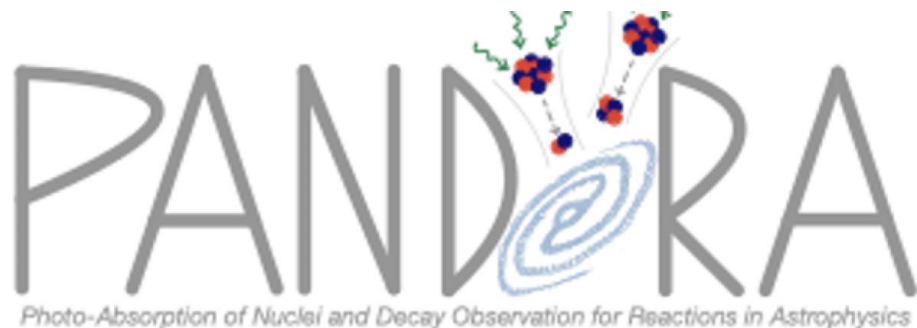


The measurement of gamma-decay in photo-nuclear reaction on ^{12}C 、 ^{13}C for the PANDORA project

The University of Osaka
Yumaro Suzuki



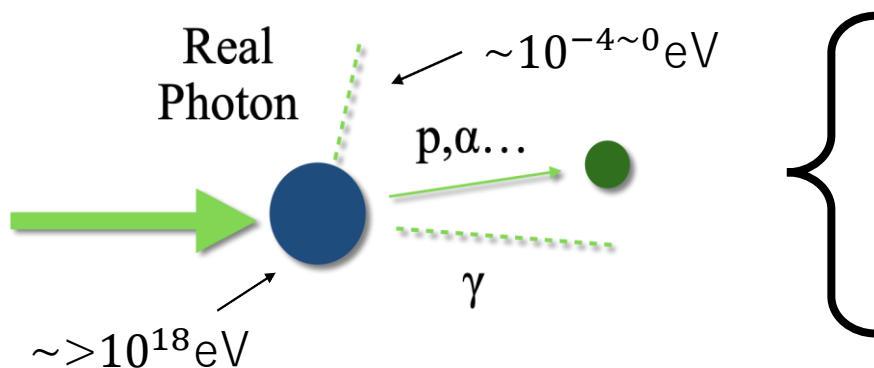
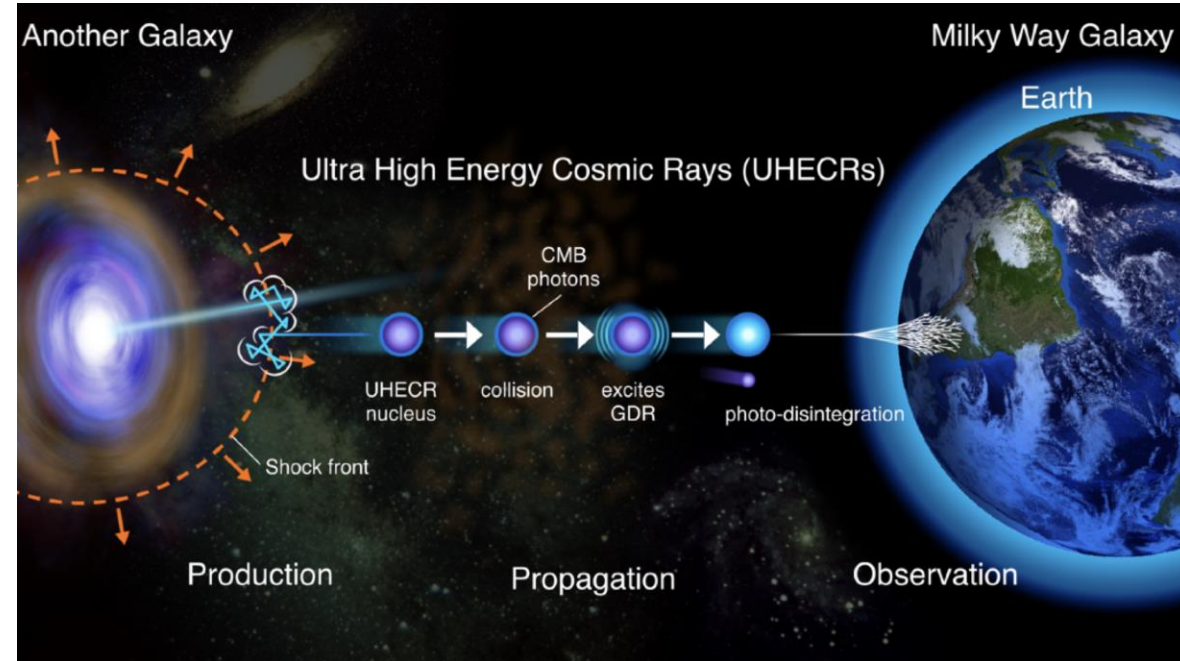
PANDORA project

Research on photonuclear reaction of light nuclei to understand Ultra-high energy cosmic rays.

➔ PANDORA project

Iso-Vector Giant Dipole Resonance (IVGDR) makes an important role in photonuclear reaction of UHECRs.

➔ Nuclei which is IVGDR decays immediately emitting particles and gamma-rays.



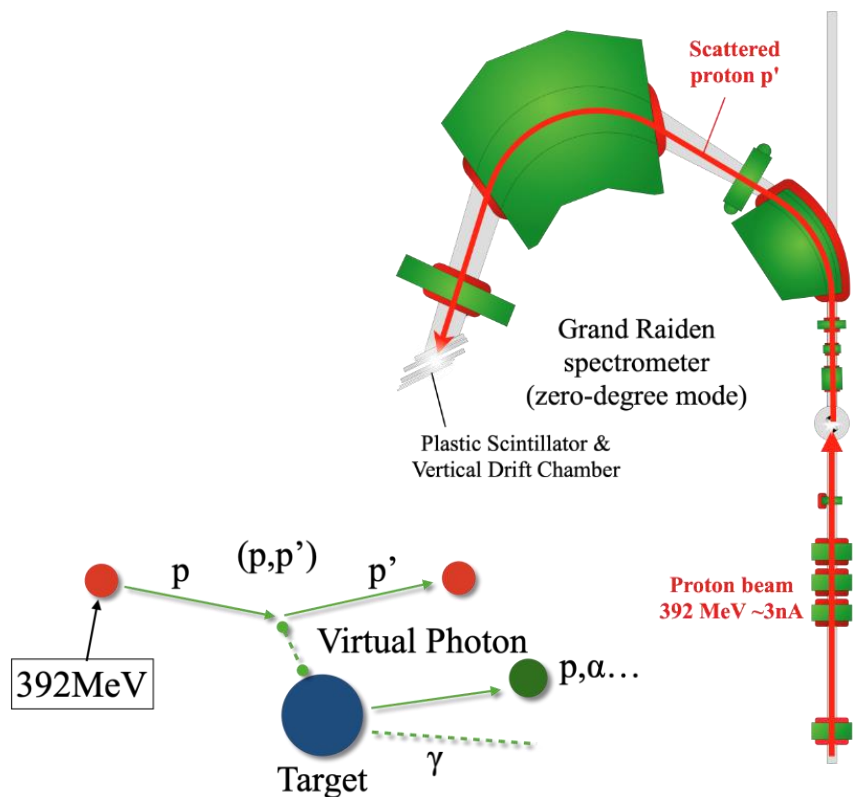
Particle decay measurement (Si detector)

Gamma decay measurement (LaBr₃ detector)

Gamma decay measurement

- Grand Raiden spectrometer (GR).

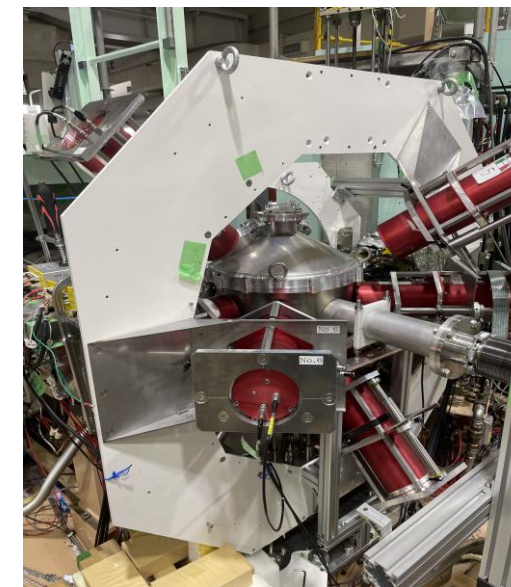
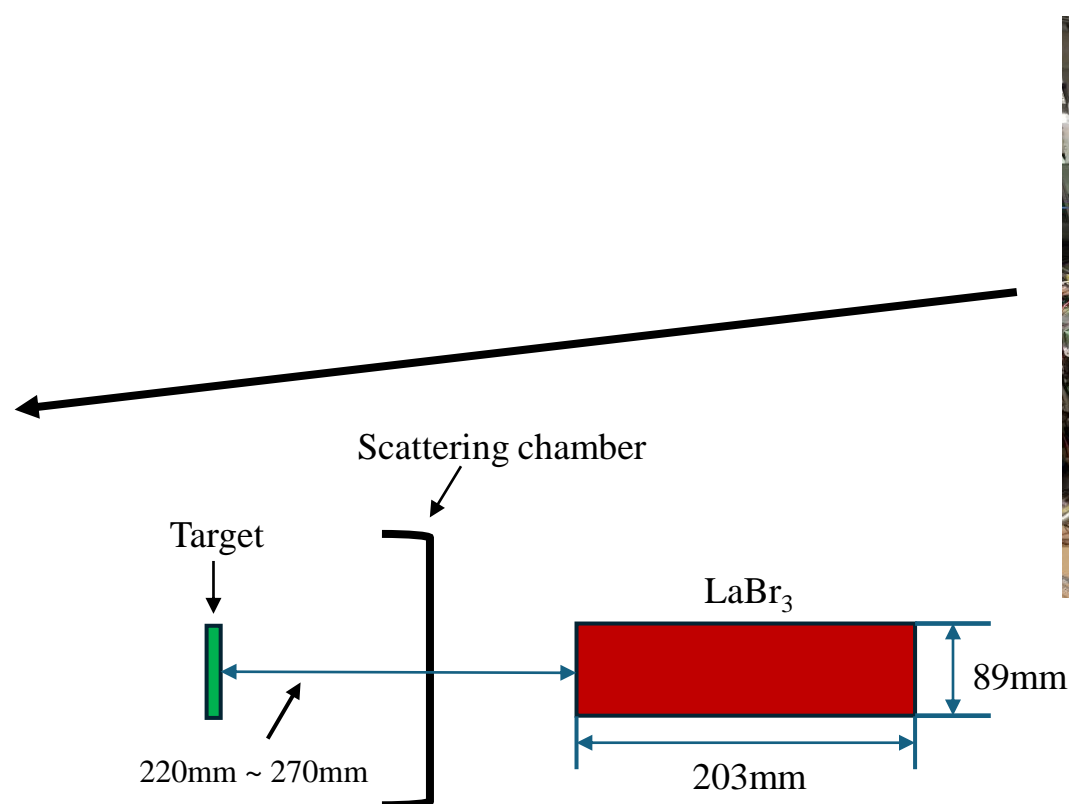
GR is a spectrometer that uses a magnetic field to separate the protons scattered after irradiation of a target by energy.



- LaBr₃ detector

LaBr₃ detector is an inorganic scintillator.

We placed eight LaBr₃ detectors around the scattering chamber to detect the γ - ray.



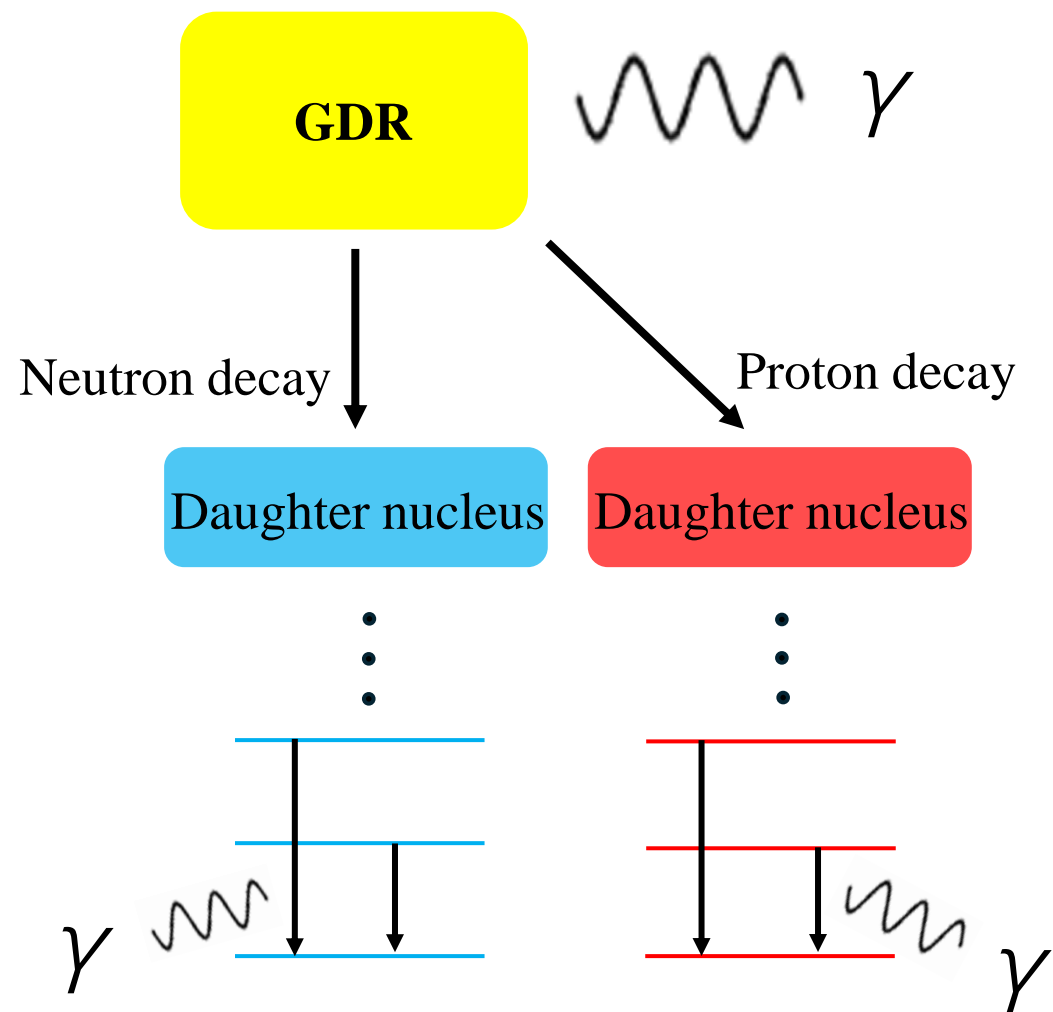
Gamma decay measurement

Gamma ray measurement

- Neutron decay branching ratios can be calculated
- Compare the proton decay result with the Si detectors
- Comprehensive study of gamma decay of Giant Resonance

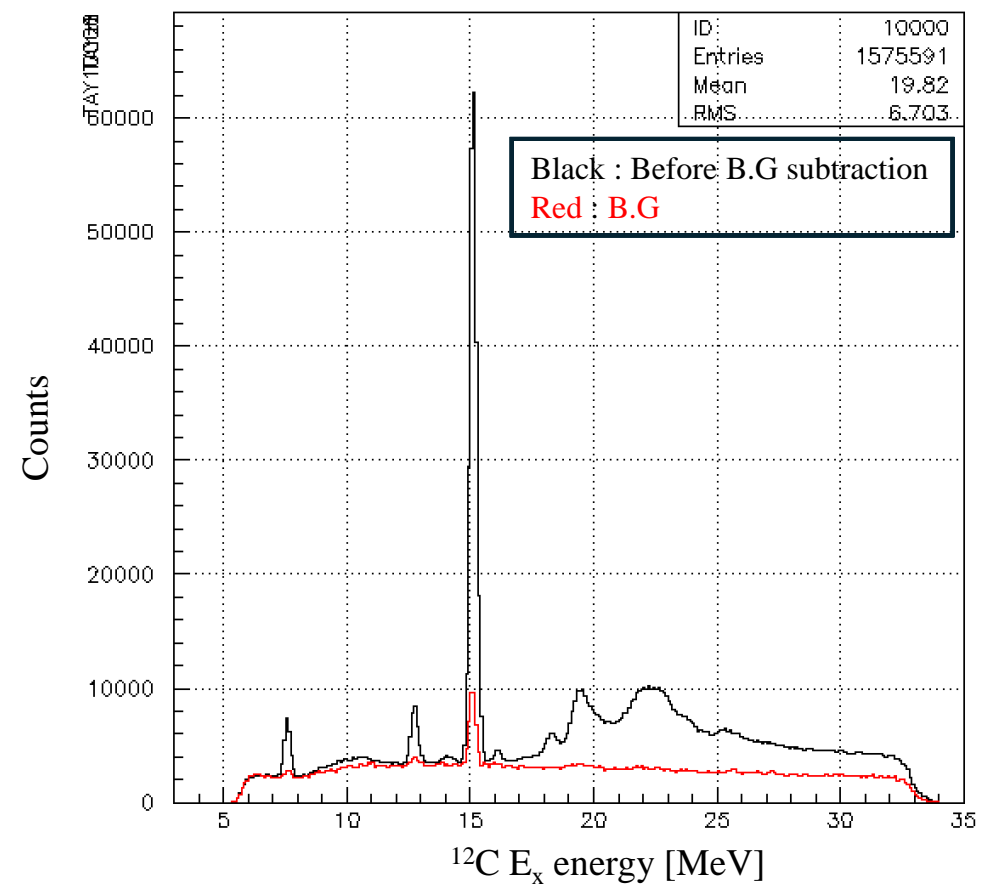
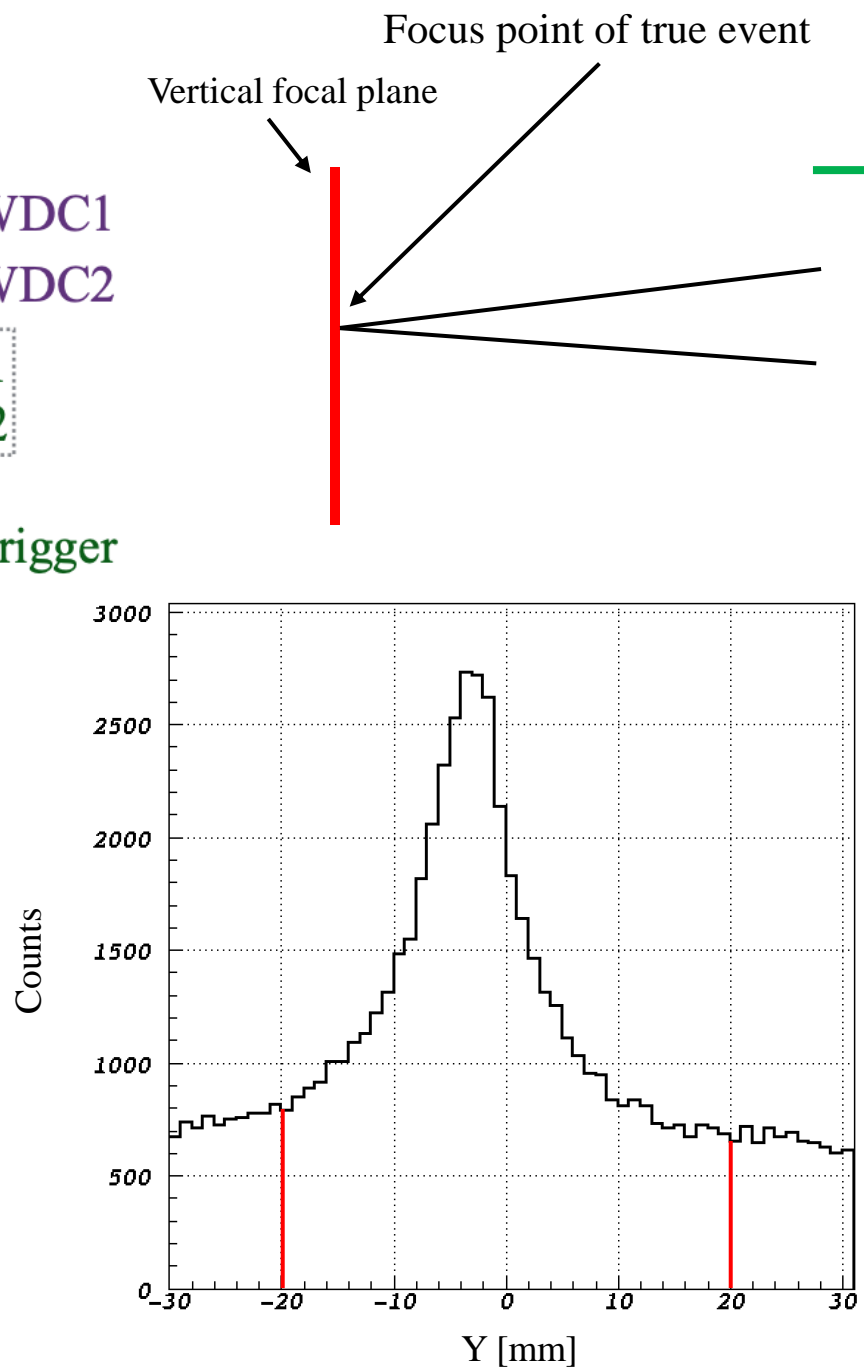
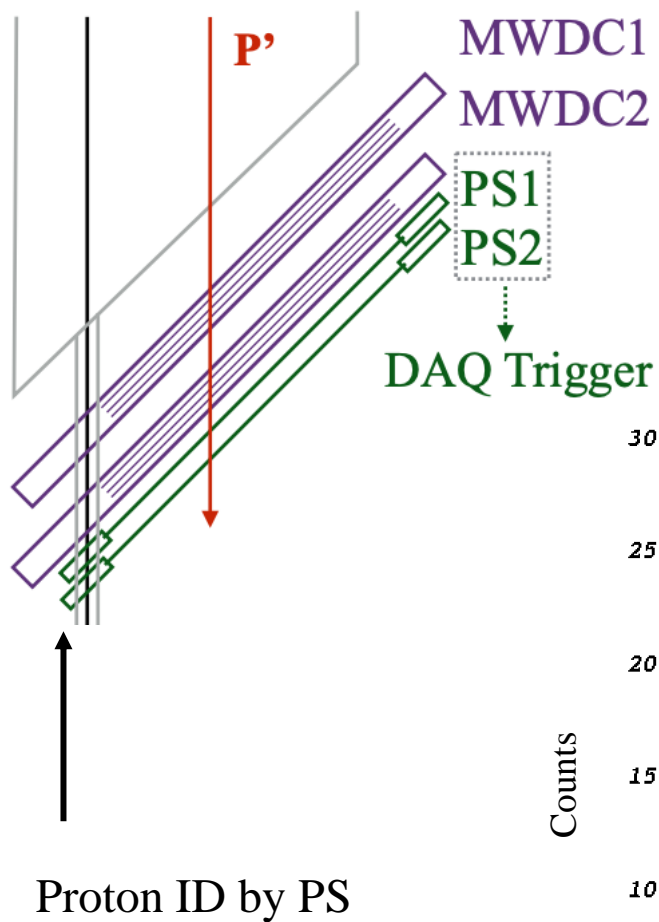
There are many gamma rays

- Direct decay from GDR
- Gamma decay after particle decays



Preliminary result

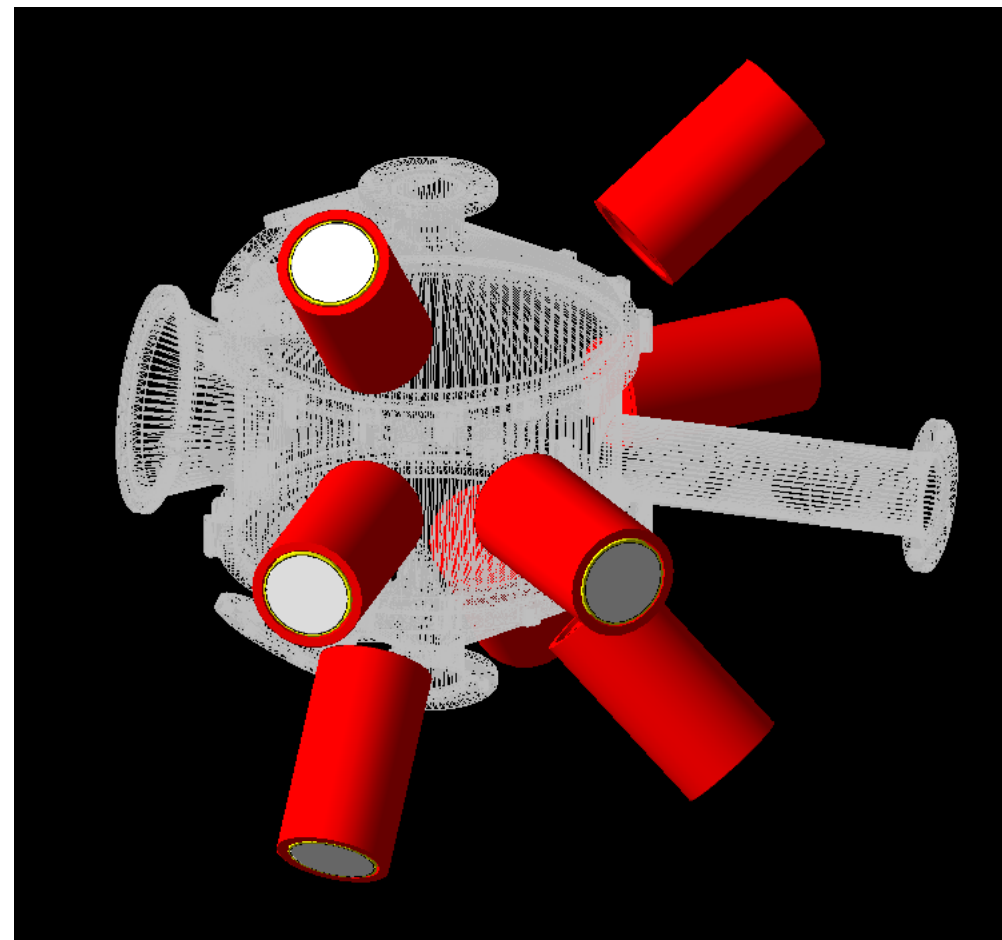
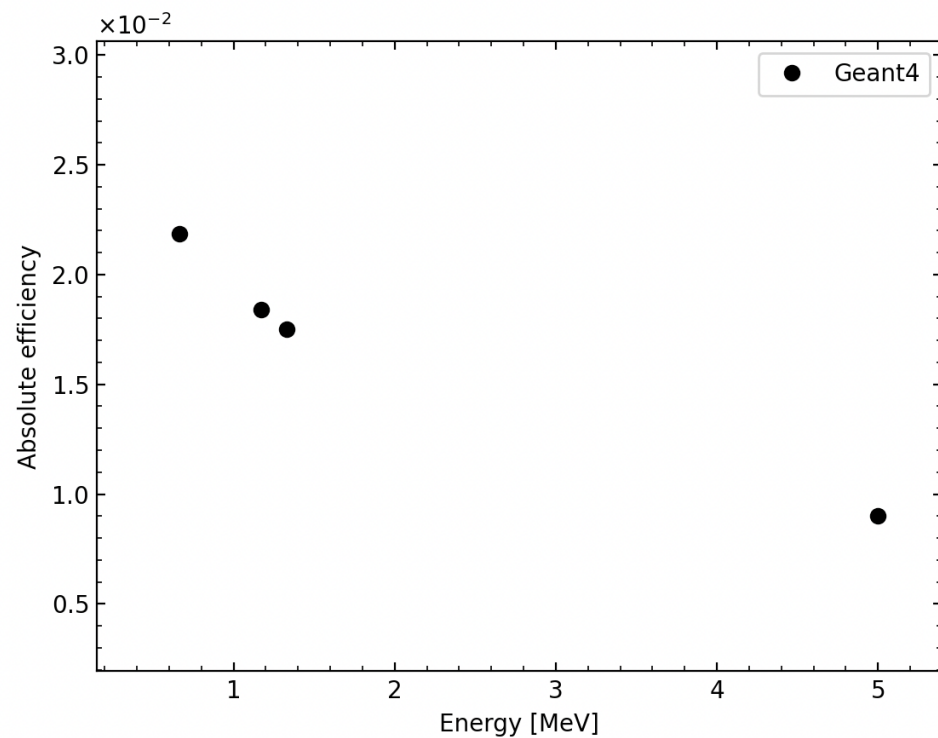
GR analysis



Branching ratio & Efficiency

Branching ratio

$$\frac{{}^A X(p, p' \gamma) \text{ counts}}{{}^A X(p, p') \text{ counts}} \times \frac{100}{\text{efficiency}} [\%]$$



Efficiency for full energy peak up to 5 MeV

^{12}C 0 degree

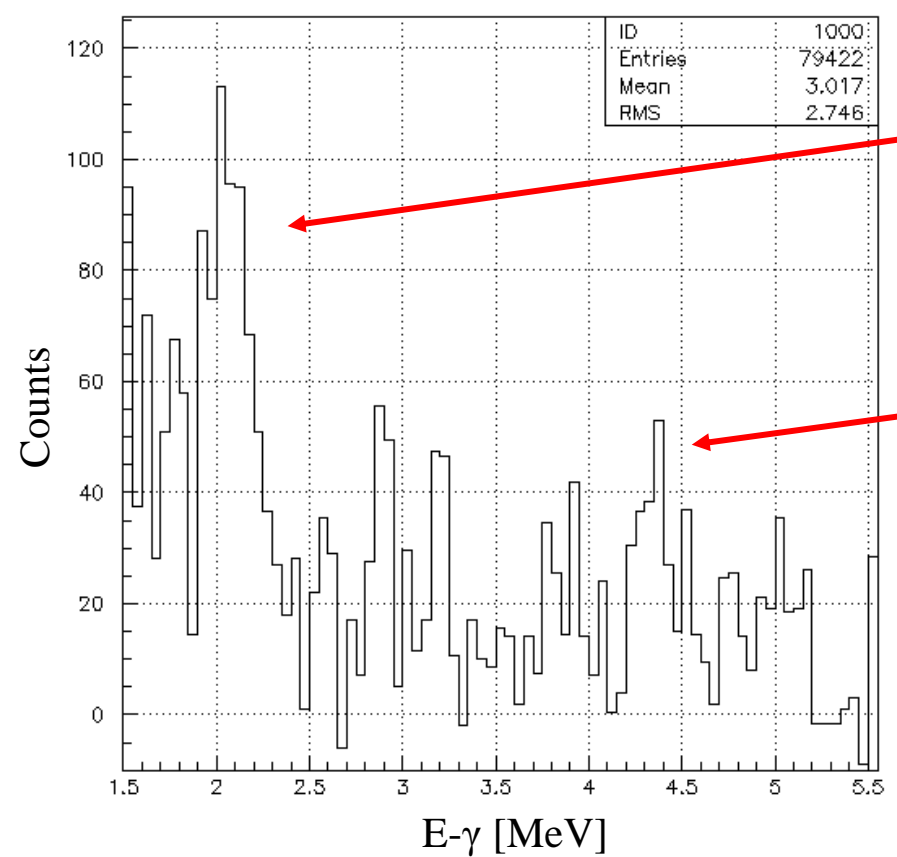
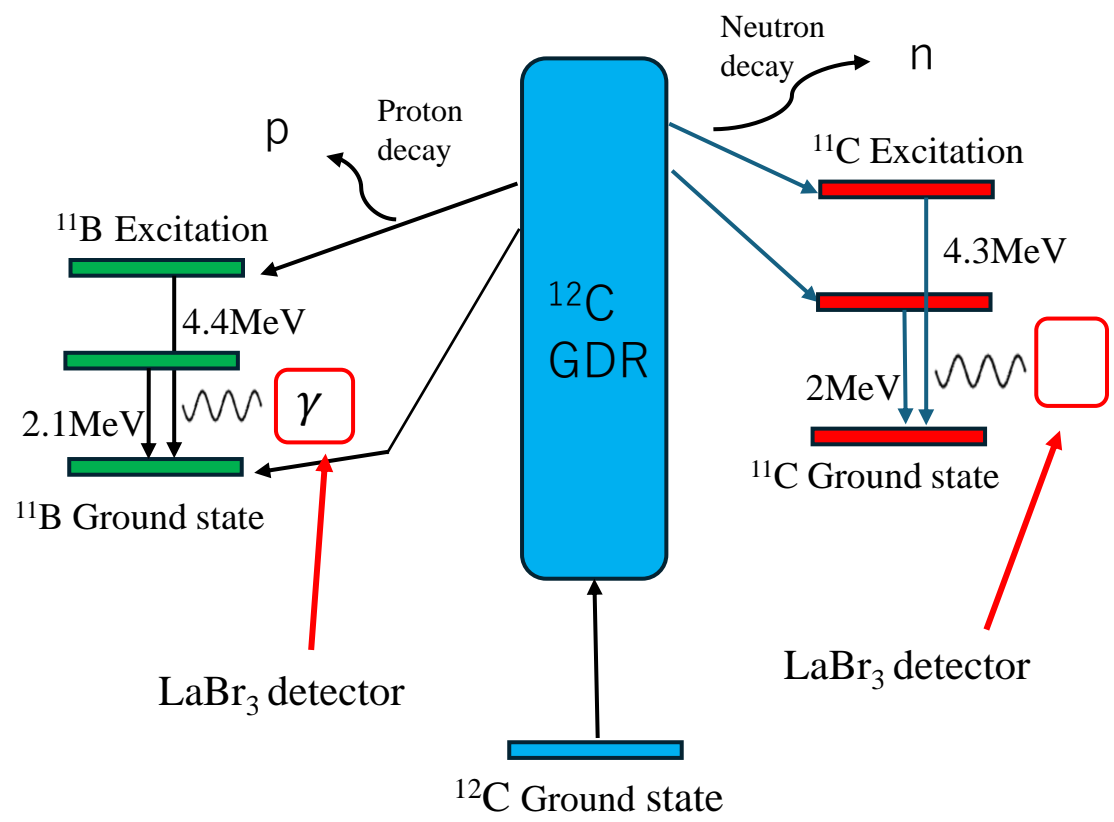
Branching ratio

$$\frac{^{12}\text{C}(p, p'\gamma) \text{ counts}}{^{12}\text{C}(p, p') \text{ counts}} \times \frac{100}{\text{efficiency}} [\%]$$

Branching ratios

→ $^{11}\text{C} + ^{11}\text{B} : E_1 = 11.0 [\%]$

→ $^{11}\text{C} + ^{11}\text{B} : E_2 = 7.2 [\%]$

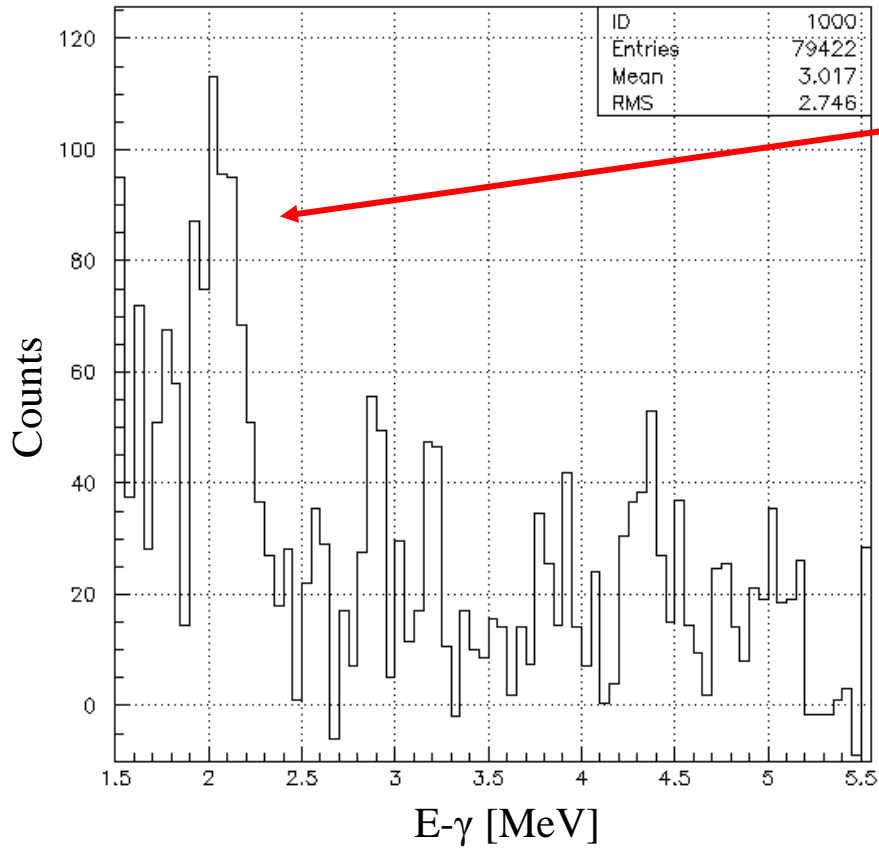


{ $^{11}\text{C} : E_1 = 2.00 [\text{MeV}]$
 $^{11}\text{B} : E_1 = 2.13 [\text{MeV}]$

{ $^{11}\text{C} : E_2 = 4.32 [\text{MeV}]$
 $^{11}\text{B} : E_2 = 4.45 [\text{MeV}]$

^{12}C 0 degree

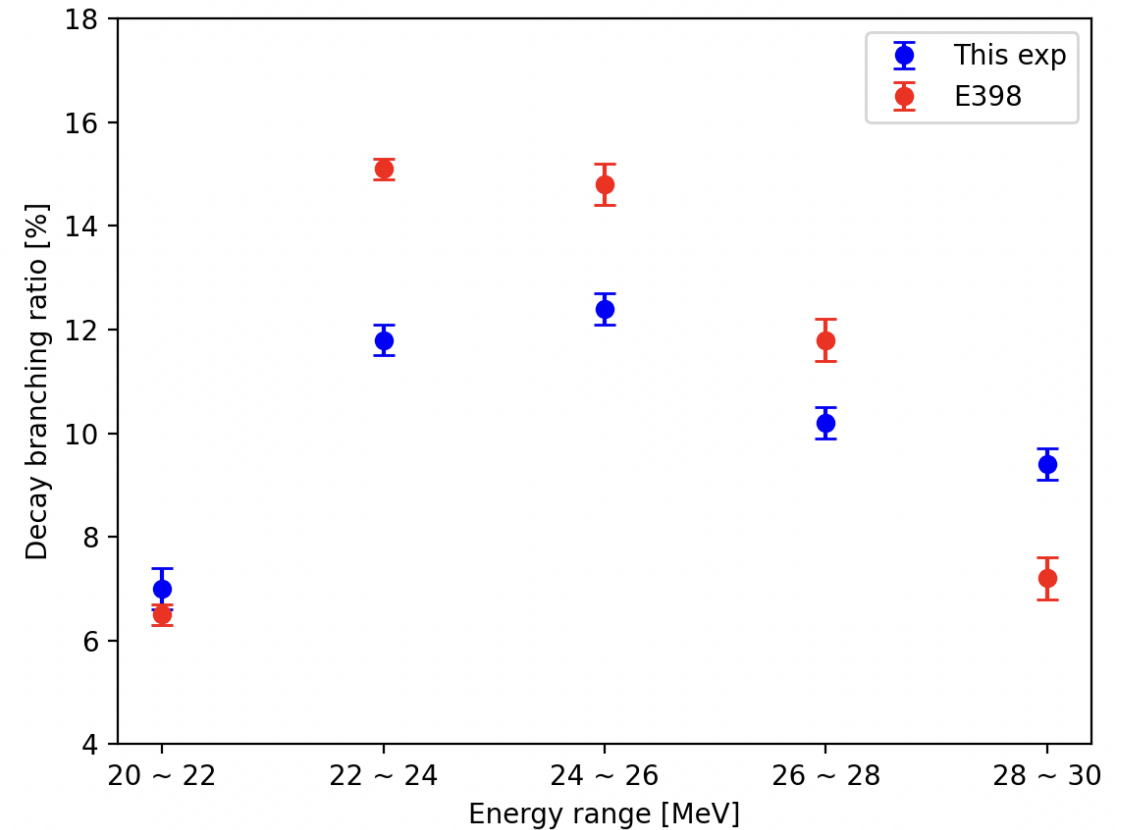
M. S. Reen^{1,*}, I. Ou¹, T. Sudo¹, D. Fukuda¹, T. Mori¹, A. Ali¹, Y. Koshio¹, M. Sakuda^{1,†}, A. Tamii^{2,‡}, N. Aoi², M. Yosoi², E. Ideguchi², T. Suzuki², T. Yamamoto², C. Iwamoto³, T. Kawabata⁴, S. Adachi⁵, M. Tsumura⁵, M. Murata⁵, T. Furuno⁵, H. Akimune⁶, T. Yano⁷, T. Suzuki⁸ and R. Dhir⁹



$^{11}\text{C} : E_1=2.00[\text{MeV}]$
 $^{11}\text{B} : E_1=2.13[\text{MeV}]$

E398

- 392 MeV 0deg (p,p')
- ^{12}C target
- Gamma-decay measurement from GDR
- NaI scintillator



¹³C 0 degree

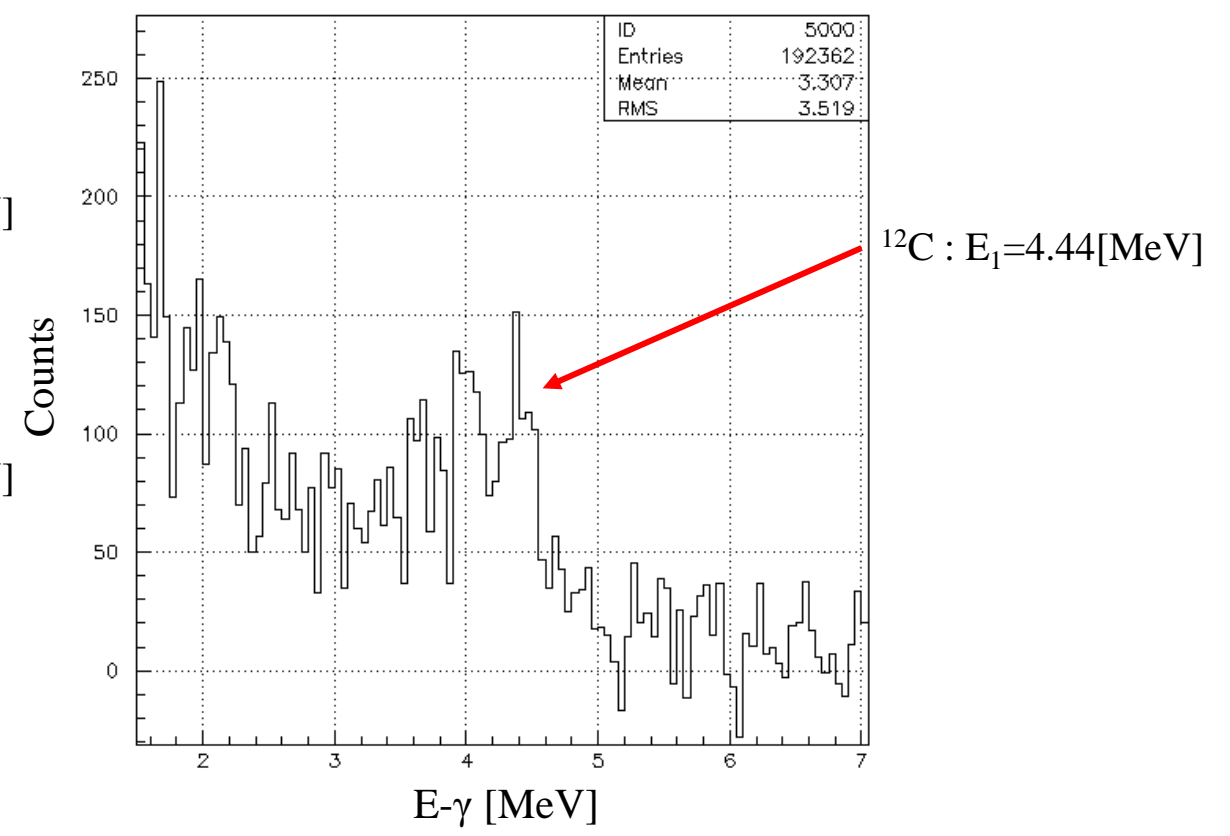
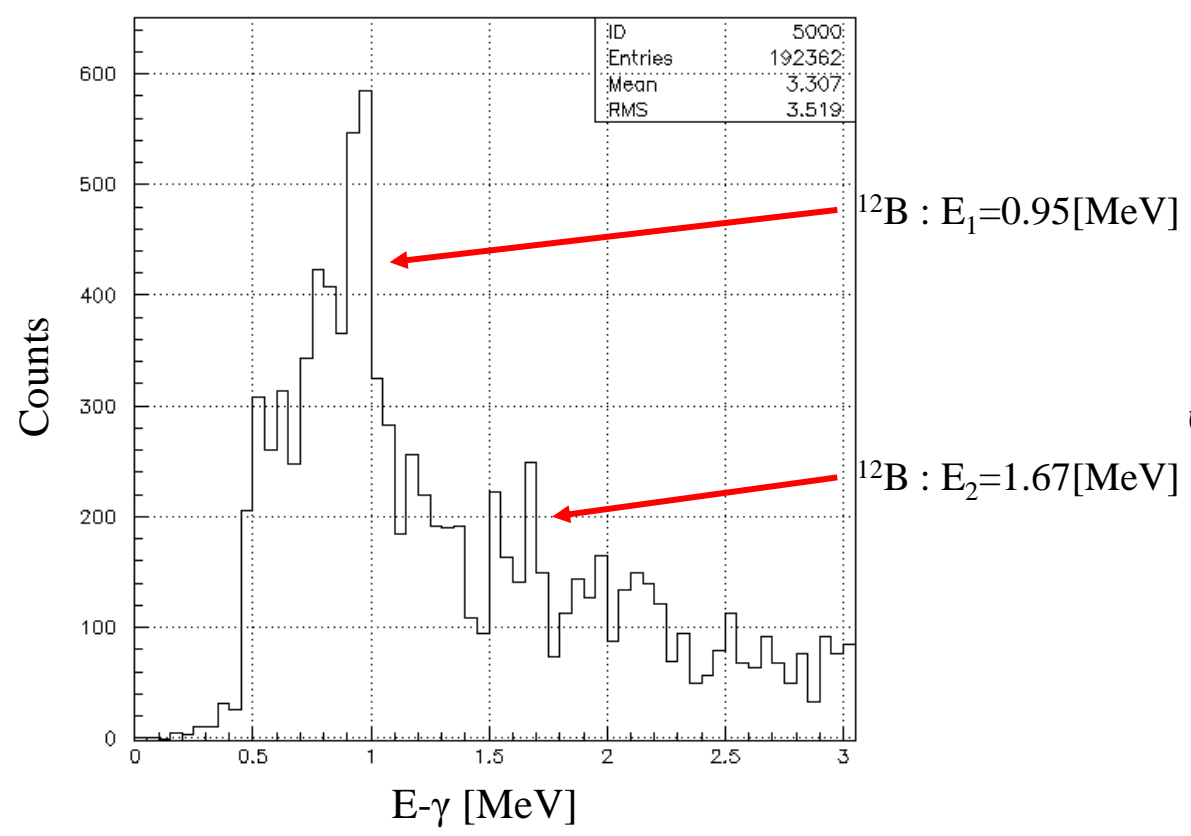
Branching ratio

$$\frac{{}^{13}\text{C}(p, p'\gamma) \text{ counts}}{{}^{13}\text{C}(p, p') \text{ counts}} \times \frac{100}{\text{efficiency}} [\%]$$

Branching ratios

→ ¹²C : E₁ = 7.6[%]

→ ¹²B : E₁ = 7.9[%] E₂ = 2.3[%]



^{13}C 0 degree

Photoneutron emission cross sections for ^{13}C

H. Utsunomiya^{1,2}, S. Goriely,³ M. Kimura,⁴ N. Shimizu⁵, Y. Utsuno,^{6,7} G. M. Tveten,^{8,9}
T. Renstrøm,^{8,9} T. Ari-izumi,^{1,10} and S. Miyamoto^{11,12}

$$\int \frac{d^2\sigma}{dEd\Omega} d\Omega = \frac{1}{E} \int \frac{dN}{d\Omega} d\Omega \cdot \sigma_{abs}$$

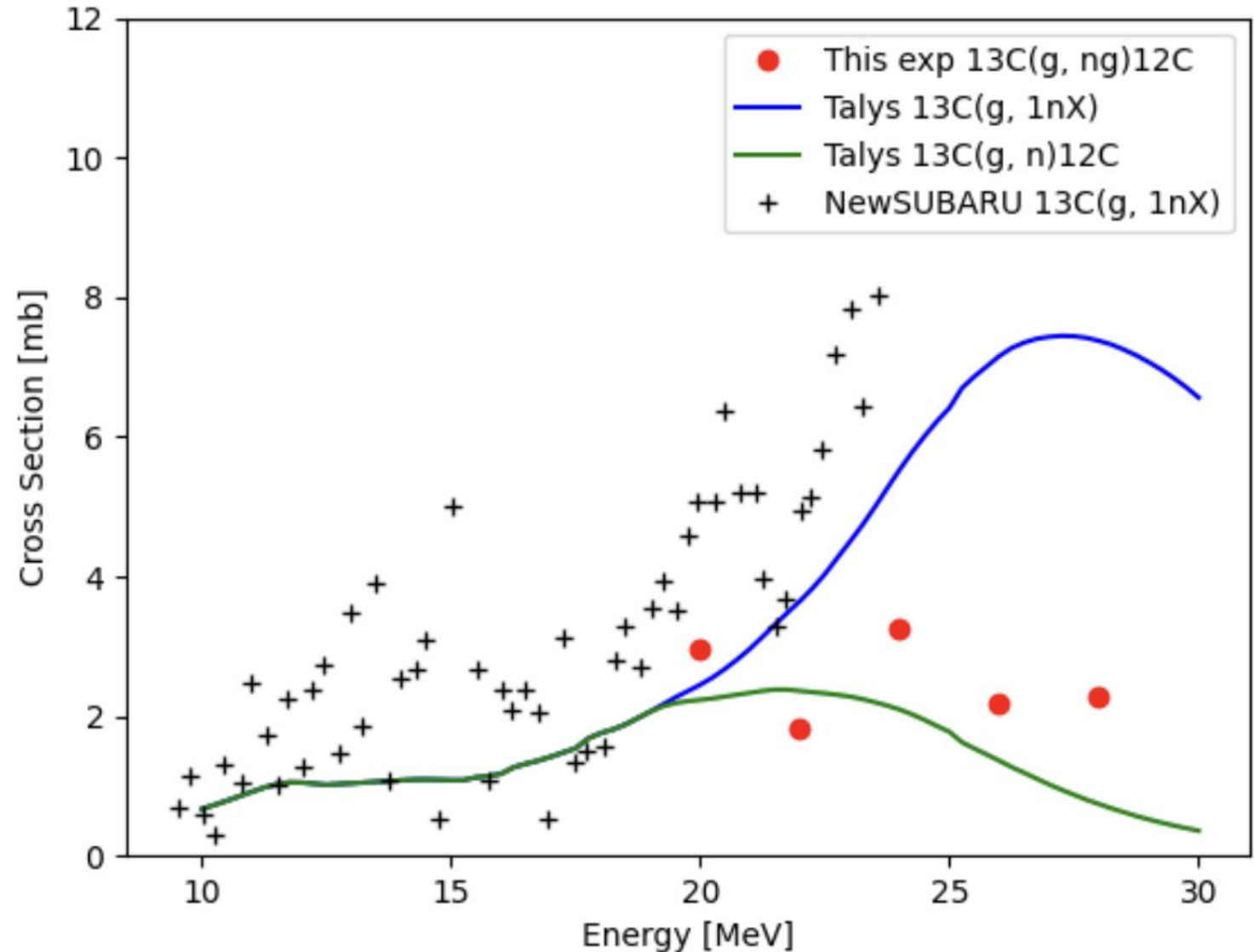
$$\sigma_{(\gamma,ng)} = \sigma_{abs} \cdot \text{Branch}(\gamma,ng)$$

$$S_{^{13}\text{C}}(\text{p}) = 17.53 \text{ MeV}$$

$$S_{^{13}\text{C}}(\text{n}) = 4.95 \text{ MeV}$$

$$S_{^{12}\text{C}}(\text{p}) = 15.60 \text{ MeV}$$

$$S_{^{12}\text{C}}(\text{n}) = 18.72 \text{ MeV}$$



Summary

Summary

- Expected gamma-ray peaks appear in the analysis using the LaBr_3 detectors.
 - Efficiency simulation on the Geant4 was done.
 - Branching ratios and cross sections are roughly calculated.
- Analysis on finite angle data and MDA is needed

Thank you

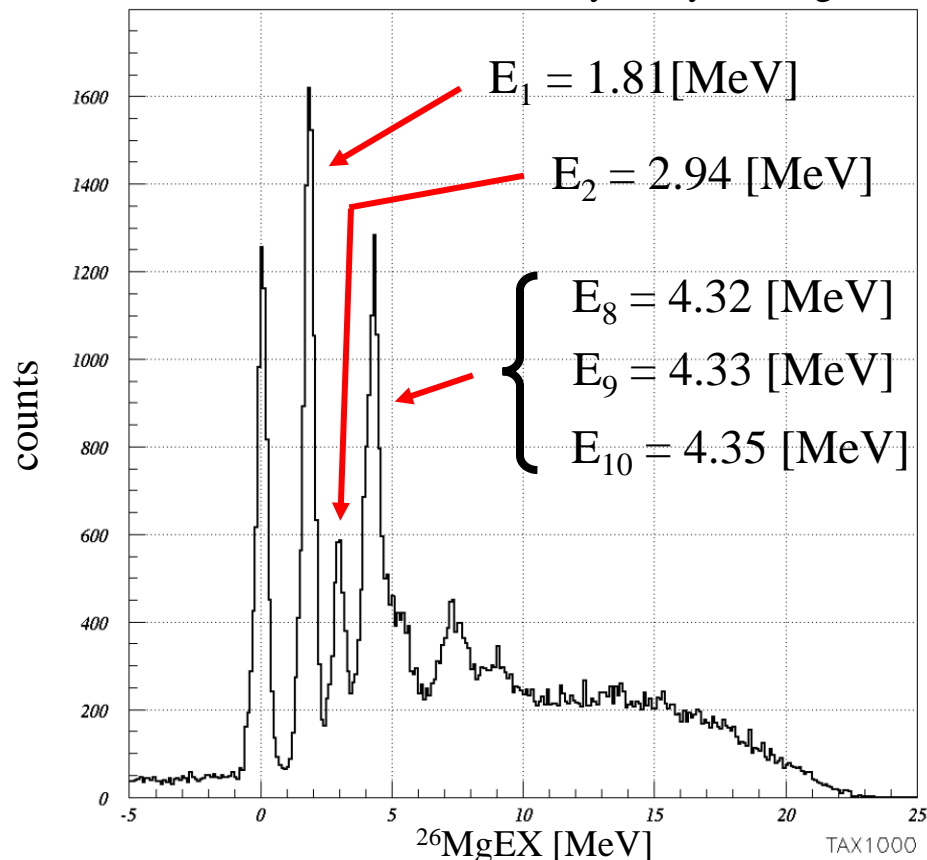
Presentation Backup

^{27}Al 0 degree

Branching ratio

$$\frac{{}^{12}C(p, p'\gamma) \text{ counts}}{{}^{12}C(p, p') \text{ counts}} \times \frac{100}{\text{efficiency}} [\%]$$

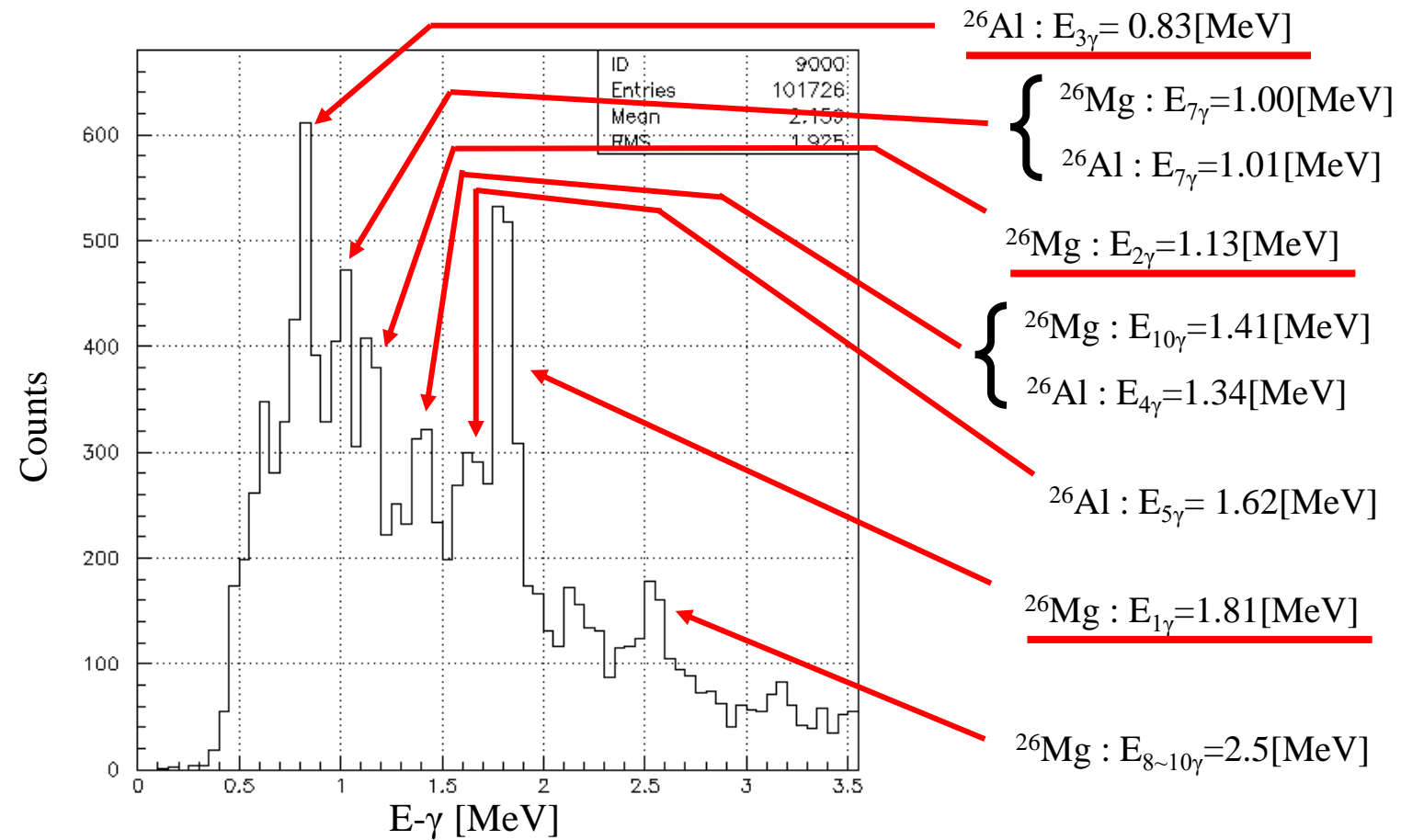
Analyzed by Y.Sasagawa



EXMGPQ Ex E - factor*Si 1FE for 27Al

Branching ratios

$$\begin{aligned} & \rightarrow {}^{26}\text{Al} : E_3 = 13.9[\%] \\ & {}^{26}\text{Mg} : E_1 = 22.8[\%] \quad E_2 = 8.4[\%] \end{aligned}$$



¹²C 4.5 degrees

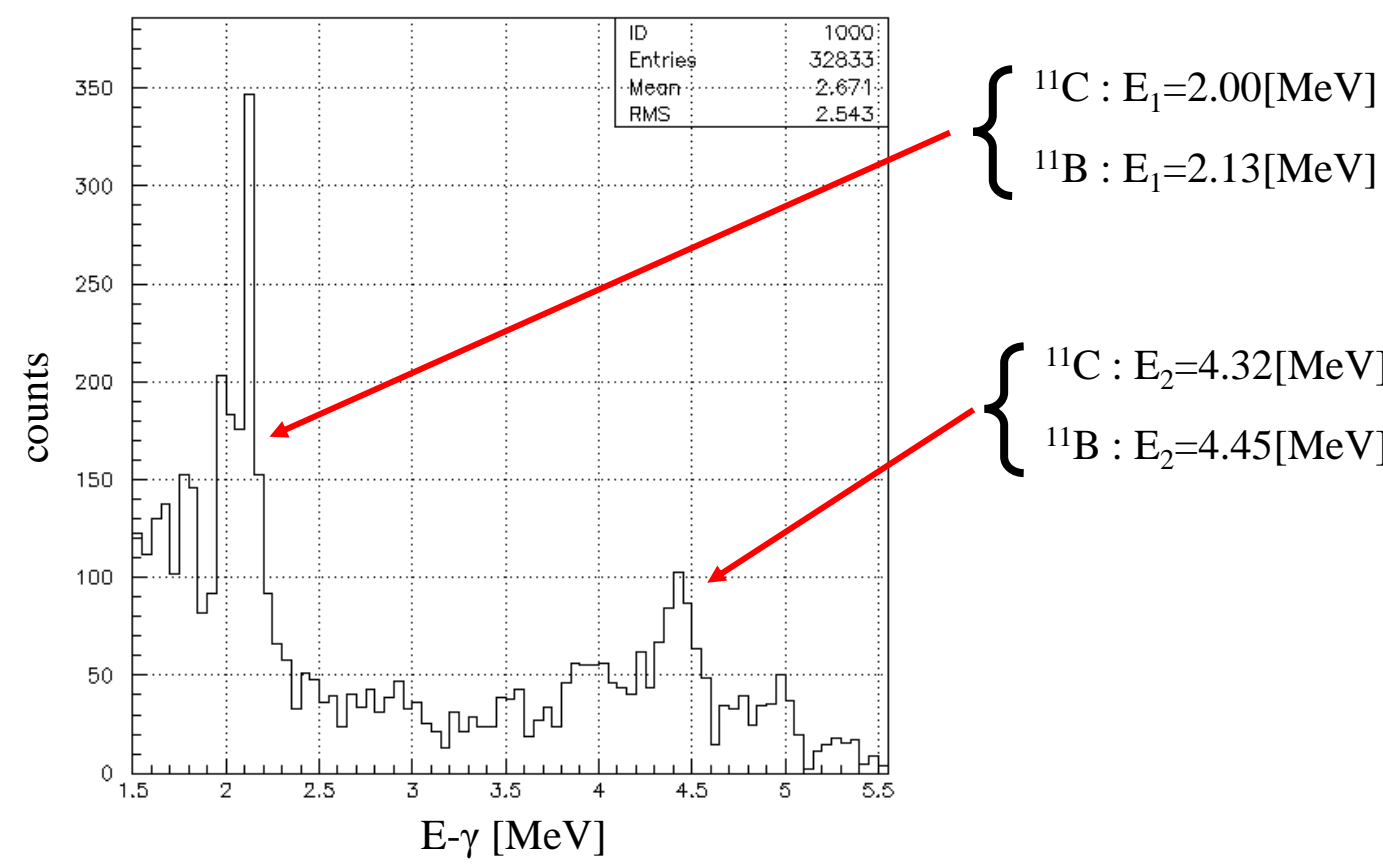
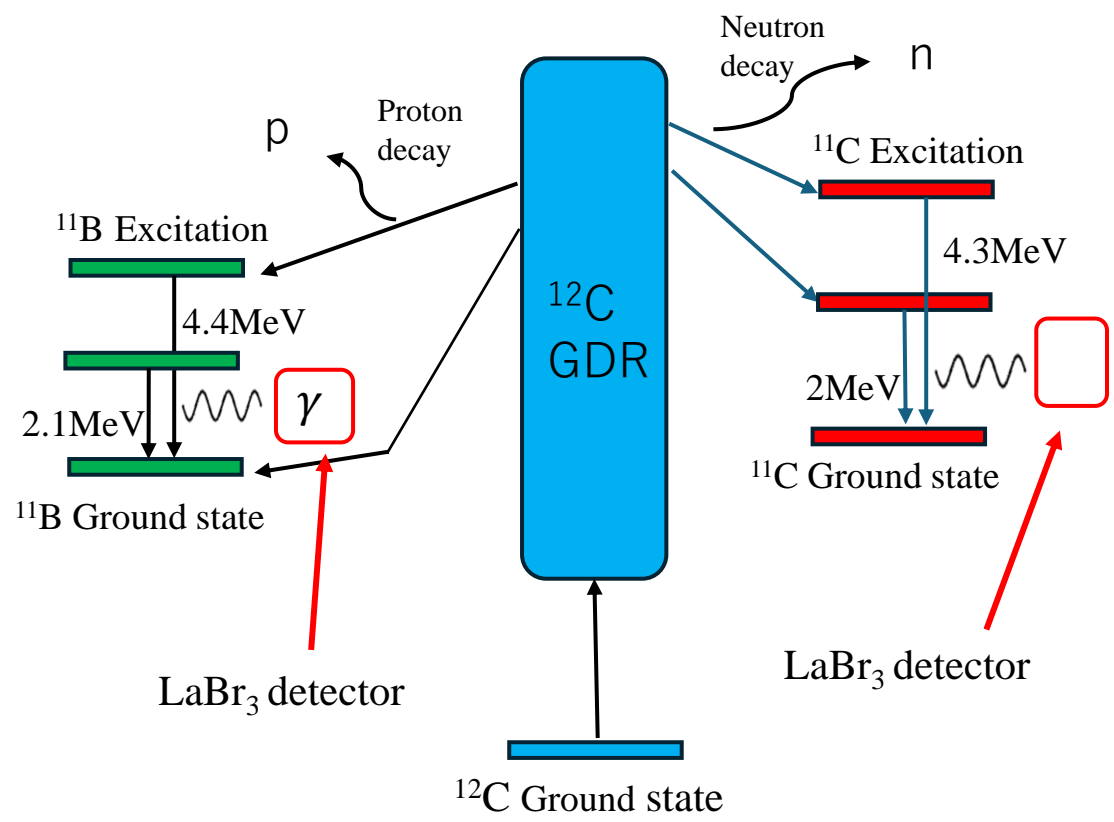
Branching ratio

$$\frac{{}^{12}\text{C}(p, p'\gamma) \text{ counts}}{{}^{12}\text{C}(p, p') \text{ counts}} \times \frac{100}{\text{efficiency}} [\%]$$

Branching ratios

→ ¹¹C + ¹¹B : E₁=15.7 [%]

→ ¹¹C + ¹¹B : E₂= 10.5[%]

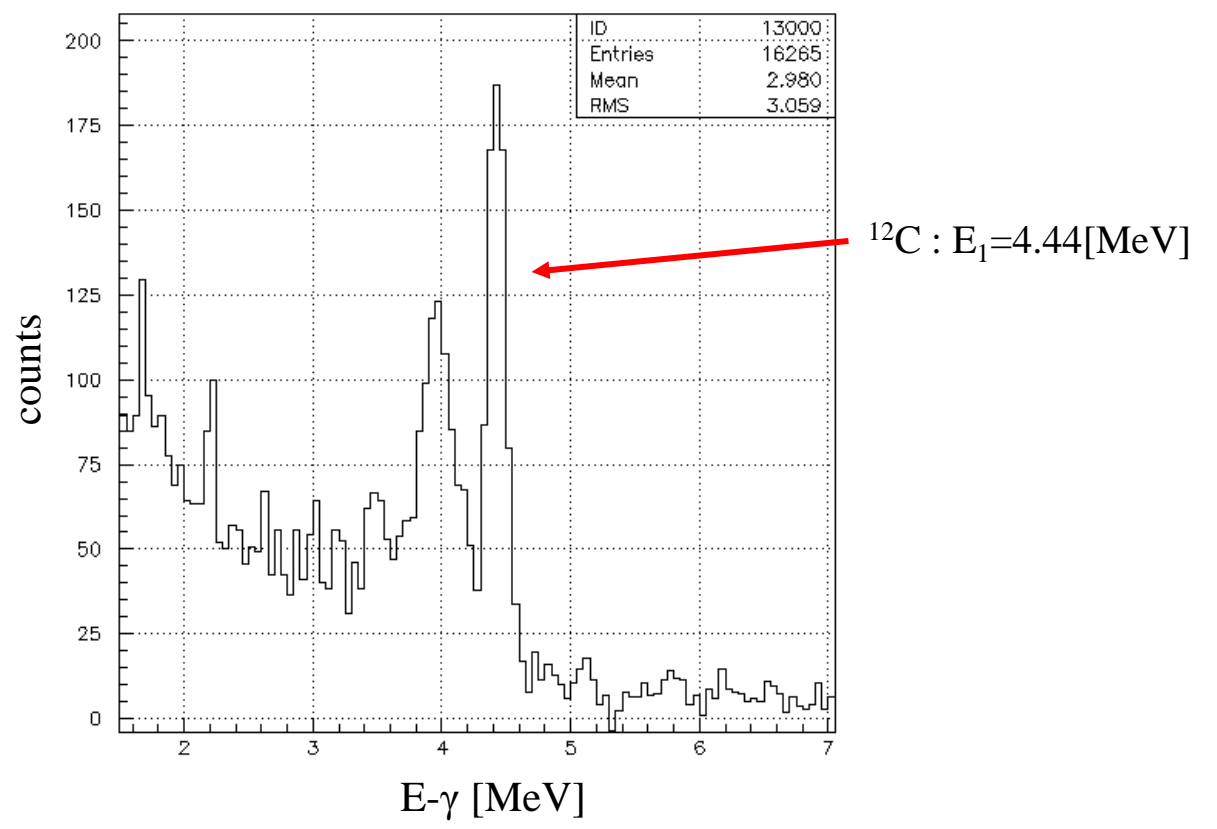
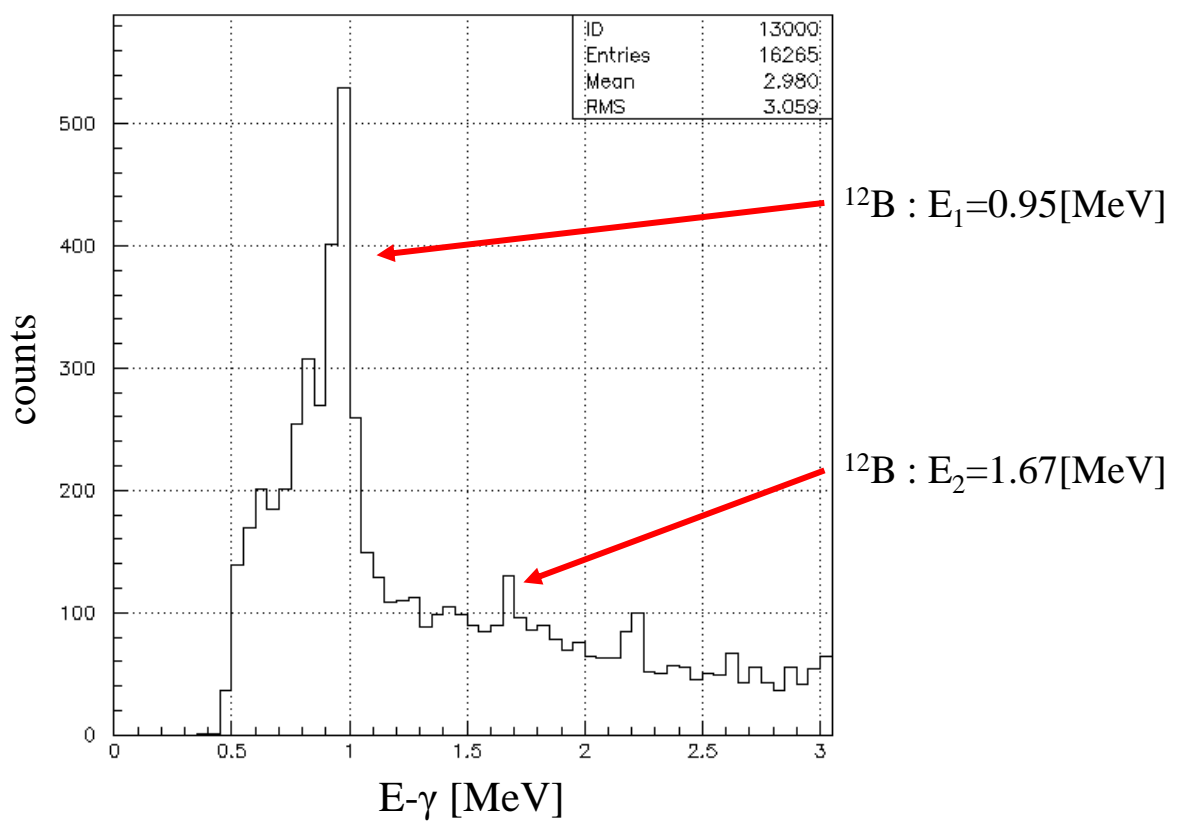


¹³C 4.5 degrees

Branching ratio

$$\frac{{}^{13}\text{C}(p, p'\gamma) \text{ counts}}{{}^{13}\text{C}(p, p') \text{ counts}} \times \frac{100}{\text{efficiency}} [\%]$$

GR analysis has not been done yet

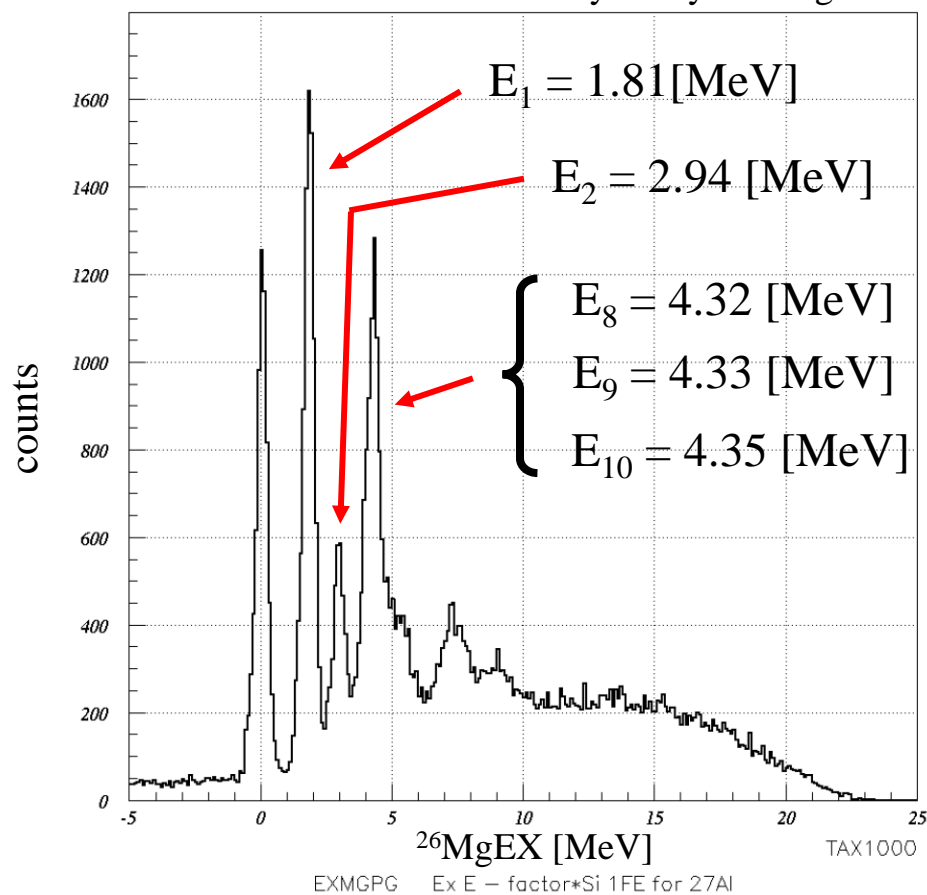


^{27}Al 4.5 degrees

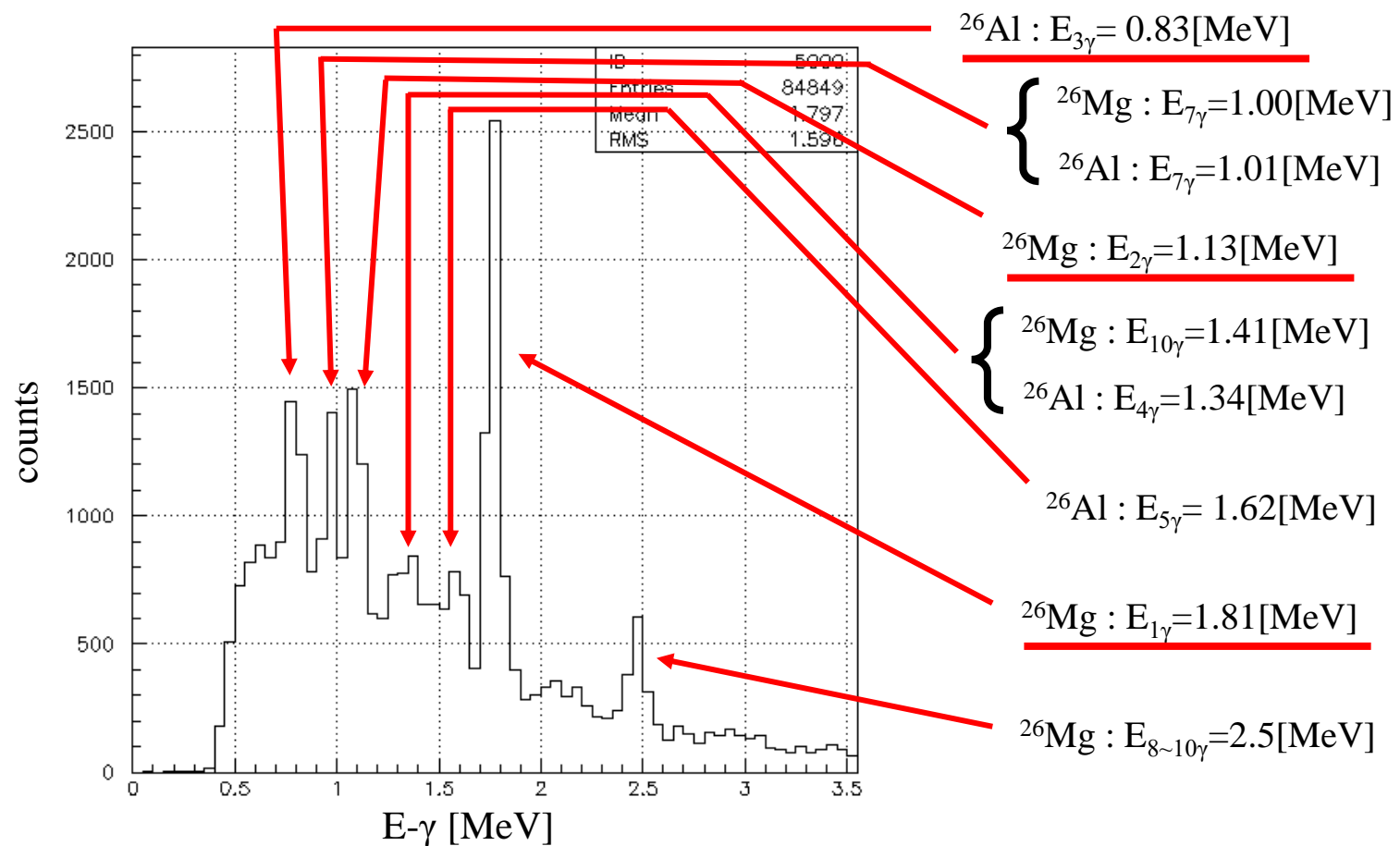
Branching ratio

$$\frac{{}^{12}C(p, p'\gamma) \text{ counts}}{{}^{12}C(p, p') \text{ counts}} \times \frac{100}{\text{efficiency}} [\%]$$

Analyzed by Y.Sasagawa



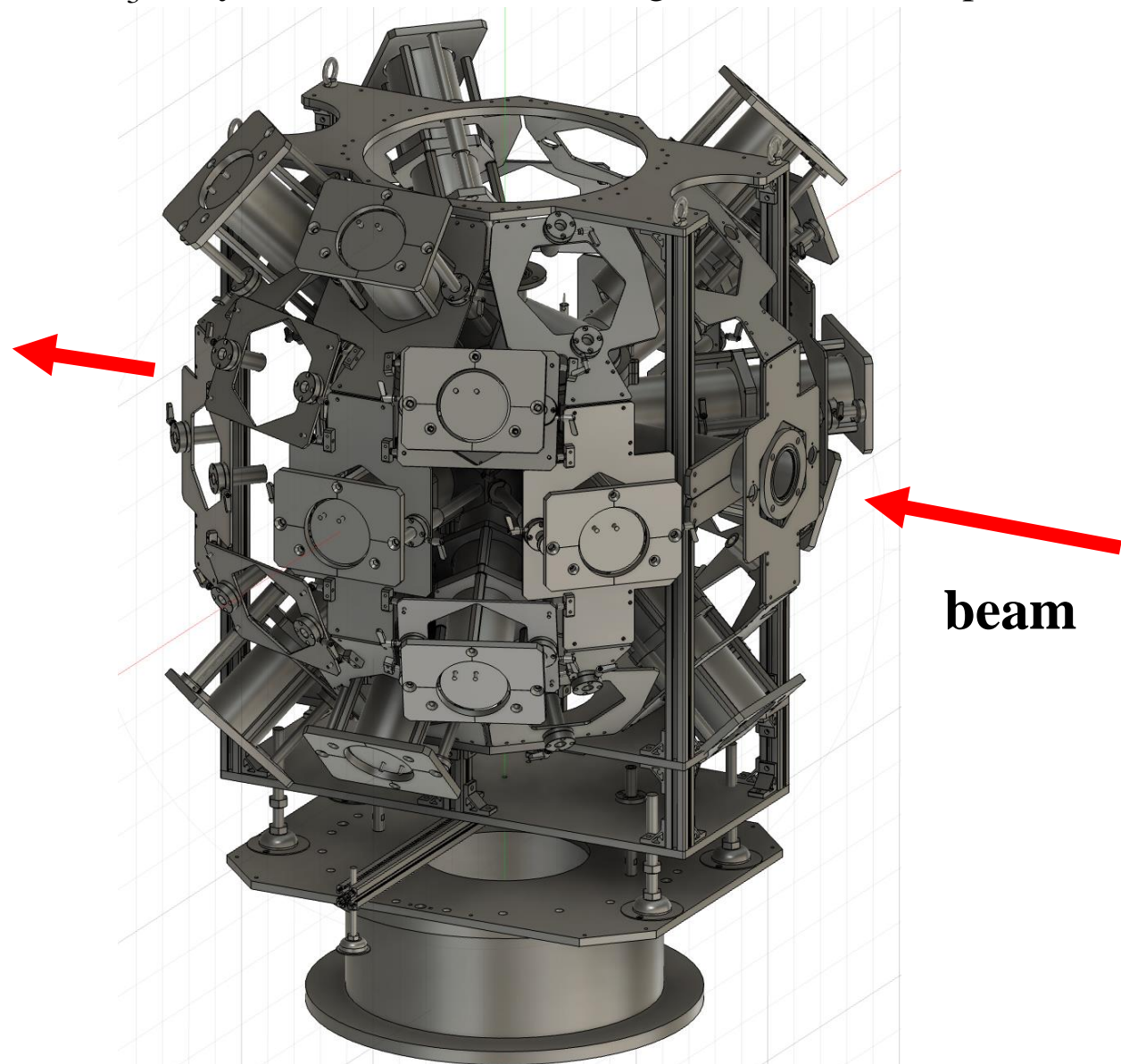
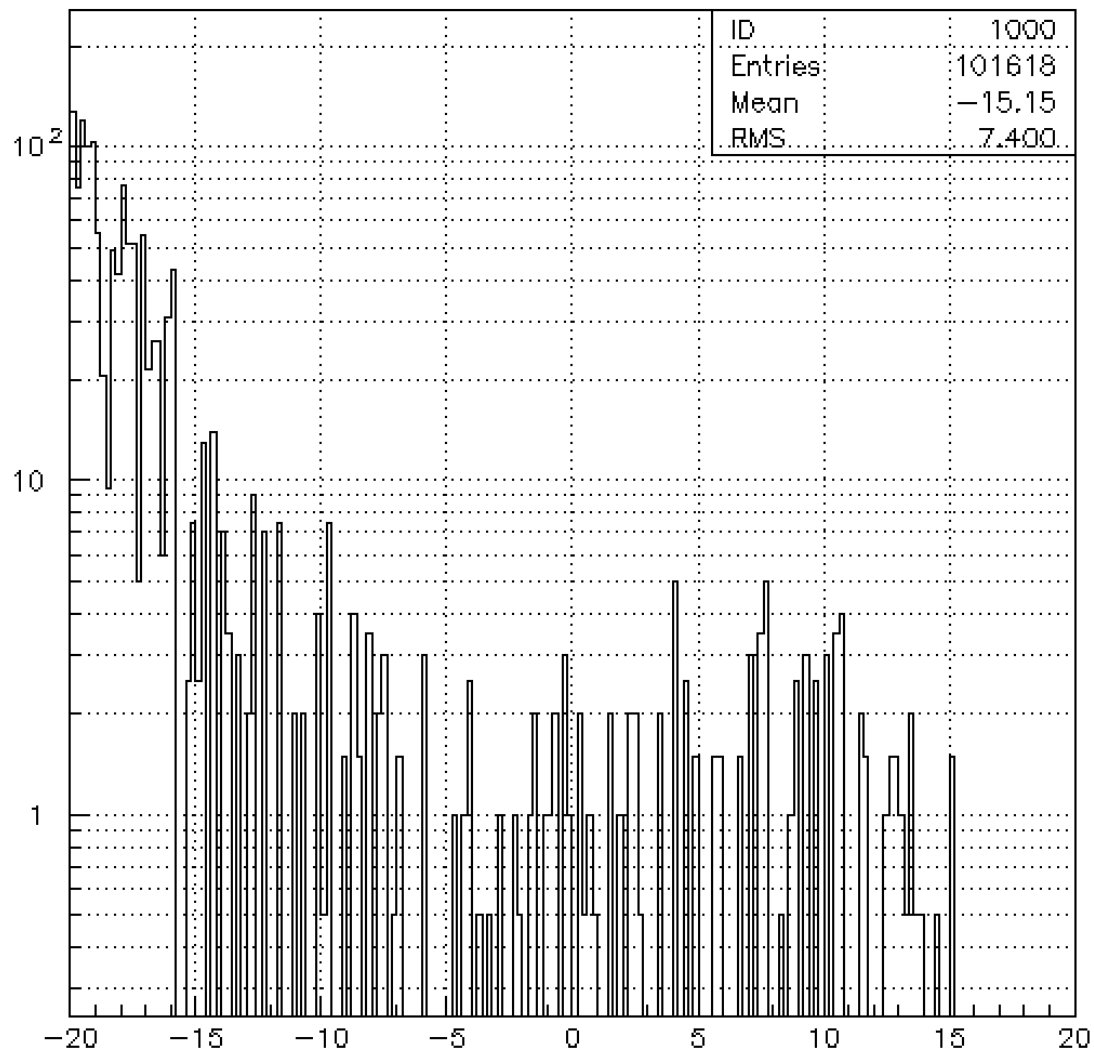
GR analysis has not been done yet



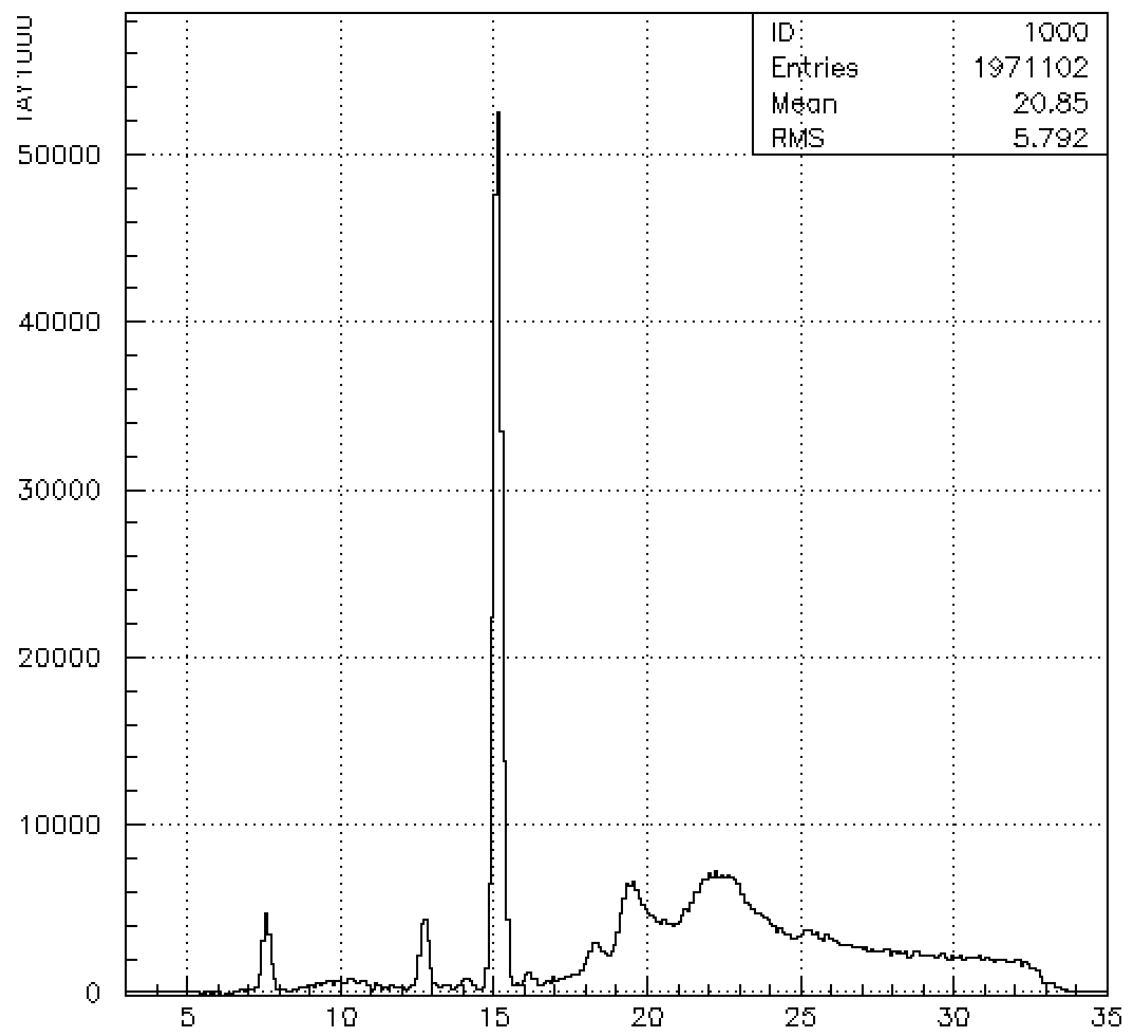
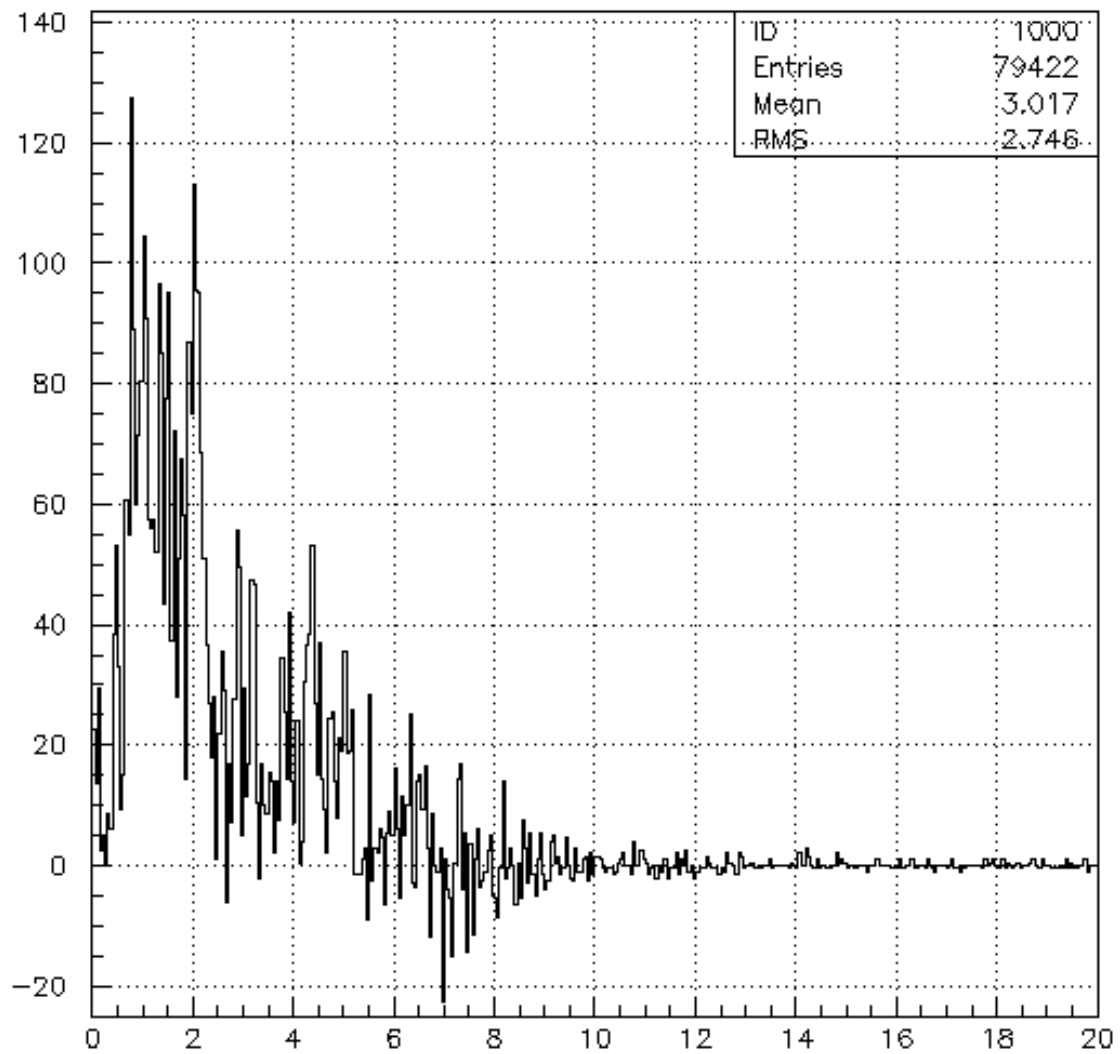
E (level) (keV)	XREF		J ⁿ (level)	T _{1/2} (level)	E (γ) (keV)	I (γ)	M (γ)	Final Levels		
0	AB DE GHI K MNOP	TU WX Za cdef	ijklm opq	STABLE						
4439.82 21	AB DE GHI K MNOP	TU WXY a cdef	ijklm opq	10.8×10 ⁻³ eV 6 % IT = 100	4438.94	100	E2	0	0+	
7654.07 19	AB DE GHIJK MNOP	TUVWX a cdef	ijkl nopqr	9.3 eV 9 % IT = 4.16×10 ⁻² % α ≈ 100	3213.79	100	E2	4439.82	2+	
9641 5	DE GHIJ MNOP	TUVWXY a cdef	ijklmnopqr	46 keV 3 % α ≈ 100 % IT < 4.1×10 ⁻⁵	9637	100	E3	0	0+	
9870 60		ef k		850 keV 85 % IT ≈ 7.1×10 ⁻⁶ % α ≈ 100						
9930 30 ?		Z f k		2710 keV 80						
10.3E+3 3 ?	AB	N	Z d i n v	(0+)	3.0 MeV 7 % α ≈ 100					
10847 4	D GH MN	TUV X a def	ijklmn pq	1-	273 keV 5 % α ≈ 100					
11836 4	D GH MNO	TUV XY a def	ijkl n p	2-	230 keV 8 % α ≈ 100 % IT > 0					
12400?		N		(5+,4-,6-,7+)	% α ≈ 100					
12710 6	AB GH MNOP	TUVWXY a cde	g ij l n p	1+	18.1 eV 28 % IT = 2.2 % α = 97.8	8267 12703	15 3 100 14	M1 M1	4439.82 0	2+ 0+
13.3E3 2 ?	G			4+	1.7 MeV 2					
13316 20	I MN	T V Y de	kl n p	4-	360 keV 43 % α ≈ 100 % IT > 0					
14079 5	GHI MN	T a de g	jkl nopq	4+	272 keV 6 % α ≈ 100					
E (level) (keV)	XREF		J ⁿ (level)	T _{1/2} (level)	E (γ) (keV)	I (γ)	M (γ)	Final Levels		
15110 3	B E MNOP	TU XY a cde	ij	1+	43.6 eV 10 % IT = 95.9 % α = 4.1	2400 4809 7453 10665 15100	1.5 4 4.2 15 2.83 36 2.49 34 100 2	M1 M1 M1 M1	12710 10.3E+3 7654.07 4439.82 0	1+ (0+) 0+ 2+ 0+
15440 40		a e hijk	v	(2+)	1.77 MeV 20 % α ≈ 100					

ELPIS

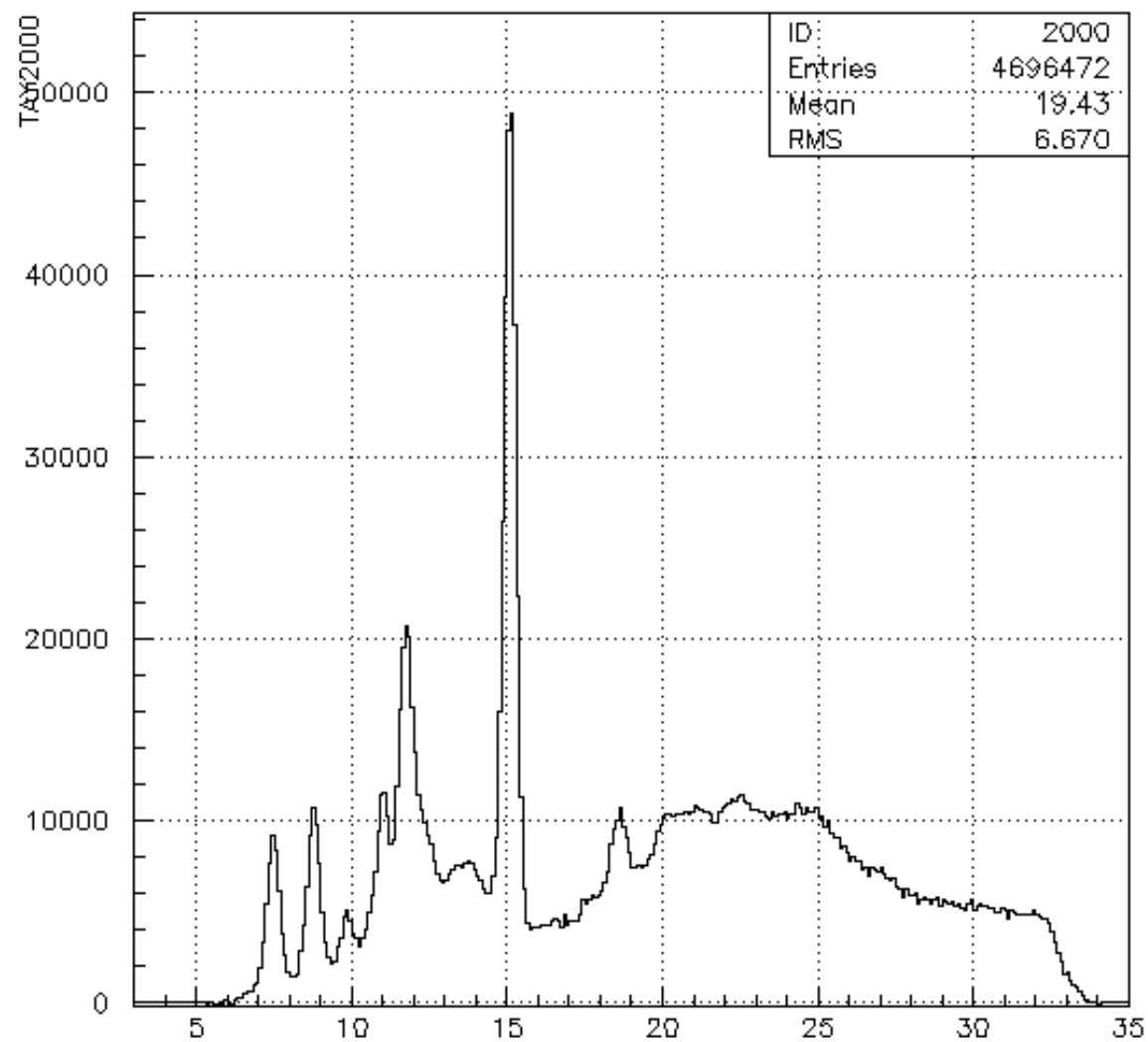
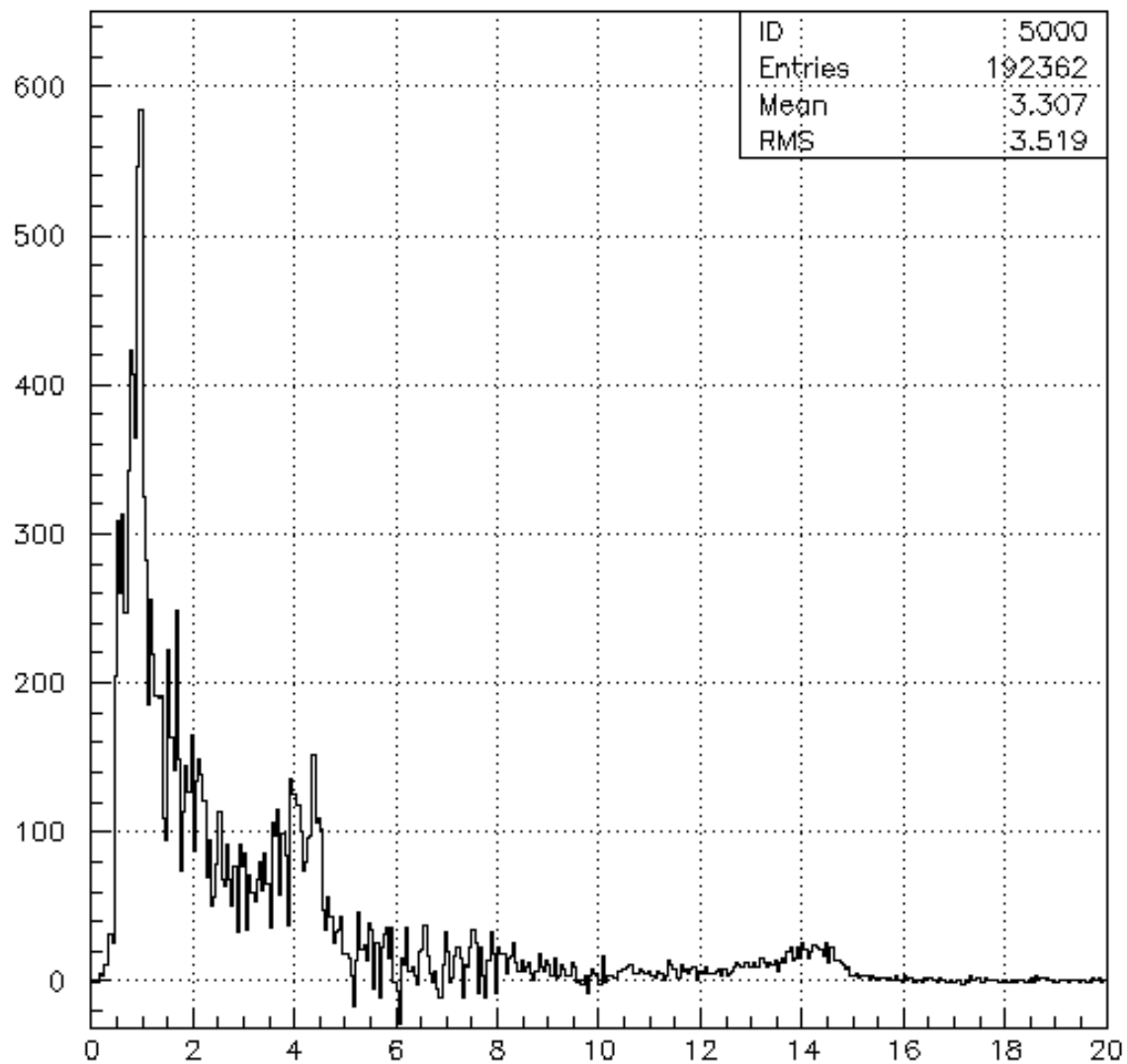
(Expanded LaBr_3 array for PANDORA Investigations with GR Spectrometer)



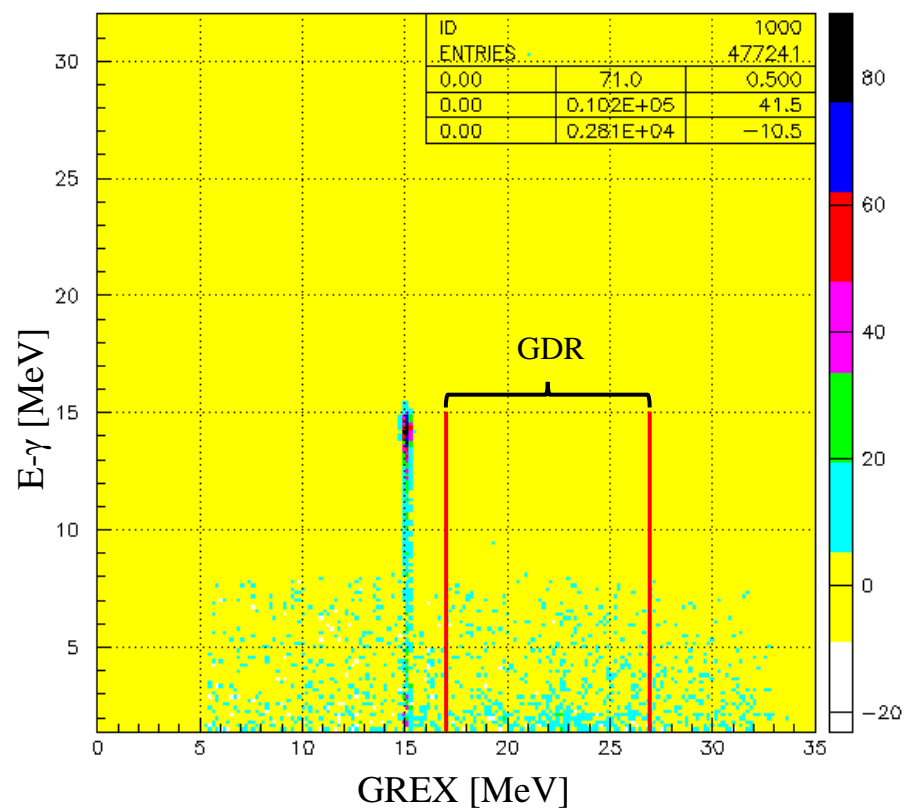
12C 0 degree



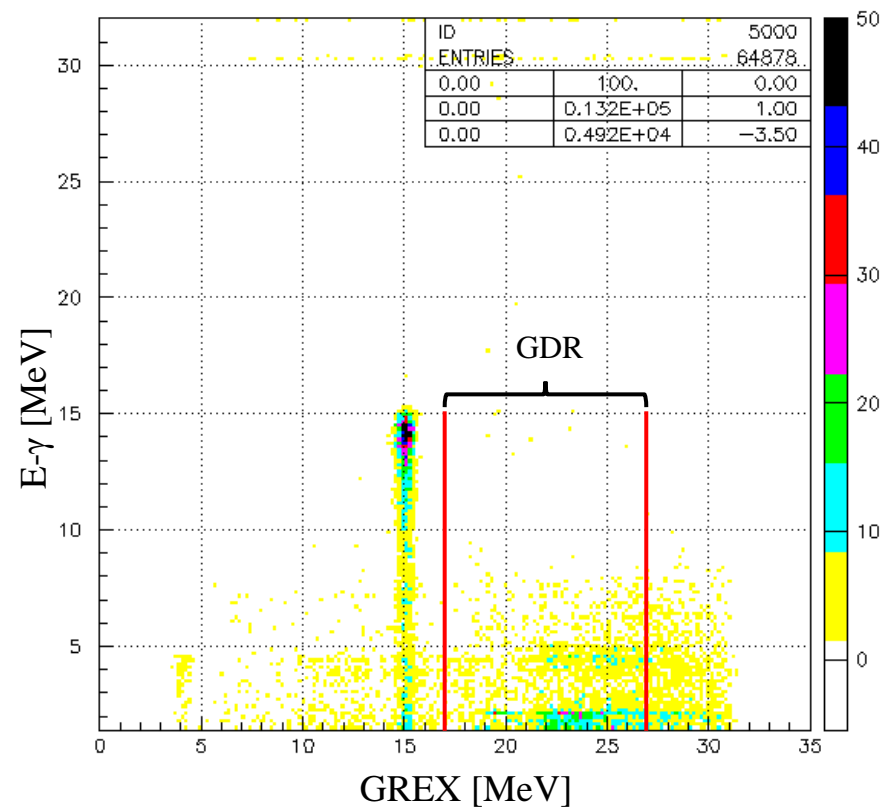
^{13}C 0 degree



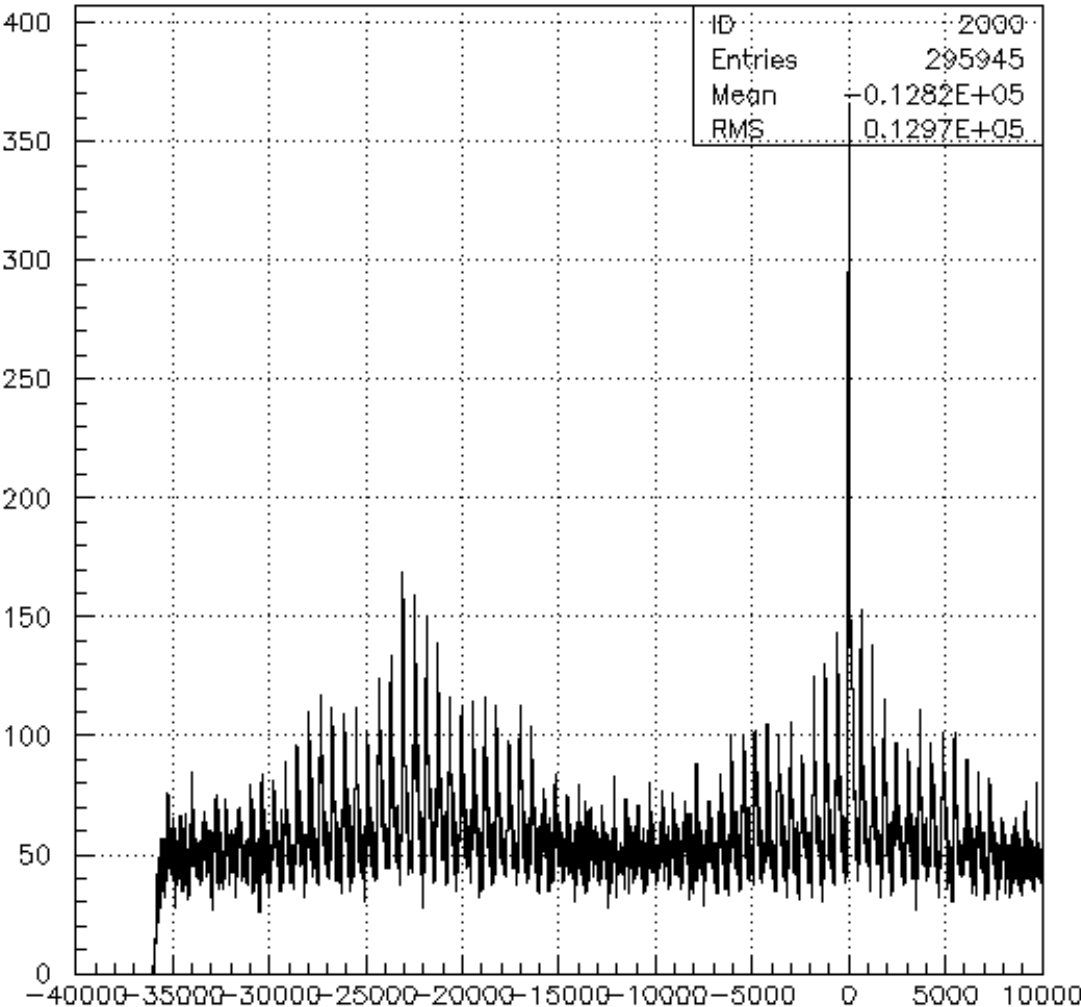
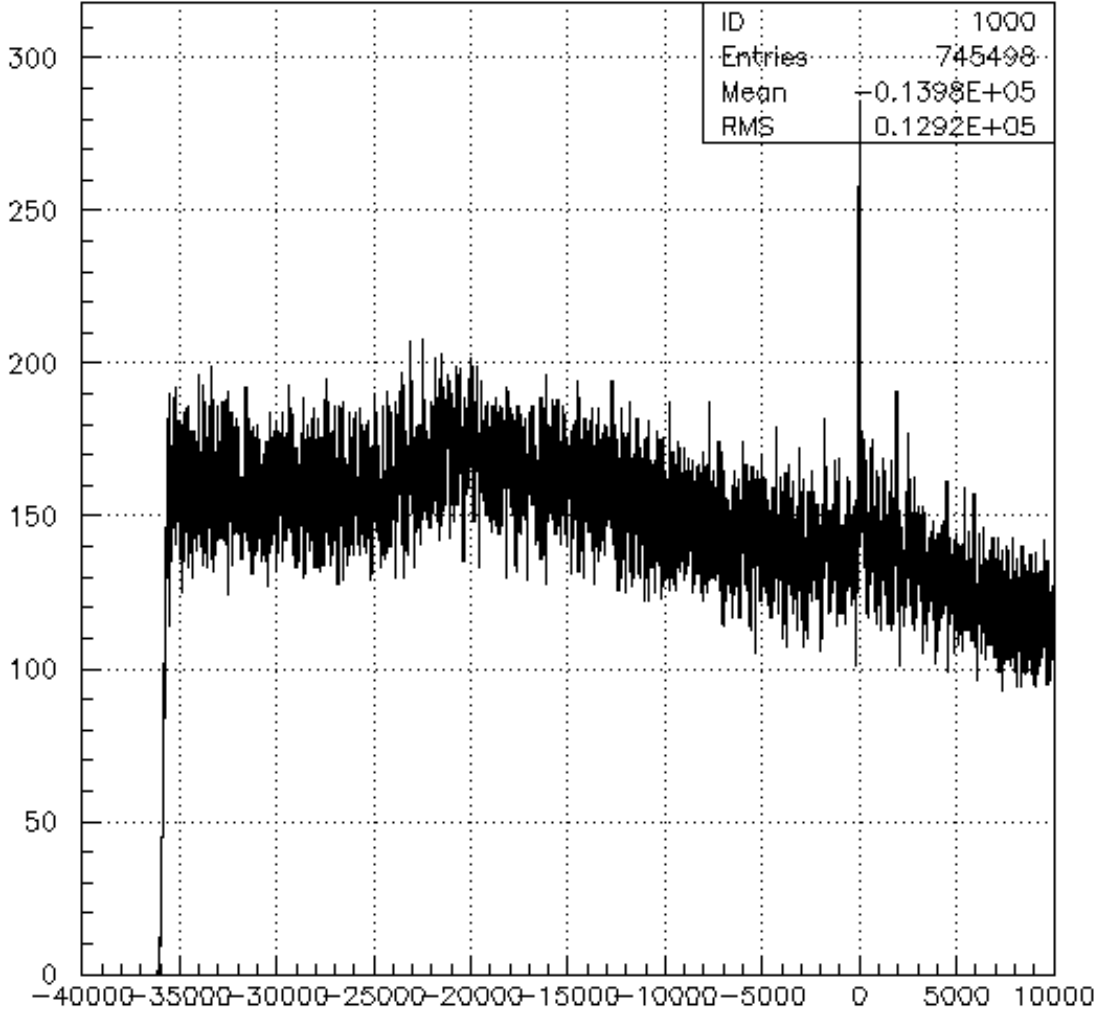
12C 0 degree



12C 0 degree



TDC data

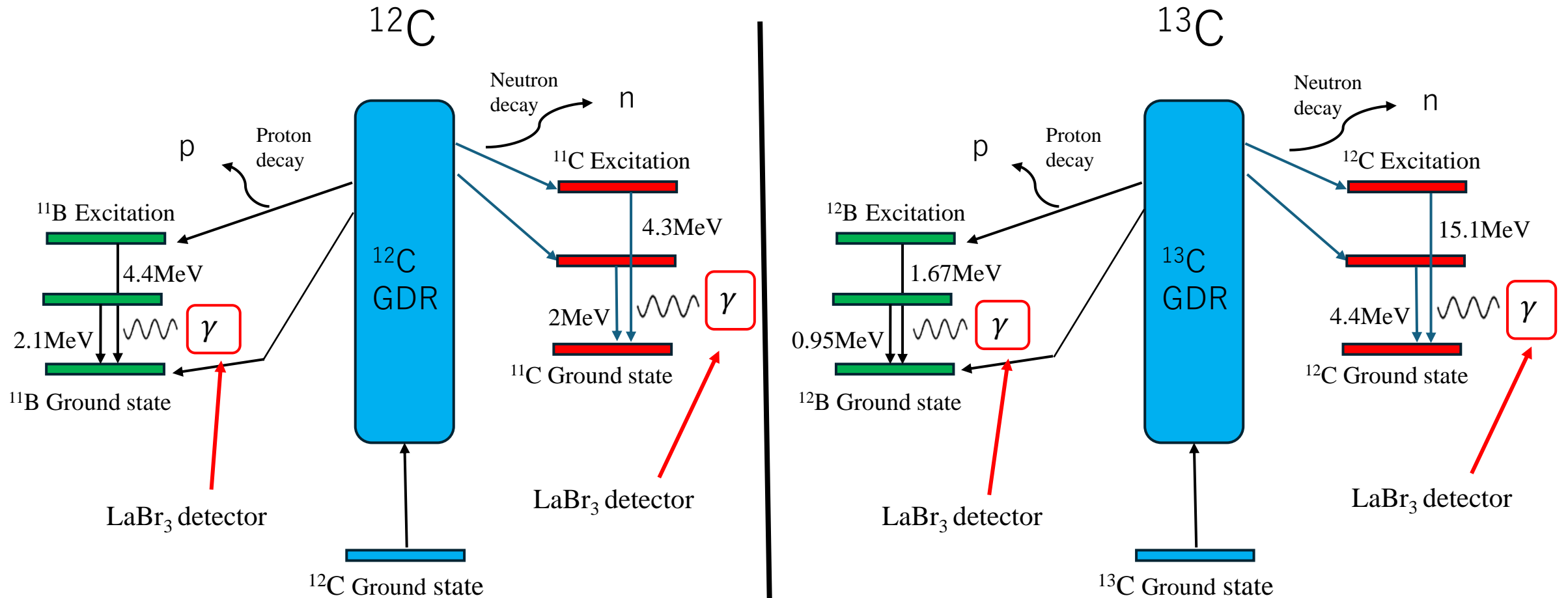


GDR & decays

Targets : ^{12}C , ^{13}C

Scattering angles : 0° , 4.5°

GDR : 17 ~ 27 MeV

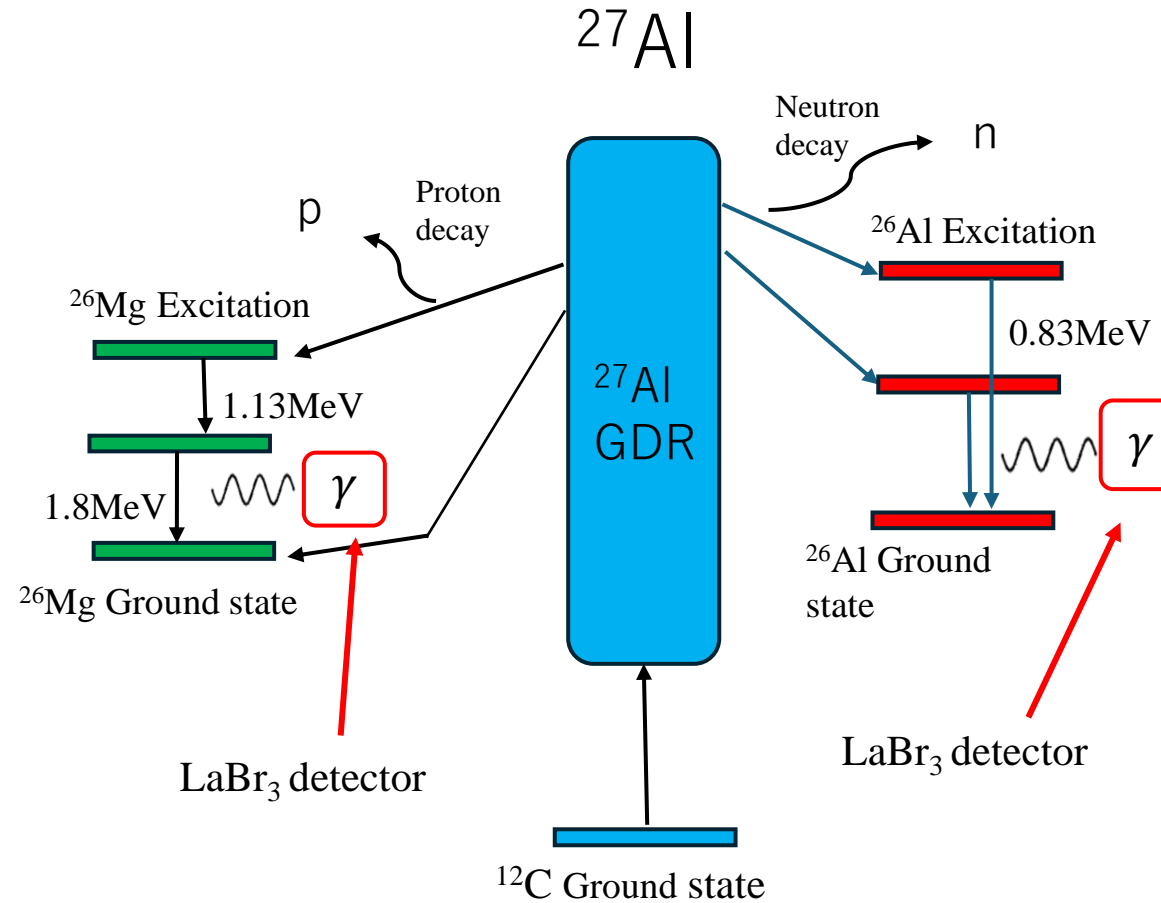


GDR & decays

Targets : ^{27}Al

Scattering angles : 0° , 4.5°

GDR : 15 ~ 25 MeV



Proton decay (^{26}Mg)

$S(p) = 8.3 \text{ MeV}$

GDR : 15 ~ 25 MeV

XREFs ✓	J^π ✓	$T_{1/2}/\text{Decay}$ ✓	$E(\gamma)$ ✓	$I(\gamma)$ ✓	$M(\gamma)$ ✓	Final Levels ✓	
E (level) (keV)	XREF	J^π (level)	$T_{1/2}$ (level)	E (γ) (keV)	I (γ)	M (γ)	Final Levels
0.0	ABCDEFGHI KLM OPQRSTU	0+	STABLE				
1808.74 4	AB DE GHI KLM OPQR TU	2+	476 fs 21	<u>1808.68 4</u>	100	E2	0.0 0+
2938.33 4	AB E GHI KLM OPQR TU	2+	141 fs 8	<u>1129.61 4</u> <u>2938.15 5</u>	100.0 6 10.7 6	M1+E2 E2	1808.74 2+ 0.0 0+
3082.9 20	E			1274.1 20	100		1808.74 2+
3420.2 17	E			1611.4 17	100	D+Q	1808.74 2+
3564.9 19	E			1756.1 19	100		1808.74 2+
3588.56 9	E GHI KLM OPQR TU	0+	6.45 ps 48	<u>1779.74 8</u>	100	E2	1808.74 2+
3941.57 4	A E GHI KL OP R TU	3+	0.83 ps 12	<u>1003.25 4</u> <u>2132.71 4</u>	100.0 16 61.3 16	M1+E2 M1	2938.33 2+ 1808.74 2+
4318.89 5	A E GHI L OPQR TU	4+	272 fs 16	1380.88 18 <u>2510.01 5</u>	1.85 11 100 2	[E2]	2938.33 2+ 1808.74 2+
4332.52 5	A E GHI LM OPQR U	2+	20 fs 3	1394.28 7 <u>2523.69 6</u> 4332.2 3	19.3 10 100.0 13 7.6 8	[E2]	2938.33 2+ 1808.74 2+ 0.0 0+
4350.09 4	A E G I L OPQR	3+	105 fs 28	409.4 5 <u>1411.72 4</u> <u>2541.18 6</u>	0.041 25 93 4 100 4	M1+E2 M1+E2	3941.57 3+ 2938.33 2+ 1808.74 2+
4644.9 13	E			2836.0 13	100	D+Q	1808.74 2+

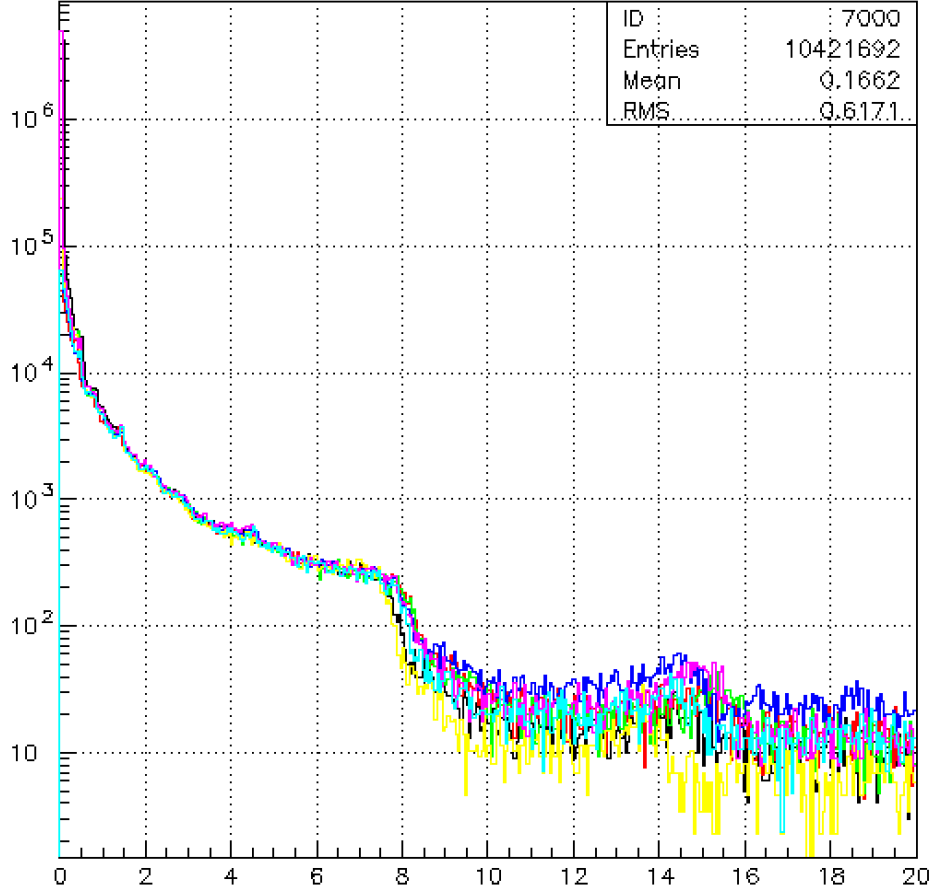
Proton decay (^{26}Al)

$$S(n) = 13.1 \text{ MeV}$$

GDR : 15 ~ 25 MeV

E (level) (keV)	XREF	J ^π (level)	T _{1/2} (level)	E (γ) (keV)	I (γ)	M (γ)	Final Levels	
0.0	ABCDEFGHIJKLMN	5+	7.17×10 ⁺⁵ y 24 % ε = 100					
228.305 13	AB DEFGHIJKLMNO	0+	6.3460 s 8 % ε = 100					
416.852 3	A D FGHJKLMN	3+	1.20 ns 1	416.848	100	[E2]	0.0	5+
1057.739 12	A CD FGHJKLMN	1+	25 fs 5	<u>829.3 4</u>	100	M1	228.305	0+
1759.034 8	A CD FGH IJ LMN P	2+	3.8 ps 6	701.285 <u>1342.145</u>	2.0 1 100.0 1		1057.739 416.852	1+ 3+
1850.62 3	A D FGH JKLMN	1+	32 fs 3	1433.73 <u>1622.0 7</u>	0.7 1 100.0 1	[E2] M1	416.852 228.305	3+ 0+
2068.86 5	CD FG J LMN	(4+)	310 fs 50	1651.95 2068.77	100.0 15 44.9 15	(M1+)E2	416.852 0.0	3+ 5+
2069.47 3	D FGH J LMN	(2+)	14 fs 2	218.85 310.43 <u>1011.71</u> 1652.56 1841.09	0.053 13 0.23 3 100.0 12 28.4 11 4.0 5	[M1] [M1] [M1] [M1] E2	1850.62 1759.034 1057.739 416.852 228.305	1+ 2+ 1+ 3+ 0+
2071.64 4	A CD FG JKLMN	1+	367 fs 69	221.02 1654.73 <u>1842.8 7</u>	0.045 11 11.9 10 100 1	[E2] M1	1850.62 416.852 228.305	1+ 3+ 0+
2365.150 18	CD FGH J LM OP	3+	0.8 ps 2	295.678 606.108 1307.375 1948.219 2365.034	100 2 2.90 10 26.9 8 64.7 20 1.75 8	[M1] [E2] [E2]	2069.47 1759.034 1057.739 416.852 0.0	(2+) 2+ 1+ 3+ 5+

Calibration



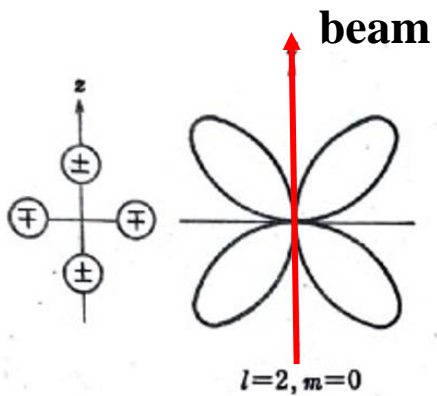
LaBr₃ setup on next experiment

2nd PANDORA experiment

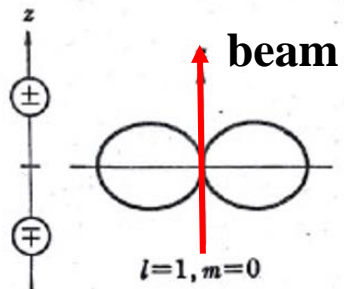
The number of LaBr₃ detectors will be 16.

Angles : 45°, 90°, 112.5°, 135°

4 detectors for each angle.



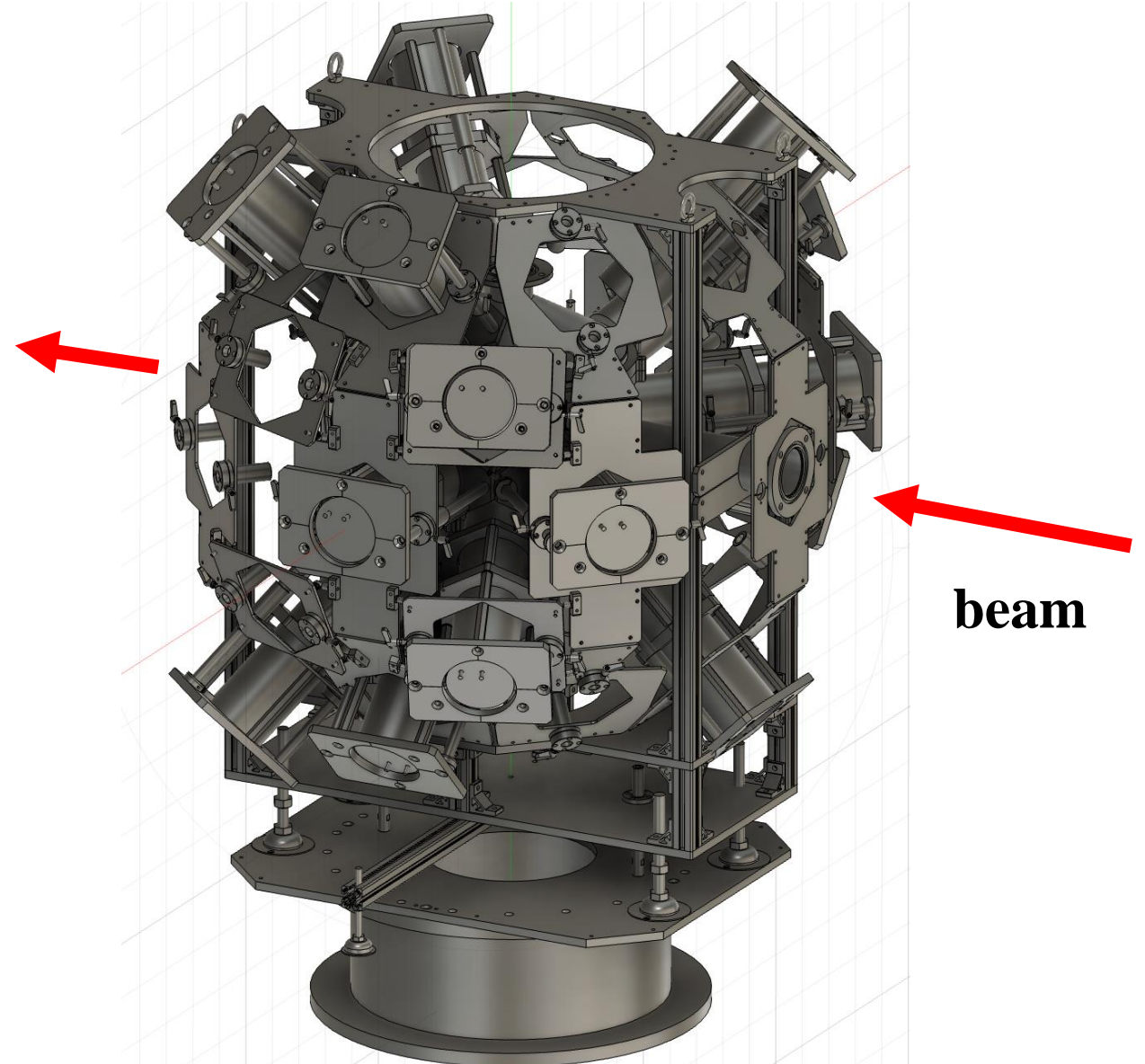
E2, M2 gamma ray



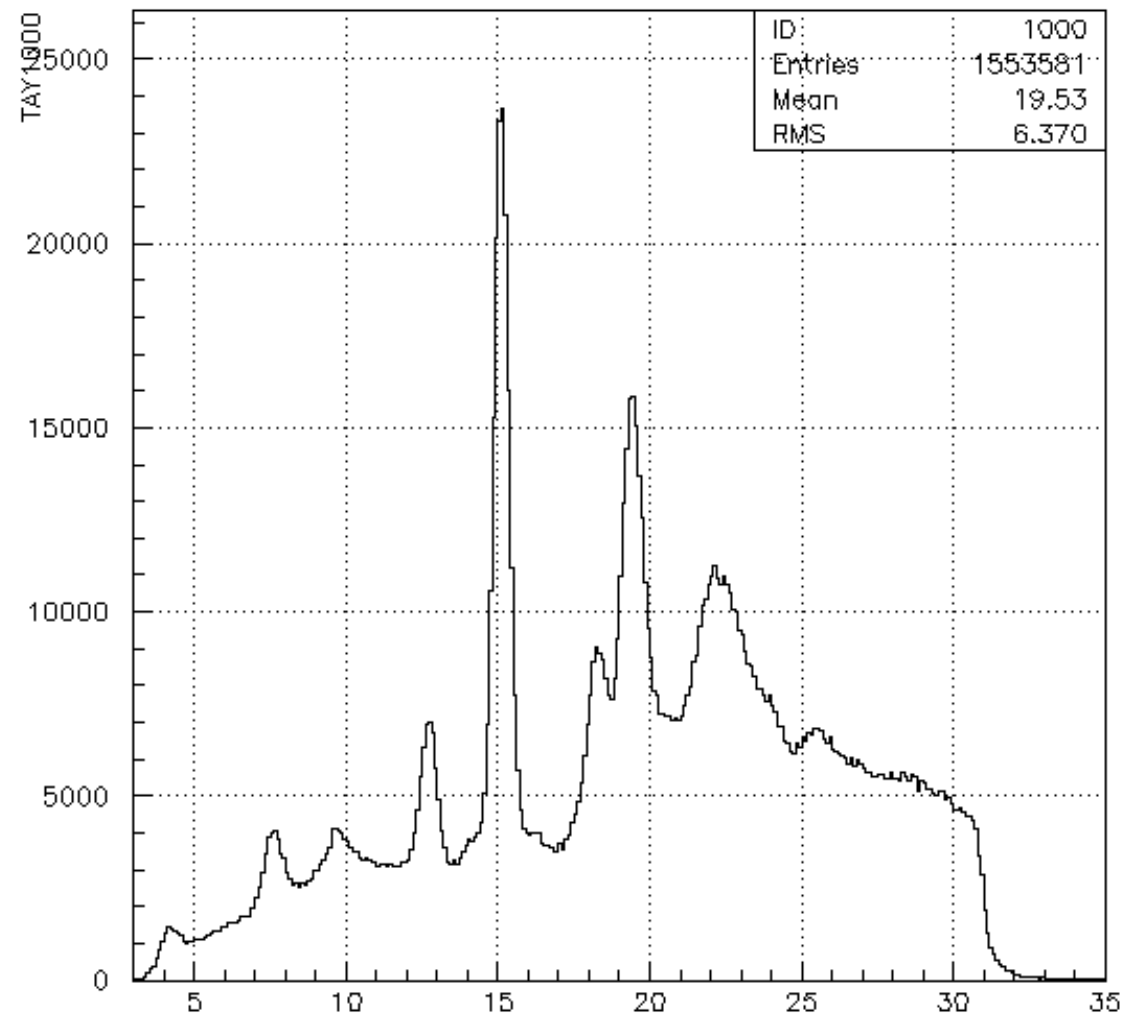
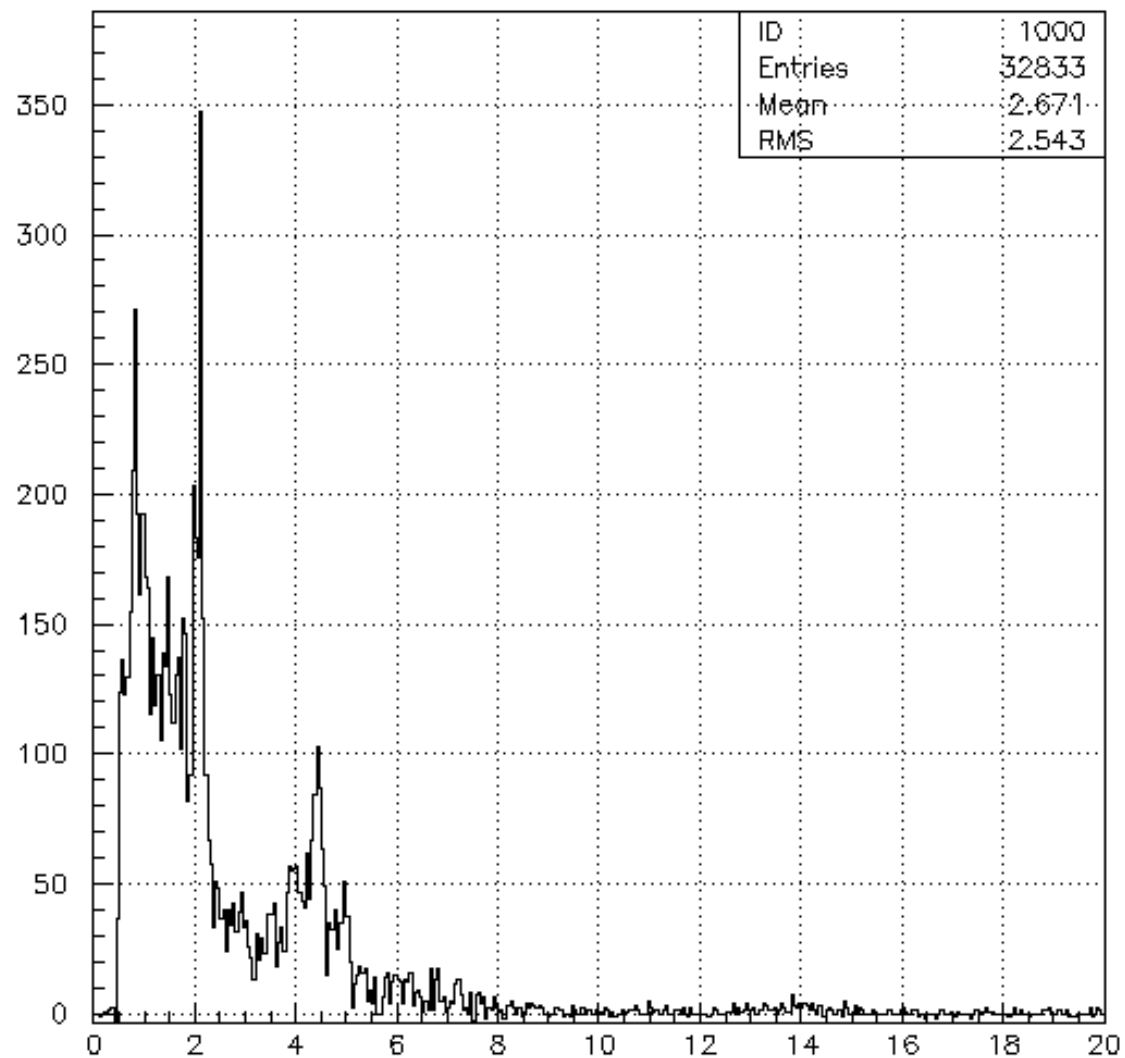
E1, M1 gamma ray

ELPIS

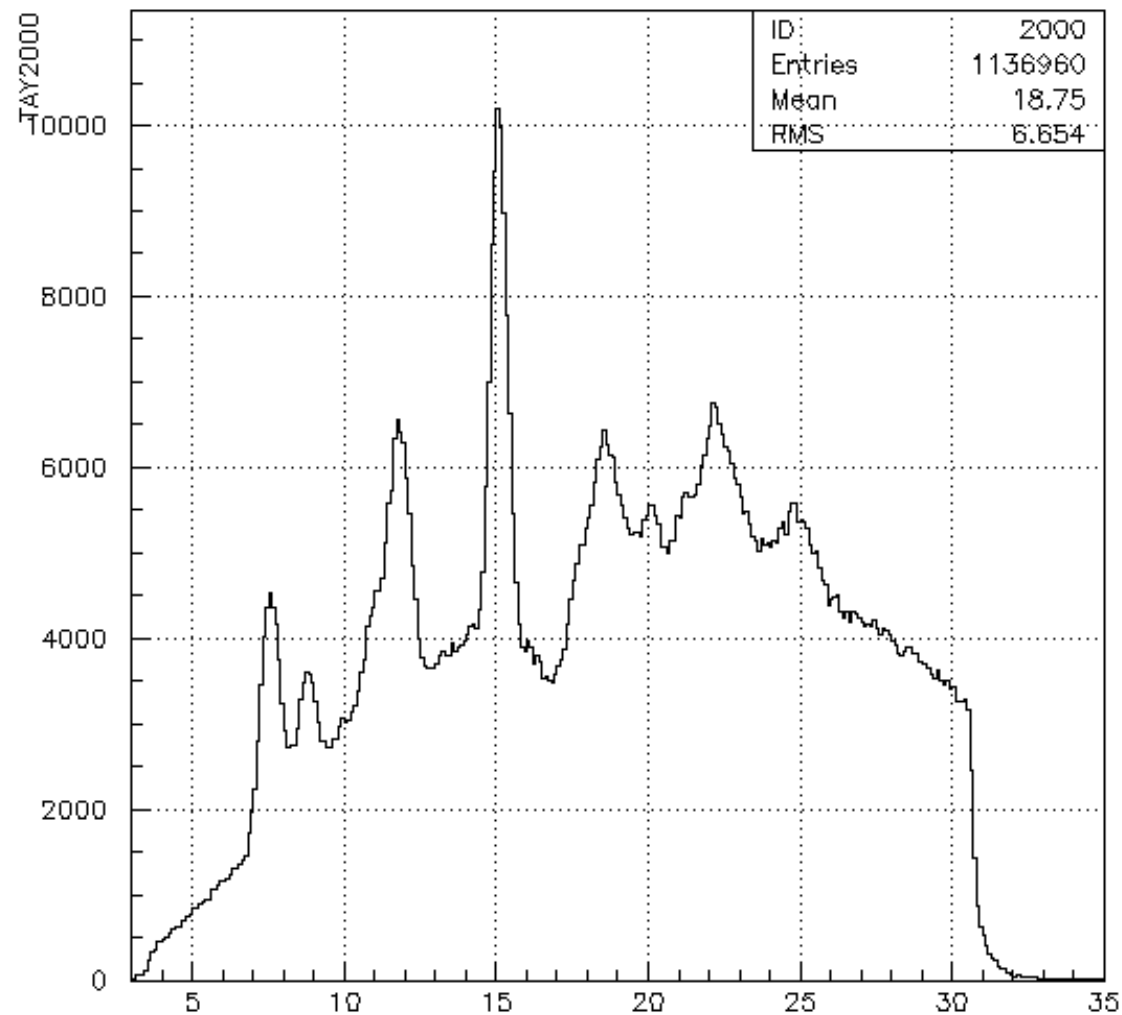
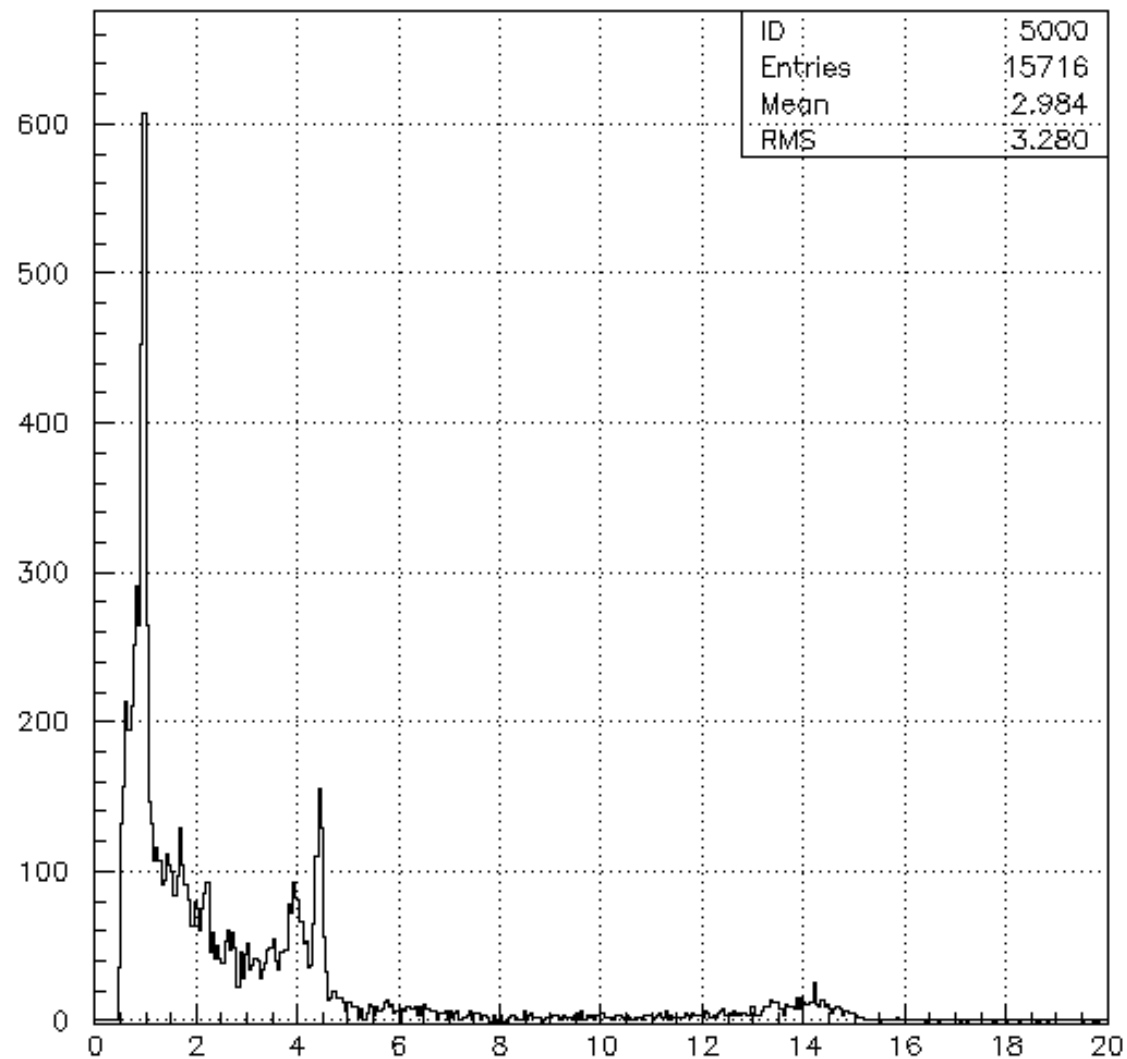
(Expanded LaBr₃ array for PANDORA Investigations with GR Spectrometer)



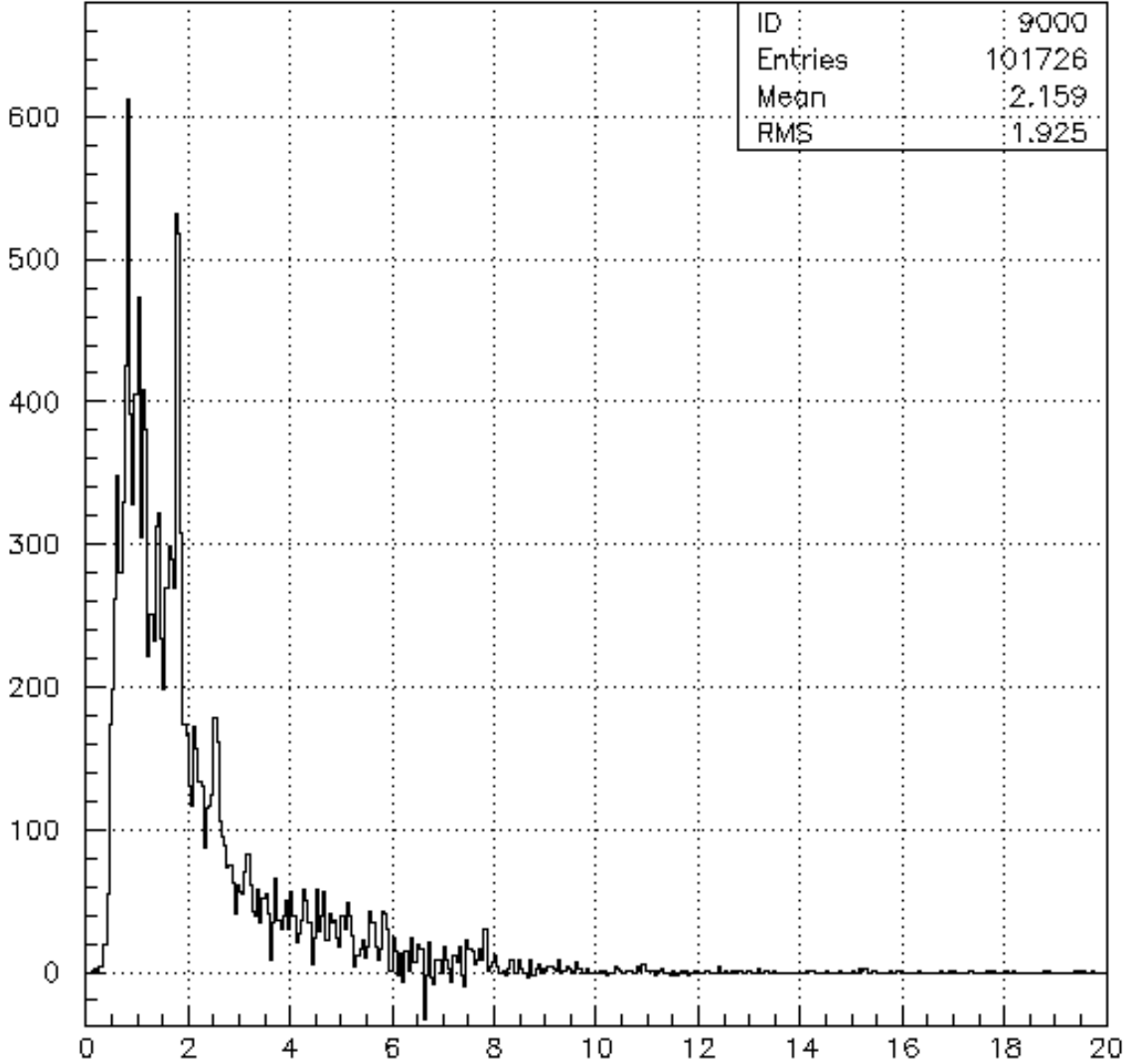
12C 4.5 degrees



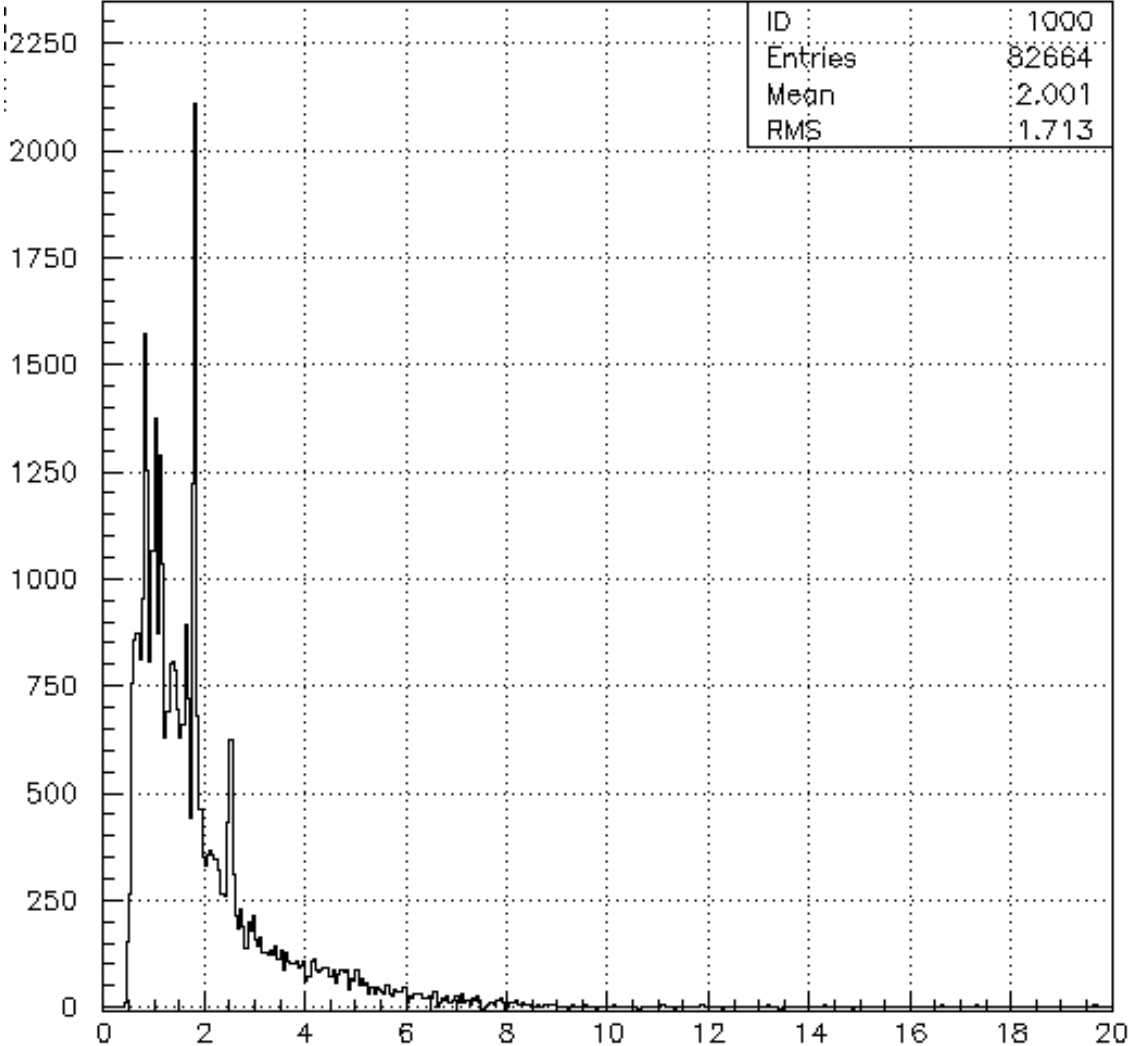
^{13}C 4.5 degrees



27Al 0 degree



27Al 4.5 degrees



Branching ratios & Efficiency

Branching ratio

$$\frac{{}^A X(p, p' \gamma) \text{ counts}}{{}^A X(p, p') \text{ counts}} \times \frac{100}{\text{efficiency}} [\%]$$

detector efficiency @ 15.1 MeV

$$\frac{\text{Number of 15.1 MeV } \gamma \text{-ray}}{\text{Number of 15.1 MeV excited state} \times 0.873} = 0.014$$

Efficiency was roughly calculated using the results of another thesis.

Renormalized multiplying by ratio

