

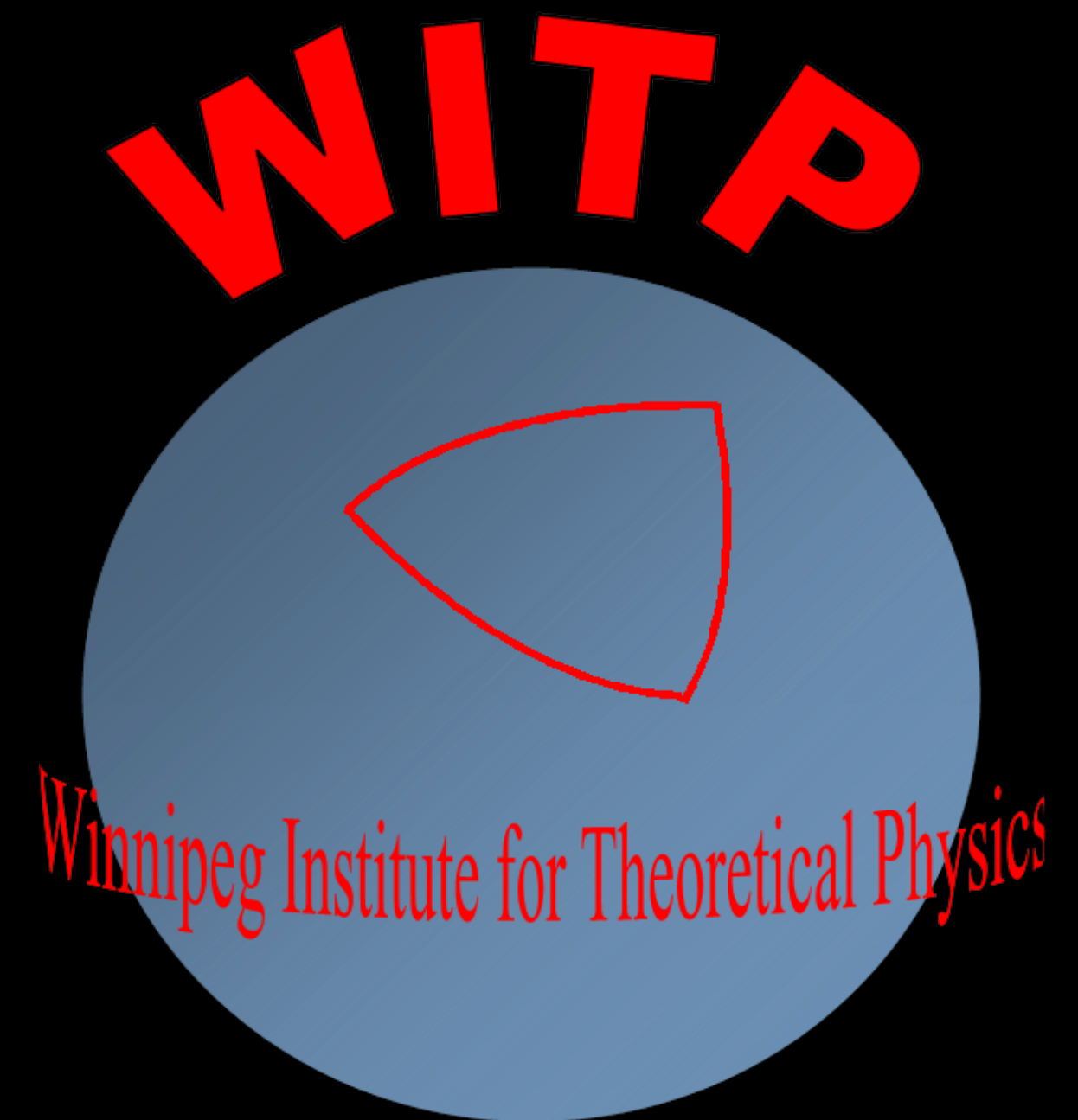
# Higgs Inflation: *Particle Factory*

Evan McDonough  
University of Winnipeg  
Department of Physics



June 22, 2026  
CAP Congress

Based on Chowdhury, Kolb, Jenks, EM  
JHEP '26



# Plan for today:

1. Particle Cosmology & Cosmic Inflation

2. Particle Production

3. Higgs Inflation: Particle Factory

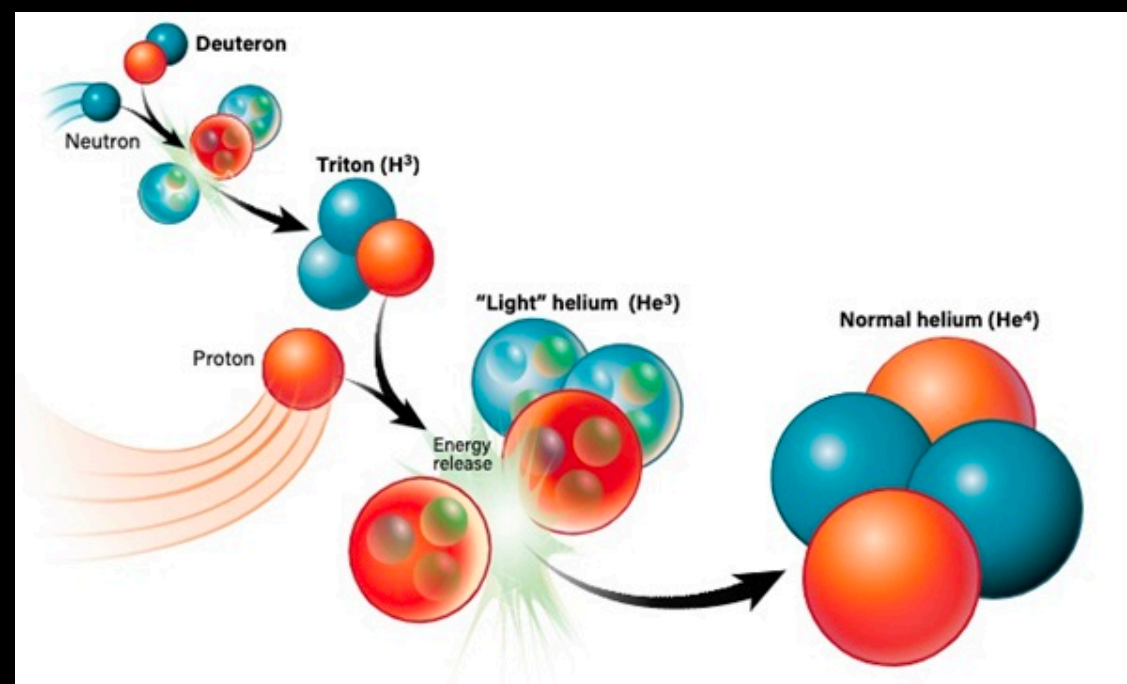
Hot off the presses: *Non-thermal Leptogenesis*

# Particle Cosmology circa 1948

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Winnipeg

## Big Bang Nucleosynthesis

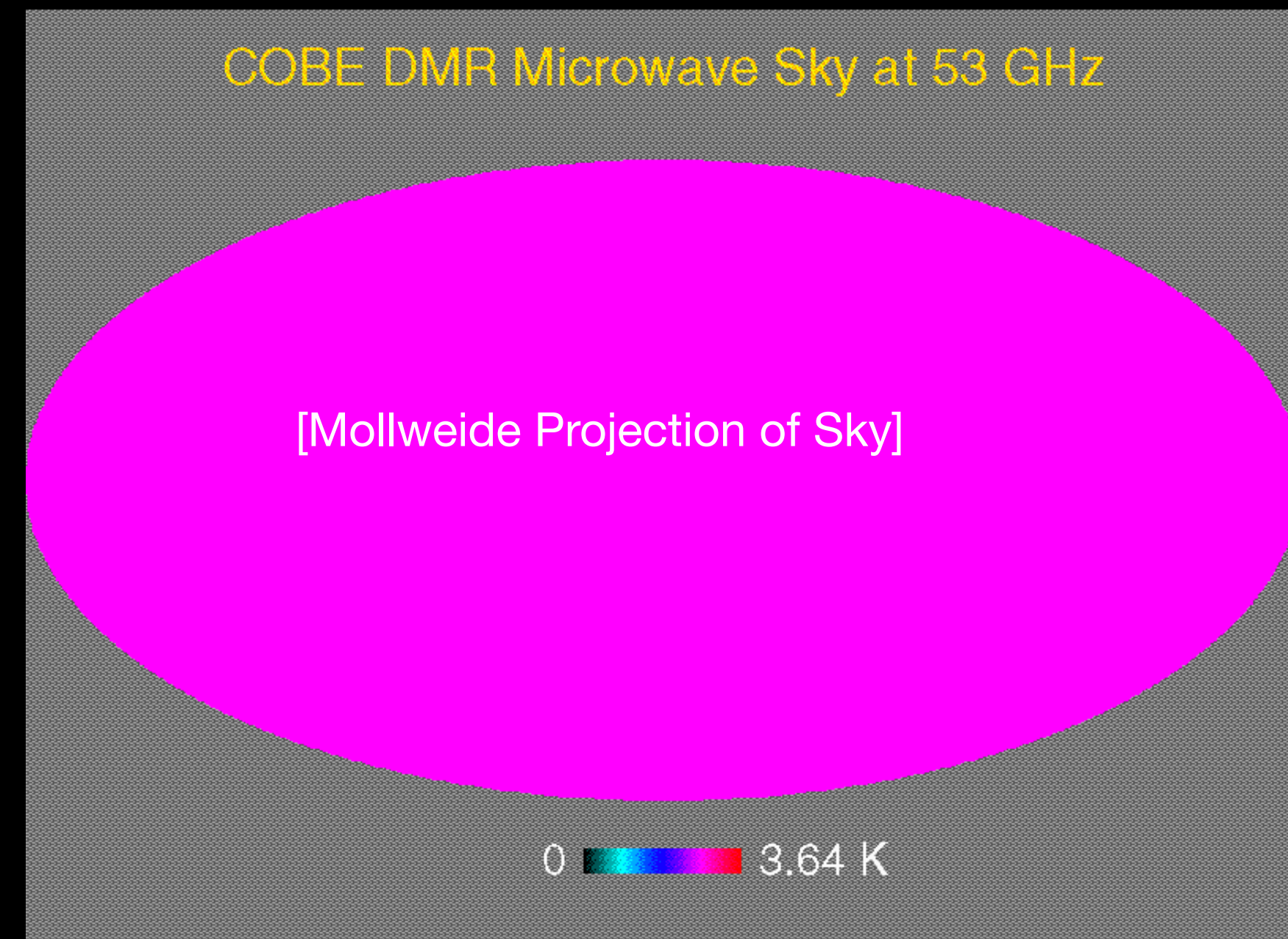
The “ $\alpha\beta\gamma$ ” paper : 1948 Alpher, Bethe, Gamow



Fantastic agreement with observations

## Cosmic Microwave Background

Predicted: 1948 Alpher & Herman  
Discovered: 1964 Penzias & Wilson.  
Nobel Prize: 1978



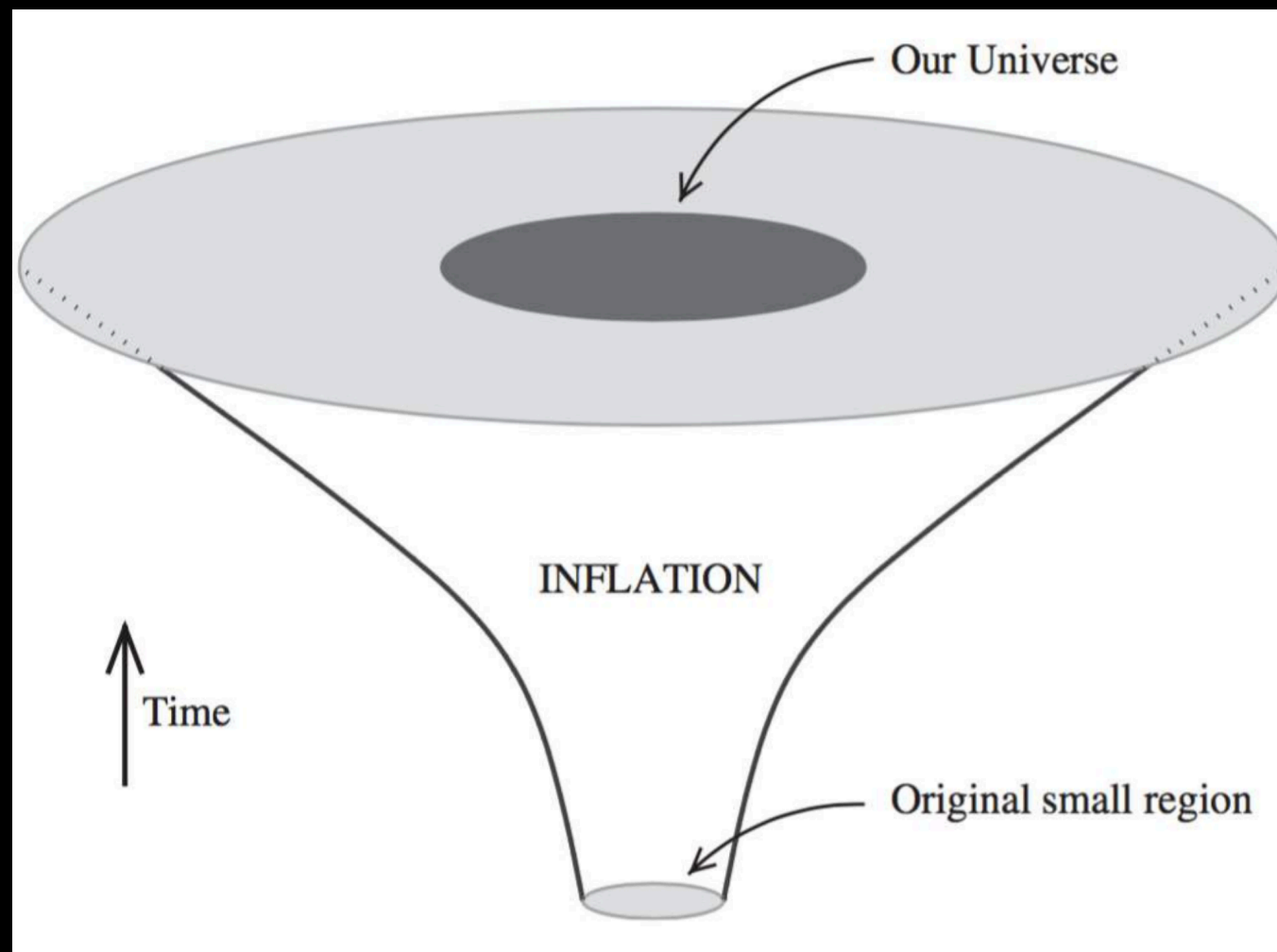
# 1980's: Start of the Modern Era

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Winnipeg

## Cosmic Inflation

1980

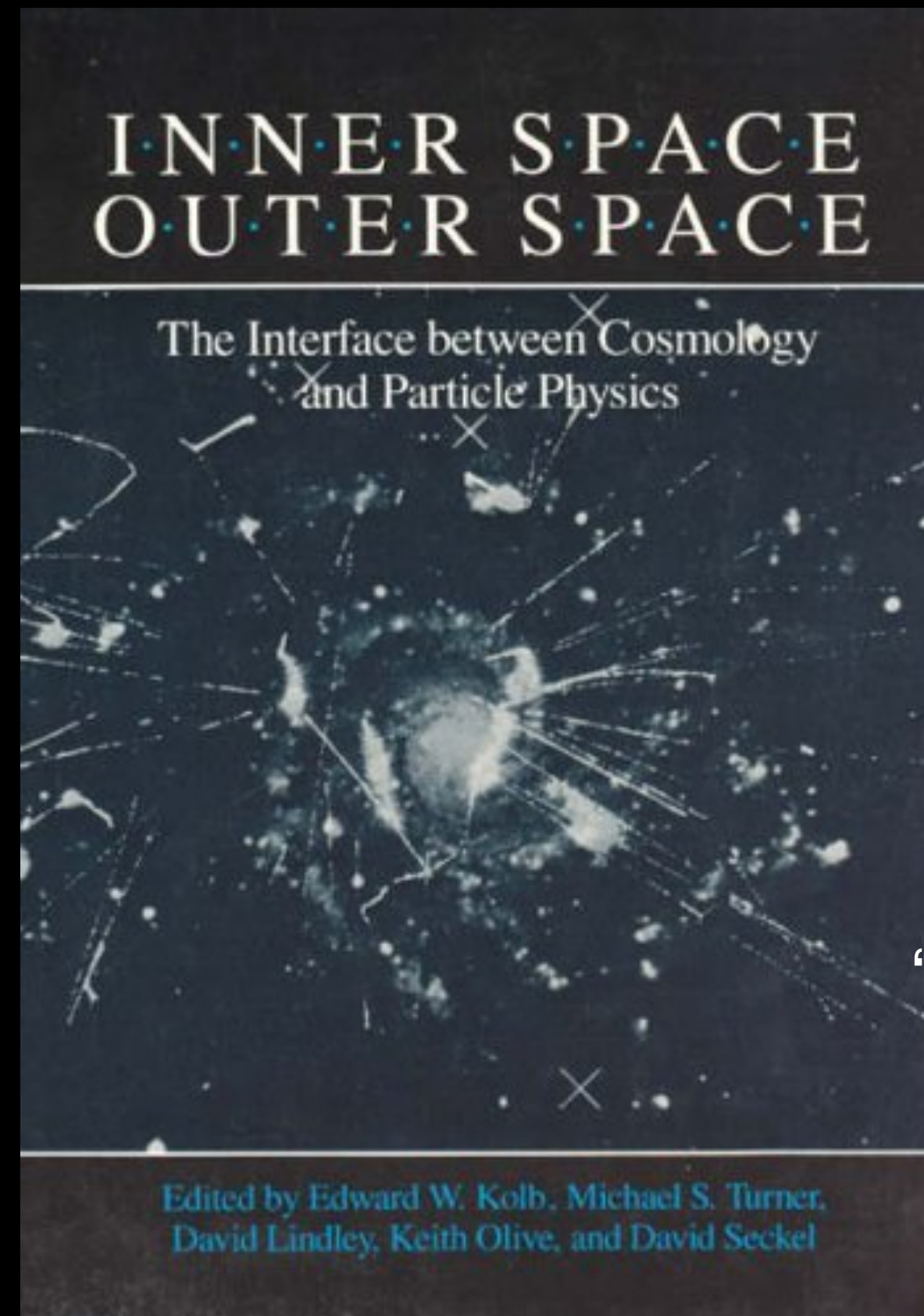
Guth, + others



## Particle Astrophysics Division, Fermilab

1983

Dir. Edward Kolb

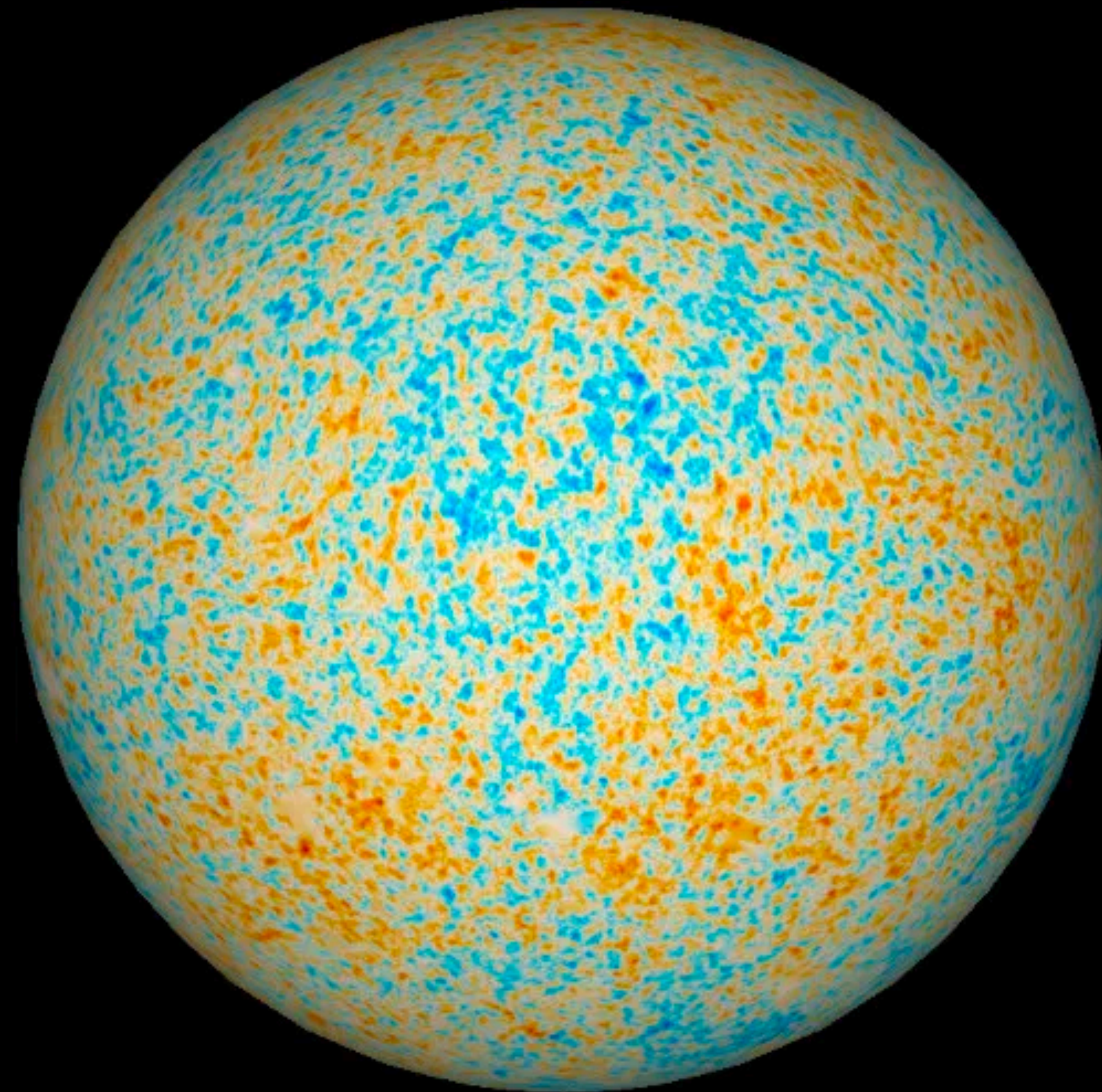


## WIMP Miracle

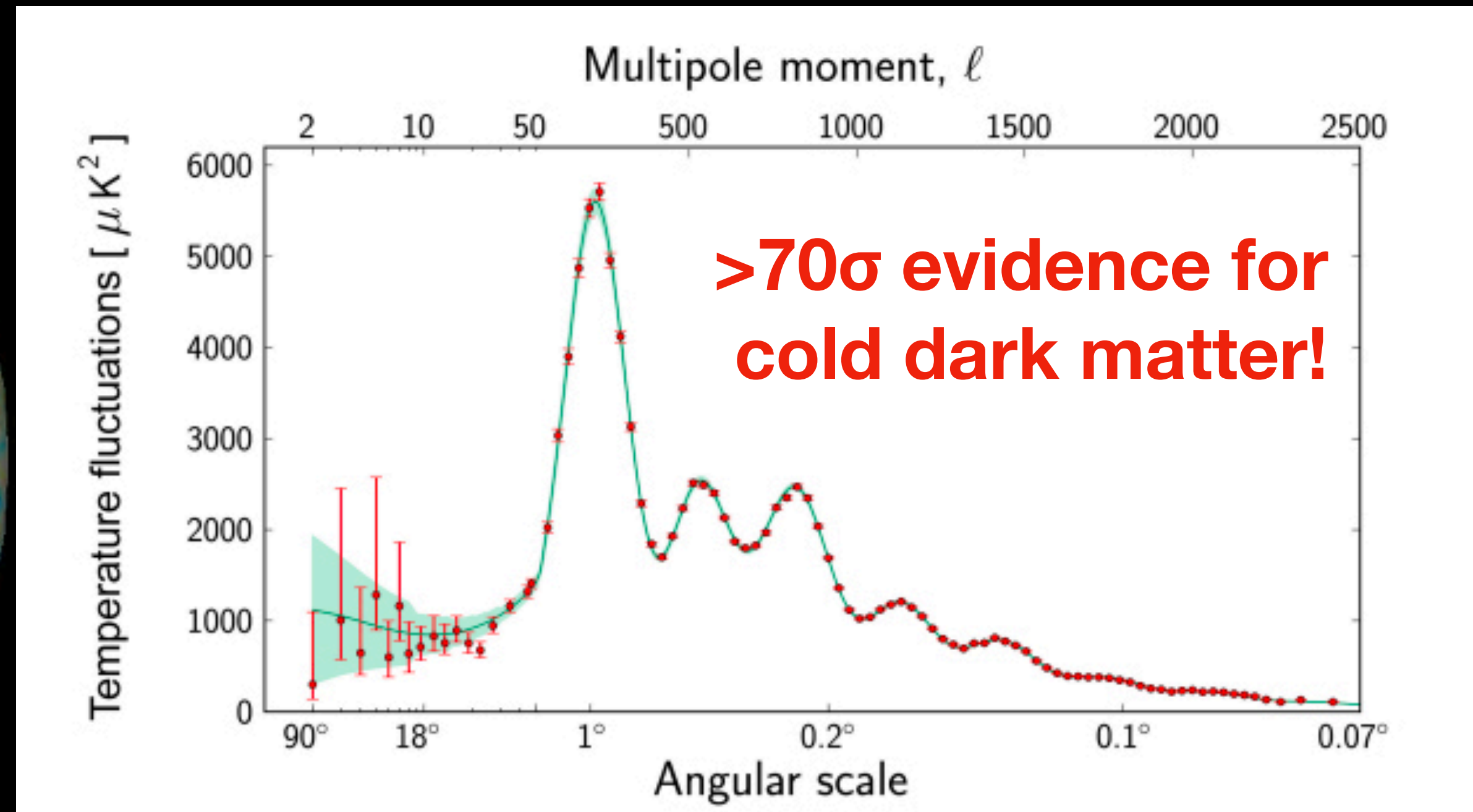
Goldberg (1983).  
Hagelin, Nanopoulos, Olive,  
Srednicki (1984).  
Steigman & Turner (1985)

1984 conference:  
“Inner Space Outer Space”

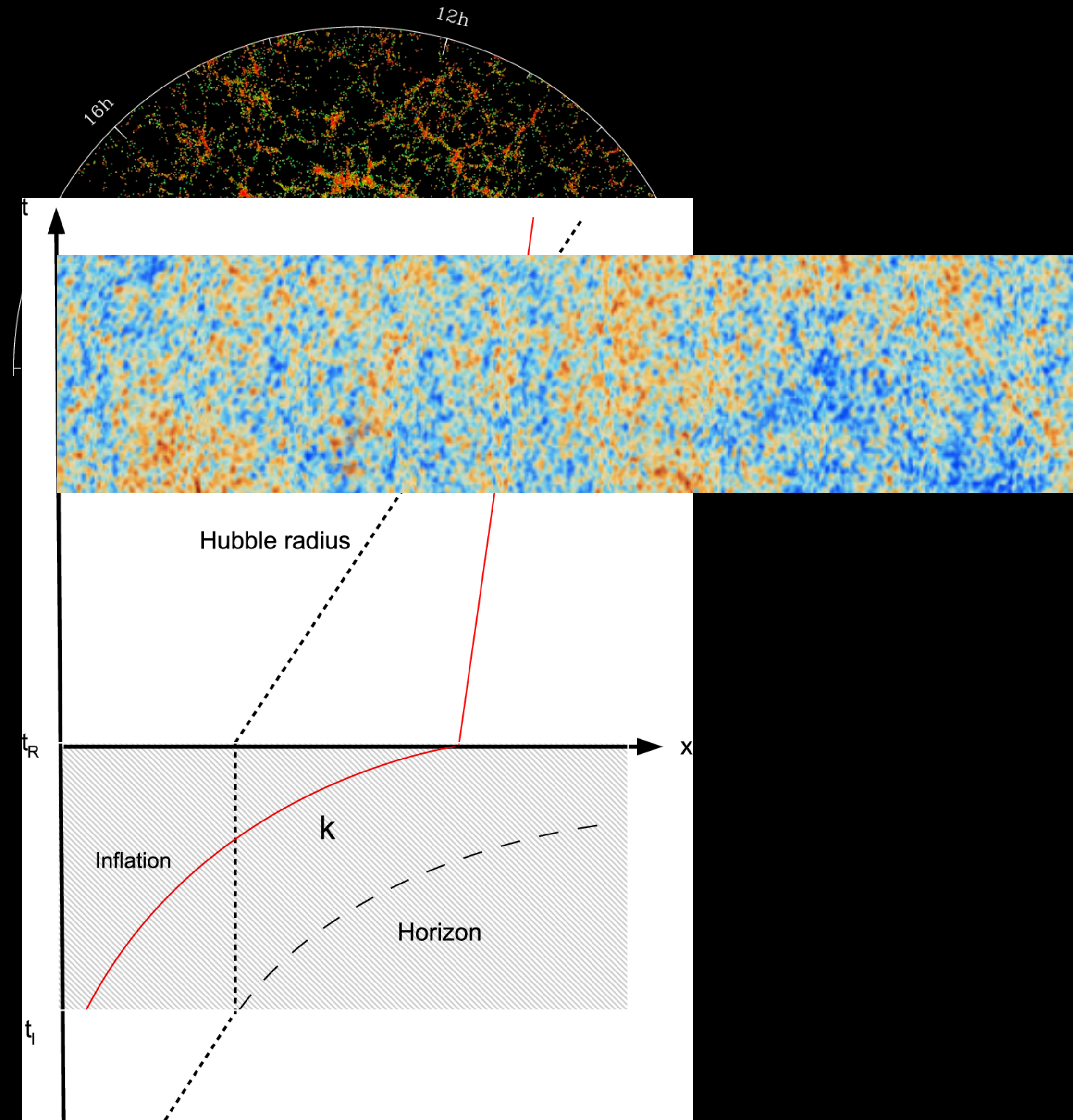
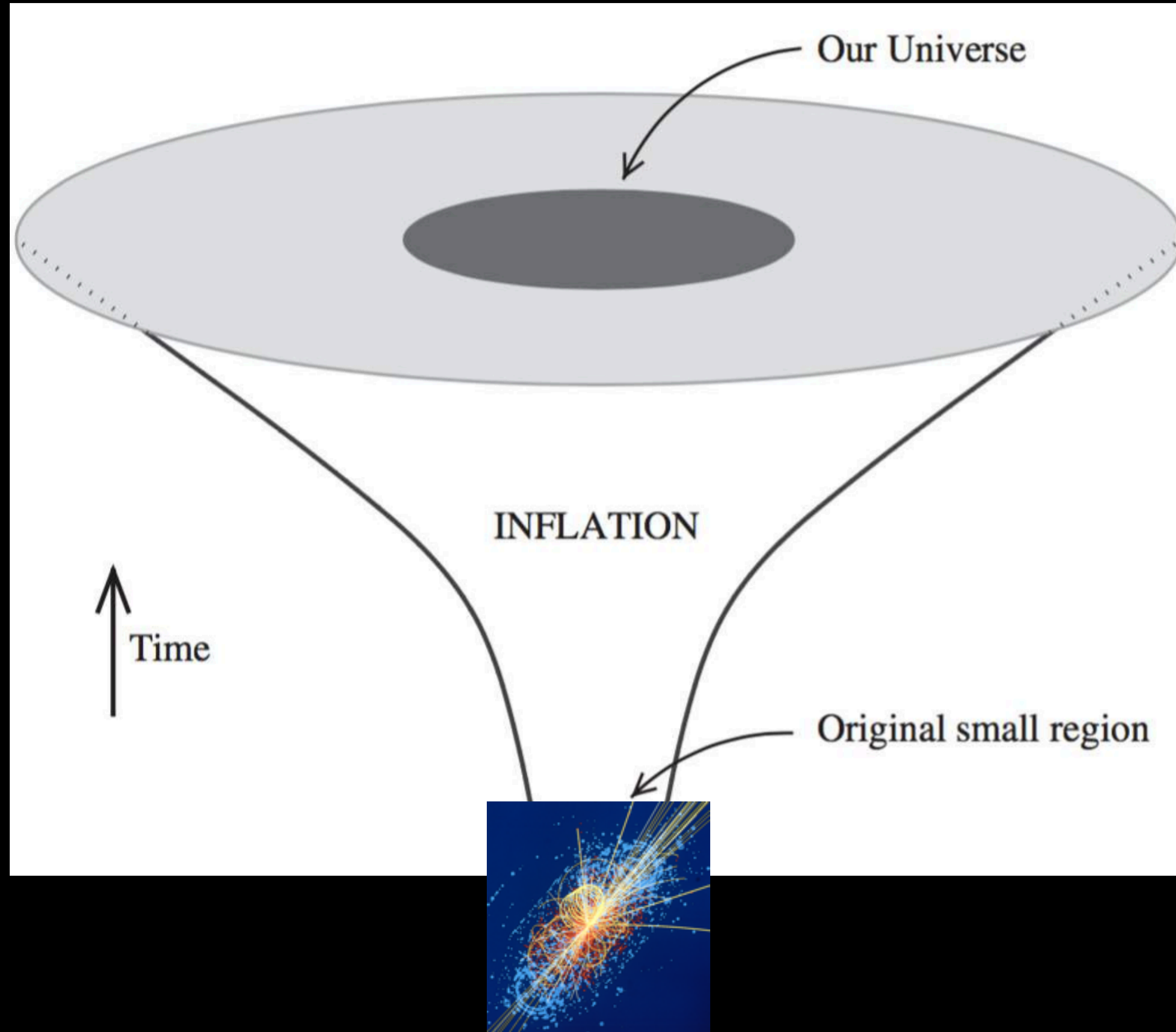
# Evidence for Dark Matter



Planck 2018



# Cosmic Inflation: Cosmological Collider



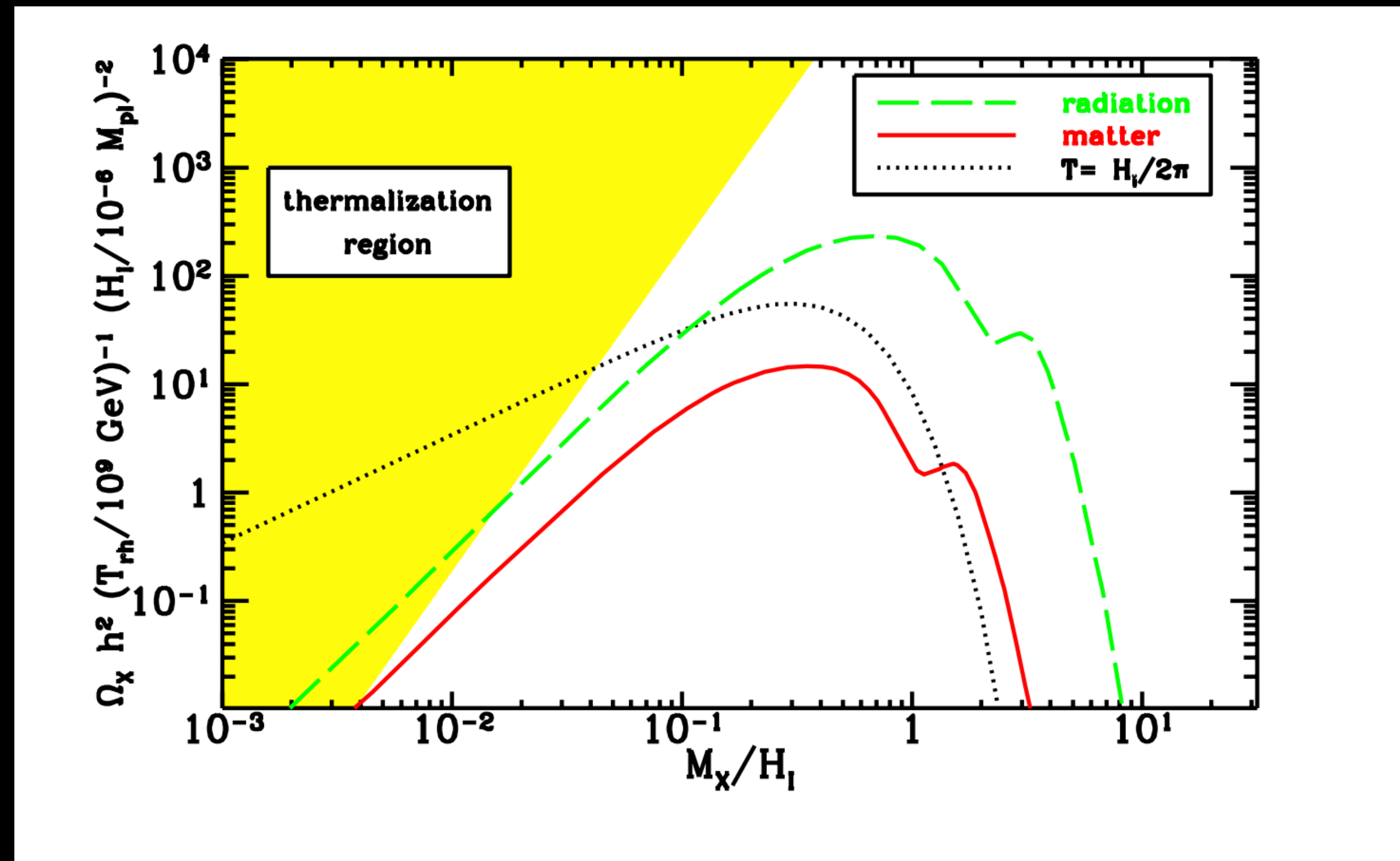
# Cosmic Inflation as Particle Factory

Evan McDonough  
Winnipeg

[Submitted on 14 Oct 1998]

## WIMPZILLAS!

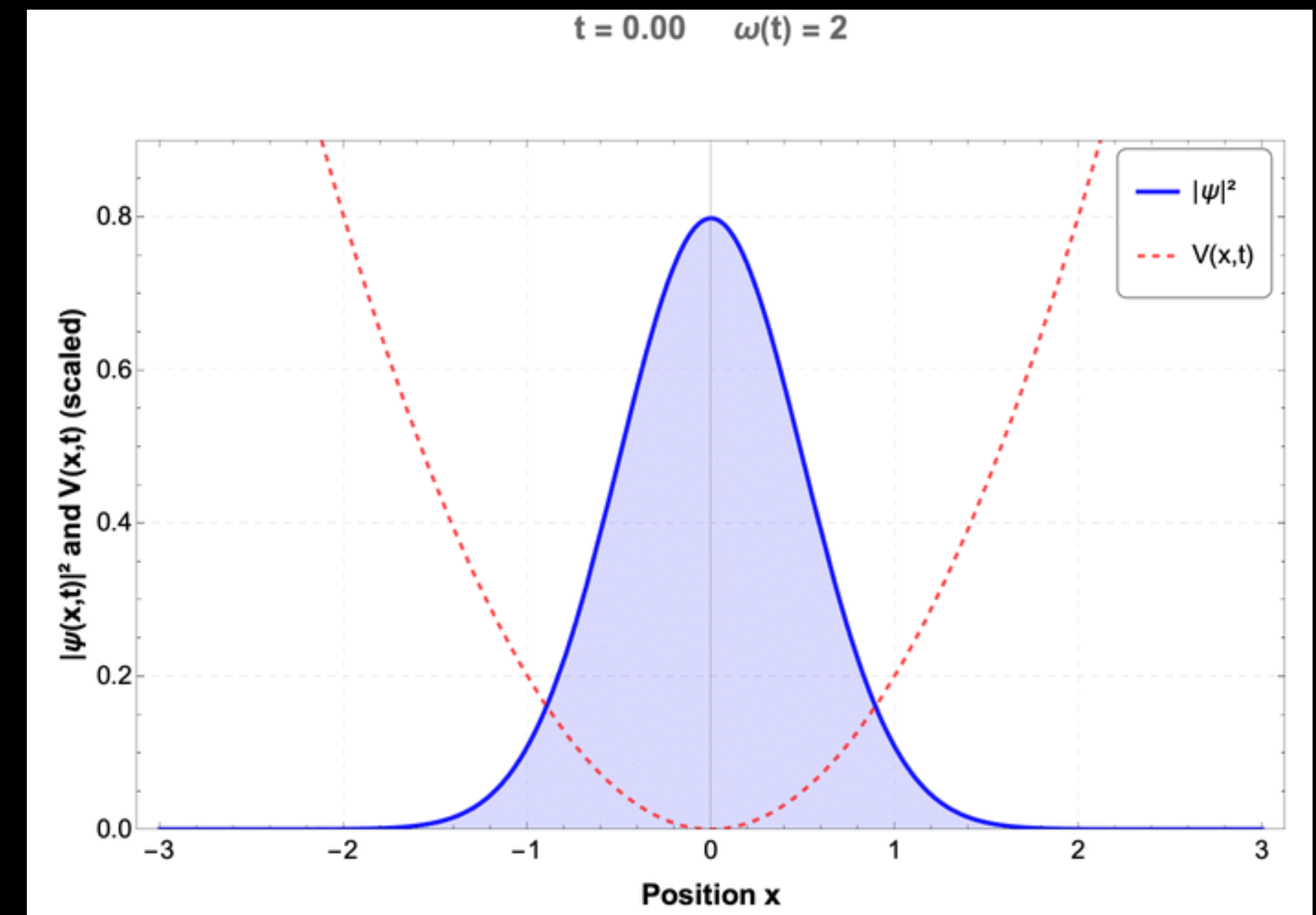
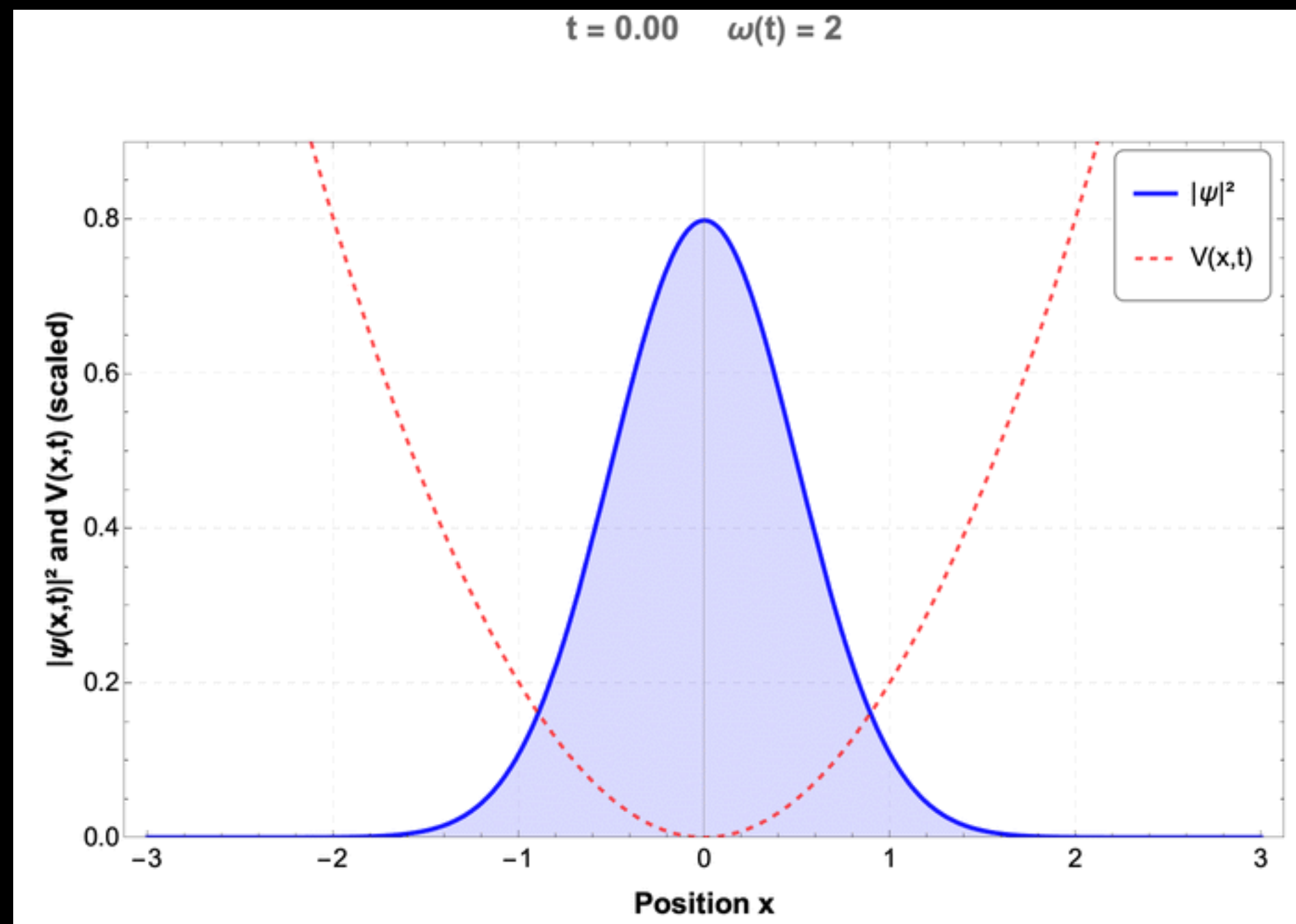
Edward W. Kolb, Daniel J. H. Chung, Antonio Riotto



# “Particle Production” in Quantum Mechanics

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Winnipeg

$$i\hbar \frac{\partial \psi}{\partial t} = -\frac{\hbar^2}{2m} \nabla^2 \psi + \frac{1}{2} m \omega(t)^2 x^2$$

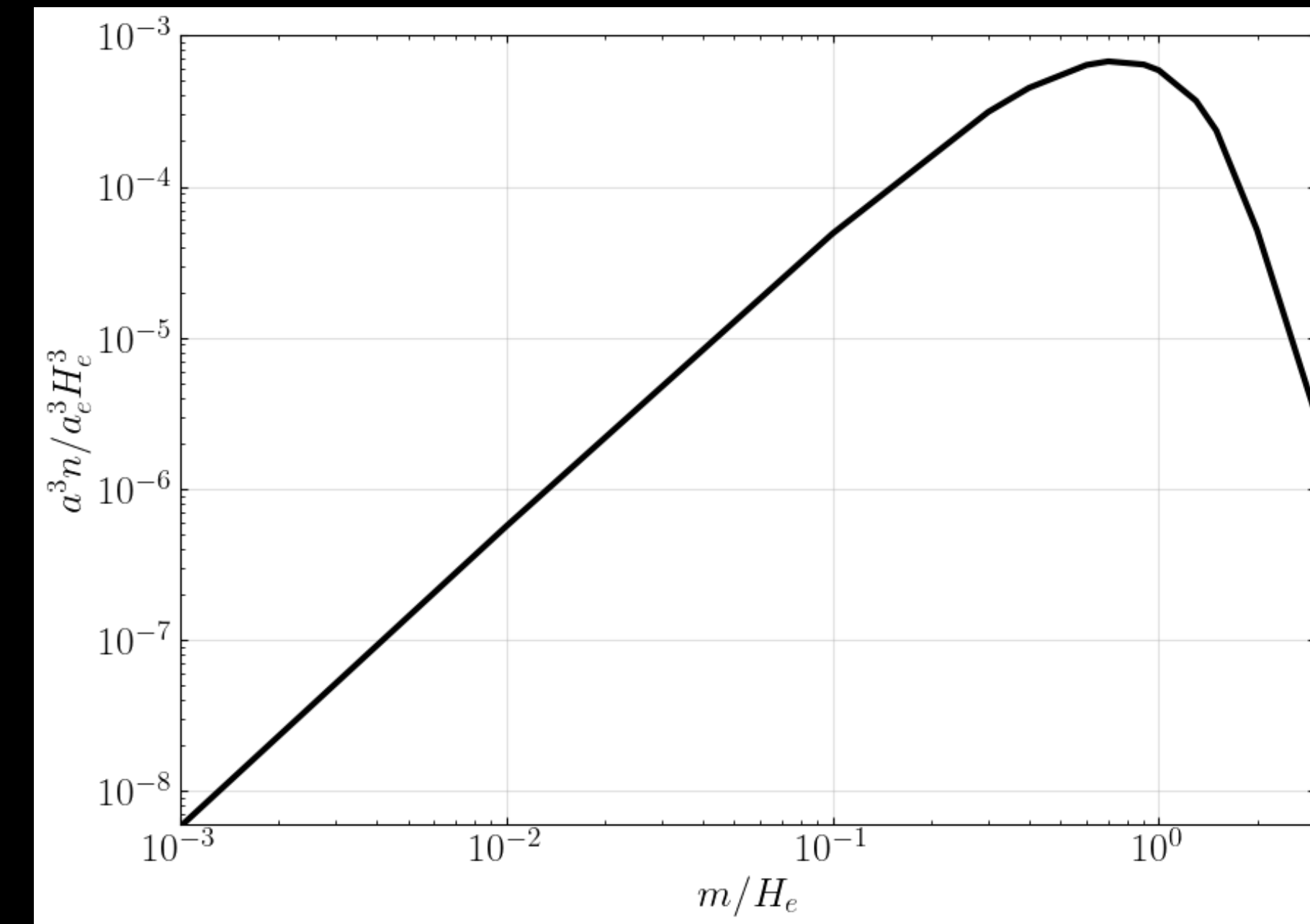
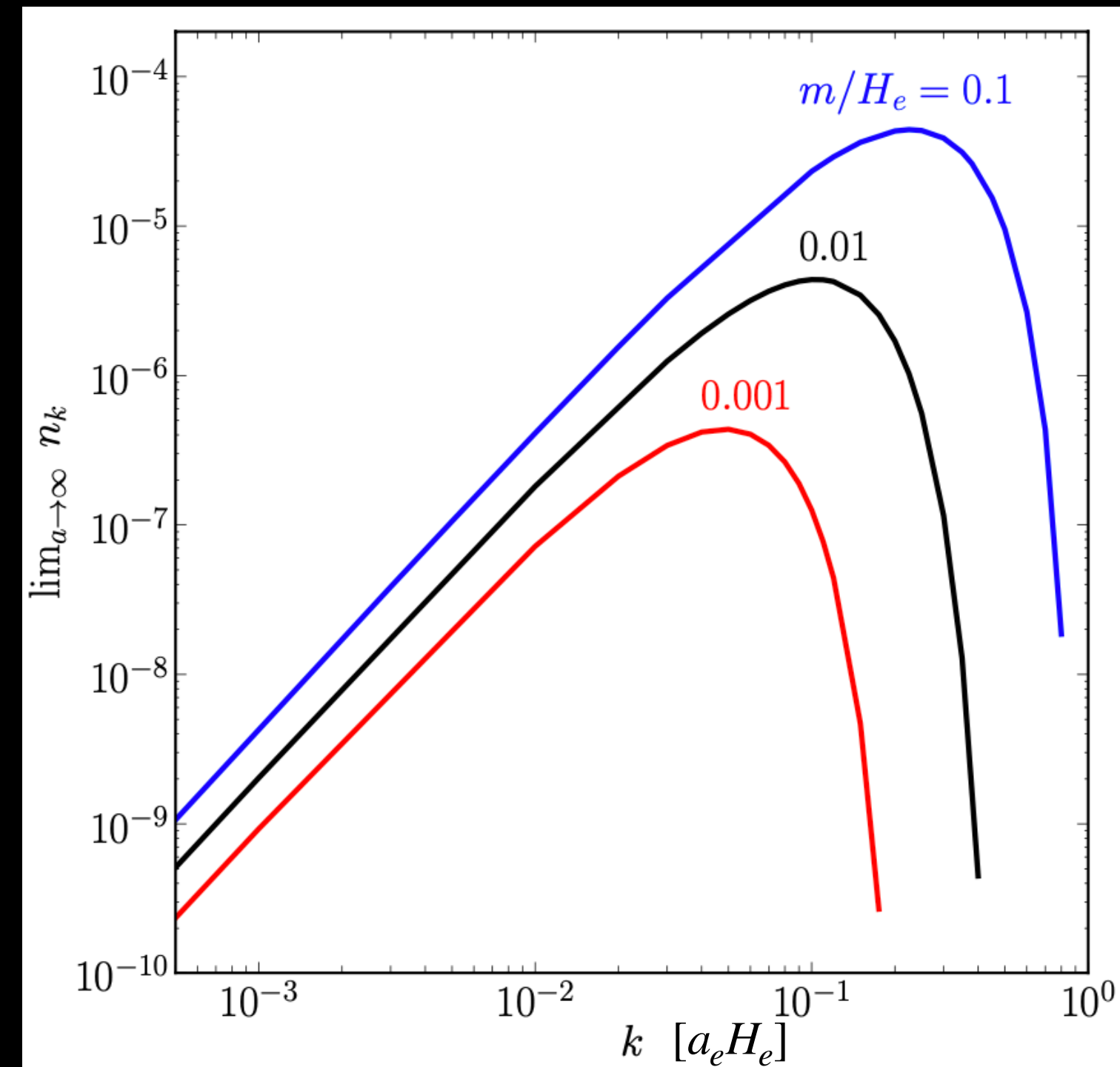
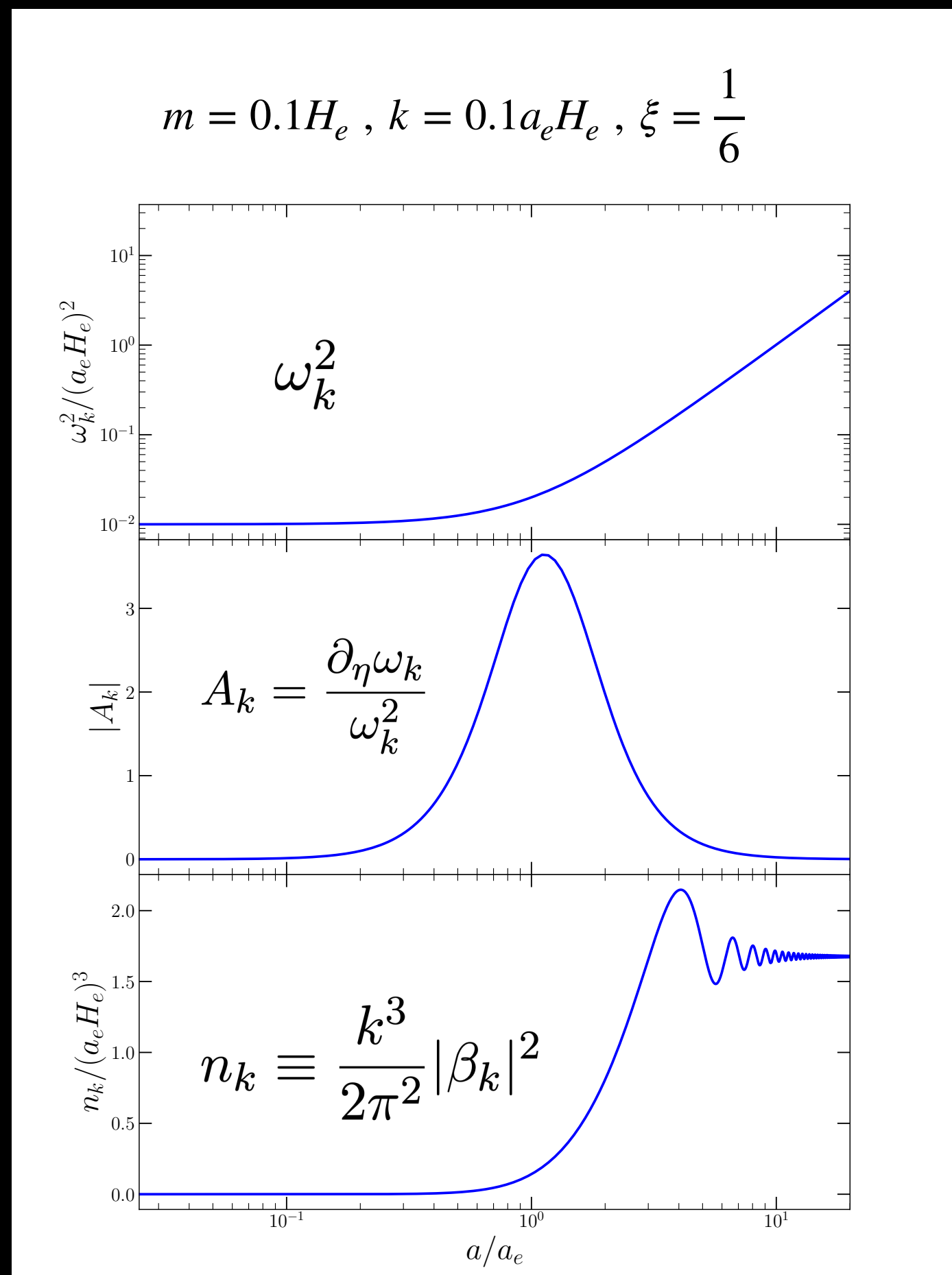


# Application to inflation: Scalar Field

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Kolb, Riotto, Chung, +...

$$\omega_k^2 = k^2 + m^2 a^2 + \left(\frac{1}{6} - \xi\right) a^2 R$$



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# Inflation: idea in search of a particle

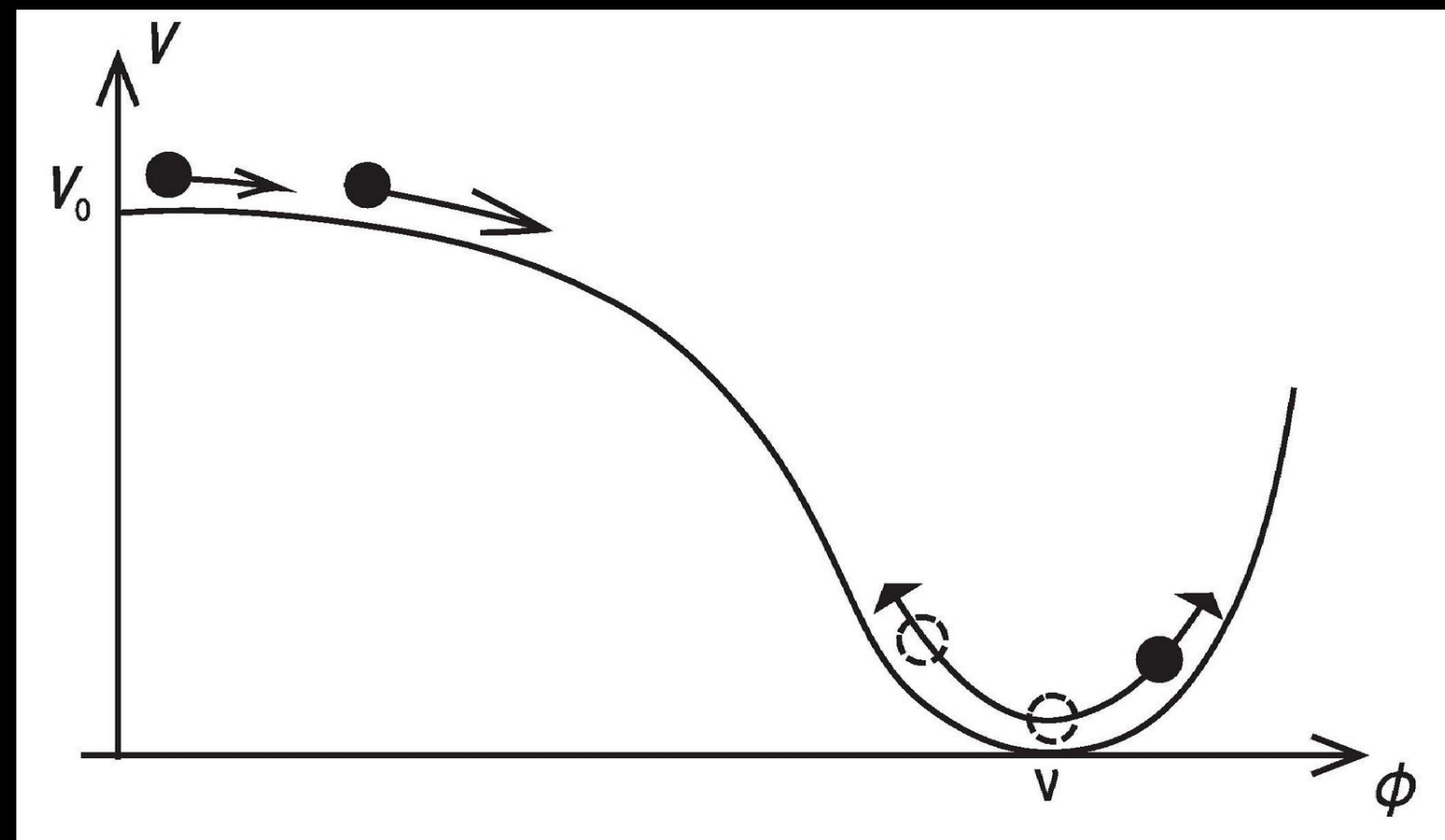
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— Many models. Most common: scalar field with  $V(\phi)$

— Many are now ruled out, e.g.

$V = \phi^2$ ,  $\phi^4$ : potential is too steep ( $V'/M$ )

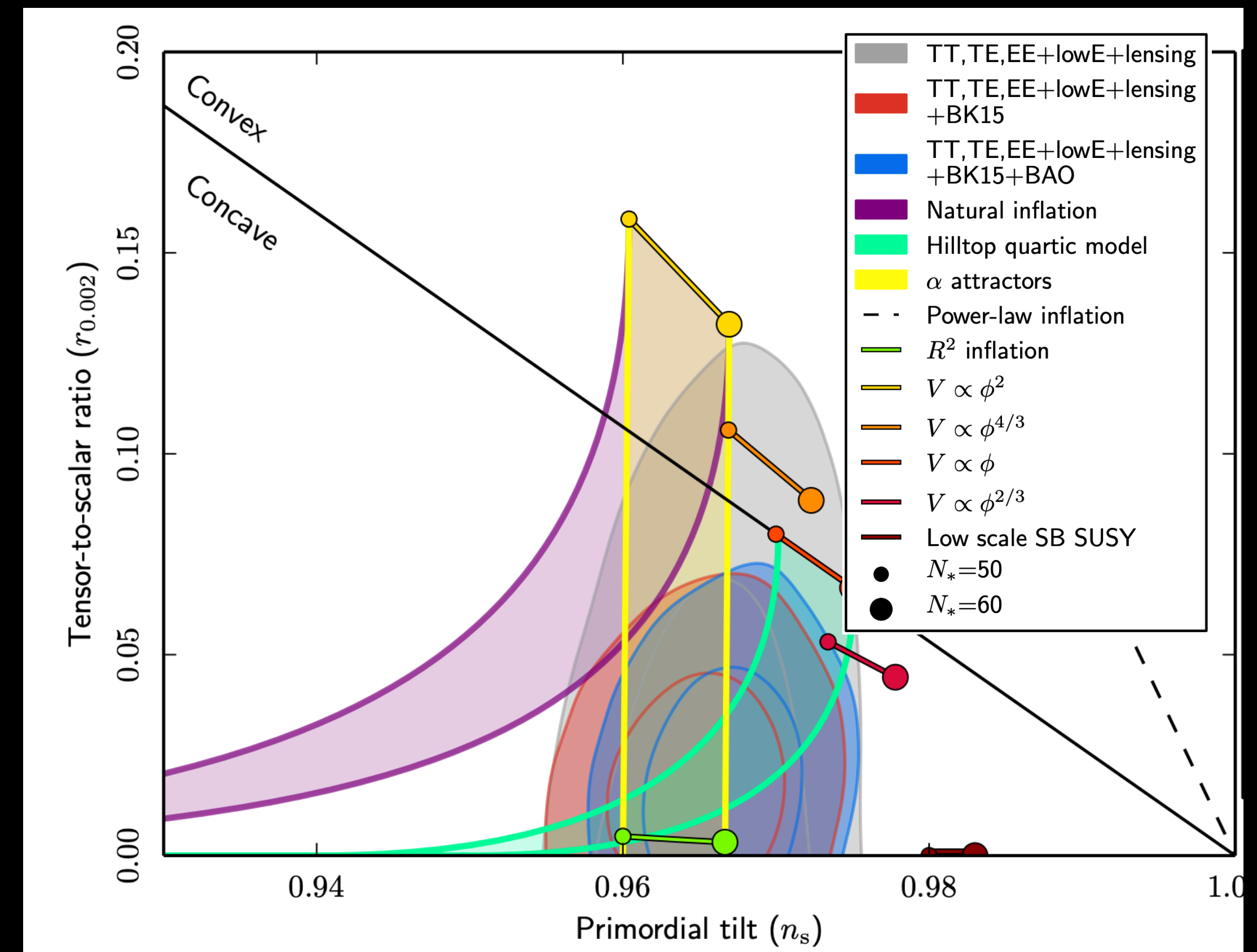
Planck 2018



Challenges:

— How connected to SM?

— Needs Quantum Gravity?



# Higgs Inflation

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Winnipeg

$$\mathcal{L} = \frac{1}{2}M_{pl}^2 R + \frac{1}{2}|D_\mu \Phi|^2 - \underbrace{\frac{\lambda}{4}(|\Phi|^2 - v^2)^2}_{\text{Too steep}} + \xi |\Phi|^2 R$$

Too steep

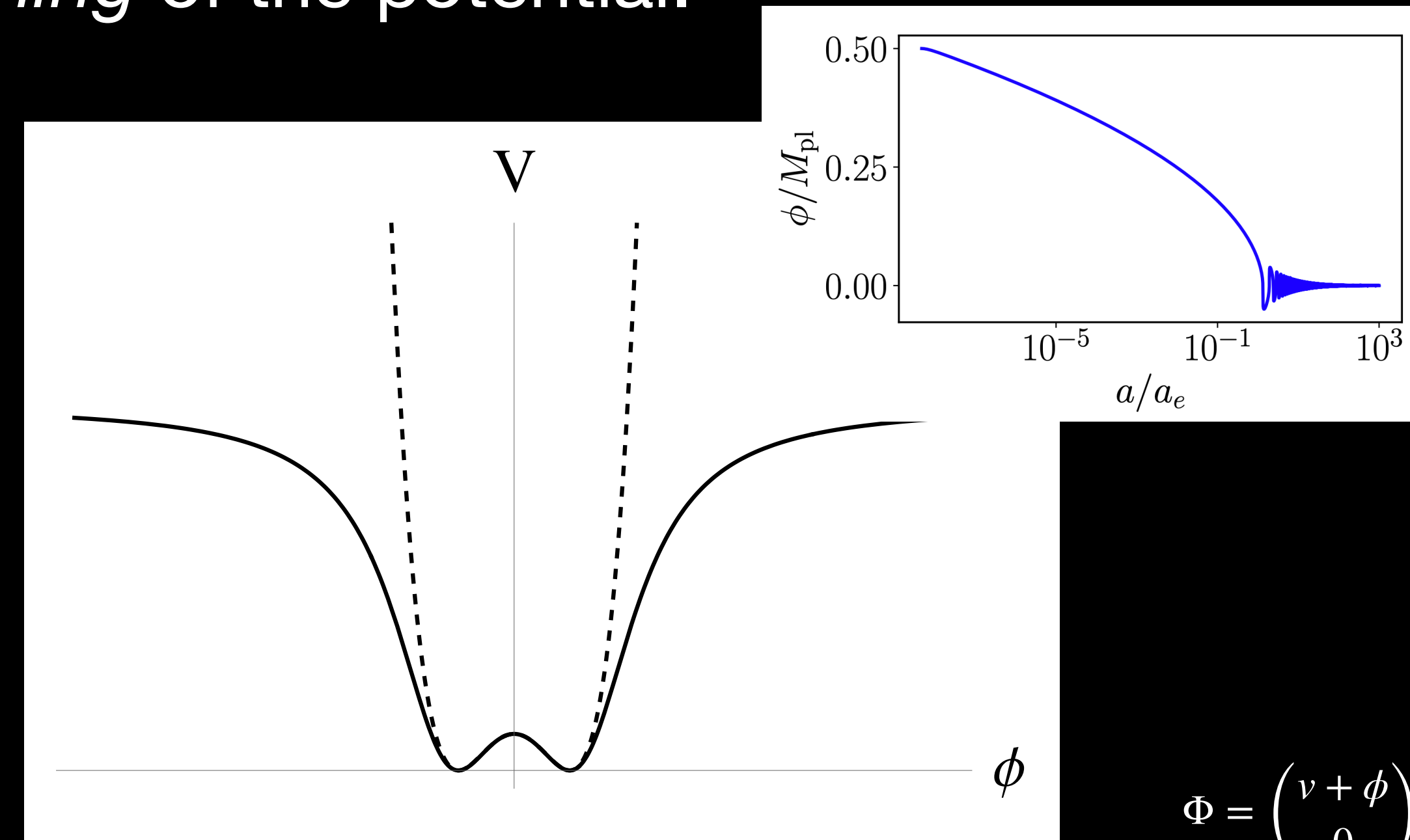
“Nonminimal coupling”

1970’s: Coleman, Parker, +...  
Required by renormalization

During inflation,  $\xi$  acts like a *stretching* of the potential.

Transform to “Einstein Frame”

$$g_{\mu\nu} \rightarrow e^{\Omega(\phi)} g_{\mu\nu} \Rightarrow V \rightarrow \frac{V}{(M_{pl}^2 + \xi |H|^2)^2}$$



$$\Phi = \begin{pmatrix} v + \phi \\ 0 \end{pmatrix}$$

# Cosmological Observables

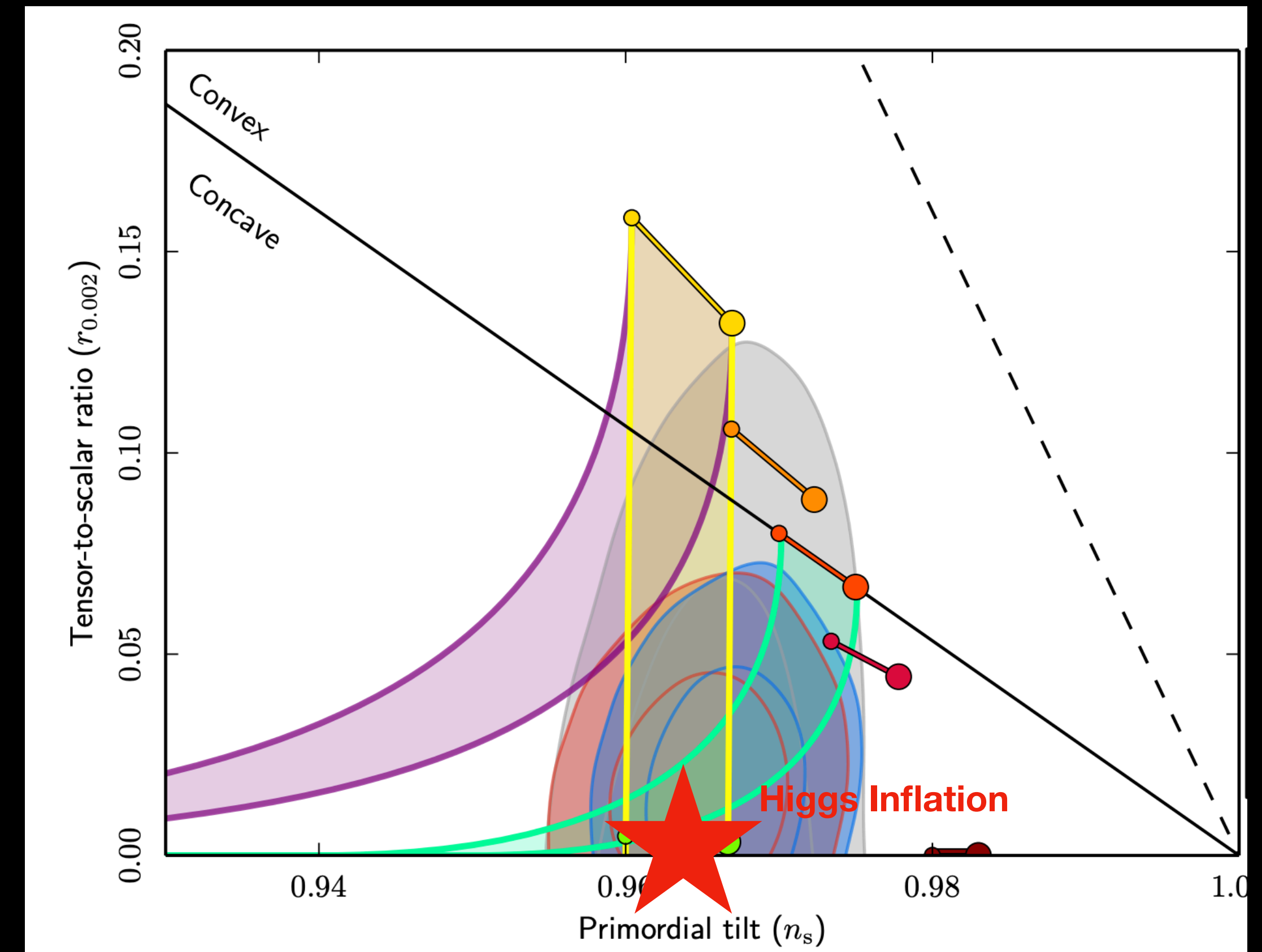
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Spectrum of scalar fluctuations:

$$\mathcal{P}(k) = A_s \left( \frac{k}{k_0} \right)^{n_s - 1}$$

$$A_s \sim \frac{\lambda}{\xi^2} \sim 10^{-10} \Rightarrow \xi \sim 10^5 \sqrt{\lambda(E_{\text{inf}})}$$

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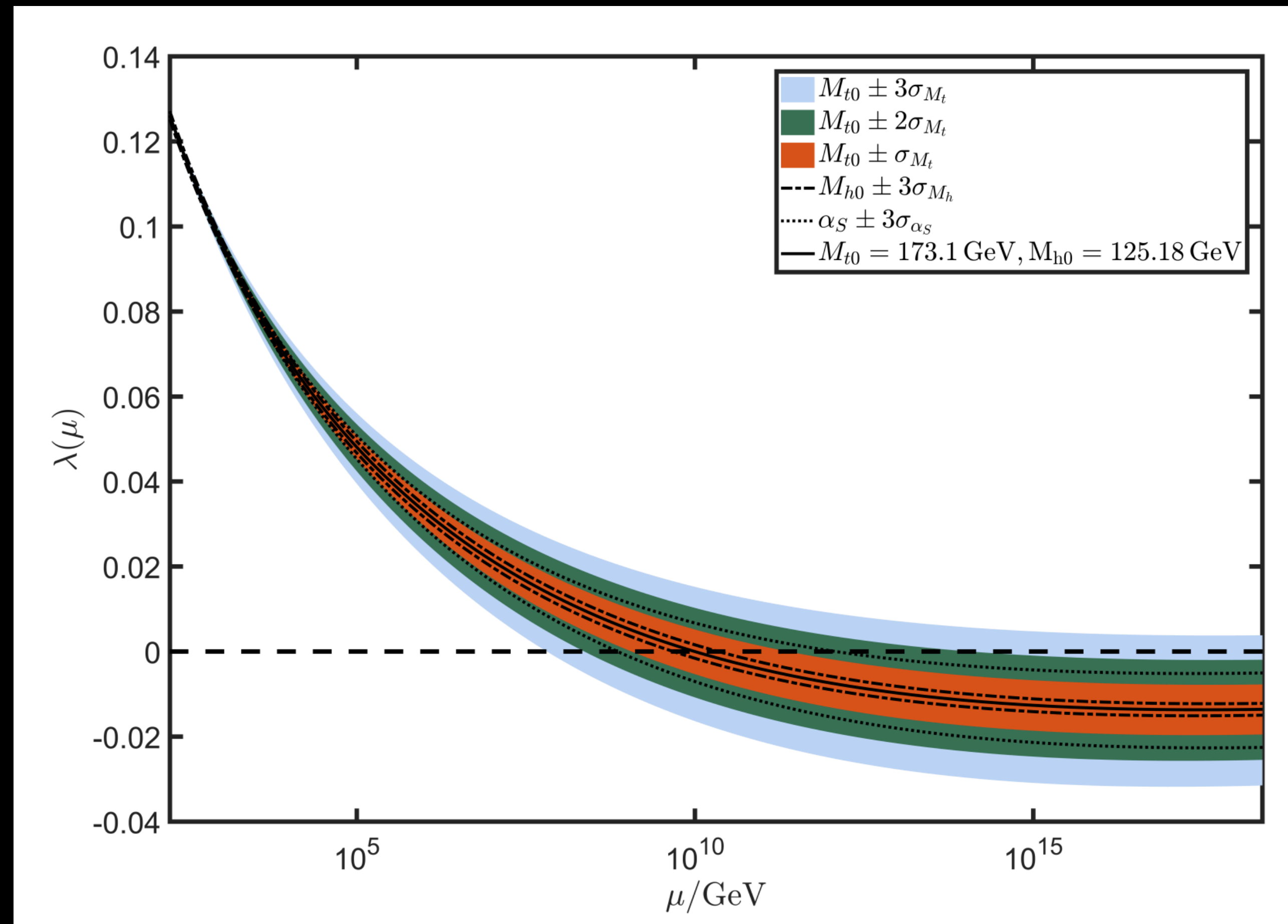


## The Planck “Sweet Spot”

# The Higgs Self-Coupling: Instability!

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Winnipeg

Energy scale of Higgs Inflation is  $\sim 10^{12}\text{GeV}$ .

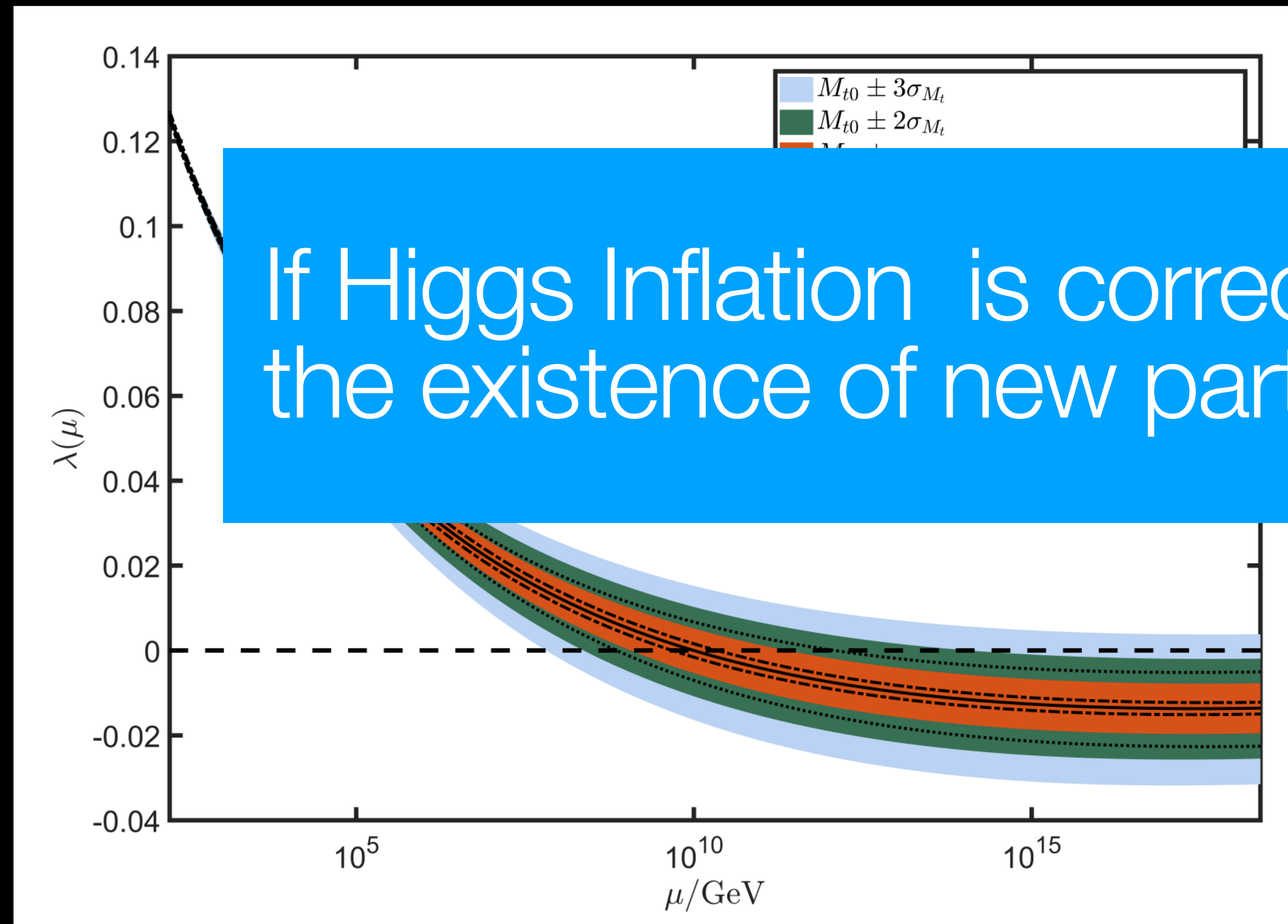


Markkanen, Rajantie, Stopyra '18

# The Higgs Self-Coupling: Instability!

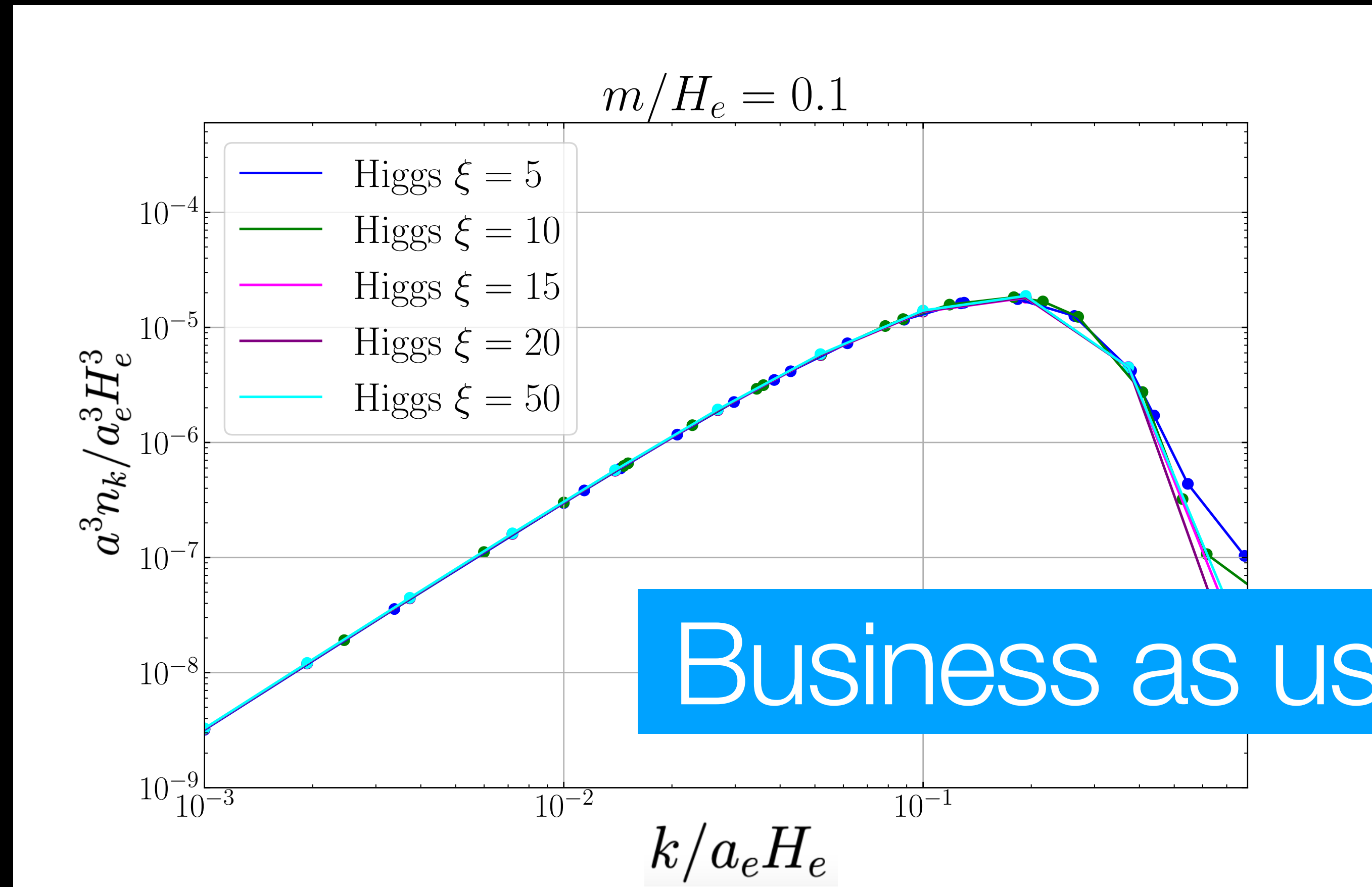
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Energy scale of Higgs Inflation is  $\sim 10^{12}\text{GeV}$ .



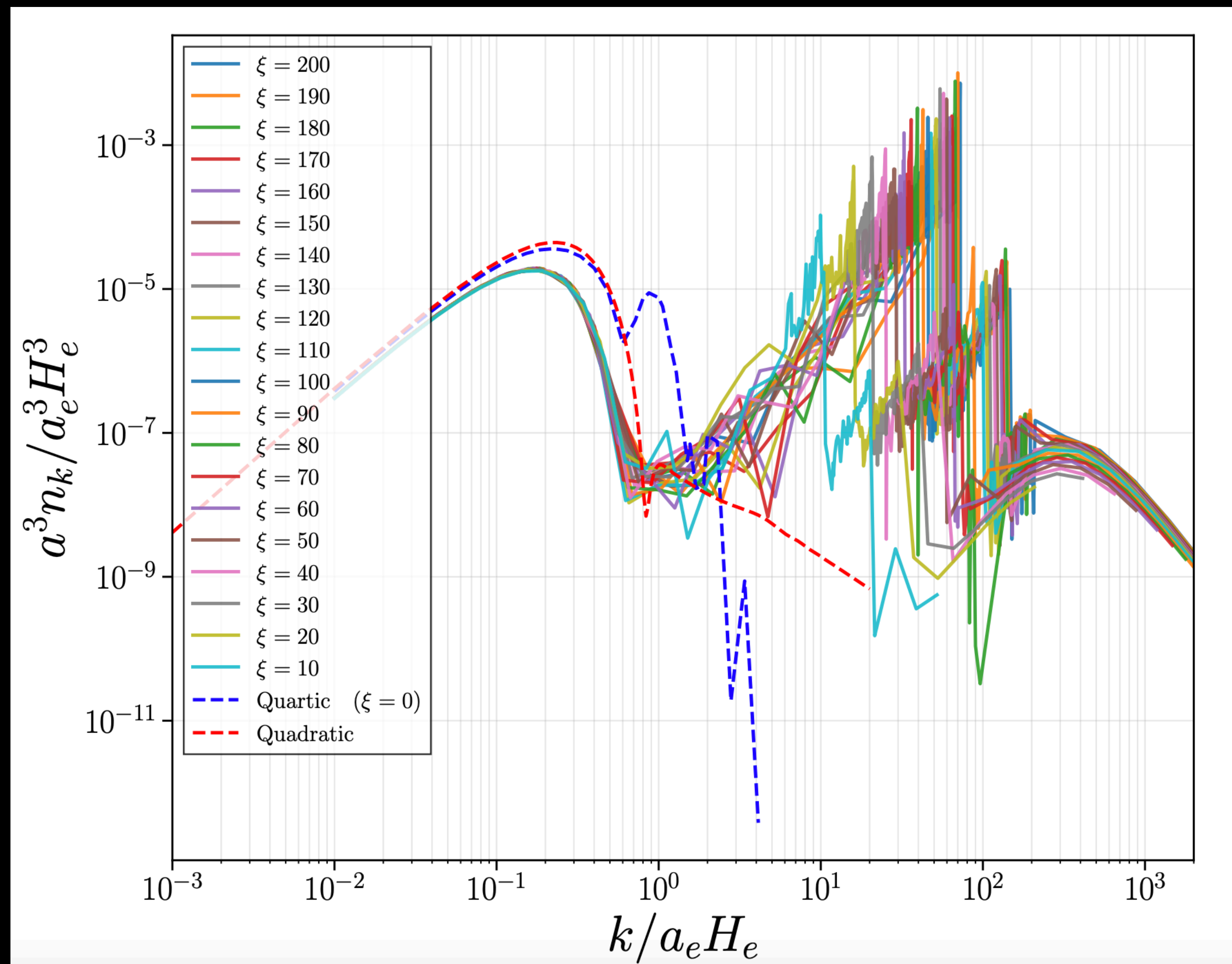
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# WIMPZillas in Higgs Inflation



# Higgs Inflation: A new peak emerges!

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Winnipeg

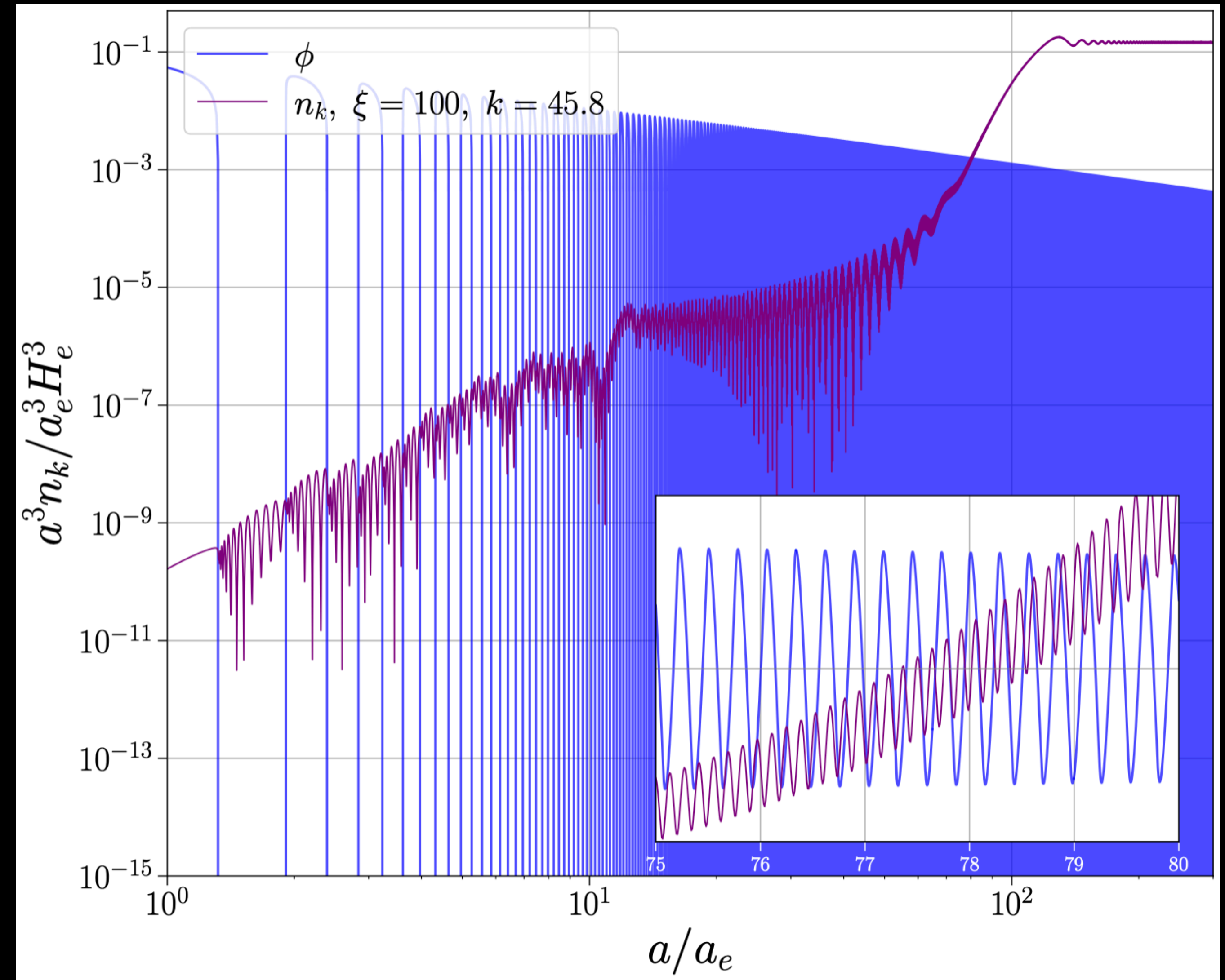
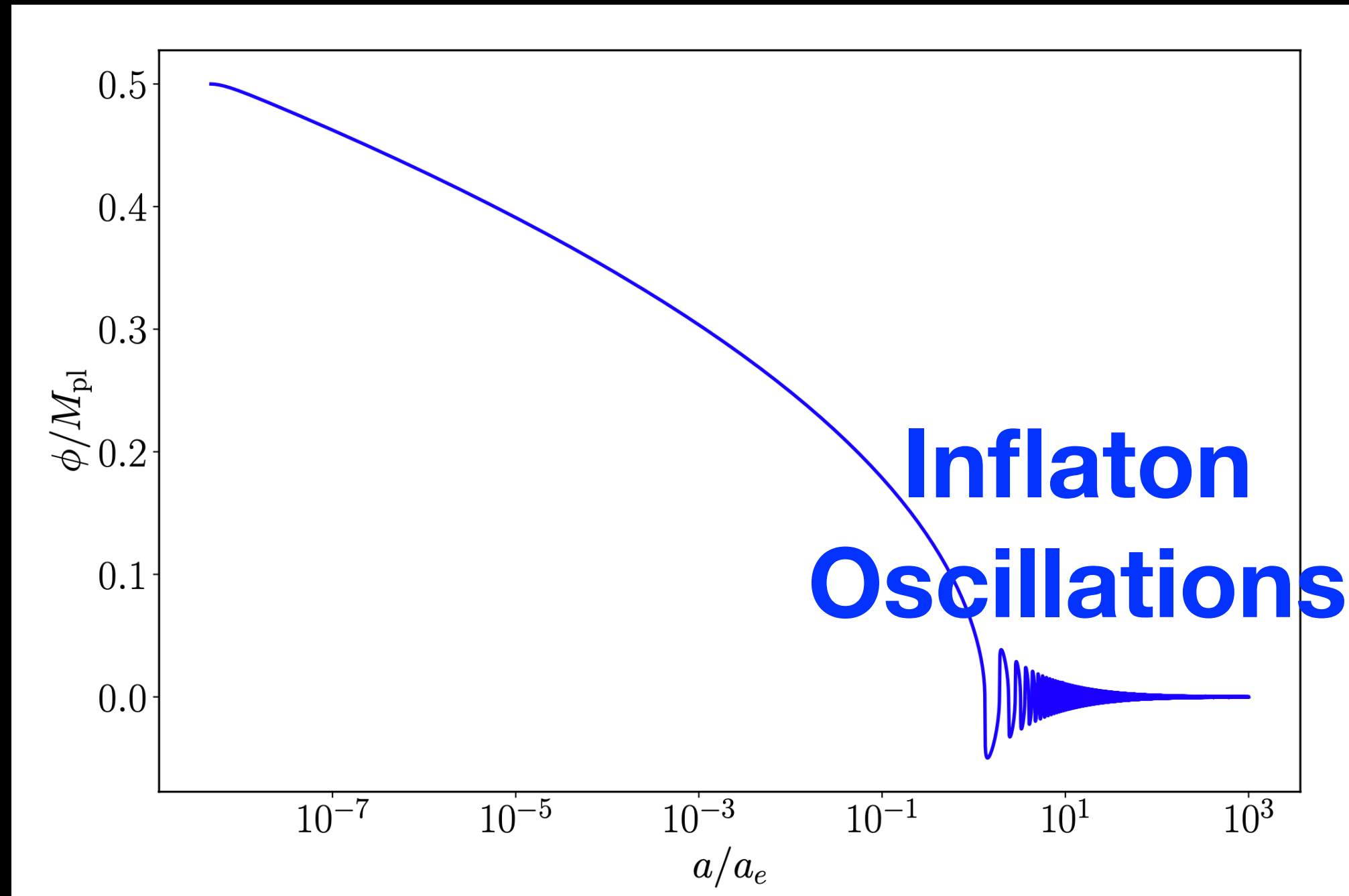


Chowdhury, Kolb, Jenks, EM  
JHEP '26

# Resonance

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## Particle Number Density

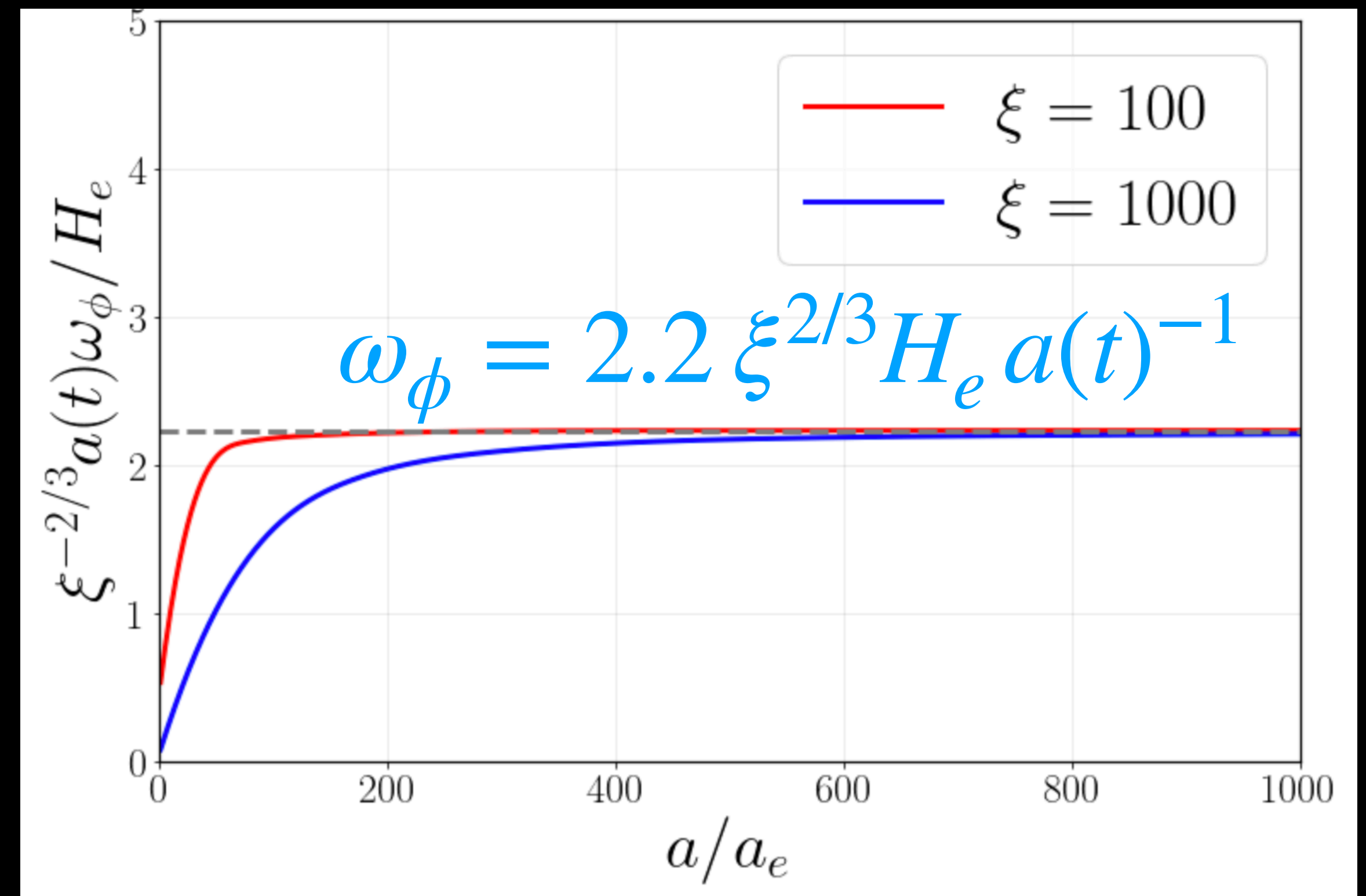
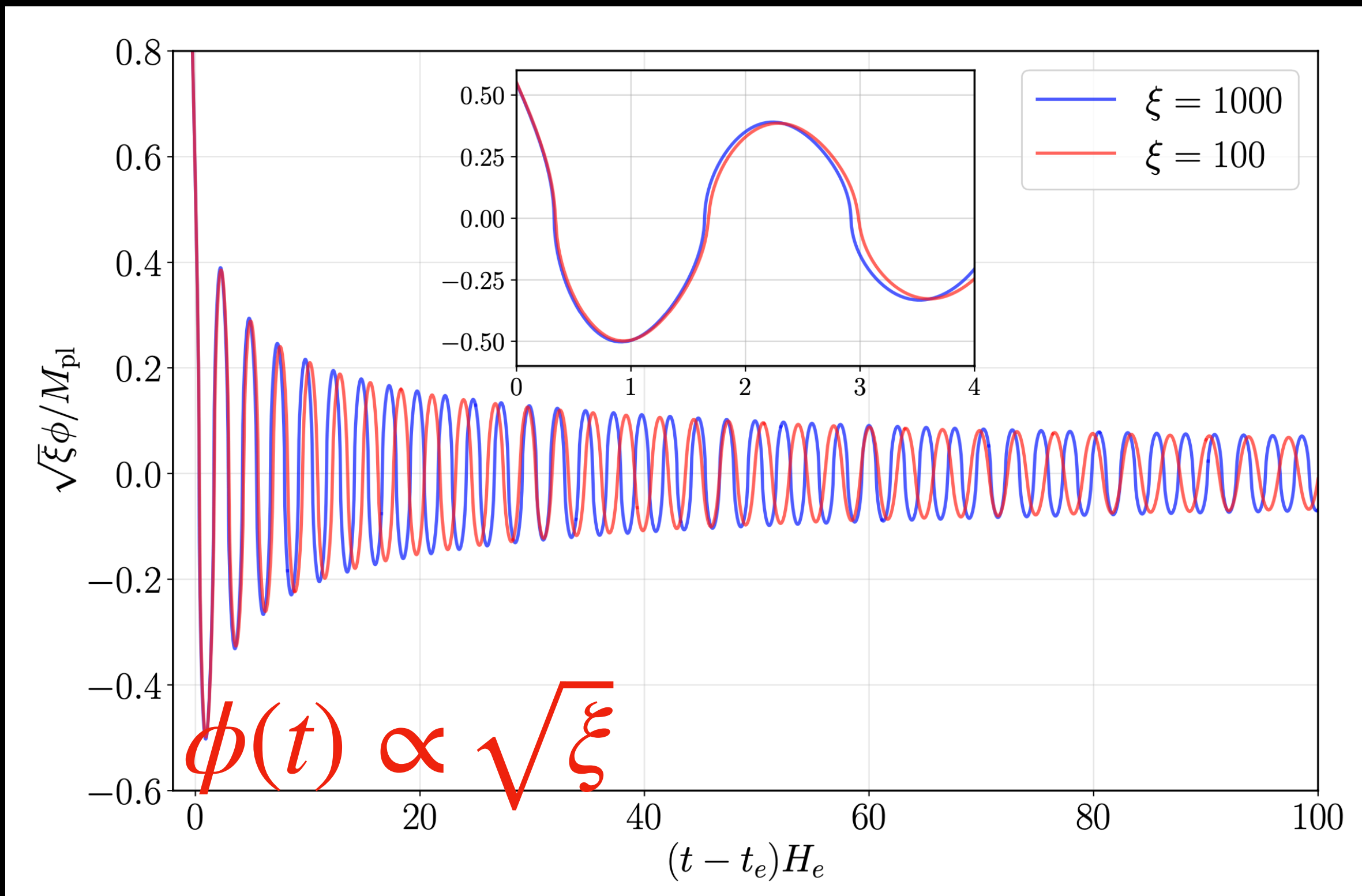


# Scaling symmetry

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Early Times:  
Universal  $\sqrt{\xi} \phi(t)$

Late times:  
Universal *frequency*

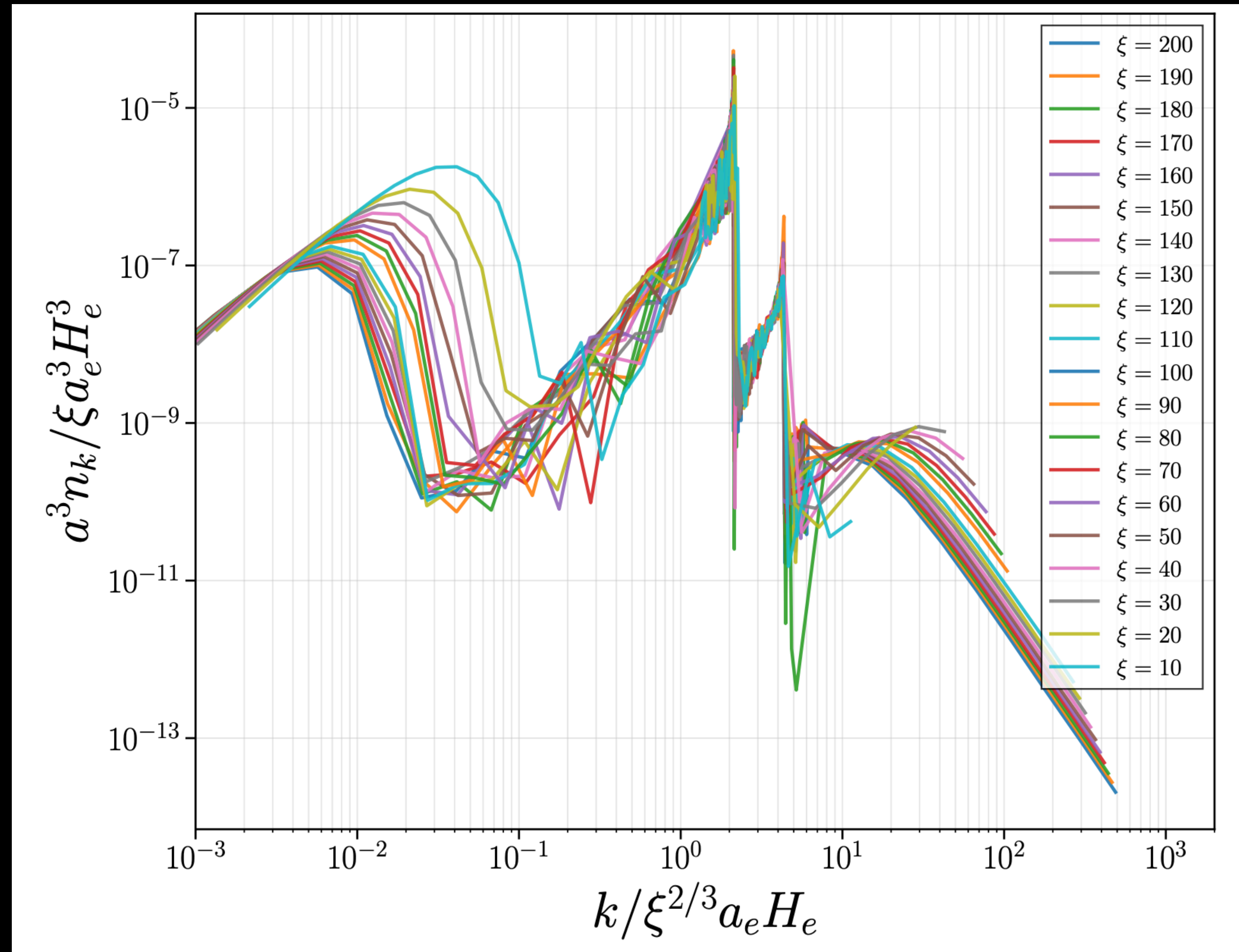


# Particle Production: Universal Structure

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Winnipeg

$$k_{\text{peak}} \approx 2 \xi^{2/3} H_e$$

Chowdhury, Kolb, Jenks, EM  
JHEP '26



# Dark Matter

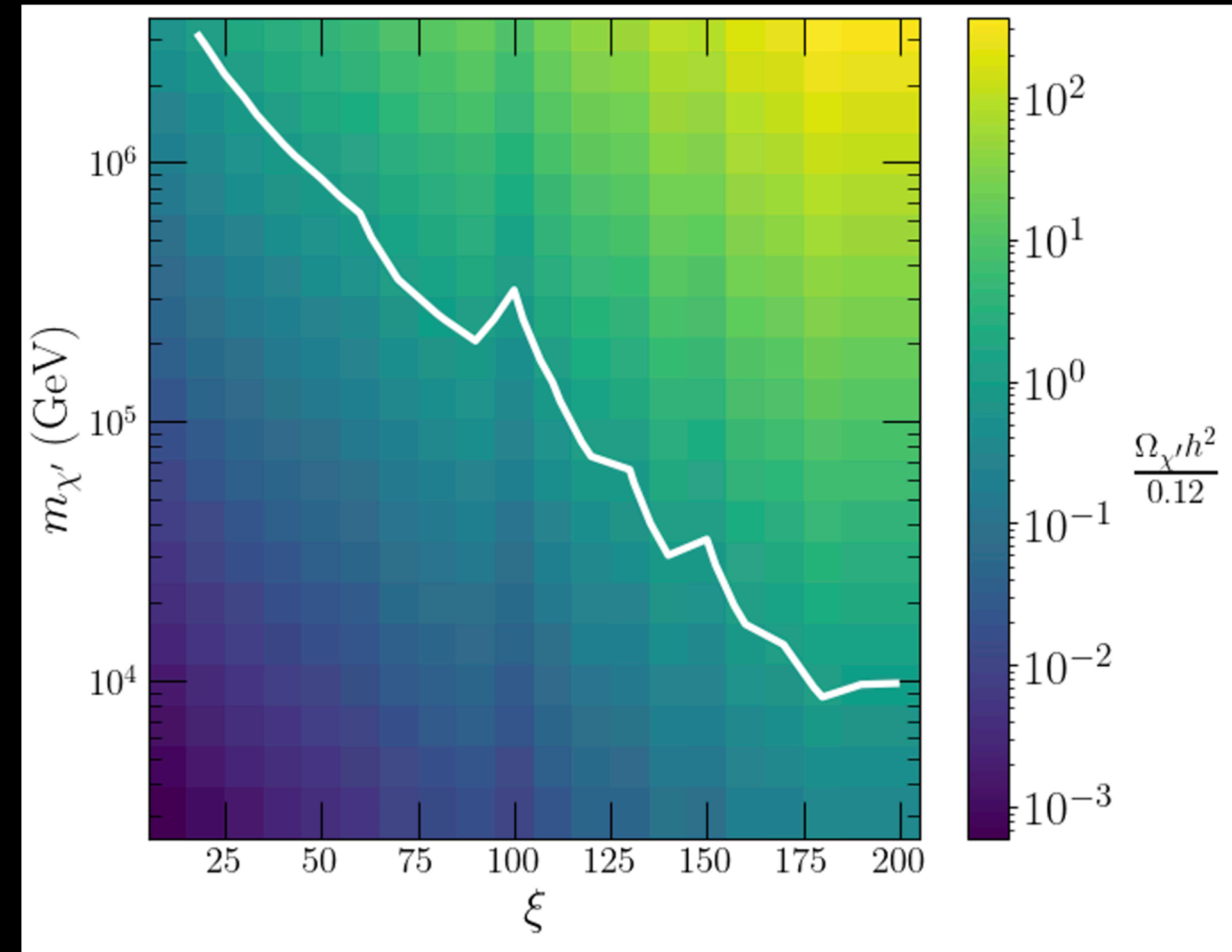
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Dark matter mass fixes  $\xi$ !

$$\frac{\Omega_\chi h^2}{0.12} = \frac{m_\chi}{H_e} \left( \frac{H_e}{10^{12} \text{GeV}} \right)^2 \left[ \frac{T_{\text{RH}}(\xi)}{10^9 \text{GeV}} \right] \frac{a^3 n}{10^{-5}},$$

+ CMB fixes  $\lambda(E_{\text{inf}})$

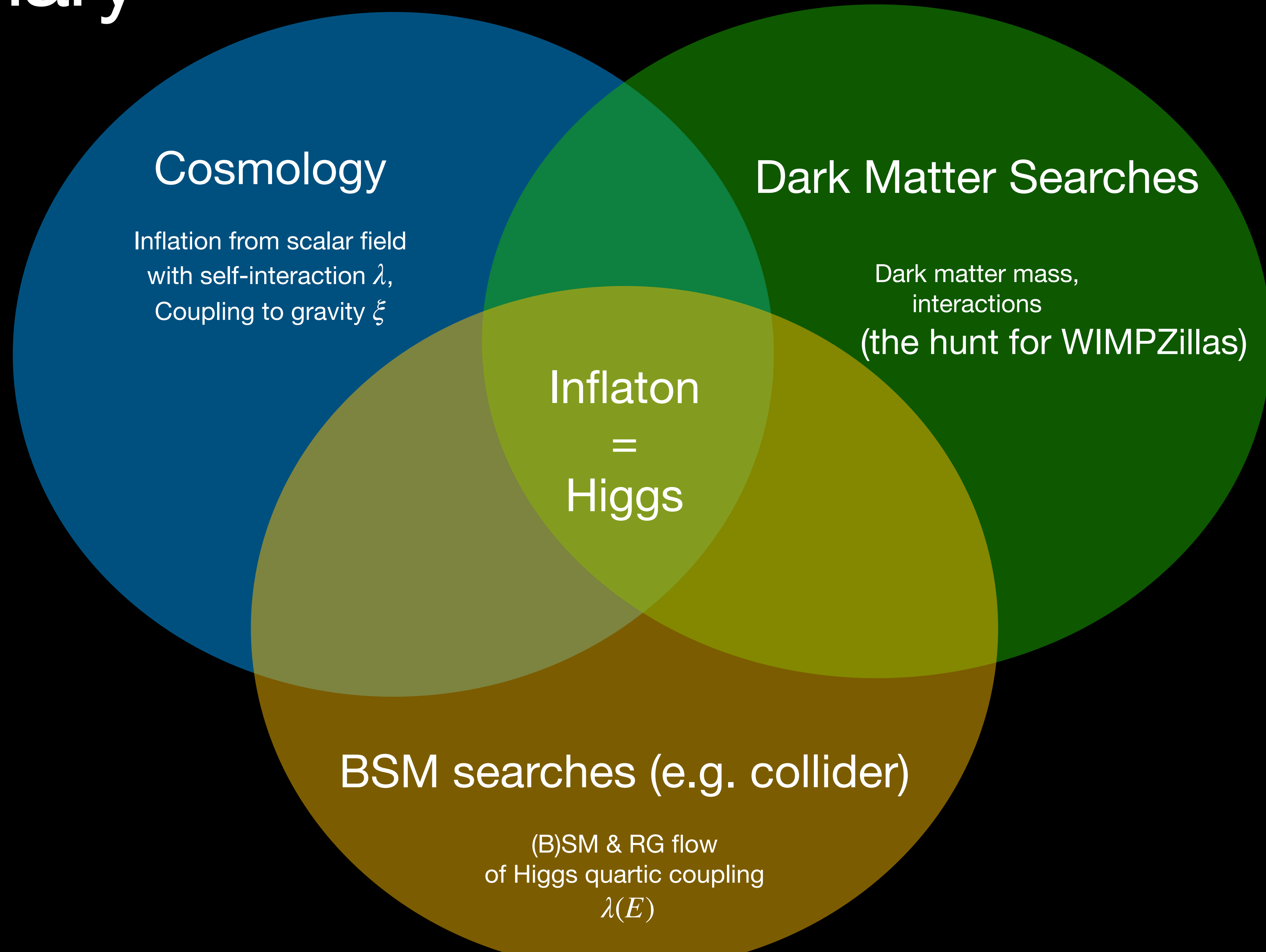
$$A_s \sim \frac{\lambda}{\xi^2} \sim 10^{-10} \Rightarrow \boxed{\xi \sim 10^5 \sqrt{\lambda(E_{\text{inf}})}}$$



Strongly restricts new physics!

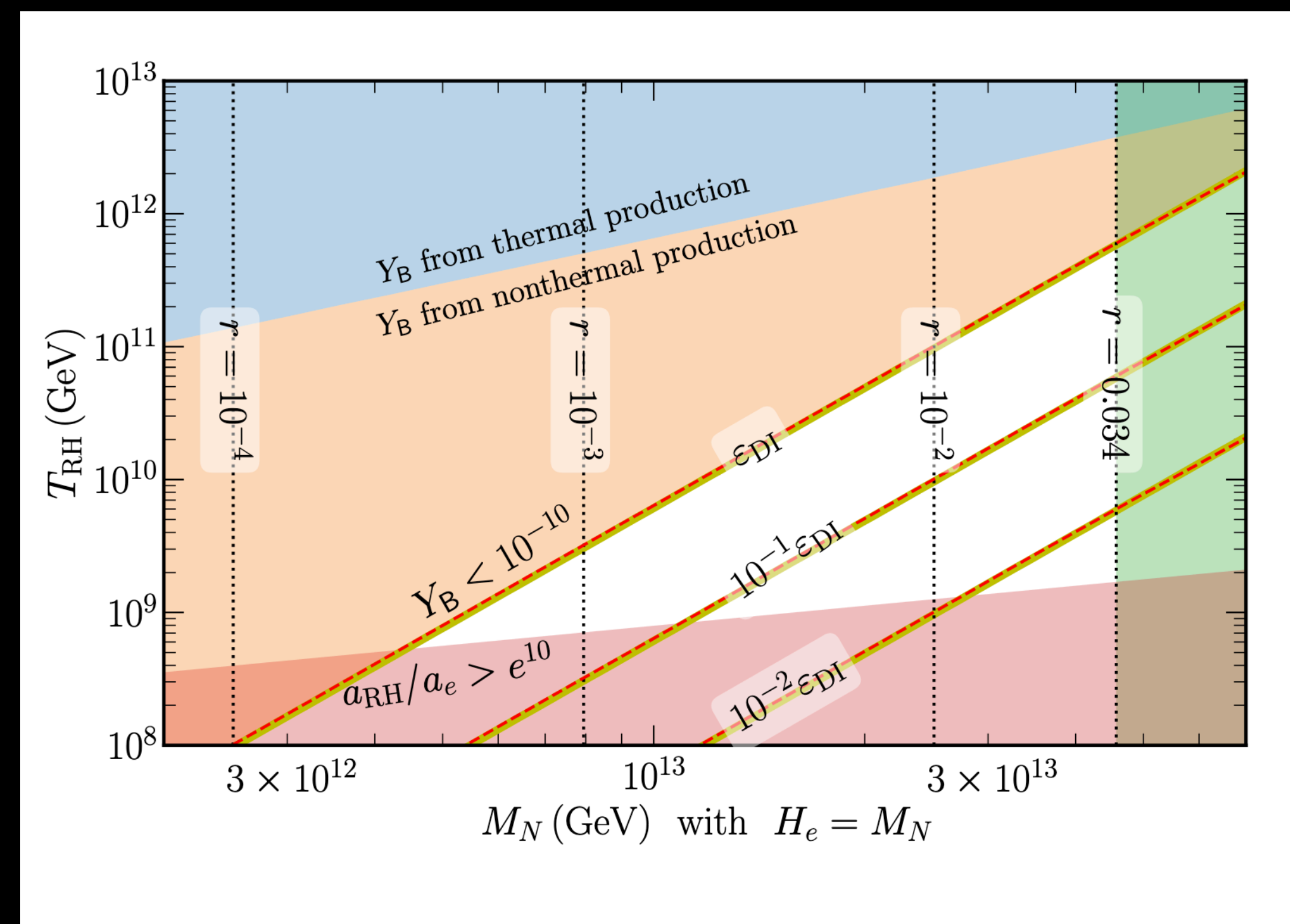
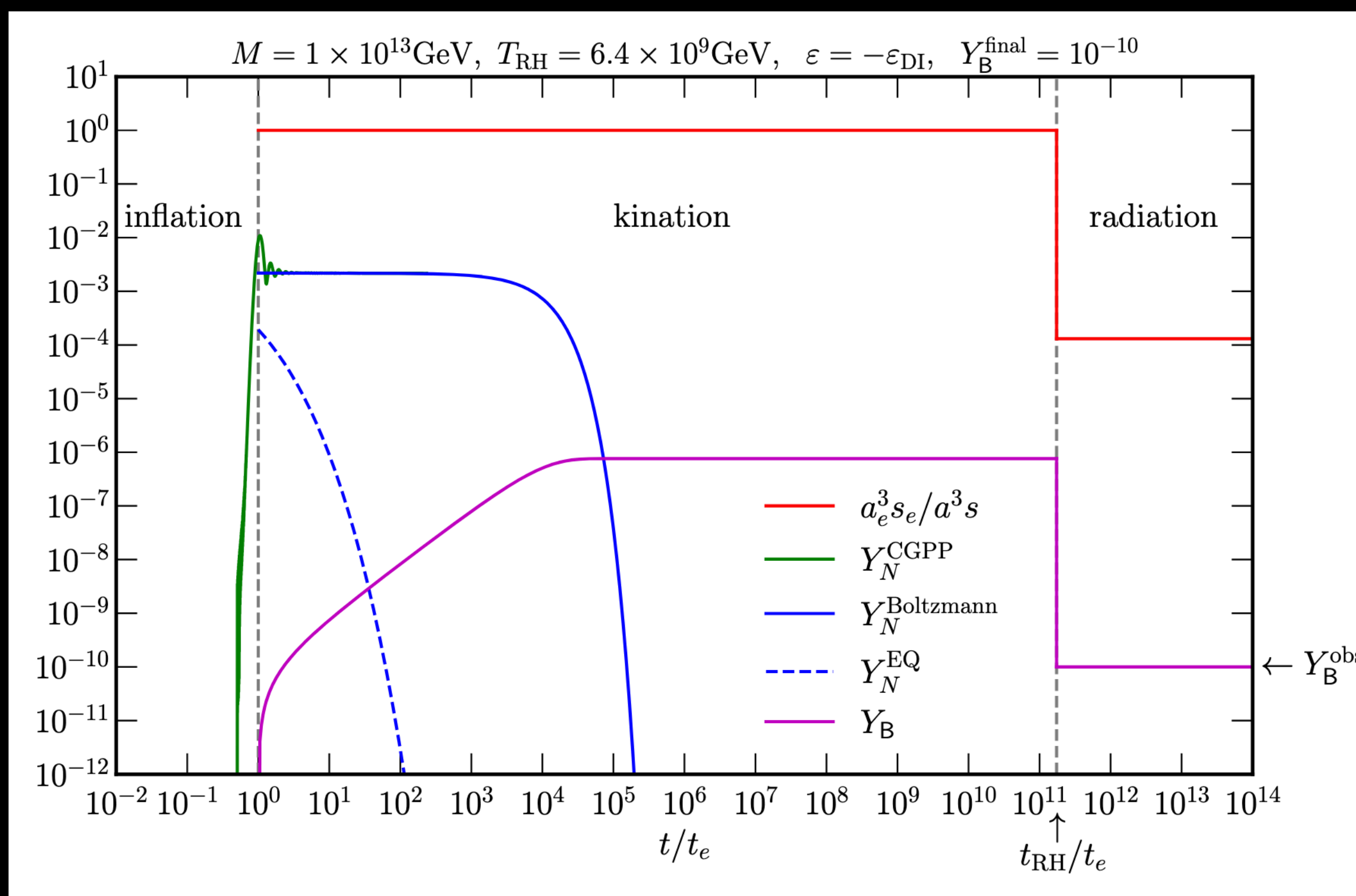
# Summary

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Winnipeg



# Hot off the press: *Non-Thermal Leptogenesis*

With Chowdhury, Jenks, Kolb, Long



# Thanks!

Evan McDonough  
UWinnipeg

