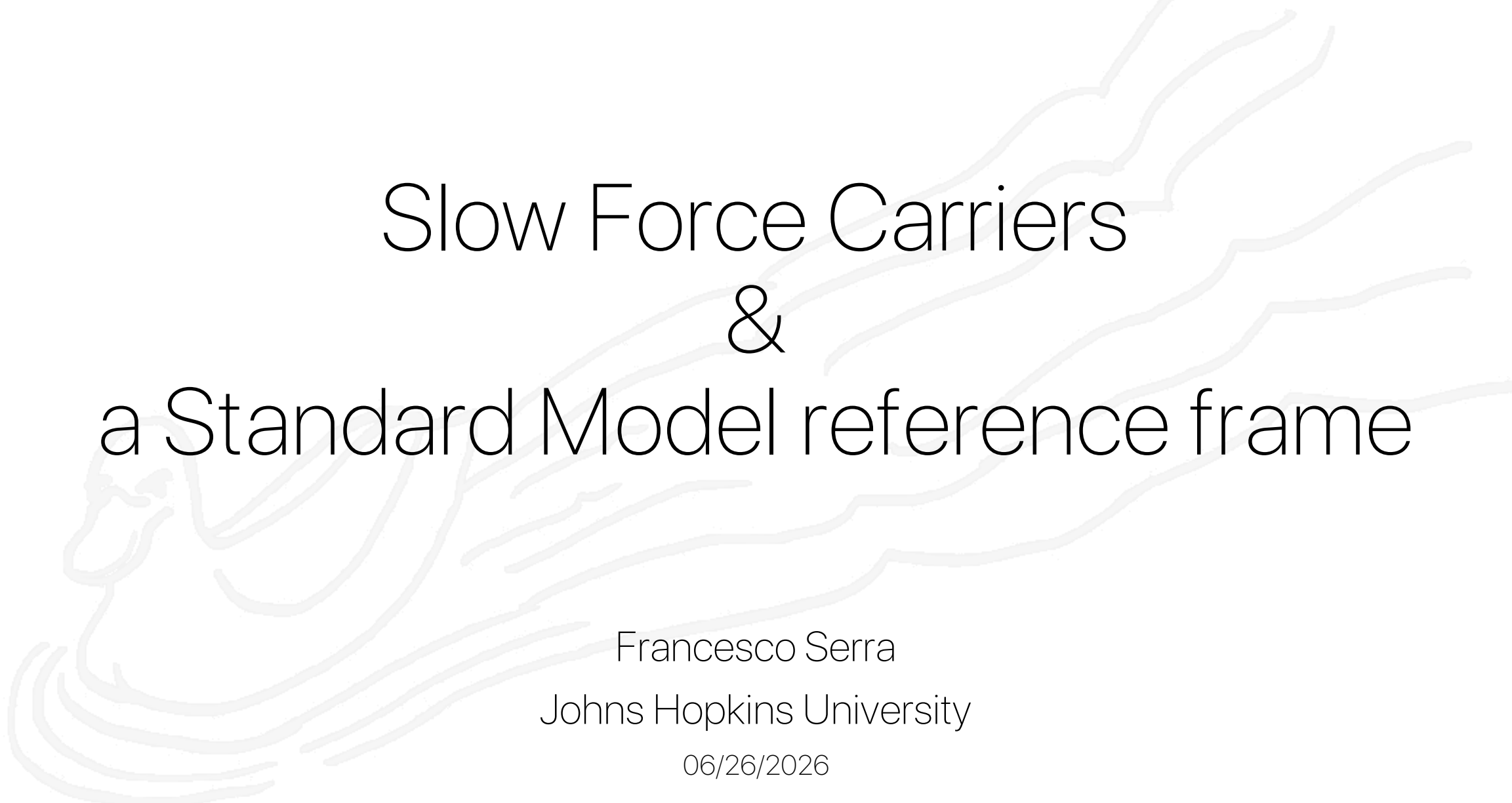




Slow Force Carriers & a Standard Model reference frame

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based on: F. Serra, 2607.xxxxx

In this talk:

A new kind of physics could be hidden within the Standard Model: slow force carriers

New regime of Lorentz breaking: Standard Model + a reference frame that you can hit

Where are the missing IR pieces of nature?

Option 1

Dark physics follows the same structures we have observed so far: new fields within Lorentz and gauge invariant theory.

Option 2

Dark physics indicates a mismatch between nature and symmetry: new exotic excitations with Lorentz-breaking spectrum.
Dark because they break the symmetry.

Breaking of Lorentz symmetry

Gravity:

Reference frame has gravitational properties;
defined by clock field (aether/khronon).

(e.g. Jacobson2004,
Blas2010)

Standard Model:

Historically

Reference frame is a passive, ethereal medium;
background effect: only change dispersion relations.

(e.g. Colladay1998)

This talk

Is there room for a material reference frame
with more active SM properties?

Symmetry point of view

Lorentz symmetry

+ (spin=1, m=0)

therefore

(little group)

gauge symmetry

$$\varepsilon_\mu(q) \rightarrow \varepsilon_\mu(q) + q_\mu$$

Dynamical view of gauge theory

Hamiltonian for gapless spin1 + local conservation law

~~$$\mathcal{L} = -\frac{1}{4}F_{\mu\nu}F^{\mu\nu}$$~~

$$H = \int d^3x \left(\frac{1}{2}\vec{E}^2 + \frac{1}{2}(\vec{\nabla} \wedge \vec{A})^2 \right) \quad A_0 = 0 \quad \{\vec{A}(\vec{x}), \vec{E}(\vec{y})\} = \delta^3(\vec{x} - \vec{y})$$



No dependence on the longitudinal part of A,
The longitudinal mode has speed=0

Local conservation law

$$\frac{d}{dt}\vec{\nabla} \cdot \vec{E} = \{\vec{\nabla} \cdot \vec{E}, H\} = 0$$

Unphysical, gauge transformations

$$\delta_\lambda F = \lambda(\vec{x})\{F, \vec{\nabla} \cdot \vec{E}\} \quad \delta_\lambda \vec{A} = \vec{\nabla} \lambda$$

Slow departures from gauge theory

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Dynamically weak departure from gauge invariance:

- Keep Hamiltonian almost unchanged;
- Add no new fields or fundamental energy scales;



give the longitudinal mode a low speed



$$H = \int d^3x \left(\frac{1}{2} \vec{E}^2 + \frac{1}{2} (\vec{\nabla} \wedge \vec{A})^2 + \frac{c_L^2}{2} (\vec{\nabla} \cdot \vec{A})^2 \right) \quad c_L \ll 1$$

(gauge theory limit: $c_L \rightarrow 0$)

Slow massless photon

$$\frac{d}{dt}\vec{E} = \vec{\nabla} \wedge (\vec{\nabla} \wedge \vec{A}) + c_L^2 \vec{\nabla} (\vec{\nabla} \cdot \vec{A}) \quad \frac{d}{dt}\vec{A} = \vec{E}$$

$$\vec{A} = \vec{A}_T + \vec{A}_L, \quad \vec{E} = \vec{E}_T + \vec{E}_L$$

usual photons

slow longitudinal gapless mode

$$\frac{d^2}{dt^2}\vec{A}_T = \vec{\nabla} \wedge (\vec{\nabla} \wedge \vec{A}_T) \quad \frac{d^2}{dt^2}\vec{A}_L = c_L^2 \vec{\nabla}^2 \vec{A}_L \quad \xrightarrow{c_L \rightarrow 0} \quad \frac{d}{dt}\vec{\nabla} \cdot \vec{E} = 0$$

$$\omega^2 = c_L^2 \vec{k}^2$$

Higher-form Galilean symmetry: mass is protected

$$A_i \rightarrow A_i + c_i + b_{ij}x^j$$

$$\frac{d}{dt} \int d^3x (c_i x^i + b_{ij} x^i x^j) (\vec{\nabla} \cdot \vec{E} - J^0) = 0$$

$$m = 0$$

Decoupling by low speed

Limit of $c_L \rightarrow 0$ gives QED.

gauge-invariant dressing $\sim$ resummation of slow modes

$$\vec{A}_L = \vec{\nabla} \phi \quad \Psi(x) = e^{-ie\phi(x)} \psi(x)$$

$$\bar{\psi} \left(i\gamma^0 \partial_t - i\vec{\gamma} \cdot \vec{\nabla} - e\vec{\gamma} \cdot \vec{A}_L - m \right) \psi = \bar{\Psi} \left(i\gamma^0 \partial_t - \gamma^0 \partial_t \phi - i\vec{\gamma} \cdot \vec{\nabla} - m \right) \Psi$$

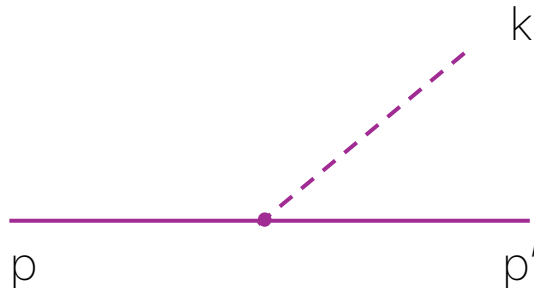
$$\partial_t \phi(\vec{k}, t) = -c_L \hat{k} \cdot \vec{A}_L(\vec{k}, t) \quad \longrightarrow \quad \text{the coupling to gauge invariant fields is speed-suppressed!!}$$

$$\mathcal{L} = \frac{1}{2} \dot{\vec{A}}^2 - \frac{1}{2} (\vec{\nabla} \wedge \vec{A})^2 - \frac{c_L^2}{2} (\vec{\nabla} \cdot \vec{A})^2 + \bar{\Psi} \left(i\gamma^0 \partial_t - i\vec{\gamma} \cdot \vec{\nabla} - e\vec{\gamma} \cdot \vec{A}_T - e\gamma^0 \partial_t \phi - m \right) \Psi$$

Cherenkov Radiation $\longrightarrow$ Hard Recoils

Very slow, gapless mode, couples to matter:

$$\begin{pmatrix} \sqrt{p^2 + m^2} \\ \vec{p} \end{pmatrix} = \begin{pmatrix} \sqrt{p'^2 + m^2} + c_L k \\ \vec{p}' + \vec{k} \end{pmatrix}$$



$$\vec{p}' \simeq \vec{p} - 2p \cos \theta \hat{k}$$



Large momentum transfer!

For a fermion:

$$\Gamma = \frac{1}{2E_p} \int \frac{d^3 p' d^3 k}{(2\pi)^6} \frac{1}{2E_{p'} 2c_L |\vec{k}|} \frac{1}{4} \sum_{spin} |\hat{k} \cdot \vec{J}(p, p')|^2 (2\pi)^4 \delta^4(p - p' - k)$$

$$\Gamma = 4\alpha c_L \left(E_p - \frac{p^2}{3E_p} \right) + O(c_L^2)$$



Rare event!

A new regime: SM + reference frame

What happens if, e.g., $c_L \lesssim 10^{-40}$?

- Modes do not propagate
- SM interactions look perfect
- Very rare, but visible recoils!



Slow modes become momentum stored in the rest frame.



The reference frame becomes a material that you can hit!

Experimental bounds

$g - 2$	$c_L \lesssim 10^{-7}$
Cosmic rays power loss	$c_L \lesssim 10^{-23}$
Recoils: Collider beams	$c_L \lesssim 10^{-29}$
Recoils: Pulsar spin-down rate	$c_L \lesssim 10^{-40}$
Recoils: Earth heat balance	$c_L \lesssim 10^{-45}$
Recoils: cosmic ray nuclei	$c_L \lesssim 10^{-48}$
Recoils: DM detection (Xe)	$c_L \lesssim 10^{-60}$

→ Ordinary observables & effects

→ Rare, Hard recoils against the reference frame!

Coming next: Cosmic Recoil Background?

So...

slow force carriers



reference frame

New IR physics within gauge theory, new effects, new ways to test SM:

New connection between gauge redundancy and soft modes/IR divergences

a whole quantum reference frame!

opportunities for new physics, especially in gravity and cosmology

is relativity fundamental?