

Magnetogenesis via plasma processes

Young Dae Yoon

Asia Pacific Center for Theoretical Physics

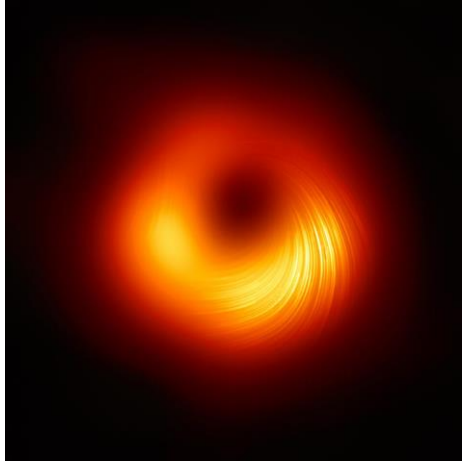
Department of Physics, Pohang University of Science and Technology

2nd APCTP-INPP Demokritos Workshop

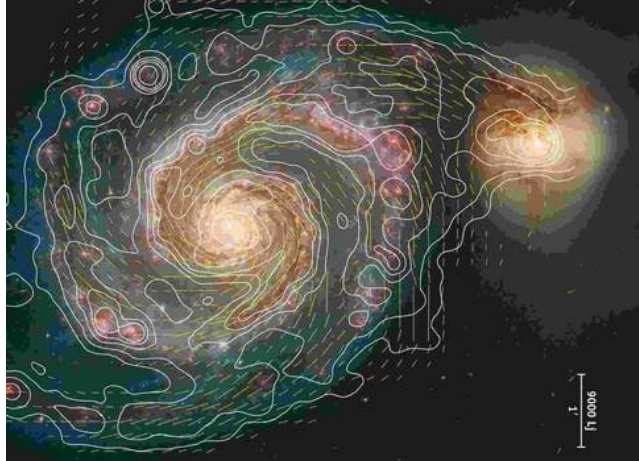
08/26/25

Magnetic fields in the Universe

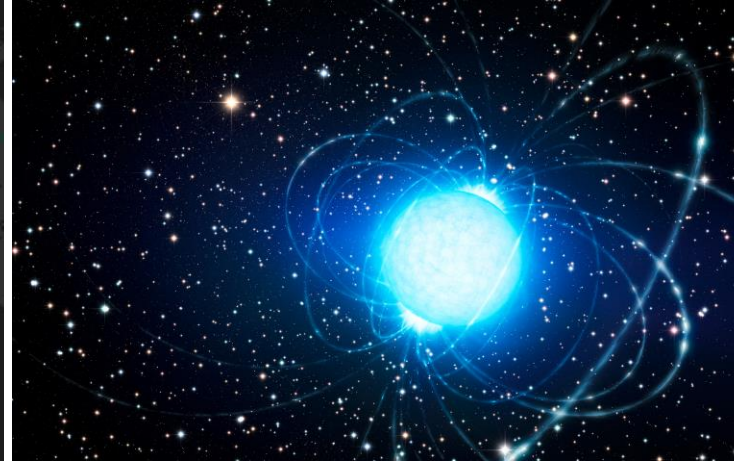
- Magnetic fields are everywhere in the Universe at all scales and strength



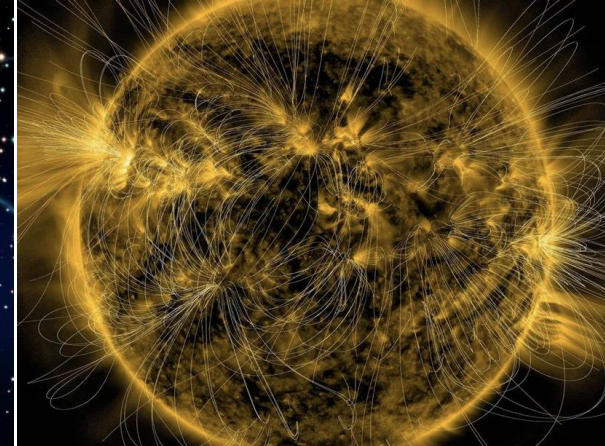
Black Holes



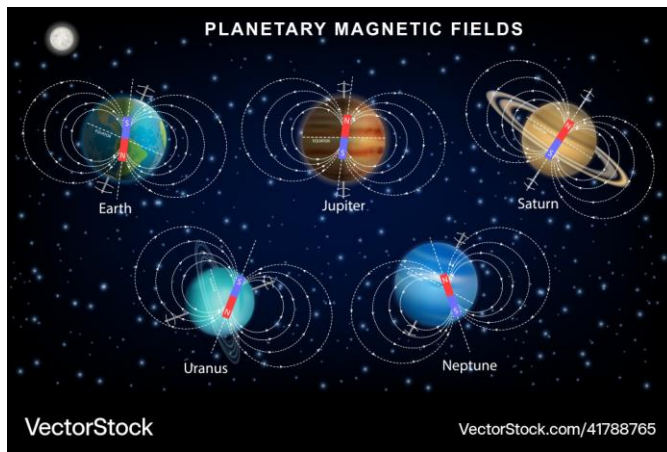
Galaxies



Neutron Stars



Stars



Planets

μG --- Interstellar Medium, Galaxies
 G --- Black Holes, Stars, Planets
 10^{10}G --- Neutron Star (Magnetars)

Standard cosmology: no magnetic fields were created at the big bang

How are cosmic/astrophysical magnetic fields generated?

- Magnetohydrodynamics (MHD) is the largest-scale description of magnetized plasmas

- One important equation:

$$\mathbf{E} + \mathbf{U} \times \mathbf{B} = 0$$

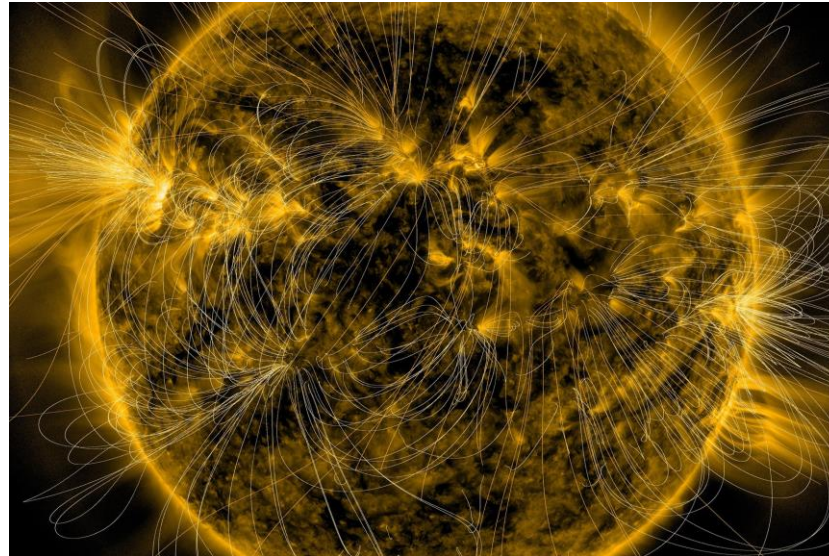
Ideal Ohm's law

- Take curl and use Faraday's law

$$\frac{\partial \mathbf{B}}{\partial t} = \nabla \times (\mathbf{U} \times \mathbf{B})$$

=

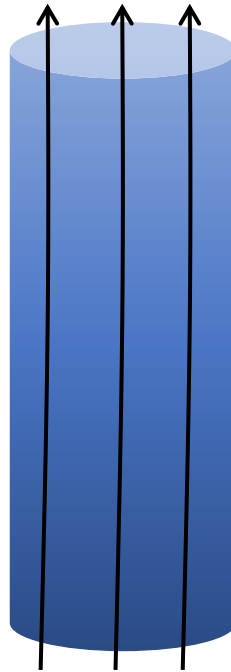
MHD induction equation



- So what is the conserved quantity in MHD?
- Magnetic Flux!

$$\frac{\partial \mathbf{B}}{\partial t} = \nabla \times (\mathbf{U} \times \mathbf{B}) \Rightarrow \frac{d}{dt} \int \mathbf{B} \cdot d\mathbf{S} = 0$$

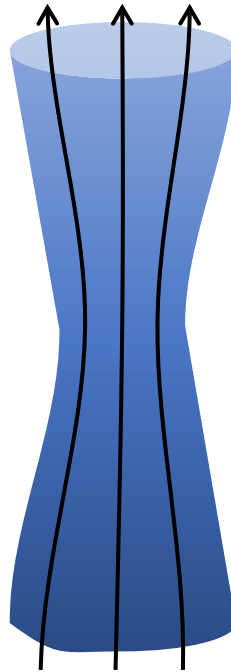
- A plasma wants to keep its magnetic flux \Leftrightarrow Magnetic fields are “frozen into” the plasma



- So what is the conserved quantity in MHD?
- Magnetic Flux!

$$\frac{\partial \mathbf{B}}{\partial t} = \nabla \times (\mathbf{U} \times \mathbf{B}) \Rightarrow \frac{d}{dt} \int \mathbf{B} \cdot d\mathbf{S} = 0$$

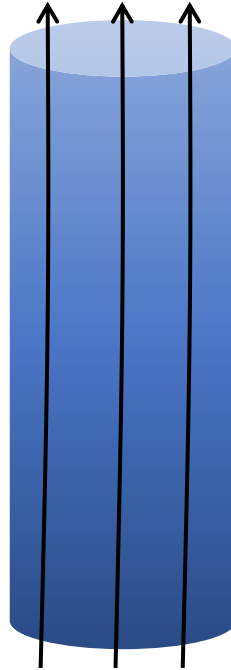
- A plasma wants to keep its magnetic flux \Leftrightarrow Magnetic fields are “frozen into” the plasma



- So what is the conserved quantity in MHD?
- Magnetic Flux!

$$\frac{\partial \mathbf{B}}{\partial t} = \nabla \times (\mathbf{U} \times \mathbf{B}) \Rightarrow \frac{d}{dt} \int \mathbf{B} \cdot d\mathbf{S} = 0$$

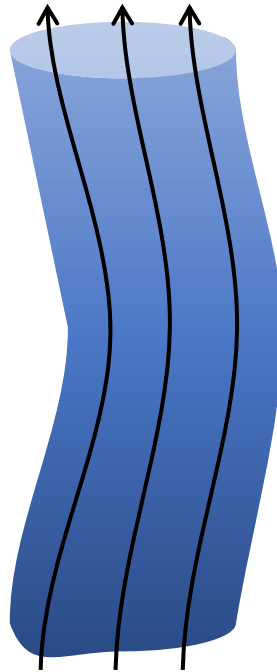
- A plasma wants to keep its magnetic flux \Leftrightarrow Magnetic fields are “frozen into” the plasma



- So what is the conserved quantity in MHD?
- Magnetic Flux!

$$\frac{\partial \mathbf{B}}{\partial t} = \nabla \times (\mathbf{U} \times \mathbf{B}) \Rightarrow \frac{d}{dt} \int \mathbf{B} \cdot d\mathbf{S} = 0$$

- A plasma wants to keep its magnetic flux \Leftrightarrow Magnetic fields are “frozen into” the plasma



- Plasma induction equation is the underlying equation

$$\frac{\partial \mathbf{B}}{\partial t} = \nabla \times (\mathbf{U} \times \mathbf{B})$$

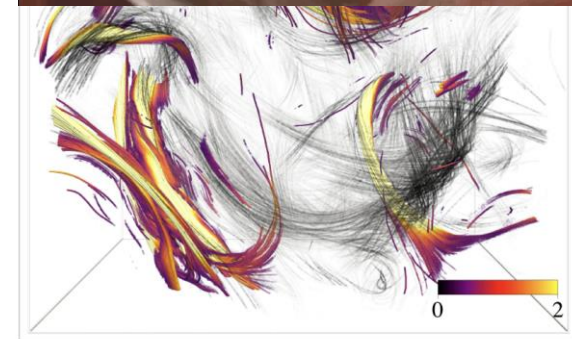
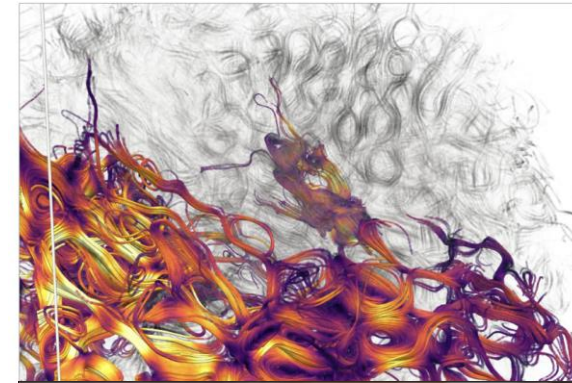
- If $\mathbf{B} = 0$ at $t = 0$,

$$\frac{\partial \mathbf{B}}{\partial t} = 0$$

- We need a source for the “seed” magnetic field!

- Has to be due to a non-conservative electric field

$$\frac{\partial \mathbf{B}}{\partial t} = -\nabla \times \mathbf{E}$$



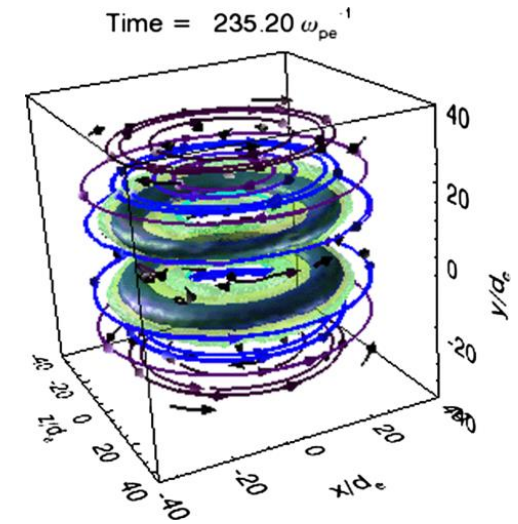
Zhou+, ApJ (2023)

Seed Magnetogenesis

- Two popular mechanisms for seed magnetogenesis
- Biermann battery [Biermann, 1950]
 - Misalignment of density and pressure

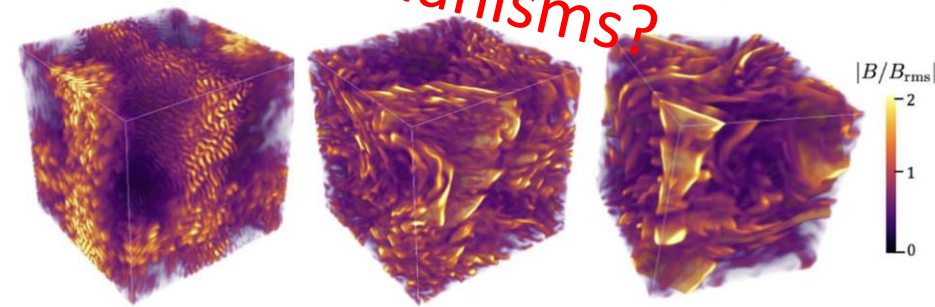
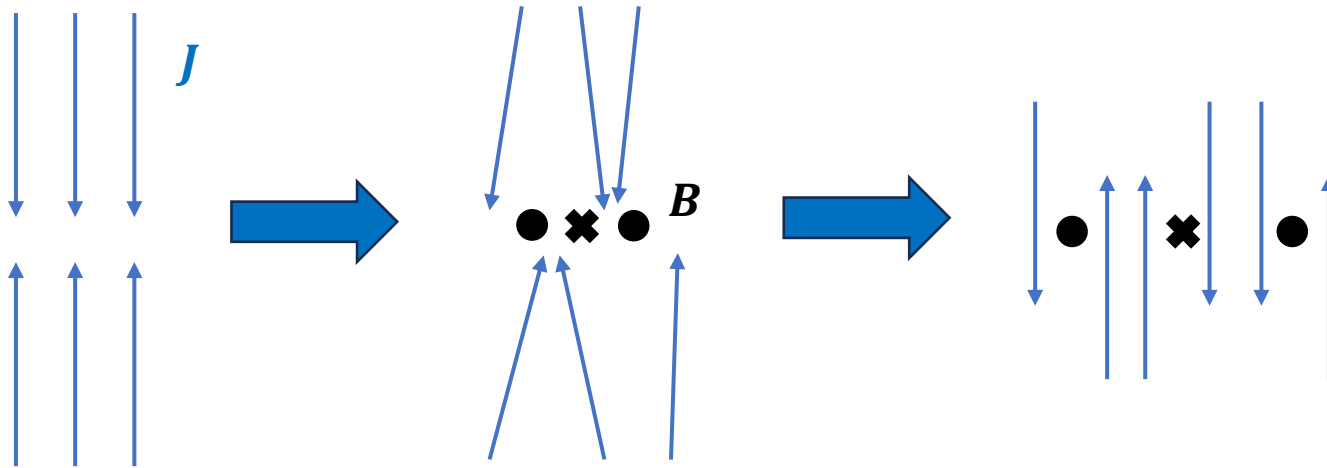
Can we generalize these mechanisms and find new mechanisms?

$$\frac{\partial \mathbf{B}}{\partial t} = -\frac{\nabla n \times \nabla p}{en^2}$$



[Schoeffler+, PRL (2014)]

- Weibel Instability [Weibel, 1959]
 - Anisotropic pressure or, equivalently, counter-propagating beams



[Zhou+, ApJ (2023)]

$$m_\sigma \frac{d\mathbf{u}_\sigma}{dt} = q_\sigma (\mathbf{E} + \mathbf{u}_\sigma \times \mathbf{B}) - \frac{\nabla \cdot \overleftrightarrow{\mathbf{p}}_\sigma}{n_\sigma}$$

$$m_{\sigma} \frac{d\mathbf{u}_{\sigma}}{dt} = q_{\sigma} (\mathbf{E} + \mathbf{u}_{\sigma} \times \mathbf{B}) - \frac{\nabla \cdot \overleftrightarrow{\mathbf{p}}_{\sigma}}{n_{\sigma}}$$

Acceleration of plasma species σ

$$m_\sigma \frac{d\mathbf{u}_\sigma}{dt} = q_\sigma (\mathbf{E} + \mathbf{u}_\sigma \times \mathbf{B}) - \frac{\nabla \cdot \overleftrightarrow{\mathbf{p}}_\sigma}{n_\sigma}$$

Lorentz force

$$m_\sigma \frac{d\mathbf{u}_\sigma}{dt} = q_\sigma (\mathbf{E} + \mathbf{u}_\sigma \times \mathbf{B}) - \frac{\nabla \cdot \overleftrightarrow{\mathbf{p}}_\sigma}{n_\sigma}$$

Pressure gradient

$$m \frac{d\mathbf{u}}{dt} = q (\mathbf{E} + \mathbf{u} \times \mathbf{B}) - \frac{\nabla \cdot \overleftrightarrow{\mathbf{p}}}{n}$$

$$m \frac{d\mathbf{u}}{dt} = q (\mathbf{E} + \mathbf{u} \times \mathbf{B}) - \frac{\nabla \cdot \overleftrightarrow{\mathbf{p}}}{n}$$

$$m \left(\frac{\partial \mathbf{u}}{\partial t} + \mathbf{u} \cdot \nabla \mathbf{u} \right) = q (\mathbf{E} + \mathbf{u} \times \mathbf{B}) - \frac{\nabla \cdot \overleftrightarrow{\mathbf{p}}}{n}$$

$$m \frac{d\mathbf{u}}{dt} = q (\mathbf{E} + \mathbf{u} \times \mathbf{B}) - \frac{\nabla \cdot \overleftrightarrow{\mathbf{p}}}{n}$$

$$m \left(\frac{\partial \mathbf{u}}{\partial t} + \mathbf{u} \cdot \nabla \mathbf{u} \right) = q (\mathbf{E} + \mathbf{u} \times \mathbf{B}) - \frac{\nabla \cdot \overleftrightarrow{\mathbf{p}}}{n}$$

$$m \nabla \times \left(\frac{\partial \mathbf{u}}{\partial t} + \mathbf{u} \cdot \nabla \mathbf{u} \right) = q \nabla \times (\mathbf{E} + \mathbf{u} \times \mathbf{B}) - \nabla \times \left(\frac{\nabla \cdot \overleftrightarrow{\mathbf{p}}}{n} \right)$$

$$m\nabla \times \left(\frac{\partial \mathbf{u}}{\partial t} + \mathbf{u} \cdot \nabla \mathbf{u} \right) = q\nabla \times (\mathbf{E} + \mathbf{u} \times \mathbf{B}) - \nabla \times \left(\frac{\nabla \cdot \overleftrightarrow{\mathbf{p}}}{n} \right)$$

Vorticity $\boldsymbol{\Omega} = \nabla \times \mathbf{u}$

$$m \nabla \times \left(\frac{\partial \mathbf{u}}{\partial t} + \mathbf{u} \cdot \nabla \mathbf{u} \right) = q \nabla \times (\mathbf{E} + \mathbf{u} \times \mathbf{B}) - \nabla \times \left(\frac{\nabla \cdot \overleftrightarrow{\mathbf{p}}}{n} \right)$$

Vorticity $\boldsymbol{\Omega} = \nabla \times \mathbf{u}$

$$m \left(\frac{\partial \boldsymbol{\Omega}}{\partial t} \right)$$

$$m \nabla \times \left(\frac{\partial \mathbf{u}}{\partial t} + \mathbf{u} \cdot \nabla \mathbf{u} \right) = q \nabla \times (\mathbf{E} + \mathbf{u} \times \mathbf{B}) - \nabla \times \left(\frac{\nabla \cdot \overleftrightarrow{\mathbf{p}}}{n} \right)$$

Vector Calc Identity $\nabla \times [\mathbf{u} \cdot \nabla \mathbf{u}] = -\nabla \times [\mathbf{u} \times \nabla \times \mathbf{u}]$

$$m \left(\frac{\partial \Omega}{\partial t} \right)$$

$$m \nabla \times \left(\frac{\partial \mathbf{u}}{\partial t} + \mathbf{u} \cdot \nabla \mathbf{u} \right) = q \nabla \times (\mathbf{E} + \mathbf{u} \times \mathbf{B}) - \nabla \times \left(\frac{\nabla \cdot \overleftrightarrow{\mathbf{p}}}{n} \right)$$

Vector Calc Identity $\nabla \times [\mathbf{u} \cdot \nabla \mathbf{u}] = -\nabla \times [\mathbf{u} \times \boldsymbol{\Omega}]$

$$m \left(\frac{\partial \boldsymbol{\Omega}}{\partial t} \right)$$

$$m \nabla \times \left(\frac{\partial \mathbf{u}}{\partial t} + \mathbf{u} \cdot \nabla \mathbf{u} \right) = q \nabla \times (\mathbf{E} + \mathbf{u} \times \mathbf{B}) - \nabla \times \left(\frac{\nabla \cdot \overleftrightarrow{\mathbf{p}}}{n} \right)$$

Vector Calc Identity $\nabla \times [\mathbf{u} \cdot \nabla \mathbf{u}] = -\nabla \times [\mathbf{u} \times \boldsymbol{\Omega}]$

$$m \left(\frac{\partial \boldsymbol{\Omega}}{\partial t} - \nabla \times [\mathbf{u} \times \boldsymbol{\Omega}] \right)$$

$$m \nabla \times \left(\frac{\partial \mathbf{u}}{\partial t} + \mathbf{u} \cdot \nabla \mathbf{u} \right) = q \nabla \times (\mathbf{E} + \mathbf{u} \times \mathbf{B}) - \nabla \times \left(\frac{\nabla \cdot \overleftrightarrow{\mathbf{p}}}{n} \right)$$

Faraday's law $\nabla \times \mathbf{E} = -\frac{\partial \mathbf{B}}{\partial t}$

$$m \left(\frac{\partial \boldsymbol{\Omega}}{\partial t} - \nabla \times [\mathbf{u} \times \boldsymbol{\Omega}] \right)$$

$$m \nabla \times \left(\frac{\partial \mathbf{u}}{\partial t} + \mathbf{u} \cdot \nabla \mathbf{u} \right) = q \nabla \times (\mathbf{E} + \mathbf{u} \times \mathbf{B}) - \nabla \times \left(\frac{\nabla \cdot \overleftrightarrow{\mathbf{p}}}{n} \right)$$

Faraday's law $\nabla \times \mathbf{E} = -\frac{\partial \mathbf{B}}{\partial t}$

$$m \left(\frac{\partial \boldsymbol{\Omega}}{\partial t} - \nabla \times [\mathbf{u} \times \boldsymbol{\Omega}] \right) = q \left(-\frac{\partial \mathbf{B}}{\partial t} \right)$$

$$m\nabla \times \left(\frac{\partial \mathbf{u}}{\partial t} + \mathbf{u} \cdot \nabla \mathbf{u} \right) = q\nabla \times (\mathbf{E} + \mathbf{u} \times \mathbf{B}) - \nabla \times \left(\frac{\nabla \cdot \overleftrightarrow{\mathbf{p}}}{n} \right)$$

$$m \left(\frac{\partial \boldsymbol{\Omega}}{\partial t} - \nabla \times [\mathbf{u} \times \boldsymbol{\Omega}] \right) = q \left(-\frac{\partial \mathbf{B}}{\partial t} + \nabla \times [\mathbf{u} \times \mathbf{B}] \right)$$

$$m\nabla \times \left(\frac{\partial \mathbf{u}}{\partial t} + \mathbf{u} \cdot \nabla \mathbf{u} \right) = q\nabla \times (\mathbf{E} + \mathbf{u} \times \mathbf{B}) - \nabla \times \left(\frac{\nabla \cdot \overleftrightarrow{\mathbf{p}}}{n} \right)$$

$$m \left(\frac{\partial \boldsymbol{\Omega}}{\partial t} - \nabla \times [\mathbf{u} \times \boldsymbol{\Omega}] \right) = q \left(-\frac{\partial \mathbf{B}}{\partial t} + \nabla \times [\mathbf{u} \times \mathbf{B}] \right) - \nabla \times \left(\frac{\nabla \cdot \overleftrightarrow{\mathbf{p}}}{n} \right)$$

$$m \nabla \times \left(\frac{\partial \mathbf{u}}{\partial t} + \mathbf{u} \cdot \nabla \mathbf{u} \right) = q \nabla \times (\mathbf{E} + \mathbf{u} \times \mathbf{B}) - \nabla \times \left(\frac{\nabla \cdot \overleftrightarrow{\mathbf{p}}}{n} \right)$$

$$m \left(\frac{\partial \boldsymbol{\Omega}}{\partial t} - \nabla \times [\mathbf{u} \times \boldsymbol{\Omega}] \right) = q \left(-\frac{\partial \mathbf{B}}{\partial t} + \nabla \times [\mathbf{u} \times \mathbf{B}] \right) - \nabla \times \left(\frac{\nabla \cdot \overleftrightarrow{\mathbf{p}}}{n} \right)$$

$$m \nabla \times \left(\frac{\partial \mathbf{u}}{\partial t} + \mathbf{u} \cdot \nabla \mathbf{u} \right) = q \nabla \times (\mathbf{E} + \mathbf{u} \times \mathbf{B}) - \nabla \times \left(\frac{\nabla \cdot \overleftrightarrow{\mathbf{p}}}{n} \right)$$

$$m \left(\frac{\partial \boldsymbol{\Omega}}{\partial t} - \nabla \times [\mathbf{u} \times \boldsymbol{\Omega}] \right) = q \left(-\frac{\partial \mathbf{B}}{\partial t} + \nabla \times [\mathbf{u} \times \mathbf{B}] \right) - \nabla \times \left(\frac{\nabla \cdot \overleftrightarrow{\mathbf{p}}}{n} \right)$$

$$\frac{\partial}{\partial t} (m \boldsymbol{\Omega} + q \mathbf{B})$$

$$m \nabla \times \left(\frac{\partial \mathbf{u}}{\partial t} + \mathbf{u} \cdot \nabla \mathbf{u} \right) = q \nabla \times (\mathbf{E} + \mathbf{u} \times \mathbf{B}) - \nabla \times \left(\frac{\nabla \cdot \overleftrightarrow{\mathbf{p}}}{n} \right)$$

$$m \left(\frac{\partial \boldsymbol{\Omega}}{\partial t} - \nabla \times [\mathbf{u} \times \boldsymbol{\Omega}] \right) = q \left(-\frac{\partial \mathbf{B}}{\partial t} + \nabla \times [\mathbf{u} \times \mathbf{B}] \right) - \nabla \times \left(\frac{\nabla \cdot \overleftrightarrow{\mathbf{p}}}{n} \right)$$

$$\frac{\partial}{\partial t} (m \boldsymbol{\Omega} + q \mathbf{B})$$

$$m \nabla \times \left(\frac{\partial \mathbf{u}}{\partial t} + \mathbf{u} \cdot \nabla \mathbf{u} \right) = q \nabla \times (\mathbf{E} + \mathbf{u} \times \mathbf{B}) - \nabla \times \left(\frac{\nabla \cdot \overleftrightarrow{\mathbf{p}}}{n} \right)$$

$$m \left(\frac{\partial \boldsymbol{\Omega}}{\partial t} - \nabla \times [\mathbf{u} \times \boldsymbol{\Omega}] \right) = q \left(-\frac{\partial \mathbf{B}}{\partial t} + \nabla \times [\mathbf{u} \times \mathbf{B}] \right) - \nabla \times \left(\frac{\nabla \cdot \overleftrightarrow{\mathbf{p}}}{n} \right)$$

$$\frac{\partial}{\partial t} (m \boldsymbol{\Omega} + q \mathbf{B}) = \nabla \times [\mathbf{u} \times (m \boldsymbol{\Omega} + q \mathbf{B})]$$

$$m\nabla \times \left(\frac{\partial \mathbf{u}}{\partial t} + \mathbf{u} \cdot \nabla \mathbf{u} \right) = q\nabla \times (\mathbf{E} + \mathbf{u} \times \mathbf{B}) - \nabla \times \left(\frac{\nabla \cdot \overleftrightarrow{\mathbf{p}}}{n} \right)$$

$$m \left(\frac{\partial \boldsymbol{\Omega}}{\partial t} - \nabla \times [\mathbf{u} \times \boldsymbol{\Omega}] \right) = q \left(-\frac{\partial \mathbf{B}}{\partial t} + \nabla \times [\mathbf{u} \times \mathbf{B}] \right) - \nabla \times \left(\frac{\nabla \cdot \overleftrightarrow{\mathbf{p}}}{n} \right)$$

$$\frac{\partial}{\partial t} (m\boldsymbol{\Omega} + q\mathbf{B}) = \nabla \times [\mathbf{u} \times (m\boldsymbol{\Omega} + q\mathbf{B})] - \nabla \times \left(\frac{\nabla \cdot \overleftrightarrow{\mathbf{p}}}{n} \right)$$

$$\frac{\partial}{\partial t} (m\mathbf{\Omega} + q\mathbf{B}) = \nabla \times [\mathbf{u} \times (m\mathbf{\Omega} + q\mathbf{B})] - \nabla \times \left(\frac{\nabla \cdot \overleftrightarrow{\mathbf{p}}}{n} \right)$$

$$\frac{\partial}{\partial t} (m\boldsymbol{\Omega} + q\mathbf{B}) = \nabla \times [\mathbf{u} \times (m\boldsymbol{\Omega} + q\mathbf{B})] - \nabla \times \left(\frac{\nabla \cdot \overleftrightarrow{\mathbf{p}}}{n} \right)$$

$$\frac{\partial}{\partial t} (m\boldsymbol{\Omega} + q\mathbf{B}) = \nabla \times [\mathbf{u} \times (m\boldsymbol{\Omega} + q\mathbf{B})] - \nabla \times \left(\frac{\nabla \cdot \overleftrightarrow{\mathbf{p}}}{n} \right)$$

Canonical Vorticity $\mathbf{Q} = m\boldsymbol{\Omega} + q\mathbf{B}$

$$\frac{\partial}{\partial t} (m\boldsymbol{\Omega} + q\mathbf{B}) = \nabla \times [\mathbf{u} \times (m\boldsymbol{\Omega} + q\mathbf{B})] - \nabla \times \left(\frac{\nabla \cdot \overleftrightarrow{\mathbf{p}}}{n} \right)$$

Canonical Vorticity $\mathbf{Q} = m\boldsymbol{\Omega} + q\mathbf{B}$

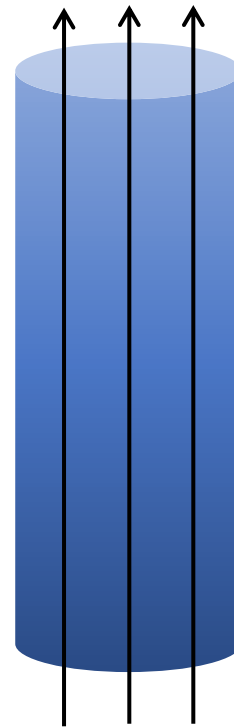
$$\frac{\partial \mathbf{Q}}{\partial t} = \nabla \times (\mathbf{u} \times \mathbf{Q}) - \nabla \times \left(\frac{\nabla \cdot \overleftrightarrow{\mathbf{p}}}{n} \right)$$

- Canonical induction equation

$$\frac{\partial \mathbf{Q}}{\partial t} = \nabla \times (\mathbf{u} \times \mathbf{Q}) - \nabla \times \left(\frac{\nabla \cdot \vec{\mathbf{p}}}{n} \right)$$

$\mathbf{Q} = 0$

Sole term responsible for magnetogenesis!



Vorticity-free
Magnetic-field-free

Pressure term

- Does our pressure term generalize both mechanisms?

$$\frac{\partial \mathbf{Q}}{\partial t} = -\nabla \times \left(\frac{\nabla \cdot \vec{\mathbf{p}}}{n} \right)$$

- Assume isotropic pressure, i.e., $\vec{\mathbf{p}} = p\vec{\mathbf{I}}$

Biermann Battery!

$$-\nabla \times \frac{\nabla \cdot \vec{\mathbf{p}}}{n} = -\nabla \times \frac{\nabla p}{n} = \frac{\nabla n \times \nabla p}{n^2}$$

- So our pressure term generalizes Biermann battery ✓
- Can it generalize Weibel instability as well?
- Assume 1D, i.e., $\nabla = \hat{x}\partial/\partial x$

$$-\left(\nabla \times \frac{\nabla \cdot \vec{\mathbf{p}}}{n} \right) \cdot \hat{z} = -\frac{1}{n} \frac{\partial^2 p_{xy}}{\partial x^2}$$

- Mixing between p_{xx} and p_{yy}

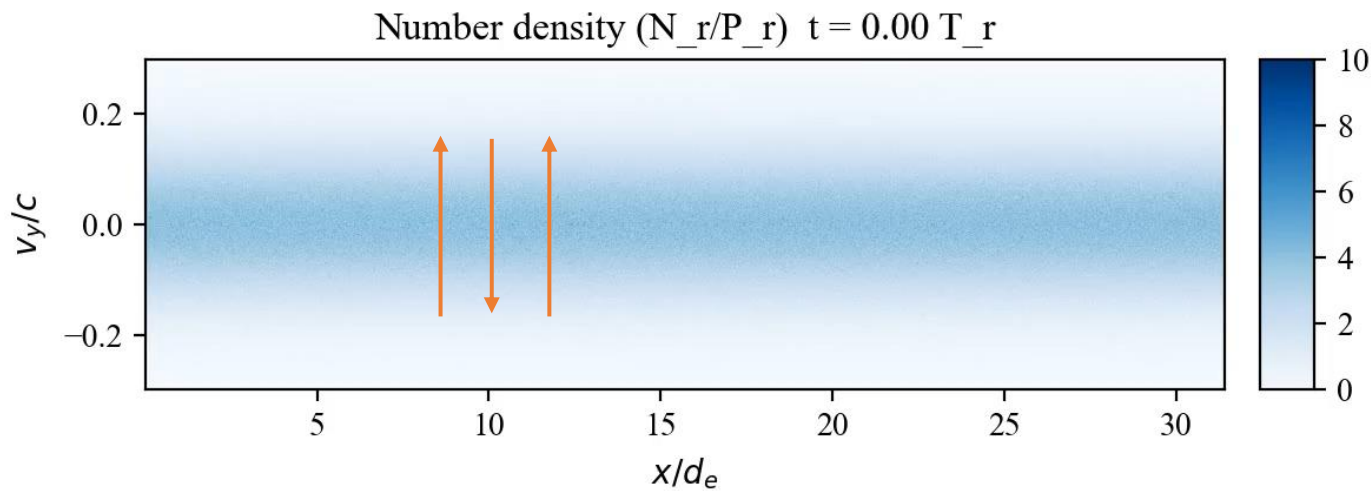
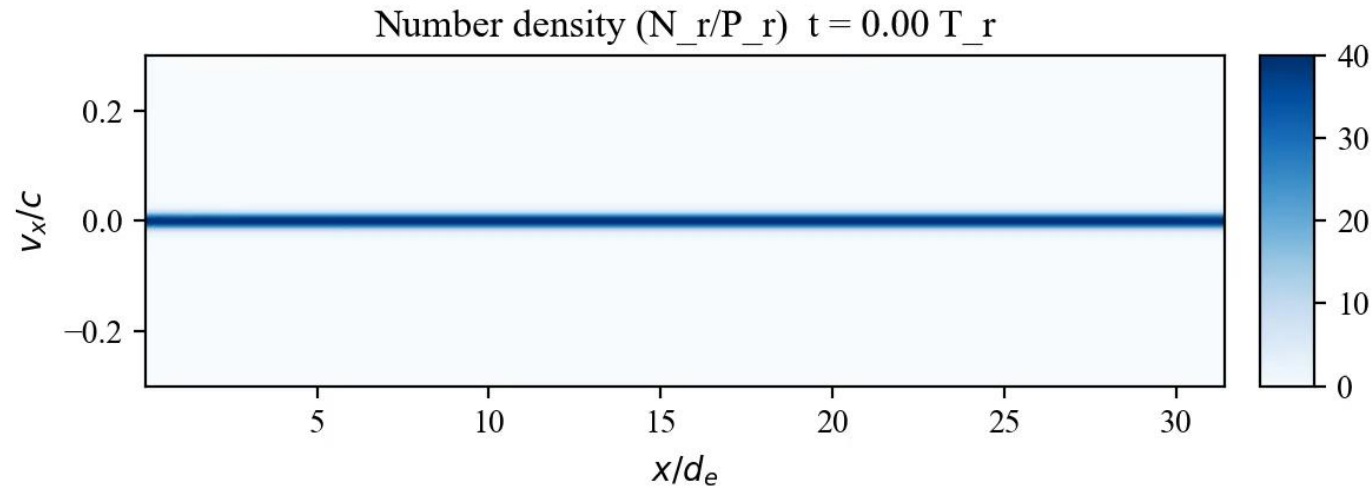
Diagonal Off-diagonal

$$\begin{pmatrix} p_{xx} & p_{xy} & p_{xz} \\ p_{xy} & p_{yy} & p_{yz} \\ p_{xz} & p_{yz} & p_{zz} \end{pmatrix}$$

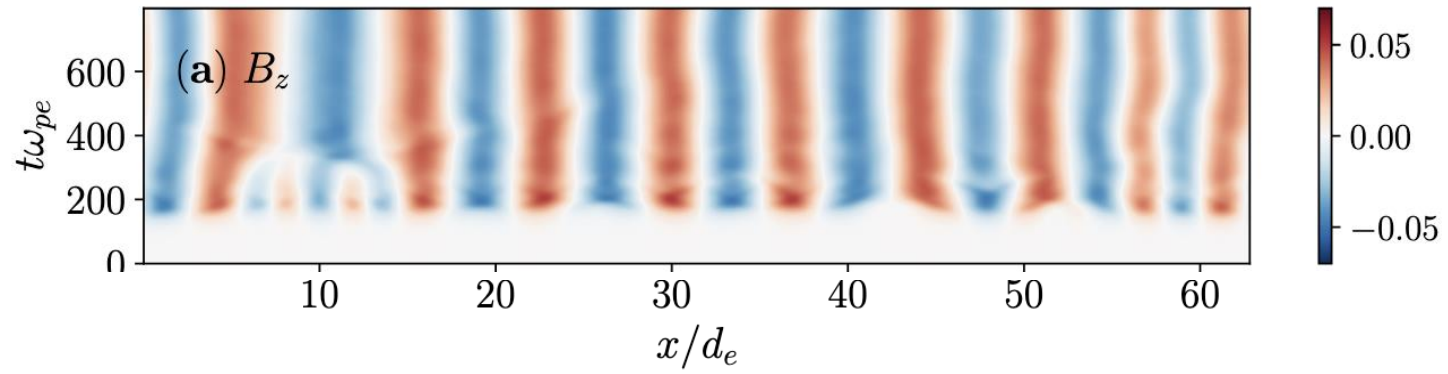
Weibel Instability

- Particle-in-cell simulations of the Weibel instability

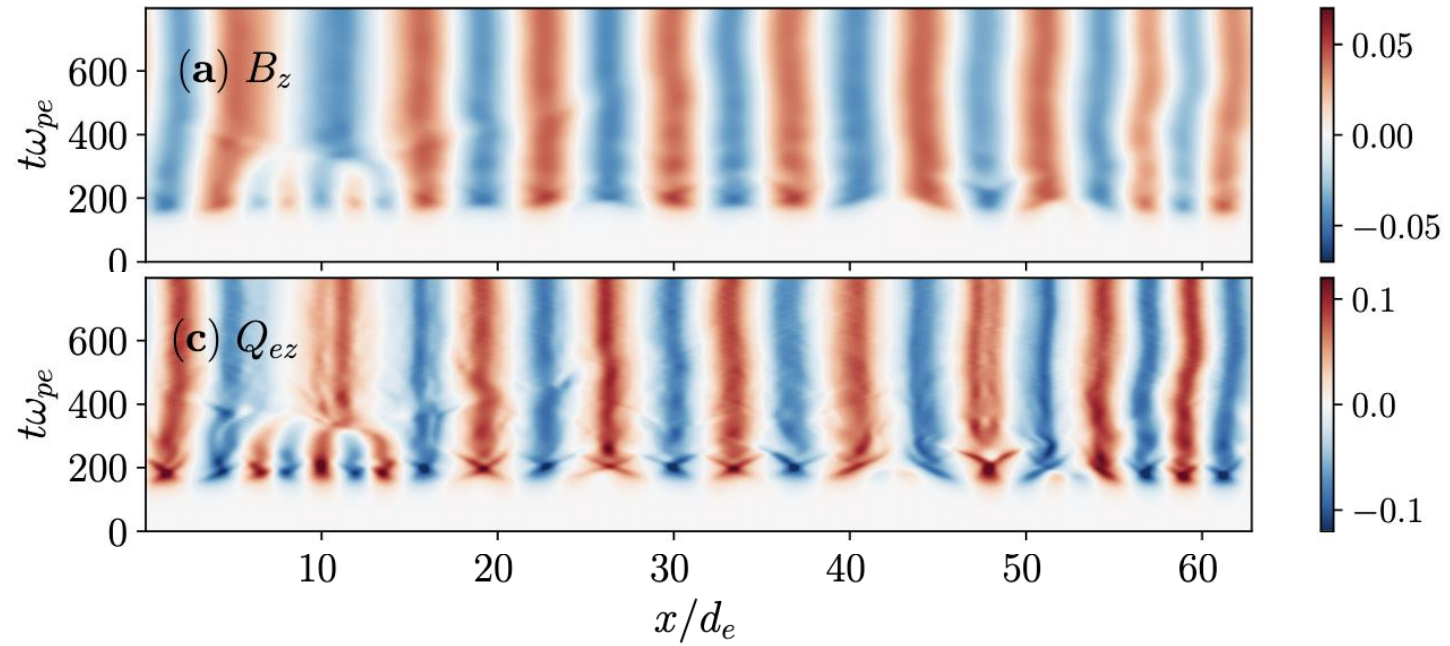
Initially, $p_{yy} \gg p_{xx}$



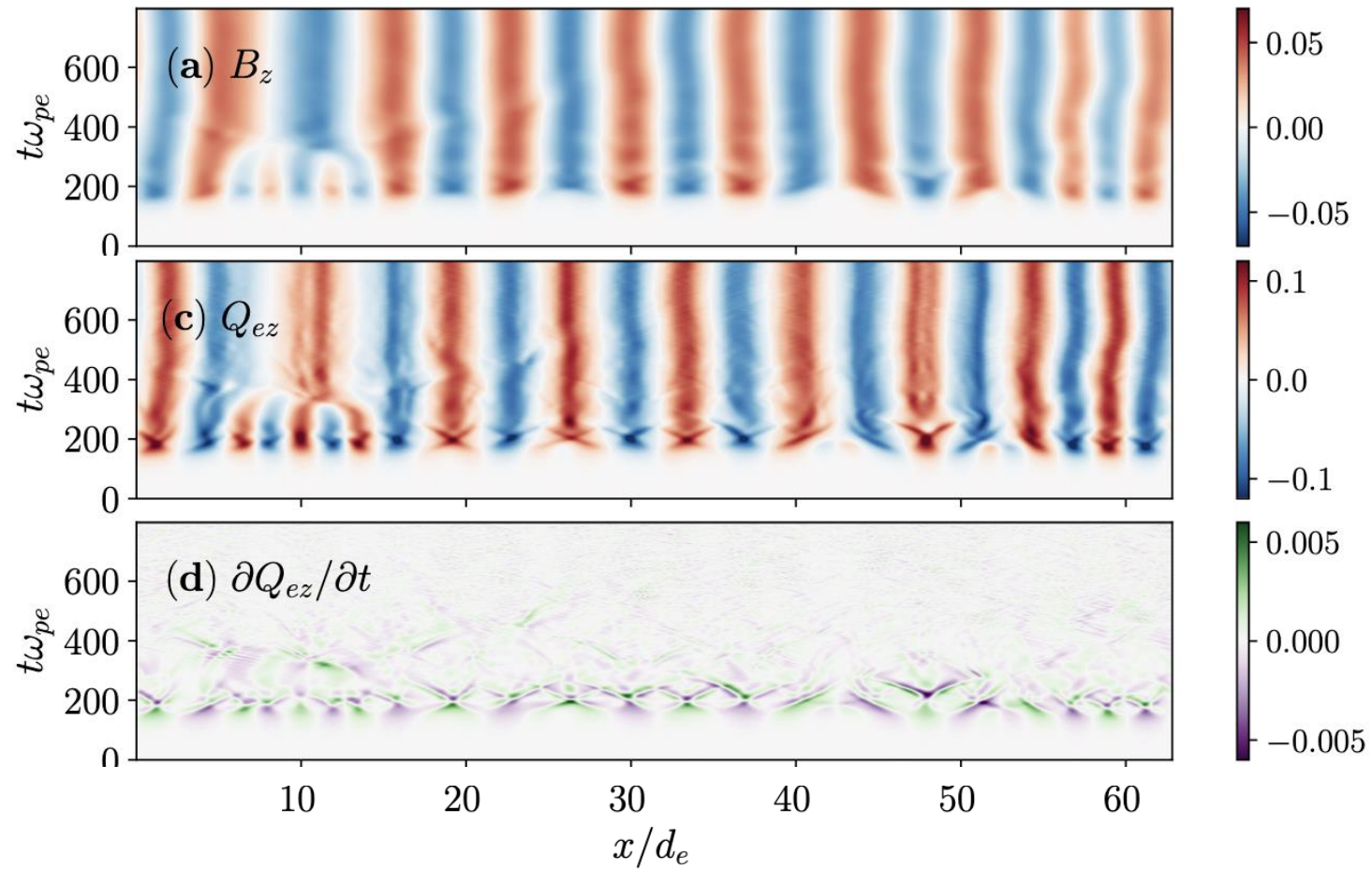
Weibel Instability



Weibel Instability

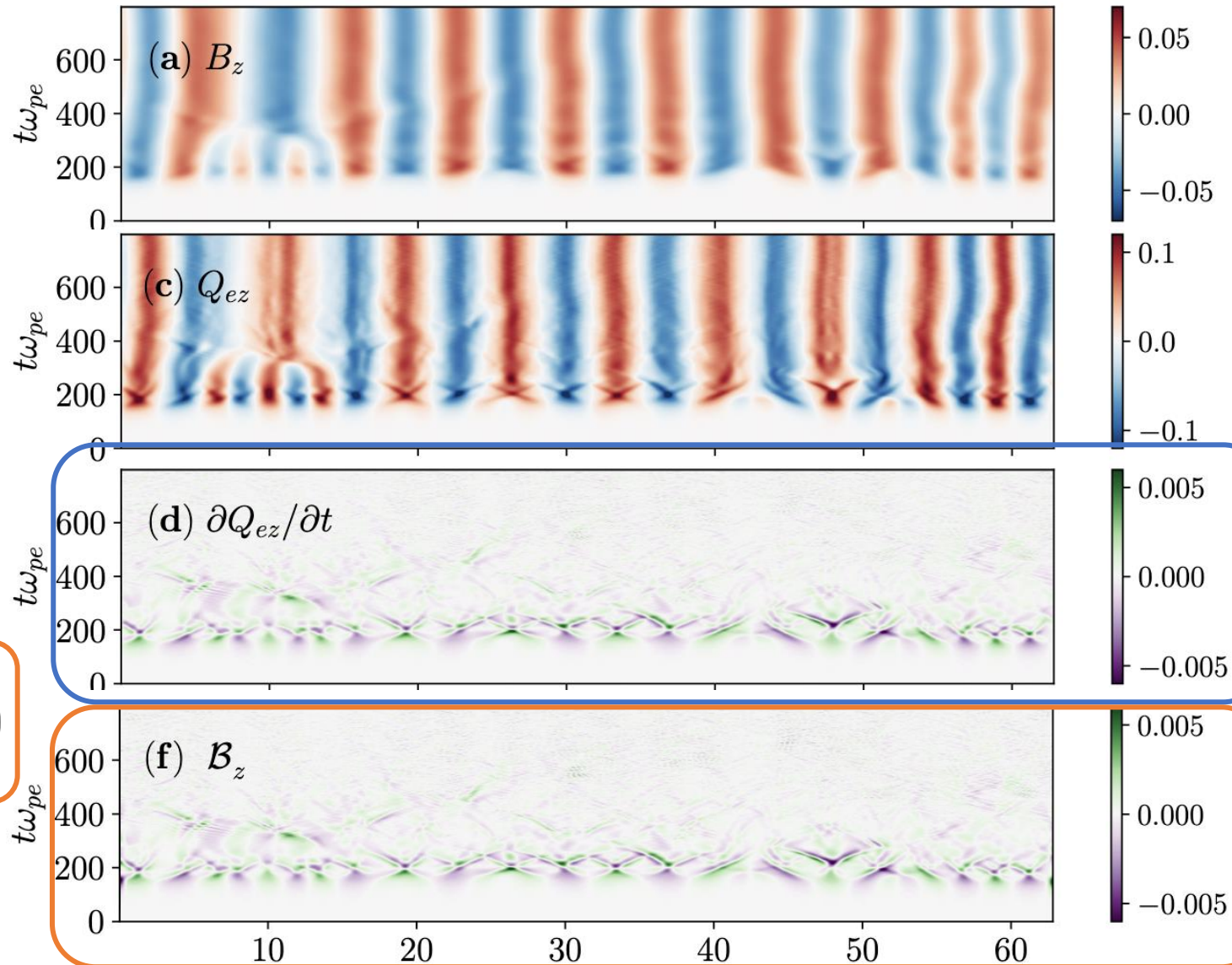


Weibel Instability



Weibel Instability

$$\frac{\partial Q}{\partial t} = -\nabla \times \left(\frac{\nabla \cdot \vec{p}}{n} \right)$$



Canonical battery is entirely responsible for magnetogenesis

New Mechanisms?

- Canonical battery generalizes Biermann battery ✓
- Canonical battery generalizes Weibel instability ✓
- Can there be other mechanisms?
- Assume 2D ($\frac{\partial}{\partial z} = 0$) and examine the z -component

$$\begin{aligned} \mathcal{B}_z &= - \left(\nabla \times \frac{\nabla \cdot \mathbf{p}_e}{n_e} \right) \cdot \hat{z} \\ &= \hat{z} \cdot \left[-\nabla \left(\frac{1}{n_e} \right) \times \nabla \cdot \mathbf{p}_e \right] \\ &\quad \text{Biermann-like term} \end{aligned}$$

New Mechanisms?

- Canonical battery generalizes Biermann battery ✓
- Canonical battery generalizes Weibel instability ✓
- Can there be other mechanisms?
- Assume 2D ($\frac{\partial}{\partial z} = 0$) and examine the z -component

$$\begin{aligned} \mathcal{B}_z &= - \left(\nabla \times \frac{\nabla \cdot \mathbf{p}_e}{n_e} \right) \cdot \hat{z} \\ &= \hat{z} \cdot \left[\underbrace{-\nabla \left(\frac{1}{n_e} \right) \times \nabla \cdot \mathbf{p}_e}_{\text{Biermann-like term}} \right] + \underbrace{\frac{1}{n_e} \left(\frac{\partial^2}{\partial y^2} - \frac{\partial^2}{\partial x^2} \right)}_{\text{Weibel term}} p_{exy} \end{aligned}$$

New Mechanisms?

- Canonical battery generalizes Biermann battery ✓
- Canonical battery generalizes Weibel instability ✓
- Can there be other mechanisms?
- Assume 2D ($\frac{\partial}{\partial z} = 0$) and examine the z-component

$$\begin{aligned} \mathcal{B}_z &= - \left(\nabla \times \frac{\nabla \cdot \mathbf{p}_e}{n_e} \right) \cdot \hat{z} \\ &= \hat{z} \cdot \left[\underbrace{-\nabla \left(\frac{1}{n_e} \right) \times \nabla \cdot \mathbf{p}_e}_{\text{Biermann-like term}} \right] + \underbrace{\frac{1}{n_e} \left(\frac{\partial^2}{\partial y^2} - \frac{\partial^2}{\partial x^2} \right)}_{\text{Weibel term}} p_{exy} + \underbrace{\frac{1}{n_e} \frac{\partial^2}{\partial x \partial y} (p_{exx} - p_{eyy})}_{\text{New term (2D-localized pressure anisotropy)}} \end{aligned}$$

- Last term

$$\mathcal{B}_z = \frac{1}{n_e} \frac{\partial^2}{\partial x \partial y} (p_{exx} - p_{eyy})$$

- Assume

$$p_{exx} - p_{eyy} \sim \exp\left(-\frac{x^2 + y^2}{\sigma^2}\right)$$

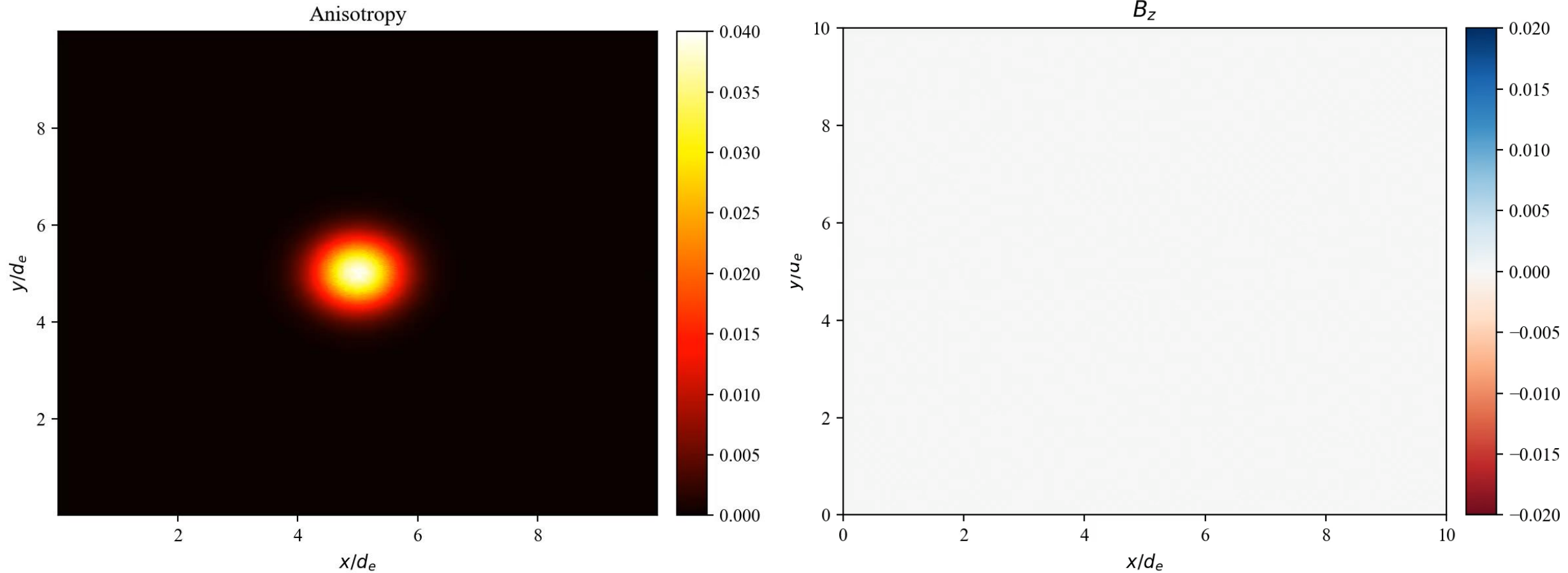
- Then

$$\mathcal{B}_z \sim xy$$

- Quadrupole magnetic fields expected

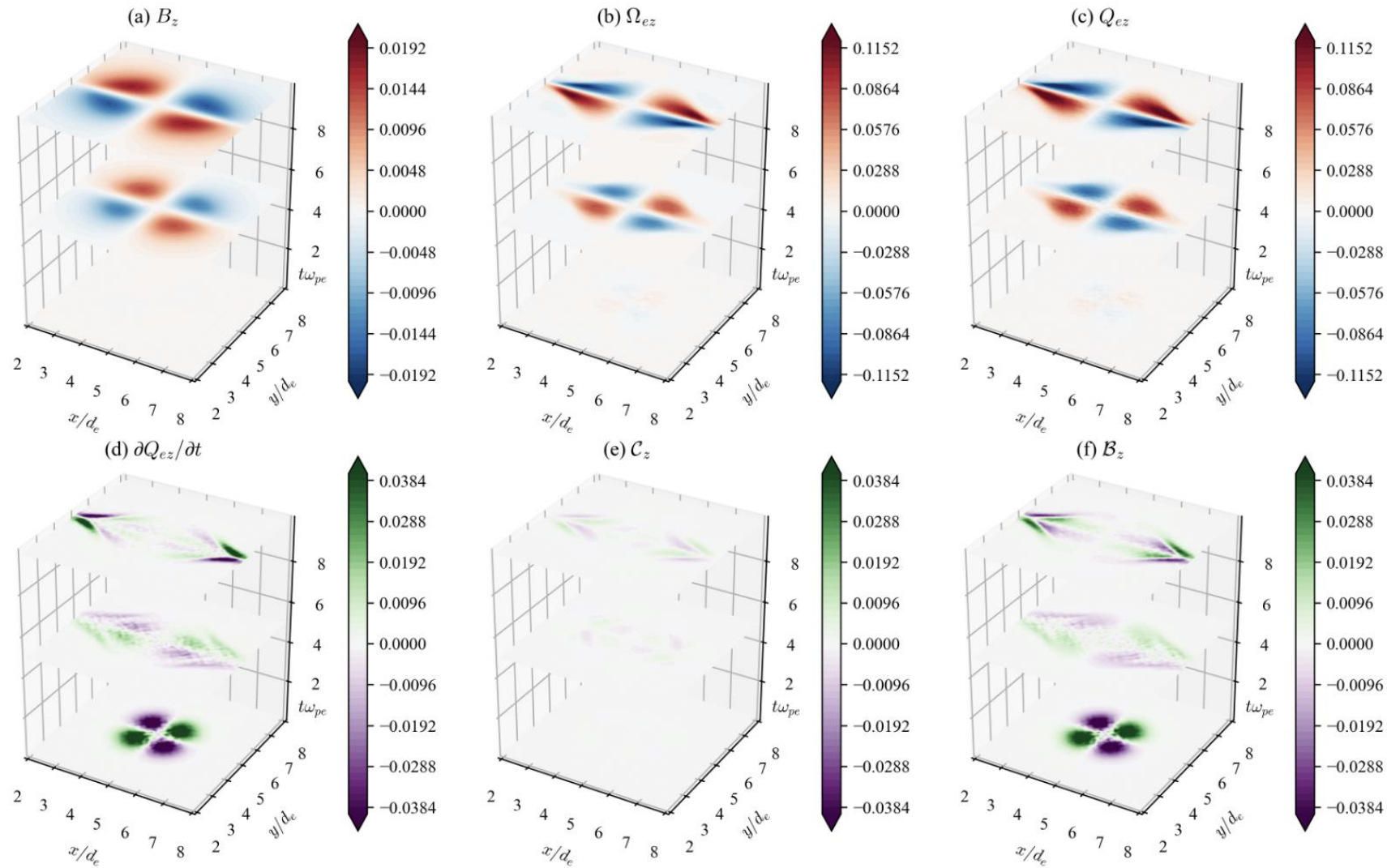
2D-localized Pressure Anisotropy

- Again, verify with particle-in-cell simulations



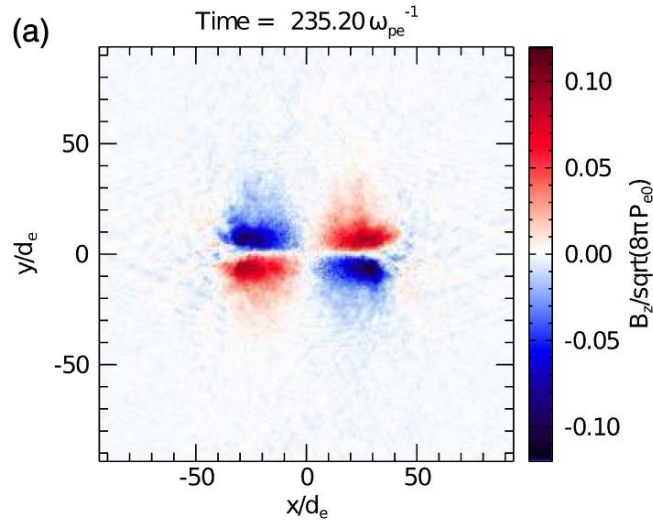
2D-localized Pressure Anisotropy

- Again, canonical battery term is confirmed to be the source!



Usage of the battery term

- Usage 1: Relative Importance among mechanisms



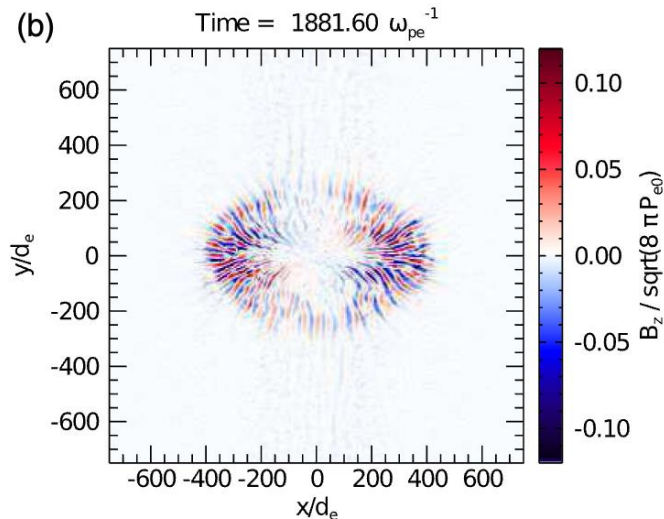
$\frac{L}{d_e} < 100$: Biermann

Biermann

Weibel

$$\left| \frac{\nabla n_e \times \nabla T_e}{n_e} \right| : \frac{1}{n_e} \frac{\partial^2 p_{exy}}{\partial x^2},$$

$$\frac{T_{diag}}{T_{off}} d_e^2 : L^2,$$



$\frac{L}{d_e} > 100$: Weibel

Typically, $T_{diag} \gg T_{off}$, so Weibel is important for $L \gg d_e$

- Usage 2: Predict other mechanisms
 - Localized density and pressure anisotropy

$$\mathcal{B}_z \sim \left(\frac{\partial}{\partial y} \frac{1}{n_e} \right) \left(\frac{\partial p_{exx}}{\partial x} \right)$$

- Localized density and uniform temperature anisotropy

$$\mathcal{B}_z \sim T_{exx} \left(\frac{\partial}{\partial y} \frac{1}{n_e} \right) \left(\frac{\partial n_e}{\partial x} \right)$$

Relativistic Battery

- In the relativistic regime, the canonical induction equation includes a term purely due to relativity

$$\frac{\partial \mathbf{Q}}{\partial t} = \nabla \times (\mathbf{u} \times \mathbf{Q}) - \nabla \times \left(\frac{\nabla \cdot \vec{\mathbf{p}}}{n} \right) + \sum_{i=x,y,z} [p_i, u_i] \quad \text{or } \epsilon_{ijk} \partial_i p_l \partial_j u_l \text{ or } \nabla p_l \times \nabla u_l$$

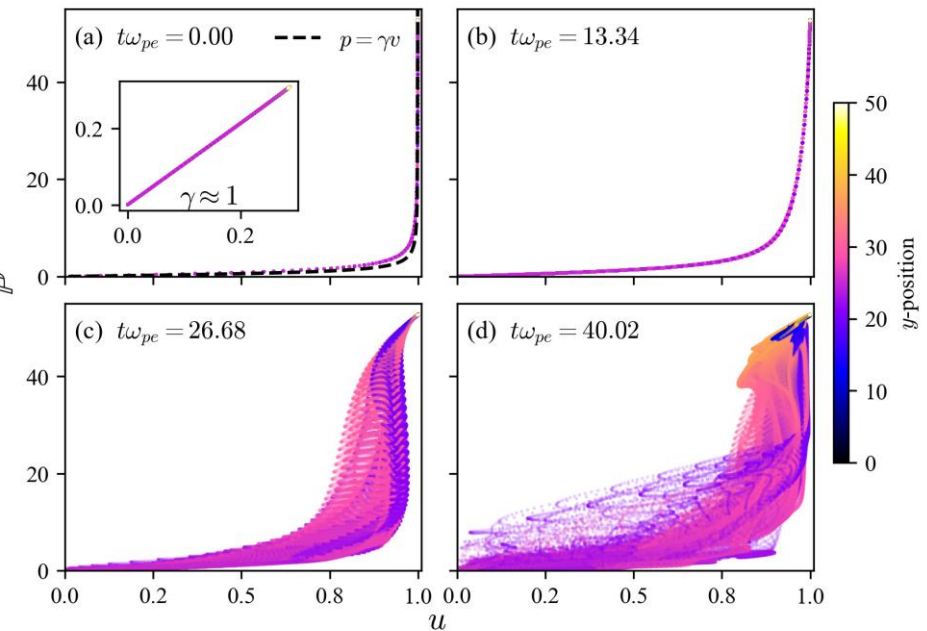
- For a single-particle, velocity and momentum commute

$$\mathbf{p} = \mathbf{p}(\mathbf{v}) = \frac{\mathbf{v}}{\sqrt{1 - v^2/c^2}}$$

- For a relativistic fluid, this is not generally the case

$$\langle \mathbf{p} \rangle \propto \int \mathbf{p} f d^3 \mathbf{p}$$

$$\mathbf{u} \propto \int \frac{\mathbf{p}}{\gamma} f d^3 \mathbf{p} = \int \frac{\mathbf{p}}{\sqrt{1 + p^2}} f d^3 \mathbf{p}$$



- Other non-ideal effects give rise to other terms

$$\frac{\partial \mathbf{Q}}{\partial t} = \nabla \times (\mathbf{u} \times \mathbf{Q}) - \nabla \times \left(\frac{\nabla \cdot \vec{\mathbf{p}}}{n} \right) + \hat{z} \sum_{i=x,y} [u_i, p_i] + \nu(\boldsymbol{\Omega} - \boldsymbol{\Omega}_\sigma) + QM + Rad$$

Relativity

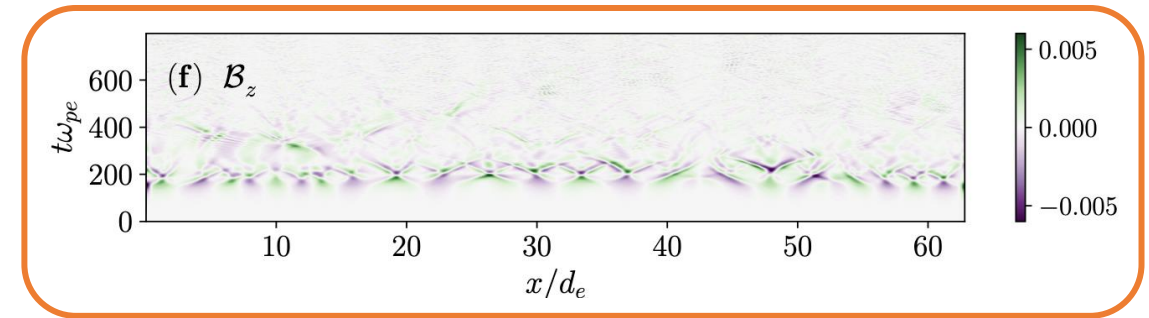
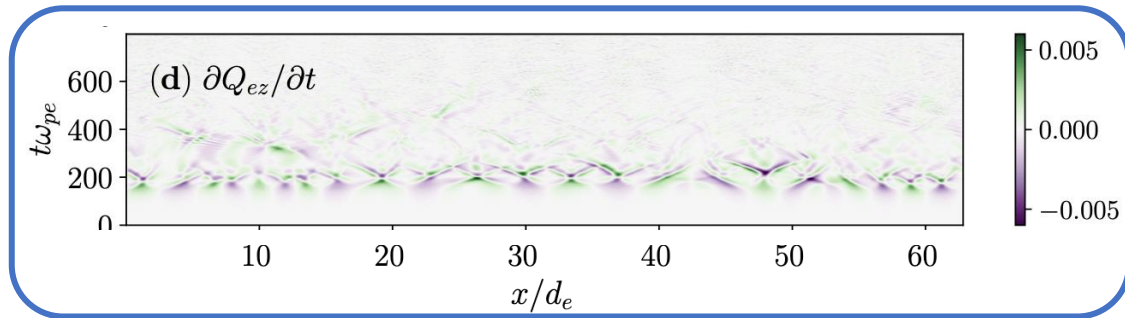
Inter-species collisions

QED effects

Radiative effects

- The pressure tensor in the plasma equation of motion turns into the canonical battery term
- This term is solely responsible for spontaneously generating canonical vorticity, which is a proxy for the magnetic field
- This term generalizes popular magnetogenesis mechanisms and predicts new ones
- Other effects may provide additional sources

$$\frac{\partial Q}{\partial t} = -\nabla \times \left(\frac{\nabla \cdot \vec{p}}{n} \right)$$



Pressure term is entirely responsible for Weibel instability ✓

The canonical induction equation generalizes both 70-year-old magnetogenesis mechanisms (and predicts new ones)