

LIV.INNO



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LIVERPOOL

Latest Collider Neutrino Measurements with the FASER detector

HEPP & APP IOP Annual Conference 2026
Thursday 9th April 2026

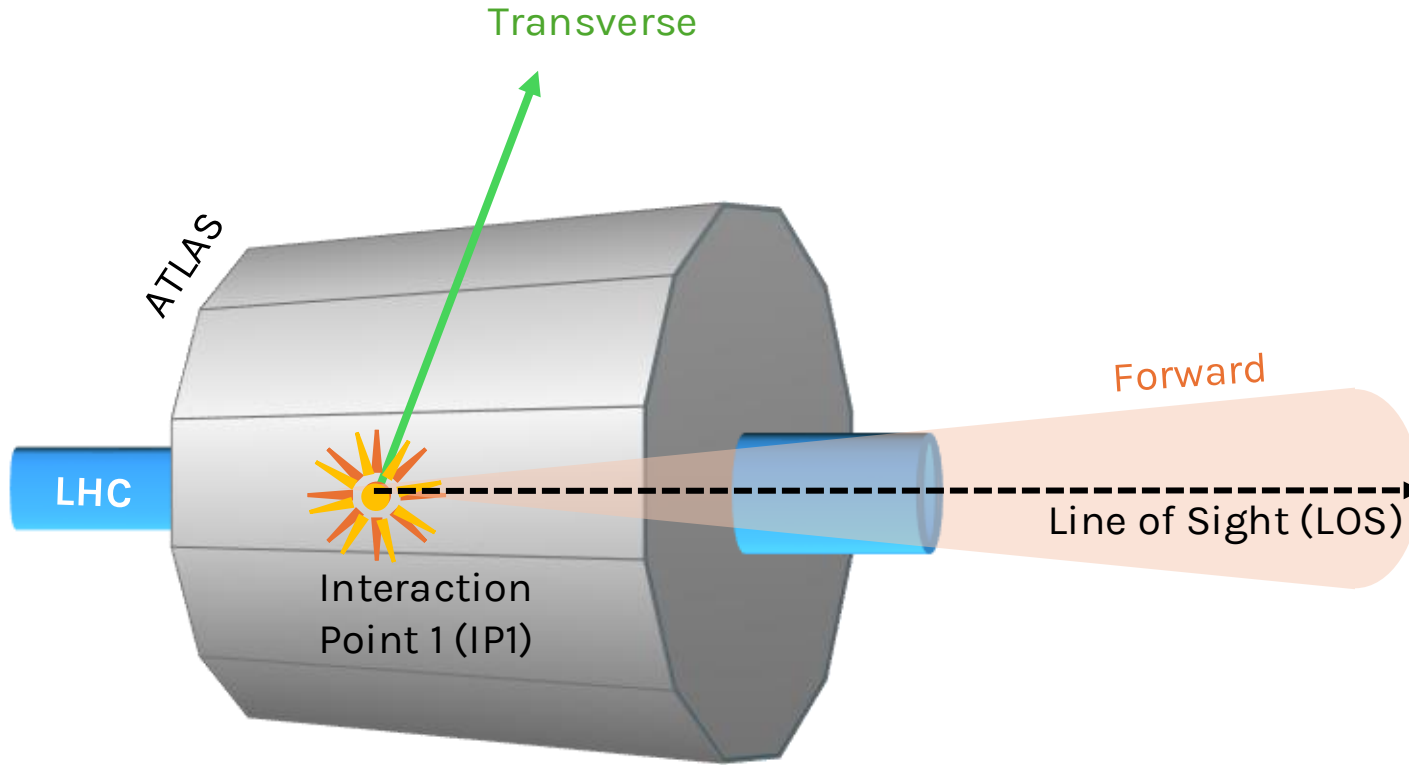
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Forward Physics at the LHC



Detectors at the LHC have **traditionally focused** on events that have:

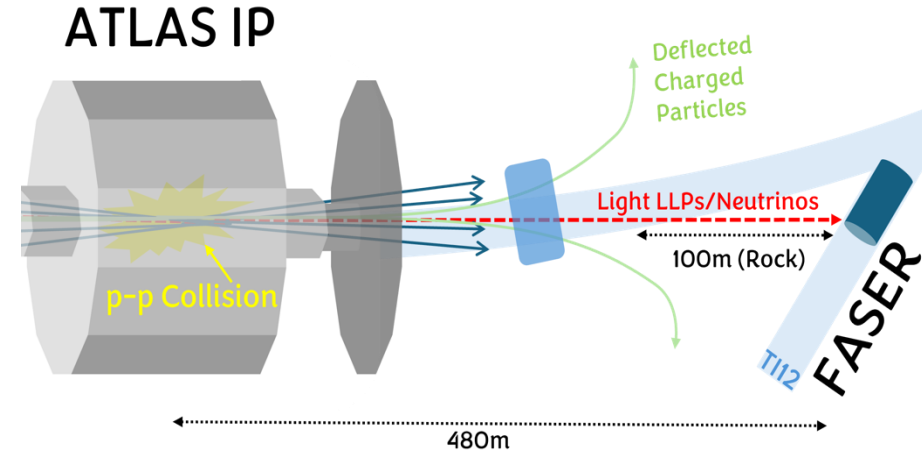
- **High** transverse momentum (p_T)
- **Strongly interacting** particles
- **Small pseudorapidity** (η)

Located around interaction points (IPs), these detectors typically **miss events** that have:

- **Low** p_T
- **Feebly interacting** particles (FIPs)
- **Large** η

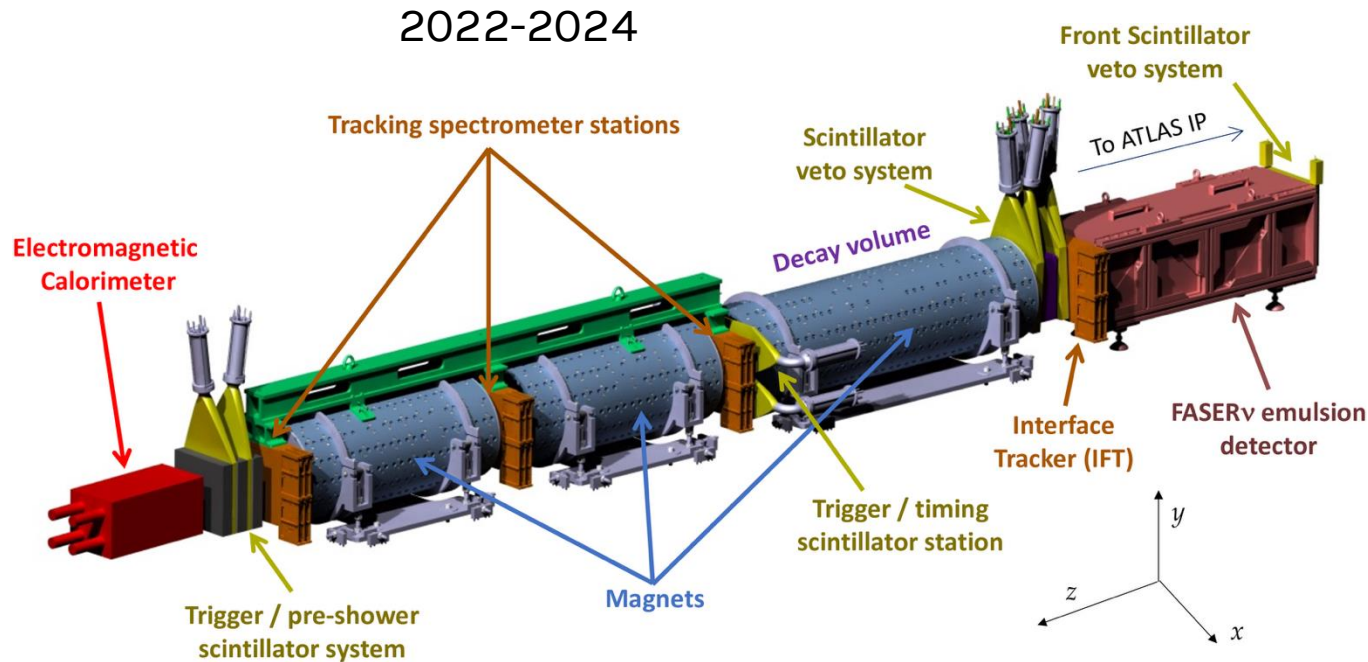
ForwArd Search ExpeRiment (FASER)

- Purposefully designed to:
 - **Search for BSM particles** i.e. Feebly-Interacting Particles (FIPS)
 - **Study** properties of collider-produced **neutrinos** at high energies
- Located **480m downstream**, on the line-of-sight (LOS) of the ATLAS IP
- Benefits from highly collimated flux of light hadrons along line of sight



FASER Detector

Detector Layout



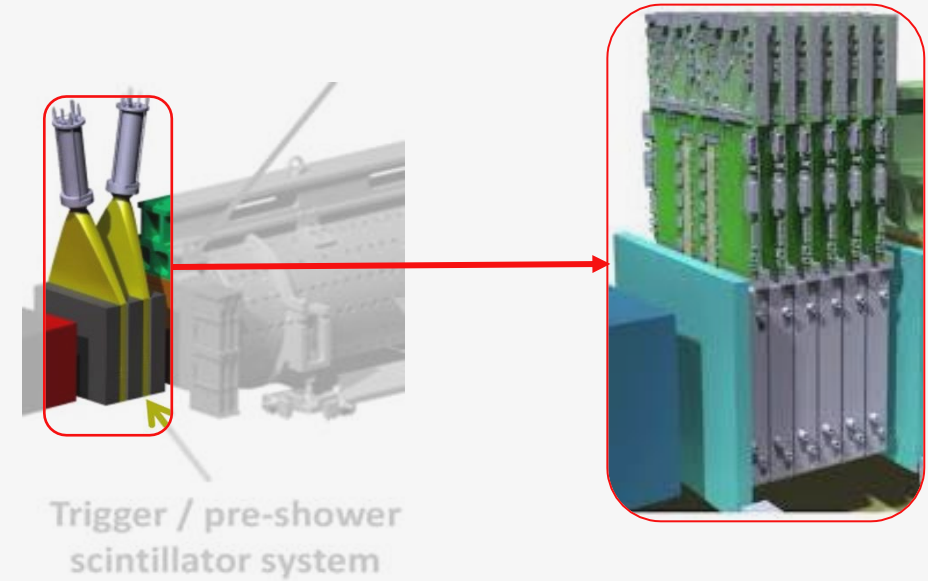
FASER detector comprised of:

- Electronic Spectrometer
- Neutrino emulsion detector (FASER ν)

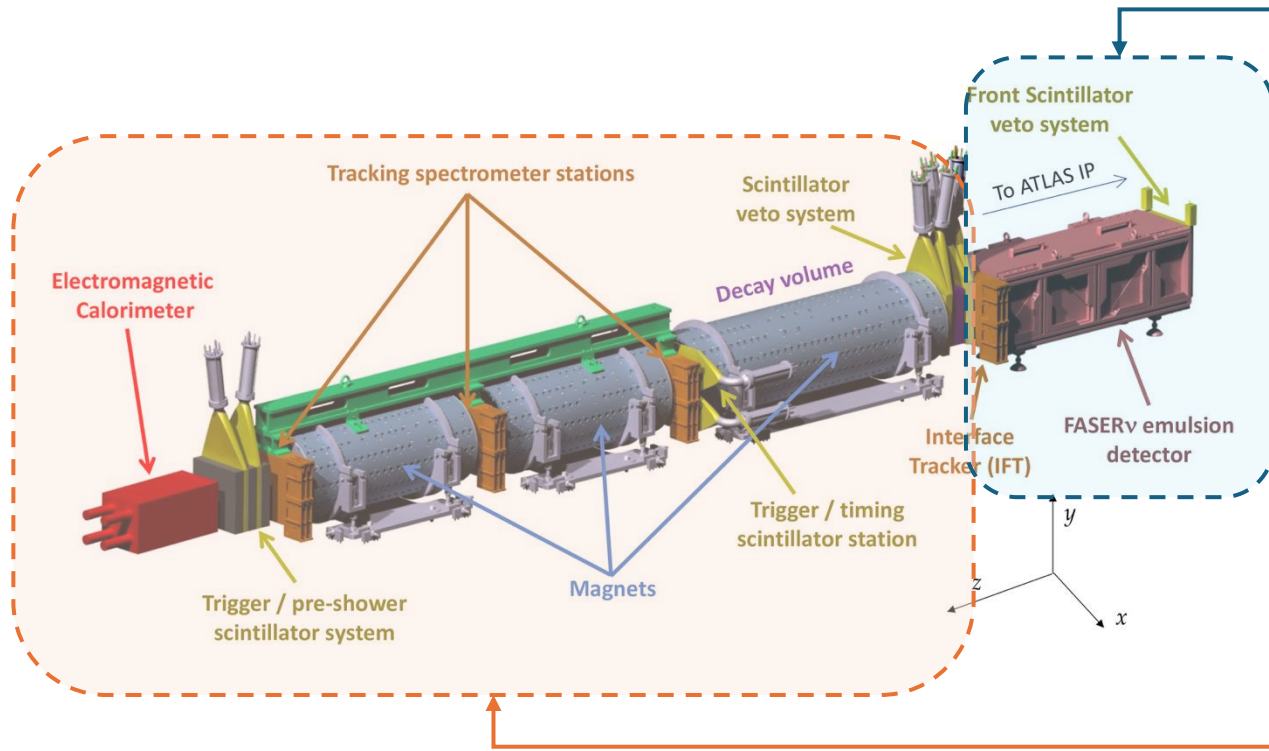
Both of these can be used to **detect neutrinos independently**

2025 - Present

Preshower upgraded with pixel-based detector



FASER Detector



Emulsion Detector

Slow to analyze

- Each film must be processed by hand and digitized before analysis

Exceptional resolution ($\sim 0.3 \mu\text{m}$)

- Emulsion films interleaved with tungsten plates

Electronic Detector

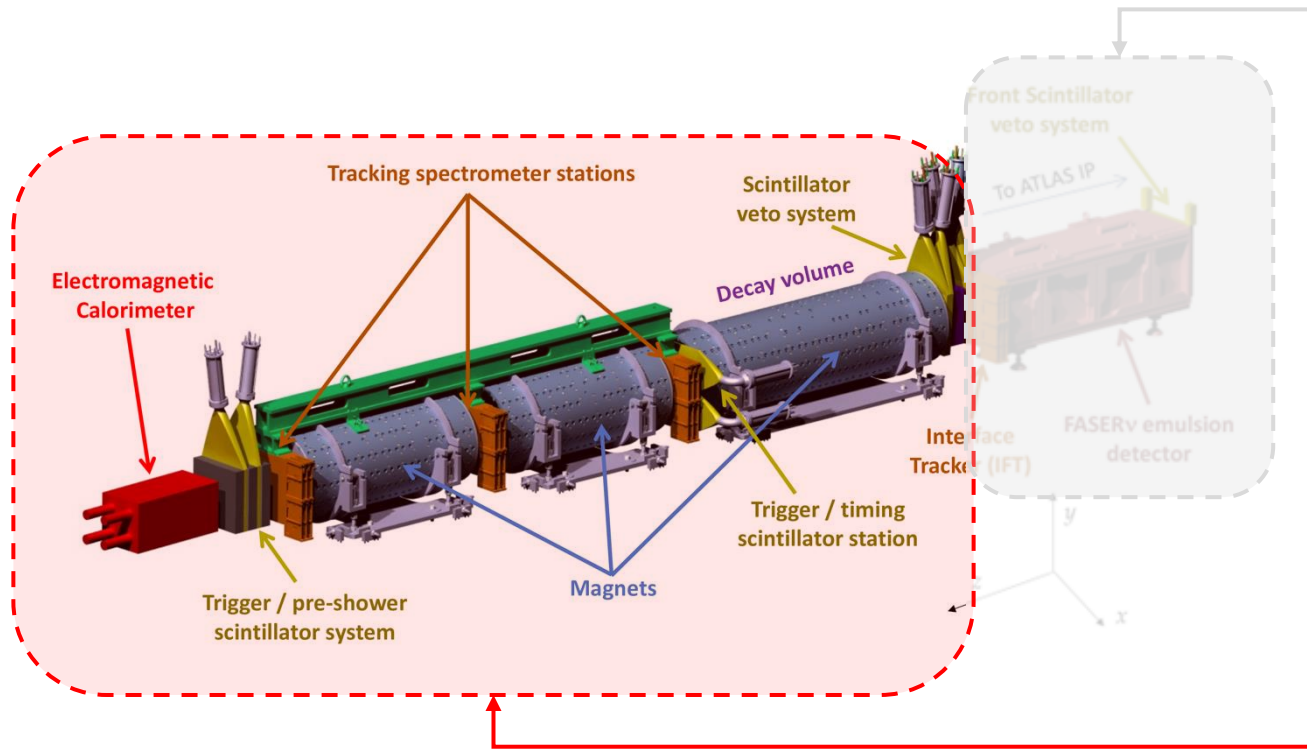
Faster to analyze

- Uses only electronic components of detector
- Benefits from prompt data reconstruction

Can **separate** neutrino decays by **charge**

- Can differentiate neutrino and anti-neutrino

FASER Detector



Emulsion Detector

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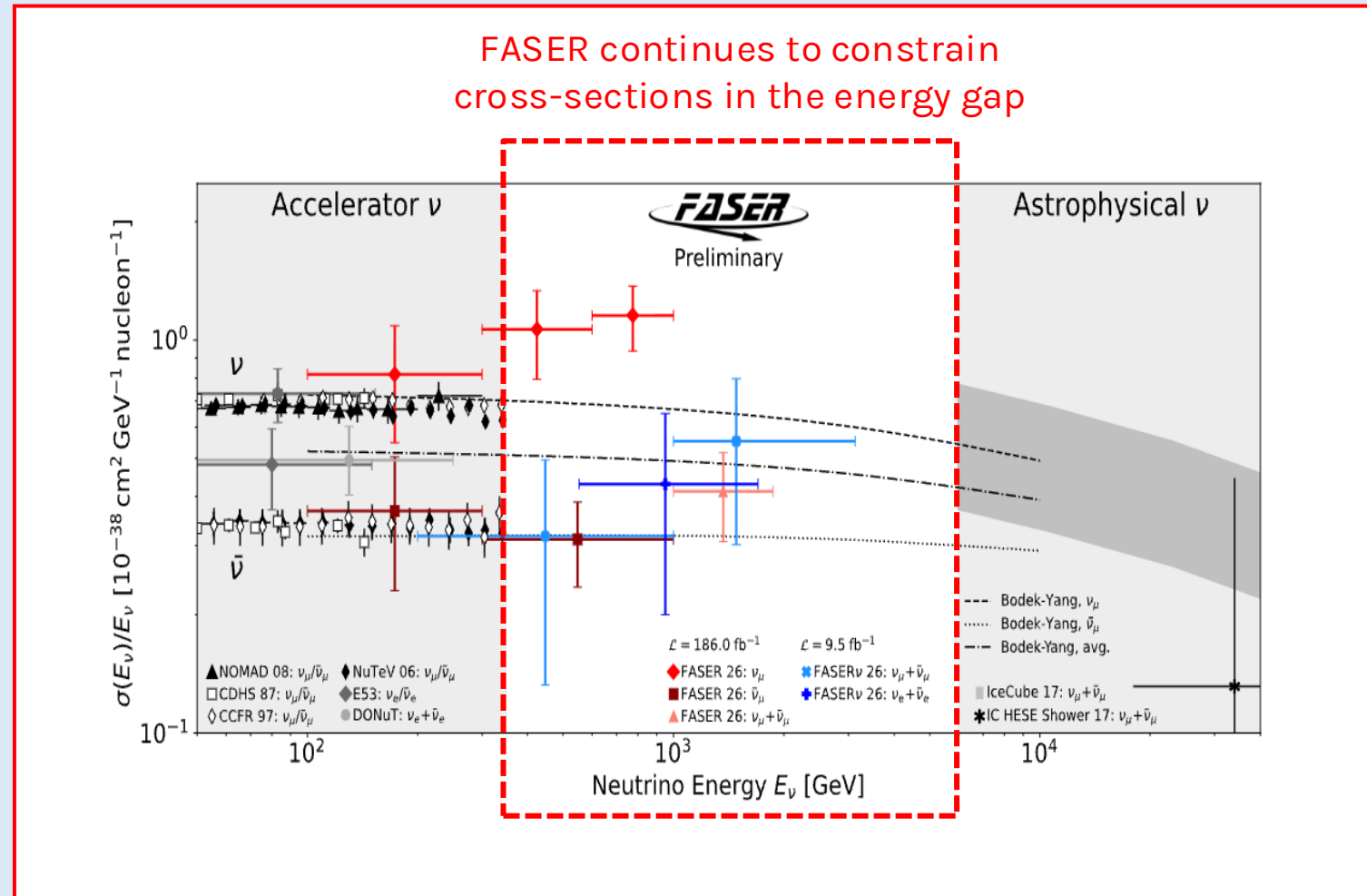
Focus of This Talk

Electronic Detector

- Faster** to analyze
 - Uses only electronic components of detector
 - Benefits from prompt data reconstruction
- Can **separate** neutrino decays by **charge**
 - Can differentiate neutrino and anti-neutrino

Neutrinos @ FASER

- Highest energy man-made neutrinos
- Produced through **light/charm decays** due to high flux of light hadrons along LOS
- Exist in the **energy gap** between neutrinos from fixed-targets and astrophysical processes
- **First direct detection by FASER (2023)**



[First Collider Neutrino Detection Paper](#)

Muon Neutrinos

Neutrinos are **produced upstream of FASER** :

- Through both light & charm **hadron decays**
- High flux in forward direction along the LOS

ν_μ **first** flavour to be studied using **only FASER's electronic detector**

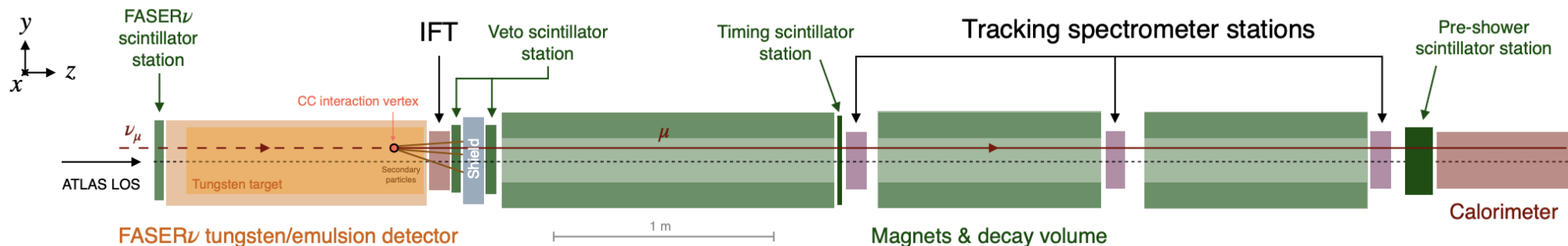
Signal

- **Interactions** occur in the **tungsten target** of FASER ν
- Resulting in a **highly collimated muon track**
- Measured by FASER's **electronic sub-detectors**

Background

Main source of background due to:

- High momentum muons that miss veto scintillators



Muon Neutrino

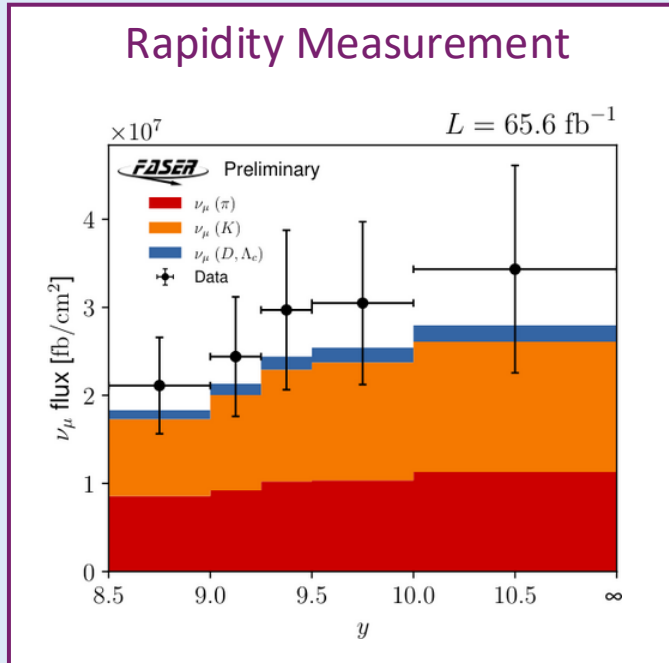
Previous Measurements

Previous measurements of ν_μ with FASER's electronic detector include:

- ν_μ flux and differential cross section measurement
- Rapidity measurement

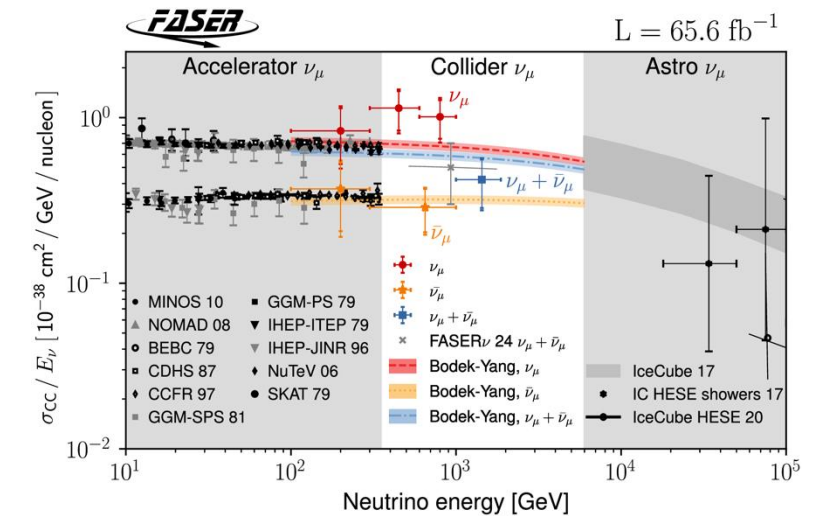
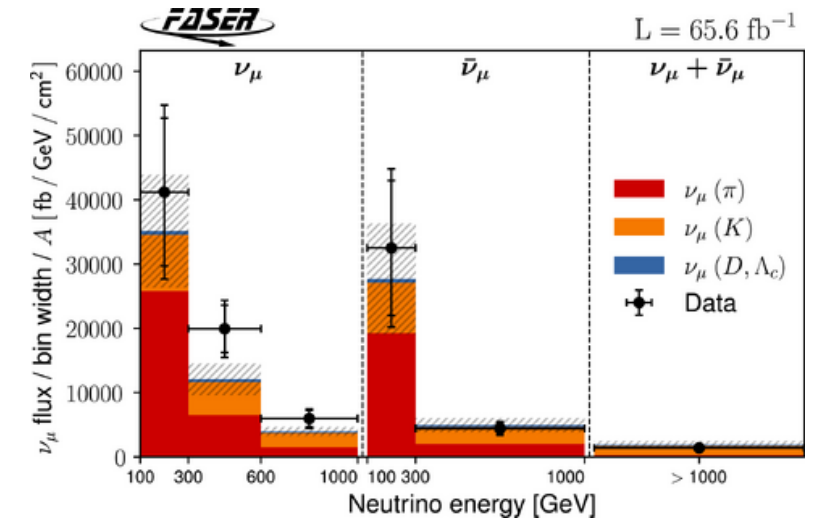
All previous measurements :

- **Measured one variable**
- Constrained the others using **MC predictions**
- Example: measuring the cross section with a fixed neutrino flux



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Flux and Differential Cross Section Measurement



CERN-EP-2024-309

Muon Neutrino Results

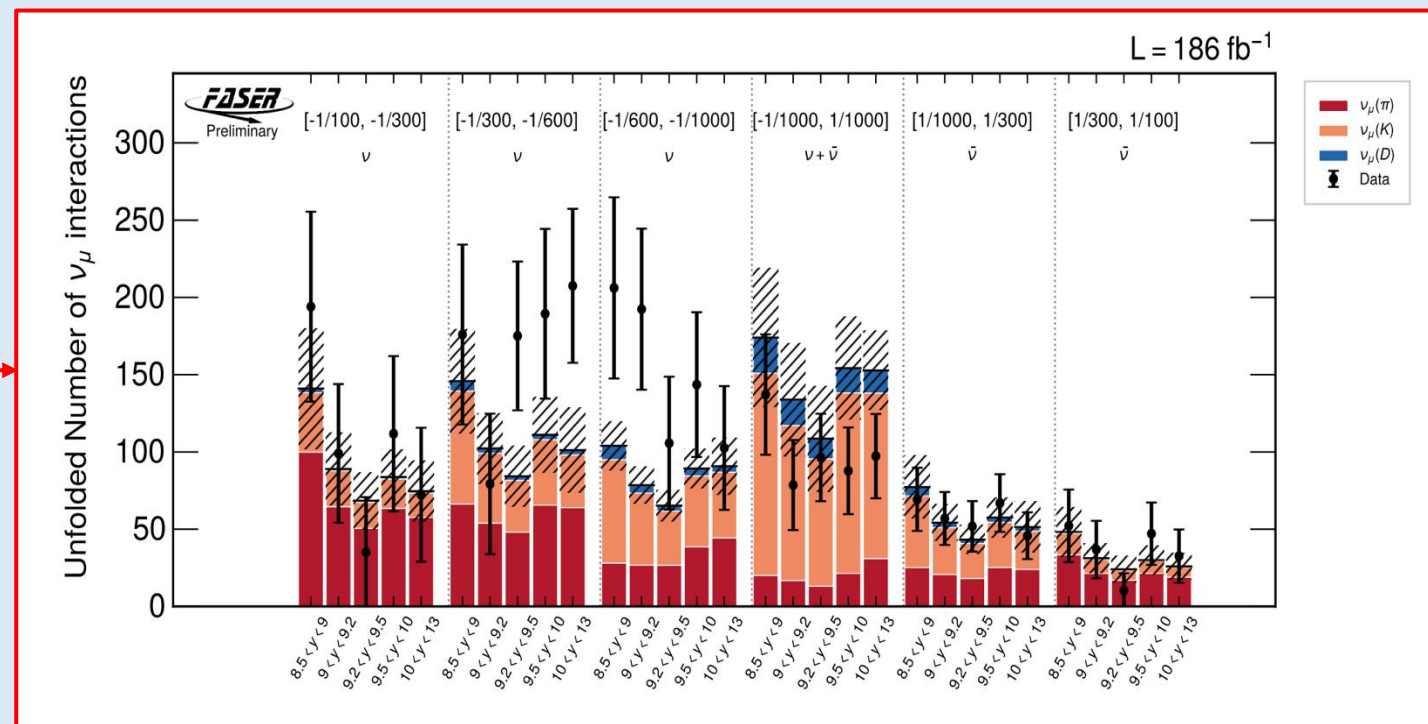
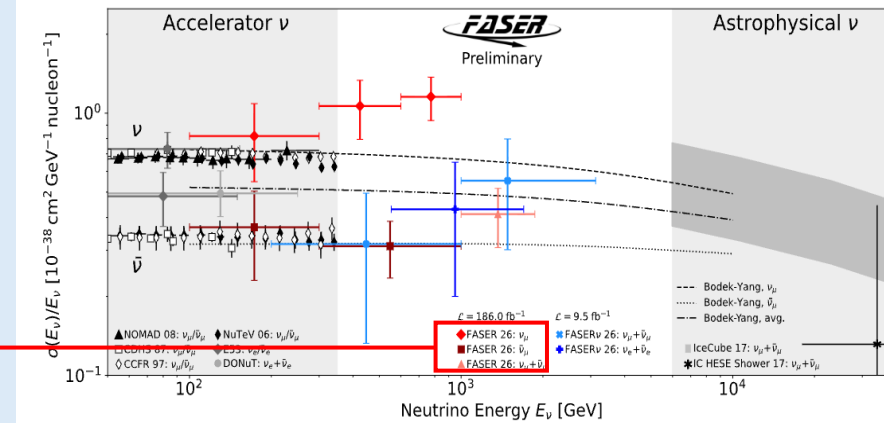
New Measurement:

- Combines previous measurements to extract a double-differential cross section
- Expressed as a function of both in energy and rapidity

Events observed :

766.8 ± 28.7 (stat.) ± 7.3 (syst.)

Result complements and expands upon measurements taken using FASER ν

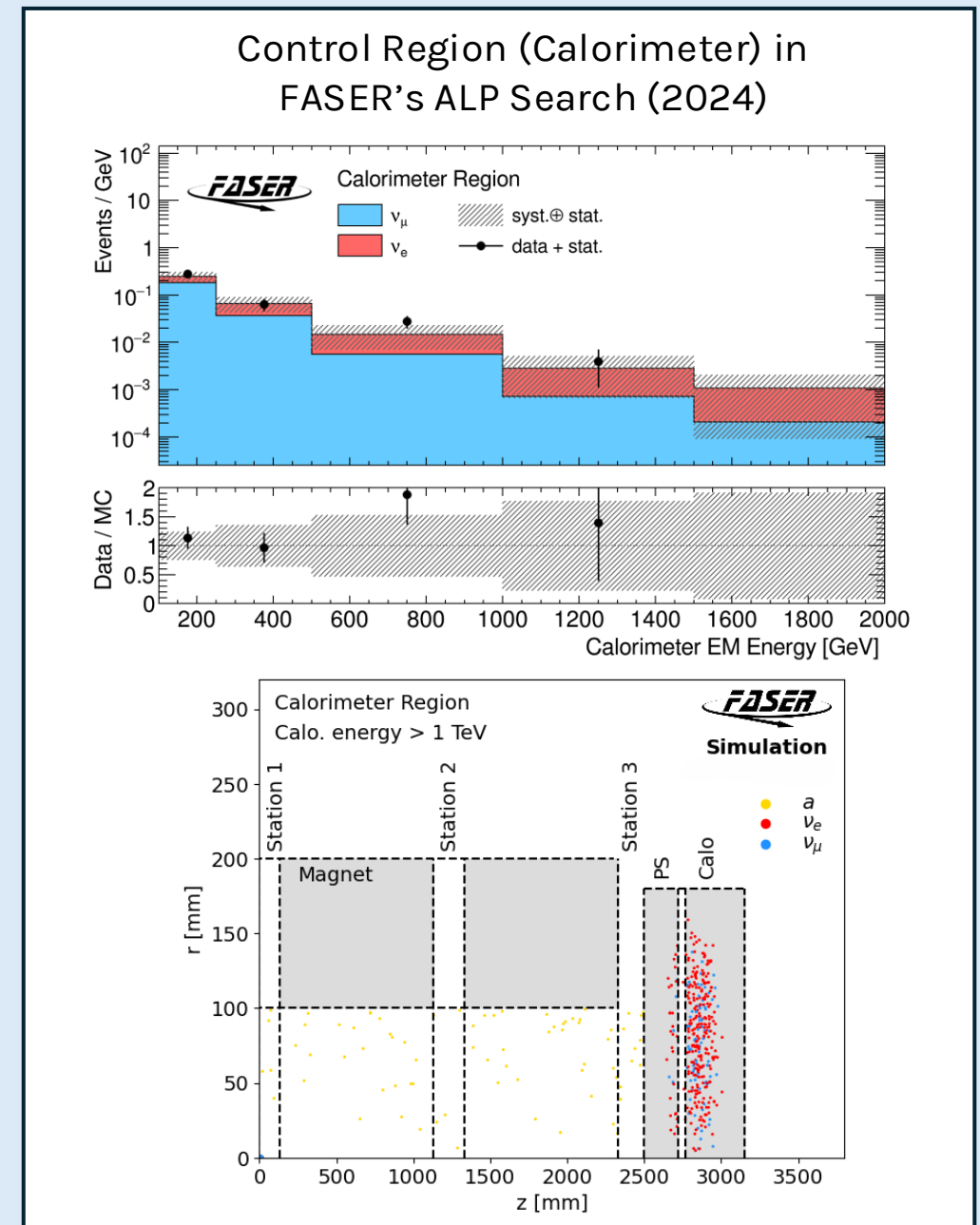


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Electron Neutrinos

- While ν_μ interactions are most abundant, other flavours are also able to be detected
- Control regions of previous BSM search for axion-like particles (ALPs) dominated by ν_e interactions
- Demonstrated potential to measure ν_e with the electronic detector
- This analysis exploits these control regions to isolate ν_e interacting in the detector

ALPS Paper



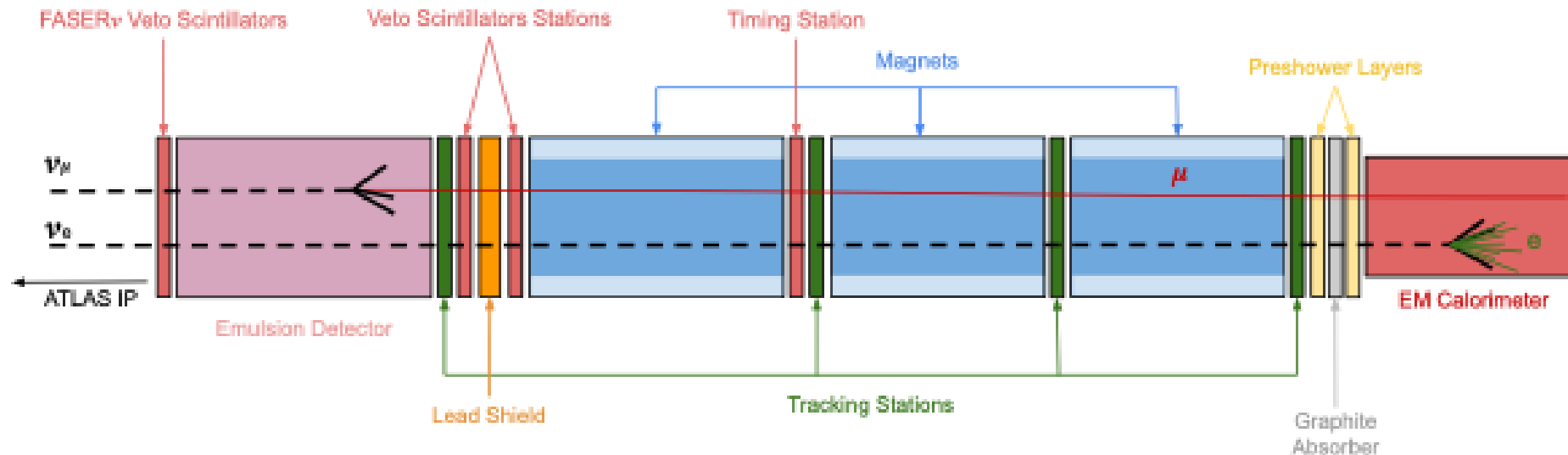
Electron Neutrinos

Event Selection

- 176.8 fb^{-1} of data
 - Collected by FASER in 2022-2024 (inclusive)
 - **Only** uses data from **electronic spectrometer**
 - **Targets** ν_e produced in both charged-current (CC) and neutral-current (NC) interactions in the calorimeter

Signal

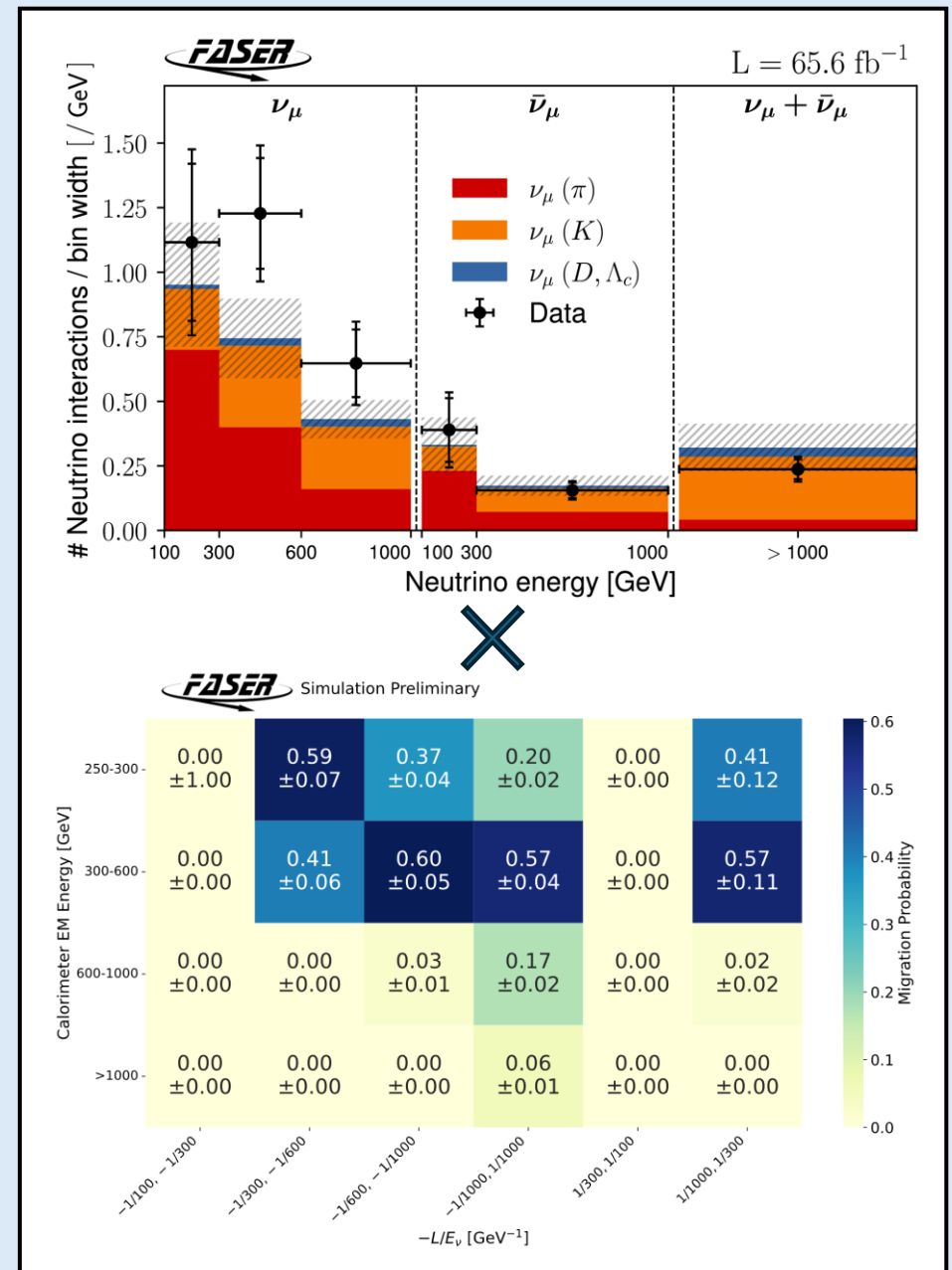
- **No charge** in either **veto** or **timing scintillators**
- **No signal** in **preshower** scintillator
- Large **energy deposit** in **calorimeter** ($>250 \text{ GeV}$)



Electron Neutrinos

Background Estimation

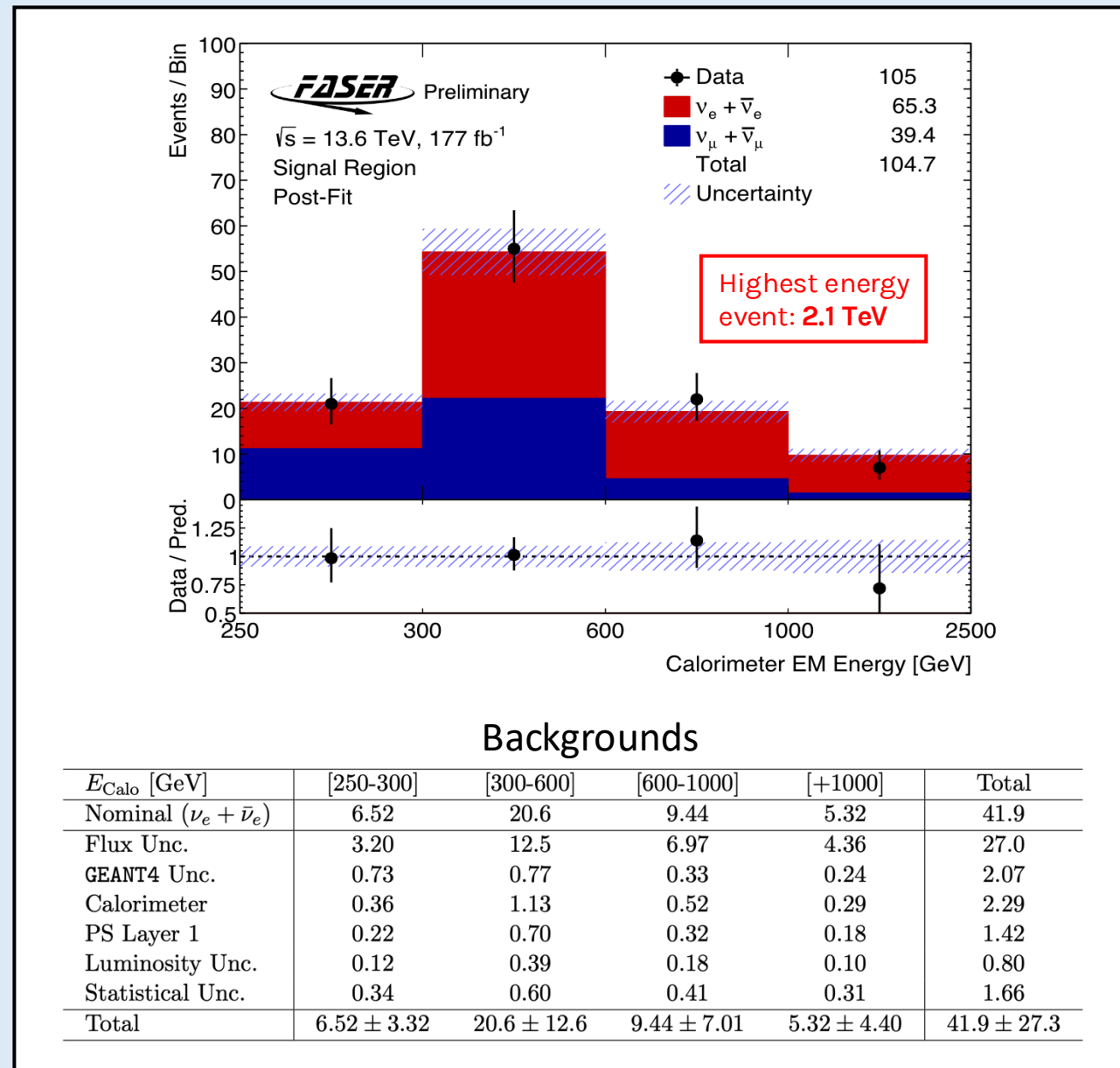
- Dominate background is muon neutrinos
- ν_μ interactions deposit less energy in calorimeter than ν_e interactions
 - Introducing a **250 GeV energy requirement** in the calorimeter significantly **suppresses the muon neutrino background**
- Using previous ν_μ cross-section measurement:
 - Background can be further constrained using data
 - Requires migration-matrix to obtain ν_μ energy from data



Electron Neutrinos Results

- **Observation** of electron neutrinos in FASER's EM calorimeter
 - Significance of 5.5σ
- Observed : $65 \pm 12 \nu_e$ (CC +NC) events
- Expected : $42 \pm 27 \nu_e$ (CC+ NC) events

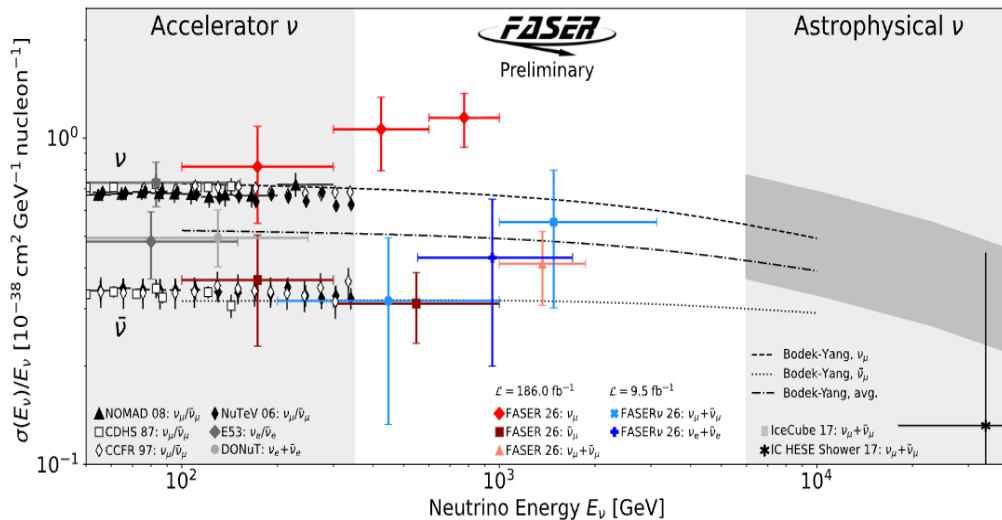
First measurement of ν_e using FASERs electronic detector !



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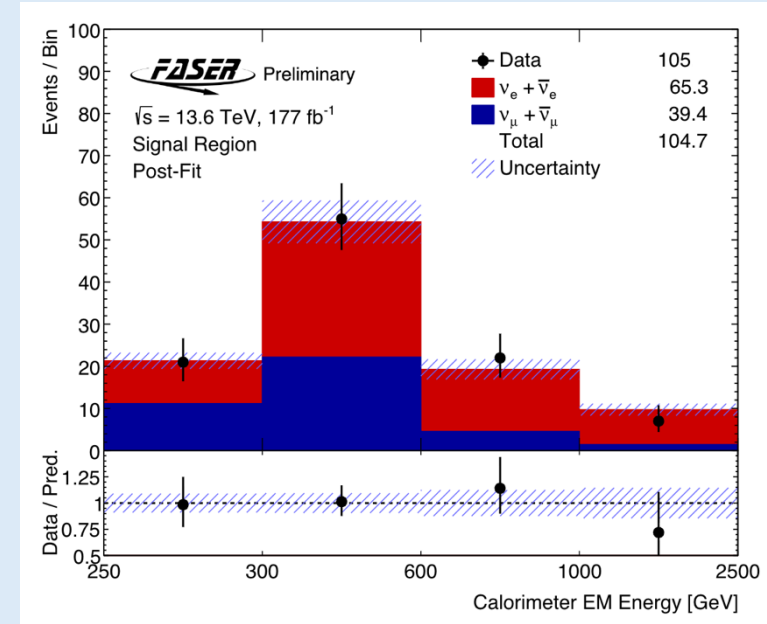
Summary

- **New measurement of muon neutrinos** with FASER's electronic detector
- **First observation of electron neutrinos achieved** solely with FASER's electronic sub-detectors
- Together with measurements from FASER's emulsion detector, this work further bridges the gap between neutrinos from fixed-target experiments and astrophysical processes

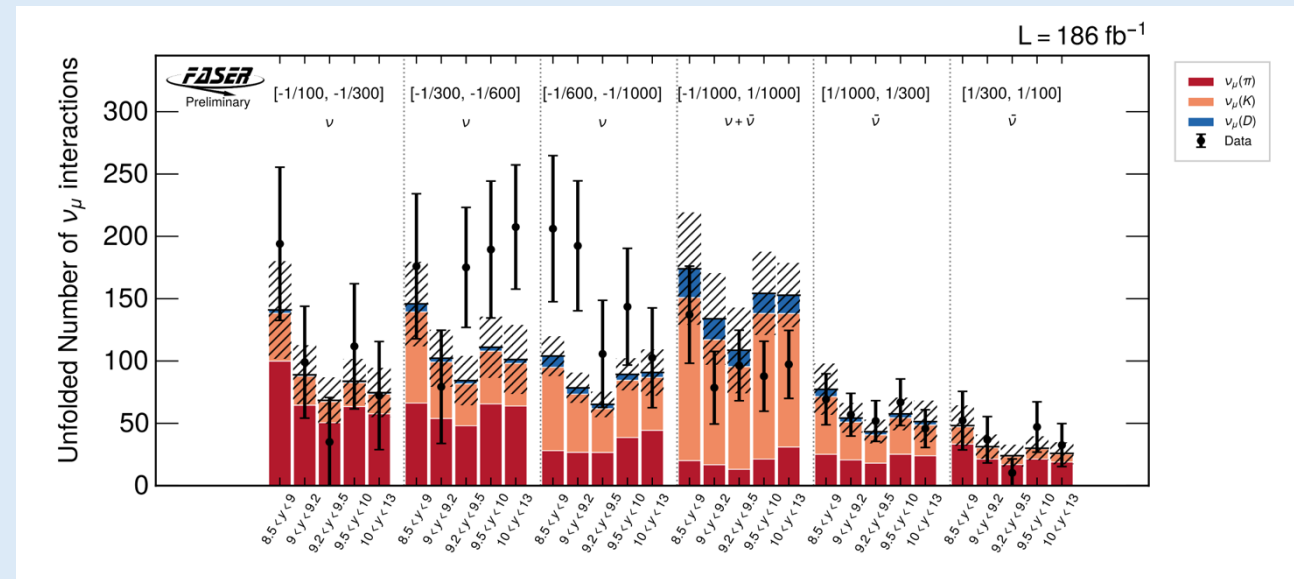


08/04/2026

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Back up

Background Estimation

Additional backgrounds include:

- Large angle muons
 - Only account for a small number of events $< 250\text{GeV}$
 - Negligible $> 250\text{ GeV}$
- Neutral hadrons
 - Negligible with signal selection applied
- Veto system inefficiency
 - Negligible
- Non-collision backgrounds
 - Beam background does not correspond to colliding bunches
 - No events observed in cosmic muon data
 - i.e. Negligible
- Tau Neutrinos
 - Expected contribution is very small

E_{Calo} [GeV]	[250-300]	[300-600]	[600-1000]	[+1000]	Total
Nominal ($\nu_e + \bar{\nu}_e$)	6.52	20.6	9.44	5.32	41.9
Flux Unc.	3.20	12.5	6.97	4.36	27.0
GEANT4 Unc.	0.73	0.77	0.33	0.24	2.07
Calorimeter	0.36	1.13	0.52	0.29	2.29
PS Layer 1	0.22	0.70	0.32	0.18	1.42
Luminosity Unc.	0.12	0.39	0.18	0.10	0.80
Statistical Unc.	0.34	0.60	0.41	0.31	1.66
Total	6.52 ± 3.32	20.6 ± 12.6	9.44 ± 7.01	5.32 ± 4.40	41.9 ± 27.3