

Modeling the nHz gravitational wave background with t-process

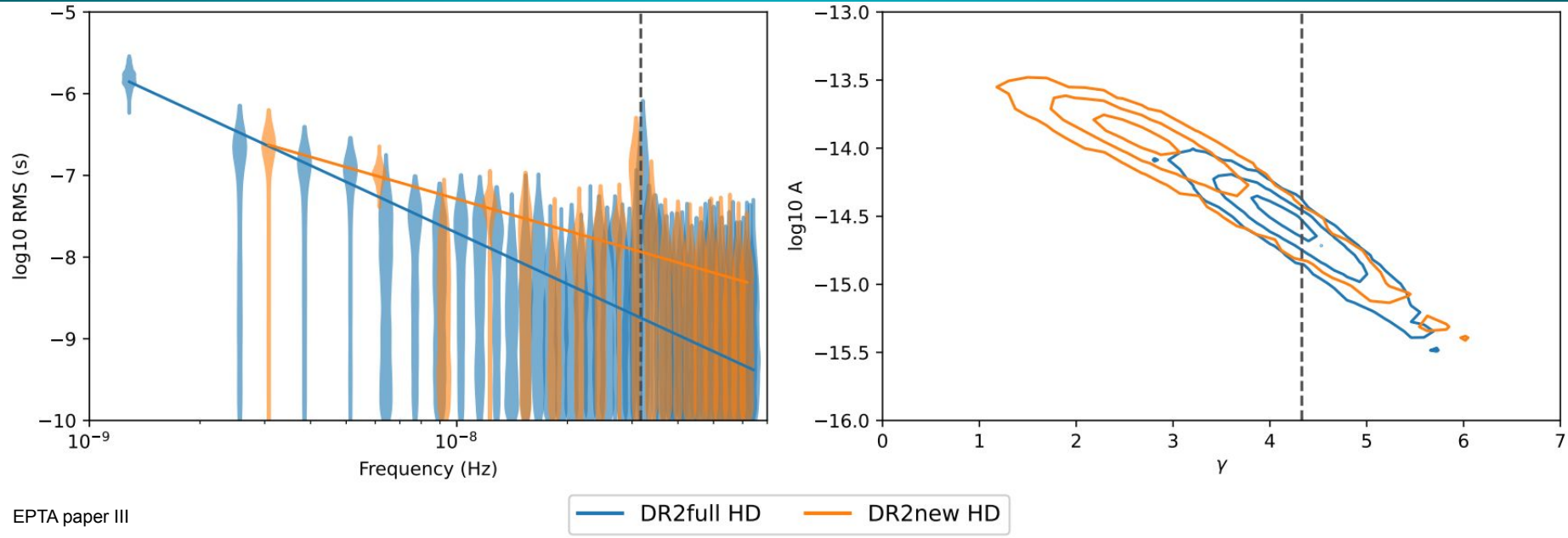
Beatrice Eleonora Moreschi

Third year PhD student

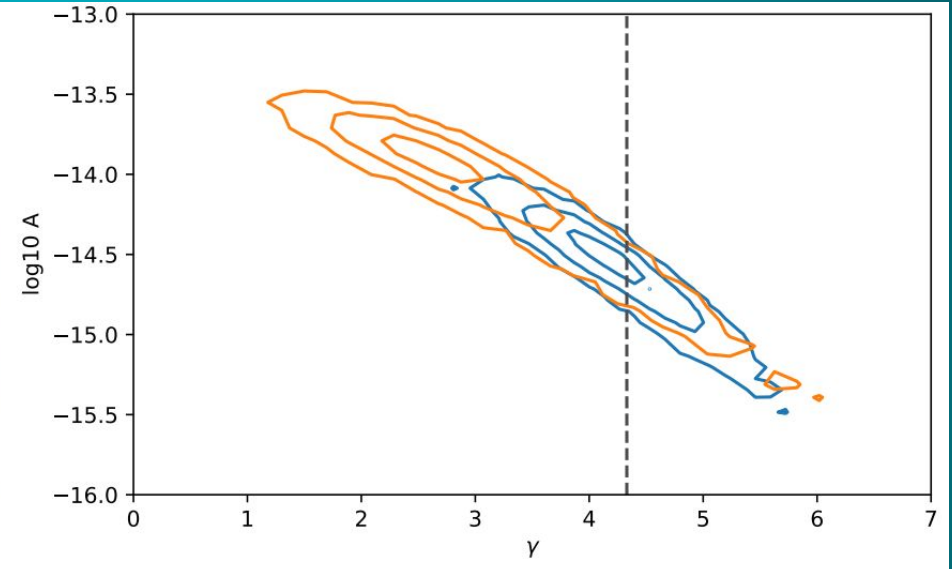
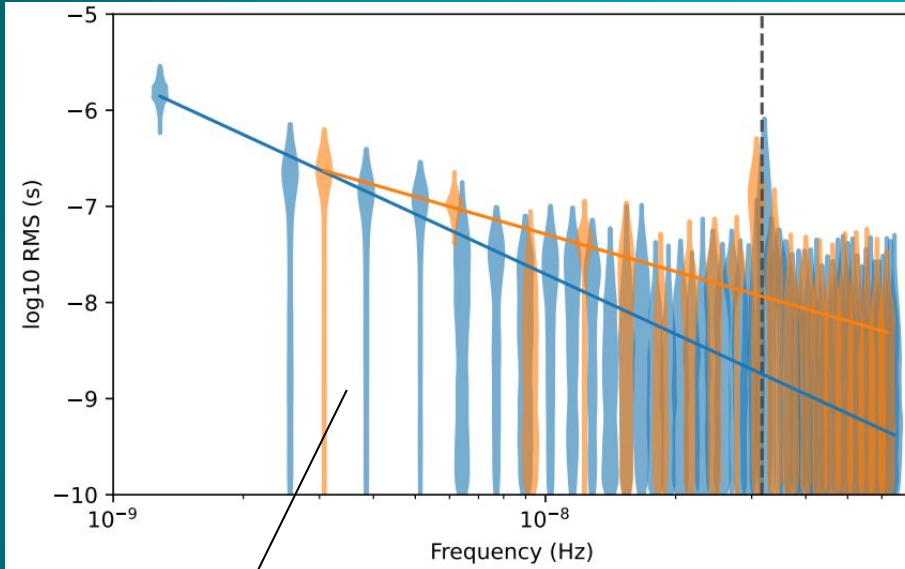
1st BiCoQ Conference: from gravity to particles

17/06/2026

What is the t-process?



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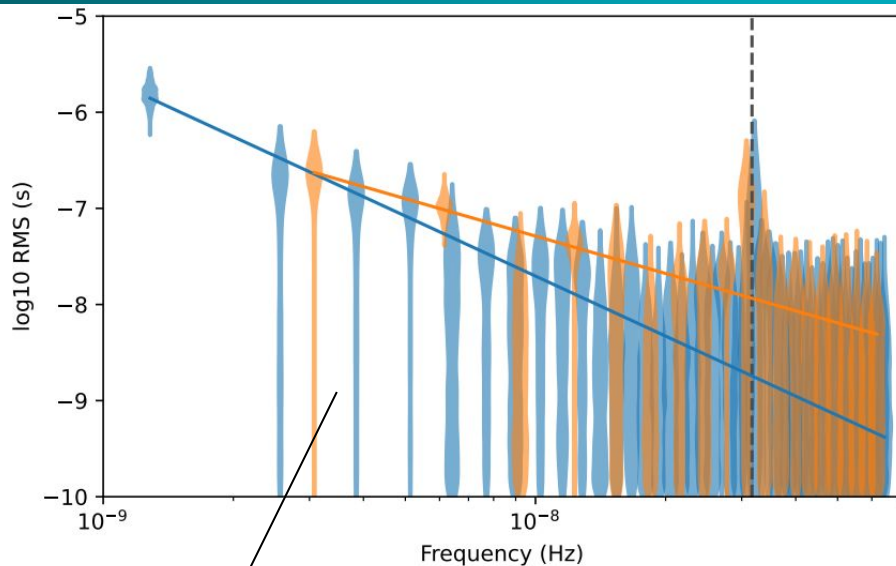
EPTA paper III

$$P_{fs}(f_i) = \frac{h_c^2(f_i)}{12\pi^2 f_i^3} df = \delta t_{\text{delay}}^2(f_i)$$

DR2new HD

Agnostic model, it allows for deviations in the PSD from a power law

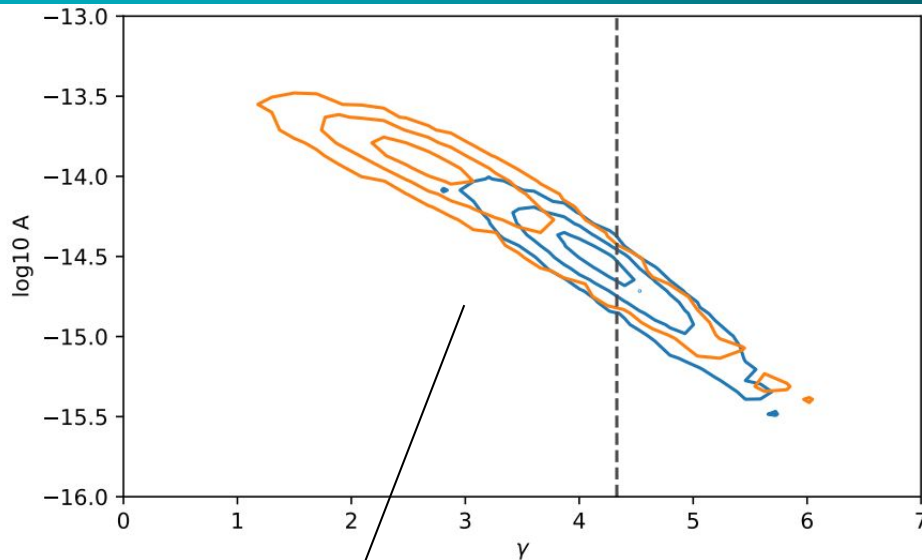
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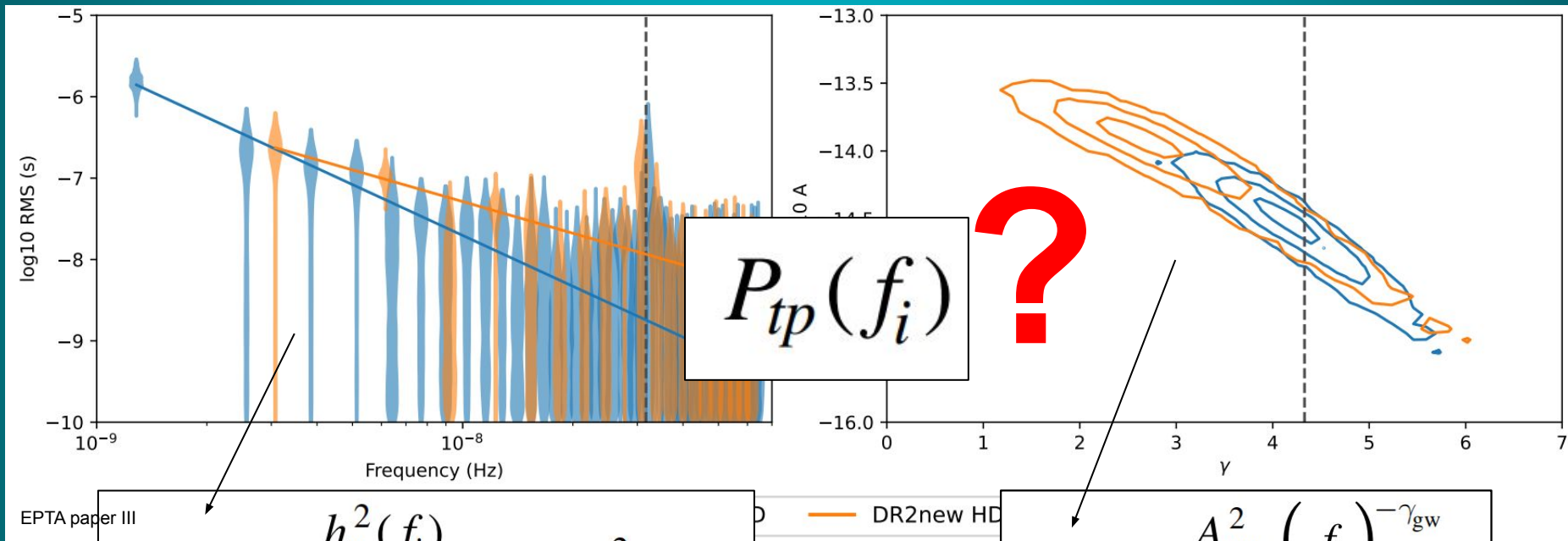


DR2new HD

$$P_{pl}(f) = \frac{A_{\text{gw}}^2}{12\pi^2} \left(\frac{f}{f_{\text{yr}}} \right)^{-\gamma_{\text{gw}}} df$$

Relationship between the power at different frequencies

What is the t-process?



$$P_{tp}(f_i)$$



EPTA paper III

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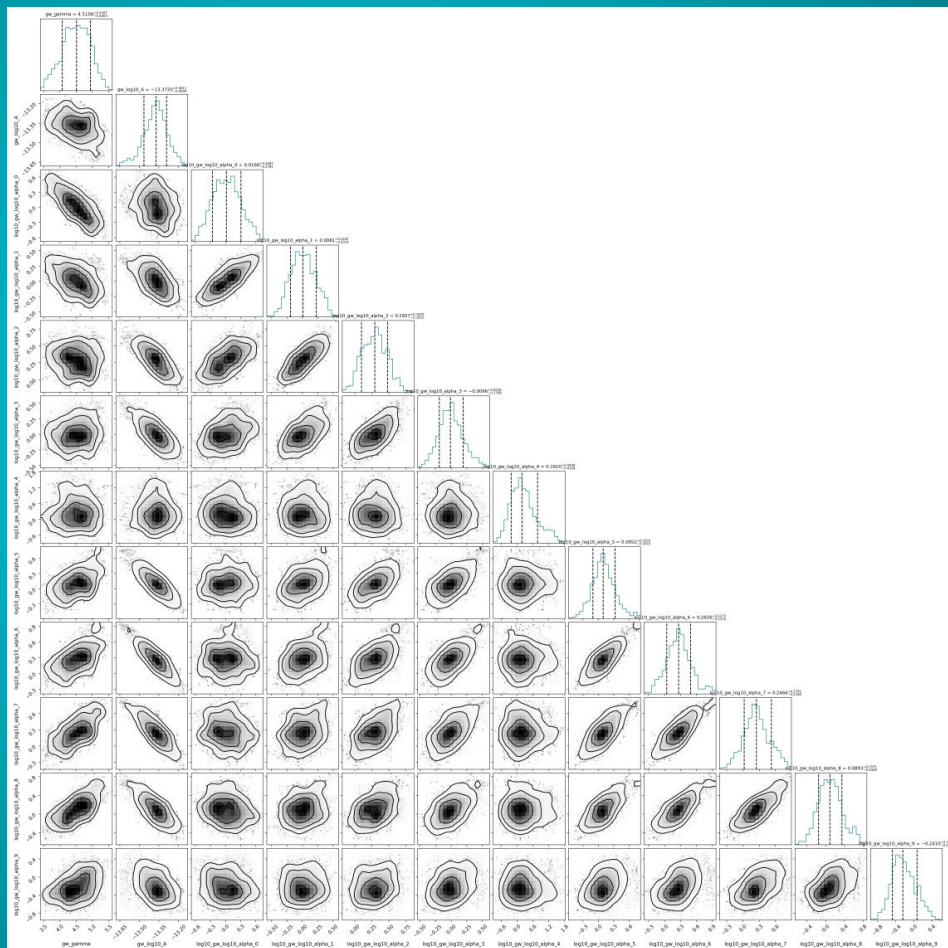
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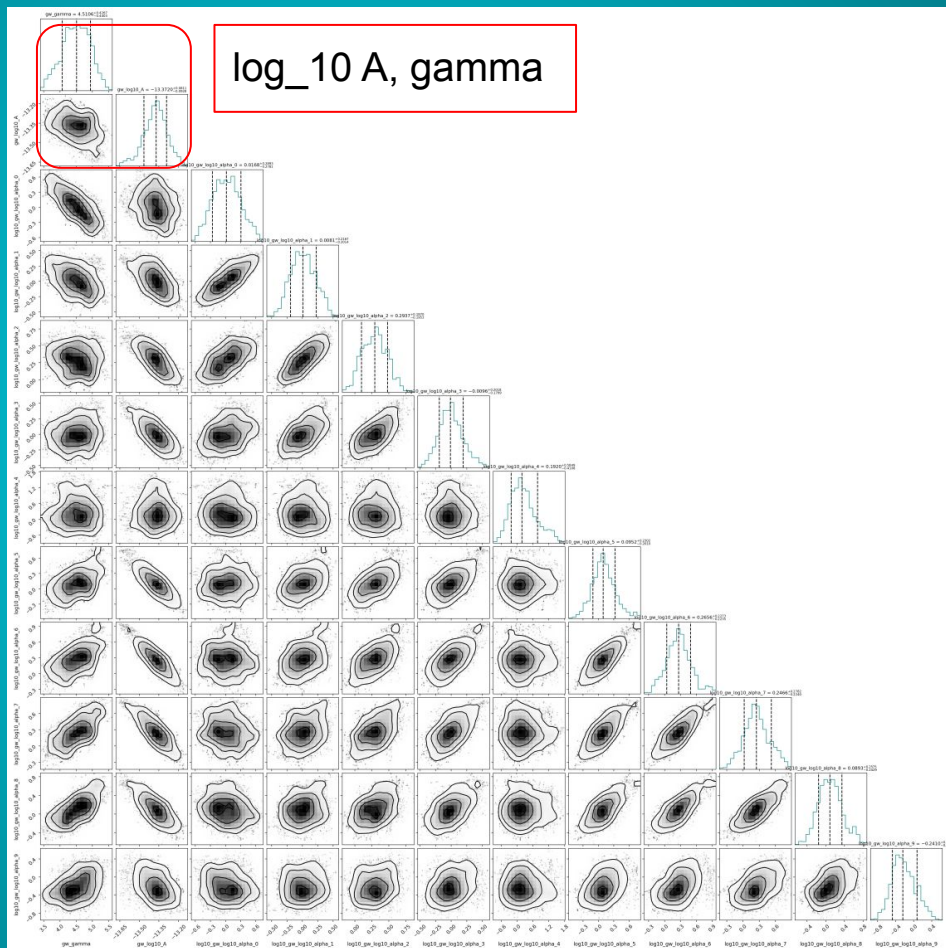
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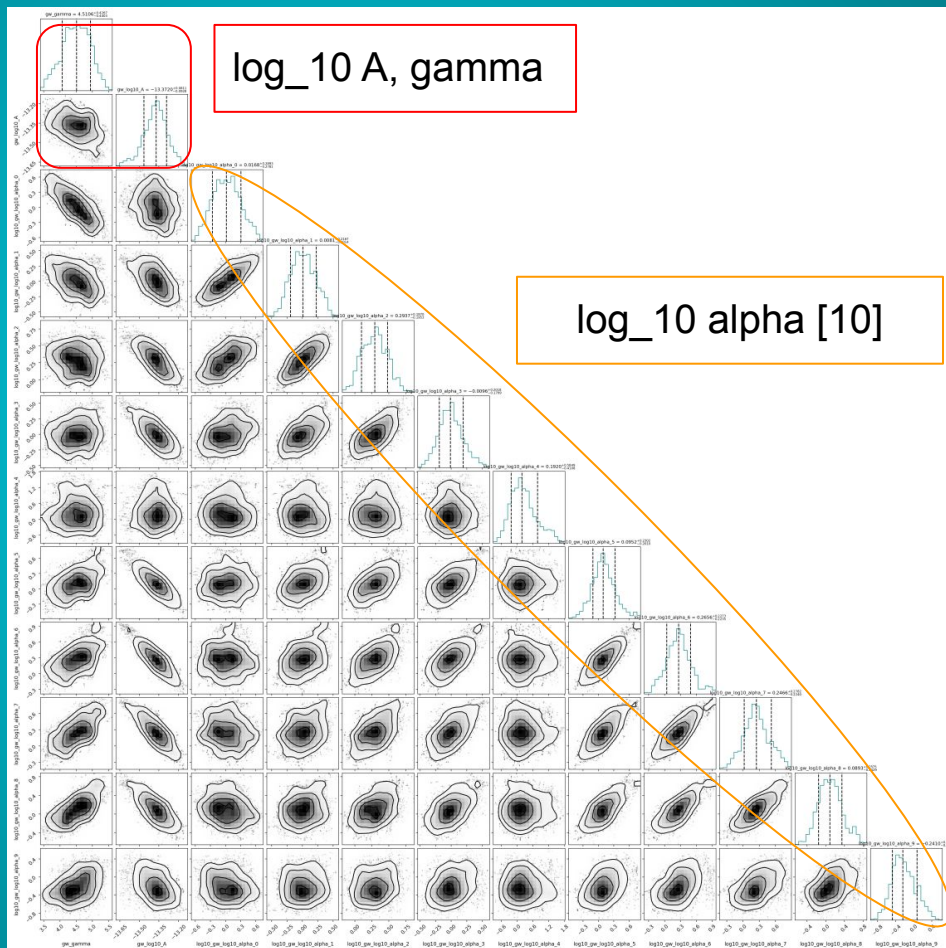
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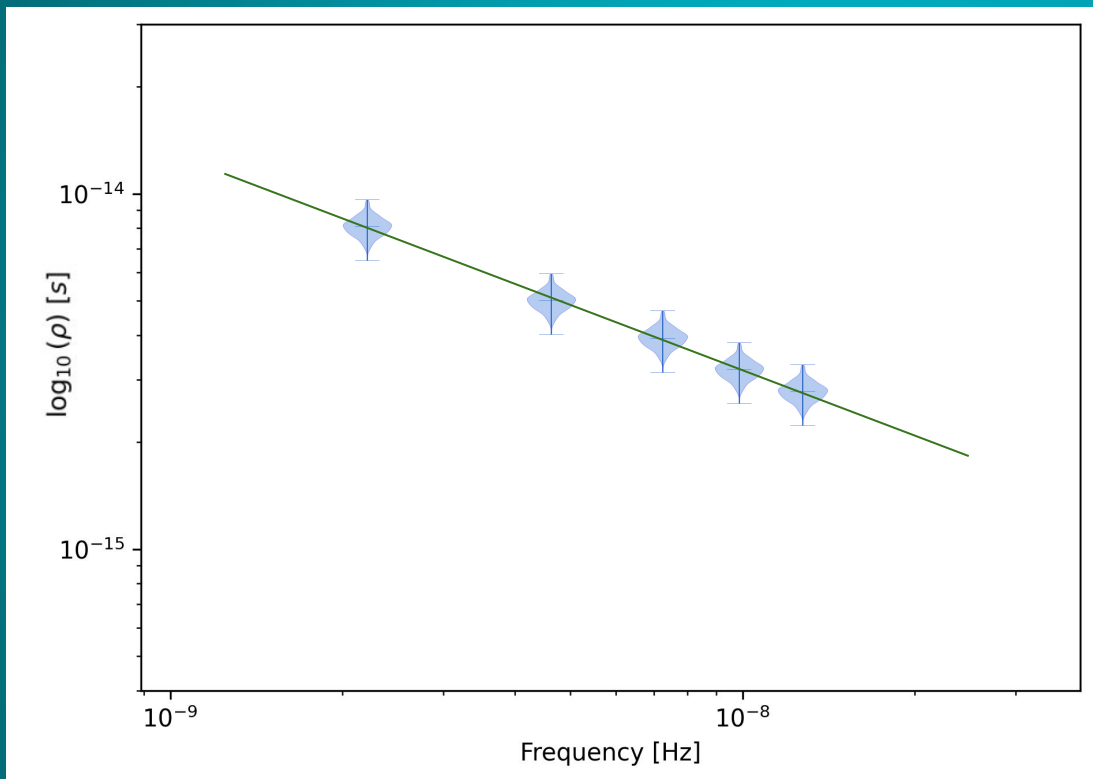


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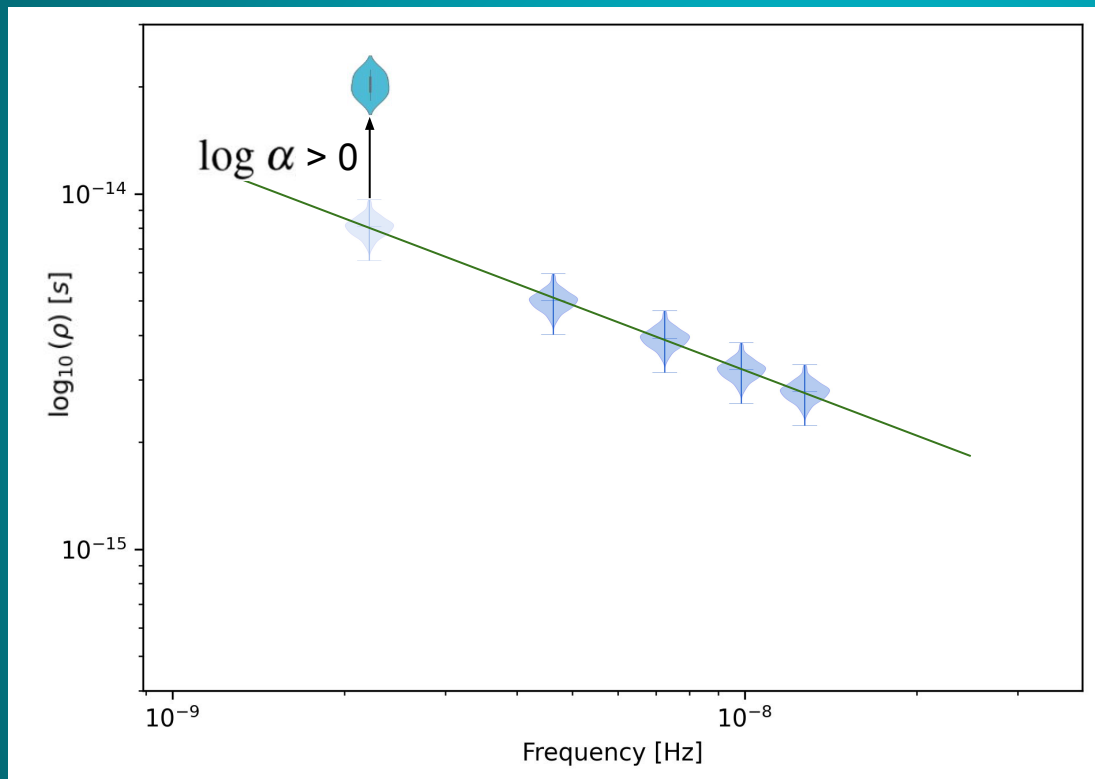
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- K. Aggarwal et al. 2019 (PSR J0613-0200 has unmodeled signal)

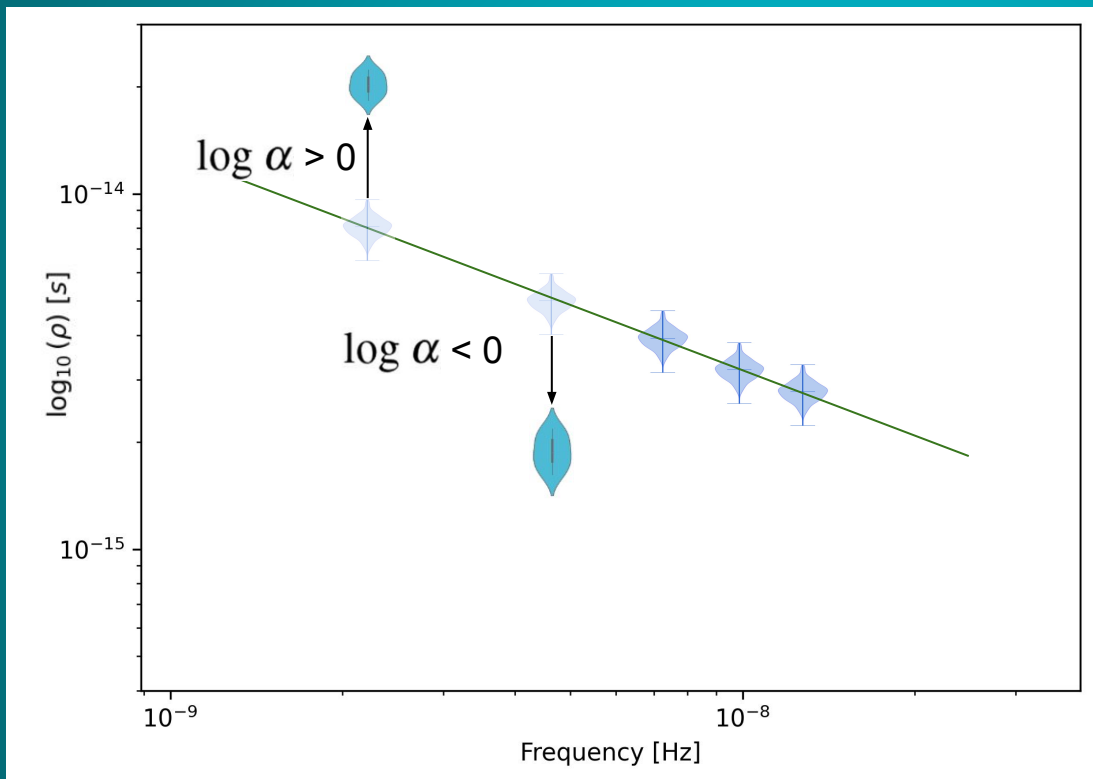
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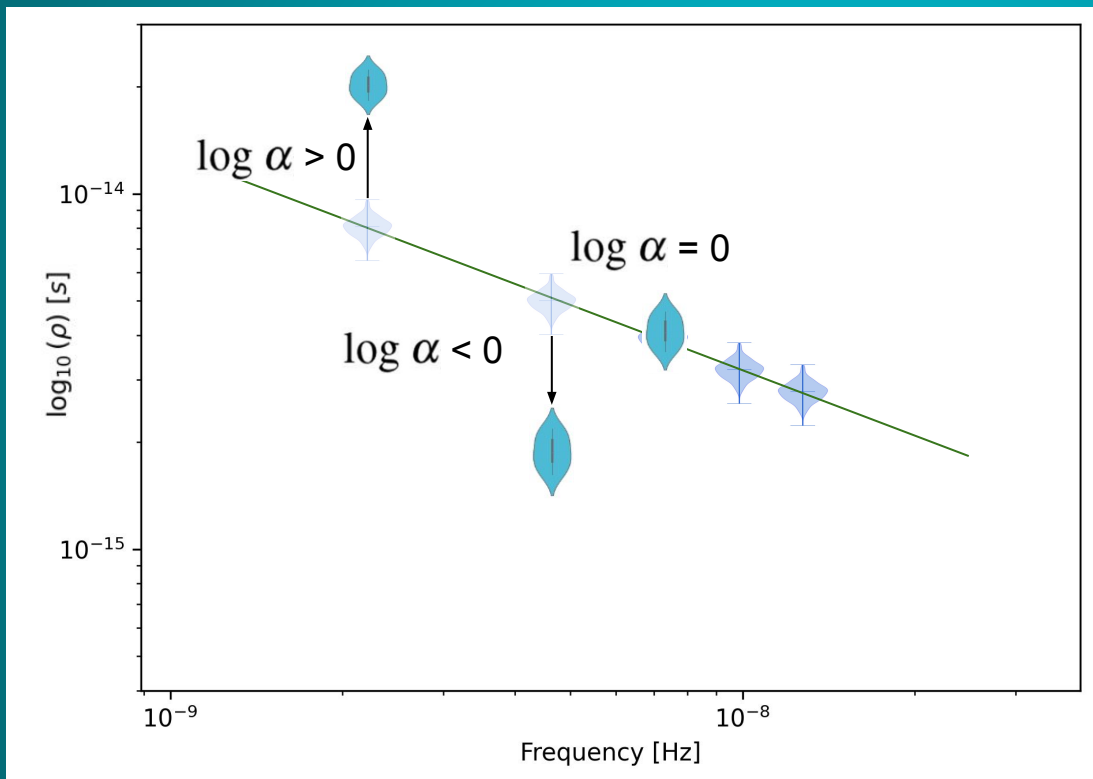
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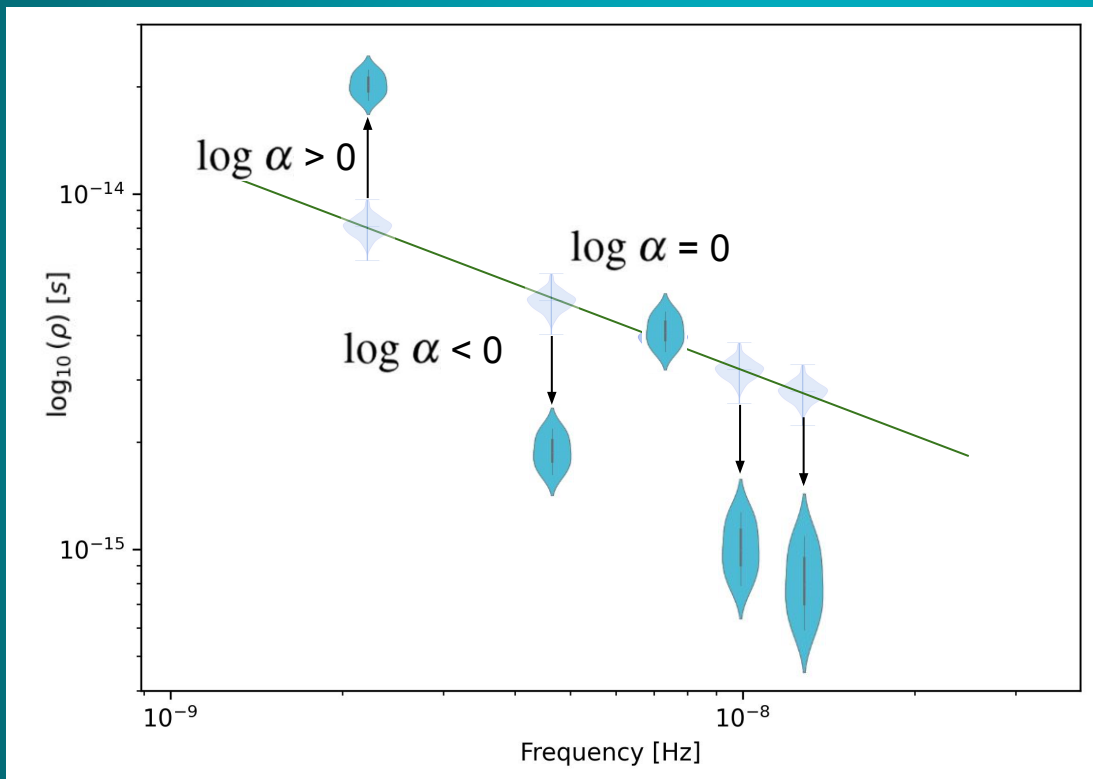
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$$P_{tp}(f_i) = \alpha_i P_{pl}(f_i)$$

$$\alpha_i \sim \text{invgam}(1, 1).$$

The inverse gamma function follows a probability of

$$\text{invgam}(\alpha_i, 1, 1) = \frac{1}{\Gamma(1)} \alpha_i^{-2} \exp\left(-\frac{1}{\alpha_i}\right),$$

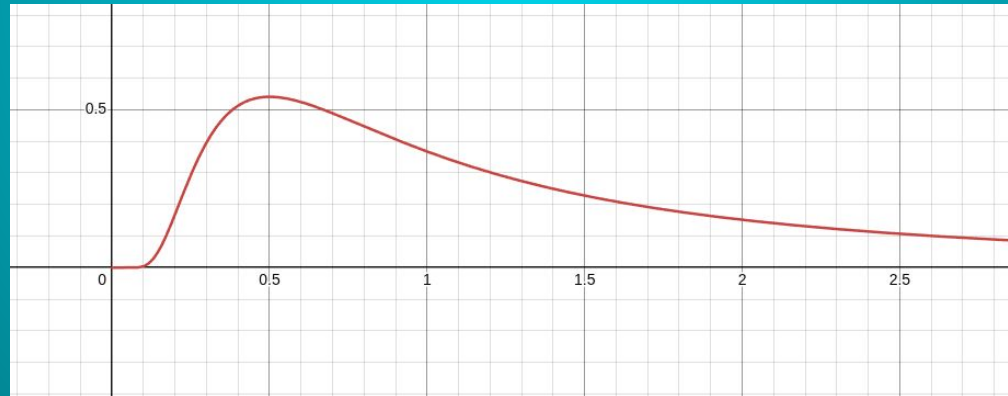
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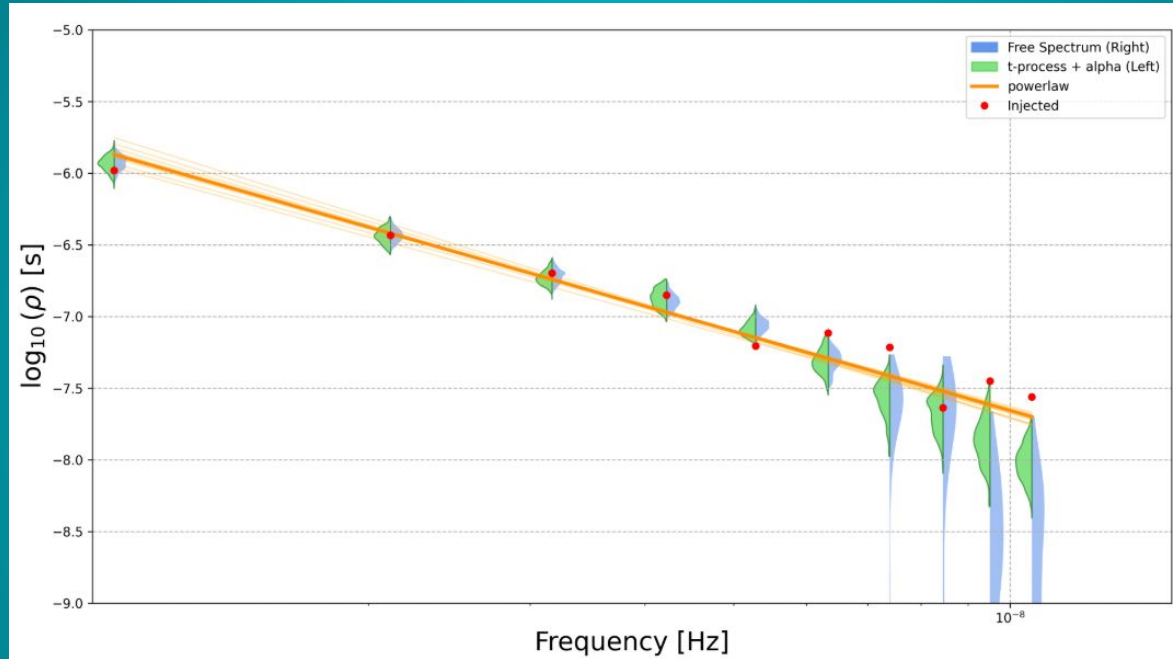
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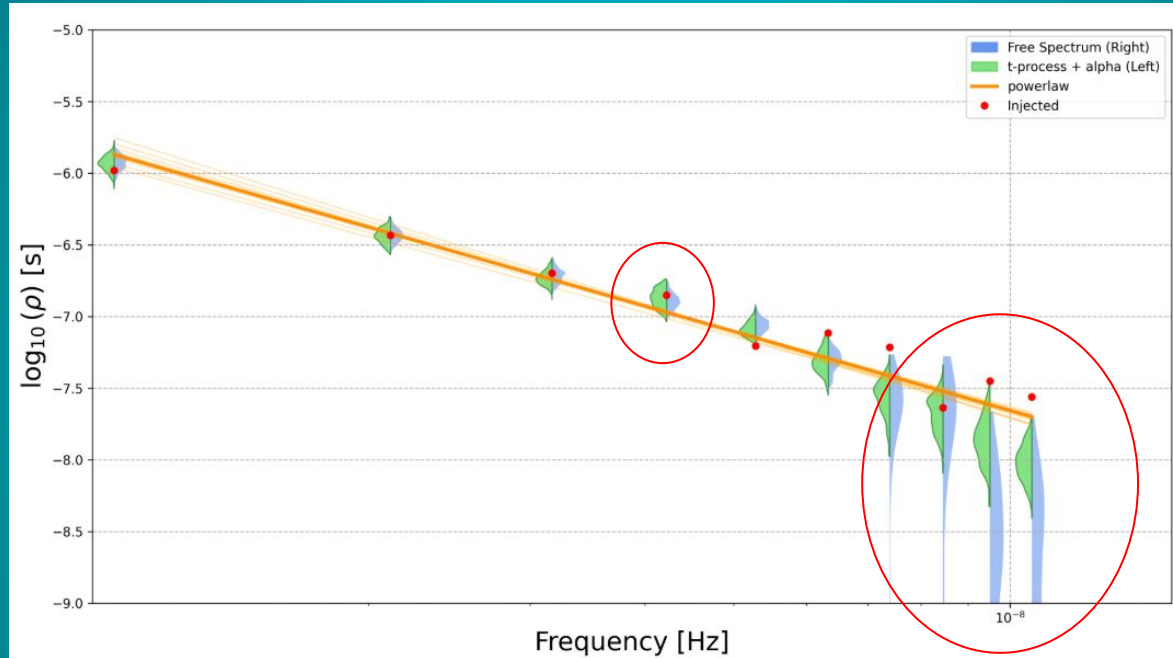
Test on simulated dataset



Real population
 $e = 0.9$
+ strong environment effects

20 pulsars
10 bins

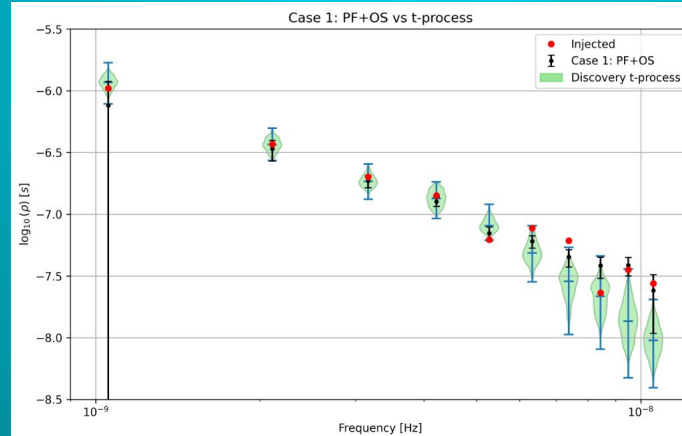
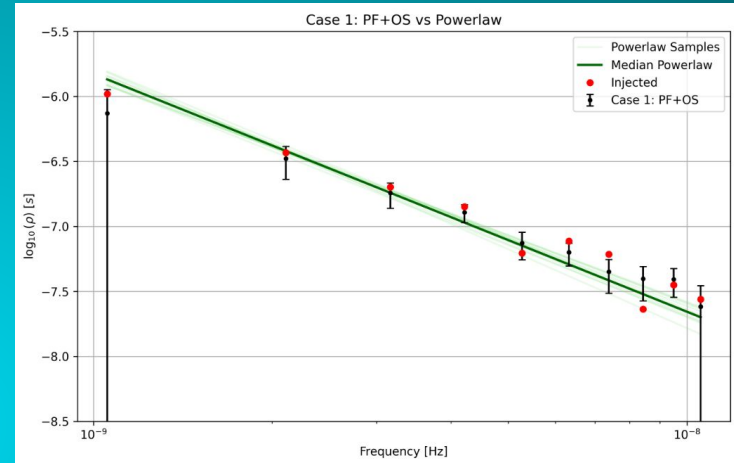
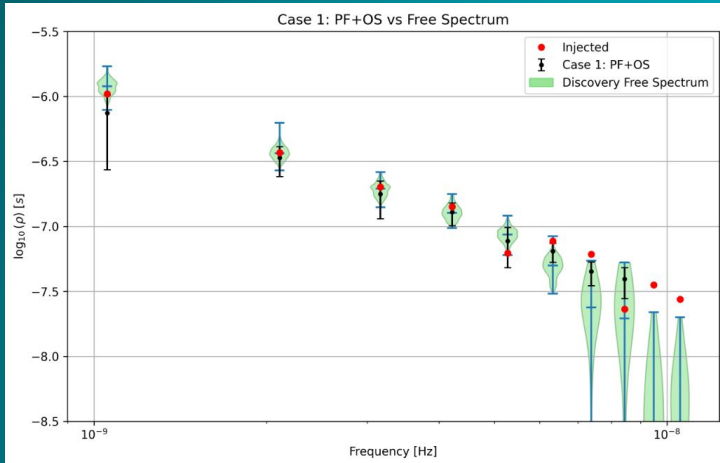
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PFOS on simulated dataset

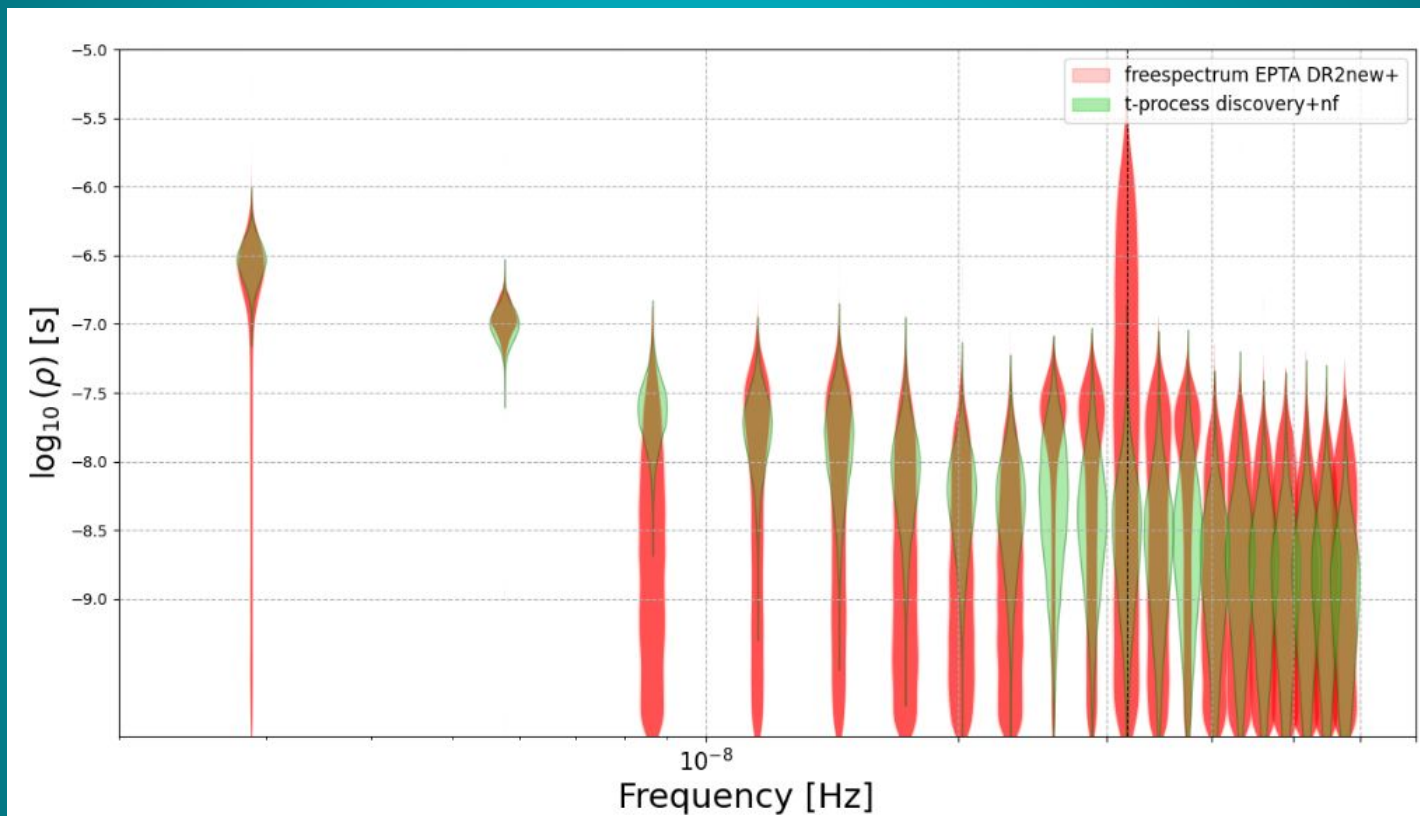


T-process and freespectrum are able to model better CGW

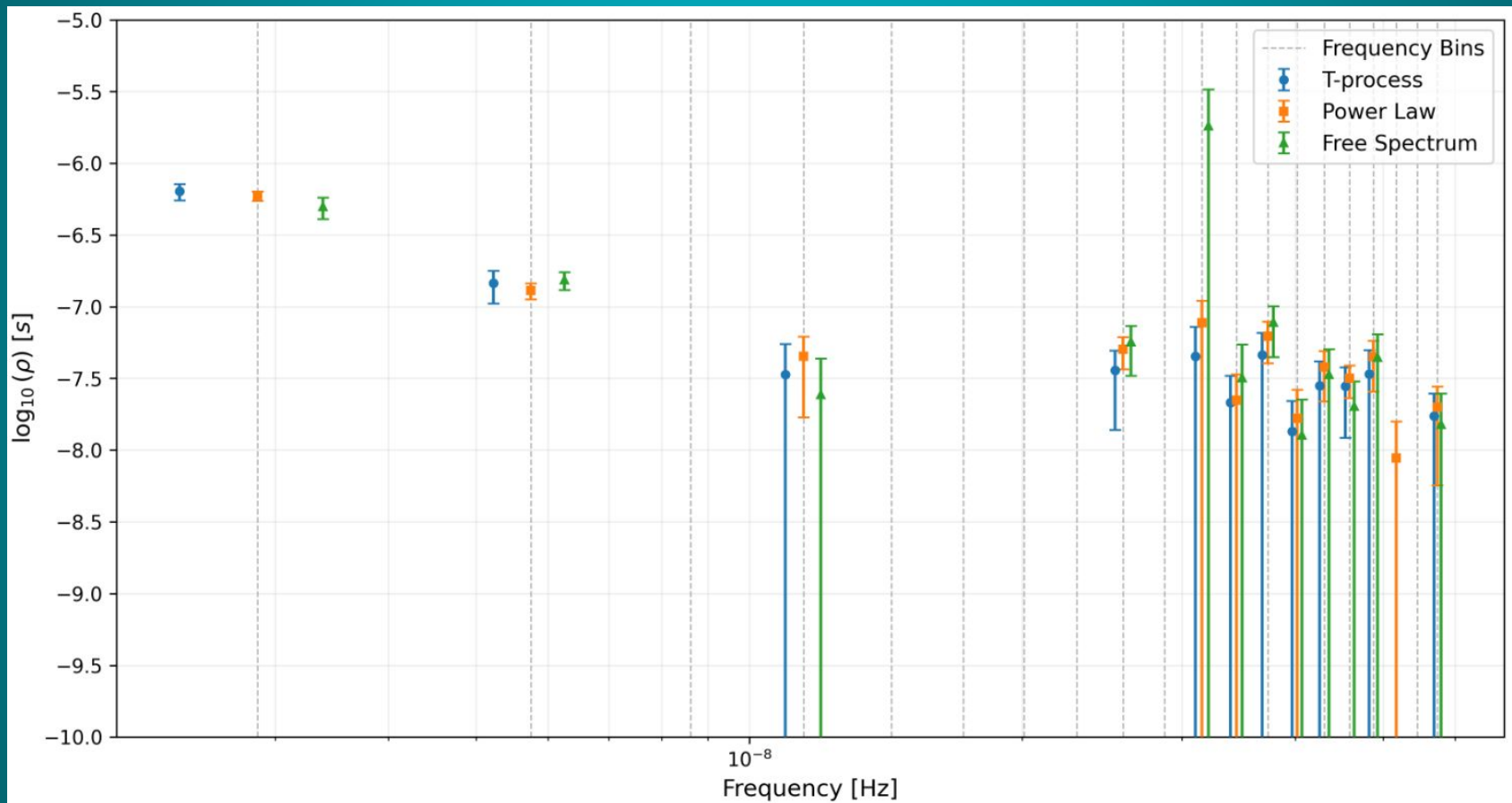
Freespectrum -> worse performance in the PFOS

T-process -> higher freq bin are more constrained

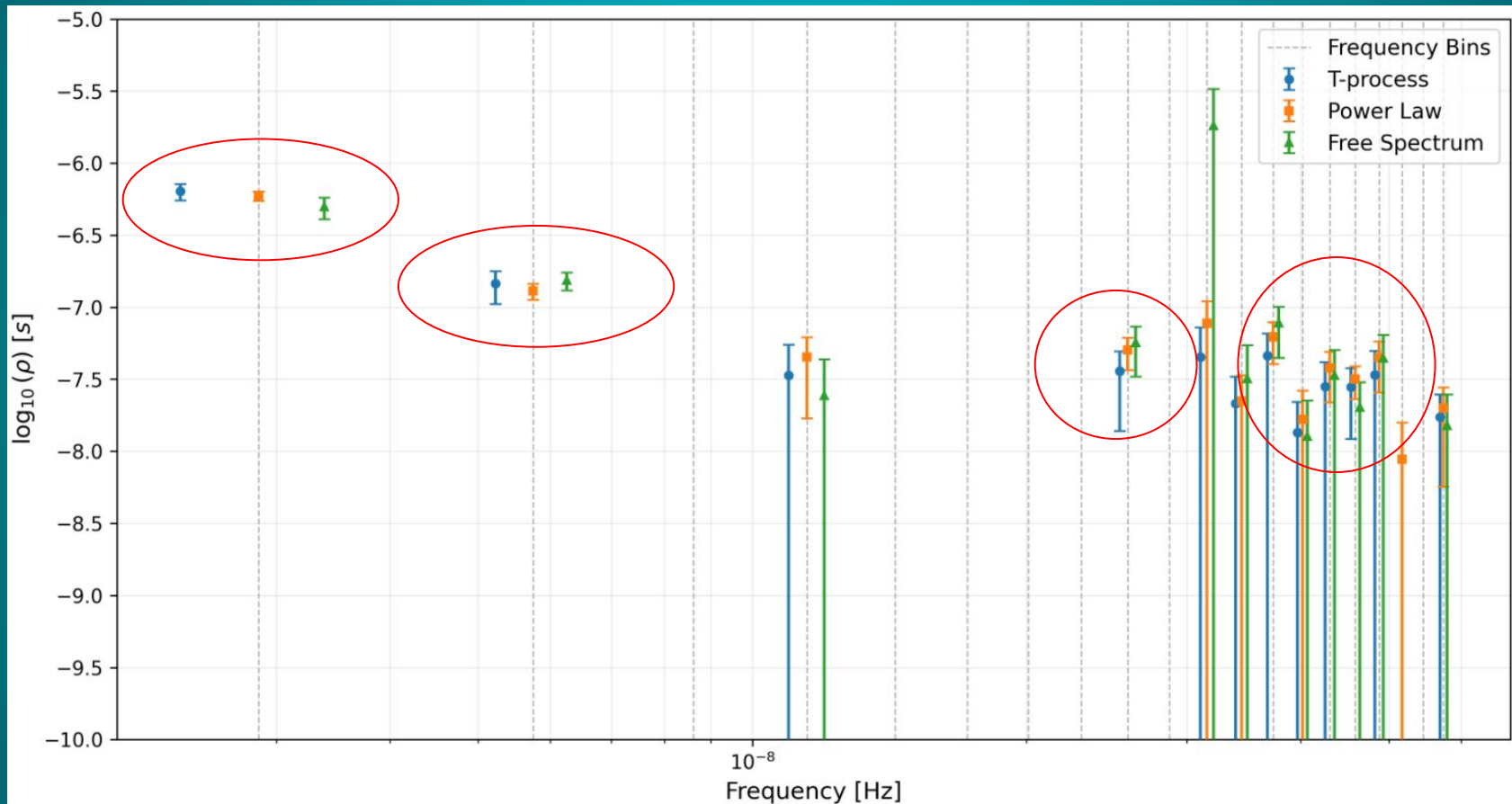
Test on EPTA DR2new+



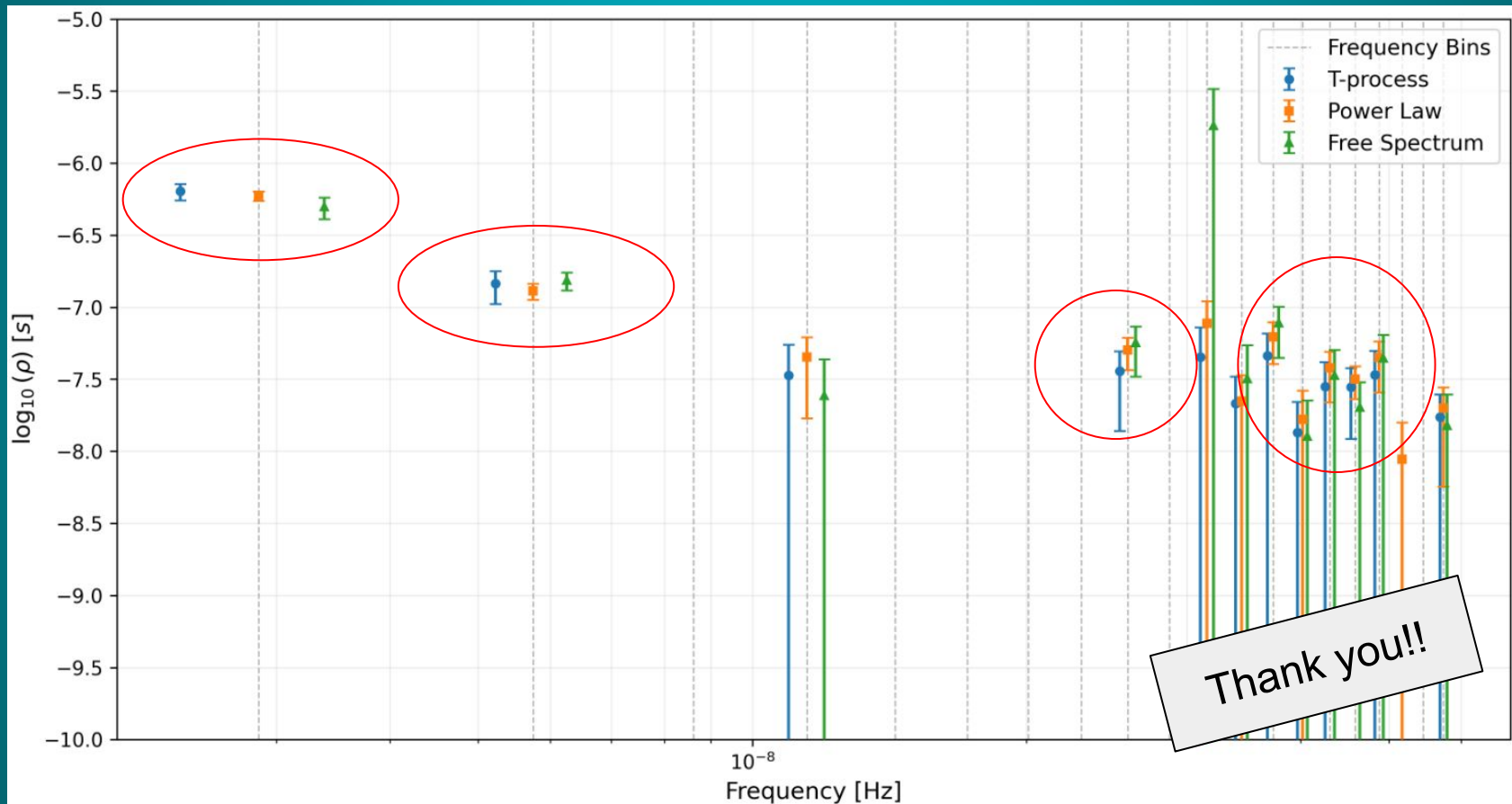
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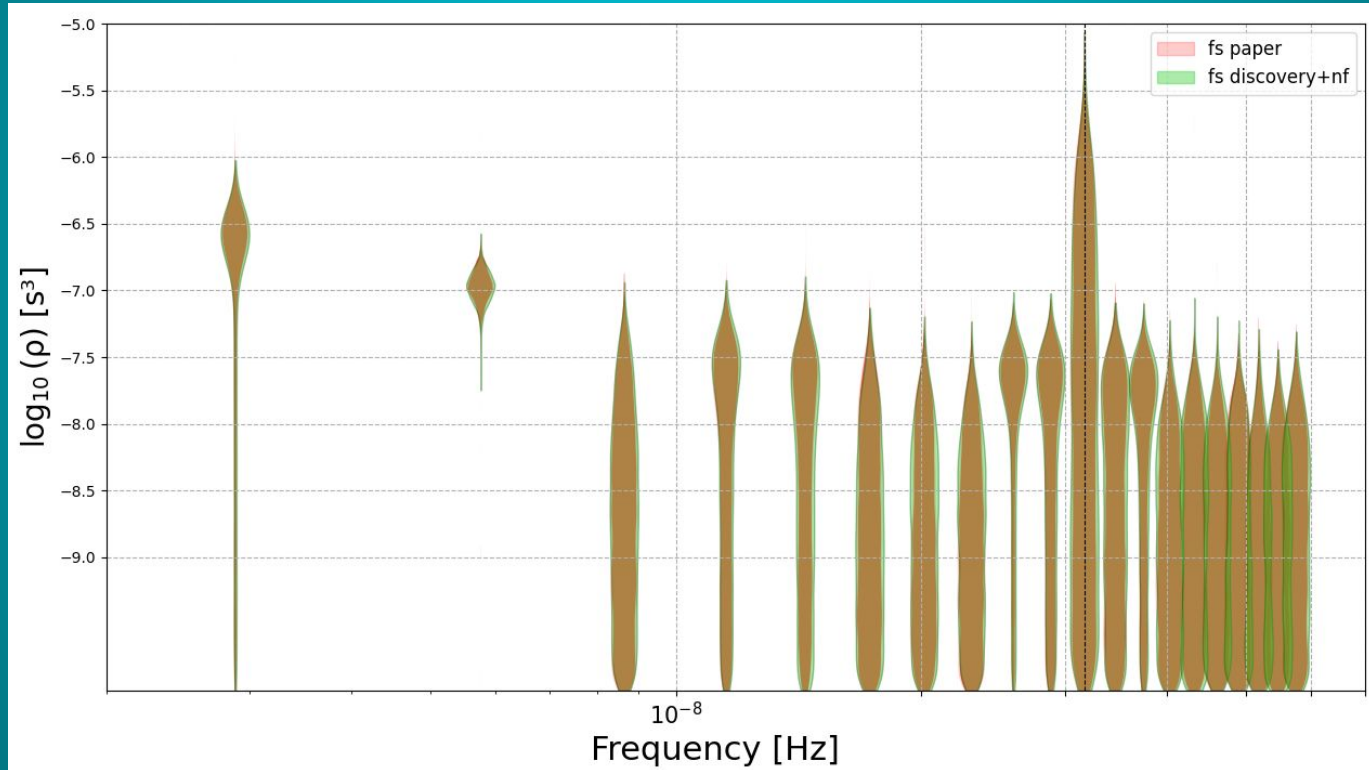


PFOS on EPTA DR2new+



bk

Comparison enterprise vs discovery with EPTA DR2 new+



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