

Secluded effects on scalar light dark matter

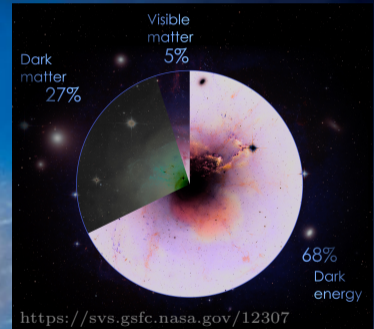
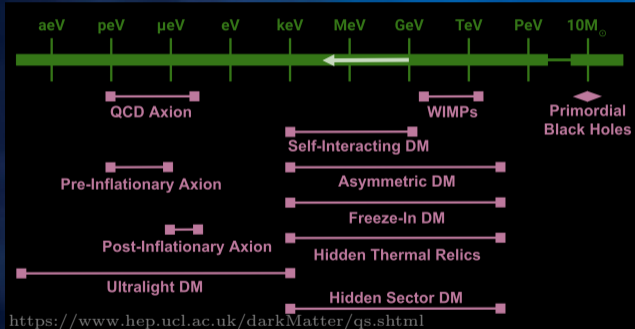
Lilianna Hariasz

Simon Fraser University

May 11, 2026

Phenomenology 2026 Symposium, May 11-13, 2026, University of Pittsburgh

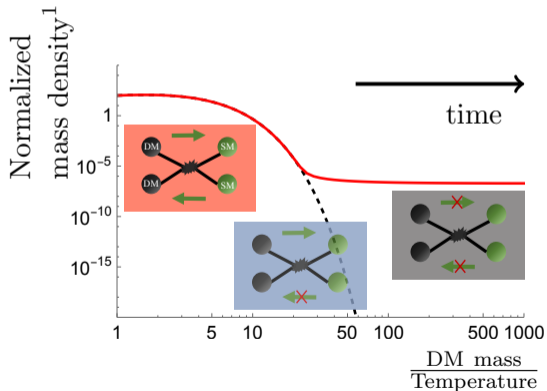
Light Dark Matter (LDM)



Extending the Standard WIMP Scenario

Weakly Interacting Massive Particle

- Thermal relic, one new fermion

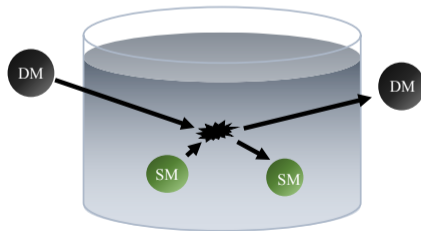
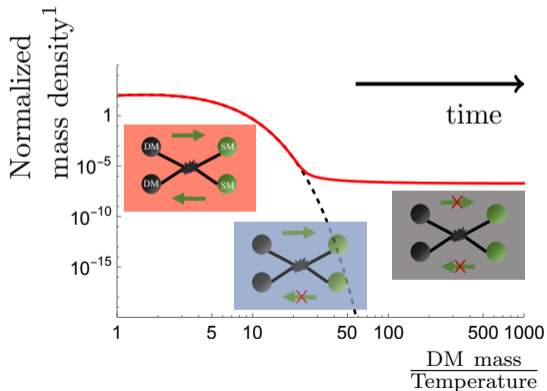


¹Normalized to equilibrium number density when mass = temperature; 100 GeV WIMP; SM = Standard Model.

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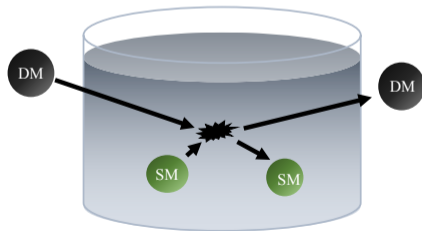
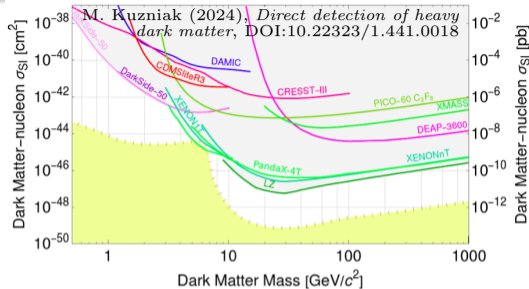
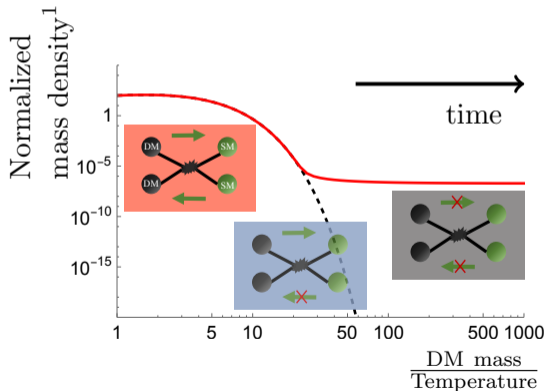


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Extending the Standard WIMP Scenario

Weakly Interacting Massive Particle

- Thermal relic, one new fermion
- Extensively studied



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A dark sector and light dark matter

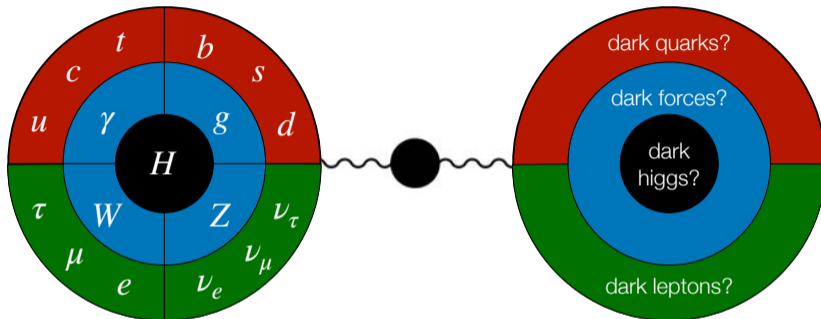
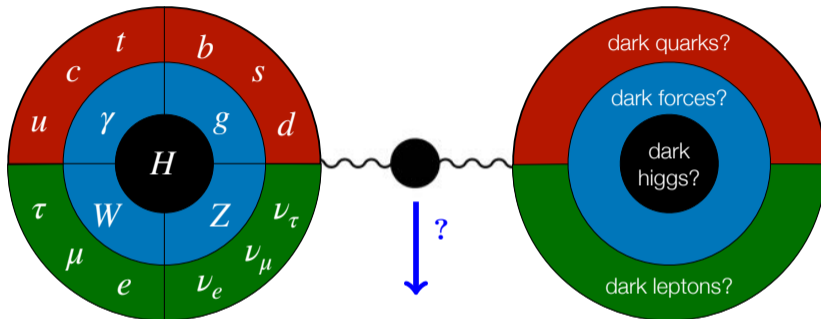


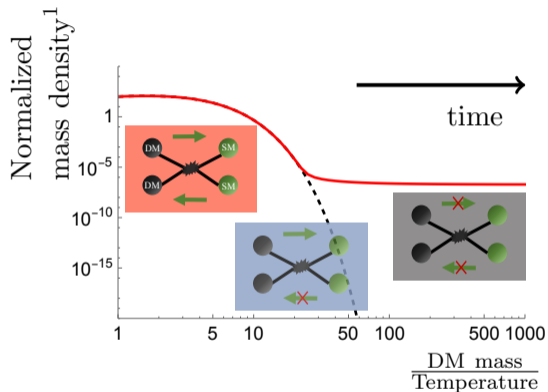
Image adapted from Gori & Williams (2022), *Dark Sector Physics at High Intensity*, arXiv:2209.04671.

A dark sector and light dark matter



New light mediator (“dark photon”)

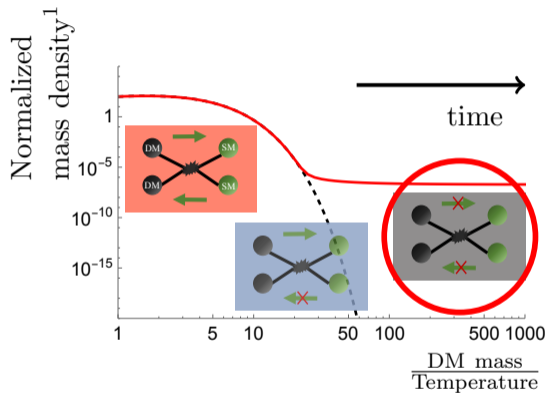
- Maintain freeze-out: matches observed relic abundance in sub-GeV range
- DM abundance provides experimentally accessible scenarios



- DM does not completely stop annihilating at freeze-out
- sub-GeV Dirac DM \implies re-ionization of hydrogen at re-combination

Some viable alternatives

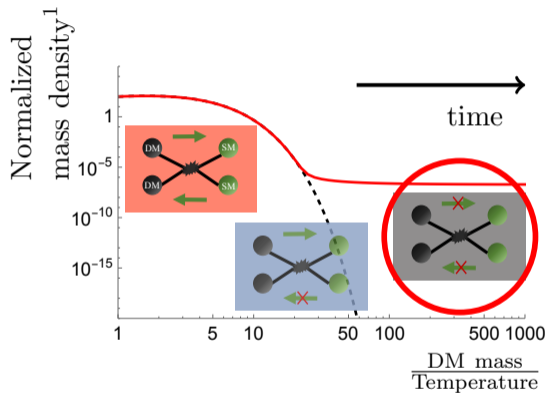
- Late-time asymmetry
- Velocity suppression:
 $\text{Complex scalar } \langle \sigma_{\text{ann}} v \rangle \propto v_{\text{rel}}^2$



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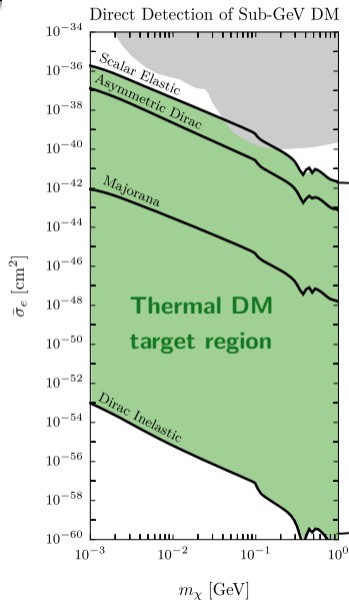


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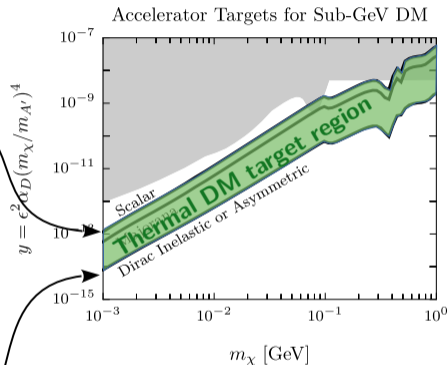
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LDM thermal relics: direct detection & accelerator targets



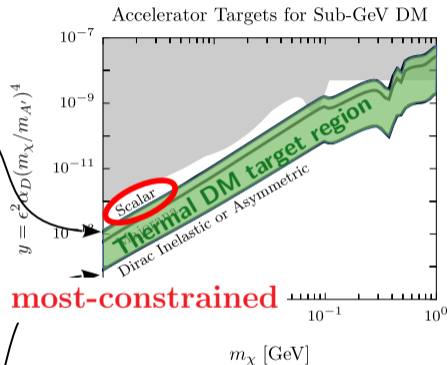
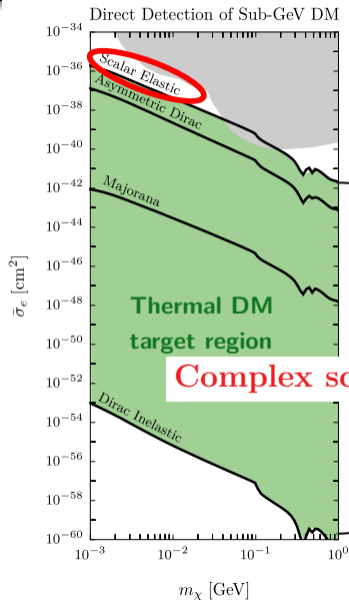
Opportunities identified at Snowmass 2021



Krnjaic et al. (2022), *A snowmass whitepaper: Dark matter production at intensity-frontier experiments.*

LDM thermal relics: direct detection & accelerator targets

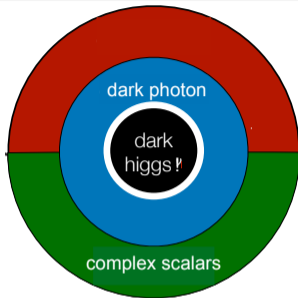
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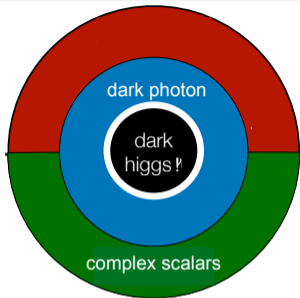
Complex scalars most-constrained

Krnjaic et al. (2022), *A snowmass whitepaper: Dark matter production at intensity-frontier experiments.*

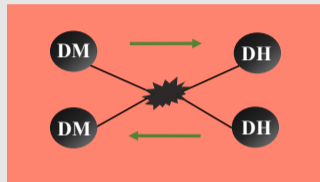
Origin of the dark photon's mass



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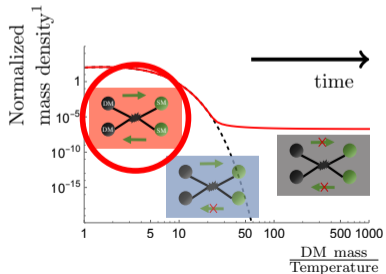
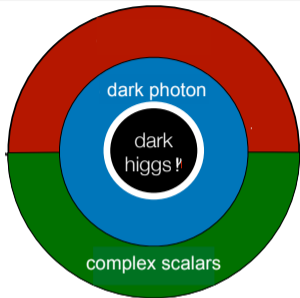


Additional dark sector interactions

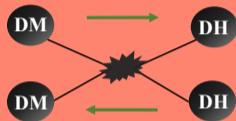


+decays, when applicable

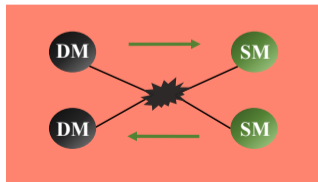
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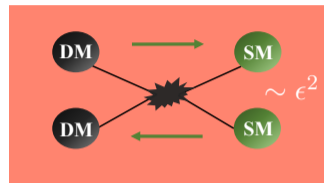
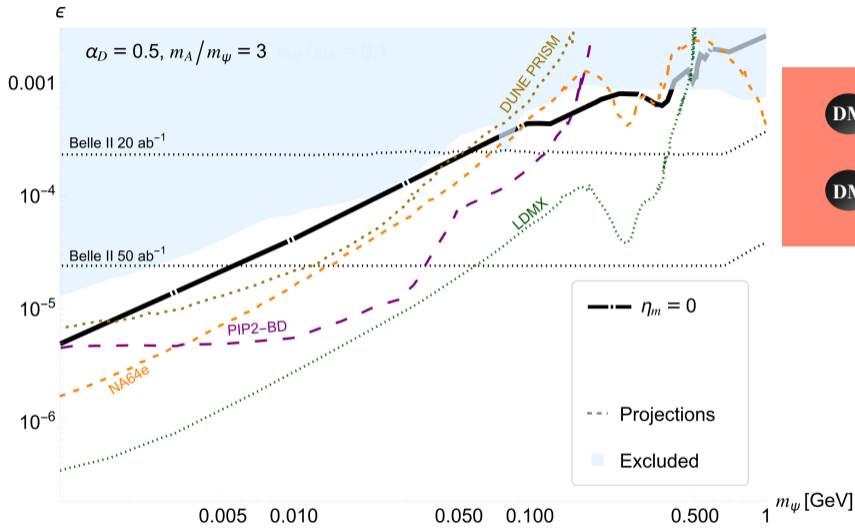
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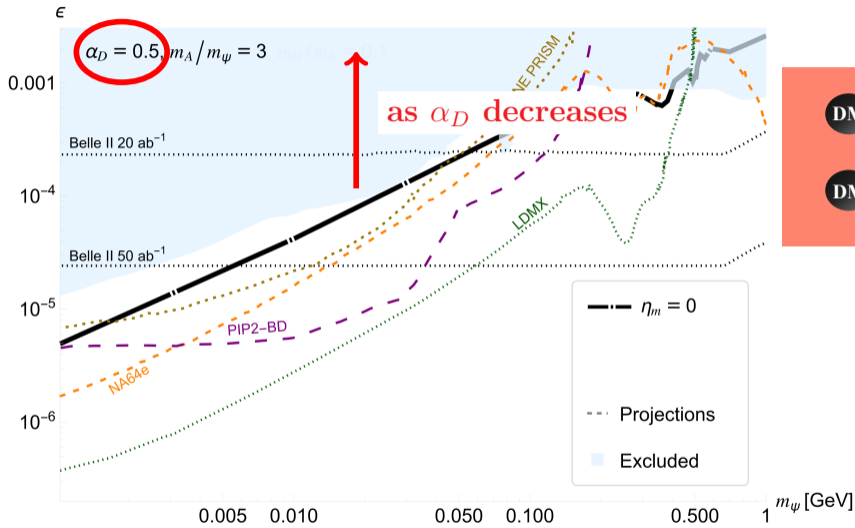


Scalar LDM & a light dark Higgs



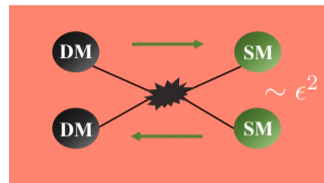
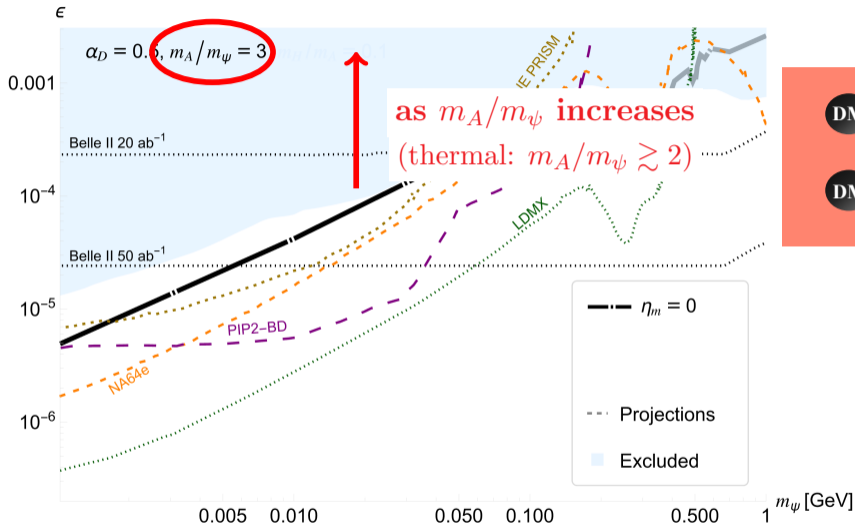
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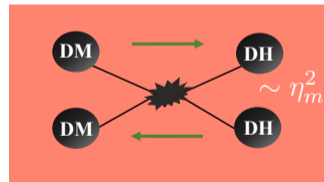
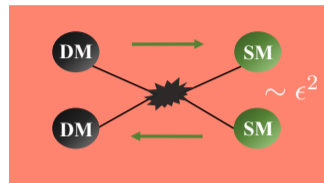
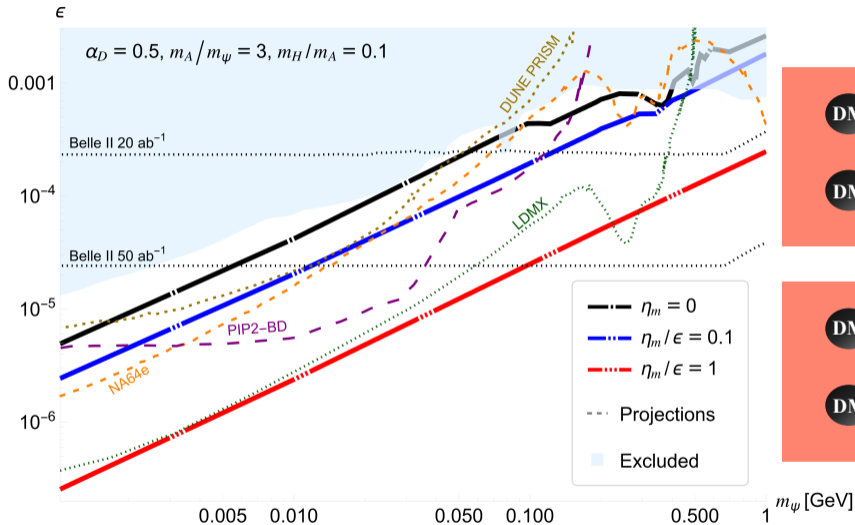
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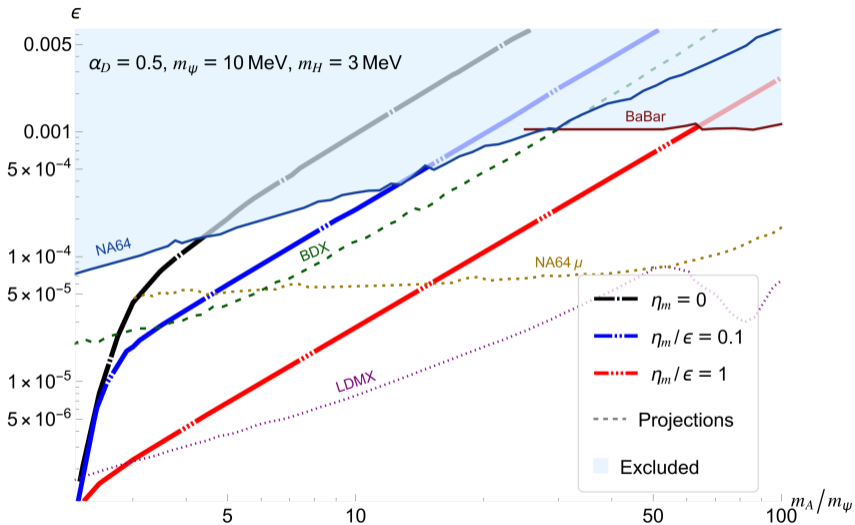
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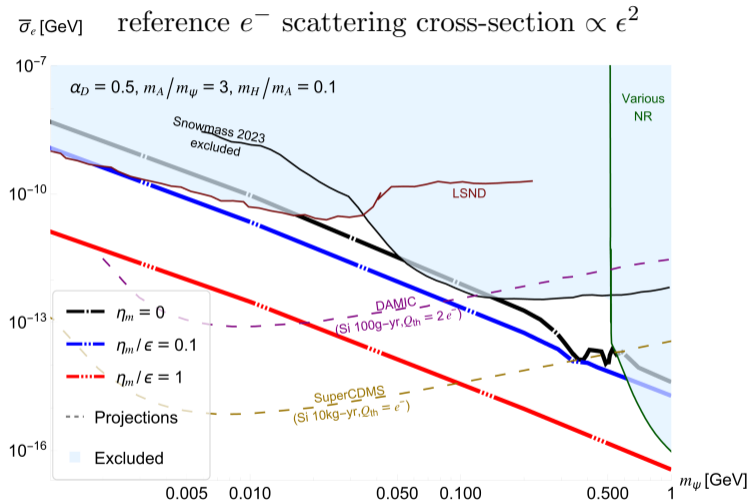
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Prospect of a heavy mediator



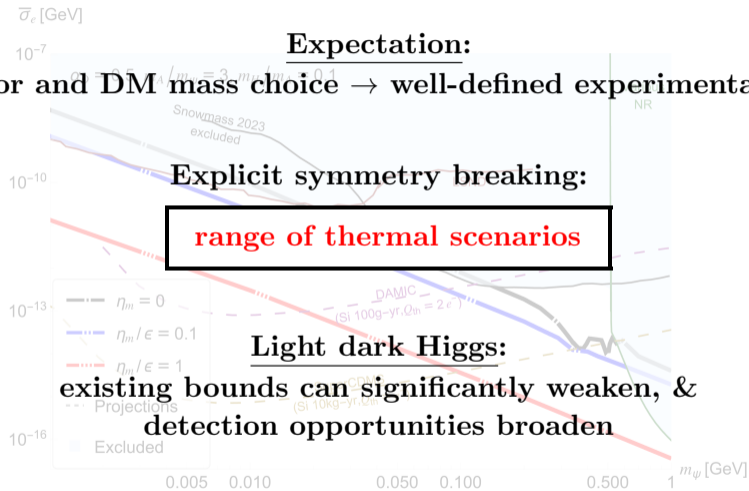
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Direct detection



LH, G. Mohlabeng, A. Mondol, & T.M.P. Tait (2026), *in prep.*. Excluded region & projections from Essig et. al (2022), *Snowmass2021 Cosmic Frontier: The landscape of low-threshold dark matter direct detection in the next decade* and Essig et. al (2016), *Direct Detection of sub-GeV Dark Matter with Semiconductor Targets*.

Expectation:
mediator and DM mass choice \rightarrow well-defined experimental target



Extra

$$\mathcal{L}_{\text{int}} \approx \epsilon e \bar{\varphi} \gamma^\mu A'_\mu \varphi. \quad (1)$$

$$\mathcal{L}_{\text{dark}} \supset |D_\mu^a \psi|^2 + |D_\mu^a \phi|^2 - V(\psi, \phi), \quad (2)$$

$$V(\psi, \phi) = \mu_\psi^2 |\psi|^2 + \lambda_\psi |\psi|^4 + (\mu_m \phi^* \psi^2 + \text{c.c.}) + \lambda_m |\phi|^2 |\psi|^2 + \mu_\phi^2 |\phi|^2 + \lambda_\phi |\phi|^4 \quad (3)$$

$$(\dot{n}_\psi + 3Hn_\psi) \left(1 + \frac{1}{3} \frac{d \ln g_s}{d \ln T}\right)^{-1} = -\frac{1}{2} \langle \sigma v \rangle \left[n_\psi^2 - (n_\psi^{\text{eq}})^2 \right] \quad (4)$$

$$\langle \sigma v \rangle_{\text{eff}} = \langle \sigma v \rangle_{\psi\psi^* \rightarrow f\bar{f}} + \langle \sigma v \rangle_{\psi\psi^* \rightarrow 2H_D} \quad (5)$$

$$\langle \sigma v_{\text{Møll}} \rangle(T) = \frac{1}{8m_\psi^4 T K_2^2 \left(\frac{m_\psi}{T}\right)} \int_{4m_\psi^2}^{\infty} ds (s - 4m_\psi^2) \sqrt{s} \sigma(s) K_1 \left(\frac{\sqrt{s}}{T}\right) \quad (6)$$

$$\sigma_{\text{SM}v_{\text{Møll}}} \approx \frac{4\pi \varepsilon^2 \alpha_D \alpha_{\text{em}} \sqrt{m_\psi^2 - m_f^2} (m_f^2 + 2m_\psi^2)}{3m_\psi \left(\Gamma_A^2 m_A^2 + (m_A^2 - 4m_\psi^2)^2 \right)} v_{\text{Møll}}^2 \quad (7)$$