

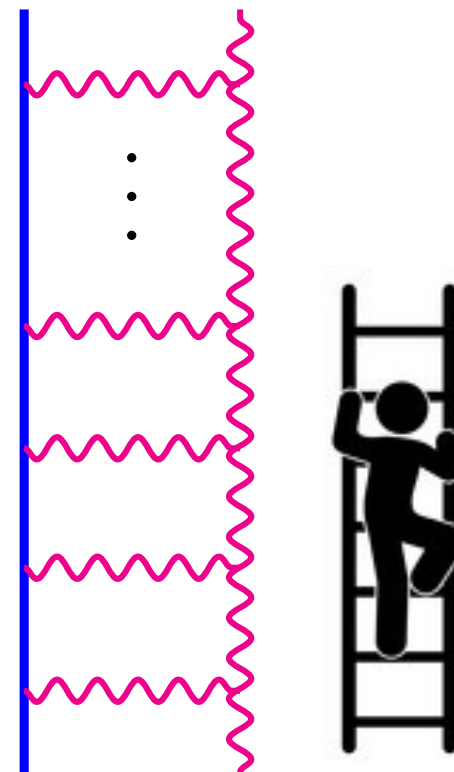
BH scattering in $N=8$ supergravity to all loops



UPPSALA
UNIVERSITET

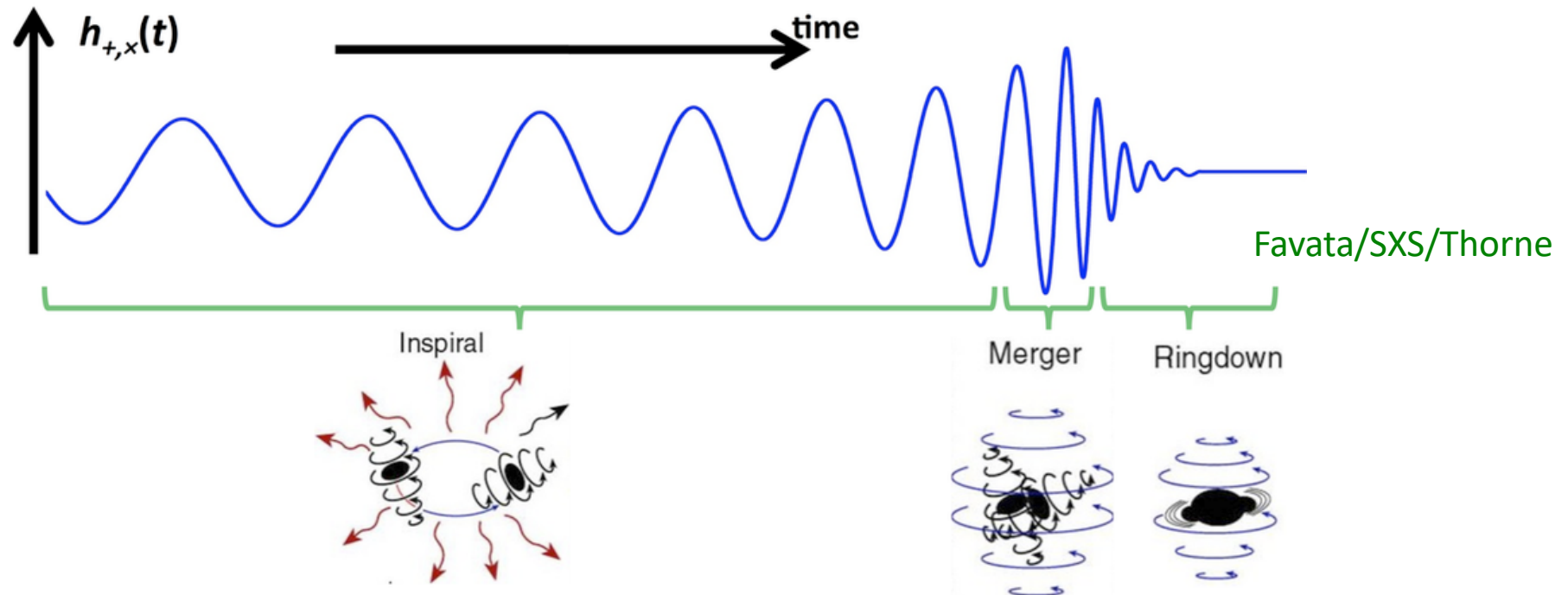
Henrik Johansson
Uppsala University

June 30, 2026



WIP: Maor Ben-Shahar, Lucile Cangemi, HJ, Chia-Hsien Shen

Motivation: gravitational waves



PM expansion: Westfahl et al.; Portilla; Ledvinka et al.; Damour; Bini, Damour; Vines; Bern, Cheung, Roiban, Shen, Solon, Zeng; Kälin, Liu, Porto; Bjerrum-Bohr, Damgaard, Festuccia, Planté, Vanhove; Ruf, Parra-Martinez; Brandhuber, Chen, Travaglini, Wen; Driesse, Jakobsen, Mogull, Plefka, Sauer, [...]

Scattering waveforms: Georgoudis, Heissenberg, Ingrid Vazquez-Holm; Rodolfo Russo, Herderschee, Roiban, Teng; Brandhuber, Brown, De Angelis, Gowdy, Chen, Travaglini; Bohnenblust, Ita, Kraus, Schlenk, [...]

See talks by: Kosower, Jakobsen, Bohnenblust, Ruf, Isabella, Chen, Shi, Brown, O'Connell

This talk: BH Compton scattering

- Compton scattering probes detailed structure of BH:

$$\mathcal{A}_{\text{Compton}} = G \left(\text{diagram 1} + \text{diagram 2} \right) + G^2 \left(\text{diagram 3} + \text{diagram 4} \right) + G^3 \left(\text{diagram 5} \right) + \dots$$

- Spin multipoles, Love/tidal, finite-size, dissip. effects (BHPT)

Caron-Huot, Miguel Correia, Isabella, Solon; Bjerrum-Bohr, Chen, Eriksen, Shah; Akpinar, del Duca, Gonzo; Akpinar; Bautista, Driesse, Haddad, Jakobsen; Ivanov, Li, Parra-Martinez, Zhou;

Bohnenblust, Eriksen, Hoogeveen, Jakobsen, Plefka; Brunello, Meo, Smith [...]

Aoude, Haddad, Helset; Bautista, Guevara, Kavanagh, Vines; Bjerrum-Bohr, Chen, Skowronek;

Cangemi, Chiodaroli, HJ, Ochirov, Pichini, Skvortsov; Ben-Shahar, Cangemi, HJ [...]

→ talks by: Jakobsen, Bohnenblust, Isabella, Chen

- Needed for $N^k\text{LO}$ calculations:



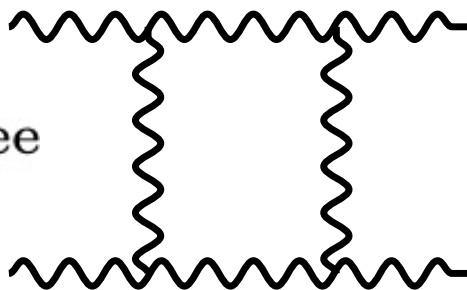
A physical toy model: $N=8$

- $N=8$ supergravity is the "best" point-particle gravity model (not an unphysical toy model)
- Quantum 4D theory: UV finite, at least up to 6 loops
- Many BPS BH solutions, natural stringy extensions
- Will focus on $\frac{1}{2}$ -BPS BH's: no-triangle property $\rightarrow$ finite size effects absent in 4D Caron-Huot, Zahraee ('18)
- BH solutions valid in $D < 11$ dimensions, goes well with dim. reg.
- Lots of experience with $N=8$ amplitudes (up to 5 loops)

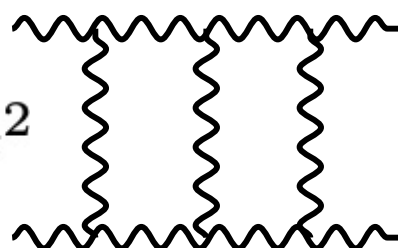
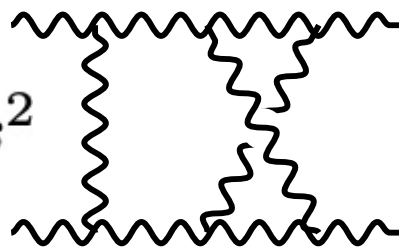
Bern, Carrasco, Chen, Edison, Johansson, Parra-Martinez, Roiban, Zeng;
Di Vecchia, Naculich, Russo, Veneziano, White; Heissenberg;
Bern, Herrmann, Roiban, Ruf, Smirnov, Smirnov, Zeng; [...]

$N=8$ supergravity amplitudes

Massless 4-graviton amplitudes:

$$stu\mathcal{M}^{\text{tree}} \left(\text{Diagram 1} \right) + \text{perms}$$


Green, Schwarz, Brink ('82)

$$stu\mathcal{M}^{\text{tree}} \left(s^2 \left(\text{Diagram 2} \right) + s^2 \left(\text{Diagram 3} \right) + \text{perms} \right)$$



Bern, Dixon, Dunbar,
Perelstein, Rozowsky ('98)

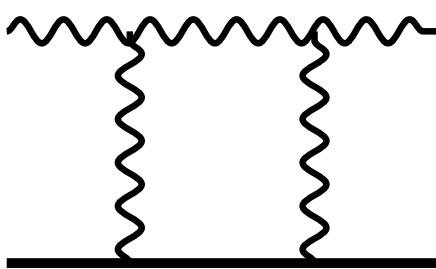
Valid in general dimension:

$$2 \leq D \leq 11$$

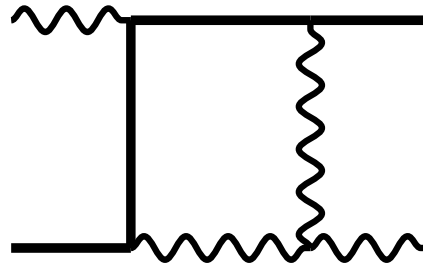
Kaluza-Klein compactification

$$SO(1, 10) \rightarrow SO(1, d) \times SO(10 - d)$$

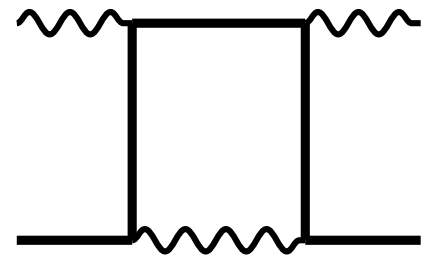
Massive multiplet $N=8$ KK momentum: $p^\mu \rightarrow (p^\mu, M, 0, \dots, 0)$



$\mathcal{O}(\hbar^0)$



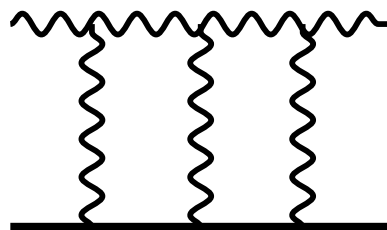
$\mathcal{O}(\hbar^1)$



$\mathcal{O}(\hbar^2)$

Classical
Limit:

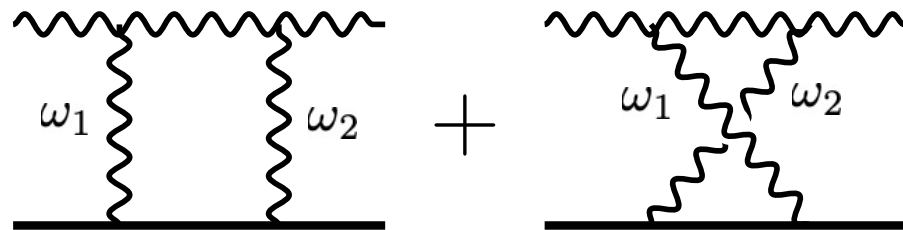
At 2 loops:



$\mathcal{O}(\hbar^0)$

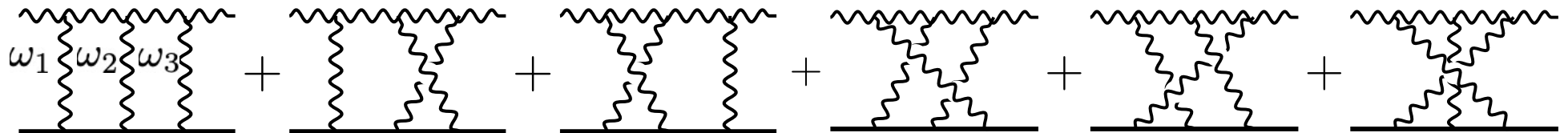
Details of classical limit

Sum over classical ladder and crossed ladder:



BH background conserves graviton energy:

$$\hat{\delta}(\omega_1 + \omega_2) \left[\frac{1}{\omega_1 - i0} + \frac{1}{\omega_2 - i0} \right] = i\hat{\delta}(\omega_1)\hat{\delta}(\omega_2)$$



$$\hat{\delta}(\omega_1 + \omega_2 + \omega_3) \left[\frac{1}{\omega_1 - i0} \frac{1}{\omega_1 + \omega_2 - i0} + \text{perm} \right] = -\hat{\delta}(\omega_1)\hat{\delta}(\omega_2)\hat{\delta}(\omega_3)$$

Saotome, Akhoury ('12)

Ladders become fan diagrams

$$\int d^D \ell \left(\text{Ladder Diagram 1} + \text{Ladder Diagram 2} \right) = \int d^d \ell \text{ Fan Diagram}$$

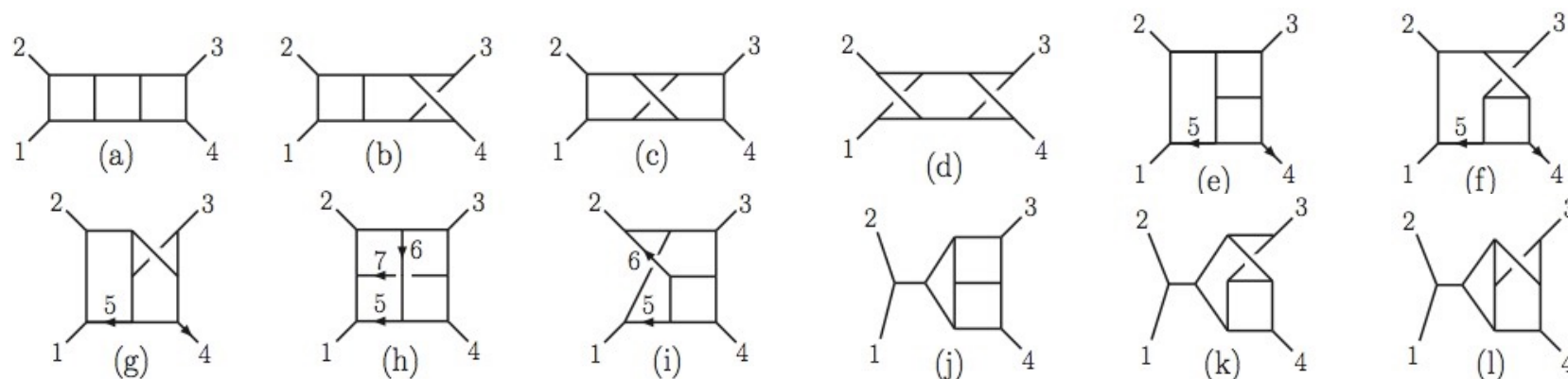
Only spatial integral: $d = D - 1$

$$\left(\begin{array}{ccc} \text{Ladder Diagram 1} & + & \text{Ladder Diagram 2} & + & \text{Ladder Diagram 3} \\ \text{Ladder Diagram 4} & + & \text{Ladder Diagram 5} & + & \text{Ladder Diagram 6} \end{array} \right) = \text{Fan Diagram}$$

Graviton energy ω $\rightarrow$ effective mass (tachyonic)

What about Numerators?

$\mathcal{N}=8$ supergravity has complicated numerators:



$\mathcal{N} = 8$ Supergravity non-BCJ numerators:
$[s^2]^2$
$[s (l_1 + k_4)^2]^2$
$(sl_{1,2}^2 + tl_{3,4}^2 - st)^2 - s^2(2(l_{1,2}^2 - t) + l_5^2)l_5^2 - t^2(2(l_{3,4}^2 - s) + l_6^2)l_6^2$ $- s^2(2l_1^2l_8^2 + 2l_2^2l_7^2 + l_1^2l_7^2 + l_2^2l_8^2) - t^2(2l_3^2l_{10}^2 + 2l_4^2l_9^2 + l_3^2l_9^2 + l_4^2l_{10}^2) + 2stl_5^2l_6^2$
$(sl_{1,2}^2 - tl_{3,4}^2)^2 - (s^2l_{1,2}^2 + t^2l_{3,4}^2 + \frac{1}{3}stu)l_5^2$

Bern, Carrasco, Dixon, HJ, Kosower, Roiban ('08)

Numerators simplify!

In the classical limit:

→ observe that all $N=8$ supergravity ladder numerators become
"rung-rule" numerators:

$$\sim stu \mathcal{M}^{\text{tree}}(s^2)^{L-1} \quad \rightarrow \quad q^2 \mathcal{M}^{\text{tree}}(2M\omega)^{2L}$$

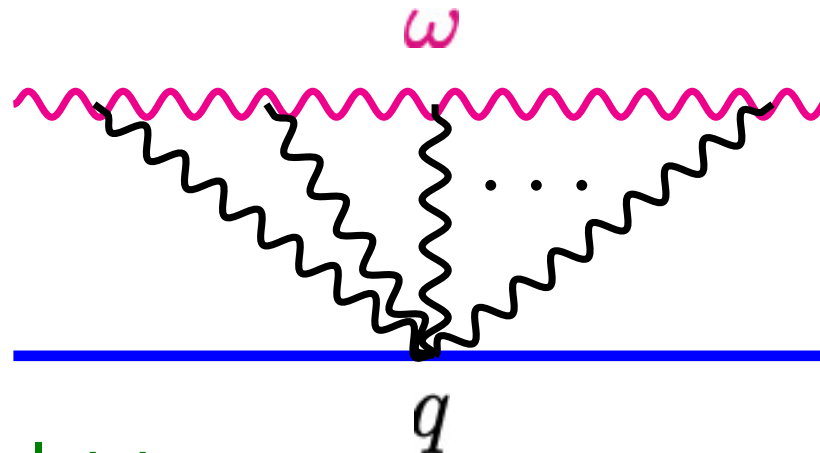
- and all non-ladder diagrams vanish (quantum)
- However, standard 3-4 loop BCJ numerators not well behaved (super-classical and quantum terms mix). [Open problem!]
- No proof that ladder "rung-rule" numerators always emerge, but one can make heuristic arguments e.g. Carrasco, Edison, HJ

Recap: $N=8$ simplifications

Ben-Shahar, Cangemi, HJ, Shen

The one diagram that rules them all:

$$q^2 \mathcal{M}^{\text{tree}} \times (4GM\omega^2)^L$$



Many options for selecting external states:

- e.g. gravitons & spinning BH:

$$\mathcal{M}^{\text{tree}}(h^+ \rightarrow h^\pm) = \frac{2GM(u \cdot \chi_\pm)^4}{q^2 \omega^2} e^{a \cdot q_\pm} + \mathcal{O}(\hbar)$$

- e.g. scalars & spherically-symmetric BH:

$$\mathcal{M}^{\text{tree}}(\phi \rightarrow \phi) = \frac{2GM\omega^2}{q^2} + \mathcal{O}(\hbar)$$

all unified by the same diagram!

$$q_\pm = k \pm k'$$

$$\chi_\pm^\mu = \{[k'|u\sigma^\mu|k], \langle k'|\sigma^\mu|k]\}$$

Integrate by solving the wave equation!

$$\square \phi = 0$$

What is the metric?

$$\frac{1}{\sqrt{-g}} \partial_\mu (\sqrt{-g} g^{\mu\nu} \partial_\nu) \phi = 0$$

Integrate by solving the wave equation!

$$\square\phi = 0$$

What is the metric?

$$\frac{1}{\sqrt{-g}}\partial_\mu(\sqrt{-g}g^{\mu\nu}\partial_\nu)\phi = 0$$

Consider a perturbative solution to: $\Delta\phi = 0$

$$\Delta = \Delta_0 + r_S\Delta_1 + r_S^2\Delta_2 + \dots$$

$$\phi = \phi_0 + \phi_1 + \phi_2 + \dots \quad \phi_0(x) = e^{-ip\cdot x}$$

$$\phi_1 = -r_S\Delta_0^{-1}\Delta_1\phi_0$$

$$\phi_2 = -r_S^2\Delta_0^{-1}\Delta_2\phi_0 + \underbrace{r_S^2\Delta_0^{-1}\Delta_1\Delta_0^{-1}\Delta_1\phi_0}_{\text{ladder/fan diagram}}$$

ladder/fan diagram

Integrate by solving the wave equation!

$$\square\phi = 0$$

What is the metric?

$$\frac{1}{\sqrt{-g}}\partial_\mu(\sqrt{-g}g^{\mu\nu}\partial_\nu)\phi = 0$$

Consider a perturbative solution to: $\Delta\phi = 0$

$$\Delta = \Delta_0 + r_S\Delta_1 + r_S^2\Delta_2 + \dots$$

$$\phi = \phi_0 + \phi_1 + \phi_2 + \dots$$

$$\phi_0(x) = e^{-ip\cdot x}$$

$$\phi_1 = -r_S\Delta_0^{-1}\Delta_1\phi_0$$

$$\phi_2 = -r_S^2\Delta_0^{-1}\Delta_2\phi_0 + \underbrace{r_S^2\Delta_0^{-1}\Delta_1\Delta_0^{-1}\Delta_1\phi_0}_{\text{ladder/fan diagram}}$$

Effective metric
linear in coupling!

$$\Delta_1 = \frac{1}{4\pi r}(-\partial_t^2) \xrightarrow{F.T.} \frac{1}{\vec{\ell}^2}\omega^2 \quad \Delta_2 = \Delta_3 = 0$$

Finding the metric!

Ben-Shahar, Cangemi, HJ, Shen

Wave equation must take the form:

$$\square\phi = \frac{1}{\sqrt{-g}}\partial_\mu(\sqrt{-g}g^{\mu\nu}\partial_\nu)\phi = \frac{1}{\sqrt{-g}}\left(\square_0 + \frac{4GM}{r}(u\cdot\partial)^2\right)\phi = 0,$$

Gives inverse metric:

$$g^{\mu\nu} = \frac{1}{\sqrt{-g}}\left(\eta^{\mu\nu} + 4GM\frac{u^\mu u^\nu}{r}\right)$$

and metric:

$$g_{\mu\nu} = \sqrt{|g|}\left(\eta_{\mu\nu} - \frac{4GM}{4GM + r}u_\mu u_\nu\right)$$

Finding the metric!

Ben-Shahar, Cangemi, HJ, Shen

Wave equation must take the form:

$$\square\phi = \frac{1}{\sqrt{-g}}\partial_\mu(\sqrt{-g}g^{\mu\nu}\partial_\nu)\phi = \frac{1}{\sqrt{-g}}\left(\square_0 + \frac{4GM}{r}(u\cdot\partial)^2\right)\phi = 0,$$

Gives inverse metric:

$$g^{\mu\nu} = \frac{1}{\sqrt{-g}}\left(\eta^{\mu\nu} + 4GM\frac{u^\mu u^\nu}{r}\right)$$

and metric:

$$g_{\mu\nu} = \sqrt{|g|}\left(\eta_{\mu\nu} - \frac{4GM}{4GM + r}u_\mu u_\nu\right)$$

line element form:

$$ds^2 = \frac{1}{\sqrt{1 + \frac{4GM}{r}}}dt^2 - \sqrt{1 + \frac{4GM}{r}}(d\vec{x}^2)$$

volume element/warp factor:

$$\sqrt{-g} = \sqrt{1 + 4GM/r}$$

The D -dimensional metric

Since the BH amplitude exists up to $D = 10$, let's generalize:

$$\square\phi = \frac{1}{\sqrt{-g}}\partial_\mu(\sqrt{-g}g^{\mu\nu}\partial_\nu)\phi = \frac{1}{\sqrt{-g}}\left(\square_0 + \frac{\lambda}{r^{D-3}}(u \cdot \partial)^2\right)\phi = 0$$

Gives inverse metric:

$$g^{\mu\nu} = \frac{1}{\sqrt{|g|}}\left(\eta^{\mu\nu} + \lambda\frac{u^\mu u^\nu}{r^{D-3}}\right)$$

and metric:

$$g_{\mu\nu} = \sqrt{|g|}\left(\eta_{\mu\nu} - \frac{\lambda}{\lambda + r^{D-3}}u_\mu u_\nu\right)$$

The D -dimensional metric

Since the BH amplitude exists up to $D = 10$, let's generalize:

$$\square\phi = \frac{1}{\sqrt{-g}}\partial_\mu(\sqrt{-g}g^{\mu\nu}\partial_\nu)\phi = \frac{1}{\sqrt{-g}}\left(\square_0 + \frac{\lambda}{r^{D-3}}(u \cdot \partial)^2\right)\phi = 0$$

Gives inverse metric:

$$g^{\mu\nu} = \frac{1}{\sqrt{|g|}}\left(\eta^{\mu\nu} + \lambda\frac{u^\mu u^\nu}{r^{D-3}}\right)$$

and metric:

$$g_{\mu\nu} = \sqrt{|g|}\left(\eta_{\mu\nu} - \frac{\lambda}{\lambda + r^{D-3}}u_\mu u_\nu\right)$$

Mass parameter:

$$\lambda = 4GM\frac{\Gamma(\frac{1}{2}(D-3))}{\pi^{\frac{1}{2}(D-3)}}$$

volume element/warp factor:

$$\sqrt{|g|} = \left(1 + \lambda/r^{D-3}\right)^{\frac{1}{D-2}}$$

This is the known (KK) metric of a D0 brane in maximal supergravity!

Solving the wave 4D equation to all loops

The BH wave equation $\left(\square_0 + \frac{4GM}{r}(u \cdot \partial)^2\right)\phi = 0$

is now equivalent to the Coloumb wave eqn in flat space

$$\left(\nabla^2 + \omega^2 + \frac{4GM\omega^2}{r}\right)\phi = 0$$

The radial equation $v(r) = rR(r)$

$$\frac{d^2v}{dr^2} + \left(\omega^2 + \frac{4GM\omega^2}{r} - \frac{\ell(\ell+1)}{r^2}\right)v = 0$$

have well-known solutions (hypergeometric/Whittaker)

Solving the wave 4D equation to all loops

The BH wave equation $\left(\square_0 + \frac{4GM}{r}(u \cdot \partial)^2\right)\phi = 0$

is now equivalent to the Coloumb wave eqn in flat space

$$\left(\nabla^2 + \omega^2 + \frac{4GM\omega^2}{r}\right)\phi = 0$$

The radial equation $v(r) = rR(r)$

$$\frac{d^2v}{dr^2} + \left(\omega^2 + \frac{4GM\omega^2}{r} - \frac{\ell(\ell+1)}{r^2}\right)v = 0$$

have well-known solutions (hypergeometric/Whittaker)

The all loop amplitude is then

$$\mathcal{M}(\theta) = \frac{1}{2i|\omega|} \sum_{\ell=0}^{\infty} (2\ell+1) (e^{2i(\delta_{\ell} + \delta_{\text{IR}})} - 1) P_{\ell}(\cos \theta)$$

with the partial-wave phase shift $e^{2i\delta_{\ell}} = \frac{\Gamma(\ell+1+i2GM\omega)}{\Gamma(\ell+1-i2GM\omega)}$

Resumming the 4D $N=8$ Compton amplitude

Ben-Shahar, Cangemi, HJ, Shen

Shift-enter directly gives:

$$\mathcal{M}(\theta) = -\frac{GM}{2 \sin^2(\theta/2)} \exp \left[-2iGM\omega \ln \sin \frac{\theta}{2} + 2i\delta_0 + 2i\delta_{\text{IR}} \right]$$

Using $\sin^2 \frac{\theta}{2} = -q^2/(4\omega^2)$ one can massage it into

$$\begin{aligned} \mathcal{M}(\theta) &= \frac{2GM\omega^2}{q^2} \exp \left[iGM\omega \ln \frac{4\omega^2}{-q^2} + 2i\delta_0 + 2i\delta_{\text{IR}} \right] \\ &= \frac{2GM\omega^2}{q^2} \exp \left[iGM\omega \left(\frac{1}{\epsilon} - \ln \frac{-q^2}{\mu^2} \right) + 2i\delta_0 \right] \end{aligned}$$

Resumming the 4D $N=8$ Compton amplitude

Ben-Shahar, Cangemi, HJ, Shen

Shift-enter directly gives:

$$\mathcal{M}(\theta) = -\frac{GM}{2 \sin^2(\theta/2)} \exp \left[-2iGM\omega \ln \sin \frac{\theta}{2} + 2i\delta_0 + 2i\delta_{\text{IR}} \right]$$

Using $\sin^2 \frac{\theta}{2} = -q^2/(4\omega^2)$ one can massage it into

$$\begin{aligned} \mathcal{M}(\theta) &= \frac{2GM\omega^2}{q^2} \exp \left[iGM\omega \ln \frac{4\omega^2}{-q^2} + 2i\delta_0 + 2i\delta_{\text{IR}} \right] \\ &= \frac{2GM\omega^2}{q^2} \exp \left[iGM\omega \left(\frac{1}{\epsilon} - \ln \frac{-q^2}{\mu^2} \right) + 2i\delta_0 \right] \\ &= \mathcal{M}^{\text{tree}} \exp \left[i \frac{\mathcal{M}^{(1)}}{\mathcal{M}^{\text{tree}}} + 2i\delta_0 \right] \end{aligned}$$

BDS-like form!

$$\delta_0 = -GM\omega\gamma_E + (GM\omega)^3 \frac{\zeta_3}{3} - (GM\omega)^5 \frac{\zeta_5}{5} + (GM\omega)^7 \frac{\zeta_7}{7} + \dots$$

IR divergence -- sleight of hand

$$\begin{aligned}\mathcal{M}(\theta) &= \frac{2GM\omega^2}{q^2} \exp \left[iGM\omega \ln \frac{4\omega^2}{-q^2} + 2i\delta_0 + 2i\delta_{\text{IR}} \right] \\ &= \frac{2GM\omega^2}{q^2} \exp \left[iGM\omega \left(\frac{1}{\epsilon} - \ln \frac{-q^2}{\mu^2} \right) + 2i\delta_0 \right]\end{aligned}$$

The 4D IR divergent phase is matched to dim. reg. by sleight of hand

$$\delta_{\text{IR}} = -GM\omega \ln(4\omega^2 R^2) \stackrel{!}{=} -\frac{GM\omega}{\epsilon} \left(\frac{\mu^2}{4\omega^2} \right)^{-\epsilon} = -GM\omega \left(\frac{1}{\epsilon} + \ln \frac{4\omega^2}{\mu^2} + \mathcal{O}(\epsilon) \right)$$

IR divergence -- sleight of hand

$$\begin{aligned}\mathcal{M}(\theta) &= \frac{2GM\omega^2}{q^2} \exp \left[iGM\omega \ln \frac{4\omega^2}{-q^2} + 2i\delta_0 + 2i\delta_{\text{IR}} \right] \\ &= \frac{2GM\omega^2}{q^2} \exp \left[iGM\omega \left(\frac{1}{\epsilon} - \ln \frac{-q^2}{\mu^2} \right) + 2i\delta_0 \right]\end{aligned}$$

The 4D IR divergent phase is matched to dim. reg. by sleight of hand

$$\delta_{\text{IR}} = -GM\omega \ln(4\omega^2 R^2) \stackrel{!}{=} -\frac{GM\omega}{\epsilon} \left(\frac{\mu^2}{4\omega^2} \right)^{-\epsilon} = -GM\omega \left(\frac{1}{\epsilon} + \ln \frac{4\omega^2}{\mu^2} + \mathcal{O}(\epsilon) \right)$$

→ Should be done more carefully!

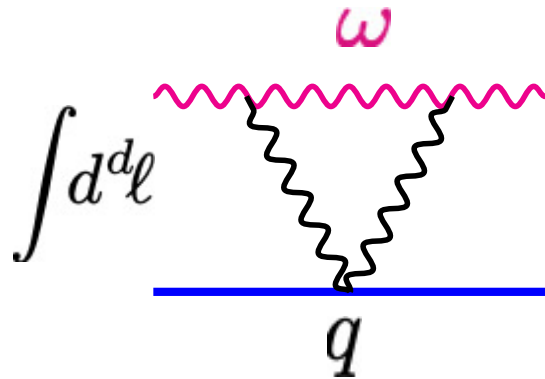
→ Proper treatment in $4 - 2\epsilon$ dimensions!

→ $N=8$ BH exists up to $D = 10$ dimensions, why not give it a go?

D-dimensional warm-up at 1-loop

Ben-Shahar, Cangemi, HJ, Shen

Consider the 1-loop classical
Compton amplitude in d -dimensions:



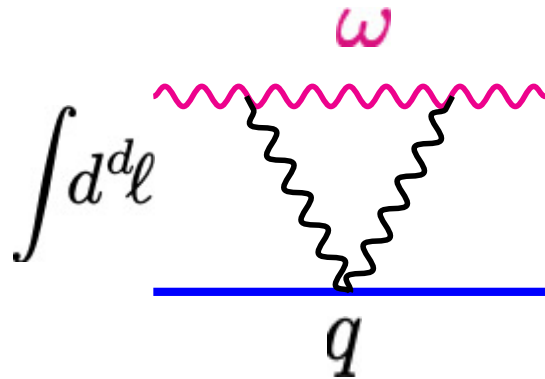
Observations:

- IR divergent: $d \leq 3$
- UV divergent: $d \geq 6$
- divide by $\Gamma(3 - d/2)$
renders it UV finite!
- evaluates to $\log()$ in integer
dimensions $d \geq 5$

D-dimensional warm-up at 1-loop

Ben-Shahar, Cangemi, HJ, Shen

Consider the 1-loop classical
Compton amplitude in d -dimensions:



Observations:

- IR divergent: $d \leq 3$
- UV divergent: $d \geq 6$
- divide by $\Gamma(3 - d/2)$
renders it UV finite!
- evaluates to $\log()$ in integer
dimensions $d \geq 5$

Work out the integer dimensional results $\rightarrow$ closed d -dimensional formula

$$\begin{aligned}
 f^{(d)} &= \frac{(1 + y^2)^{(3-d)/2}}{d - 4} \left[(y^2)^{d/2-2} \left(\ln(-y^2) + \sum_{k=1}^{(d-5)/2} \frac{(1 + y^2)^k}{k} \right) \right. \\
 &\quad \left. - i^{1-d} 2^{(d-3)/2} \sqrt{\pi} \Gamma(d/2 - 1) \left(\frac{(-y^2)^{d/2-2}}{(d-4)!!} - \sum_{k=0}^{(d-5)/2} \frac{(-1)^k (1 + y^2)^k}{2^k k! (d-4-2k)!!} \right) \right] \\
 &\propto \left(\frac{4\omega^2}{q^2} \right)^{d/2-2} \Phi\left(1 + \frac{4\omega^2}{q^2}, 1, \frac{d-3}{2}\right) - B\left(\frac{1}{2}, \frac{d}{2} - 1\right) {}_2F_1\left(\frac{1}{2}; 1; \frac{d-1}{2}, 1 + \frac{4\omega^2}{q^2}\right)
 \end{aligned}$$

Partial-wave amplitude at 1-loop

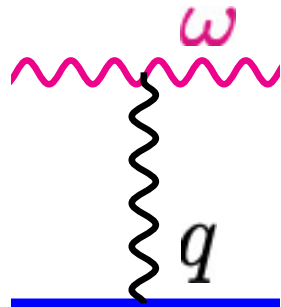
In general d Gegenbauer polynomials describe amplitude

$$\mathcal{M}(\theta) = \frac{1}{2i|\omega|} \sum_{\ell=0}^{\infty} \frac{2\ell + d - 2}{d - 2} (e^{2i(\delta_{\ell} + \delta_{\text{IR}})} - 1) C_{\ell}^{(d/2-1)}(\cos \theta)$$

For fixed ℓ, d it is easy to transform to partial waves:

First, at tree-level we get that the q propagator transform to

$$e^{2i(\delta_{\ell} + \delta_{\text{IR}})} = 1 - \frac{2iGM|\omega|\ell!}{(d-3)(d-1)_{\ell-1}} + \mathcal{O}(G^2)$$



IR divergent in $d \leq 3$

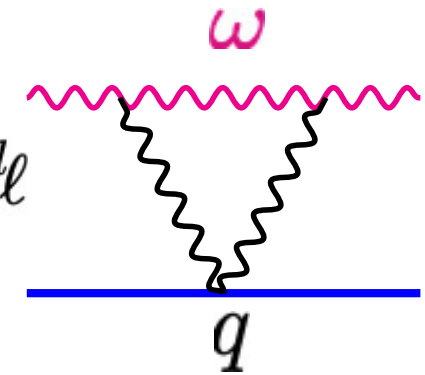
Partial-wave amplitude at 1-loop

In general d Gegenbauer polynomials describe amplitude

$$\mathcal{M}(\theta) = \frac{1}{2i|\omega|} \sum_{\ell=0}^{\infty} \frac{2\ell + d - 2}{d - 2} (e^{2i(\delta_{\ell} + \delta_{\text{IR}})} - 1) C_{\ell}^{(d/2-1)}(\cos \theta)$$

For fixed ℓ, d it is easy to transform to partial waves:

At 1-loop level we get that triangle transforms to

$$\frac{d-2}{2\ell+d-2} \frac{\mathcal{M}_{\ell}^{(1)}}{G^2 M^2 \omega^2} = - \frac{(-i)^d 2^{2d-11}}{(d-4)(d-2)} \left[\frac{\ell!}{(d-3)(d-1)_{\ell-1}} \right]^2 \int d^d \ell$$


$$+ (-1)^{\ell} 2^{d-\ell-8} B\left(\frac{1}{2}, \frac{d}{2} - 1\right) \frac{\Gamma(d-4)\Gamma(d/2)\Gamma(2d-5)}{\Gamma(d+\ell-2)^2 \Gamma(3d/2+\ell-4)} \mathcal{P}^{(\ell)}(d)$$

Recursion formula for polynomial:

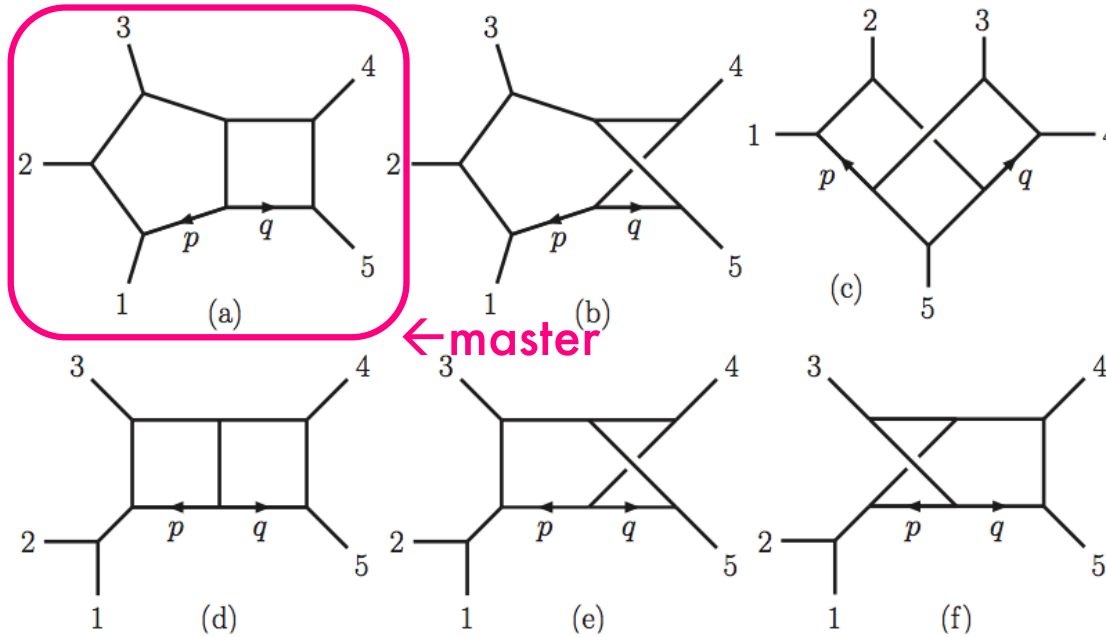
$$\mathcal{P}^{(\ell+2)} = (2\ell + d)(d^2 - 2\ell^2 - 2d(\ell + 5) - 16)\mathcal{P}^{(\ell+1)} - (\ell + 1)^2(\ell + d - 2)^2(2\ell - d + 4)(2\ell + 3d + 8)\mathcal{P}^{(\ell)},$$

Conclusion

- $N = 8$ supergravity is a testbed for advanced classical calc.
- Compton ampl. is dramatically simpler than: $2 \rightarrow 2$, waveform
- Showed that all-loop resummation is possible for $\frac{1}{2}$ BPS BH
- Identified the metric with D0 brane (11D KK mode)
- BH has no horizon, no finite-size corrections. Exact PP.
- Calculation captures infinitely many ladder diagrams of GR
- Further work required for higher-loop D -dimensional solution
- Outlook: Analyzing spin multipoles; UV divergences/finite-size corrections in higher dimensions, connection to $\frac{1}{4}$ -BPS BH.

Extra slides

Example: 2-loop 5-pts $\mathcal{N}=4$ SYM & $\mathcal{N}=8$ SG



Carrasco, HJ
1106.4711

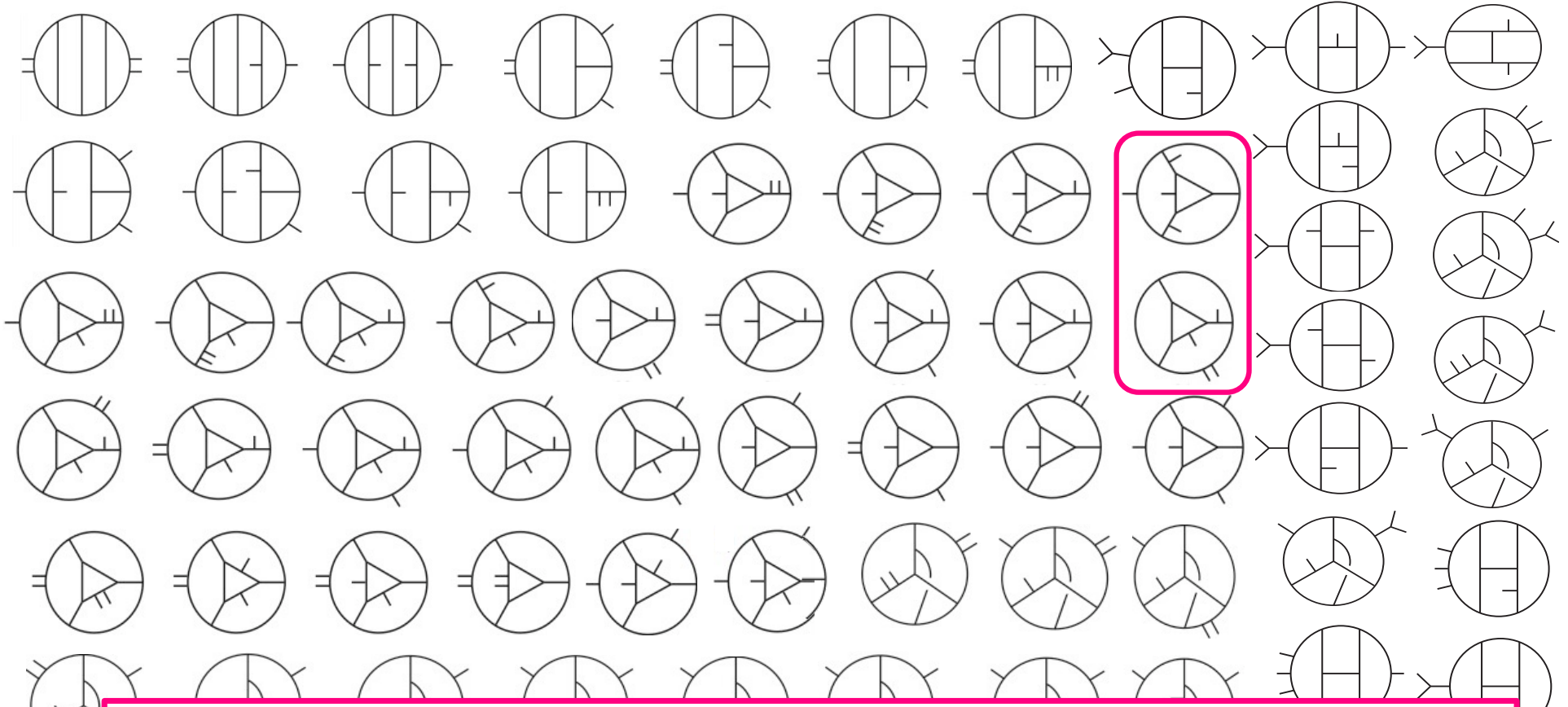
color-kinematics duality
+ unitarity method
→ simple numerators

$\mathcal{I}^{(x)}$	$\mathcal{N} = 4$ Super-Yang-Mills ($\sqrt{\mathcal{N}} = 8$ supergravity) numerator
(a),(b)	$\frac{1}{4} \left(\gamma_{12}(2s_{45} - s_{12} + \tau_{2p} - \tau_{1p}) + \gamma_{23}(s_{45} + 2s_{12} - \tau_{2p} + \tau_{3p}) \right. \\ \left. + 2\gamma_{45}(\tau_{5p} - \tau_{4p}) + \gamma_{13}(s_{12} + s_{45} - \tau_{1p} + \tau_{3p}) \right)$
(c)	$\frac{1}{4} \left(\gamma_{15}(\tau_{5p} - \tau_{1p}) + \gamma_{25}(s_{12} - \tau_{2p} + \tau_{5p}) + \gamma_{12}(s_{34} + \tau_{2p} - \tau_{1p} + 2s_{15} + 2\tau_{1q} - 2\tau_{2q}) \right. \\ \left. + \gamma_{45}(\tau_{4q} - \tau_{5q}) - \gamma_{35}(s_{34} - \tau_{3q} + \tau_{5q}) + \gamma_{34}(s_{12} + \tau_{3q} - \tau_{4q} + 2s_{45} + 2\tau_{4p} - 2\tau_{3p}) \right)$
(d)-(f)	$\gamma_{12}s_{45} - \frac{1}{4} \left(2\gamma_{12} + \gamma_{13} - \gamma_{23} \right) s_{12}$

$$\tau_{ip} = 2k_i \cdot p$$

direct calculation in
GR would naively
give $\sim 10^{18}$ terms

4-loops: 85 diagrams, 2 masters



$$\begin{aligned}
 N_{18} &= \frac{1}{4}(6u^2\tau_{25} + u(2s(5\tau_{25} + 2\tau_{26}) - \tau_{15}(7\tau_{16} + 6t)) \\
 &\quad + t(\tau_{15}\tau_{26} - \tau_{25}(\tau_{16} + 7\tau_{26})) + s(4\tau_{15}(t - \tau_{26}) + 6\tau_{36}(\tau_{35} - \tau_{45}) \\
 &\quad - \tau_{16}(4t + 5\tau_{25}) - \tau_{46}(5\tau_{35} + \tau_{45})) + 2s^2(t + \tau_{26} - \tau_{35} + \tau_{36} + \tau_{56})), \\
 N_{28} &= \frac{1}{4}(s(2\tau_{15}t + \tau_{16}(2t - 5\tau_{25} + \tau_{35}) + 5\tau_{35}(\tau_{26} + \tau_{36}) + 2t(2\tau_{46} - \tau_{56}) - 10u\tau_{25} \\
 &\quad - 4s^2\tau_{25} - 6u(\tau_{46}(t - \tau_{25} + \tau_{45}) + \tau_{25}\tau_{26}) - t(\tau_{15}(4\tau_{36} + 5\tau_{46}) + 5\tau_{25}\tau_{36})).
 \end{aligned}$$

Kerr/root-Kerr Effective theories

From matching to Kerr/root-Kerr amplitudes one can infer EFTs:

EFTs	$s = 1/2$	$s = 1$	$s = 3/2$	$s = 2$	$s = 5/2$	$s \geq 3$
Kerr	Major.	Proca	Rar.-Sch.	KK grav.	HS	HS
$\sqrt{\text{Kerr}}$	Dirac	W -boson	gravitino	HS	HS	HS

Cangemi,
Chiodaroli,HJ,
Ochirov,
Pichini,
Skvortsov

Massless limit exist for $s \leq 2$ **Kerr** and $s \leq 1$ **root-Kerr**

Kerr and **root-Kerr** related by double copy:

$$M(1\phi^s, 2\bar{\phi}^s, 3h^\pm) = iA(1\phi^{s_L}, 2\bar{\phi}^{s_L}, 3A^\pm)A(1\phi^{s_R}, 2\bar{\phi}^{s_R}, 3A^\pm)$$

For $s > 2$ **Kerr dynamics = higher-spin and higher-derivative EFTs**