

Phenomenological Modelling of Light Dark Matter with Optically Levitated Sensors

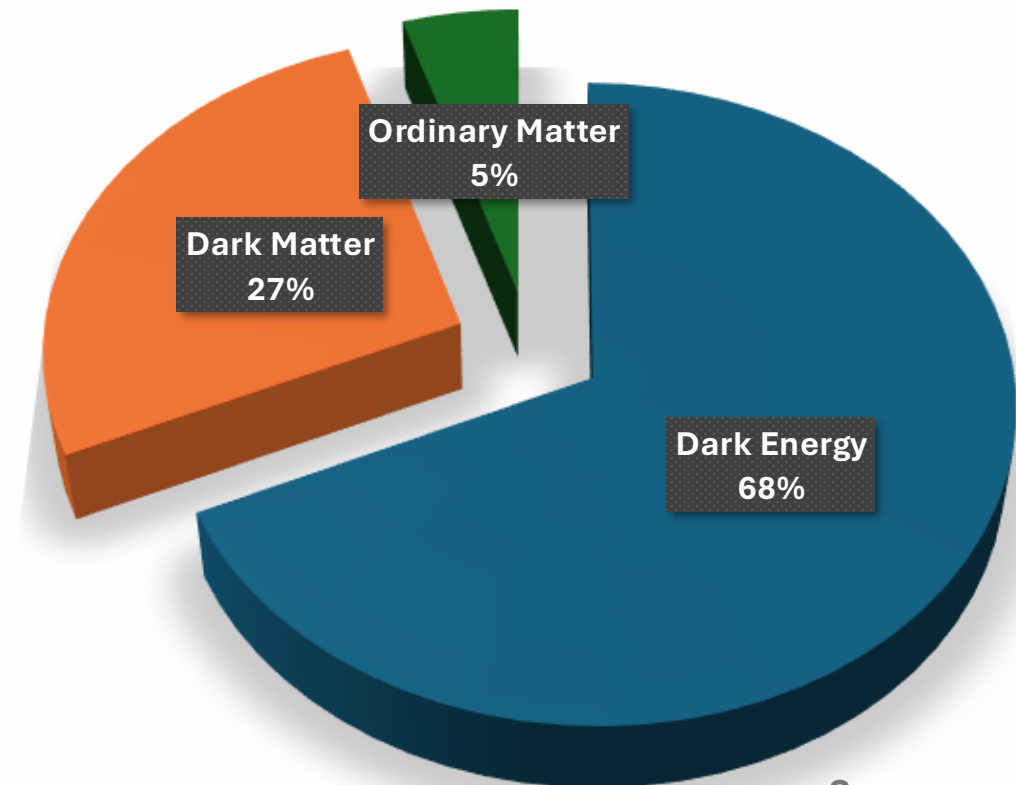
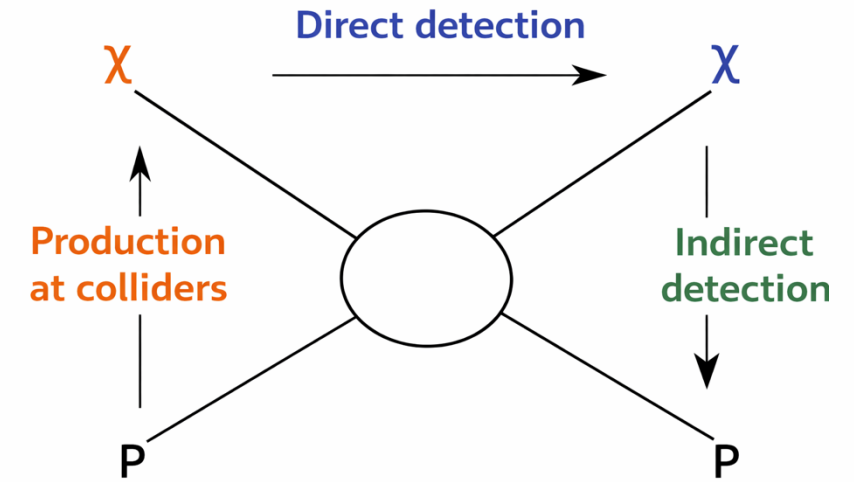
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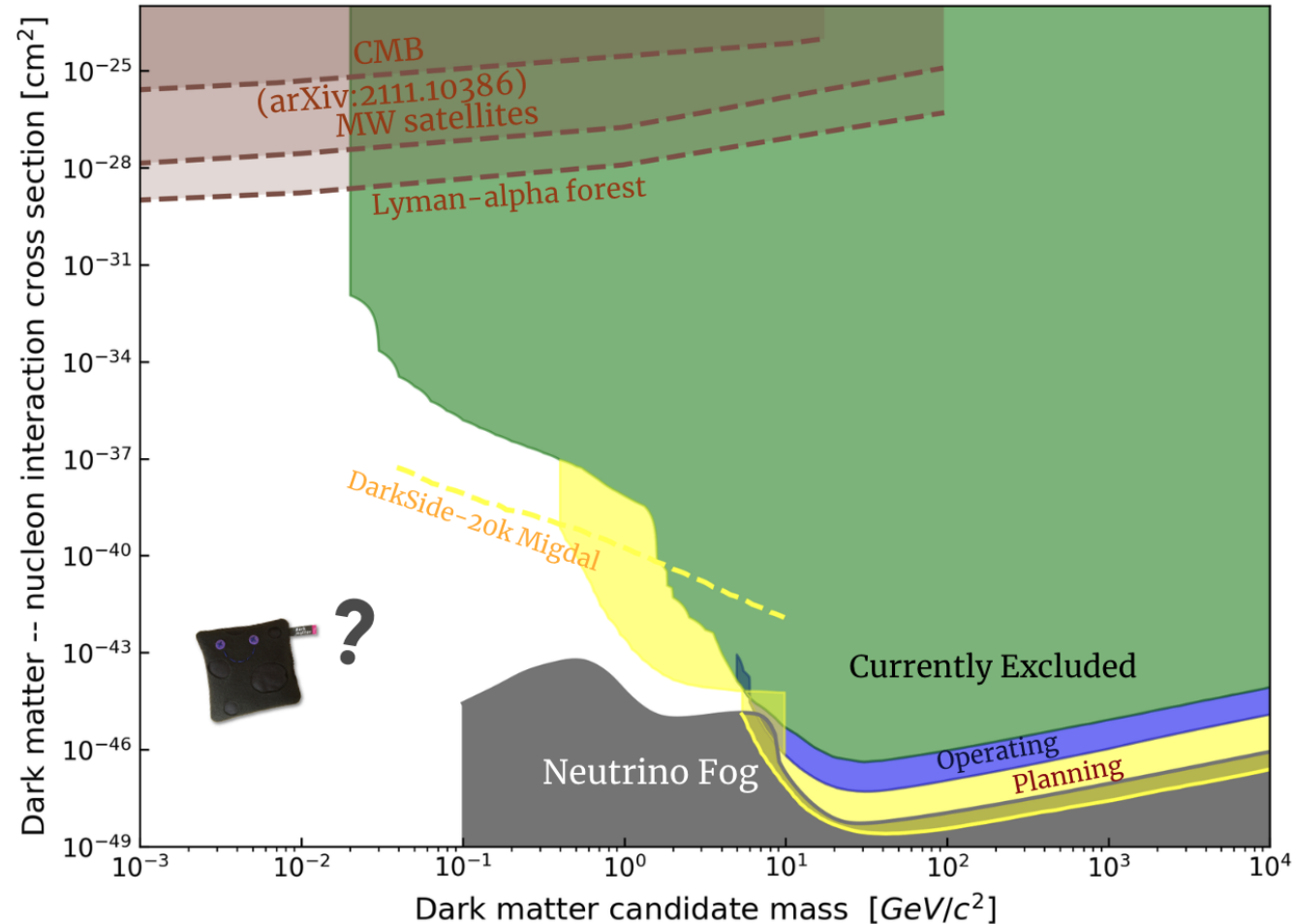
Dark matter

- Most of the matter in the Universe is **non-luminous** and **non-baryonic**
- Evidence from multiple independent probes:
 - **Galactic dynamics**
 - **CMB anisotropies**
 - **Structure formation**
- Viable candidates include:
 - Axions / dark photons / fuzzy dark matter
 - **WIMPs** and other particle candidates
 - MACHOs / primordial black holes



Beyond the WIMP Paradigm

- WIMPs have been the leading dark matter paradigm for decades
- Direct detection now probes deeply into the parameter space above \sim GeV masses, nearing the **neutrino floor**
- The **absence** of a signal places increasing **pressure** on conventional **WIMP** parameter space
- This motivates a broader experimental program targeting **lower-mass** dark matter with new detection strategies



Optomechanically levitating nanoparticles

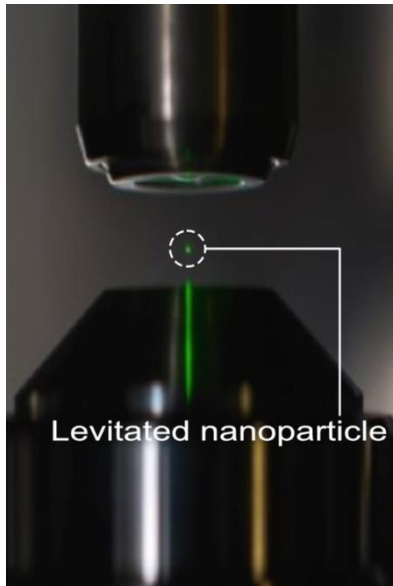


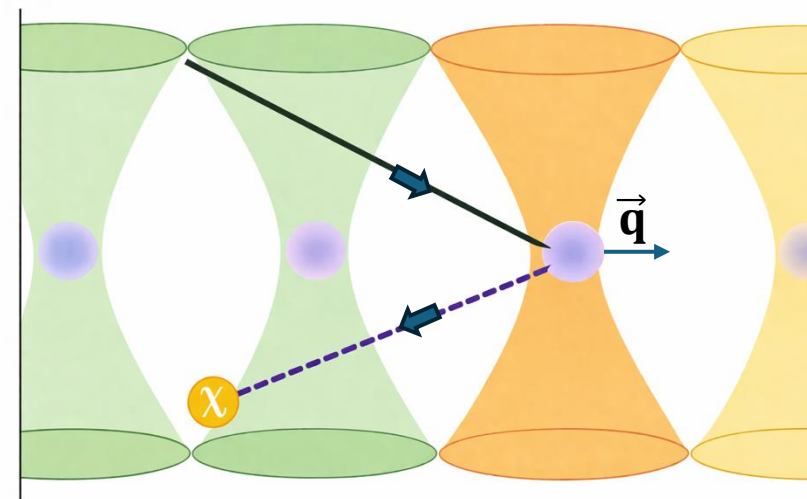
Image courtesy of:
Jayadev Vijayan

- Optical traps **levitate** particles with masses of order **femtograms** and radii of a few to a few hundred nanometres
- Continuous position readout with **nanometre (or better) precision**
- The system behaves as a **damped harmonic oscillator**
- Ultra-high vacuum setups, with **minimal backgrounds**

Several optomechanical levitation experiments are currently under development in **Manchester**

Key advantages for dark matter detection:

- Singal is a **momentum transfer** ($p \propto v$), not energy deposit ($E_R \propto v^2$)
- Enhanced sensitivity to low-velocity/ **low-mass dark matter**



From Detector Response to Dark Matter Limits



Detector model:

- Inputs:
 - Experimental parameters
- Output:
 - Synthetic displacement time series
 - **Impulse response template**
 - Calibrated PSD model

Time-domain reconstruction:

- Inputs:
 - Inject impulse
 - **Ringdown simulation**
 - **AR whitening**
 - **Matched filtering**
 - Peak SNR extraction
- Output:
 - **Detection efficiency**
 - **Amplitude-recovery map**
 - Null/threshold validation

Detector folding:

- Apply efficiency
- Convolve with reconstruction resolution
- Integrate over analysis window and exposure

$$N_{\text{sig}}(m_{\chi}, m_{\phi}, \alpha_n) = T_{\text{exp}} \int dq \frac{dR}{dq}(q) \epsilon(q) P_{\text{rec}}(q)$$

- **dR/dq: true recoil spectrum** from DM module
- $\epsilon(q)$: **detection efficiency** from impulse characterization
- $P_{\text{rec}}(q)$: migration into the selected reconstruction window

Theoretical Description of the Sensor

Goal: detect **impulsive momentum transfer** to the centre-of-mass mode of a levitated nanosphere.

Observable: $q \equiv \Delta p = m_s \Delta v$

With m_s as the mass of the sensor and Δv is the impulse-induced **velocity step**

Equation of motion:

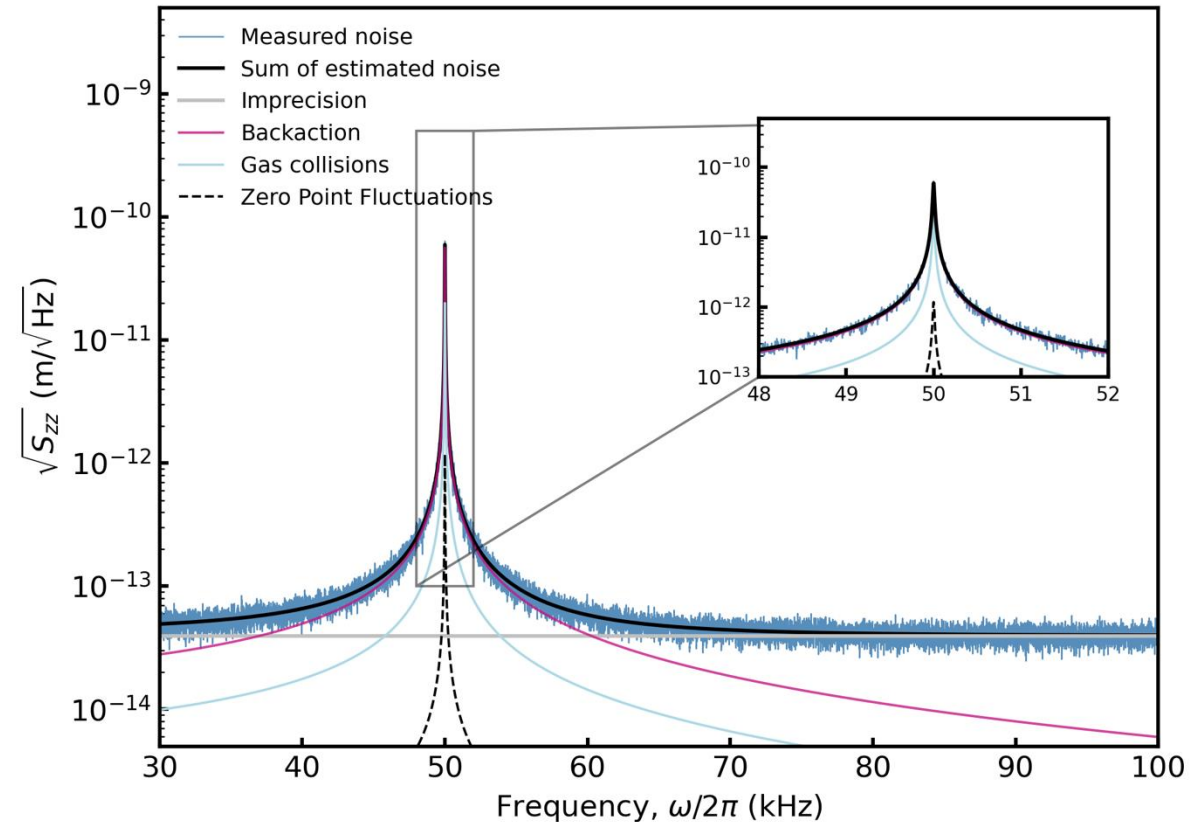
$$m_s \ddot{z} + m_s \gamma \dot{z} + m_s \omega_0^2 z = F_{\text{noise}}(t) + q \delta(t - t_0)$$

Mechanical susceptibility:

$$\chi(\omega) = \frac{1}{m_s (\omega_0^2 - \omega^2 - i\gamma\omega)}$$

Displacement-noise model:

$$S_{zz}(\omega) = |\chi(\omega)|^2 (S_{FF}^{\text{ba}} + S_{FF}^{\text{gas}}) + S_{zz}^{\text{imp}} + S_{zz}^{\text{zpf}}$$



- Silica nanosphere: $m_s \simeq 4.8$ fg, $r \simeq 166$ nm
- Centre-of-mass z-mode at ~ 50 kHz
- Readout efficiency: $\eta \sim 0.1$
- Near resonance, the noise is backaction-dominated, with subleading gas contribution

Time-Domain of the Sensor

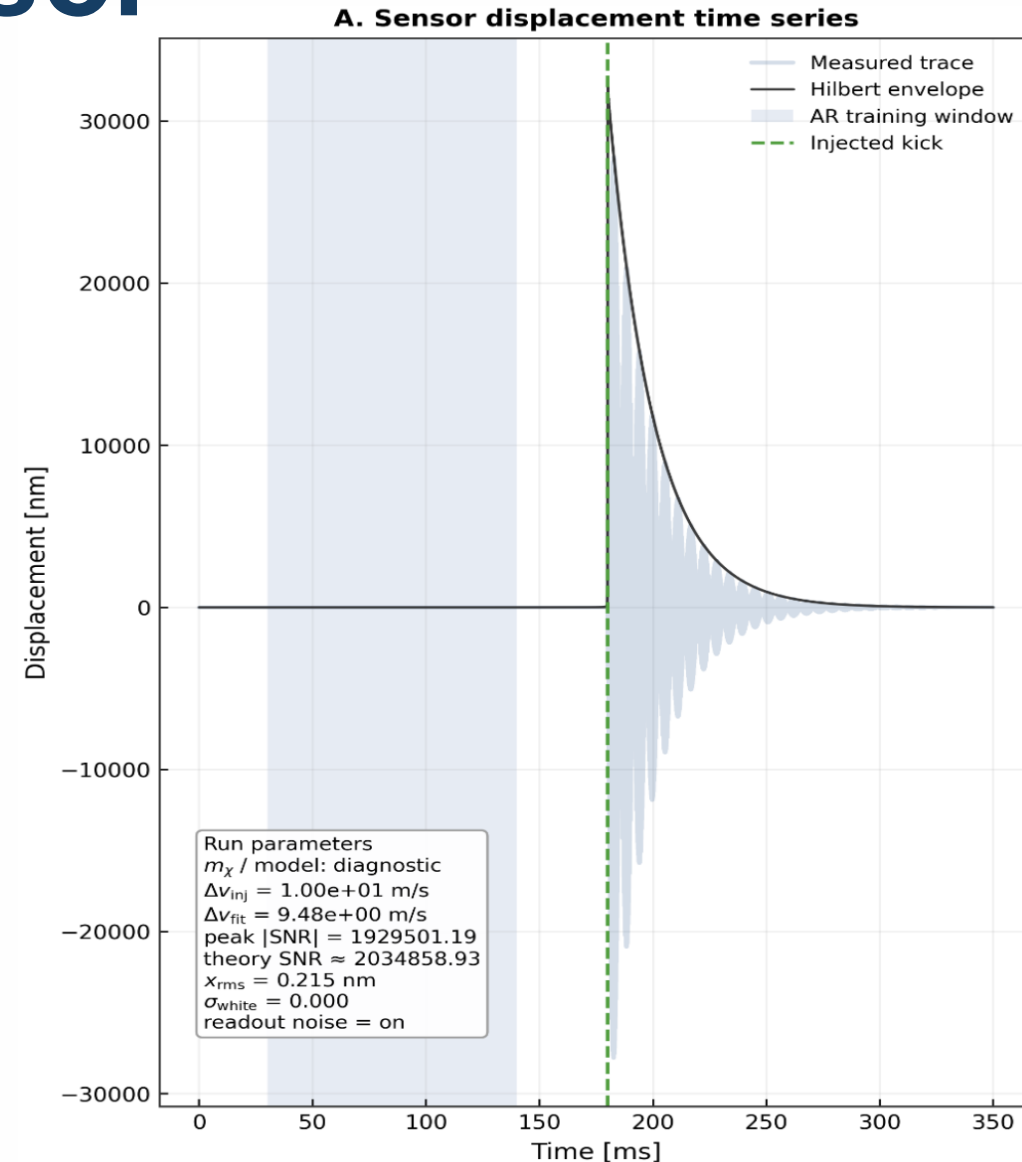
Equation of motion (Langevin form):

$$m_s \ddot{z} + m_s \Gamma \dot{z} + m_s \Omega_0^2 z = F_{\text{th}}(t) + F_{\text{ba}}(t) + F_{\text{tech}}(t) + q \delta(t - t_0)$$

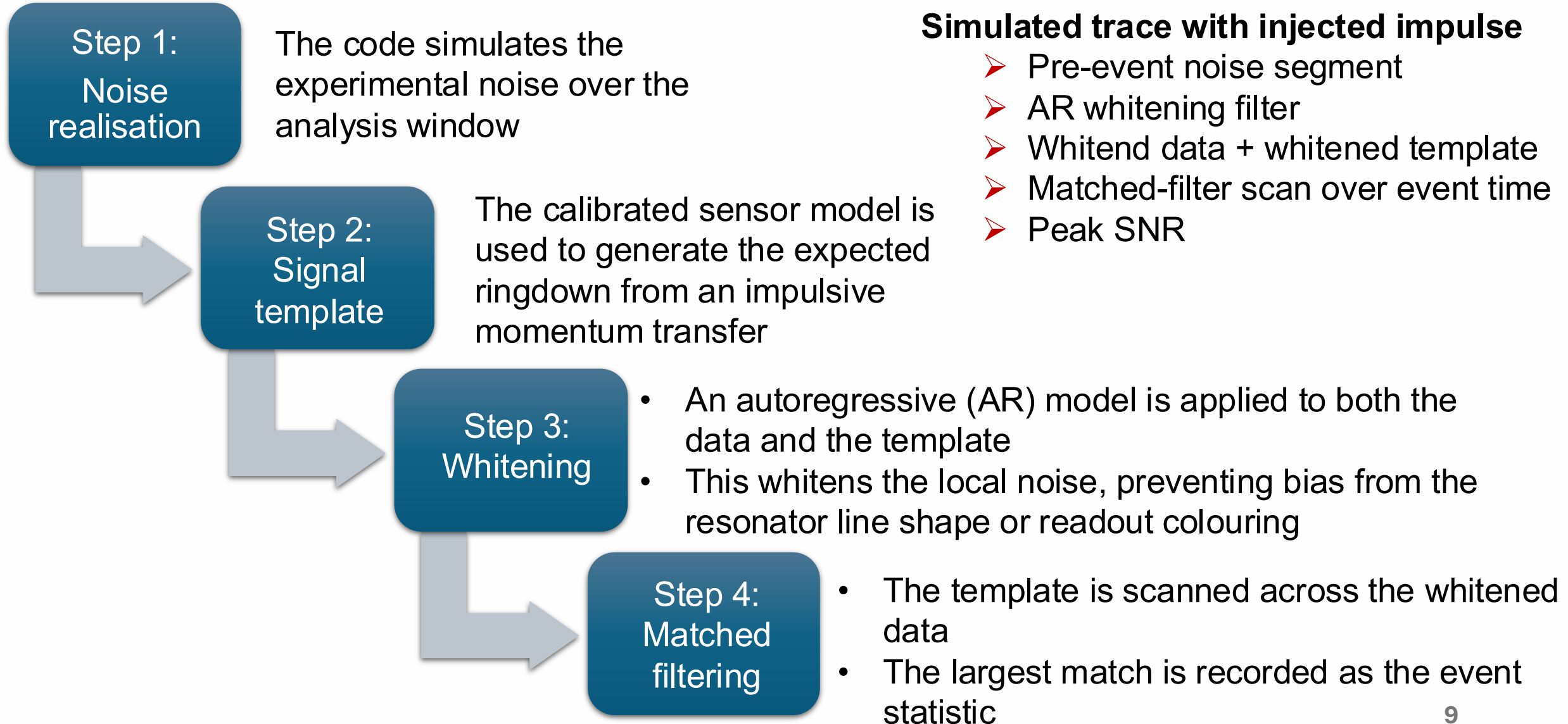
- m_s : sensor mass
- Ω_0 : resonance frequency of the readout mode
- Γ : effective linewidth
- $q = \Delta p = m_s \Delta v$ $\dot{z}(t_0^+) - \dot{z}(t_0^-) = \Delta v$ $\omega_d = \sqrt{\Omega_0^2 - \Gamma^2/4}$

$$z_{\text{sig}}(t) = \Theta(t - t_0) \frac{q}{m_s \omega_d} e^{-\Gamma(t-t_0)/2} \sin[\omega_d(t - t_0)]$$

The readout mode is modelled as a noisy **damped harmonic oscillator**. A dark-matter interaction is represented by an **instantaneous momentum transfer q** , equivalently a velocity step $\Delta v = q/m_s$. The observable is the resulting **ringdown**, simulated in the time domain with stochastic force noise and readout imprecision, and used as the template for **matched-filter reconstruction**.



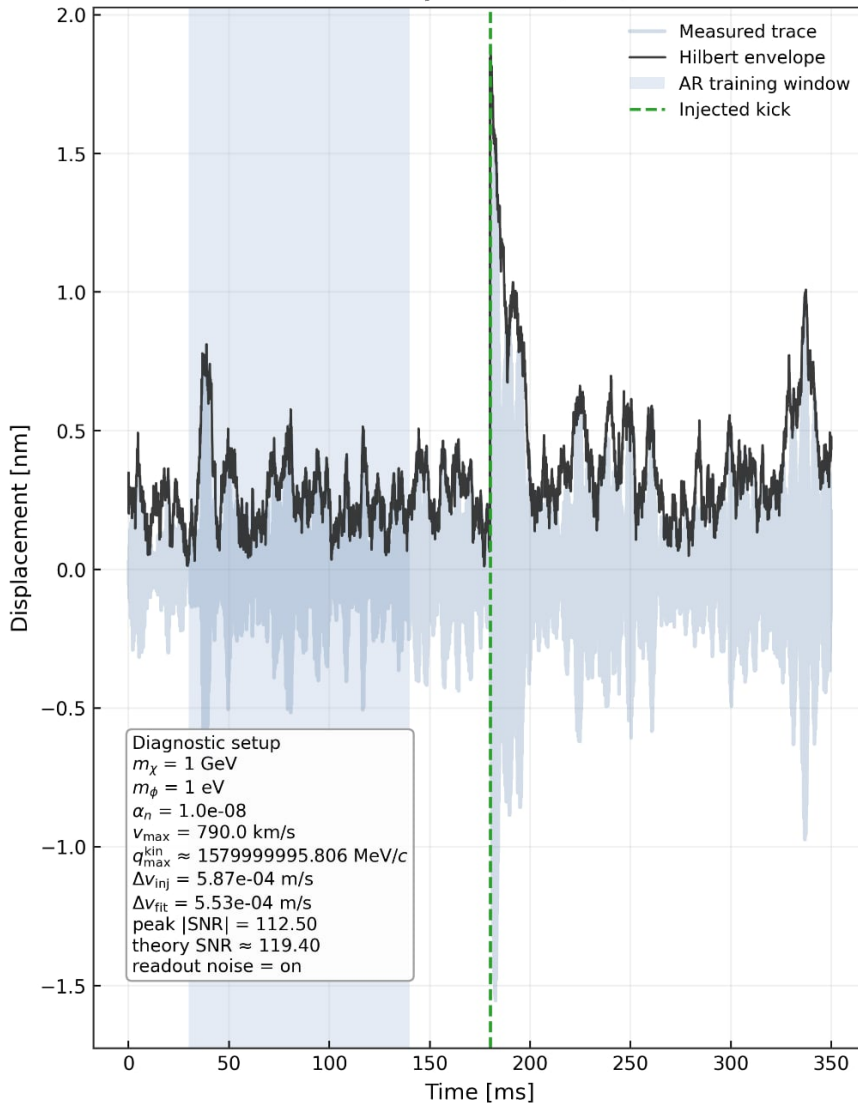
Impulse Reconstruction Algorithm



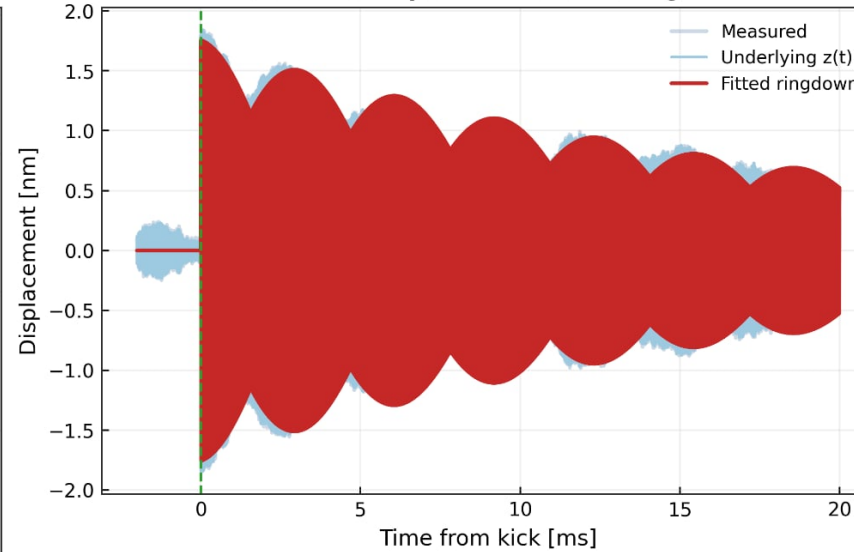
Impulse reconstruction

Dark-matter impulse detection verification at the halo-speed kinematic limit

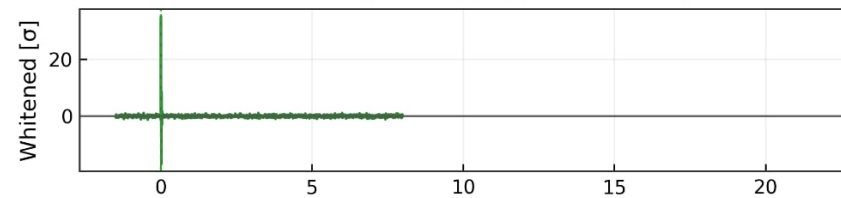
A. Sensor displacement time series



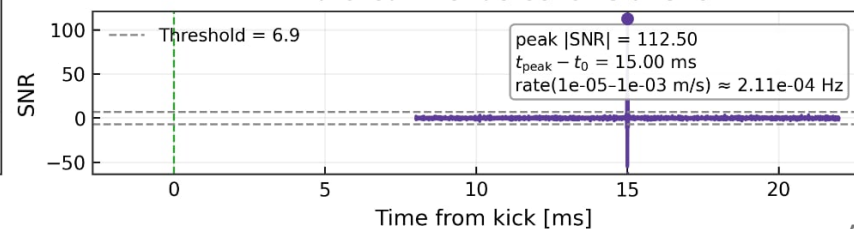
B. Zoomed response with fitted ringdown



C. Whitenened time series near the event



D. Matched-filter detection statistic



Simulated trace with injected impulse

- Pre-event noise segment
- AR whitening filter
- Whitenend data + whitenend template
- Matched-filter scan over event time
- Peak SNR

Validation of the Reconstruction Pipeline

Plot A: Representative null matched-filter trace

- Even in the absence of an injected signal, local peaks are present
- The selected peak remains well below the analysis threshold, SNR = 6.9

Plot B: Distribution of the null statistics

- Distribution of the maximum matched-filter statistics obtained from each null trial
- Among the 10000 trials, none exceeds the analysis threshold

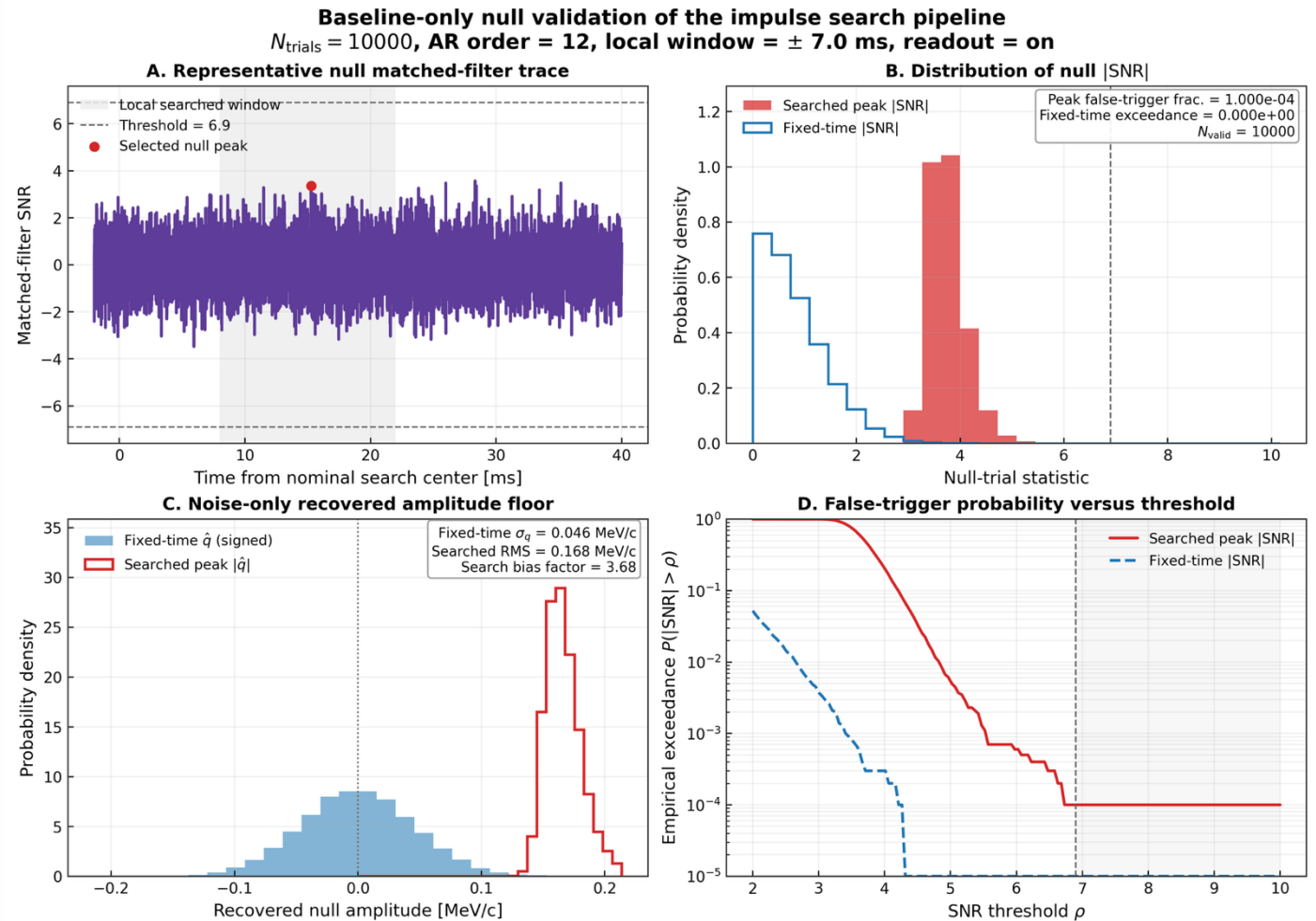
Plot C: Noise-only recovered amplitude

- **Blue histogram:** reconstructed amplitude at a fixed time
- **Red histogram:** magnitude of the best-fit amplitude for the largest match

Plot D: False-trigger probability

- Exceedance probability for the fixed-time and search-peak statistics
- The SNR threshold is way above the observed null background

When scanning a finite local time window, how large is the maximum fluctuation expected under the null hypothesis?



Yukawa-Mediated Dark Matter Levitated Sensors

Interaction Model: $\mathcal{L} \supset g_\chi \bar{\chi}\chi\phi + g_n \bar{n}n\phi, \quad \alpha_n \equiv \frac{g_\chi g_n}{4\pi}, \quad \lambda = \frac{\hbar}{m_\phi c}.$

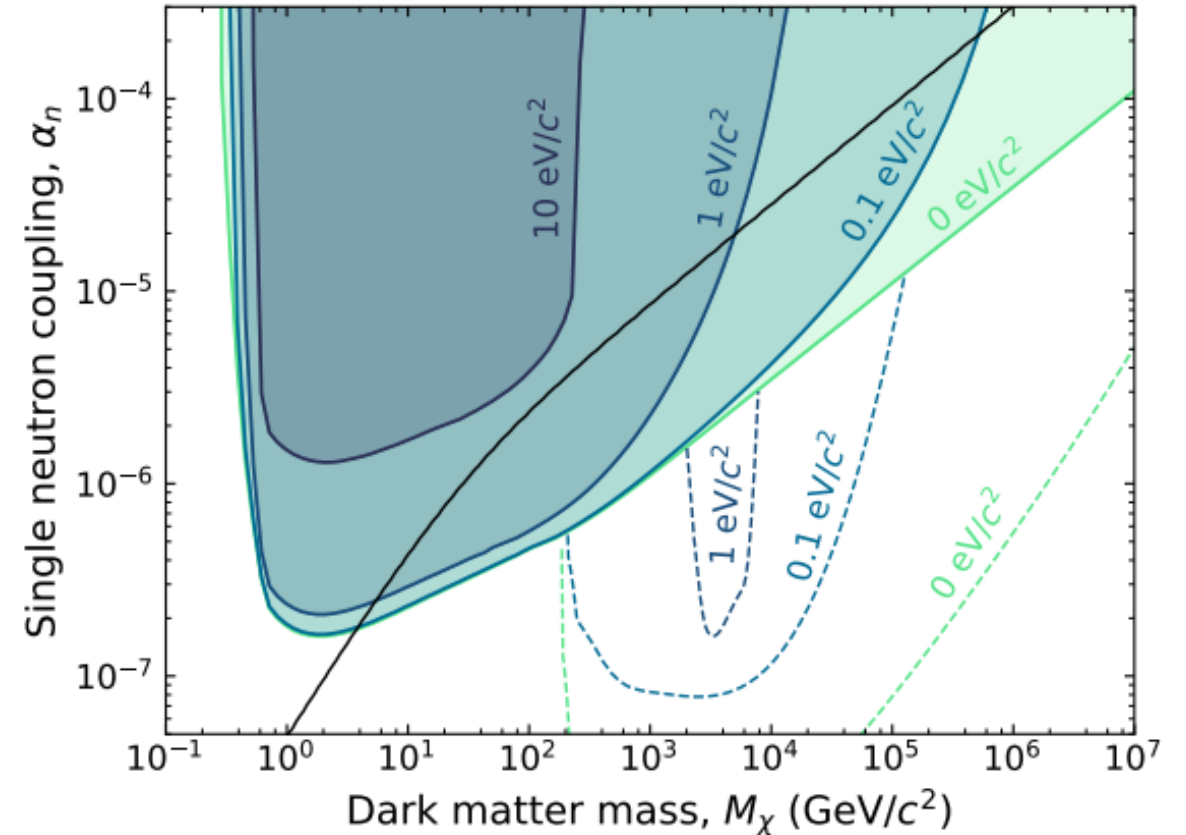
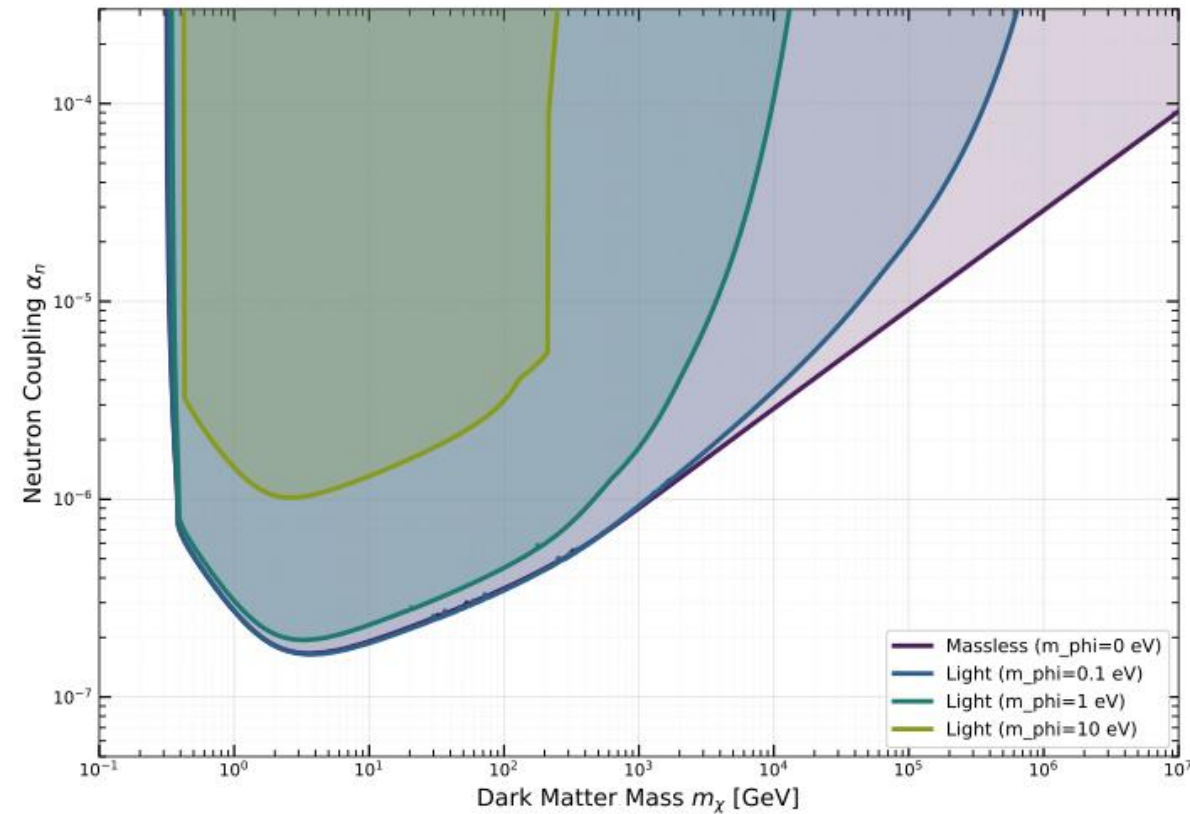
This gives a Yukawa potential: $V(r) \propto \alpha_n \frac{e^{-r/\lambda}}{r}$

Sensitivity to the new "fifth force":

- Dark matter particle scatters on the sensor via a light mediator
- Coherent nanosphere response $\rightarrow N^2$ enhancement
- Impulse-sensitive readout \rightarrow Strong reach for soft, low-q scattering
- Classical scattering regime, not Born
- $dR/d(\Delta v) \otimes \varepsilon(\Delta v) \rightarrow$ limits on α_n

The following results correspond to a **specific** experimental configuration, used here for **validation** and **exploratory** studies. The **modularity** of the framework will allow future extensions aimed at identifying the **optimal setup** for such searches.

Computational Reproduction of the Experimental Yukawa Limits



Near-quantitative agreement between the **computational reconstruction** (left) and the published **experimental exclusion curves** (right).

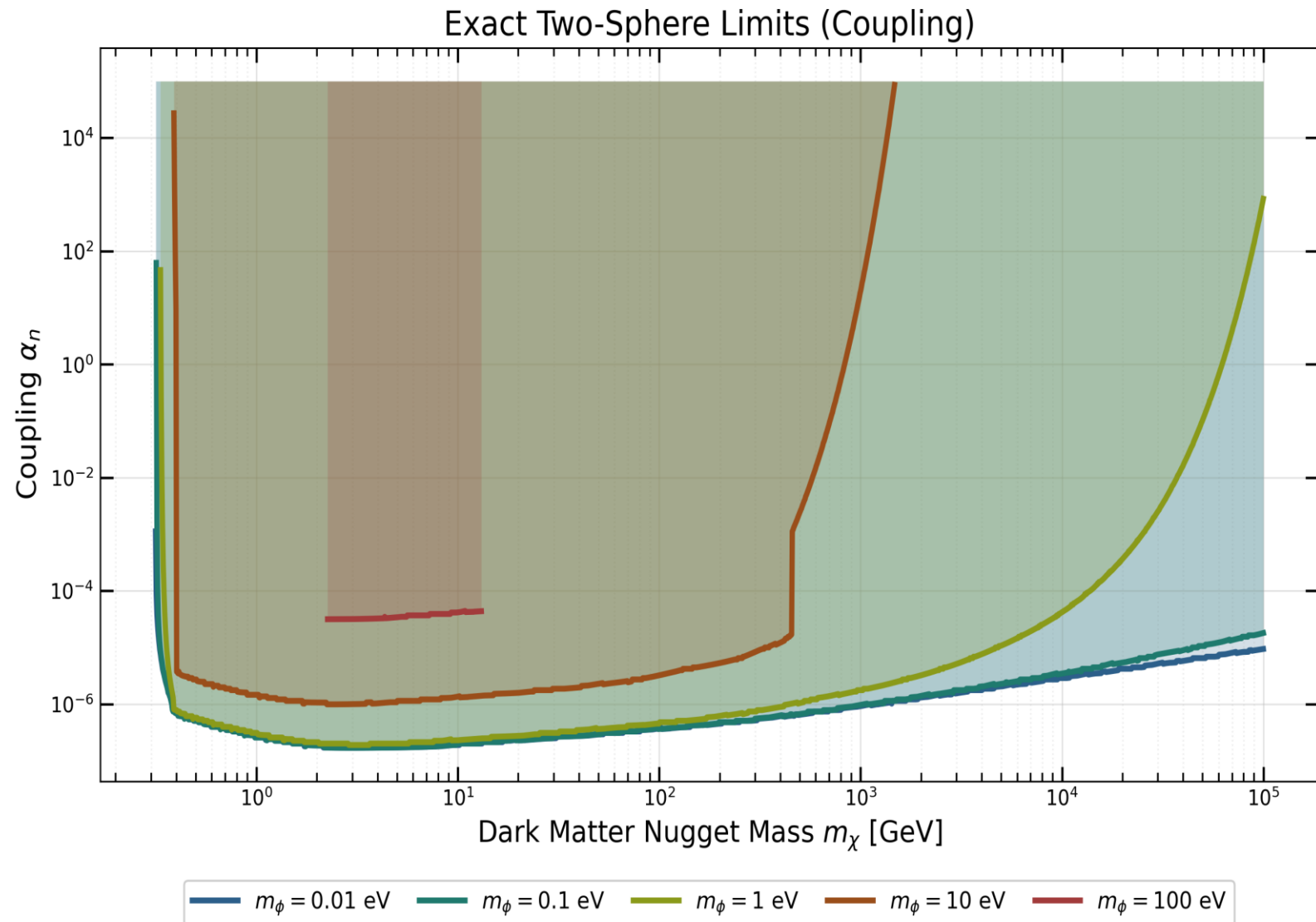
Tseng et al., *Search for Dark Matter Scattering from Optically Levitated Nanoparticles*, **PRX Quantum** **6** (2025) **040367**, arXiv:2508.00815 [hep-ex]

Puffy Dark Matter coupling limits

“Puffy” DM (composite DM) refers to DM particles that are not elementary nor point like, this are **composite microscopic** bound states with a finite spatial Radius.

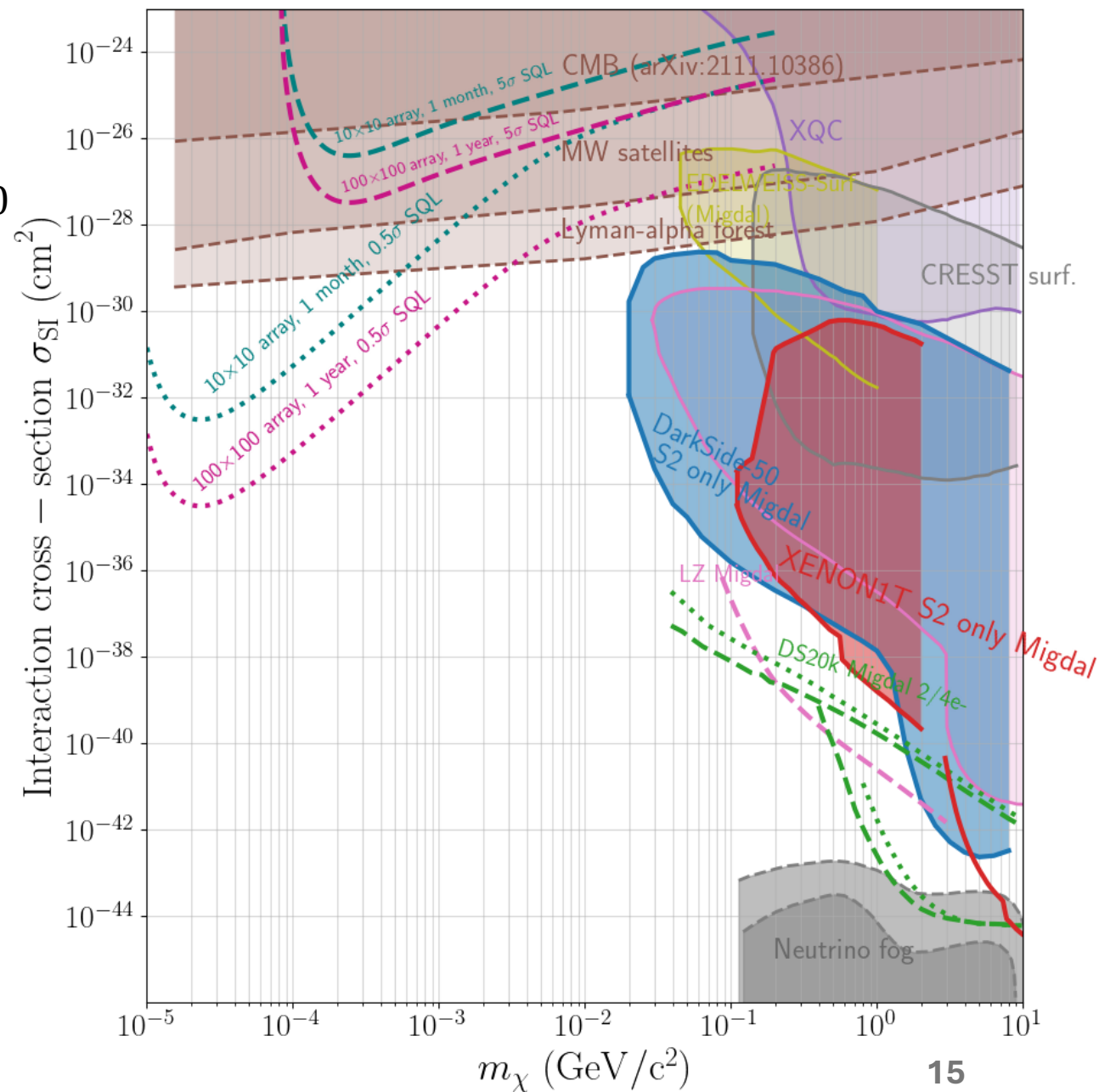
The central technical step is an exact real-space force law between two uniform spheres, obtained via a **1D Fourier–Bessel integral**, coupled to a trajectory inversion that handles the non-monotonic mapping between impact parameter and momentum transfer.

The graph presents **future experimental** constraints on the α_n coupling constant.



Levitating sensors for low SI searches

- **Theory best-case projections** for coherent low-mass DM scattering with levitated-sensor arrays: 10×10 and 100×100 sensors, shown for the corresponding benchmark exposures alongside existing DM experiments.
- Sensitivity shown for a **specific** target configuration consisting of **SiO₂ spheres** with radius of 15nm.
- **Background-free / SQL-limited benchmark:** expected backgrounds < 1 over the exposure.
- **Coherent low-mass DM scattering** extends sensitivity to very small $m_\chi < 10\text{MeV}$.
- Probes **low-mass SI** parameter space **beyond current large-scale TPCs**.
- **A further goal is to identify the optimal experimental configuration for this study.**



Conclusion

- Optically levitated sensors provide a distinctive probe of previously **unexplored** regions of low-mass dark-matter parameter space
- These sensors offer a comparatively cost-effective experimental platform
- The simulation framework developed here enables modular and efficient exploration of a broad class of dark-matter models
- It streamlines the connection between detector calibration, sensor response, and phenomenological rate predictions across different experimental configurations
- The long-term goal is to identify experimentally optimal configurations for probing well-motivated dark-matter scenarios
- **With the possible construction of the experiments by the Manchester group!**

**Thank you for listening
Any Questions?**

Near-Threshold Detection Efficiency

How does the search behave in the presence of a real signal close to the detection threshold?

Plot A: Near-Threshold detection efficiency

- For very small injected impulses, the probability of detection is small, as the signal remains below the noise floor

Plot B: Detection statistics across threshold

- The panel shows how the event statistics evolves as the true injected signal increases

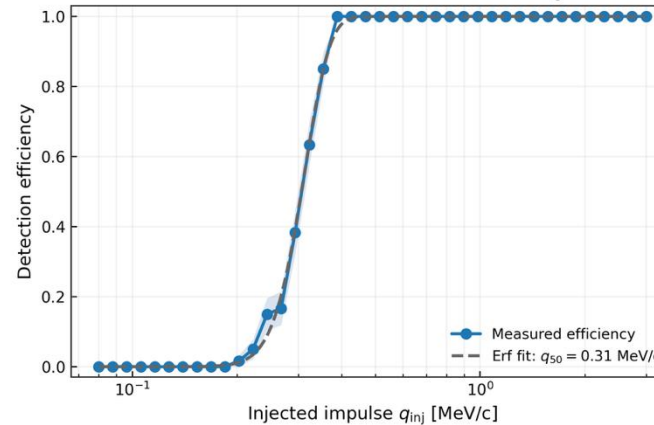
Plot C: Amplitude recovery and search bias

- Below threshold, the search often locks on the largest fluctuation, rather than the true injected signal → This produces the low-amplitude flattening and a non-zero recovered scale even under the null

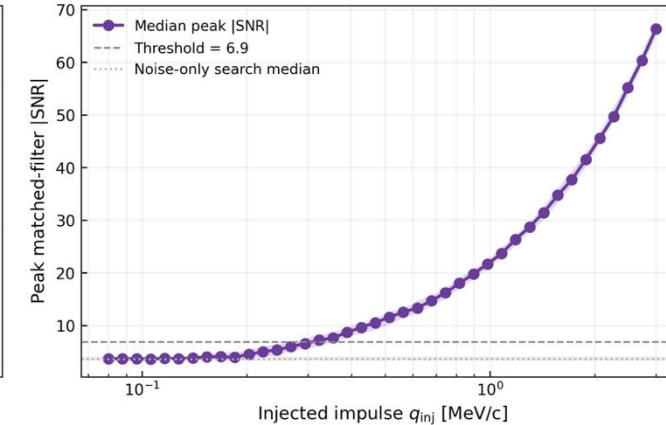
Plot D: Reconstruction width versus floor

- This panel quantifies the reconstruction width
- Once the signal is sufficiently above threshold, the 68% half-width approaches the baseline fixed time amplitude
- Above threshold, the pipeline is therefore not only efficient, but also reasonably precise

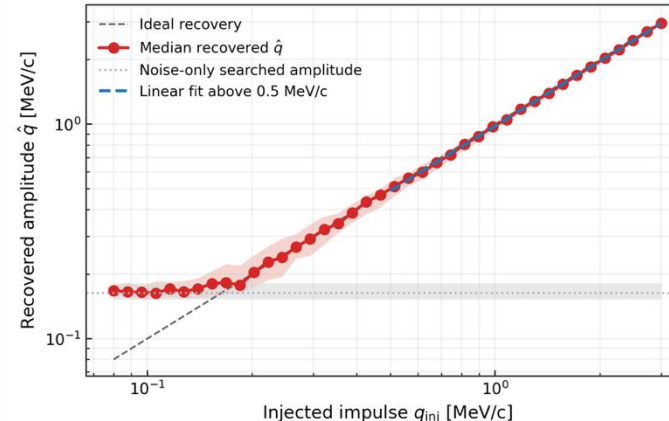
Near-threshold validation of the impulse-reconstruction pipeline
 $N_{\text{trials}} = 60$ per amplitude, $N_{\text{noise}} = 180$, AR order = 12, threshold = 6.9, readout = on
A. Near-threshold detection efficiency



B. Detection statistic across threshold



C. Amplitude recovery and search bias



D. Near-threshold width versus baseline floor

