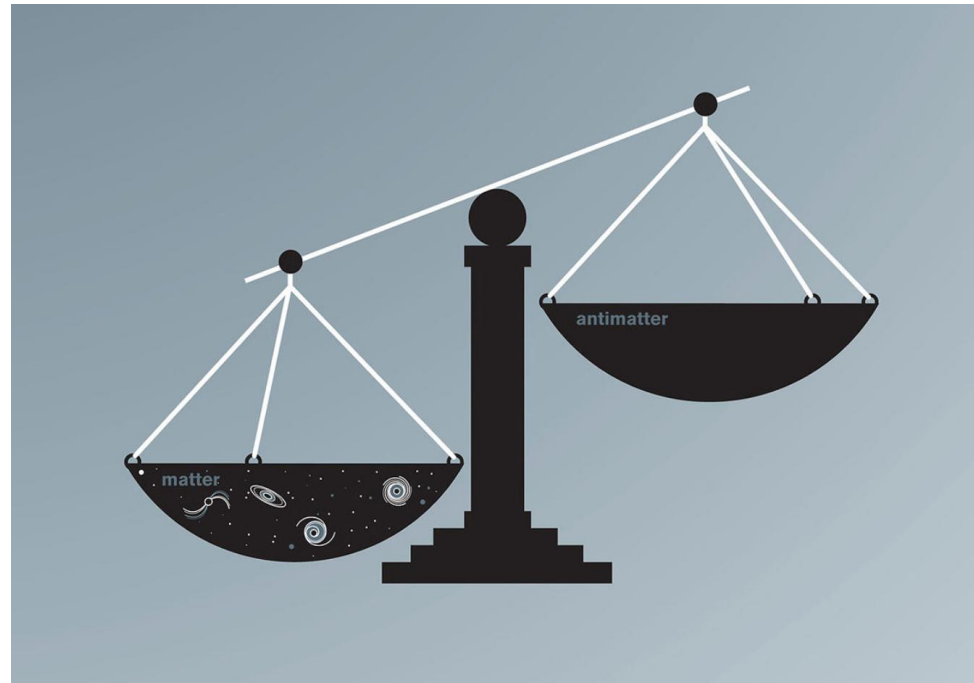


Latest Results From The CUORE Experiment

Vivek Sharma
University of Pittsburgh
PHENO 2026

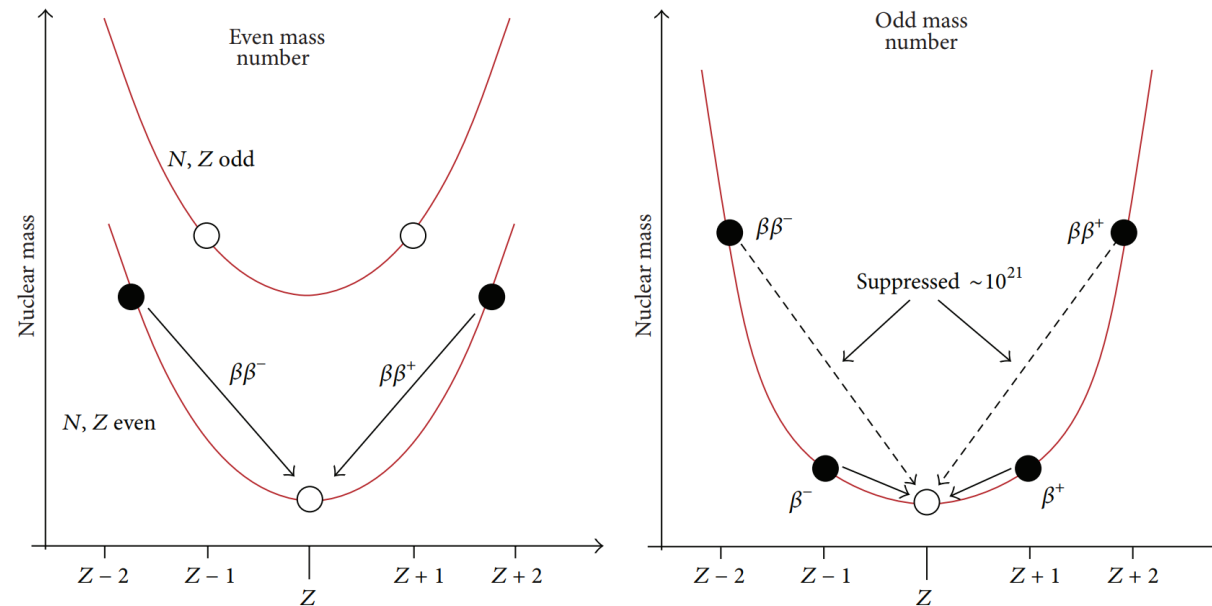
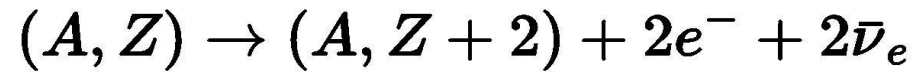
Lepton Number Violation

- **Lepton number** — the count of leptons minus antileptons — is conserved in all Standard Model interactions... *or is it?*
- **Leptogenesis** — in the early universe, heavy Majorana right-handed neutrinos decaying out of equilibrium can generate a lepton asymmetry
 - Can lead to baryon asymmetry through electroweak sphaleron processes
- $0\nu\beta\beta$ decay is a hypothetical process that **violates lepton number by 2 units**, indicating new physics beyond the Standard Model



2ν DOUBLE BETA DECAY

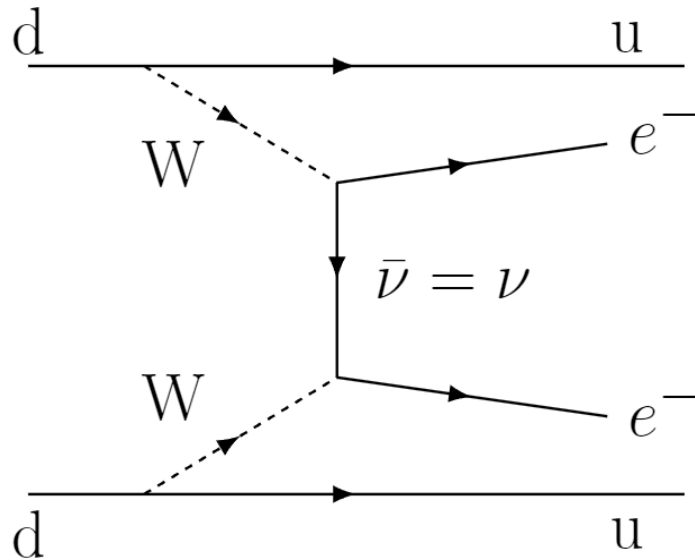
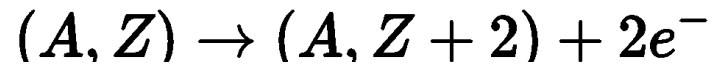
- Standard model 2nd order weak transition, extremely rare (half-life of **10¹⁹-10²² yr**)
- Observable when beta decay is forbidden or suppressed



Stefano Dell'Oro et al. "Neutrinoless Double Beta Decay: 2015 Review"

NEUTRINOLESS DOUBLE BETA DECAY

- Can occur if neutrinos are Majorana particles
- Lepton number violating process**
- Potentially impact understanding of origins of matter/anti-matter asymmetry
- Constrains **neutrino mass hierarchy, scale**
- Limit on $\langle m_{\beta\beta} \rangle$ can help rule out **Inverted Hierarchy** (model dependent)



Light neutrino exchange model

$$\frac{1}{T_{1/2}^{0\nu\beta\beta}} = G_{0\nu}(Q_{\beta\beta}, Z) |M_{0\nu}|^2 \frac{\langle m_{\beta\beta} \rangle^2}{m_e^2}$$

Phase space factor
Nuclear Matrix Elements

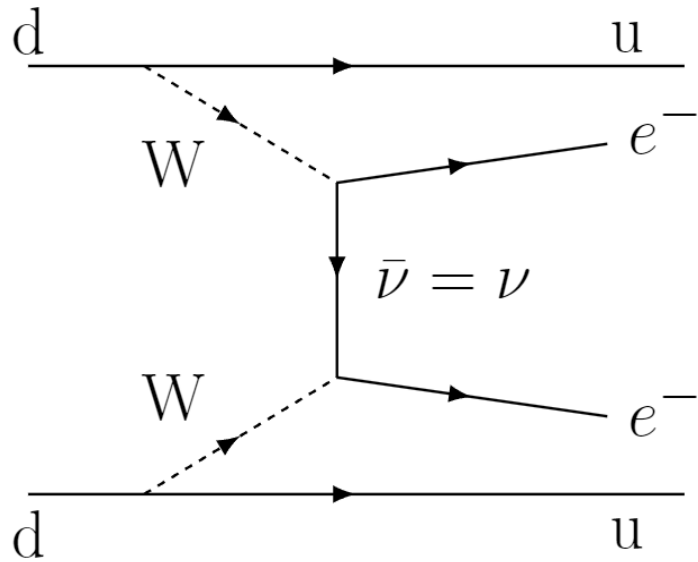
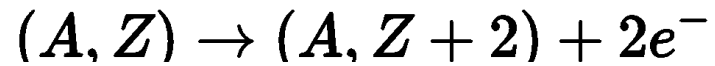
$$m_{\beta\beta} = \left| \sum_{i=1}^3 |U_{ei}|^2 m_i e^{j\alpha_i} \right|$$

Effective Majorana mass
Majorana phases

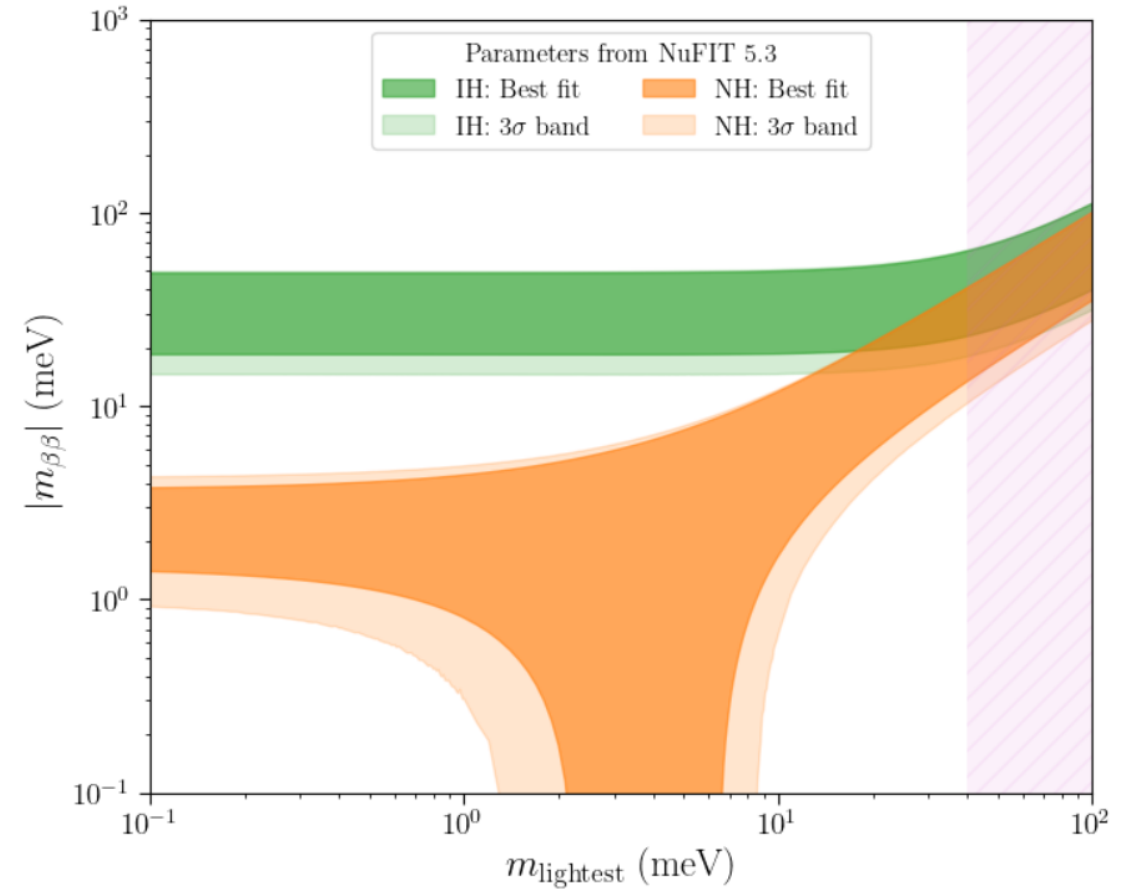
PMNS matrix
Individual neutrino masses

NEUTRINOLESS DOUBLE BETA DECAY

- Can occur if neutrinos are Majorana particles
- Lepton number violating process**
- Potentially impact understanding of origins of matter/anti-matter asymmetry
- Constrains **neutrino mass hierarchy, scale**
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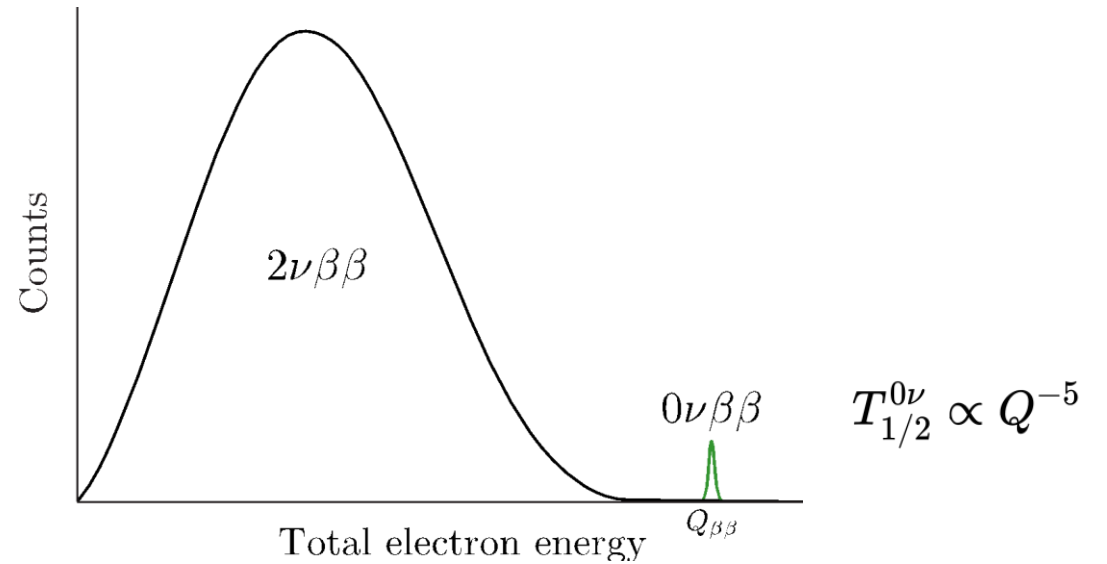


Light neutrino exchange model



DETECTION CHALLENGES

- $0\nu\beta\beta$ signal is the summed electron energy at Q-value
 - **Exceptional resolution** essential to differentiate between the $0\nu\beta\beta$ and $2\nu\beta\beta$ spectra
 - **High Q-value** for practically observable half-life and mitigate gamma-ray backgrounds
- **Low background** required for a detectable signal, going underground is necessary to mitigate cosmic ray bkg
- **Exposure** needs to be maximized
 - Large detector mass
 - Efficient duty cycle to lengthen livetime
- **Choice of isotope** should be compatible with detector technique



$$T_{1/2}^{0\nu}(1\sigma) = \ln(2) \frac{a\epsilon N_A \eta}{W} \sqrt{\frac{mt}{b\Delta E}}$$

Isotopic abundance

1-sigma sensitivity

Efficiency

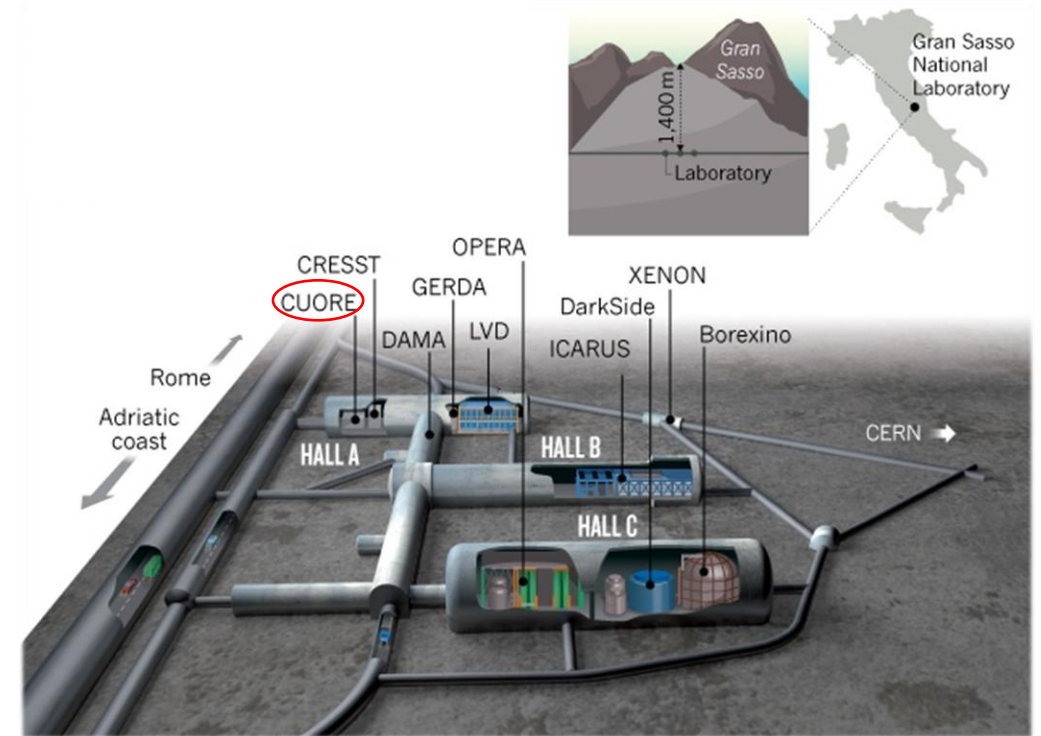
Detector mass

Bkg rate/energy/mass

Energy Resolution

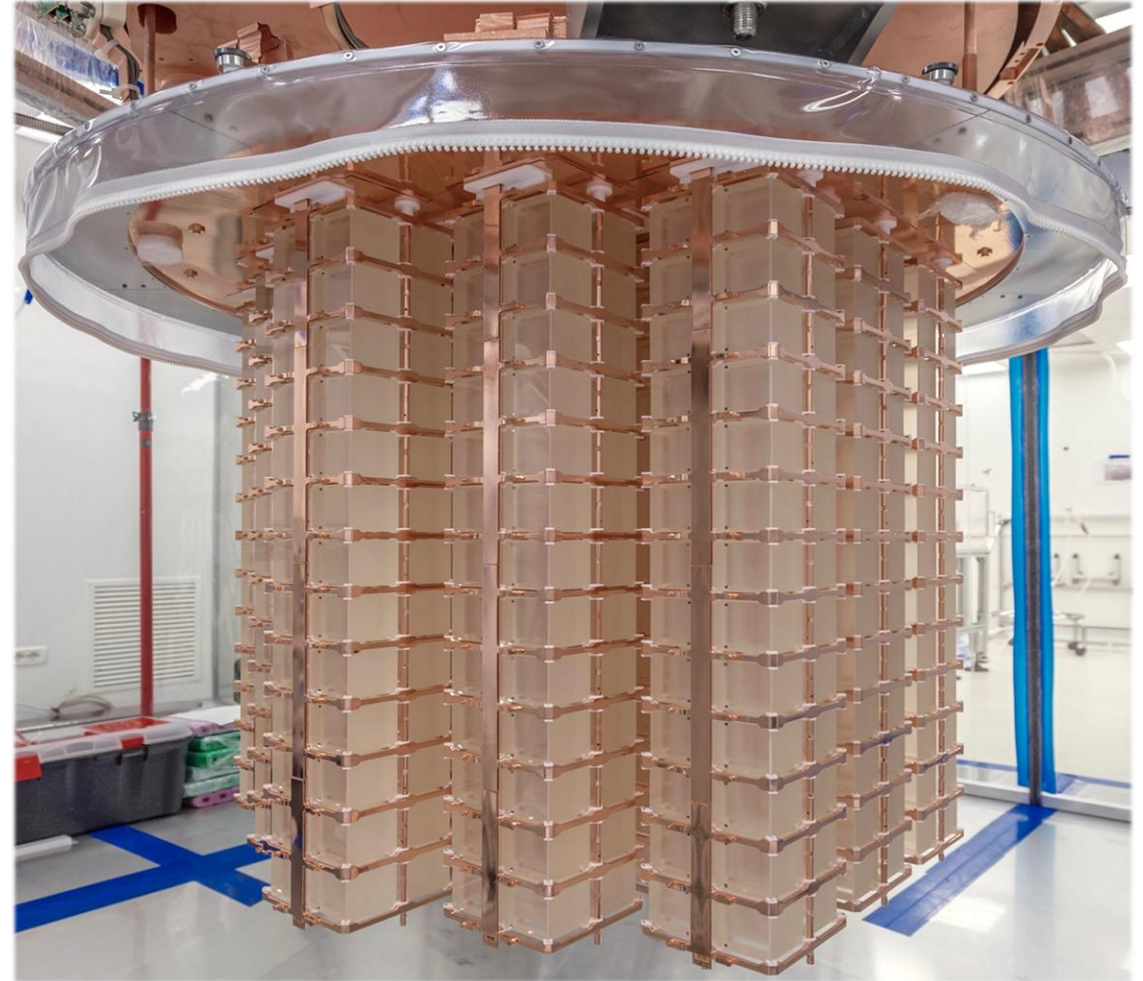
Livetime

- **Cryogenic Underground Observatory for Rare Events**
- Located at Hall A of Gran Sasso National Laboratory
- **~3600 m.w.e** of overburden, muon rate 6 orders of magnitude less than surface, extensive shielding
- ^{130}Te has a $\beta\beta$ Q-value of **2527.5 keV**
- **742 kg** TeO_2 , **206 kg** ^{130}Te (**34%** natural abundance)



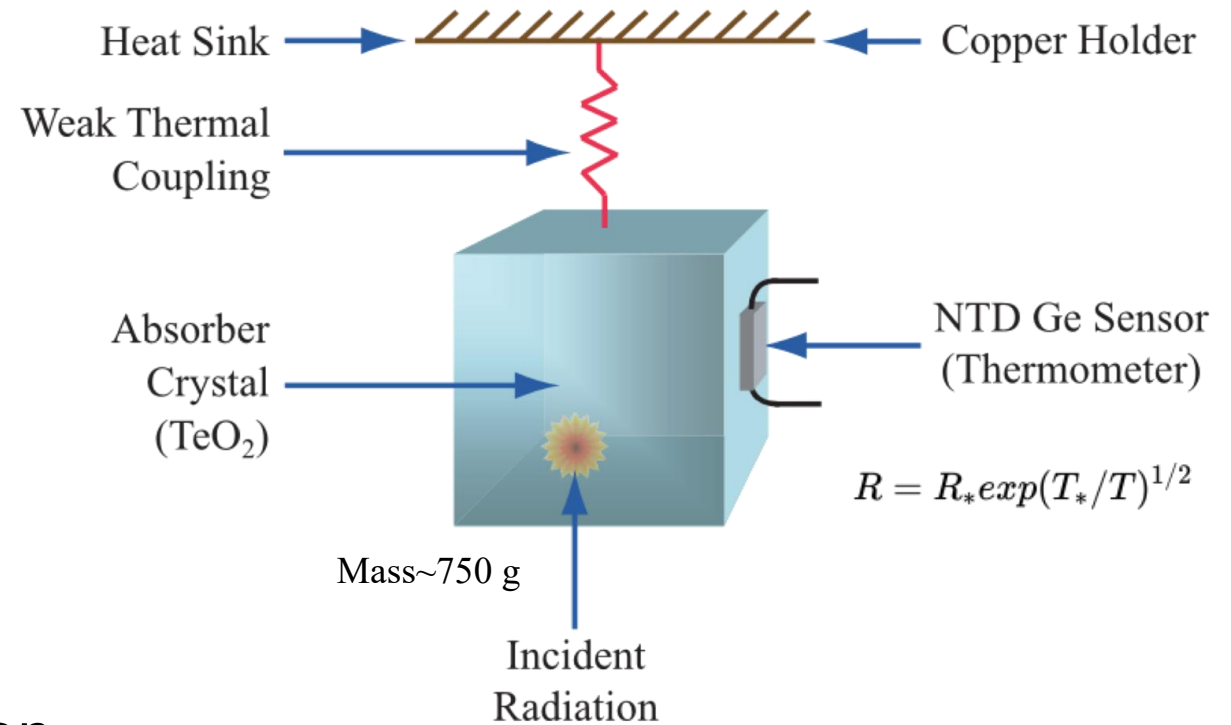
Images courtesy of LNGS: <https://www.lngs.infn.it/en>

- 988 natural TeO_2 crystals
 - Total mass: **742 kg**
 - ^{130}Te mass: **206 kg**
 - **$5 \times 5 \times 5 \text{ cm}^3$** , arranged in 19 towers
- Housed in copper frame and held in place by PTFE spacers
 - Copper linked to thermal bath
- Tightly spaced crystals allow for coincidences to be exploited for background reduction



DETECTION PRINCIPLE - CUORE

- 988 TeO_2 crystals operated as bolometers; energy deposited is registered as temperature change
 - Read out by a NTD (**N**eutron **T**ransmutation **D**oped) Ge thermometer
- Signal amplitude and detector resolution depend strongly on temperature (Debye's Law)
 - $C \propto T^3$
 - Detector operated at **~15 mK**
- In CUORE, we observe an average resolution of **~7.5 keV FWHM at 2615 keV[†]**



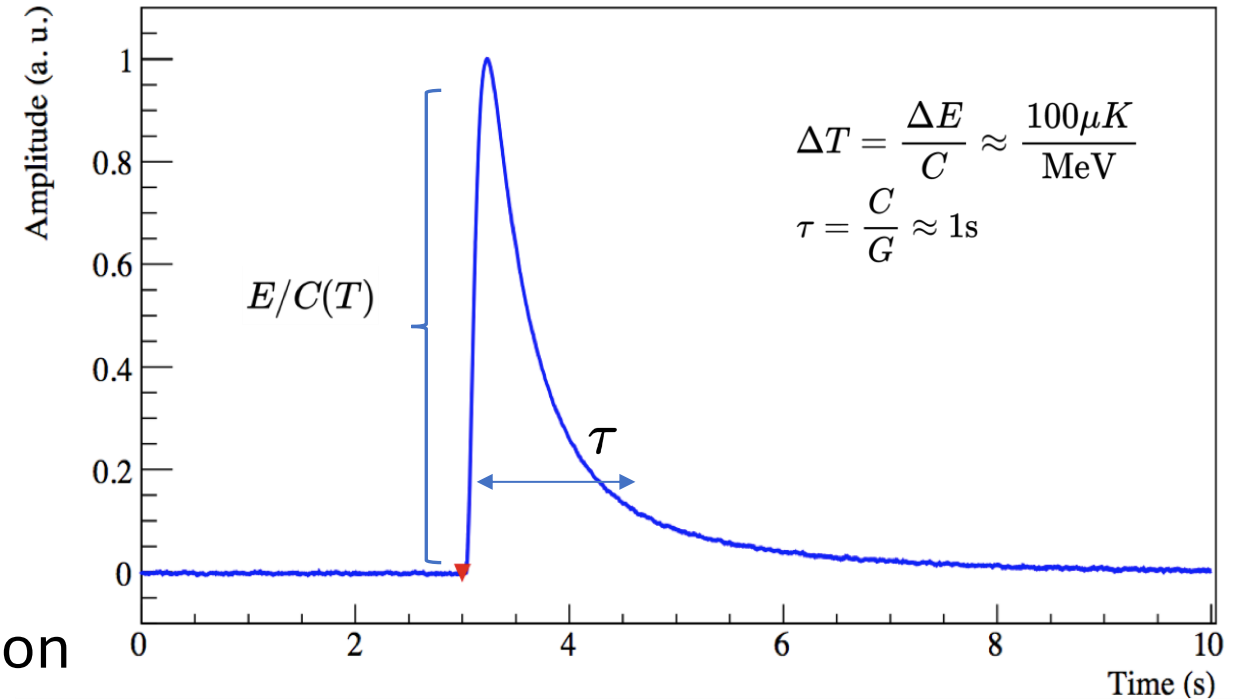
Schematic of a CUORE bolometer.
(cuore-lngs.infn.it)

[†]CUORE collaboration

<https://doi.org/10.48550/arXiv.2404.04453>

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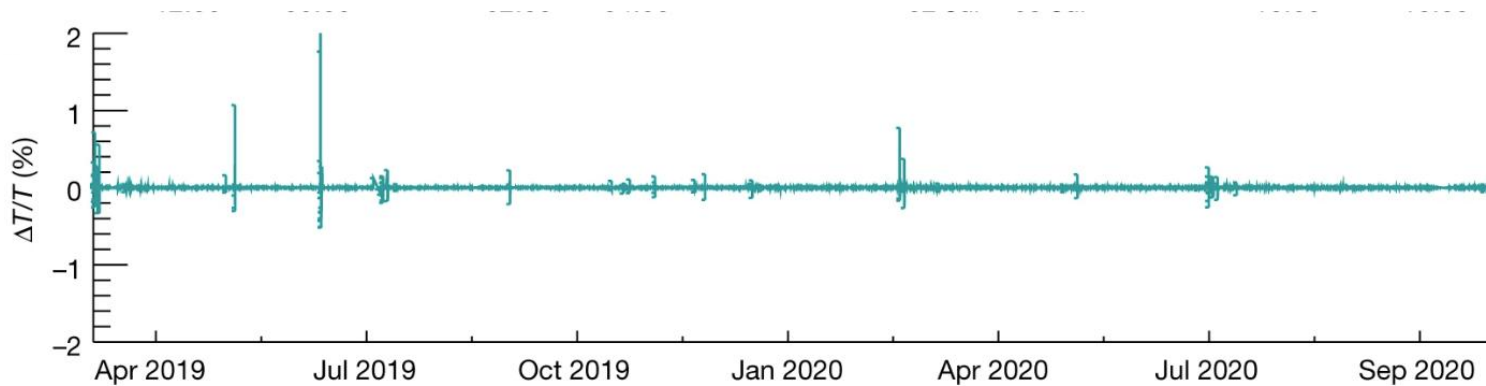


Example of a signal pulse

[†]CUORE collaboration

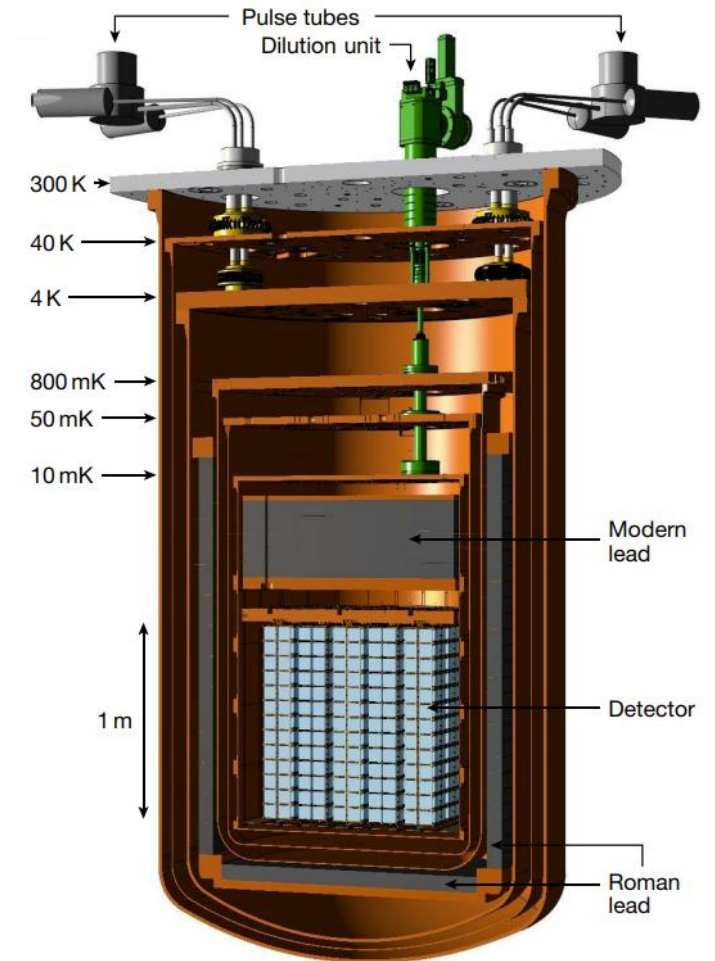
<https://doi.org/10.48550/arXiv.2404.04453>

- Operating temperature of **~15 mK** is achieved via multistage cryogen-free dilution refrigerator
 - Multistage design shields from thermal radiation
 - Experimental volume of $\sim 1 \text{ m}^3$ and payload of ~ 1.5 tonne
 - Demonstrated stability over years of data taking
- Multiple layers of shielding for background mitigation



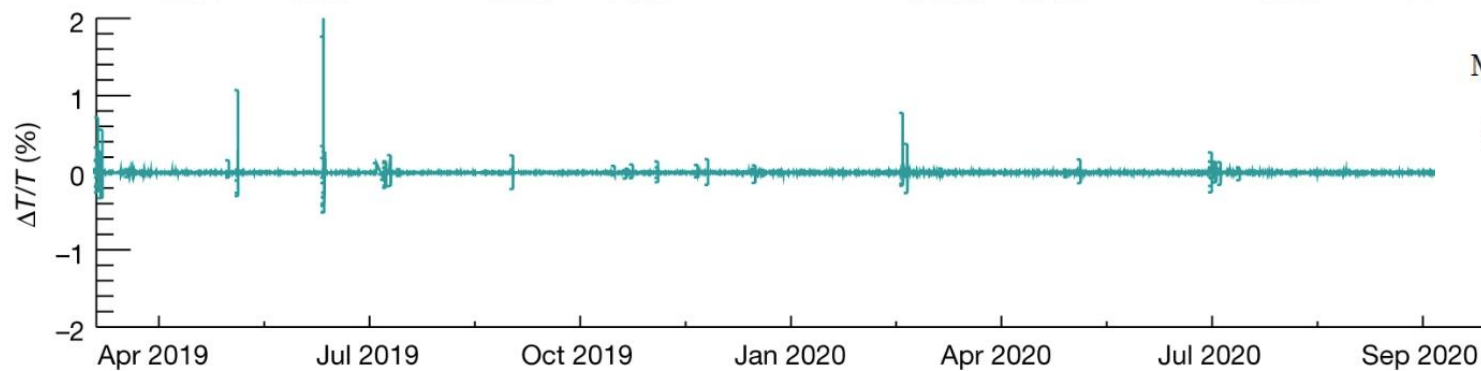
CUORE relative temperature fluctuation over time*

*CUORE collaboration
<https://www.nature.com/articles/s41586-022-04497-4.pdf>



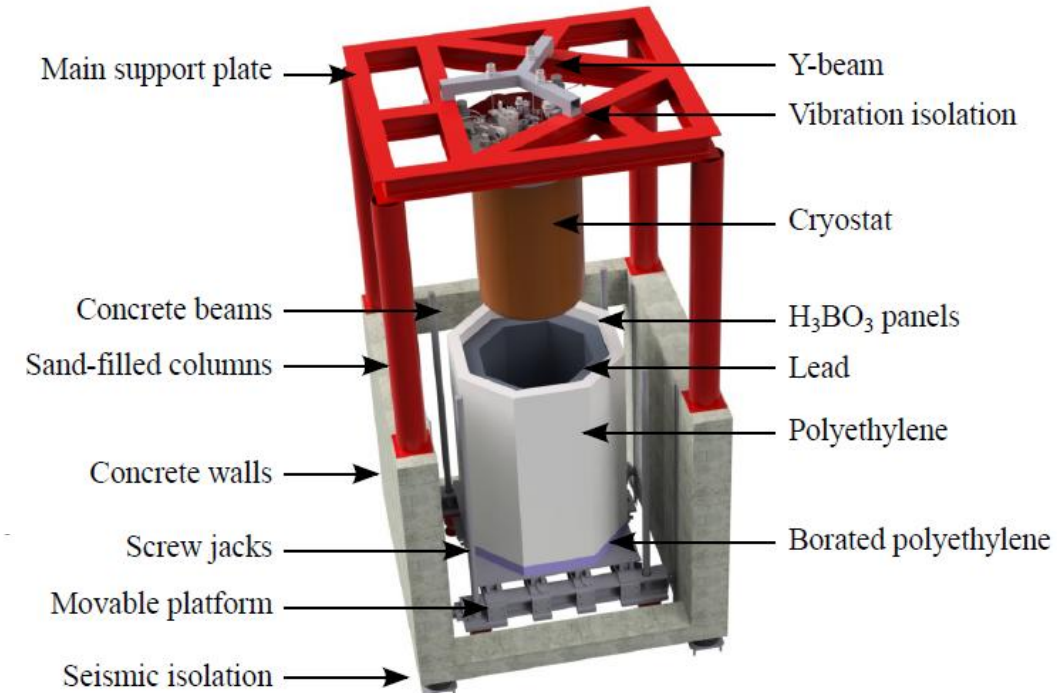
The CUORE cryostat*

- Operating temperature of **~15 mK** is achieved via multistage cryogen-free dilution refrigerator
 - Multistage design shields from thermal radiation
 - Experimental volume of $\sim 1 \text{ m}^3$ and payload of ~ 1.5 tonne
 - Demonstrated stability over years of data taking
- Multiple layers of shielding for background mitigation



CUORE relative temperature fluctuation over time*

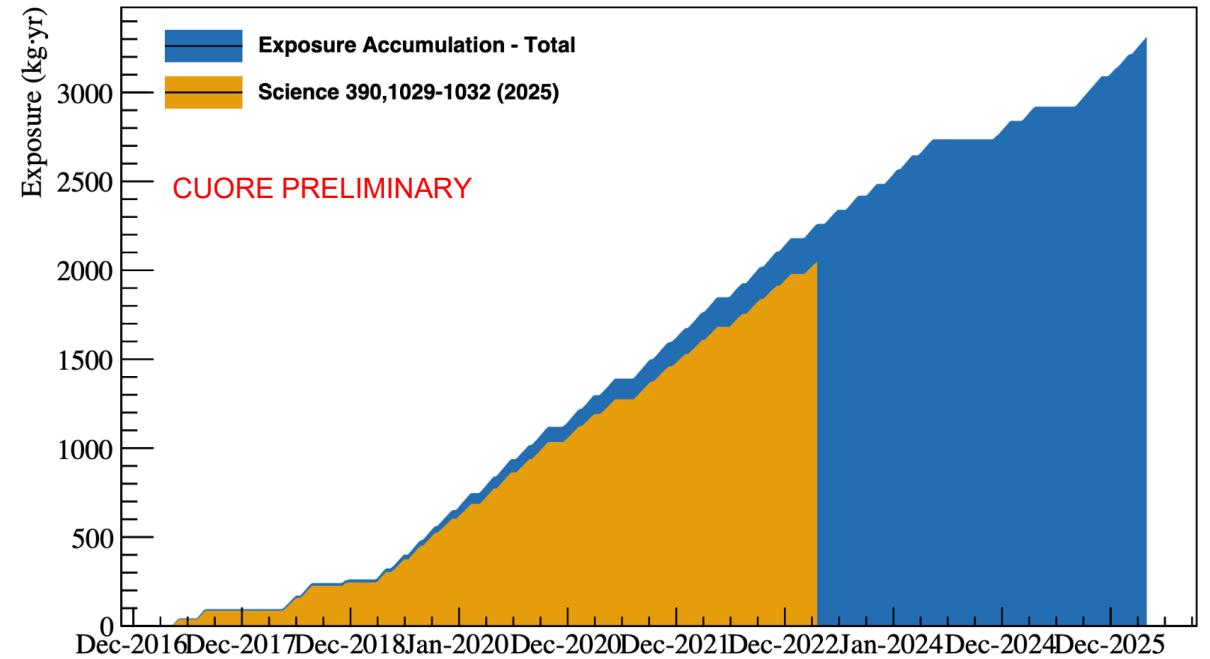
*CUORE collaboration
<https://www.nature.com/articles/s41586-022-04497-4.pdf>



External structure of CUORE

<https://doi.org/10.1016/j.cryogenics.2019.06.011>

- Data taking began in 2017
 - Software and hardware optimizations since have improved stability of data taking
- Steady data taking since 2019 with $>\sim 90\%$ uptime
- Data cycle: 1-2 months of "physics" data between ~ 4 days of calibration



2 TON YEAR RESULTS

2nd ton.yr results:

- Blinded analysis to prevent bias
- ROI (2465, 2575) keV salted with ^{208}Tl 2615 keV peak
- Fit ROI events with: ^{130}Te $0\nu\beta\beta$ peak, ^{60}Co peak, linear background
- Finalize model parameters before unblinding

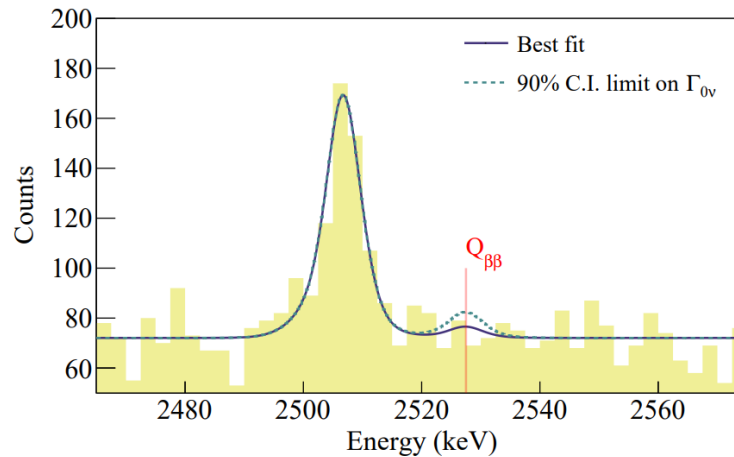
Analyzed exposure: 2039.0 kg.yr

Half-life limit: $T_{1/2} > 3.8 \times 10^{25}$ yr (90% C.I.)

Median Exclusion Sensitivity: 4.4×10^{25} yr (90% C.I.)

$m_{\beta\beta} < 70 - 250$ meV (90% C.I.)

- Light Majorana neutrino exchange model



Limits on other isotopes:

GERDA Collaboration, Phys. Rev. Lett. 125, 252502 (2020)

<https://doi.org/10.1103/PhysRevLett.125.252502>

CUPID-Mo Collaboration <https://journals.aps.org/prl/abstract/10.1103/PhysRevLett.126.181802>

CUPID-0 Collaboration, Phys. Rev. Lett. 123, 032501 (2019)

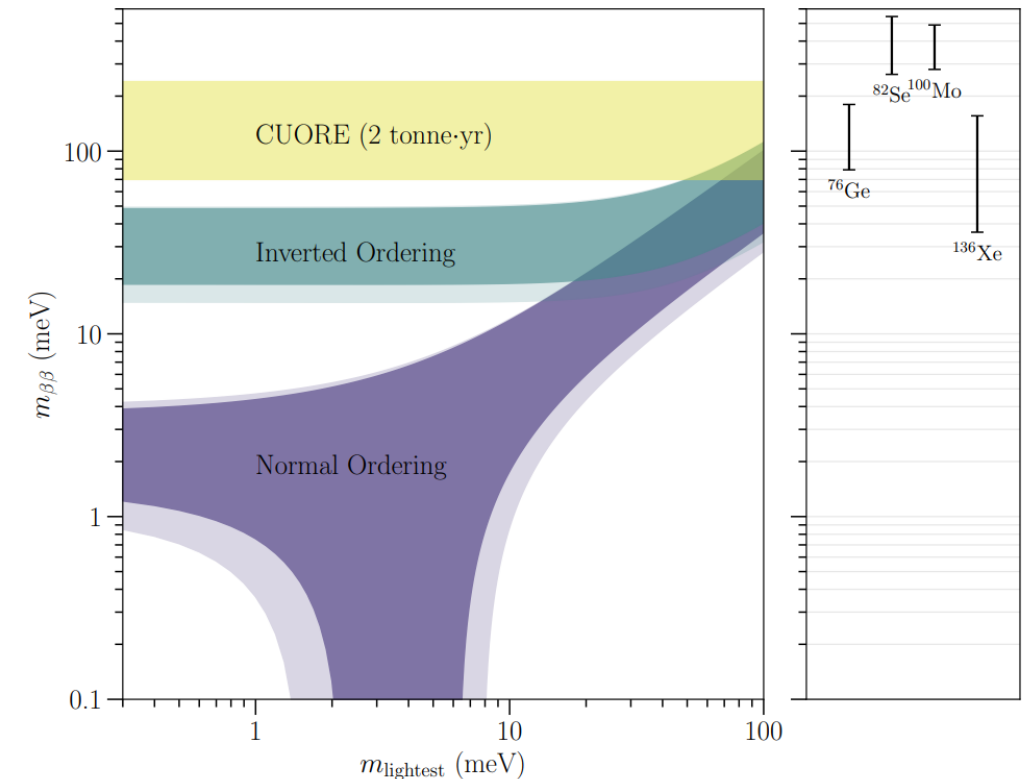
<https://doi.org/10.1103/PhysRevLett.123.032501>

KamLAND-Zen Collaboration, Phys. Rev. Lett. 117, 082503 (2016)

<https://doi.org/10.1103/PhysRevLett.117.082503>

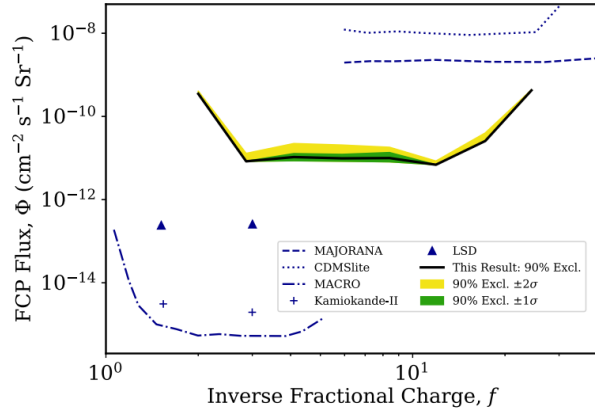
Oscillation parameters:

Esteban, I. et al., J. High En. Phys. 2020 (178) [https://doi.org/10.1007/JHEP09\(2020\)178](https://doi.org/10.1007/JHEP09(2020)178)



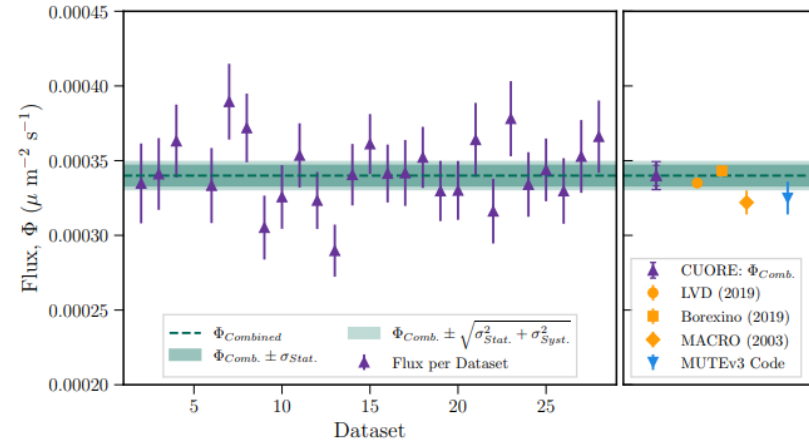
†CUORE collaboration
<https://doi.org/10.48550/arXiv.2404.04453>

Secondary Physics Studies



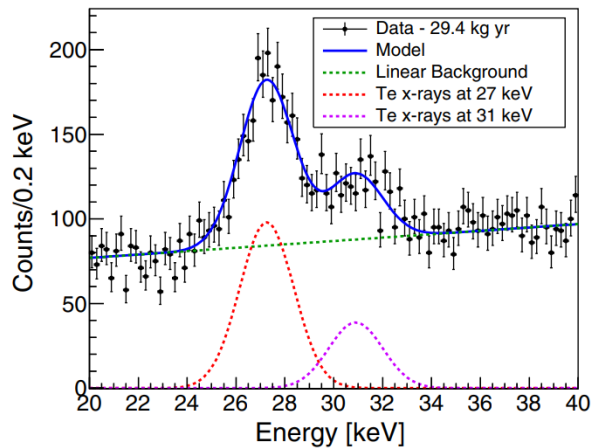
Multi-spectral search for fractionally charged particles

Phys. Rev. Lett. 133, 241801



Muon Event Reconstruction

arxiv:2509.05528

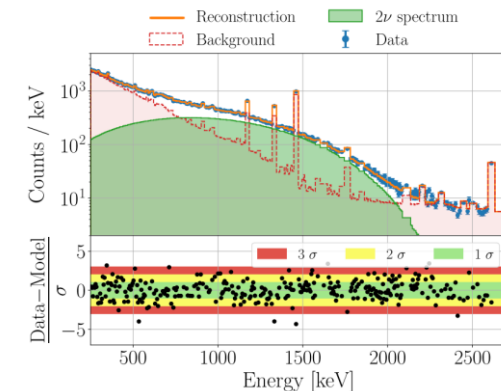


Low energy analyses to optimize sensitivity for solar axions, WIMP searches

Phys. Rev. D 113, 012012

Other ββ searches

- ^{130}Te SM-allowed $2\nu\beta\beta$ (**arXiv:2503.24137v1**)
- ^{130}Te decay to 1st excited state (**Eur. Phys. J. C 81, 567**)
- ^{128}Te $0\nu\beta\beta$ decay to ground state (**Phys. Rev. Lett. 129, 222501**)
- ^{120}Te $0\nu\beta^+$ EC decay to ground state (**Phys. Rev. C 105, 065504**)



THANK YOU!

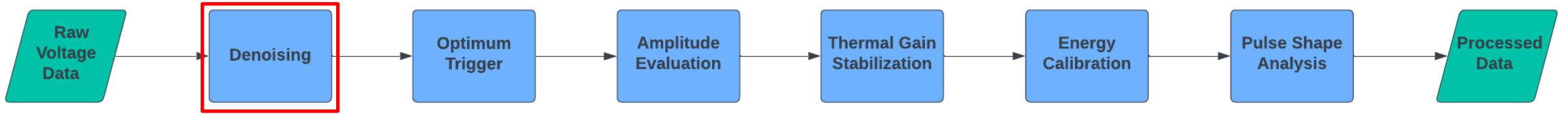


5/11/2026

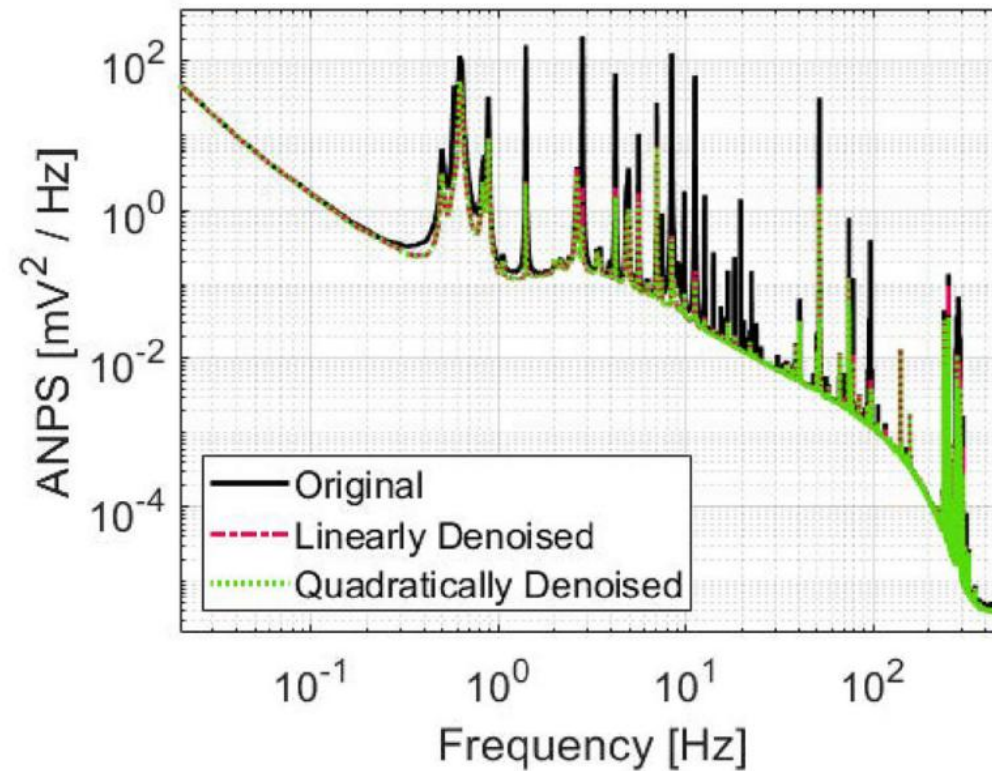
Vivek Sharma - PHENO 2026

EXTRA SLIDES

DATA PROCESSING

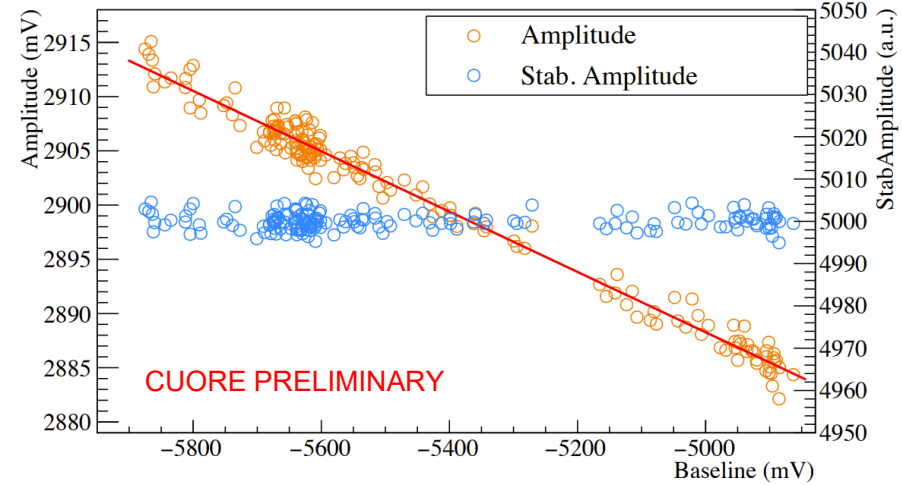
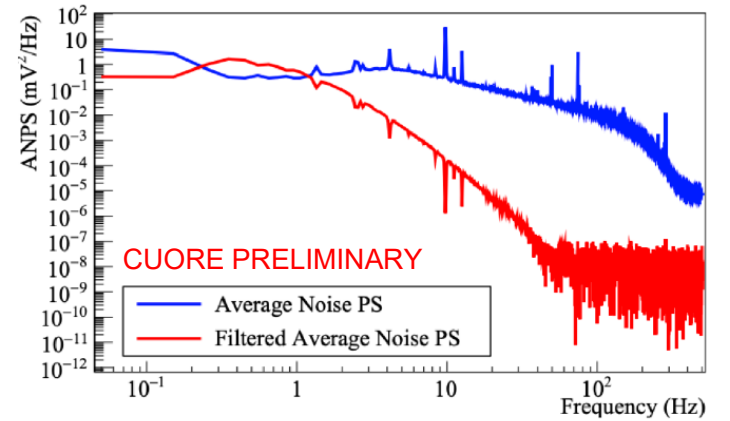
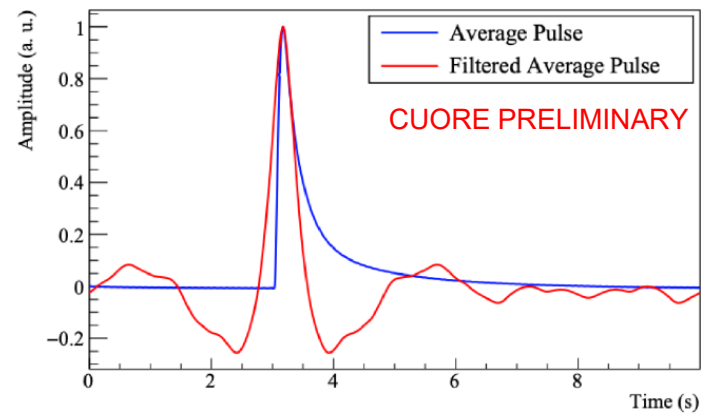
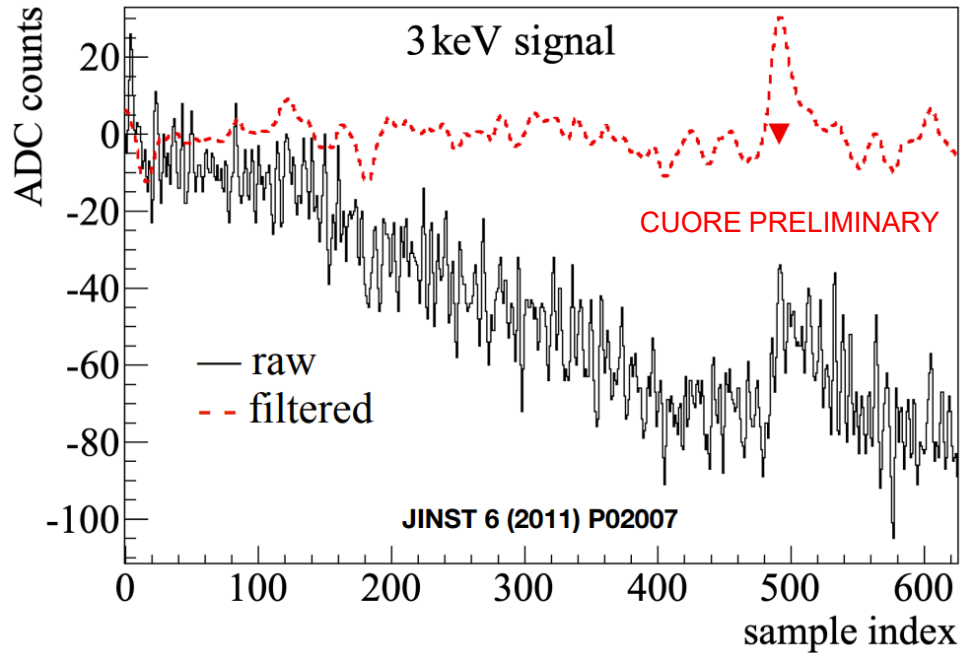
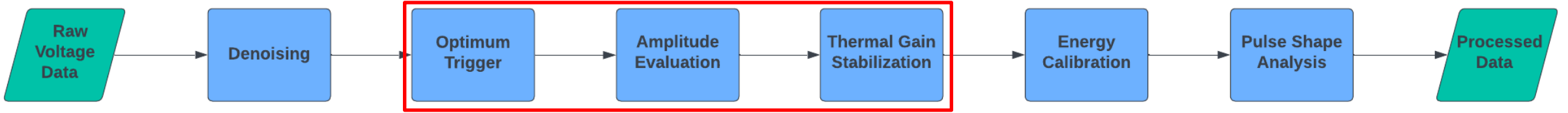


- Predict noise contribution by correlating noise between auxiliary devices and correlating with calorimeter noise
 - Produce less noisy data with improved resolution

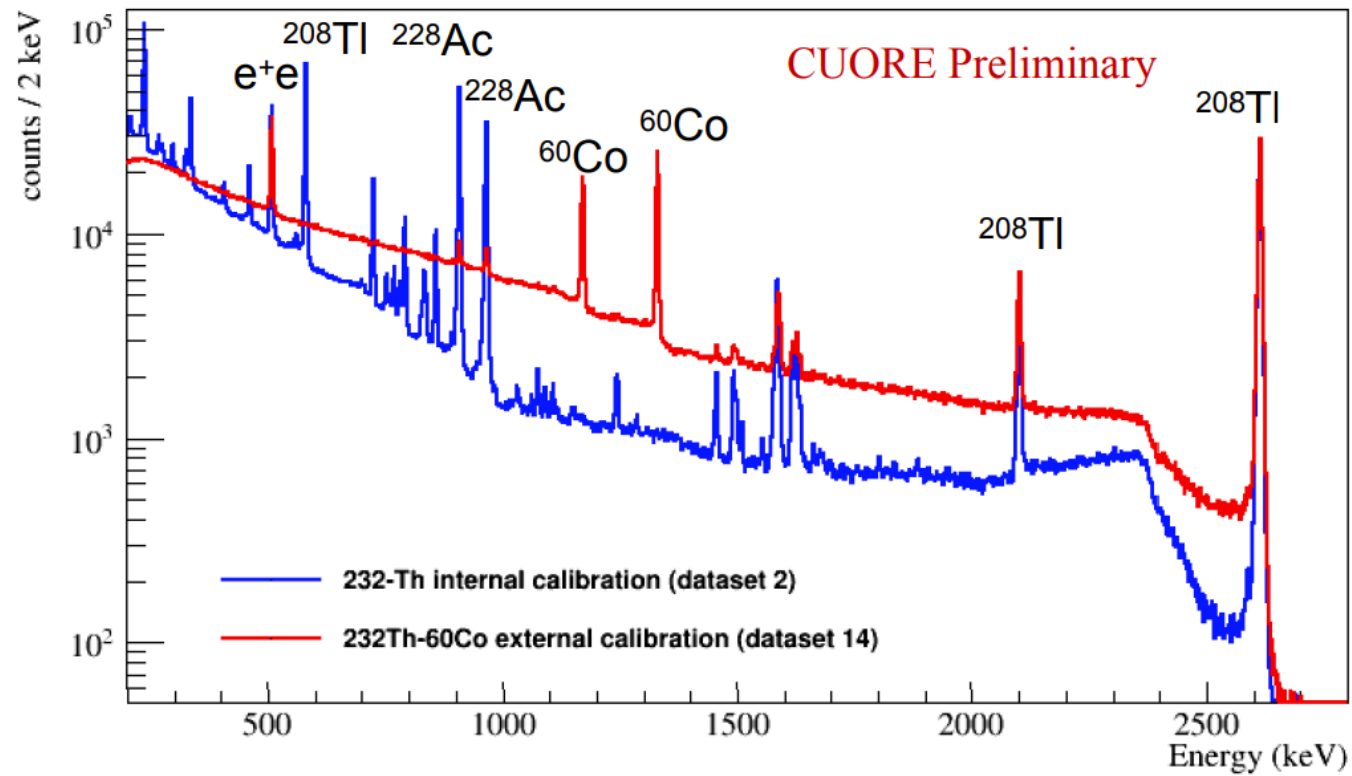
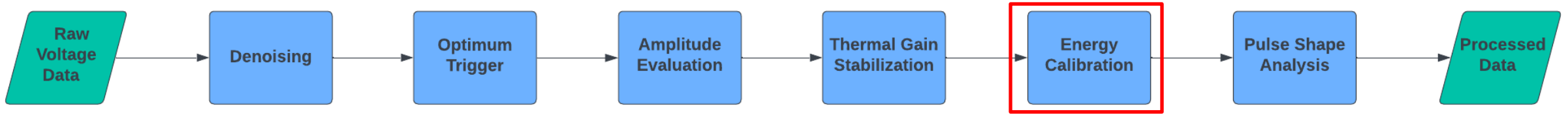


<https://doi.org/10.1140/epjc/s10052-024-12595-y>

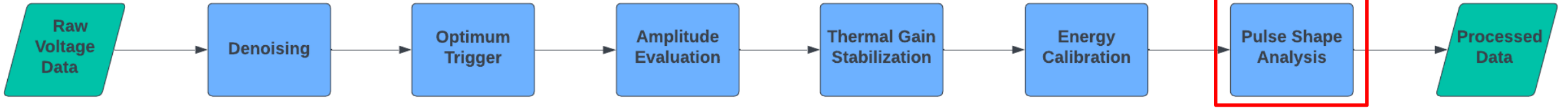
DATA PROCESSING



DATA PROCESSING

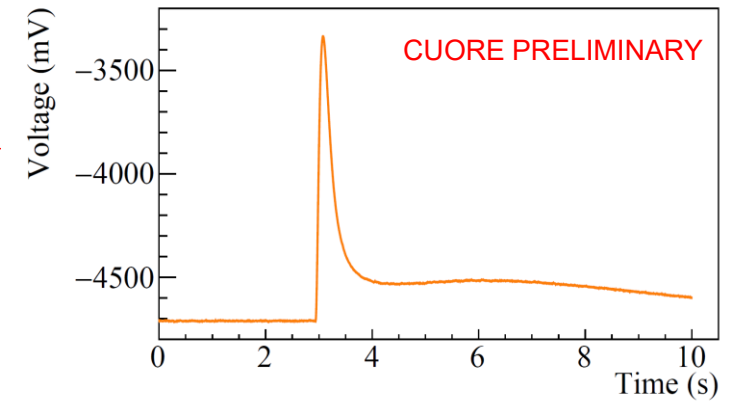
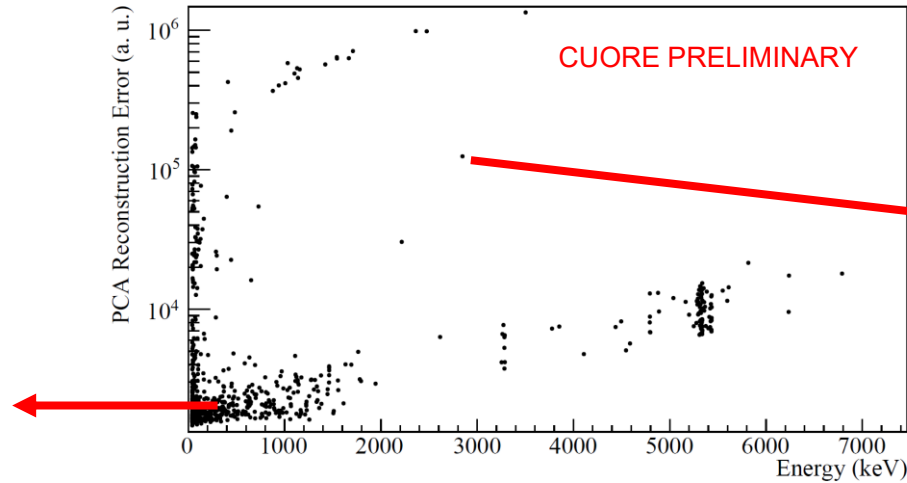
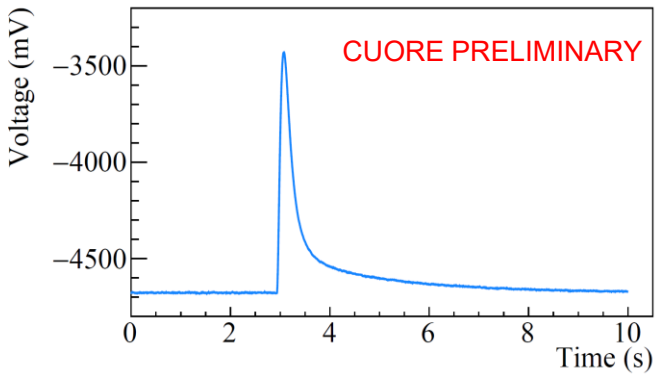


DATA PROCESSING

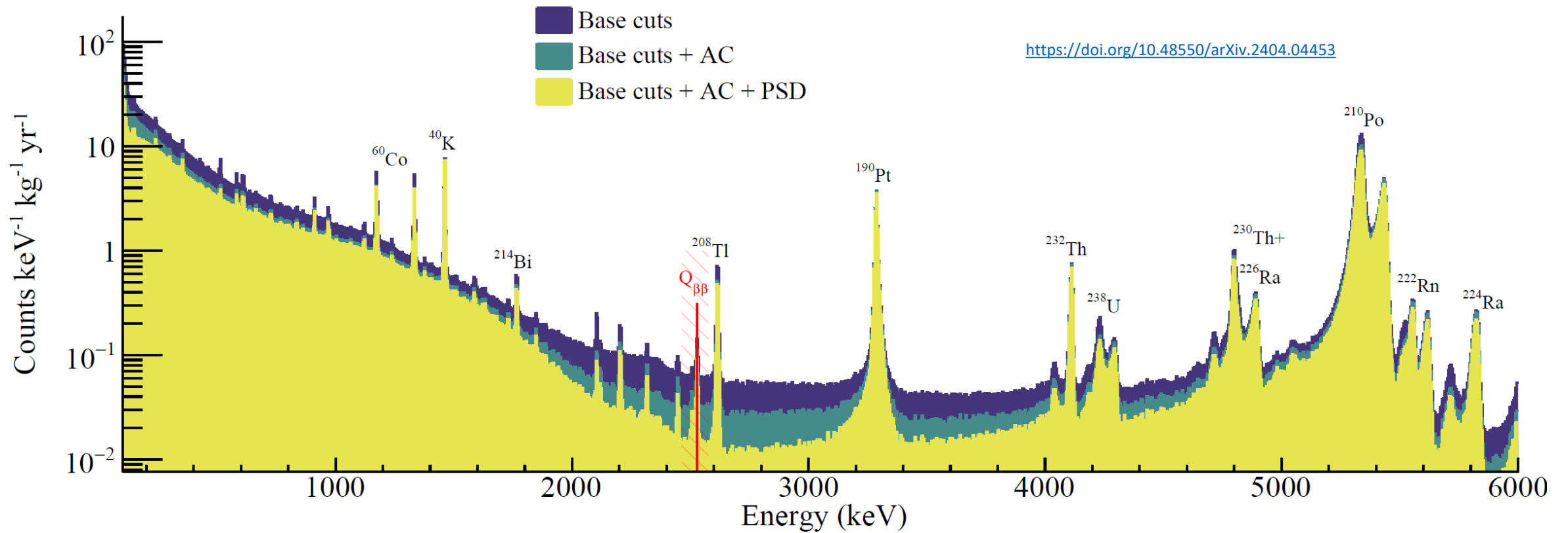
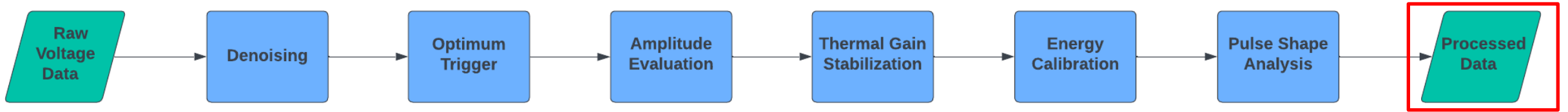


$$RE = \sqrt{\sum_{d=1}^D [x_d - (\mathbf{x} \cdot \mathbf{u}_{AP})u_d]^2}$$

Reconstruction Error



DATA PROCESSING



NEUTRINOLESS DOUBLE BETA DECAY

- Experimental observable is **decay rate**
 - Depends on effective Majorana mass ($m_{\beta\beta}$)
 - Directly related to absolute neutrino mass scale
- Additional dependency on various nuclear and atomic effects.
- Limit on ($m_{\beta\beta}$) can help rule out Inverted Hierarchy (model dependent)

$$\frac{1}{T_{1/2}^{0\nu\beta\beta}} = G_{0\nu}(Q_{\beta\beta}, Z) |M_{0\nu}|^2 \frac{\langle m_{\beta\beta} \rangle^2}{m_e^2}$$

Phase space factor
Nuclear Matrix Elements

$$m_{\beta\beta} = \left| \sum_{i=1}^3 |U_{ei}|^2 m_i e^{j\alpha_i} \right|$$

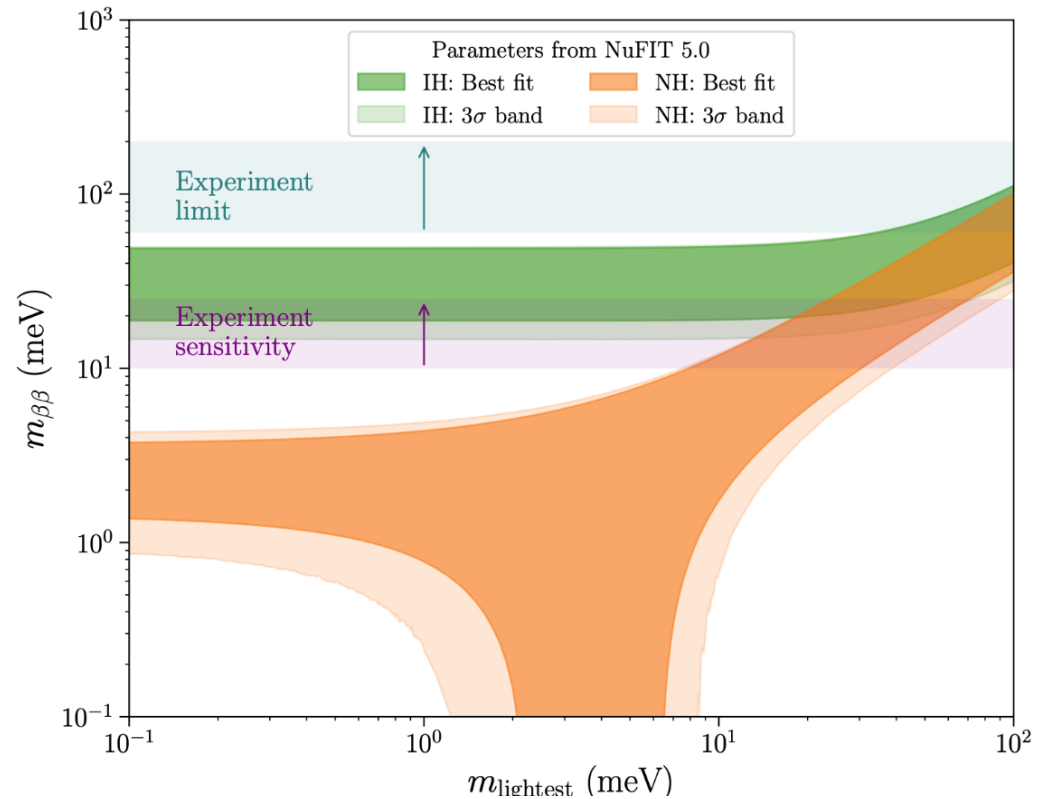
Effective Majorana mass
Majorana phases

PMNS matrix
Individual neutrino masses

<https://github.com/toej93/LobsterPlot>

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<https://github.com/toej93/LobsterPlot>

$$T_{1/2}^{0\nu}(1\sigma) = \ln(2) \frac{a\epsilon N_A \eta}{W} \sqrt{\frac{mt}{b\Delta E}}$$

1-sigma sensitivity


Isotopic abundance

Detector mass

- Active Mass: 742 kg TeO₂
- Isotopic abundance: 34%, 206 kg ¹³⁰Te

$$T_{1/2}^{0\nu}(1\sigma) = \ln(2) \frac{a\epsilon N_A \eta}{W} \sqrt{\frac{mt}{b\Delta E}}$$

Livetime



1-sigma sensitivity

- Greater than 90% duty cycle
- Cryogen-free cryostat (No regular downtime for invasive cryogenic maintenance)

$$T_{1/2}^{0\nu}(1\sigma) = \ln(2) \frac{a\epsilon N_A \eta}{W} \sqrt{\frac{mt}{b\Delta E}}$$

1-sigma sensitivity

Bkg rate/energy/mass

- Radiopurity controls on materials and assembly
 - ~3600 m.w.e overburden
 - Extensive shielding
- BI: $\sim 10^{-2}$ counts keV⁻¹ kg⁻¹ yr⁻¹

$$T_{1/2}^{0\nu}(1\sigma) = \ln(2) \frac{a\epsilon N_A \eta}{W} \sqrt{\frac{mt}{b\Delta E}}$$

1-sigma sensitivity

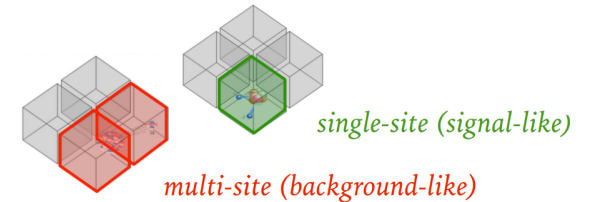
Energy
Resolution

- Crystals operated at ~15 mK
 - Denoising techniques
- $Q_{\beta\beta}$ relative energy resolution: 0.3%

$$T_{1/2}^{0\nu}(1\sigma) = \ln(2) \frac{a\epsilon N_A \eta}{W} \sqrt{\frac{mt}{b\Delta E}}$$

1-sigma sensitivity

Efficiency

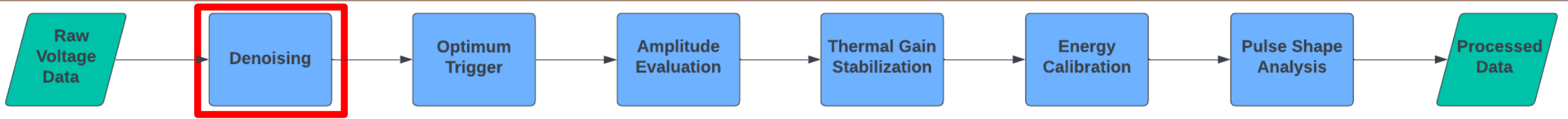


88% of $0\nu\beta\beta$ events occur in a single crystal

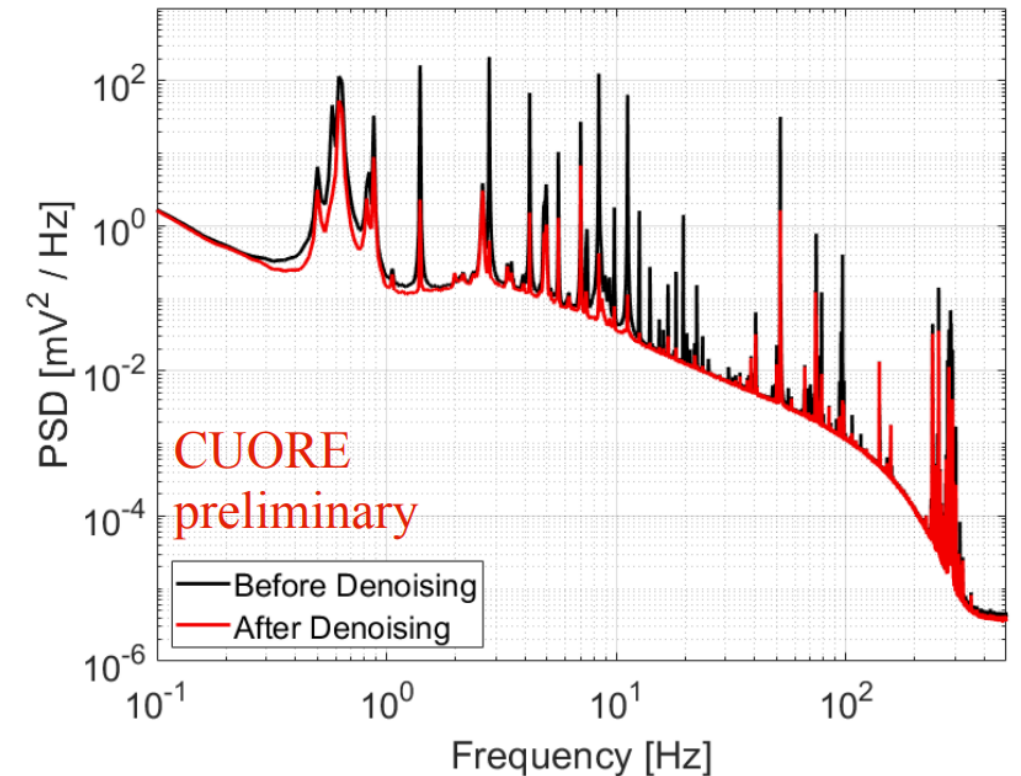
- Source = Detector
- ~88% containment efficiency
- ~93% data quality cut efficiency

<https://doi.org/10.48550/arXiv.2404.04453>

DATA PROCESSING

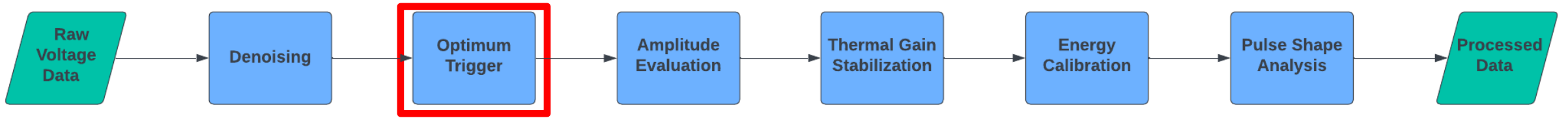


- Denoising
 - Multivariate noise cancelling algorithm to improve detector performance*
 - Uses data from auxiliary devices: Microphones, accelerometers, seismometers

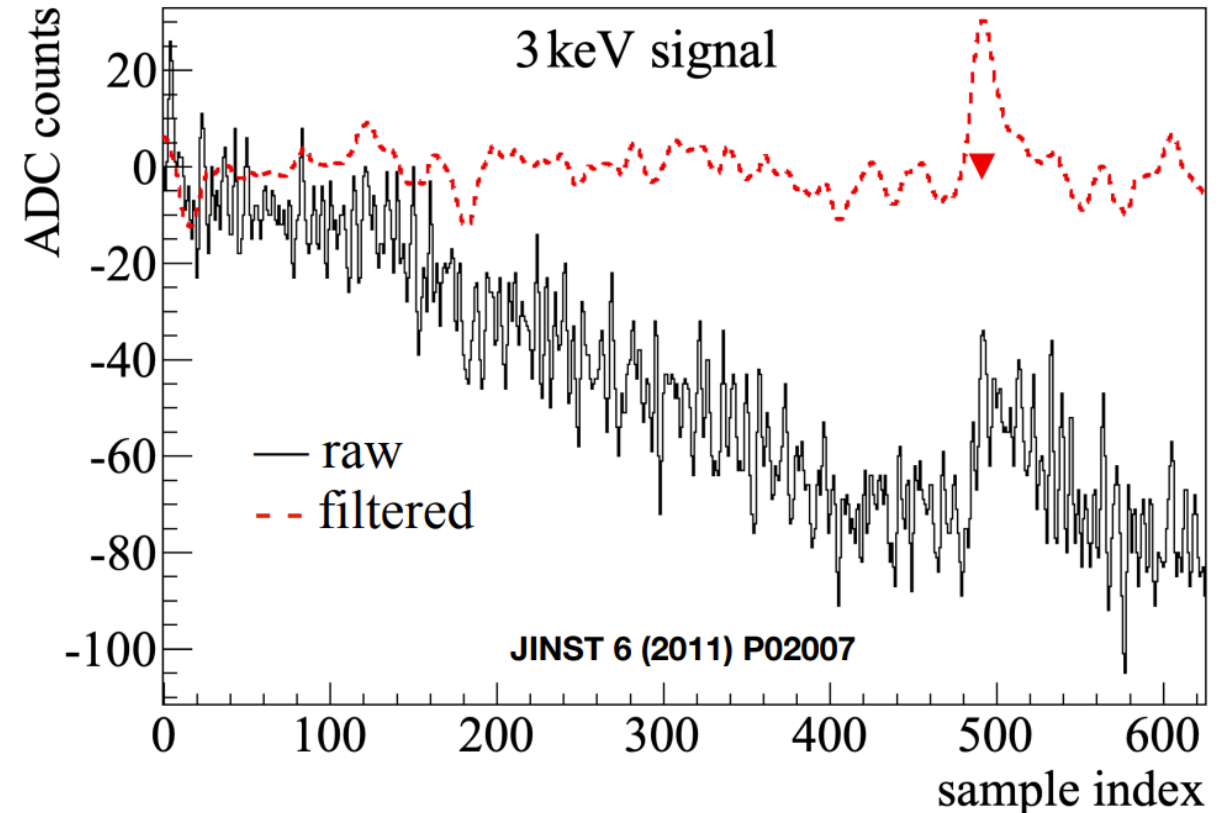


*<https://arxiv.org/pdf/2311.01131.pdf>

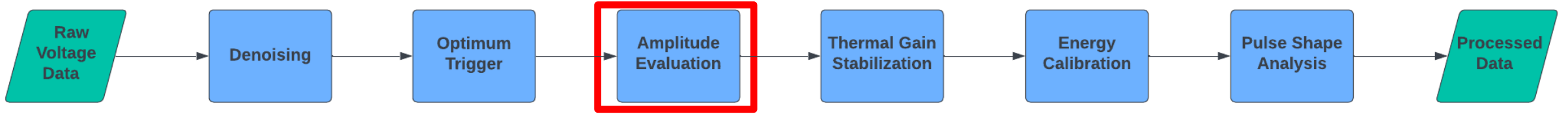
DATA PROCESSING - TRIGGER



- Optimum Trigger
 - Triggers on pulses after filtering waveform
 - Matched filter to maximize signal to noise
 - Matched filter to adjust for baseline drifts and noise
 - Lowers energy detection thresholds



DATA PROCESSING

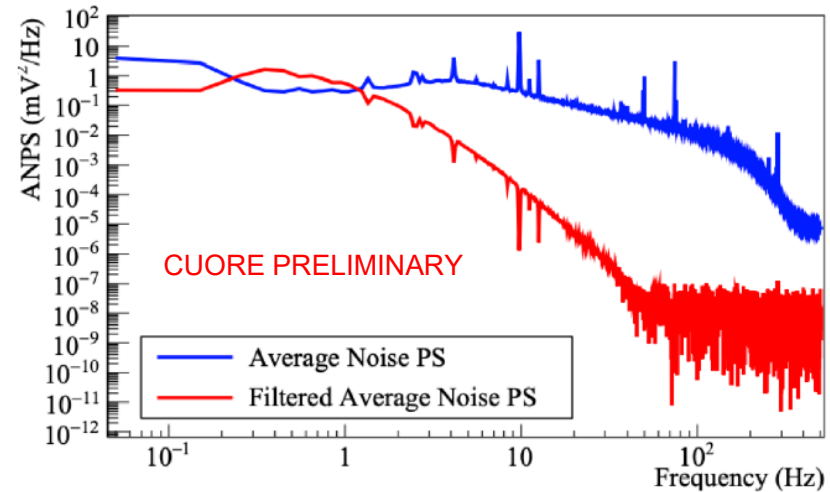
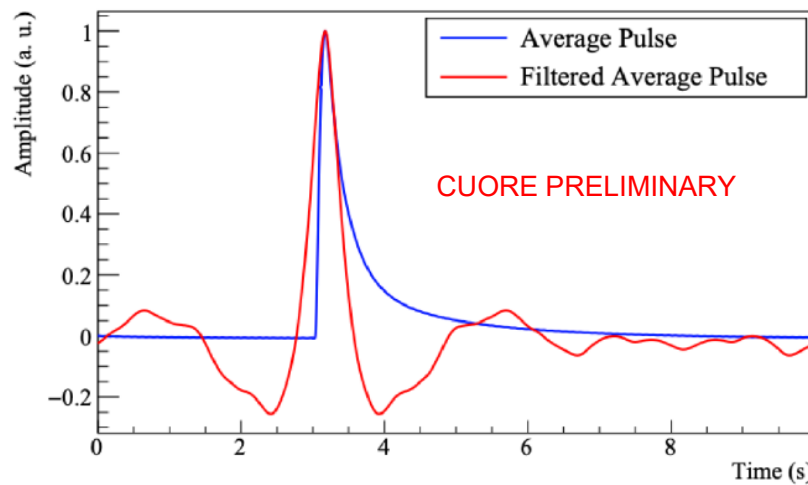


- Amplitude Evaluation

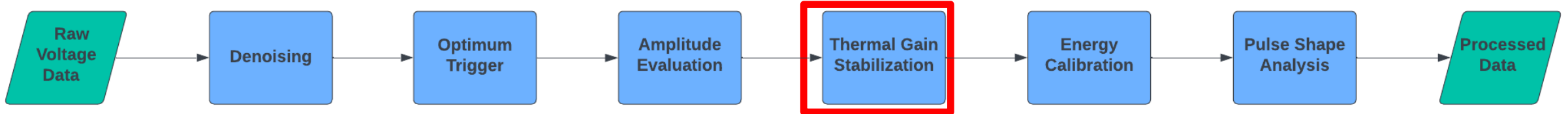
- Using optimum waveform filter to estimate amplitude of pulses
- Transfer function of each channel evaluated using average pulse and average noise power spectrum

$$H(\omega_k) \propto e^{i\omega_k t_{max}} \frac{s^*(\omega_k)}{N(\omega_k)} h(\omega_k)$$

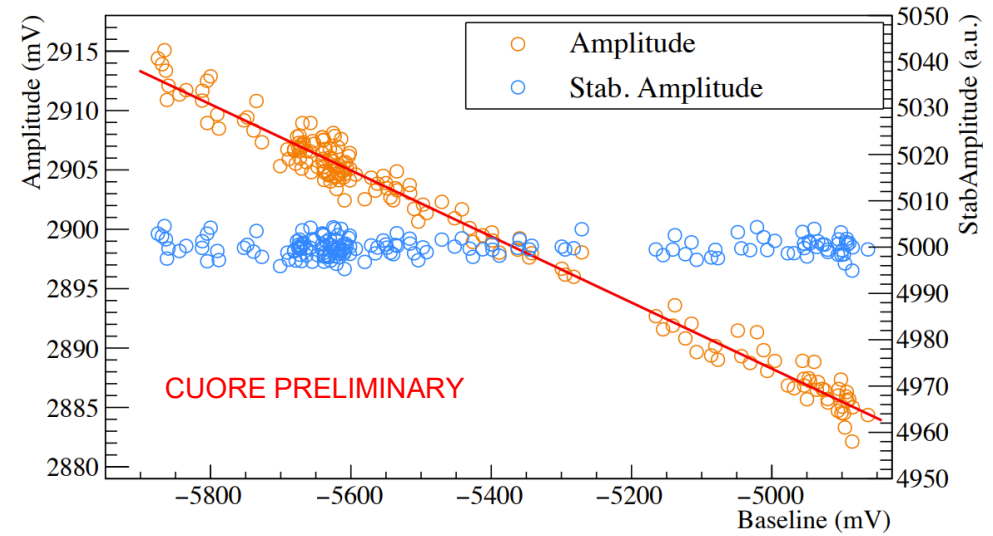
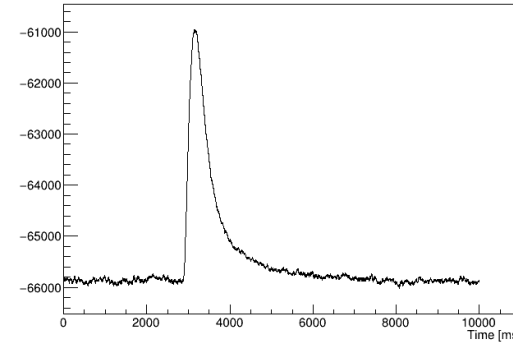
Average Pulse (pointing to $s^*(\omega_k)$)
Average Noise Power Spectrum (pointing to $N(\omega_k)$)



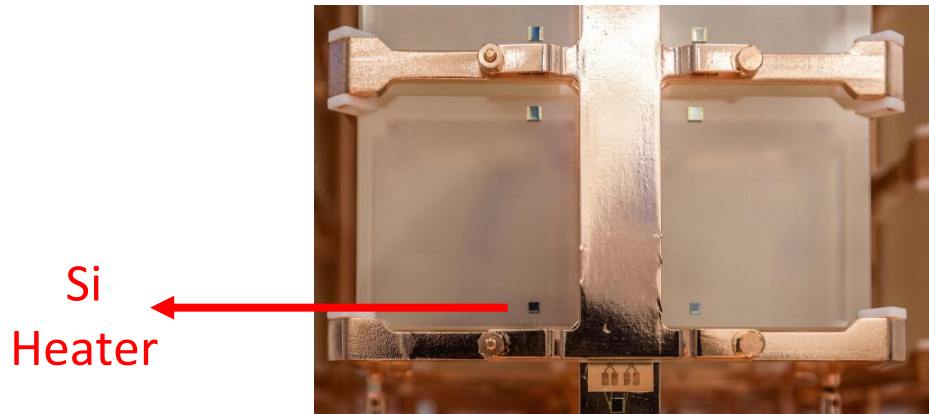
DATA PROCESSING



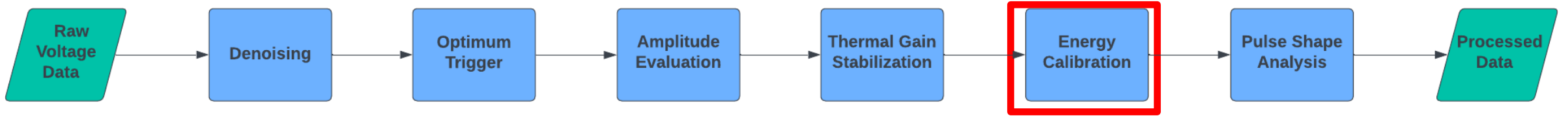
- Gain Stabilization
 - Minimizing gain dependence on temperature using periodically injected pulses



$$\text{Stabilized Amplitude} = 5000 \times \frac{\text{Amplitude}}{\text{StabPulser Amplitude}}$$

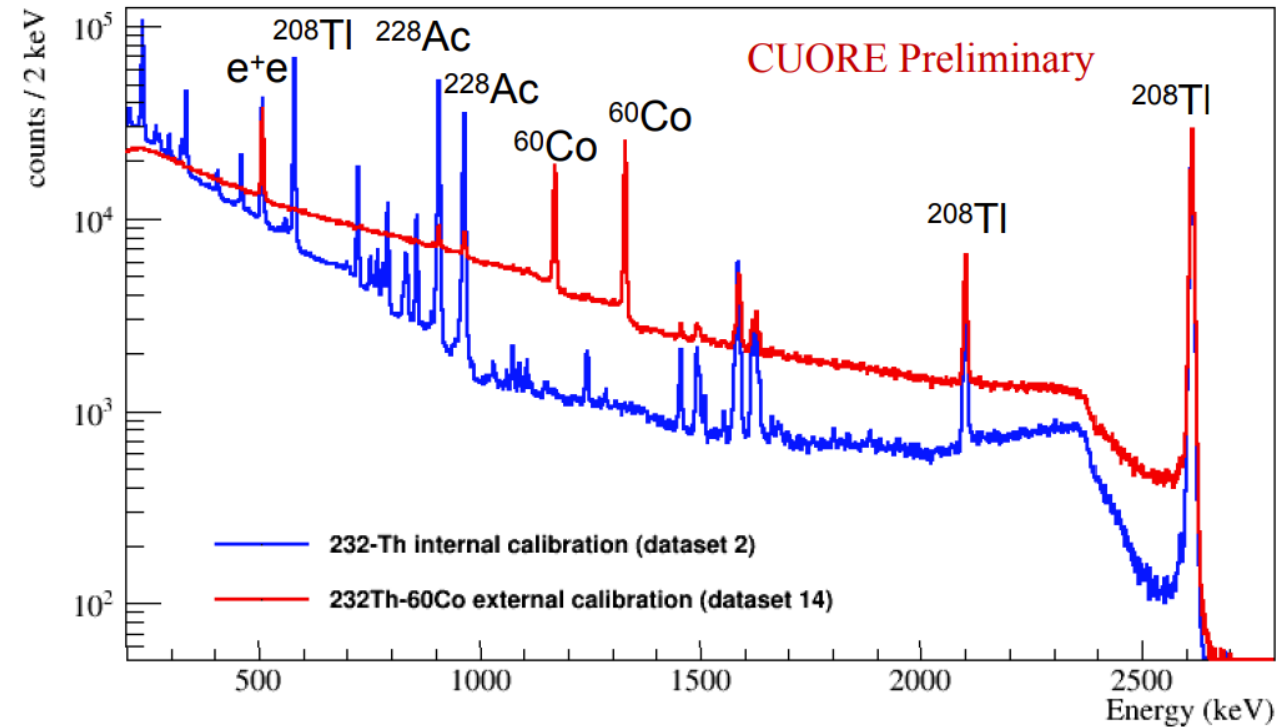


DATA PROCESSING

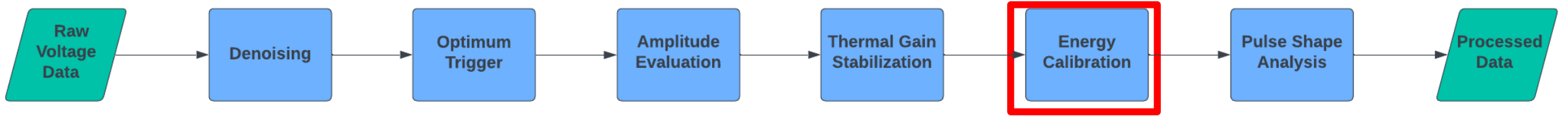


■ Calibration

- First 3 datasets used internal ^{232}Th source
- Later datasets calibrated with external ^{232}Th - ^{60}Co source
- Iterative calibration algorithm
 - Iterative seeding for crystal ball fits
 - Decision between gaussian and crystal ball fits based on various fit quality metrics
 - Tools to monitor fit quality

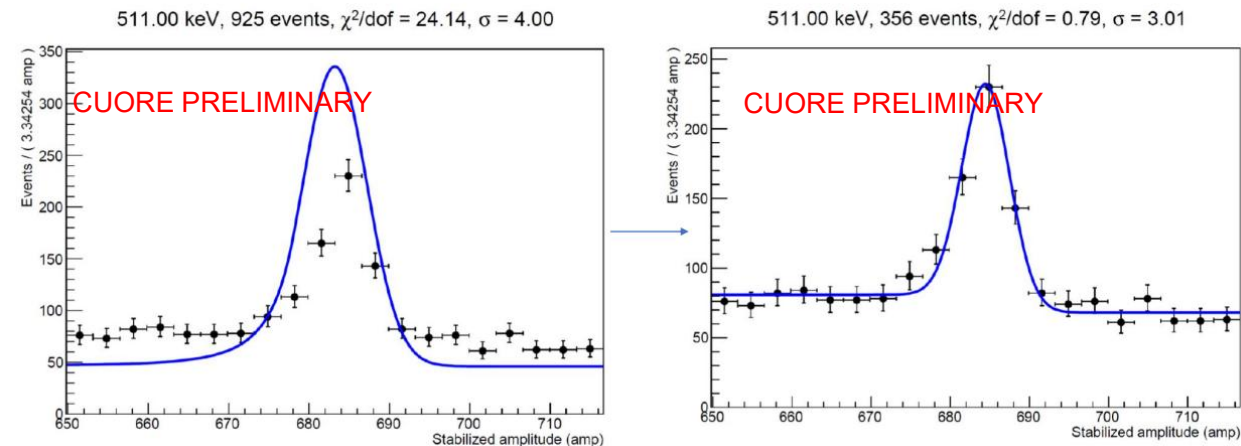


DATA PROCESSING



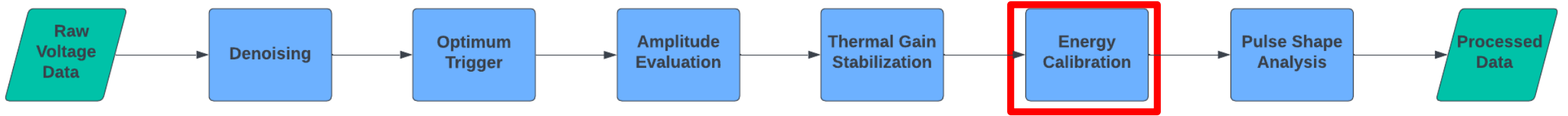
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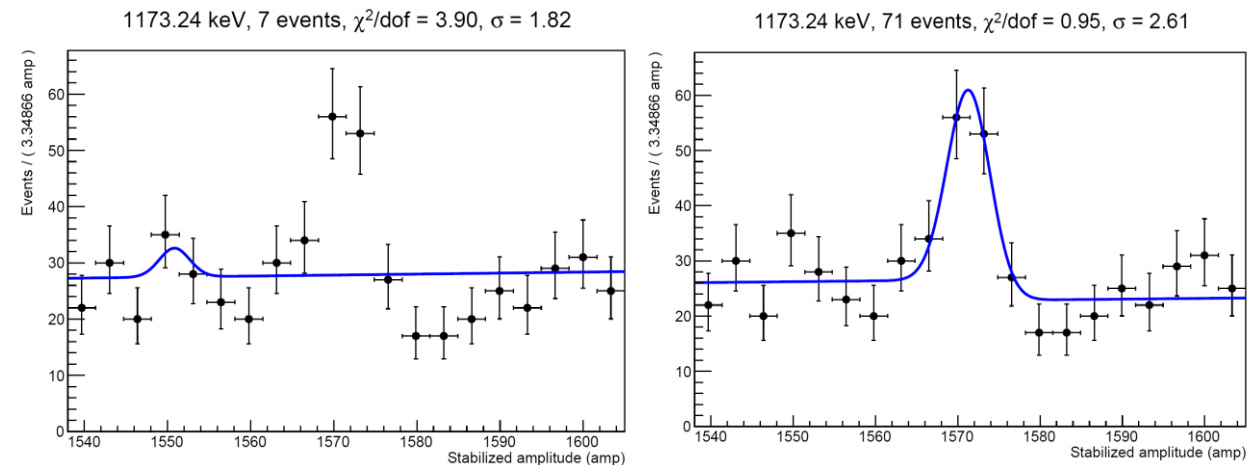
A calibration peak fit with default seeds (left), and fit with iterative algorithm (right)

DATA PROCESSING



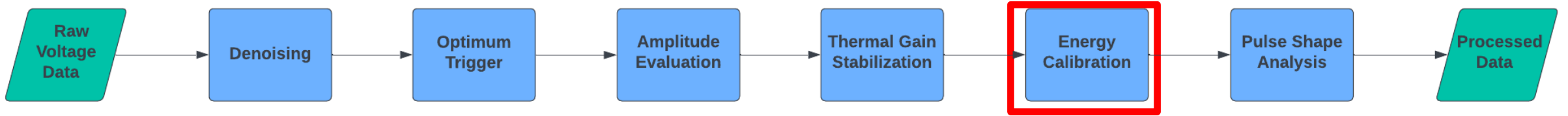
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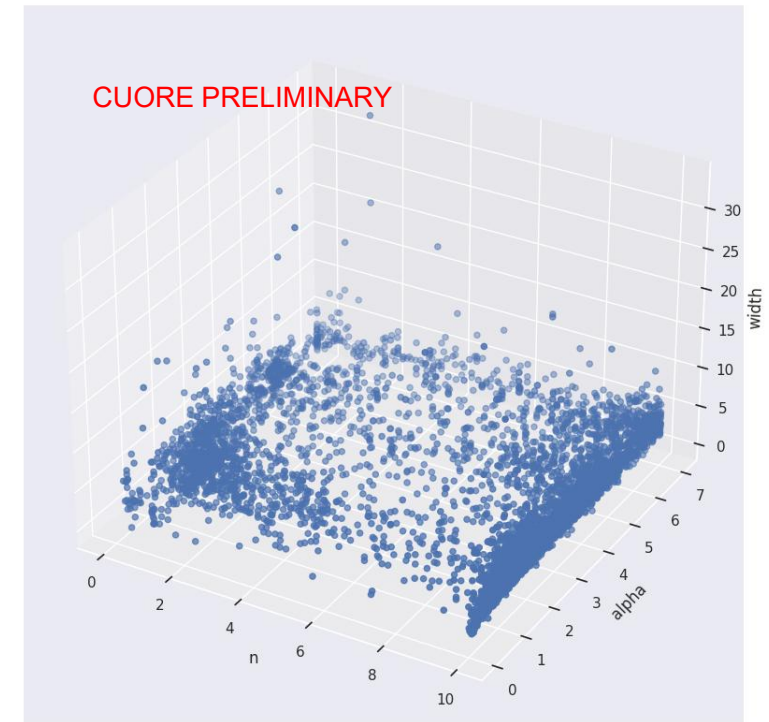
A calibration peak fit by crystal ball function (left), and gaussian (right); algorithm chose the gaussian as the better fit

DATA PROCESSING



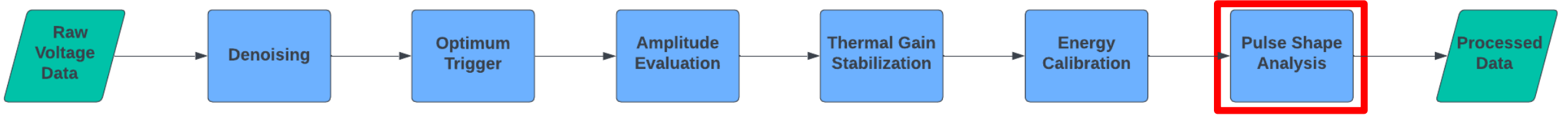
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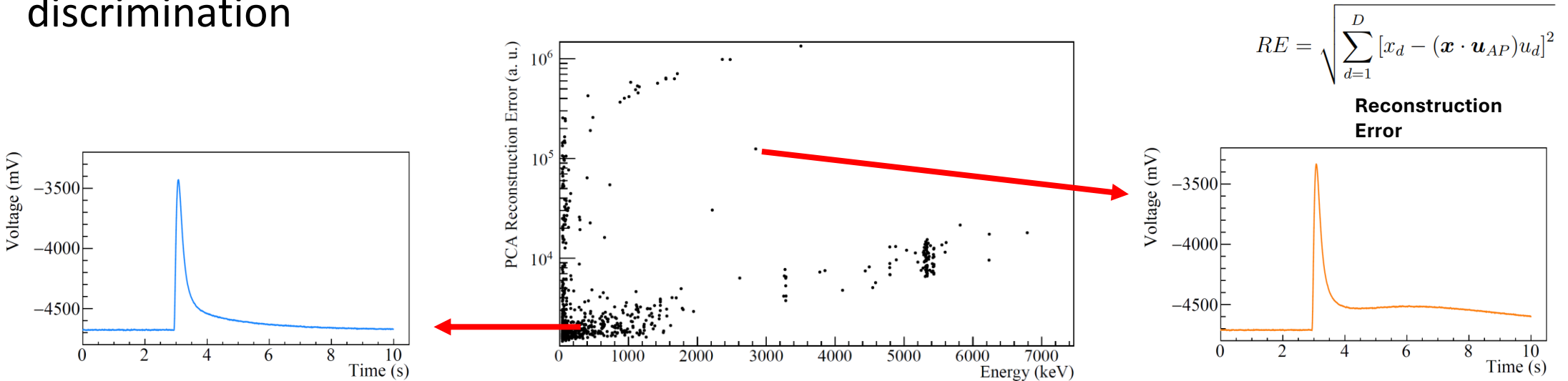
Distribution of fit parameters for a dataset

DATA PROCESSING



- Pulse Shape Discrimination

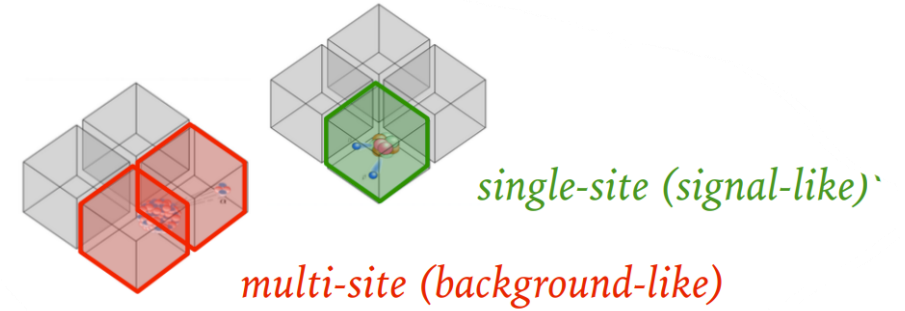
- Using PCA (Principal Component Analysis) to eliminate pulses with a non-physical shape
- Extract salient features of pulses to calculate reconstruction error for pulse shape discrimination



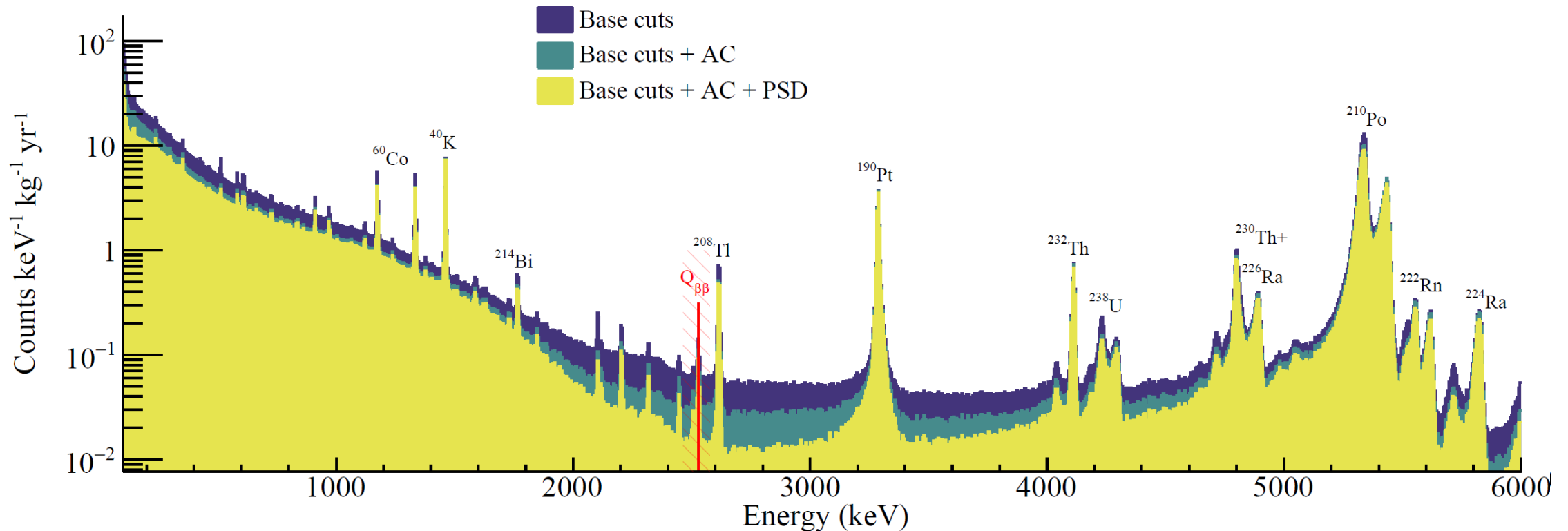
- Uncertainty in the total analysis efficiency. (93.4(2)%)
- Uncertainty in the PCA efficiency. (98.1(2)%)
- Uncertainty in the containment efficiency (88.35 ± 0.090)%
- Uncertainty in the $Q\beta\beta$ (2575.5 ± 0.013) keV
- Tellurium natural isotopic abundance (34.167 ± 0.002)%
- Uncertainties in the bias and resolution scaling

2 TONNE-YR SPECTRUM

- Blinded analysis to prevent bias
 - ROI (2465, 2575) keV salted with ^{208}Tl 2615 keV peak
 - Fit ROI events with: ^{130}Te $0\nu\beta\beta$ peak, ^{60}Co peak, linear background
 - Finalize model parameters before unblinding

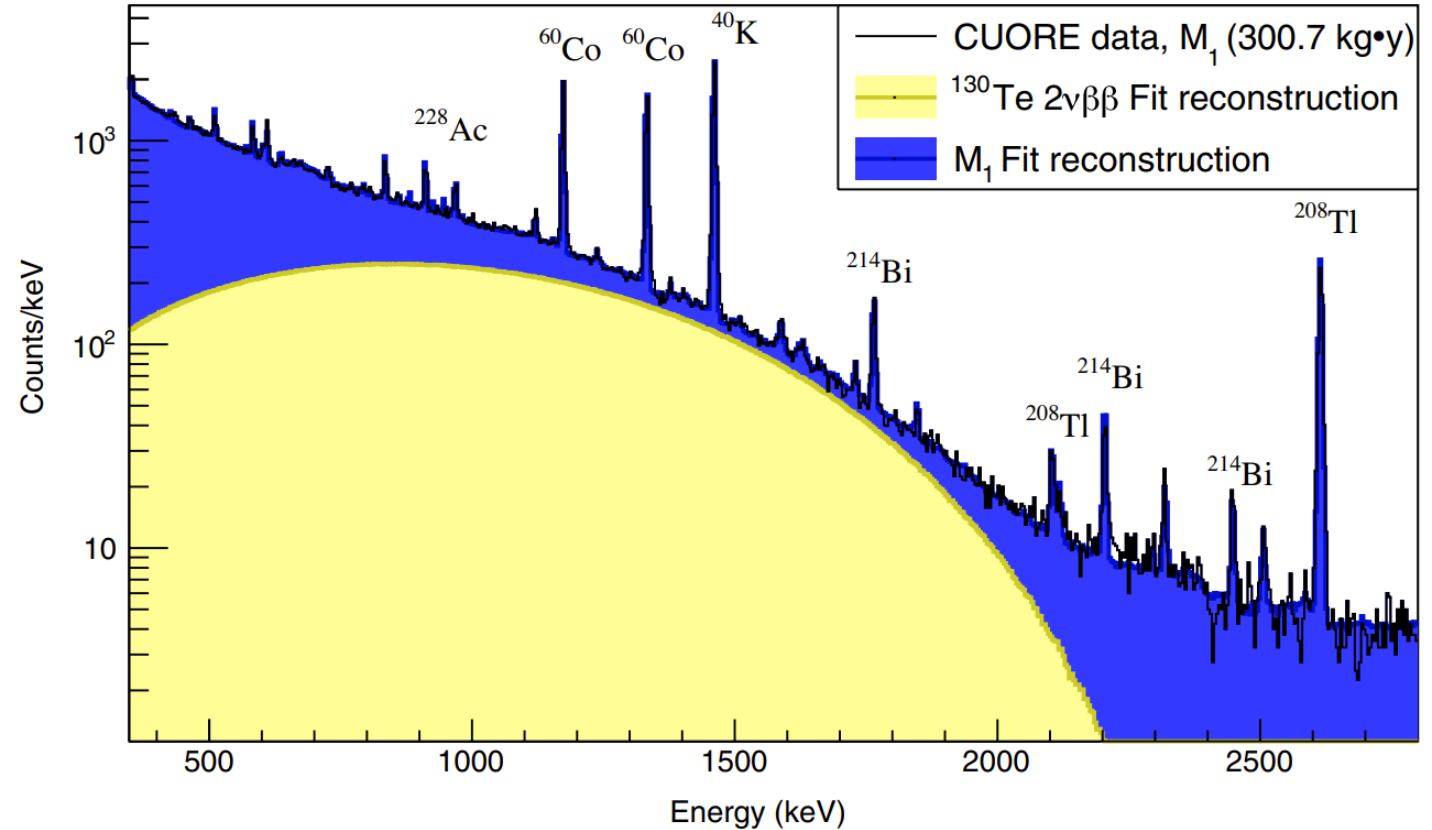


88% of $0\nu\beta\beta$ events occur in a single crystal



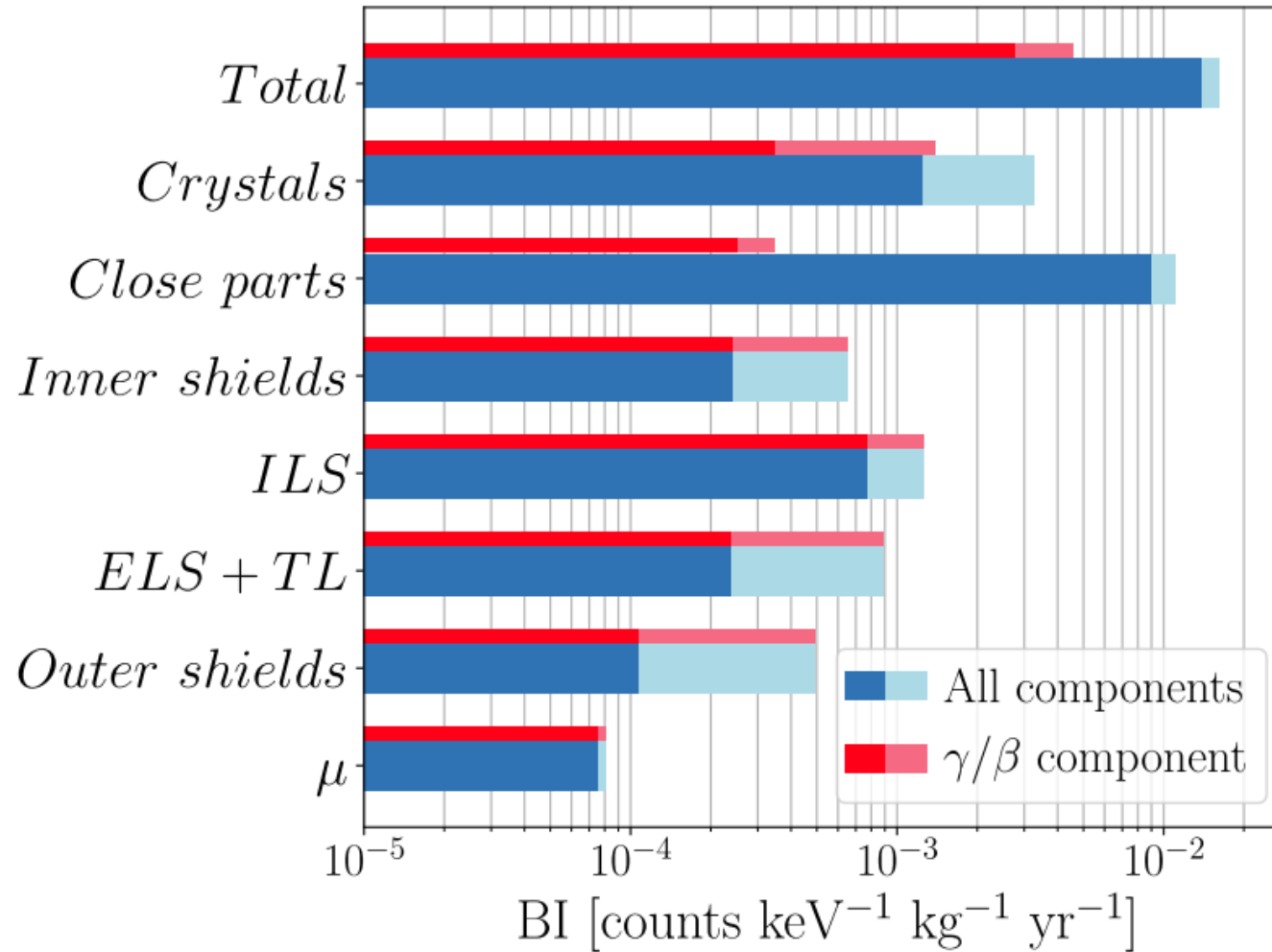
DOUBLE BETA DECAY RESULTS

- Double beta decay simulated in Geant4 with CUORE geometry and detector response
- Spectrum reconstructed by simultaneous fit of data with 62 MC simulated sources ($2\nu\beta\beta$ + surface and bulk contaminations + muons)
 - MCMC Bayesian approach
 - Uniform prior for sources except muons
- For 900-2000 keV, more than 50% counts are $2\nu\beta\beta$ events



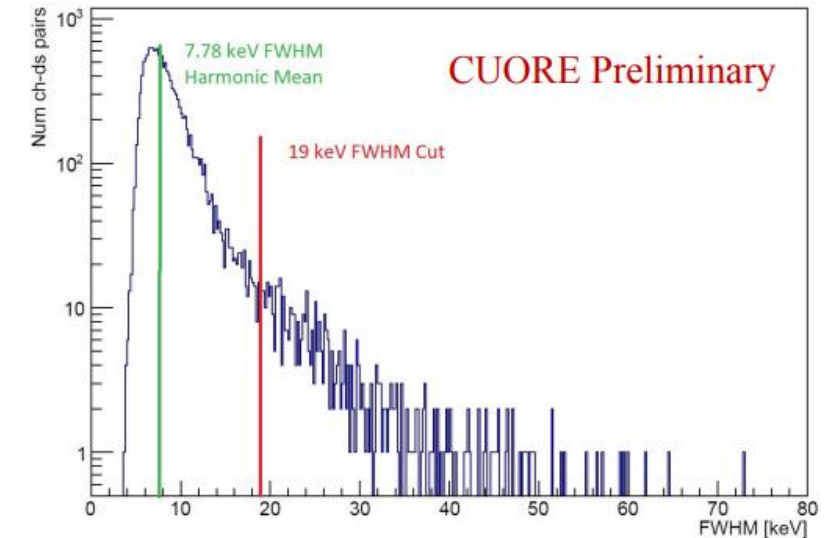
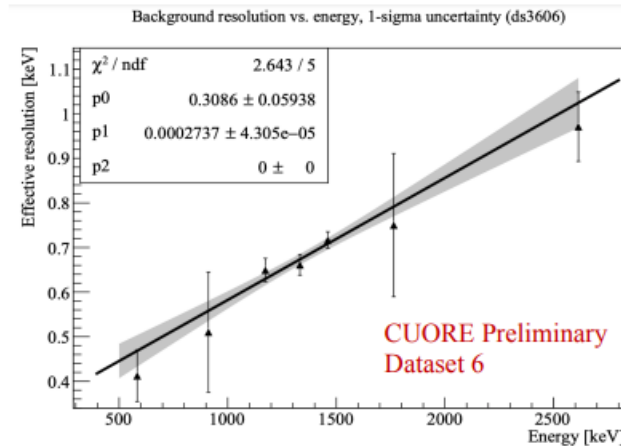
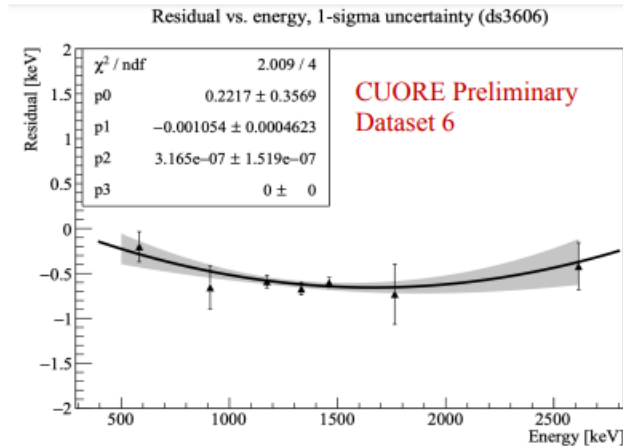
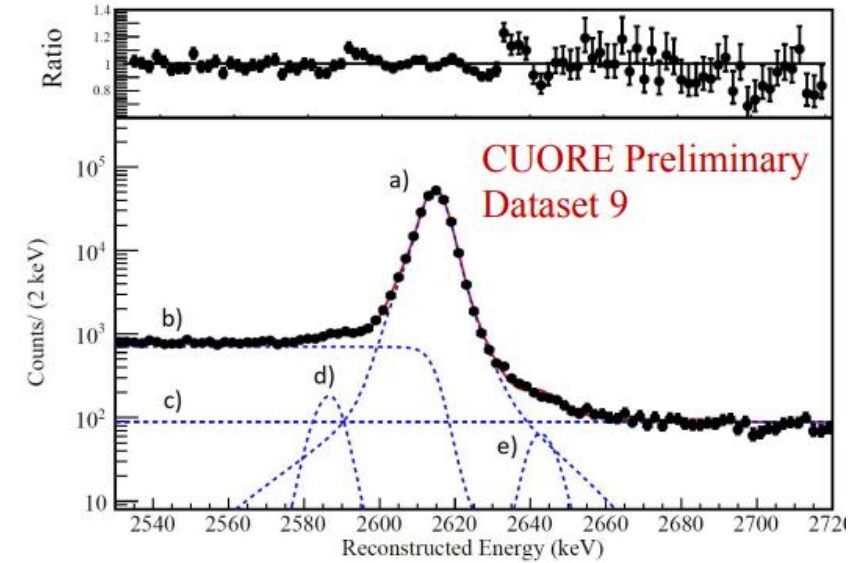
$$T_{1/2}^{2\nu} = 7.71_{-0.06}^{+0.08}(\text{stat.})_{-0.15}^{+0.12}(\text{syst.}) \times 10^{20} \text{ yr}$$

BACKGROUND BUDGET



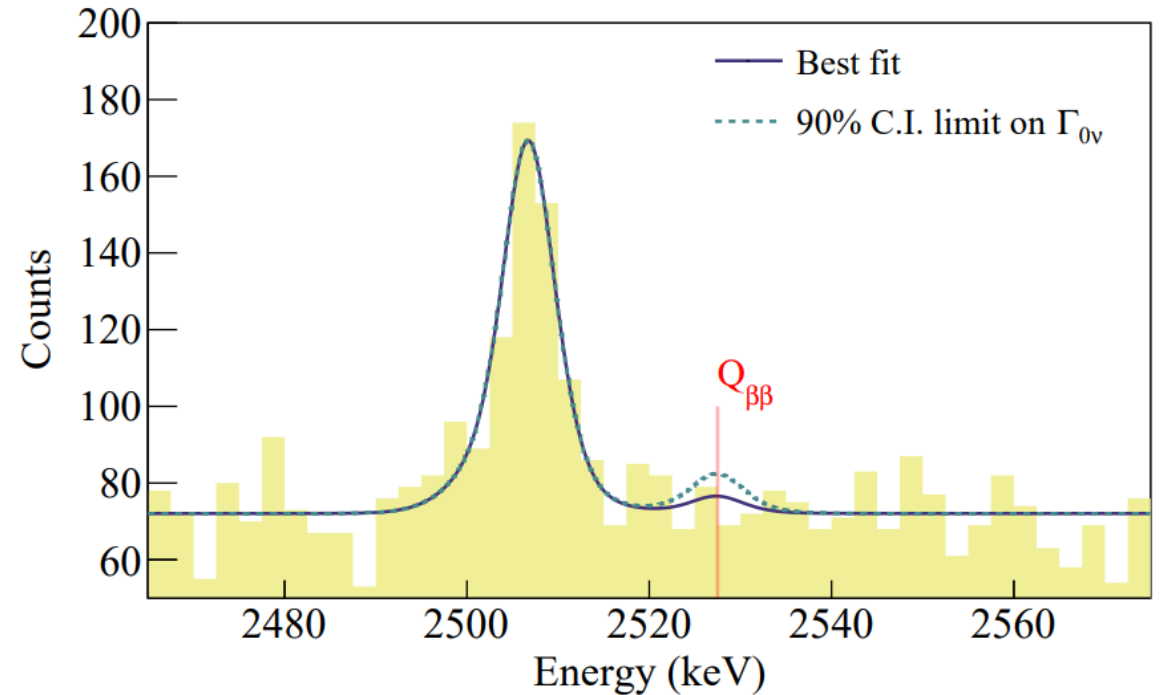
DETECTOR RESPONSE

- Fit 2615 keV calibration peak for each channel
 - 3-Gaussian signal peak
 - Compton background
 - Flat background
 - 30 keV X-ray escape peak
 - 30 keV X-ray sum peak
- Scale detector response from 2615 keV calibration fit to peaks in physics data



BACKGROUND IN ROI

- Alpha region:
 - Flat background in [2650, 3100] keV
 - $1.40(2) \times 10^{-2}$ counts/(keV kg yr)*
- $Q_{\beta\beta}$ region
 - Flat background + ^{60}Co peak in [2490, 2575] keV
 - $1.49(4) \times 10^{-2}$ counts/(keV kg yr)*
- Background dominated by degraded alpha energy depositions (90%)



*CUORE collaboration
<https://www.nature.com/articles/s41586-022-04497-4.pdf>