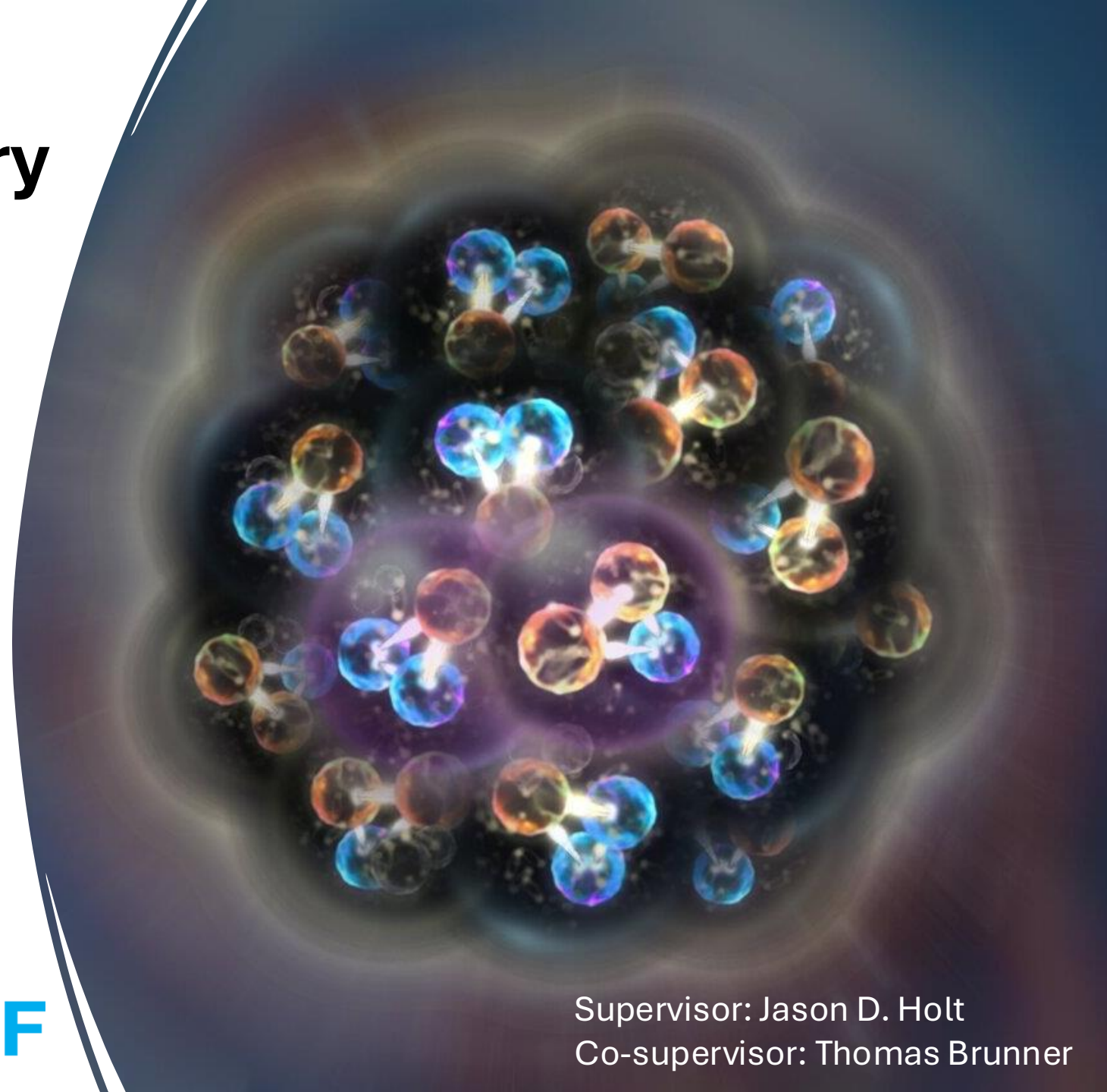


Ab Initio Nuclear Theory for Neutrinoless $\beta\beta$ Decay

Short-range Operators and Exotic
Mechanisms

Alex Todd (McGill University)

CAP Congress 2026



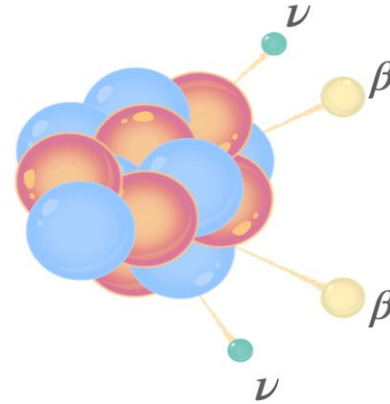
Supervisor: Jason D. Holt
Co-supervisor: Thomas Brunner

What is neutrinoless $\beta\beta$ decay?

$0\nu\beta\beta$ probes:

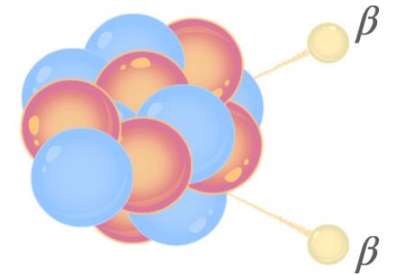
- Majorana/Dirac nature of neutrinos
- Lepton-number violation
 - Baryon asymmetry of universe
- Absolute neutrino mass scale
- Exotic BSM mechanisms
 - Heavy neutrinos? Seesaw mechanisms? Sterile neutrinos?

$2\nu\beta\beta$



(Observed)

$0\nu\beta\beta$



vs

(Hypothetical)

What is neutrinoless $\beta\beta$ decay?

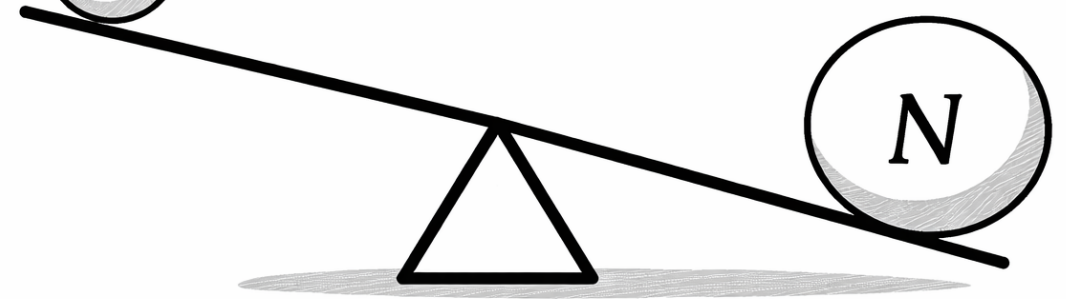
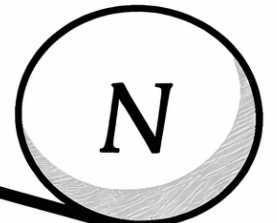
$0\nu\beta\beta$ probes:

- Majorana/Dirac nature of neutrinos
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Standard Model
neutrinos



Hypothetical heavy
neutrinos



What is neutrinoless $\beta\beta$ decay?

$0\nu\beta\beta$ probes:

- Majorana/Dirac nature of neutrinos
- Lepton-number violation
 - Baryon asymmetry of universe
- Absolute neutrino mass scale
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 - Heavy neutrinos? Seesaw mechanisms? Sterile neutrinos?

An observation of $0\nu\beta\beta$ would immediately confirm these



Interpretations rely on nuclear theory

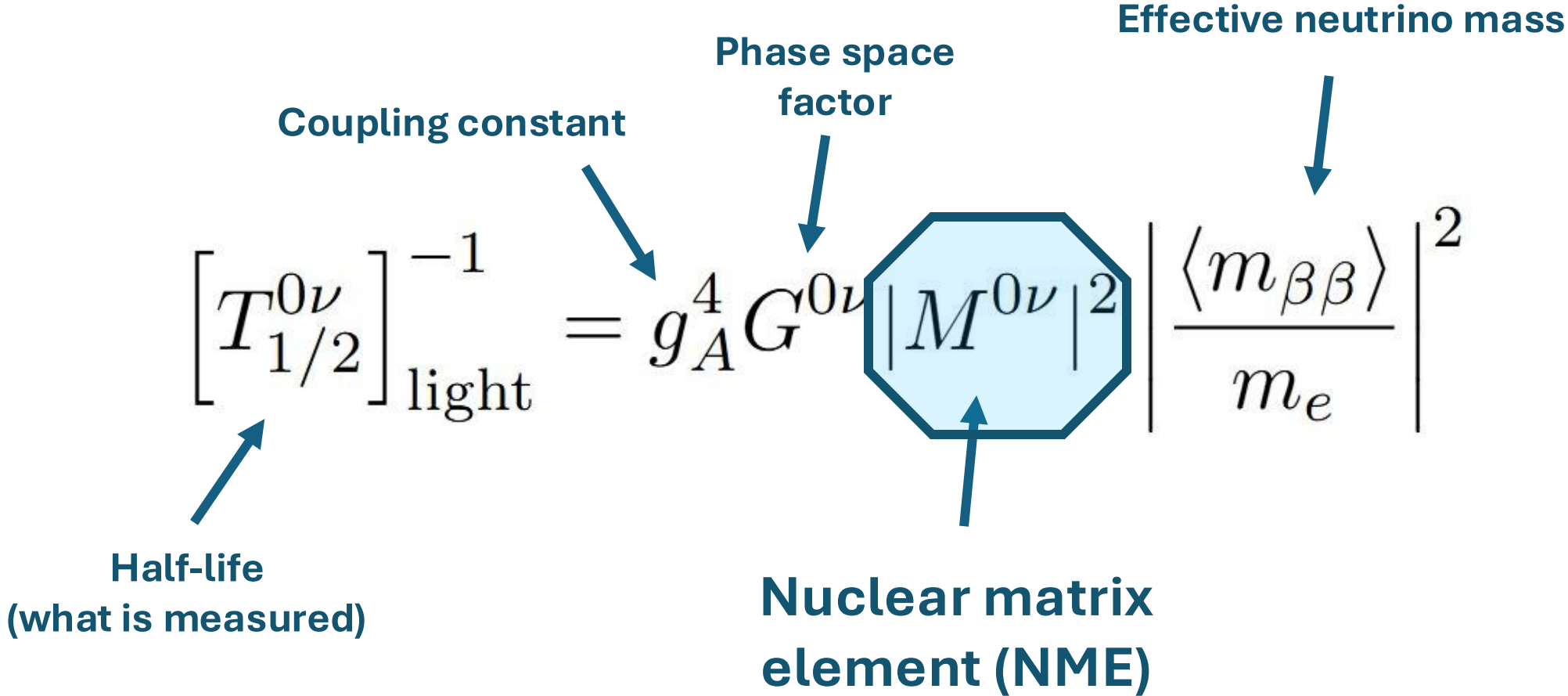


Decay rate

$$\left[T_{1/2}^{0\nu} \right]_{\text{light}}^{-1} = g_A^4 G^{0\nu} |M^{0\nu}|^2 \left| \frac{\langle m_{\beta\beta} \rangle}{m_e} \right|^2$$

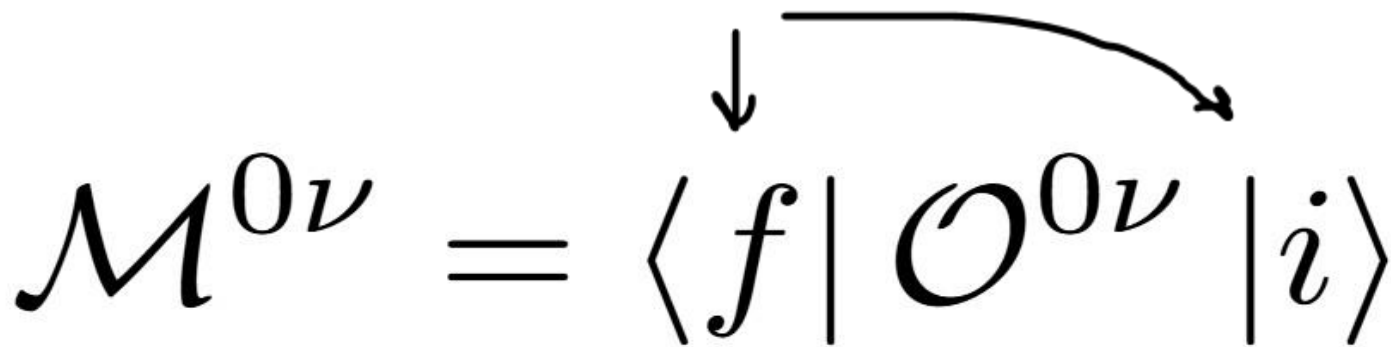
Decay rate

$$\langle m_{\beta\beta} \rangle = \left| \sum_{i=1}^3 U_{ei} m_i \right|$$



Nuclear Matrix Element

Initial and final states

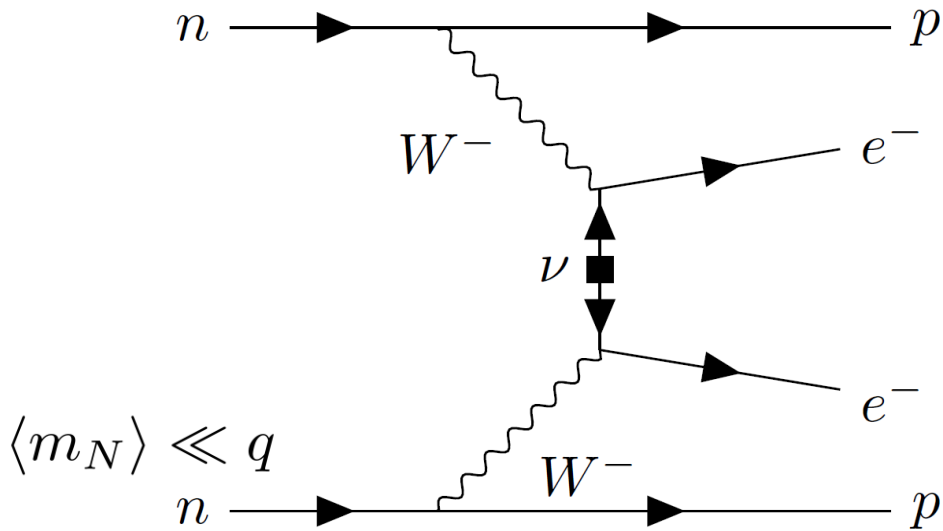
$$\mathcal{M}^{0\nu} = \langle f | \mathcal{O}^{0\nu} | i \rangle$$


Decay operator encodes specific mechanisms

Decay Mechanism Example

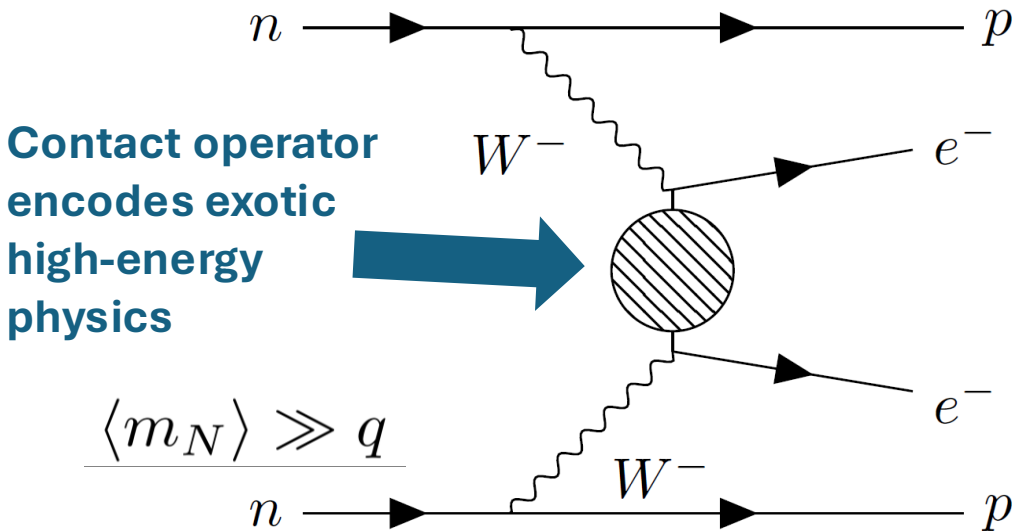
Light neutrino exchange
(long-range)

$$\left[T_{1/2}^{0\nu} \right]_{\text{light}}^{-1} = g_A^4 G^{0\nu} |M^{0\nu}|^2 \left(\frac{m_{\beta\beta}}{m_e} \right)^2$$



Heavy neutrino exchange
(short-range)

$$\left[T_{1/2}^{0\nu} \right]_{\text{heavy}}^{-1} = \text{More complicated expressions}$$



Decay Mechanism Example

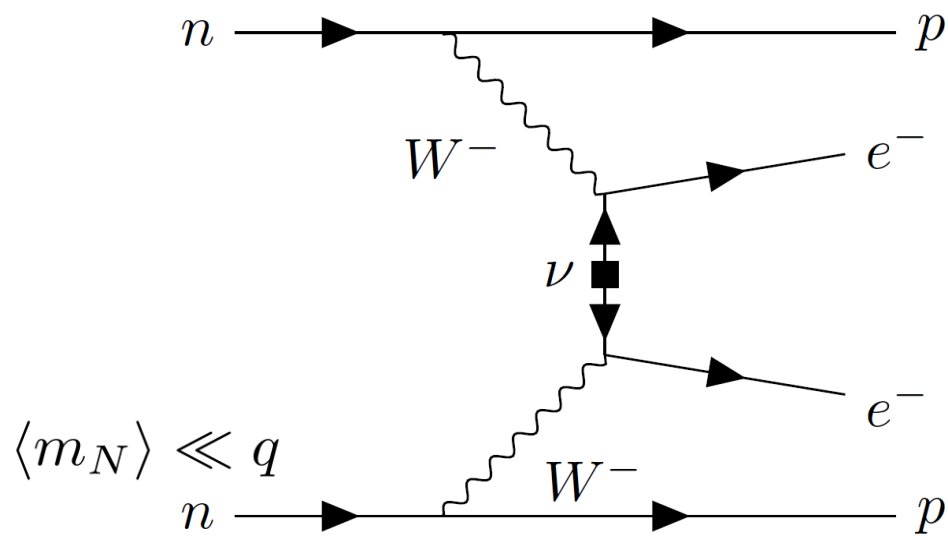
Relatively unexplored in nuclear theory

Light neutrino exchange
(long-range)

$$\left[T_{1/2}^{0\nu} \right]_{\text{light}}^{-1} = g_A^4 G^{0\nu} |M^{0\nu}|^2 \left(\frac{m_{\beta\beta}}{m_e} \right)^2$$

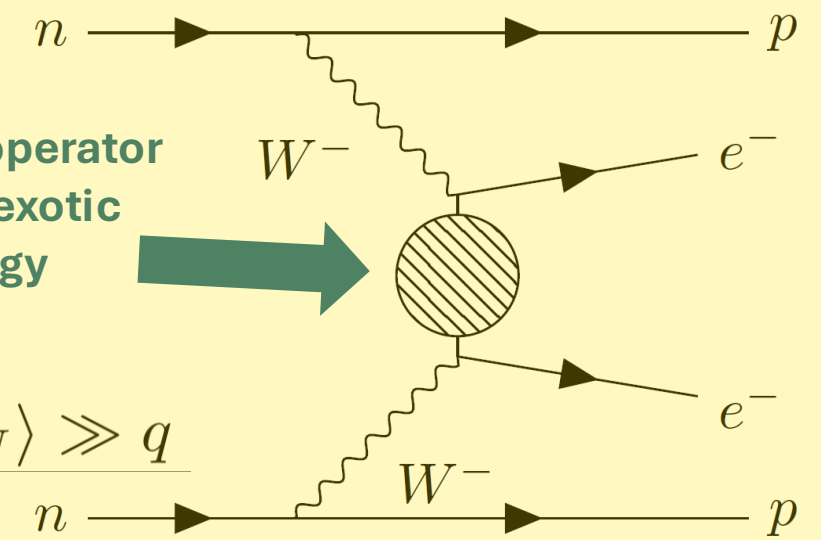
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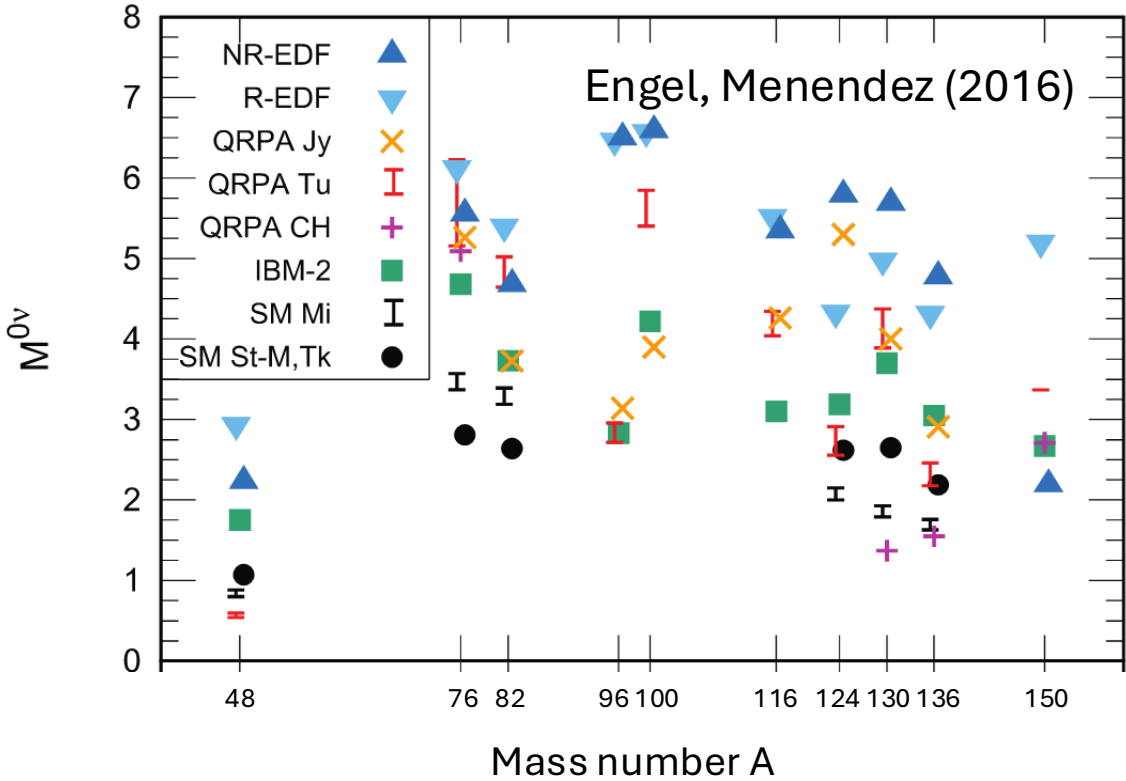
Contact operator encodes exotic high-energy physics

$$\langle m_N \rangle \gg q$$



NME uncertainty

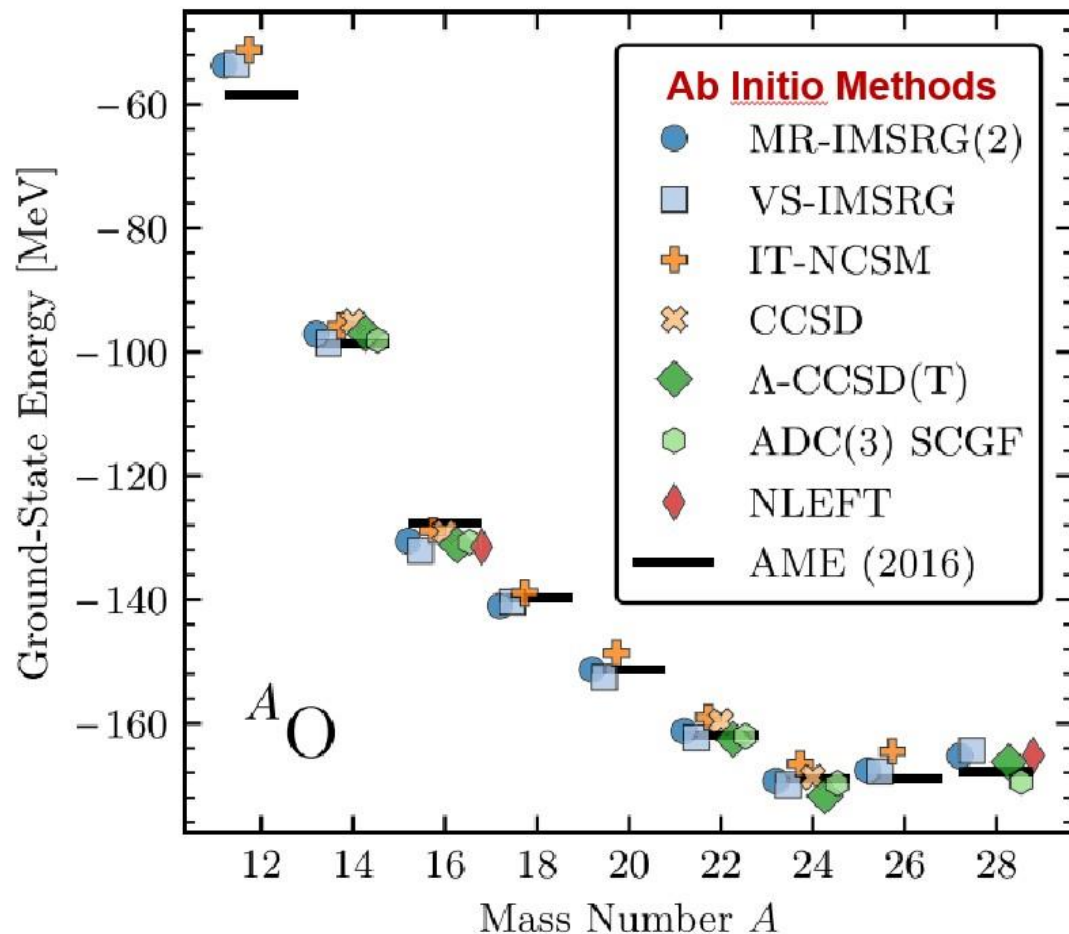
Phenomenological Methods (fit to data)



Large NME uncertainty is not good for interpreting BSM physics

NME uncertainty

Ab initio Nuclear Theory (from first-principles)

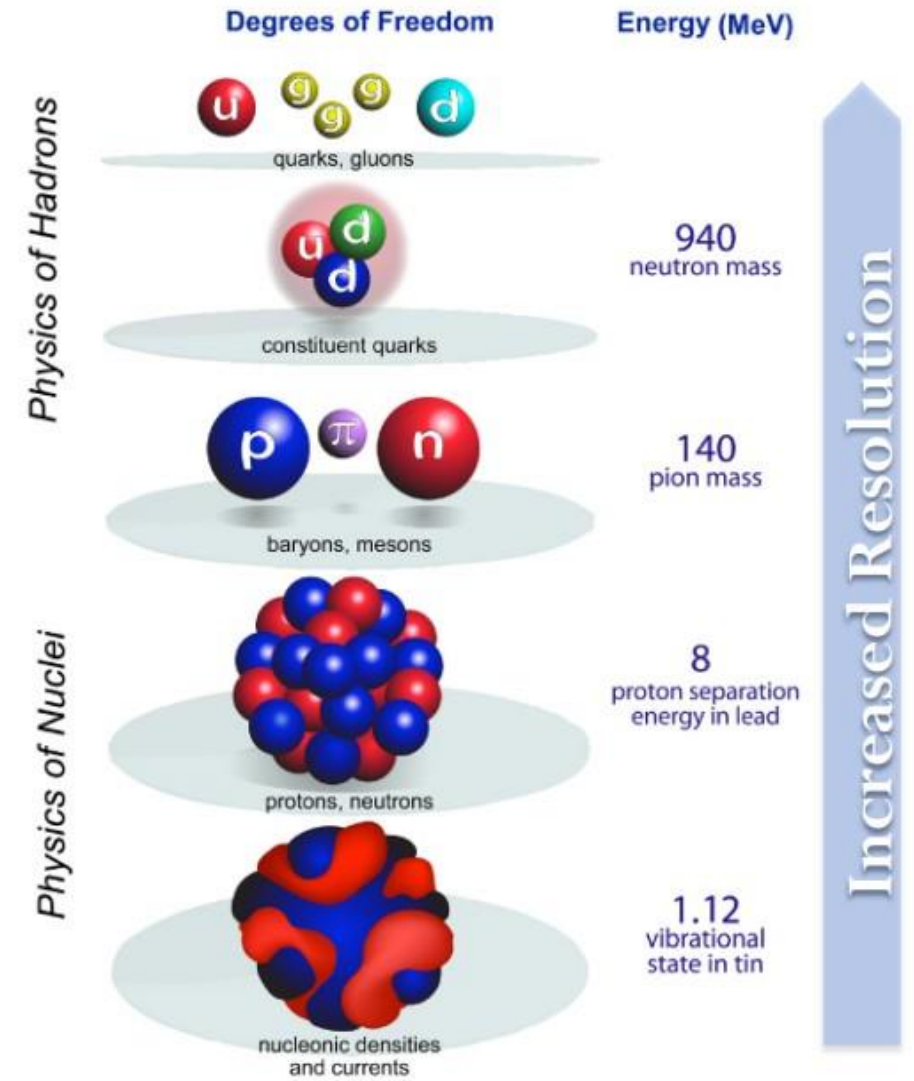


Nuclear Matrix Element



Why are the NMEs so hard to calculate?

The nucleus is complicated



Phenomenological nuclear models

Fit Hamiltonian to data,
then predict

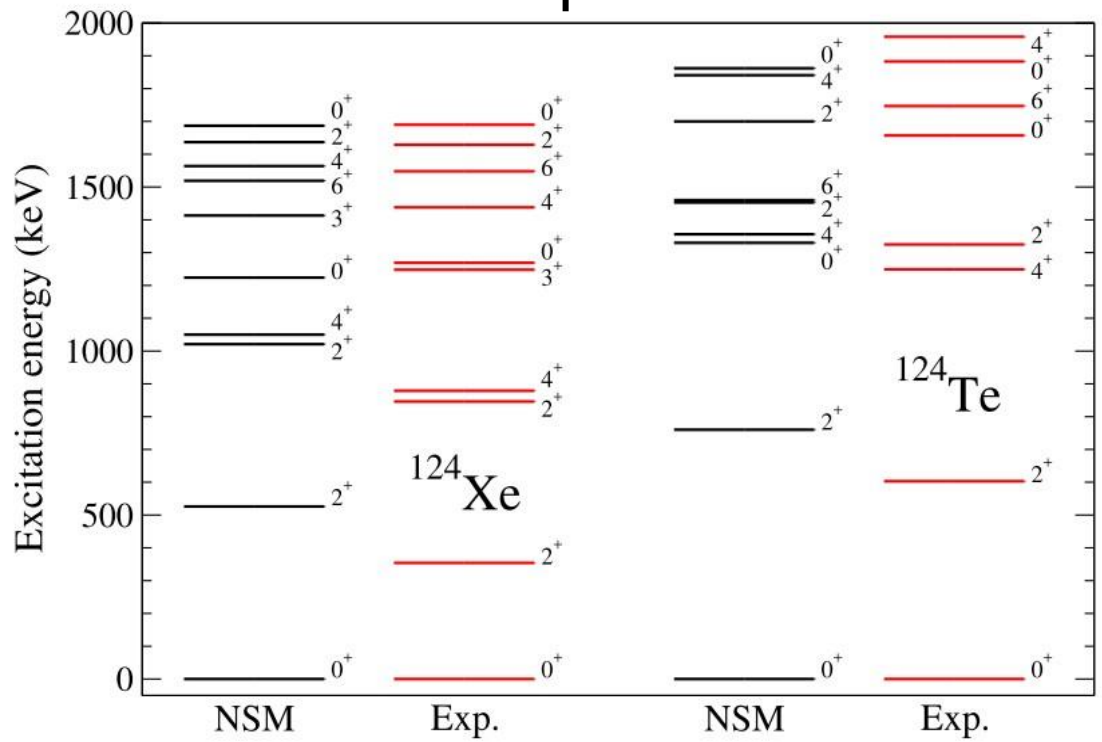
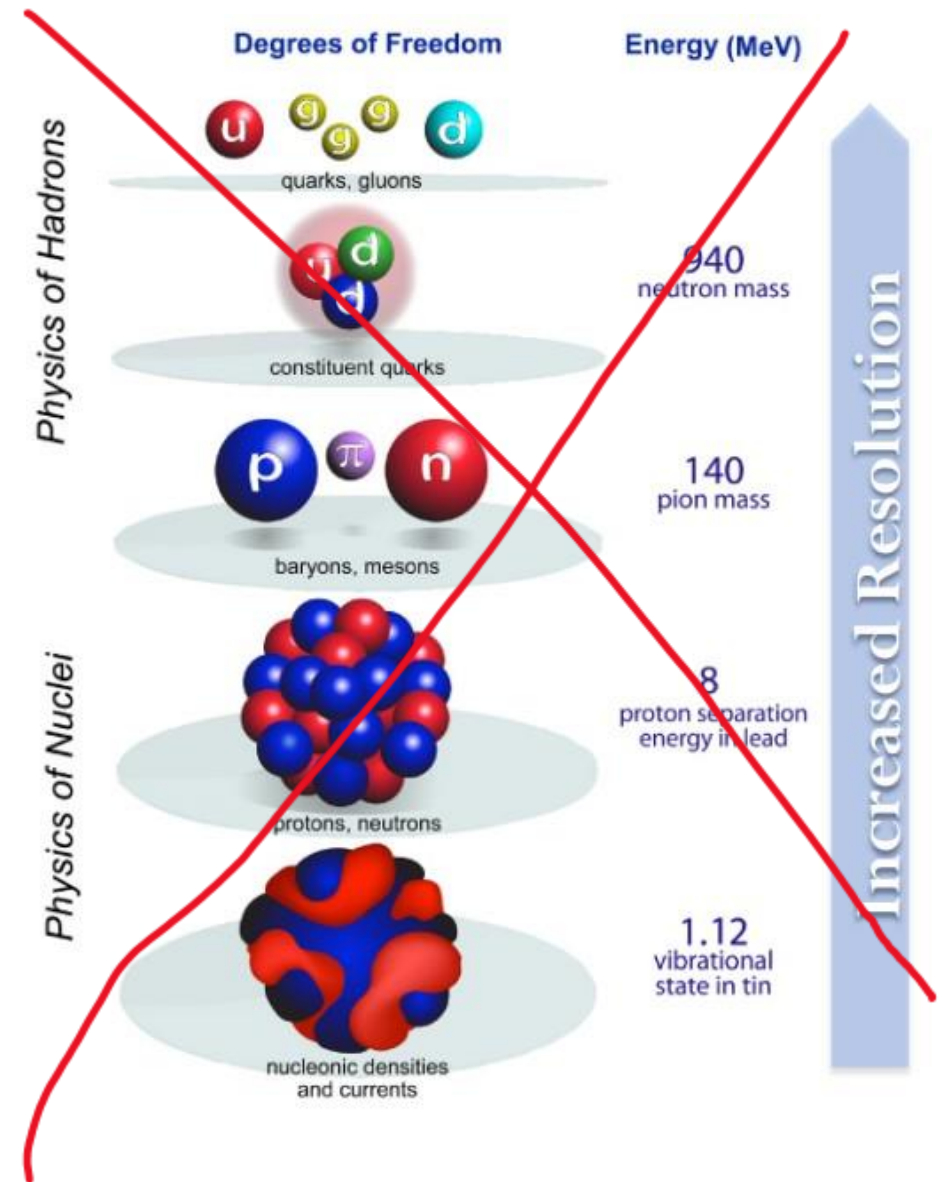


Fig. 1. ^{124}Xe and ^{124}Te excitation spectra obtained by the nuclear shell model (NSM) compared to experiment [74].

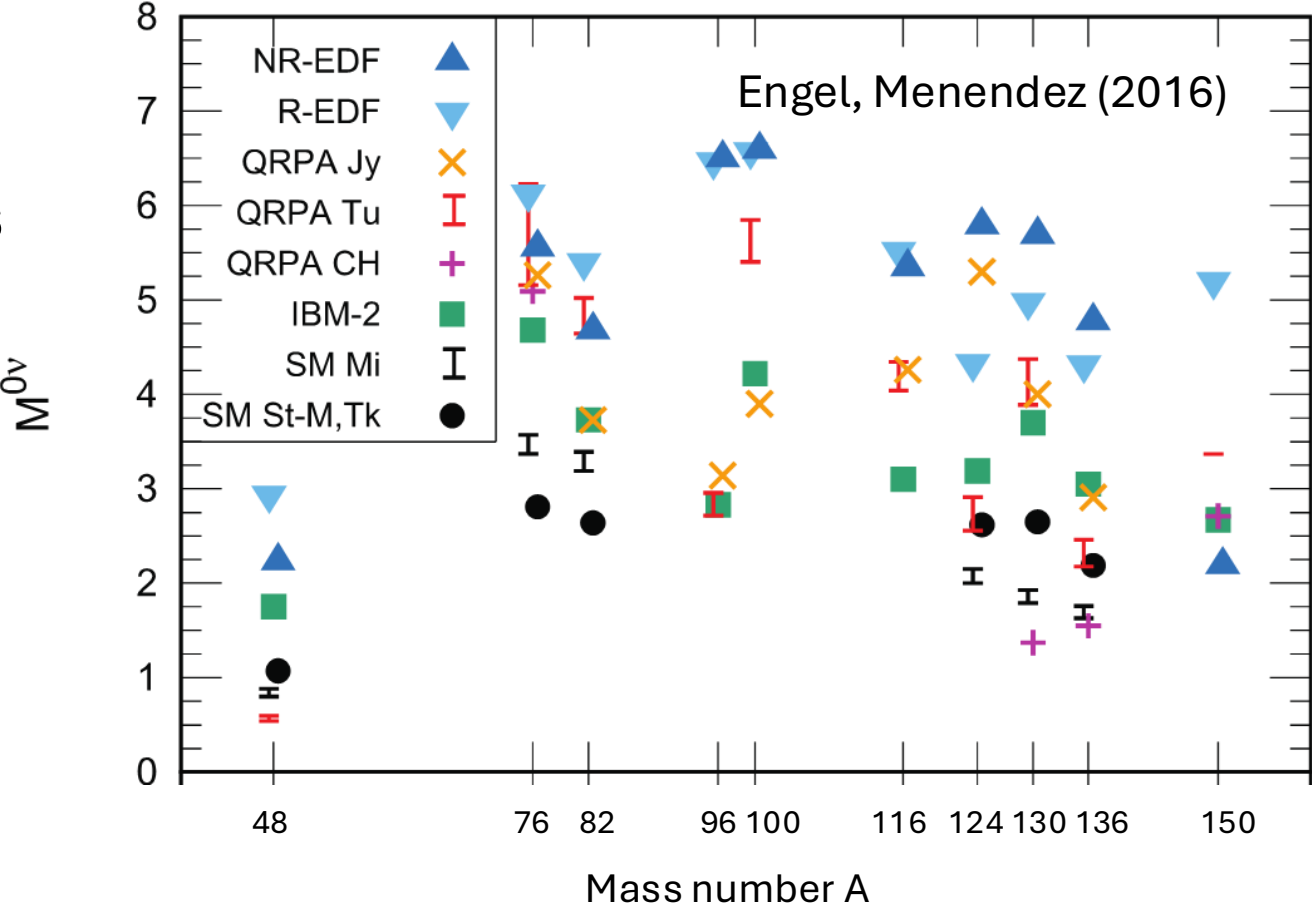
E.A. Coello Pérez, J. Menéndez, A. Schwenk (2019)



Issues with phenomenological methods

Methods are **inappropriately extrapolated** to where there is no data...

- Large spread in results
- No path for understanding uncertainty

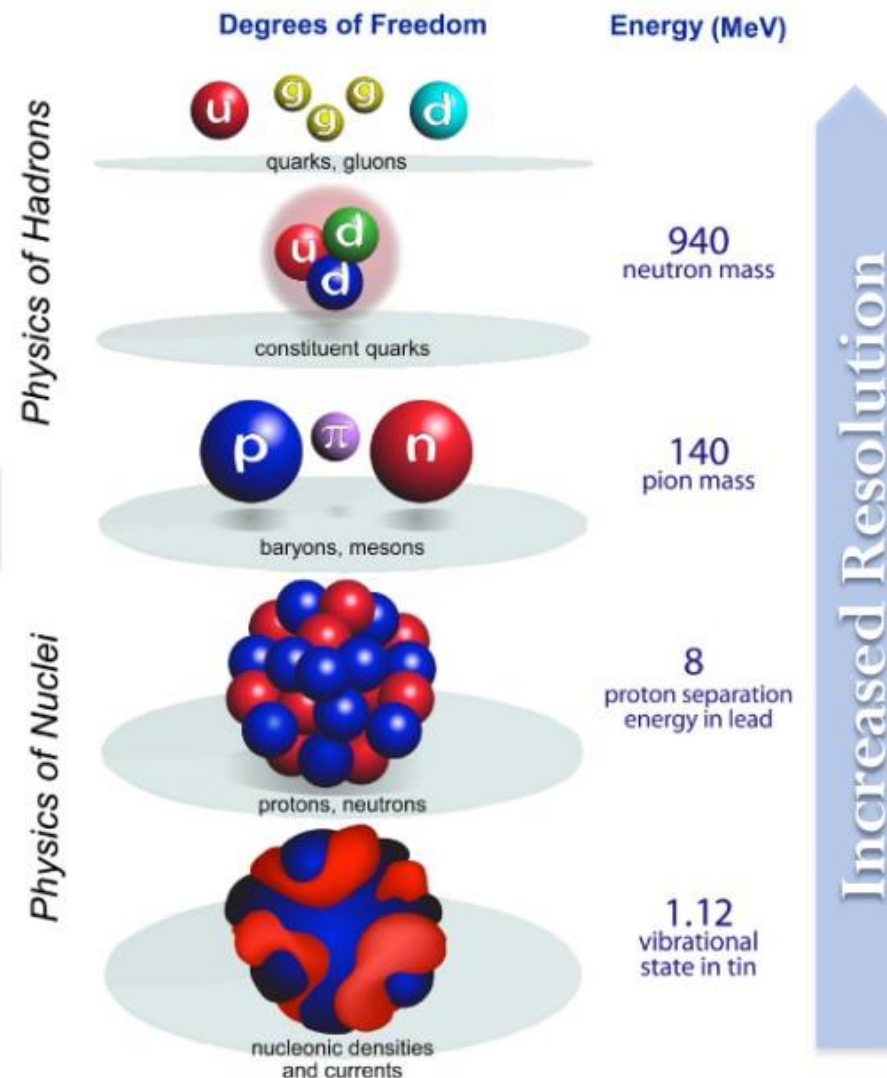


We can do better!

Renormalization group:

Effective description of higher-scale physics

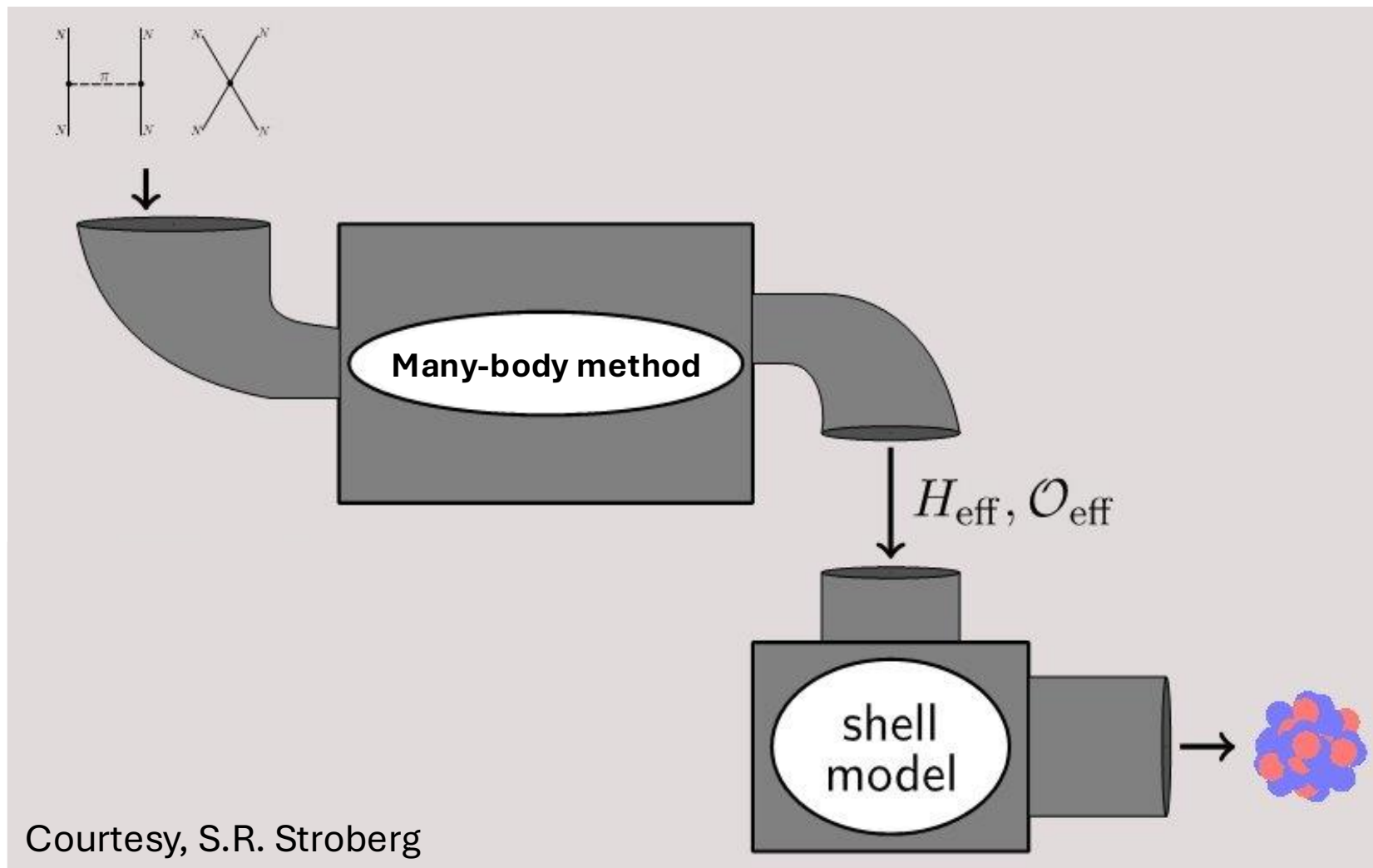
$$H(\Lambda) = T + V_{NN}(\Lambda) + V_{3N}(\Lambda) + V_{4N}(\Lambda) + \dots$$



Ab initio nuclear theory

= chiral EFT interactions
+ a (polynomially scaling)
many-body method

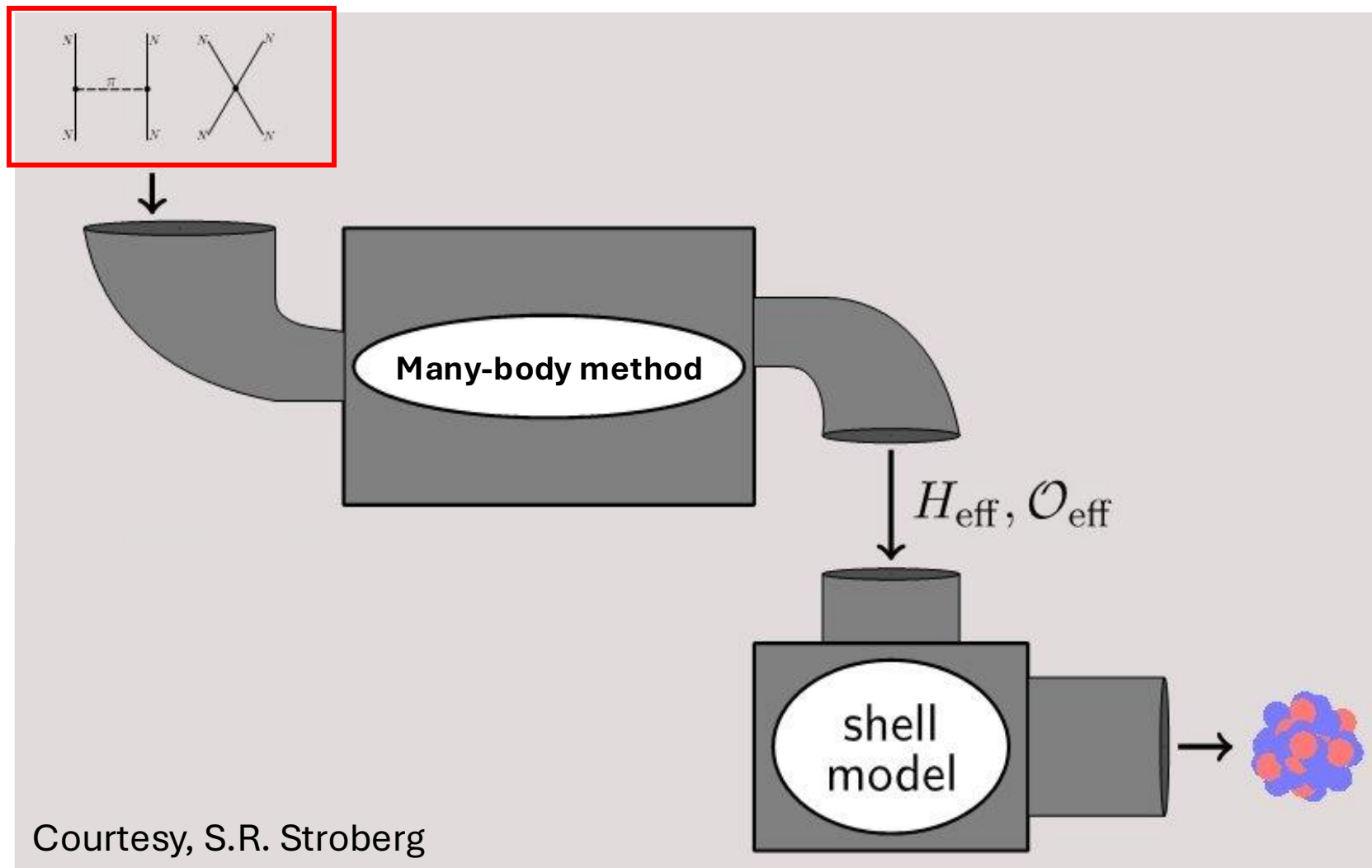
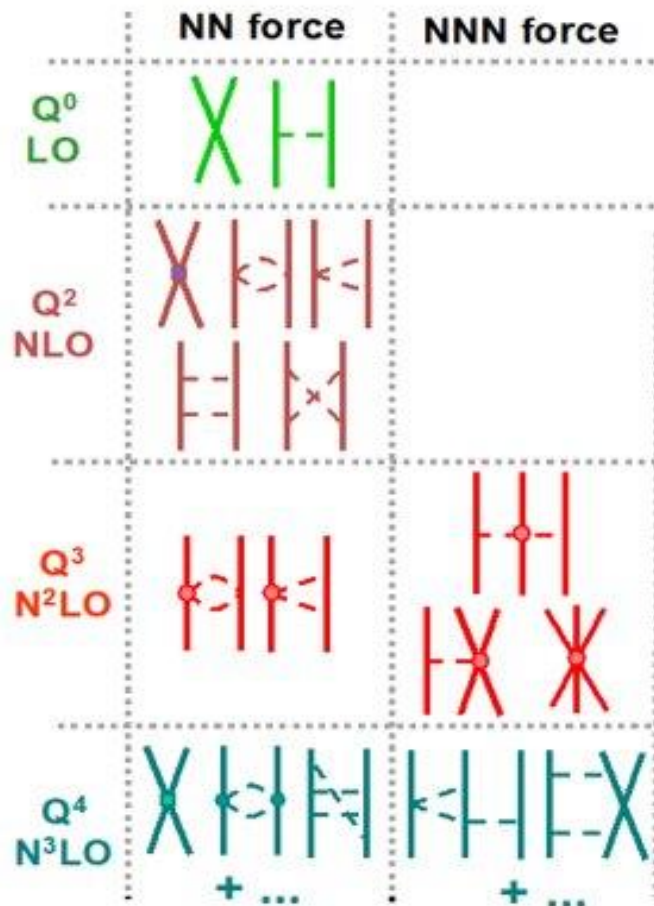
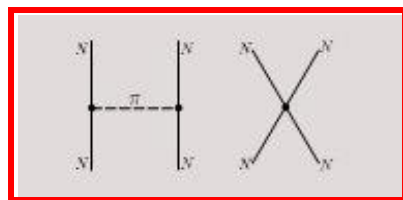
$$H\psi_n = E_n\psi_n$$



Ab initio nuclear theory

= **chiral EFT interactions**
 + a (polynomially scaling)
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$$H\psi_n = E_n\psi_n$$



Ab initio nuclear theory

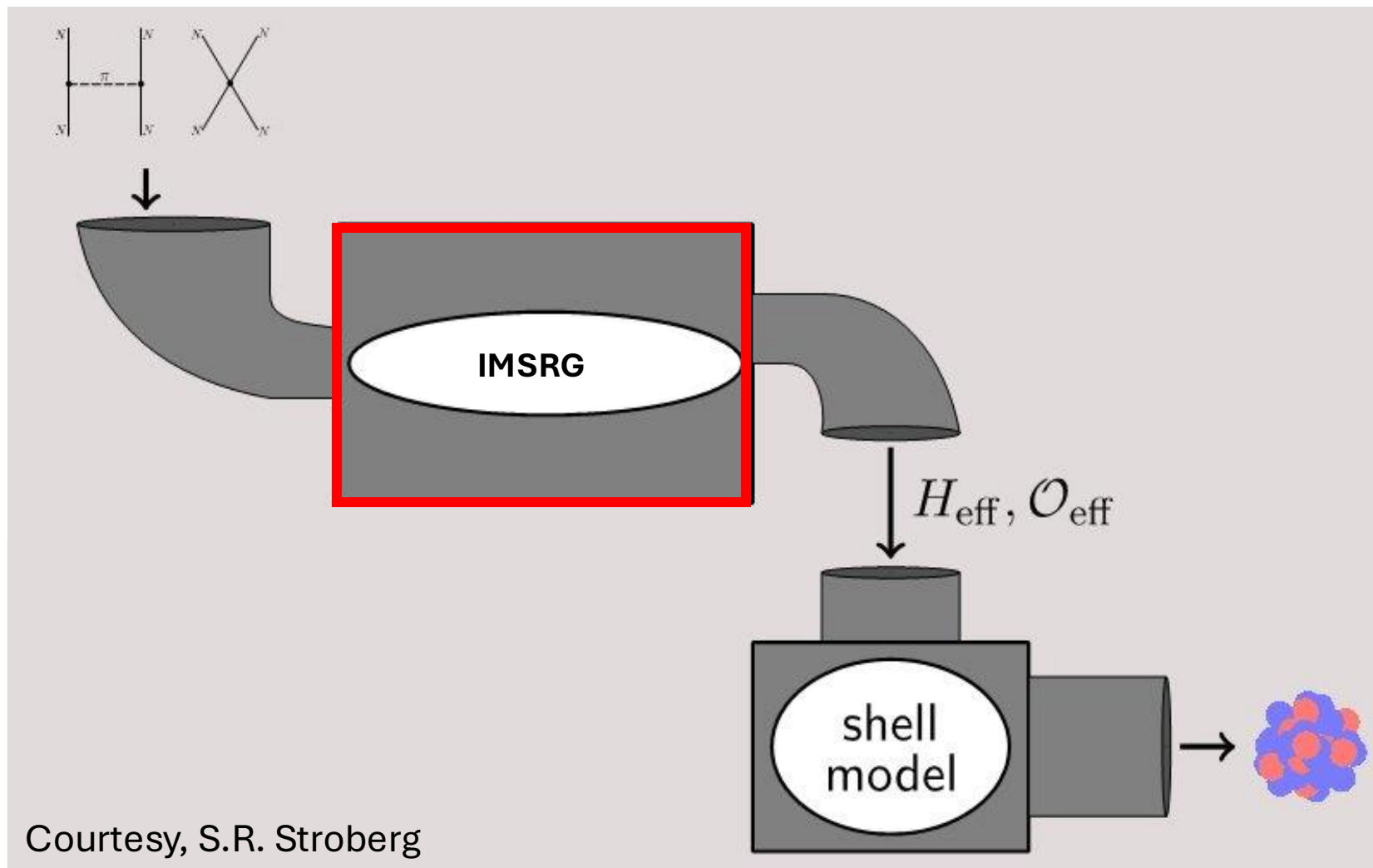
= chiral EFT interactions
+ **a (polynomially scaling)
many-body method**

$$H(s) = e^{\Omega(s)} H e^{-\Omega(s)}$$

$$\mathcal{O}(s) = e^{\Omega(s)} \mathcal{O} e^{-\Omega(s)}$$

**Continuous unitary
transformations of
Hamiltonian**

$$H\psi_n = E_n\psi_n$$

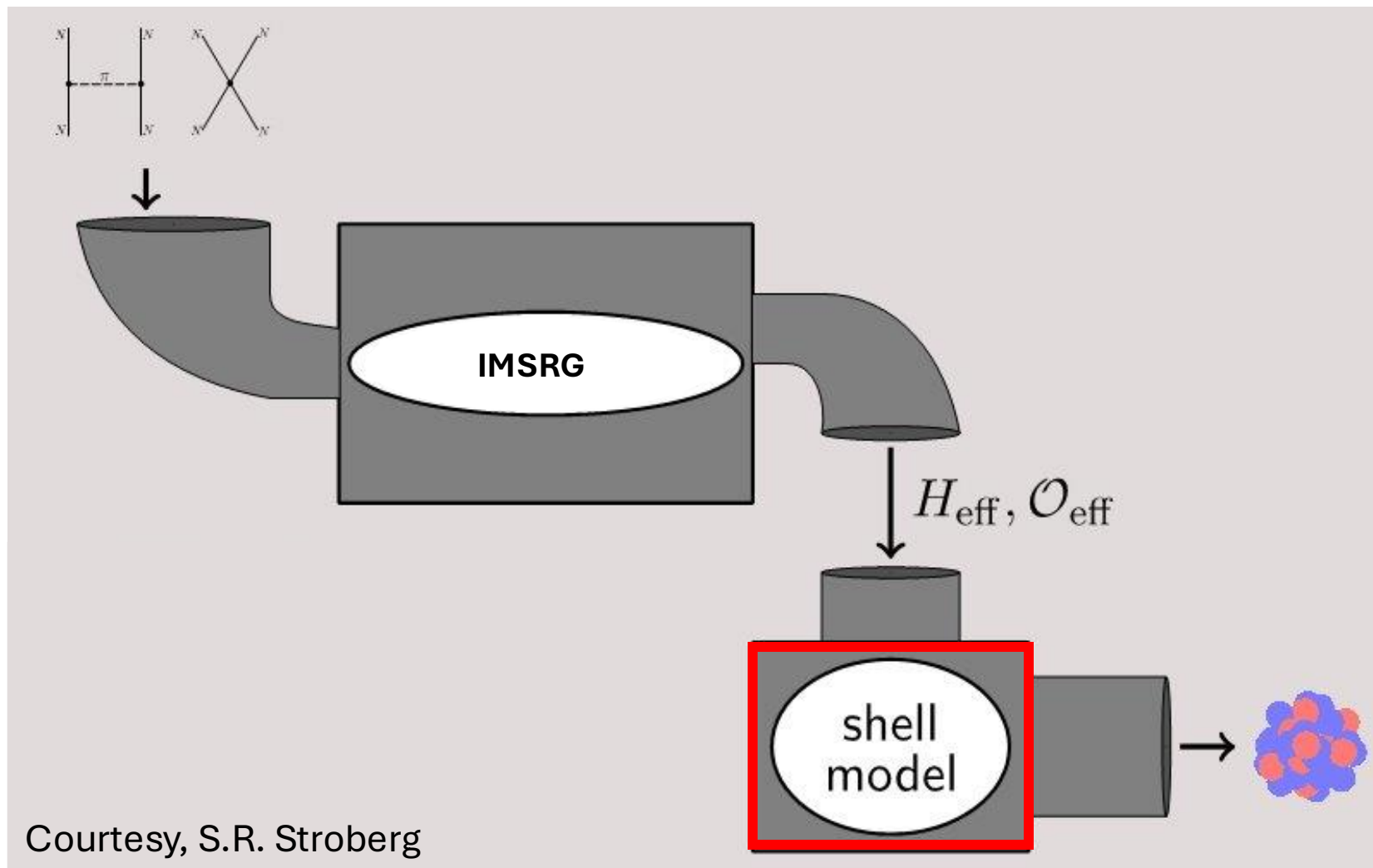


Ab initio nuclear theory

= chiral EFT interactions
+ a (polynomially scaling)
many-body method

Diagonalize \mathcal{H} !!

$$H\psi_n = E_n\psi_n$$



Short-range nuclear matrix elements results

arXiv:2604.22727

Ab initio short-range nuclear matrix elements for neutrinoless double-beta decay

A. Todd,^{1,2} T. Shickele,^{1,3} A. Belley,^{4,1,3} L. Jokiniemi,^{5,6,1} and J. D. Holt^{1,2}

¹*TRIUMF, 4004 Wesbrook Mall, Vancouver, BC V6T 2A3, Canada*

²*Department of Physics, McGill University, 3600 Rue University, Montréal, QC H3A 2T8, Canada*

³*Department of Physics & Astronomy, University of British Columbia, Vancouver, BC V6T 1Z1, Canada*

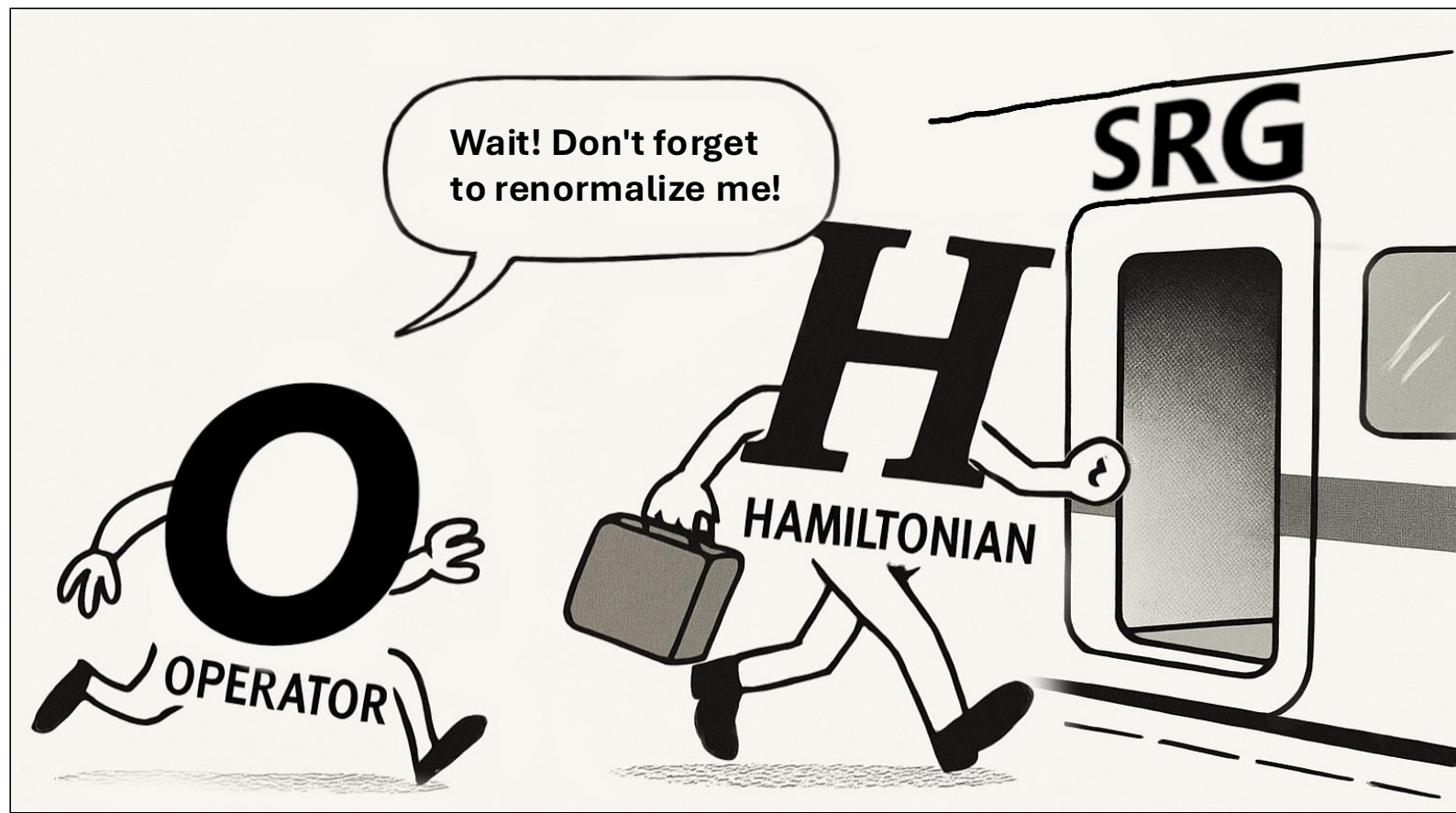
⁴*Massachusetts Institute of Technology, Cambridge, Massachusetts 02139, USA*

⁵*Technische Universität Darmstadt, Department of Physics, D-64289 Darmstadt, Germany*

⁶*ExtreMe Matter Institute EMMI, GSI Helmholtzzentrum für Schwerionenforschung GmbH, D-64291 Darmstadt, Germany*

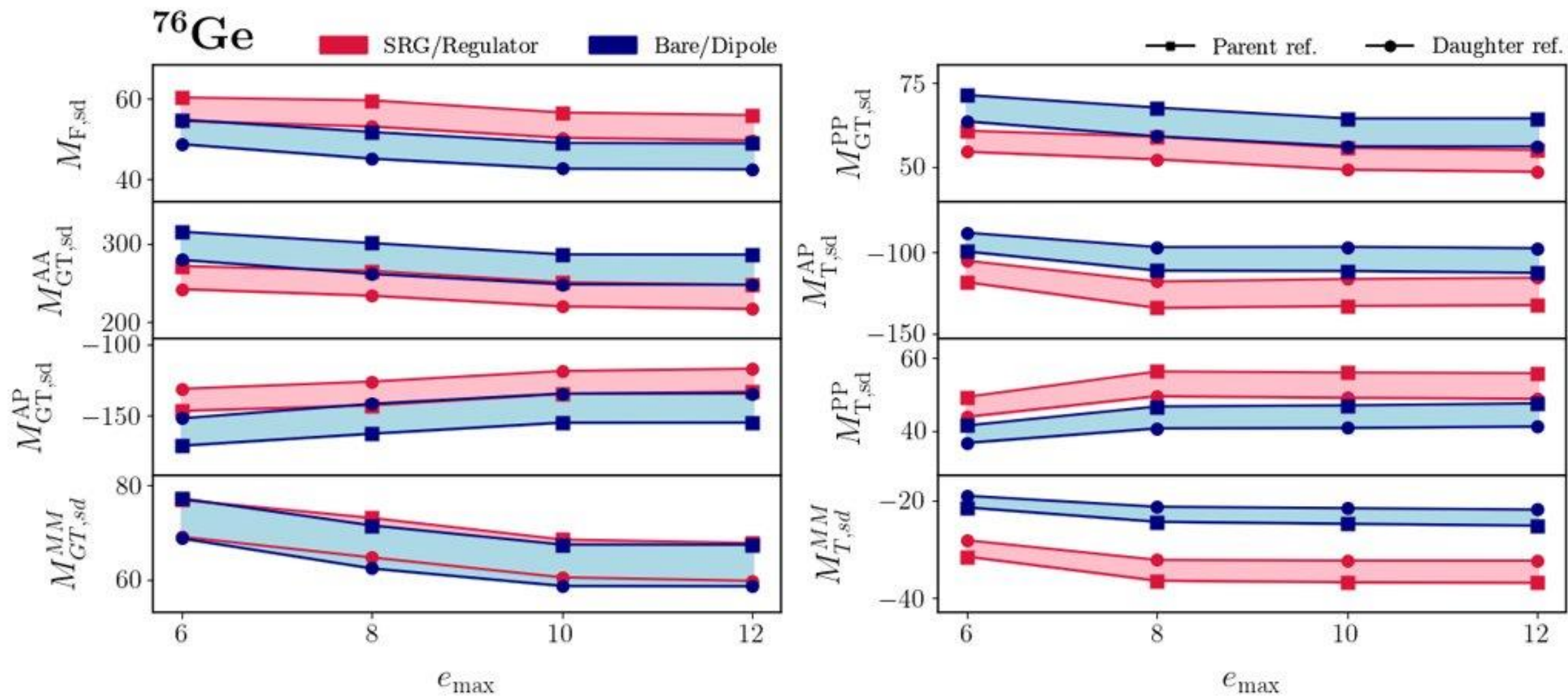
We present converged ab initio calculations of short-range neutrinoless double-beta ($0\nu\beta\beta$) decay nuclear matrix elements for the key experimental isotopes ^{76}Ge , ^{82}Se , ^{130}Te and ^{136}Xe . Starting from different nuclear forces derived from chiral effective field theory, we apply the in-medium similarity renormalization group to obtain an effective valence-space Hamiltonian along with consistently transformed $0\nu\beta\beta$ -decay operators. We then obtain a range of values for the matrix elements that is consistent with, but generally smaller than, those from phenomenology. Finally, we combine our results with current limits from $0\nu\beta\beta$ -decay searches to obtain constraints for the sterile-neutrino mixing-mass parameter space when considering the inclusion of a fourth sterile neutrino.

Be careful with renormalization

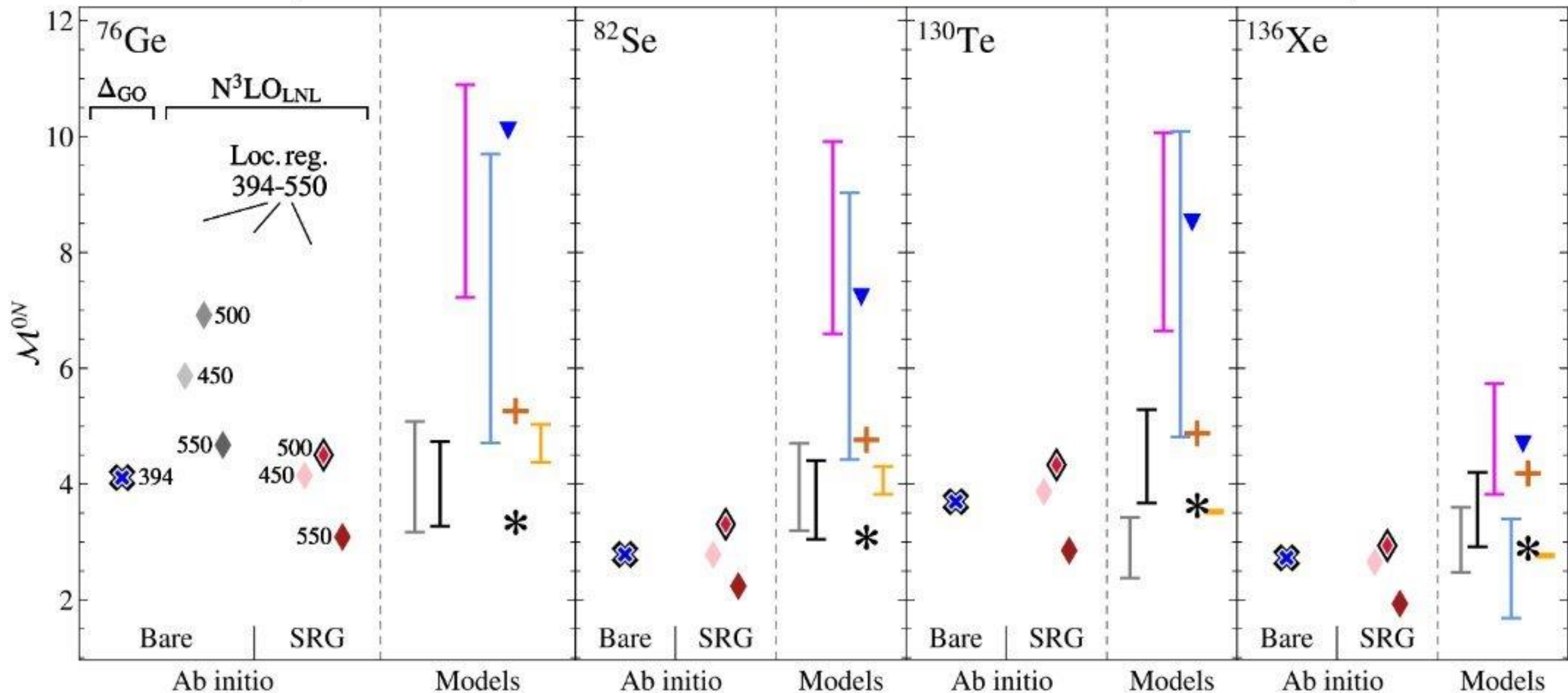


Usually we neglect this; but consistency is important for short-range operators

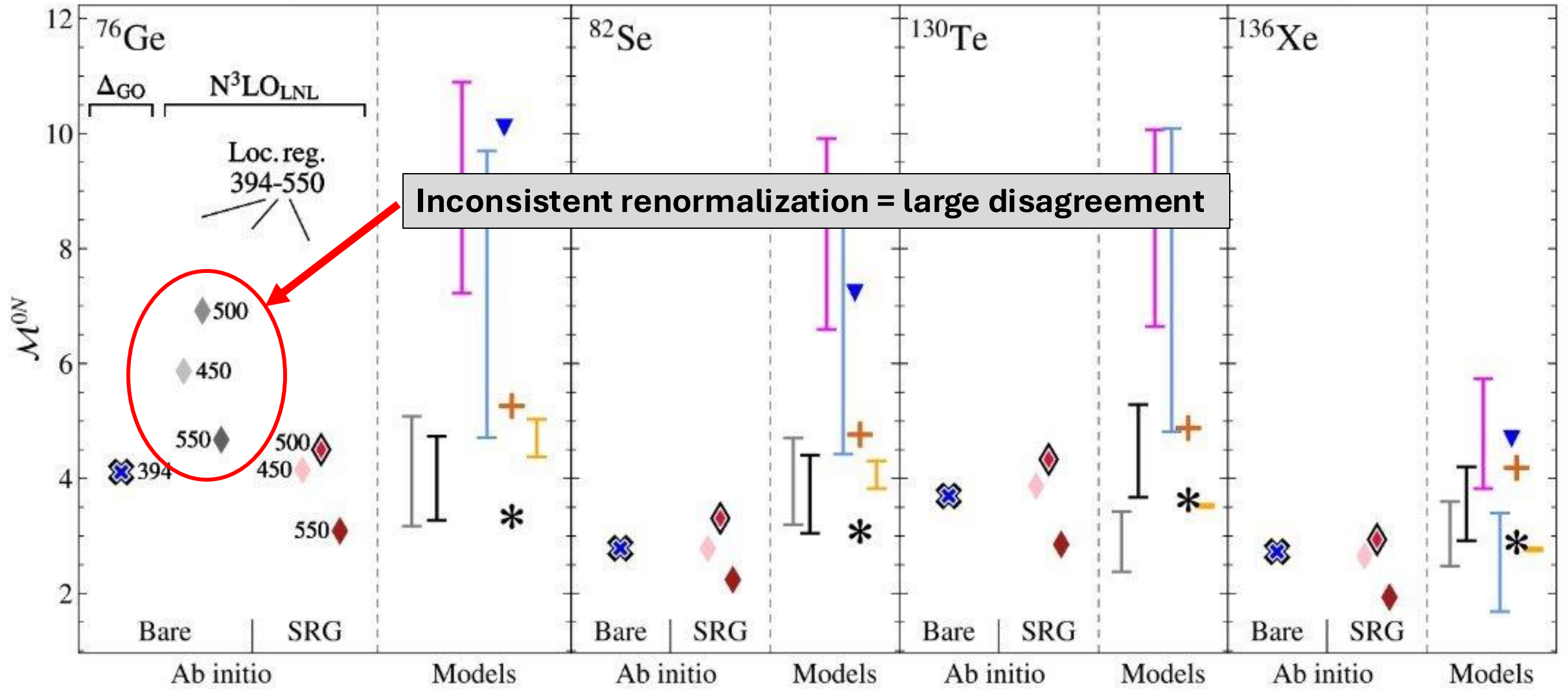
Convergence in single-particle excitations



Final short-range NMEs

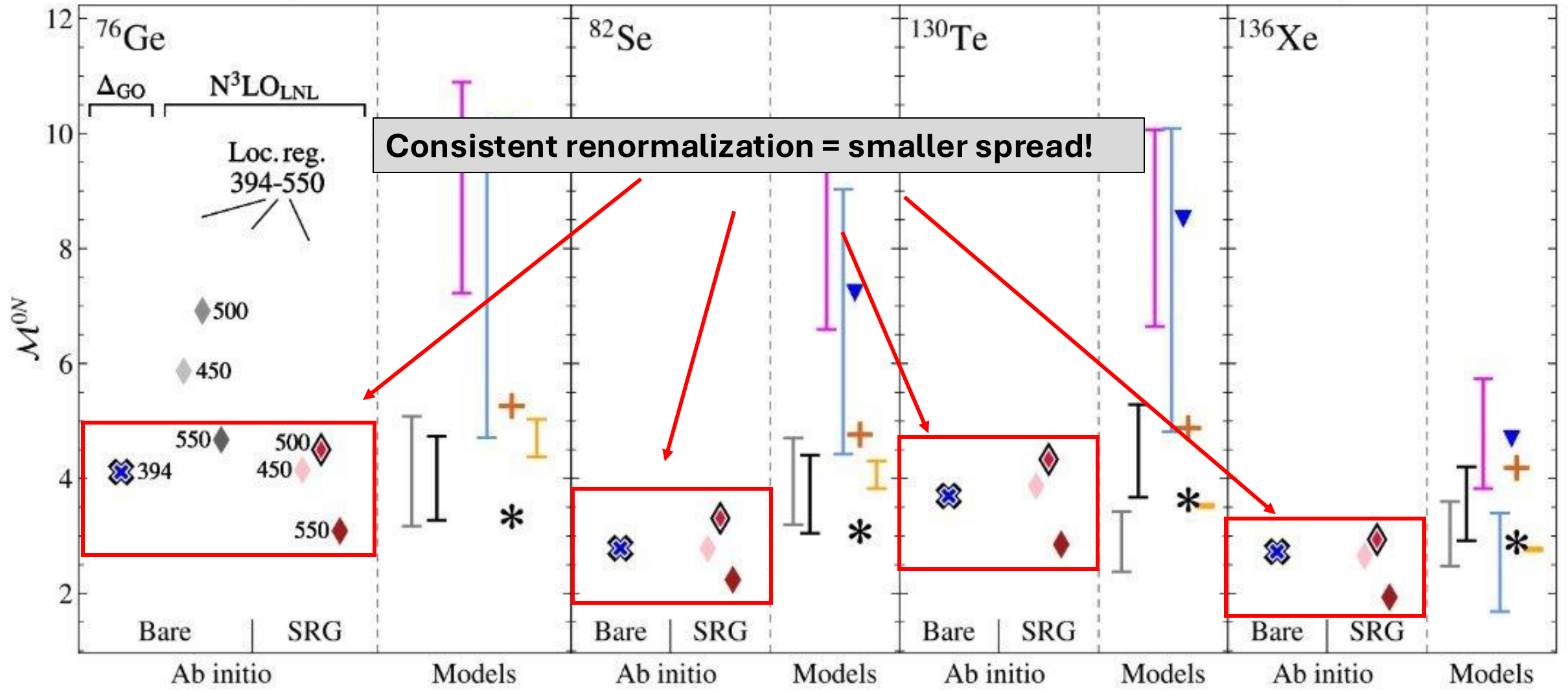


Final short-range NMEs



Inconsistent renormalization = large disagreement

Final short-range NMEs



Connection to BSM mechanisms

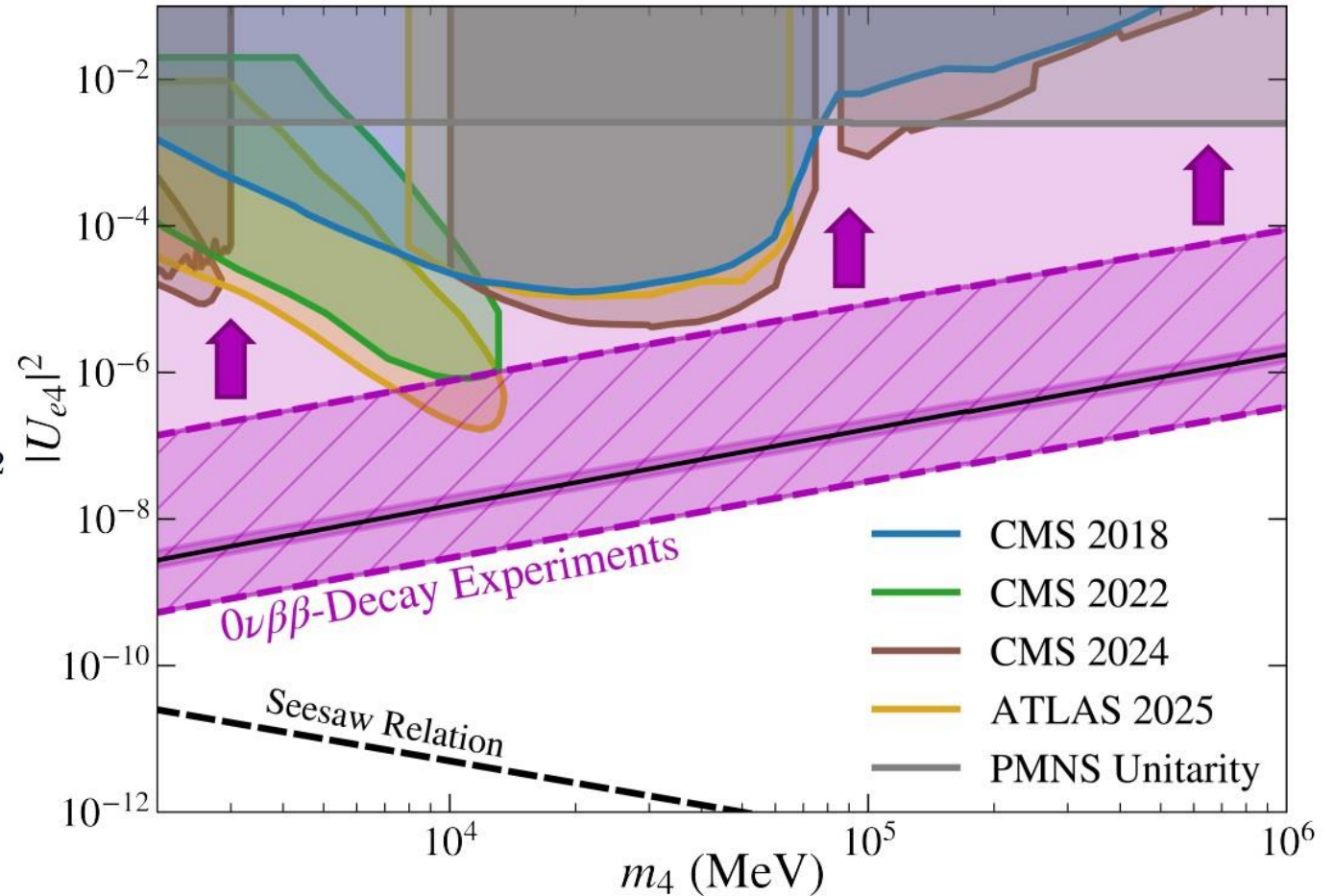
Assume 3+1 model, with heavy neutrino-exchange dominating the amplitude:

$$\begin{aligned} \left[T_{1/2}^{0\nu} \right]^{-1} &= 4g_A^4 G_{01} V_{ud}^4 \eta(\mu, m_4)^2 |U_{e4}|^4 \frac{m_\pi^4}{m_e^2 m_4^2} \\ &\times \left[\frac{5}{6} g_1^{\pi\pi} M_{sd}^{PP} + \frac{g_1^{\pi N}}{2} M_{sd}^{AP} + 2g_1^{NN} M_{F,sd} \right]^2 \end{aligned}$$

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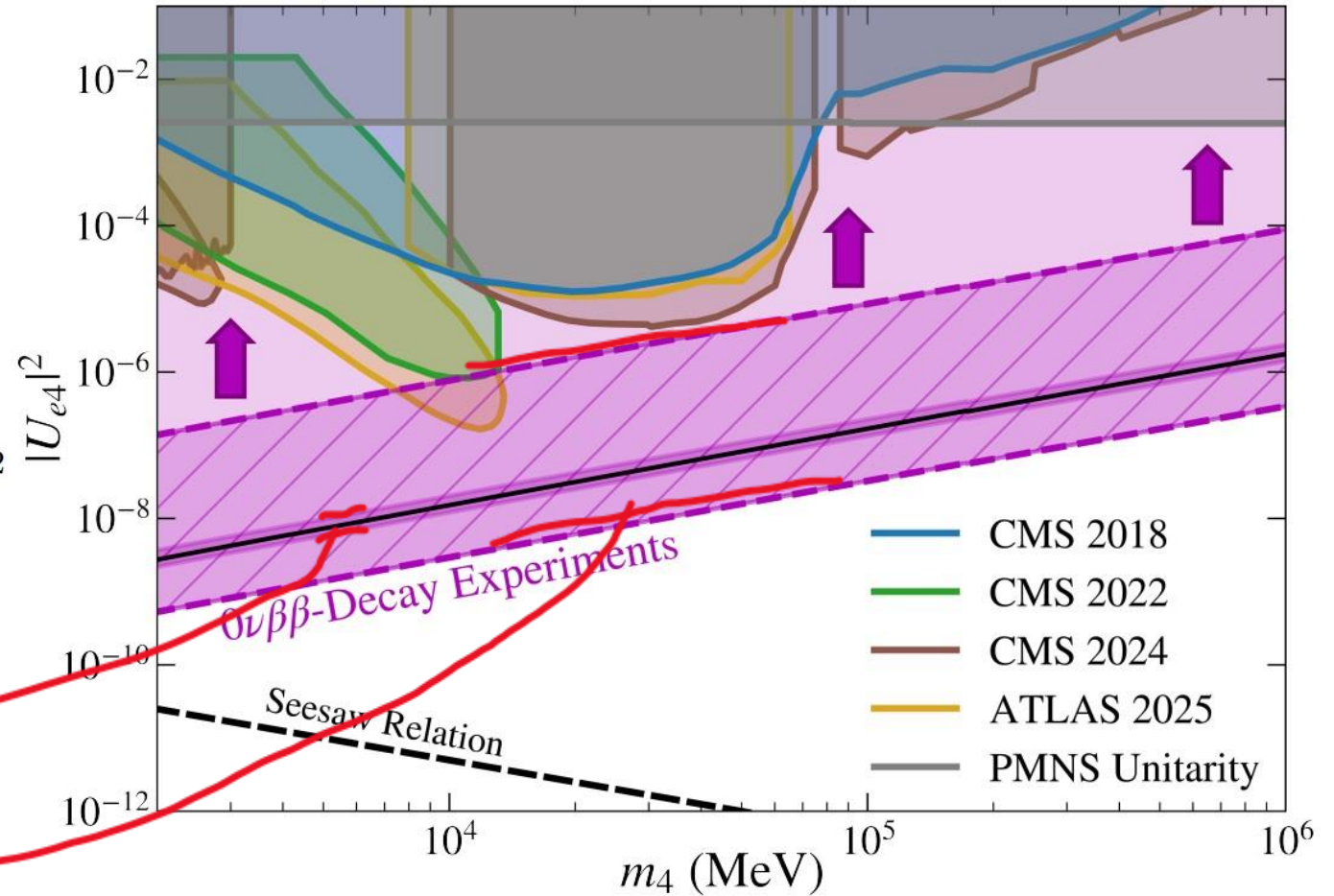
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NME uncertainty

LEC uncertainty



Conclusions:

- Short-range operators are sensitive to different schemes: consistent regulators and SRG transformations are required
- Overall much smaller spread than phenomenological approaches. This is encouraging given the radically different approach.
- 3+1 model: Current $0\nu\beta\beta$ limits competitively probe heavy neutrinos

Outlook:

- Uncertainty quantification
- Taking higher-orders within our method: 3-body operators, N2LO operator contributions

Collaborators: Taiki Shickele, Antoine Belley,
Lotta Jokiniemi, Jason Holt

Conclusions:

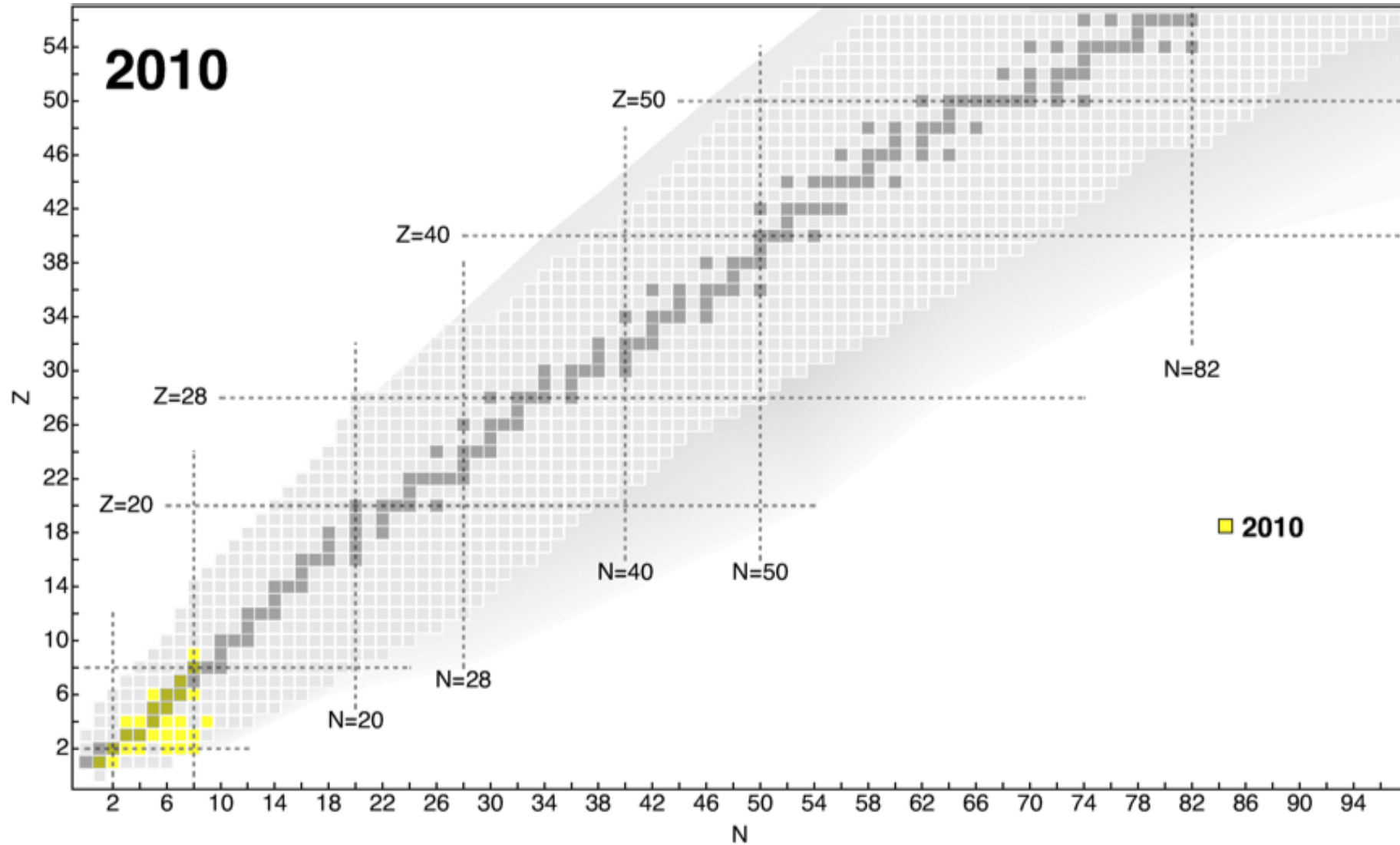
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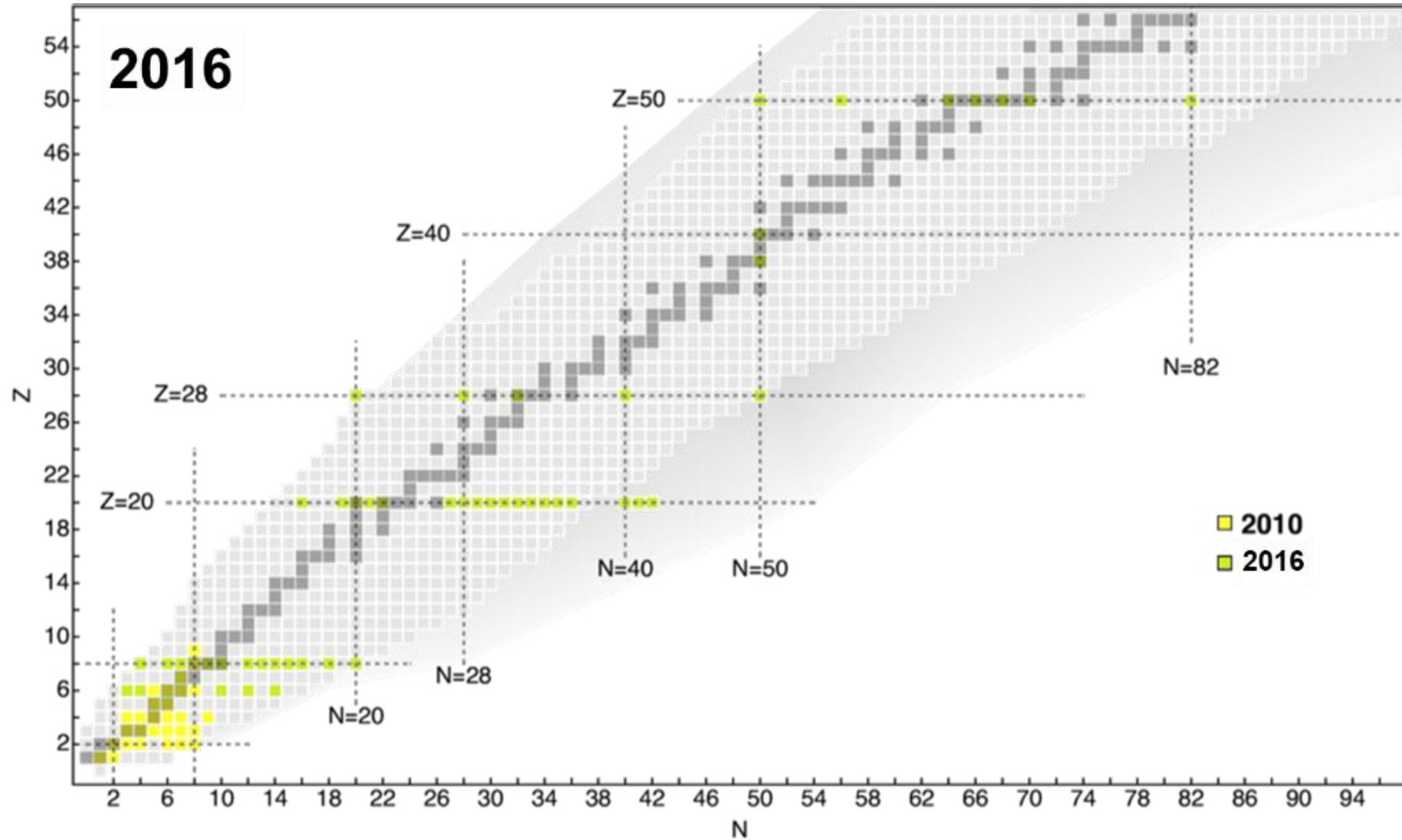
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Backup Slides

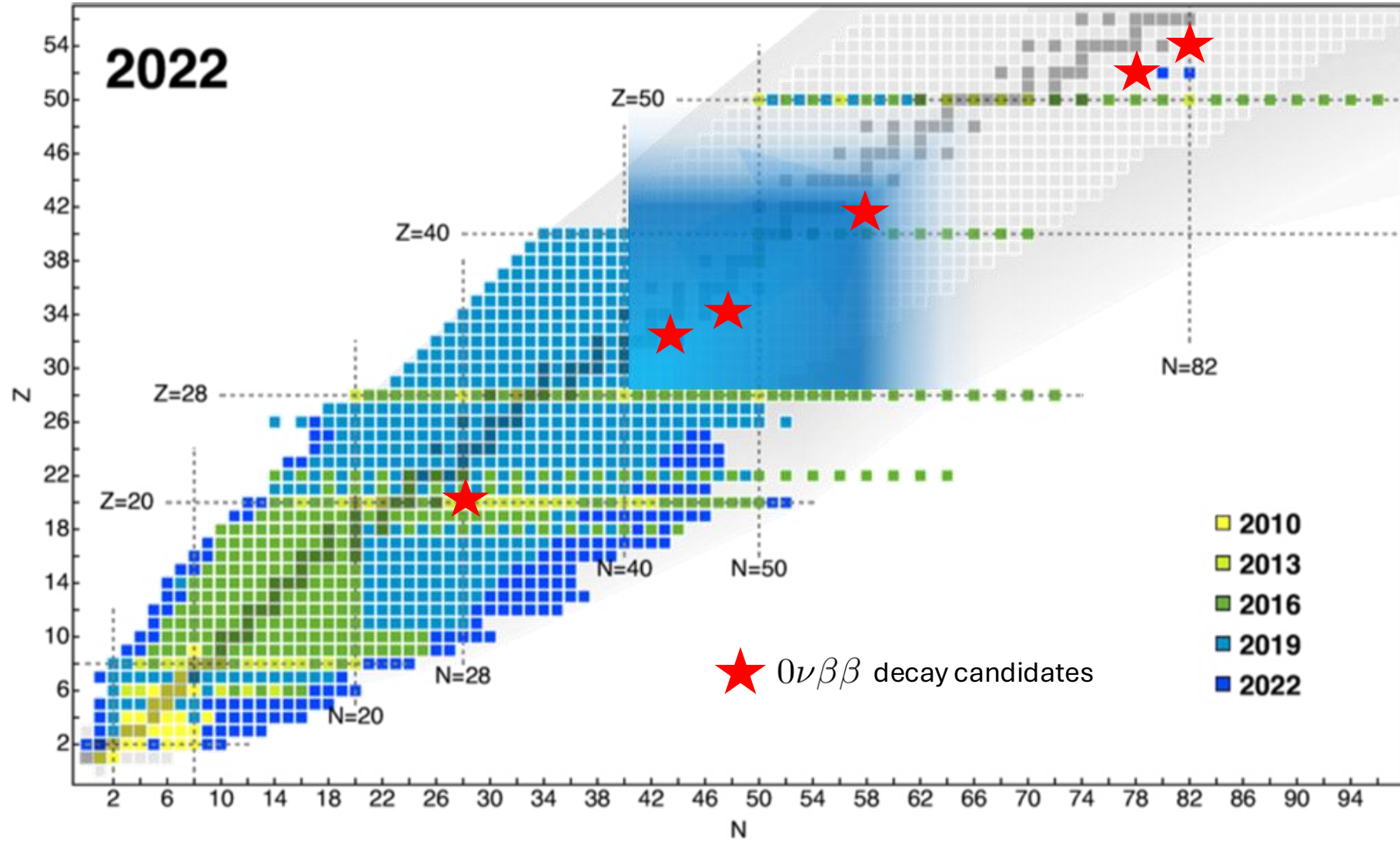
Ab initio progress



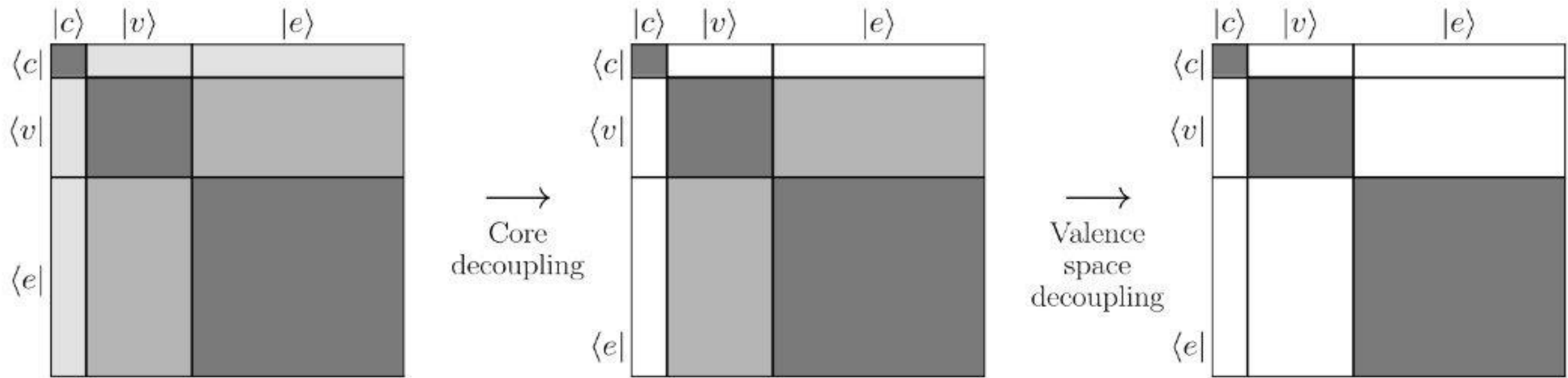
Ab initio progress



Ab initio progress

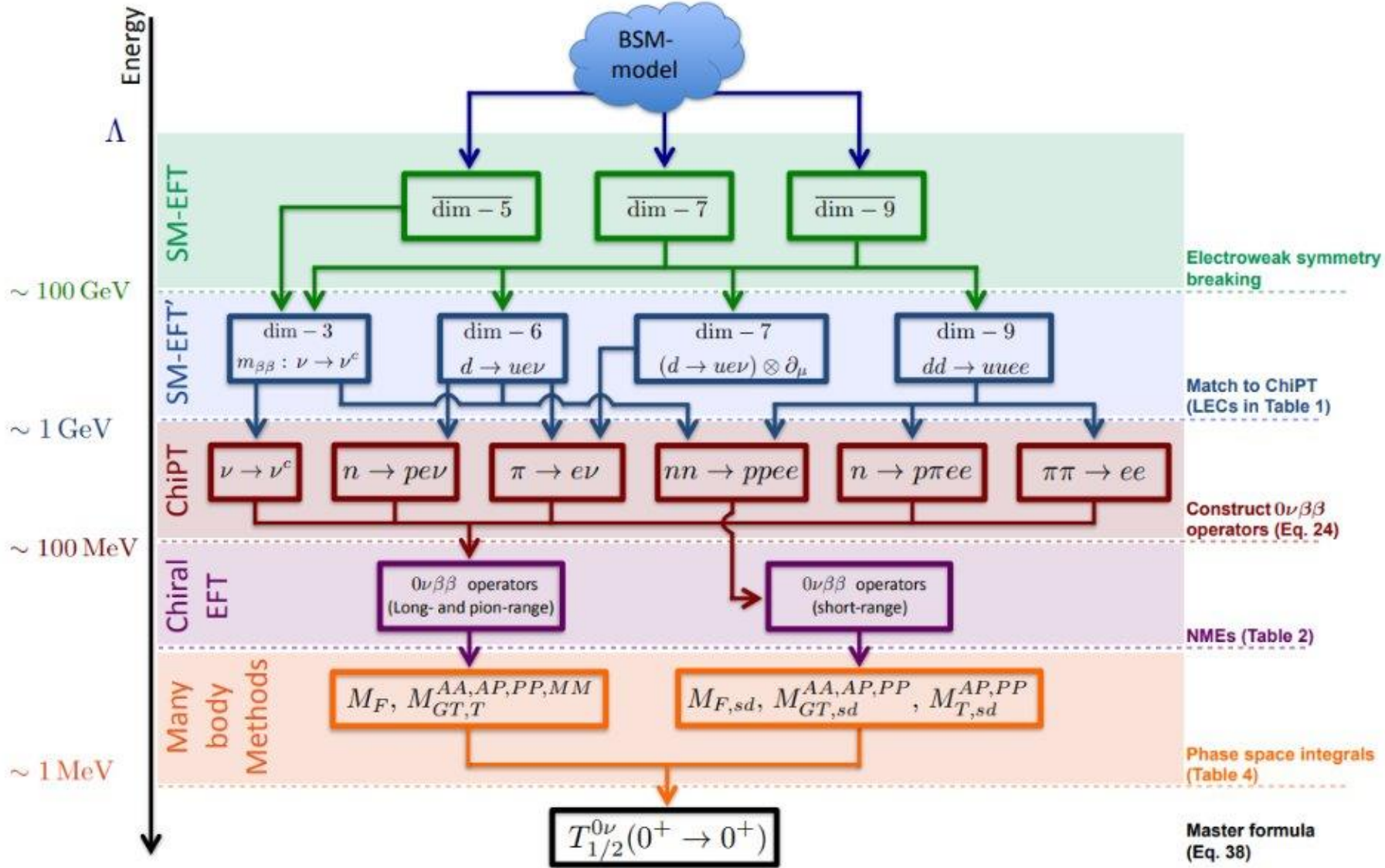


Valence-space In-medium Similarity Renormalization Group (VS-IMSRG)



$$H(s) = e^{\Omega(s)} H(0) e^{-\Omega(s)} = \sum_{k=0}^{\infty} \frac{1}{k!} [\Omega(s), H(0)]^{(k)} = H + [\Omega, H] + \frac{1}{2} [\Omega, [\Omega, H]] + \dots$$

$0\nu\beta\beta$ "master formula"



Model independence

Leading order decay rate for any BSM mechanism can be built out of **9 long-range NMEs** and **8 short-range NMEs**

$$M_F, M_{GT}^{AA}, M_{GT}^{AP}, M_{GT}^{PP}, M_{GT}^{MM}, M_T^{AA}, M_T^{AP}, M_T^{PP}, M_T^{MM}$$

$$M_{F,sd}, M_{GT,sd}^{AA}, M_{GT,sd}^{AP}, M_{GT,sd}^{PP}, M_{GT,sd}^{MM}, M_{T,sd}^{AP}, M_{T,sd}^{PP}, M_{T,sd}^{MM}$$