

Gravitational Waves from Early-Universe Turbulent Sources

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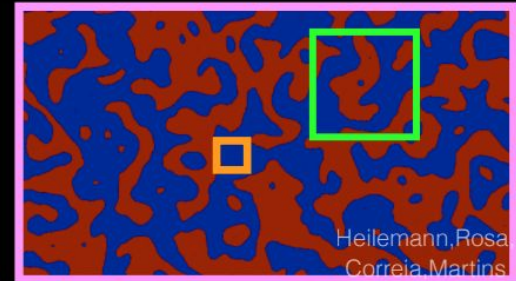
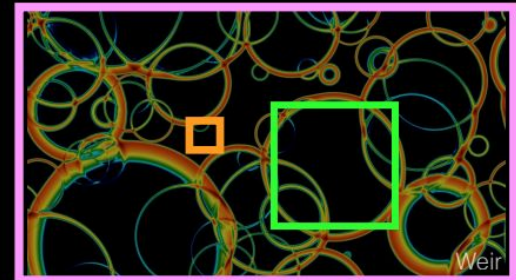
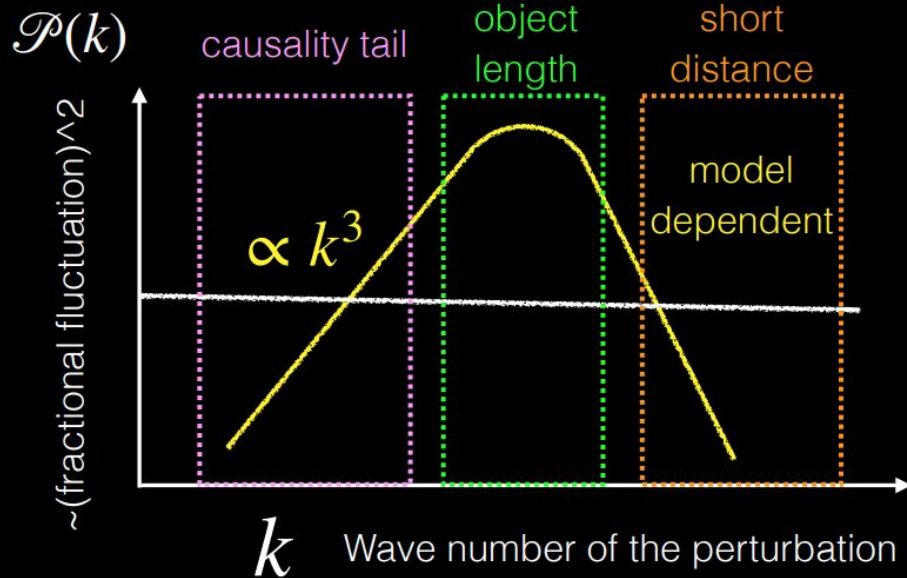


Introduction

- **1980s:** Witten (1984), Hogan (1986) – FOPTs produce GWs.
- **1992–1994:** Kosowsky, Turner & Watkins; Kamionkowski, Kosowsky & Turner – Envelope approximation, analytical parametrization framework.
- **2002:** Kosowsky, Mack & Kahniashvili – Turbulence as a GW source.
- **2007–2009:** Gogoberidze, Kahniashvili & Kosowsky; Caprini, Durrer & Servant – MHD refinements, analytic templates.
- **2010s–today:** Hindmarsh et al.; Roper Pol et al.; Auclair et al. – Numerical simulations.

Slide from Yuhsin Tsai, this Morning's Presentation

The length scale of the **peak** corresponds to the **correlation length** of the objects; **height** $\sim O(0.1)$ fluctuation scaled by energy density ratio



From Cosmological Perturbation Theory to GWs

$$h''_A + 2\mathcal{H}h'_A + k^2 h_A = 16\pi G a^2 \sigma_A$$

- **LHS:** Wave equation for first-order TT tensor perturbations with Hubble friction in conformal variables.
- **RHS:** TT anisotropic stress of relevant process which stirs cosmic fluids.

Parameters of the Model

- α : Strength of the phase transition.
- β : Inverse duration of the phase transition.
- T_* : Temperature of the phase transition.
- κ : Efficiency of energy transfer.
- v_w : Wall velocity of the bubbles.

With simple dimensional analysis one can generate a rough estimate of the total energy density of the GW radiation and the relevant frequency scales for the signal.

[Caprini & Figueroa (2018)][Kosowsky et. al 1992]

$$\beta^2 h \propto 16\pi G a^2 \sigma$$

$$h \propto 16 \frac{\pi G a^2 \sigma}{\beta}$$

$$\rho_{GW} \propto \frac{1}{32\pi G} \langle \dot{h} \dot{h} \rangle$$

$$f \sim 1.6 \times 10^{-5} [\text{Hz}] \left(\frac{\beta}{H_*} \right) \left(\frac{g_*}{100} \right)^{1/6} \left(\frac{T_*}{100 [\text{GeV}]} \right)$$

$$h^2 \Omega_{GW} = 1.6 \times 10^{-5} \left(\frac{100}{g_*} \right)^{1/3} \left(\frac{H_*}{\beta} \right)^2 \left(\frac{\kappa \alpha}{1 + \alpha} \right)^2$$

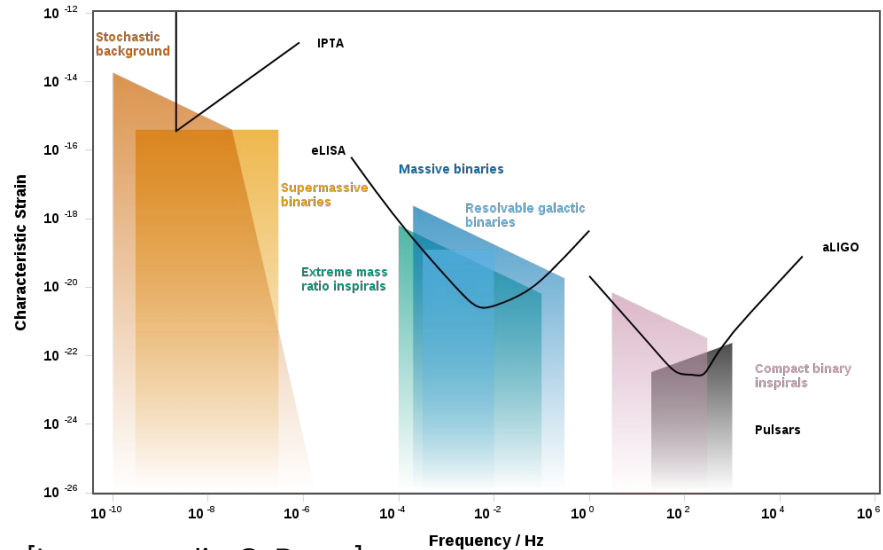
Parameters of the Model

With these equations one can determine the relevance for different physical processes.

- EWPT ~ 100 GeV
- QCDPT ~ 150 MeV

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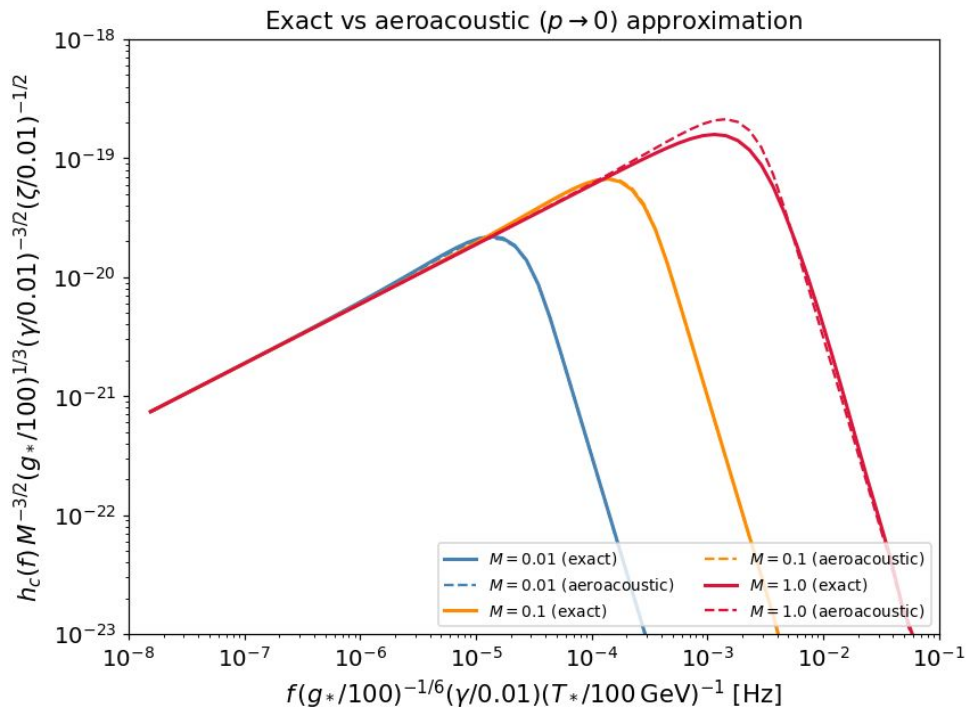


[Image credit: C. Berry]

Gravitational Waves from MHD

Case I: Stationary turbulence

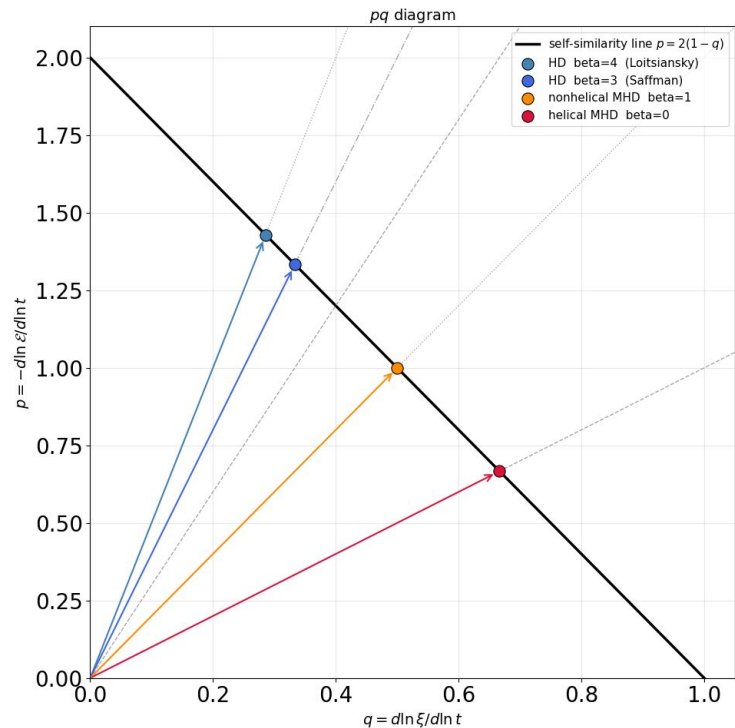
The modeling assumes fully developed stationary **Kolmogorov turbulence** and with **Kraichnan decoration** functions. [Gogoberidze et. al. 2007]



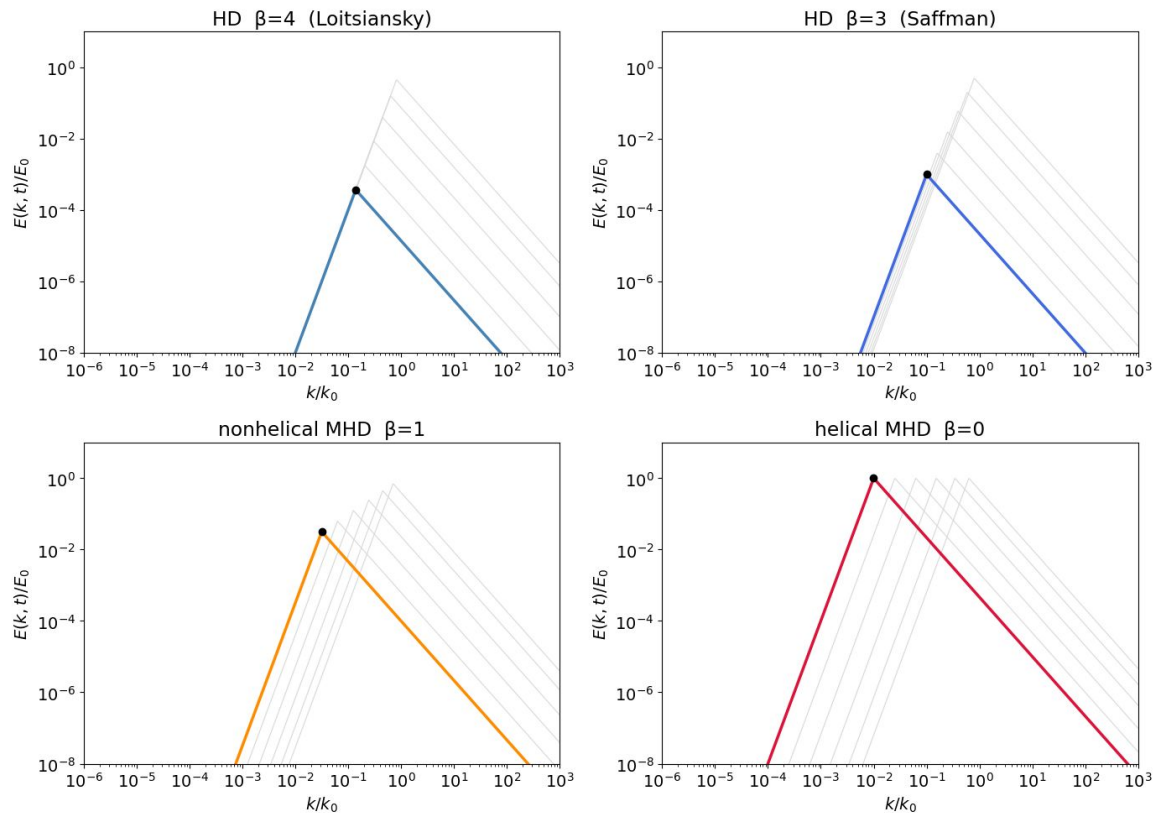
Gravitational Waves from MHD

Case II: Decaying turbulence

To introduce decaying we impose analytical ansatz of power law decay of both energy and coherence length scale. [Brandenburg & Kahnishvili 2016]



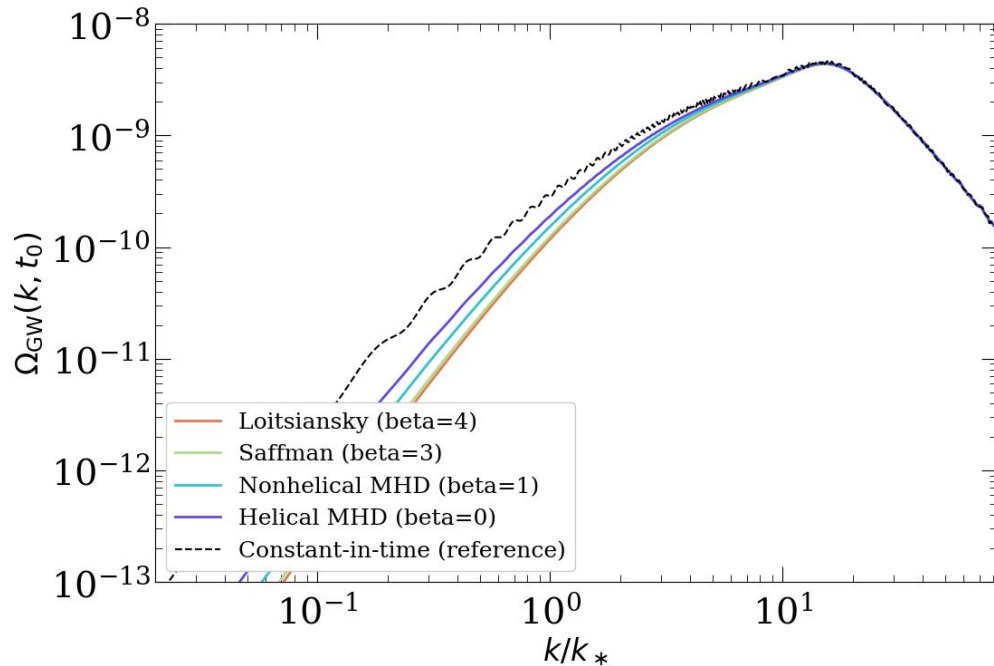
Gravitational Waves from MHD



Gravitational Waves from MHD

Case II: Decaying turbulence

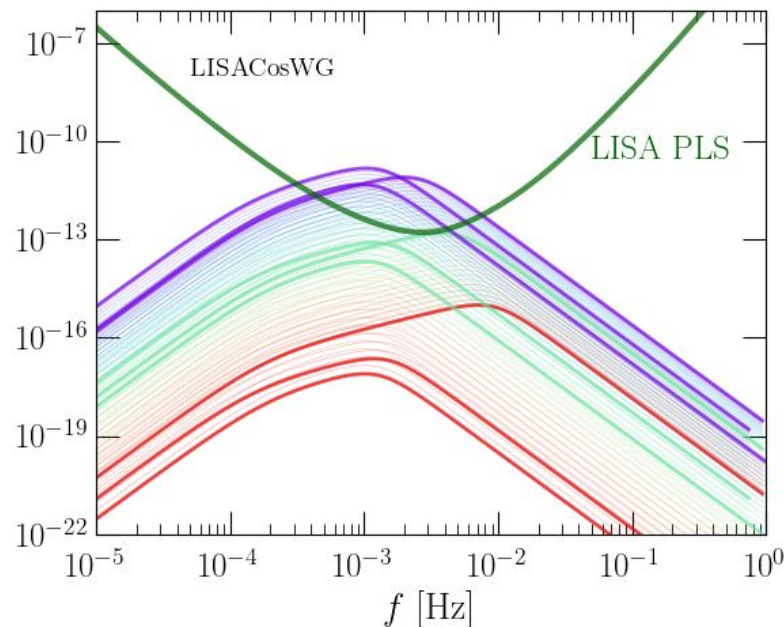
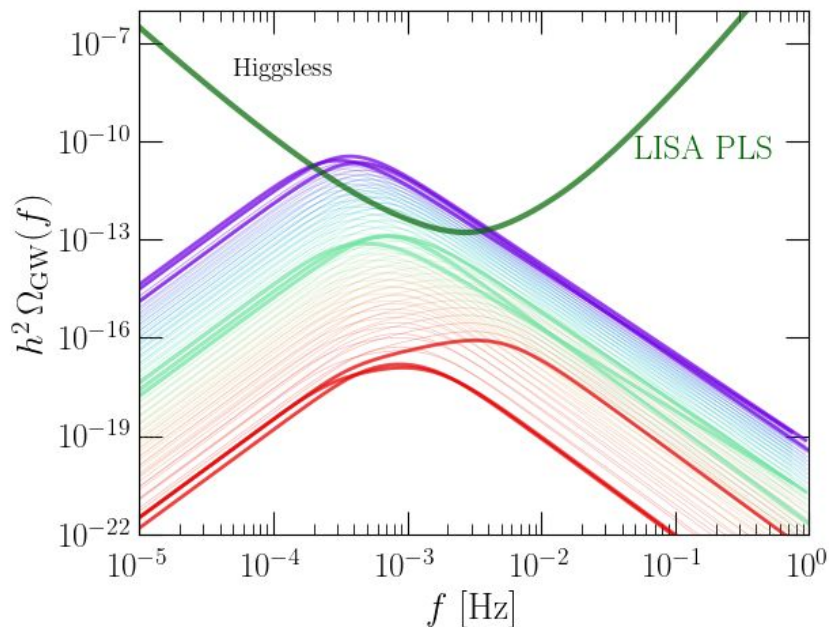
To introduce decaying we impose analytical ansatz of power law decay of both energy and coherence length scale. [A. Roper Pol et. al 2025]



Gravitational Waves from Numerical Simulations

CosmosGW

(Soon on Github: <https://github.com/CosmoGW>)



Conclusions

Two routes, both incomplete

- analytical: transparent and cheap, but closure dependent
- numerical: full physics, but expensive and code dependent

CosmoGW, paper this month, release in two weeks

- one hub for Pencil Code, Cosmos Lattice, Higgsless
- common analytical pipeline, consistent redshifting
- simulations and models on the same footing

Outlook

- feed simulated spectra into the analytical pipeline, test where closures break

Acknowledgements

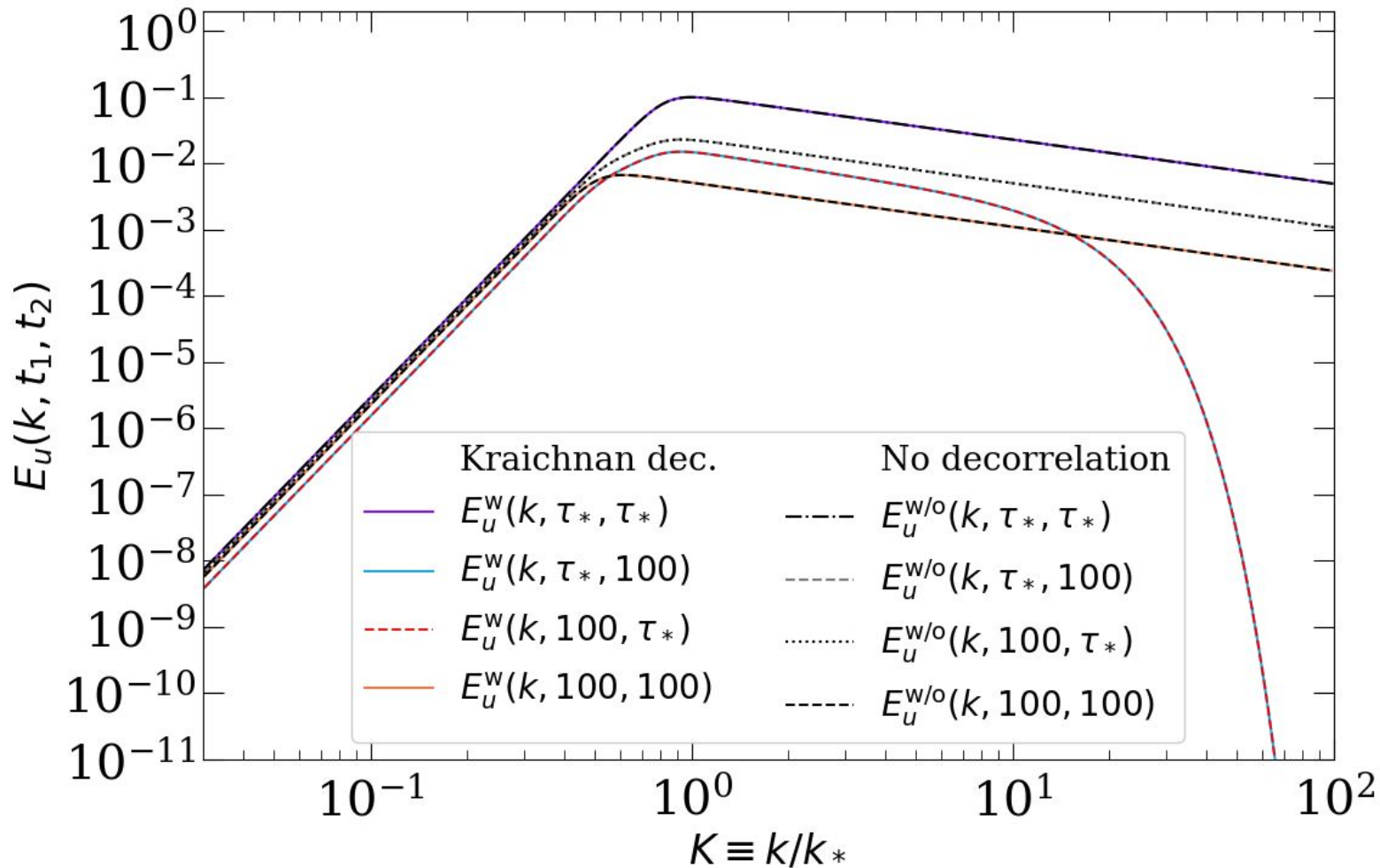
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Backup

