

Higher Order Radiative Corrections and Experimental Activities: Examples from BESIII

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The BESIII Experiment



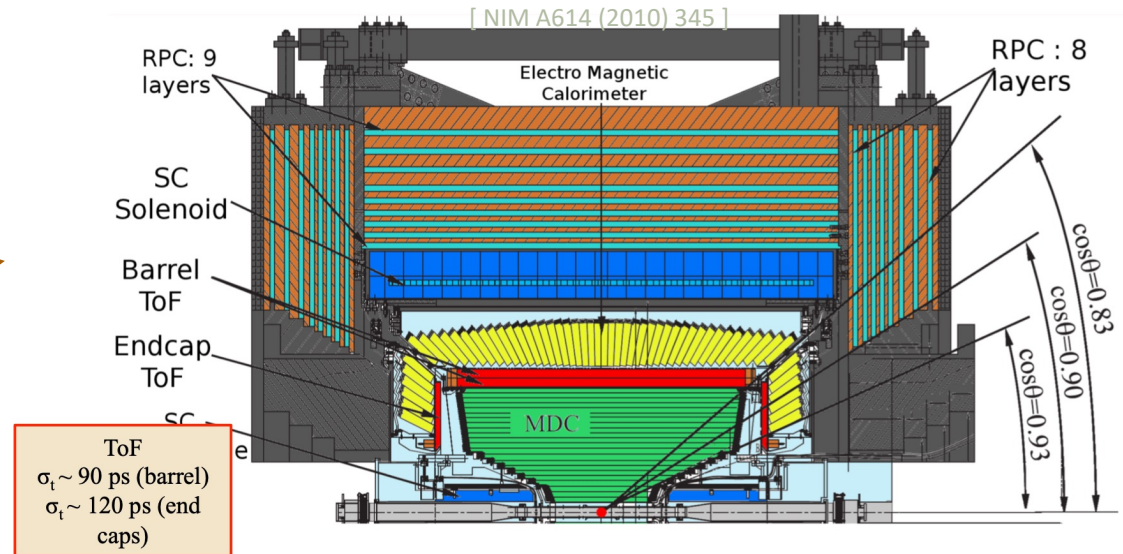
Located at the BEPCII collider
(Beijing, China)

Symmetric e^+e^- beams

ECM between 1.8-5 GeV

Maximum luminosity: $> 1 \text{ nb}^{-1}/\text{s}$

93% coverage of the solid angle



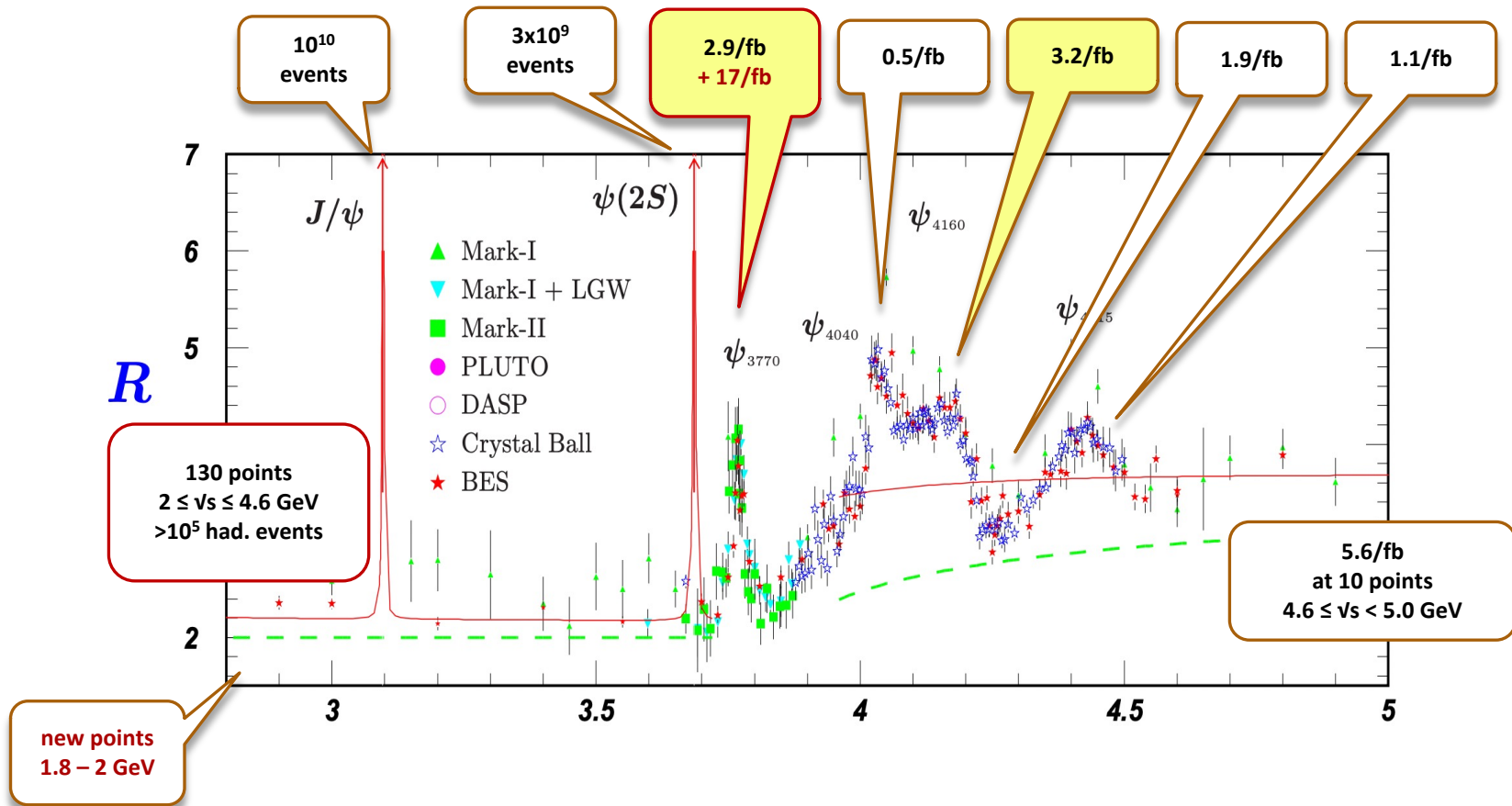
ToF
 $\sigma_t \sim 90 \text{ ps}$ (barrel)
 $\sigma_t \sim 120 \text{ ps}$ (end caps)

Drift Chamber
 $\sigma_{r\phi} \sim 130 \mu\text{m}$ (single wire)
 $\sigma_{p_t}/p_t \sim 0.5 \%$ @ 1 GeV

Electromagnetic CsI(Tl) Calorimeter
 $\sigma_E/E < 2.5\%$ @ 1 GeV (barrel)
 $\sigma_E/E < 5\%$ @ 1 GeV (end caps)
 $\sigma_{xy} \sim (6 \text{ mm})/E^{1/2}$ @ 1 GeV

RPC Muon Detector
 $\Delta\Omega/4\pi = 93\%$

BESIII Contributions to HVP

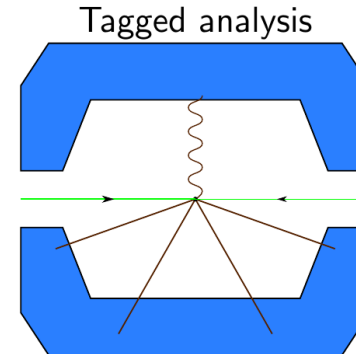
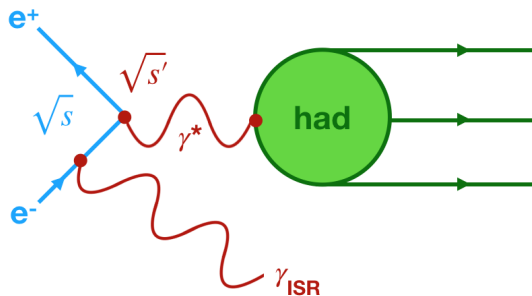


World largest τ -charm dataset in e^+e^- annihilation ...

... and still **growing**

Pion Form Factor: impact of beyond NLO contributions

Measurement Strategy: Old and New



- Employ initial state radiation to access relevant energy range
- Radiated photon at large angle – kinematic threshold
- Two charged tracks, identified as pions (PID), originating from IP

Published: [Phys.Lett.B753 (2016) 629]

- Detect ISR photon
- 4C Kinematic Fit ($\pi\pi\gamma$)
- Phokhara **Accuracy: 0.9%**

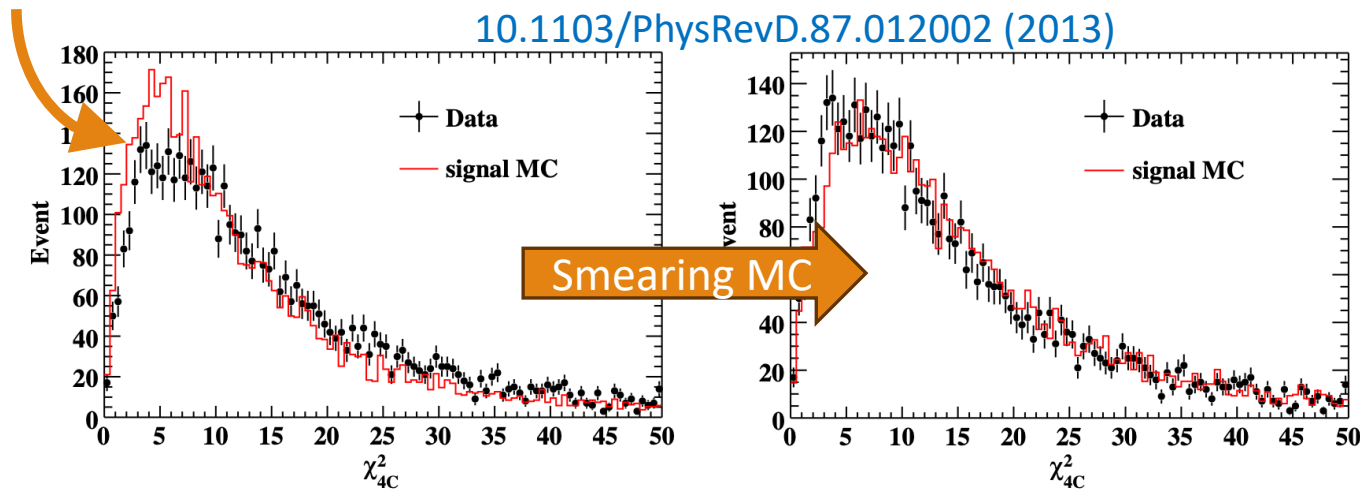
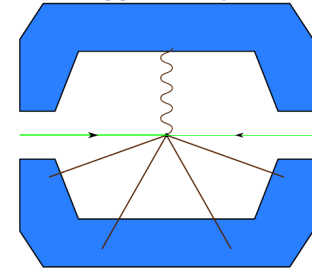
New:

Aiming to 0.7%

- Do not look for ISR photon
- 1C Kinematic Fit ($\pi\pi\gamma$)
- Babayaga@NLO (new), Phokhara

Kinematic Fit Corrections

- Small differences between detector response in data and MC expected
- Correction/calibration factor to take this into account
- Two components for kinematic fit:
 - Photon detection efficiency (only 4C)
 - Track momentum resolution $\rightarrow \chi^2$ (1C & 4C)



- Use control sample to correct for the difference

Photon Detection Efficiency

Using radiative $ee \rightarrow \mu\mu$ events:

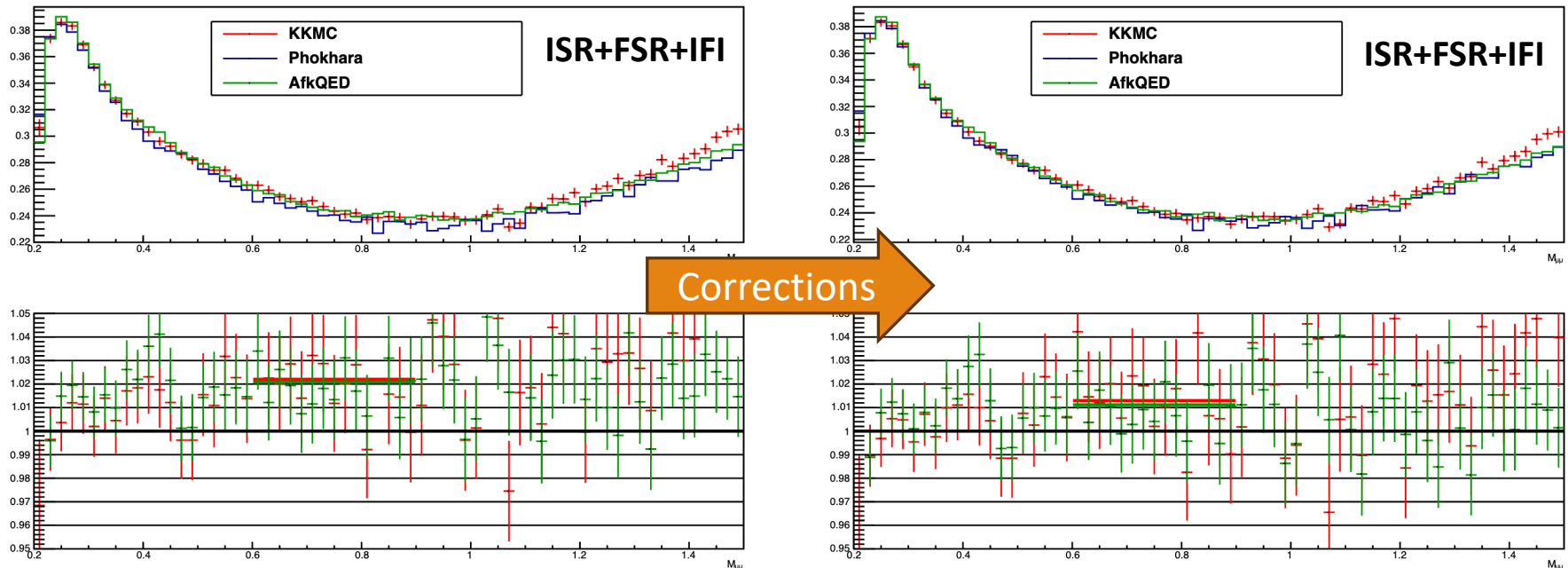
- Perform **1C Kinematic fit** (**photon momentum prediction**, $\chi^2 < 10$)
- **Look for a photon** closing the kinematics (**4C Kinematic fit**)
- **Determine** efficiency (and relative **corrections**) for data and MC in bins of energy and polar angle

Such procedure corrects for differences in:

- Detection efficiency
- Resolution
- **Missing radiative corrections, as long as accepted by 1C**

Tested using KKMC/AfkQED as data and correcting Phokhara

Impact of Correction



Corrections have an impact on the agreement!

Discrepancy reduced from 2% to 1%

➤ Still **1% of higher order contributions lost** by kinematic fit

New inputs: Babayaga@NLO

New version of **Babayaga@NLO** provides **now full NLO + PS** matching for $\mu\mu/\pi\pi\gamma$:

- Full set of corrections, includes correctly VP
- **Can verify pion/muon differences**

Compare Phokhara 10.0 vs Babayaga@NLO (2026)

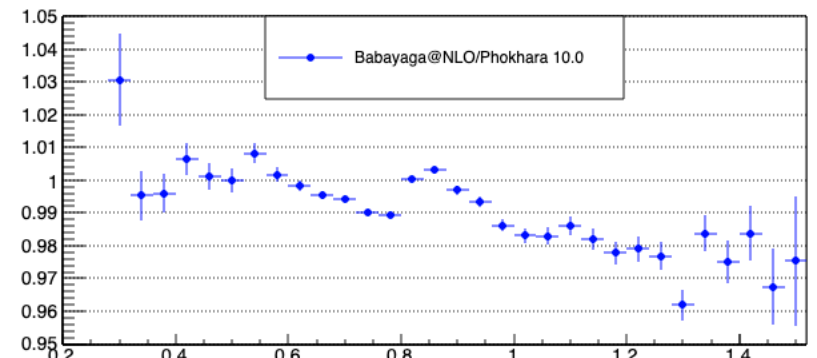
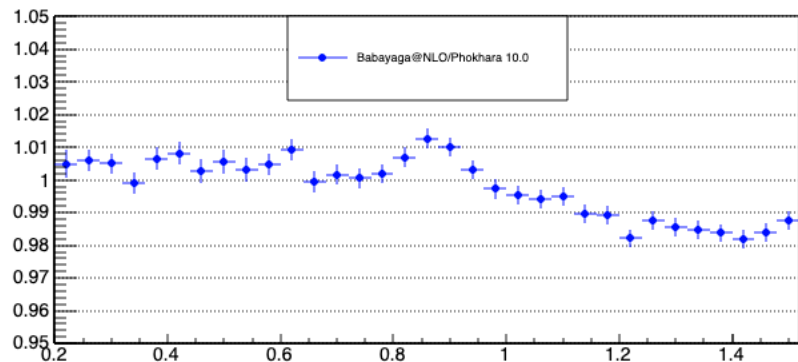
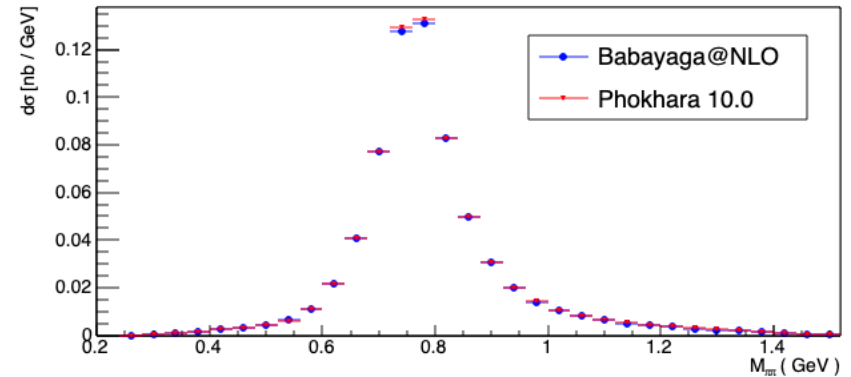
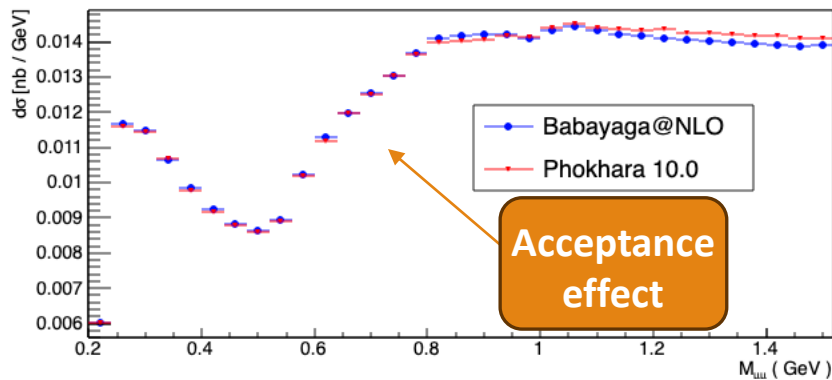
Generator settings:

- $E_{cm} = 3.773$ GeV
- Charged particles in $[15^\circ, 165^\circ]$
- Photons in $[0^\circ, 180^\circ]$
- Pion FF: Phokhara (“new GS” in Phokhara 10.0)
- VP: nsk (Babayaga), hadr5 (Phokhara)
- Unweighted events

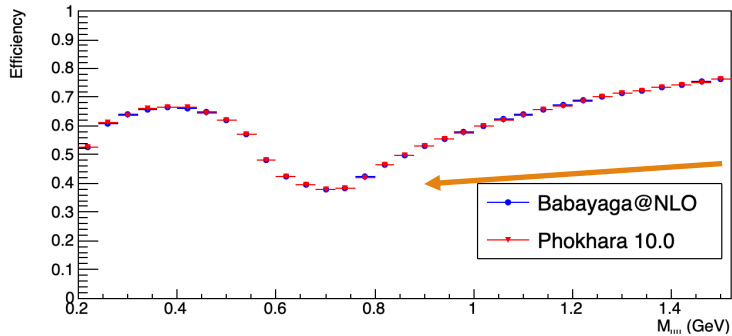
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New version of Babayaga provides now full NLO + PS matching for $\mu\mu\gamma/\pi\pi\gamma$:

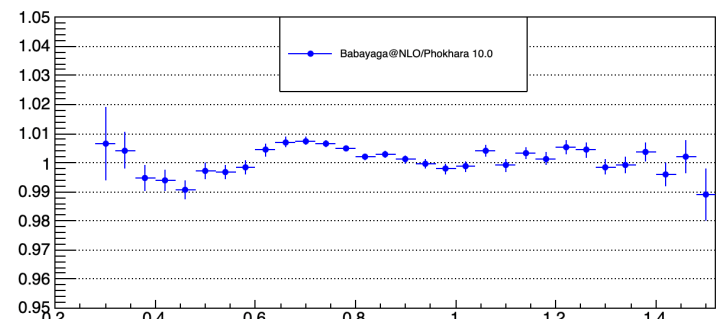
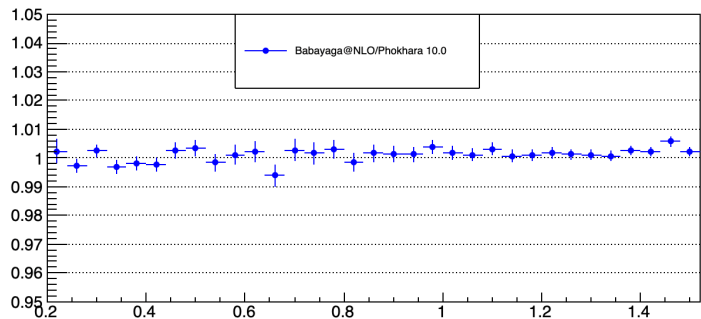
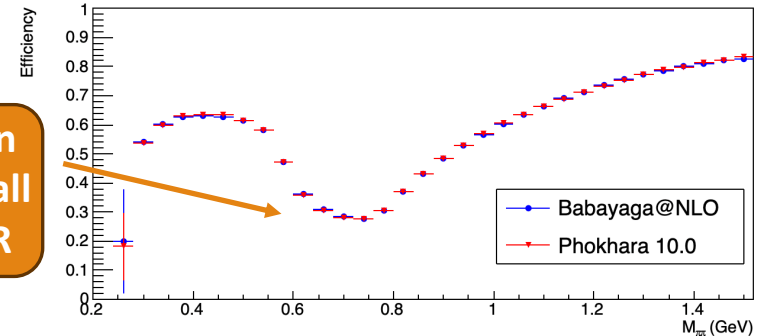
- Full set of corrections, includes correctly VP
- Can verify pion/muon differences



Geometrical Selection

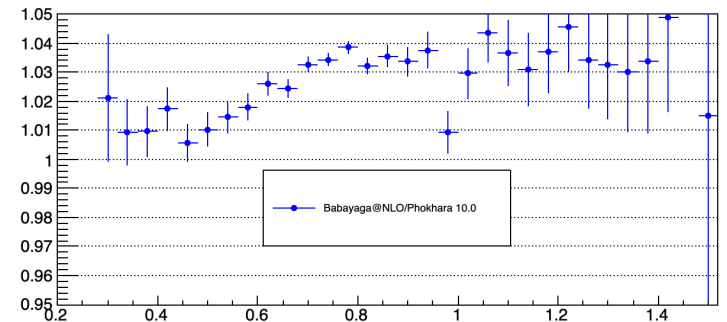
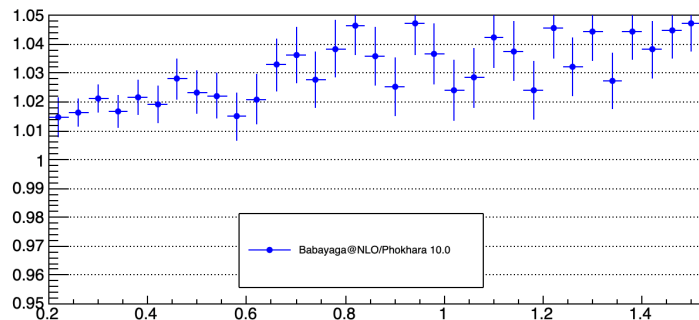
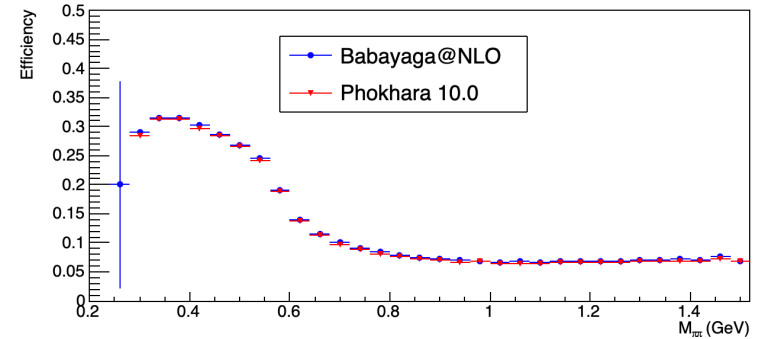
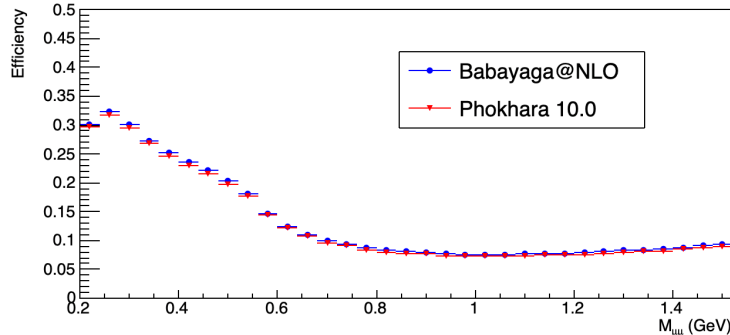


Transition
Large-Small
Angle ISR



- Small differences in the cross section do not propagate to the efficiency ($\mu\mu$)
 - Better than 0.5% agreement!
- A larger effect is observed for pion, still well within 1%

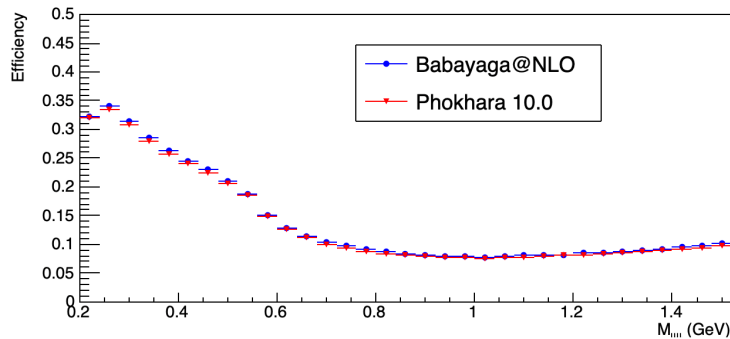
Introduce Kinematic Fit: 4C



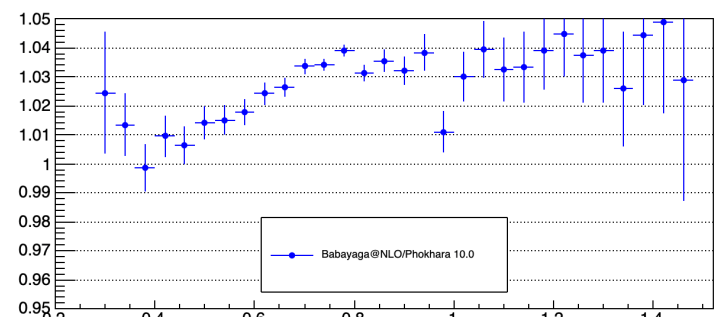
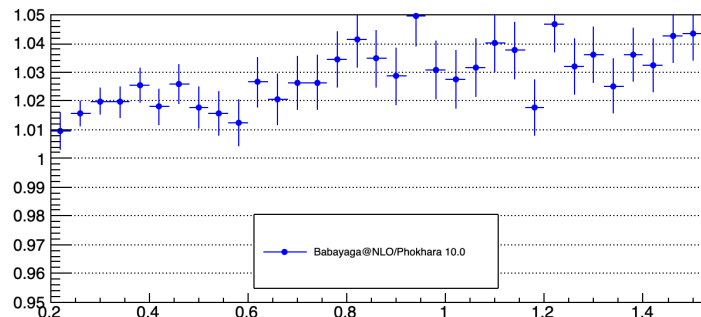
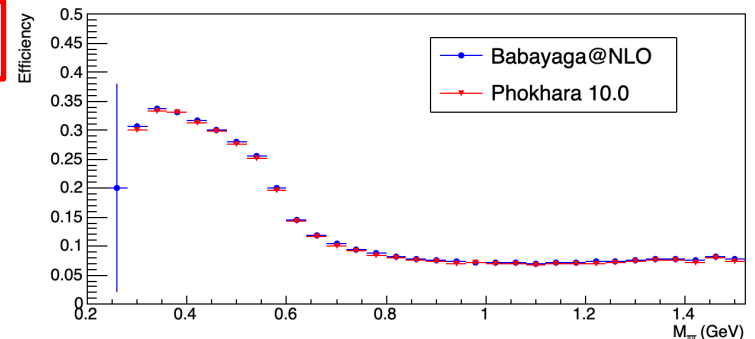
The imposition of the kinematic fit has a significant impact!

- Discrepancies in the 3% to 4% level close to rho, larger than expected
- Possible differences when approaching the threshold

Photon Corrections: 1C Kin. Fit



$\chi^2 < 10$

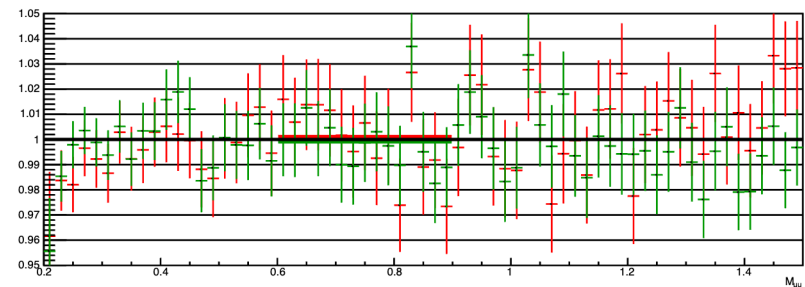
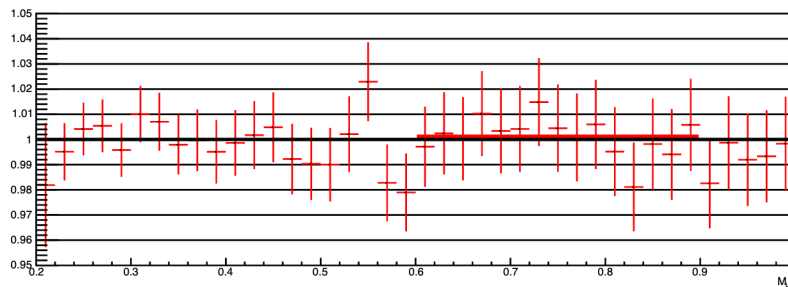
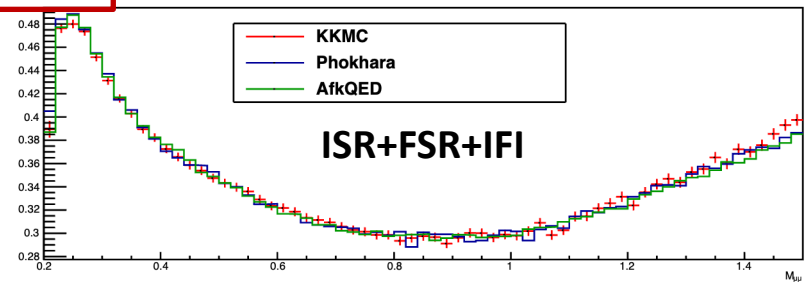
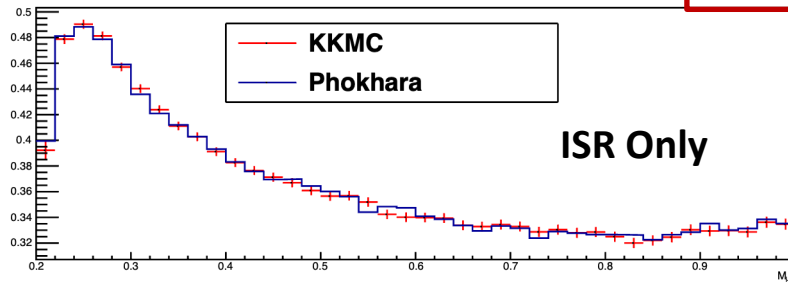


Large discrepancies are still observed!

- The tight selection does not allow to recover the missing contributions
- Different conclusion wrt to previous comparisons
 - ! Effect goes in the wrong direction, higher efficiency \rightarrow smaller FF !

Reducing Constraints: 1C Fit

$$20 < \theta_\gamma < 160$$



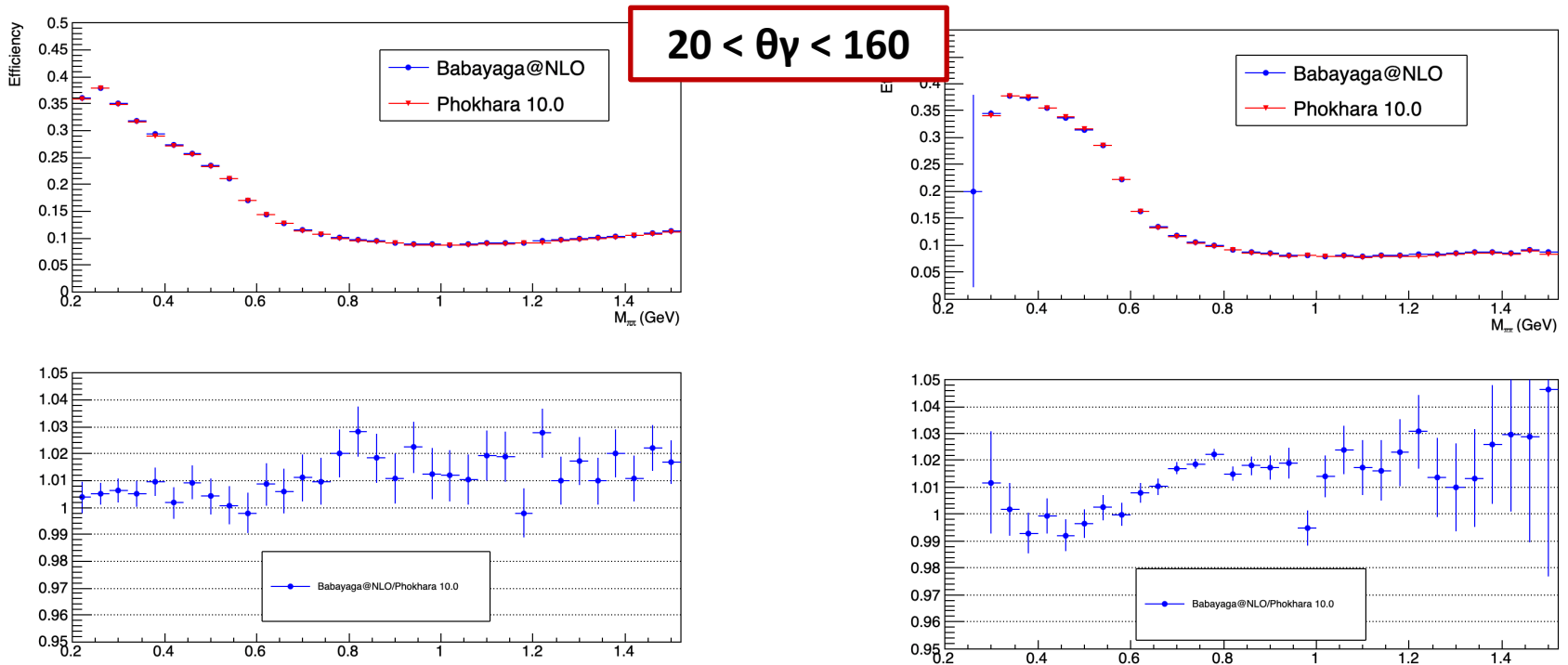
Reduce the sensitivity to additional radiation (new measurement)

Looser χ^2 cut with respect to photon efficiency (40 vs 10)

➤ **Good agreement** between Phokhara, KKMC, and AfkQED (at **0.1%**)

Allow for cross check and determination of eventual bias of published result

Reducing Constraints: 1C Fit



Different picture than estimated before!

- Differences at the 1-2% level
- Similar behavior for muons and pions, but unclear close to threshold
- Smaller than 4C fit, still a very useful cross checks

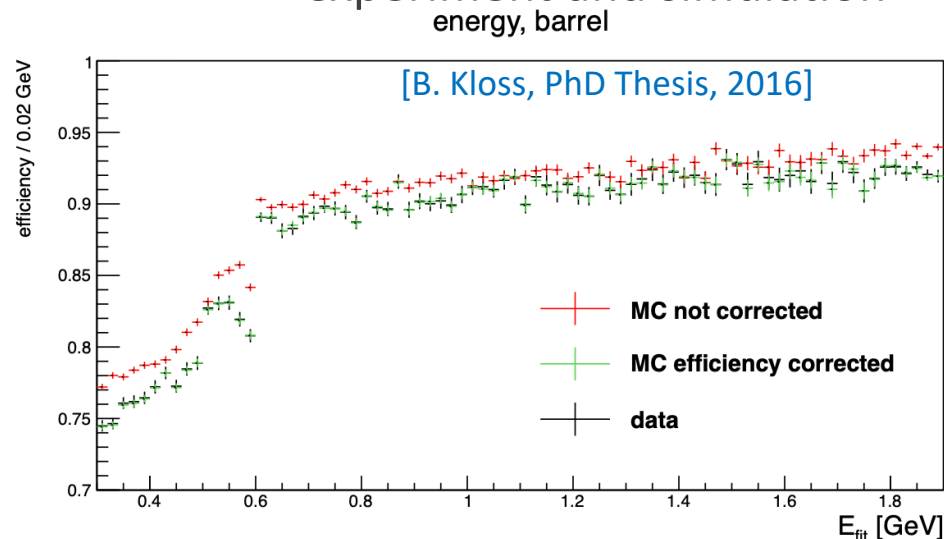
Summary

- **First measurement** of the pion FF by **BESIII** published in 2015
- Claimed **accuracy: 0.9%**
- Impact of Phokhara limitations on the result seems to be much larger than expected **O(3%)** introduced by the **4C Kinematic Fit**
- **New analysis** on going
 - **First result** based on about **6 fb⁻¹** of data **aiming at 0.7% accuracy** (soon)
 - **Final measurement** based on **20 fb⁻¹** **with 0.5% precision** (normalization to muons)
- Looser constraint on kinematics (**1C Kinematic Fit**)
- Still, **we observe** differences at the 1-2% level between NLO and NLO+PS!
- **Looking forward to further MC generator advancements:** Phokhara (+YFS or @NNLO), ...

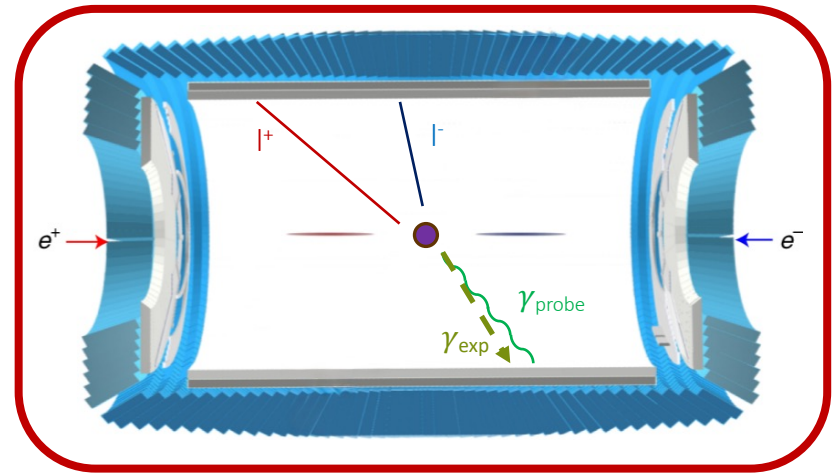
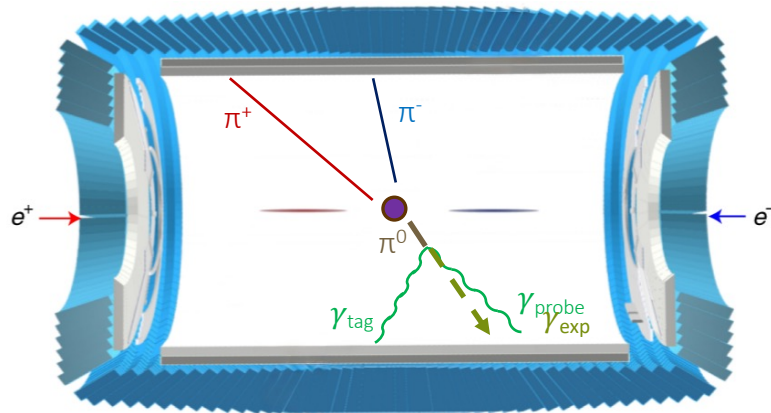
Radiative Corrections and Photon Detection Efficiency

Why do we need corrections?

- **MC** simulations are **critical for precision** measurements:
 - Efficiency determination, background subtraction, radiative corrections, ...
 - Geant 4 is an excellent tool for particle interaction with matter, but
 - Small gaps/displacement not foreseen in detector design, electronics, aging, ...
- **Need to verify and (eventually) account for differences** between real experiment and simulation



How to compute?



A **tag and probe strategy** is applied:

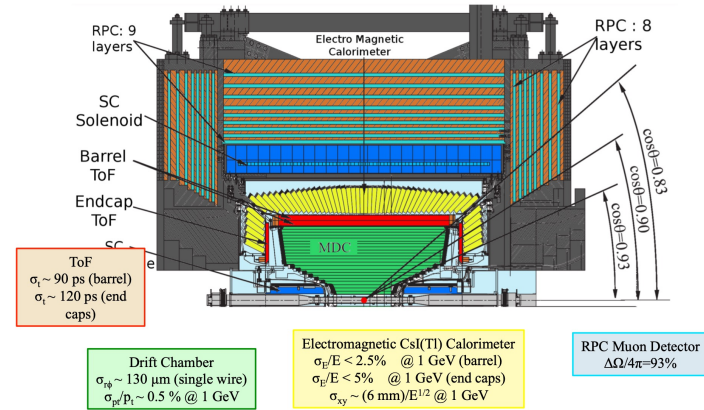
- Select processes for which the production of a photon can be tagged
- Use available information to **predict the photon** energy and direction
- **Look in the detector** for signals compatible with the predicted photon
- **Compare** how often the photon is detected in **data and MC** (efficiency)

In our work, we use $e^+ e^- \rightarrow l^+ l^- \gamma$ events

Selection and Generators

Event Selection:

- Exactly **2 charged particles** from the IP
- Not back-to-back (angle $< 170^\circ$)
- Both **identified as muons (EMC&MUC) or electrons (EMC)**
- **Kinematic fit (1C)** in the assumption of $ee\gamma/\mu\mu\gamma$
- Look for photon closing the kinematics (**4C fit**)



Generators and Radiative Corrections

Radiative process	$ee\gamma$	$\mu\mu\gamma$
PS	Babayaga 3.5	Babayaga 3.5
LO + PS	Babayaga@NLO	Babayaga@NLO
NLO	Babayaga@NLO (2026) ¹	Phokhara
NLO + PS	Babayaga@NLO (2026) ²	Babayaga@NLO (2026)

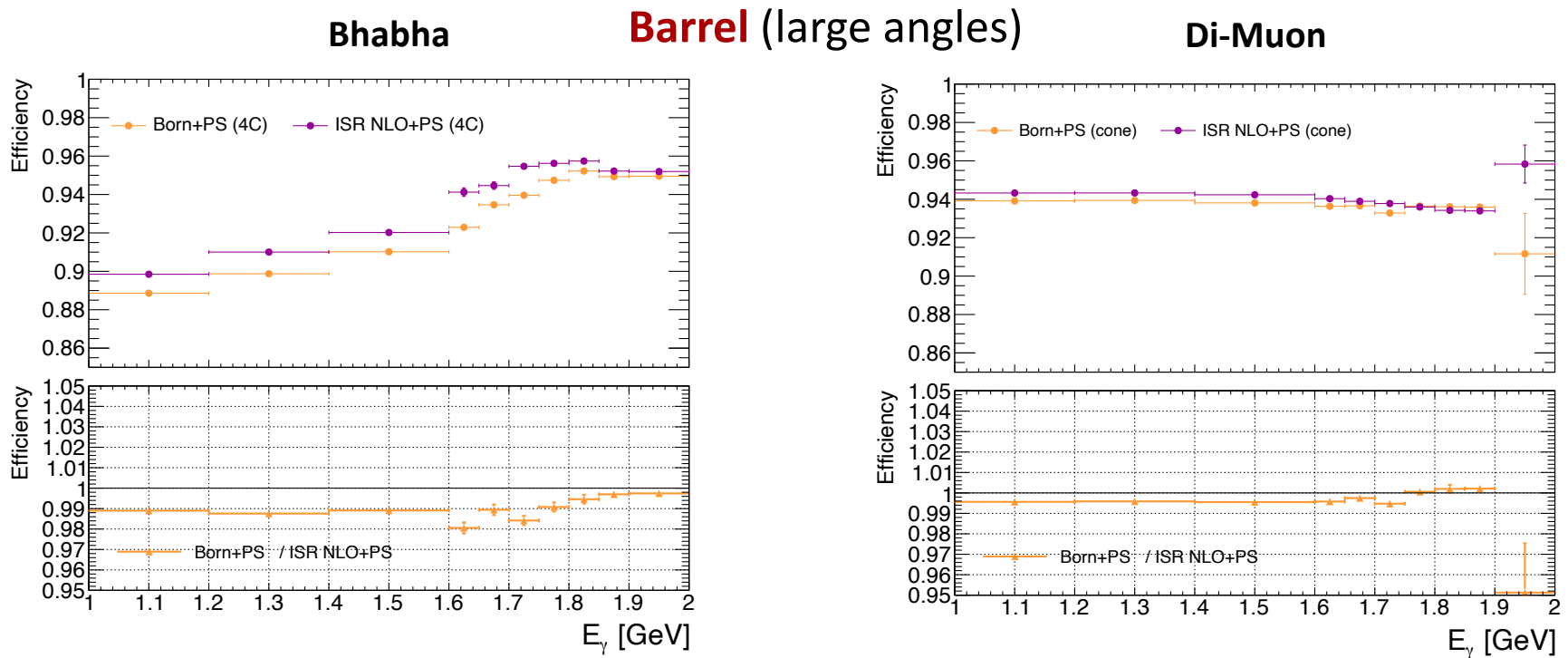
1: private version

2: pre-alpha test version, not validated!!!

Efficiency at different orders

Assume Babayaga@NLO (2026) perfectly reproduce data:

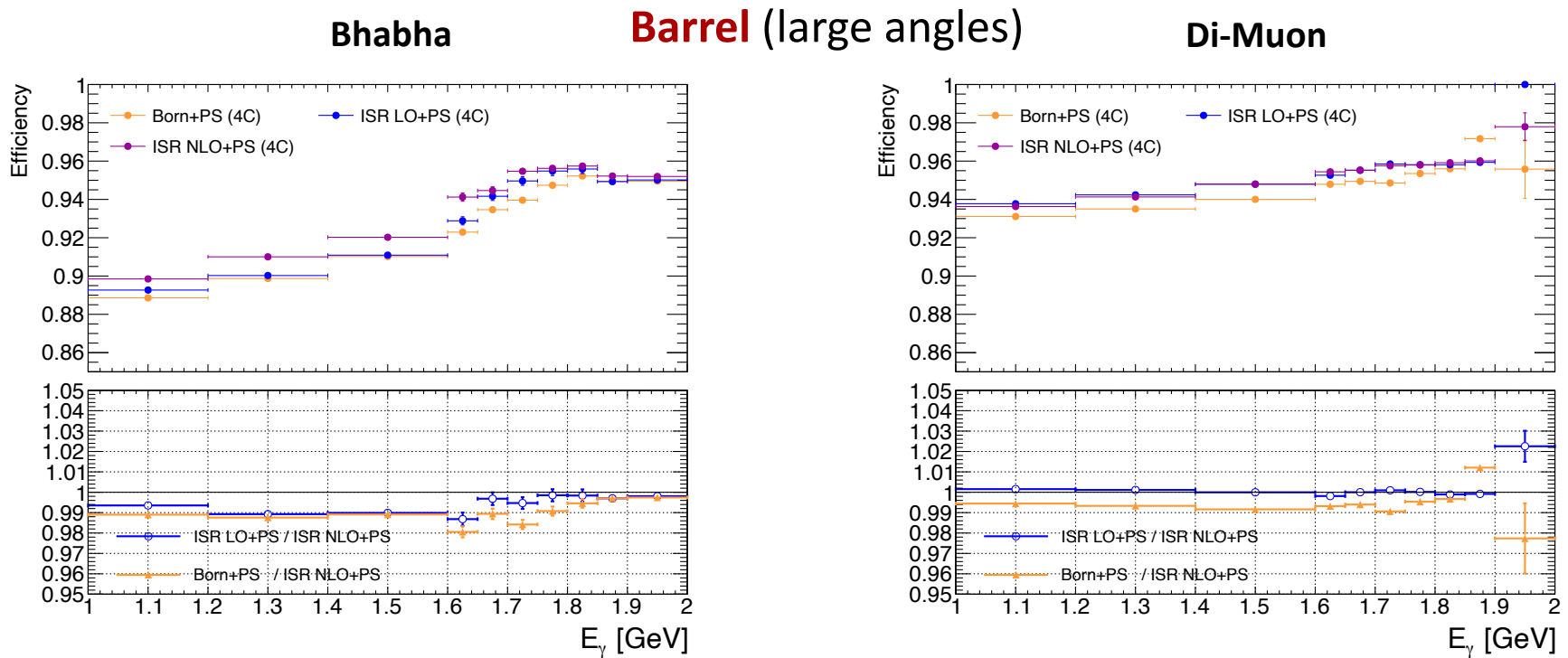
- Compare impact of missing corrections



Efficiency at different orders

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Efficiency at different orders

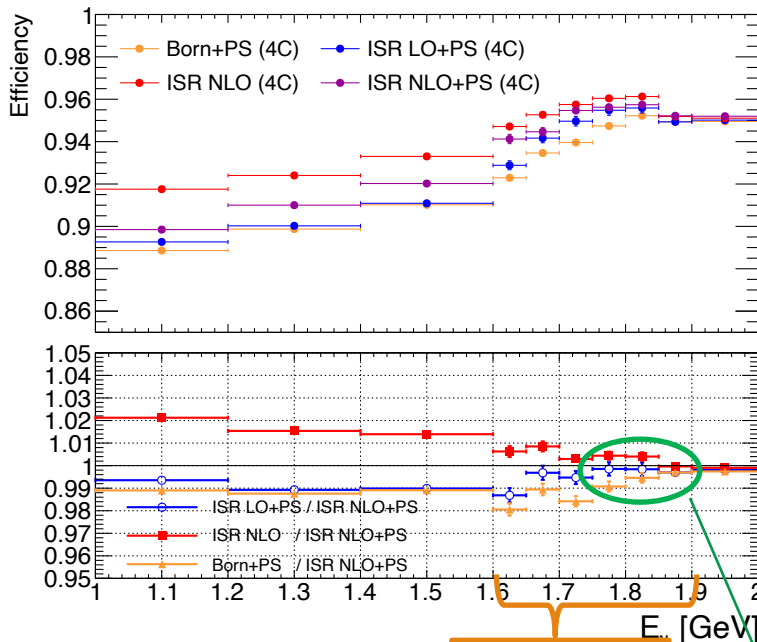
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Bhabha

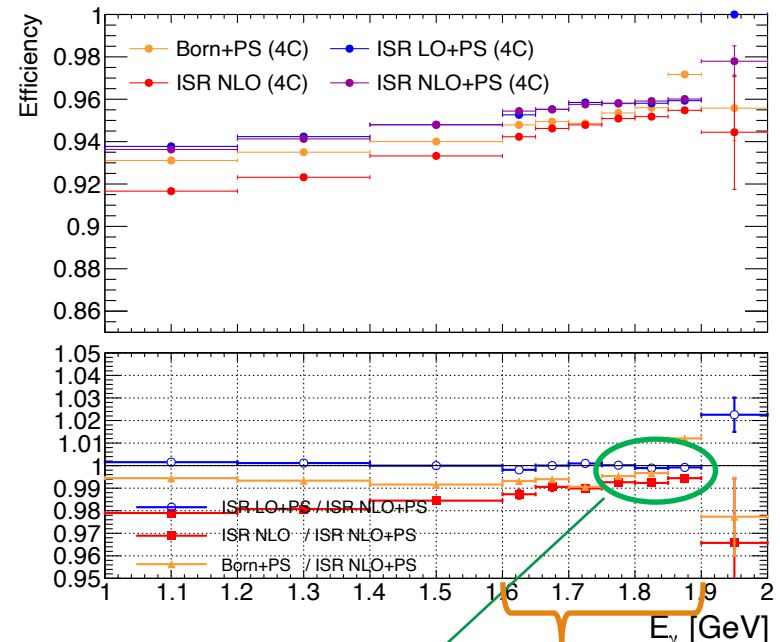
Barrel (large angles)

Di-Muon



$0.2 < \sqrt{s'} < 1.5$

Agreement at the per-mil level



$0.2 < \sqrt{s'} < 1.5$

Efficiency at different orders

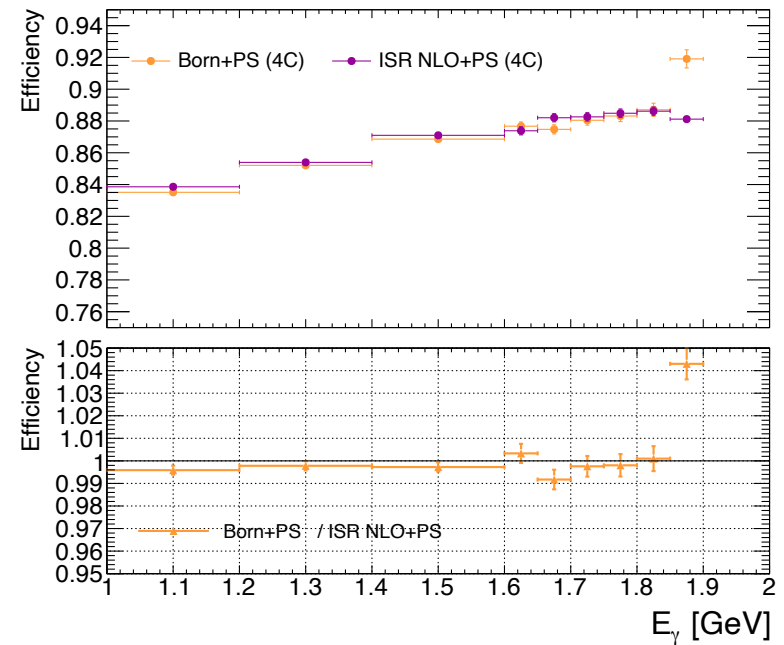
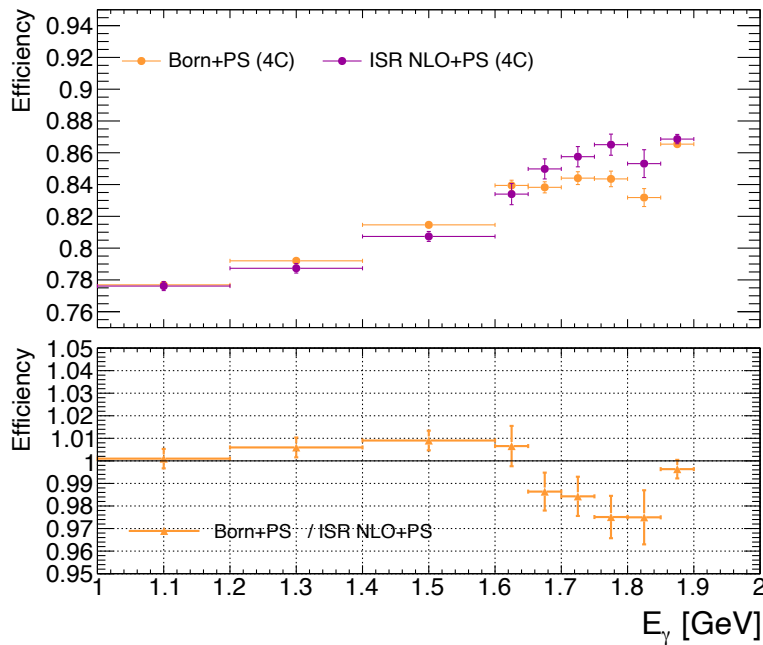
Assume Babayaga@NLO (2026) perfectly reproduce data:

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Bhabha

End Caps (smaller angles)

Di-Muon



Efficiency at different orders

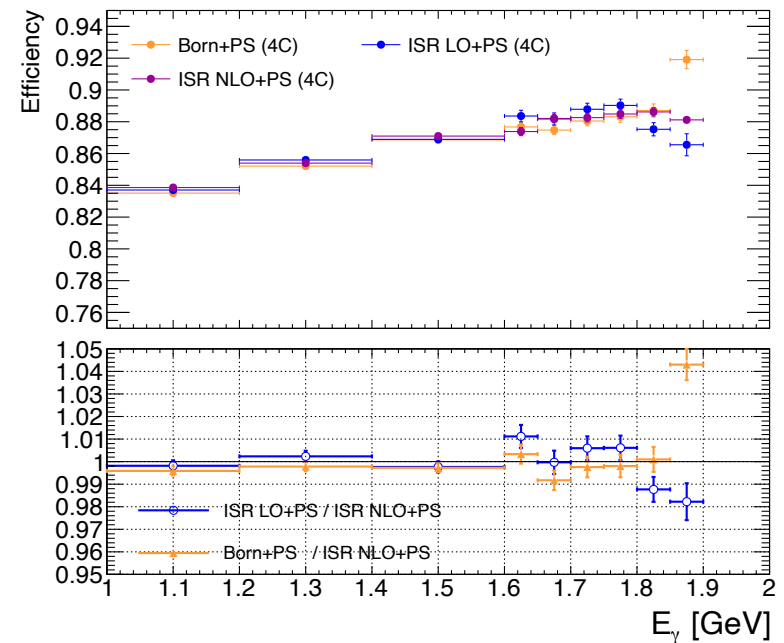
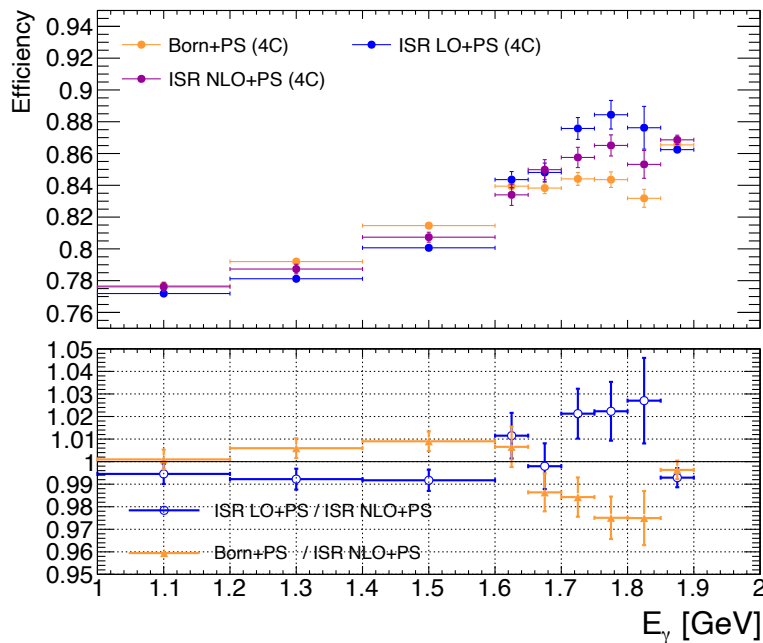
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End Caps (smaller angles)

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Efficiency at different orders

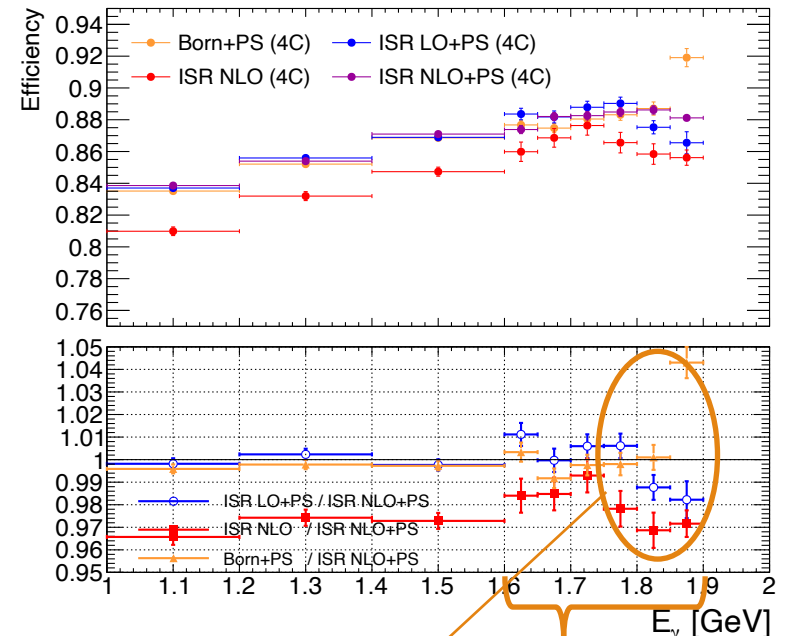
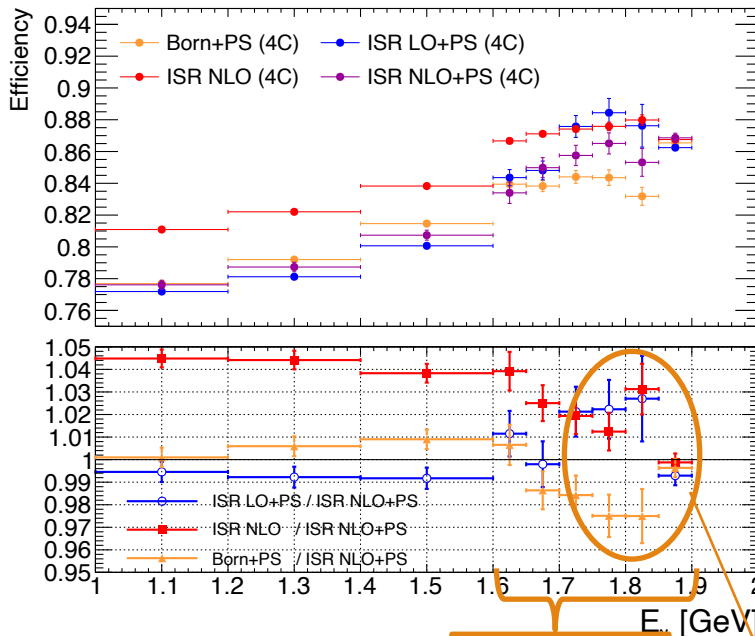
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- Compare impact of missing corrections

Bhabha

End Caps (smaller angles)

Di-Muon



0.2 < vs' < 1.5

0.2 < vs' < 1.5

Larger discrepancies

What do we learn?

- Photon detection efficiency is highly sensitive to radiative corrections!
- PS matching has a much larger impact than radiator function!
- Di-muon production seems to converge faster than Bhabha
- The barrel region is less sensitive to higher order effect
 - We look forward to the release of NLO+PS for radiative Bhabha
 - Cross check for radiative Bhabha scattering beyond NLO would be extremely important

We thank the Babayaga team for the help and availability!