



Probing Non-Standard Interactions in NOvA with Bayesian MCMC Methods

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The NOvA Experiment

- ▶ Long-baseline ν oscillation experiment at Fermilab, operational since 2014
- ▶ Two liquid-scintillator detectors **810 km** apart: 0.3 kt Near Detector (FNAL) and 14 kt Far Detector (Ash River, MN)
- ▶ 14.6 mrad off-axis beam \Rightarrow narrow peak at **1.8 GeV**
- ▶ Long baseline \Rightarrow large **MSW matter effect**, sensitive to new physics



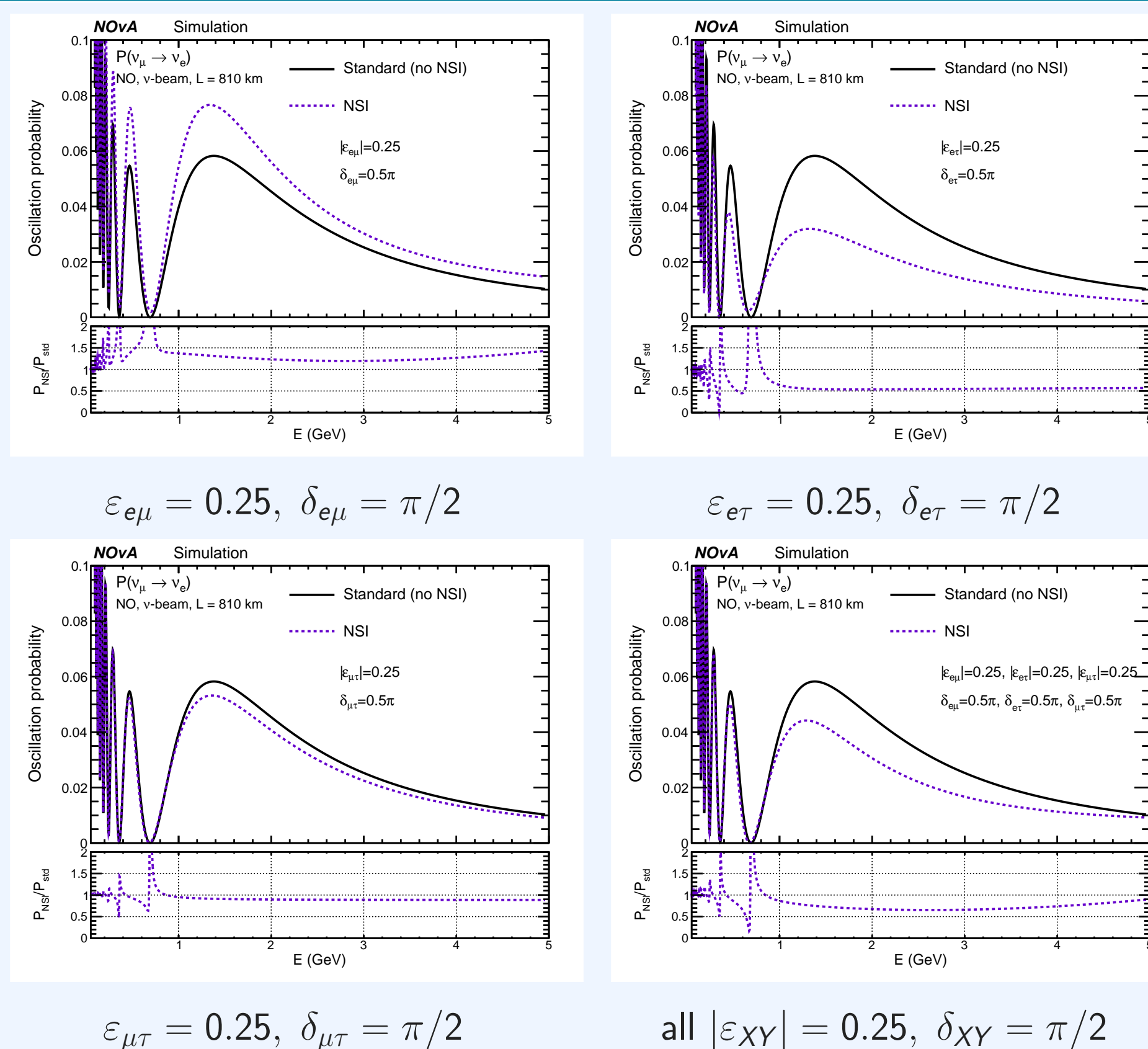
Non-Standard Interactions (NSI)

NSI add beyond-SM contact interaction terms to the matter Hamiltonian:

$$\mathcal{H} = U \text{diag}(0, \Delta_{21}, \Delta_{31}) U^\dagger + V_e \begin{pmatrix} 1 + \varepsilon_{ee} & \varepsilon_{e\mu} & \varepsilon_{e\tau} \\ \varepsilon_{e\mu}^* & \varepsilon_{\mu\mu} & \varepsilon_{\mu\tau} \\ \varepsilon_{e\tau}^* & \varepsilon_{\mu\tau}^* & \varepsilon_{\tau\tau} \end{pmatrix}$$

- ▶ Off-diagonal $\varepsilon_{\alpha\beta}$ are **complex**: magnitude $|\varepsilon|$ + CP-violating phase δ ; SM limit $\varepsilon_{\alpha\beta} = 0$
- ▶ NOvA is sensitive to $\varepsilon_{e\mu}$ and $\varepsilon_{e\tau}$ via $\nu_\mu \rightarrow \nu_e$ appearance and the MSW matter effect

NSI Effect on $\nu_\mu \rightarrow \nu_e$ Appearance



- ▶ Different NSI sectors **could** change oscillation probability in opposing ways
- ▶ A frequentist profile-likelihood treats each sector independently; Bayesian MCMC allows **all sectors to vary simultaneously** in a single fit

Bayesian MCMC: Metropolis-Hastings

$P(\theta | \text{data}) \propto \mathcal{L}(\text{data} | \theta) \pi(\theta)$, $\mathcal{L} \propto e^{-\frac{1}{2}\chi^2_{\text{NOvA}}}$ (joint $\nu_e + \nu_\mu$, FHC+RHC); NO+IO in one chain; systematics **marginalized**.

Mechanism:

Propose (Gaussian step, all params) \rightarrow **Score** ($\log \mathcal{L}' = -\frac{1}{2}\chi^2 + \log \pi$) \rightarrow **Accept** ($\alpha = \min(1, e^{\Delta \log \mathcal{L}'})$) or **stay** \rightarrow \odot

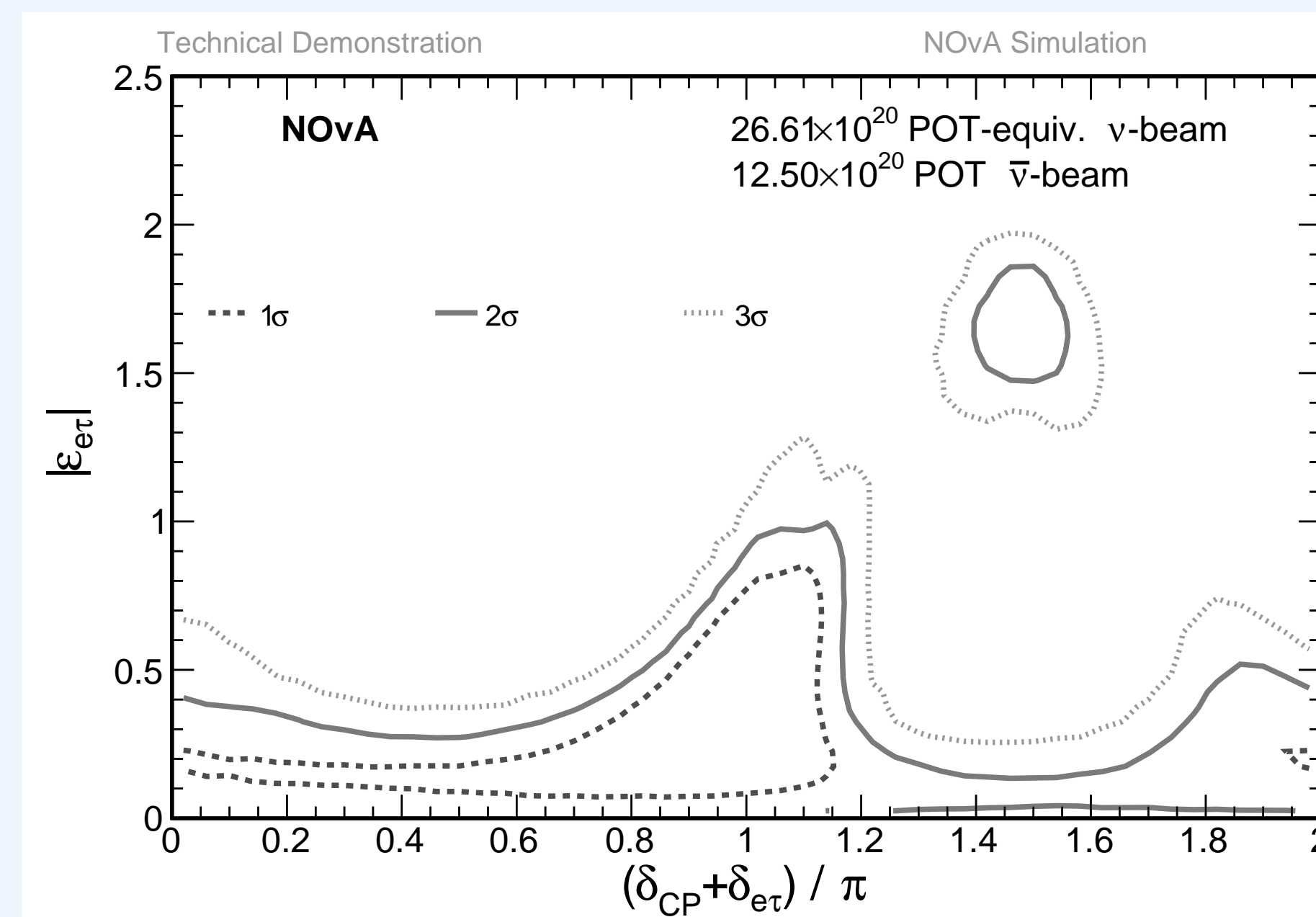
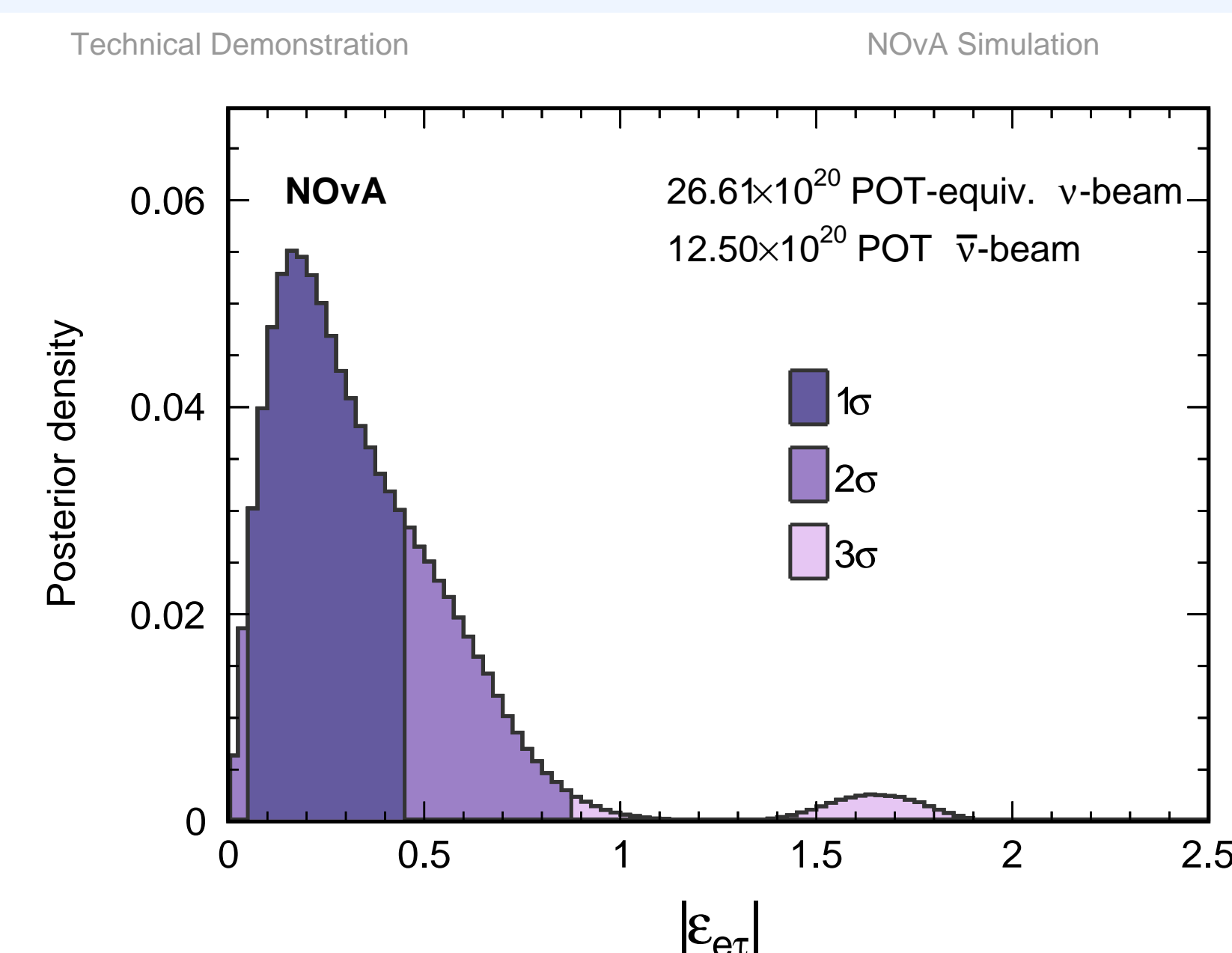
The **prior** $\pi(\theta)$ for complex NSI parameters $\varepsilon_{\alpha\beta}$ is non-trivial — see analysis below \downarrow

Prior Sensitivity in NSI: Three Approaches ($\varepsilon_{e\tau}$ example)

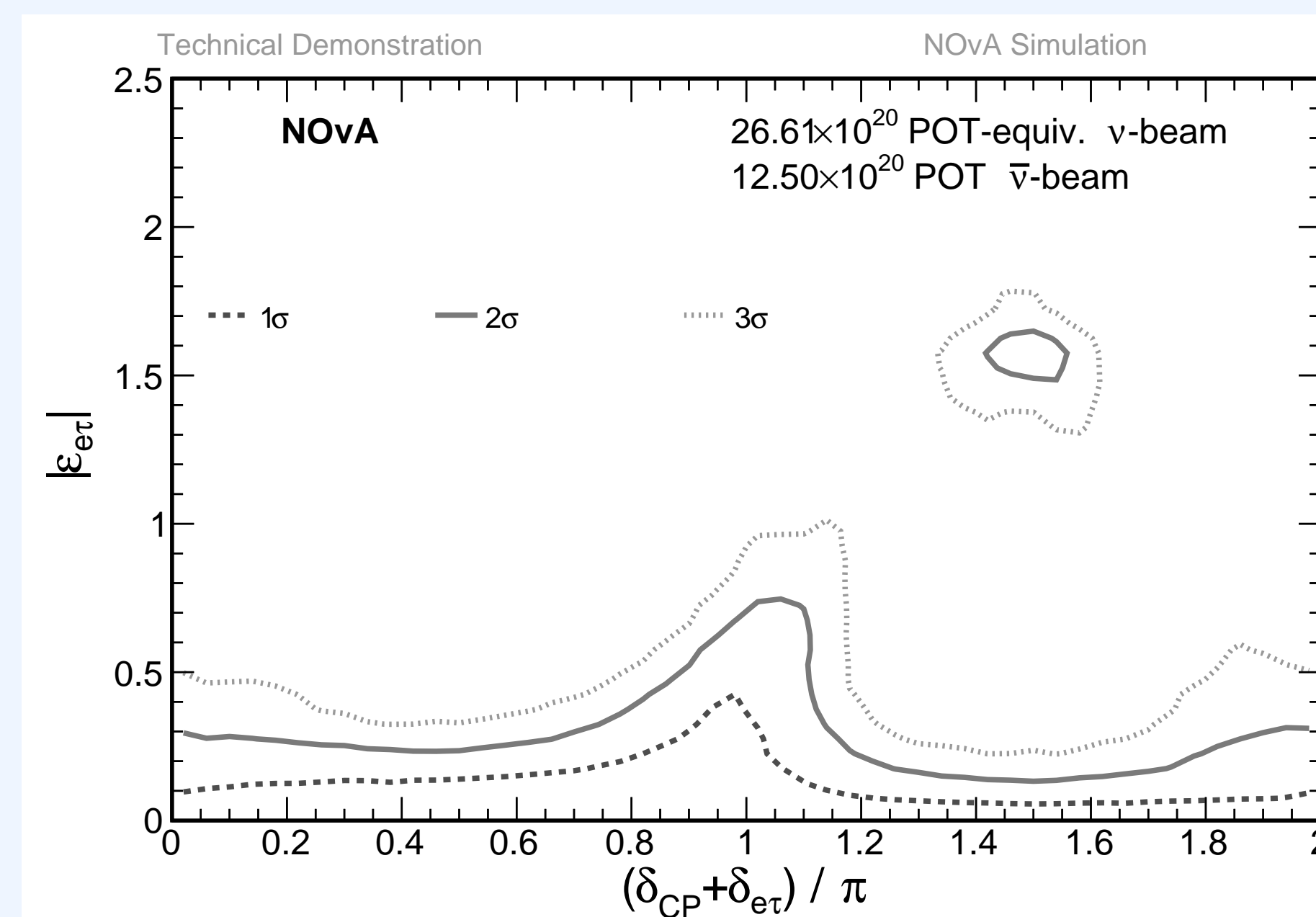
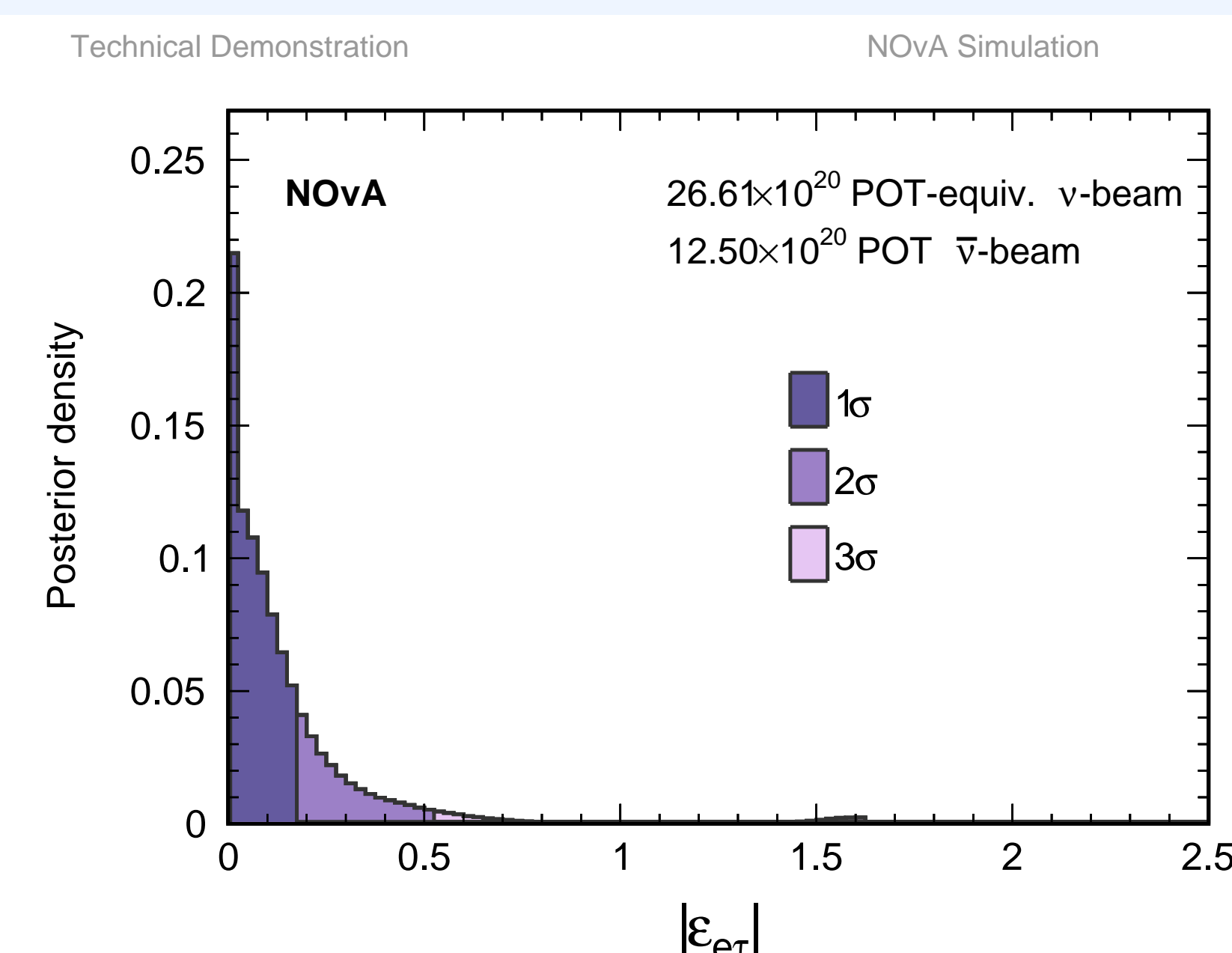
1D posterior, $|\varepsilon_{e\tau}|$ (NO+IO)

2D posterior, $|\varepsilon_{e\tau}|$ vs $\delta_{CP} + \delta_{e\tau}$

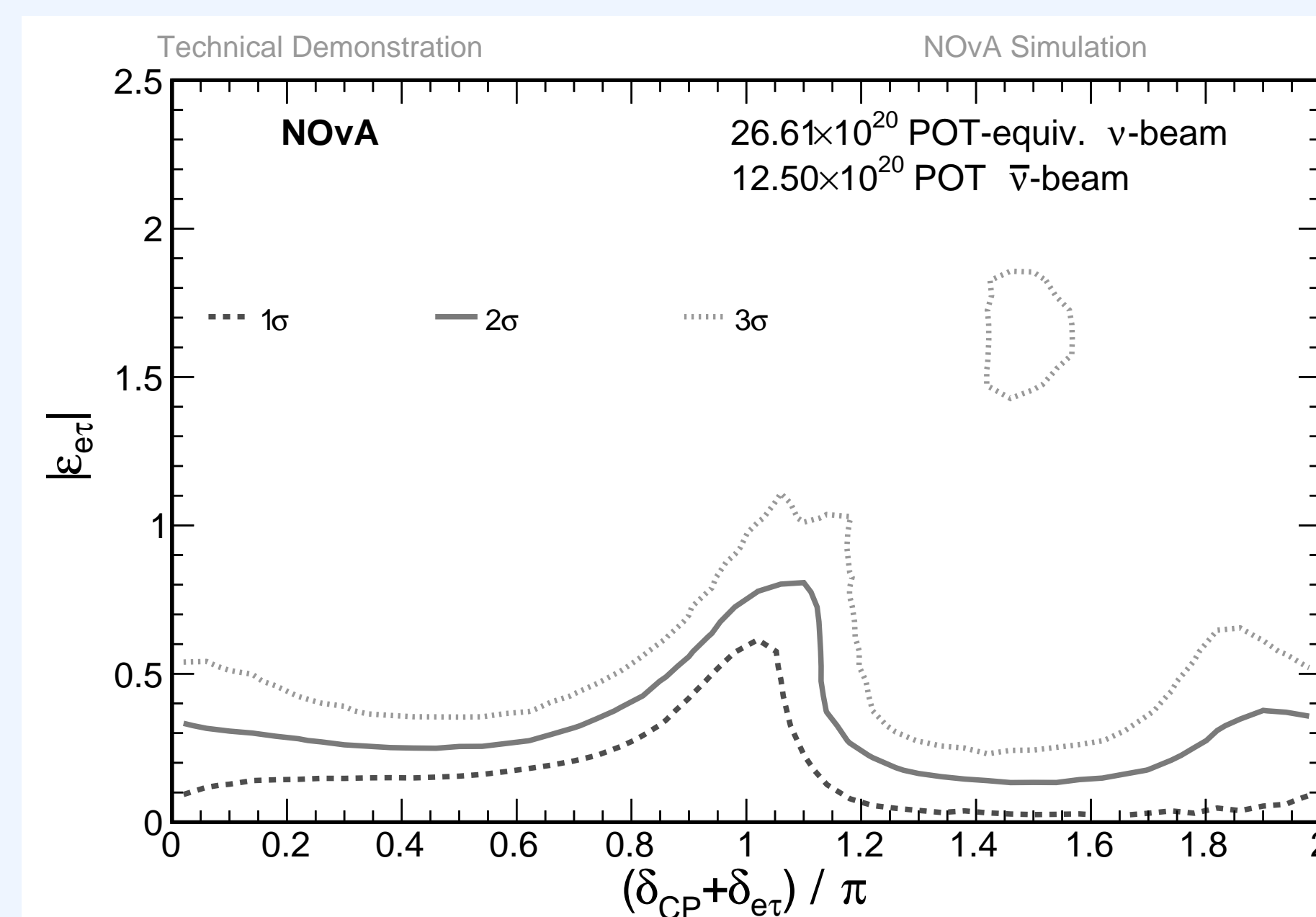
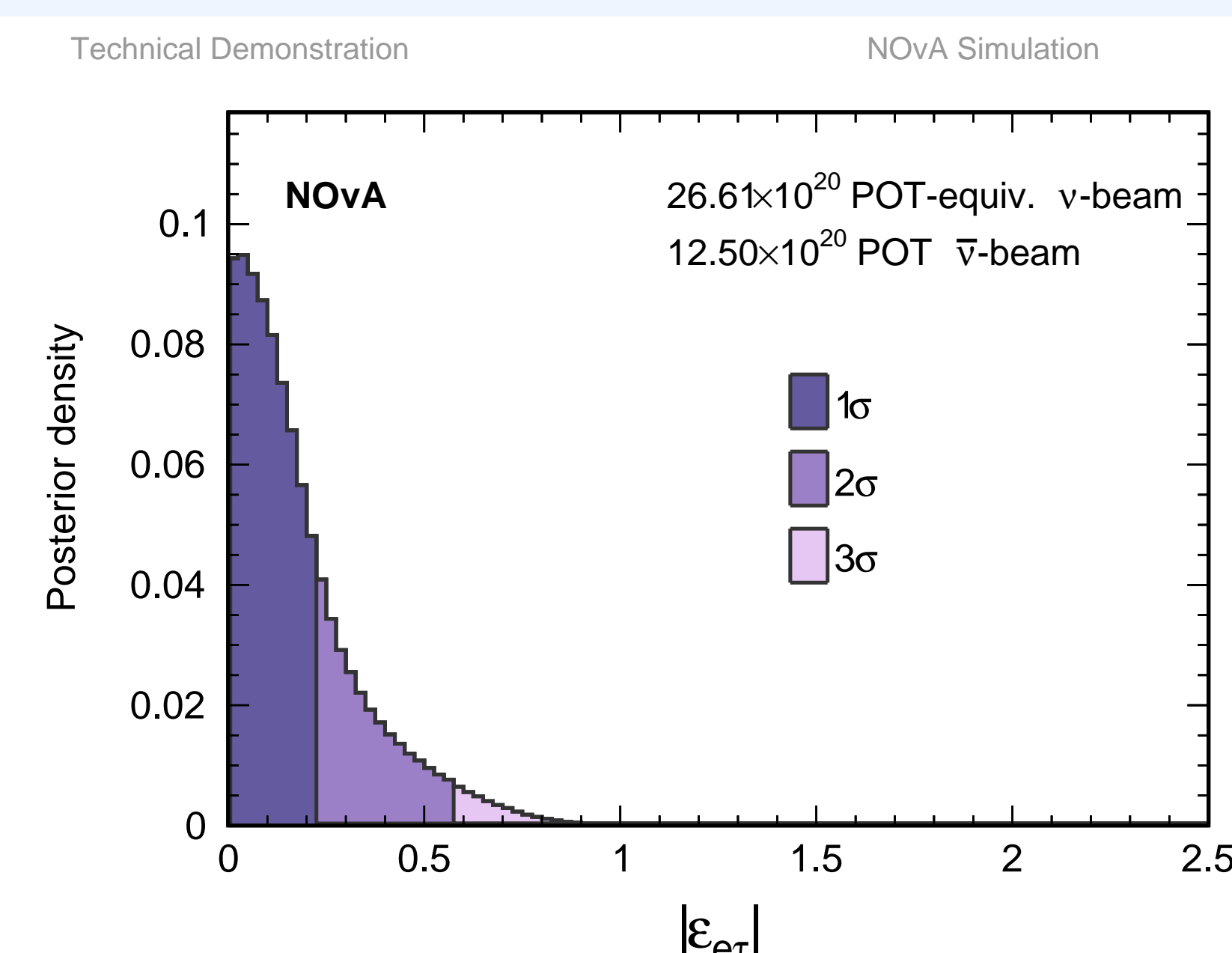
A: Flat Cartesian — $\pi(x, y) = \text{const} \Rightarrow \pi(|\varepsilon|) \propto |\varepsilon|$



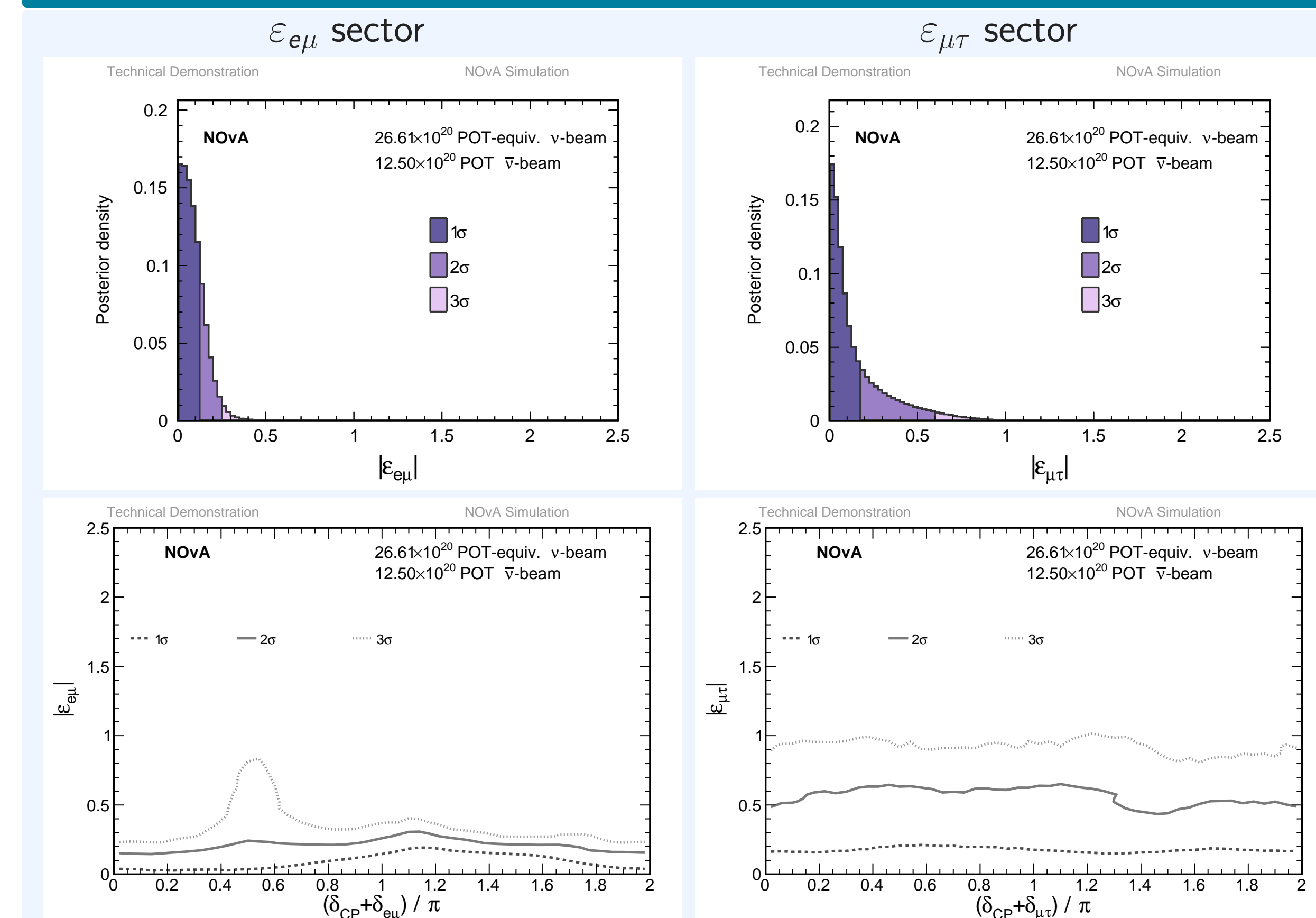
B: Mixture Prior — $\pi(|\varepsilon|) = 0.1 \cdot \mathcal{HN}(\sigma=0.1) + 0.9 \cdot \mathcal{HN}(\sigma=1.0)$



C: Jacobian Correction \star — $\log \pi(x, y) = -\log(|\varepsilon| + \varepsilon_{\text{reg}})$, $\varepsilon_{\text{reg}} = 10^{-4}$ [Preferred]



Jacobian Correction Prior for Other NSI Sectors



Analysis and Summary of Prior Selections

- ▶ **Method A** (flat Cartesian): $\pi(x, y) = \text{const}$ induces $\pi(|\varepsilon|) \propto |\varepsilon|$ via the polar Jacobian — the prior assigns zero probability to $|\varepsilon| = 0$, creating a **dip** at $|\varepsilon| \rightarrow 0$ and suppressing the SM limit even when the data have no NSI signal.
- ▶ **Method B** (mixture prior): the narrow HN component boosts probability near $|\varepsilon| \approx 0$, but the prior shape dominates the statistics near zero, creating a **spike** at $|\varepsilon| = 0$ — not suitable as a final result.
- ▶ **Method C** (Jacobian correction, **recommended**): $-\log(|\varepsilon| + \varepsilon_{\text{reg}})$ added to the log-prior cancels the $|\varepsilon|$ Jacobian, restoring a **flat marginal prior** in $|\varepsilon|$. Dip and spike are both removed; results show good qualitative agreement with the frequentist profile likelihood.

Conclusions

- ▶ Bayesian NSI analysis is a **unique challenge**: complex $\varepsilon_{\alpha\beta}$ makes prior choice non-trivial; **Method C (Jacobian correction)** is the recommended approach
- ▶ These results are a **technical demonstration** toward full validation relative to the NOvA frequentist NSI approach

Acknowledgments & References

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[1] M. A. Acero et al. (NOvA Collaboration), *Phys. Rev. D* **110**, 012005 (2024), arXiv:2311.07835.

[2] S. Brooks et al. (eds.), *Handbook of Markov Chain Monte Carlo*, Chapman & Hall/CRC (2011).