

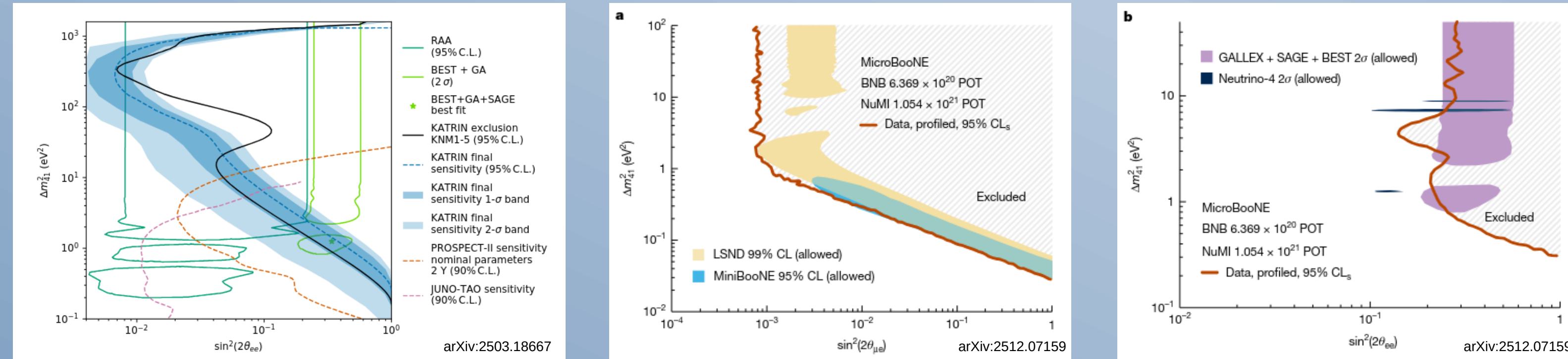
# Sterile Neutrino search with KM3NeT/ORCA6-18 dataset

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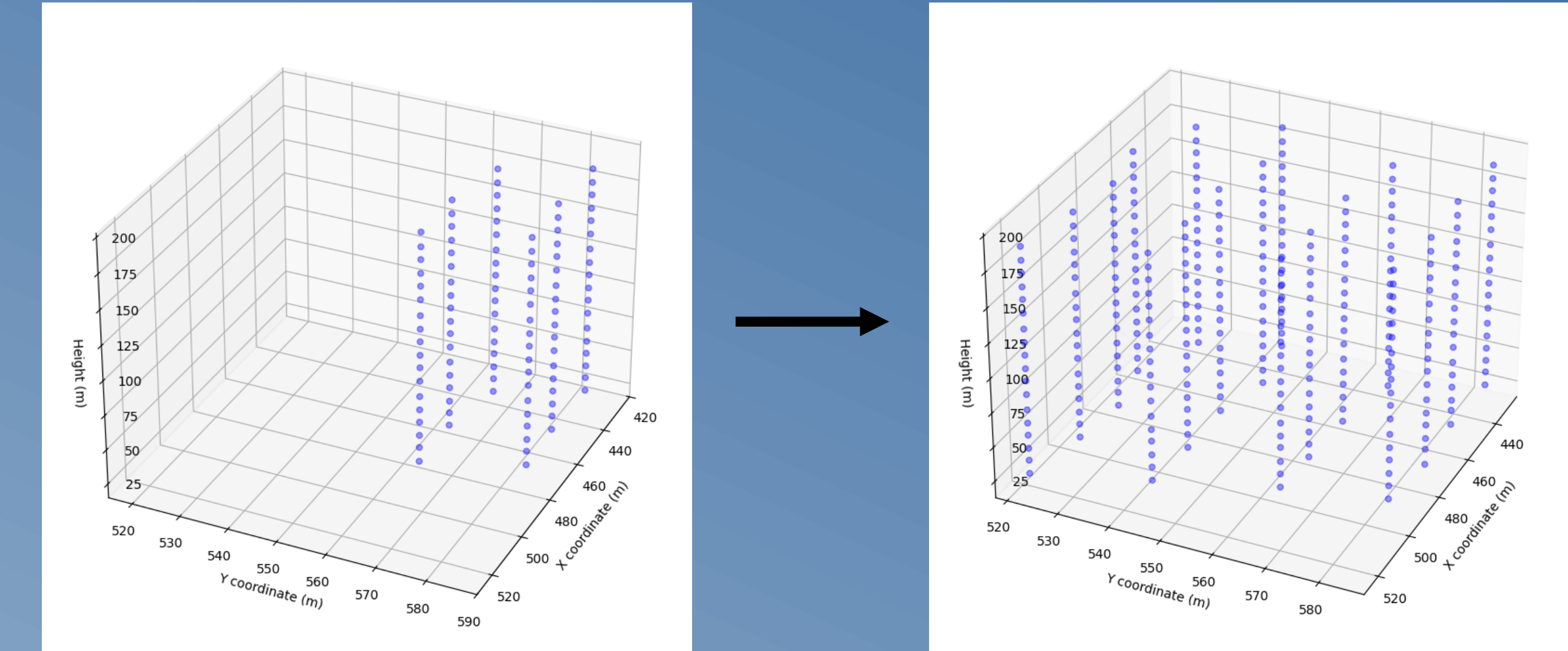
## The History

- Historic results from LSND, MiniBooNE, reactor anomalies indicate sterile neutrino signal with  $\Delta m_{41}^2 \sim 1 \text{ eV}^2$
- Recent results from MicroBooNE, KATRIN show no indication of sterile
- **Atmospheric neutrinos allow check on different mixing matrix elements  $U_{\alpha 4}$  for the same sterile mass splitting**



## The Detector

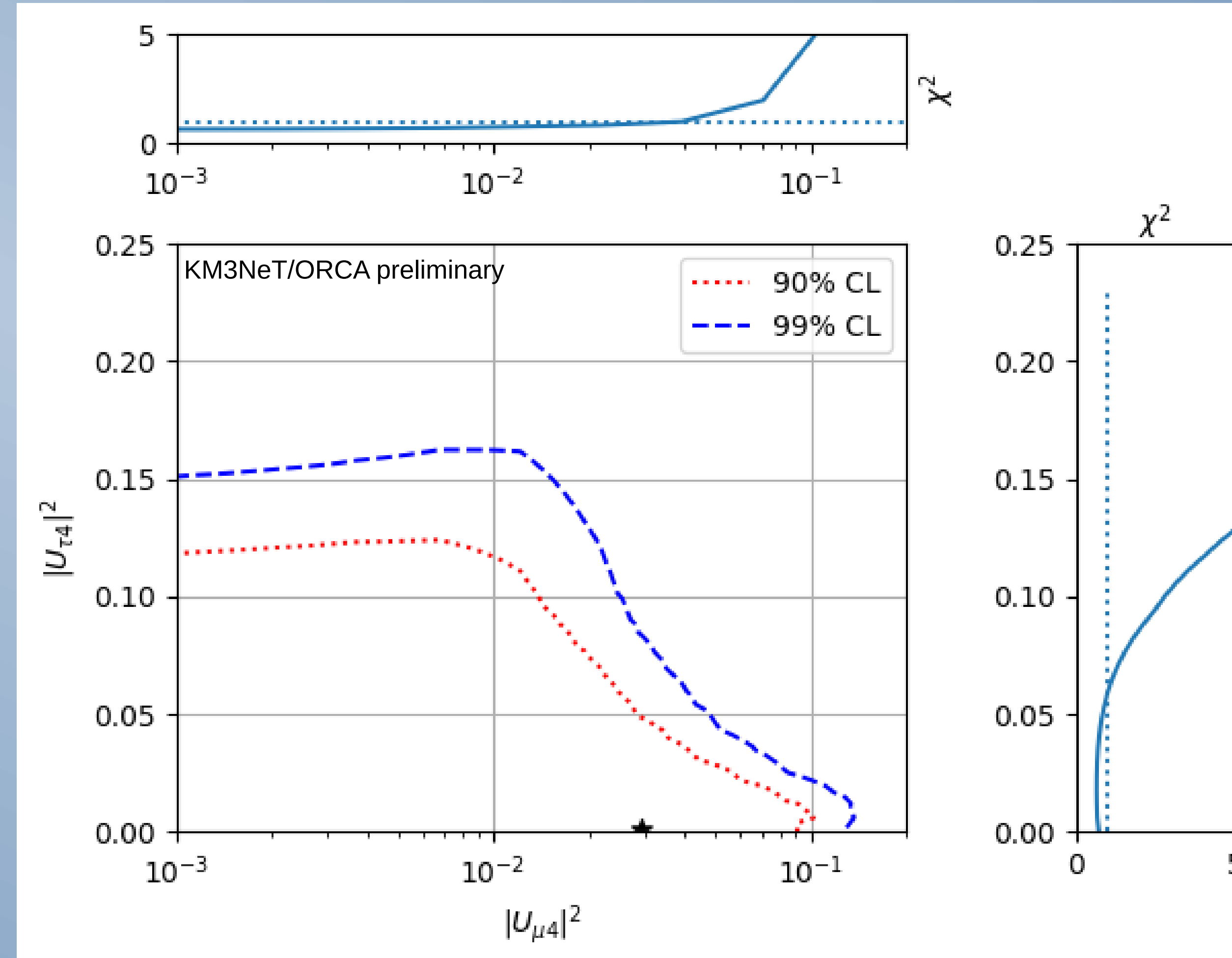
- KM3NeT is a water cherenkov neutrino detector under construction in the Mediterranean Sea (Toulon, FR and Sicily, IT)
- KM3NeT/ORCA is optimised for energies **2 GeV – 100 GeV** for studies on **atmospheric neutrinos**
- New dataset covering **~1.4Mton-yr of data**, 3.3x increase over ORCA6 analysis
- 18 strings / 108 total – exposure increasing non-linearly!



## The Analysis Method

Perform multi-dimensional binned log-likelihood fit in

- **Energy** ( $> 2 \text{ GeV}$ )
- **Zenith angle** (Upgoing only)
- **Detector class** (ORCA67 + ORCA11-18)
- **Reconstructed class** (Track + Intermediate + Shower)
- Model systematics for **Standard Oscillations** ( $\Delta m_{31}^2, \theta_{23}$ ), **Flux Model, Cross Section, Detector Response**



## The Results

Scan over  $\theta_{24} / \theta_{34}$  space to profile the 90% / 99% contour in  $(|U_{\mu 4}|^2 / |U_{\tau 4}|^2)$

$$|U_{\mu 4}|^2 = \cos^2 \theta_{24}$$

$$|U_{\tau 4}|^2 = \sin^2 \theta_{24} \cos^2 \theta_{34}$$

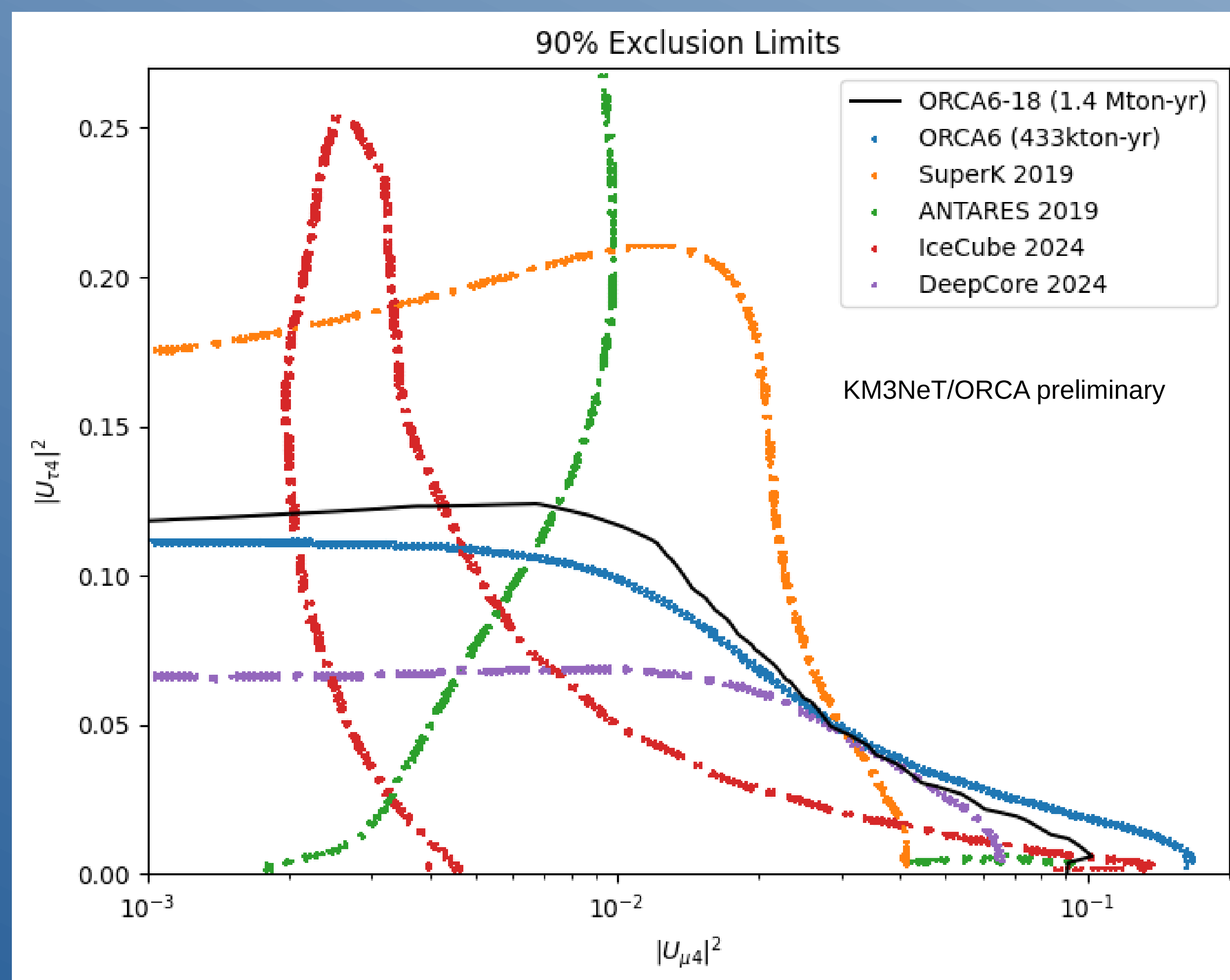
Fit consistent with no sterile signal and prior 433 kton-yr result

**BEST FIT ( $< 90\% \text{ 1D confidence limit}$ )**

$$|U_{\mu 4}|^2 = 0.029, < 0.094$$

$$|U_{\tau 4}|^2 = 8.8 \times 10^{-4}, < 0.11$$

All systematics compatible with nominal (NuFit6.1 for osc params) at  $< 2\sigma$



KM3NeT/ORCA preliminary Event Type	ORCA6/7 tracks	ORCA6/7 interm.	ORCA6/7 showers	ORCA11-18 tracks	ORCA11-18 interm.	ORCA11-18 showers
$\nu_e \text{CC} + \bar{\nu}_e \text{CC}$	50	336	682	87	646	911
$\nu_\mu \text{CC} + \bar{\nu}_\mu \text{CC}$	3869	748	493	5641	1151	557
$\nu_\tau \text{CC} + \bar{\nu}_\tau \text{CC}$	36	80	176	53	141	212
$\nu \text{NC} + \bar{\nu} \text{NC}$	19	112	276	31	219	430
atmospheric $\mu + \bar{\mu}$	29	22	30	31	20	21
Total MC	4003	1298	1657	5843	2176	2131
Total Data	4098	1361	1586	6329	2236	2016

## The Global Picture

- Consistent and comparable with results from other atmospheric neutrino experiments (Super-K, IceCube, DeepCore, ANTARES)
- Changes from prior ORCA6 analysis come from **increase in exposure and improved systematic understanding**

Parameter	Best fit value	Upper Limit	Lower Limit
$\theta_{24}$	0.172 <sup>+0.018</sup> <sub>-0.018</sub>	0.251	0.0
$\theta_{34}$	0.03017 <sup>+0.0001</sup> <sub>-0.0001</sub>	0.071	0.0
Flux			
$\Phi_{\text{GSF1}}$	0.0984 <sup>+0.0009</sup> <sub>-0.0009</sub>	0.123	0.0795
$\Phi_{\text{K-2P}}$	1.39 <sup>+0.14</sup> <sub>-0.14</sub>	1.58	1.20
$\Phi_{\text{K-31G}}$	-0.0420 <sup>+0.008</sup> <sub>-0.008</sub>	-0.04	-0.04
$\Phi_{\text{K-158G}}$	0.0384 <sup>+0.008</sup> <sub>-0.008</sub>	0.04	0.03
$\Phi_{\text{K-20T}}$	1.382 <sup>+0.13</sup> <sub>-0.13</sub>	1.58	1.20
$\Phi_{\text{K-31G}}$	-0.185 <sup>+0.08</sup> <sub>-0.08</sub>	-0.18	-0.18
$\Phi_{\text{K-158G}}$	0.187 <sup>+0.017</sup> <sub>-0.017</sub>	0.19	0.18
$\Phi_{\text{K-2P}}$	-0.235 <sup>+0.016</sup> <sub>-0.016</sub>	-0.23	-0.23
$\Phi_{\text{K-20T}}$	-0.764 <sup>+0.12</sup> <sub>-0.12</sub>	-0.76	-0.76
$\Phi_{\text{K-31G}}$	-0.0447 <sup>+0.008</sup> <sub>-0.008</sub>	-0.04	-0.04
$\Phi_{\text{K-158G}}$	-0.173 <sup>+0.011</sup> <sub>-0.011</sub>	-0.17	-0.17
$\Phi_{\text{K-2P}}$	-0.160e-01 <sup>+0.011</sup> <sub>-0.011</sub>	-0.16	-0.16
$\Phi_{\text{K-20T}}$	-0.794 <sup>+0.12</sup> <sub>-0.12</sub>	-0.79	-0.79
$\Phi_{\text{K-31G}}$	0.102 <sup>+0.009</sup> <sub>-0.009</sub>	0.10	0.10
$\Phi_{\text{K-158G}}$	-0.203 <sup>+0.011</sup> <sub>-0.011</sub>	-0.20	-0.20
$\Phi_{\text{K-2P}}$	0.0819 <sup>+0.008</sup> <sub>-0.008</sub>	0.08	0.08
$\Phi_{\text{K-158G}}$	0.00564 <sup>+0.0005</sup> <sub>-0.0005</sub>	0.0056	0.0056
$\Phi_{\text{K-2P}}$	0.171 <sup>+0.018</sup> <sub>-0.018</sub>	0.18	0.17
$\Phi_{\text{K-158G}}$	0.0400 <sup>+0.0004</sup> <sub>-0.0004</sub>	0.04	0.04
Detector Systematics			
$E_{\text{scale}}$	-0.0480 <sup>+0.0012</sup> <sub>-0.0012</sub>	-0.04	-0.04
Trigger Efficiency	-1.378 <sup>+0.051</sup> <sub>-0.051</sub>	-1.37	-1.37
Cross Section			
$S_{\text{NC}}$	1.21 <sup>+0.39</sup> <sub>-0.39</sub>	1.21	1.21
$S_{\text{T}}$	0.987 <sup>+0.14</sup> <sub>-0.14</sub>	0.98	0.98
$\sigma_{\text{QE}}$	0.545 <sup>+0.48</sup> <sub>-0.48</sub>	0.54	0.54
$\sigma_{\text{RES}}$	0.387 <sup>+0.27</sup> <sub>-0.27</sub>	0.38	0.38
Oscillation Parameters			
$\delta_{24}$	0.812 <sup>+0.14</sup> <sub>-0.14</sub>	0.81	0.81
$\theta_{23}$	0.708 <sup>+0.062</sup> <sub>-0.062</sub>	0.70	0.70
$\Delta m_{31}^2$	2.192 <sup>+0.29</sup> <sub>-0.29</sub> $\times 10^{-03}$	2.19	2.19

## Systematic Effects

- Most important systematics on POIs are
- Standard oscillation parameters ( $\Delta m_{31}^2, \theta_{23}$ )
- Pion 20T uncertainty in flux model
- Kaon + 2P production uncertainty in flux model