

COMMON ENVELOPE EVOLUTION

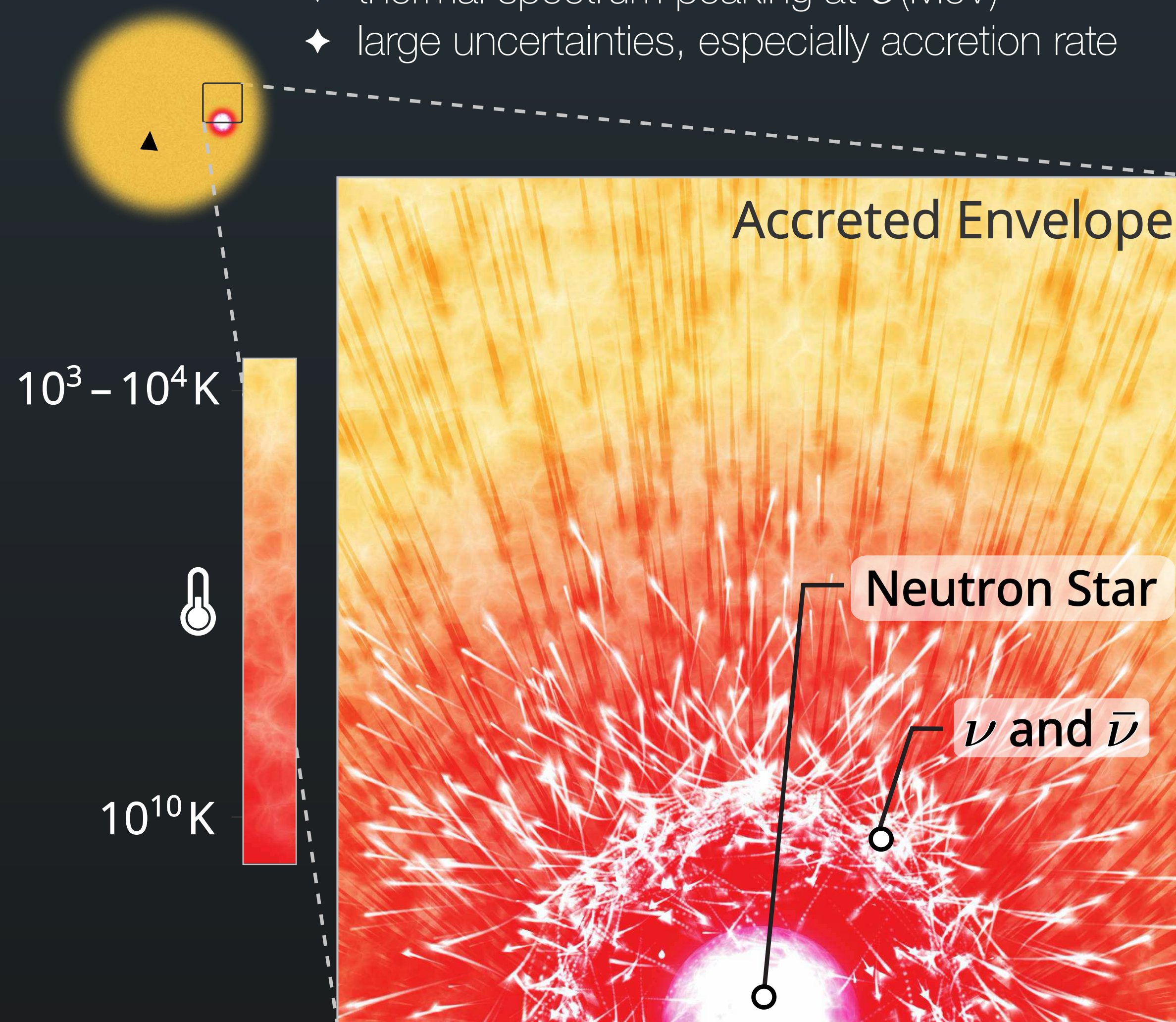
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Abstract

Common-envelope evolution, where a **neutron star or black hole is engulfed by a companion**, is a critical step in the **formation pathways for gravitational-wave sources and X-ray binaries**. But it has never been directly observed.. We show that the hypothesized super-Eddington accretion during a common-envelope event would produce **months-long neutrino signals in the MeV range that are within reach of present and planned detectors**. While there are substantial uncertainties on the rate of such events (0.01–1/century in the Milky Way) and the neutrino luminosity, a search for such a signal in archival or upcoming data could lead to significant new insights into the astrophysics of common-envelope evolution, thereby answering long-standing open questions in astrophysics.

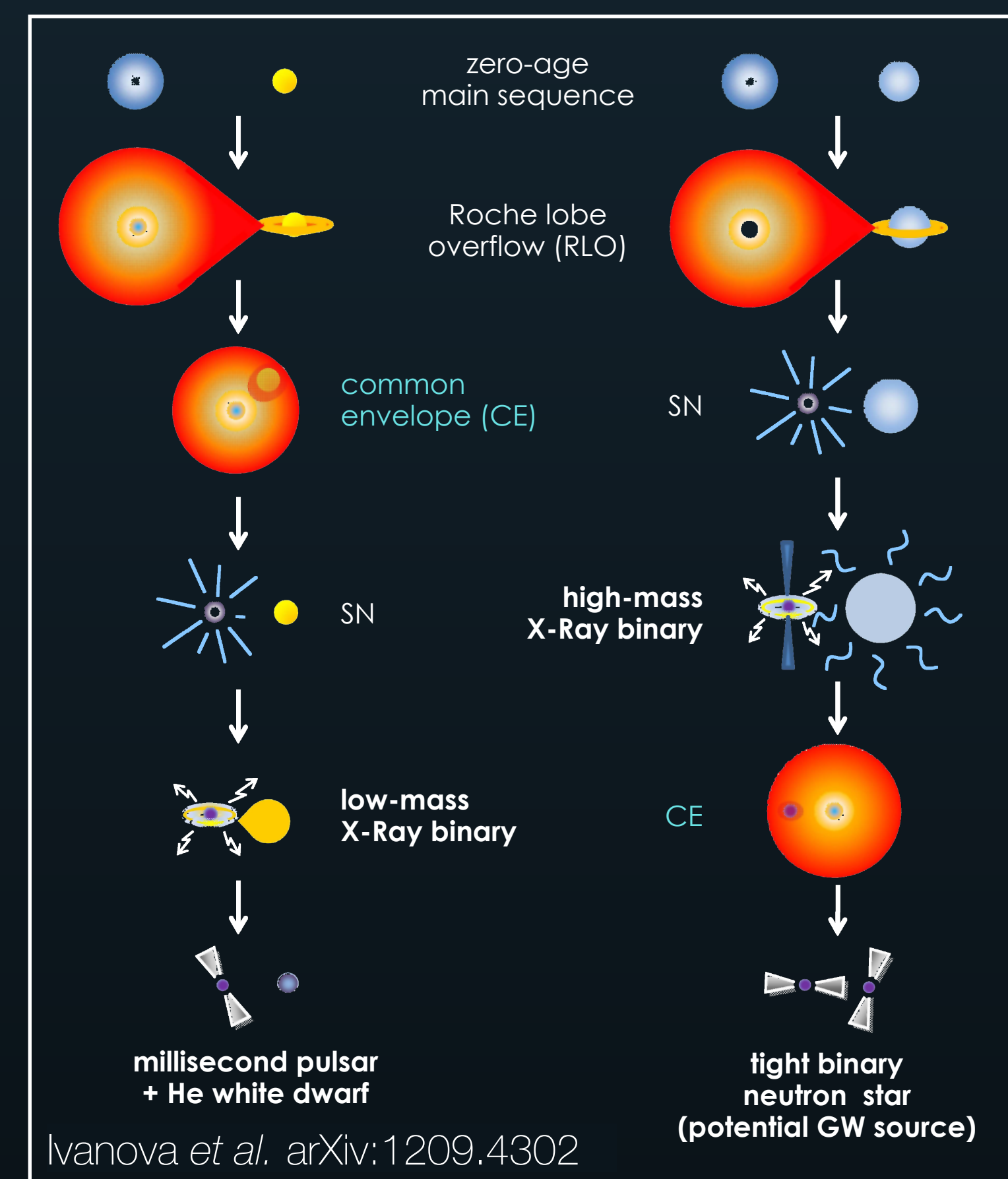
Neutrino Emission

- ◆ huge accretion rates during CEE (up to $0.1 M_{\odot}/\text{yr}$)
- ◆ **gravitational energy release** leads to heating
- ◆ photons trapped \Rightarrow not an efficient cooling channel
- ◆ **neutrino cooling!** ($e^+e^- \rightarrow \nu\bar{\nu}$, $\gamma^* \rightarrow \nu\bar{\nu}$)
 - ◆ in addition: kinetic energy release (explosion?)
 - ◆ thermal spectrum peaking at $\mathcal{O}(\text{MeV})$
 - ◆ large uncertainties, especially accretion rate

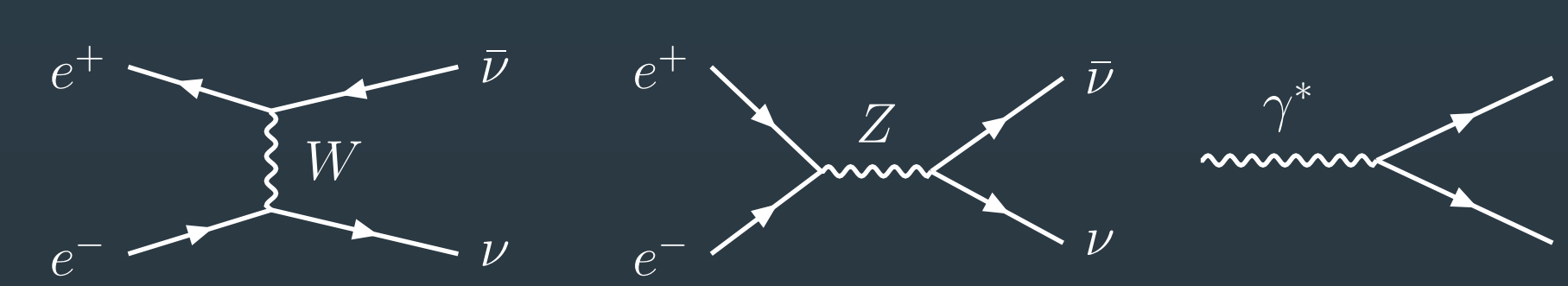


Astrophysical Context

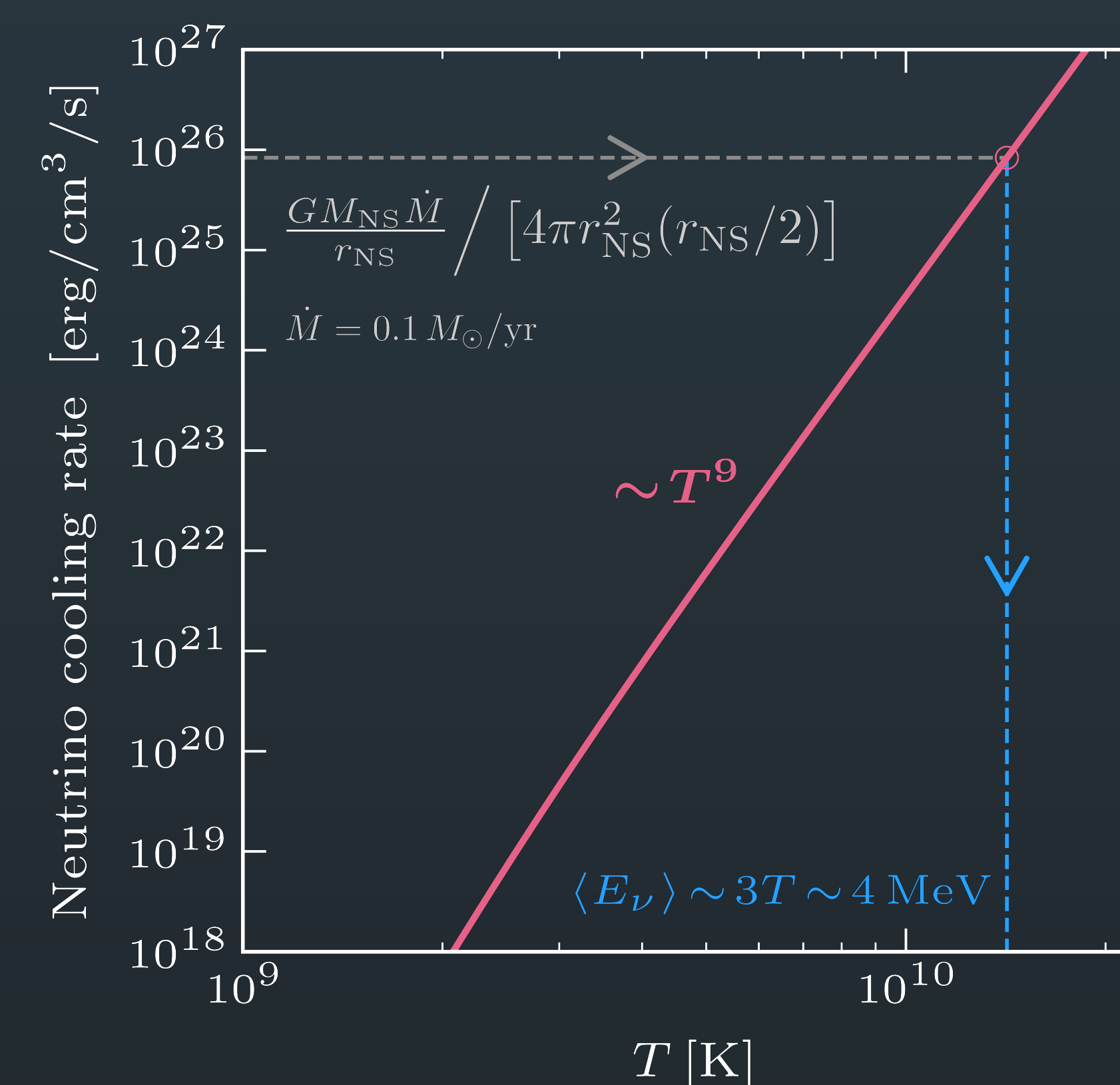
- ◆ critical step in the formation of **X-ray binaries** and **GW sources**
- ◆ theoretically poorly understood
- ◆ **never directly observed**
- ◆ duration $\mathcal{O}(\text{months-years})$
- ◆ very rare: (0.01–1 / century)



Neutrino Spectrum

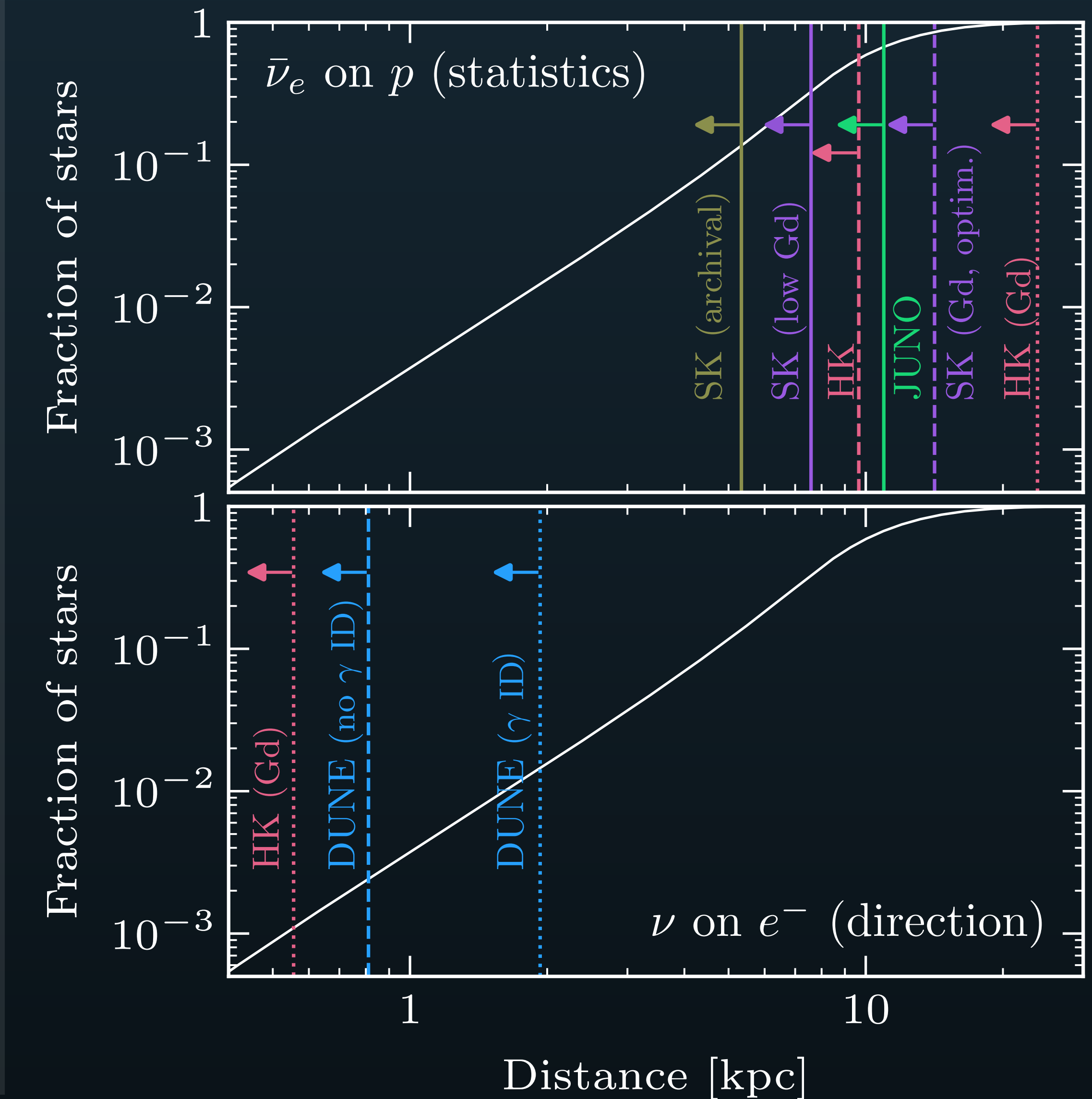


- ◆ thermal spectrum
- ◆ accretion rate determines temperature
- ◆ neutrino flux \propto accretion rate



Distance Reach

3σ sensitivity (Normal Ordering)

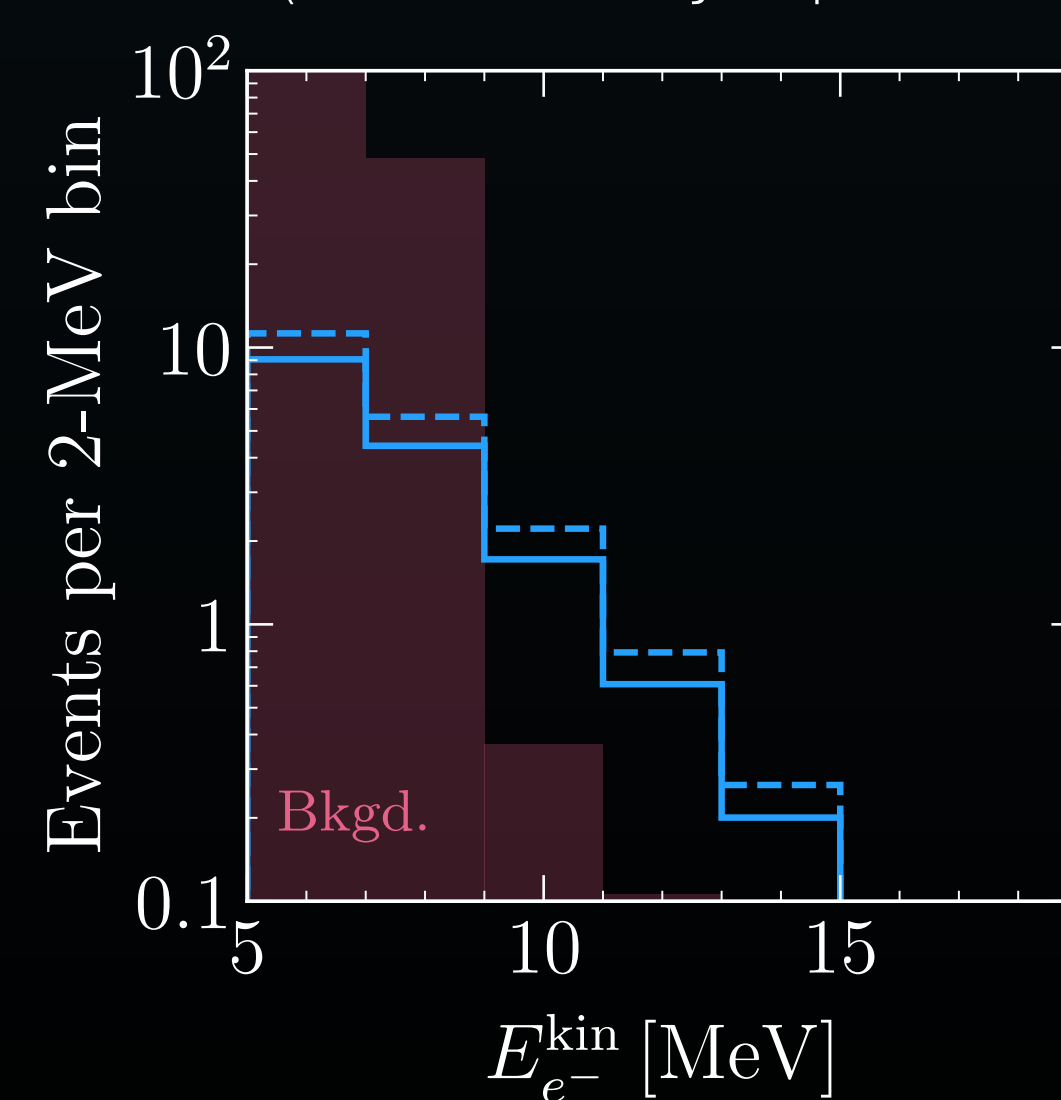


Detection

DUNE
 ν - e scattering (*directional*)

Backgrounds

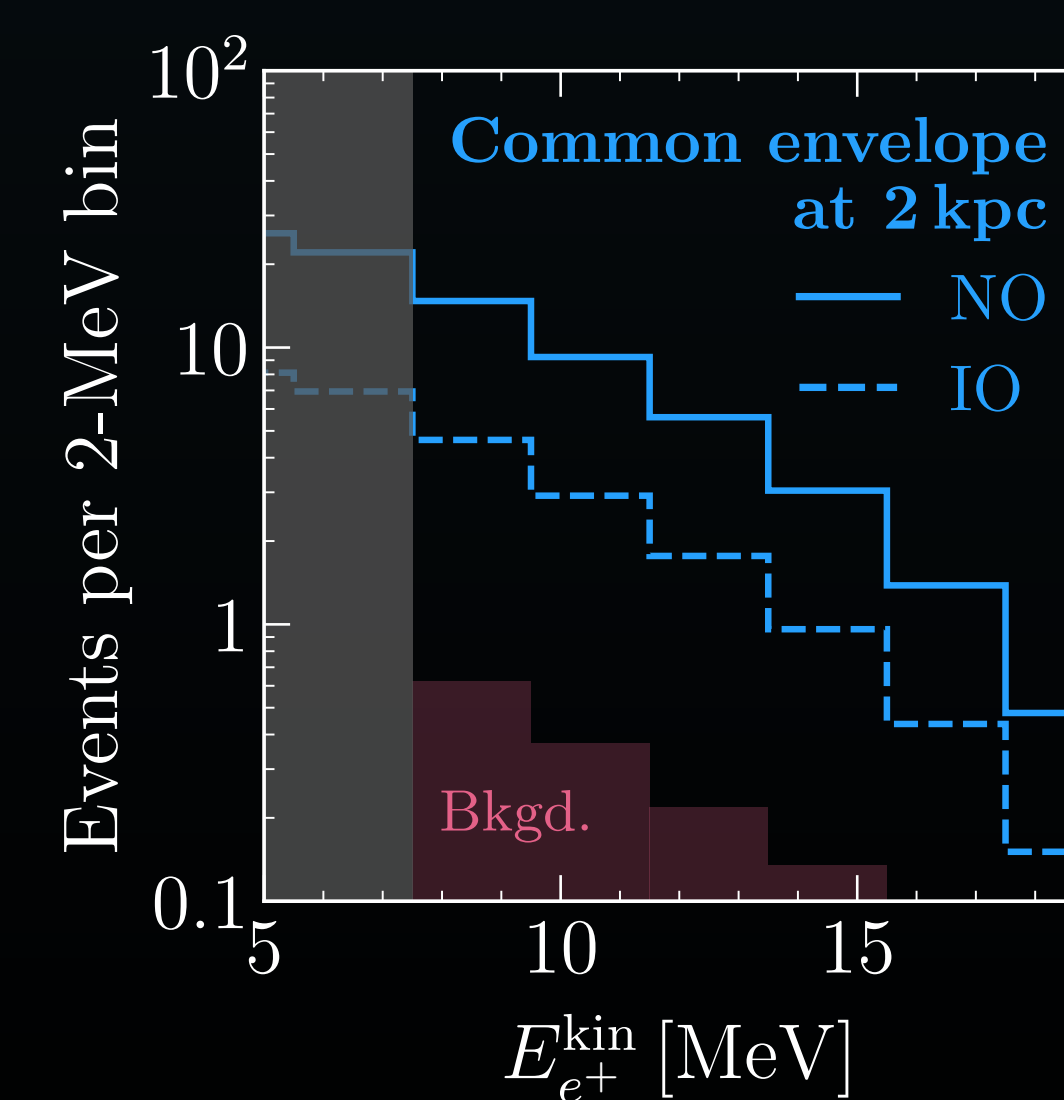
- ◆ neutrons
- ◆ spallation products
- ◆ solar ν must be suppressed (directionality / γ from $^{40}\text{Ar}^*$)



SuperK (archival w/o Gd)
inv. β decay (*non-directional*)

Backgrounds

- ◆ accidental
- ◆ ^9Li from spallation
- ◆ atmospheric ν via NC

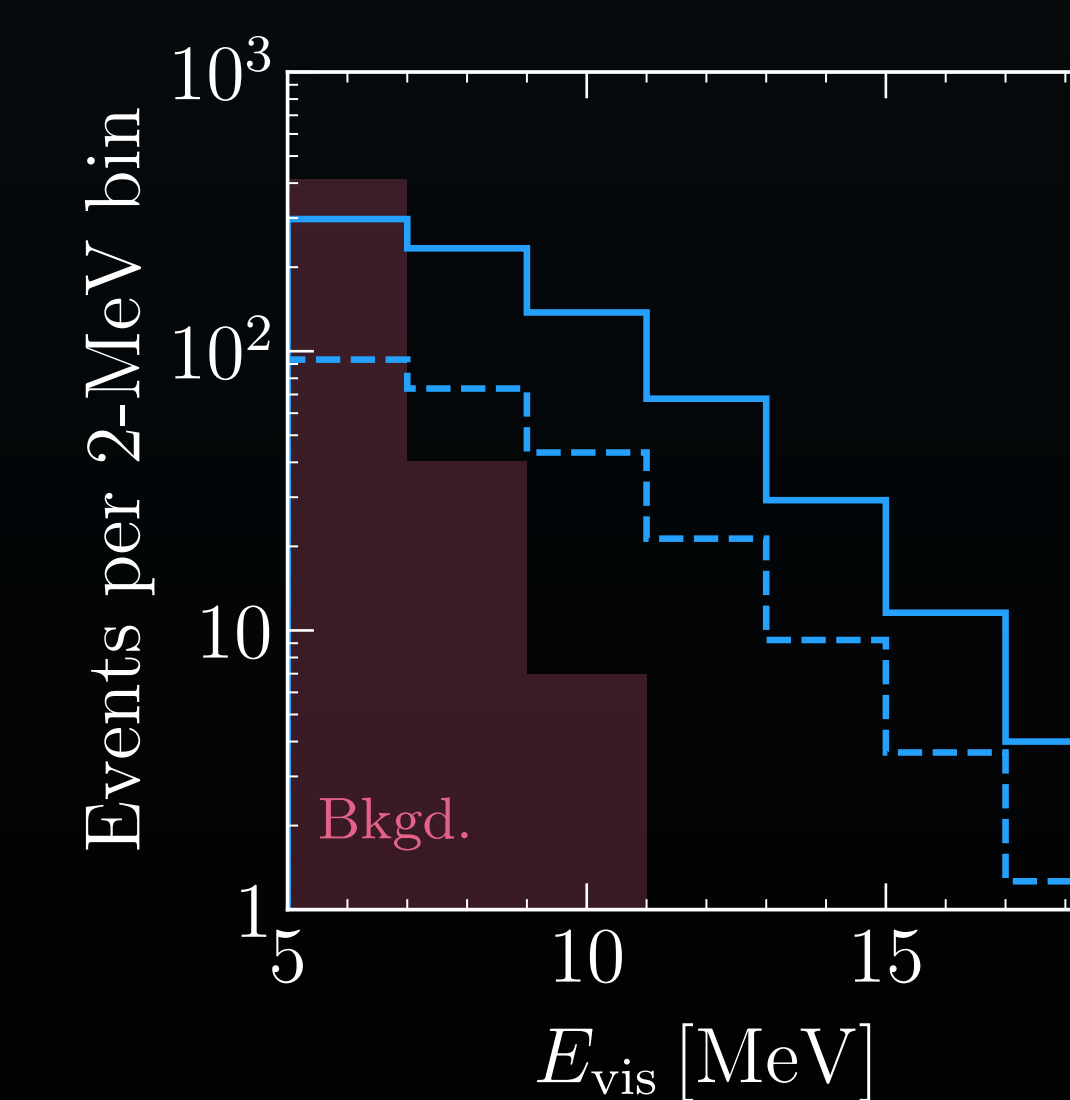


JUNO

inv. β decay (*non-directional*)

Backgrounds

- ◆ reactor ν
- ◆ ^9Li and ^8He from spallation
- ◆ atmospheric ν via NC



Take-Home Message

- ◆ **new astrophysical neutrino source**
- ◆ **detectable anywhere in our galaxy**
- ◆ very **rare**, but payoff of detection would be huge
- ◆ many astrophysical uncertainties, which neutrino observations could help resolve
- ◆ **requires dedicated search strategy** (excess in ν event rate lasting \sim months)
- ◆ signal could be hidden in existing data

Ivan Esteban, John Beacom, JK, *Detectable MeV Neutrino Signals from Neutron-Star Common-Envelope Systems*, Phys. Rev. Lett. 134 (2025) 18, 181003, [arXiv:2310.19868](https://arxiv.org/abs/2310.19868)

code for computing hydrodynamic profiles and neutrino spectra available on **GitHub**: <https://github.com/ivan-esteban-phys/common-envelope-thermal>

