



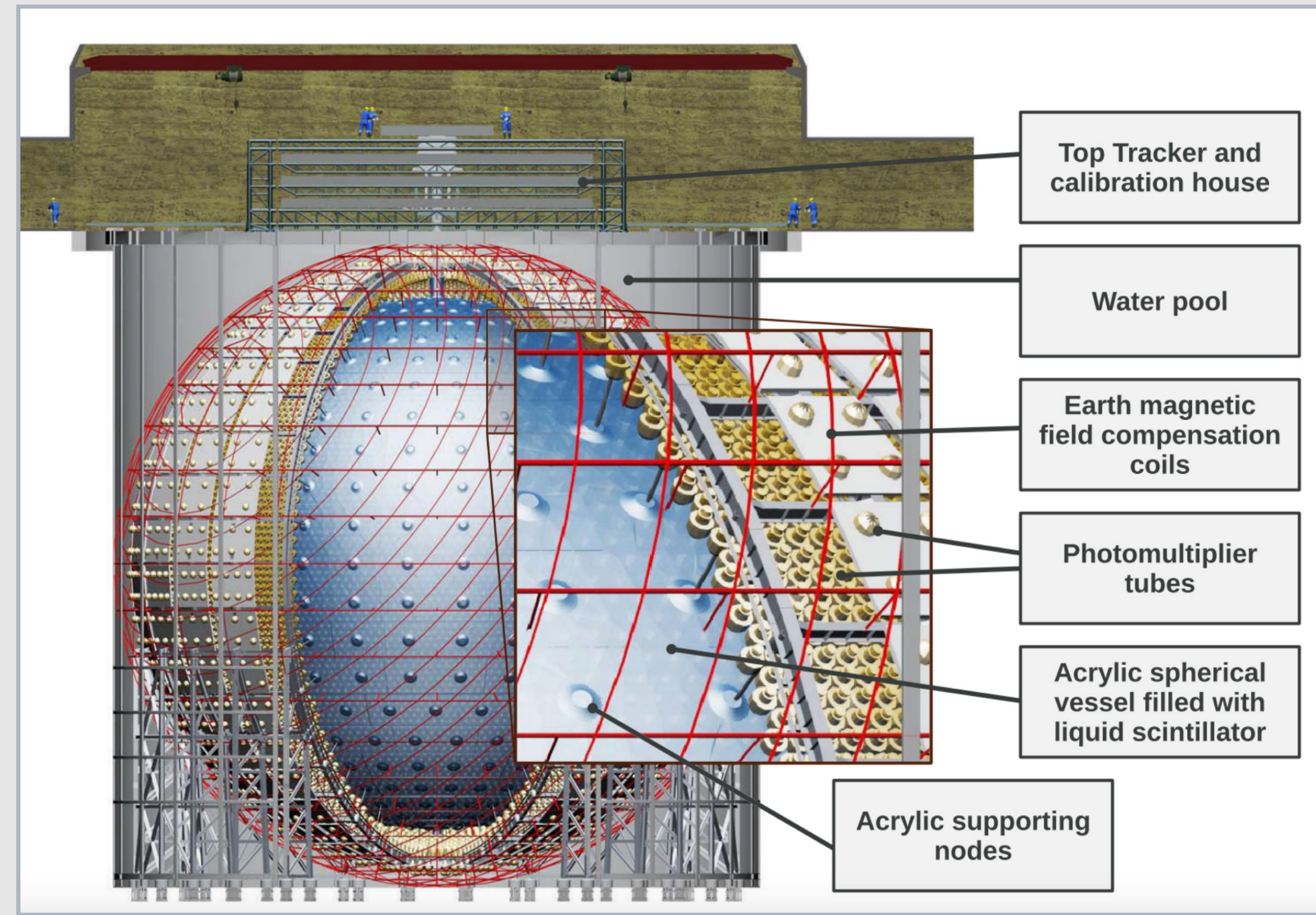
# Energy Response Model of the JUNO Central Detector

Yilin Liao, Shanghai Jiao Tong University ♦ xll123@sjtu.edu.cn  
Roberto Mandujano, University of California, Irvine ♦ rmanduj@uci.edu  
(on behalf of the JUNO collaboration)



## 1. Jiangmen Underground Neutrino Observatory [1]

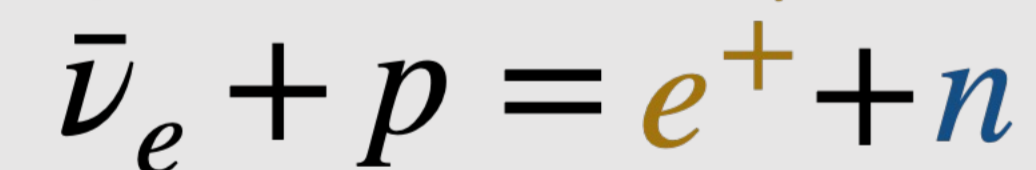
- Located in Southern China next to multiple reactors providing large  $\bar{\nu}_e$  flux
- 20 kton of liquid scintillator instrumented with 2 sets of PMTs reaching ~78% photocoverage
- Primarily designed for neutrino mass ordering determination through precise measurement of the  $\bar{\nu}_e$  energy spectrum



## 2. Antineutrino Detection

Reactor antineutrinos detected through Inverse Beta Decay (IBD)

**Prompt:**  $e^+e^-$  annihilation  $\gamma$



**Delayed:** n capture on H

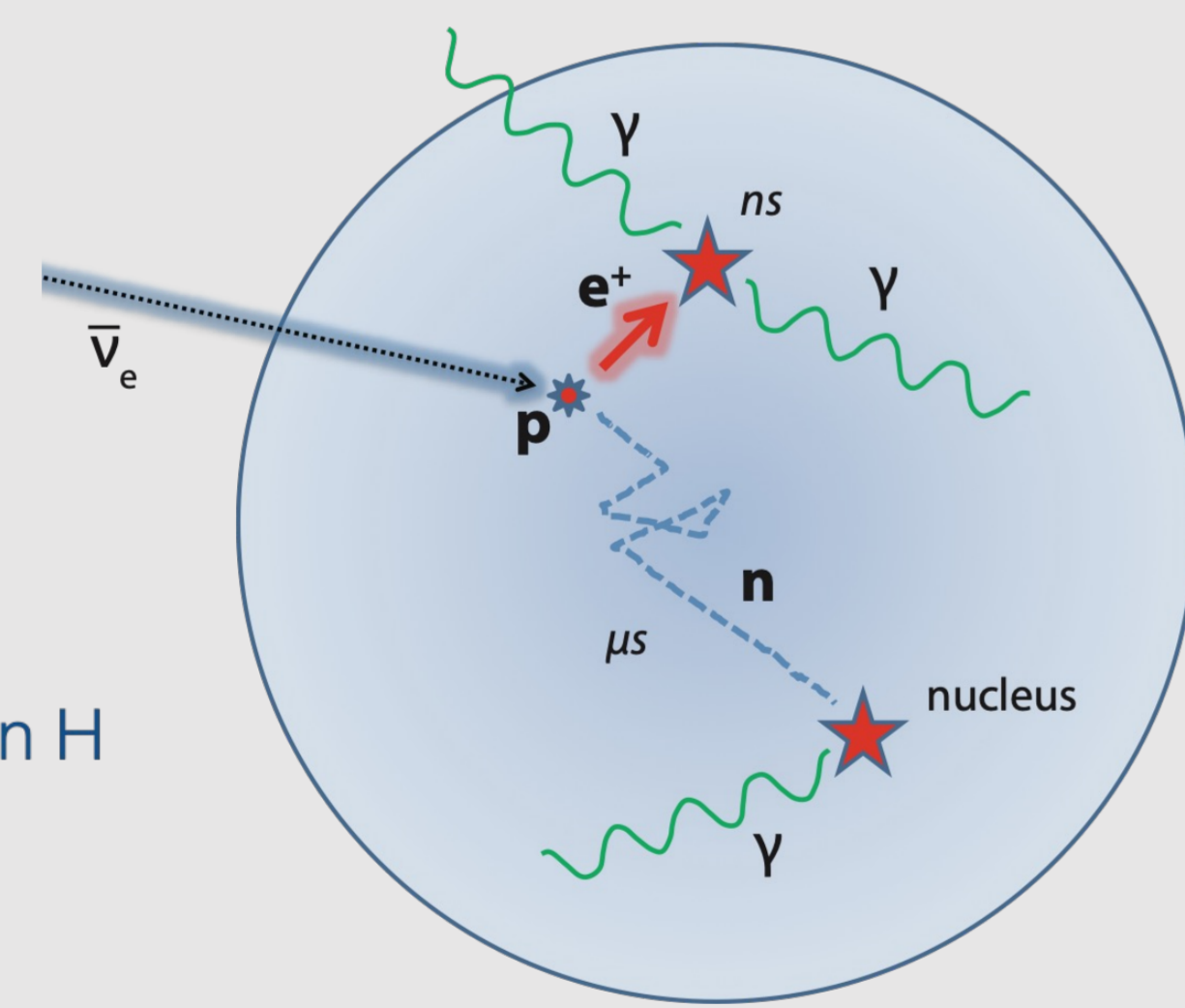


Figure of IBD process

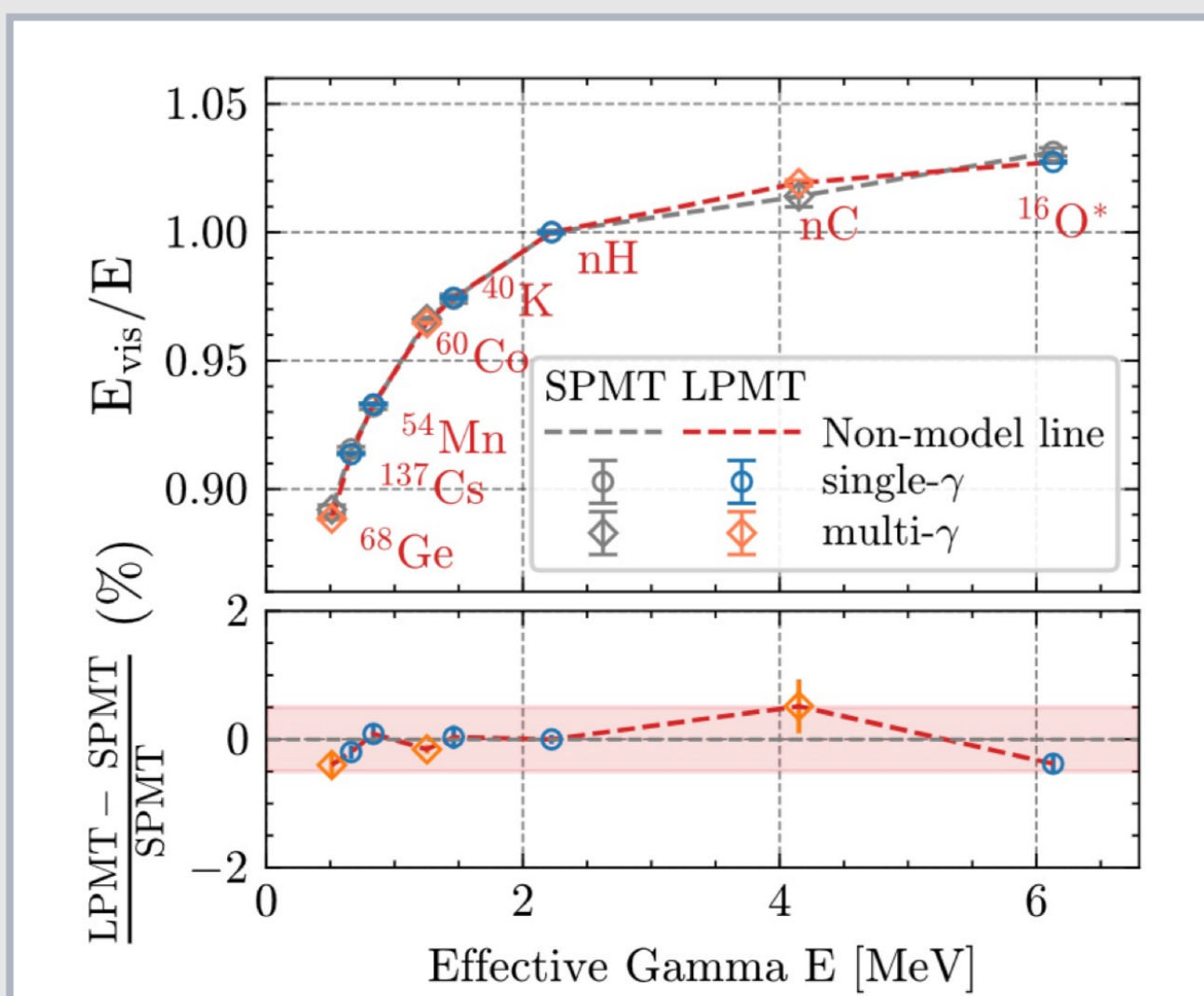
Spatial-temporal coincidence of prompt and delayed events provide IBD signal of high purity

## 3. Energy Non-linearity [1]

Since  $E_{e^+} \propto E_\nu$ , mapping the non-linear relationship between  $E_{vis}$  &  $E_{dep}$  is key to JUNO's oscillation program

Energy non-linearity depends on:

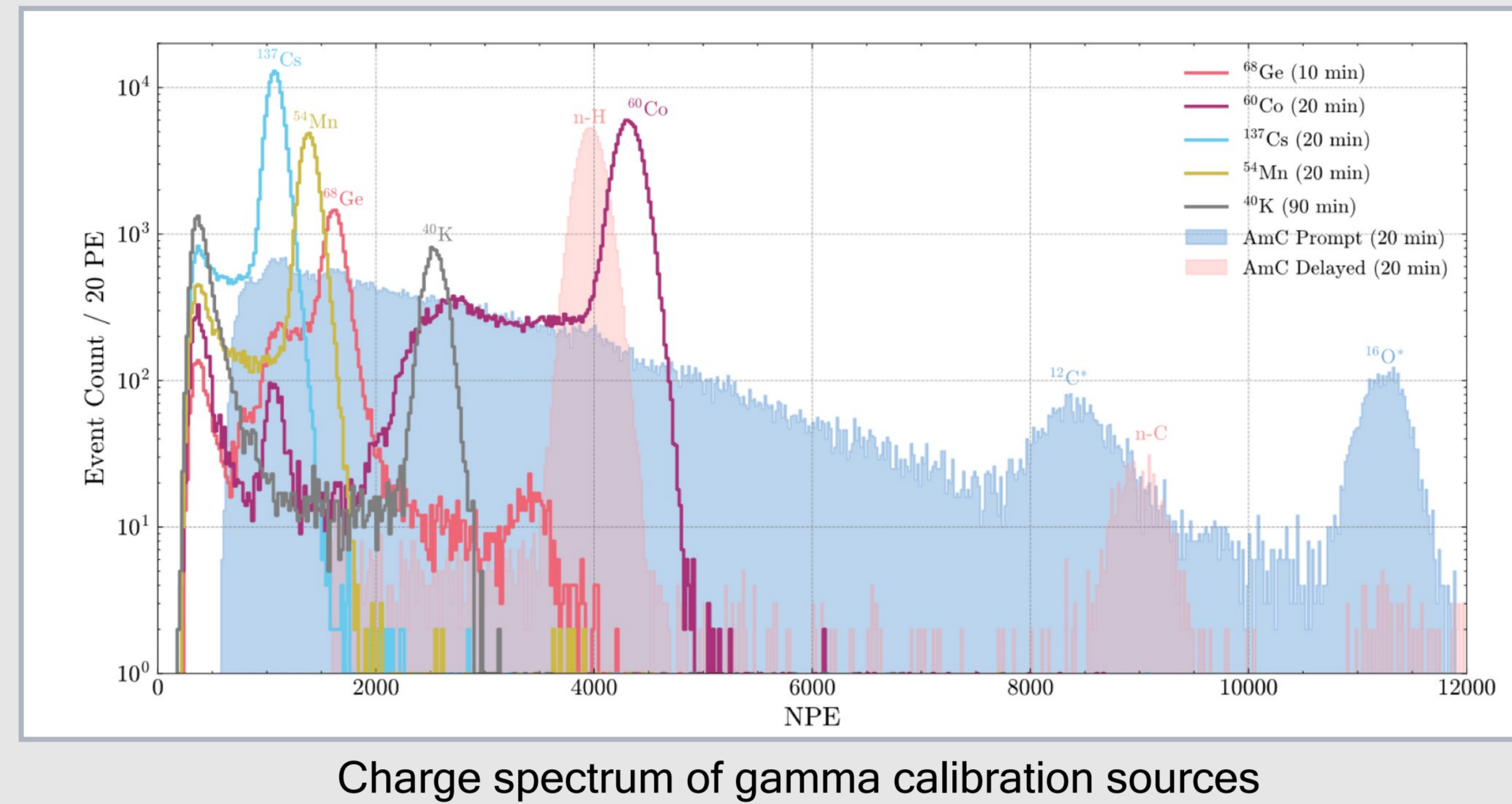
- Quenching (low energies)
- Cherenkov light (high energies)
- Instrumental Response



## 4. Calibration Sources [1,5]

- Deployed radioactive gamma sources cover energies from 1-8 MeV
- Cosmogenic isotopes from muon spallation directly constrain  $e^\pm$
- 16.5 m radius fiducial volume chosen

Cosmogenic Isotope	Emission
$^{10}\text{C}$	$e^+$
$^{11}\text{C}$	$e^+$
$^{12}\text{B}$	$e^-$
$^{12}\text{N}$	$e^+$



## 5. Modeling [2,3,4,5]

We present 2 models used to describe the JUNO detector response:  
a. COMEER: Uses  $dE, dx, \beta$  step information from Geant4 to calculate the number of photoelectrons (PE) in an event for different particles:

$$NPE = S * \sum_{step} \left( \frac{dE}{1 + k_B * \frac{dE}{dx}} + p * \frac{1}{1 - n^2 * \beta^2} \right) dx$$

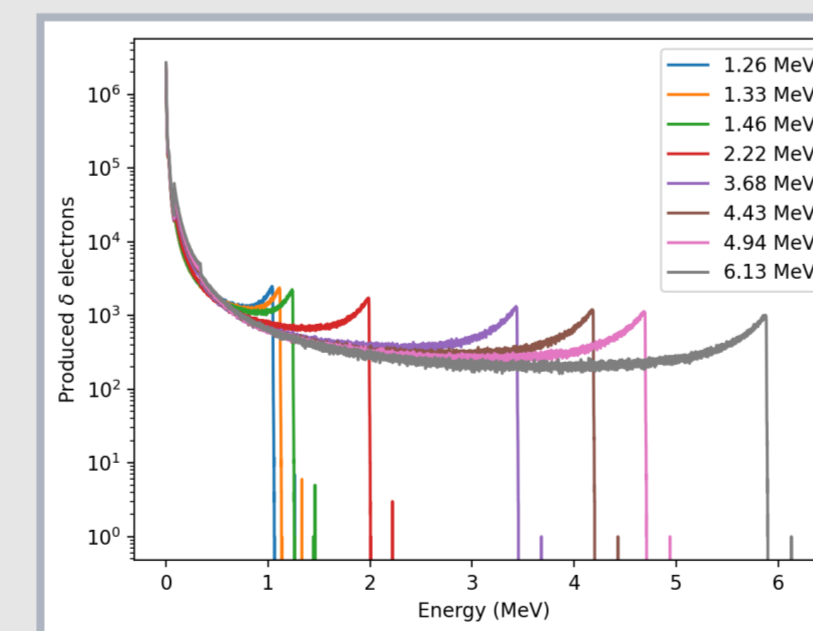
and a statistical resolution term:  $Res_{stat} = \sqrt{a^2 * NPE + NPE_{dark}}$

b. GEPRIS: Gamma-Electron-Positron Response in Scintillator calculates the electron non-linearity  $f_{NL}^e$  as:

$$f_{NL}^e \equiv \frac{E_{vis}}{E_{dep}} = f_{quench} \left( E, \frac{dE}{dx}, kB \right) + f_{Cherenkov}(E) + f_{instNL}$$

Gammas are described as a function of their delta electron production  $P(E^e)$ , simulated in Geant4:

$$\frac{E_{vis}^Y}{E_{dep}^Y} = \frac{\int_0^{E_{max}^e} P(E^e) * E^e * f_{NL}^e dE^e}{\int_0^{E_{max}^e} P(E^e) * E^e dE^e}$$



Simulated delta electron energy distributions for gammas of various energies

## References

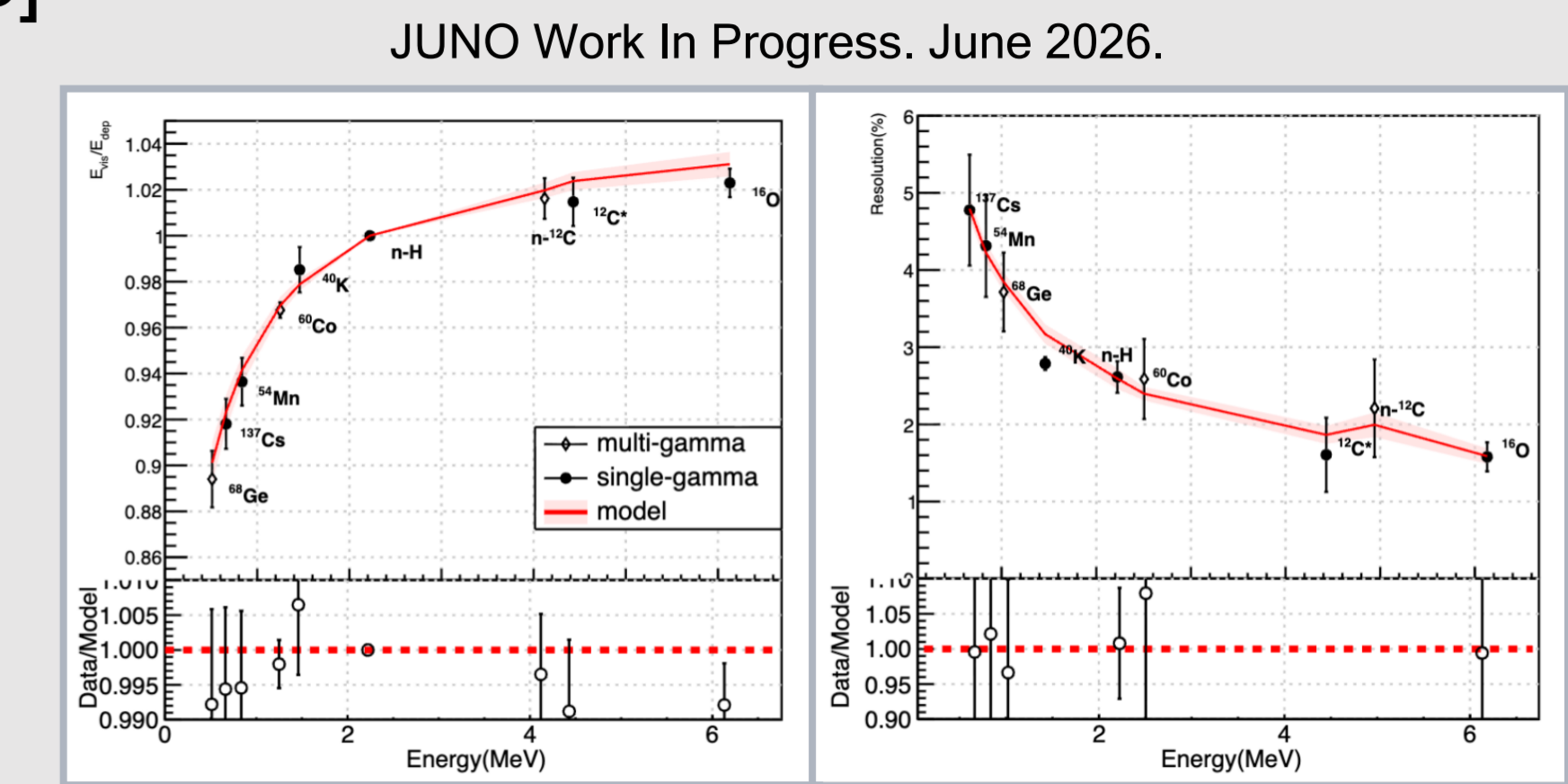
- [1] Abusleme et al, Initial performance results of the JUNO detector. Chinese Phys. C 50 (2026) 043001
- [2] <https://github.com/robx97/GEPRIS>
- [3] Mougeot, X. BetaShape: Calculation of Beta and Neutrino Spectra, Mean Energies and Log fit values. Laboratoire National Henri Becquerel (LNE-LNHB), CEA-Saclay.
- [4] Berger, M. J. ESTAR, PSTAR, and ASTAR: Computer Programs for Calculating Stopping-Power and Range Tables for Electrons, Protons, and Helium Ions. NISTIR 4999, National Institute of Standards and Technology.
- [5] The JUNO collaboration., Abusleme, A., Adam, T. et al. Calibration strategy of the JUNO experiment. J. High Energy. Phys.2021, 4 (2021)

## Acknowledgements

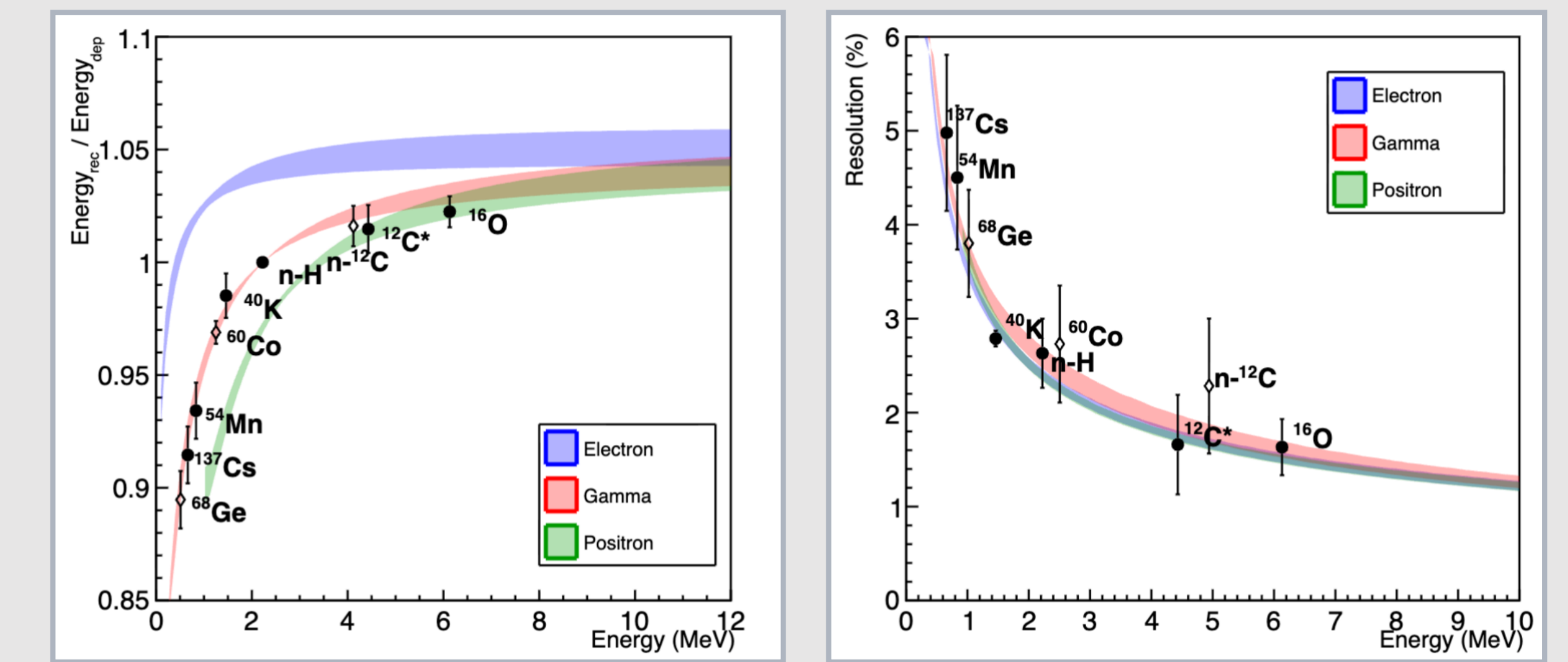
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## 6. COMEER Results [5]

- Volume weighted average of resolutions was calculated
- Fits to calibration data at different detector positions were performed separately



Discrepancy between model and data is within 1%



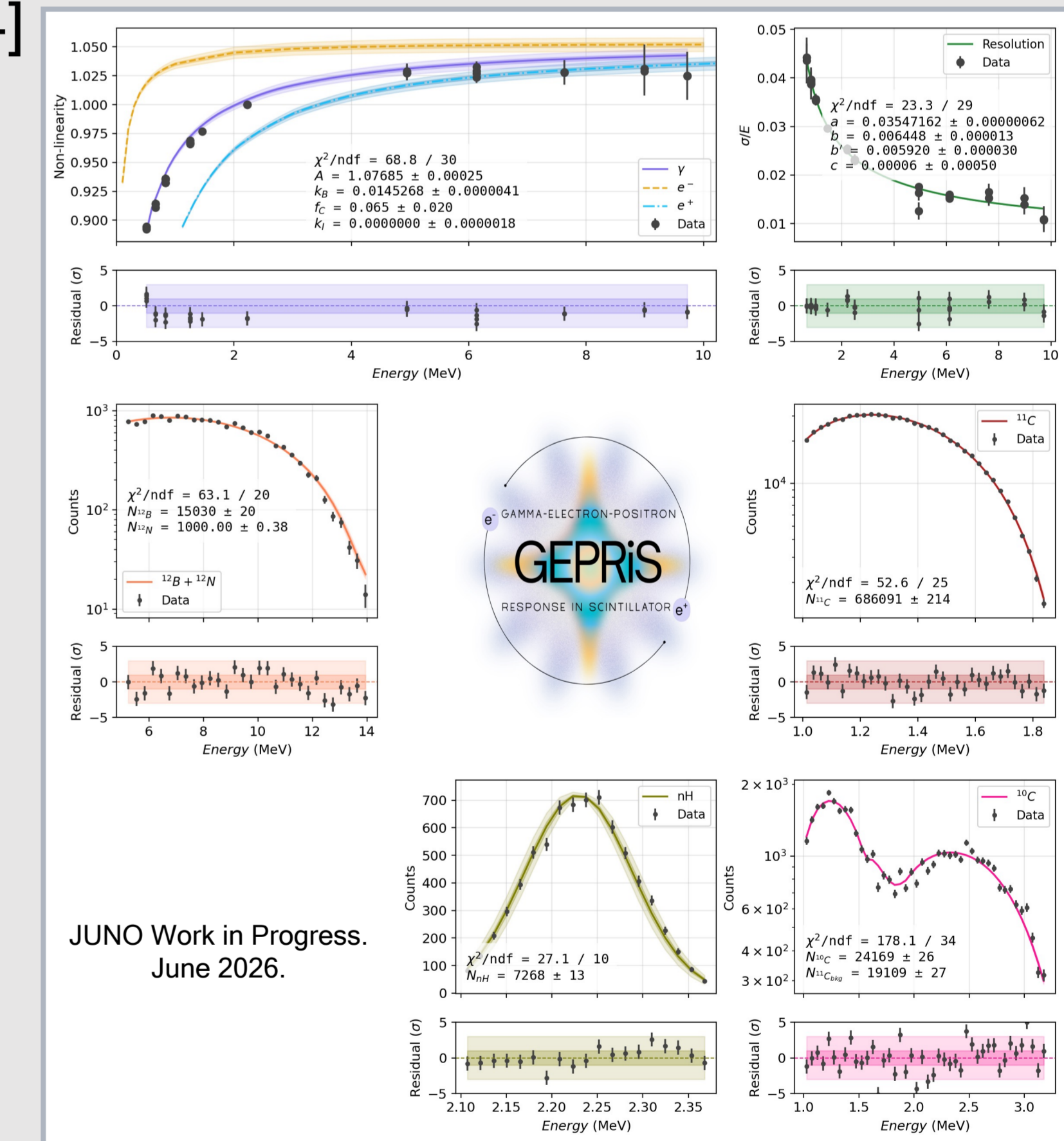
COMEER predictions for gammas, positrons and electrons

## 7. GEPRIS Results [2,3,4]

- Joint fit of gamma points, their resolution and several cosmogenic beta spectra taken between August 2025 and May 2026

- Additional parameter b' added to capture larger resolution in uniformly distributed samples

- Gamma points from different rounds of calibration fit concurrently



## 7. Comparison and Outlook

- Positron energy non-linearity is the main focus, mapping positron to neutrino energy in oscillation analysis
- Positron results between models agree within  $1\sigma$
- Further study to include larger fiducial volume (within 17.2m) planned in the future

