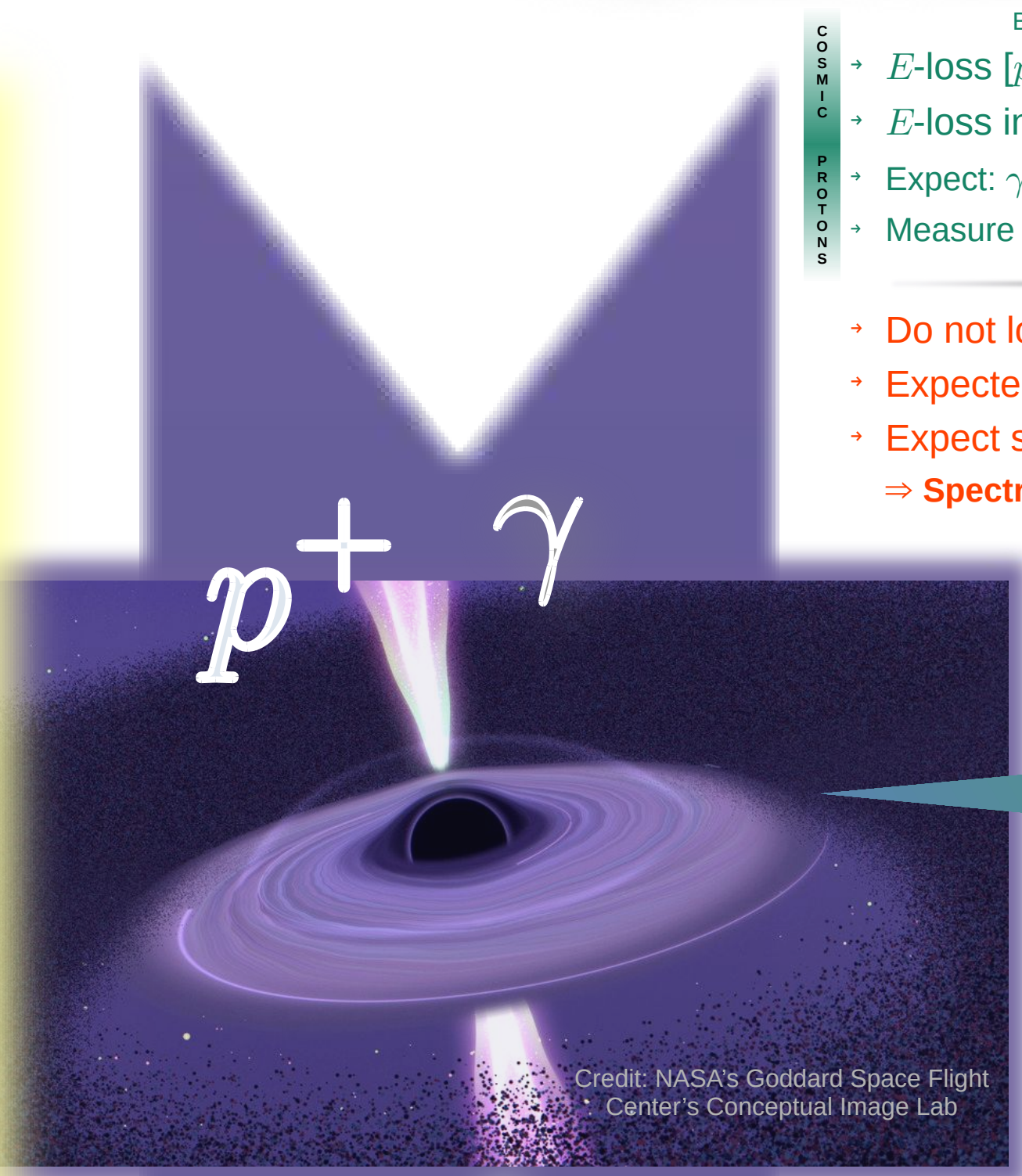


Spectral Signatures of Multi-Mediator Neutrino Self-Interactions at Ultra-High Energies

Ayushi Kaushik[†], Atri Bhattacharya, Kenji Nishiwaki

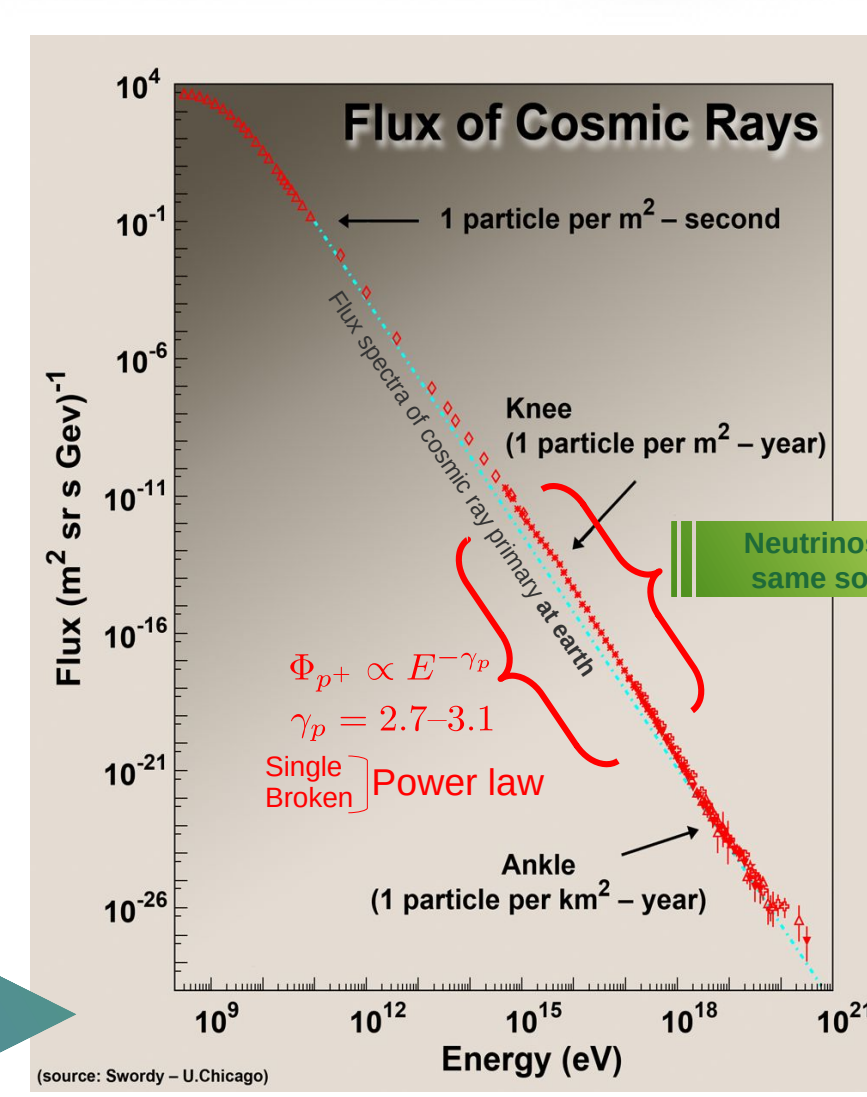
SHIV NADAR
INSTITUTION OF EMINENCE DEEMED TO BE
UNIVERSITY
DELHI NCR

PRODUCTION



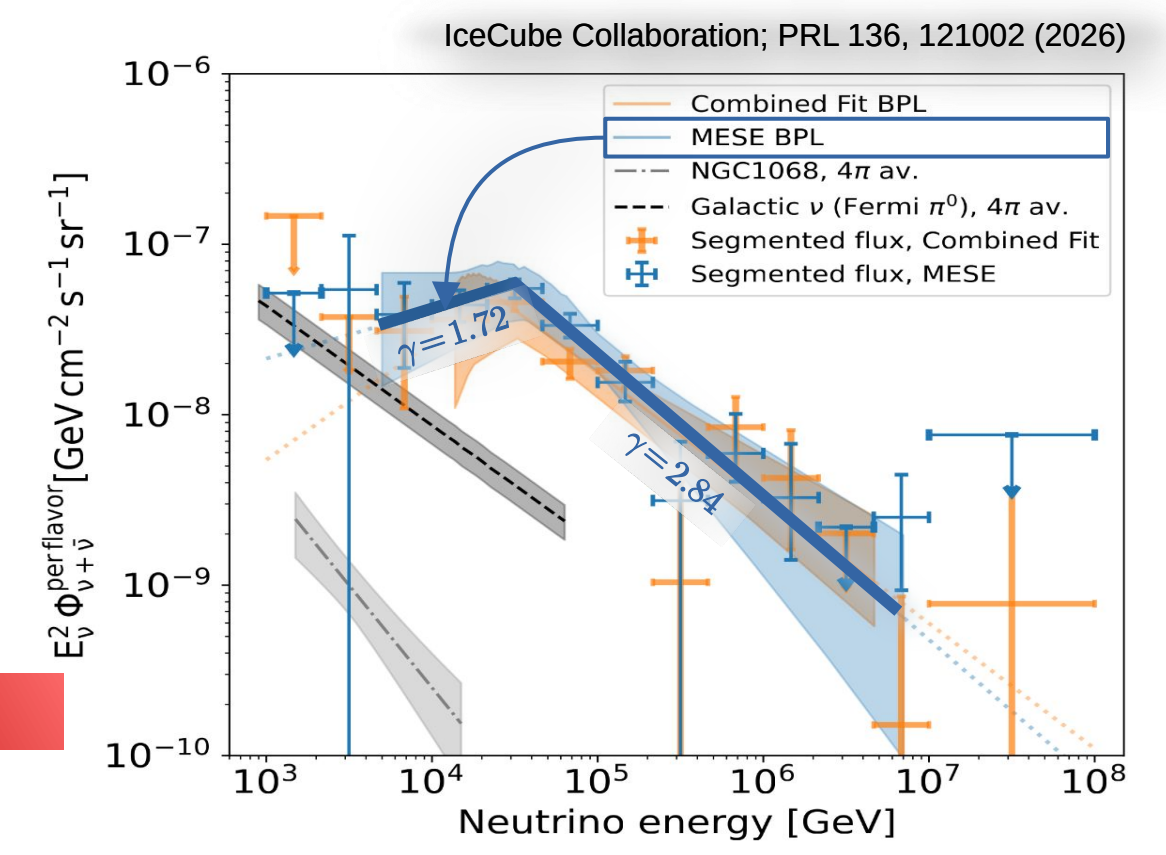
- Berezinsky, et al. PRD 74, 043005 (2006)
- E -loss [$p^+ + \gamma_{\text{CMB}}$] during propagation
- E -loss increases with p energies
- Expect: $\gamma_p @ \text{Earth} > \gamma_p @ \text{Source}$
- Measure $\gamma_p \approx 2.7-3.1 \Rightarrow \text{Source } \gamma_p \approx 2.0-2.4$

- Do not lose energy (red-shift only)
- Expected from same sources
- Expect source $\gamma_p \approx 2.0-2.4$
- $\Rightarrow \text{Spectral index @ Earth: } \gamma_p \approx 2.0-2.4$



Diffuse Astrophysical ν Flux

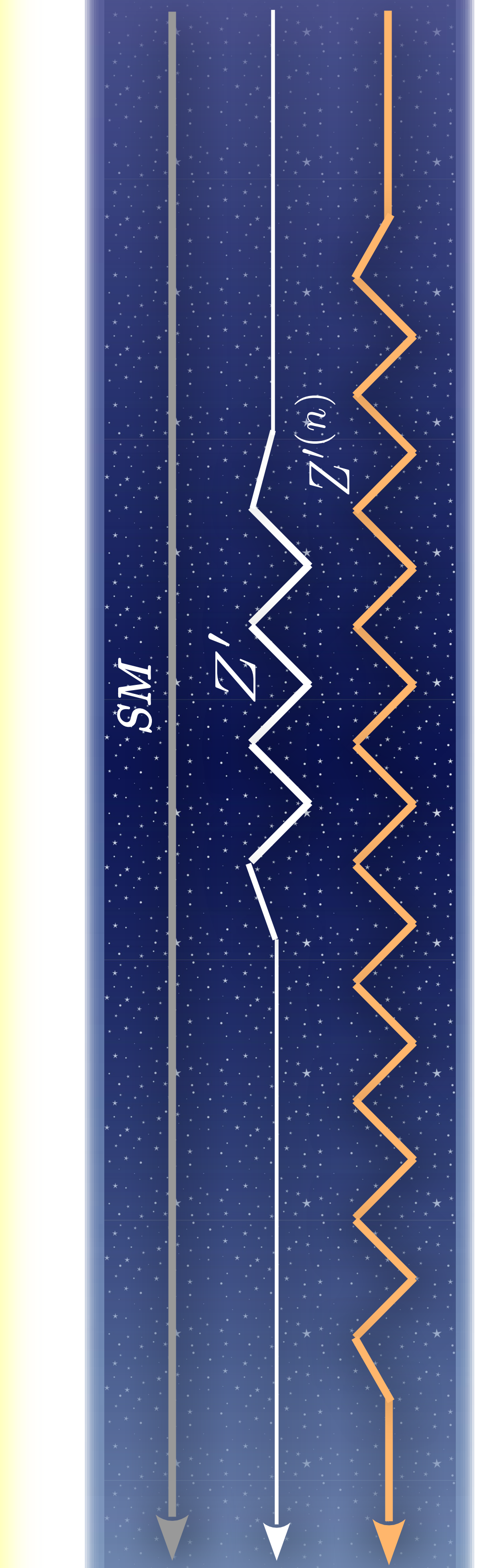
Expectation	Reality
$\nu_e : \nu_\mu : \nu_\tau \approx 1 : 1 : 1$	✓
Isotropic flux	✓
$\Phi_\nu \propto E^{-\gamma}$ power-law	✗ SPL: $\gamma \sim 2.8$
$\gamma \sim 2$ spectral index @source	✓ BPL
E -loss $\Rightarrow \gamma \sim 2 @ \text{Earth}$	Single/Broken Power Law



MOTIVATION

- Assume: Hard ν spectra at source: $\gamma \sim 2$
- Attenuation during propagation $\Rightarrow$ Spectral softening to $\gamma \sim 2.8$
- Target: CvB relic ν
- Problem: Heavy SM med $M_Z = 91.2 \text{ GeV}$
- Solution: BSM physics with light Z'

PROPAGATION



ATTENUATION AGAINST COSMIC ν BACKGROUND

Standard Model

- Propagation of UHE ν through CvB happens as if through vacuum
- Vacuum oscillation between ν flavours
- No attenuation of flux $\Rightarrow$ No spectral impact

The Hunt for Low-Mass Gauge Bosons Broken $U(1)_{L\mu-L\tau}$ gauge symmetry in 4d

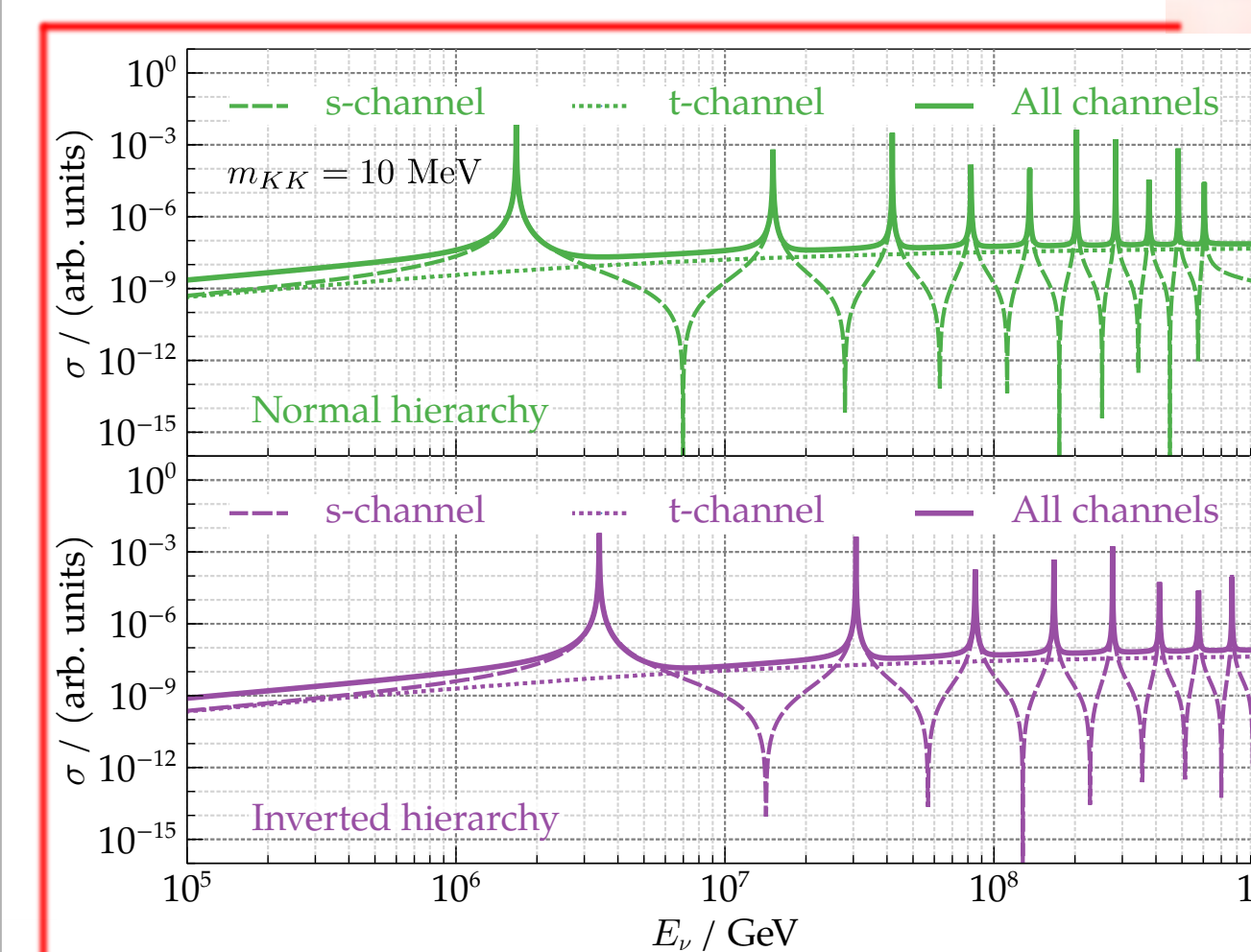
- Anomaly free 4d BSM with single goldstone boson
- Symmetry breaking $\Rightarrow$ One new massive boson [Z']
- MeV-scale masses for Z' feasible
 - $\Rightarrow$ PeV-energy ν_{ast} scattering against ν_{CvB}
 - $\Rightarrow$ On-shell Z' from s-channel resonance
 - $\Rightarrow$ Sensitive to ν mass Hierarchy: Normal vs Inverted
- Possible to probe at UHE ν detectors like IceCube...
 - $\Rightarrow$...not enough attenuation to soften PeV spectra
 - $\Rightarrow$ Only three peaks for any given Z' mass
 - $\Rightarrow$ Requires soft ν flux spectrum at source

Multiple MeV-scale Gauge Bosons Broken $U(1)_{L\mu-L\tau}$ gauge symmetry in KK 5d

Single 5d gauge field $\Rightarrow$ Tower of gauge bosons in 4d

$$m_{Z^{(n)}} = (2n-1)m_{KK}$$

$n=1, 2, 3, \dots$ m_{KK} values with increasing n [log(E) scale]

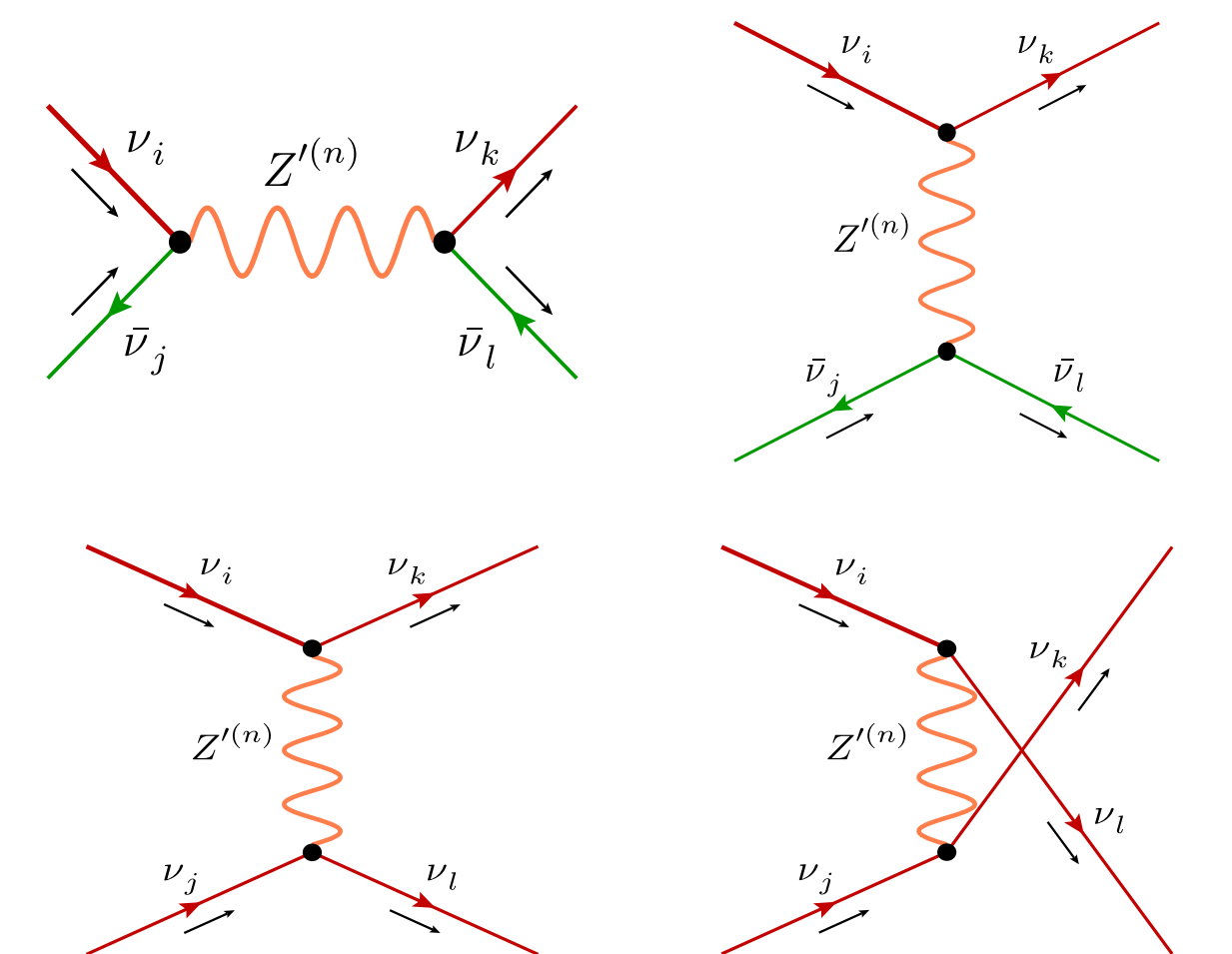


- Multiple resonances in $\sigma(\nu_{\text{ast}} \nu_i \rightarrow \nu_j \nu_k)$
- Interference: Amp mediated by $Z^{(n)}$ $n=1, 2, \dots$
- Destructive interference between s-channel resonances
- Dips in cross-section not seen in single Z' case (only single s-channel)
- Non-resonant t/u channels wipe out interference dips
- Must include all possible t/u channel diagrams

Evolving ν flux using Boltzmann equation

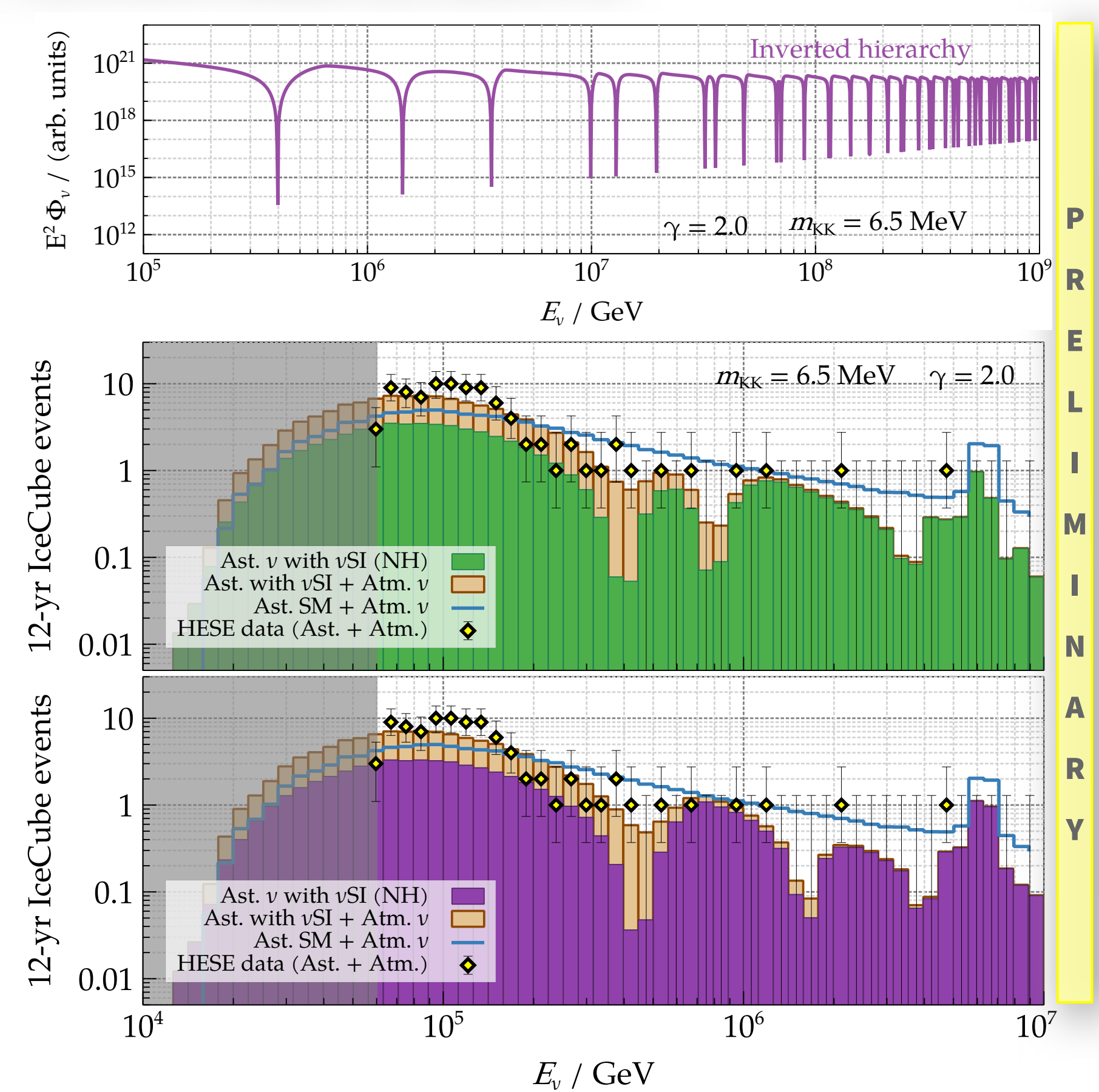
$$-(1+z) \frac{H(z)}{c} \frac{d\phi_i}{dz} = \left\{ \begin{array}{l} \text{Flux injection } \gamma \propto 2 \\ J_i(E_0, z) - \phi_i \sum_j \langle n_{\nu_i}(z) \sigma_{ij}(E_0, z) \rangle \\ + \int_{E_0}^{\infty} dE' \sum_{j,k} \phi_k \left\langle n_{\nu_j}(z) \frac{d\sigma_{kj}(E', z)}{dE_0} \right\rangle \end{array} \right.$$

Neutrino flux (blue arrow), Flux injection (purple box), Annihilation (red box), Regeneration (green box), Red-shift (orange arrow).



EFFECT ON ICECUBE HIGH ENERGY STARTING EVENTS

- *Scattering via multiple MeV-mass mediators
- *Tower of gauge bosons
 - $\Rightarrow$ Multiple s-channel resonances in x-section
 - $\Rightarrow$ Multiple dips in flux spectrum
- *Narrower resonances with increasing Z' mass
- *Repeated resonant attenuation with increasing astrophysical ν energy
 - $\Rightarrow$ Natural softening of spectral shape at 1 PeV+
 - $\Rightarrow$ Consistency with HESE data even when originating with hard spectral index $\gamma \approx 2$
- *Rich phenomenology with impact on other searches: MuSIC, NA64, $F\mu\text{BD}$ etc. ...



Ref

Main: A. Bhattacharya, A. Kaushik, and K. Nishiwaki;
Work-in-progress

Related: Hooper et al, PRD 108,023007 (2023); nuSIProp PRD 104, 123014 (2021); D. Chakraborty et al PRD 110, 095030 (2024) and PRD 113, 115038 (2026)

[†]ak356@snu.edu.in