



Oskar Klein
centre



Rymdstyrelsen
Swedish National Space Agency

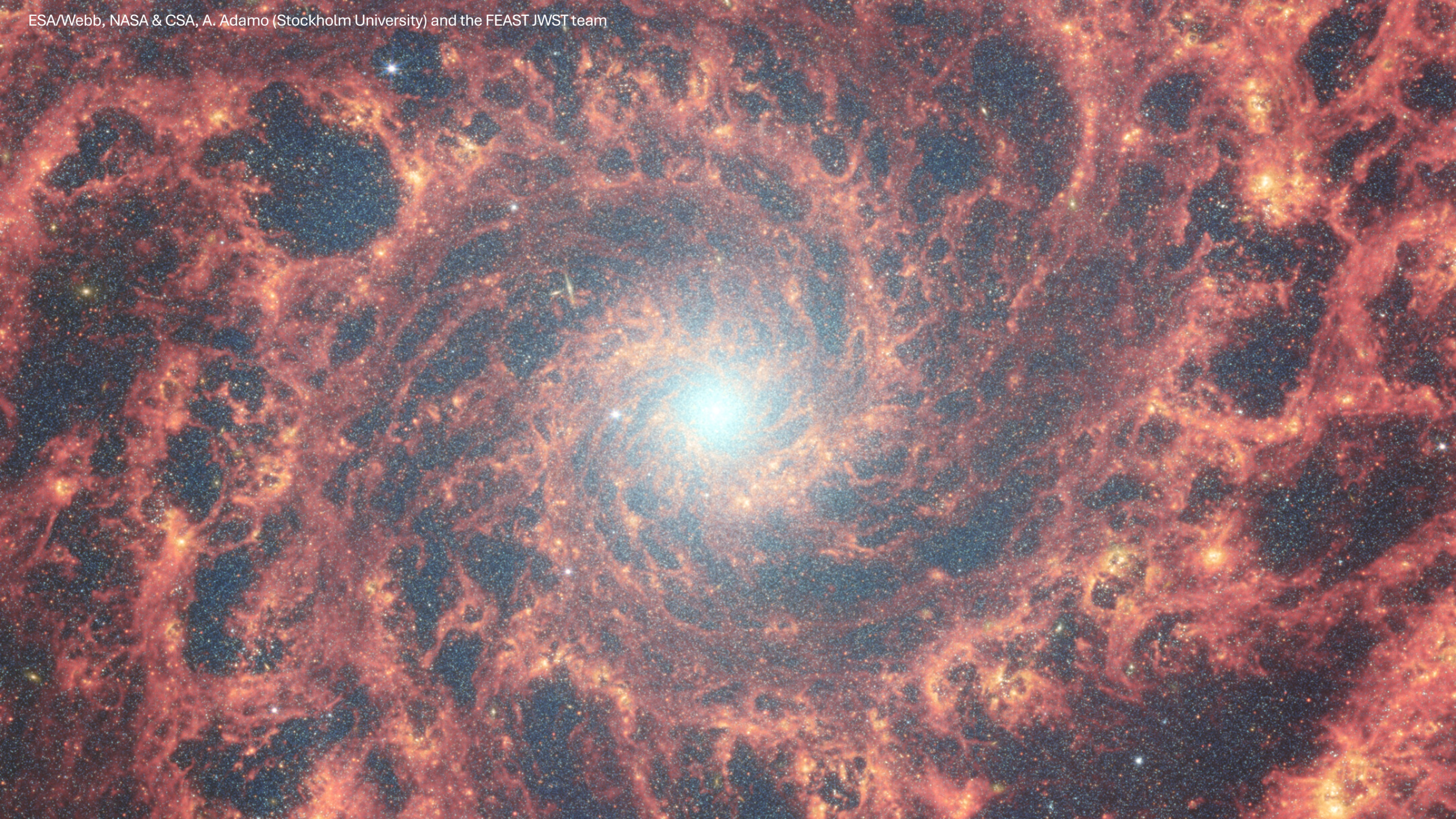


Stockholm
University

FEAST: a JWST/NIRSpec survey of emerging young star clusters in NGC628

Helena Faustino Vieira

with Angela Adamo, Arjan Bik, Neville Shane, Linda Smith, Alex Pedrini, Giacomo Bortollini
+ FEAST collaboration





Star formation

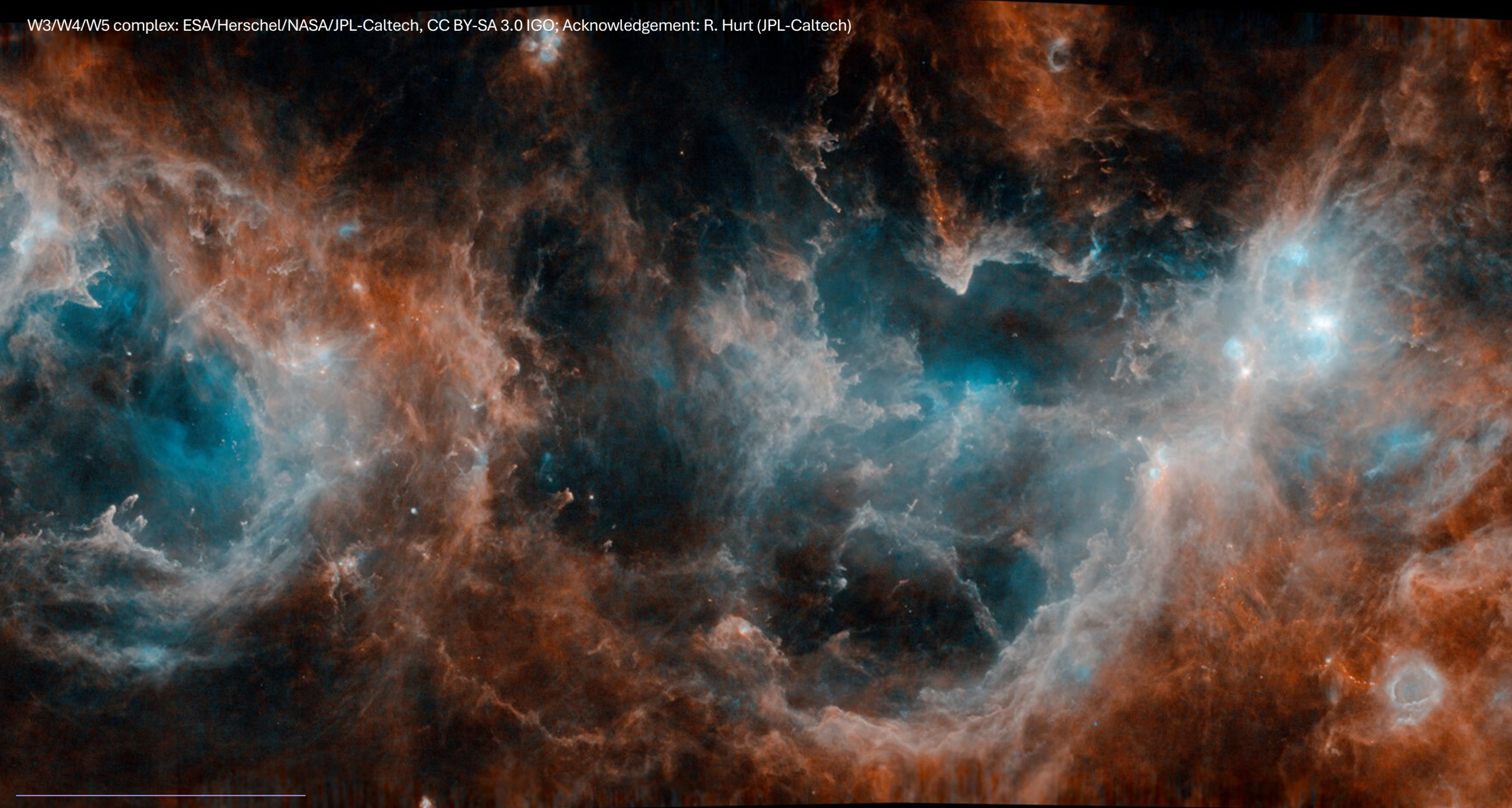
Stellar feedback

- Stellar winds
- Radiation
- Supernova

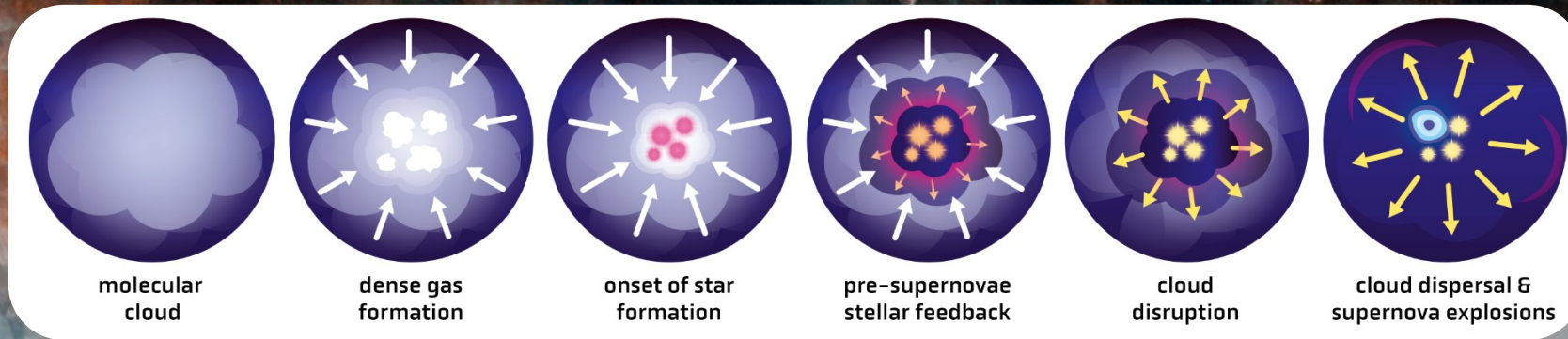


Star formation





Different wavelengths trace different stages in the star formation process

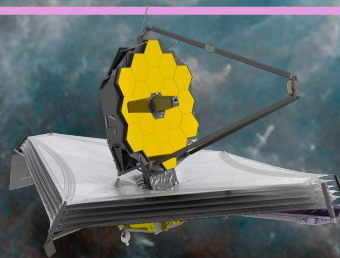


from Schinnerer&Leroy 2024

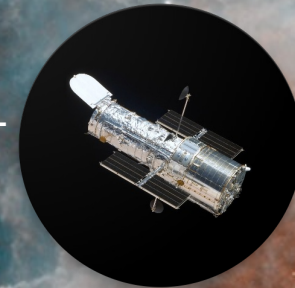
ALMA



JWST



HST



RADIO

MICROWAVE

INFRARED

VISIBLE

ULTRAVIOLET

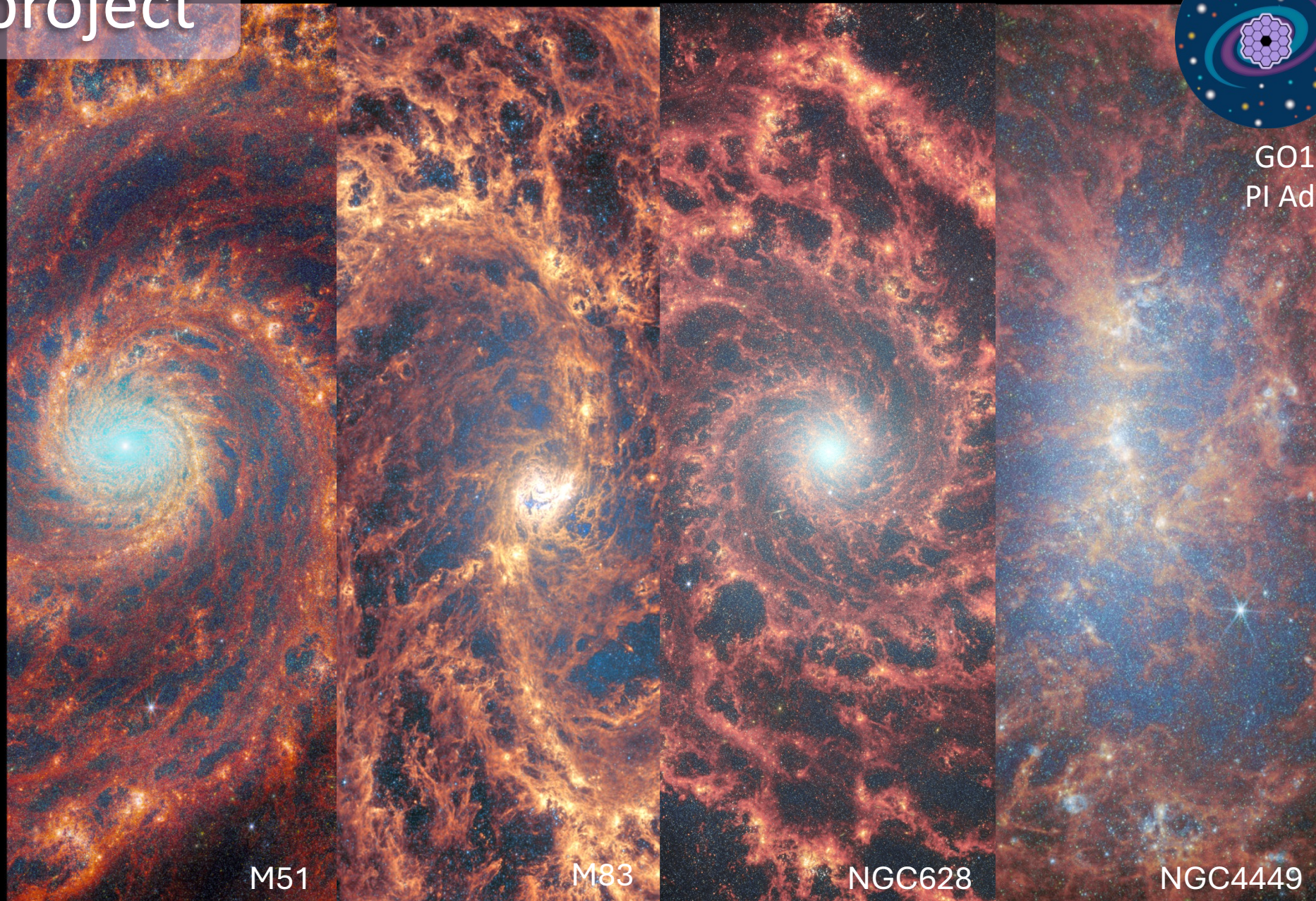
The FEAST project

Feedback in
Emerging
extragalactic Star
clusters

feast-survey.github.io



GO1783
PI Adamo

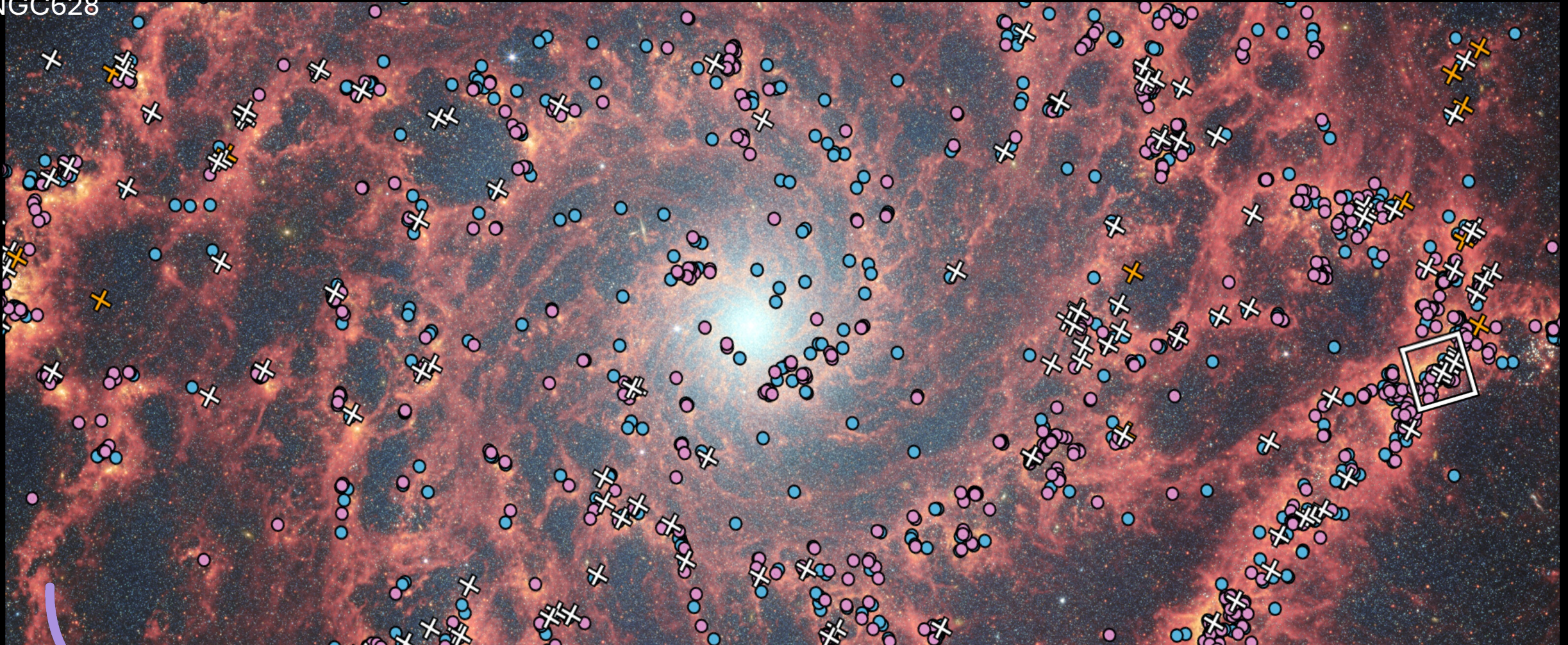


M51

M83

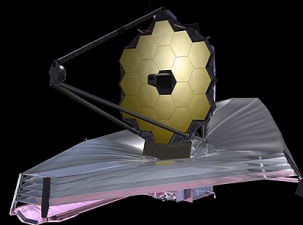
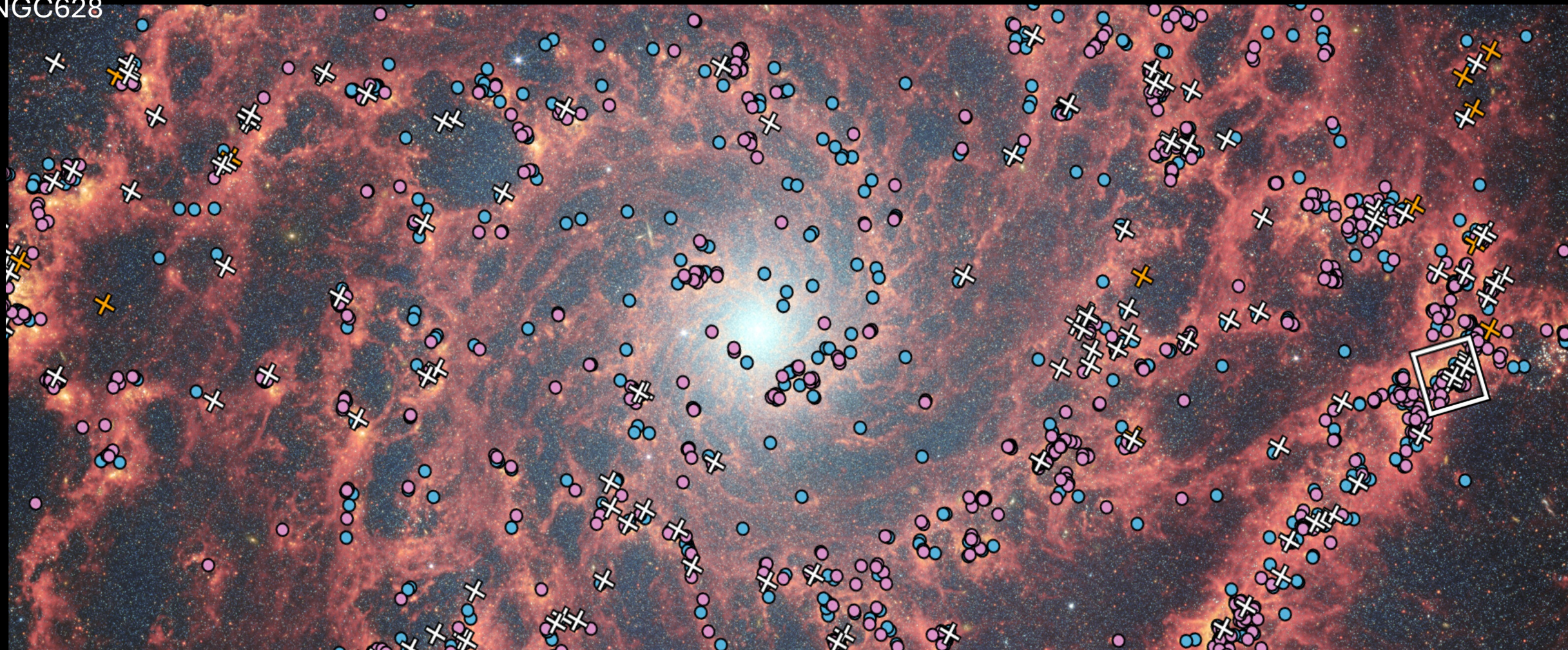
NGC628

NGC4449

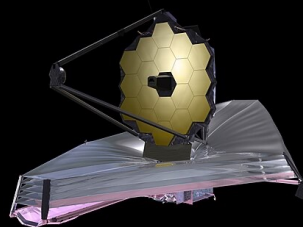
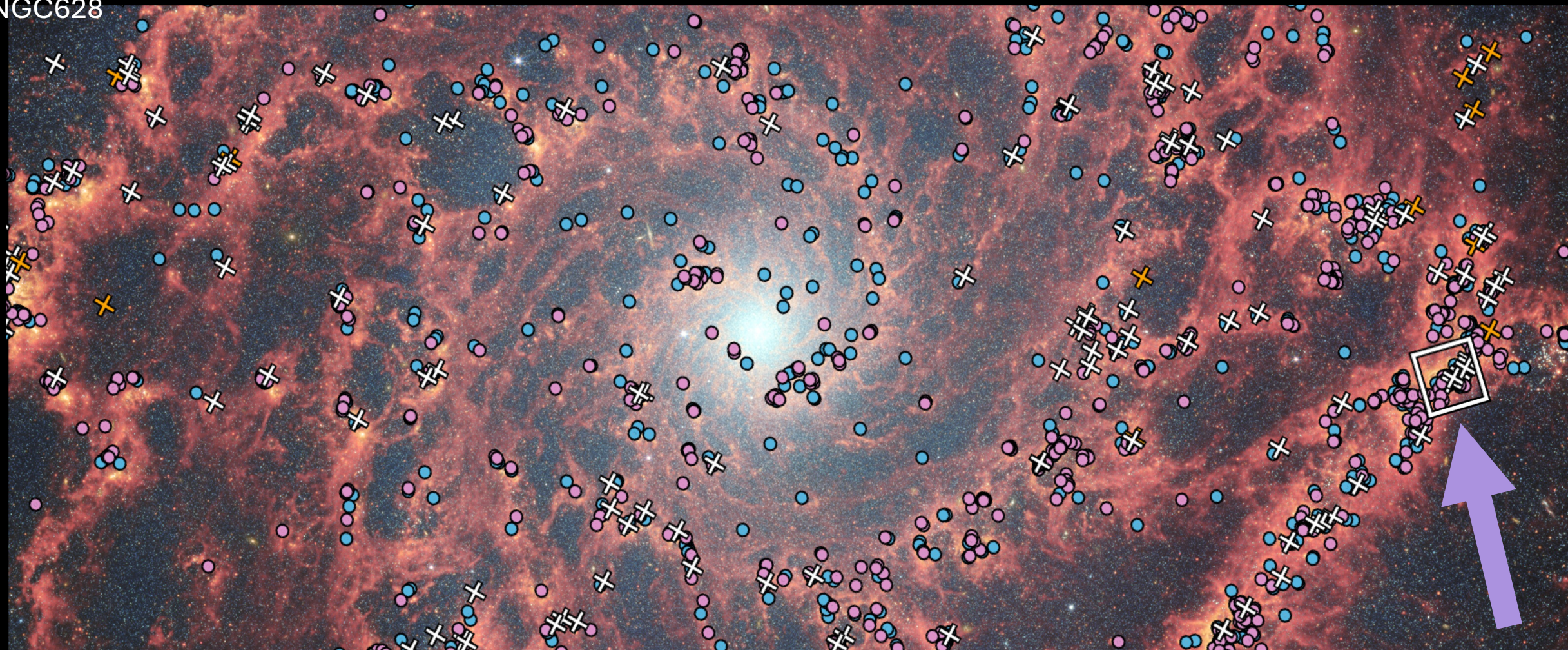


Emerging young star clusters (eYSCs) – compact peaks of $\text{Pa}\alpha$ and/or $3.3\mu\text{m}$ PAH

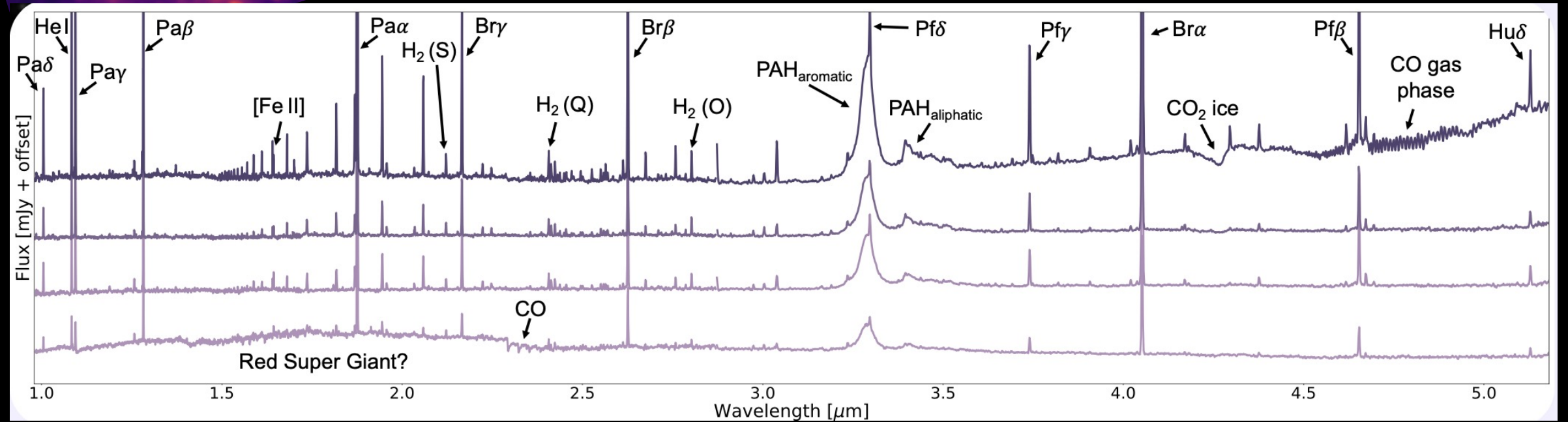
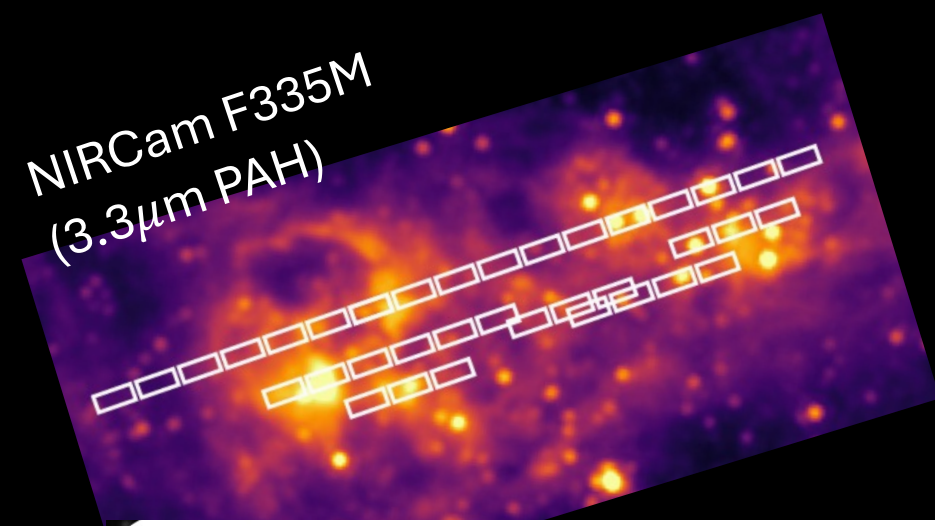
Not visible in the optical yet: eYSCs are still emerging/clearing their natal molecular cloud, and thus actively driving stellar feedback (Pedrini+2024,2025; Knutas+2025, Adamo+in prep)



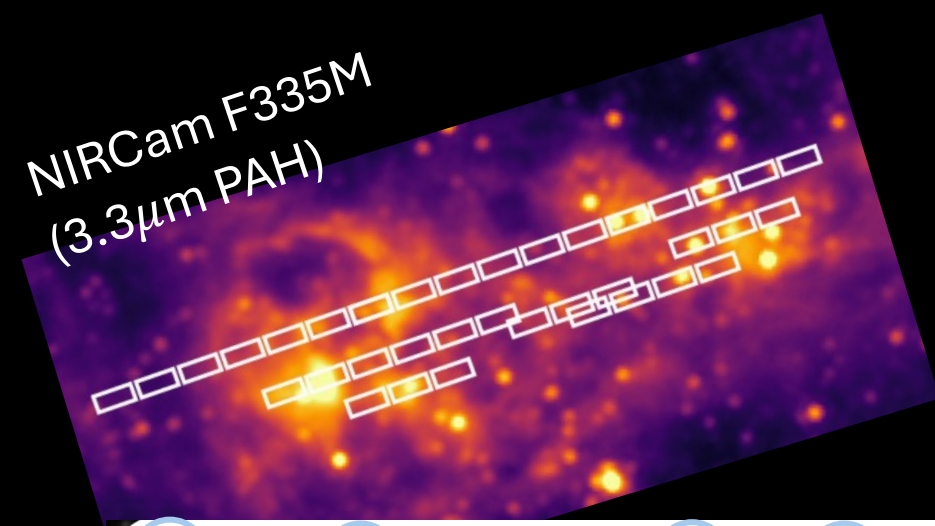
Observed with NIRSpect/MOS (GO3503)
– spectroscopy in the 1 – 5 μ m range



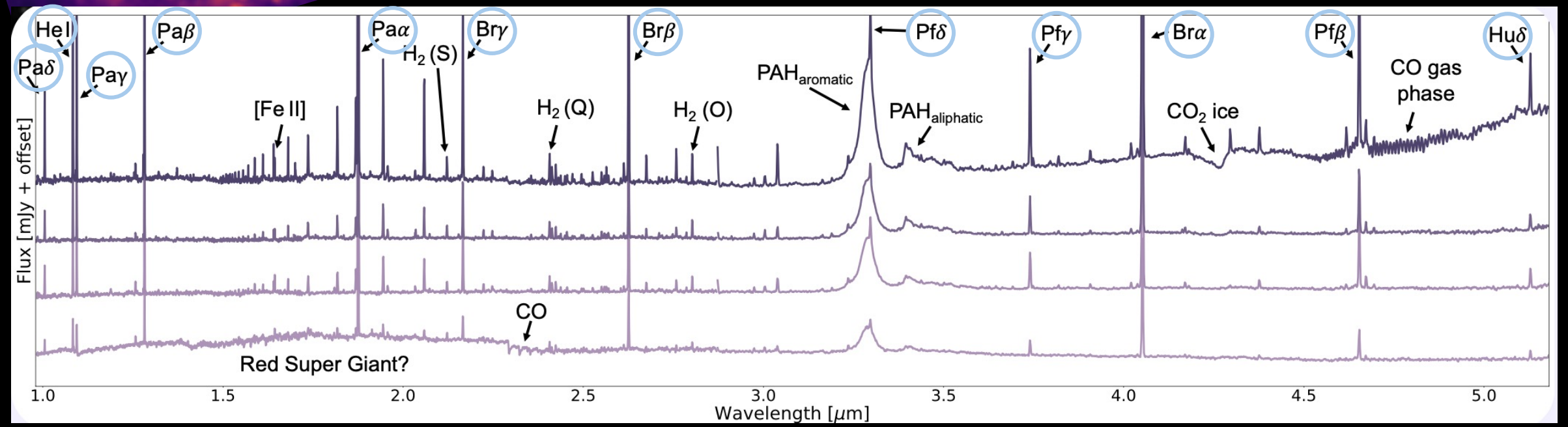
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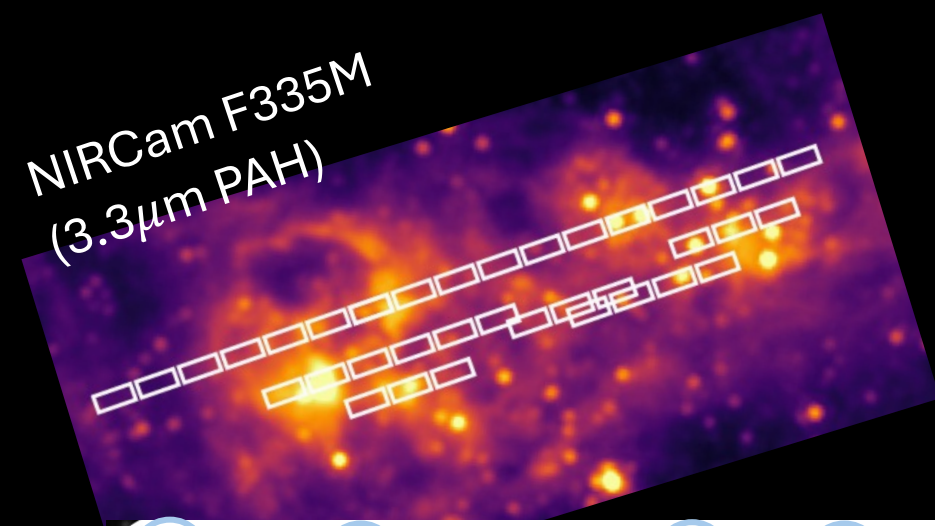


Faustino Vieira+2026

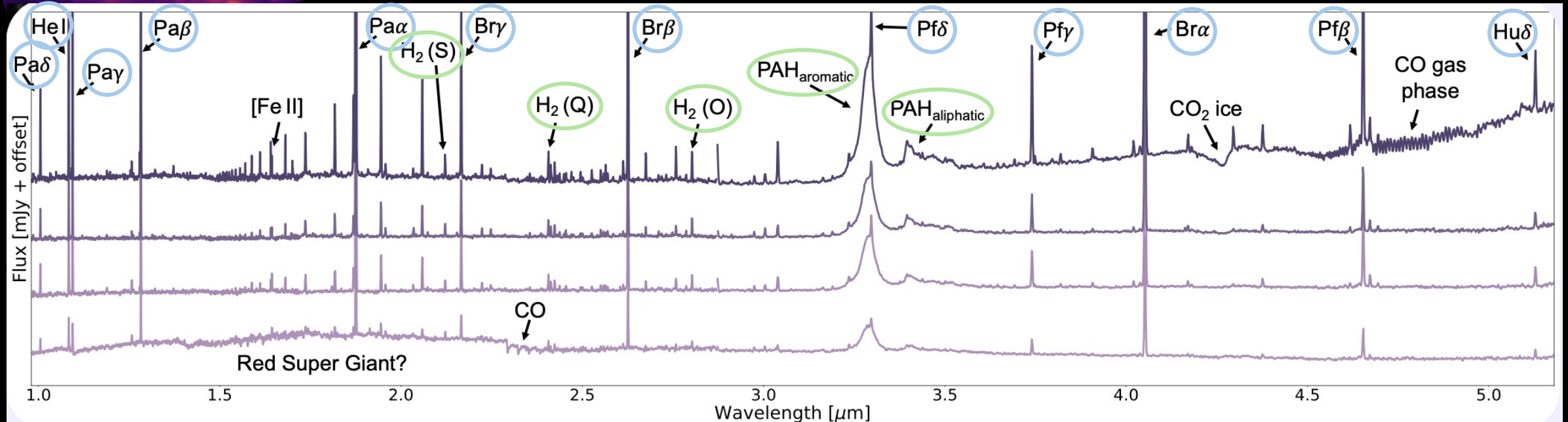


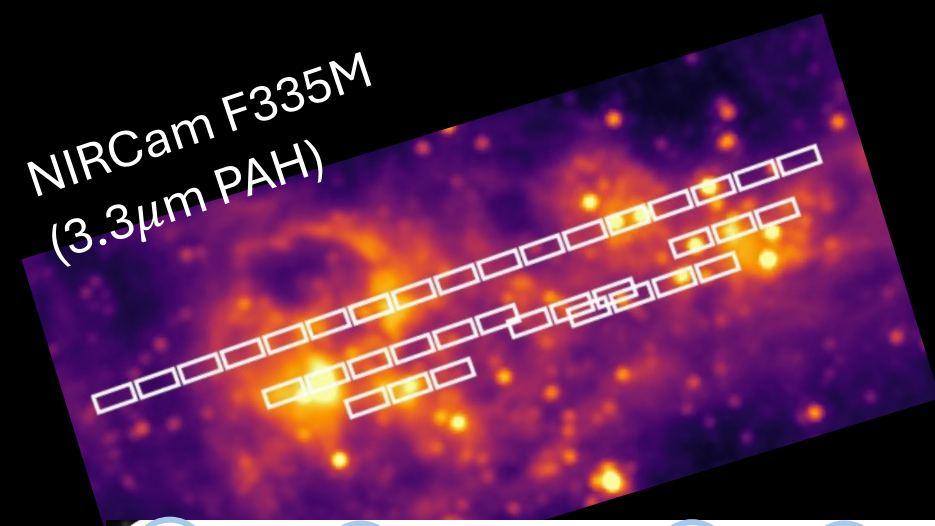
Ionization – multiple H and He lines



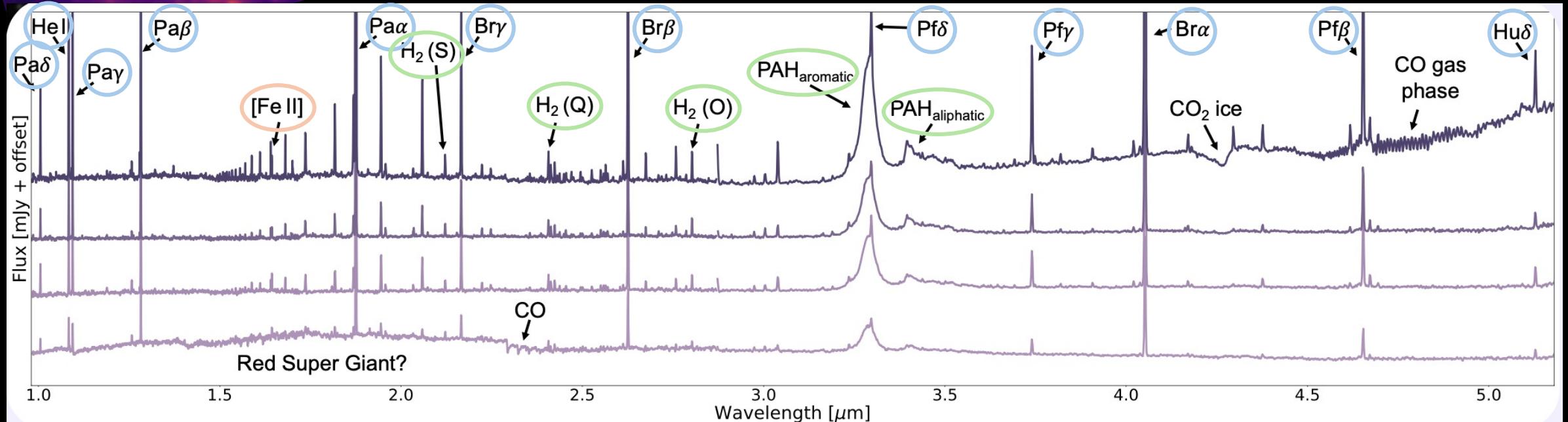


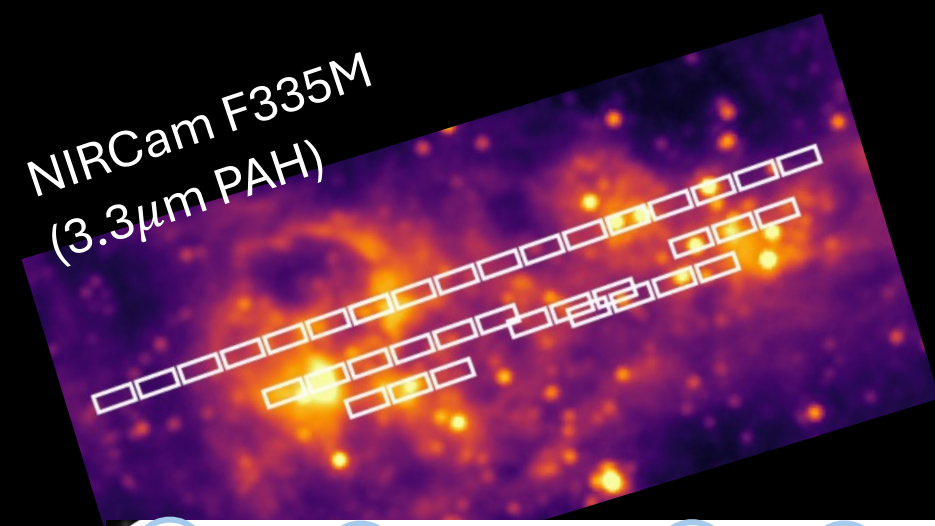
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 Gas/Dust – forest of H₂ lines; 3.3 μ m PAH emission





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 Stellar feedback diagnostics – [FeII]/Br γ , H₂ ratios



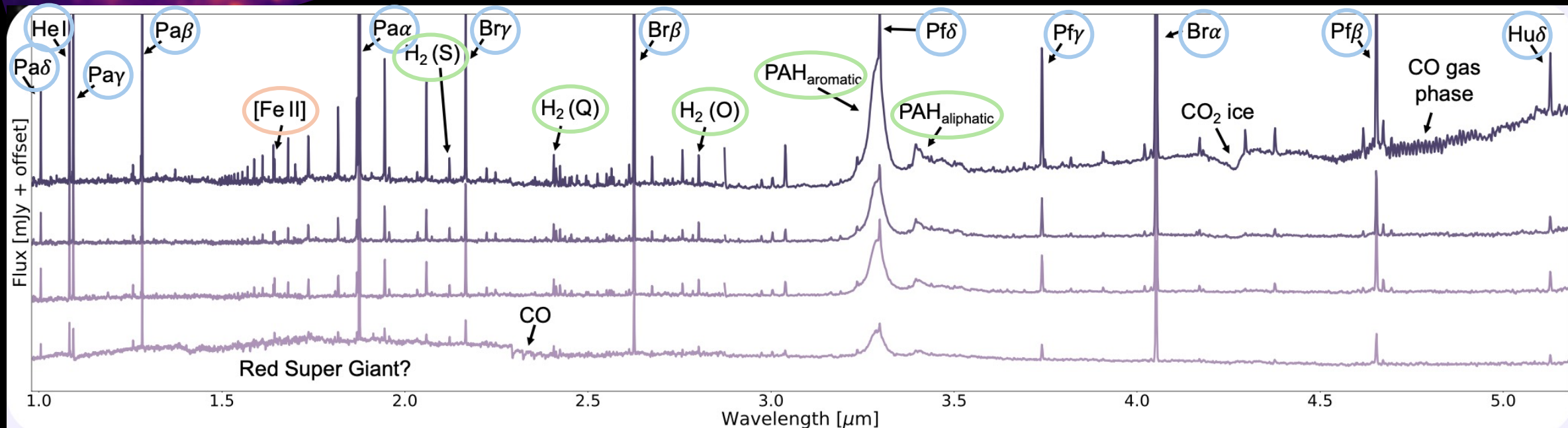


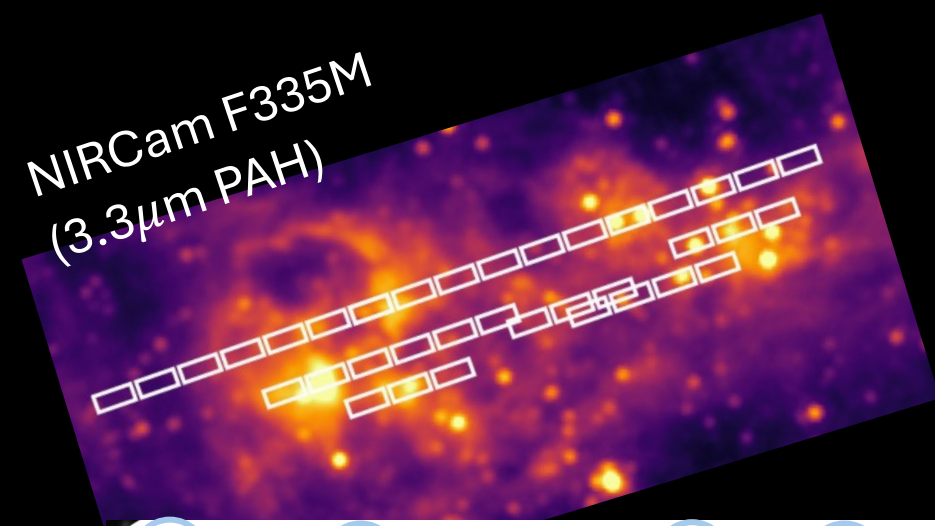
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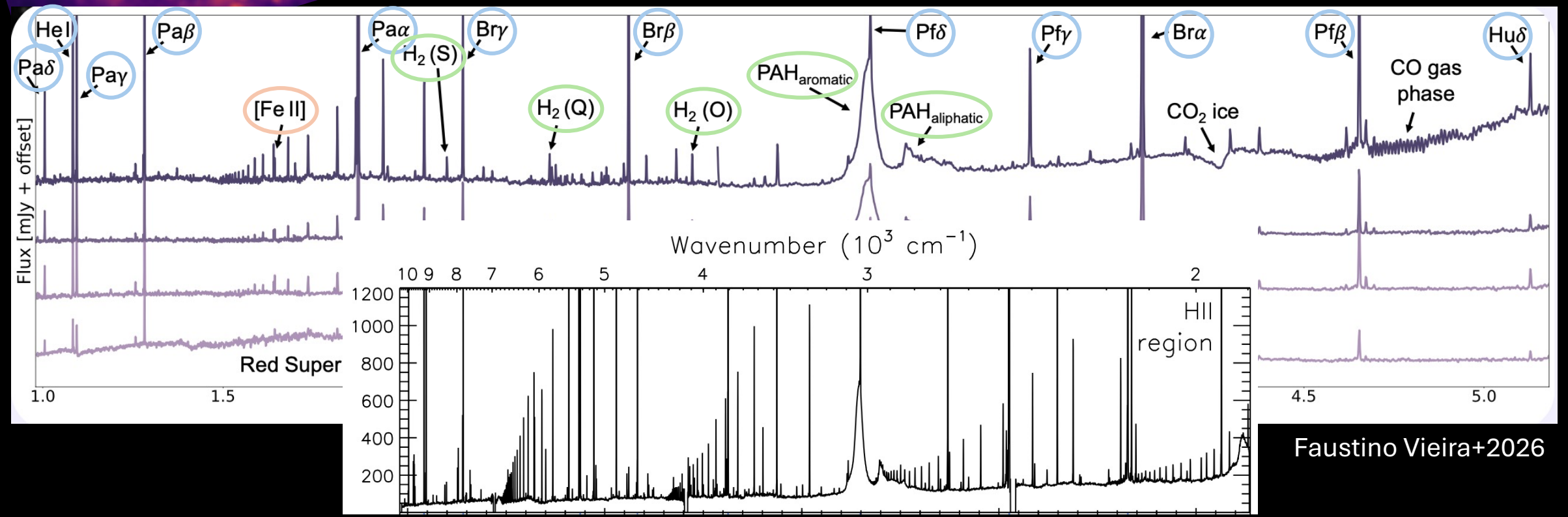
Stellar feedback diagnostics – [FeII]/Br γ , H₂ ratios

and other interesting features: gas-phase CO, ices, etc





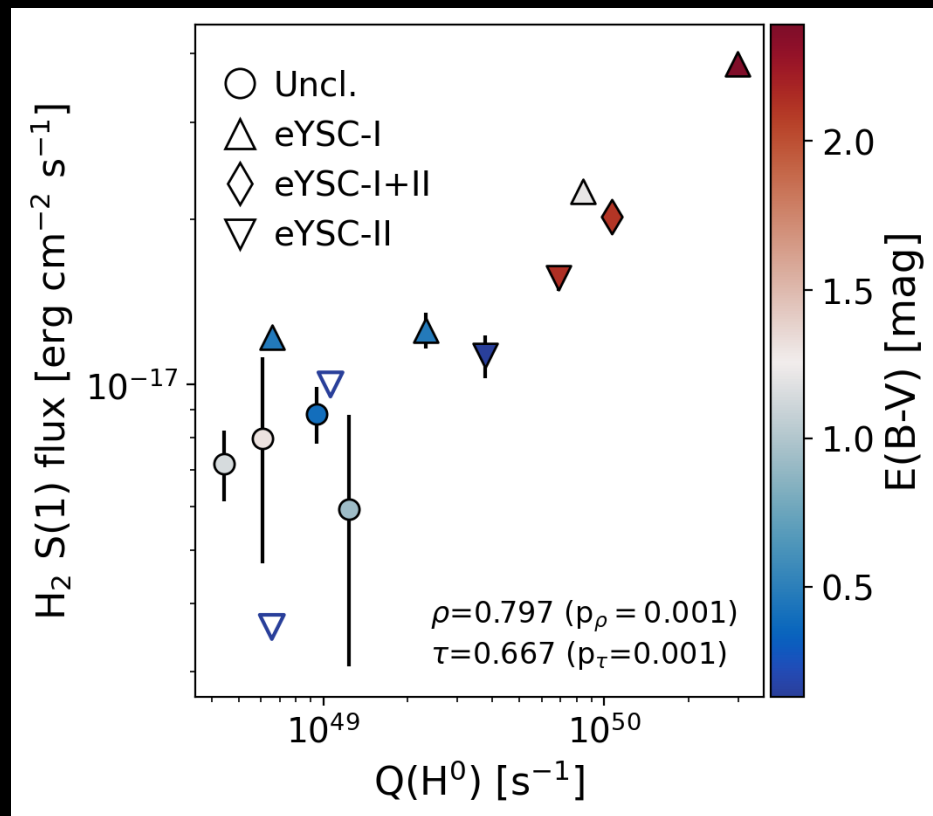
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Spectra of an HII region in the Orion Bar from Peeters+2024

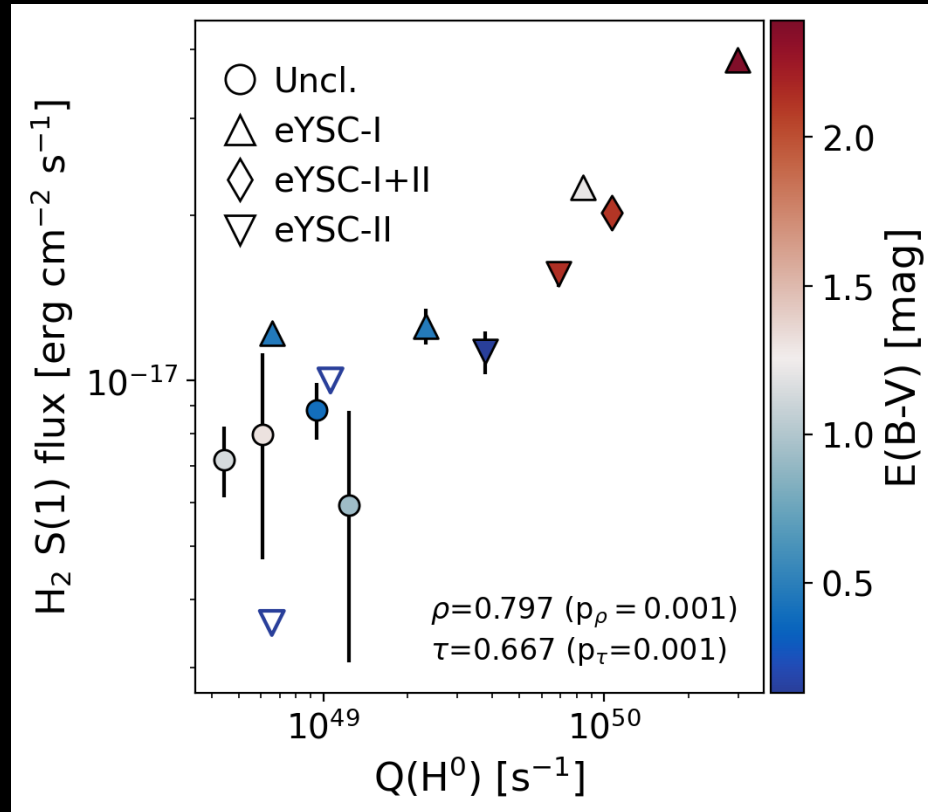
Faustino Vieira+2026

Probing the photodissociation regions



Older ← ————— Younger

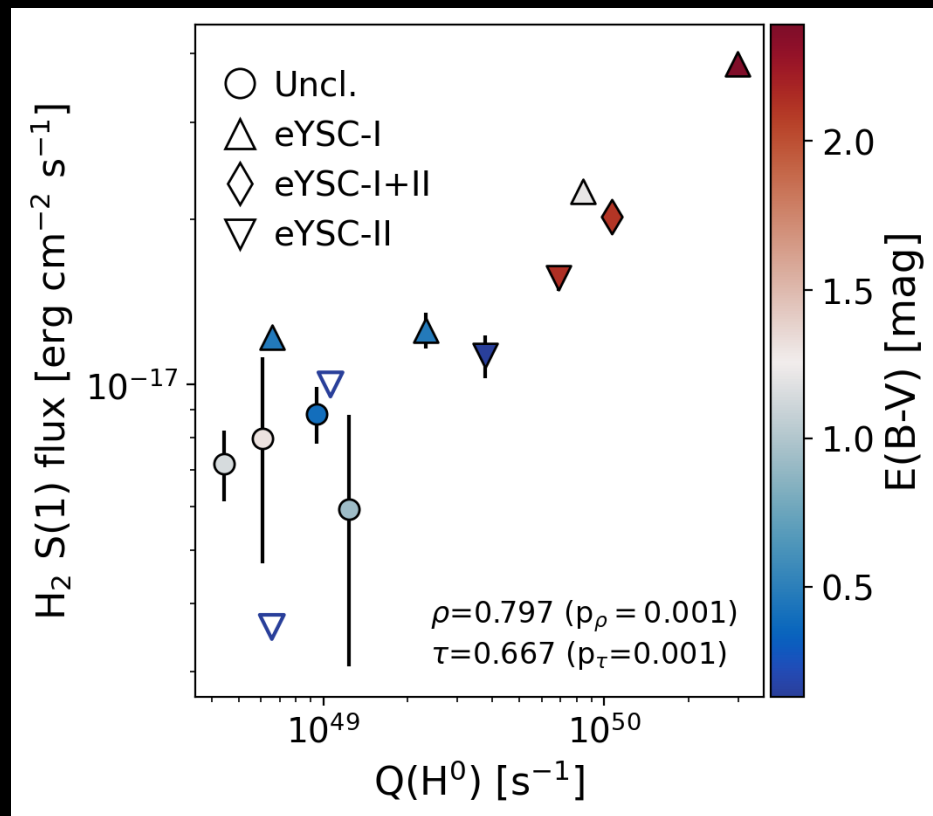
Probing the photodissociation regions



Older ← ————— Younger

As clusters evolve and emerge from their cloud, the PDR also evolves

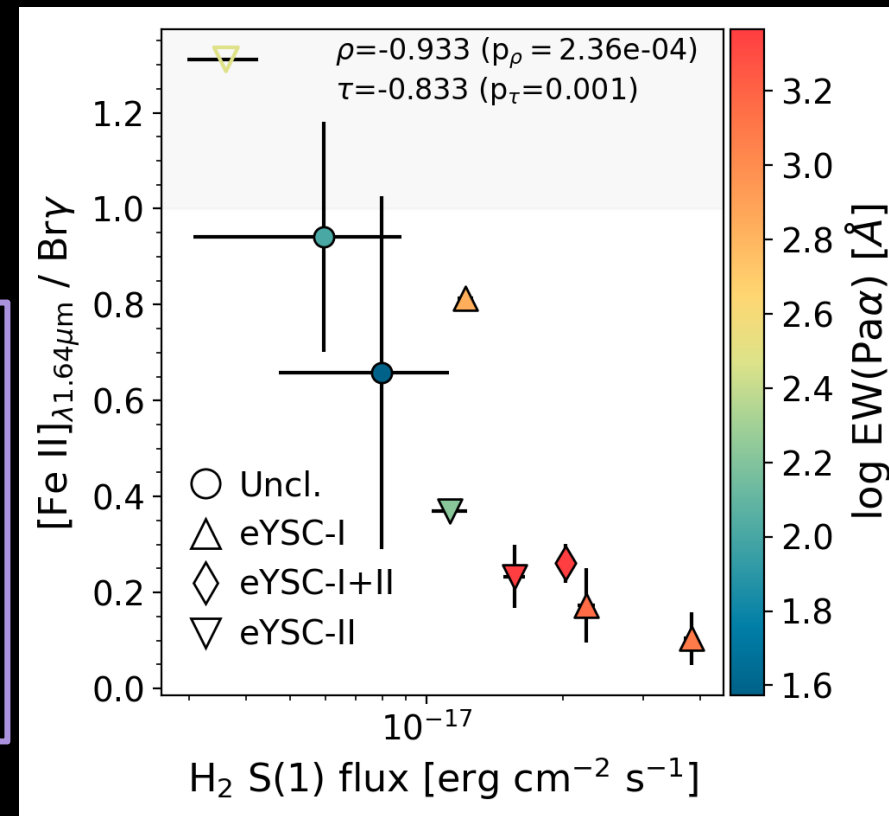
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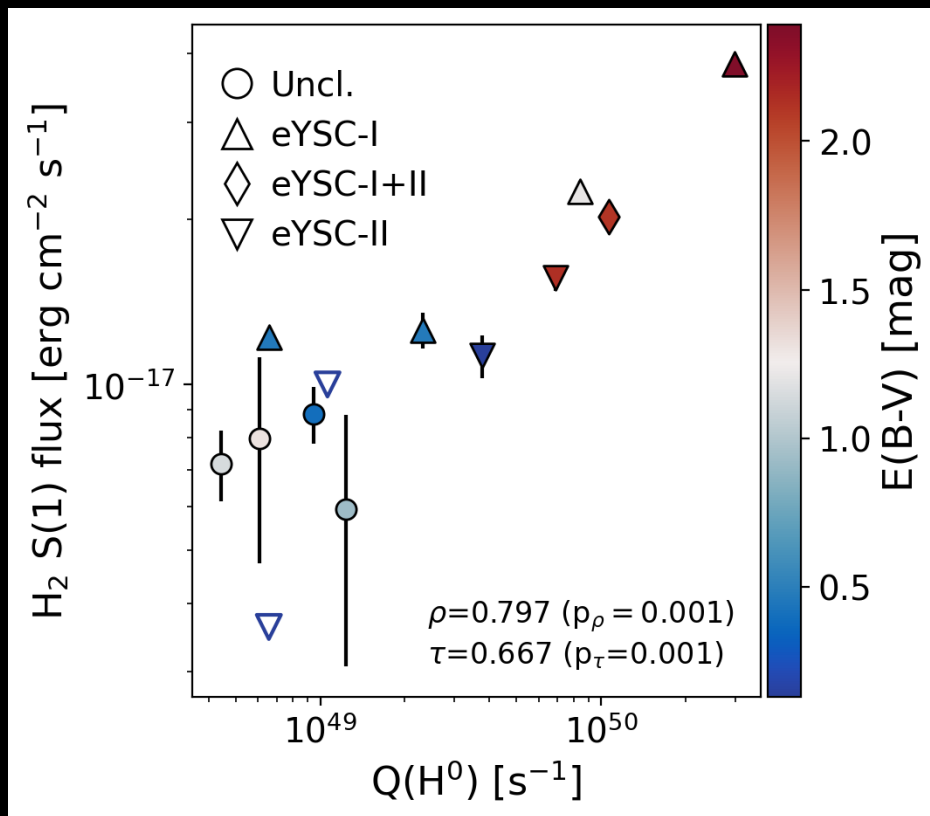
Diagnosing the stellar feedback



Photoionization

Shocks

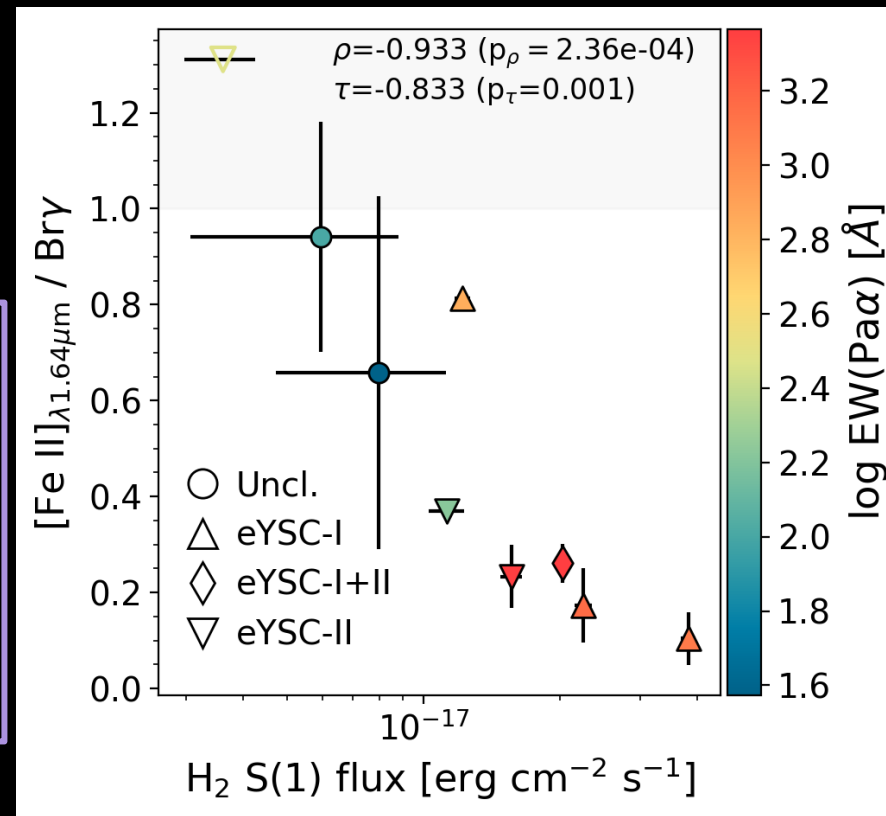
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Diagnosing the stellar feedback

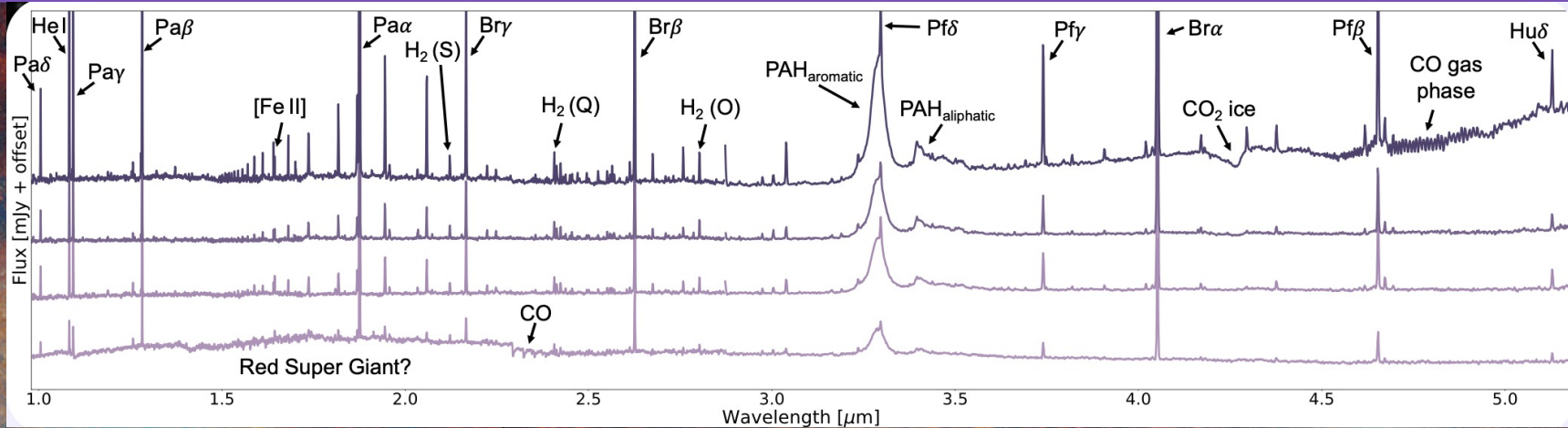


Photoionization

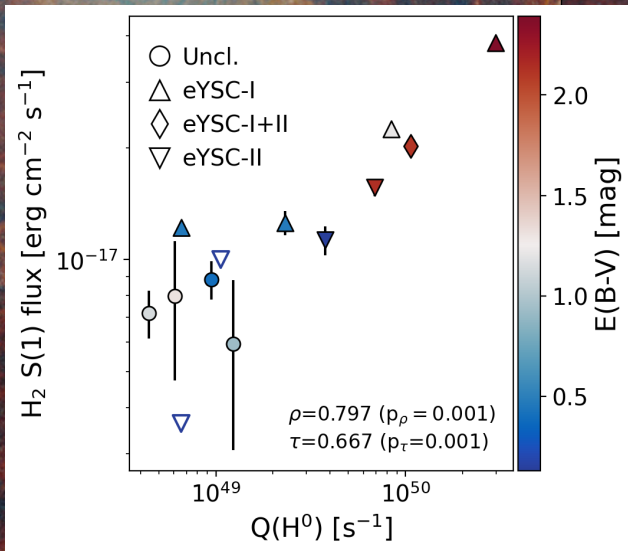
Shocks

Photoionization is likely the dominant stellar feedback mechanism for these eYSCs

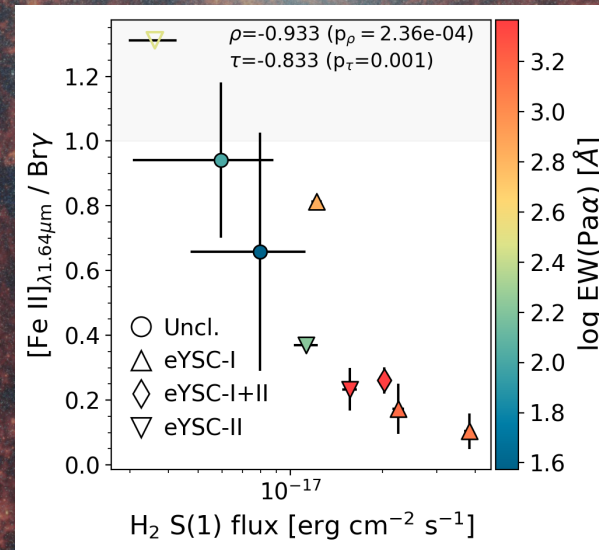
JWST allow us to probe emerging young star cluster at unprecedented detail in nearby galaxies



Scan for paper!



The surrounding ISM is evolving alongside young star clusters, as they emerge from their natal cloud



Pre-supernova feedback is already important in helping eYSCs clear molecular clouds, before the first supernova