



Observatoire
de Paris

PSL 



LTE

 Université
de Lille

 SORBONNE
UNIVERSITÉ

LABORATOIRE
NATIONAL
DE MÉTROLOGIE
ET D'ESSAIS 

Laboratoire Temps Espace

An optical lattice clock with a bosonic isotope of mercury

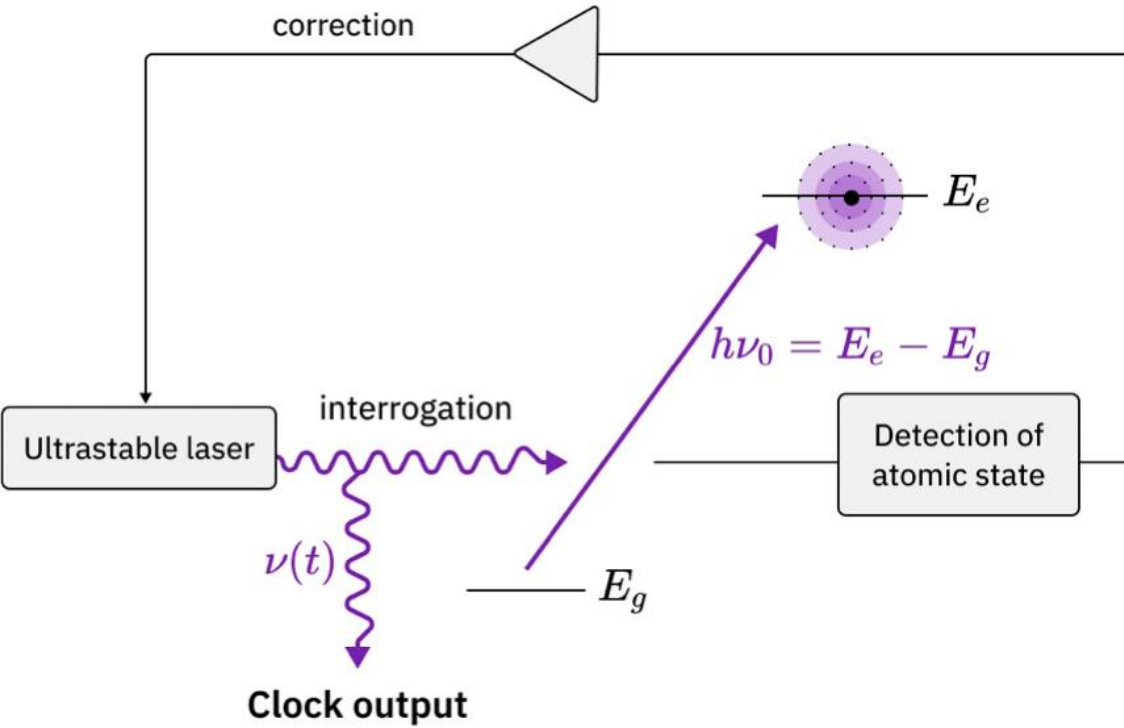
International Conference on Precision Physics of Simple Atomic Systems, Vienna

May, 21st 2026

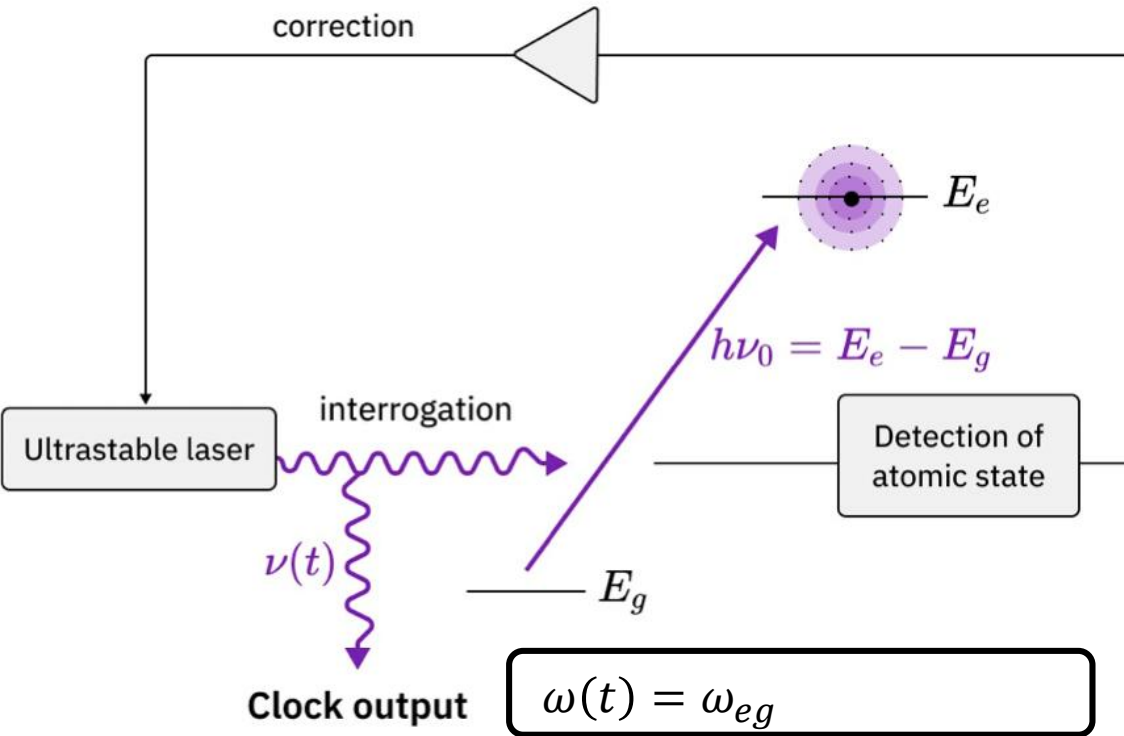
Ashley Béguin

 ppas 2026

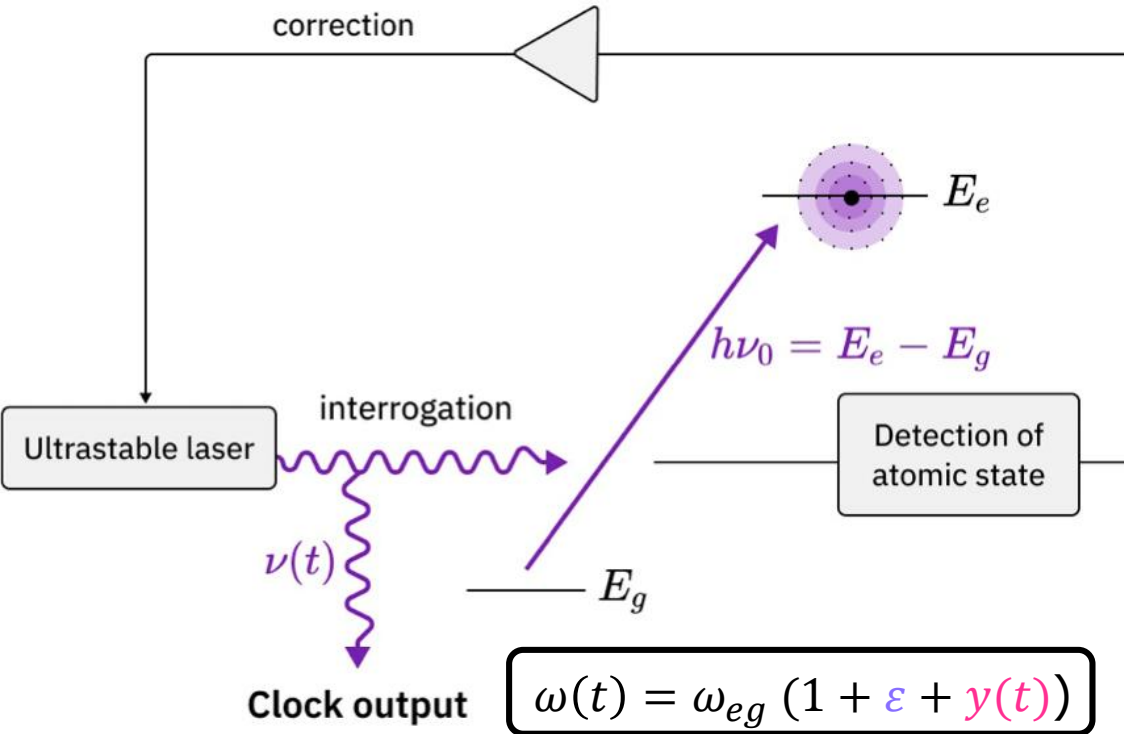
Clock principle



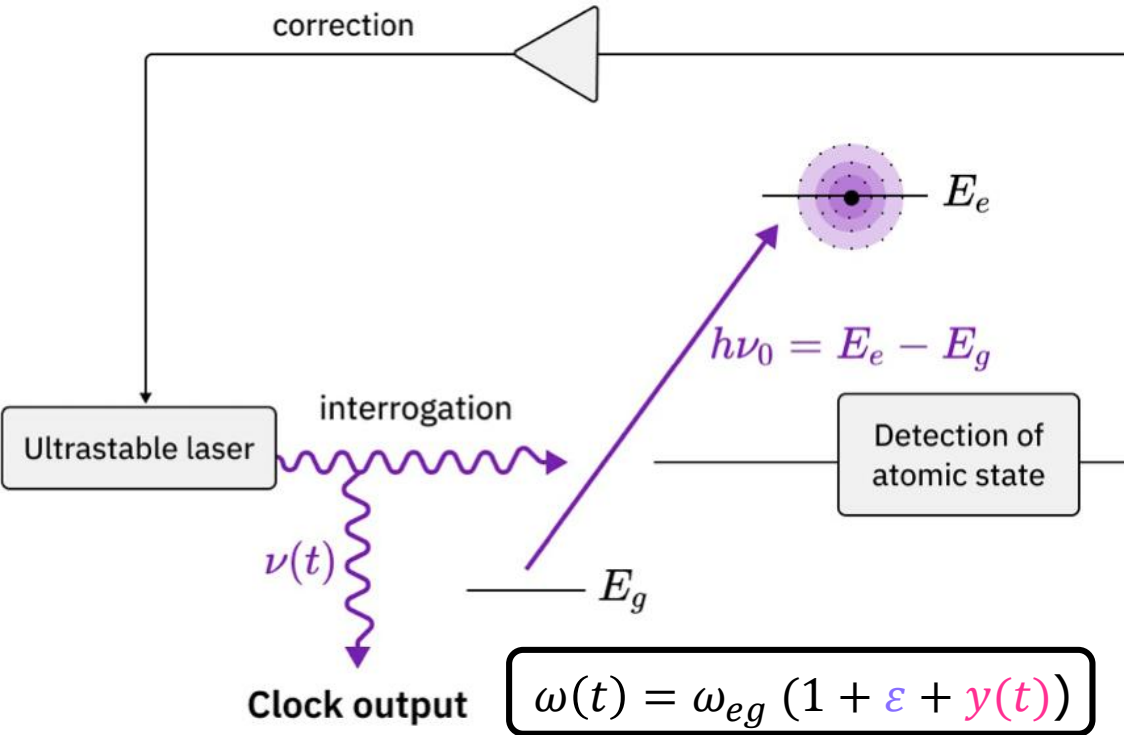
Clock principle



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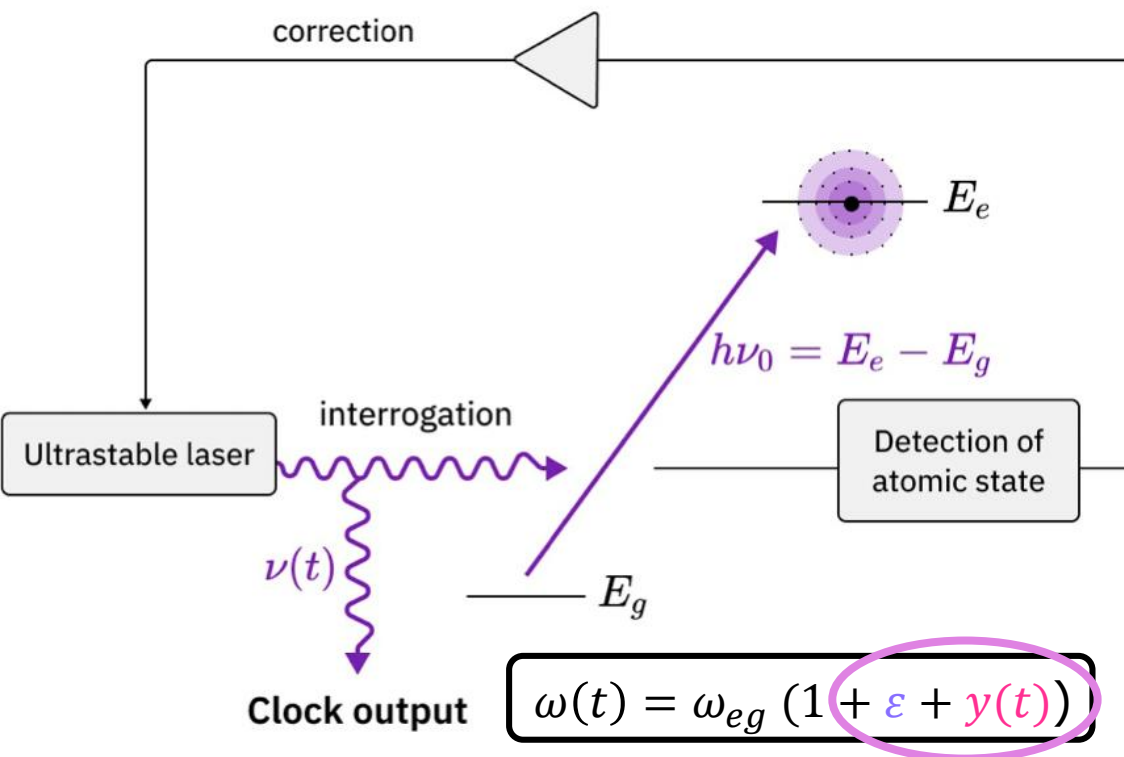
Clock principle



Performances:

- ❑ High **accuracy**: overall uncertainty on ε
- ❑ High **stability**: statistical properties of $y(t)$
- ❑ Low sensitivity to external perturbations
- ❑ Reproducible

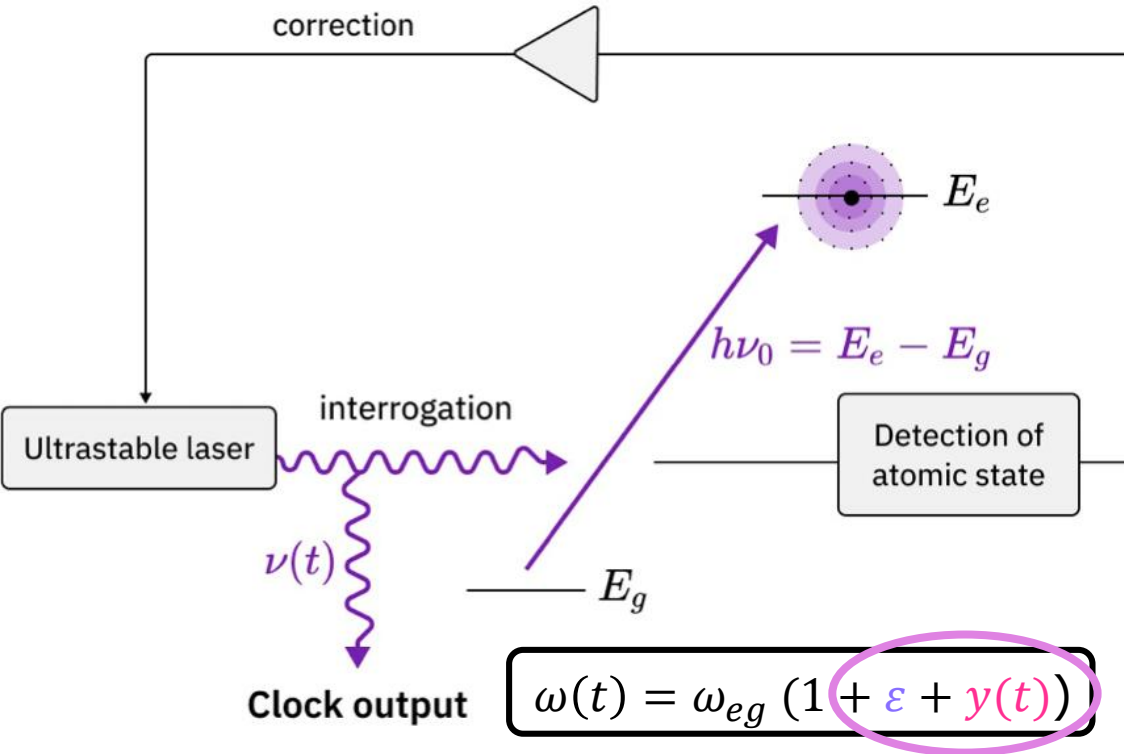
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Clock principle



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Physical effects inducing noise and shifts:

- Main sources of noise
 - Laser
 - Detection
- Effects of trapping
- Effects of thermal radiation
- Interactions between atoms
- Many other physical and technical effects

Applications

□ Metrology

- Better time & frequency references
- Redefinition of the SI second

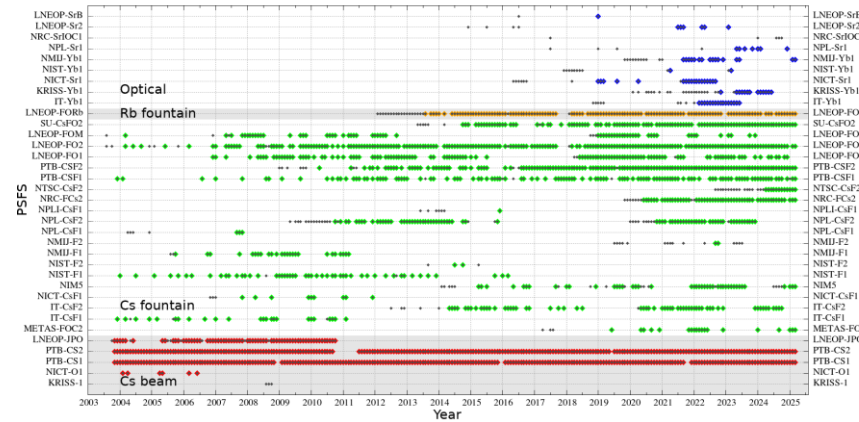
□ Tests of fundamental laws

- Test of GR *Living Rev. Relativ. 17, 4 (2014)*
- Stability of fundamental constants *Rev. Mod. Phys. 90, 025008 (2018)*
- Search for dark matter *Living Rev Relativ 28, 6 (2025)*
- Probe nuclear properties via isotope shifts *Nat Rev Phys 7, 119 (2025)*

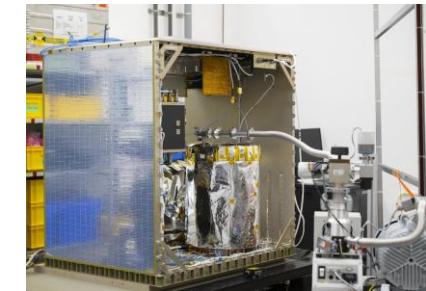
□ Chronometric geodesy

- Transportable optical clocks
- Optical fiber and free-space links

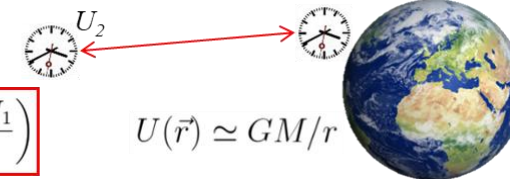
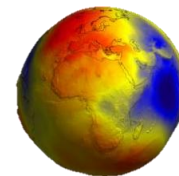
Graphical representation of all evaluations of Primary and Secondary Frequency Standards reported since Circular T 190. Enhanced color dots indicate evaluations carried out within the month of TAI computation.



Optica 13, 143 (2026)



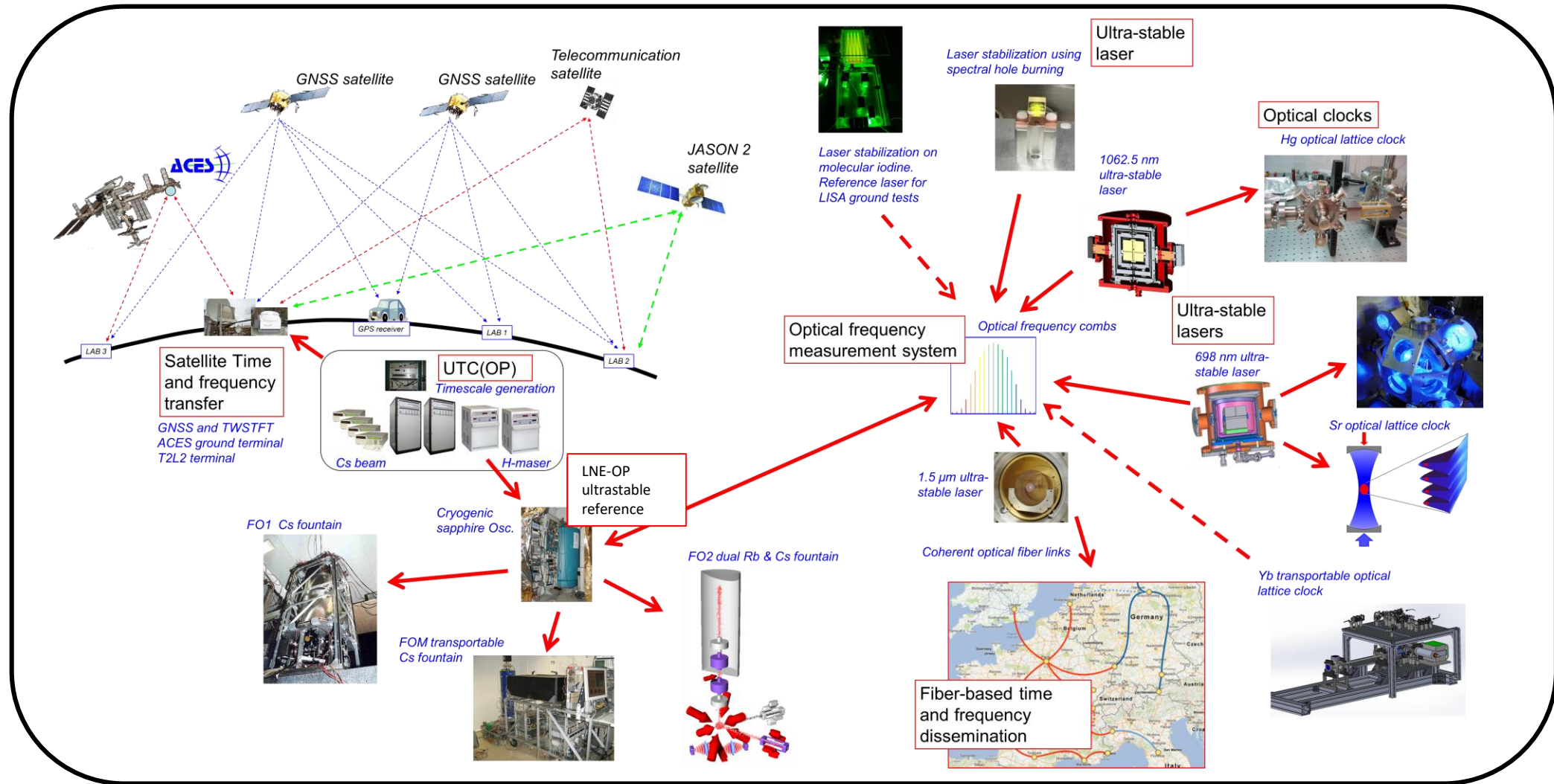
ACES



$$\frac{\nu_2}{\nu_1} = \left(1 - \frac{U_2 - U_1}{c^2}\right)$$

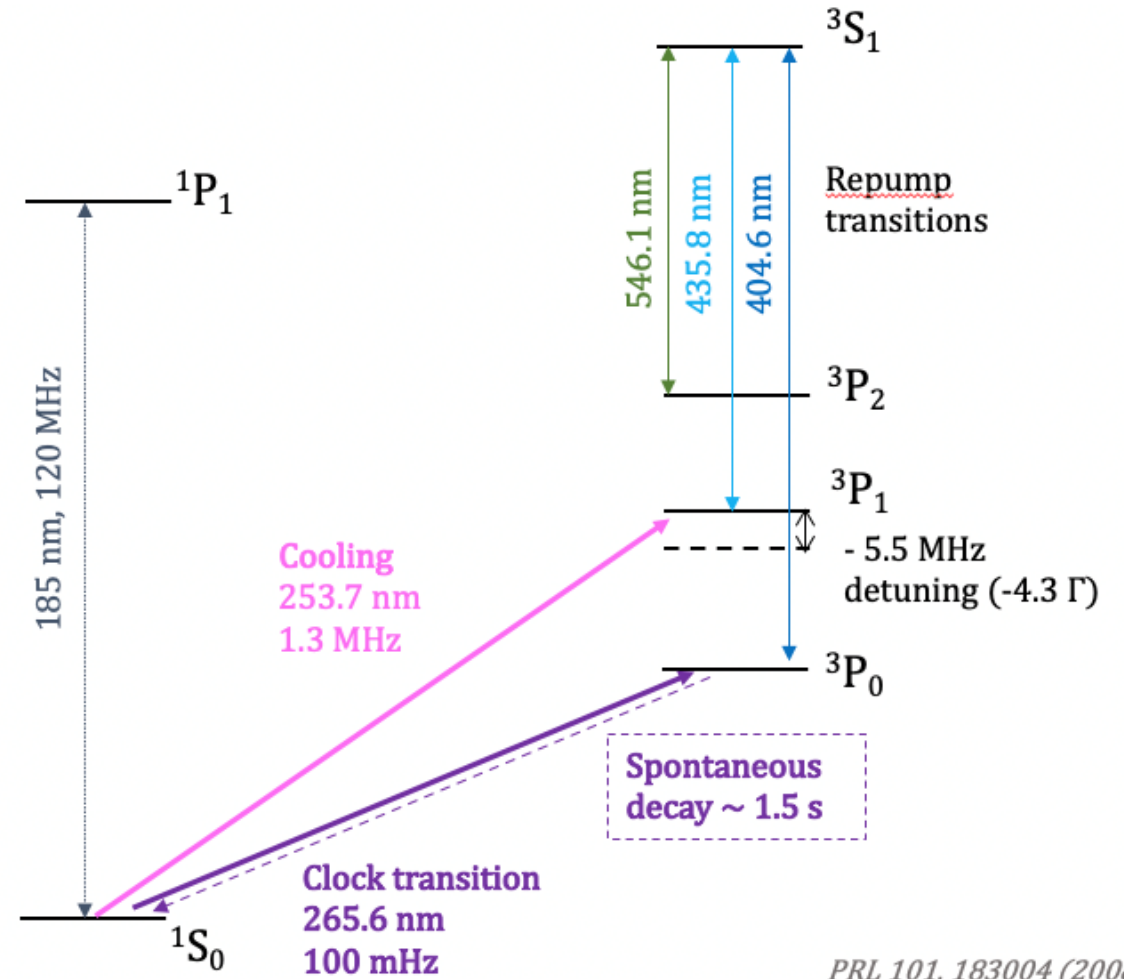
$$U(\vec{r}) \simeq GM/r$$

Laboratoire Temps Espace (LTE) clock ensemble



Why using mercury (Hg) for an optical clock?

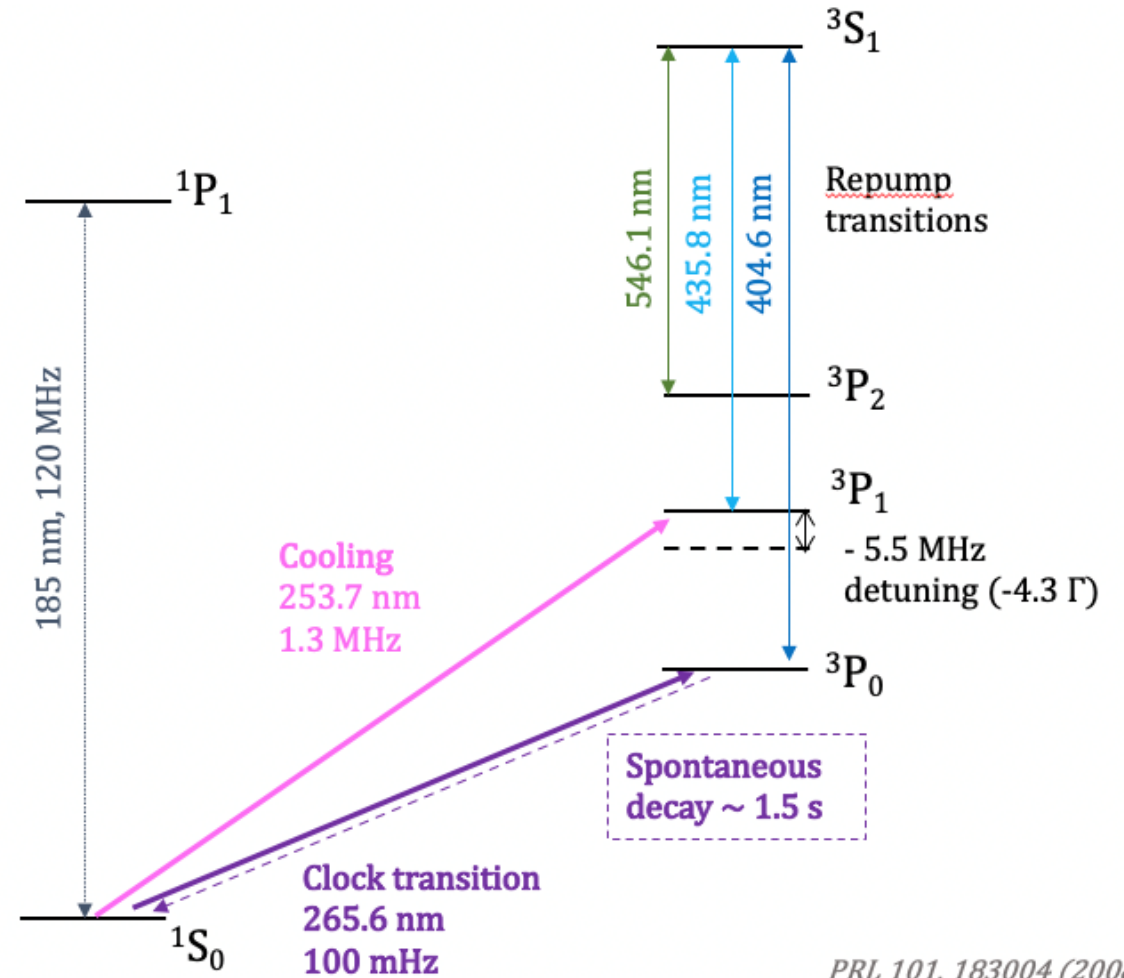
- ❑ Low sensitivity to black body radiation (Yb/16, Sr/30)
- ❑ High vapour pressure (no oven)
- ❑ Simple transition structure (7 stable isotopes)
- ❑ High sensitivity to possible variations of constant α



PRL 101, 183004 (2008)

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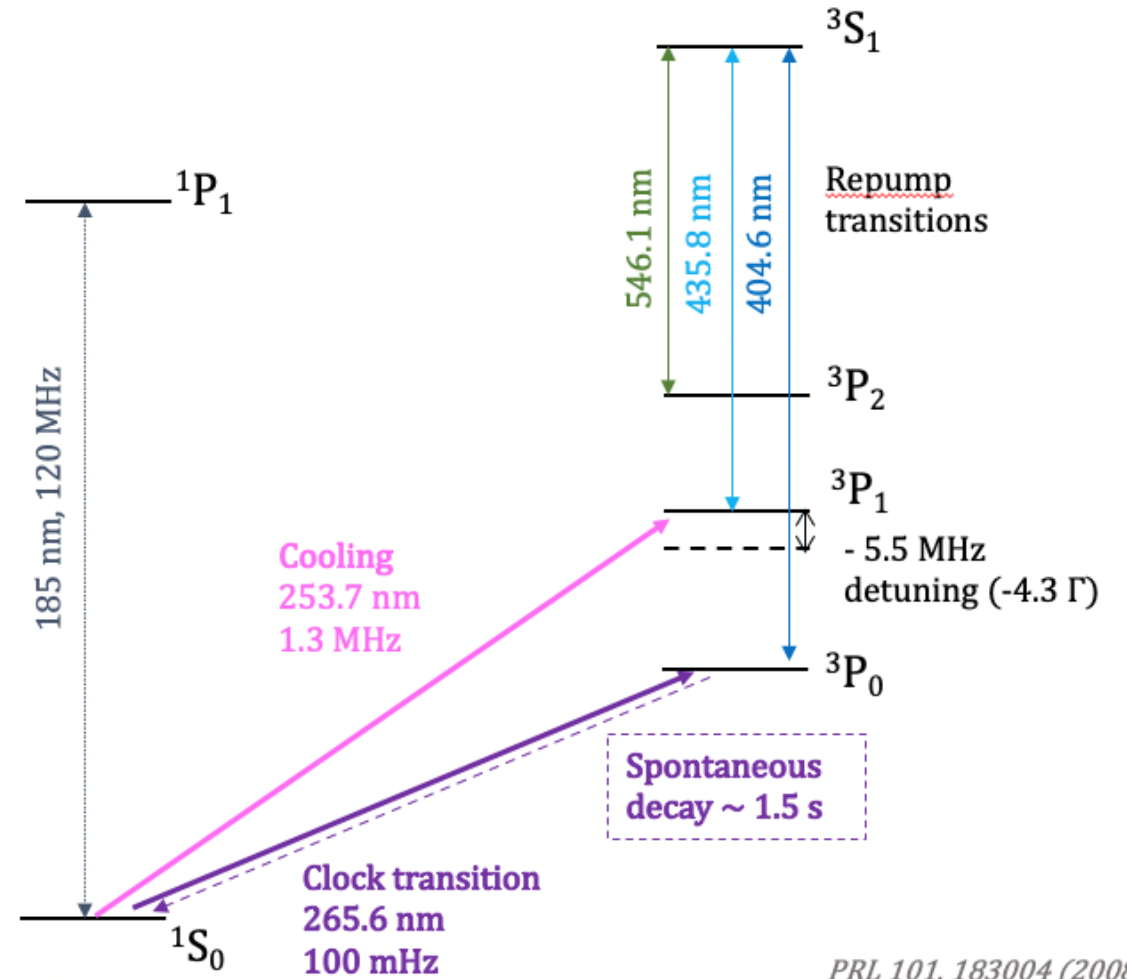


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- So far only ^{199}Hg isotope:**
- ❑ Secondary representation of the second
 - ❑ Performances reached @ LTE:
 - Systematic effects uncertainty : $\sim 5 \times 10^{-17}$ [1]
 - Stability : $8\text{-}9 \times 10^{-16}$ @ 1s [2][3]



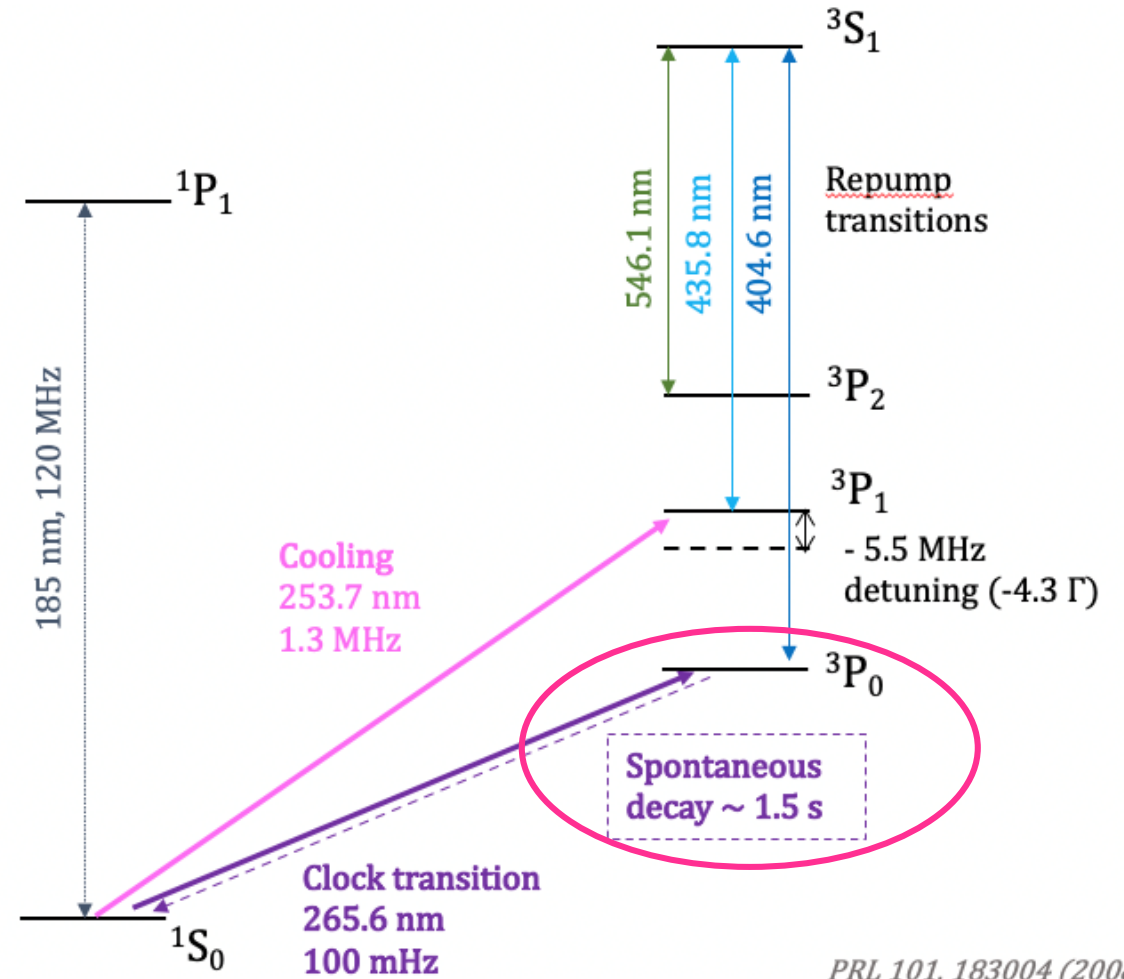
PRL 101, 183004 (2008)

[1] 2024, Zyskind, PhD Thesis
[2] 2016, Tyumenev, NJP18, 113002
[3] 2023, Guo, PRA107, 033116

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- Transition width soon a limitation vs new generation of ultrastable lasers



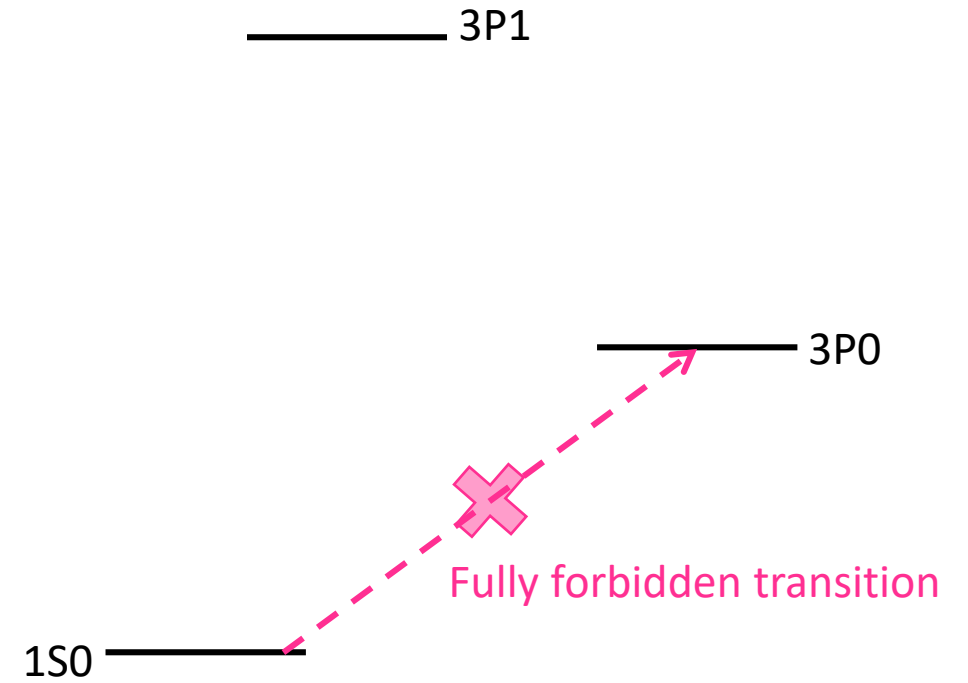
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Towards Hg bosonic isotopes

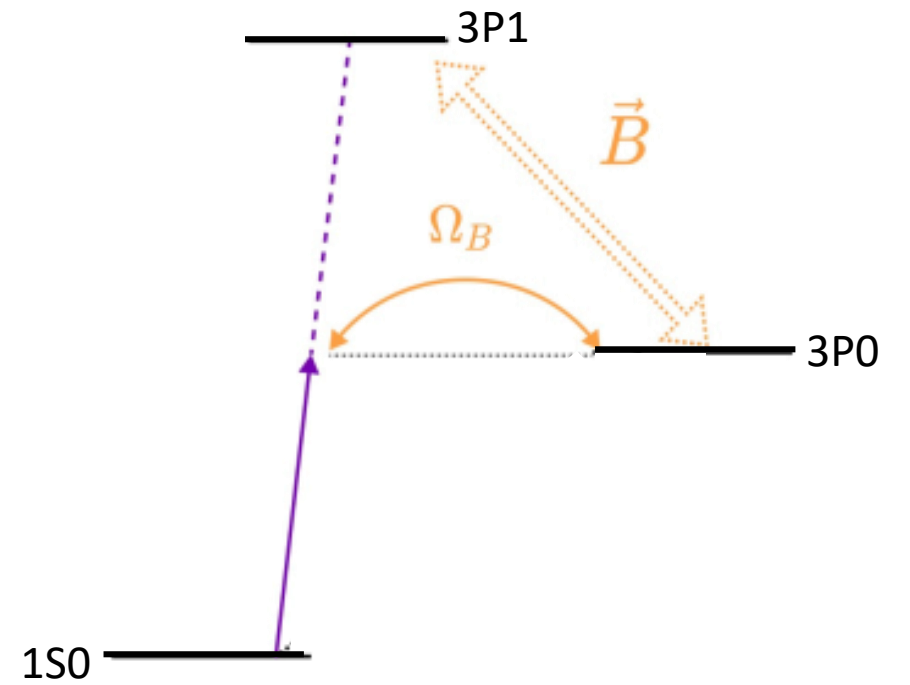


Bosonic isotopes: no nuclear spin



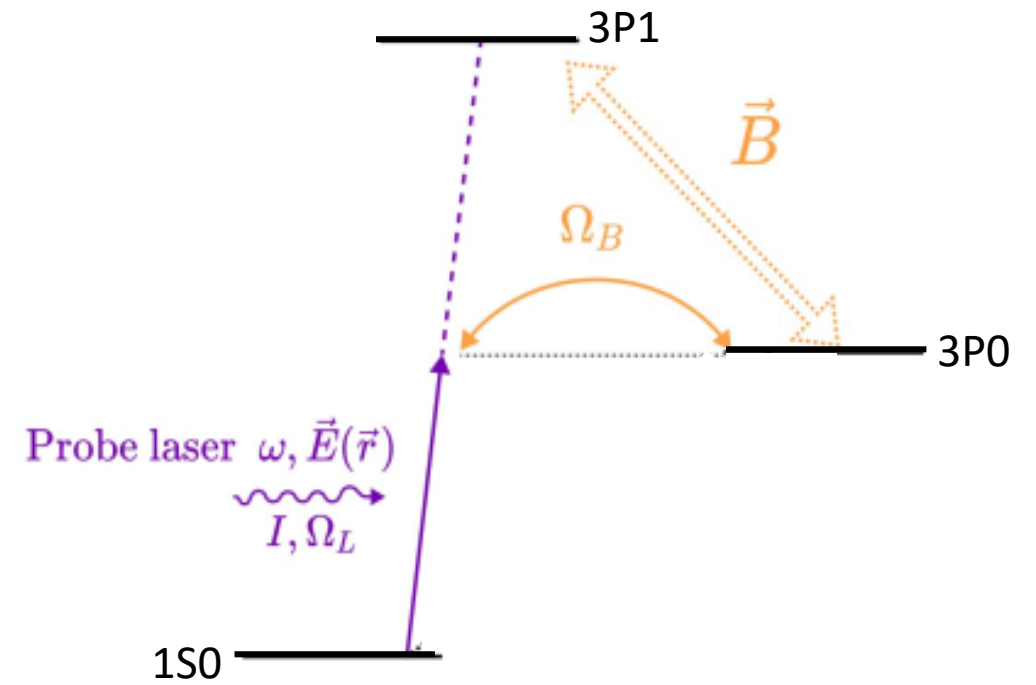
Bosonic isotopes: no nuclear spin

Quenching scheme



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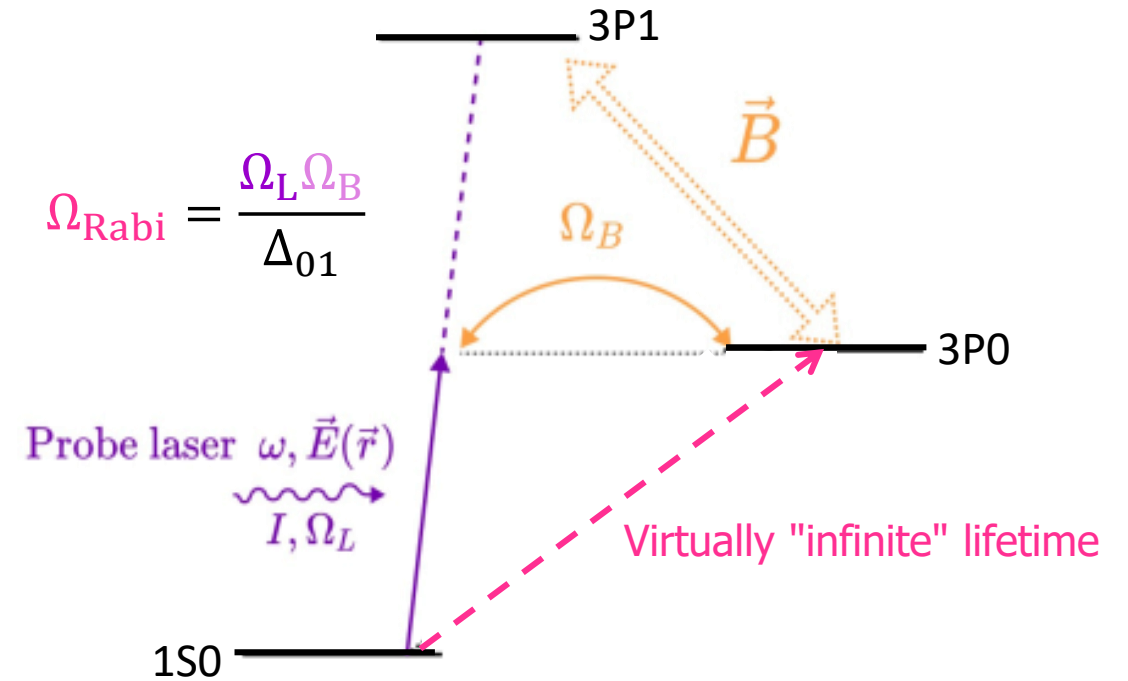
Towards Hg bosonic isotopes



Bosonic isotopes: no nuclear spin

- ❑ Still inducible via state mixing using a strong magnetic field [4]
- ❑ Already demonstrated in ^{88}Sr and ^{174}Yb

Quenching scheme



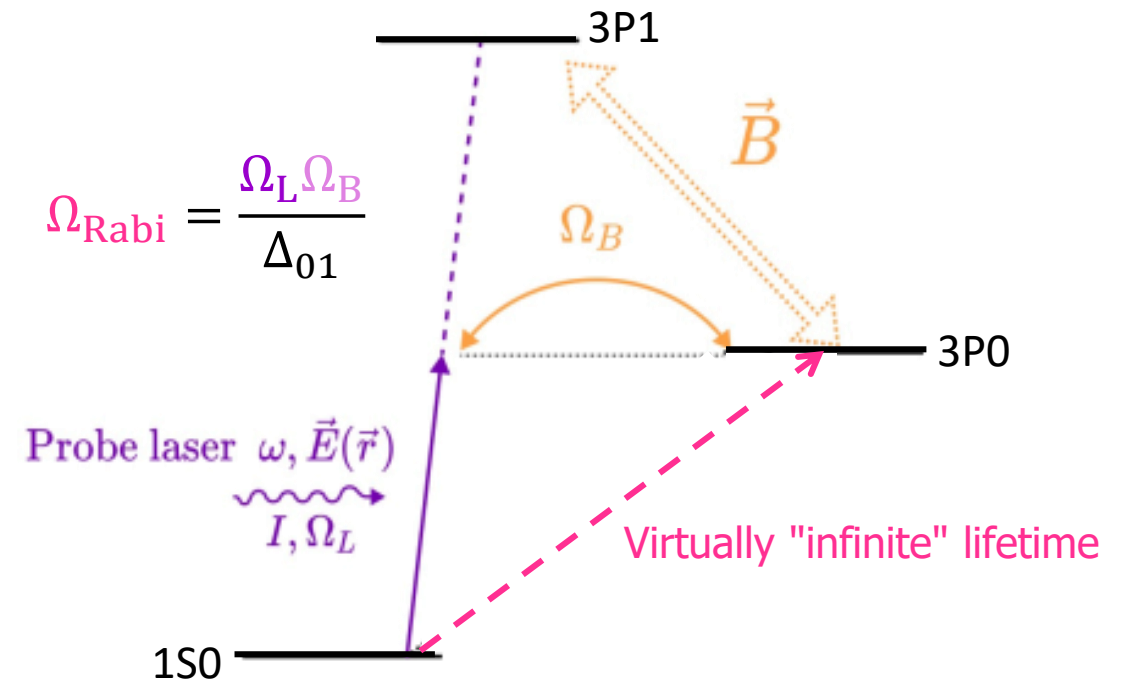
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Weak coupling:

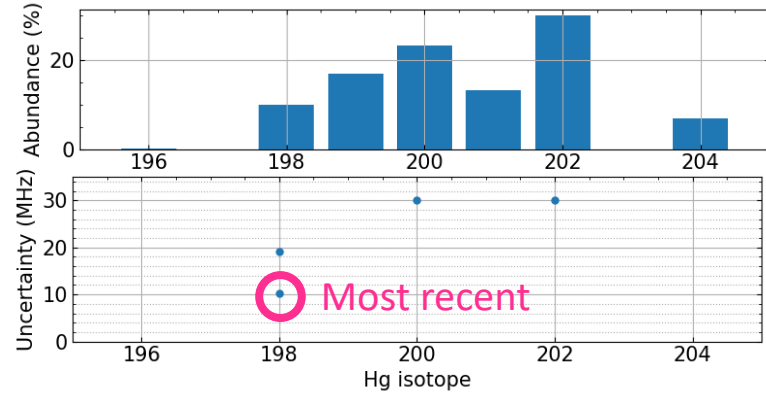
- ❑ Strong magnetic field \Rightarrow Quadratic Zeeman Shift
- ❑ Strong probe power \Rightarrow Probe Light Shift

Quenching scheme



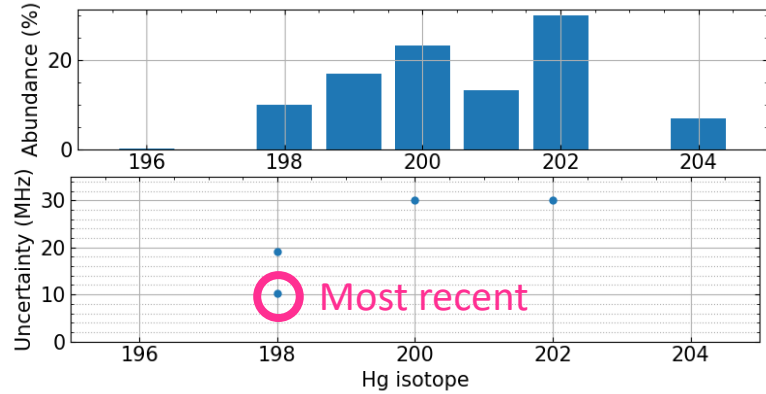
^{198}Hg bosonic isotope

Isotope choice:

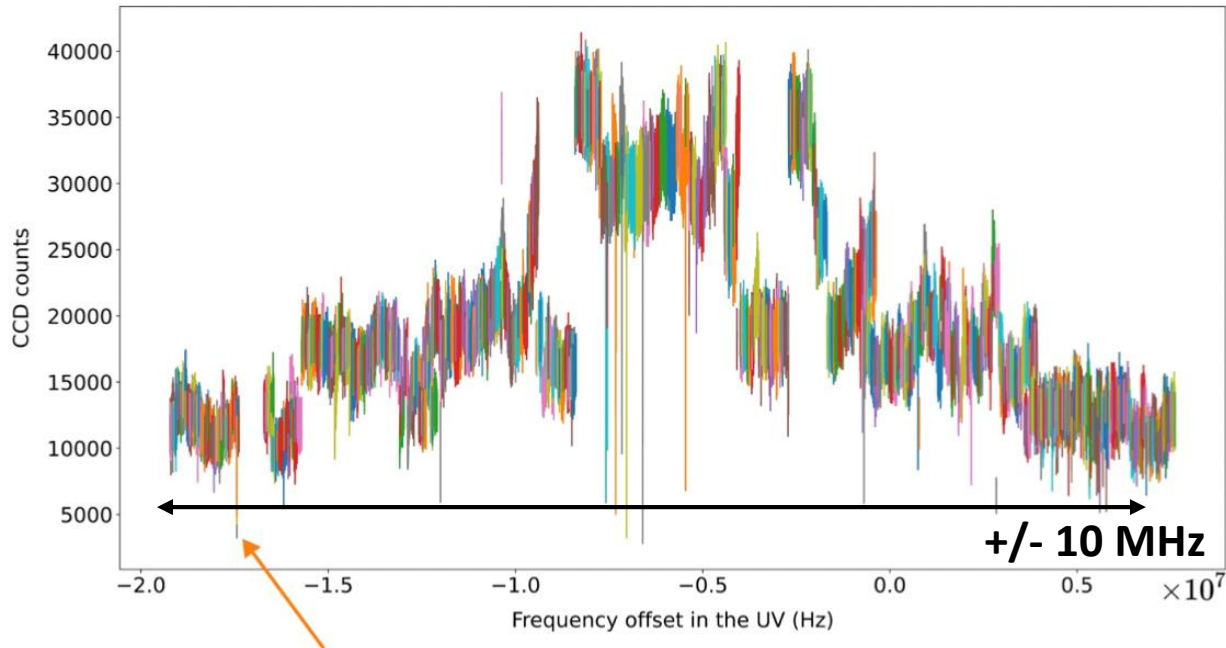


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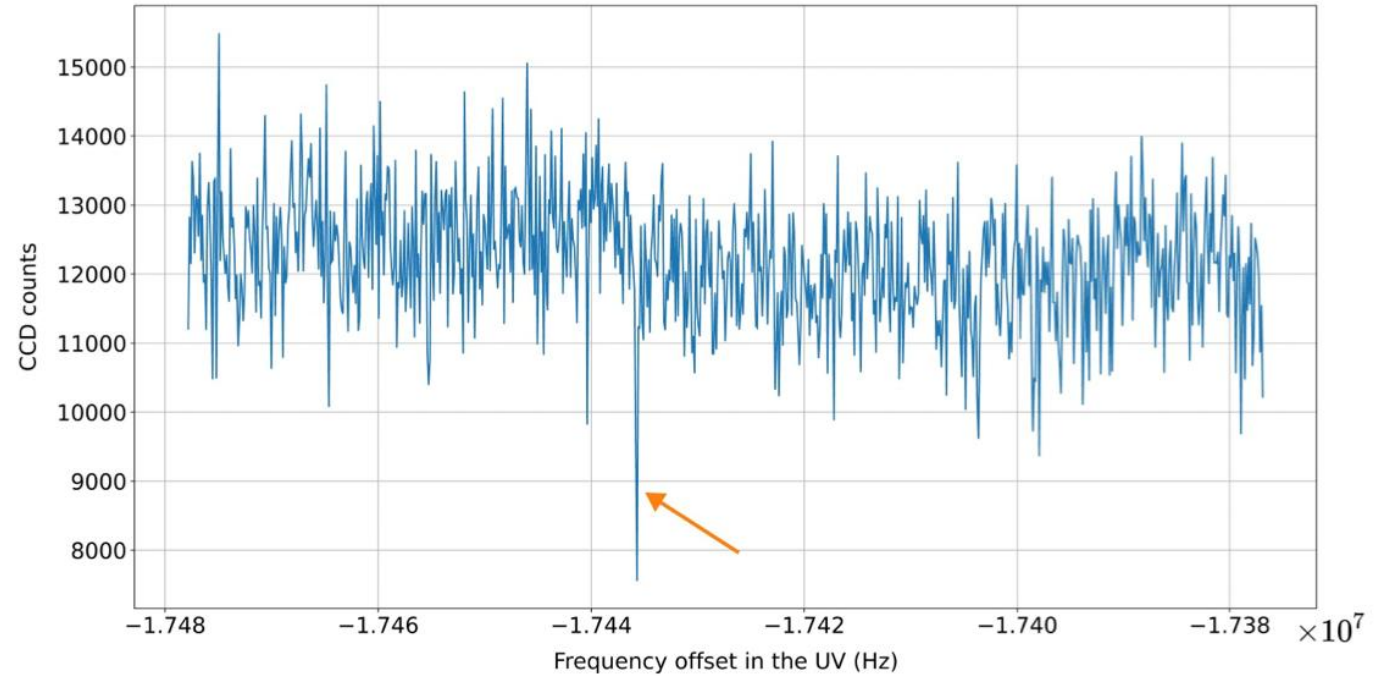
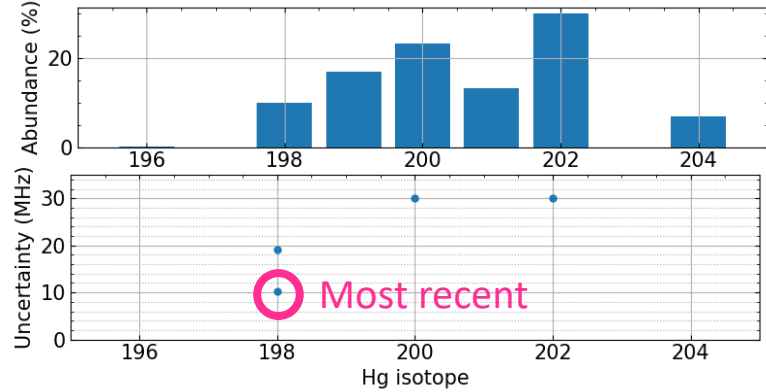


Probing the transition: Transition width of \sim Hz over ± 10 MHz range!

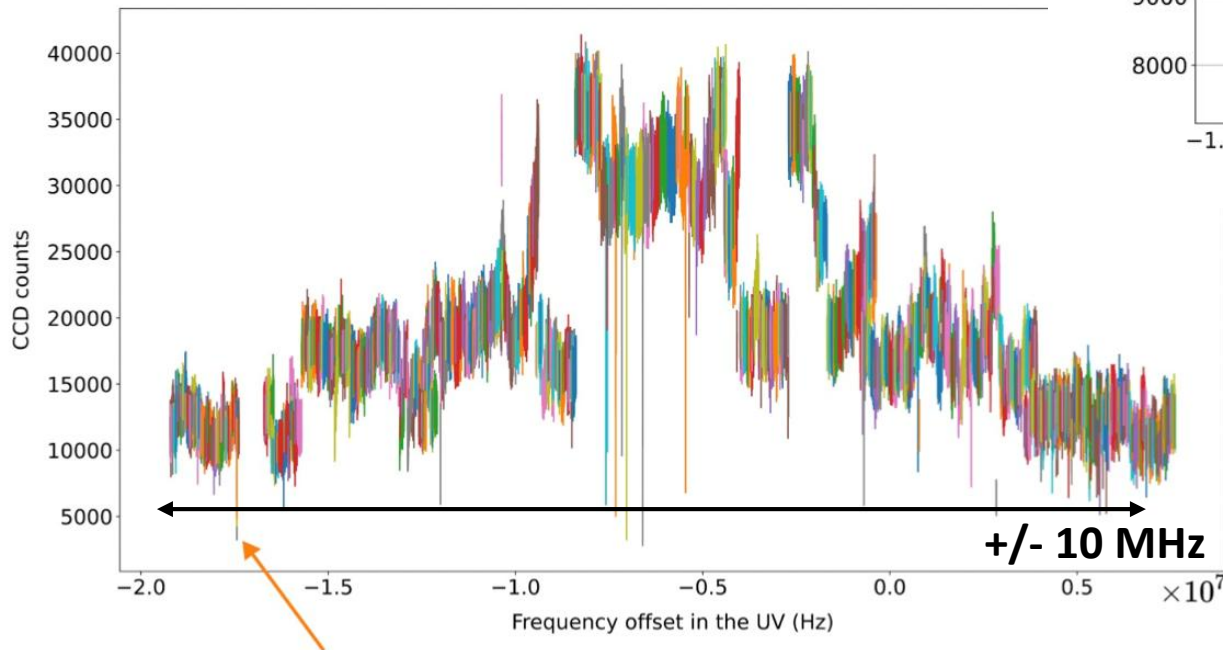


198Hg bosonic isotope

Isotope choice:



Probing the transition:



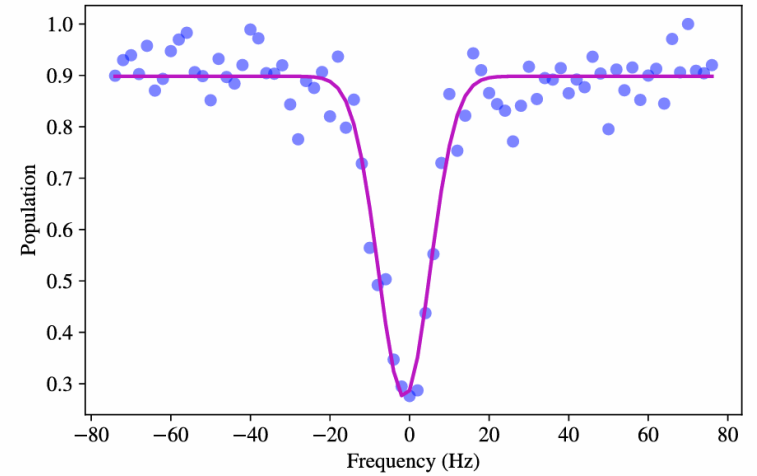
Transition found in 2024! [1]

198Hg bosonic isotope clock

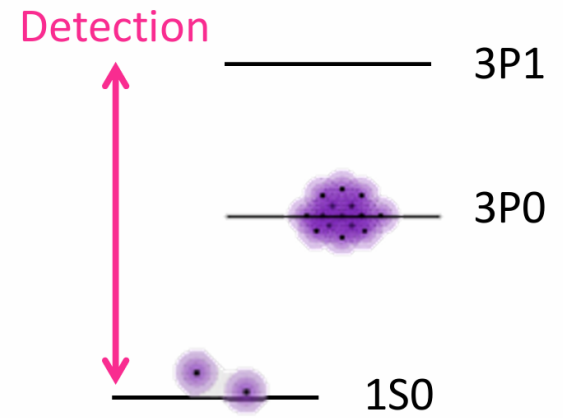
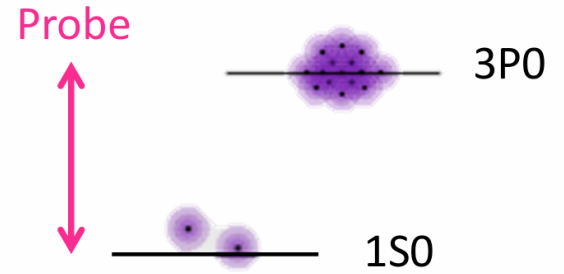
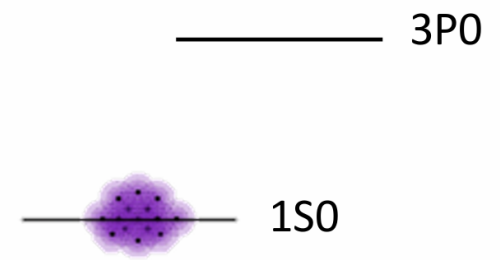


Lattice:
Trapping depth $\sim 92.4 E_r$
Waist $\sim 69 \mu\text{m}$

Ultrastable laser:
Linewidth $\sim 200 \text{ mHz}$
Mean drift rate $\sim -16.9 \text{ mHz/s}$



2D MOT
3D MOT



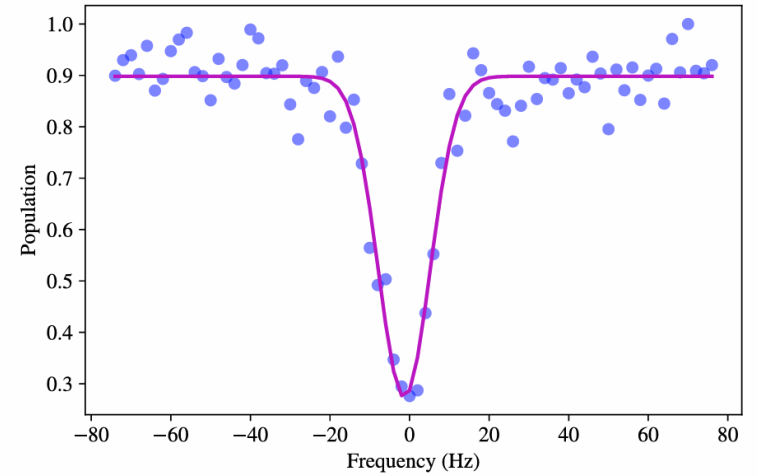
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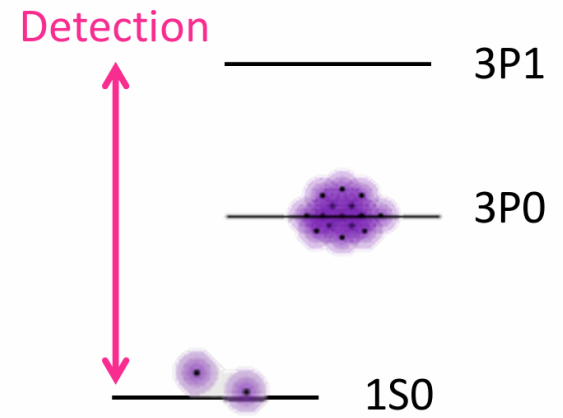
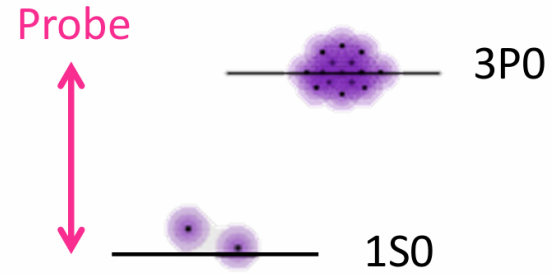
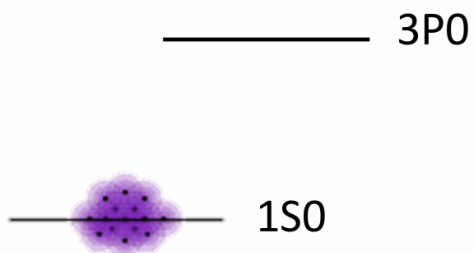
How to get a reliable clock?

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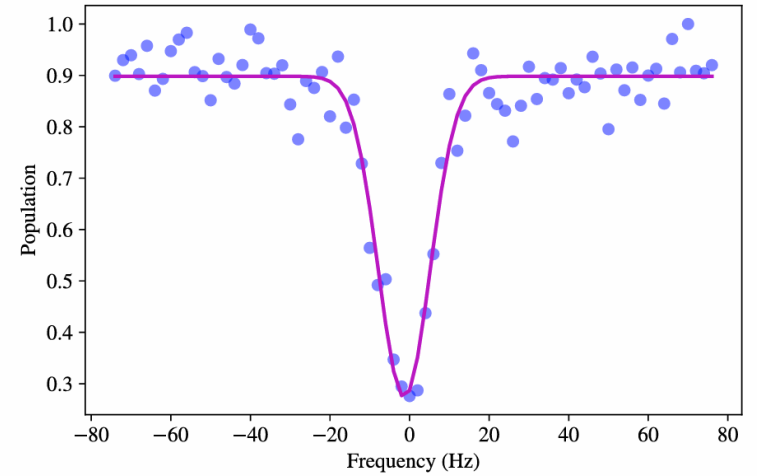
198Hg bosonic isotope clock

How to get a reliable clock?

- ❑ High stability → Statistics
- ❑ High accuracy → Systematics

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Ultrastable laser:
Linewidth ~ 200 mHz
Mean drift rate ~ -16.9 mHz/s



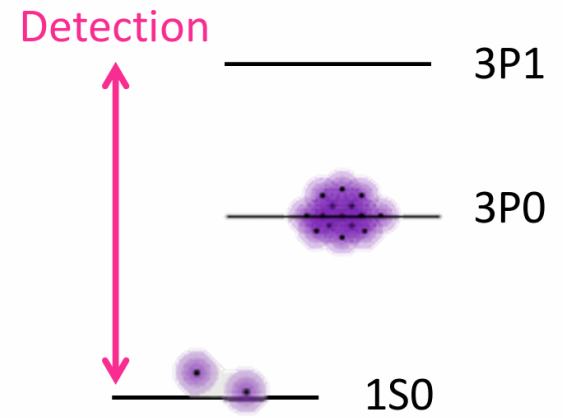
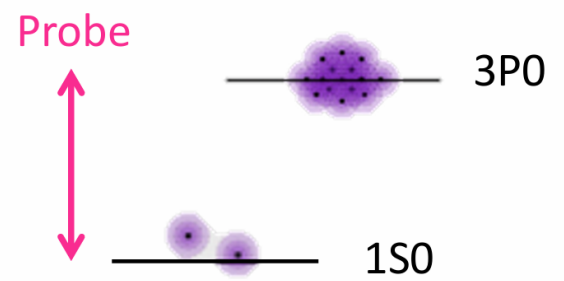
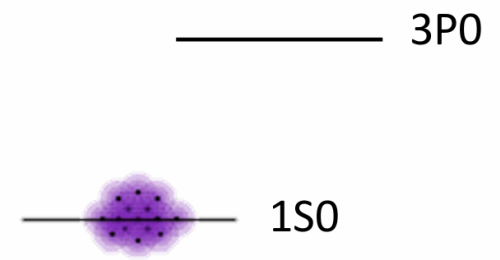
Cooling

Lattice trapping

Clock interrogation

Detection

2D MOT
3D MOT



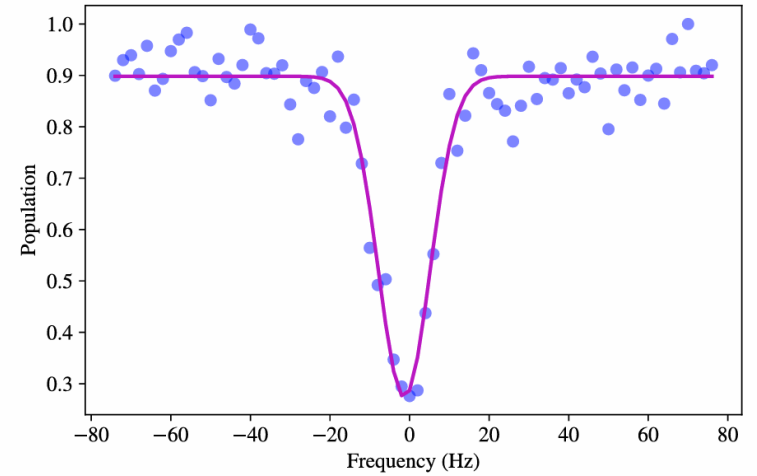
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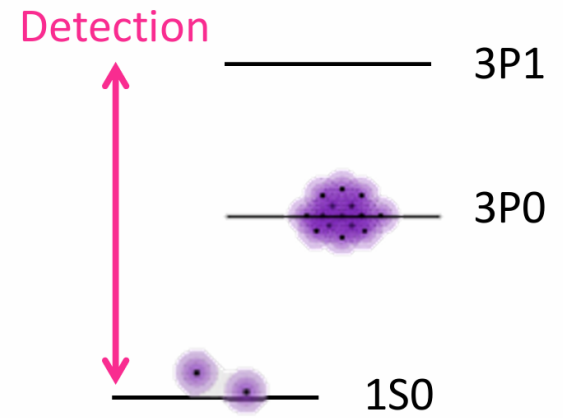
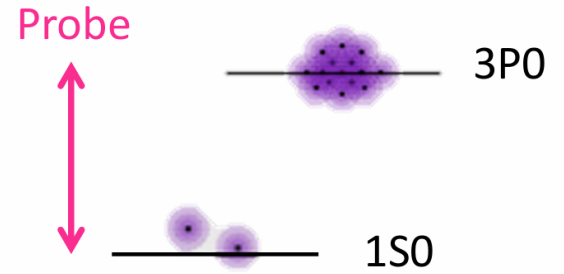
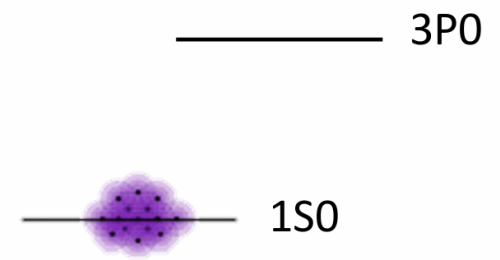
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2D MOT
3D MOT



Lattice linear lightshift

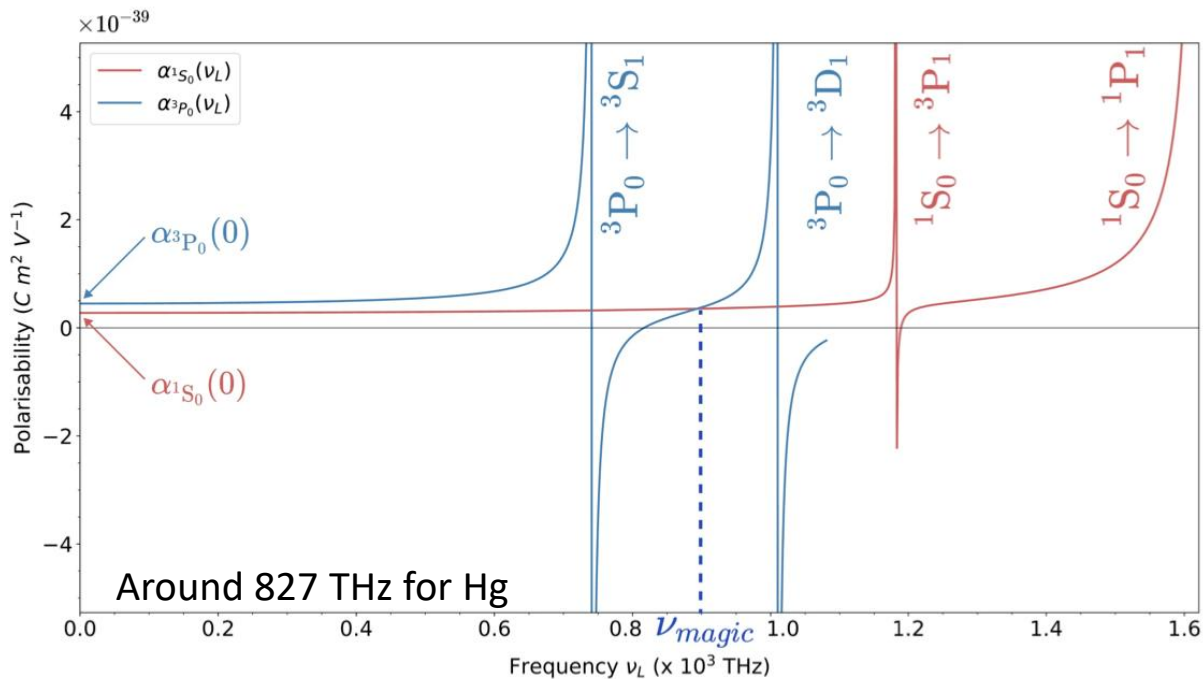
- Total shift on atomic clock transition: $\Delta\nu = \left(\alpha_{1S_0}(\nu) - \alpha_{3P_0}(\nu) \right) \frac{|E_L|^2}{8\pi\hbar}$

198Hg systematics characterization



Lattice linear lightshift

- Total shift on atomic clock transition: $\Delta\nu = \left(\alpha_{1S_0}(\nu) - \alpha_{3P_0}(\nu) \right) \frac{|E_L|^2}{8\pi\hbar}$
- Magic Wavelength: $\alpha_{1S_0}(\nu) = \alpha_{3P_0}(\nu)$



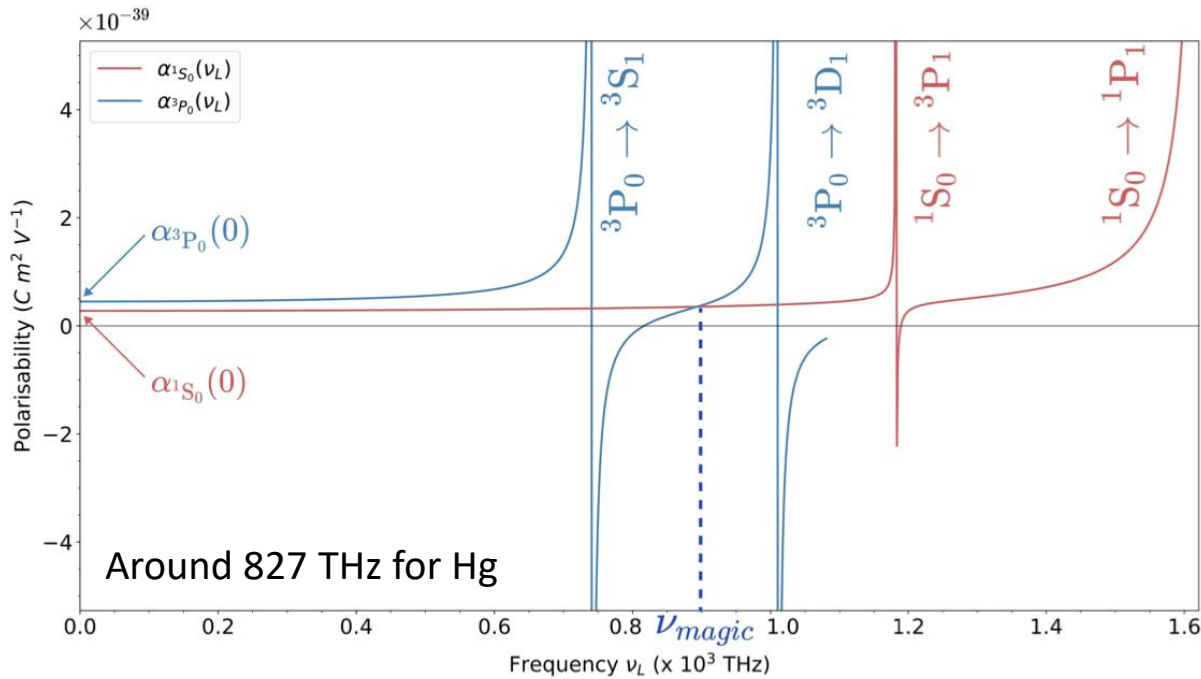
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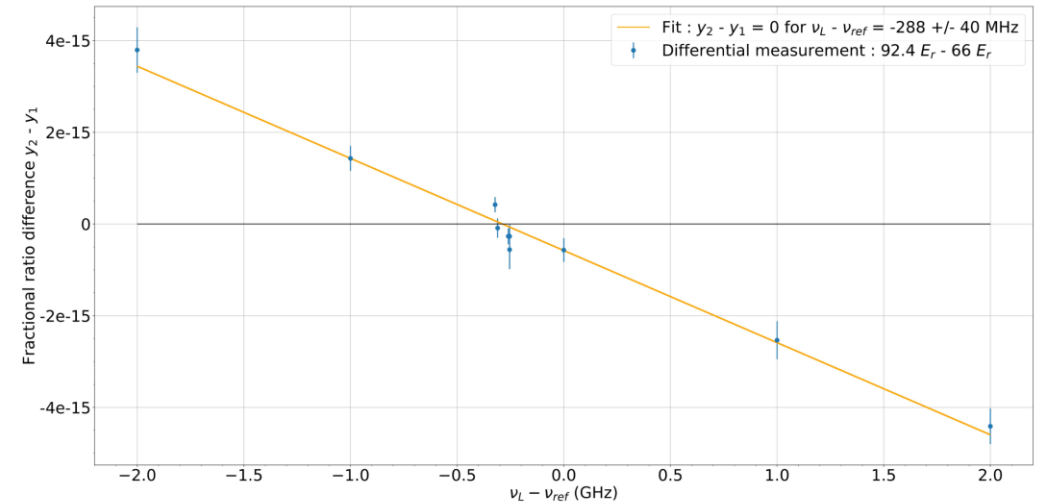
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Magic Wavelength calibration

$$\Delta_{\text{magic}} = K\nu_L(\nu_L - \nu_{\text{magic}})U_L$$

- Differential measurements using different lattice depths



$$\nu_{\text{magic}} = 826.855\,228(28)\text{ THz}$$

$$\frac{\sigma\Delta}{\nu_0} = \pm 3.2 \times 10^{-16}$$

Systematics calibration of clock transition

Cold collisional shift

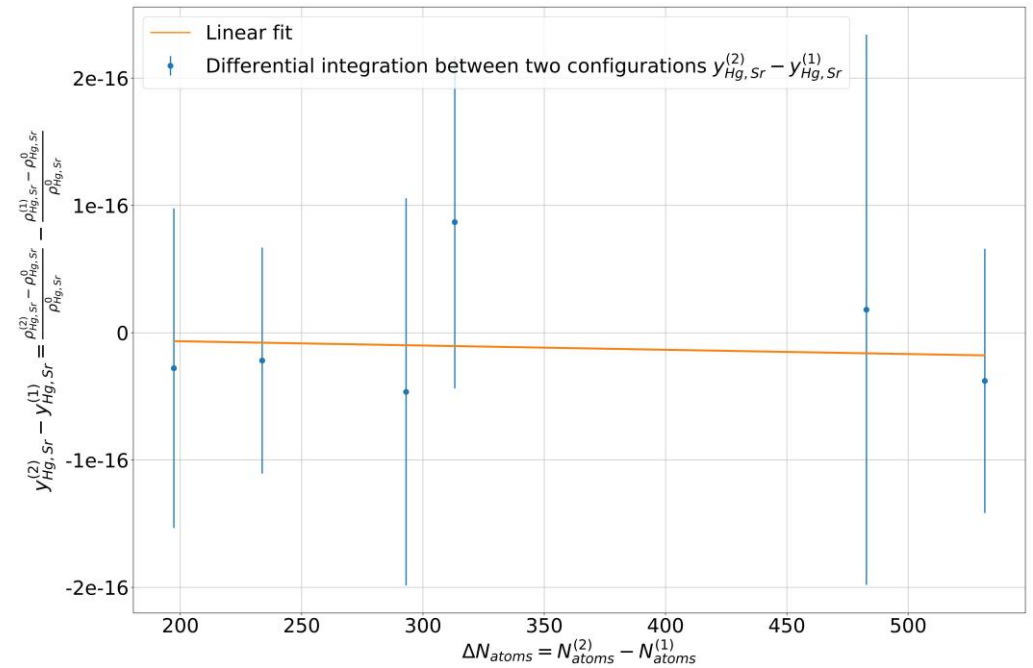
- ❑ Induced by interactions between cold atoms

$$\Delta_{\text{CCS}} = K_{\text{CCS}} N_{\text{atom}}$$

- ❑ Differential measurements with different atom number

$$\frac{\sigma_{\Delta}}{\nu_0} = \pm 1.4 \times 10^{-16} / 1000 \text{ atoms}$$

Cold collisional shift calibration



Systematics calibration of clock transition

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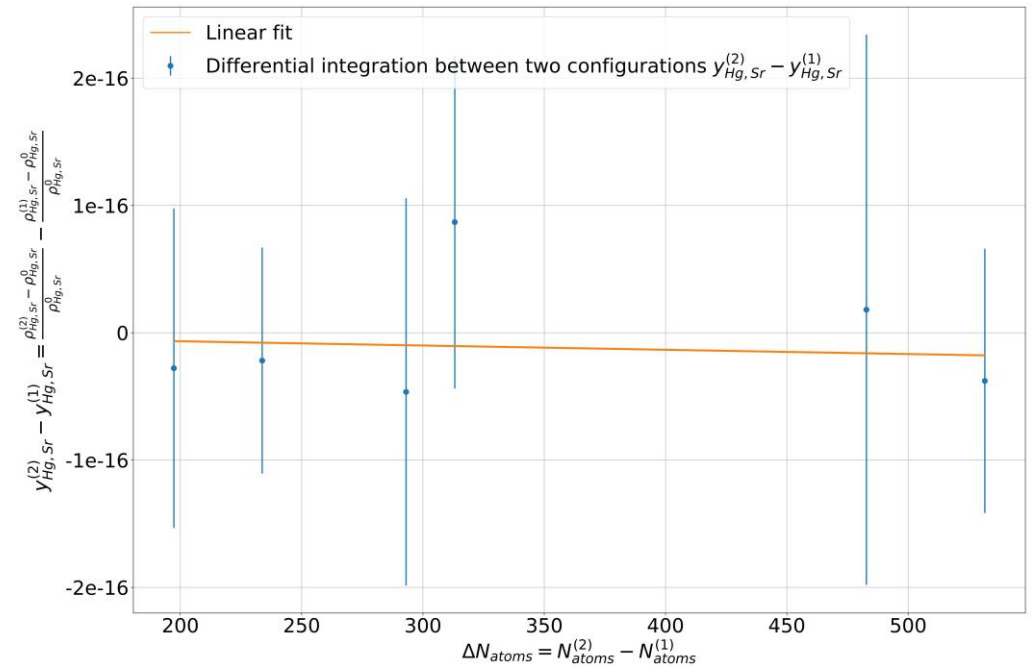
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Our experiment: ~ 500-1000 atoms

Cold collisional shift calibration

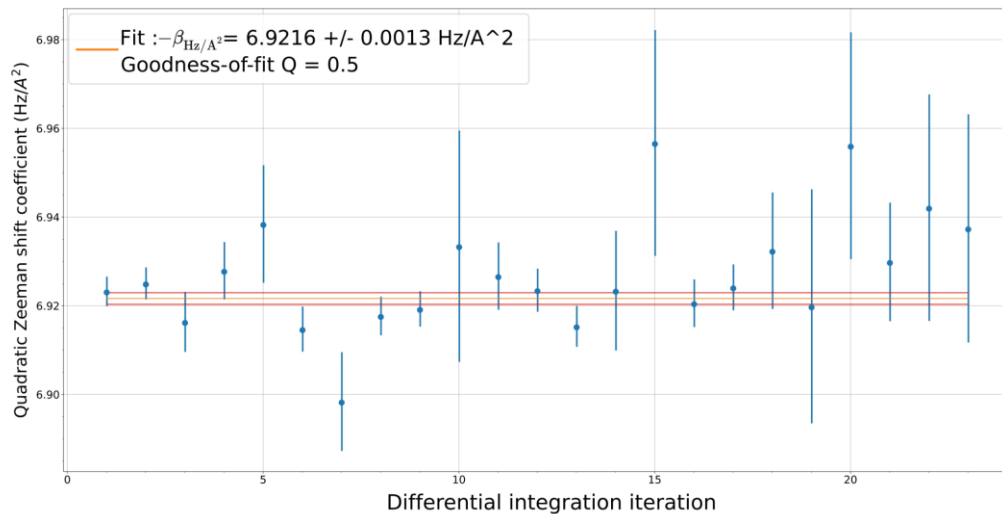


Systematics calibration of clock transition

Quadratic Zeeman Shift (QZS)

$$\square \Delta_{QZS} = \beta \left\| \vec{B} \right\|^2$$

- Differential measurements with different values of B



$$\beta_{exp} = -2\pi \cdot 3.0 \pm 0.2 \text{ MHz/T}^2$$

198Hg systematics characterization

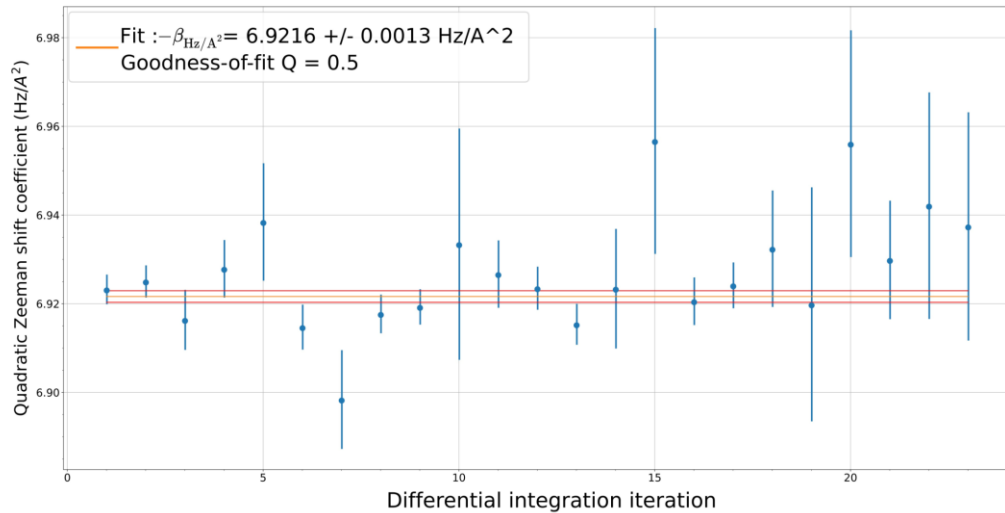


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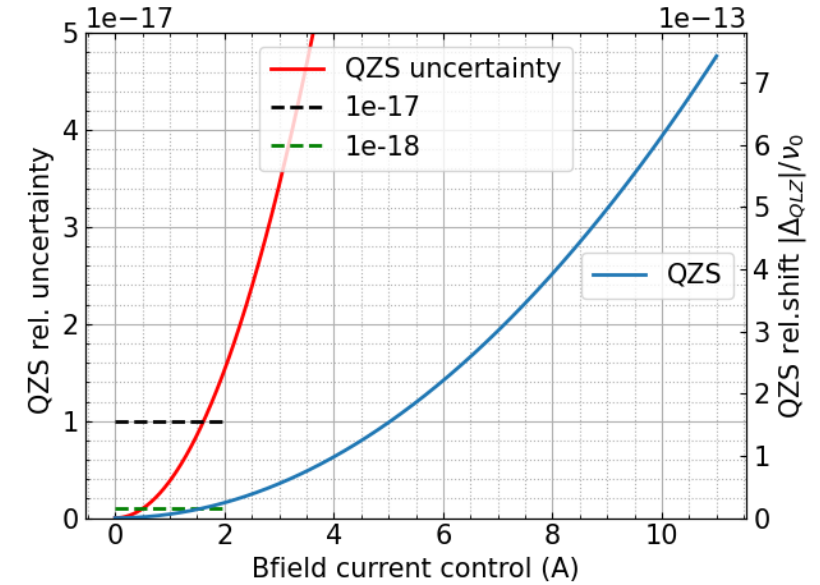
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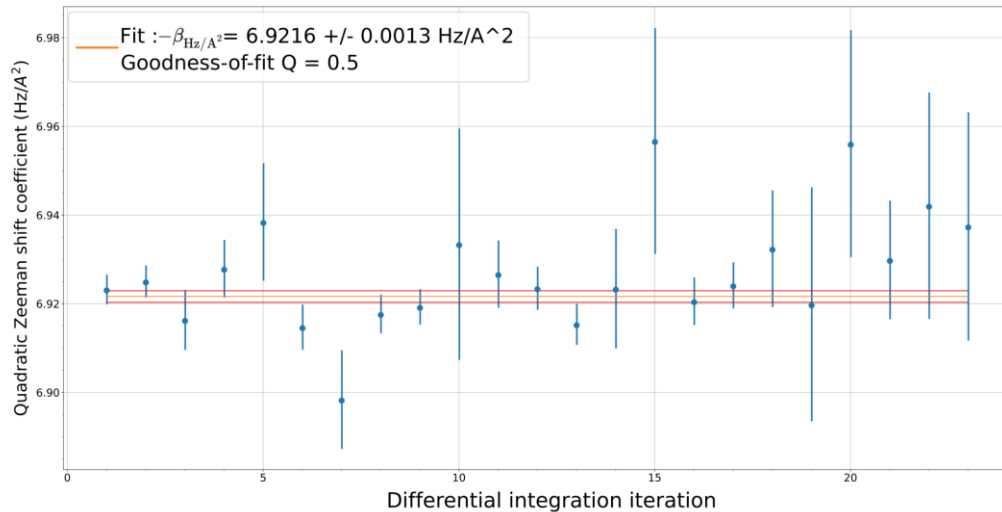


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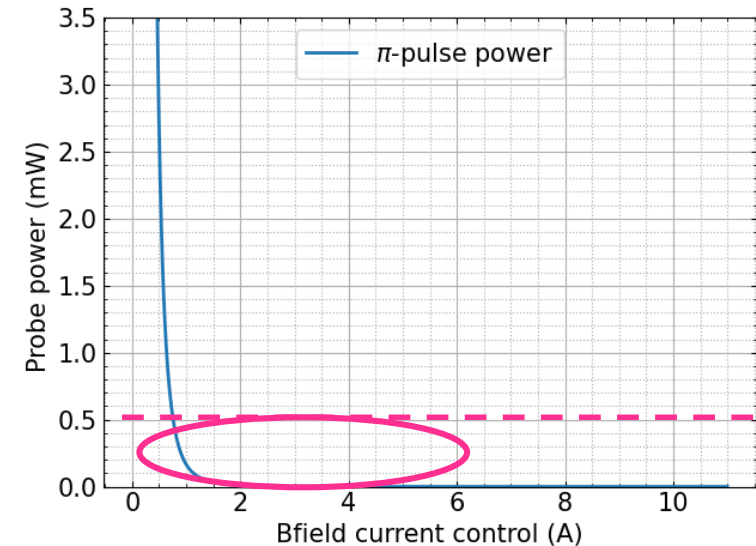
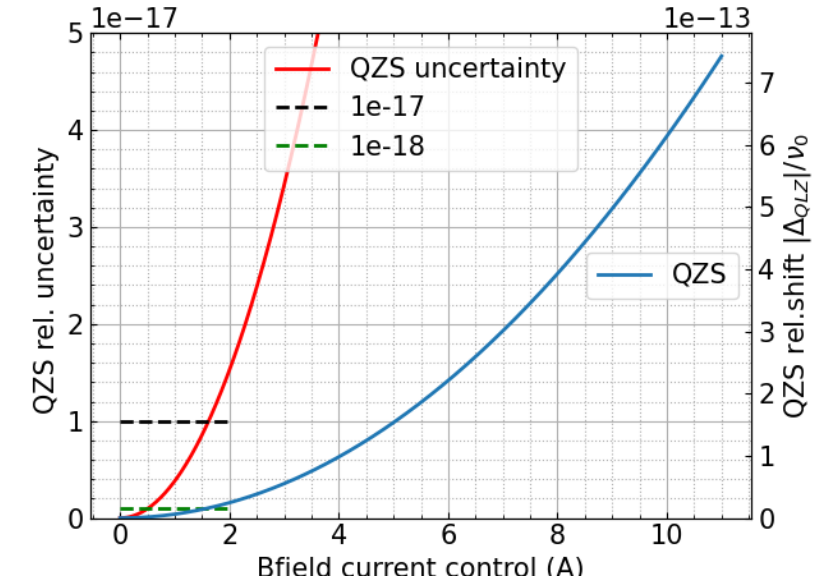
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maximum power $\sim 500 \mu\text{W}$

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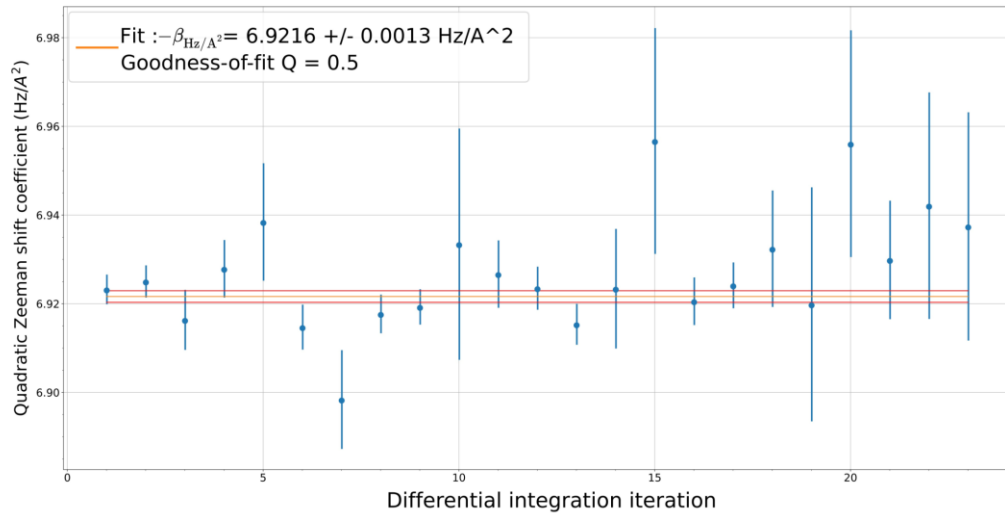


Systematics calibration of clock transition

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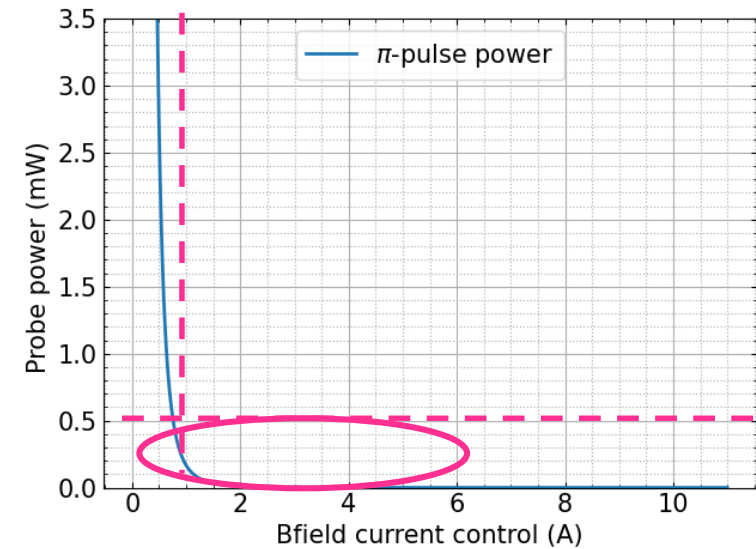
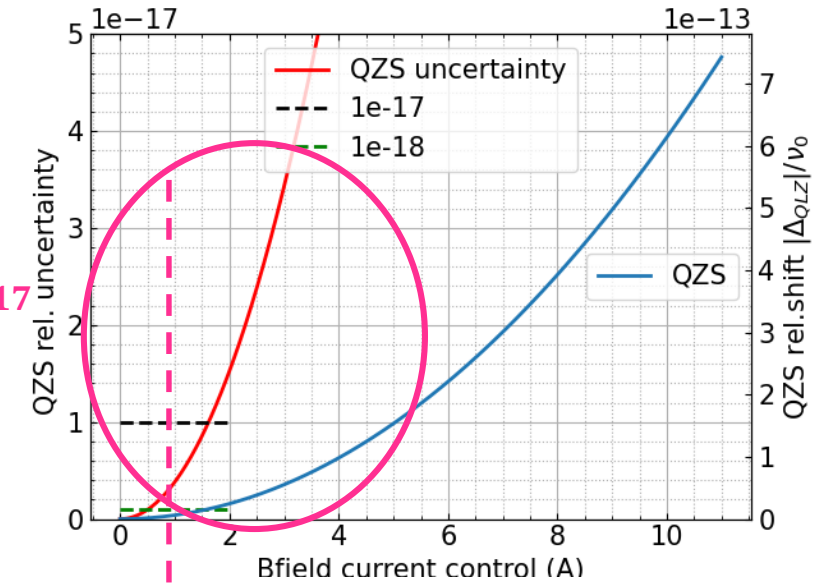
$$\Delta_{QZS} = \beta \|\vec{B}\|^2$$

- Differential measurements with different values of B



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Systematics calibration of clock transition

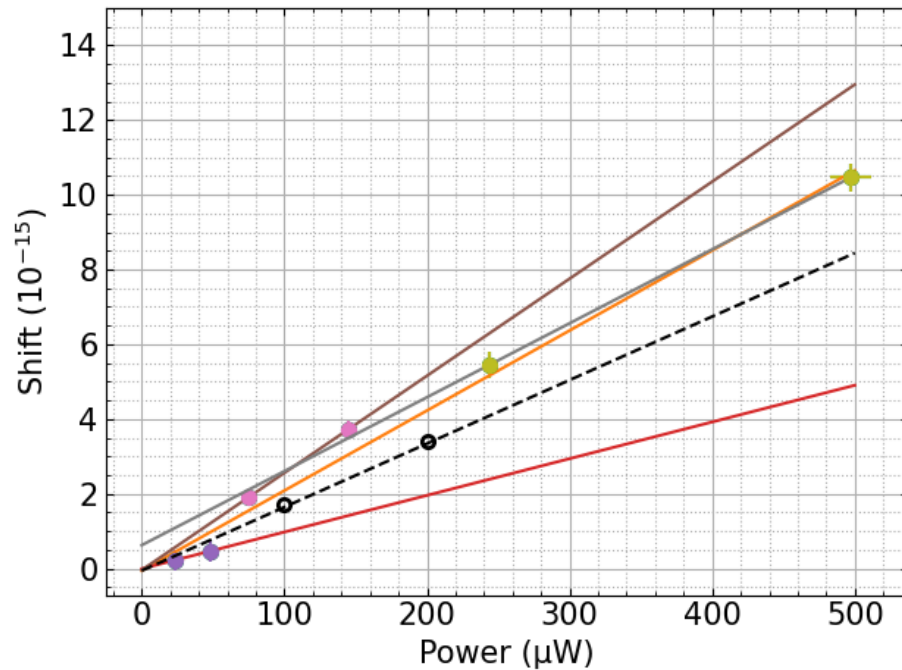
Probe Light Shift (PLS)

$$\square \Delta_{PLS} = \alpha P$$

Systematics calibration of clock transition

Probe Light Shift (PLS)

- $\Delta_{PLS} = \alpha P$
- Differential measurements at different power

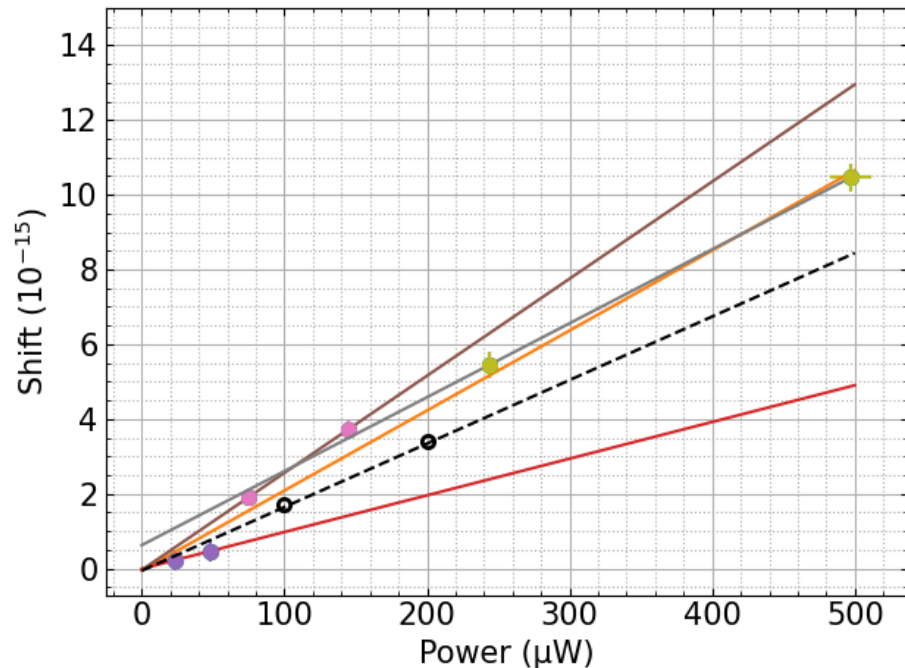


Systematics calibration of clock transition

Probe Light Shift (PLS)

□ $\Delta_{PLS} = \alpha P$

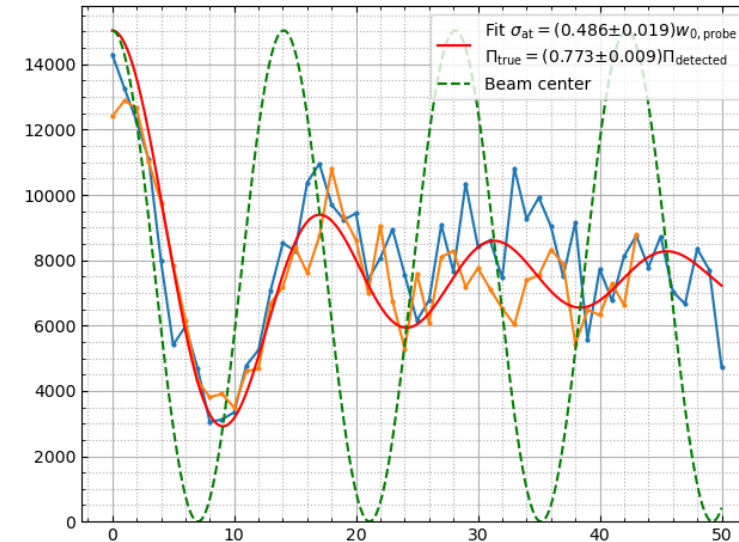
□ Differential measurements at different power



Maximum power $\sim 500 \mu\text{W}$:

- Waist of $\sim 63.5 \mu\text{m}$
- $\sigma_{\text{at}} \sim 50 \mu\text{m}$

➔ Strong spatial inhomogeneities



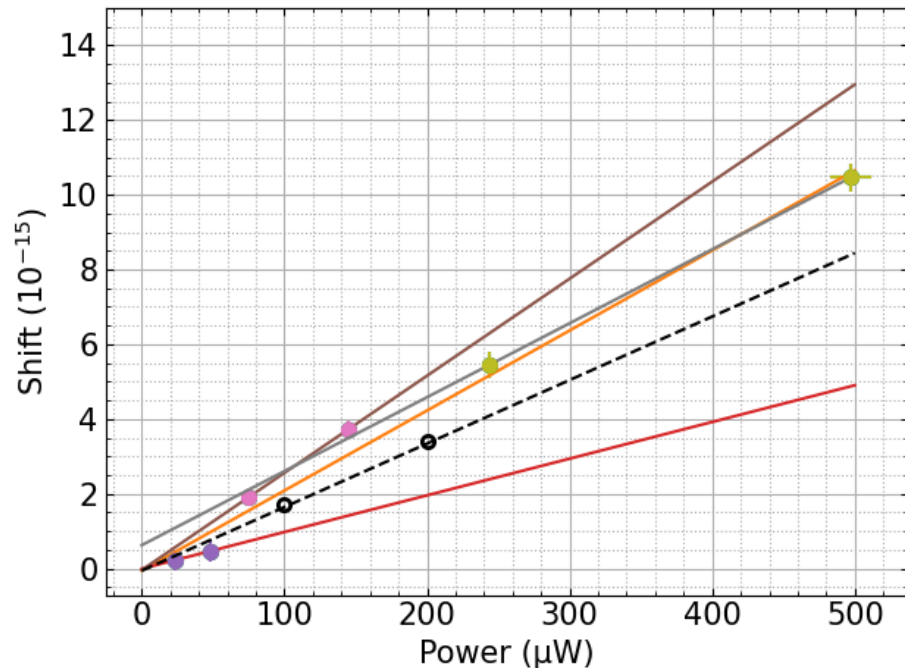
198Hg systematics characterization

Systematics calibration of clock transition

Probe Light Shift (PLS)

□ $\Delta_{PLS} = \alpha P$

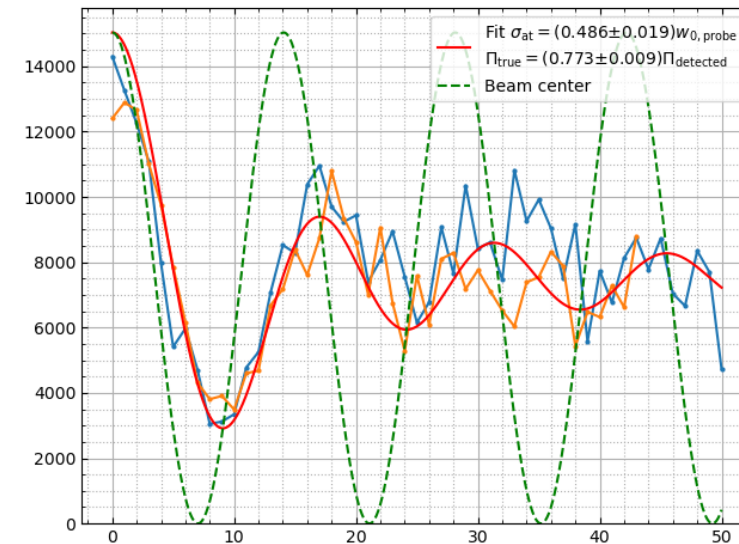
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➔ Biases on PLS estimation

$$\frac{\Delta_{\nu}}{\nu_0} = \pm 5.8 \times 10^{-18} / \mu\text{W}$$

198Hg / 87Sr Comparison



Zyskind et al., arXiv:2512.04920

2 month local comparison campaign:

- ❑ Total relative systematic uncertainty: 6.9×10^{-16}
- ❑ Statistics $< 3 \times 10^{-17}$
- ❑ ^{87}Sr : 4.6×10^{-17}

$$\frac{\nu_{198\text{Hg}}}{\nu_{87\text{Sr}}} = 2.629\,315\,734\,684\,118\,1\,(18)$$

$$\nu_{198\text{Hg}} = 1\,128\,575\,945\,288\,666.3 \pm 0.78 \text{ Hz}$$

Effect	Correction ($\times 10^{-17}$)	Uncertainty ($\times 10^{-17}$)
Quadratic Zeeman effect	7868.7	4.9
Magnetic field fluctuations	0	2.9
Residual magnetic field	0	4.3
Cold collisional shift	4.1	17.3
Linear lattice light shift	0	32.3
Nonlinear lattice light shift	2.1	2.1
Blackbody radiation	14.6	1.5
Background gas collisions	3.0	3.0
Probe light shift	-291.8	58.4
Total	7600.6	69.4

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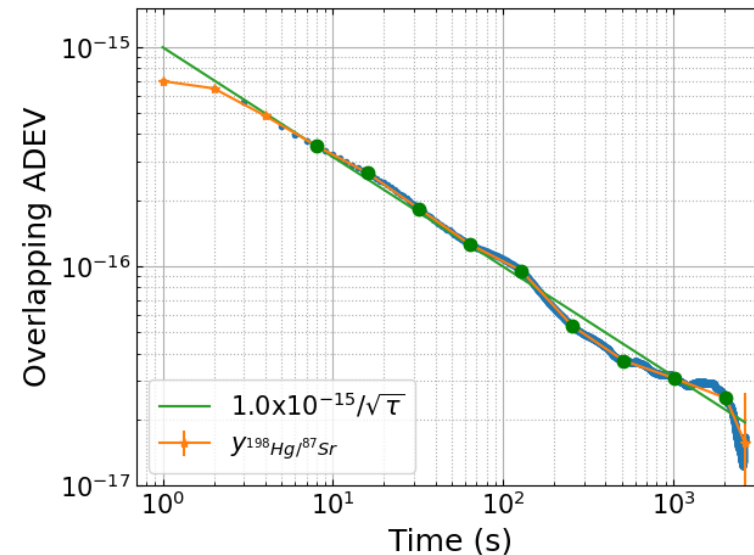
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Stability estimation:

With $\sim 1\text{s}$ cycle/200ms probe:

$$\sigma_{^{198}\text{Hg}}(\tau) \sim 6 - 7 \times 10^{-16} / \sqrt{\tau}$$

- ❑ Already better than fermionic isotope
- ❑ Limited by **Dick contribution**



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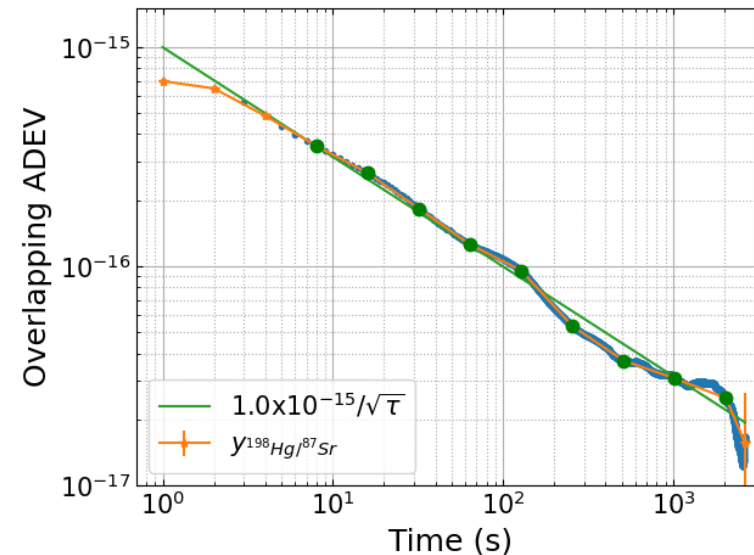
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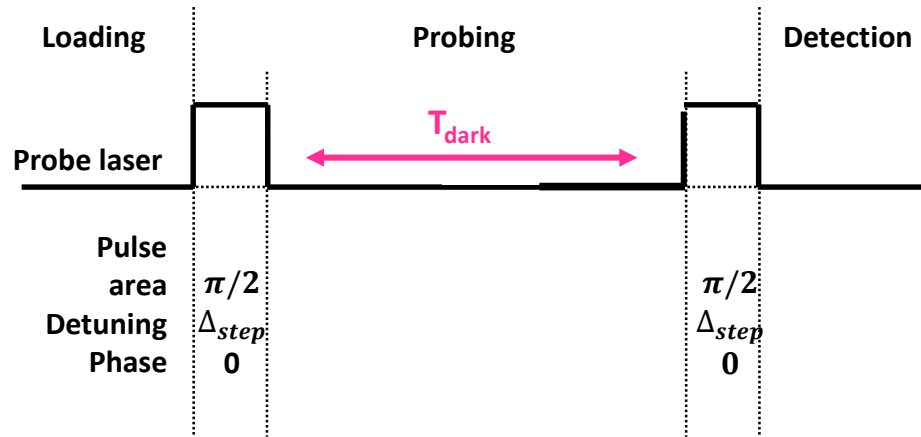
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Hyper-Ramsey Spectroscopy

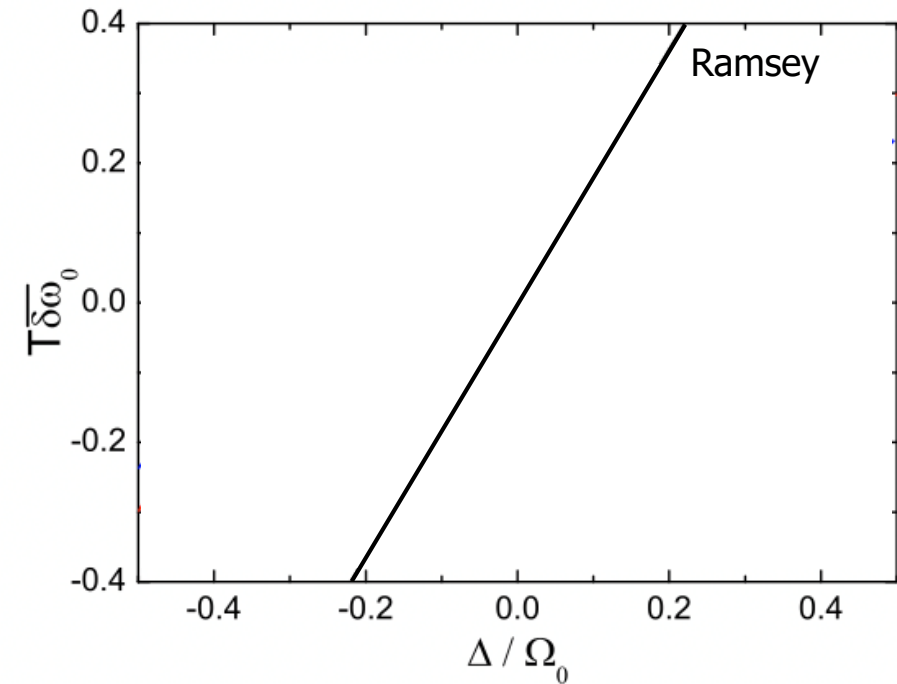
Ramsey Interferometer:



$$\phi = -\delta T_{\text{dark}} + \Delta\tau$$

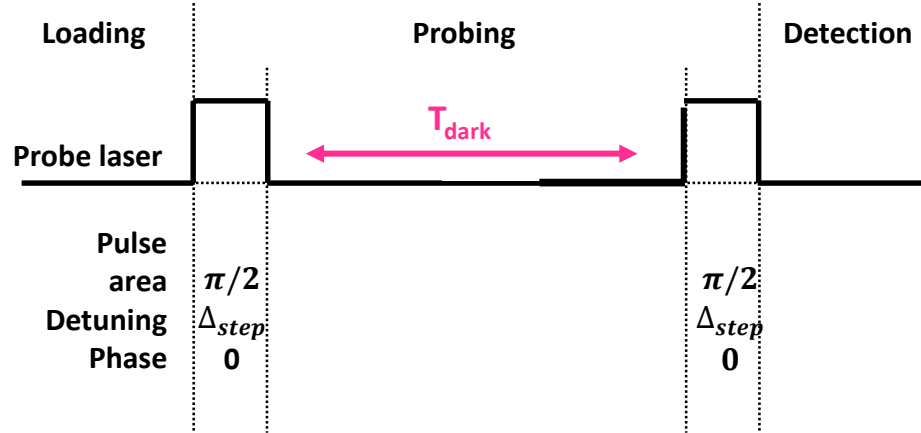
Sensitive to light shift errors:

- Shift $\delta\omega_0$
- Broadening



Hyper-Ramsey Spectroscopy

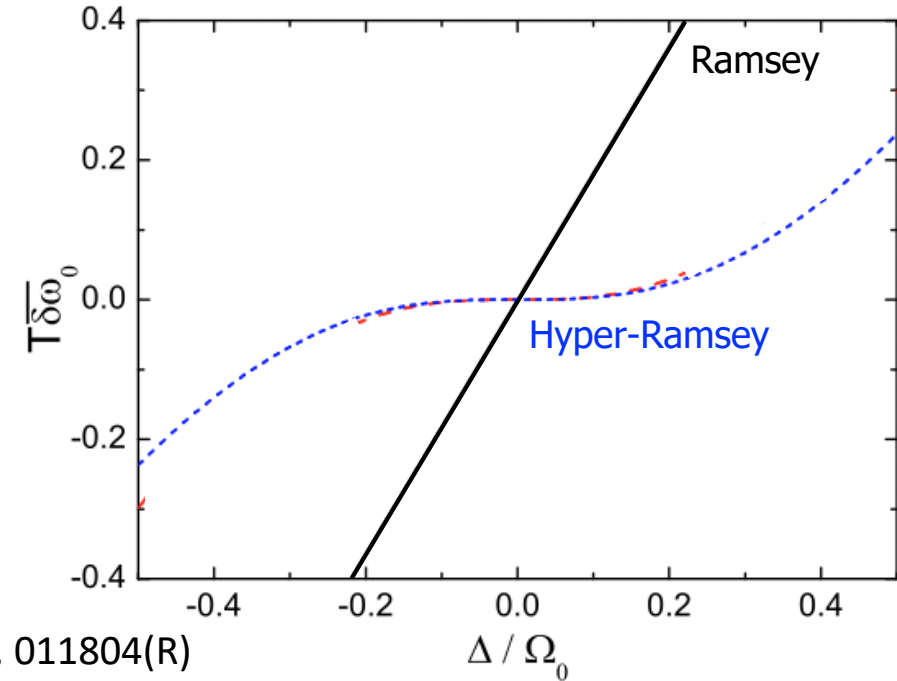
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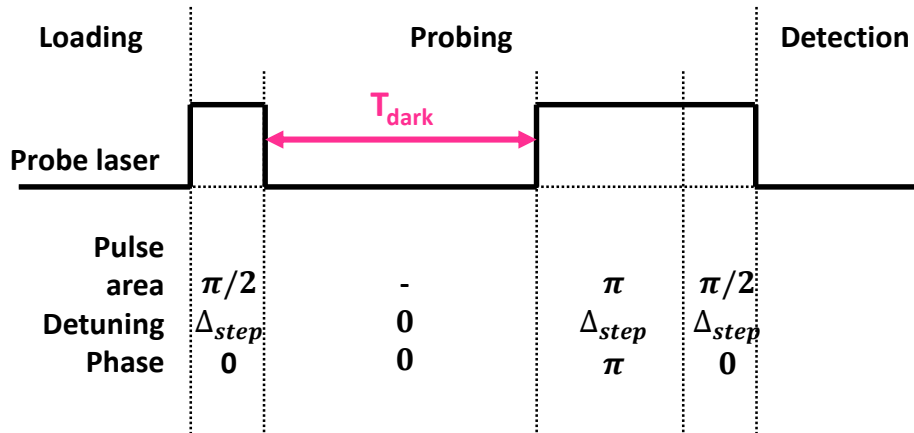
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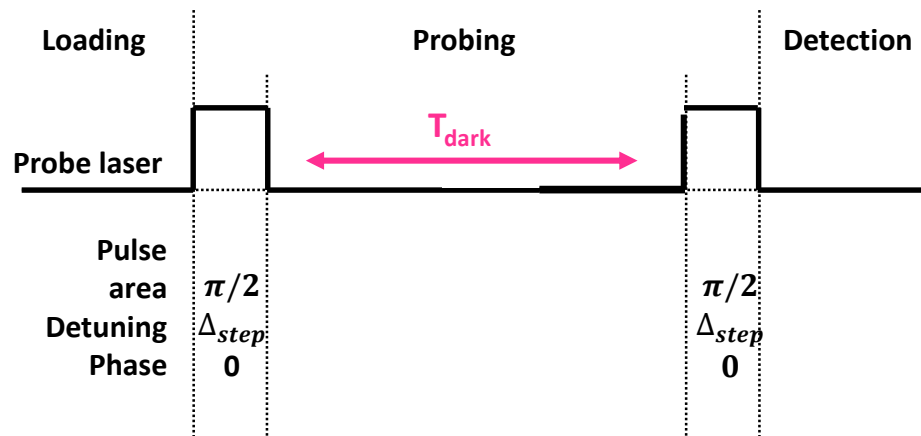
Phys. Rev. A **82**, 011804(R)

Hyper-Ramsey Interferometer:

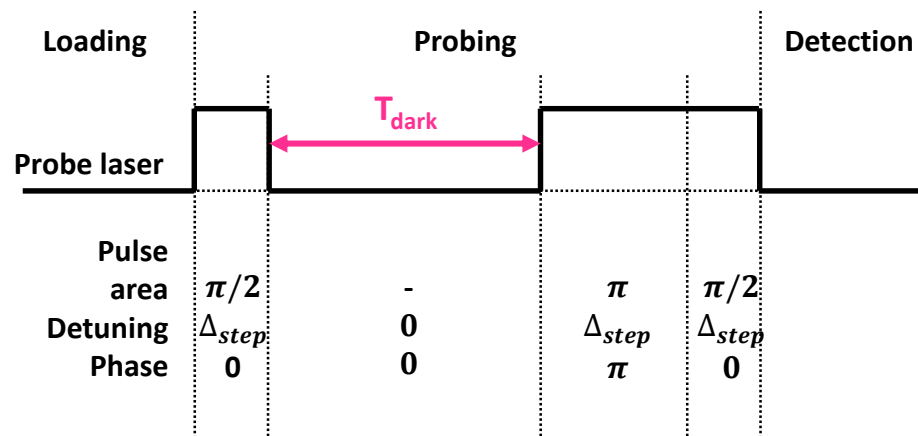


Hyper-Ramsey Spectroscopy

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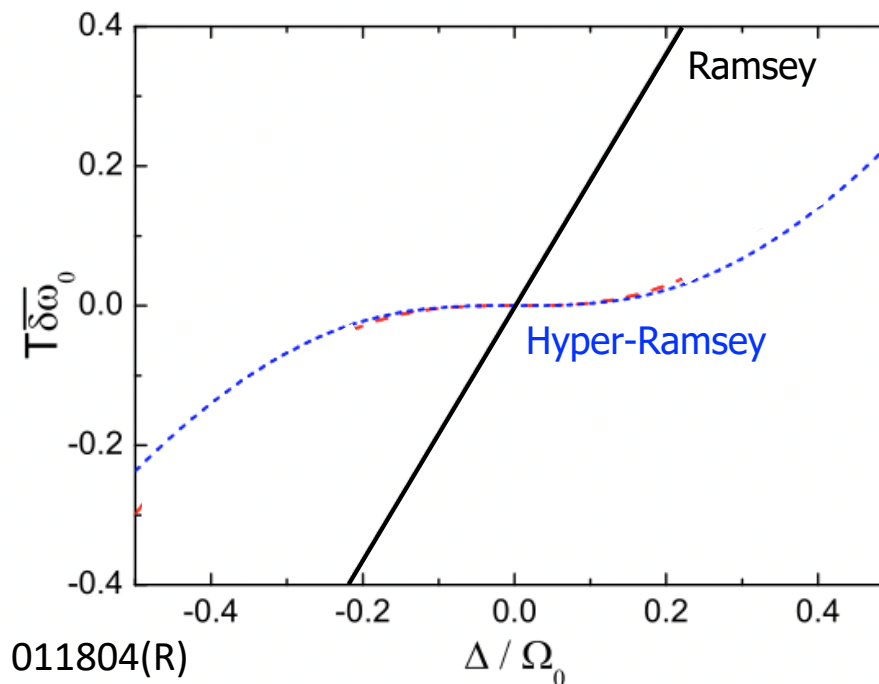
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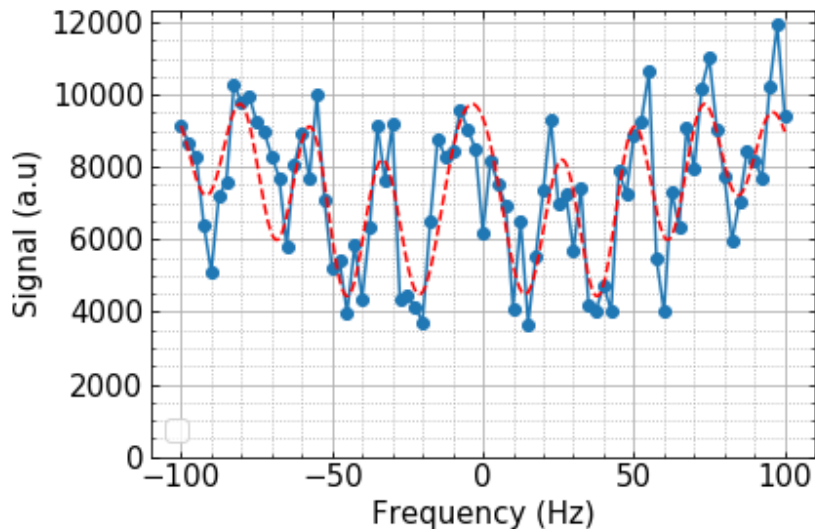
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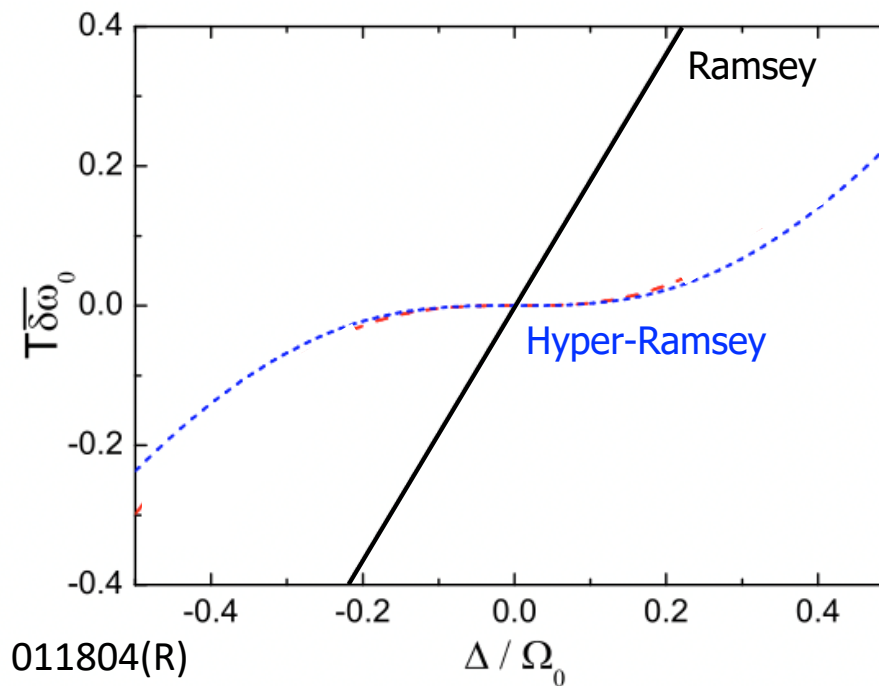
Phys. Rev. A **82**, 011804(R)

- ☐ Expected reduction on PLS $< 1/1000!$
- ☐ Insensitive to pulse area errors (spatial inhomogeneities)

Hyper-Ramsey Spectroscopy

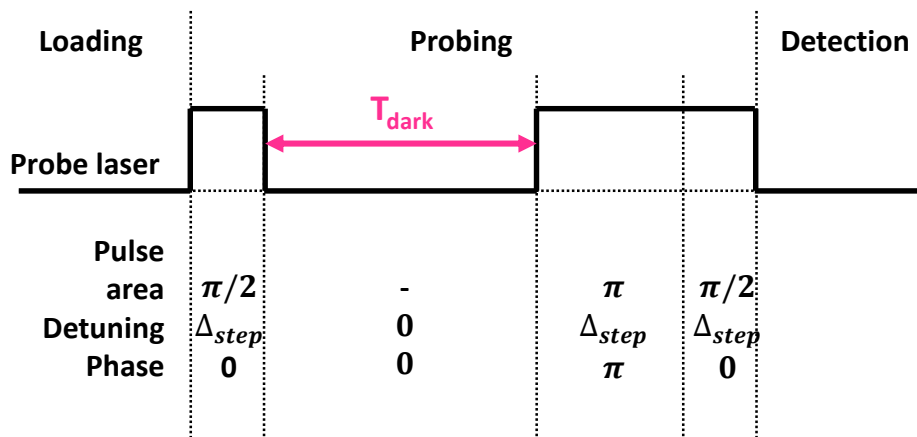


Successful implementation
1st integration tests on-going



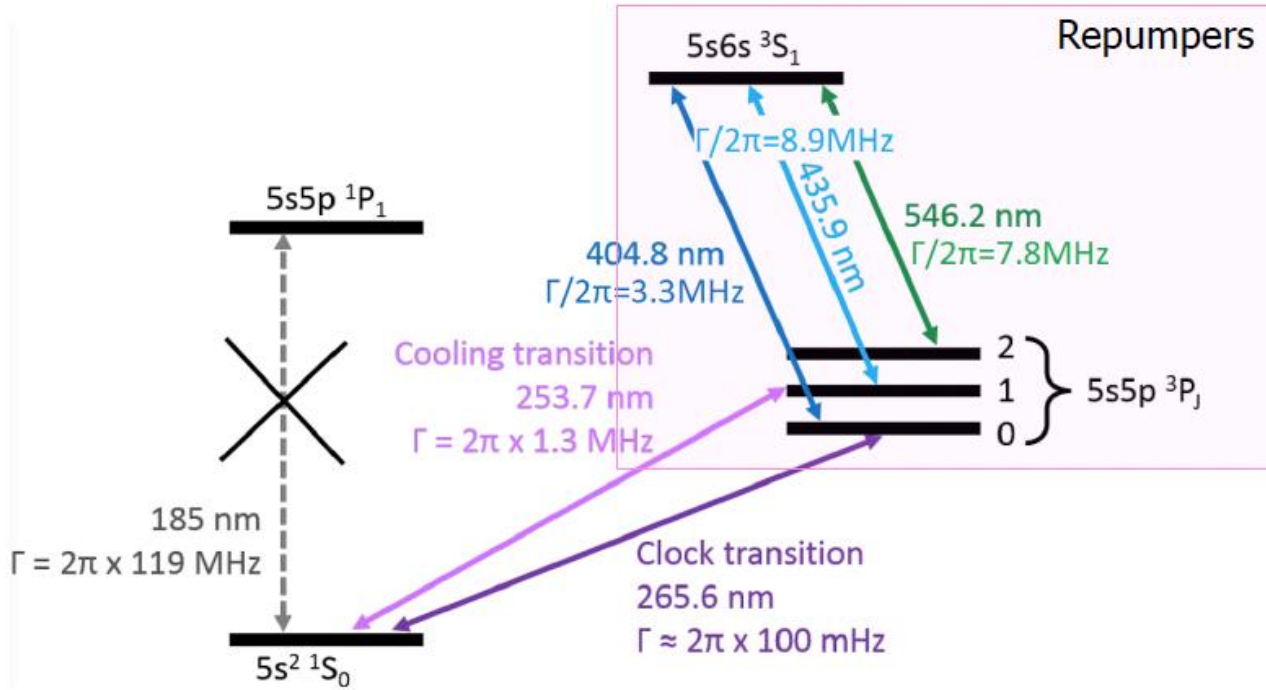
Phys. Rev. A **82**, 011804(R)

Hyper-Ramsey Interferometer:

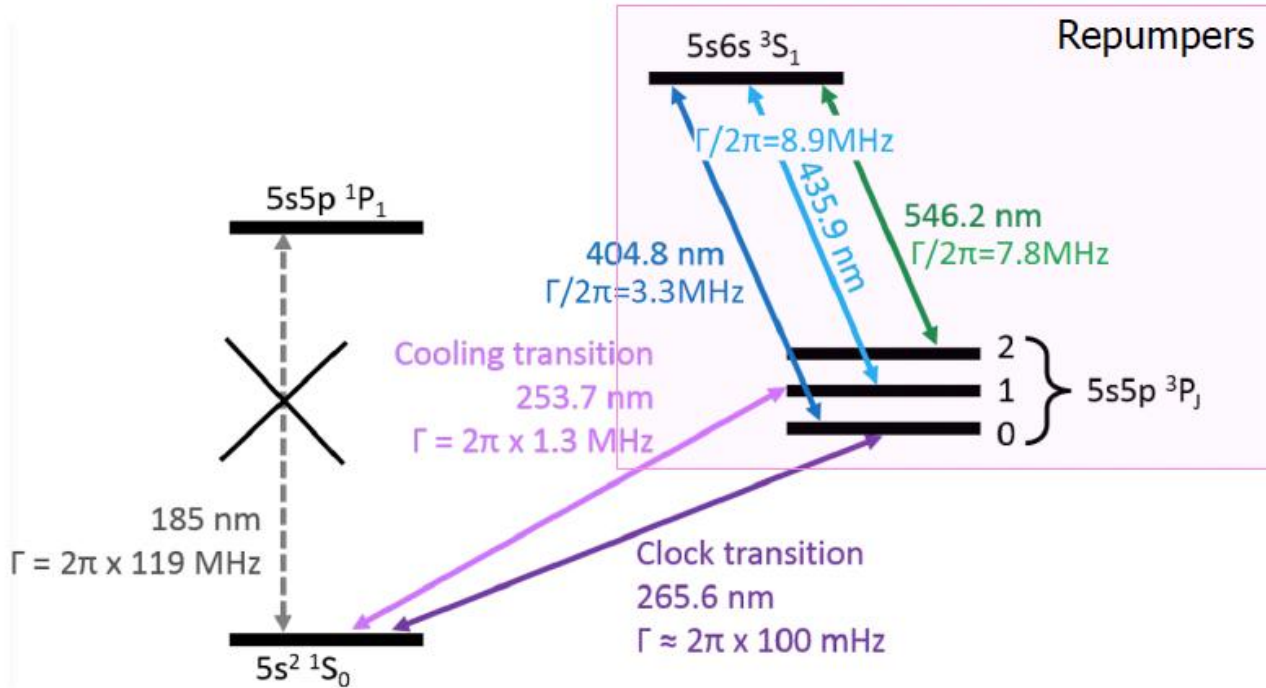


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Using of Repumpers



Using of Repumpers

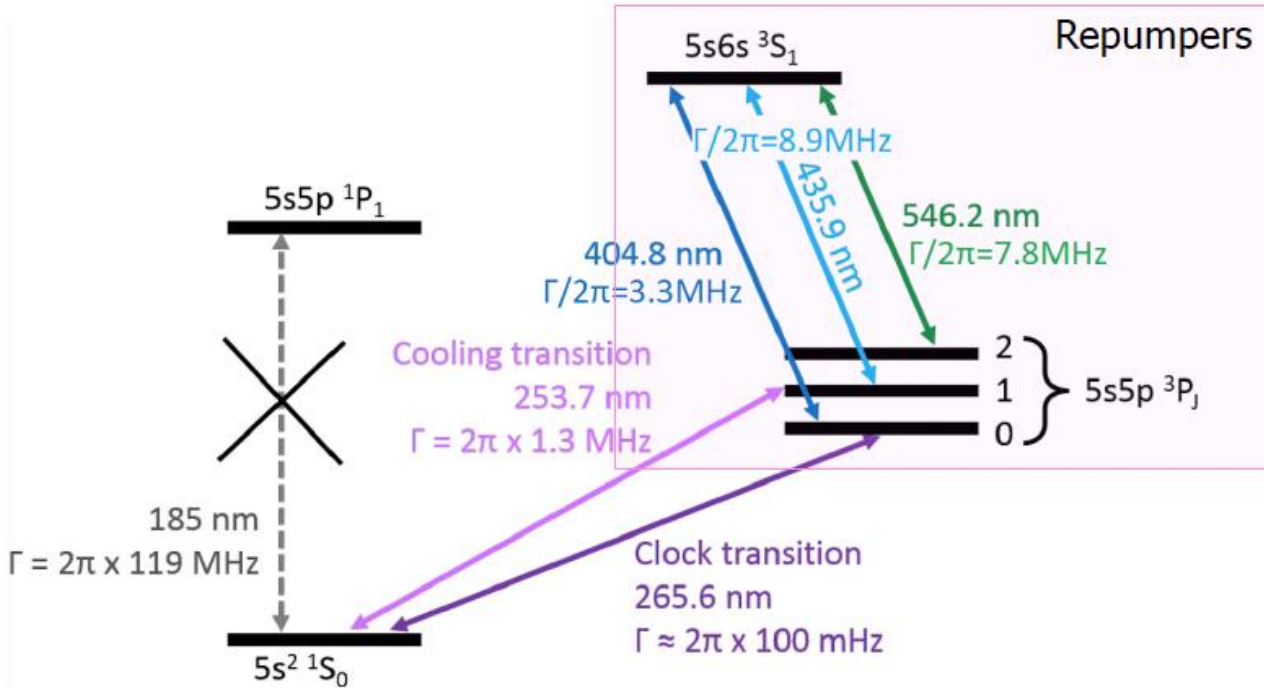


198Hg repumpers:

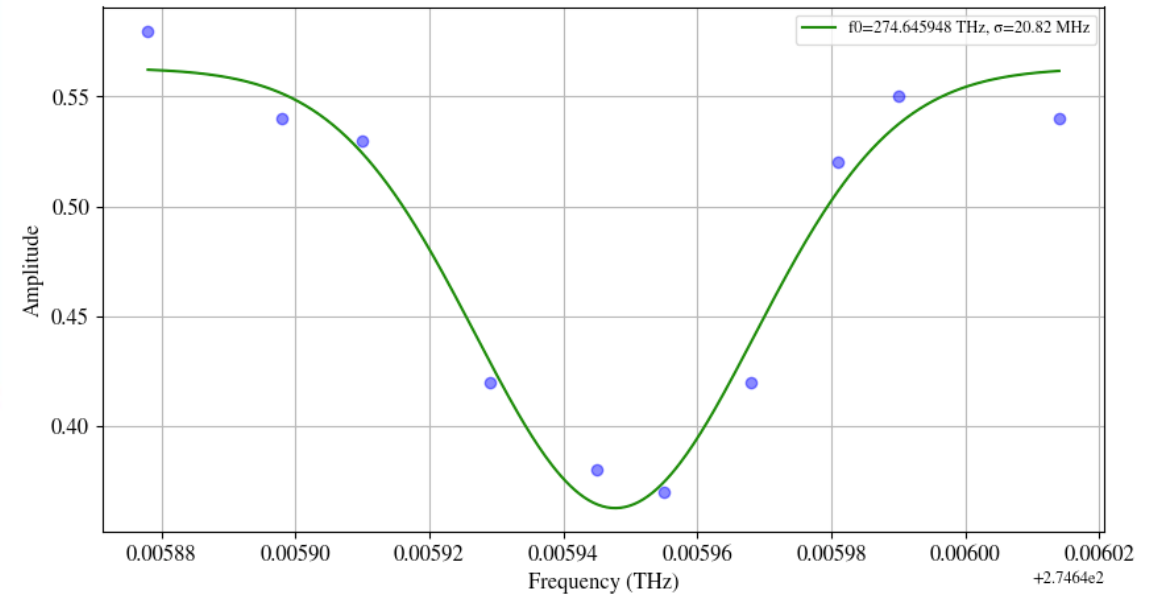
- Sideband cooling
 - Decrease atomic cloud size

- Normalized detection

Using of Repumpers



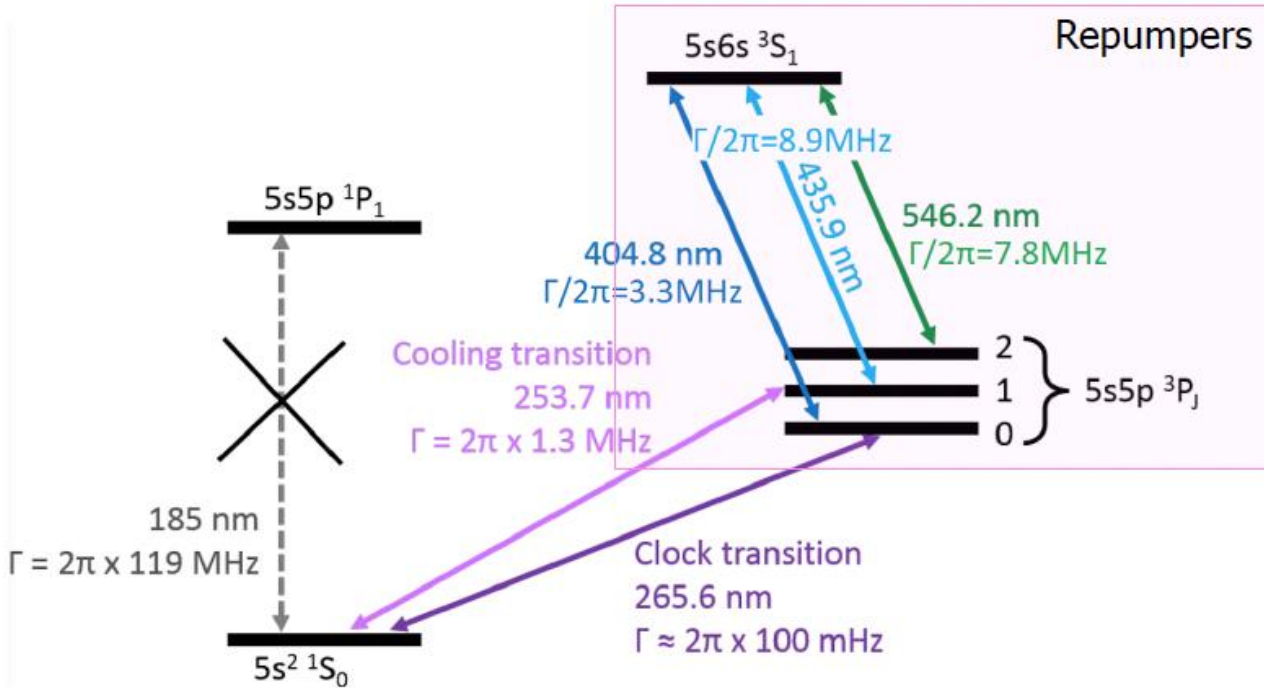
405 transition



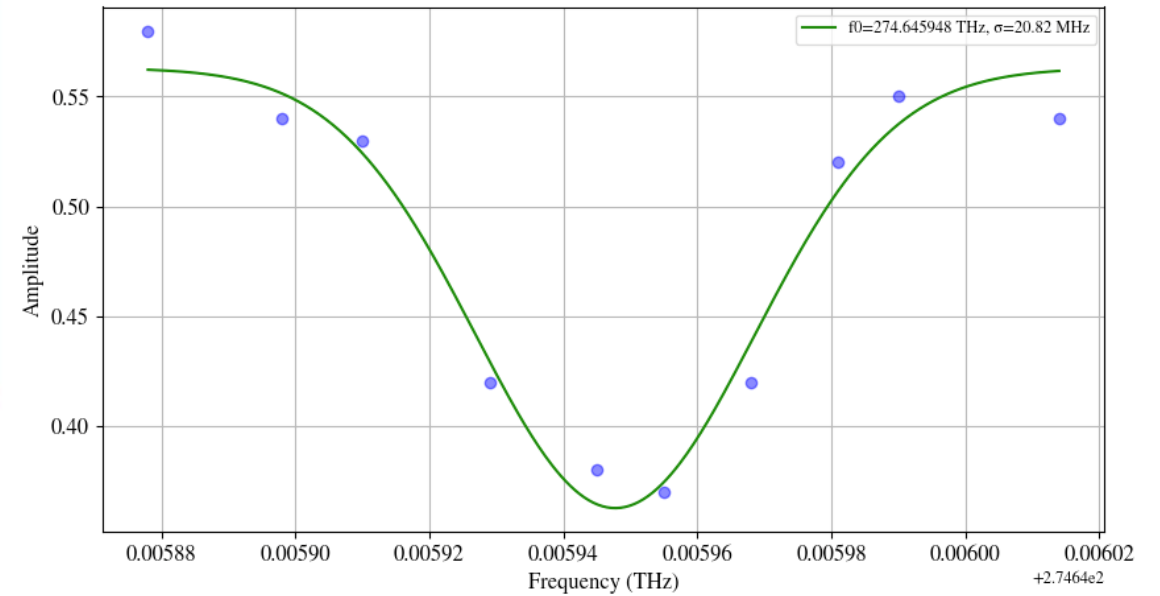
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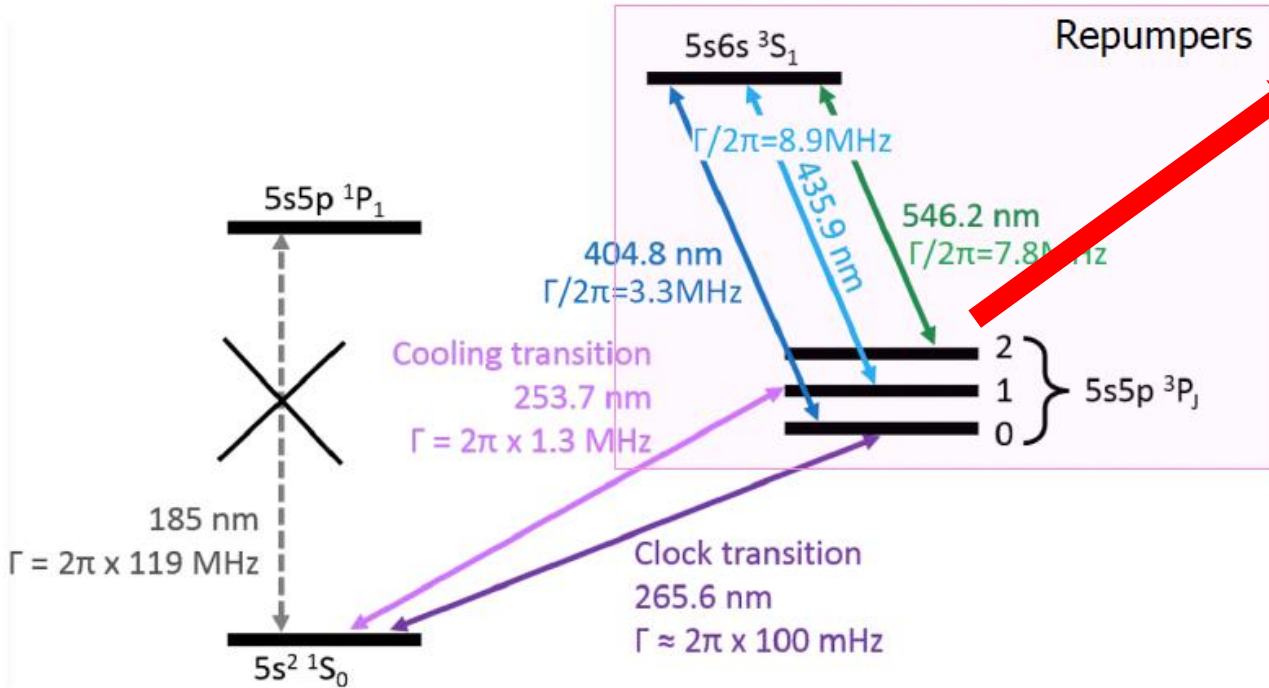
405 transition



No signal of 546 transition so far

- 198Hg repumpers:
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Using of Repumpers



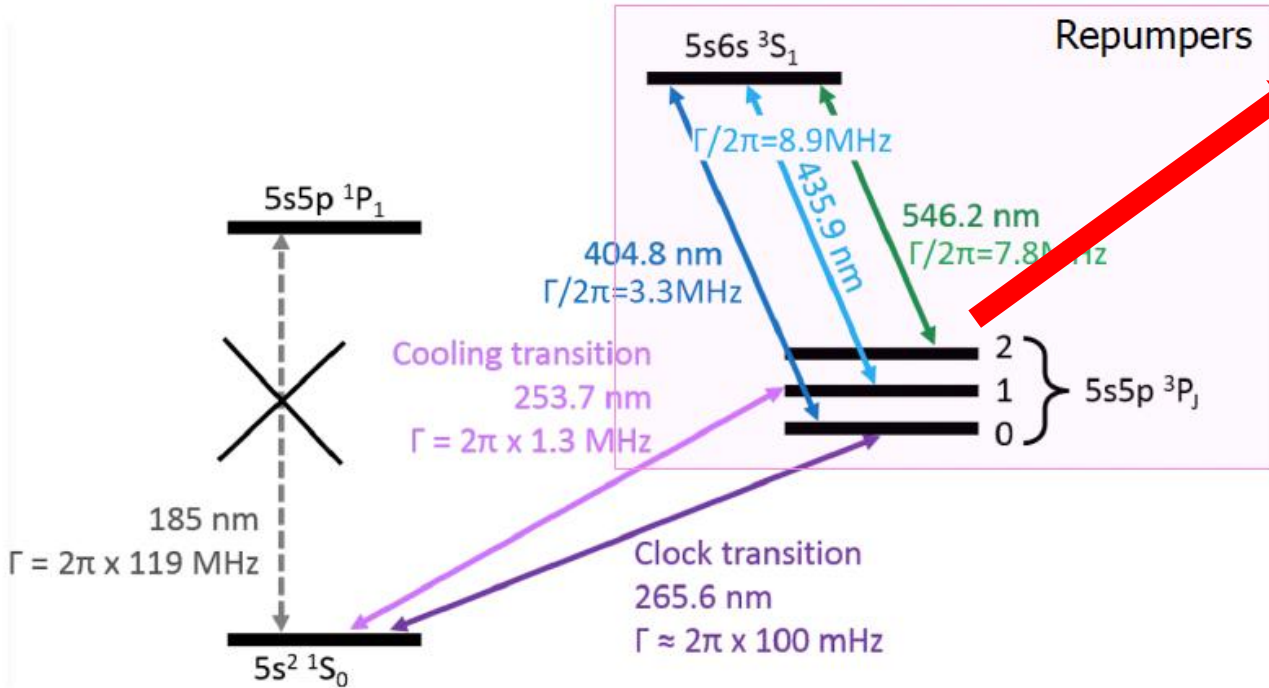
- $1D_2$ (Intensity: 2000) 366.3nm
- $3D_2$ (Intensity: 3000) 365.5nm
- $3D_3$ (Intensity: 9000) 365.01nm

Optical lattice: 362.5nm

No signal of 546 transition so far

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 - Normalized detection

Using of Repumpers



- $1D_2$ (Intensity: 2000) 366.3nm
- $3D_2$ (Intensity: 3000) 365.5nm
- $3D_3$ (Intensity: 9000) 365.01nm

Optical lattice: 362.5nm

- Blue detuned lattice w/ dipole force $\sim 100x$ larger
- Should loose atoms in 3P2 in a few ms
- Very short pulses of 546nm: need more power

No signal of 546 transition so far

- 198Hg repumpers:
- Sideband cooling
 - Decrease atomic cloud size
 - Normalized detection

Conclusion

□ 1st experimental measurement of the clock transition

in a bosonic isotope of Hg:

$$\frac{\nu_{198\text{Hg}}}{\nu_{87\text{Sr}}} = 2.629\,315\,734\,684\,118\,1 \quad \text{w/ } \sim 7 \times 10^{-16} \text{ uncertainty}$$

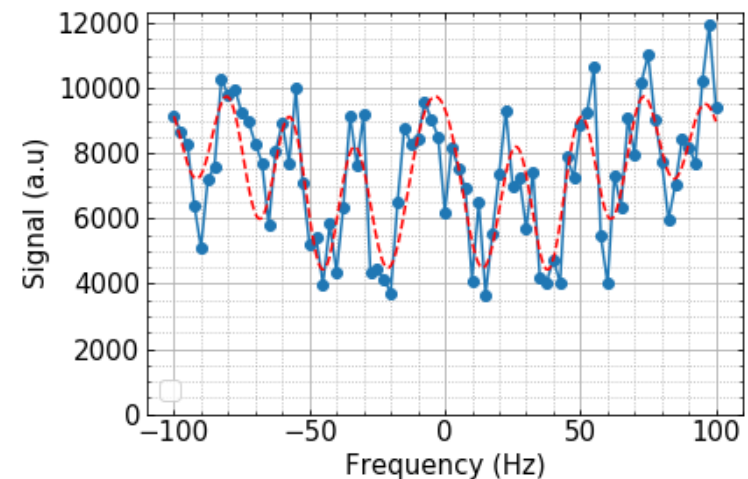
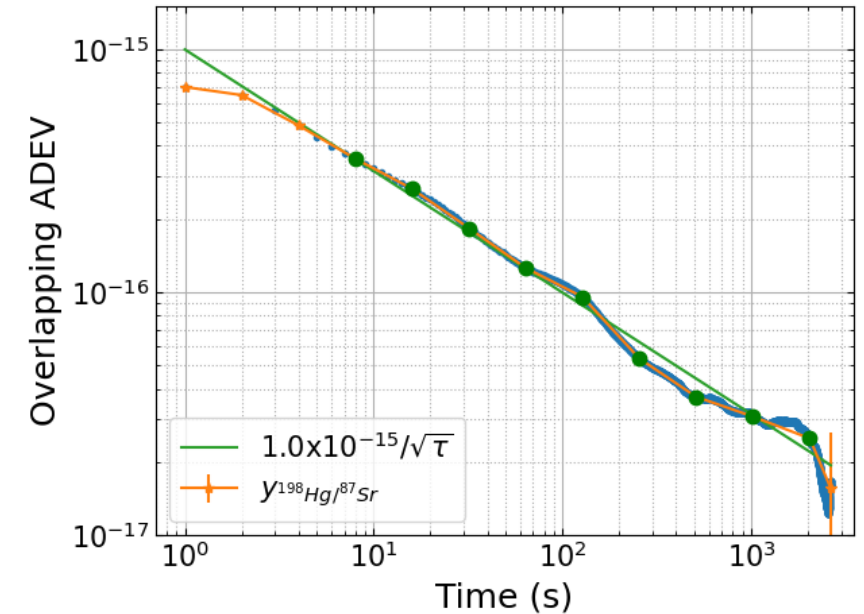
□ Stability $6 - 7 \times 10^{-16} / \sqrt{\tau}$

(Paper submitted on PRL)

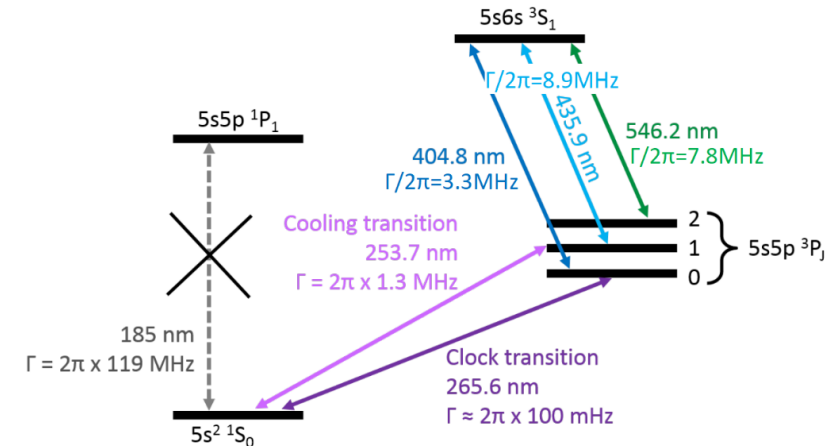
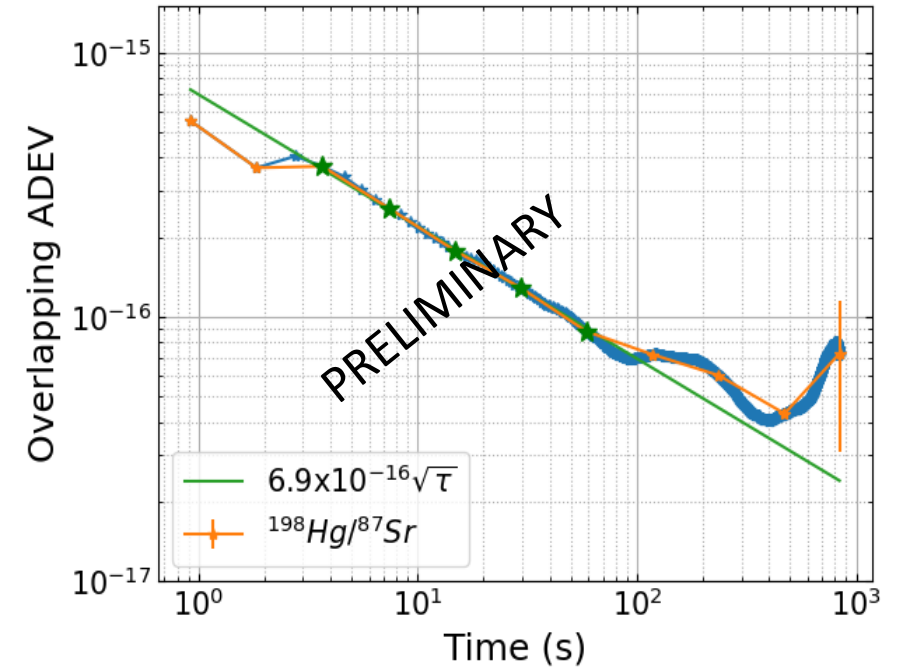
□ Uncertainty dominated by probe light shift uncertainty

Insensitive hyper-Ramsey interrogation

□ Setting up repumpers



- ❑ Investigating purity transfer and synchronization of Hg/Sr clocks
(cancelling Dick contribution)
- ❑ Continue working on implementing repumping
- ❑ New campaign of systematics calibration
(lattice light shift, cold collision calibration...)
- ❑ New vacuum chamber for improved lifetime ($\sim 600\text{ms}$)



Thanks!



Mercury Lab

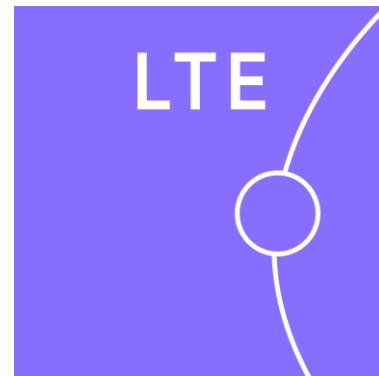
Sébastien Bize
Thomas Lauprêtre
Carla Zyskind

Sroutium Lab

Jérôme Lodewyck

Frequency Comb Lab

Benjamin Pointard
Rodolphe Le Targat

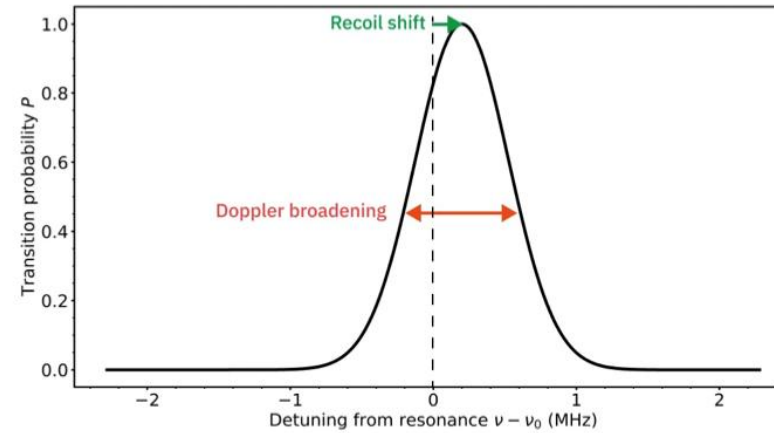


Laboratoire Temps Espace

198Hg systematics characterization

Lattice linear lightshift

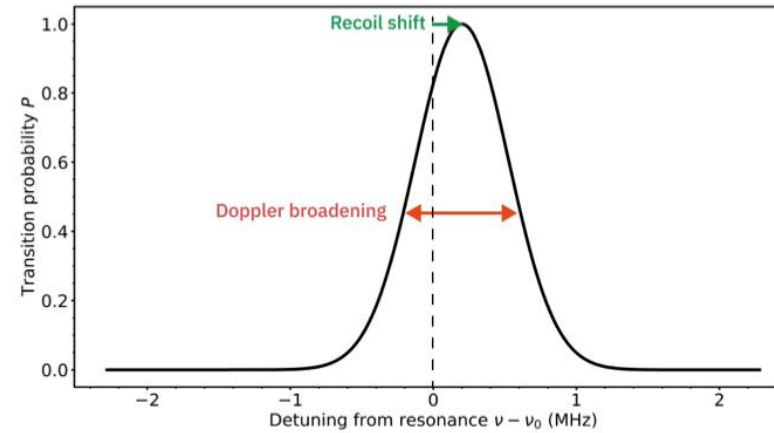
- Unbound atoms: Doppler broadening + recoil shift



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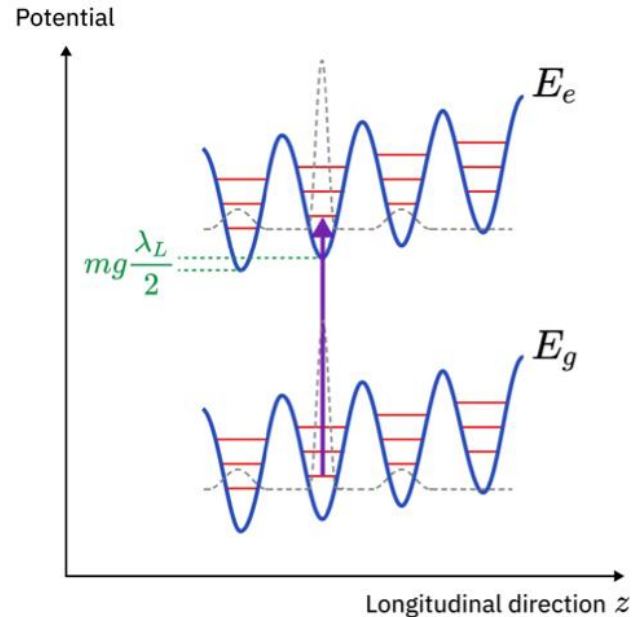
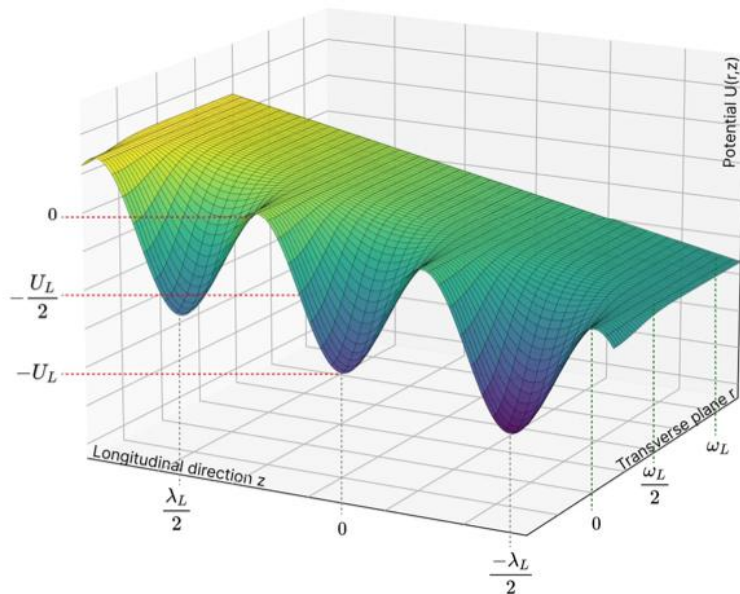
- Unbound atoms: Doppler broadening + recoil shift
- Trapped atoms: suppressing motional effects



Lattice linear lightshift

- ❑ Unbound atoms: Doppler broadening + recoil shift
- ❑ Trapped atoms: suppressing motional effects

Vertical lattice trap:



Confinement in Lamb-Dicke regime:

- ❑ Intracavity power $\sim 9\text{W}$
- ❑ Waist $\sim 69\ \mu\text{m}$

➔ **Strong Stark shifts on the two clock levels**