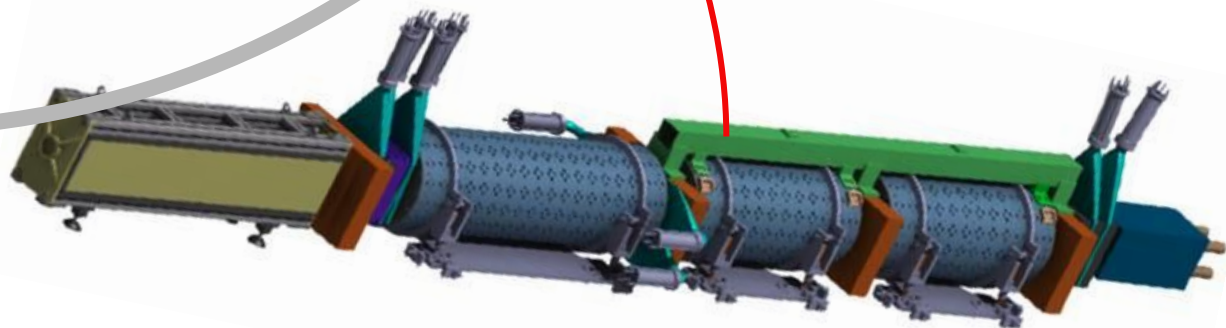
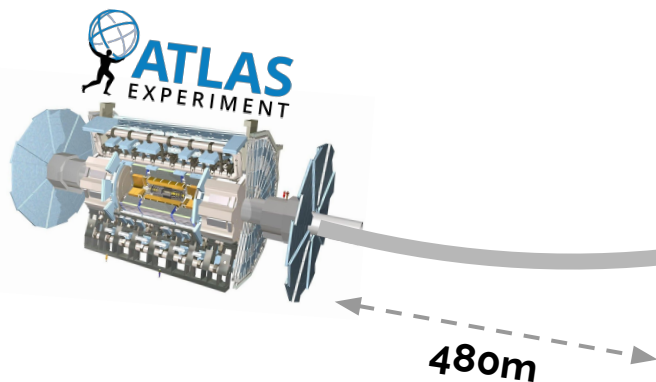


480m



## Electronic spectrometer:

- **Fast** analysis (only using electronic components)
- **Neutrino/anti-neutrino separation**
- **EM Calorimeter**



# FASER : Emulsion Detector Events

FASER

Preliminary

$e$

1000  $\mu\text{m}$

FASER

Preliminary

$e$

100  $\mu\text{m}$

FASER

Preliminary

$\mu$

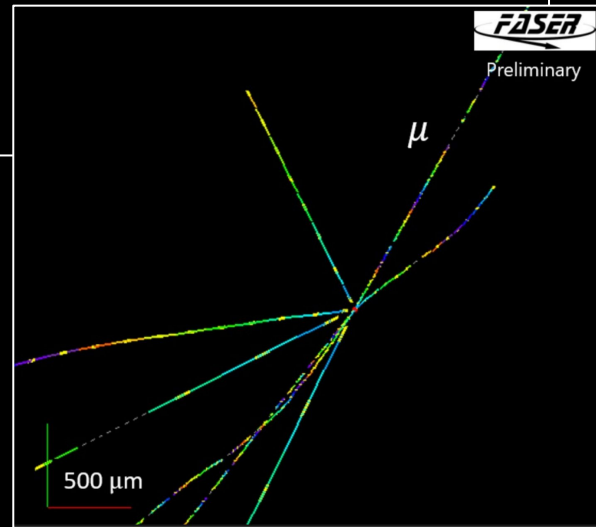
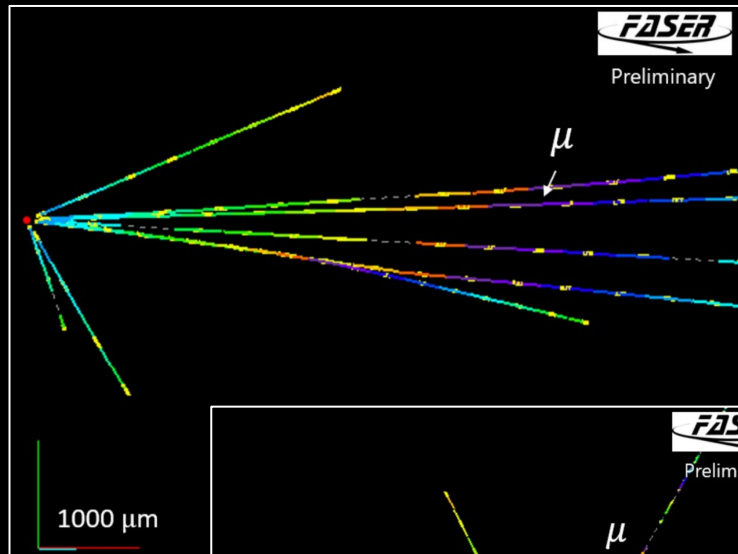
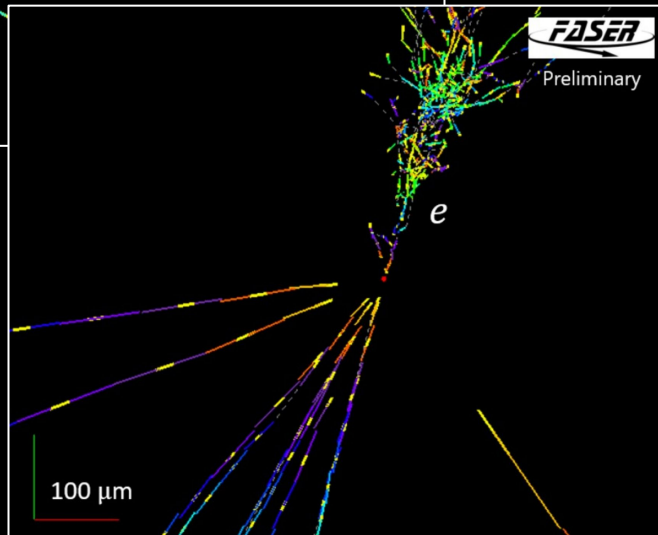
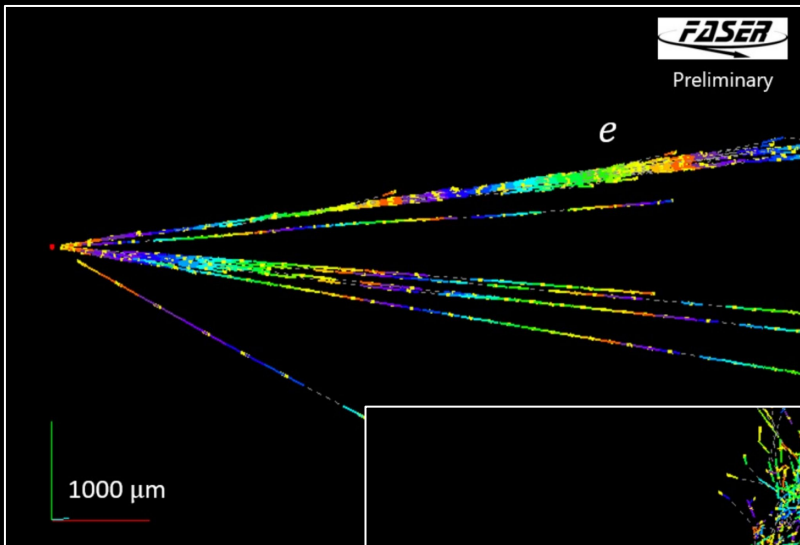
1000  $\mu\text{m}$

FASER

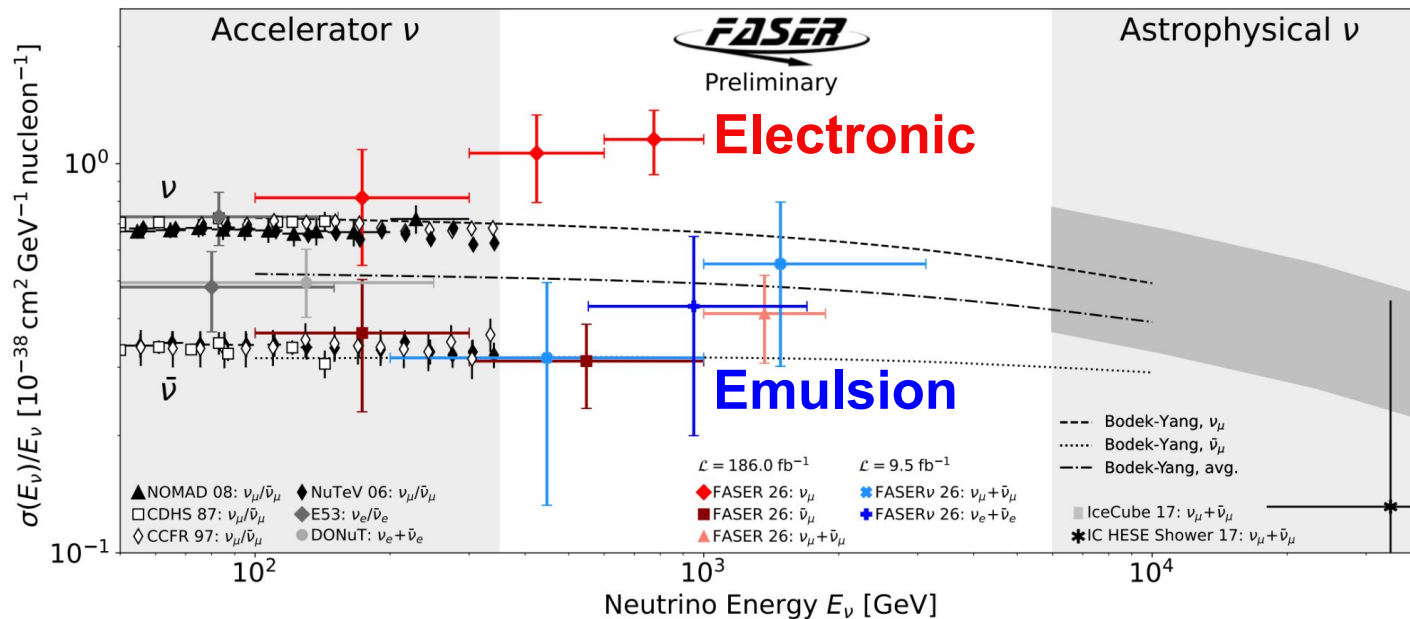
Preliminary

$\mu$

500  $\mu\text{m}$



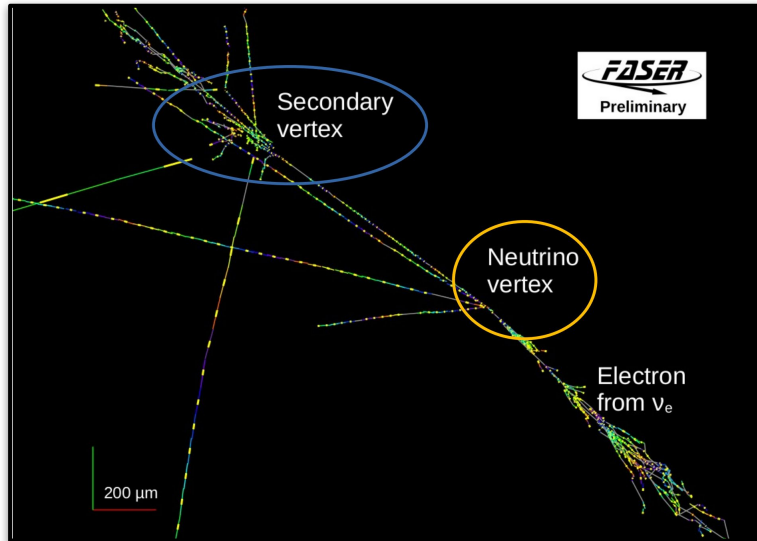
**~x2.8 increase in electronic data analyzed**



**~x5 increase in emulsion mass studied**

## First Search for **Neutrino-Induced Charm Production** CERN-FASER-CONF-2026-003

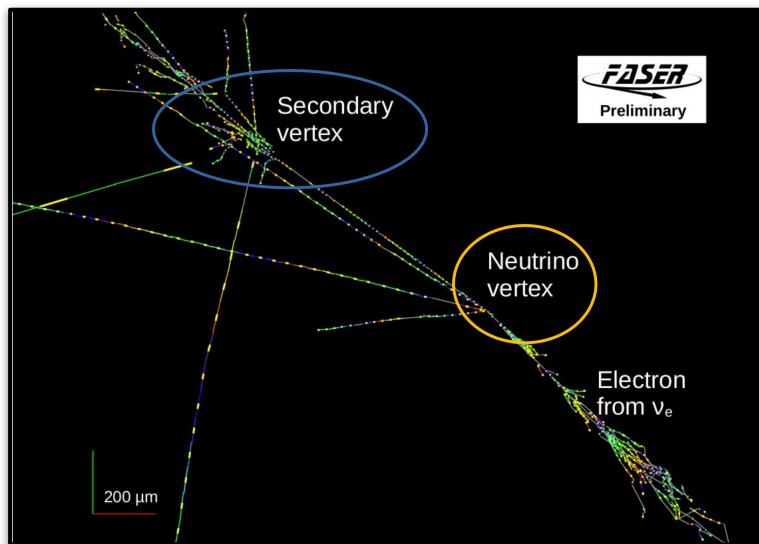
- Direct handle on charged charm decays
- Benchmark for  **$\tau$  reconstruction**



**Displaced secondary vertex**

## First Search for Neutrino-Induced Charm Production CERN-FASER-CONF-2026-003

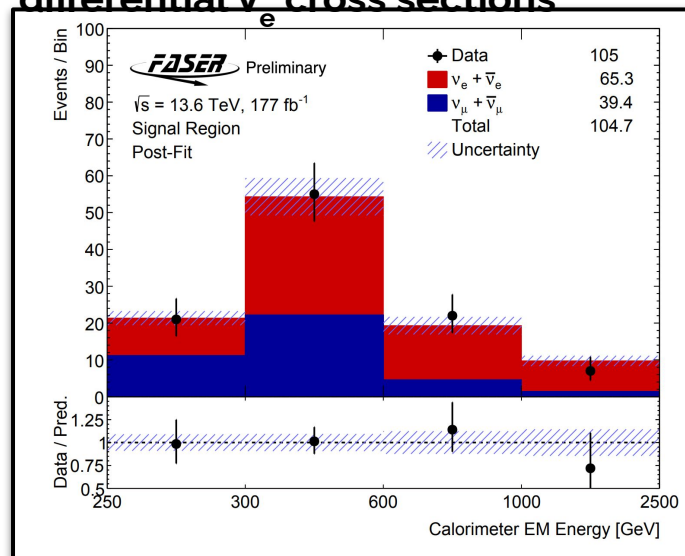
- Direct handle on charged charm decays
- Benchmark for  $\tau$  reconstruction

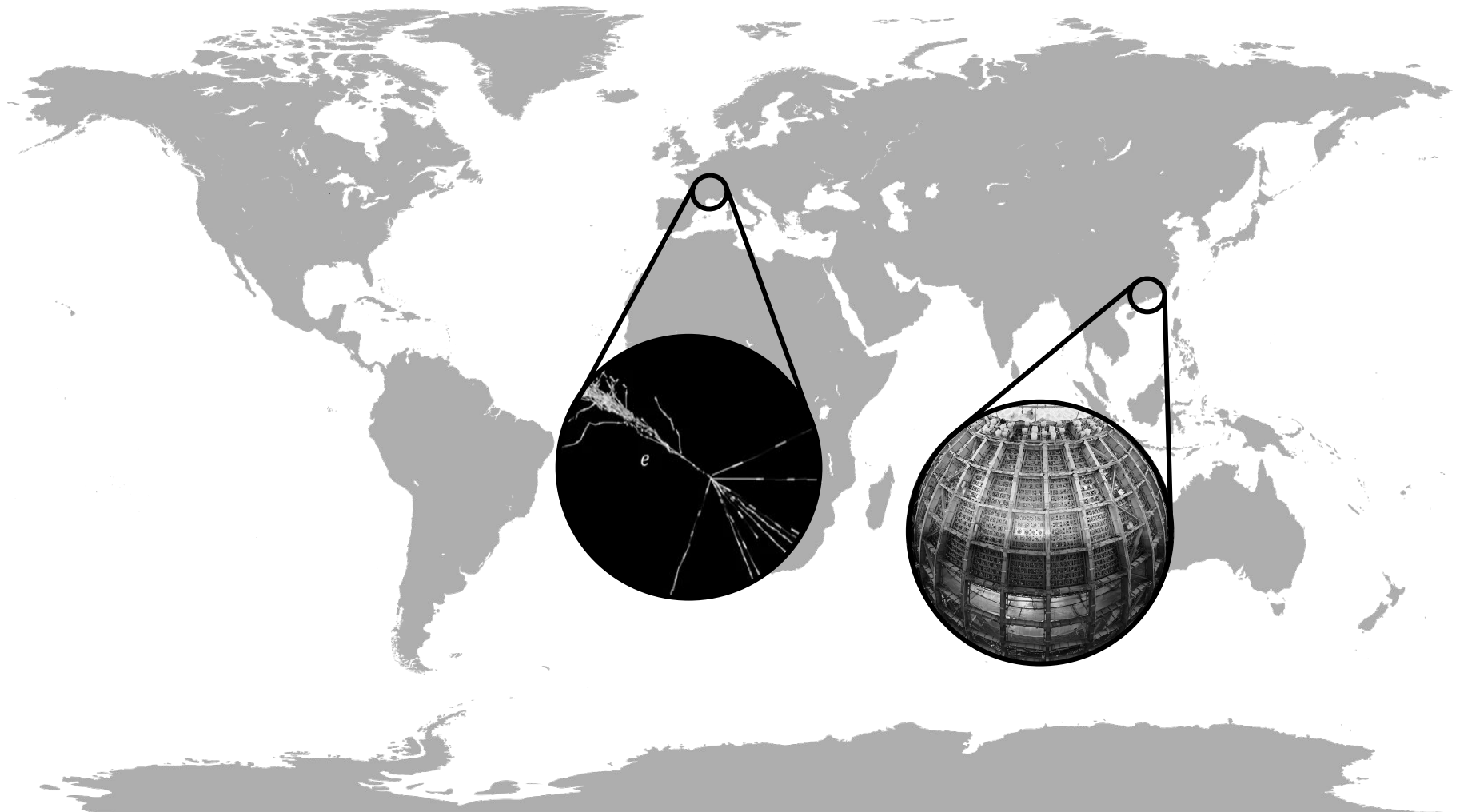


Displaced secondary vertex

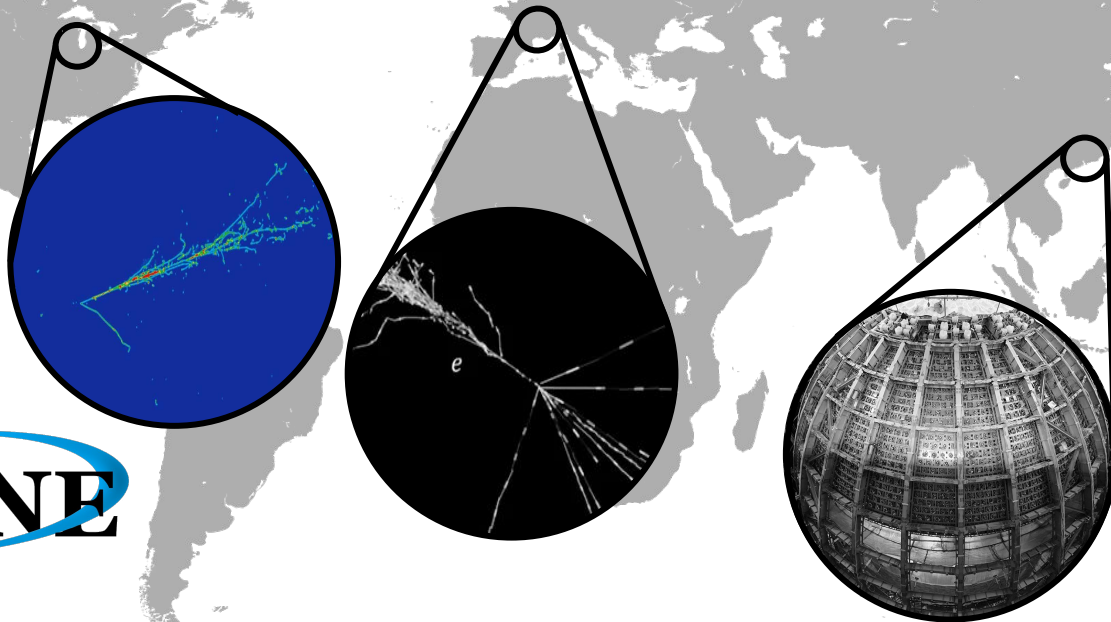
## First $\nu_e$ identified with FASER electronic components (not emulsion) CERN-FASER-CONF-2026-004

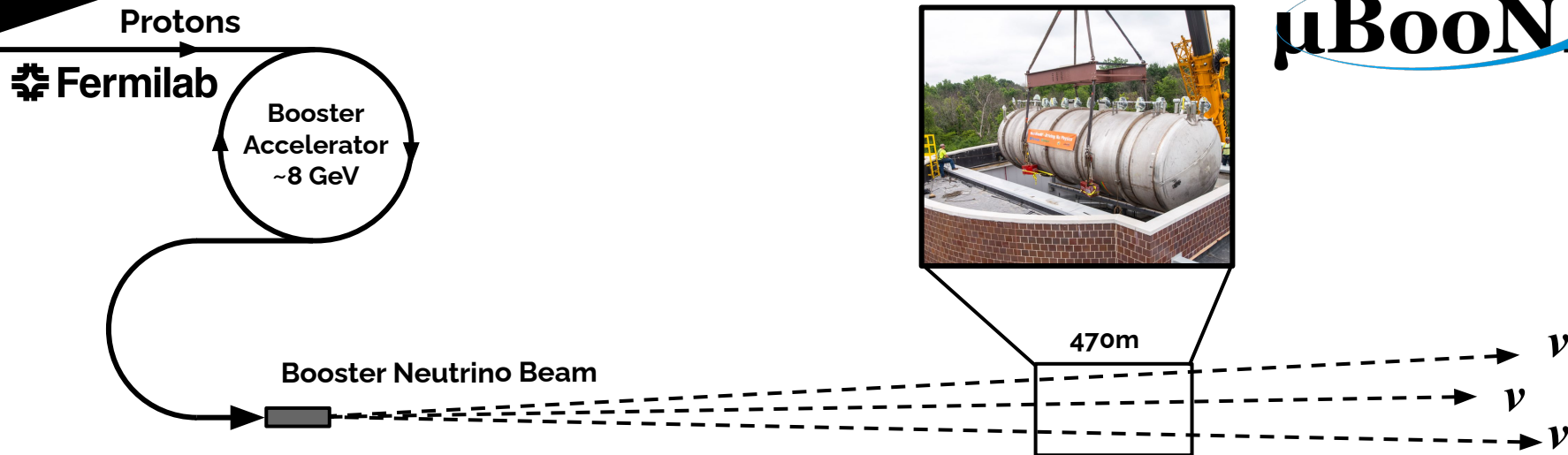
- $\sim 65 \nu_e$  events observed above an expected  $\sim 40$  background (**5.5 $\sigma$  significance**)
- Milestone towards Energy dependent differential  $\nu_e$  cross sections





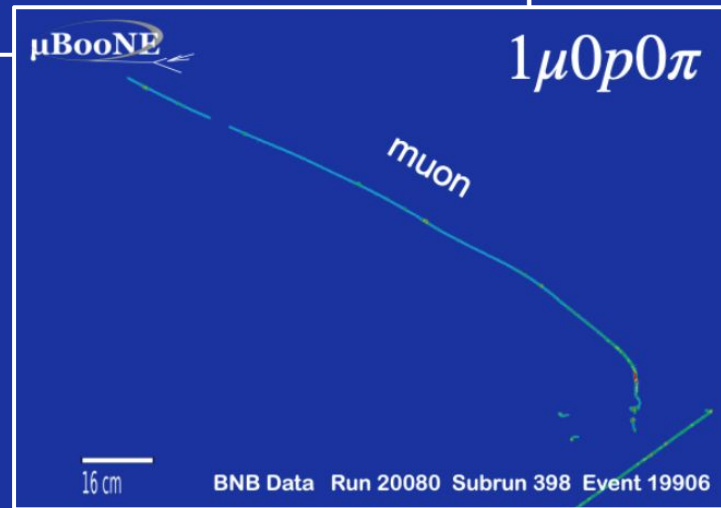
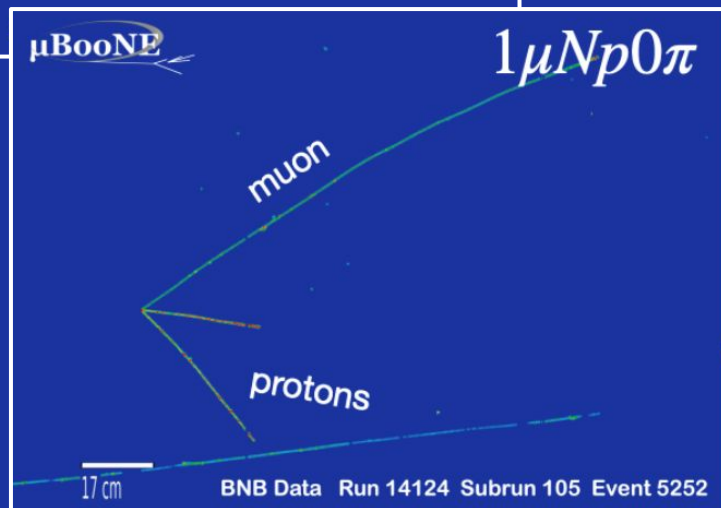
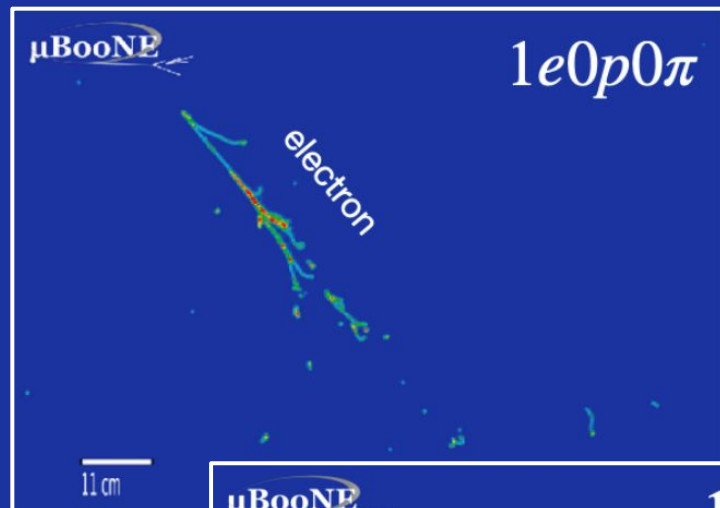
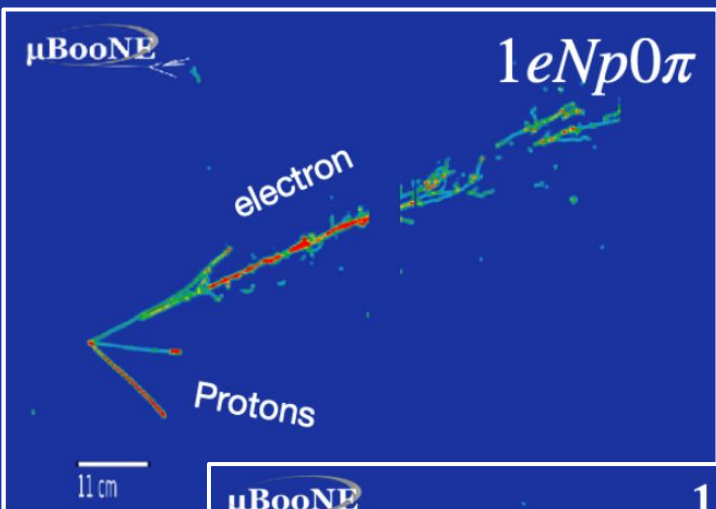
**$\mu$ BooNE**



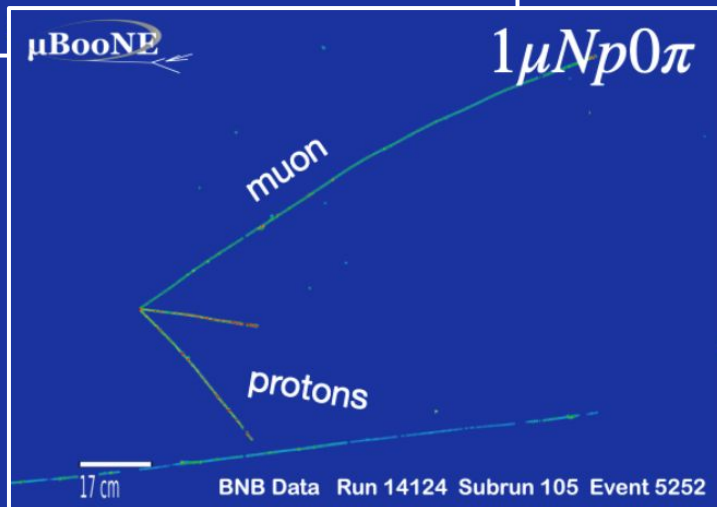
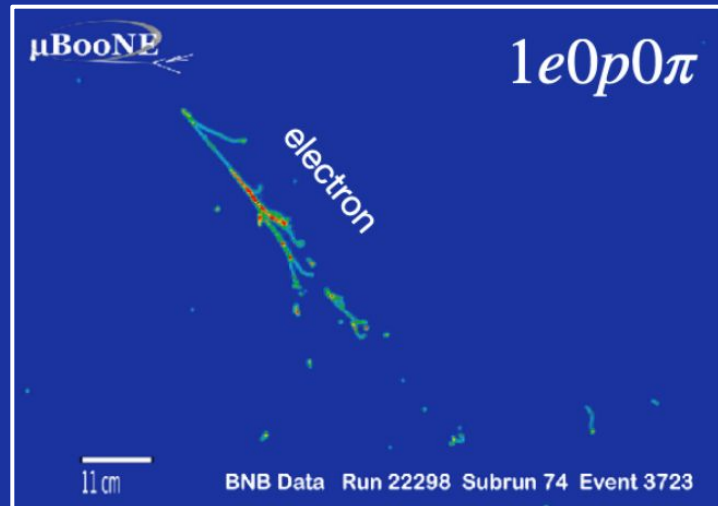
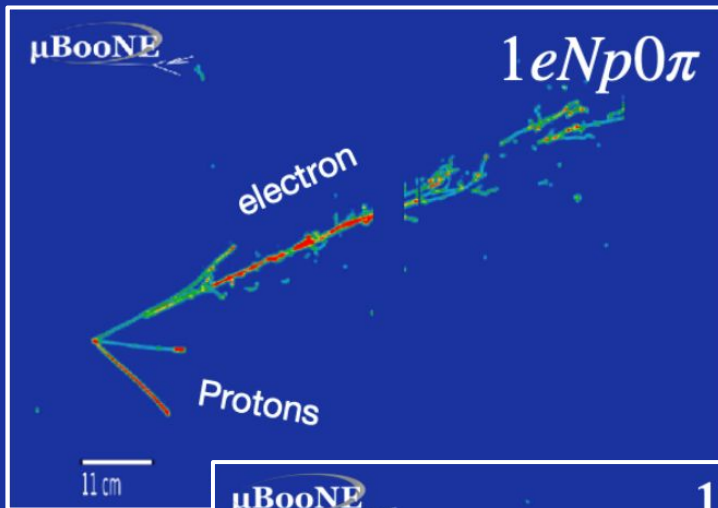


**MicroBooNE** is a **90 metric ton LArTPC** sitting in the **Booster Neutrino Beam** (BNB) at Fermilab, taking data **2015-2021**

# MicroBooNE



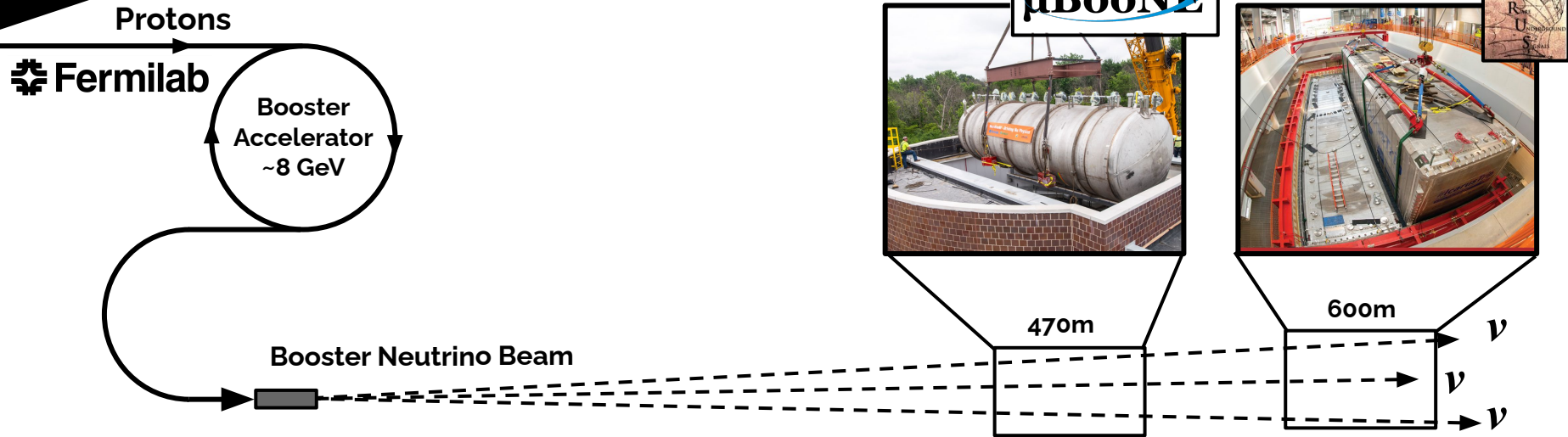
# MicroBooNE



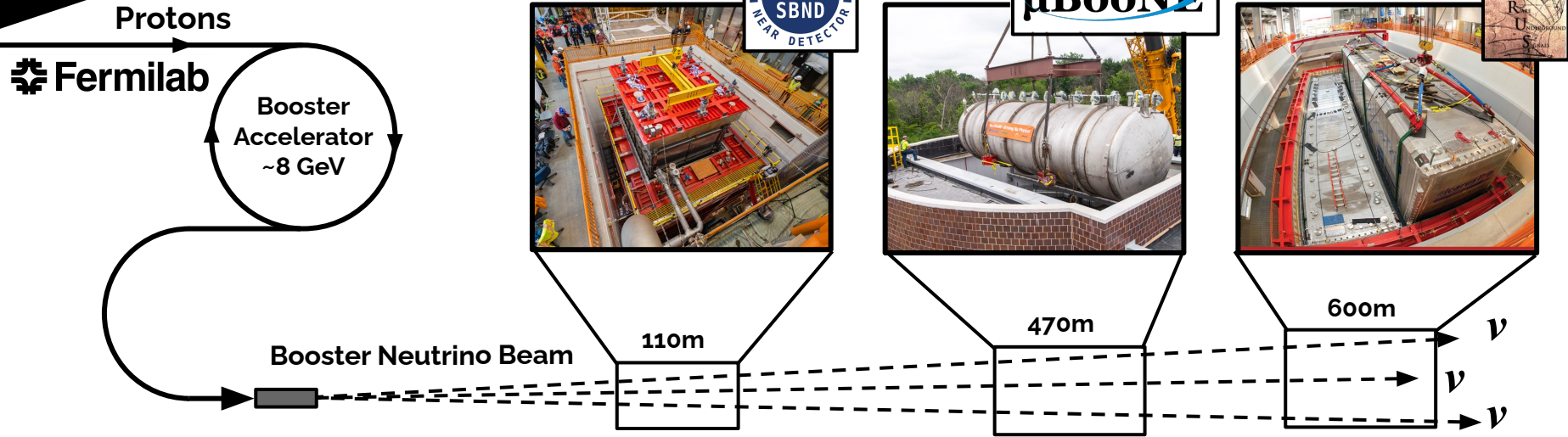
LArTPC Technology enables

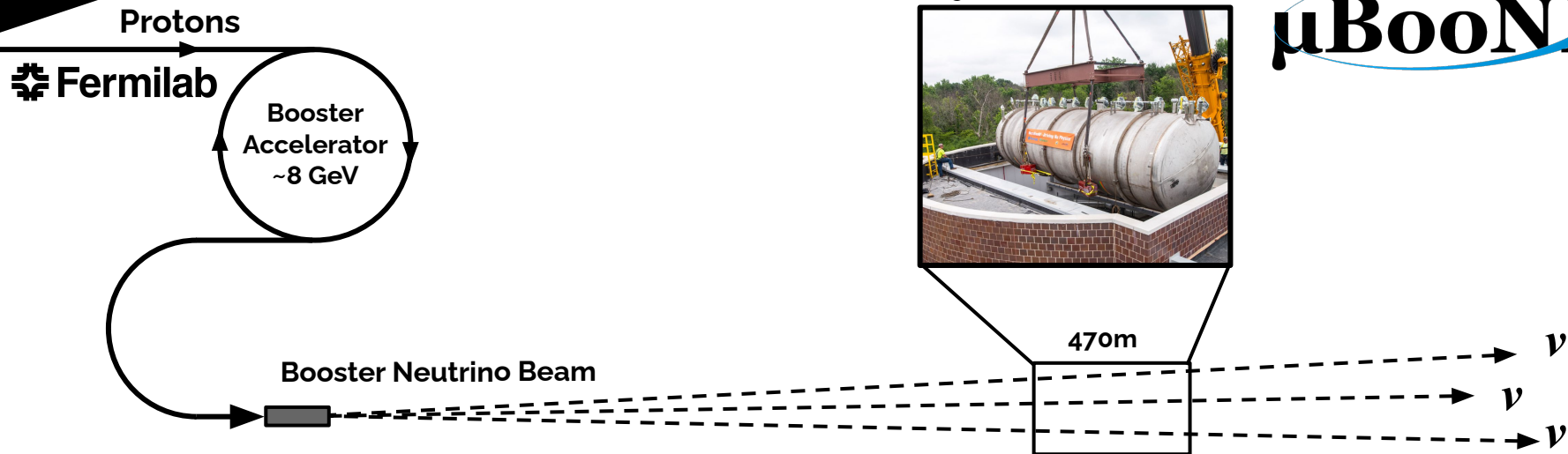
- 3 Millimeter Scale Resolution
- Calorimetry & Energy Measurement
- Powerful Particle ID

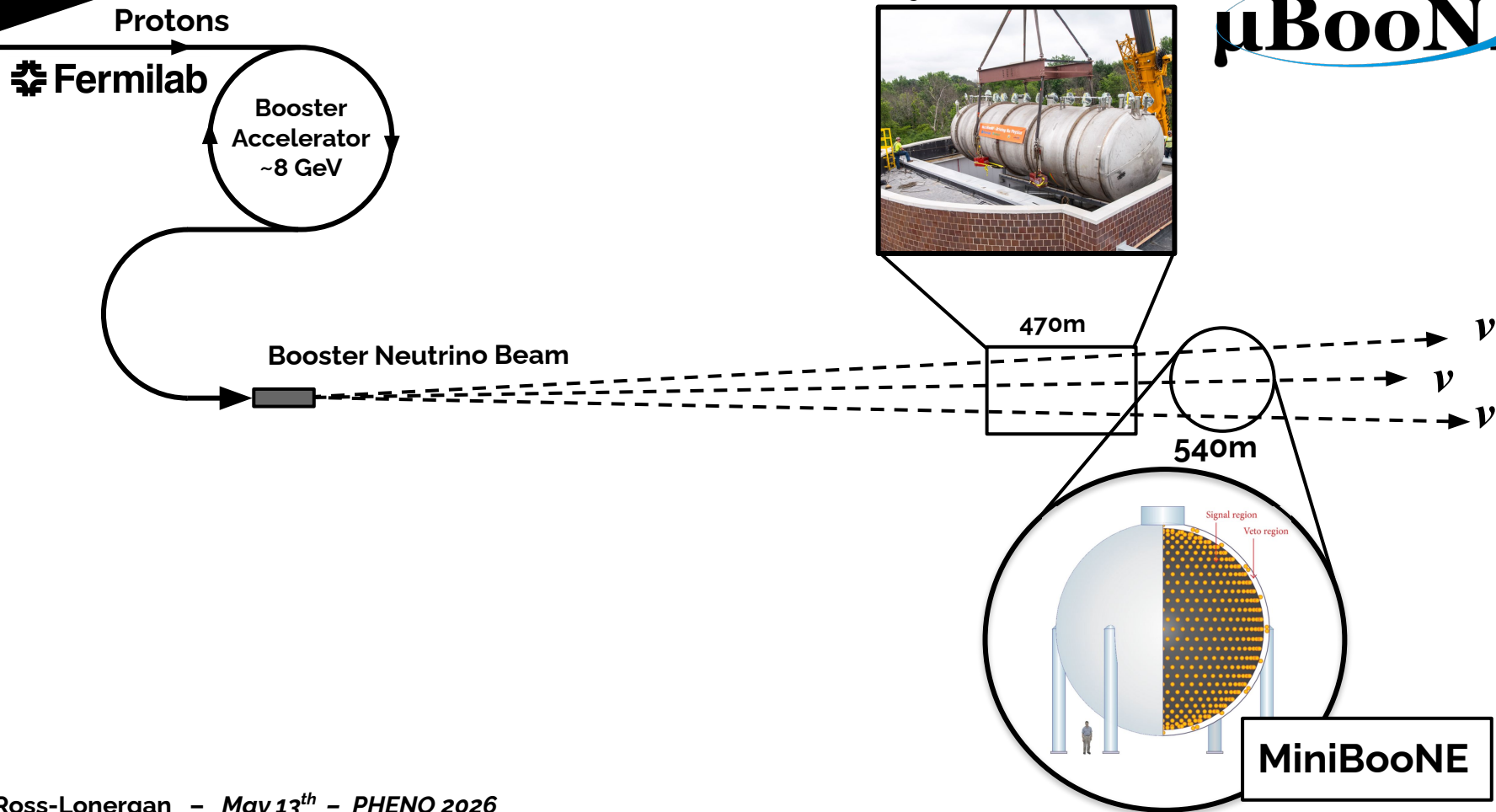
# SBN Program

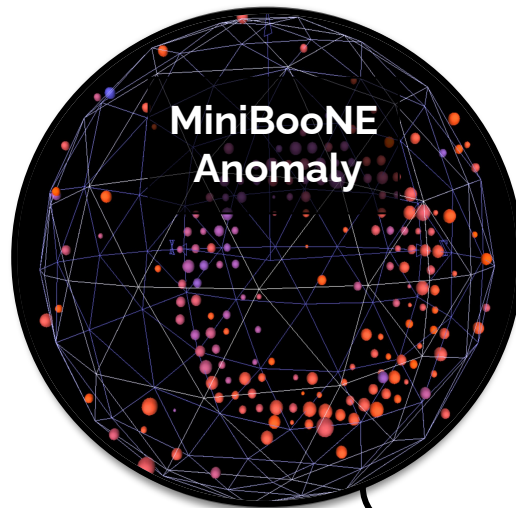


# SBN Program

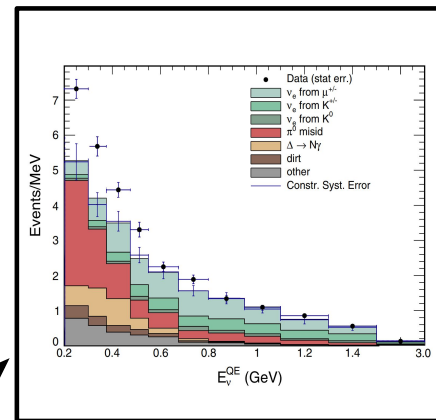


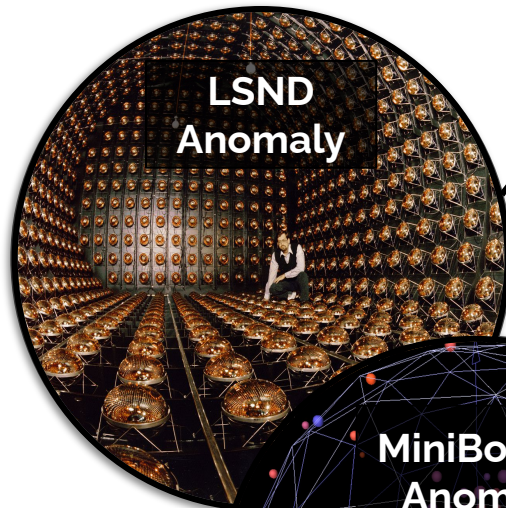




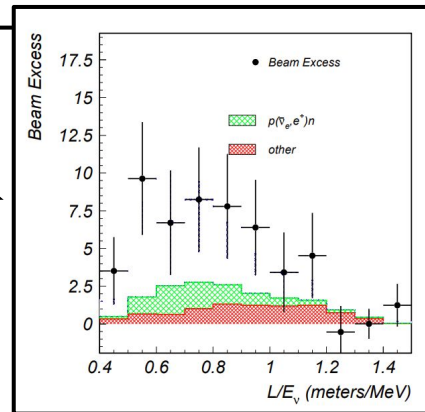


**4.8 $\sigma$   
Excess**

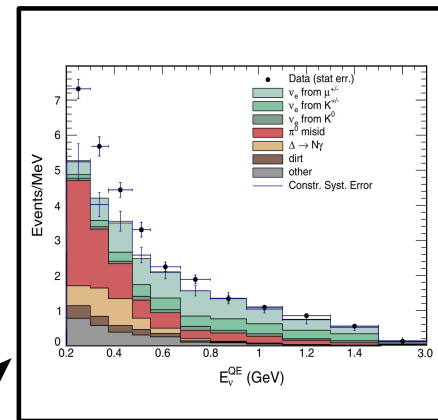




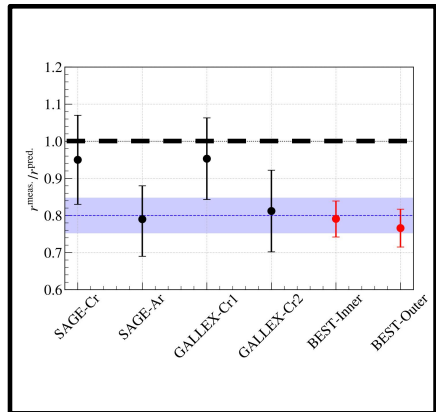
$3.8\sigma$   
Excess



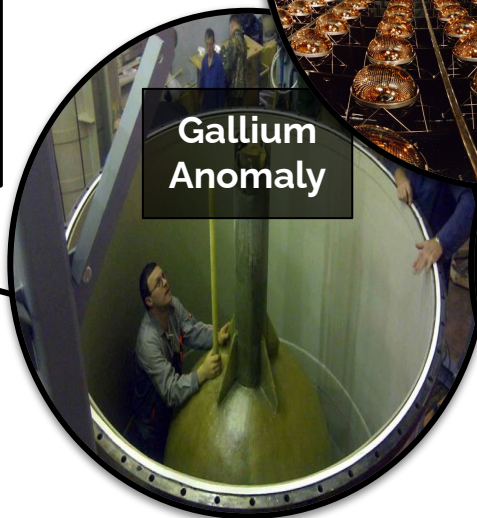
$4.8\sigma$   
Excess



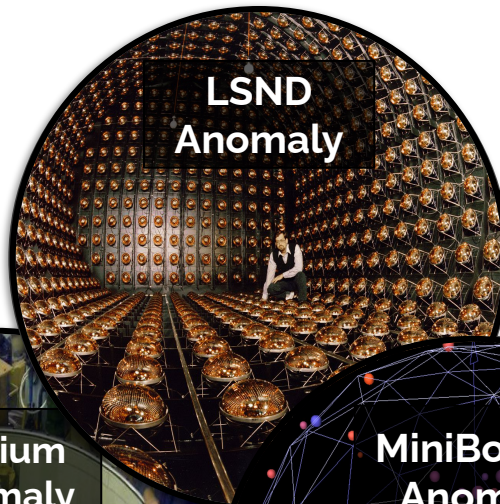
# Gallium Anomaly



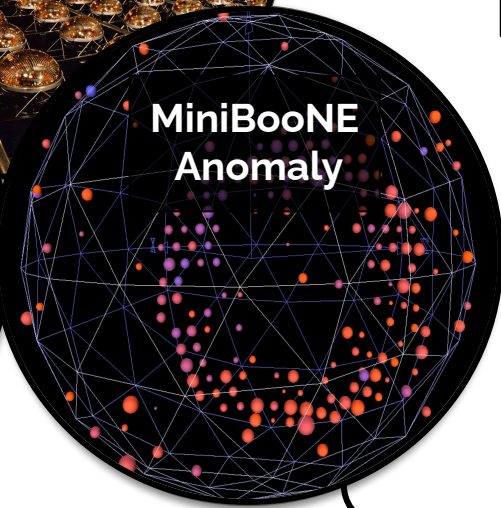
**4.0 $\sigma$   
Deficit**



**Gallium  
Anomaly**

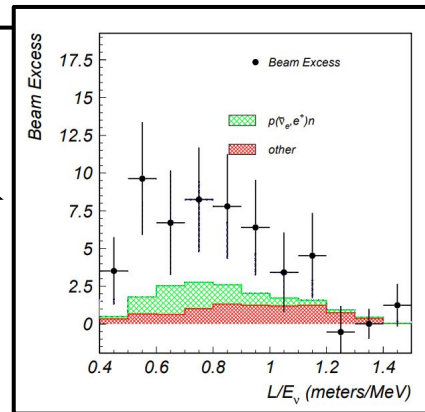


**LSND  
Anomaly**

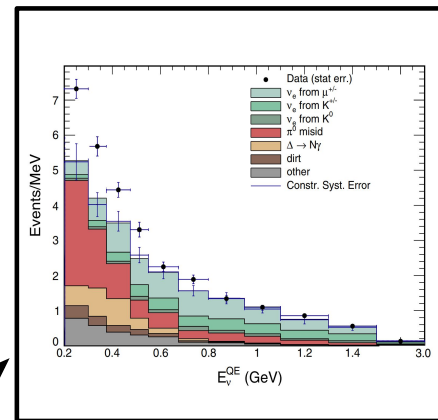


**MiniBooNE  
Anomaly**

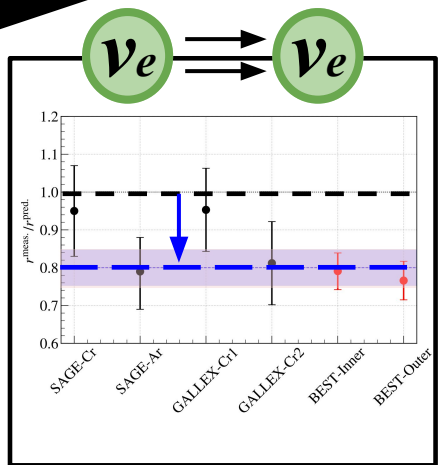
**3.8 $\sigma$   
Excess**



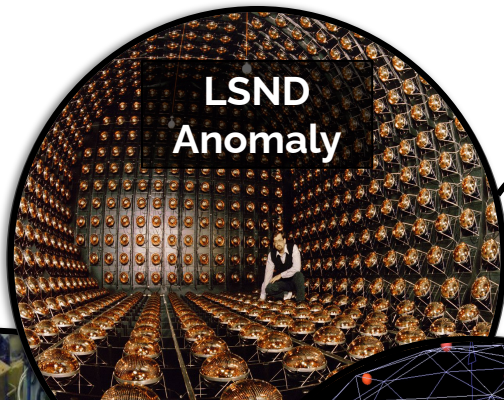
**4.8 $\sigma$   
Excess**



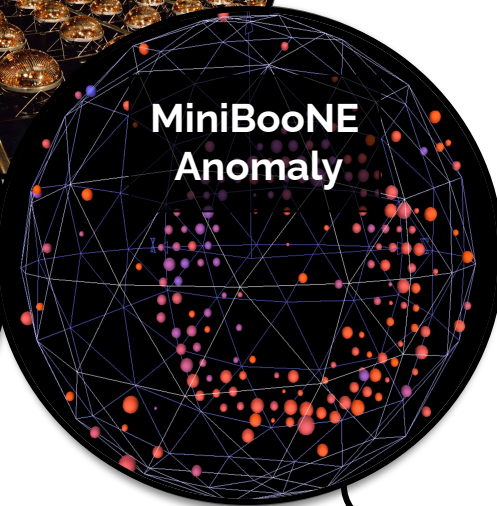
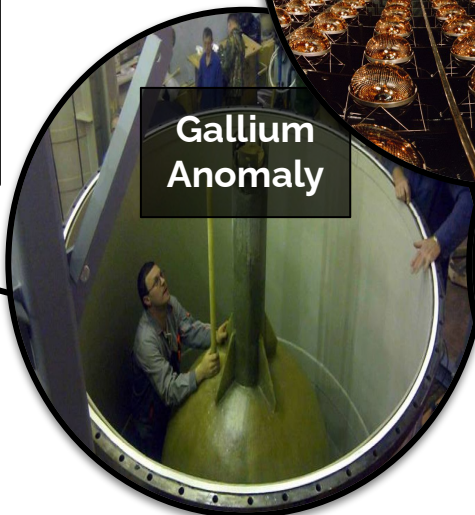
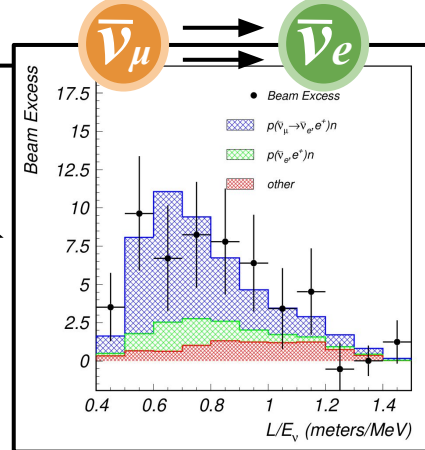
# Short Baseline Neutrino Anomalies



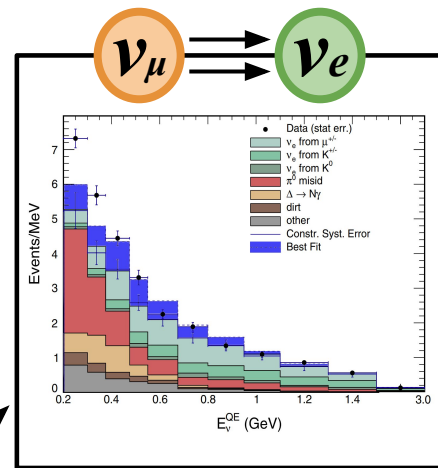
**4.0 $\sigma$   
Deficit**



**3.8 $\sigma$   
Excess**

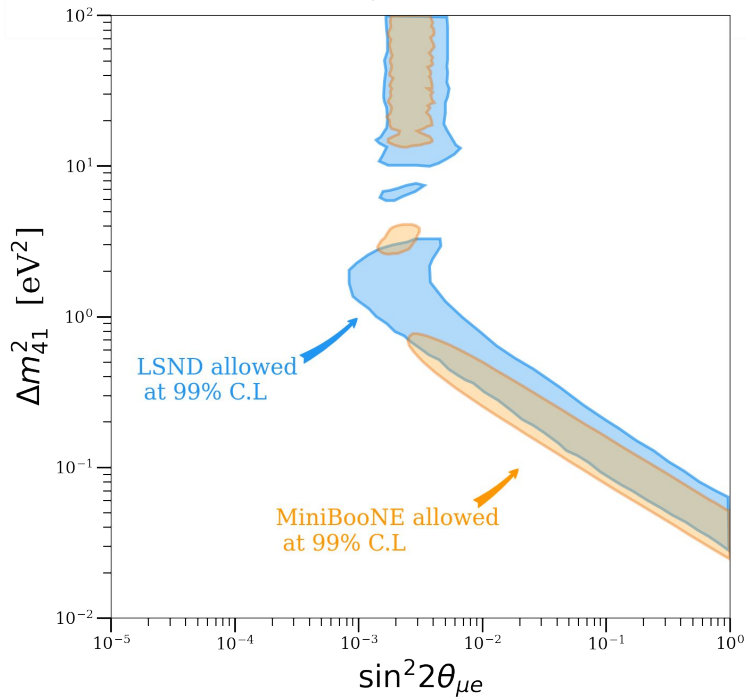


**4.8 $\sigma$   
Excess**

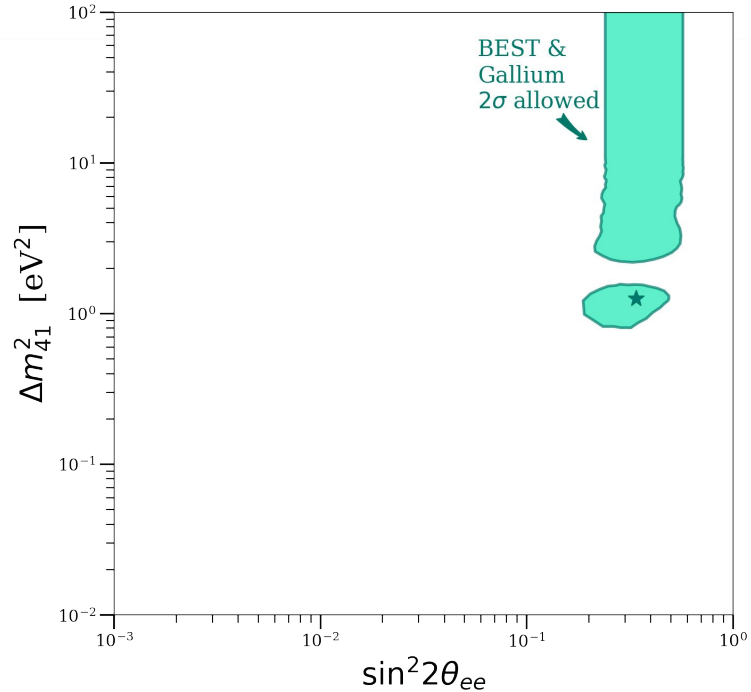


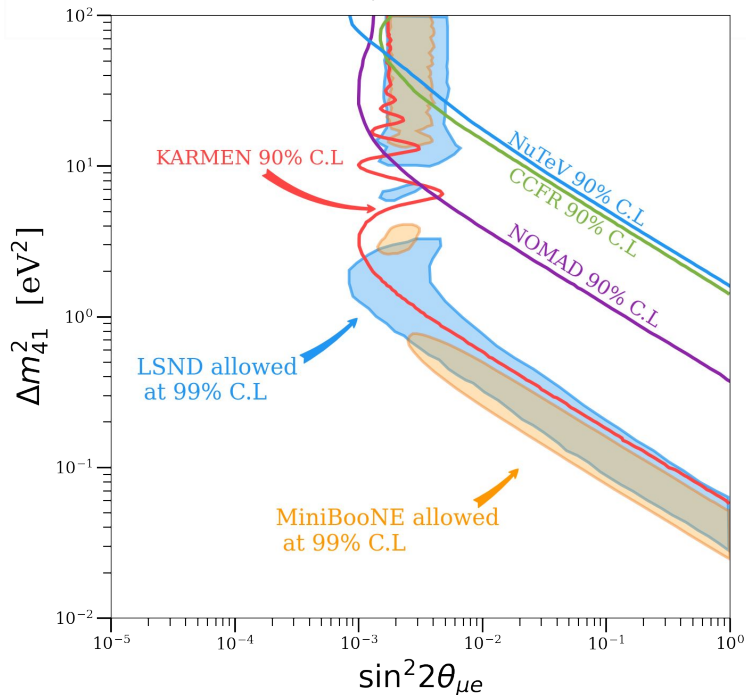
**Individually** consistent with  $eV^2$  sterile neutrinos

$\nu_\mu \rightarrow \nu_e$

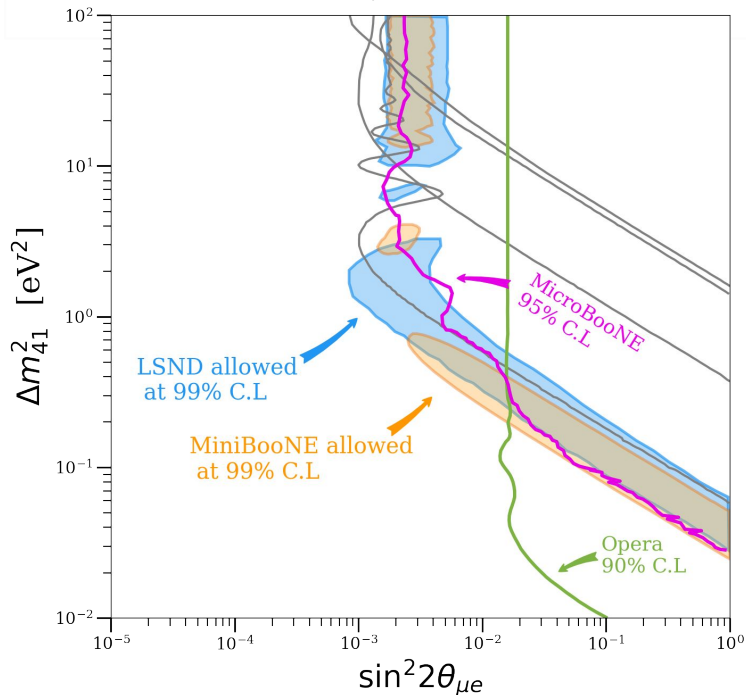


$\nu_e \rightarrow \nu_e$





- **CCFR [1984] @ Fermilab**
  - PhysRevLett.52:10.1103, 1984
- **KARMEN [2002] @ RAL (UK)**
  - Phys.Rev.D65:112001,2002
- **NuTeV [2002] @ Fermilab**
  - Phys.Rev.Lett.89:011804,2002
- **NOMAD [2003] @ CERN**
  - Phys.Lett.B570:19-31,2003

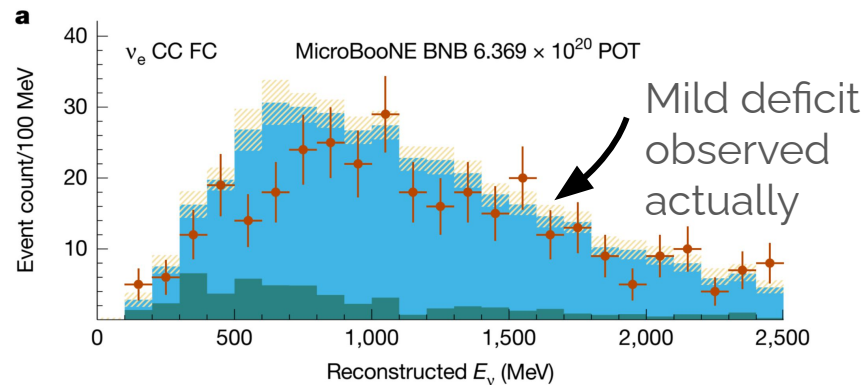


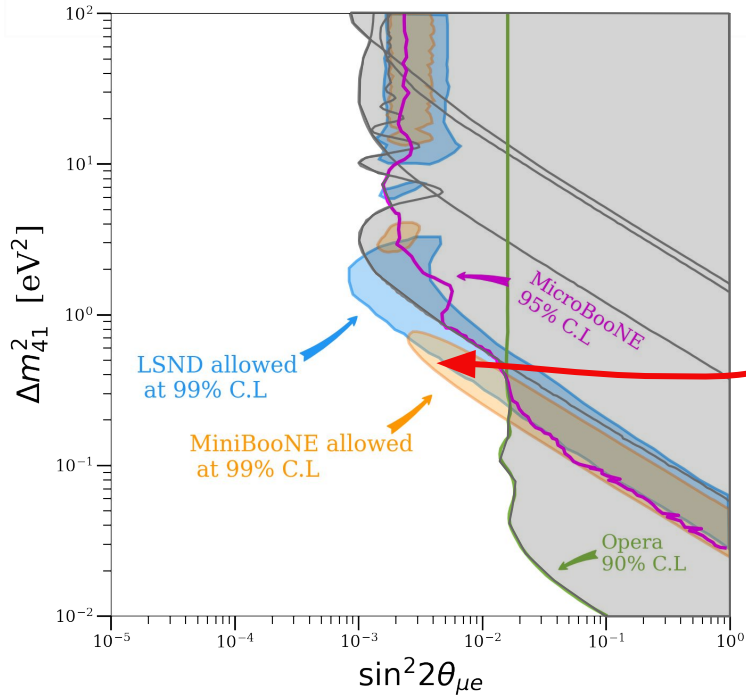
- **OPERA [2019] @ Gran Sasso**

- Phys. Rev. D 100, 051301 (2019)

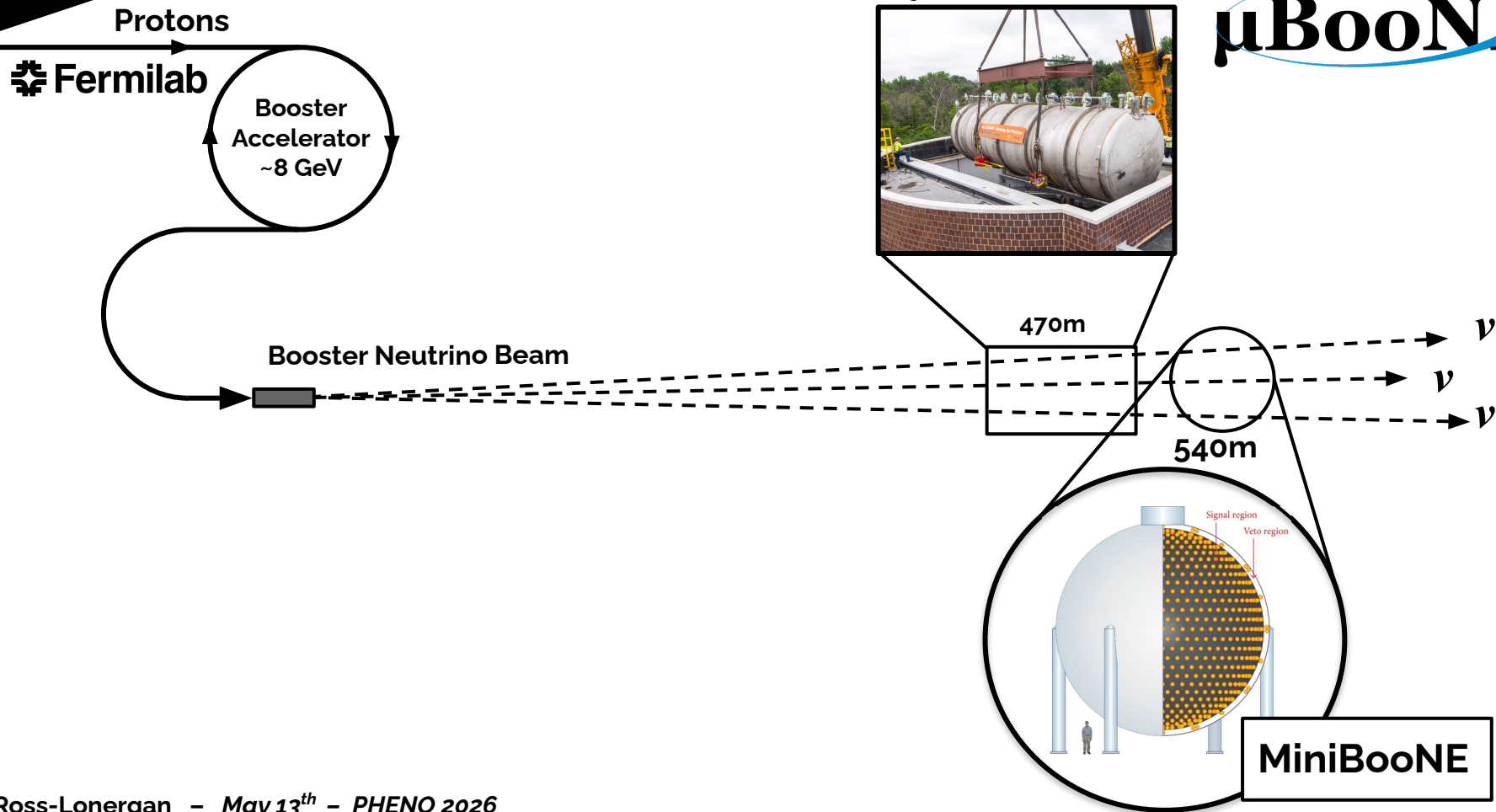
- **MicroBooNE [2023] @ Fermilab**

- Phys. Rev. Lett. 130, 011801 (2023)

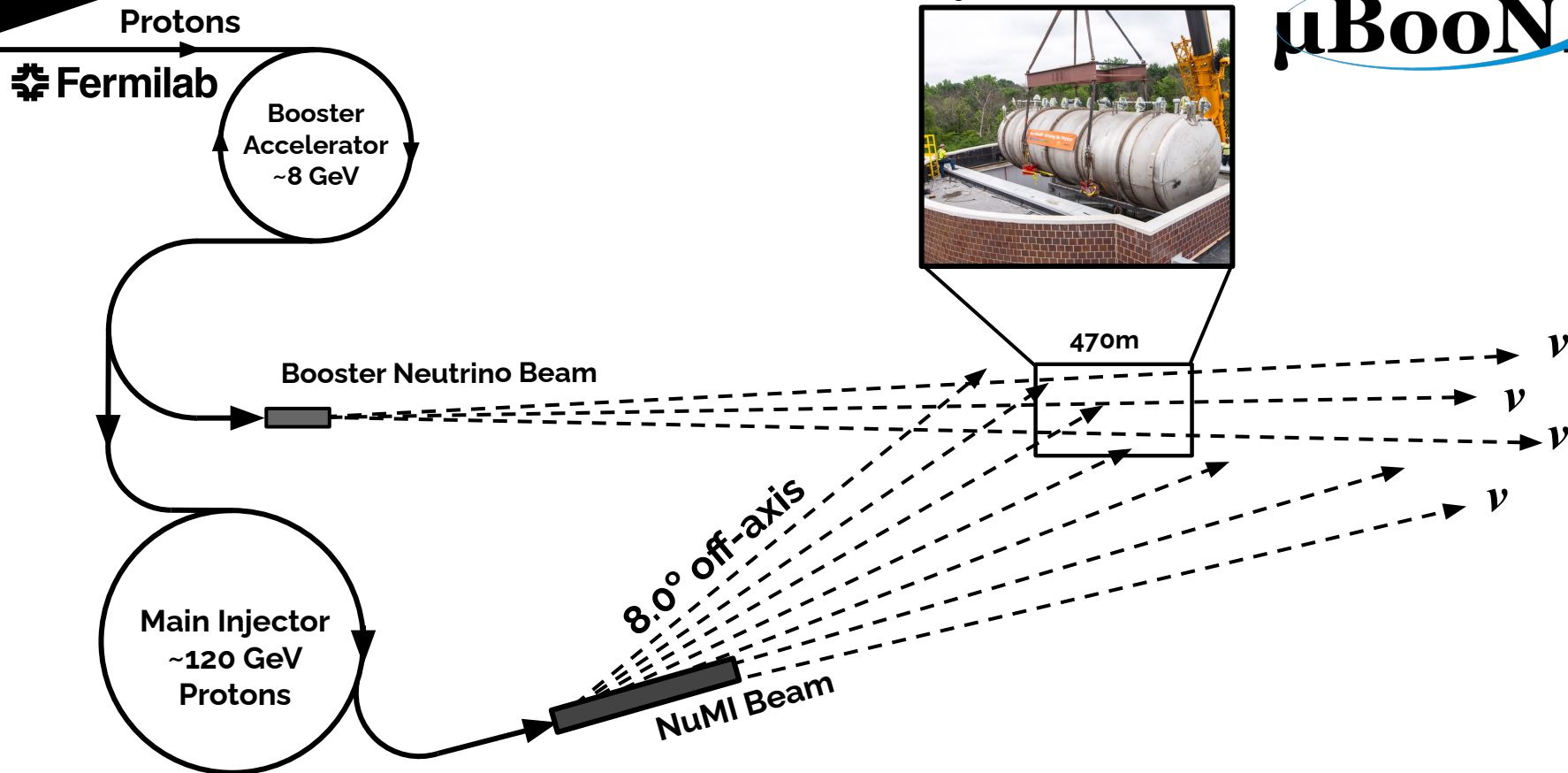




As of **PHENO 2025**, pesky region unprobed..

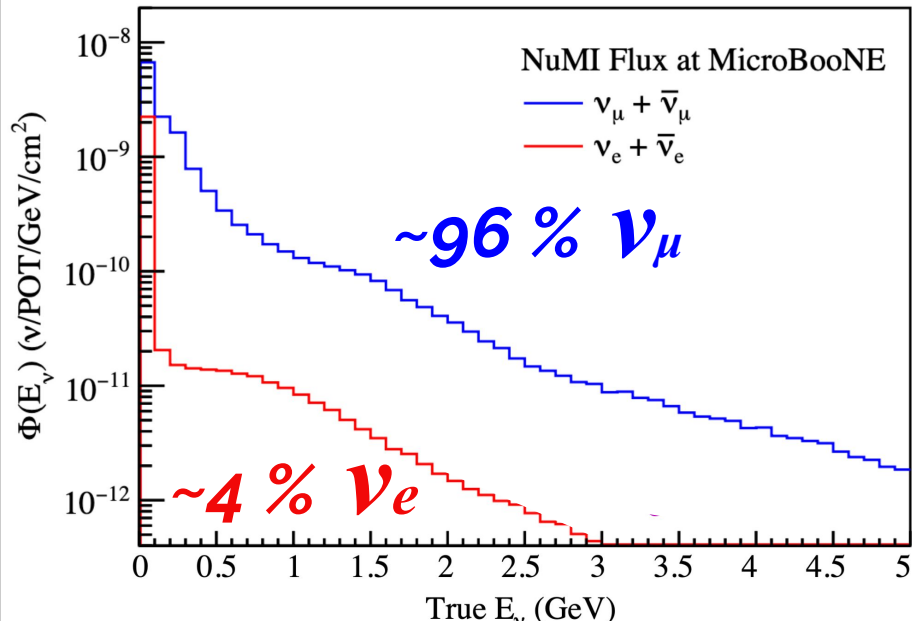
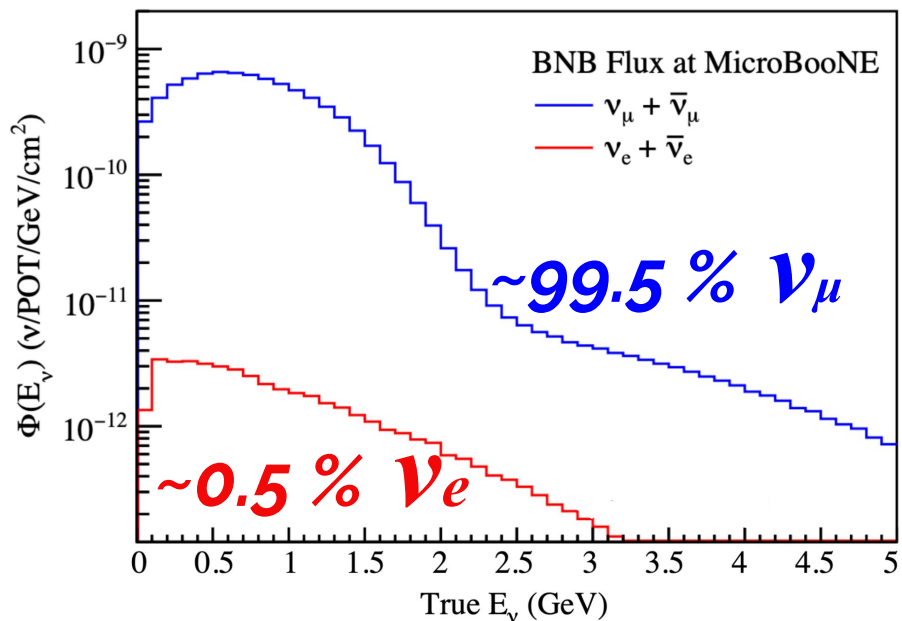


# MicroBooNE and the NuMI Beam

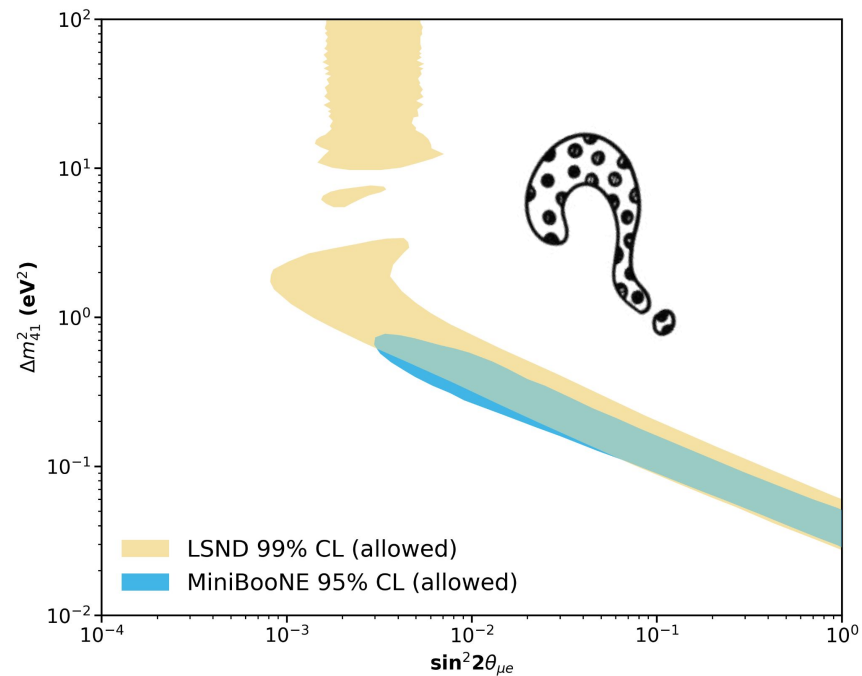
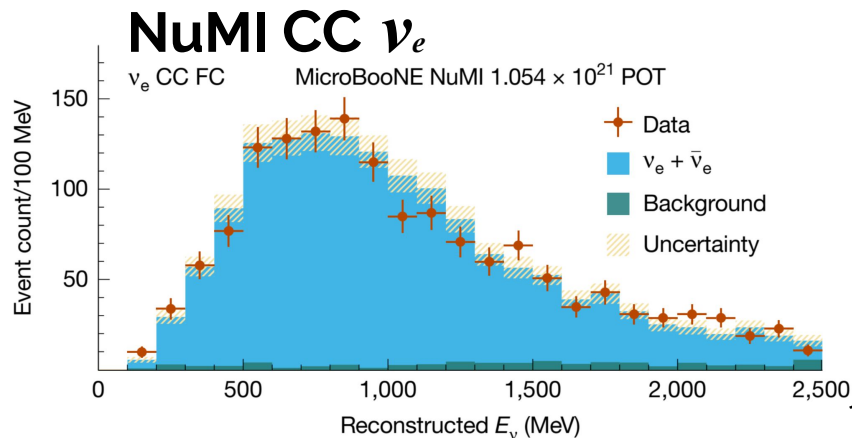
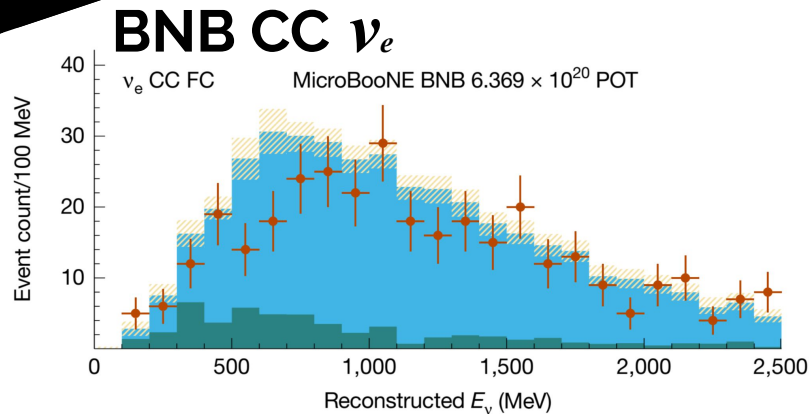


# BNB On Axis

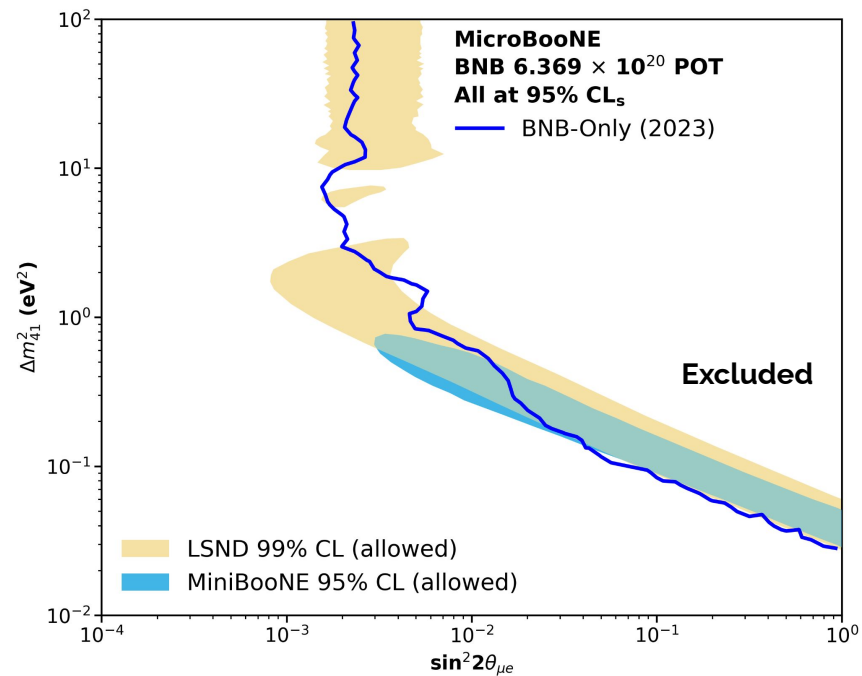
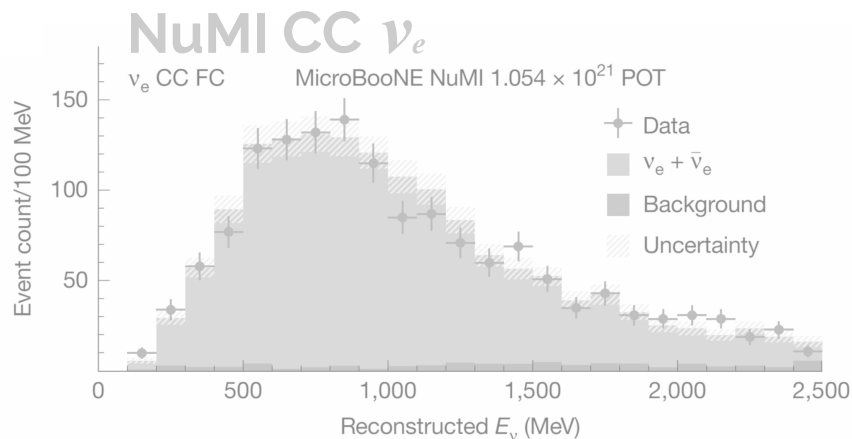
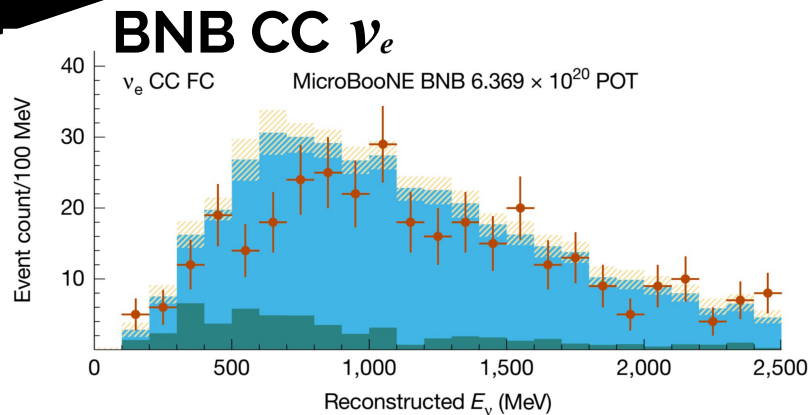
# NuMI 8° off-axis



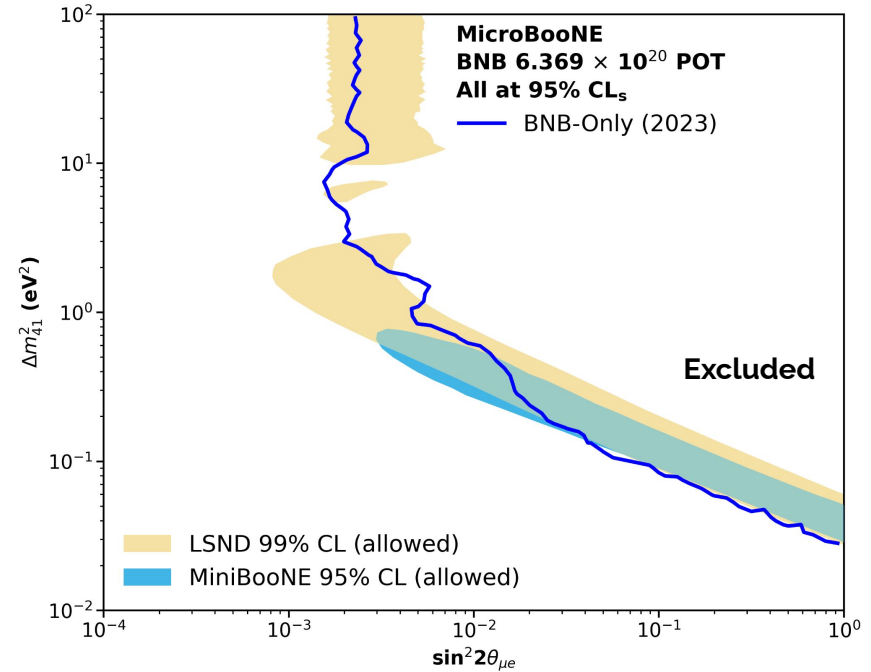
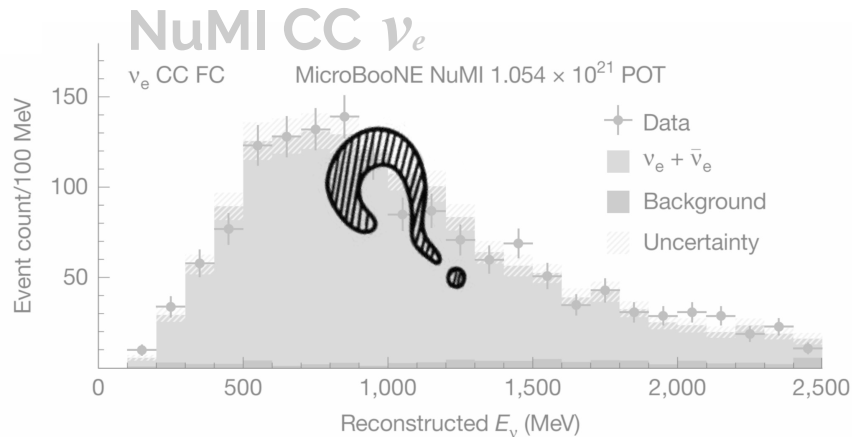
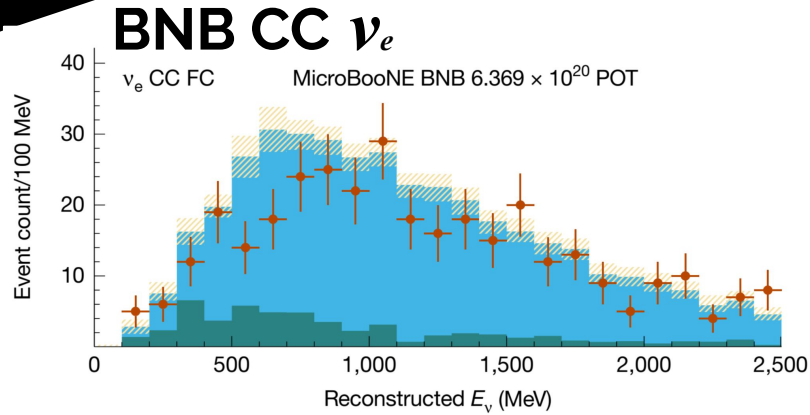
# MicroBooNE and the NuMI Beam



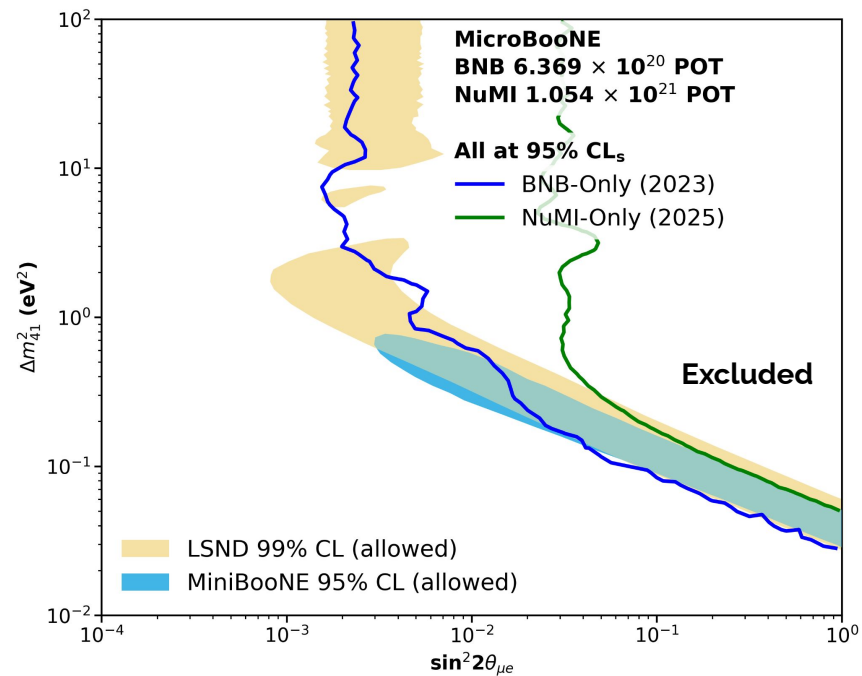
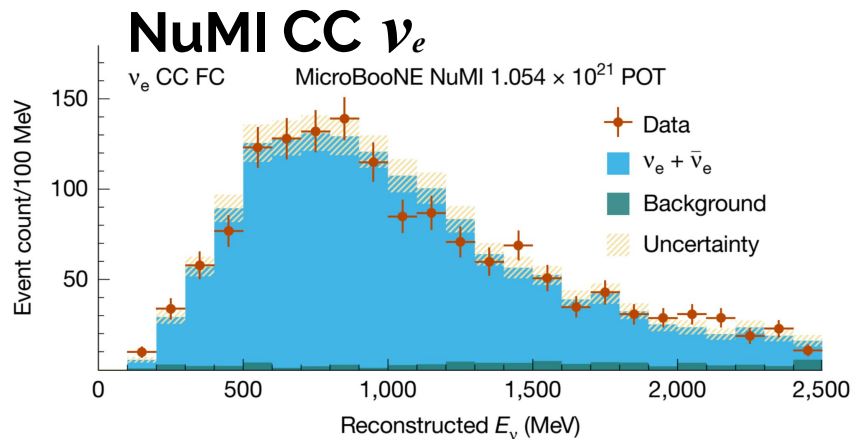
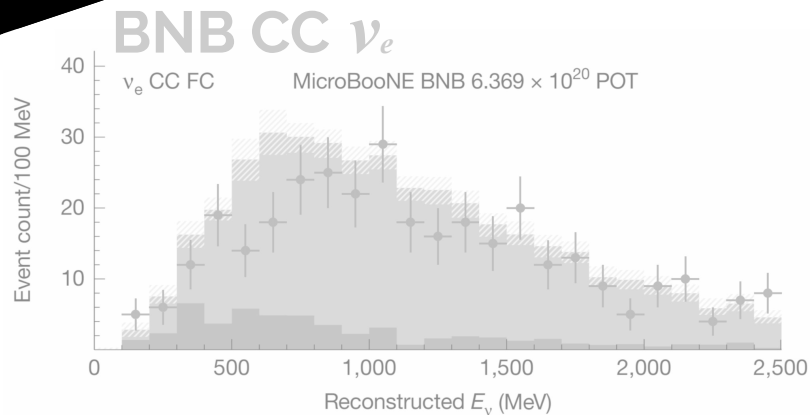
# MicroBooNE BNB Only Result (2023)



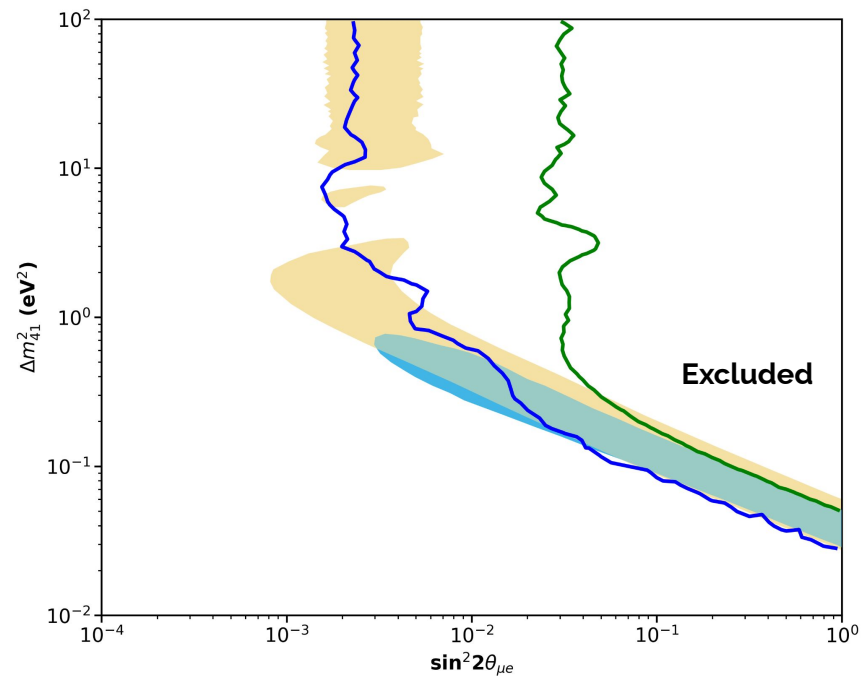
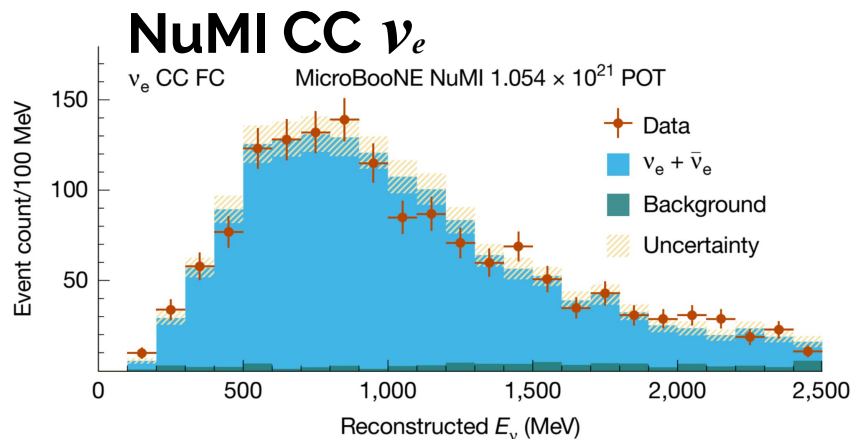
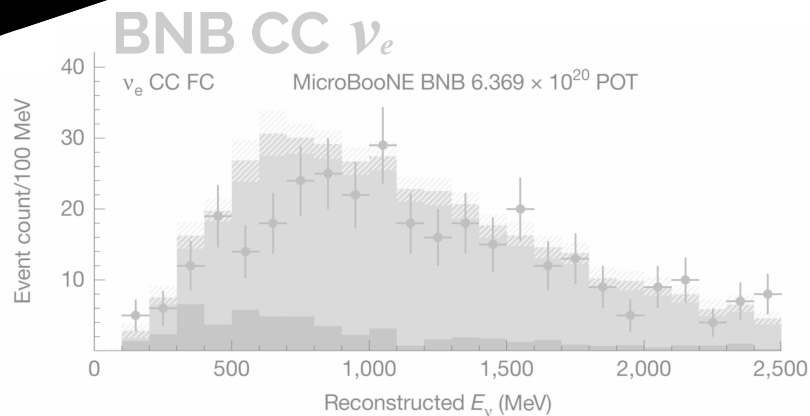
# MicroBooNE BNB Only Result (2023)



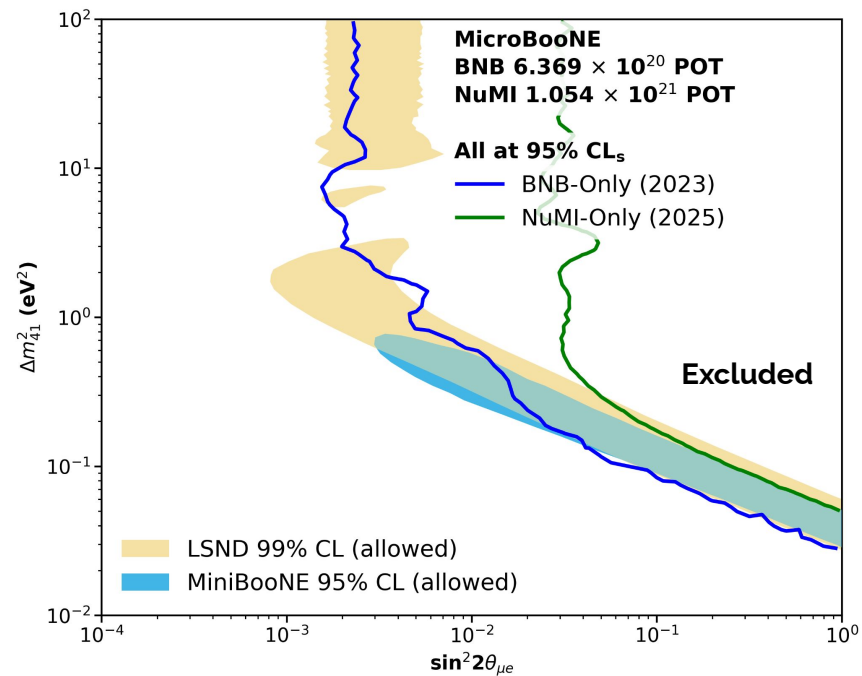
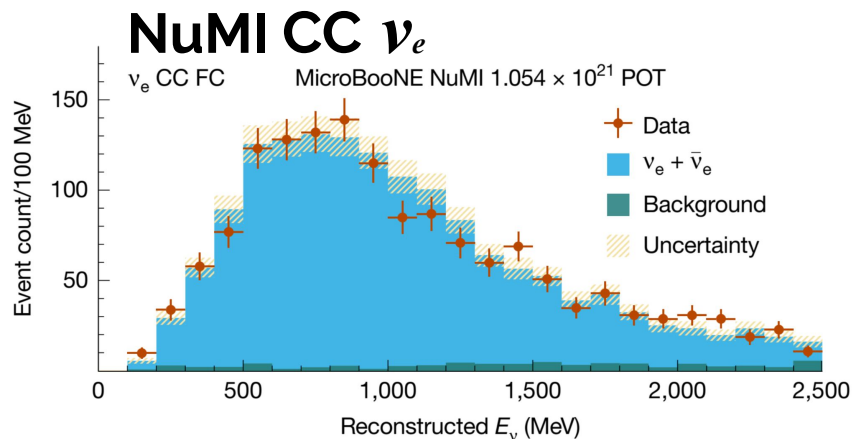
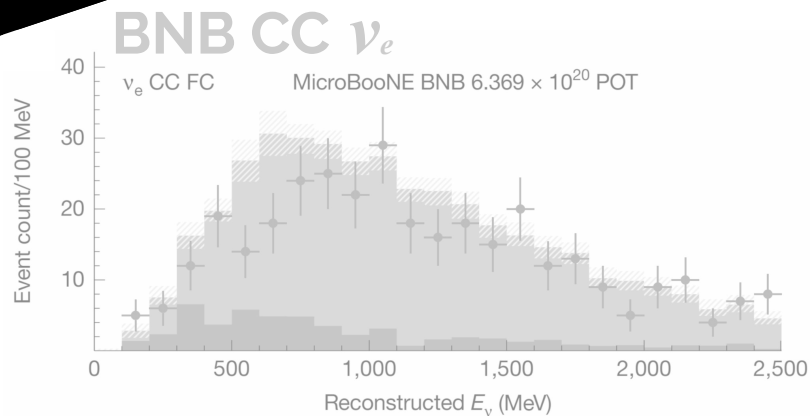
# MicroBooNE NuMI Only Result



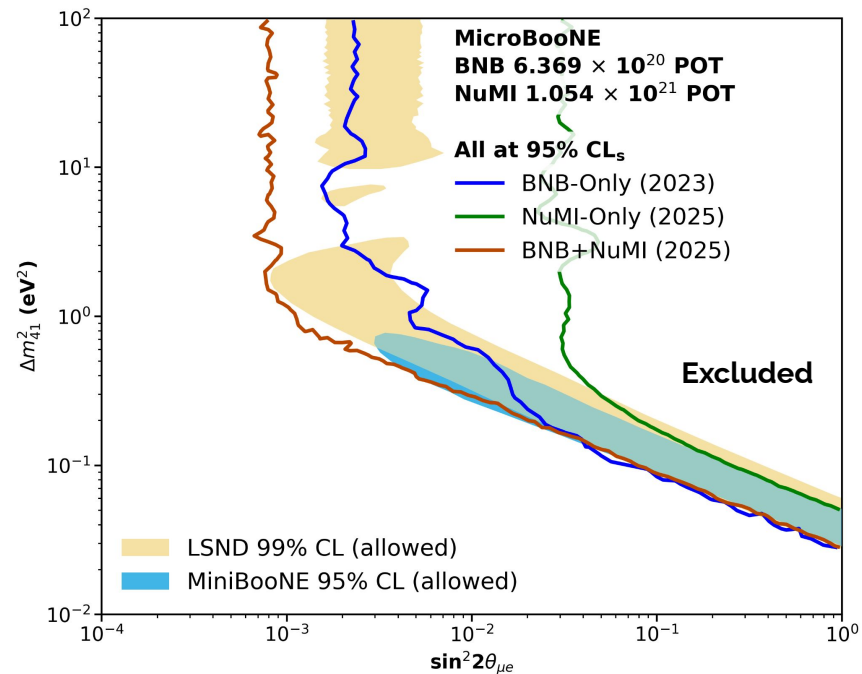
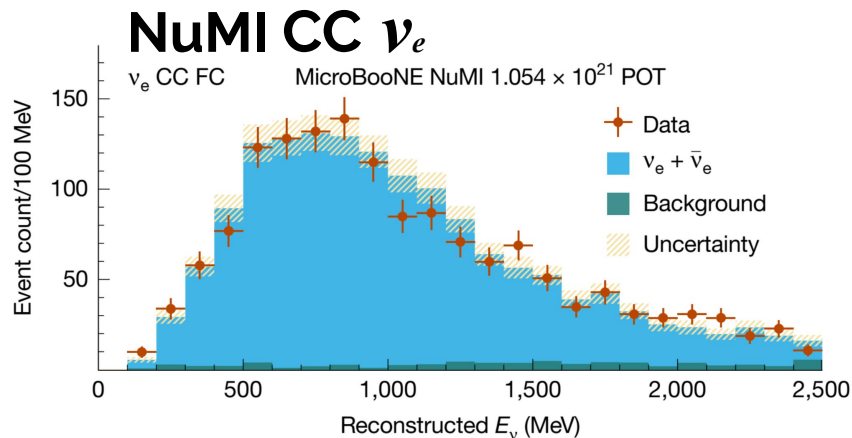
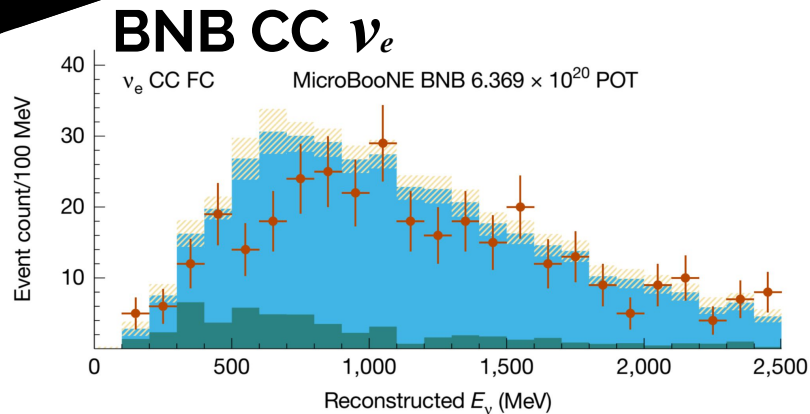
# MicroBooNE NuMI Only Result



# MicroBooNE NuMI Only Result



# MicroBooNE BNB+NuMI Dual Beam Result (2026)

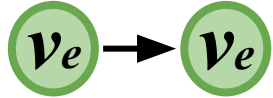


# MicroBooNE 3+1 Degeneracy

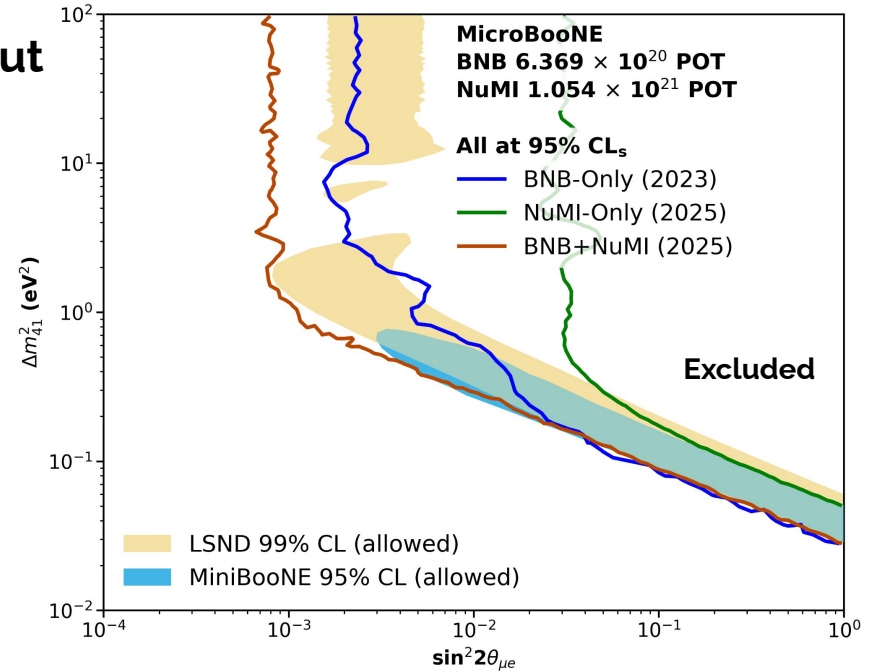
$\nu_e$  appearance



$\nu_e$  disappearance



Not independent  
Can't have appearance without  
associate disappearance



# MicroBooNE 3+1 Degeneracy

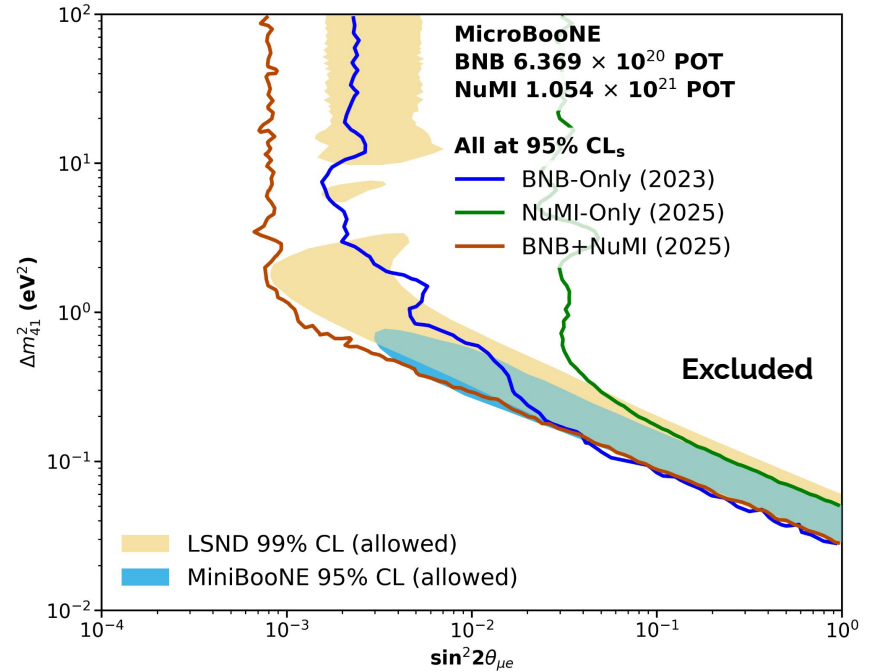
$\nu_e$  appearance



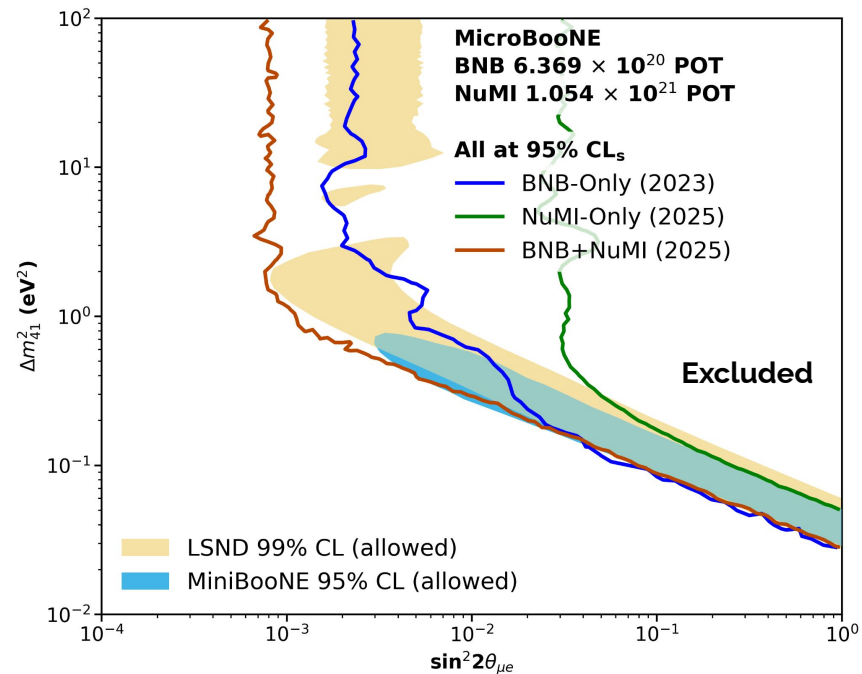
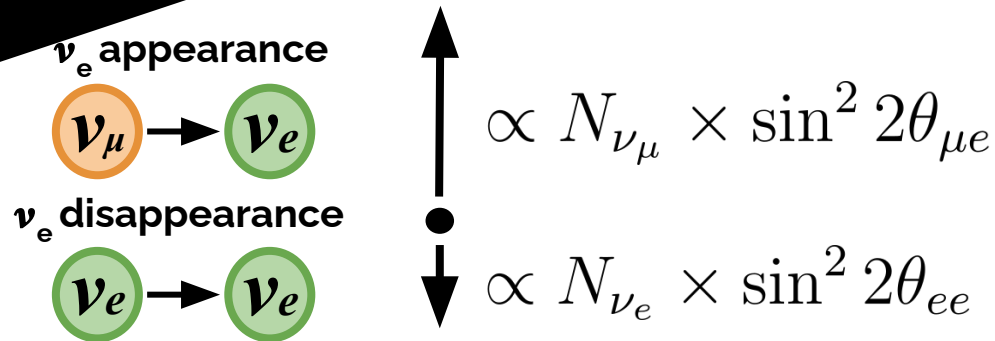
$\nu_e$  disappearance



$$\propto N_{\nu\mu} \times \sin^2 2\theta_{\mu e}$$



# MicroBooNE 3+1 Degeneracy



# MicroBooNE 3+1 Degeneracy

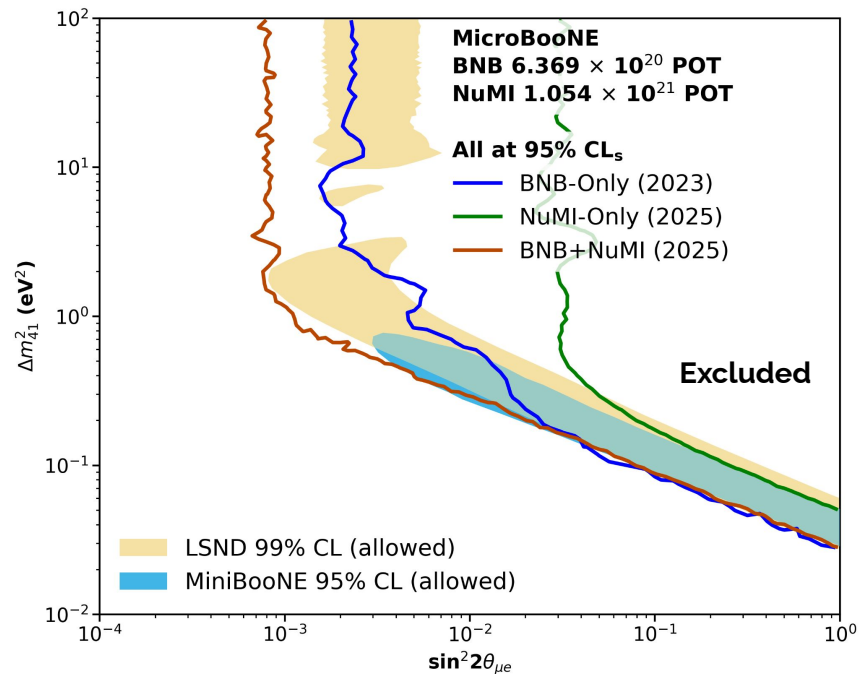
$\nu_e$  appearance



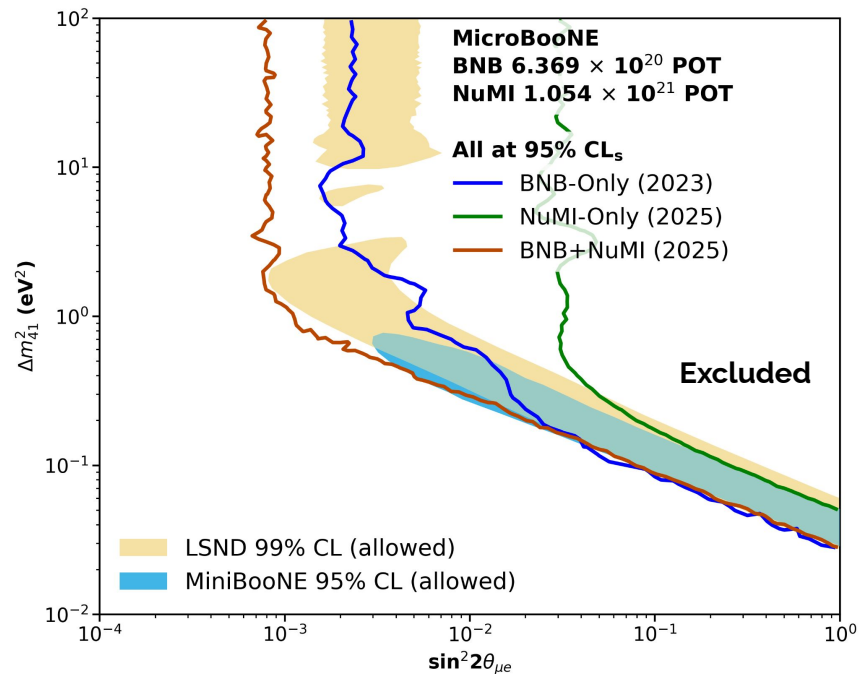
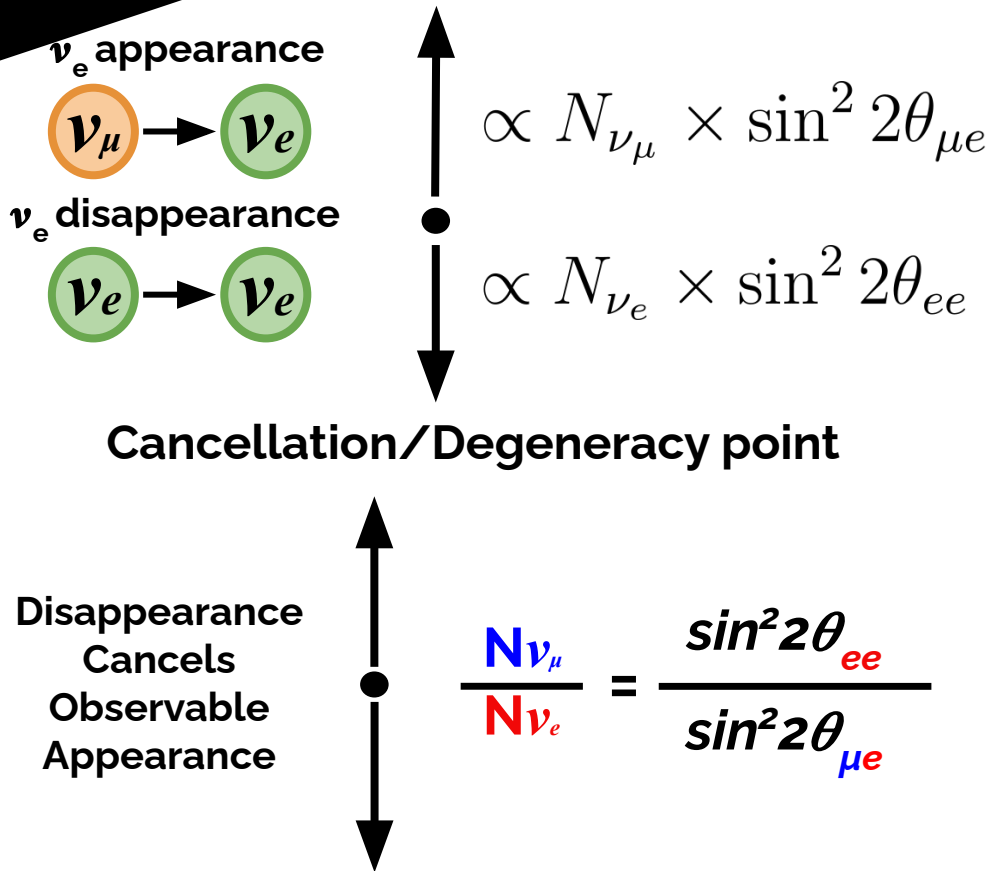
$\nu_e$  disappearance



$$\begin{aligned} &\uparrow \propto N_{\nu_\mu} \times \sin^2 2\theta_{\mu e} \\ &\bullet \\ &\downarrow \propto N_{\nu_e} \times \sin^2 2\theta_{ee} \end{aligned}$$



# MicroBooNE 3+1 Degeneracy

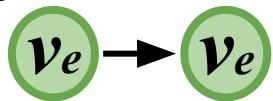


# MicroBooNE 3+1 Degeneracy

$\nu_e$  appearance



$\nu_e$  disappearance



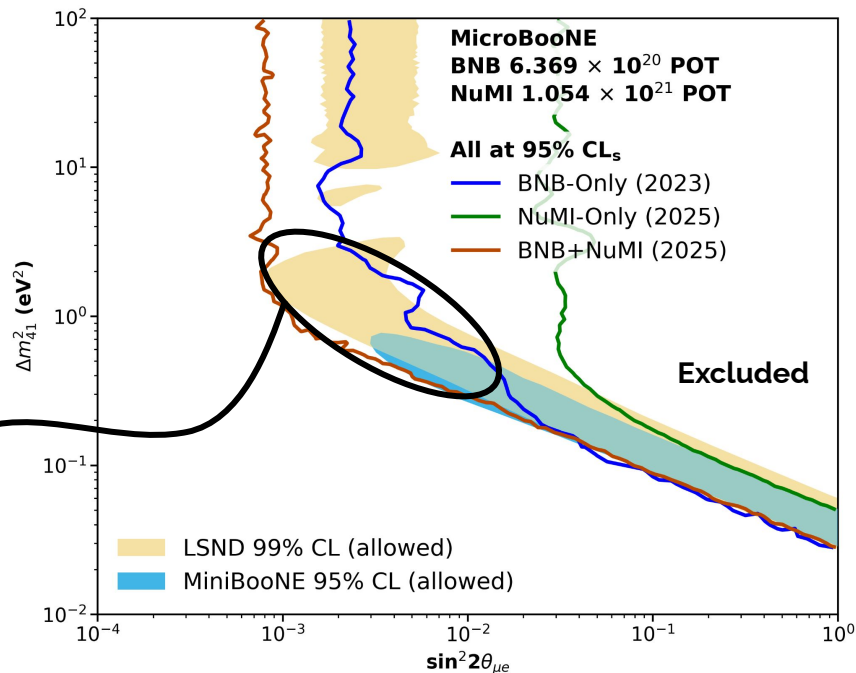
$$\propto N_{\nu_\mu} \times \sin^2 2\theta_{\mu e}$$

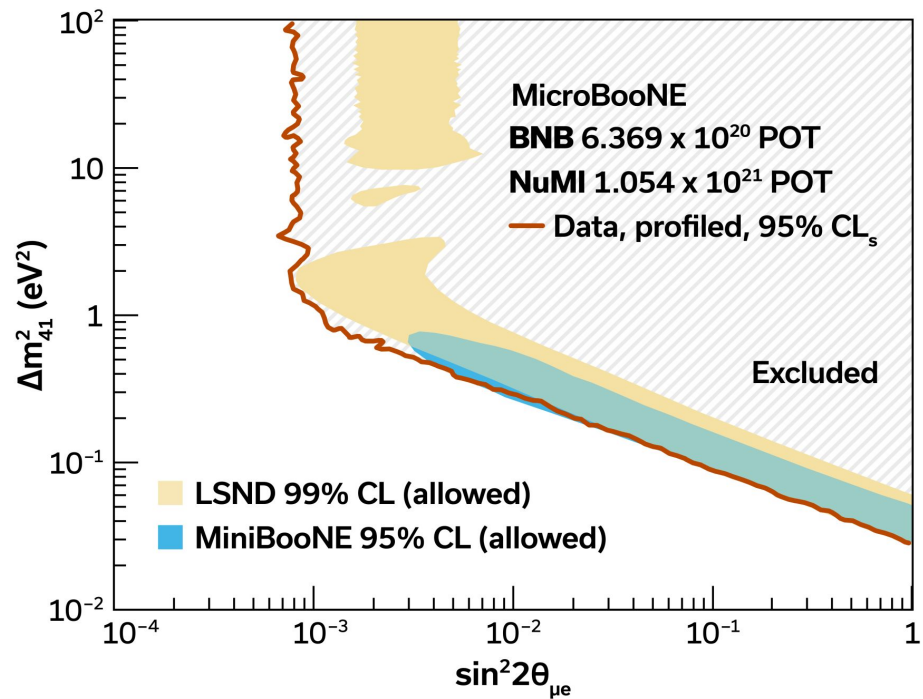
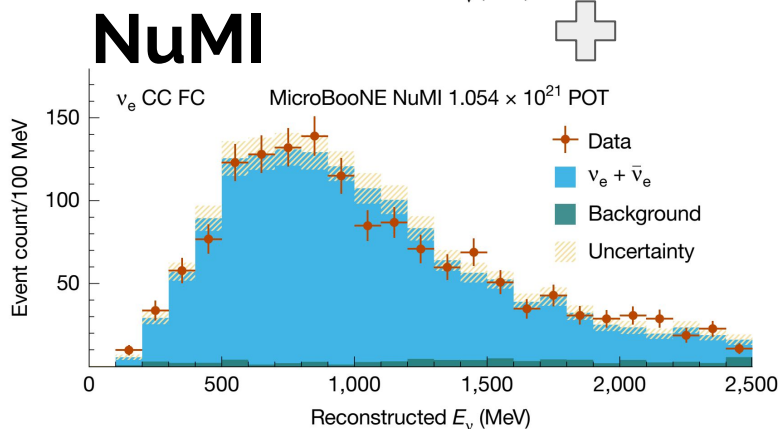
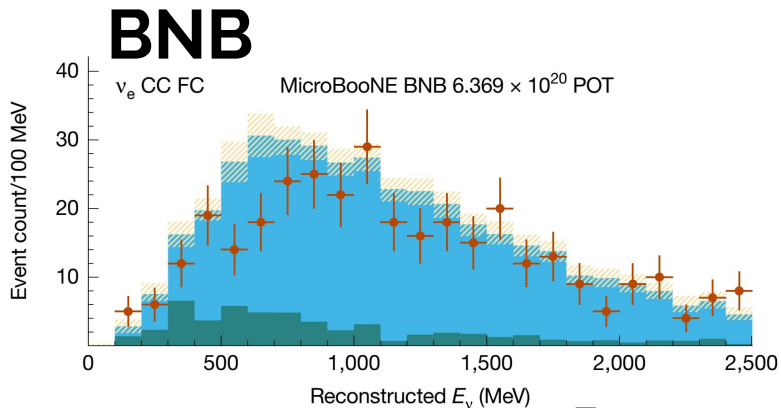
$$\propto N_{\nu_e} \times \sin^2 2\theta_{ee}$$

Cancellation/Degeneracy point

Disappearance  
Cancels  
Observable  
Appearance

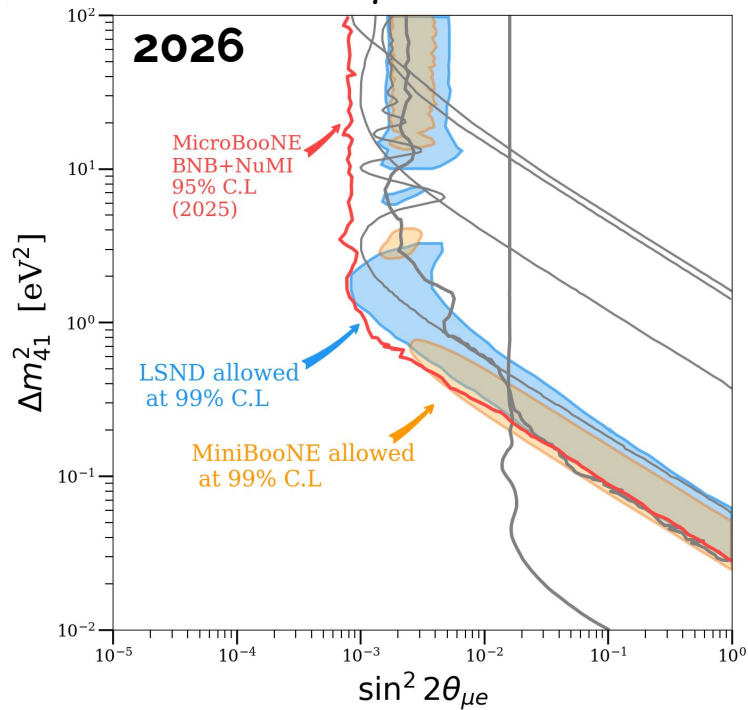
$$\frac{N_{\nu_\mu}}{N_{\nu_e}} = \frac{\sin^2 2\theta_{ee}}{\sin^2 2\theta_{\mu e}}$$

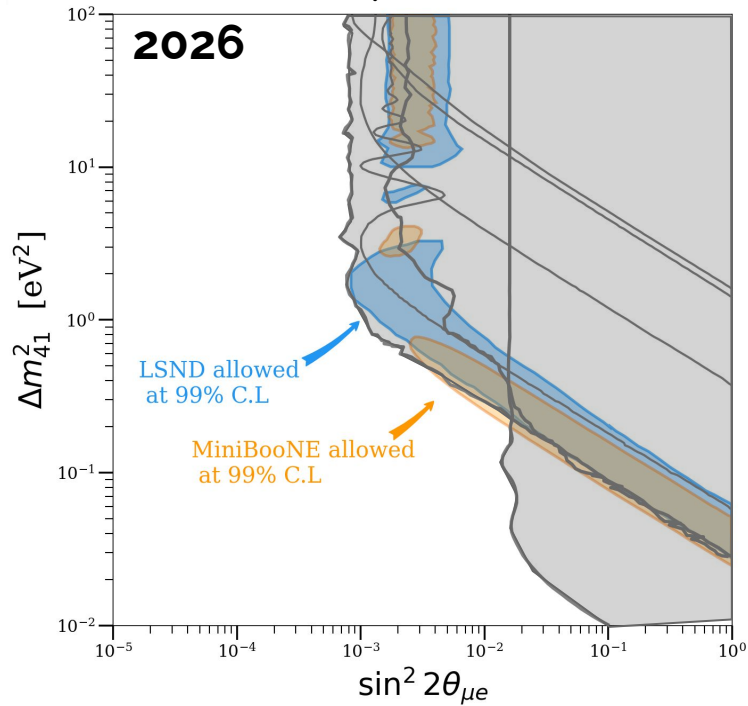




MicroBooNE 95% CL result

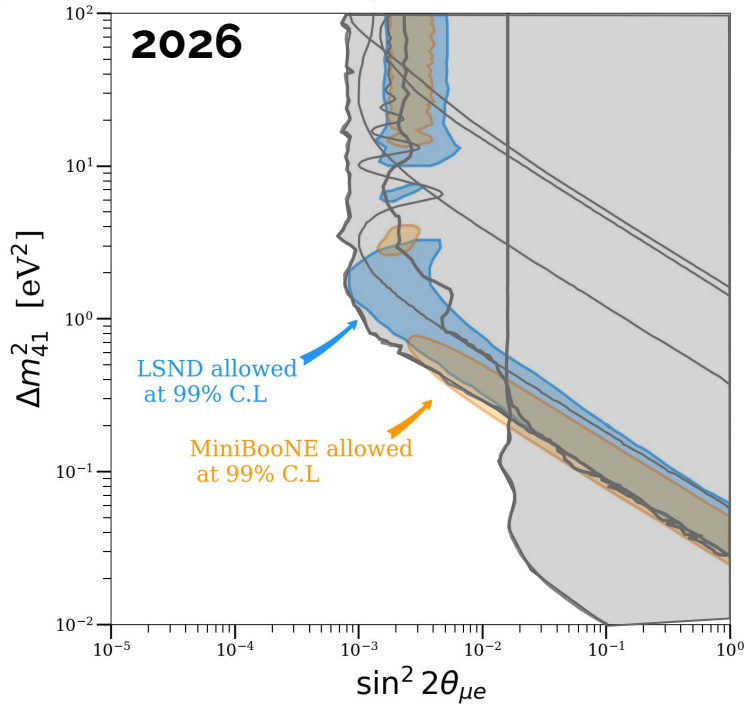
- Excludes entire LSND 99% allowed region
- Excludes vast majority of MiniBooNE 95% allowed region



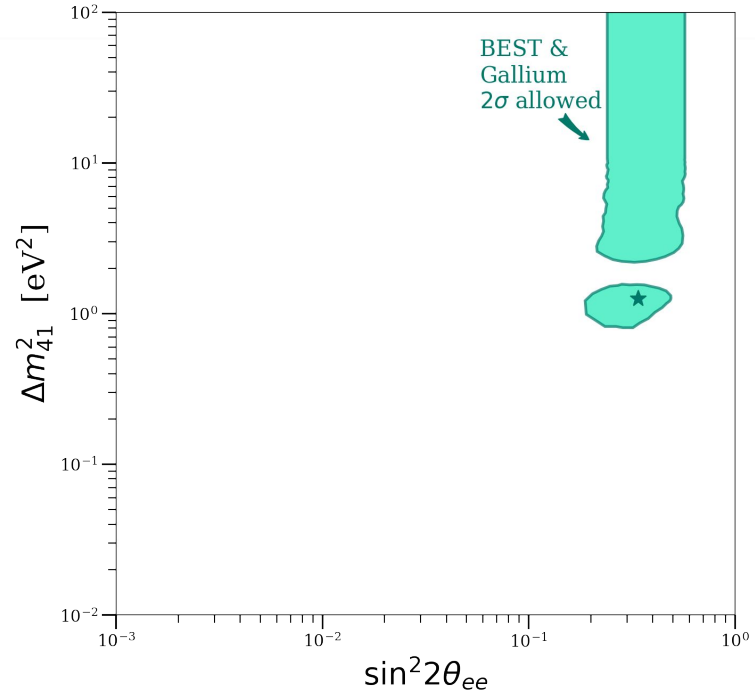


# $\nu_e$ Disappearance

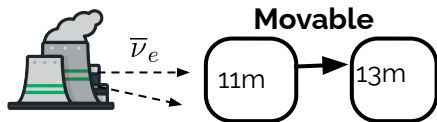
## $\nu_\mu \rightarrow \nu_e$



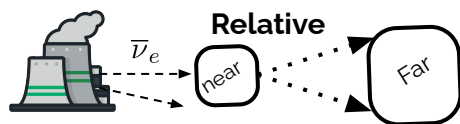
## $\nu_e \rightarrow \nu_e$



# $\bar{\nu}_e$ Disappearance

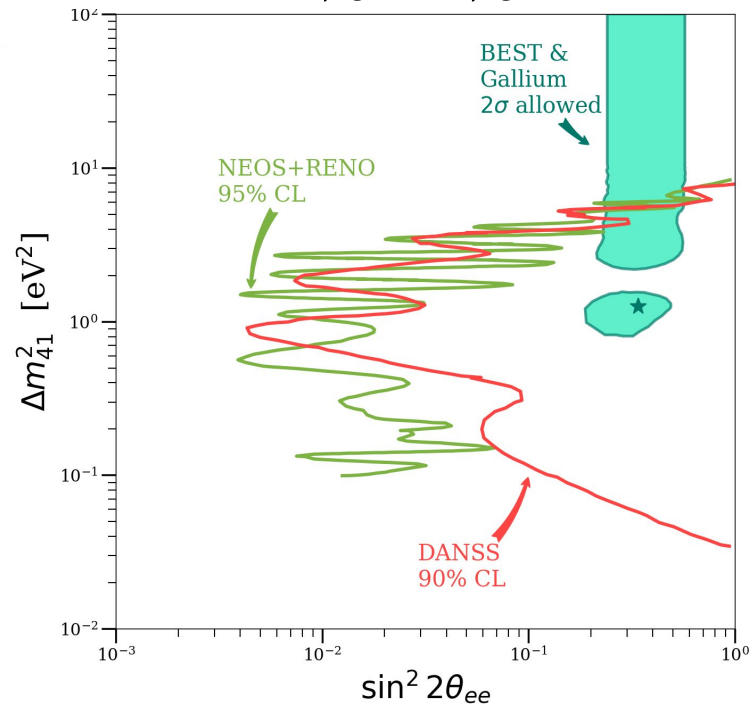


**DANSS**, Movable detector in Russia,  
[Phys.Lett.B 787\(2018\)56-63](#)



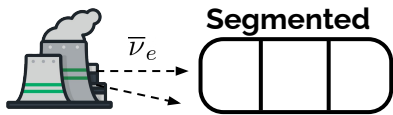
**NEOS+RENO**, relative 419 m RENO and 24m  
NEOS detectors in Korea  
[Phys. Rev. D 105, L111101 \(2022\)](#)

$\nu_e \rightarrow \nu_e$



**Need to get closer to reactors!**

# $\bar{\nu}_e$ Disappearance

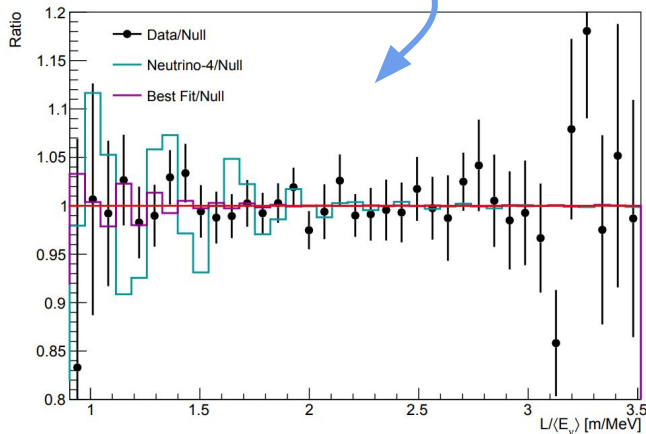


**STEREO** (ILL, Grenoble, France)

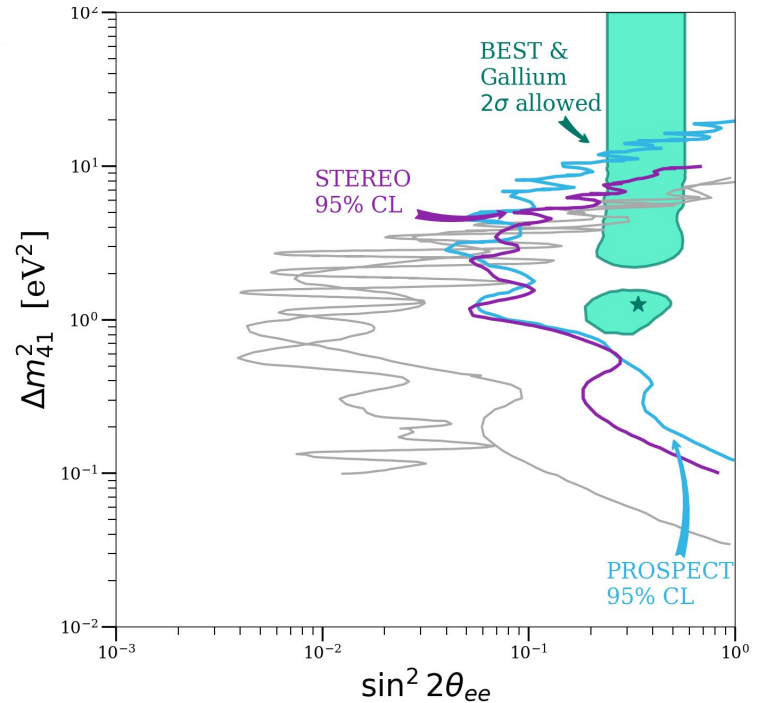
- [Nature 613, 257–261 \(2023\)](#)
- **9.4 to 11.1 m Baselines**

**PROSPECT** (Oak Ridge National Lab, USA)

- [arXiv:2406.10408 \[hep-ex\] \(2024\)](#)
- **7-9 m Baselines**



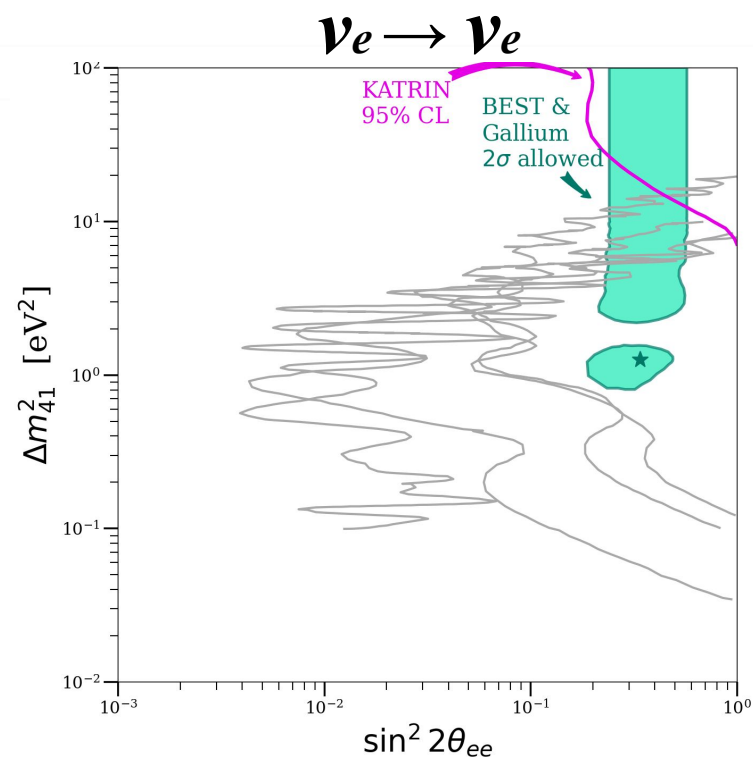
# $\nu_e \rightarrow \nu_e$



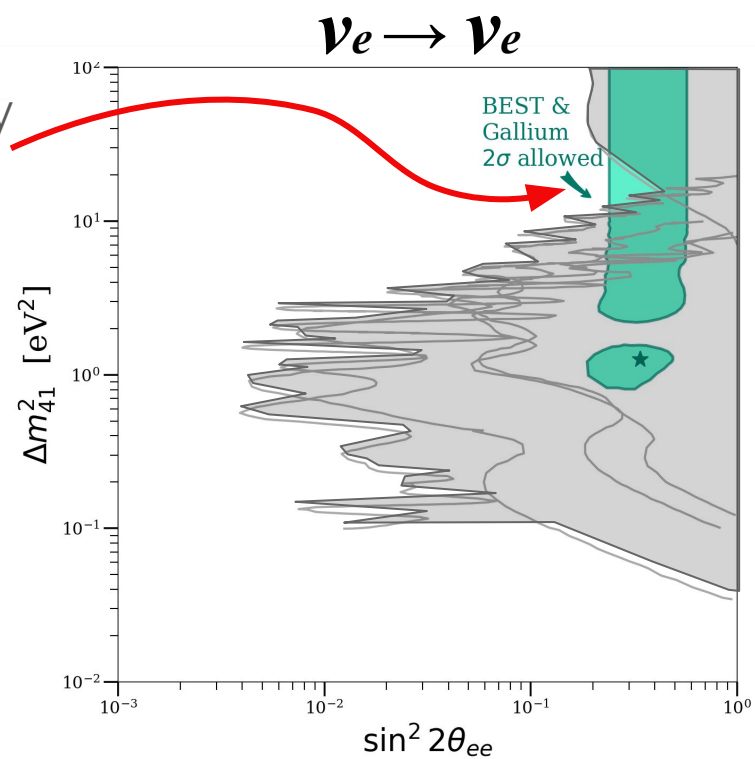
# $\nu_e$ Disappearance

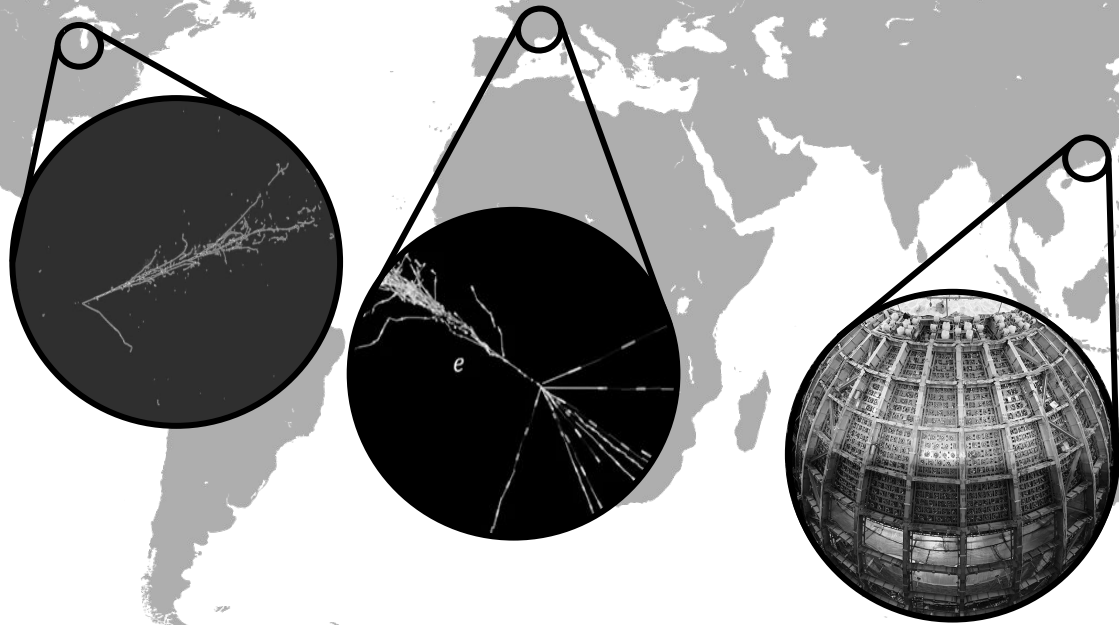
**KATRIN** in Karlsruhe Germany  
Designed to determine the **neutrino mass** from kinematic measurements of the  **$\beta$ -decay of tritium**  
[PhysRevD.105.072004 \(2022\)](#)

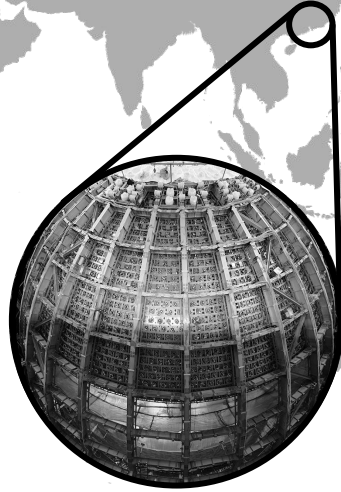
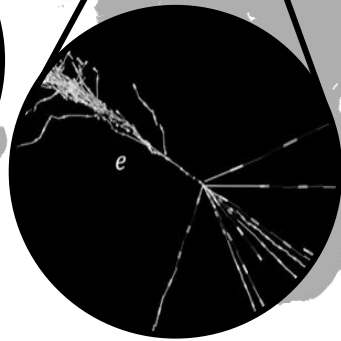
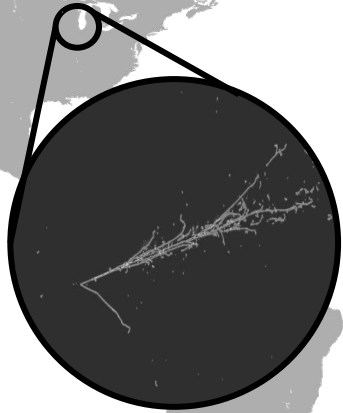
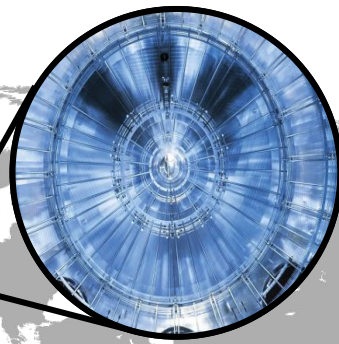
**Unique!** KATRIN is a **non-oscillatory** lab experiment sensitive to relevant solutions at large  $\Delta m^2$  values!

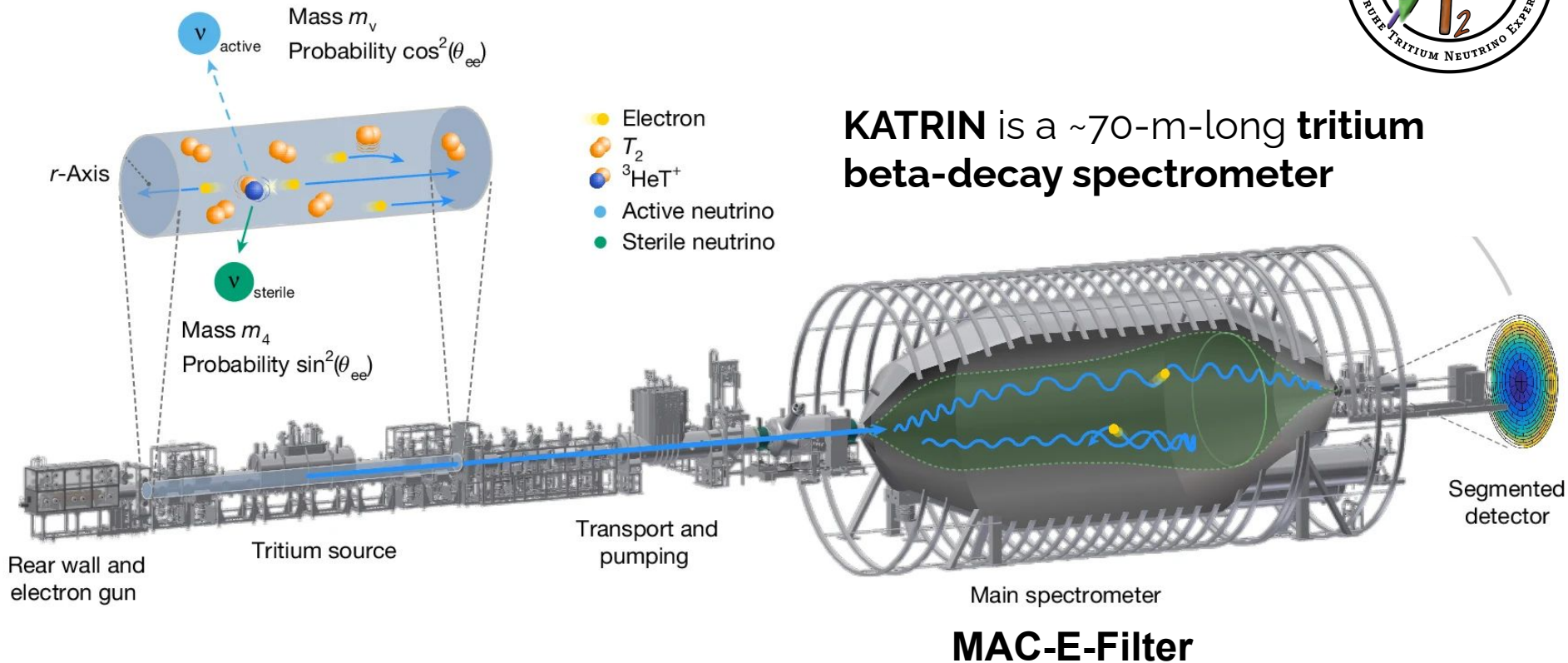


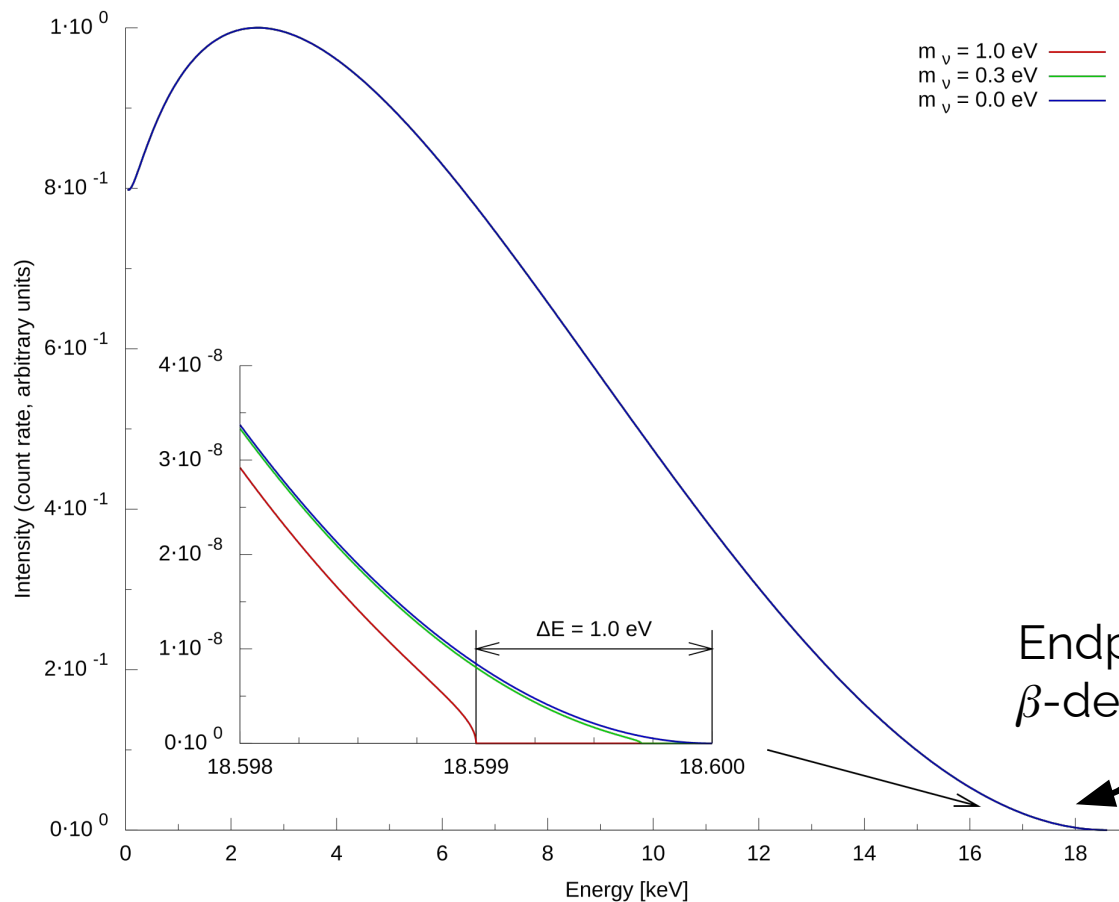
As of **PHENO 2025**, pesky  
“**Gap**” region unprobed..







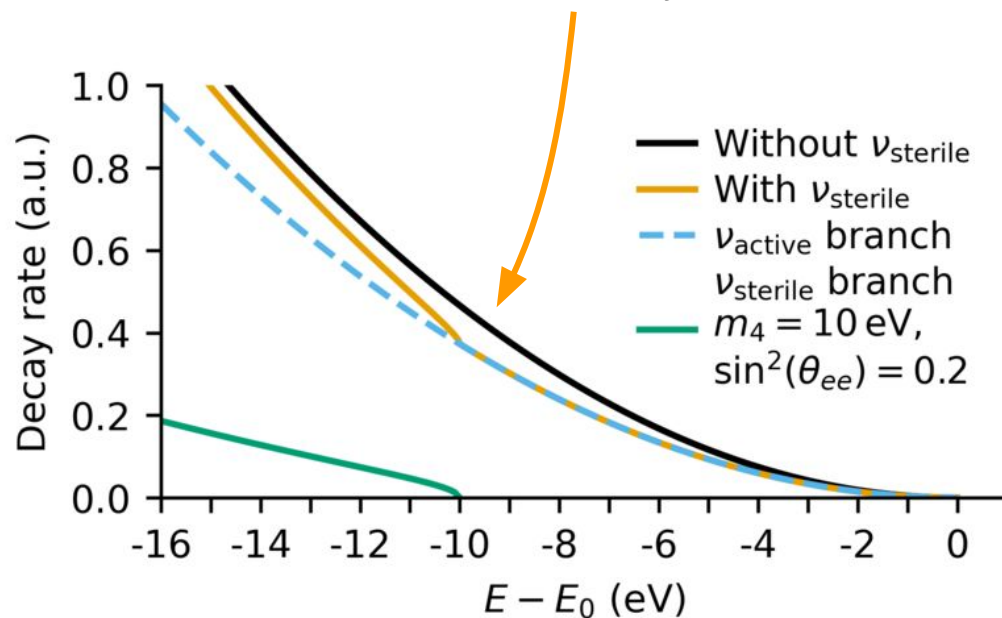


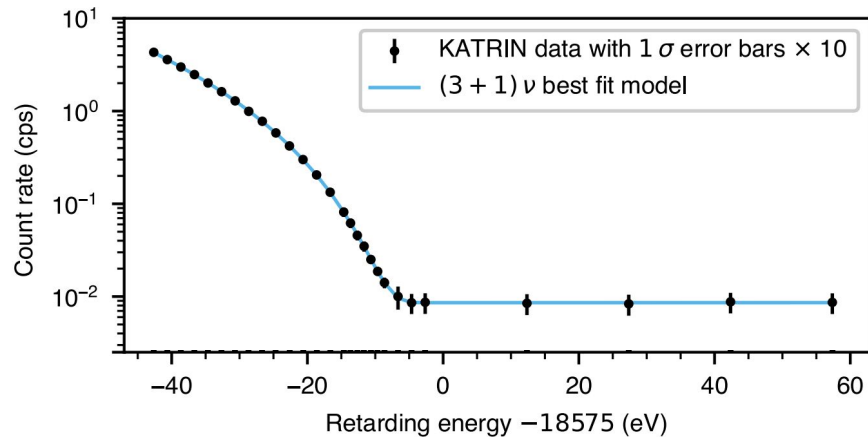


Endpoint of Tritium  $\beta$ -decay  $\sim$  **neutrino mass**



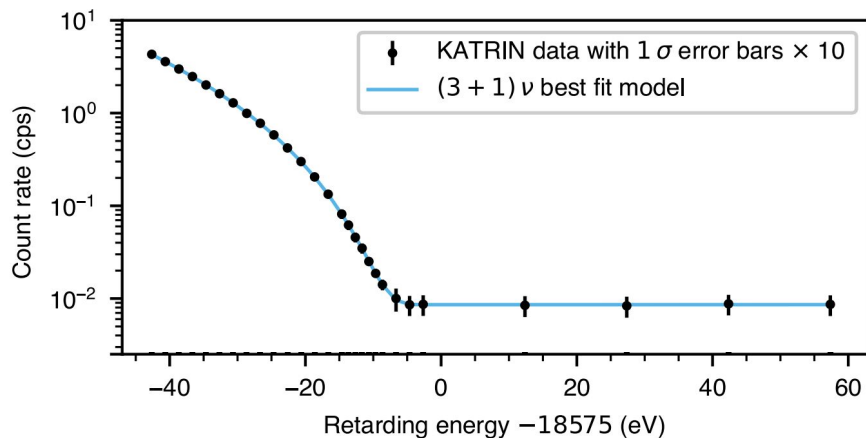
“Kink” further up from endpoint of Tritium  $\beta$ -decay ~ **Sterile neutrinos**





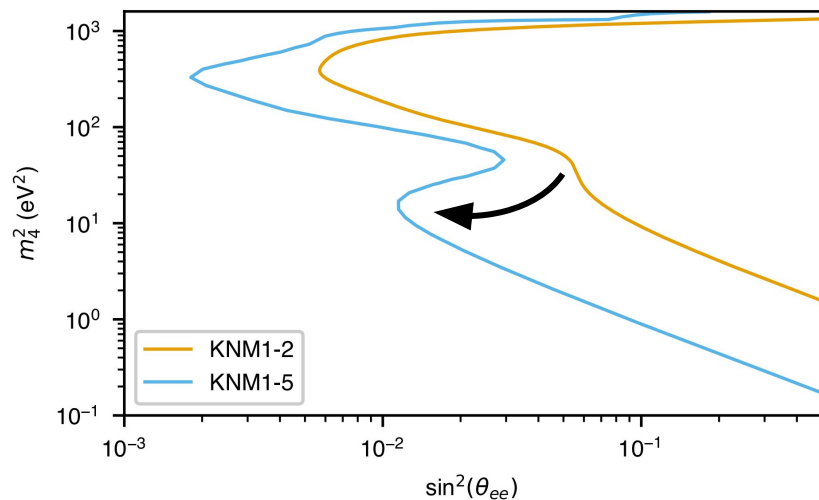
Analyzed **36 million tritium  $\beta$ -decay electrons** in a data set taken over **259 days**

- **A sixfold increase in statistics** over 2022 results
- **Substantial improvements** understanding of **systematics**.

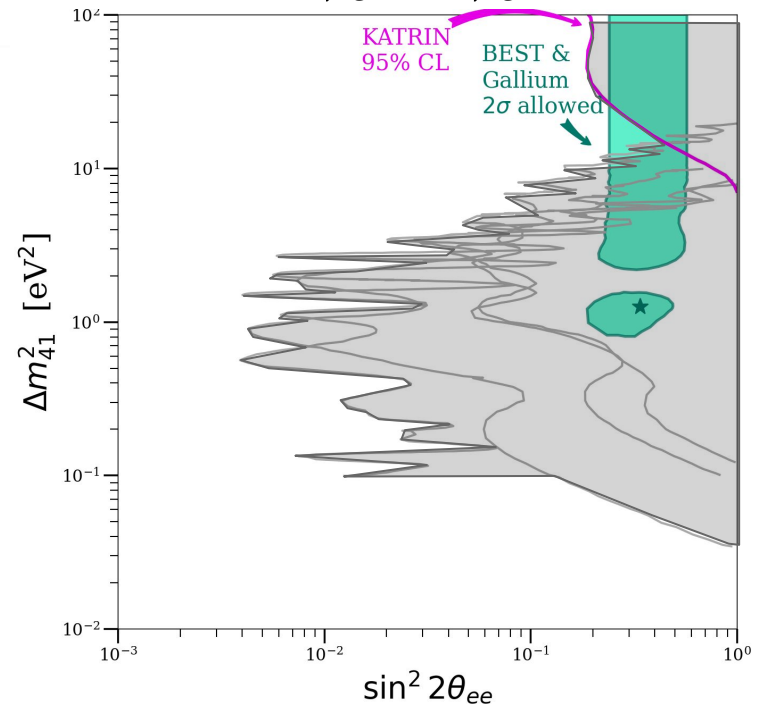


Analyzed **36 million tritium  $\beta$ -decay electrons** in a data set taken over **259 days**

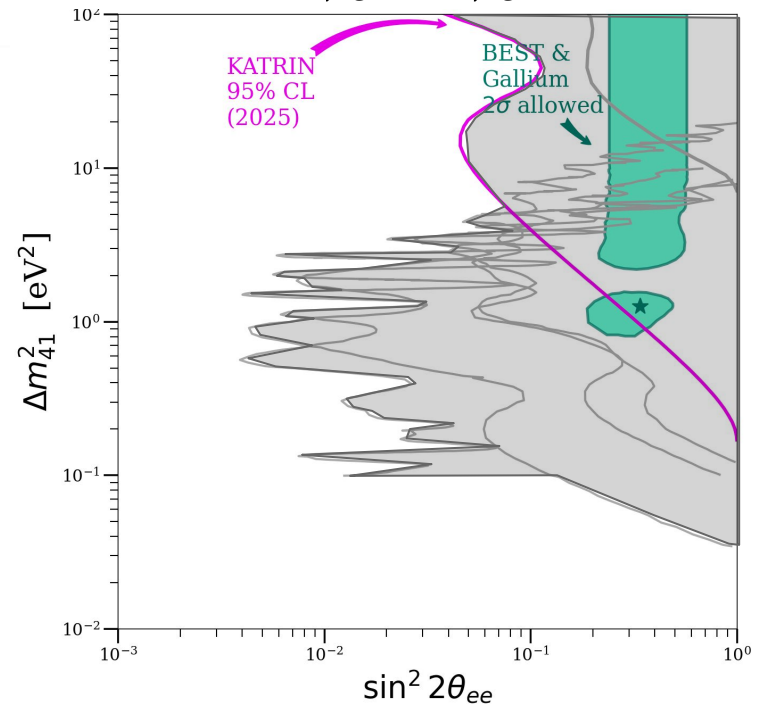
- **A sixfold increase in statistics** over 2022 results
- **Substantial improvements** understanding of **systematics**.



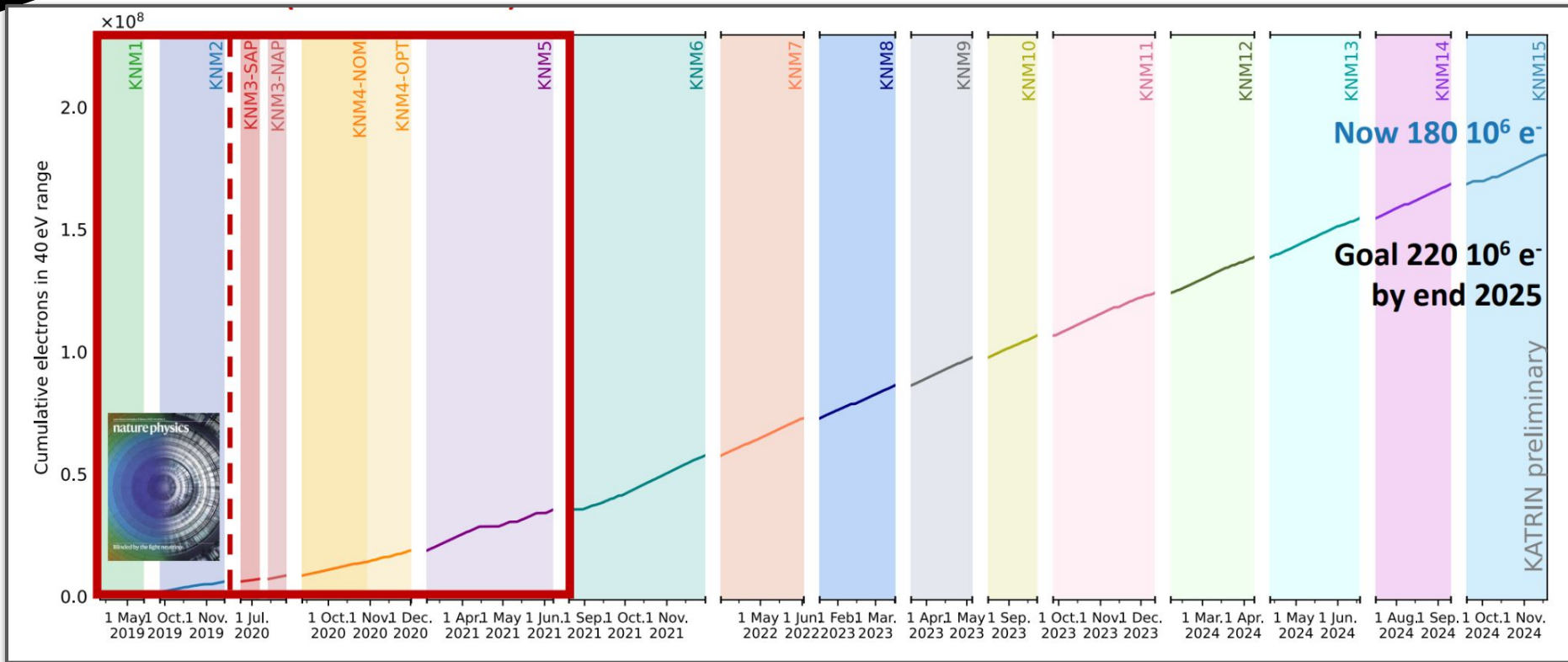
$\nu_e \rightarrow \nu_e$



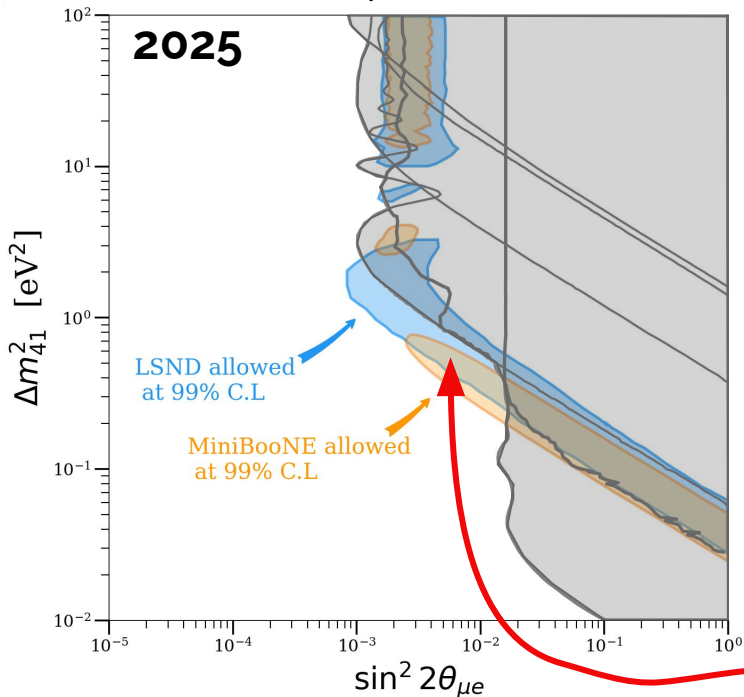
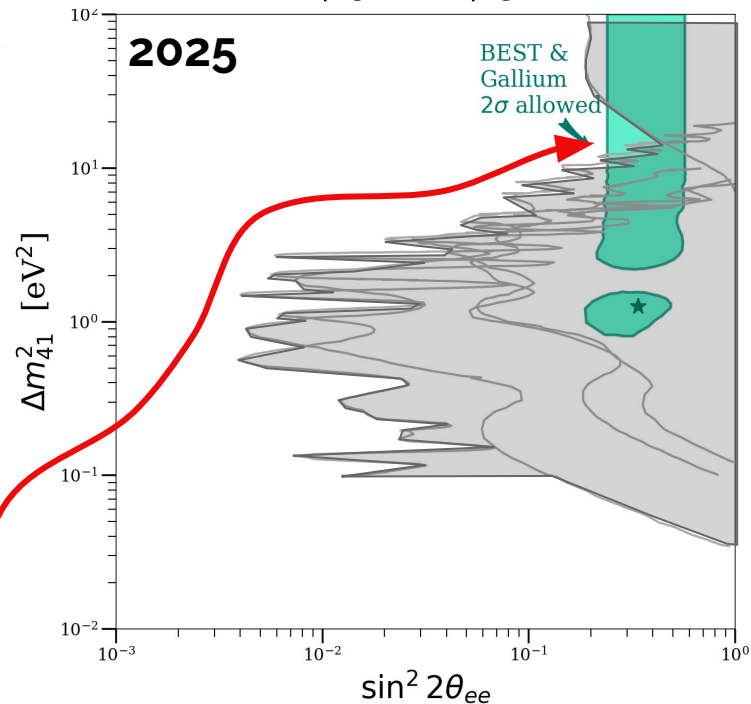
$\nu_e \rightarrow \nu_e$



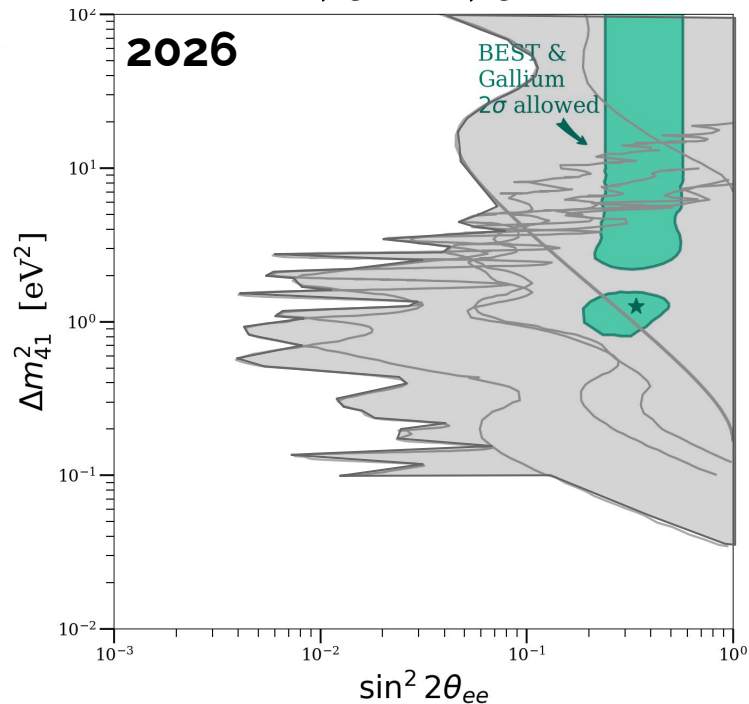
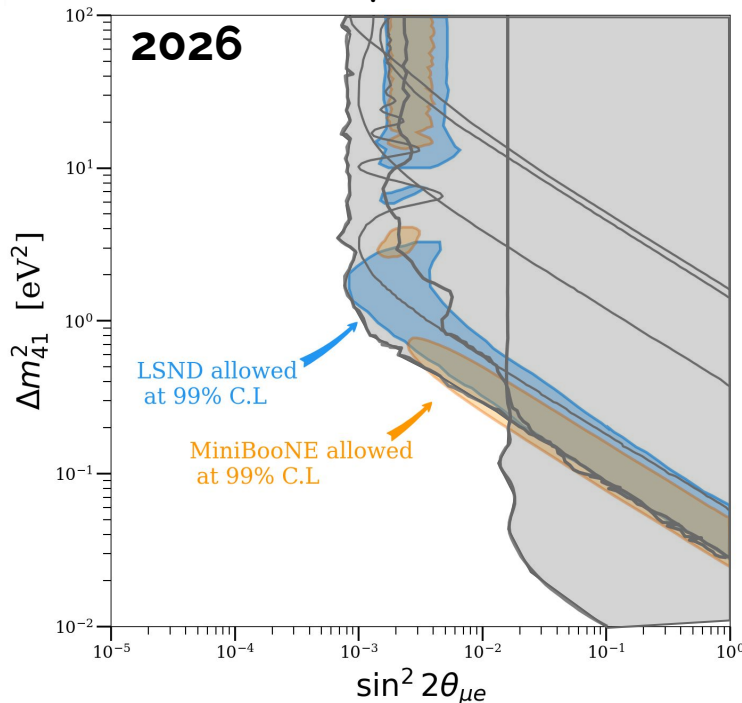
# A Lot More Data to Come



Slide from Thierry Lasserre

$\nu_\mu \rightarrow \nu_e$  $\nu_e \rightarrow \nu_e$ 

As of **PHENO 2025**  
pesky "**Gap**" regions  
unprobed..



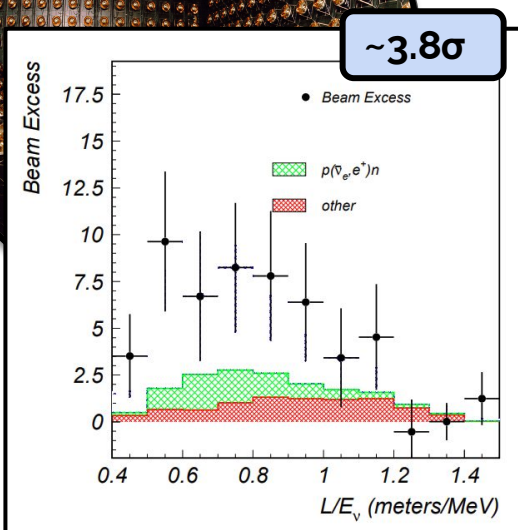
**Last ~6 months have been an impressive time** for advancing our understanding of light sterile neutrinos!

A single light eV scale sterile is **no longer a viable solution** for short-baseline anomalies

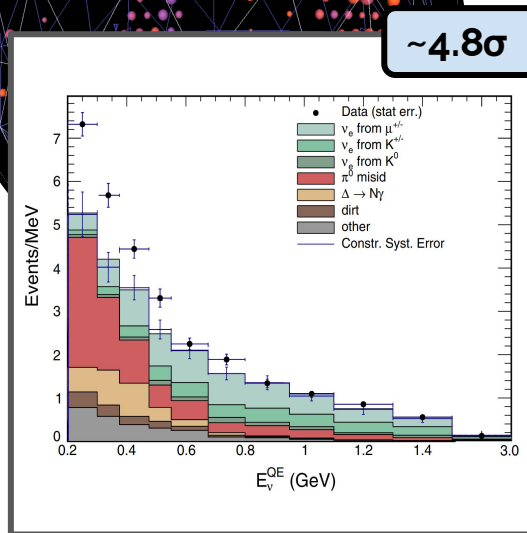
# If a model fails to explain the data *throw out the model, not the data.*

If not  $eV^2$  scale sterile neutrinos, then what?

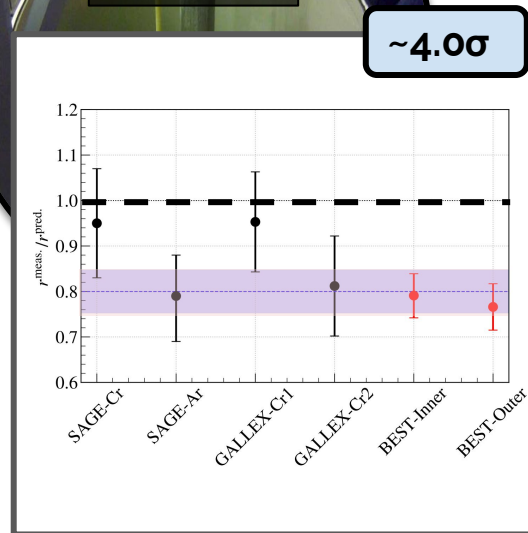
LSND  
Anomaly



MiniBooNE  
Anomaly



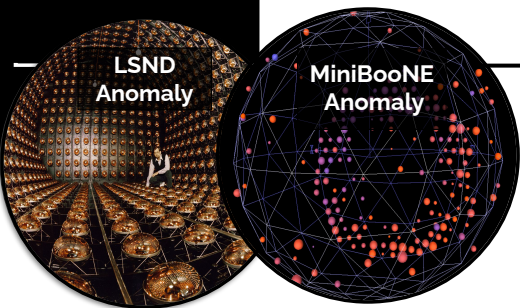
Gallium  
Anomaly



## Additions/Modifications to light sterile neutrino oscillations\* I.e 3+1 + something else

- **3+2** and **3+3** (more light steriles)
  - C. Giunti *et al* Phys.Rev.D 84 (2011) 9, 073008
- **3+1+Non-Standard interactions** (NSI)
  - J. Liao *et al* Phys.Rev.D 99 (2019) 1, 015016
- **3+1+Decay**
  - Z.Moss *et al* Phys.Rev.D 97, (2018) 055017
- **3+1+Decoherence**
  - C.A. Argüelles *et al* C.Phys. Rev. D 107 (2023), 036004
- **3+1+Resonance Matter Effects**
  - D.Alves *et al* JHEP08 (2022) 034

Lots of effort in this direction, to **keep the core 3+1 light oscillating sterile neutrino model** and modify slightly to **help reduce tension**



**BSM explanations** beyond “oscillating” steriles have grown in popularity.

Many exploiting the fact that the **MiniBooNE excess could in fact be photons or e+e- pairs** and not electrons

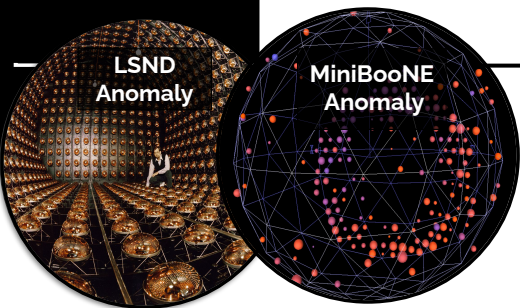
Category	Model	Signature	Anomalie		References
			LSND	MiniBooNE	
Flavor transitions Secs. 3.1.1-3.1.3, 3.1.5	(3+1) oscillations	$\nu$ oscillations	✓	✓	Reviews and global fits [93, 103, 105, 106]
	(3+1) w/ invisible sterile decay	oscillations w/ $\nu_4$ invisible decay	✓	✓	[151, 155]
	(3+1) w/ sterile decay	$\nu_4 \rightarrow \phi \nu_e$	✓	✓	[159–162, 270]
Matter effects Secs. 3.1.4, 3.1.7	(3+1) w/ anomalous matter effects	$\nu_\mu \rightarrow \nu_e$ via matter effects	✓	✓	[143, 147, 271–273]
	(3+1) w/ quasi-sterile neutrinos	$\nu_\mu \rightarrow \nu_e$ w/ resonant $\nu_s$ matter effects	✓	✓	[148]
Flavor violation Sec. 3.1.6	Lepton-flavor-violating $\mu$ decays	$\mu \rightarrow e \nu_\mu \nu_e$	✓	✗	[174, 175, 274]
	neutrino-flavor-changing bremsstrahlung	$\nu_\mu A \rightarrow \nu_e A$	✓	✓	[275]
Decays in flight Sec. 3.2.3	Transition magnetic mom., heavy $\nu$ decay	$N \rightarrow \nu \gamma$	✗	✓	[207]
	Dark sector heavy neutrino decay	$N \rightarrow e^+ \nu(X)$ $N \rightarrow \nu(X)$	✗	✓	[208]
Neutrino Scattering Secs. 3.2.1, 3.2.2	neutrino-induced upscattering	$\nu \rightarrow \nu e^+ e^-$ $N \rightarrow \nu e^+ e^-$	✓	✓	[205, 206, 209–216]
	neutrino dipole upscattering	$\nu \rightarrow \nu e^+ e^-$	✓	✓	[40, 185, 187, 188, 190, 193, 233, 276]
Dark Matter Scattering Sec. 3.2.4	dark particle-induced upscattering	$e^- e^- \rightarrow e^- e^-$	✗	✓	[217]
	dark particle-induced inverse Primakoff	$\gamma \rightarrow \gamma$	✓	✓	[217]

**Signals**

- Electron
- Photon
- Di-Photon
- e+e-

**30+ dark-sector models in last few years with Incredibly rich and varied phenomenology**

Table modified from Snowmass [White Paper](#) on Light Sterile Neutrino Searches and Related Phenomenology



**BSM explanations** beyond “oscillating” steriles have grown in popularity.

Many exploiting the fact that the **MiniBooNE excess could in fact be photons or  $e^+e^-$  pairs** and not electrons

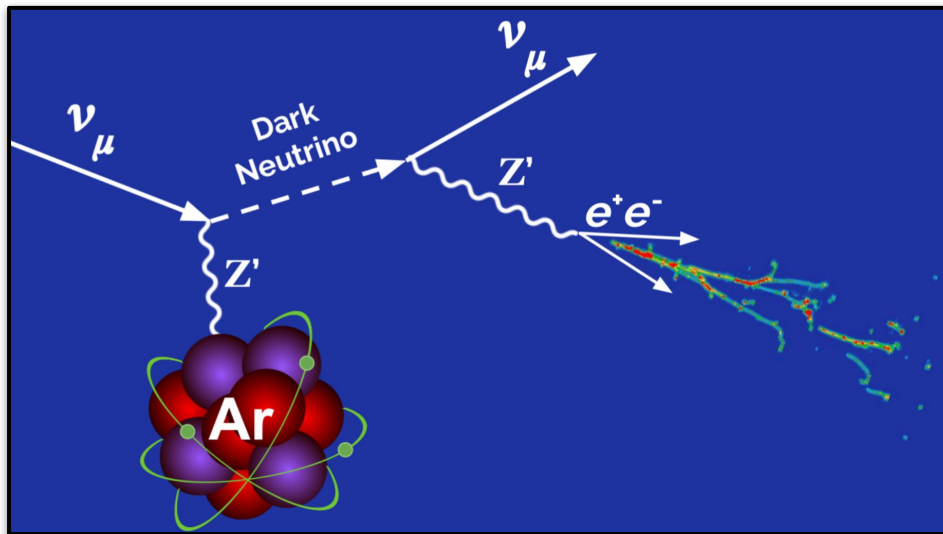
	Model	Signature	Anomalie		References
			LSND	MiniBooNE	
Flavor Secs. 3.1.1, 3.1.2	(3+1) oscillations	$e^-$ oscillations	✓	✓	Reviews and global fits [93, 103, 105, 106]
	(3+1) w/ invisible sterile decay	oscillations w/ $\nu_4$ invisible decay	✓	✓	[151, 155]
	w/ sterile decay	$\nu_4 \rightarrow \phi \nu_e$	✓	✓	[159–162, 270]
Matter effects Secs. 3.1.4, 3.1.7	w/ anomalous matter effects	$\nu_\mu \rightarrow \nu_e$ via matter effects	✓	✓	[143, 147, 271–273]
	quasi-sterile $\nu_s$ resonant $\nu_s$ matter effects	$\nu_\mu \rightarrow \nu_e$ w/ resonant $\nu_s$ matter effects	✓	✓	[148]
Flavor violation Sec. 3.1.6	Lepton number-violating $\mu$ decays	$\nu_\mu A \rightarrow \nu_e A + e^-$	✓	✗	[174, 175, 274]
	neutrino-flavor-changing bremsstrahlung	$\nu_\mu A \rightarrow \nu_e A + \gamma$	✓	✓	[275]
Decays in flight Sec. 3.2.3	Transition magnetic mom., heavy $\nu$ decay	$N \rightarrow \nu \gamma$	✗	✓	[207]
	Dark sector heavy neutrino decay	$N \rightarrow \nu e^+ e^-$ $N \rightarrow \nu \gamma \gamma$	✗	✓	[208]
Neutrino Scattering Secs. 3.1, 3.2.2	neutrino-induced upscattering	$e^- \rightarrow \nu e^- \gamma$ $N \rightarrow \nu e^- \gamma$	✓	✓	[205, 206, 209–216]
	neutrino dipole upscattering	$e^- \rightarrow \nu e^- \gamma$	✓	✓	[40, 185, 187, 188, 190, 193, 233, 276]
Dark Matter Scattering Sec. 3.2.4	dark particle-induced upscattering	$e^- \rightarrow \nu e^- \gamma$	✗	✓	[217]
	dark particle-induced inverse Primakoff	$\gamma \rightarrow \nu e^- e^+$	✓	✓	[217]

### Signals

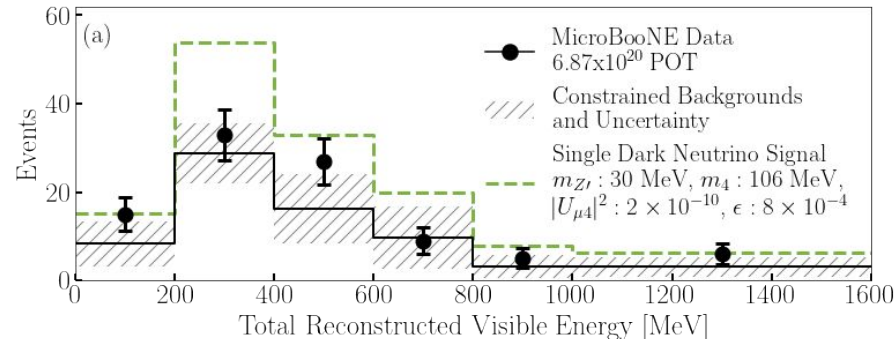
- $e^-$  Electron
- $\gamma$  Photon
- $\gamma \gamma$  Di-Photon
- $e^+ e^-$   $e^+e^-$

**30+ dark-sector models in last few years with Incredibly rich and varied phenomenology**

Table modified from Snowmass [White Paper](#) on Light Sterile Neutrino Searches and Related Phenomenology

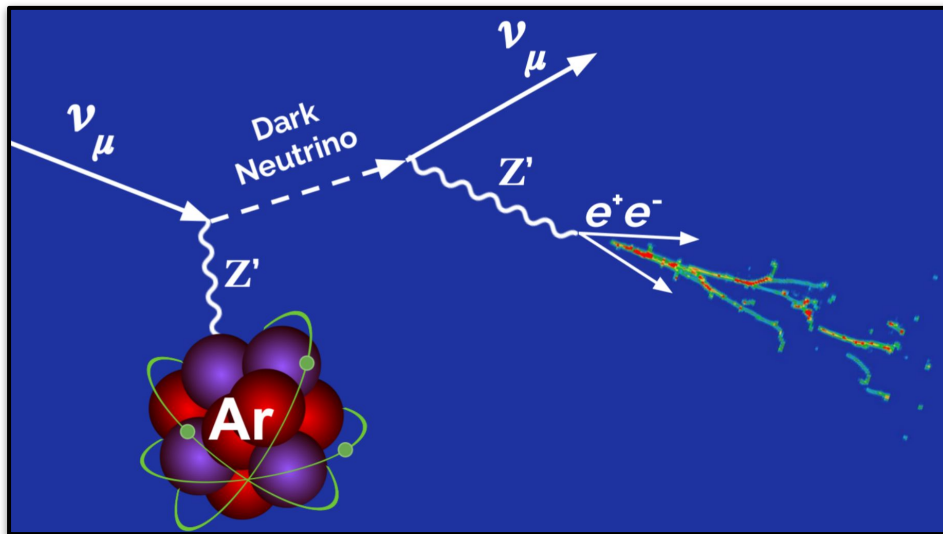


Search for forward going  $e^+e^-$  pairs from neutrino induced upscattering producing an unstable HNL (Dark neutrino)

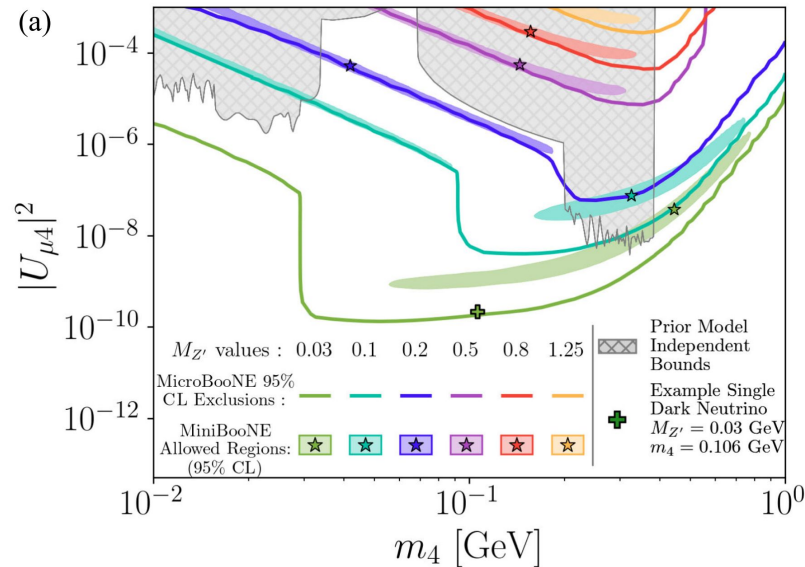


Bkg Prediction :  $69.7 \pm 17.3$  events

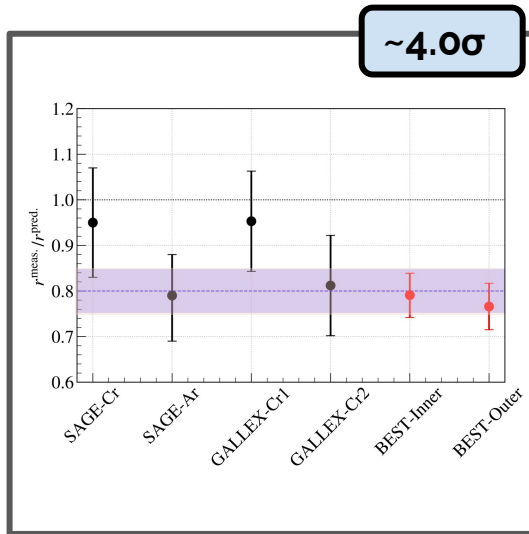
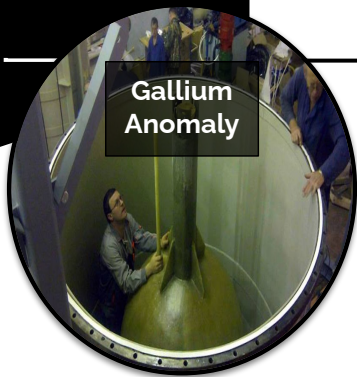
Observed Data : **95** events



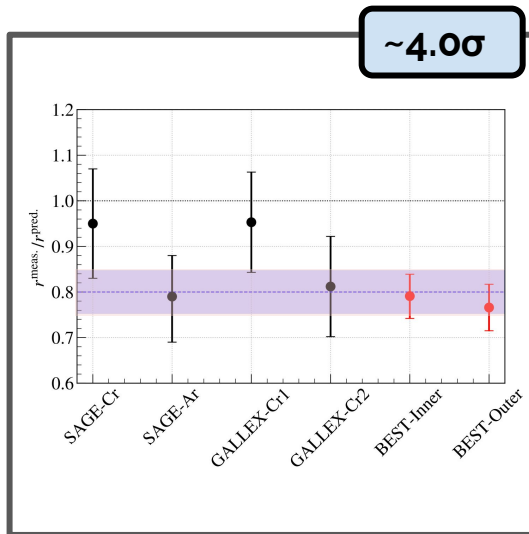
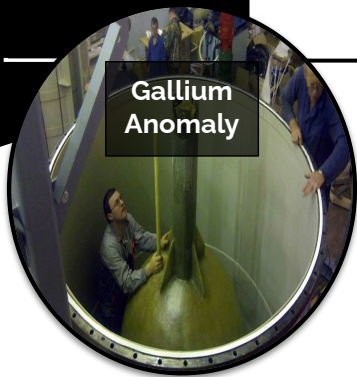
Search for forward going  $e^+e^-$  pairs from neutrino induced upscattering producing an unstable HNL (Dark neutrino)



The **world's first direct limits** on these dark sector models and, at the 95% confidence level, **excludes the majority of the parameter space viable as a solution to the MiniBooNE anomaly**



A **~20% disappearance effect** in electron neutrinos is very difficult to engineer in BSM models without breaking many LEP and EW bounds

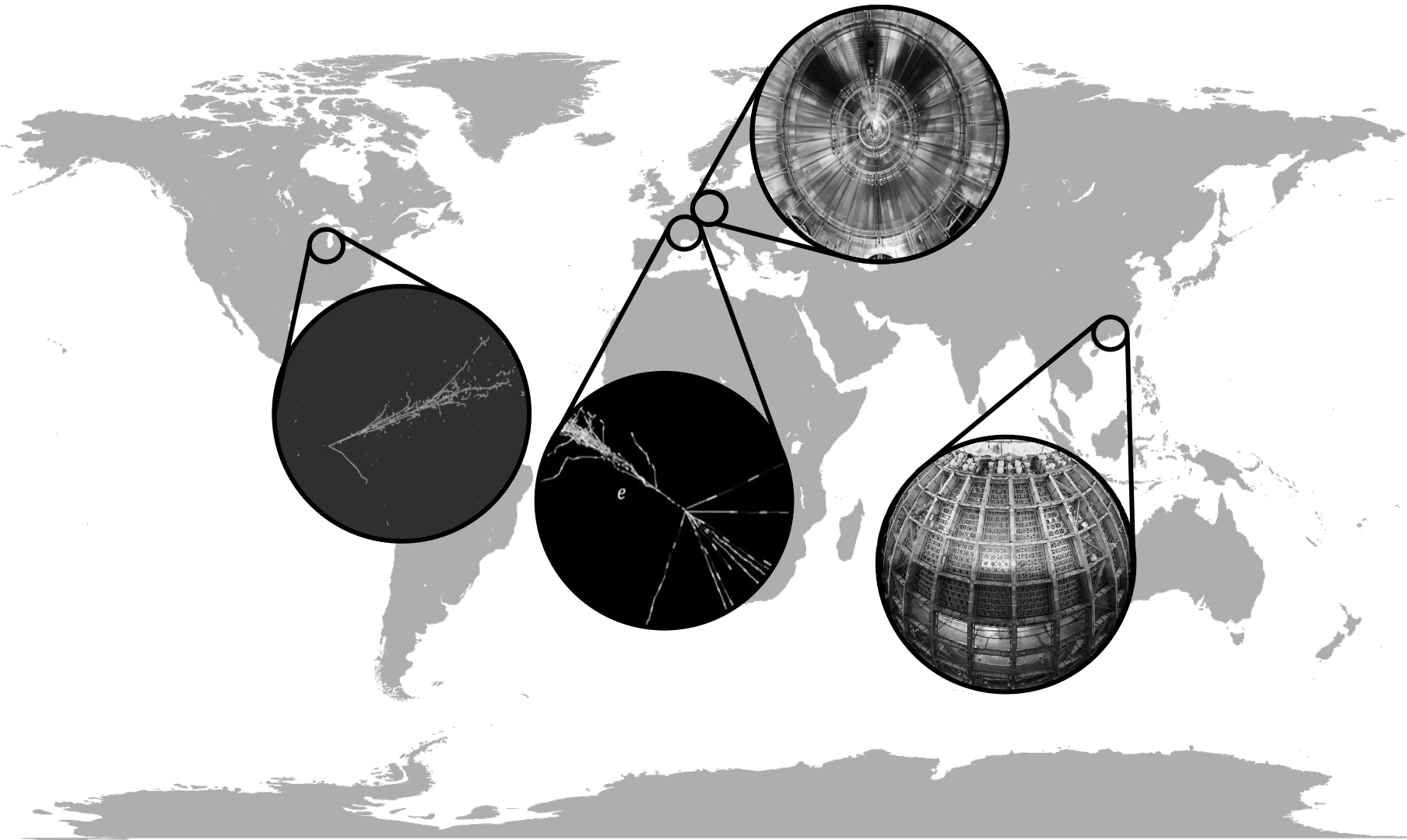


Recent paper

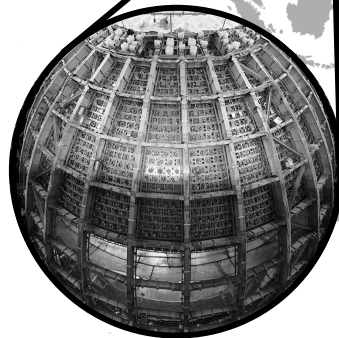
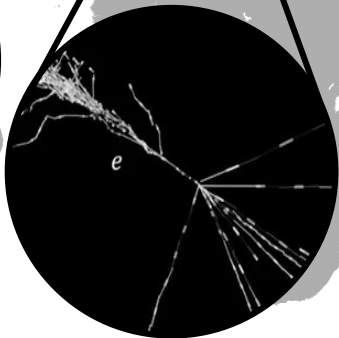
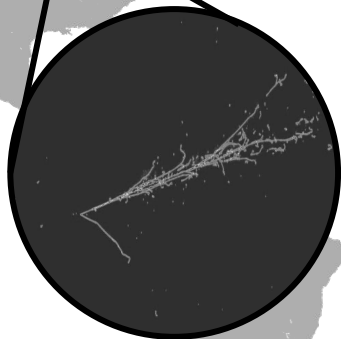
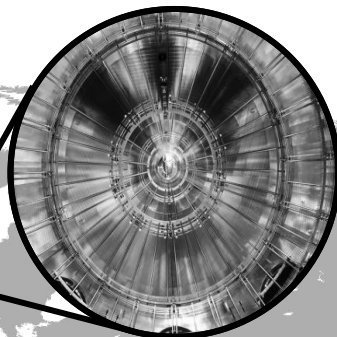
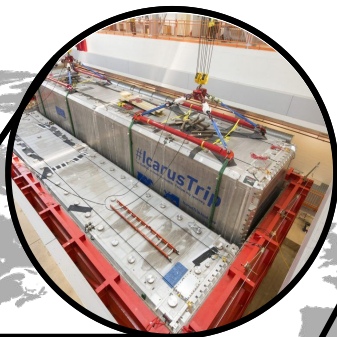
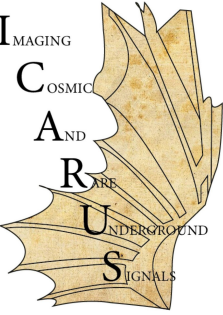
**Cadeddu et al.**, "A possible solution to the gallium anomaly moving beyond the leptonic wave function factorization" *arXiv:2512.20560 [hep-ph]*

- Questions the **detailed-balance link between  $\nu$ - $^{71}\text{Ga}$  capture and  $^{71}\text{Ge}$  electron capture**, valid only under factorization of nuclear and leptonic matrix elements (i.e. lepton wave functions constant over the nuclear volume).
- Drop that factorization and the cross-section *potentially* falls by up to ~20%

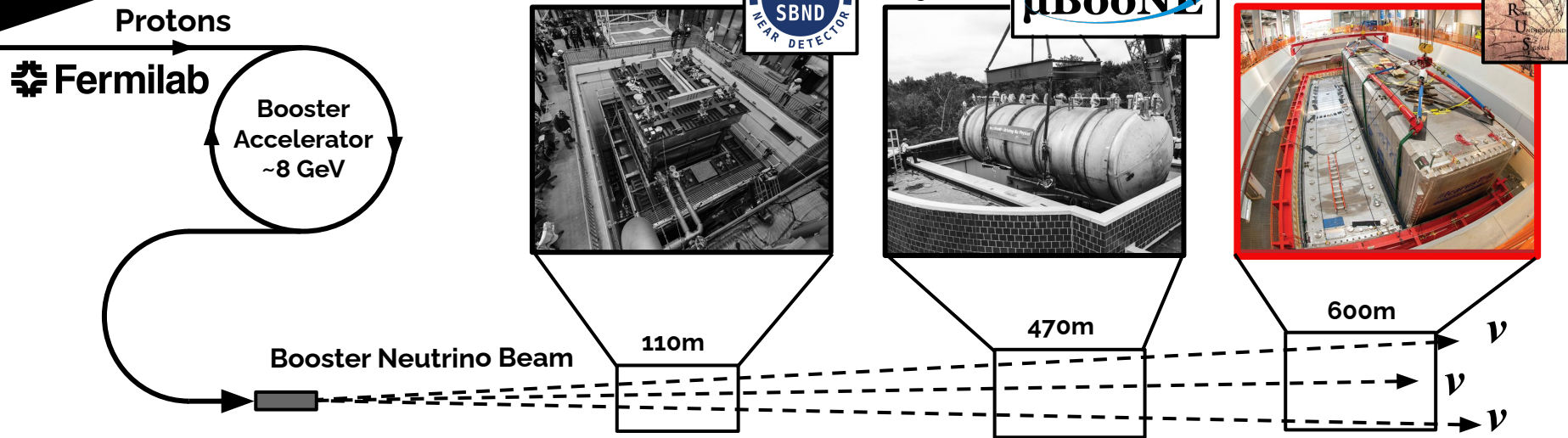
A **~20% disappearance effect** in electron neutrinos is very difficult to engineer in BSM models without breaking many LEP and EW bounds

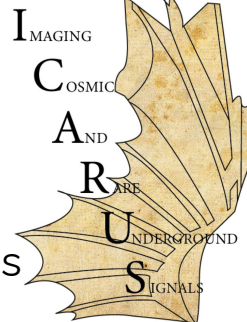


IMAGING  
COSMIC  
AND  
RARE  
UNDERGROUND  
SIGNALS

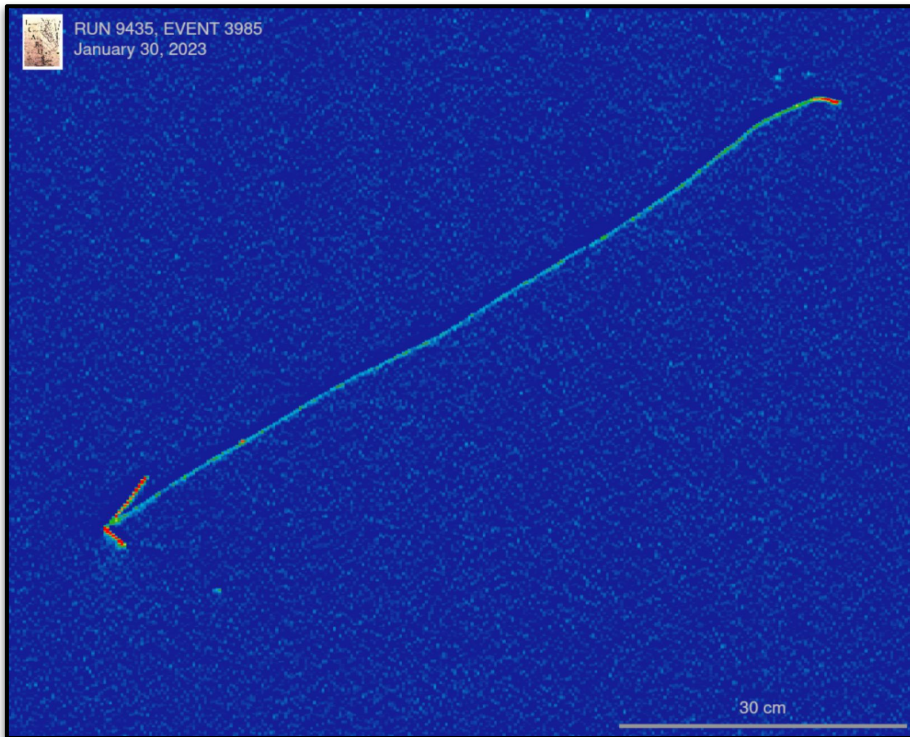


# ICARUS Far Detector



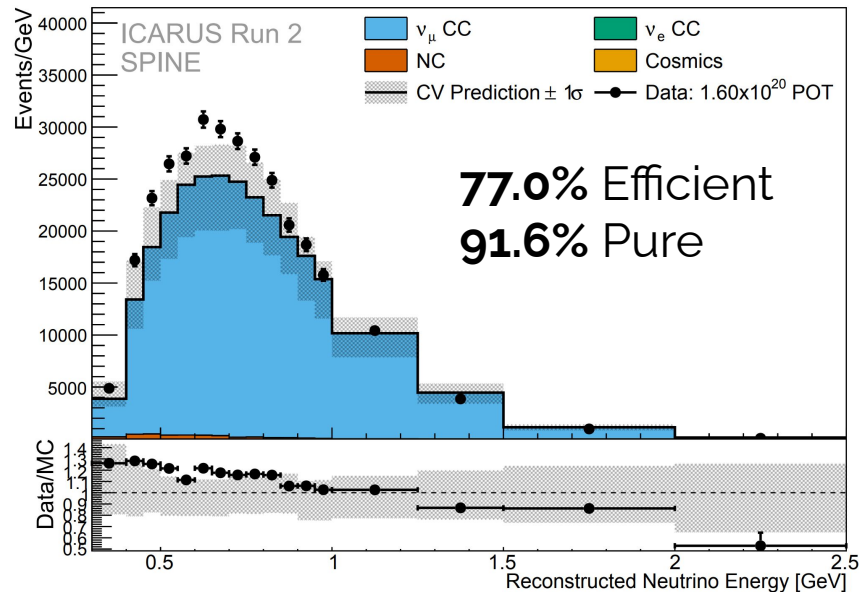


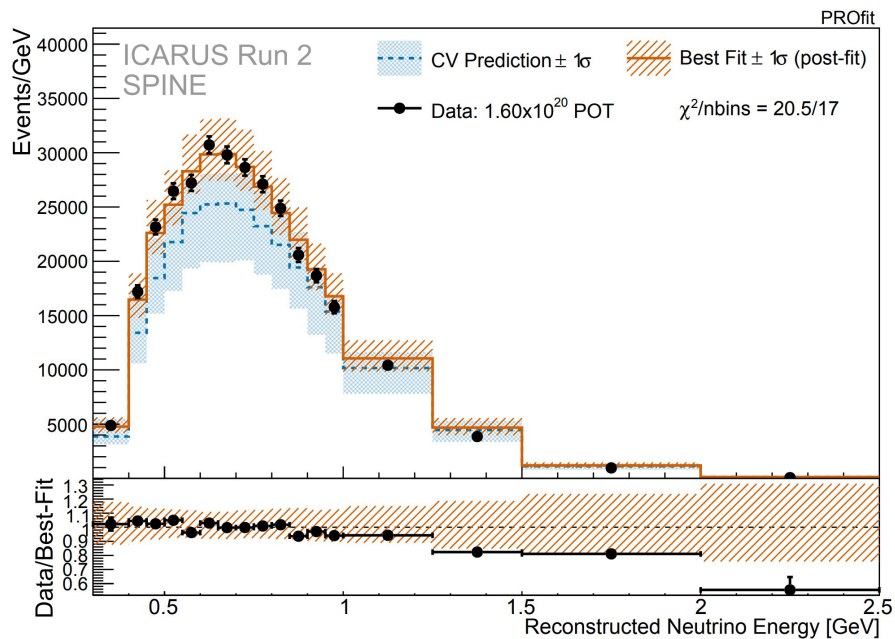
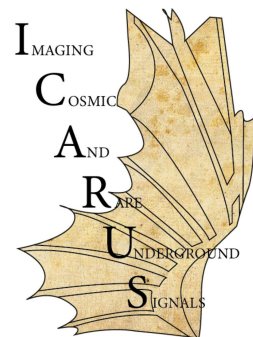
## Search for muon neutrino disappearance



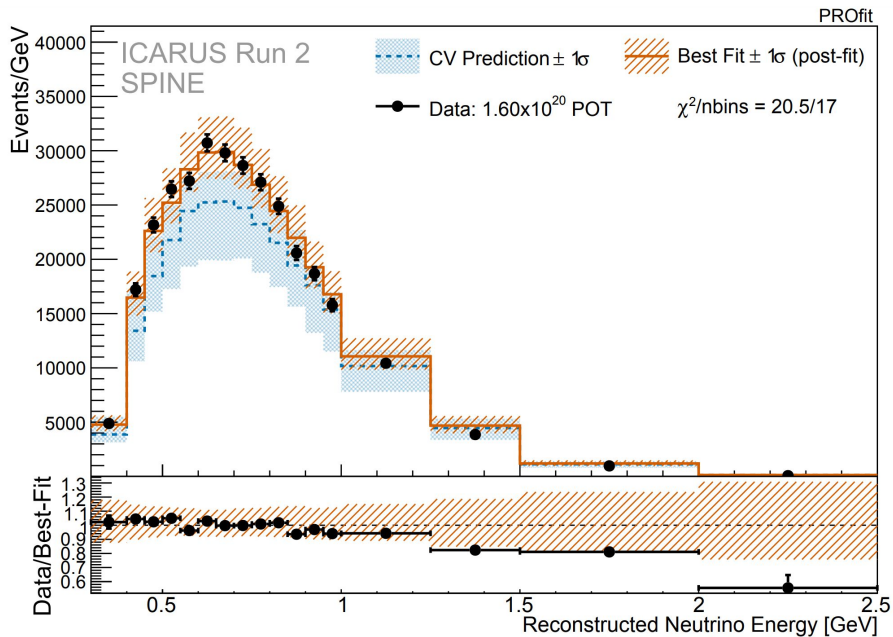
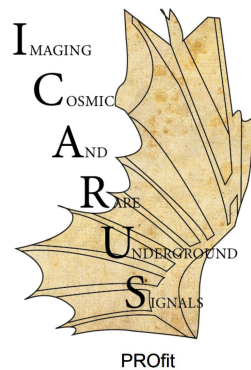
Two separate, independent analyses

- Pandora and **SPINE** (shown)



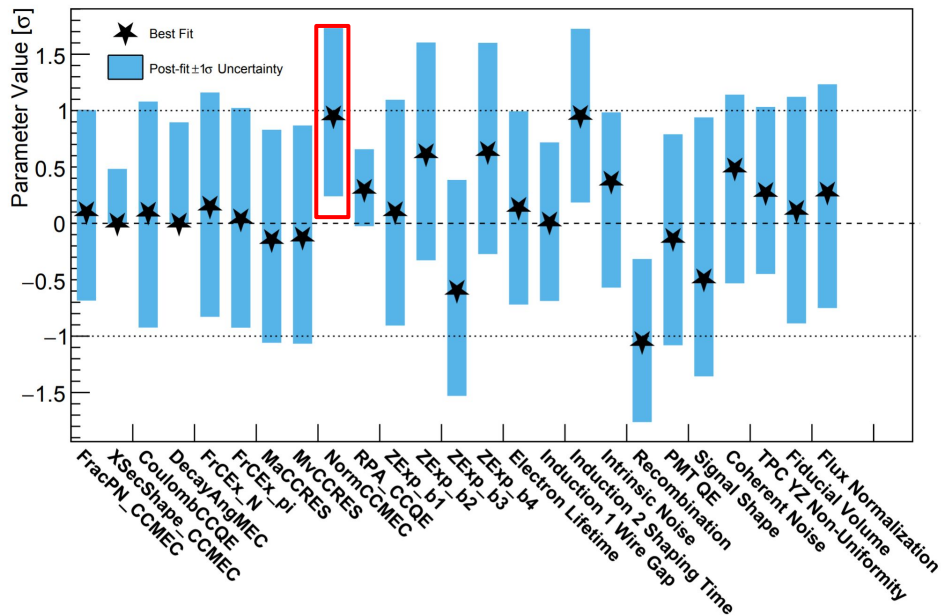


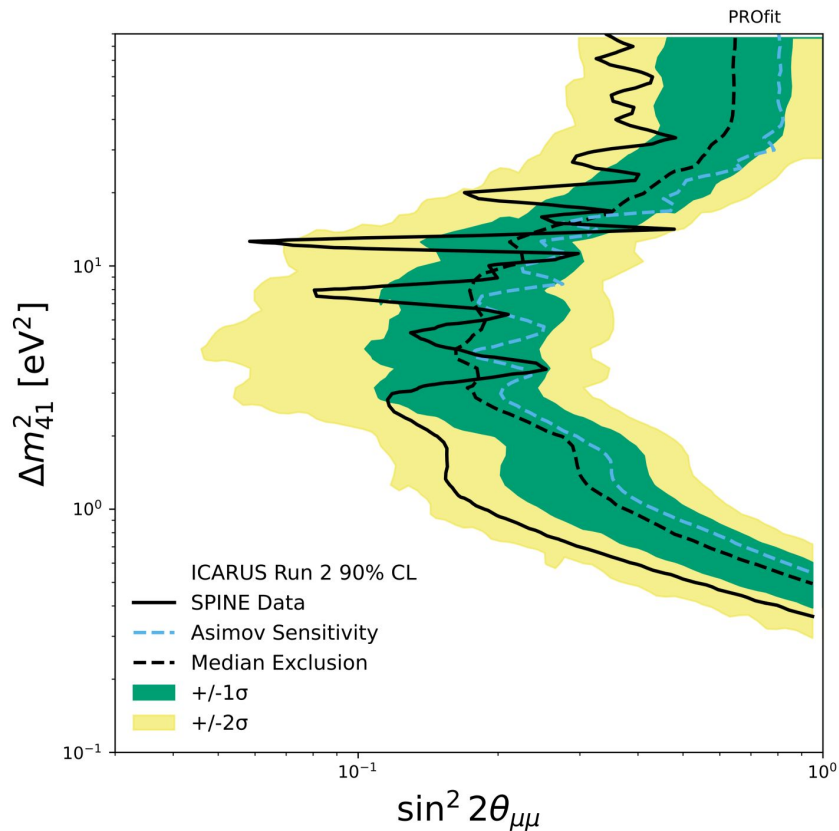
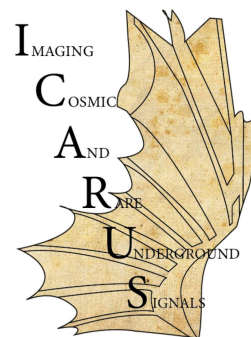
Fit to Background Systematics only



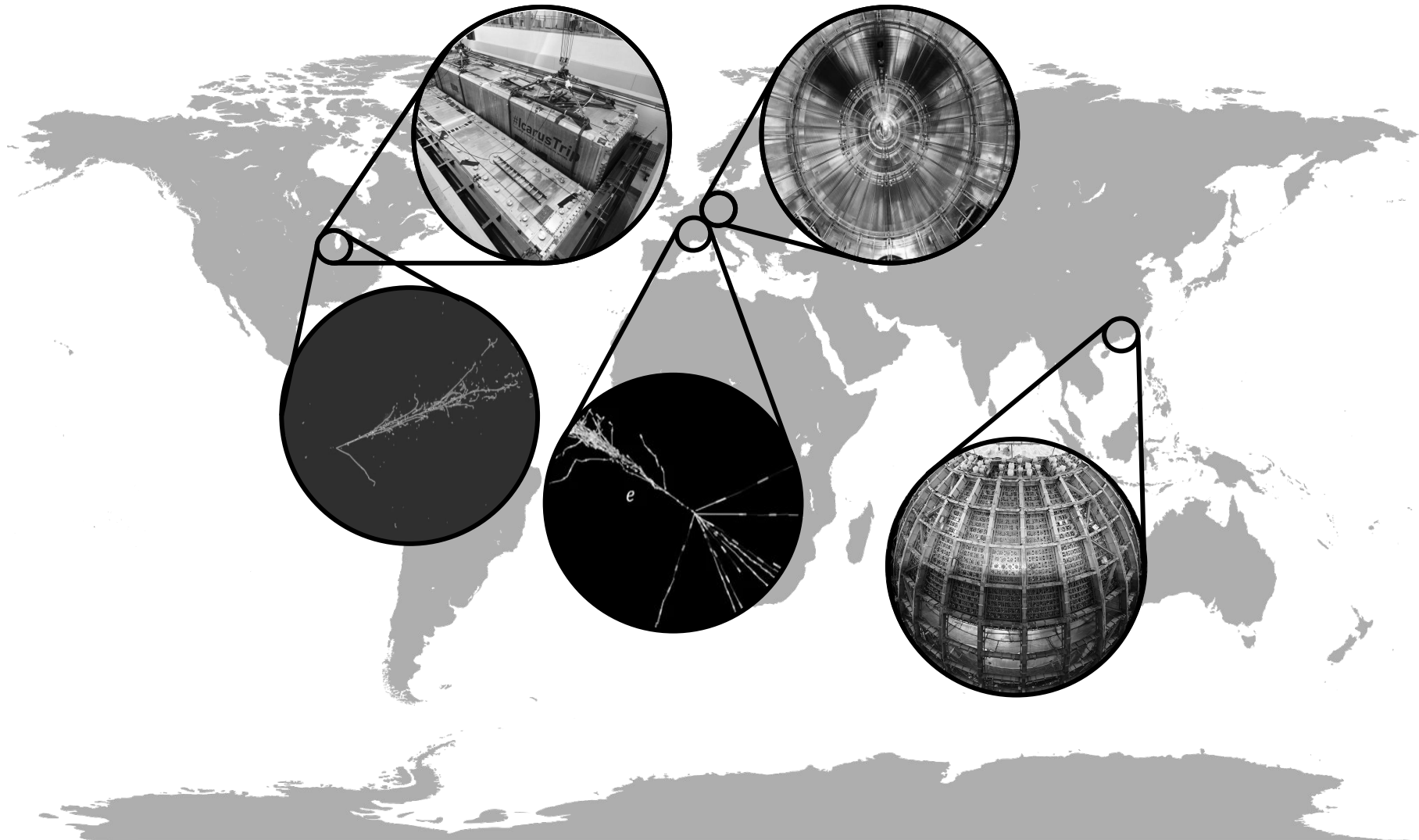
Fit to Background Systematics only

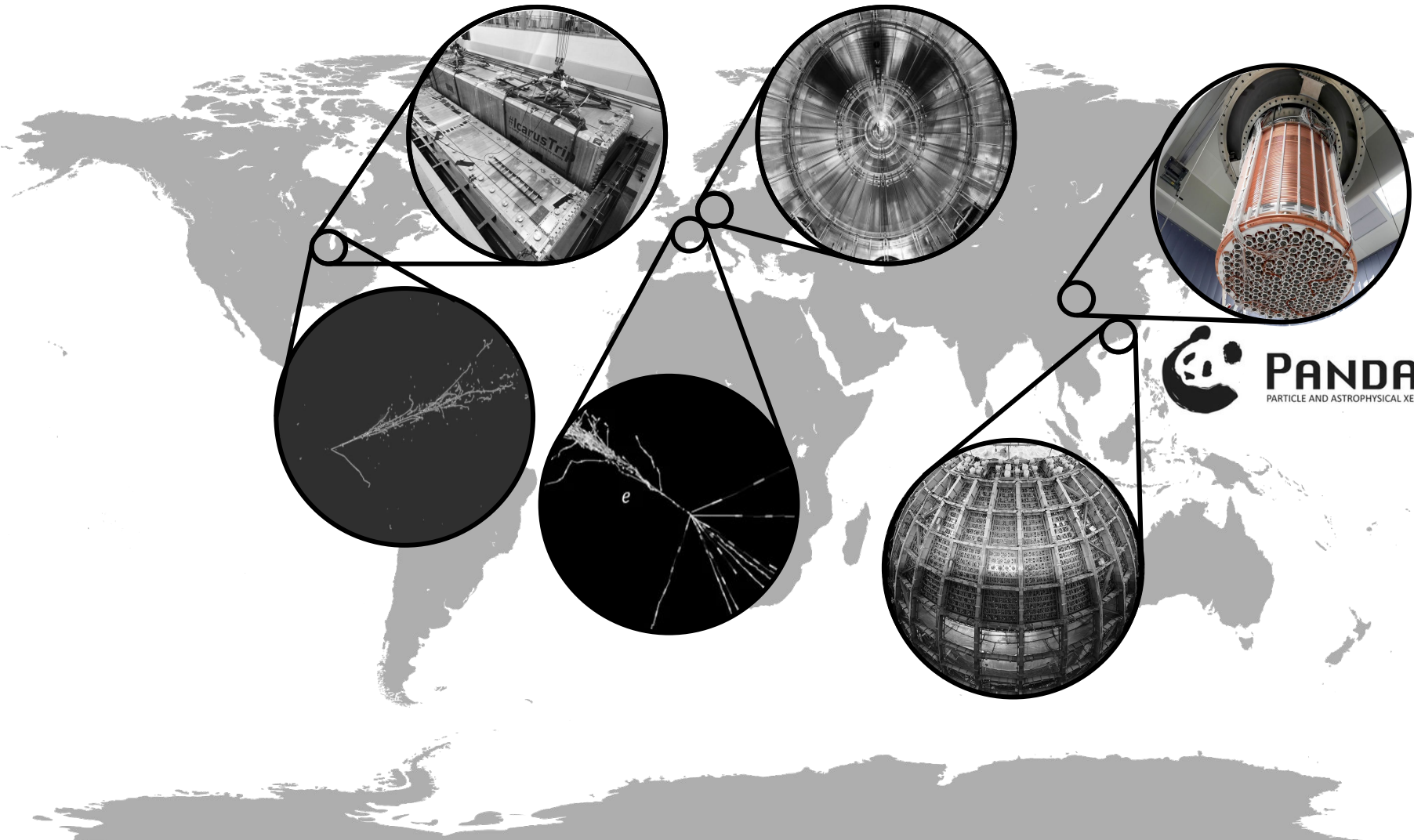
## CC MEC Normalization: $+1\sigma$



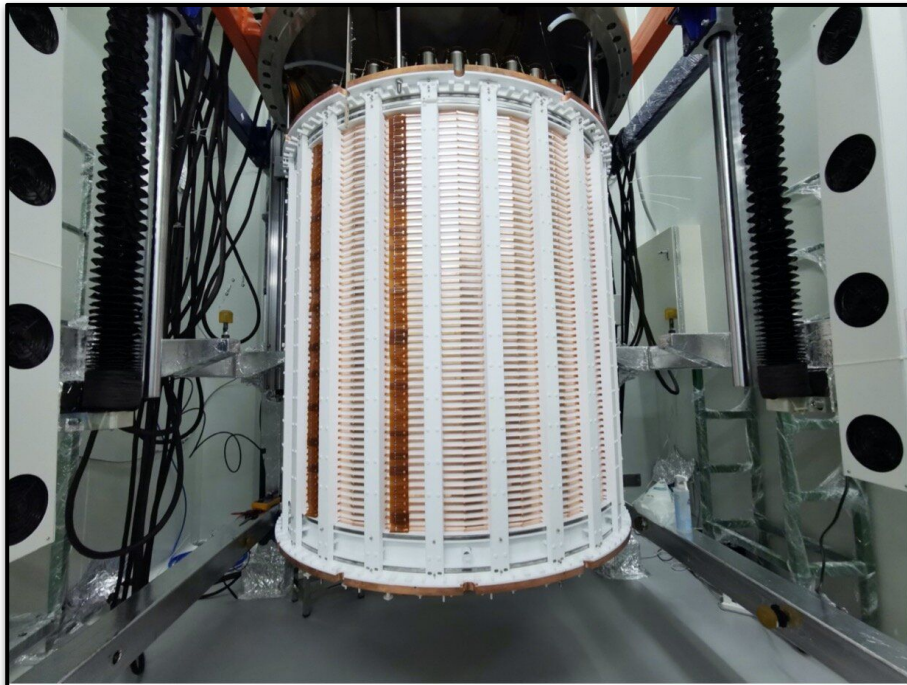


- Full end-to-end oscillation analysis exercised
- **No evidence for  $\nu_\mu$  disappearance.**
- Extremely Systematics Limited
  - Joint fit with **SBND** near detector will **reduce systematics** from **~20% to few percent level**

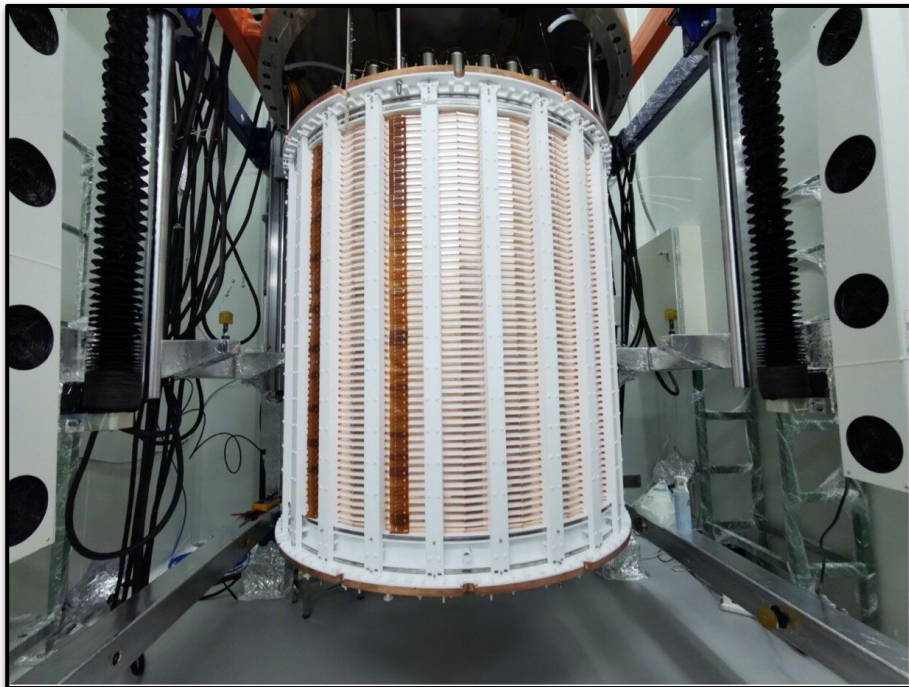




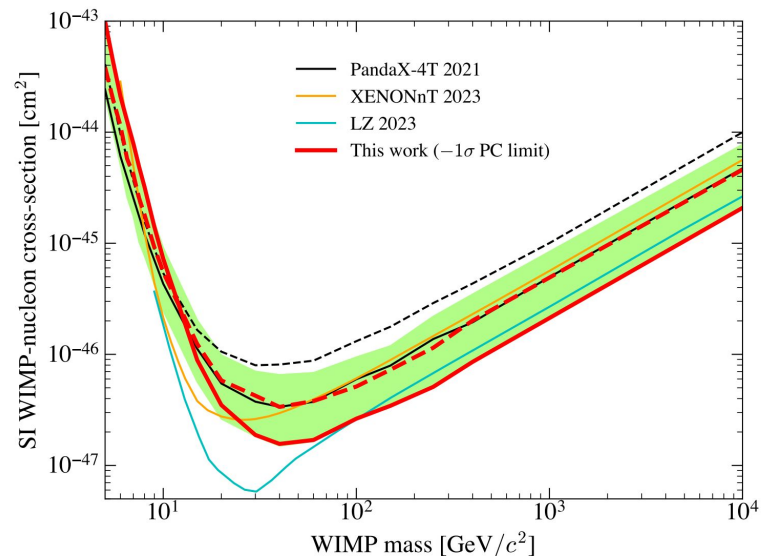
**PANDA X**  
PARTICLE AND ASTROPHYSICAL XENON TPC



**Dual-phase liquid xenon TPC** China  
Jinping Underground Laboratory with  
2,400 m rock overburden, **deepest lab  
in the world**

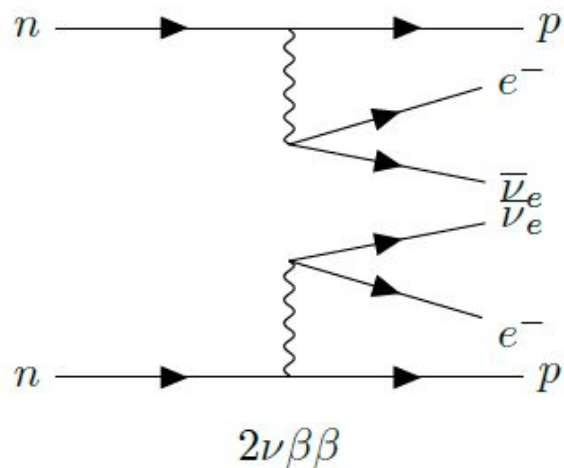


**Dual-phase liquid xenon TPC China**  
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**in the world**



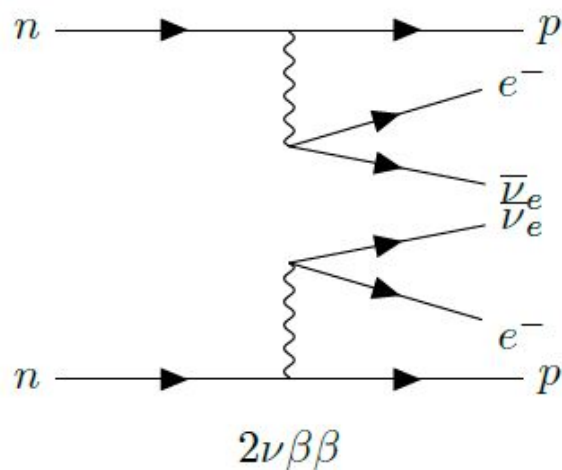


3.7 tonnes of natural xenon of which  
**8.58% is naturally abundant  $^{136}\text{Xe}$**

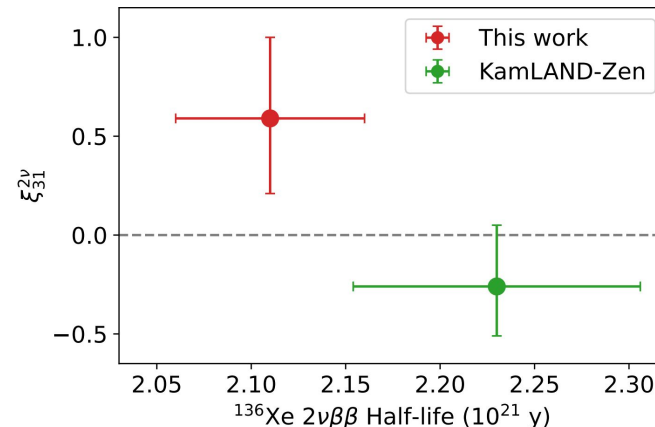




3.7 tonnes of natural xenon of which  
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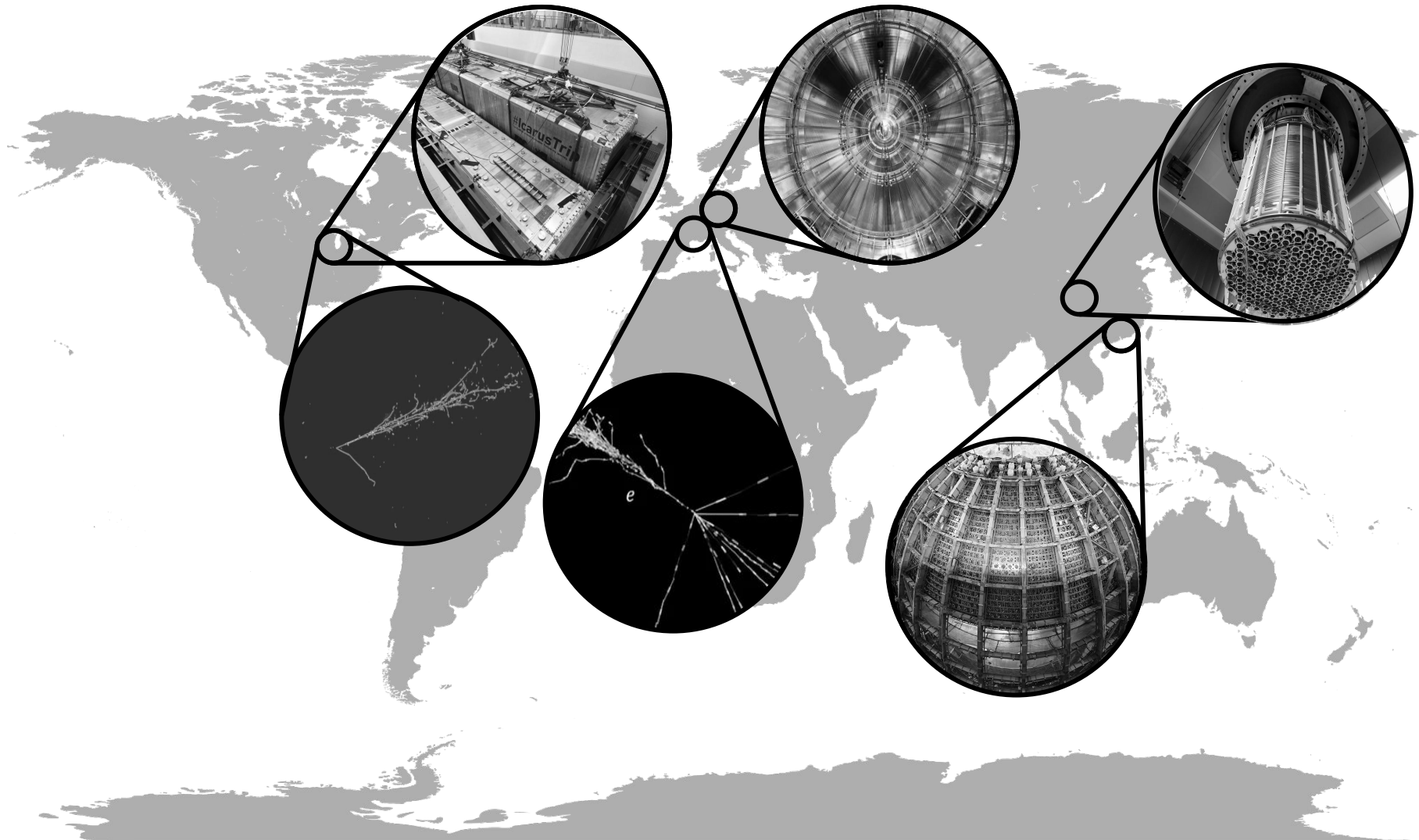


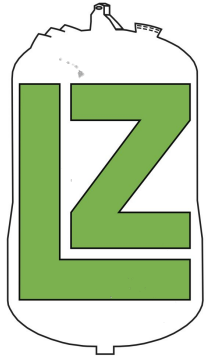
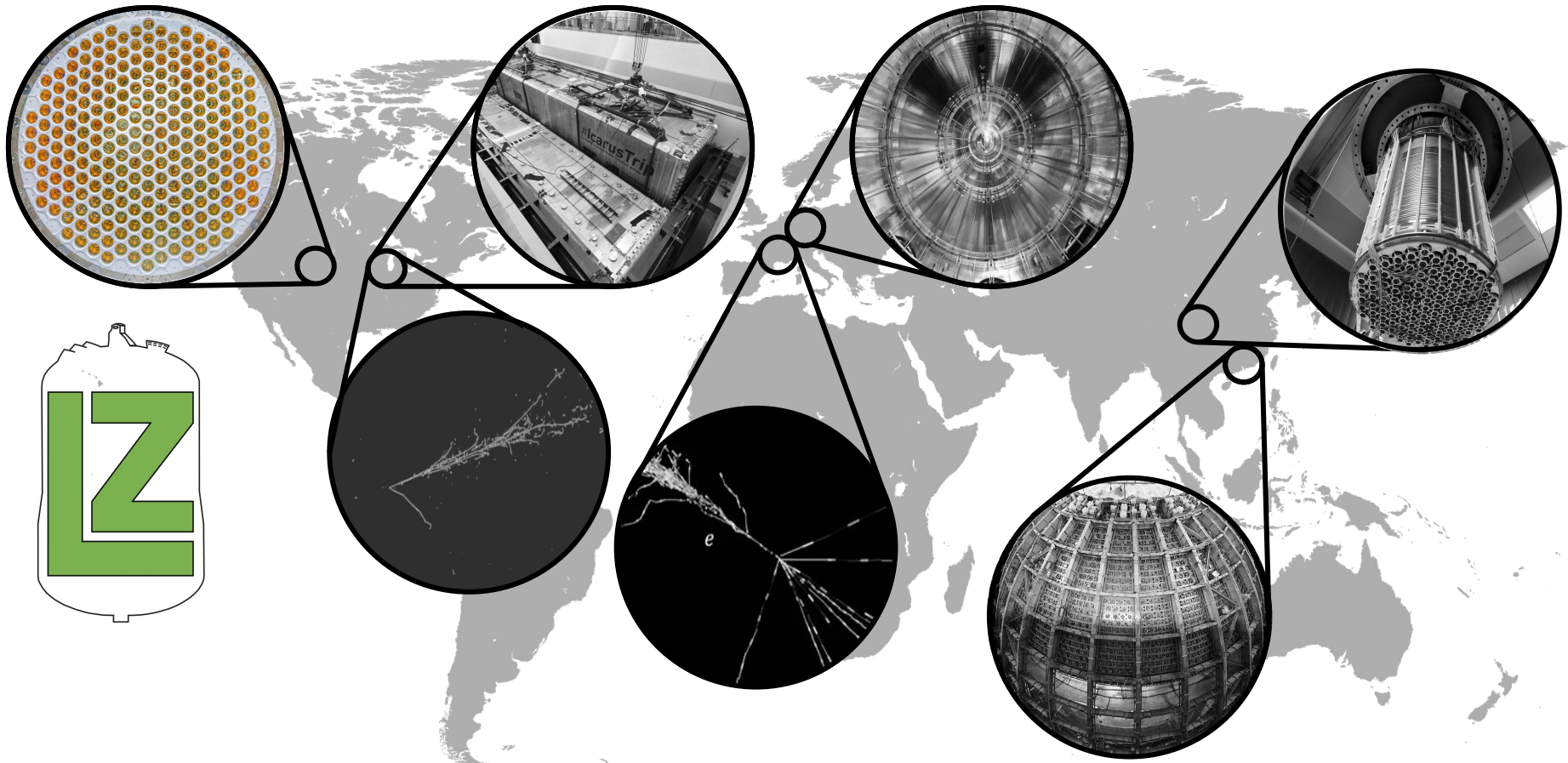
**258 Days of data (~25% of  
 planned exposure)**

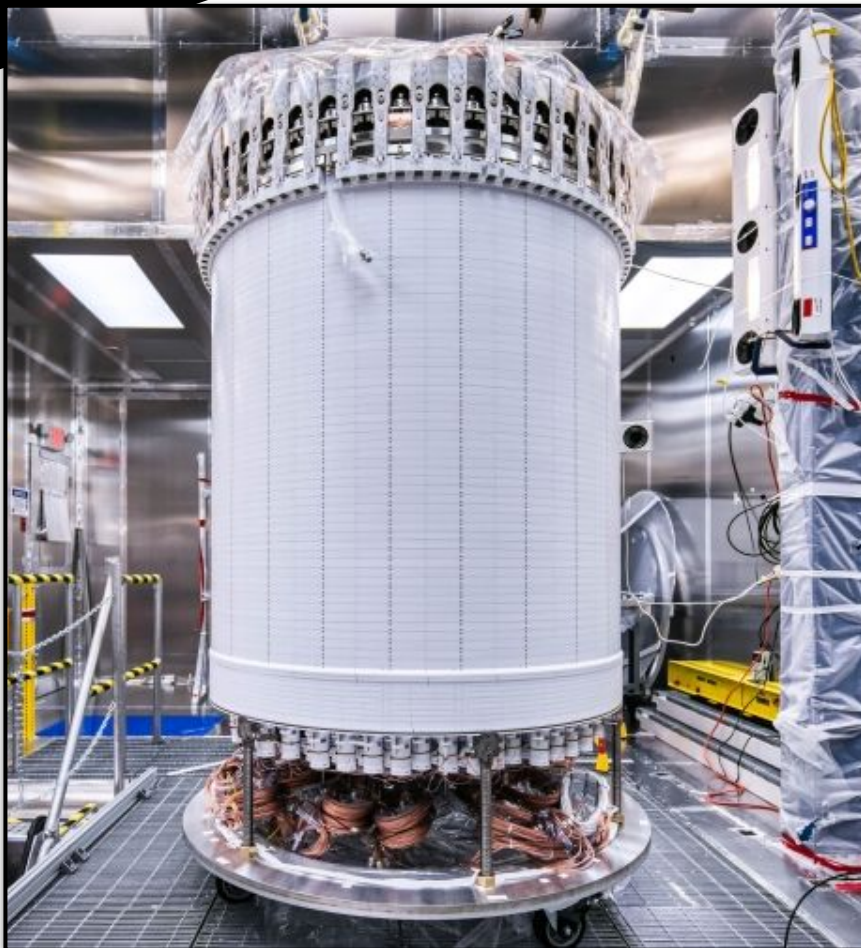
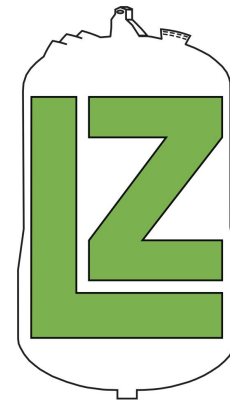


$$T_{1/2}(2\nu\beta\beta) = (2.14 \pm 0.05) \times 10^{21} \text{ yr}$$

PandaX-4T is now the world's most **precise  $^{136}\text{Xe}$   $2\nu\beta\beta$  spectrometer** and is starting to **constrain the same nuclear matrix elements** that  $0\nu\beta\beta$ -to- $m_{\beta\beta}$  conversions depend on.







**7-tonne dual-phase liquid-xenon TPC** dark-matter experiment located 4,850 ft underground in SURF, South Dakota

**See talks by Hugh Lippincott and Jack Genovesi Monday!**



Progress and Prospects in Dark Matter Direct Detection  
Hugh Lippincott, UCSB



May 11-13, 2026 - PHENO



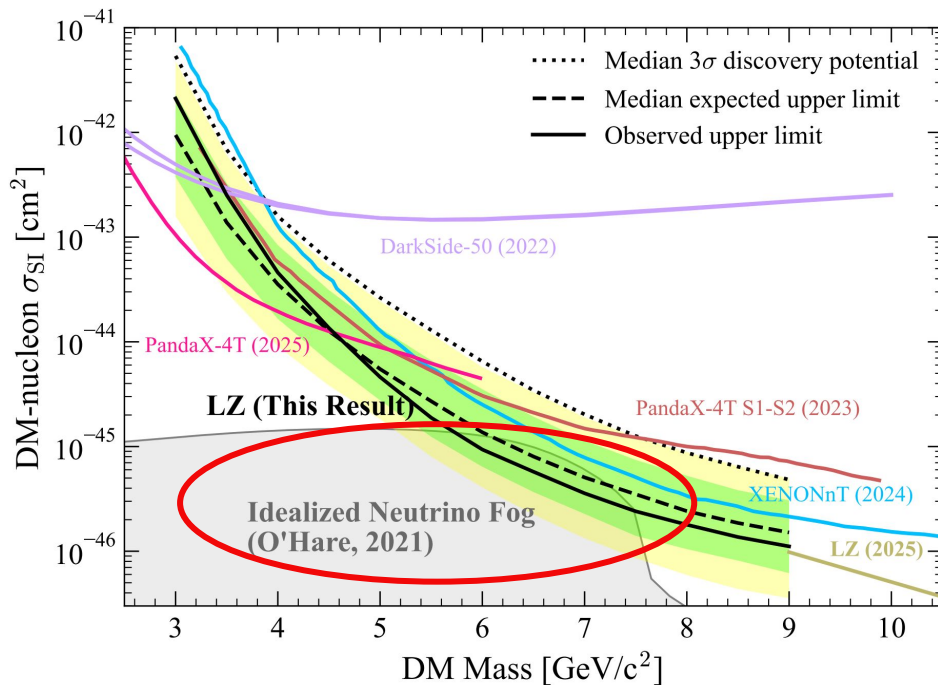
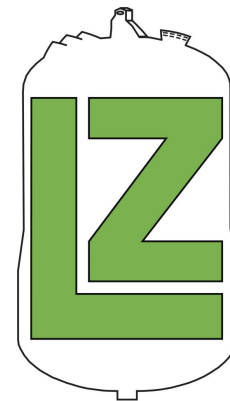
Latest Results of the LUX-ZEPLIN (LZ) Dark Matter Experiment

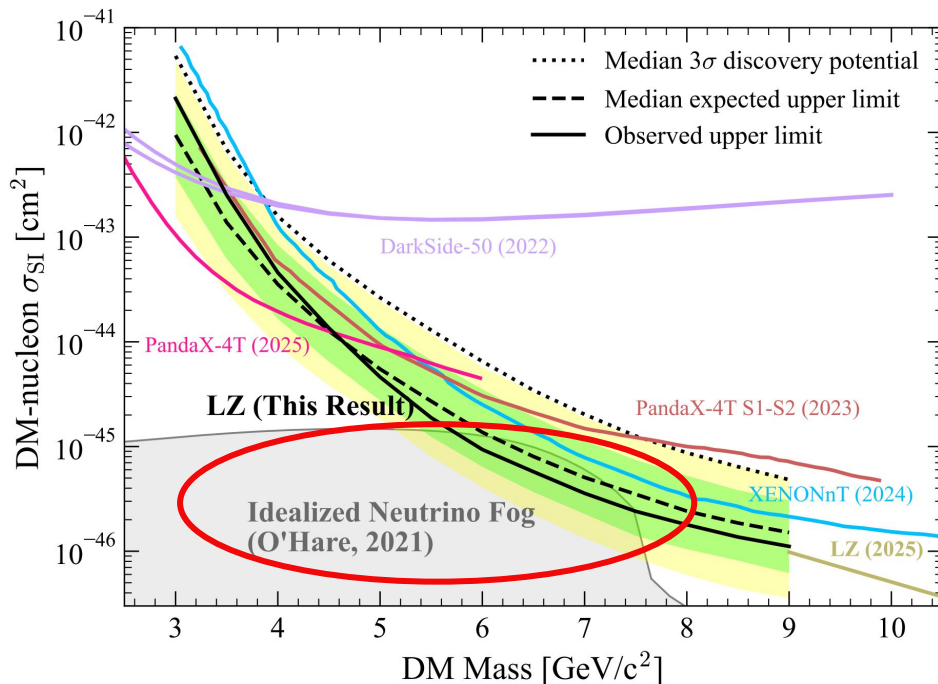
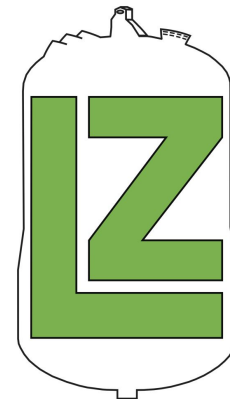


PennState

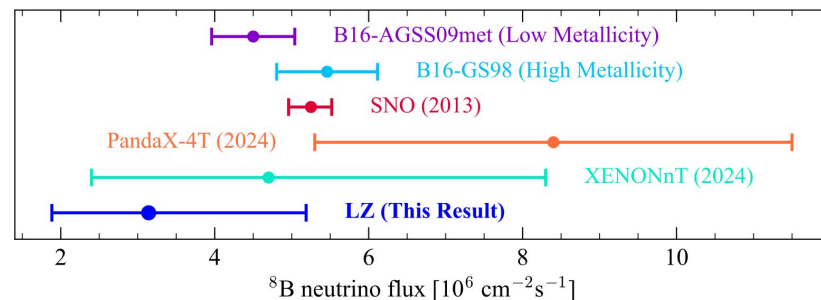
Jack Genovesi  
on behalf of the LUX-ZEPLIN (LZ) Collaboration  
PHENO, 2026





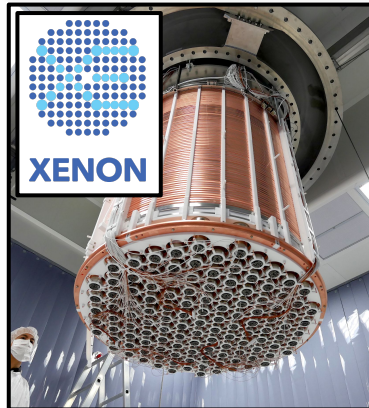
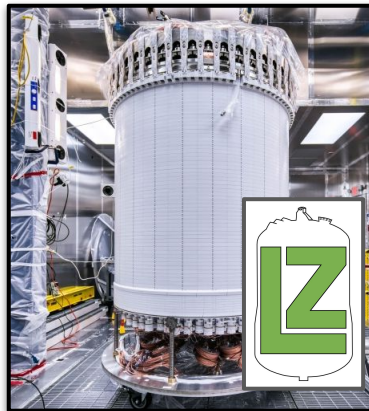


“..observe a signal consistent with **Solar  $^8\text{B}$  CEvNS** events, corresponding to a  **$4.5\sigma$**  statistical significance”



The same CEvNS events that *limit* dark matter searches are a **new measurement of the solar  $^8\text{B}$  flux** independent of SNO/Super-K.

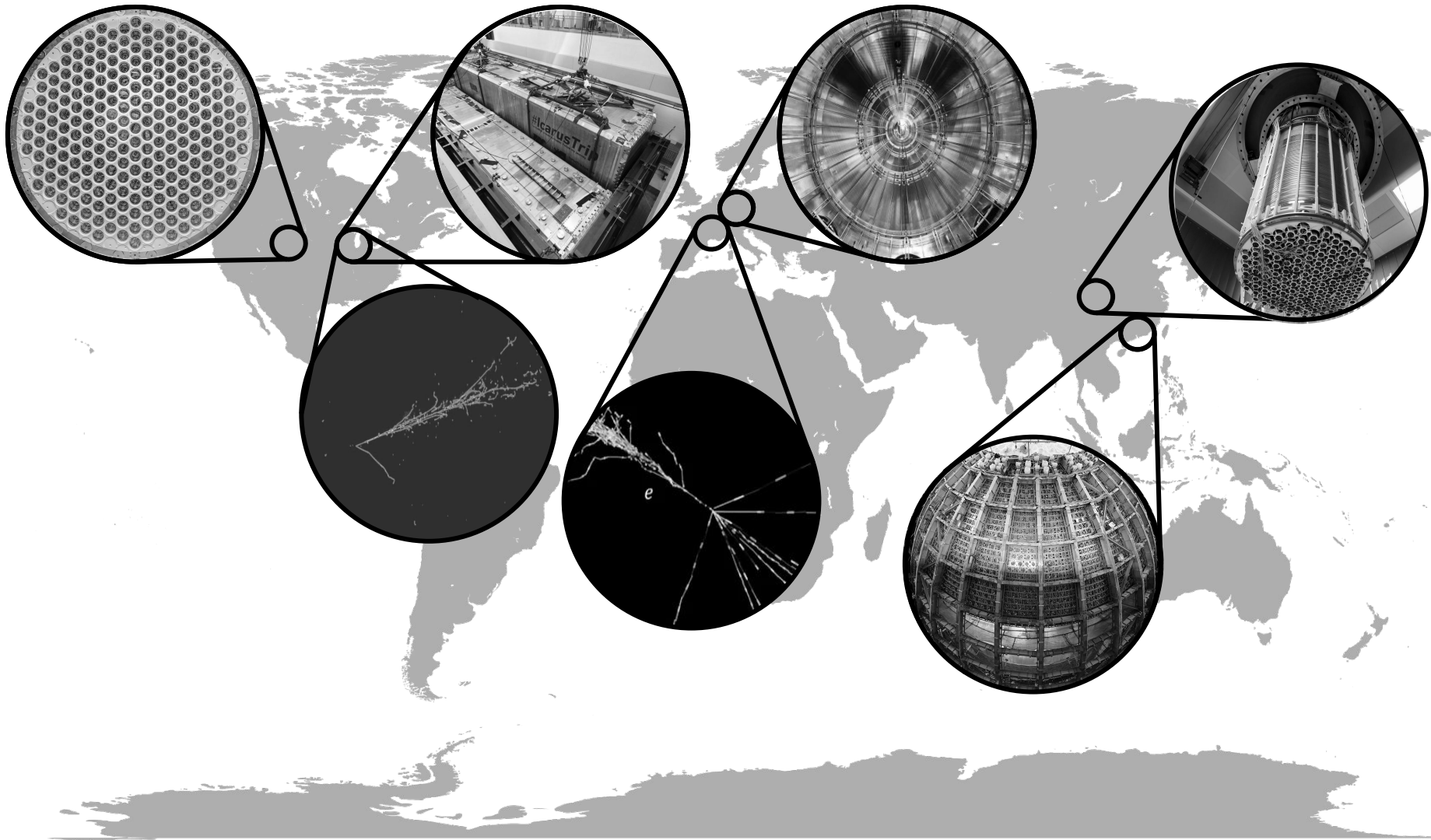
As exposure grows they become a clean low- $Q^2$  probe of NSI, neutrino magnetic moments, light mediators, and the weak mixing angle.

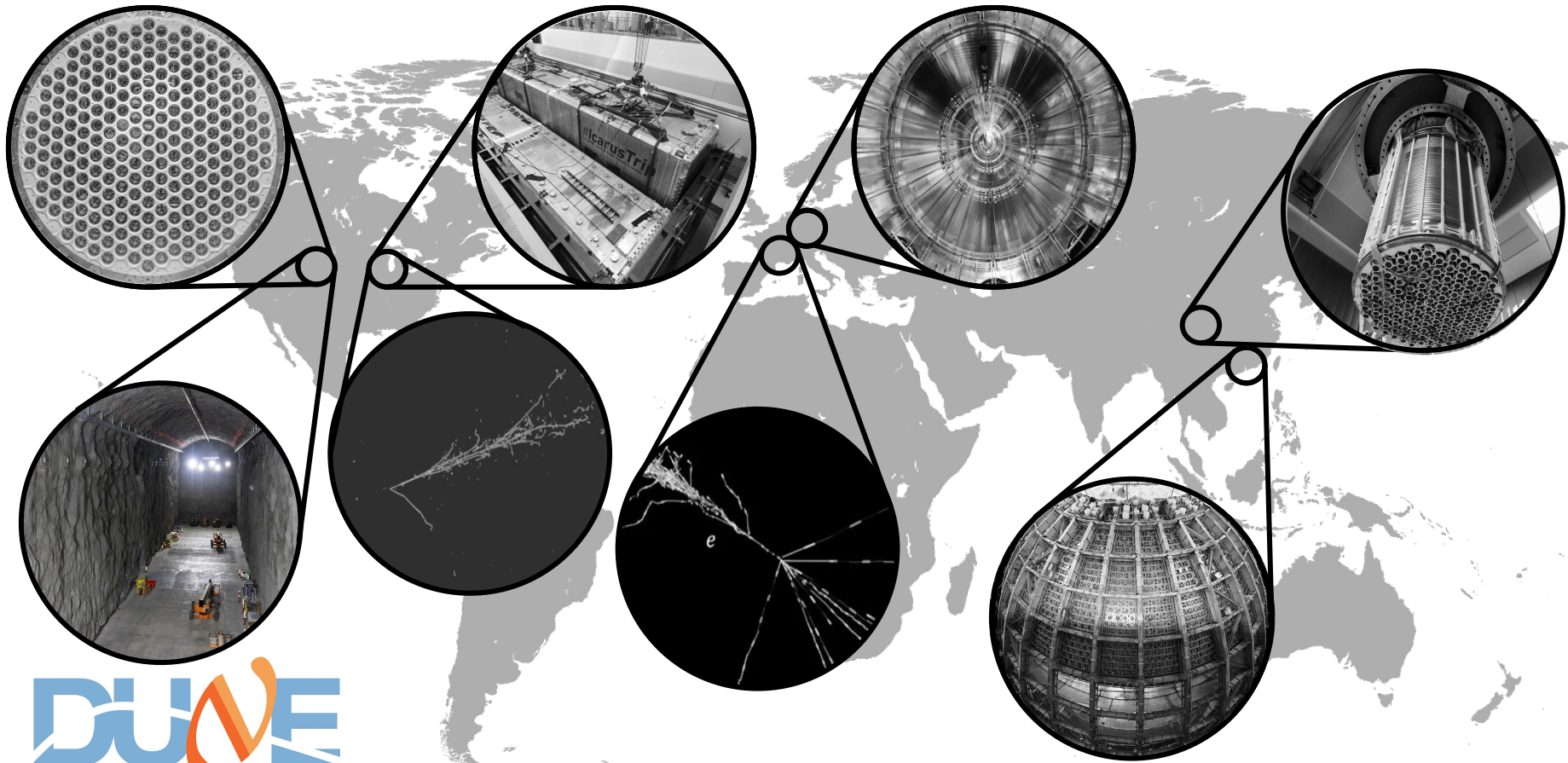


+

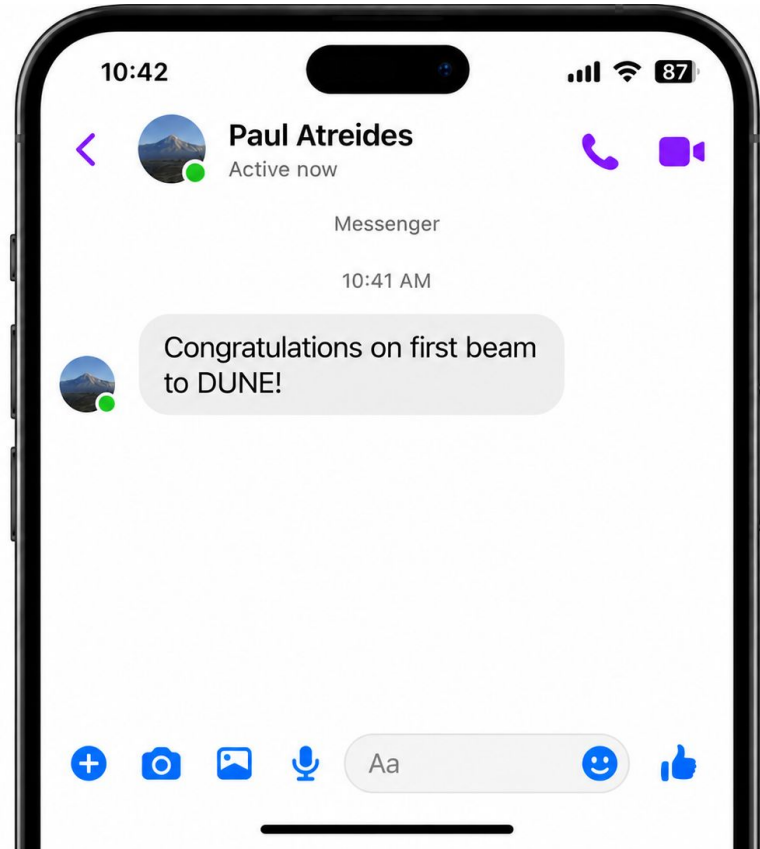
- DarkSide 20 (LAr)
- XLZD
- PandaX-xT
- ARGO (LAr)
- ...

Tonne-scale+ DM detectors have effectively become solar neutrino observatories in the past 18 months

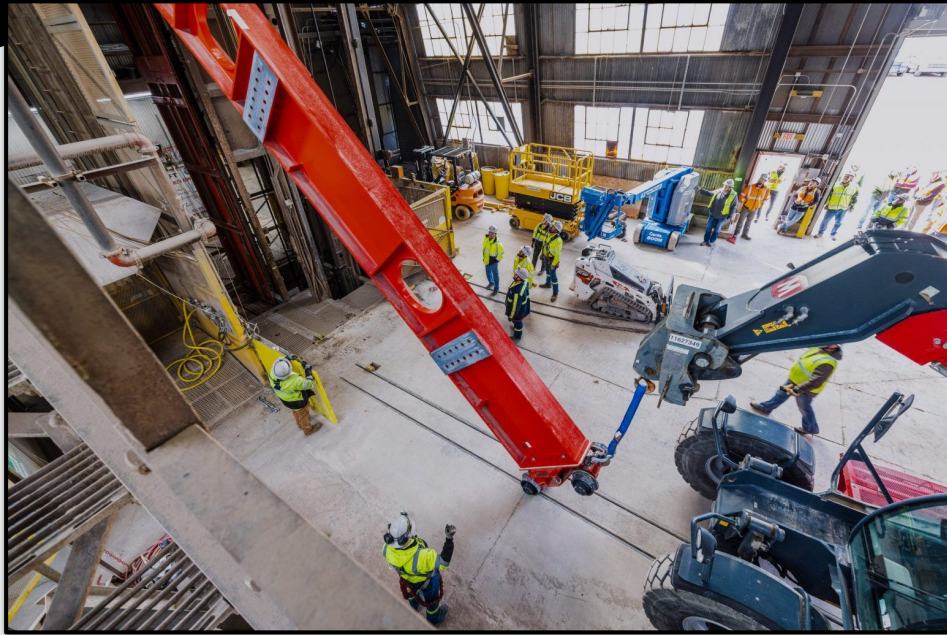




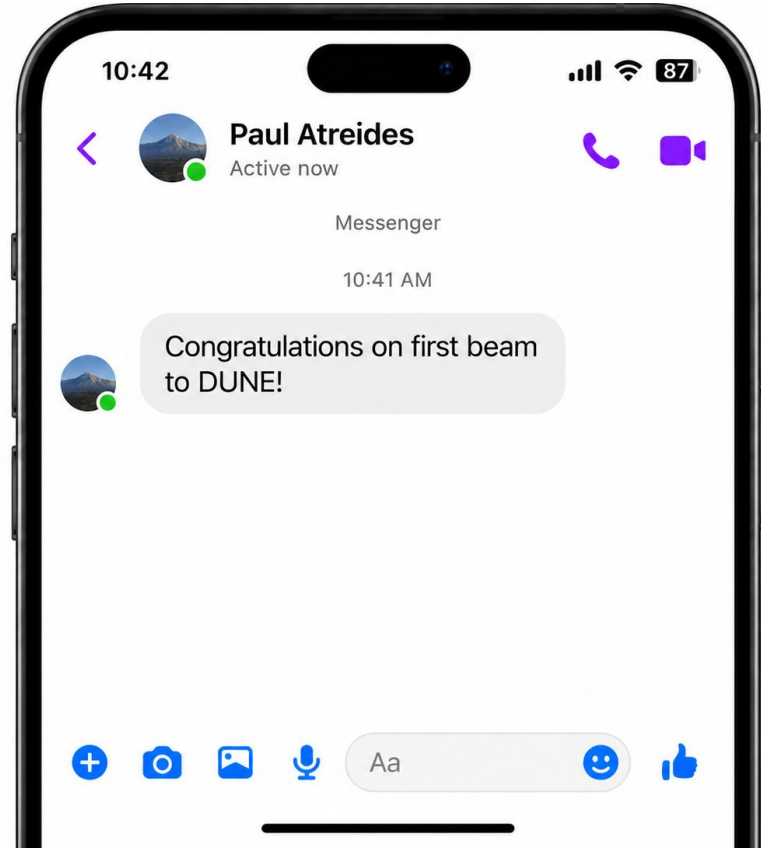
**DUNE**  
DEEP UNDERGROUND  
NEUTRINO EXPERIMENT



# DUNE: Deep Underground Neutrino Experiment



First of of 1,360 steel beams for cryostat construction lowered into DUNE caverns (4,500 Metric Tonnes of steel) on **May 7<sup>th</sup>**



# DUNE: Deep Underground Neutrino Experiment



All caverns complete in 2024,  
first *beams* lowered mark the  
**start of detector construction**





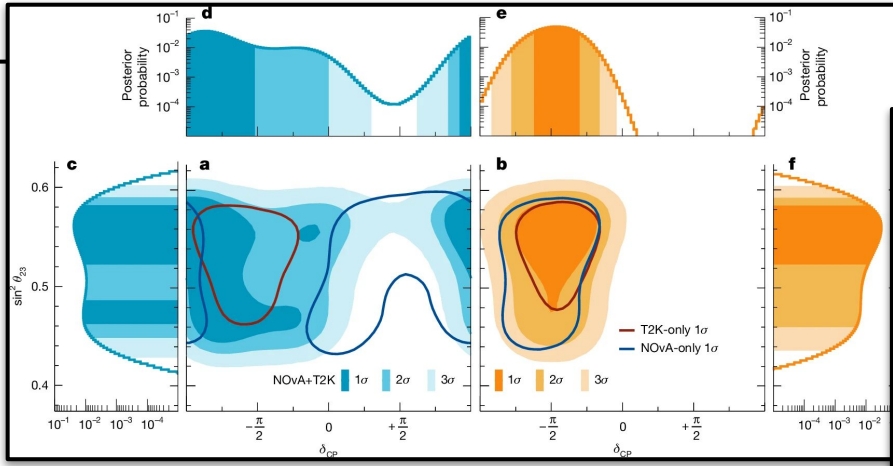
Last ~6 months

Plus many many more..

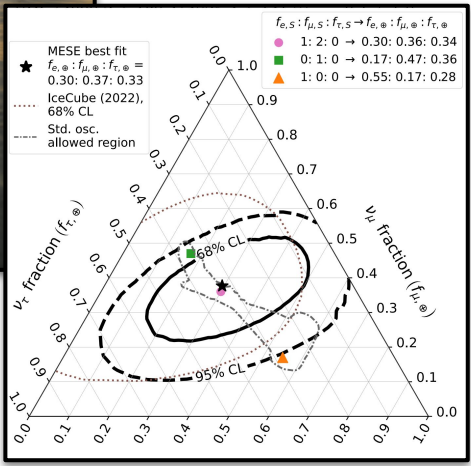


The KM3NeT Collaboration  
Nature volume 638,  
376–382 (2025)

M. Ross-Lonergan – May 13<sup>th</sup> – PHENO 2026

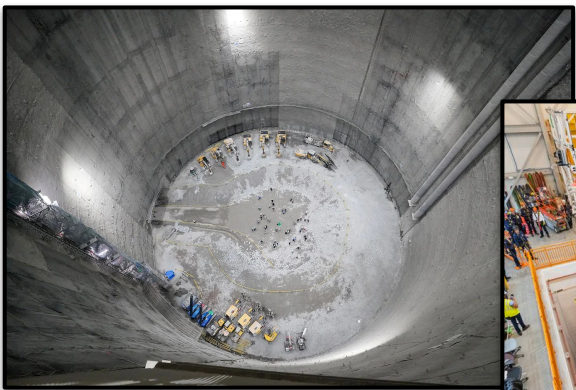
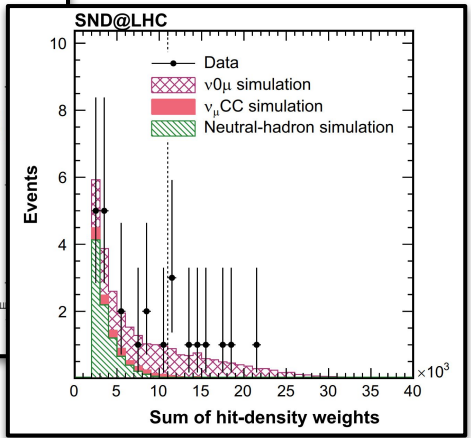


NoVA & T2K Joint Fit Nature volume 646, 818–824 (2025)



IceCube: arXiv:2510.24957

SND@LHC Phys. Rev. Lett. 134, 231802



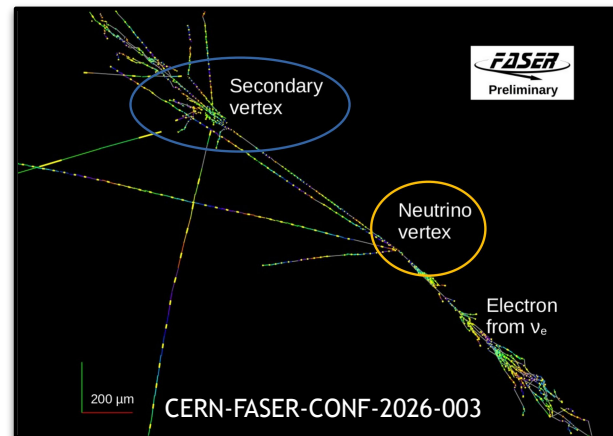
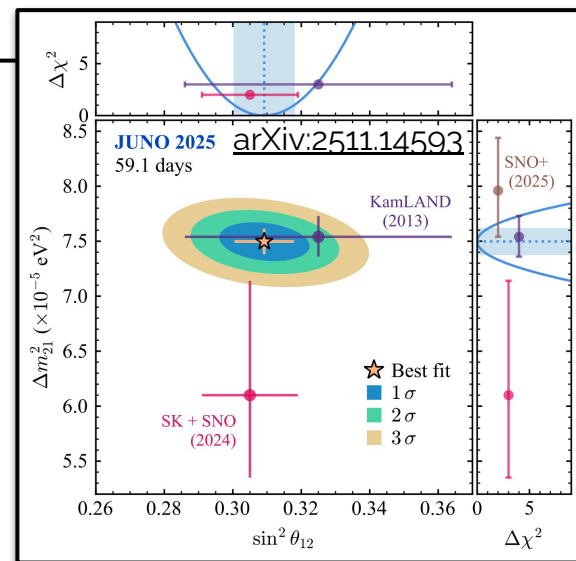
HyperK Cavern Complete (Aug 2025)



SBND fully operational

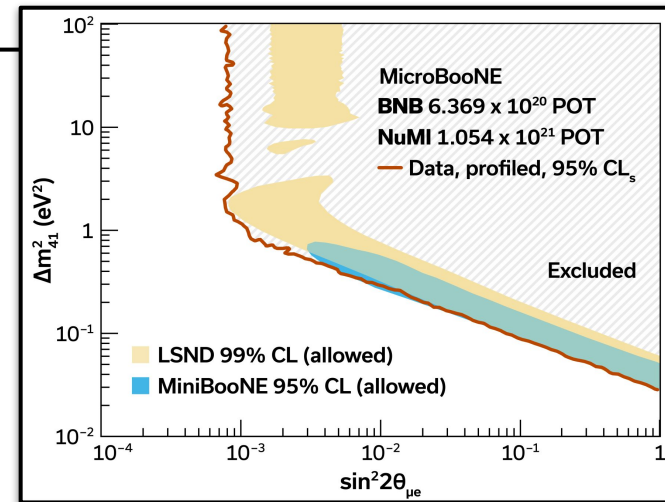
# Conclusions I

- Next generation **precision measurement experiments** are here with **JUNO** successful commissioning and operation
- Significant ramp up of **collider neutrino data and capabilities** which, alongside **xenon dark-matter TPCs**, are rapidly what we call neutrino experiments at the TeV and MeV scale

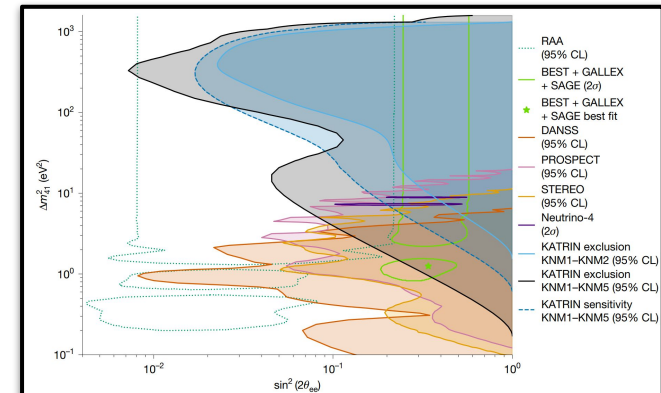


# Conclusions II

- Next generation **precision measurement experiments** are here with **JUNO** successful commissioning and operation
- Significant ramp up of **collider neutrino data and capabilities** which, alongside **xenon dark-matter TPCs**, are rapidly what we call neutrino experiments at the TeV and MeV scale
- There has been **extreme tension** in global data for a “light 3+1 sterile” for years. **Not just “global tension” anymore**, with new results **directly ruling out** almost the entire viable parameter space of various anomalies.
  - *But the anomalies themselves remain and need explanations!*

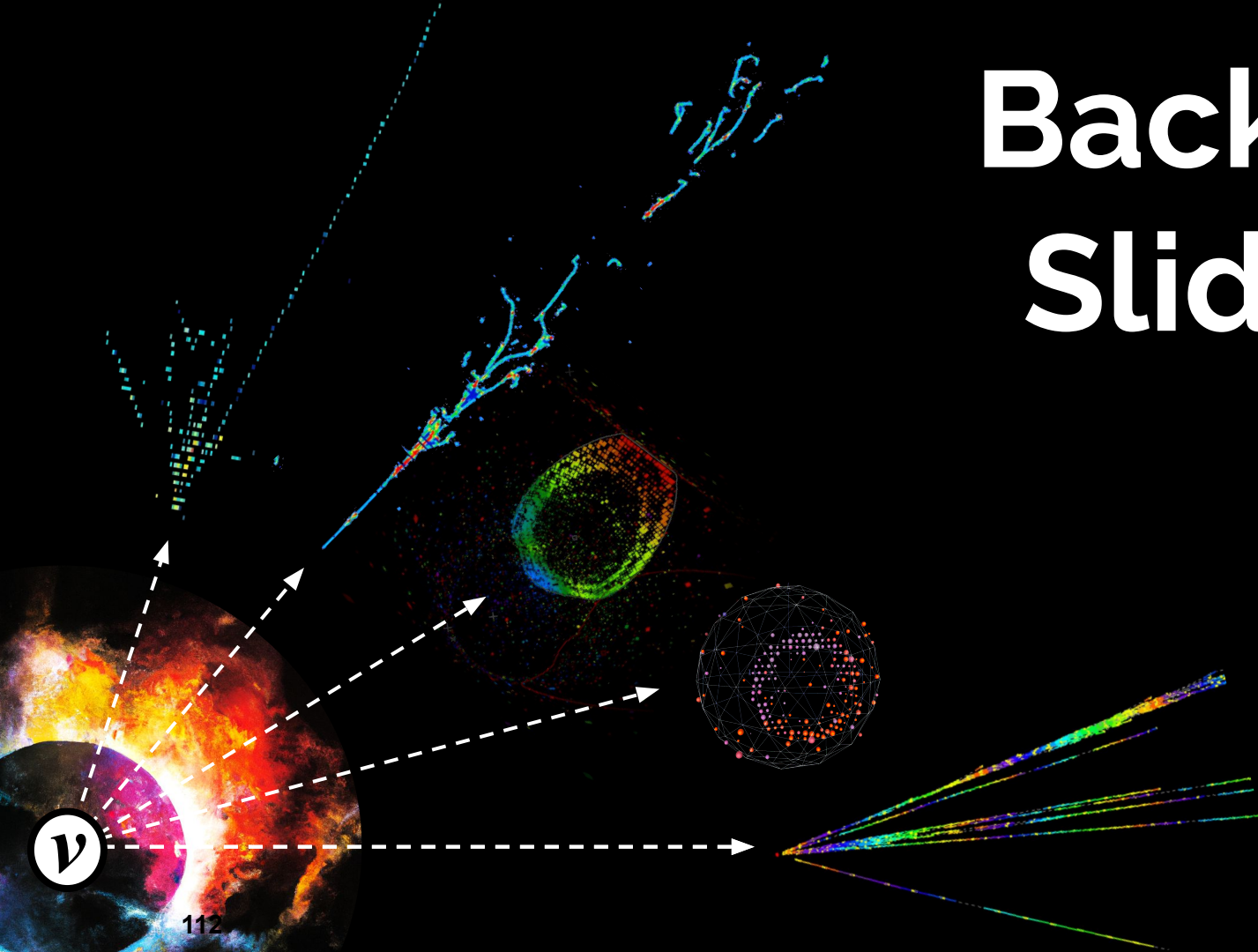


MicroBooNE: Nature 648, 64–69 (2025)



KATRIN: Nature 648, 70–75 (2025)

# Backup Slides



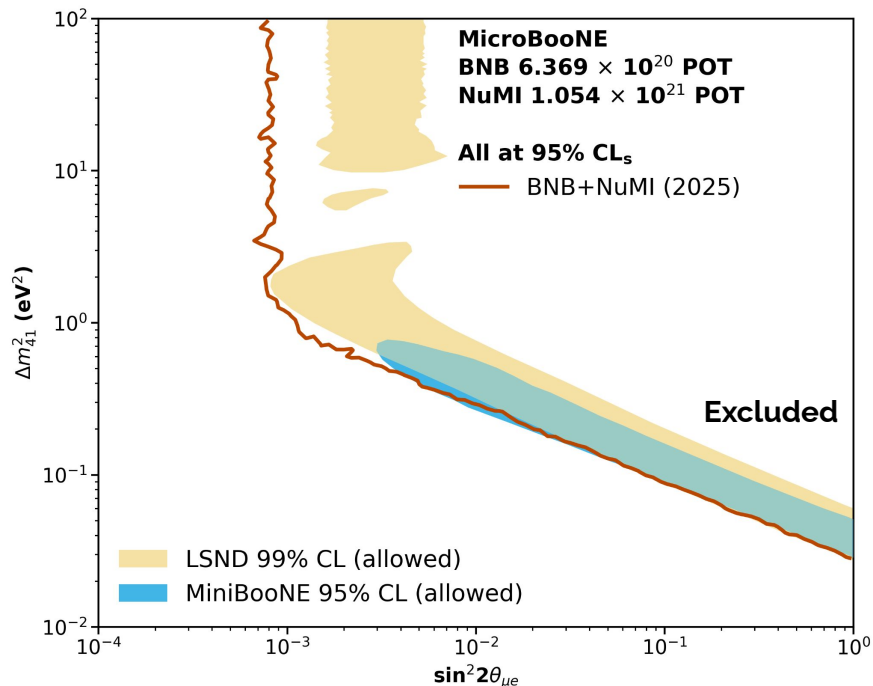


Sources/Processes	Type	Radiation
$^{137}\text{Cs}$	$\gamma$	0.662 MeV
$^{54}\text{Mn}$	$\gamma$	0.835 MeV
$^{60}\text{Co}$	$\gamma$	1.173 + 1.333 MeV
$^{40}\text{K}$	$\gamma$	1.461 MeV
$^{68}\text{Ge}$	$e^+$	annihilation 0.511 + 0.511 MeV
$^{241}\text{Am-Be}$	n, $\gamma$	neutron + 4.43 MeV ( $^{12}\text{C}^*$ )
$^{241}\text{Am-}^{13}\text{C}$	n, $\gamma$	neutron + 6.13 MeV ( $^{16}\text{O}^*$ )
(n, $\gamma$ )p	$\gamma$	2.22 MeV
(n, $\gamma$ ) $^{12}\text{C}$	$\gamma$	4.94 MeV or 3.68 + 1.26 MeV

# First Dual Beam Sterile Neutrino Constraints

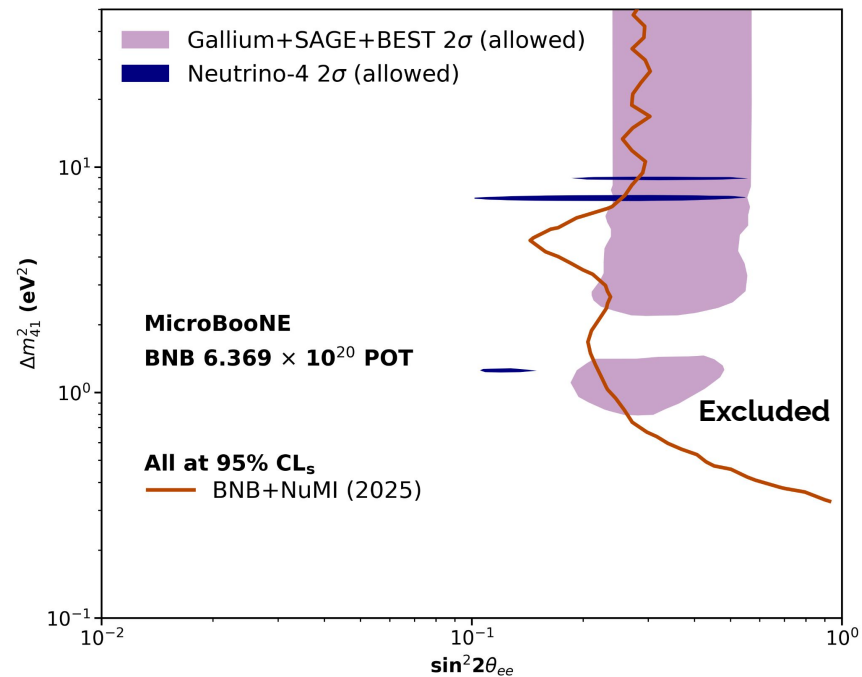
MicroBooNE: Nature volume 648, 64–69 (2025)

$$\nu_\mu \rightarrow \nu_e$$



- Excludes LSND 99% allowed region
- Excludes vast majority of MiniBooNE 95% allowed region

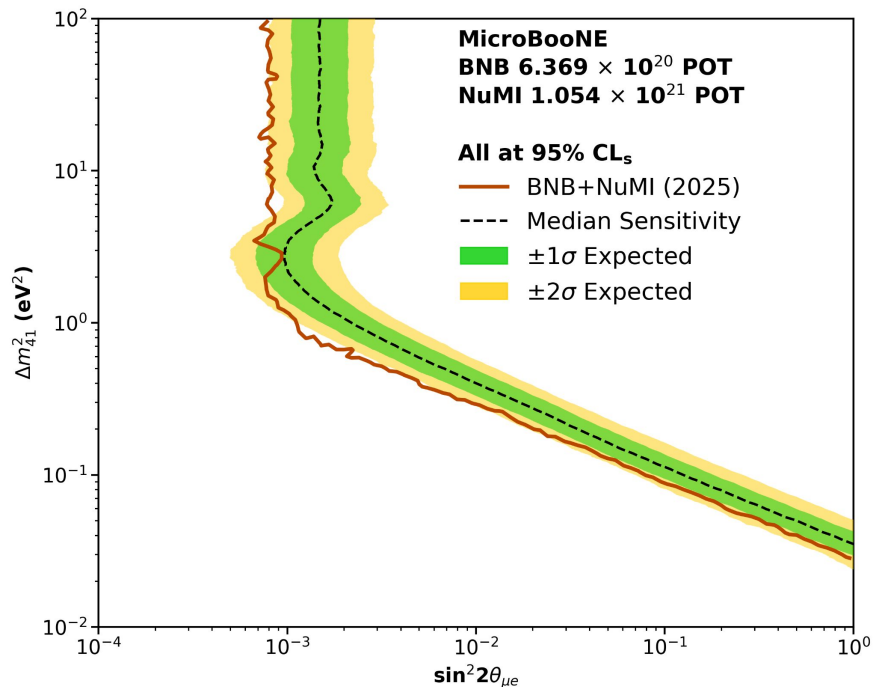
$$\nu_e \rightarrow \nu_e$$



- Provides an **internal, self-consistent  $\nu_e$  disappearance constraint**
- Covers most of Gallium allowed region and part of Neutrino-4 allowed region

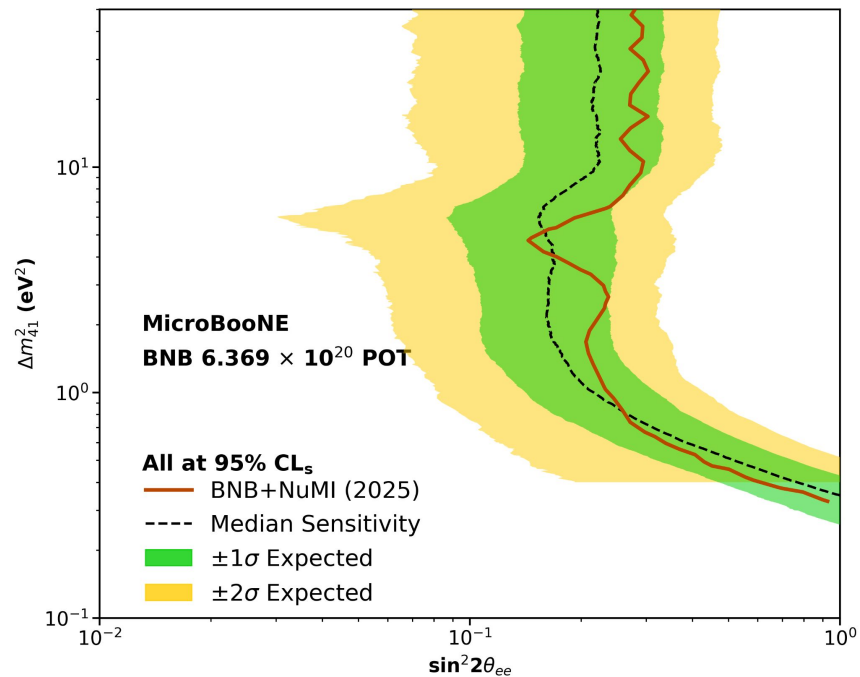
# Predicted Sensitivities

$\nu_\mu \rightarrow \nu_e$



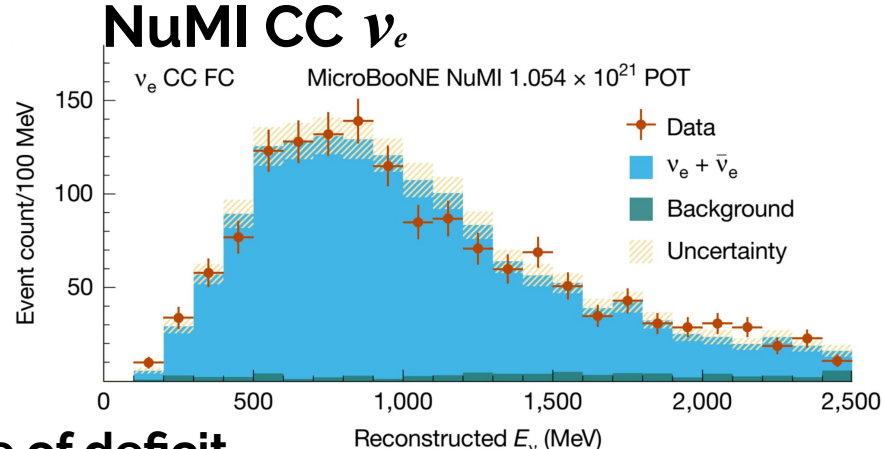
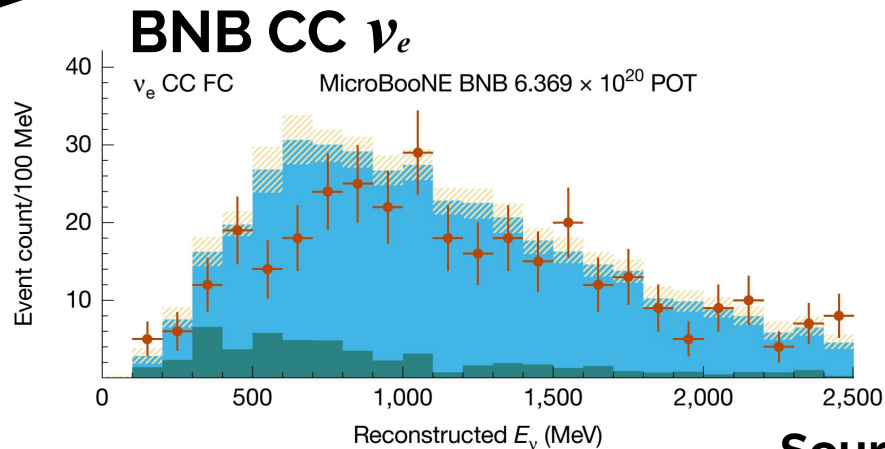
**Data exclusion is stronger than median sensitivity.**  
Lies around the  $2\sigma$  expected experiment

$\nu_e \rightarrow \nu_e$



Data exclusion is slightly weaker than median sensitivity.

# Strength of Result Driven by BNB $\nu_e$



## Source of deficit

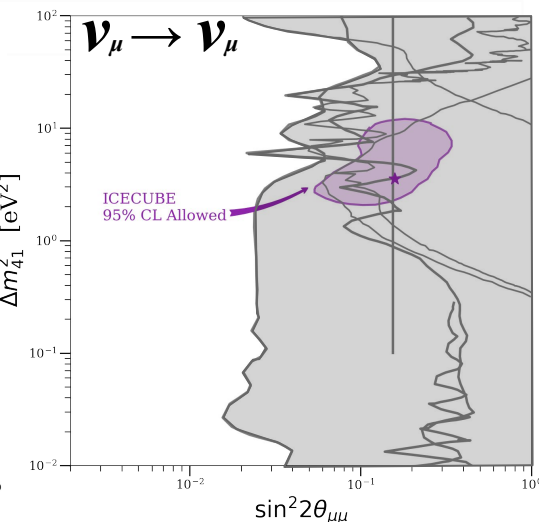
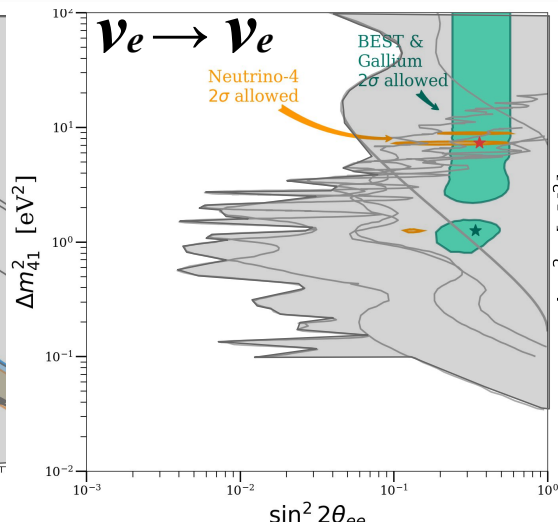
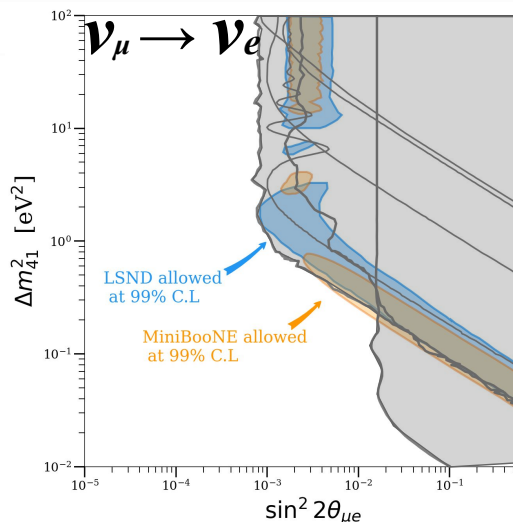
- Interaction Modelling? **Same Model**
- Detector Modelling? **Same Detector/treatment**
- Reconstruction Effect? **Same Reco Framework**
- Flux Modelling? **Different Fluxes/Beams**

- **Statistical**

- **New Physics Driven**

Investigations  
Continue  
~Double the Data

# Where we Stand?



## Positive Hints

- LSND
- MiniBooNE

## Null Results

- Karmen
- NuTeV
- CCFR
- NOMAD
- OPERA
- MicroBooNE

## Positive Hints

- Gallium
- Neutrino-4

## Null Results

- NEOS+RENO
- DANSS
- PROSPECT
- STEREO
- KATRIN

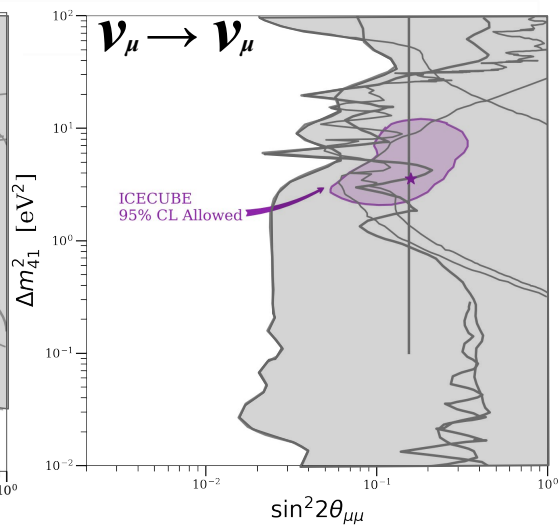
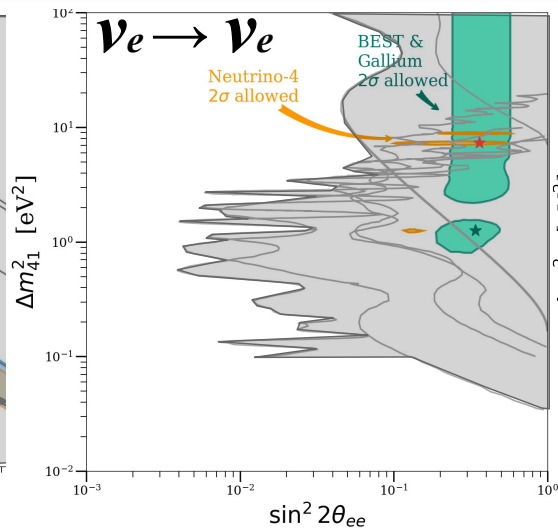
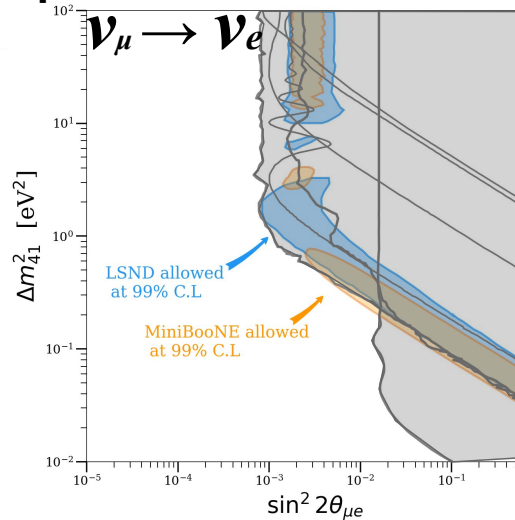
## Positive Hints

- IceCube

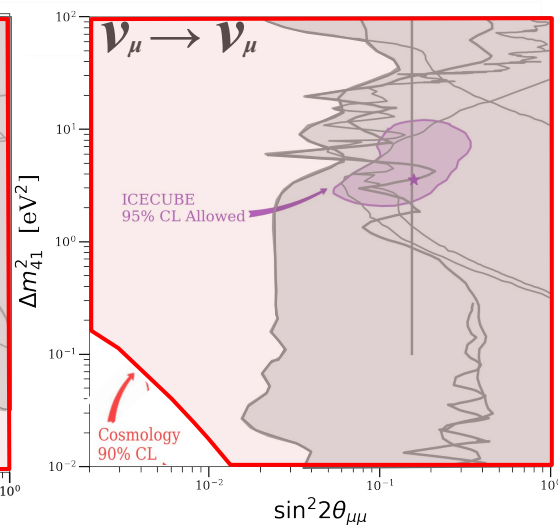
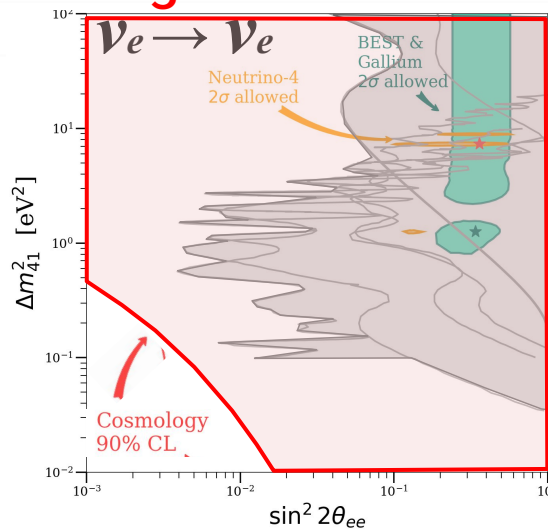
## Null Results

- MiniBooNE
- CDHS
- CCFR
- T2K
- NoVA
- MINOS
- Super-K

# The Elephant in the Room...

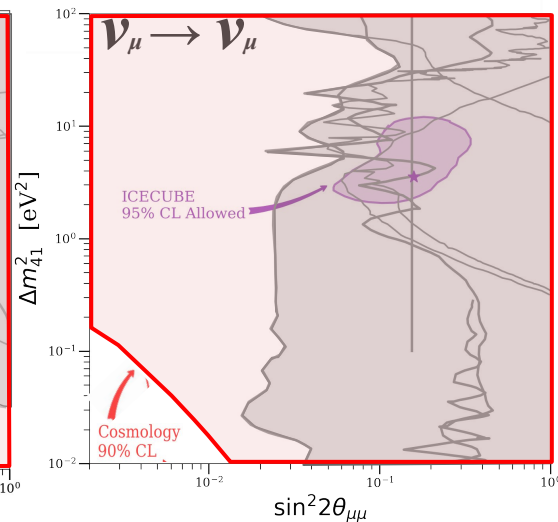
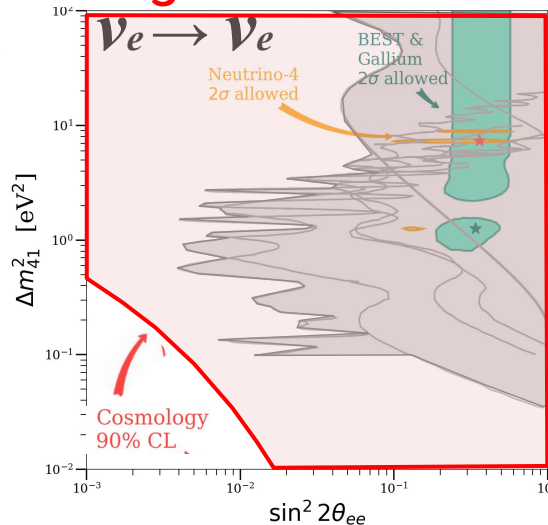


**HUGE tension**  
between  
cosmological  
bounds and *vanilla*  
**3+1 short-baseline  
steriles**



S. Hagstotz et al, [PhysRevD.104.123524](https://arxiv.org/abs/1907.08727)

**HUGE tension**  
between  
cosmological  
bounds and **vanilla**  
**3+1 short-baseline  
steriles**



Energy density of **Relativistic Neutrinos**

S. Hagstotz et al, [PhysRevD.104.123524](https://arxiv.org/abs/1708.07401)

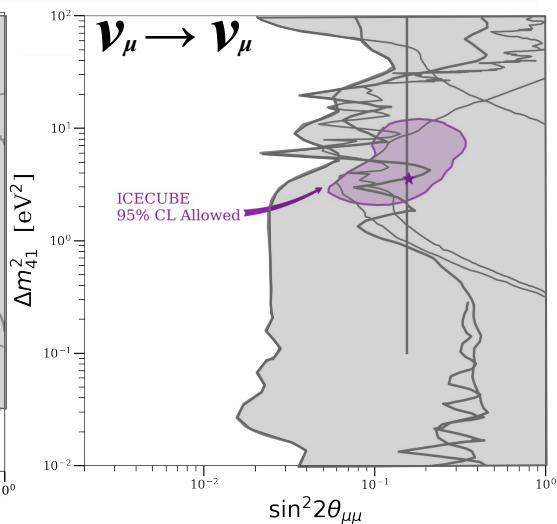
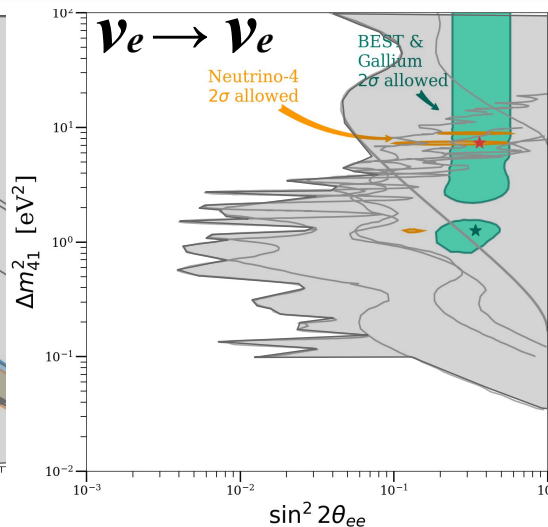
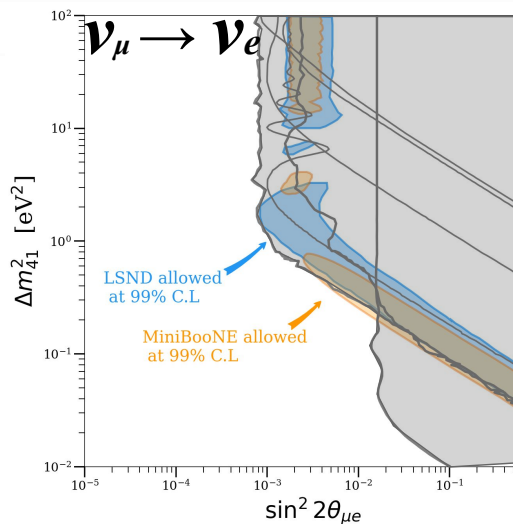
$$N_{\text{eff}}^{\nu} = \frac{8}{7} \left( \frac{11}{4} \right)^{4/3} \frac{\rho_{\nu}}{\rho_{\gamma}}$$

$$N_{\text{eff}} = 2.99 \pm 0.17 \text{ (CMB+lens+BAO) } \text{ [Plank2018](#) }$$

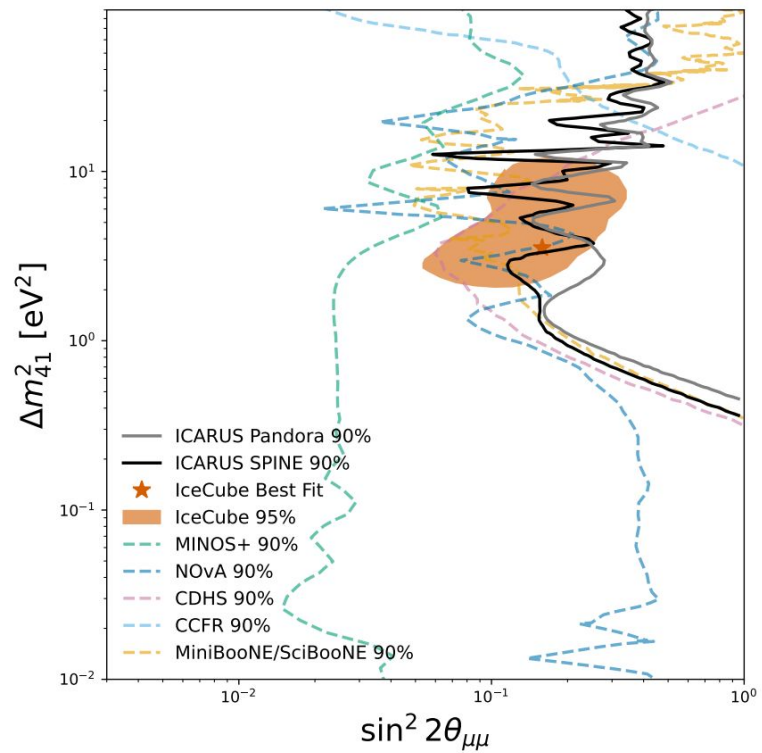
**Can this be avoided? Is this  
dependent on specific models  
of  $\Lambda$ CDM?**

**Possibly, but you have to add  
more new physics! It's no longer a  
simple vanilla "3+1" model**

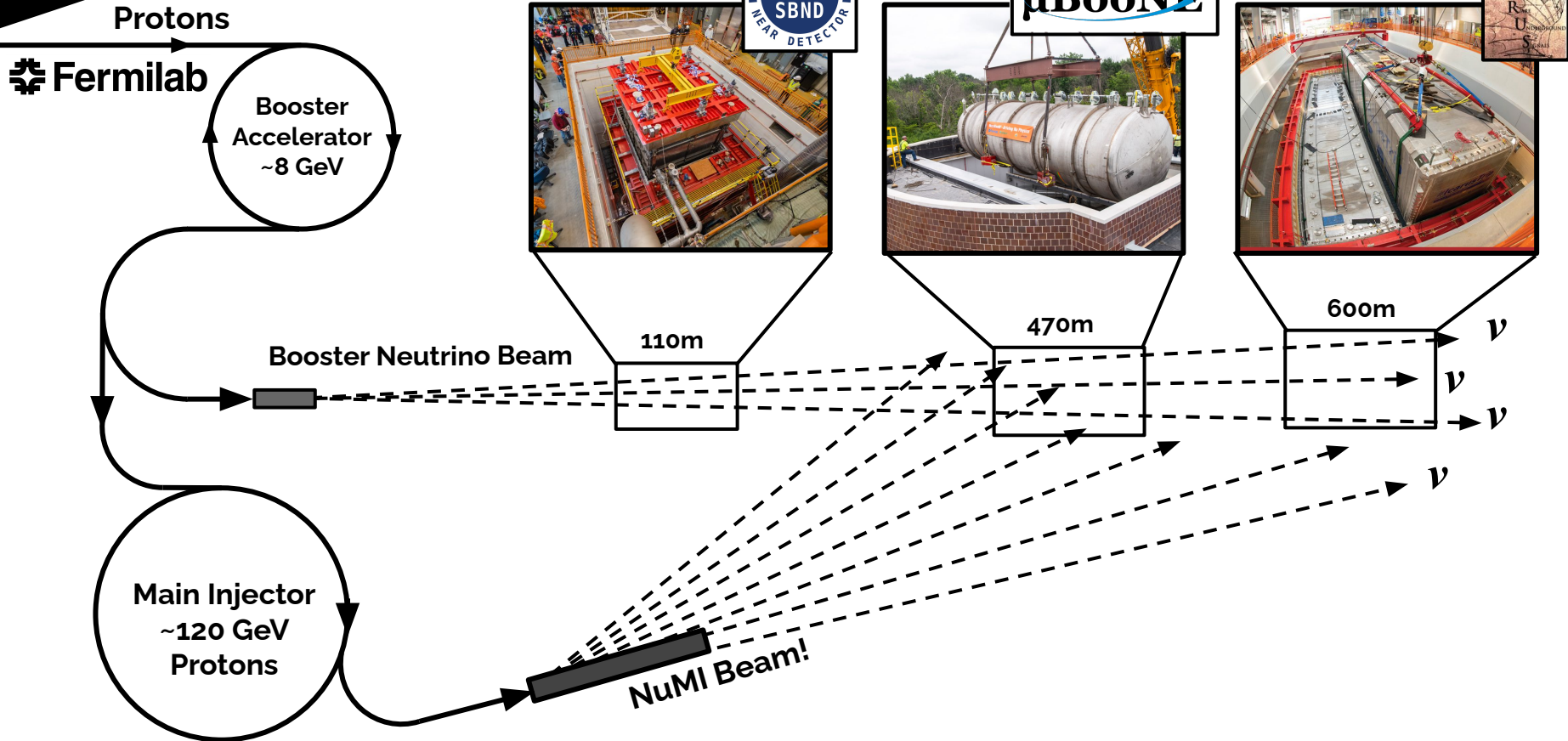
# Where We Stand – 2026



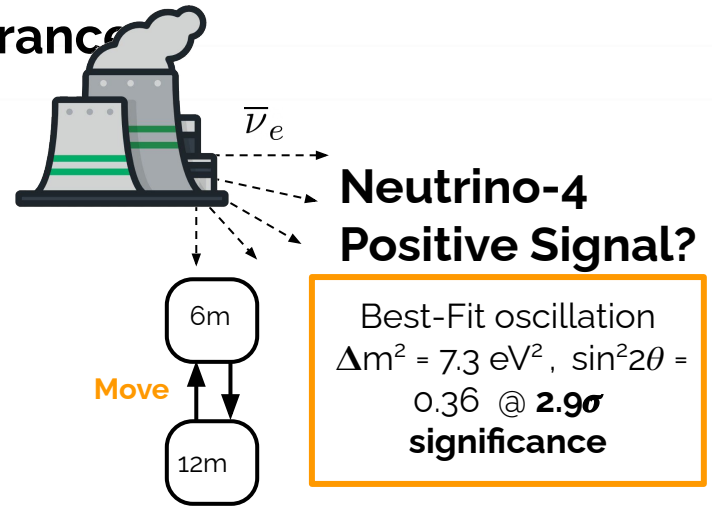
- There has been **EXTREME tension** in global data for a “light 3+1 oscillating sterile neutrinos” for years, gets *much worse* with **cosmology** included



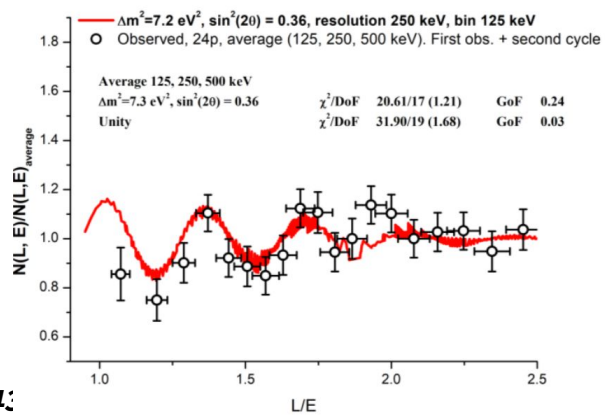
# SBN Programme



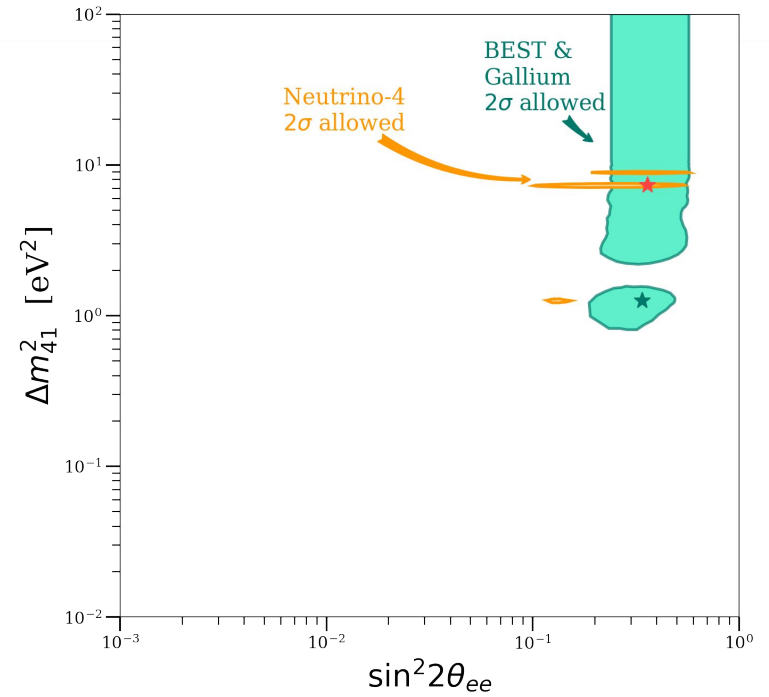
# $\bar{\nu}_e$ Disappearance

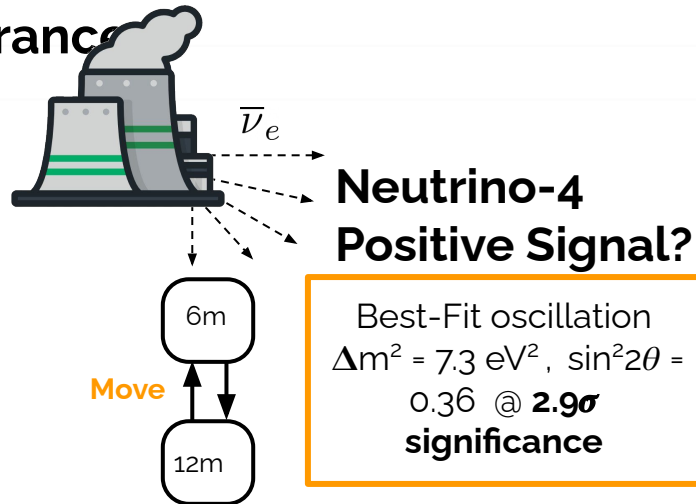


- Commercial reactor, 100 MW<sub>th</sub> in Russia



# $\nu_e \rightarrow \nu_e$

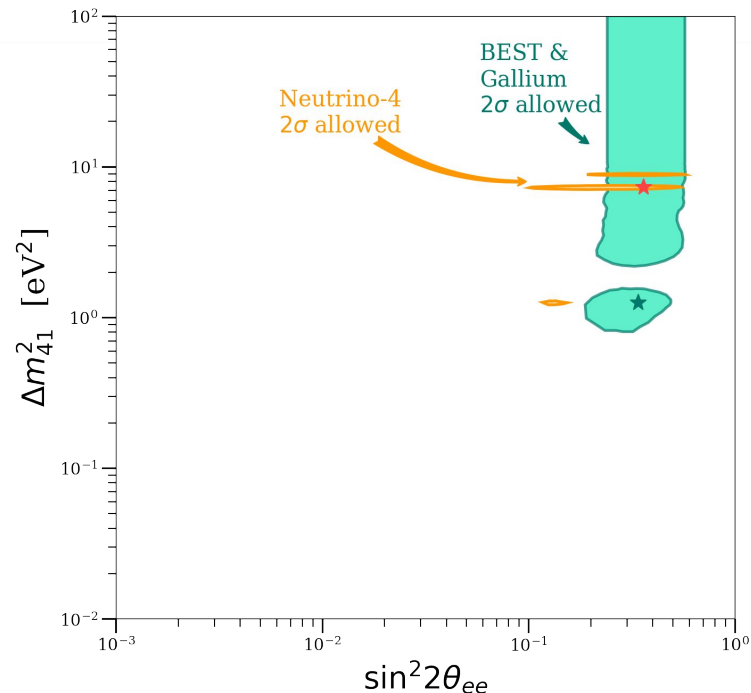




Several questions have been raised by community so far, including the **effects of the statistical approach used**, and impact of **systematics** and **expected backgrounds**

- H. Almazán et al, [arXiv:2006.13147](https://arxiv.org/abs/2006.13147)
- C. Giunti et al [j.physletb.2021.136214](https://arxiv.org/abs/2021.136214)

“Therefore, we conclude that the claimed Neutrino-4 indication in favor of short-baseline neutrino oscillations with very large mixing is rather doubtful.”

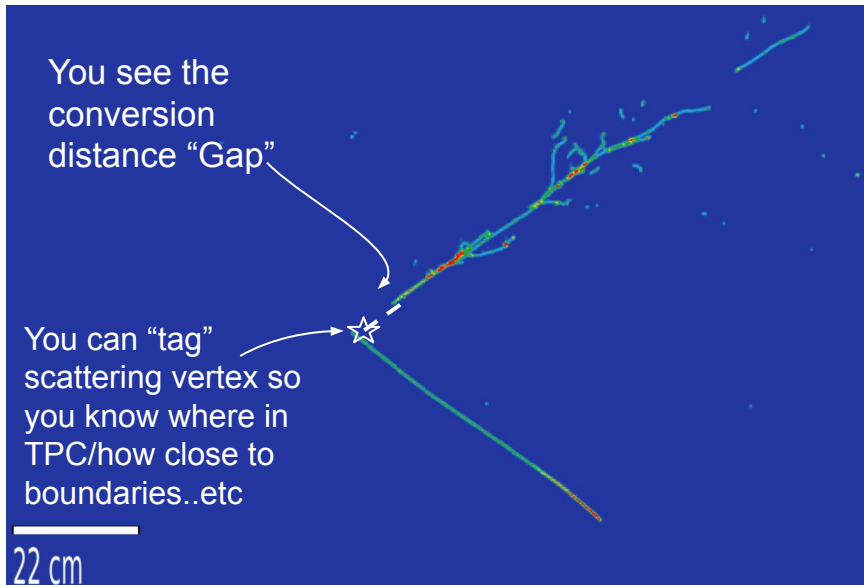




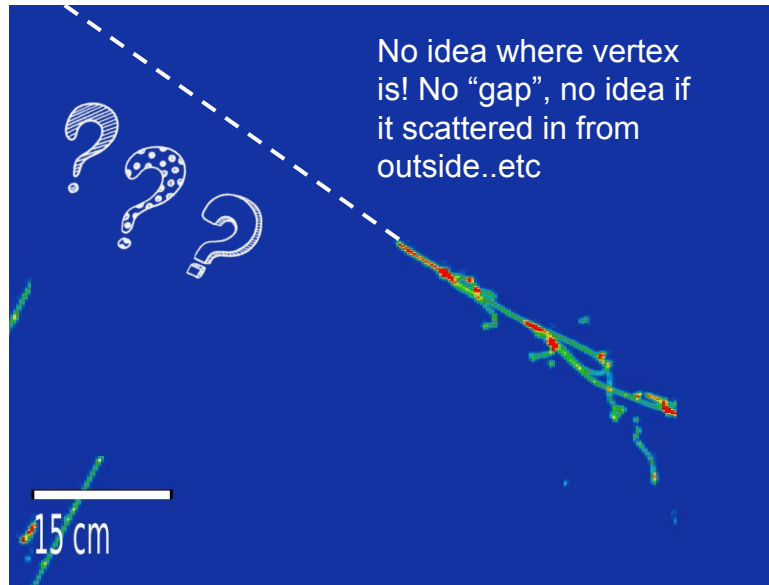
Two key facts to understand photon/ $e^+e^-$  searches

1.  $\pi^0$ 's are the dominant, and often irreducible, background
2. **Broadly speaking two categories of searches, photons with and without hadronic/proton activity**

## Photon + Protons



## Photon + No Protons

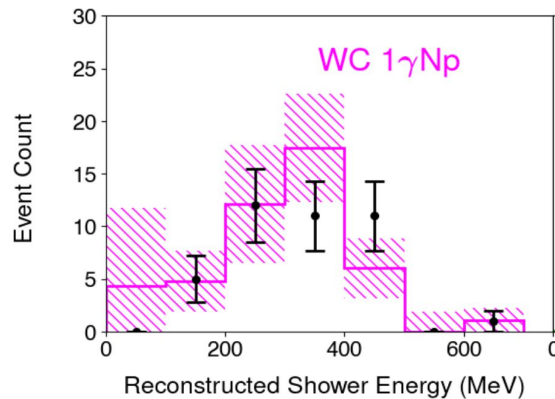
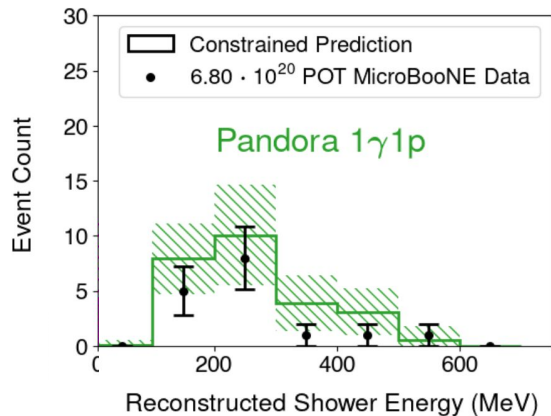


# Single-Photon with Protons



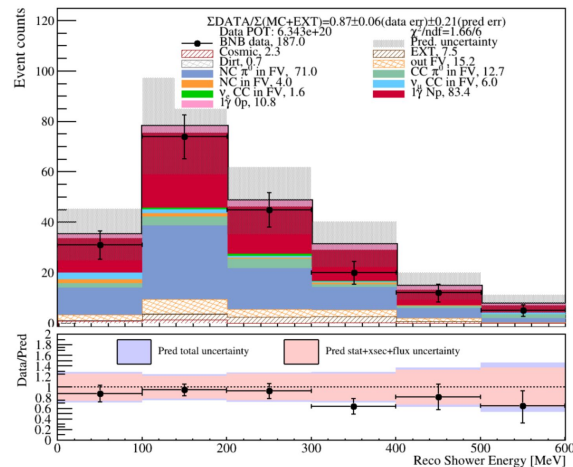
Enhanced Search for Neutral Current  $\Delta$  Radiative  
Single-Photon Production in MicroBooNE

○ [arXiv:2502.05750 \[hep-ex\] \(2025\)](https://arxiv.org/abs/2502.05750)



Inclusive Search for Anomalous  
Single-Photon Production in MicroBooNE

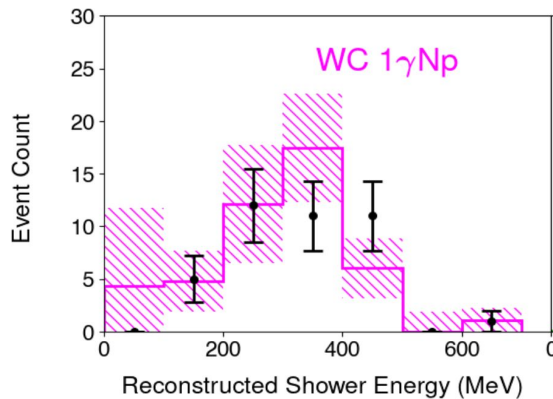
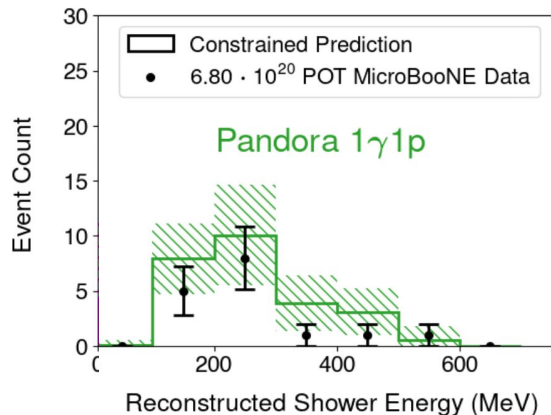
○ [arXiv:2502.06064 \[hep-ex\] \(2025\)](https://arxiv.org/abs/2502.06064)





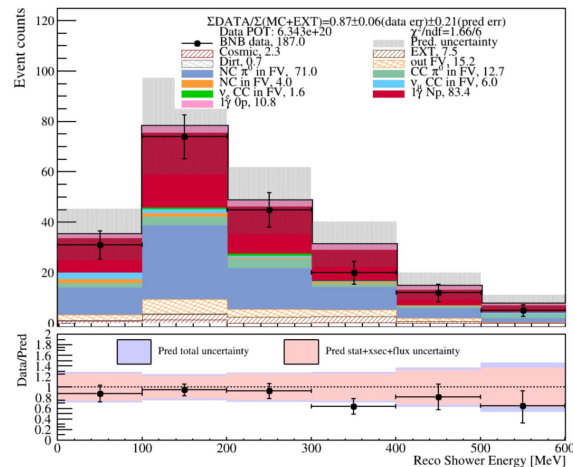
Enhanced Search for Neutral Current  $\Delta$  Radiative  
Single-Photon Production in MicroBooNE

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Inclusive Search for Anomalous  
Single-Photon Production in MicroBooNE

○ [arXiv:2502.06064 \[hep-ex\] \(2025\)](https://arxiv.org/abs/2502.06064)



- In all cases, consistent picture, good agreement with predictions
  - **No Anomalous Excess**
- Backgrounds  $\sim 70 \rightarrow 80\%$  NC  $\pi^0$  but we would expect to be sensitive to a MiniBooNE size signal,

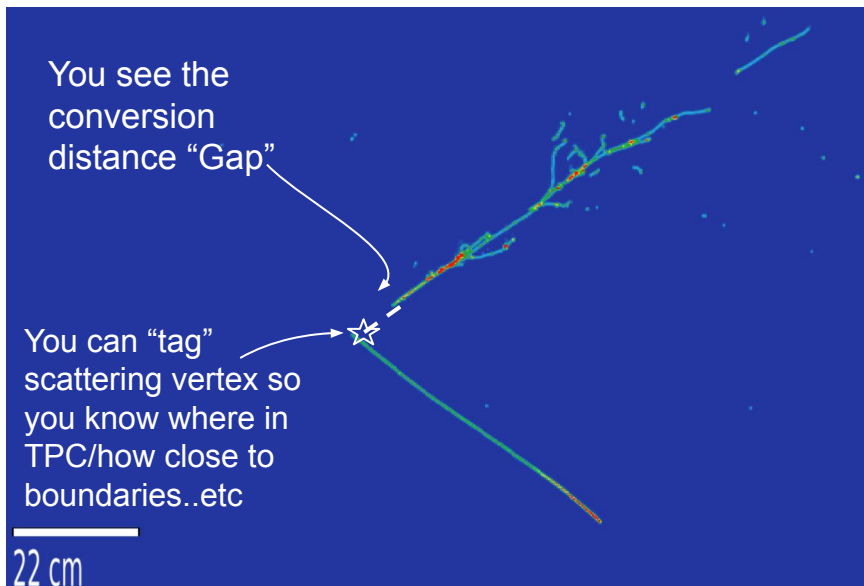




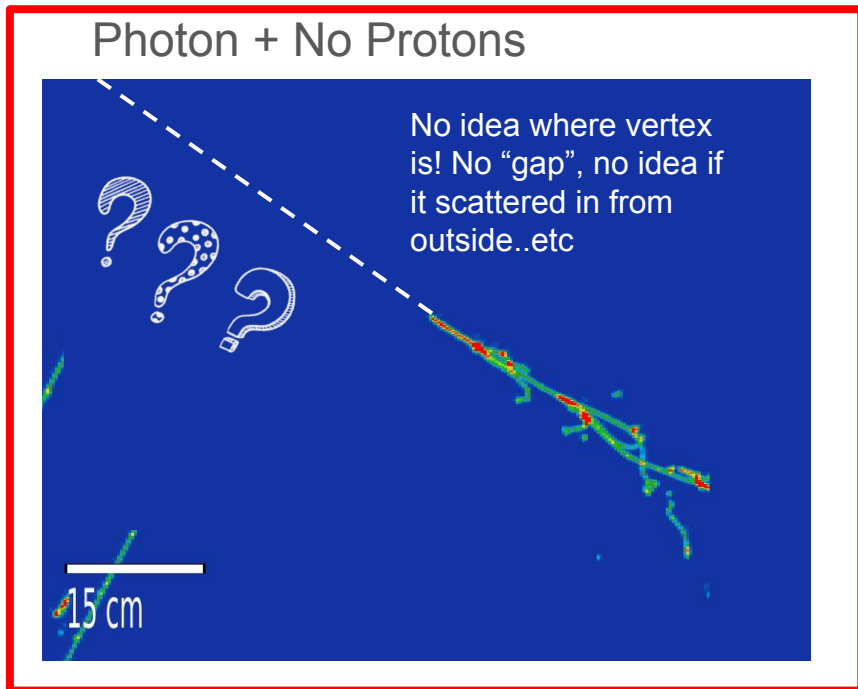
Two key facts to understand photon/e+e- searches

1.  $\pi^0$ 's are dominant, and often irreducible, background
2. **Broadly speaking two categories of searches, photons with and without hadronic/proton activity**

## Photon + Protons



## Photon + No Protons



# Single-Photon **without** Protons



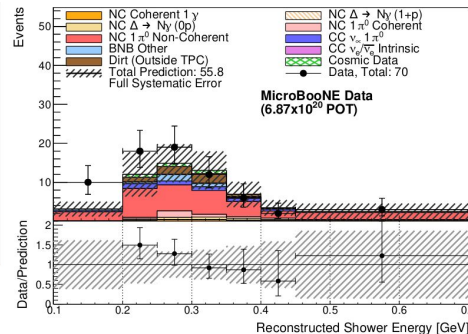
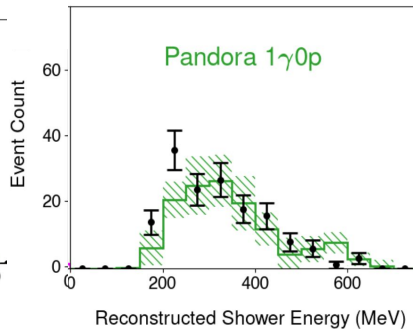
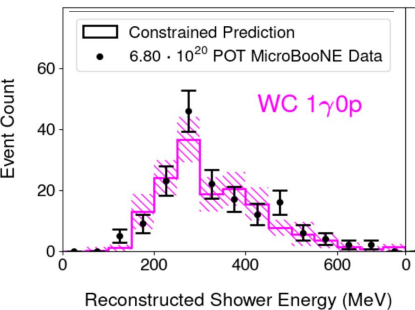
Enhanced Search for Neutral Current  $\Delta$  Rad  
Single-Photon Production in MicroBooNE

○ [arXiv:2502.05750 \[hep-ex\] \(2025\)](https://arxiv.org/abs/2502.05750)



First Search for Neutral Current **Coh**  
Single-Photon Production in MicroBooNE

○ [arXiv:2502.06091 \[hep-ex\] \(2025\)](https://arxiv.org/abs/2502.06091)



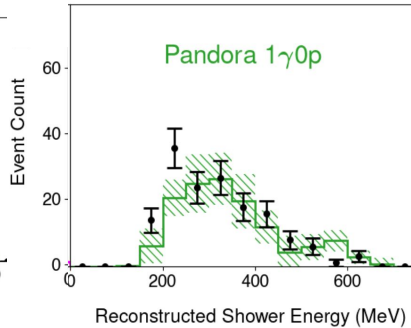
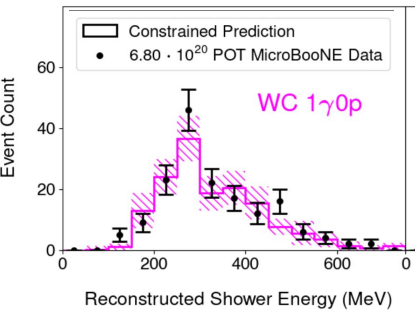
- *Broadly* good agreement with predictions...
- A few local bumps ~insignificant, and crucially **none of them were particularly sensitive!** NC  $\pi^0$  backgrounds much harder with no vertex!

# Single-Photon **without** Protons



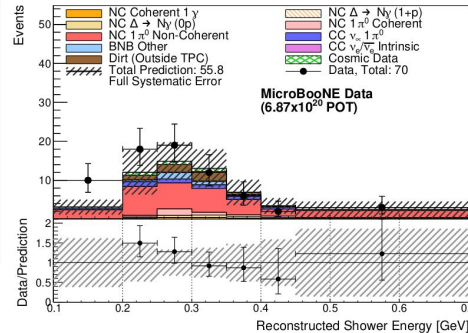
Enhanced Search for Neutral Current **Δ Rad** Single-Photon Production in MicroBooNE

- [arXiv:2502.05750 \[hep-ex\] \(2025\)](https://arxiv.org/abs/2502.05750)



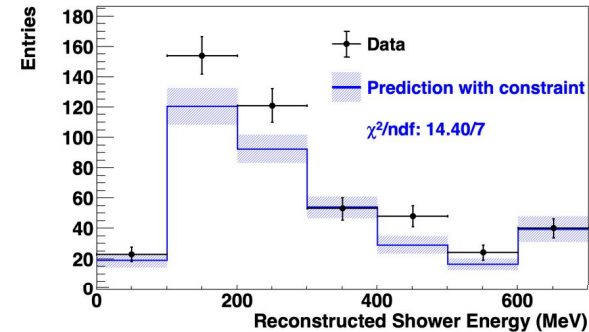
First Search for Neutral Current **Coh** Single-Photon Production in MicroBooNE

- [arXiv:2502.06091 \[hep-ex\] \(2025\)](https://arxiv.org/abs/2502.06091)



**Inclusive** Search for Anomalous Single-Photon Production in MicroBooNE

- [arXiv:2502.06064 \[hep-ex\] \(2025\)](https://arxiv.org/abs/2502.06064)



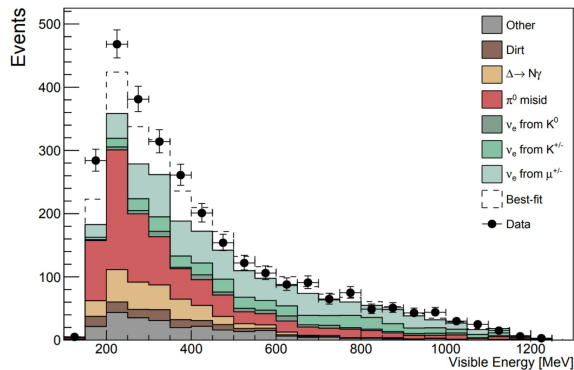
Excess of **93 ± 22(stat.) ± 35(syst.)** data events (**2.2σ**) from 0-600 MeV

- *Broadly* good agreement with predictions...
- A few local bumps ~insignificant, and crucially **none of them were particularly sensitive!** NC  $\pi^0$  backgrounds much harder with no vertex!
- The one analysis that *is more sensitive*, the inclusive analysis, sees a mild **2.2σ** excess, contained below 600 MeV

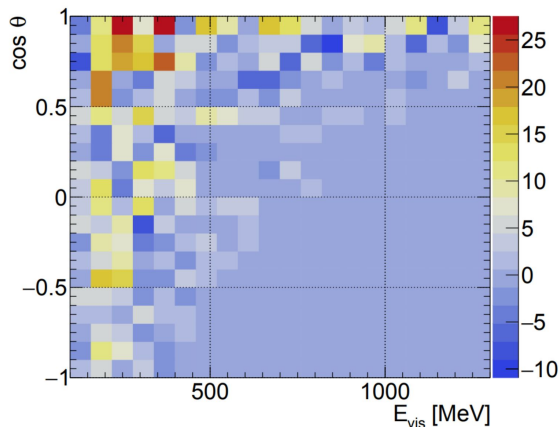
# Overall Kinematic Agreement

MiniBoONE

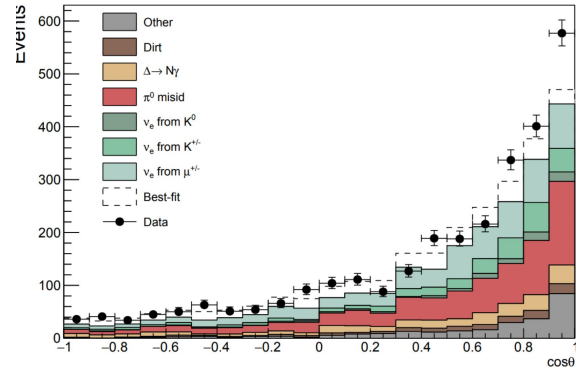
## Shower/Visible Energy



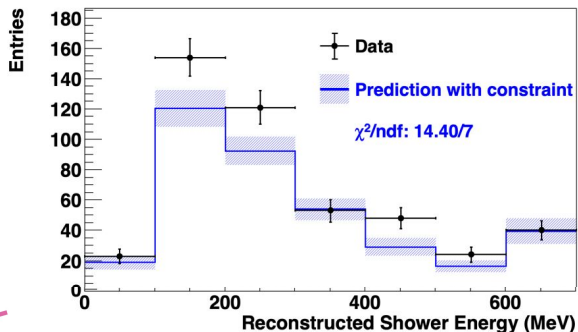
Excess



## Shower/Visible Angle



MicroBoONE



## Shower Energy vs Angle - Excess Events

