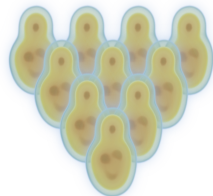


The unexpected uses of a bowling pin

Exploiting ^{20}Ne isotopes for precision characterizations
of collectivity in small systems

Govert Nijs

July 1, 2026



Based on:

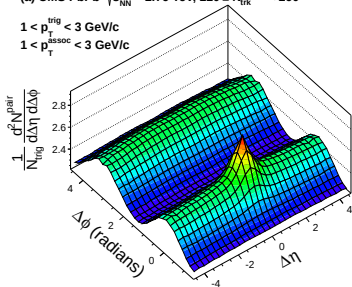
- Giacalone, GN, van der Schee, 2606.03993
- Giacalone, Bally, GN, Shen, Duguet, Ebran, Elhatisari, Frosini, Lähde, Lee, Lu, Ma, Meißner, Noronha-Hostler, Plumberg, Rodríguez, Roth, van der Schee, Somà, 2402.05995



One fluid to rule them all?

(a) CMS PbPb $\sqrt{s_{NN}} = 2.76$ TeV, $220 \leq N_{\text{Trk}}^{\text{offline}} < 260$

$1 < p_{\text{T}}^{\text{trig}} < 3$ GeV/c
 $1 < p_{\text{T}}^{\text{assoc}} < 3$ GeV/c



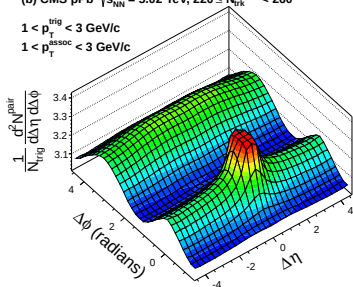
- Collective flow is one of the main characteristics of AA collisions.
 - We interpret this as the hydrodynamic expansion of a QGP formed in the collision.
 - Our understanding is quantitative, and models describe the data with ever-improving precision.
- Collective flow has also been seen in smaller systems:
 - High multiplicity pPb,
 - High multiplicity pp,
 - Inside high multiplicity jets,
 - ...
- Does a QGP form in these systems?
- Are small systems hydrodynamic in nature?



One fluid to rule them all?

(b) CMS pPb $\sqrt{s_{NN}} = 5.02 \text{ TeV}$, $220 \leq N_{\text{trk}}^{\text{offline}} < 260$

$1 < p_T^{\text{trig}} < 3 \text{ GeV/c}$
 $1 < p_T^{\text{assoc}} < 3 \text{ GeV/c}$

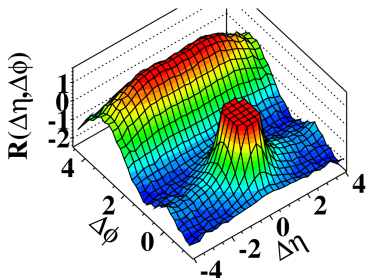


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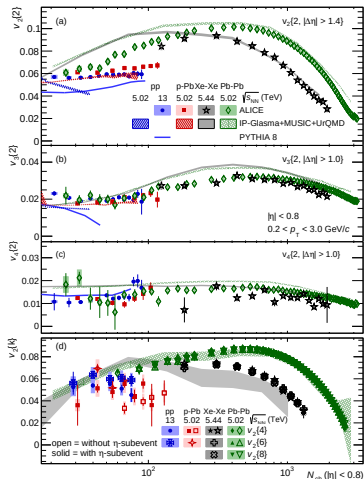
(d) CMS $N \geq 110$, $1.0\text{GeV}/c < p_T < 3.0\text{GeV}/c$



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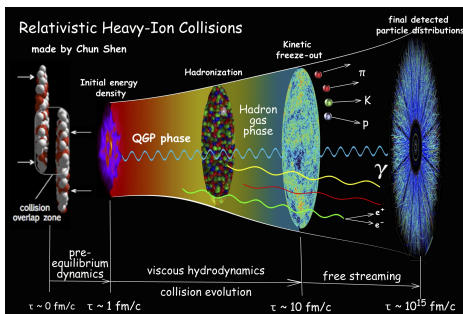
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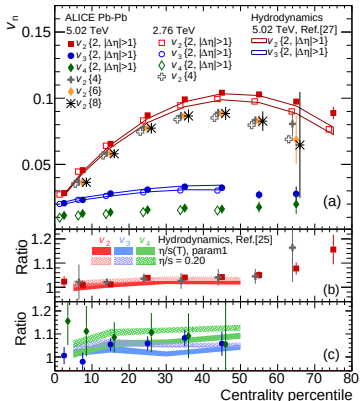
A broad picture of heavy ion collisions



- A heavy ion collision proceeds in several stages:
 - An initial collision.
 - The approach to local equilibrium.
 - Hydrodynamic expansion.
 - Freeze-out as the fluid cools and falls out of equilibrium.
- The first two stages are the least understood.
- If small systems follow this paradigm, they spend relatively more time in these early stages!



Recap: why do we believe PbPb is hydrodynamic?

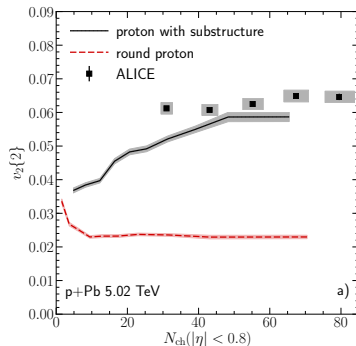


- Not just the presence of $v_n\{k\}$.
- We understand where the $v_n\{k\}$ come from!
 - Hydrodynamics converts initial state anisotropic geometry into final state momentum anisotropy.
 - We can control the initial geometry by binning in centrality.
- For $p\text{Pb}$ this is not the case.
 - We do not understand the initial geometry.
 - No clear connection between initial state and final state.
- Why is the geometry of $p\text{Pb}$ so hard to control?



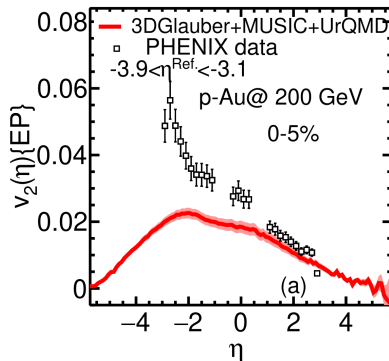
Reason 1: substructure

- In a $p\text{Pb}$ collision, one of the projectiles is a single proton.
- Proton geometry at low x is largely unknown.
- In models, substructure is usually modelled as a proton consisting of several constituents.
- With substructure, the collision geometry becomes significantly less round.
- This has a large effect on the v_2 .
 - We can describe the data...
 - ... but only when making the 'right' choice for proton substructure.
- We expect better theoretical control if none of the projectiles are a proton.



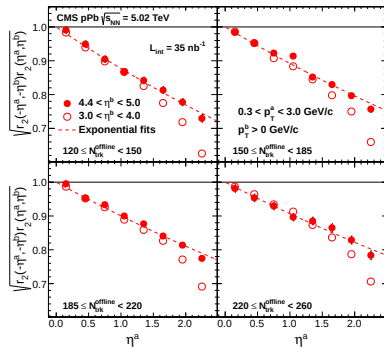
Reason 2: longitudinal structure

- In $p\text{Pb}$, $v_2(\eta)$ is not flat at midrapidity.
- Also, the flow decorrelates across rapidities more than in PbPb.
- This violates the assumption of boost invariance present in most models.
 - Dropping this assumption is an active area of study.
 - Can probably be overcome in the future with better tuned models.
 - However, current descriptions are not precise enough to reliably control this aspect.
- Symmetric systems avoid this problem.

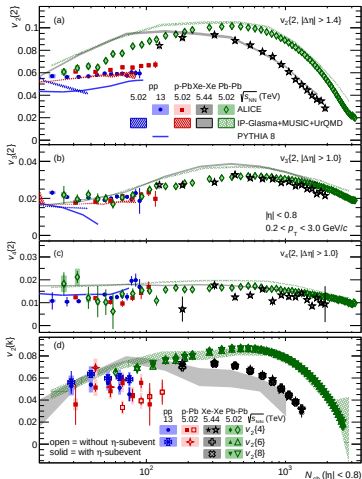


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Posing a precise question

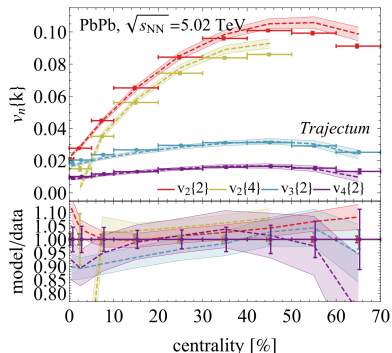


- Can we describe a small system using a hydrodynamical model *which also describes PbPb*?
 - Hydro model used should describe a wide range of PbPb observables.
 - We want to use the remaining freedom in the model after demanding agreement with PbPb data as an uncertainty.
- Can we find a quantity to predict which does not suffer from huge theoretical uncertainties? Wishlist:
 - Small sensitivity to proton substructure.
 - No longitudinal structure issues.
 - Quantifiable and small theory uncertainty.



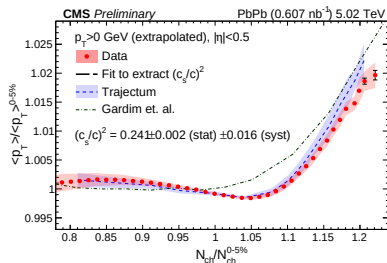
Interlude: Bayesian analysis & systematic uncertainties

- We use the *Trajectum* model, and apply Bayesian analysis to fit to PbPb observables.
- We fit to 670 individual data points at 2.76 and 5.02 TeV.
 - $v_2\{2\}$ is shown as an example.
- The resulting posterior distribution allows us to predict observables, *including systematic theory uncertainties*.
 - The ultracentral $\langle p_T \rangle$ is shown as an example.



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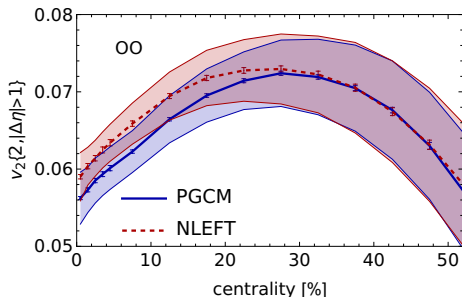


Predictions for $^{16}\text{O}^{16}\text{O}$ at the LHC

- Using the parameters obtained by Bayesian analysis, we predict the v_2 of $^{16}\text{O}^{16}\text{O}$ collisions at LHC energies.

- The bands represent the systematic uncertainty.
- Uncertainty varies between 4% and 14%.
- Different nuclear structure models give slightly different but consistent answers.

- Virtually no dependence on proton substructure.
- Longitudinal decorrelation is much less of an issue compared to $p\text{Pb}$.

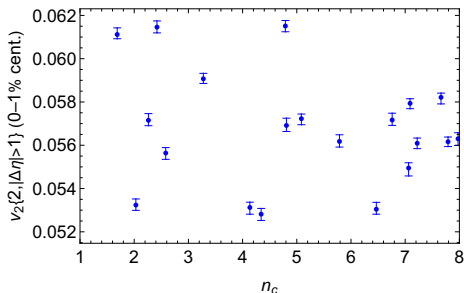


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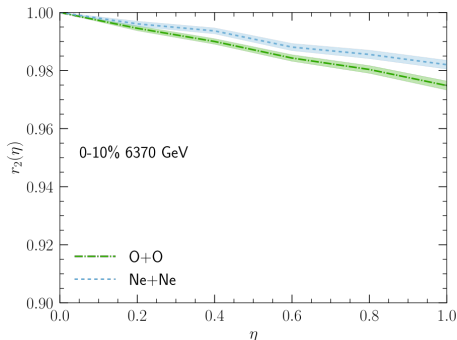
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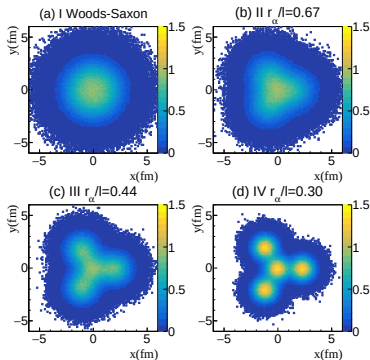


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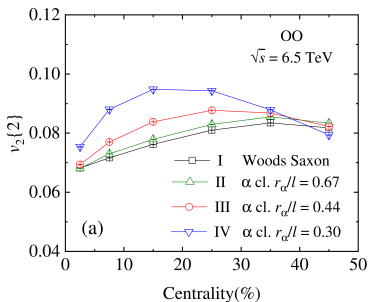
Dependence on ~~nucleon substructure~~ nuclear structure



- We removed the dependence on nucleon substructure.
 - ^{16}O nucleus is much larger than a single proton.
 - Makes sense that proton substructure becomes subleading.
- But we also introduced dependence on nuclear structure.
 - Presence of α -clusters matters.
 - Size of α -clusters matters.
 - Placement of α -clusters matters.
- Are we just exchanging one problem for another?
- No. Nuclear structure can be reliably computed for small nuclei.



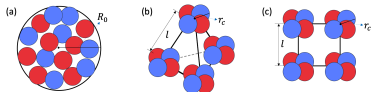
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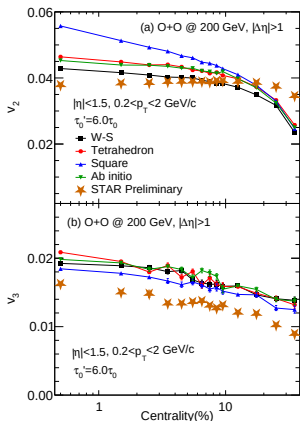
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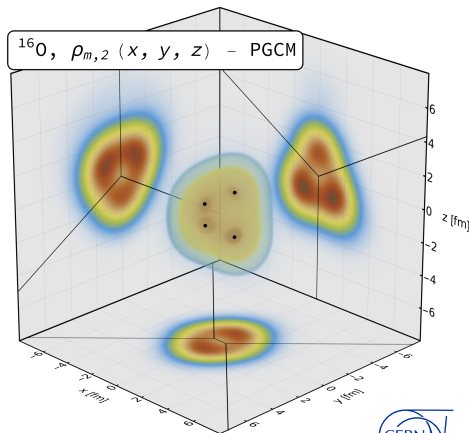


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Ab initio nuclear structure models

- PGCM and NLEFT are both state-of-the-art nuclear structure computations.
 - PGCM provides a nucleon density.
 - NLEFT provides configurations.
- Both frameworks yield results that are largely consistent with each other.
- ^{16}O consists of 4 α -clusters.
- The α -clusters are arranged in an irregular tetrahedron.



Revisiting the wishlist

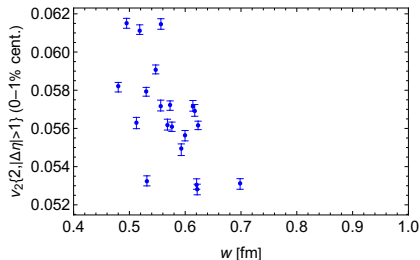
$v_n\{k\}$ in	$p\text{Pb}$	OO
Small sensitivity to proton substructure	✗	✓
No longitudinal structure issues	✗	✓
Quantifiable theory uncertainty	✗	✓
Small theory uncertainty	✗	$\geq 4\%$

- Theory has a much better handle on $^{16}\text{O}^{16}\text{O}$ compared to $p\text{Pb}$.
- Uncertainties are still substantial.
 - This limits our ability to test theory models.
- Can we do better?



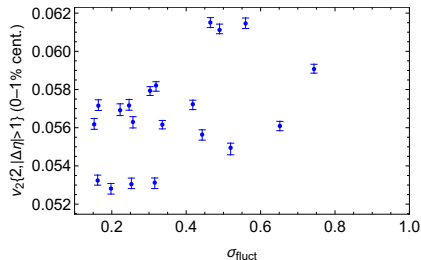
Where are these uncertainties coming from?

- There are still substantial geometric uncertainties in the $^{16}\text{O}^{16}\text{O}$ prediction for v_2 :
 - Increasing the nucleon size w decreases v_2 .
 - Increasing initial state fluctuations σ_{fluct} increases v_2 .
- Systems close in size to $^{16}\text{O}^{16}\text{O}$ should share these dependencies.
 - Can we exploit this?



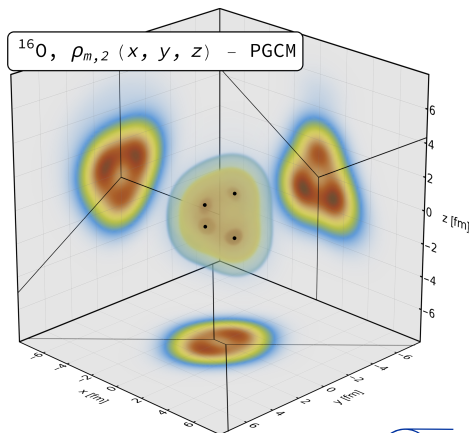
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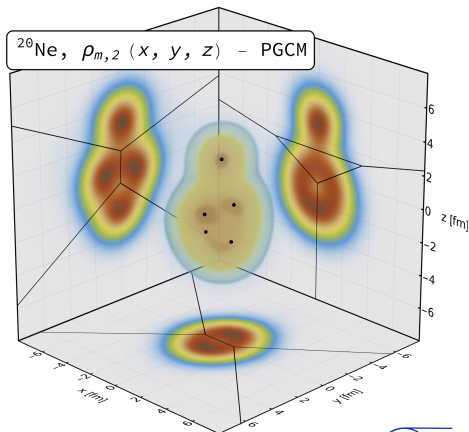
The nuclear bowling pin: ^{20}Ne

- We now turn around the argument.
 - ~~Using ratios of observables to extract the difference in shape.~~
 - Using a known shape difference to predict ratios of observables.
- ^{20}Ne is close in size to ^{16}O .
 - Remaining geometric uncertainties should largely cancel.
- ^{20}Ne looks like ^{16}O , but with an extra α -cluster on top.
 - Whereas ^{16}O is close to spherical, ^{20}Ne is the most deformed nucleus in the Segre chart.
 - Central v_2 ratio NeNe/OO should have a large signal.



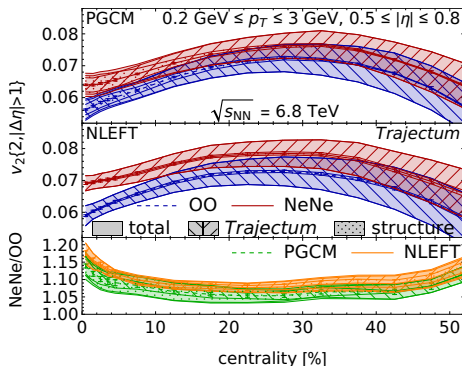
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Comparing ^{20}Ne to ^{16}O significantly reduces errors!

- Central NeNe/OO v_2 ratio has a large enhancement.
 - Ratio of $v_2\{2\}$ reaches percent level precision from 5% to 20% centrality!
 - Geometric uncertainties indeed largely cancel.
- Difference of $\rho(v_2\{2\}^2, \langle p_T \rangle)$ has uncertainty reduced by up to a factor 5!

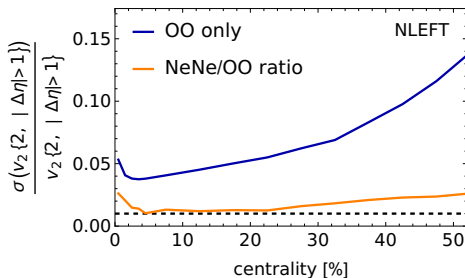


0–1%	$v_2\{2\}_{\text{NeNe}}/v_2\{2\}_{\text{OO}}$	$\rho_{2,\text{NeNe}} - \rho_{2,\text{OO}}$
NLEFT	1.174(8) _{stat.} (31) _{syst.} ^{Traj.} (4) _{syst.} ^{str.}	−0.124(14) _{stat.} (10) _{syst.} ^{Traj.} (7) _{syst.} ^{str.}
PGCM	1.139(6) _{stat.} (27) _{syst.} ^{Traj.} (28) _{syst.} ^{str.}	−0.124(10) _{stat.} (10) _{syst.} ^{Traj.} (29) _{syst.} ^{str.}



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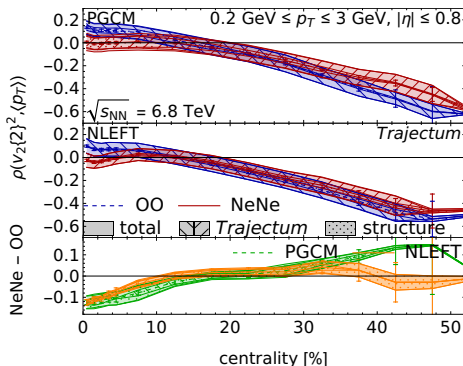


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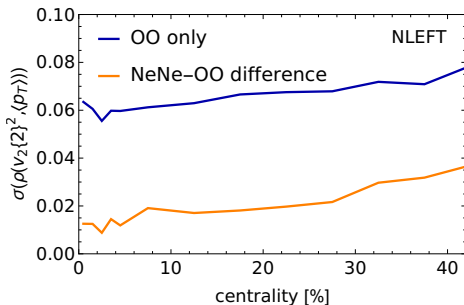


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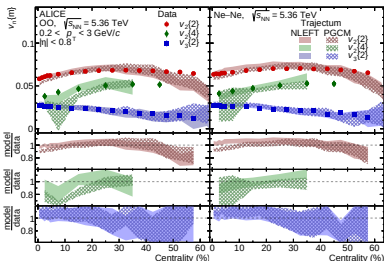
Revisiting the wishlist again

$v_n\{k\}$ in	$p\text{Pb}$	OO	NeNe/OO
Small sensitivity to proton substructure	✗	✓	✓
No longitudinal structure issues	✗	✓	✓
Quantifiable theory uncertainty	✗	✓	✓
Small theory uncertainty	✗	$\geq 4\%$	$\geq 1\%$

- Theory has a much better handle on $^{16}\text{O}^{16}\text{O}$ compared to $p\text{Pb}$.
- Theory uncertainties can be substantially reduced further by supplementing $^{16}\text{O}^{16}\text{O}$ collisions with $^{20}\text{Ne}^{20}\text{Ne}$ collisions.
 - We can bring the initial geometry under control to the percent level!
 - We can use this as a probe for non-hydro contribution to collectivity!



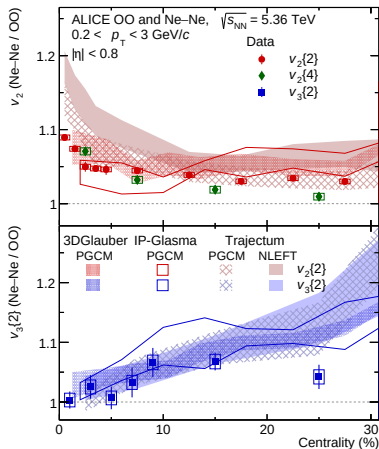
How do these predictions hold up?



- We answer some questions...
 - Hydro agrees broadly with data.
- ...and we also got new questions!
 - NLEFT *Trajectory* describes OO, but has tensions in the ratio.
 - PGCM *Trajectory* describes the ratio better, but has tensions in OO.
 - IP-Glasma is compatible with the ratio, but does not have the central enhancement and does not describe OO (see ATLAS results).
 - No model describes all the data!
 - Without NeNe data, there would not be enough precision to see these tensions.



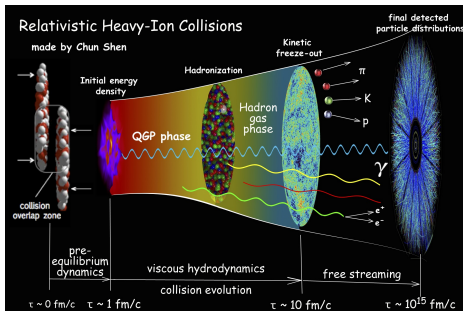
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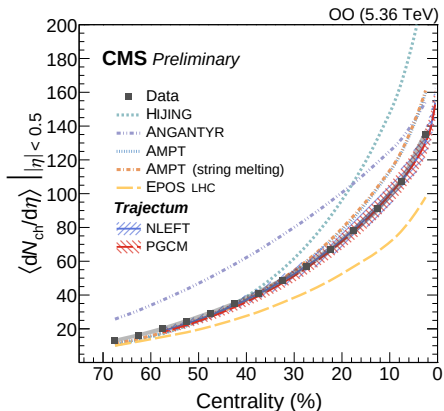
Where do these models differ?



- Details of the initial state and pre-hydrodynamic stage matter!
- Opportunity to learn about the stages we understand least!
- We can fit to the new data, with the aim of revealing precisely these aspects.



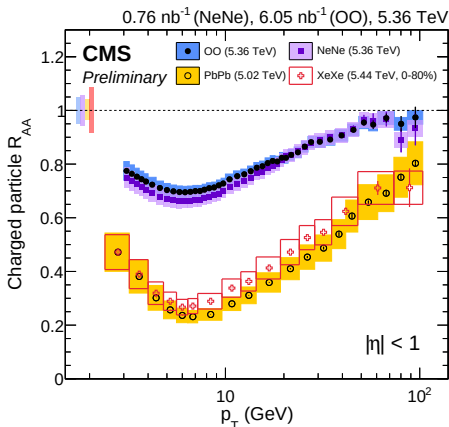
We got lots of other results too!



- Multiplicity agrees very well with *Trajectum*, and shows large differences between models!
- R_{AA} shows a difference between NeNe and OO.
 - Surprisingly well described by a path length based scaling of the PbPb R_{AA} .



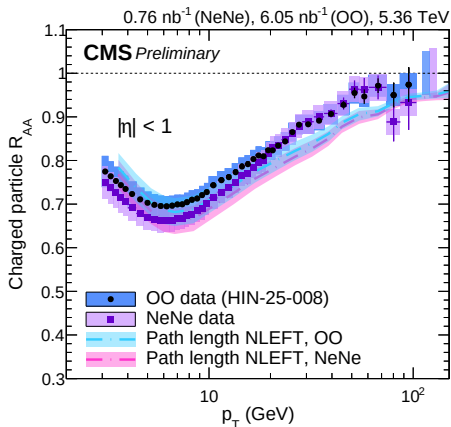
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The deformation of ^{129}Xe

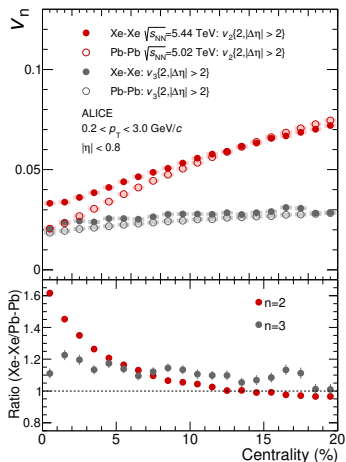
- We describe ^{129}Xe with a *deformed* Woods-Saxon distribution:

$$\rho_{\text{WS}}(r) \propto \frac{1}{1 + \exp\left(\frac{r - R(\theta, \phi)}{a}\right)},$$

$$R(\theta, \phi) = R \left(1 + \beta_2 \cos \gamma Y_2^0 + \beta_2 \sin \gamma Y_2^2 \right),$$

with Y_l^m spherical harmonics.

- We do not distinguish protons and neutrons.
- Observables in the central region are sensitive to β_2 and γ :
 - v_2 is sensitive to β_2 .
 - ρ_2 is sensitive to β_2 and γ .



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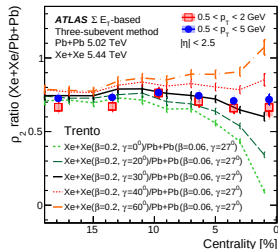
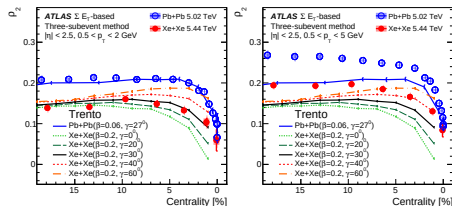
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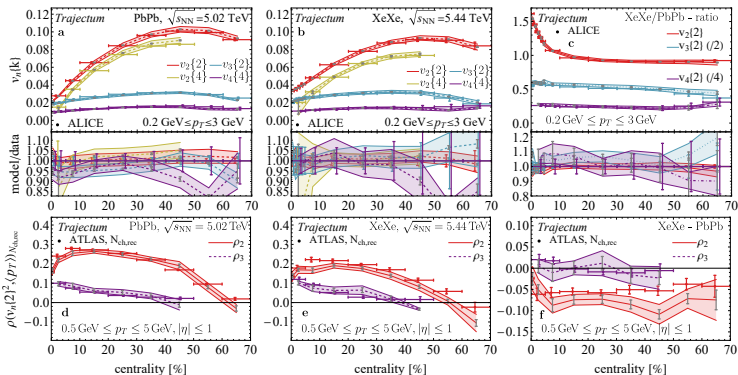
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A selection of posterior observables

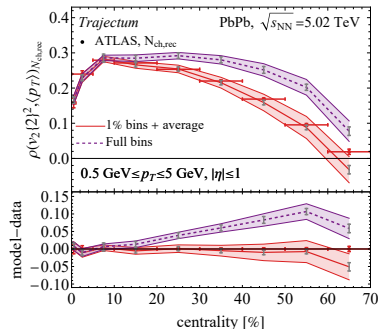


- Precise description of $v_n\{k\}$ and ρ_2 .
 - We fit PbPb as well as XeXe data.
 - For ρ_2 we also fit to the difference of XeXe compared to PbPb, for $v_n\{k\}$ we fit to the ratio.



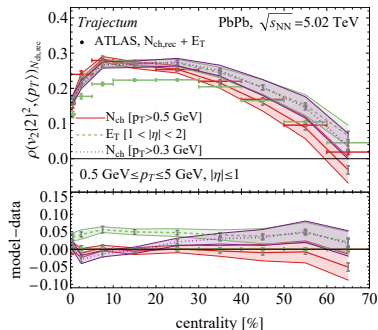
Details are important!

- Some observables, such as ρ_2 , depend sensitively on the measurement details.
- Bin size matters:
 - Measurement of ρ is done in small centrality bins.
 - Should compute in small bins, then average.
- Definition of centrality:
 - Different definitions give different answers.
 - In this case, many definitions are available.
 - Should compare to a measurement which uses a definition that our model can reliably compute.
- Ignoring these details can result in biased answers!



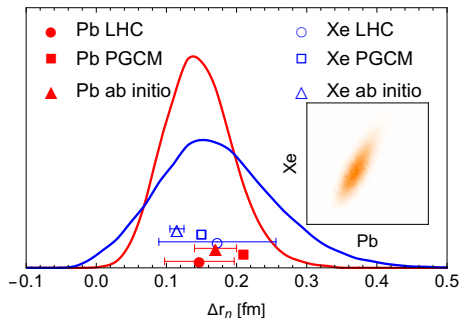
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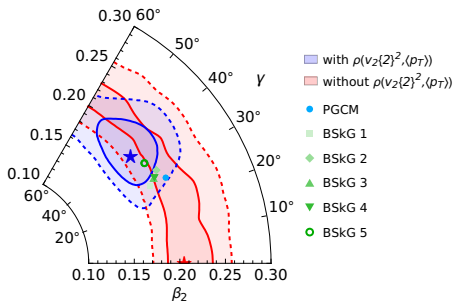
Bayesian analysis result using LHC data

- We are able to extract the skin depth a .
 - Using the known charge radius, we can convert to neutron skin.
 - First experimental measurement of the neutron skin of ^{129}Xe .
 - The neutron skins of ^{208}Pb and ^{129}Xe are correlated.
- The deformation parameters β_2 and γ are well constrained.
- ^{129}Xe is truly triaxial, shaped like a kiwi.
- Potential applications to $0\nu\beta\beta$ decay.



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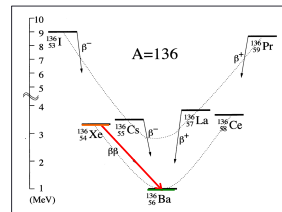


Neutrinoless double beta decay

- For certain nuclei, single beta decay is kinematically forbidden, but double beta decay is allowed.
- If the neutrino is a Majorana particle, then the two neutrinos can annihilate: neutrinoless double beta decay.
 - In the spectrum for the electrons, this would show up as a peak at the end.
- The decay rate is given by:

$$\Gamma \propto |M|^2 m_{\beta\beta}^2.$$

- $m_{\beta\beta}$ is the *effective Majorana mass* of the neutrinos.
- $|M|$ is a *nuclear matrix element* for this decay.
- $|M|$ is very complicated to compute, and cannot be measured by a different process.

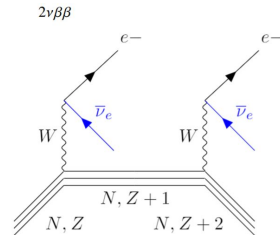


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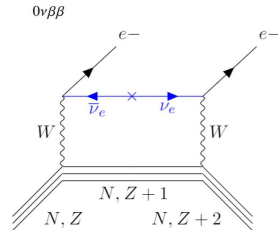


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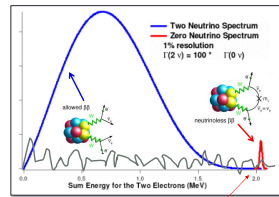


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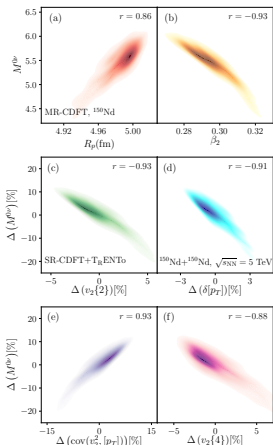
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Nuclear matrix elements from heavy ion collisions



- Nuclear matrix elements are very correlated with the *difference* in shape of the nuclei.
 - This leaves imprints in heavy ion collisions of these nuclei.
 - The nuclei involved are isobar pairs, making extracting the difference in shape easier.
 - The ‘shape’ can be made well-defined using multinucleon correlators.
- We are currently expanding on this, and will conduct a sensitivity study for $^{76}\text{Ge} \mapsto ^{76}\text{Se}$.



The power of short runs

- The ^{129}Xe run took just 6h of data.
 - We can quantitatively extract its shape.
 - Proof of concept for similar future studies with big physics impact.
- The ^{20}Ne run took 12h.
 - Broad agreement with hydro.
 - Tensions provide new questions.
 - Should we next go smaller? Larger?
- What can we learn from future short runs?
 - ^{76}Ge and ^{76}Se
 - ^{48}Ca and ^{40}Ca
 - ^4He
 - ^{24}Mg
- We have the opportunity to shape the future of the LHC ion program!



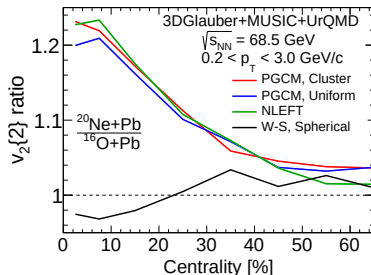
Backup

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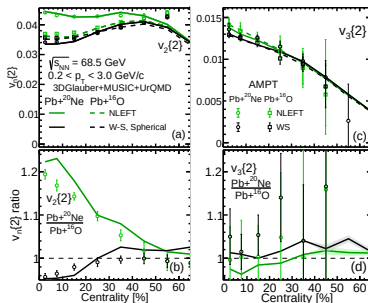
Opportunities at LHCb

- The SMOG2 system at LHCb is able to inject gas into the beampipe.
 - This allows to perform collisions at $\sqrt{s} = 68.5$ GeV.
- Similar to the NeNe/OPb ratio, one can take a NePb/OPb v_2 ratio.
 - A large signal is predicted by multiple models.
- Challenges:
 - Injected Ne gas is not isotopically pure.
 - Injecting O has not been approved.
- ^{16}O and ^{20}Ne beams would allow for several new collision systems:
 - $^{16}\text{O}^{16}\text{O}$ (if O is approved for SMOG2),
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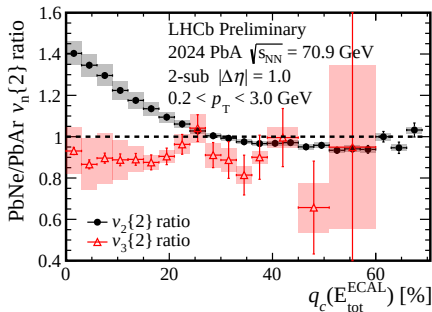
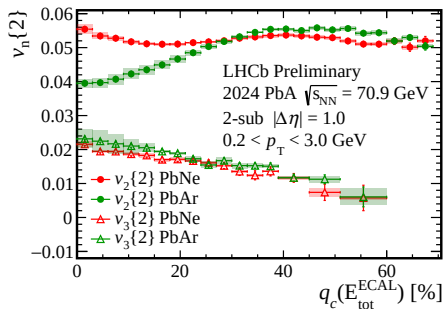


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How do predictions hold up?



- Measured v_2 , v_3 ratios between NeNe and OO agree well with model predictions.

