

Understanding Misaligned Condensates with Yukawa Couplings

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Outline

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2. Fermion Effects: Dynamical Dressing & Particle Production
3. Unitary Equations of Motion for Dynamical Condensates
4. Asymptotic Stationary States & Other Implications
5. Conclusions

1. Condensates and the Effective Potential

Condensates of Quantum Fields

- **Coherent Field States:** Special quantum mechanical state that behaves as “classically” as possible. In bosonic systems, it is an eigenstate of the annihilation operator.
- **Field Condensates:** Non-zero expectation value of a quantum field operator. Indicates the system is “filled” with macroscopic occupation of some mode.
- **Mean Field:** An approximation where a quantum field operator is decomposed into a classical (c-number) field plus a fluctuation (field operator)



- Condensates are ubiquitous in modern particle physics (e.g. Inflation, Axions/ALPs, Higgs)

- How do we consistently describe the time-evolution of the classical object together with its quantum fluctuations?

The Effective Potential (Zero-Temp)

- The effective potential origins: how do radiative corrections modify spontaneous symmetry breaking?
- Defined as the generating functional of single particle irreducible Green's functions at zero momentum transfer.
- Useful in understanding phase transitions in quantum field theories.
- While originally computed using Feynman diagrams or functional methods, Symanzik (1970) gave a more expedient and intuitive Hamiltonian derivation (for zero temperature):

$$V_{eff}(\varphi) = \frac{1}{\mathcal{V}} \langle \Phi | H | \Phi \rangle$$

Volume   Coherent State

The Model: Scalar Condensate with Yukawa Couplings

- Consider a real scalar coupled to N_f identical massless fermions in Minkowski spacetime:

$$\mathcal{L} = \frac{1}{2} \partial_\mu \phi \partial^\mu \phi - V(\phi) + \sum_{i=1}^{N_f} \bar{\psi}_i \left(i \not{\partial} - \frac{y_0}{\sqrt{N_f}} \phi \right) \psi_i \quad V(\hat{\phi}) = V_0 + \frac{1}{2} m_0^2 \hat{\phi}^2 + \frac{\lambda_0}{4N_f} \hat{\phi}^4$$

- Which has the associated Hamiltonian:

$$H = \int d^3x \left\{ \frac{\hat{\pi}^2}{2} + \frac{(\nabla \hat{\phi})^2}{2} + V(\hat{\phi}) + \sum_{i=1}^{N_f} \psi_i^\dagger \left(-i \vec{\alpha} \cdot \vec{\nabla} + \gamma^0 \frac{y_0}{\sqrt{N_f}} \hat{\phi} \right) \psi_i \right\}$$

The Dynamical Case

- Increasingly, phenomenologists have used the effective potential to describe the time evolution of the expectation values of homogeneous fields.
- The idea is to use the equation of motion:

$$\ddot{\varphi}(t) + \frac{d}{d\varphi} V_{eff}(\varphi(t)) = 0$$

- But is this ultimately justified?
- Consider the Hamiltonian but with conditions: $\langle \Phi; 0_F | \hat{\phi}(\vec{x}, t) | \Phi; 0_F \rangle = \sqrt{N_f} \varphi(t)$; $\langle \Phi; 0_F | \hat{\pi}(\vec{x}, t) | \Phi; 0_F \rangle = \sqrt{N_f} \dot{\varphi}(t)$

$$\hat{\phi}(\vec{x}, t) \equiv \sqrt{N_f} \varphi(t) + \hat{\delta}(\vec{x}, t) \quad ; \quad \hat{\pi}(\vec{x}, t) = \sqrt{N_f} \dot{\varphi}(t) + \hat{\pi}_\delta(\vec{x}, t)$$

Time-varying mean field

Quantum Fluctuation

The Dynamical Case

- The fluctuation is subject to the constraints: $\langle \Phi; 0_F | \hat{\delta}(\vec{x}, t) | \Phi; 0_F \rangle = 0$; $\langle \Phi; 0_F | \hat{\pi}_\delta(\vec{x}, t) | \Phi; 0_F \rangle = 0$
- Obtain Heisenberg equations of motion up to order δ^3 :

$$\sqrt{N_f} \left[\ddot{\varphi}(t) + v'(\varphi(t)) + y_0 \bar{\psi}\psi \right] + \ddot{\hat{\delta}}(\vec{x}, t) - \vec{\nabla}^2 \hat{\delta}(\vec{x}, t) + \mathcal{M}^2(\varphi(t)) \hat{\delta}(\vec{x}, t) + \frac{v'''(\varphi(t))}{2\sqrt{N_f}} \hat{\delta}^2(\vec{x}, t) + \dots = 0$$

$$i \frac{\partial}{\partial t} \psi_i(\vec{x}, t) = \left(-i\vec{\alpha} \cdot \vec{\nabla} + \gamma^0 m_F(\varphi(t)) \right) \psi_i(\vec{x}, t) + \dots \quad \bar{\psi}\psi \equiv \frac{1}{N_f} \sum_{i=1}^{N_f} \bar{\psi}_i(\vec{x}, t) \psi_i(\vec{x}, t)$$

- Spatial homogeneity \rightarrow Express $\hat{\delta}$ and $\hat{\psi}$ in terms of Fourier modes.

The Dynamical Case

- Inserting into the Hamiltonian one obtains the energy density:

$$\mathcal{E} = \frac{1}{\mathcal{V}} \langle \Phi; 0_F | H | \Phi; 0_F \rangle = N_f \left\{ \underbrace{\frac{\dot{\varphi}^2}{2}}_{\text{Kinetic}} + \underbrace{v(\varphi)}_{\text{Classical Potential}} + \underbrace{\frac{1}{\mathcal{V}} \sum_{\vec{k},s} V_{-\vec{k},s}^\dagger(t) \left[\vec{\alpha} \cdot \vec{k} + \gamma^0 m_F(\varphi(t)) \right] V_{-\vec{k},s}(t)}_{\text{Fermions}} + \underbrace{\frac{1}{2N_f \mathcal{V}} \sum_{\vec{k}} \left[|\dot{g}_k(t)|^2 + \omega_k^2(t) |g_k(t)|^2 \right]}_{\text{Scalar Fluctuation}} \right\}$$

- Kinetic
- Classical Potential
- Scalar Fluctuation
- Fermions

Recall: In Minkowski Spacetime manifest energy conservation.

$$\dot{\mathcal{E}} = 0$$

We can deduce the concomitant equations of motion from this condition:

$$\dot{\mathcal{E}} = \left(\sqrt{N_f} \dot{\varphi} \right) \sqrt{N_f} \left[\ddot{\varphi}(t) + v'(\varphi(t)) + y_0 \frac{1}{\mathcal{V}} \sum_{\vec{k},s} V_{-\vec{k},s}^\dagger(t) \gamma^0 V_{-\vec{k},s}(t) + \frac{v'''(\varphi(t))}{2N_f} \frac{1}{\mathcal{V}} \sum_{\vec{k}} |g_k(t)|^2 \right]$$

- Large N limit can suppress scalar fluctuation ($N_f \rightarrow \infty$)

The Dynamical Case: Large N Limit

- Inserting into the Hamiltonian one obtains the energy density:

$$\mathcal{E} = \frac{1}{\mathcal{V}} \langle \Phi; 0_F | H | \Phi; 0_F \rangle = N_f \left\{ \frac{\dot{\varphi}^2}{2} + v(\varphi) + \frac{1}{\mathcal{V}} \sum_{\vec{k},s} V_{-\vec{k},s}^\dagger(t) \left[\vec{\alpha} \cdot \vec{k} + \gamma^0 m_F(\varphi(t)) \right] V_{-\vec{k},s}(t) + \frac{1}{2N_f \mathcal{V}} \sum_{\vec{k}} \left[|\dot{\varphi}_{\vec{k}}(t)|^2 + \omega_{\vec{k}}^2(t) |\varphi_{\vec{k}}(t)|^2 \right] \right\}$$

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- Large N limit can suppress scalar fluctuation ($N_f \rightarrow \infty$)

2. Fermion Effects: Dynamical Dressing & Particle Production

$$\ddot{\varphi}(t) + v'(\varphi(t)) + y_0 \sum_s \int V_{-\vec{k},s}^\dagger(t) \gamma^0 V_{-\vec{k},s}(t) \frac{d^3 k}{(2\pi)^3} = 0$$

$$\left[i \gamma^0 \partial_t - \vec{\gamma} \cdot \vec{k} - m_F(t) \right] U_{\vec{k},s}(t) = 0$$

$$\left[i \gamma^0 \partial_t - \vec{\gamma} \cdot \vec{k} - m_F(t) \right] V_{-\vec{k},s}(t) = 0$$

Quasi-Static/Adiabatic Approximation of Dirac Eqn.

- Express spinors in terms of mode functions and time-independent spinors

$$\begin{aligned}
 U_{\vec{k},s}(t) &= \left[i \gamma^0 \partial_t - \vec{\gamma} \cdot \vec{k} + m_F(t) \right] f_k(t) u_s \\
 V_{-\vec{k},s}(t) &= \left[i \gamma^0 \partial_t - \vec{\gamma} \cdot \vec{k} + m_F(t) \right] h_k(t) v_s
 \end{aligned}
 \quad \longrightarrow \quad
 \begin{aligned}
 \left[\frac{d^2}{dt^2} + E_k^2(t) - i \dot{m}_F(t) \right] f_k(t) &= 0 \\
 \left[\frac{d^2}{dt^2} + E_k^2(t) + i \dot{m}_F(t) \right] h_k(t) &= 0
 \end{aligned}
 \quad E_k(t) = \sqrt{k^2 + m_F^2(t)}$$

- WKB Ansatz: $f_k(t) = e^{-i \int_0^t \Omega_k(t') dt'}$ $\Omega_k(t) = \Omega_{Rk}(t) + i \Omega_{Ik}(t)$

- Adiabatic Expansion:

$$\begin{aligned}
 \Omega_{Rk}(t) &= E_k(t) - \frac{1}{4} \left(\frac{\ddot{E}_k(t)}{E_k^2(t)} - \frac{3}{2} \frac{\dot{E}_k^2(t)}{E_k^2(t)} \right) + \frac{1}{2} \left(\frac{\dot{m}_F^2(t)}{4E_k^3(t)} + \frac{\dot{m}_F^2(t) m_F(t)}{E_k^4(t)} \right) - \frac{\ddot{m}_F(t)}{4E_k^2(t)} \dots \\
 \Omega_{Ik}(t) &= -\frac{\dot{m}_F(t)}{2E_k(t)} \left[1 + \frac{m_F(t)}{E_k(t)} \right] + \dots
 \end{aligned}$$

Quasi-Static/Adiabatic Approximation of Dirac Eqn.

- Anticipating particle production, we change basis to the *zeroth order adiabatic spinors* (\mathcal{U}, \mathcal{V}):

$$U_{\vec{k},s}(t) = A_{k,s}(t) \mathcal{U}_{\vec{k},s}(t) + B_{k,s}(t) \mathcal{V}_{-\vec{k},s}(t)$$

$$V_{-\vec{k},s}(t) = C_{k,s}(t) \mathcal{V}_{-\vec{k},s}(t) + D_{k,s}(t) \mathcal{U}_{\vec{k},s}(t)$$

- The price we pay for using this basis are time-dependent Bogoliubov coefficients.

Energy Conservation and Equation of Motion

- The time-dependent fermion particle number: $n_k(t) = \bar{n}_k(t) = |B_{k,s}(t)|^2 = \frac{N_k^2 k^2 |if_k(t) - E_k(t)f_k(t)|^2}{[2E_k(t)(E_k(t) + m_F(t))]}^2$
- Thus the condensate's energy density takes the following form:

$$\mathcal{E} = N_f \left\{ \frac{\dot{\varphi}^2}{2} + \underbrace{v(\varphi)}_{v_{eff}(\varphi)} - 2 \int E_k(t) \frac{d^3k}{(2\pi)^3} + 2 \int E_k(t) \underbrace{[n_k(t) + \bar{n}_k(t)]}_{2|B_{k,s}(t)|^2} \frac{d^3k}{(2\pi)^3} \right\}$$

- Energy conservation entails the corresponding equations of motion $\dot{\mathcal{E}} = 0$

$$\ddot{\varphi} + \frac{d}{d\varphi} v_{eff}(\varphi) + 4y_0 \int k^2 \left[N_k^2 |f_k(t)|^2 - \frac{1}{2E_k(t)(E_k(t) + m_F(t))} \right] \frac{d^3k}{(2\pi)^3} = 0$$

Particle and Entropy Production

- By constructing the system's density matrix in the adiabatic basis, evolving in time one obtains:

$$\rho_S(t) = \prod_{\vec{k},s} \prod_{\vec{p},s'} \sum_{n_{\vec{k},s}=0}^1 \sum_{m_{\vec{p},s'}=0}^1 C_{m_{\vec{p},s'}}^*(p) C_{n_{\vec{k},s}}(k) |n_{\vec{k},s}; \bar{n}_{-\vec{k},s}\rangle \langle m_{\vec{p},s'}; \bar{m}_{-\vec{p},s'}| e^{2i(m_{\vec{p},s'} E_p(\infty) - n_{\vec{k},s} E_k(\infty))(t-t^*)}$$

- Off-diagonal terms average out in long time limit (decoherence via dephasing):

$$\rho_S^{(d)} = \prod_{\vec{k},s} [\cos^2(\theta_k)] \sum_{n_{\vec{k},s}=0}^1 \left(\tan^2(\theta_k) \right)^{n_{\vec{k},s}} |n_{\vec{k},s}; \bar{n}_{-\vec{k},s}\rangle \langle n_{\vec{k},s}; \bar{n}_{-\vec{k},s}|$$

Particle and Entropy Production

- This produces an effective *mixed state* with concomitant entanglement entropy:

$$S = -\text{Tr} \rho_S^{(d)} \ln \rho_S^{(d)} = - \sum_{\vec{k},s} \sum_{n_{k,s}=0}^1 p_{n_{k,s}} \ln p_{n_{k,s}}$$
$$S = -2 \sum_{\vec{k}} \left\{ \left(1 - n_k(\infty)\right) \ln \left(1 - n_k(\infty)\right) + n_k(\infty) \ln n_k(\infty) \right\}$$

- Same form as a thermal fermi gas!
 - However, distribution functions given by asymptotic particle states rather than LTE distributions.
- Why entropy from unitary evolution?
 - Decoherence in density matrix yields mixed state—Halmark of the *eigenstate thermalization hypothesis*
- Using a phenomenologically-inspired effective potential approach misses this “thermalization”.

Energy Conservation and Equation of Motion

$$\mathcal{E} = N_f \left\{ \frac{\dot{\varphi}^2}{2} + \underbrace{v(\varphi) - 2 \int E_k(t) \frac{d^3 k}{(2\pi)^3}}_{v_{eff}(\varphi)} + 2 \int E_k(t) \underbrace{\left[n_k(t) + \bar{n}_k(t) \right]}_{2|B_{k,s}(t)|^2} \frac{d^3 k}{(2\pi)^3} \right\}$$

$$\ddot{\varphi} + \frac{d}{d\varphi} v_{eff}(\varphi) + 4y_0 \int k^2 \left[N_k^2 |f_k(t)|^2 - \frac{1}{2E_k(t)(E_k(t) + m_F(t))} \right] \frac{d^3 k}{(2\pi)^3} = 0$$

- The zeroth order adiabatic contributions furnish the corrections in the effective potential.
- The higher order terms (2nd and higher) can be interpreted as inducing production of (adiabatic) fermions.
- Particle production contribution to energy density is manifestly positive!
 - Suppose condensate is “rolling down” the effective potential $\rightarrow \dot{\varphi}$ increases as does particle production!
 - This then “damps” the dynamics of the condensate!

3. Unitary Equations of Motion for Dynamical Condensates

The Dynamical Case: N=1

- Inserting into the Hamiltonian one obtains the energy density:

$$\mathcal{E} = \frac{1}{\mathcal{V}} \langle \Phi; 0_F | H | \Phi; 0_F \rangle = N_f \left\{ \underbrace{\frac{\dot{\varphi}^2}{2}}_{\text{Kinetic}} + \underbrace{v(\varphi)}_{\text{Classical Potential}} + \underbrace{\frac{1}{\mathcal{V}} \sum_{\vec{k},s} V_{-\vec{k},s}^\dagger(t) \left[\vec{\alpha} \cdot \vec{k} + \gamma^0 m_F(\varphi(t)) \right] V_{-\vec{k},s}(t)}_{\text{Fermions}} + \underbrace{\frac{1}{2N_f \mathcal{V}} \sum_{\vec{k}} \left[|\dot{g}_k(t)|^2 + \omega_k^2(t) |g_k(t)|^2 \right]}_{\text{Scalar Fluctuation}} \right\}$$

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We can deduce the concomitant equations of motion from this condition:

$$\dot{\mathcal{E}} = \left(\sqrt{N_f} \dot{\varphi} \right) \sqrt{N_f} \left[\ddot{\varphi}(t) + v'(\varphi(t)) + y_0 \frac{1}{\mathcal{V}} \sum_{\vec{k},s} V_{-\vec{k},s}^\dagger(t) \gamma^0 V_{-\vec{k},s}(t) + \frac{v'''(\varphi(t))}{2N_f} \frac{1}{\mathcal{V}} \sum_{\vec{k}} |g_k(t)|^2 \right]$$

Renormalized Equations—Energy Density

- We can now construct a fully renormalized (conserved) energy density:

$$\mathcal{E} = \underbrace{\frac{\dot{\varphi}_R^2}{2}}_{\text{1-Loop Effective Energy}} + v_{eff}(\varphi_R) + \underbrace{4 \int E_k(t) n_k(t) \frac{d^3 k}{(2\pi)^3} \Big|_R}_{\text{Renormalized Fermion Particle Production}} + \underbrace{\int \omega_k(t) \tilde{\mathcal{N}}_k(t) \frac{d^3 k}{(2\pi)^3}}_{\text{Scalar Particle Production}} = v_{eff}(\varphi_R(0))$$

$$4 \int E_k(t) n_k(t) \frac{d^3 k}{(2\pi)^3} \Big|_R = 4 \int E_k(t) n_k(t) \frac{d^3 k}{(2\pi)^3} - \underbrace{\frac{1}{2} \dot{\varphi}^2(t) \frac{y_0^2}{4\pi^2} \mathcal{I}\left[\frac{\Lambda}{\mu}; 0\right]}_{\text{Divergent "Counter Term"}}$$

4. Asymptotic Stationary States & Other Implications

Asymptotic Stationary States

- In the large N_f limit, the fermion production dominates. Energy is gradually drained from condensate.
 - Possible dynamical fixed point: $\varphi(t \rightarrow \infty) = \varphi(\infty); \dot{\varphi}(\infty) = \ddot{\varphi}(\infty) = 0$
 - By energy conservation: $v_{eff}(\varphi_R(\infty)) + 4 \int E_k(\infty) n_k(\infty) \frac{d^3k}{(2\pi)^3} = v_{eff}(\varphi_R(0))$
 - Particle Production Contribution: $\int E_k(\infty) n_k(\infty) \frac{d^3k}{(2\pi)^3} = \frac{1}{4} (v_{eff}(\varphi_R(0)) - v_{eff}(\varphi_R(\infty))) = \text{finite} > 0$
- Interpretation: The condensate produces a “bath” of fermions and settles into fixed location on the (effective) potential landscape.
 - This behavior is not possible in the phenomenologically-inspired equation of motion: $\ddot{\varphi}(t) + \frac{d}{d\varphi} V_{eff}(\varphi(t)) = 0$

Anomalous Spontaneous Symmetry Breaking?

- Dissipation by fermion production could “freeze” a condensate $|\varphi_R(\infty) \neq 0|$
 - Stable VEV w/out SSB violating minima in tree-level potential.
- This is a *possible* novel symmetry breaking mechanism which is a confluence of two effects:
 1. Radiatively induced maxima of static effective potential (Coleman-Weinberg) produces instability
 2. Misalignment dynamics with concomitant profuse particle production + conserved energy yields asymptotic stationary state
- Unlike finite temp. effects (where highly excited states restore symmetry),
 - More entanglement entropy \rightarrow more fermion/anti-fermion pairs \rightarrow larger $|\varphi_R(\infty)|$
- This is a unique result of dynamical scalar condensates coupled to (large numbers) of fermion fields.

Conclusions

- The usual effective potential **does not** correctly capture the dynamics of a *dynamical mean field*.
- However, the 0th order adiabatic furnishes a natural basis for constructing equations of motion for condensate evolution.
 - Higher order terms encode spontaneous particle production of coupled fermions or scalar fluctuations.
 - Decoherence of asymptotic density matrices suggest emergence of entanglement entropy.
 - This production mechanism efficiently transfers energy from dynamical condensate suggesting asymptotic dynamical fixed points.
 - Though this basis is used as a “pointer” the resulting equations are valid to **all adiabatic orders**.
 - Equations are energy conserving, renormalizable, and unitary.
- Real time dynamics highlights the natural emergence of wave function renormalization through “dynamical dressing” of bare condensate.
- Linger Questions:
 - Condensate evolution in cosmological spacetimes? Possible interaction with cosmological particle production? Does the classical misalignment mechanism *only* produce “cold dark matter”?
 - Gauge boson behavior?
 - Possible implications for Higgs sector?