

IAS PROGRAM

# Fundamental Physics

January 12-16, 2026

## Applying strong dynamics to dark matter phenomenology

*Xiaoyong Chu (ICTP-AP, UCAS)*

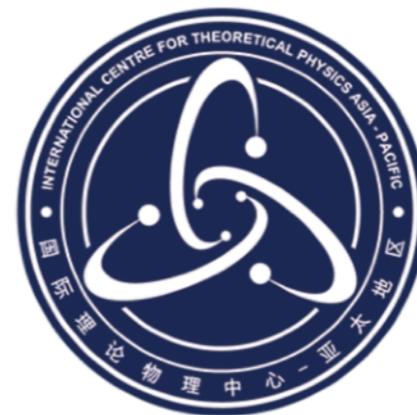
With Josef Pradler, Daris Samart 2512.08517



中国科学院大学  
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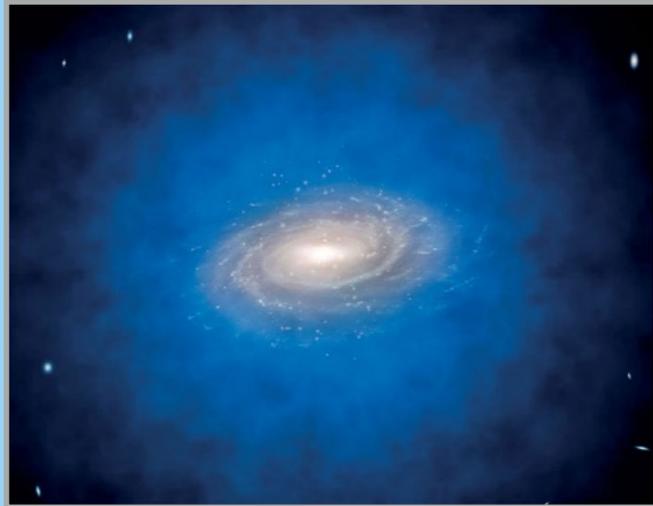
### ICTP-AP

International Centre  
for Theoretical Physics Asia-Pacific  
国际理论物理中心-亚太地区

# 1. If **strong dynamics** in dark sector

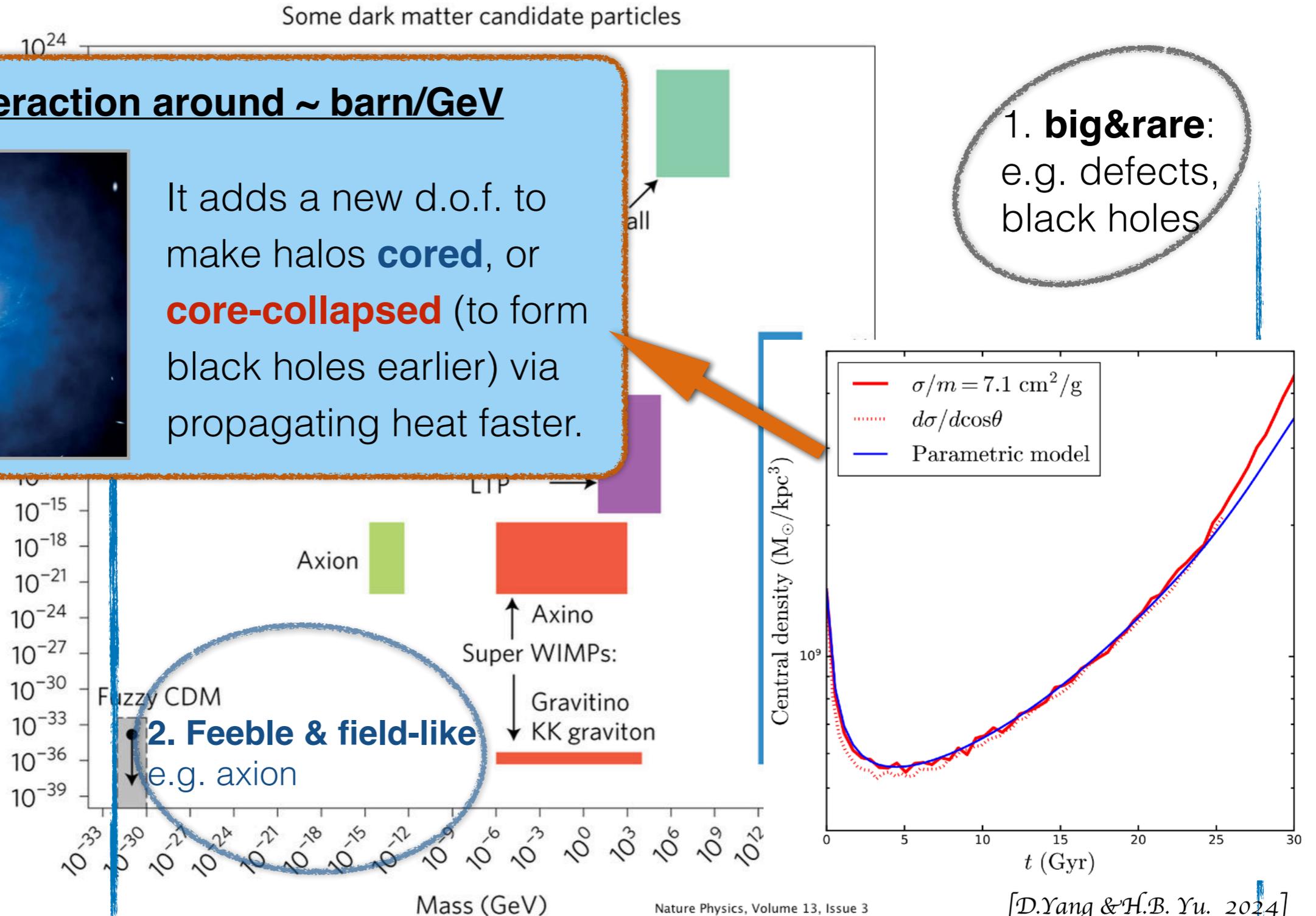
# Paradigm shift in BSM phenomenology: dark sector

## 3. Self-interaction around $\sim$ barn/GeV



It adds a new d.o.f. to make halos **cored**, or **core-collapsed** (to form black holes earlier) via propagating heat faster.

1. **big&rare**:  
e.g. defects,  
black holes



Nature Physics, Volume 13, Issue 3

[D. Yang & H. B. Yu. 2024]

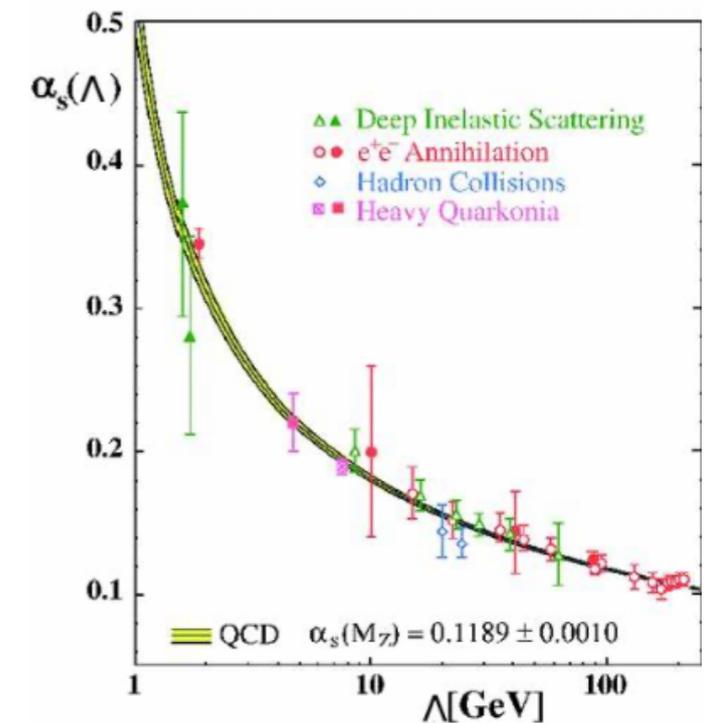
# Dark strong dynamics has rich phenomenology

## A dark-QCD DM can be the stable composite state

- dark baryon (if it is like the SM);
- dark pion (light-quark limit);
- dark glueball (heavy-quark limit);
- large-N nuclei, dark atom,....

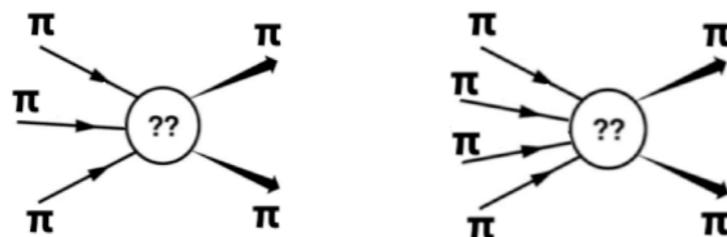
*J. M. Cline, Z. Liu, et al., 2013,*

*G. D. Kribs & E.T. Neil 2016*



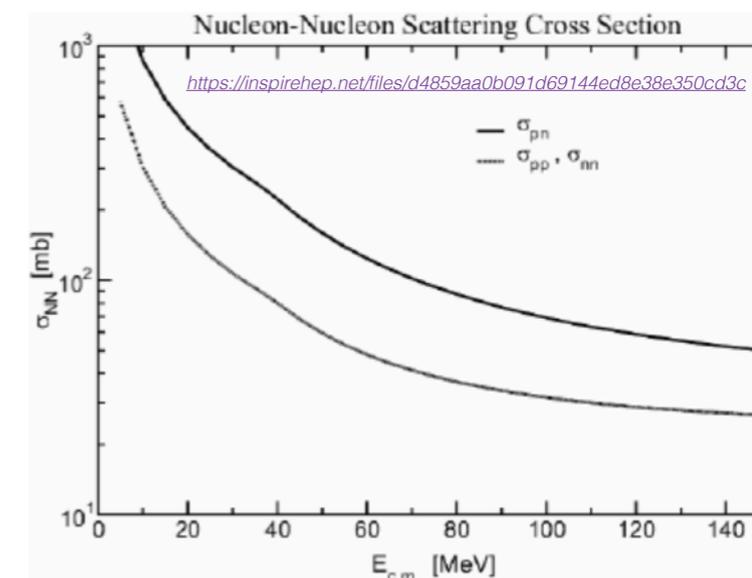
## Production mechanism:

- DM could be asymmetric, to prevent significant self-annihilation.  
5 GeV to explain coincidence, or other co-genesis?
- Or one needs N-body annihilation ( $N > 2$ )



**Self-interaction:** exchanging one or more dark pions.

Interesting SIDM energy-scale: sub-GeV QCD phase-transition



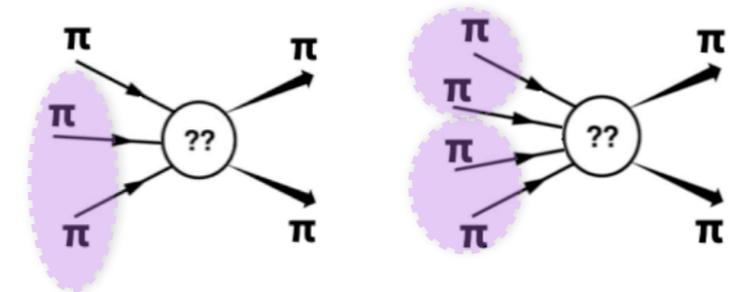
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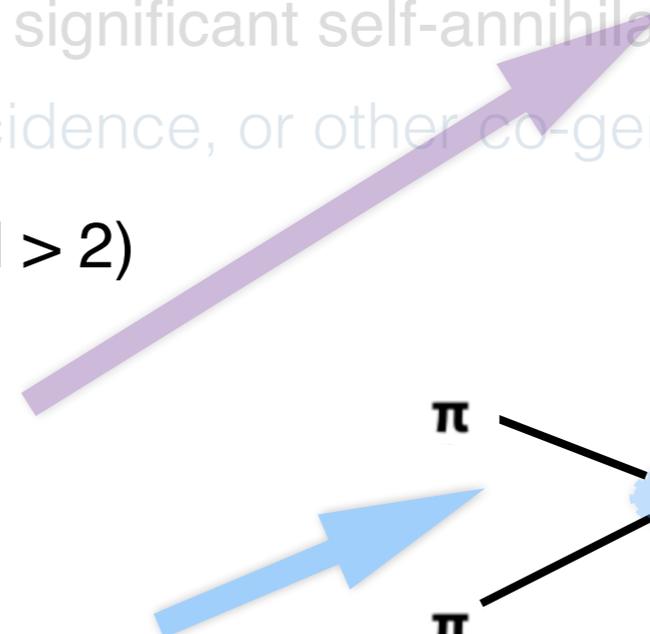
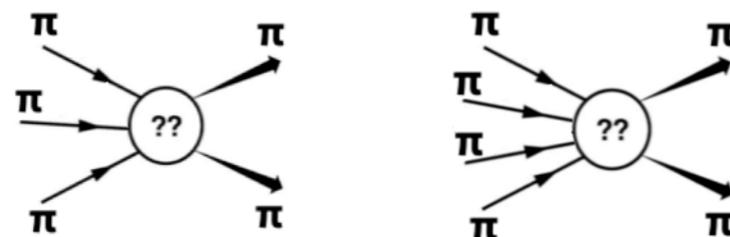
*G. D. Kribs & E.T. Neil 2016*



*[XC, M. Nikolic, J. Pradler 2024]*

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*[XC, C. Garcia-Cely, H. Murayama 2018]*

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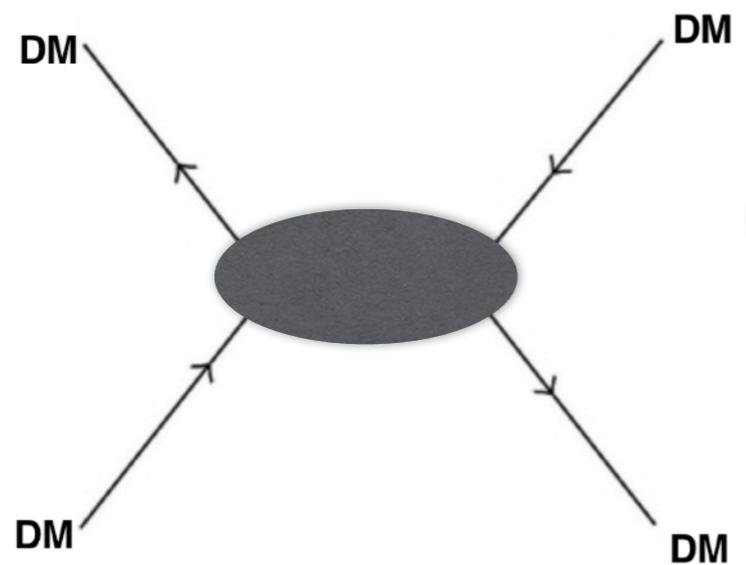
# 3. **Dark self-interactions:** universally

# Dark strong dynamics has **rich phenomenology**

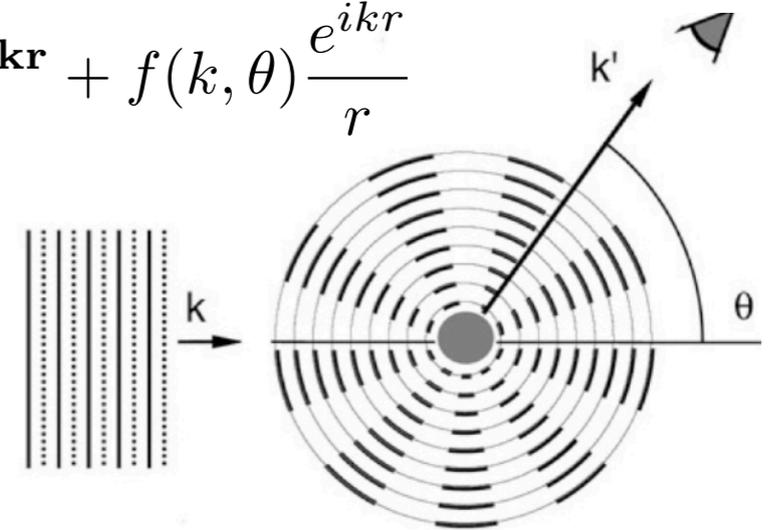
Quantitative statements need lattice, and depend on concrete models.

**How far can we go with universal/analytical methods?**

[XC, C. García-Cely, H. Murayama 2019]



$$\psi(r) \simeq e^{i\mathbf{k}\mathbf{r}} + f(k, \theta) \frac{e^{ikr}}{r}$$



$$f_\ell(k) \equiv \frac{e^{2i\delta_\ell(k)} - 1}{2ik} = \frac{1}{k(\cot \delta_\ell(k) - i)}$$

**boundary conditions of finite-range potential** allows an expansion of the phase shift at very low velocities [Schwinger, Blatt&Jackson, Bethe, 1940s] :

$$k \cot \delta_0 = -\frac{1}{a_0} + \frac{k^2}{2r_{e,0}} + \mathcal{O}(k^4)$$

(strong non-relativistic scattering, suggesting  $\text{Im}[\delta_0] \ll \text{Re}[\delta_0]$ , and neglected)

# A quantitative way?

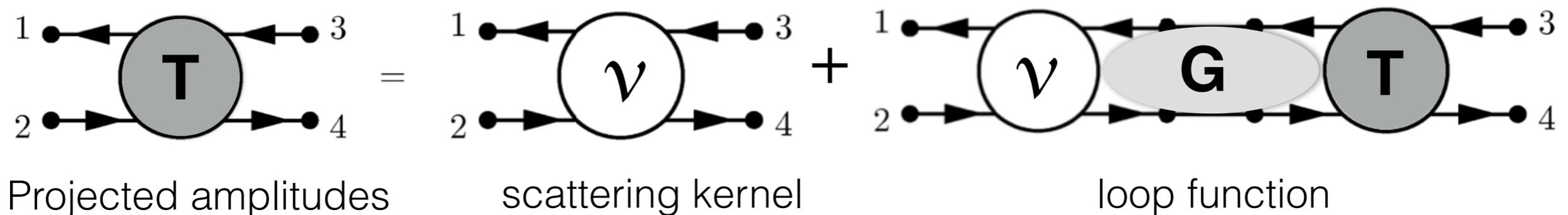
[XC, Josef Pradler, Daris Samart 2025]

Adopt various **unitarization methods** (N/D, IAM etc.) and impose s/p-wave projections, but essentially:

$$T(s) = \frac{v(s)}{1 + v(s)G(s)}$$

(Sign depends on convention)

Similar to Lippmann-Schwinger equation, it is simply



At non-relativistic limit  $\tilde{G}(E) \equiv \int d^3p \frac{\Theta(\Lambda - p)}{E - 2m_\pi - \frac{\vec{p}^2}{m_\pi}}$

# A quantitative way?

[XC, Josef Pradler, Daris Samart 2025]

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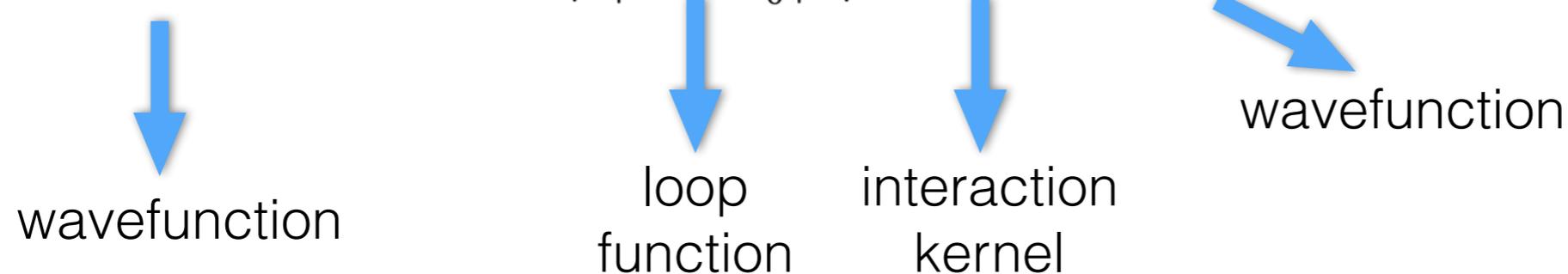
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For **bound state-dominated** scattering amplitudes:

[following D. Gamermann, J. Nieves, E. Oset, & E. Ruiz Arriola, 2009]

1. first work out the bound wavefunction

$$\langle \vec{p} | \Psi \rangle = \int d^3k \int d^3k' \langle \vec{p} | \frac{1}{E - \hat{H}_0} | \vec{k} \rangle \langle \vec{k} | \hat{V} | \vec{k}' \rangle \langle \vec{k}' | \Psi \rangle$$



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↓  
wavefunction

↓  
loop  
function

↓  
interaction  
kernel

↓  
wavefunction



$$|\Psi(0)| = \frac{|\tilde{G}(E)|}{(2\pi)^{3/2}} \left| \frac{d\tilde{G}(E)}{dE} \right|^{-1/2} \quad \text{at} \quad E = 2m_\pi - E_B$$

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1. first work out the bound wavefunction;
2. then take another form of the amplitude;

$$T(E^2) = \frac{g_{X\pi\pi}^2}{E - (2m_\pi - E_B)}$$

Both  $1/T$  and  $E - (2m_\pi - E_B)$  go to zero at the BS pole (**L'Hôpital's rule**)

$$g_{X\pi\pi}^2 = 32\pi^3 m_\pi^2 \left( \frac{d\tilde{G}(E)}{dE} \right)^{-1} \Big|_{E=2m_\pi - E_B}$$



# A quantitative way?

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**Nature choice of parameters** gives astrophysically interesting values

(sub-GeV SIDM): (whatever dark group structures)

$$\frac{\langle \sigma_{\text{self}} \rangle}{m_{\pi}} \simeq 1 \text{ cm}^2/\text{g} \left( \frac{0.1}{\kappa} \right)^2 \left( \frac{0.5 \text{ GeV}}{m_{\pi}} \right)^3 \left( \frac{|\Psi(0)|}{0.14 m_{\pi}^{3/2}} \right)^4 \left( \frac{0.02}{|G(s)|_{s=(2m_{\pi}-\kappa m_{\pi})^2}} \right)^4$$

# 3. **Dark self-interactions:** a toy model

# A quantitative way?

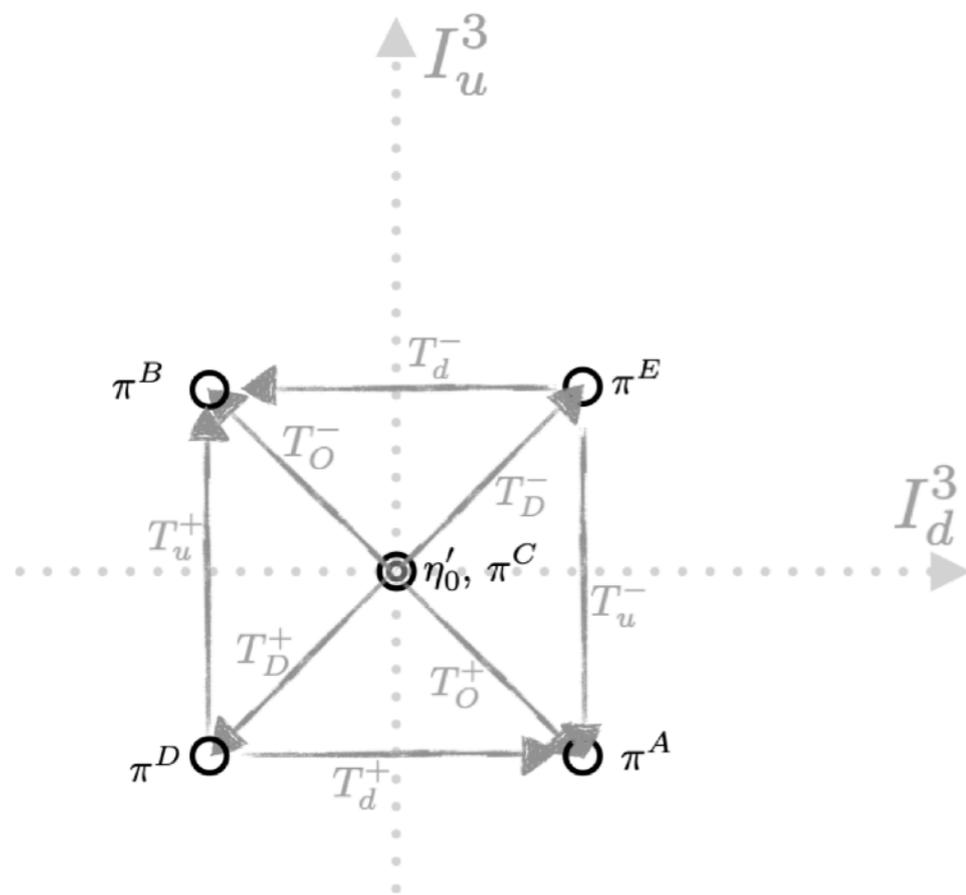
[XC, Josef Pradler, Daris Samart 2025]

Take SIMP motivated: 2-quark case, global symmetry  $SU(4) \rightarrow Sp(4)$

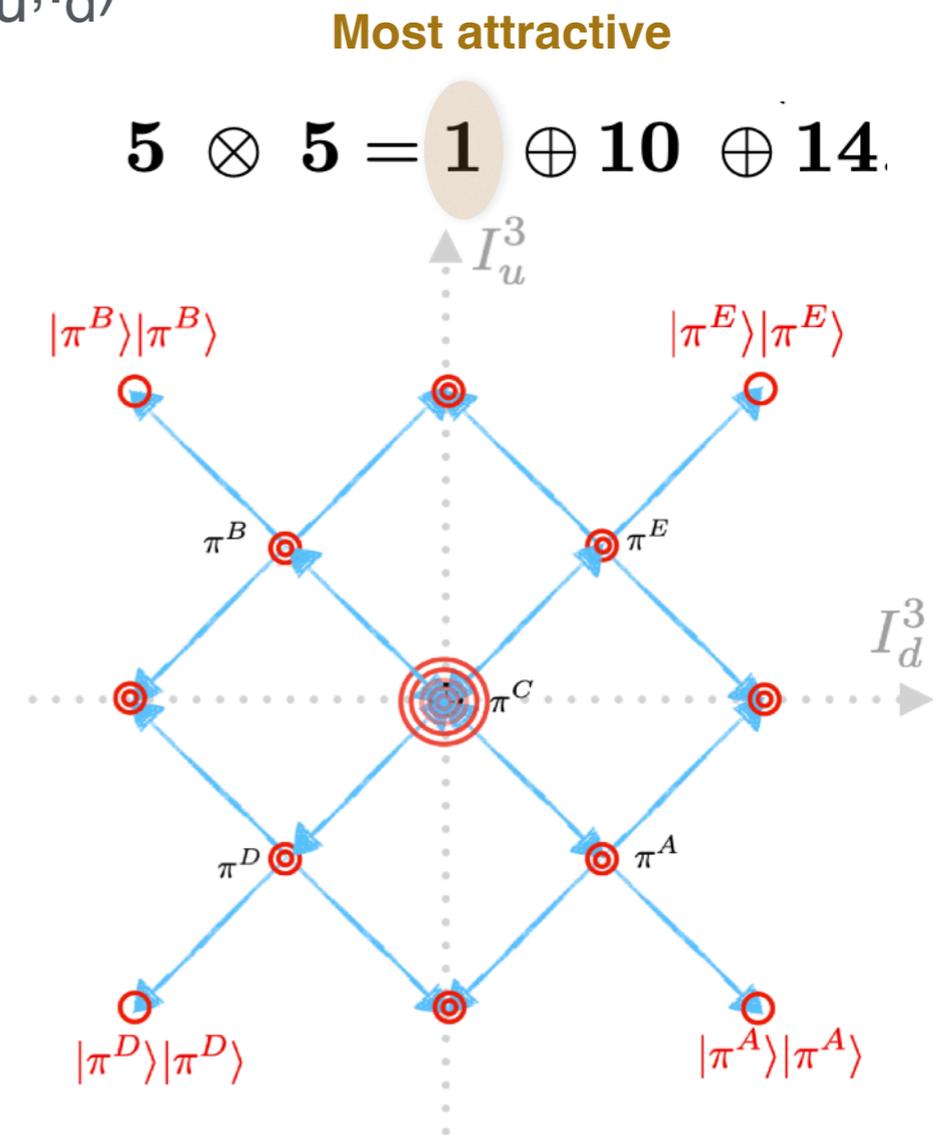
(five dark pions as dark matter)

For rank-2 group  $Sp(4)$ , weight vectors labeled by  $(I_u, I_d)$

**Traceless** part of anti-symmetric rank-2 tensor **5**



$$\pi^{a_1}(p_1) + \pi^{a_2}(p_2)$$



# A quantitative way?

[XC, Josef Pradler, Daris Samart 2025]

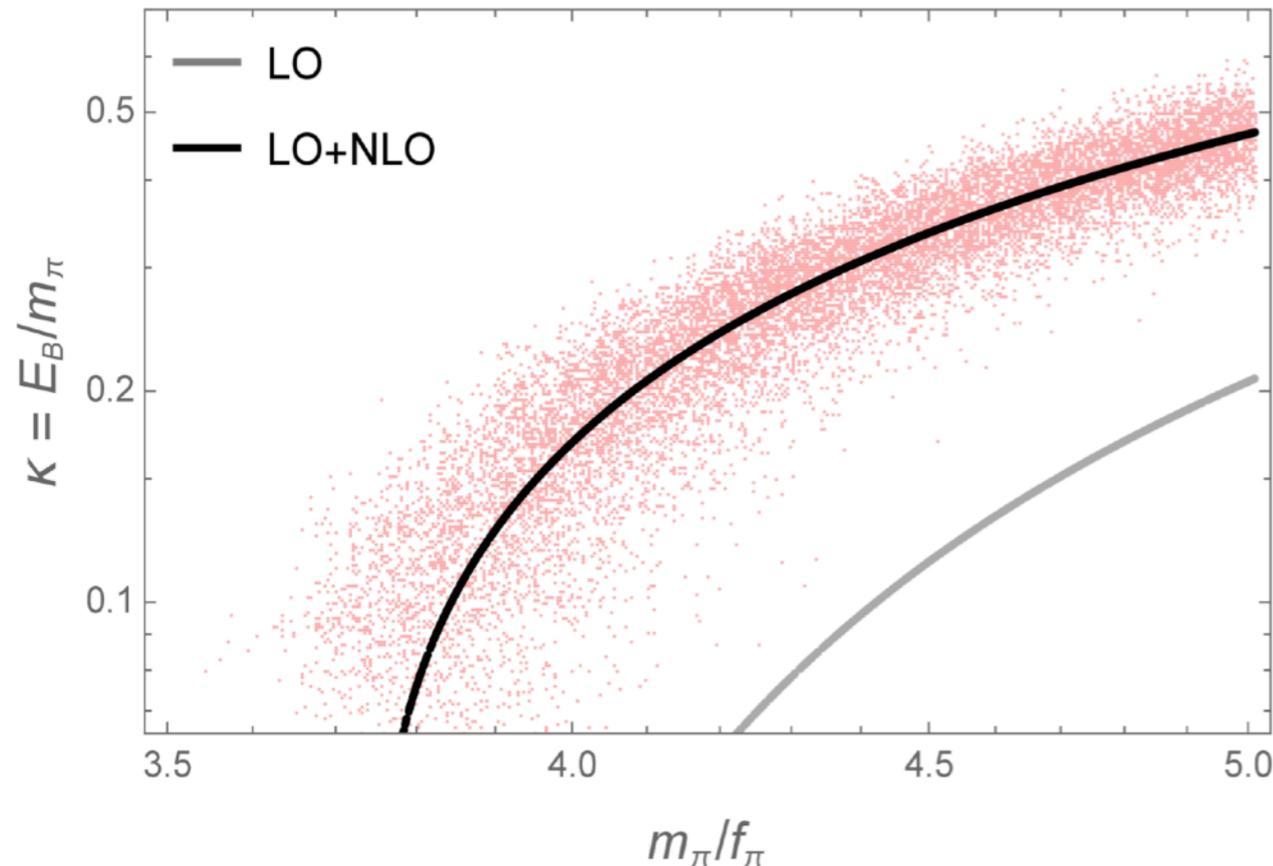
Take SIMP motivated: 2-quark case, global symmetry  $SU(4) \rightarrow Sp(4)$

(five dark pions as dark matter)

→ Singlet (isoscalar) channel of  $\pi^{a_1}(p_1) + \pi^{a_2}(p_2) \rightarrow \pi^{a_3}(p_3) + \pi^{a_4}(p_4)$

use amplitudes from *M. Albaladejo & J. A. Oller 2012*, to work out the s-wave projected interaction kernel to NLO.

Only s-wave pole exists (sigma meson becomes a **bound state**, rho meson cannot)



# A quantitative way?

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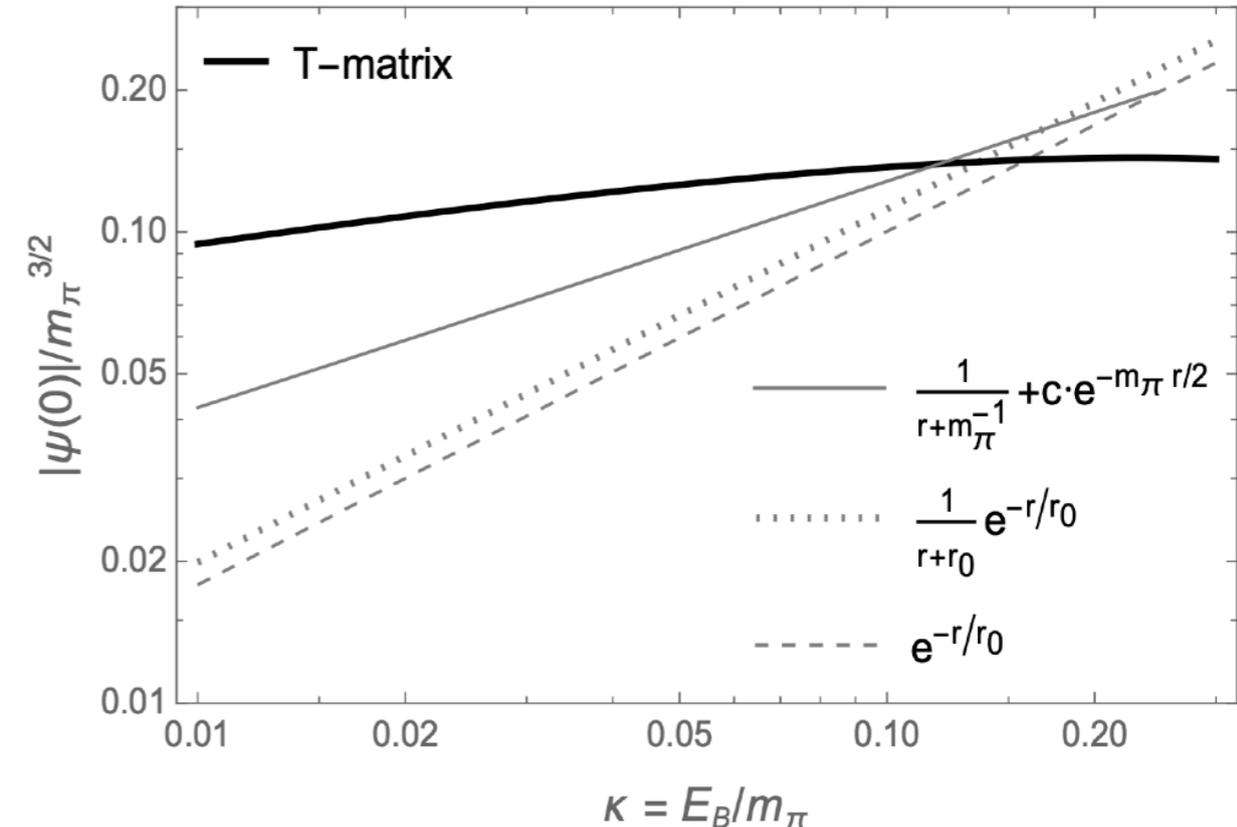
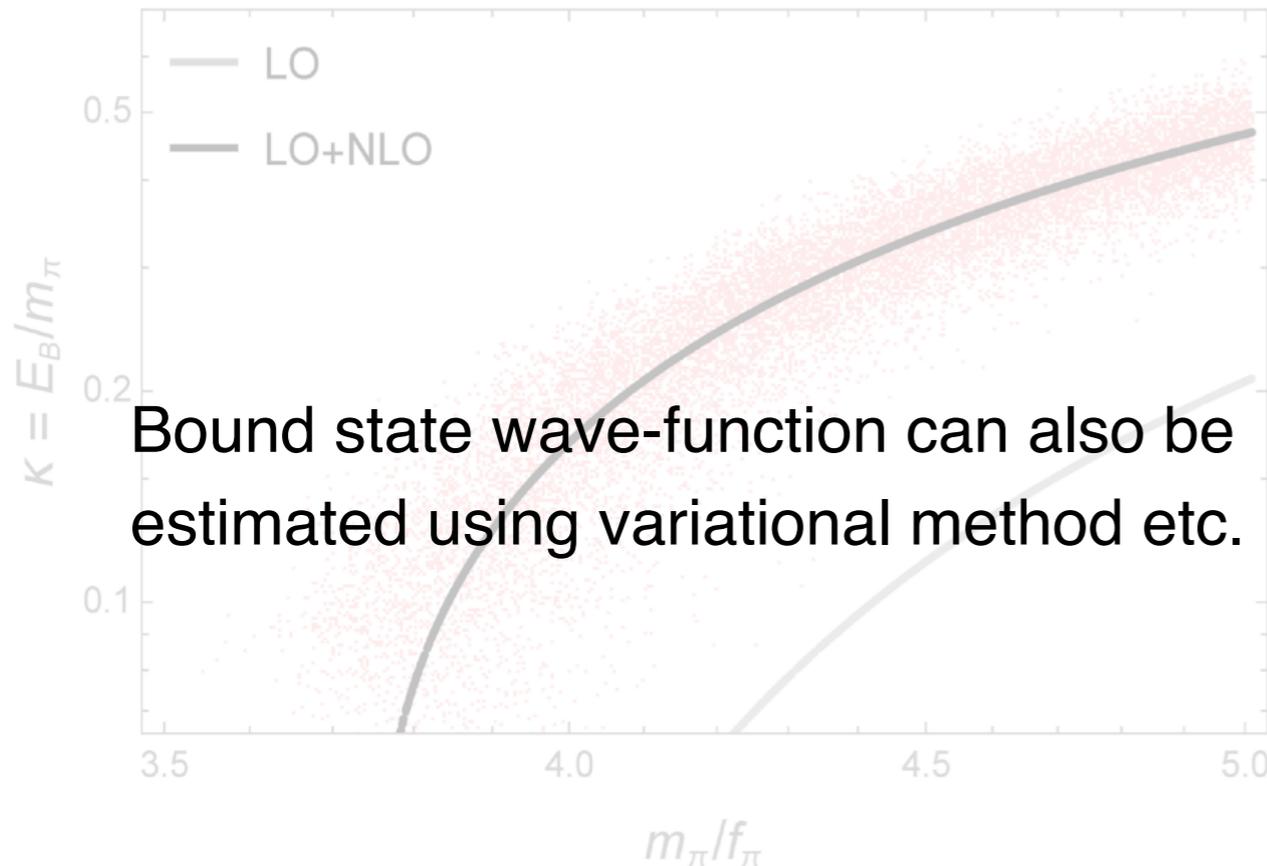
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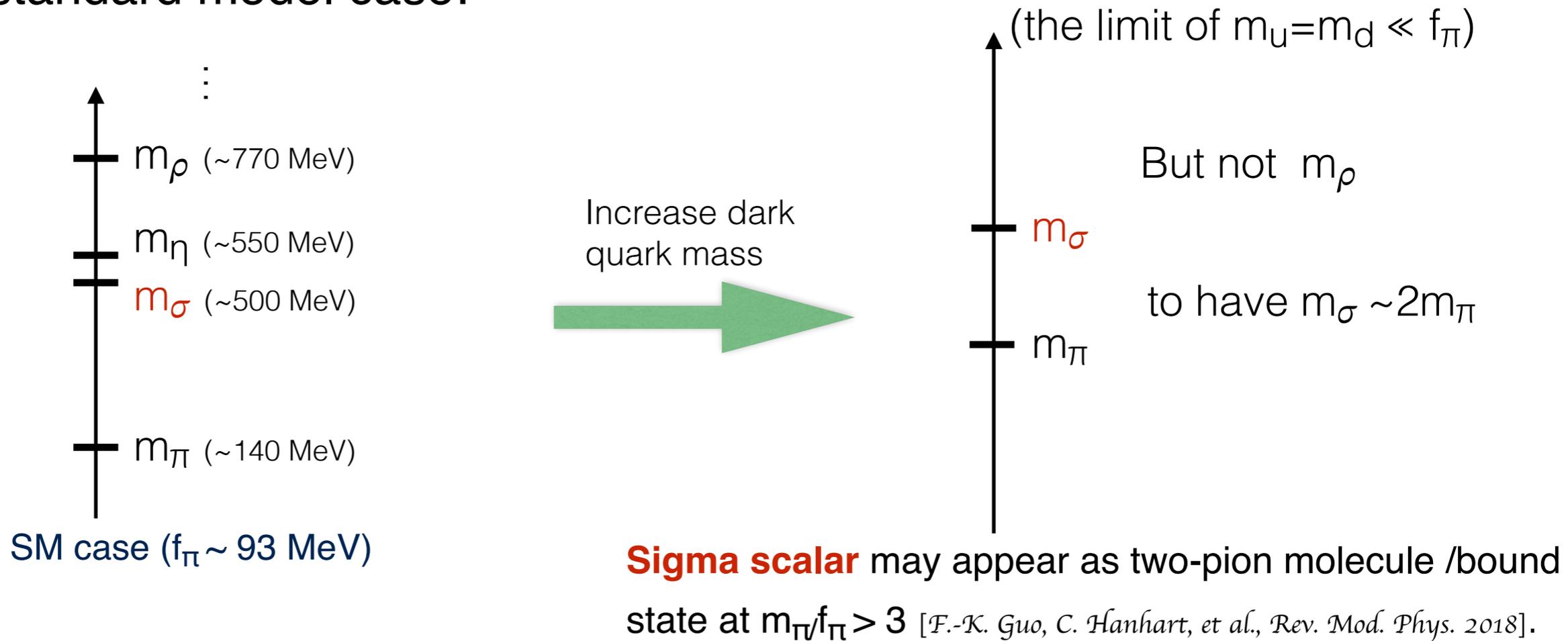
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In the standard model case:



These results of  $SU(4)/Sp(4)$  are not a surprise.

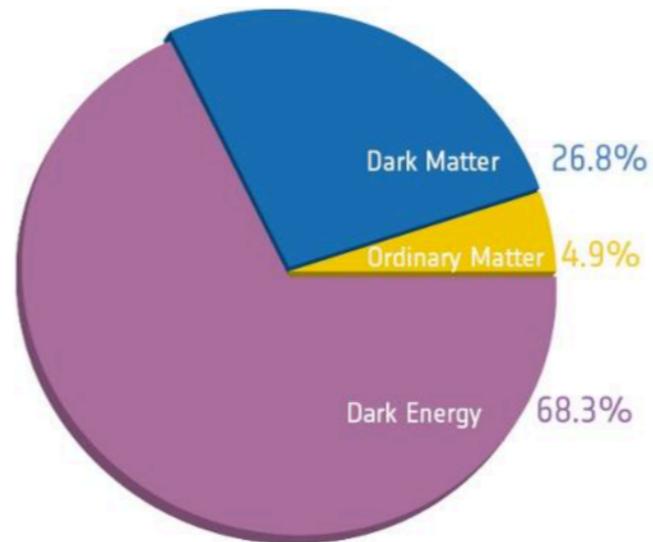
Actually **the attraction is weaker** in this group structure.

# 5. Perspectives

# In the end

**A dark-QCD DM** can be the stable composite state

**Precisely-measured** relic abundance



**Without** known particle properties

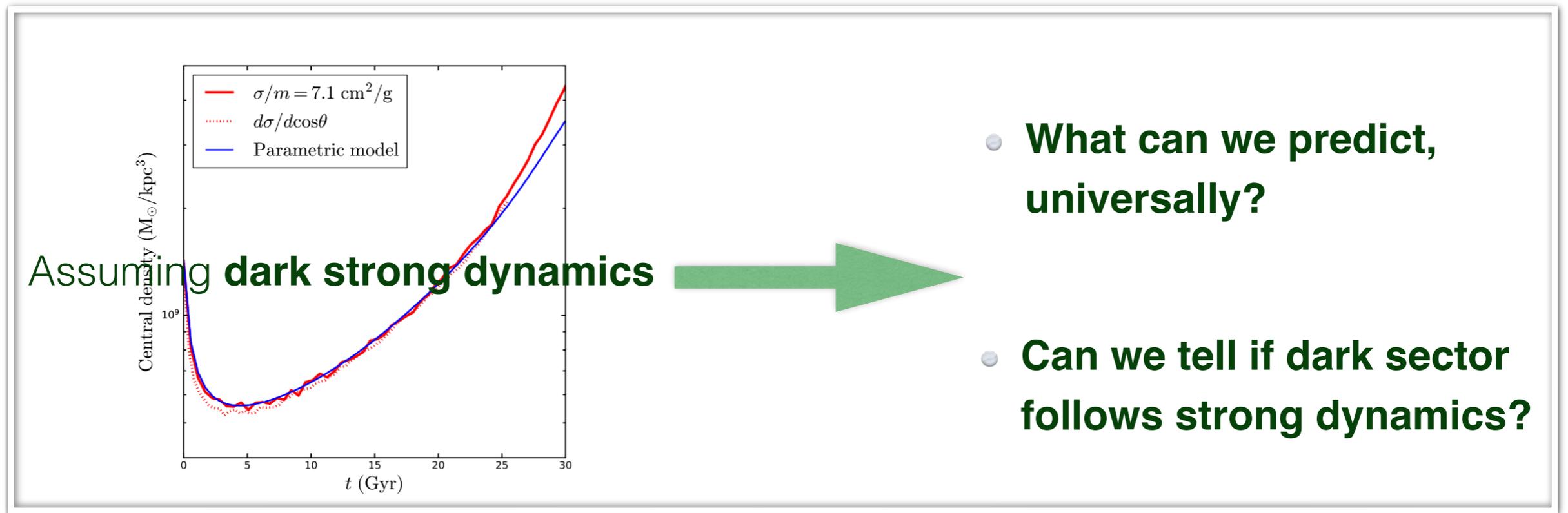
DARK MATTER

$J = ?$

Mass  $m = ?$   
Mean life  $\tau = ?$

DECAY MODES	Fraction ( $\Gamma_j/\Gamma$ )	Confidence level	$\rho$ (MeV/c)
?	?	?	?

Borrowed from A.Ibarra



- **What can we predict, universally?**
- **Can we tell if dark sector follows strong dynamics?**

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***Thanks!***