

Sequential SBI for the LHC

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with Henning Bahl

based on arXiv:2610.XXXX



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Alexander von
HUMBOLDT
STIFTUNG

SBI for the LHC

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Inferring couplings

$$|\mathcal{M}(x|\theta)|^2 = \sum_a w_a(\theta) f_a(x)$$

Polynomial structure in couplings θ

SBI for the LHC

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$$\log \Lambda(\theta, \theta_0) = \sum_{i=1}^n \log \left[\sum_a c_a(\theta) r_a(x_i) \right]$$

The total log-likelihood can be reconstructed from likelihood-ratio estimators trained on a **finite** set of basis processes or coupling structures.

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References

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Brehmer et al. : arXiv:1907.10621
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Bahl et al. : arXiv:2509.05409v2
Chatterjee et al. : arXiv:2107.10859v2
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Feasible since one does not need generate separate simulations for the continuous coupling values θ

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No need to learn parameterized likelihood ratios

SBI for the LHC

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Inferring mass, width

$$\left| M(x | m, \Gamma) \right|^2 \propto \frac{1}{(s - m^2)^2 + m^2 \Gamma^2},$$

Dependence comes through propagators which don't have a polynomial structure.

SBI for the LHC

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$$r(x | \theta) = \frac{p_S(x | \theta)}{p_B(x)}, \quad \theta = (m_1, \Gamma_1, m_2, \Gamma_2, \dots)$$

Need to learn a parameterized likelihood-ratio from simulations covering this continuous parameter space.

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References

- **arXiv:2002.04699: Resonance Searches with Machine Learned Likelihood Ratios**
- **arXiv:2506.00119: Generator Based Inference (GBI)**
- **arXiv:2506.06438: Data-Driven High-Dimensional Statistical Inference with Generative Models**

SBI for the LHC

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$$\left| M(x \mid m, \Gamma) \right|^2 \propto \frac{1}{(s - m^2)^2 + m^2 \Gamma^2},$$

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Generating simulations over a continuous parameter space can become a significant bottleneck:

- expensive detector simulations
- higher-order corrections
- low efficiencies after cuts
- Difficulty in inferring s (reconstructing the invariant mass)

Sequential SBI

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Start with simulations
from a few
parameters from a
proposal distribution

$$q(\theta)$$

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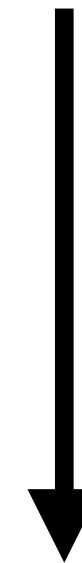
Train a conditional
classifier on the
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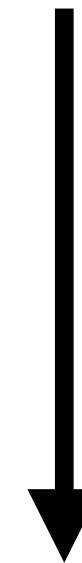
Using the classifier,
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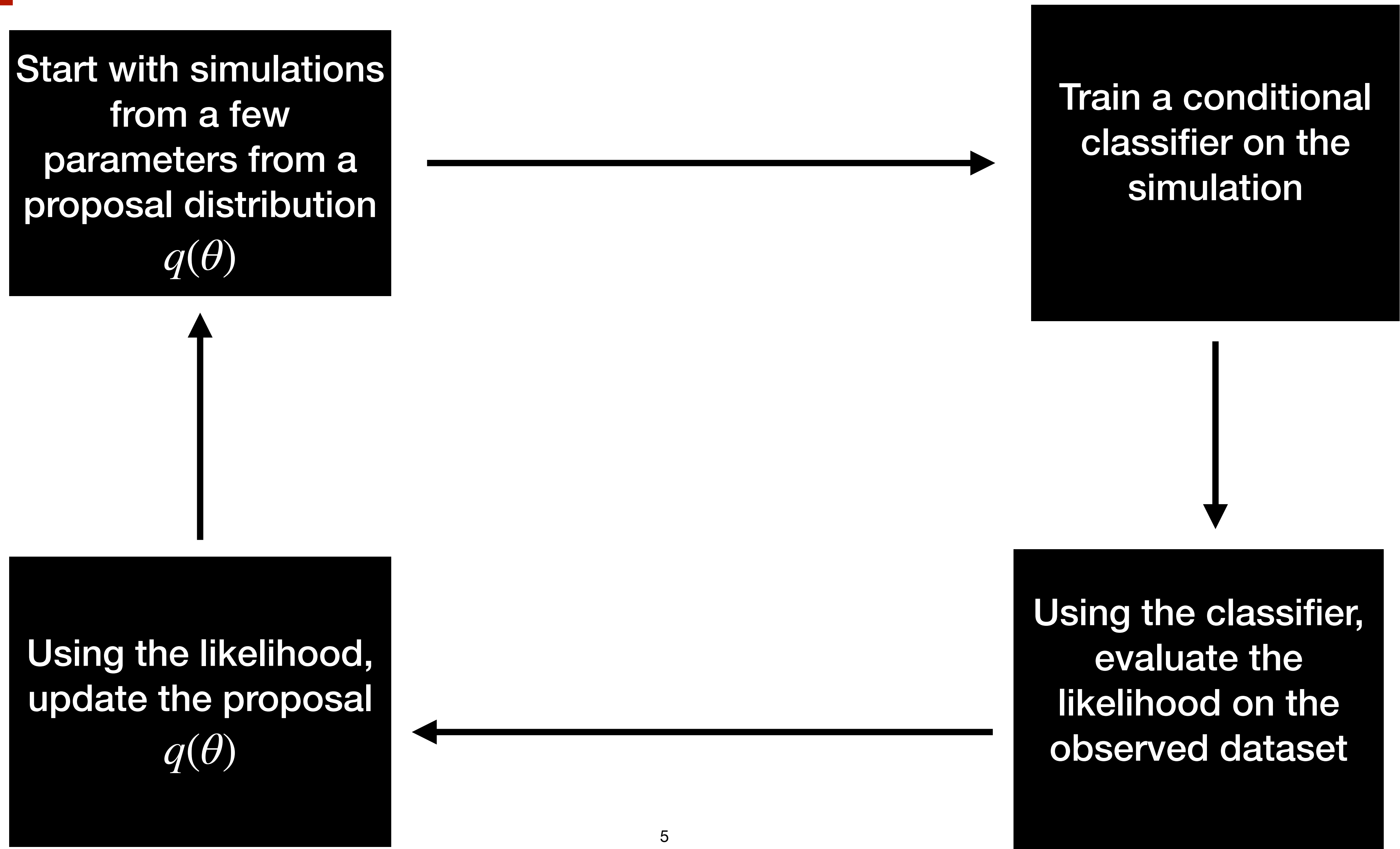


Using the likelihood,
update the proposal
 $q(\theta)$

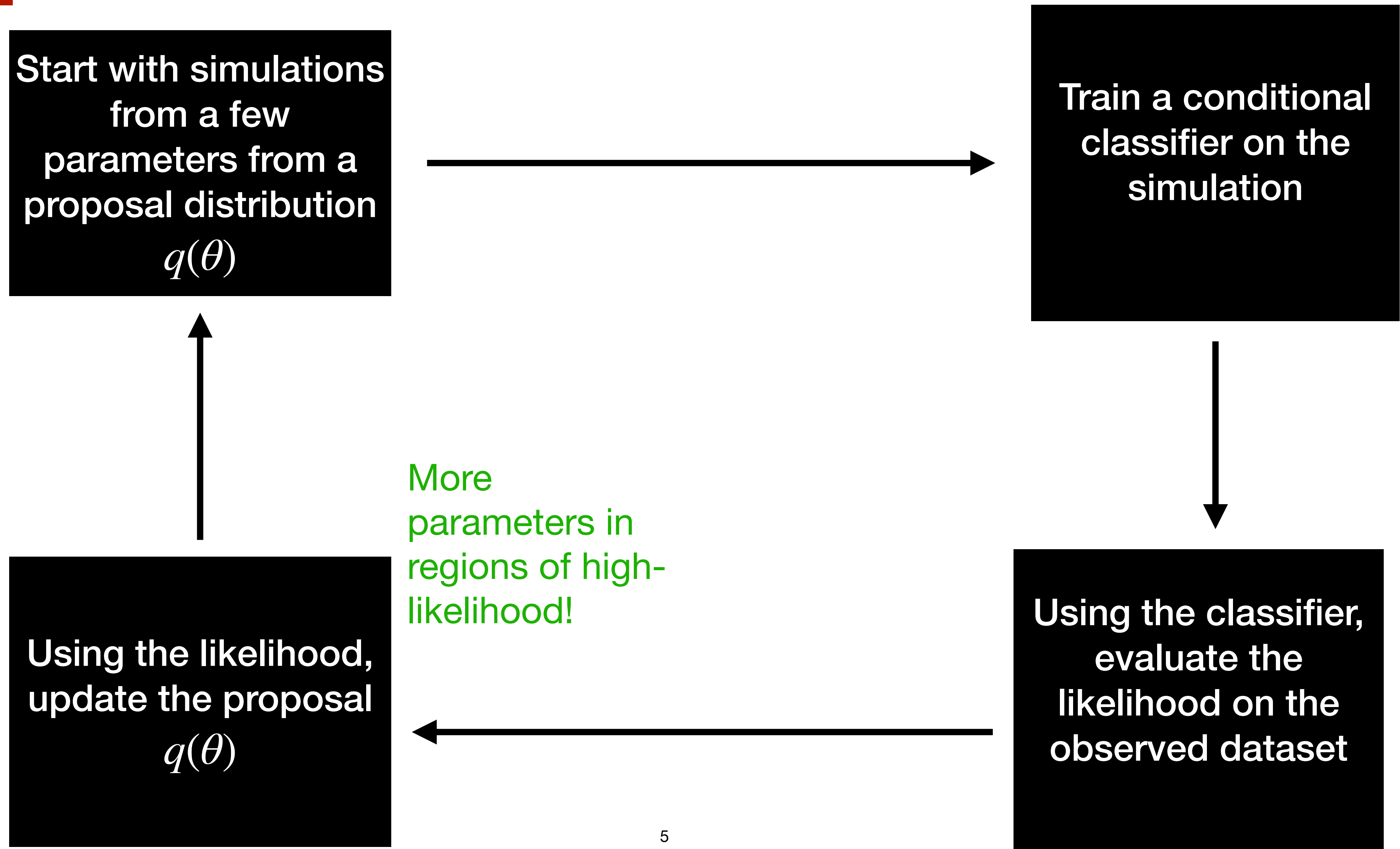


Using the classifier,
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Sequential SBI



Sequential SBI

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$$q_{base}(\theta)$$

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$$\theta_i^{(t)} \sim q_t(\theta)$$

$$x \sim p_S(x \mid \theta_i^{(t)})$$

Simulations
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$$r^{(t)}(x \mid \theta) = \frac{p_S(x \mid \theta)}{p_B(x)}$$

A conditional
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$$\log \Lambda_{\text{shape}}(\theta, B) = \sum_{i=1}^n \log \left[1 + \frac{\sigma_S(\theta)}{\sigma_B} r(x_i | \theta) \right]$$
$$-n \log \left[1 + \frac{\sigma_S(\theta)}{\sigma_B} \right]$$

Evaluate the likelihood on the
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$$q_{t+1}(\theta) \propto q_{\text{base}}(\theta) [\Lambda^{(t)}(\theta)]^\gamma$$

Update proposal

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Evaluate the likelihood on the
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Sequential SBI

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Start with a base distribution as the proposal

With iterations, the proposal focuses on higher likelihood regions

$$\begin{aligned}\theta_i^{(t)} &\sim q_t(\theta) \\ x &\sim p_S(x \mid \theta_i^{(t)})\end{aligned}$$

Simulations from a proposal distribution

$$r^{(t)}(x \mid \theta) = \frac{p_S(x \mid \theta)}{p_B(x)}$$

A conditional classifier

$$\begin{aligned}\log \Lambda_{\text{shape}}(\theta, B) &= \sum_{i=1}^n \log \left[1 + \frac{\sigma_S(\theta)}{\sigma_B} r(x_i \mid \theta) \right] \\ &\quad - n \log \left[1 + \frac{\sigma_S(\theta)}{\sigma_B} \right]\end{aligned}$$

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Evaluate the likelihood on the observed dataset

$0 < \gamma < 1$ encourages exploration beyond the high likelihood regions

Sequential SBI

References

arXiv:1605.06376 : Fast ϵ -free Inference of Simulation Models with Bayesian Conditional Density Estimation

arXiv:1905.07488 : Automatic Posterior Transformation for Likelihood-Free Inference

arXiv:1805.07226 : Sequential Neural Likelihood: Fast Likelihood-free Inference with Autoregressive Flows

arXiv:2002.03712 : On Contrastive Learning for Likelihood-free Inference

arXiv:2107.01214 : Truncated Marginal Neural Ratio Estimation

arXiv:1903.00007 : Fast likelihood-free cosmology with neural density estimators and active learning

arXiv:2111.08030 : Fast and Credible Likelihood-Free Cosmology with Truncated Marginal Neural Ratio Estimation

arXiv:2209.06733 : SICRET: Supernova Ia Cosmology with truncated marginal neural Ratio Estimation

arXiv:2304.02035 : Peregrine: Sequential simulation-based inference for gravitational wave signals

Toy example: Inference at parton level

$$r_{\text{ME}}(x | \theta) = \frac{|\mathcal{M}_S(x | \theta)|^2}{|\mathcal{M}_B(x)|^2}$$

Matrix element ratio at each m, Γ can be obtained from Madgraph

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$$\log \Lambda_{\text{shape}}(g, \theta) = \sum_{i=1}^n \log [1 + g^8 r_{\text{ME}}(x_i | \theta)] - n \log \left[1 + \frac{g^8 \sigma_S(\theta)}{\sigma_B} \right]$$

Exact log likelihood ratio from Madgraph

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Matrix element ratio at each m, Γ can be obtained from Madgraph

Compare:

$$\log \Lambda_{\text{shape}}(g, \theta) = \sum_{i=1}^n \log [1 + g^8 r_{\text{ME}}(x_i | \theta)] - n \log \left[1 + \frac{g^8 \sigma_S(\theta)}{\sigma_B} \right]$$

Exact log likelihood ratio from Madgraph

$$\log \Lambda_{\text{shape}}(g, \theta) = \sum_{i=1}^n \log \left[1 + \frac{g^8 \sigma_S(\theta)}{\sigma_B} r(x_i | \theta) \right] - n \log \left[1 + \frac{g^8 \sigma_S(\theta)}{\sigma_B} \right]$$

Classifier based log-likelihood ratio

Toy example: Inference at parton level

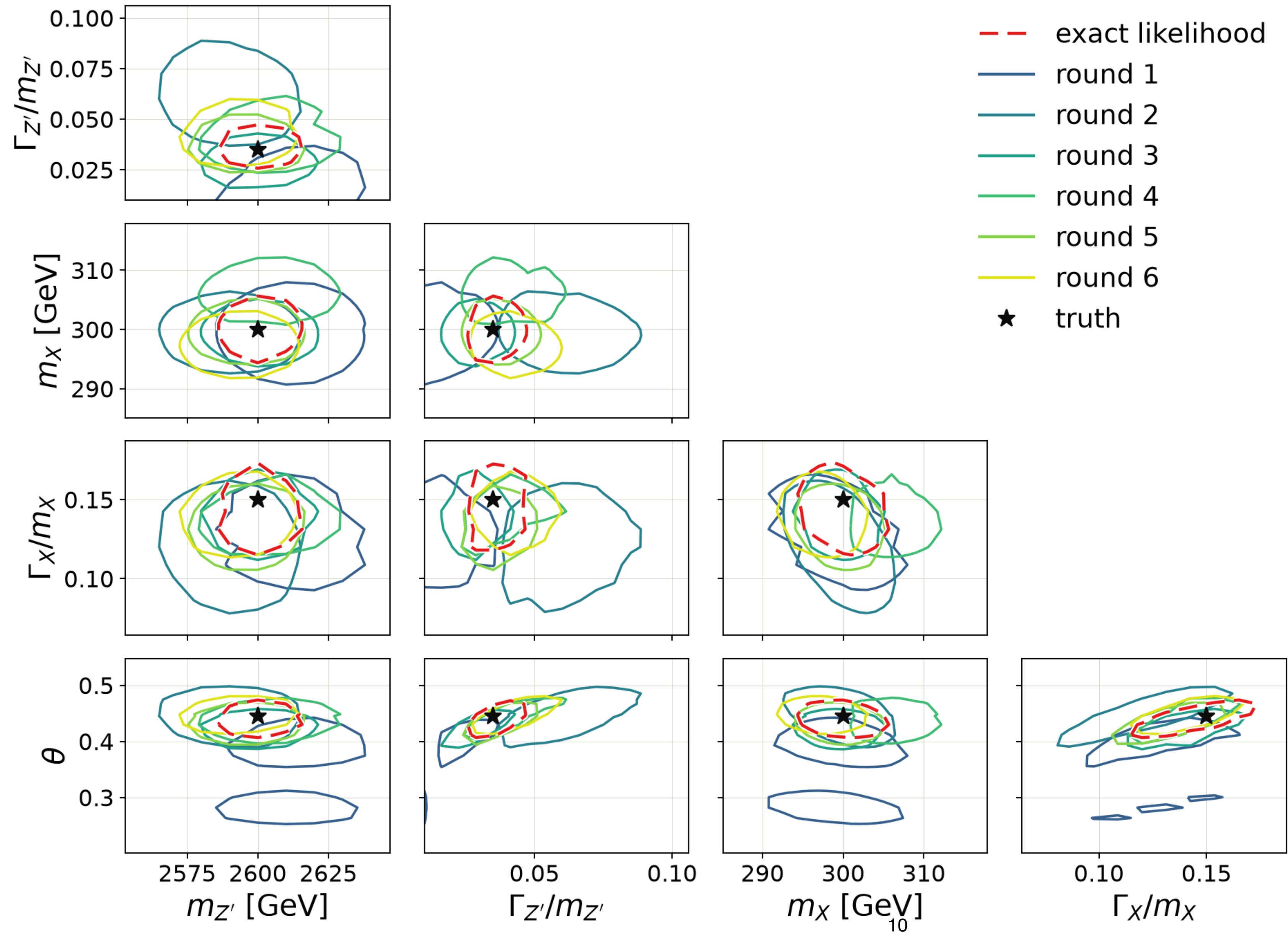
Signal: $u\bar{u} \rightarrow Z' \rightarrow X\bar{X} \rightarrow (u\bar{d})(\bar{u}d)$

Background: $u\bar{u} \rightarrow (u\bar{d})(\bar{u}d)$

$\theta : \theta_{shape} = (m_{Z'}, \Gamma_{Z'}, m_X, \Gamma_X)$ and g

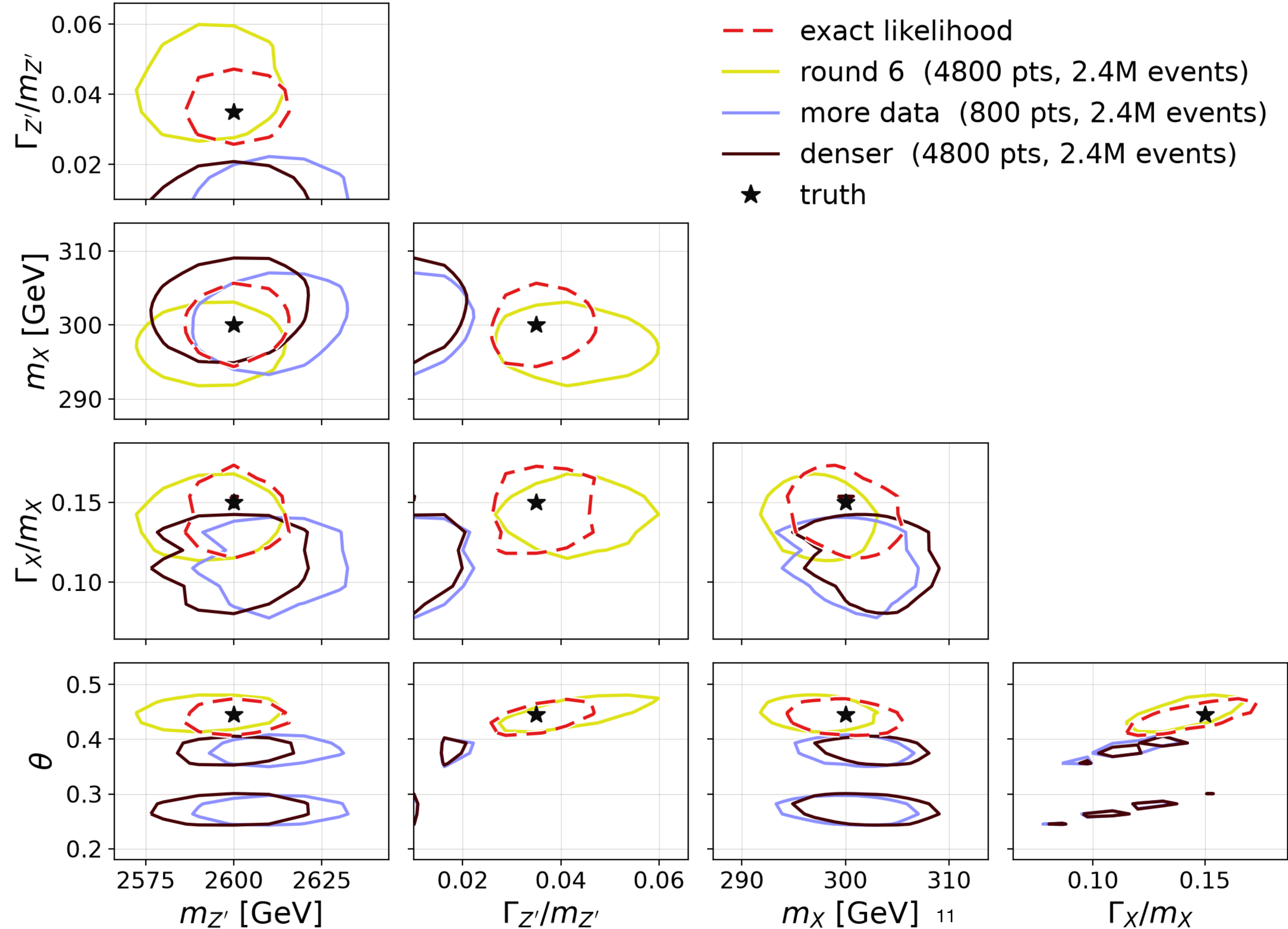
$$m_{Z'} = 2600 \text{ GeV}, \quad \frac{\Gamma_{Z'}}{m_{Z'}} = 0.035, \quad m_X = 300 \text{ GeV}, \quad \frac{\Gamma_X}{m_X} = 0.15$$

Toy example: Inference at parton level



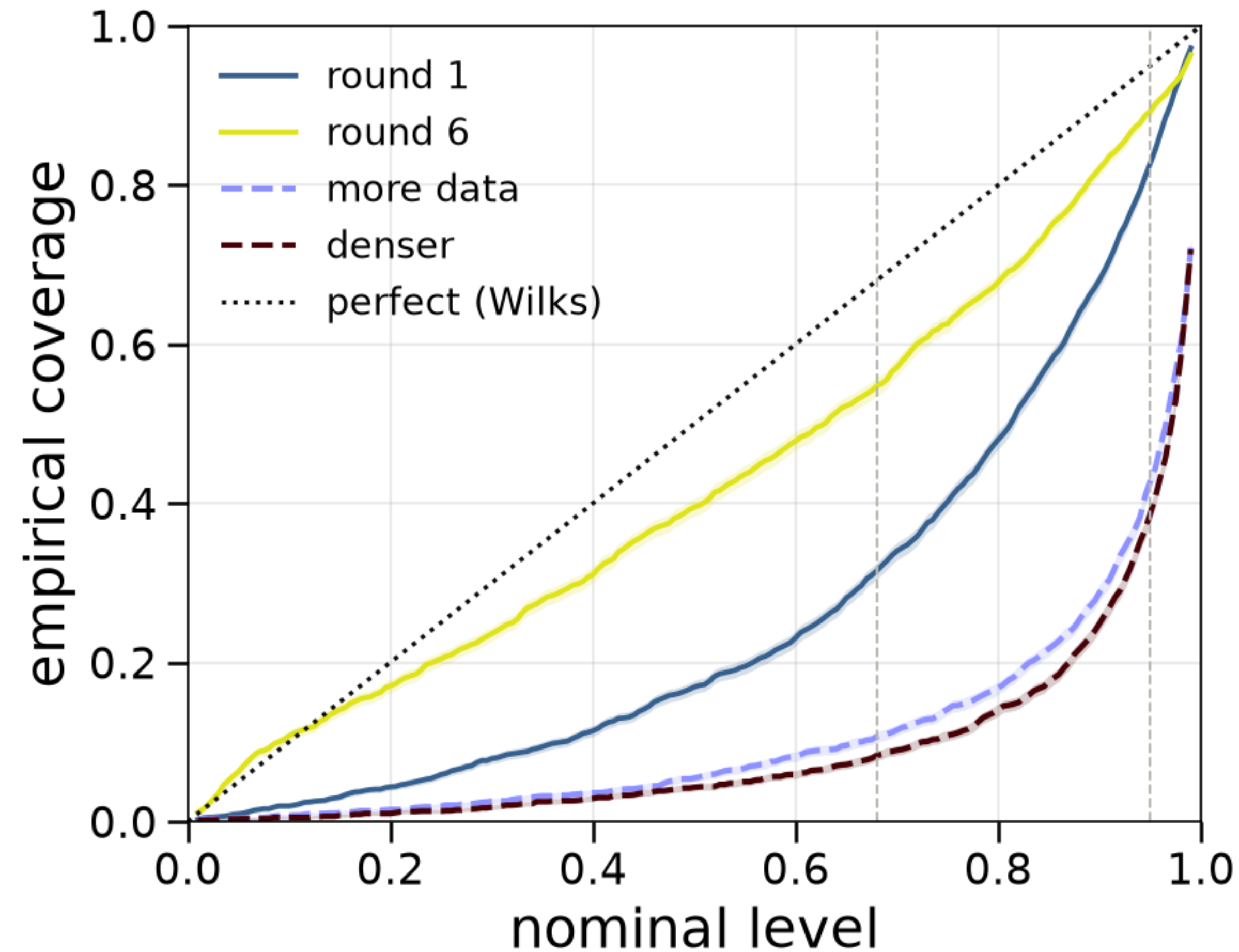
Rounds converge
to true confidence
regions!

Toy example: Inference at parton level



Iterative rounds do better single round with the same simulation budget!

Toy example: Inference at parton level



Coverage improves with rounds!

Inference at reco-level

Signal: $pp \rightarrow Y \rightarrow HH \rightarrow (t\bar{t})(t\bar{t})$

Background: $pp \rightarrow$ QCD Di-jet

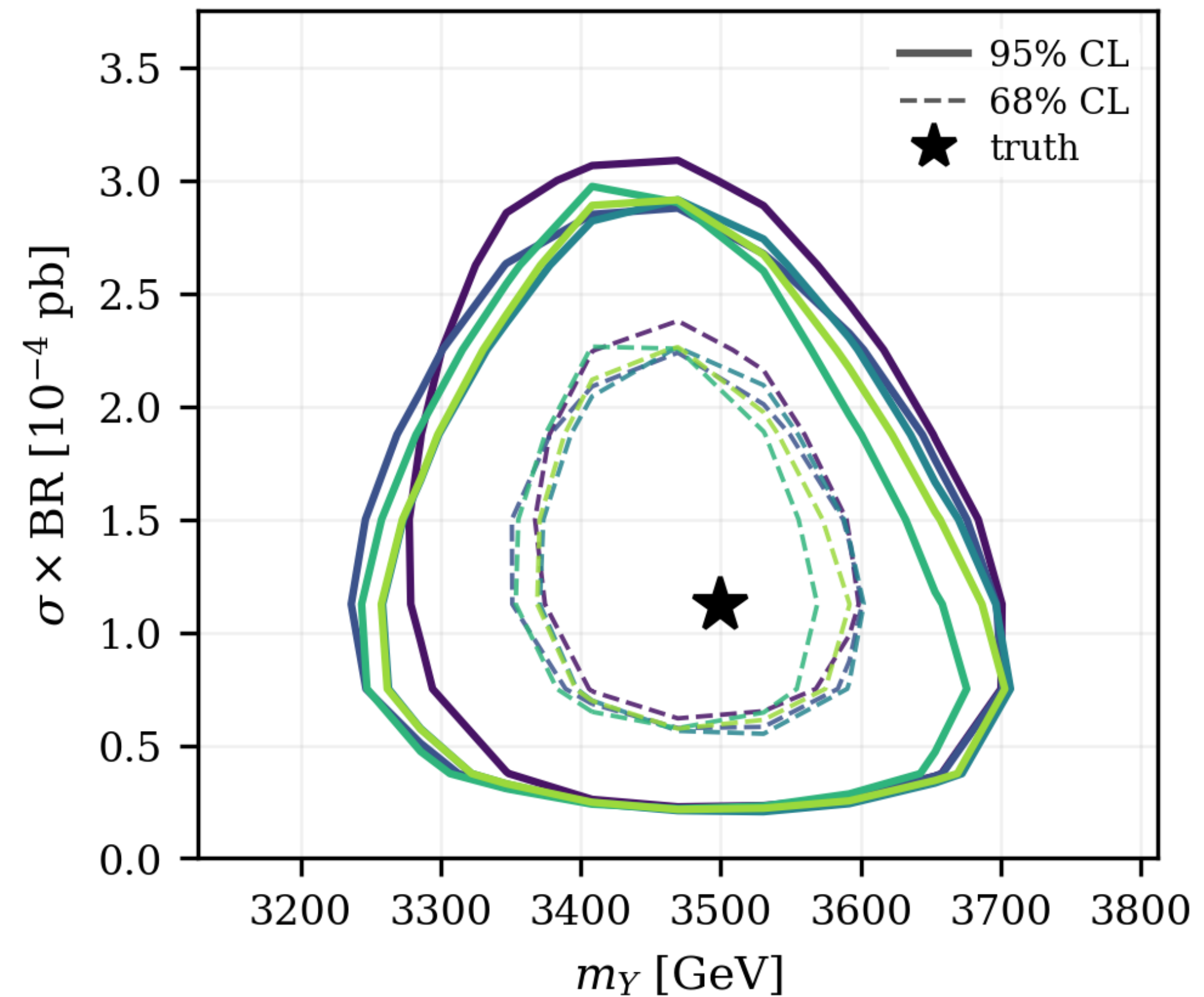
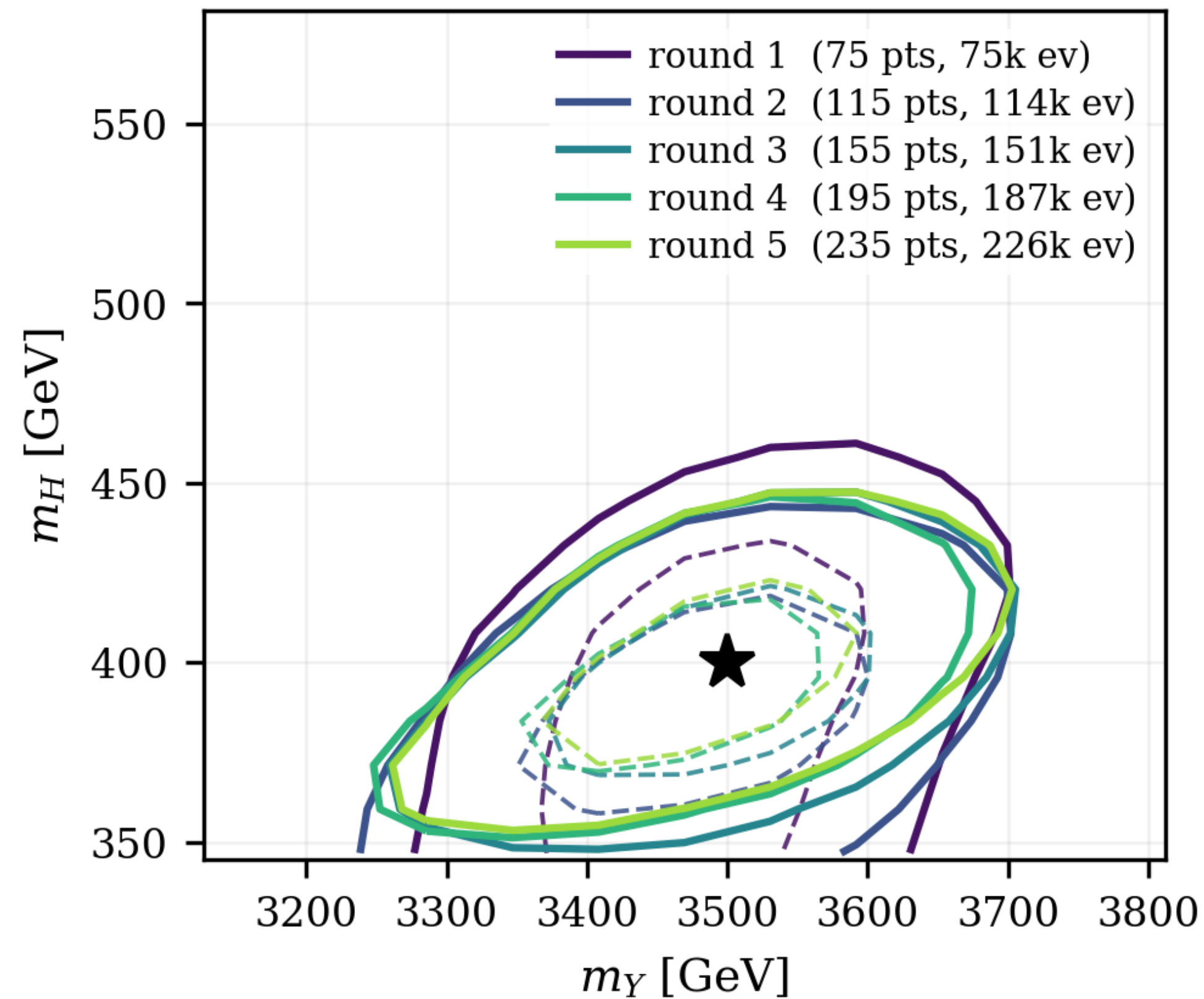
$\theta : \theta_{shape} = (m_Y, m_H)$ and σ_{sig}

$m_Y = 3500 \text{ GeV}, \quad m_H = 400 \text{ GeV} \quad \sigma_{sig} = 0.010928 \text{ pb}$

Parameterized classifier trained on full reconstructed jets

L-GATr-slim (arXiv:2512.17011)

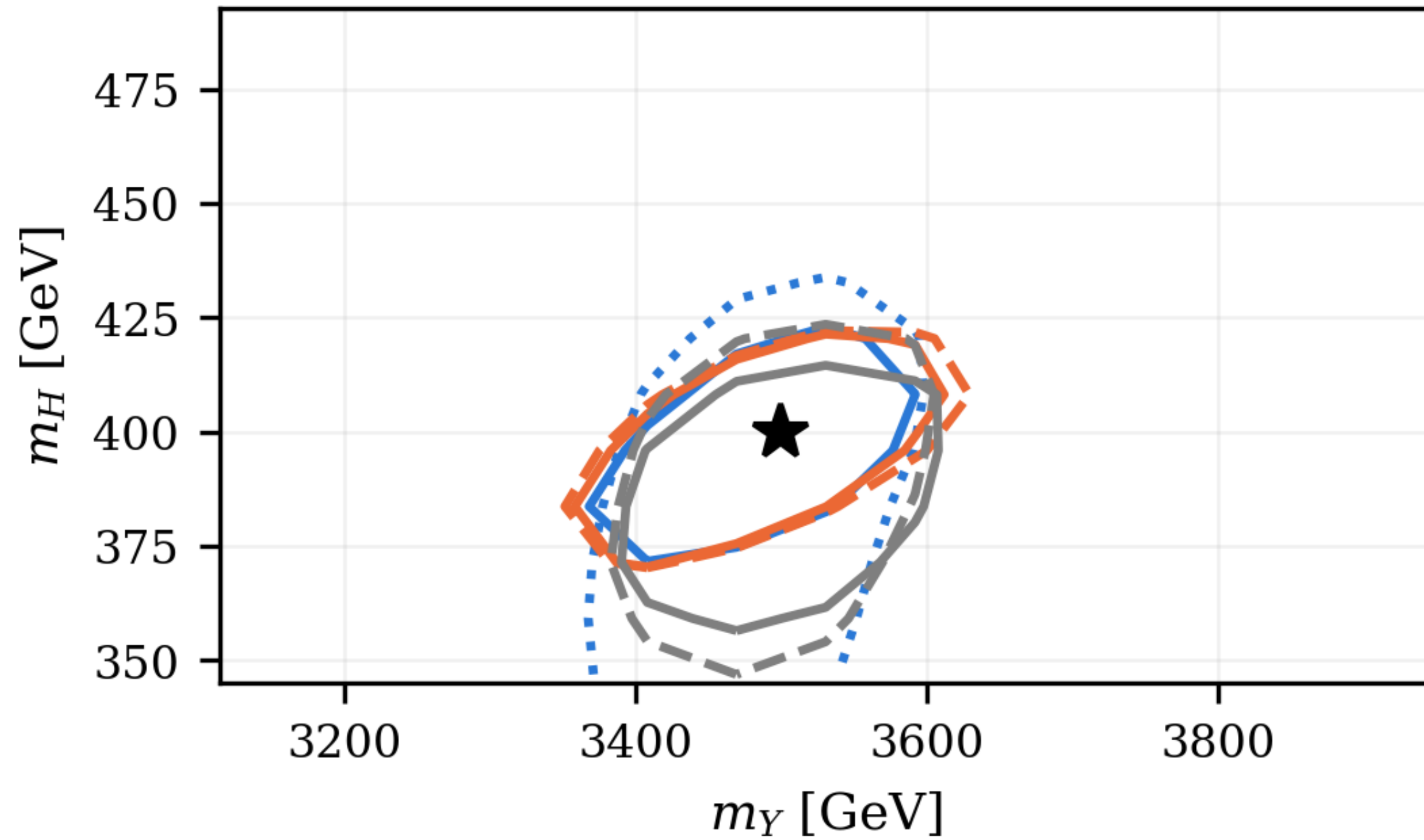
Inference at reco-level (Preliminary Results)



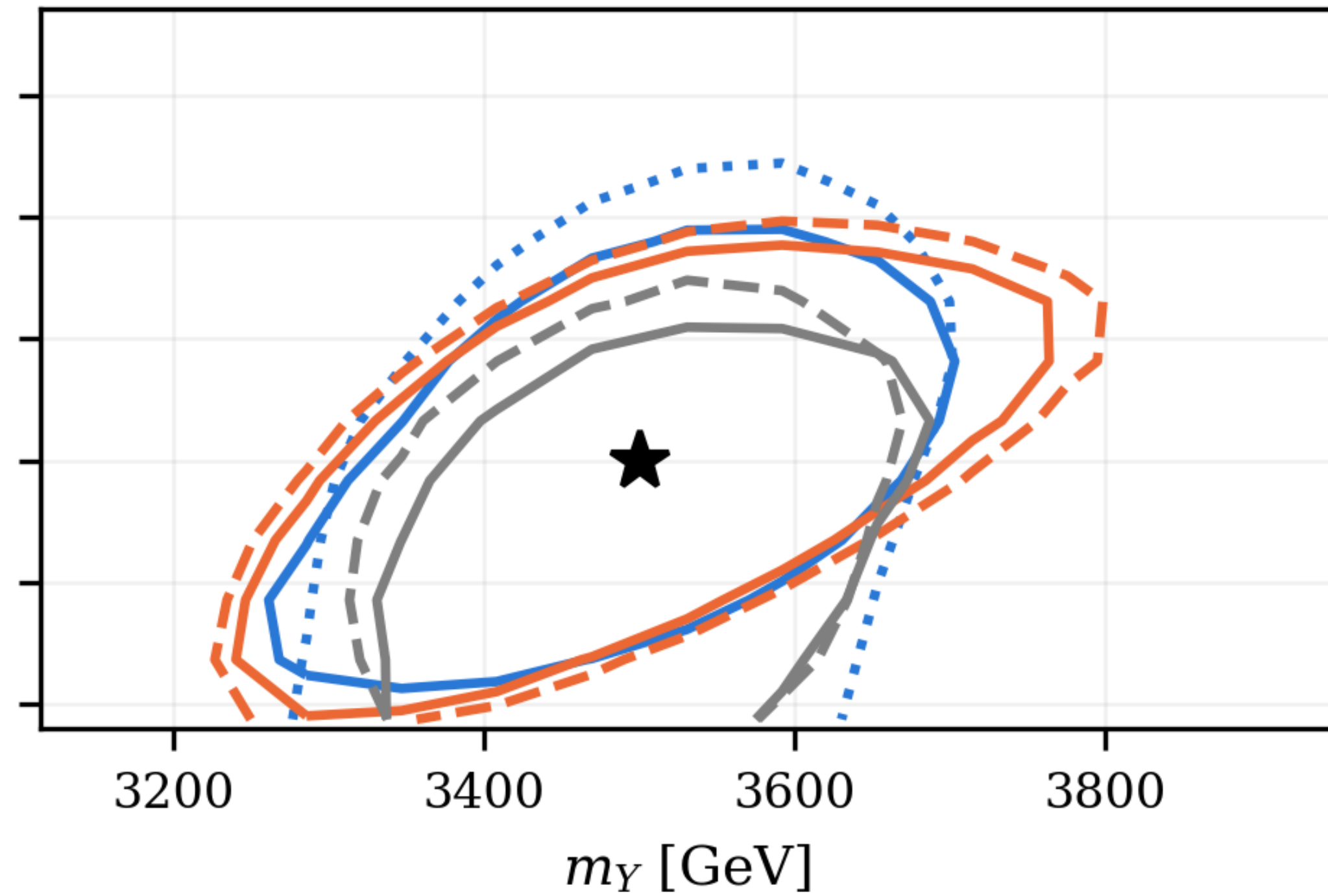
Rounds converge to tighter 68% confidence regions

Inference at reco-level (Preliminary Results)

1σ region (68% CL)



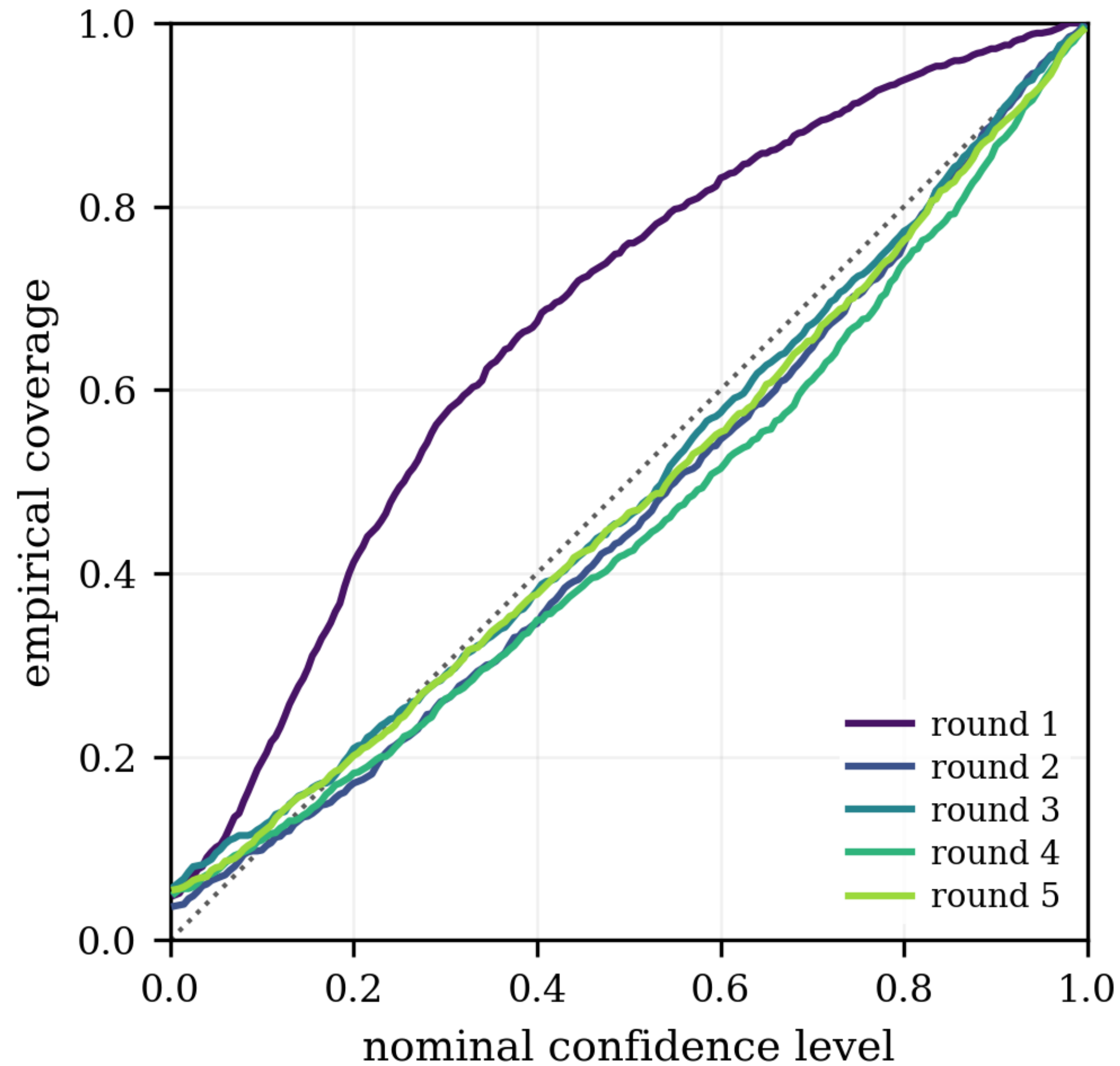
2σ region (95% CL)



- | | | |
|---------------------------------|-------------------------------------|-----------------------------------|
| — iterative, round 5 (226k ev) | — 30 × 25 grid, 1k/pt (699k ev) | — round-1 grid, 10k/pt (699k ev) |
| ... iterative, round 1 (75k ev) | - - 30 × 25 grid, 0.5k/pt (365k ev) | - - round-1 grid, 5k/pt (375k ev) |

Tighter 68% and 95% confidence regions than single round with the more simulations

Inference at reco-level (Preliminary Results)



Iterations don't make the coverage under-confident or over-confident.

Conclusions

- For inferring mass/widths generating simulations can be a bottleneck.
- Sequential-SBI generates simulations for parameters closer to high likelihood regions, resulting in more efficient simulations
- For parton level and reco-level, iterative SBI results in better confidence regions as well as improved coverage

THANK YOU