



SRG/eROSITA results on the Western Galactic hemisphere



Miriam E. Ramos Ceja



High Energy Astrophysics and Cosmology in the era of all-sky surveys

Yerevan, June 2026

eROSITA X-ray telescope on SRG (Спектр-РГ)

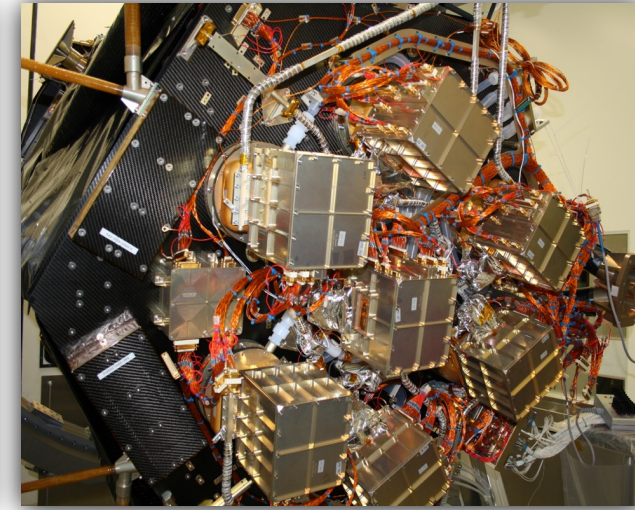
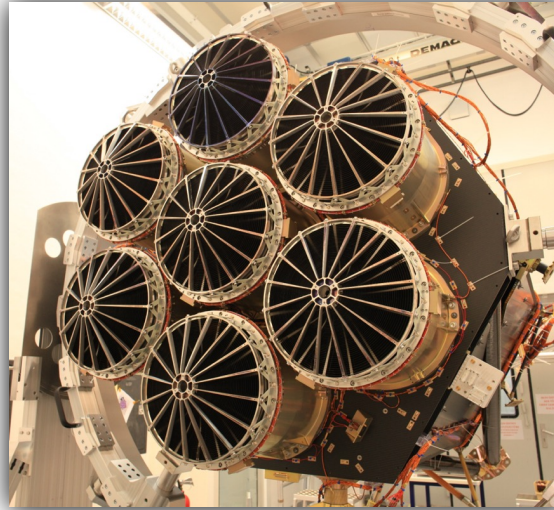


- extended ROentgen Survey with an Imaging Telescope Array
- Spectrum-Roentgen-Gamma (SRG) observatory
- Launch: July 13 2019, from Baikonur cosmodrome
- Orbit: Lagrange 2 (L2) point

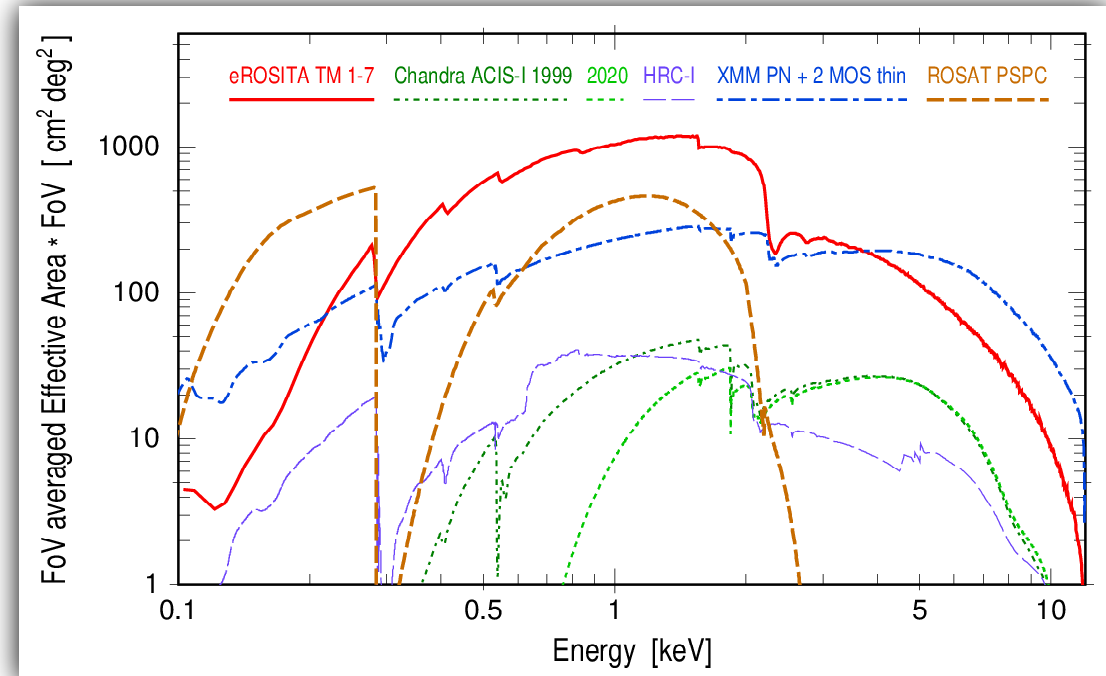


eROSITA in numbers

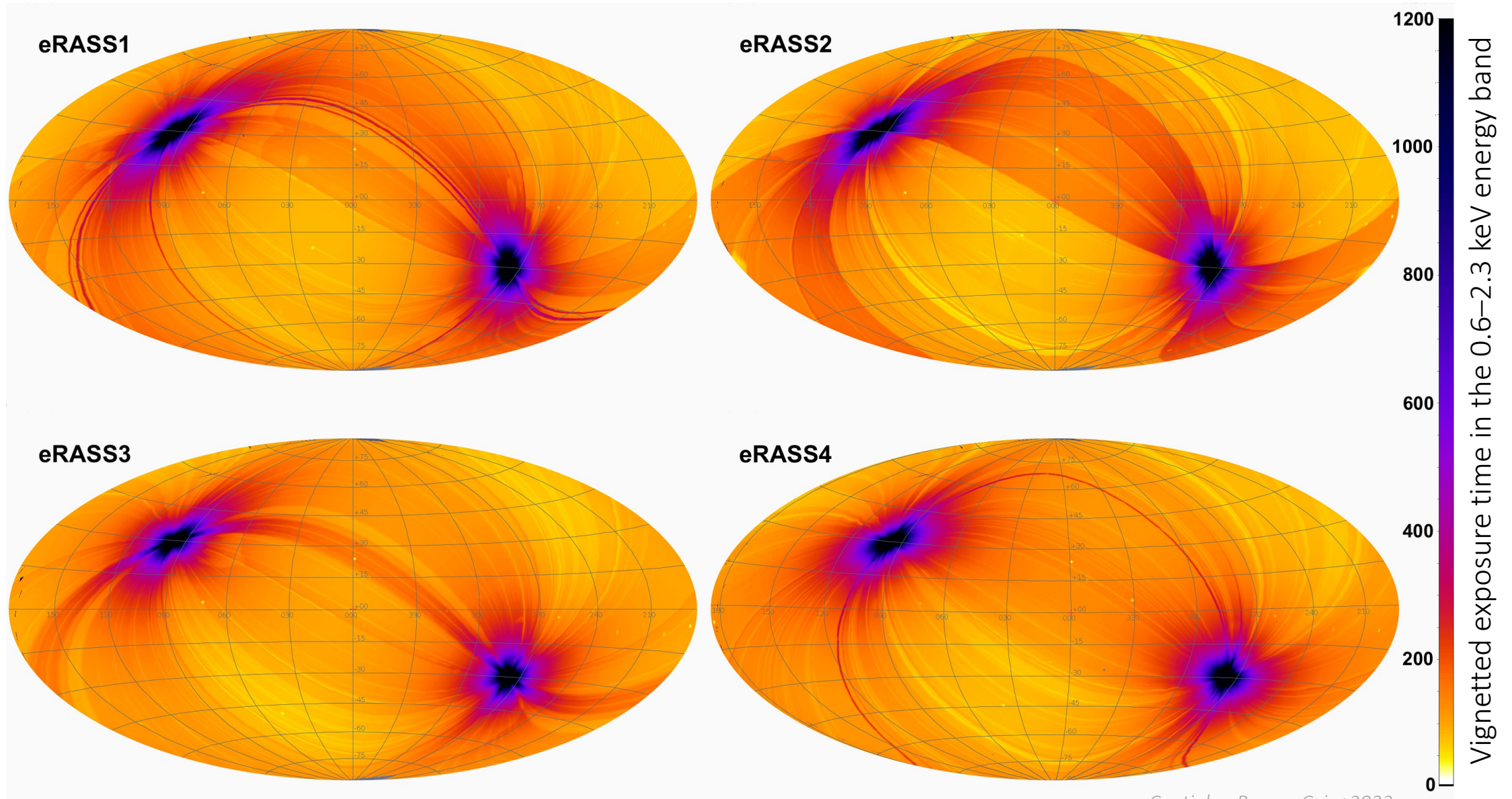
- 7 Wolter-I telescopes
- Each with 54 nested mirror shells
- Energy range: 0.1 – 10 keV
- Field of view: ~ 1 degree
- Effective area: 1237 cm^2 @ 1keV, 139 cm^2 @ 5 keV
- Half-Energy width (HEW): $\sim 18''$ (on-axis, pointed), $\sim 30''$ (FoV averaged, survey)
- Positional accuracy: $\sim 4.5''$ (1σ)
- Completed 4.4 all-sky surveys in ~ 2.5 yr



Predehl+2021

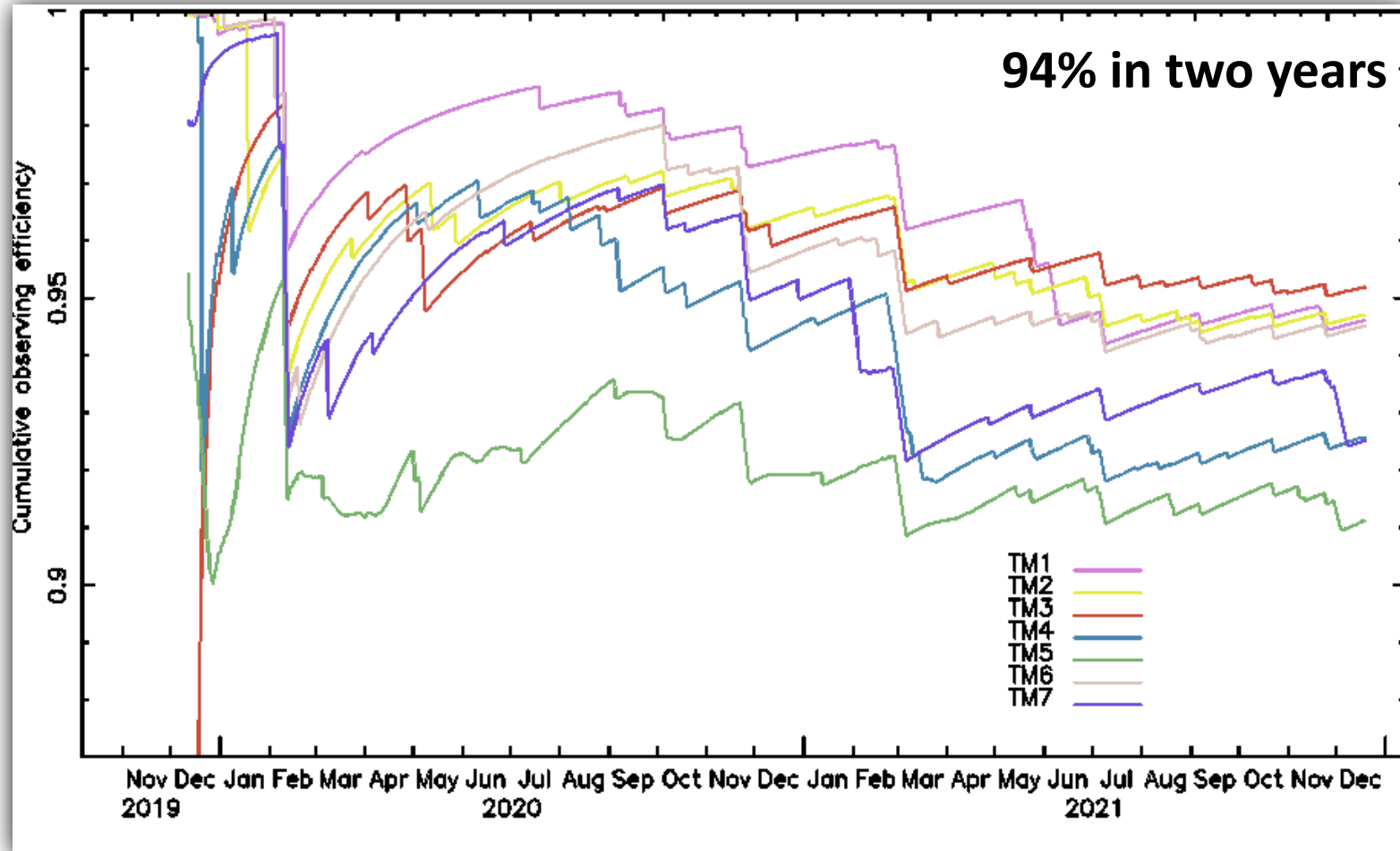


eROSITA observation efficiency



Coutinho, Ramos-Ceja+2022

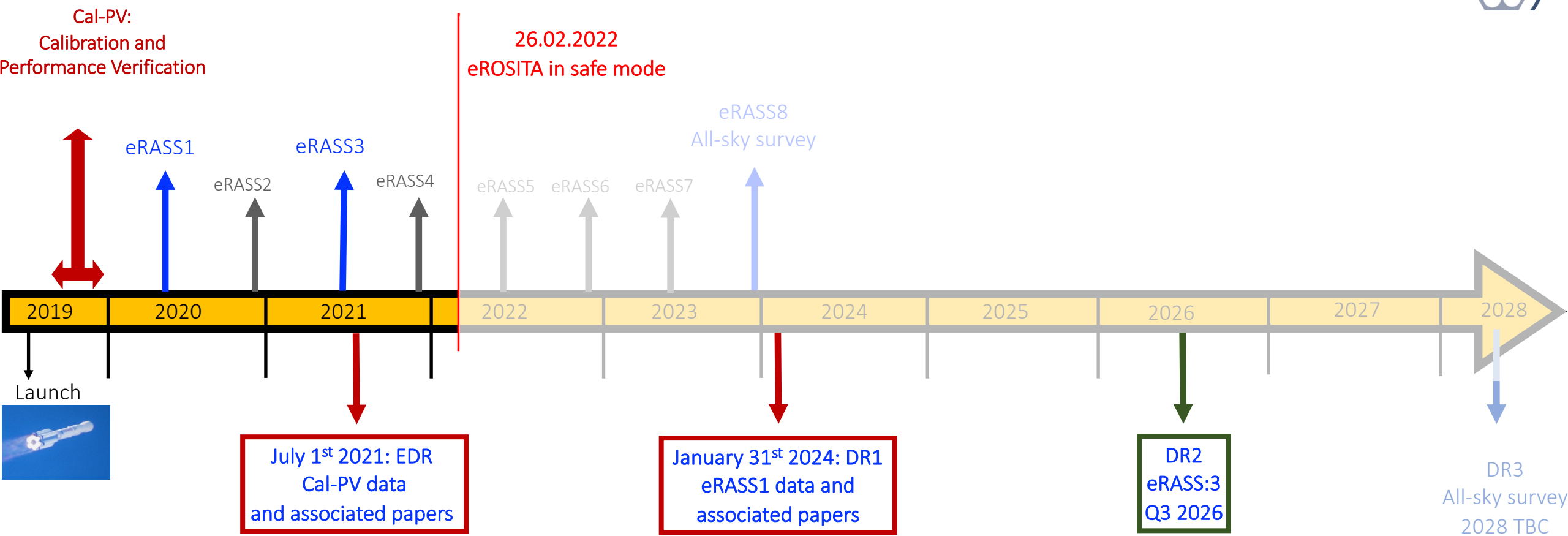
eROSITA observation efficiency



- Camera electronics and Interface and Thermal Controller anomalies
- Orbit corrections
- Light leak in telescope modules 5 and 7

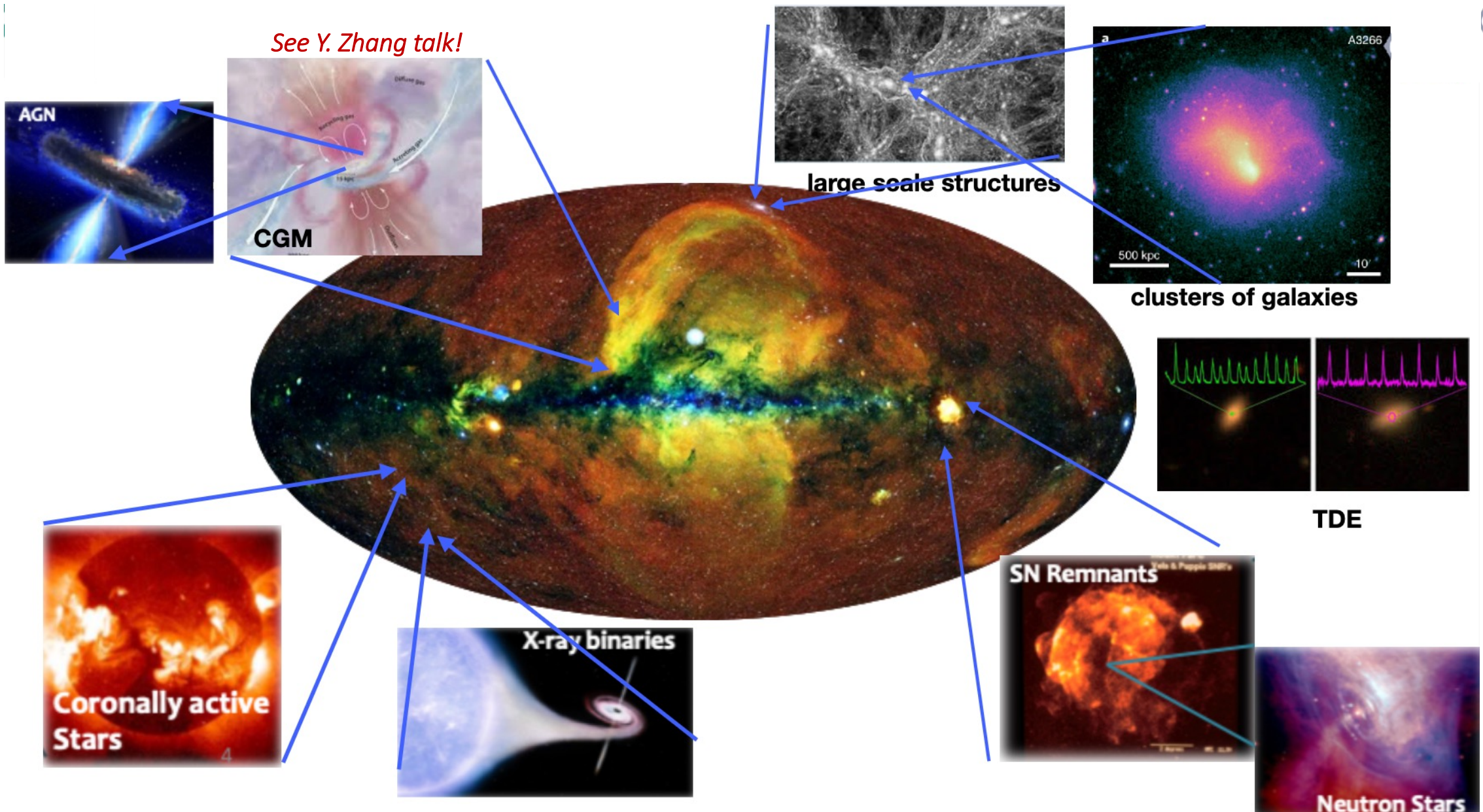
Coutinho, Ramos-Ceja+2022

eROSITA programmatics



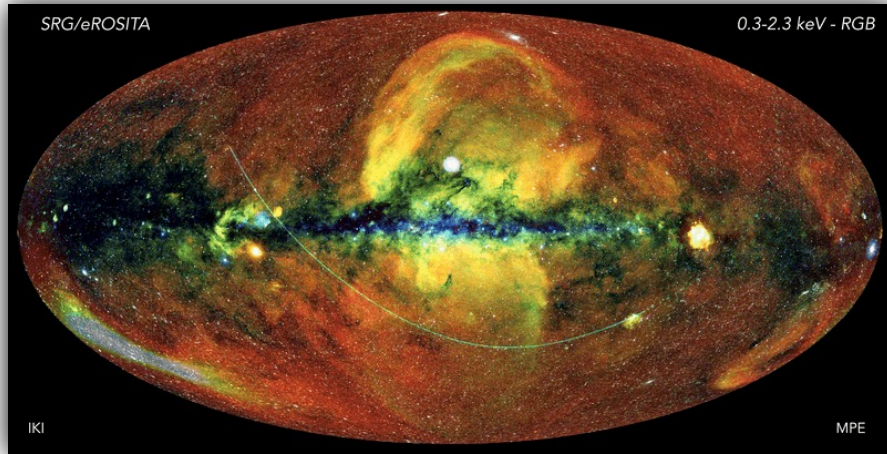
- Early Data Release (EDR) in 2021: several fields, including eFEDS mini-survey <https://erosita.mpe.mpg.de/edr>
- Data Release 1 (DR1) in 2024: eRASS1 data, including source catalogues <https://erosita.mpe.mpg.de/dr1>

eROSITA: Illuminating the Hot, Energetic Universe

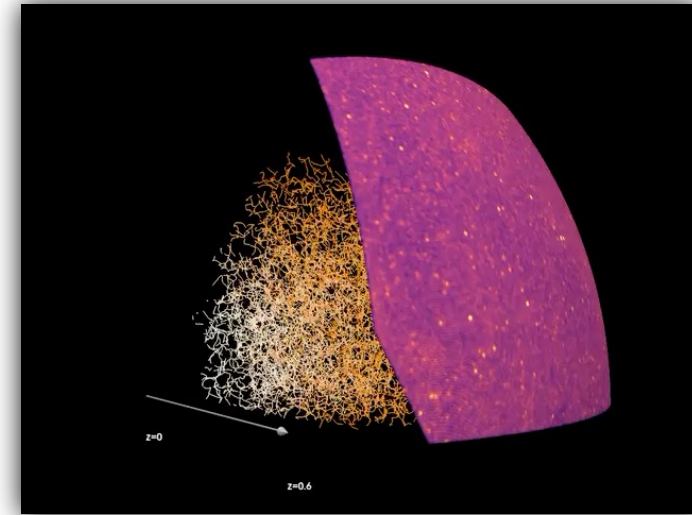


See Y. Zhang talk!

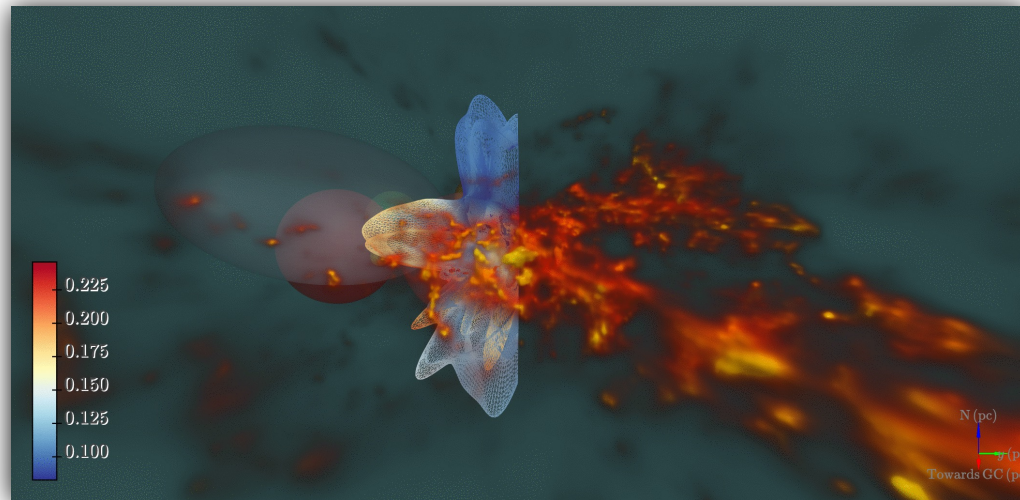
Selected science highlights since DR1



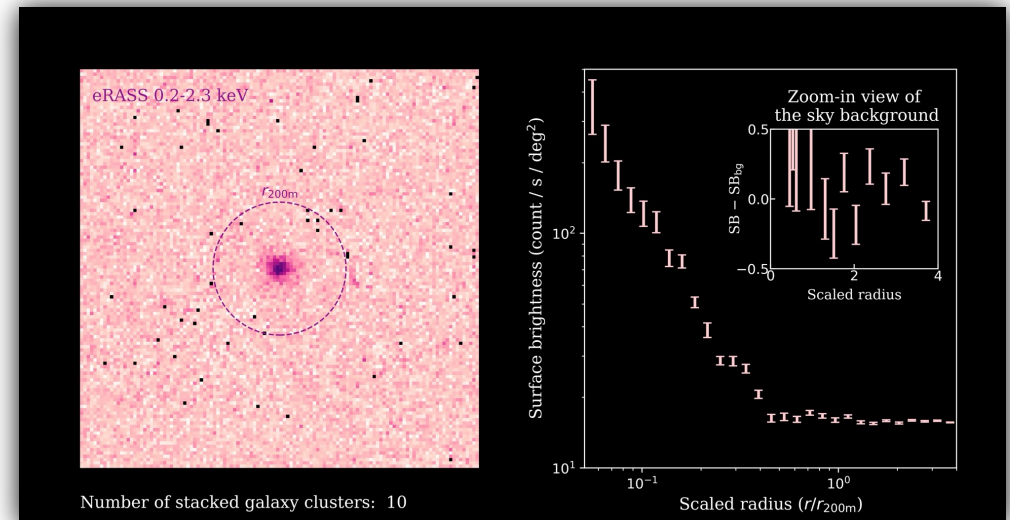
Dennerl+2026



Zhang+2024



Yeung+2024

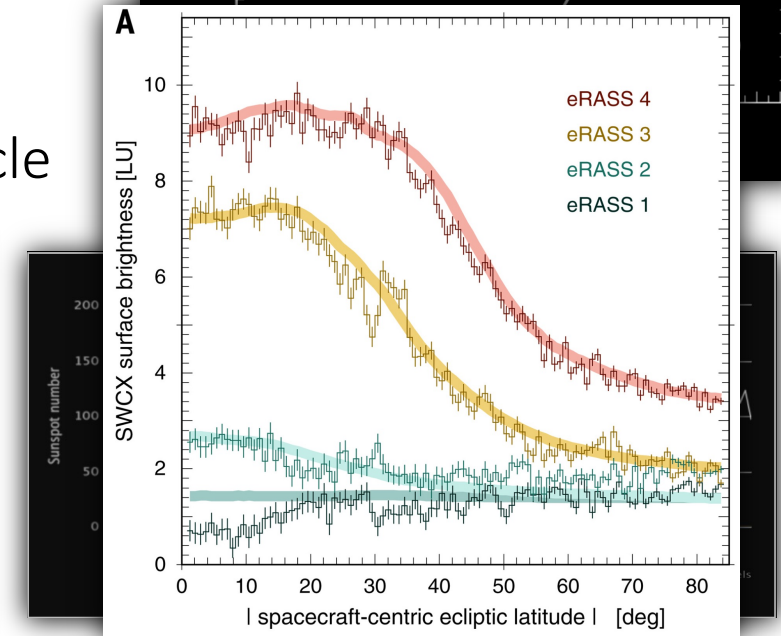
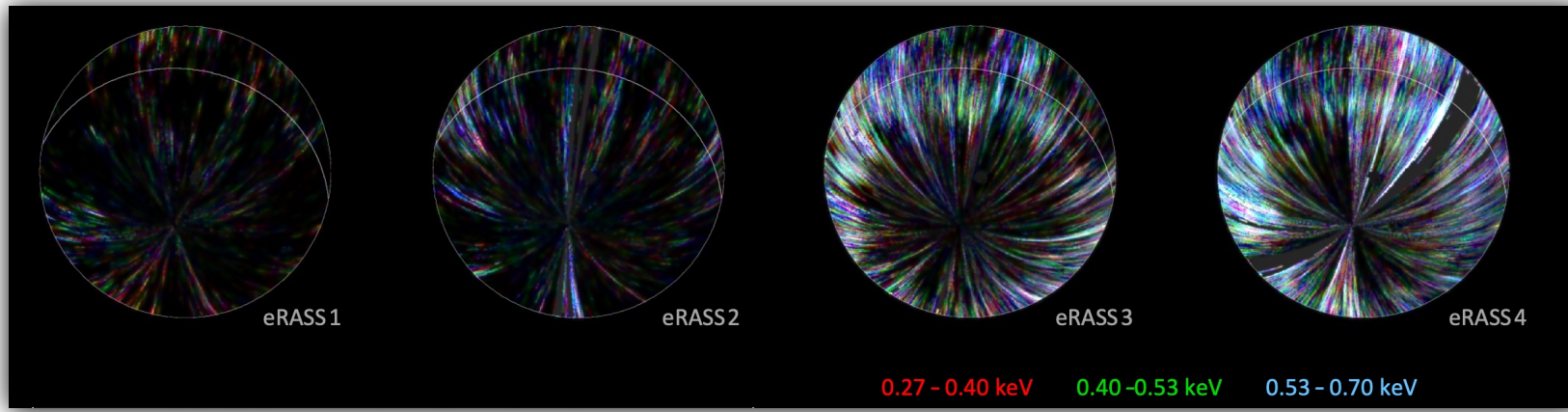
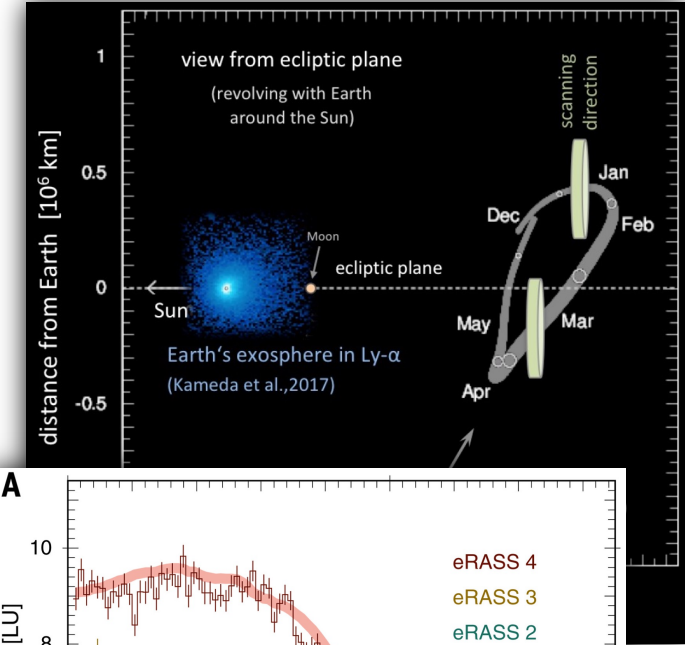


Zhang+2026

Solar System contribution to the soft X-ray sky



- X-rays are produced by ions picking up electrons from neutral gases, e.g., geocorona and heliosphere: solar-wind charge exchange (SWCX) emission
- SRG/eROSITA: first X-ray satellite which avoids geocoronal emission
- Isolated the heliospheric component and reconstructed the soft X-ray sky as it would appear when observed from outside the Solar System
- Clear evolution of heliospheric X-ray emission across the solar cycle

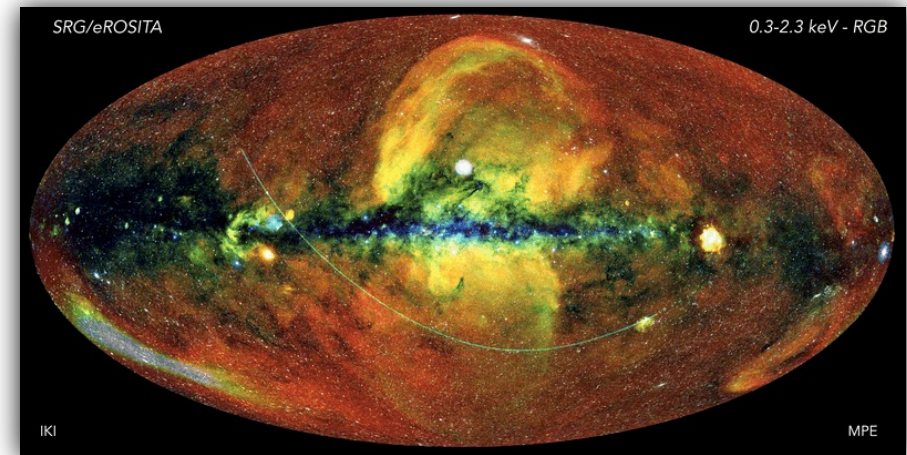
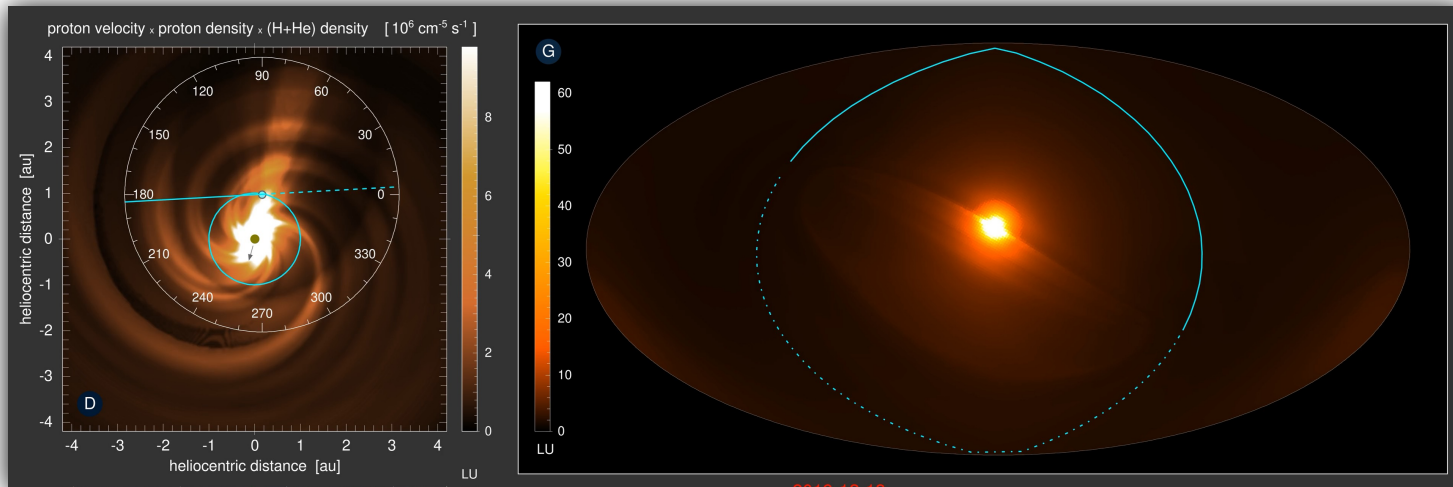


K. Dennerl+2026

Solar System contribution to the soft x-ray sky



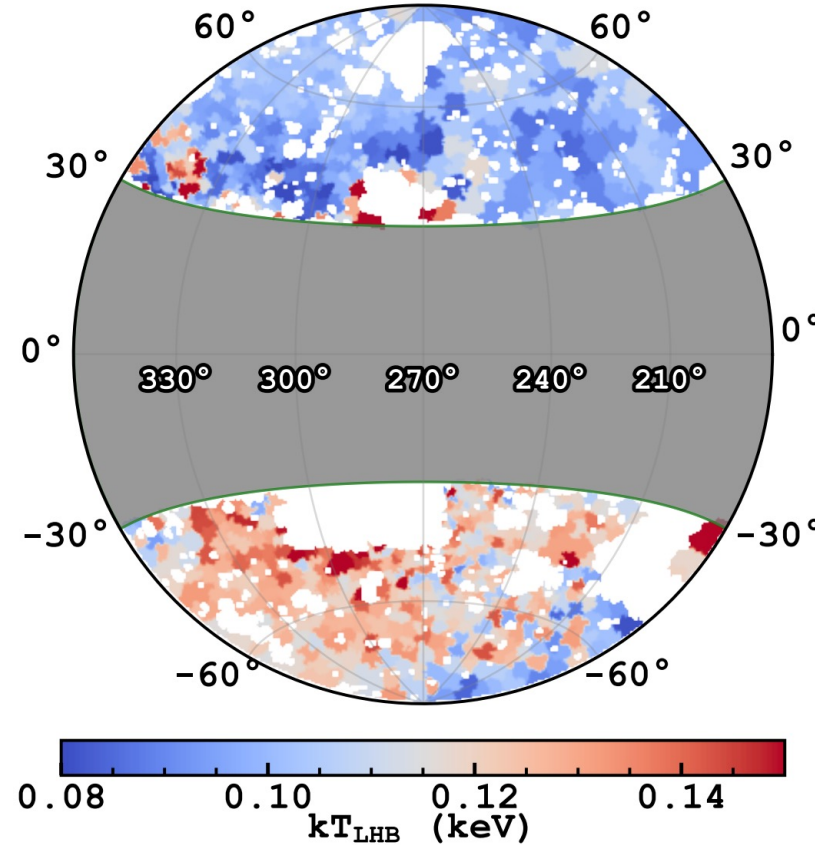
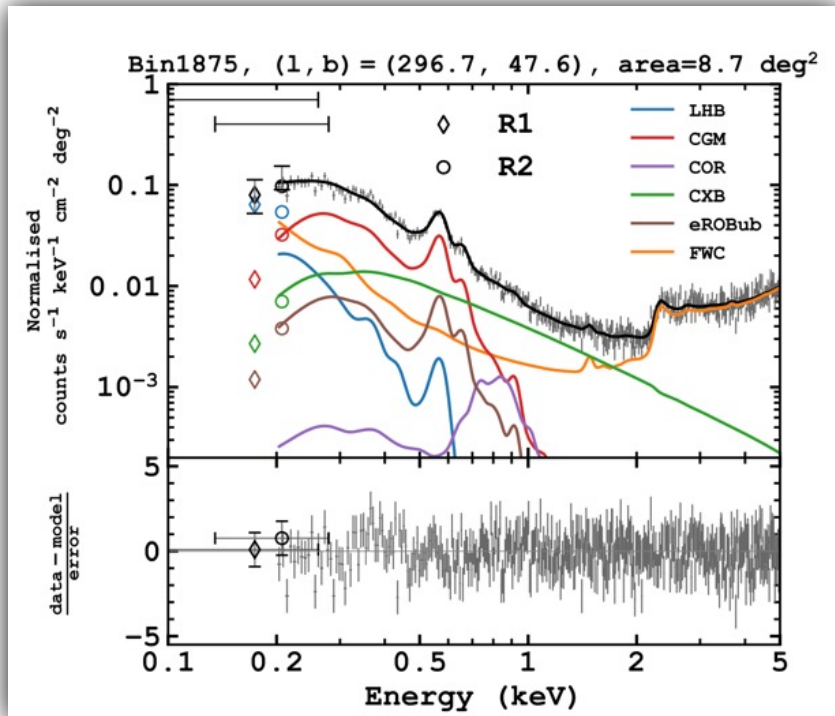
- Localized region of enhanced X-ray emission that does not revolve around the Sun
- Solar System moves through the Galaxy, gas containing helium atoms flows through it
 - Sun's gravity bends their trajectories, creating the helium focusing cone (predicted in the 70's)
- 3D model using solar wind measurements and interstellar matter data
- Emission predominantly originates from spiral structures produced by varying solar wind speeds



K. Dennerl+2026

The Local Hot Bubble

- eRASS1 (less SWXC contaminated) study of the diffuse soft X-ray background (SXRb): Local Hot Bubble
- North-South temperature difference and gradient towards low latitudes



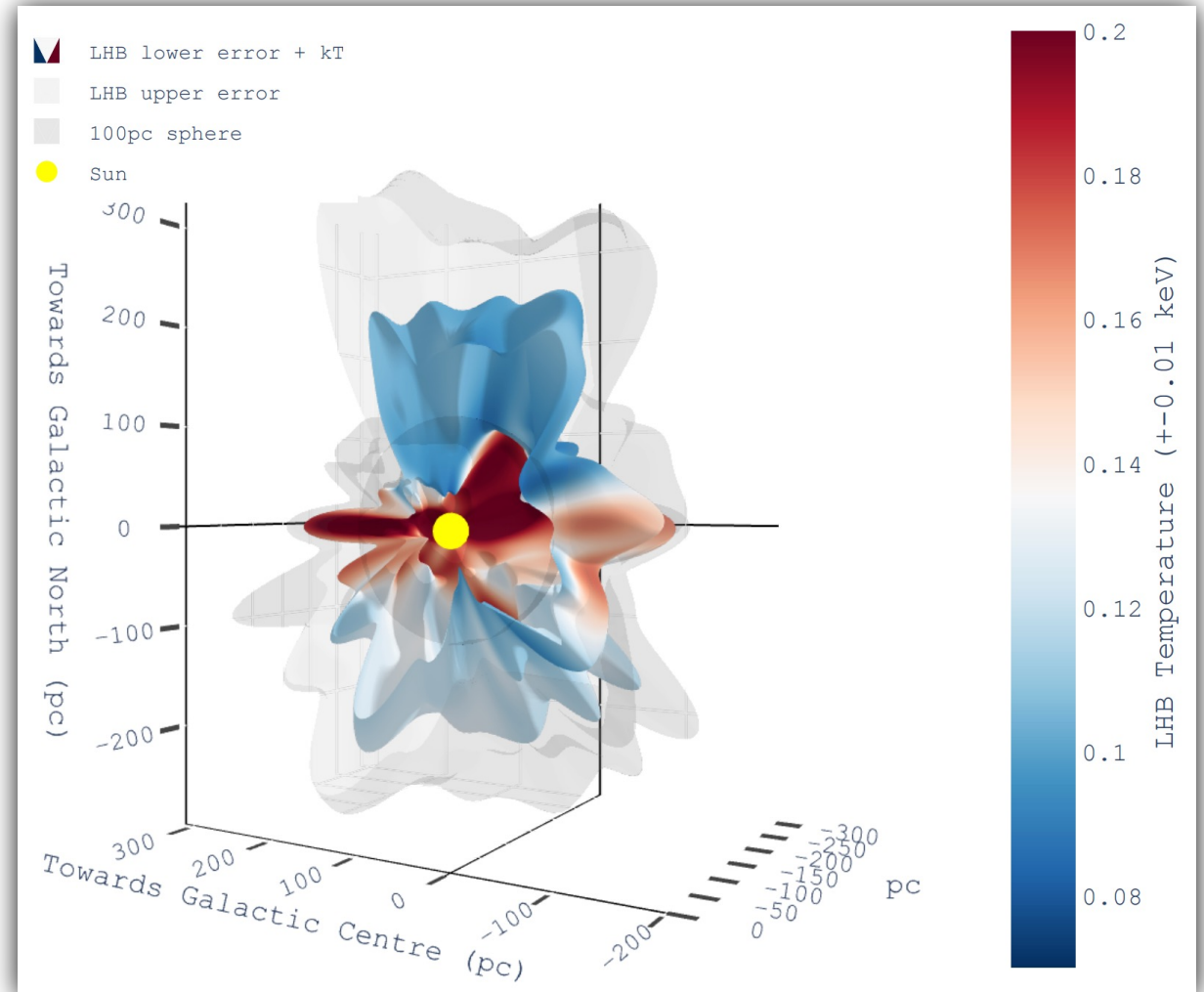
$$kT_N = 100.8 \pm 0.5 \text{ eV}$$

$$kT_S = 121.8 \pm 0.6 \text{ eV}$$

M. C. H. Yeung+2024

The Local Hot Bubble

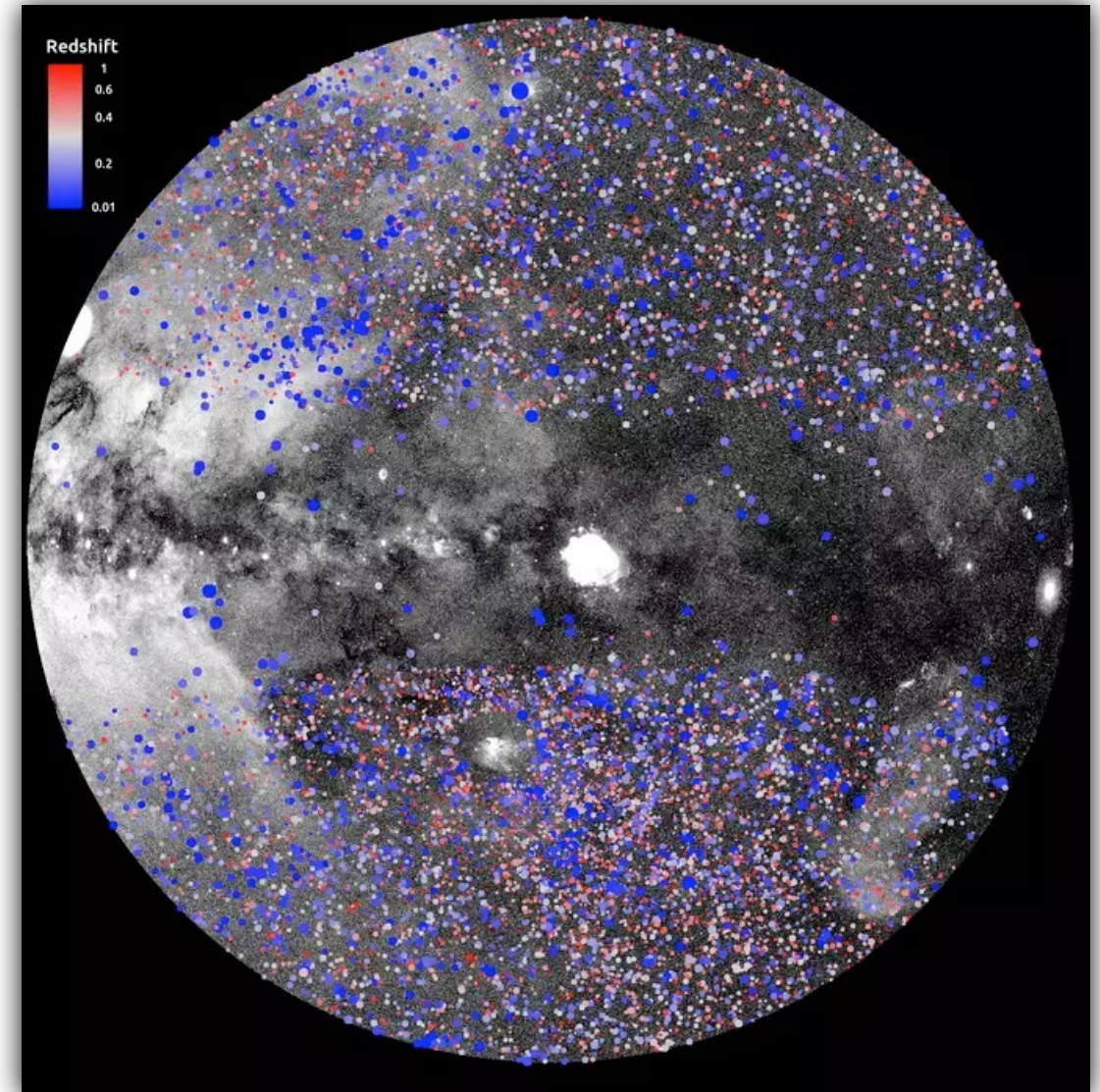
- eRASS1 (less SWXC contaminated) study of the diffuse soft X-ray background (SXRb): Local Hot Bubble
- North-South temperature difference and gradient towards low latitudes
- 3D map of the local hot bubble, constructed from emission measure assuming constant density



M. C. H. Yeung+2024

Galaxy clusters and groups in eRASS1

- Primary sample ($L_{\text{ext}} > 3$):
 - Total effective area: 13.1k deg²
 - 12,247 clusters with redshift measurements
 - 86% purity
- Weak lensing survey coverage
 - DES, HSC, and KiDS surveys
 - Coverage: 4968 deg²

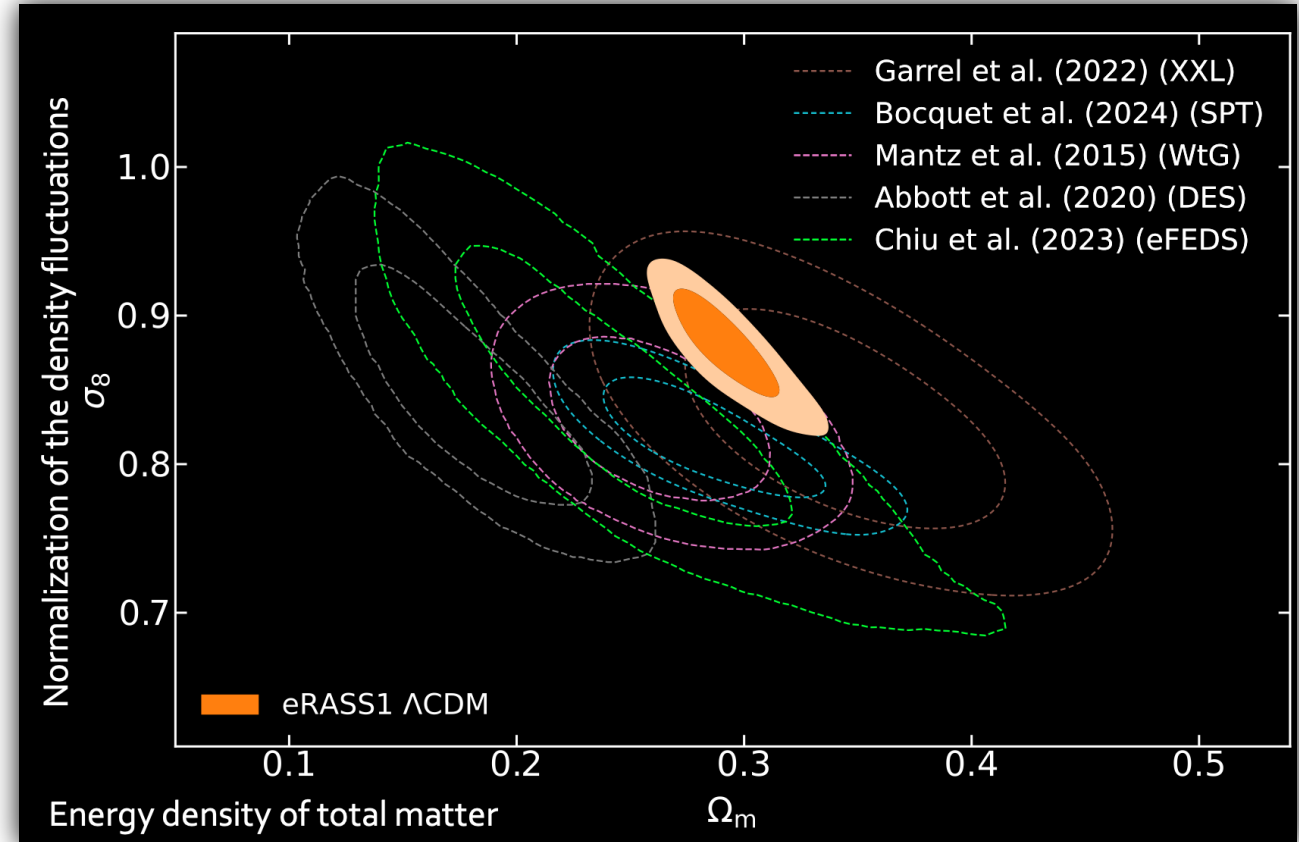


E. Bulbul+2024, A. Liu+2024, M. Kluge+2024

Galaxy clusters and groups in eRASS1

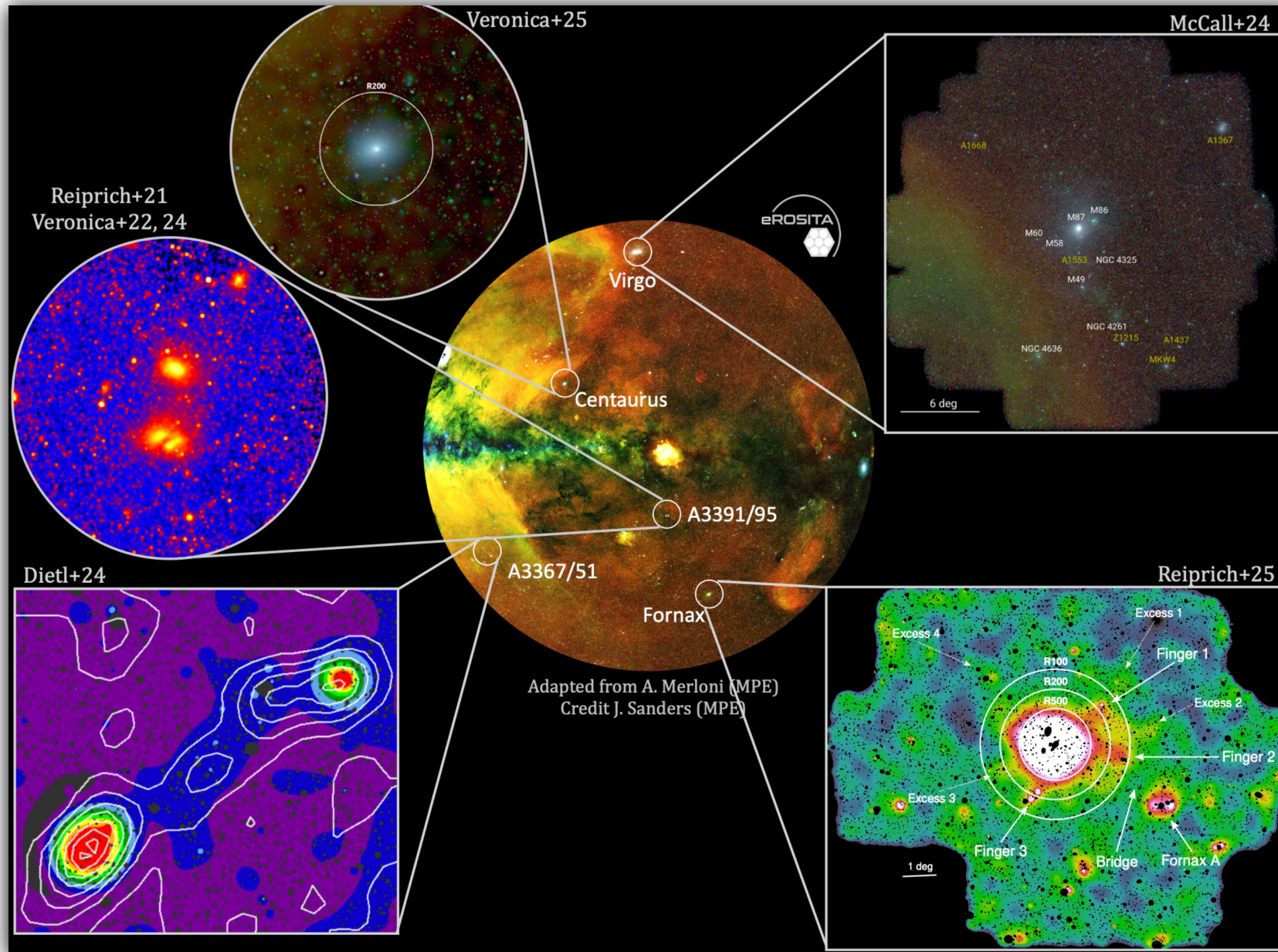


- Primary sample ($L_{\text{ext}} > 3$):
 - Total effective area: 13.1k deg²
 - 12,247 clusters with redshift measurements
 - 86% purity
- Weak lensing survey coverage
 - DES, HSC, and KiDS surveys
 - Coverage: 4968 deg²



E. Bulbul+2024, A. Liu+2024, M. Kluge+2024

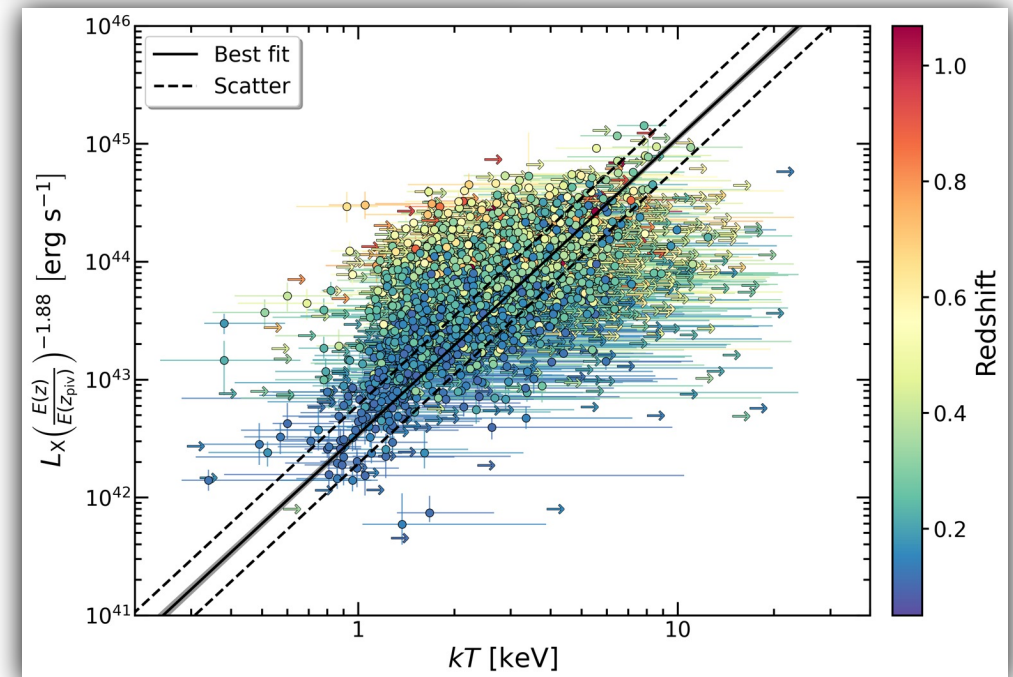
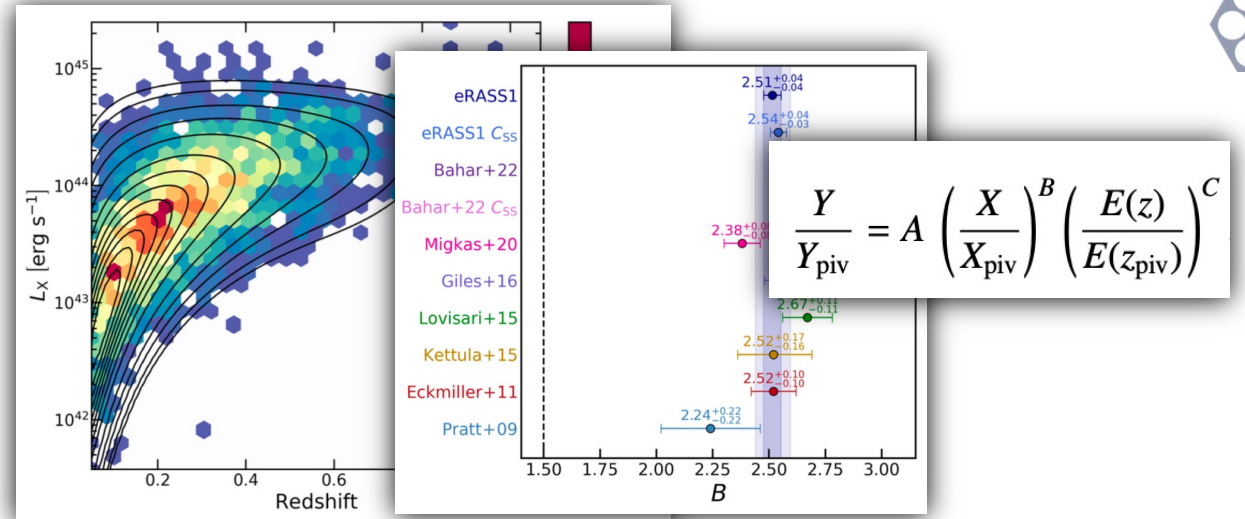
eROSITA results on individual galaxy clusters



X-ray scaling relations of galaxy groups and clusters



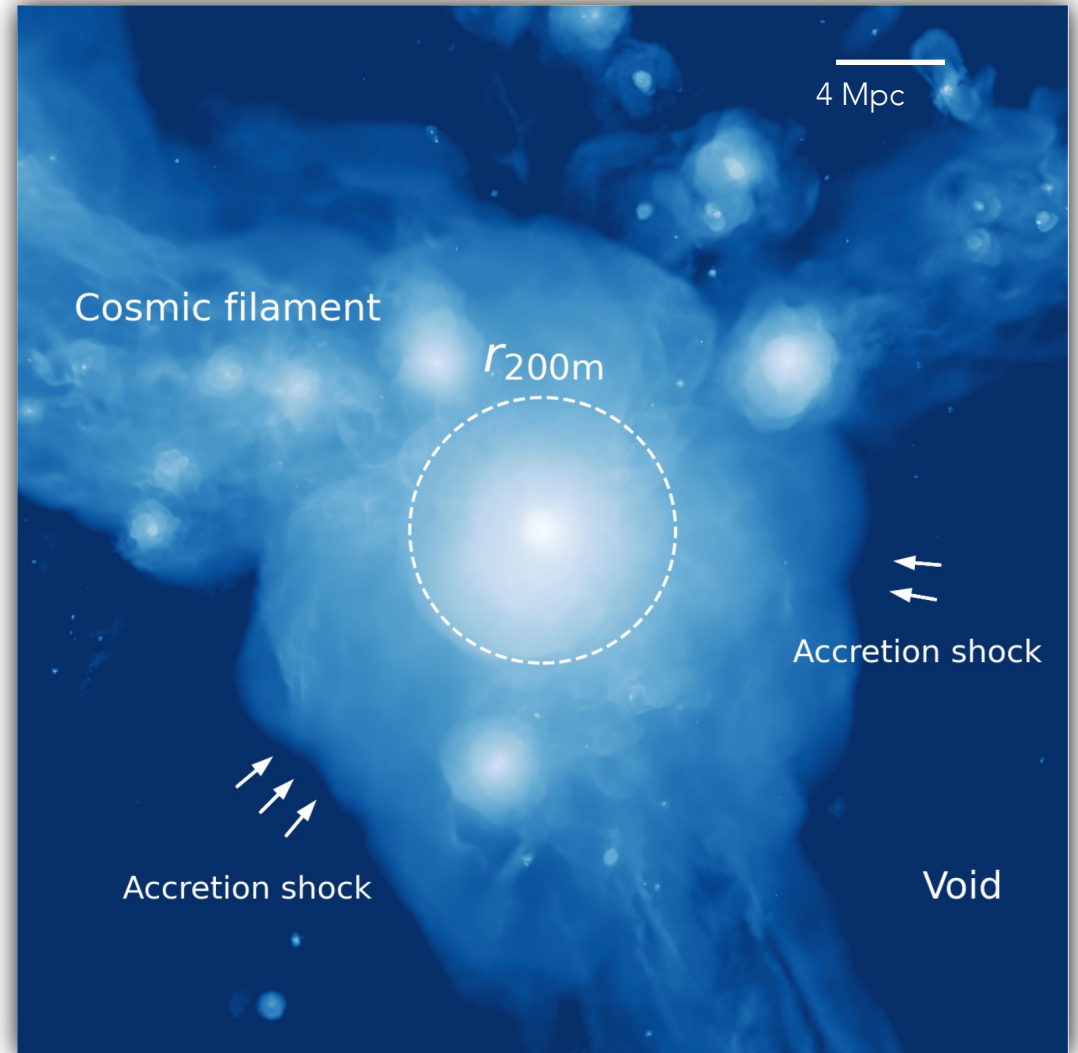
- 3061 galaxy groups and clusters from eRASS1
- Redshift range: 0.05 – 1.07
- Mass range: $1.1 \times 10^{13} - 1.6 \times 10^{15} M_{\odot}$
- Selection effects fully modeled using simulations
- $L_X - T$, $L_X - M_{\text{gas}}$, $L_X - Y_X$, and $M_{\text{gas}} - T$
- Scaling relation slopes deviate from self-similar predictions
- Redshift evolution matches self-similar expectations



M. E. Ramos-Ceja, L. Fiorino+2025 (A&A submitted)

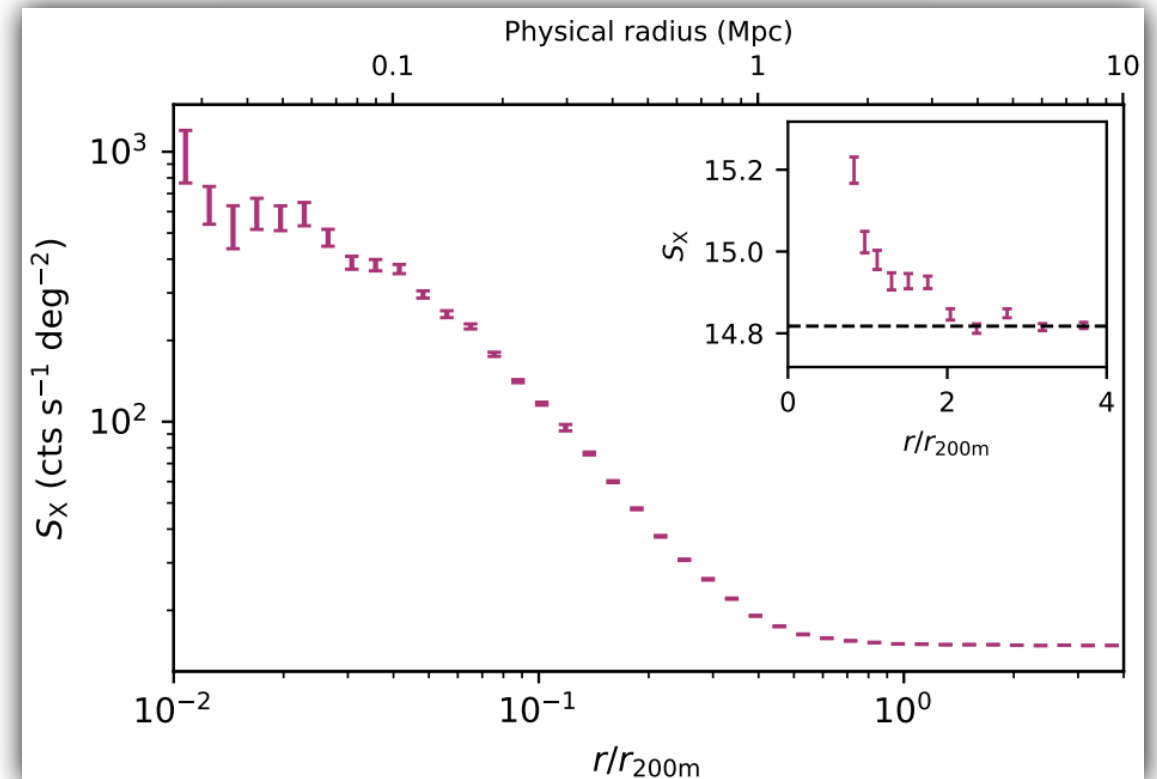
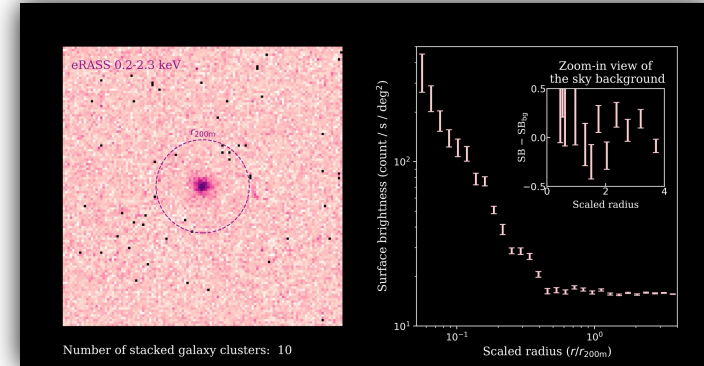
Detection of baryons beyond the virial radius

- Distribution of shock-heated gas from the outskirts of galaxy clusters into the large-scale accretion regions
- eRASS:4 data, high-fidelity imaging of nearby structures
- Stacking of 680 massive clusters ($M_{500} > 10^{14} M$) at $0.03 < z < 0.2$
- $\sim 3M$ total photon counts



Detection of baryons beyond the virial radius

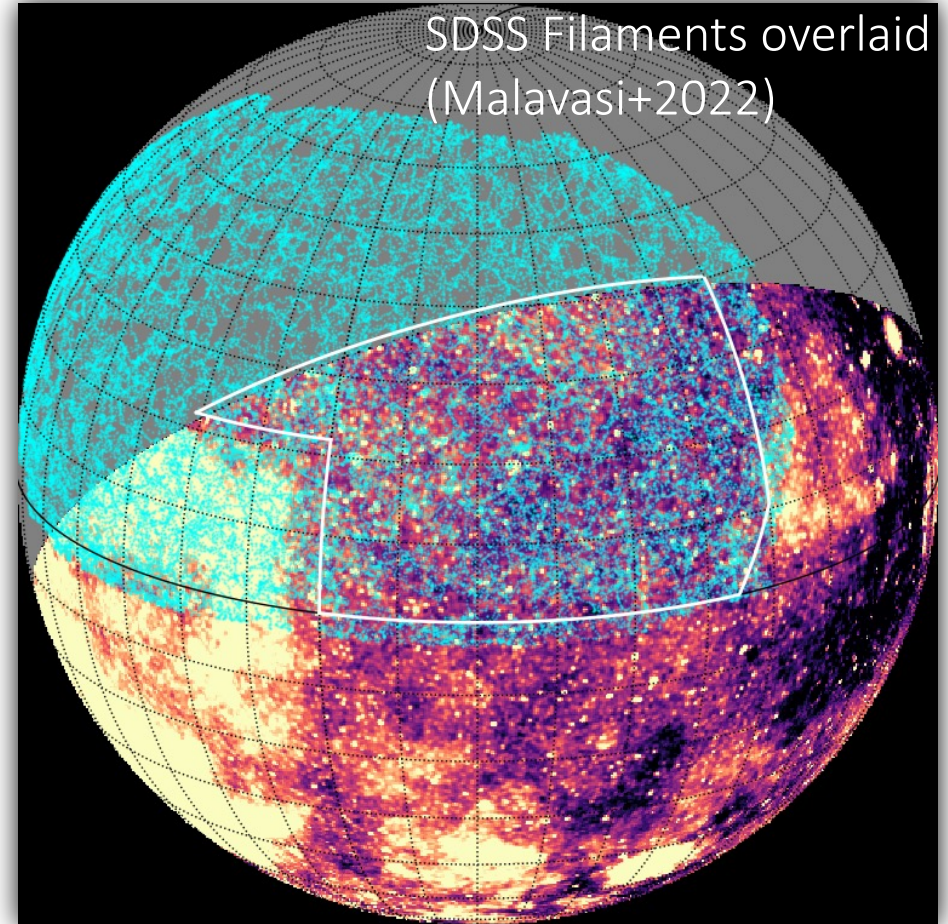
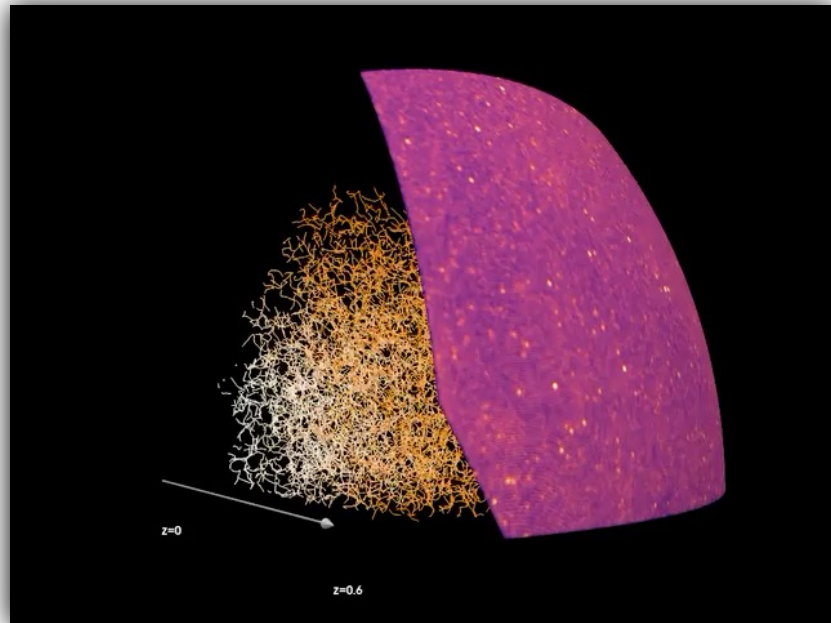
- Cluster outskirts & emission:
 - Extended to ~ 4.5 Mpc ($2R_{200m}$): 12σ
 - Baryon overdensity: $\Delta = 30$
 - Transition Radius (R_{200m}): 40% of emission originates from filaments or unresolved haloes
 - Help to disentangle feedback models
 - Gas mass fraction is reduced if residual clumping is added



X-ray emission from the warm-hot phase gas in long cosmic filaments



- The elusive warm-hot intergalactic medium
- eRASS:4 data
- 7817 filaments stacked in $0.2 < z < 0.6$ (from SDSS)
- Physical length $20 \text{ Mpc} < l < 100 \text{ Mpc}$

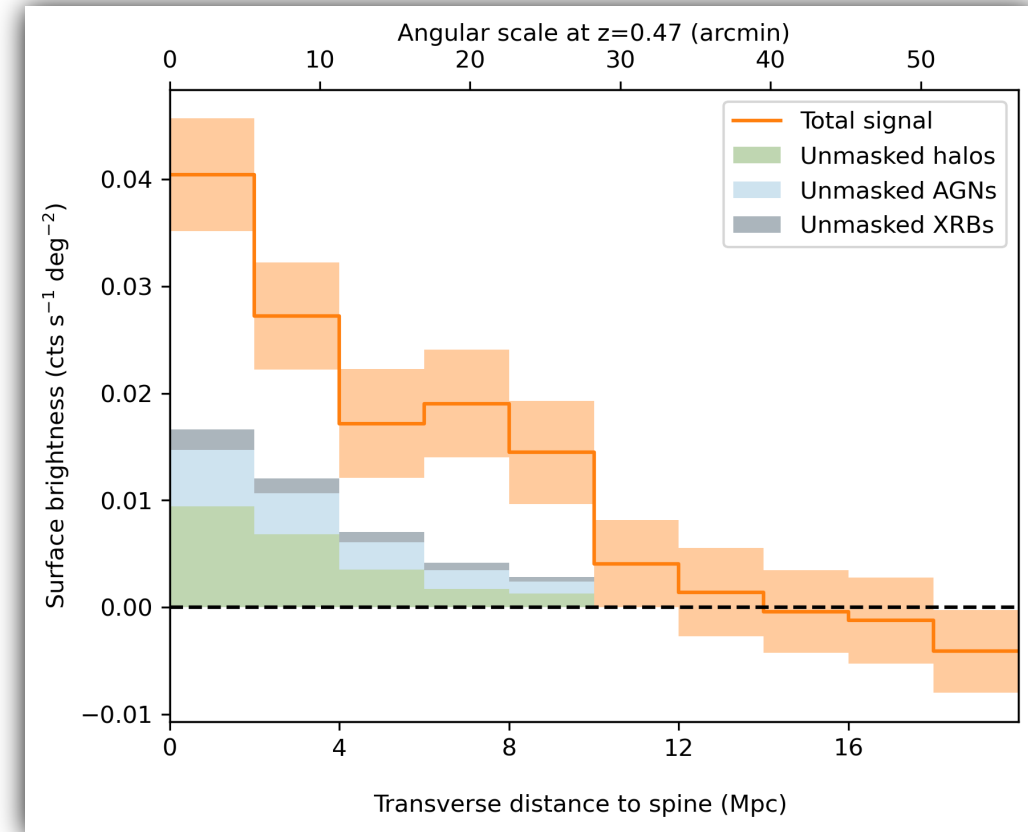
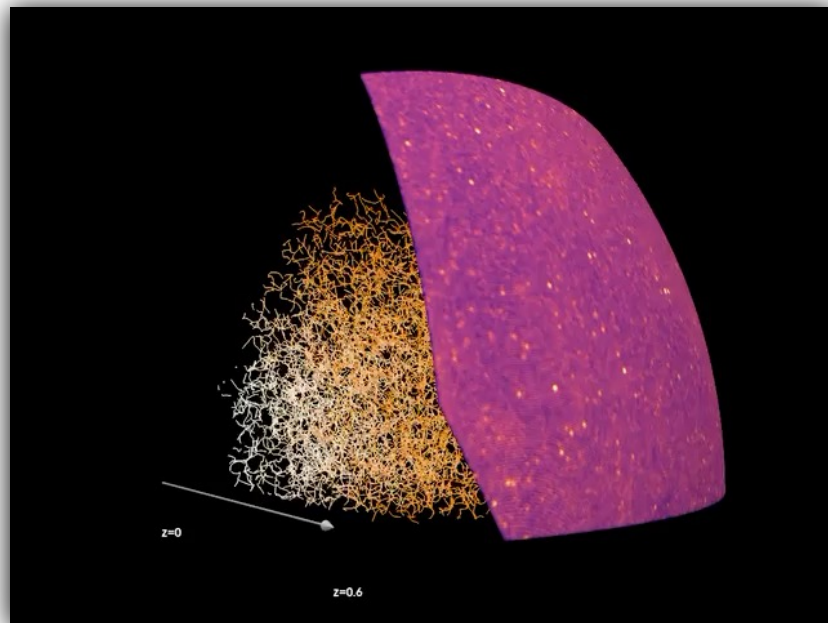


X. Zhang+2024

X-ray emission from the warm-hot phase gas in long cosmic filaments



- Stacked 9σ detection of WHIM (0.3–1.2 keV)
- $\sim 40\%$ could be contamination from unmasked sources
- $\sim 60\%$ from the hottest and densest part of the WHIM
- 5.4σ detection of WHIM



X. Zhang+2024

eROSITA-DE publication list



2026

<https://erosita.mpe.mpg.de/>

Yeung, M. C. H., et al., [The SRG/eROSITA diffuse soft X-ray background II. spectra and morphology of the eROSITA bubbles in the western Galactic hemisphere](#), A&A, submitted (2026).

Sbarrato, T., et al., [BLAZ4R and the eROSITA view of \$z>4\$ blazars](#), A&A, submitted (2026).

Buchner, J., et al., [A large population of over-massive black hole quasars at \$z=0.3-0.8\$ revealed by eROSITA](#), A&A, submitted (2026).

Ota, N., et al., [Bulk and turbulent gas motions in the interacting galaxy cluster Abell 3395 South observed with XRISM](#), PASJ, submitted (2026).

Gatuzz, E., et al., [Probing habitable regions with SRG/eROSITA](#), A&A, submitted (2026).

Qi, N., et al., [Investigating central star formation in local AGN host galaxies: is there tension between coeval growth and AGN feedback?](#), A&A, accepted (2026).

Markowitz, A., et al., [A Changing-Look Seyfert Discovered by eROSITA Reveals a Two-Component Broad-Line Region](#), A&A, accepted (2026).

Zheng, X., et al., [The average X-ray spectrum of the volume-complete M-, F-, G-, and K-type star sample within 10 pc of the Sun](#), A&A, 709, 275A (2026).

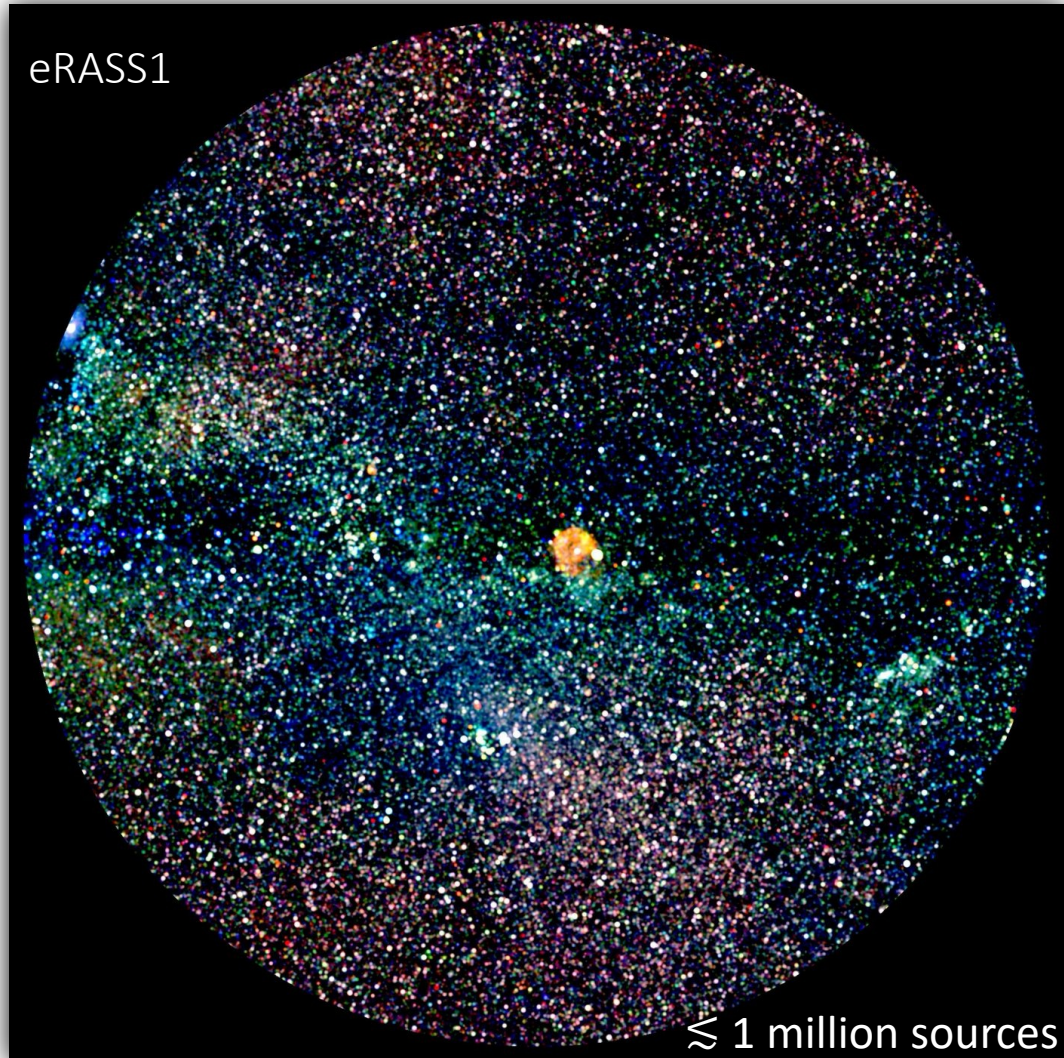
Zheng, Y.-L., et al., [Unveiling the Sources of X-Ray Luminosity in DESI Galaxy Groups: Insights from the SRG/eROSITA All-sky Survey](#), ApJ, 997, 111 (2026).

Zhang, X., et al., [The SRG/eROSITA All-Sky Survey. Detection of shock-heated gas beyond the halo boundary into the accretion region](#), A&A, 709, 72A (2026).

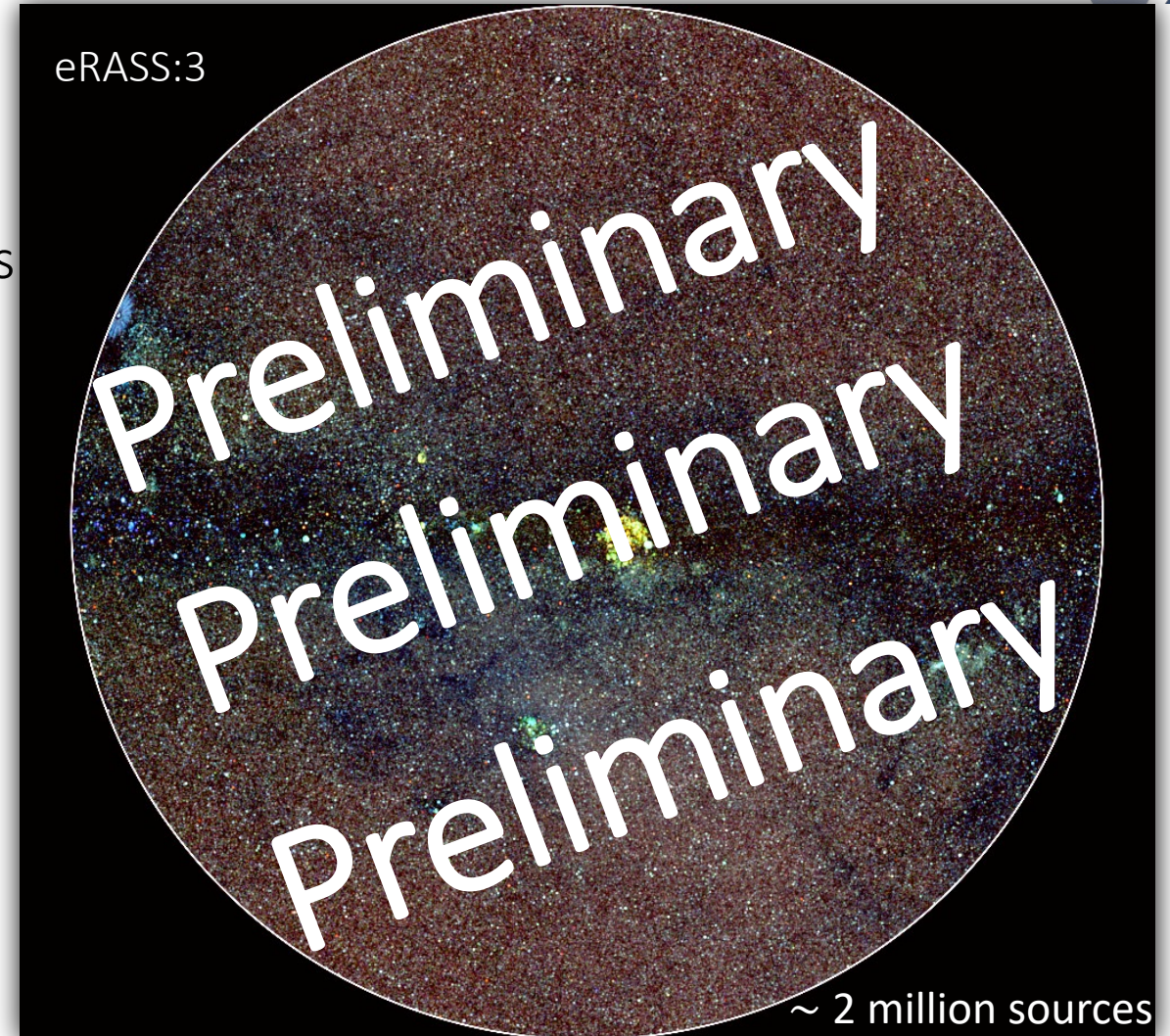
Haemmerich, S., et al., [BlazEr1: The eROSITA Blazar Catalog. Blazars and Blazar Candidates in the First eROSITA Survey](#), A&A, 709, 37A (2026).

Dennerl, K., et al., [Determination of the Solar System contribution to the soft x-ray sky](#), Science, 392, 285 (2026).

eROSITA-DE Data Release 2 (DR2)



Point
sources
only

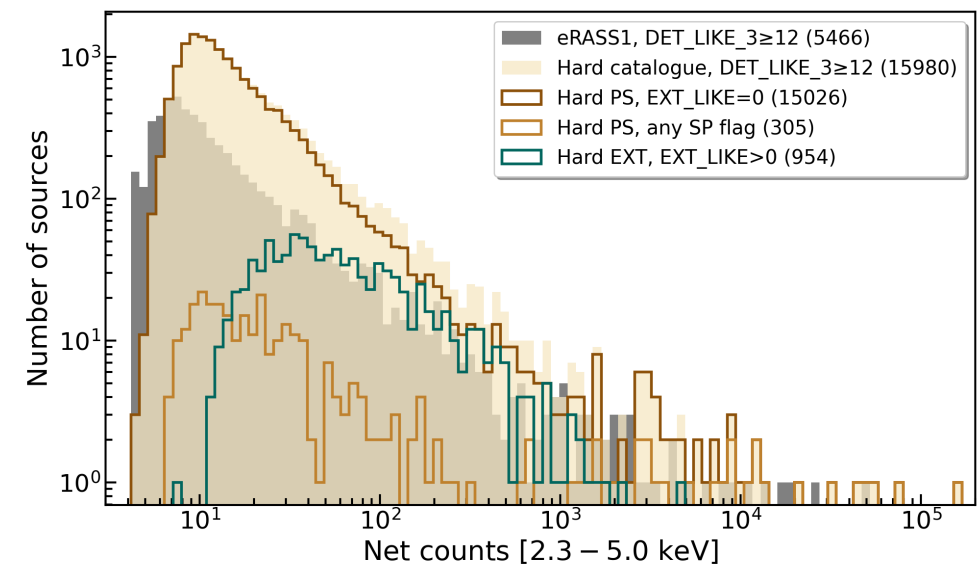
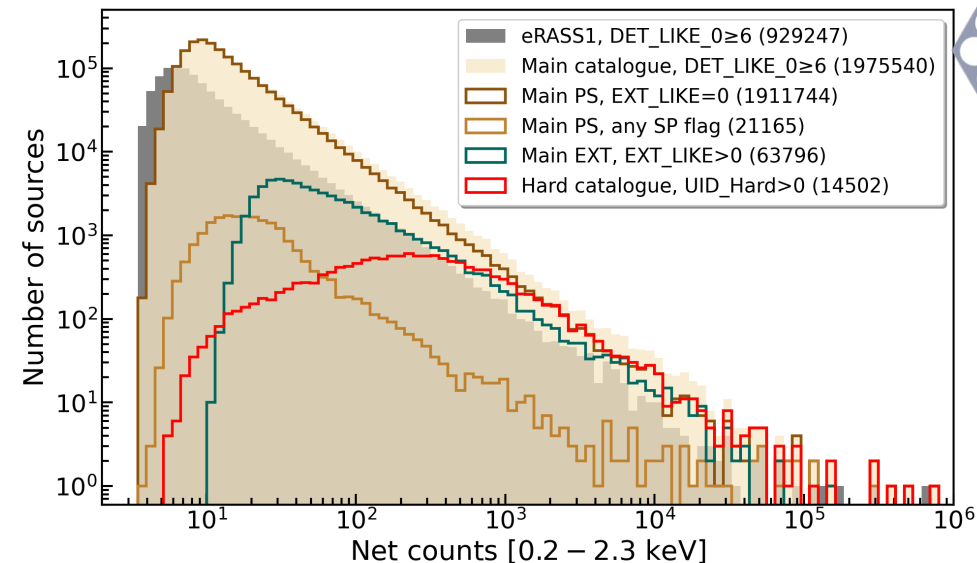


M. E. Ramos-Ceja+2026 (A&A submitted)

Data Release 2 (DR2)

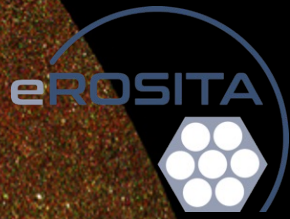


- Based on eRASS:3 data (1.5 years of observations).
- **Only** X-ray catalogues are released:
 - Main catalogue (0.2–2.3 keV): ~2 million sources (~1.9M point-like, ~64k extended).
 - Hard catalogue (2.3–5.0 keV): ~15,000 sources.
- Extensive comparisons with eRASS1, 2RXS, 5XMM and CSC 2.1.
- Additionally:
 - Six catalogues that identify and classify the optical and infrared counterparts of the point-sources (LS10, CW20, GDR3).
 - Majority of identified sources are extragalactic.



M. E. Ramos-Ceja+2026 (A&A submitted)

Summary



- X-ray astronomy plays a fundamental role in exploring the Universe by revealing both common and extreme astrophysical phenomena.
- eROSITA aboard SRG is the most powerful wide-field X-ray telescope developed to date. Since beginning operations in 2019, it has completed 4.4 all-sky surveys over more than two years of observations.
- Its large grasp, stable instrumental background, and regular observing cadence enable eROSITA to probe previously unexplored regions of X-ray parameter space.



Thank you for your attention!